

# CP VIOLATION AND LEPTOGENESIS

NNNI5 and UD2

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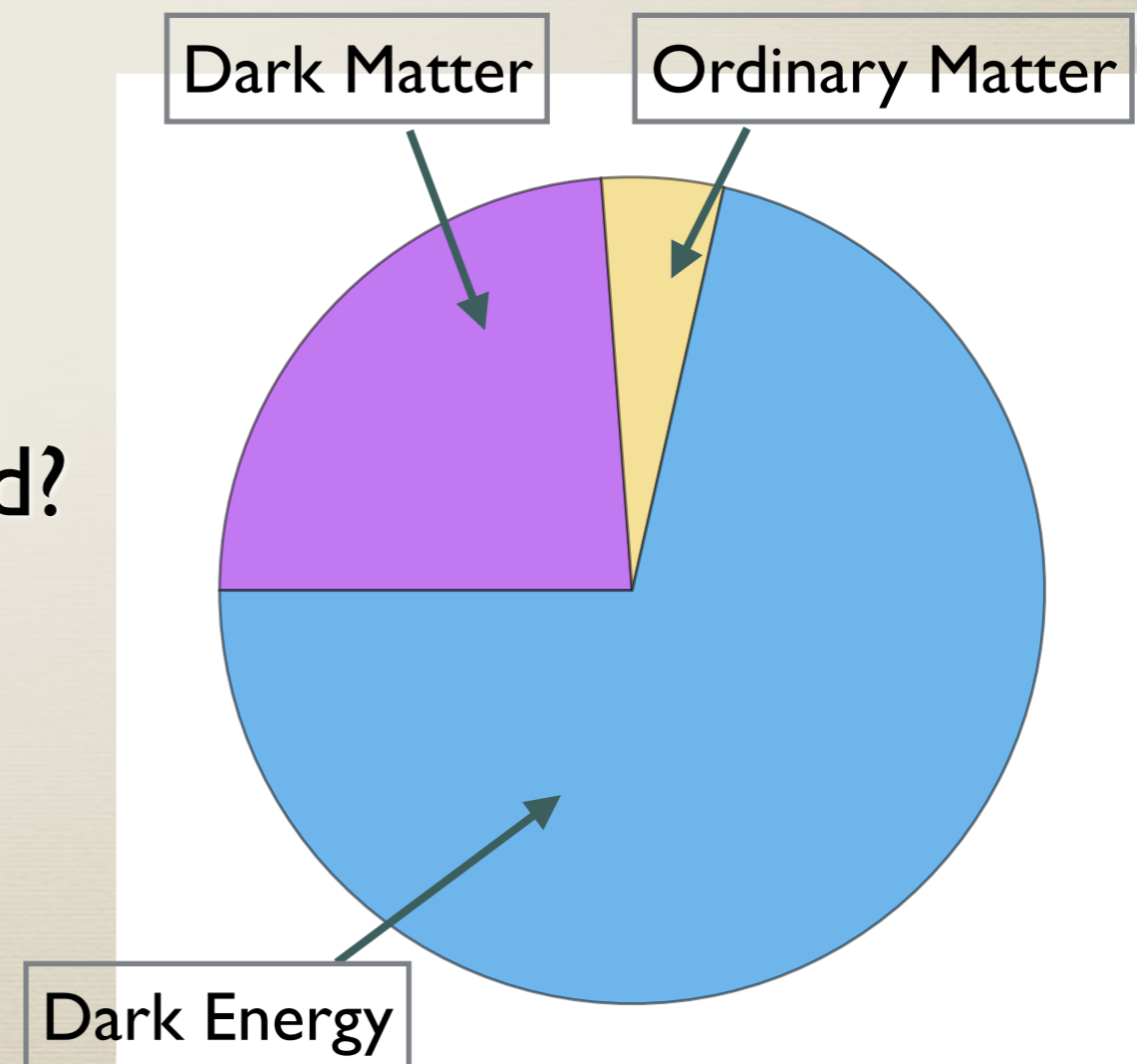


# Outline

- What is the baryon asymmetry and why is it important?
- How do we measure it?
- Review Sakharov Conditions
- Leptogenesis
- Connecting leptonic CP Violation to the baryon asymmetry

# Content of the Universe

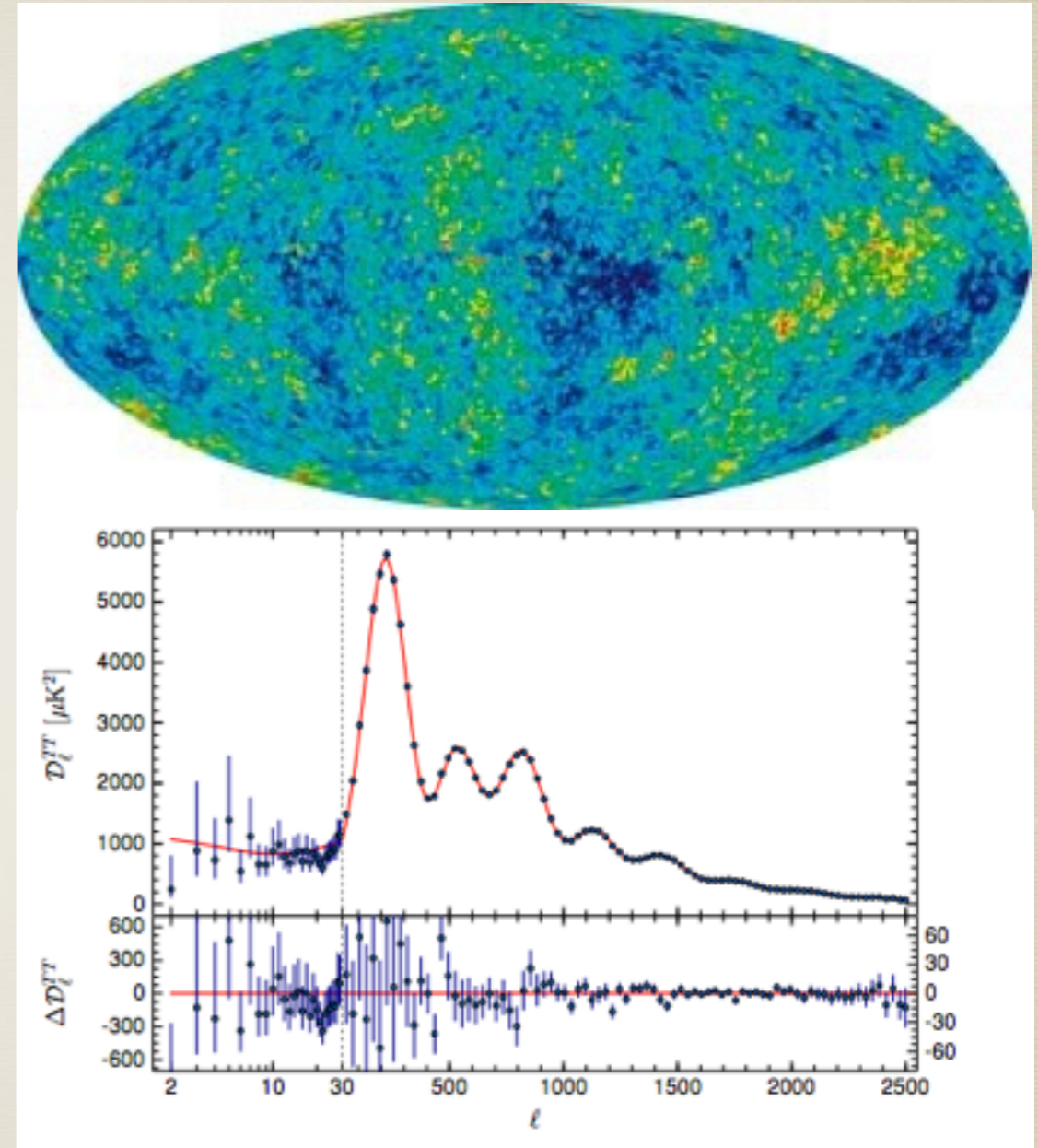
- Stars and Galaxies~0.5%, Neutrinos~0.1-1.5%.
- Mostly proton, neutrons and electron ~4.4%.
- Anti-Matter~0%
- Fortunately for us there exist a matter anti-matter asymmetry
- How was this asymmetry created?
- Initial Condition or dynamical process?



# Measuring the Baryon Asymmetry

## PLANCK Data

- Acoustic peaks in CMB anisotropy are due to oscillations in photon baryon fluid ( $T \sim 0.25\text{eV}$ )
- Position and shape of peaks in the power spectrum affected by amount of baryons.

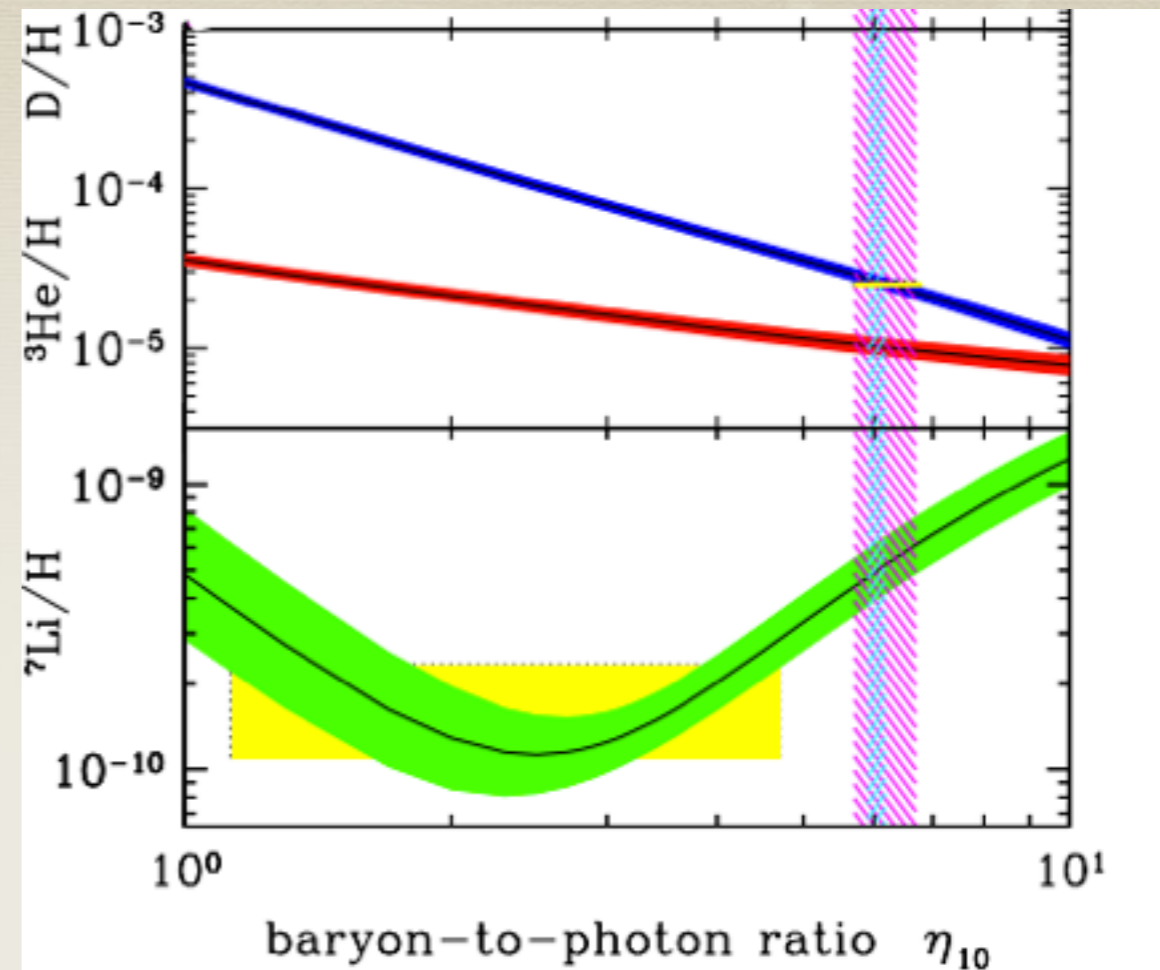


Planck 2015 (1502.01589)

$$\text{baryon/photon} \equiv \eta_{CMB} = (6.08 \pm 0.06) \times 10^{-10}$$

# Big Bang Nucleosynthesis

- BBN sensitive to era  $T < 1 \text{ MeV}$  (age  $> 1 \text{ s}$ ).  $p$  and  $n$  fuse to form light elements.
- Light elemental abundance dependent on baryon to photon
- D and H abundance determined using **high resolution of spectra** from quasars.



BBN PDG, 2012

$$\text{baryon/photon} \equiv \eta_{BBN} = (6.23 \pm 0.17) \times 10^{-10}$$

Remarkable agreement between  $\eta_{\text{CMB}}$  and  $\eta_{\text{BBN}}$

# Sakharov Conditions

In order to **dynamically generate** the baryon asymmetry in the early Universe, **Sakharov's Conditions** must be satisfied.

Sakharov, 1967

1. **Baryon (and/or Lepton) Number Violation.**

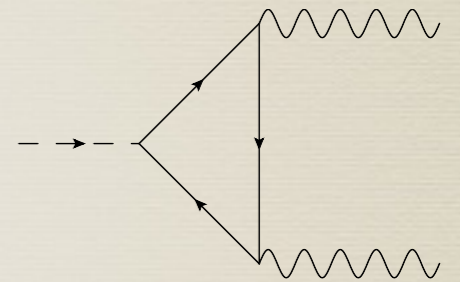
2. **C and CP Violation.**

3. **Out of equilibrium interactions.**

# Baryon Number Violation

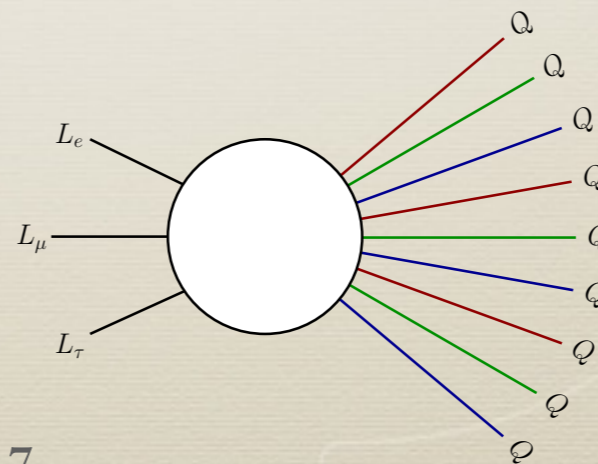
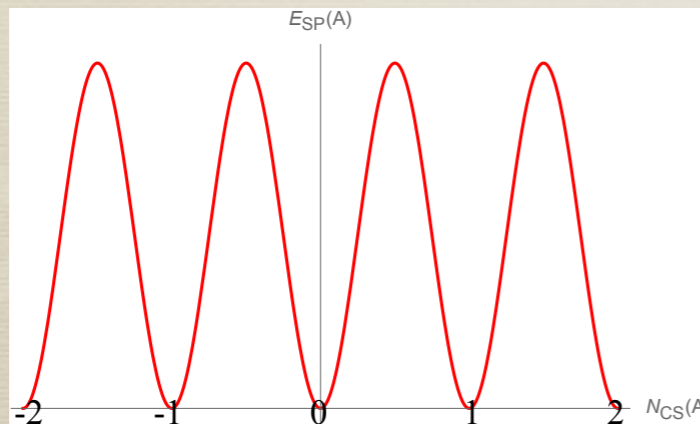
- If B number is conserved, then baryon asymmetry would reflect highly tuned initial condition.
- B, L are **accidental symmetries** of the SM. Conserved at **tree level** but violated at **loop-level**. t'Hooft, 1976

$$\partial_\nu J^\nu_B = \partial_\nu J^\nu_L = \frac{N_f}{32\pi^2} \left( \underbrace{g^2 W_{\nu\mu}^p W^{\tilde{p}\nu\mu}}_{SU(2) \text{ field strengths}} - g'^2 \underbrace{B_{\nu\mu} B^{\nu\mu}}_{U(1) \text{ field strength}} \right)$$



**B+L violated and B-L conserved**

- At high T, transition between different vacua can occur by jumping over barrier: B+L violation via sphaleron.



**At  $T > T_{EW} \sim 100 \text{ GeV}$**

# CP-Violation

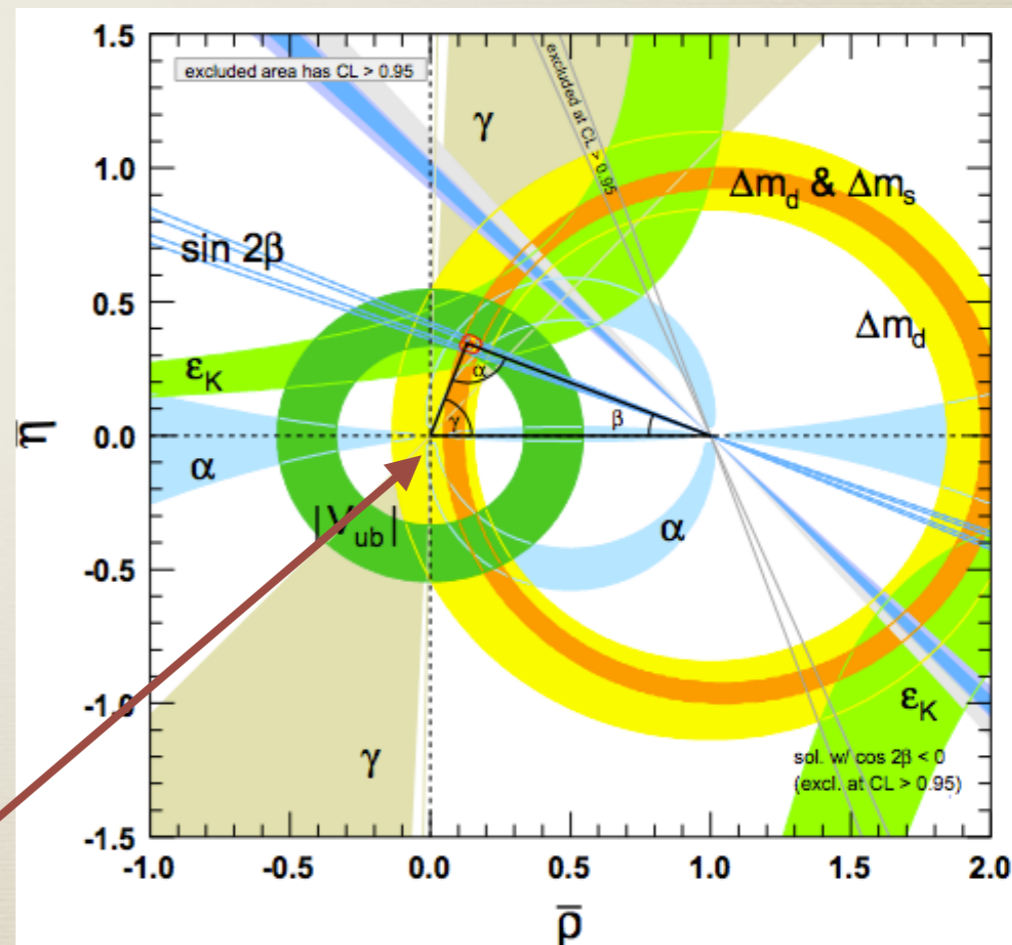
- CPV necessary otherwise same rate for particle and anti-particle  $\rightarrow$  no excess.
- **CP is violated** in the **quark sector** and emerges in kaon and b-systems. There is 1 CPV phase.

$$\frac{n_B - n_{\bar{B}}}{n_\gamma} \propto \frac{J \times \Delta m_{tcb}^2 \times \Delta m_{dsb}^2}{T_{EW}^{12}} \approx 10^{-17}$$

7 orders of magnitude too small.

- Need **new sources of CPV**. Gavela et al Nucl. Phys B430  
In this talk, we focus on **leptonic CPV**.

area of triangle = Jarlskog invariant/2



CKM PDG, 2012



# Out of Equilibrium

- In equilibrium:  $\Gamma_{process} \equiv \Gamma_{inverse}$
- Even if **B** and **CP** violating processes occur their effect is **cancelled** by their inverse processes.
- Thanks to the expansion of the Universe, processes can come out of equilibrium when  $\Gamma \sim H$ 
  - interaction rate
  - expansion rate
- In **leptogenesis** (with type I see-saw) heavy **sterile neutrino** decays out of equilibrium.

# Leptogenesis

- Leptogenesis can take place in the context of a see-saw model. As the Universe expands,  $N$  decays **out of equilibrium**.  $N$  decays generate a **lepton asymmetry** which is converted to a **baryon asymmetry** via **sphaleron** processes.

Fukugita, Yanagida (PLB 174), see others Covi, Roulet, Vissani, Buchmuller, Plumacher

- **L - violation**
- **C and CP-Violation.**
- **Out of equilibrium:** Expansion of Universe.

# See-Saw Mechanism

- Introduce a RH neutrino (SM singlet) and couple it to the Higgs.



$$\mathcal{L} = -Y_\nu \overline{N} L H - \frac{1}{2} \overline{N^c} M_N N$$

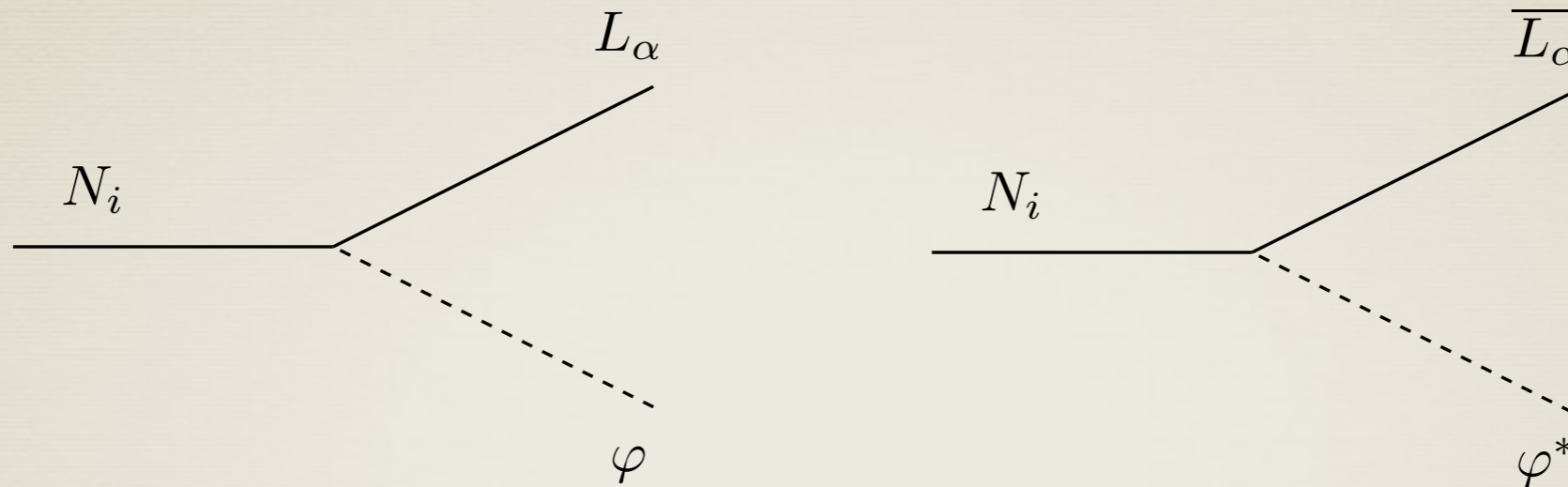
CPV from irremovable phases

Lepton Number is violated

$$\begin{pmatrix} 0 & m_D \\ m_D^T & M_N \end{pmatrix} \xrightarrow{Y_{\nu\nu}} m_\nu = \frac{Y_\nu^2 v_H^2}{M_N} \sim \frac{1\text{GeV}^2}{10^{10}\text{GeV}} \sim 0.1\text{eV}$$

# Basic Mechanism

- Heavy, singlet Majorana neutrinos decay out of equilibrium



- Occurs at  $T \sim M_{N_1}$
- Inverse decay rates **faster** (**slower**) than expansion of Universe H, **strong** (**weak**) washout.

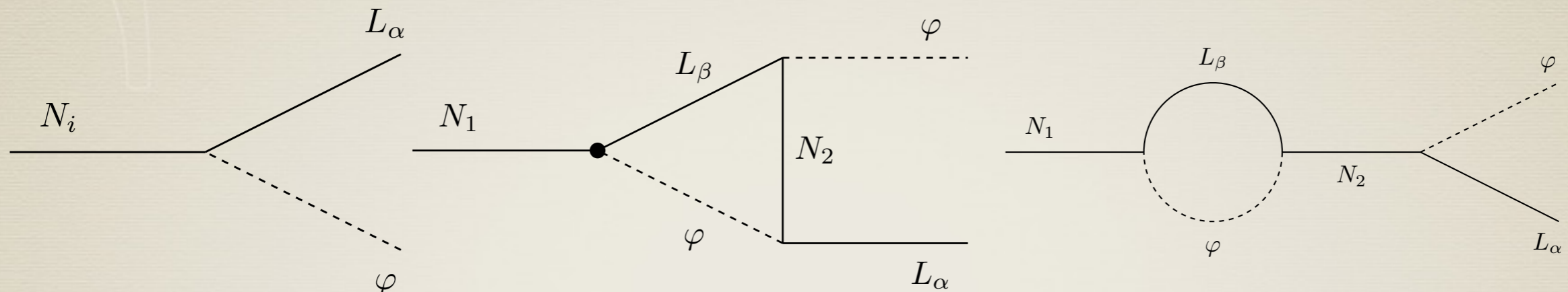
lepton  
asymmetry

sphaleron

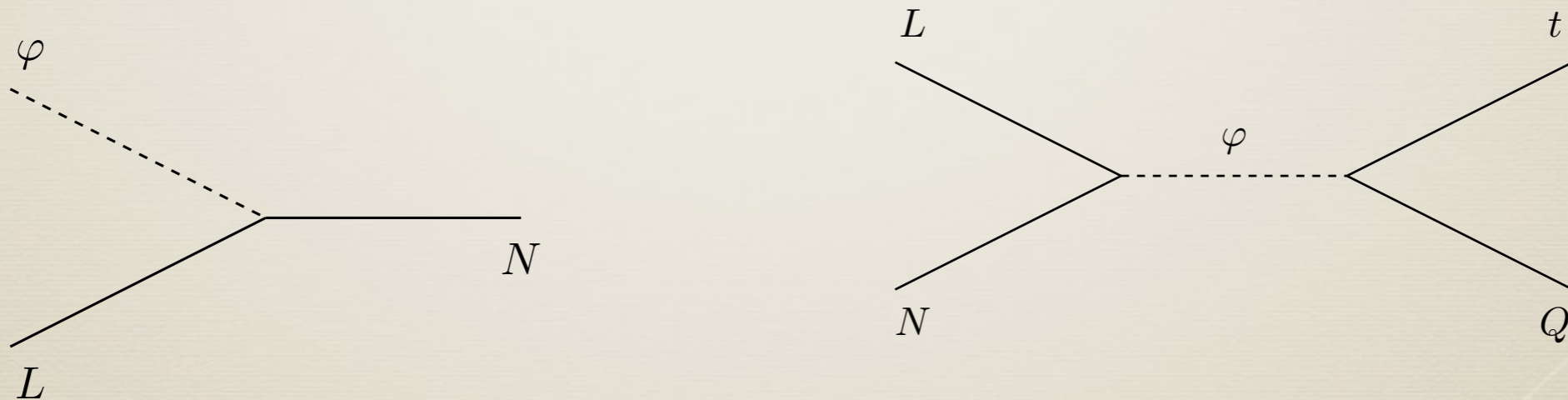
baryon  
asymmetry

- Decay asymmetry ( $\epsilon$ ) arises from interference between tree and loop diagrams.

$$\epsilon = \frac{\Gamma(N_1 \longrightarrow \phi L) - \Gamma(N_1 \longrightarrow \phi^\dagger \bar{L})}{\Gamma(N_1 \longrightarrow \phi L) + \Gamma(N_1 \longrightarrow \phi^\dagger \bar{L})}$$



- Washout ( $k$ ) can occur due to diagrams such as



$$Y_B = \text{const} \times C_{SP} \times k \times \epsilon$$

sphaleron conversion factor

13

washout

decay asymmetry

# CPV and Leptogenesis

- Assume Type-I see-saw with 3 hierarchical N

$$RR^T = I$$

$$\epsilon_l = \frac{\sum_l [\Gamma(N_1 \rightarrow HL_l) - \Gamma(N_1 \rightarrow \overline{HL}_l)]}{\sum_l [\Gamma(N_1 \rightarrow HL_l) + \Gamma(N_1 \rightarrow \overline{HL}_l)]}$$

high energy  
untestable  
parameters

$$Y = \frac{1}{v_H} \sqrt{M} R \sqrt{m} U^\dagger$$

low energy  
parameters

$$\epsilon_l = -\frac{3M_1}{16\pi v^2} \frac{\text{Im} \left( \sum_{l\beta\rho} m_\beta^{\frac{1}{2}} m_\rho^{\frac{3}{2}} U_{l\beta}^* U_{l\rho} R_{1\beta} R_{1\rho} \right)}{\sum_\beta m_\beta |R_{1\beta}|^2}$$

1 flavour ( $T > 10^{12}$  GeV)

$$\epsilon_1 = \frac{3M_1}{16\pi v^2} \frac{\text{Im}(\sum_\rho m_\rho^2 R_{1\rho}^2)}{\sum_\beta m_\beta |R_{1\beta}|^2}$$

Independent of  $U_{\text{PMNS}}$

2 flavour ( $10^{10} < T < 10^{12}$  GeV)

$$\epsilon_\tau \quad \epsilon_2$$

$$U_{l\beta}^* U_{l\rho} R_{1\beta} R_{1\rho}$$

Flavour effects allow a link between low energy leptonic observables and leptogenesis.

CPV high scale

$Y_\nu$

CPV low scale

$\delta, \alpha_{21}, \alpha_{31}$

CP

CP

CPV

CPV

CPV

CP

Establish that CPV phases  $U$  could be responsible for leptogenesis

Pascoli, Petcov, Riotto, PRD and NPB 2007.

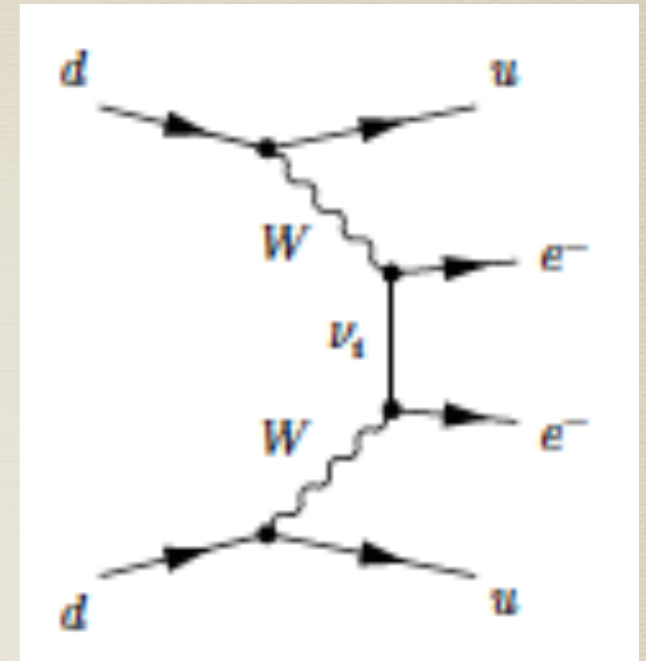
Can we test the ingredients that could establish a link between leptonic observables (low scale) and the baryon asymmetry (high scale)?



Yes we can!  
Are neutrinos Majorana?  
Is there CPV in the lepton sector?

# Lepton Number Violation

- Implementation of see-saw mechanism, the **light neutrinos** generated are **Majorana**
- $\beta 0\nu\nu$  can determine the nature of the neutrino via rare decay:



$$(A, Z) \longrightarrow (A, Z + 2) + 2e^{-} (\Delta L = 2)$$

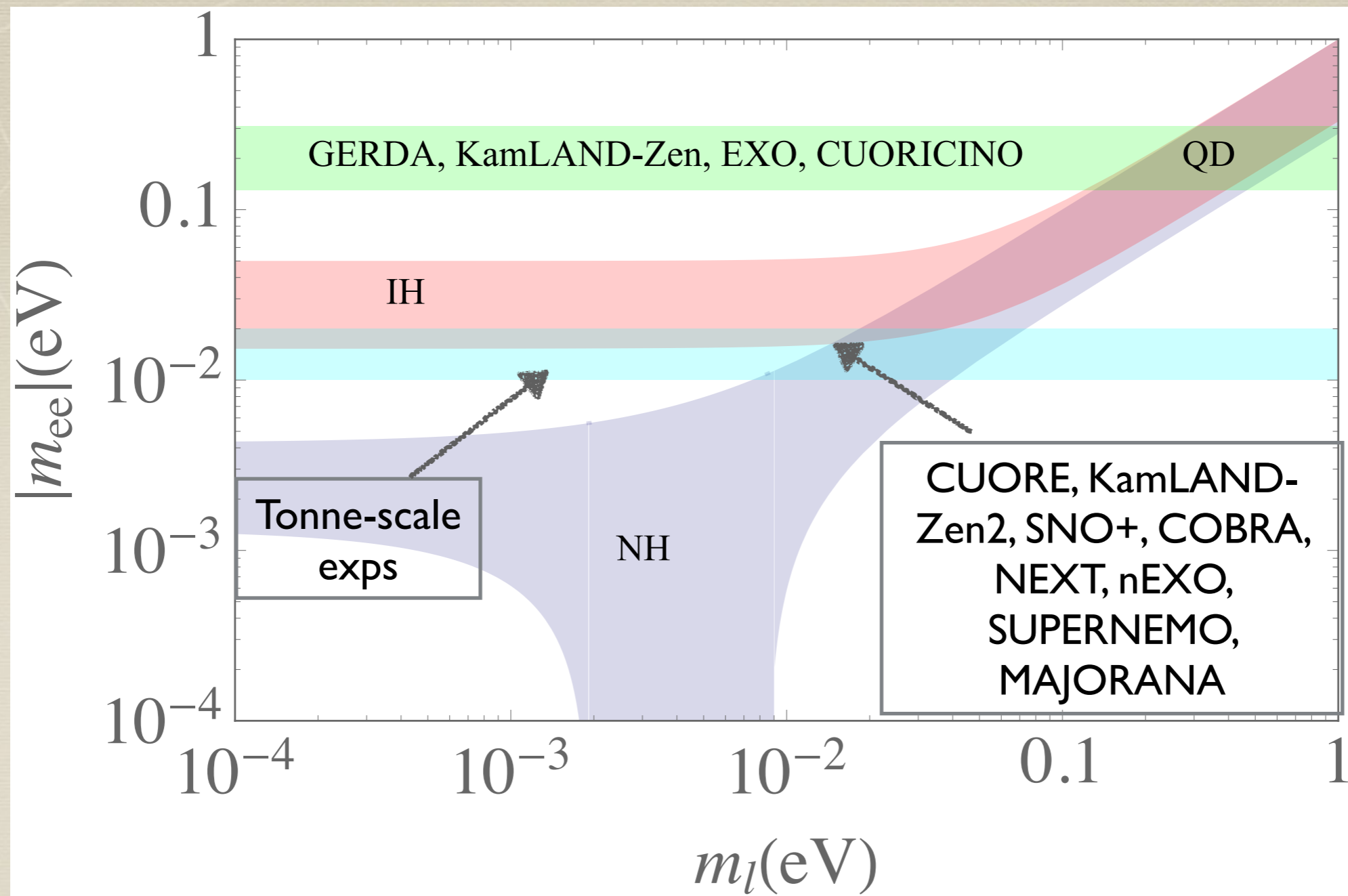
- The half life of rare decay is proportional to effective Majorana mass:

$$|m_{ee}| = |m_1 \cos^2 \theta_{12} \cos^2 \theta_{13} + m_2 \sin^2 \theta_{12} \cos^2 \theta_{13} e^{i\alpha_{21}} + m_3 \sin^2 \theta_{13} e^{i(\alpha_{31} - 2\delta)}|$$

partly known

known

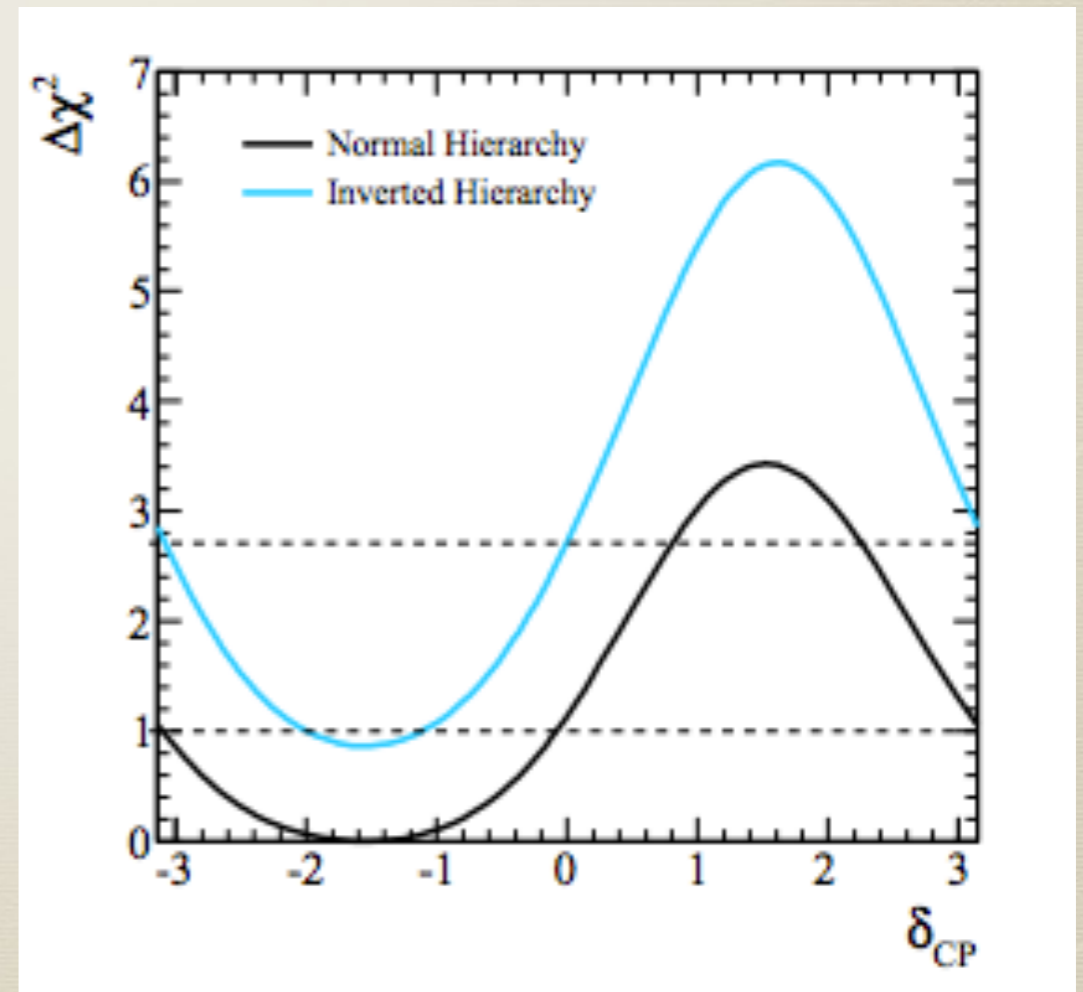
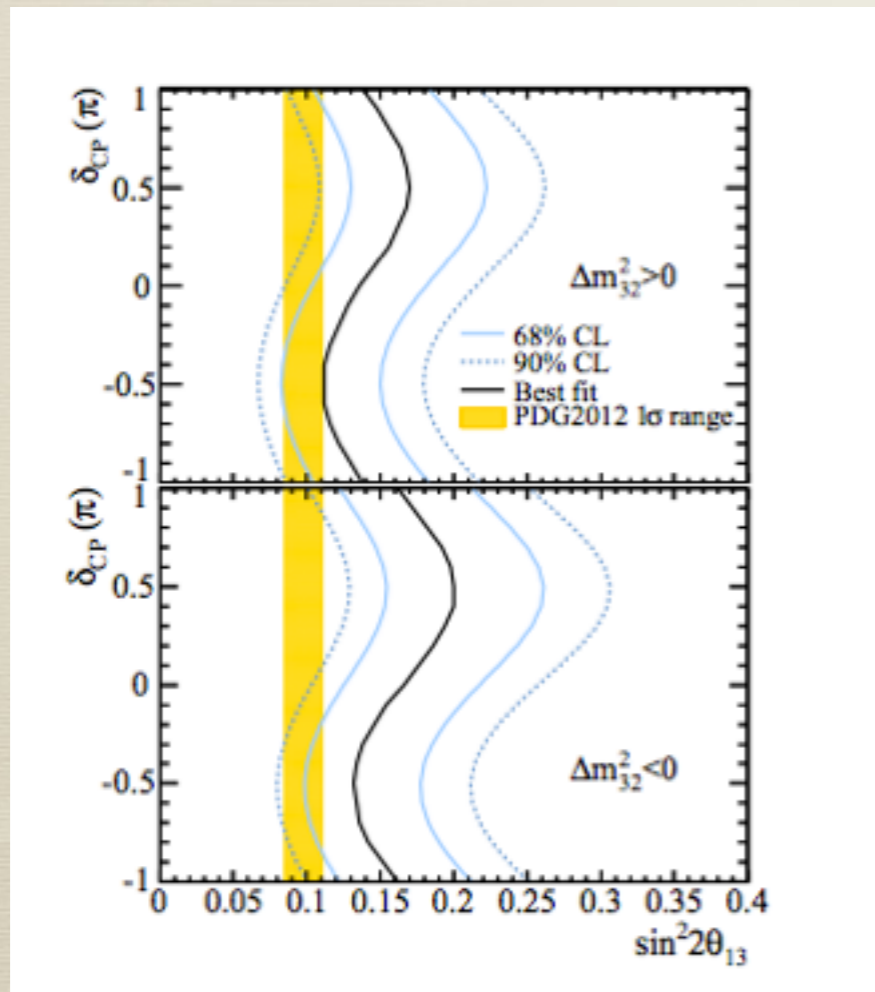
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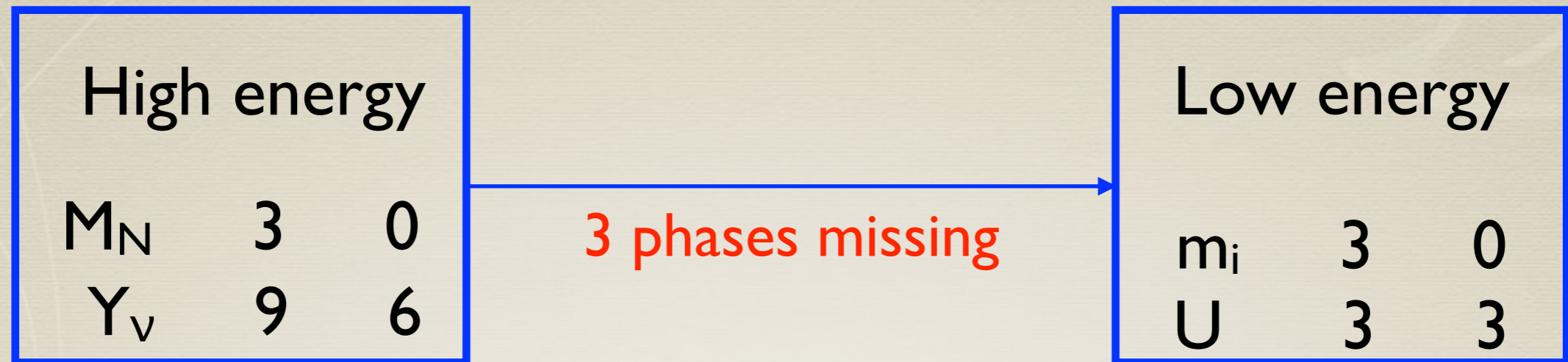
- Wide range of future experiments: **positive signal** would mean **L-violation**.
- $\beta 0\nu\nu$  can probe neutrino mass ordering, absolute mass scale and CPV phases (in principle).

# CPV in the lepton sector

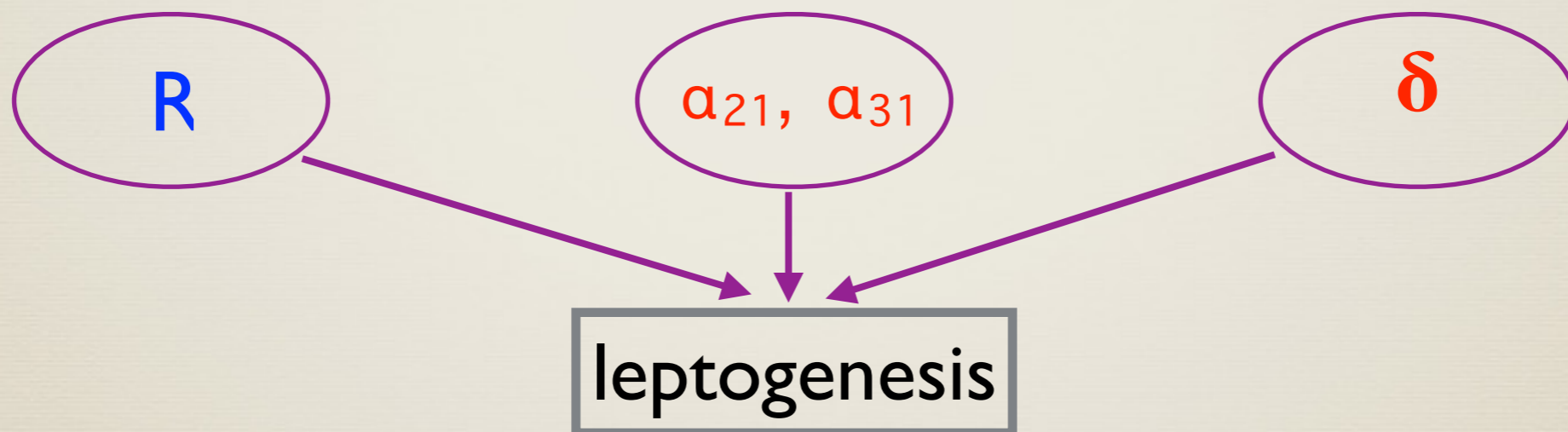
- Is **CPV** in  $\delta$  sufficient to produce the **baryon asymmetry**?
- T2K and reactor data have shown a **slight preference** for **maximal CPV**.



# Consider type-I see saw with 3 hierarchical N



$$\epsilon_l = -\frac{3M_1}{16\pi v^2} \frac{\sum_{l\beta\rho} m_\beta^{\frac{1}{2}} m_\rho^{\frac{3}{2}} U_{l\beta}^* U_{l\rho} R_{1\beta} R_{1\rho}}{\sum_\beta m_\beta |R_{1\beta}|^2}$$

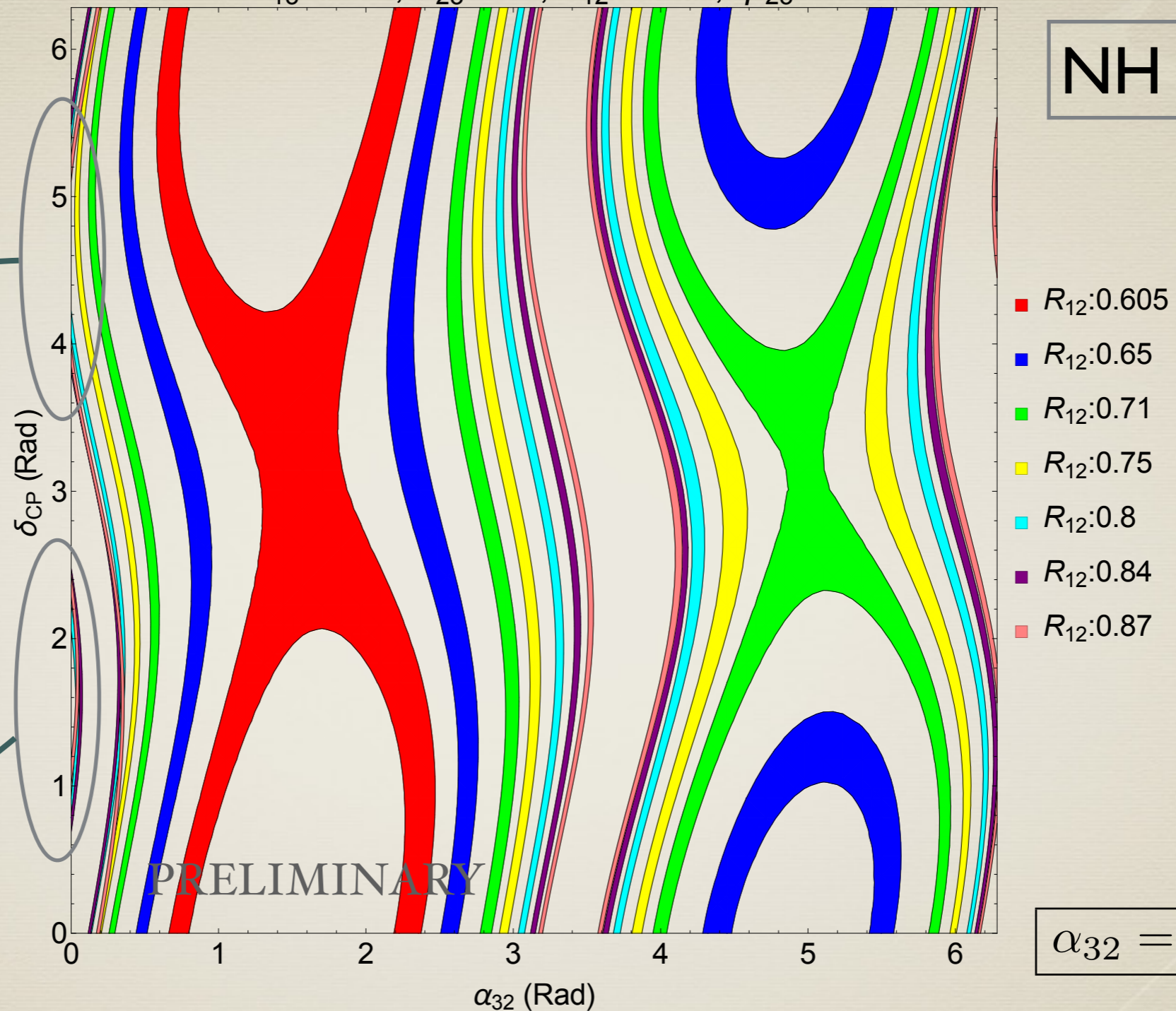


In collaboration  
with Mark  
Ross-Lonergan,  
Silvia Pascoli  
and Serguey  
Petcov

$|Y_B|$  consistent with experiment  
 $\theta_{13} 8.5^\circ, \theta_{23} 45^\circ, \theta_{12} 33.5^\circ, \beta_{23} 0^\circ$

$$M_N = 5 \times 10^{11} \text{ GeV}$$

NH spectra



$\delta$  alone can  
produce  
baryon  
asymmetry

$$\alpha_{32} = \alpha_{31} - \alpha_{21}$$

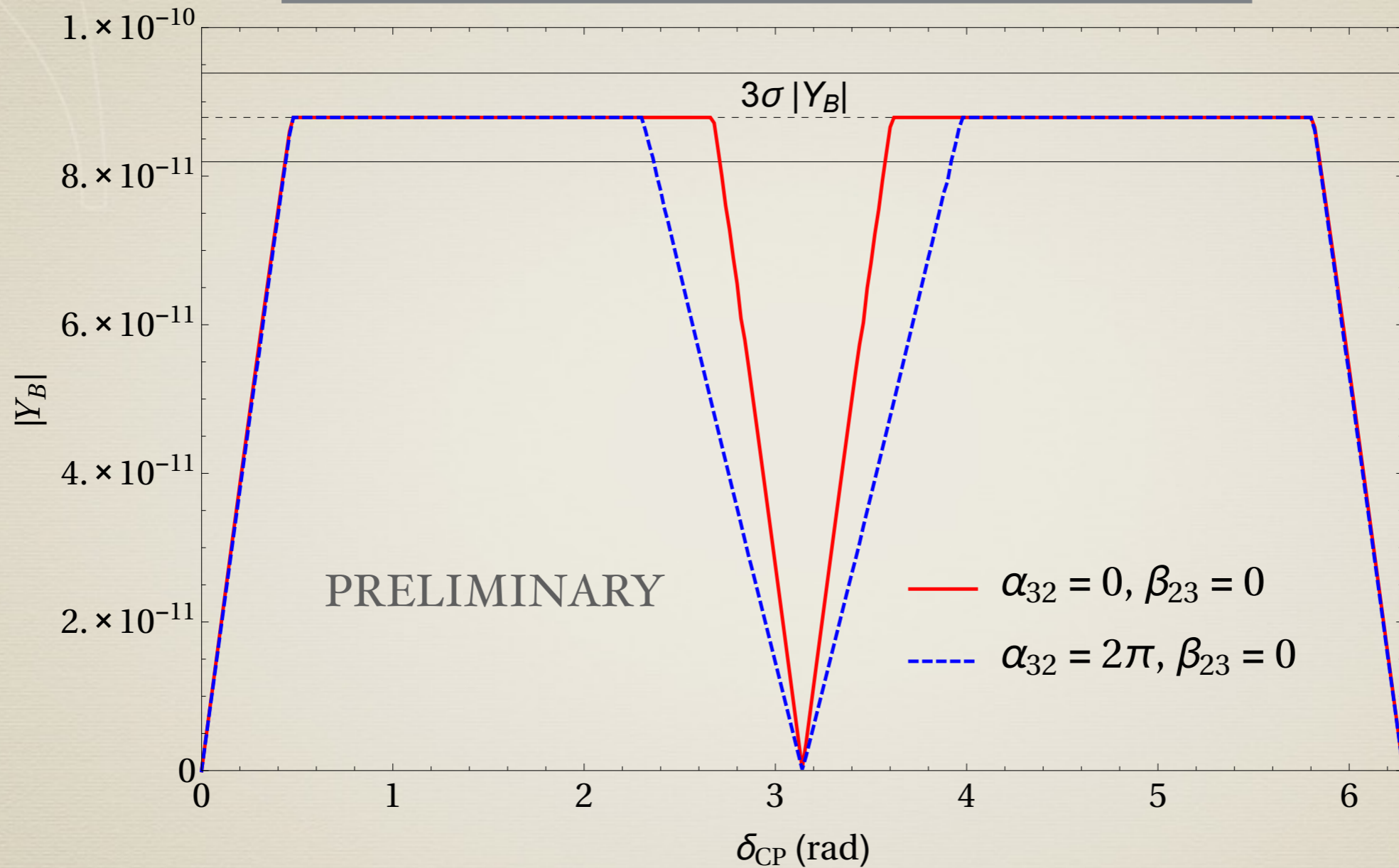
$\alpha_{32}$  is dominant but contribution from  $\delta$  can produce baryon  
asymmetry

$$\alpha_{32} = \alpha_{31} - \alpha_{21}$$

$$M_N = 5 \times 10^{11} \text{ GeV}$$

all other parameters have been marginalised

NH spectra



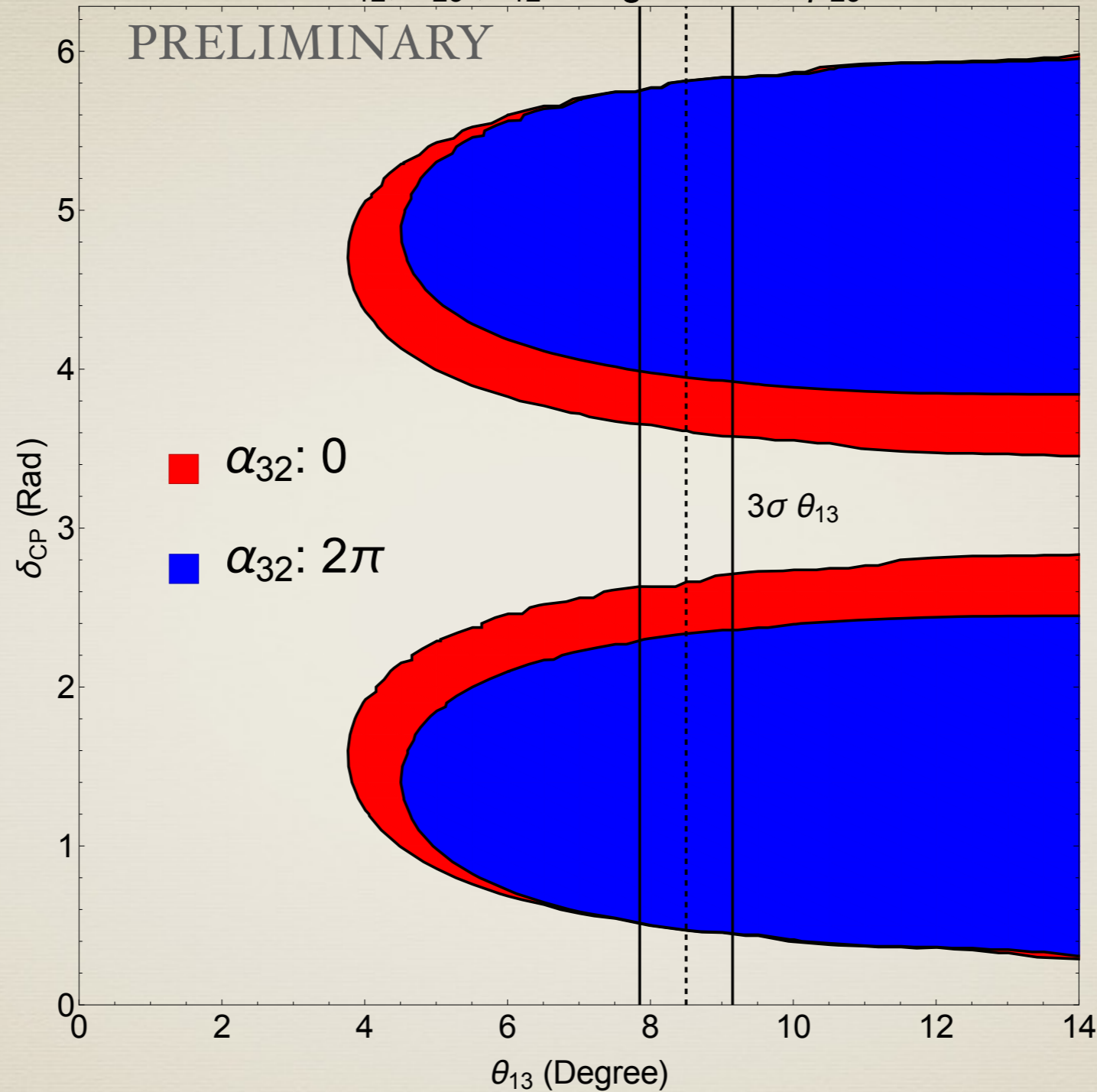
$\delta$  can produce baryon asymmetry

$|Y_B|$  consistent with experiment  
with  $R_{12}, \theta_{23}, \theta_{12}$  Marginalised,  $\beta_{23} 0^\circ$

$$M_N = 5 \times 10^{11} \text{ GeV}$$

$$\alpha_{32} = \alpha_{31} - \alpha_{21}$$

NH spectra



Need non-zero value of  $\theta_{13}$  to produce baryon asymmetry



# Conclusions

- CP violation and L violation are key ingredients for leptogenesis to occur.
- Interesting hints of maximal CPV from LBL experiments. Future LBL can make precision measurements of  $\delta$ .
- $\beta 0\nu\nu$  could indicate L-violation and Majorana CPV. Indication of link between lepton sector and leptogenesis.

In the context of see-saw type I and flavour effects, the CPV phase could generate the baryon asymmetry