



SIDIS - Working Group

June 16, 2021

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for the Pavia group

Standard TMD formula

$$f_1^q(x, b_T; \mu) = \sum_i (C_{qi} \otimes f_1^i)(x, b_*; \mu_b) e^{S(b_*; \mu_b, \mu)} e^{g_K(b_T) \log \mu / \mu_0} f_{\text{NP}}^q(x, b_T)$$

Standard TMD formula

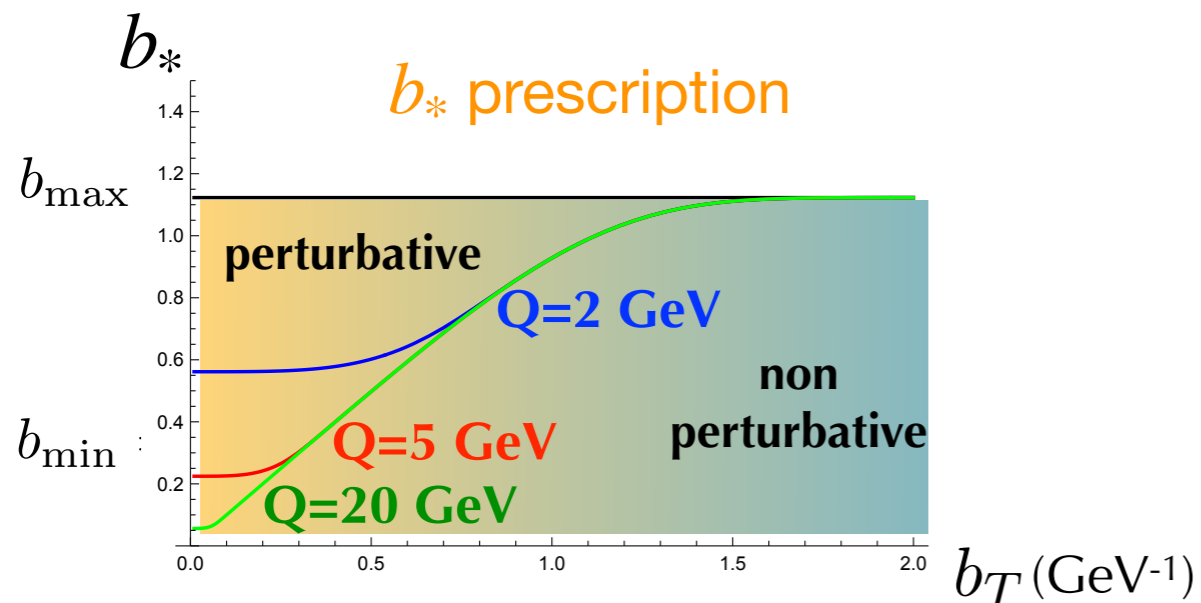
Perturbative part :

OPE matching coefficients C_{qi}

collinear PDFs at scale $\mu_b = \frac{2e^{-\gamma_E}}{b_*(b_T)}$

Sudakov factor \rightarrow perturbative evolution

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$$\left. \begin{aligned} b_{\max} &= 2e^{-\gamma_E} \\ b_{\min} &= \frac{2e^{-\gamma_E}}{Q} \end{aligned} \right\} 1 \leq \mu_b \leq Q$$

For $\mu = \mu_0 = 1$ GeV $\text{TMD}(x, b_T; \mu_0) = \text{PDF}(x; \mu_0) f_{\text{NP}}(x, b_T)$

other prescriptions possible..

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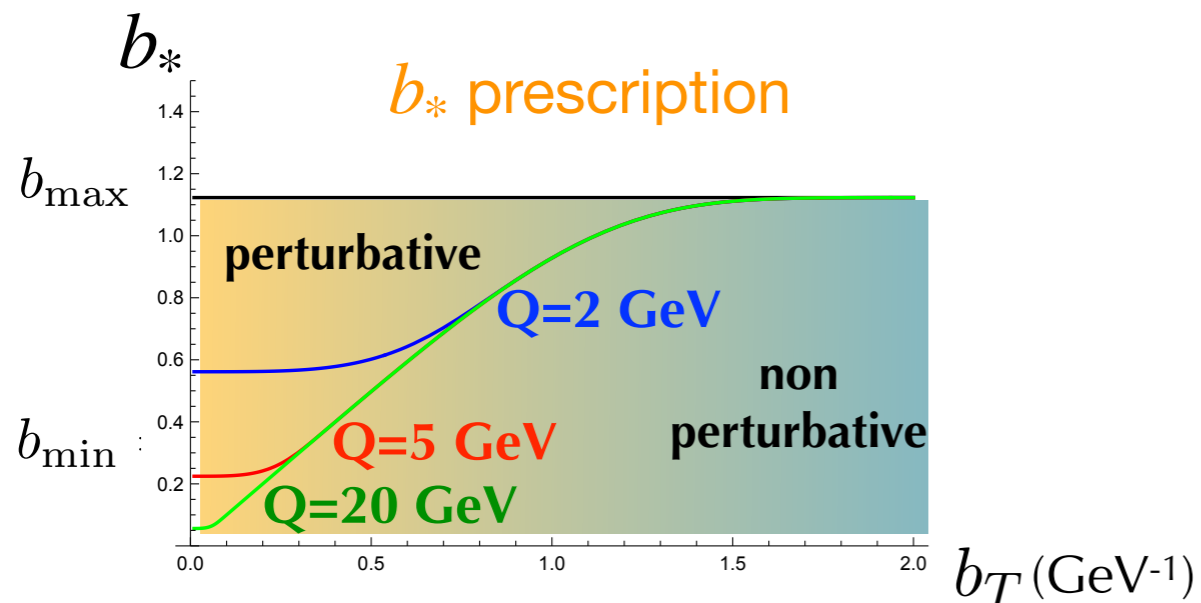
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Focus on NonPerturbative part : evolution

intrinsic wave function



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NonPerturbative functional form

PV17 fit

A. Bacchetta *et al.*, JHEP06 (2017) 081, arXiv:1703.10157

intrinsic wave function

$$f_{\text{NP}}(x, \mathbf{k}_T^2) = \frac{1}{\pi} \frac{1 + \lambda \mathbf{k}_T^2}{g_1 + \lambda g_1^2} e^{-\mathbf{k}_T^2/g_1}$$

$$g_1(x) = N_1 \frac{(1-x)^\alpha x^\sigma}{(1-\hat{x})^\alpha \hat{x}^\sigma}$$

$\hat{x} = 0.1$

evolution

$$g_K(b_T) = -g_2 \frac{b_T^2}{4}$$

similar for TMD FF

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$g_2 \rightarrow$ nonperturbative evolution

$N_1 = g_1(\hat{x}) \rightarrow$ mid- x width of TMD

$\sigma \rightarrow$ low- x width of TMD

$\alpha \rightarrow$ high- x width of TMD

$\lambda \rightarrow$ weight of second Gaussian

} not much constrained by fit

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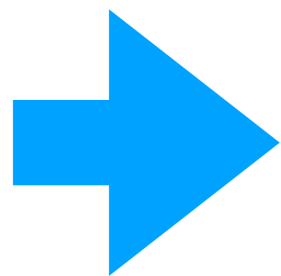
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PV17 nonperturbative parameters

PV17 fit

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Data from:

SIDIS
Hermes , Compass

Drell-Yan
FermiLab

Z-boson production
Tevatron: CDF , D0

8059 pts , 11 parameters , $\chi^2/\text{dof} = 1.55 \pm 0.05$

g_2	N_1 [GeV ²]	σ	α	λ [GeV ²]
0.13 ± 0.01	0.28 ± 0.06	0.17 ± 0.02	2.95 ± 0.05	0.86 ± 0.78

$$\bar{g}_2 \pm \Delta g_2$$

$$\bar{N}_1 \pm \Delta N_1$$

$$\bar{\sigma} \pm \Delta\sigma$$

$$\bar{\alpha} \pm \Delta\alpha$$

$$\bar{\lambda} \pm \Delta\lambda$$



all standard deviations at 68% confidence level

Sensitivity coefficients of PV17 parameters

Yellow Report R. Abdul Khalek *et al.*, arXiv:2103.05419

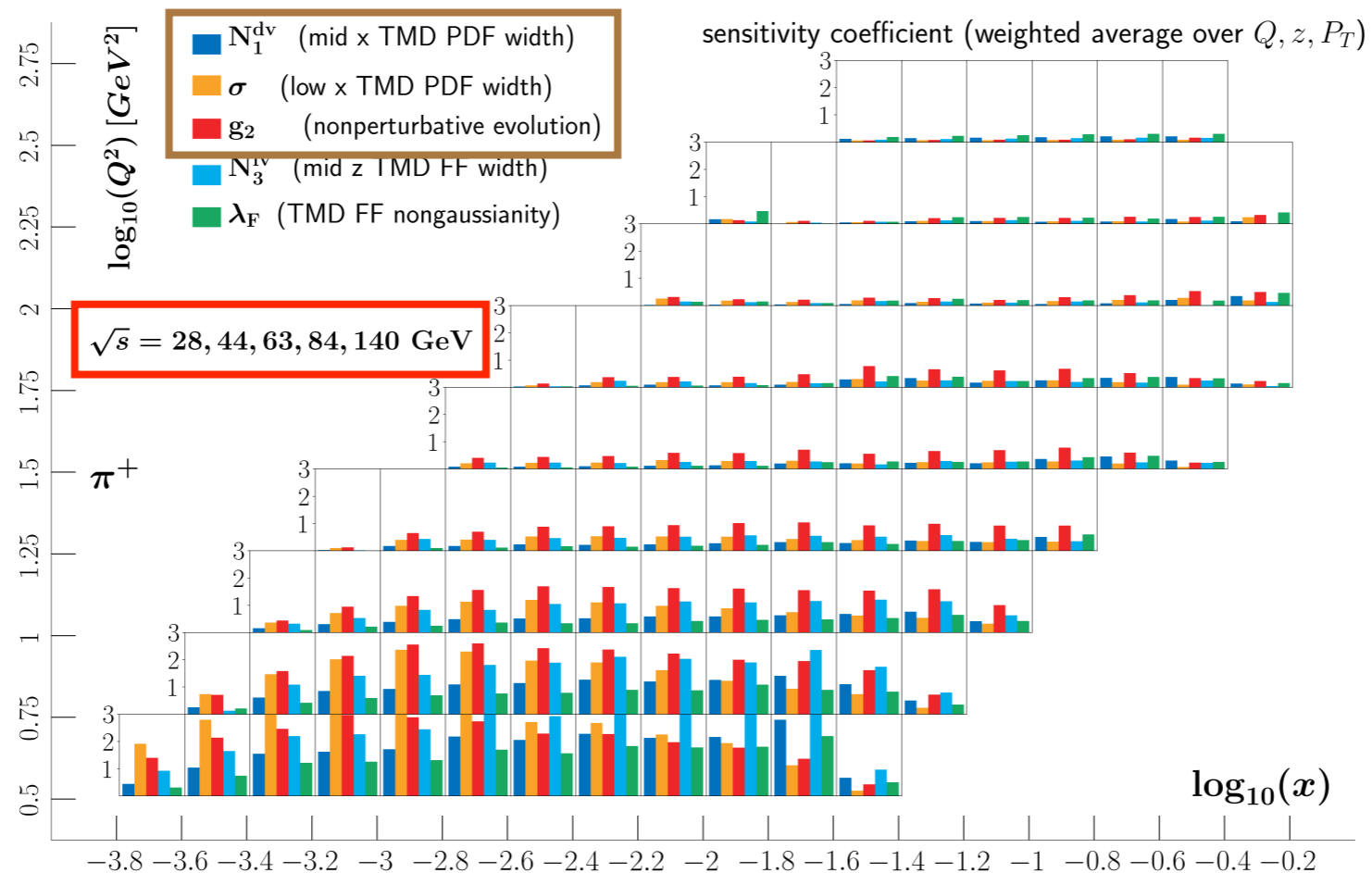


Figure 8.30: Expected sensitivities to various TMD PDF and FF parameters, as well as the TMD evolution as shown for the various collision energy options and for detected final-state positive pions. The impact has been averaged over final state hadron transverse momentum and fractional energy for better visibility.

$$S(\langle x \rangle_{\text{bin}}, \langle Q^2 \rangle_{\text{bin}}) \text{ with weighted average over } z, P_T, \sqrt{s}$$

(weights proportional to $1/\text{error of pseudodata}$)

Sensitivity coefficient and standard deviation

sensitivity coefficient S of object f w.r.t. observable O

$$S(O, f) = \frac{\langle Of \rangle - \langle O \rangle \langle f \rangle}{\delta O \Delta f}$$

experimental error for O

standard deviation for f

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standard deviation for f

1. from PV17, we know a parameter A with error ΔA
2. if we perform a new measurement that produces on A an error equal to its initial standard deviation, $\delta A = \Delta A$, we expect the error on A to scale as $1/\sqrt{2}$. We postulate that this corresponds to $S(A) = 1$

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3. in fact, if A can be ideally considered as parameter and observable, then

$$S(A, A) = \frac{\langle A A \rangle - \langle A \rangle \langle A \rangle}{\delta A \Delta A} = \frac{(\Delta A)^2}{\Delta A \Delta A} = 1$$

4. the error on A scales as $1/\sqrt{2} = 1/\sqrt{1 + (S = 1)}$. If the new measurement is more precise, then $S > 1$ and the error is further reduced; viceversa, for $S < 1$

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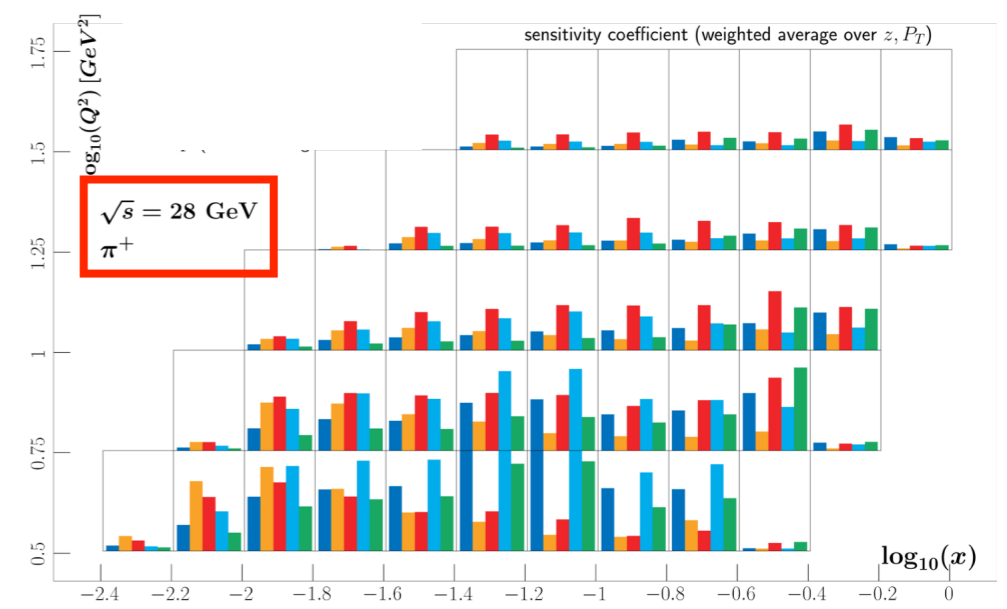
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5. for n measurements, the error on A should scale as $1/\sqrt{1 + S_1 + \dots + S_n}$

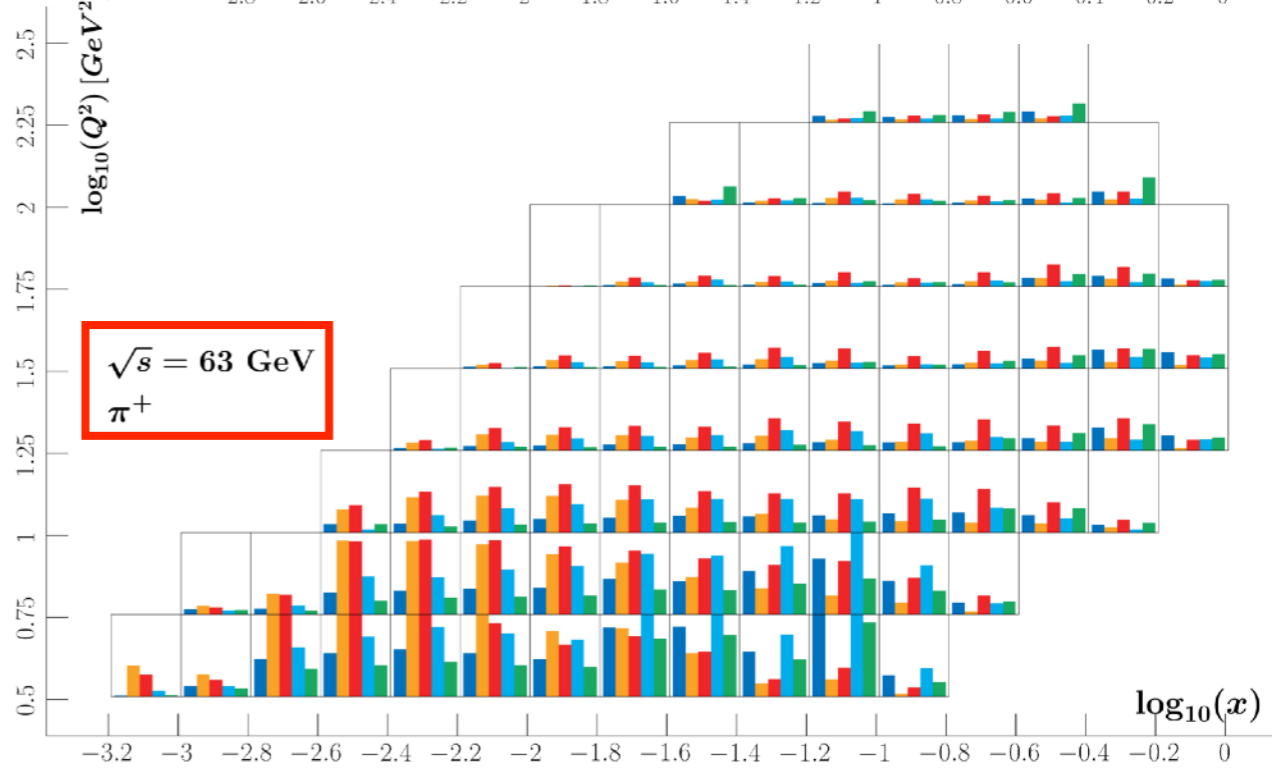
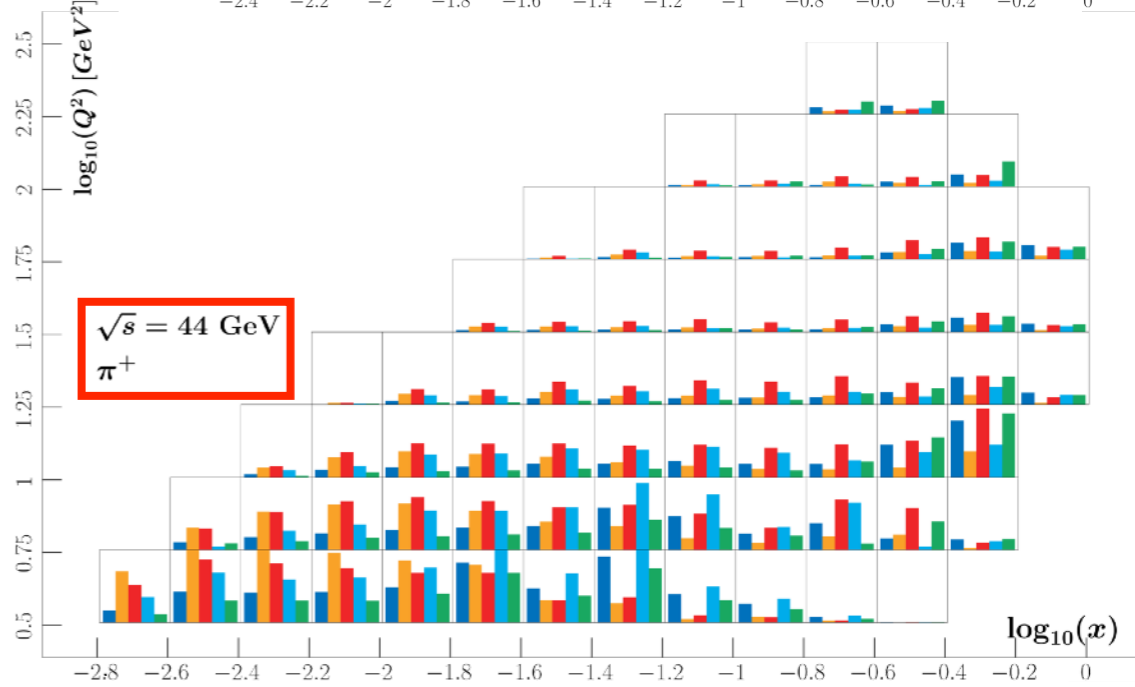
$S(\langle x \rangle_{\text{bin}}, \langle Q^2 \rangle_{\text{bin}})$ with
weighted average over z, P_T

- N_1^{dv} (mid x TMD PDF width)
- σ (low x TMD PDF width)
- g_2 (nonperturbative evolution)
- N_3^{fv} (mid z TMD FF width)
- λ_F (TMD FF nongaussianity)



$S(\langle x \rangle_{\text{bin}}, \langle Q^2 \rangle_{\text{bin}})$ with
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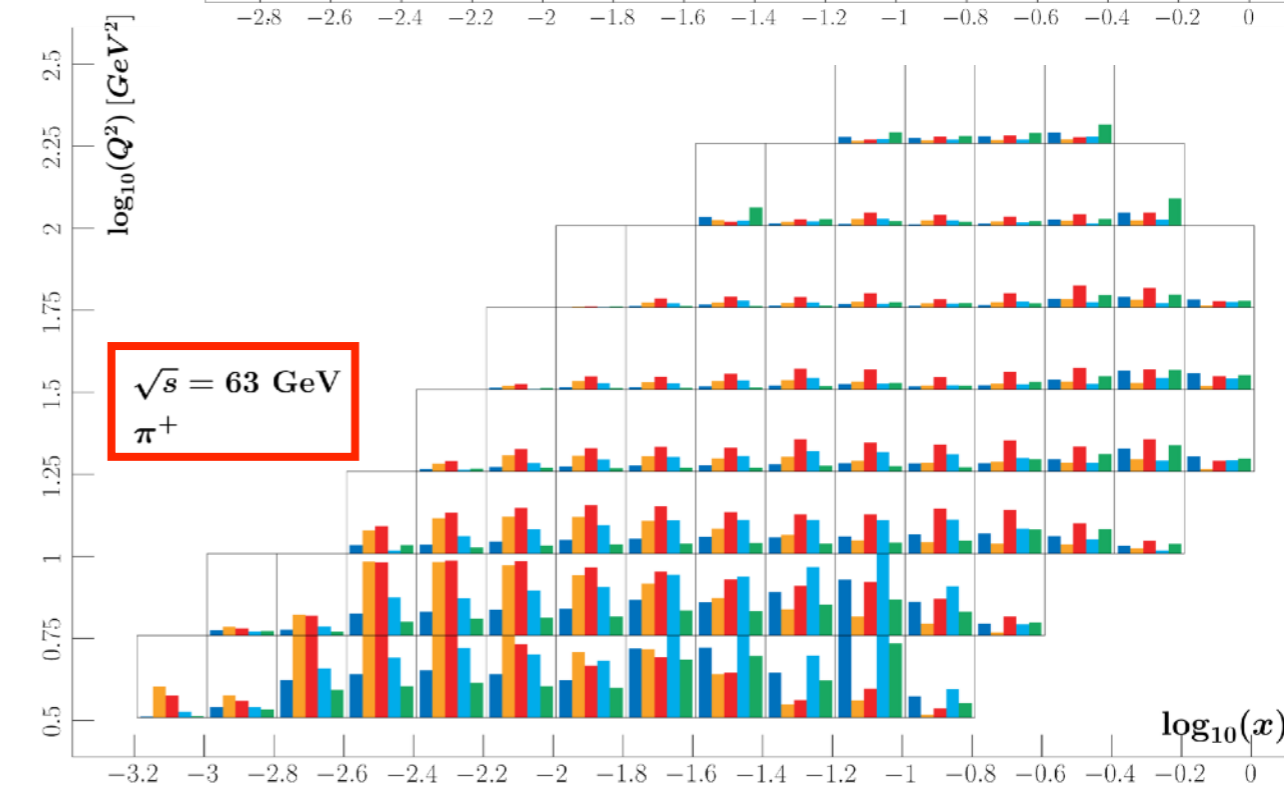
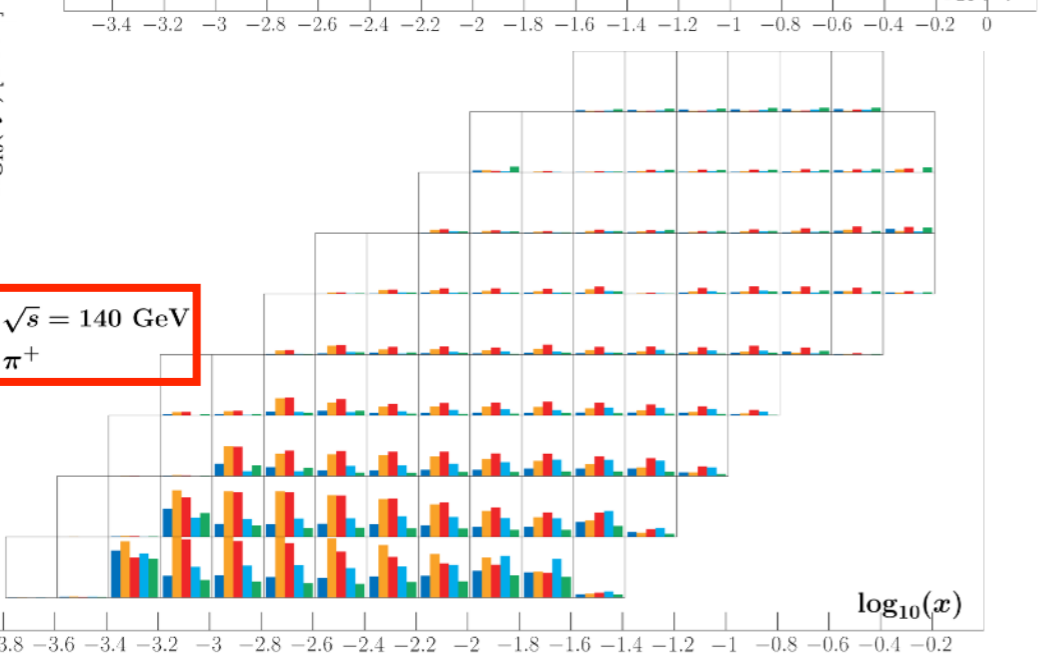
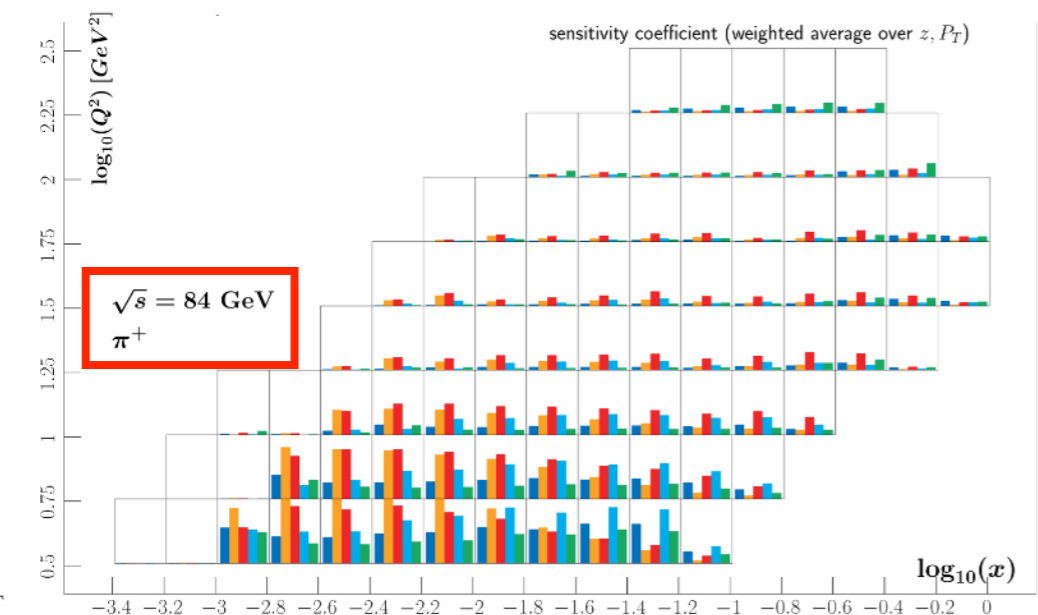
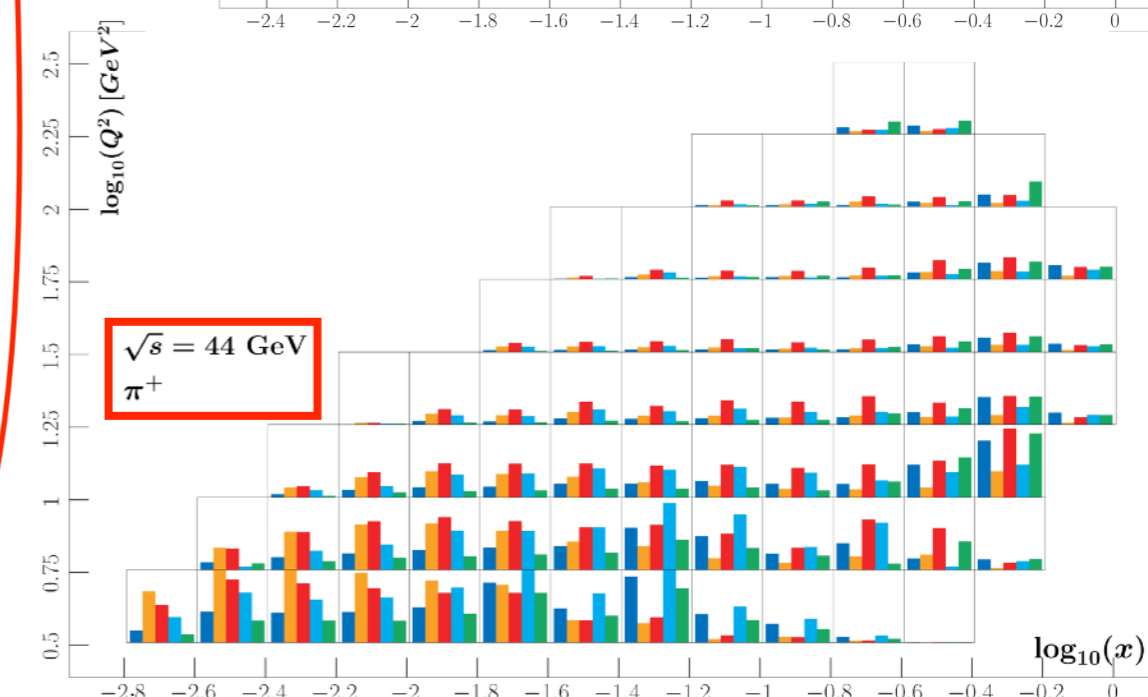
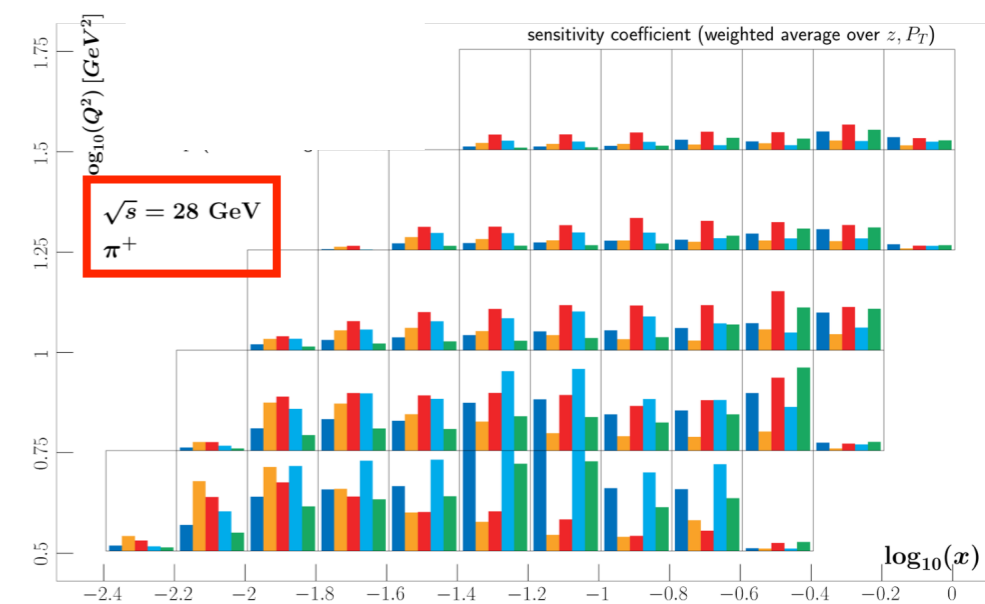
- N_1^{dv} (mid x TMD PDF width)
- σ (low x TMD PDF width)
- g_2 (nonperturbative evolution)
- N_3^{IV} (mid z TMD FF width)
- λ_F (TMD FF nongaussianity)



$S(\langle x \rangle_{\text{bin}}, \langle Q^2 \rangle_{\text{bin}})$ with
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higher $\sqrt{s} \rightarrow$ push to lower x , higher Q^2 ,
but larger errors \rightarrow lower coefficients S



NonPerturbative evolution g_2

summing over all (x, Q^2) bins

PV17 fit: $\Delta g_2 = 0.01 \longrightarrow$	run at $\sqrt{s} = 28$ GeV, π^+	$\longrightarrow 0.00155$	$R(g_2) = 6.45$
	run at $\sqrt{s} = 44$ GeV, π^+	$\longrightarrow 0.00120$	$R(g_2) = 8.33$
	run at $\sqrt{s} = 63$ GeV, π^+	$\longrightarrow 0.00108$	$R(g_2) = 9.26$
	run at $\sqrt{s} = 84$ GeV, π^+	$\longrightarrow 0.00105$	$R(g_2) = 9.52$
	run at $\sqrt{s} = 140$ GeV, π^+	$\longrightarrow 0.00096$	$R(g_2) = 10.36$

consistent trend:
larger $\sqrt{s} \longrightarrow$ larger covered (x, Q^2)
 \longrightarrow more stringent constraint

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Caveat

- optimistic upper limit in reduction coeff. R (no correlation between measurements in different bins)
- exercise biased by rigidity of PV17 functional form; future fits with (many) thousands points could demand more flexible forms

Max error reduction R for PV17 parameters

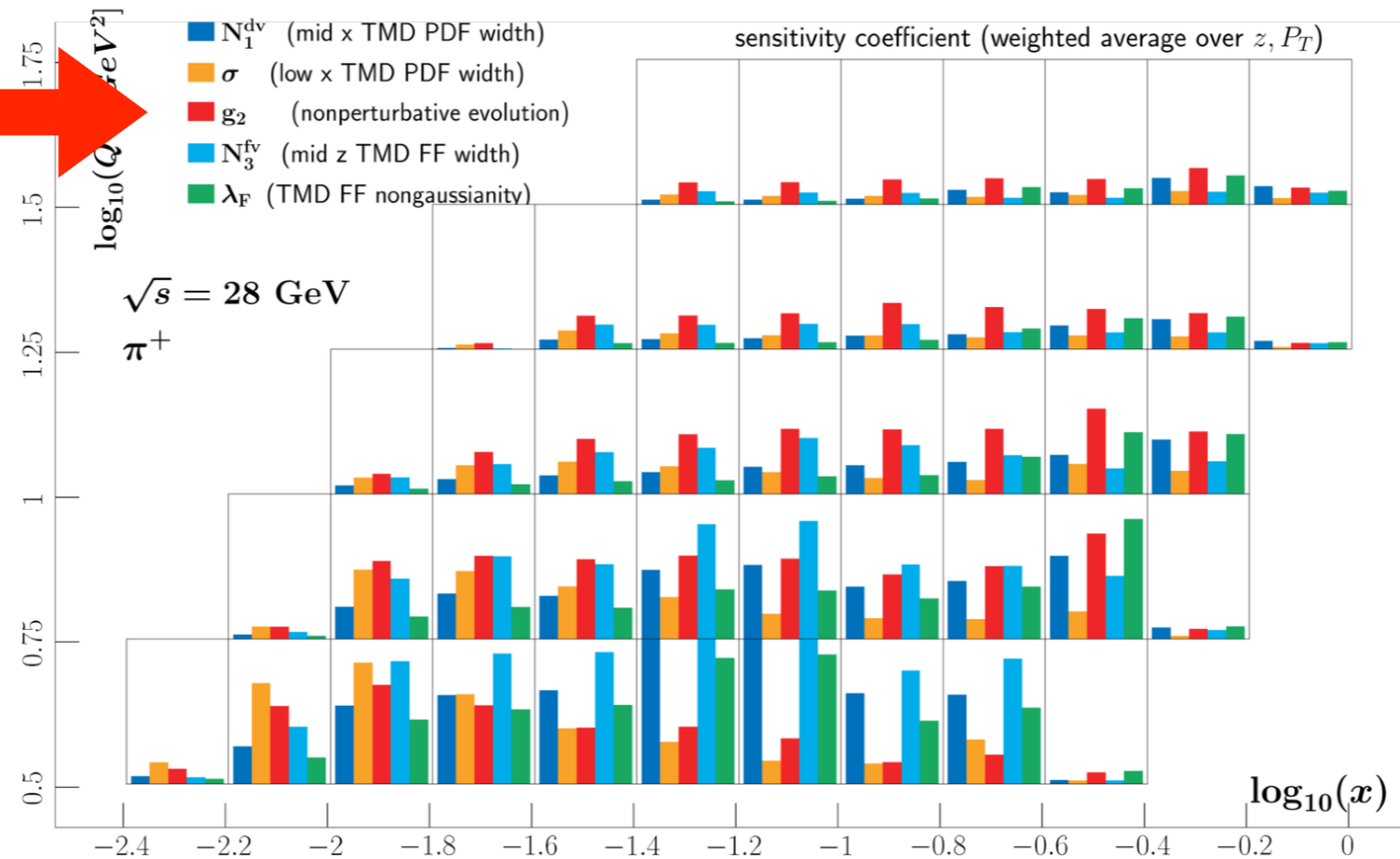
Energies Parameters	$\sqrt{s} = 28$	$\sqrt{s} = 44$	$\sqrt{s} = 63$	$\sqrt{s} = 84$	$\sqrt{s} = 140$
g_2 nonperturbative evolution	6.45	8.33	9.26	9.52	10.36
N_1 mid- x TMD width	5.94	6.52	6.96	6.80	6.73
σ low- x TMD width	5.05	6.85	8.00	8.55	10.00

Most sensitive (x, Q^2) bins: g_2

PV17 fit: $\Delta g_2 = 0.01 \longrightarrow$ summing over all (x, Q^2) bins $\longrightarrow 0.00155$

$R(g_2) = 6.45$

$\sqrt{s} = 28 \text{ GeV}, \pi^+$



Most sensitive (x, Q^2) bins: g_2

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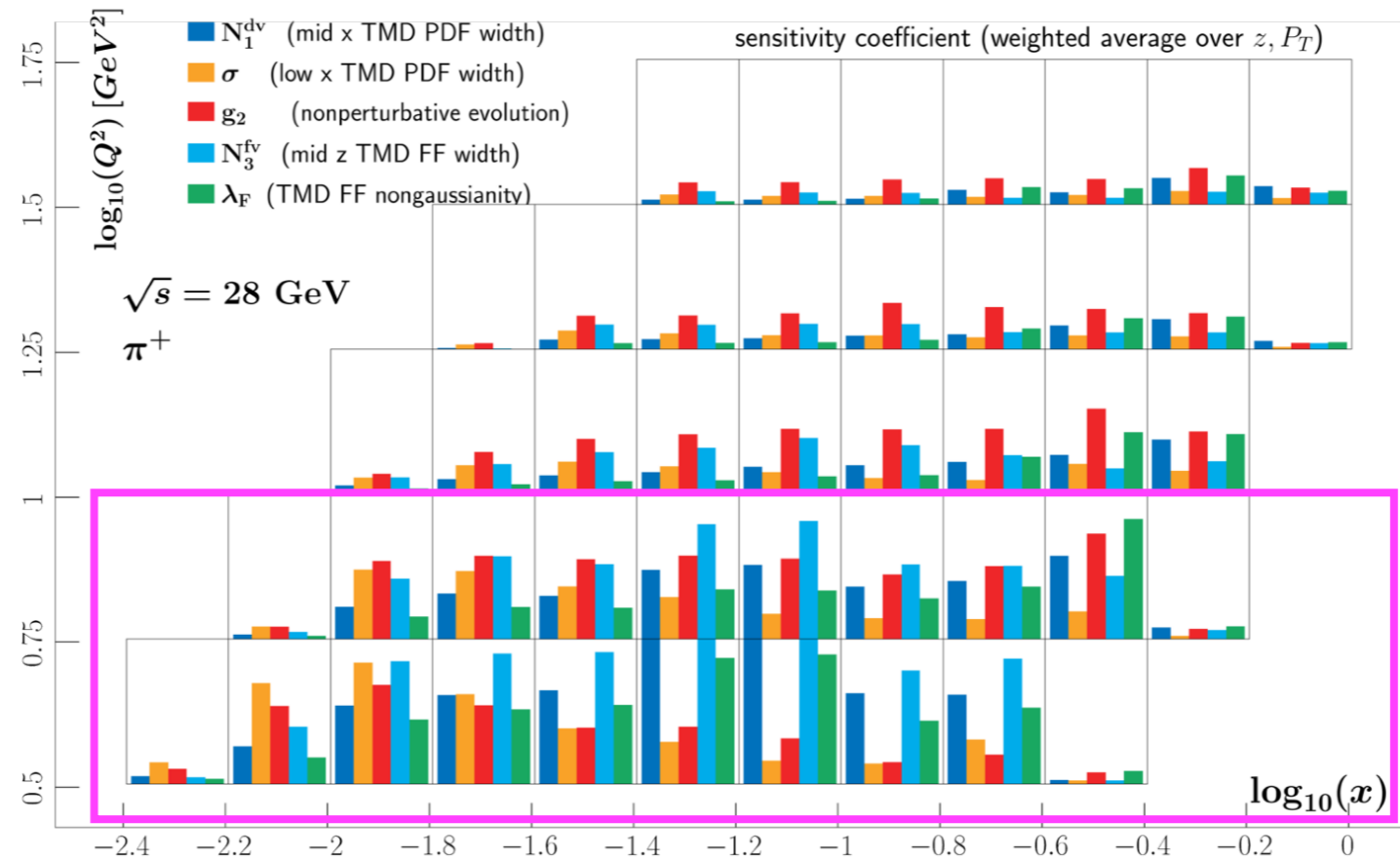
summing over all (x, Q^2) bins $\longrightarrow 0.00155$

$R(g_2) = 6.45$

summing over all $Q^2 \leq 10$ bins $\longrightarrow 0.00218$

71% R

$\sqrt{s} = 28$ GeV, π^+

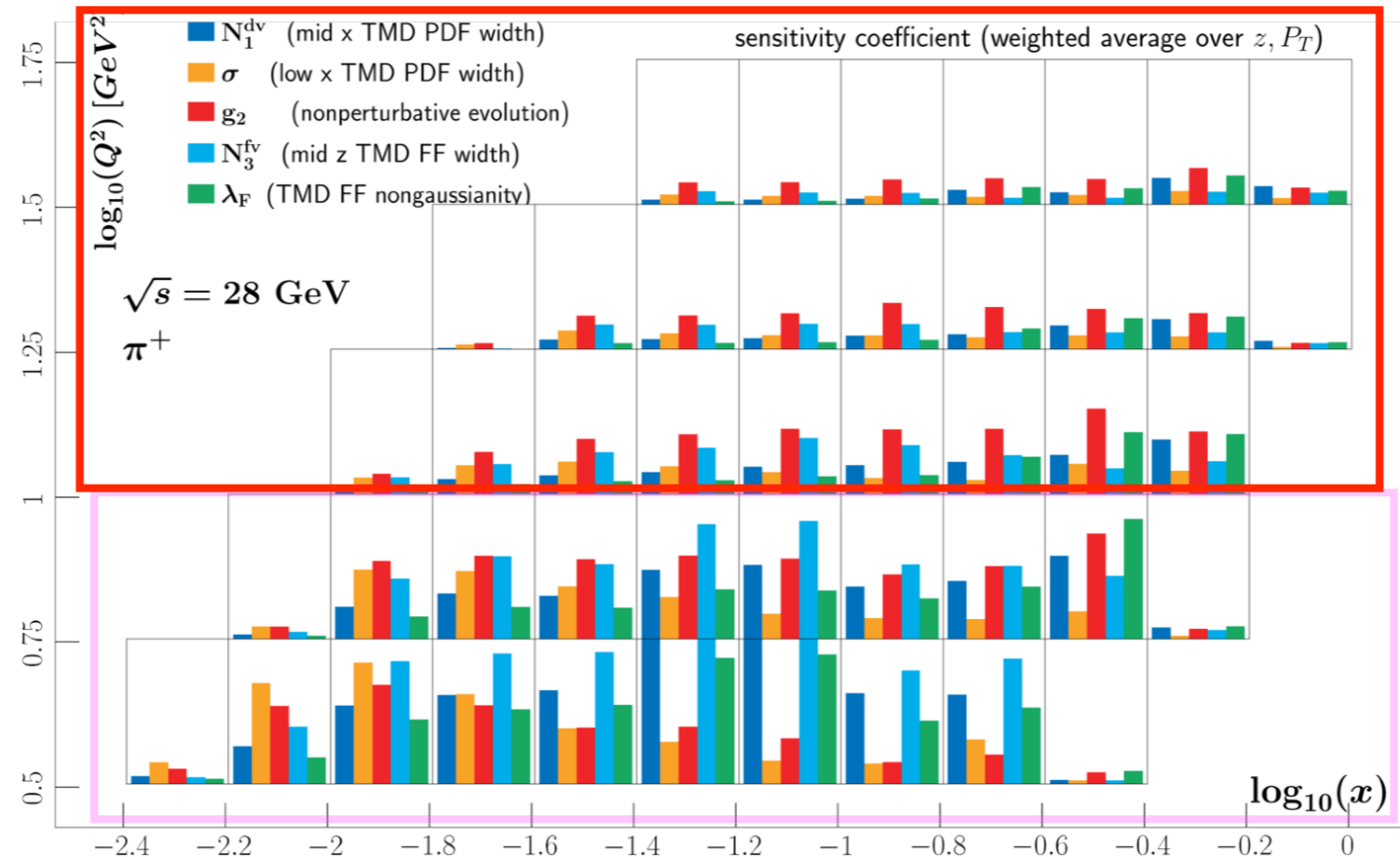


Most sensitive (x, Q^2) bins: g_2

PV17 fit: $\Delta g_2 = 0.01 \longrightarrow$

summing over all (x, Q^2) bins \longrightarrow 0.00155	$R(g_2) = 6.45$
summing over all $Q^2 \leq 10$ bins \longrightarrow 0.00218	71% R
summing over all $10 \leq Q^2 \leq 100$ bins \longrightarrow 0.00216	72% R

$\sqrt{s} = 28$ GeV, π^+

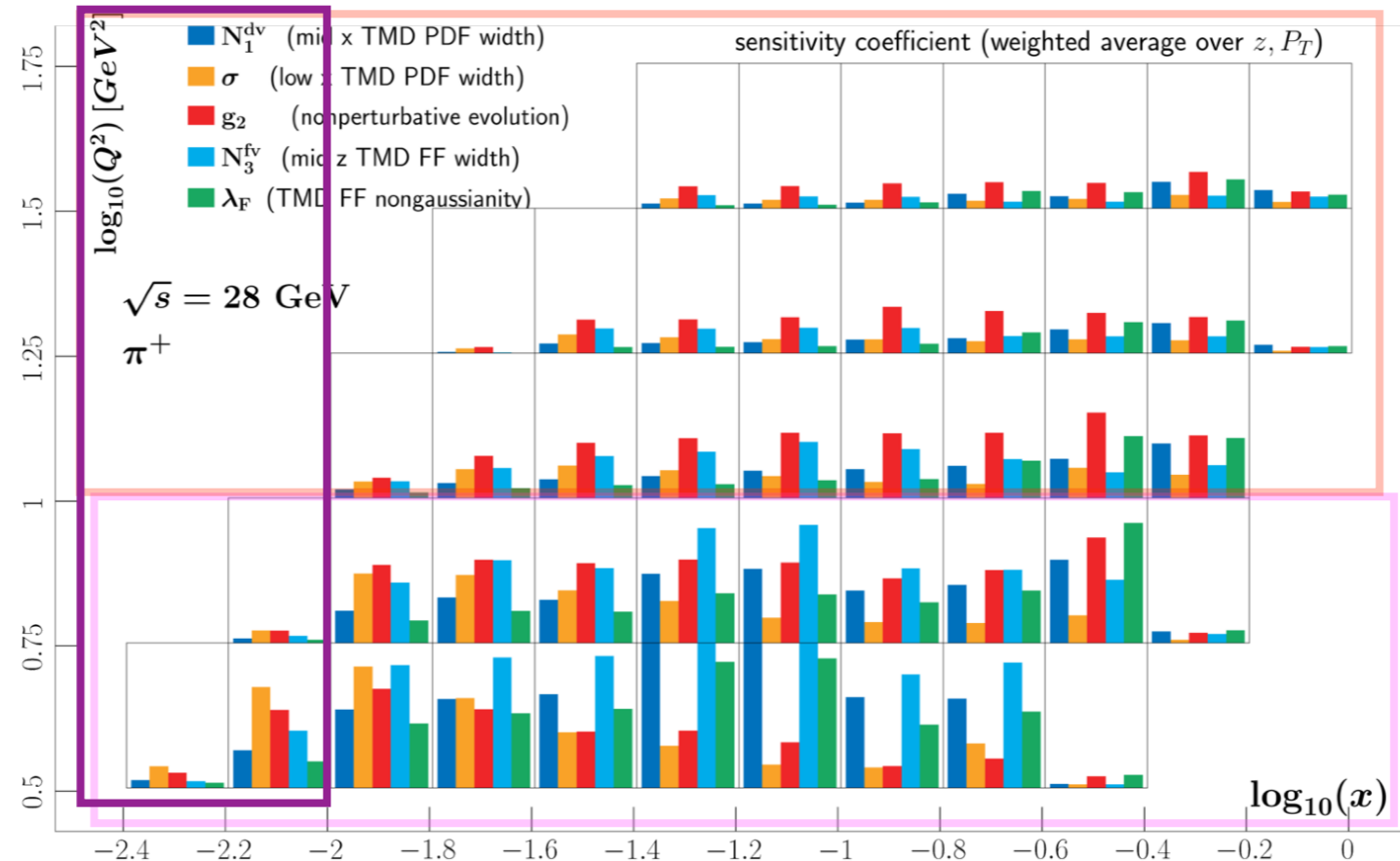


Most sensitive (x, Q^2) bins: g_2

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summing over all $10 \leq Q^2 \leq 100$ bins \longrightarrow 0.00216	72% R
summing over all $x \leq 0.01$ bins \longrightarrow 0.00306	51% R

$\sqrt{s} = 28$ GeV, π^+

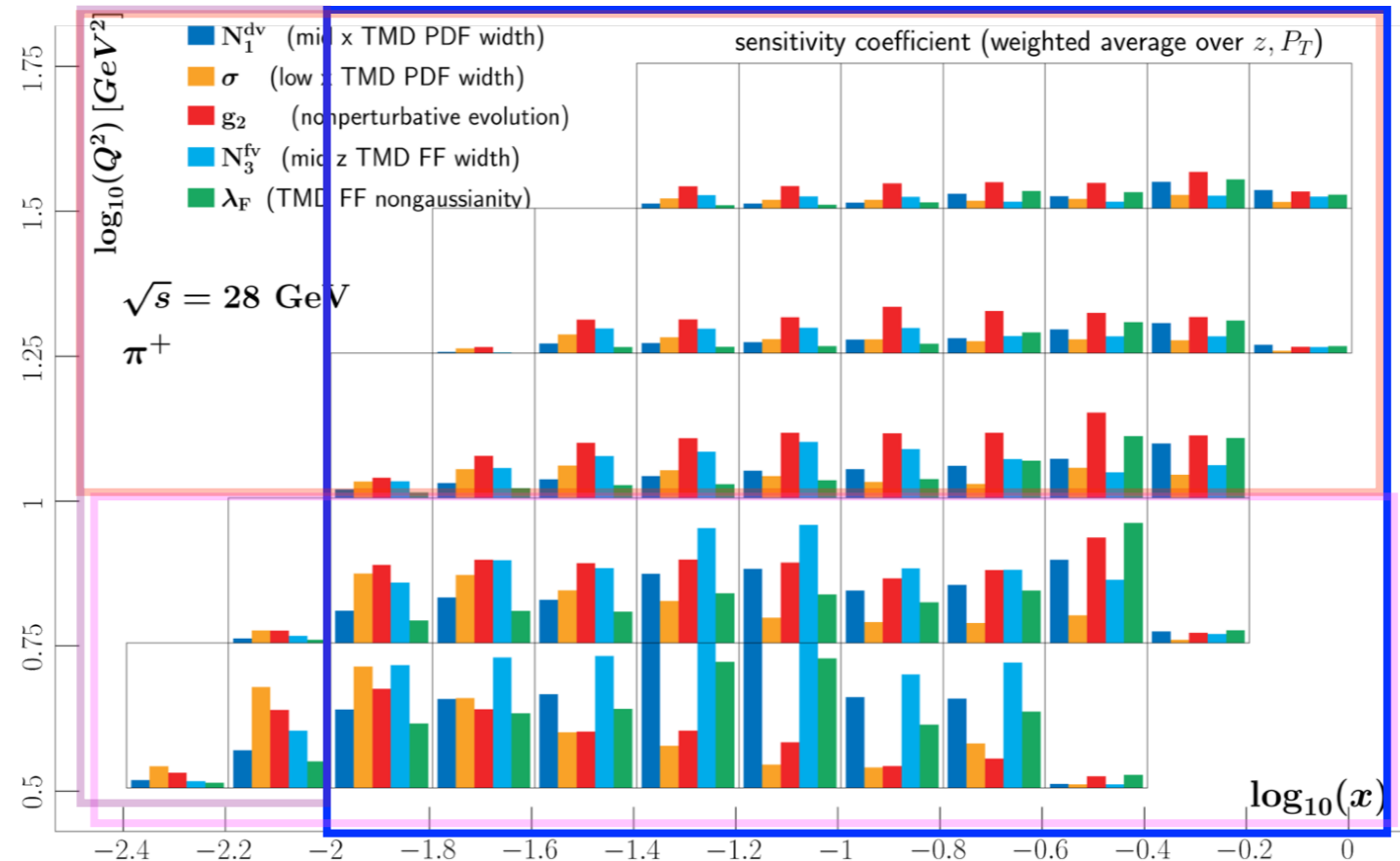


Most sensitive (x, Q^2) bins: g_2

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summing over all $10 \leq Q^2 \leq 100$ bins \longrightarrow	0.00216	72% R
summing over all $x \leq 0.01$ bins \longrightarrow	0.00306	51% R
summing over all $x \geq 0.01$ bins \longrightarrow	0.00177	88% R

$\sqrt{s} = 28$ GeV, π^+



Most sensitive (x, Q^2) bins: g_2

bins	All bins	$Q^2 \leq 10$	$10 \leq Q^2 \leq 100$	$x \leq 0.01$	$0.01 < x \leq 1$
Energies					
$\sqrt{s} = 28$	$R = 6.45$	71% R	72% R	51% R	88% R
$\sqrt{s} = 44$	$R = 8.33$	71% R	69% R	48% R	89% R
$\sqrt{s} = 63$	$R = 9.26$	72% R	67% R	59% R	82% R
$\sqrt{s} = 84$	$R = 9.52$	73% R	66% R	64% R	77% R
$\sqrt{s} = 140$	$R = 10.36$	75% R	63% R	80% R	61% R

Most sensitive (x, Q^2) bins: g_2

good options at $\sqrt{s} = 140, x \leq 0.01$ but also at $\sqrt{s} = 44, 0.01 < x \leq 1$

bins	All bins	$Q^2 \leq 10$	$10 \leq Q^2 \leq 100$	$x \leq 0.01$	$0.01 < x \leq 1$
Energies					
$\sqrt{s} = 28$	$R = 6.45$	71% R	72% R	51% R	88% R
$\sqrt{s} = 44$	$R = 8.33$	71% R	69% R	48% R	89% R
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$\sqrt{s} = 84$	$R = 9.52$	73% R	66% R	64% R	77% R
$\sqrt{s} = 140$	$R = 10.36$	75% R	63% R	80% R	61% R

Most sensitive (x, Q^2) bins: N_1

good options at medium-large x and/or at low Q^2

bins	All bins	$Q^2 \leq 10$	$10 \leq Q^2 \leq 100$	$x \leq 0.01$	$0.01 < x \leq 1$
Energies					
$\sqrt{s} = 28$	$R = 5.94$	83% R	58% R	43% R	92% R
$\sqrt{s} = 44$	$R = 6.52$	77% R	62% R	41% R	92% R
$\sqrt{s} = 63$	$R = 6.96$	81% R	56% R	50% R	88% R
$\sqrt{s} = 84$	$R = 6.80$	83% R	52% R	61% R	81% R
$\sqrt{s} = 140$	$R = 6.73$	83% R	53% R	79% R	63% R

Most sensitive (x, Q^2) bins: σ

consistent trend: good option at $\sqrt{s} = 140, x \leq 0.01$ but doesn't need large Q^2

bins	All bins	$Q^2 \leq 10$	$10 \leq Q^2 \leq 100$	$x \leq 0.01$	$0.01 < x \leq 1$
Energies					
$\sqrt{s} = 28$	$R = 5.05$	83% R	58% R	49% R	90% R
$\sqrt{s} = 44$	$R = 6.85$	82% R	57% R	62% R	79% R
$\sqrt{s} = 63$	$R = 8.00$	83% R	55% R	72% R	71% R
$\sqrt{s} = 84$	$R = 8.55$	83% R	55% R	76% R	66% R
$\sqrt{s} = 140$	$R = 10.00$	83% R	54% R	86% R	52% R