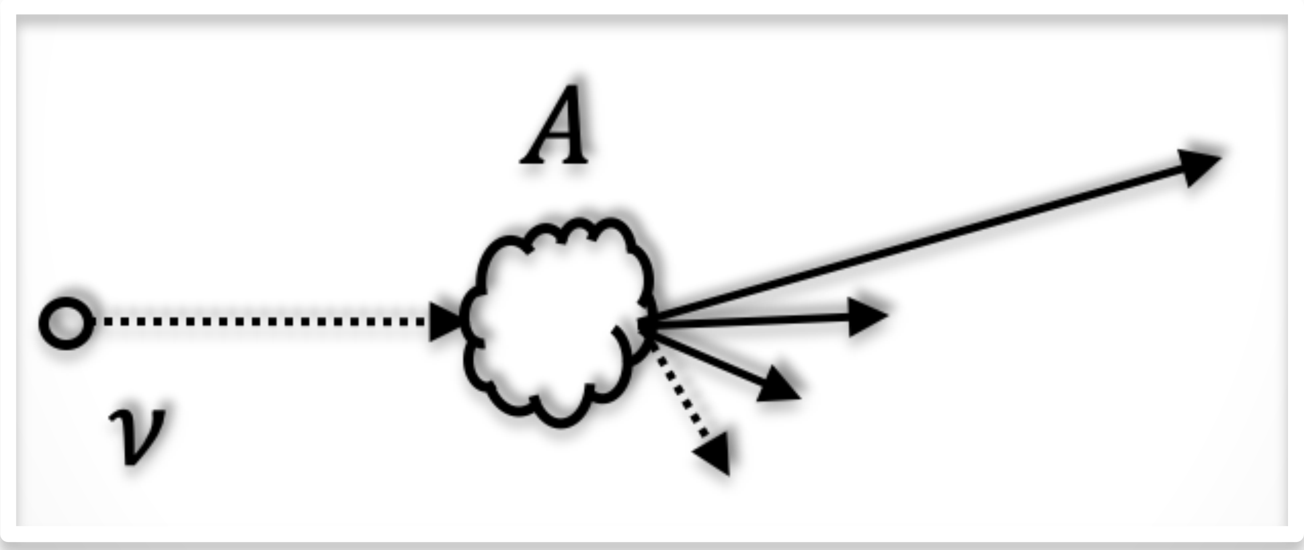
Neutrino-Nucleus Scattering in Neutrino Oscillation Experiments



Shirley Li (Fermilab)

3-Flavor Neutrino Oscillations

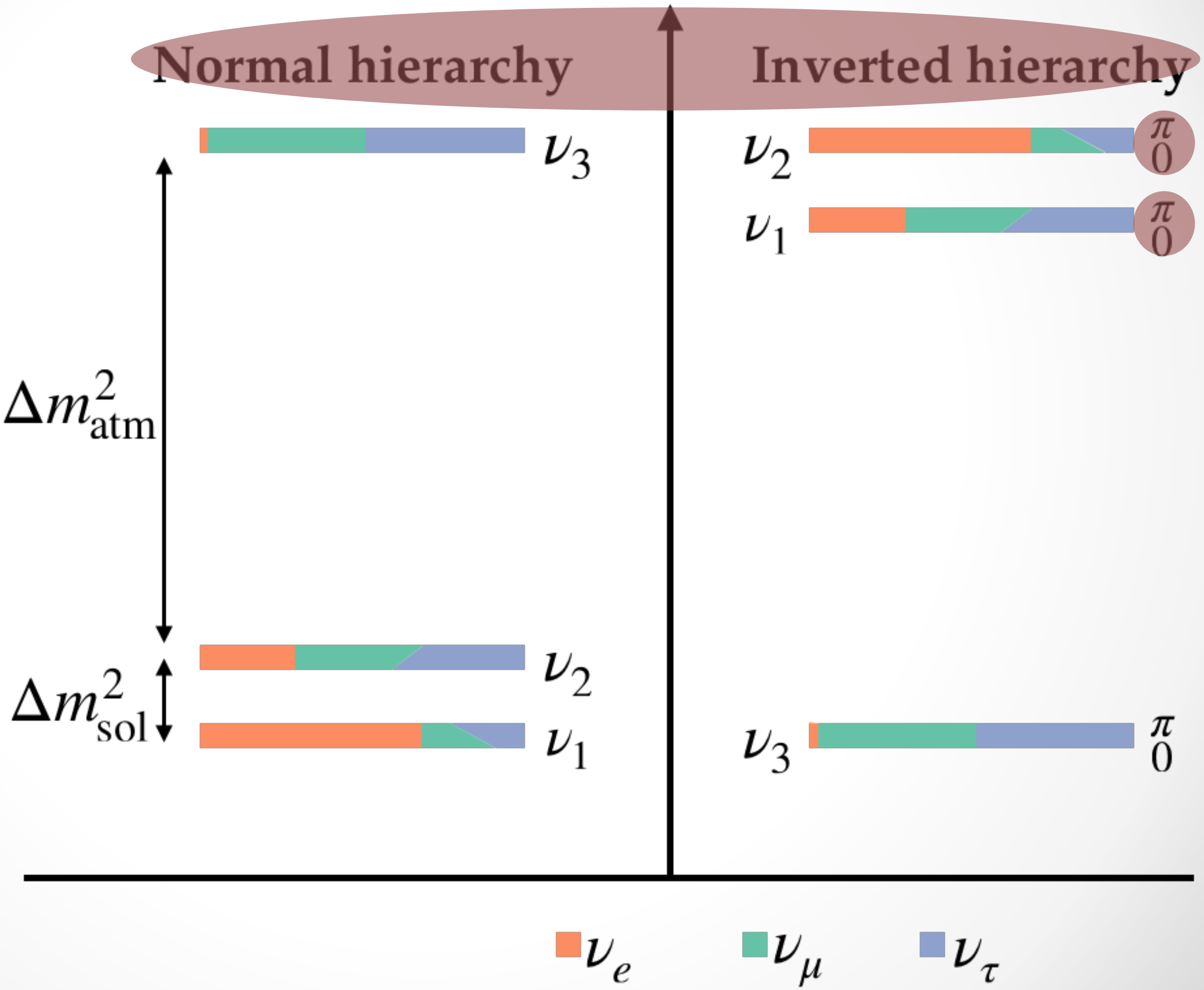
Unknown:

δ_{CP}, MH

A Few %:

$\theta_{12}, \theta_{13}, \theta_{23}$,

$\Delta m^2_{21}, \Delta m^2_{32}$



3-Flavor Neutrino Oscillations

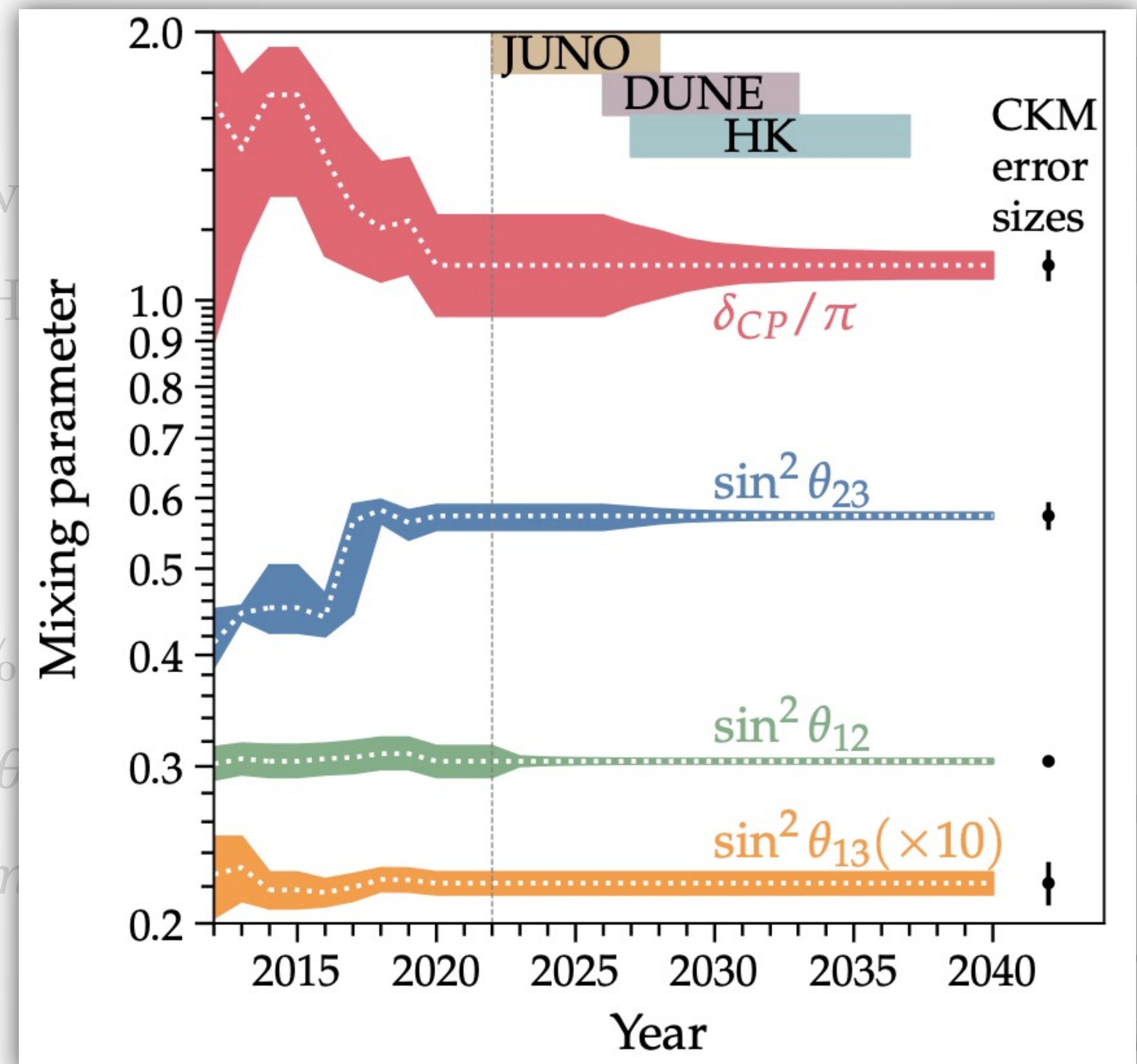


Figure modified from Song et al., 20

Uncovering Neutrino-Nucleus Cross Section Problems

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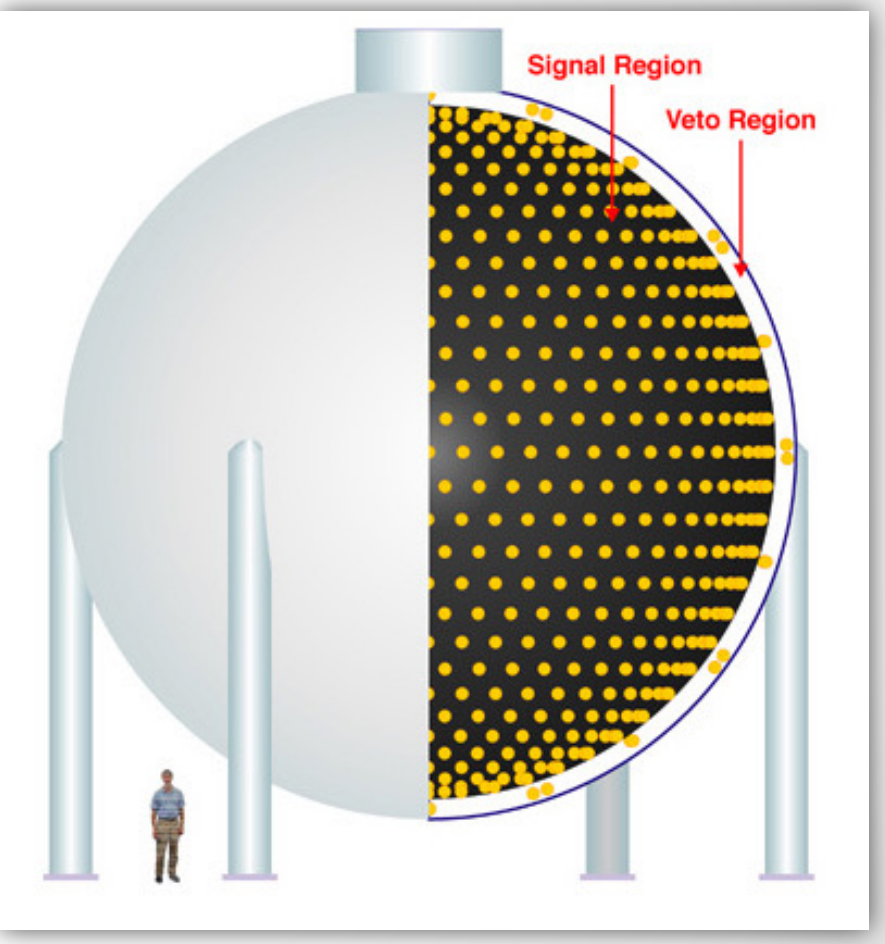
MiniBooNE

- Started to Produce Results in 2007
- Designed to Test Oscillations of eV Sterile Neutrinos
- 800 MeV @ 500 m

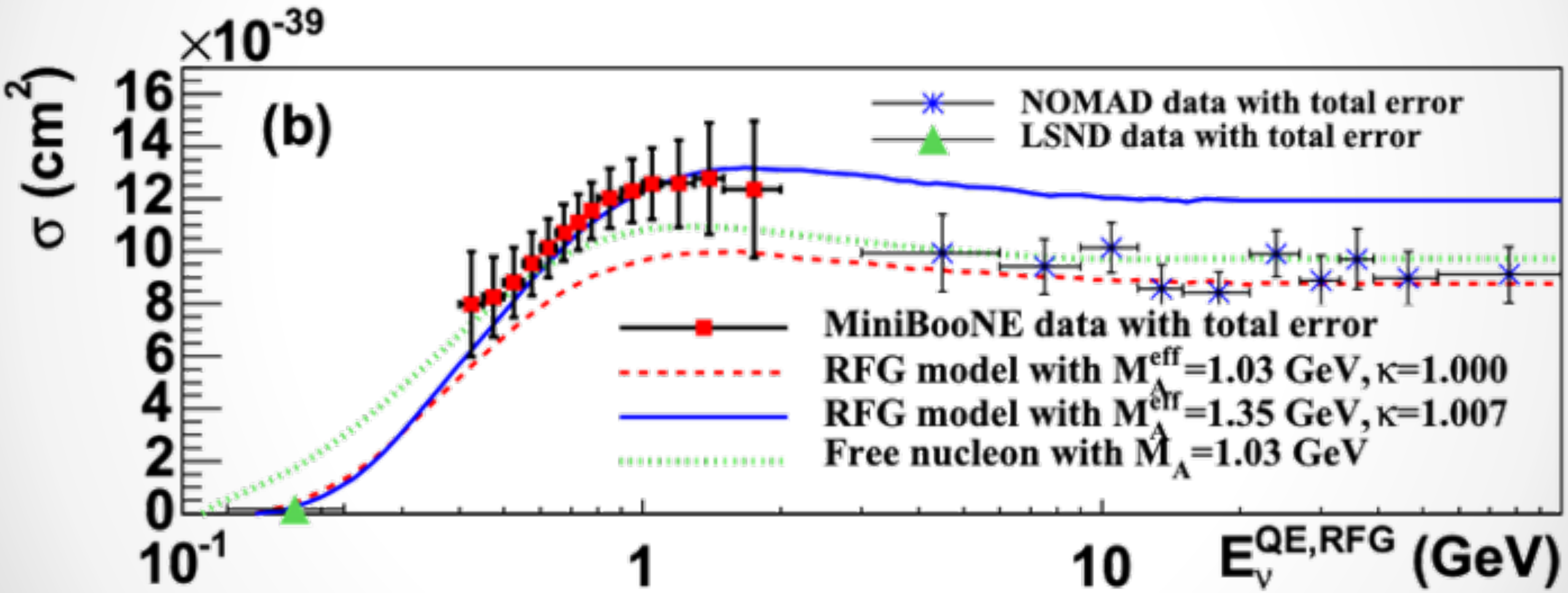
MiniBooNE, 09



Axial Mass m_A Measurement

$$\langle N' | J_W^\mu | N \rangle \supset \bar{\mu}'_N \, \gamma^\mu \, \gamma^5 F_A \mu_N, \qquad F_A(Q^2) = \frac{F_A(0)}{\left(1+\frac{Q^2}{m_A^2}\right)^2}$$

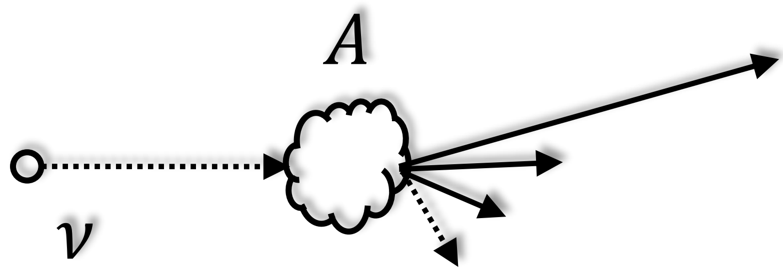
MiniBooNE, 10



Results Disagree with Previous Experiments

Outline

Neutrino-Nucleus Cross Sections



How Do They Affect Measurements

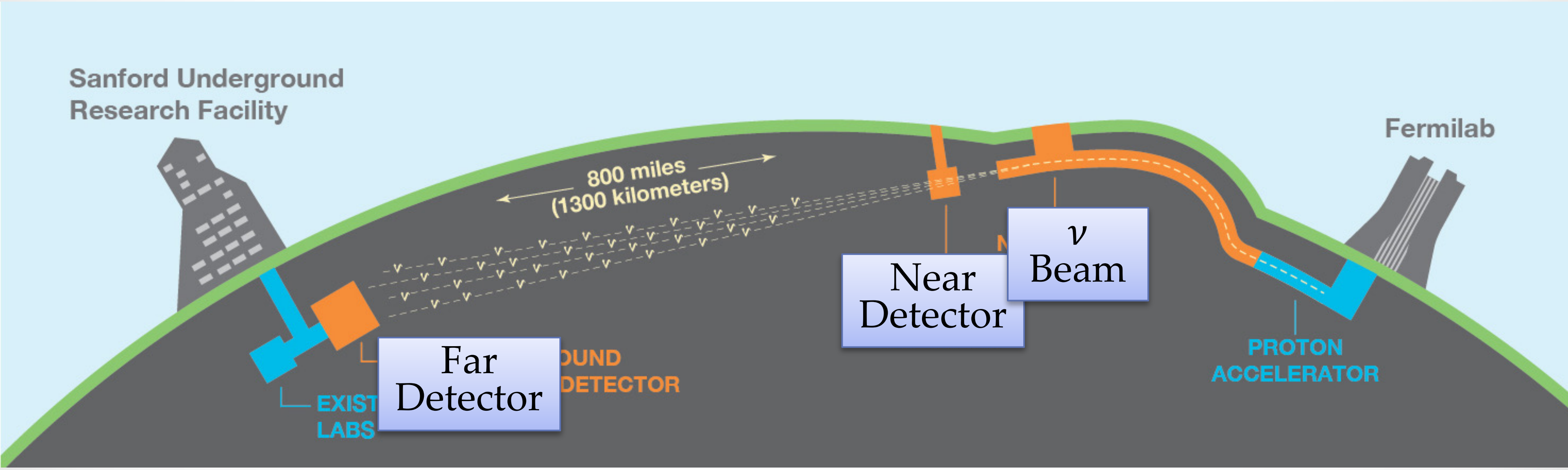
How to Improve

How Do We Compute the Cross Sections?

Long-Baseline Experiments

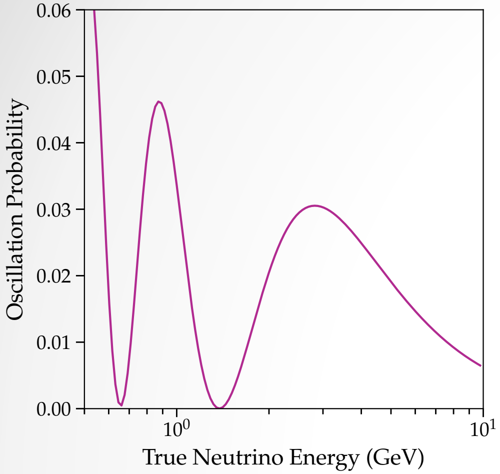
$$0.5\text{—}5\ \text{GeV}\ \nu_\mu \xrightarrow{\sim 1000\ \text{km}} \nu_e$$

DUNE 15



1. Discussions Focus on DUNE and NOvA
2. Details Differ for Hyper-K and T2K

Measuring Neutrino Oscillation



Without Near Detector

Cross Section

Flux

$$P\left(\nu_\mu \rightarrow \nu_e, E_\nu, l = L\right) = \frac{f_e(E_\nu, L)}{f_\mu(E_\nu, 0)} = \frac{N_e(E_\nu, L)/\sigma_e(E_\nu)}{f_\mu(E_\nu, 0)}$$

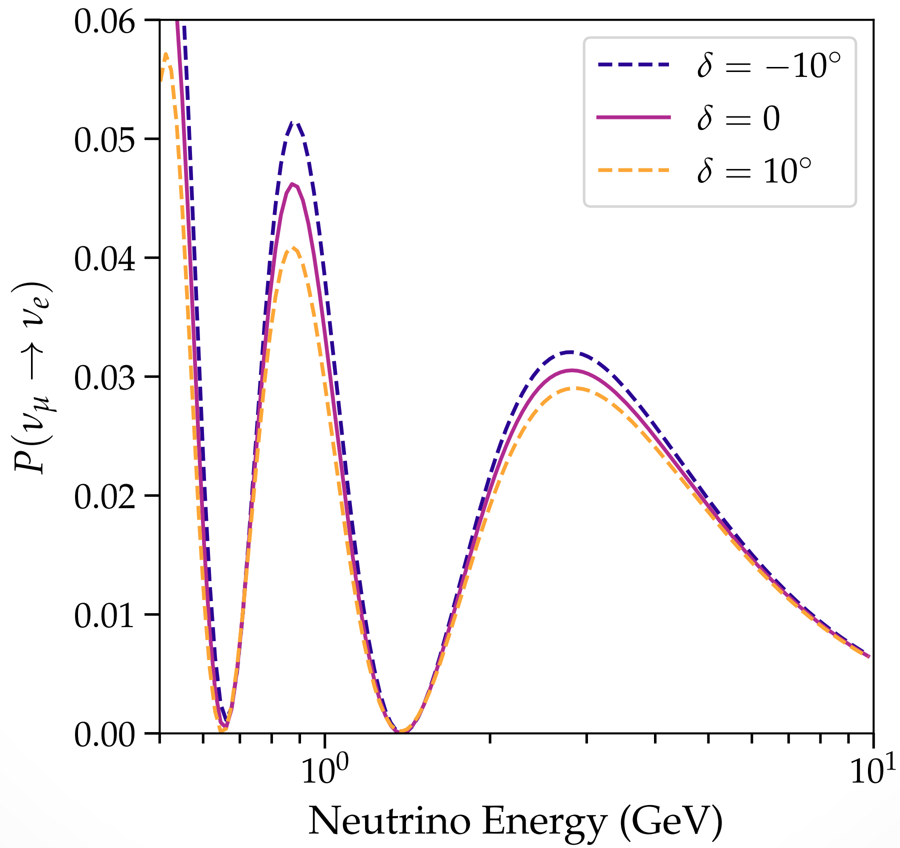
Neutrino Energy

$\sigma_e(E_\nu)$: $\nu_e + {}^{40}\mathrm{Ar} \rightarrow e^- + X$

$$\delta P/P\left(\nu_\mu \rightarrow \nu_e, E_\nu\right) \propto \delta\sigma_e/\sigma_e(E_\nu)$$

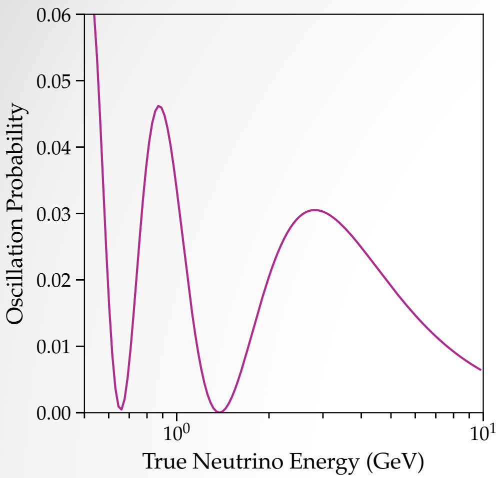
Expected Accuracy

DUNE Nominal Accuracy on δ_{CP} as An Example



Need to Measure $P(E_\nu)$/Predict $\sigma(E_\nu)$ Accurately (≤ 5%)

Measuring Neutrino Oscillation



With Near Detector

Far Detector

$$P\left(\nu_\mu \to \nu_e, E_\nu, l = L\right) = \frac{f_e(E_\nu, L)}{f_\mu(E_\nu, 0)} = \frac{N_e(E_\nu, L)/\sigma_e(E_\nu)}{N_\mu(E_\nu, L)/\sigma_\mu(E_\nu)}$$

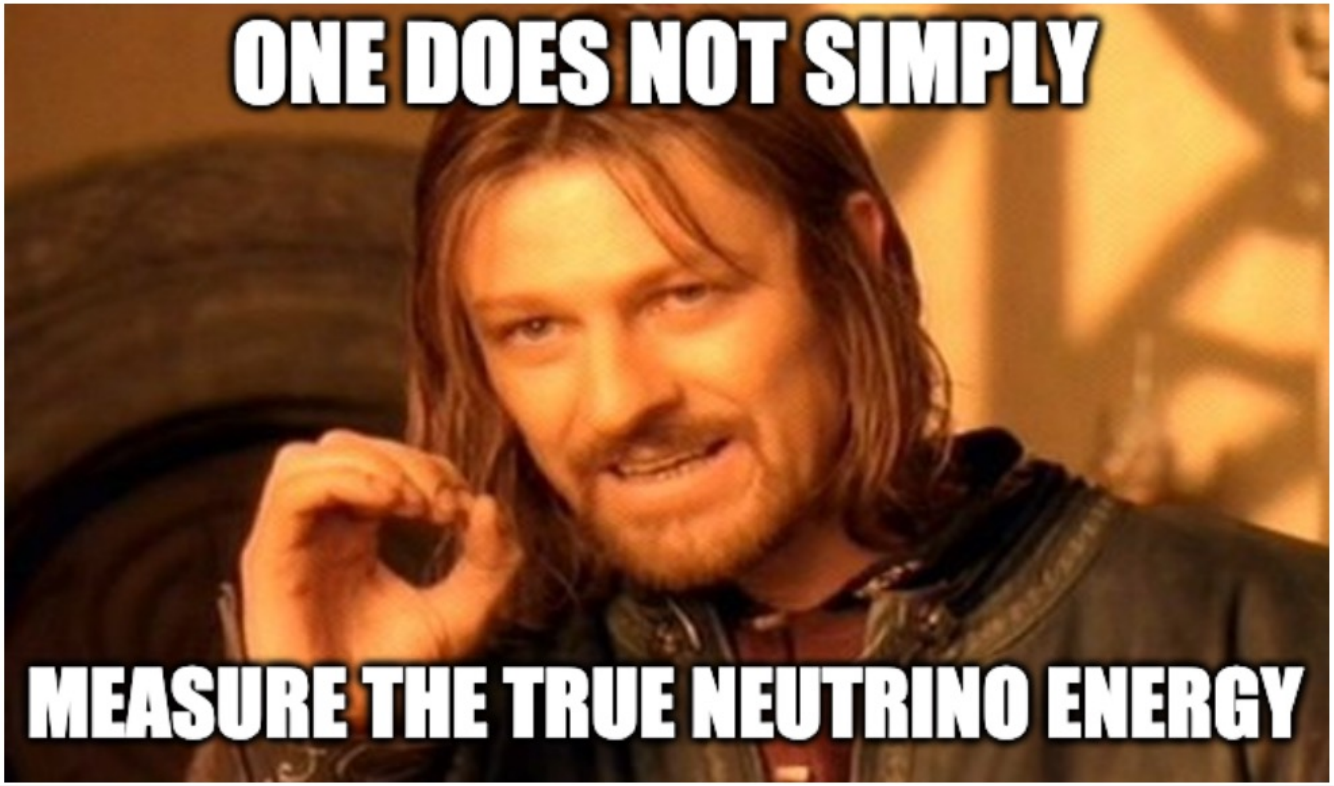
Near Detector

*Difference between σ_e and σ_μ is well appreciated

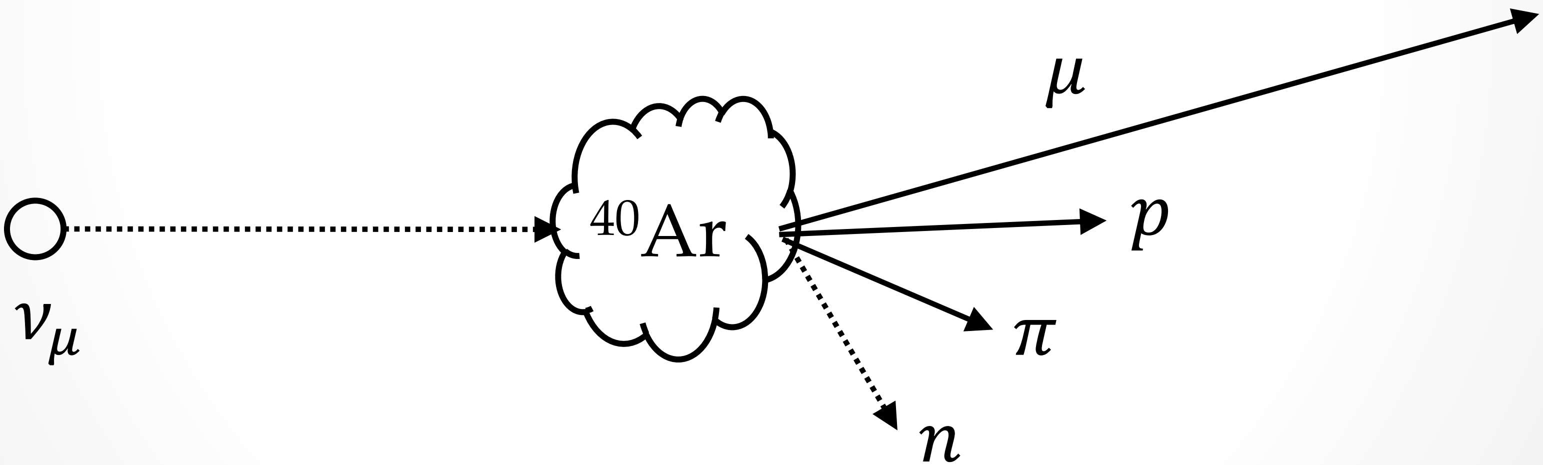
Cross Section Predictions No Longer Play A Role??*

Near/Far Cancellation?

$$P(E_\nu) = \frac{N_e(E_\nu, L)/\sigma_e(E_\nu)}{N_\mu(E_\nu, L)/\sigma_\mu(E_\nu)}$$

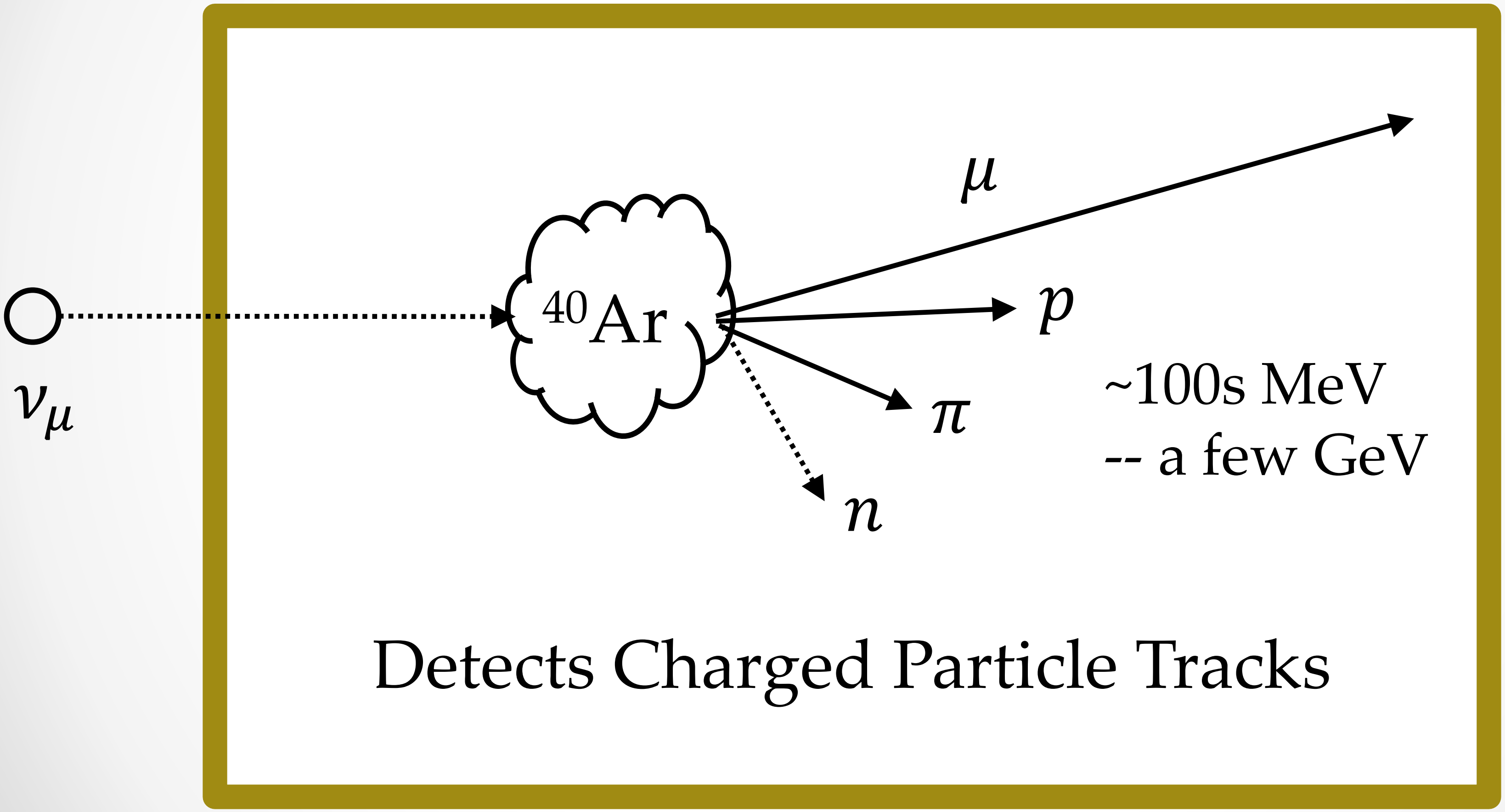


How Do Cross Section Calculations Impact Neutrino Energy Reconstruction?

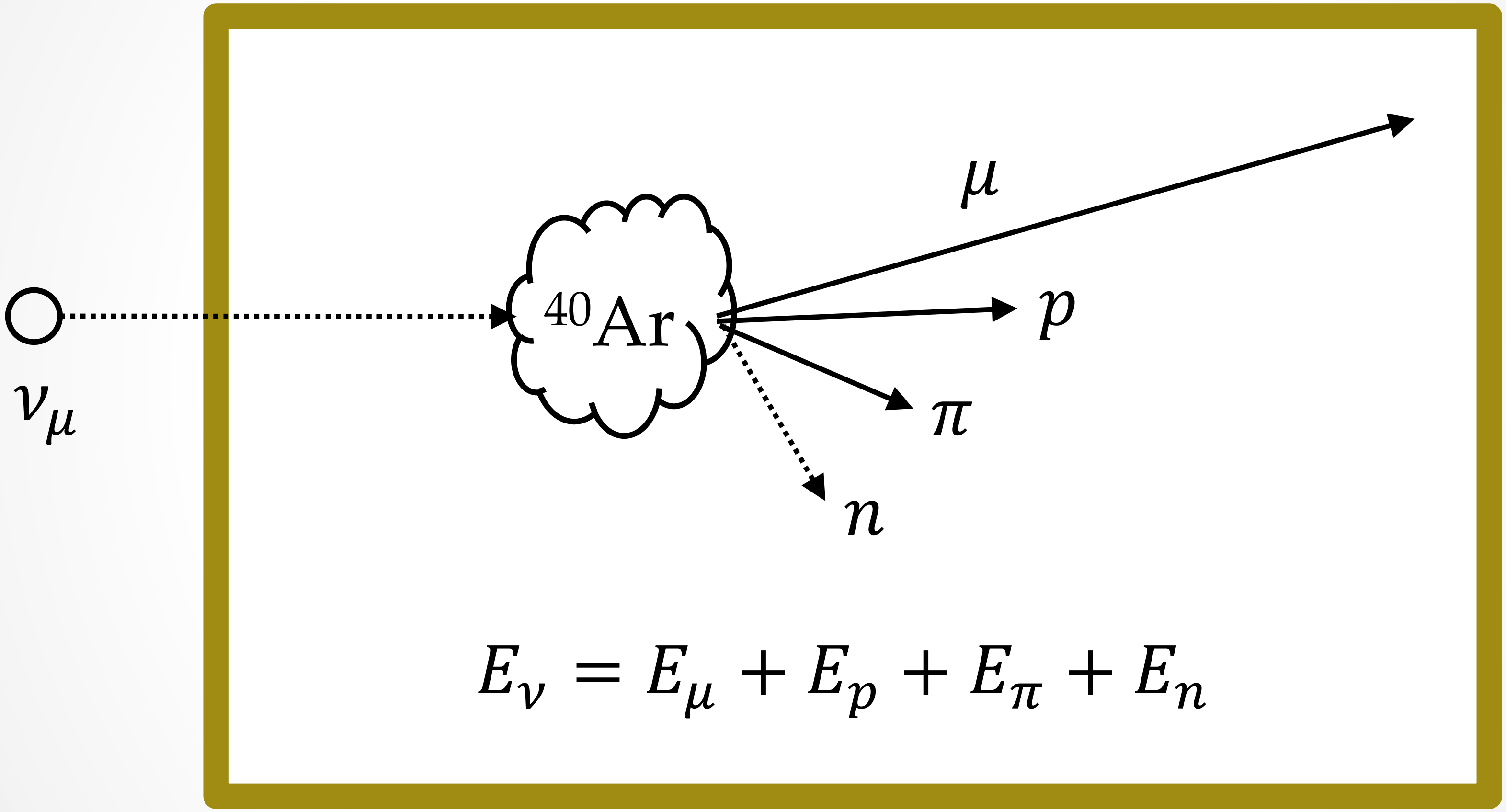


How Neutrinos Are Detected

DUNE: Liquid Argon Time-Projection Chamber

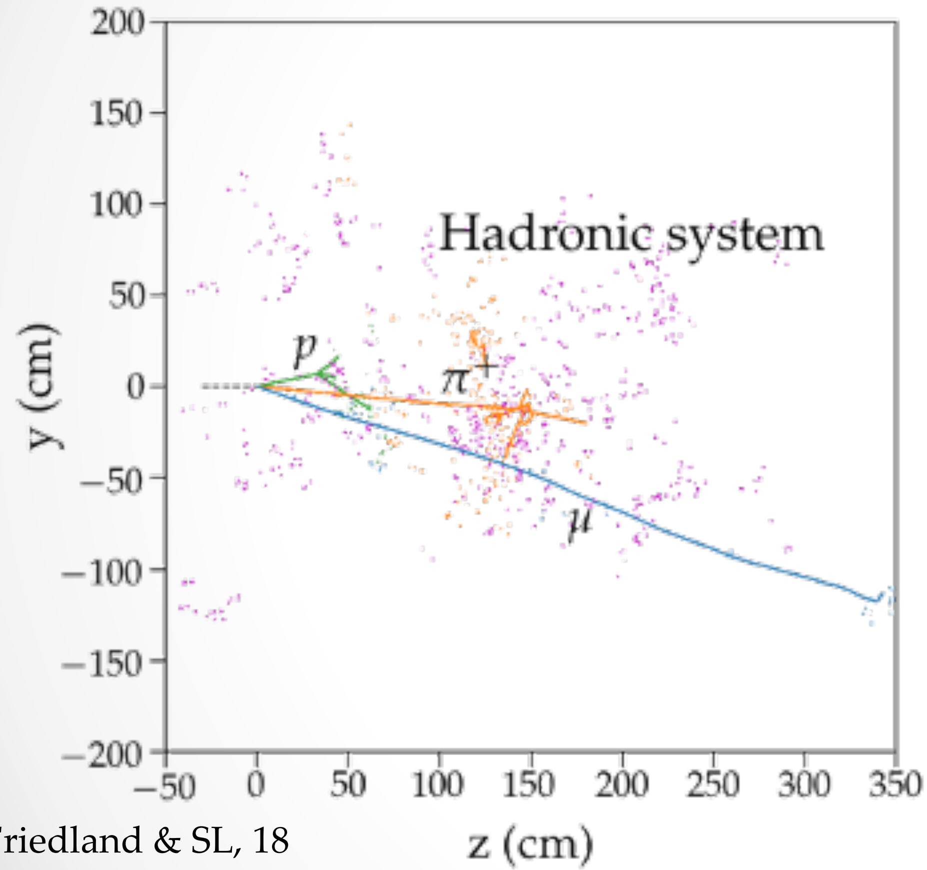


A Theorist's View of a Neutrino Event



Only Predictions of Neutron Fraction Are Important

A Simulated Neutrino Event



Friedland & SL, 18

- Proton vs. Pion: Quenching
- Spectrum: Thresholds
- Number of Final-State Particles: Nuclear Breakup Energy

All Exclusive Final States Play A Role

The Cross Section Predictions That We Need:

$$\frac{\mathrm{d}\sigma}{\mathrm{d}E_1 \mathrm{d}E_2 \ldots \mathrm{d}E_n}$$

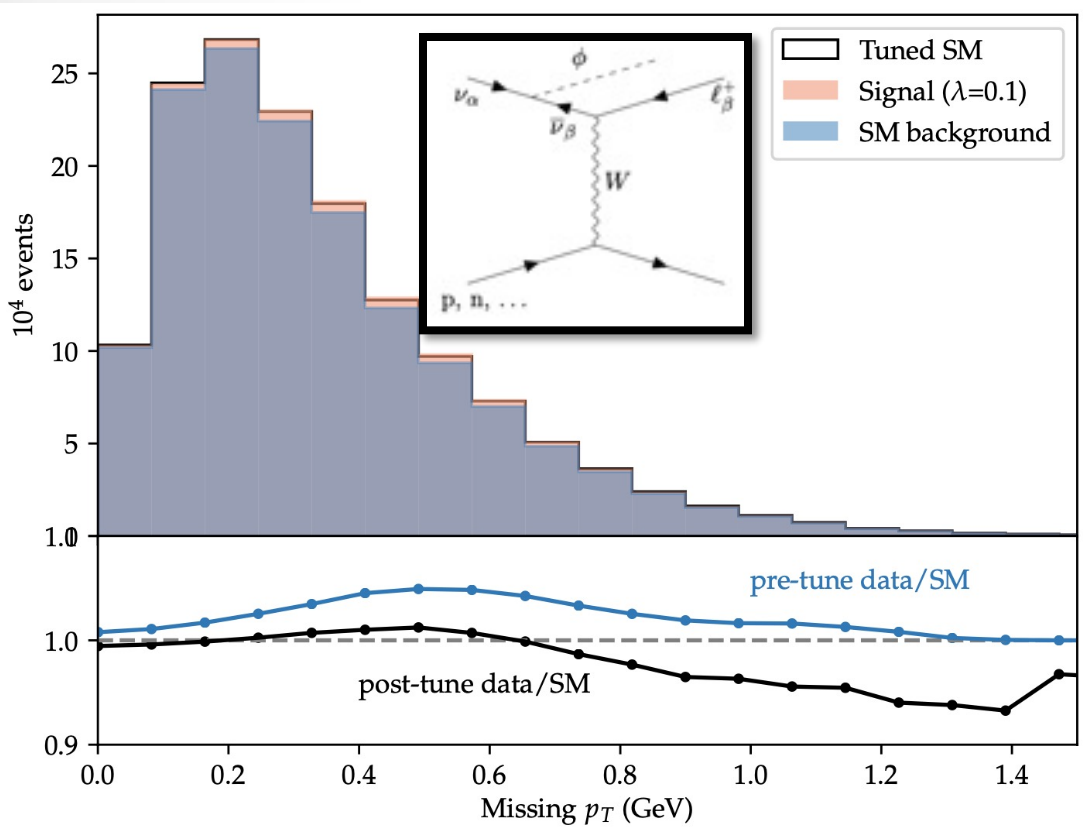
Not So Much:

$$\sigma(E_\nu)$$

Impact on BSM Searches

Case Study 1: Searching for Missing p_T

Coyle, SL, Machado, in prep



Signature:
Large Missing p_T

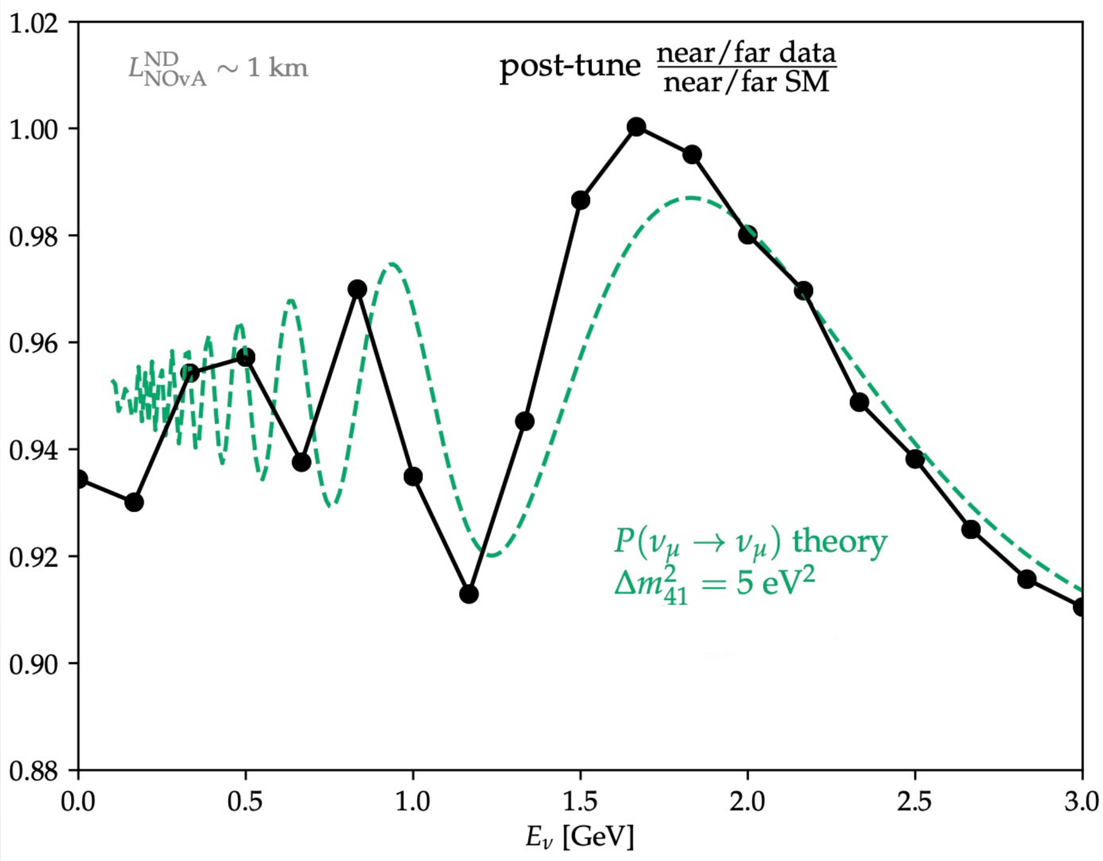
E.g., $\mathcal{O} = \frac{(L_\alpha H)(L_\beta H)}{\Lambda^2_{\alpha\beta}}\phi \to \frac{1}{2}\lambda_{\alpha\beta}\nu_\alpha\nu_\beta\phi$

After Tune:
Signature Gone

Impact on BSM Searches

Case Study 2: Sterile Neutrinos

Coyle, SL, Machado, in prep

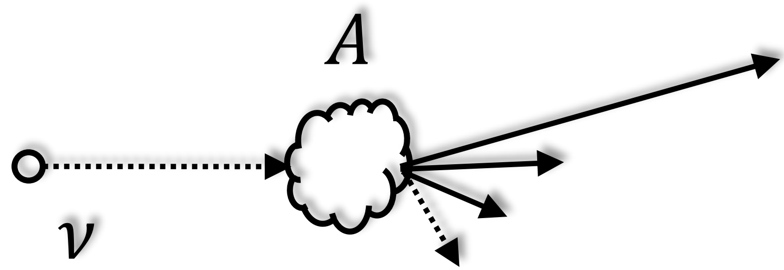


Signature:
Wiggles in Near/Far Ratios

After Tune:
Slight Shift,
Signature Persists

Outline

Neutrino-Nucleus Cross Sections



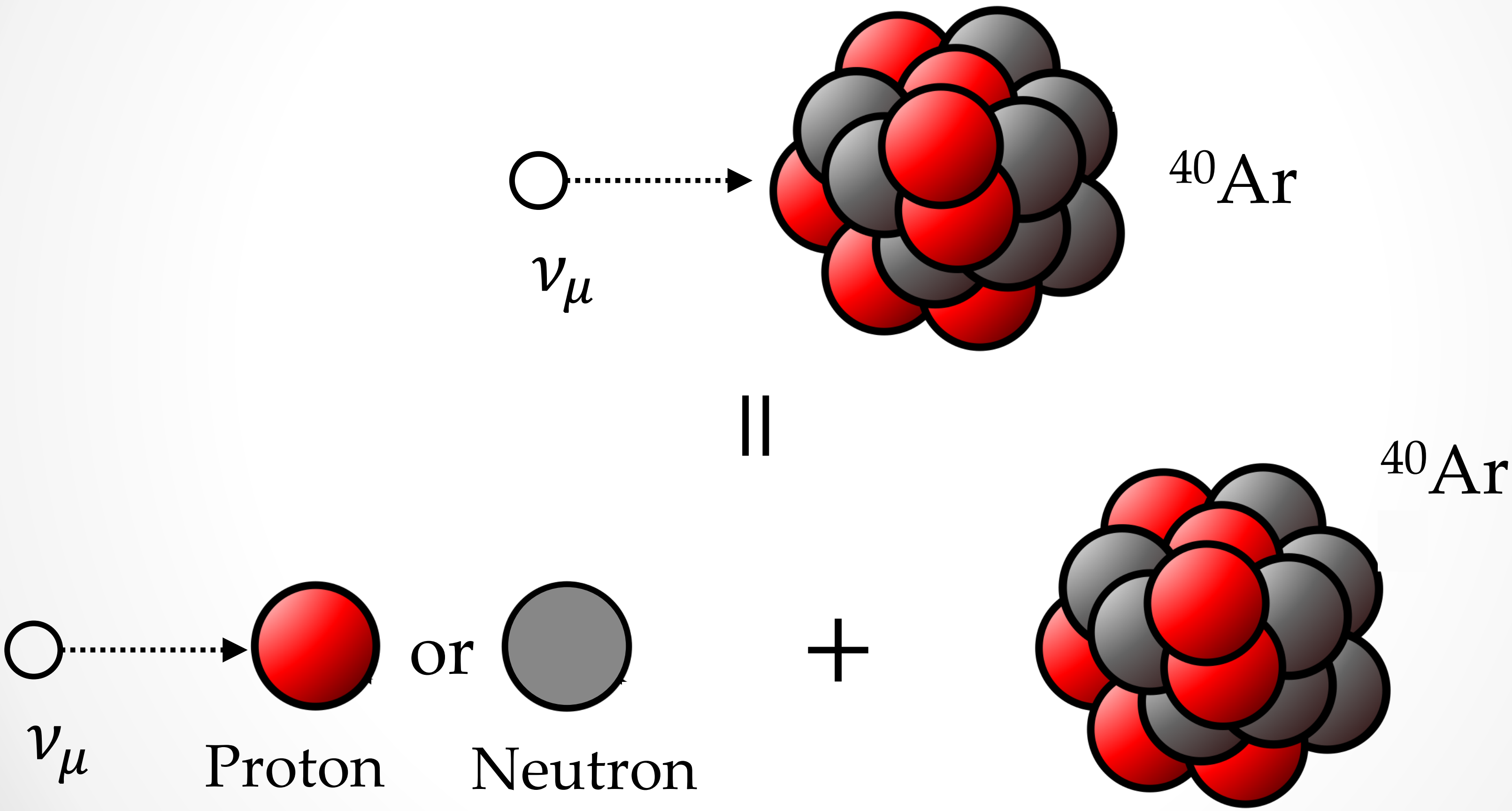
How Do They Affect Measurements

How to Improve

How Do We Compute the Cross Sections?

ν-Nucleus Cross Sections

Beam Energy: 0.5 GeV – 5 GeV

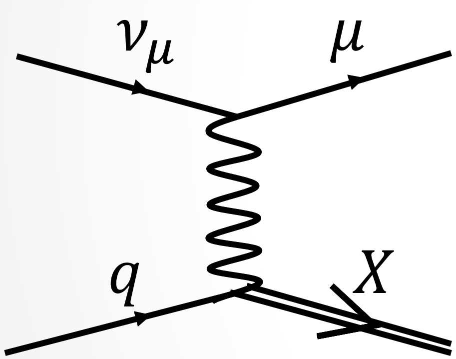


^{12}C, ^{16}O for Other Experiments

ν-Nucleon Cross Sections

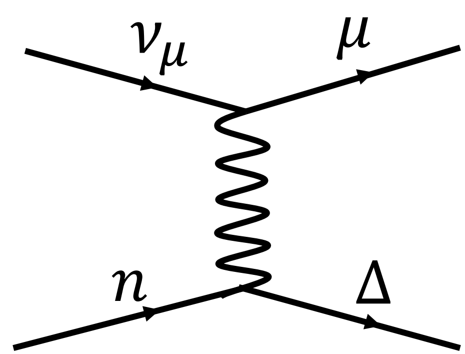
Beam Energy: 0.5 GeV – 5 GeV

Deep Inelastic Scattering



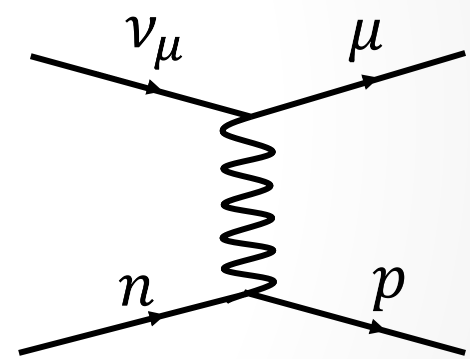
Multiple Particles

Resonance Production



Nucleon + Meson

Quasi-Elastic Scattering



Single Nucleon

No Controlled Expansion

$Q^2 \cong 1\ \text{GeV}^2$

Kinematic Region

Ankowski, Friedland, SL, in prep

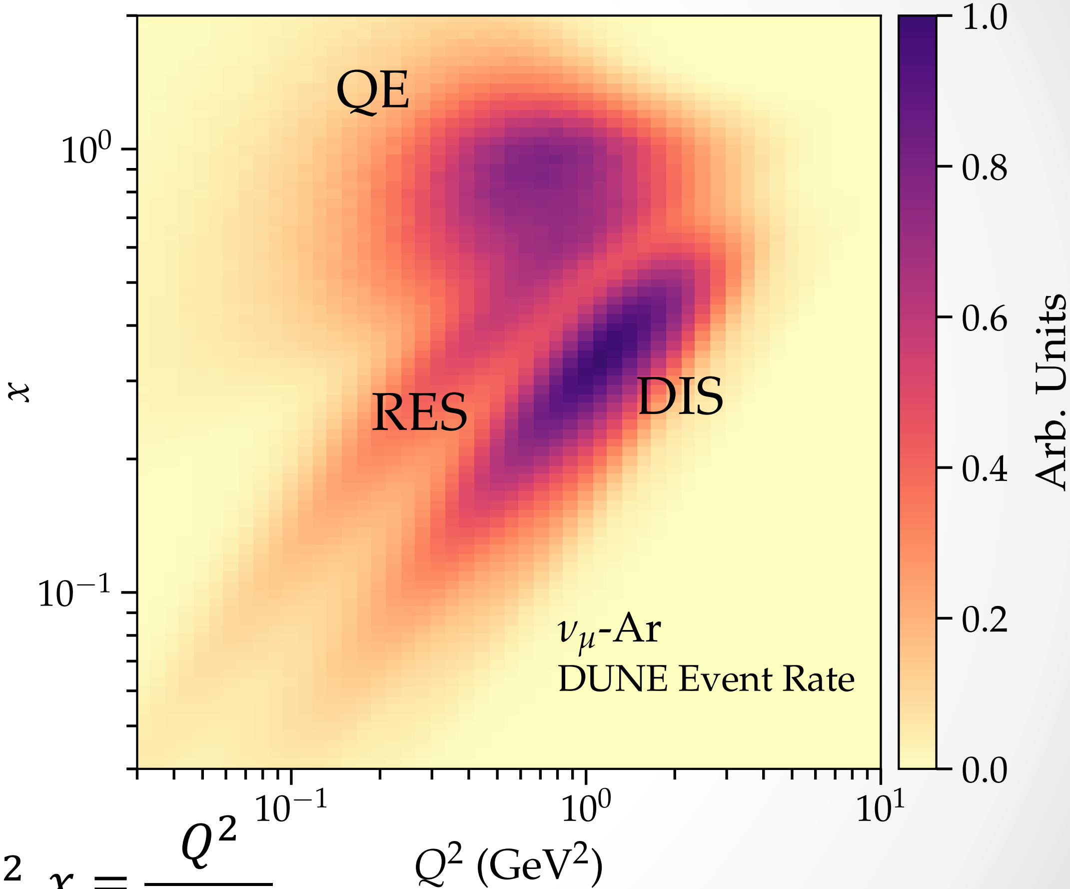
- DIS: $\nu_\mu + d \rightarrow \mu + u$

  Energy Too Low for Factorization

- RES: $\nu_\mu + n \rightarrow \mu + \Delta$

  30+ Resonance States, Not Enough Data

- QE: $\nu_\mu + n \rightarrow \mu + p$



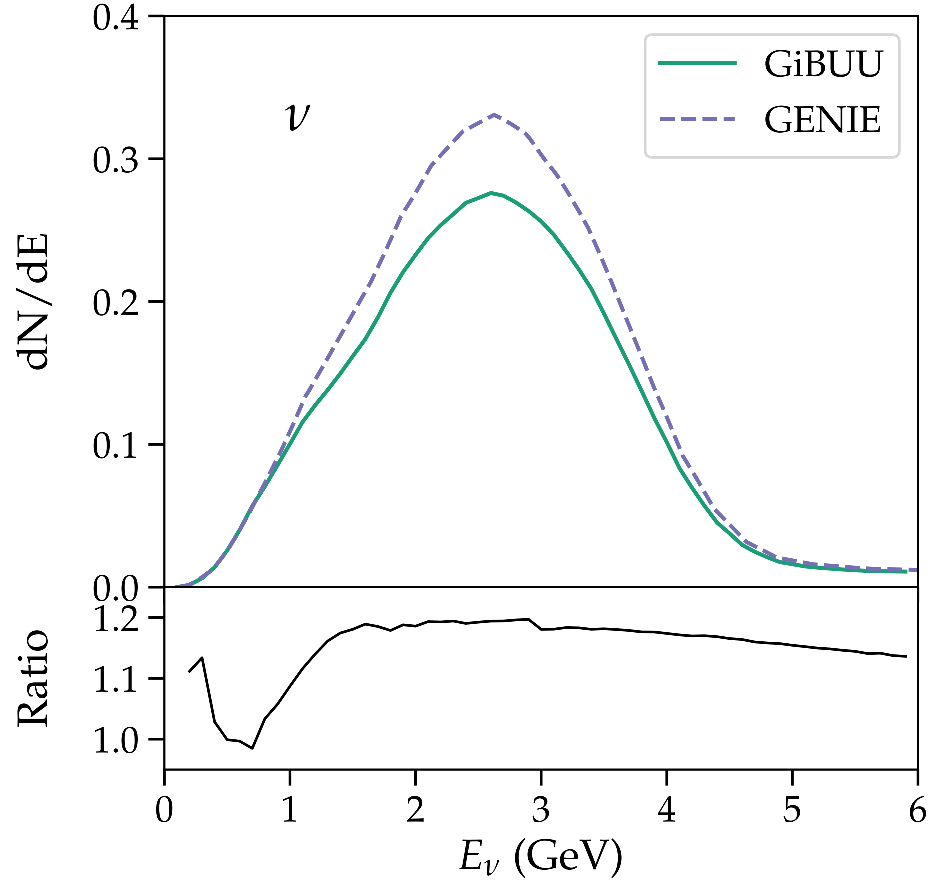
$$Q^2 = -q^2, x = \frac{Q^2}{2p \cdot q}$$

All Channels Are Important

Generator vs. Generator

Ankowski, Friedland, SL, in prep

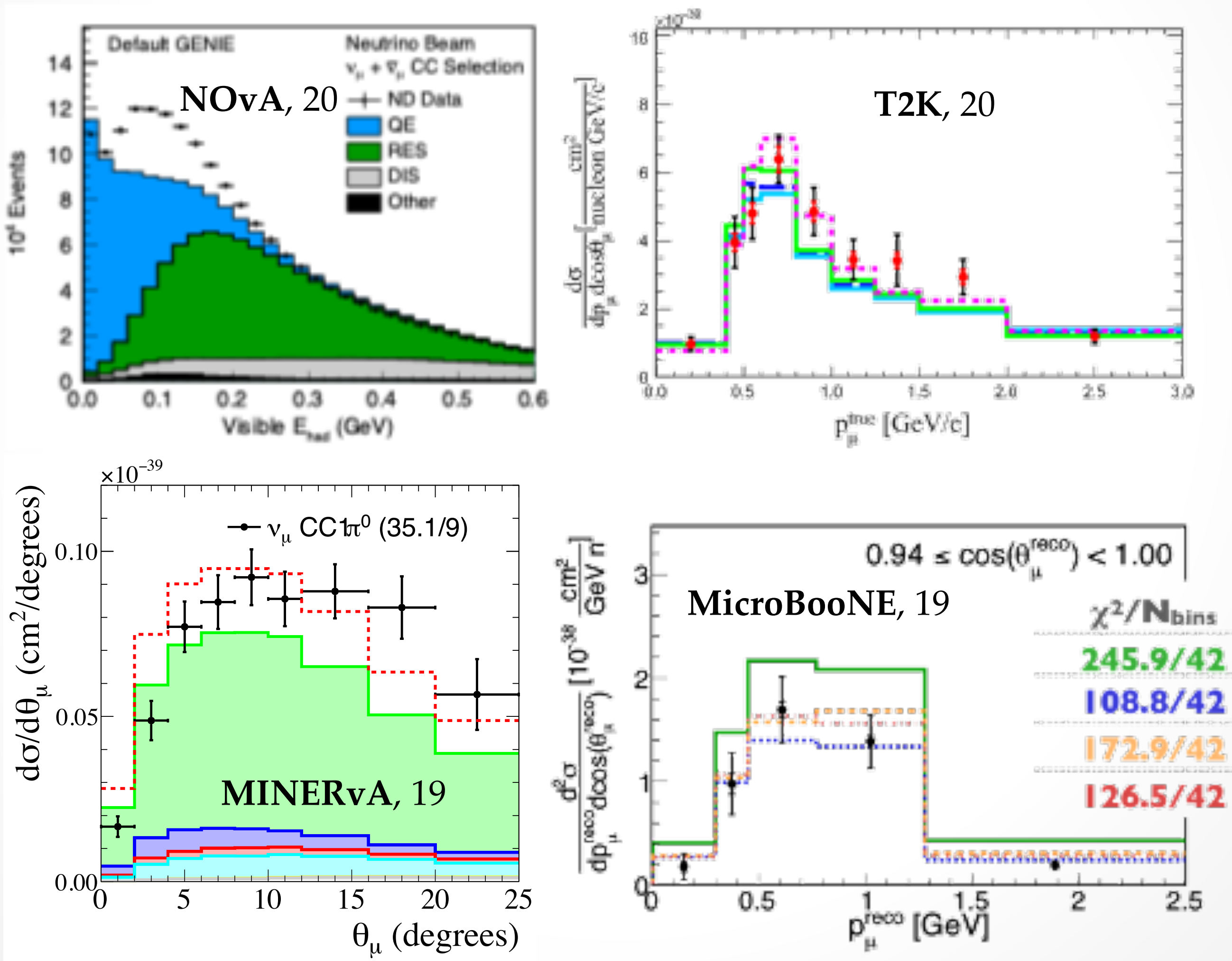
Compare $\sigma(E_\nu)$



Informative, But Not Adequate

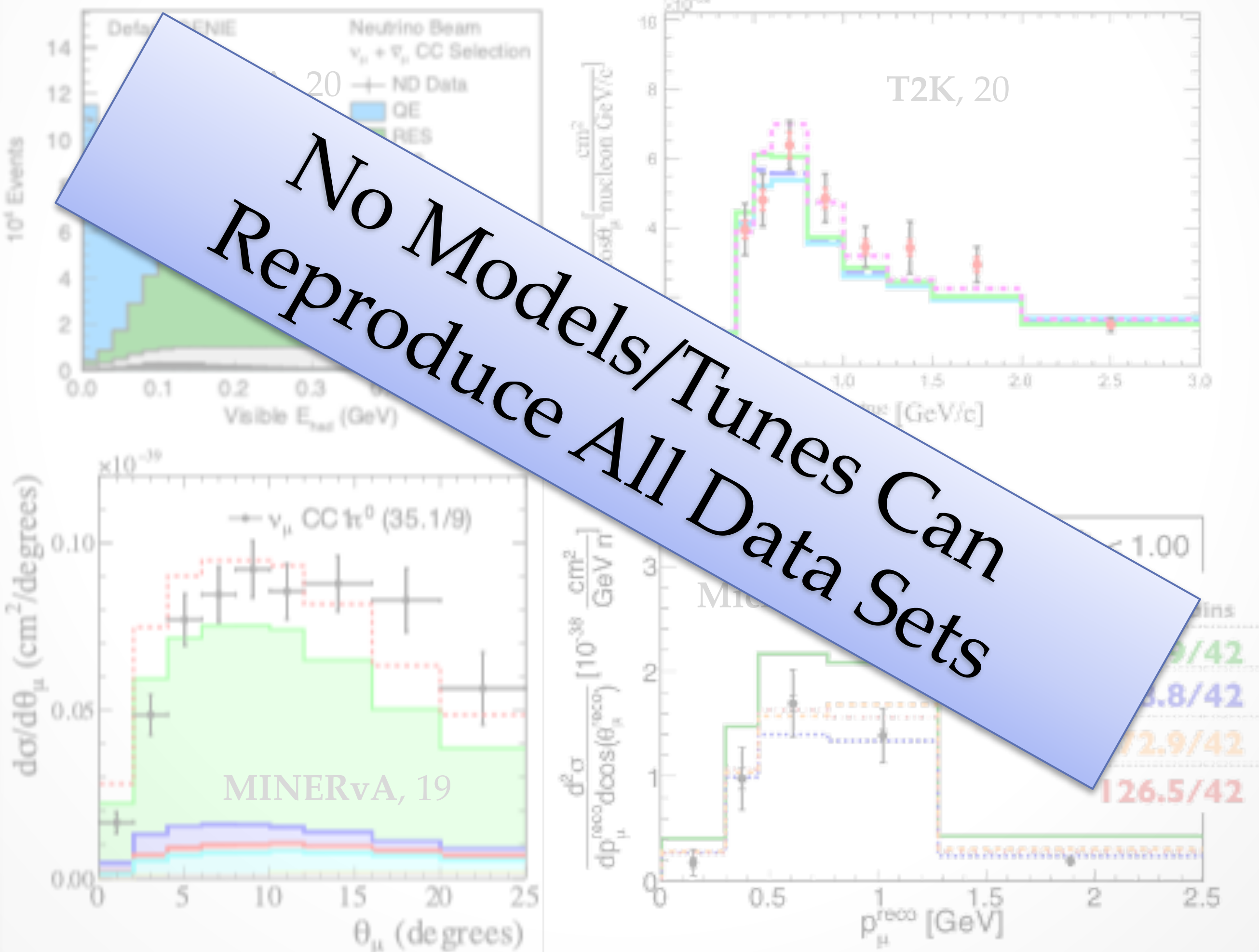
Generator vs. ν Data

Experiments Use Generators, Mostly GENIE



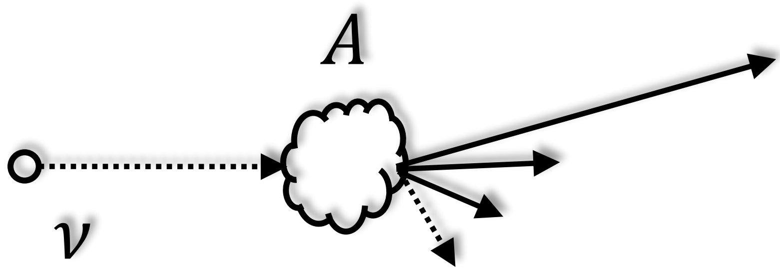
Generator vs. ν Data

Experiments Use Generators, Mostly GENIE



Outline

Neutrino-Nucleus Cross Sections



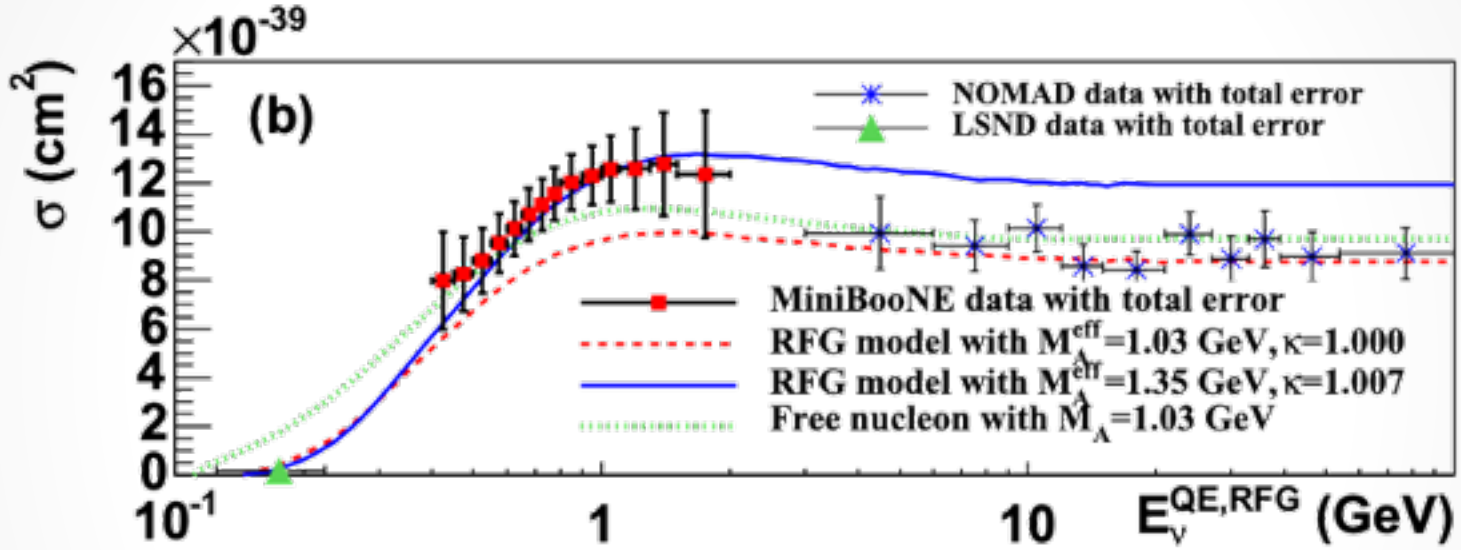
How Do They Affect
Measurements

How to Improve

How Do We Compute
the Cross Sections?

MiniBooNE Setting the Tone

MiniBooNE, 10

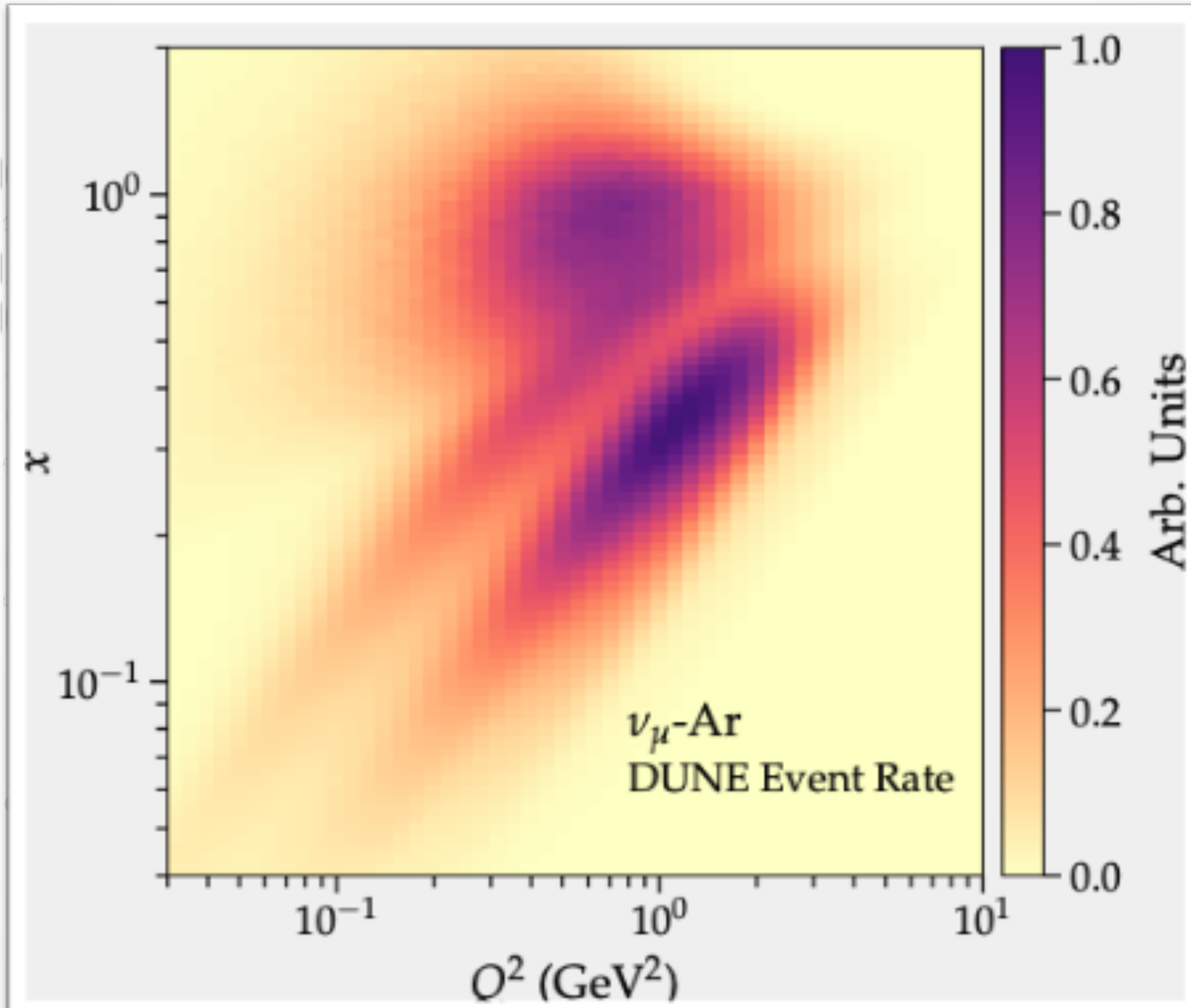


Quasi-Elastic And Meson-Exchange Current Channels Are Important

What About Resonance Production and DIS?

MiniBooNE Setting the Tone

MiniBooNE, 10

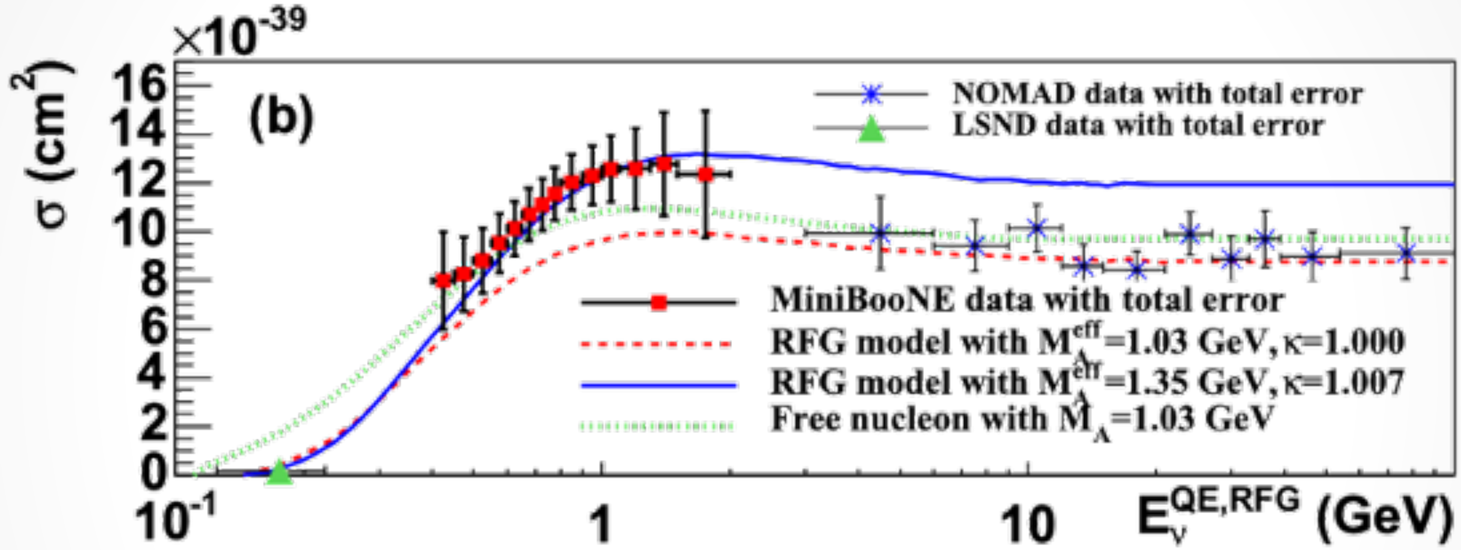


Quasi-El... Channels

What About Resonance Production and DIS?

MiniBooNE Setting the Tone

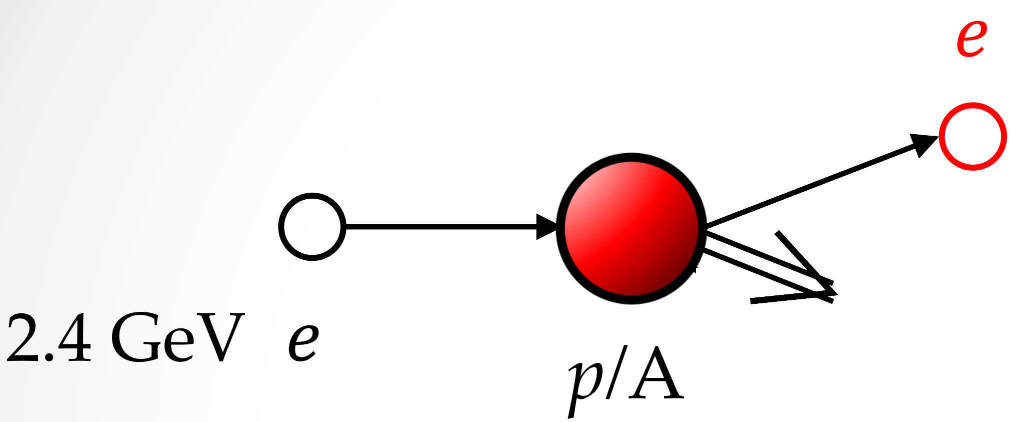
MiniBooNE, 10

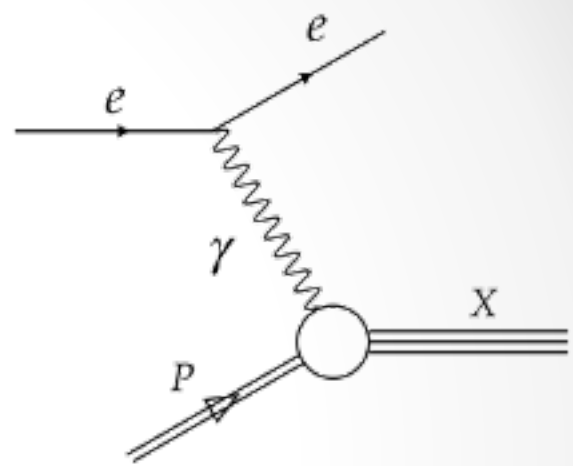


Quasi-Elastic And Meson-Exchange Current Channels Are Important

What About Resonance Production and DIS?

ν-A Generators vs. e-A Data





- Same Primary Vertex Models, Only Vector Couplings
- Same Final-State Interactions / Nucleon Distributions

E.g., Elastic Scattering:

$e-p$

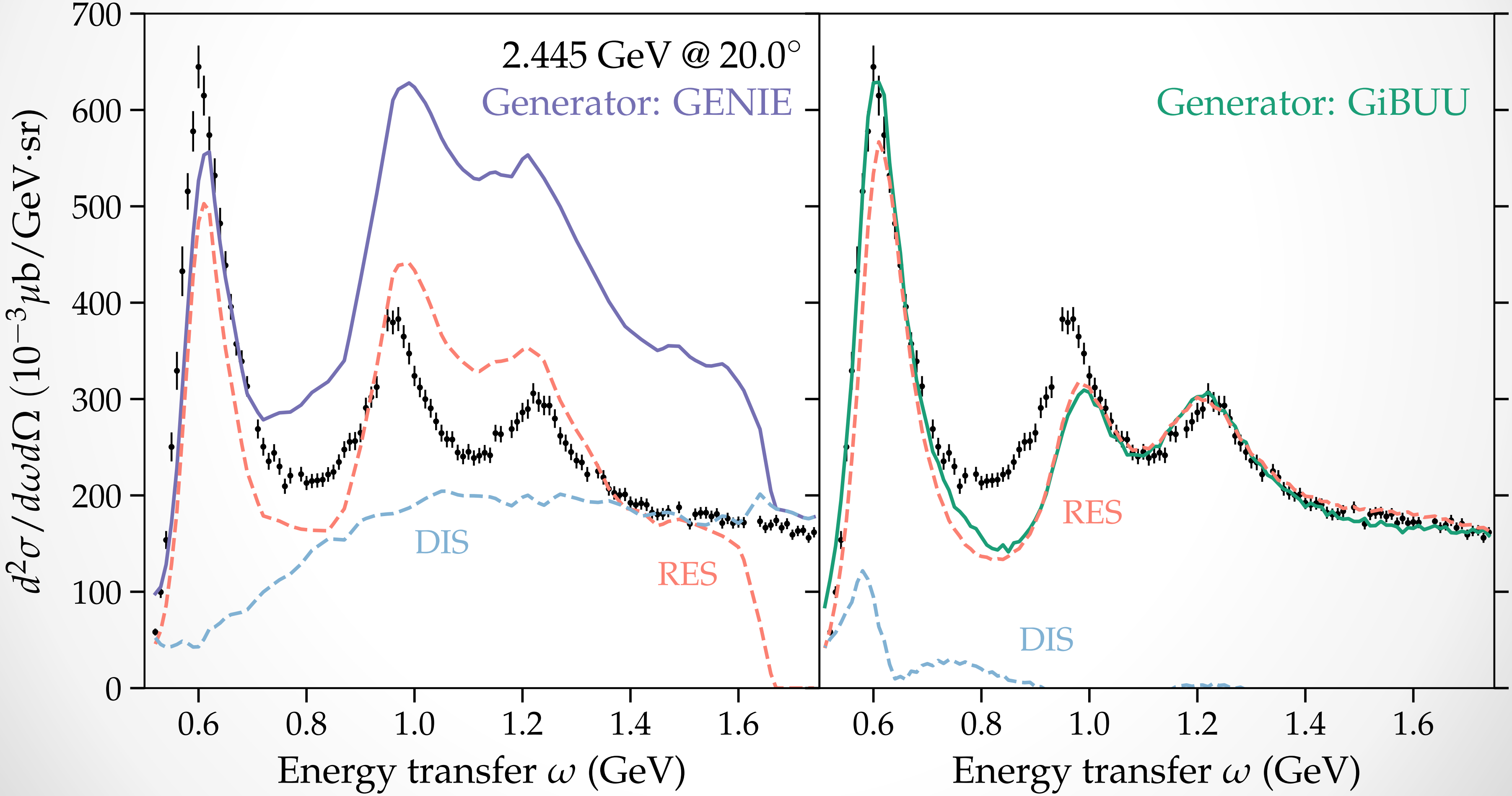
$$\langle N' | J^{\mu}_{\gamma} | N \rangle = \bar{\mu}'_N \left\{ \gamma^{\mu} F_1 + \frac{i\sigma^{\mu\nu} q_{\nu}}{2M} F_2 \right\} \mu_N$$

$\nu-p$

$$\langle N' | J^{\mu}_{W} | N \rangle = \bar{\mu}'_N \left\{ \gamma^{\mu} F_1 + \frac{i\sigma^{\mu\nu} q_{\nu}}{2M} F_2 + \gamma^{\mu} \gamma^5 F_A + \frac{q^{\mu} \gamma^5}{M} F_P \right\} \mu_N$$

e-p Scattering

Neutrino Generators vs. Data



Ankowski, Friedland, SL, in prep
Data from Niculescu et al., 00

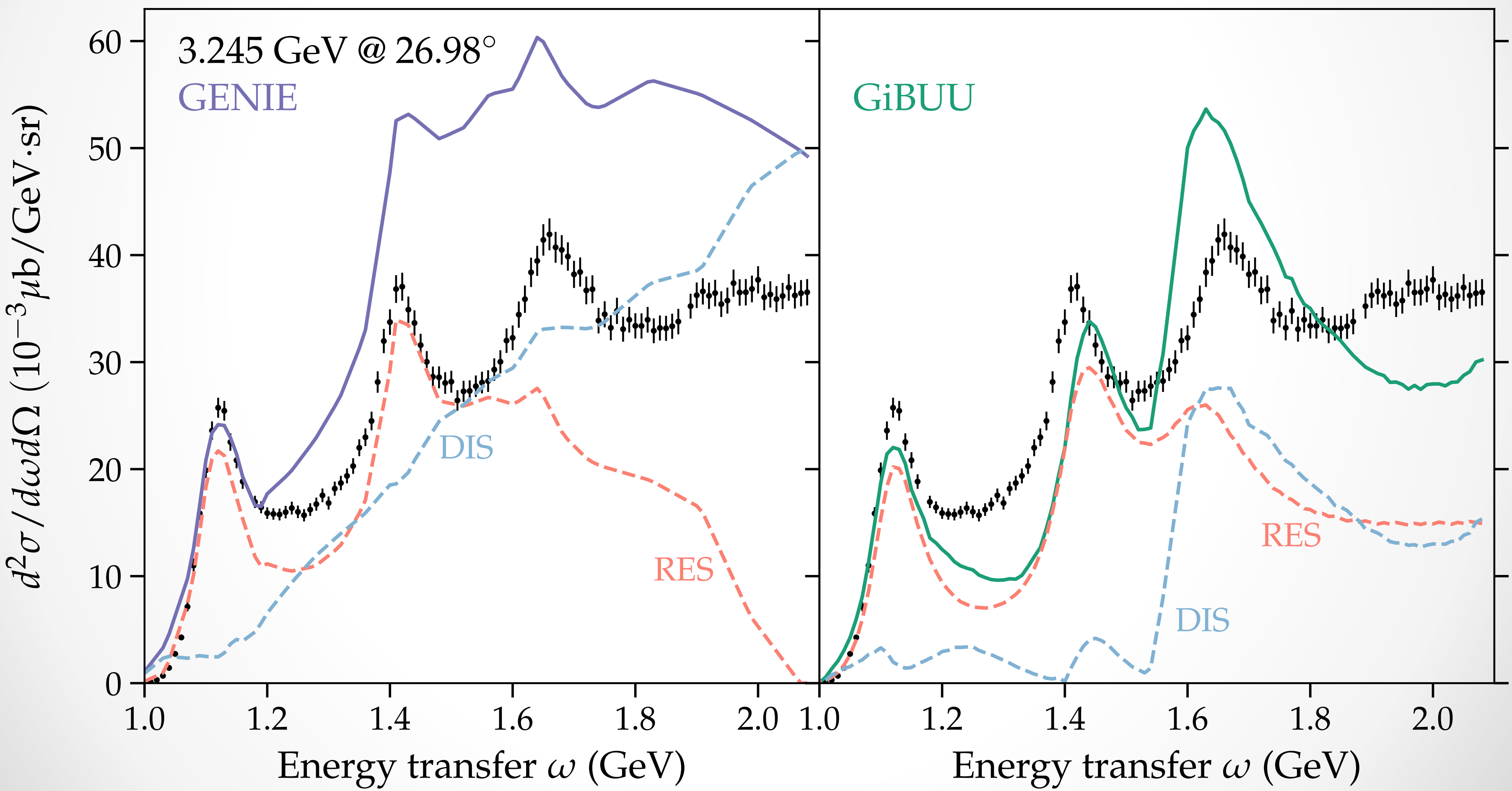
$$\omega = E_e - E_{e'}$$

Generator Treatment

	GENIE	GiBUU
RES	relativistic quark model	fit for form factors from single-pion prod. data $\langle\Delta\|J^\mu\|N\rangle = \bar{u}^\alpha_\Delta[\gamma_5(\not{q} g^{\alpha\mu} - q^\alpha\gamma^\mu)G_1 + \gamma_5(q^\alpha p'^\alpha - q\cdot p' g^{\alpha\mu})G_2 + \gamma_5(q^\alpha q^\mu - q^2 g^{\alpha\mu})G_3]u_N$
DIS	PDF modified from GV98 that aims to work down to $Q^2 \to 0$	PYTHIA

e-*p* Scattering

Neutrino Generators vs. Data

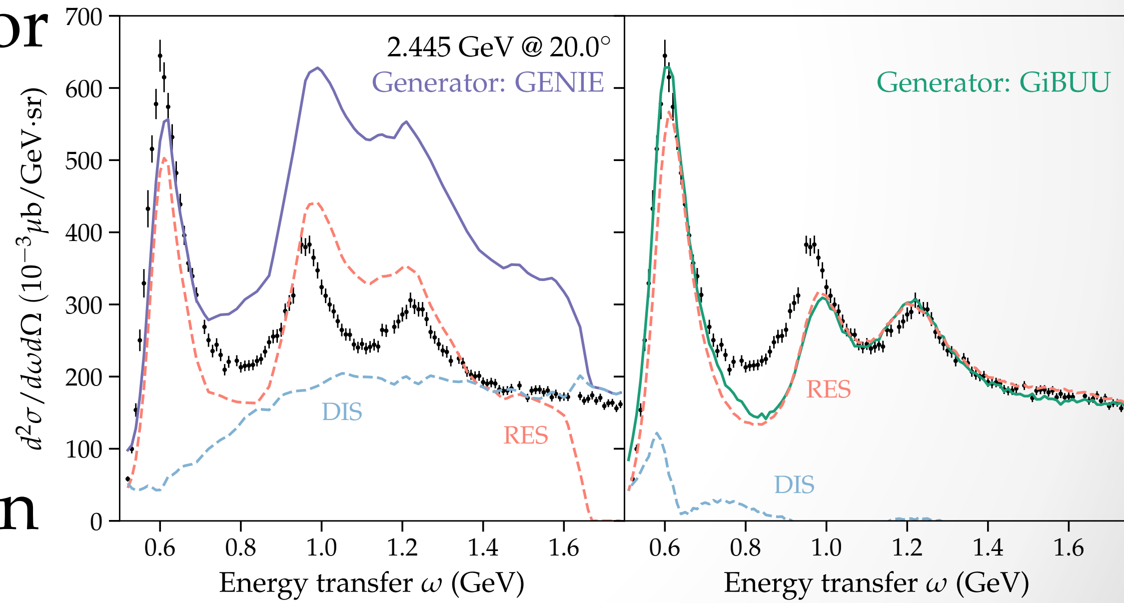


Ankowski, Friedland, SL, in prep
Data from Niculescu et al., 00

e-*p* Scattering

Neutrino Generators vs. Data

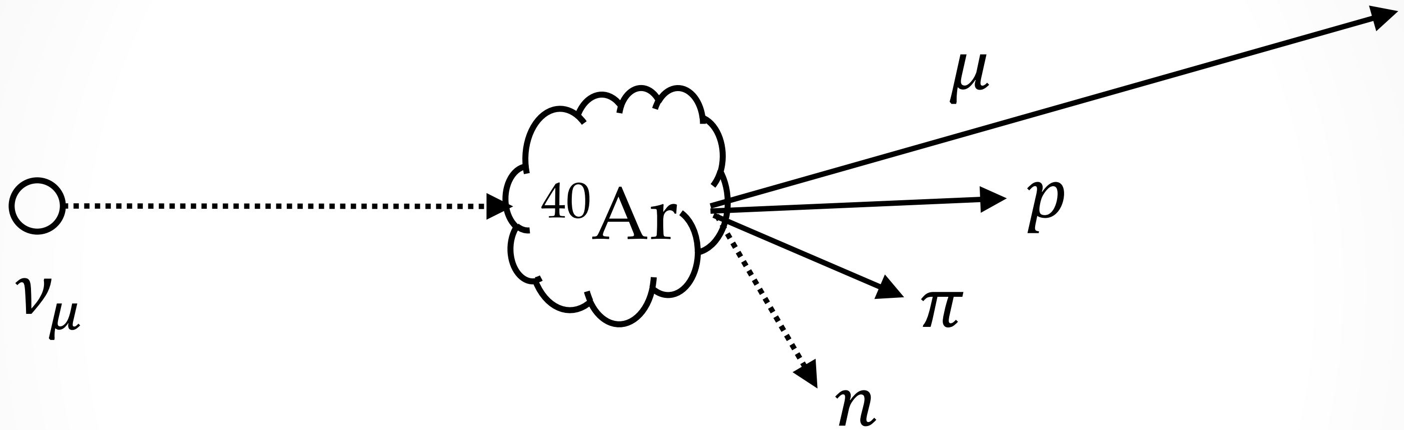
- GENIE: 50—80% error
  GiBUU: 20—30% error
- RES vs. DIS: Not a Good Separation
- Major Implementation Errors



Ankowski, Friedland, SL, in prep
Data from Niculescu et al., 00

Cross Section Calculations

Observables Are High-Dimensional



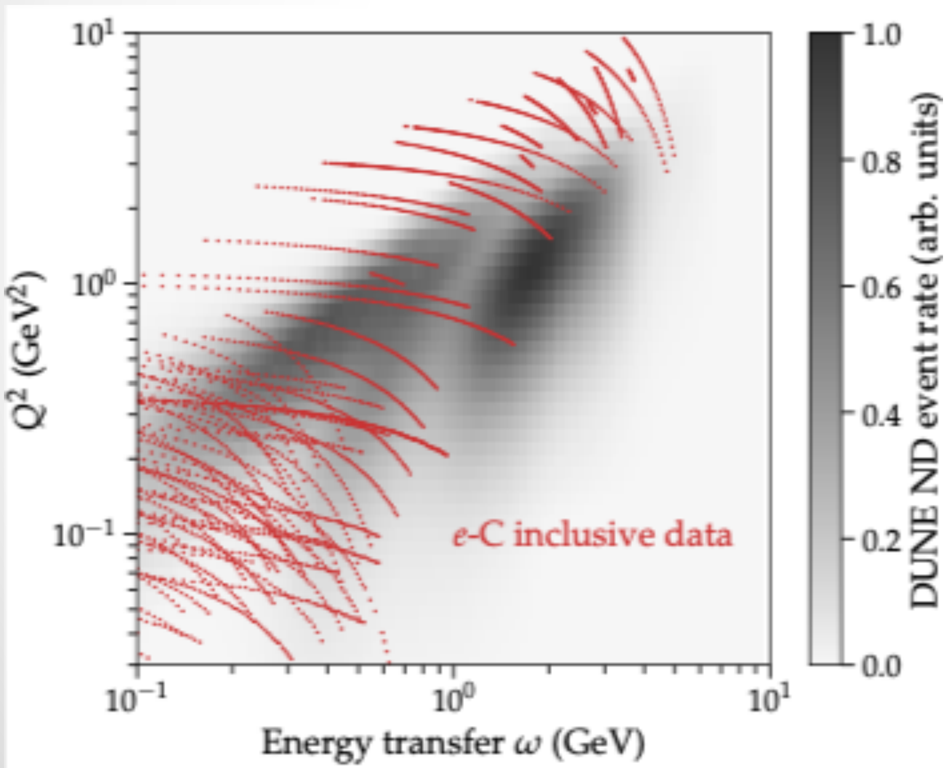
$$\frac{\mathrm{d}\sigma}{\mathrm{d}E_1 \mathrm{d}E_2 \ldots \mathrm{d}E_n} \text{ (Exclusive), not } \frac{\mathrm{d}\sigma}{\mathrm{d}E} \text{ (Inclusive)}$$

Much Harder to Compute Exclusive Cross Sections

More Data Is Needed!

Ankowski et al., 19

Existing Coverage Is Poor

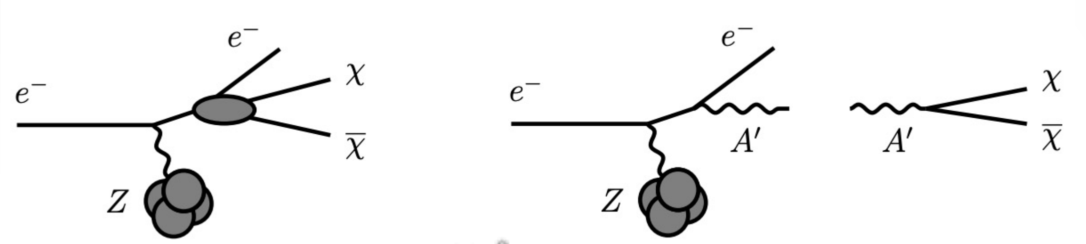


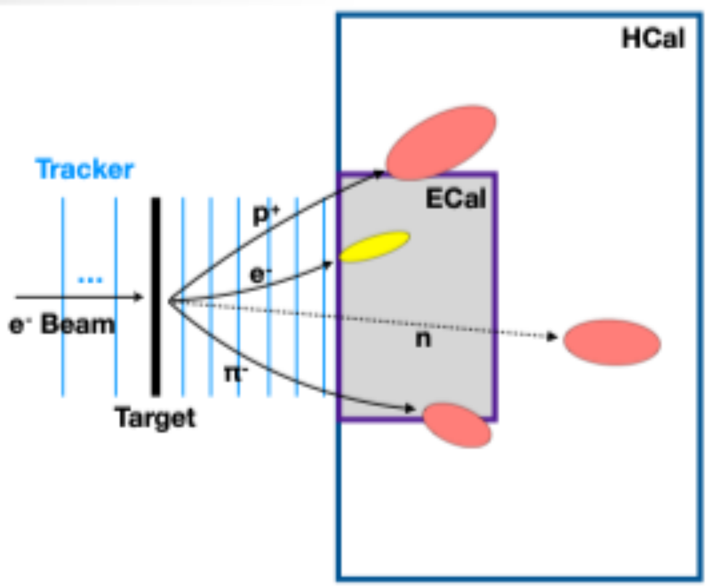
- Only 1 Set of Argon Data
- Little Coverage of Phase Space Important for DUNE
- Only Inclusive Data

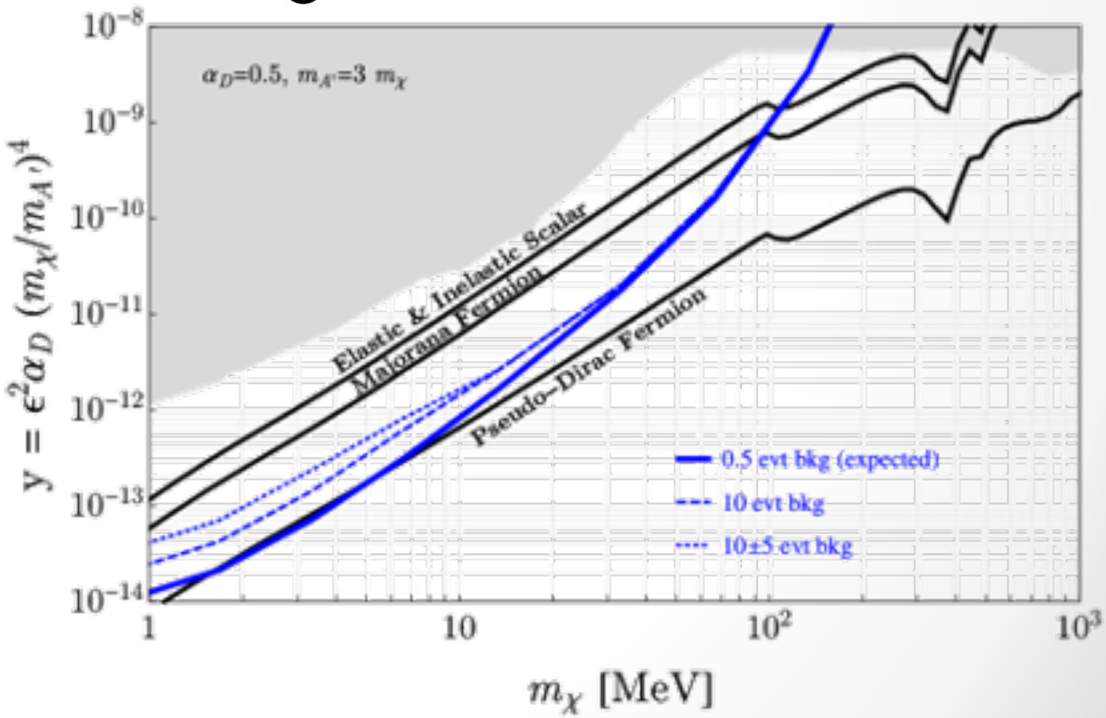
LDMX

Ankowski et al., 19

Light Dark Matter eXperiment





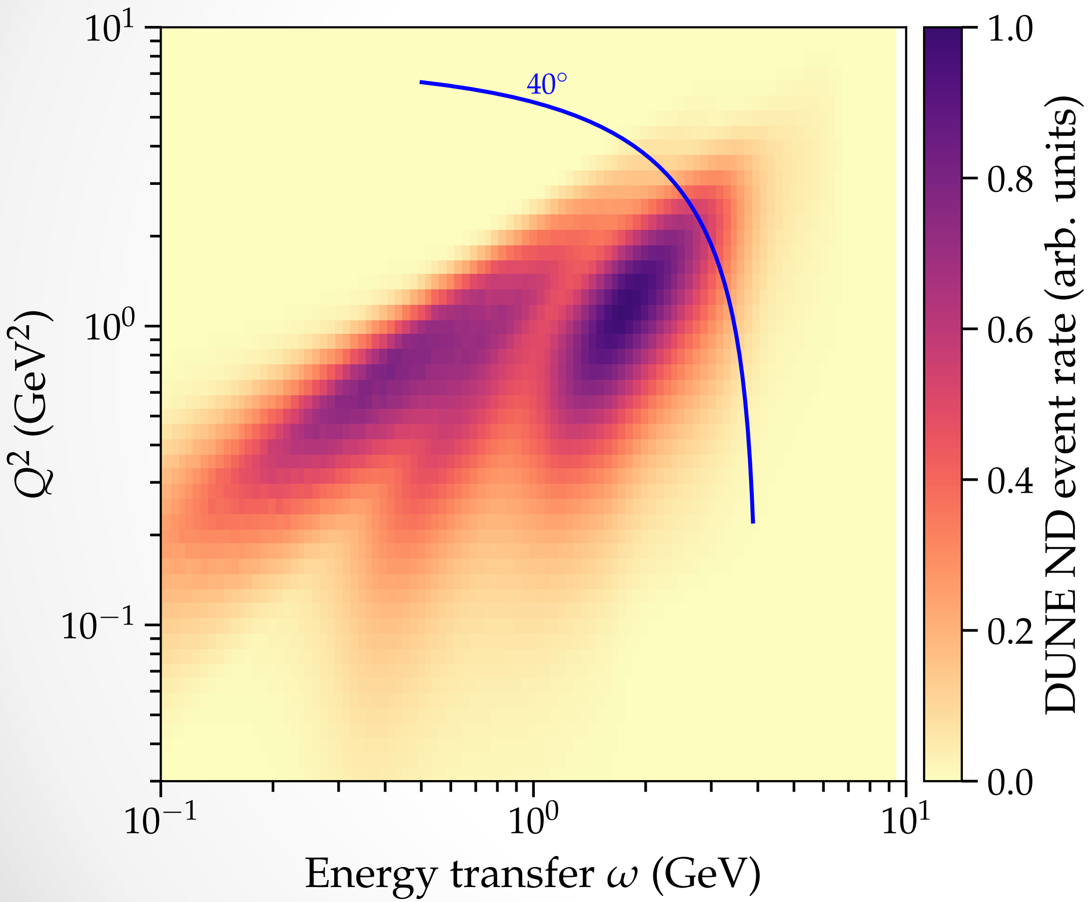


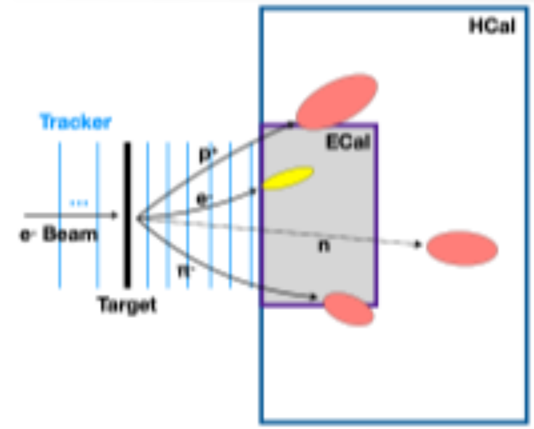
Planning Stage Funded!

LDMX, 19

Important Design Features

Ankowski et al., 19
Also: CLAS12





- Good Coverage
- Different Nuclei
- All Outgoing Particles in the Forward Region

Large Overlap with DUNE Phase Space

Conclusions

1. GeV Neutrino-Nucleus Scattering is Crucial to the Success of Long-Baseline Neutrino Experiments

2. No Complete Theoretical Framework Available; Difficult to Assess Uncertainties

3. More Scattering Data is Needed

4. New Theoretical Ideas Are Needed

Thank you

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