Inclusive DIS Variable Reconstruction at ep Colliders: some thoughts based on HERA

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[With thanks to many colleagues who made the plots shown here]

Similarities and Differences

HERA was:

- A high energy electron-proton collider with polarised electron/positron beams

HERA was not:

- An electron-ion collider
- A polarised target machine
- A high luminosity collider

... useful to compare but not necessarily to follow ...

Disclaimer:

- I worked on H1, so examples taken from there.
- ZEUS is broadly similar
- HERMES is a different talk entirely

H1 Detector and some immediate comments



- HERA detectors were (initially) built to focus more on high Q² / BSM searches and less on low x/Q² physics
- Reality turned out a bit differently
- 'Backward calorimeter' and
 MWPC later replaced with
 SPACAL with electromagnetic and
 hadronic sections + Backward
 Drift Chambers

- There was beamline electron tagger ($Q^2 < 0.01 \text{ GeV}^2$), but then a gap in tagged electron acceptance until $Q^2 \sim 1 \text{ GeV}^2$, only partially / temporarily fixed later.

- Locating main HCAL inside coil improved hadronic response (obviously₃ limited by magnet bore size)

Inclusive Reconstruction Basics



- x, Q² (via y, Q²) can be reconstructed from any two of E_e , θ_e , E_h , θ_h (see later for details)
- Hadronic final state kinematics also important for background rejection
- Starting point is therefore electron identification & reconstruction, plus inclusive hadronic final state measurement.

Scattered Electron Identification

- For high electron energies (>~ 10 GeV or 1/3 beam energy), choosing highest energy or highest p_T electromagnetic calo cluster is already efficient and almost background free

- At smaller energies, misidentification and 'photoproduction' background become important.



Scattered Electron Identification

- Particle ID at H1 was very limited (basically only dE/dx of tracker)
- Additional requirements improve selection efficiency and suppress most of the background:
 - ... compactness & isolation of cluster (radius, depth, HCAL fraction)
 - ... link to inner track (spatially and in E/p ratio)
 - ... overall event kinematics: total E-pz (electrons+hadrons) = $2E_e$

Energy E'_e of scattered electron candidate	> 3.4 GeV
Transverse size R_{log} of candidate cluster	< 5 cm
Hadronic energy fraction behind the cluster	$<15\%$ of ${E_e}'$
Transverse distance between cluster and linked track	< 6 cm
$E - p_z$	> 35 GeV
z position of interaction vertex	$ z_v < 35 \text{ cm}$

Table 1: Criteria applied to select DIS events at high inelasticity y.



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- Residual background subtraction controlled through comparisons with 'wrong-charge' clusters & subsample with tagged photoproduction electron
- Measurements down to $E_e \sim 3$ GeV (1/10 beam energy) were made.

Inclusive Hadronic Final State Reconstruction

- Reconstructing the inclusive hadronic final state (in general - not only high p_T jets) essential for photoproduction background rejection and kinematic reconstruction beyond electron-only methods.

- Use of hadronic final state p_T and E-pz as basic variables minimises impact of missing energy from proton remnant (which has E \approx pz)
- Energy flow algorithms developed to reconstruct hadronic final state by combining calorimeter and tracking information making optimal use of both
- Suppression of calorimeter noise at low energies is very important
- Hadronic final state measurements `easily' calibrated using pT and E-pz balance versus scattered electron in NC events



Why not just reconstruct NC kinematics using the electron method?

 $y_e = 1 - \frac{E'_e}{E_e} \sin^2 \frac{\theta}{2}$

Electron method resolution in y (~1/x) degrades as 1/y ... [E_e' getting large, towards the 'kinematic peak']



[Plots from Yellow Report]

... serious limitation on measurements at high x, where PDFs poorly known → important part of EIC programme

Figure 8.17: Resolutions, defined as (reconstructed - true)/true, for kinematic variables in NC 18x275 GeV events. The ineleasticity is require to be y < 0.95

A further complication: Initial State Radiation corrections

ISR corrections explode as $y \rightarrow 1$ (i.e. at low x)



... calculable in principle, but with uncertainties due to PDFs etc

Kinematic Variable Reconstruction Methods

Any combination of E_e , θ_e , E_h , θ_h can be used

- 1) Electron only method (NC)
- 2) Hadron only method (CC)

Even for inclusive NC processes, it is possible to do better by mixing 1) and 2).

3) Double Angle and 'DA-pT' methods (θ_e , θ_h)

 \rightarrow insensitive to calorimeter energy resolution

4) Sigma method and e-Sigma method $(E_e, \theta_e, (E - p_z)_h)$

 \rightarrow insensitive to initial state radiation

The best choice depends on kinematic region and details of detector performance. Common feature is improved resolution at low y 10

Sigma Method

 $y_{\Sigma} = \frac{\Sigma_{h}}{E - P_{Z}} , \qquad Q_{\Sigma}^{2} = \frac{P_{T,e}^{2}}{1 - y_{\Sigma}} , \qquad x_{\Sigma} = \frac{Q_{\Sigma}^{2}}{sy_{\Sigma}} . \qquad \text{where } \Sigma_{h} = (E - p_{Z}) \text{ of hadrons}$ $E - P_{Z} = E'_{e}(1 - \cos \theta_{e}) + \sum_{i} (E_{i} - p_{Z,i}) = \Sigma_{e} + \Sigma_{h} ,$

e-Sigma Method

$$y_{e\Sigma} = \frac{Q_e^2}{sx_{\Sigma}} = \frac{2E_e}{E - P_Z} y_{\Sigma} , \qquad Q_{e\Sigma}^2 = Q_e^2 , \qquad x_{e\Sigma} = x_{\Sigma} .$$

Double Angle Method

 $y_{DA} = \frac{\tan(\theta_h/2)}{\tan(\theta_e/2) + \tan(\theta_h/2)}, \quad Q_{DA}^2 = 4E_e^2 \cdot \frac{\cot(\theta_e/2)}{\tan(\theta_e/2) + \tan(\theta_h/2)}, \quad x_{DA} = \frac{Q_{DA}^2}{sy_{DA}}. \quad \text{where} \quad \tan\frac{\theta_h}{2} = \frac{\Sigma_h}{P_{T,h}}.$

Double Angle / pT Method

Replace θ_h with θ_{PT} where

$$\tan \frac{\theta_{PT}}{2} = \frac{\Sigma_{PT}}{P_{T,e}}, \quad \text{where} \quad \Sigma_{PT} = 2E_e \frac{C(\theta_h, P_{T,h}, \delta_{PT}) \cdot \Sigma_h}{\Sigma_e + C(\theta_h, P_{T,h}, \delta_{PT}) \cdot \Sigma_h}.$$

Examples of Improved Performance

Low y resolution

High y radiative corrections



Data used in Final HERA paper

Data Set		xB; Grid		O ² [GeV ²] Grid		£	e ⁺ /e ⁻	\sqrt{s}	$x_{\rm Bi}O^2$ from	Ref.	
		from	to	from	to	pb ⁻¹	0.62	GeV	equations	2011/2012	
HERA I $E_n = 820$ GeV and $E_n = 920$ GeV data sets											
H1 svx-mb[2]	95-00	0.000005	0.02	0.2	12	2.1	e ⁺ p	301.319	13,17,18	[3]	
H1 low $Q^{2}[2]$	96-00	0.0002	0.1	12	150	22	e ⁺ p	301,319	13,17,18	[4]	
H1 NC	94-97	0.0032	0.65	150	30000	35.6	e ⁺ p	301	19	[5]	
H1 CC	94-97	0.013	0.40	300	15000	35.6	e^+p	301	14	[5]	
H1 NC	98-99	0.0032	0.65	150	30000	16.4	e p	319	19	[0]	
H1 CC	98-99	0.013	0.40	300	15000	16.4	e p	319	14	[6]	
H1 NC HY	98-99	0.0013	0.01	100	800	16.4	e p	319	13	[7]	
H1 NC	99-00	0.0013	0.65	100	30000	65.2	e^+p	319	19	[7]	
H1 CC	99-00	0.013	0.40	300	15000	65.2	e^+p	319	14	[7]	
ZEUS BPC	95	0.000002	0.00006	0.11	0.65	1.65	e ⁺ p	300	13	[11]	
ZEUS BPT	97	0.0000006	0.001	0.045	0.65	3.9	e ⁺ p	300	13, 19	[12]	
ZEUS SVX	95	0.000012	0.0019	0.6	17	0.2	e^+p	300	13	[13]	
ZEUS NC[2] high/low Q2	96-97	0.00006	0.65	2.7	30000	30.0	e ⁺ p	300	21	[14]	
ZEUS CC	94-97	0.015	0.42	280	17000	47.7	e ⁺ p	300	14	[15]	
ZEUS NC	98-99	0.005	0.65	200	30000	15.9	e ⁻ p	318	20	[16]	
ZEUS CC	98-99	0.015	0.42	280	30000	16.4	e ⁻ p	318	14	[17]	
ZEUS NC	99-00	0.005	0.65	200	30000	63.2	e ⁺ p	318	20	[18]	
ZEUS CC	99-00	0.008	0.42	280	17000	60.9	e ⁺ p	318	14	[19]	
HERA II $E_p = 920 \text{ GeV}$ data sets											
H1 NC 1.5p	03-07	0.0008	0.65	60	30000	182	e ⁺ p	319	13, 19	[8]1	
H1 CC 1.5p	03-07	0.008	0.40	300	15000	182	e ⁺ p	319	14	[8]1	
H1 NC 1.5p	03-07	0.0008	0.65	60	50000	151.7	e p	319	13, 19	[8]1	
H1 CC 1.5p	03-07	0.008	0.40	300	30000	151.7	e p	319	14	[8]1	
H1 NC med O2 *y.5	03-07	0.0000986	0.005	8.5	90	97.6	e ⁺ p	319	13	[10]	
H1 NC low 02 *9.5	03-07	0.000029	0.00032	2.5	12	5.9	e ⁺ n	319	13	[10]	
ZEUS NC	06-07	0.005	0.65	200	30000	135.5	e ⁺ p	318	13,14,20	[22]	
ZEUS CC 1.5p	06-07	0.0078	0.42	280	30000	132	e ⁺ n	318	14	[23]	
ZEUS NC 1.5	05-06	0.005	0.65	200	30000	169.9	e ⁻ n	318	20	[20]	
ZEUS CC 1.5	04-06	0.015	0.65	280	30000	175	e p	318	14	[21]	
ZEUS NC nominal *9	06-07	0.000092	0.008343	7	110	44.5	e ⁺ n	318	13	[24]	
ZEUS NC satellite *	06-07	0.000071	0.008343	5	110	44.5	e ⁺ n	318	13	[24]	
HERA IF $E = 575$ GeV days sets											
HINC high Q^2	07	0.00065	0.65	35	800	5.4	e ⁺ n	252	13 10	101	
HI NC low O^2	07	0.0000279	0.0148	15	90	5.9	e p e ⁺ p	252	13, 19	[10]	
ZEUS NC nominal	07	0.000147	0.013340	7	110	7.1	e p e ⁺ p	251	13	[24]	
ZEUS NC nonlinal	07	0.000125	0.013349	5	110	7.1	e p	251	13	[24]	
HERA II $E = -60$ GeV data sets											
H1NC bigh 0 ² 07 0.00081 0.65 35 800 118 e ⁺ p 225 13.10 [0]											
H1 NC low O^2	07	0.00081	0.03	15	006	12.2	e p	225	13, 19	[9]	
TELIS NG seminal	07	0.0000348	0.0146	1.5	90	12.2	e p	225	13	[10]	
ZEUS NC nominal	07	0.000184	0.016686		110	13.9	e p	225	13	[24]	
ZEUS NC satellite	0/	0.000143	0.016686	5	110	13.9	e p	225	15	[24]	

Input data to final HERA combination

Executive NC summary:

- e-method (13) used in limited phase space regions only

- Σ , e Σ (17,18, 19) used extensively by H1 at low y (<~ 0.1-0.2)

- DA, pT (20,21) used extensively by ZEUS

Summary / Questions

Beyond the (excellent) material in the Yellow Report... some thoughts on things we may still want to investigate in ATHENA inclusive group ...

[all depend on details of detector; requires simulation of proposed solutions and reconstruction algorithms based on multiple components]

- What can be gained in scattered electron selection / background rejection from sophisticated requirements including cluster compactness, isolation, overall event E-pz etc?
- What level of performance is needed / can be obtained in overall hadronic final state reconstruction (via energy flow algorithms using multiple detector components)
- How much can we improve on NC kinematic reconstruction by trying sigma and double-angle methods?

 \rightarrow Possibly significant implications for detector design ...