



Status of India-based Neutrino Observatory



Md. Naimuddin
University of Delhi

NNN15

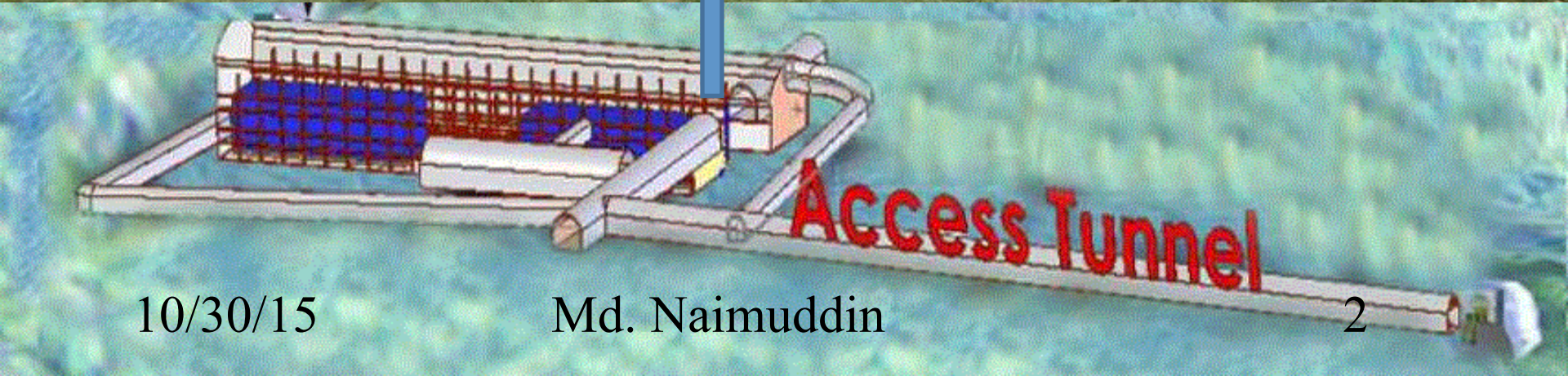
October 28 – 31, 2015

Stony Brook University, USA

Md. Naimuddin

10/30/15

1



10/30/15

Md. Naimuddin

2



INO and ICAL Detector



- At the beginning of this year, we received approval for building the INO facility near Madurai in south India.
- A cavern of dimensions $132\text{m} \times 26\text{m} \times 32.5\text{m}$ will be constructed at the end of a 1.91 km long tunnel.
- INO will have a 50 kilotons magnetized Iron Calorimeter (ICAL) to detect the atmospheric muonic neutrinos and anti neutrinos interactions.
- Uniqueness of this experiment is its capability to differentiate between a positive charged muon and a negatively charged muon and thus between a muon neutrino and a muon anti-neutrino that produces it.



The INO-ICAL White Paper



Physics Potential of the ICAL detector at the India-based Neutrino Observatory (INO)

The ICAL Collaboration

arXiv:1505.07380v1 [physics.ins-det] 27 May 2015



Physics with INO



- Atmospheric neutrinos provide a wider range for E and L than any artificial neutrino source.
- An ability to discriminate between neutrinos and anti-neutrinos enables efficient determination of neutrino mass ordering independent of CP phase.
- Accurate determination of the atmospheric parameters (θ_{23} octant, deviation of θ_{23} from maximality)
- Determination of neutrino mass hierarchy (large θ_{13} helps)

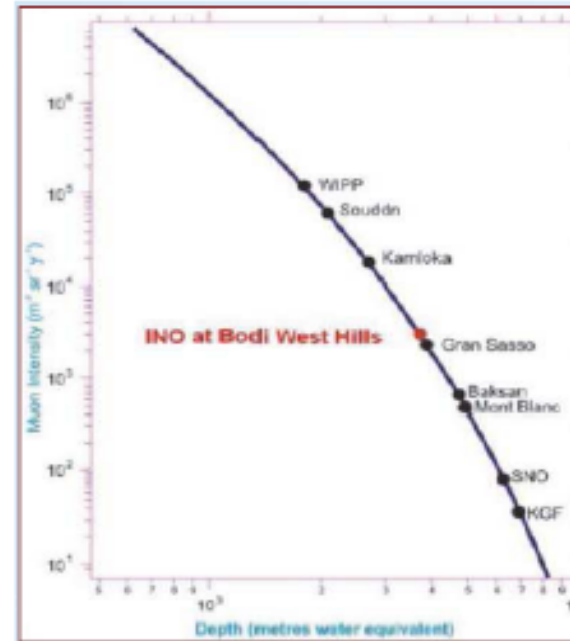
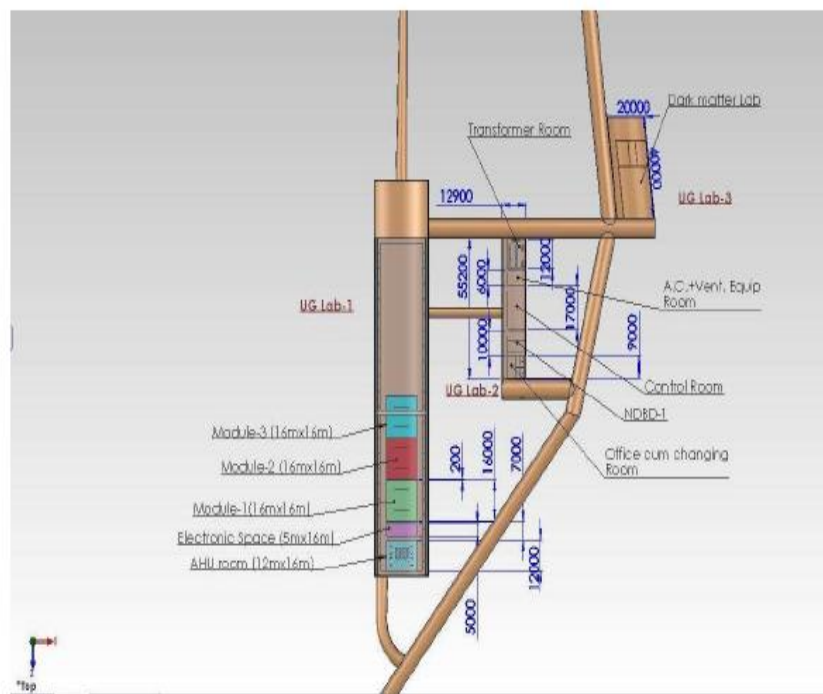


Physics with INO



- Determination of CP violation in the lepton sector
- Nonstandard interactions, CPT violation, long range forces, ultrahigh energy muon fluxes, ...
- Hadron shower reconstruction allows access to neutrino energy and high energy cosmic rays

INO: Site at a Glance



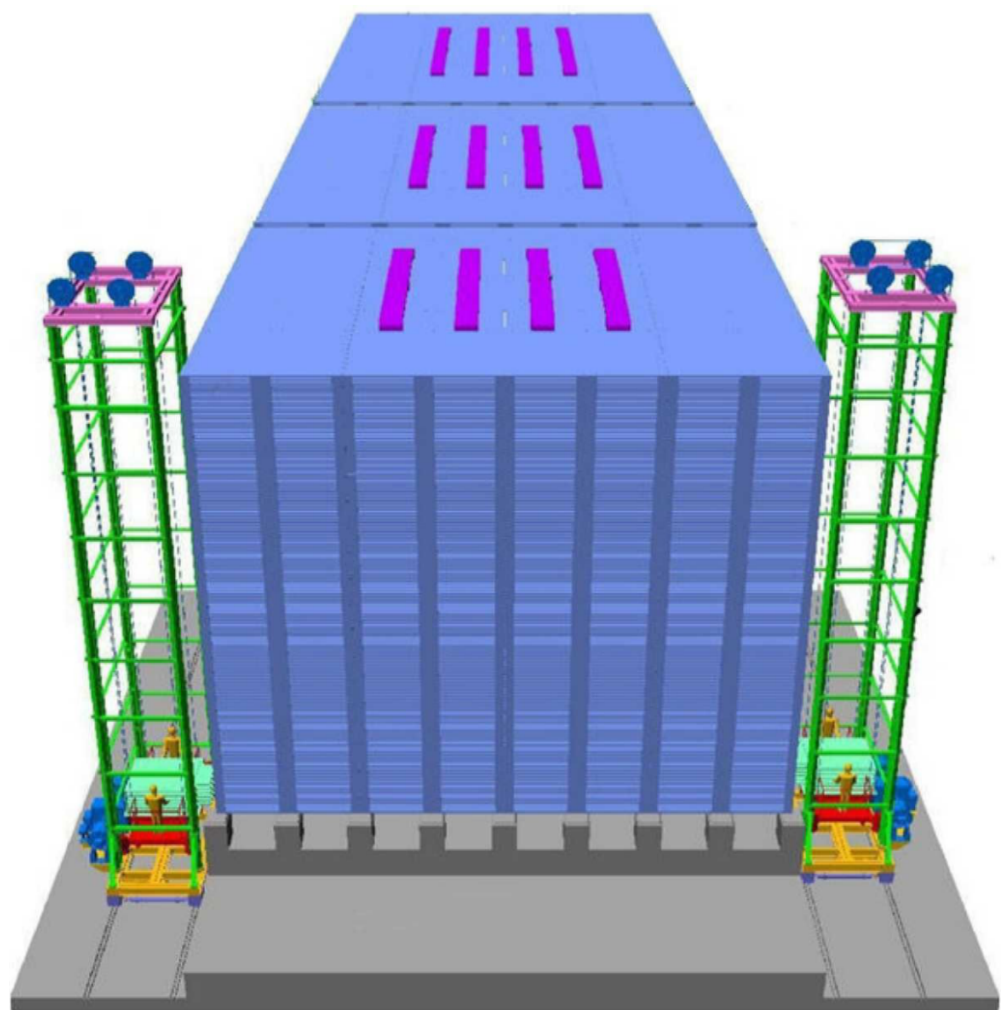
- Cavern set in Charkonite Rock under the 1589 m peak;
- Vertical cover 1289 m;
- Accessible through a 2 km tunnel
- Cavern 1 will host 50 kt ICAL (space for 100 kt);
- Other caverns for multiple experiments ($0\nu\beta\beta$, DM)

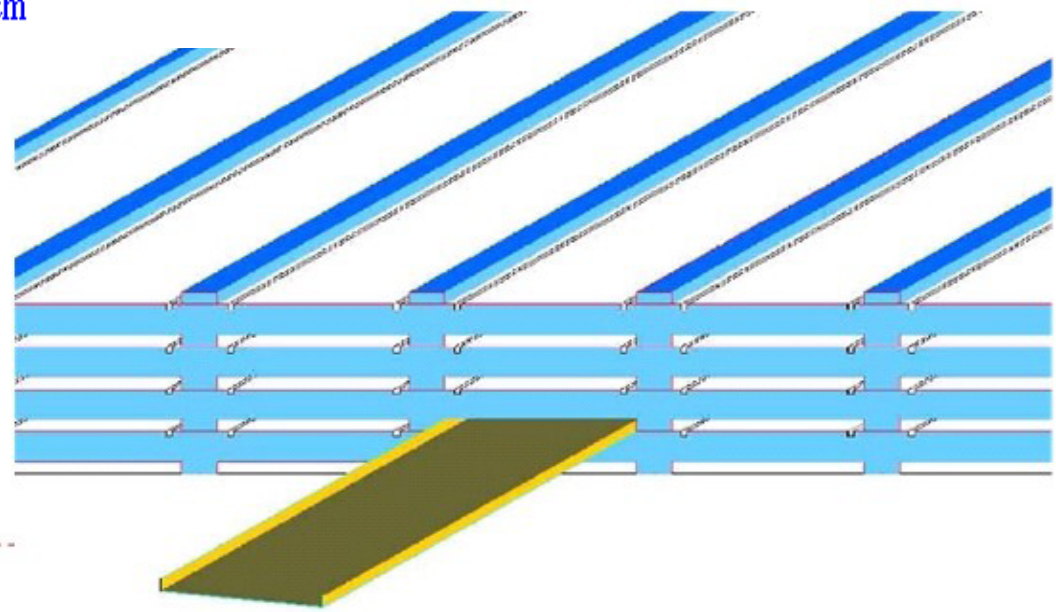
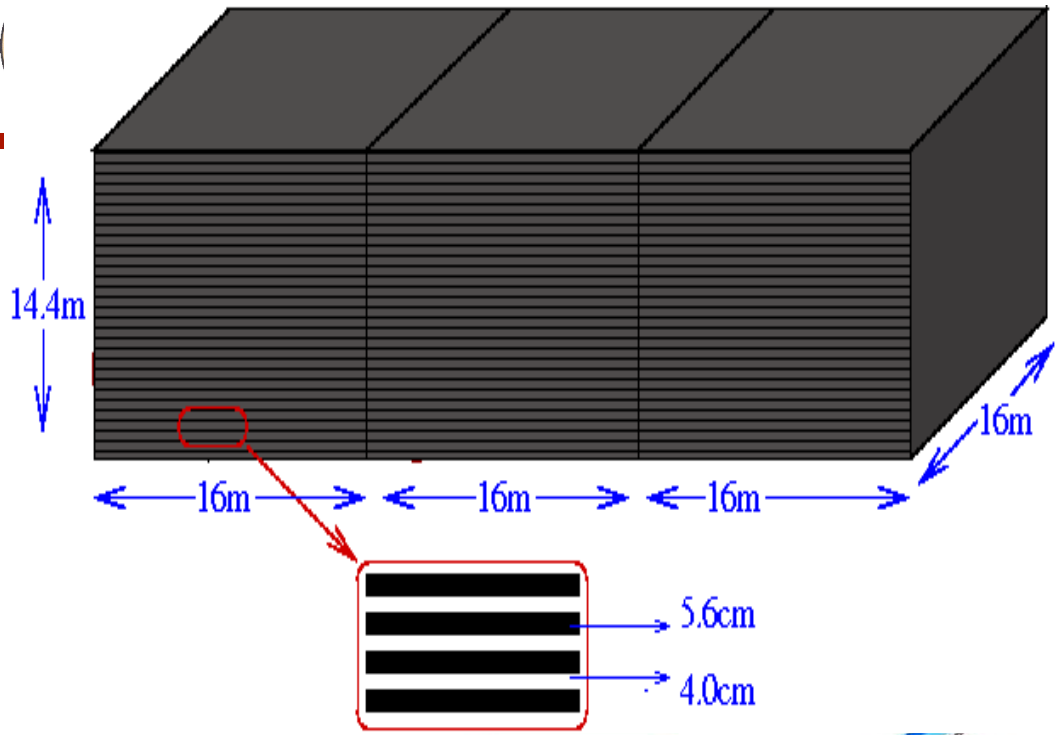
10/30/15

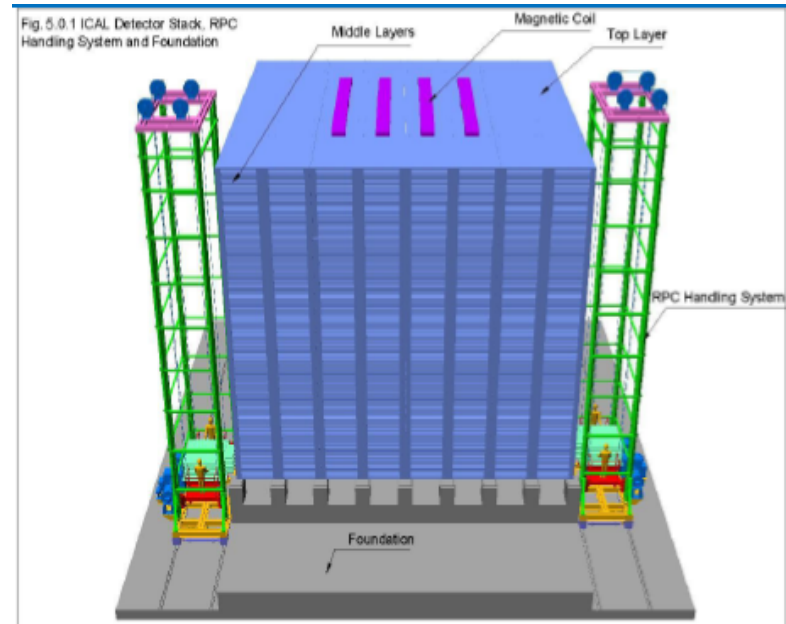
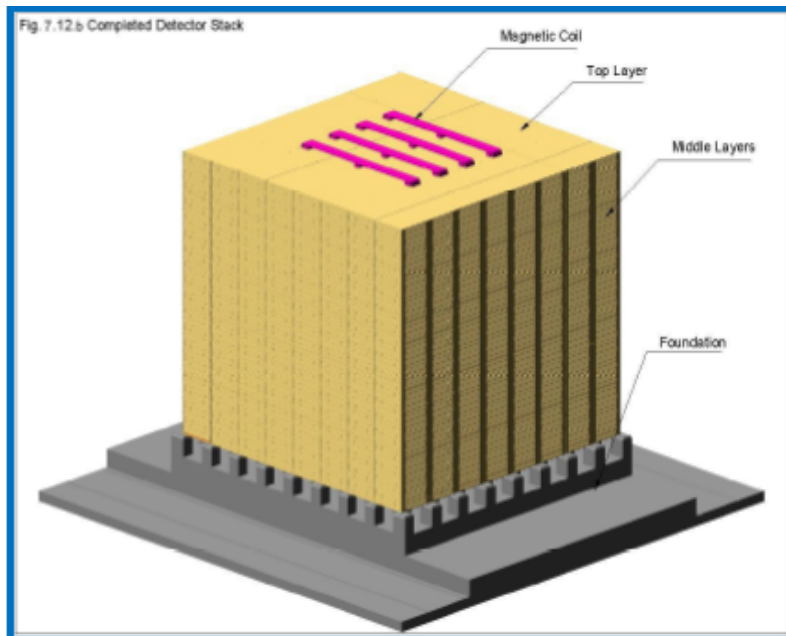
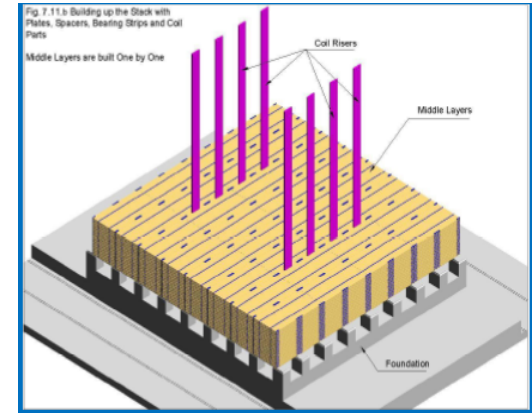
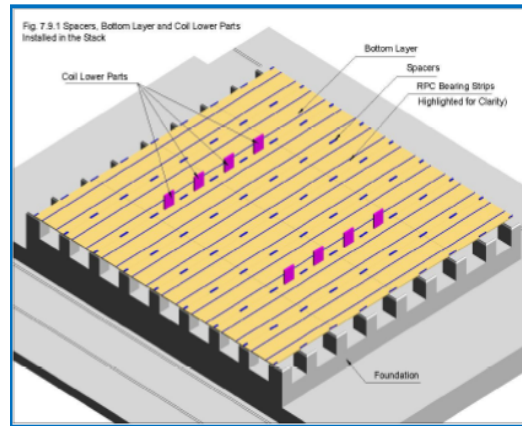
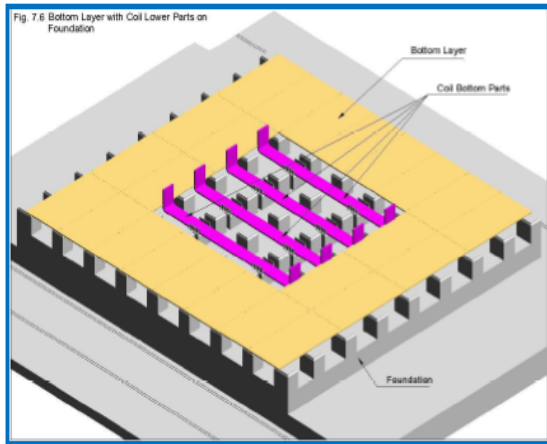
Md. Naimuddin

7

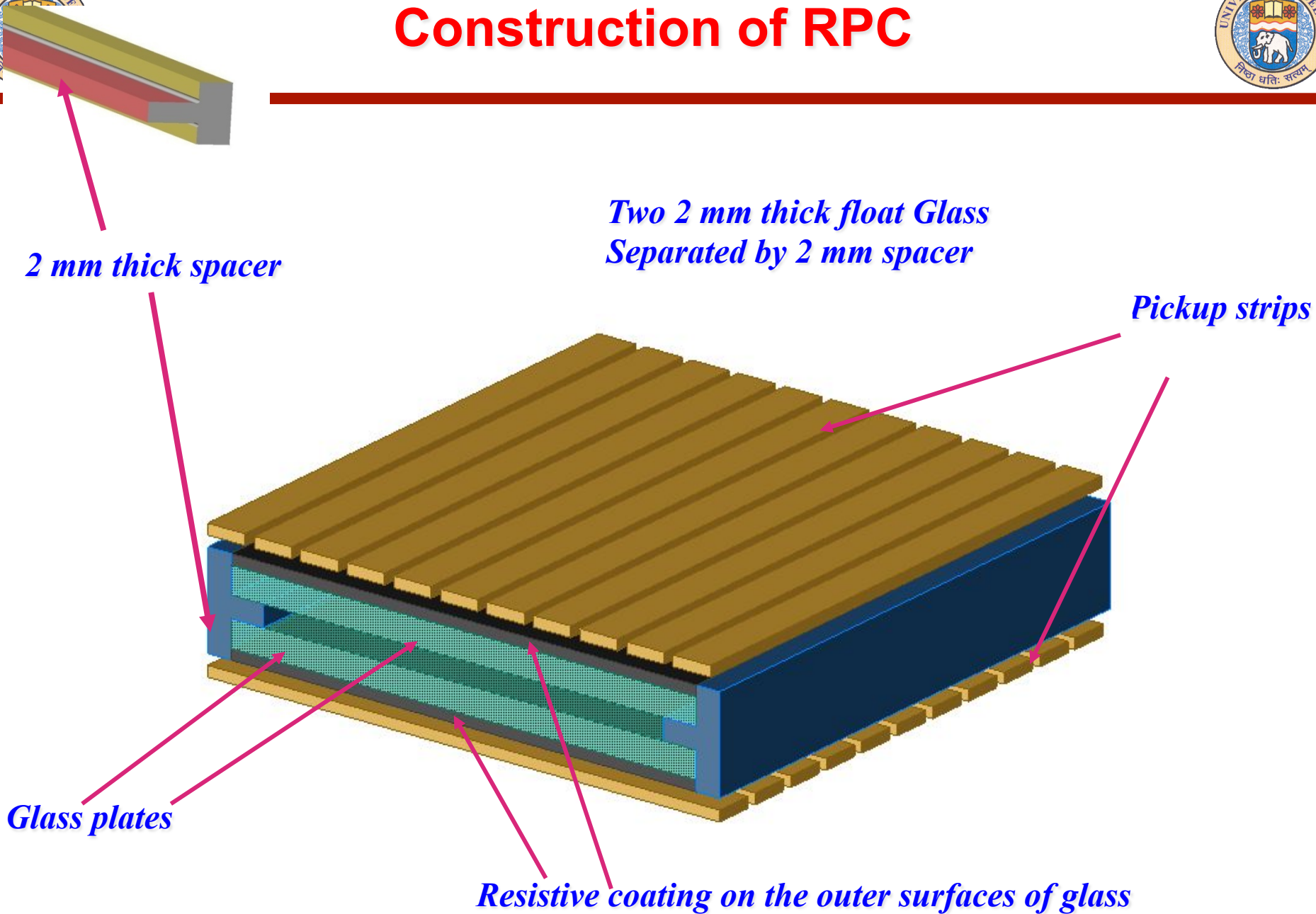
- Three modules, each of size $16\text{m} \times 16\text{m} \times 14.4\text{m}$.
- In each module 151 layers of iron plates and RPC.
- 5.6 cm Thick iron plates are separated by 4.0cm gap for RPC, act as active detector element.
- Total mass of 51kton.
- Magnetic field applied $1 \sim 1.5\text{T}$
- The readout of RPC is performed by external orthogonal pick up strips(X and Y strips).

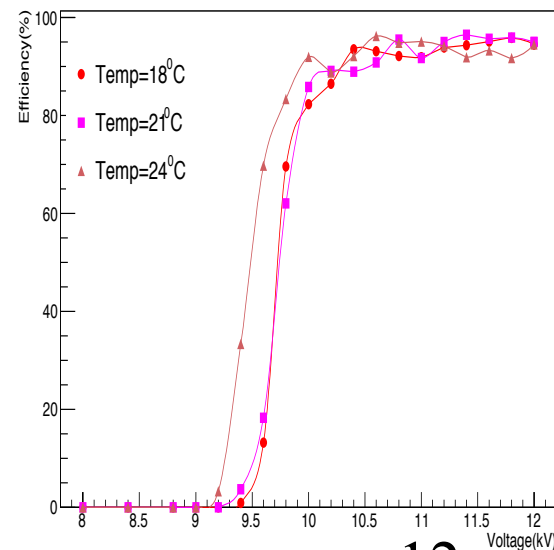
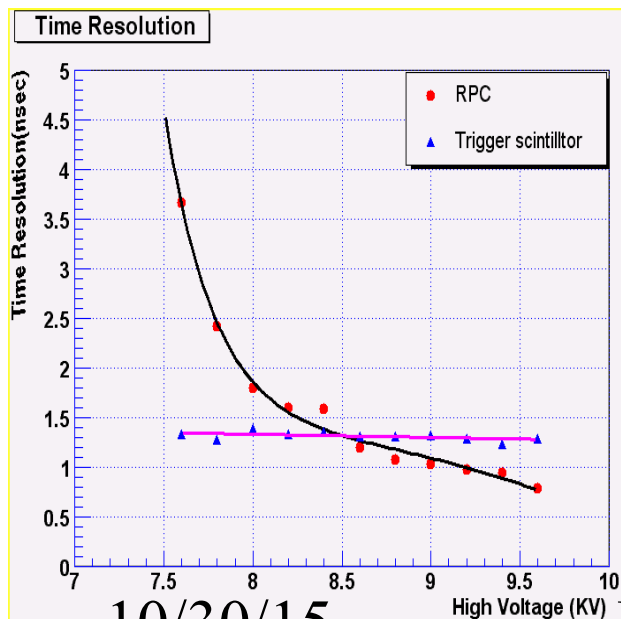
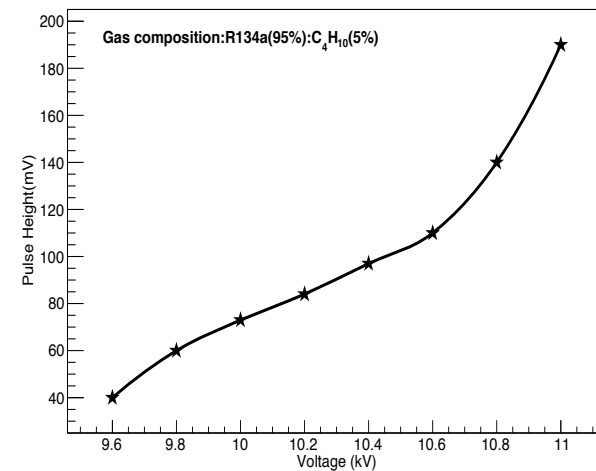
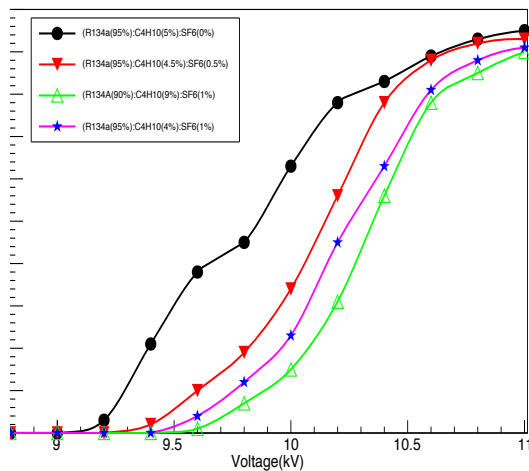
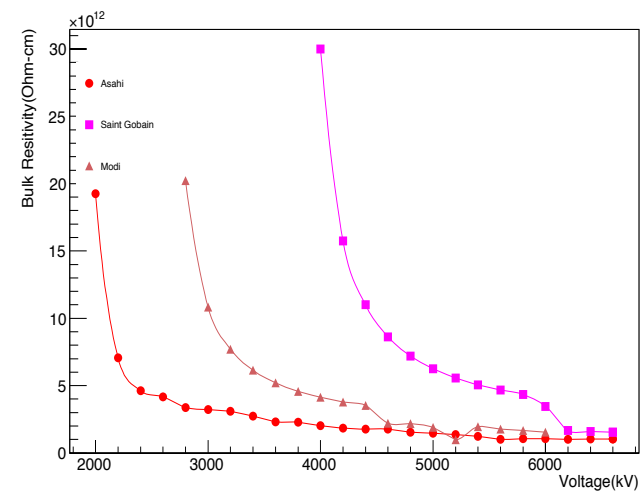






Construction of RPC







Simulation Framework



NUANCE

Neutrino Event Generation

$$\nu_\ell + N \rightarrow \ell + X .$$

Generates particles that result from a random interaction of a neutrino with matter using theoretical models for both neutrino fluxes and cross-sections.

Output:

- (i) Reaction Channel
- (ii) Vertex and time information
- (iii) Energy and momentum of all final state particles

GEANT

Event Simulation

$\ell + X$ through simulated ICAL
Simulates propagation of particles through the ICAL detector with RPCs and magnetic field.

Output:

- (i) x, y, z, t of the particles as they propagate through detector
- (ii) Energy deposited
- (iii) Momentum information

DIGITISATION

Event Digitisation

(X, Y, Z, T) of final states on including noise and detector efficiency

Add detector efficiency and noise to the hits.

Output:

- (i) Digitised output of the previous stage

ANALYSIS

Event Reconstruction

(E, \vec{p}) of ℓ, X (total hadrons)
Fit the muon tracks using Kalman filter techniques to reconstruct muon energy and momentum; use hits in hadron shower to reconstruct hadron information.

Output:

- (i) Energy and momentum of muons and hadrons, for use in physics analyses.



Particle Detection in ICAL

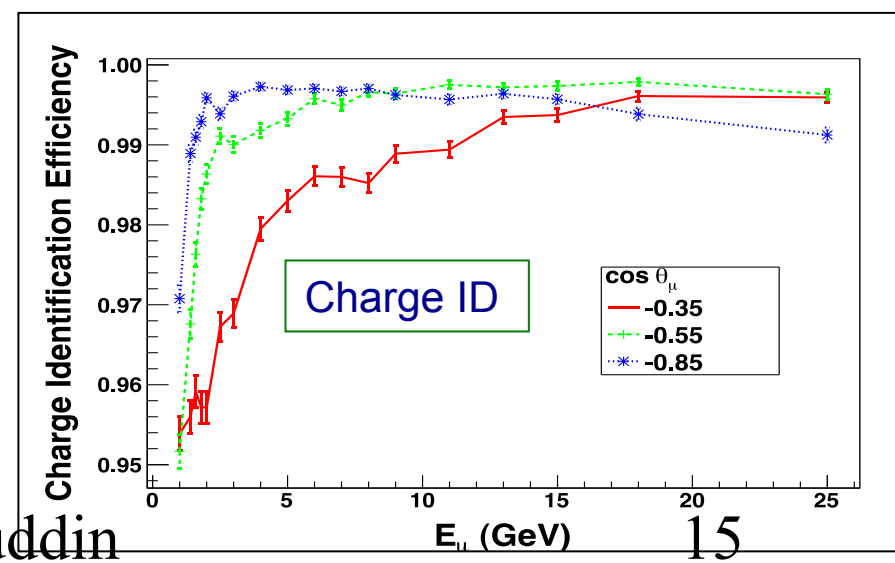
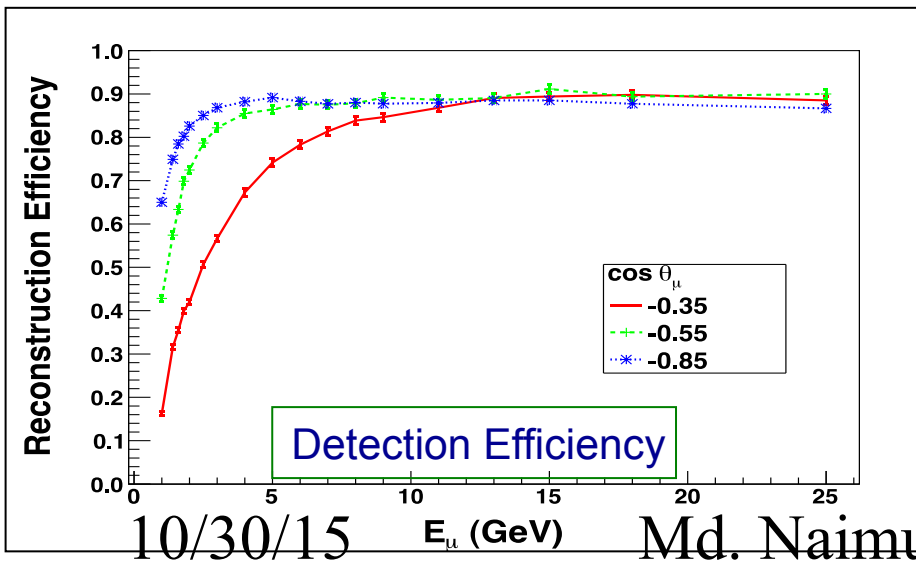
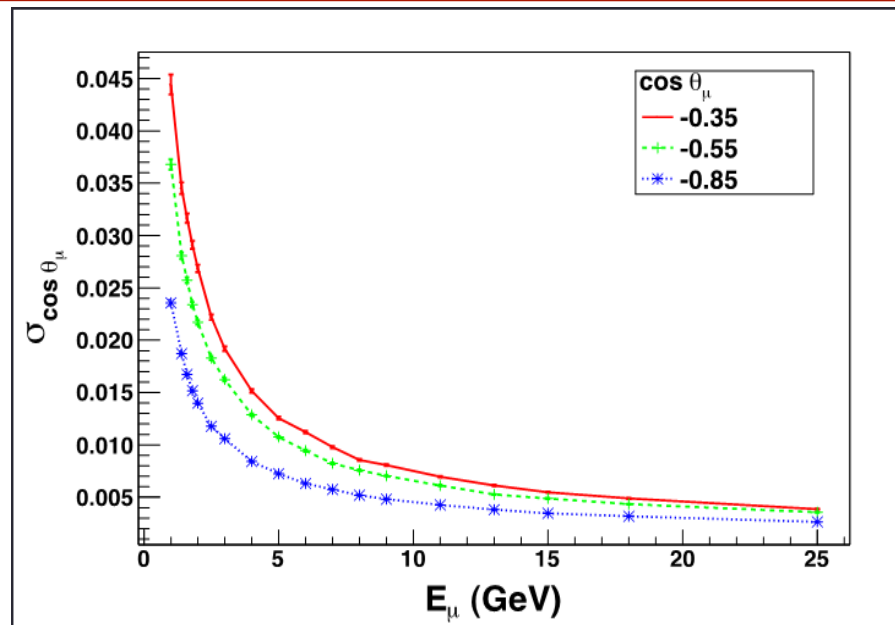
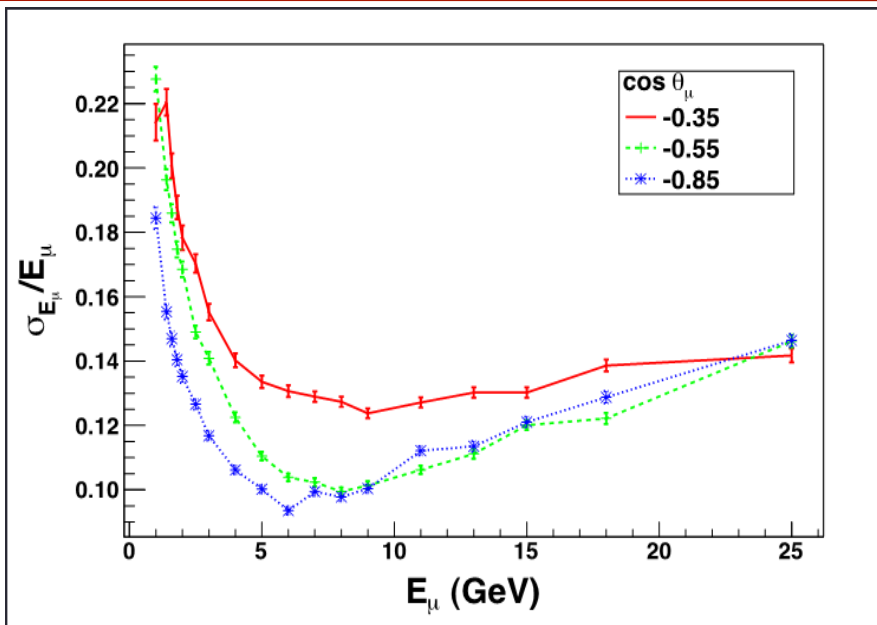
- Neutrinos interact within ICAL detector and produce associated lepton and hadronic shower.

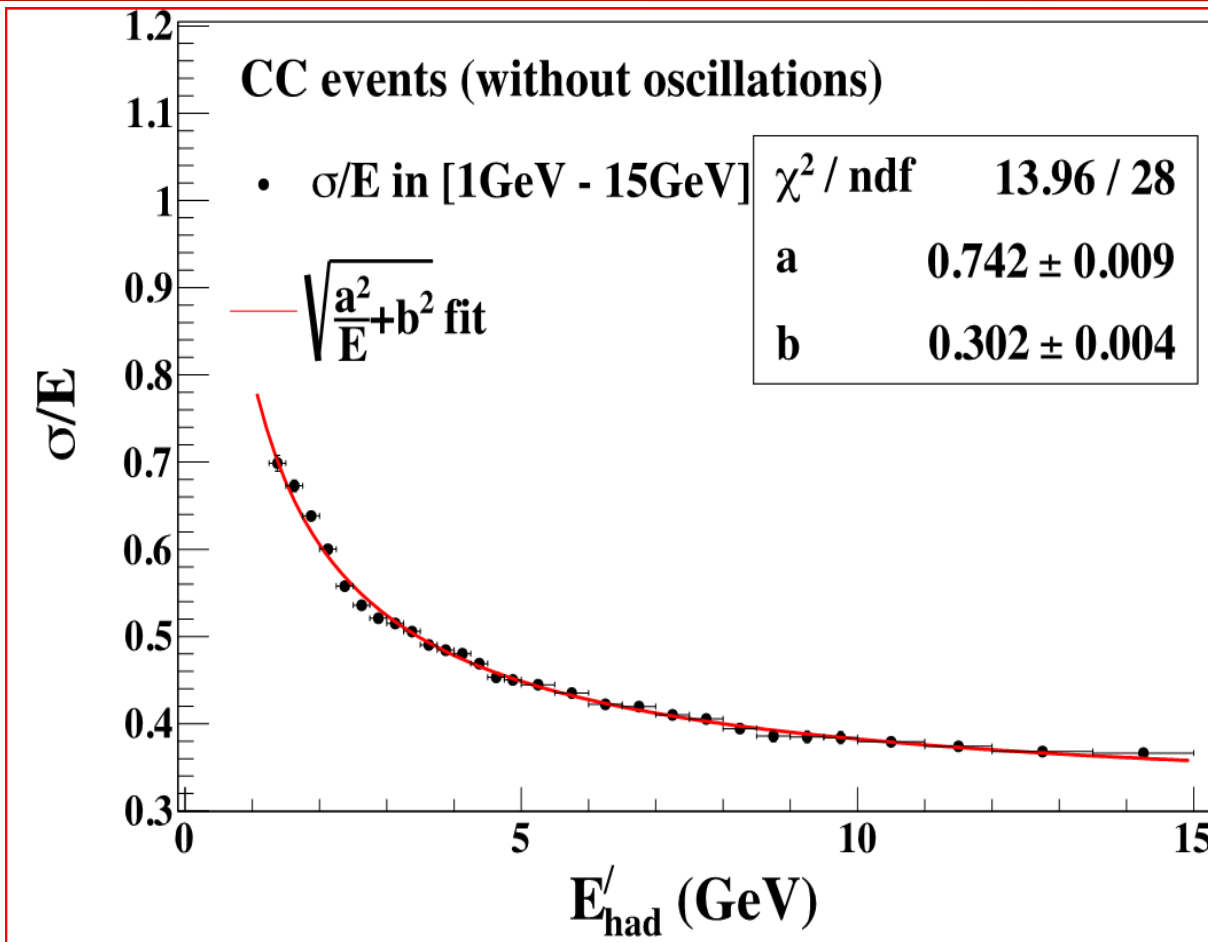
- For muon neutrino , neutrino energy (E_ν) will be the sum of muon energy (E_μ) and hadronic energy (E_h):

$$E_\nu = E_\mu + E_h$$

- To reconstruct the E_ν precisely both muon energy and hadron energy have to be measured very precisely
- Muons give a clear track inside detector, Energy of muons can be reconstructed from the track length in the detector.
- The energy of hadrons can be calibrated as a function of number of total hits.

Muon Resolutions

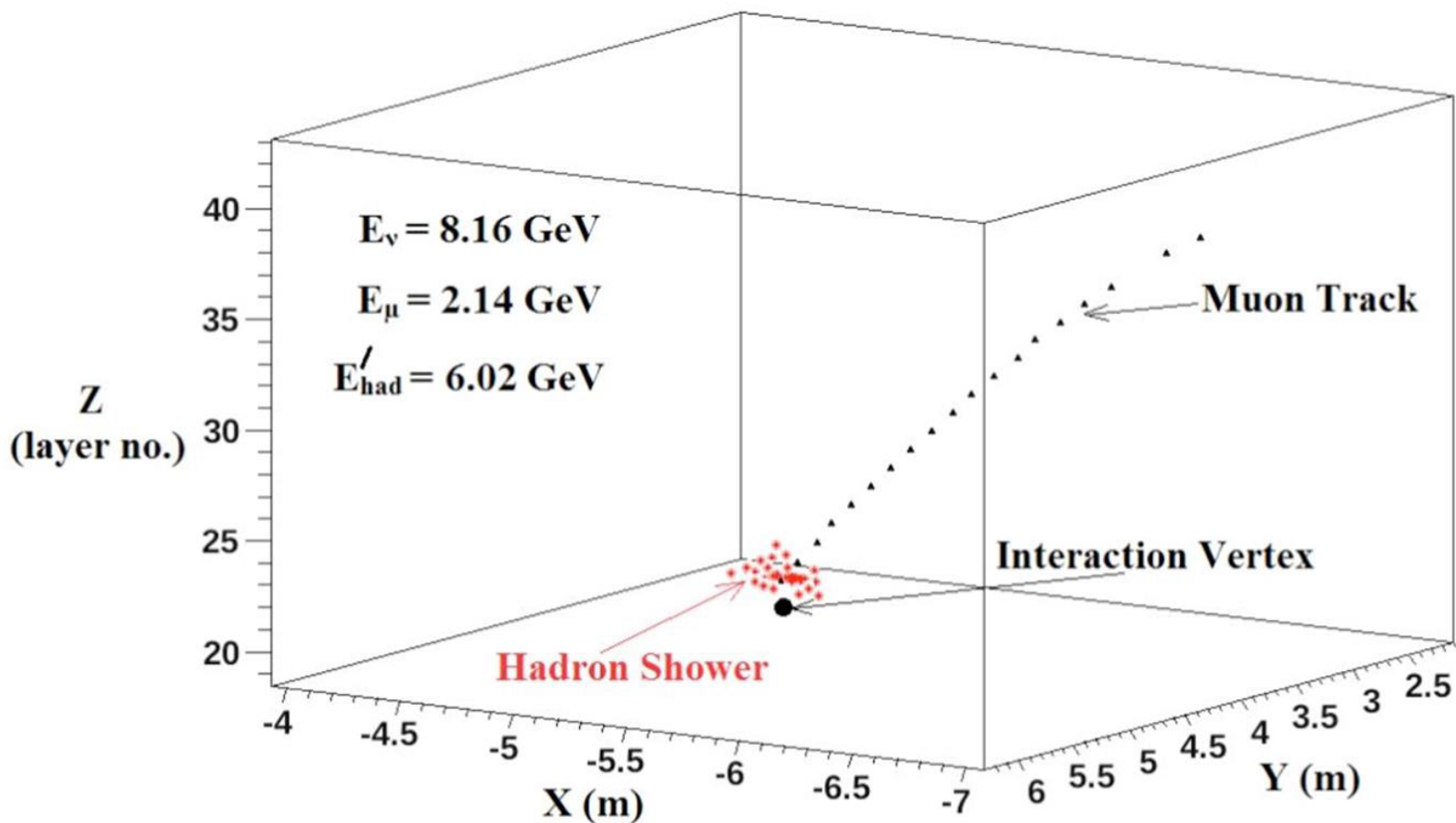




Resolution for 2GeV energy is approximately 60% and 15 GeV is approx. 36%.

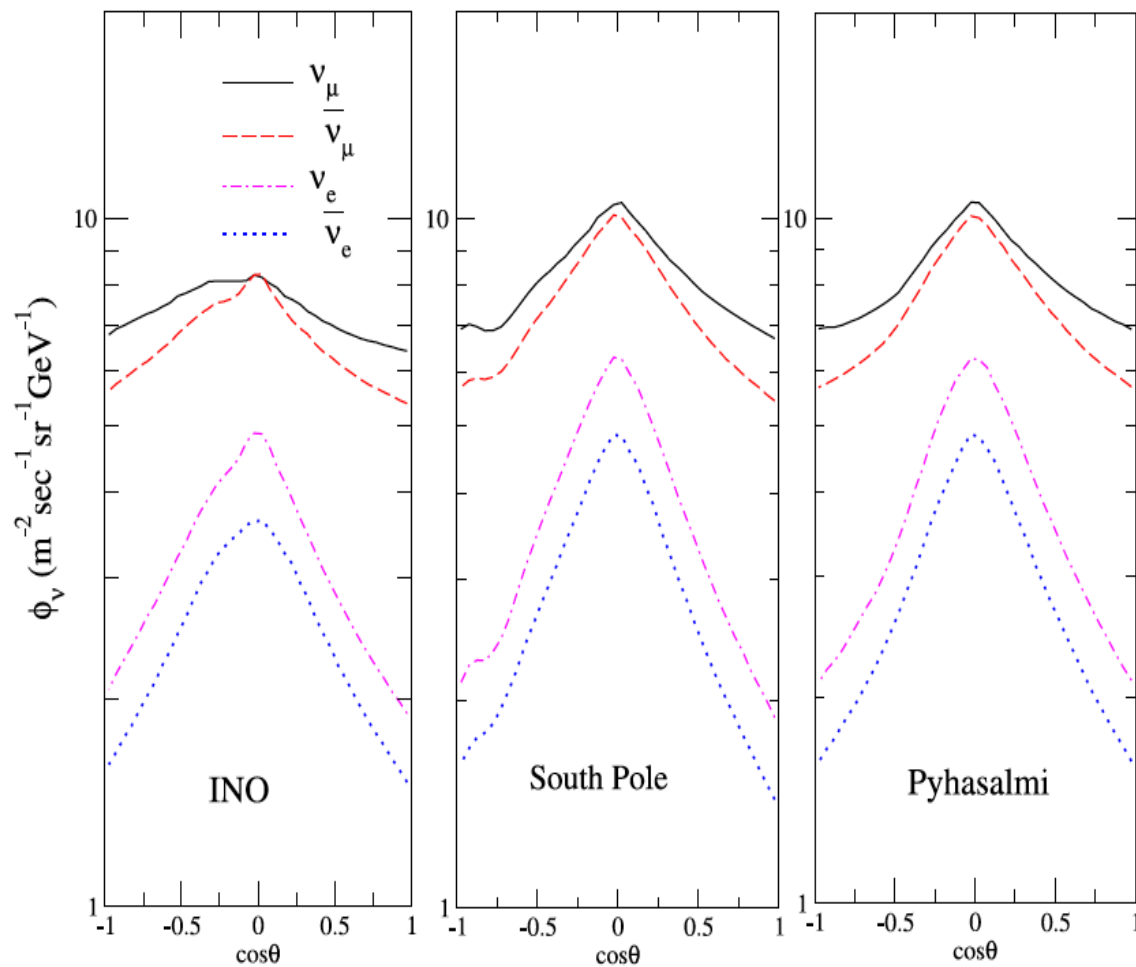
$$E'_h = E_\nu - E_\mu \text{ (from hadron hit calibration)}$$

A Typical Event in ICAL



From GEANT4 Simulation

- Atmospheric neutrino flux has been generated with NUANCE using Honda 3d fluxes for the Kamioka site in Japan.
- The Honda atmospheric fluxes at the INO site to be finalized soon.



Athar, Honda, Kajita, Kasahara, Midorikawa,
[arXiv:1210.5154 \[hep-ph\]](https://arxiv.org/abs/1210.5154)



The χ^2 Analysis



We define the Poissonian χ^2_- for μ^- events as :

$$\chi^2_- = \min_{\xi_l} \sum_{i=1}^{N_{E'_{\text{had}}}} \sum_{j=1}^{N_{E_\mu}} \sum_{k=1}^{N_{\cos \theta_\mu}} \left[2(N_{ijk}^{\text{theory}} - N_{ijk}^{\text{data}}) - 2N_{ijk}^{\text{data}} \ln \left(\frac{N_{ijk}^{\text{theory}}}{N_{ijk}^{\text{data}}} \right) \right] + \sum_{l=1}^5 \xi_l^2,$$

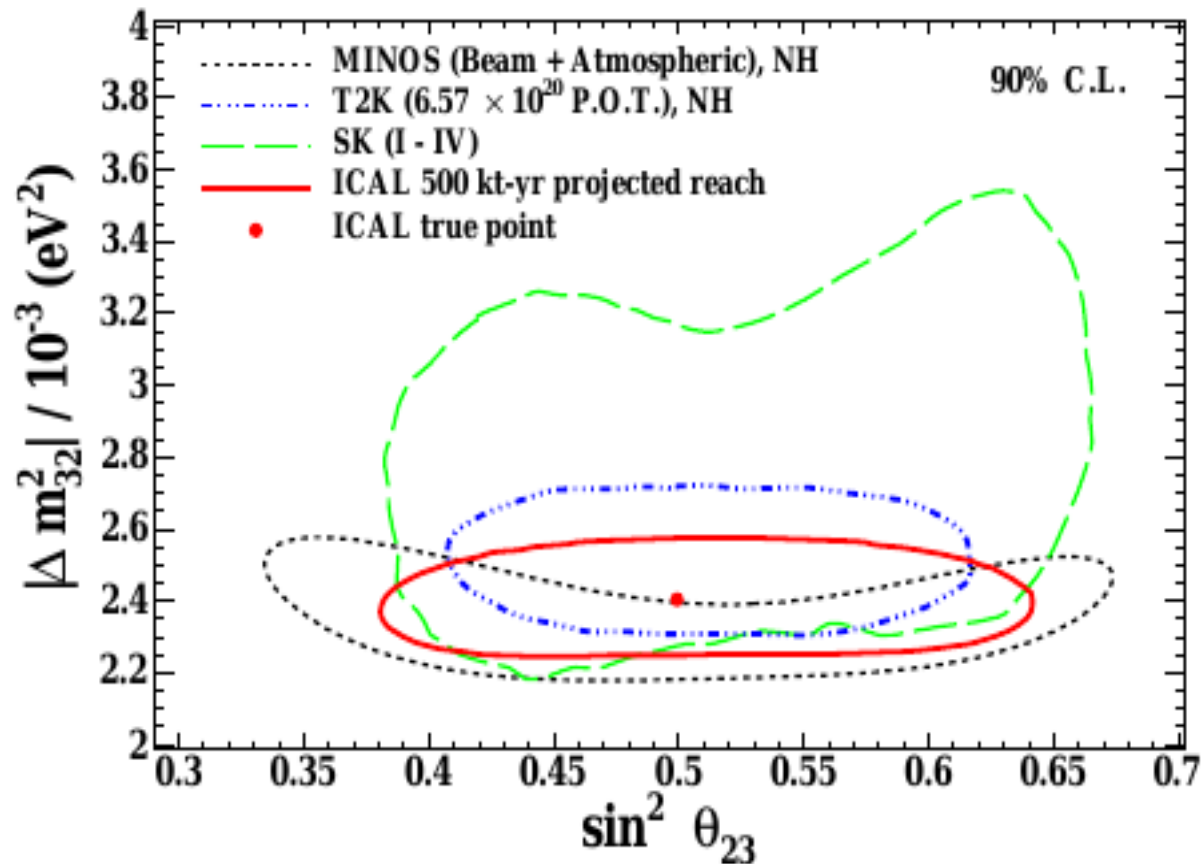
where

$$N_{ijk}^{\text{theory}} = N_{ijk}^0 \left(1 + \sum_{l=1}^5 \pi_{ijk}^l \xi_l \right).$$

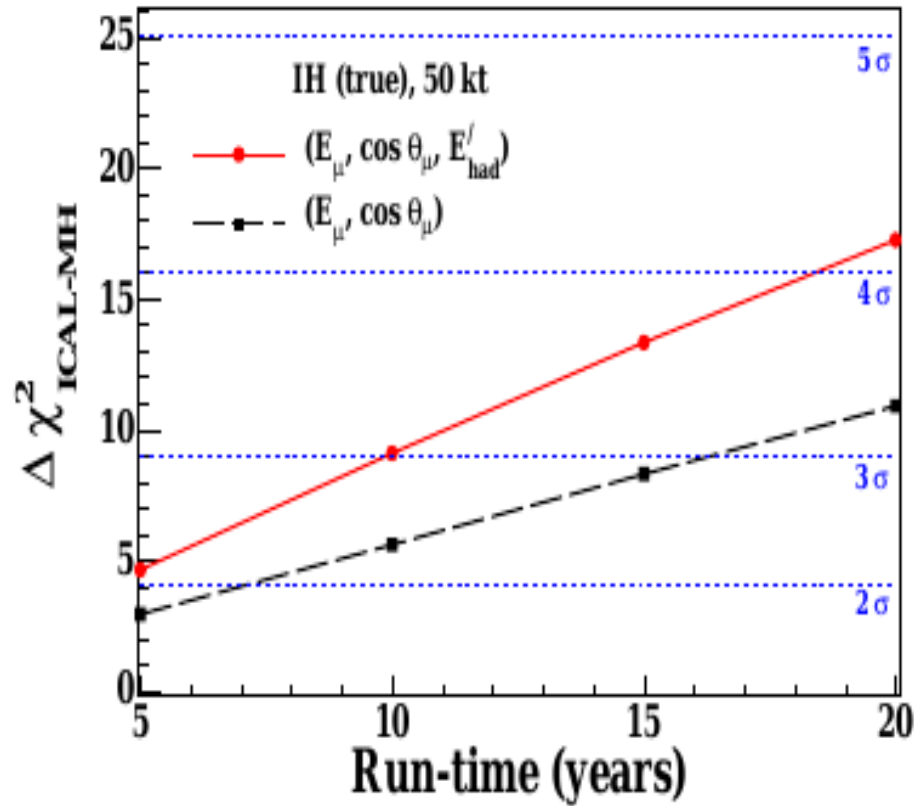
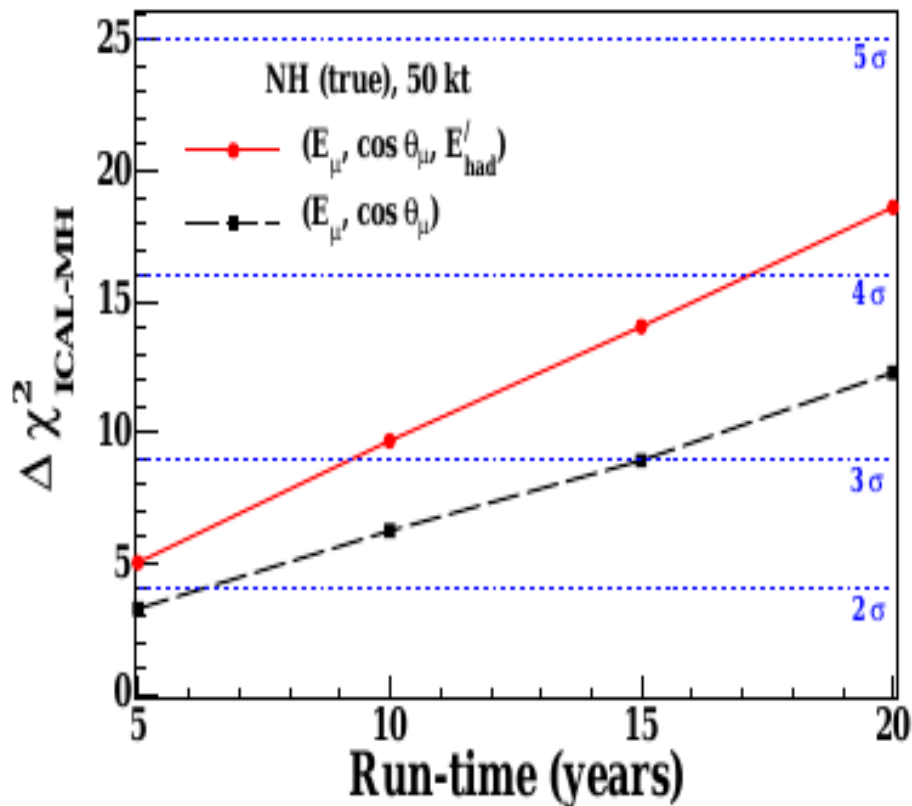
- 1) Overall 5% systematic uncertainty
- 2) Overall flux normalization: 20%
- 3) Overall cross-section normalization: 10%
- 4) 5% uncertainty on the zenith angle dependence of the fluxes
- 5) Energy dependent tilt factor:

$$\Phi_\delta(E) = \Phi_0(E) [E/E_0]^\delta \approx \Phi_0(E) [1 + \delta \ln E/E_0]$$

where $E_0 = 2$ GeV and δ is the 1σ systematic error of 5%

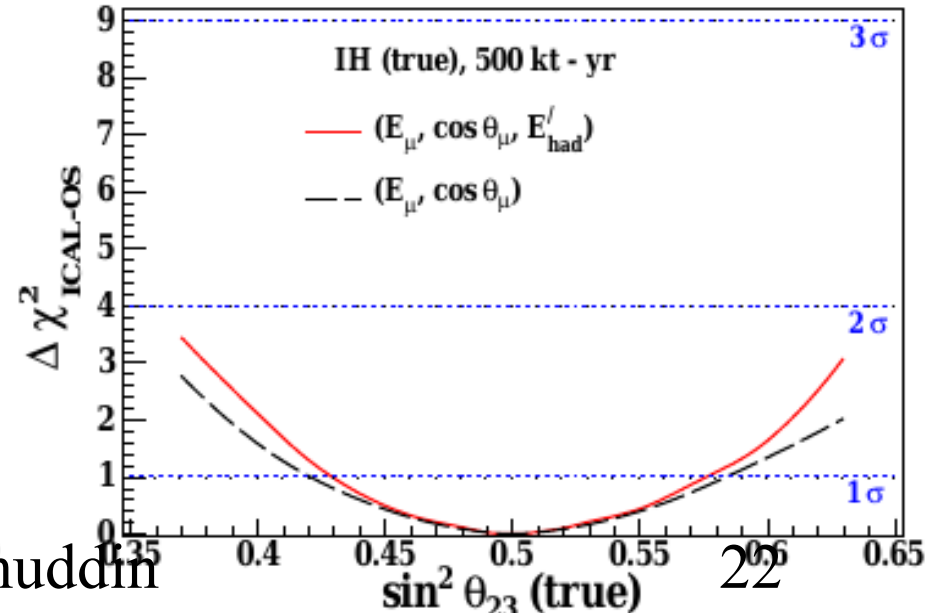
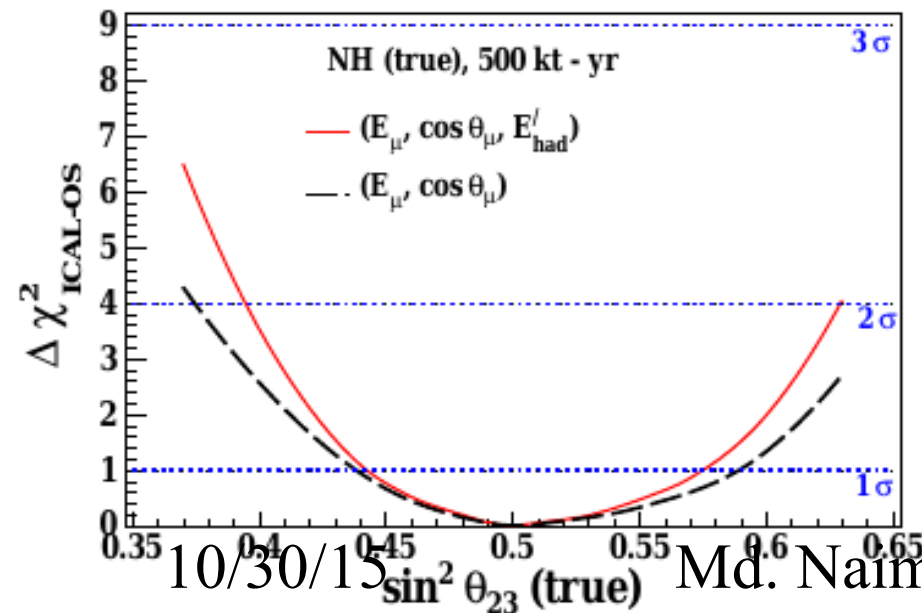
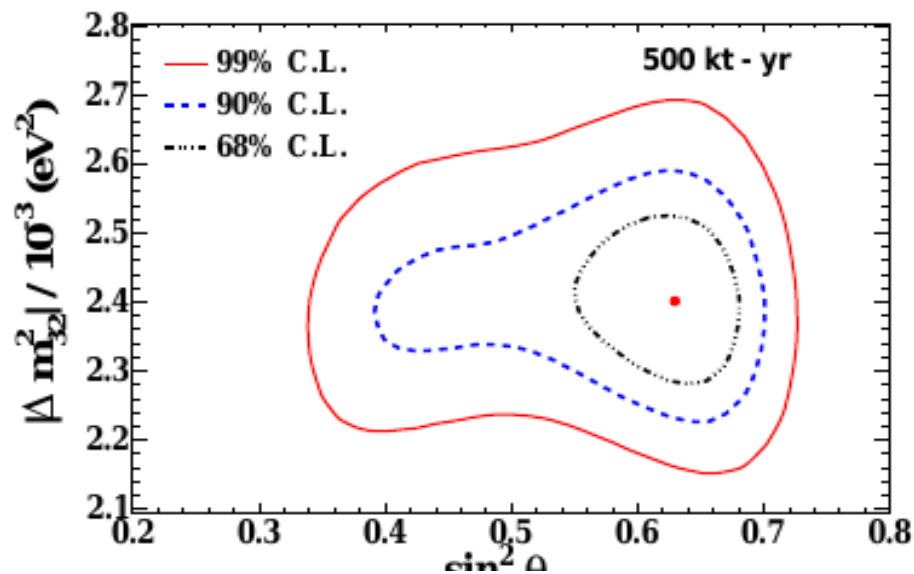
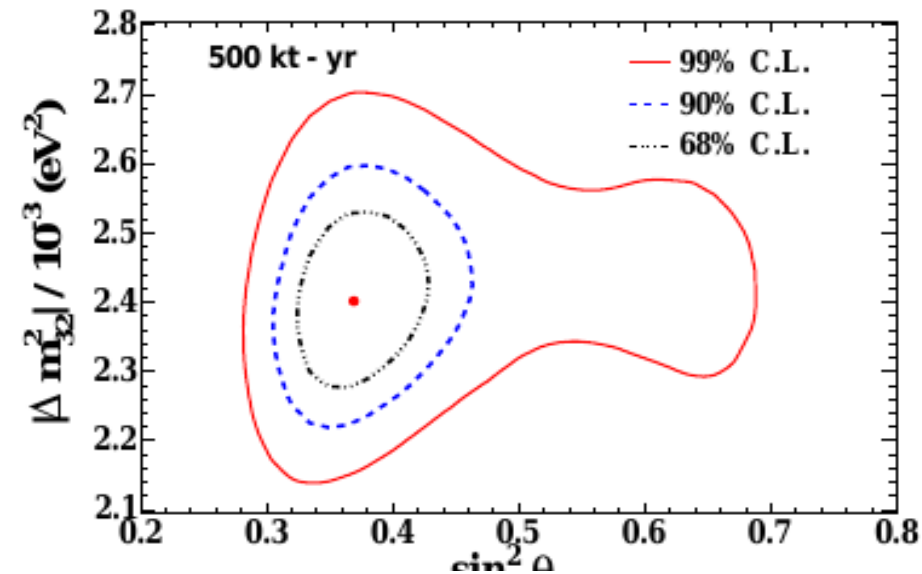


→ Use priors on $|\Delta m^2_{\text{atm}}|$, θ_{23} , θ_{13} from LBL+ reactors projected reach

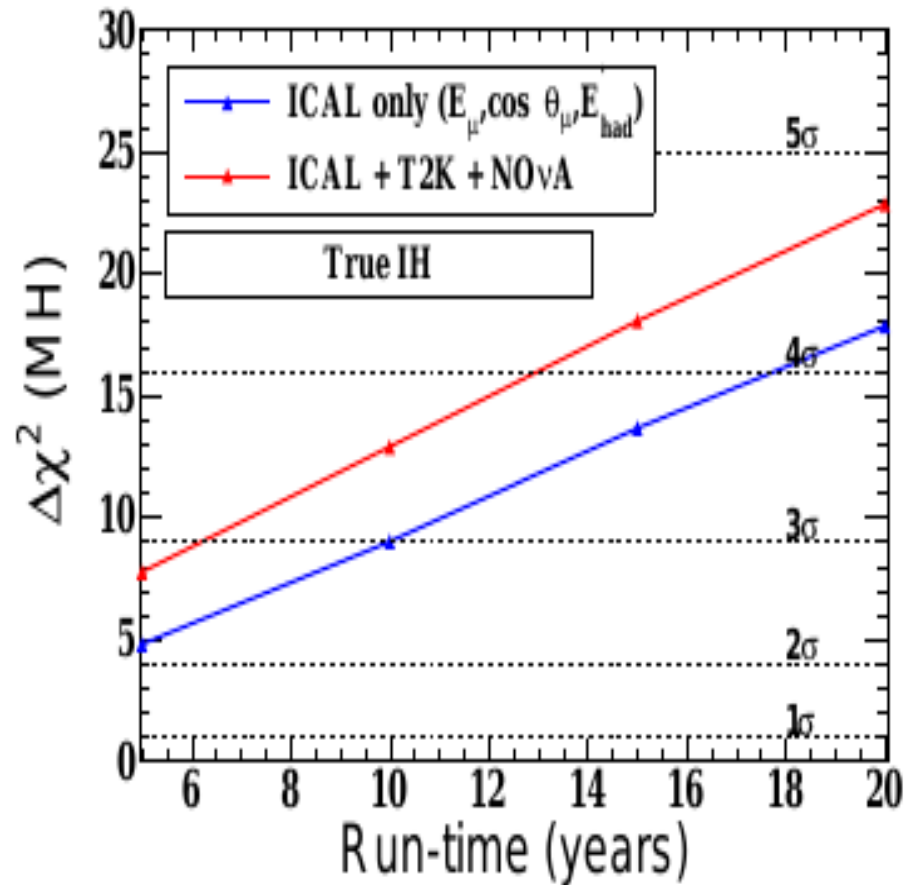
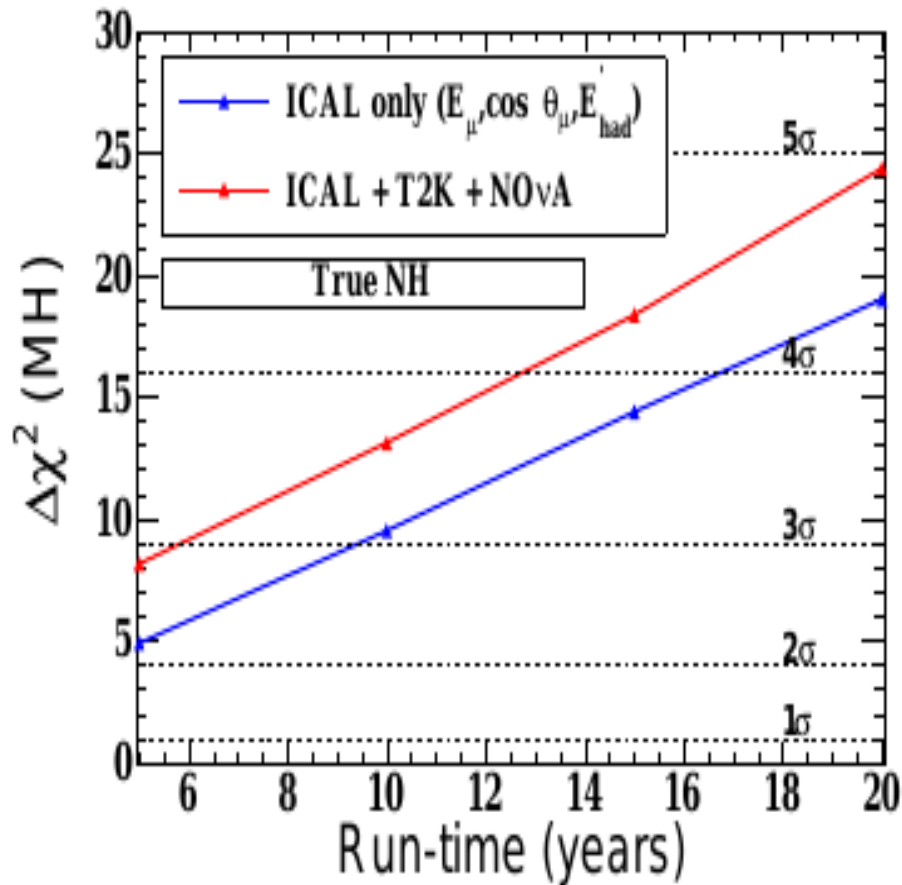


~2.3 σ sensitivity for $\sin^2\theta_{23}=0.5$, $\sin^22\theta_{13}=0.1$ by 2025 (5 yrs)
 ~3 σ sensitivity for $\sin^2\theta_{23}=0.5$, $\sin^22\theta_{13}=0.1$ by 2030 (10 yrs)

Octant Sensitivity

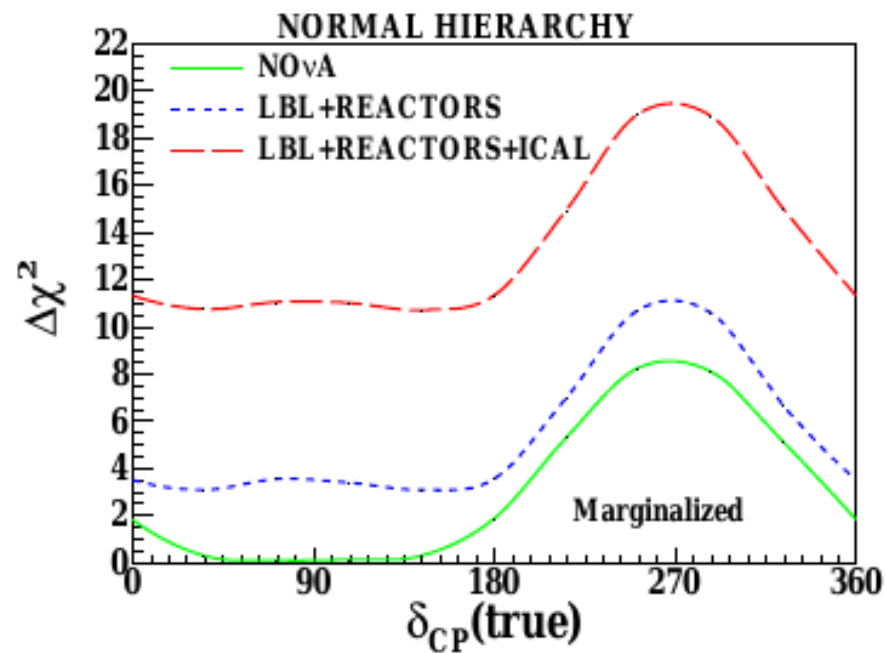
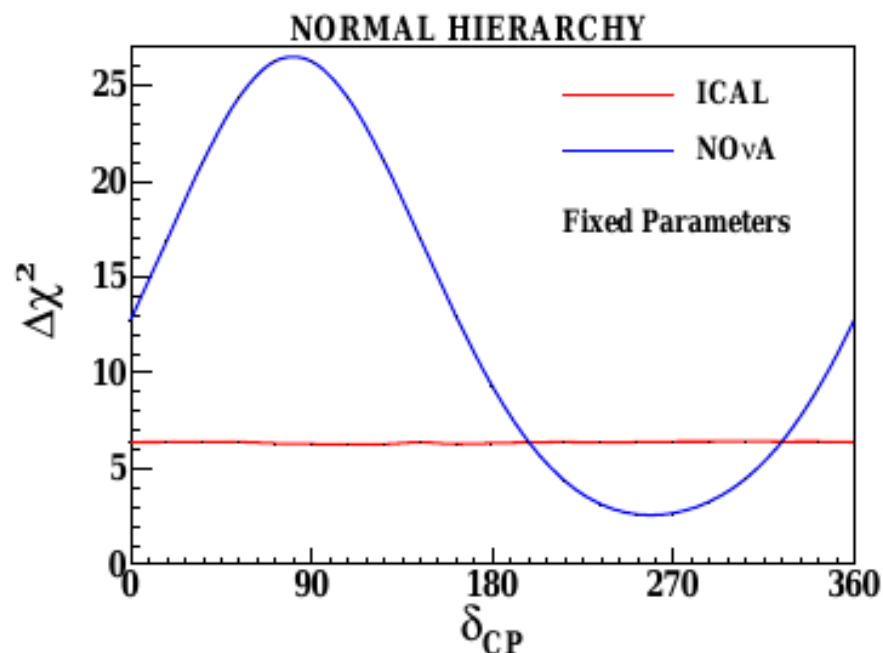


INO Combined with T2K and NOvA



~3 σ sensitivity for $\sin^2\theta_{23}=0.5$, $\sin^2\theta_{13}=0.1$ in 6 yrs for NH.

~4 σ sensitivity for $\sin^2\theta_{23}=0.5$, $\sin^2\theta_{13}=0.1$ in 13 yrs for NH.



- Though ICAL itself is rather insensitive to δ_{CP} , data from ICAL can still improve the determination of δ_{CP} itself, by providing input on mass hierarchy.
- This is especially crucial in the range $0 \leq \delta_{CP} \leq \pi$, precisely where the ICAL data would also improve the hierarchy discrimination of NOvA and other experiments



Other Analysis in Progress at ICAL



- ✓ Search for sterile neutrinos
- ✓ CPT violation and Non-Standard Interactions
- ✓ Search for magnetic monopoles
- ✓ Search for dark matter from the Sun
- ✓ Long range forces
- ✓ Exploiting NC events
- ✓ Possibilities of electron detection



Summary

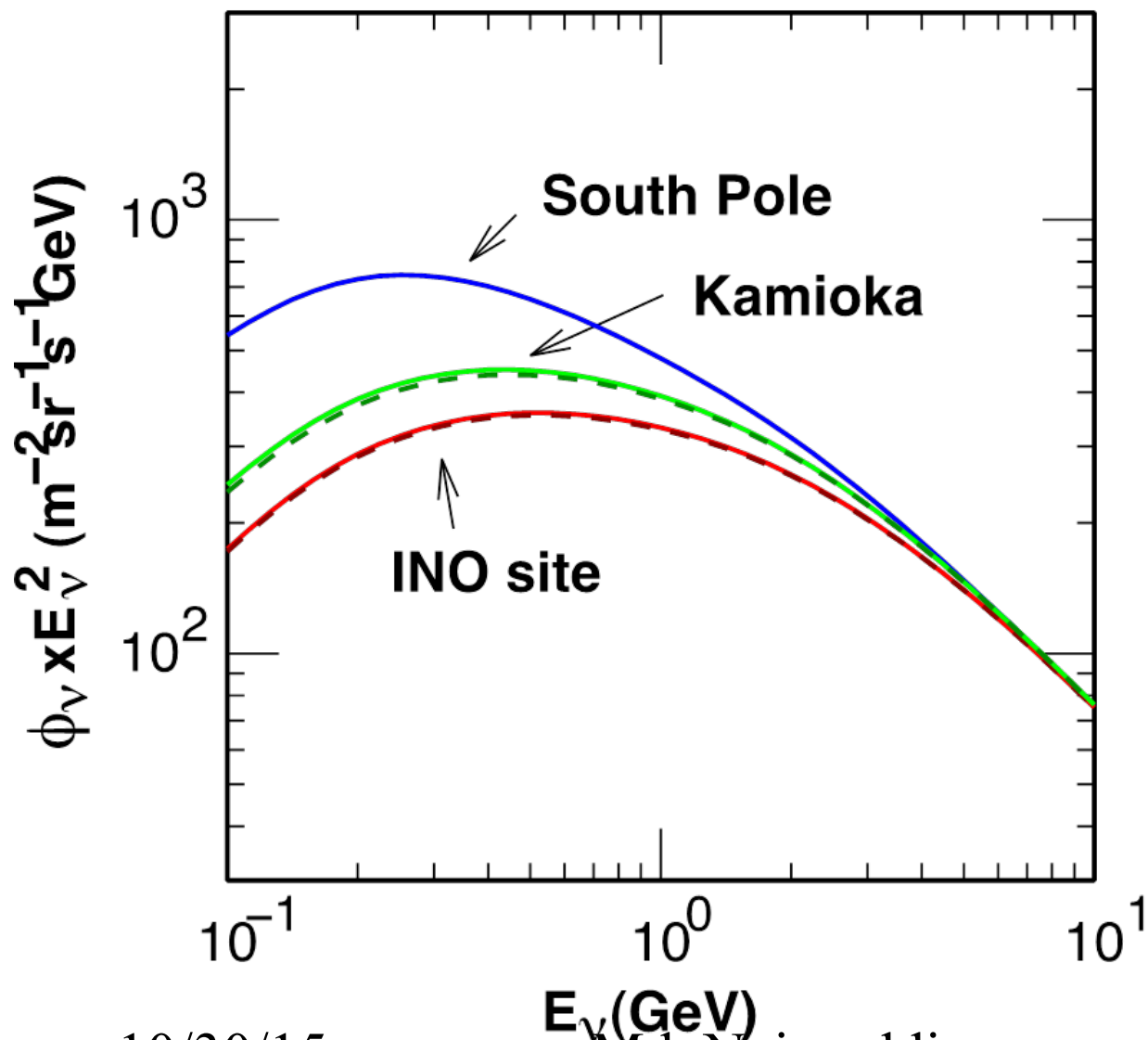


- ✓ The land for the experimental site and surface facility at Madurai have been acquired.
- ✓ Detector R&D is almost complete for the base design. Further improvements are being pursued.
- ✓ Construction of an engineering module 8m X 8m X 2.1m is being initiated.
- ✓ The work is interrupted due to some litigations pending in courts of law regarding certain clearances.
- ✓ Still a long way to go

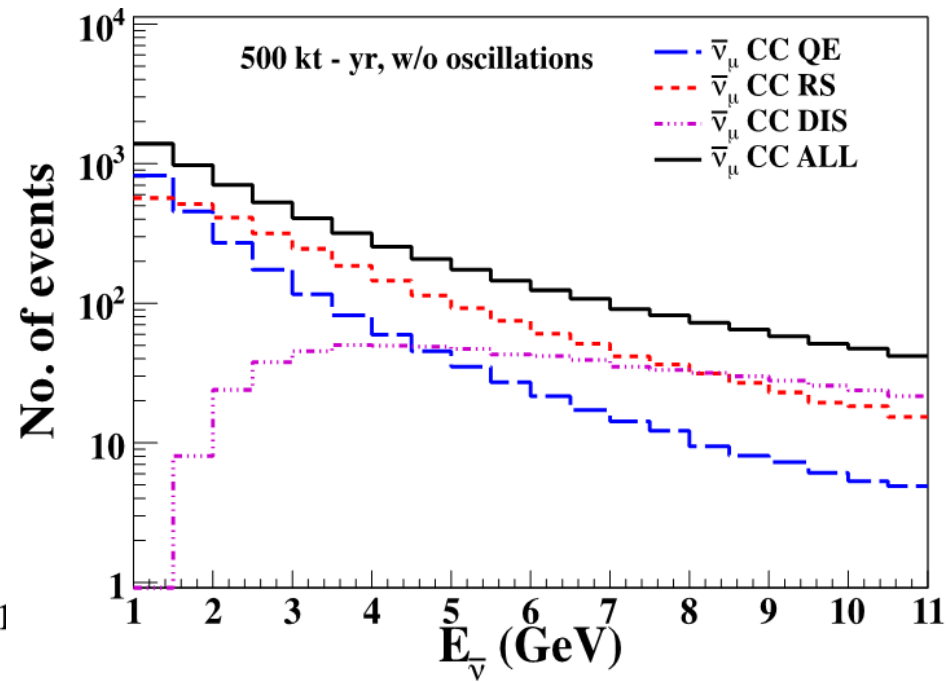
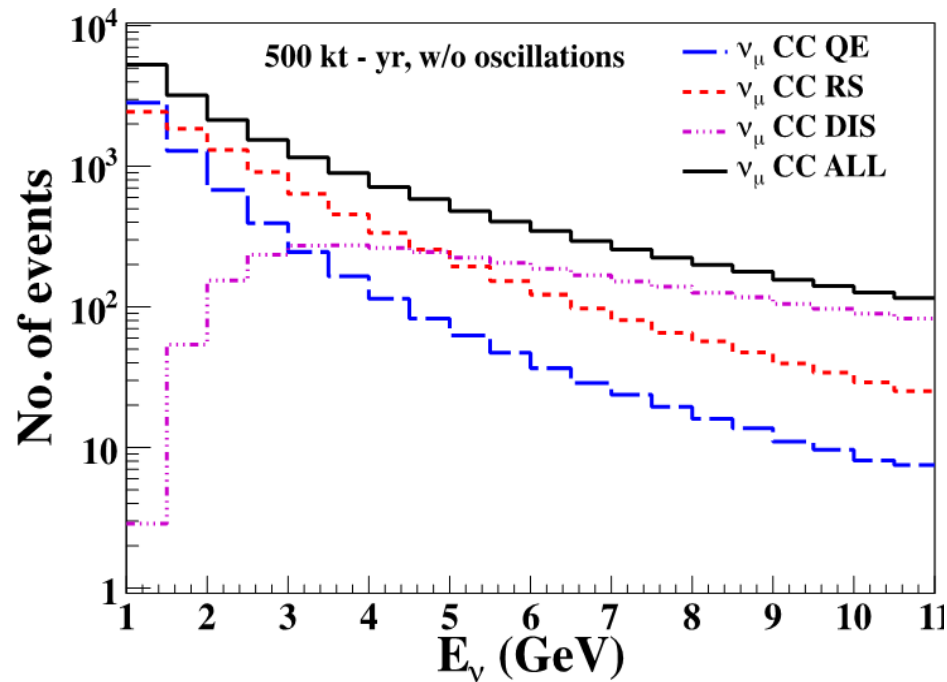


Thank you!

Atmospheric Flux



Averaged over all directions
Summed over all flavors of neutrino and anti-neutrino



Relative contributions of three cross-section processes to the total events in the absence of oscillation and without detector efficiency and resolutions