



Computing Challenges in ~~Nuclear~~ and Particle Physics

Paul Laycock

August 3rd 2021

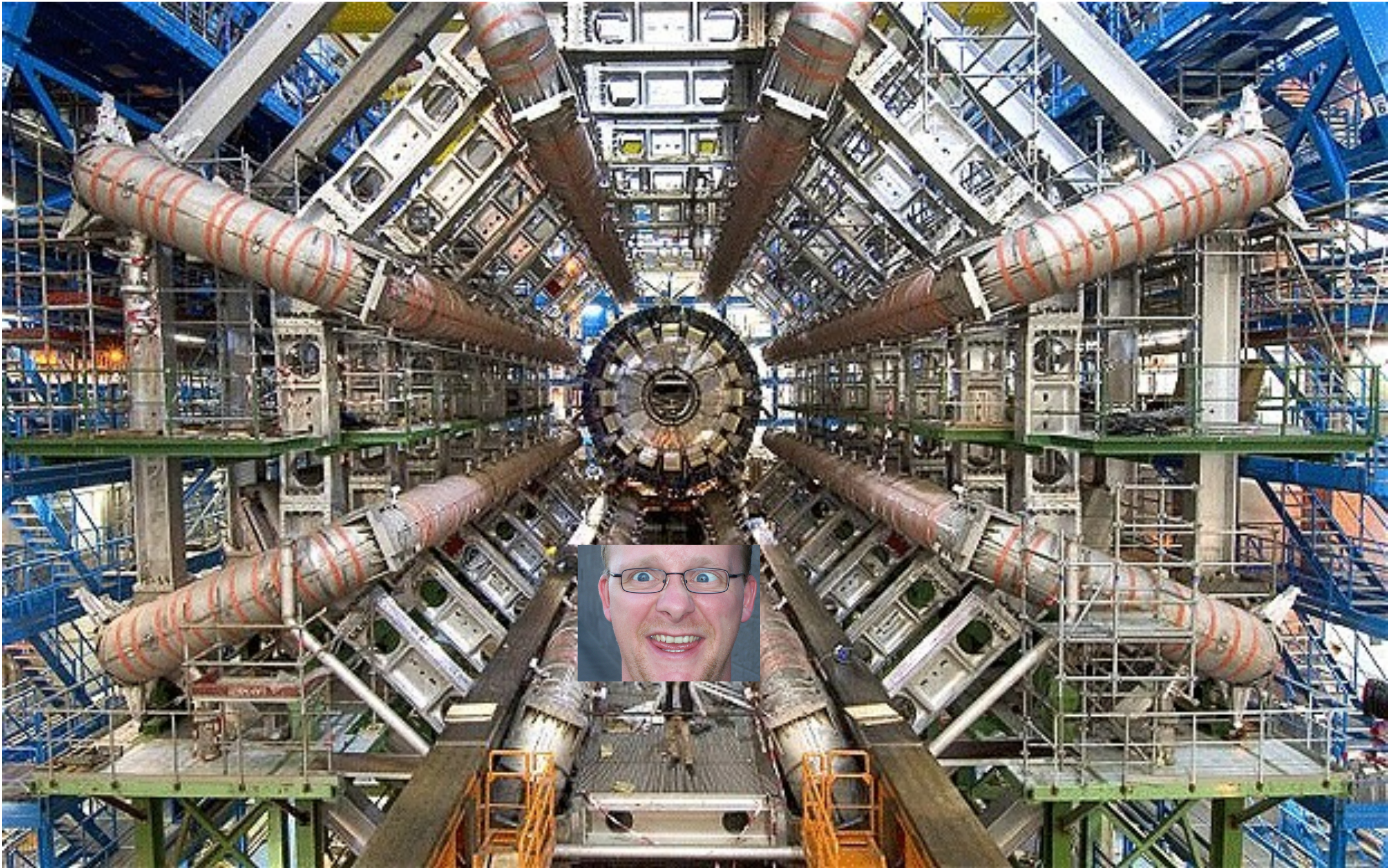


Particle physics and data

	BaBar	H1	ZEUS	HERMES	Belle	BESIII	CDF	DØ
End of data taking	07.04.08	30.06.07	30.06.07	30.06.07	30.06.10	2017	30.09.11	30.09.11
Type of data to be preserved	RAW data Sim/rec level Data skims in ROOT	RAW data Sim/rec level Analysis level ROOT data	Flat ROOT based ntuples	RAW data Sim/rec level Analysis level ROOT data	RAW data Sim/rec level	RAW data Sim/rec level ROOT data	RAW data Rec. level ROOT files (data+MC)	Raw data Rec. level ROOT files (data+MC)
Data Volume	2 PB	0.5 PB	0.2 PB	0.5 PB	4 PB	6 PB	9 PB	8.5 PB
Desired longevity of long term analysis	Unlimited	At least 10 years	At least 20 years	5-10 years	5 years	15 years	Unlimited	10 years

The ATLAS and CMS experiments already have 100s of petabytes of data

Particle Physics and me



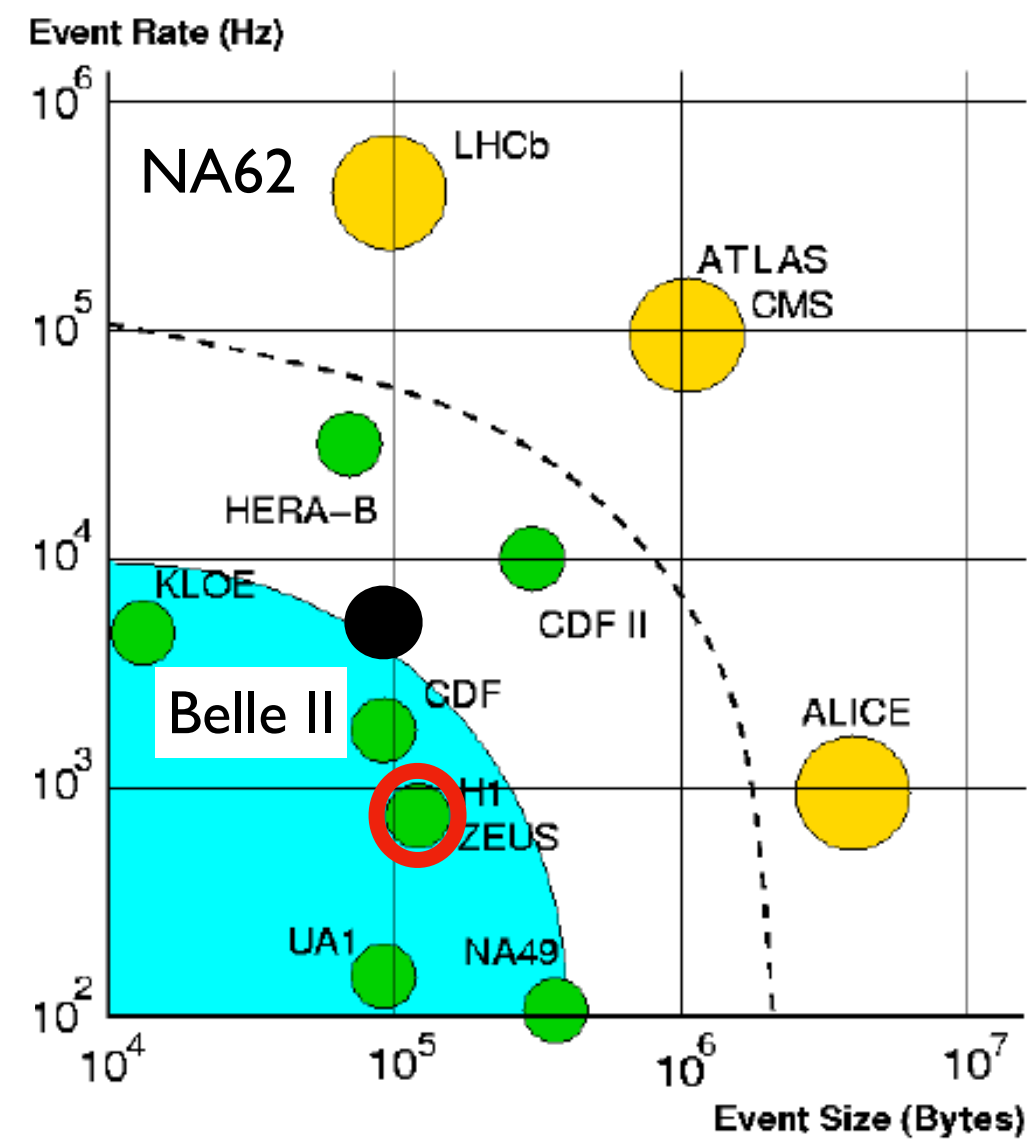
Physics Data

Plot modified from:
“GridPP: development of the UK
computing Grid for particle physics.”

H1

- Proton structure and QCD
- small event sizes and rates

**DAQ throughput =
event rate * event size**



Physics Data

Plot modified from:
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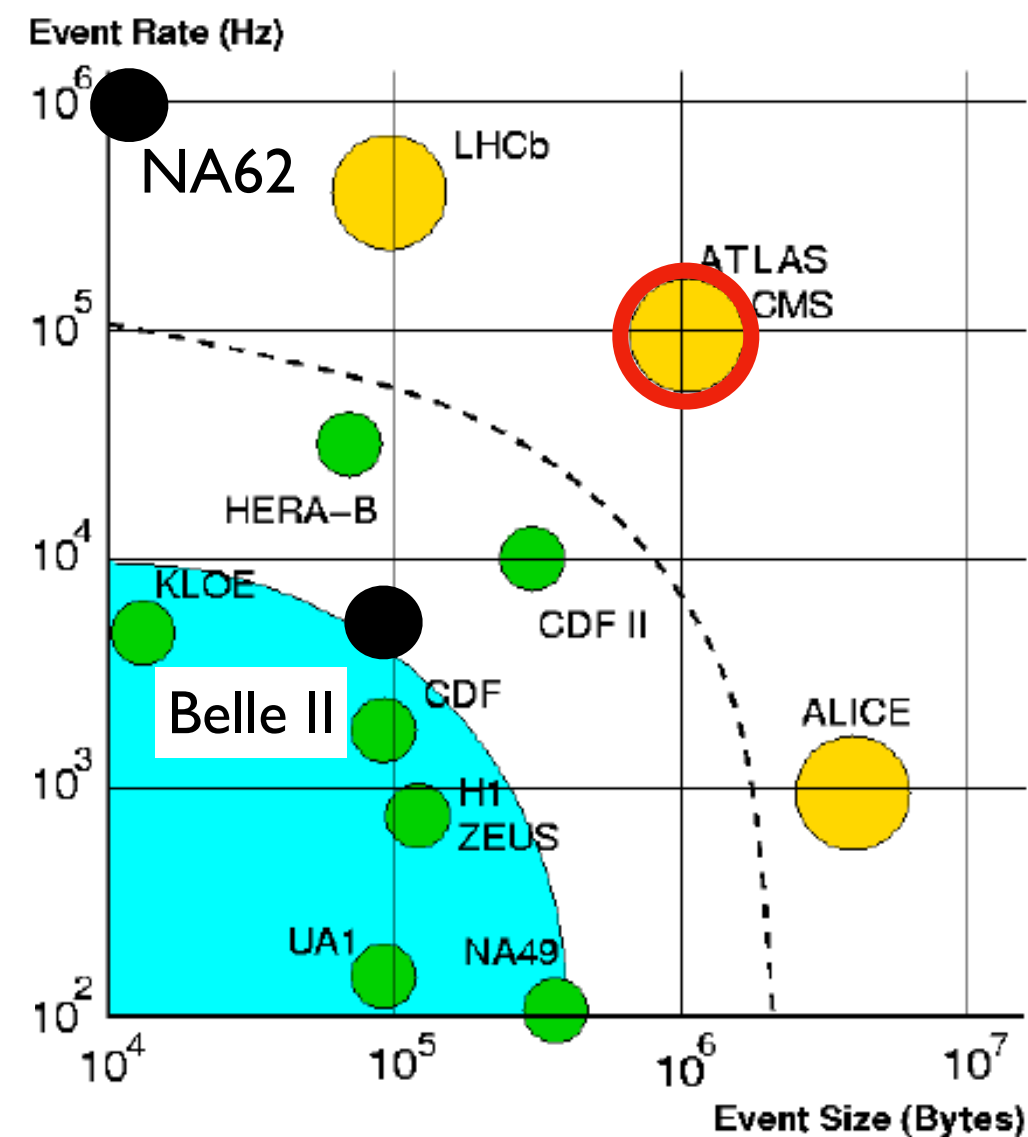
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ATLAS

- Higgs, searches for new physics
- big event sizes and rates

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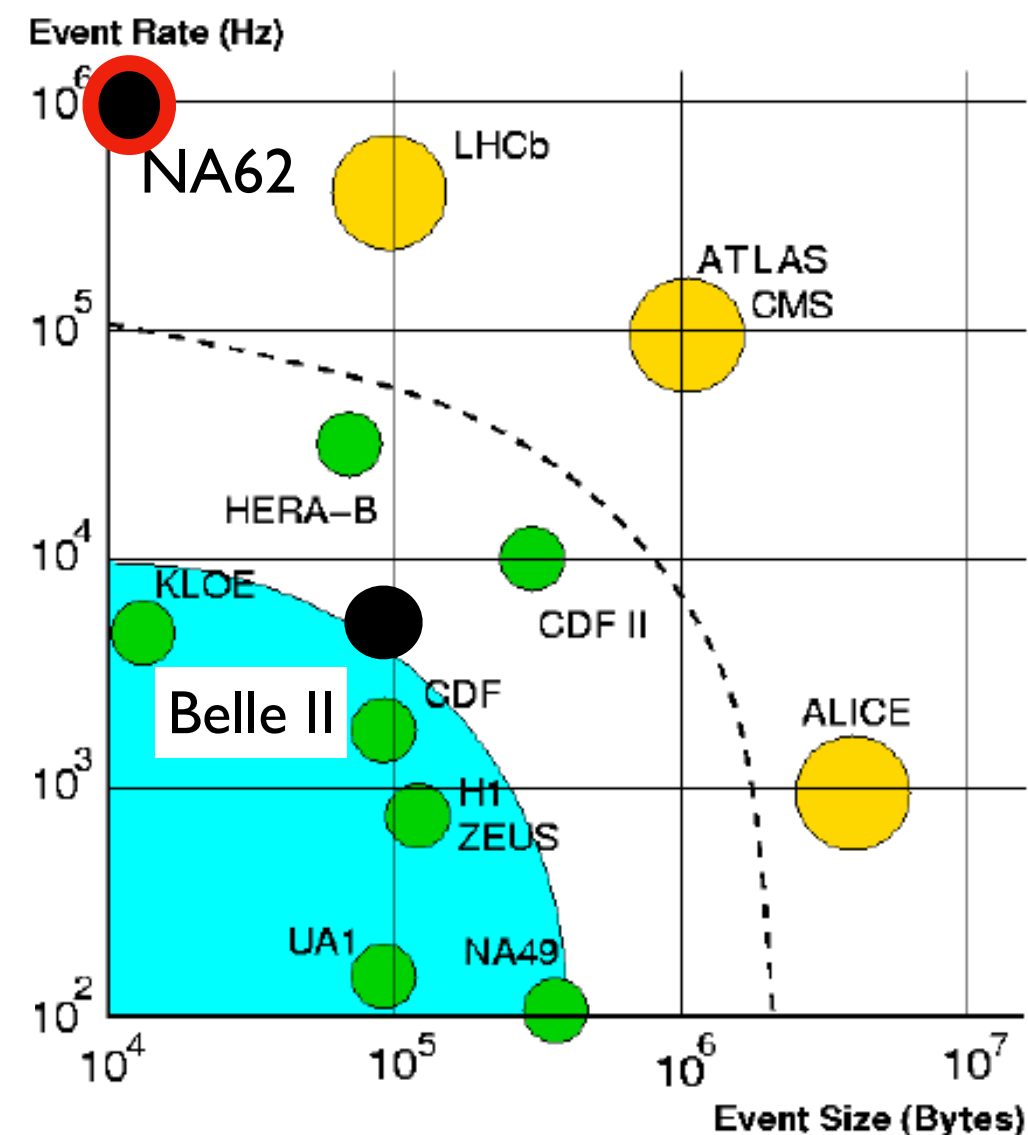
ATLAS

- Higgs, searches for new physics
- big event sizes and rates

NA62

- Ultra-rare kaon decays
- huge rates of small size events

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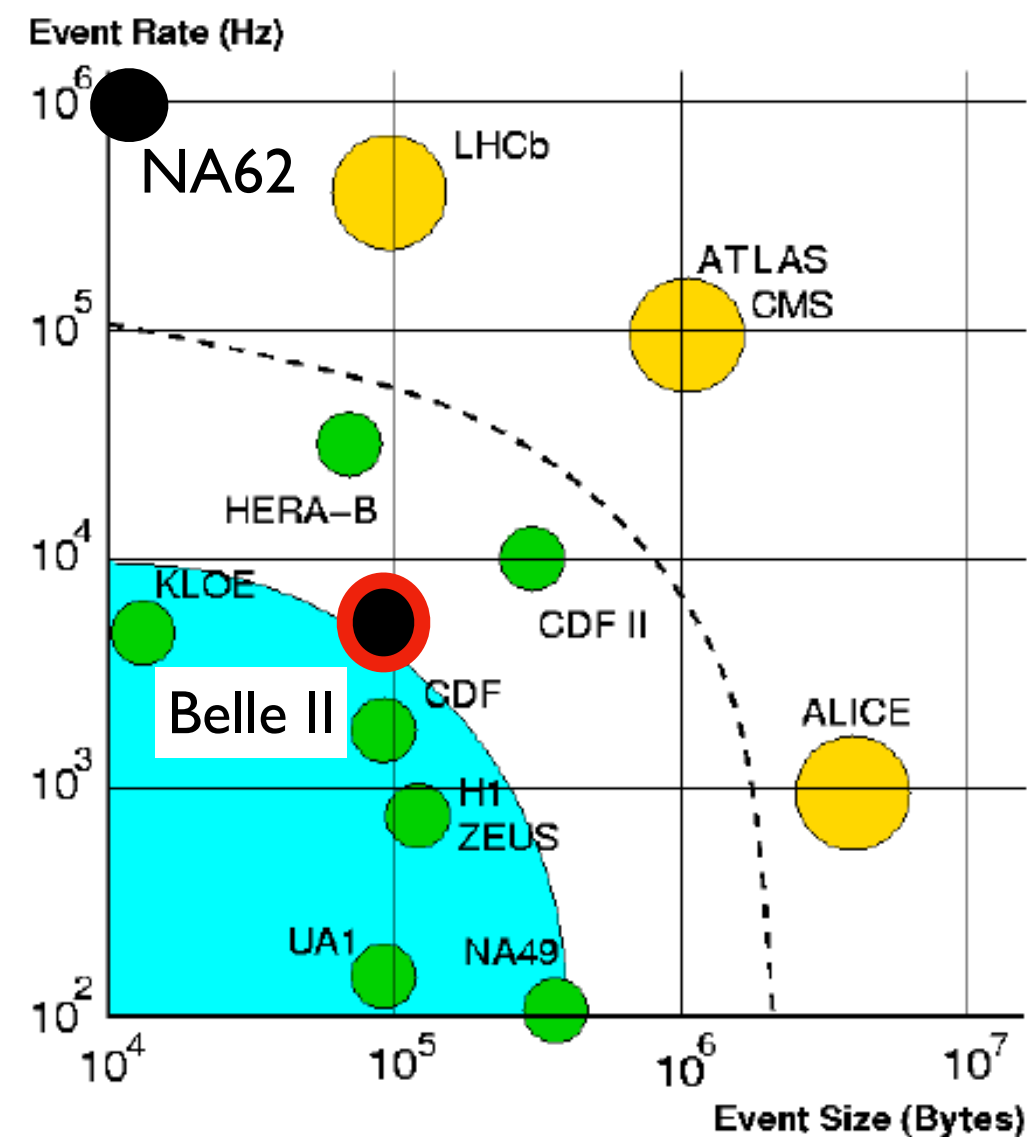
NA62

- Ultra-rare kaon decays
- huge rates of small size events

Belle II

- Ultra-rare B decays
- modest event sizes and rates

**DAQ throughput =
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Physics Data

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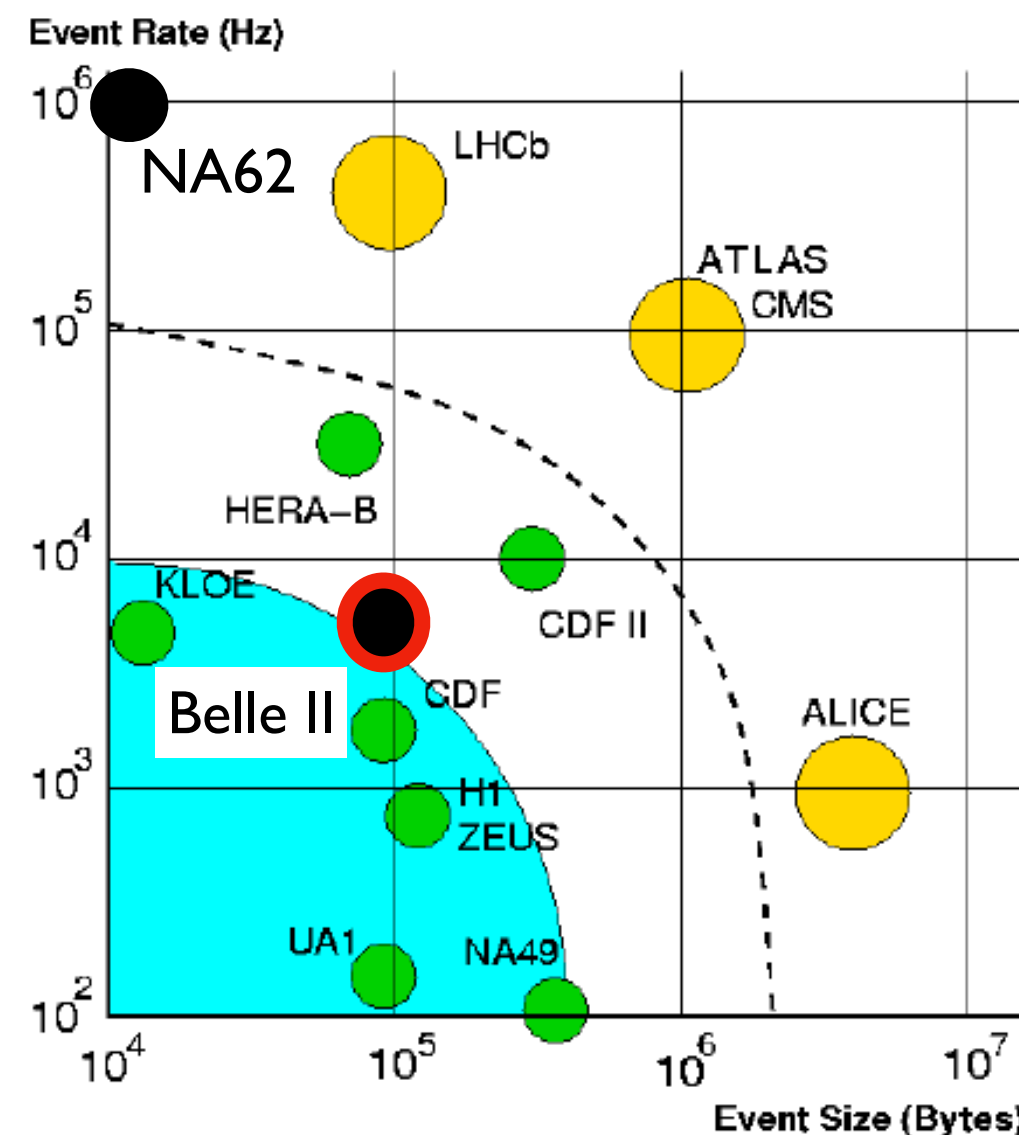
- Ultra-rare kaon decays
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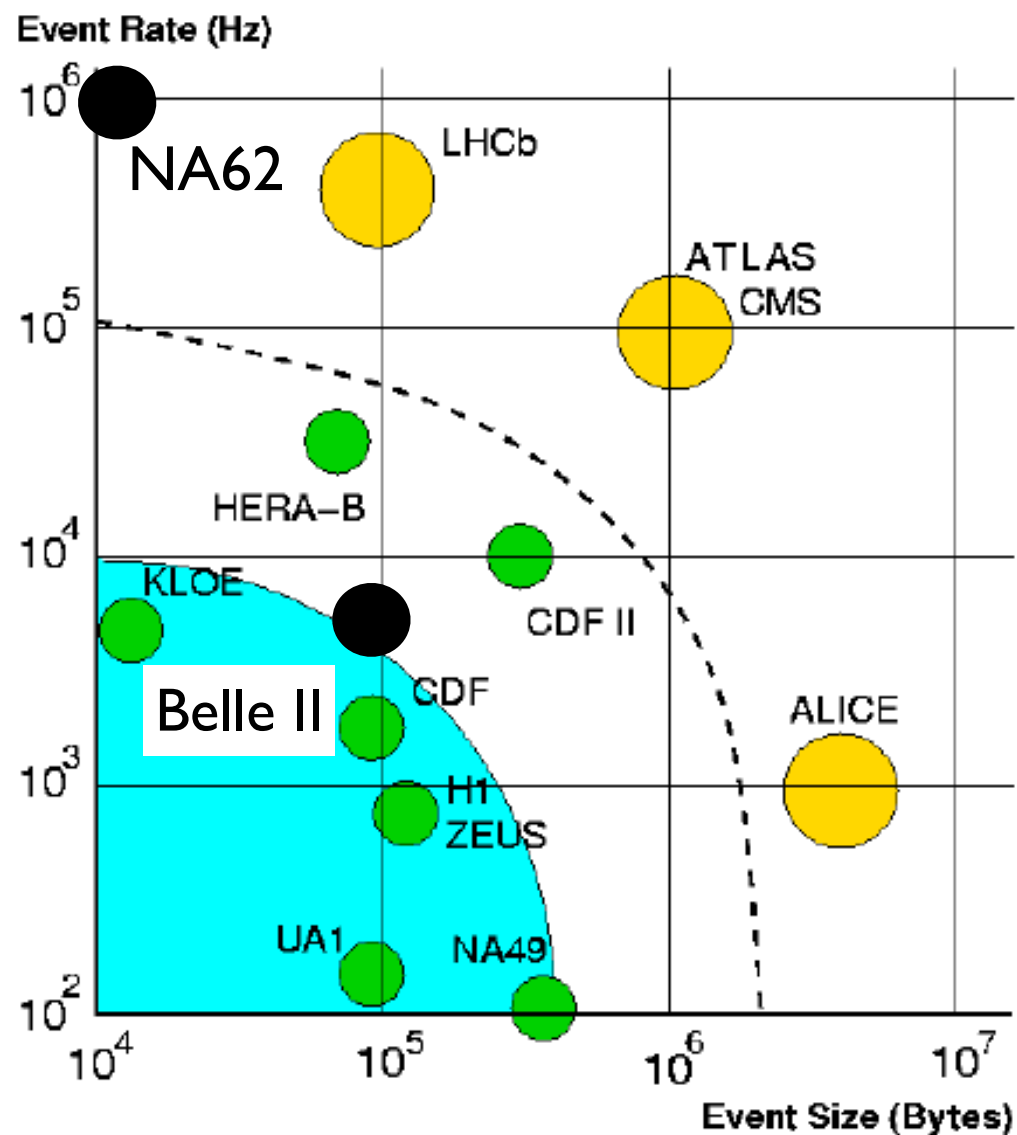
Belle II

- Ultra-rare B decays
- modest event sizes and rates

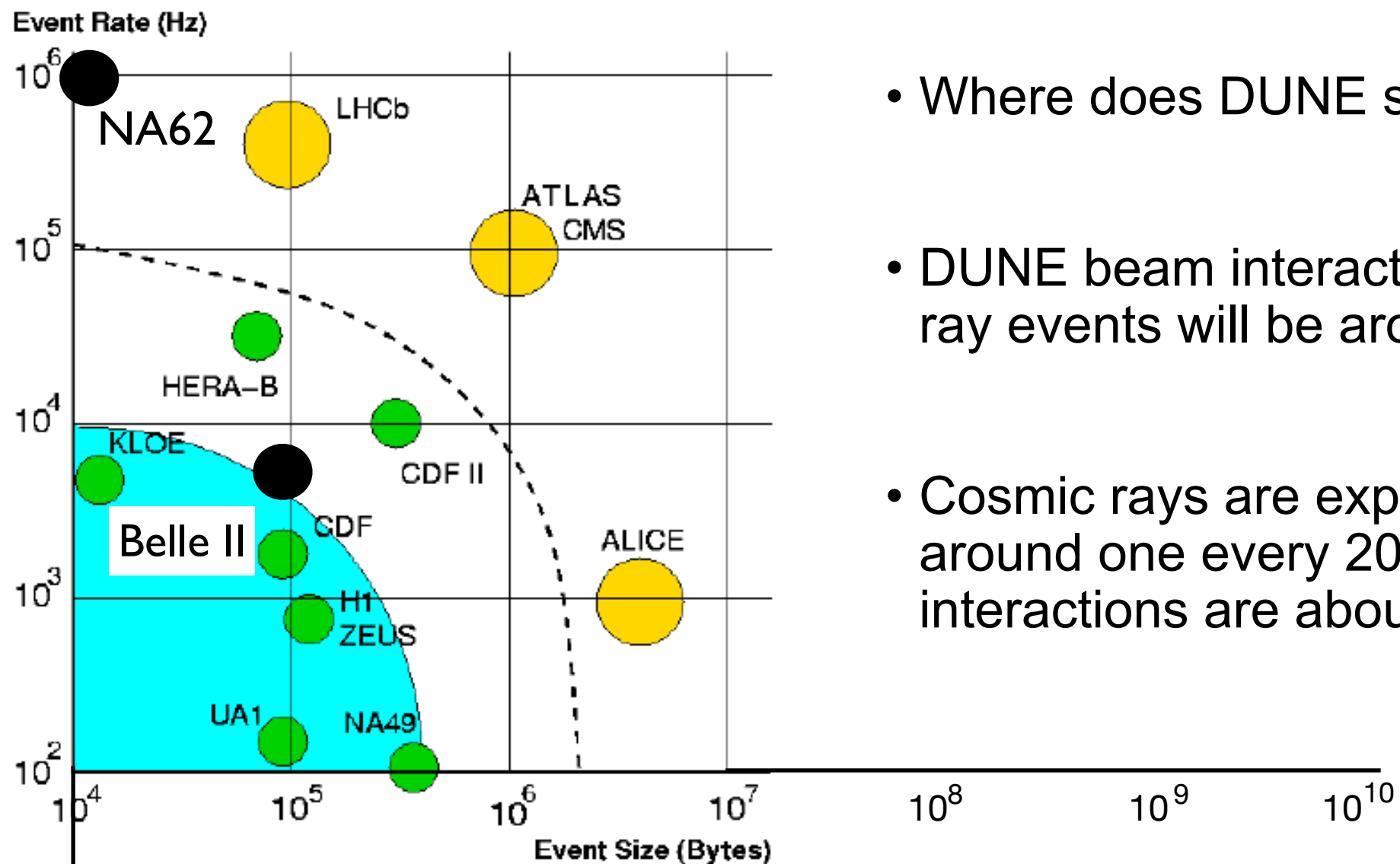
Question: What about DUNE??

**DAQ throughput =
event rate * event size**





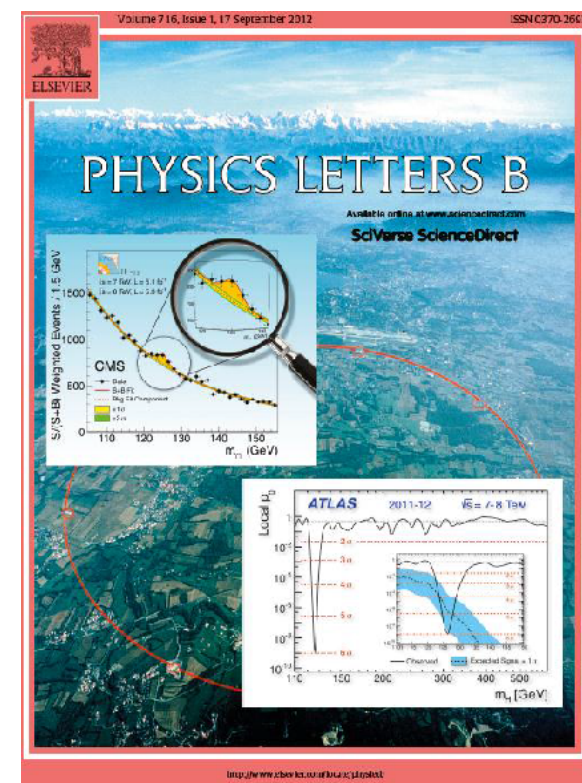
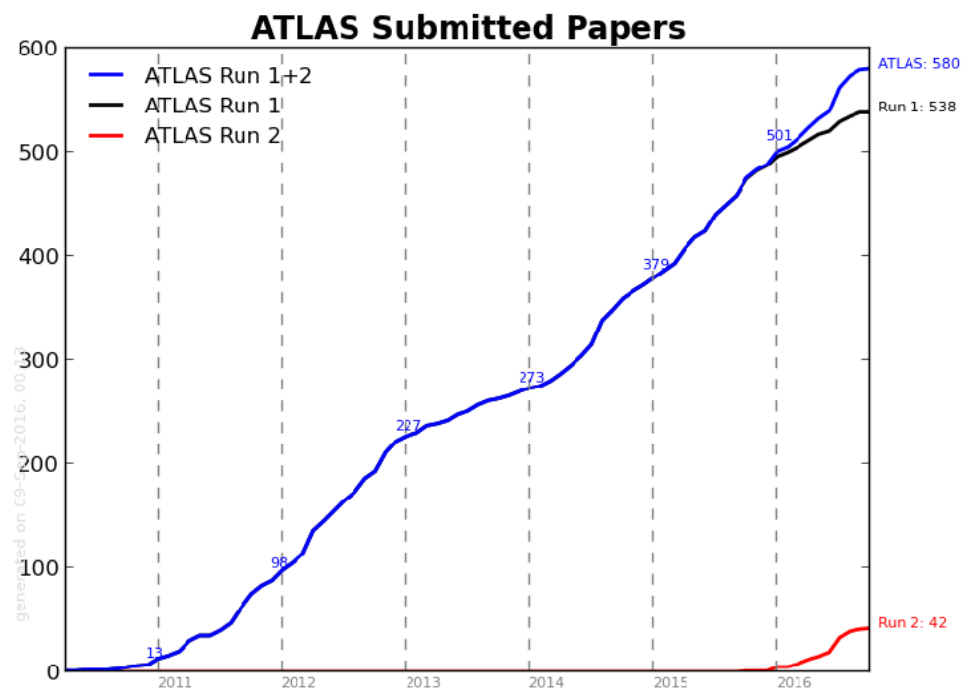
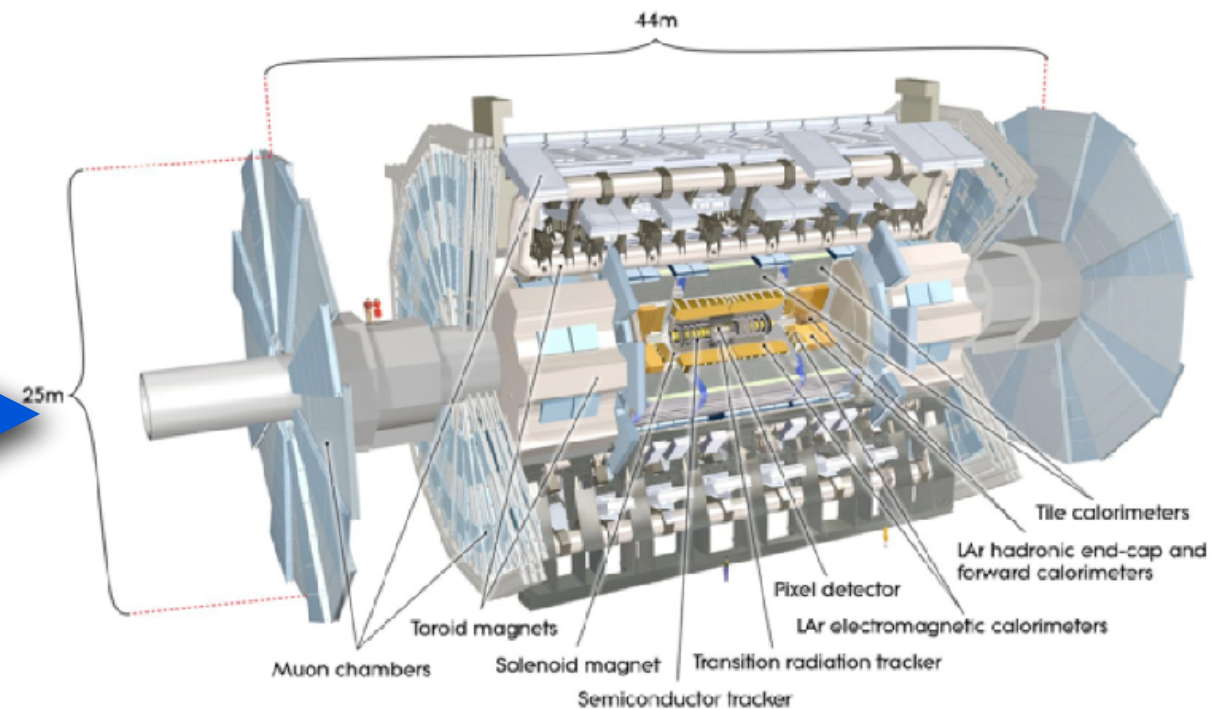
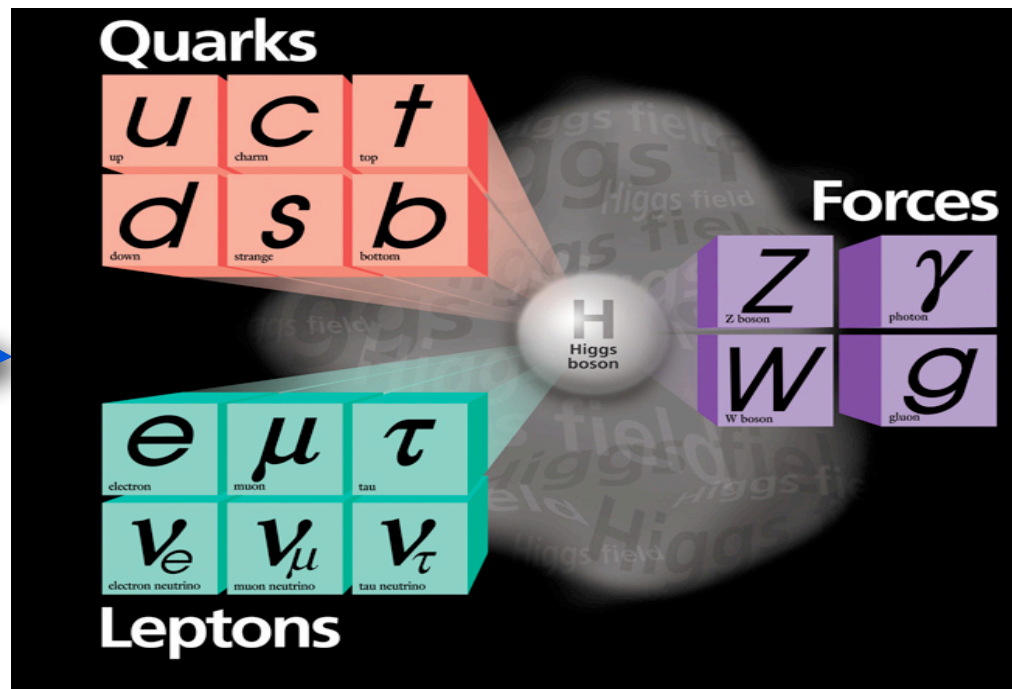
- Where does DUNE sit on this figure?
- DUNE beam interactions and cosmic ray events will be ??? Bytes in size
- Cosmic rays are expected at a rate of ????



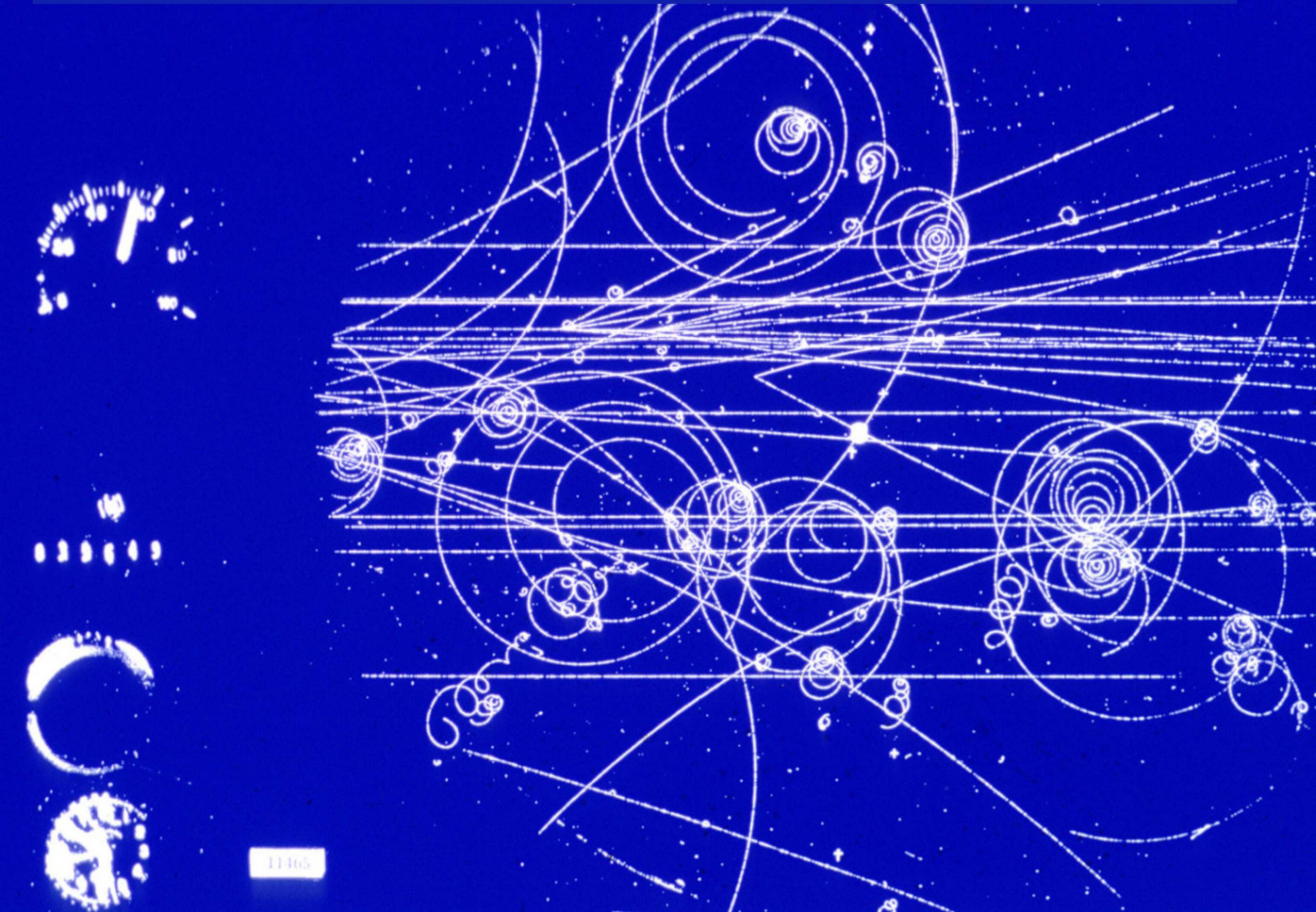
- Where does DUNE sit on this figure?
- DUNE beam interactions and cosmic ray events will be around 6GB in size
- Cosmic rays are expected at a rate of around one every 20s (beam interactions are about 40 / day)



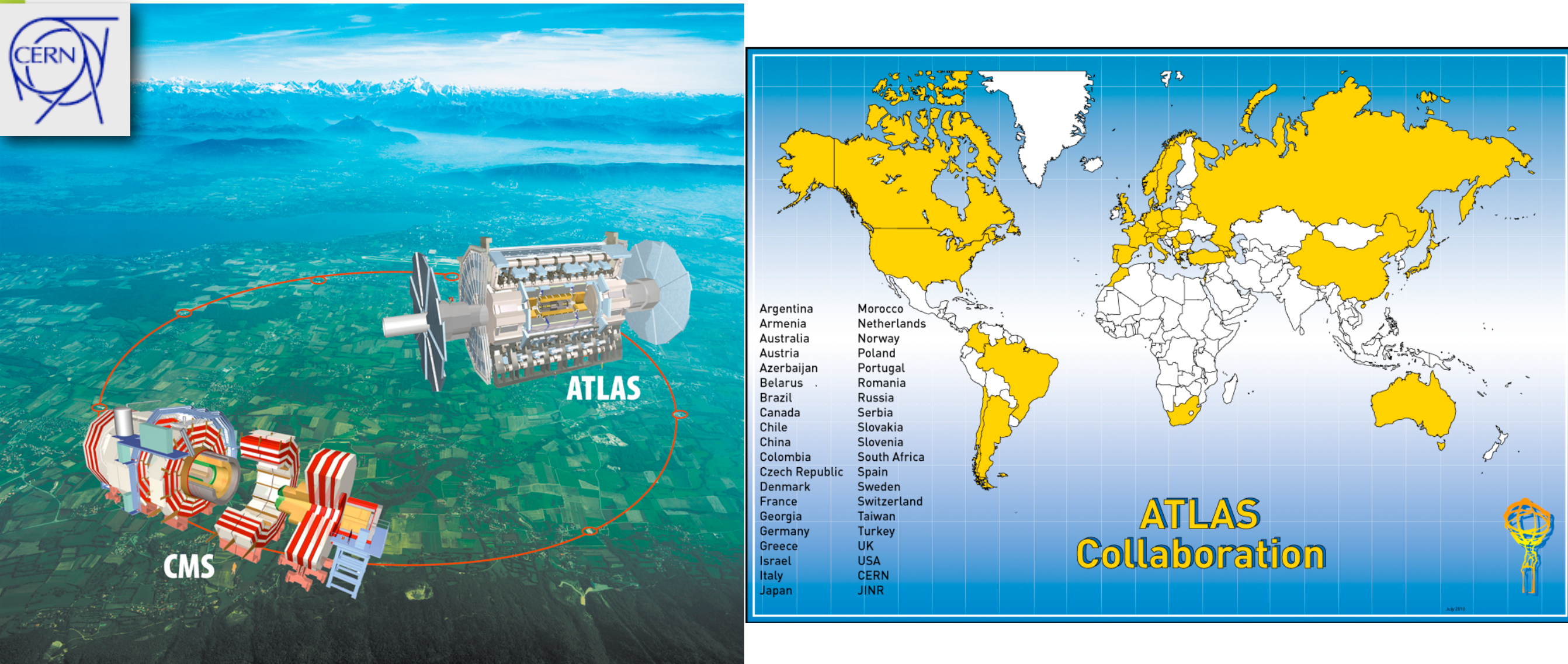
The particle physics cycle



... data analysis used to involve a person looking at pictures



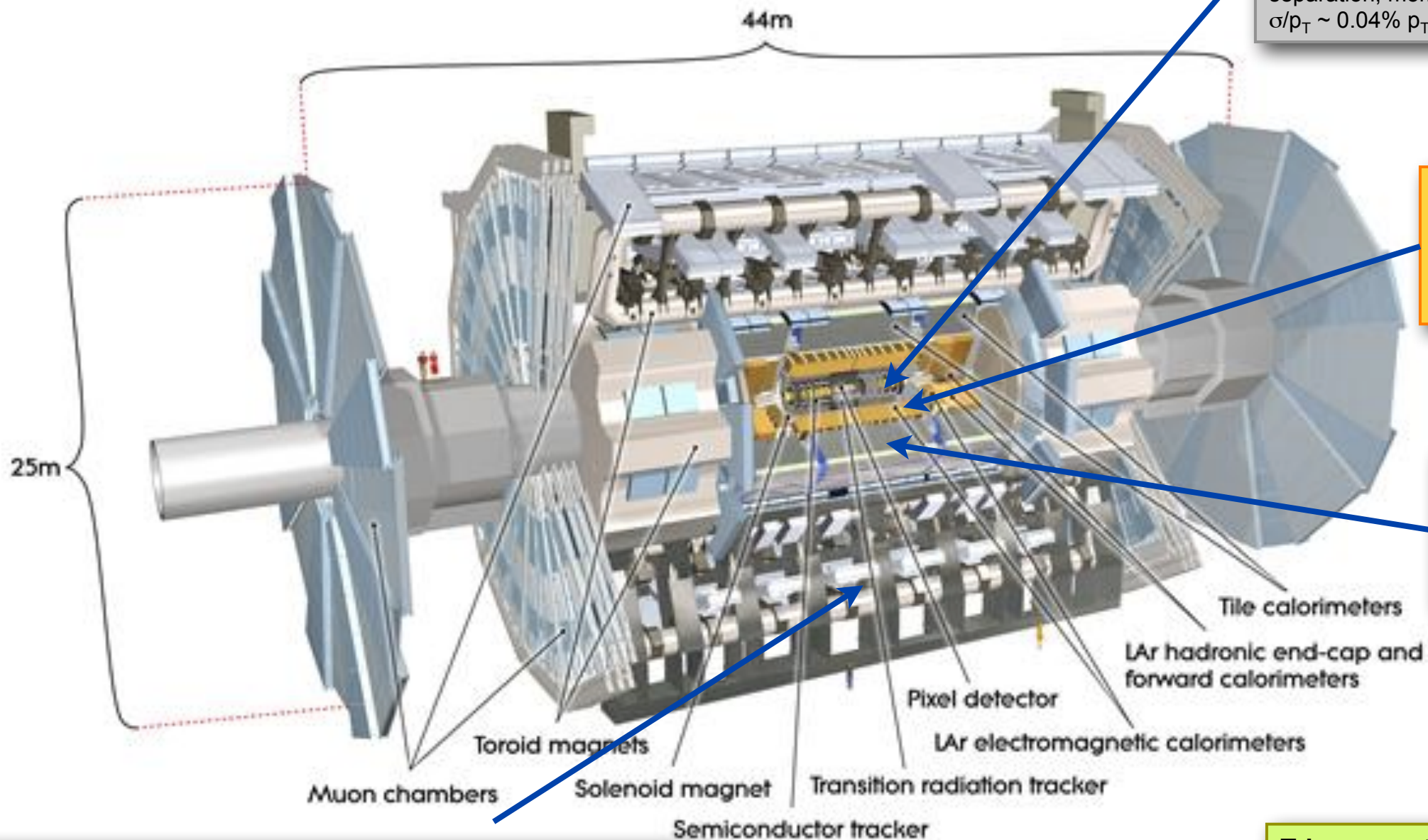
ATLAS and the Large Hadron Collider



I'll use ATLAS as an example of a modern experiment
similar things are done on other experiments

The ATLAS Detector @ LHC

L ~ 46 m, \varnothing ~ 22 m, 7000 tons
~ 10^8 electronic channels



Inner Tracker ($|\eta| < 2.5$, $B=2T$):
Si Pixels, Si strips, Trans. Rad. Det.
Precise tracking and vertexing, e/π
separation, momentum resolution:
 $\sigma/p_T \sim 0.04\% p_T (\text{GeV}) \oplus 1.5\%$

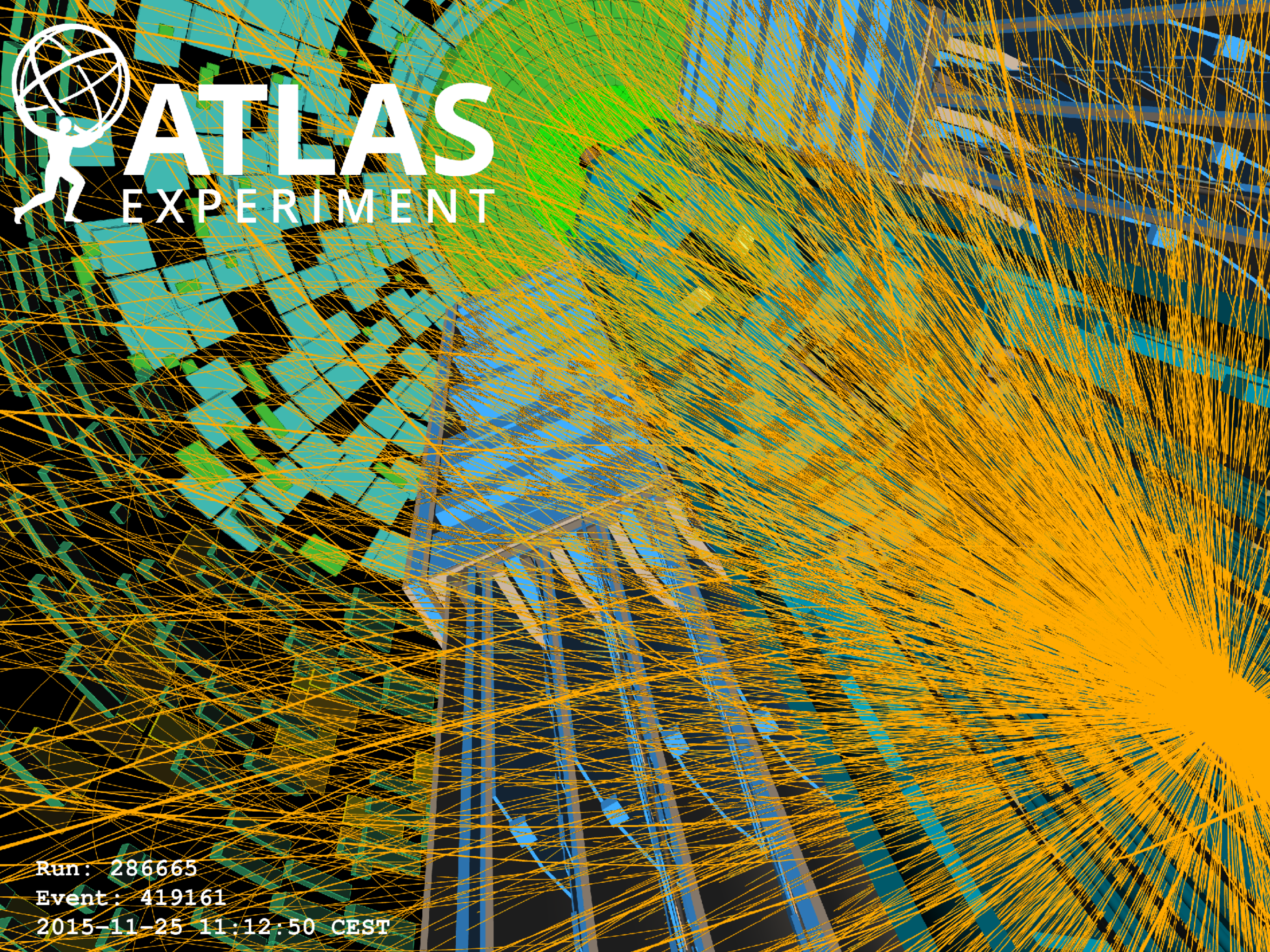
EM calorimeter:
Pb-LAr Accordion, e/γ
trigger, id. and meas.,
energy res.: $\sigma/E \sim$
 $10\%/\sqrt{E} \oplus 0.7\%$

HAD calorimetry ($|\eta| < 5$): Fe/
scintillator Tiles (cen), Cu/W-LAr
(fwd). trigger and meas. of jets
and $E_{T, \text{miss}}$, energy res.: $\sigma/E \sim$
 $50\%/\sqrt{E} \oplus 3\%$

Muon Spectrometer: air-core toroids with gas-based muon chambers.
trigger and meas. with momentum resolution $< 10\%$ up to $E_\mu \sim 1 \text{ TeV}$

Trigger system: 3-levels reducing
the IA rate from 40 MHz to ~200 Hz

Millions of detector readout channels read out to reconstruct one “event”



ATLAS

EXPERIMENT

Run: 286665

Event: 419161

2015-11-25 11:12:50 CEST

Analysing a lot of data

ATLAS has about 500 PBs of data

ATLAS data

Now analyse it !



Question:

How long would it take to read all of the ATLAS data?

*(Assume for simplicity you have off-the-shelf SSDs
with read speed ~500MB/s*

Hint: there are about 30 million seconds / year)

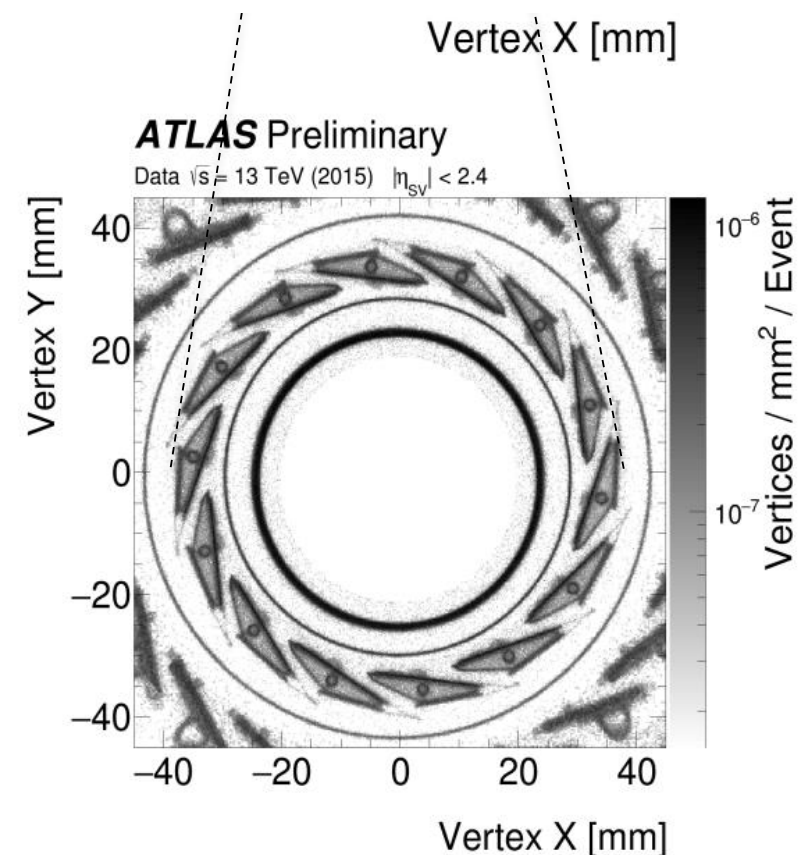
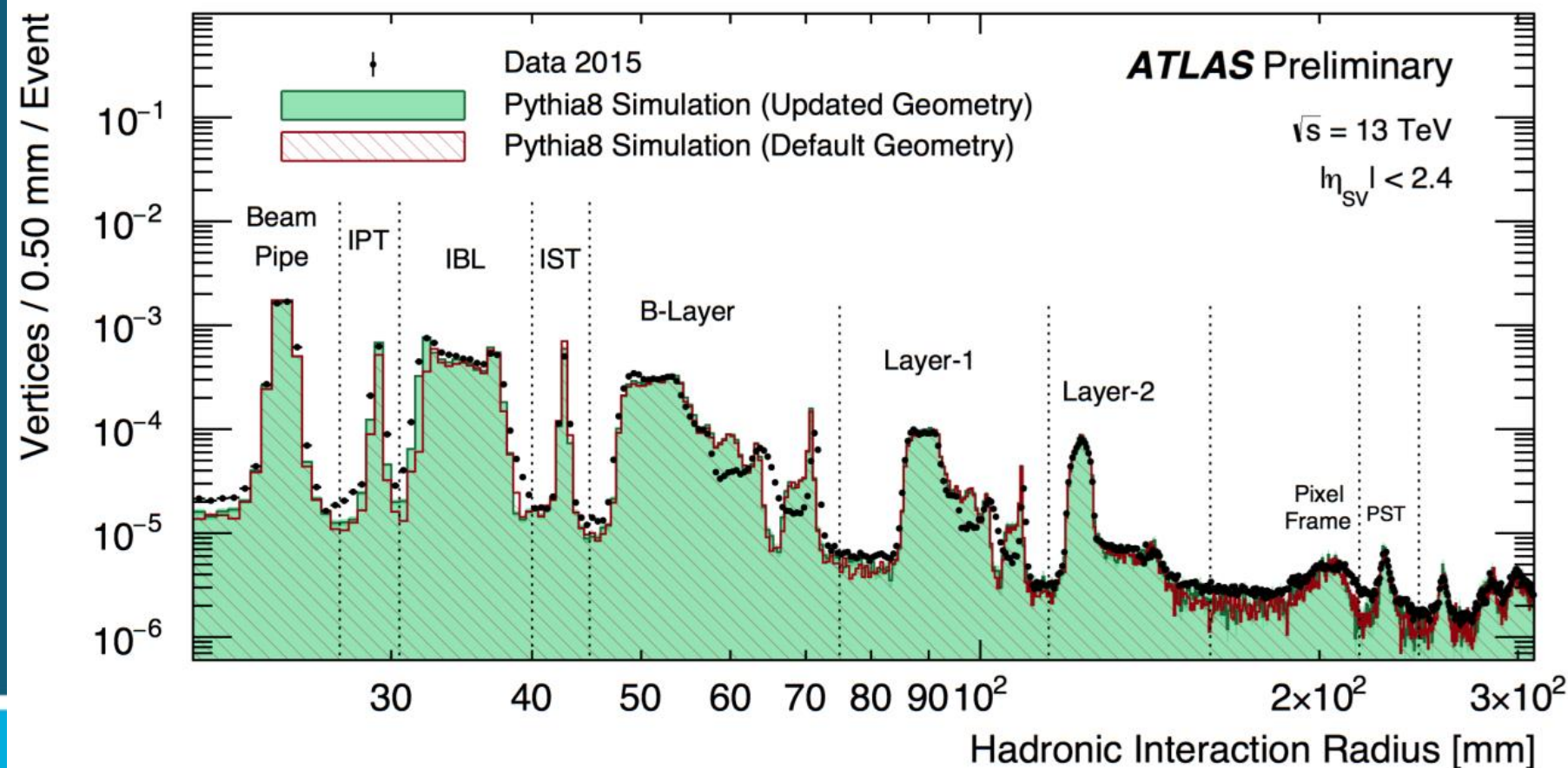
Simulation and understanding detectors

We use **simulations** to model the detector as **accurately** and **precisely** as possible

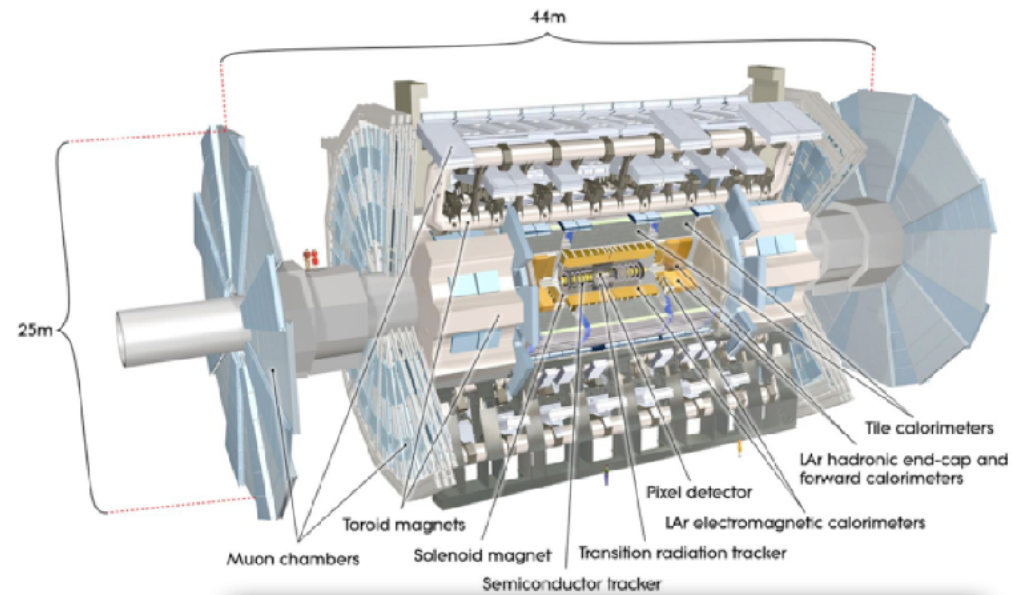
We then **test** that our simulations are accurate **using real data**

We correct our simulations if necessary

Once our simulation is an **accurate model** of our detector, we can use it to **correct the data for detector response**

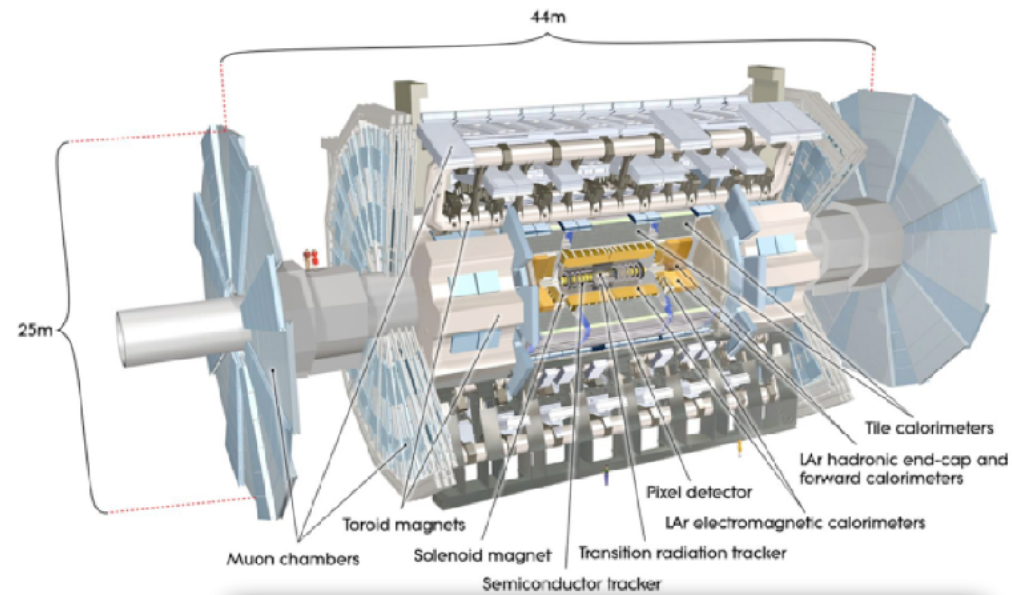


Exabyte-scale physics analysis

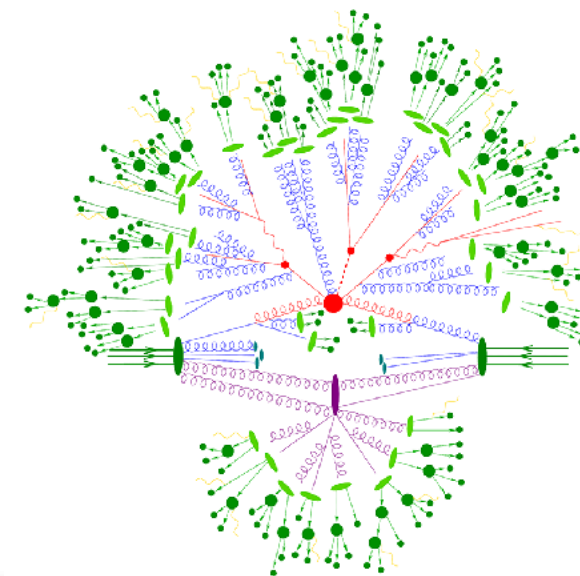


Exabytes of Data

Exabyte-scale physics analysis

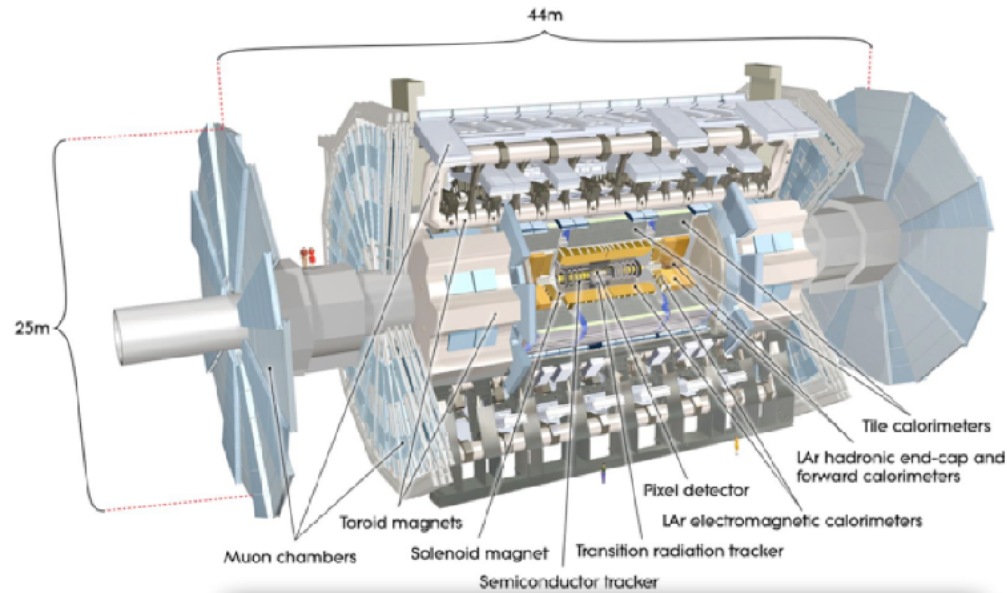


Exabytes of Data

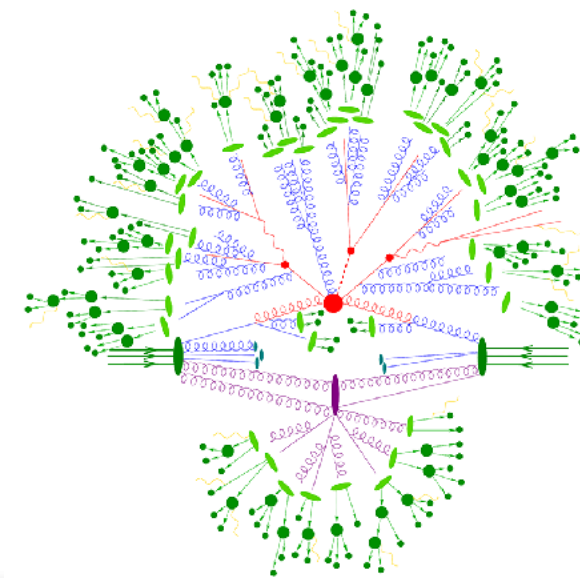


Exabytes of Simulation

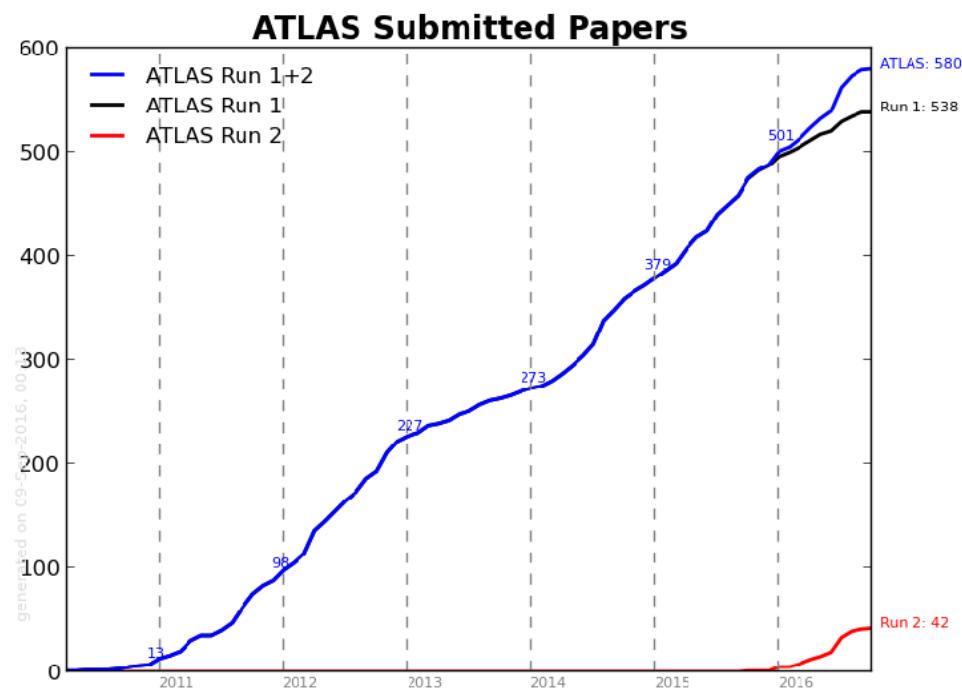
Exabyte-scale physics analysis



Exabytes of Data

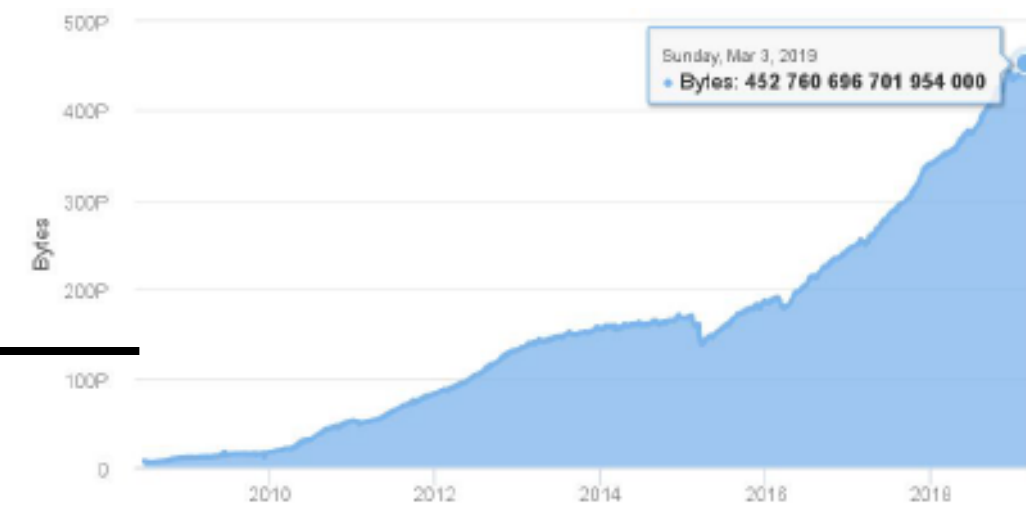
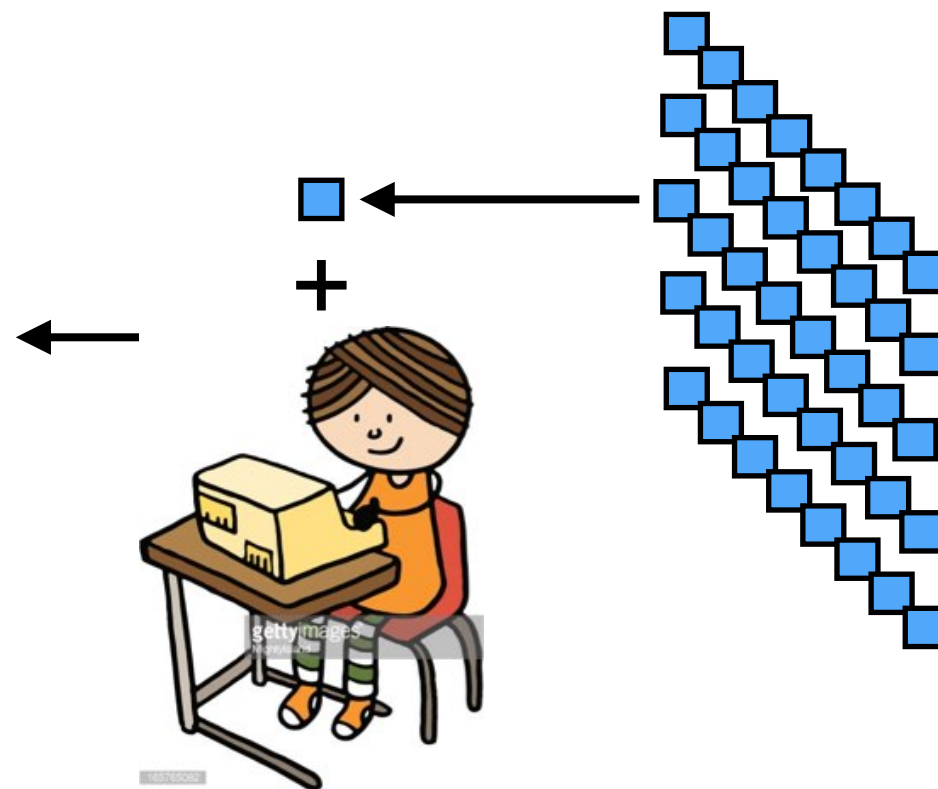
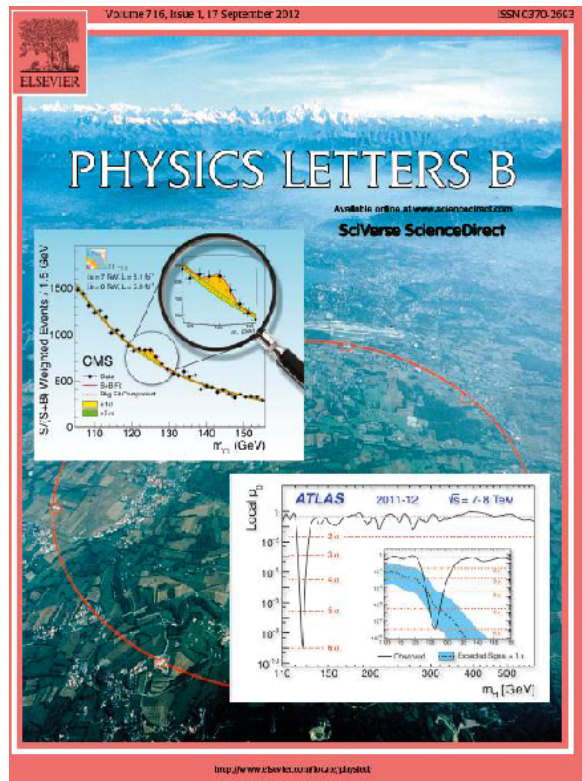


Exabytes of Simulation



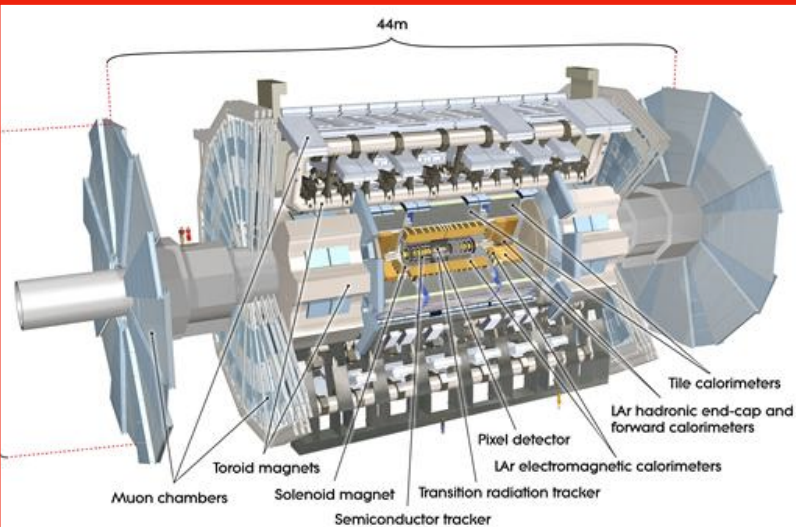
Publish!

Data analysis



***Analysis is performed on a fraction of the data
(more on that later!)***

Data's journey



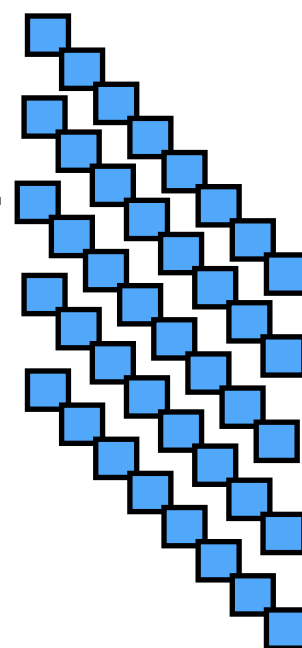
Trigger/DAQ



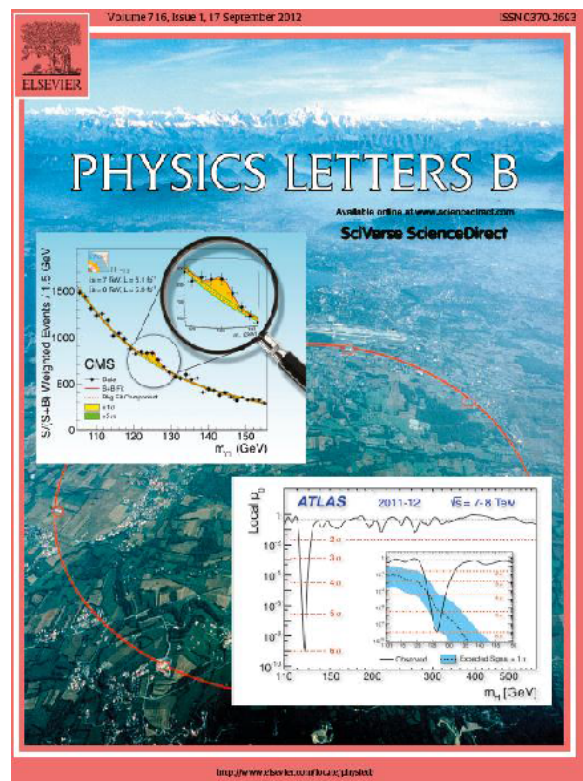
Data
Preparation



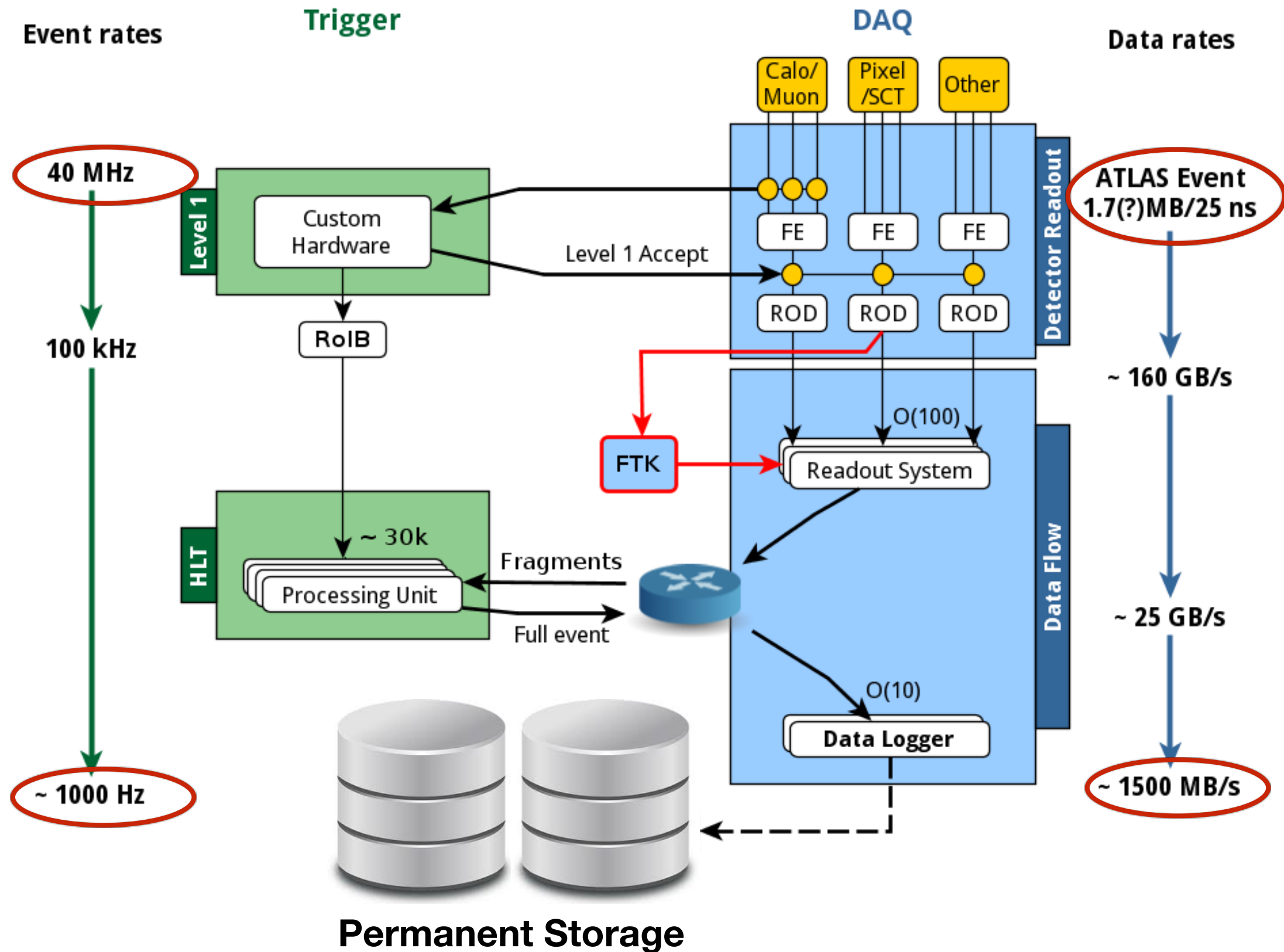
Distributed
computing



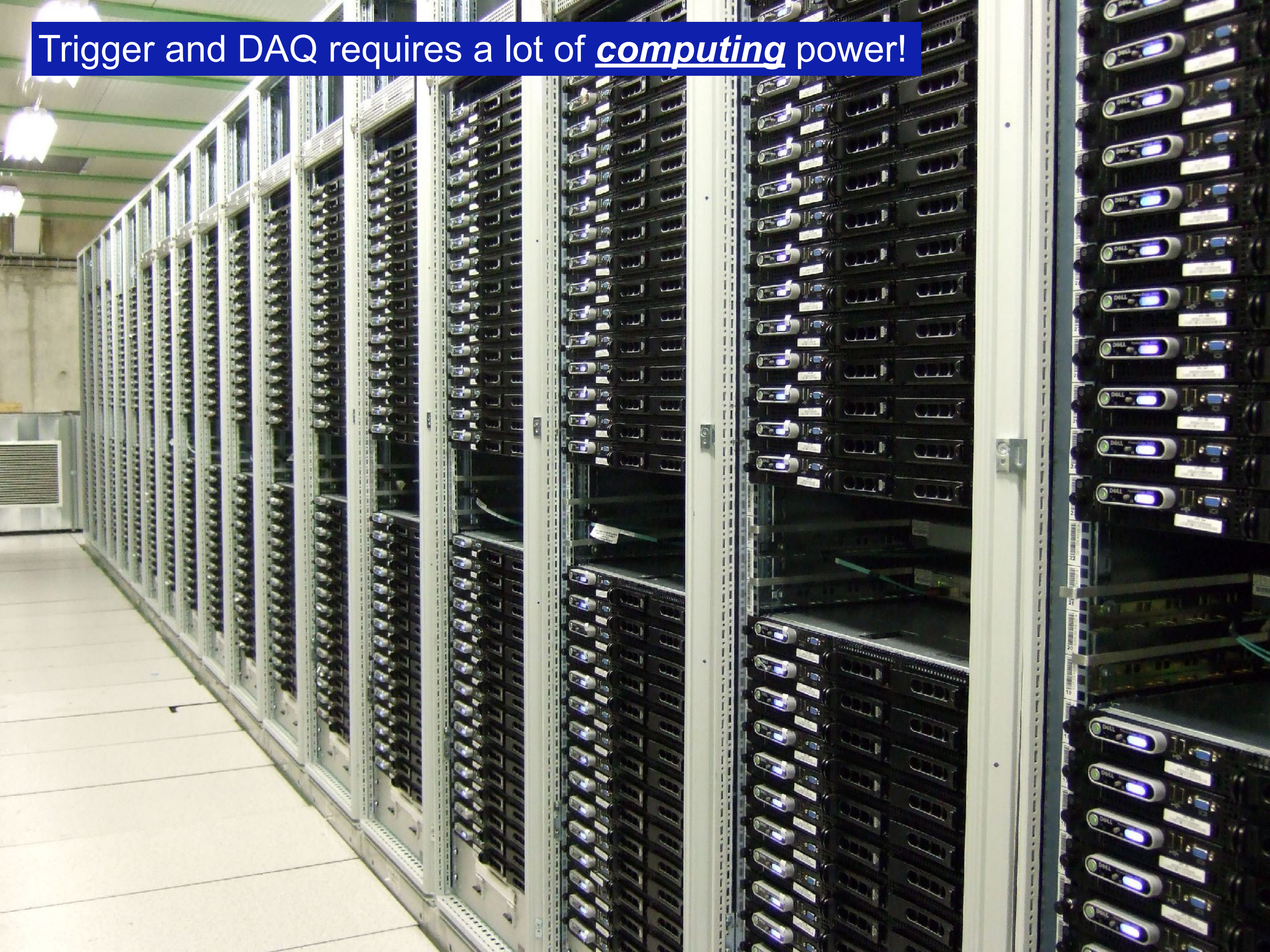
+



The Atlas Trigger and DAQ



Trigger and DAQ requires a lot of computing power!





Data Preparation

Three major steps to **prepare data for physics analysis** and achieve

- reliable, high quality data (yes, we ***reject*** low quality data)
- the ***best performance*** from our detectors
- readiness for ***physics analysis***

1. **Reconstruct physics signals** from the data

- Produce information like how many muons does the event have?

Muon
Spectrometer

Hadronic
Calorimeter

Electromagnetic
Calorimeter

Solenoid magnet

Tracking

Transition
Radiation
Tracker

Pixel/SCT detector

Proton

Neutron

Muon

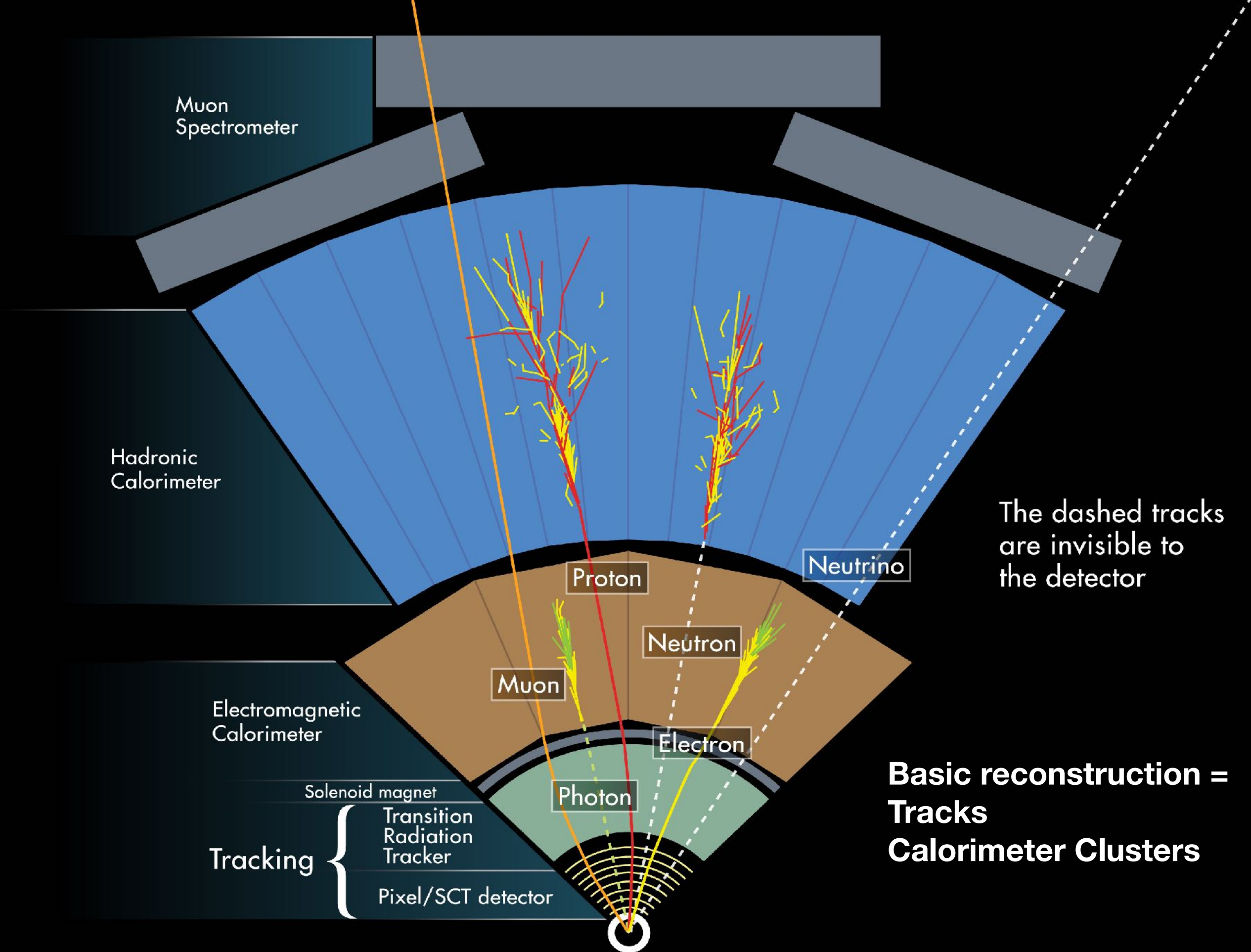
Electron

Photon

Neutrino

The dashed tracks
are invisible to
the detector

**Basic reconstruction =
Tracks
Calorimeter Clusters**



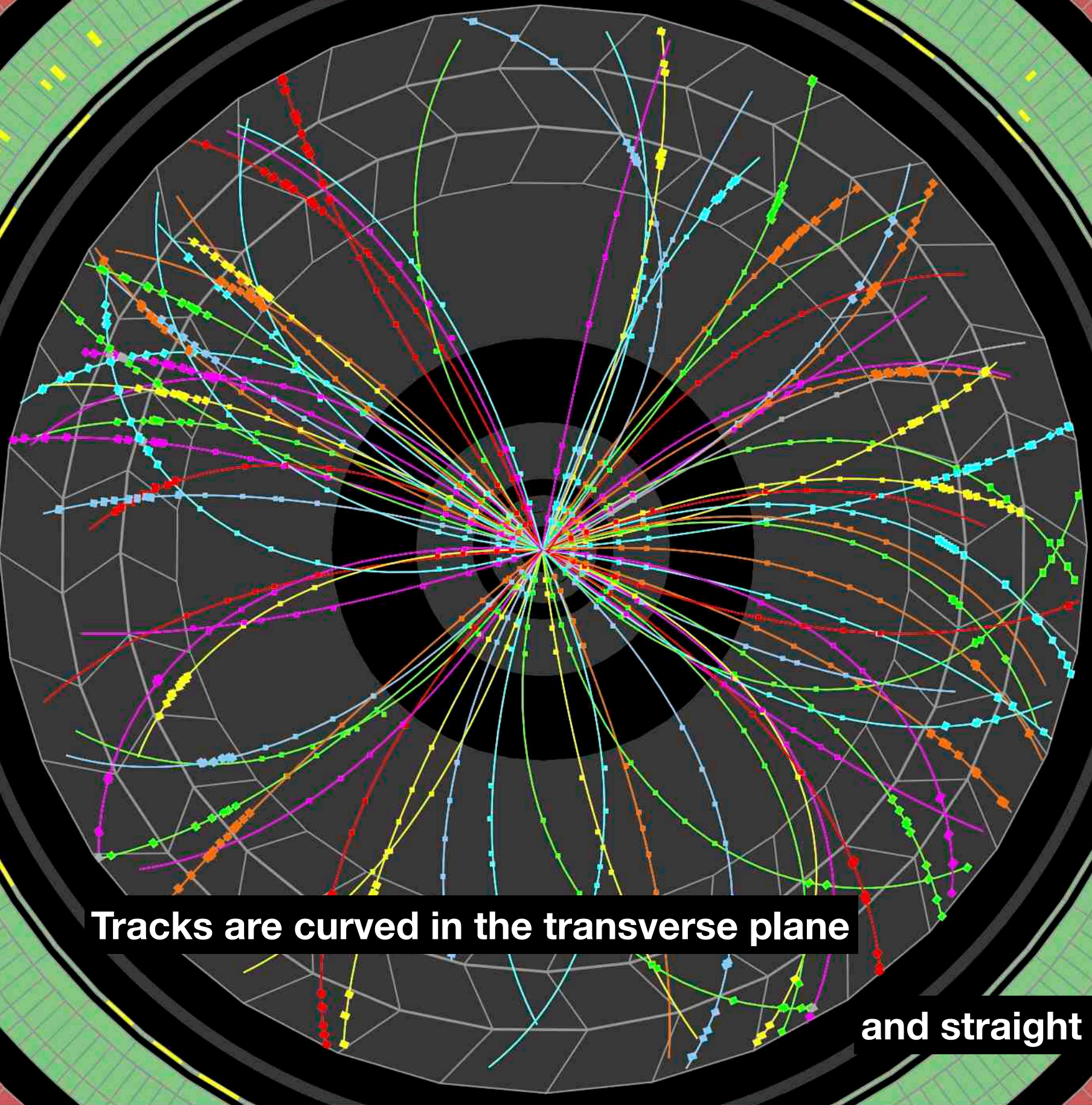


ATLAS

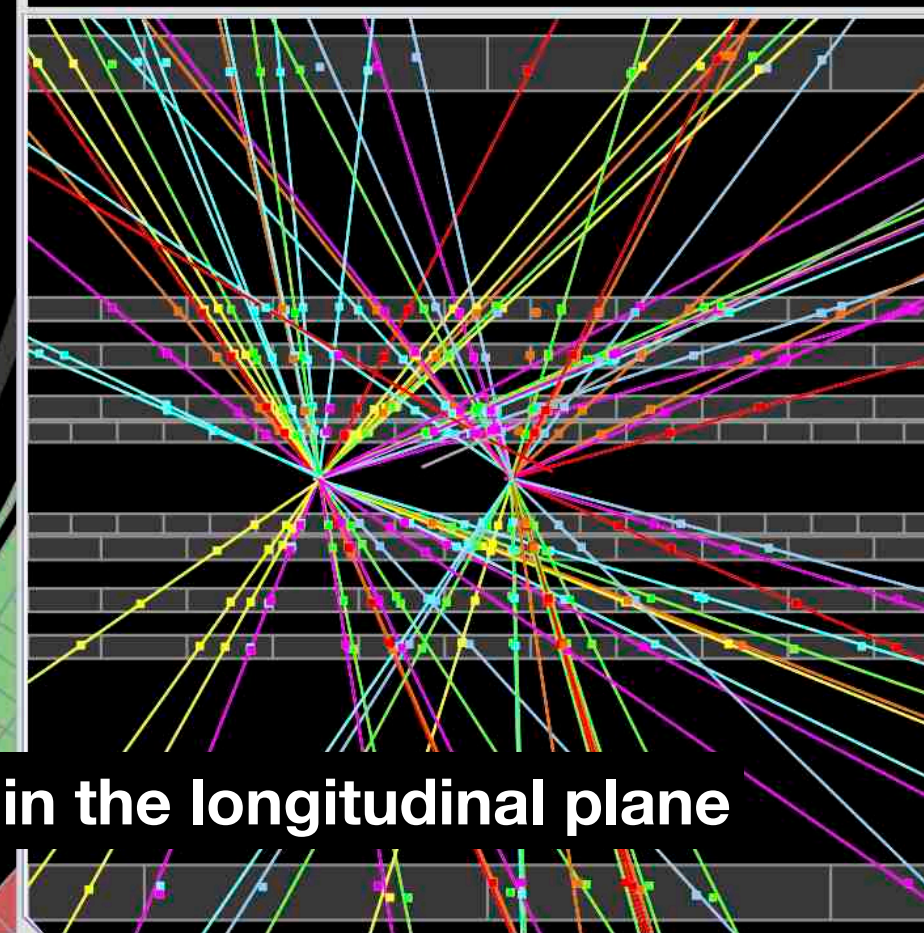
EXPERIMENT

Run Number: 265545, Event Number: 5720351

Date: 2015-05-21 10:39:54 CEST



Tracks are curved in the transverse plane



and straight in the longitudinal plane

Track fitting

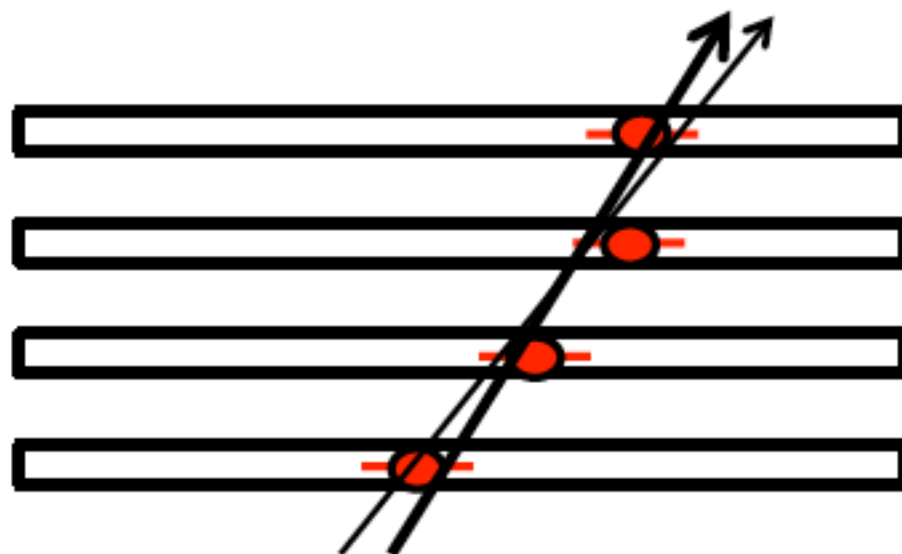
◎ Perfect measurement – ideal



◎ Imperfect measurement – reality



◎ Small errors and more points help to constrain the possibilities




◎ Quantitatively:

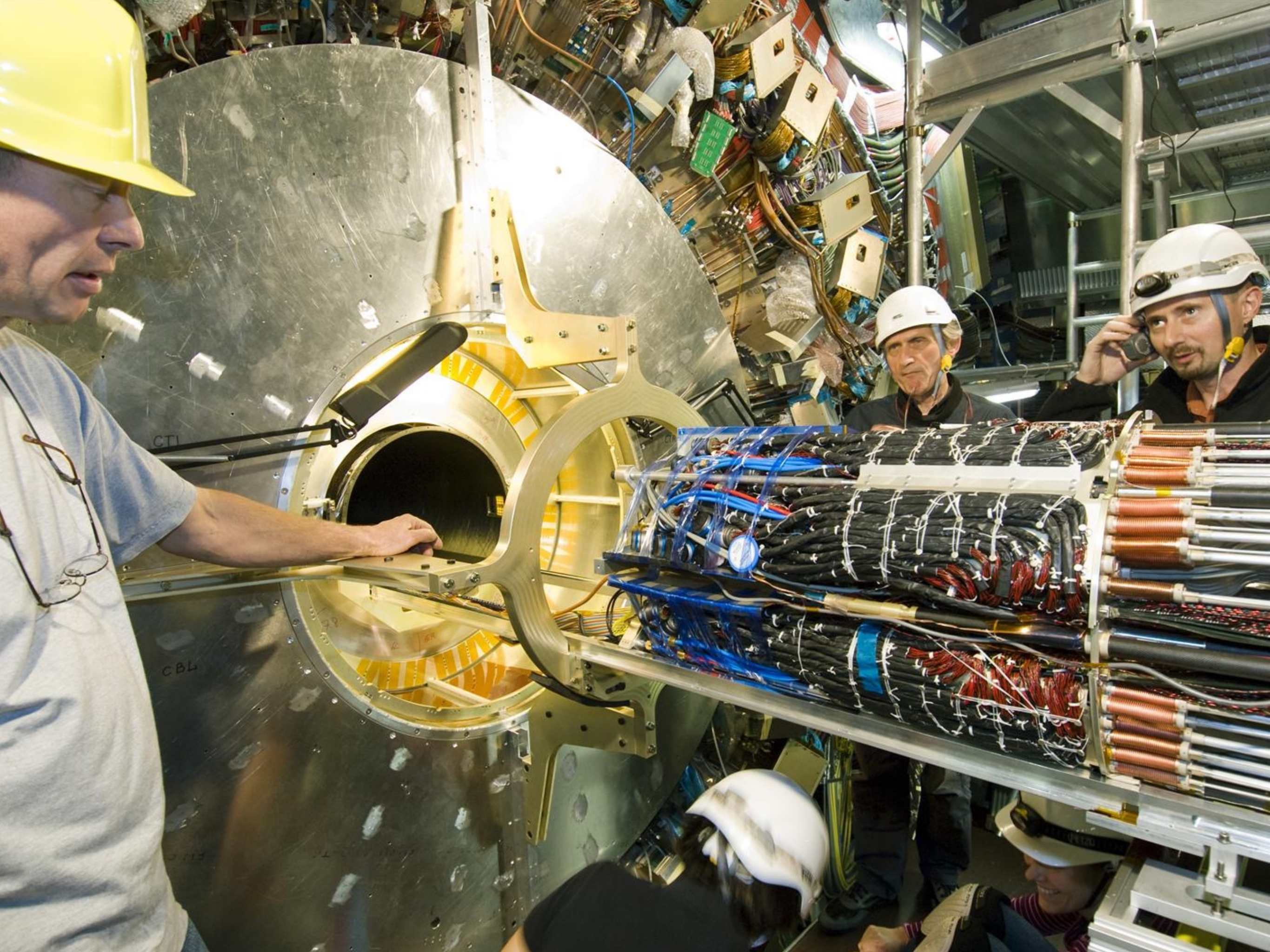
- ◎ Parameterize the track;
- ◎ Find parameters by Least-Squares-Minimization;
- ◎ Obtain also uncertainties on the track parameters.

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 - Produce information like how many muons does the event have?
2. **Calibrate** the detectors
 - Correct imperfections, account for changes over time...



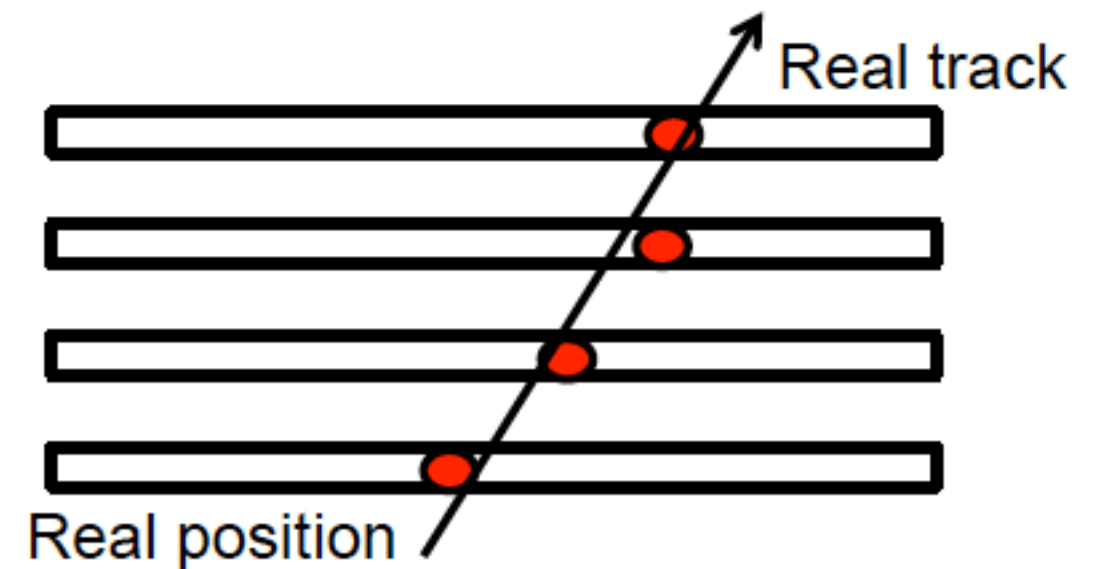
Real detector effects

⊙ Presence of Material

- ⊙ Coulomb scattering off the core of atoms
- ⊙ Energy loss due to ionization
- ⊙ Bremsstrahlung
- ⊙ Hadronic interaction

⊙ Misalignment

- ⊙ Detector elements not positioned in space with perfect accuracy.
- ⊙ Alignment corrections derived from data and applied in track reconstruction.



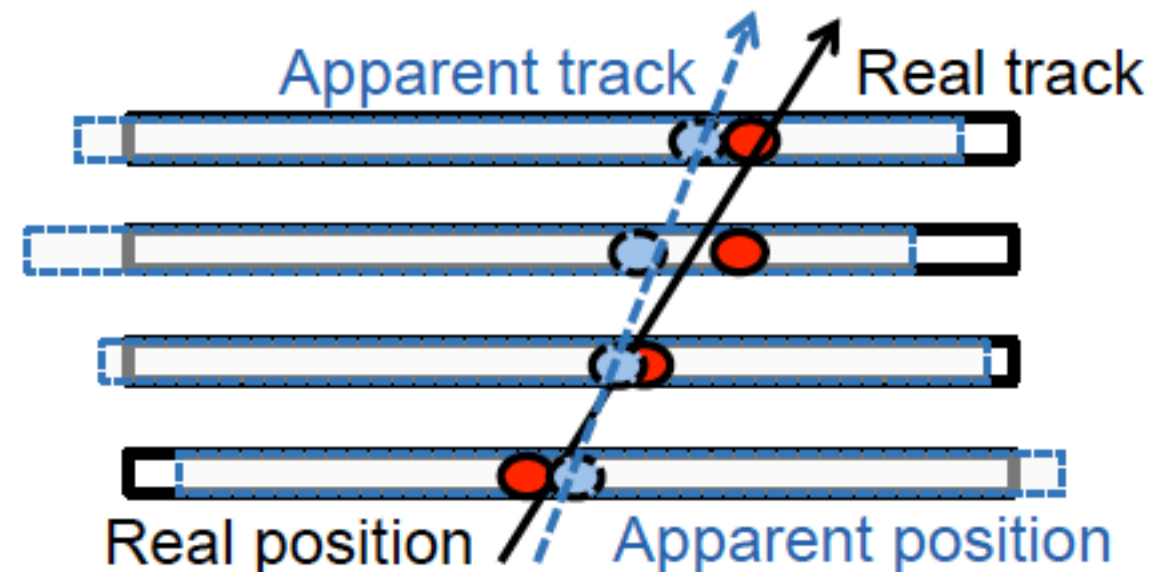
Correcting detector effects - calibration

◎ Presence of Material

- ◎ Coulomb scattering off the core of atoms
- ◎ Energy loss due to ionization
- ◎ Bremsstrahlung
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- Correct imperfections, account for changes over time...

3. Make sure that the **data quality** is excellent, also in real time

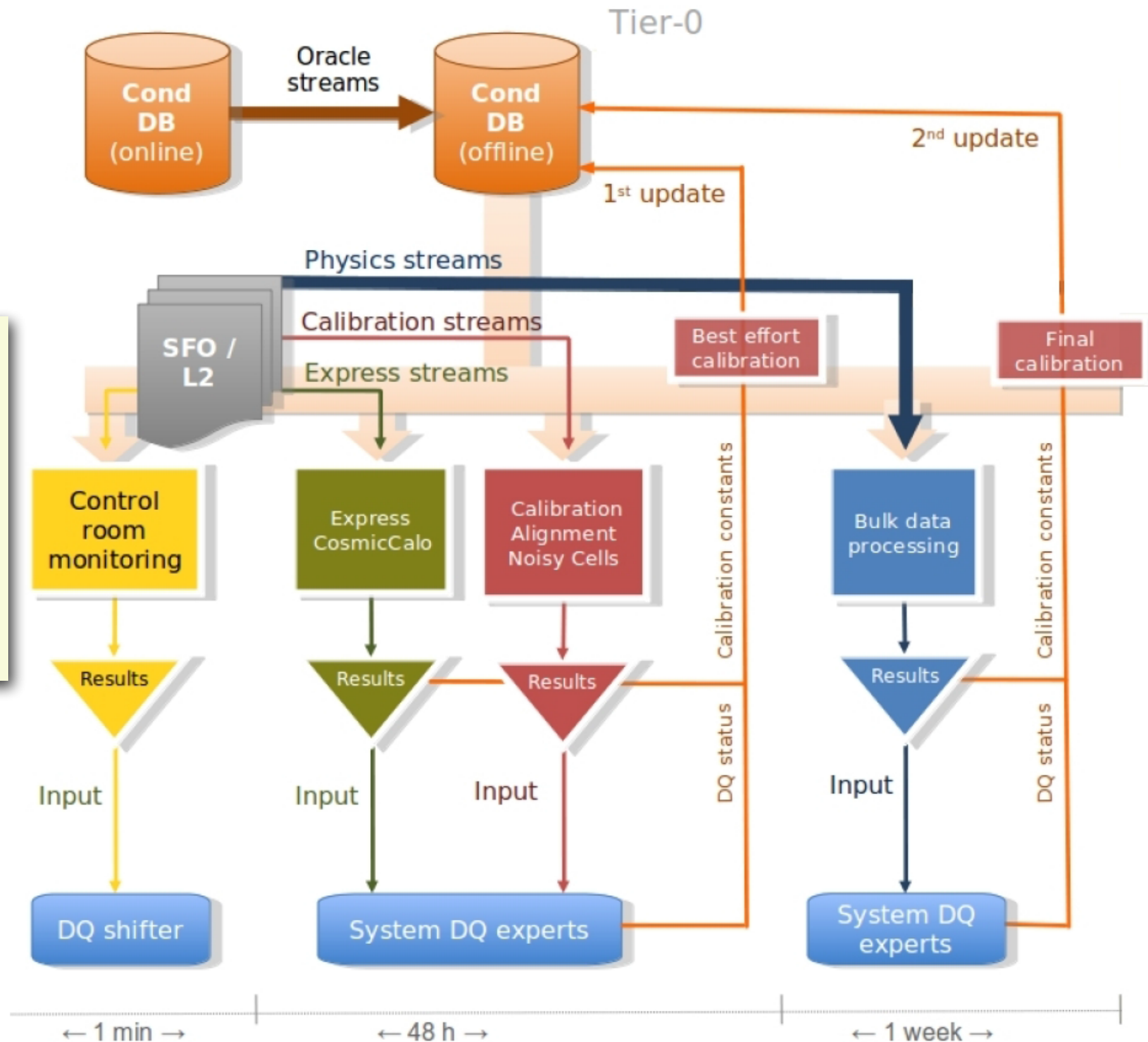
- Maximise the amount of useful data

Data Quality

Check during data taking

Check a fraction of the data with a quick calibration

Check all of the data with the best calibration - publish this data !!



Data Preparation

Three major steps to **prepare data for physics analysis** and achieve

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1. **Reconstruct physics signals** from the data

- Produce information like how many muons does the event have?

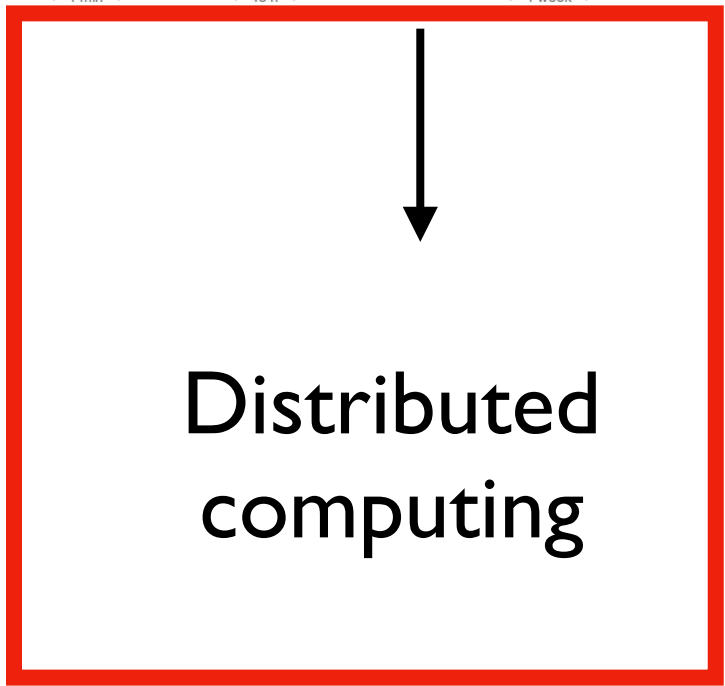
2. **Calibrate** the detectors

- Correct imperfections, account for changes over time...

3. Make sure that the **data quality** is excellent, also in real time

- Maximise the amount of useful data







The Worldwide LHC Computing Grid

- Now the data has been ***prepared for physics analysis***, it's time to extract our favourite physics signal!
- Many experiments, particularly those at the **LHC**, use computing sites all over the world via **the grid** to
 - harness all of that ***computing power***
 - enable collaborators ***worldwide*** to access the data

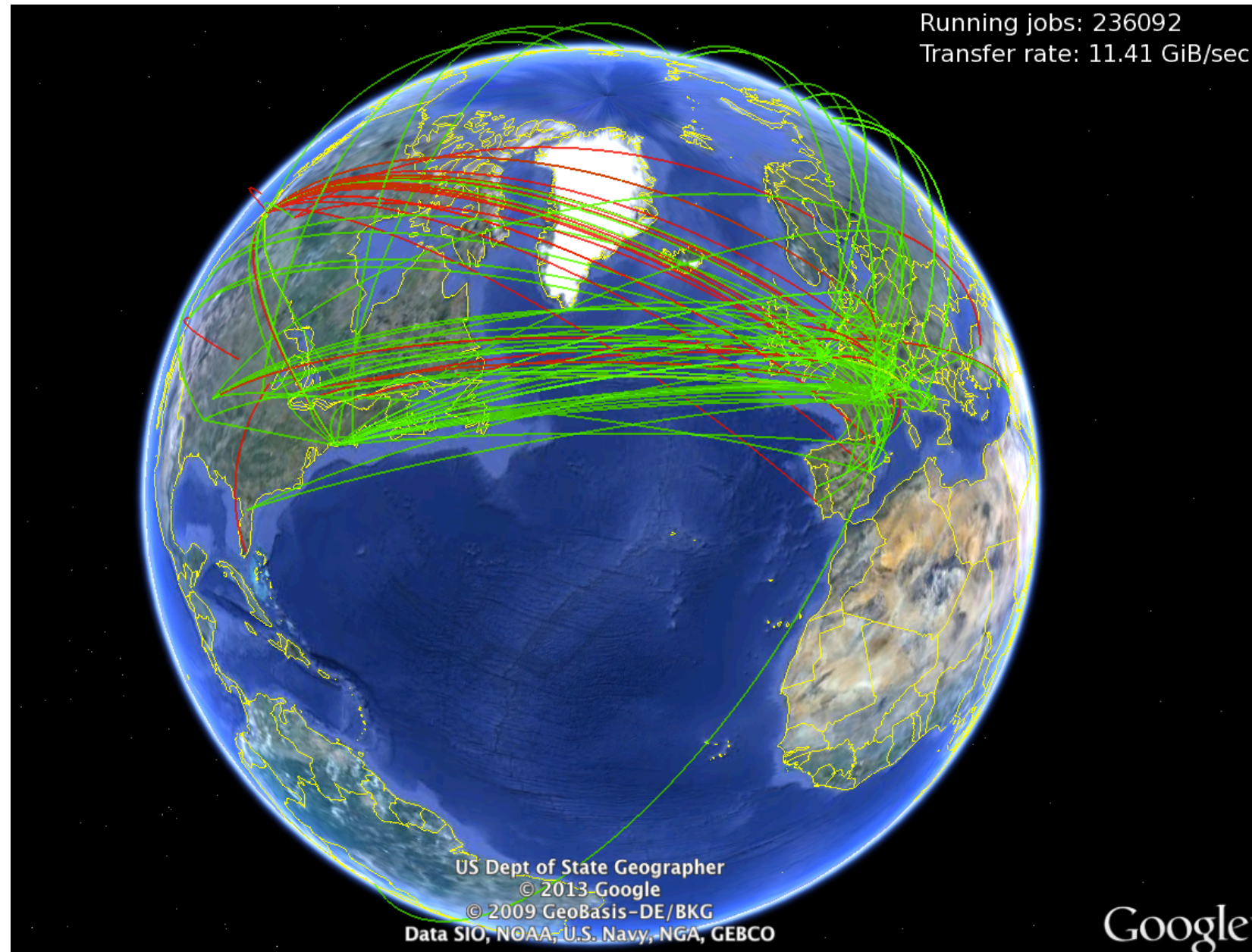


Image taken from the **WLCG** GoogleEarth Dashboard

- <http://wlcg.web.cern.ch/wlcg-google-earth-dashboard>



Needles in haystacks

- There are billions of events and the ones we are really interested in are **very rare**
- Often the interesting events are also **very difficult to distinguish** from background
 - Requires **high precision detectors**, which means **lots of data** for each event
- The data are structured but each event is different - **unique data science challenge**

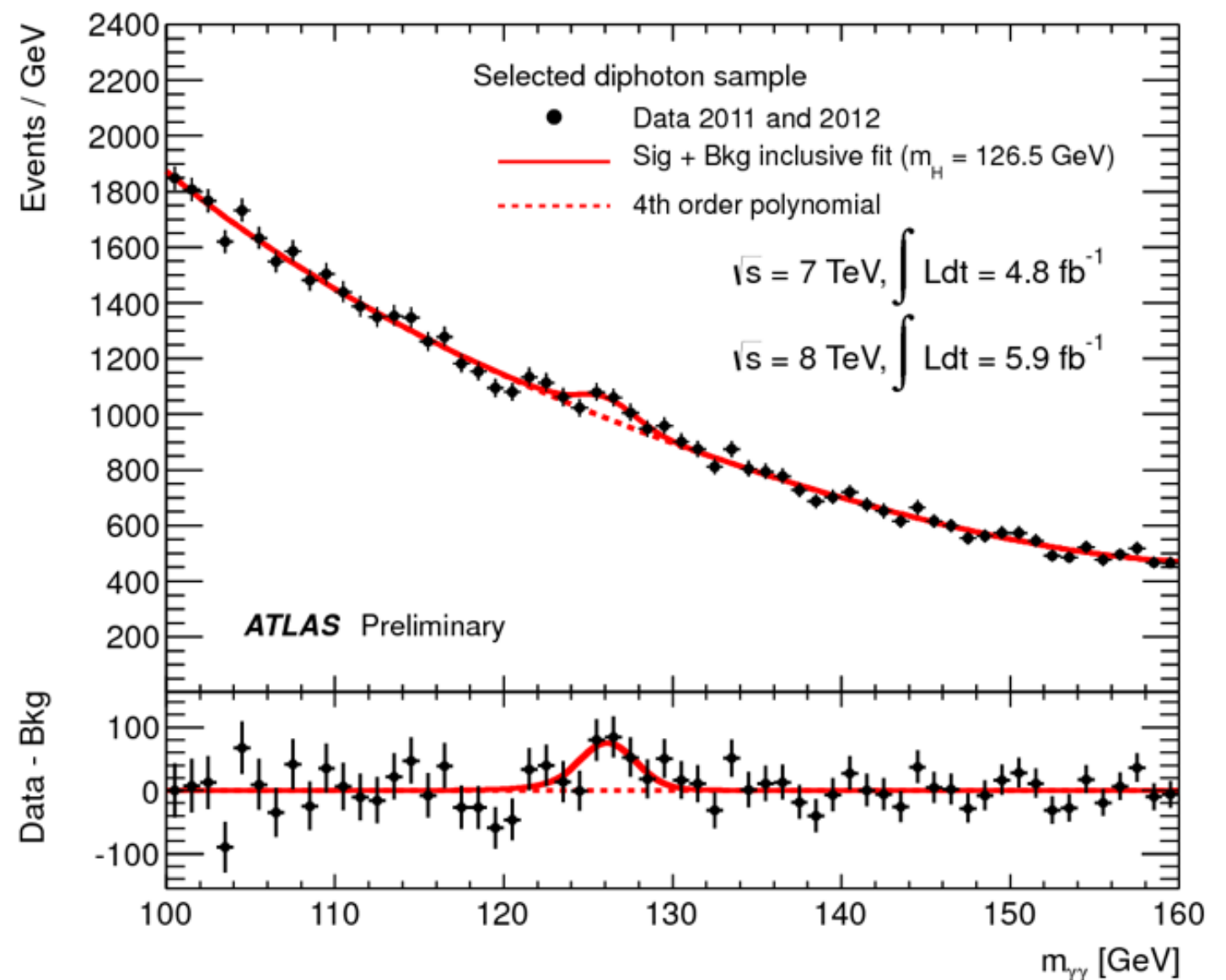
Data reduction proceeds via a two-pronged approach:

Select only the events that you are interested in

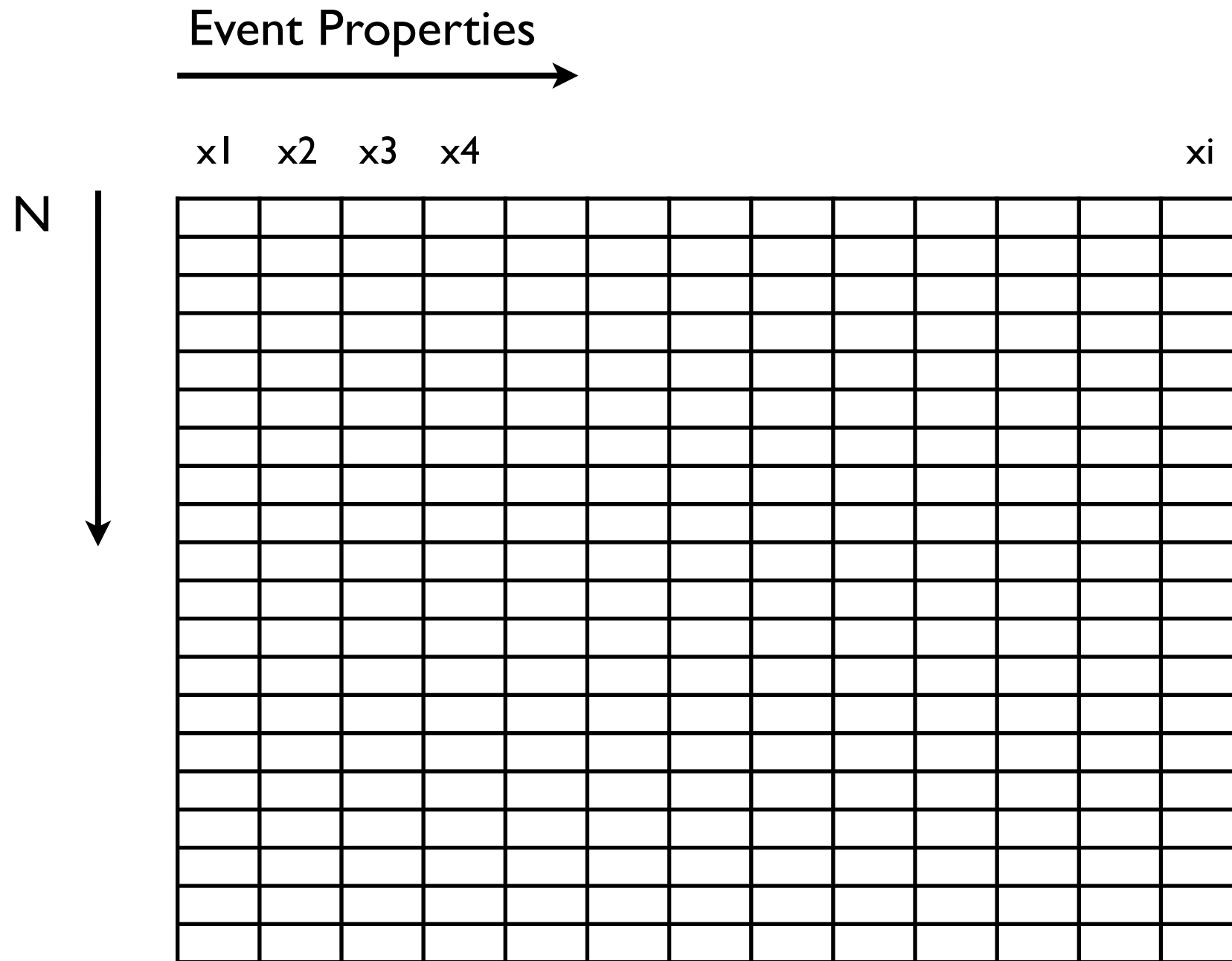
- e.g. events with two photons

Keep only the information you need

Final statistical inference is only performed on the reduced data

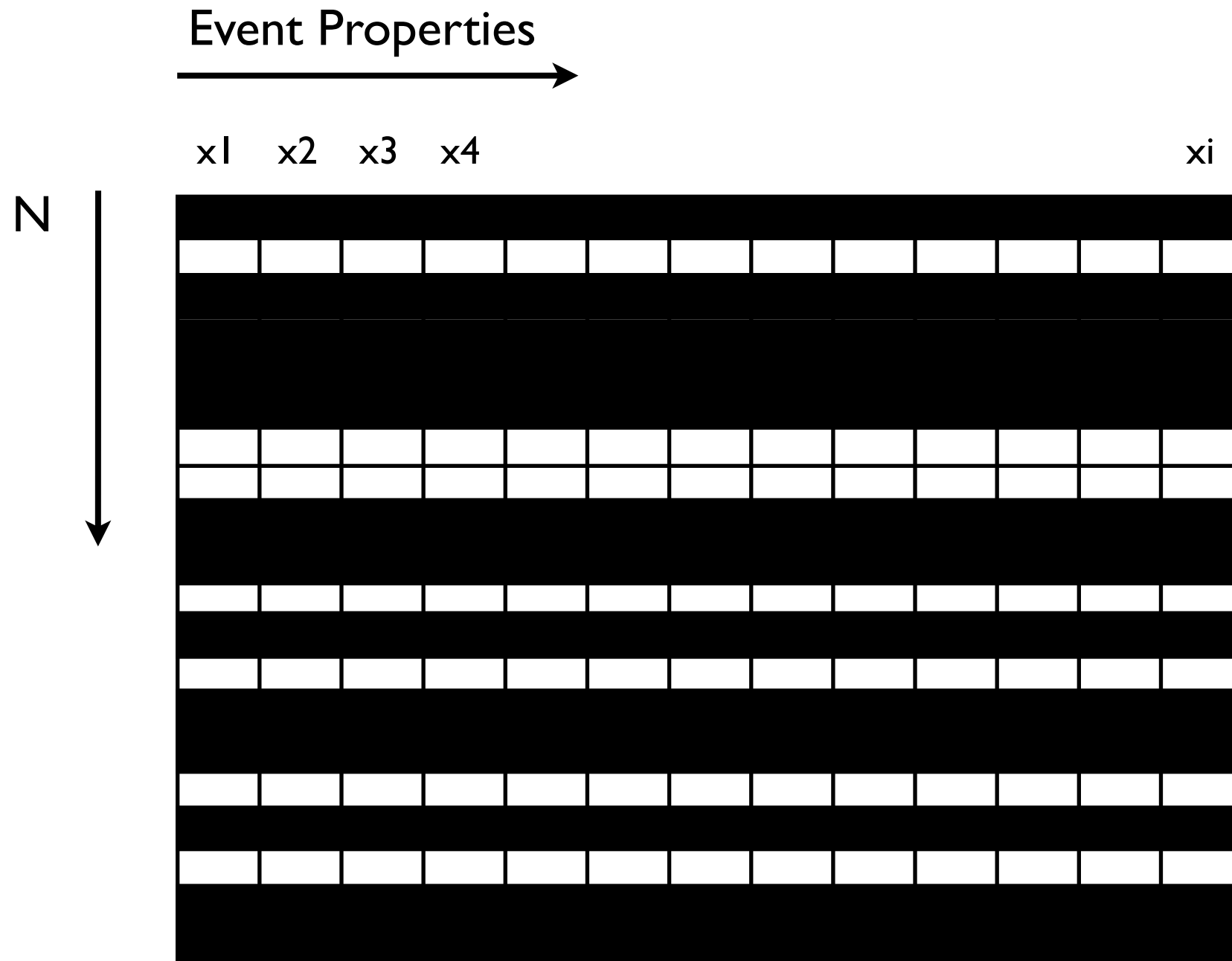


Slice and dice - data reduction



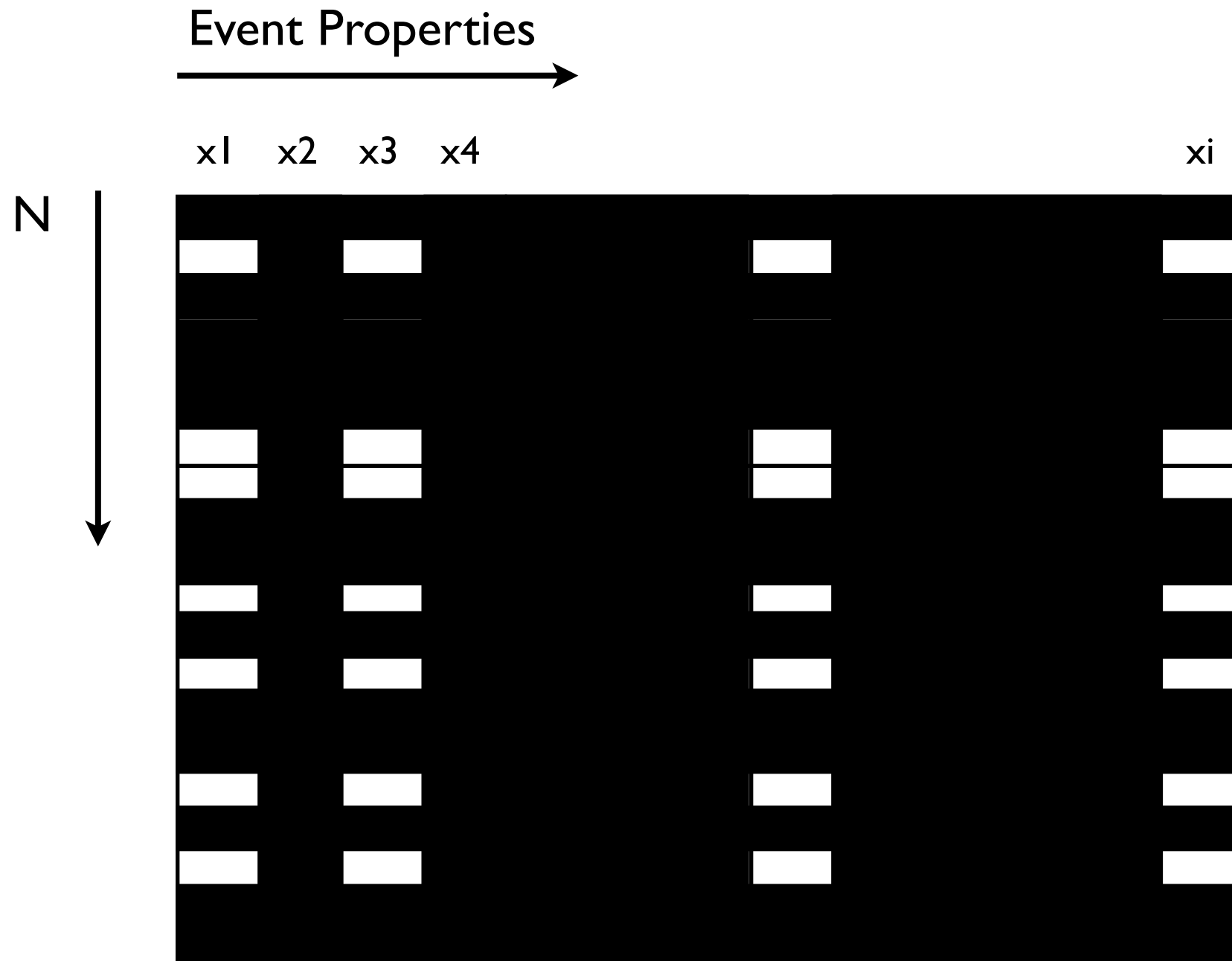
- N events * i event properties makes for a LOT of data !

Slice and dice - data reduction



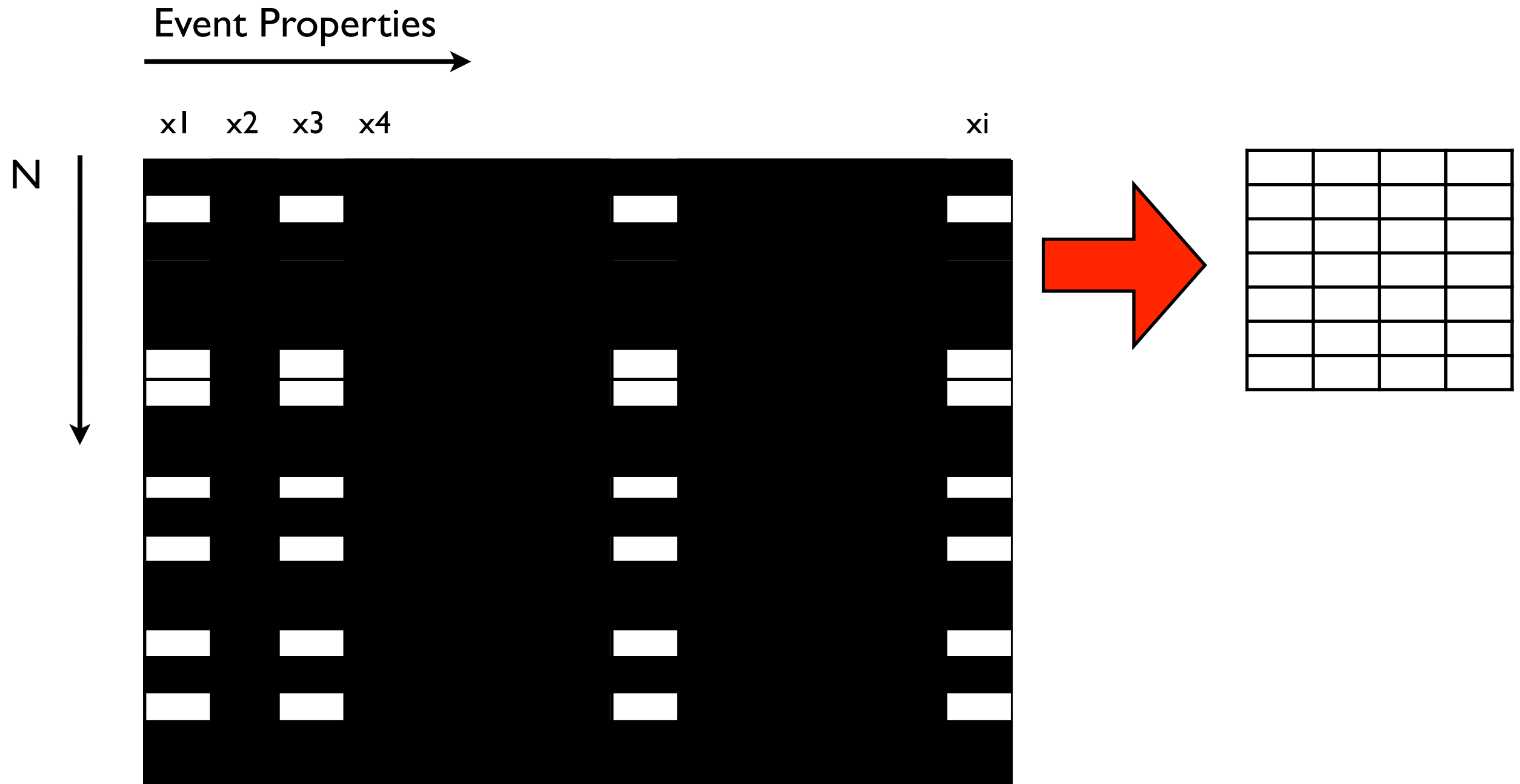
- Select *only the interesting* events

Slice and dice - data reduction



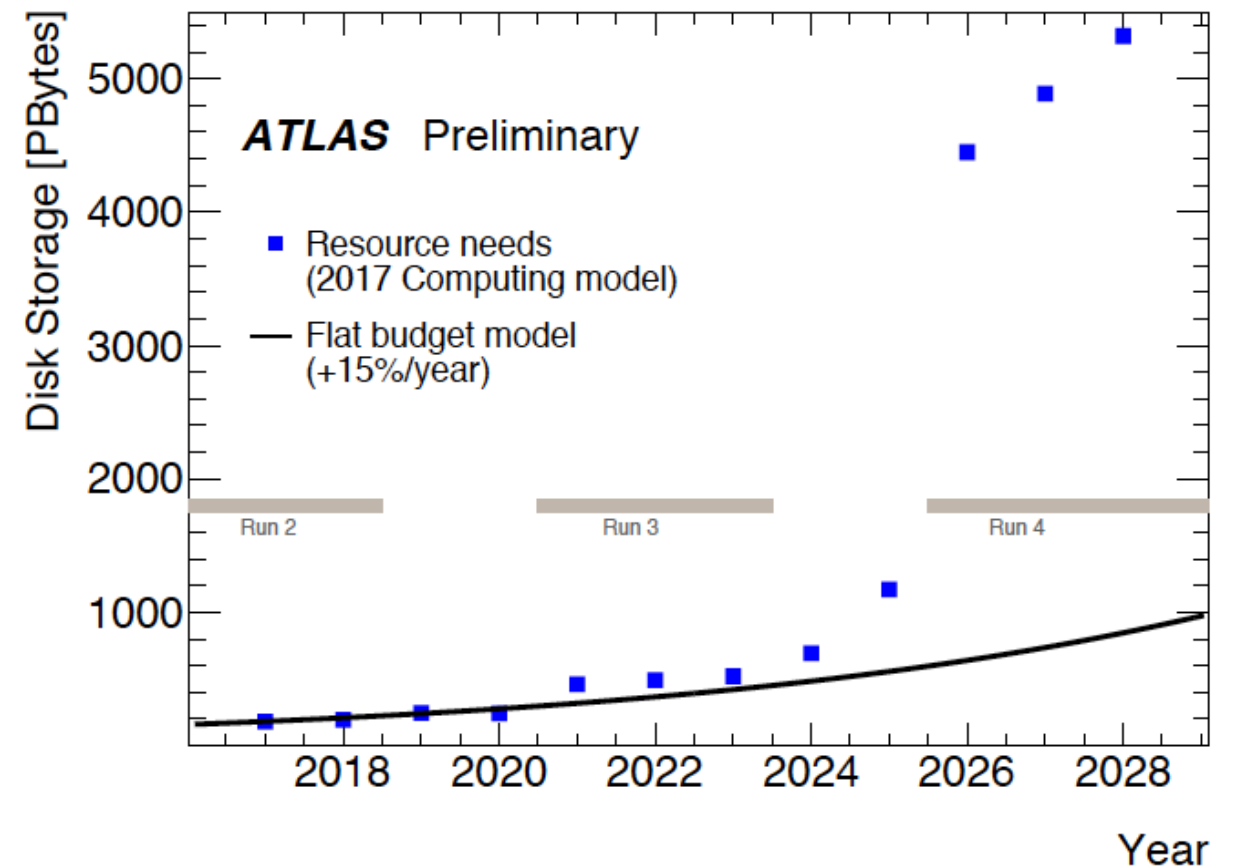
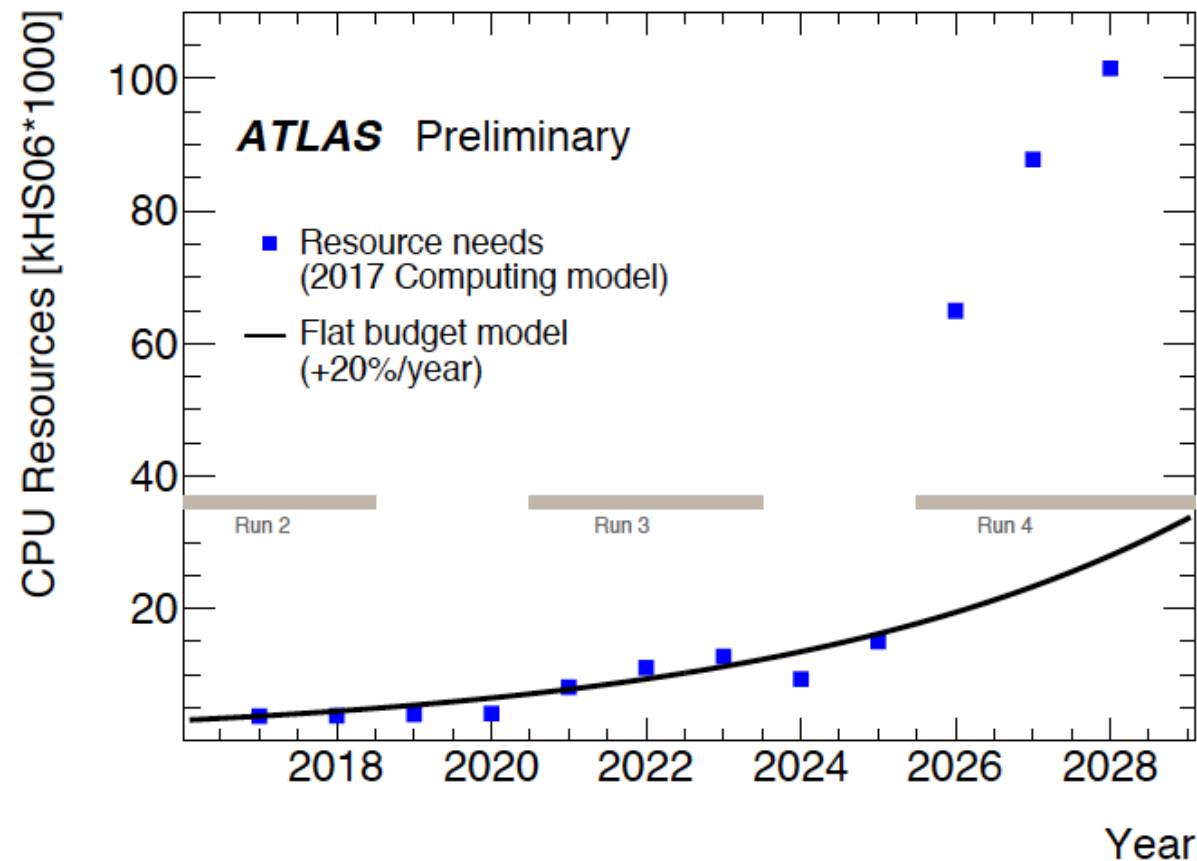
- Select ***only the interesting*** event properties

Slice and dice - data reduction



- From now on you only need to process a smaller amount of data

Future software and computing



The data challenge is going to get a lot harder than standard progress can handle, c.f. HEP Software Foundation paper on software and computing R&D for the 2020s:

<http://arxiv.org/pdf/1712.06982.pdf>

Conclusions

Physics experiments produce a lot of data!

- **Triggers** reduce the raw event rate
- **Data Preparation** is crucial to get the best quality reconstructed and calibrated data
- **Data reduction** is crucial to enable analysis, otherwise there's simply too much data to process

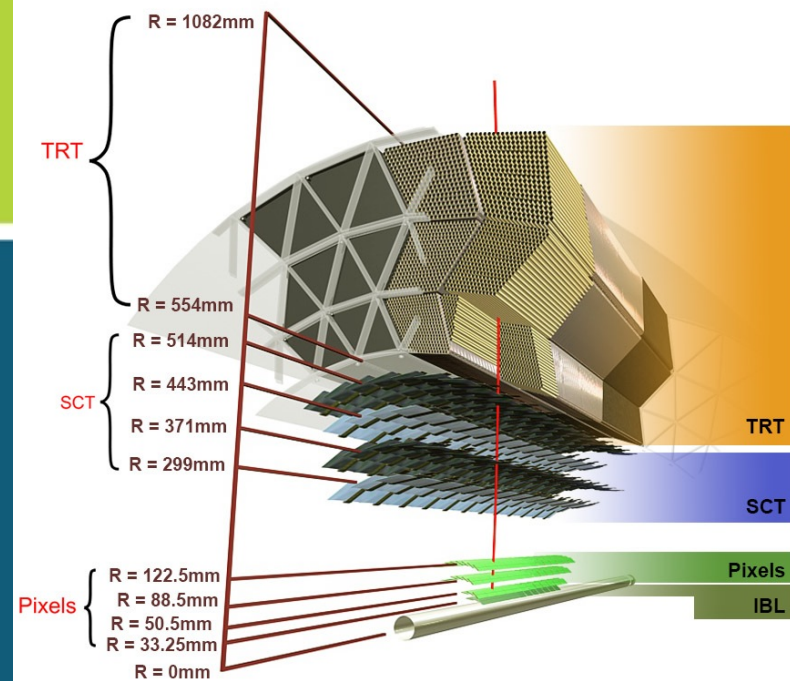
All of this requires *computing, computing, computing!*

- and not only hardware, we need great software too

The future will bring even more computing challenges that will need to be solved

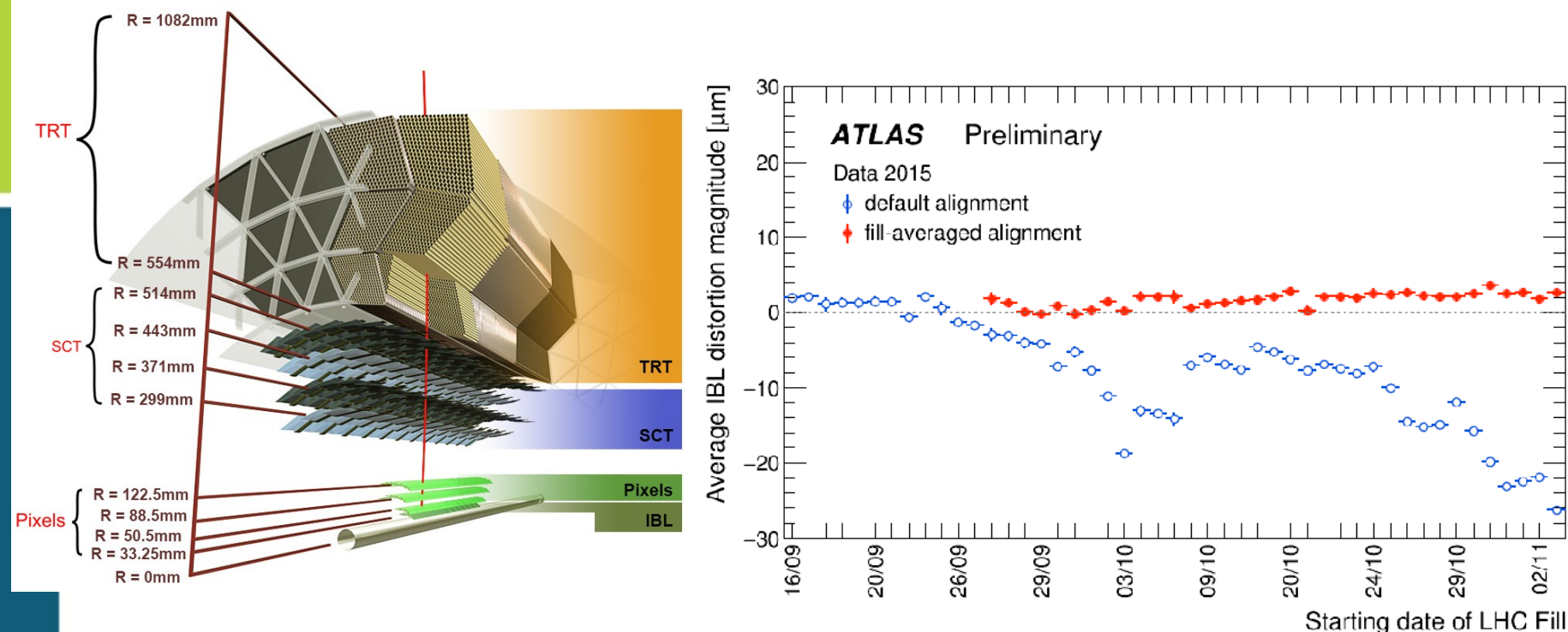
Backup

Calibration



During the break between Run 1 and Run 2, ATLAS inserted the IBL, an extra layer of silicon tracker close to the beam pipe

Calibration

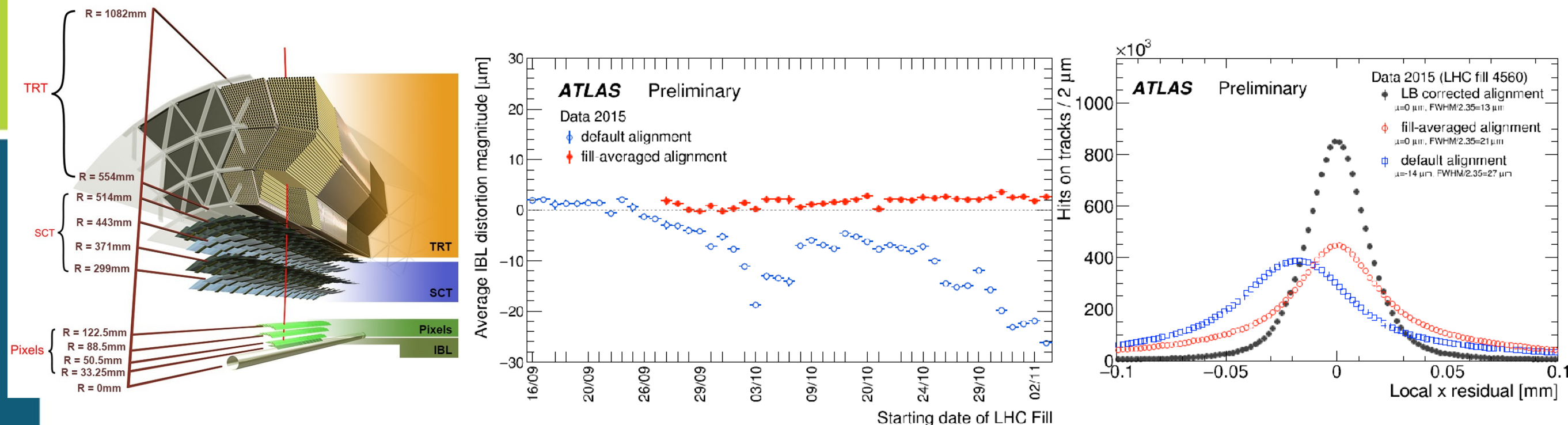


During the break between Run 1 and Run 2, ATLAS inserted the IBL, an extra layer of silicon tracker close to the beam pipe

At the start of data taking in Run 2, it started to move

As time went on, the movement was very significant, much more than the detector precision so the movement could really be seen in physics distributions and data quality

Calibration



During the break between Run 1 and Run 2, ATLAS inserted the IBL, an extra layer of silicon tracker close to the beam pipe

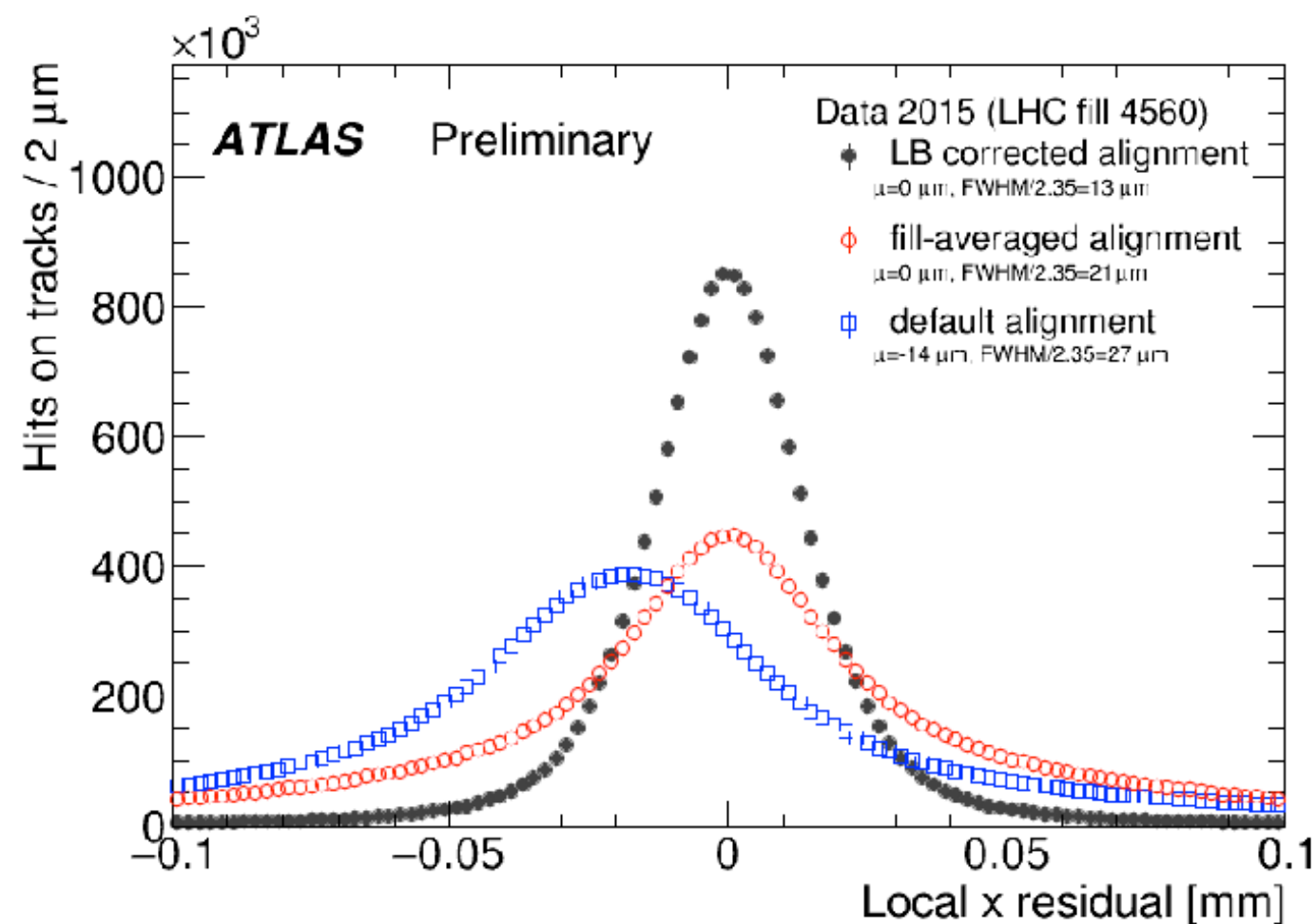
At the start of data taking in Run 2, it started to move

As time went on, the movement was very significant, much more than the detector precision so the movement could really be seen in physics distributions and data quality

ATLAS quickly implemented and commissioned a correction procedure as part of its calibration process

Following the correction the performance of the detector was back to nominal

What makes good data quality?

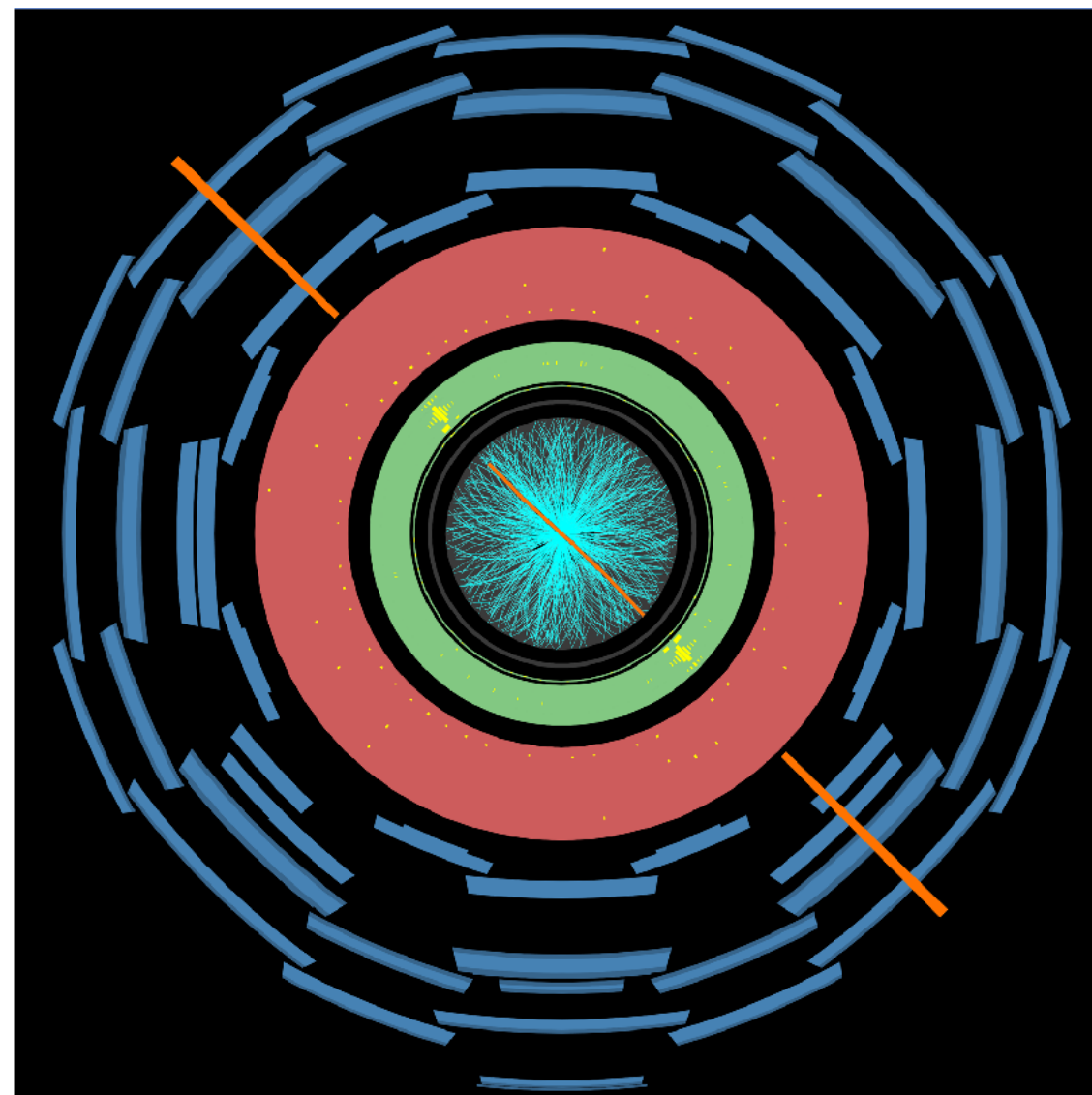
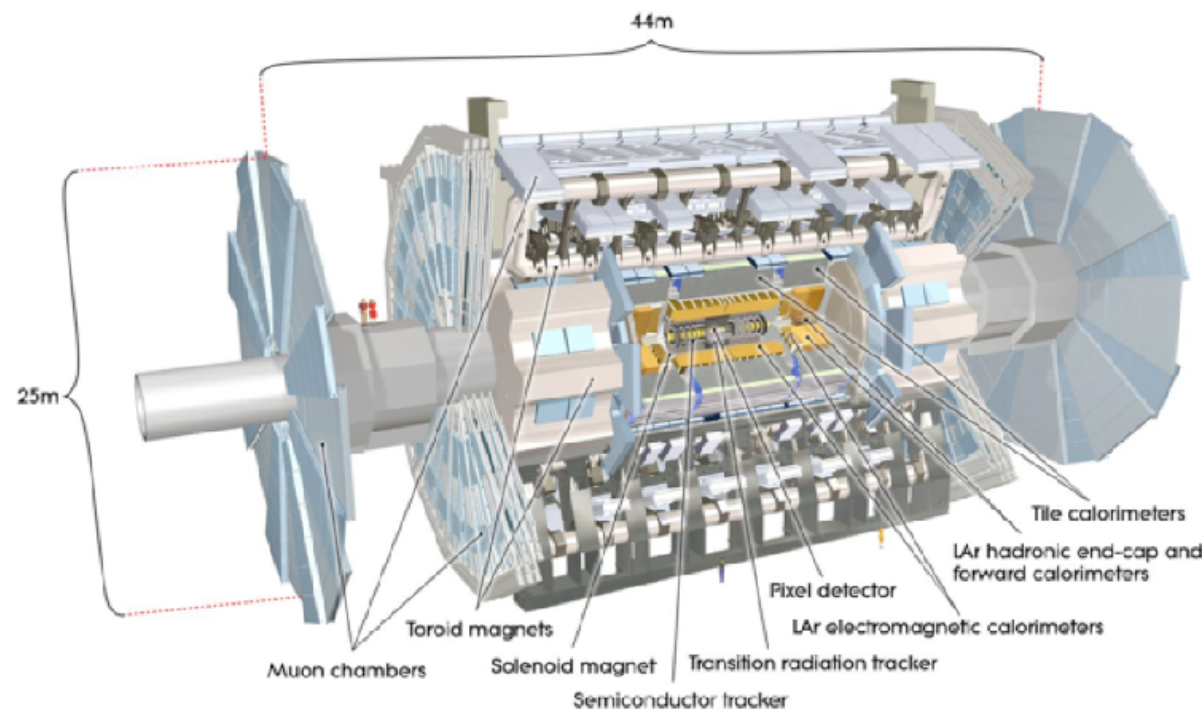


The **ATLAS IBL** is a good example of a data quality plot

We need a reference, here that would be the **black** histogram, how we expect the data to look

If the data quality shifter sees the **blue** or **red** histogram, they will raise the alarm!

What am I?

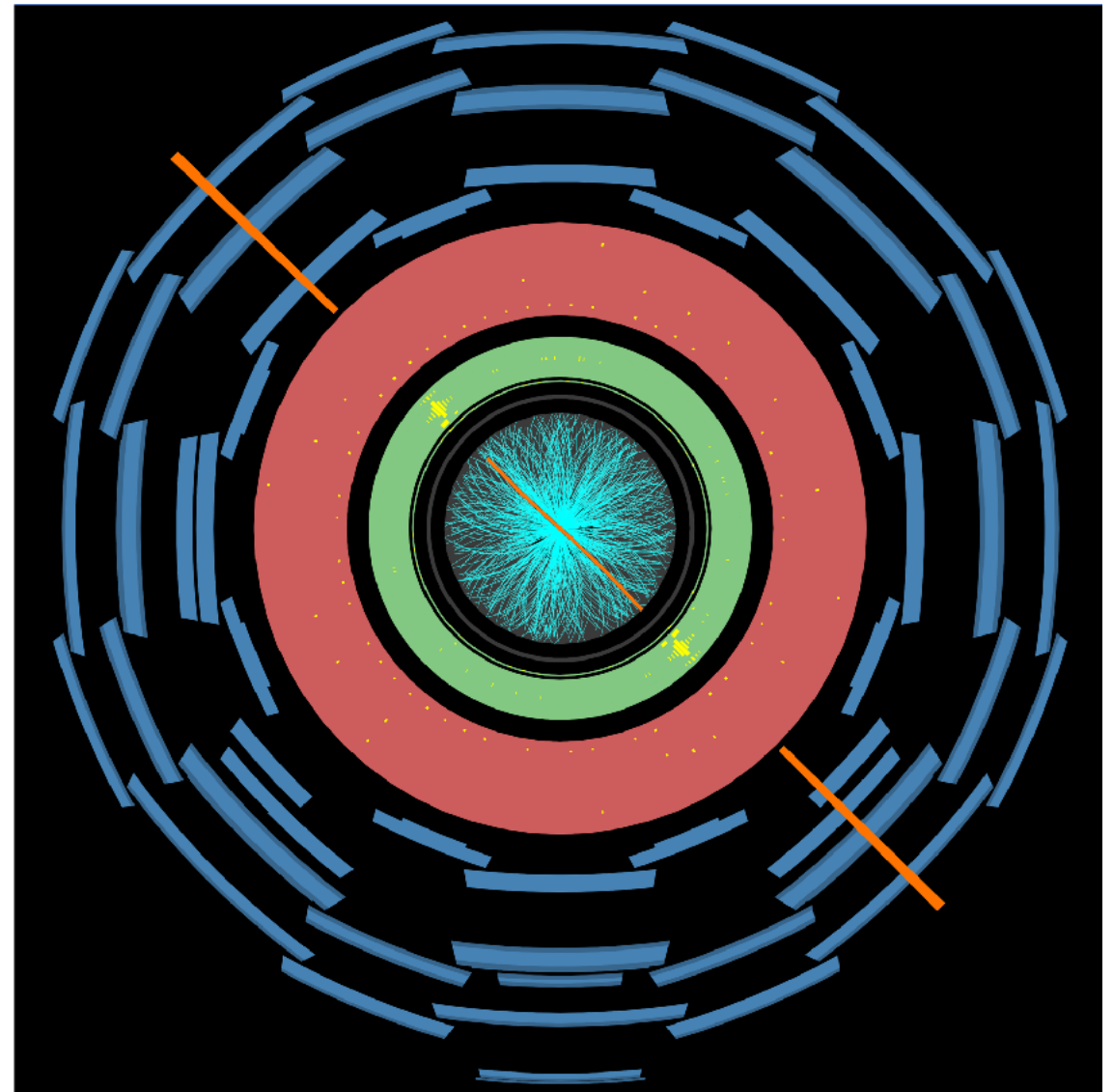
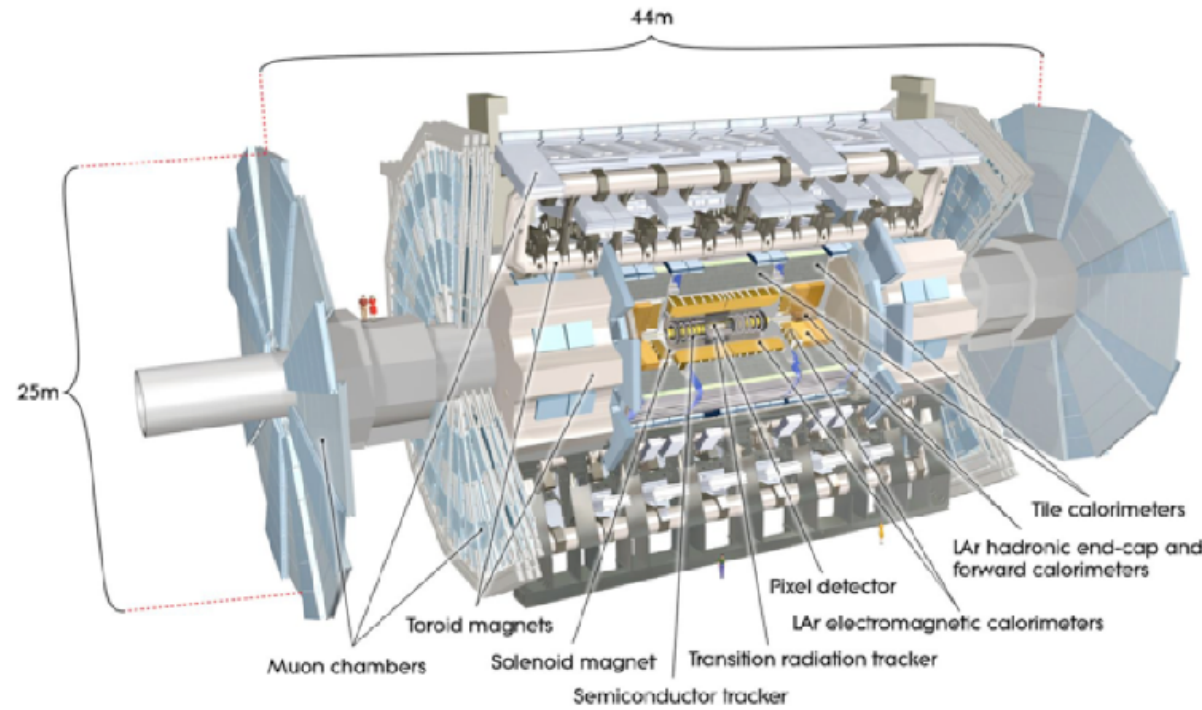


In the event display, right, we are looking down the beam pipe, which we call ***the transverse plane*** here

Question: What is the sum of the momentum in this plane **before** and **after** the particle collision

Question: What physics process are we observing in this event?

Neutrinos



- **Question:** How would this event display look if we had produced a ***W* boson** instead of a ***Z* boson** ?

Muon Spectrometer

Hadronic Calorimeter

Electromagnetic Calorimeter

Solenoid magnet

Tracking

Transition Radiation Tracker

Pixel/SCT detector

Proton

Neutron

Muon

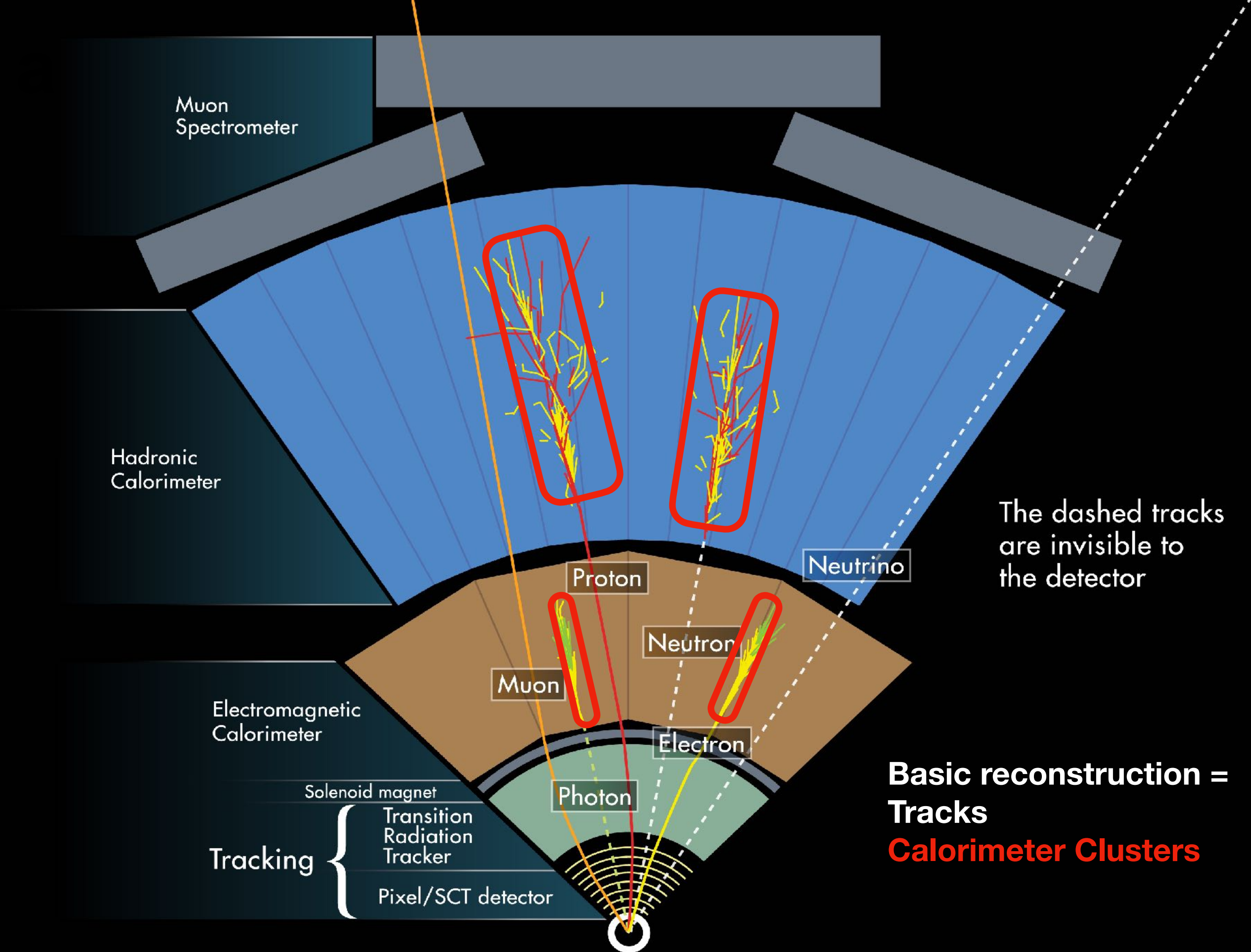
Electron

Photon

Neutrino

The dashed tracks are invisible to the detector

Basic reconstruction =
Tracks
Calorimeter Clusters



Muon
Spectrometer

Hadronic
Calorimeter

Electromagnetic
Calorimeter

Tracking

Solenoid magnet

Transition
Radiation
Tracker

Pixel/SCT detector

Proton

Neutron

Muon

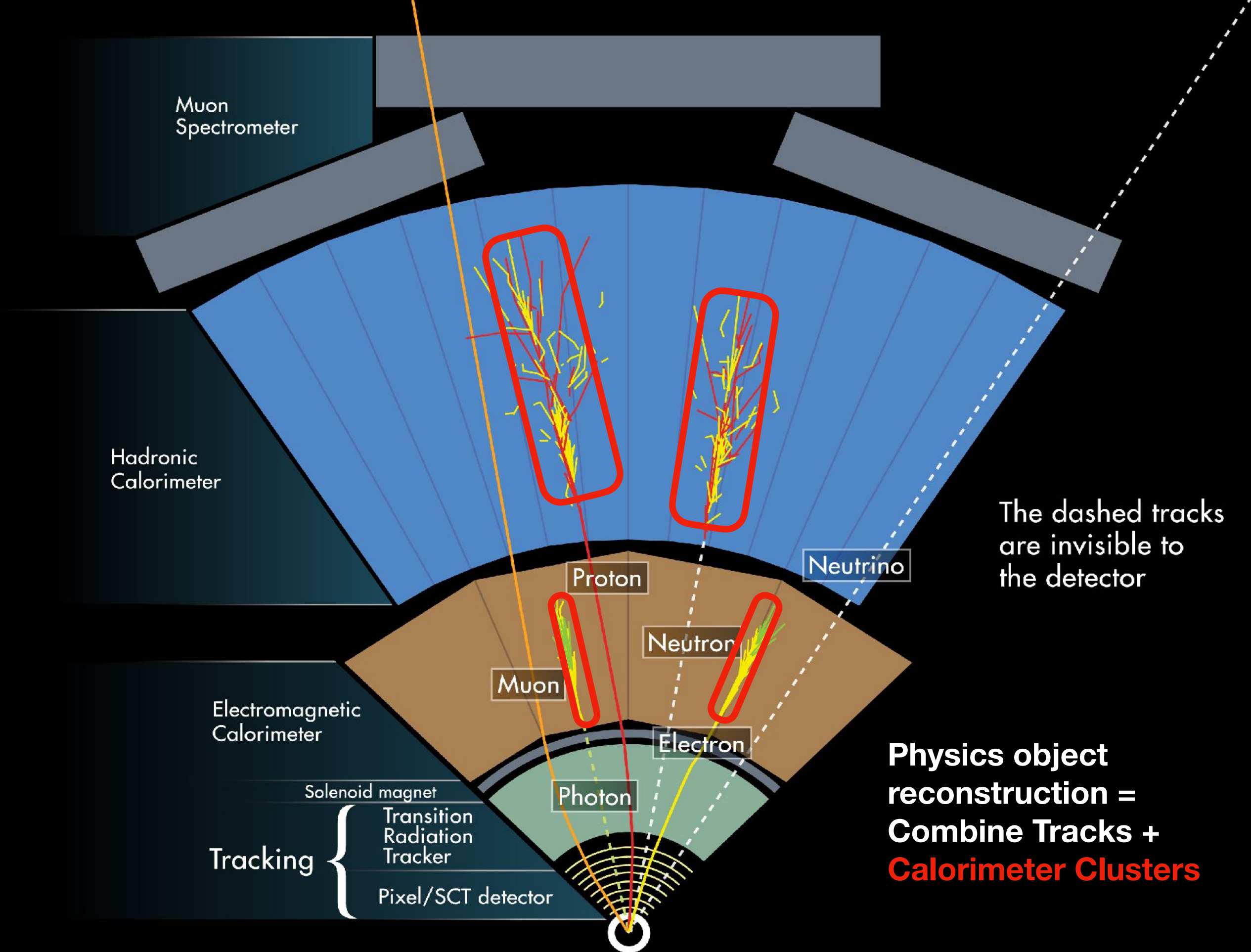
Electron

Photon

Neutrino

The dashed tracks
are invisible to
the detector

**Physics object
reconstruction =
Combine Tracks +
Calorimeter Clusters**



Here be dragons... and muons

Muon Spectrometer

Hadronic Calorimeter

Electromagnetic Calorimeter

Solenoid magnet

Tracking

Transition
Radiation
Tracker

Pixel/SCT detector

Proton

Neutron

Muon

Electron

Photon

Neutrino

The dashed tracks
are invisible to
the detector

**Muon reconstruction =
Track reconstruction
+ muon spectrometer hits**

