



Recent strangeness results from the RHIC beam energy scan

Xianglei Zhu

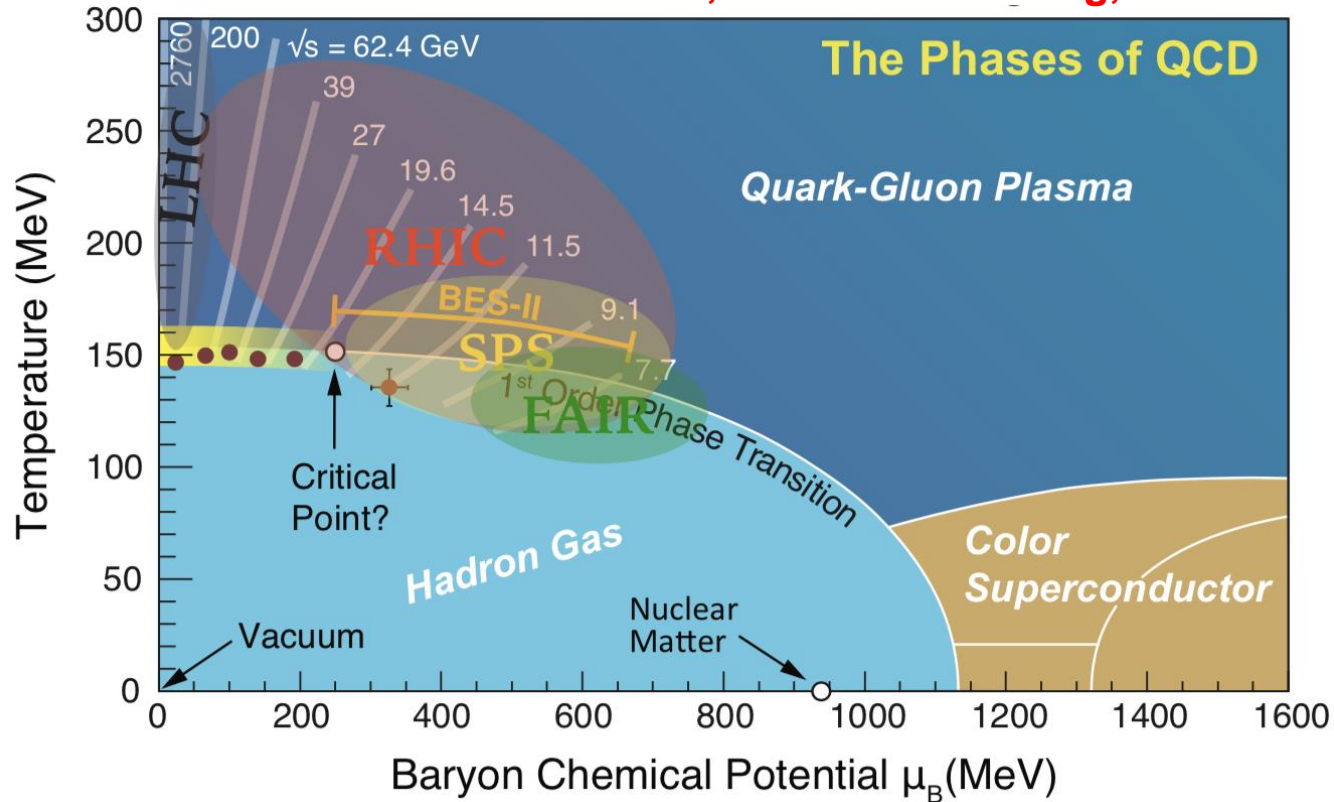
Tsinghua University

5/11/2021

BNL nuclear physics seminar

QCD phase diagram

B. Müller, BEST Col. Meeting, 2016

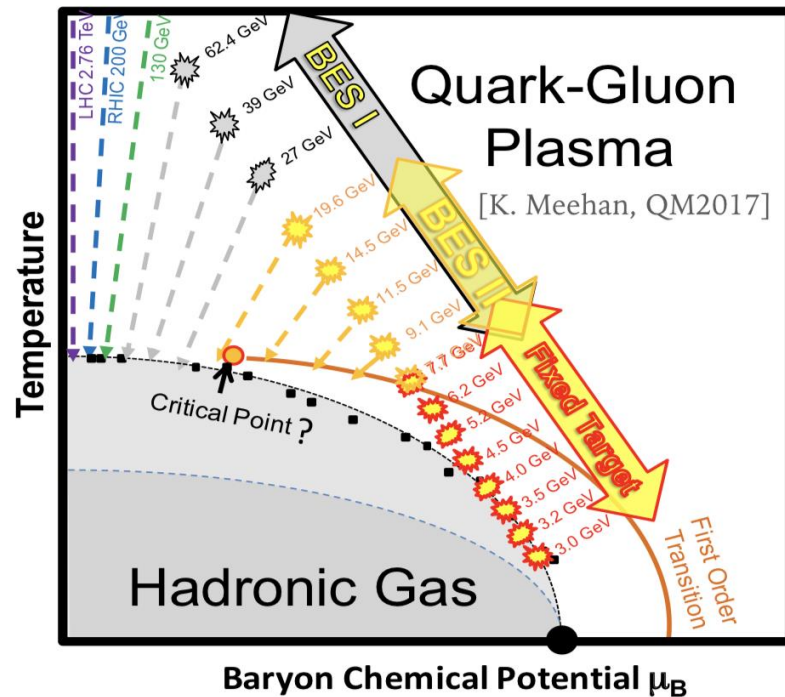
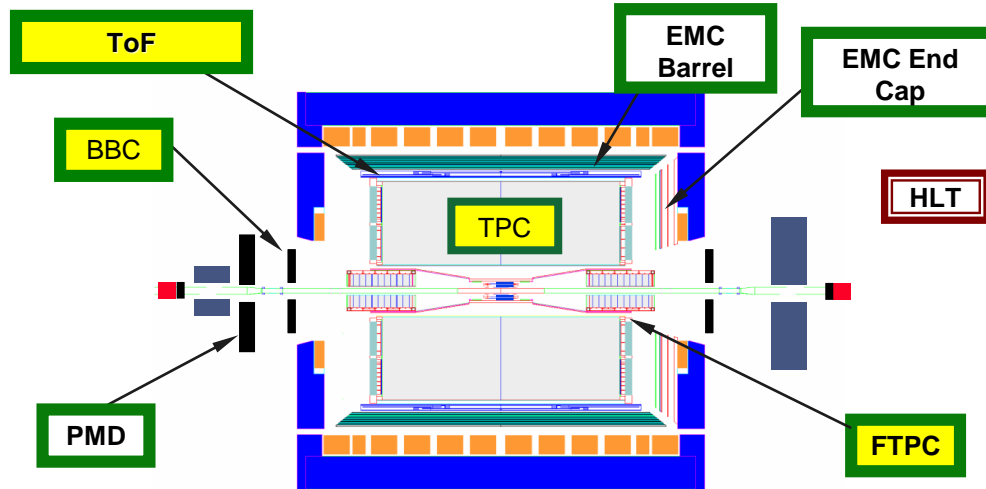


- **RHIC Beam Energy Scan**

Cover the intermediate baryon density region

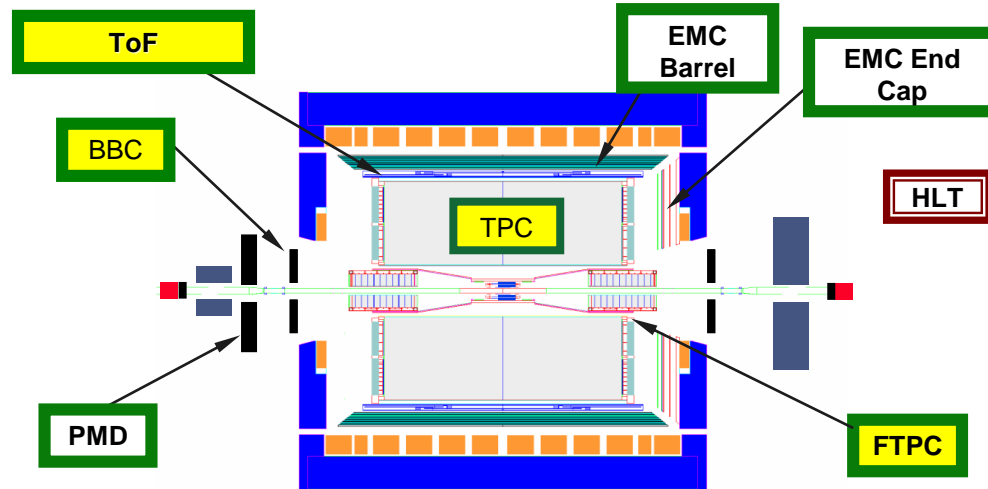
Look for **onset of de-confinement, phase boundary** and critical point

STAR BES



- STAR:
Collider experiment at RHIC
- full azimuthal coverage at mid-rapidity
- BES-I (completed)
 $\text{Au+Au } \sqrt{s_{NN}} = 62.4 - 7.7 \text{ GeV}$
- BES-II (on-going)
 $\text{Au+Au } \sqrt{s_{NN}} = 19.6 - 7.7 \text{ GeV}$
- Fixed-target (on-going)
 $\text{Au+Au } \sqrt{s_{NN}} = 7.7 - 3.0 \text{ GeV}$
 μ_B up to 721 MeV

STAR BES-I datasets



- STAR:
Collider experiment at RHIC
- full azimuthal coverage at mid-rapidity
- BES-I (completed)
 $\text{Au+Au } \sqrt{s_{NN}} = 62.4 - 7.7 \text{ GeV}$

Year	<i>Collisions</i>	$\sqrt{s_{NN}}$ (GeV)	MB events
2010	Au+Au	7.7	~ 4 M
2010	Au+Au	11.5	~ 12 M
2014	Au+Au	14.5	~ 13 M
2011	Au+Au	19.6	~ 36 M
2011 / 2018	Au+Au	27	~ 70 M / ~ 560 M
2010	Au+Au	39	~ 130 M
2017	Au+Au	54.4	~ 556 M
2010	Au+Au	62.4	~ 46 M

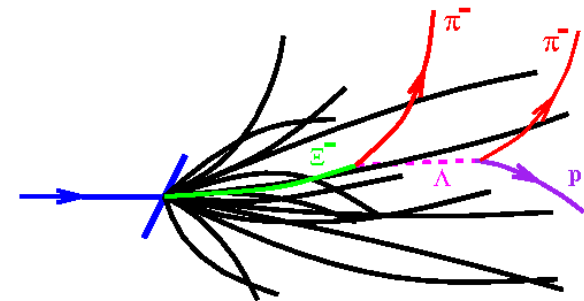
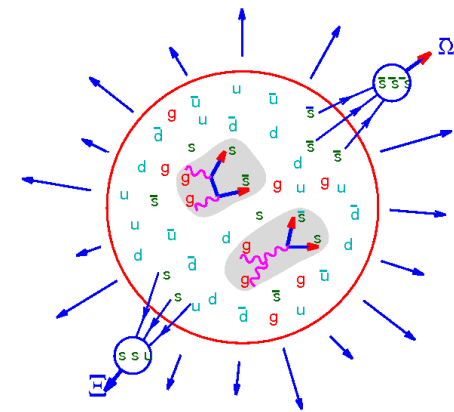
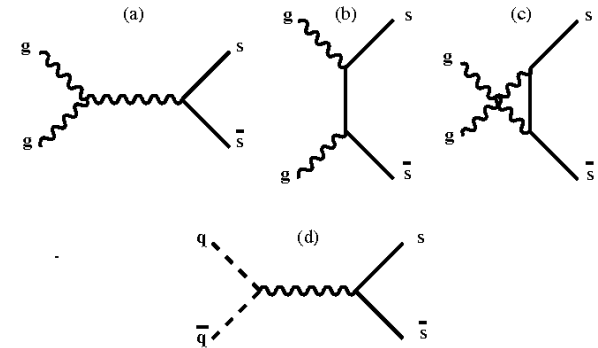
Why strangeness?

Rafelski & Müller, 1982

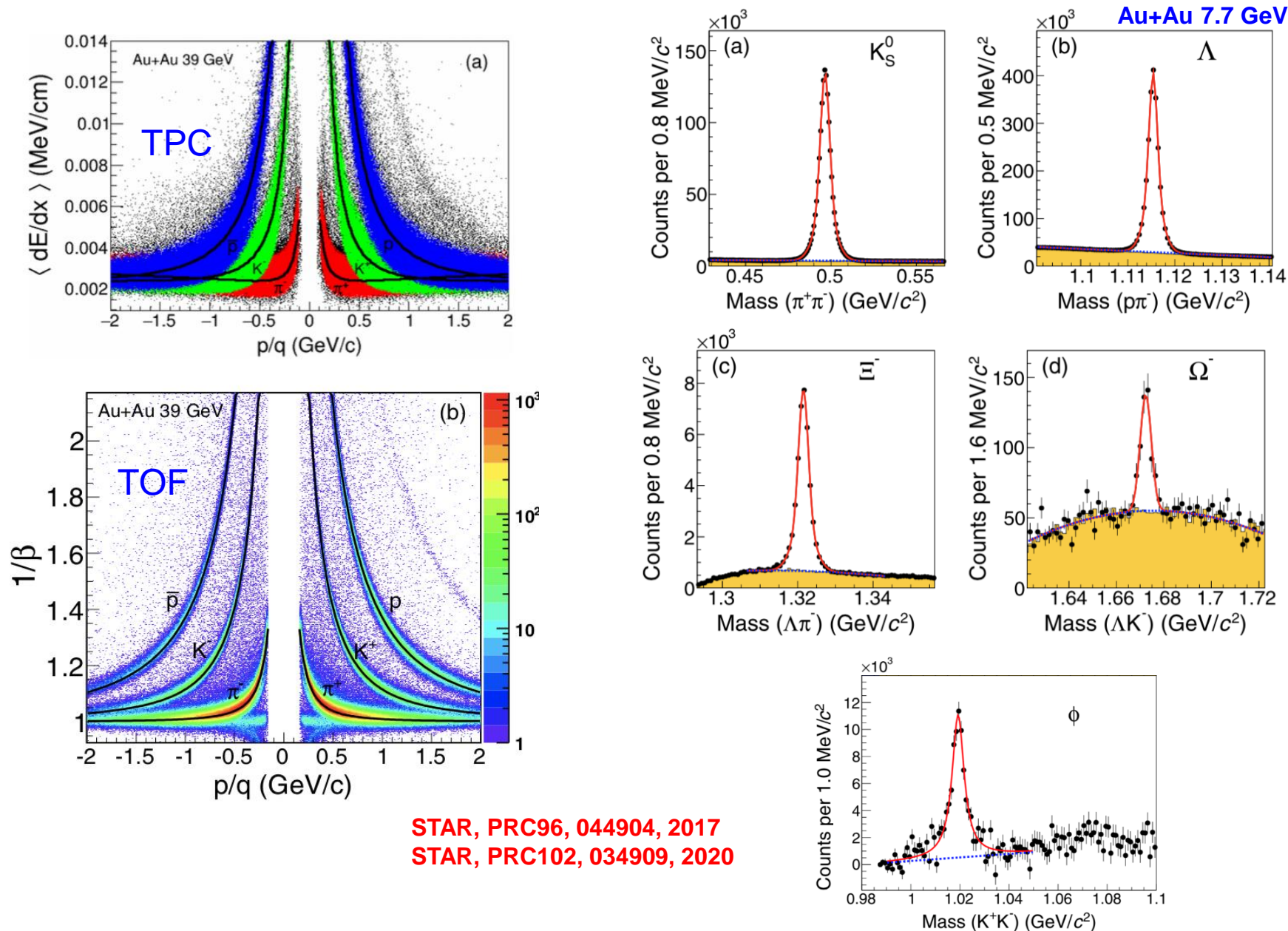
- Strange quarks
 - Not exist in colliding nuclei
 - Current mass $\sim 100 \text{ MeV} < T_c$
 - Easily pair-produced in de-confined QGP medium

→ **Strangeness enhancement !**

- Hadrons with (multiple) strange quarks
 - Small hadronic cross section
 - Sensitive to the early stage dynamics of the medium
 - Can be easily reconstructed and identified in experiment, up to high p_T !

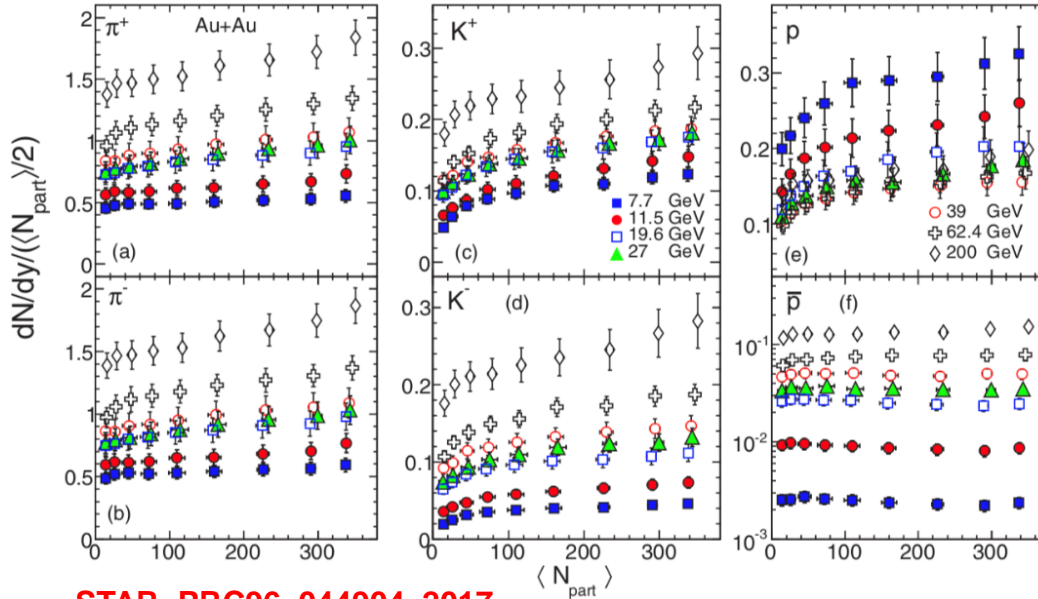


Particle identification and reconstruction

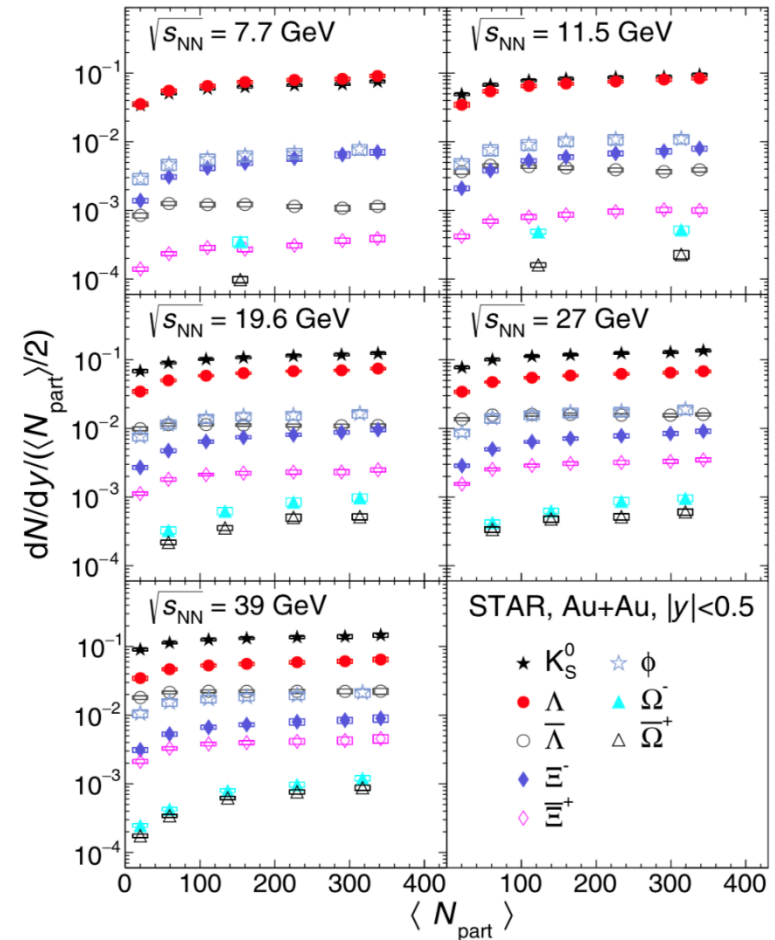


STAR, PRC96, 044904, 2017
 STAR, PRC102, 034909, 2020

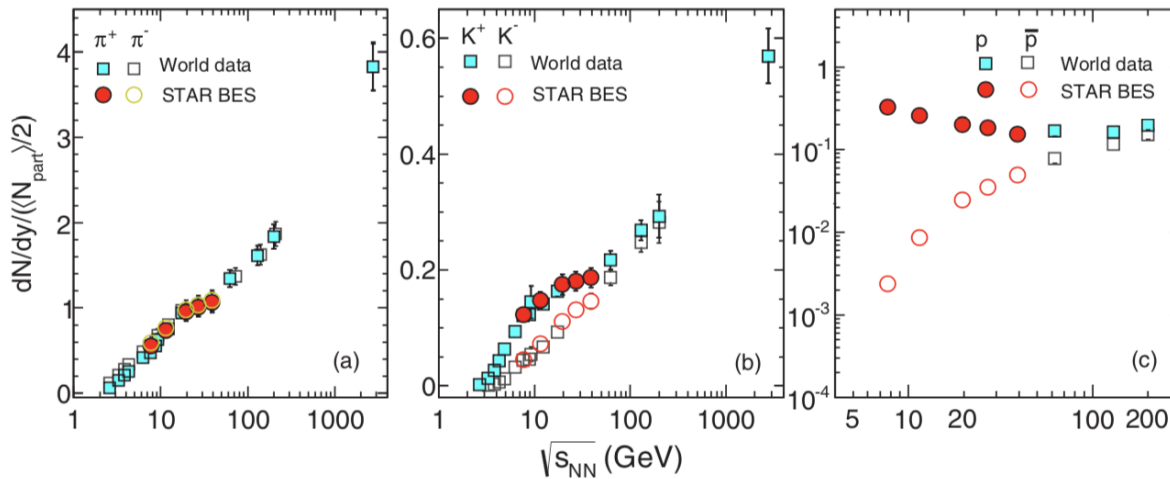
Particle yields, dN/dy , at mid- y



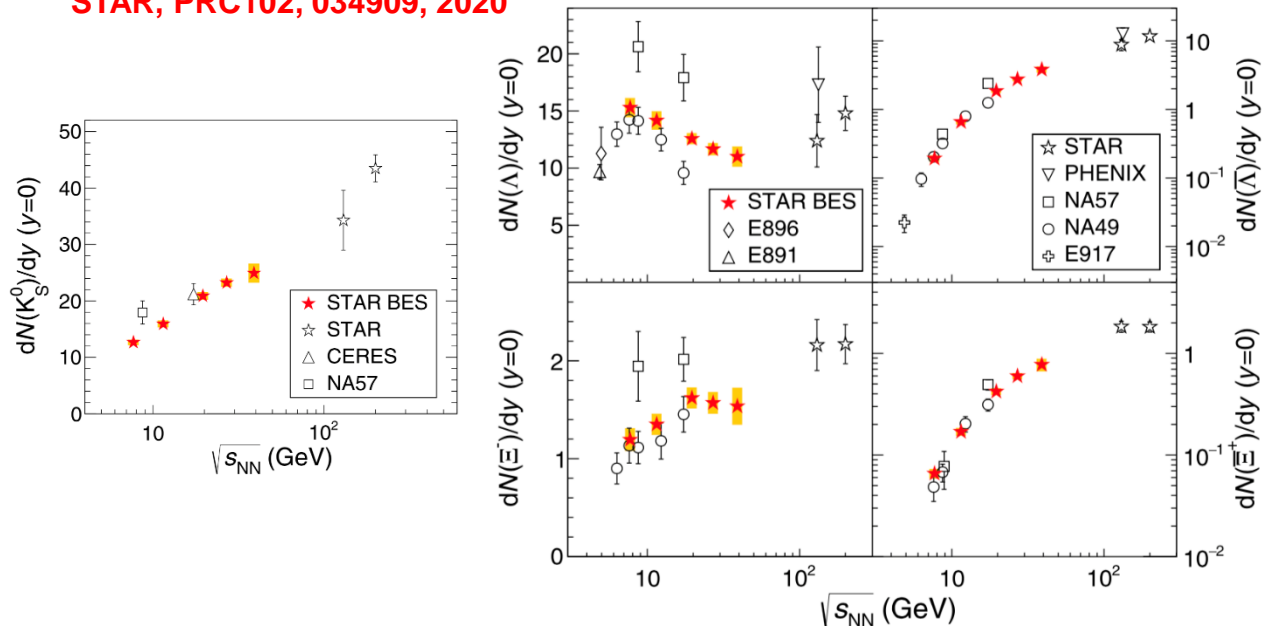
- Yield per participating pair increases towards central and higher energies in general
- Exceptions:
 - p and Λ yields decrease towards higher energy
 - \bar{p} and $\bar{\Lambda}$ has weak centrality dependence



Particle yields in central collisions

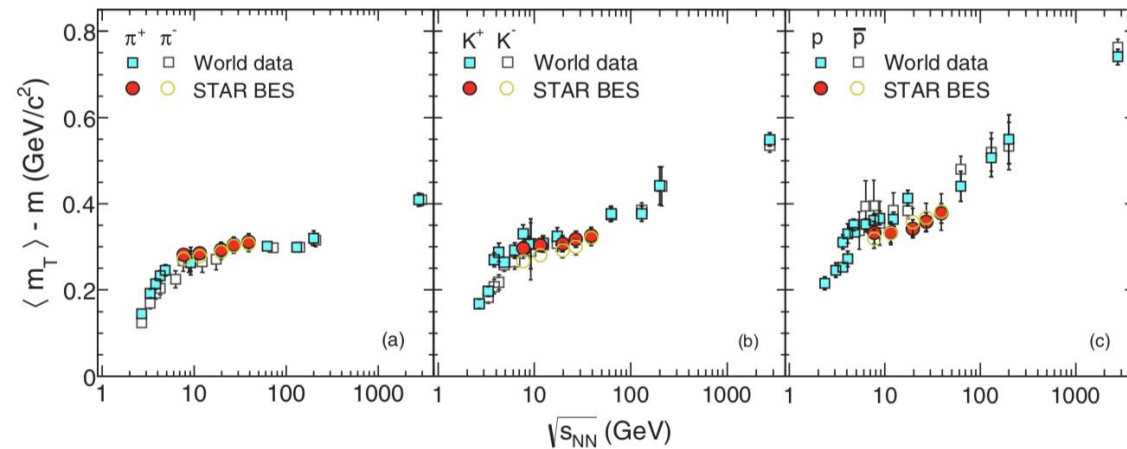


STAR, PRC96, 044904, 2017
 STAR, PRC102, 034909, 2020



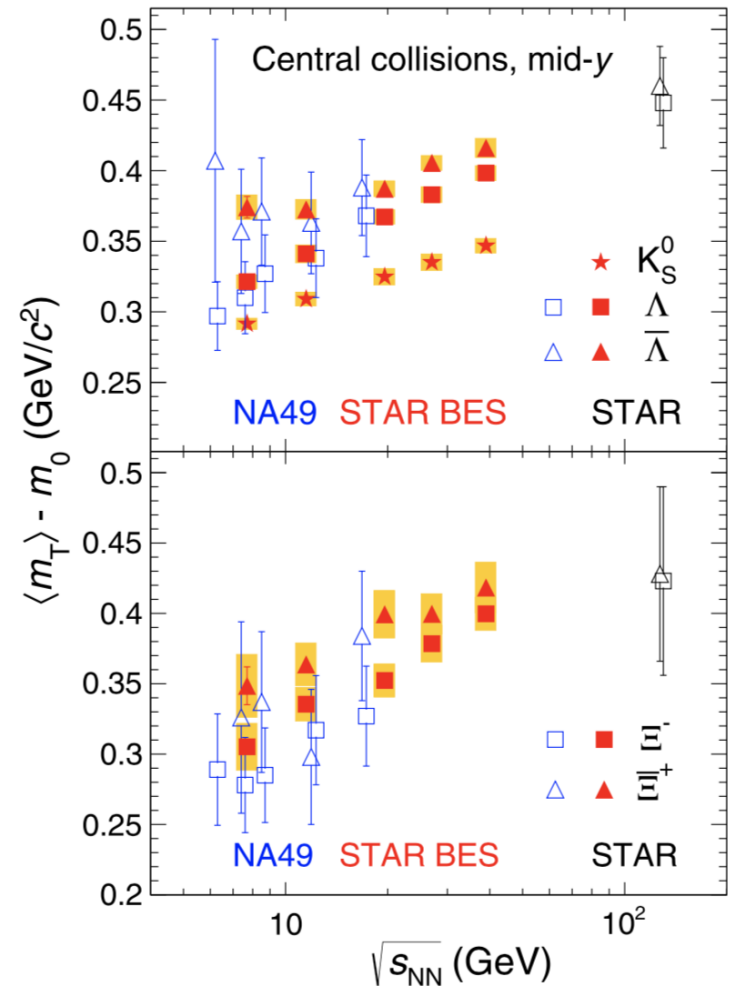
- STAR BES-I data consistent with published data in general
- Rich structure in these excitation functions
- p and Λ yields reach minimum at 39 GeV:
 interplay of baryon transport and pair production

Average transverse mass

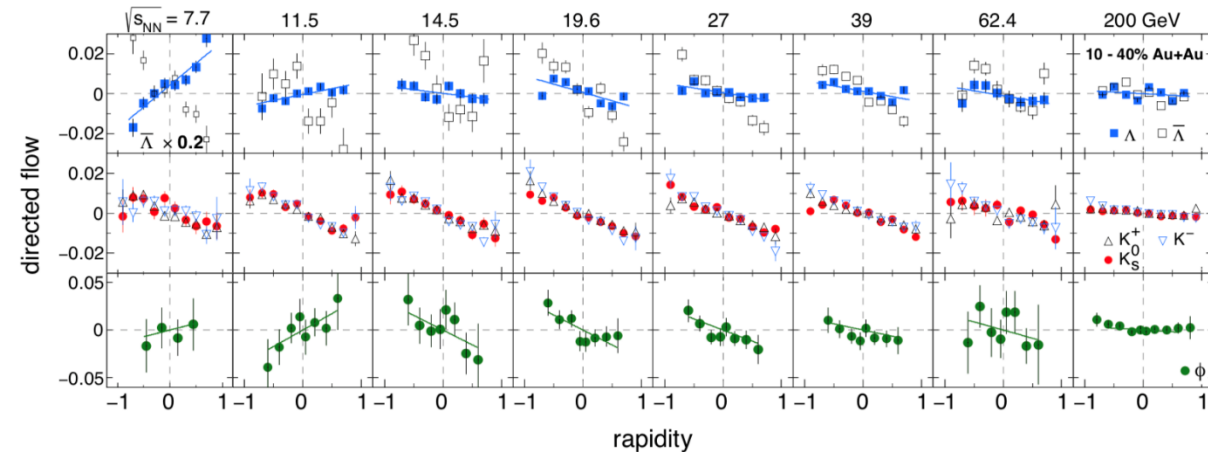


STAR, PRC96, 044904, 2017
STAR, PRC102, 034909, 2020

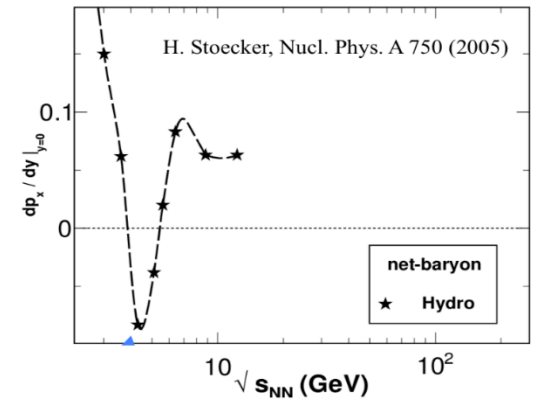
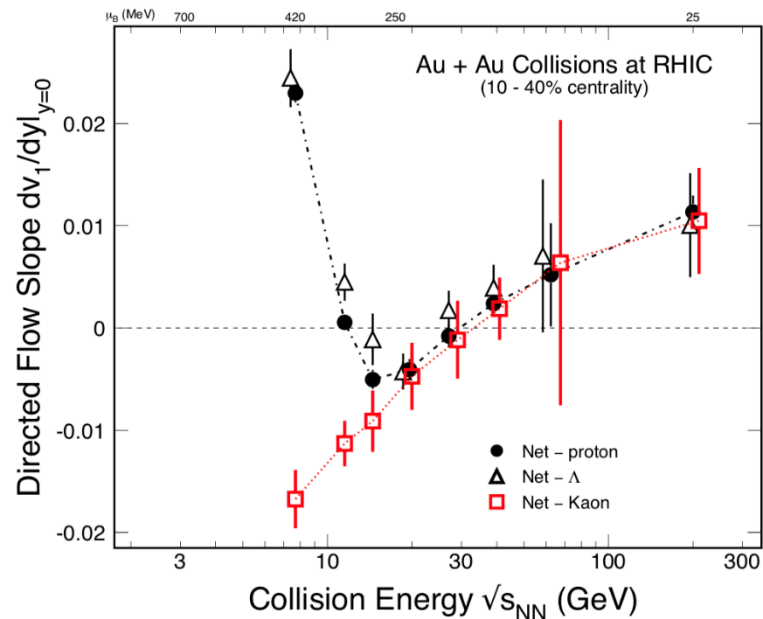
- A step-like structure can be seen in the energy dependence
- Λ and $\bar{\Lambda}$ show split at lower energies might be due to baryon-antibaryon annihilations at high baryon density



Directed flow

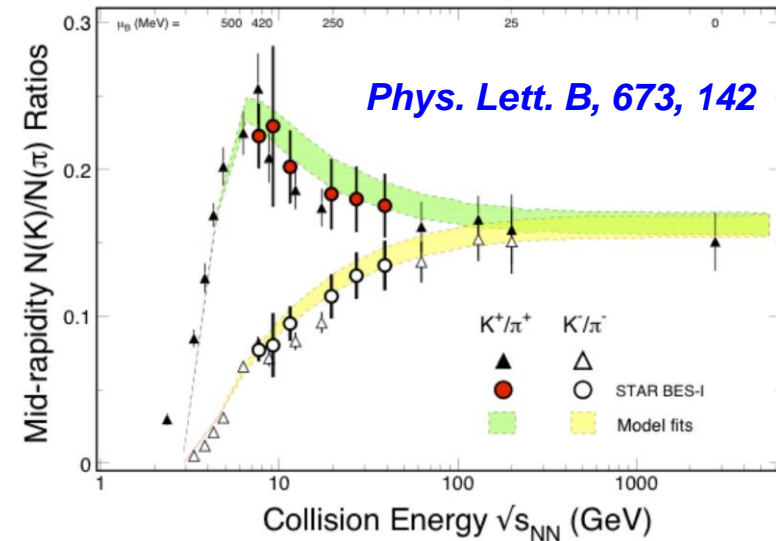


STAR, PRL112, 162301, 2014
STAR, PRL120, 062301, 2018

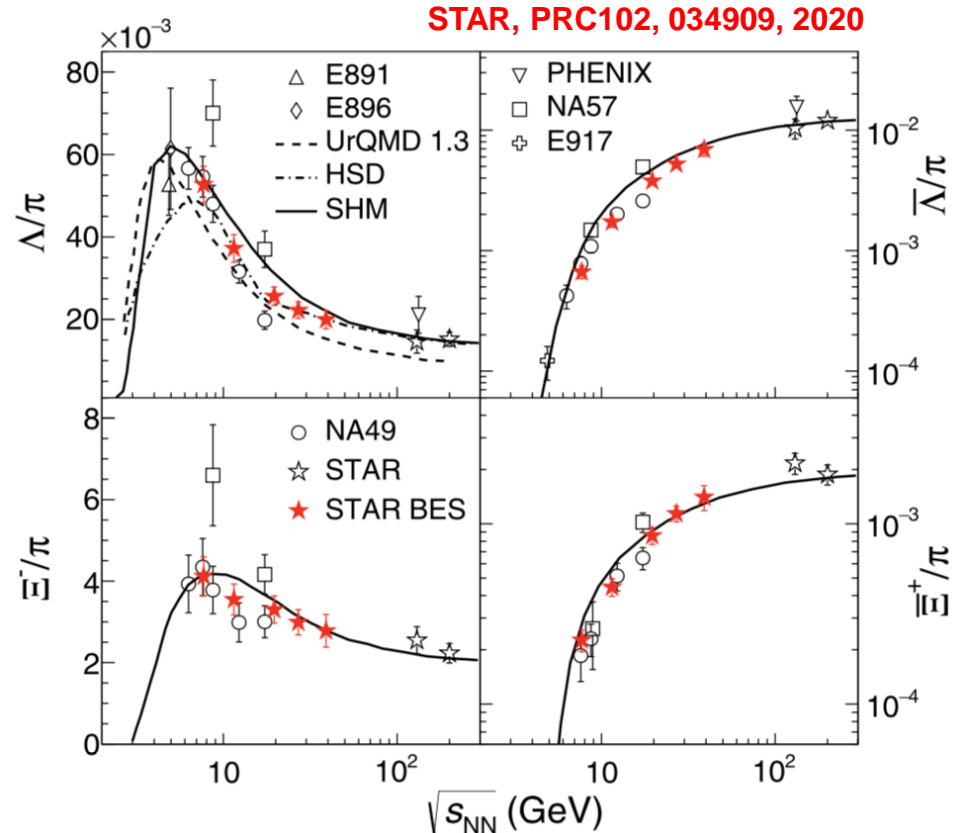
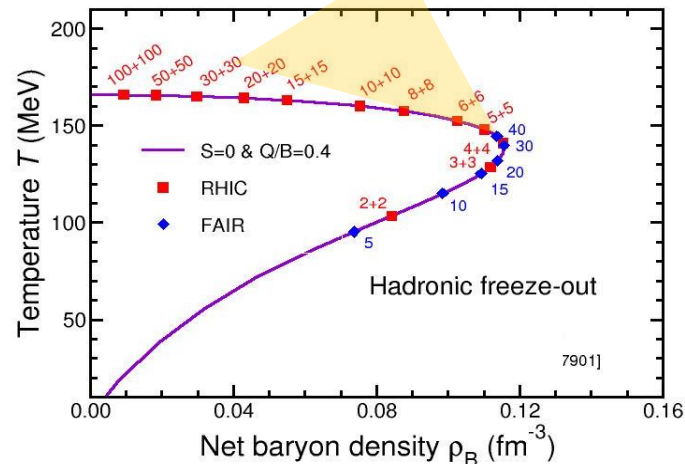


- Sign change of proton dv_1/dy , softening of EOS, first-order phase transition
- Double sign change seen in net-proton, net- Λ , not seen in net-kaon
- Need theory to explain

Strange hadron to pion ratio

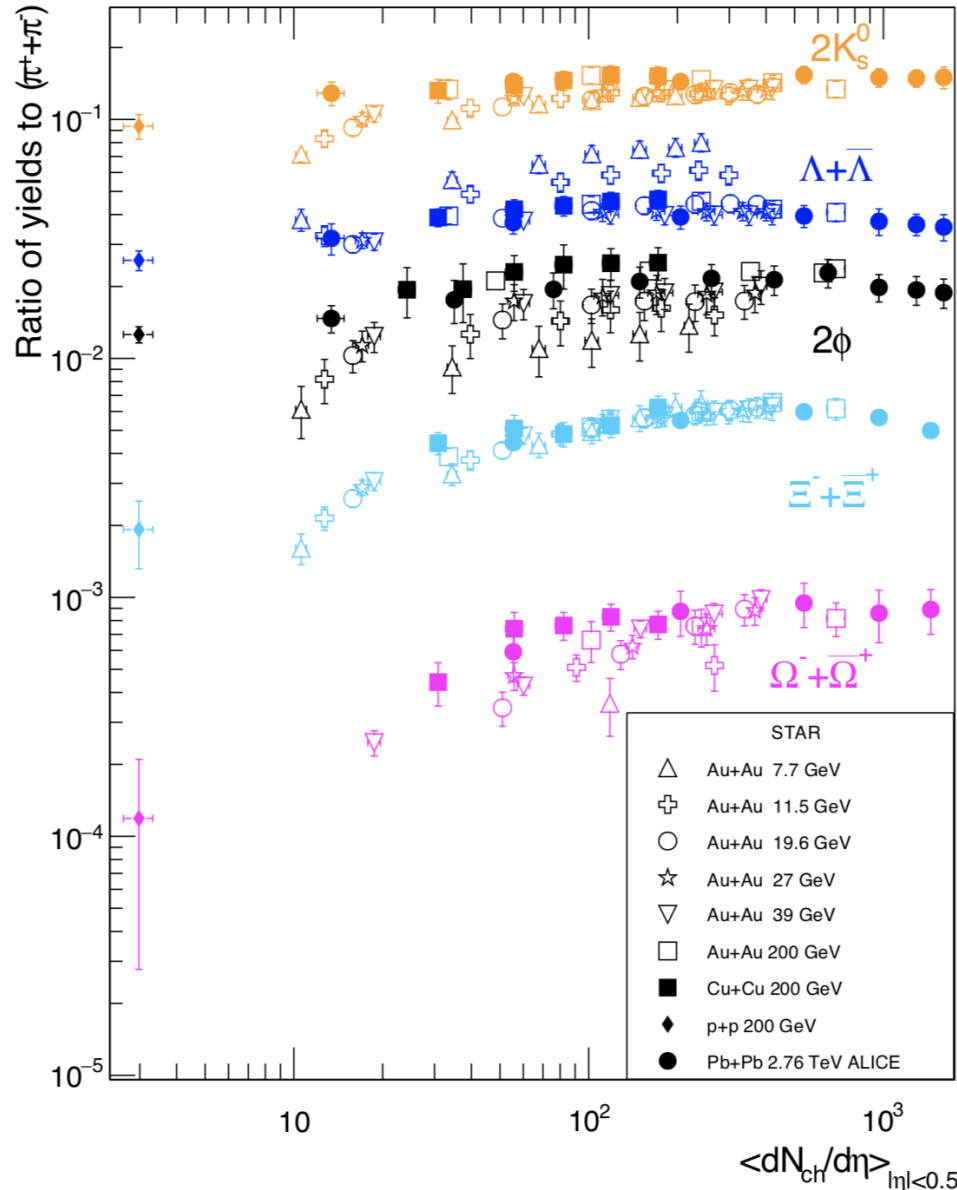


RHIC BES



- Particle ratios consistent with NA49, consistent with the picture of a **maximum net-baryon density around $\sqrt{s_{NN}} \sim 8$ GeV at freeze-out**

Strange hadron to pion ratio vs $dN_{ch}/d\eta$



Yan Huang, APS April Meeting 2021, SQM2021

STAR, PRC96, 044904, 2017

STAR, PRC102, 034909, 2020

ALICE, PRC88, 044910, 2013

$$\frac{dn}{dy} = \frac{\sqrt{M(1 + \sinh^2 y)} \frac{dn}{dy}}{\sqrt{1 + M \sinh^2 y} \frac{dn}{dy}},$$

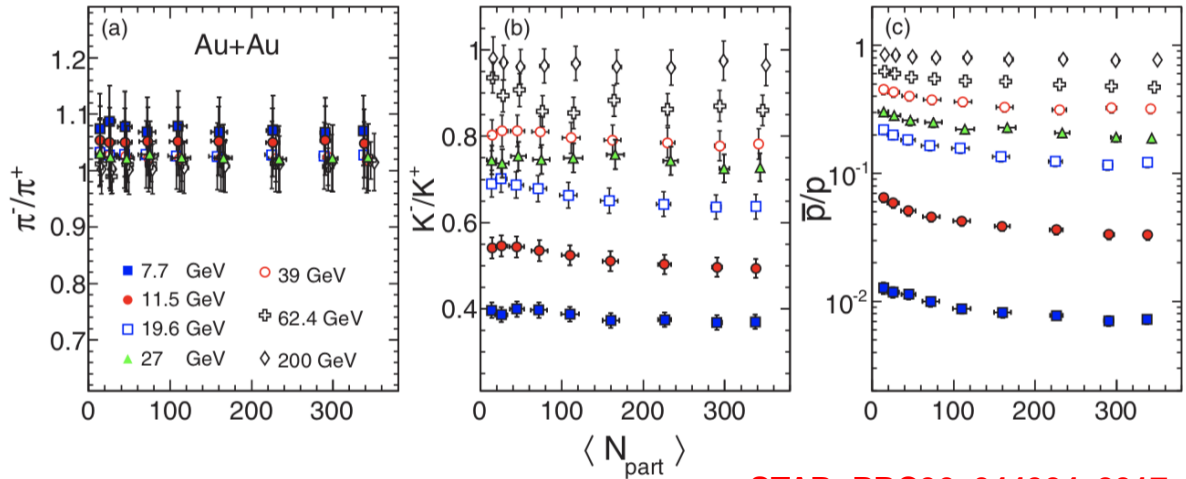
$$\text{where } M = 1 + m^2/p_t^2$$

$$dN_{ch}/d\eta = \sum dN_{ch}/d\eta (k^\pm, \pi^\pm, p, \bar{p})$$

$$dN_{ch}/d\eta(\eta = 0) \sim dN_{ch}/d\eta(|\eta| < 0.5)$$

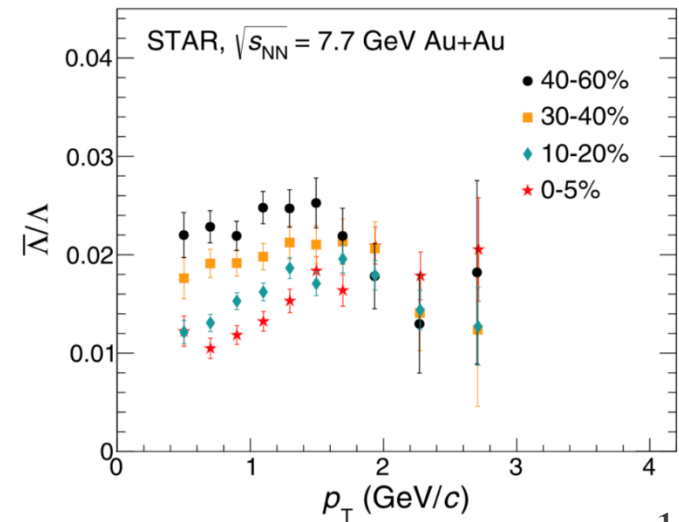
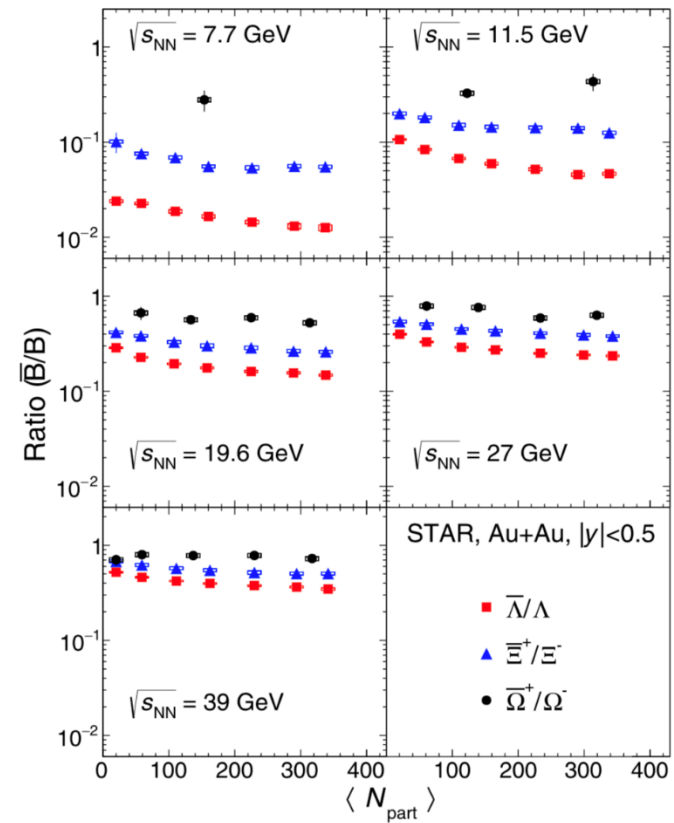
- The ratios at different energies/centralities/systems mainly depend on charged hadrons multiplicity, except for Λ and ϕ
- The ratios saturate at large charged hadrons multiplicity

Anti-hadron to hadron ratio

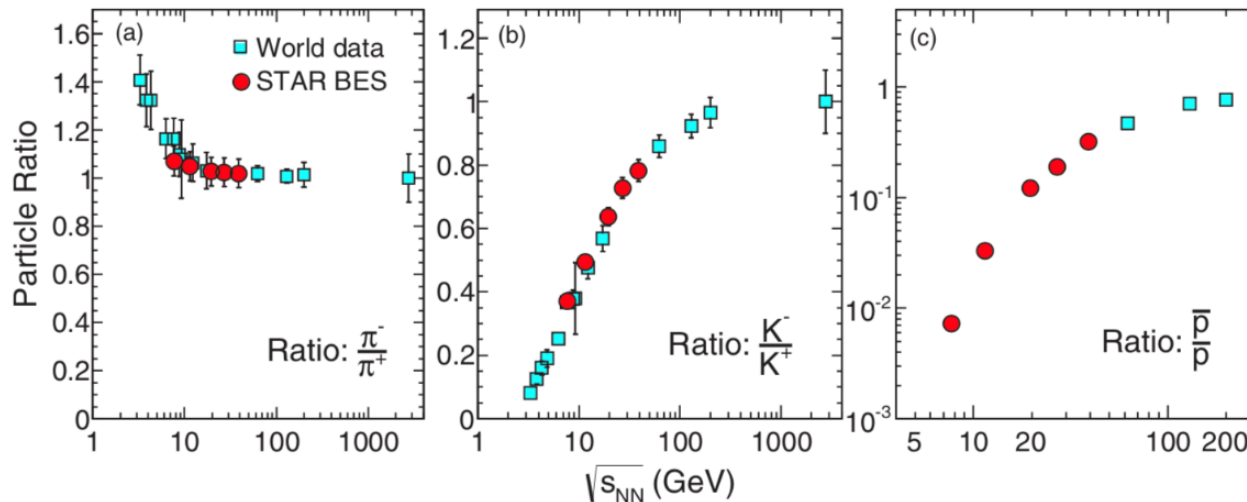


STAR, PRC96, 044904, 2017
STAR, PRC102, 034909, 2020

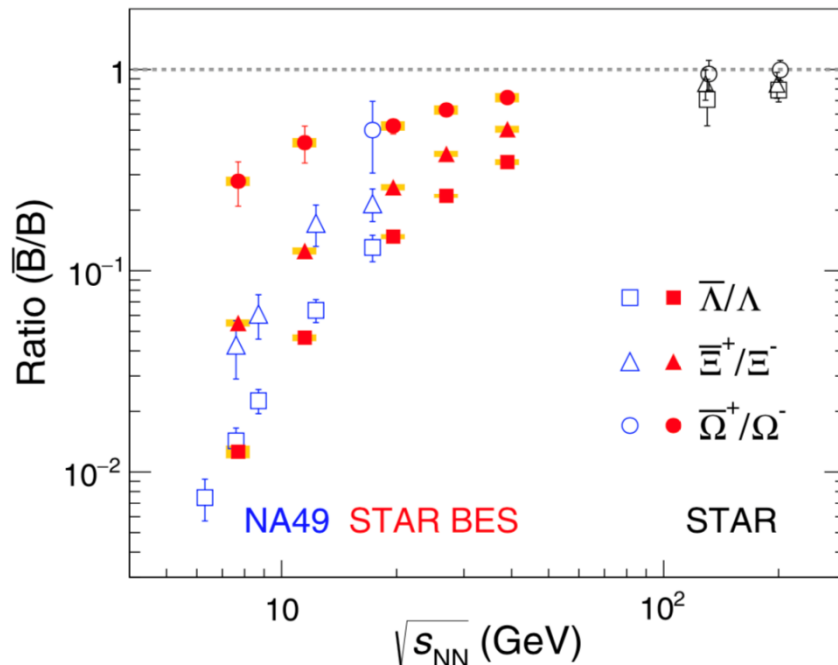
- Centrality dependence of \bar{B}/B ratios:
peripheral > central
- This effect is more prominent at lower energies.
baryon stopping and/or anti-baryon absorption
- Loss of low p_T $\bar{\Lambda}$ in central collisions



Anti-hadron to hadron ratio



STAR, PRC96, 044904, 2017
STAR, PRC102, 034909, 2020



- STAR BES data lie in a trend with NA49 data
- \bar{B}/B ratios increase with number of strange quarks at low energies
 $\bar{\Omega}^+/\Omega^- > \bar{\Xi}^+/\Xi^- > \bar{\Lambda}/\Lambda > \bar{p}/p$

Anti-hyperon to hyperon ratio

$$n_i = \frac{g_i}{(2\pi^2)} \gamma_S^{|S_i|} m_i^2 T K_2(m_i/T) \exp(\mu_i/T)$$

$$\frac{\bar{\Lambda}}{\Lambda} = \exp\left(-\frac{2\mu_B}{T} + \frac{2\mu_S}{T}\right)$$

$$\ln\left(\frac{\bar{\Lambda}}{\Lambda}\right) = -\frac{2\mu_B}{T} + \frac{2\mu_S}{T}$$

$$\frac{\bar{\Xi}^+}{\Xi^-} = \exp\left(-\frac{2\mu_B}{T} + \frac{4\mu_S}{T}\right)$$



$$\ln\left(\frac{\bar{\Xi}^+}{\Xi^-}\right) = -\frac{2\mu_B}{T} + \frac{4\mu_S}{T}$$

$$\frac{\bar{\Omega}^+}{\Omega^-} = \exp\left(-\frac{2\mu_B}{T} + \frac{6\mu_S}{T}\right)$$

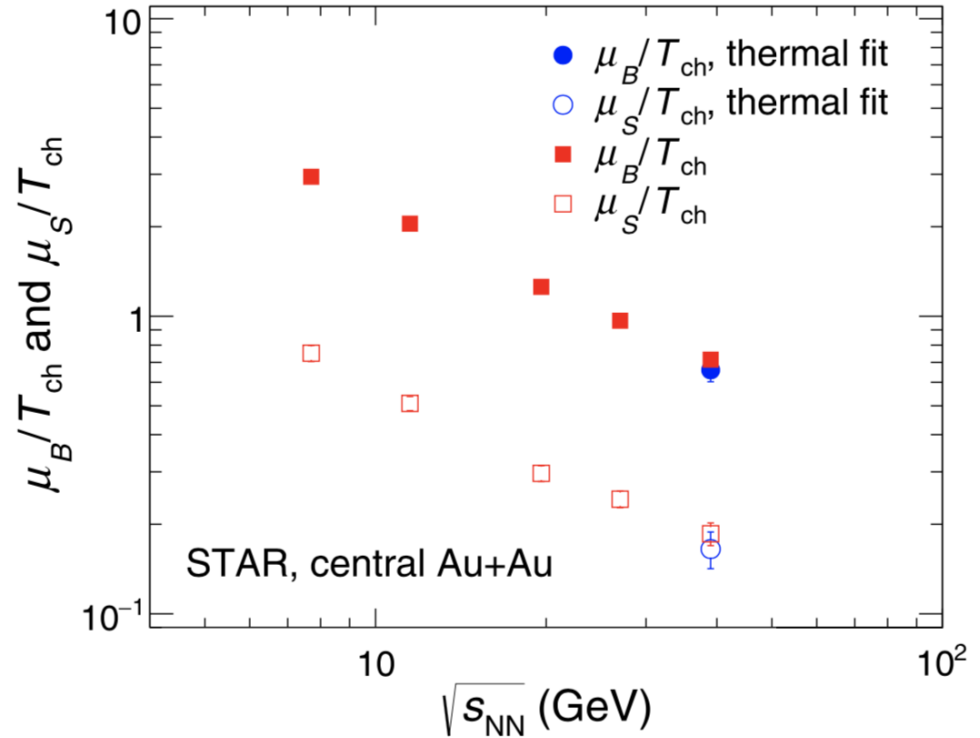
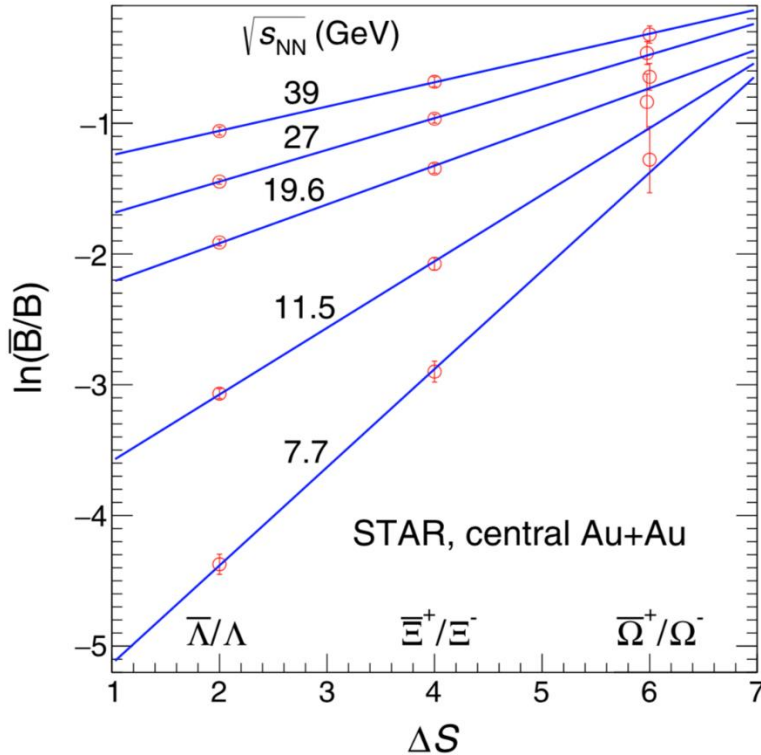
$$\ln\left(\frac{\bar{\Omega}^+}{\Omega^-}\right) = -\frac{2\mu_B}{T} + \frac{6\mu_S}{T}$$

- T is the temperature.
- μ_B is the baryon chemical potential.
- μ_S is the strangeness chemical potential.

(arXiv:nucl-th/9704046v1 by J.Cleymans & Phys. Rev. C 71(2005)054901)

μ_S/T_{ch} and μ_B/T_{ch}

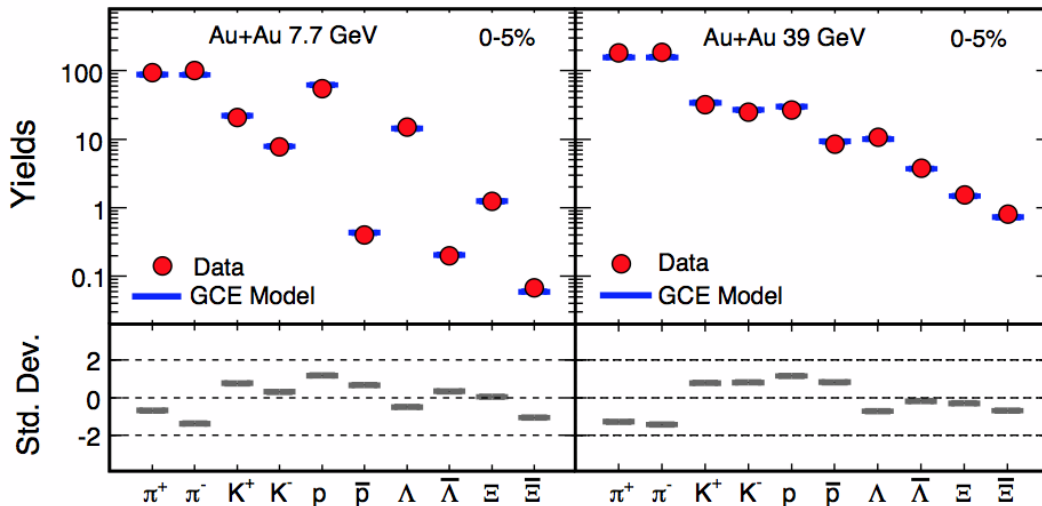
STAR, PRC102, 034909, 2020



- Anti-hyperon to hyperon ratios are fit well with statistical thermal model
- Chemical freeze-out parameters, μ_S/T_{ch} and μ_B/T_{ch} , are extracted

Chemical freeze-out parameters: T_{ch} vs. μ_B

STAR, Phys. Rev. C 96, 044904, 2017



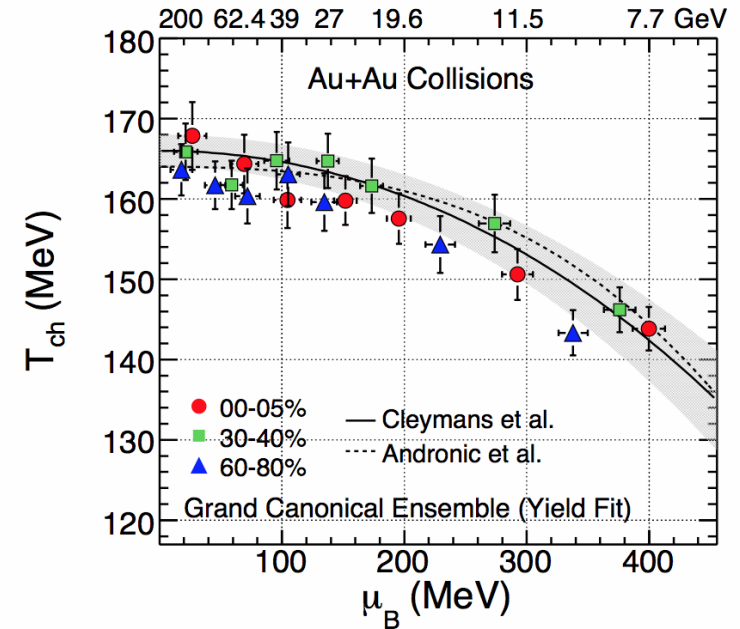
✓ Particles used : π , K , p , Λ , Ξ

✓ Ensemble used:

Grand canonical (GCE)

✓ Fit parameters:

T_{ch} , μ_B , μ_s and γ_s



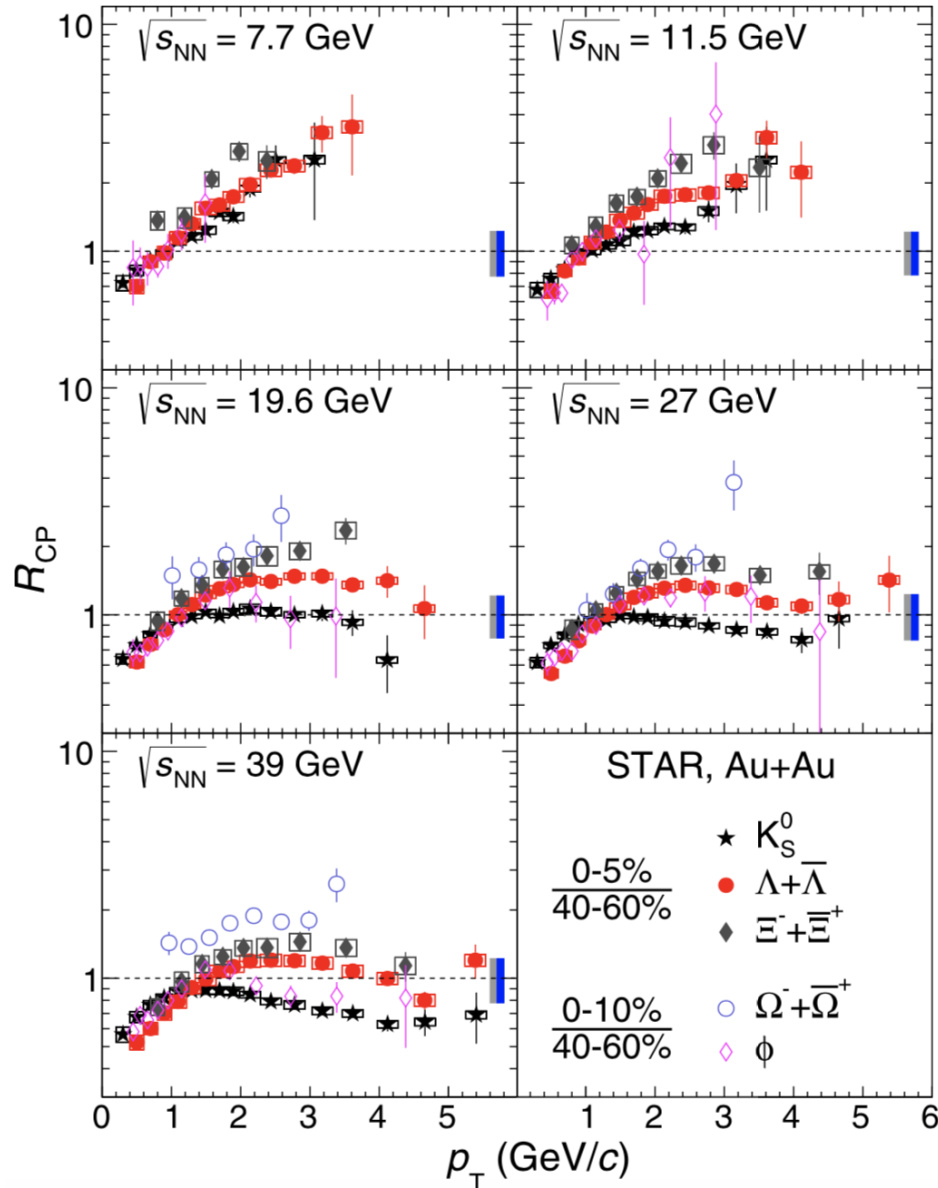
Andronic: NPA 834 (2010) 237

Cleymans: PRC 73 (2006) 034905

Au+Au 200 GeV : Phys. Rev. C **83** (2011) 24901

Thermus, S. Wheaton & J. Cleymans, Comput. Phys. Commun. 180: 84-106, 2009.

Nuclear modification factors R_{CP}

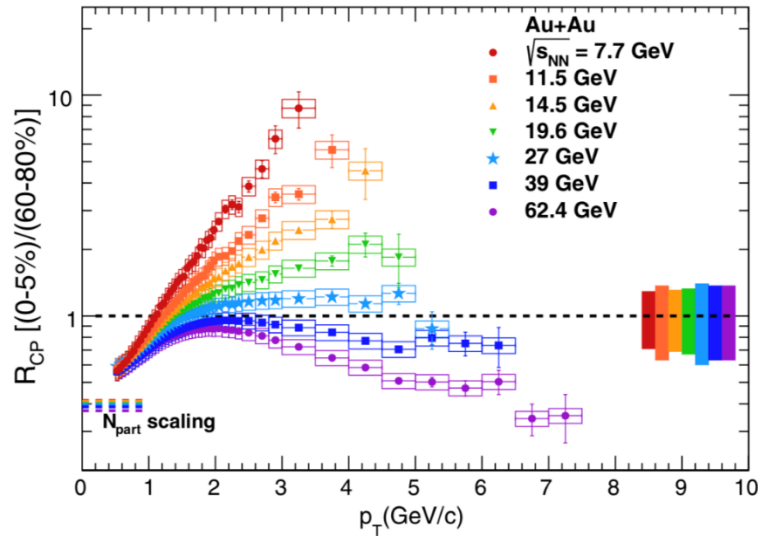


$$R_{CP}(p_T) = \frac{[d^2\sigma/(N_{bin}p_T dp_T dy)]_{central}}{[d^2\sigma/(N_{bin}p_T dp_T dy)]_{peripheral}}$$

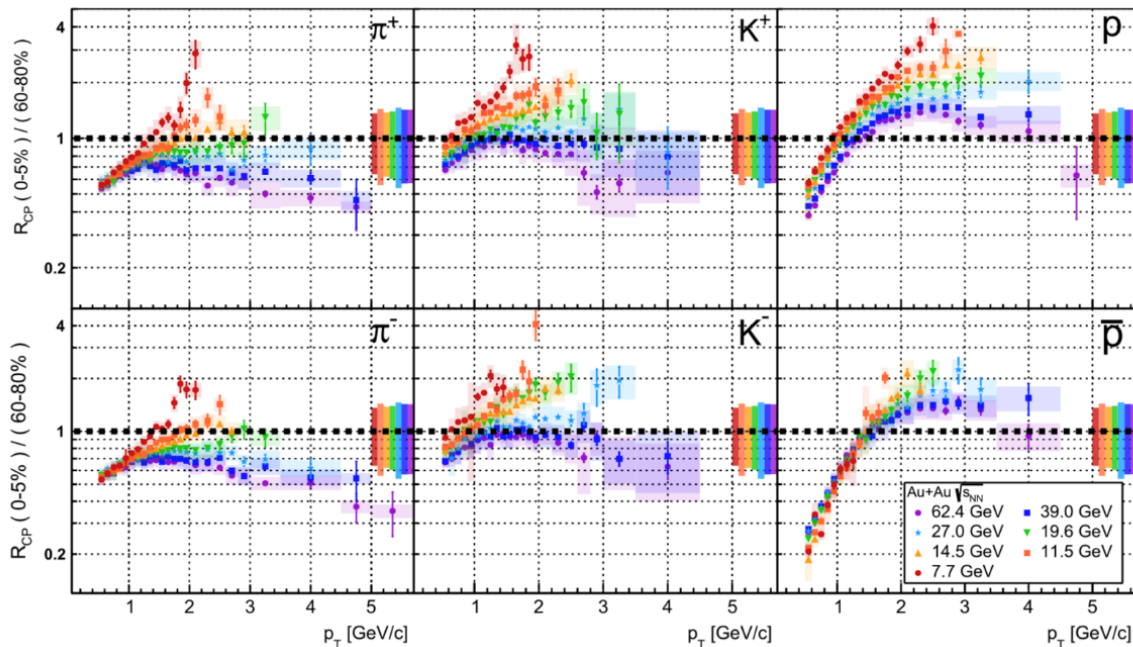
- No K_S^0 suppression in Au+Au 7.7 and 11.5 GeV
- Cronin effect and other effects (radial flow) compete with partonic energy loss
- Intermediate p_T , particle R_{CP} difference becomes smaller @ 7.7 and 11.5 GeV

STAR, PRC102, 034909, 2020

Nuclear modification factors R_{CP}

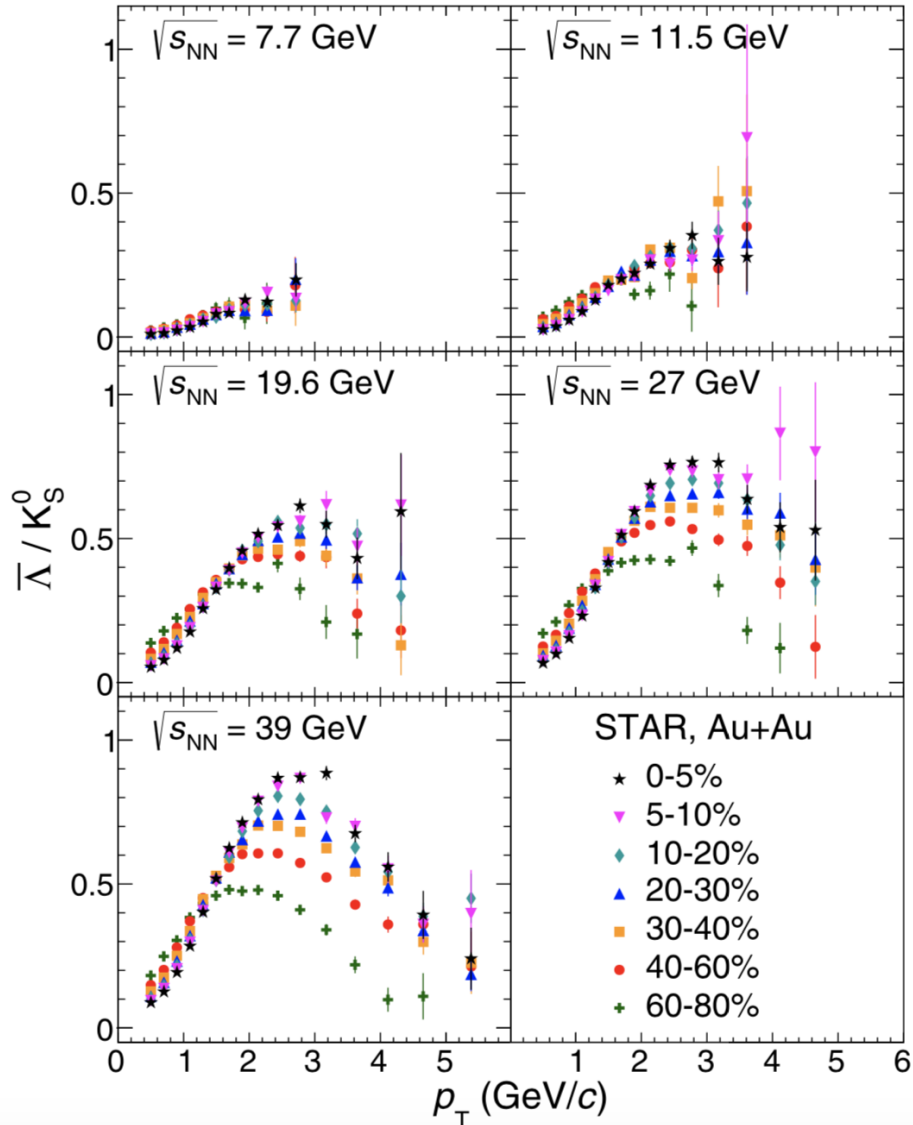


- No suppression for lower energies
- Cronin effect and other effects (radial flow) compete with partonic energy loss



STAR, PRL121, 032301, 2018

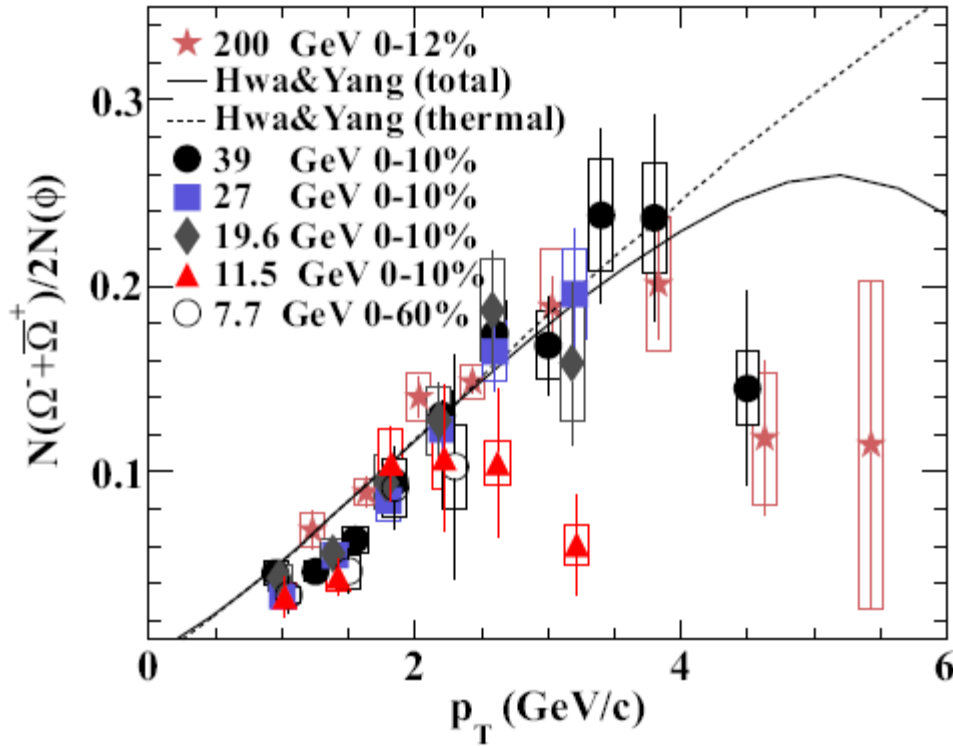
Baryon to meson ratio: $\bar{\Lambda}/K_S^0$



STAR, PRC102, 034909, 2020

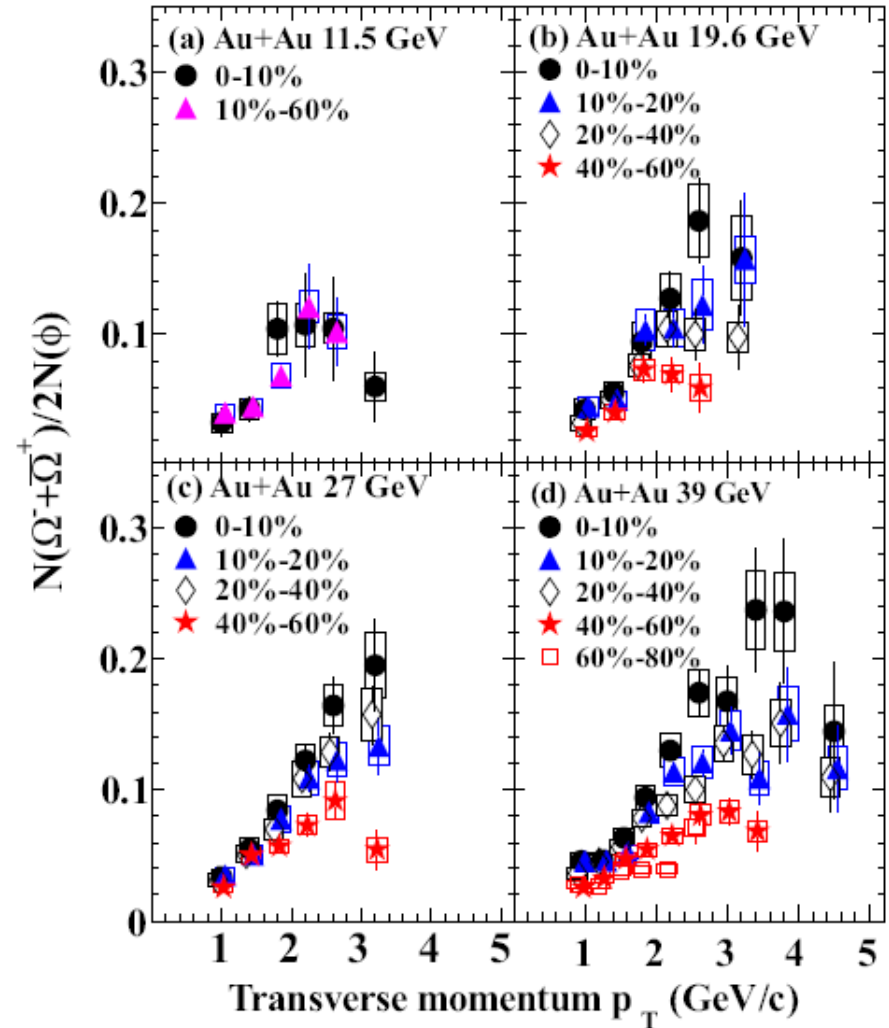
$\sqrt{s_{NN}} < 19.6$ GeV, at intermediate p_T , the separation of central (0-5%) and peripheral (40-60%) collisions in $\bar{\Lambda}/K_S^0$ becomes less significant

Ω / ϕ ratio

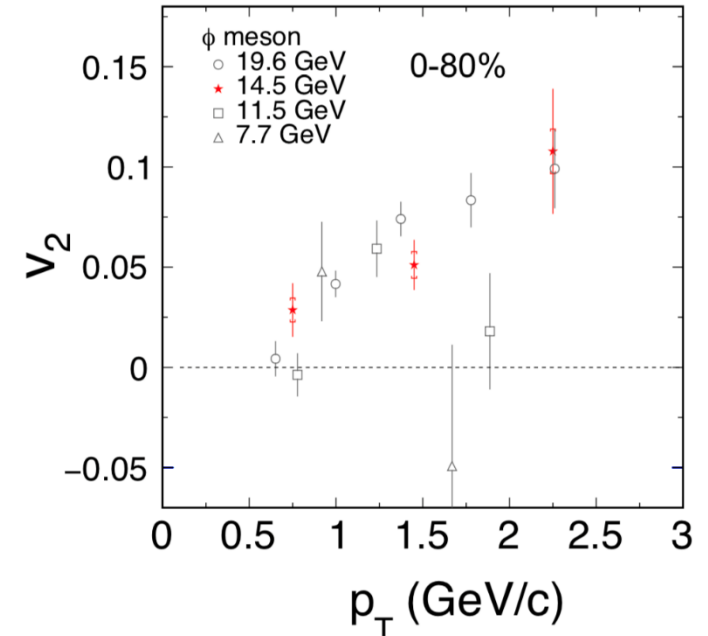
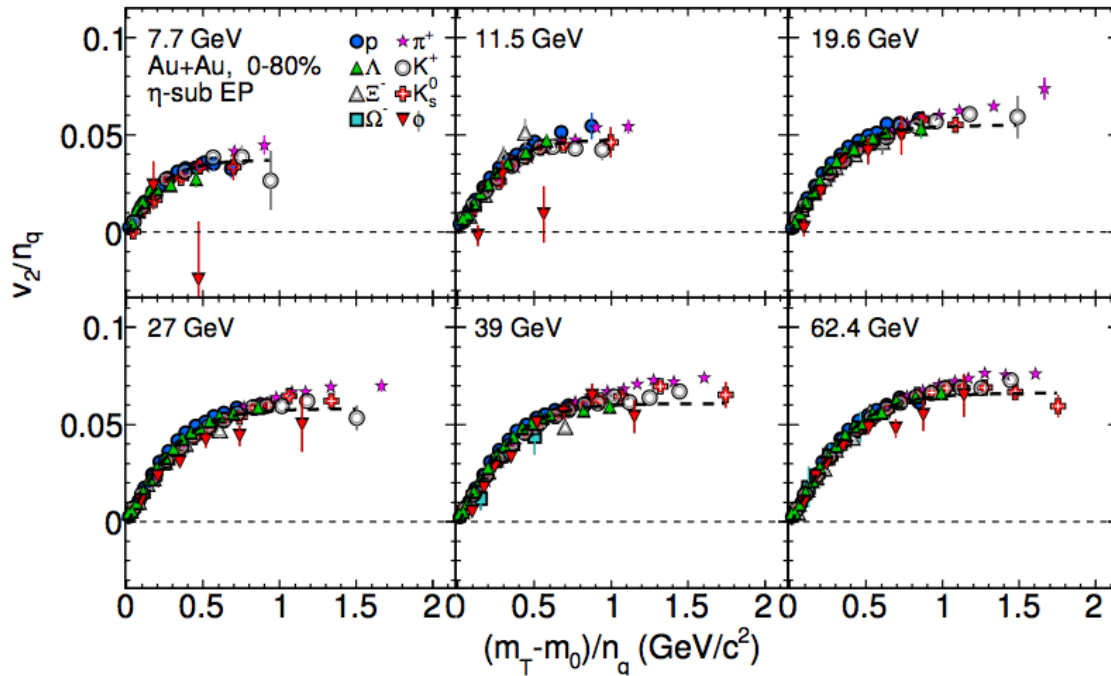


STAR, Phys. Rev. C 93, 021903 (R), 2016

- Intermediate p_T Ω/ϕ ratios:
Indication of separation between ≥ 19.6 and 11.5 GeV
- Ω/ϕ ratios: 40%-60% peripheral $<$ 0-10% central for 19.6, 27 and 39 GeV



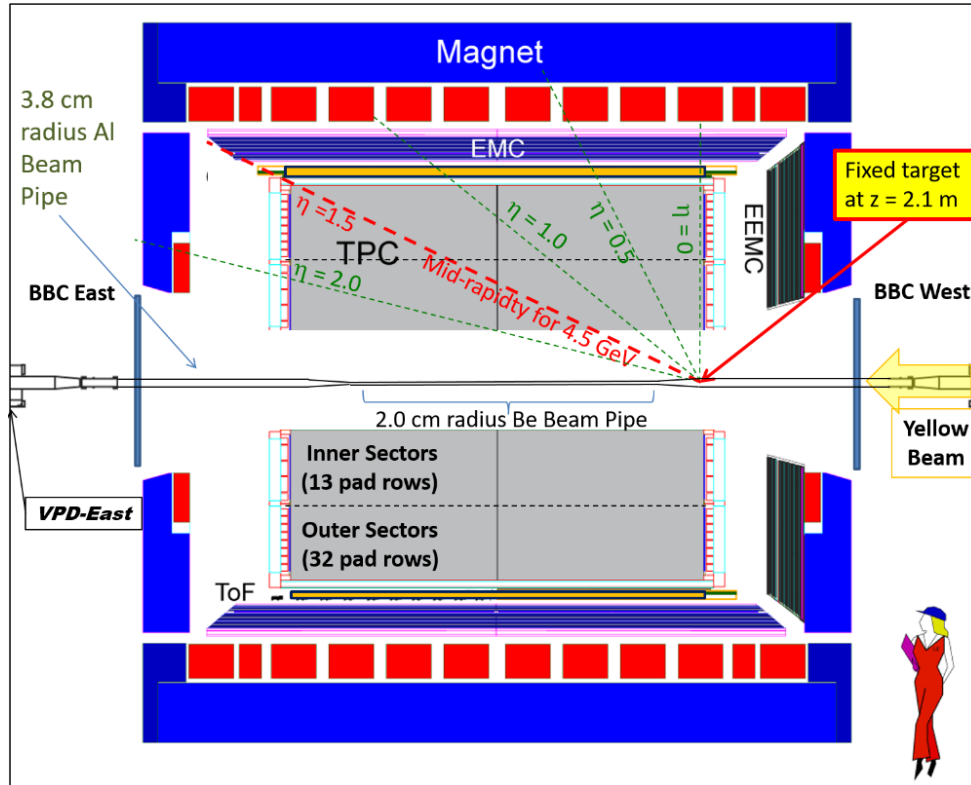
Elliptic flow



STAR, Phys. Rev. C88 (2013) 014902; Phys. Rev. C93 (2016) 014907

- NCQ scaling holds with 10% for selected particles in all energies
- 14.5 GeV: Sizable ϕ meson v_2 , comparable to 19.6 GeV
- High statistics and more collision energies below 20 GeV needed!

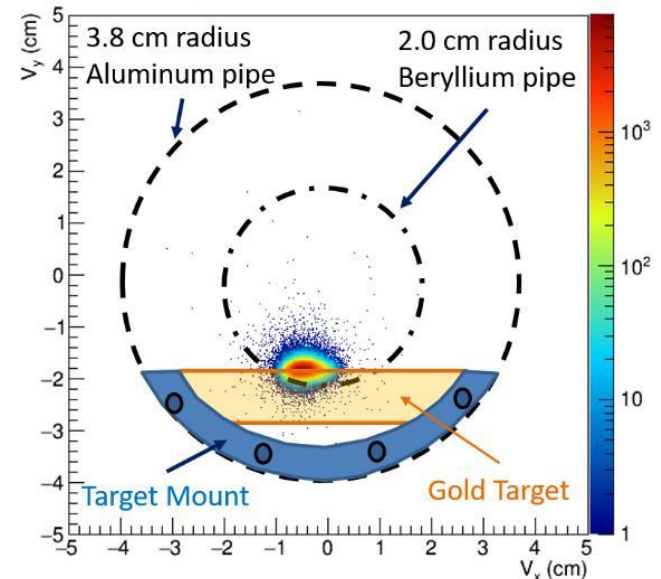
The STAR fixed-target program



1 mm thick (4% inter. prob.) gold target (2015)
 → 1/4 mm thick gold target in FXT phys. program



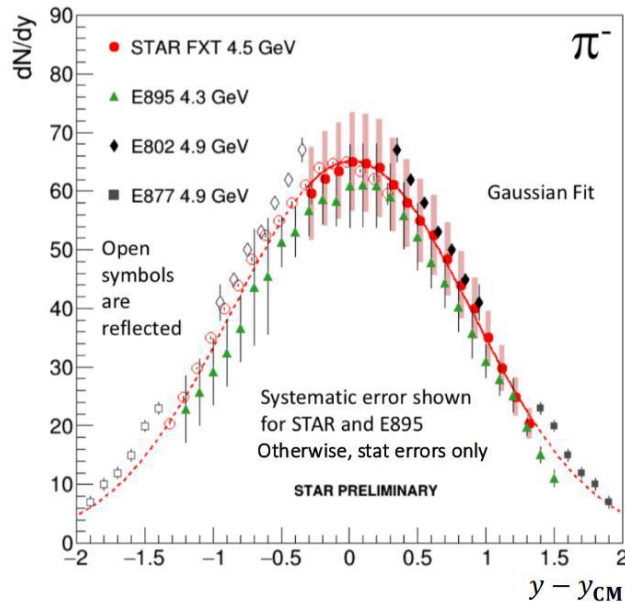
V_y vs. V_x Distribution



1.3M events from half hour **test run**, top 30% central trigger, Au+Au $\sqrt{s_{NN}}=4.5$ GeV

3.4M events from two hour **test run**, top 30% central trigger, Al+Au $\sqrt{s_{NN}}=4.9$ GeV

Hadron spectra and dN/dy in Au+Au $\sqrt{s_{NN}}=4.5$ GeV

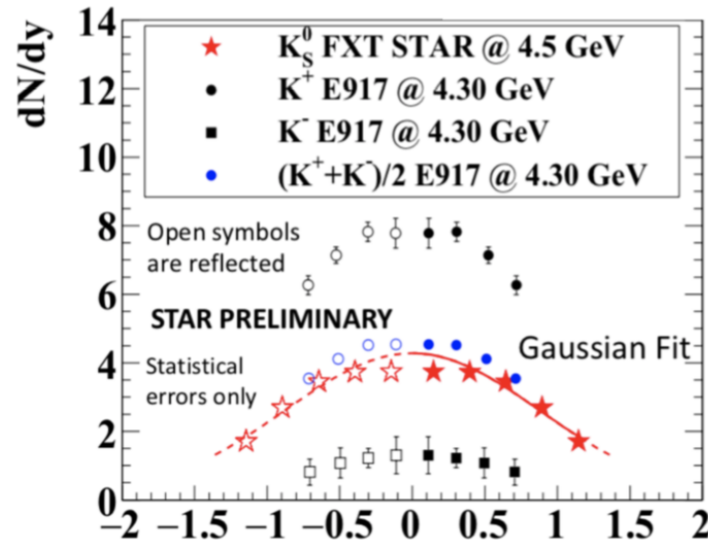


E895. Phys. Rev. C 68 (2003) 054905

E802. Phys. Rev. C 57 (1998) R466

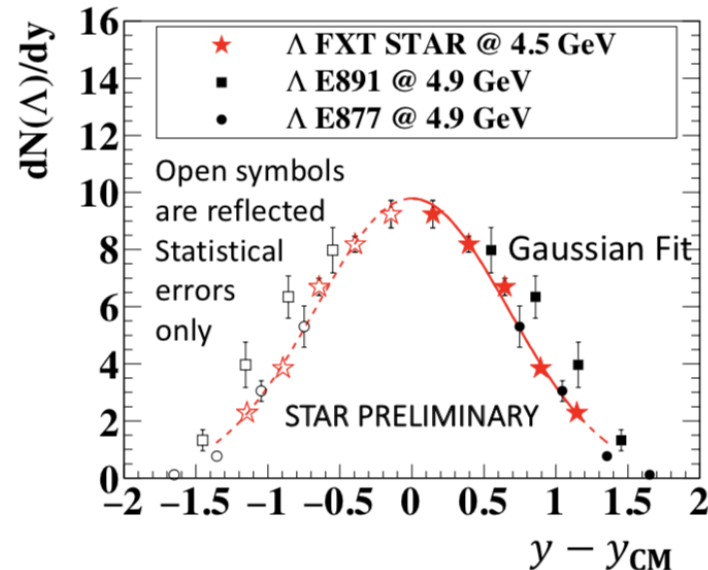
E877. Phys. Rev. C 62 (2000) 024901

- Amplitude and width of rapidity densities are consistent with AGS experiments
- $m_T - m_0$ and y range can be extended by eTOF and iTPC upgrades



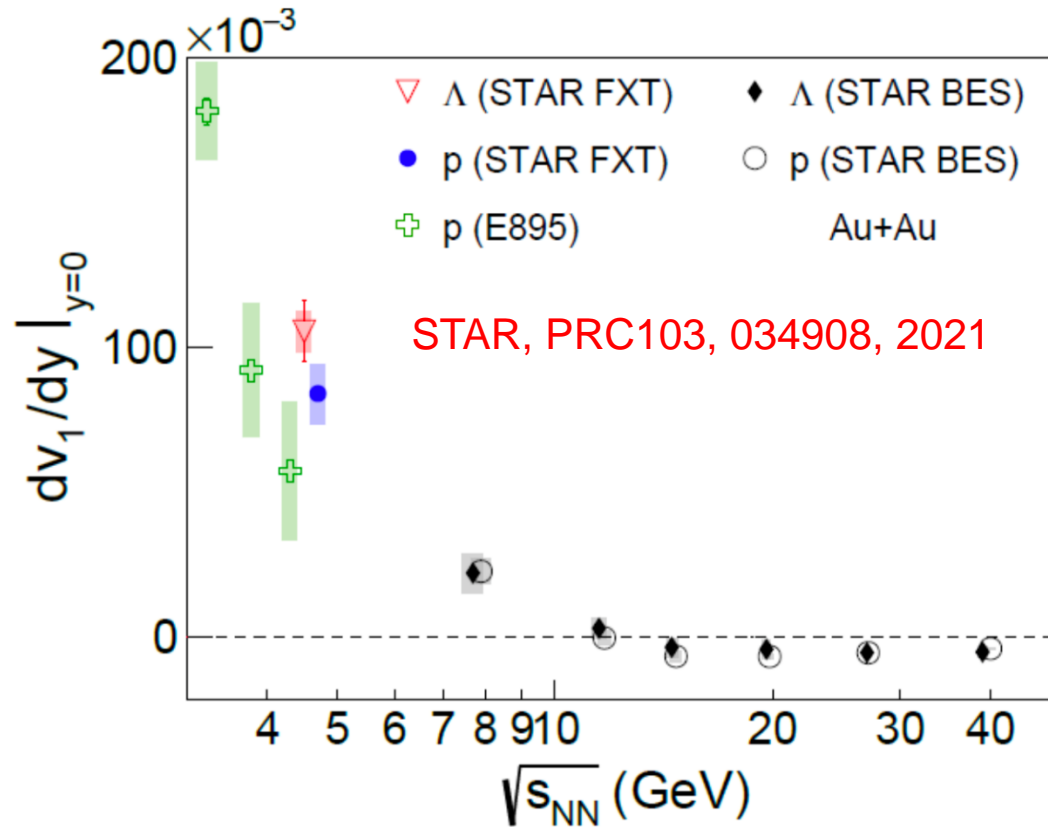
*Y. Wu, QM2018
Top 5%*

*M.-U. Ashraf,
QM2019*



Directed flow in Au+Au $\sqrt{s_{NN}}=4.5$ GeV

E895. Phys. Rev. Lett. 84 (2000) 005488
STAR . Phys. Rev. Lett. 112 (2014) 162301



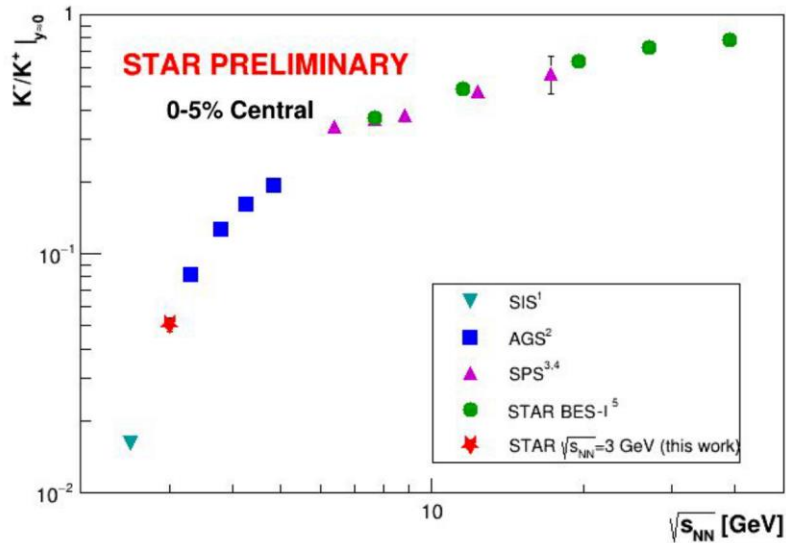
Baryon v_1 slope is consistent with E895 at 4.3 GeV

The STAR fixed-target physics program (2018-2021)

Year	<i>Collisions</i>	$\sqrt{s_{NN}}$ (GeV)	Good events
2018	Au+Au	3.0	~ 258 M
2018	Au+Au	7.2	~ 155 M
2019	Au+Au	3.9	~ 53 M
2019	Au+Au	3.2	~ 200 M
2019 / 2020	Au+Au	7.7	~ 164 M
2020	Au+Au	4.5	~ 108 M
2020	Au+Au	6.2	~ 118 M
2020	Au+Au	5.2	~ 103 M
2020	Au+Au	3.9	~ 117 M
2020	Au+Au	3.5	~ 116 M
2021	Au+Au	3.0, 9.2, 11.5,13.7...	On-going

D. Cebra, APS April Meeting, 2021

Charged kaon production in Au+Au $\sqrt{s_{NN}}=3.0$ GeV



B. Kimelman, CPD2021

SIS: *J. Phys. G* 28, 2011

AGS: *Phys. Lett. B* 490, 53;

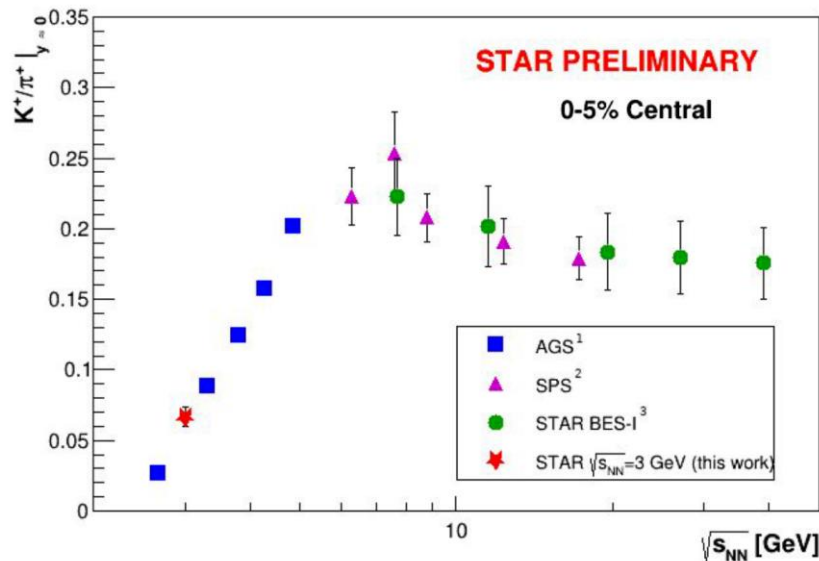
Phys. Lett. B 476, 1

SPS: *Phys. Rev. C* 77, 024903;

Phys. Rev. C 66, 054902;

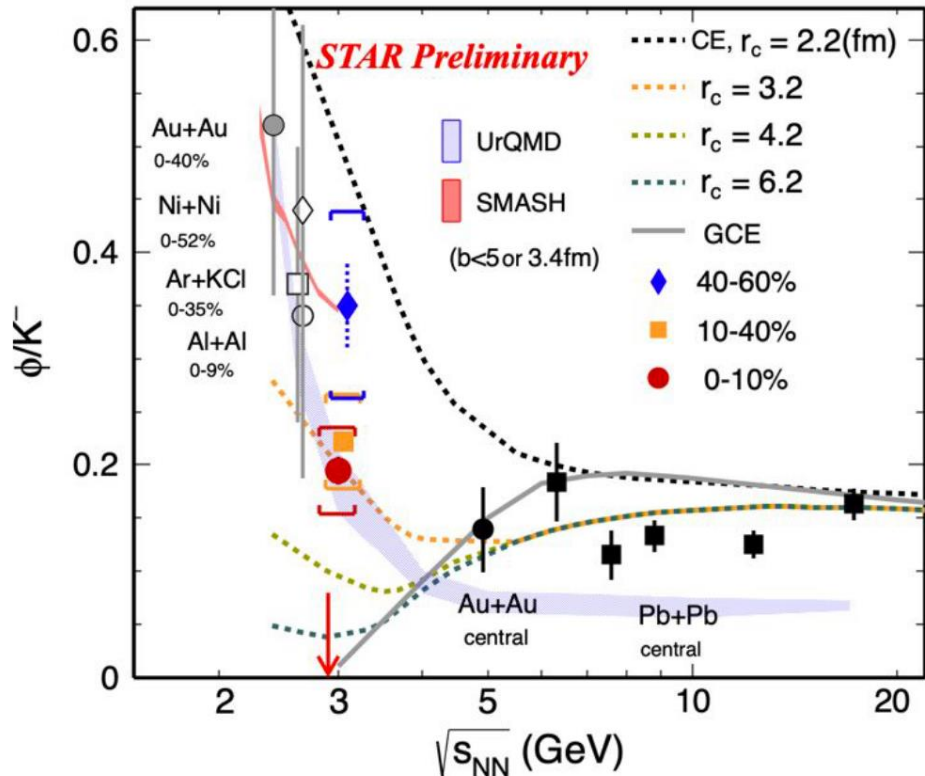
Phys. Rev. C 77, 024903

STAR: *Phys. Rev. C* 96, 044904



- K^+/π^+ and K^-/K^+ ratios follow world trend
- K^-/K^+ ratio drops at lower energies from associated production of K^+ ($NN \rightarrow N\Lambda K$)

ϕ production in Au+Au $\sqrt{s_{NN}}=3.0$ GeV



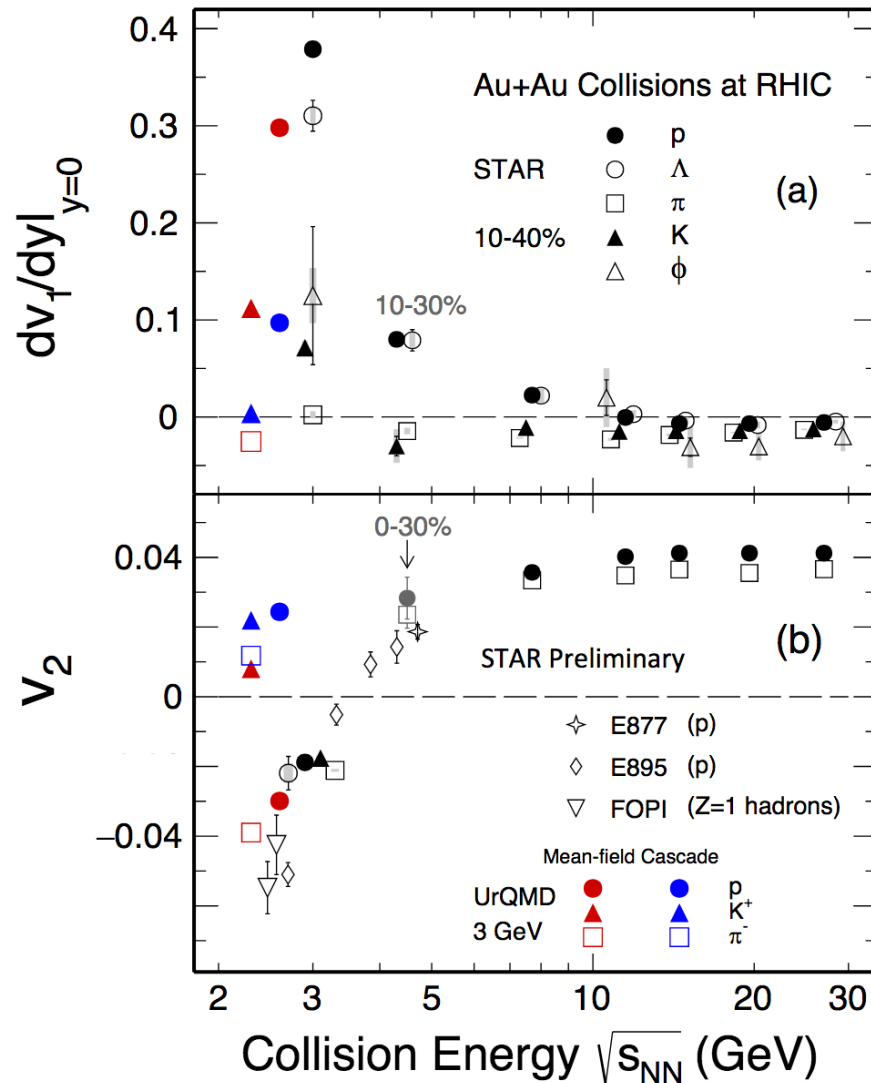
World data: *Phys. Lett. B* 778, 403-407, *Phys. Rev. C* 80.025209;
Phys. Rev. C 69.054901;
Phys. Rev. C 78, 044907; *Phys. Rev. C* 77,
 024903, *Phys. Rev. C* 66, 054902

Models: *Nucl. Phys. A* 772, 167; *Phys. Lett. B*
 603, 146
J. Phys. G: Nucl. Part. Phys. 43 (2016) 015104
Phys. Rev. C 99, 064908

G. Xie, S. Radhakrishnan, APS April Meeting, 2021

- Low energies, small systems: local strangeness conservation
- Canonical instead of Grand Canonical Ensemble describe statistical production → reduced phase space → “Canonical suppression”
- Data favors a CE calculation with a correlation length $r_c = 3.2$ fm in 0-10%
- Data strongly disfavors GCE, results $\sim 5\sigma$ away from zero (for 0-10%)

v_1 and v_2 in Au+Au $\sqrt{s_{NN}}=3.0$ GeV



S. Radhakrishnan, APS April Meeting, 2021

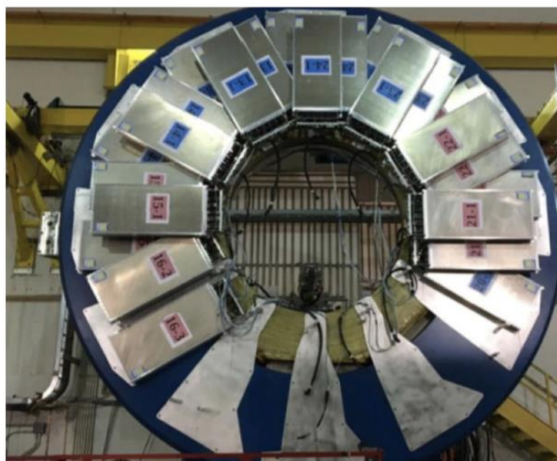
- Negative v_1 slope and large positive v_2 at high energy collisions
- Positive v_1 slope and negative v_2 for all measured particles in 3 GeV collisions
- Positive v_1 slope observed for kaons and ϕ mesons for the first time
- Results from UrQMD with baryonic mean-field interactions qualitatively describe the data

E877: *Phys. Rev. C* 56, 3254-3264

E895: *Phys. Rev. Lett.* 85, 940

FOPI: *Phys. Lett. B* 612, 173

STAR detector upgrades for BES-II



iTPC:

- Improves dE/dx
- Extends η coverage from 1.0 to 1.5
- Lowers p_T cut-in from 125 MeV/c to 60 MeV/c
- Ready in 2019

EndCap TOF:

- Forward rapidity coverage is critical
- PID at $\eta = 0.9$ to 1.5
- Improves the fixed target program
- Provided by CBM-FAIR
- Ready in 2019

EPD:

- Improves trigger
 - Reduces background
 - Allows a better centrality and reaction plane measurement
- Ready in 2018

iTPC: <https://drupal.star.bnl.gov/STAR/starnotes/public/sn0619>

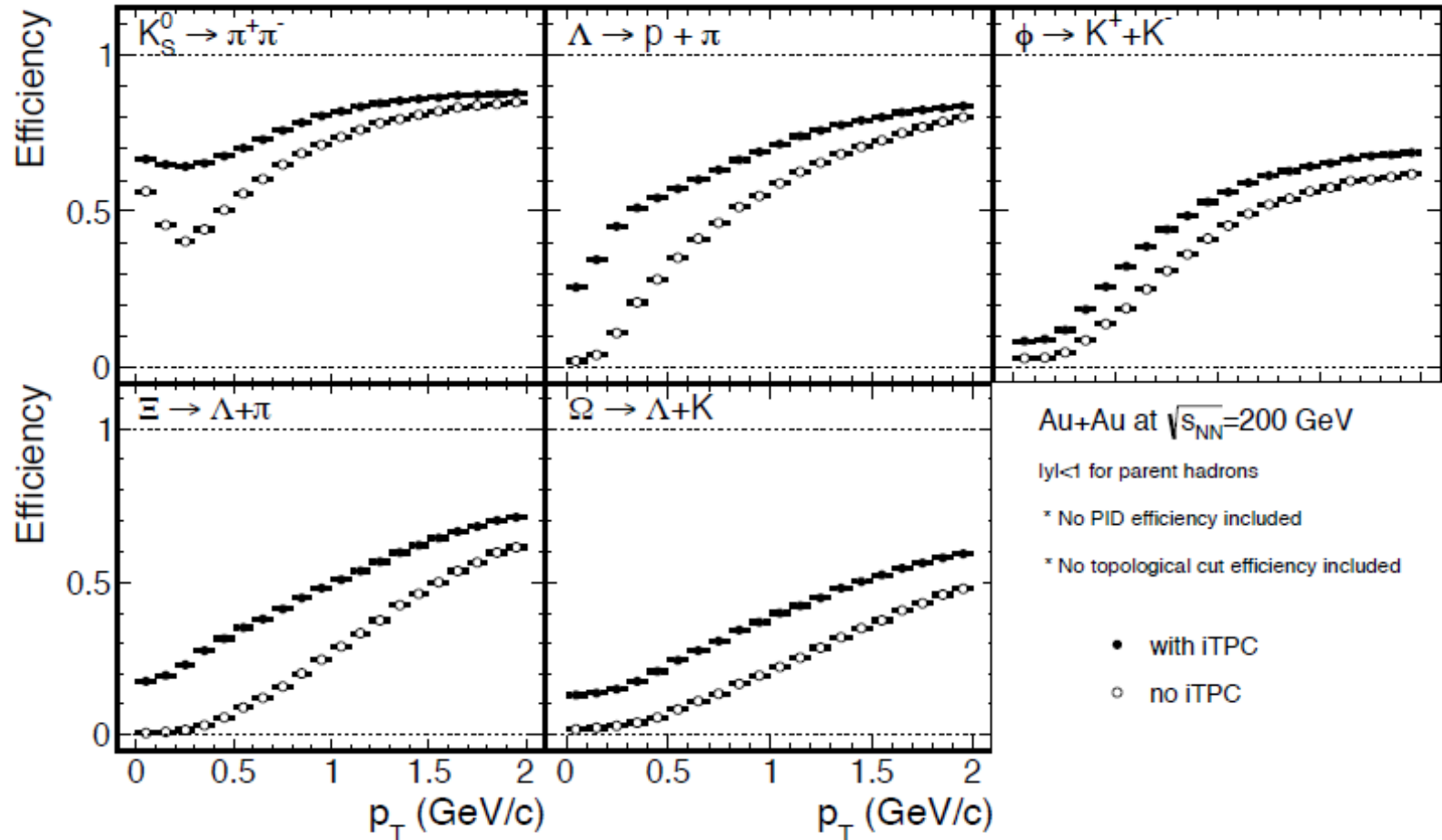
eTOF: STAR and CBM eTOF group, arXiv: 1609.05102

EPD: J. Adams, et al. Nucl. Instr. Meth. A 968, 163970 (2020)

- 1) **Enlarge rapidity acceptance**
- 2) **Improve particle identification**
- 3) **Enhance centrality/EP resolution**

All 3 detectors fully installed before Run-19.

iTPC improves strangeness reconstruction in BES II



- Significant improvement of efficiency especially for Ξ , Ω

H. Masui, A. Schmah / LBNL

STAR BES-II data samples

Year	<i>Collisions</i>	$\sqrt{s_{NN}}$ (GeV)	Good events
2019	Au+Au	19.6	~ 582 M
2019	Au+Au	14.5	~ 324 M
2020	Au+Au	11.5	~ 235 M
2020	Au+Au	9.2	~ 162 M
2021	Au+Au	7.7	~ 100 M (goal)
2021	Au+Au	17.1	~ 250 M (goal)

Data taking goes smoothly and reached the goal of statistics.

Summary & outlook

- STAR BES-I have measured systematically the production of strangeness at intermediate baryon density.
- Many structures are visible in strangeness related observables in this energy range.
- In particular, QGP signatures appear to turn off at lower collision energies ($<20\text{GeV}$), but need more statistics to confirm
- The almost completed STAR BES-II with detector upgrade (iTPC, eTOF, EPD) and larger data samples allow **precise** measurement of the matter properties at intermediate to high baryon density (μ_B up to 721 MeV)