

A no-lose theorem for discovering the new physics of $(g - 2)_\mu$

Brookhaven Forum (Virtual)
4 Nov 2021

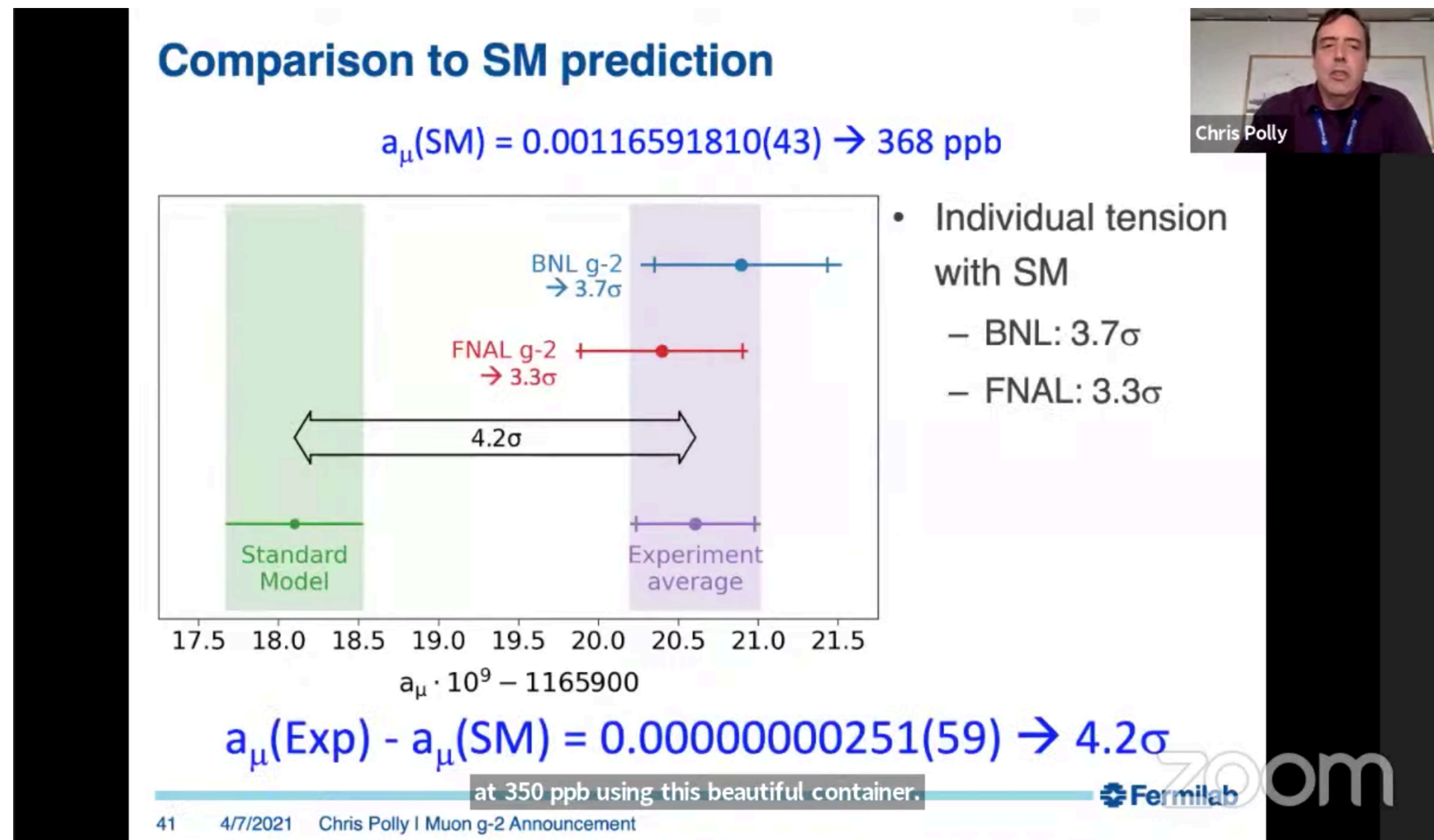
David Curtin
University of Toronto

based on 2006.16277, 2101.10334 with
Rodolfo Capdevilla, Yonatan Kahn, Gordan Krnjaic



Exciting!

7 April 2021:



Upshot:

BNL + Fermilab say

$$\Delta a_\mu = (2.51 \pm 0.59) \times 10^{-9}$$

which is disagreeing with SM prediction at 4.2σ level.

If real:

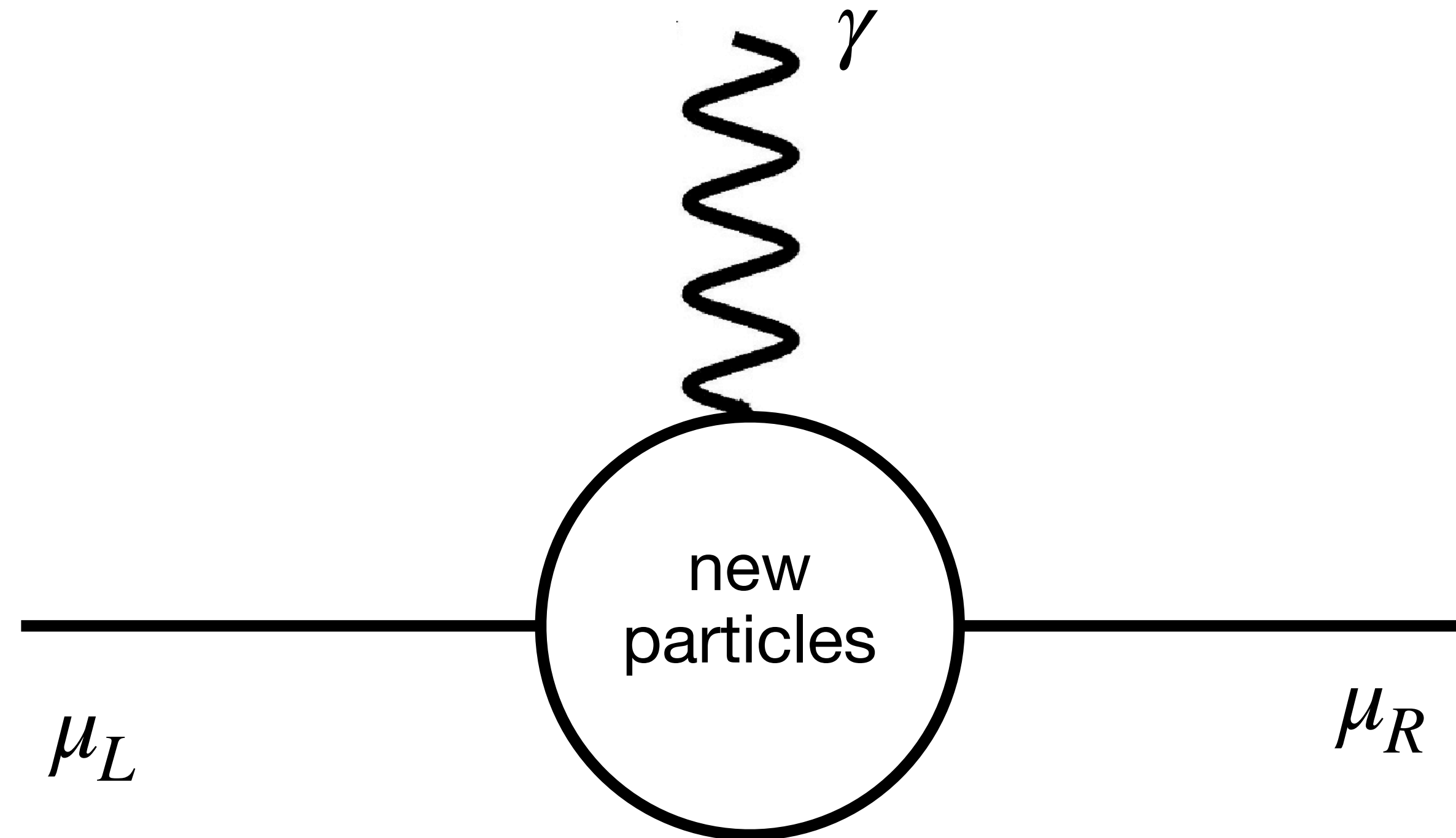
BSM physics talks to the muon!

→ Muon physics program from ~ GeV - 10 TeV will find new physics!

2006.16277, 2101.10334

BSM Physics in $(g - 2)_\mu$

Very simple:



$\gtrsim O(10^3)$ papers
over past decades

Could be almost anything, as long as it couples to muons

Could be connected to dark matter, SUSY, axions, any other new physics motivation...

This is such a general new physics contribution that it could be embedded within almost any BSM theory.

Ask a simpler question...

What would it take to *guarantee* we discover this new physics, *regardless* of the complete theory?

Model Exhaustive Approach

2006.16277, 2101.10334

Rodolfo Capdevilla, DC, Yonatan Kahn, Gordan Krnjaic

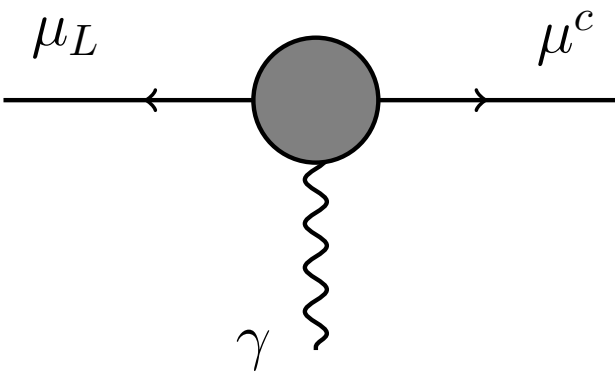
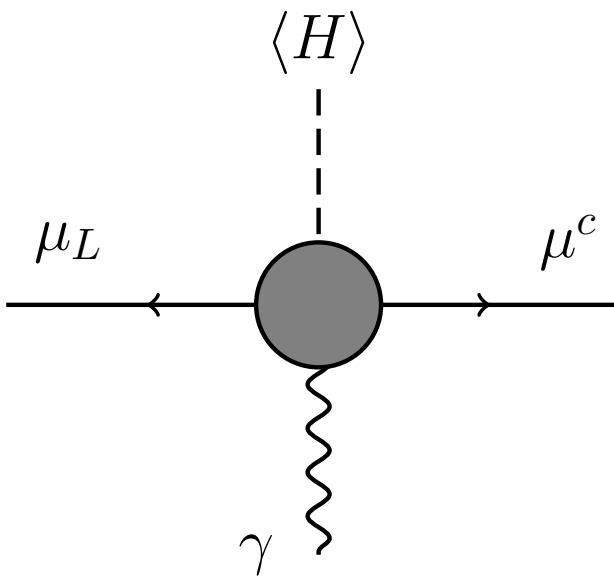
General BSM analysis of $(g - 2)_\mu$

| | | |
|-------------------------------|---|--|
| Assumptions | $\Delta a_\mu = a_\mu^{\text{obs}}$ $U(1)_{em}$ gauge invariance | $\Delta a_\mu = a_\mu^{\text{obs}}$ SM gauge invariance |
| $(g - 2)_\mu$ diagram | | |
| How to predict new signatures | $\frac{1}{M}(\mu_L \sigma^{\nu\rho} \mu^c) F_{\nu\rho}$ | $\frac{1}{M^2} H^\dagger (L \sigma^{\nu\rho} \mu^c) F_{\nu\rho}$ |

Model-Independent

EFT analysis suggests $M \lesssim 250 \text{ TeV} \dots$ Really?

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Model-Independent

We would love to discover this new physics **DIRECTLY**.

Where the new particles at??

EFT analysis suggests $M \lesssim 250$ TeV.... Really?

General BSM analysis of $(g - 2)_\mu$

| | | | |
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| Assumptions | $\Delta a_\mu = a_\mu^{\text{obs}}$ $U(1)_{em}$ gauge invariance | $\Delta a_\mu = a_\mu^{\text{obs}}$ SM gauge invariance | $\Delta a_\mu = a_\mu^{\text{obs}}$ SM gauge invariance Perturbativity |
| $(g - 2)_\mu$ diagram | | | |
| How to predict new signatures | $\frac{1}{M}(\mu_L \sigma^{\nu\rho} \mu^c) F_{\nu\rho}$ | $\frac{1}{M^2} H^\dagger (L \sigma^{\nu\rho} \mu^c) F_{\nu\rho}$ | Specific choices of BSM particles and their SM quantum numbers in loop |

Model-Independent

“Model-Exhaustive”

If we assume perturbative unitarity, we can look inside the 4-point function!

Can we do this in full generality?

Model-Exhaustive Analysis

Assume new physics obeys perturbative unitarity.*

Assume new $(g - 2)_\mu$ contribution arises at one-loop.**

Then consider:

- all possible $SU(2)_L \otimes U(1)_Y$ gauge representations for the new particles
- all possible Lorentz group representations*** for the new particles
- arbitrary multiplicity N_{BSM} of new particles
- all possible masses & couplings that generate Δa_μ^{exp}

Then ask: what are some irreducible experimental signatures?

*pushing couplings right up to unitarity limit should capture parametrics of non-perturbative solutions, they still have to obey gauge invariance.

** higher loop contributions require lower BSM mass scales, should be discoverable with the experiments we consider

*** Spin 0, 1/2 and 1. Higher spin g-2 contributions highly suppressed, 2104.03231.

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Divide BSM Theory Space into two classes

Singlet Scenarios:

new physics in $(g - 2)_\mu$ is **SM singlets only**

simple theory space, more complicated phenomenology

Electroweak Scenarios:

everything else: i.e. new particles with non-trivial EW representations in loop

complicated theory space, simple phenomenology (new charged particles!)

Model Exhaustive Approach

Singlet Scenarios

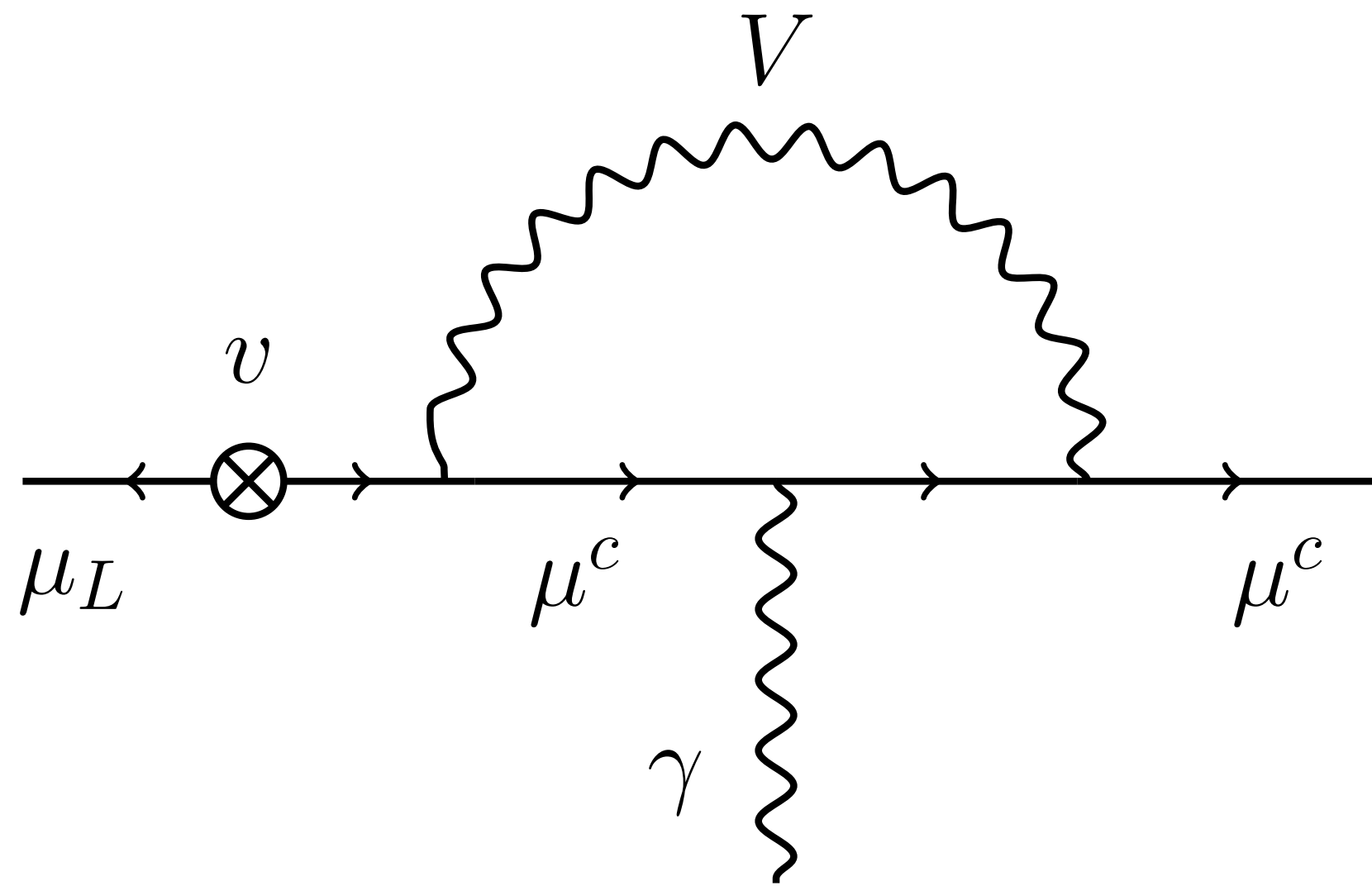
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Singlet Scenarios

Vector

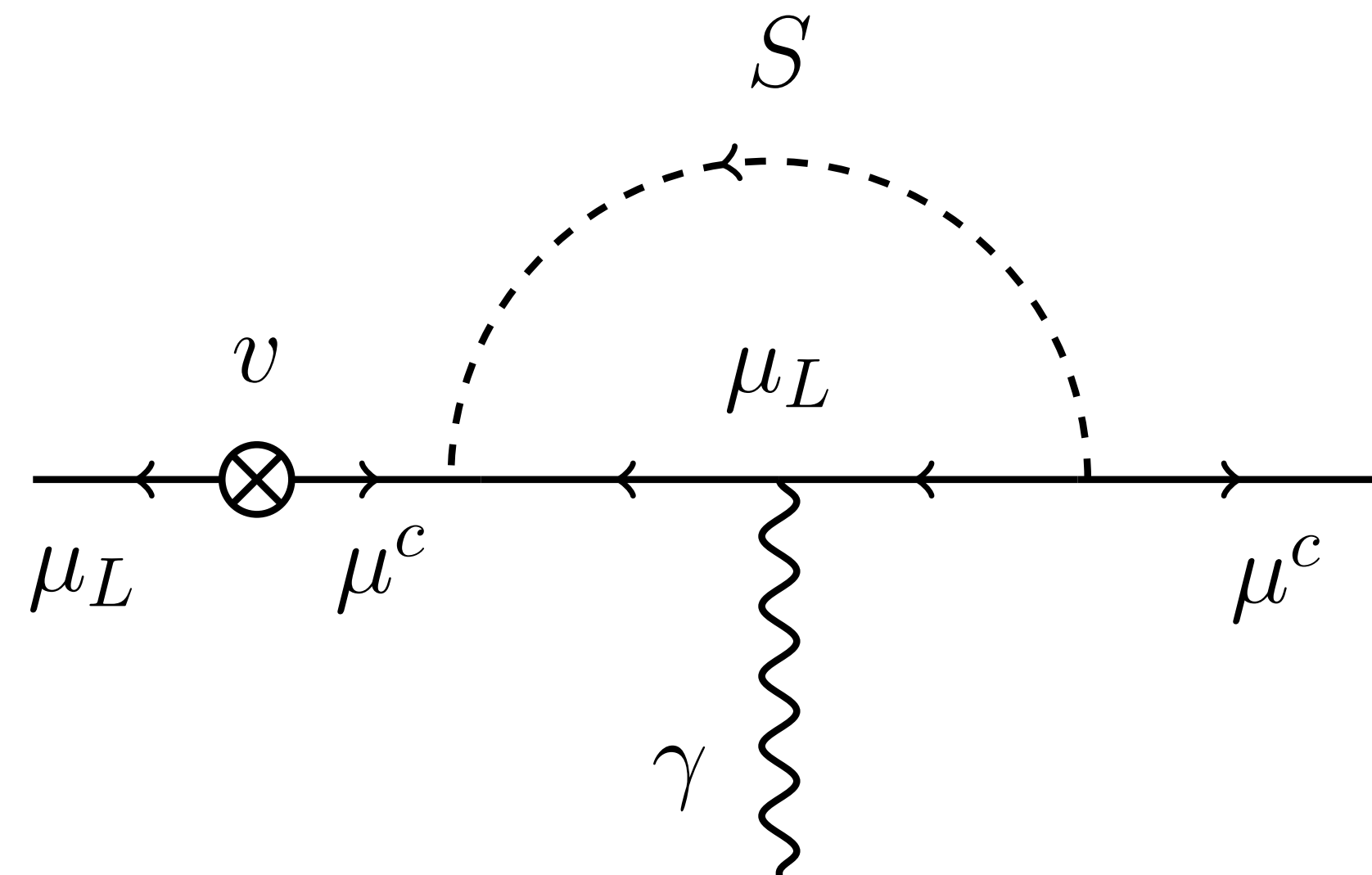
$$\mathcal{L}_V \supset g_V V_\alpha (\mu_L^\dagger \bar{\sigma}^\alpha \mu_L + \mu^{c\dagger} \bar{\sigma}^\alpha \mu^c) + \frac{m_V^2}{2} V_\alpha V^\alpha$$



$$\Delta a_\mu^V / \Delta a_\mu \approx N_{\text{BSM}} g_V^2 \left(\frac{200 \text{ GeV}}{m_V} \right)^2$$

Scalar

$$\mathcal{L}_S \supset - (g_S S \mu_L \mu^c + \text{h.c.}) - \frac{1}{2} m_S^2 S^2$$

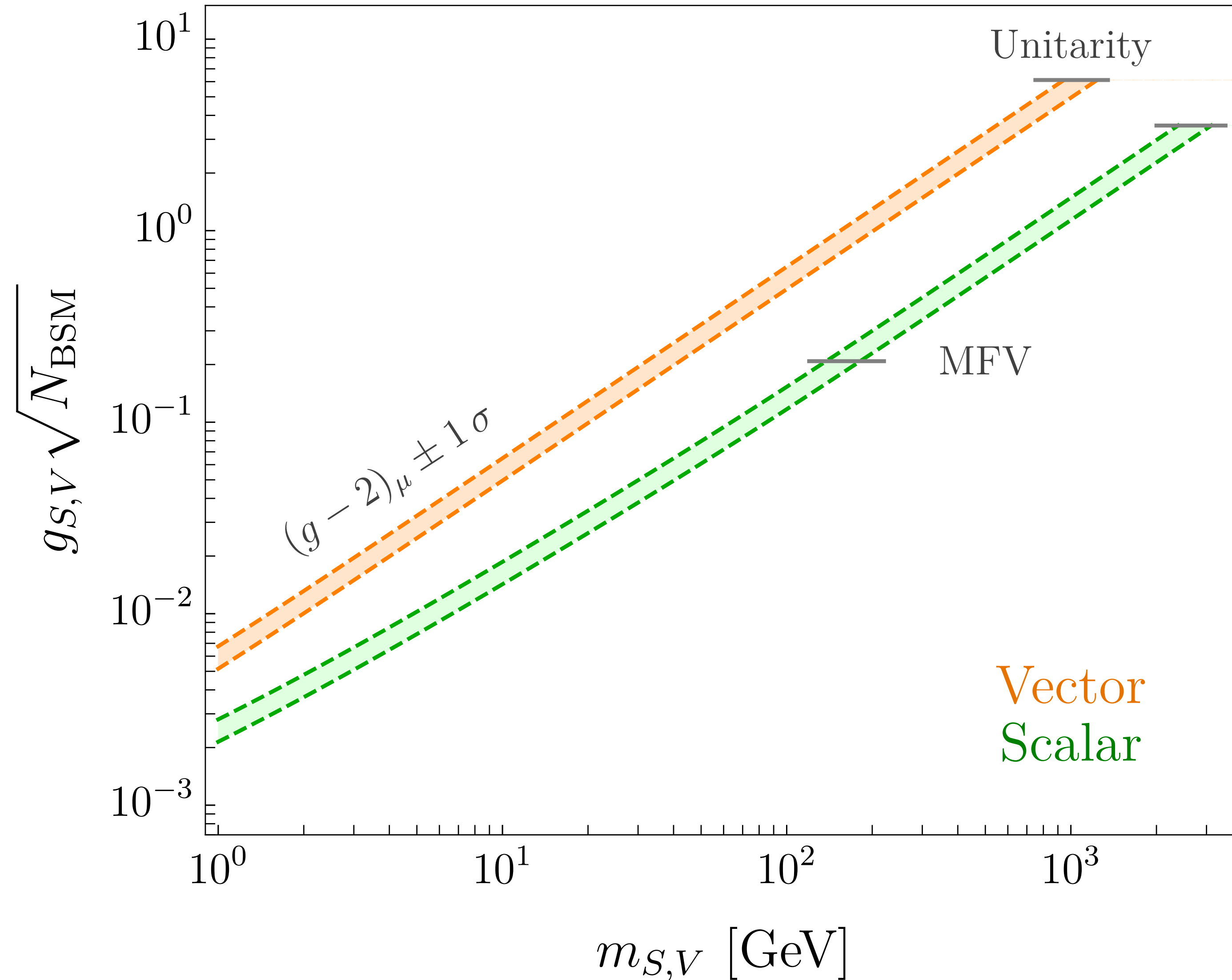


$$\Delta a_\mu^S / \Delta a_\mu \approx N_{\text{BSM}} g_S^2 \left(\frac{700 \text{ GeV}}{m_S} \right)^2$$

$(g - 2)_\mu$ contributions from RHN-type singlet fermion is suppressed by m_ν and too small.

Interesting edge case: $a F_{\mu\nu} \tilde{F}_D^{\mu\nu}$ axion-dark-photon contribution (2104.03276), but is also discoverable.

Singlet Scenarios



Requires singlet below 3 TeV

couples to muon $g_S \propto m_S$

Model Exhaustive Approach

Electroweak Scenarios

2006.16277, 2101.10334

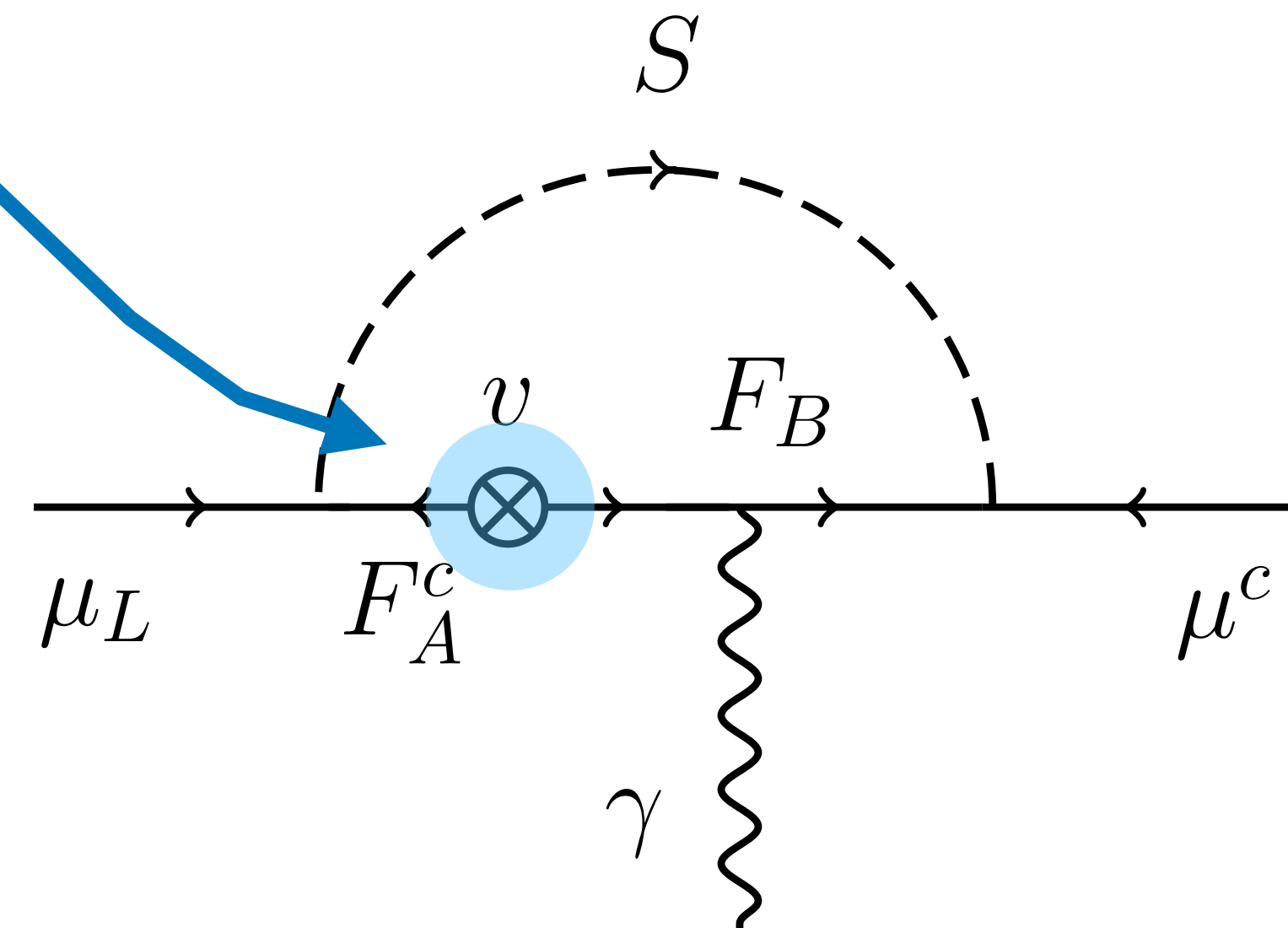
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Electroweak Scenarios

In general a complicated model space: all non-singlet one-loop possibilities!

Can generate $(g - 2)_\mu$ for much higher BSM masses due to **large Higgs vev / chirality flip**

$$\Delta a_\mu \propto \frac{m_\mu v g_{BSM}^3}{m_{BSM}^2}$$



But perhaps the experimental signatures are simpler: **new charged particles!**

New Charged Particles

Those are the “easiest to discover”:

- guaranteed Drell-Yan Production
- have to leave some visible signal in your detector

Main question: how much collider energy \sqrt{s} do I need to produce at least the **lightest** BSM charged state?

New Charged Particles

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Main question: how much collider energy \sqrt{s} do I need to produce at least the **lightest** BSM charged state?

$$M_{\text{BSM,charged}}^{\text{max}} \equiv \max_{\substack{\text{BSM theory space} \\ \Delta a_{\mu} = \Delta a_{\mu}^{\text{obs}}}} \left\{ \min_{i \in \text{BSM spectrum}} \left(m_{\text{charged}}^{(i)} \right) \right\}$$

We can brute-force min-max across all the models' parameter space to find the **highest possible mass the lightest BSM charged state can have to be consistent with g-2.**

Electroweak Simplified Models

Model-exhaustive analyses are not a new idea, but this **theory space maximization** to find the **largest possible BSM charged mass** is non-trivial.

We will define some **simplified models** which are engineered to produce **the heaviest possible BSM charged masses** while explaining $(g-2)$!

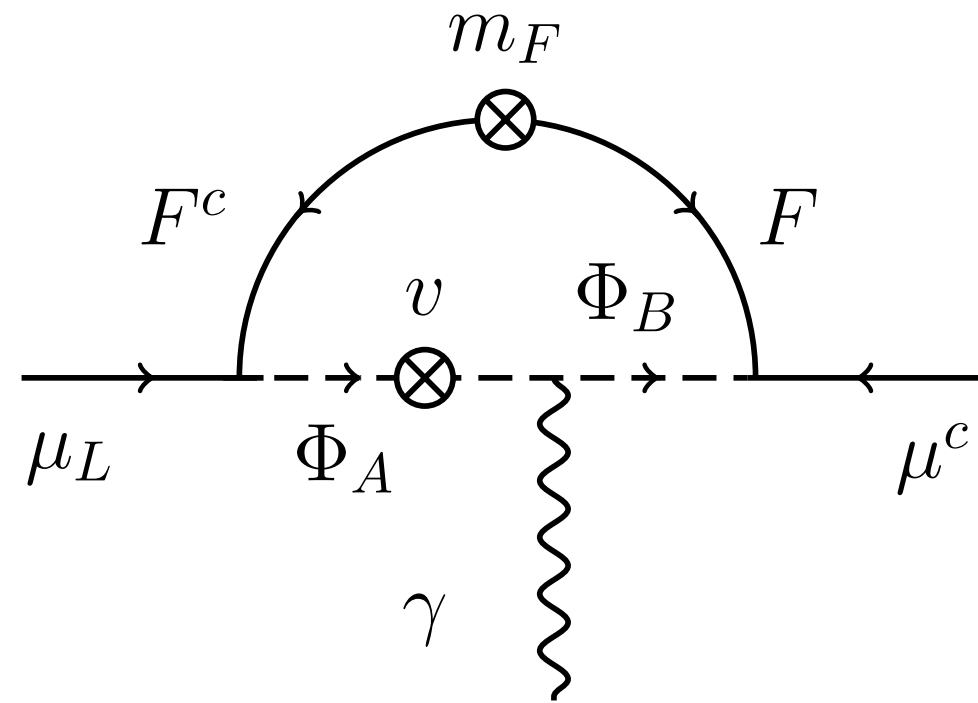
Maximizing over the space of those simplified models will give us our answer!

Engineering specs:

- need BSM (i.e. large) **chiral flip insertion**
- need BSM (i.e. large) **Higgs vev insertion**
- **need three new fields** (boson, fermion, and two of something)
- **no new sources of EWSB** (those have their own lower-mass signatures)

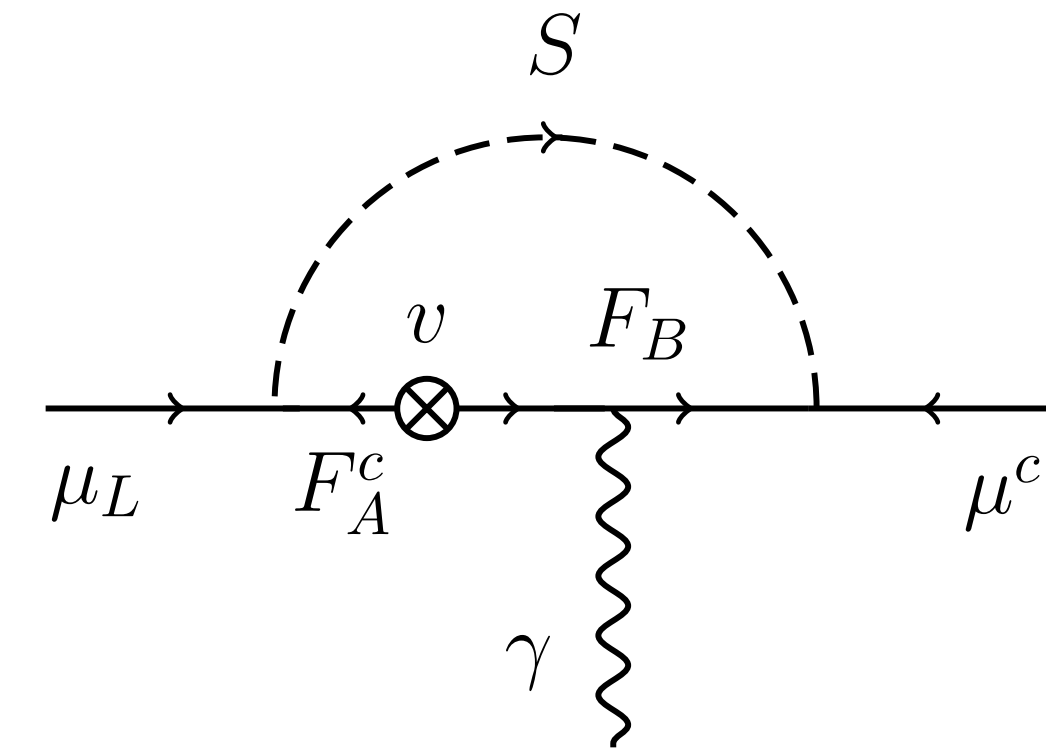
Electroweak Simplified Models

SSF



$$\mathcal{L}_{\text{SSF}} \supset -y_1 F^c L_{(\mu)} \Phi_A^* - y_2 F \mu^c \Phi_B - \kappa H \Phi_A^* \Phi_B - m_A^2 |\Phi_A|^2 - m_B^2 |\Phi_B|^2 - m_F F F^c + \text{h.c.}$$

FFS



$$\mathcal{L}_{\text{FFS}} \supset -y_1 F_A^c L_{(\mu)} \Phi^* - y_2 F_B \mu^c \Phi - y_{12} H F_A^c F_B - y'_{12} H^\dagger F_A F_B^c - m_A F_A F_A^c - m_B F_B F_B^c - m_S^2 |\Phi|^2 + \text{h.c.}$$

New complex scalars and vector-like fermions that acquire some mixing after EWSB.

Consider all possible choices of $SU(2)_L \otimes U(1)_Y$ representations $R \leq 3, Q \leq 2$.

Arbitrary number of BSM degrees of freedom (copies) N_{BSM} .

We have checked that other simplified models with fewer BSM fields, or involving Majorana fermions, new vectors, etc give smaller $\Delta a_\mu \rightarrow$ lower masses for new charged states \rightarrow do not affect theory space maximization

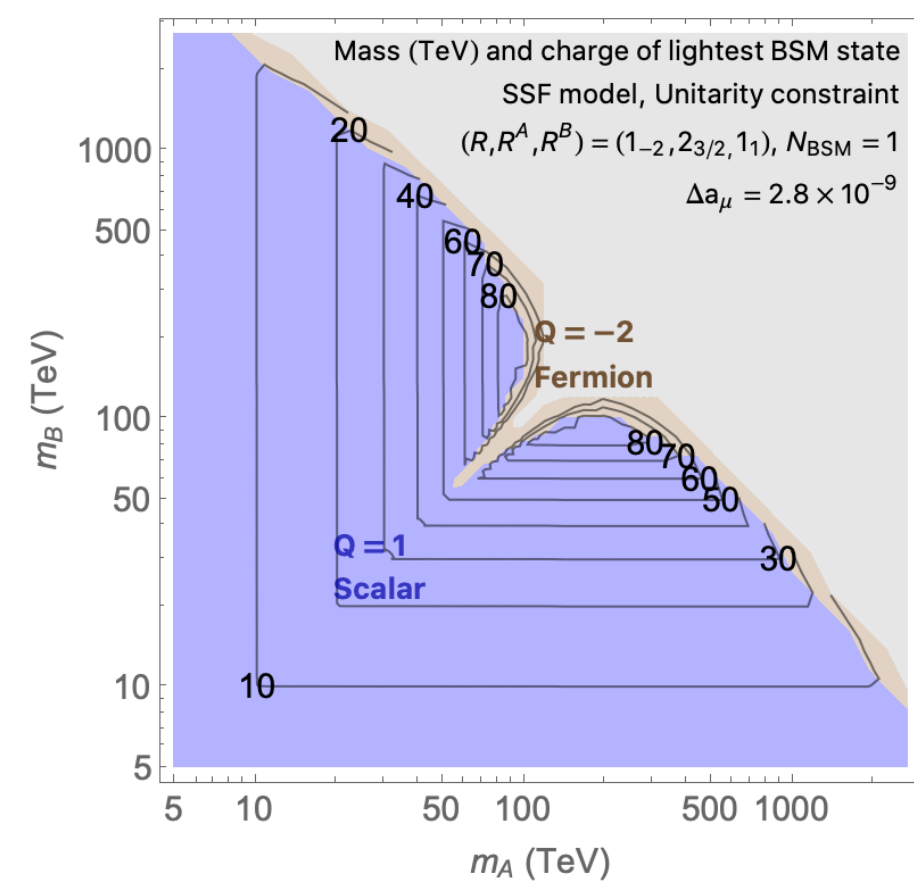
What's the result?

$$M_{\text{BSM,charged}}^{\text{max}} \equiv \max_{\text{BSM theory space}} \left\{ \min_{i \in \text{BSM spectrum}} \left(m_{\text{charged}}^{(i)} \right) \right\}$$

$$\Delta a_{\mu} = \Delta a_{\mu}^{\text{obs}}$$

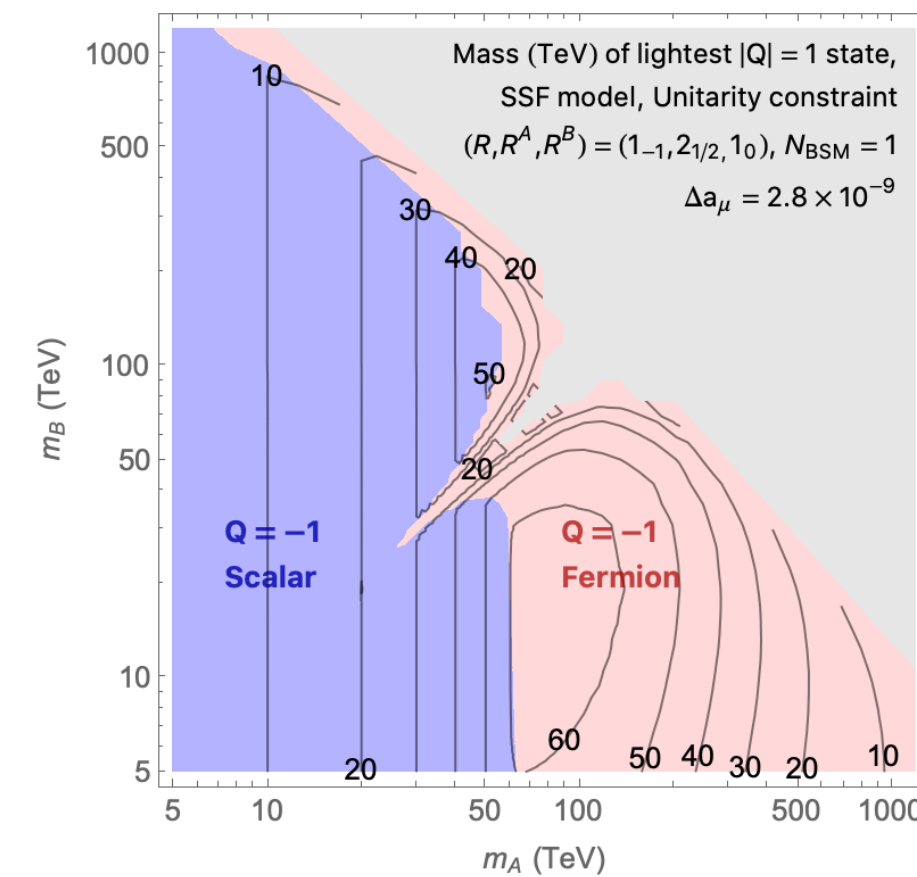
Example of parameter space plot for two EW models, showing lightest BSM charged particle mass with unitarity constraints only

SSF, all BSM fields charged



$$M_{\text{BSM,charged}}^{\text{max,unitarity}} = 86 \text{ TeV}$$

SSF, charged and neutral fields



$$M_{\text{BSM,charged}}^{\text{max,unitarity}} = 65 \text{ TeV}$$

EW Model Results

$$M_{\text{BSM,charged}}^{\text{max,X}} \approx \left(\frac{2.8 \times 10^{-9}}{\Delta a_{\mu}^{\text{obs}}} \right)^{\frac{1}{2}} \times \begin{cases} (100 \text{ TeV}) N_{\text{BSM}}^{1/2} & \text{for } X = (\text{unitarity}^*), \\ (20 \text{ TeV}) N_{\text{BSM}}^{1/2} & \text{for } X = (\text{unitarity} + \text{MFV}), \\ (20 \text{ TeV}) N_{\text{BSM}}^{1/6} & \text{for } X = (\text{unitarity} + \text{naturalness}^*), \\ (9 \text{ TeV}) N_{\text{BSM}}^{1/6} & \text{for } X = (\text{unitarity} + \text{naturalness} + \text{MFV}). \end{cases}$$

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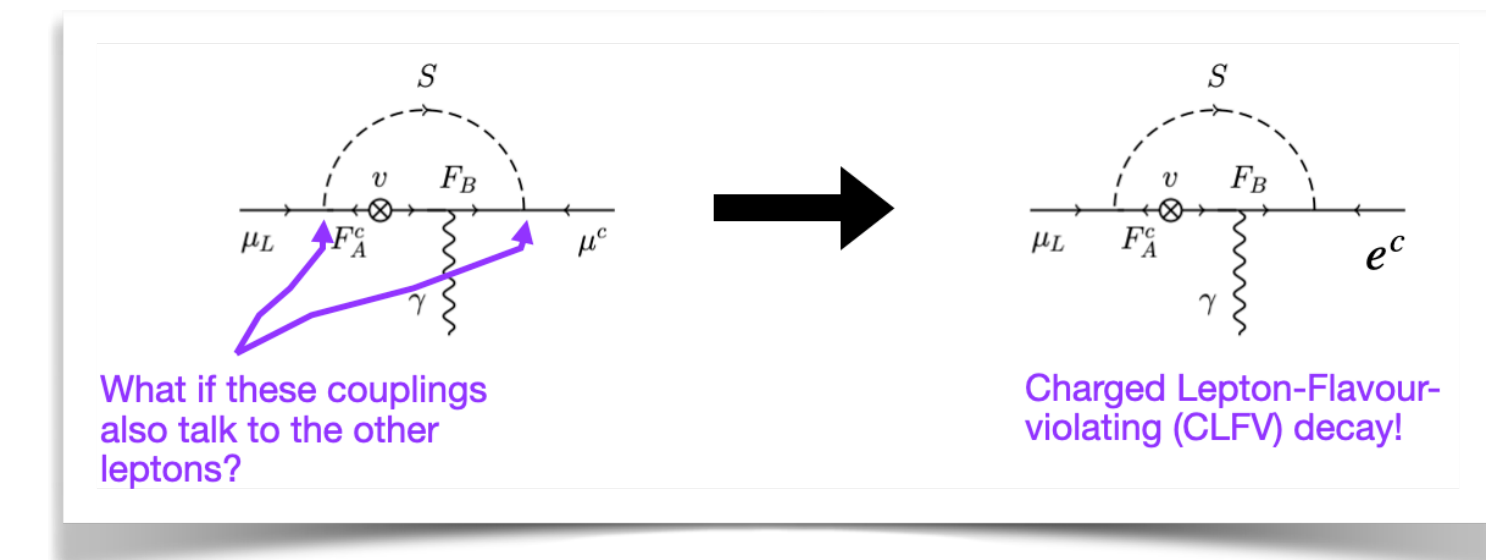
Imposing only unitarity constraints on all couplings.

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Imposing MFV on couplings to avoid CLFV bounds.



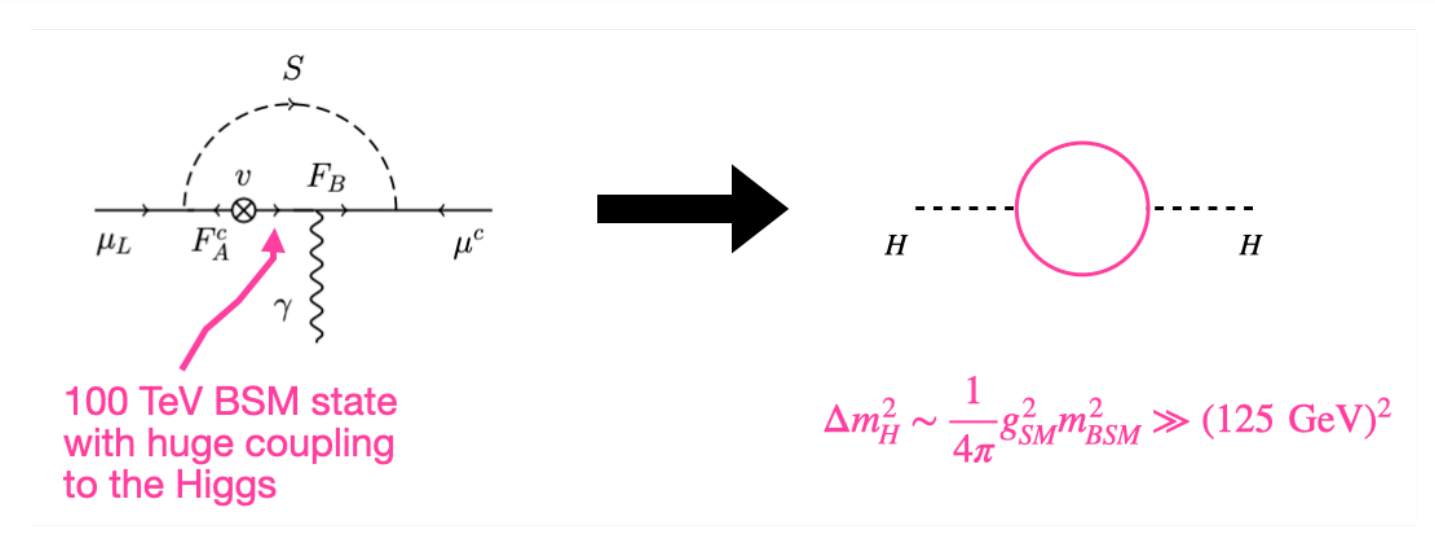
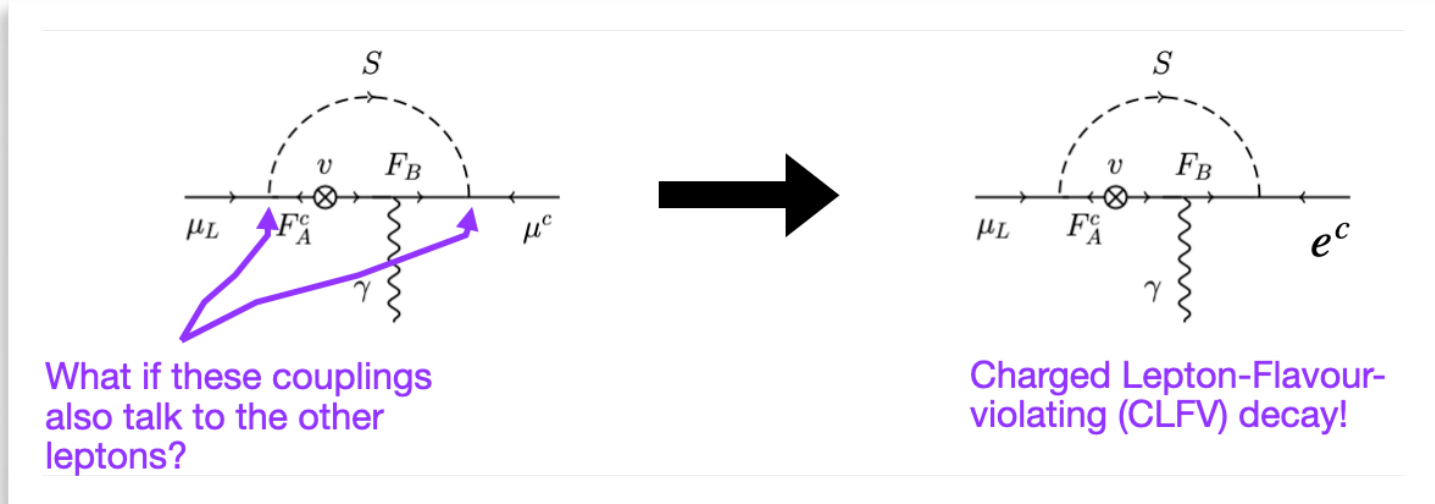
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Imposing $\Delta < 100$ Naturalness constraint, since BSM particles generate **calculable and large** corrections to Higgs & μ mass.



EW Model Results

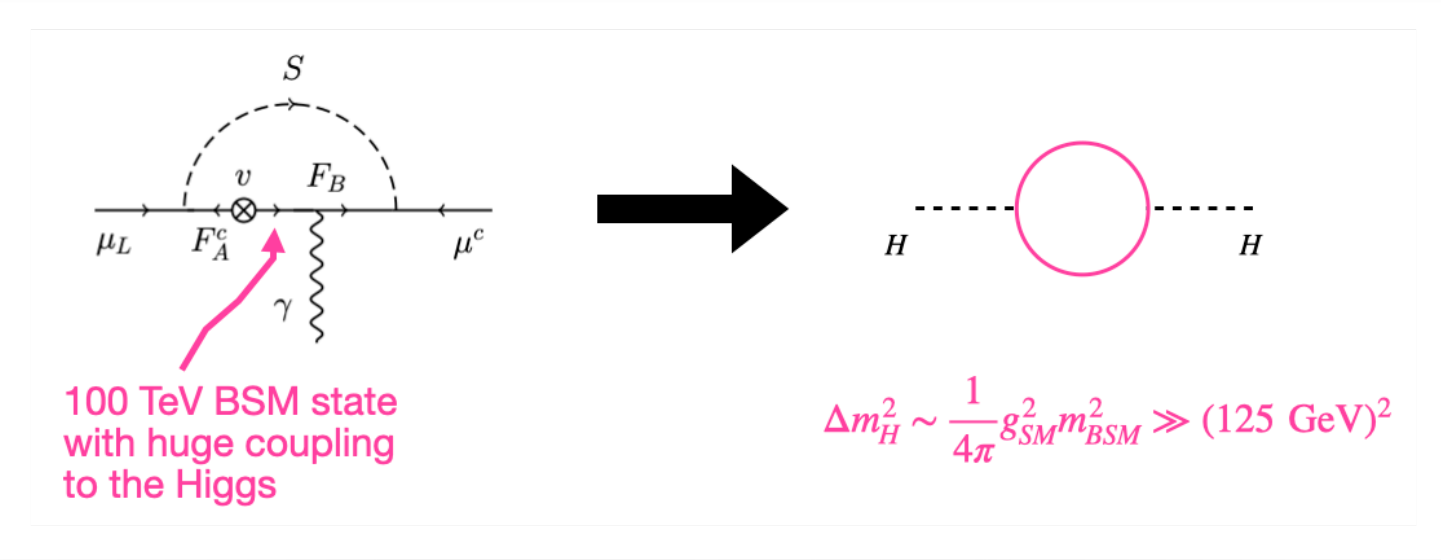
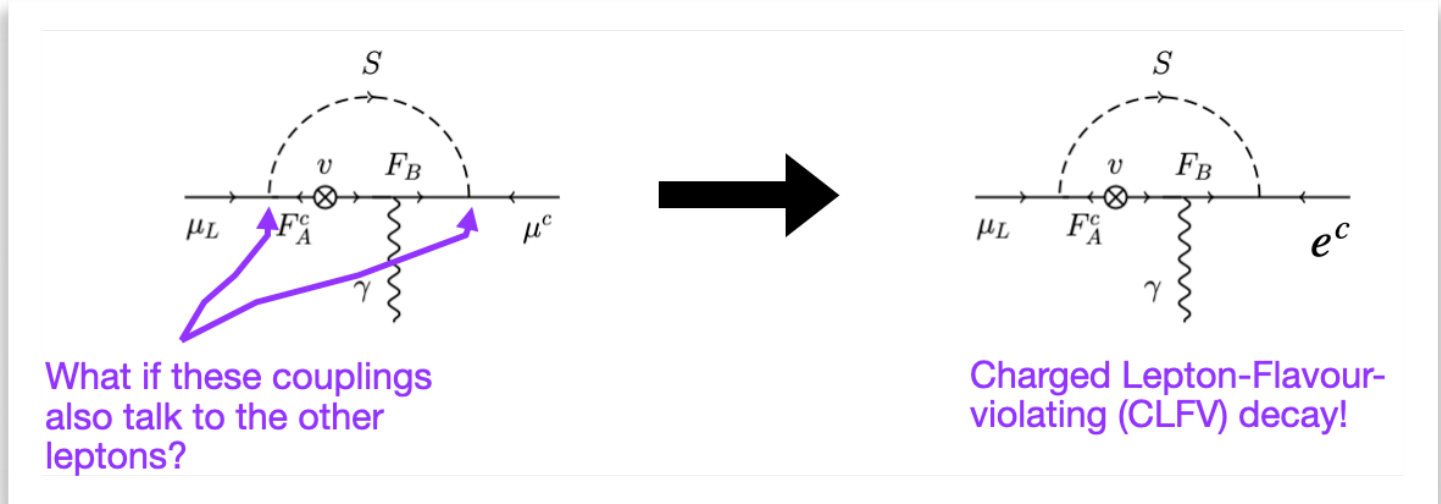
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Imposing only unitarity constraints on all couplings.

Imposing MFV on couplings to avoid CLFV bounds.

Imposing $\Delta < 100$ Naturalness constraint, since BSM particles generate **calculable and large** corrections to Higgs & μ mass.

MFV + naturalness: the most “reasonable” upper bound!



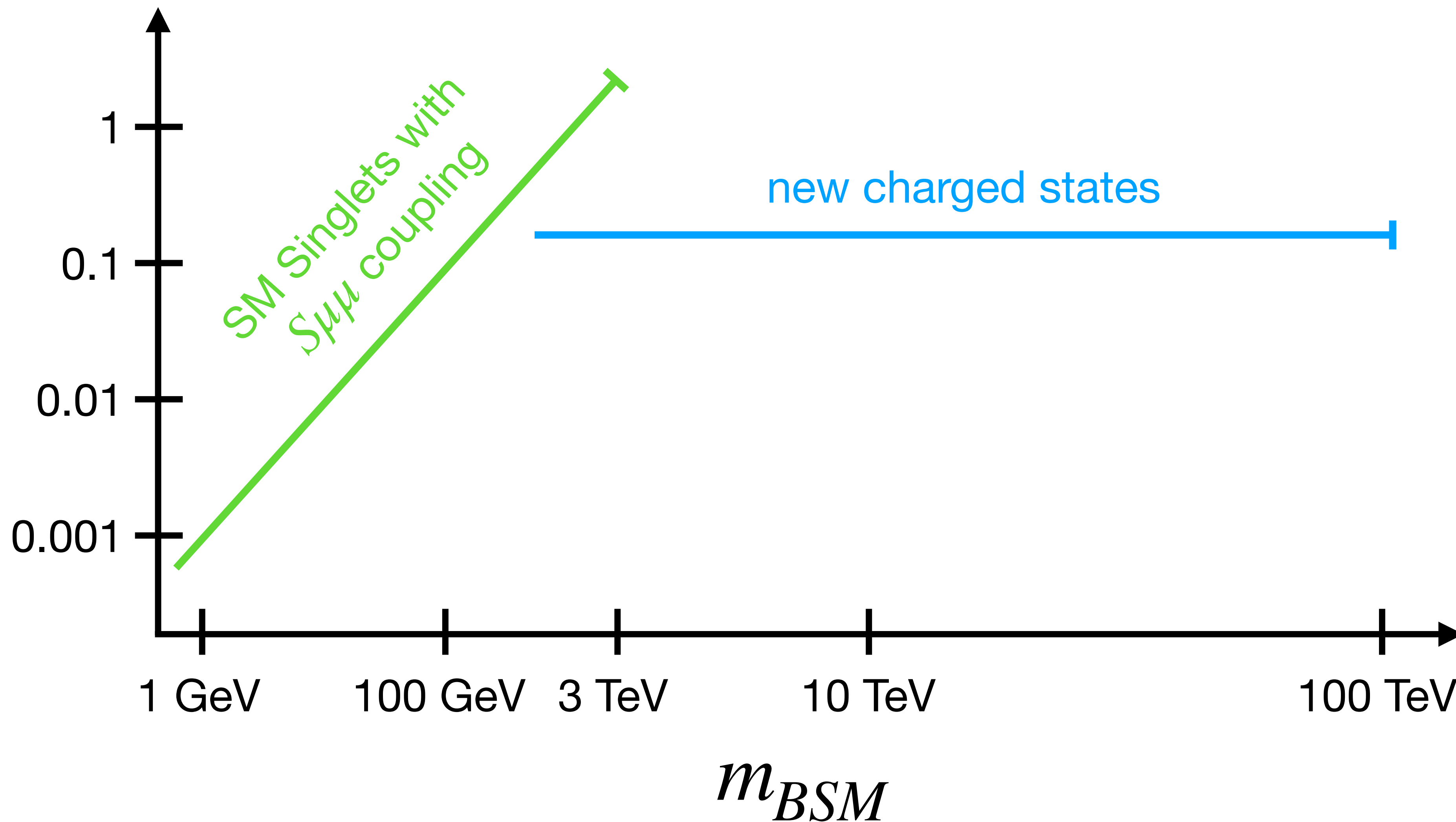
Experimental Target for discovering BSM

2006.16277, 2101.10334

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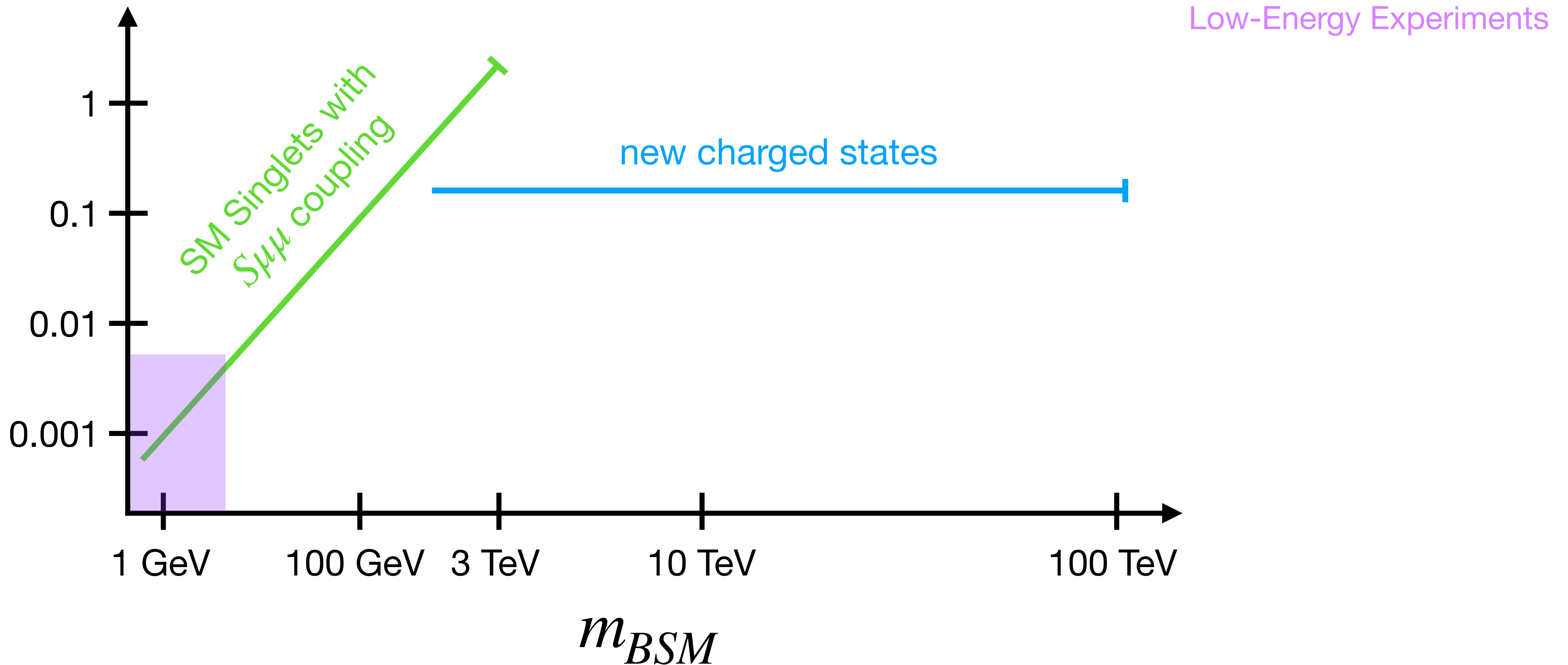
Irreducible Signatures of $(g - 2)_\mu$ Solutions

Relevant coupling



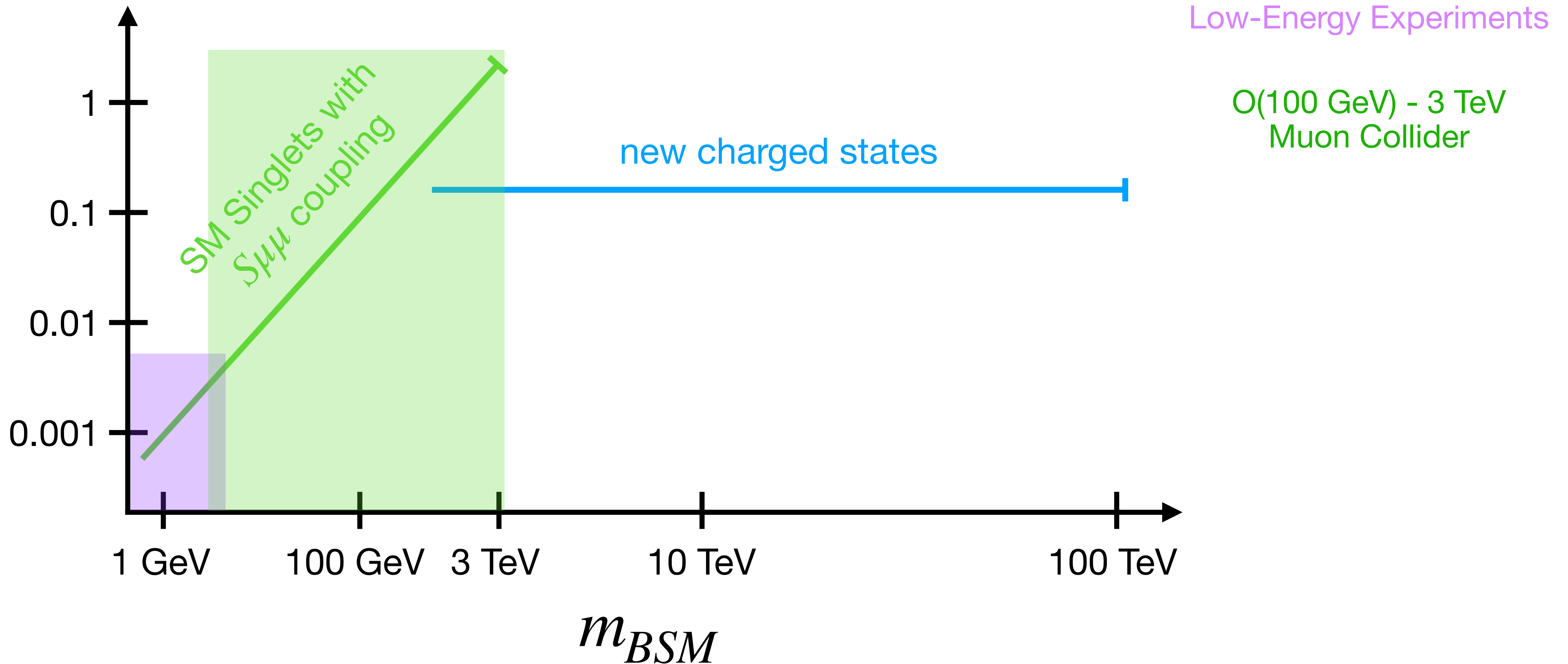
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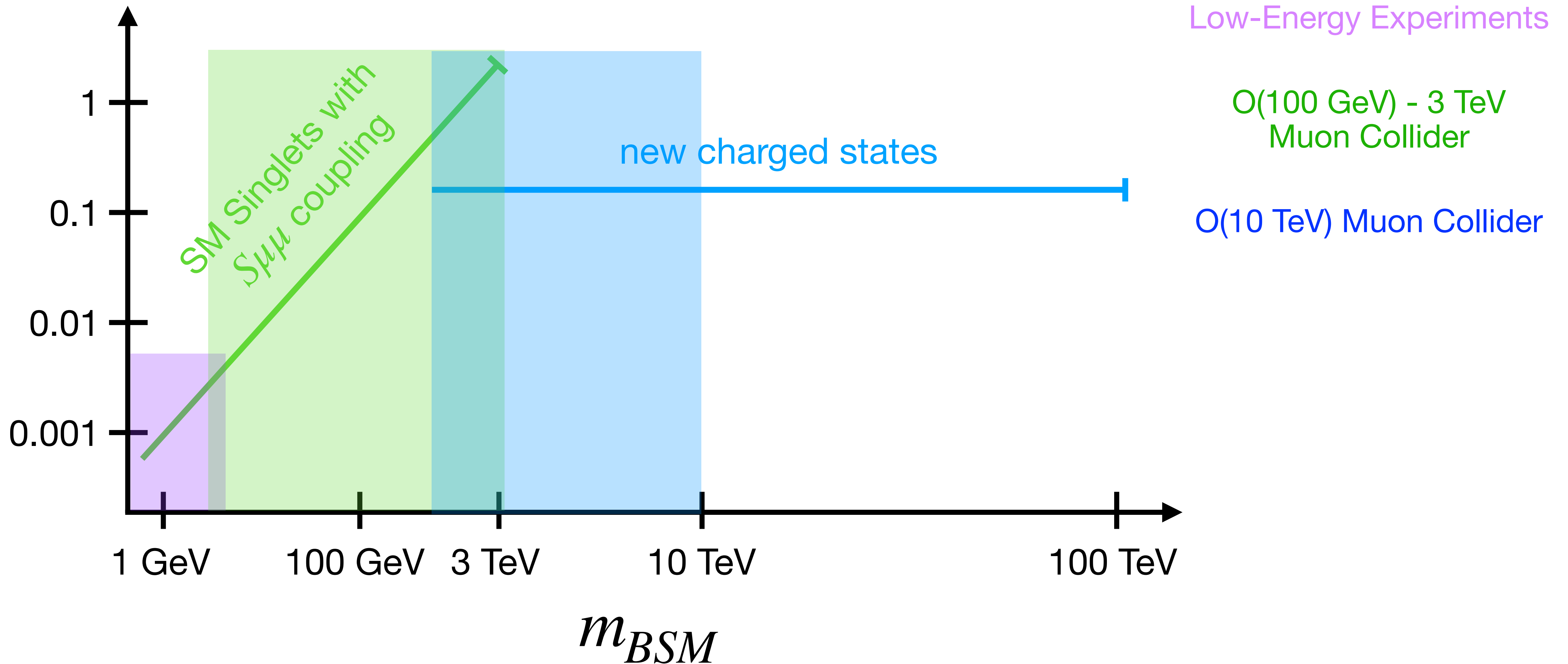
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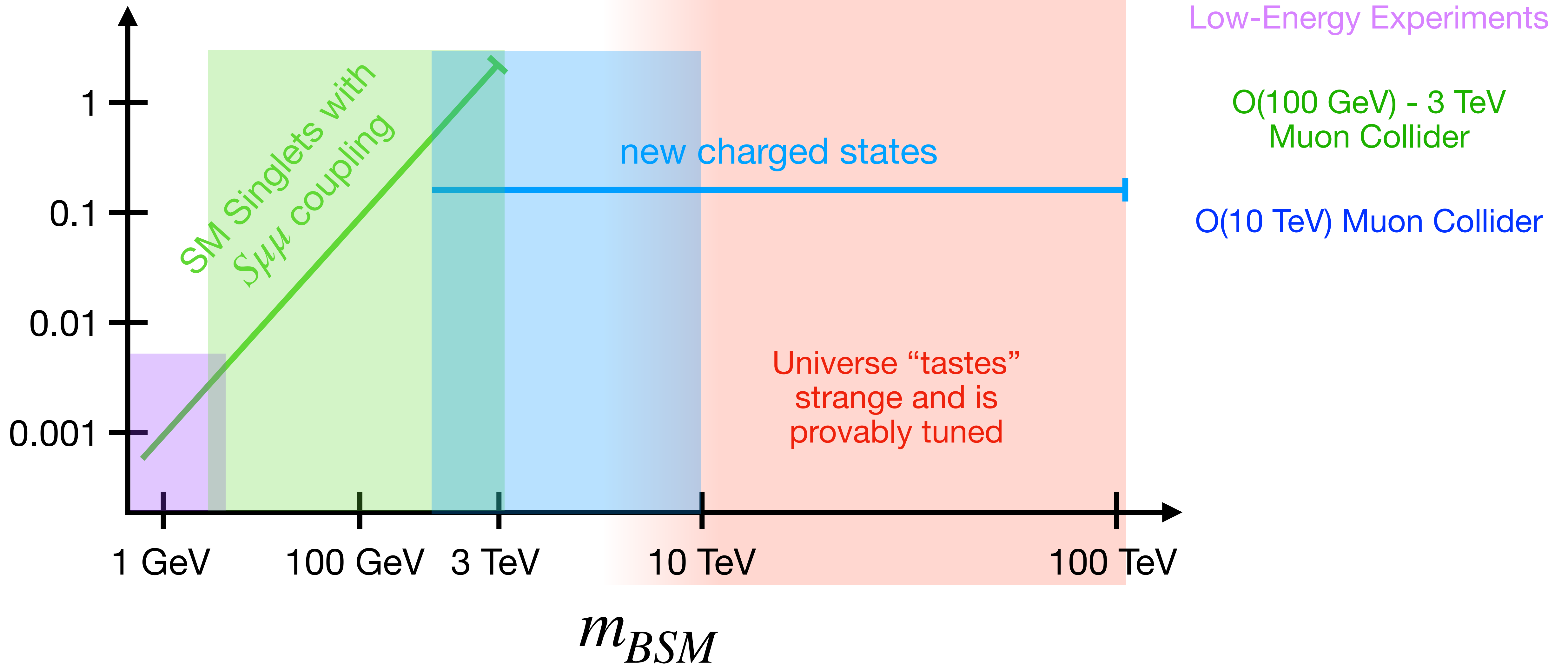
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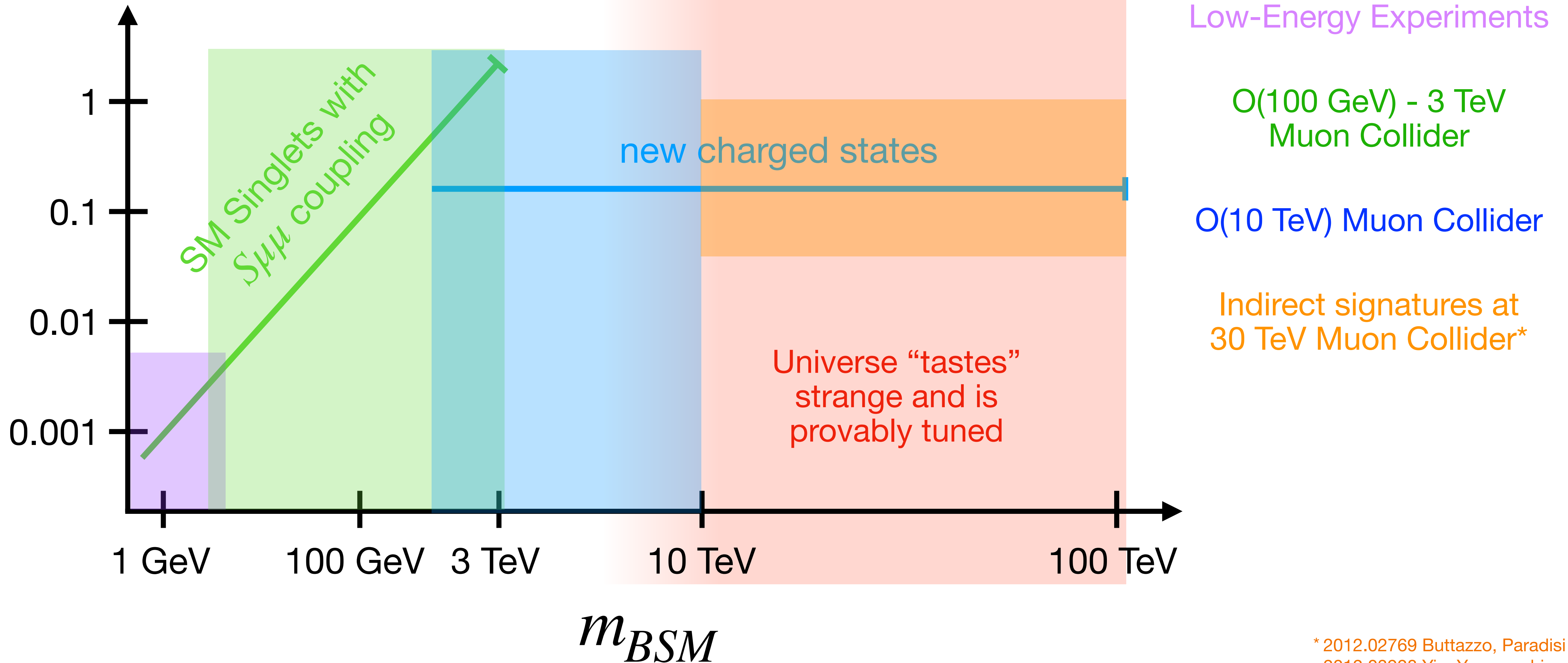
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* 2012.02769 Buttazzo, Paradisi
2012.03928 Yin, Yamaguchi

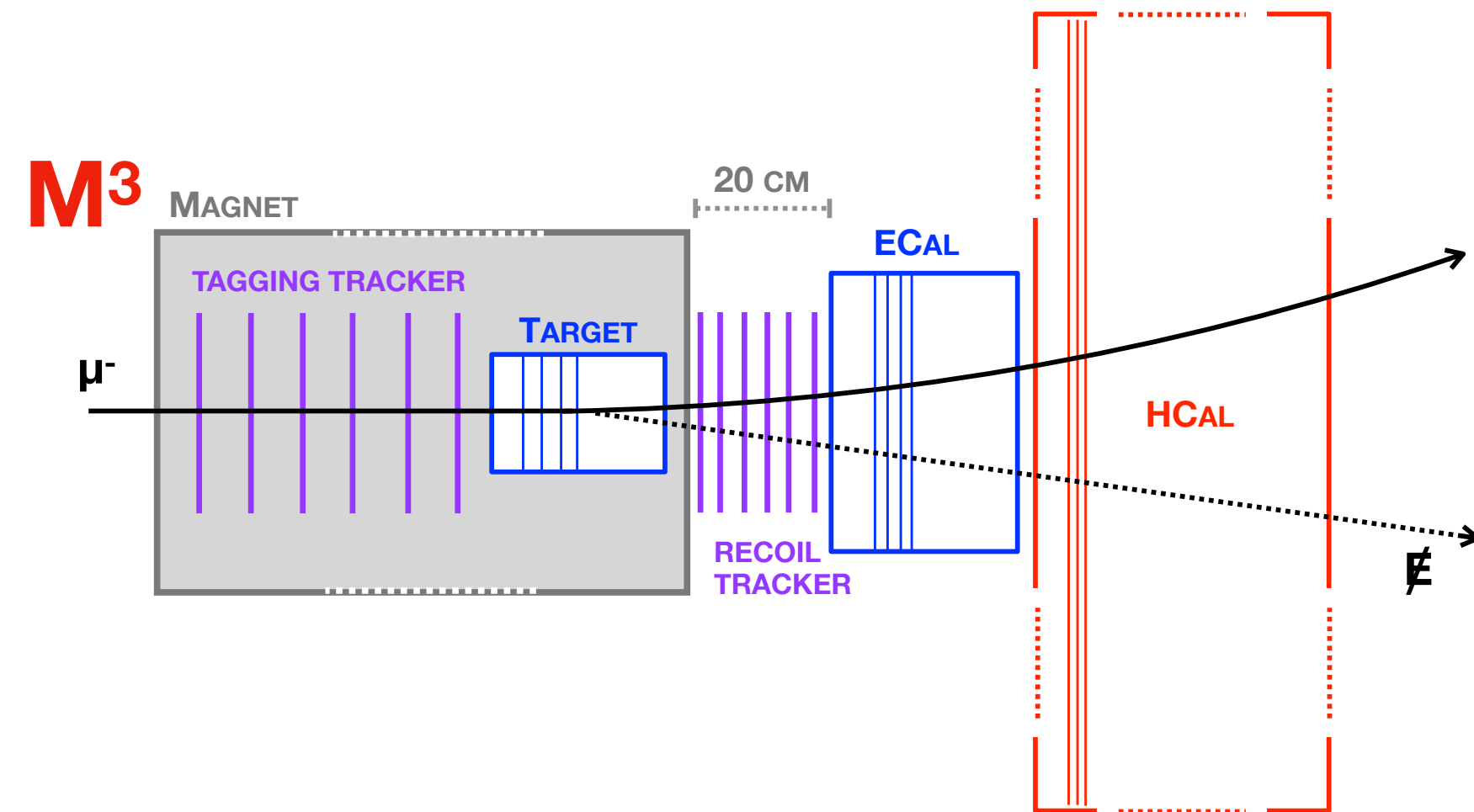
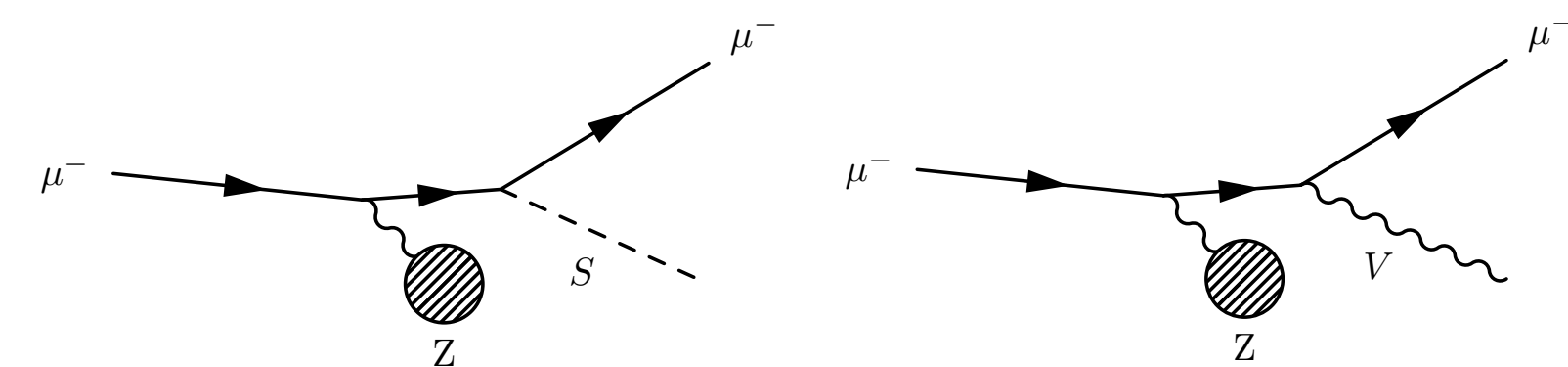
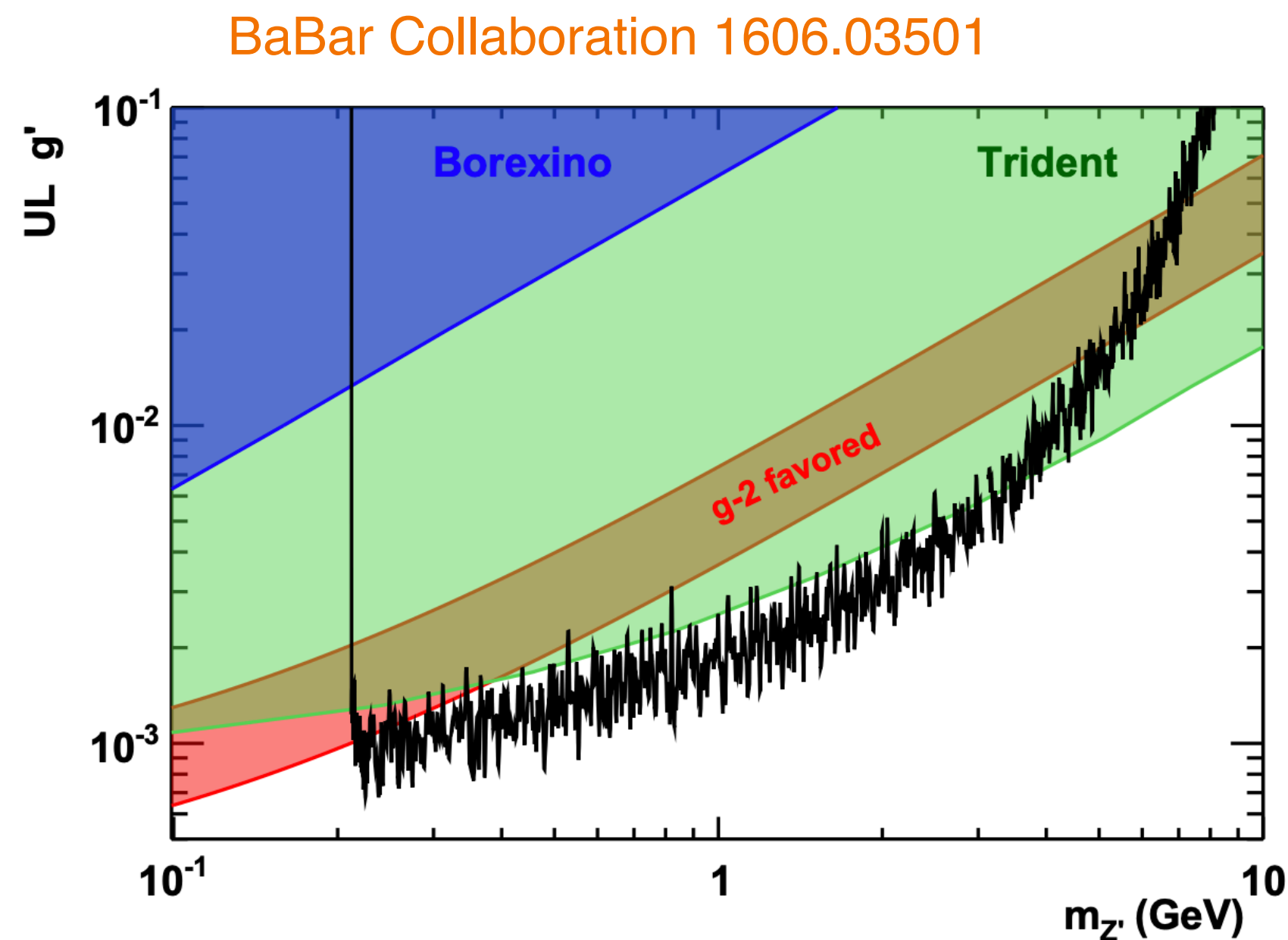
Low Energy Experiments

Intensity Frontier Experiments

A lot of Singlet Scenario parameter space is already excluded below a few GeV.

Remaining space can be fully covered by **Muon Fixed Target experiments:**

M³ proposal at Fermilab / NA64 μ at CERN



See also e.g.:

- Mohlabeng 1809.07768
- Dark Sector Community Report 1707.04591
- SHiP physics case 1504.04855
- Krnjaic 1512.04119
- Batell, Freitas, Ismail, McKeen 1712.10022
- Chen, Pospelov, Zhong, 1701.07437
- Bauer, Foldenauer, Jaeckel, 1803.05466

S.N. Gninenko, N.V. Krasnikov, M.M. Kirsanov, D.V. Kirpichnikov 1604.08432
Kahn, Krnjaic, Tran, Whitbeck, 1804.03144

**A muon fixed-target experiment would allow
fully inclusive coverage for \lesssim GeV
solutions of the $(g - 2)_\mu$ anomaly.**

**Very important near-term experimental
opportunity!**

Muon Colliders

Muon colliders an incredibly attractive path to explore high energy physics.

Muon Colliders

1901.06150

The Muon Collider Working Group

Jean Pierre Delahaye¹, Marcella Diemoz², Ken Long³, Bruno Mansoulié⁴, Nadia Pastrone⁵ (chair), Lenny Rivkin⁶, Daniel Schulte¹, Alexander Skrinsky⁷, Andrea Wulzer^{1,8}

The Muon Smasher's Guide 2103.14043

Hind Al Ali¹, Nima Arkani-Hamed², Ian Banta¹, Sean Benevedes¹, Dario Buttazzo³, Tianji Cai¹, Junyi Cheng¹, Timothy Cohen⁴, Nathaniel Craig¹, Majid Ekhterachian⁵, JiJi Fan⁶, Matthew Forsslund⁷, Isabel Garcia Garcia⁸, Samuel Homiller⁹, Seth Koren¹⁰, Giacomo Koszegi¹, Zhen Liu^{5,11}, Qianshu Lu⁹, Kun-Feng Lyu¹², Alberto Mariotti¹³, Amara McCune¹, Patrick Meade⁷, Isobel Ojalvo¹⁴, Umut Oktem¹, Diego Redigolo^{15,16}, Matthew Reece⁹, Filippo Sala¹⁷, Raman Sundrum⁵, Dave Sutherland¹⁸, Andrea Tesi^{16,19}, Timothy Trott¹, Chris Tully¹⁴, Lian-Tao Wang¹⁰, and Menghang Wang¹

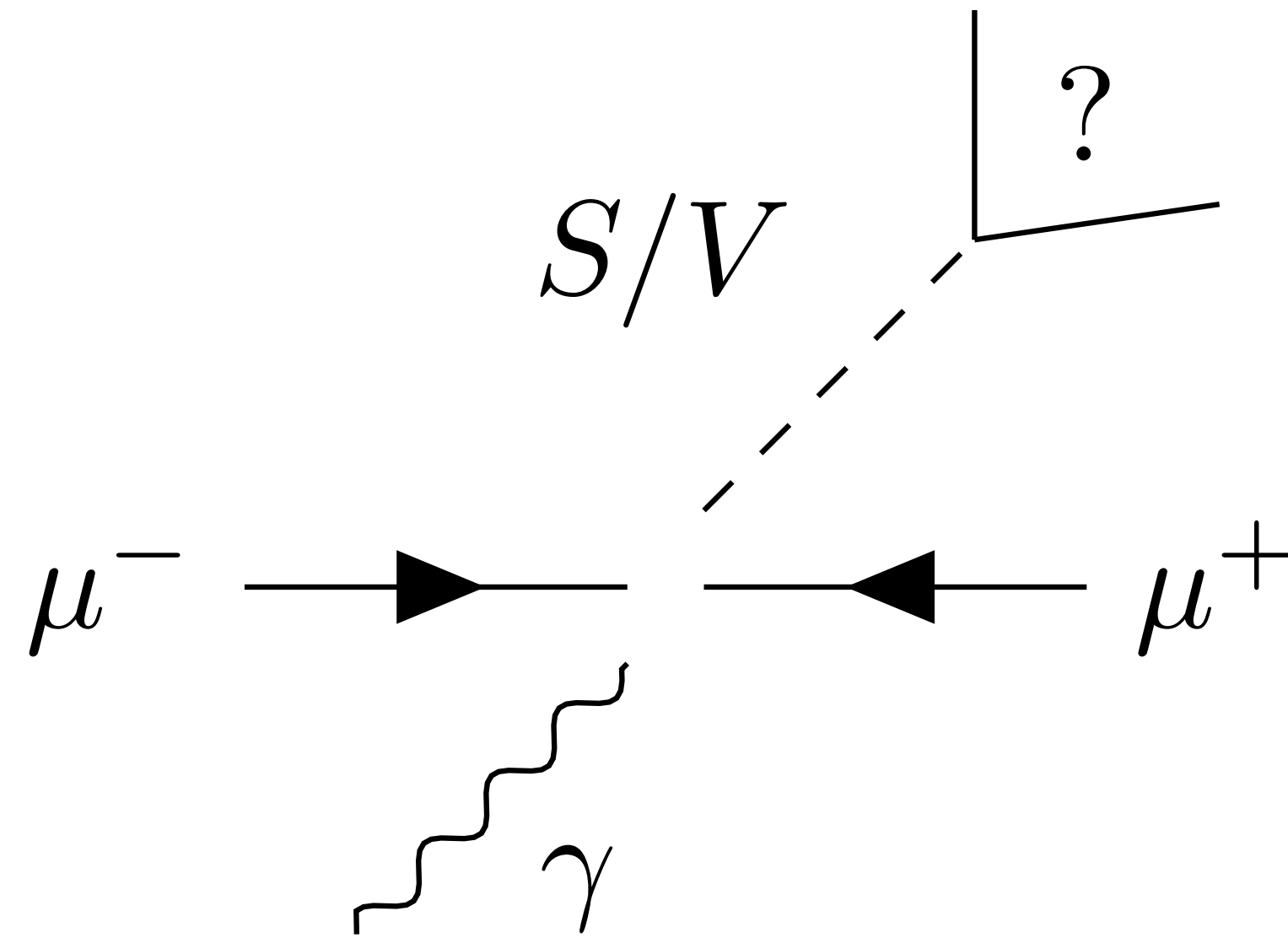
2005.10289 Constantini, De Lillo, Maltoni, Mantani, Mattelaer, Ruiz, Zhao

Bonus:

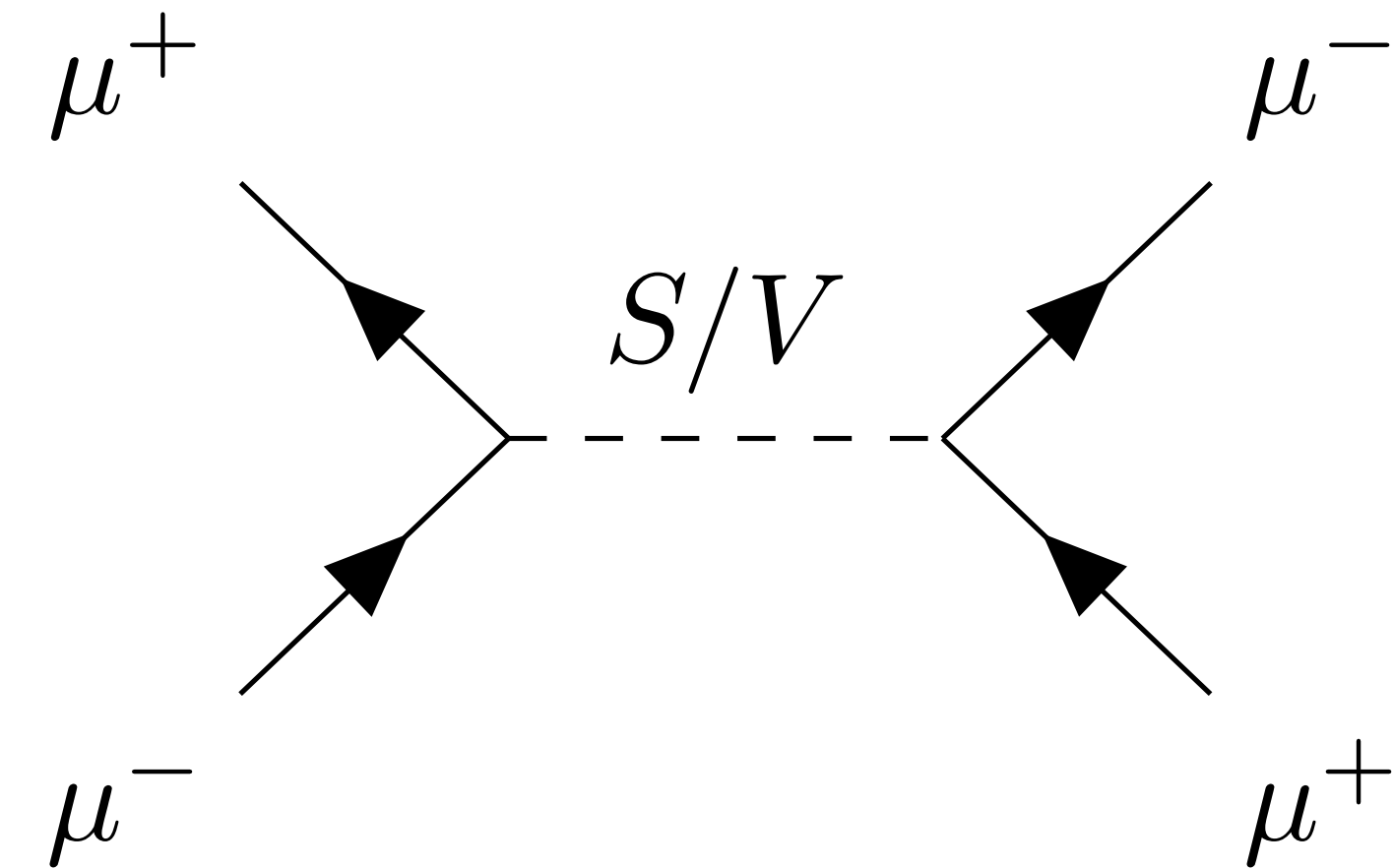
They are also “guaranteed” to discover the new physics of $(g - 2)_\mu$

$\sqrt{s} \sim 200 \text{ GeV} - 3 \text{ TeV}$: Discovering Singlet Scenarios

Discovering the singlet production in
fully inclusive search:
mono-photon + anything



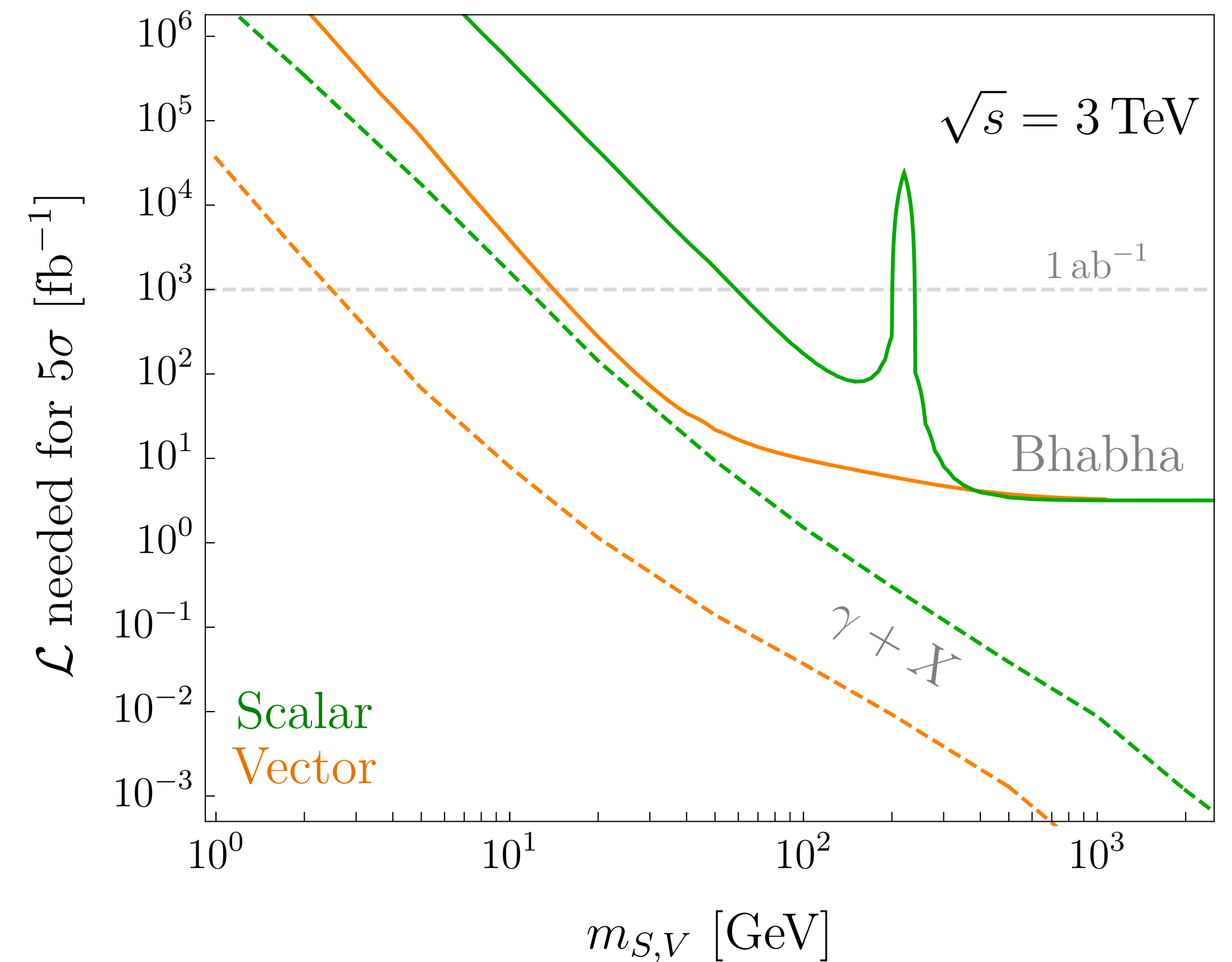
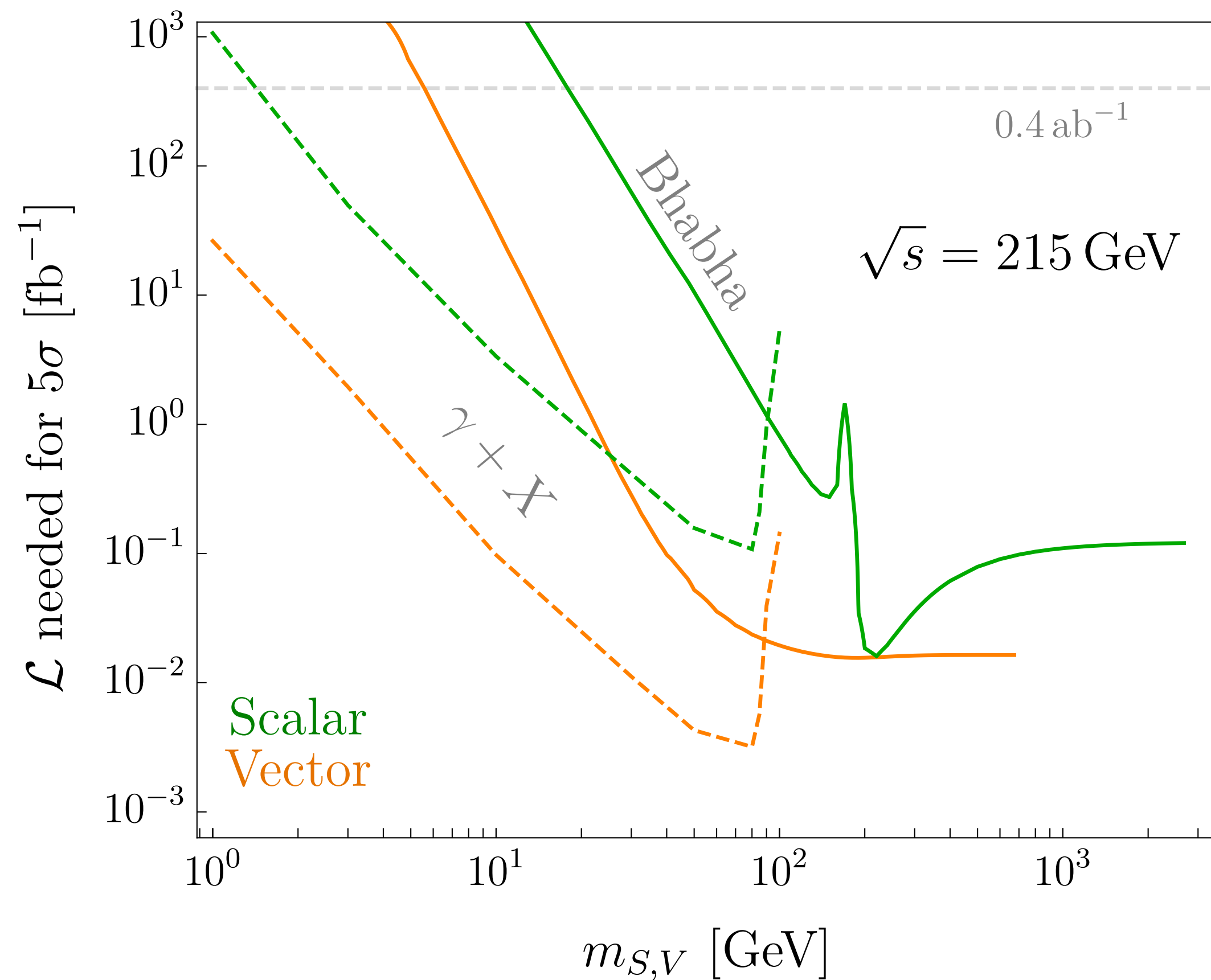
Indirect observation:
corrections to
Bhabha scattering



Only *guaranteed* coupling of singlet is to muons: Muon Collider is special!

$\sqrt{s} \sim 200 \text{ GeV} - 3 \text{ TeV}$: Discovering Singlet Scenarios

Collider study including conservative detector effects shows lumi needed for discovery



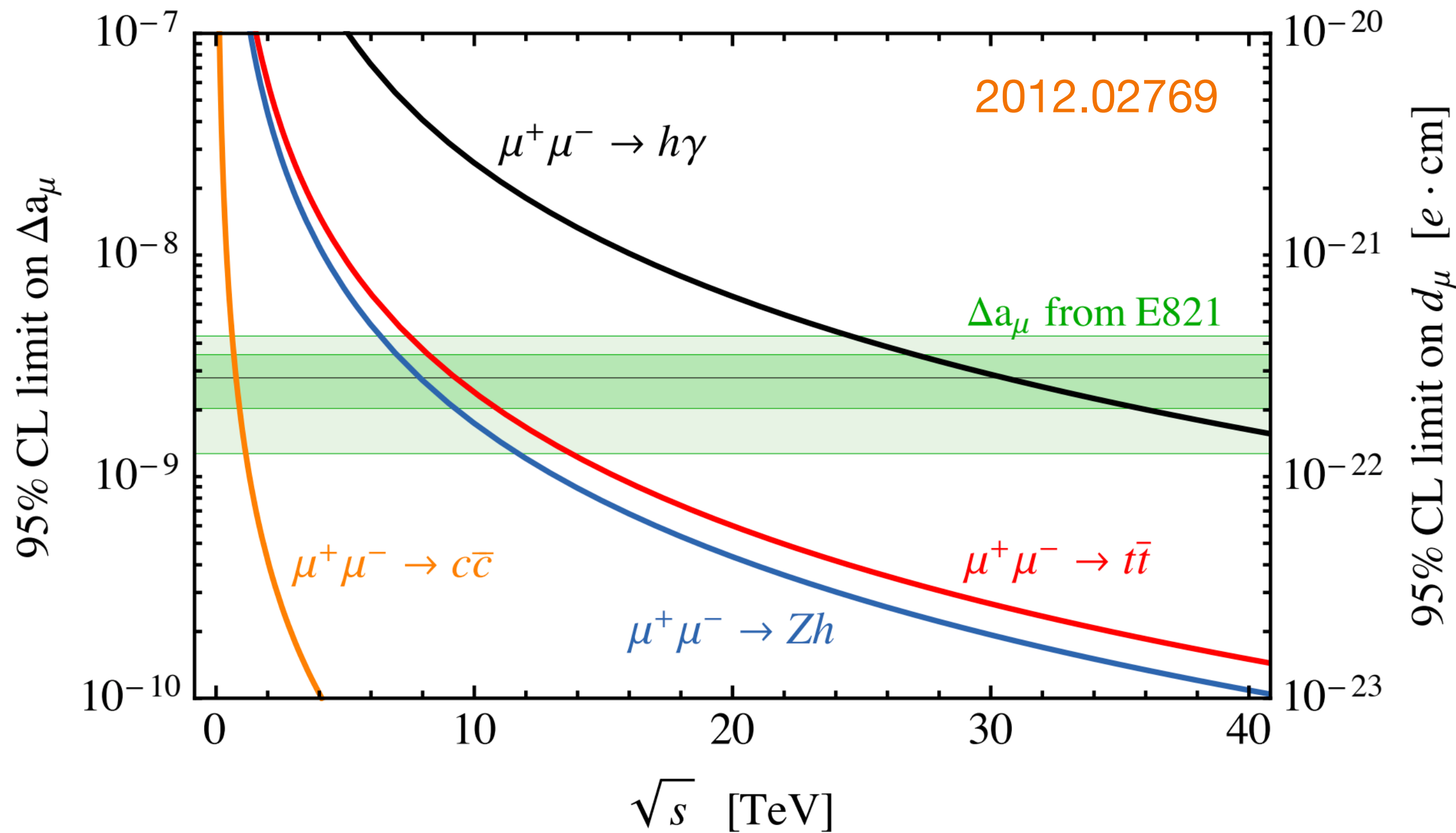
A TeV-scale muon collider program would discover all **Singlet solutions** to the $(g - 2)_\mu$ anomaly.

A 30 TeV muon collider will discover all “reasonable” **EW Scenarios** that account for the $(g - 2)_\mu$ anomaly.

But what if you don't see anything?

$\sqrt{s} \sim 10$ TeV: Indirect $h\gamma$ Signal

2012.02769 Buttazzo, Paradisi
2012.03928 Yin, Yamaguchi



If the new physics is heavier than 15 TeV, a 30 TeV muon collider could still see the

$$\mu\mu \rightarrow h\gamma$$

signal produced by the same operator

$$\frac{1}{M^2} H^\dagger (L \sigma^{\nu\rho} \mu^c) F_{\nu\rho}$$

A 30 TeV muon collider will see either new charged states, and/or the indirect

$$\mu\mu \rightarrow h\gamma \text{ signal!}$$

Therefore, if you don't see new charged states, you **know**** the states are there but at higher masses.**

→ Proof of a very weird and tuned universe!

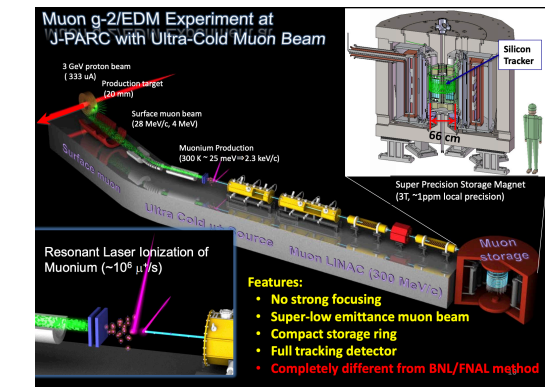
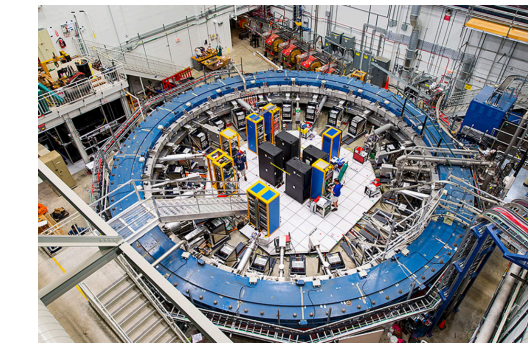
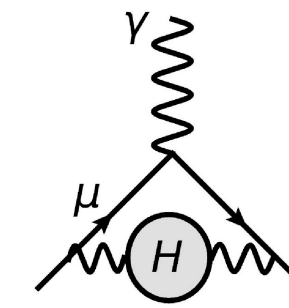
A no-lose theorem for $(g - 2)_\mu$

2006.16277, 2101.10334

Rodolfo Capdevilla, DC, Yonatan Kahn, Gordan Krnjaic

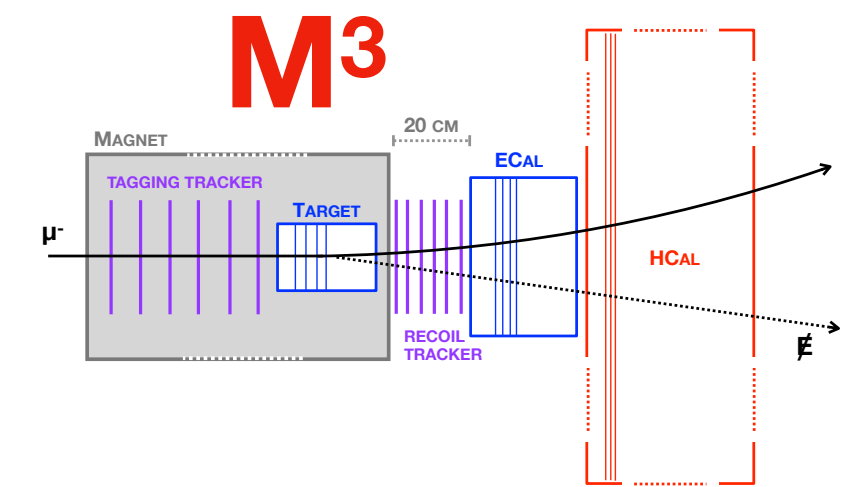
No-Lose Theorem for $(g - 2)_\mu$

1. Confirm the $(g - 2)_\mu$ anomaly is real.

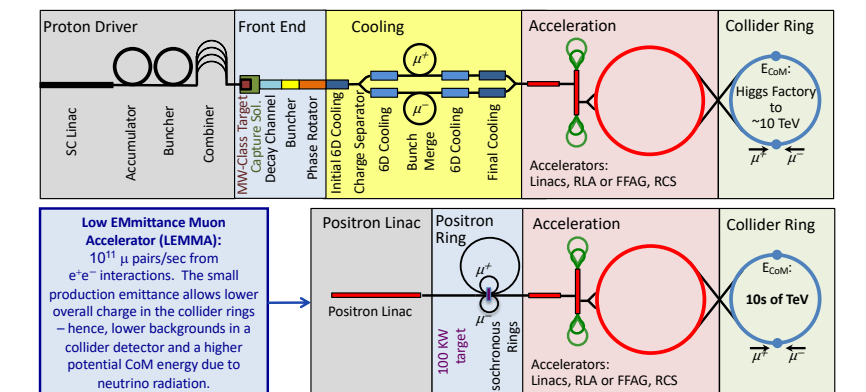


2. Look for \lesssim GeV Singlet Scenarios in μ fixed target experiments.

3. Build a TeV-scale muon collider. Discover **all** Singlet solutions (and probe deep into EW Scenario parameter space as well).



4. Build a 10-TeV-scale muon collider. Discover **all** “reasonable” Electroweak solutions, and/or observe $h\gamma$ signal.



5. **Either find new particles, or prove the universe is explicitly, calculably fine-tuned with weird flavour physics.**

Either way, a comprehensive muon program revolutionizes our understanding of the universe.

Thank you!