Cone-size dependence of jet suppression from LHC to RHC

Daniel Pablos - INFN Torino



This project has received funding from the European Union's Horizon 2020 research and innovation programme under the Marie Sklodowska-Curie grant agreement n. 754496.

RBRC Workshop: Predictions for sPHENIX 20th July 2022





Istituto Nazionale di Fisica Nucleare





 η_1

 p_T

Daniel Pablos

Jets in pp

η_2 p_T



2



Collimated structure enforced through collinear divergences & color coherence.

Daniel Pablos

Jets in pp

Parton density evolution described via DGLAP: $t\frac{\partial}{\partial t}f(x,t) = \int \frac{dz}{z} \frac{\alpha_s}{2\pi} P_+(z) f\left(\frac{x}{z},t\right)$





Collimated structure enforced through collinear divergences & color coherence.

Daniel Pablos

Jets in pp

Parton density evolution described via DGLAP: $t\frac{\partial}{\partial t}f(x,t) = \int \frac{dz}{z} \frac{\alpha_s}{2\pi} P_+(z) f\left(\frac{x}{z},t\right)$ η_2 p_T

 Jets are defined with clustering algorithm, reconstruction radius R.









Collimated structure enforced through collinear divergences & color coherence.

-> Degree of jet activity determines, e.g., out-of-cone radiation (causes dijet asymmetry in pp). $p_{T,1} > p_{T,2}$

Jets in pp

 Jets are defined with clustering algorithm, reconstruction radius R.











Lesson example: total energy loss proportional to jet activity (more energy loss sources): -> Causes dijet asymmetry in AA (for same traversed length). Milhano & Zapp - EPJ '16 Causes selection bias towards narrower jets. Hadron vs. Jet Suppression Casalderrey, Hulcher, Milhano, DP, Rajagopal - PRC '19

Daniel Pablos

Crucial insights obtained with jet quenching MCs.









Jet suppression beyond MC:

How to describe energy loss off an extended, multi-partonic object?



Outline

Daniel Pablos

Jets in AA





Jet suppression beyond MC:

How to describe energy loss off an extended, multi-partonic object?



Outline

Daniel Pablos

Jets in AA

Non-perturbative modelling of long wavelength jet modes: Where does "lost" energy go to?





Barata, Mehtar-Tani - JHEP '20







Barata, Mehtar-Tani - JHEP '20







Barata, Mehtar-Tani - JHEP '20







Barata, Mehtar-Tani - JHEP '20



Daniel Pablos





Out of Cone Radiation

Only emissions that end up out of the cone R should be accounted for:

Multiplicative
$$\omega \frac{\mathrm{d}I_{>}}{\mathrm{d}\omega} = \int_{(\omega R)^2}^{\infty} \mathrm{d}k_{\perp}^2 \,\omega \frac{\mathrm{d}I}{\mathrm{d}\omega \,\mathrm{d}k_{\perp}^2} \simeq B\left(\omega R; Q_{\mathrm{broad}}^2\right) \times \omega \frac{\mathrm{d}I}{\mathrm{d}\omega}$$

Ansatz:

 $B(\omega R; Q_{\rm bro}^2)$ Mehtar-Tani, DP, Tywoniuk - PRL '21

$$\mathcal{P}(\boldsymbol{k}) \simeq egin{cases} rac{4\pi}{Q_s^2} \mathrm{e}^{-\boldsymbol{k}^2/Q_s^2} & k_\perp^2 \ll Q_{\mathrm{med}}^2 \ rac{4\pi Q_s^2}{\boldsymbol{k}^4} & k_\perp^2 \gg Q_{\mathrm{med}}^2 \end{cases}$$

Use Molière expansion around multiple soft scatterings (a.k.a. IOE). Barata et al. - PRD '21

Can be improved with fully differential spectrum. Barata et al. - JHEP '21

$$\begin{array}{rcl} \text{oad} \end{array} & = & \frac{Q_{\text{broad}}^2}{4\pi} \int_y^\infty \mathrm{d}x \, \mathcal{P}(x) & y &= (\omega R)^2 / Q_{\text{broad}}^2 \\ \end{array}$$

$$\frac{\partial}{\partial L} \mathcal{P}(\boldsymbol{k}, L) = C_R \int_{\boldsymbol{q}} \gamma(\boldsymbol{q}) \left[\mathcal{P}(\boldsymbol{k} - \boldsymbol{q}, L) - \mathcal{P}(\boldsymbol{k}) \right]$$



Bare Quenching Factor

Baier, Dokshitzer, Mueller - JHEP '01 For steeply falling spectrum and small energy loss: Salgado, Wiedemann - PRD '03 $\sim \sim$

$$rac{\mathrm{d}\sigma_{\mathrm{med}}}{\mathrm{d}p_T} = \int_0^\infty \mathrm{d}\epsilon \, \mathcal{P}(\epsilon) \left. rac{\mathrm{d}\sigma_{\mathrm{vac}}}{\mathrm{d}p_T'}
ight|_{p_T'=p}$$

Mehtar-Tani, DP, Tywoniuk - PRL '21

Quenching factor of a single parton for multiple independent emissions (R dependent):

$$Q_{\rm rad}^{(0)}(p_T) = \exp\left[-\int_{\omega_s}^{\infty} d\omega \, \frac{dI_{>}}{d\omega} \left(1 - e^{-\nu\omega}\right) - \int_{T}^{\omega_s} d\omega \, \frac{dI^{(0)}}{d\omega} \left(1 - e^{-\nu\omega(1 - \left(\frac{R}{R_{\rm rec}}\right)^2)}\right) \right]$$
$$\nu \equiv \frac{n}{p_T}$$
Full out-of-cone spectrum
$$\omega_s \equiv (g_{\rm med}^2 N_c / (2\pi)^2)^2 \pi \hat{q}_0 L^2$$

for semi-hard emissions

Note that:

$$\Delta E = \left(1 - \left(\frac{R}{R_{\rm rec}}\right)^2\right) \int_T^{w_s} dw \, w \frac{dI^{(0)}}{dw} = -\frac{d}{d\nu} Q_{\rm rad}^{(0),\rm turb}(p_T)$$

Daniel Pablos

$$_{p_T+\epsilon} \approx \frac{\mathrm{d}\sigma_{\mathrm{vac}}}{\mathrm{d}p_T} \underbrace{\int_0^\infty \mathrm{d}\epsilon \,\mathcal{P}(\epsilon) \mathrm{e}^{-\epsilon \frac{n}{p_T}}}_{Q(p_T)}$$

 \rightarrow O(1) emission probability

undergo turbulent cascade, thermalise

if uniformly distributed in jet hemisphere $R_{\rm rec} \sim \pi$

INFN Torino

 $\nu = 0$





Quenched Phase Space of a Jet



Daniel Pablos

Only those jet modes that:

are formed inside the medium, and,

$$t_f < L$$

are resolved by the medium,

 $t_f < t_d$

contribute to double-logarithmic enhancement of quenched phase space:

$$\mathbf{PS}_{\text{in}} = \bar{\alpha} \int_{t_{\text{f}} < t_{\text{d}} < L} \frac{\mathrm{d}\theta}{\theta} \int \frac{\mathrm{d}z}{z} \equiv \bar{\alpha} \ln \frac{R}{\theta_c} \left(\ln \frac{p_T}{\omega_c} + \frac{2}{3} \right)$$

Mehtar-Tani, Tywoniuk - PRD '18 see also Caucal, Iancu, Mueller, Soyez - PRL '18







Jet Suppression: Framework

Use microjet distributions derived using Generating Functional (GF) framework:

Vacuum evol. obeys DGLAP:

$$\frac{df_{j/i}^{\text{incl}}(z,t)}{dt} = \sum_{k} \int_{z}^{1} \frac{dz'}{z'} P_{jk}(z') f_{k/i}^{\text{incl}}(z/z',t)$$
Dasgupta et al. - JHEP "

Extend GF in the medium to resum energy loss effects due to multi-particle nature of jet:

$$egin{aligned} & rac{\partial Q_i(p, heta)}{\partial \ln heta} = \int_0^1 \mathrm{d}z \, rac{lpha_s(k_\perp)}{2\pi} p_{ji}^{(k)}(z) \Theta_{\mathrm{res}}(z) \ & imes [Q_j(zp, heta)Q_k((1-z)p, heta)-] \end{aligned}$$

Energy loss versus R displays non-monotonic behaviour. Competing effects: Increasing R means larger quenched phase space: more quenching.

Daniel Pablos



- Increasing R means more likely to retain emitted (or thermalised) quanta: less quenching.

Mehtar-Tani, DP, Tywoniuk - PRL '21





Jet Suppression at LHC

17



Daniel Pablos

Mehtar-Tani, DP, Tywoniuk - PRL '21

- Use PYTHIA8 to generate spectrum at initial angle $R_0=1$ (with nuclear PDFs EPS09 LO in medium case)
- Evolve microjets using DGLAP down to jet R.
- Compute resummed quenching factors for each jet p_T and R:
 - ->> Bare quenching factor requires knowledge of event-by-event, centrality dependent QGP properties:
 - Embed framework into realistic heavy-ion environment:
 - Glauber sampling, random azimuthal orientation.
 - Compute event-by-event relevant quantities, e.g.: (in local fluid rest frame)

$$= \int_{\Gamma(t)} dx_F \qquad \hat{q}_0 \propto \frac{1}{L} \int_{\Gamma(t)} dx_F T^3(x) \left(\frac{p \cdot u(x)}{p^0}\right)$$

Path of jet through hydro. profile (VISHNU) down to T_c



Jet Suppression at LHC



Daniel Pablos

Mehtar-Tani, DP, Tywoniuk - PRL '21

Only two unconstrained parameters:

- $\rightarrow R_{\rm rec}$ varied between $\begin{array}{l} R_{\rm rec} = \pi/2 & ({\rm isotropic}) \\ R_{\rm rec} = (5/6) \pi/2 & ({\rm wake inspired}) \end{array}$ Casalderrey, Milhano, DP, Rajagopal, Yao - JHEP '21 $\rightarrow g_{med}$ fit to ATLAS R=0.4 around p_T~120 GeV at 0-10% • $g_{\text{med}} \in \{2.2, 2.3\}$ $\langle \hat{q}_0 \rangle \simeq 0.41 \text{ GeV}^2/\text{fm}$ in 0-10%
 - $\hat{q} = 2.46 \text{ GeV}^2/\text{fm}$ $\omega_c \approx 65 \text{ GeV}$ due to logarithmic corrections.

Good description of both centrality and jet p_T suppression.







R Dependence & Modelling Uncertainties



Mild R dependence, in agreement with CMS data.

Mehtar-Tani, DP, Tywoniuk - PRL '21

Daniel Pablos

Modelling sensitivity at p_T=110 GeV for **R** between **0.2 and 0.6**:

Parameter	Variation	Effect
θ_c	$[heta_c/2, 2 heta_c]$	$ \lesssim 20\%$
IOE	LO/NLO	$\sim 2\%$
n	± 1	$\sim 10\%$
$R_{ m rec}$	$[1, \infty]$	$\lesssim 10\%$
ω_s	$[\omega_s/2, 2\omega_s]$	$ \lesssim 8\%$

- > NLO contribution very small (hard emissions tend to be collinear).
- Modelling of fate of lost energy relatively small.
- Determination of quenched phase space relatively large. Improvable in pQCD.

Need to improve perturbative sector before non-perturbative becomes relevant (for R<0.6!)



RHIC vs LHC Vacuum Spectra



Steeper spectrum at RHIC energies, will push R_{AA} down.

Daniel Pablos

Jet p_T [GeV]

Spectra increases with increasing R due to recapturing out-of-cone radiation.





RHIC vs LHC Vacuum Quark Fraction



Larger quark-initiated jet fraction at RHIC, should push total R_{AA} up.

Daniel Pablos

Quark-initiated jet fraction decreases with increasing R, as gluon-initiated jets are more active.

	_	
	-	
	-	
	-	
	-	
)0	1000	

RHIC vs LHC Quenching Factor

RHIC

60



 $\langle \hat{q}_0 \rangle^{\text{RHIC}} \simeq 0.25 \,\text{GeV}^2/\text{fm}$ $\langle \hat{q} \rangle^{\text{RHIC}} \simeq 1.22 \,\text{GeV}^2/\text{fm}$

 $\langle L \rangle^{\rm RHIC} \simeq 4.5 \, {\rm fm}$

Similar quenching factors between LHC and RHIC (@ $\sim p_T$) when considering medium properties.

Resummation more significant at LHC due to larger phase space.

 $\langle \hat{q}_0 \rangle^{\rm LHC} \simeq 0.44 \, {\rm GeV}^2 / {\rm fm}$ $\langle \hat{q} \rangle^{\text{LHC}} \simeq 2.34 \,\text{GeV}^2/\text{fm}$

$$\langle L \rangle^{\rm LHC} \simeq 5.6 \, {\rm fm}$$

INFN Torino



LHC

800 900 1000





Daniel Pablos

Effect of Nuclear PDF

Jet RAA at RHIC

5 - 10%

30-40%

35

Jet p_T [GeV]



Daniel Pablos





Jet v₂ at RHIC

 Interesting grouping in v₂ for different R.

R=0.3, and especially
 R=0.4, migrate as a
 function of centrality.



Daniel Pablos



Coherence vs. Centrality



Daniel Pablos

Different critical angle dists. for different lengths,

i.e. different centralities.

5–10% 40-50% 60-70%



Jet RAA at RHIC - No Coherence

 $\theta_c \to 0$

 Remove effects of coherence.

Larger R tend to be
 more suppressed
 than w/ coherence.



Daniel Pablos

27

INFN Torino

Jet v₂ at RHIC - No Coherence

 $\theta_c \to 0$ Less grouping, especially for smaller R.

No migration of
 R=0.3 or R=0.4.



Daniel Pablos

Jet v₂ & Coherence Effects

Effect up to differences in v₂ of 0.015.

Jet v₂ & Coherence Effects

sPH-TRG-2022-001

Effect up to differences in v₂ of 0.015.

30

At strong coupling:

Modification of stress-energy tensor due to supersonic quark contains sound and diffusive modes.

Effective source for hydro corresponds to drag force on the quark.

Agreement between hydrodynamics & wake of a quark even for small distances $\sim 1/T$.

> Fulfils Energy-Momentum Conservation in the Jet+Plasma Interplay.

The Wake of a Quark

Chesler & Yaffe - PRD '07

Estimation of the Hadrons from the Wake

Assuming:

->> small perturbations on top of Bjorken flow.

-> perturbation stays localised near jet's rapidity.

Expand Cooper-Frye spectrum to first order in perturbations:

Velocity pert.

Daniel Pablos

$$= {s \, \tau \over c_s^2} \int d\eta \, d^2 x_\perp \, {\delta T \over T}$$

Temperature pert.

32

Fully constrained by energy-momentum conservation.

Computationally efficient.

Neglects important effects from local flow.

The Diffusion Wake

33

No wake: Increasing R leads to more energy loss.

QGP ridge only: Partial recovery of energy yields mild R dependence.

Daniel Pablos

w/ QGP trough also: extra suppression due to diffusion wake.

The Effect of the Recoiling Jet

Daniel Pablos

ading	<pr>> density of wake hadrons</pr> w.r.t leading jet axis.
Ω.	Aligned in rapidity
20 20 20	Subleading jet's QGP trough hits leading jet.
.0	Separated in rapidity
-10	Subleading jet's QGP trough misses leading jet.

 $p_T^L > 250 \text{ GeV}$ $p_T^S > 80 \text{ GeV}$ $\Delta \phi_D > 2\pi/3$

35

differential in $|\eta_D| \equiv |\eta_L - \eta_S|$

Jet RAA at SPHENIX

Analytic, but over-simplified medium response needs to be improved: ->> Starting point: linearised hydro eqs. for perturbations on top of viscous Bjorken flow.

Daniel Pablos

Linearized Hydrodynamics

• $\partial \epsilon$: energy density pert.

Wavefront structure (Mach cone) diffuses due to viscosity.

• g_x : momentum pert.

Wing shaped structure diffuses due to viscosity (diffusion wake).

Casalderrey, Milhano, DP, Rajagopal, Yao - JHEP '21 **INFN** Torino

Sound modes make wake energy spread in rapidity with time.

Jet breaks long. boost invariance.

Casalderrey, Milhano, DP, Rajagopal, Yao - JHEP '21 Daniel Pablos

Linearized Hydrodynamics

Vortex ring around jet direction.

 \rightarrow Imprints on Λ polarisation? Serenone et al. - PLB '21

Efficient Computations of the Wake

Casalderrey, Milhano, DP, Rajagopal, Yao - in preparation

(a) Energy-momentum deposition for $E_i = 10$ GeV (filled arrow) and $E_i = 50$ GeV (empty arrow).

Daniel Pablos

Hadrons from the wake depend on evolution time, local flow... thousands of possibilities.

> 3+1D **Viscous Hydrodynamics** ~ 2 hours

Superposition of **Bjorken flow solutions** Trans., Rot., & Boosts ~ 3 seconds

Efficient Computations of the Wake

40

Daniel Pablos

Average distribution (~ 50 events) becomes *harder* and more *collimated* than without transverse flow.

Conclusions & Outlook

- Jet energy loss requires accounting for several effects:
 Resummation of multi-particle energy loss inside R from high-virtuality shower.
 Determination of the resolved phase-space. Improvable in pQCD.
 Modelling of hydrodynamized component. Important for larger R and smaller p_T.
- Jet azimuthal anisotropy especially sensitive to coherence physics:
 Effect measurable with sPHENIX.
- Higher p_T jets at RHIC kinematics with moderate R are sensitive to diffusion wake of recoiling jet:
 - Effect measurable with sPHENIX.
- Computationally efficient jet-by-jet, flow-dependent wakes:
 - Insights into devising tailored observables to reveal wake signatures.
 - Useful for massive medium response Monte Carlo simulations.
 - Can be supplemented as non-perturbative component to analytical computations.

Daniel Pablos

Thanks for your attention!

Further Improvements on Single Charge Energy Loss

- All order resummation of medium induced radiation spectrum.
- Resummed Opacity Expansion (ROE) to cover Bethe-Heitler regime.

In-medium fragmentation of hard parton in QGP through effective kinetic theory.

 \rightarrow Includes 1 \rightarrow 2 and 2 \rightarrow 2 processes.

-----> Features turbulent cascade, modified chemistry around the jet.

Detailed analysis of dynamics, can account for medium response.

Resummed Quenching Factor

Daniel Pablos

Bare quenching factors (dashed):

less quenching for larger R.

-> Easier to keep (recover) the emitted (thermalised) modes.

Resummed quenching factors (solid):

Iarger R can lead to more quenching.

Interplay between energy recovery and size of quenched phase space.

Mehtar-Tani, DP, Tywoniuk - PRL '21

More Data Comparison

Good description of ALICE at lower p_T.

Absolute value shows some tension with CMS, originates from CMS/ATLAS tension.

Accounting for rapidity cuts

W/DGLAP: fit pp spectrum at $R_0=1$, evolve down to R. w/o DGLAP: fit pp spectrum directly at $\eta = 0.9 - R$.

Quark fraction depends on:

Rapidity, via PDF.

----> Cone R, via microjet evolution.

Daniel Pablos

- Redoing calculation with ALICE rapidity cuts makes no difference.
- Quark fraction change small in ratio.
- -> Effect of spectral index change negligible.

Jet RAA at LHC

Daniel Pablos

Jet v₂ at LHC

Daniel Pablos

Jet suppression increases with increasing R.

Daniel Pablos

- Include non-eq. contribution only, i.e. jet particles that did not hydrodynamize:

 - Wider jets "lose" more energy, more energy loss sources.

Include both non-eq. and QGP "ridge" contributions:

- Energy is progressively recovered with increasing R.
- nPDF effect sets an upper limit on R_{AA} at very high p_T .

Daniel Pablos

Include non-eq., QGP "ridge" and QGP trough contribution:

- QGP trough amounts to jet suppression; over-subtraction effect.
- Effect increases with increasing R.

Daniel Pablos

Competition of effects that yield, overall, a very mild evolution from small to large R.

Leading Jet Suppression vs. Ind

DP - PRL '20

A new observable.

R = 0.4

leading jet area easy to miss; small effect from QGP trough.

R = 1.0

strong dependence on $|\eta_D|$; knee visible when $\eta_D \sim R$.

$$p_T^L > 250 \text{ GeV}$$

 $p_T^S > 80~{
m GeV}$ $\Delta \phi_D > 2\pi/3$

differential in $|\eta_D| \equiv |\eta_L - \eta_S|$

Hybrid Strong/Weak Coupling Model

States States

Strongly coupled energy loss (hydrodynamization rate)

PYTHIA8 down to hadro. scale (formation time argument for spacetime picture)

Daniel Pablos

$$E\frac{d\Delta N}{d^3p} = \frac{1}{32\pi} \frac{m_T}{T^5} \cosh(y - y_j) \exp\left[-\frac{m_T}{T} \cosh(y - y_j)\right] \exp\left[-\frac{m_T}{T} \cosh(y - y_j) + \frac{1}{3}m_T \Delta M_T \cosh(y - y_j)\right]$$

Hadrons from the hydro. wake (medium response)

Casalderrey-Solana, Gulhan, Milhano, DP, Rajagopal JHEP '15, '16, '17

Modified spectra after Cooper-Frye:

$$\frac{\mathrm{d}\Delta N}{p_T \,\mathrm{d}p_T \,\mathrm{d}\phi \,\mathrm{d}y} = \frac{1}{(2\pi)^3} \int \tau \,\mathrm{d}x \,\mathrm{d}y \,\mathrm{d}\eta_s \,m_T \cosh(\mathbf{y} - \eta_s) \Big[f\Big(\frac{u^{\mu}p_{\mu}}{T_f + \delta T}\Big) - f\Big(\frac{u^{\mu}_0 p_{\mu}}{T_f}\Big) \Big]$$

With wake

Ansatz to account for radial flow:

 \rightarrow Boost particle momentum to frame in which fluid cell moves with $u^{\mu} = \gamma_{\perp}(\cosh \eta, v_x, v_y, \sinh \eta)$.

$$p^{\mu}_{
m cell} = \Lambda^{\mu}_{\
u}(oldsymbol{v}_{
m cell})p^{
u}_{
m lab}$$

Use flow profile at freeze out from 2+1D hydro. simulations such as VISHNU.

Accounting for Radial Flow

Casalderrey, Milhano, DP, Rajagopal, Yao - JHEP '21

