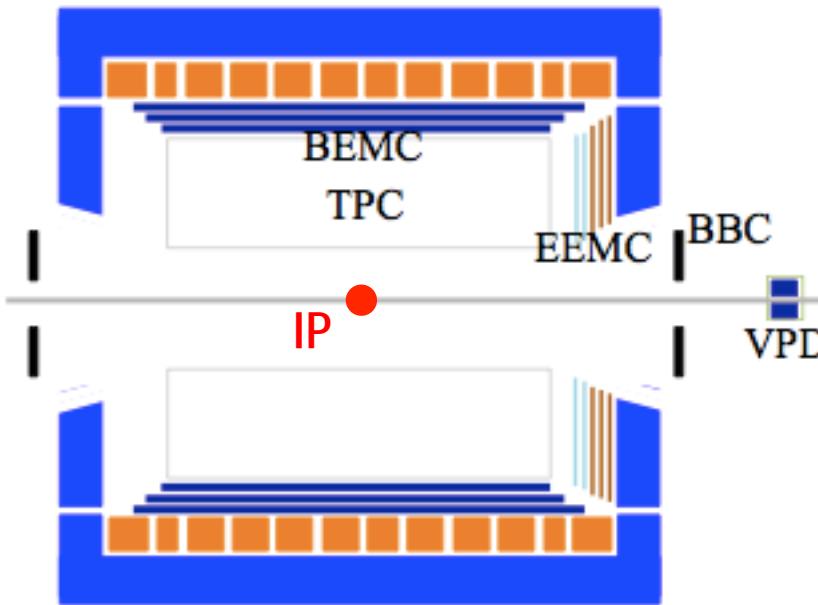


Highlights and future plan of RHICf

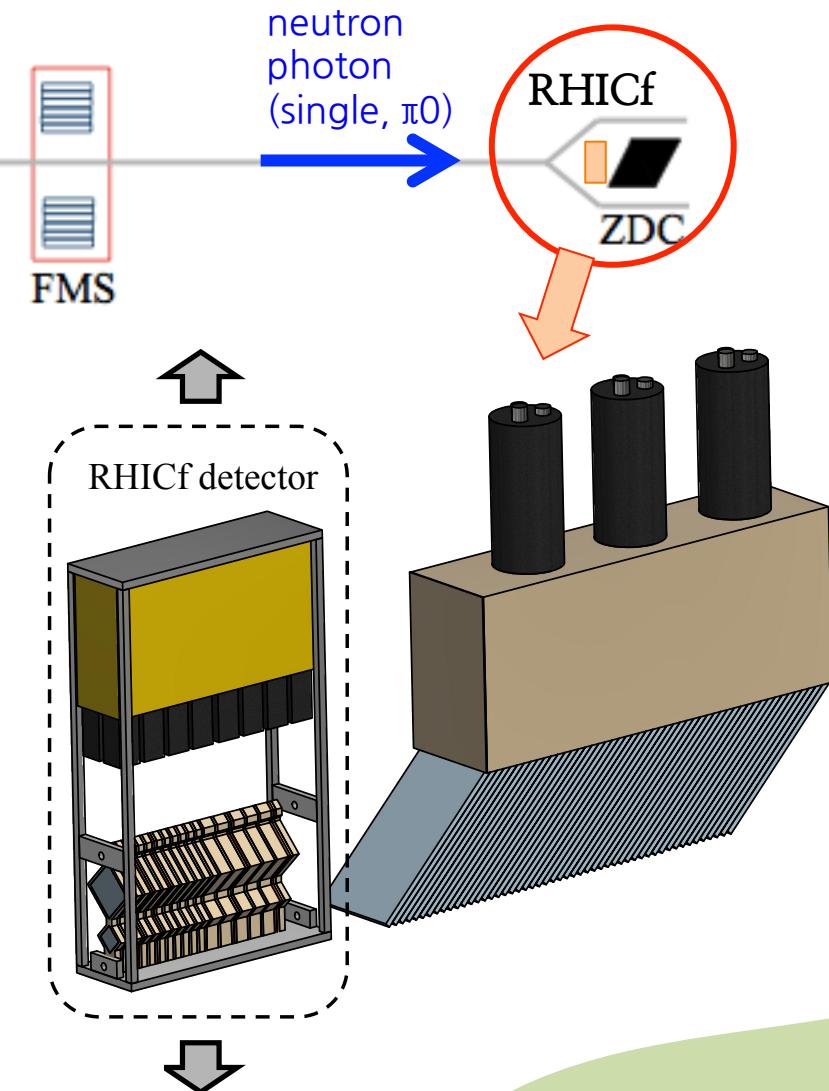
**Minho Kim (RIKEN BNL Research Center)
on behalf of the RHICf collaboration**

RHIC forward (RHICf) experiment

STAR experiment

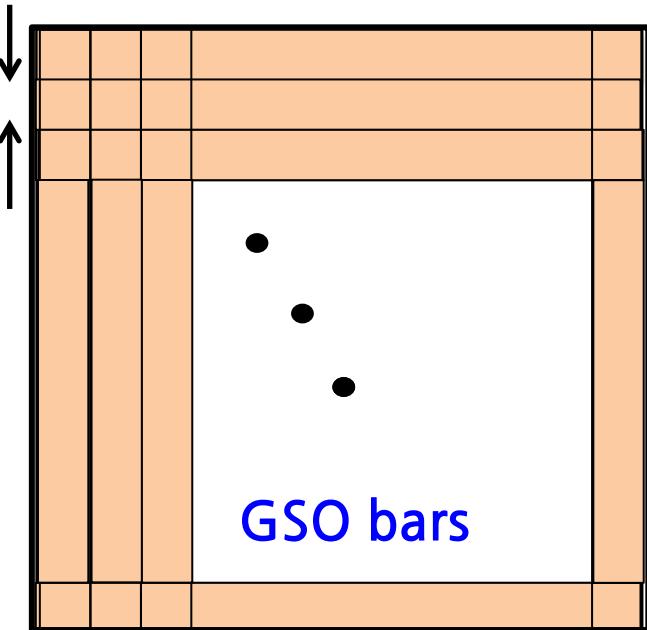
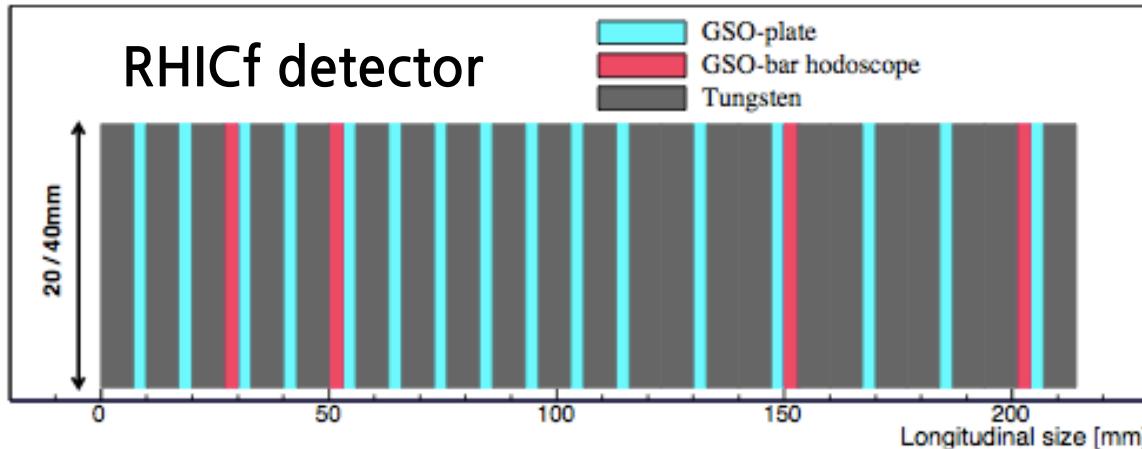


- Operated at STAR in polarized $p + p$ collisions at $\sqrt{s} = 510$ GeV in June 2017.
- RHICf detector was installed in front of the ZDC.
- $\eta > 6$,
 $0.2 < x_F < 1.0$, and
 $0.0 < p_T < 1.0$ GeV/c.

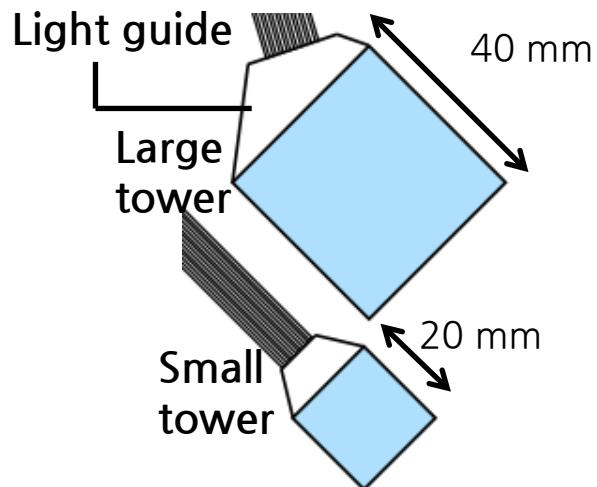


RHICf detector

Side view

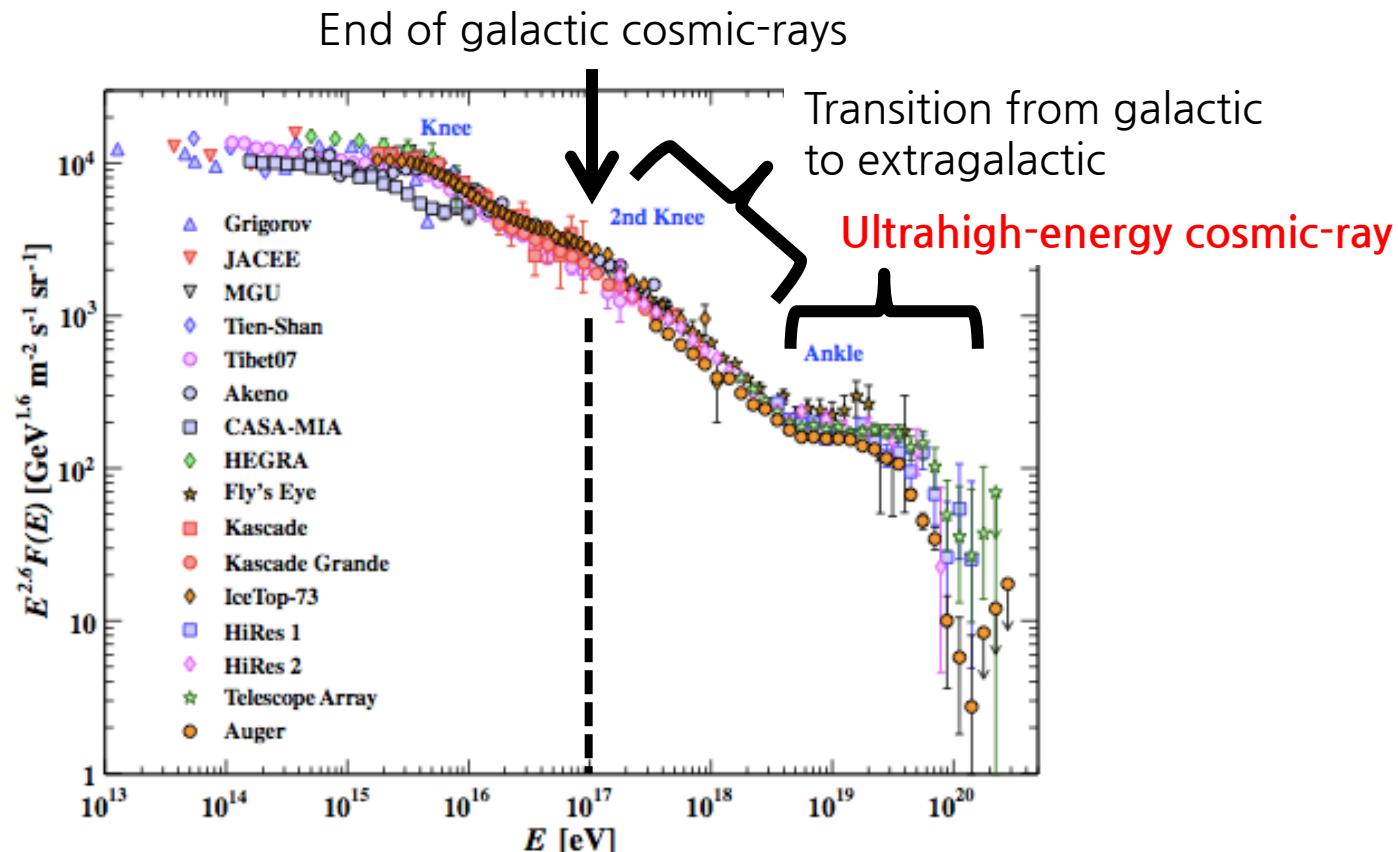


Front view



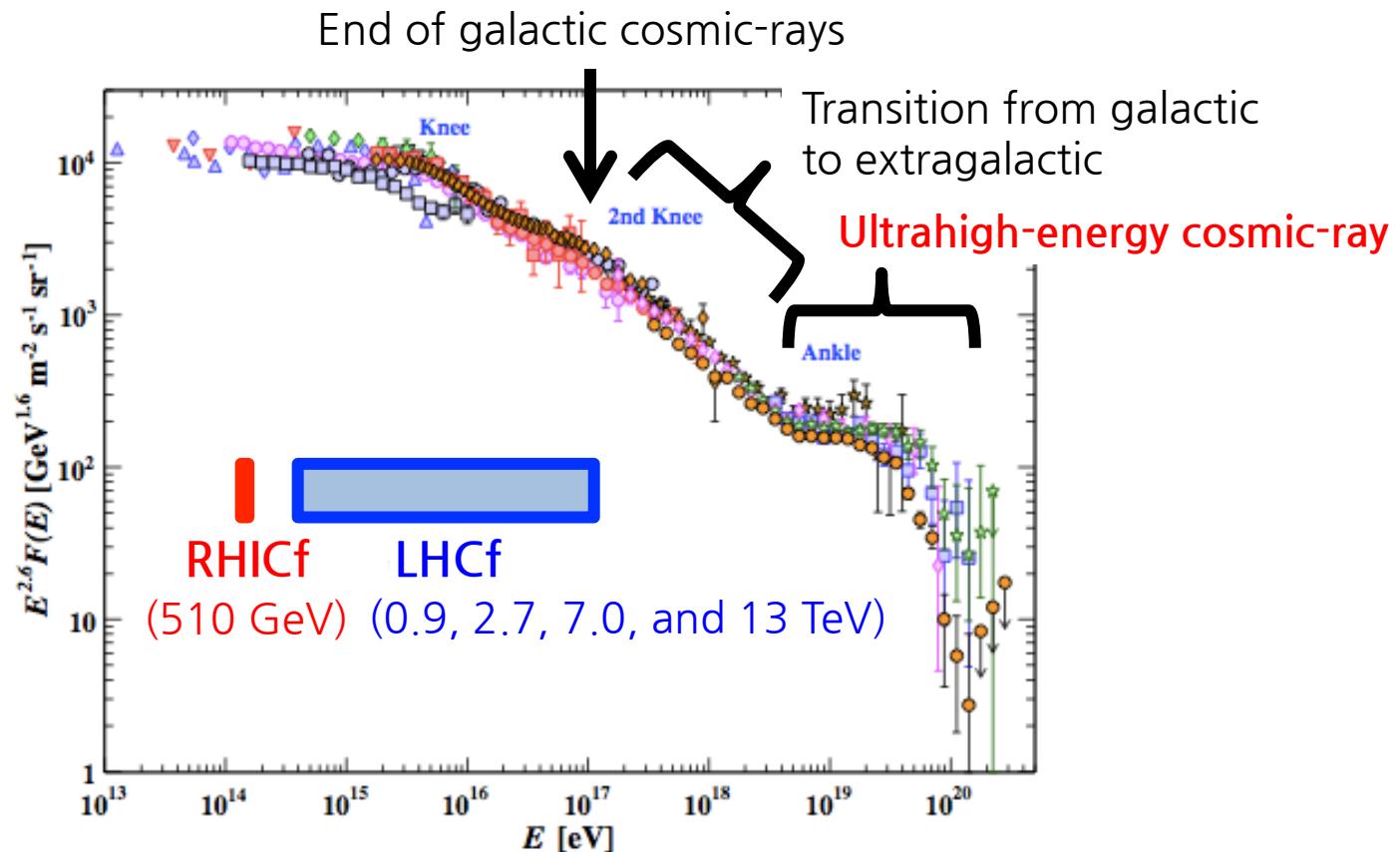
- 17 tungsten absorbers ($44 X_0$, $1.6 \lambda_{int}$), 16 GSO plates, and 4 layers of GSO bars.
- $\sigma_x \sim 0.1 \text{ mm}$, $\sigma_E \sim 2\%$ for 250 GeV photon.
- $\sigma_x \sim 1.0 \text{ mm}$, $\sigma_E \sim 35\%$ for 250 GeV neutron.
- Shower is triggered using energy deposits in three successive layers.

1. Cross section measurement



- Cosmic-ray above 10^{18} eV is called ultrahigh-energy cosmic ray (UHECR).
- Different interaction model predicts different acceleration mechanism for the origin of the UHECRs.
 - Cross section data can constrain and tune the existing models.

1. Cross section measurement

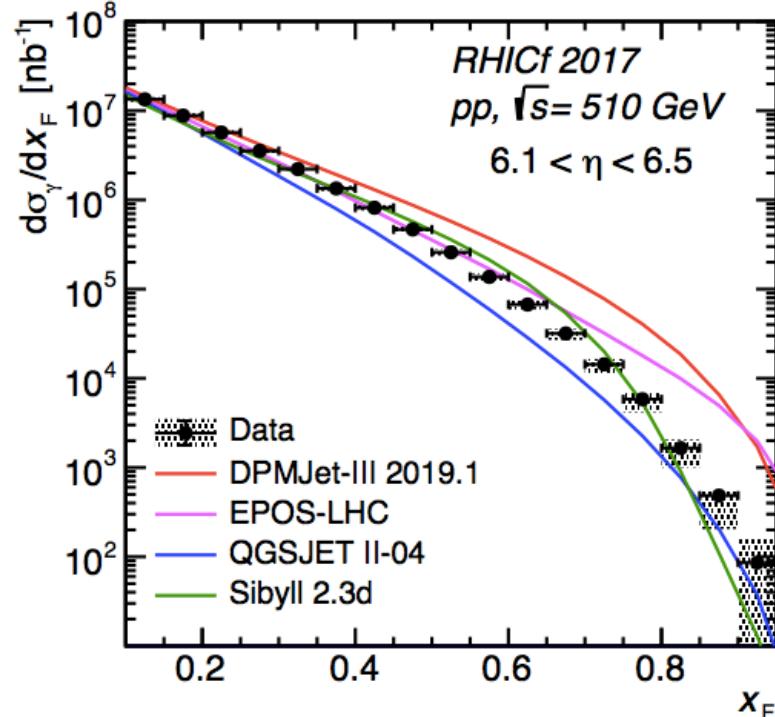
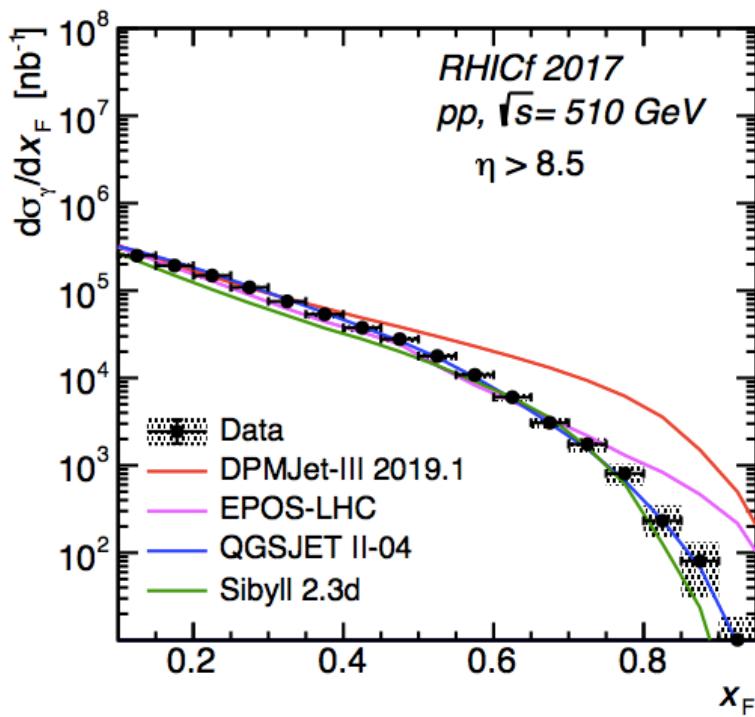


- Cross sections of neutron, photon, and π^0 have been measured by the LHCf experiment up to $\sqrt{s} = 13$ TeV.
- \sqrt{s} dependence of the forward particle production will be tested in a wide energy range by comparing the RHICf with LHCf data.

New result!

Photon cross section

arXiv:2203.15416

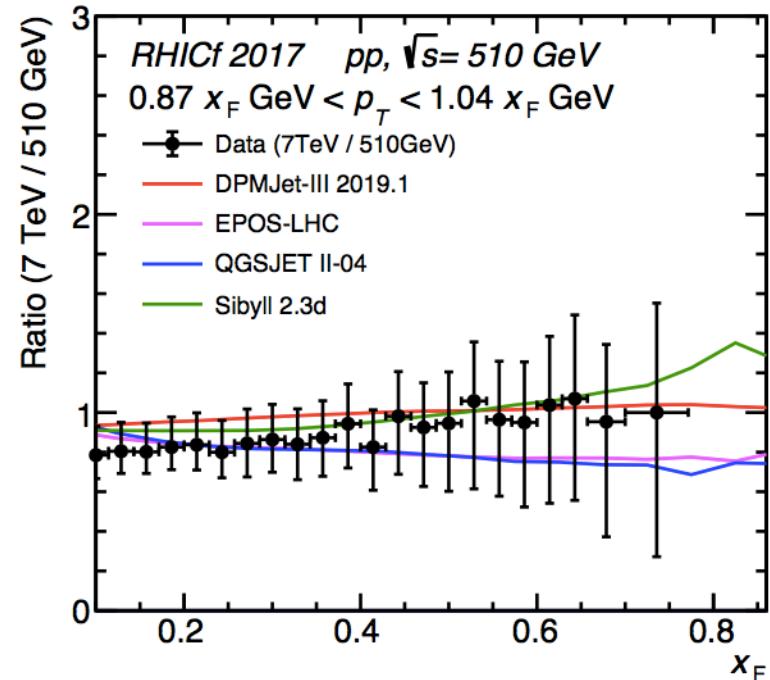
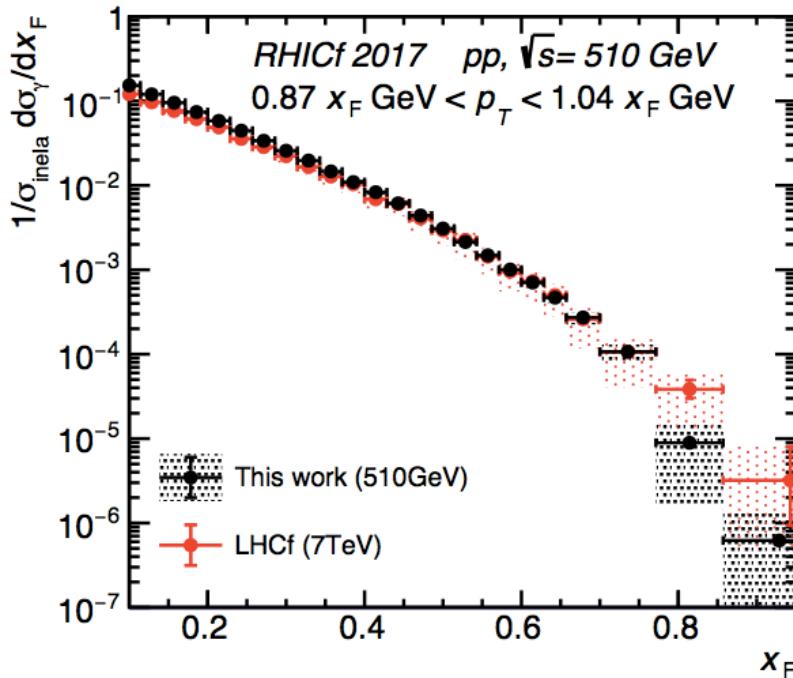


- DPMJet-III 2019.1 and EPOS-LHC well reproduce the data in the lower x_F region, but predict larger flux in the higher x_F .
- QGSJET-II 04 and Sibyll 2.3d show good agreement with data in the higher η , but show softer and harder slope in the lower η .

New result!

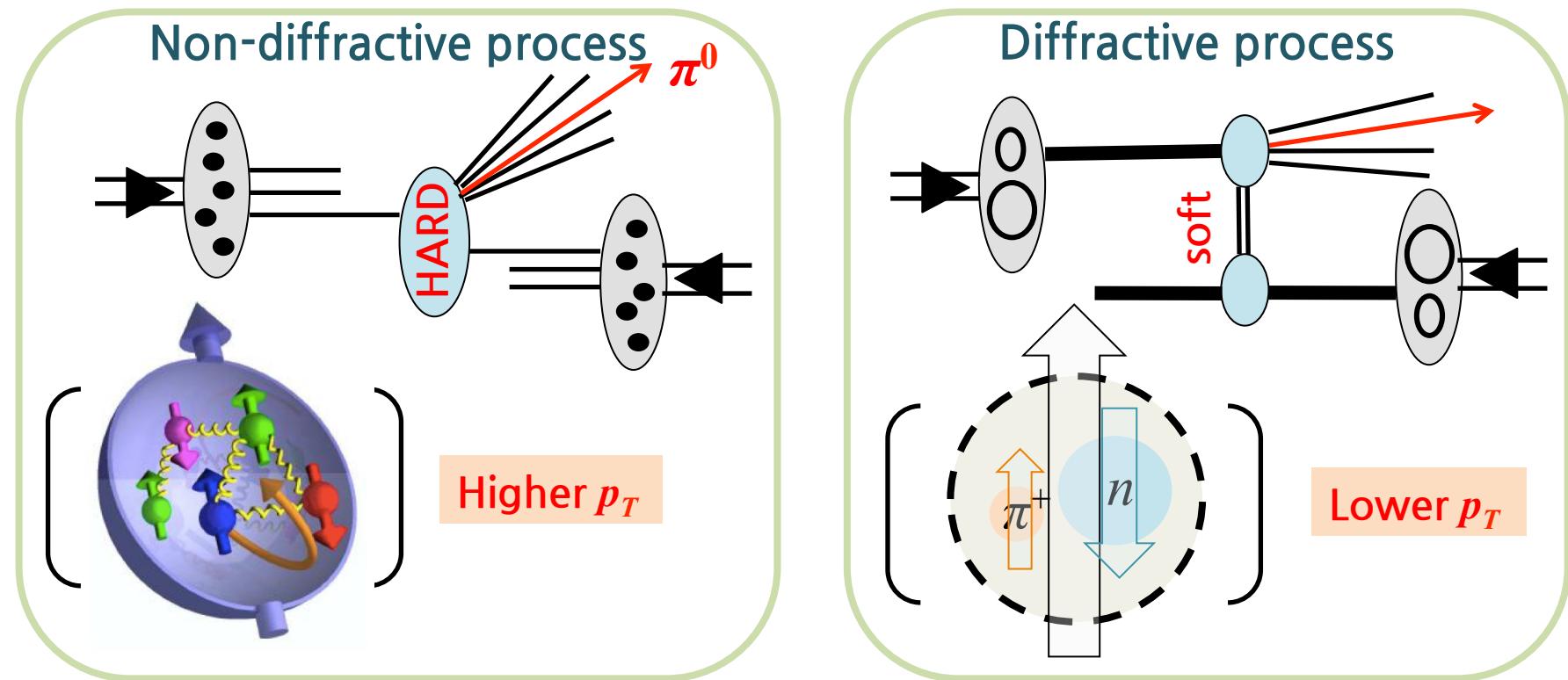
Photon cross section

arXiv:2203.15416



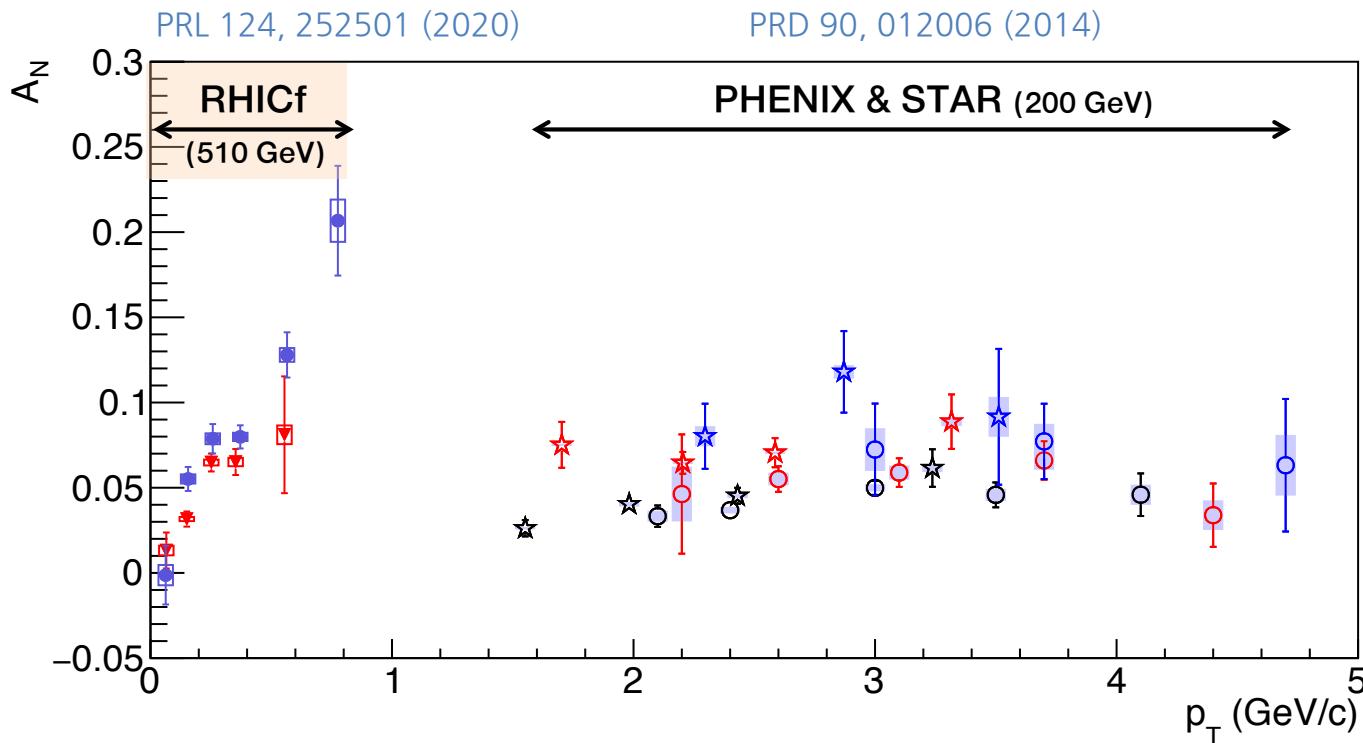
- RHICf result is consistent with LHCf result within the uncertainty.
- However, if the x_F scaling raw works or there is still a weak x_F dependence is not clear due to the uncertainty. → Will be more clear in the future publications.
- Comparison for neutron will also be studied soon.

2. A_N measurement (neutral pion)



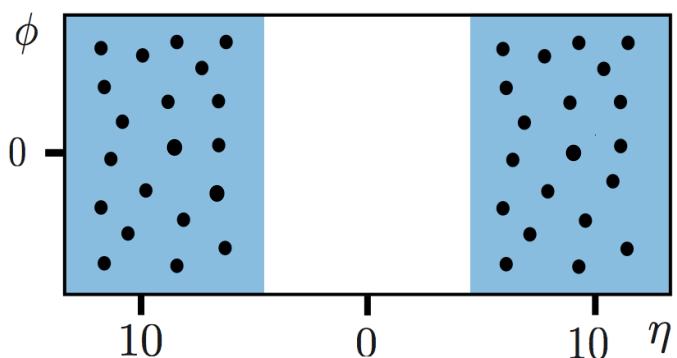
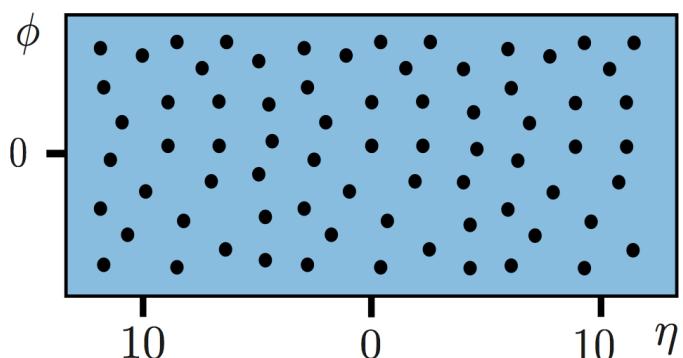
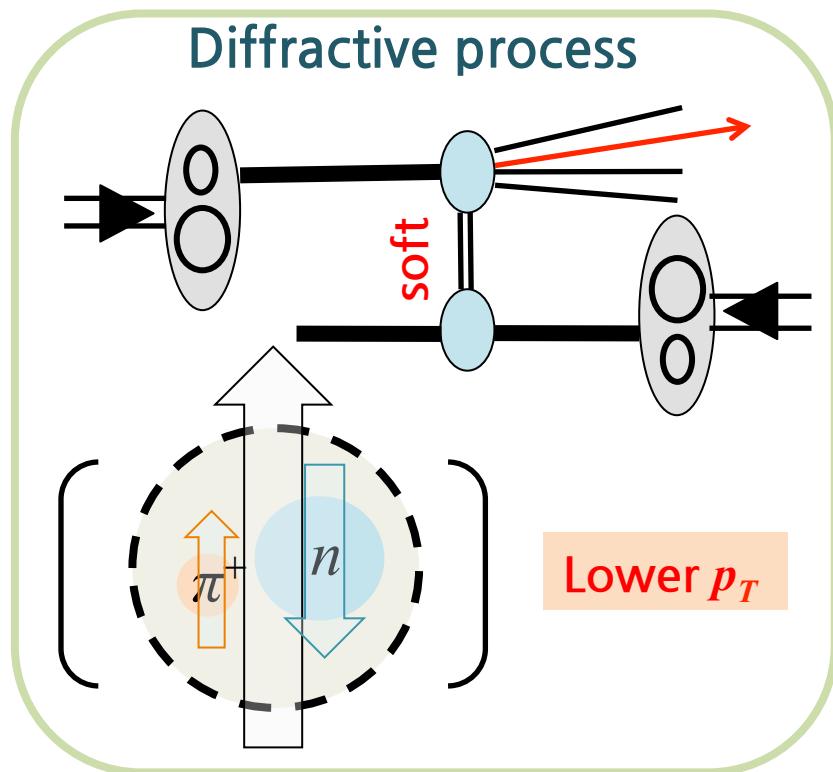
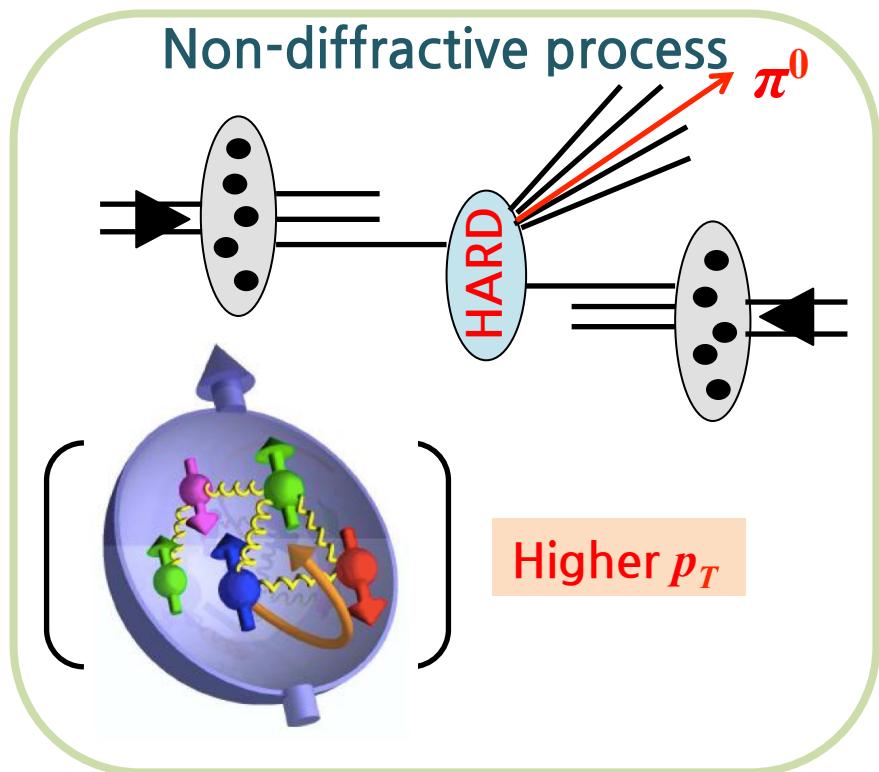
- The π^0 asymmetry has been measured and studied only in the quarks and gluons' degrees of freedom.
- In order to study a possible diffractive contribution to the π^0 asymmetry, the RHICf experiment measured the A_N for very forward π^0 production.

2. A_N measurement (neutral pion)



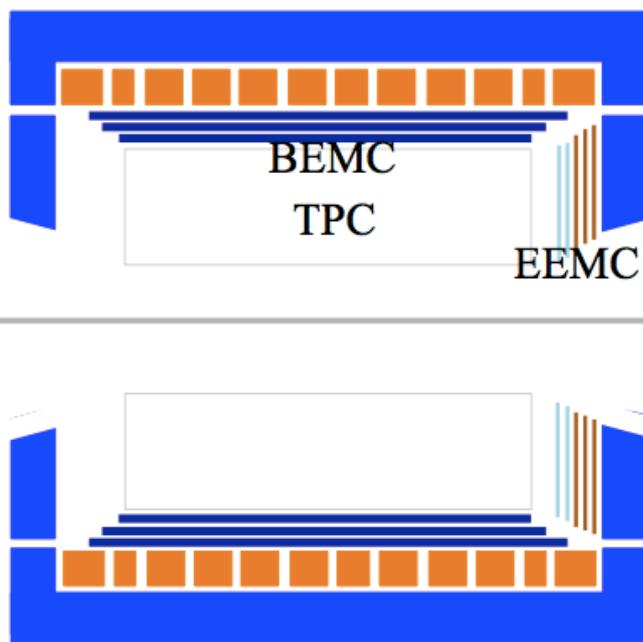
- Very forward ($\eta > 6$) neutral particle A_N is an important observable to study the spin-involved diffractive particle production mechanism.
- RHICf data showed a possible diffractive contribution to the finite $\pi^0 A_N$.
- Multi-dimensional analysis will be started using STAR detectors.

2. A_N measurement (neutral pion)

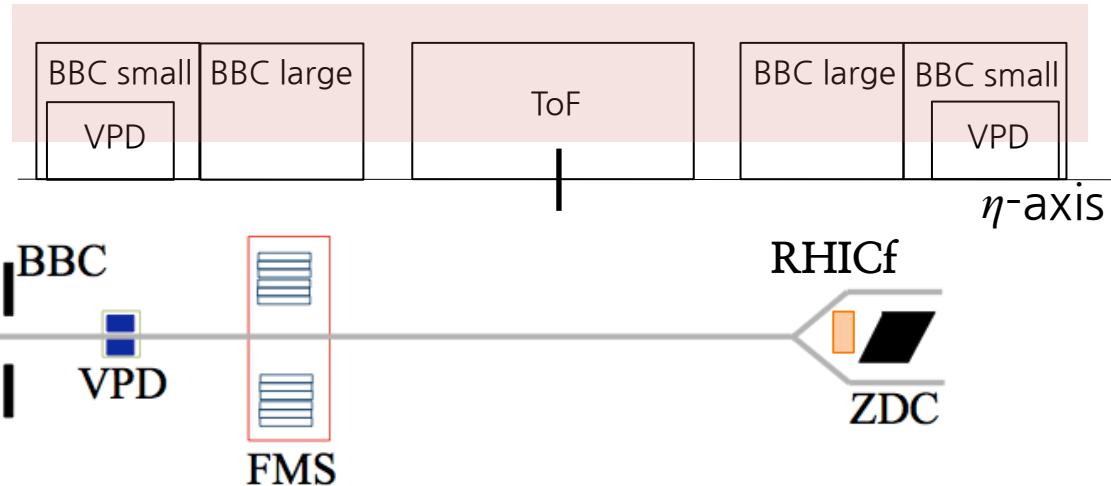


RHICf-STAR combined analysis

Central detectors



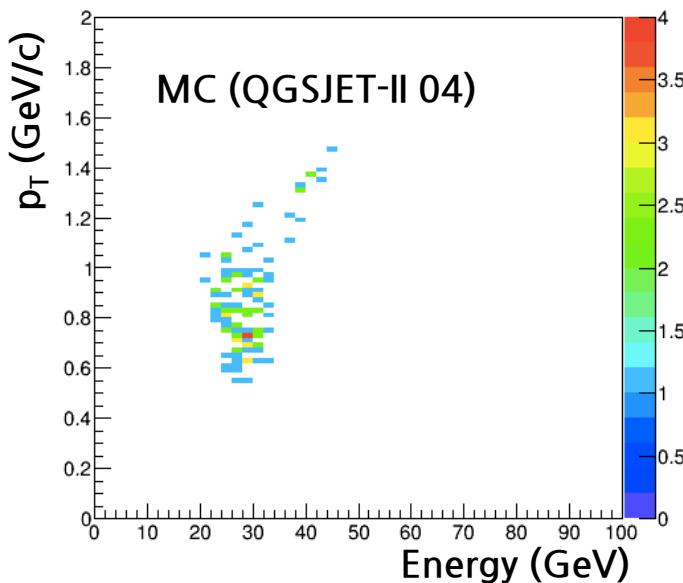
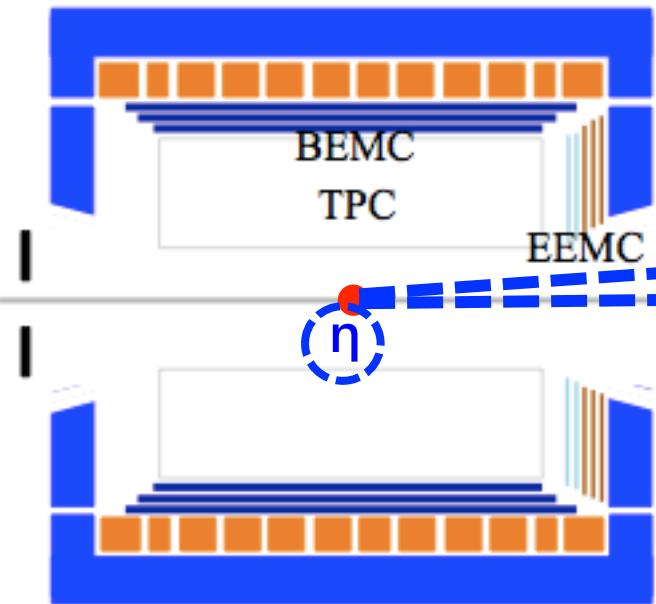
Non diffractive event



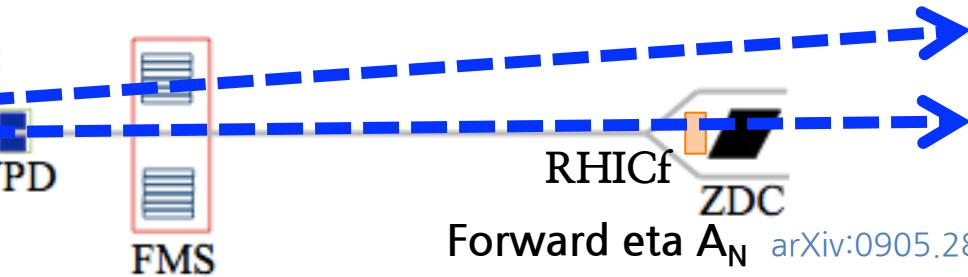
- Using STAR ToF, BBC, and VPD, we can study the detector correlation or event type dependence for the very forward $\pi^0 A_N$.
- For example, there should be signals in the TOF, BBC, and VPD if a π^0 comes non-diffractive event.
- Roman pot is also crucial for studying the event type dependence.

RHICf-STAR combined analysis

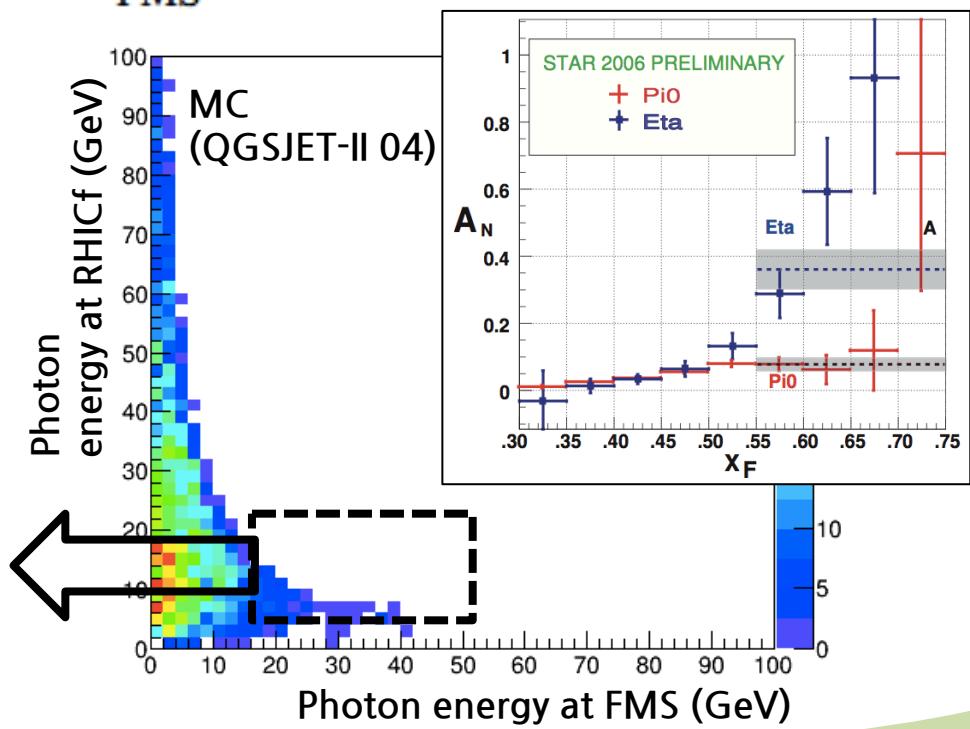
STAR experiment



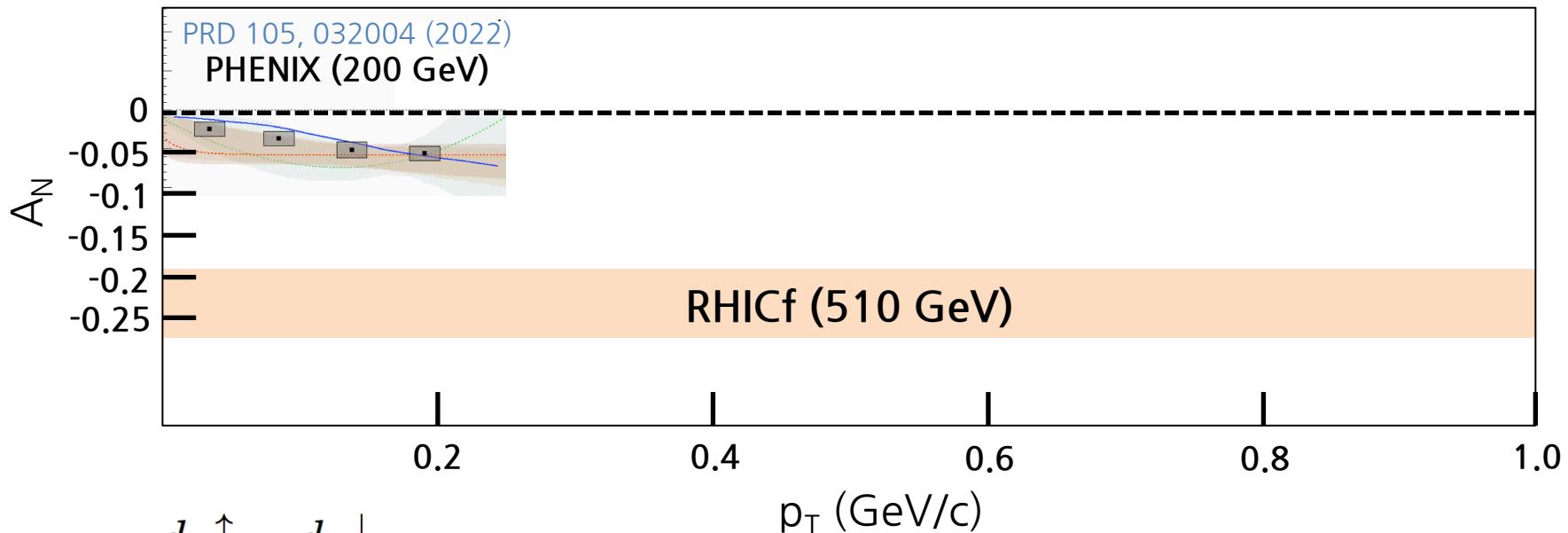
- We can possibly study the eta meson in a kinematic range where both diffractive and diffractive interactions are available.



Forward eta A_N [arXiv:0905.2840v1](https://arxiv.org/abs/0905.2840v1)



2. A_N measurement (neutron)



$$\begin{aligned}
 A_N &= \frac{d\sigma^\uparrow - d\sigma^\downarrow}{d\sigma^\uparrow + d\sigma^\downarrow} \\
 &= \frac{\sum_X |\langle cX|T|\uparrow\rangle|^2 - \sum_X |\langle cX|T|\downarrow\rangle|^2}{\sum_X |\langle cX|T|\uparrow\rangle|^2 + \sum_X |\langle cX|T|\downarrow\rangle|^2} \\
 &= \frac{-2\text{Im} \sum_X \langle cX|T|-\rangle\langle +|T^\dagger|cX\rangle}{\sum_X |\langle cX|T|+\rangle|^2 + \sum_X |\langle cX|T|-\rangle|^2}
 \end{aligned}$$

π exchange: **spin flip**

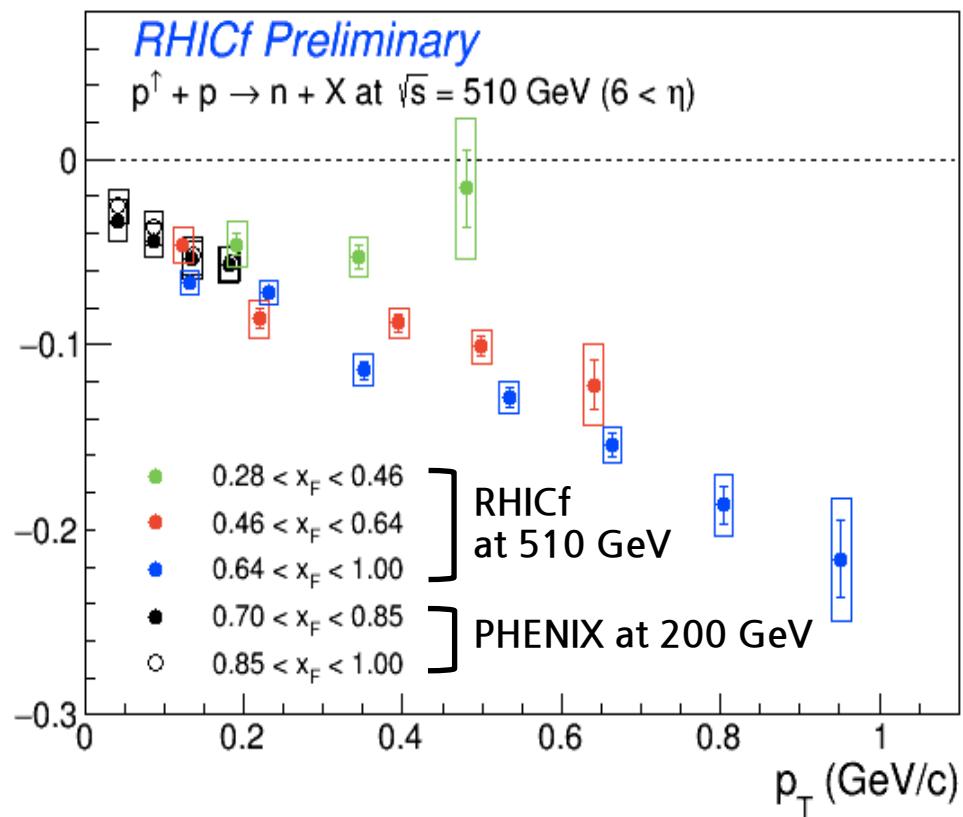
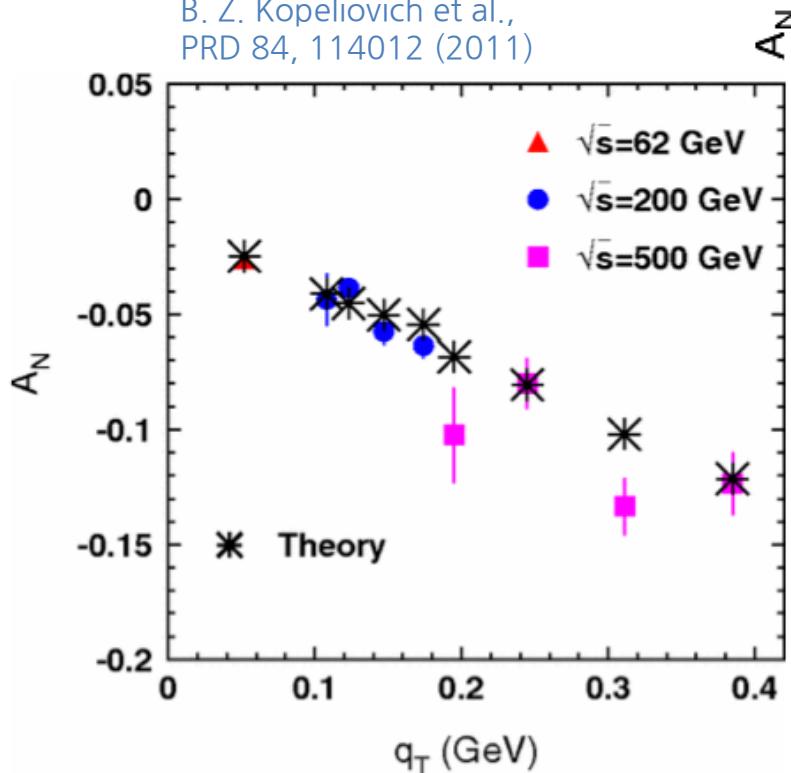
a_1 exchange: **spin non-flip**

- Neutron A_N can be explained by an interference between spin flip and non-flip amplitudes.
- RHICf data can test the π and a_1 exchange framework in a wide p_T region.

New preliminary result!

Neutron A_N

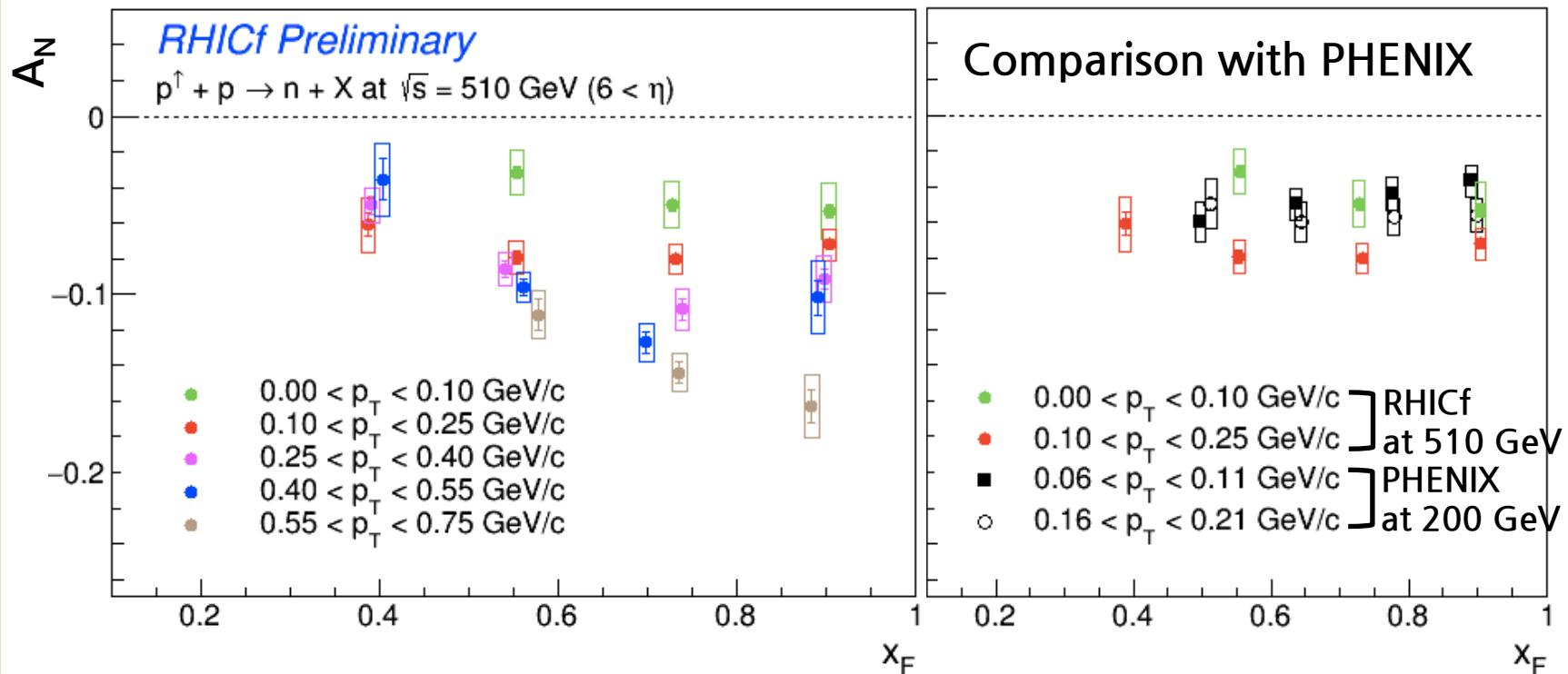
B. Z. Kopeliovich et al.,
PRD 84, 114012 (2011)



- π and a_1 exchange model predicts that the A_N increases in magnitude with p_T without x_F dependence.
- In higher x_F , the A_N increases in magnitude with p_T up to 1 GeV/c.
- There seems a x_F dependence especially in the higher p_T region.

New preliminary result!

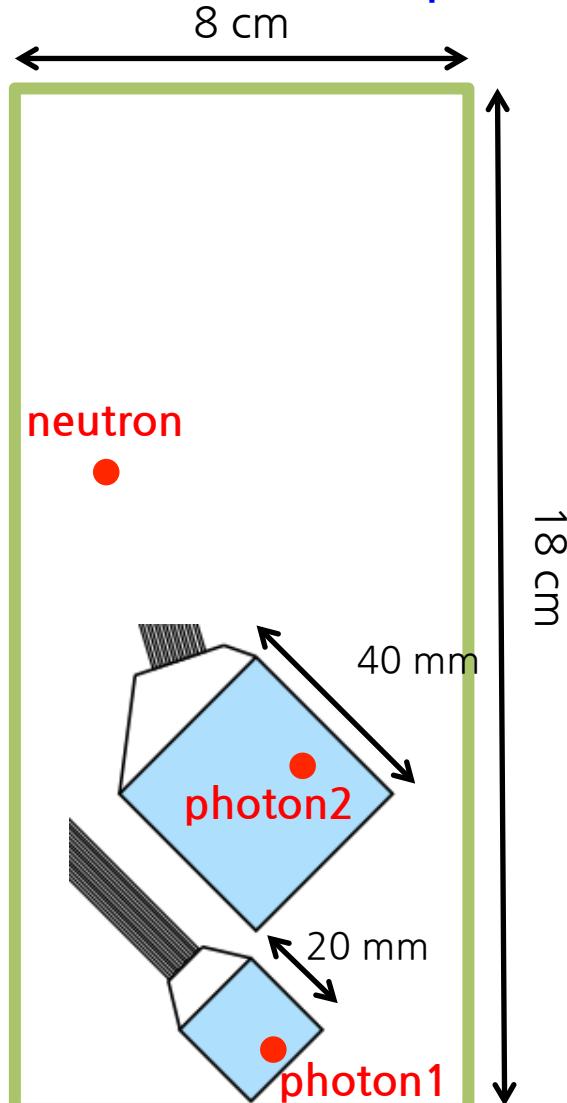
Neutron A_N



- In the lower p_T region, the A_N s are flat showing no x_F dependence.
- In the higher p_T region, it seems that there is a x_F dependence.
- Whether the x_F dependence is true or not will become clear with more precise background estimation.

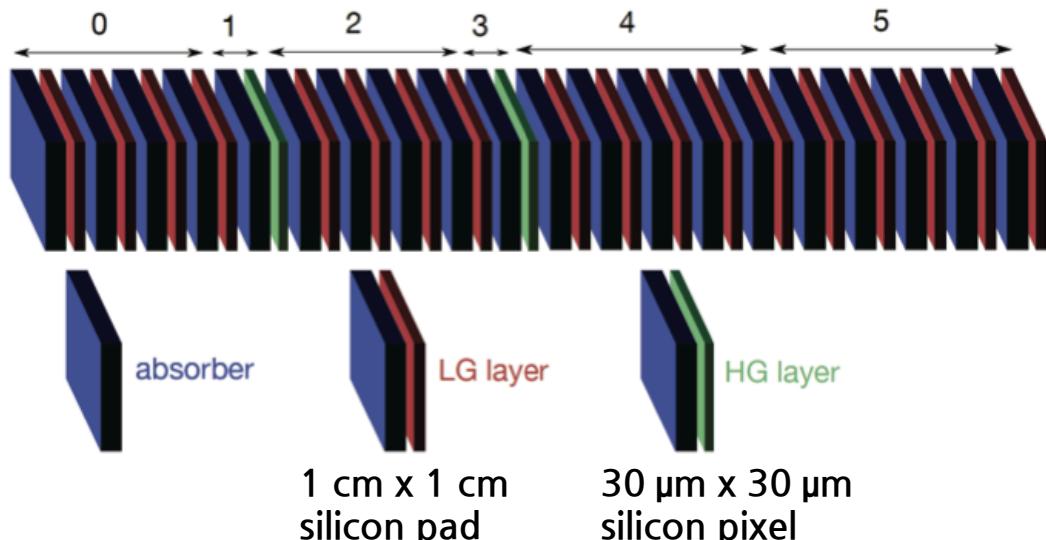
RHICf-II experiment

RHICf-II detector acceptance



ALICE FoCal-E

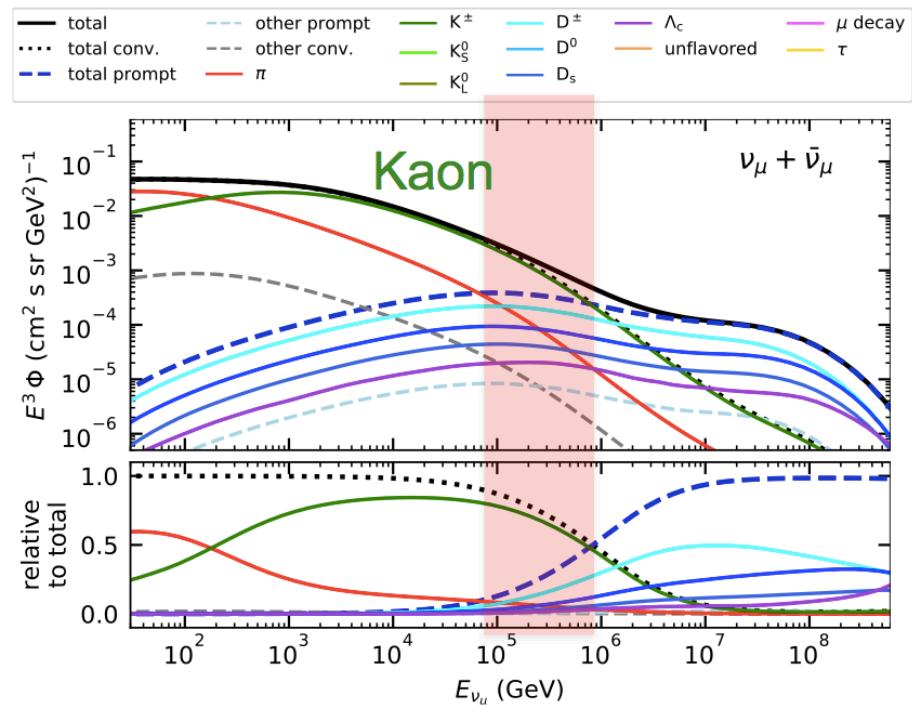
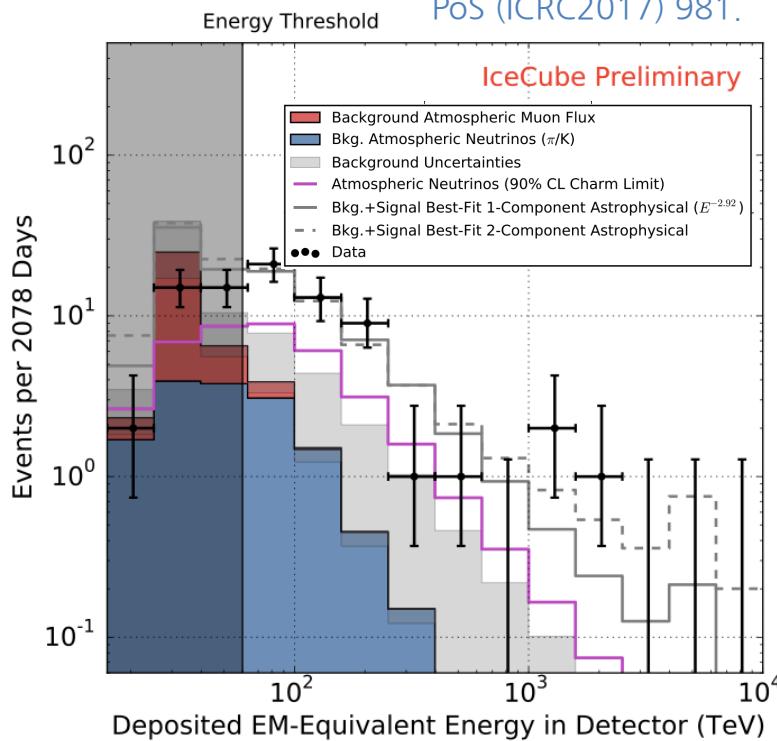
Longitudinal segmentation



- RHICf-II experiment is to measure more various particles with larger detector.
- Technology of the ALICE FoCal-E is transferred to the RHICf-II detector.
- Unfortunately, a possible RHICf-II operation schedule at STAR was canceled due to lack of human resources.

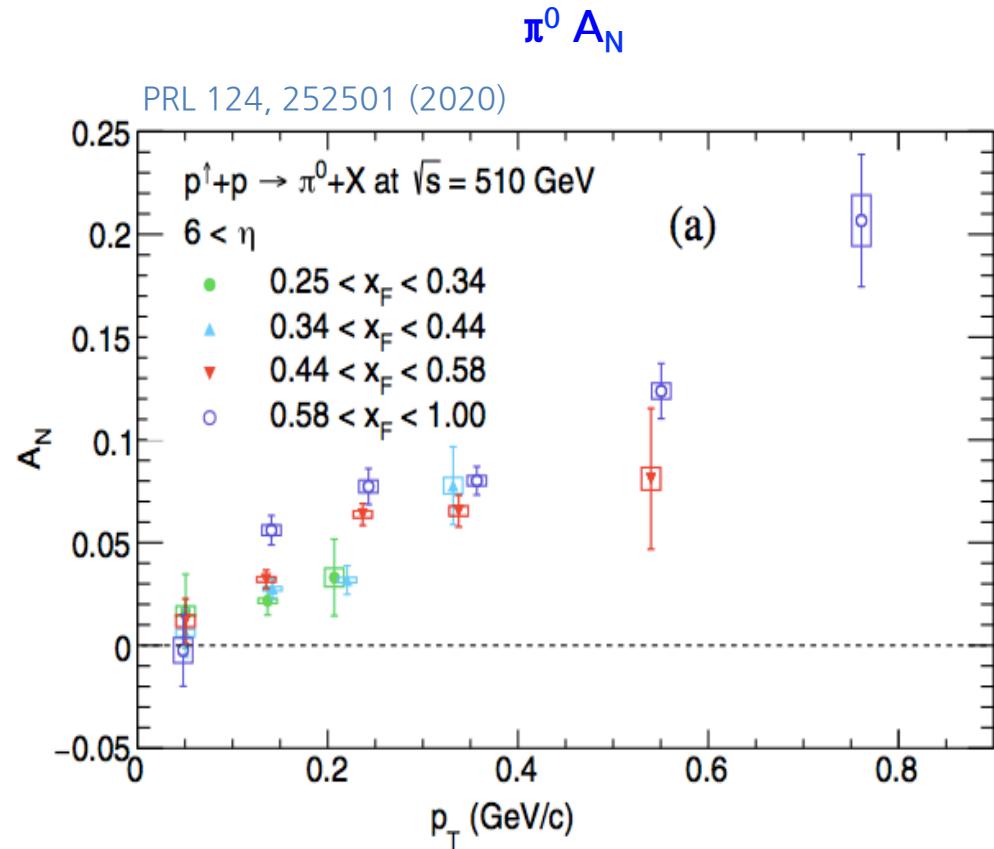
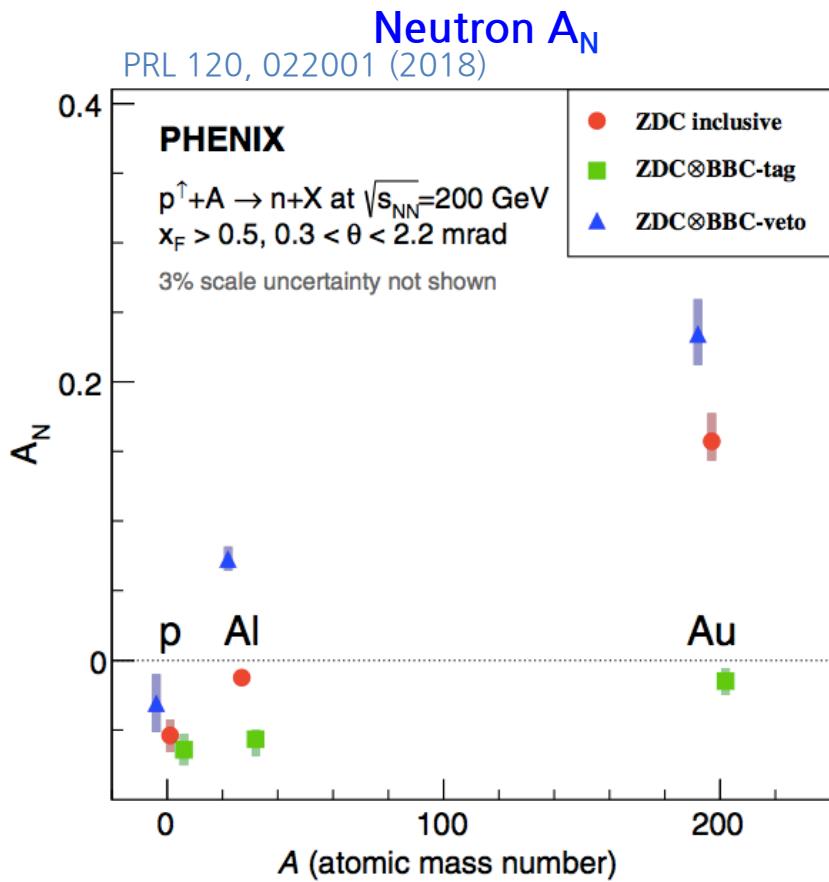
Kaon study at RHICf-II

PoS (ICRC2017) 981.



- Astrophysical neutrinos can enable us to understand the acceleration mechanism and source of the UHECRs.
- Atmospheric neutrino study is important to determine the background for the astrophysical neutrino search.
- One of the main sources for the atmospheric neutrino flux is kaon.

A_N study at RHICf-II



- With pA collisions, we can study the A dependence of the neutron A_N in more detail and also measure that of π^0 .
- A_N of Λ and K_S^0 can be measured to broaden our knowledge for the particle production mechanism and a possible origin of the RHICf π^0 result.

Summary

- In June 2017, the RHICf experiment measured the cross section and the A_N for very forward particles.
- Within the uncertainties, the RHICf photon cross section at $\sqrt{s} = 510$ GeV showed a x_F scaling with that of LHCf at $\sqrt{s} = 7$ TeV.
- There seems to be a x_F dependence in the neutron A_N especially in the higher p_T region.
- RHICf-STAR detector correlation will be studied to better understand the A_N for very forward π^0 (and neutron) production.
- A possible RHICf-II schedule was canceled by the lack of human resources.