

# Overview of Classical Machine Learning in HEP

Benjamin Nachman

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[bpnachman.com](http://bpnachman.com)

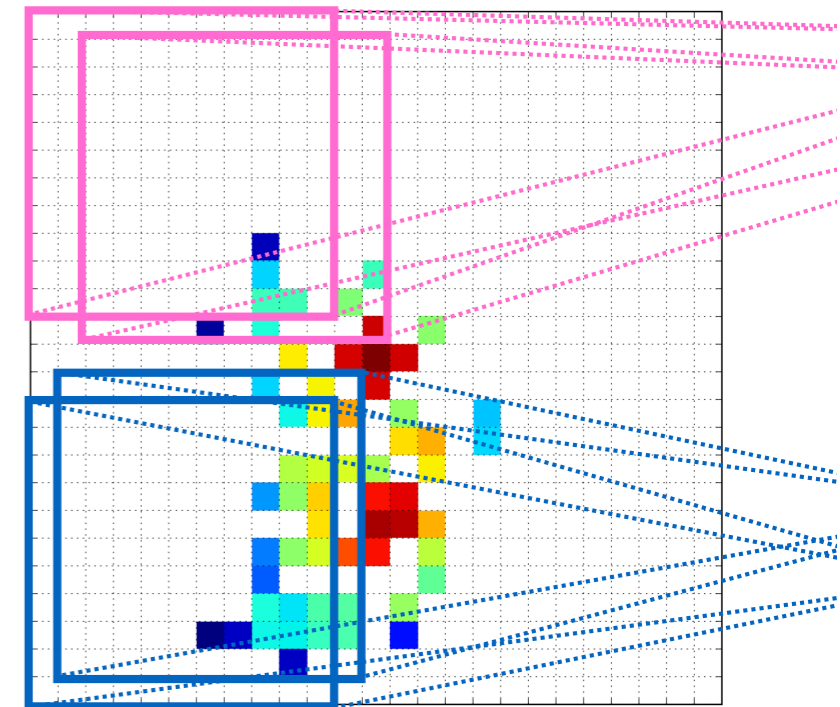
[bpnachman@lbl.gov](mailto:bpnachman@lbl.gov)



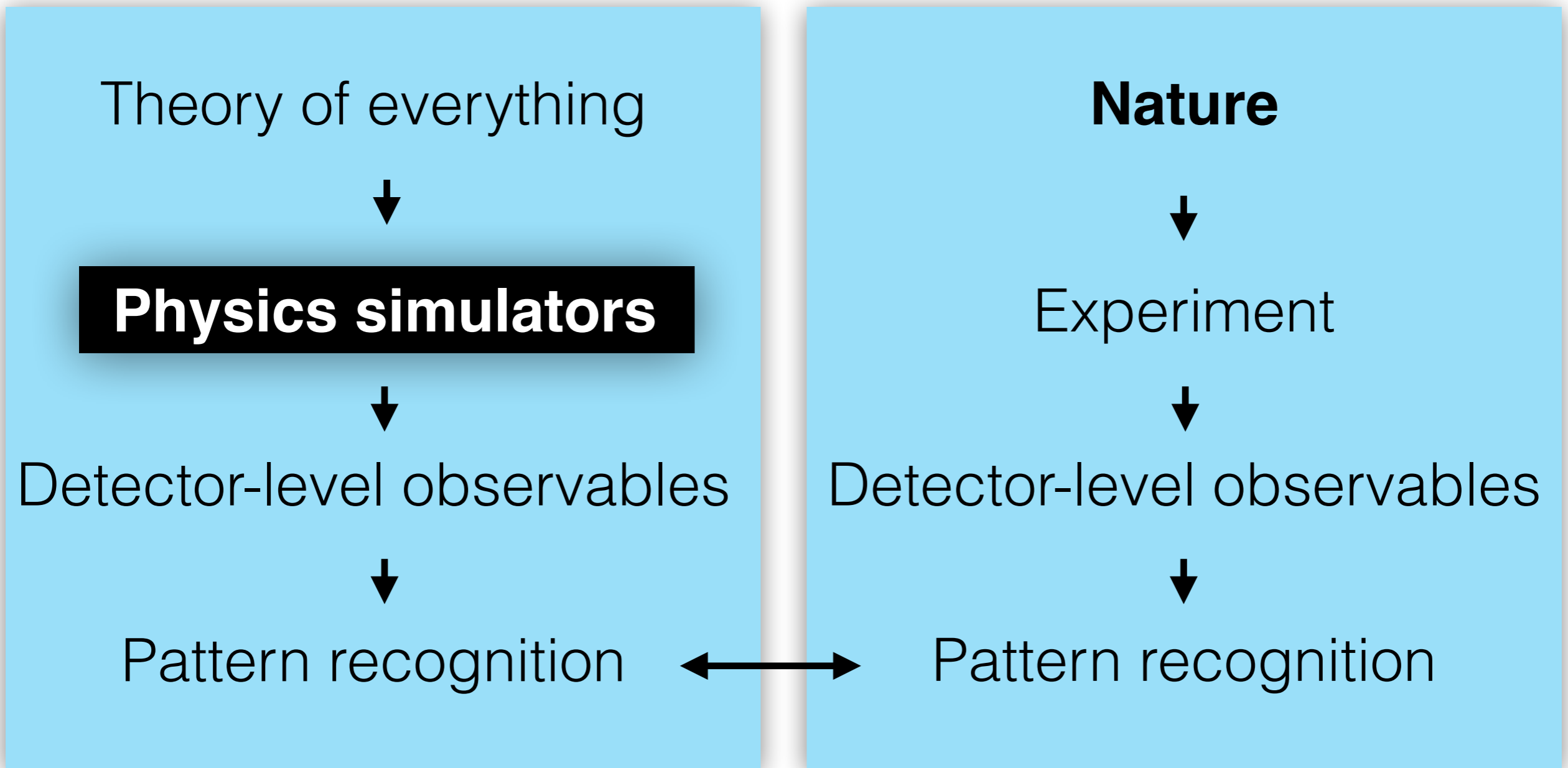
@bpnachman



bnachman

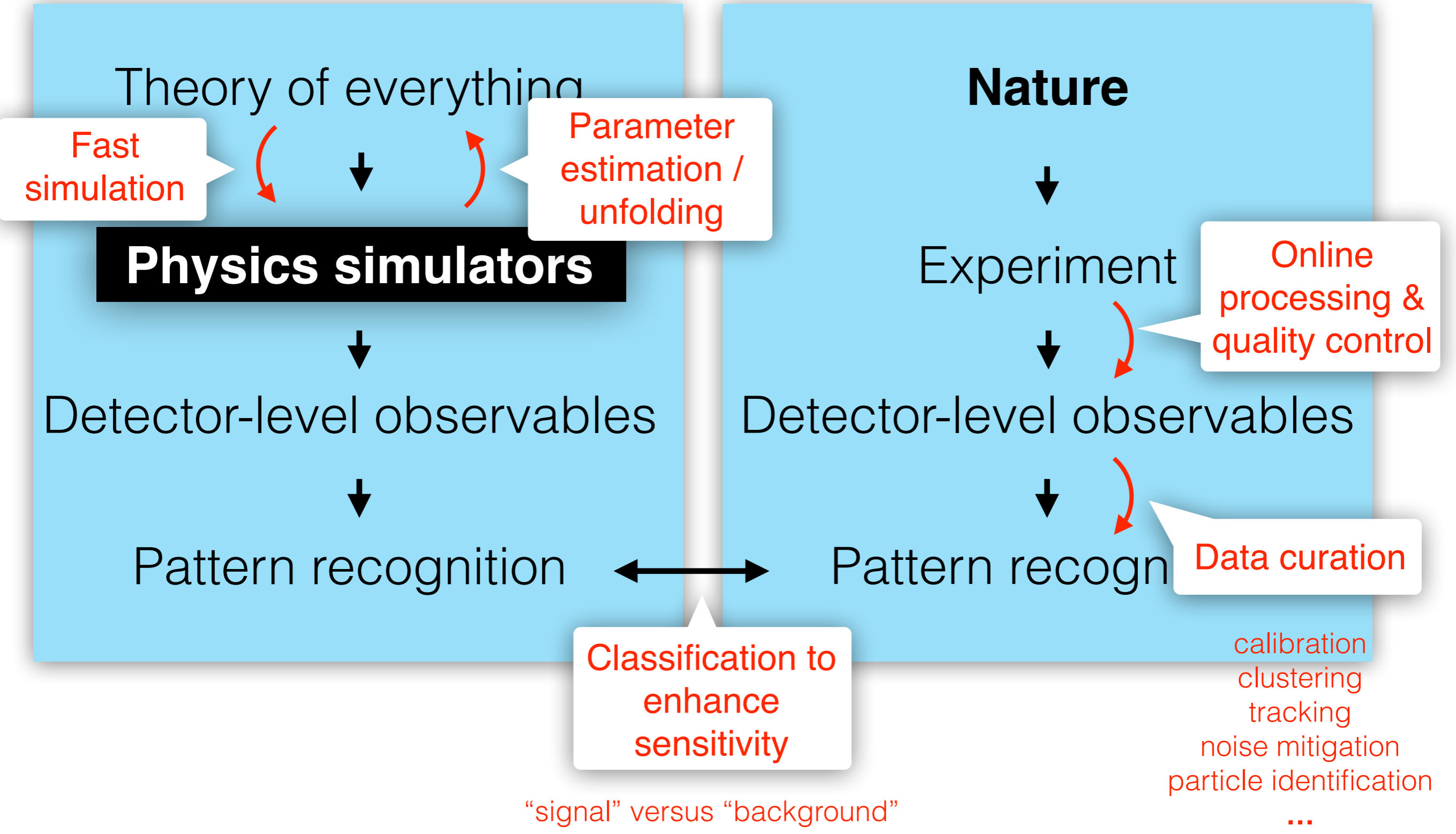


QC/QML for HEP  
June 2022

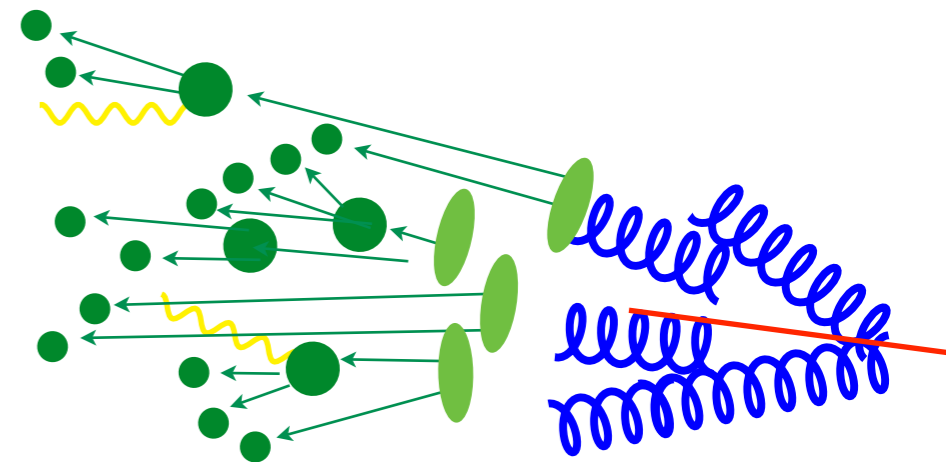




# Data analysis in HEP + Machine Learning



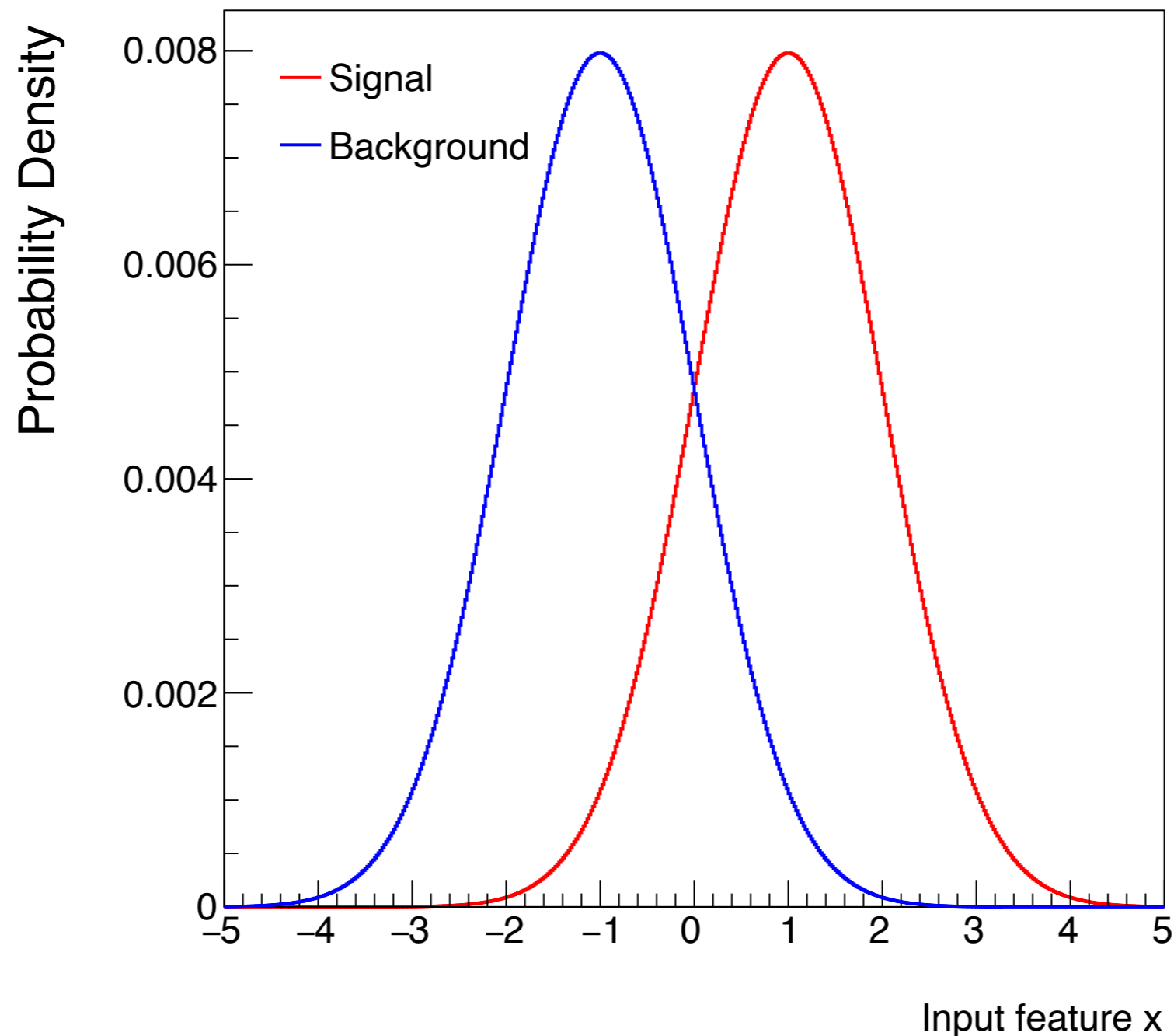
- Classification
  - Machine Learning and Optimality
  - HEP images
  - Other architectures for HEP
- Regression, Generative Models / likelihood-free approaches, Anomaly Detection



# Machine learning and optimality

5

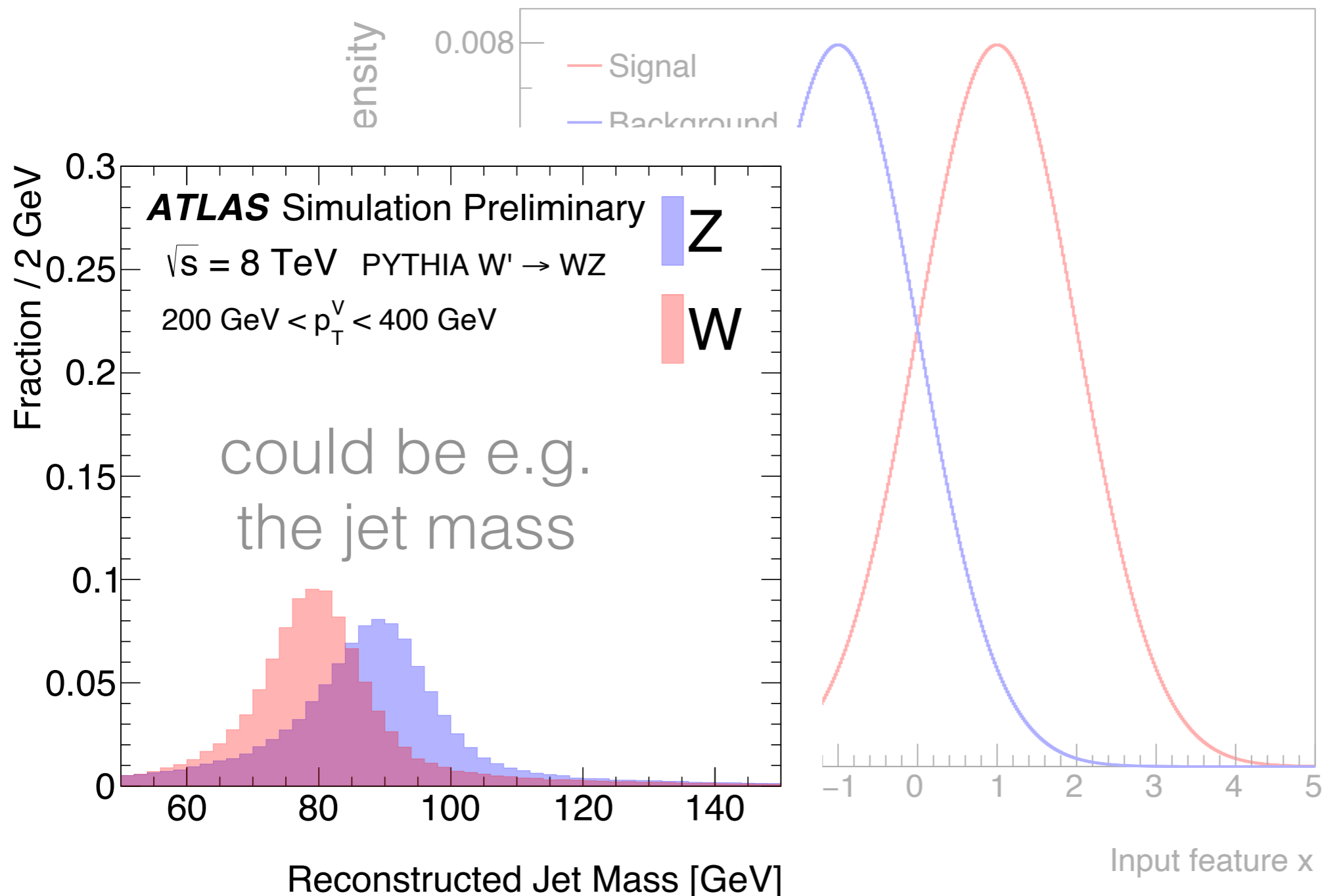
Let's consider an important special case:  
binary classification in 1D



# Machine learning and optimality



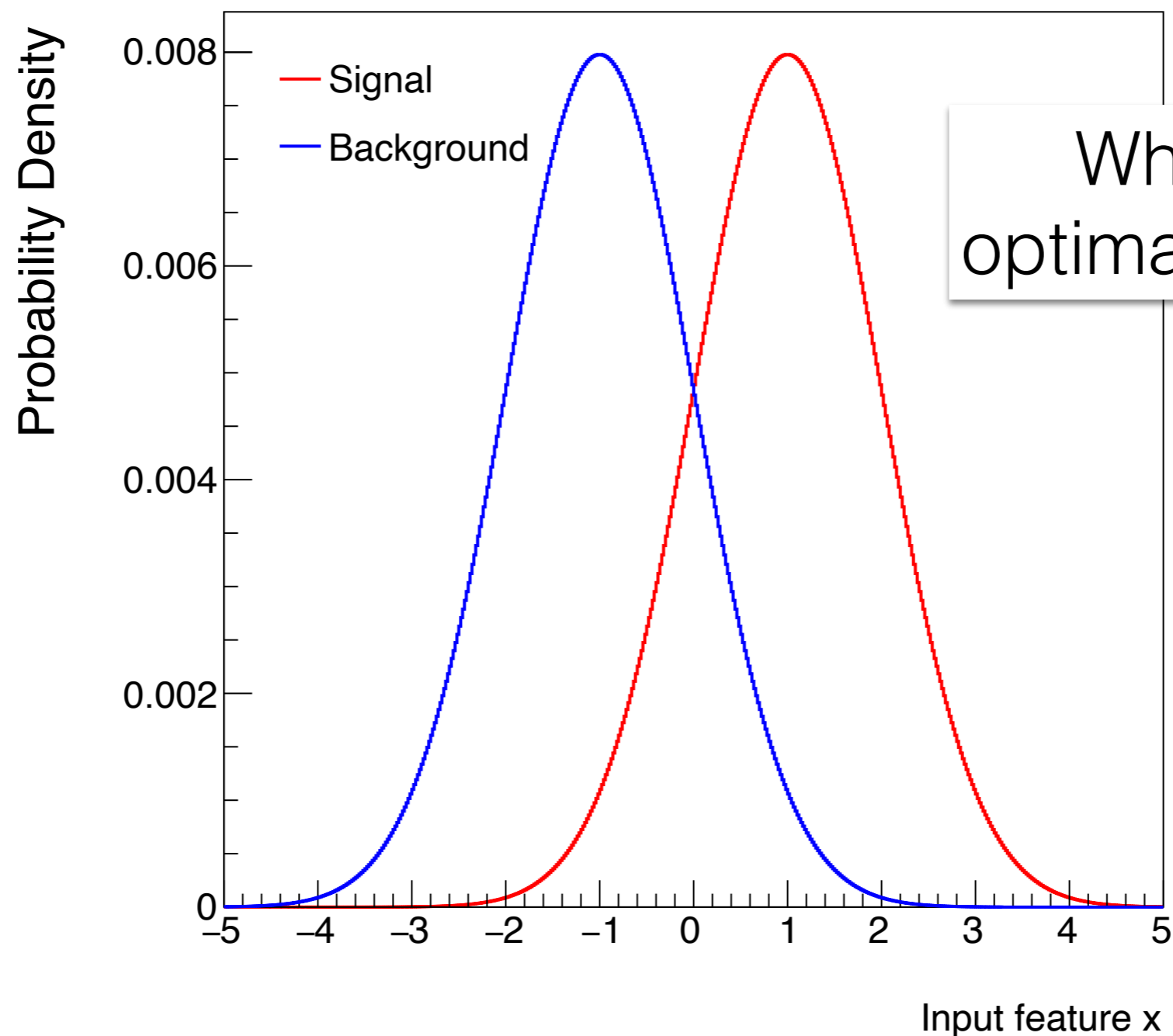
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# Machine learning and optimality



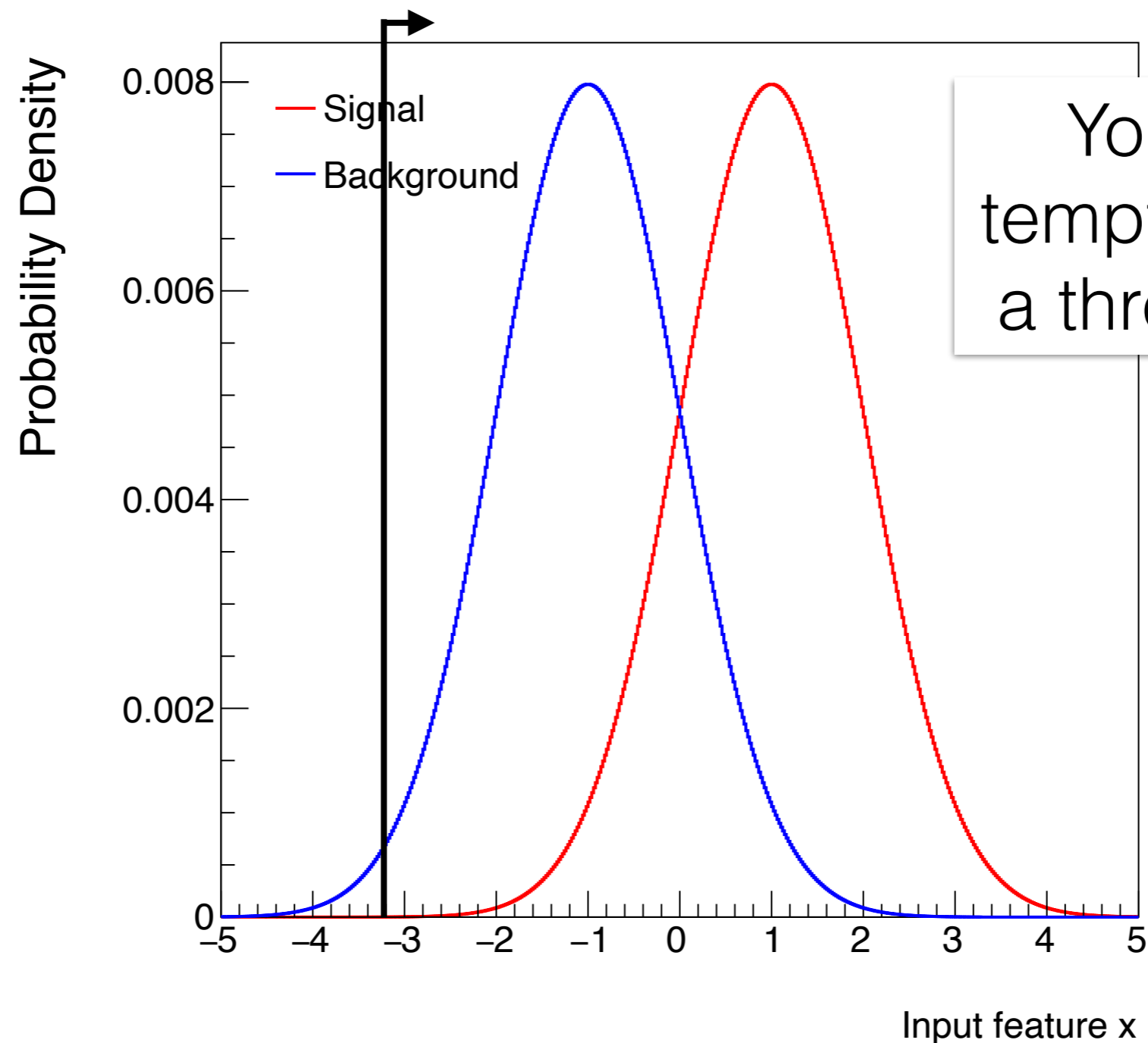
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# Machine learning and optimality



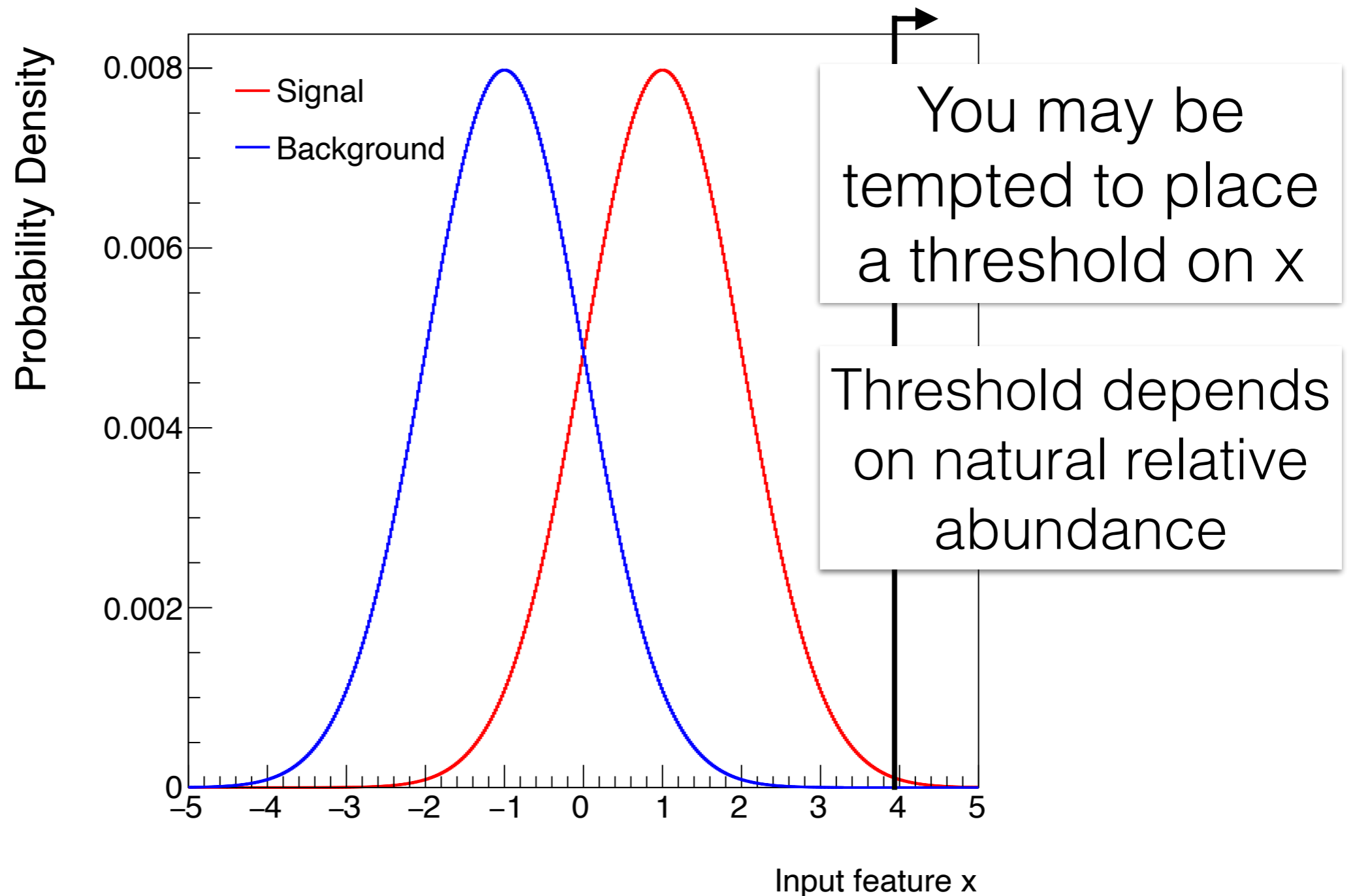
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# Machine learning and optimality

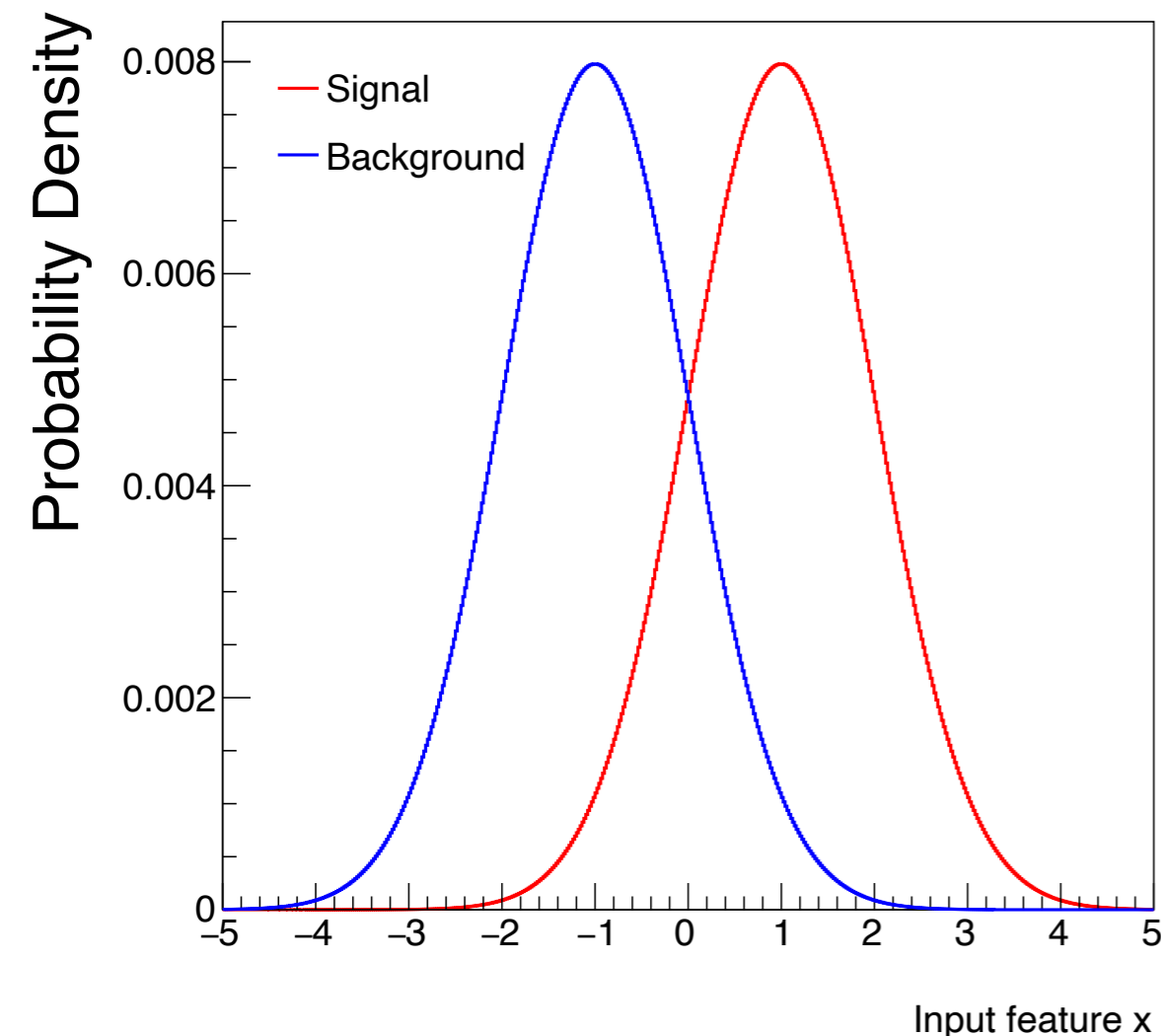


Let's consider an important special case:  
binary classification in 1D



# Machine learning and optimality

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Is the simple threshold cut **optimal**?

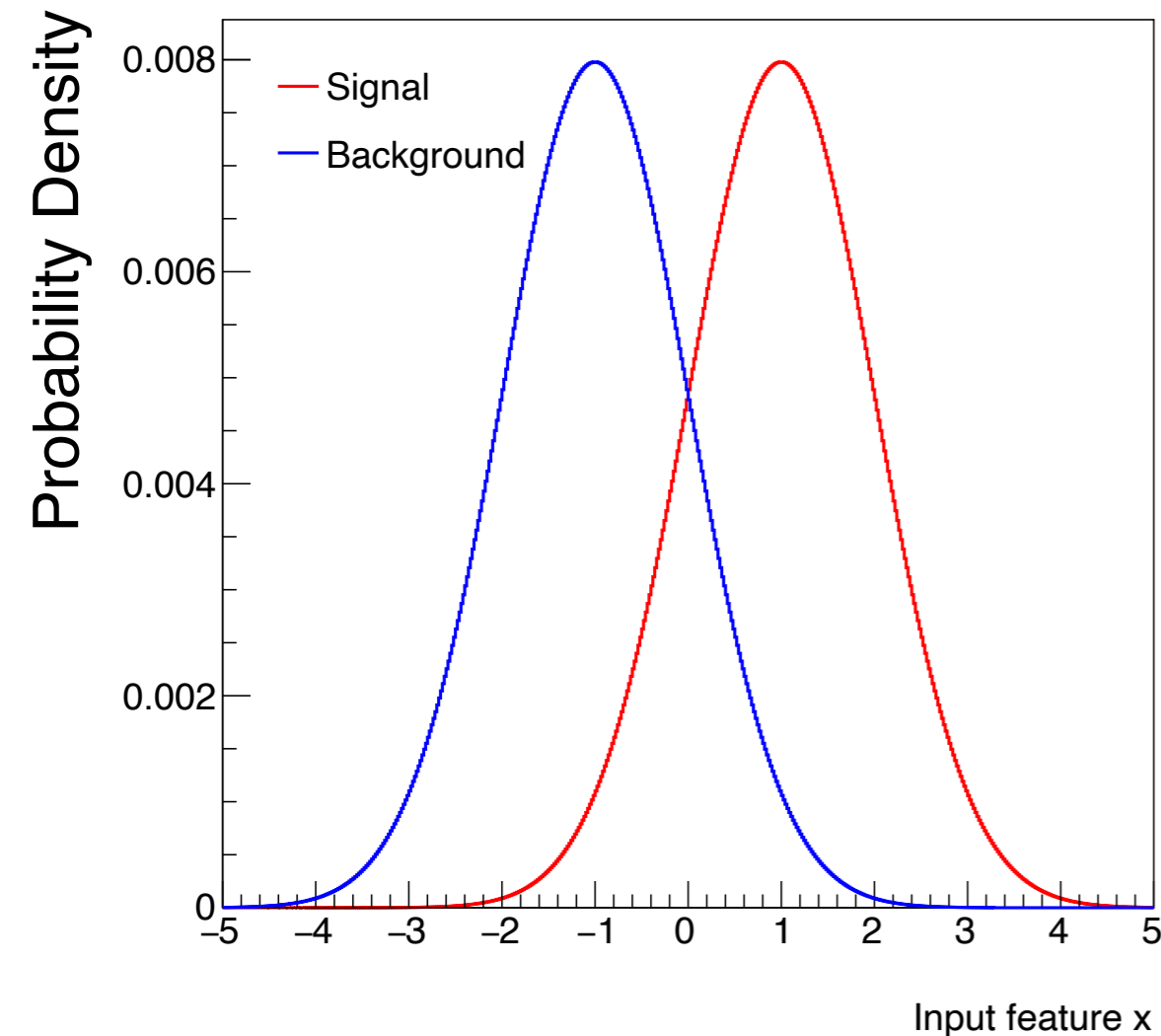
TPR = true positive rate or  
“signal efficiency”  
FPR = false positive rate or  
1 - “background rejection”

*See Neyman-Pearson lemma*

Fact 1: The classifier that results in the lowest FPR for a given TPR is a cut on the **likelihood ratio (LR)**.

$$LR(x) > c, \quad LR(x) = p(x|\text{signal}) / p(x|\text{background})$$





Is the simple threshold cut **optimal**?

TPR = true positive rate or “signal efficiency”  
FPR = false positive rate or 1 - “background rejection”

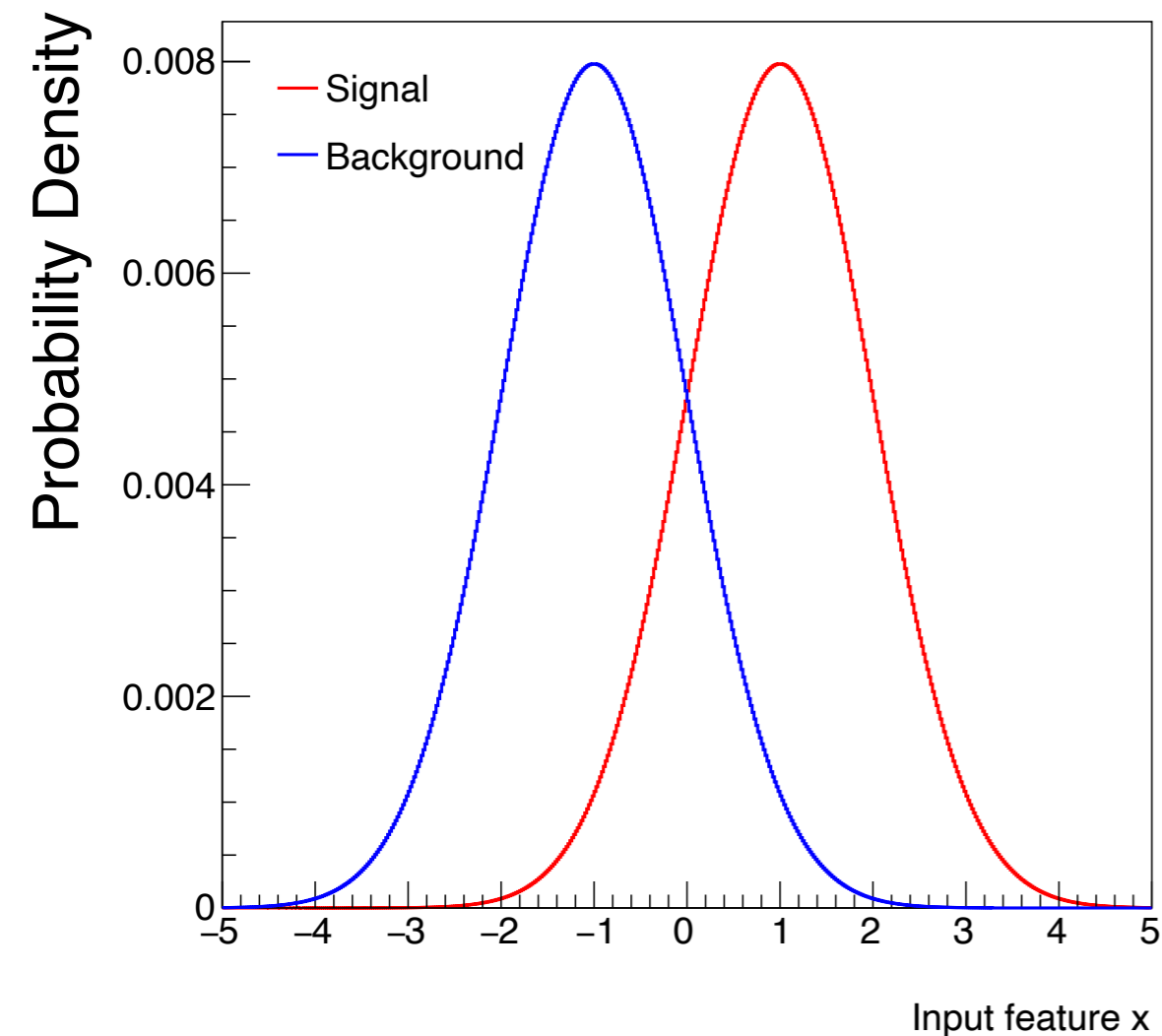
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Fact 1: The classifier that results in the lowest FPR for a given TPR is a cut on the **likelihood ratio (LR)**.

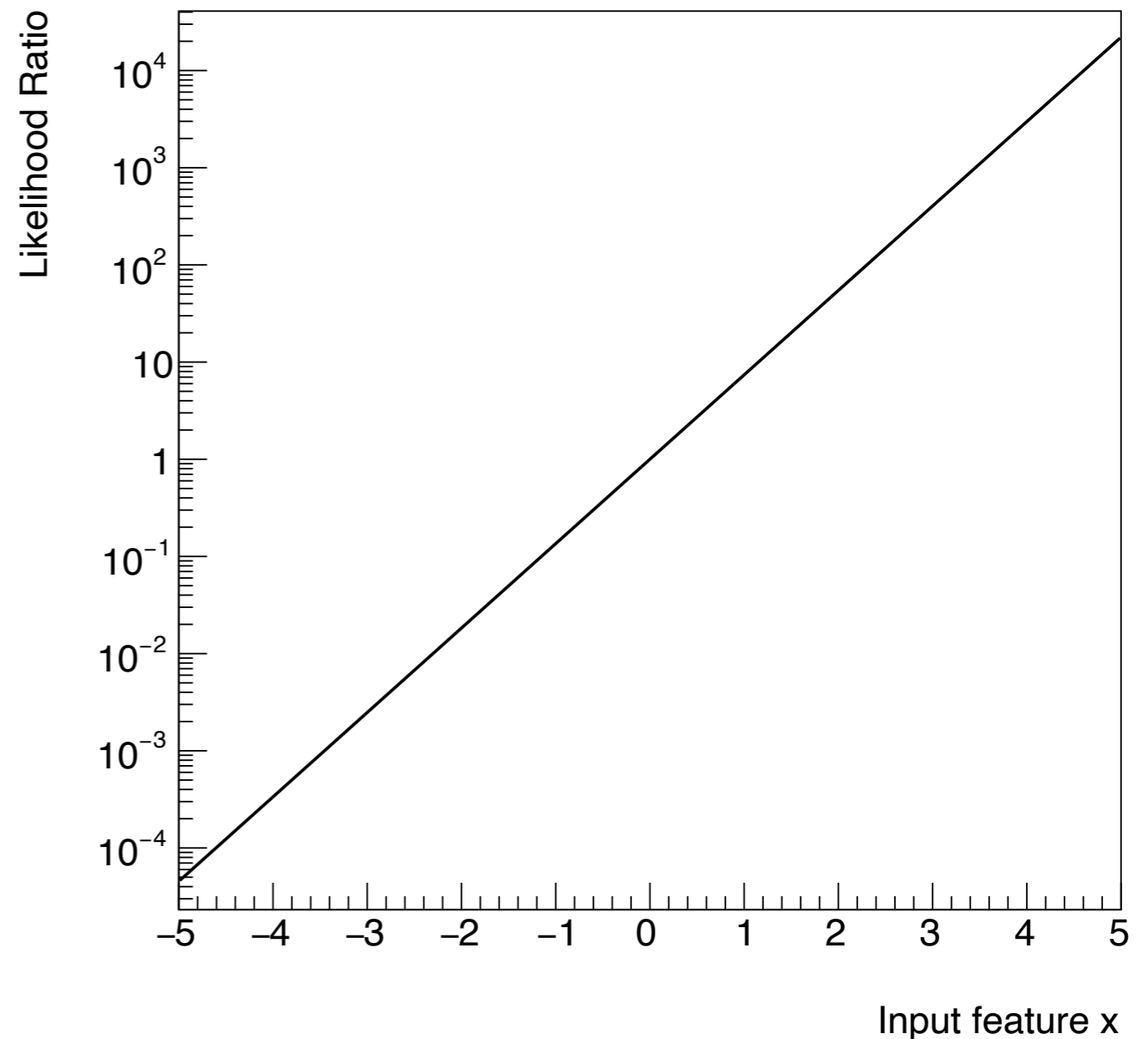
Fact 2: Two classifiers that are related by a **monotonic transformation** result in the same performance.

# Machine learning and optimality

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Is the simple threshold cut **optimal**?

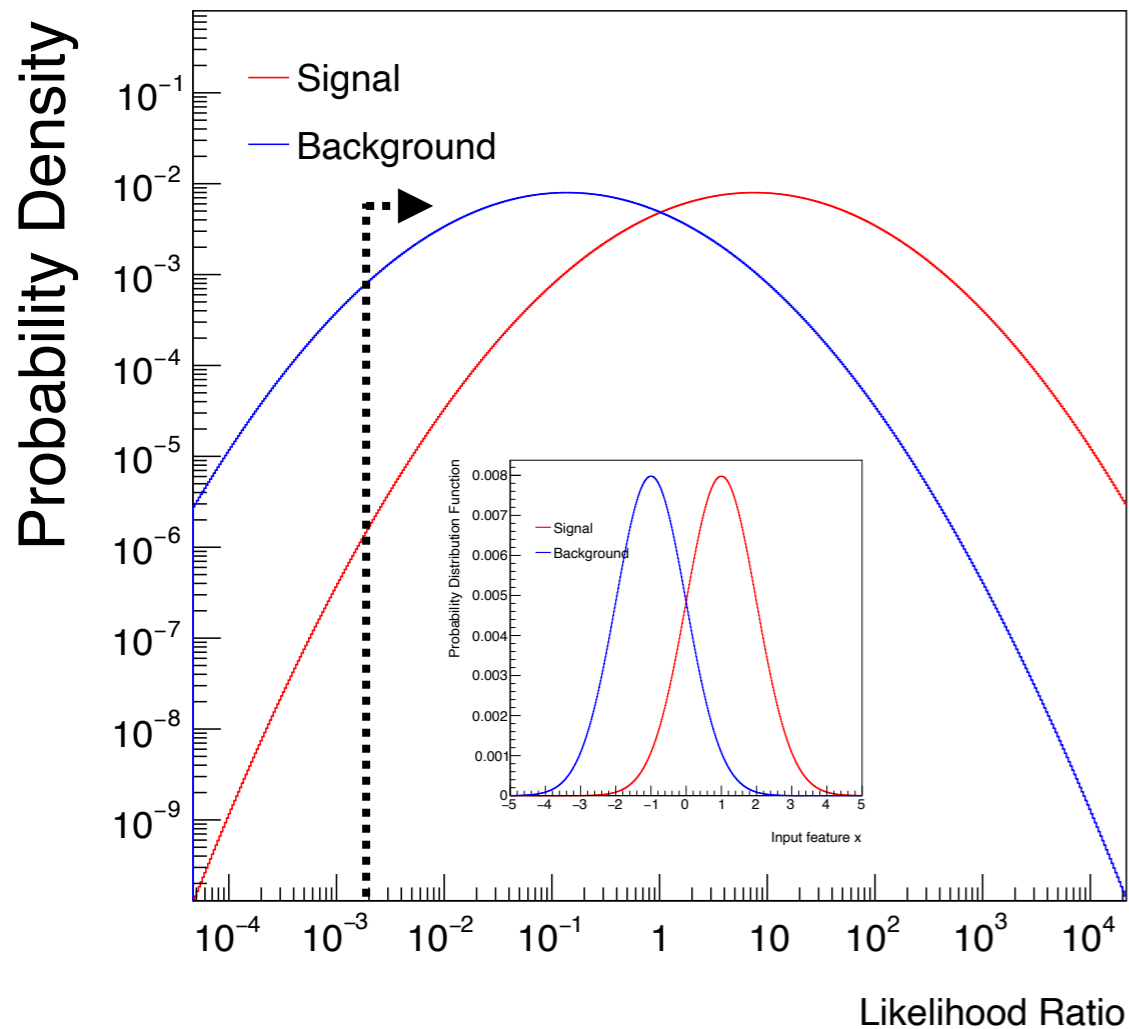


In this simple case, the log LR is proportional to x:  
**no need for non-linearities!**

*Threshold cut is optimal*

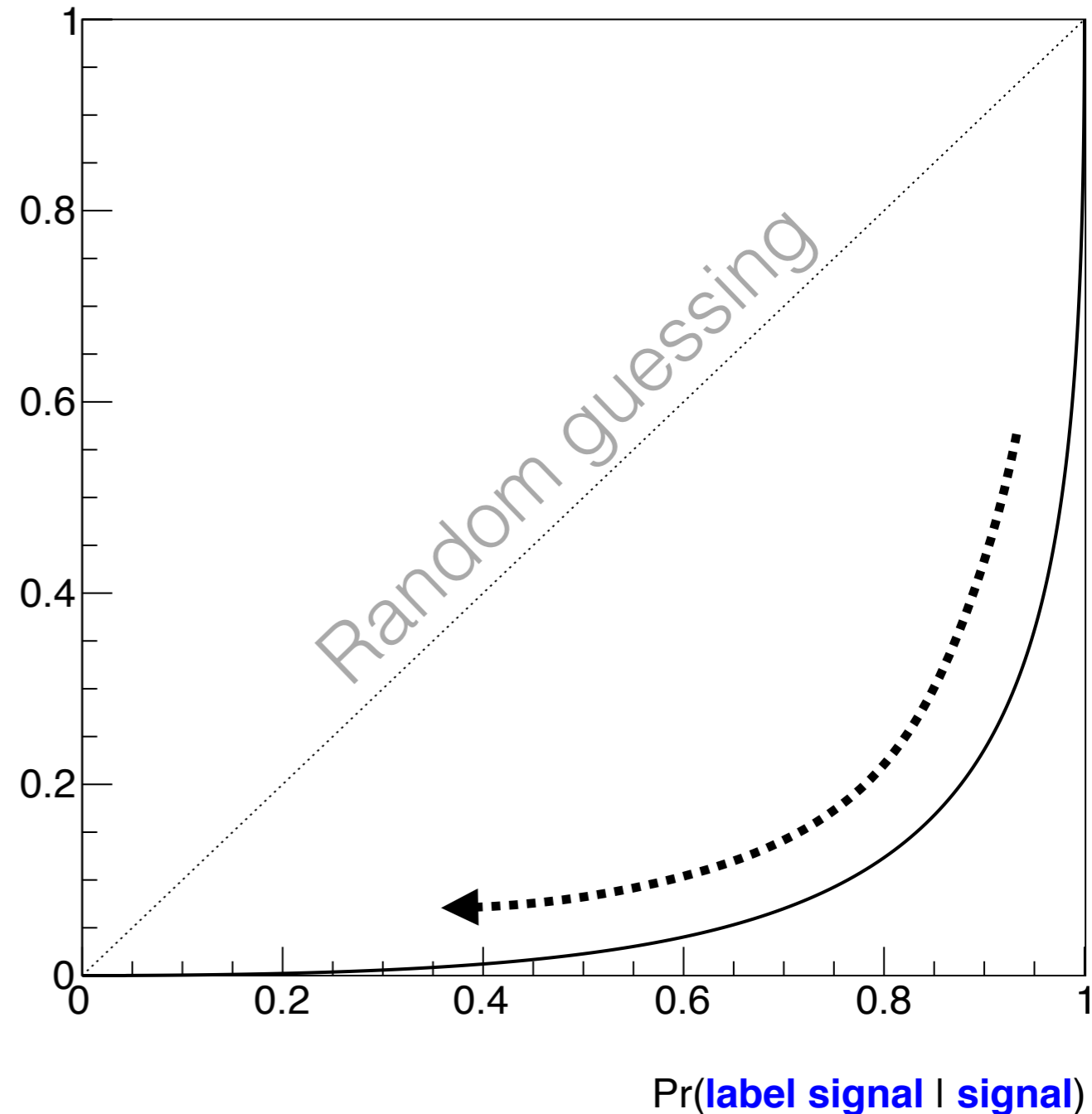
# Machine learning and optimality

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$\Pr(\text{label signal} \mid \text{background})$

The optimal procedure is a threshold on the LR



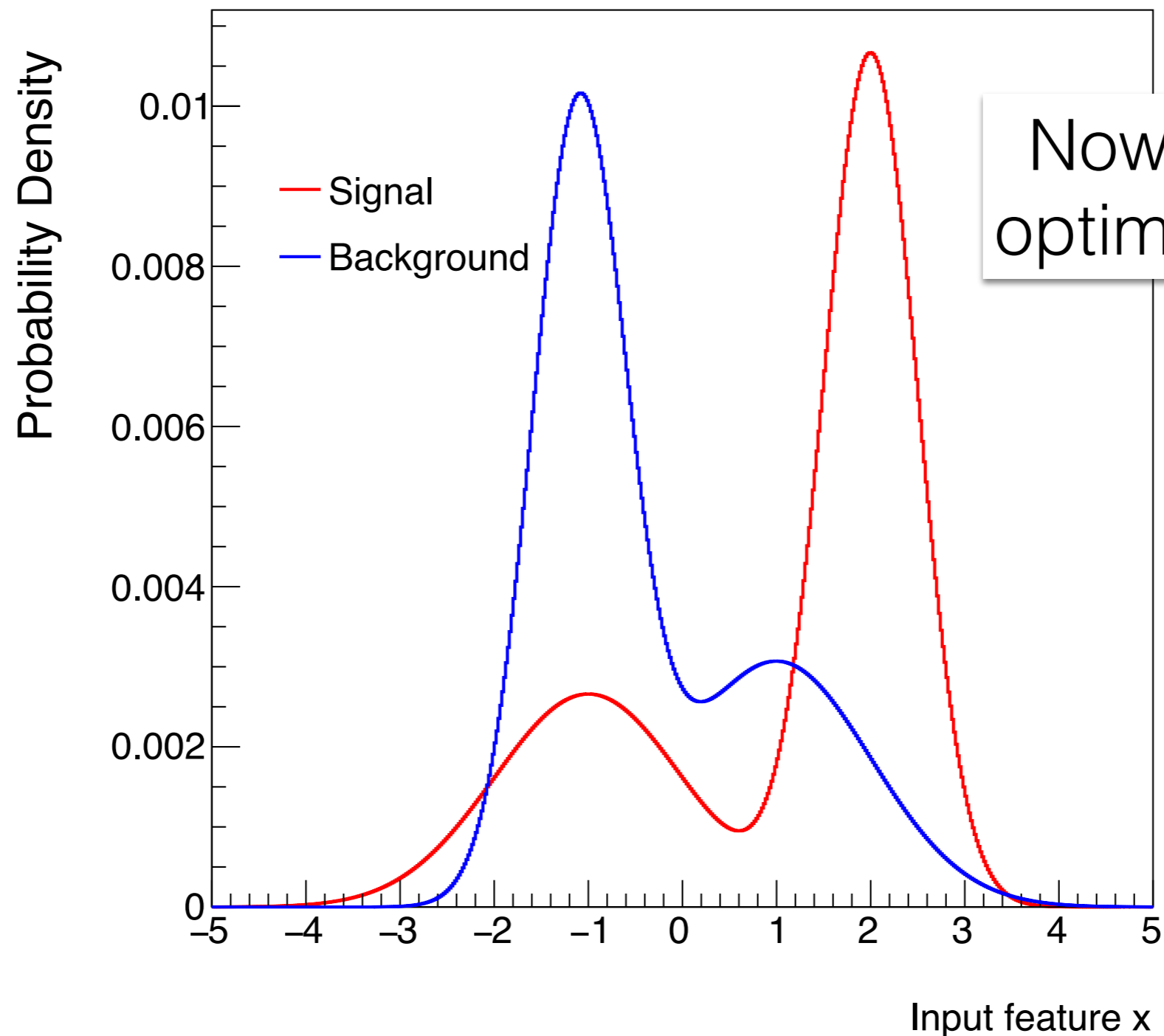
“Receiver Operating Characteristic” (**ROC**) Curve

# Machine learning and optimality

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What if the distribution of  $x$  is complicated?

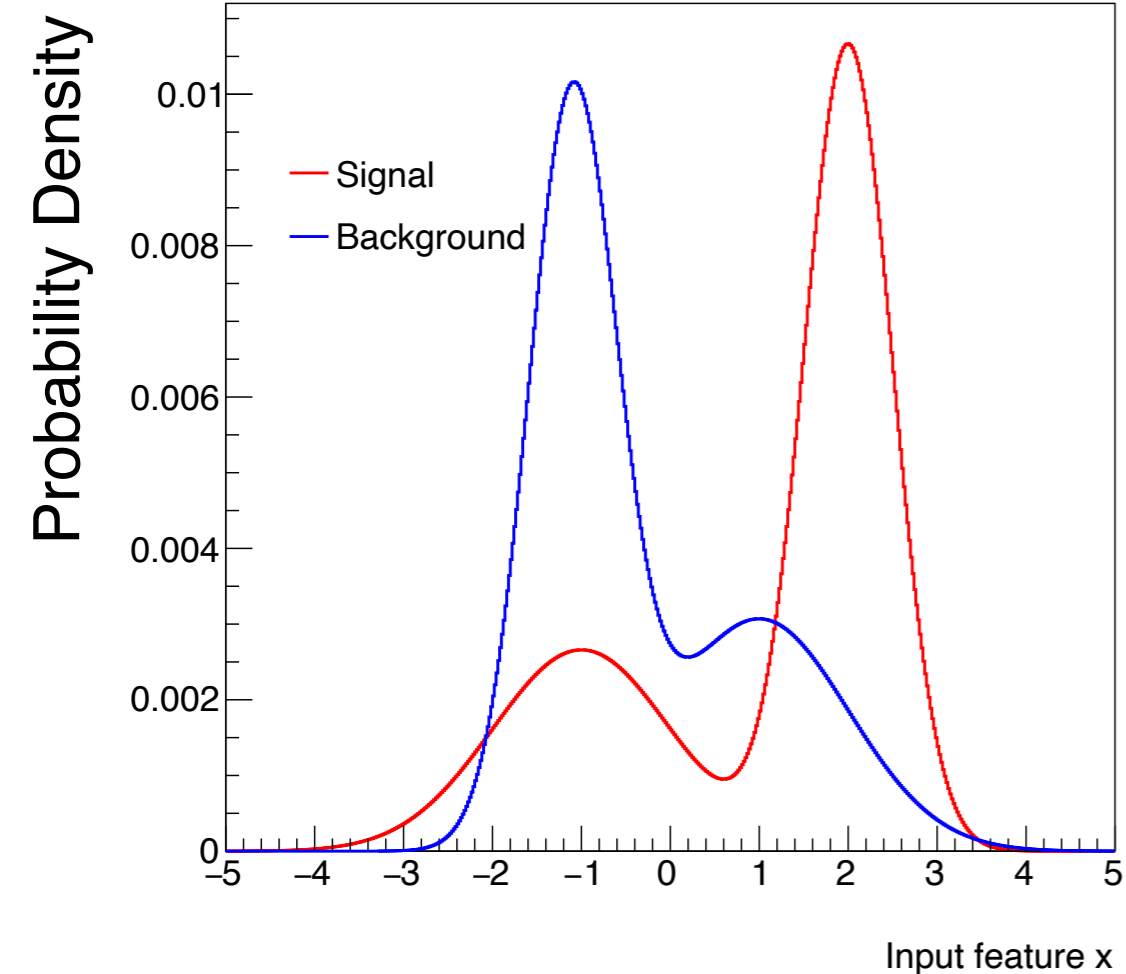
Real life is complicated!



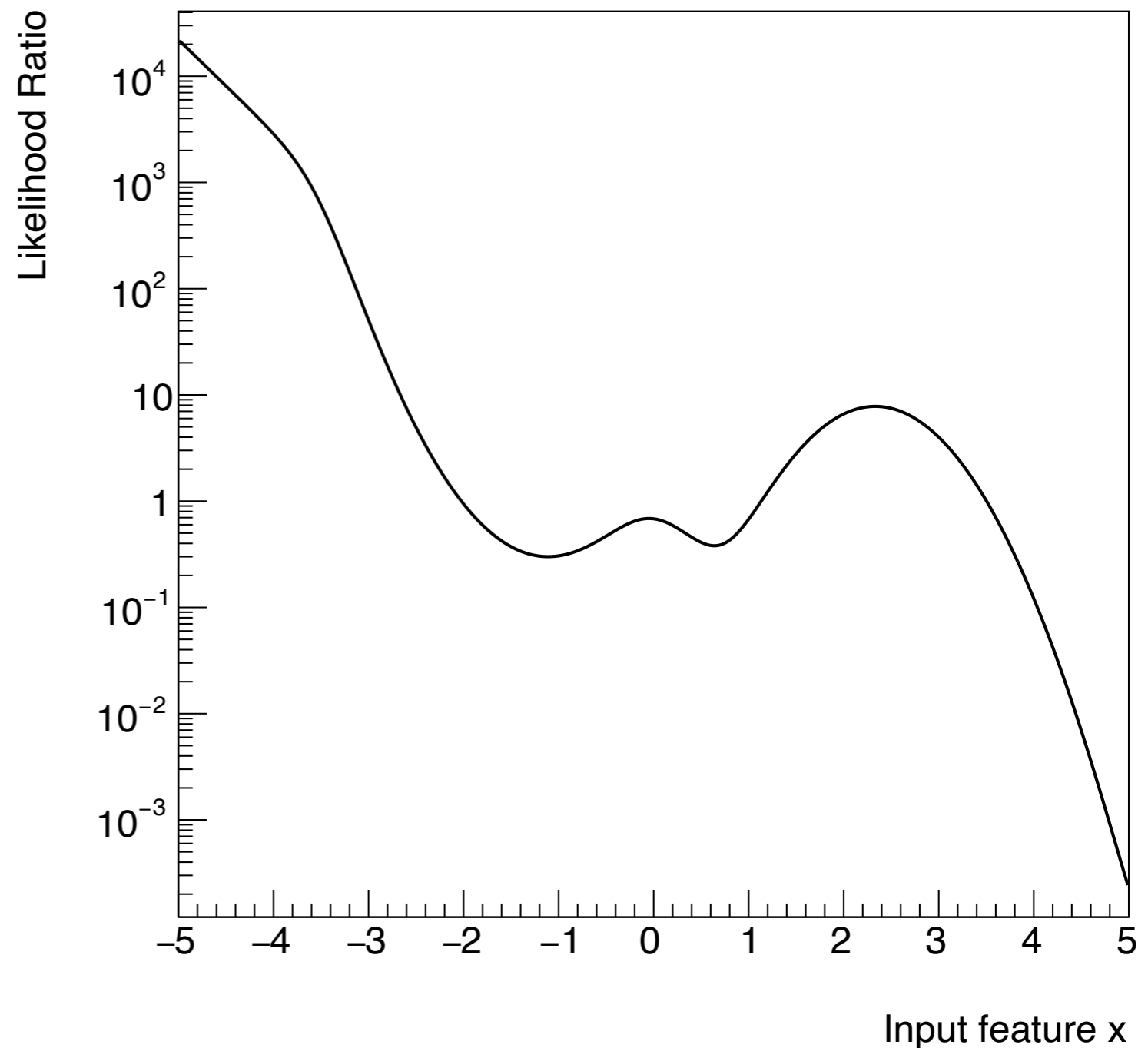
Now what is the optimal classifier?

# Machine learning and optimality

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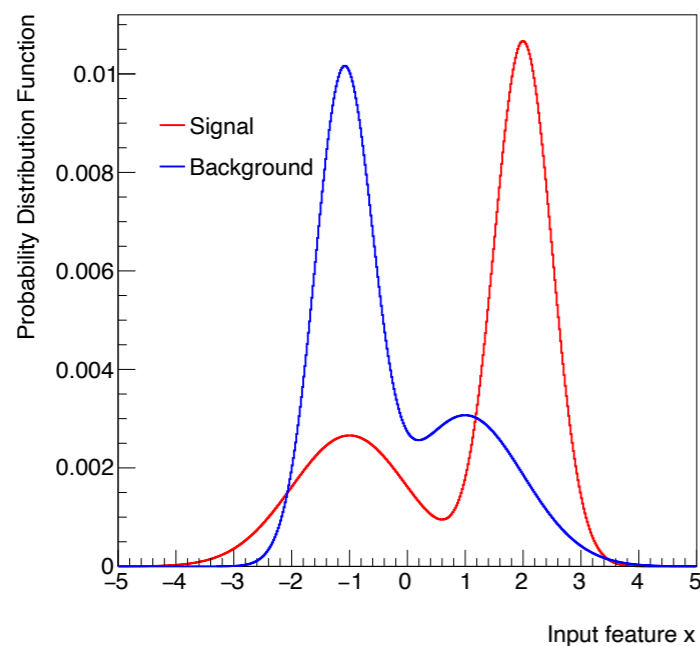
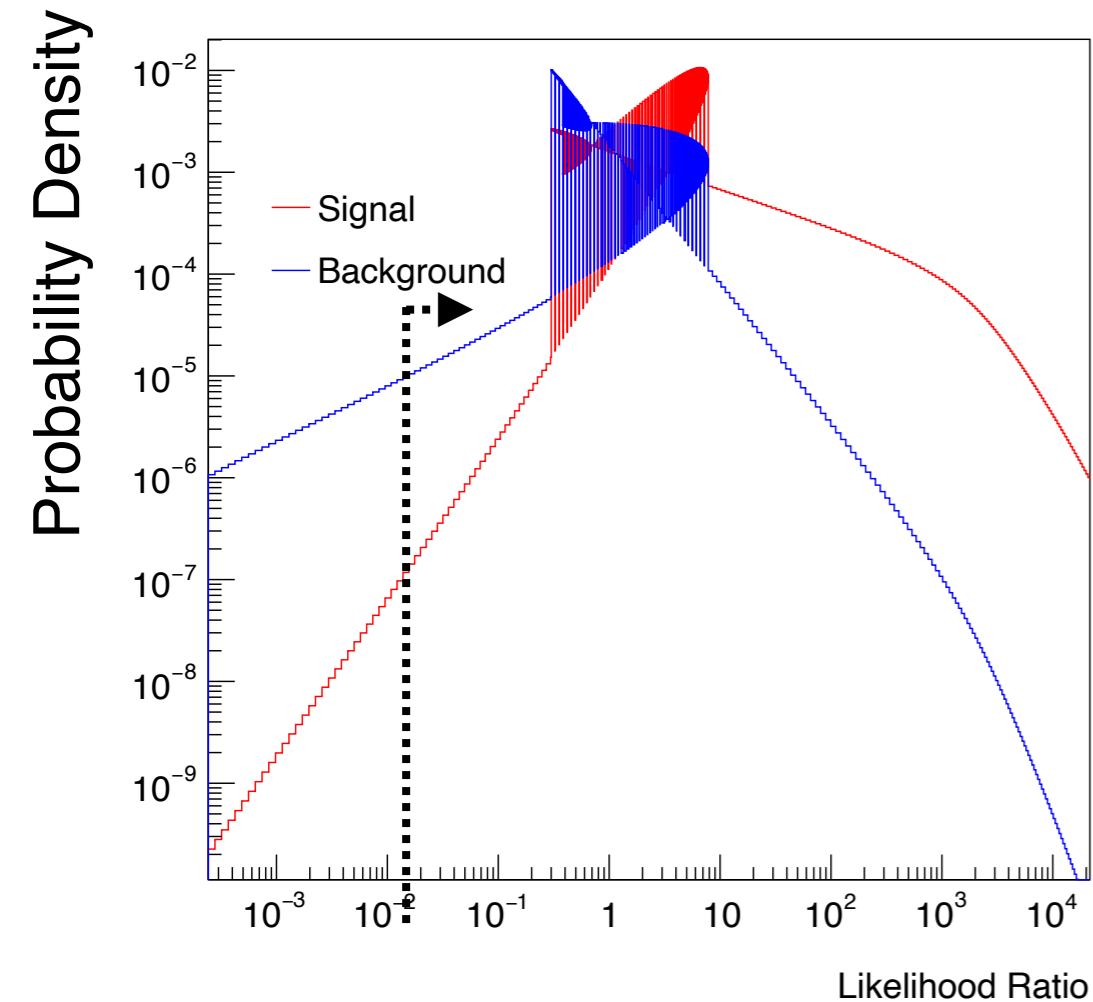
In this case, LR is highly non-linear (**non-monotonic**) function of  $x$



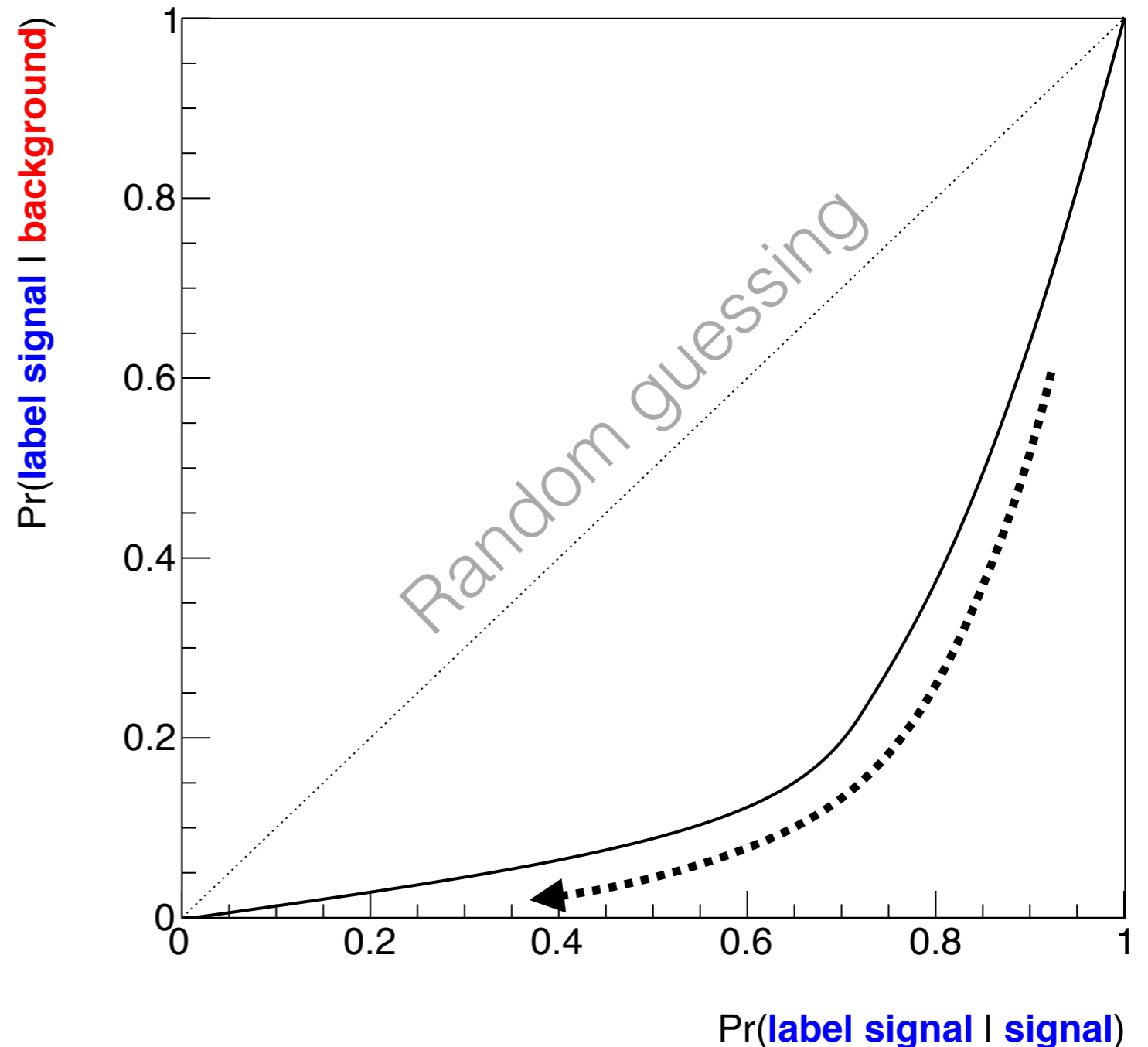
A threshold on  $x$   
would be sub-optimal

# Machine learning and optimality

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ROC worse than the Gaussians, as expected since more PDF overlap



Why don't we always just compute the optimal classifier?

In the last slides, we had to estimate the likelihood ratio - this required binning the PDF  
binning works in 1D, but intractable as feature dimension  $\gg 1$  ("curse of dimensionality")

machine learning for classification is simply  
**the art of estimating the likelihood ratio  
with limited training examples**

# HEP Tools for (Classical) Classification

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= tools for likelihood ratio estimation

- “Histogramming”
- Nearest Neighbors
- Support Vector Machines (SVM)
- (Boosted) Decision Trees
- (Deep) Neural Networks
- ...

Not widely used; only useful if decision boundary is ‘simple’

has most things and ROOT-compatible but the community base is **much** smaller than the other ones

Software: TMVA, scikit-learn, XGBoost, tensorflow, pytorch...

does “everything” except DNNs

Data formats: .root, .npy, .hdf5

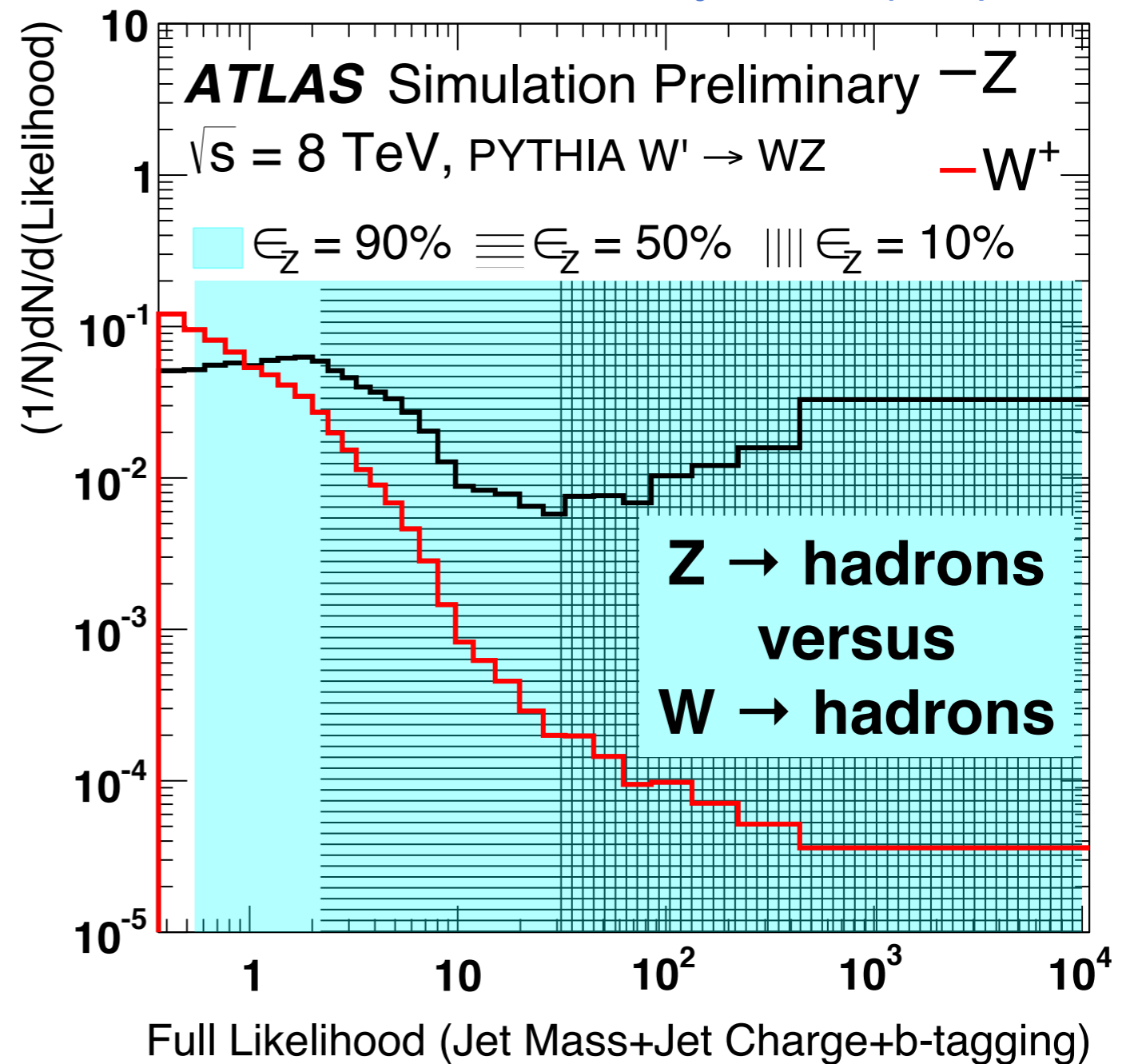


# Histogramming

If you have a 1D problem, look no further!

If your problem can be decomposed into a product/sum of 1D problems...look no further!

If these do not apply... look elsewhere.

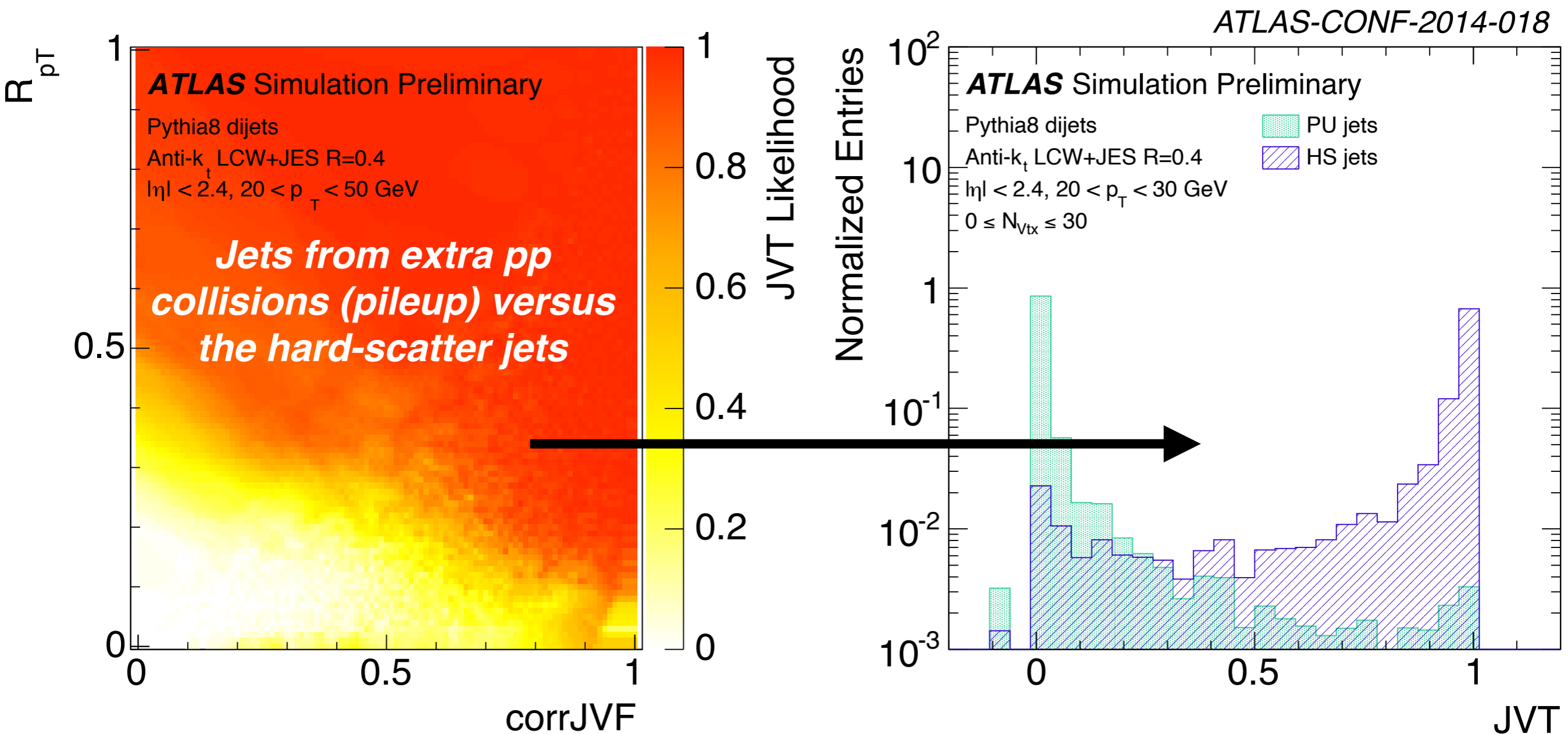


$$p(M, Q, B|V) = \sum_{\mathcal{F}} \Pr(\mathcal{F}|V) p(M|\mathcal{F}, V) p(Q|\mathcal{F}, V) \Pr(B|\mathcal{F}, V),$$

# Nearest Neighbors

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In 2D, a nice extension of histogramming is to estimate the likelihood ratio based on the number of S and B points nearby.



# Boosted Decision Trees (BDTs)

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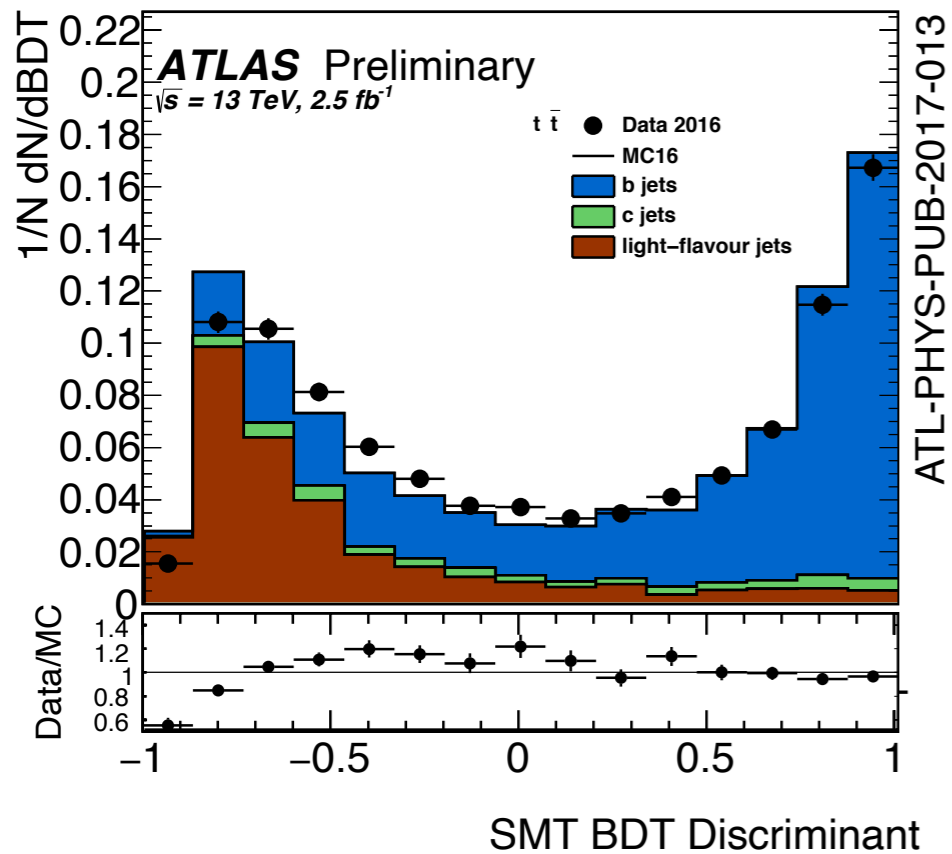
A decision tree is a partition of the feature space.  
One tree is a set of binary “cuts”.

Boosting makes an ensemble classifier. For example, a community favorite AdaBoost, applies weights to the misclassified events.

XGBoost is becoming more the de facto standard. Actually, this method was popularized because of the Higgs Kaggle Challenge!

N.B. BDTs are not differentiable

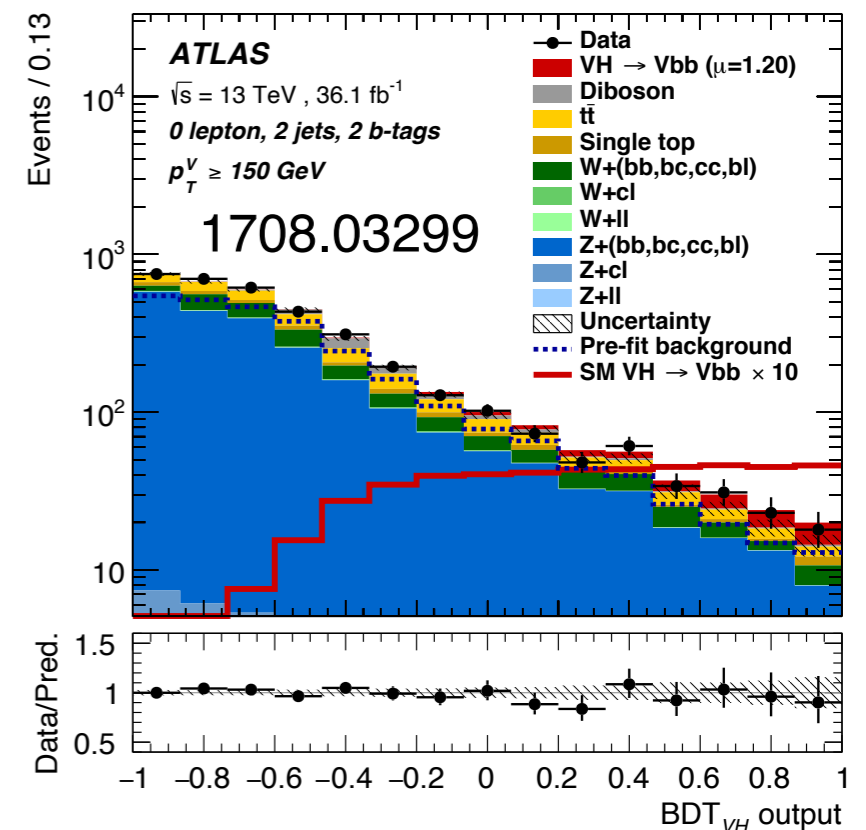
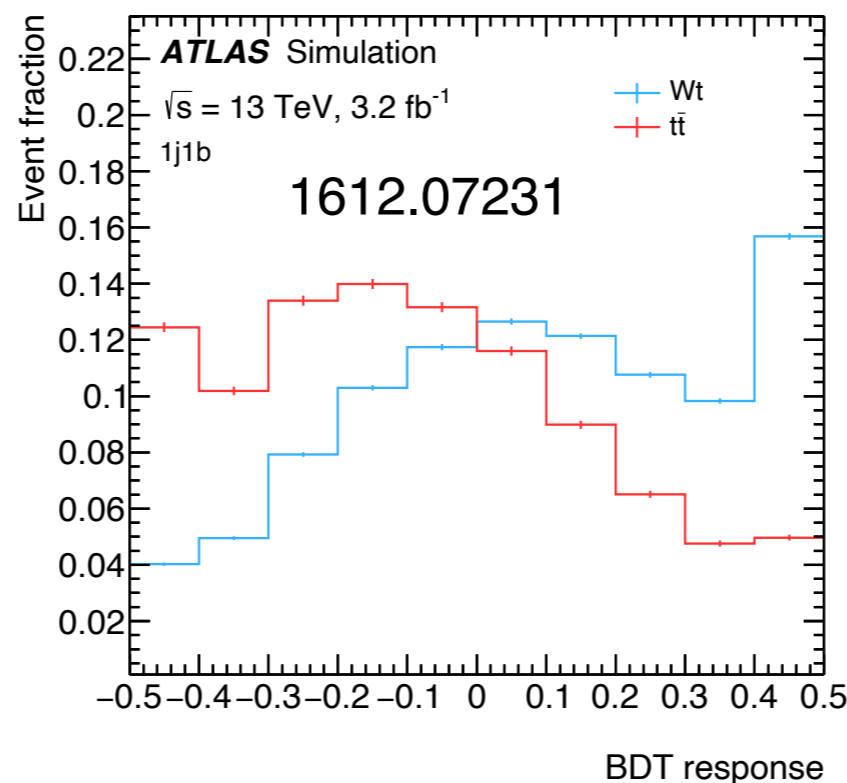
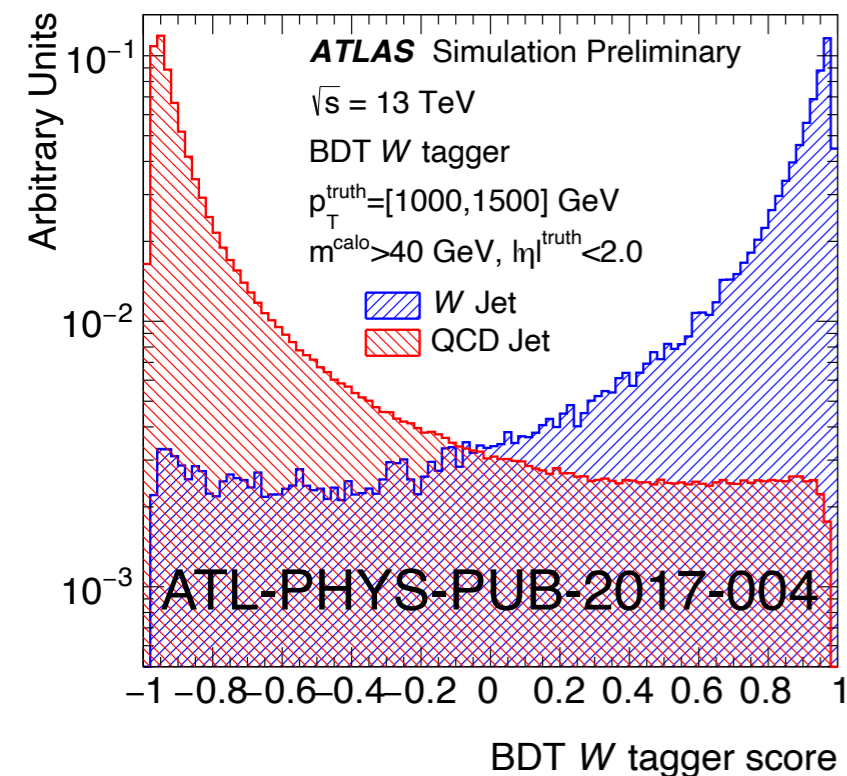
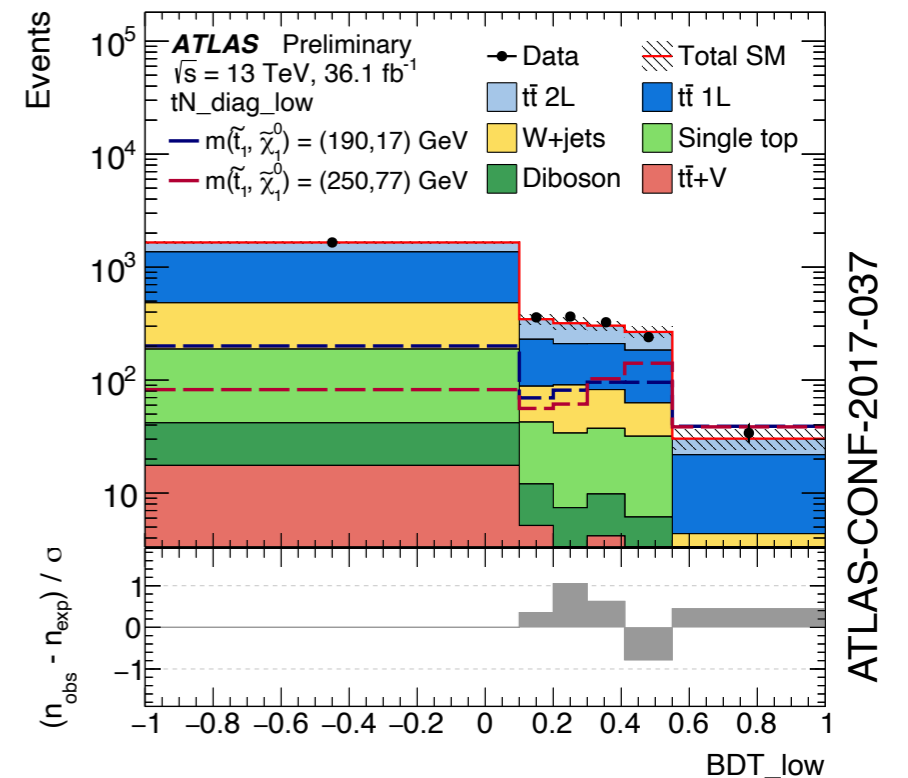
# Boosted Decision Trees (BDTs)



We love BDTs.

If  $3 < \dim(\text{feature vector}) < O(10)$

this is probably right for you!



# Boosted Decision Trees (BDTs)

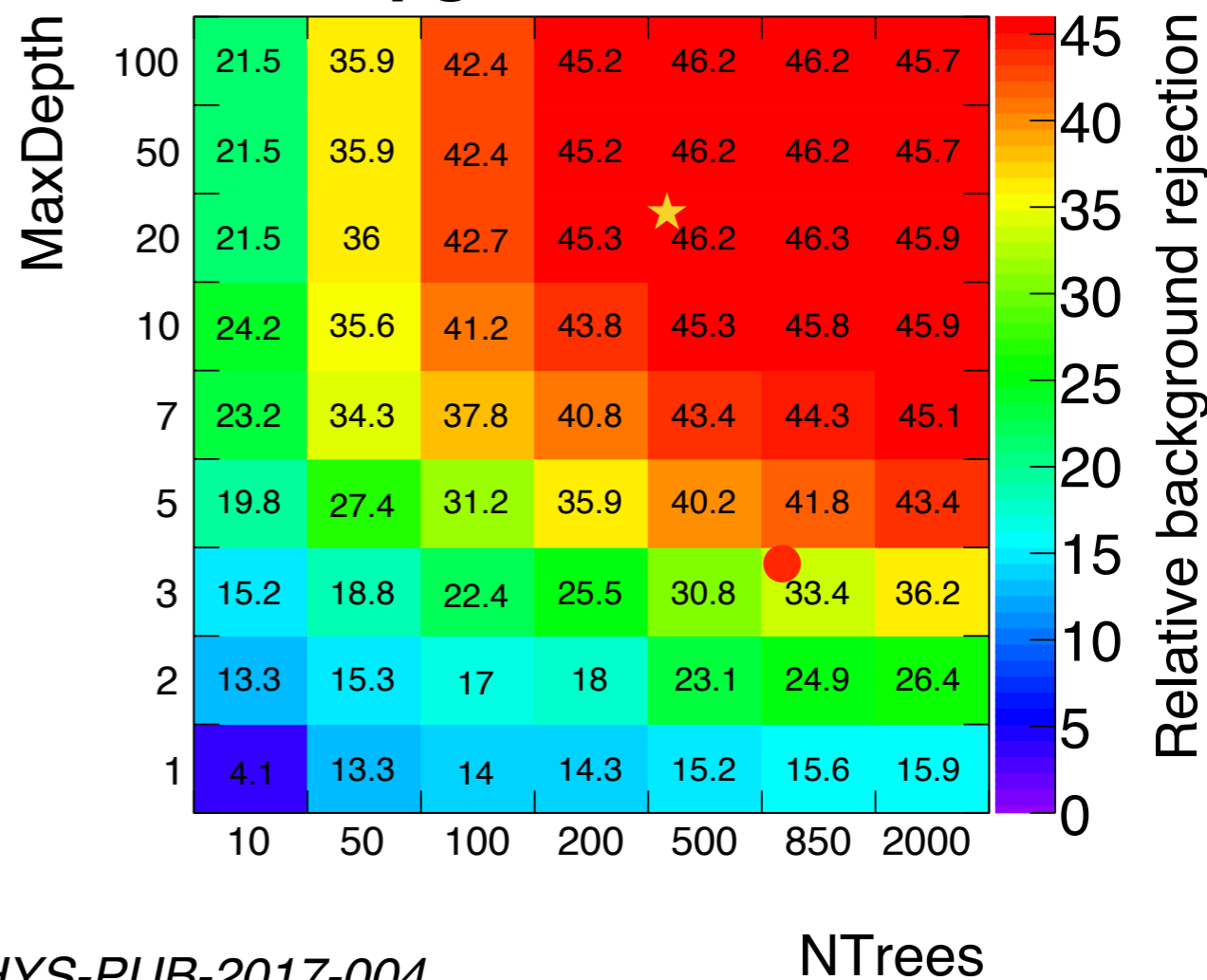
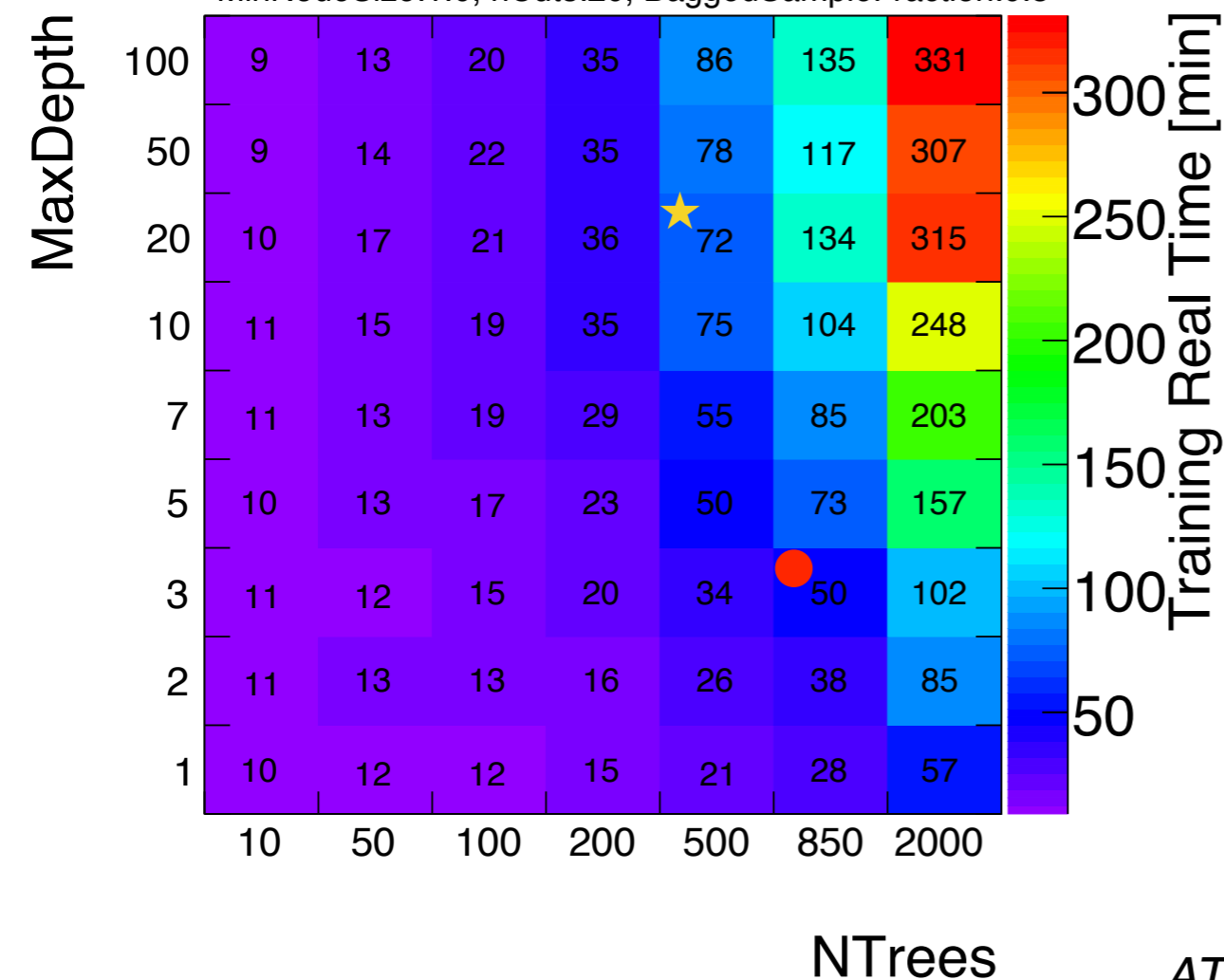


We love BDTs because they are fast to train, are close to “cuts”, and do not have very many parameters. They are also rather robust to *overtraining*.

## ATLAS Simulation Preliminary

$\sqrt{s}=13\text{TeV}$ , BDT  $W$  Tagging,  $\epsilon_{\text{sig}}^{\text{rel}}=50\%$   
 $W$  Jet,  $p_{\text{T}}^{\text{truth}}=[200,2000]$  GeV,  $m^{\text{calo}}>40$  GeV,  $|\eta|^{\text{truth}}<2.0$   
 MinNodeSize:1.0, nCuts:20, BaggedSampleFraction:0.5

## $W \rightarrow \text{hadrons}$ versus $q/g \rightarrow \text{hadrons}$

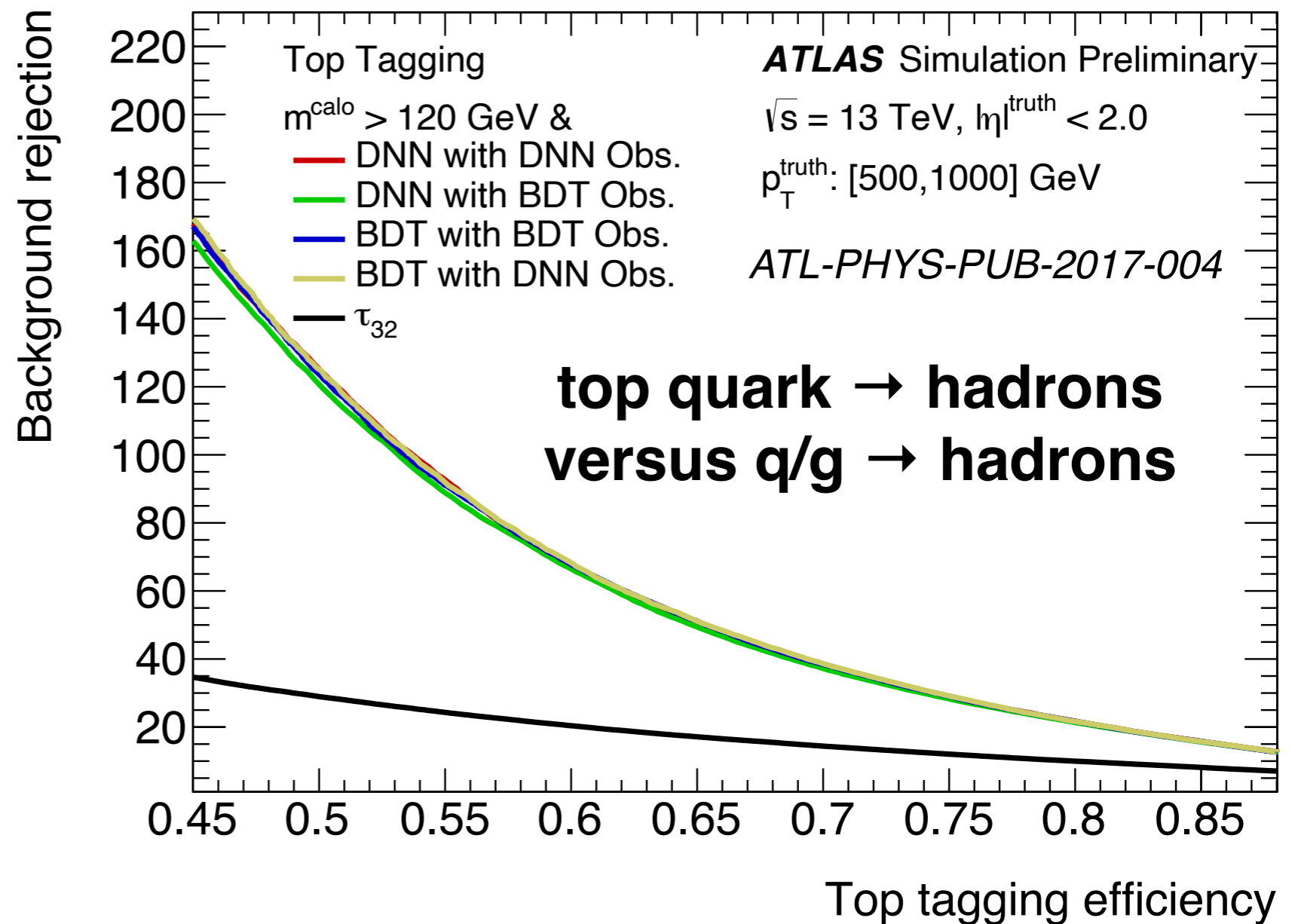


# Boosted Decision Trees (BDTs)

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There is really not a good reason to use a (D)NN with  $\ll O(100)$  dimensions.

However, they are becoming increasingly easy to train ...



Neural Networks were popular at LEP and then mostly fell out of favor until the deep learning revolution.

## **Finding Gluon Jets with a Neural Trigger**

Leif Lönnblad<sup>1</sup>, Carsten Peterson<sup>2</sup> and Thorsteinn Rognvaldsson<sup>3</sup>

Department of Theoretical Physics, University of Lund  
Sölvegatan 14A, S-22362 Lund, Sweden

*Phys. Rev. Lett.* 65 (1990) 1321



Neural Networks were popular at LEP and then mostly fell out of favor until the deep learning revolution.

The NN's of the 90s are rather different than those of today! With advanced in hardware (GPUs), architectures (dropout, ReLU), etc. the DNNs of today are qualitatively different and more powerful.



# What is a Neural Network learning?

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Consider the popular binary cross entropy:

$$\text{loss}(f(x)) = - \sum_{i \in S} \log f(x_i) - \sum_{i \in B} \log(1 - f(x_i))$$

# What is a Neural Network learning?

28

Consider the popular binary cross entropy:

$$\text{loss}(f(x)) = - \sum_{i \in S} \log f(x_i) - \sum_{i \in B} \log(1 - f(x_i))$$

If  $f$  is optimal, what will it learn?

One can show that **asymptotically**,

$$f(x) \approx \Pr(S | X = x)$$

(this is monotonic with the likelihood ratio)

# What is a Neural Network le

Consider the popular bin

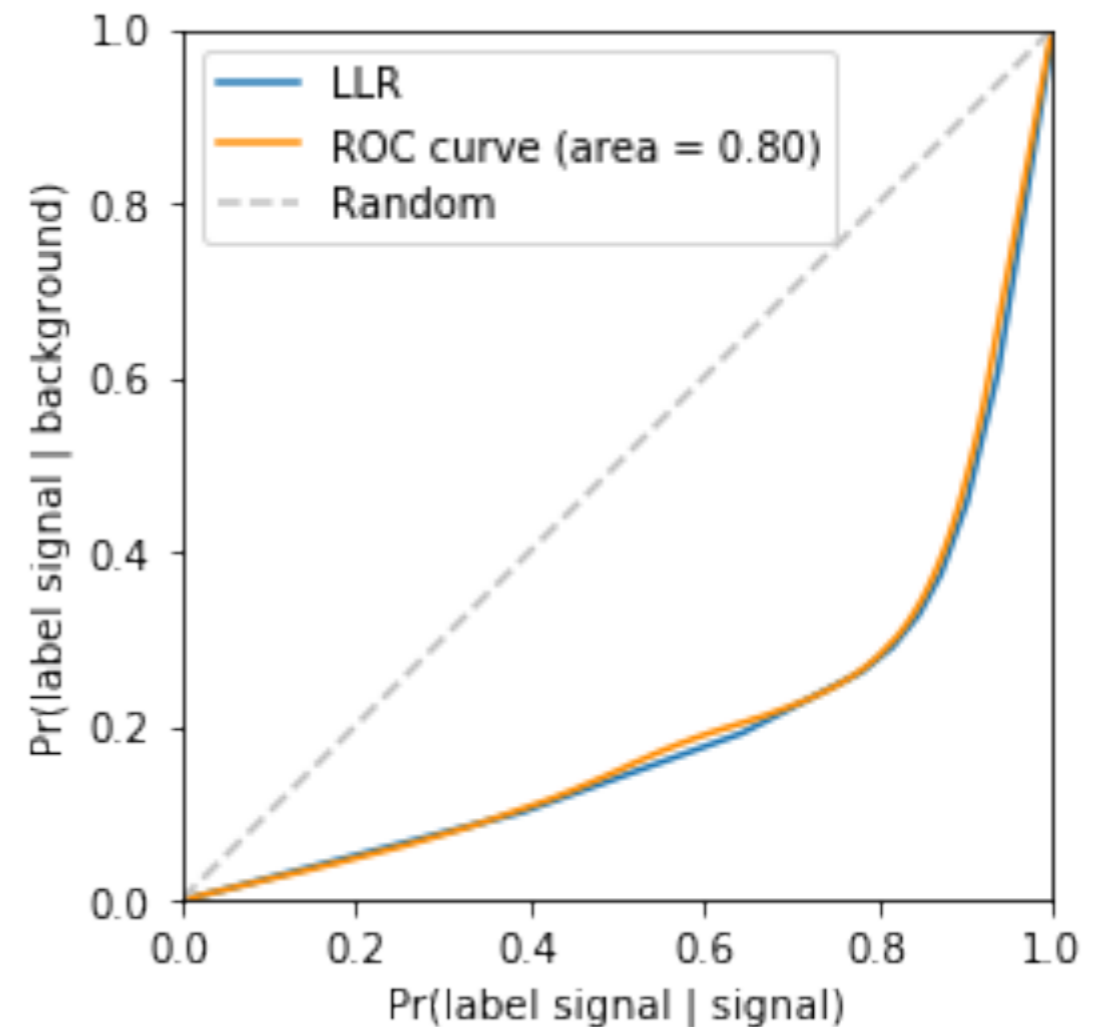
$$\text{loss}(f(x)) = - \sum_{i \in S} \log f(x_i)$$

If  $f$  is optimal, what

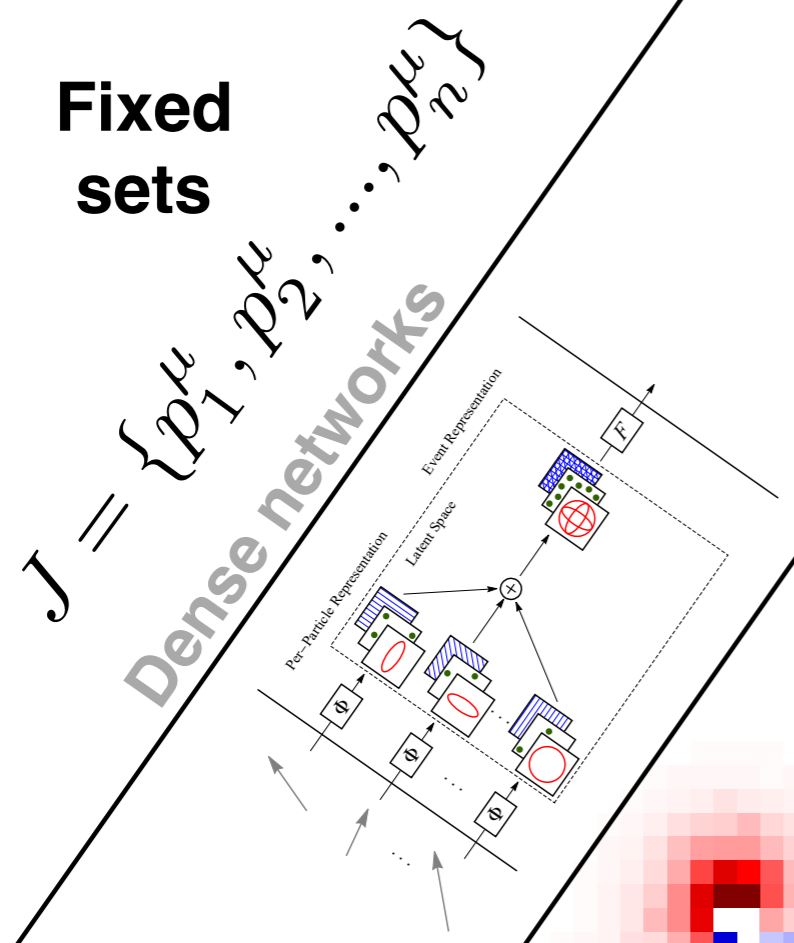
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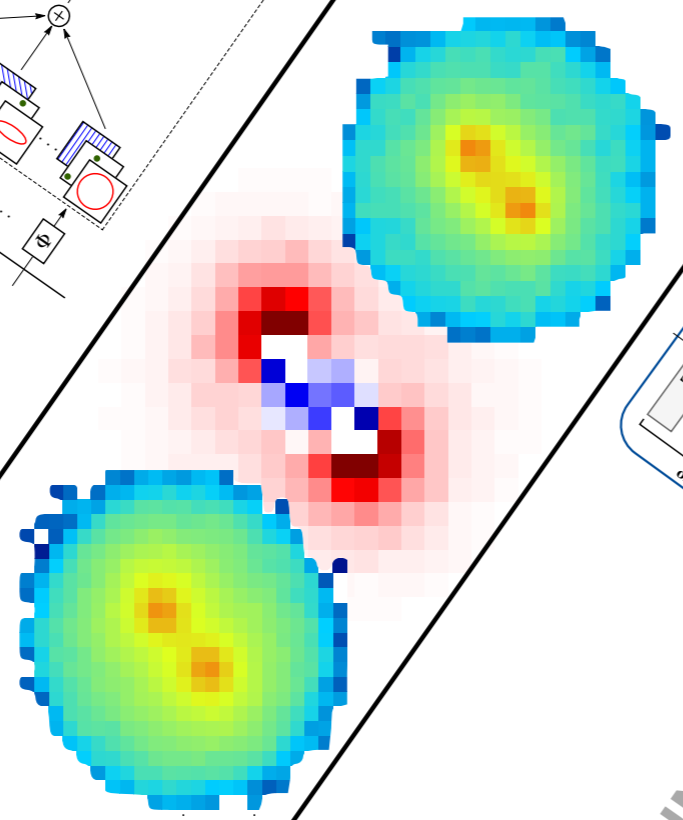
# Deep neural networks for HEP classification



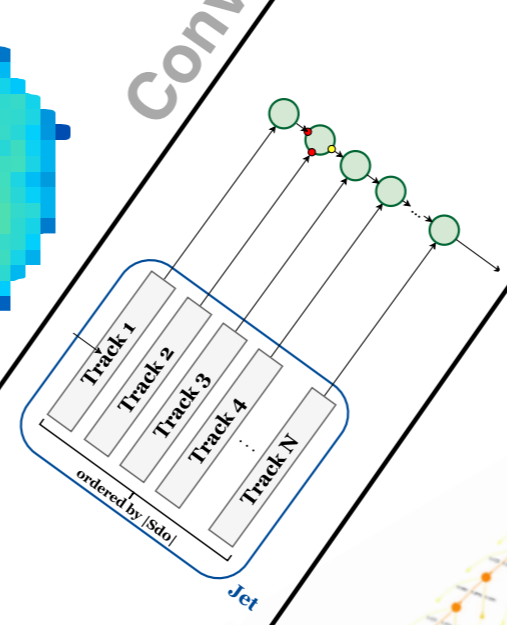
**Variable sets**

**Deep sets**

**Images**



**Sequences**



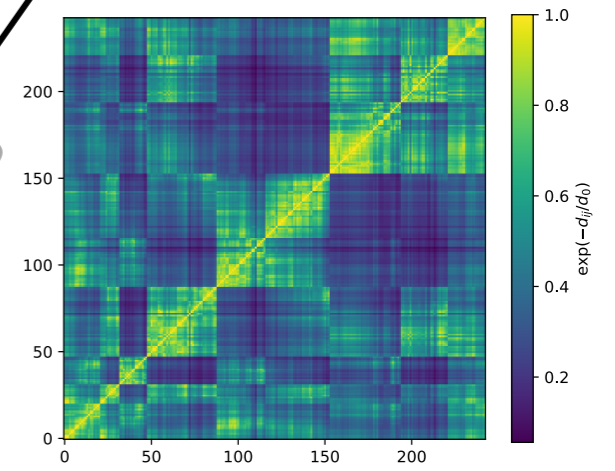
**Recurrent NNs**

**Trees**

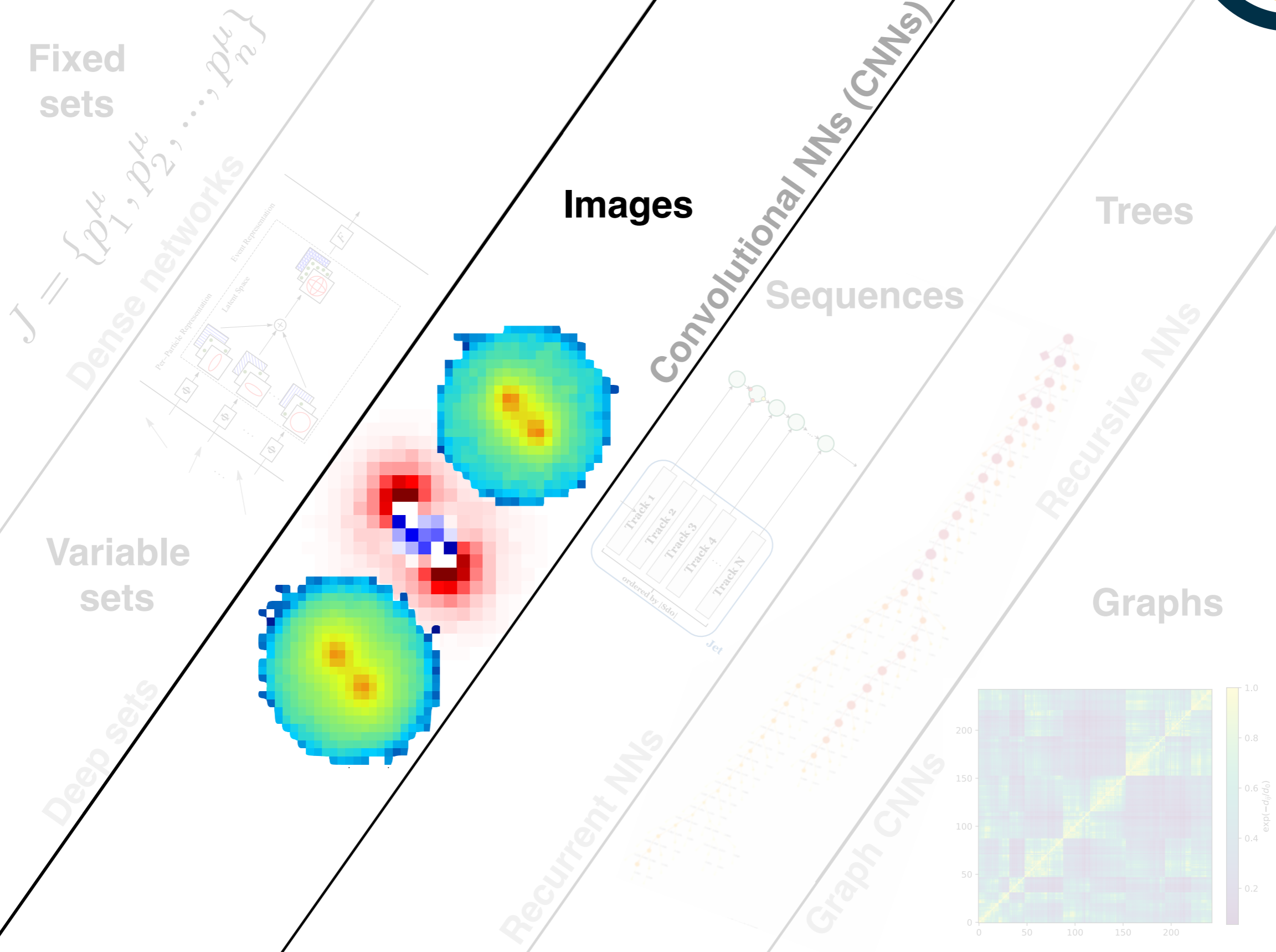


**Graphs**

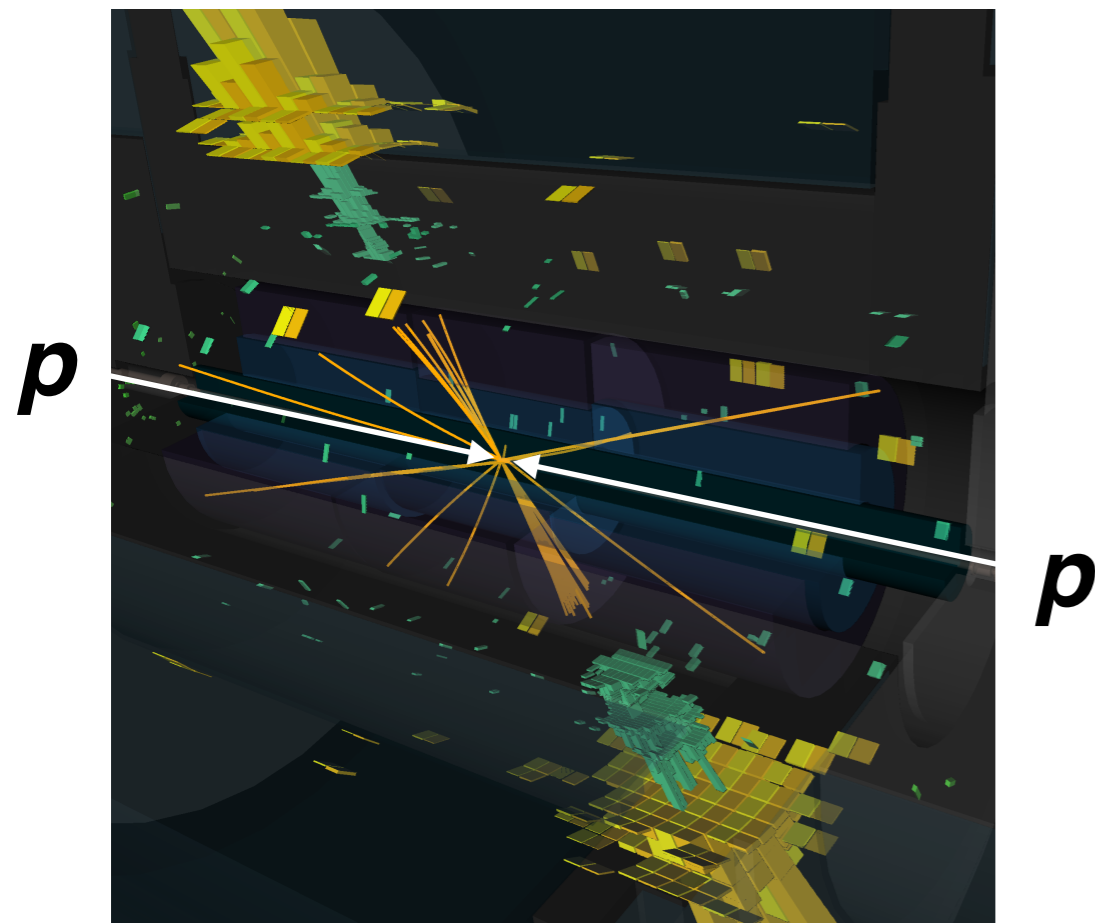
**Graph CNNs**



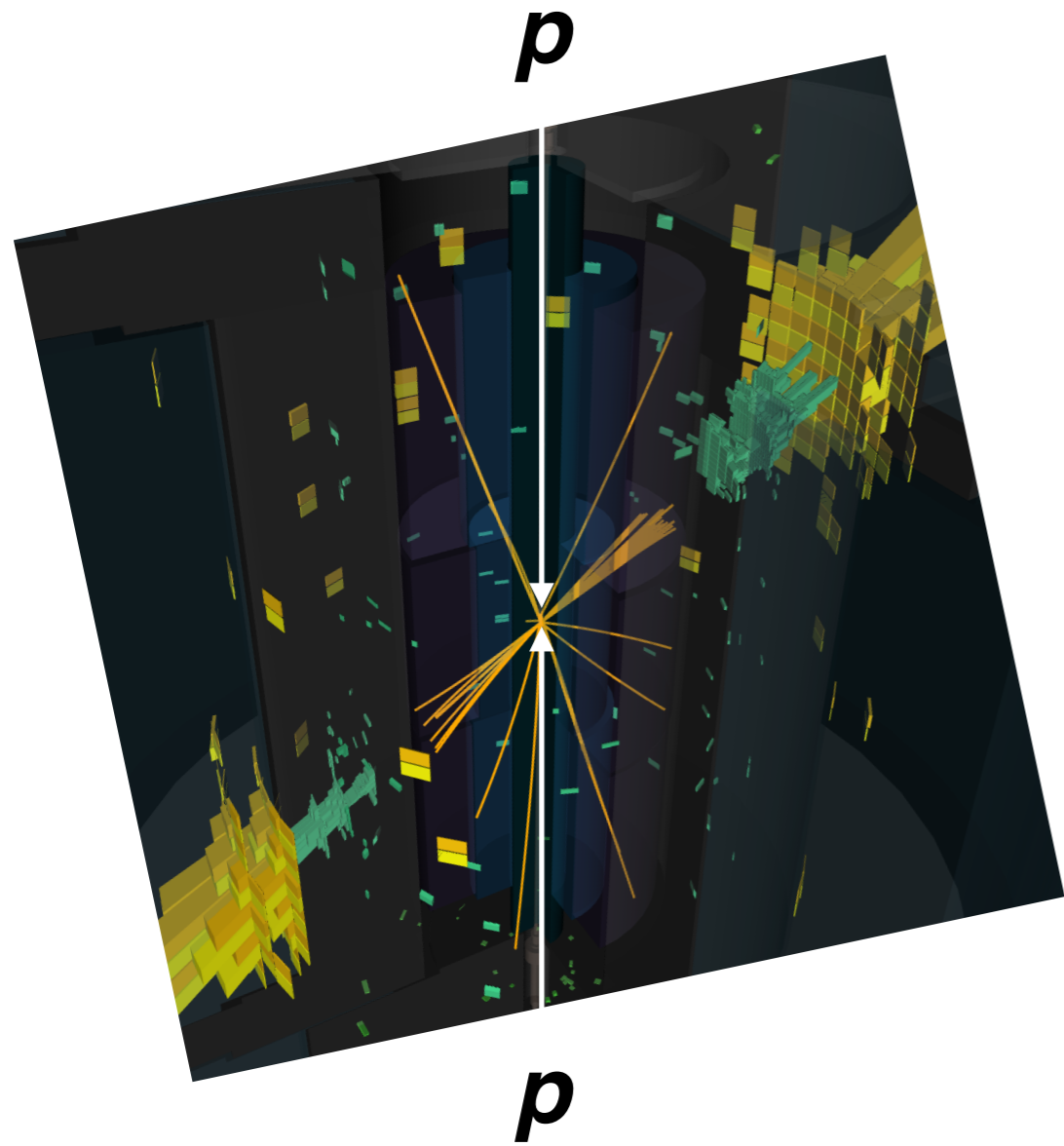
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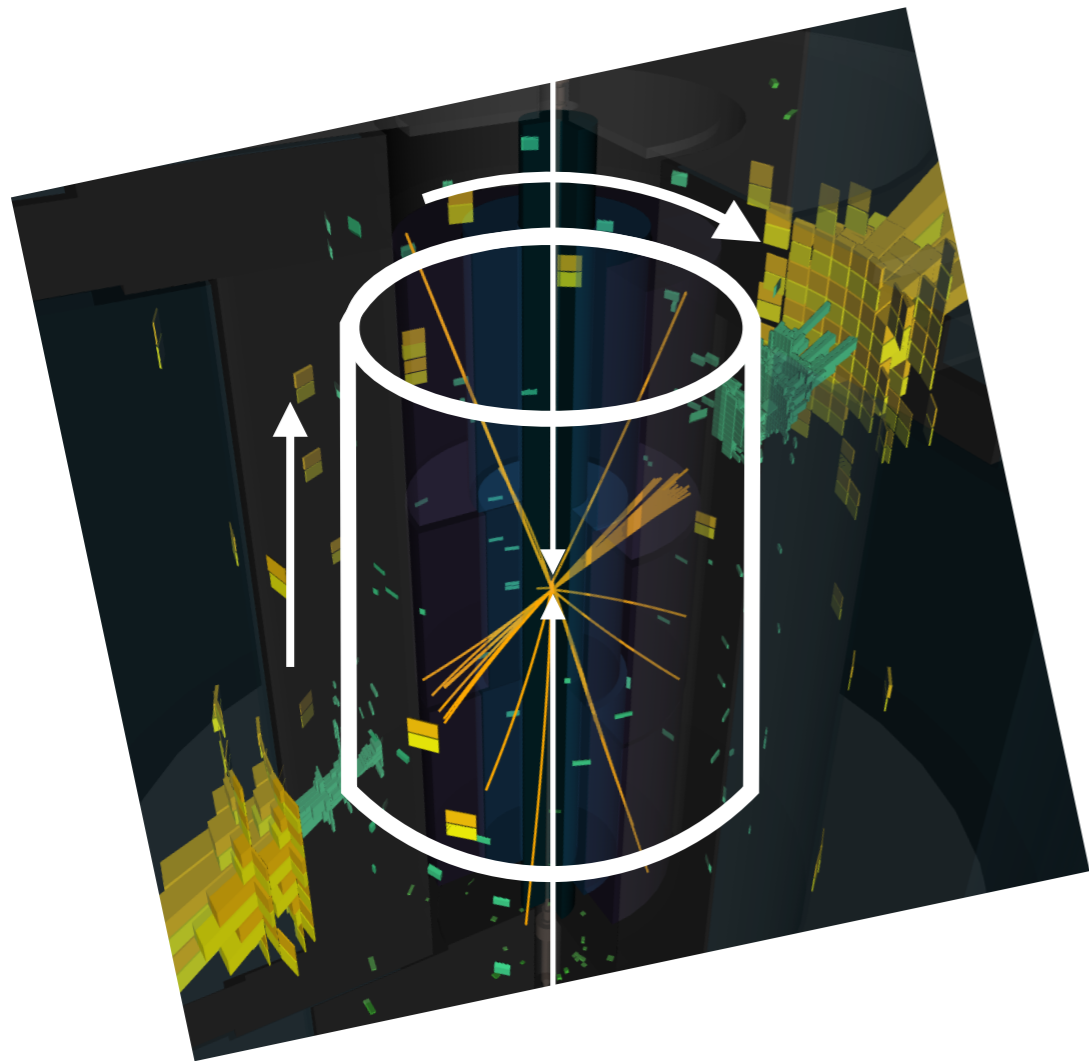
# HEP data as an image



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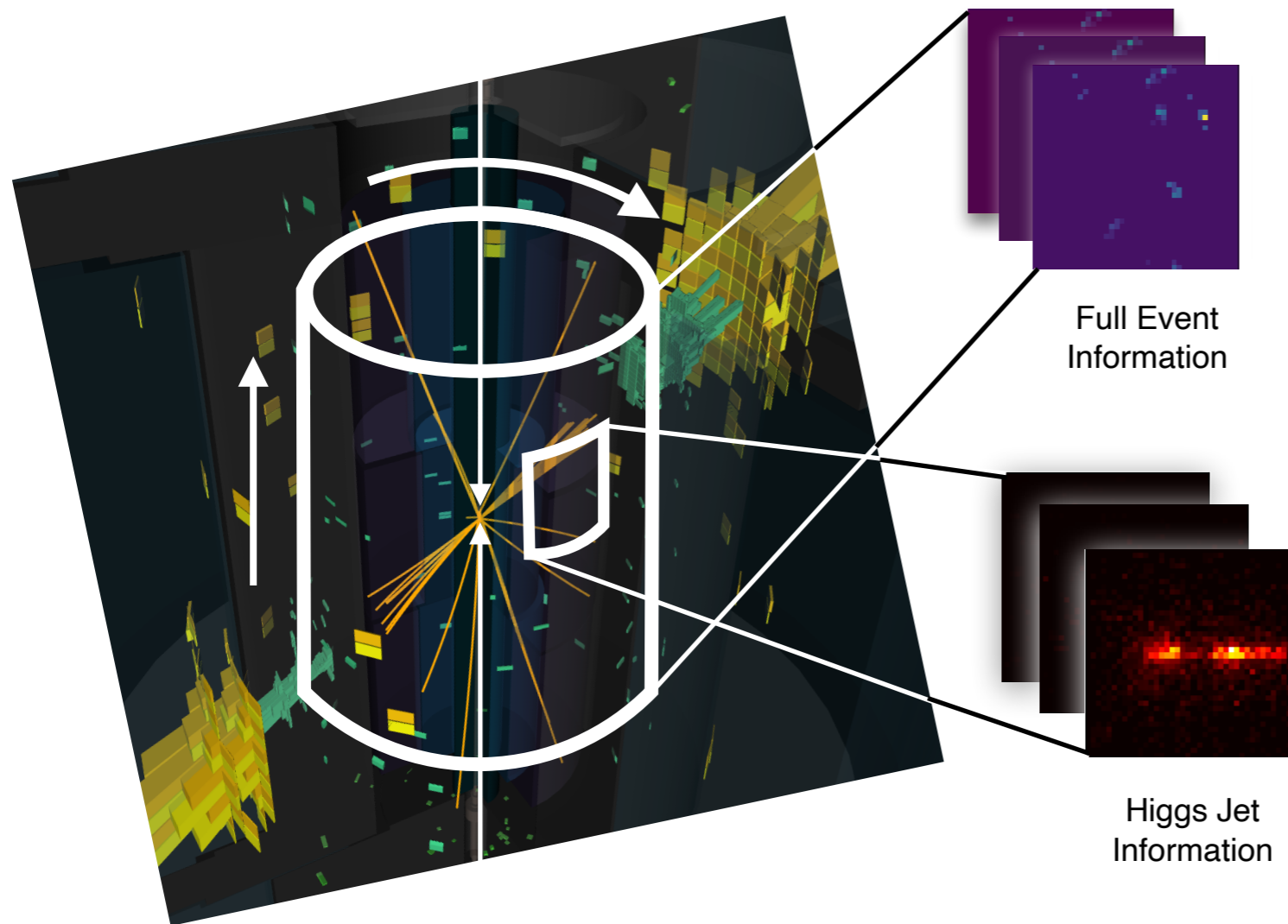


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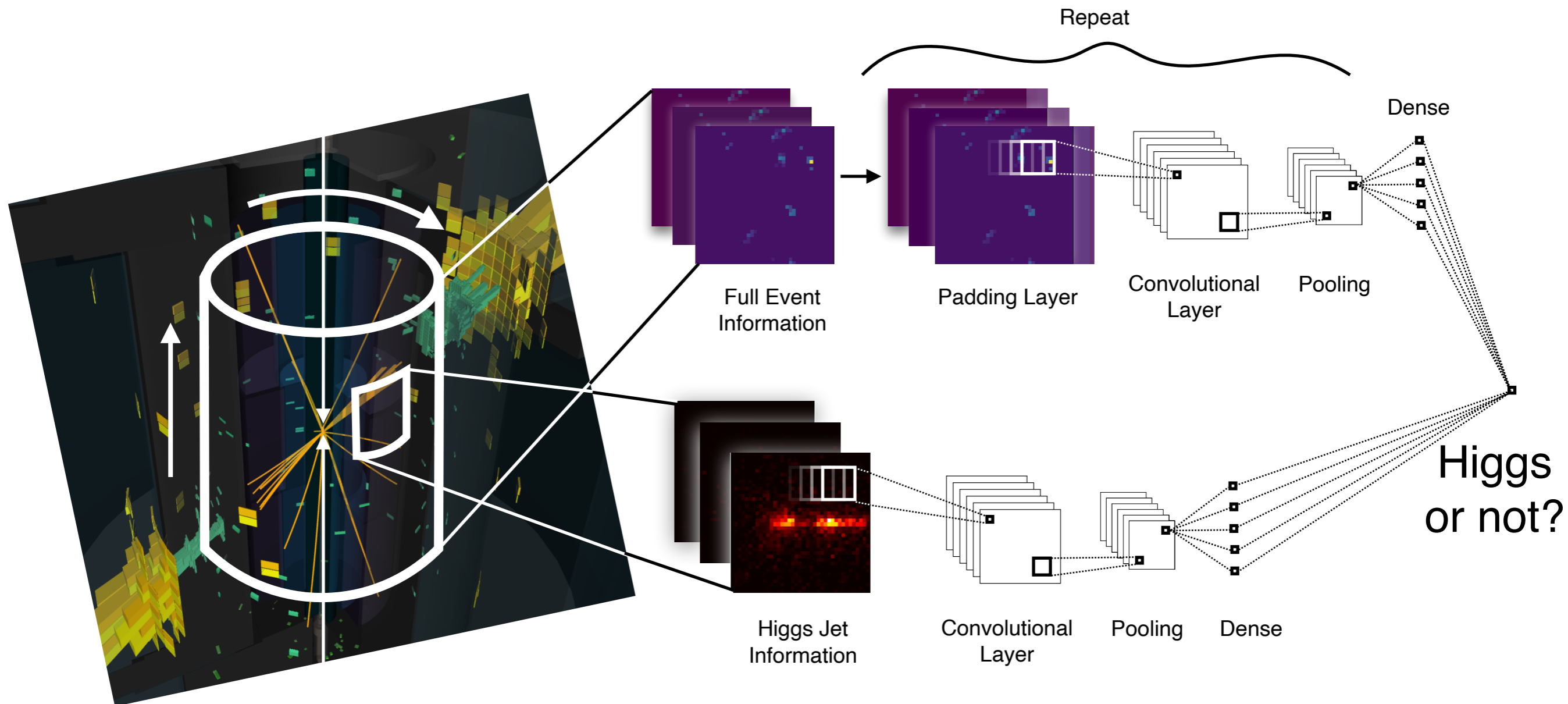


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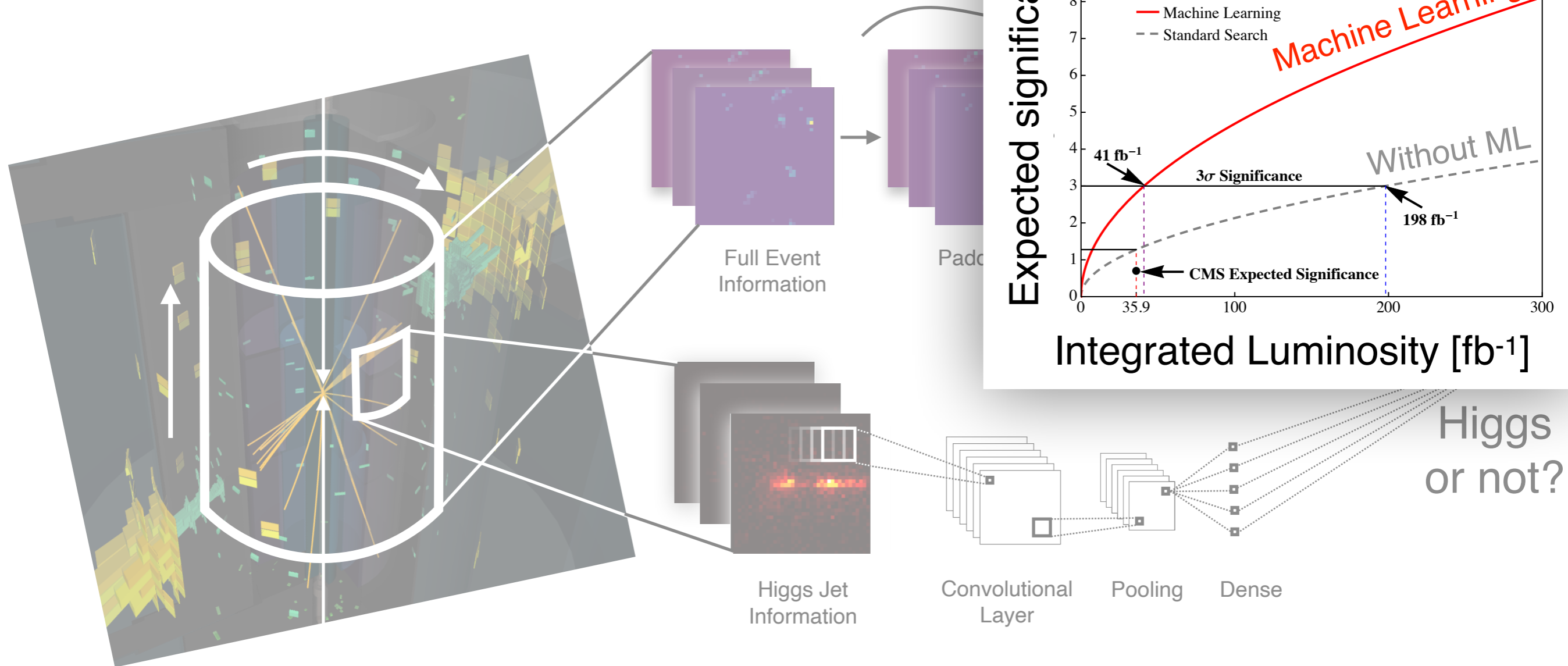
# HEP data as an image

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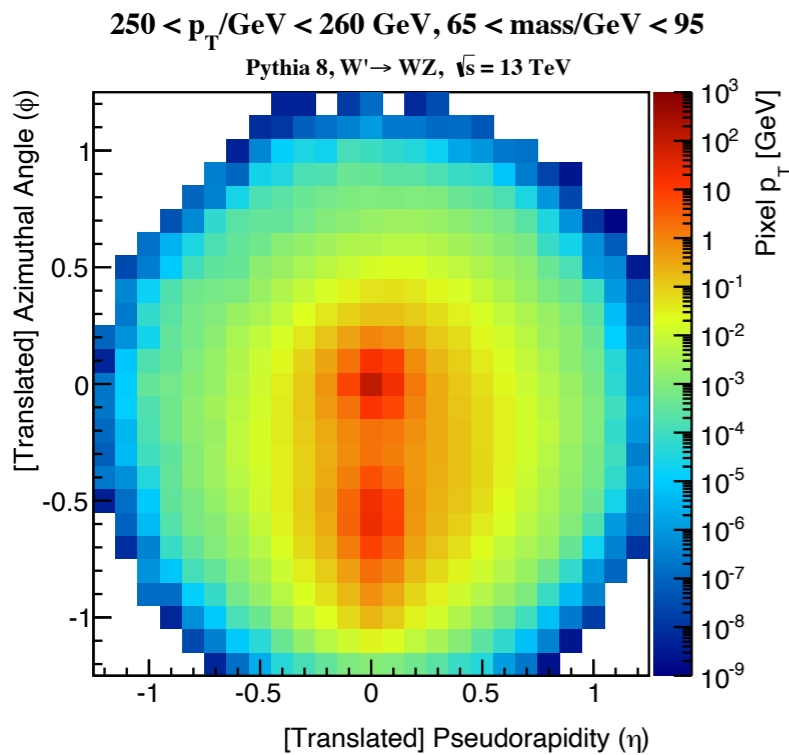
Can combine local and global information from jet images and “event” images.

# HEP data as an image



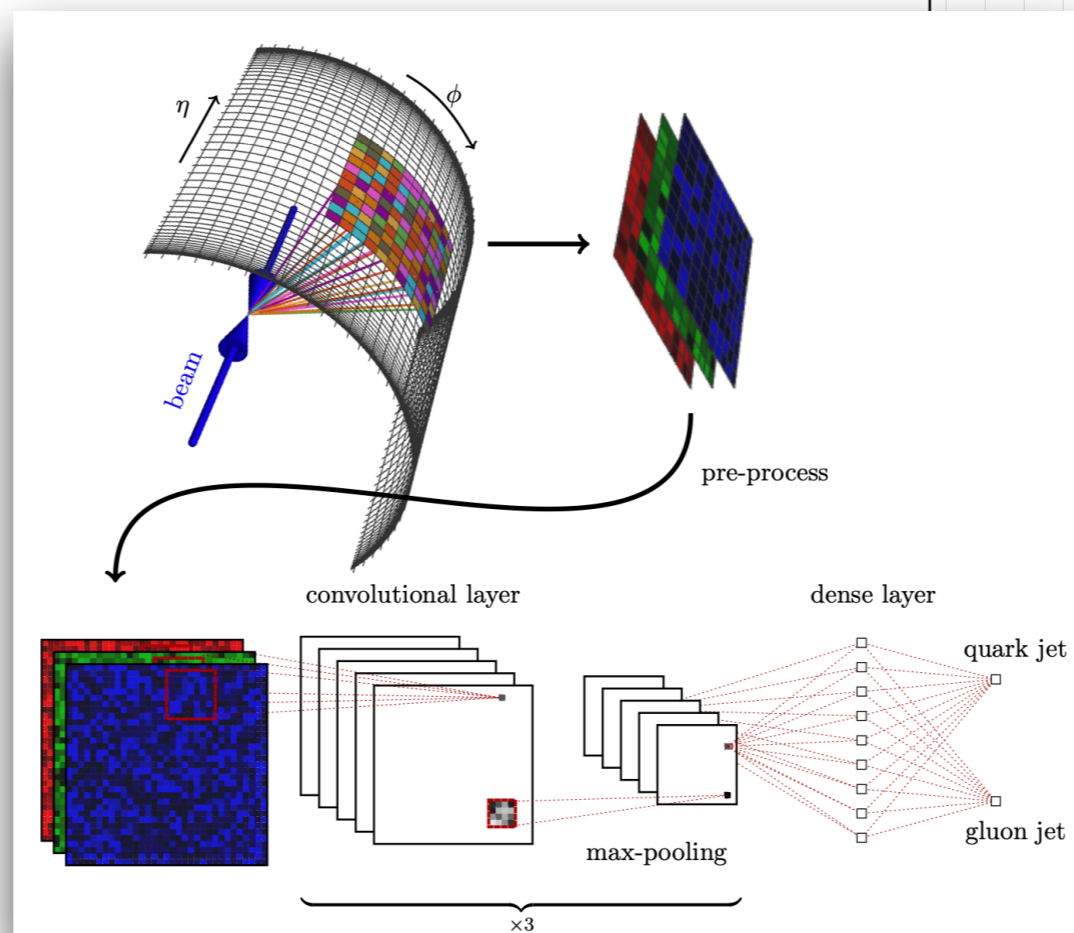
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# More HEP images

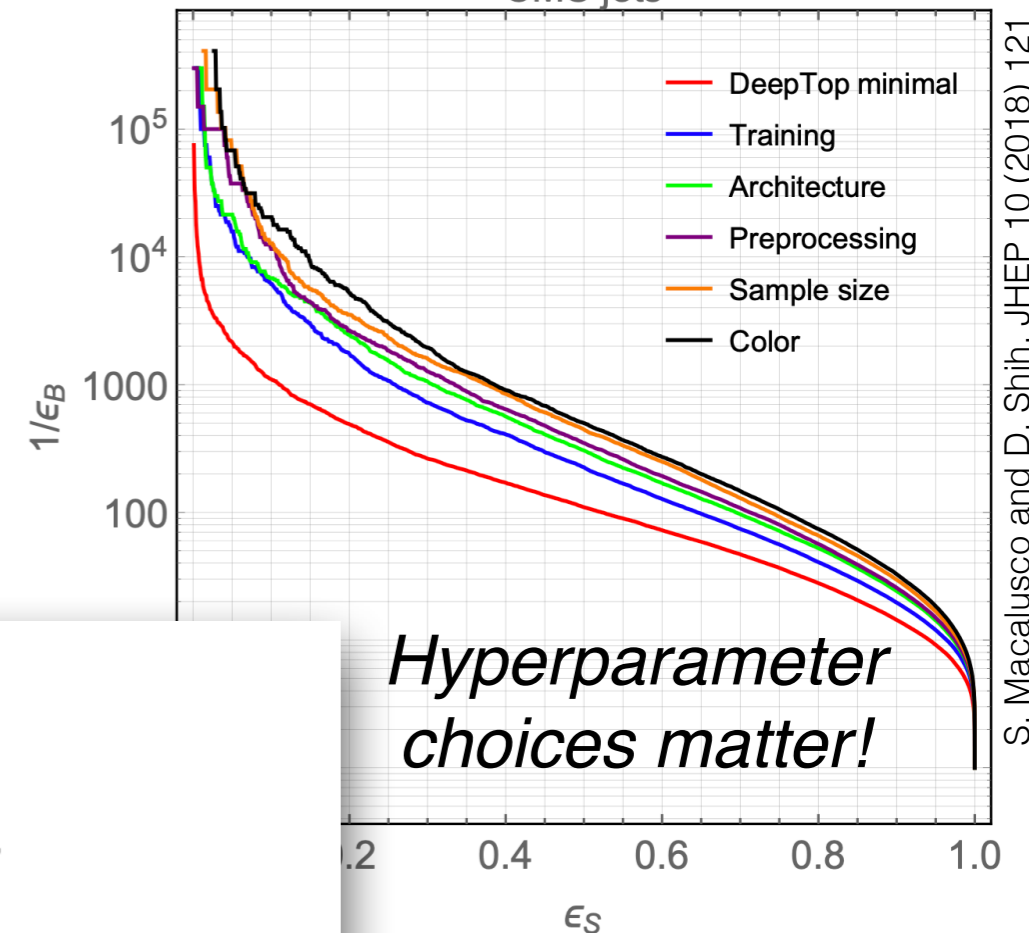


L. De Oliveira et al., JHEP 07 (2016) 069

Many topologies:  
 top quarks, W/Z/H  
 bosons, BSM  
 particles, q/g, etc.



P. Komiske, E. Metodiev, M Schwartz, JHEP 01 (2017) 110

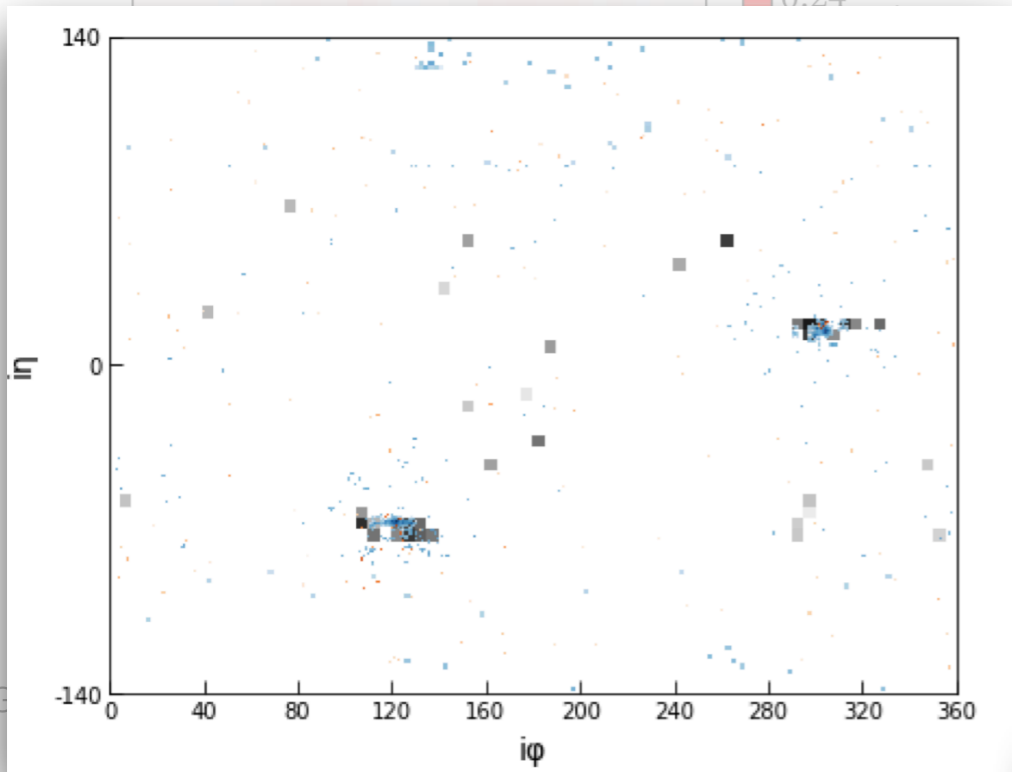


*Hyperparameter choices matter!*

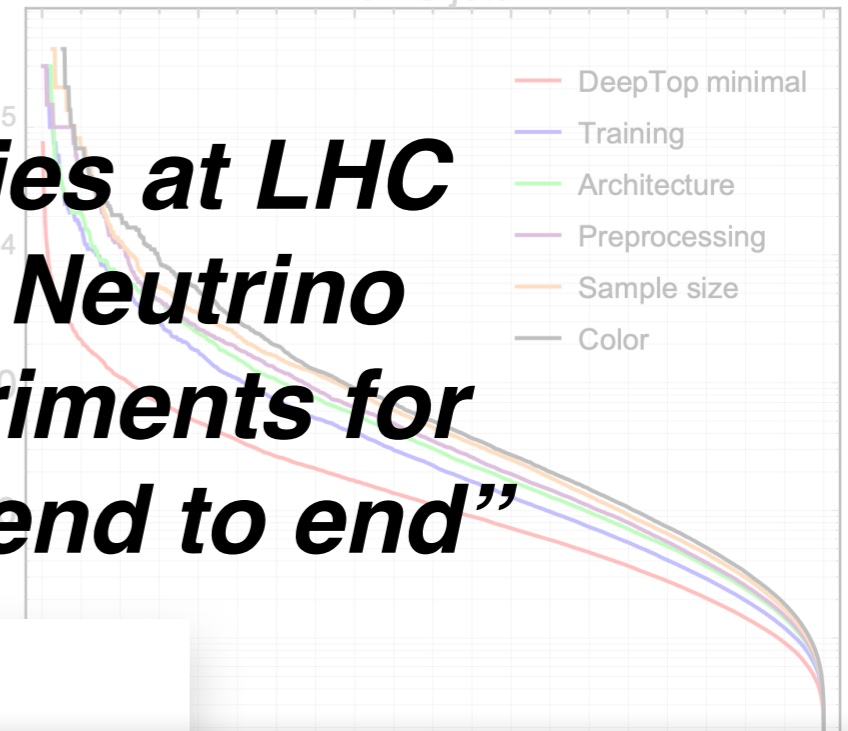
Multi-channel:  
 use calorimeter  
 & tracking  
 information to  
 make **RGB**  
 image.

# More HEP images

M. Andrews et al., <https://arxiv.org/abs/1902.08276>

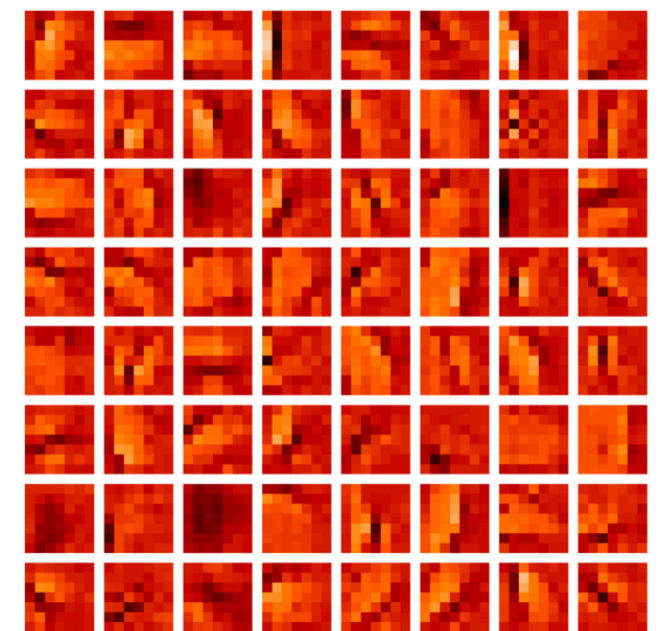
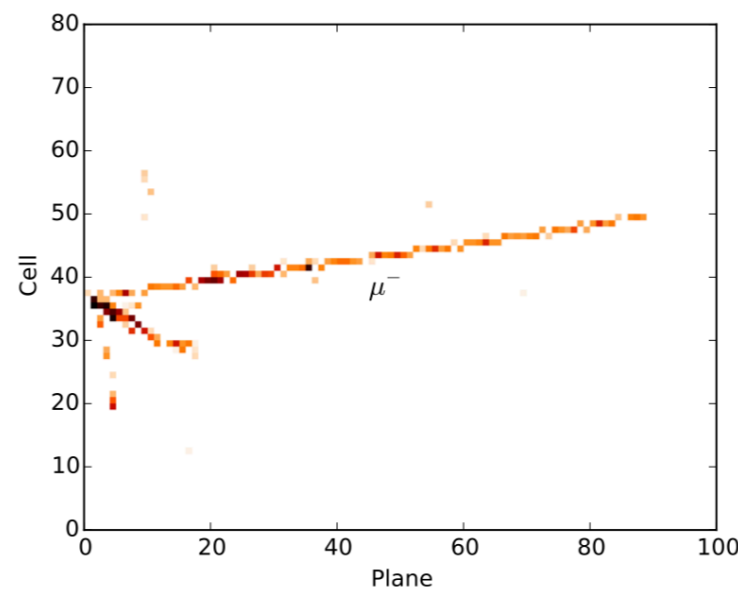


**Studies at LHC  
and Neutrino  
experiments for  
fully “end to end”**



S. Macaluso and D. Shih, JHEP 10 (2018) 121

A. Aurisano, et al., JINST 11 (2016) P09001

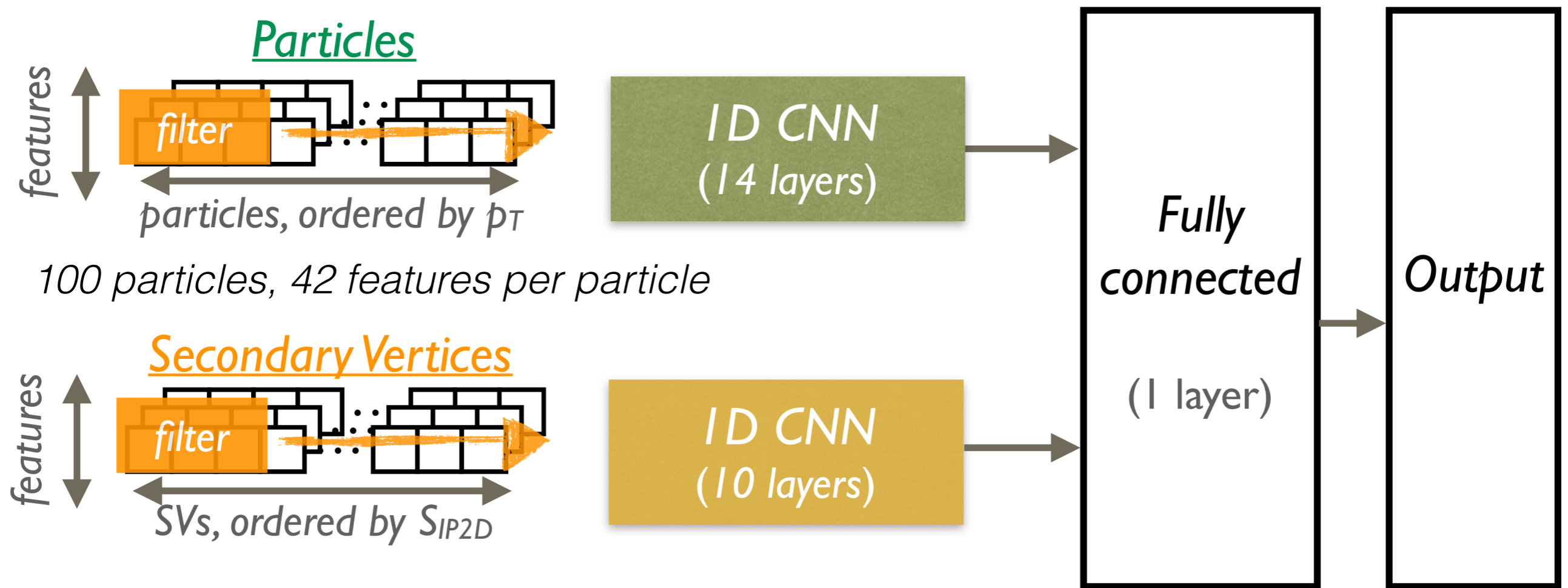


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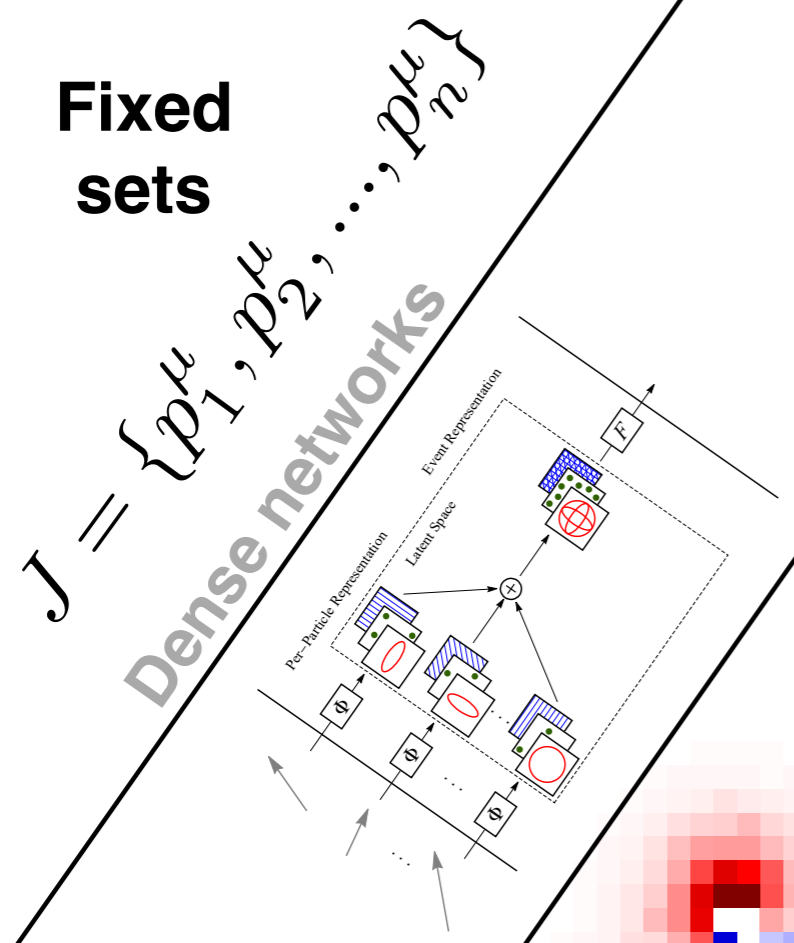
# A last word about CNNs for now

One can use CNNs as automated “feature extractors” even if the inputs are not images.



This is the structure of the CMS Collaboration Deep AK8 jet classifier.

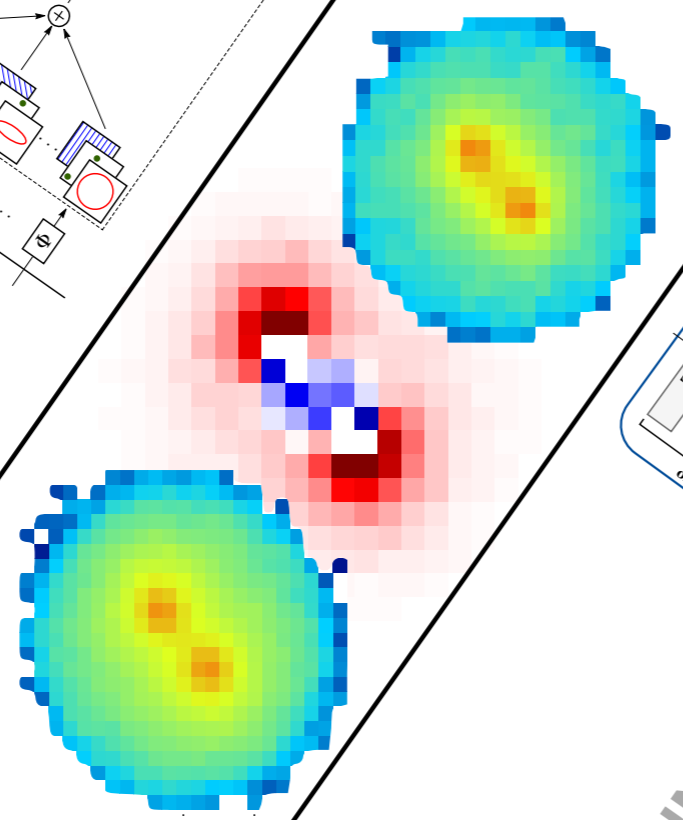
# Deep neural networks for HEP classification



**Variable sets**

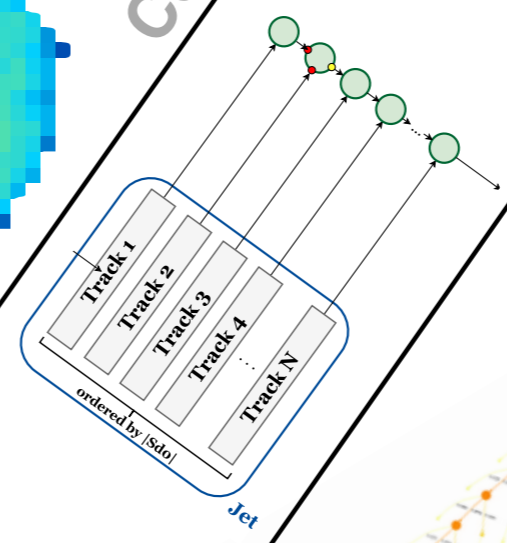
**Deep sets**

**Images**



**Convolutional NNs (CNNs)**

**Sequences**



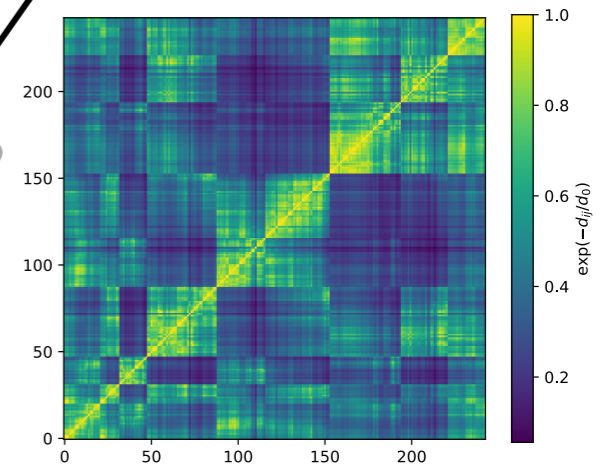
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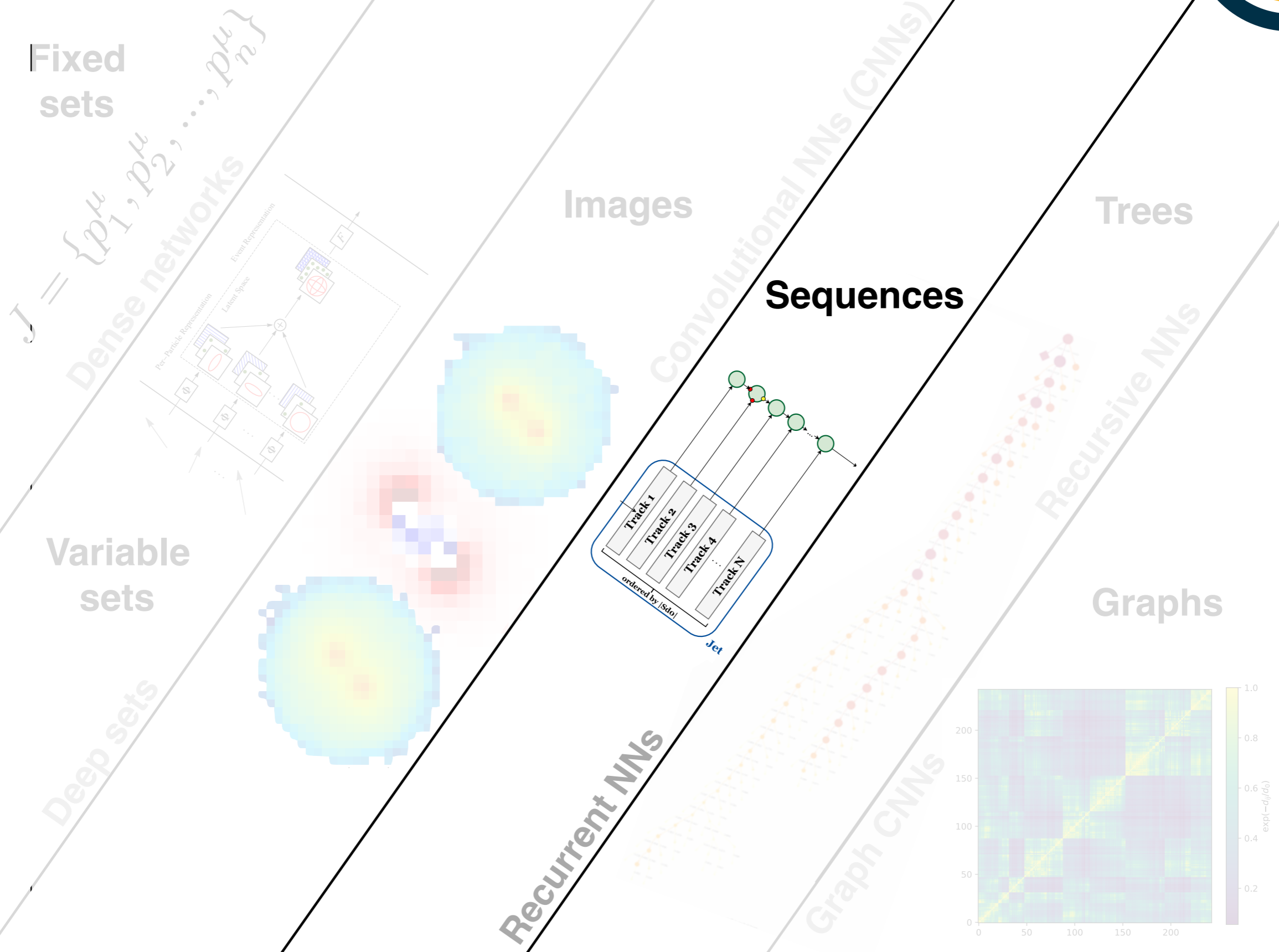


**Graphs**

**Graph CNNs**



# Deep neural networks for HEP classification





One key challenge with images is that they have a fixed size.

*In many contexts, this is ideal, because the data also have a fixed size. However, this is not always the case.*

For example, events / jets have a variable number of particles.

One can represent these particles as a sequence in order to apply variable-length approaches that can access the full feature granularity.

# Sequence learning with RNNs

Flavor tagging (classify jets from b-quark or not) has a long history of ML. Use features of the charged-particle tracks inside jets.

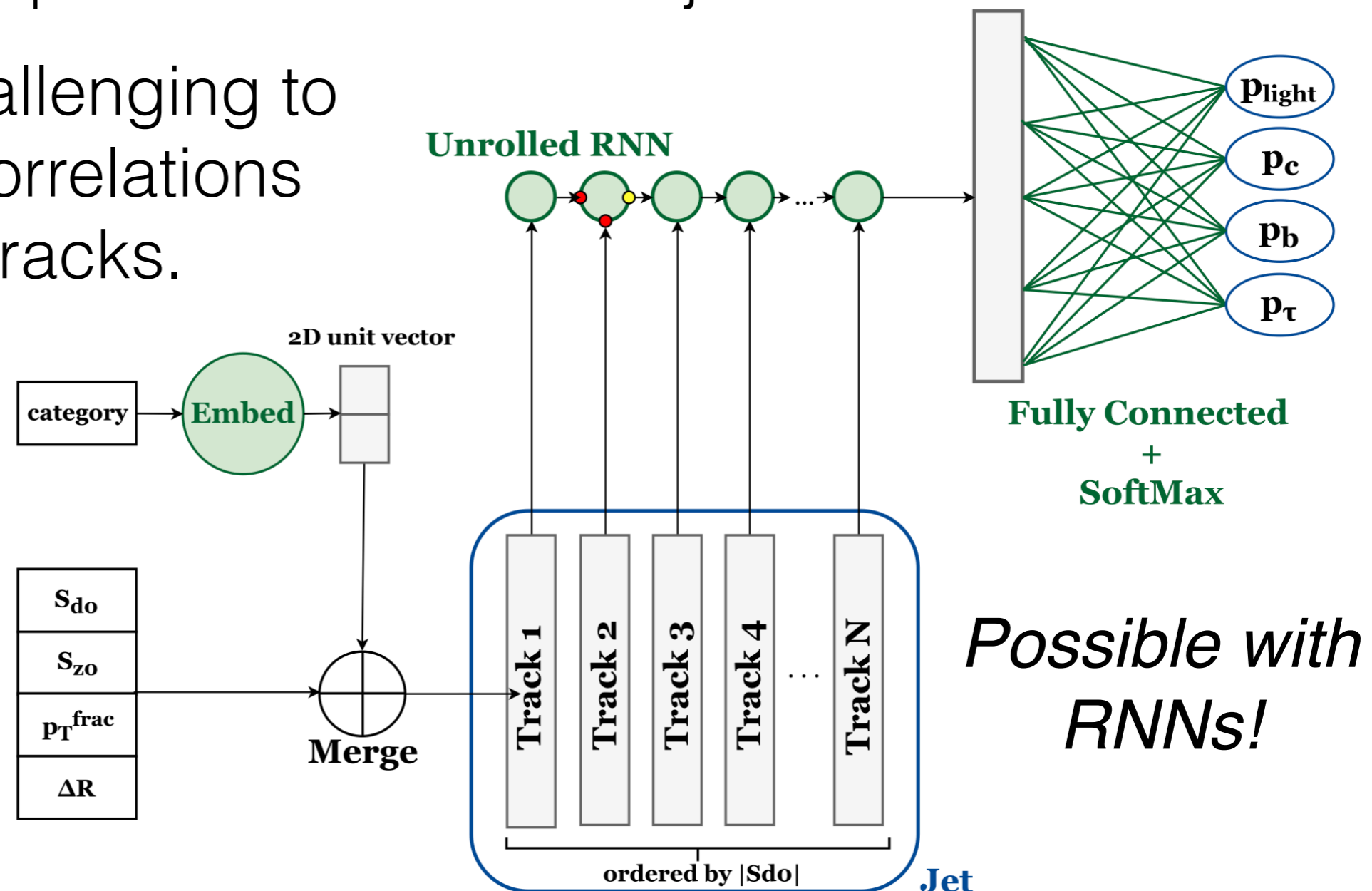
In the past, challenging to incorporate correlations between tracks.

# Sequence learning with RNNs

45

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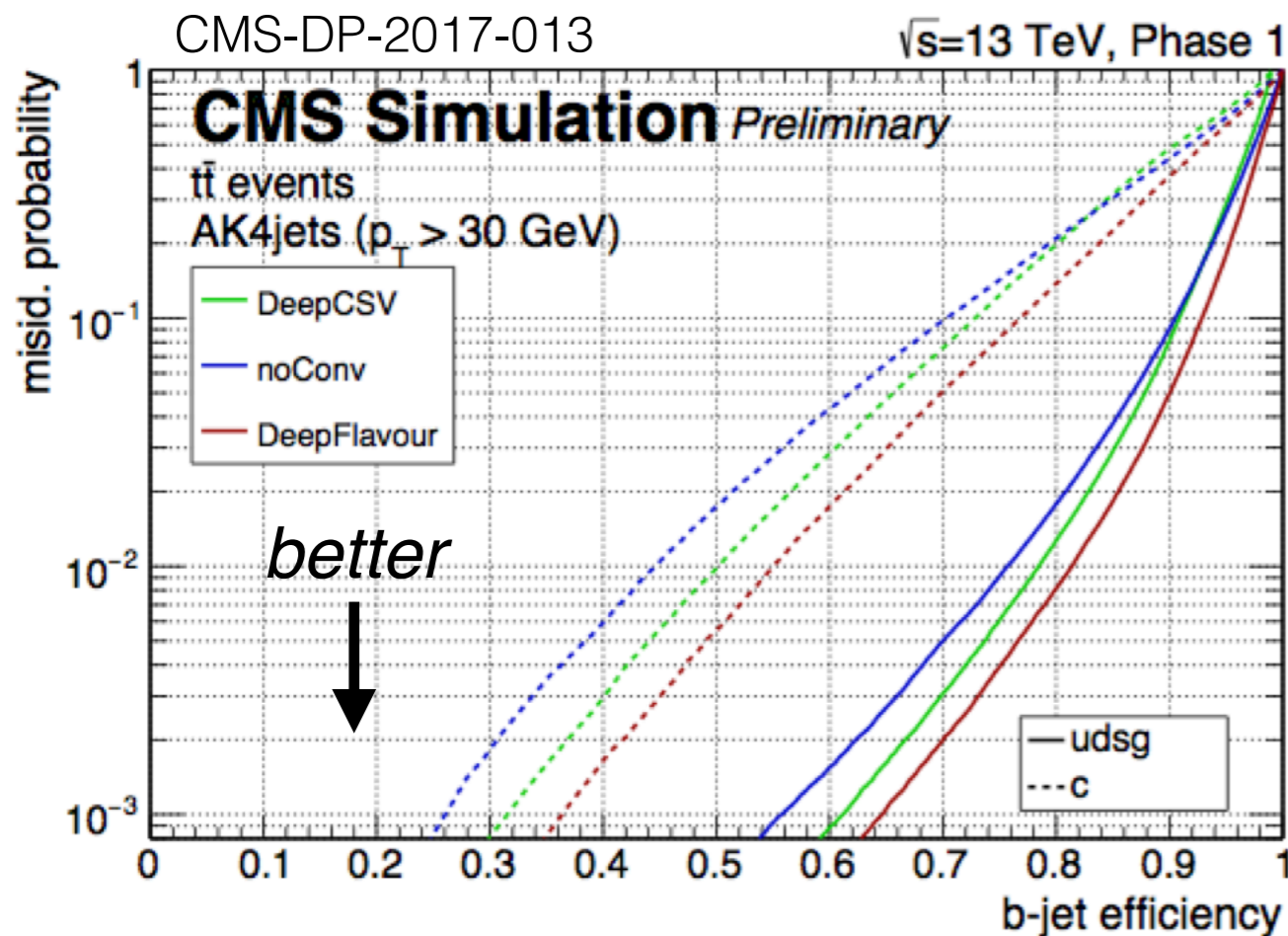
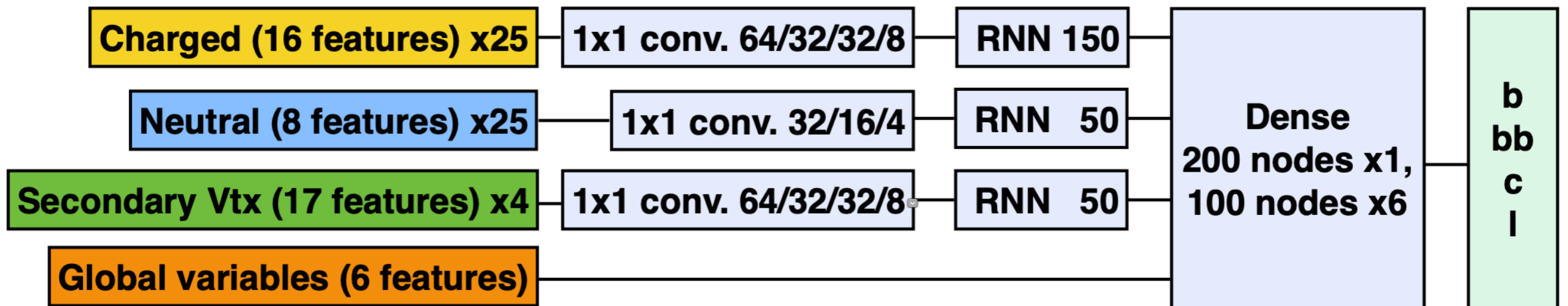
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*Possible with  
RNNs!*

# Hybrid methods

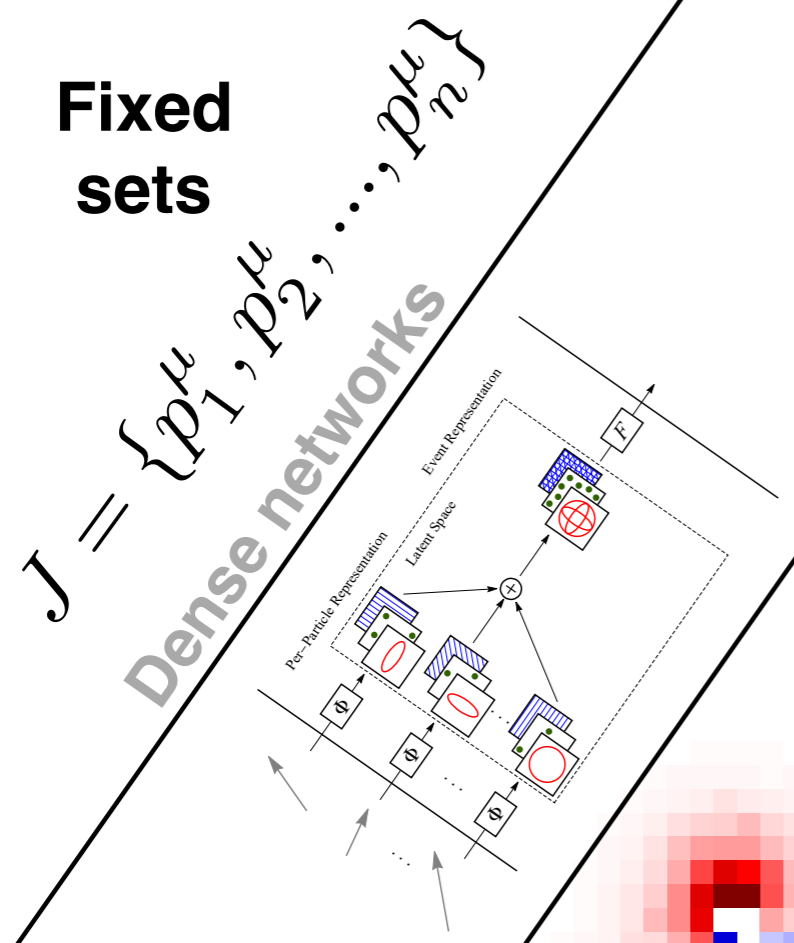
46



RNN + 1x1 CNNs for dimensionality reduction.

This reduction improved the performance of the overall classifier.

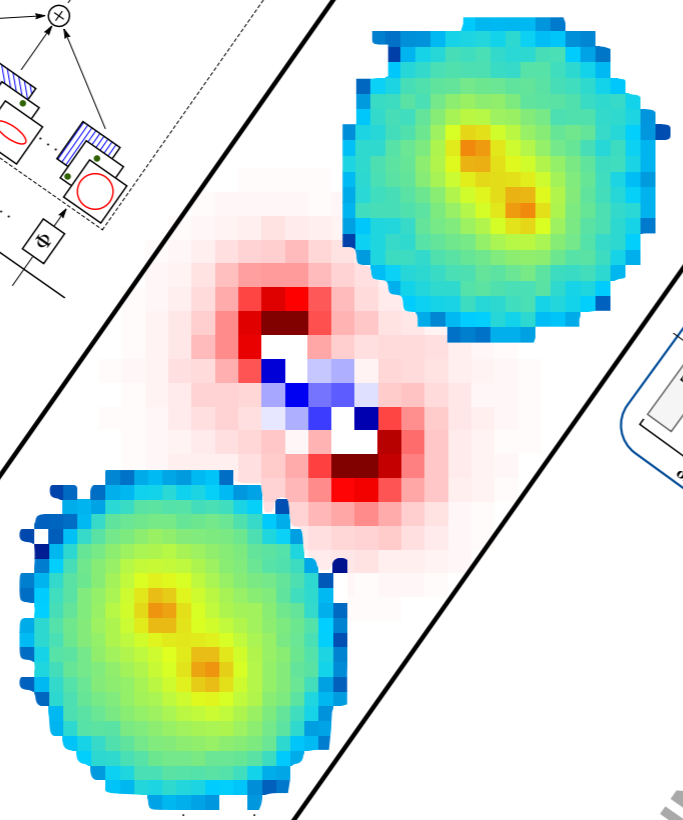
# Deep neural networks for HEP classification



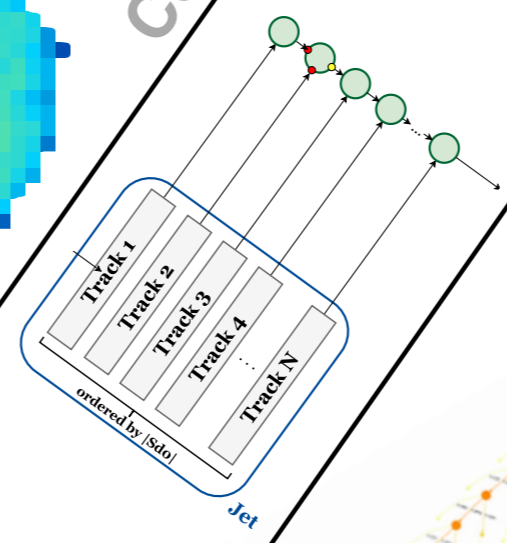
**Variable sets**

**Deep sets**

**Images**



**Sequences**



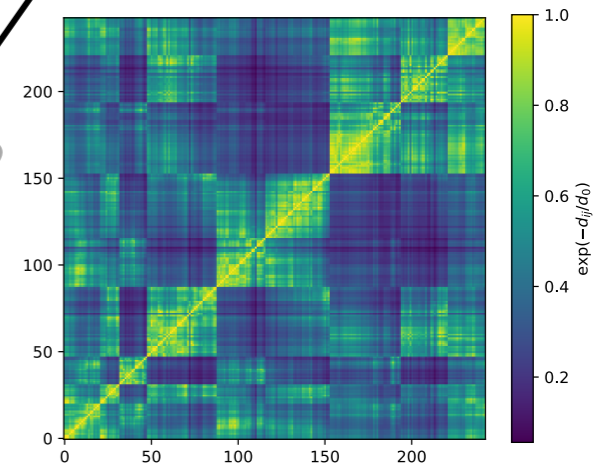
**Recurrent NNs**

**Trees**



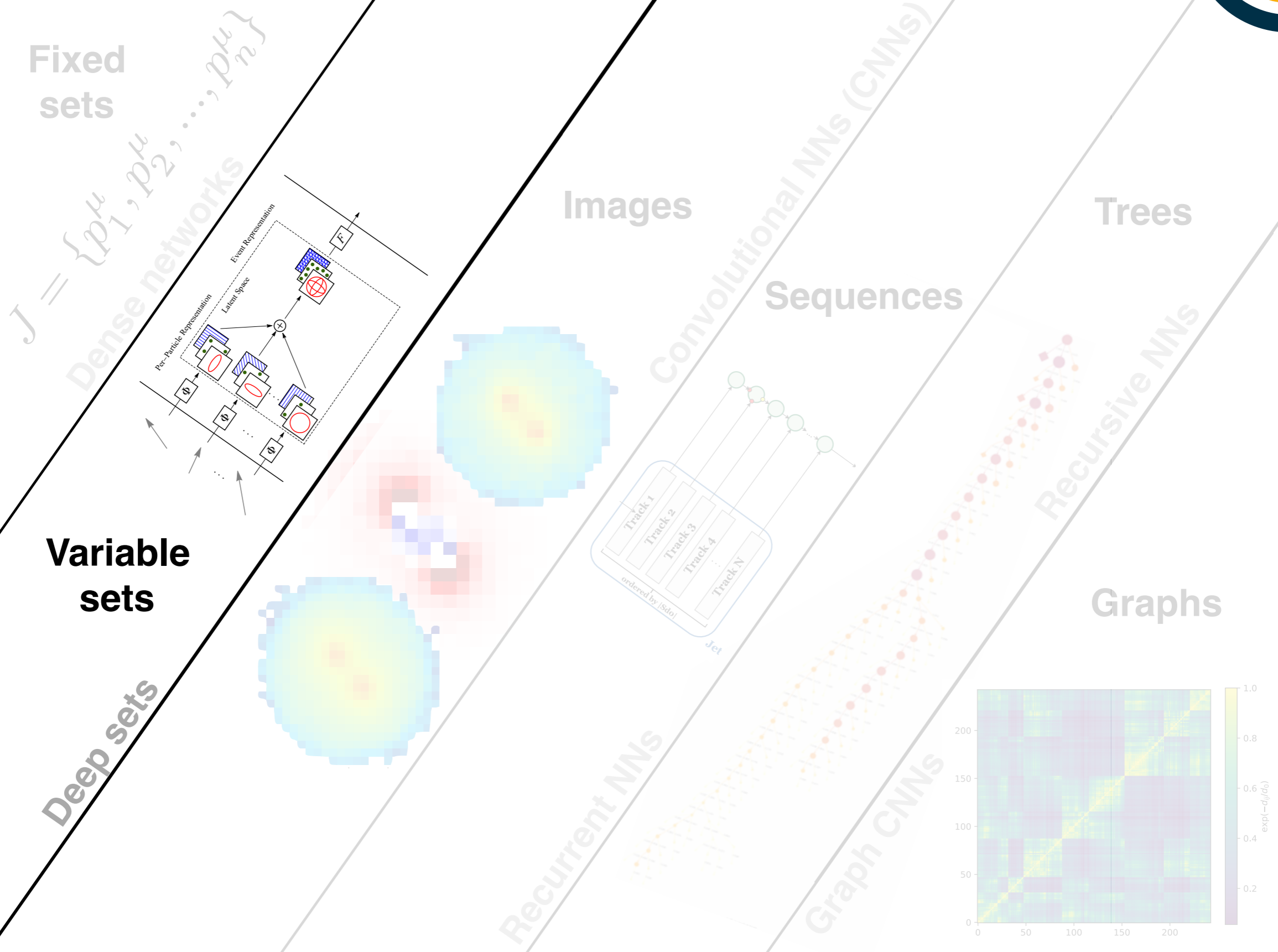
**Graphs**

**Graph CNNs**





# Deep neural networks for HEP classification



A challenge with sequence learning is that thanks to quantum mechanics, there is often no unique order.

A common scenario is that we have a variable-length **SET** of particles and we would like to learn from them directly.

Solution: set learning / point cloud approaches

# Solution 1: Deep sets / Particle flow Networks

50

Factorize the problem into two networks: one that **embeds the set into a fixed-length latent space** and one **that acts on a permutation invariant operation** on that latent space:

$$f(\{x_1, \dots, x_M\}) = F \left( \sum_{i=1}^M \Phi(x_i) \right)$$

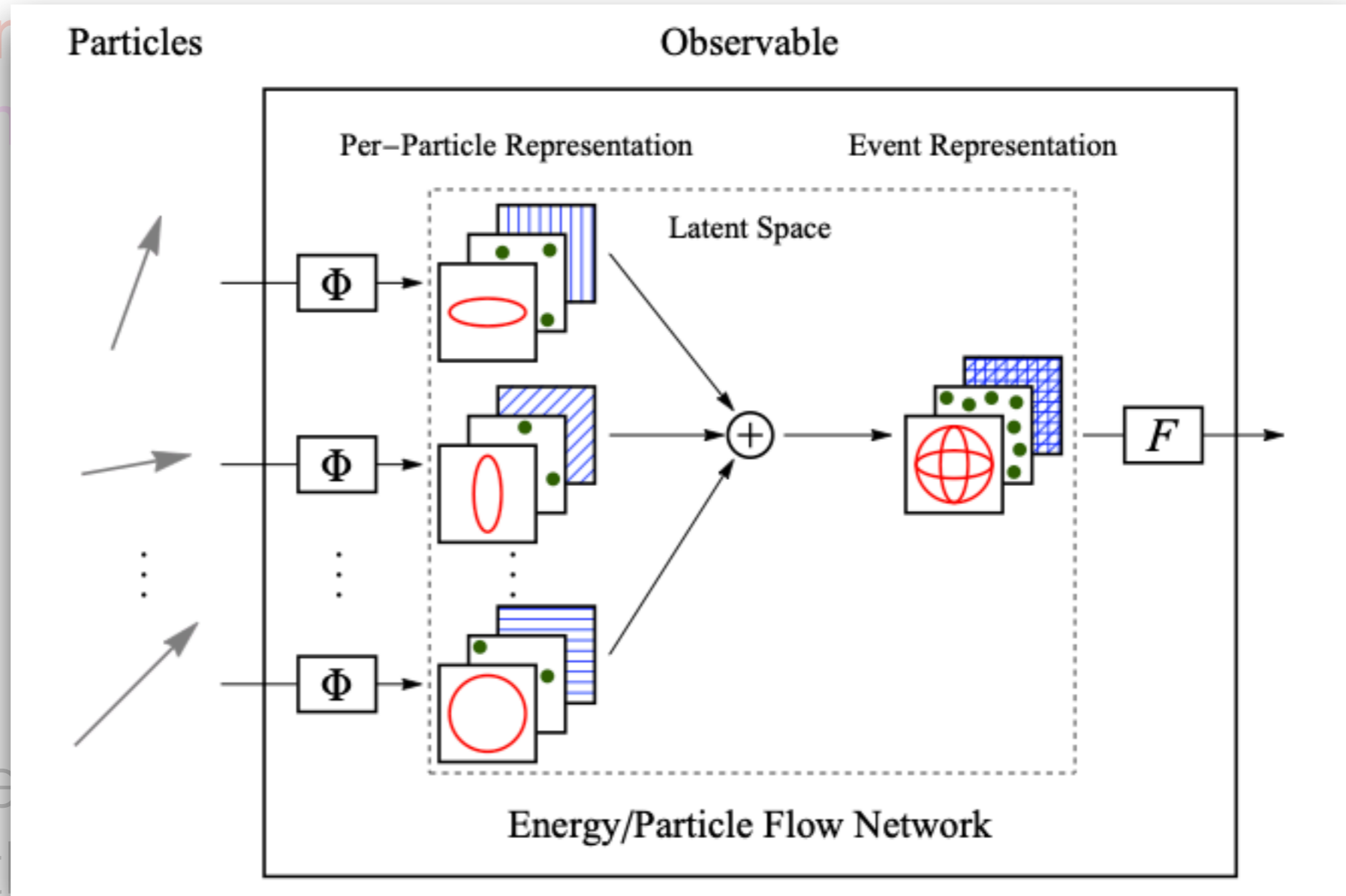
Due to the sum, this structure can operate on any length set and the order of the inputs doesn't matter.



# Solution 1: Deep sets / Particle flow Networks

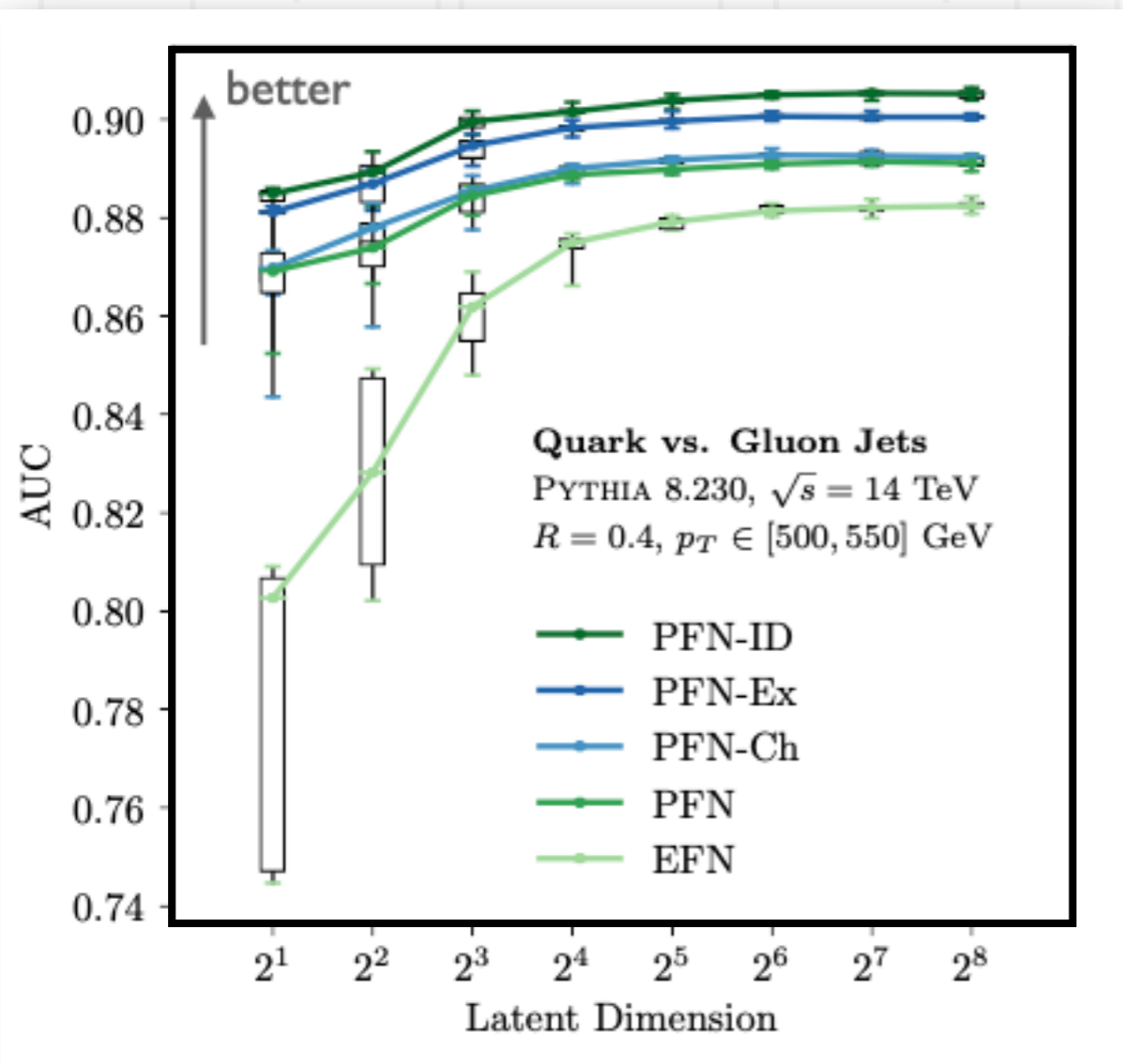
51

Factorize the problem into two networks: one that **embeds the set in a permutation-invariant latent space** that acts on the space:



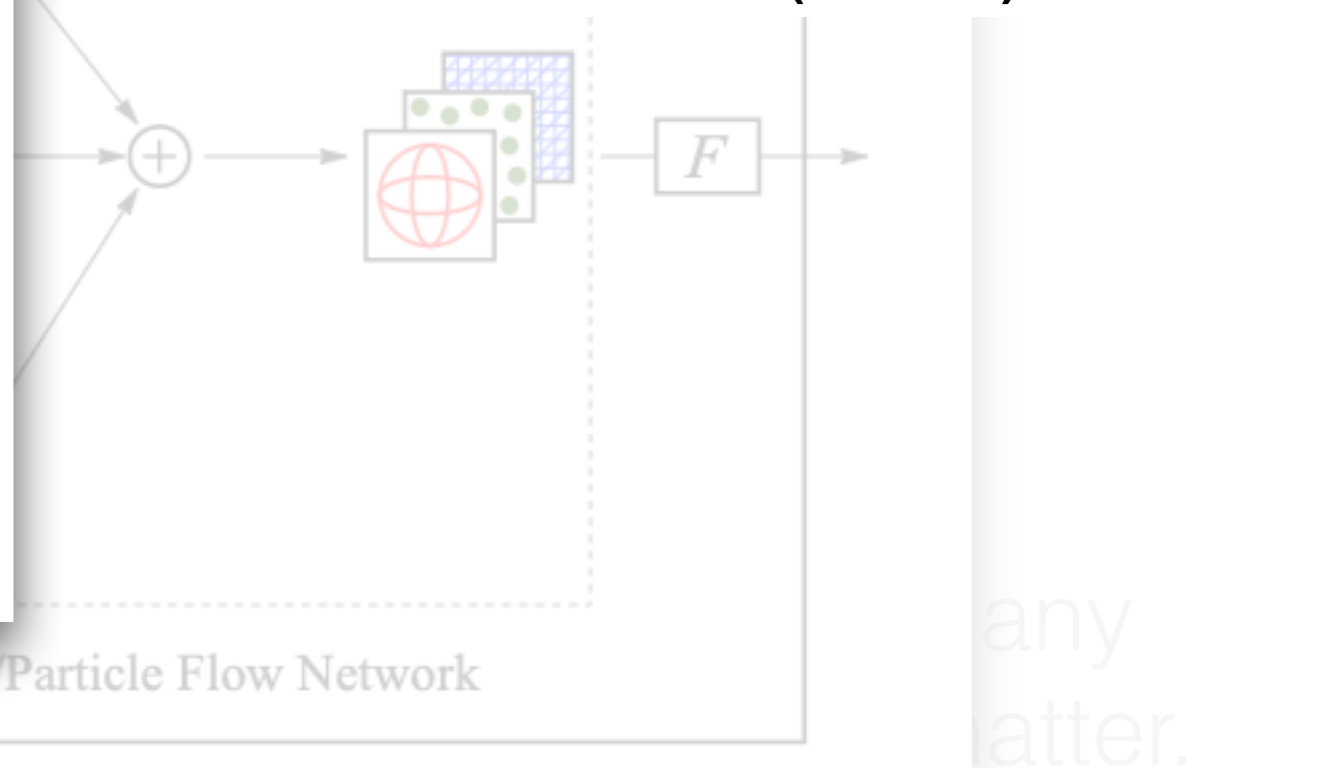
# Solution 1: Deep sets / Particle flow Networks

52



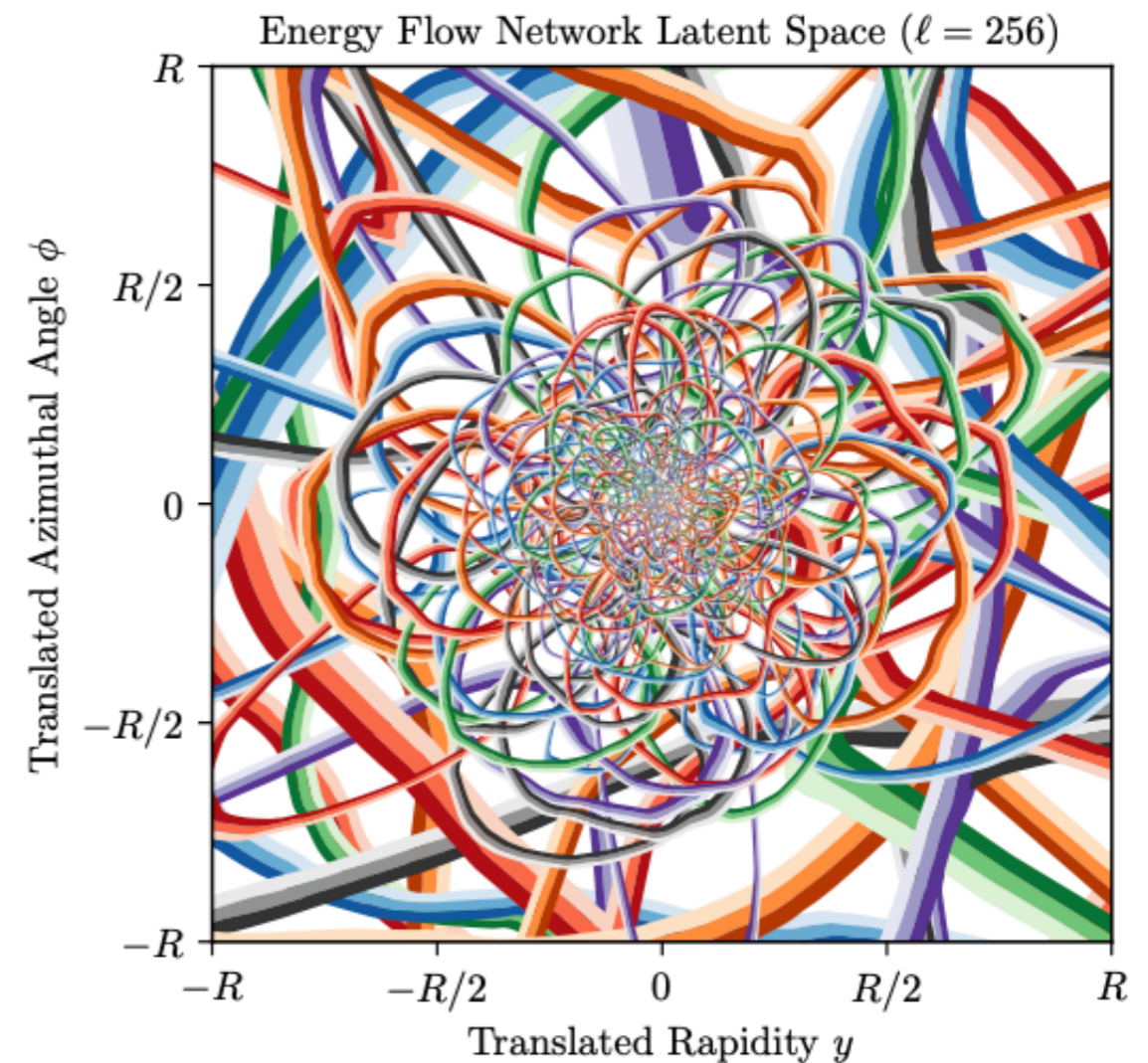
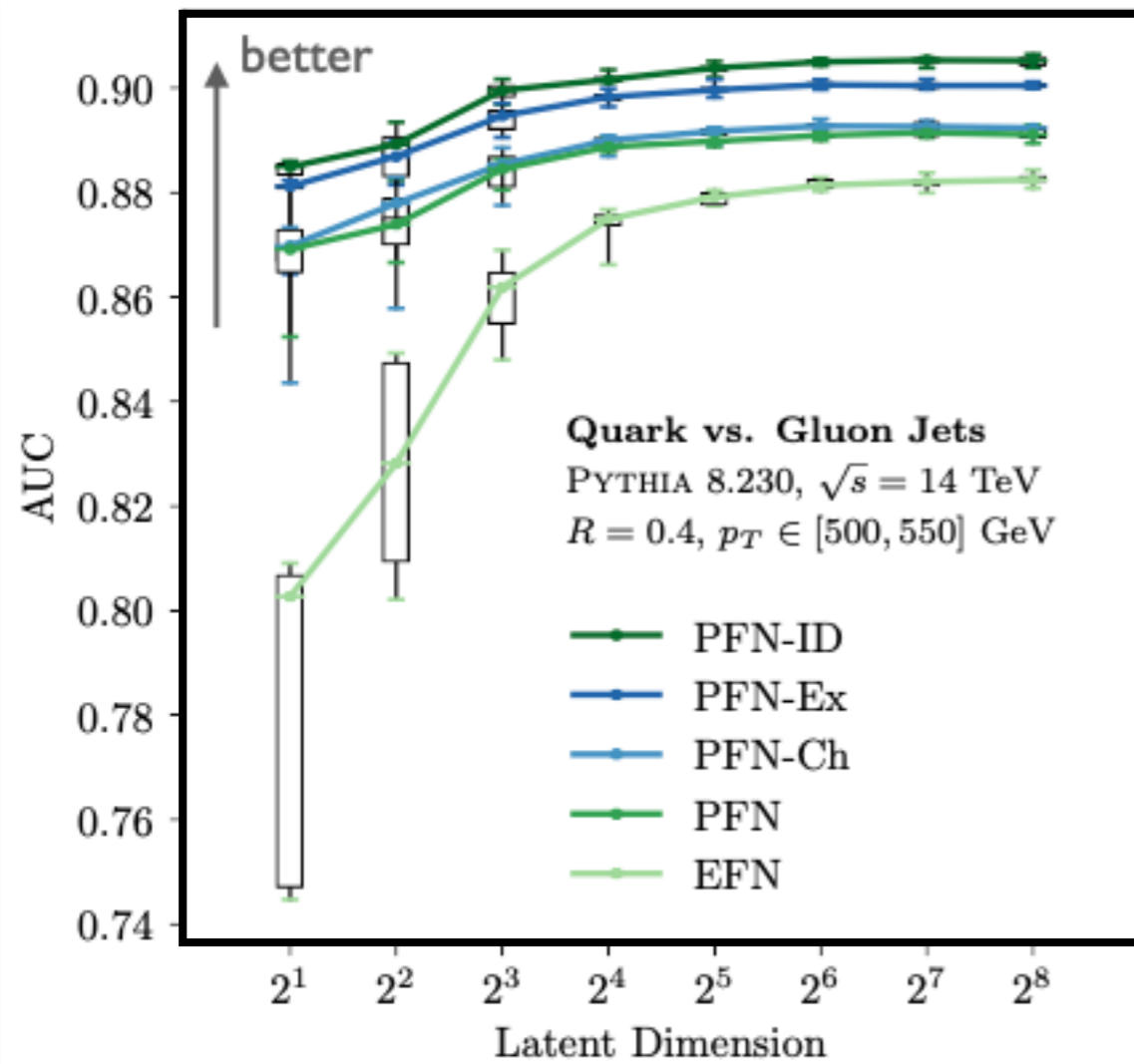
two networks: one that embeds

- Can readily incorporate per-particle features
- Can be made infrared and collinear safe (EFN) safe



# Solution 1: Deep sets / Particle flow Networks

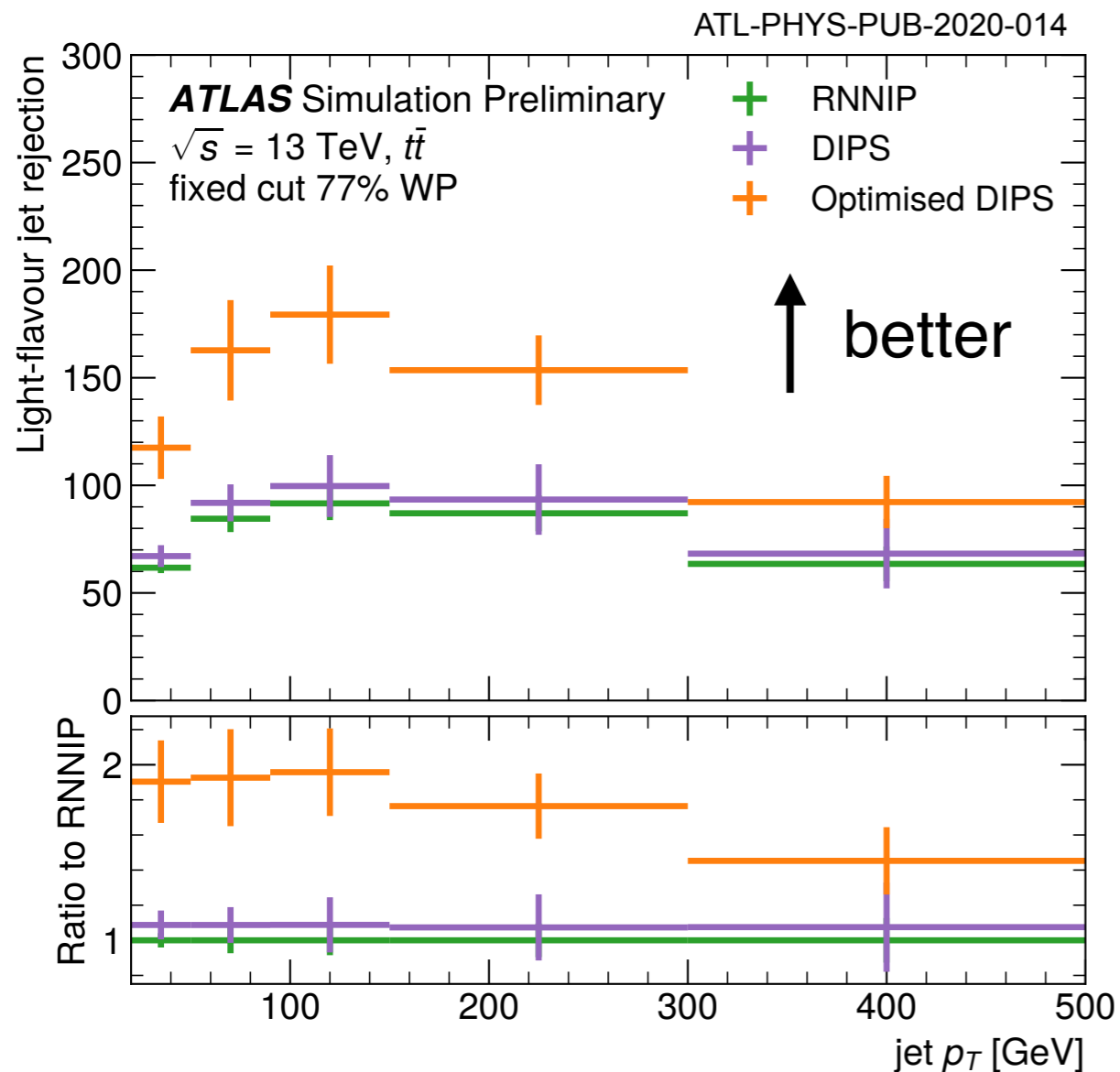
53



Latent space in IRC safe case is interpretable (and predictable!)

# Solution 1: Deep sets / Particle flow Networks

54



Faster to train than RNN so can do R&D on input features to improve overall performance.

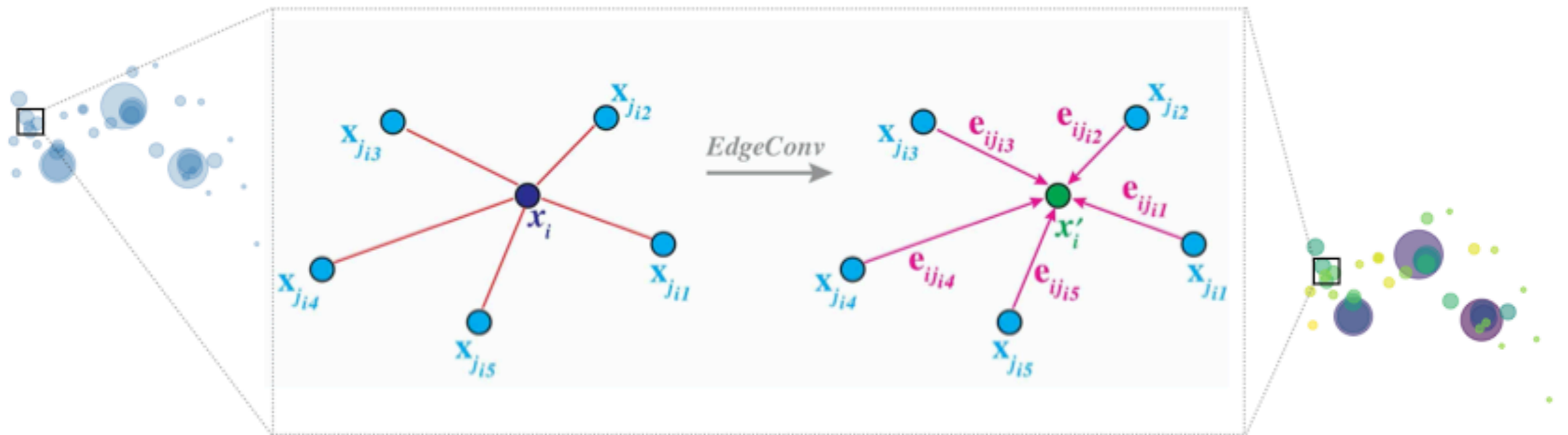
Latent space in IRC safe case is interpretable (and predictable!)

# Solution 2: Graph methods

55

Classic CNNs operate on a fixed grid and are not invariant under the permutation of points

Can generalize CNNs to act on graphs



Need to define distances using particle properties



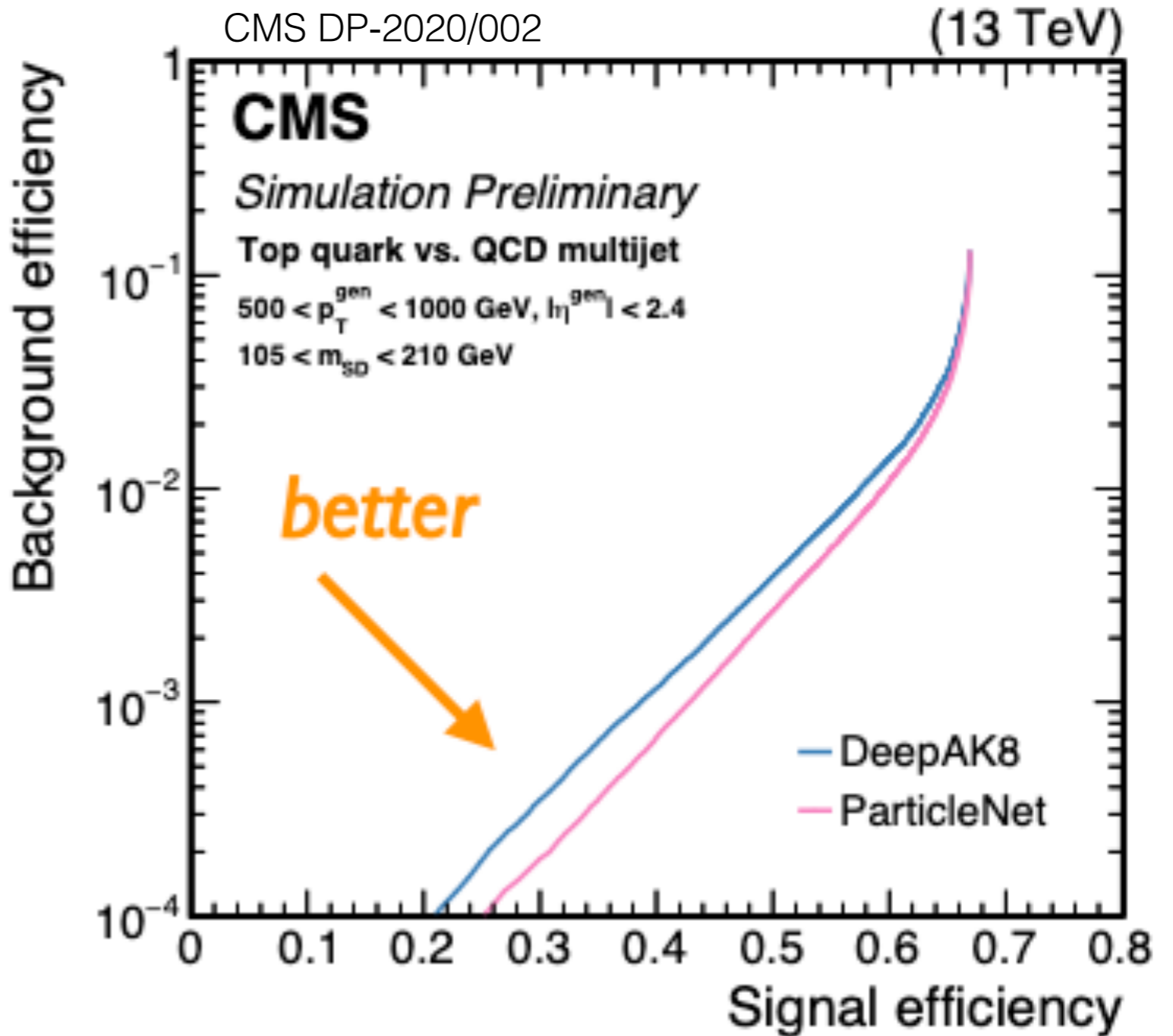
# Solution 2: Graph methods

Classic CNNs are not invariant under rotation

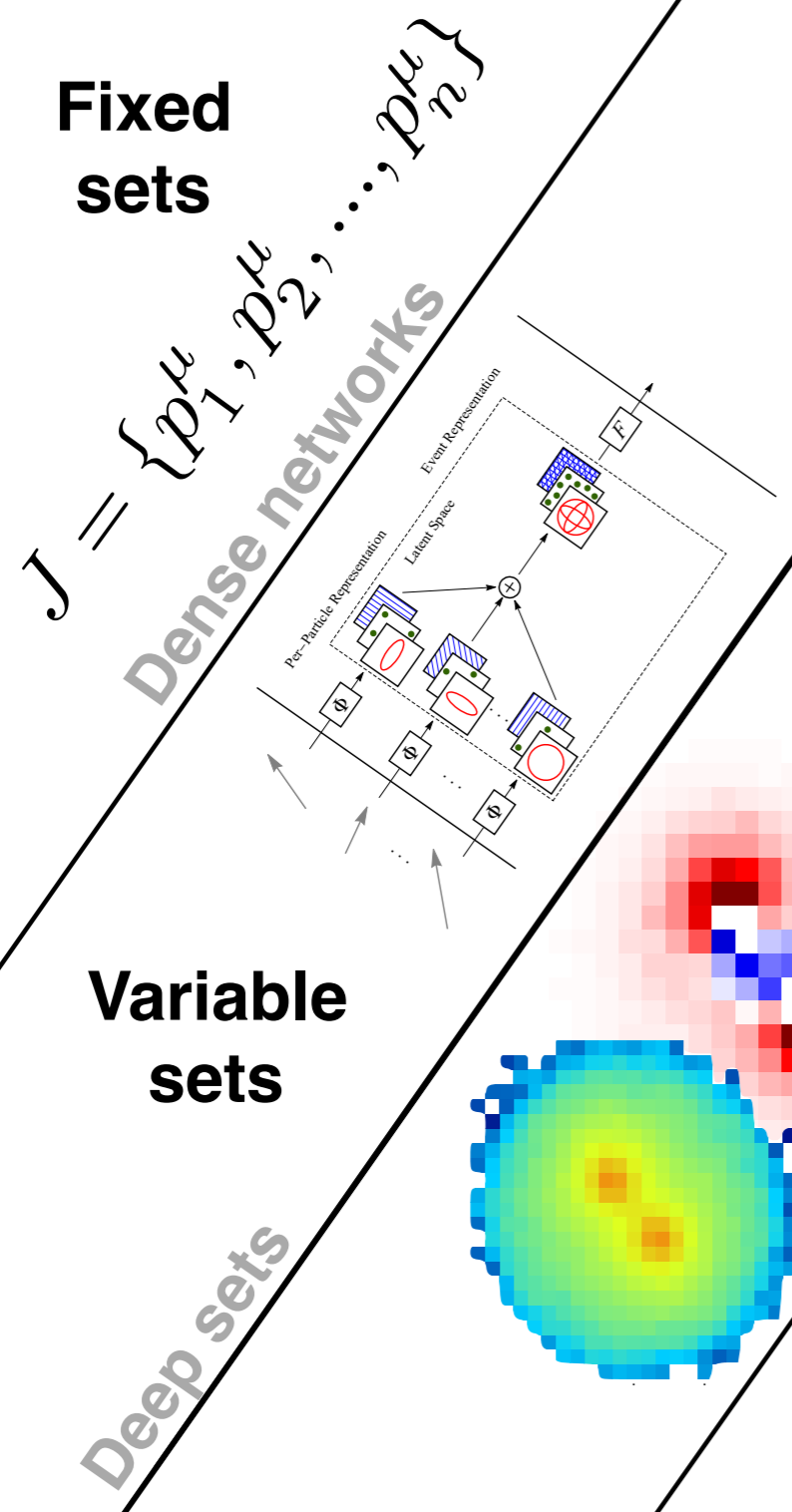
Can generalize to new classes

Competitive performance to other state-of-the-art methods

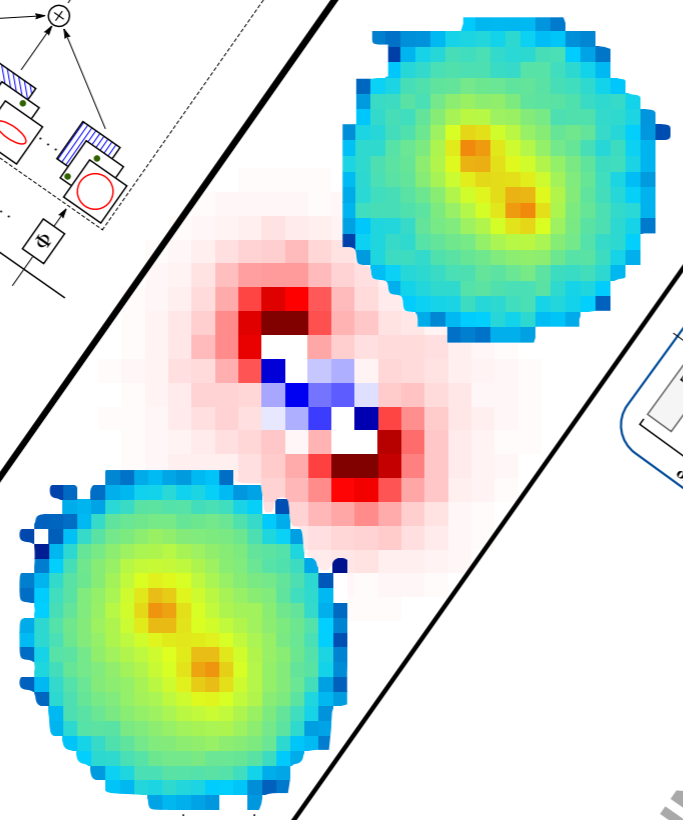
Need to define distance



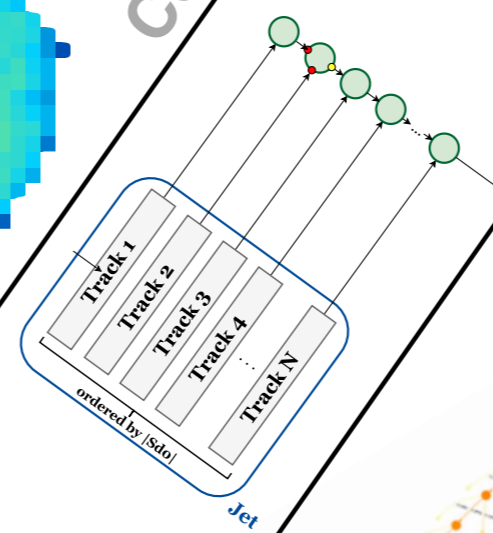
# Deep neural networks for HEP classification



**Images**



**Sequences**



**Recurrent NNs**

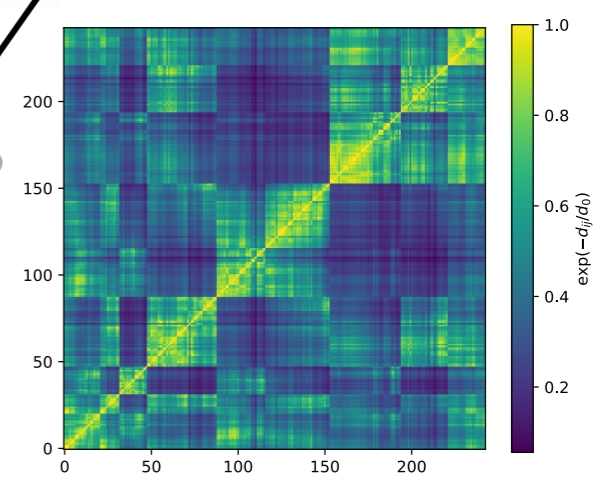
**Convolutional NNs (CNNs)**

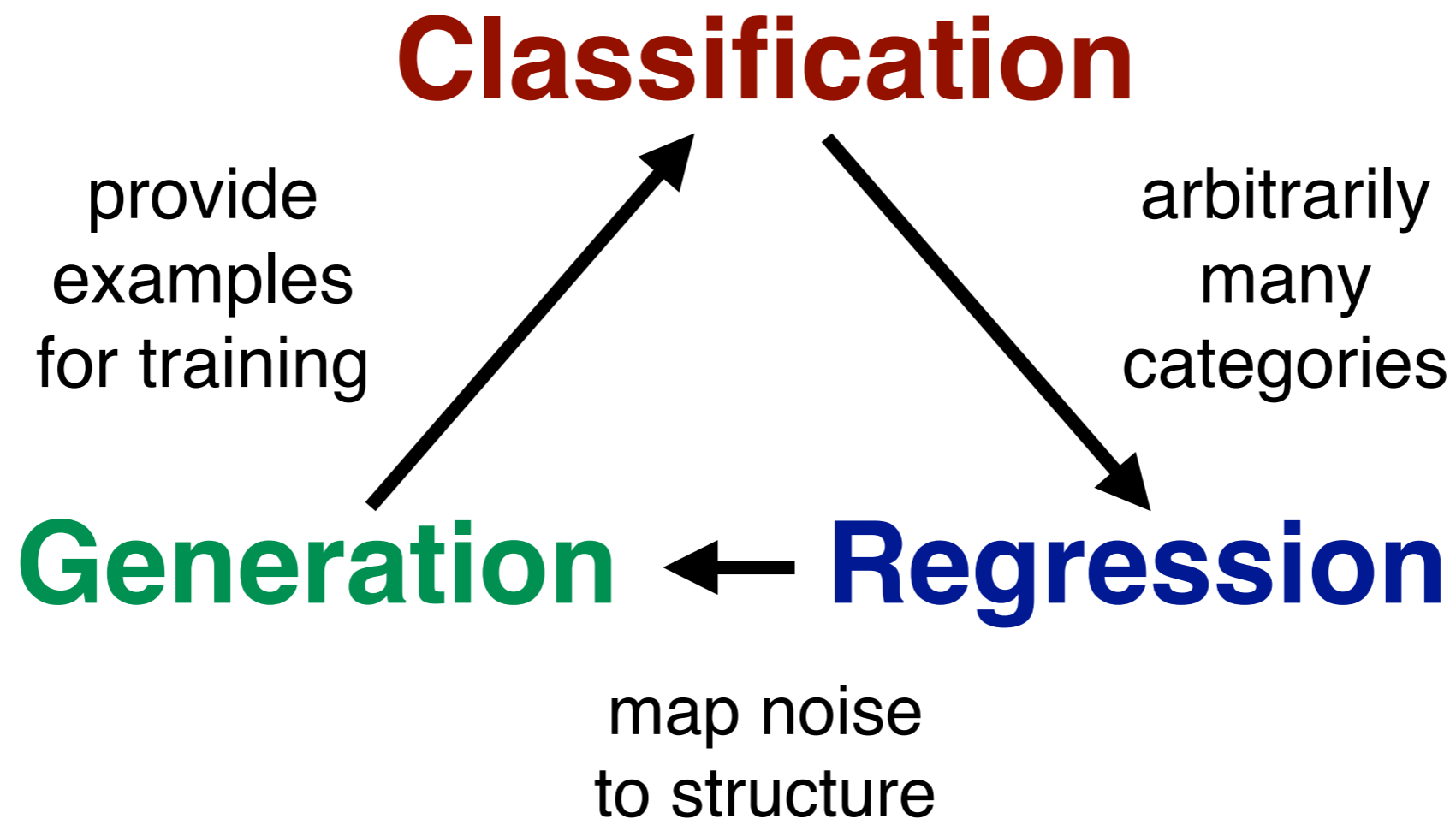
**Graph CNNs**

**Trees**



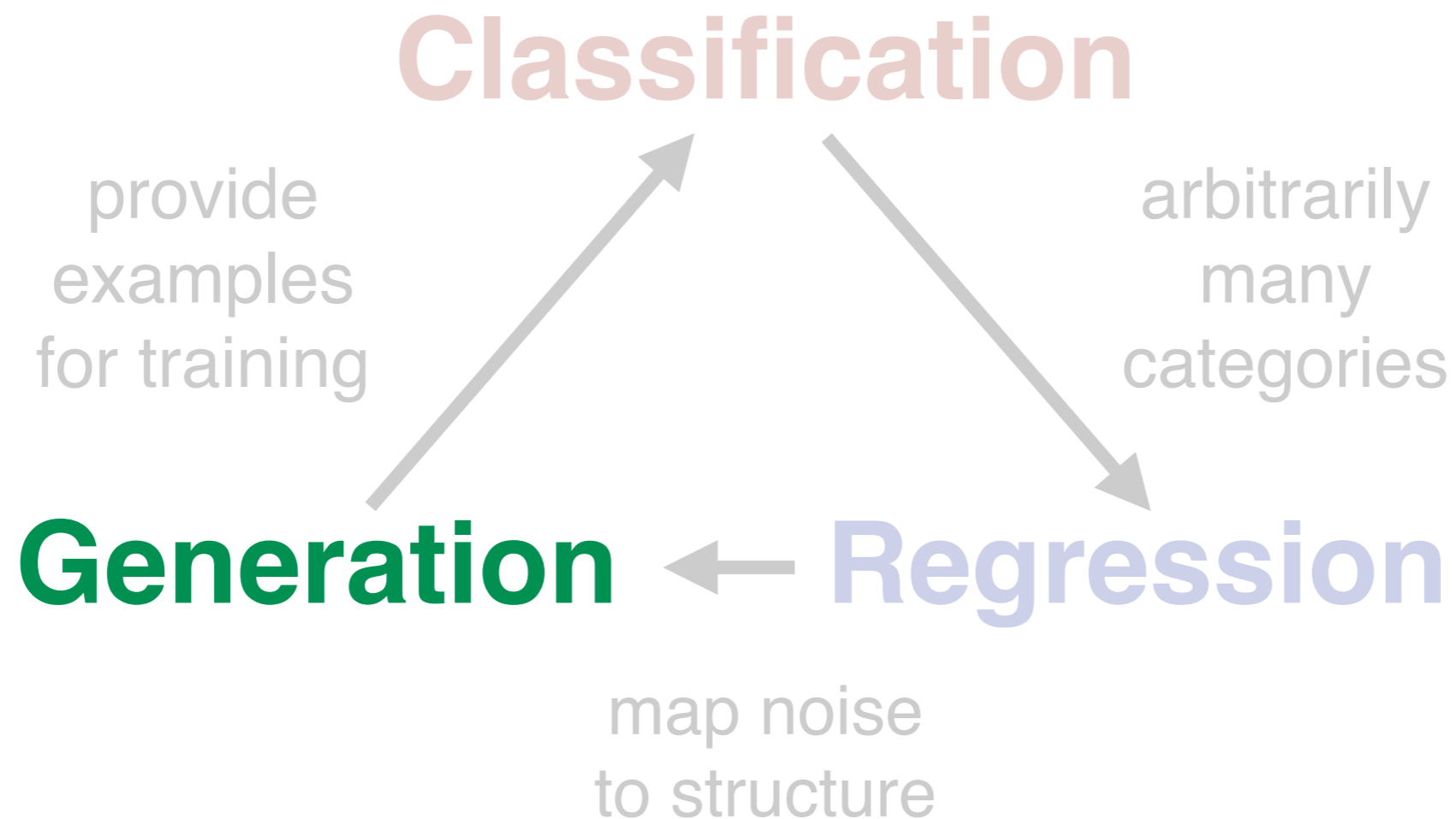
**Graphs**





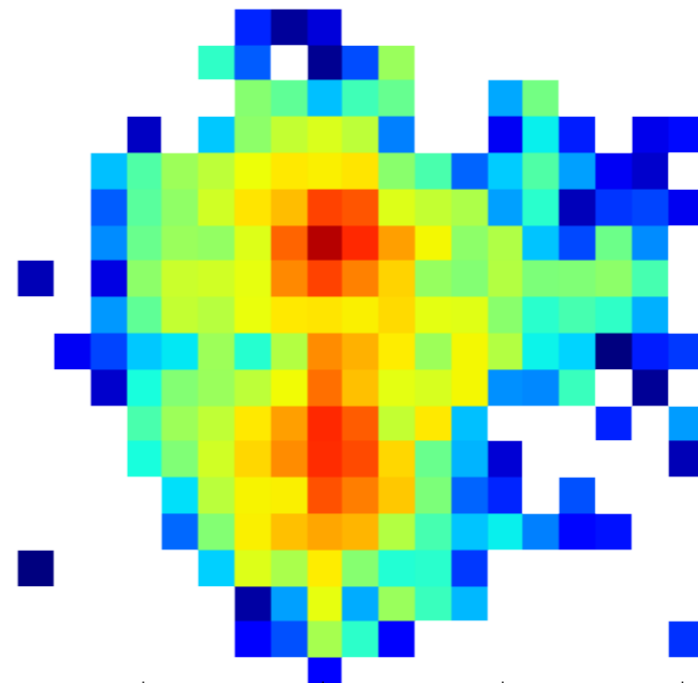
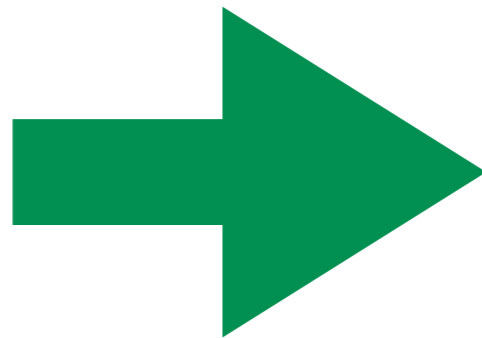
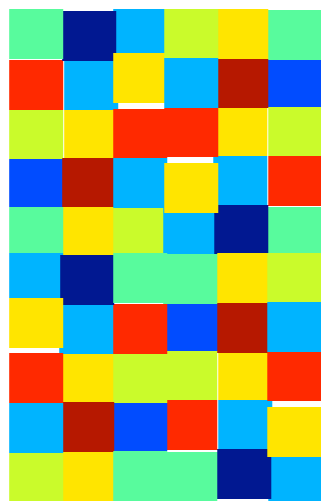
+related topics like anomaly detection, simulation-based inference, ...





+related topics like anomaly detection, simulation-based inference, ...

A **generator** is nothing other than a function that maps random numbers to structure.



Deep generative models: the map is a deep neural network.

## **GANs**

*Generative Adversarial Networks*

## **NFs**

*Normalizing Flows*

## **VAEs**

*Variational Autoencoders*

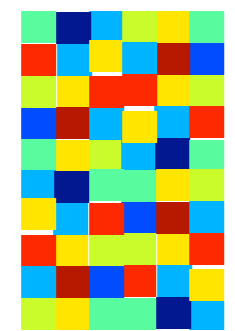
Deep generative models: the map is a deep neural network.

# Introduction: GANs

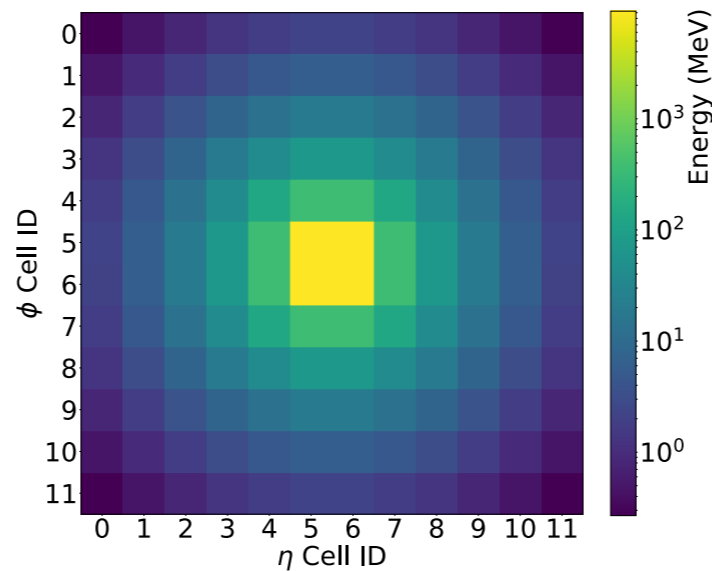
62

Generative Adversarial Networks (GANs):

*A two-network game where one **maps noise to structure** and one **classifies images as fake or real**.*

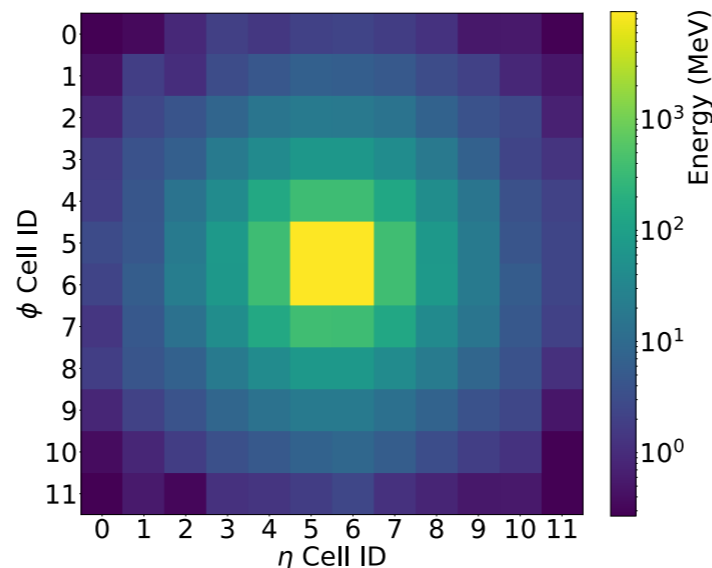


noise



{real, fake}

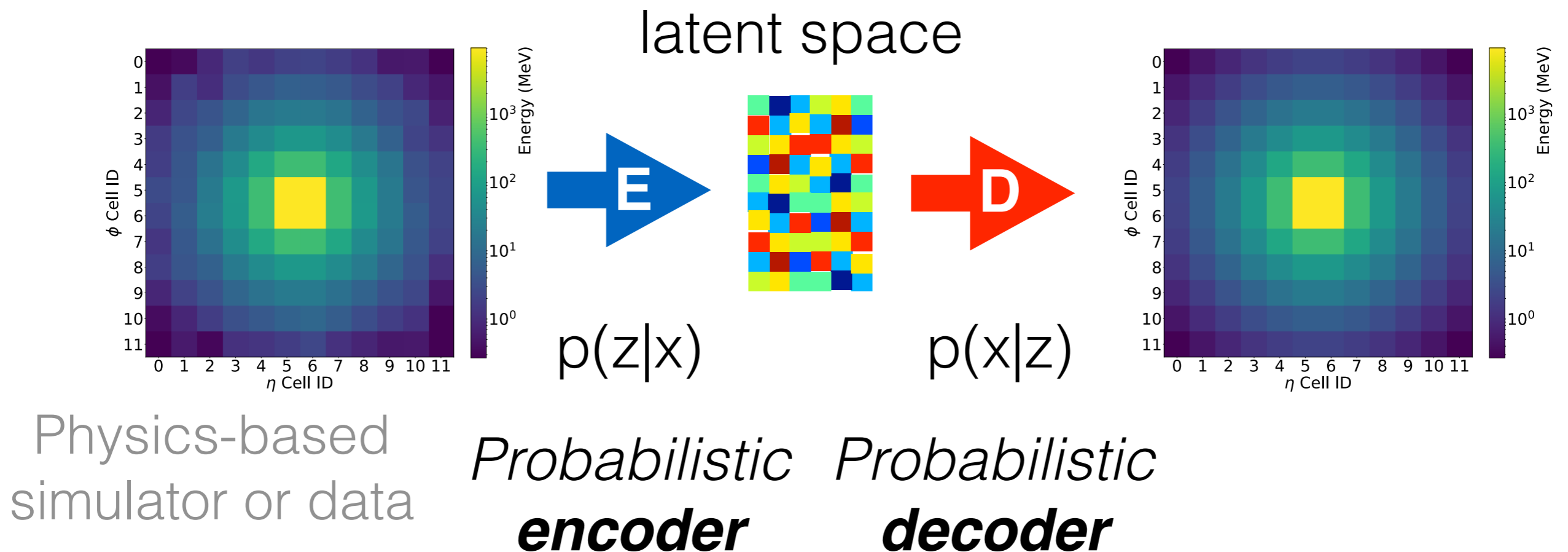
When **D** is maximally confused, **G** will be a good generator



Physics-based simulator or data

# Introduction: VAEs

Variational Autoencoders (VAEs):  
*A pair of networks that embed the data into a latent space with a given prior and decode back to the data space.*

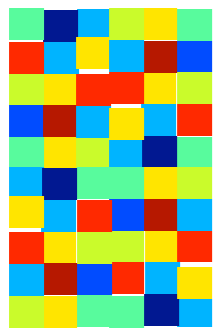


# Introduction: NFs

64

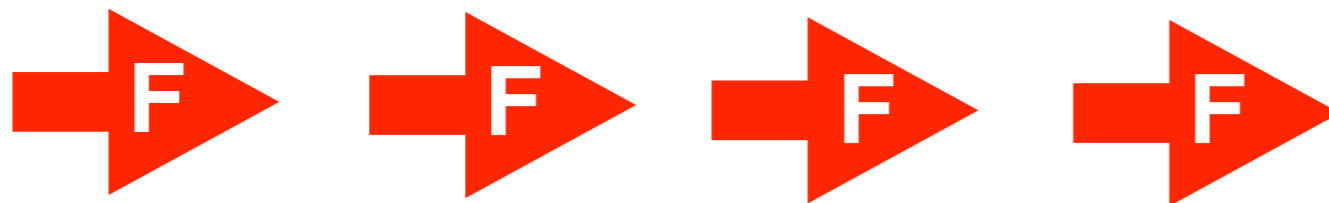
Normalizing Flows (NFs):

*A series of invertible transformations mapping a known density into the data density.*



latent space

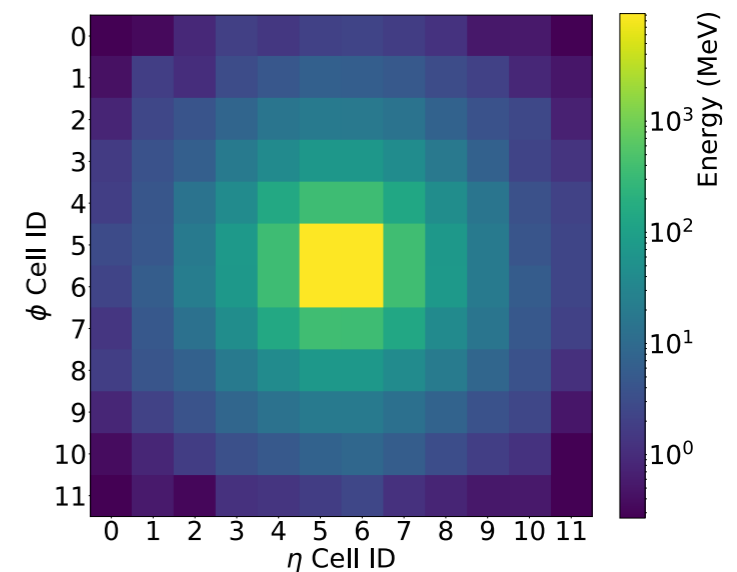
$p(z)$



*Invertible transformations with tractable **Jacobians***

$$p(x) = p(z) \left| \frac{dF^{-1}}{dx} \right|$$

Optimize via maximum likelihood



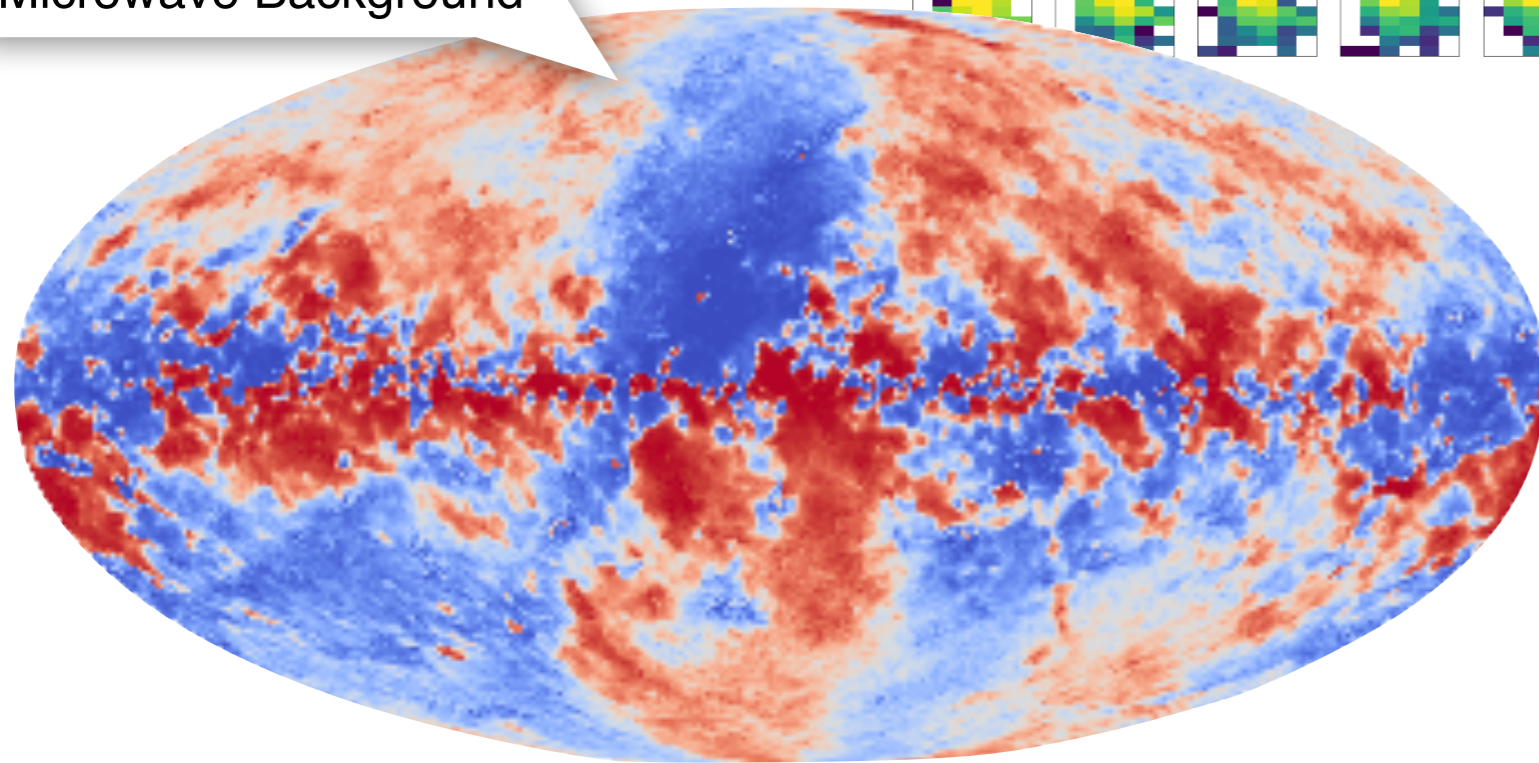
$p(x)$



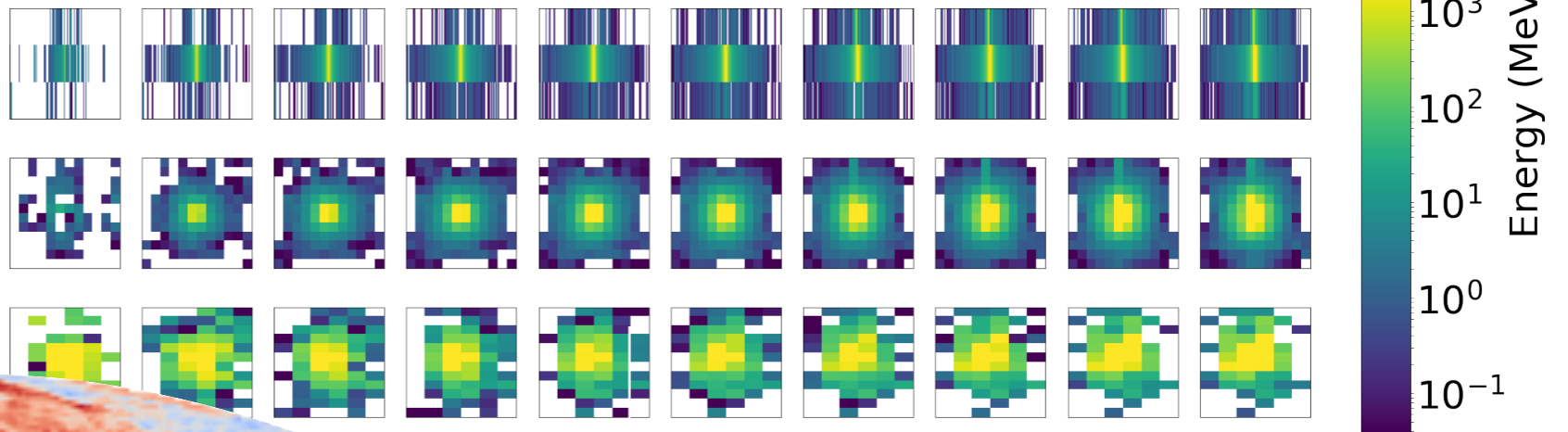
# Generative Models for Particle/Nuclear/Astro

**All of these pictures are fake!**

Synthetic Galactic radiation for Cosmic Microwave Background

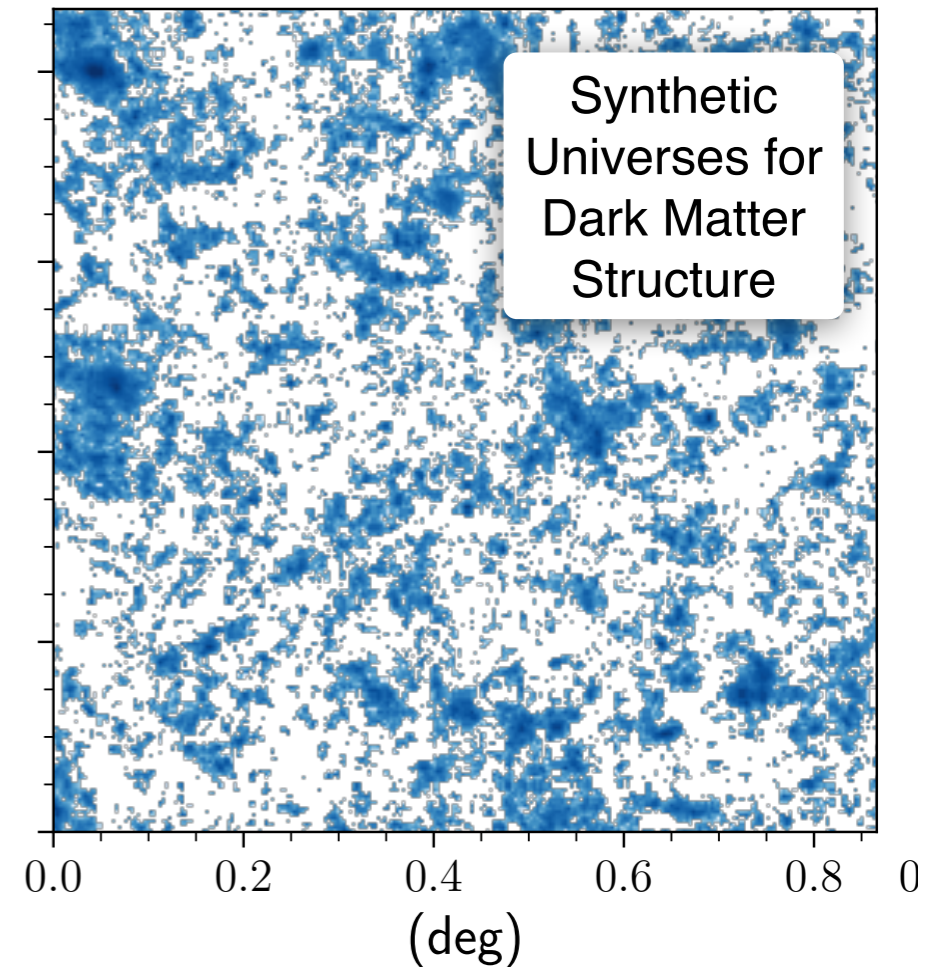


Material Interactions with High Energy Particles

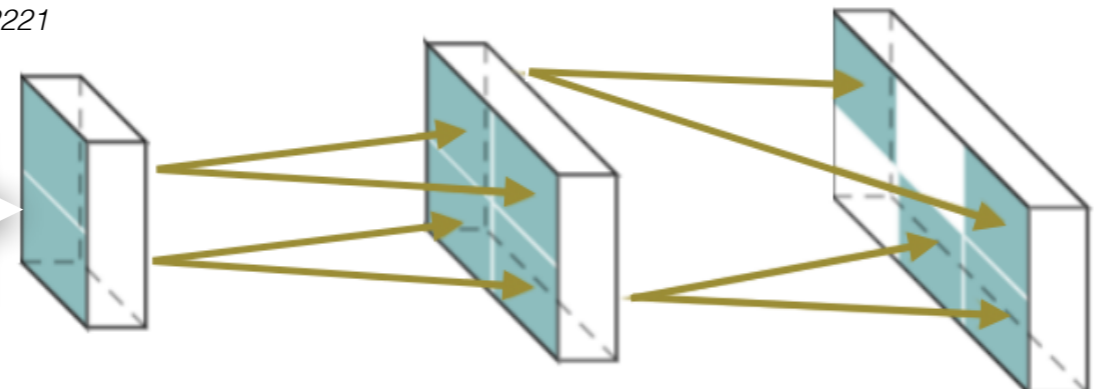


M. Paganini, L. De Oliveira, B. Nachman, Phys. Rev. Lett. 120 (2018) 042003

Synthetic Universes for Dark Matter Structure



The Structure of Radiation in the Quantum Strong Force



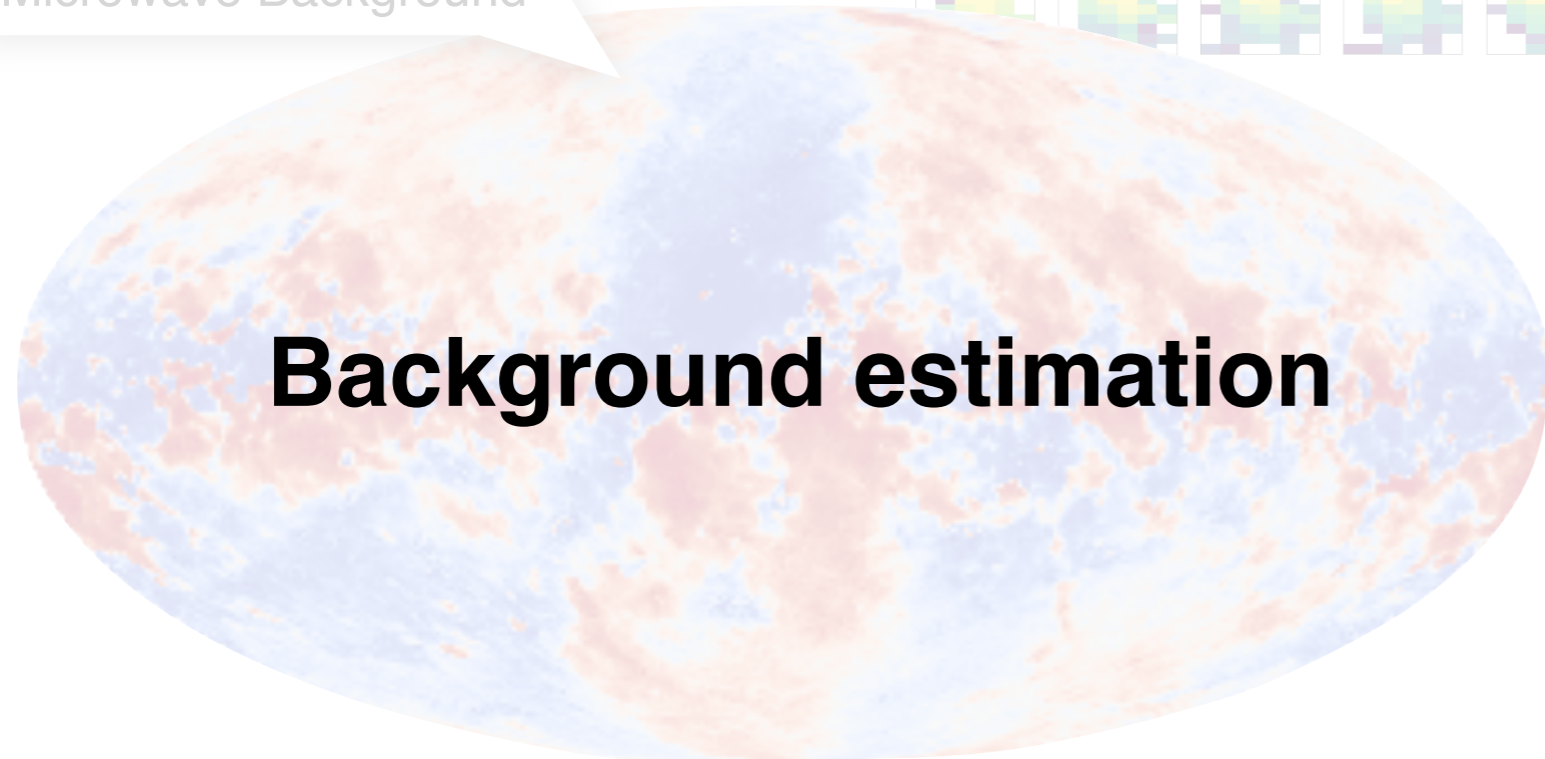
Y. S. Lai, D. Neill, M. Płoskoń, F. Ringer, arXiv:2012.06582

M. Mustafa, et al., Comp. Astrophysics and Cosmology 6 (2019)

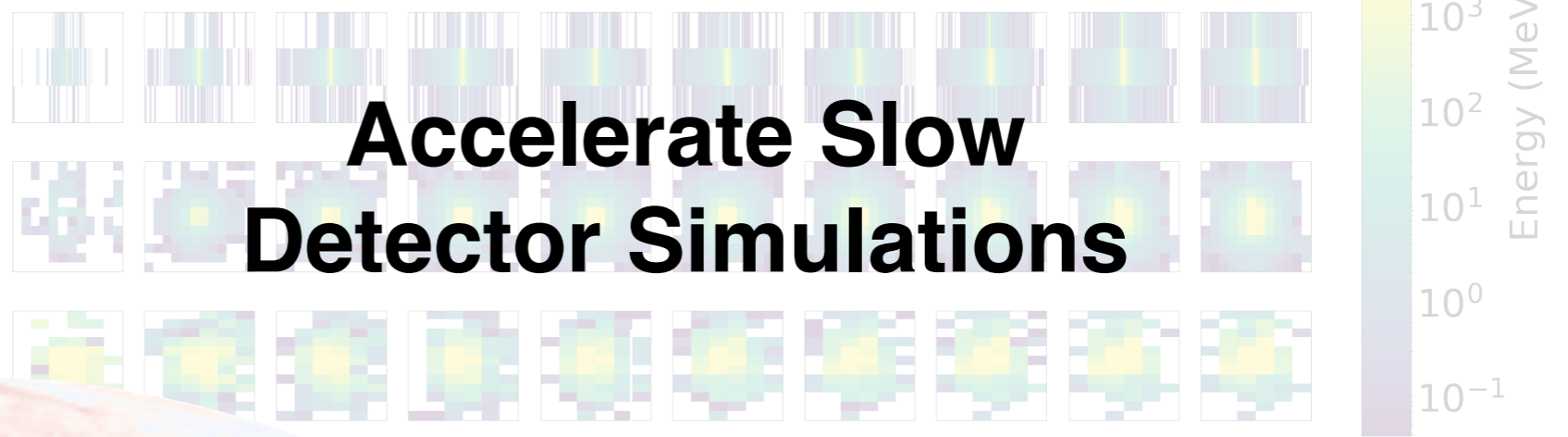
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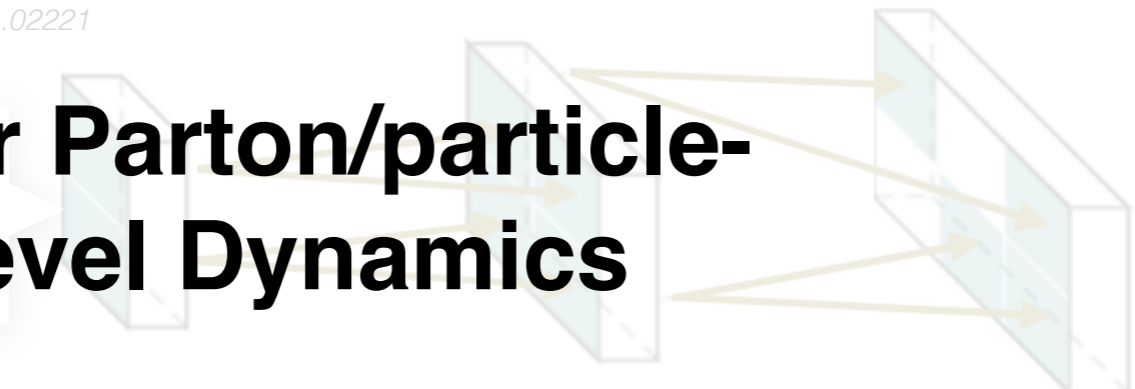


**Accelerate Slow Detector Simulations**

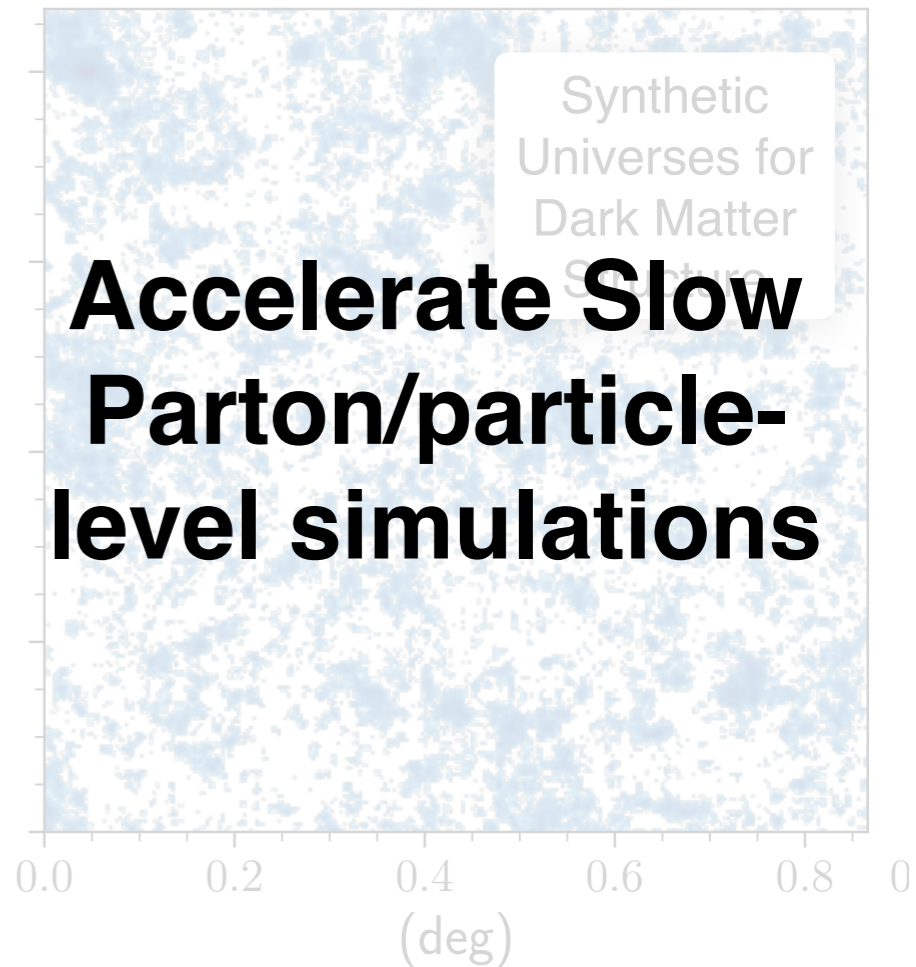
M. Paganini, L. De Oliveira, B. Nachman, Phys. Rev. Lett. 120 (2018) 042003

The Structure of Radiation in the Quantum Strong Force

**Infer Parton/particle-level Dynamics**



Y. S. Lai, D. Neill, M. Płoskoń, F. Ringer, arXiv:2012.06582



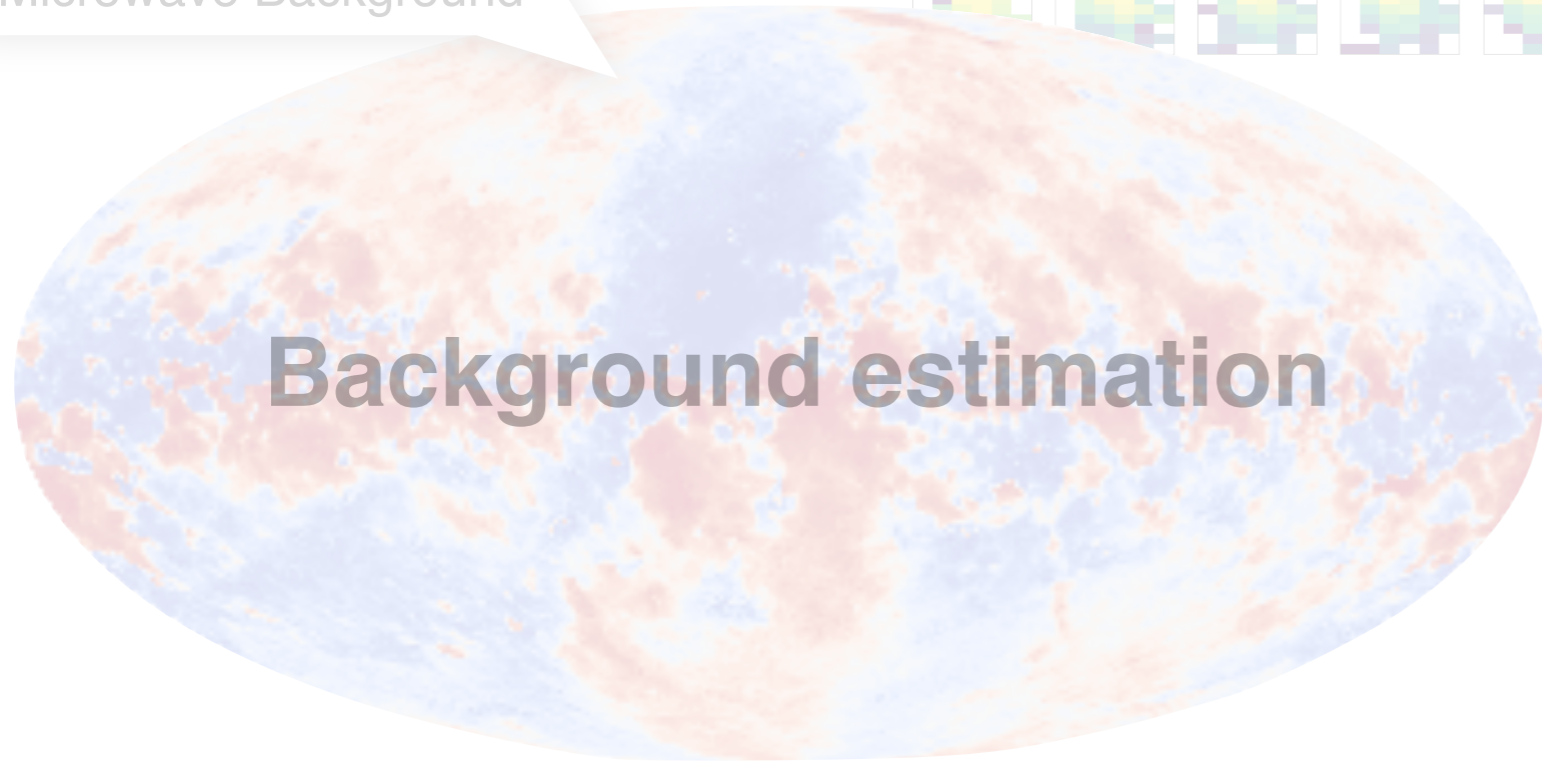
M. Mustafa, et al., Comp. Astrophysics and Cosmology 6 (2019)



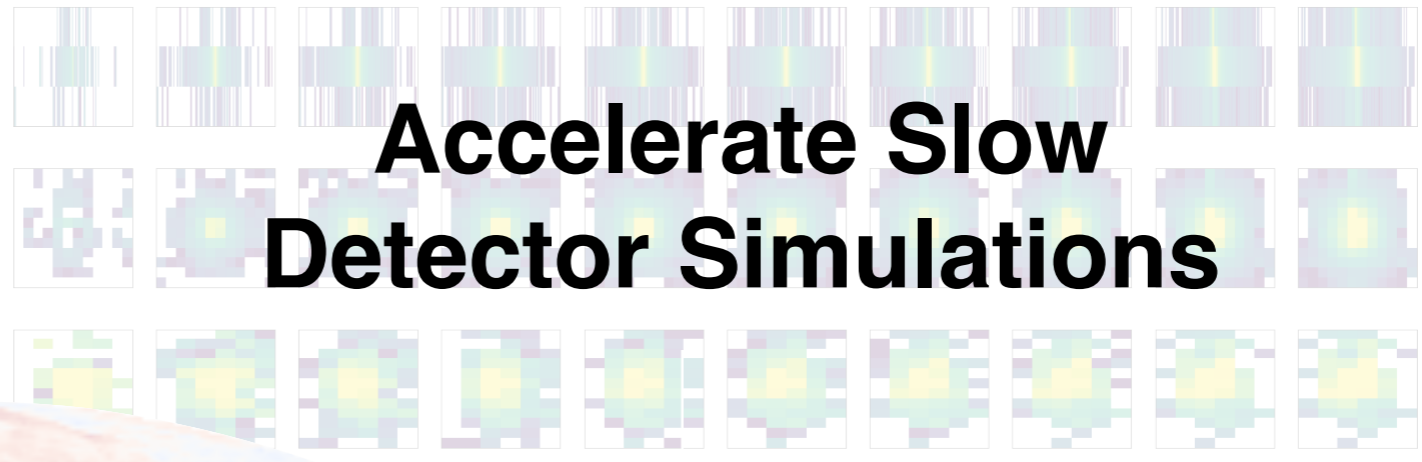
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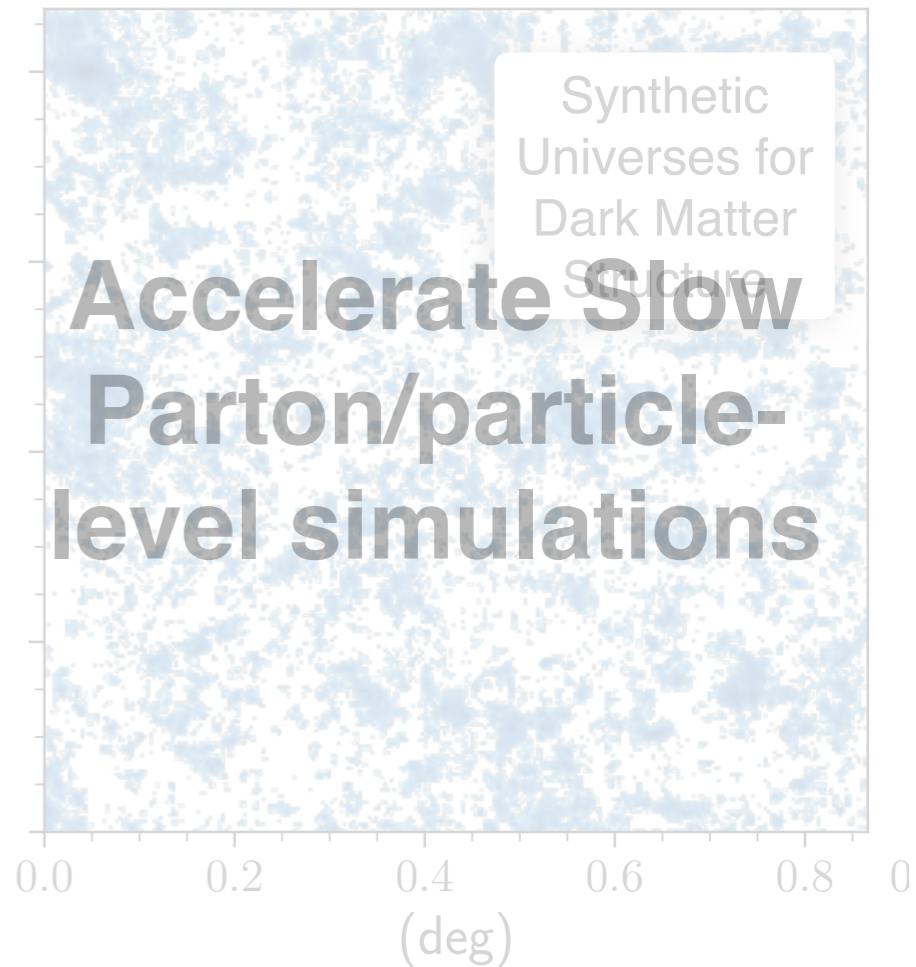
Synthetic Galactic radiation for Cosmic Microwave Background



Material Interactions with High Energy Particles



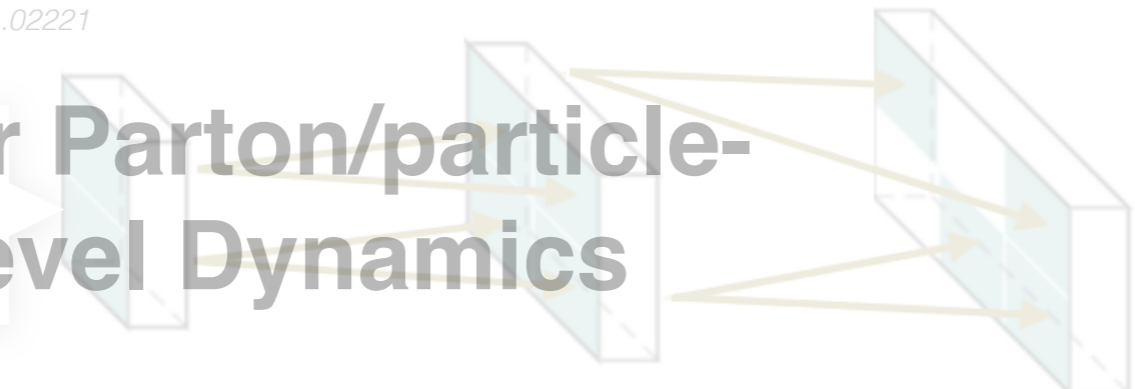
M. Paganini, L. De Oliveira, B. Nachman, Phys. Rev. Lett. 120 (2018) 042003



M. Mustafa, et al., Comp. Astrophysics and Cosmology 6 (2019)

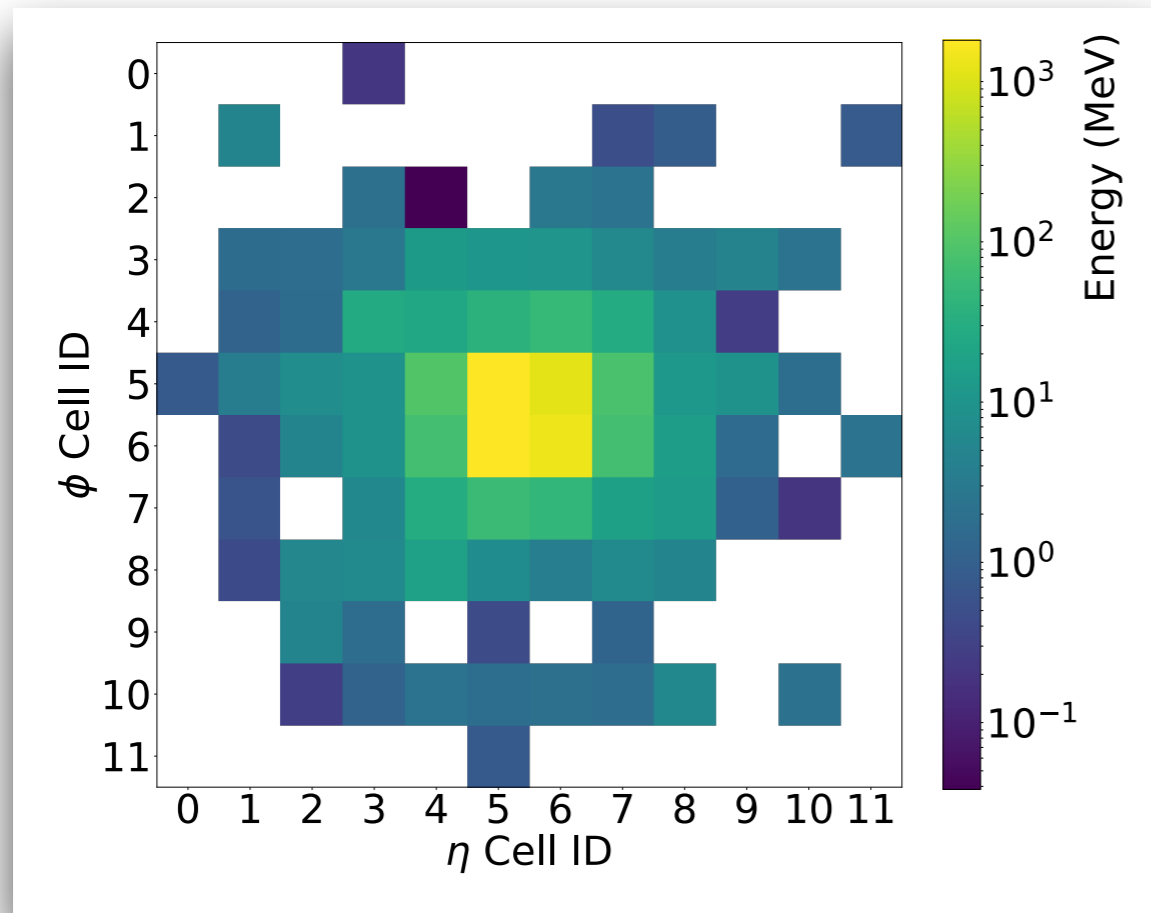
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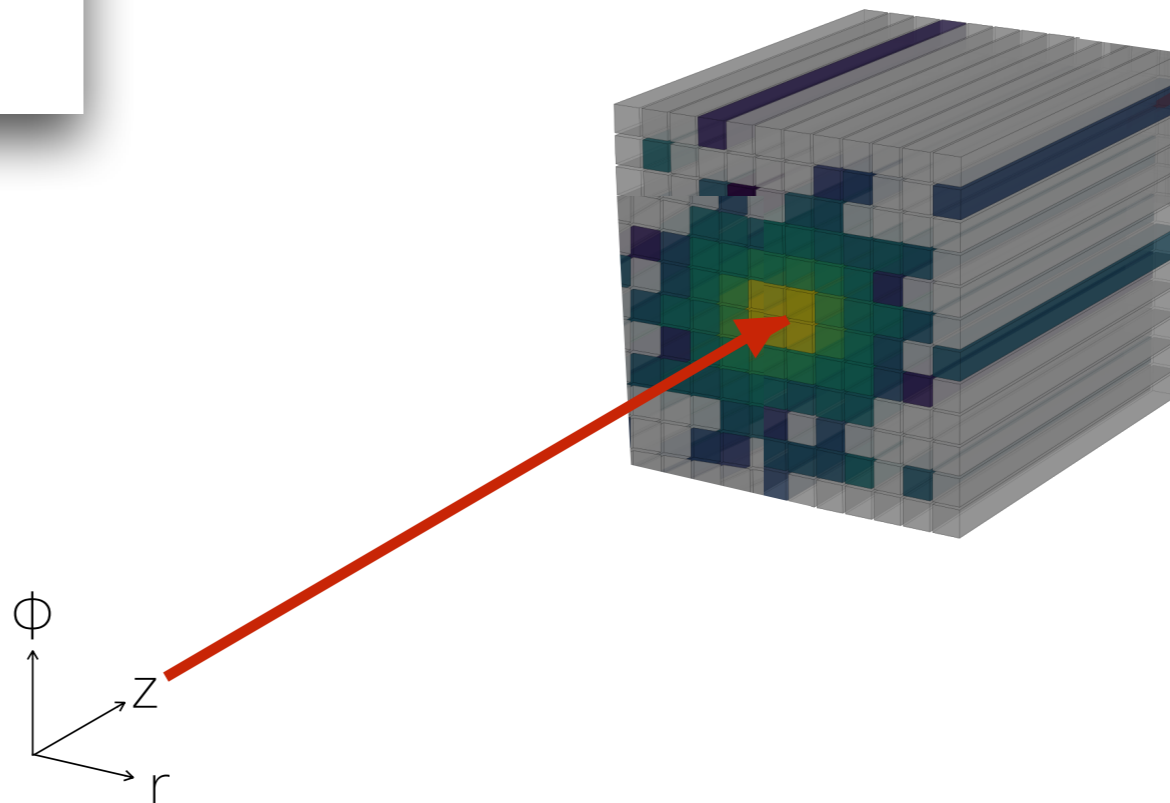
Y. S. Lai, D. Neill, M. Płoskoń, F. Ringer, arXiv:2012.06582

N. Krachmalnicoff and G. Puglisi, arXiv:2011.02221



Calorimeters are often the slowest to simulate  
*stopping particles requires simulating interactions of all energies*

Grayscale images:  
Pixel intensity =  
energy deposited

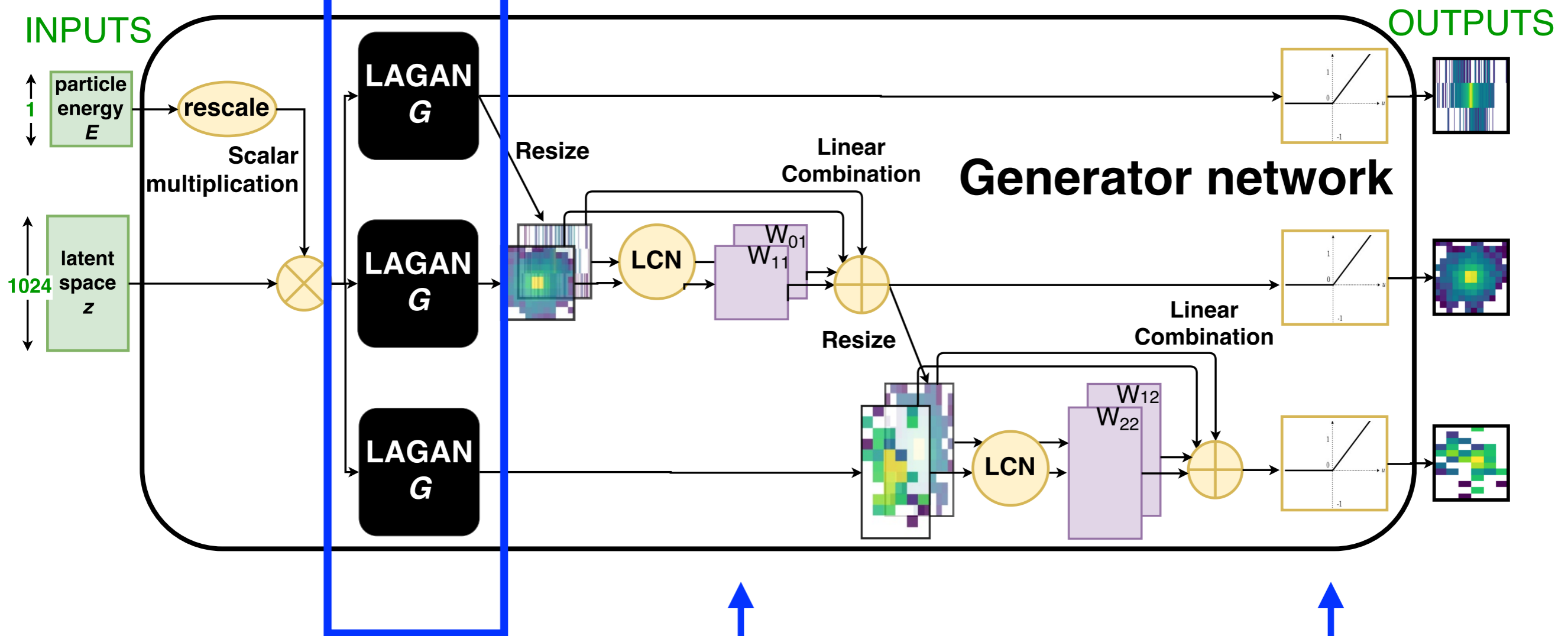




# Introducing CaloGAN

One image per calo layer

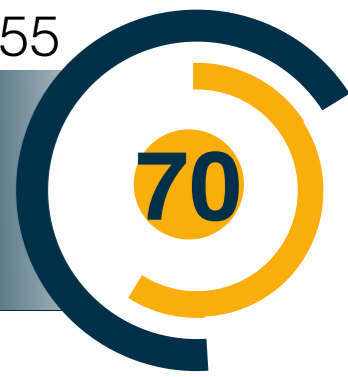
One network per particle type; input particle energy



LA = Locally Aware, like a CNN

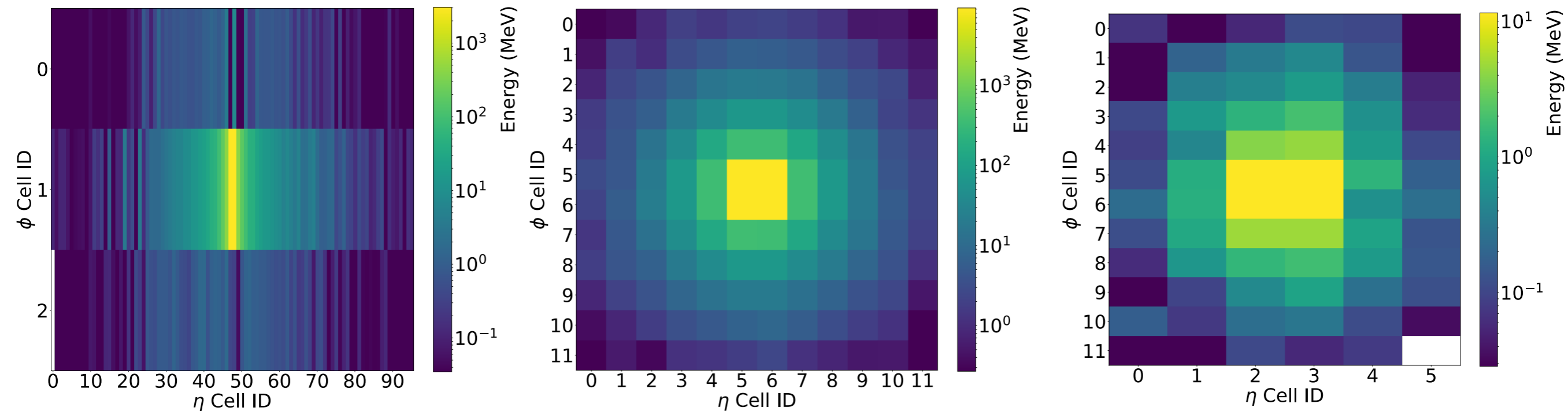
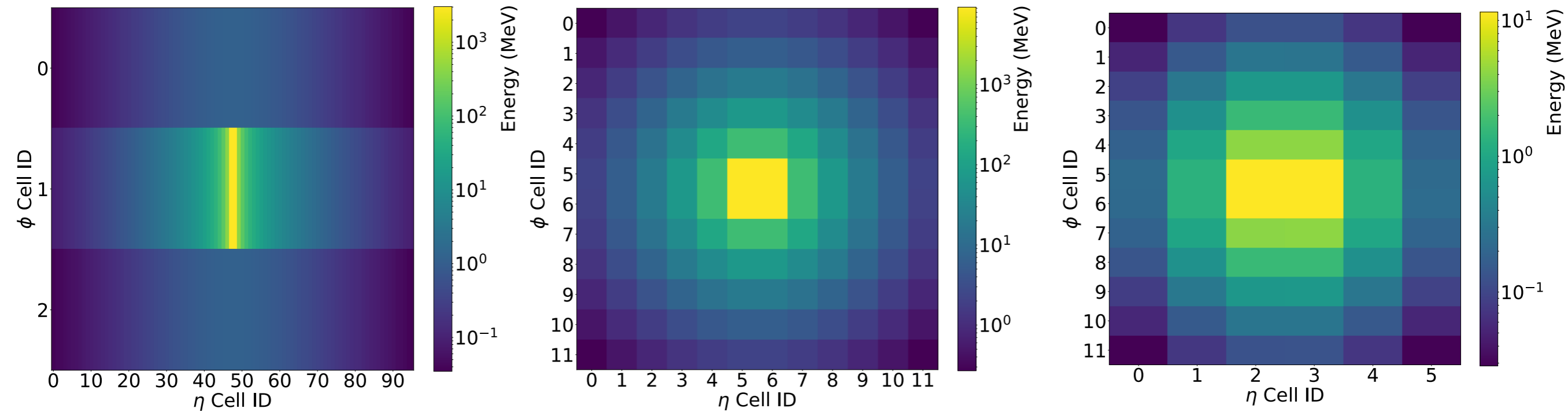
use layer  $i$  as input to layer  $i+1$

ReLU to encourage sparsity



# Performance: average images

## Geant4

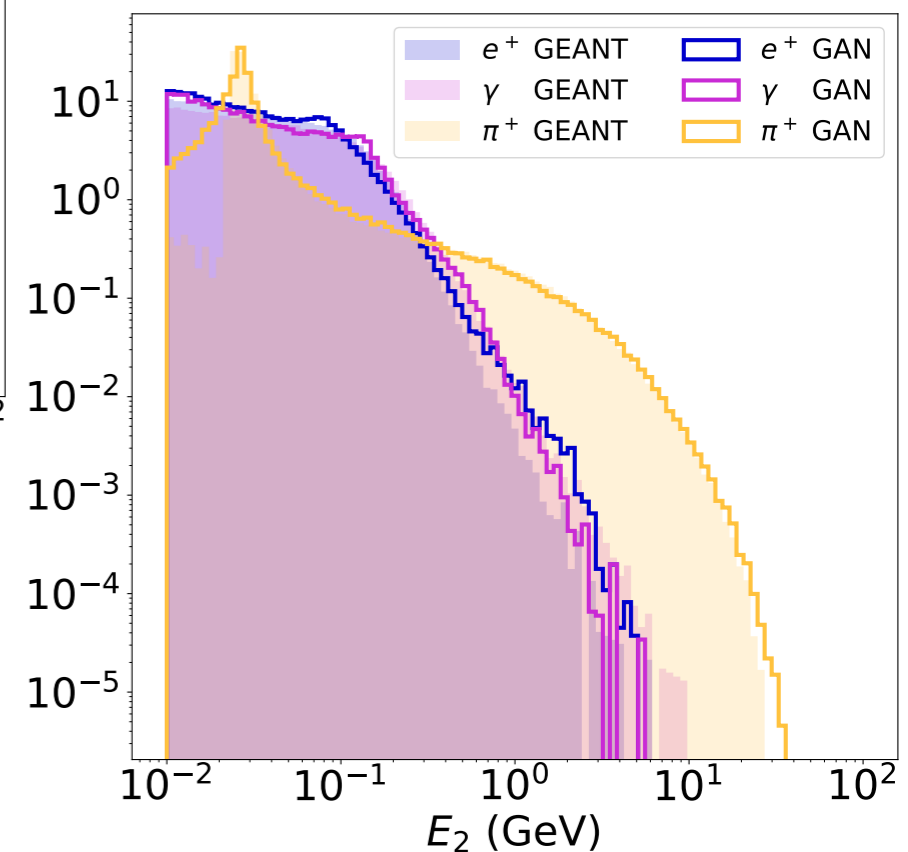
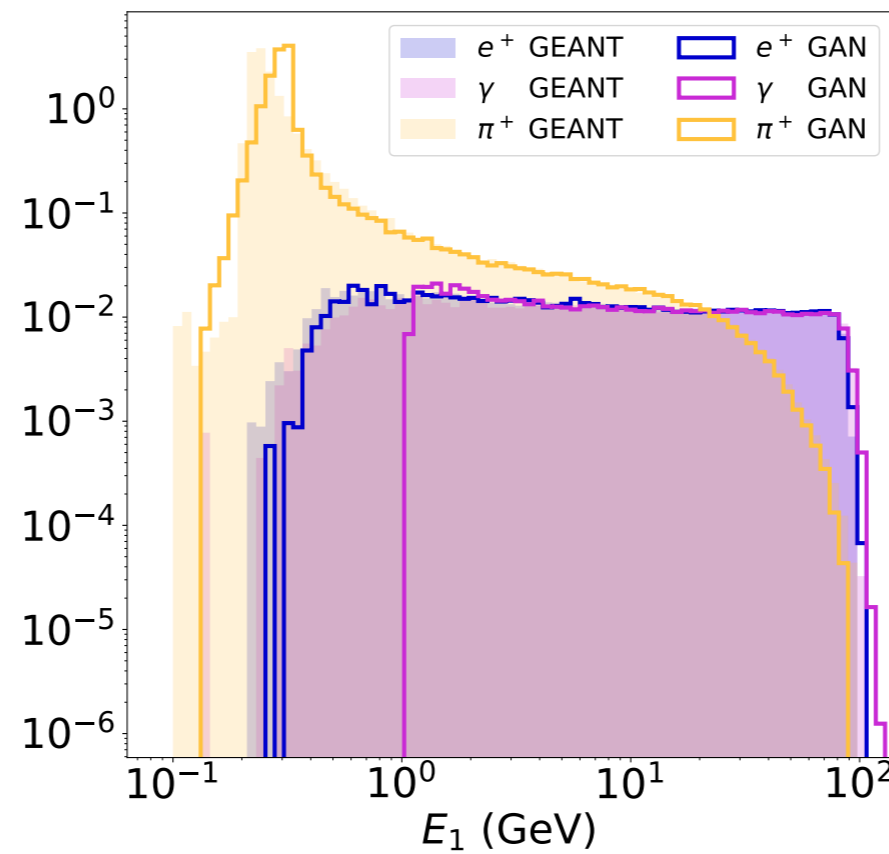
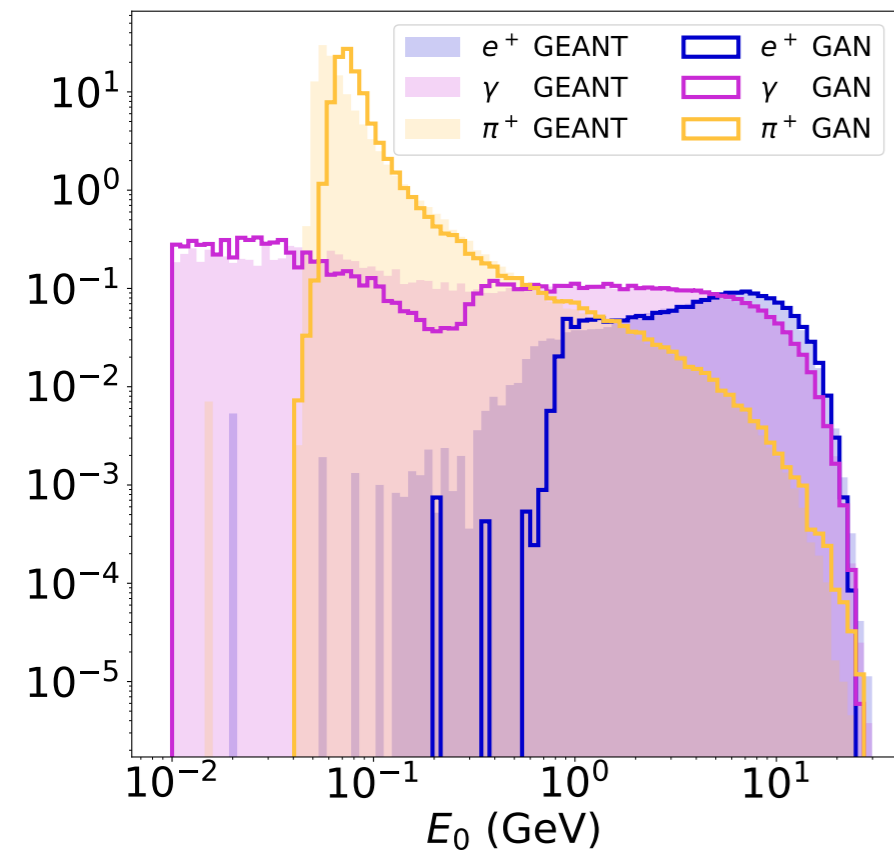


## CaloGAN



# Performance: energy per layer

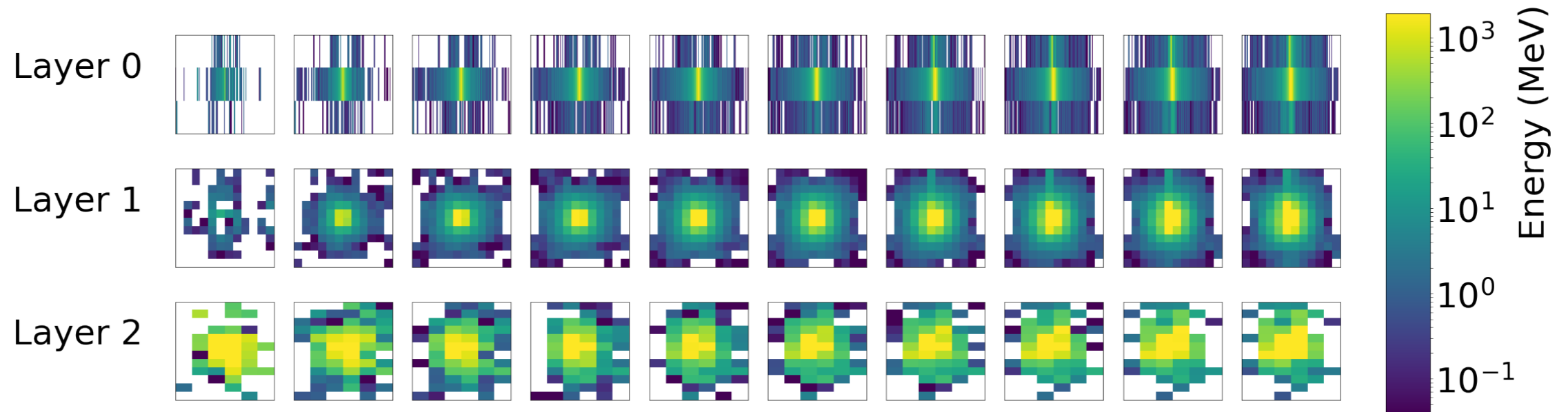
Pions deposit much less energy in the first layers; leave the calorimeter with significant energy



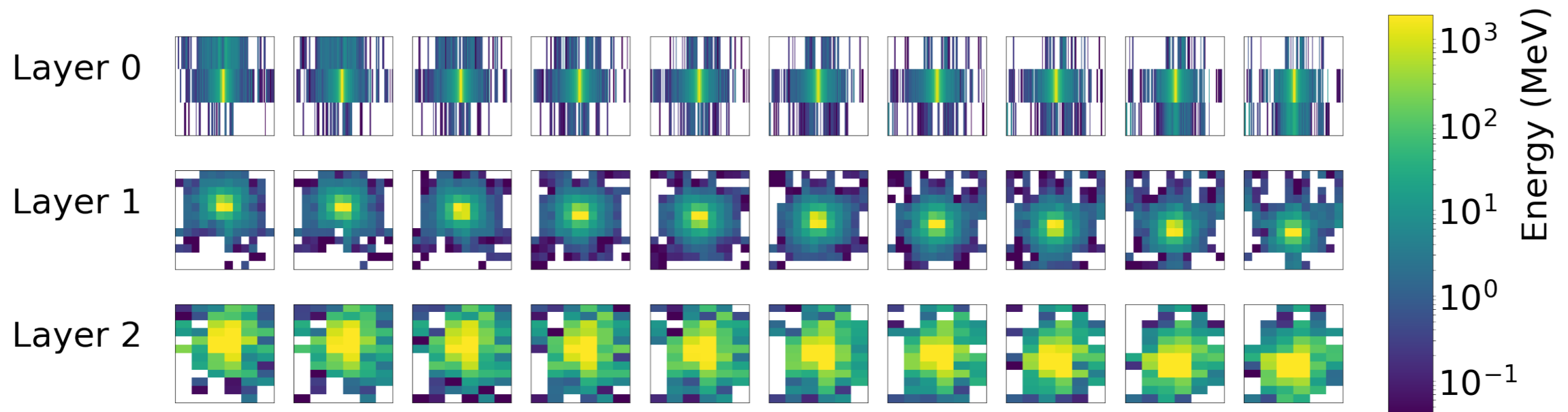


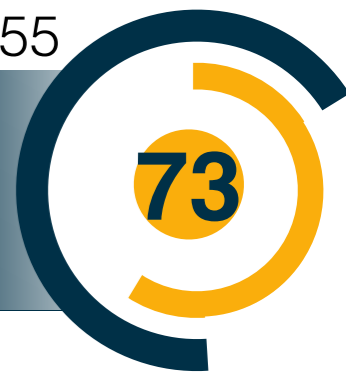
# Conditioning

Fix noise, scan latent variable corresponding to energy



Fix noise, scan latent variable corresponding to x-position



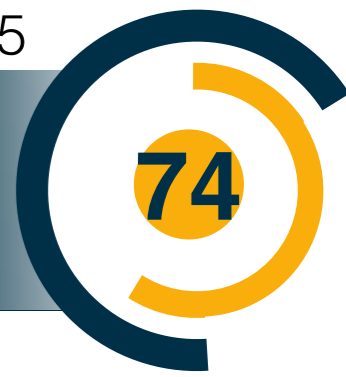


# Timing

Generation Method	Hardware	Batch Size	milliseconds/shower
GEANT4	CPU	N/A	1772 ←
CALOGAN	CPU <i>Intel Xeon E5-2670</i>	1	13.1
		10	5.11
		128	2.19
		1024	2.03
	GPU <i>NVIDIA K80</i>	1	14.5
		4	3.68
		128	0.021
		512	0.014
		1024	0.012 ←

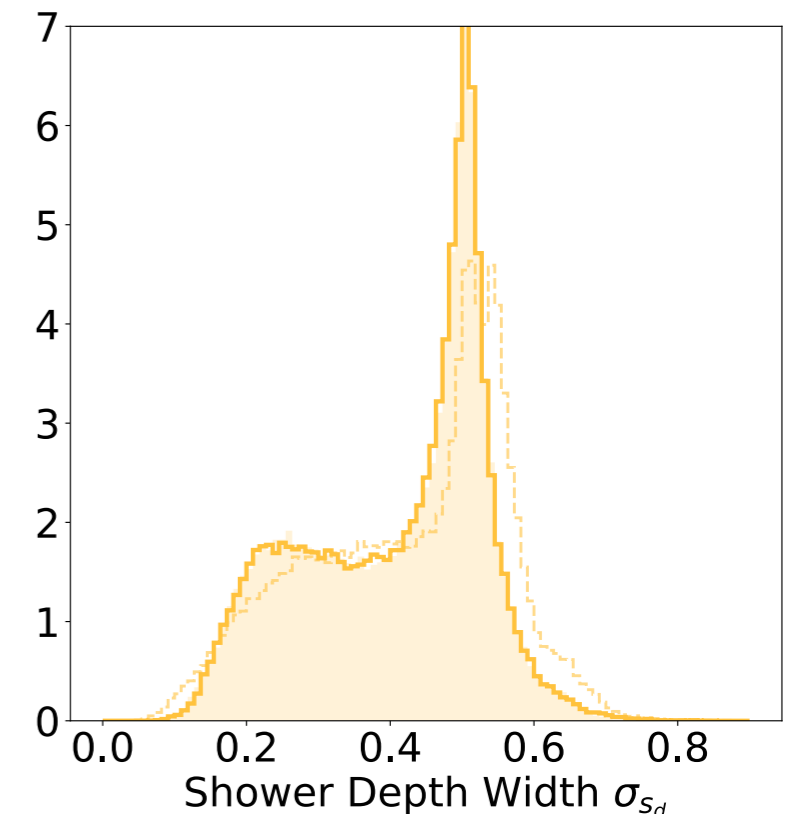
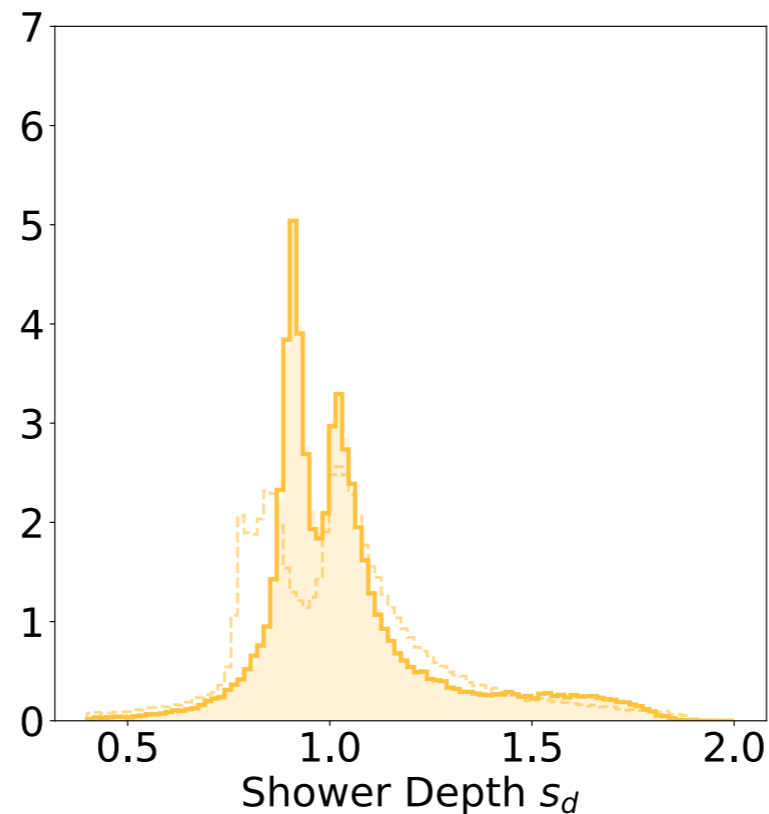
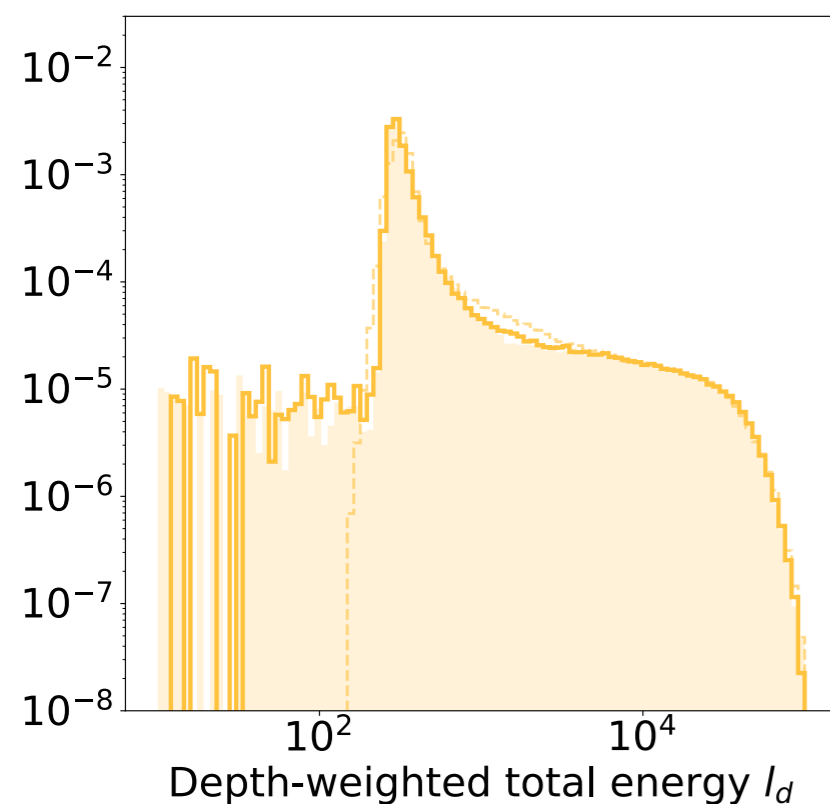
(clearly these numbers have changed as both technologies have improved - this is simply meant to be qualitative & motivating!)





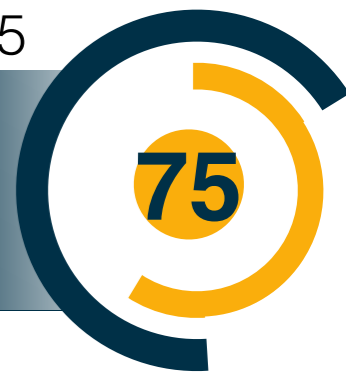
# Current State of the art

Generative models have gotten much better; **flow models** are particularly promising. Added bonus: have an explicit density.



$\pi^+$  GEANT
 
 $\pi^+$  CaloGAN
 
 $\pi^+$  CaloFlow

many other papers - see Living Review, 2102.02770



# Current State of the art

Generative models have gotten much better: **flow models** are

AUC / JSD		DNN	
		vs. CALOGAN	vs. CALOFlow
$e^+$	unnormalized	1.000(0) / 0.993(1)	0.847(8) / 0.345(12)
	normalized	1.000(0) / 0.997(0)	0.869(2) / 0.376(4)
$\gamma$	unnormalized	1.000(0) / 0.996(1)	0.660(6) / 0.067(4)
	normalized	1.000(0) / 0.994(1)	0.794(4) / 0.213(7)
$\pi^+$	unnormalized	1.000(0) / 0.988(1)	0.632(2) / 0.048(1)
	normalized	1.000(0) / 0.997(0)	0.751(4) / 0.148(4)

Output is nearly indistinguishable from Geant4 !

*AUC = 1 means easily distinguishable, AUC = 0.5 means not distinguishable*

Depth-weighted total energy  $l_d$

Shower Depth  $s_d$

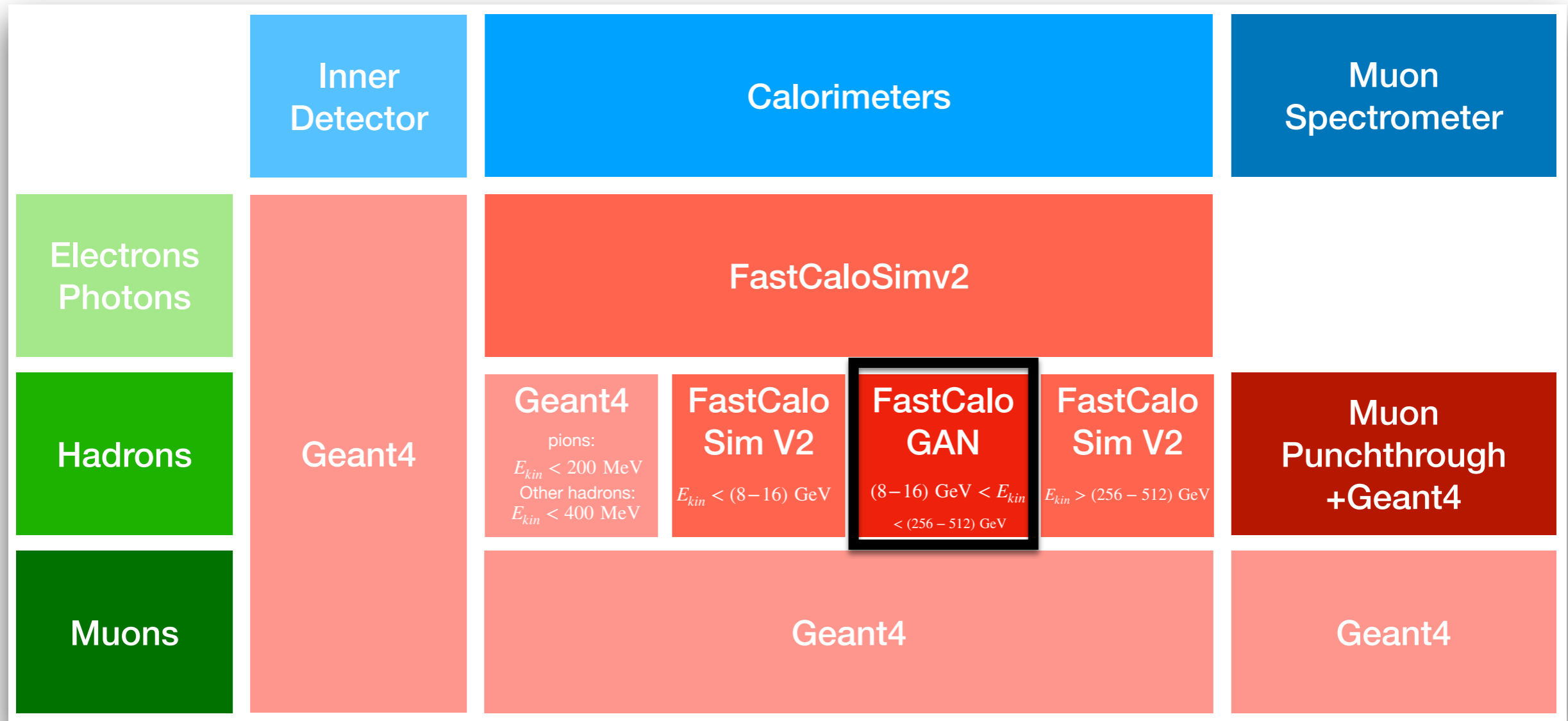
Shower Depth Width  $\sigma_{s_d}$

   $\pi^+$  GEANT    
    $\pi^+$  CaloGAN    
    $\pi^+$  CaloFlow

many other papers - see Living Review, 2102.02770



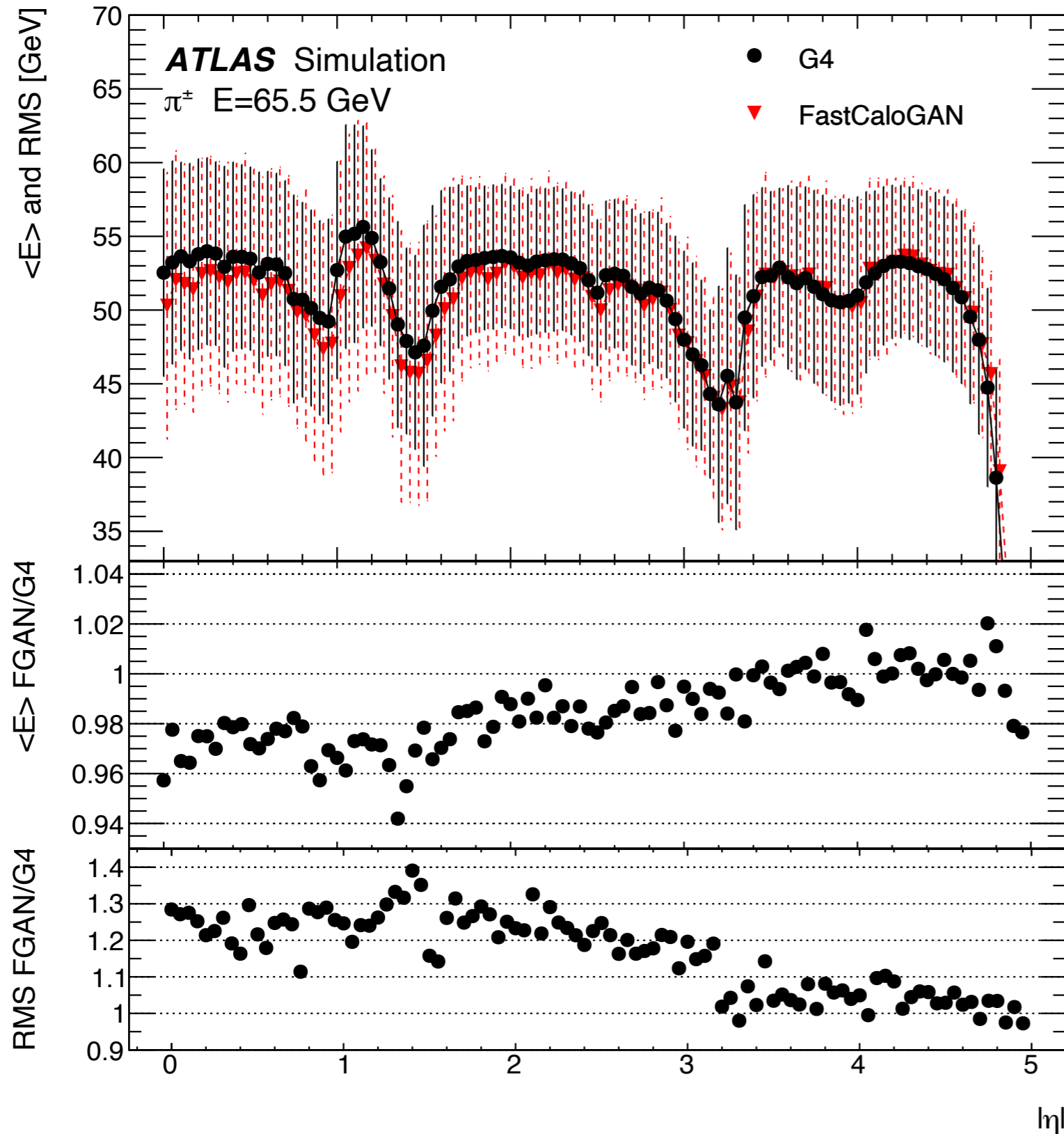
# Integration into real detector sim.



The ATLAS Collaboration fast simulation (AF3) now includes a GAN at intermediate energies for pions



# Integration into real detector sim.

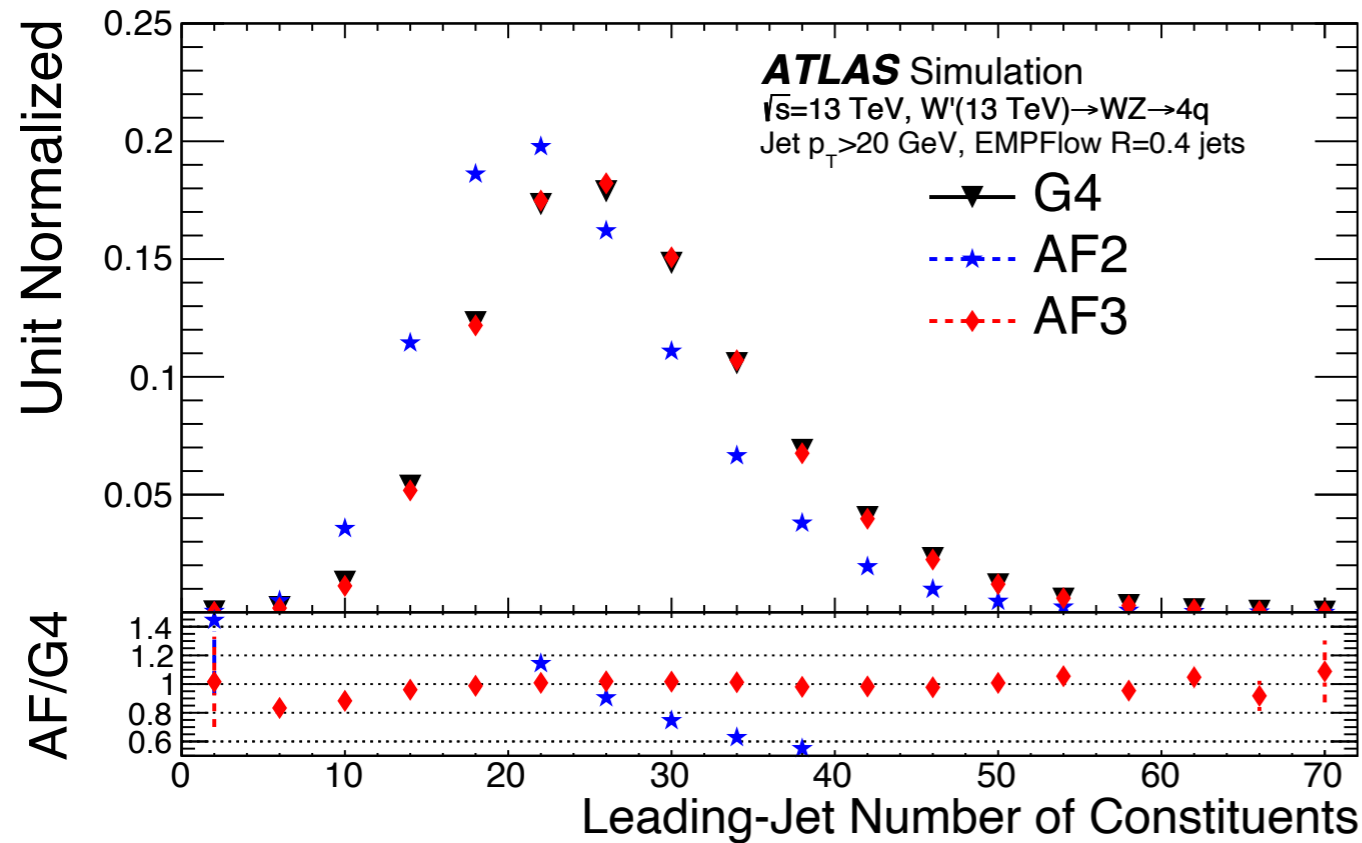


The GAN architecture is relatively simple, but it is able to match the energy scale and resolution well.

There is one GAN per  $\eta$  slice

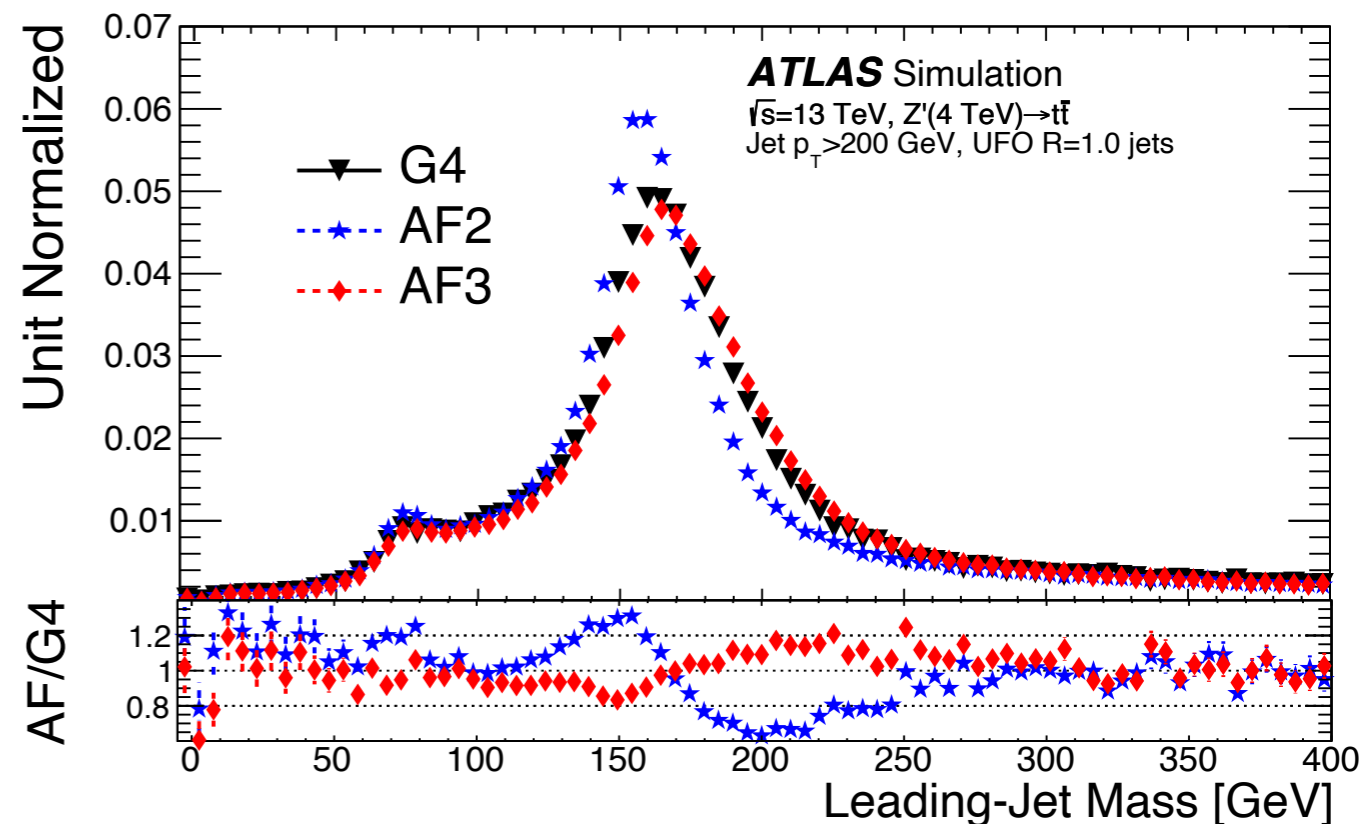


# Integration into real detector sim.



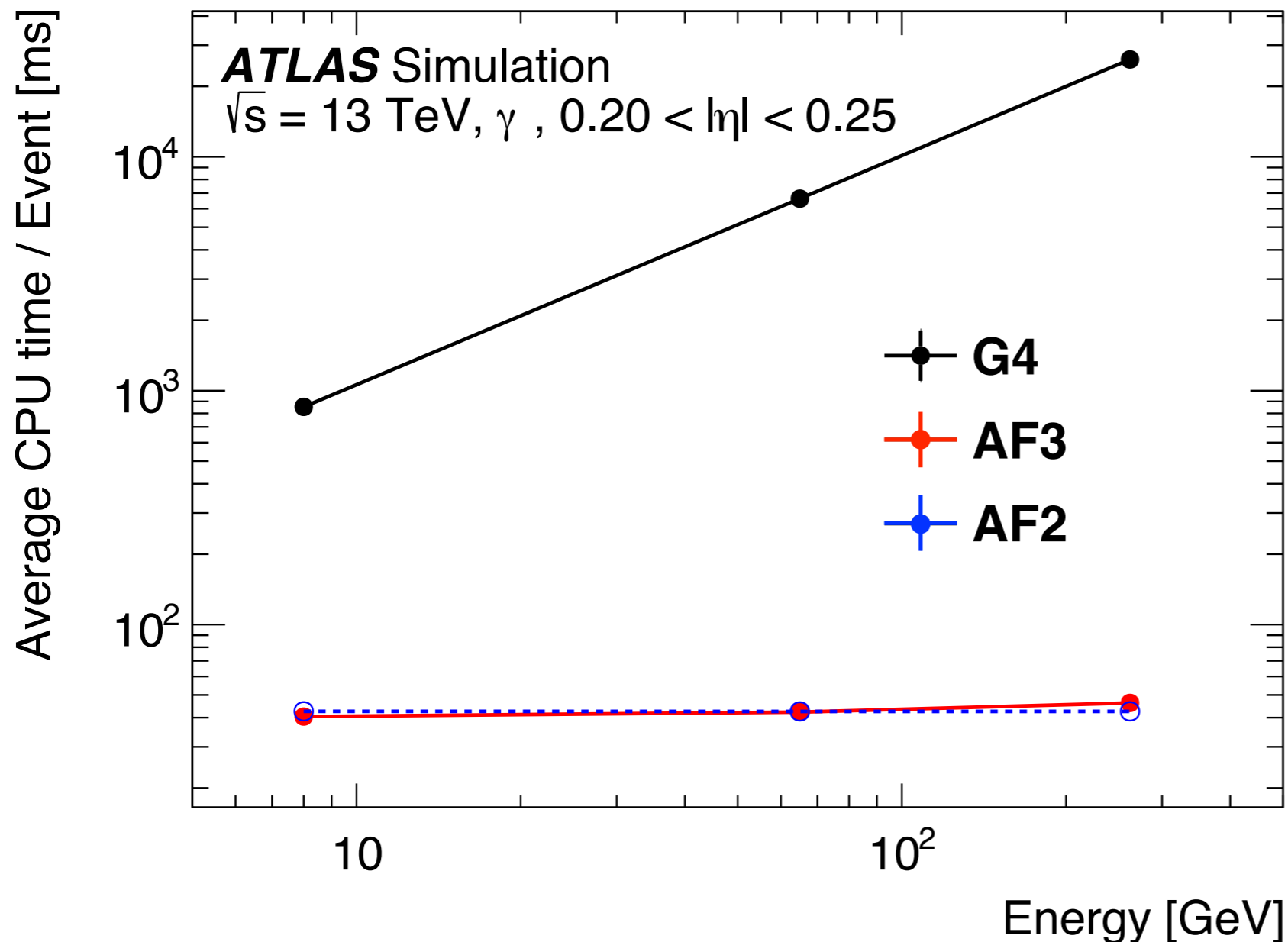
The new fast simulation (**AF3**) significantly improves jet substructure with respect to the older one (**AF2**)

Ideally, the same calibrations derived for full sim. (Geant4-based) can be applied to the fast sim.





# Integration into real detector sim.

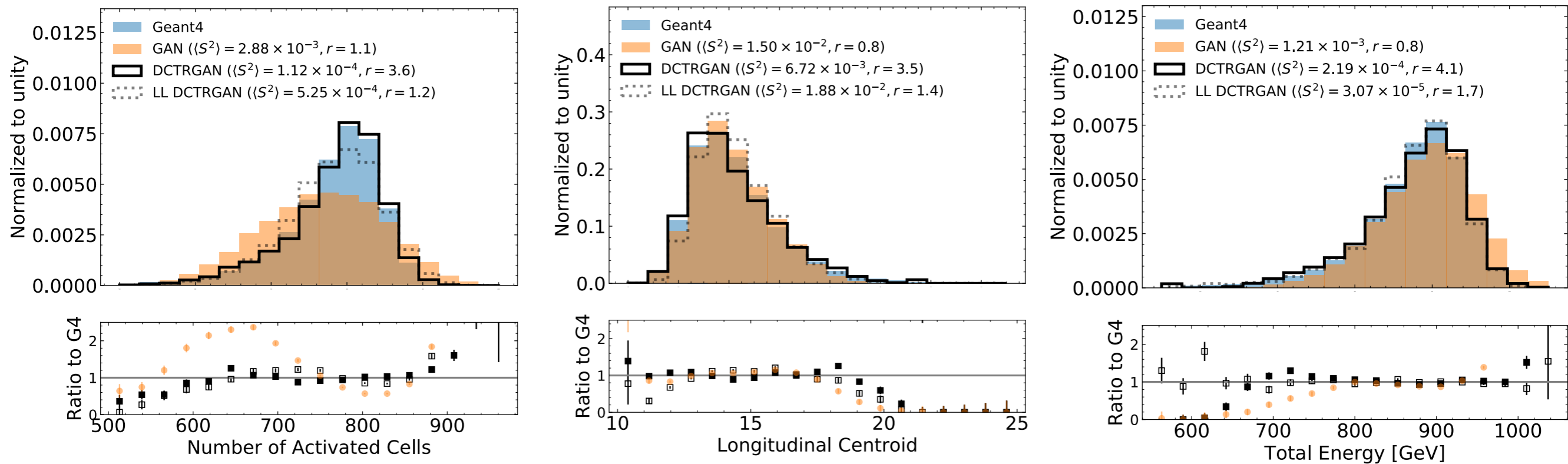


As expected, the fast sim. timing is independent of energy, while Geant4 requires more time for higher energy.



# Refining Simulations

As we move towards precision, we may need to complement primary generative models with post-hoc correction models (e.g. via reweighting)



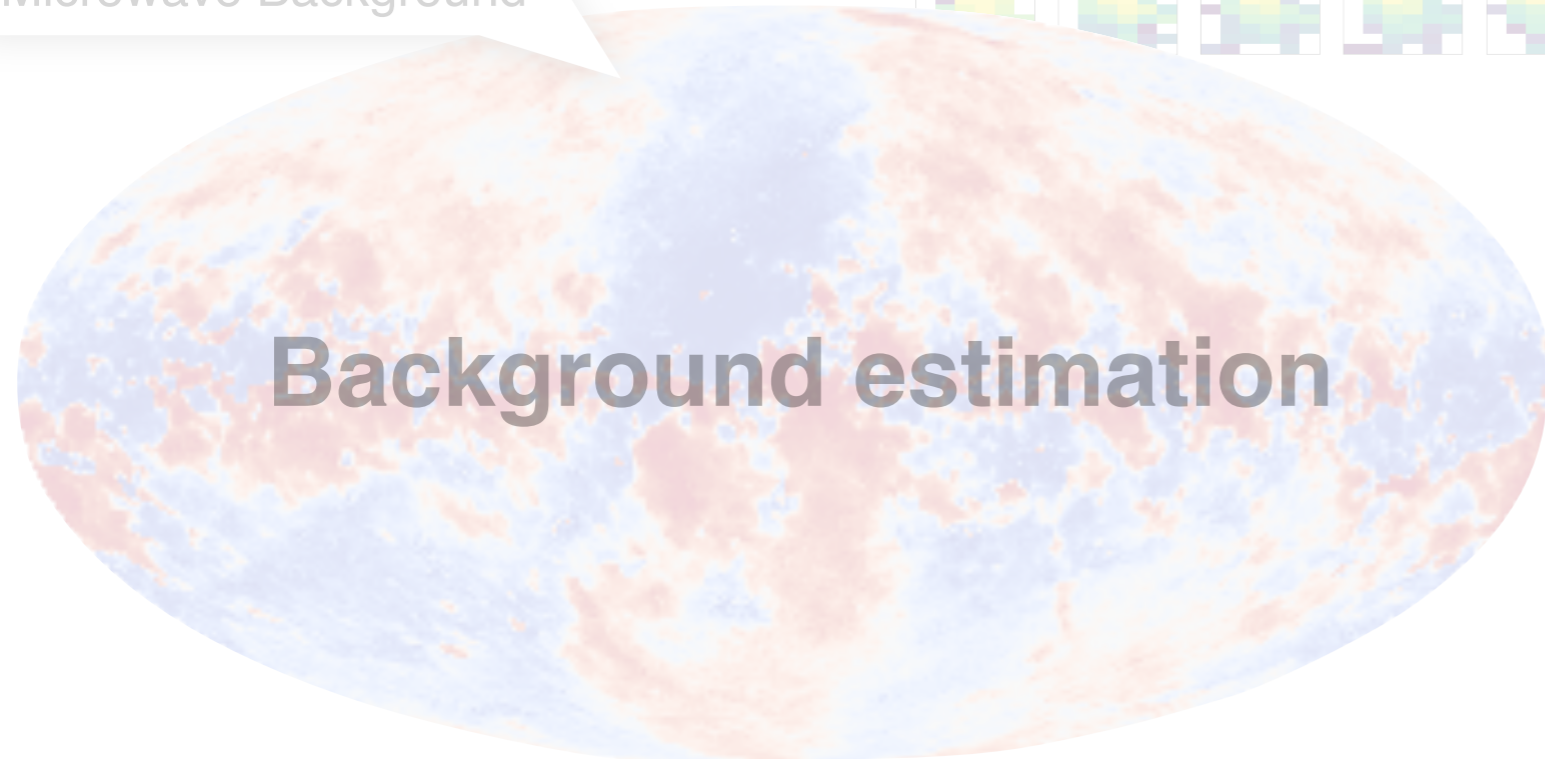
See also 2106.00792 (“LaSeR”) and 2107.08648 (optimal transport-based)



# Generative Models for Particle/Nuclear/Astro

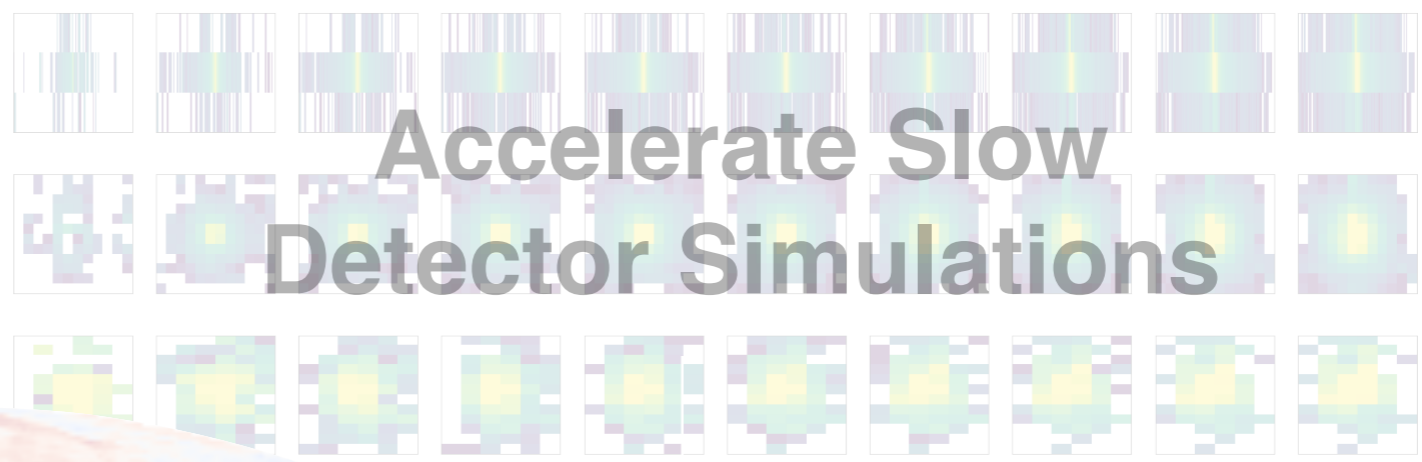
All of these pictures are fake!

Synthetic Galactic radiation for Cosmic Microwave Background



Background estimation

Material Interactions with High Energy Particles

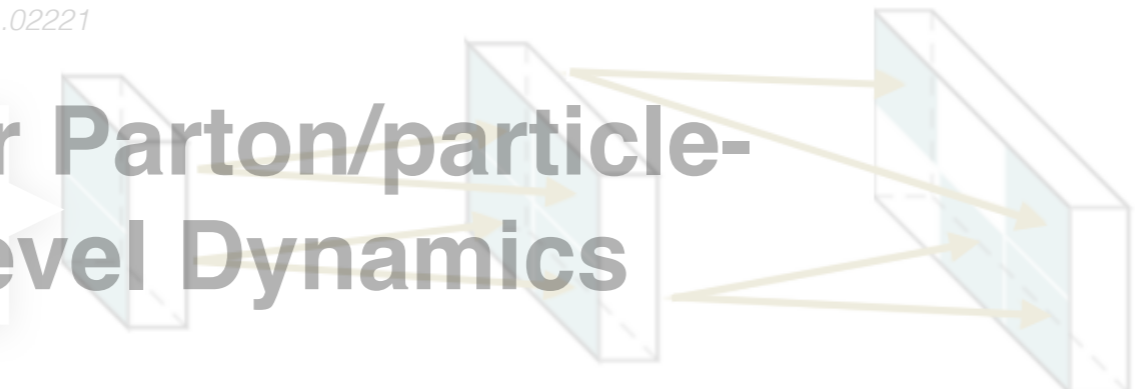


Accelerate Slow Detector Simulations

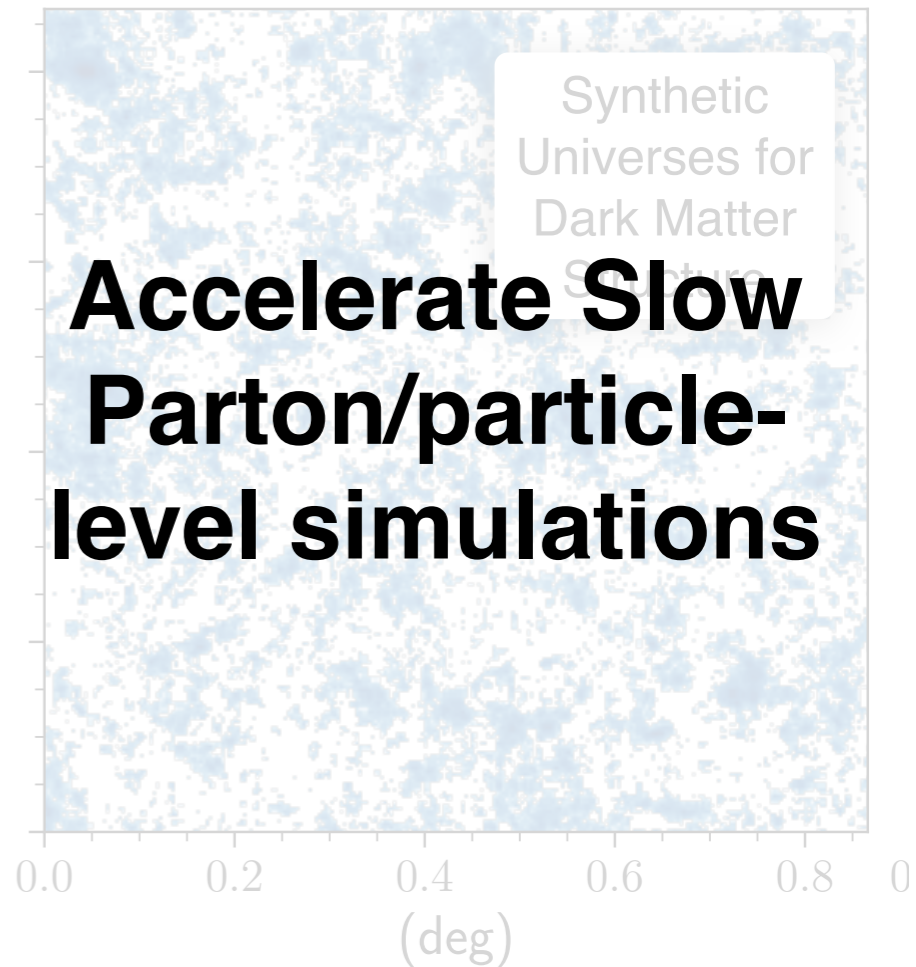
M. Paganini, L. De Oliveira, B. Nachman, Phys. Rev. Lett. 120 (2018) 042003

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M. Mustafa, et al., Comp. Astrophysics and Cosmology 6 (2019)

## Flat jet images with GANs

M. Paganini, L. de Oliveira, B. Nachman

1701.05927

\*these are just representative examples - see Living Review, 2102.02770

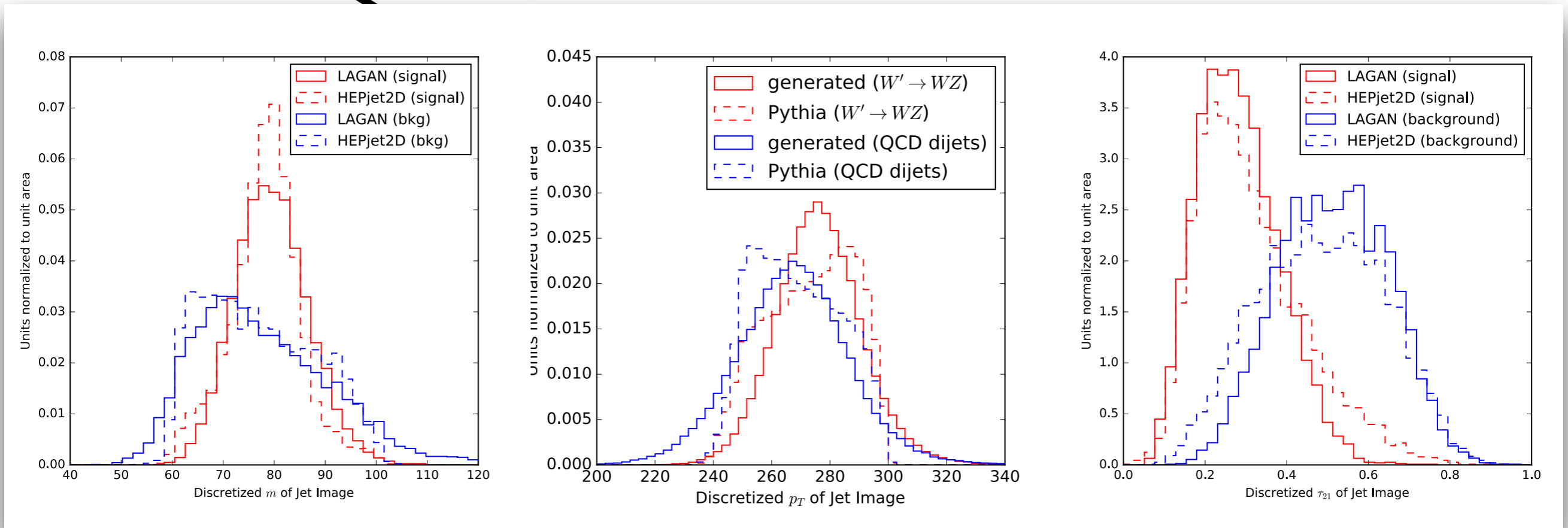
# Accelerating Parton/Particle Sim.\*

83

## Flat jet images with GANs

M. Paganini, L. de Oliveira, B. Nachman

1701.05927



LA = Locally aware; somewhere between a DNN and a CNN

**Weight sharing across space**

\*these are just representative examples - see Living Review, 2102.02770

# Accelerating Parton/Particle

Flat jet images with GANs

M. Paganini, L. de Oliveira, B. Nachman

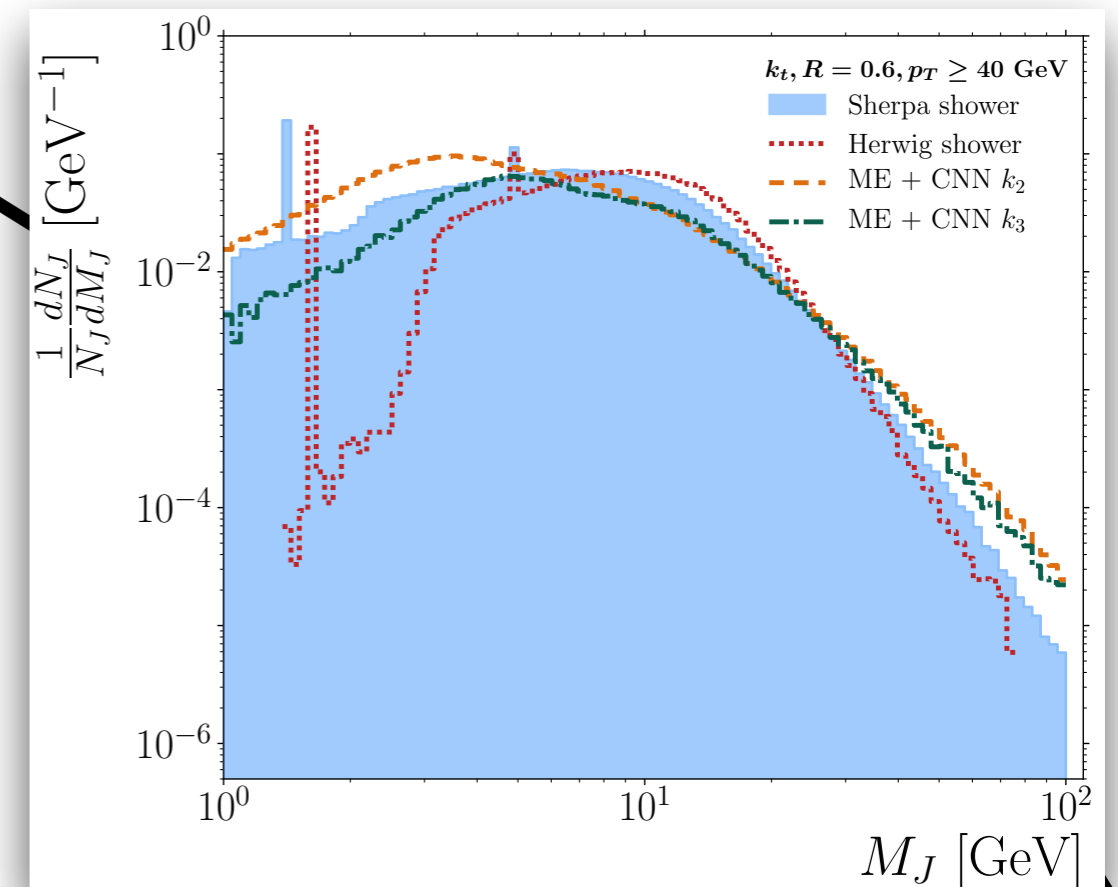
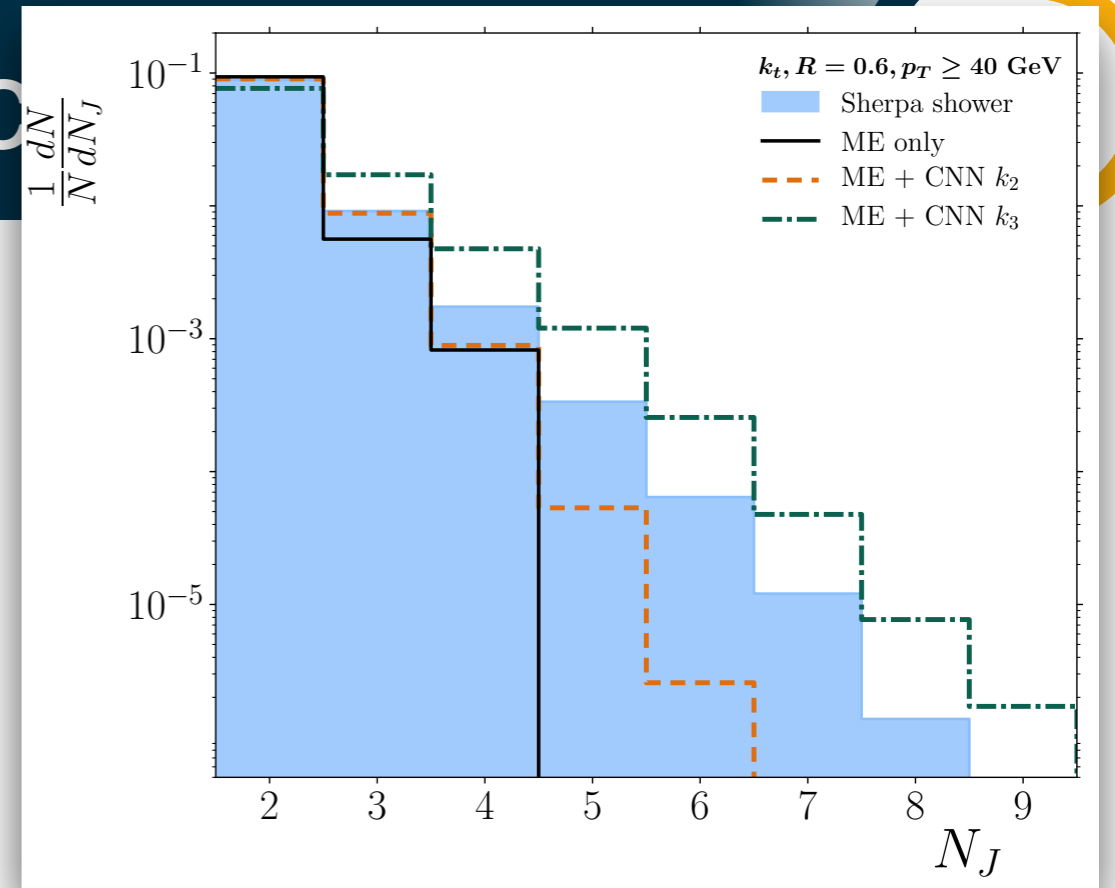
1701.05927

Scale invariant  
images with AEs

J. Monk

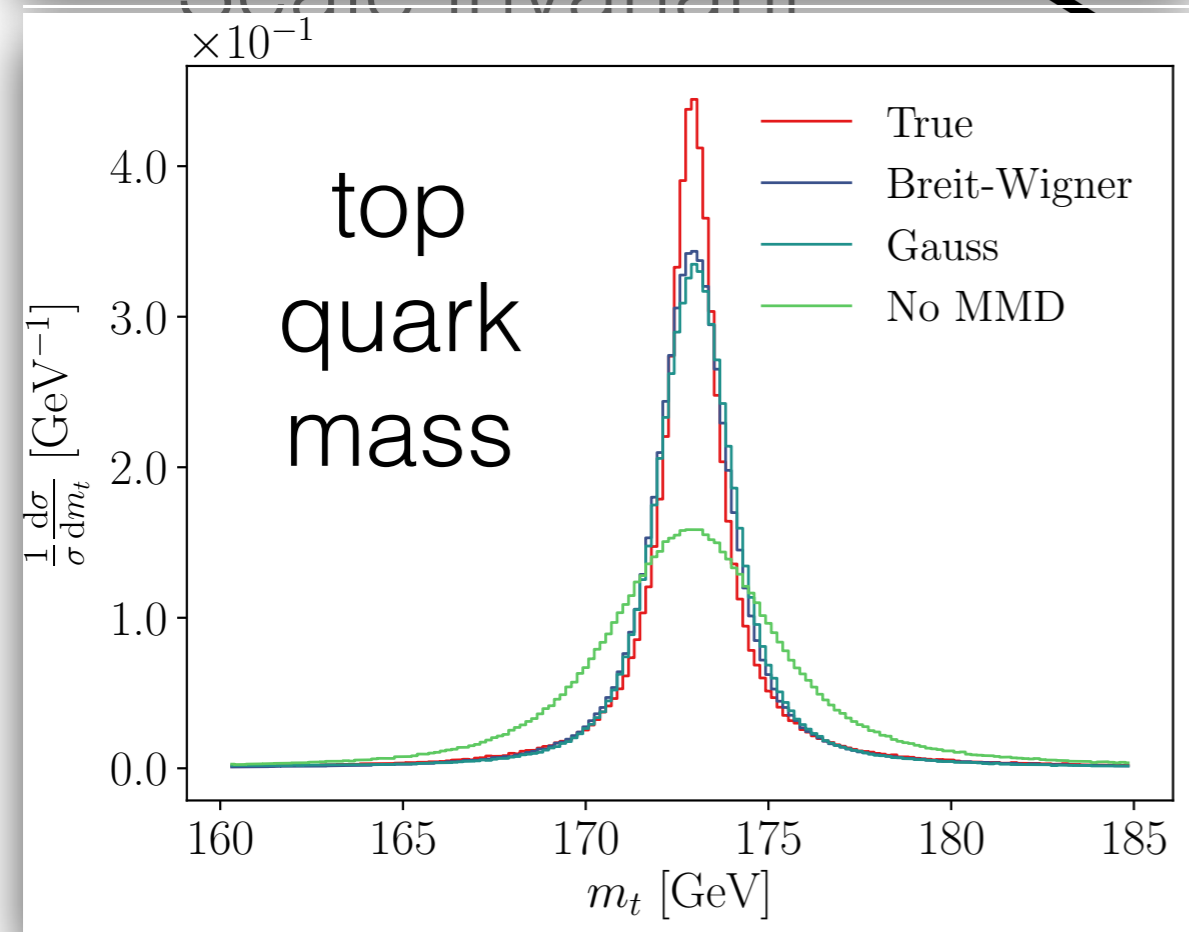
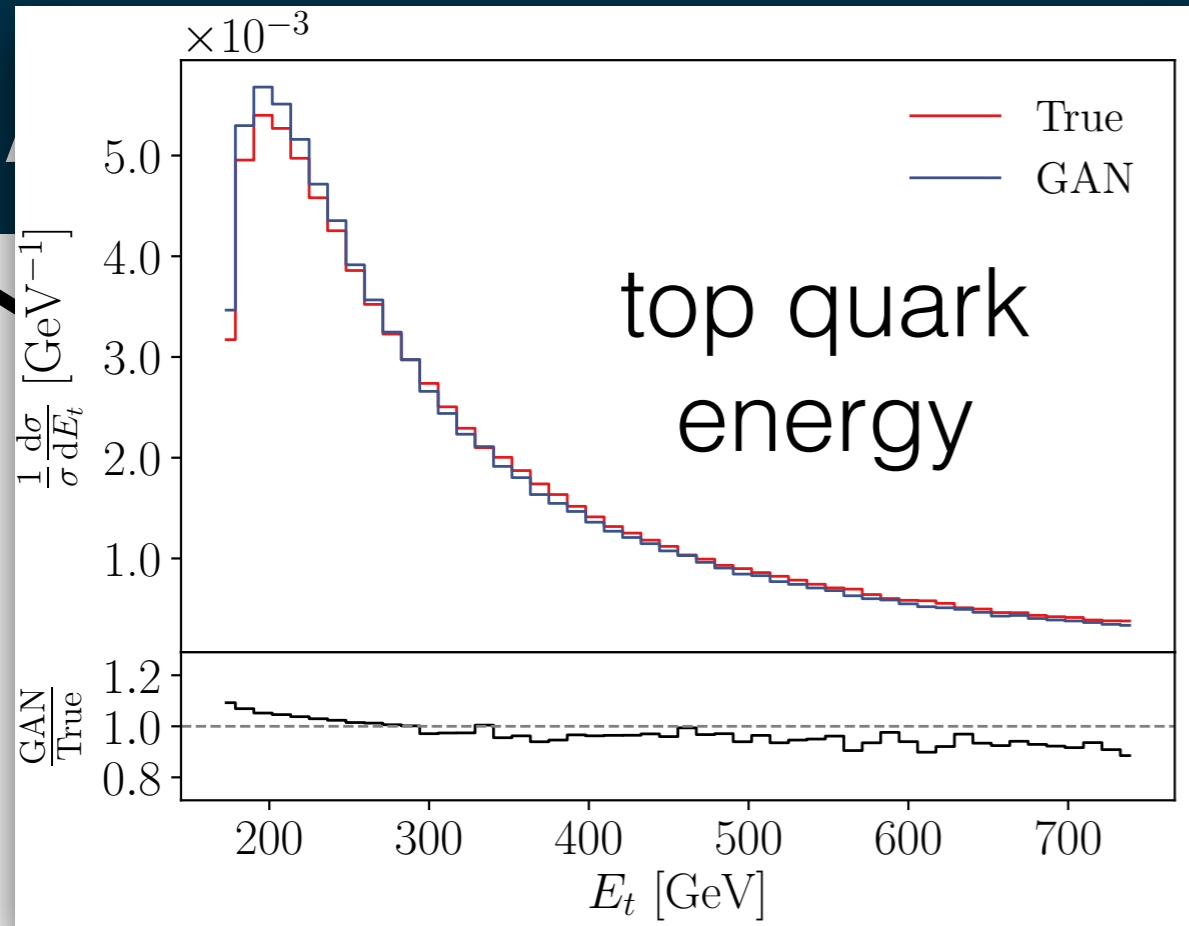
1807.03685

**Weight sharing across space + “time”**



\*these are just representative examples - see Living Review, 2102.02770

# Particle Sim.\*



MMD = maximum mean discrepancy

Fixed number of 4-vectors, allow for intermediate resonances

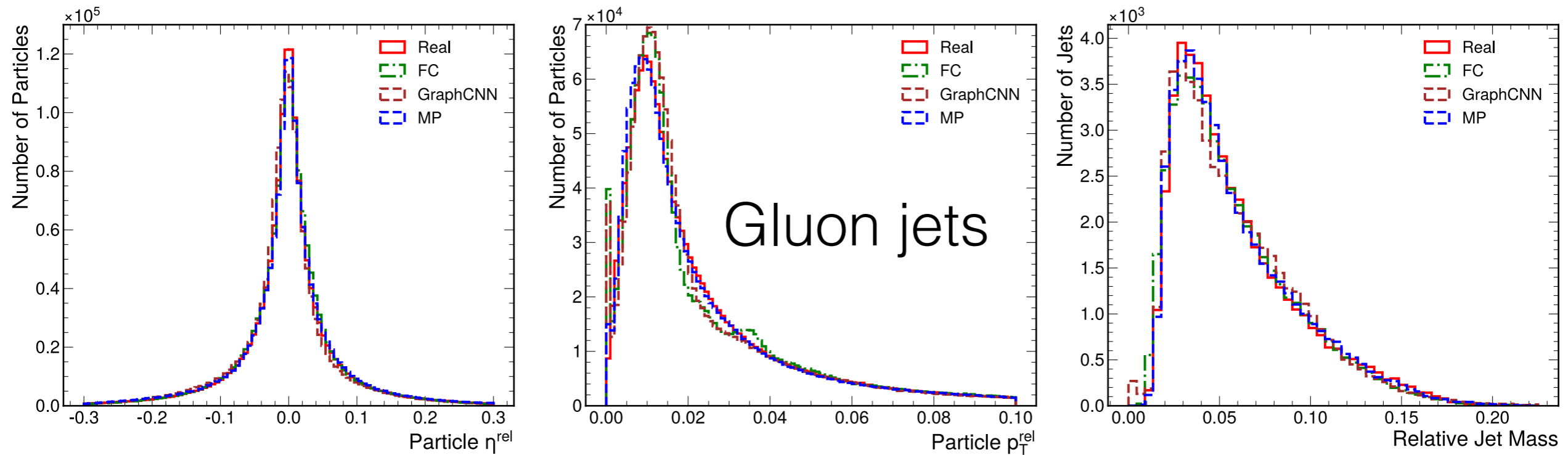
A. Butter, T. Plehn, R. Winterhalder

1907.03764

\*these are just representative examples - see Living Review, 2102.02770

## Flat jet images with GANs

M. Paganini, L. de Oliveira, B. Nachman



Variable-length  
output with graphs

R. Kansal et al.

2106.11535

\*these are just representative examples - see Living Review, 2102.02770

# Accelerating Parton/Particle Sim.\*

87

Flat jet images with GANs

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1701.05927

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Variable-length  
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R. Kansal et al.

2106.11535

?

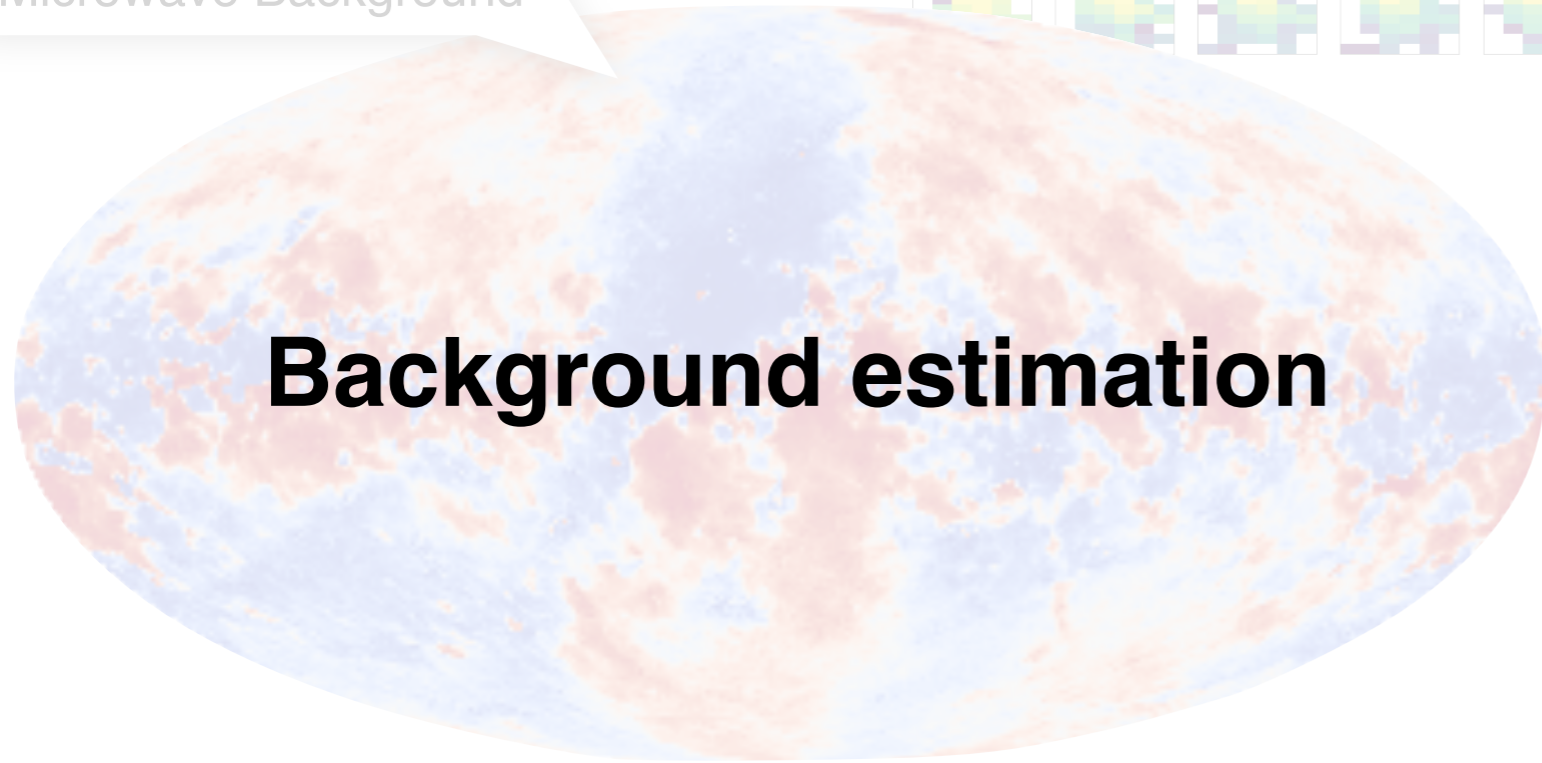
\*these are just representative examples - see Living Review, 2102.02770



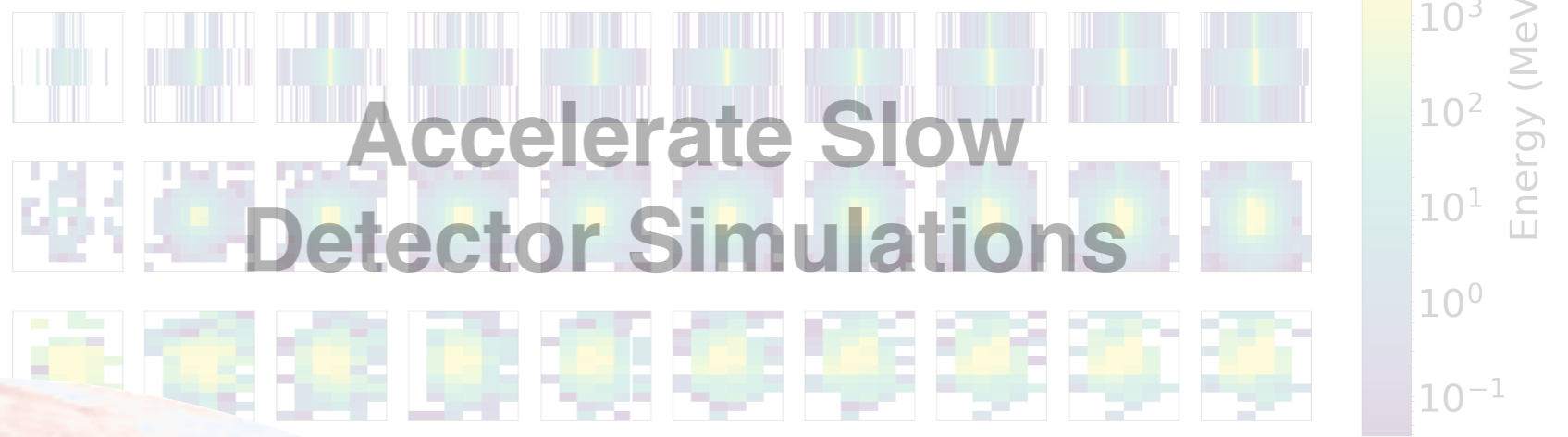
# Generative Models for Particle/Nuclear/Astro

All of these pictures are fake!

Synthetic Galactic radiation for Cosmic Microwave Background



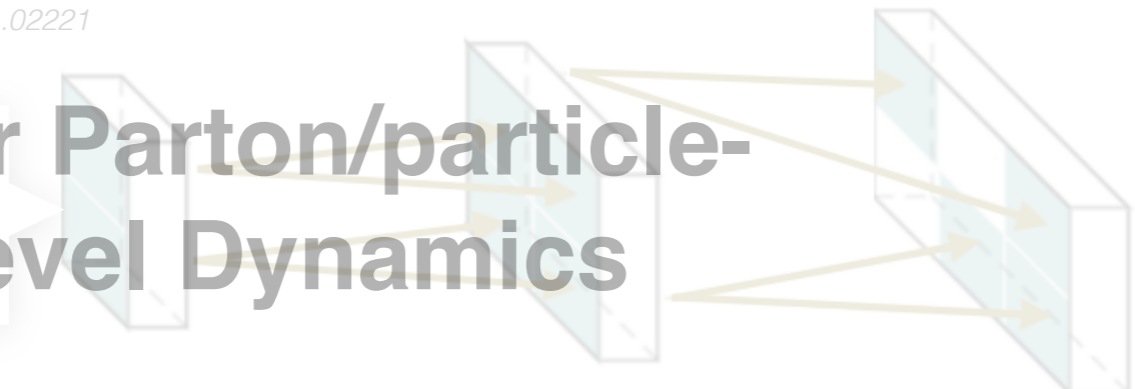
Material Interactions with High Energy Particles



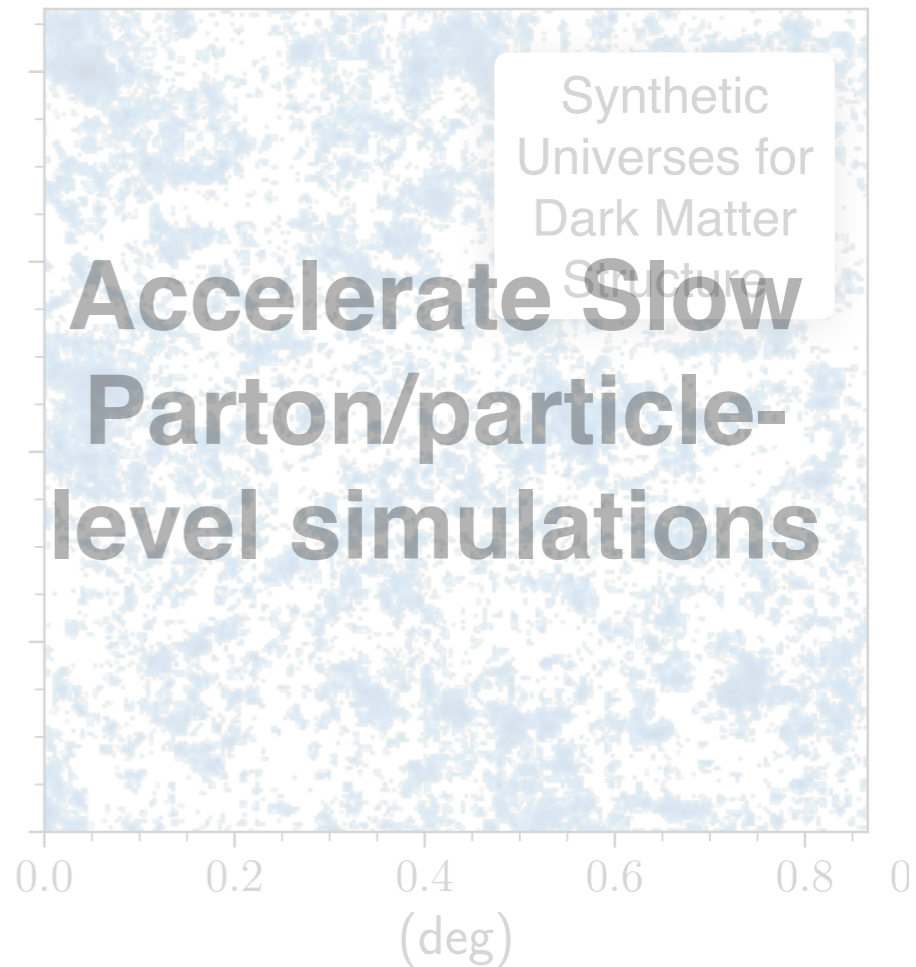
M. Paganini, L. De Oliveira, B. Nachman, Phys. Rev. Lett. 120 (2018) 042003

The Structure of Radiation in the Quantum Strong Force

**Infer Parton/particle-level Dynamics**



Y. S. Lai, D. Neill, M. Płoskoń, F. Ringer, arXiv:2012.06582



M. Mustafa, et al., Comp. Astrophysics and Cosmology 6 (2019)

# Background Estimation

89

Many cases where simulation is **not good enough** - can train generate models to act as **simulation surrogates**

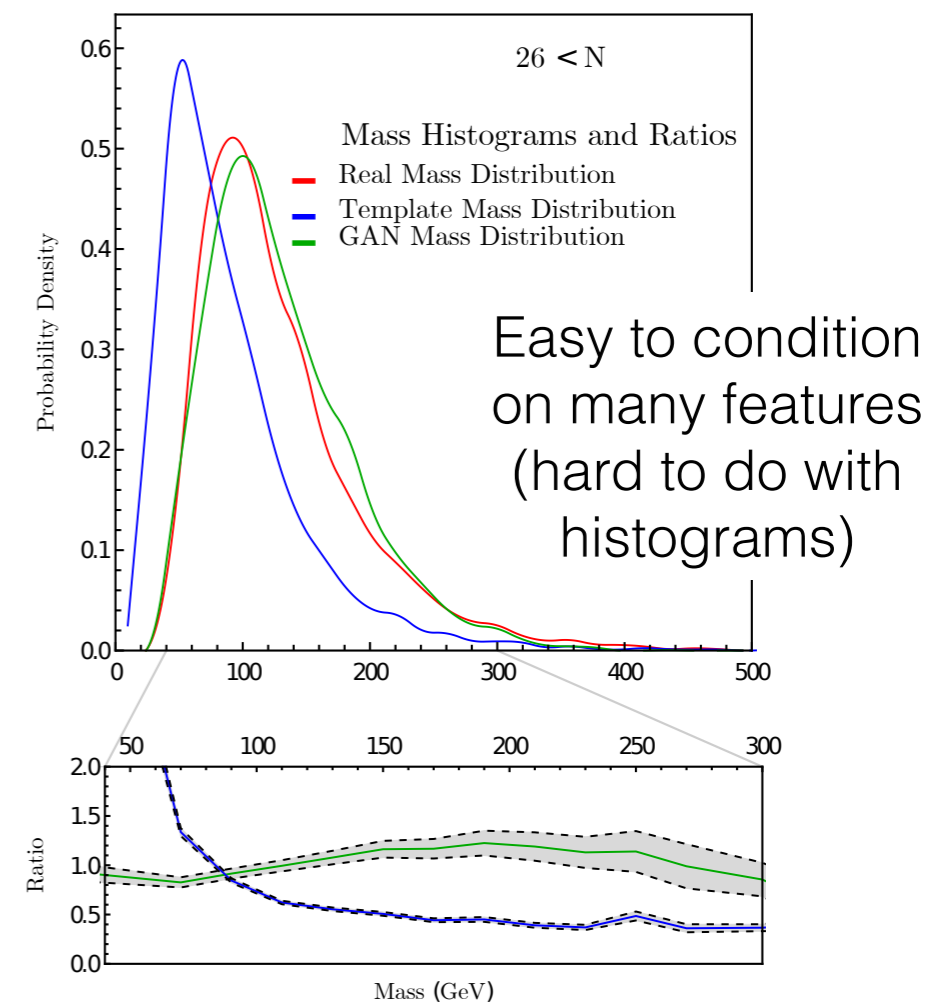
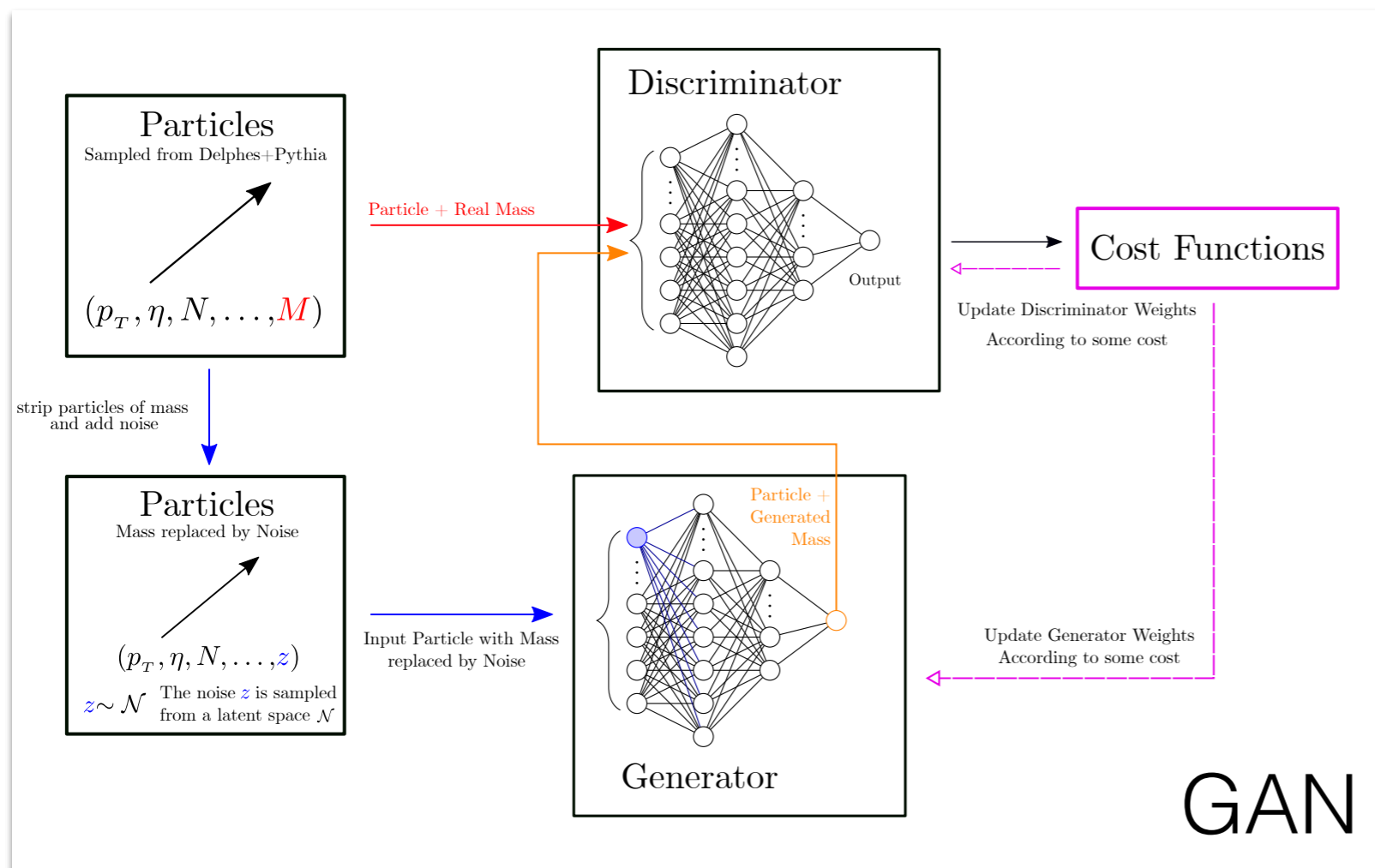
N.B. everything in I've shown before this,  
we trained on simulation, not on data (!)



# Background Estimation

Many cases where simulation is **not good enough** - can train generate models to act as **simulation surrogates**

Example 1: unbinned templates for QCD jets to extrapolate in jet multiplicity

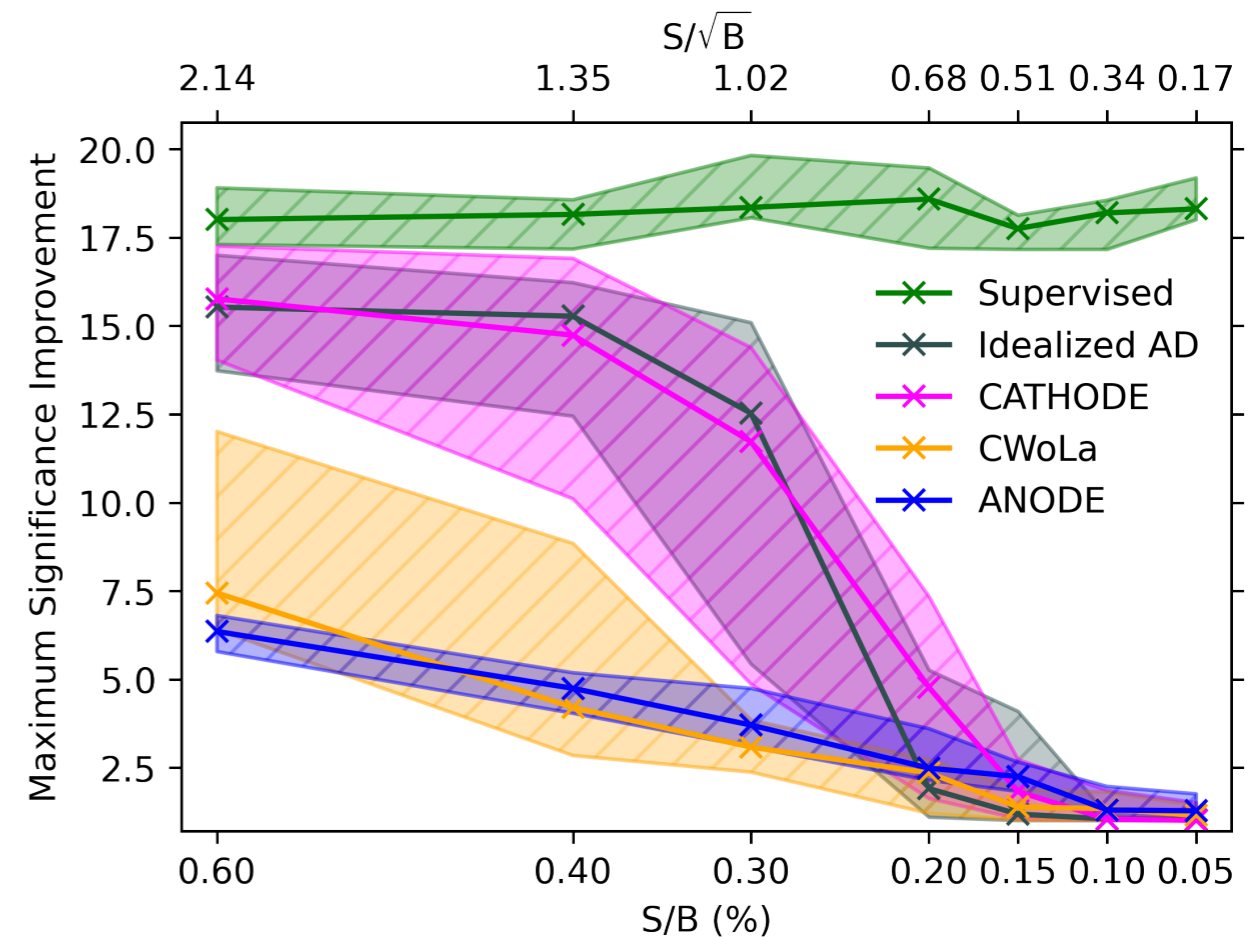
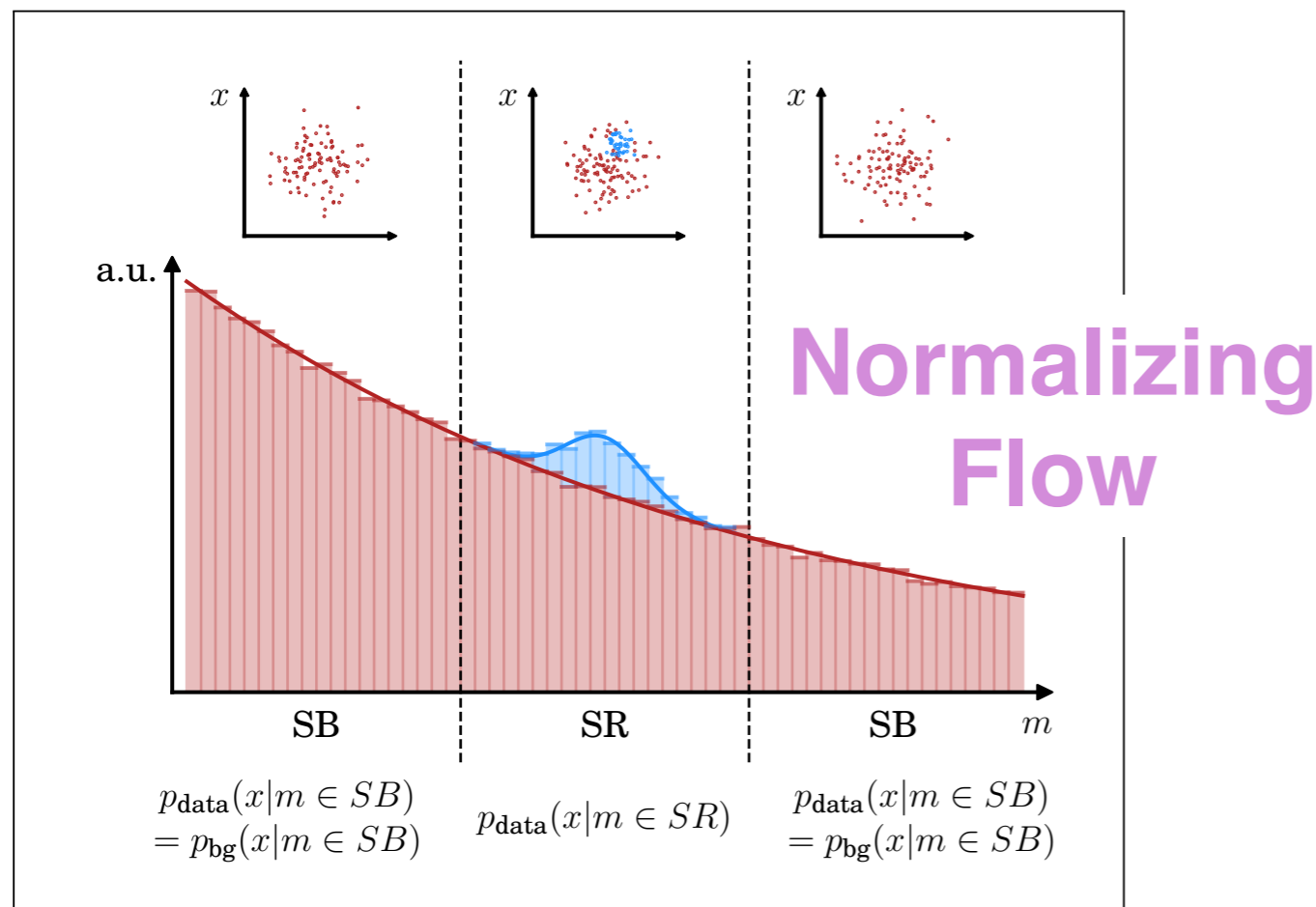




# Background Estimation

Many cases where simulation is **not good enough** - can train generate models to act as **simulation surrogates**

Example 2: unbinned templates for QCD jets to extrapolate in dijet mass

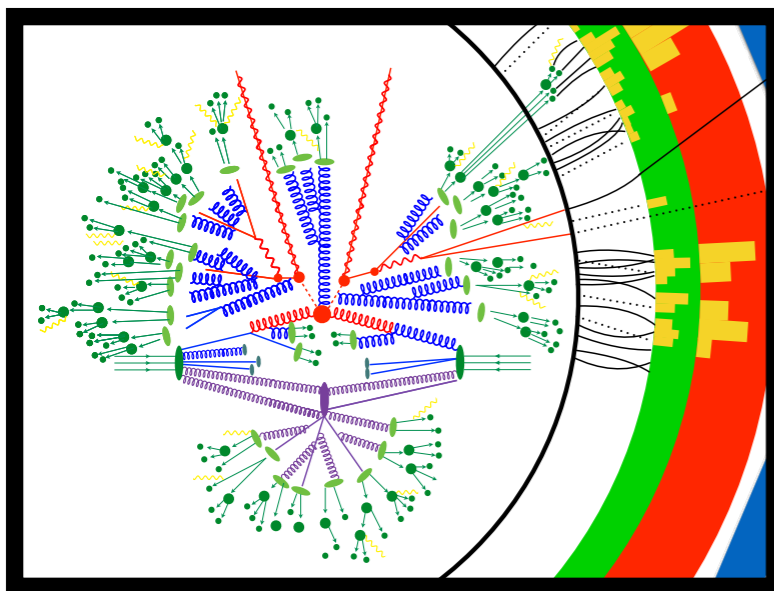


## Anomaly detection

# Related: Data Compression

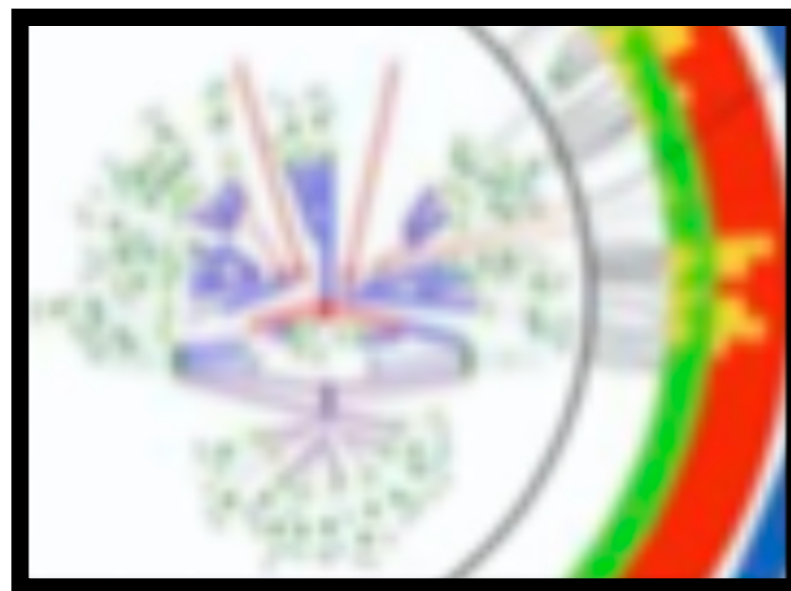
You can think of surrogate models as compressing the data into the parameters of the neural network.

No compression



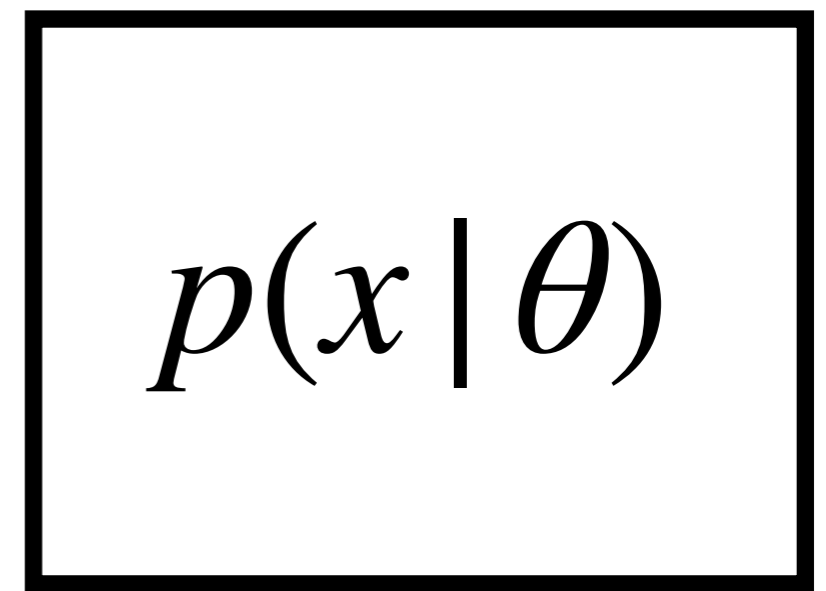
Many numbers per **event**

Compress per event



Small set of numbers per **event**

Compress entire dataset



Small set of numbers per **dataset**

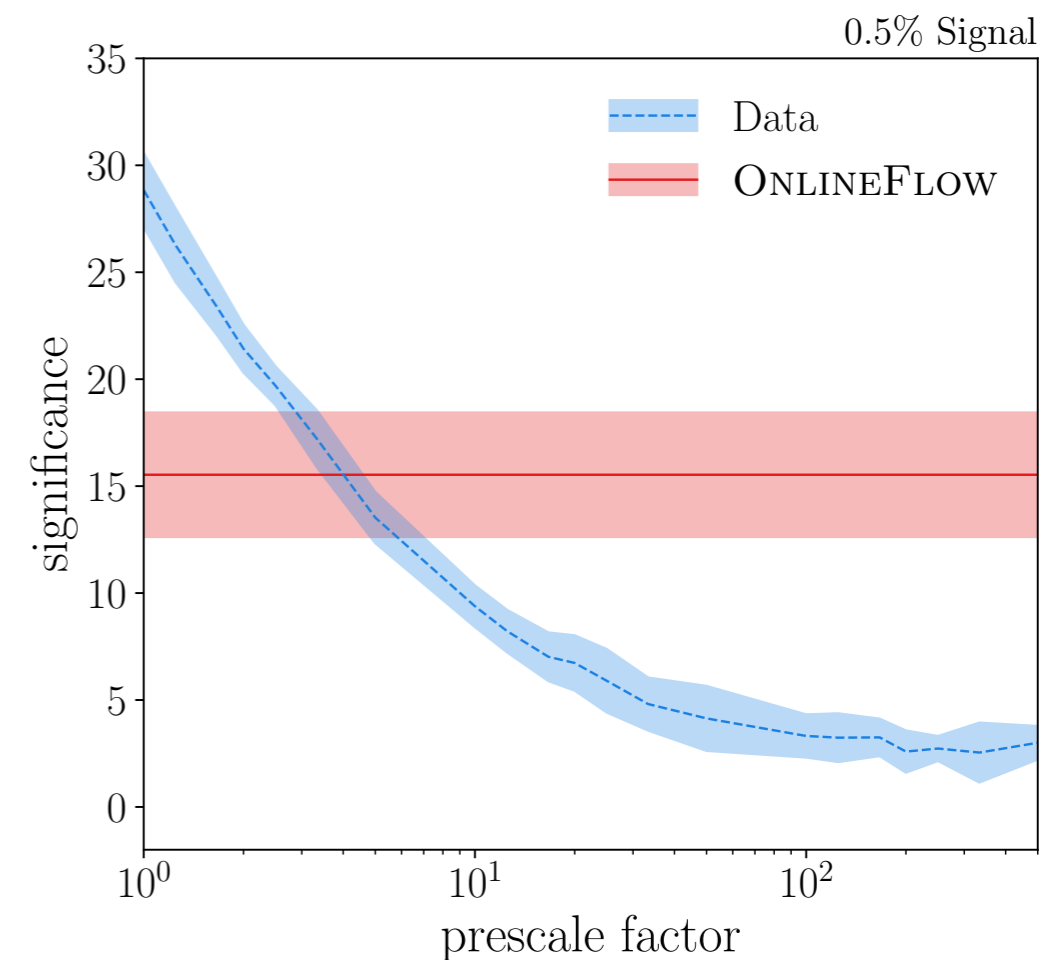
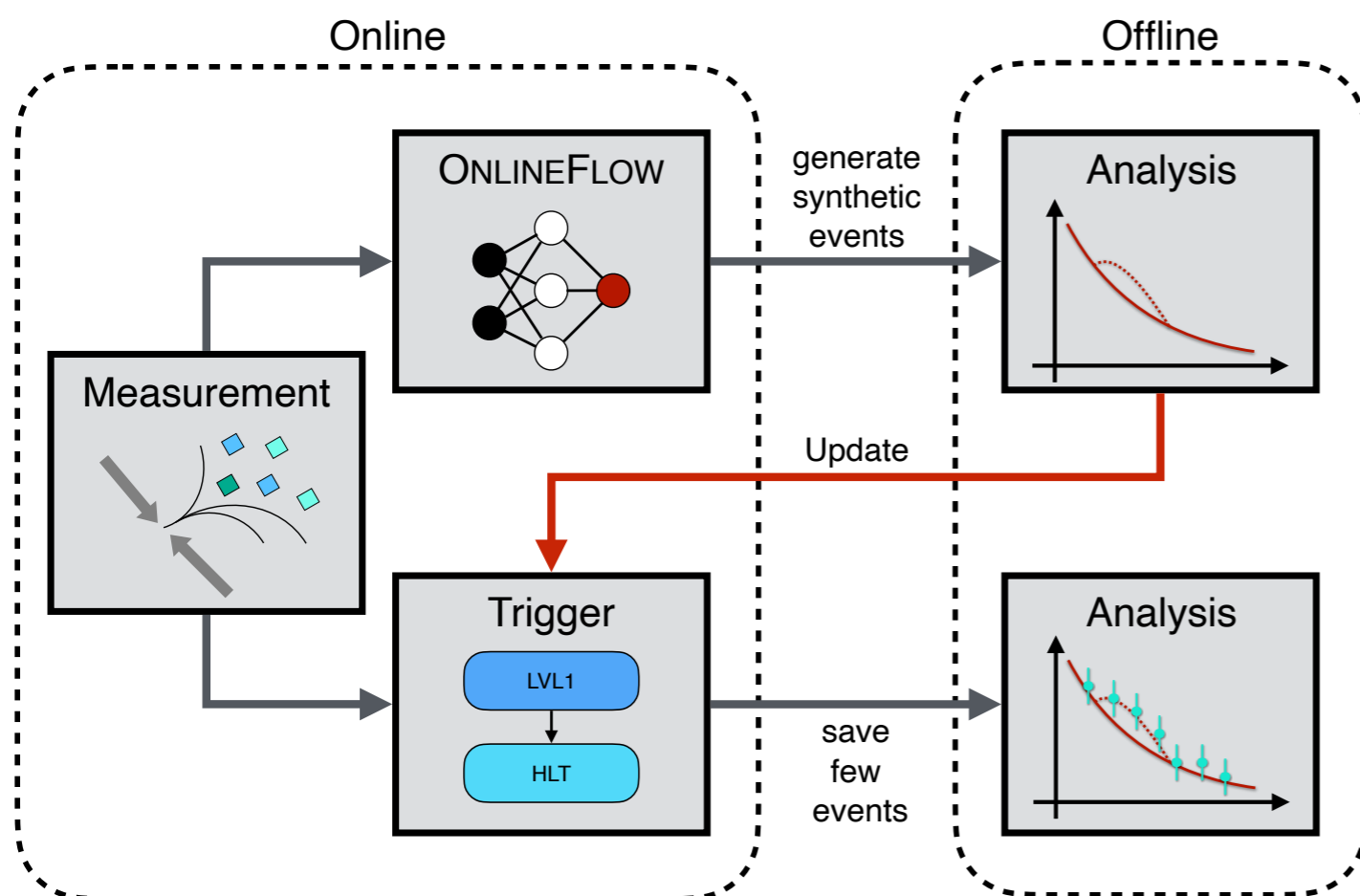




# Related: Data Compression

You can think of surrogate models as compressing the data into the parameters of the neural network.

*Can this also be used for anomaly detection?*

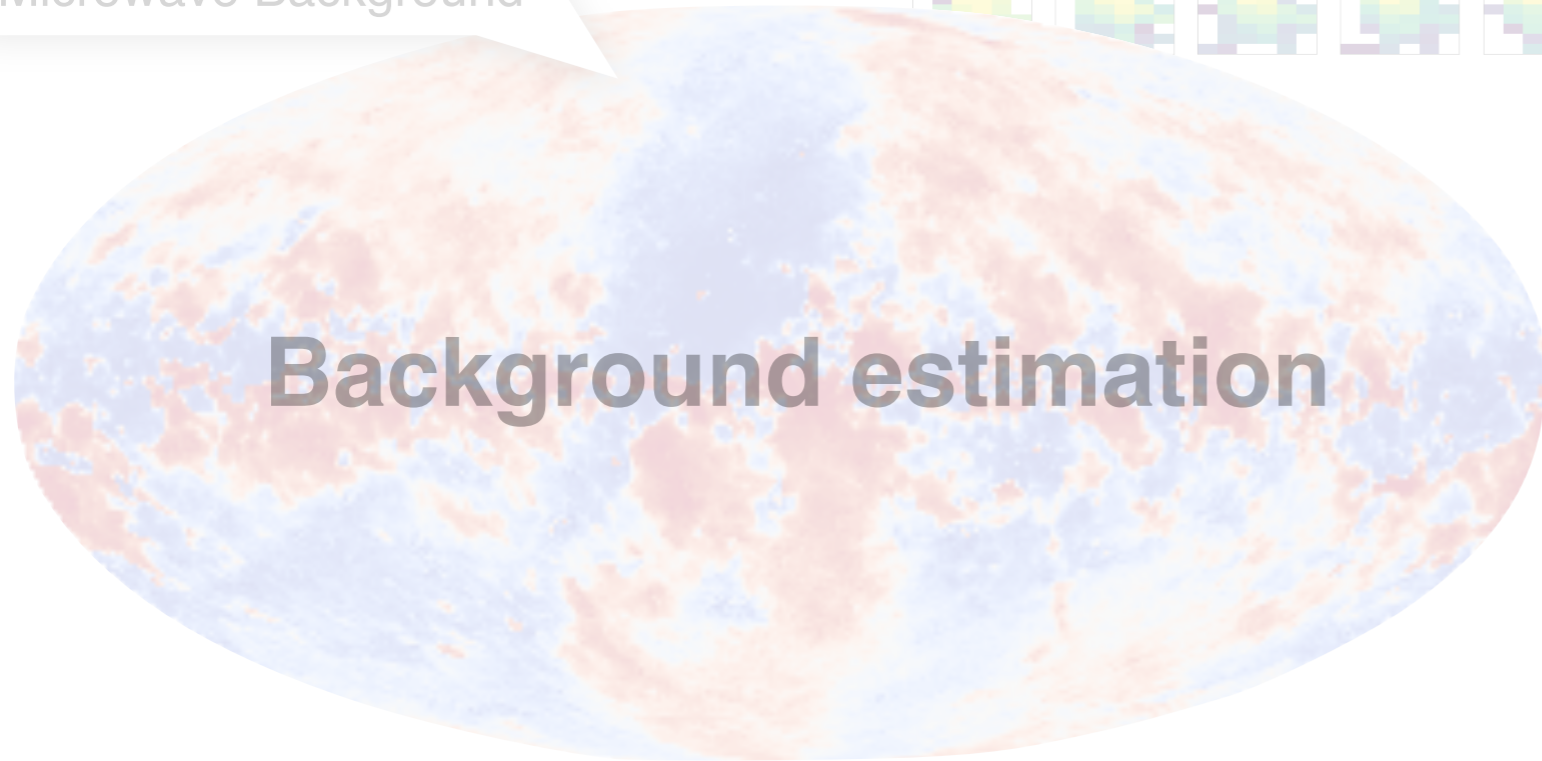


(amount of data discarded by standard trigger)

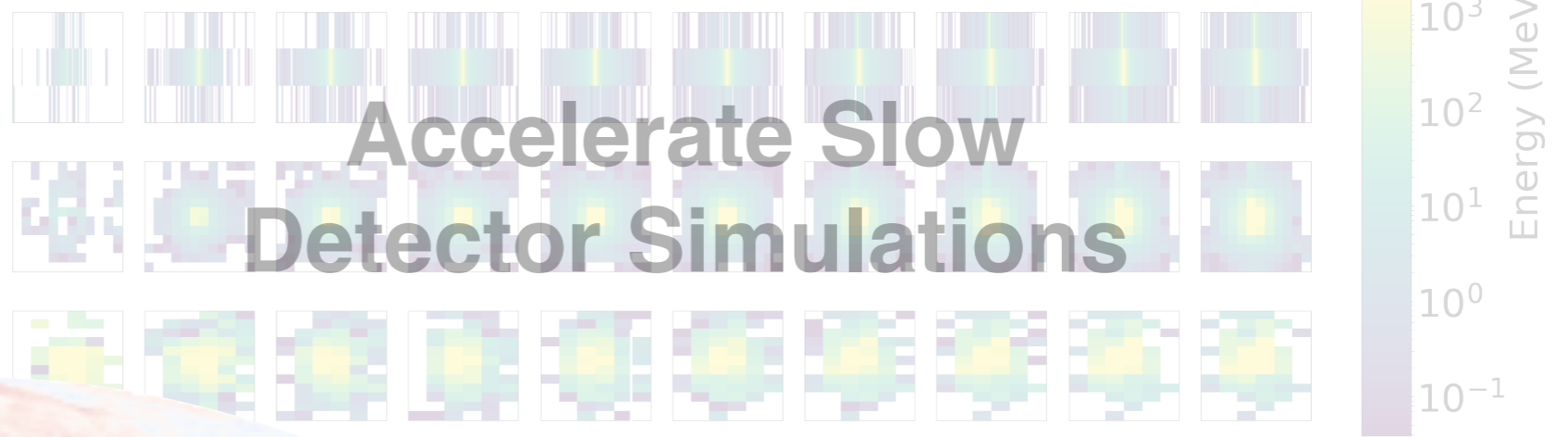
# Generative Models for Particle/Nuclear/Astro

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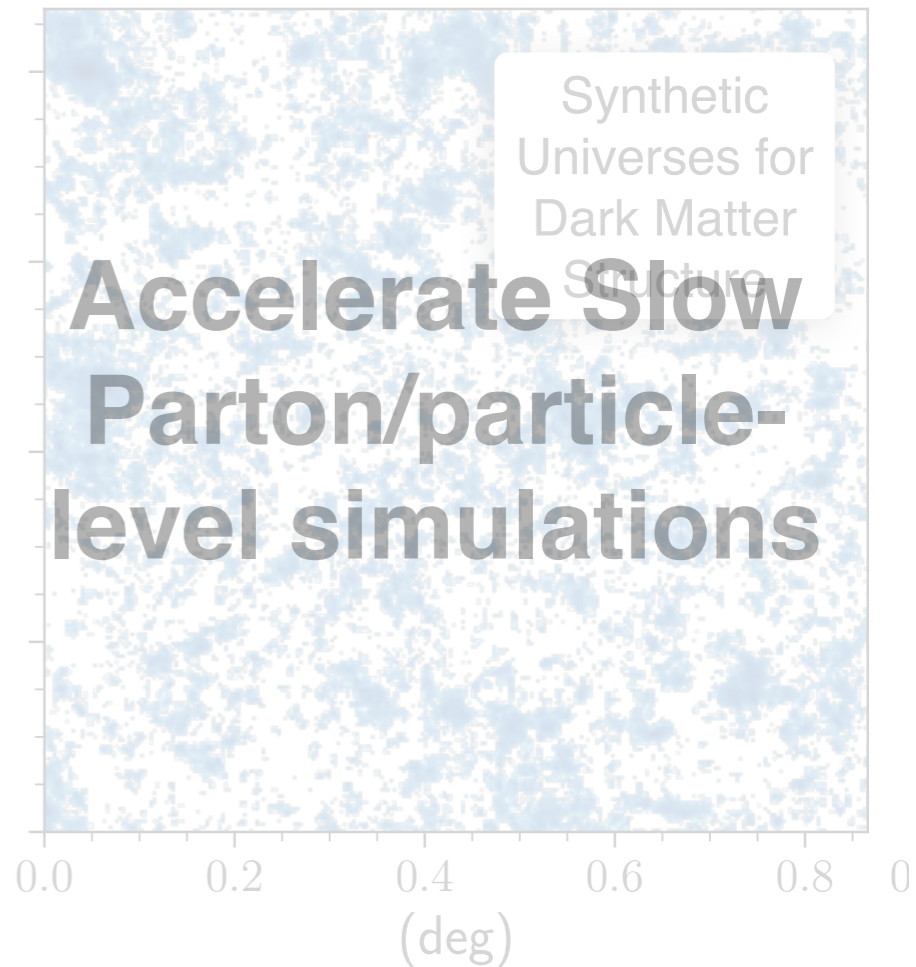
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M. Paganini, L. De Oliveira, B. Nachman, Phys. Rev. Lett. 120 (2018) 042003

The Structure of Radiation in the Quantum Strong Force

Infer Parton/particle-level Dynamics



M. Mustafa, et al., Comp. Astrophysics and Cosmology 6 (2019)



# Inferring Parton/particle-level Dynamics



95

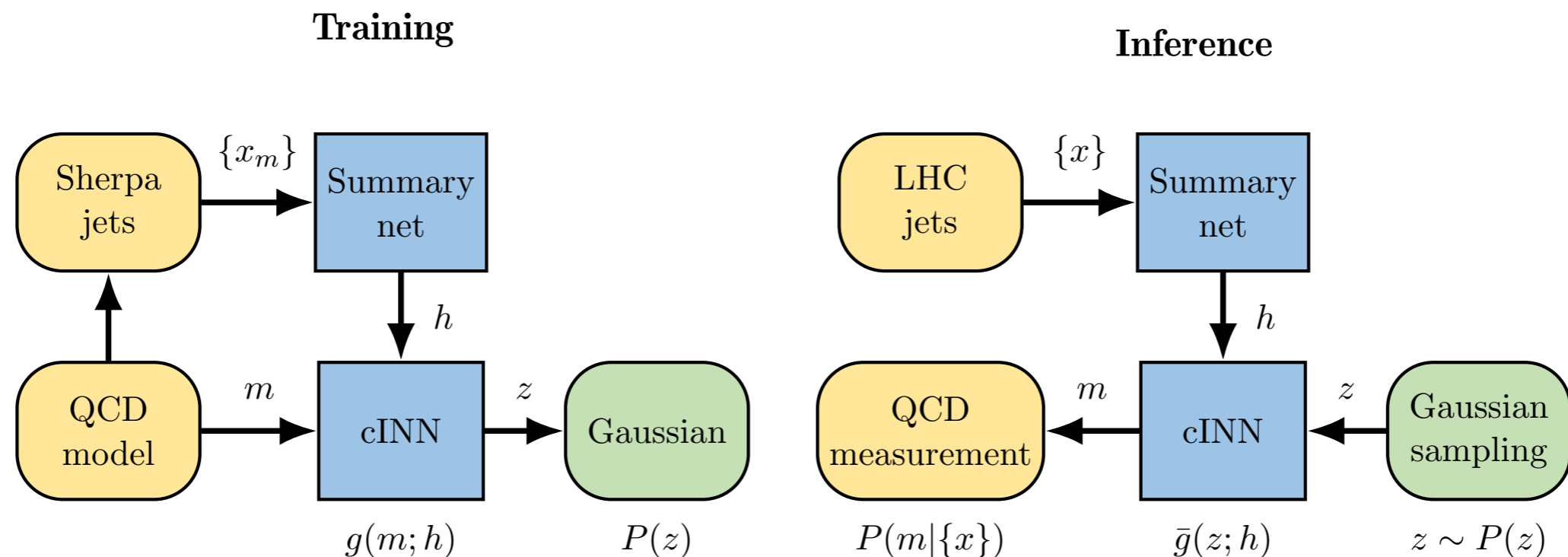
Can we use generative models directly for inference?  
(and not “just” for augmenting/accelerating simulation)



# Inferring Parton/particle-level Dynamics

Can we use generative models directly for inference?  
(and not “just” for augmenting/accelerating simulation)

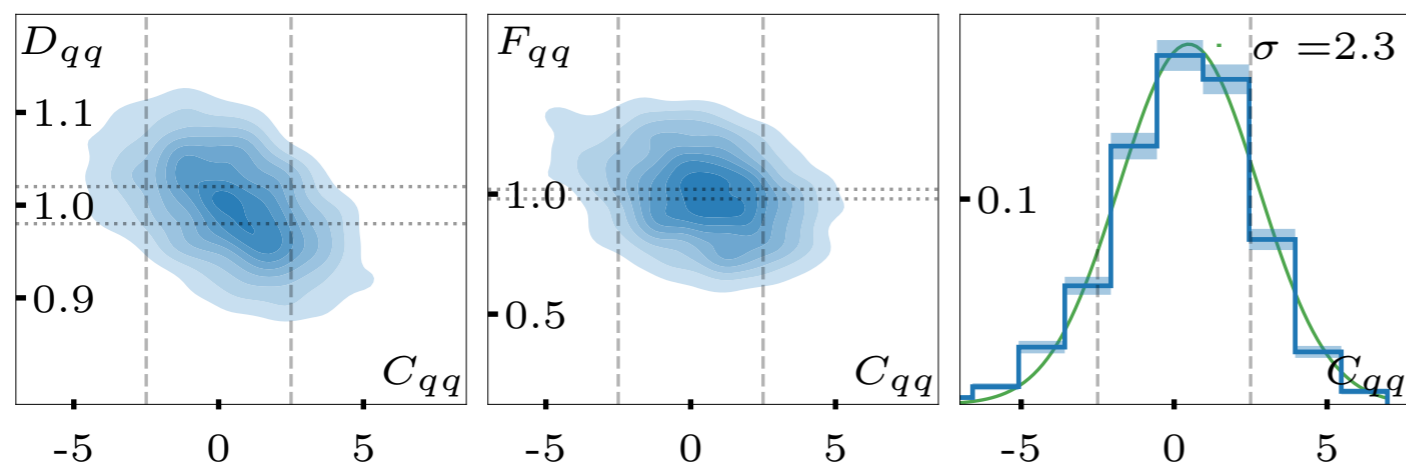
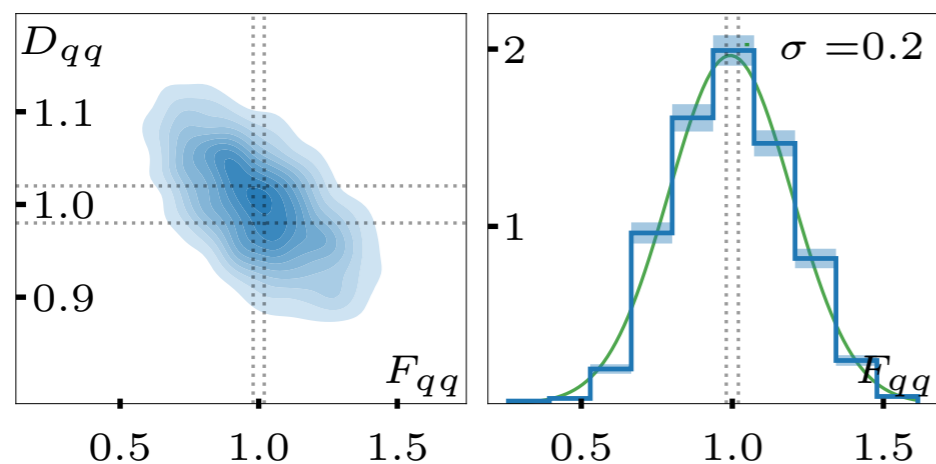
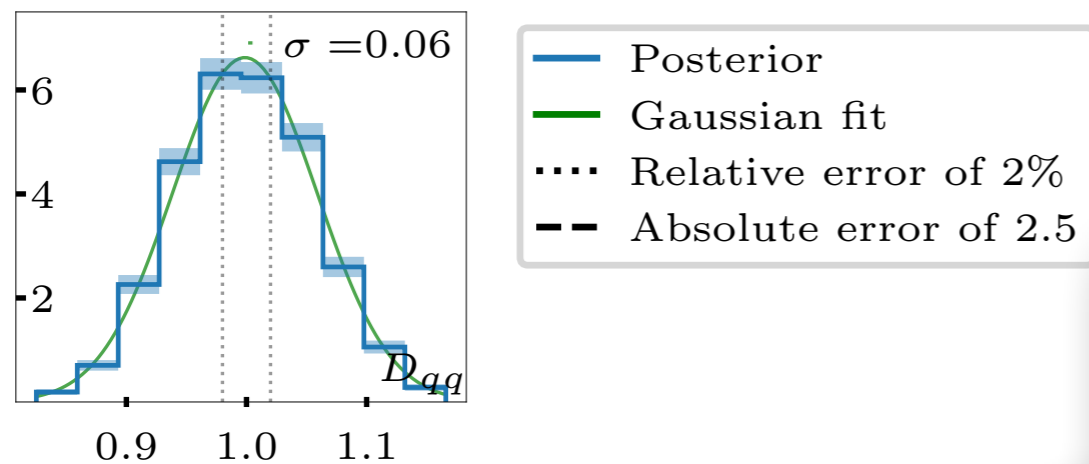
Example 1: Inferring fragmentation functions



See also 1804.09720 (“JUNIPR”) and 2012.06582 (GAN-based)

# Inferring Parton/particle-level Dynamics

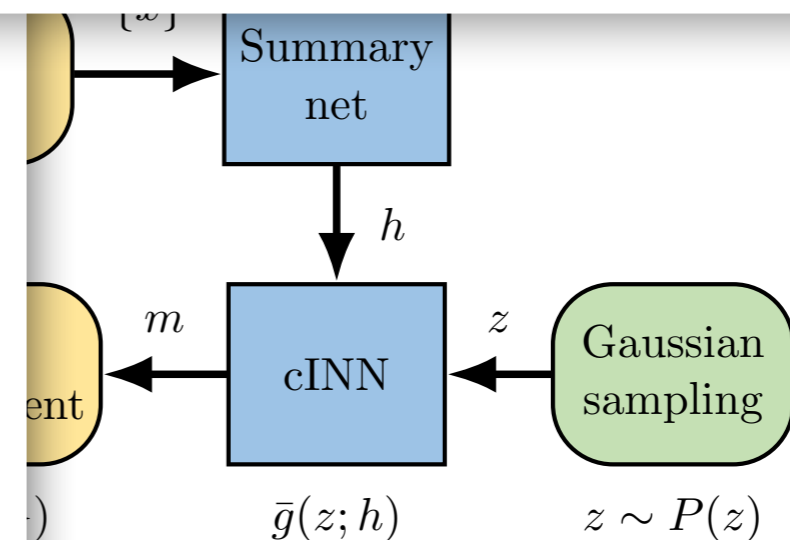
directly for inference?  
(accelerating simulation)



$$P_{qq}(z, y) = C_F \left[ D_{qq} \frac{2z(1-y)}{1-z(1-y)} + F_{qq}(1-z) + C_{qq}yz(1-z) \right]$$

$$P_{gq}(z, y) = T_R \left[ F_{qq} (z^2 + (1-z)^2) + C_{gq}yz(1-z) \right]$$

$$P_{gg}(z, y) = 2C_A \left[ D_{gg} \left( \frac{z(1-y)}{1-z(1-y)} + \frac{(1-z)(1-y)}{1-(1-z)(1-y)} \right) + F_{gg}z(1-z) + C_{gg}yz(1-z) \right]$$



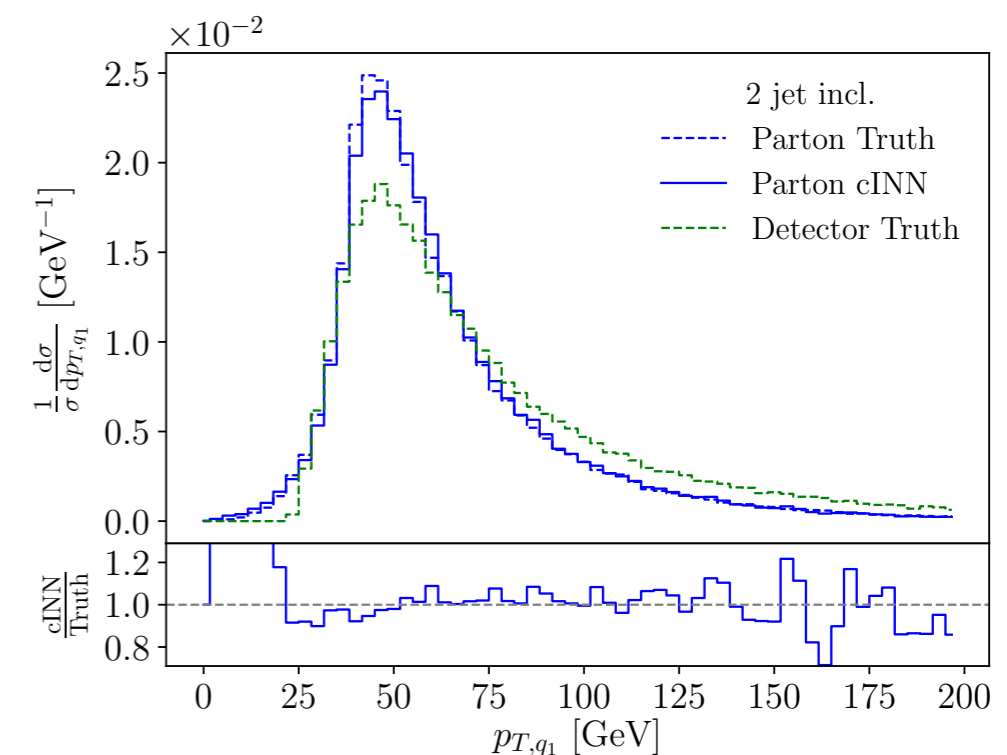
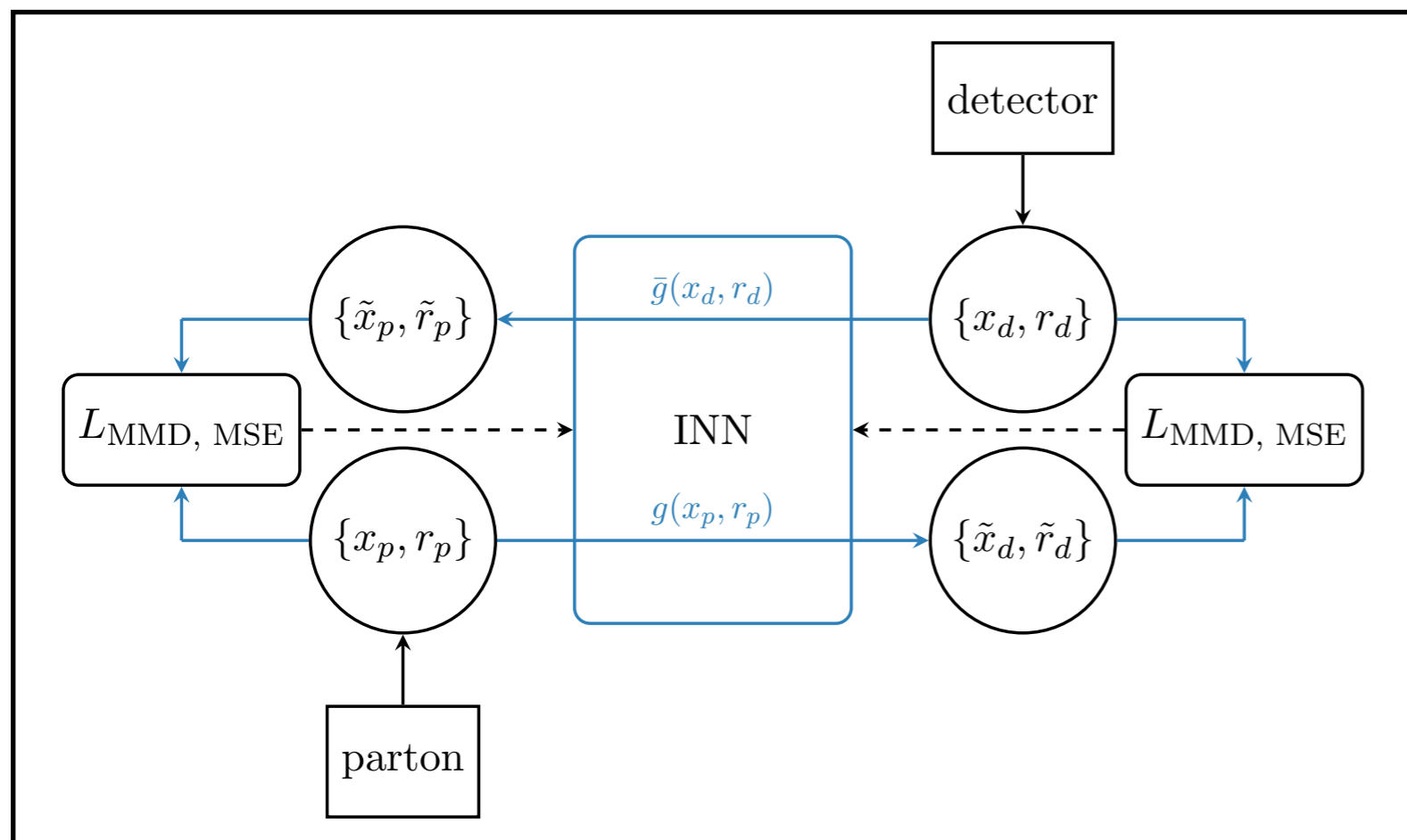
See also 1804.09720 (“JUNIPR”) and 2012.06582 (GAN-based)



# Infering Parton/particle-level Dynamics

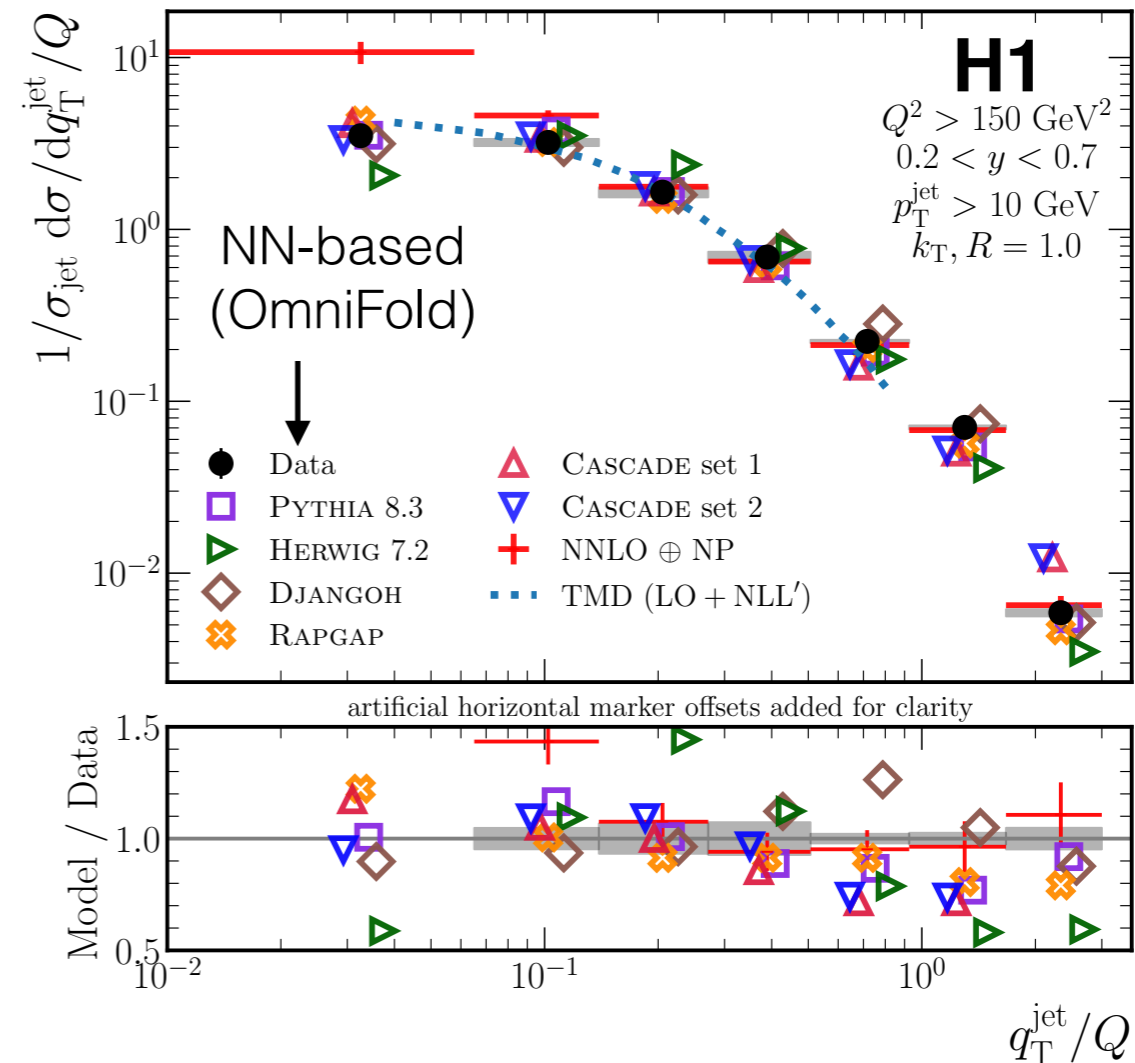
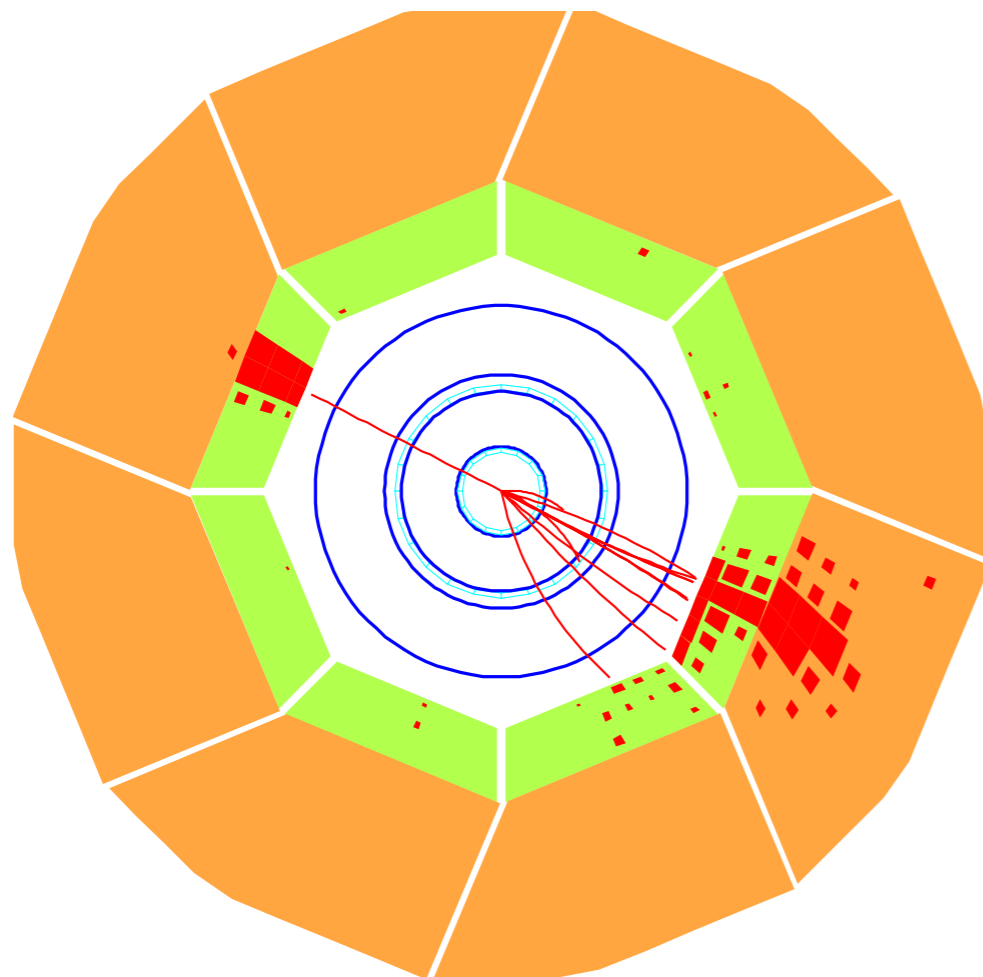
Can we use generative models directly for inference?  
(and not “just” for augmenting/accelerating simulation)

## Example 2: Unfolding



See also 1911.09107 (“OmniFold”) and 2101.08944 (“OTUS”)

# Infering Parton/particle-level Dynamics



These (and related) methodologies are being studied for  $ep$  collisions!

See also 1911.09107 (“OmniFold”) and 2101.08944 (“OTUS”)

el.  
Truth  
cINN  
or Truth

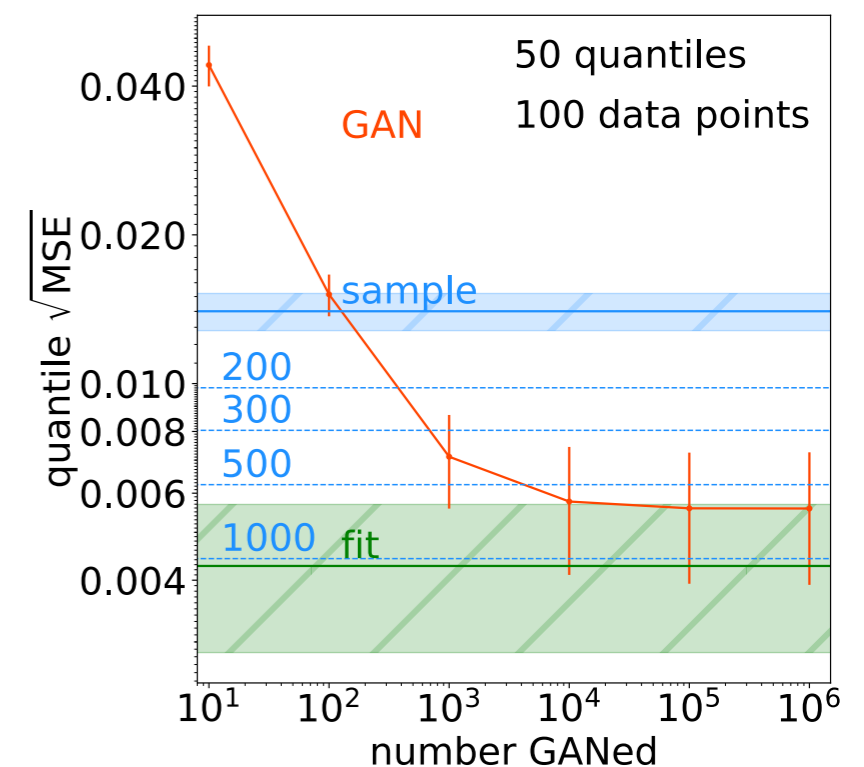
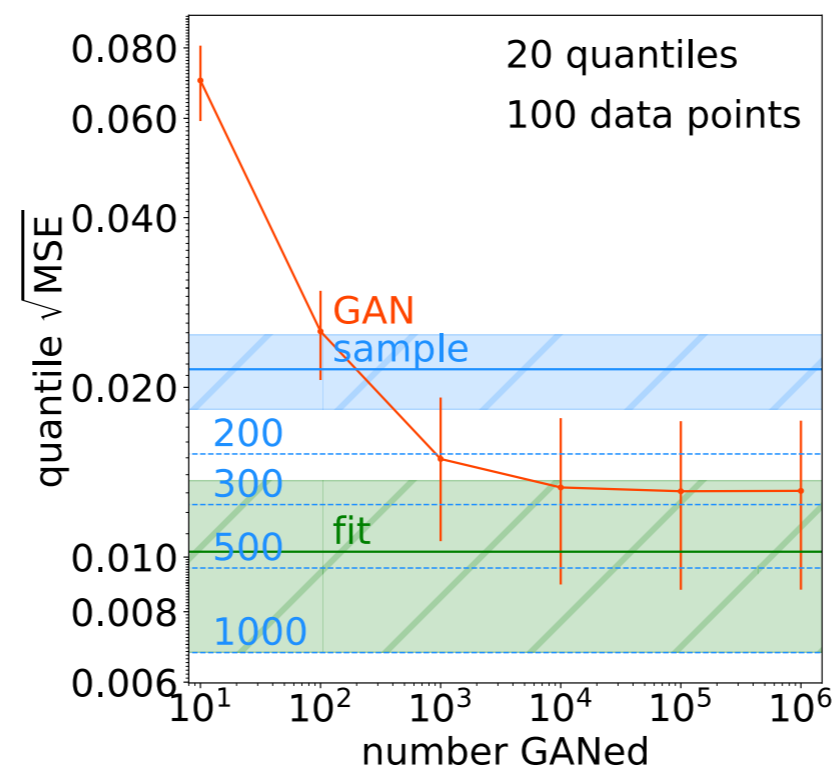
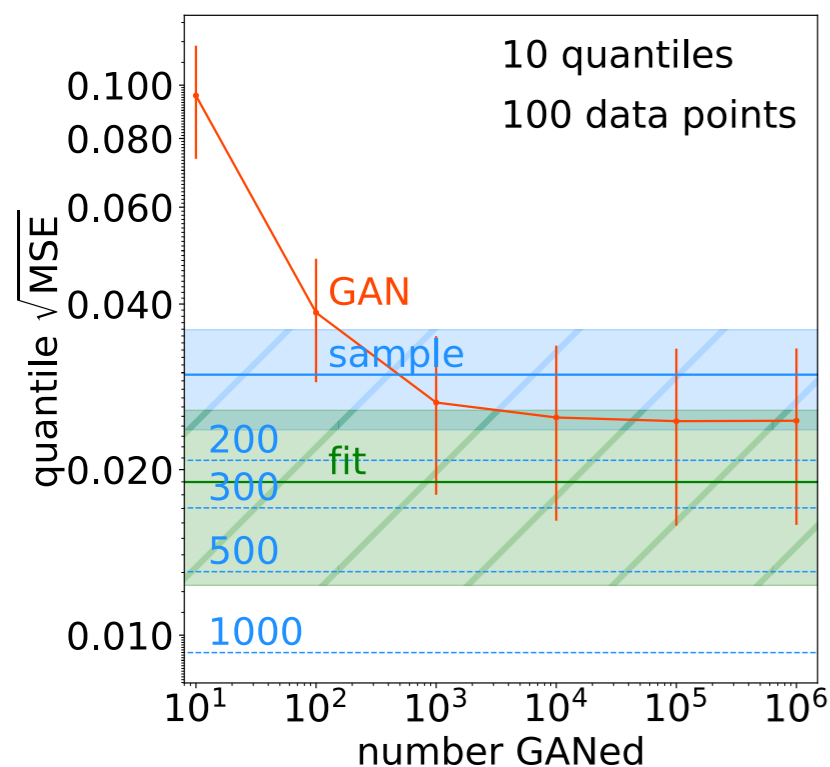
175 200



# Uncertainties

Performance continues to improve on many fronts. As we integrate these tools into our workflows, we need to think about uncertainties.

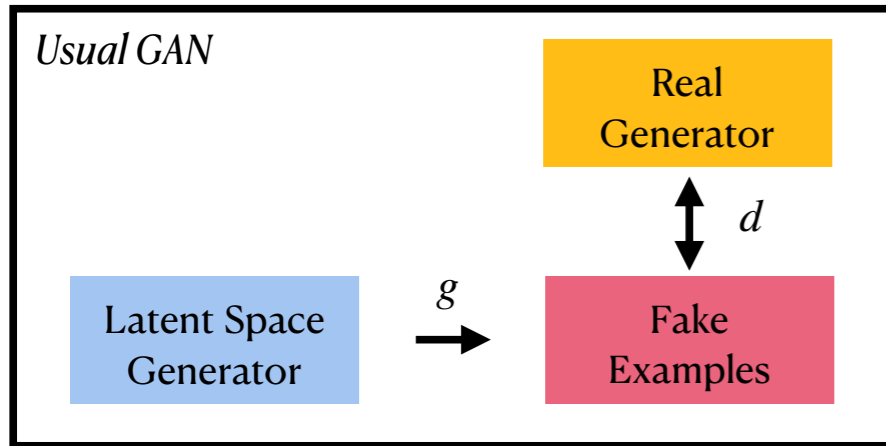
One question is about the **statistical power** of samples from a generative model. This depends on the implicit or explicit information we encode in the networks.



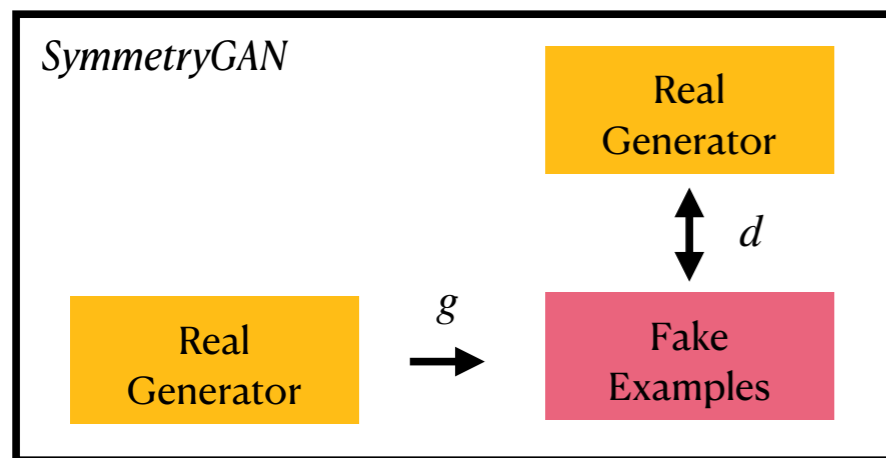
See also 1909.03081, 2002.06307, 2104.04543 (Generative Bayesian NNs), and 2107.08979 (“resampling”)



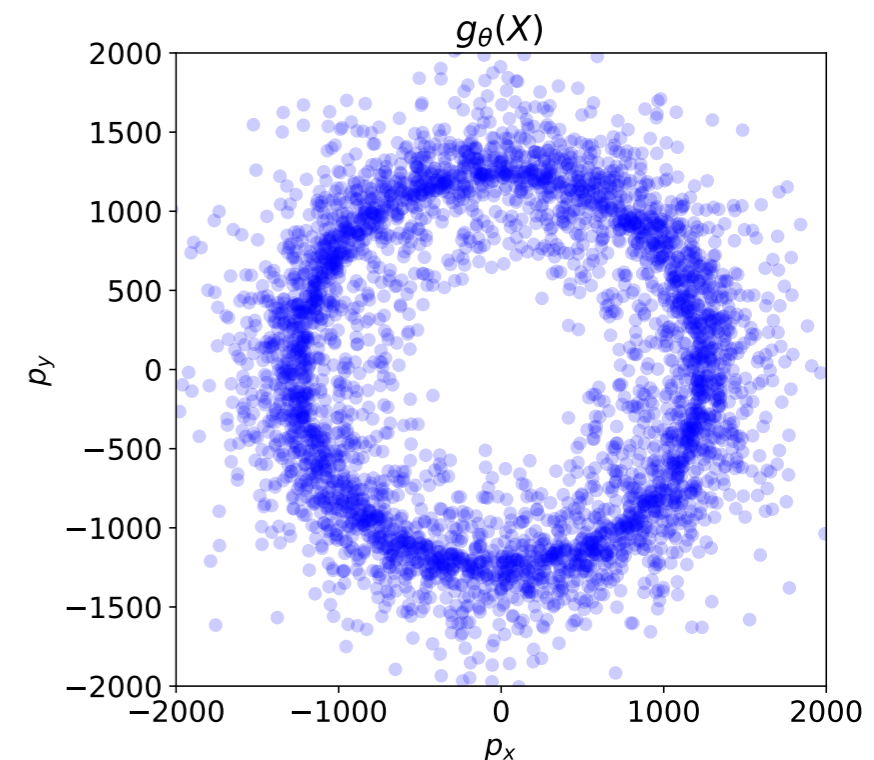
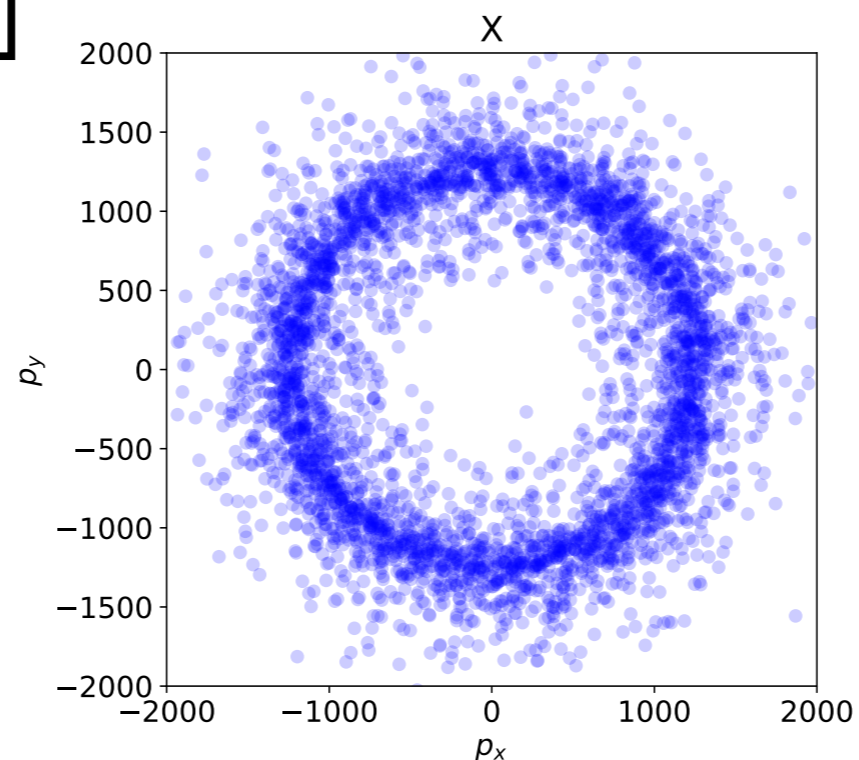
# What else can we do?



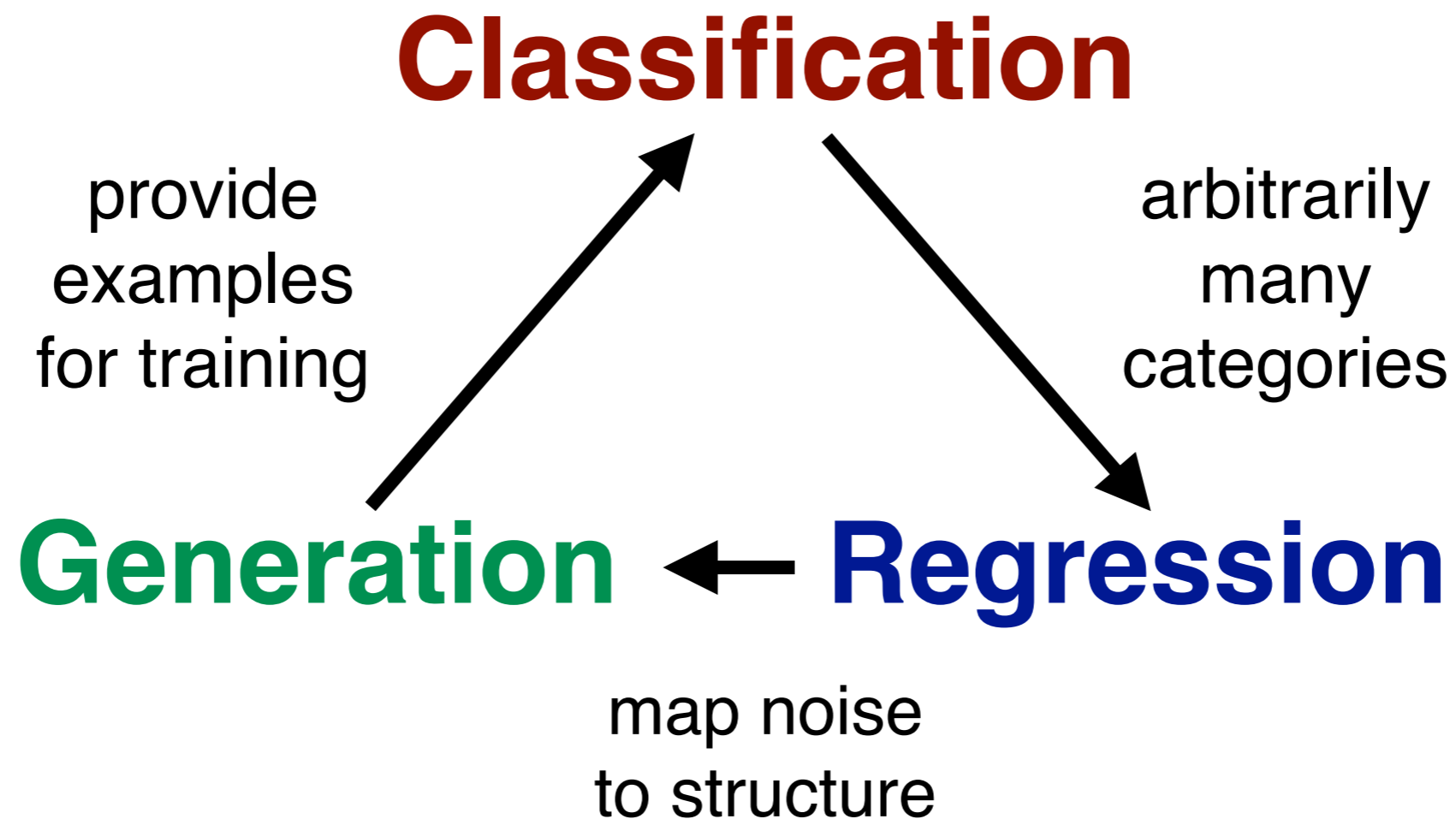
The framework of generative models is quite flexible and we can do more than generate events.



For example, can **discover symmetries** in data!







+related topics like anomaly detection, simulation-based inference, ...

<https://iml-wg.github.io/HEPML-LivingReview/>

## HEPML-LivingReview

### A Living Review of Machine Learning for Particle Physics

*Modern machine learning techniques, including deep learning, is rapidly being applied, adapted, and developed for high energy physics. The goal of this document is to provide a nearly comprehensive list of citations for those developing and applying these approaches to experimental, phenomenological, or theoretical analyses. As a living document, it will be updated as often as possible to incorporate the latest developments. A list of proper (unchanging) reviews can be found within. Papers are grouped into a small set of topics to be as useful as possible. Suggestions are most welcome.*

[download](#) [review](#)

See also <https://arxiv.org/abs/2102.02770>

Now we'll hear quite a bit about Quantum Machine Learning

It seems like the jury is still out about the near term utility of QML  
(see also Sulaiman's talk tomorrow for a counter point to many recent papers in this area)

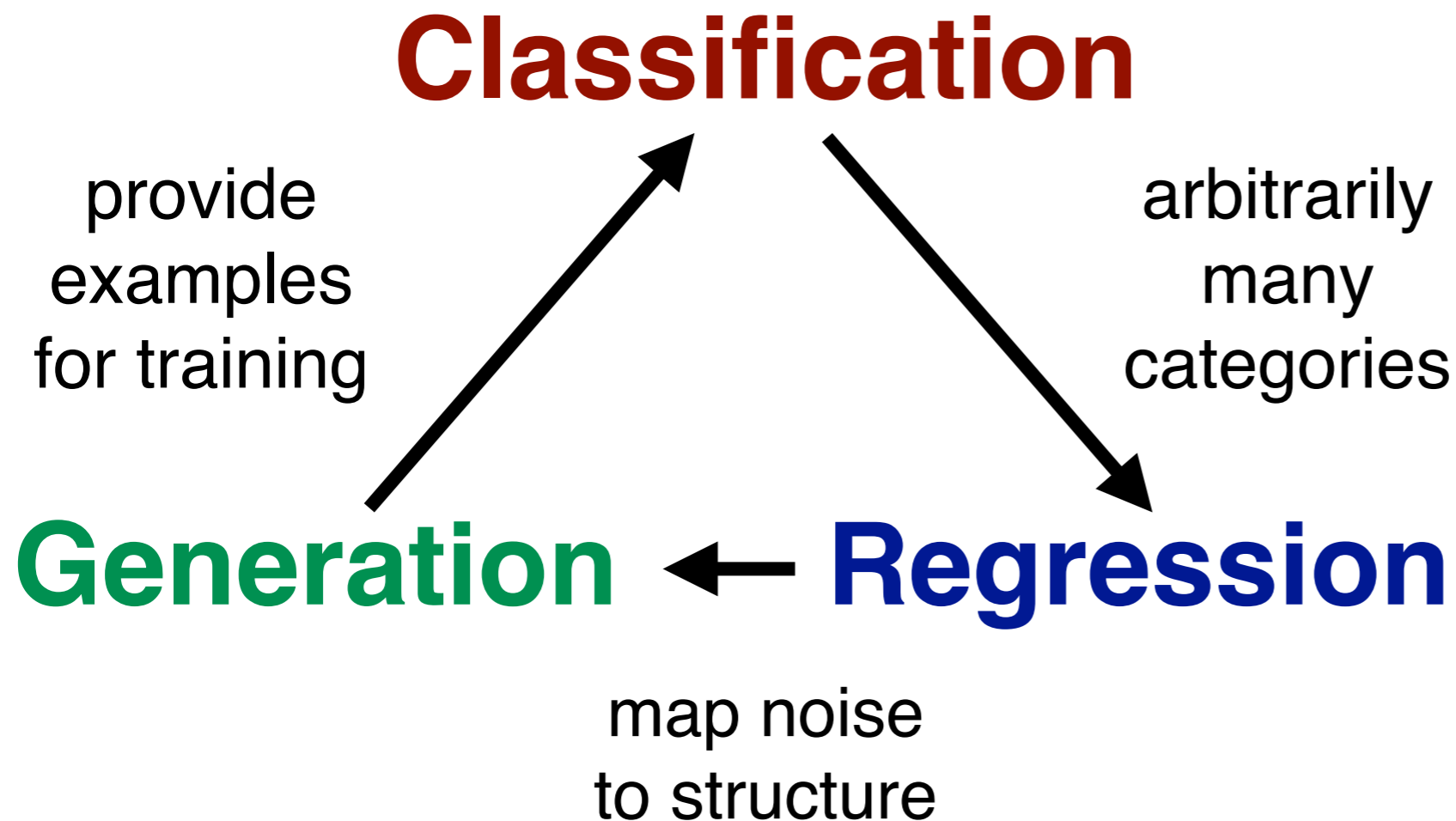
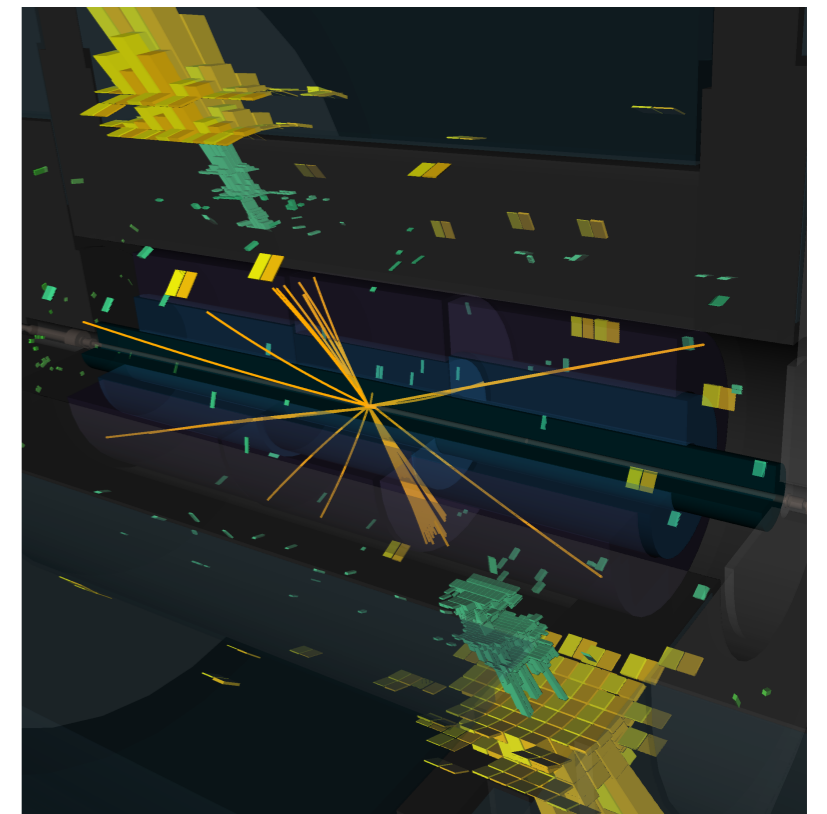
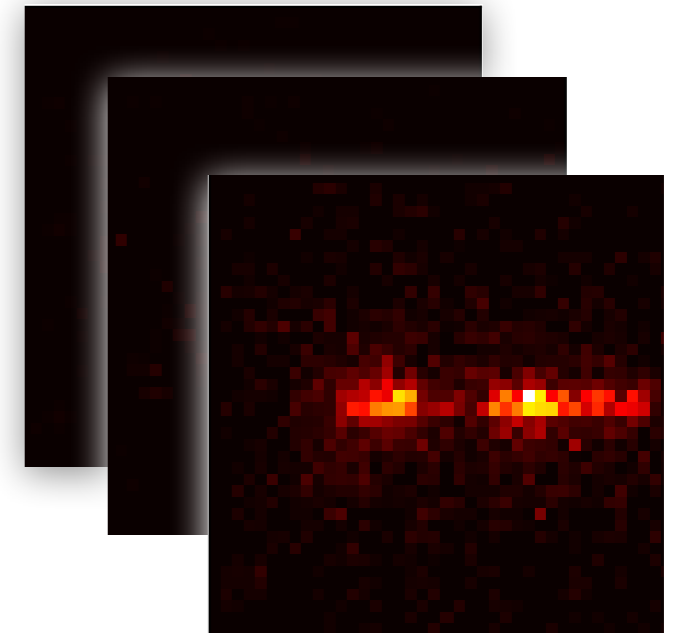
*QML is an exciting tool and I'll certainly watch developments with interest.  
I look forward to the results and discussion today and tomorrow!*

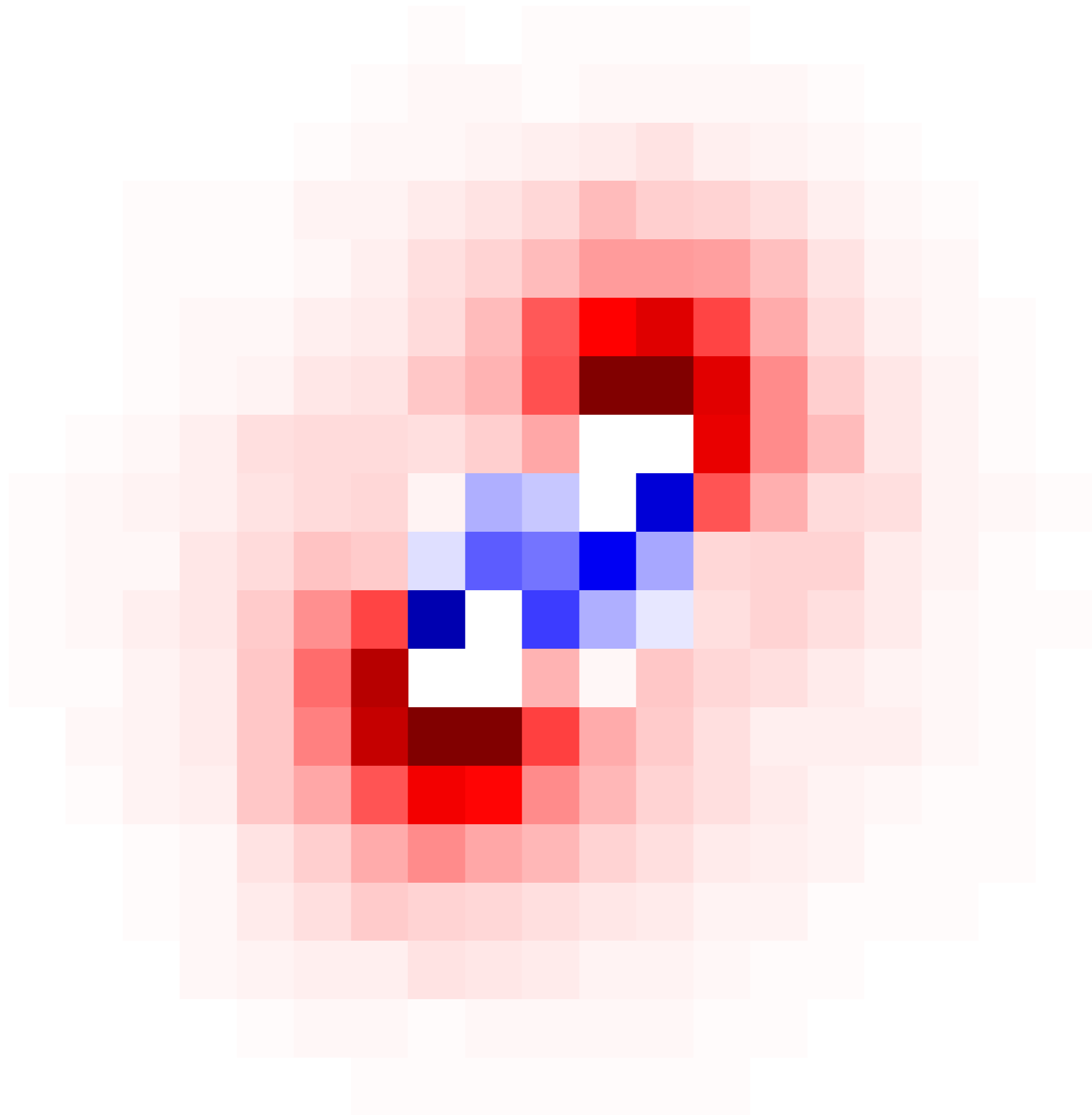
**I'll stress that classical ML is already impacting HEP science and it will continue to grow in importance (!)**

# Conclusions and Outlook

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Particle physics data are complex and unique - modern machine learning tools will help us them to the fullest potential to discover something new and fundamental!





Fin.

Backup

