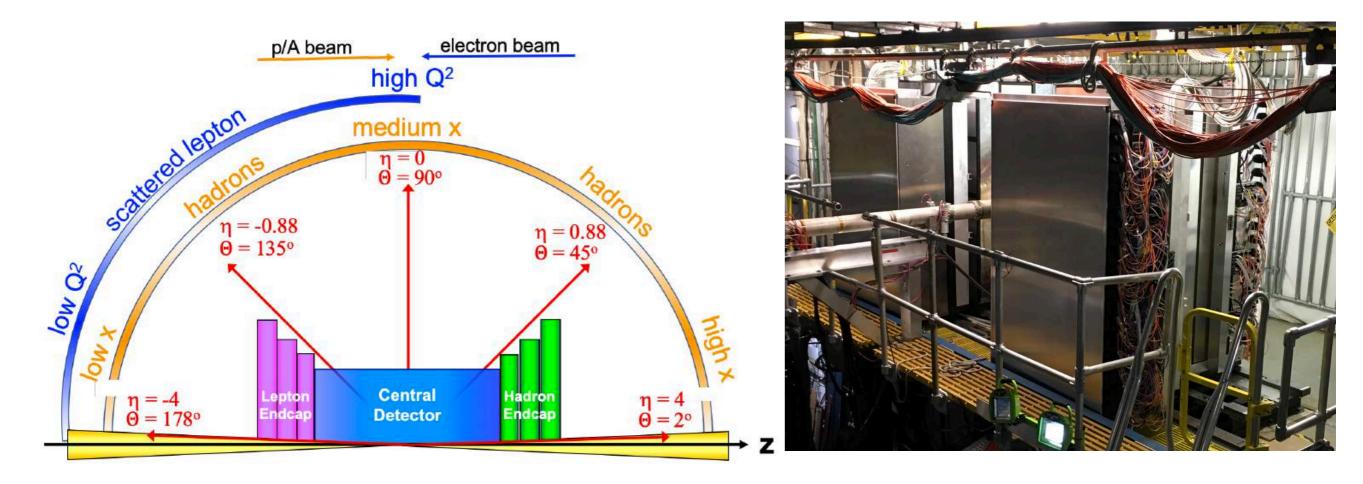
Hadronic Calorimetry. O.Tsai (UCLA)

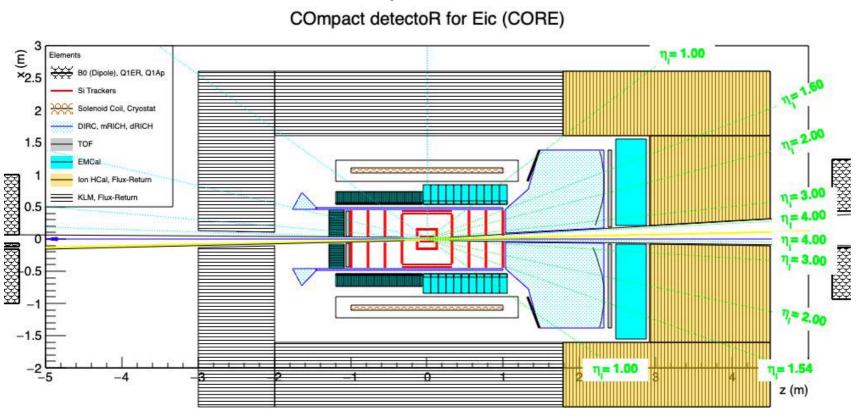


- eRD1/STAR Optimization of forward calorimeters system for EIC reference detector.
- Construction of STAR Forward Calorimeter System.

Synergy between STAR FCS and EIC Calorimetry R&D lead to:

- development of EIC reference detector concept and technologies.
- helped to ensure these technologies are now well established within EIC user community.

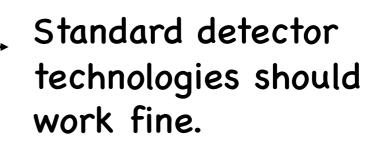
Hadron Side EndCap.

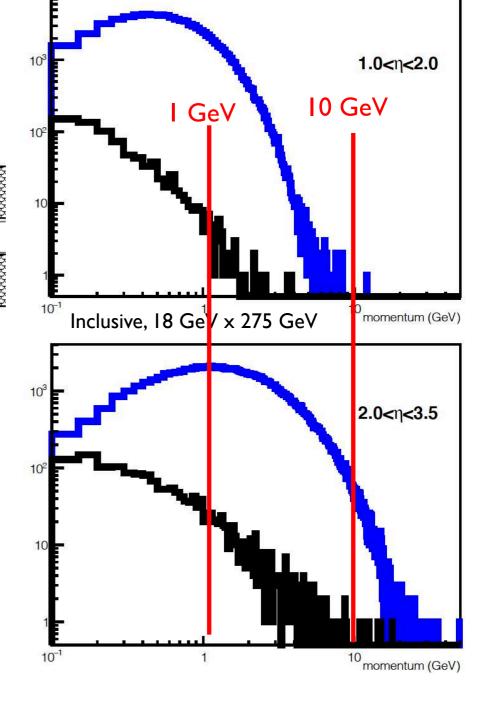


- Requirements in YR, resolution 50%/√E + 10%. 6% constant term for eta>3 is desired
- Desired as good as possible 35%//E + X%. (N.B. there is no discussion in YR text to support these numbers.)
- Requires outstanding Hcal/Ecal system to achieve this.

Conditions at EIC Hadron EndCap:

- Particles Energy low, difficult for calorimeters
- Interaction Rate low, < 500kHz
- Occupancy low
- Radiation Exposure low
- Neutron Fluxes some concern.
- Beam Pipe Hole limit acceptance to eta ~ 3, hadron showers are wide, will leak into beampipe up to 40% lost at high eta (<u>https://indico.phy.ornl.gov/event/38/</u> talk by F. Bock)

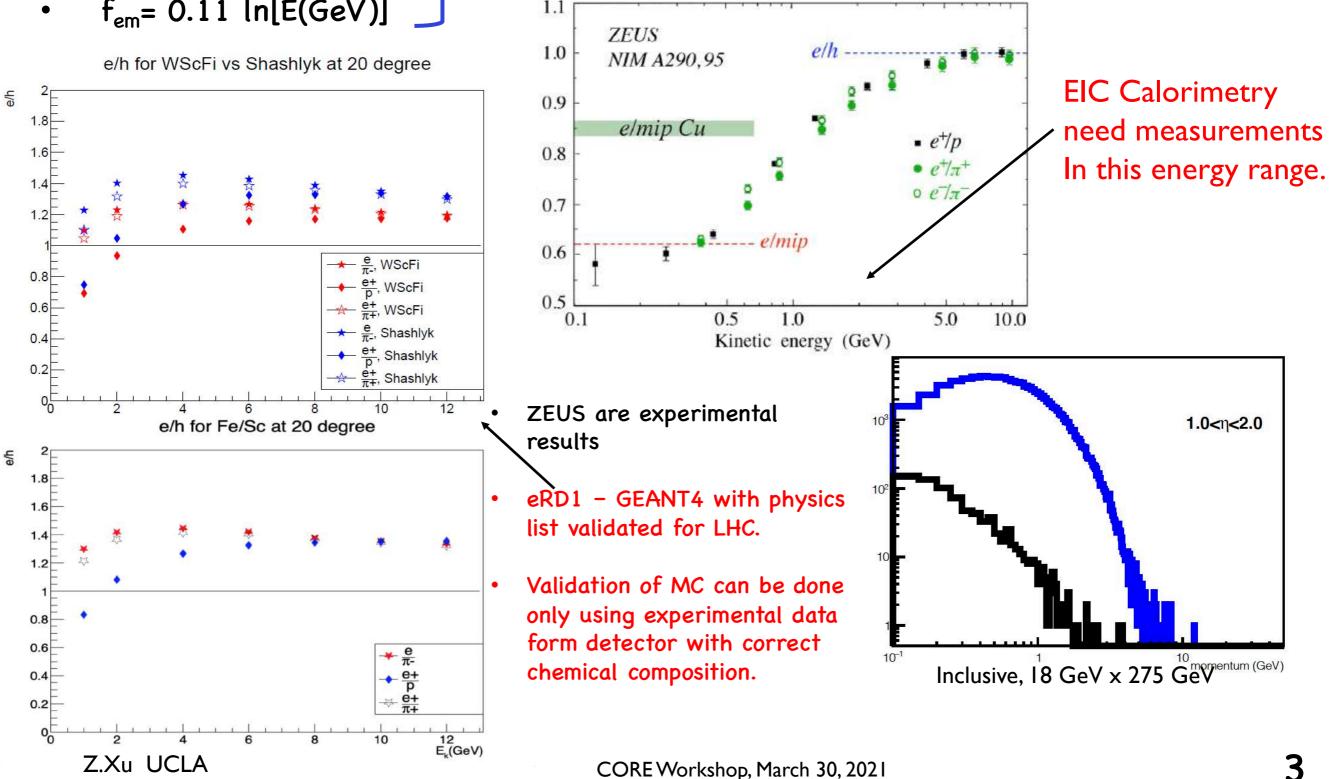




Hadron calorimeter systems. EIC Challenges.

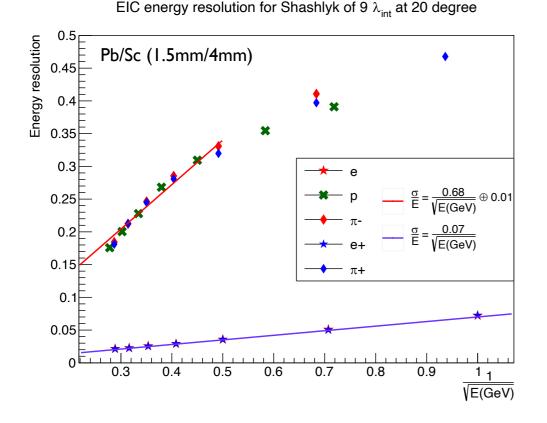
- **e/h** ≠ 1
- $e/h_ecal \neq e/h_hcal$
- e/h = f(E)
- $e/p \neq e/\pi$
- $f_{em} = 0.11 \ln[E(GeV)]$

Jet energy resolution is always poorer than for a single hadron. Despite ~ 20% of jet energy (em) measured very accurately by Ecal.

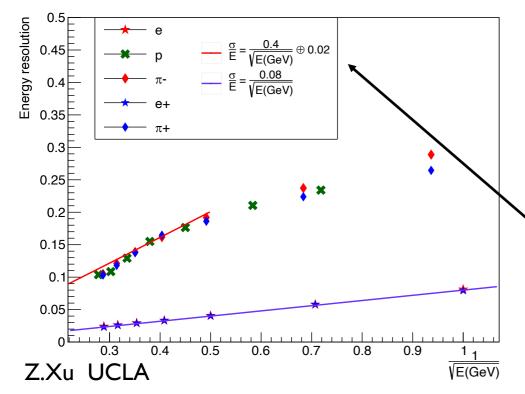


How important to tune e/h value? Hypothetical Configurations.

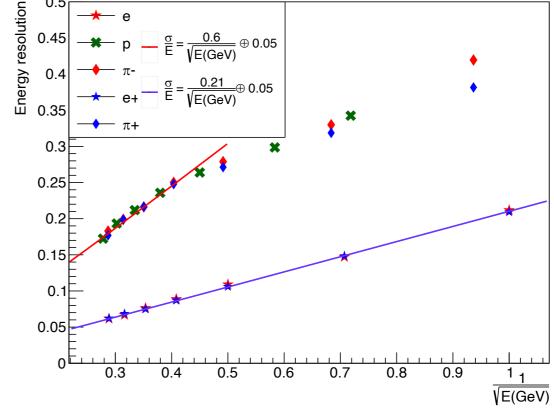
0.5



EIC energy resolution for W/ScFi of 9 λ_{int} at 20 degree



EIC energy resolution for Fe/Sc 20/3mm of 9 λ_{int} at 20 degree

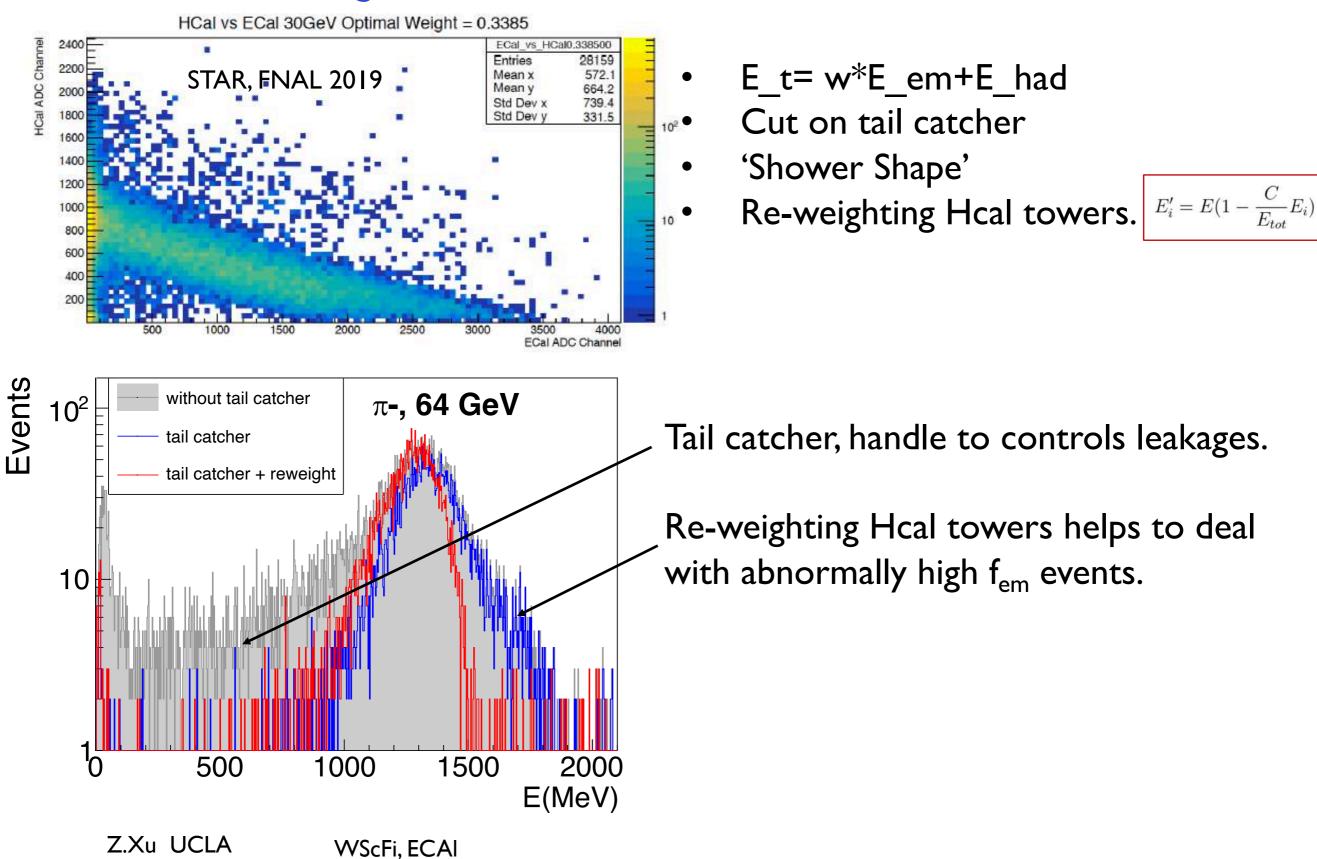


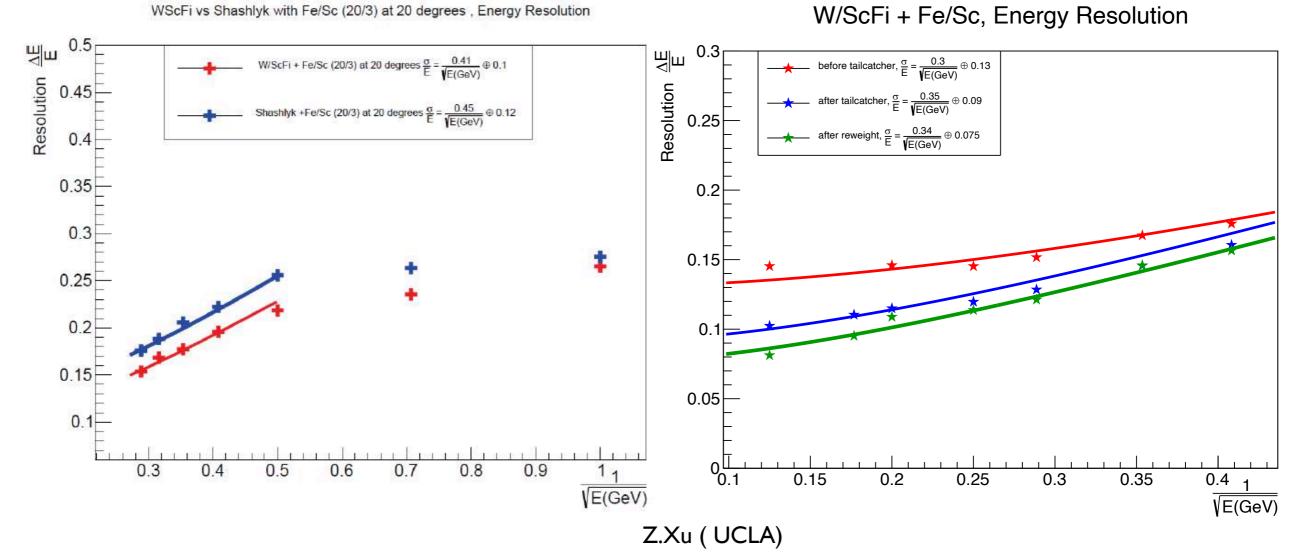
Hypothetical variant, 9 interaction lengths long calorimeters. Same structure for Ecal and Hcal sections. Three different technologies:

- SHASHLYK (Phenix, STAR Forward)
- WScFi (STAR Forward 2014)
- Fe/Sc (STAR Forward 2020)

Proper detector composition required for good hadronic resolution. I.e. desired to keep e/h as close as practically possible to 1. N.B. these are MC not an experimental results.

Realistic Configurations, i.e. binary systems Em + HAD.





At lower energies (EndCap eta range I-2) stochastic term almost always will dominate.

N.B. no cuts on tail-catcher or re-weighting was applied here.

WScFI 23 X0 vs SHASHLYK 18X0 (both depth and better e/h plays role) At higher energies (EndCap eta range 2-3) constant term start to dominate.

Cut on tail catcher + re-weighting Hcal towers

With cut on tail catcher and reweighting Hcal towers GEANT4 resolution looks very good with stochastic term at ~35% and constant term at ~7% (N.B. efficiency, fit).

Does deeper Ecal helps to reconstruct low energy hadrons?

N N 10³ π-@ 1 GeV √ π- @ 8 GeV @ 2 GeV -EM C 10 π-@ 10 GeV HCe 0.7 - π- @ 4 GeV 10 × π-@ 12 GeV _____ π- @ 6 GeV 0.6 λ_m+4λ GeV 2 10³ 10² 0.5 10 0.4 100 0.5λ 100 λ_144 GeV 4 160 10³ 0.3 10² 10 0.2 0.57 +4.57 10³ 0.1 10² 0.5 0.6 8.0 0.7 0.9 λ_{WScFi} Z.Xu UCLA 10 200 0.77 SOO EIC energy resolution for different interaction length of W/ScFi 200 300 0.57, 44.57 GeV 8 0.67 300 0.83.44.20 10³ 0.35 resolution @ 0.51 \ 10² ⊕ 0.09 10 E(GeV) 0.35 0.09 energy 0.3 VE(GeV) 0.72 +4.3 200 10³ π - @ 1.02 10² 0.25 10 Increased depth of ECal 0.2 Increased depth of Ecal does help a bit to improve energy 0.15 resolution of the system for low energy hadrons. Assuming we 0 0.3 0.8 have good PID, additional e/h (TRD etc.,) you can do that. 0.4 0.5 0.6 07 0.9 11

energy ratio of WScFi vs percentage of interaction length

E(GeV)

What about advanced HCals ? (no PFA, calorimeters only).

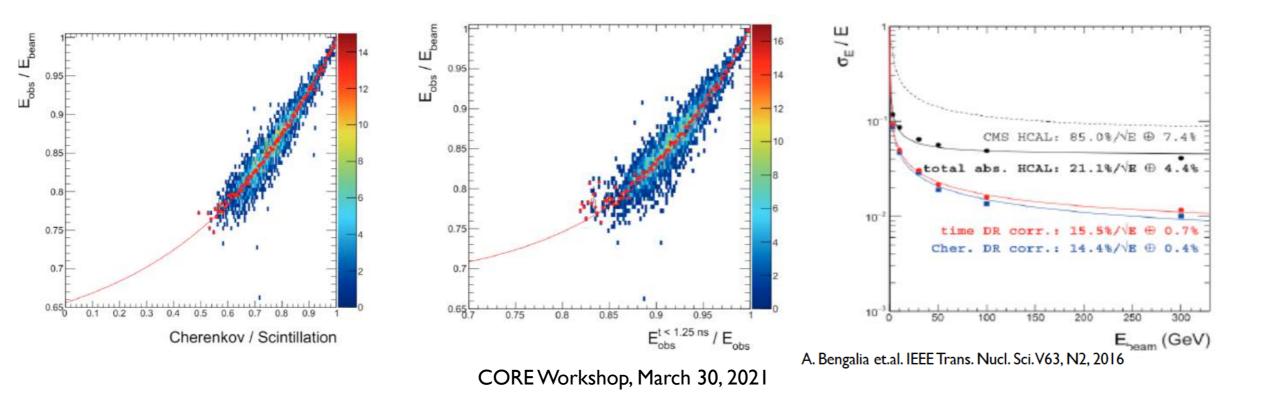
Compensated detectors. Record holder is SPACAL ~ 30%/sqrt(E) But, below 10 GeV compensation does not work, and looks like there is no solution (at least known)

- Dual readout methods developed by HEP during past 20 years or so, DREAM project.
- There is no real detector implementations beyond prototypes.
- There are few talks related to DR at latest CPAD <u>https://indico.fnal.gov/event/46746/</u>
- There are also 3D imaging concepts CALICE, IDEA and such (100s M of channels + timing)

What is DR?

- Find observable which correlates with number of neutrons (kinetic energy of n correlates with `invisible' energy). Observables can be C/S, Time, Spatial characteristics of shower etc. Unlike compensation dose not require precise tuning of chemical composition of detector.
- E-by-E correct detected energy using this observable. Or 3D imaging + ML

Theoretically hadronic resolution can be as good as 20%/sqrt(E) (Fe/Sc (20/3)), i.e. em type resolution On practice very difficult to realize and costly.



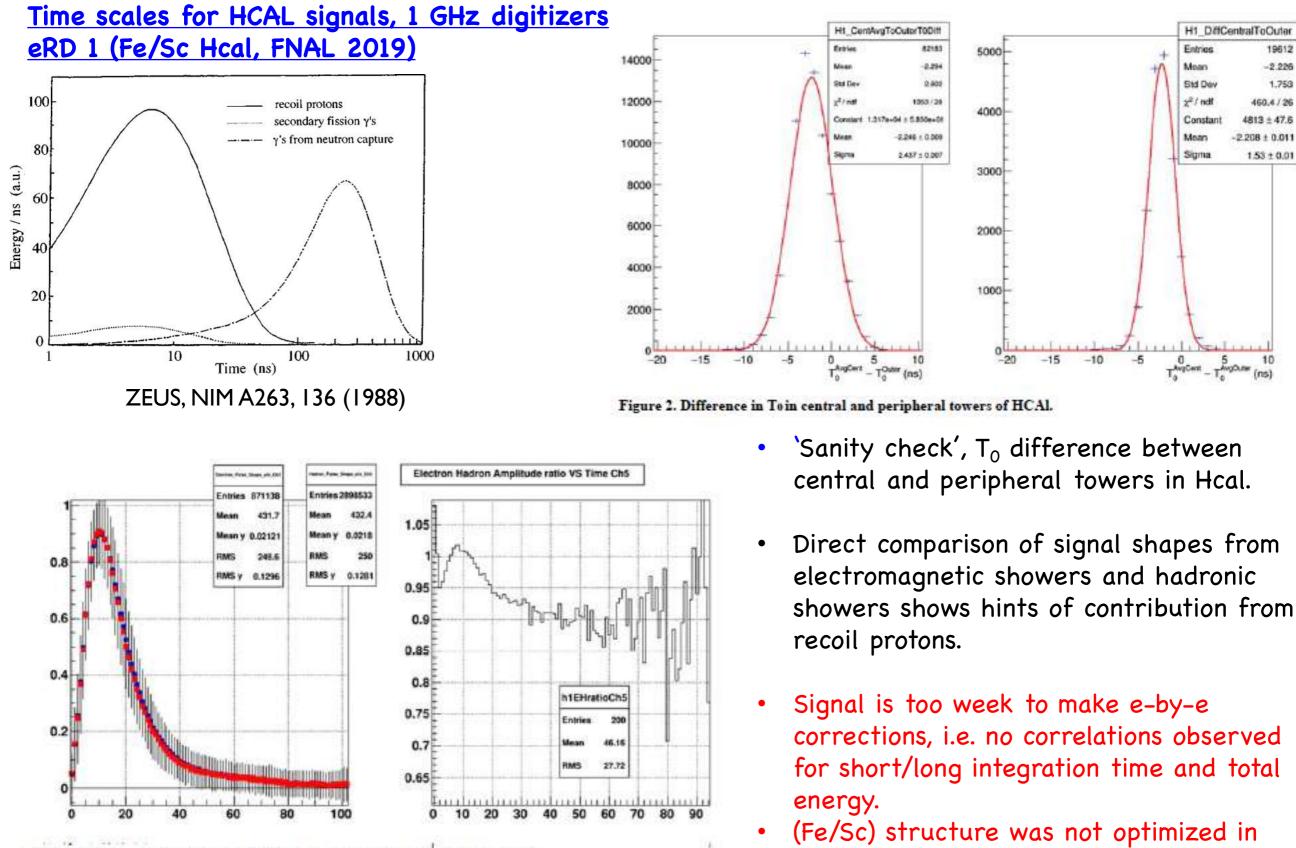


Figure 3. Shape of signals for 20 GeV electrons and pions. X axis bin size 1 ns.

any sense for such purposes.

dependence.

by A. Kiselev – found very weak

Pb/Sc structure (17mm/6mm) was studied

IMHO, realization of DR methods at EIC will be difficult.

- Timing will not work for Tile/WLS well no matter what type of absorber will be used (eRD1 conclusion after 2019 tests)
- Timing may work with SPACAL type, no MC tried. Need high Z to generate enough neutrons, but then had to think about flux return. Too many neutrons may be problematic for SiPMs. There are some problems with fiber type DREAM detectors (see latest CPAD, Tile configuration DREAM approach).
- Shower shape may be difficult due to relatively low energy of hadrons. And it is not clear how well it will work in compact detector at forward rapidities for jets. Magic with ML, AI ③, may be?
- C/S may work with SPACAL types. HCal is a flux return, had to be made from steel.

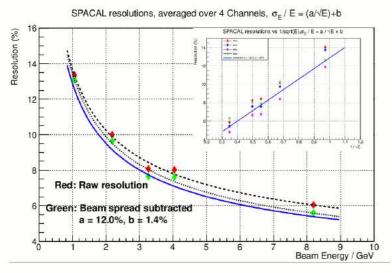
Practical realization of any DR approaches are challenging: fast digitizers, C/S separation, high granularity, integration issues...

With CD2 in Jan 2023, they are out of reach. For long term upgrades it is possible they may be worked out (SciGlass ?)

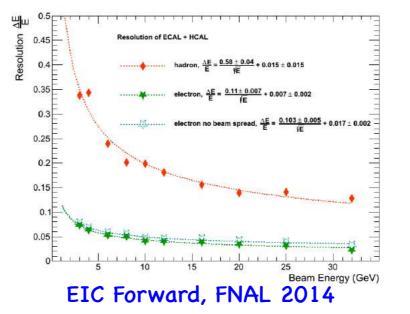
<u>Reference EIC Hadron EndCap studies summary (eRD1 reports 2019–2020):</u>

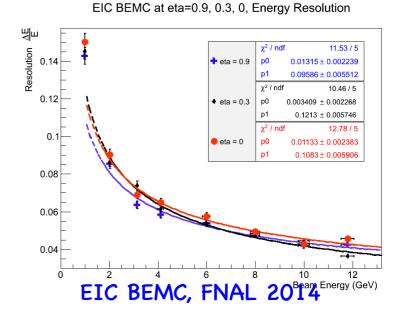
- There are some tricks which may help improve resolution of reference hadron endcap system, may be by 20%, without escalating the cost.
- It is little to no room to improve e/h for Fe/Sc section.
- WScFi, e/h good as it is (may be improved a little, but need experimental data).
- Tail catcher will allow to control leakages from the back. (Easy to integrate).
- Dead material between Ecal and Hcal is not an issue, because it is not needed.
- Different Instrumental effects, like light collection non-uniformities in hcal section has little effect on resolution, checked with gSTAR.
- With cut on tail catcher and re-weighting Hcal towers GEANT4 resolution looks very good with stochastic term at ~35% and constant term at ~7% (N.B. efficiency, fit).
- Need to think a bit more about increasing depth of Ecal that may be important for Barrel, due to magnet coils between Ecal and Hcal.

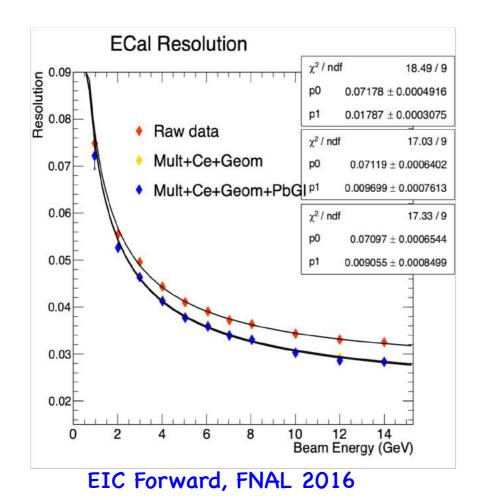
Transition to targeted R&D...

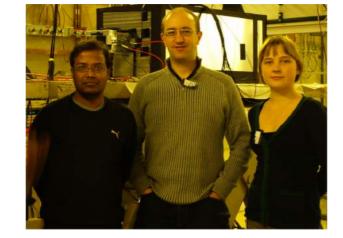


Proof of principle. FNAL 2012

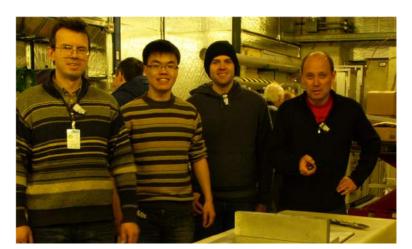














Test Runs 2012 -2016

STAR Forward Calorimeter System (FCS), 2020



STAR Collaborators, Members of UC EIC Consortia Assembling FCS in Dec. 2020, BNL

Large group of STAR collaborators actively engaged in all aspects of the project: ACU, BNL, UCLA, UCR, Indiana University CEEM, UKU, OSU, Rutgers U., Temple U., Texas A&M U., Valparaiso U.

Forward Calorimeter System (FCS)

- ECal 1496 channels ~ 8 tons
- HCal 520 channels ~ 30 tons.
- SiPM Readout Bias ~ 67V
- New digitizers + Trigger FPGA = DEP_{boards}





Detector R&D Advisory Committee Meeting. BNL, March 25, 2021

From generic R&D and YR to targeted R&D.

- Technologies for WScFi and Fe/Sc (construction method) are well established and spread in community (STAR and sPHENIX). Developed during generic EIC detector R&D.
- Performance of reference detector Hadron EndCap is very good on paper. Well exceed requirements of YR.

What we need to do before CD2 (Jan. 2023)?

A full scale prototype WScFi + Fe/Sc with transverse size 0.6m x 0.6m, with integrated tail catcher for hadron endcap.
a) HCAL part is IP independent.

b) HCAL part is endcap independent (e or h side)

• A test beam or two (FTBF at FNAL may be OK, BNL A2 will be nice to revive)

Timescale is doable. Construction of prototype will take 1 or 1.5 years, cost ~ \$300k There are few small R&D topics which has to be finished (light collection efficiency and such) these are already funded by EIC generic detector R&D (Funds for FY2020 have not been received yet).

Thanks!