# Resolving the Scales of the Quark-Gluon Plasma with Energy Correlators



Carlota Andres, Fabio Dominguez, Raghav Kunnawalkam Elayavalli, JH, Cyrille Marquet, and Ian Moult





21/07/2022

#### Outline

- 1. Introduction
- 2. Correlation functions of  $\mathcal{E}(\vec{n})$
- 3. The observable analytically
- 4. The observable numerically

#### Part 1: Introduction



Time

We want to study the QGP in HIC.

- 20 years of HIC at RHIC, 10 years of HIC at the LHC, sPHENIX coming soon.

We need to study the QGP with sensitive QCD probes with good theoretical control...

21/07/2022



21/07/2022



Problem:

Jet quenching is a multi-scale process. It is difficult to unambiguously resolve the scales/properties of the QGP involved within current approaches.



There has been a lot of work introducing observables.

1512.08107, 1710.03237, 1812.05111, 2010.00028, and more

We would like to present a new approach to add to this body of work.



Polchinski: There is a lot of QCD data... can they see this scaling?

Feb 2009

Maldacena: People do not do this, I haven't figured out why they don't. I think they just haven't thought about this. Can you resolve separate **jets** well enough to study the small angles?

> Well, this is the point - here you **don't have to talk about jets!**

in the

TE PINTE

-----

#### Recap

•

- Correlation functions in statistics:
  - $\operatorname{Corr}_2(X, Y) = \langle XY \rangle \langle X \rangle \langle Y \rangle$  (also just the covariance)
  - $\operatorname{Corr}_{3}(X, Y, Z) = \langle XYZ \rangle \langle X \rangle (\langle Y \rangle \langle Z \rangle \operatorname{Corr}_{2}(Y, Z))$
- In physics we usually refer to  $\langle X_1 \dots X_n \rangle$  as an *n* point correlator. This is just conventional and has origins in that often  $\langle X_i \rangle = 0$ .
- QFT correlators (propogators) relate back to these statistical correlators through the path integral and statistical mechanics...

- Generally one can define correlators of any quantum charge or conserved quantity.
- For QCD, correlators of energy flux are usually of most interest these naturally remove soft physics without grooming.

$$\mathcal{E}(\vec{n}) = \lim_{r \to \infty} \int_{0}^{\infty} dt \ r^2 n^i T_{0i}(t, r\vec{n})$$
  
$$\mathcal{E}(\vec{n}) \simeq \int_{0}^{\infty} dt \ \mathcal{E}_{\text{flux through AD}}$$

 $\left( \right)$ 

 $\bigwedge$ 

$$\mathcal{E}(\vec{n}) = \lim_{r \to \infty} \int dt \ r^2 n^i T_{0i}(t, r\vec{n}) + \int_{r \to \infty} \int dt \ r^2 n^i T_{0i}(t, r\vec{n}) + \int_{r \to \infty} \int dt \ r^2 n^i T_{0i}(t, r\vec{n}) + \int_{r \to \infty} \int dt \ r^2 n^i T_{0i}(t, r\vec{n}) + \int_{r \to \infty} \int dt \ r^2 n^i t +$$

$$\mathcal{E}(\vec{n}) = \lim_{r \to \infty} \int^{\infty} dt \ r^2 n^i T_{0i}(t, r\vec{n})$$

In fact, one can show that all collider observables can be expressed as a weighted sum over energy correlators:

$$\langle 0 \rangle = \sum_{i} C_{i}(0) \int d\vec{n}_{1,\dots,i} \langle \mathcal{E}(\vec{n}_{1}) \dots \mathcal{E}(\vec{n}_{i}) \rangle.$$
 2004.11381, 2205.06818

Perhaps not surprising when one thinks of a  $\mathcal{E}(\vec{n})$  as providing the idealised output of a calorimeter.

Also intuitively, higher point correlators are more differential and so provide more information on the process at hand.

 $\infty$ 

$$\begin{aligned} \mathcal{E}(\vec{n}) &= \lim_{r \to \infty} \int_{0}^{\infty} dt \ r^{2} n^{i} T_{0i}(t, r\vec{n}) \\ \frac{\langle \mathcal{E}^{n}(\vec{n}_{1}) \mathcal{E}^{n}(\vec{n}_{2}) \rangle}{Q^{2n}} &= \frac{1}{\sigma} \sum_{ij} \int \frac{\mathrm{d}\sigma_{ij}}{\mathrm{d}\vec{n}_{i} \mathrm{d}\vec{n}_{j}} \frac{E_{i}^{n} E_{j}^{n}}{Q^{2n}} \delta^{(2)}(\vec{n}_{i} - \vec{n}_{1}) \delta^{(2)}(\vec{n}_{j} - \vec{n}_{2}) \end{aligned}$$



Pros:

- Defined on inclusive cross-sections and can be made insentive to soft radiation. Textbook example of where pp CSS factorisation can be used without any violation.
  - $\frac{d\sigma}{d\zeta} = \int dE_J E_J^2 H(E_J) J_{\text{EEC}}(\zeta, E_J) + \text{power corrections}, \ \underline{2109.03665}$
- Well studied by CFT community. Powerfull techniques exist for calculations: light-ray OPE, celestial Blocks, lorentzian inversion. 2202.04085
- Including the only calculation of jet substructure at strong coupling. <u>0803.1467</u> Cons:
- Tend to be reliant on high stats.
- Not event-by-event so cannot be directly used to tag.

#### Case Study: vacuum jets

In summary: The small angle behavior of the energy correlation functions is determined by the spin j = 3 non-local operators that appear in the OPE

0803.1467

$$\langle \mathcal{E}(\theta_1)\mathcal{E}(\theta_2)\cdots \rangle \sim \sum_n |\theta_{12}|^{\tau_n-4} \langle \mathcal{U}_{3-1,n}(\theta_2)\cdots$$

 $\tau_n(j) = 2 - \gamma(j)$ , in the formula above j = 3

In a CFT  $\gamma(j)$  is a constant whilst in QCD the running coupling causes  $\gamma(j)$  to have logarithmic scale dependence but at LL the structure is otherwise unchanged.

An n-point correlator has a  $\tau(j = n + 1) = 2 - \gamma(j = n + 1)$  scaling

 $R_L \sim \frac{\Lambda_{\rm QCD}}{p_{T\,\rm jet}}$  breaks the OPE scaling of approx assympttically free dynamics.



Which correlation function is the one for us?

- In the previous slide the 2-point correlator gives a sentive pobe of hadronisation.
- In <u>2201.08393</u> the 3-point provided a sentive probe to the top mass.

Look to what is currently done and sucessful.

- *R<sub>AA</sub>* can be expressed as a function of one-point correlators + corrections:
  - $R_{AA} = \langle N_{AA} \rangle / (\langle N_{Coll} \rangle \langle N_{pp} \rangle)$ .  $\langle N \rangle$  is the one point correlator of the number opertator and due to momentum conservation  $\langle N \rangle \approx \langle \mathcal{E} \rangle / \langle Q \rangle$ .
- In effect, *R<sub>AA</sub>* gives access to the simplest but also least sensitive correlator. Let us increase the sentivity (at the expense of a little more complexity) by looking directly at the 2-point correlator.

#### Part 3: The observable analytically



#### The observable analytically

Equiv. to  $t_f < L$ 

Vacuum  $\theta \ll 1$  resummation

 $\theta > (EL)^{-1/2}$  Medium induced quenching

 $\frac{\mathrm{d}\Sigma^{(1)}}{\mathrm{d}\theta} \sim \frac{1}{\theta^{1-\gamma(3)}} + \mathcal{O}(\theta^0)$ 

$$\begin{aligned} \frac{\mathrm{d}\Sigma^{(n)}}{\mathrm{d}\theta} \Big|_{\theta \gtrsim \theta_L} &= \frac{1}{\sigma_{qg}} \int \mathrm{d}z \frac{\mathrm{d}\sigma_{qg}}{\mathrm{d}\theta \mathrm{d}z} z^n (1-z)^n \\ &\times \left( 1 + \mathcal{O}(\alpha_{\mathrm{s}} \ln \theta_L^{-1}) + \mathcal{O}\left(\alpha_{\mathrm{s}} \frac{\mu_{\mathrm{s}}^n}{E^n}\right) \right) \\ \frac{\mathrm{d}\sigma_{qg}}{\mathrm{d}\theta \mathrm{d}z} &= \frac{\mathrm{d}\sigma_{qg}^{\mathrm{vac}}}{\mathrm{d}\theta \mathrm{d}z} (1 + F_{\mathrm{med}}(z, \theta, \hat{q}, L)) \xrightarrow{1907.03653, 2107.02542} \\ \frac{\mathrm{d}\sigma_{qg}^{\mathrm{vac}}}{\mathrm{d}\theta \mathrm{d}z} \approx \frac{\alpha_{\mathrm{s}} C_{\mathrm{F}} \sigma}{\pi} \frac{1 + (1-z)^2}{z \theta} + \mathcal{O}(\alpha_{\mathrm{s}}^2, \theta^0) \end{aligned}$$

#### The observable analytically

$$\frac{\mathrm{d}\sigma_{qg}}{\mathrm{d}\theta\mathrm{d}z} = \frac{\mathrm{d}\sigma_{qg}^{\mathrm{vac}}}{\mathrm{d}\theta\mathrm{d}z} (1 + F_{\mathrm{med}}(z,\theta,\hat{q},L))$$

Static brick medium: length L, transport coefficient  $\hat{q}$ .

We assume up to one hard splitting occurs in the medium: q->qg.

The initial quark is considered to have large light-cone energy, as do both its decay products. Formally,  $E \rightarrow \infty$  and  $0 \ll z \ll 1$ .

All three particles undergo broadening and energy loss by interacting with the medium. Broadening and energy loss are resumed in the BDMPS-Z formalism with a harmonic oscillator potential.



We study quark jets with substructure formed from the q->qg process. The diagrams are drawn for  $\gamma$ ->qq processes without meaningful loss of information: 2107.02542 considered  $\gamma$ ->qq, q->qg, and g->gg processes.

#### The observable analytically

Vacuum  $\theta \ll 1$  resummation  $\theta > (EL)^{-1/2}$  Medium induced quenching

$$\frac{\mathrm{d}\Sigma^{(n)}}{\mathrm{d}\theta} = \int \frac{\mathrm{d}z}{\sigma_{qg}} \frac{\mathrm{d}\sigma_{qg}^{\mathrm{vac}}}{\mathrm{d}\theta\mathrm{d}z} (g^{(n)}(\theta,\alpha_{\mathrm{s}}) + F_{\mathrm{med}}(z,\theta,\hat{q},L)) \quad (7)$$
$$\times z^{n} (1-z)^{n} \left( 1 + \mathcal{O}(\alpha_{\mathrm{s}}\ln\theta_{L}^{-1}) + \mathcal{O}\left(\alpha_{\mathrm{s}}\frac{\mu_{\mathrm{s}}^{n}}{E^{n}}\right) \right),$$

where  $g^{(1)} = \theta^{\gamma(3)}$  at fixed coupling The expression for  $g^{(n)}$  with n > 1 is more complicated. However, crucially one still has that  $g^{(n)} \to 1$  as  $\alpha_{\rm s} \ln \theta^{-1} \to 0$ .

#### Part 4: The observable numerically

$$F_{\rm med} = 2 \int_0^L \frac{\mathrm{d}t_1}{t_{\rm f}} \left[ \int_{t_1}^L \frac{\mathrm{d}t_2}{t_{\rm f}} \cos\left(\frac{t_2 - t_1}{t_{\rm f}}\right) \mathcal{C}^{(4)}(L, t_2) \mathcal{C}^{(3)}(t_2, t_1) - \sin\left(\frac{L - t_1}{t_{\rm f}}\right) \mathcal{C}^{(3)}(L, t_1) \right]$$

$$\mathcal{C}_{gq}^{(3)}(t_2,t_1) = \frac{1}{N_c^2 - 1} \left\langle \mathrm{tr}[V_2^{\dagger}V_1] \, \mathrm{tr}[V_0^{\dagger}V_2] - \frac{1}{N_c} \, \mathrm{tr}[V_0^{\dagger}V_1] \right\rangle \,.$$

$$\begin{split} \mathcal{C}_{gq}^{(3)}(t_2,t_1) &= \mathrm{e}^{-\frac{1}{2}\int_{t_1}^{t_2}\mathrm{d}s\,n(s)[N_c(\sigma_{02}+\sigma_{12})-\frac{1}{N_c}\sigma_{01}]} \\ &= \mathrm{e}^{-\frac{1}{12}\hat{q}(t_2-t_1)^3\theta^2\left(1+z^2+\frac{2z}{N_c^2-1}\right)}. \end{split}$$

$$\begin{split} \mathcal{C}_{gq}^{(4)}(L,t_2) &= \frac{1}{N_c^2 - 1} \left\langle \operatorname{tr}[V_1^{\dagger} V_1 V_2^{\dagger} V_2] \operatorname{tr}[V_2^{\dagger} V_2] - \frac{1}{N_c} \operatorname{tr}[V_1^{\dagger} V_1] \right\rangle, \\ &\qquad \qquad \frac{1}{N_c^2} \langle \operatorname{tr}[V_1 V_2^{\dagger} V_2 V_1^{\dagger}] \operatorname{tr}[V_2 V_2^{\dagger}] \rangle \simeq \mathrm{e}^{-\frac{1}{4} \hat{q} \theta^2 (t-t_2) (t_2 - t_1)^2 (1 - 2z + 3z^2)} \\ &\qquad \qquad \times \left( 1 - \frac{1}{2} \hat{q} \theta^2 z (1 - z) (t_2 - t_1)^2 \int_{t_2}^t \mathrm{ds} \, \mathrm{e}^{-\frac{1}{12} \hat{q} \theta^2 [(s-t_2)^2 (2s - 3t_1 + t_2) + 6z (1 - z) (s-t_2) (t_2 - t_1)^2} \right] \end{split}$$



#### The observable numerically $\theta_{c} \gg \theta_{L}$ $E \gg \hat{q}L^{2}$



For angles  $\theta_c \gg \theta \gg \theta_L$ , the quark jet undergoes some energy loss but the substructure is not resolved.



Initial splitting can be resolved by the medium when  $\theta \gg \theta_L$ . Broadening and energy loss occur.

21/07/2022











#### An analysis in JEWEL is now also under way.





Early results indicate the main features of the curves are resilient against a hadron pt cut  $p_t \gtrsim 2$  GeV.

Complimentarity between be measurement at sPHENIX and LHC.

#### Conclusions

Energy Correlators are cool and fun!



- Our early results suggest properties of the QGP can be resolved by using energy correlators for jet substructure.
- Our initial analysis uses the BDMPS-Z model for the numerics. However, the basic features should be model independent, they are set by formation times and uncertainty relations. Could not be explained by changing q/g fraction.
- Correlators are naturally insentive to low scale physics hadronisation, background amd soft corrections typically are sub-leading.
- We are optimistic for a future measurement at sPHENIX and are studying feasibility in JEWEL.

#### Part N/A: Supplemental Material

$$\mathcal{M}_{\gamma \to q\bar{q}} = \frac{e}{E} e^{i\frac{\boldsymbol{p}_1^2}{2zE}L + i\frac{\boldsymbol{p}_2^2}{2(1-z)E}L} \int_0^\infty dt \int_{\boldsymbol{k}_1, \boldsymbol{k}_2} \left[ \mathcal{G}(\boldsymbol{p}_1, L; \boldsymbol{k}_1, t | zE) \, \bar{\mathcal{G}}(\boldsymbol{p}_2, L; \boldsymbol{k}_2, t | (1-z)E) \right]_{ij}$$
$$\times \gamma_{\lambda, s, s'}(z) \boldsymbol{k} \cdot \boldsymbol{\epsilon}_{\lambda}^* \, \mathcal{G}_0(\boldsymbol{k}_1 + \boldsymbol{k}_2, t | E)$$

$$\begin{aligned} \mathcal{G}(\boldsymbol{p}_{1},t_{1};\boldsymbol{p}_{0},t_{0}) &= \int_{\boldsymbol{x}_{1},\boldsymbol{x}_{2}} \mathrm{e}^{-i\boldsymbol{p}_{1}\cdot\boldsymbol{x}_{1}+i\boldsymbol{p}_{0}\cdot\boldsymbol{x}_{0}} \mathcal{G}(\vec{x}_{1},\vec{x}_{0}) \\ \mathcal{G}(\vec{x}_{1},\vec{x}_{0}) &= \int_{\boldsymbol{r}(t_{0})=\boldsymbol{x}_{0}}^{\boldsymbol{r}(t_{1})=\boldsymbol{x}_{1}} \mathcal{D}\boldsymbol{r} \exp\left[i\frac{E}{2}\int_{t_{0}}^{t_{1}}\mathrm{d}s\,\dot{\boldsymbol{r}}^{2}\right]V(t_{1},t_{0};[\boldsymbol{r}]) \\ V(t_{1},t_{0};[\boldsymbol{r}]) &= \mathcal{P}\exp\left[ig\int_{t_{0}}^{t_{1}}\mathrm{d}t\,\mathbf{t}^{a}A^{-,a}(t,\boldsymbol{r}(t))\right] \\ \frac{\mathrm{d}N^{\mathrm{med}}}{\mathrm{d}z\mathrm{d}\boldsymbol{p}^{2}} &= \frac{1}{4(2\pi)^{2}\,z(1-z)}\langle|\mathcal{M}_{\gamma\to q\bar{q}}|^{2}\rangle = \frac{1}{4(2\pi)^{2}\,z(1-z)}\langle|\mathcal{M}_{\gamma\to q\bar{q}}^{\mathrm{i}}+\mathcal{M}_{\gamma\to q\bar{q}}^{\mathrm{out}}|^{2}\rangle \end{aligned}$$

29

#### Part N/A: Supplemental Material

 $\frac{\mathrm{d}\sigma_{qg}}{\mathrm{d}\theta\mathrm{d}z} = \frac{\mathrm{d}\sigma_{qg}^{\mathrm{vac}}}{\mathrm{d}\theta\mathrm{d}z} (1 + F_{\mathrm{med}}(z,\theta,\hat{q},L))$ 

$$F_{\rm med} = 2 \int_0^L \frac{\mathrm{d}t_1}{t_{\rm f}} \left[ \int_{t_1}^L \frac{\mathrm{d}t_2}{t_{\rm f}} \cos\left(\frac{t_2 - t_1}{t_{\rm f}}\right) \mathcal{C}^{(4)}(L, t_2) \mathcal{C}^{(3)}(t_2, t_1) - \sin\left(\frac{L - t_1}{t_{\rm f}}\right) \mathcal{C}^{(3)}(L, t_1) \right]$$

$$\mathcal{C}_{gq}^{(3)}(t_2, t_1) = \frac{1}{N_c^2 - 1} \left\langle \operatorname{tr}[V_2^{\dagger} V_1] \operatorname{tr}[V_0^{\dagger} V_2] - \frac{1}{N_c} \operatorname{tr}[V_0^{\dagger} V_1] \right\rangle. \qquad \mathcal{C}_{gq}^{(3)}(t_2, t_1) = e^{-\frac{1}{2} \int_{t_1}^{t_2} \mathrm{d}s \, n(s) [N_c(\sigma_{02} + \sigma_{12}) - \frac{1}{N_c} \sigma_{01}]} = e^{-\frac{1}{12} \hat{q}(t_2 - t_1)^3 \theta^2 \left(1 + z^2 + \frac{2z}{N_c^2 - 1}\right)}.$$