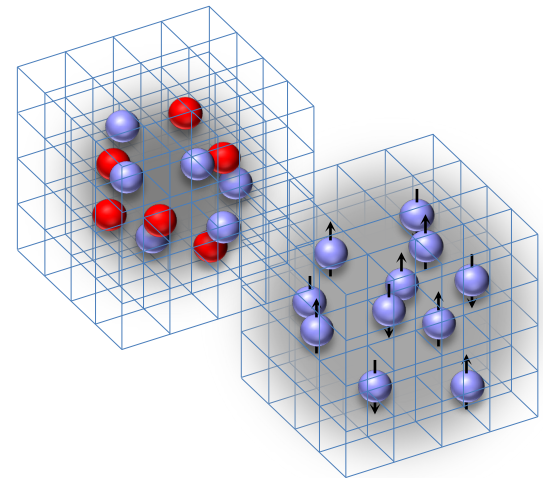


# *Ab initio* nuclear structure theory and connections to heavy-ion collisions

Dean Lee  
Facility for Rare Isotope Beams  
Michigan State University  
Nuclear Lattice EFT Collaboration

RBRC Workshop:  
Physics Opportunities from the RHIC Isobar Run  
January 25 – 28, 2022



# Outline

*Ab initio* methods

Lattice effective field theory

Hidden spin-isospin exchange symmetry

Essential elements for nuclear binding

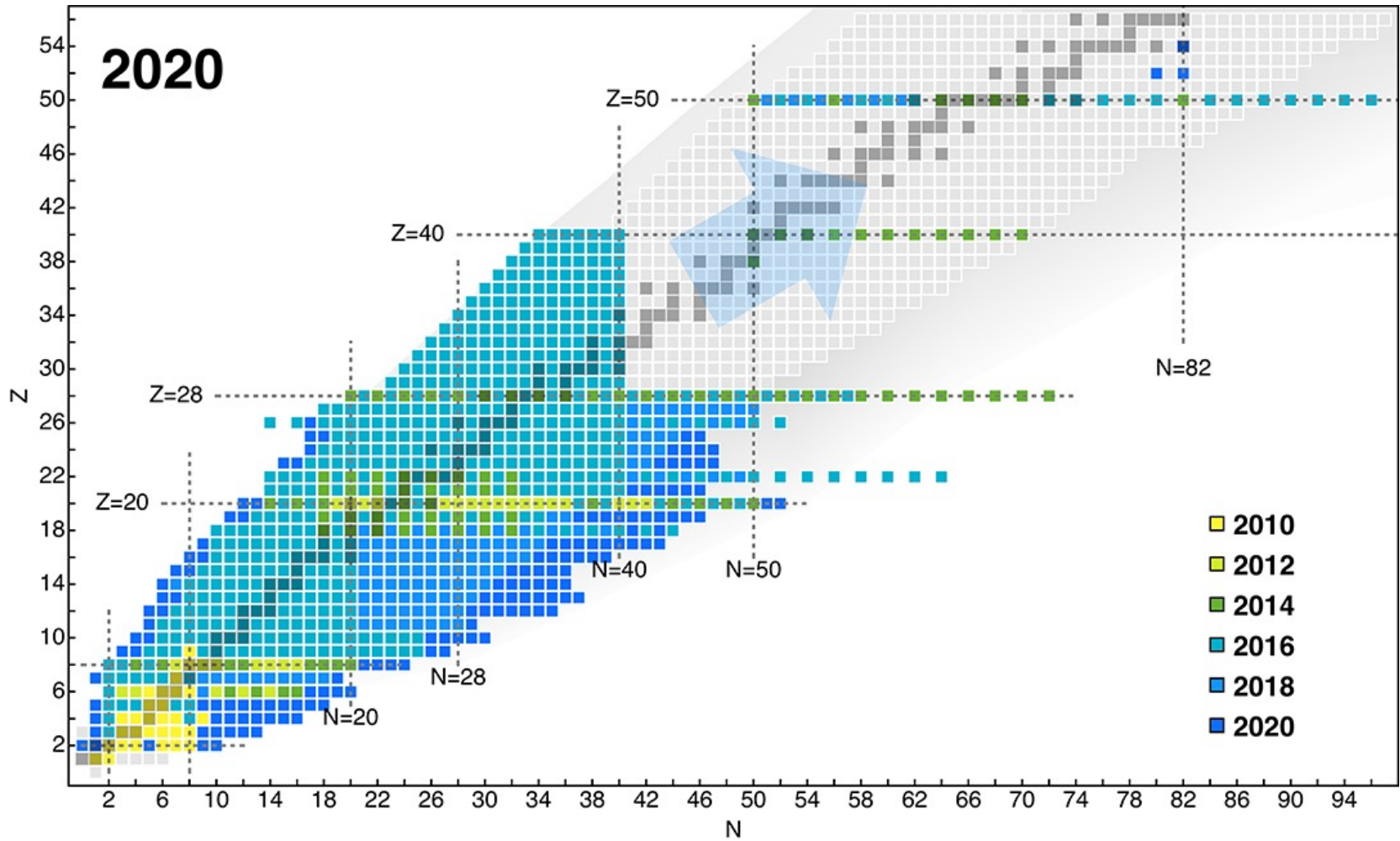
$^{16}\text{O}$   $^{16}\text{O}$  collisions at RHIC and LHC energies

Nuclear thermodynamics

Structure and spectrum of  $^{12}\text{C}$

Outlook

# *Ab initio* methods



[Hergert, Front. Phys. 8, 379 (2020)]

# Many-Body Perturbation Theory

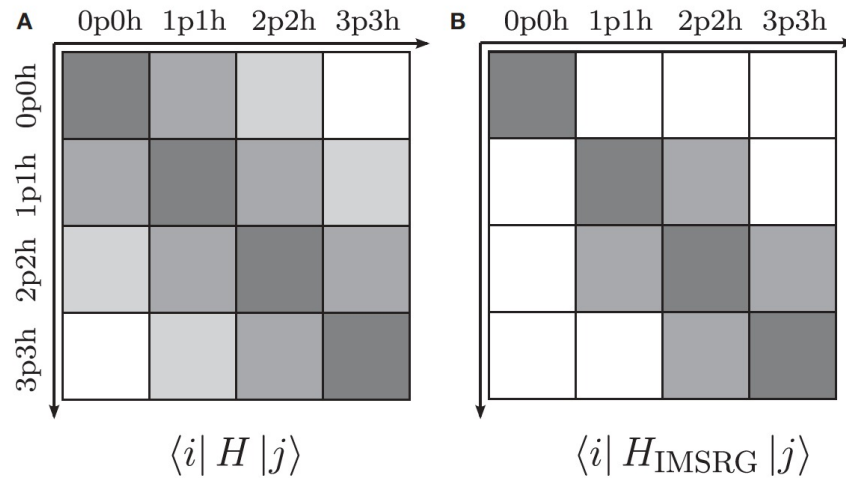
[Roth et al., Phys. Lett. B683, 272 (2010); Tichai et al., Phys. B756, 283 (2016)]

$$|\Psi\rangle = |\Phi\rangle + \sum_{n=1}^{\infty} \left( \frac{1}{H_0 - E^{(0)}} H_I \right)^n |\Phi\rangle$$

$$E = E^{(0)} + \langle \Phi | \sum_{n=1}^{\infty} H_I \left( \frac{1}{H_0 - E^{(0)}} H_I \right)^n | \Phi \rangle$$

# In-Medium Similarity Renormalization Group

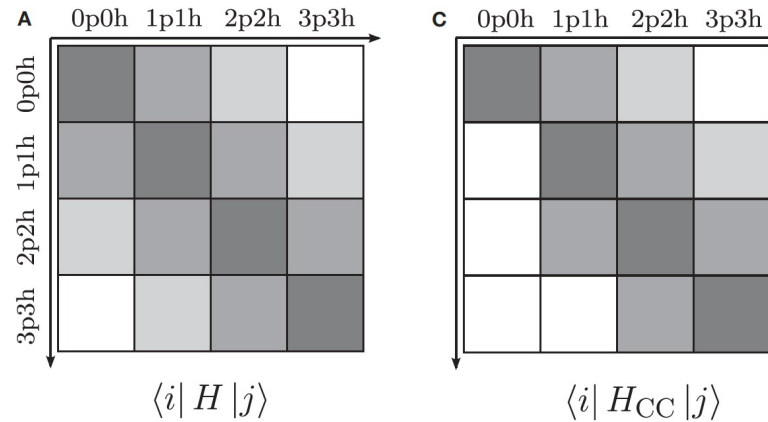
[Hergert et al., Phys. Rev. 621, 165 (2016); Stroberg et al., Phys. Rev. Lett. 118, 032502 (2017)]



[Hergert, Front. Phys. 8, 379 (2020)]

## Coupled Cluster Methods

[Hagen et al., Rept. Prog. Phys. 77, 096302 (2014); Duguet et al., Phys. Rev. C 91, 064320 (2015)]



## Self-Consistent Green's Functions

[Dickhoff et al., Prog. Part. Nucl. Phys. 52, 377 (2004), Soma et al. Phys. Rev. C 101, 014318 (2020)]

$$g_{pq\dots rs} = \langle \Psi_0^A | T [a_p(t_p) a_q(t_q) \cdots a_s^\dagger(t_s) a_r^\dagger(t_r)] | \Psi_0^A \rangle$$

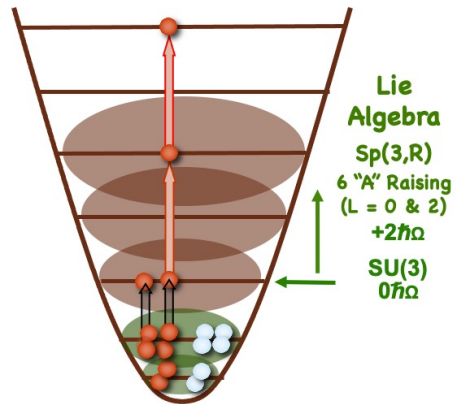
## No-Core Configuration Interaction

[Barrett et al., Prog. Part. Nucl. Phys. 69, 131 (2013); Navratil et al., Phys. Scripta. 91 053002, (2016)]

$$|\Psi\rangle = |\Psi\rangle_{\text{core}} \otimes |\Psi\rangle_{\text{valence}} \rightarrow |\Psi\rangle_{\text{all valence}}$$

## Symmetry-Adapted No-Core Configuration Interaction

[Launey et al., Prog. Part. Nucl. Phys. 89, 101 (2016); Dytrych et al. Phys. Rev. Lett. 124, 042501 (2020)]



[From Jerry Draayer's talk]

## Monte Carlo Shell Model

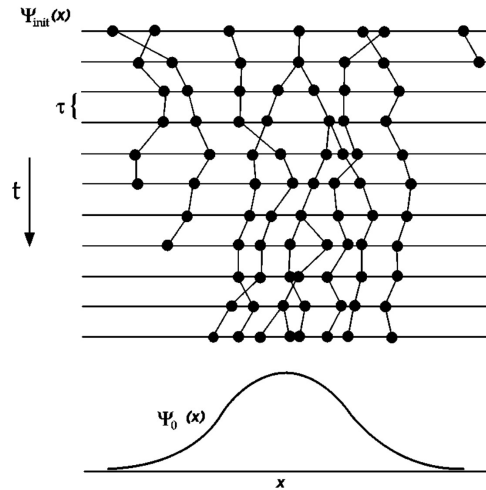
[Otsuka et al., Prog. Part. Nucl. Phys. 47, 319 (2001), Shimizu Phys. Scripta. 92, 063001 (2017)]

$$|\Phi_i(J, M, \pi)\rangle = \sum_{K=-J}^J g_K P_{M,K}^J P^\pi |\phi_i\rangle$$

$$|\Psi(J, M, \pi)\rangle = \sum_{i=1}^{N_{\text{basis}}} f_i |\Phi_i(J, M, \pi)\rangle$$

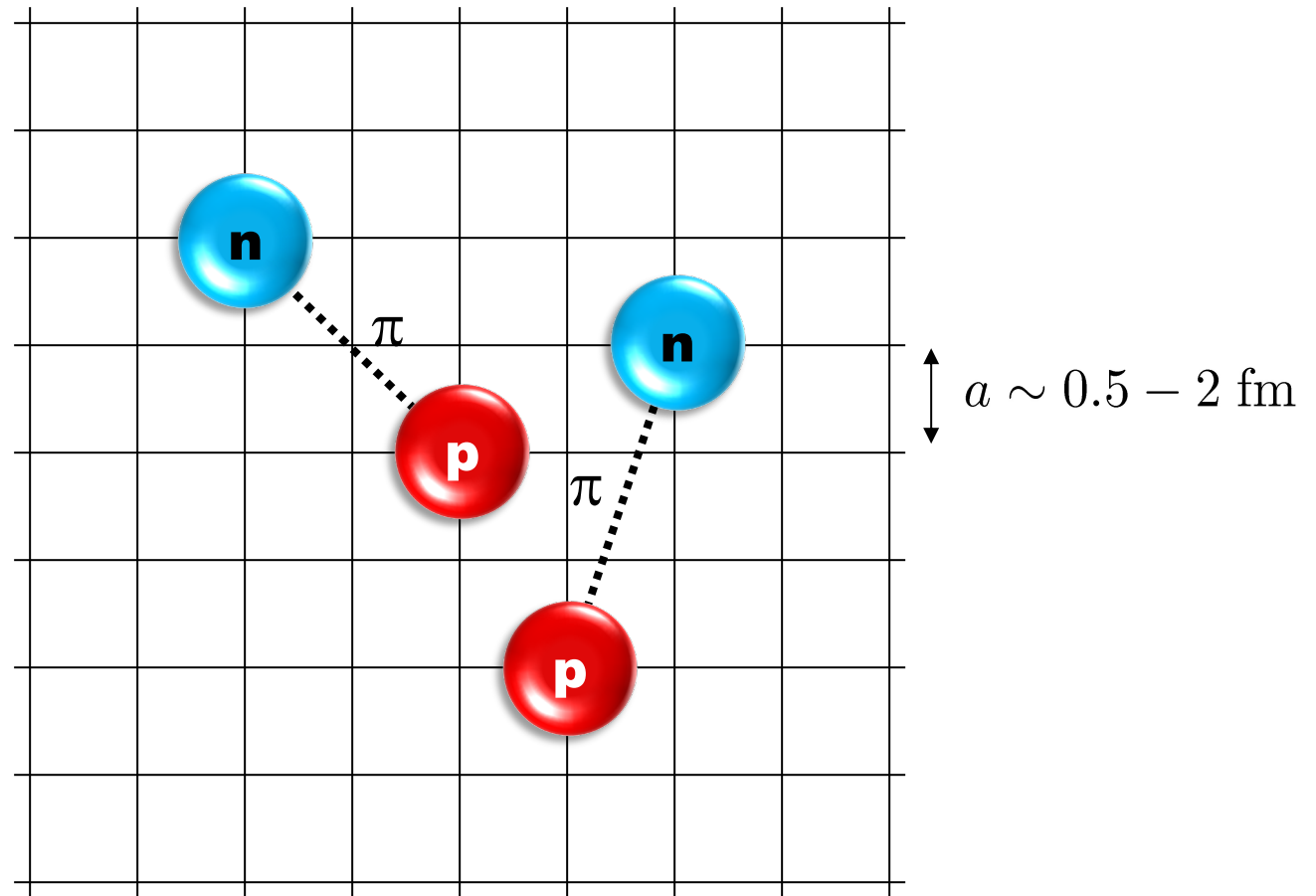
## Quantum Monte Carlo

[Carlson et al., Rev. Mod. Phys. 87, 1067 (2015); Gandolfi et al, Front. Phys. 8, 117 (2020)]



[Foulkes et al., Rev. Mod. Phys. 73, 1 (2001)]

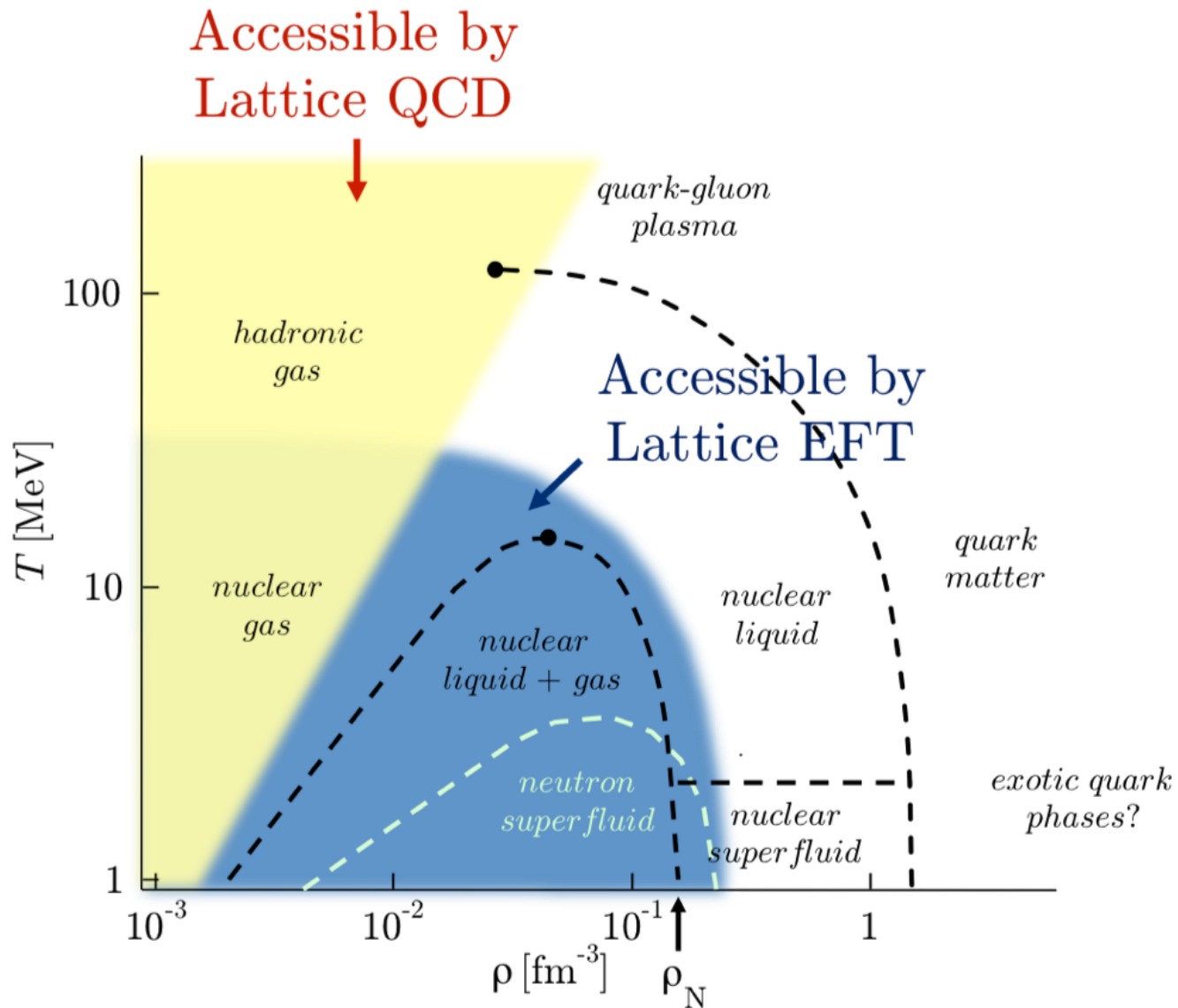
## Lattice effective field theory



[D.L, Prog. Part. Nucl. Phys. 63 117-154 (2009)]

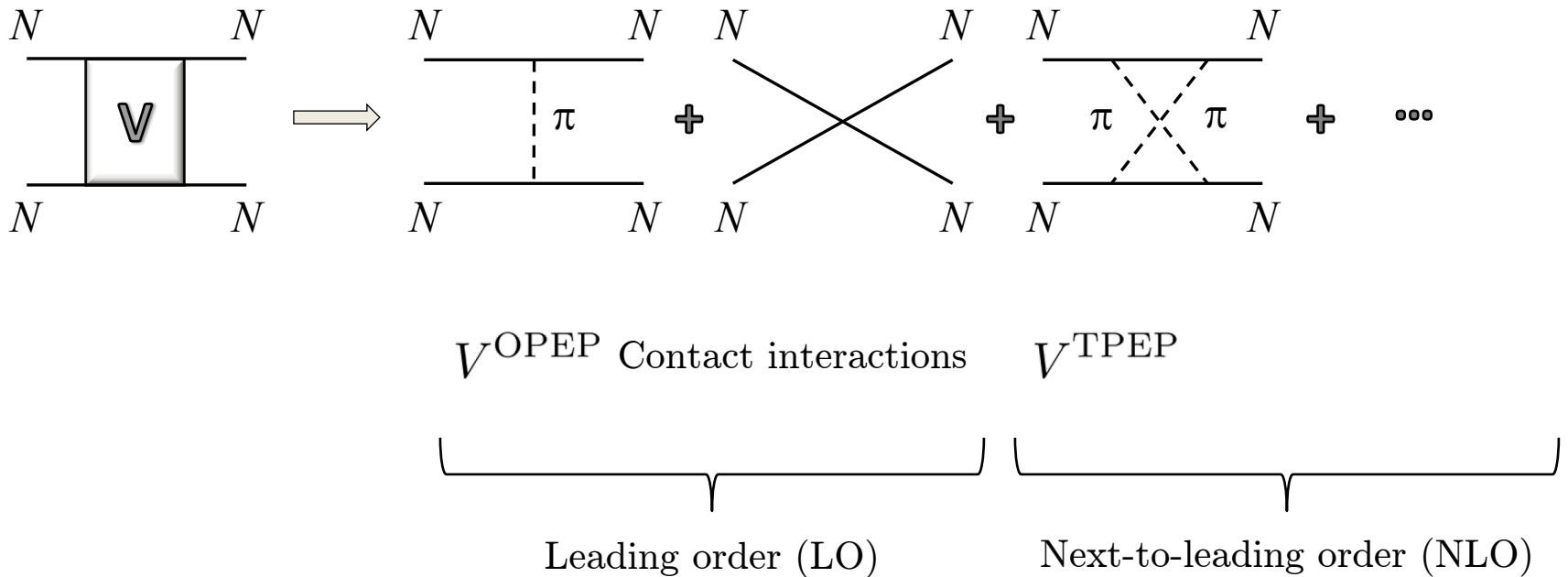
[Lähde, Meißner, Nuclear Lattice Effective Field Theory (2019), Springer]



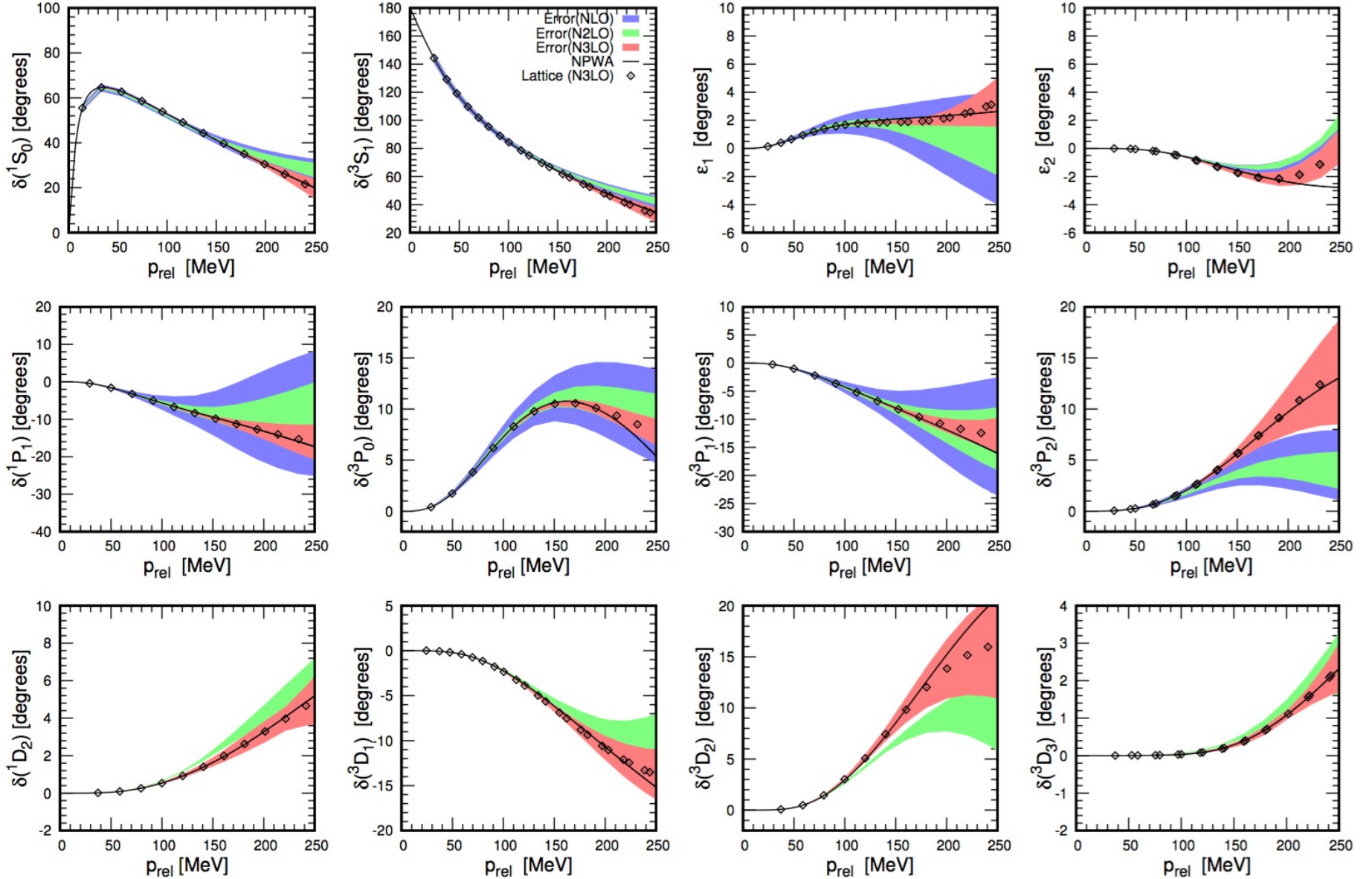


# Chiral effective field theory

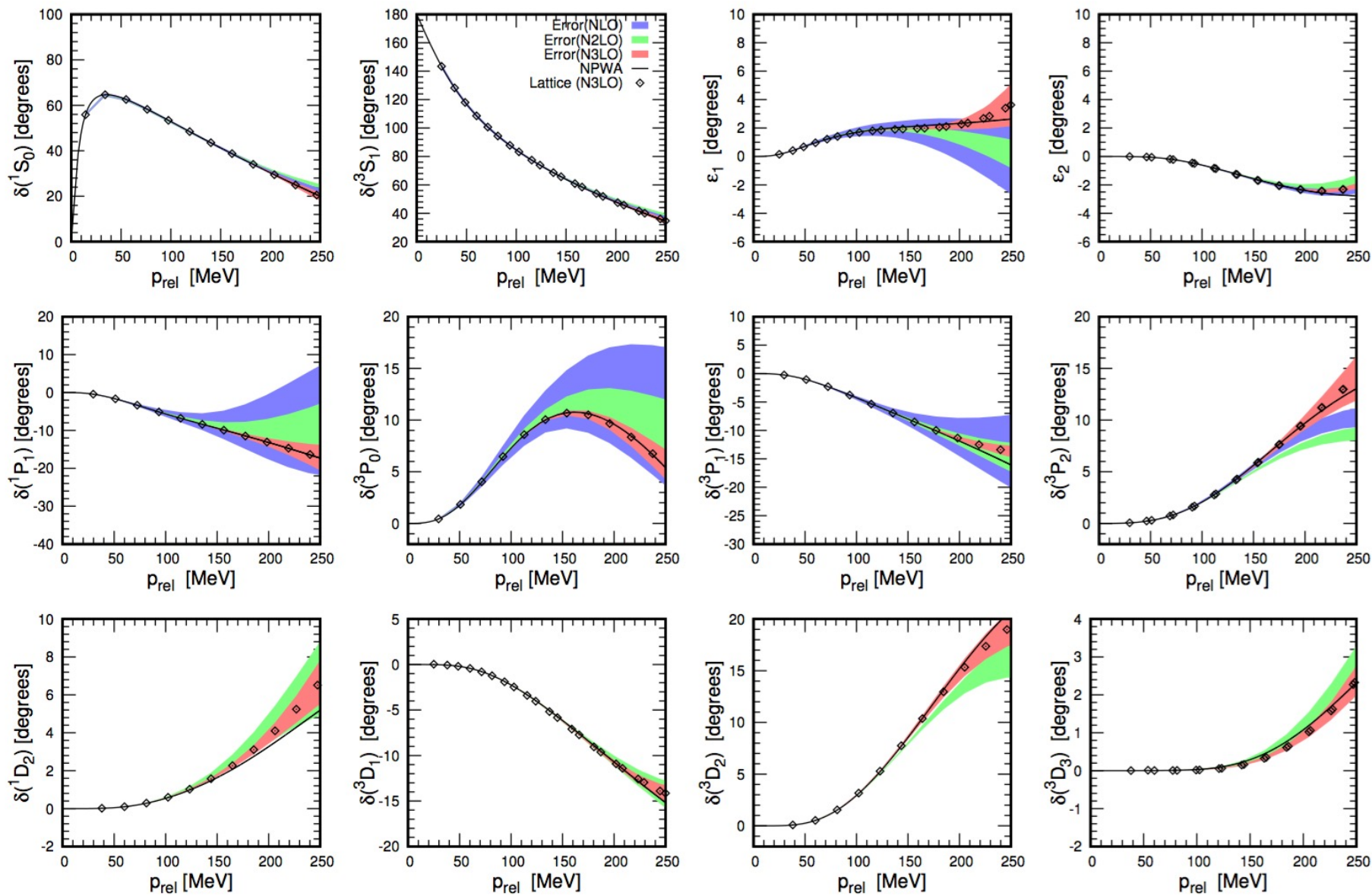
Construct the effective potential order by order



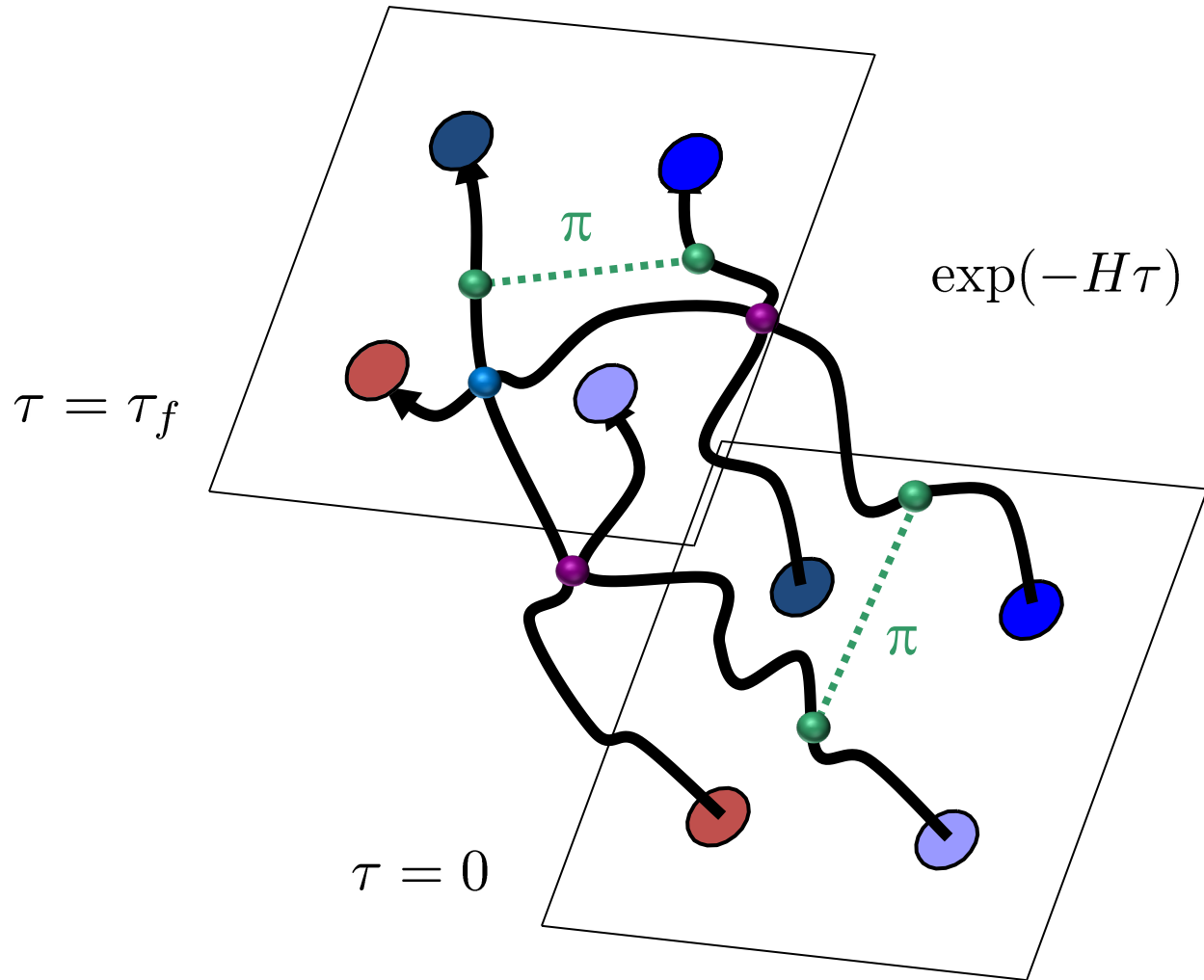
$a = 1.315$  fm



$a = 0.987 \text{ fm}$



# Euclidean time projection

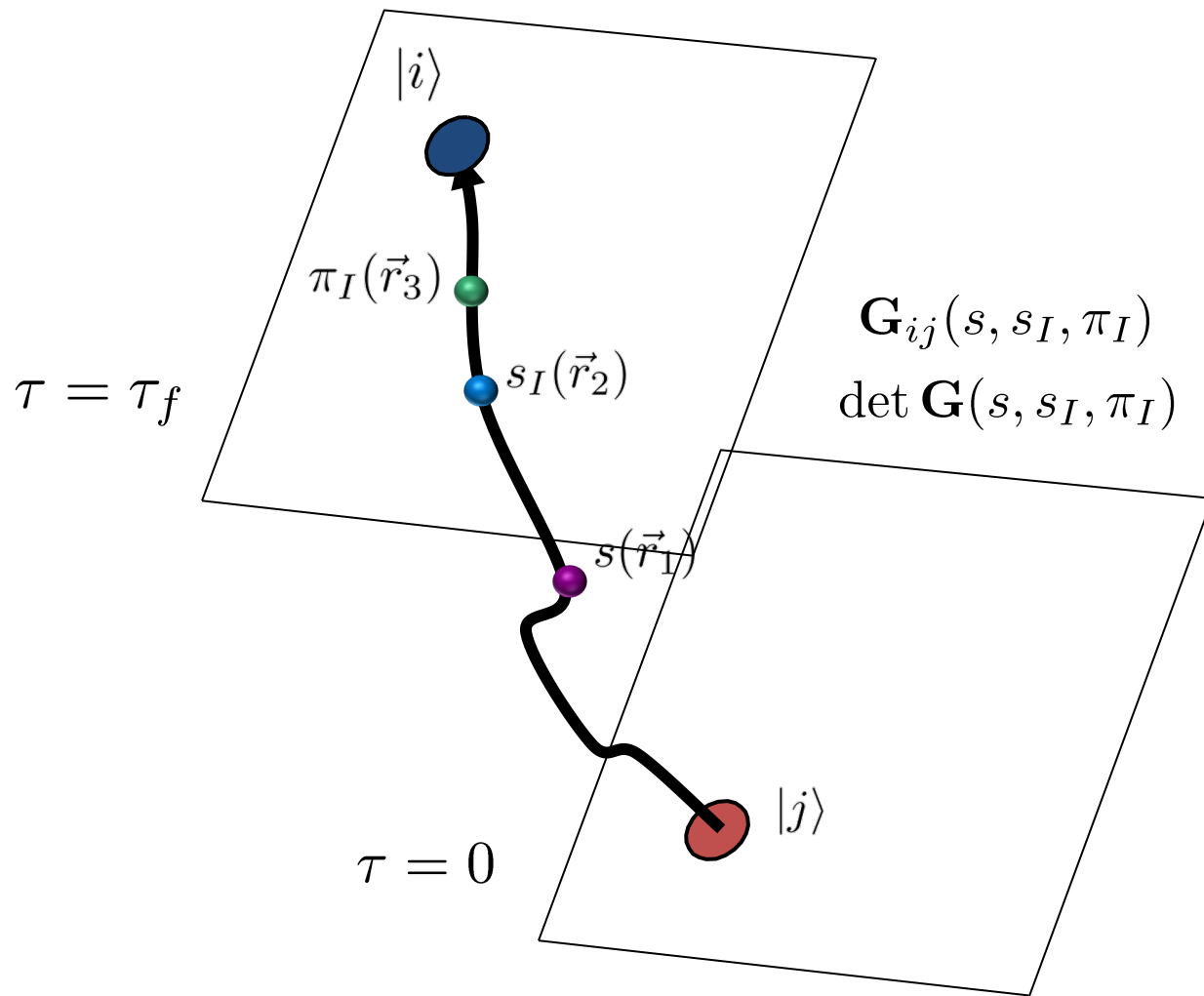


## Auxiliary field method

We can write exponentials of the interaction using a Gaussian integral identity

$$\begin{aligned} & \exp \left[ -\frac{C}{2} (N^\dagger N)^2 \right] \quad \diagdown \quad (N^\dagger N)^2 \\ & = \sqrt{\frac{1}{2\pi}} \int_{-\infty}^{\infty} ds \exp \left[ -\frac{1}{2} s^2 + \sqrt{-C} s (N^\dagger N) \right] \quad \diagup \quad s N^\dagger N \end{aligned}$$

We remove the interaction between nucleons and replace it with the interactions of each nucleon with a background field.



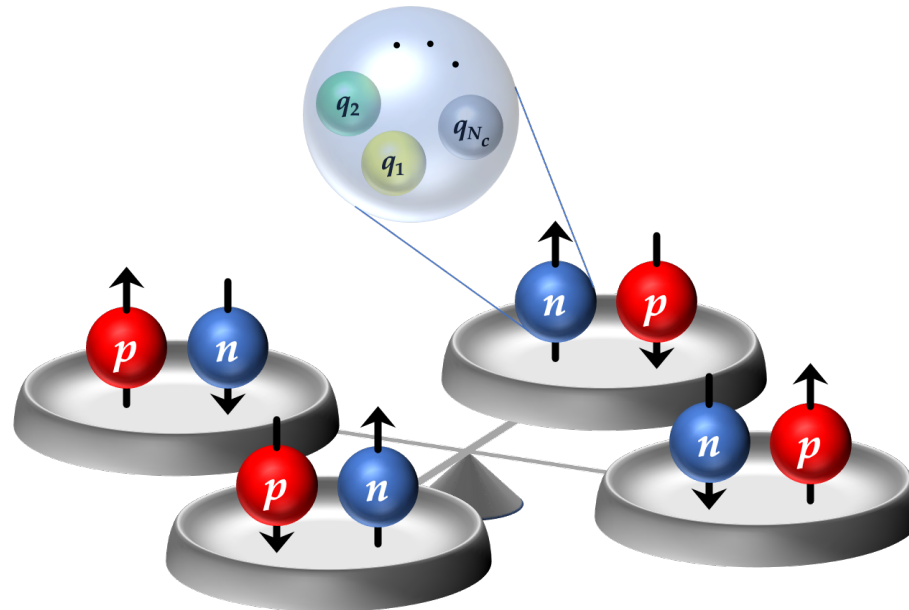
## Hidden spin-isospin exchange symmetry

Kaplan, Savage, PLB 365, 244 (1996)

Calle Gordon, Arriola, PRC 80, 014002 (2009)

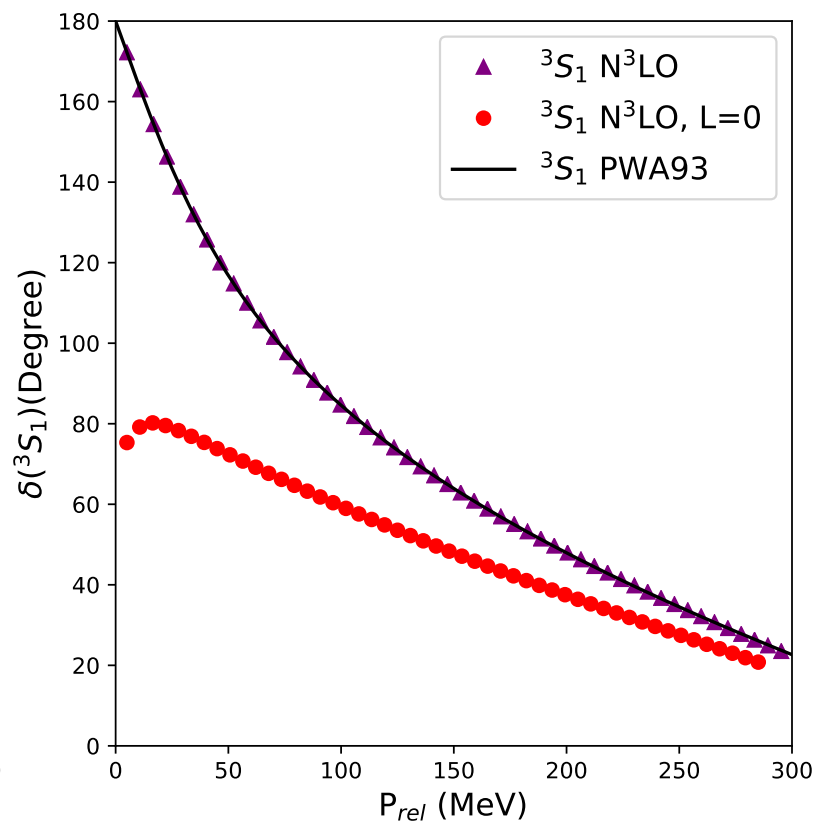
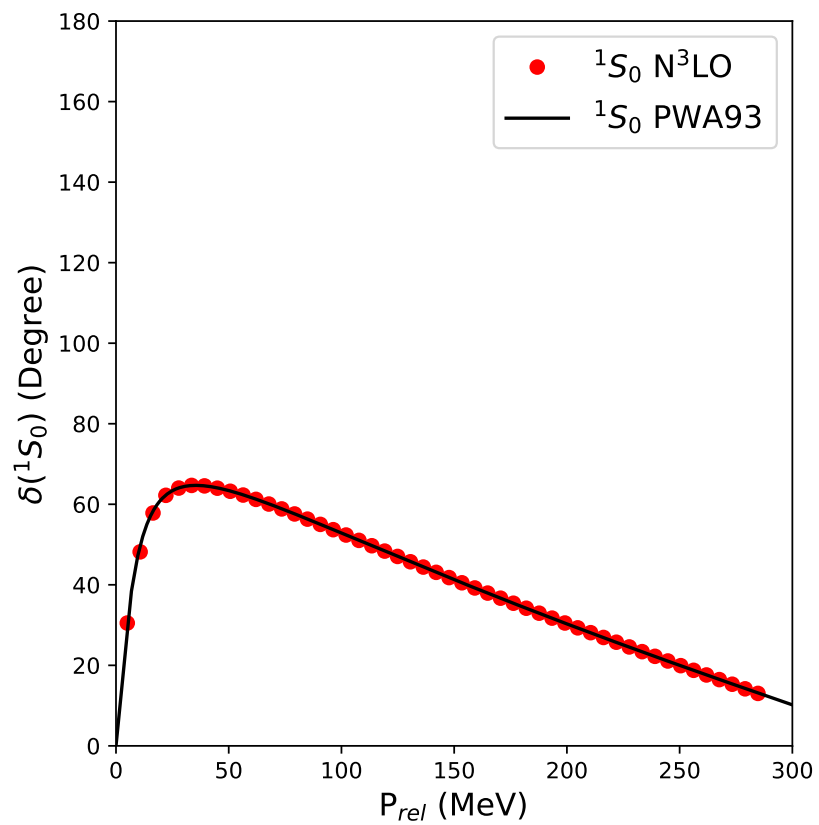
Kaplan, Manohar, PRC 56, 76 (1997)

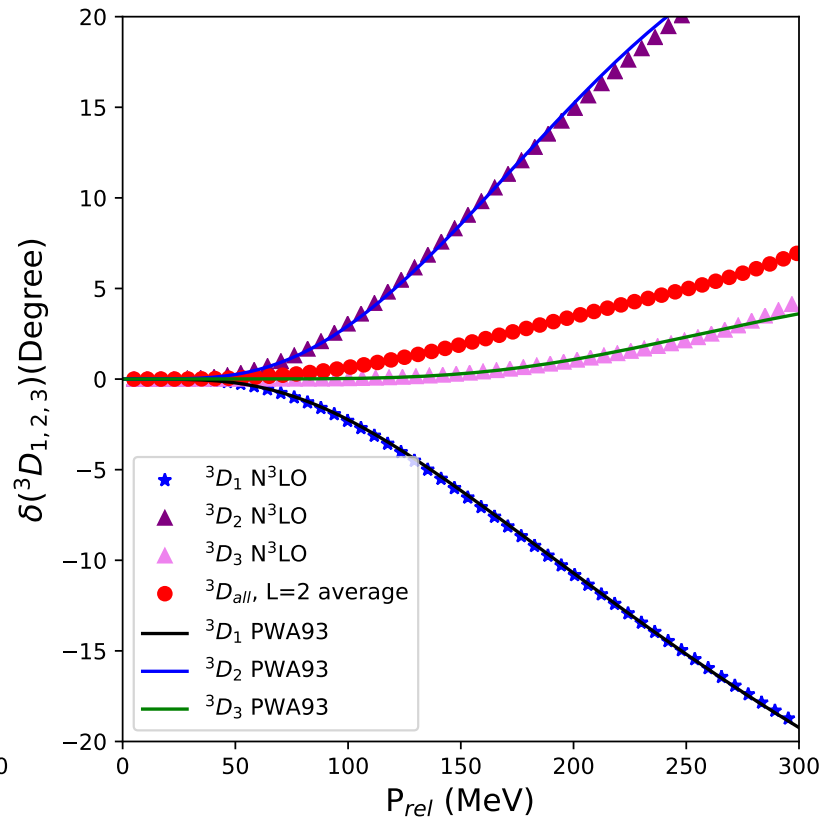
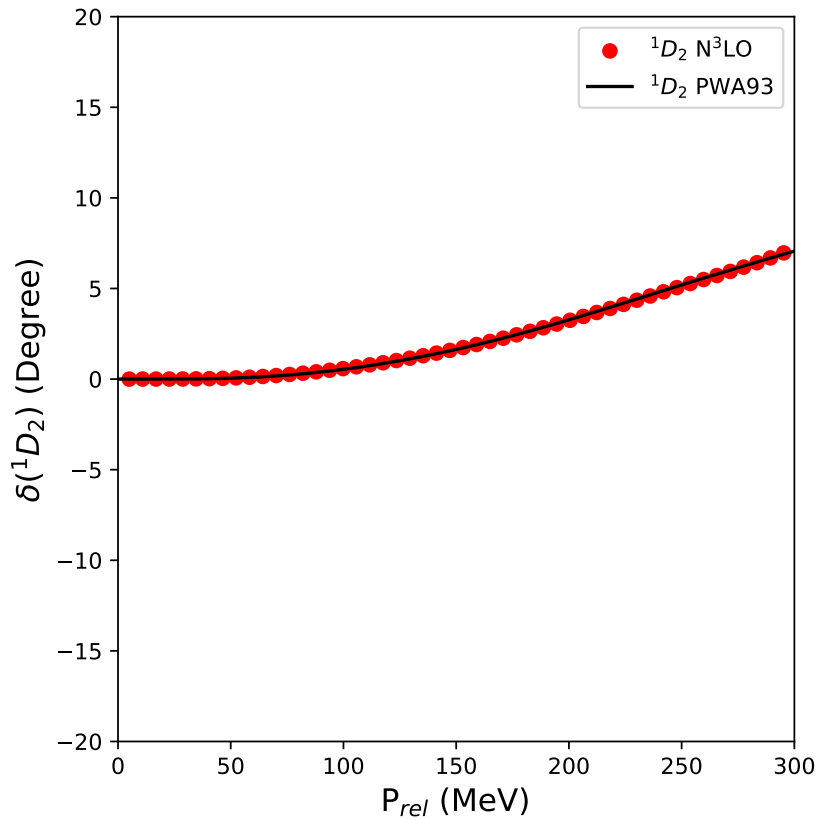
$$V_{\text{large-}N_c}^{2N} = V_C + \vec{\sigma}_1 \cdot \vec{\sigma}_2 \vec{\tau}_1 \cdot \vec{\tau}_2 W_S + (3\hat{r} \cdot \vec{\sigma}_1 \hat{r} \cdot \vec{\sigma}_2 - \vec{\sigma}_1 \cdot \vec{\sigma}_2) \vec{\tau}_1 \cdot \vec{\tau}_2 W_T$$



[D.L., Bogner, Brown, Elhatisari, Epelbaum, Hergert, Hjorth-Jensen, Krebs, Li, Lu, Meißner, PRL 127, 062501 (2021)]







[D.L., Bogner, Brown, Elhatisari, Epelbaum, Hergert, Hjorth-Jensen, Krebs, Li, Lu, Meißner, PRL 127, 062501 (2021)]

## Essential elements for nuclear binding

What is the minimal nuclear interaction that can reproduce the ground state properties of light nuclei, medium-mass nuclei, and neutron matter simultaneously with no more than a few percent error in the energies and charge radii?

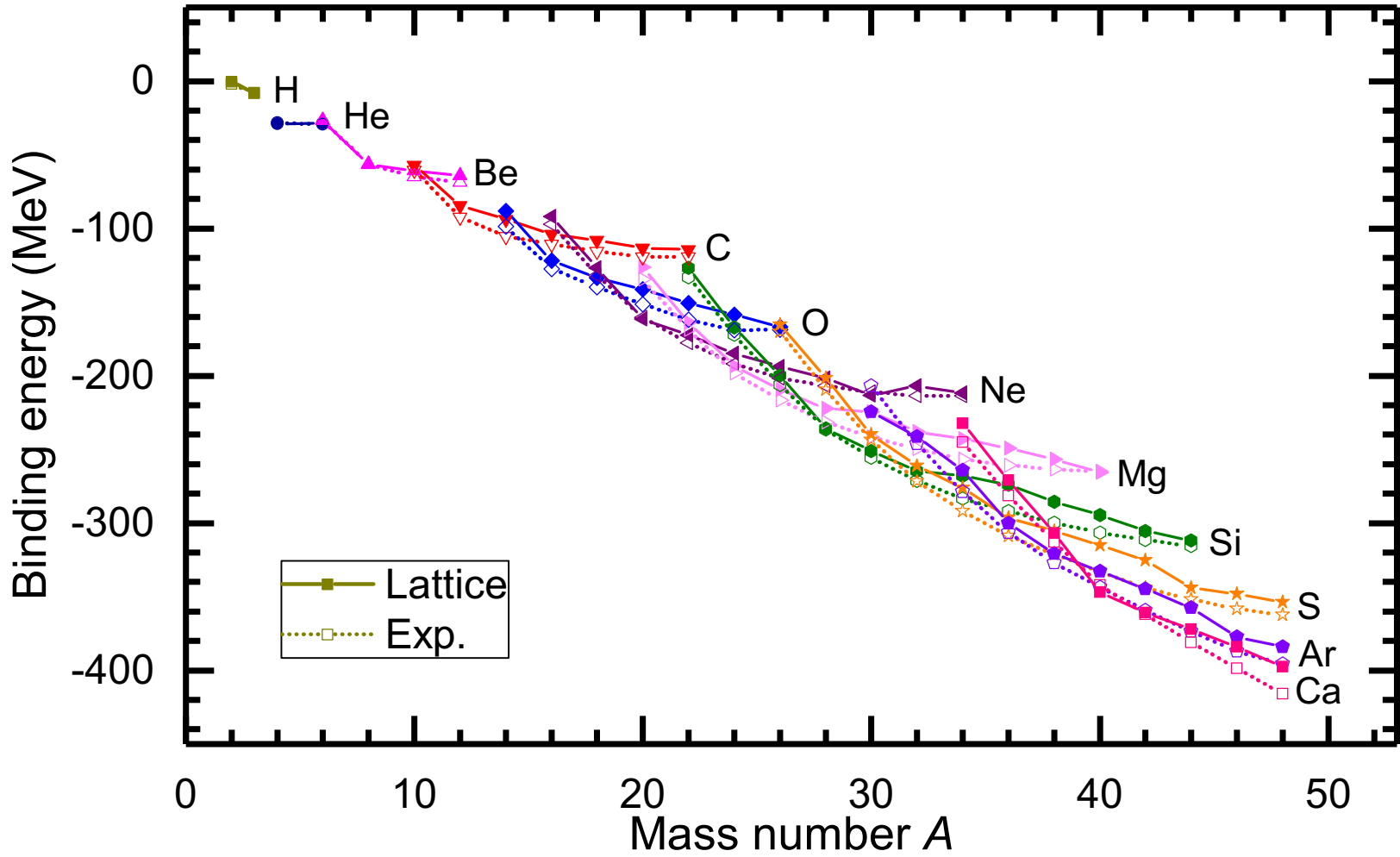
We construct an interaction with only four parameters.

1. Strength of the two-nucleon  $S$ -wave interaction
2. Range of the two-nucleon  $S$ -wave interaction
3. Strength of three-nucleon contact interaction

fit to  
 $A = 2, 3$  systems

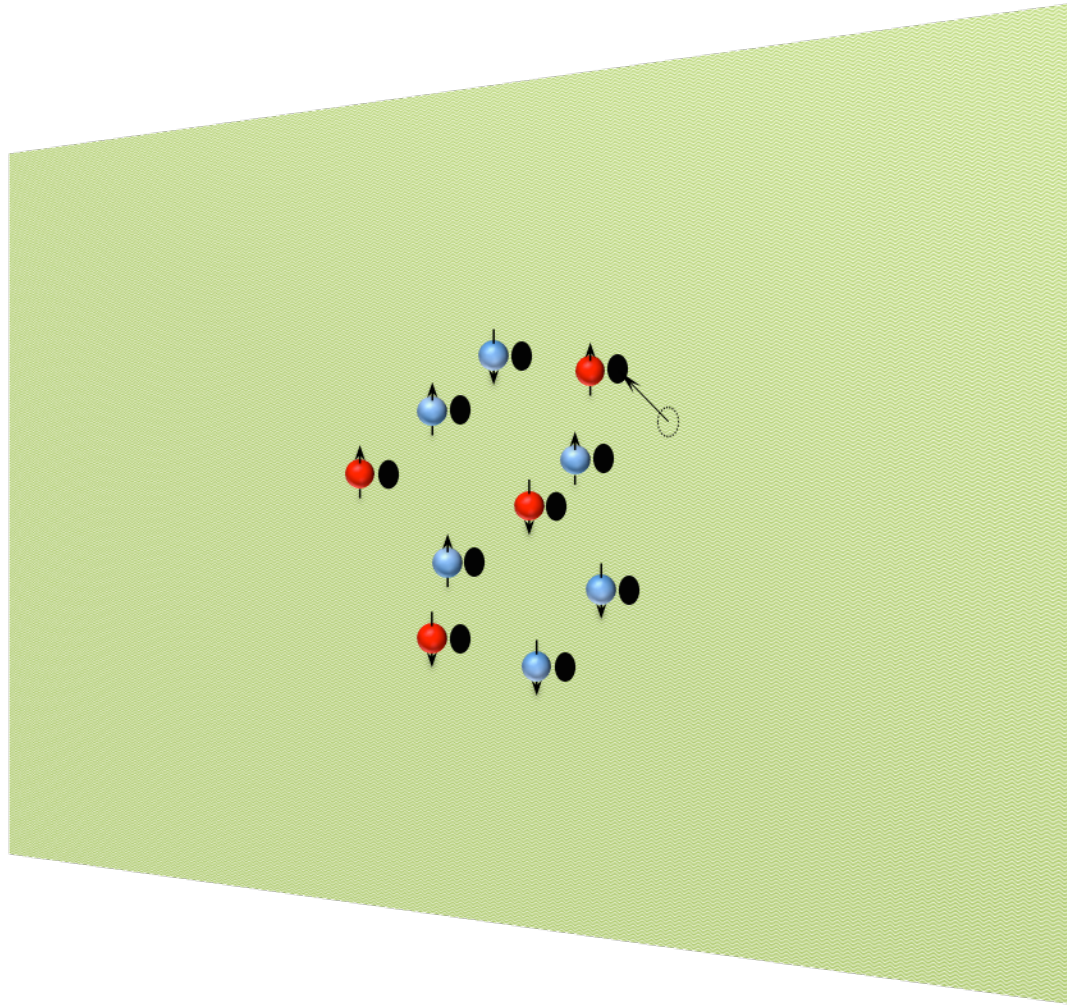
4. Range of the local part of the two-nucleon interaction

fit to  $A > 3$



	$B$	Exp.	$R_{\text{ch}}$	Exp.
${}^3\text{H}$	8.48(2)(0)	8.48	1.90(1)(1)	1.76
${}^3\text{He}$	7.75(2)(0)	7.72	1.99(1)(1)	1.97
${}^4\text{He}$	28.89(1)(1)	28.3	1.72(1)(3)	1.68
${}^{16}\text{O}$	121.9(1)(3)	127.6	2.74(1)(1)	2.70
${}^{20}\text{Ne}$	161.6(1)(1)	160.6	2.95(1)(1)	3.01
${}^{24}\text{Mg}$	193.5(02)(17)	198.3	3.13(1)(2)	3.06
${}^{28}\text{Si}$	235.8(04)(17)	236.5	3.26(1)(1)	3.12
${}^{40}\text{Ca}$	346.8(6)(5)	342.1	3.42(1)(3)	3.48

# Pinhole algorithm



## Seeing Structure with Pinholes

Consider the density operator for nucleon with spin  $i$  and isospin  $j$

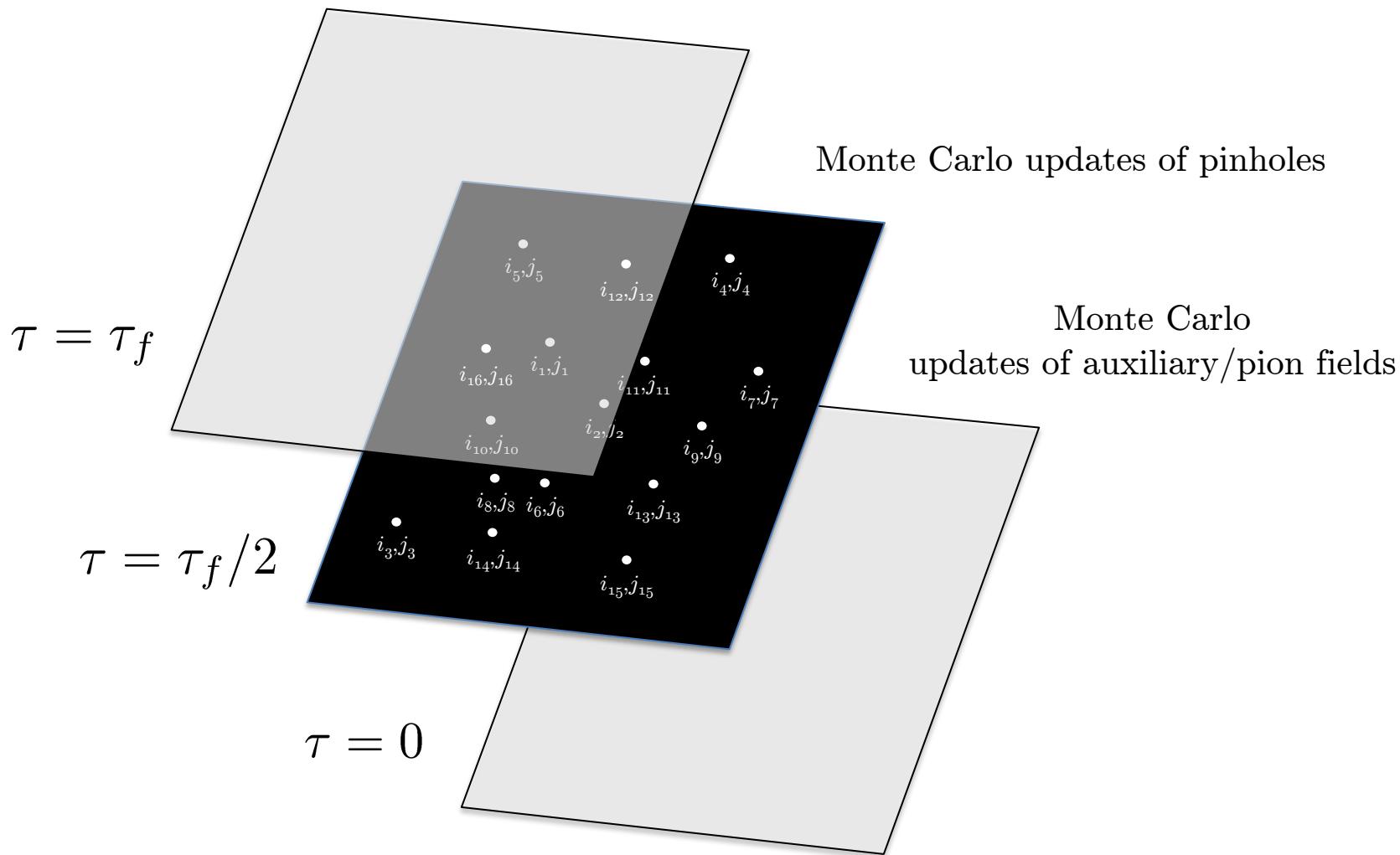
$$\rho_{i,j}(\mathbf{n}) = a_{i,j}^\dagger(\mathbf{n})a_{i,j}(\mathbf{n})$$

We construct the normal-ordered  $A$ -body density operator

$$\rho_{i_1,j_1,\dots,i_A,j_A}(\mathbf{n}_1,\dots,\mathbf{n}_A) = : \rho_{i_1,j_1}(\mathbf{n}_1) \cdots \rho_{i_A,j_A}(\mathbf{n}_A) :$$

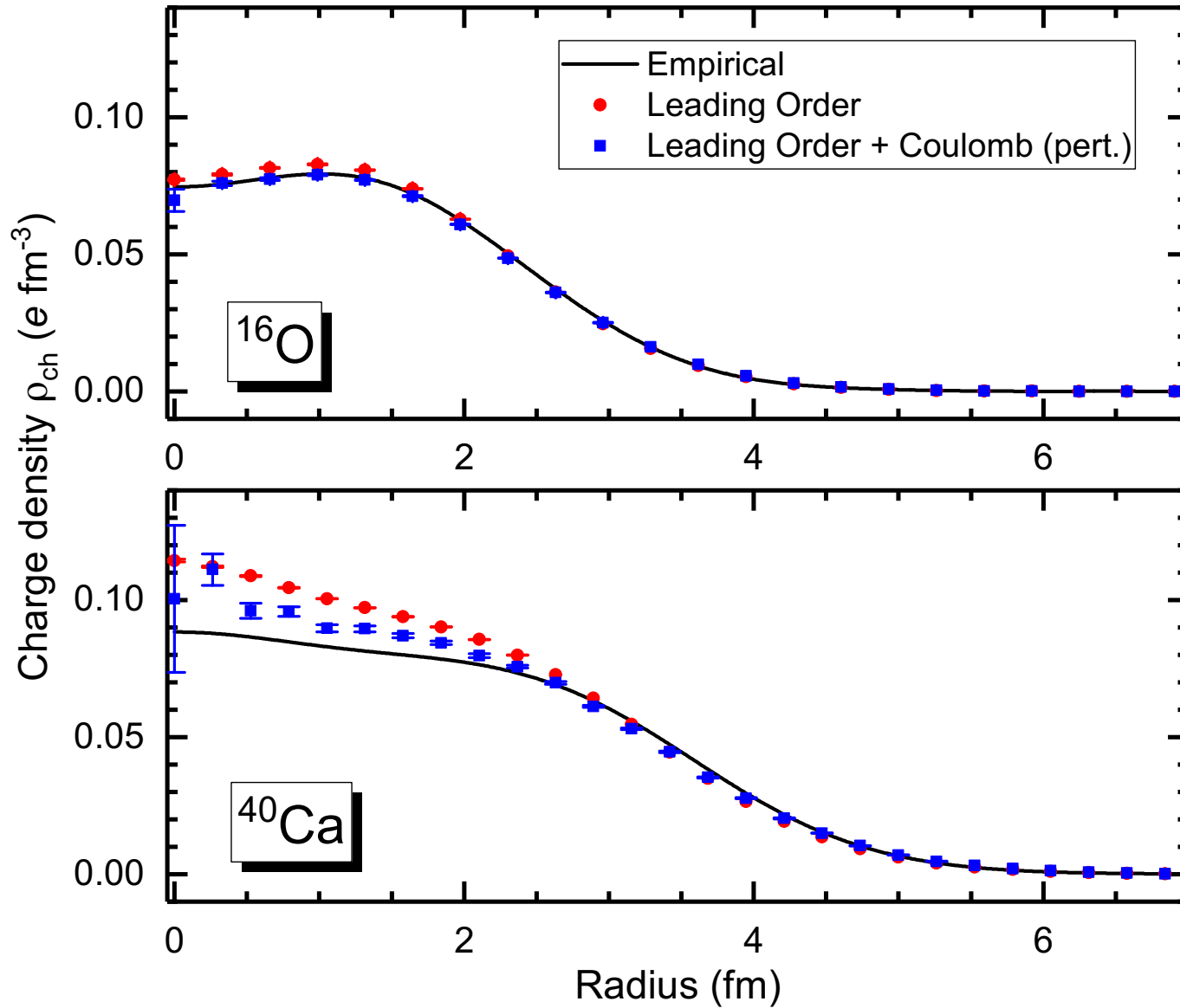
In the simulations we do Monte Carlo sampling of the amplitude

$$A_{i_1,j_1,\dots,i_A,j_A}(\mathbf{n}_1,\dots,\mathbf{n}_A,t) = \langle \Psi_I | e^{-Ht/2} \rho_{i_1,j_1,\dots,i_A,j_A}(\mathbf{n}_1,\dots,\mathbf{n}_A) e^{-Ht/2} | \Psi_I \rangle$$

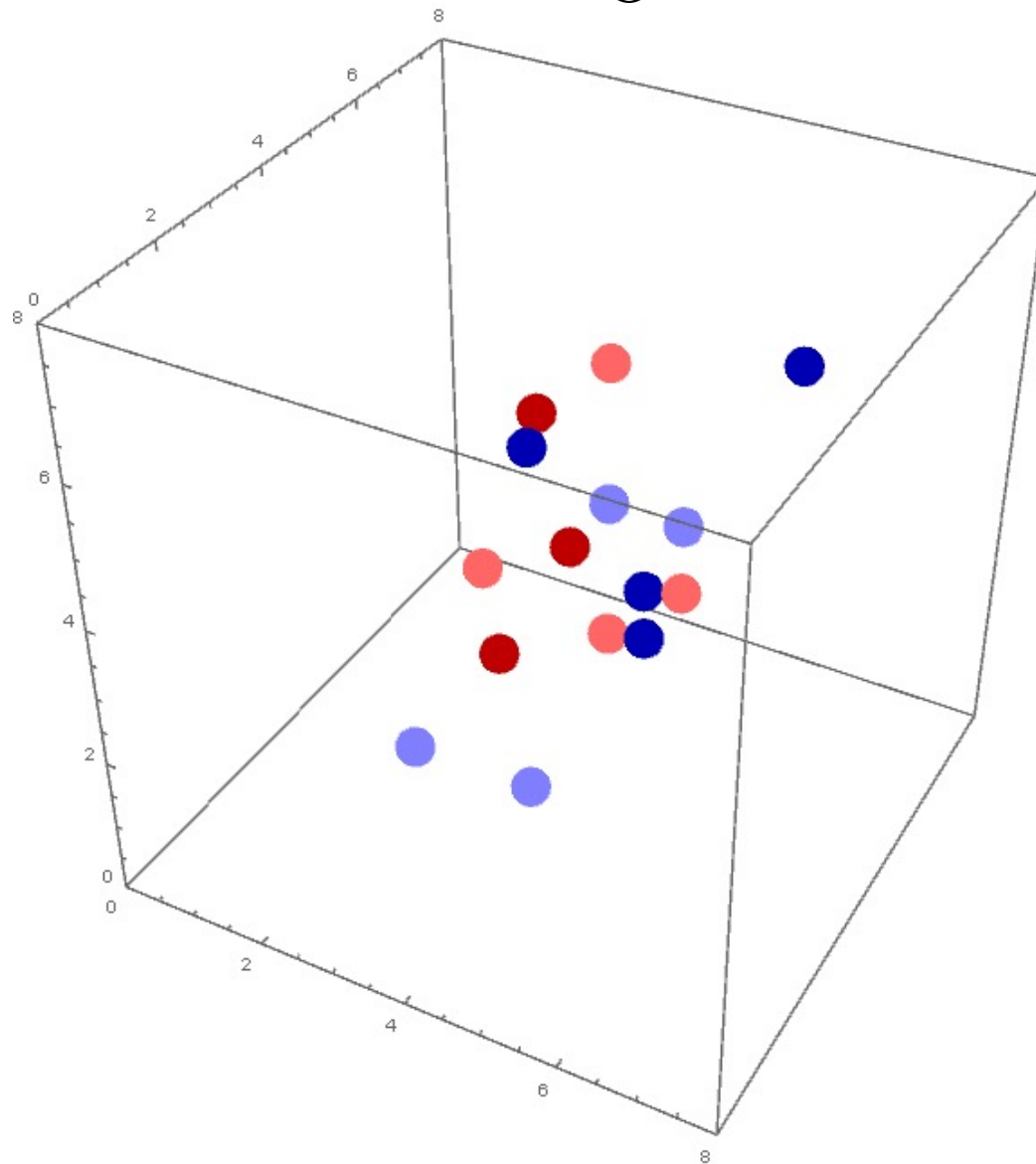


[Elhatisari, Epelbaum, Krebs, Lähde, D.L., Li, Lu, Meißner, Rupak, PRL 119, 222505 (2017)]





$^{16}\text{O}$



- proton up
- proton down
- neutron up
- neutron down

## $^{16}\text{O}$ $^{16}\text{O}$ collisions at RHIC and LHC energies

[Summerfield, Lu, Plumberg, D.L., Noronha-Hostler, Timmins PRC 104, L041901 (2021)]

$^{16}\text{O}$   $^{16}\text{O}$  collisions have been proposed at RHIC and LHC to probe dependence on initial states of intermediate size, where alpha clustering is expected to be significant.

For  $^{16}\text{O}$   $^{16}\text{O}$  collisions, the system is quite small, and we are pushing the limits of hydrodynamics. We use the Duke Bayesian tune of iEBE-VISHNU package to  $p\text{Pb}$  and  $\text{PbPb}$  collisions at the LHC.

[Bernhard et al., Nat. Phys. 15 1113 (2019), Moreland et al., Phys. Rev. C 101, 024911 (2020)]

We use the  $\text{T}_{\text{R}}\text{ENTo}$  model to general the initial entropy distribution.

[Moreland et al., Phys. Rev. C 92, 044903 (2015)]

The initial entropy distribution is then passed through a free-streaming phase of duration  $0.37 \text{ fm}/c$  and used to initialize the hydrodynamics evolution.

We compute the following cumulants of the flow harmonics  $v_n$ :

$$v_n\{2\} = [\langle v_n^2 \rangle]^{\frac{1}{2}}$$
$$v_n\{4\} = \left[ 2 \langle v_n^2 \rangle^2 - \langle v_n^4 \rangle \right]^{\frac{1}{4}}$$

We first compute results taking the initial density as a Woods-Saxon potential with density, radius, and diffusivity fitted to empirical values. We then consider the same Woods-Saxon potential, taking into account the quark substructure of the nucleons. Lastly, we consider using the nucleon distribution from the lattice effective field theory calculations. This will incorporate correlations such as alpha clustering.

[Summerfield, Lu, Plumberg, D.L., Noronha-Hostler, Timmins PRC 104, L041901 (2021)]

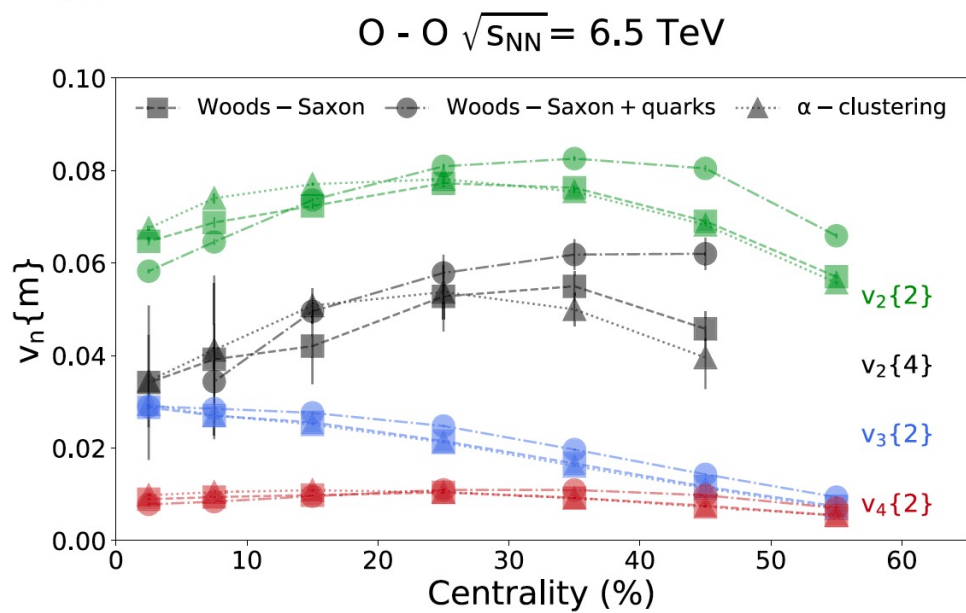
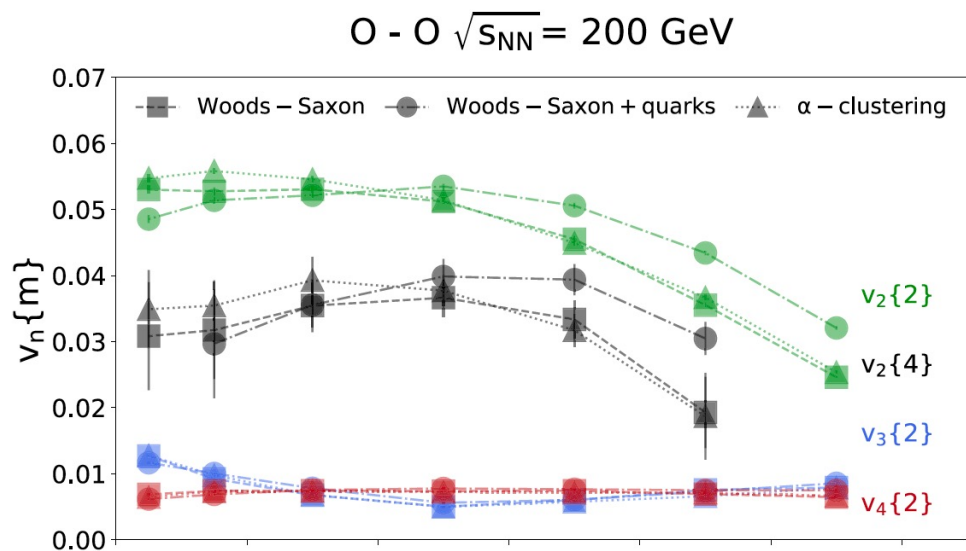
How much alpha clustering?

$$\langle \rho^2 \rangle_{16\text{O}} = 4.59(11) \langle \rho^2 \rangle_{4\text{He}}$$

$$\langle \rho^3 \rangle_{16\text{O}} = 4.67(23) \langle \rho^3 \rangle_{4\text{He}}$$

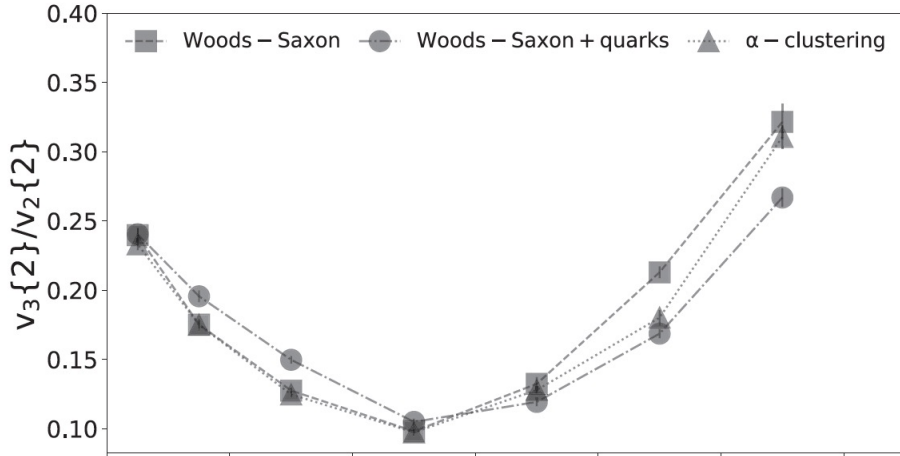
$$\langle \rho^4 \rangle_{16\text{O}} = 4.44(27) \langle \rho^4 \rangle_{4\text{He}}$$

About 10% greater than the simple tensor product of four alpha clusters. The excess is due to entanglement of the alpha cluster wave functions.

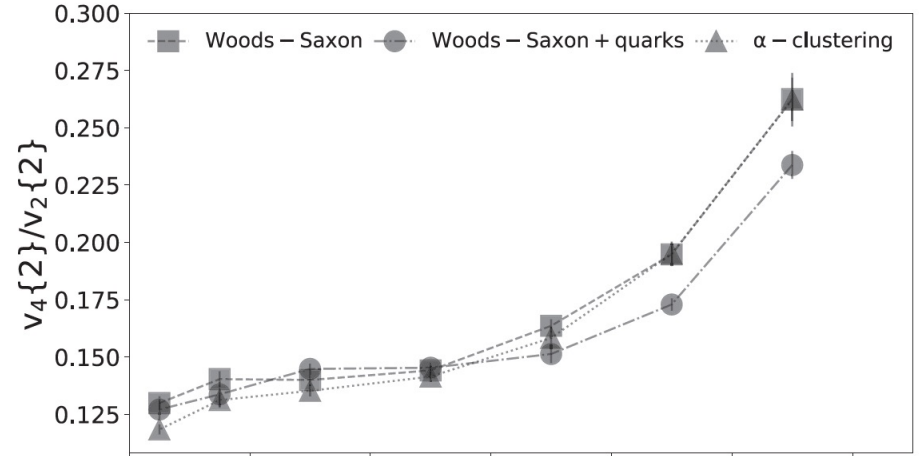


[Summerfield, Lu, Plumberg, D.L., Noronha-Hostler, Timmins PRC 104, L041901 (2021)]

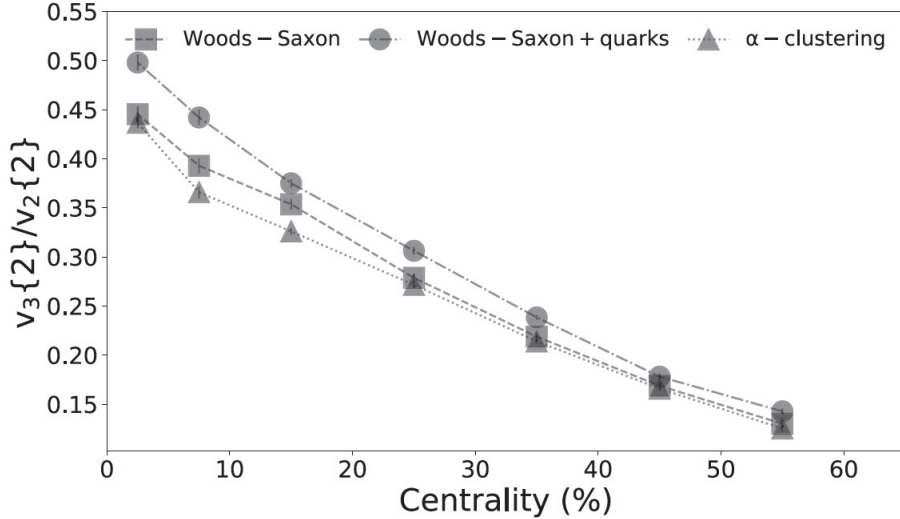
O - O  $\sqrt{s_{NN}} = 200$  GeV



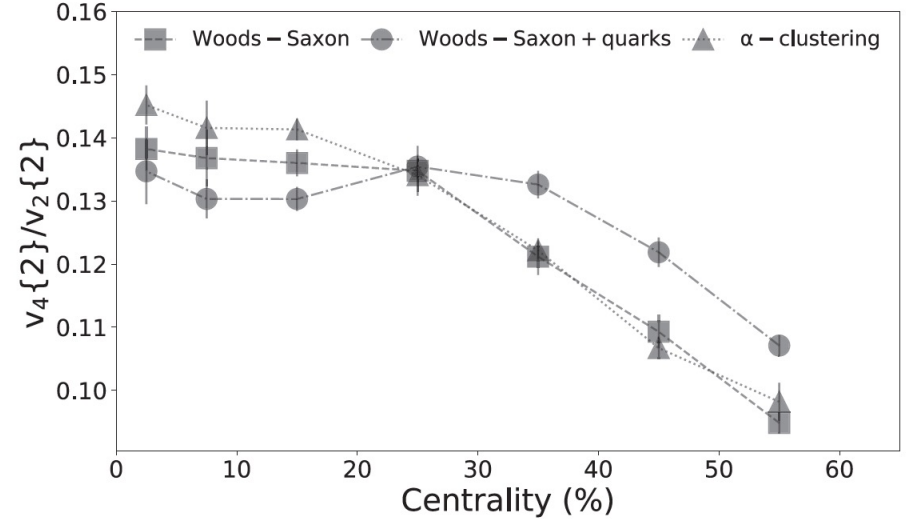
O - O  $\sqrt{s_{NN}} = 200$  GeV



O - O  $\sqrt{s_{NN}} = 6.5$  TeV

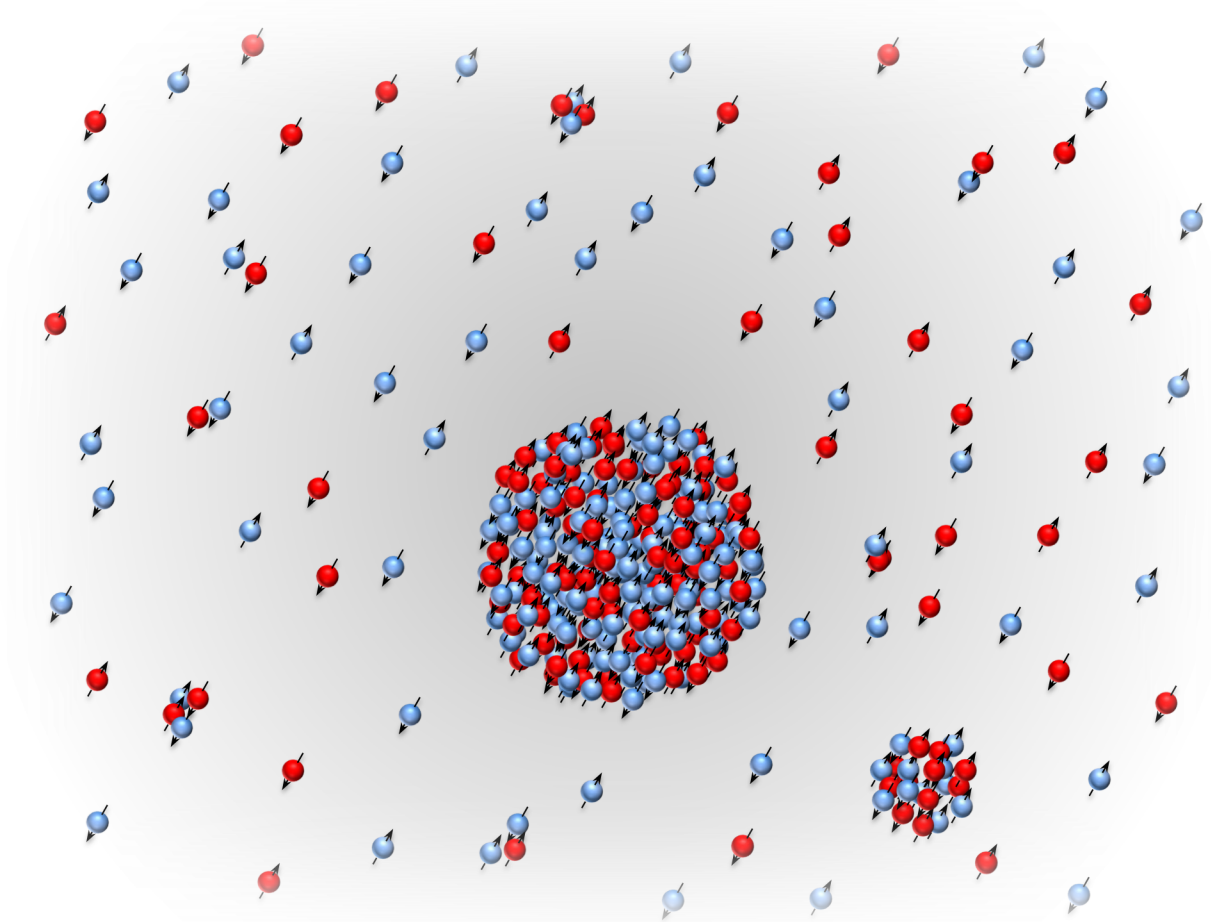


O - O  $\sqrt{s_{NN}} = 6.5$  TeV



[Summerfield, Lu, Plumberg, D.L., Noronha-Hostler, Timmins PRC 104, L041901 (2021)]

# Ab initio nuclear thermodynamics





## *Ab initio* nuclear thermodynamics

In order to compute thermodynamic properties of finite nuclei, nuclear matter, and neutron matter, we need to compute the partition function

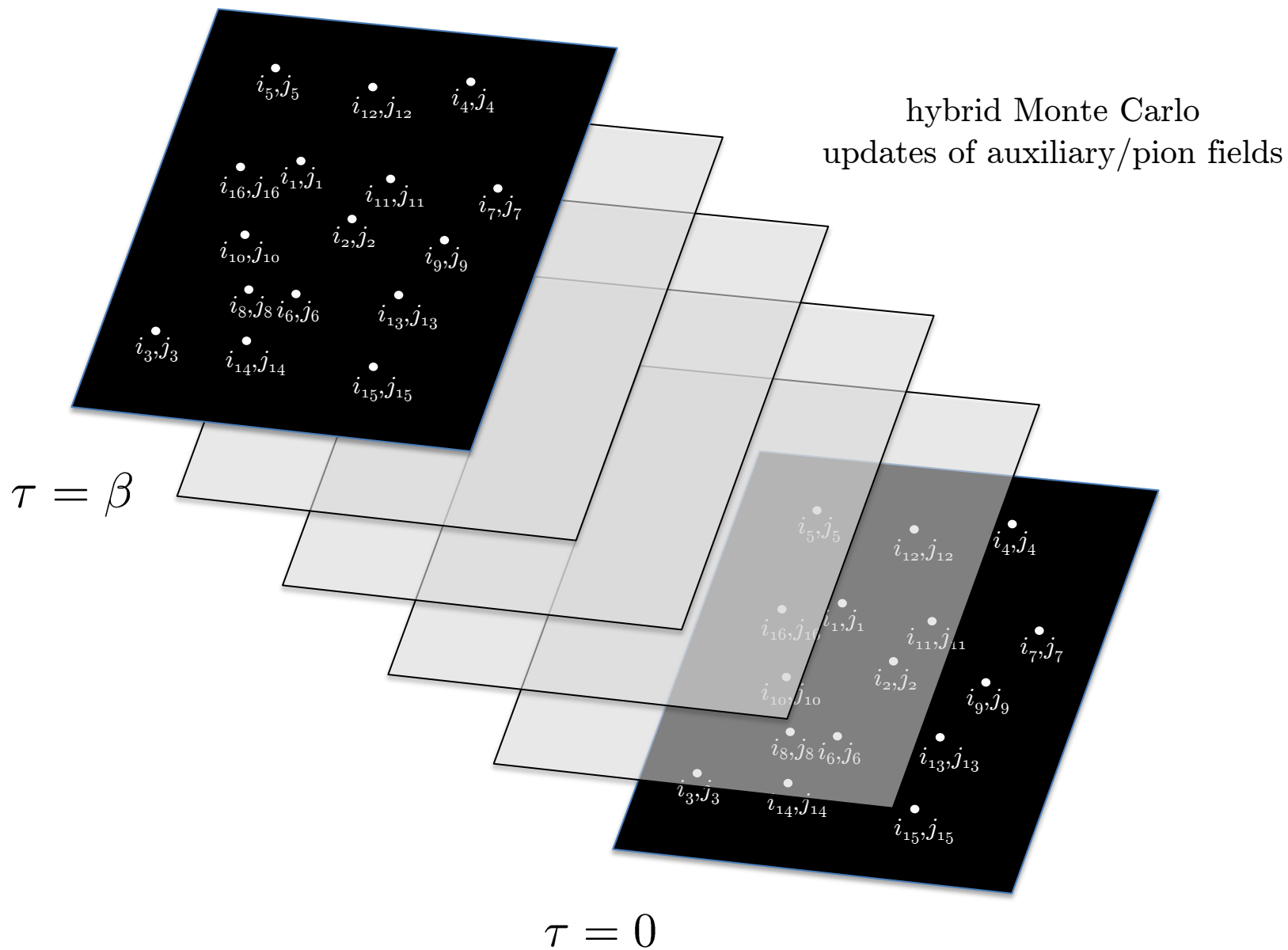
$$\text{Tr} \exp(-\beta H)$$

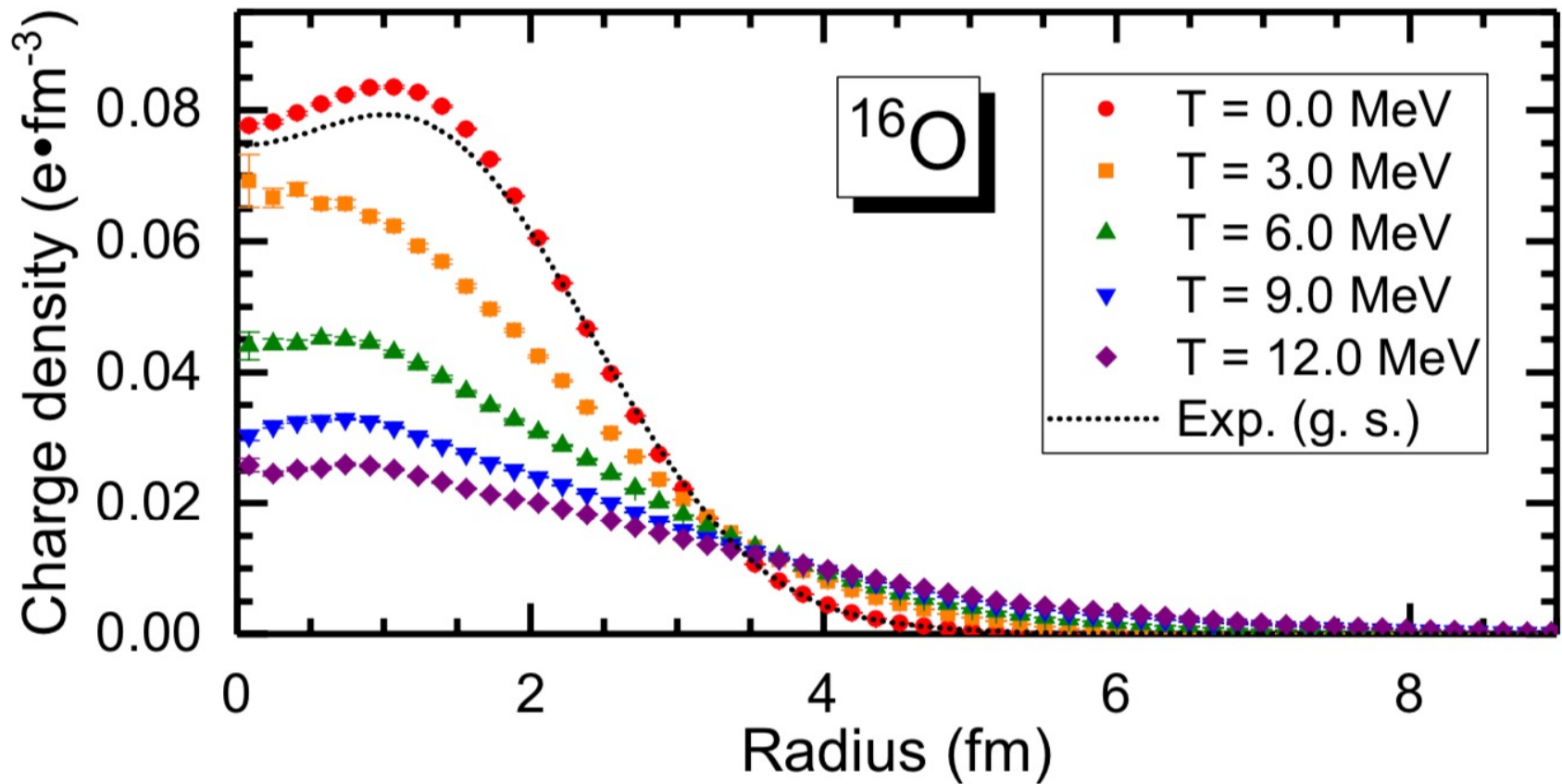
We compute the quantum mechanical trace over  $A$ -nucleon states by summing over pinholes (position eigenstates) for the initial and final states

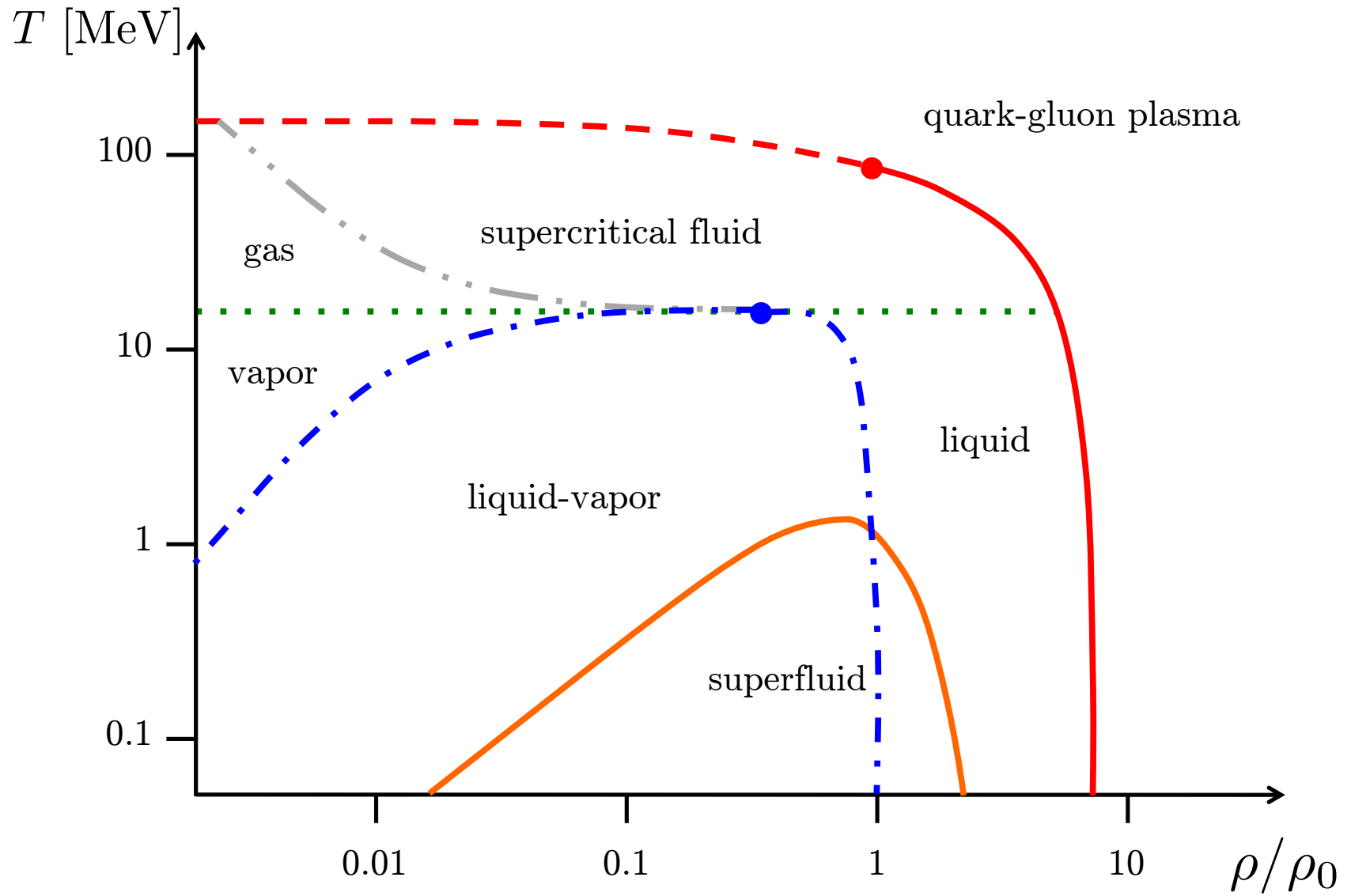
$$\begin{aligned} & \text{Tr} O \\ &= \frac{1}{A!} \sum_{i_1 \cdots i_A, j_1 \cdots j_A, \mathbf{n}_1 \cdots \mathbf{n}_A} \langle 0 | a_{i_A, j_A}(\mathbf{n}_A) \cdots a_{i_1, j_1}(\mathbf{n}_1) O a_{i_1, j_1}^\dagger(\mathbf{n}_1) \cdots a_{i_A, j_A}^\dagger(\mathbf{n}_A) | 0 \rangle \end{aligned}$$

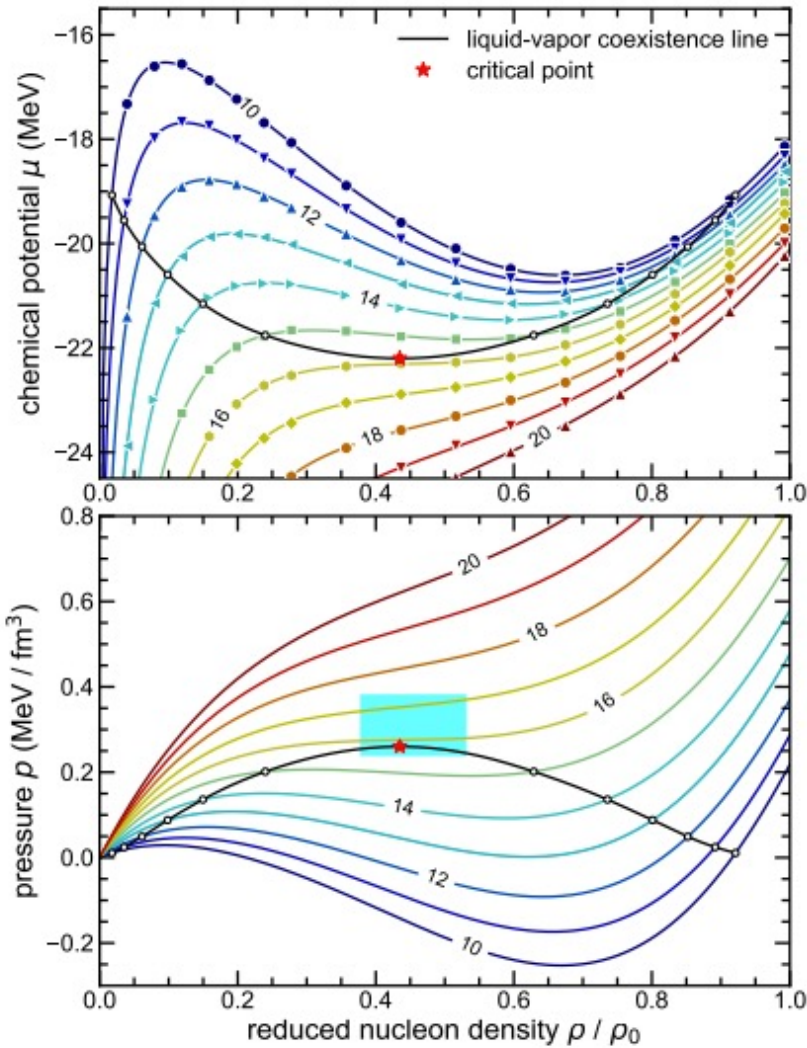
This can be used to calculate the partition function in the canonical ensemble.

# Metropolis updates of pinholes









$$T_c = 15.80(0.32)(1.60) \text{ MeV}$$

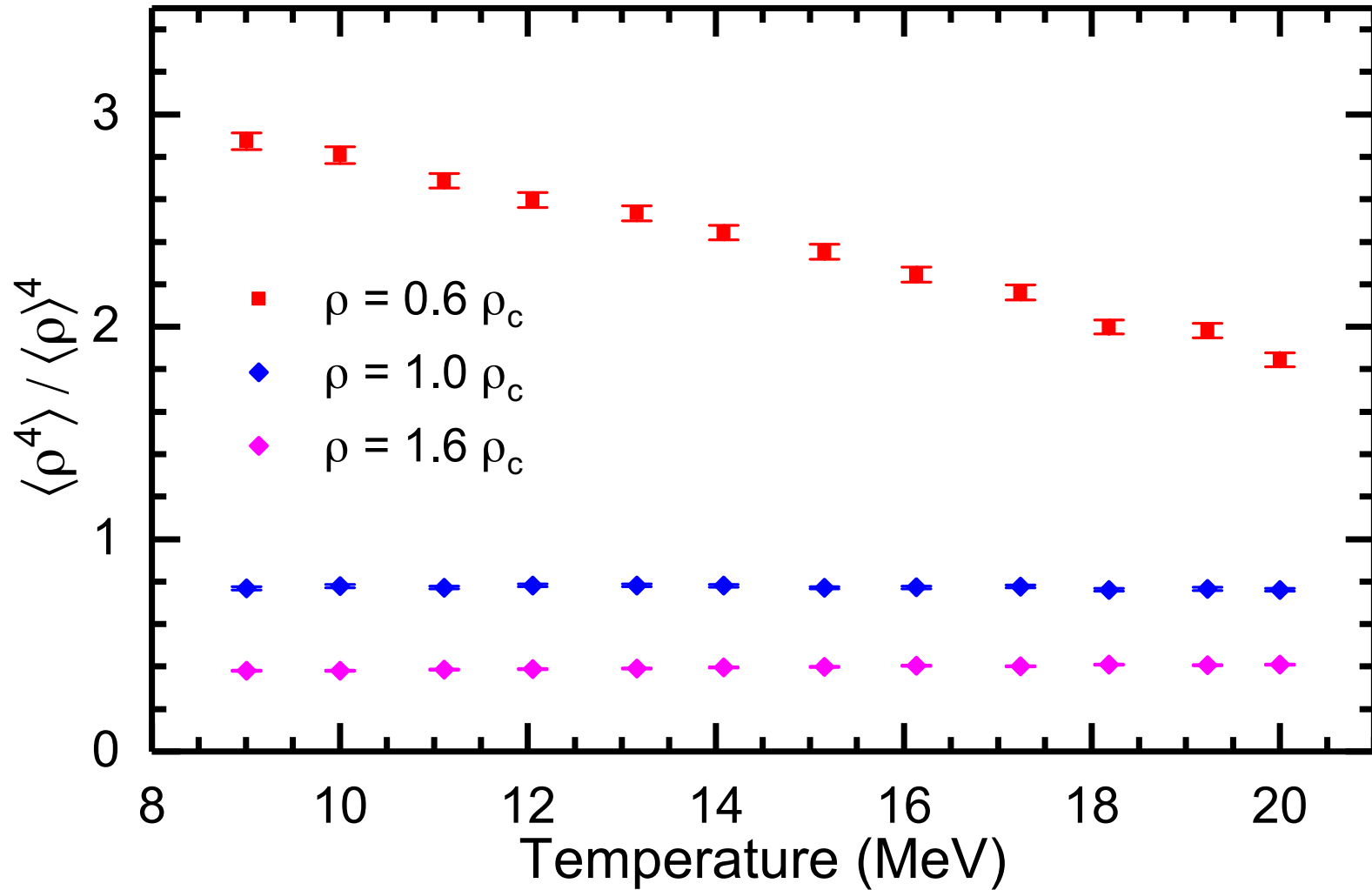
$$\rho_c = 0.089(04)(18) \text{ fm}^{-3}$$

$$\mu_c = -22.20(0.44)(2.20) \text{ MeV}$$

$$P_c = 0.260(05)(30) \text{ MeV fm}^{-3}$$

[Lu, Li, Elhatisari, D.L., Drut, Lähde, Epelbaum, Meißner, PRL 125, 192502 (2020)]

## Alpha clustering as function of density and temperature

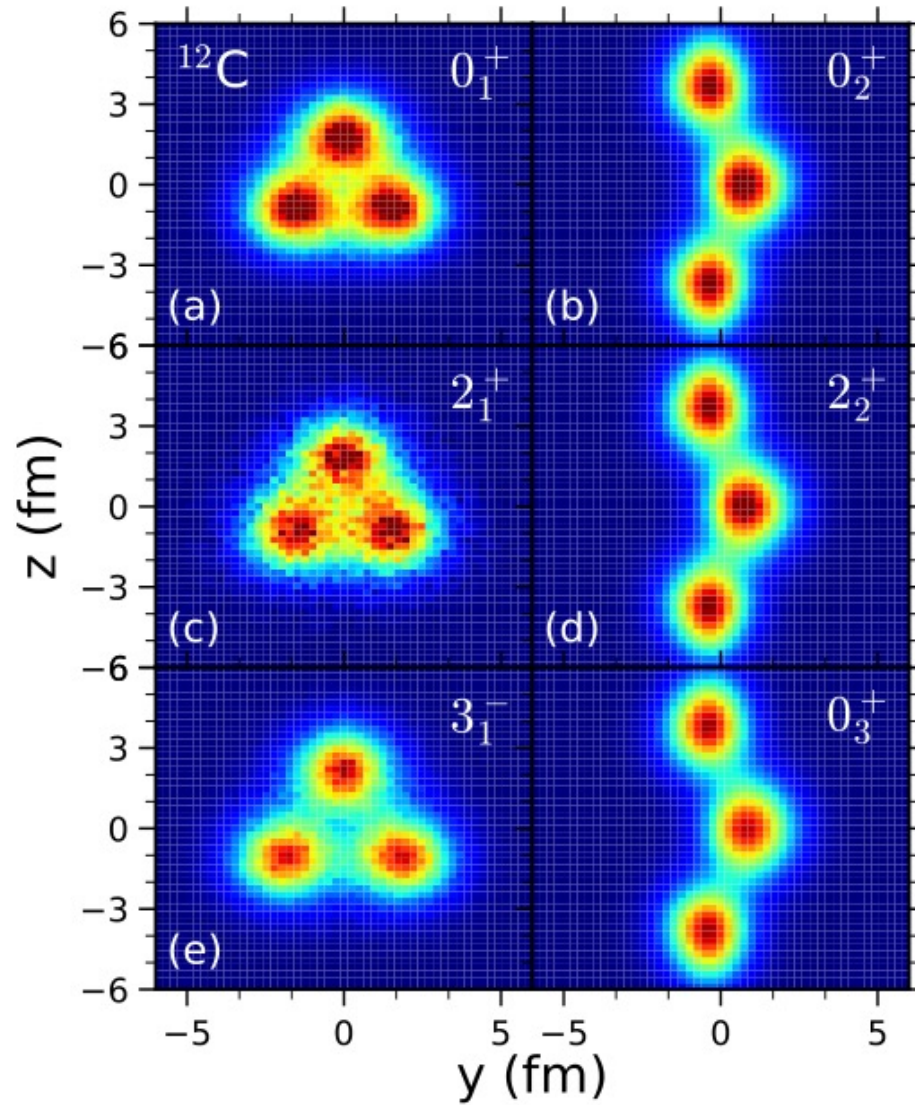


[Lu, Li, Elhatisari, D.L., Drut, Lähde, Epelbaum, Meißner, PRL 125, 192502 (2020)]

## Structure and spectrum of $^{12}\text{C}$

[Shen, Lähde, D.L. Meißner, EPJA 57, 276 (2021)]

State	$a = 1.97 \text{ fm}$	$a = 1.64 \text{ fm}$	Experiment
$0_1^+$	−92.15(3)	−92.12(4)	−92.162
$2_1^+$	−88.87(4)	−88.19(17)	−87.722
$0_2^+$	−85.20(15)	−85.23(22)	−84.508
$3_1^-$	−84.9(2)	−83.3(5)	−82.521(5)
$2_2^+$	−83.5(2)	−83.1(5)	−82.29(6)
$0_3^+$	−80.0(3)	−79.2(6)	−81.9(3)
$1_1^-$	−81.5(4)	−79.7(4)	−81.315(4)
$2_1^-$	−78.6(2)	−76.1(2)	−80.326(4)
$1_1^+$	−79.67(11)	−78.14(24)	−79.452(6)
$4_1^-$	−78.1(2)	−75.5(5)	−78.846(20)
$4_1^+$	−80.99(11)	−79.1(6)	−78.083(5)
$2_3^+$	−79.9(4)	−77.9(2)	−76.056
$0_4^+$	−79.25(11)	−76.94(18)	−74.402



[Work in progress: Shen et al.]



## Outlook

This was an overview of some current capabilities for *ab initio* nuclear structure calculations and how they can interface with studies of relativistic heavy ion collisions.

Over the next decade, it will likely be possible to do *ab initio* calculations of nuclear states across the entire nuclear chart. What interesting problems can the nuclear structure community and the relativistic heavy-ion community address together in the future?