LAr Doping for low-energy sensitivity of LArTPCs

Fernanda Psihas





Photosensitive dopants (dopants that convert light into charge) might change what is achievable for small signals in LAr experiments.



Improvements to energy resolution at the MeV scale would also improve the physics sensitivity of lowenergy physics signals of interest in current LArTPC experiments.



A low-energy LArTPC R&D program can expand the physics reach of future LArTPCs and by enabling new physics below the 10s of MeV.

LARTPO PHYSICS ENERGY RANGE

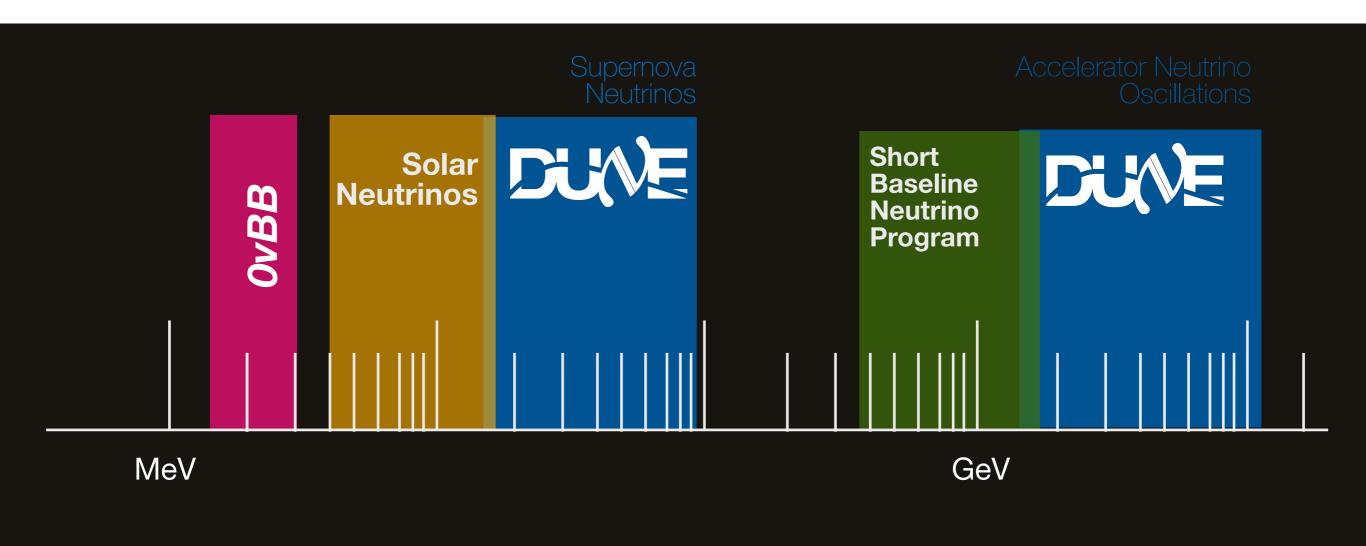
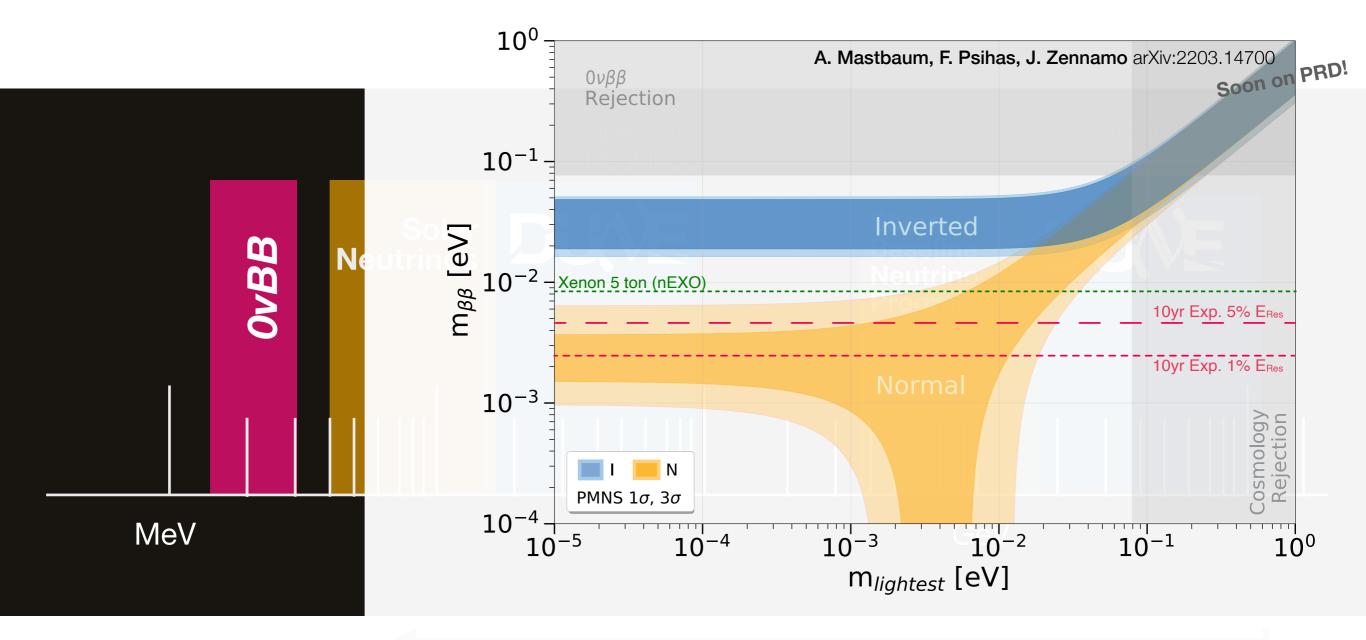


Photo-sensitive dopants can improve

→ Photo-sensitive dopants can enable

A OVBB SENSITIVE KTON LARTPC



DUNE-β Concept ——

Xenon doping at 2%



! Depleted argon

Fernanda Psihas

<3% energy resolution</p>

THIS TALK _____

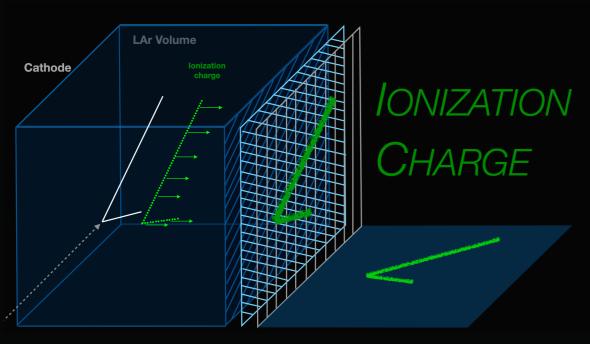
Photosensitive dopants

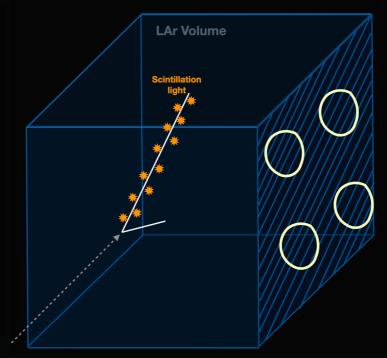






SIGNALS IN LARTPCS





SCINTILLATION LIGHT



Directional



Very slow



Information about trajectory and energy



Isotropic



Very fast



Information about timing and

*possibly also energy



-X- Collected very efficiently



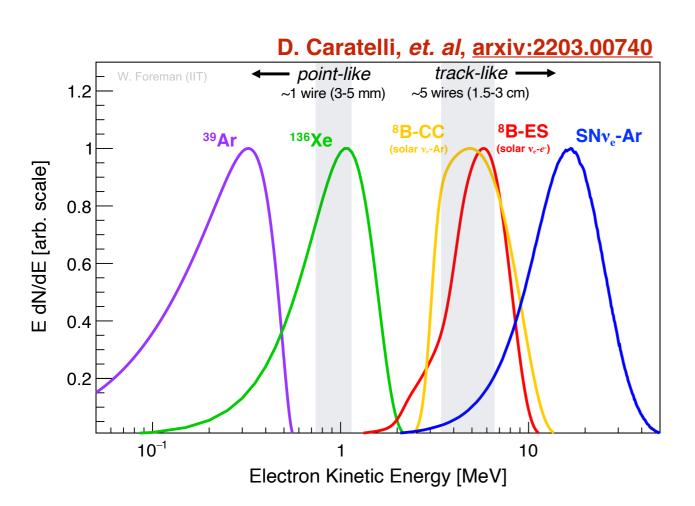
Collected with less efficiency

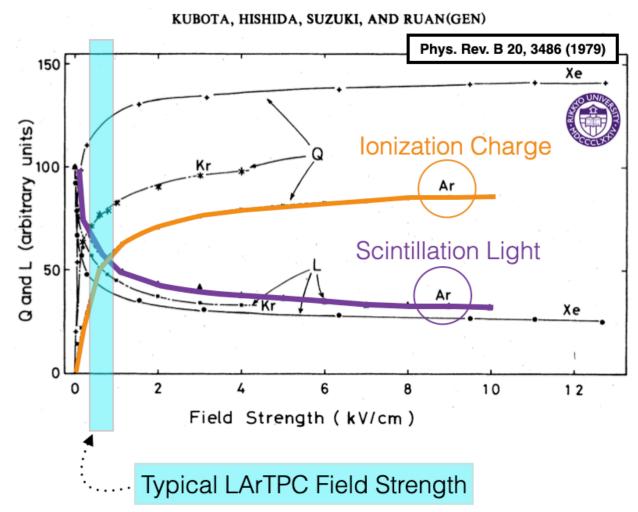
MATTERS A LOT AT LOW ENERGIES

WHY LIGHT WILL MATTER FOR SMALL SIGNALS

Charge + Light = Constant

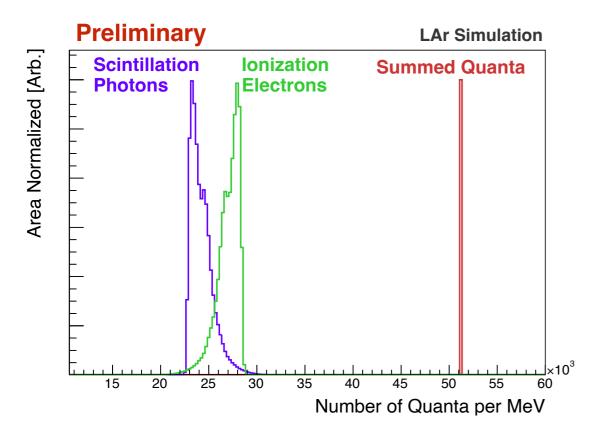
On DUNE, we'll expect ~50/50 charge to light breakdown.





This ratio is sufficient for the needs of GeV physics but will impact our ability to do physics at the MeV scale

CHARGE + LIGHT = BETTER RESOLUTION AT LOW ENERGIES



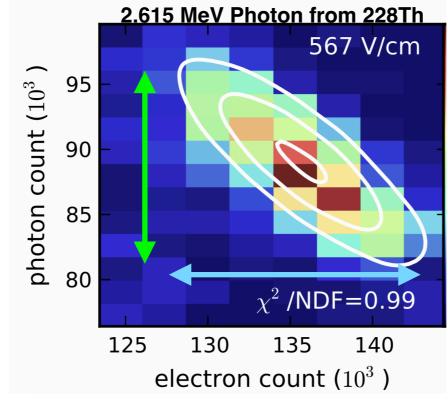
Anti-correlated light and charge signals

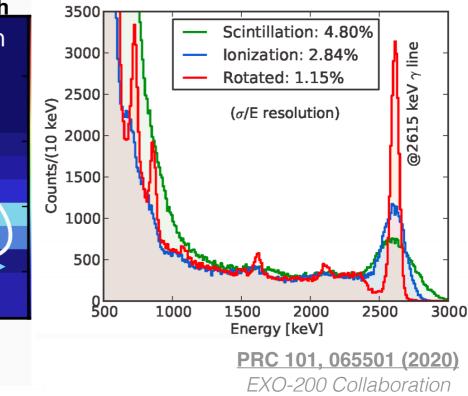
Using information from both signals can yield a more precise energy measurement.

EXO-200, a LXeTPC searching for 0vββ, explored the anti-correlation between light and charge signals

By combining light and charge they were able to improve their energy resolution by 3x, to ~1%

To achieve this they collected 30,000 γ/MeV





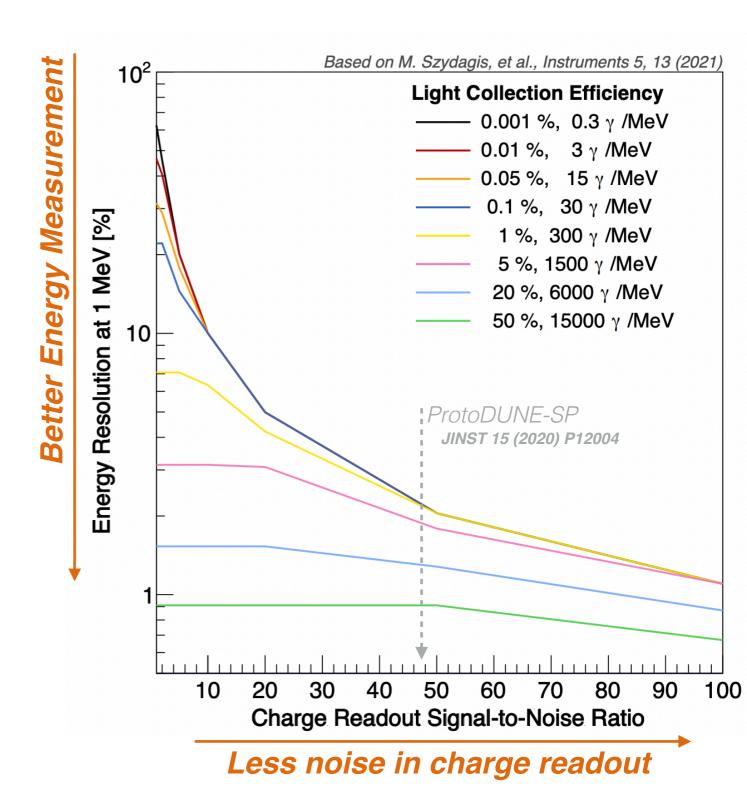
HOW MUCH LIGHT DO WE NEED?

NEST^[*] models the microphysics of energy deposits in noble liquids and gases.

Explored the energy resolution for 1 MeV electrons in LAr for detectors with various efficiency and noise conditions

Achieving the best possible energy resolution need to collect at least 6000 photons per MeV

[*] Noble Elements Simulation Technique, http://nest.physics.ucdavis.edu/



Traditionally light collected at anode plane

DUNE FD Module 5,600 cm by 1,200 cm



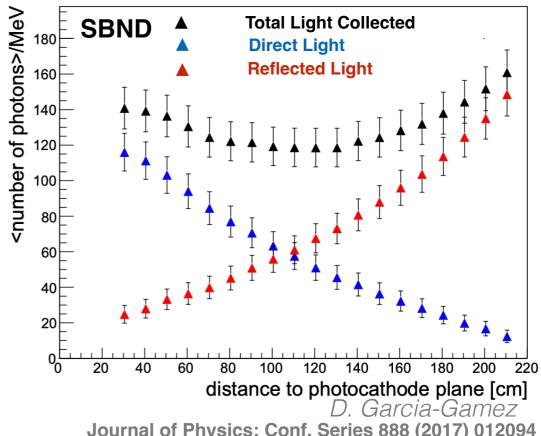
Increasing light collection on large LArTPCs is a challenge:

- Scintillation photons have to travel large distances.
- Low photon detection coverage by design.

The best light collection efficiency has been accomplished on SBND

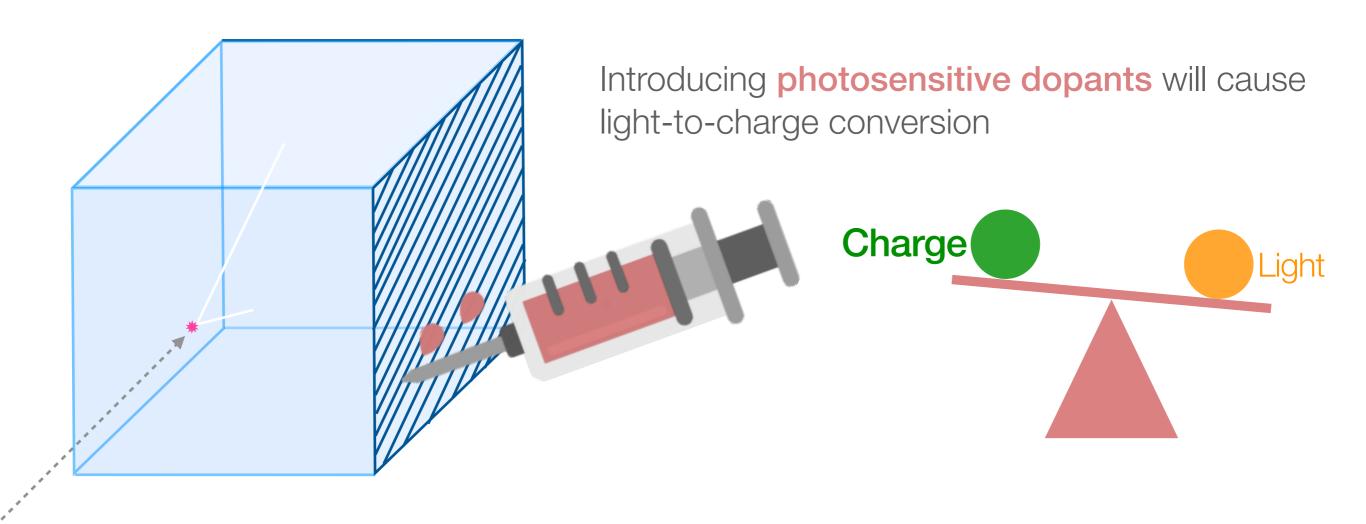
Best LArTPC

Light collection < 160 photons/MeV << 6000 photons/MeV



Journal of Physics: Conf. Series 888 (2017) 012094

PHOTOSENSITIVE DOPANT CONCEPT



What we know:



Good indications that this is a promising avenue of R&D

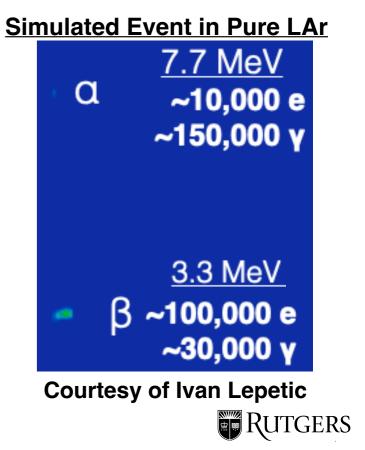
R&D Questions

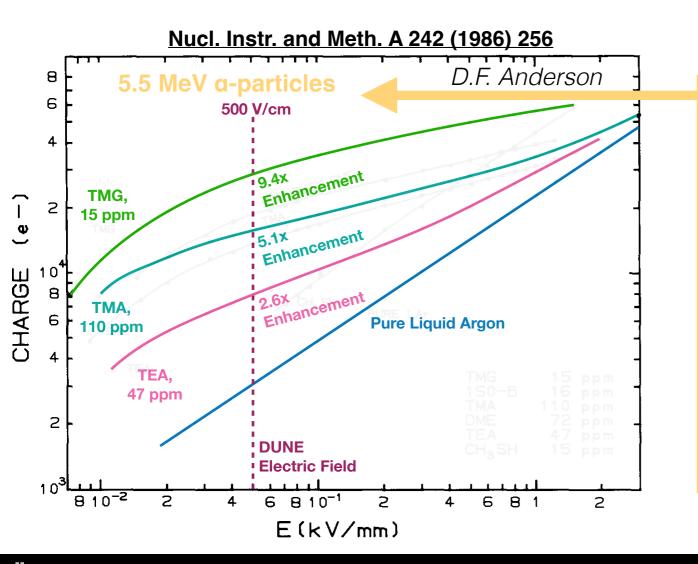
Lot's of productive and impactful R&D for the coming years.

R&D QUESTIONS FOR DOPANTS

Small test stands explored a variety of chemicals and found an increase in charge for highly scintillating particles.

The most commonly used have ionization energies of 7-9 eV: Tetramethylgermane (**TMG**), (CH₃)₄Ge, Trimethylamine (**TMA**), N(CH₃)₃, Triethylamine (**TEA**), N(CH₂CH₃)₃





Does this extend to β's below 5 MeV?

What does the improvement look like and what fraction (if any) of the light survives after doping?

What energy resolutions are achievable below 5MeV with the addition of dopants?

EXPECTED IMPROVEMENTS

ICARUS doped a 3-ton prototype LArTPC with TMG to the few ppm level

TMG was selected because it didn't react with their filter material and was easily purified

After introducing TMG observed:

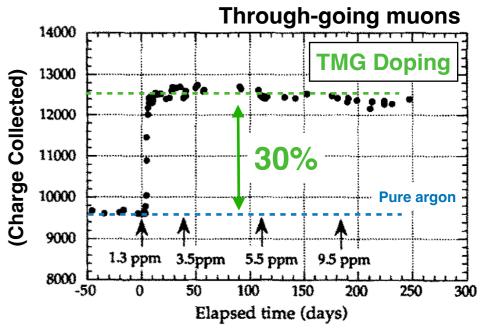
30% increase in muon charge signals

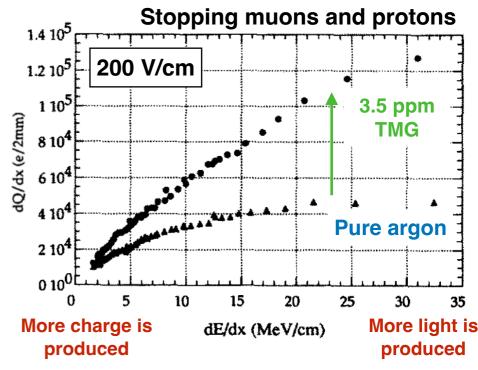
Stable operation for 250 days

Found a more linear detector response for highly ionizing particles

Nucl. Instrum. Methods. Phys. Res. B 355, 660 (1995).

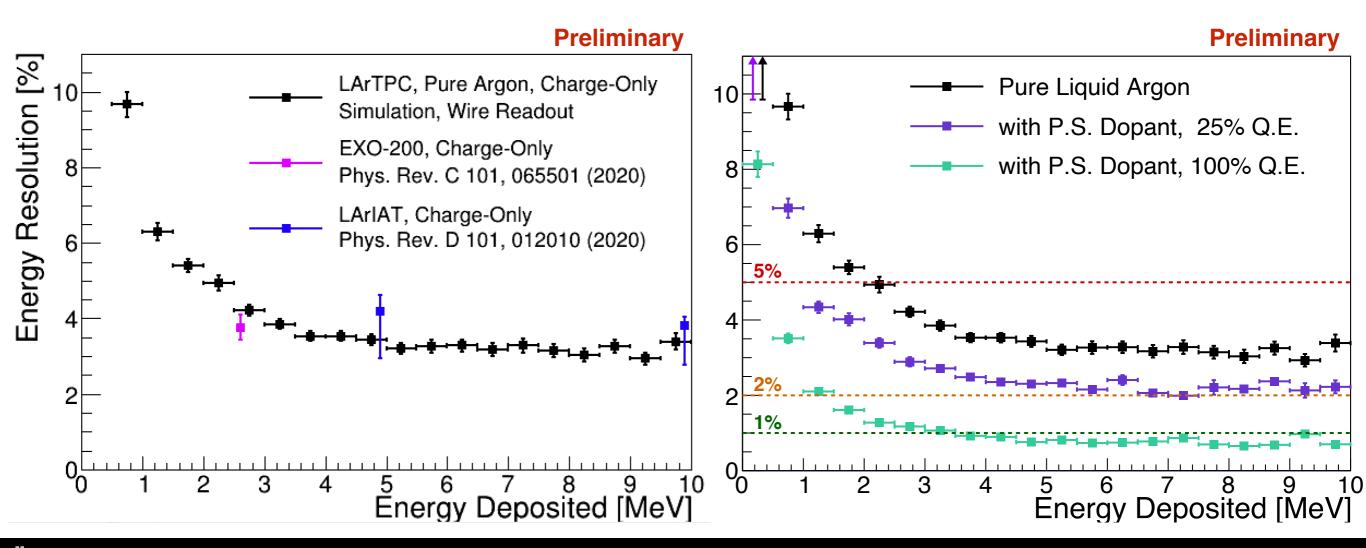
ICARUS Collaboration





Studied improved electron response with simulation of dopants

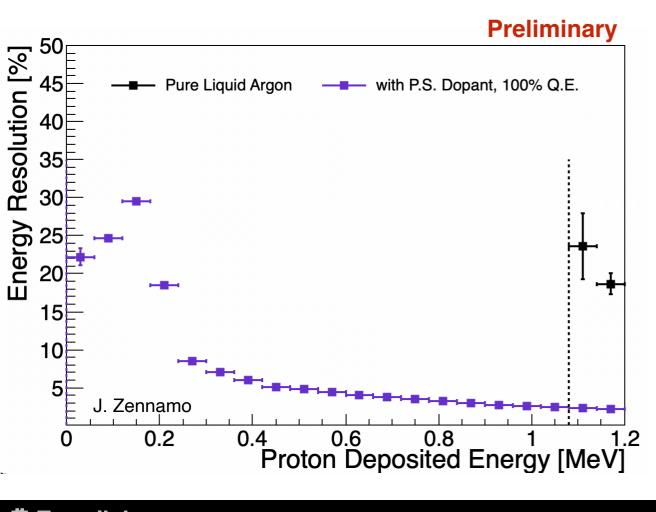
- Converts scintillation light to ionization charge, fully integrated into LArSoft
- Performed a full large LArTPC detector simulation
 - Included wire noise (~350 ENC, ~40 SNR for MIPs), microphysical effects, detector response, noise filtering, signal processing, and energy reconstruction

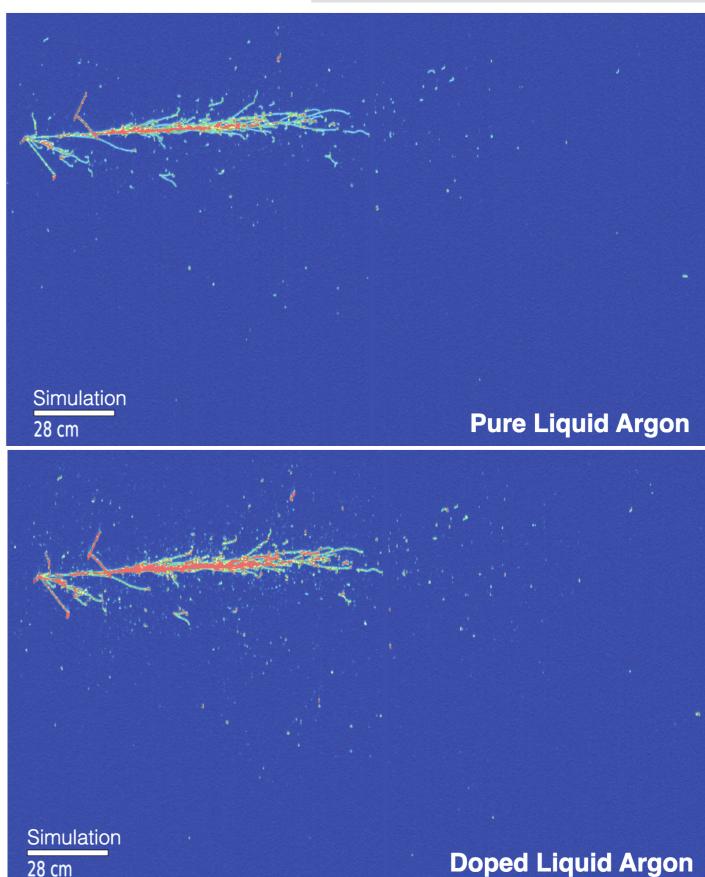


SIMULATION-EXPECTED IMPROVEMENTS

Converting the light to charge:

- Lowers threshold 1.1 MeV to 10 keV
- Improves charge only energy resolution





TINYTPO @ FERMILAB

Low-energy LArTPC test-stand for photo-sensitive dopants:

Stage 1:

Operate TinyTPC with LArPix v2 pixel readout in a cryostat at Fermilab. (Min bias data run)

Stage 2:

Introduce radioactive sources and benchmark LArPix v2 performance at low energies

Stage 3:

Introduce dopants and demonstrate charge enhancement

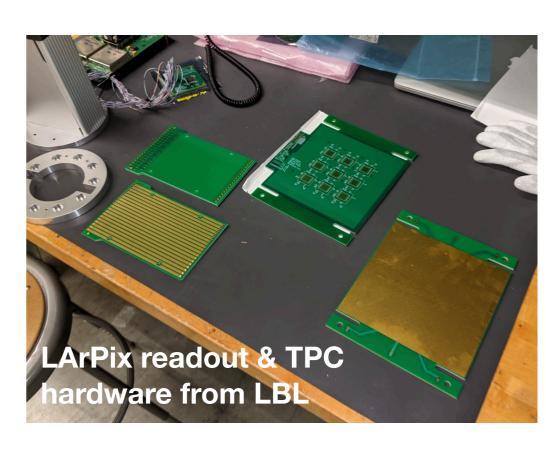
Stage 4:

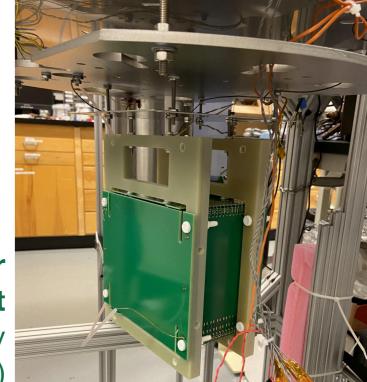
Explore performance enhancements for low energy signals and optimal doping strategies

GOALS

*Demonstrate that dopants produce charge enhancement for low energy signals. *Characterize performance of pixel readouts for low energy signals with and without doping.

Plan for deployment (post assembly & bench tests)





TINYTPC @ FERMILAB

Raofa Raisa 💠

SIST INTERN

Panos Englezos

Mechanical support design & prototyping



Assembly at Fermilab



Stage 1 deployment expected in early 2023.

Operate TinyTPC with LArPix v2 in cryostat for min bias data

Thank you for the support from Fermilab cryo facilities and staff.

Funding from DOE, FNAL New Initiatives Grant and the FNAL Neutrino Division

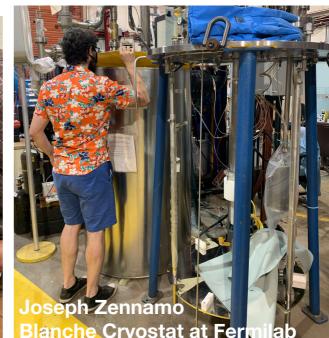
SIST Internship & GEM internship program at Fermilab

Thank you yo LBNL for parts and electronics!

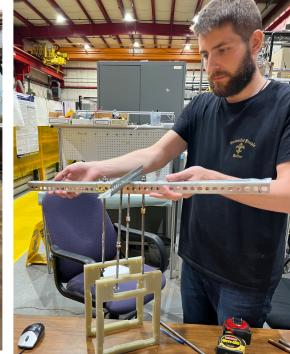
Bench testing







Stay tuned!



WIDER R&D PROGRAM

Extend demonstrations of dopant effects at energies below 5 MeV

Demonstrations of feasibility at kton-scale

Searches for & design of **optimal doping scenarios** for desired light-to-charge ratios

Studies of the interaction of dopants with:

- other dopants (i.e. Xe)
- filtration systems
- fluid dynamics in the cryostat

What is the impact on the LArTPC physics capabilities?

- Timing in a light-less DUNE
- Enhancement of low energy components of GeV events
- Improvements to other low energy signal sensitivities

THOUGHT-PROVOKING IDEAS FOR PHOTOSENSITIVE DOPANTS

Photo-sensitive dopant R&D could change how we think of low energy physics with LArTPCs



LArTPC runs with reduced light, enhanced charge and ~1% energy resolution at 1 MeV.

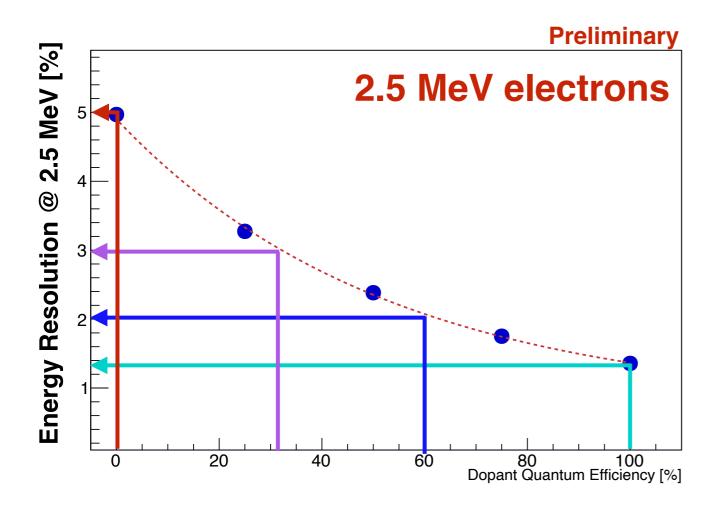
Large LArTPC data runs with interchangeable doping strategies.

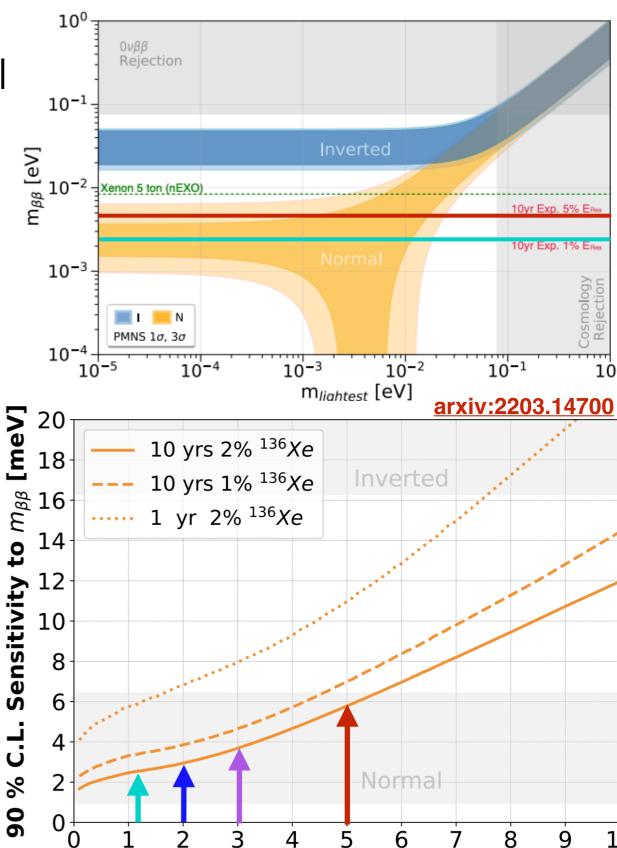
Your low-energy analysis idea enabled by the ability to alter LArTPC light-to-charge ratio



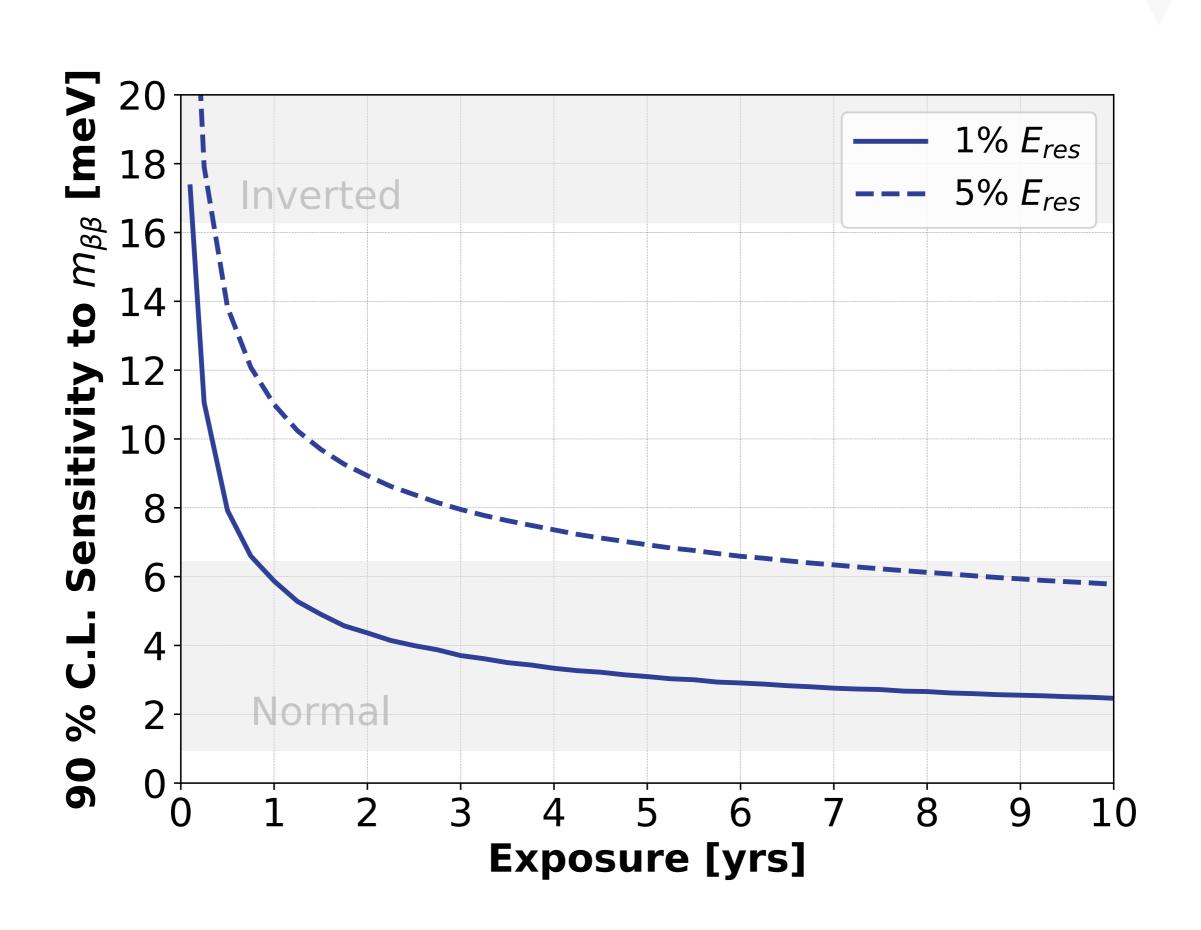
The achievable energy resolution will have a direct impact on how strong of a 0vββ search we can perform

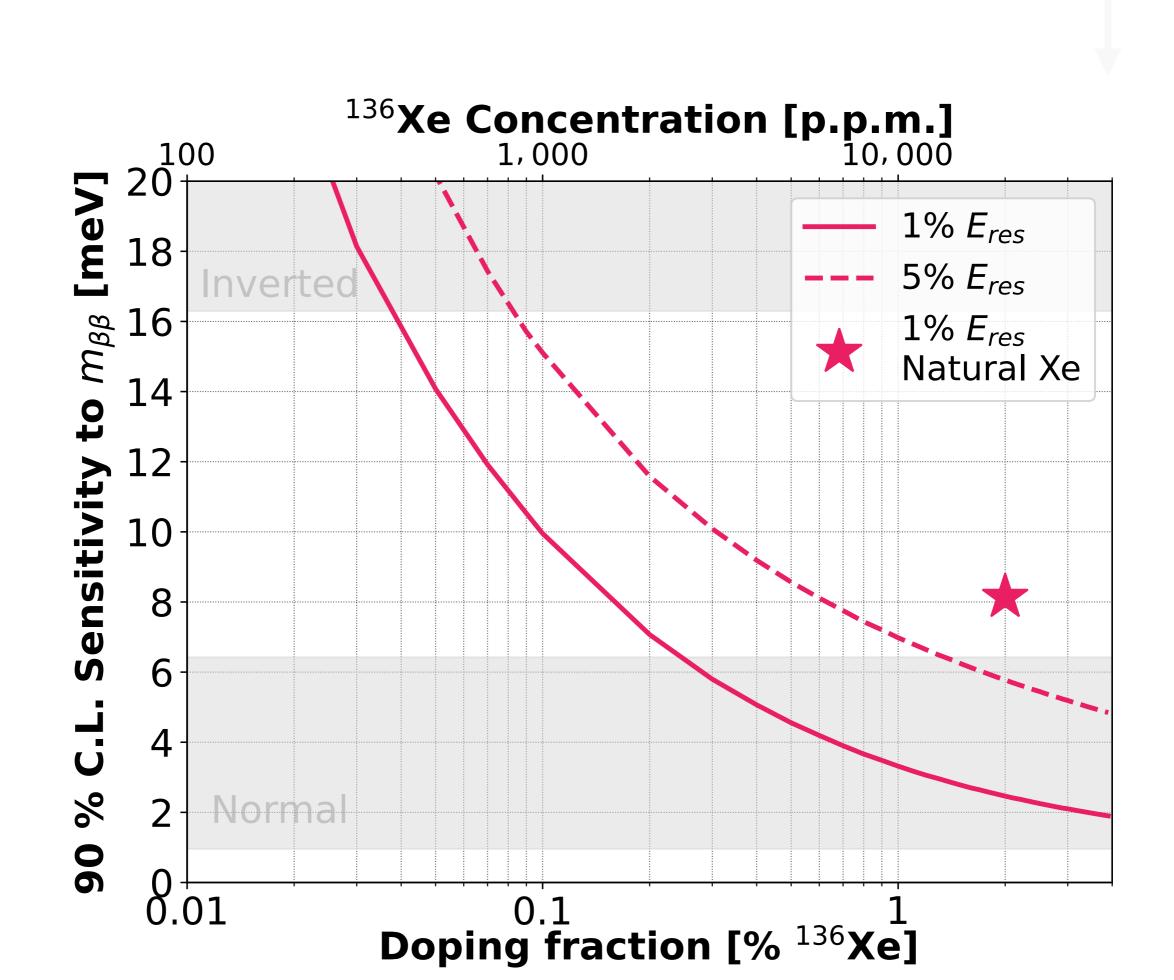
Even without doping, DUNE could reach into the normal heirarchy

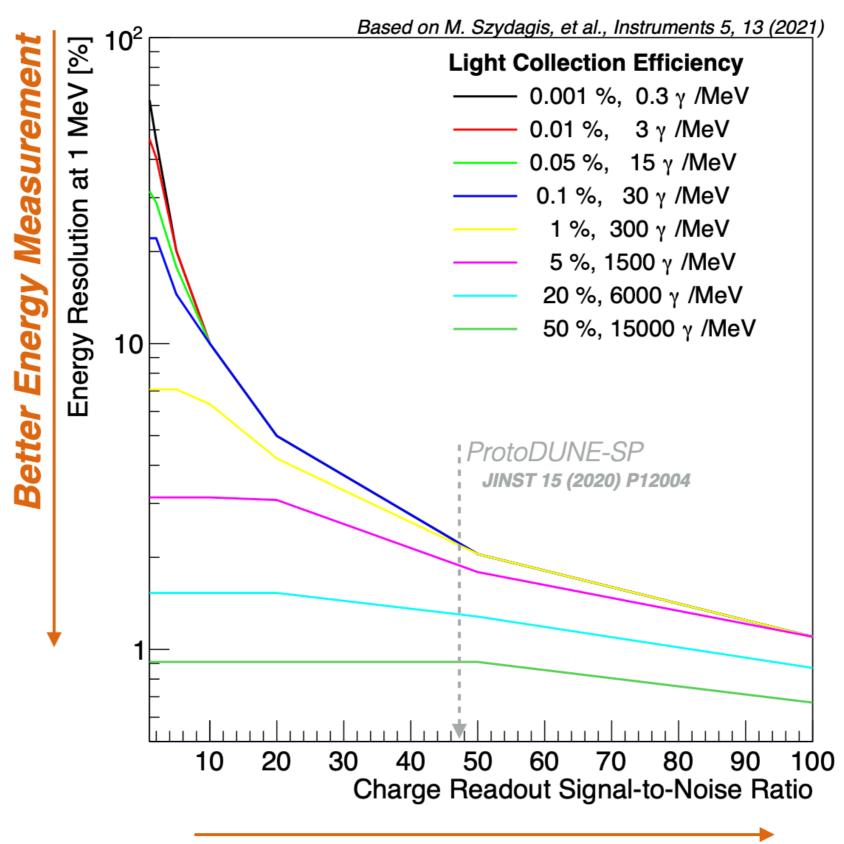


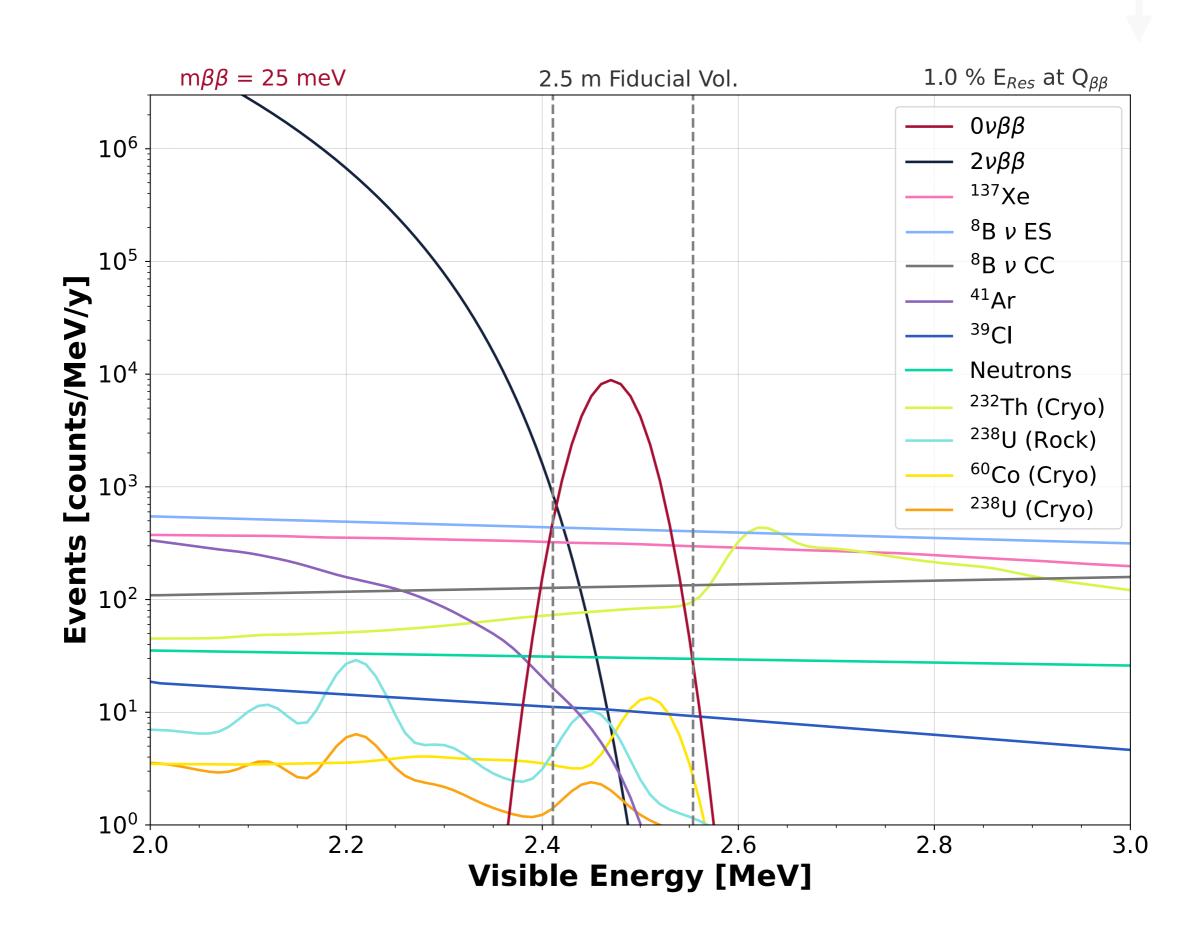


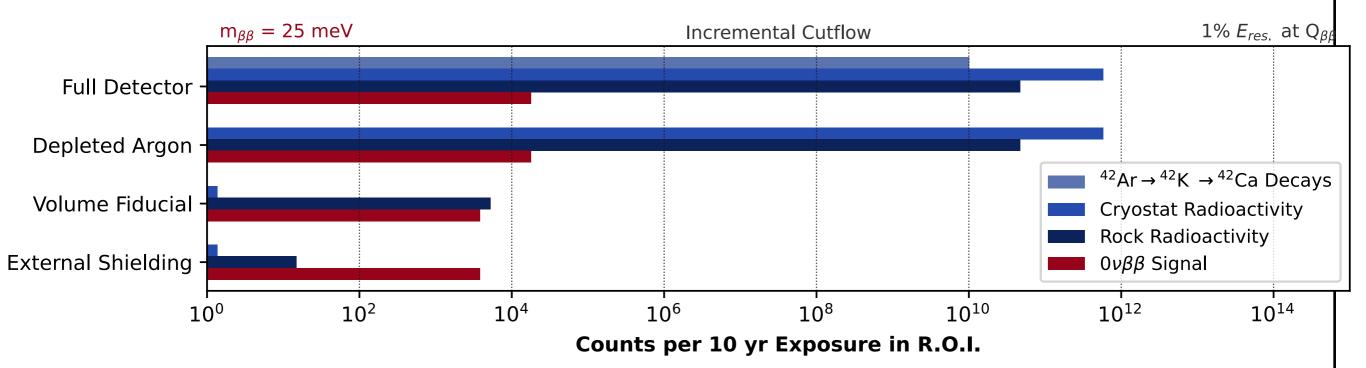
Energy Res [%]











BACKGROUND MITIGATION

Background	Activity	Events in ROI	Mitigation strategy	
Isotope Intrinsic				And paper on arxiv
$^{136}\mathrm{Xe},2 uetaeta$	2% , $T_{1/2} = 2.165 \times 10^{21}$ years [61]	130.28	None	<u>arxiv 2203.10147</u>
Environmental Radiole				
²³² Th, Rock	3.34 ppm [8, 52]	10.71	Passive Shielding	Demonstrated the feasibility
²³⁸ U, Rock	3.34 ppm [8, 52] 7.11 ppm [8, 52]	340.71	Passive Shielding	of reconstructing the MeV-
²³² Th, Steel	0.1 ppb [50]		Fiducialization	<u> </u>
²³⁸ U, Steel	1 ppb [50]	2.24	Fiducialization	scale signal 214Bi-214Po
⁶⁰ Co, Steel	0.013 mBq/g [50]	10.09	Fiducialization	topology in a large-scale
³⁹ Ar, LAr	1 Bq/kg [62]	Negligible	Energy threshold	wire-readout LArTPC
²²² Rn, LAr	$10 \text{ mBq/m}^3 [8]$		Coincident ²¹⁴ Po Tag	
⁴² Ar, LAr	Negligible [63]	Negligible	Use of ⁴² Ar depleted LAr	
Solar Neutrinos				
8 B ν Elastic Scatters	Standard Solar Model Flux [64]	662.04		
8 B ν_e Charged Curren	t Standard Solar Model Flux [64]	196.00	Photon Coincidence Tag	
Spallation Products				Veto photon coincidence
^{32}P	$34 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$		Photon Coincidence Tag	within 32 cm of signal
^{39}Cl	$150 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$		Coincident Muon Timing	
$^{41}\mathrm{Ar}$	$1600 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$		Photon Coincidence Tag	candidates
$^{137}\mathrm{Xe}$	$3.8 \text{ day}^{-1} (10 \text{ kton})^{-1} [65]$	449.43	Photon Coincidence Tag	
$^{16}\mathrm{N}$	$0.033 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$	0 0	Coincident Muon Timing	
30 Al	$1.4 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$		Coincident Muon Timing	
40 Cl	$27 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$		Coincident Muon Timing	\ /atai.a al aitla i.a
20 F	$2 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$	0 0	Coincident Muon Timing	Veto window within
^{34}P	$12 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$	0 0	Coincident Muon Timing	2m and 60sec of all
³⁸ Cl	$110 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$		None	
³⁶ Cl	$110 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$		None	muon tracks.
37 Ar	$110 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$	0 0	Photon Coincidence Tag	
^{33}P	$34 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$	0 0	Photon Coincidence Tag	
$^{11}\mathrm{Be}$	$0.34 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$	Negligible	Coincident Muon Timing	

See Andy Mastbaum's

talk from this meeting

BACKGROUND MITIGATION

Background	Activity	Events in ROI	Mitigation strategy				
Isotope Intrinsic							
136 Xe, $2\nu\beta\beta$	2% , $T_{1/2} = 2.165 \times 10^{21}$ years [61]	130.28	None	2νββ			
Environmental Radiological Backgrounds							
²³² Th, Rock	3.34 ppm [8, 52] 7.11 ppm [8, 52]	16 71	Passive Shielding				
²³⁸ U, Rock	7.11 ppm [8, 52]	J40.71	Passive Shielding				
²³² Th, Steel	0.1 ppb [50]	117.80	Fiducialization	Radioactivity			
²³⁸ U, Steel	1 ppb [50]	2.24	Fiducialization				
⁶⁰ Co, Steel	0.013 mBq/g [50]		Fiducialization				
$^{39}\mathrm{Ar,\ LAr}$	1 Bq/kg [62]		Energy threshold				
$^{222}\mathrm{Rn,\ LAr}$	$10 \text{ mBq/m}^3 [8]$	Negligible	Coincident ²¹⁴ Po Tag				
42 Ar, LAr	Negligible [63]	Negligible	Use of ⁴² Ar depleted LAr				
$Solar\ Neutrinos$							
8 B ν Elastic Scatters	Standard Solar Model Flux [64]	662.04		Solar			
8 B ν_{e} Charged Current	Standard Solar Model Flux [64]	196.00	Photon Coincidence Tag	Colai			
Spallation Products							
^{32}P	$34 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$	Negligible	Photon Coincidence Tag				
³⁹ Cl	$150 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$		Coincident Muon Timing				
⁴¹ Ar	$1600 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$		Photon Coincidence Tag	Spallation			
137 Xe	$3.8 \text{ day}^{-1} (10 \text{ kton})^{-1} [65]$		Photon Coincidence Tag	•			
^{16}N	$0.033 \mathrm{day^{-1}} (10 \mathrm{kton})^{-1} [59]$	Negligible	Coincident Muon Timing				
30 Al	$1.4 \mathrm{day^{-1}} (10 \mathrm{kton})^{-1} [59]$		Coincident Muon Timing				
40 Cl	$27 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$		Coincident Muon Timing				
$^{20}\mathrm{F}$	$2 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$	Negligible	Coincident Muon Timing				
^{34}P	$12 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$	Negligible	Coincident Muon Timing				
38 Cl	$110 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$	Negligible	None				
$^{36}\mathrm{Cl}$	$110 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$		None				
$^{37}\mathrm{Ar}$	$110 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$		Photon Coincidence Tag				
^{33}P	$34 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$	Negligible	Photon Coincidence Tag				
¹¹ Be	$0.34 \text{ day}^{-1} (10 \text{ kton})^{-1} [59]$	Negligible	Coincident Muon Timing				