

Probing the Hydrodynamics of Strongly Coupled QFTs

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arXiv:2206.04765 with Romatschke

arXiv:2104.02024 with Cohen, Lamm, Yamauchi

arXiv:2111.08158

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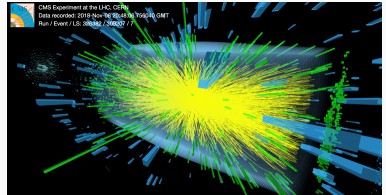
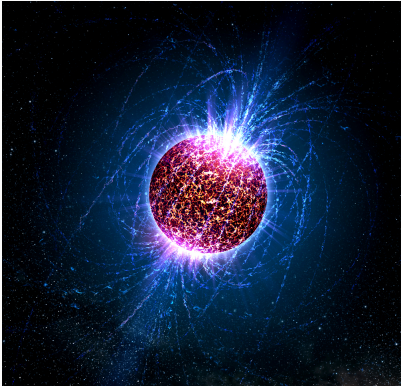
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The universe has not yet equilibrated

Heat death of the universe is not expected for $\sim 10^{100}$ more years.



For now: need nonperturbative methods for nonequilibrium physics.

Hydrodynamics

Finite-temperature quantum matter (**hard to simulate!**), when “zoomed out”, is described by *classical* hydrodynamics (“**easy**”).

Navier-Stokes:

$$\rho \frac{du_i}{dt} + \partial_i p = \eta \left(\frac{1}{3} \partial_i \partial_j u_j + \partial_j^2 u_i \right) + \zeta \dots$$

More systematic (and relativistic): $\partial_\mu T^{\mu\nu} = 0$, expand $T^{\mu\nu}$ in ∇ :

$$T^{\mu\nu} = T_{(0)}^{\mu\nu} - \underbrace{2\eta \nabla^{\langle\mu} u^{\nu\rangle} - \zeta \Delta^{\mu\nu} \nabla_\lambda^\perp u^\lambda}_{T_{(1)}^{\mu\nu}} + \underbrace{\kappa [R^{\langle ij \rangle} - 2R^{t\langle ij \rangle t}]}_{T_{(2)}^{\mu\nu}} + \dots + \dots$$

Gradient expansion: long distances, long times

Transport coefficients: LECs through which quantum effects can appear

Sound waves

$$\int dx \sin kx \langle T^{00}(x, t) T^{00}(0, 0) \rangle \sim \exp \left[i c_s k t - \left(\frac{\zeta + \frac{2(d-1)}{d} \eta}{\epsilon + P} \right) k^2 t \right]$$

Shear waves

$$\int dx \sin kx \langle T^{01}(x, t) T^{01}(0, 0) \rangle \sim e^{-\frac{\eta}{\epsilon + P} k^2 t}$$

“Shear channel”

$$\langle T^{12}(\omega, k) T^{12}(\omega, k) \rangle = P - i\eta\omega + O(\omega^2) + O(k)$$

Paradox: $\eta_{\text{atm}} \sim 10^{-5} \frac{\text{kg}}{\text{m s}}$ results in a sound attenuation time on the order of days, but everyday experience says it should be seconds.

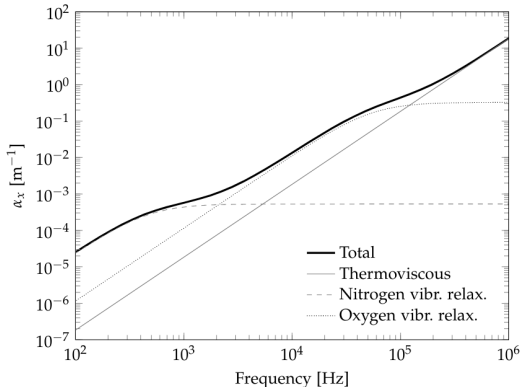


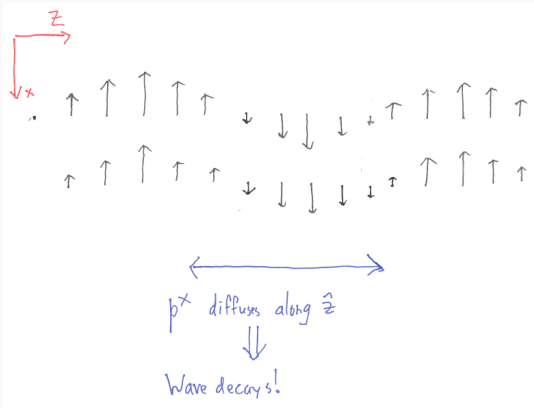
Figure 2.7: Absorption coefficient and its composition for air at 293.15 K, atmospheric pressure and 70% relative humidity. Rotational relaxation for nitrogen and oxygen is represented as bulk viscosity.

(Figure from E. M. Viggien's thesis.)

Shear viscosity

Two equivalent phrasings of what η represents:

Transport of x -momentum along \hat{z} OR decay of shear waves



Shear waves in a free theory

Impart some momentum at $x = 0$. After time t , what is the amplitude of the shear wave?

$$C(t) \sim \int d^3v \underbrace{\rho(v)}_{\sim e^{-v^2}} \cos(kv_z t) \sim e^{-t^2}$$

This decay is *gaussian*, faster than any exponential. $\eta_{\text{free}} = \infty$.

Physically: the transport of momentum (carried by individual particles) is entirely unobstructed.

Another example: in a rigid body, $\eta = \infty$. (Proof: pick up a pen.)

A nontrivial lower bound on transport?

In a free theory, particles are free, transport momentum efficiently. Sound waves decay super-exponentially.

In a rigid body, particles can't move at all. Phonons are free, and don't decay, so they transport momentum efficiently.

In a rigid body, phonons are efficient transporters *because* particles can't move.

Blocking one mechanism of momentum transport opens up another one (by allowing sound waves to propagate without decaying).

A more careful version of this argument is in Kovtun, Moore, Romatschke, "The Stickiness of Sound".

Status of the KSS conjecture

KSS Conjecture: $\frac{\eta}{s} \geq \frac{1}{4\pi}$ in “all” theories.

Counterexample from T. Cohen (arXiv:0702136) involves a large number of **weakly interacting** massive species.

In this family of models $\frac{\eta}{s}$ is made arbitrarily small, but shear waves still decay quickly.

Moral: bound the decay constant, not $\frac{\eta}{s}$!

See e.g. arXiv:2111.08158 (Lawrence) and arXiv:2005.06482 (Baggioli and Li)

Large- N_f expansions

Nonrelativistic fermions in three dimensions:

$$S = \int_0^\beta d\tau \int d^3\mathbf{x} \left[\sum_f \psi_{s,f}^\dagger(\mathbf{x}) \left(\partial_\tau - \frac{\nabla^2}{2m} - \mu \right) \psi_{s,f}(\mathbf{x}) + \sum_{f,f'} \frac{4\pi a_s}{mN} \psi_{\uparrow,f}^\dagger(\mathbf{x}) \psi_{\downarrow,f'}^\dagger(\mathbf{x}) \psi_{\downarrow,f}(\mathbf{x}) \psi_{\uparrow,f'}(\mathbf{x}) \right]$$

(The limit $a_s \rightarrow -\infty$ is the unitary Fermi gas.)

To expand in powers of N^{-1} , introduce an auxiliary field to make N a parameter.

$$S_{\text{eff}} = -N \left[\log \det \left(\partial_\tau - \sigma_z \frac{\nabla^2}{2m} - \sigma_z \mu + i\zeta^* \sigma_- - i\zeta \sigma_+ \right) - \int \frac{m\zeta\zeta^*}{4\pi a_s} \right]$$

Evaluate the path integral $Z = \int \mathcal{D}\zeta e^{-S_{\text{eff}}}$ with a saddle-point expansion.

Thermodynamic transport

Not all LECs in the hydrodynamic expansion are specific to out-of-equilibrium physics.

- Pressure
- Gravitational wave-to-matter coupling (κ)

These appear when the spacetime metric undergoes a time-independent perturbation.

Equivalently, these are detectable from fluctuations in thermodynamic equilibrium.

Gravitational wave-to-matter coupling of the unitary Fermi gas

$$T^{ij} \supset \kappa [R^{(ij)} - 2R^{t(ij)t}]$$

Thermodynamic transport can be seen from *Euclidean* correlators:

$$\kappa = \left. \frac{\partial^2}{\partial k^2} \langle T^{12} T^{12} \rangle (\omega = 0, k) \right|_{k=0}$$

For nonrelativistic fermions, the stress-energy tensor is:

$$T^{12} = \frac{1}{4m} \left[\partial_1 \Psi^\dagger \sigma_z \partial_2 \Psi + \partial_2 \Psi^\dagger \sigma_z \partial_1 \Psi - \partial_1 \partial_2 \Psi^\dagger \sigma_z \Psi - \Psi^\dagger \sigma_z \partial_1 \partial_2 \Psi \right] - \frac{is}{4m} \partial_k \Sigma_k$$

$$\text{where } \Sigma_3 = \Psi \sigma_x (\partial_1 - i\partial_2) \Psi + \Psi^\dagger \sigma_x (\partial_1 + i\partial_2) \Psi^\dagger$$

Evaluated at the saddle point:

$$C(k) = -\frac{2}{m^2} \int \frac{d^4 p}{(2\pi)^4} \frac{p_1^2 p_2^2 \left[(\epsilon_{\mathbf{p}-\mu})(\epsilon_{\mathbf{k}+\mathbf{p}-\mu}) - \omega^2 - \Delta^2 \right]}{\left[(\epsilon_{\mathbf{p}-\mu})^2 + \omega^2 + \Delta^2 \right] \left[(\epsilon_{\mathbf{p}+\mathbf{k}-\mu})^2 + \omega^2 + \Delta^2 \right]} + \frac{s^2 k^2}{2m^2} \int \frac{d^4 p}{(2\pi)^4} p_1^2 \text{tr}[\sigma_x G(\omega, \mathbf{p}) \sigma_x G(-\omega, -\mathbf{p})]$$

Integrating:

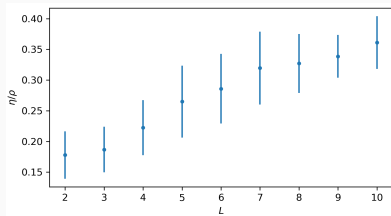
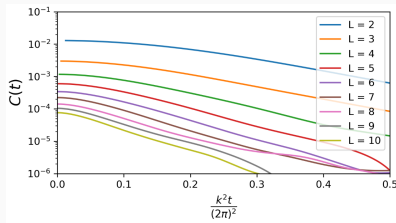
$$\lim_{a_s \rightarrow -\infty} = \frac{(2m\mu)^{\frac{3}{2}}}{3\pi^2 m} \frac{1}{\xi^{\frac{3}{2}}} \frac{N}{12} = \frac{n}{12m}$$

Viscosity from molecular dynamics

From an equilibrated MD simulation, compute a time-series:

$$f(T) = \sum_n \rho_x(T) \sin k_z x_z(T)$$

Plot the autocorrelation and fit the decay to $f(T) \sim e^{-\frac{\eta k^2}{\rho} t}$:



(Inter-particle potential $V(r) \sim e^{-2r}$)

Viscosity from a quantum computer

The best model of a cat is another cat. (*Norbert Wiener*)

Quantum simulations are *conceptually* simple.

- Physical Hilbert space \mathcal{H}_P ; qubit Hilbert space $\mathcal{H}_C \approx \mathbb{C}^{2^Q}$
- Define an injective map $\mathcal{H}_P \rightarrow \mathcal{H}_C$
- Decompose operators H and \mathcal{O} in 1- and 2-Pauli terms
- Time-evolve, possibly with $H = H(t)$, and then measure!

Either measure T^{01} repeatedly and look at autocorrelation (best on large systems), or explicitly measure $\langle T^{01} T^{01} \rangle$ via linear response (better on small systems).

See arXiv:2104.02024 (Cohen, Lamm, Lawrence, Yamauchi) for details.

Real-time path integrals

Let's derive a (lattice) path integral for time-separated correlators.

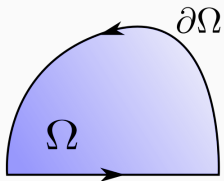
$$\begin{aligned} \langle \mathcal{O}(t)\mathcal{O}(0) \rangle &\propto \text{Tr}(e^{-\delta H})^{\beta/\delta} (e^{iH\delta})^{t/\delta} \mathcal{O} (e^{-iH\delta})^{t/\delta} \mathcal{O} \\ &\propto \int \mathcal{D}\phi \underbrace{\langle \phi | e^{-\delta H} | \phi' \rangle \cdots \langle \phi' | e^{iHt} | \phi'' \rangle \cdots \langle \phi'' | e^{-iHt} | \phi''' \rangle}_{\text{Complex!}} \mathcal{O}(\phi'') \mathcal{O}(\phi''') \end{aligned}$$

The “probability distribution” is complex!

Moreover, if you replace the complex $\rho(\phi)$ with $|\rho(\phi)|$, almost all of the weight cancels. ($\int |\rho| \gg \int \rho$)

(One can also study real-time correlators by analytic continuation. I have nothing to say about this here.)

Cauchy's Theorem

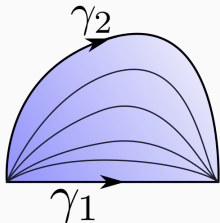


For any holomorphic function $f(z)$:

$$0 = \int_{\partial\Omega} f dz$$

$$\int_{\gamma_1} f dz = \int_{\gamma_2} f dz$$

If we can continuously deform γ_1 to γ_2 , “tracing out” Ω .



Physics is unchanged, but fluctuations can be removed ($\int |p(\phi)|$ changes)

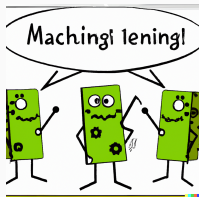
Stokes' Theorem

$$\int \mathcal{D}\phi e^{-S(\phi)} = \int e^{-S(\phi) + \nabla \cdot v(\phi)}$$

This is a strict generalization of the Cauchy's theorem based methods.

Any sign problem can be removed by some appropriate v .

The vector field v is responsible for representing the “pure fluctuation” part of the integrand.



The transverse Ising model

$$H = -\mu \sum_r \sigma_x(r) - \sum_{\langle rr' \rangle} \sigma_z(r) \sigma_z(r')$$

In one dimension, this is (dual to via Jordan-Wigner) a theory of free fermions. Tuning $\mu \rightarrow \mu_c$ makes the fermions massless. This is a free CFT.

In two dimensions, this theory is interacting. Tuning to $\mu_c \approx 3.044$ yields the $O(1)$ (Ising) CFT; also accessible from scalar field theory.

Thermodynamics and RG flows around the $O(N)$ models are very well understood (in $2 + 1$ dimensions):

- Bootstrap
- Monte carlo
- ϵ expansion

Simulating the transverse Ising model

Ising spin \leftrightarrow qubit

On a universal quantum computer, time-evolve via Suzuki-Trotter:

$$e^{-i(A+B)\epsilon} \approx e^{-iA\epsilon} e^{-iB\epsilon}$$

TIM is the first model with nontrivial hydrodynamics likely to be accessed.

In both cases, evaluate $\langle T^{00}(t, x) T^{00}(0, 0) \rangle$ and observe the decay.

Exponential-cost classical algorithm

Start with a Haar-random $|\Psi\rangle$.

Apply $e^{-\frac{\beta}{2}H}$ to obtain a thermal state. (Agrees with canonical ensemble in the large-volume limit.)

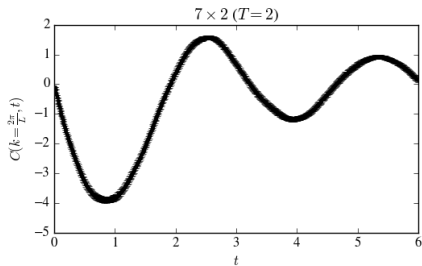
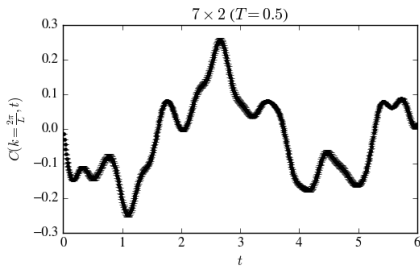
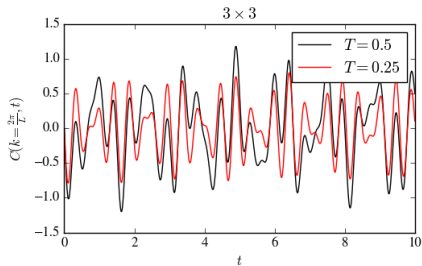
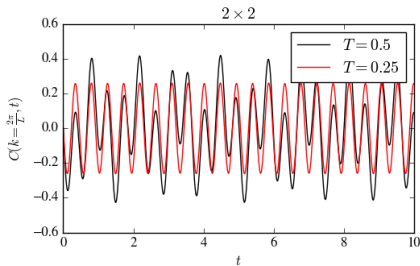
Apply operator \mathcal{O} .

Apply $(1 - i\epsilon H)^{\frac{t}{\epsilon}}$ to time-evolve.

Apply \mathcal{O} again to get expectation value.

Sound attenuation in the transverse Ising model

In progress!



Large-N calculation in the $O(N)$ CFT

From arXiv:2104.06435 (Romatschke):

$$\frac{1}{N} \left(\frac{\eta T}{\epsilon + P} \right)_{O(N)} = 0.42(1) + O(N^{-1})$$

Recklessly extrapolating gives an Ising shear viscosity of

$$4\pi \frac{\eta T}{\epsilon + P} = 5.28(12).$$

Do other methods agree?

The Shape of Things To Come

- Most analytic methods are limited to narrow regimes. (Weak coupling; large N ; holographic theories.)
- Scalable quantum computers will come online *eventually*; will they come before efficient classical methods are available?
- Machine learning for sign problems
- Bootstrap methods

How slow can transport be?