

BREAD Gigahertz Pilot

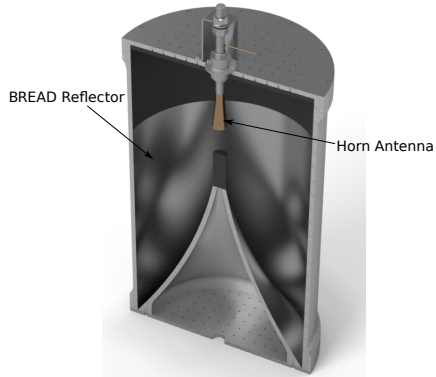
CPAD Workshop 2022

Gabe Hoshino

November 30, 2022

GigaBREAD

- ▶ GigaBREAD is the GHz BREAD pilot designed to look for axions and dark photons in the 10-15 GHz range
- ▶ In the GHz regime, the reflector can be coupled to a microwave horn antenna

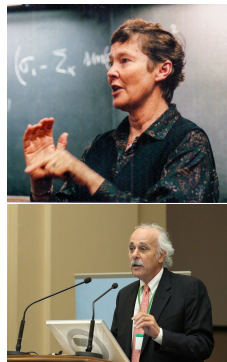


The QCD Axion

- ▶ The strong force is expected to violate CP symmetry via the following term in the QCD Lagrangian:

$$\lambda_{\text{QCD}} \supset \frac{\theta g^2}{32\pi^2} G\tilde{G}$$

- ▶ This would be observable through a neutron EDM, but neutron EDM experiments constrain θ to be very small.
- ▶ A solution, proposed by Helen Quinn and Roberto Peccei, promotes θ to a quantum field which becomes the axion.
- ▶ Couplings between the axion and standard model particles can be feeble, making it a good dark matter candidate.



The BREAD Reflector Concept

- ▶ Axion field modifies Maxwell's equations such that photons are emitted perpendicular to the walls of the outer cylinder in the presence of a strong B -field.
- ▶ A parabolic reflector is placed in the middle of the cylinder to focus the photons onto a point.

Axion induced
 E -field:

$$\mathbf{E}_a = -\frac{1}{\epsilon} g_{a\gamma\gamma} \mathbf{B}_{\text{ext}} a$$

Coupling through
Primakoff Effect:

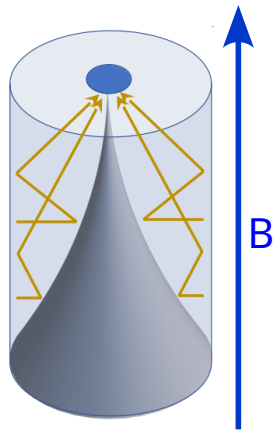
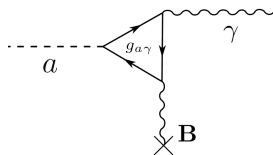


Image from Stefan Knirck

The BREAD Reflector Concept

- ▶ Geometry of the reflector is not very frequency dependent and thus can be in broadband dark photon and axion searches
- ▶ Volume of the detector is not as frequency-dependent as for resonant cavity searches which makes BREAD particularly compelling for higher mass axion searches.

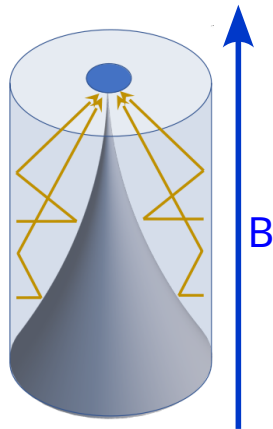
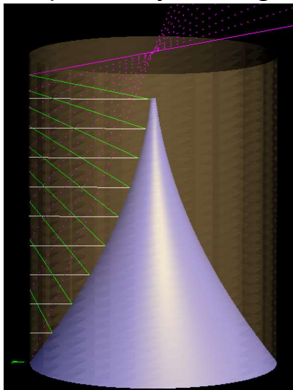


Image from Stefan Knirck

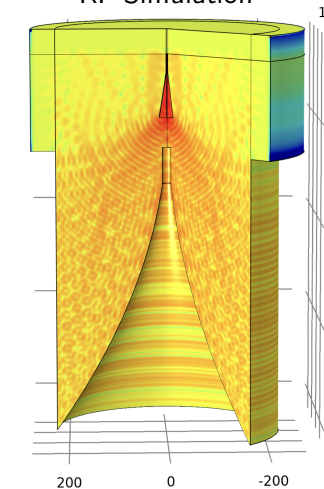
BREAD in Optical and RF Frequency Regimes

Optical Ray-Tracing

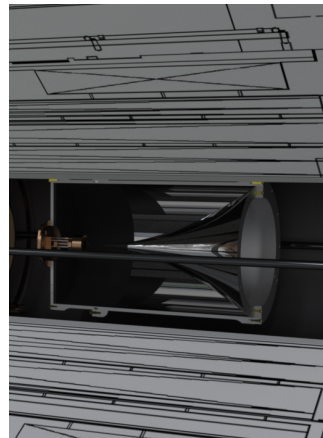


See Christina Wang's talk for more
information on InfraBREAD

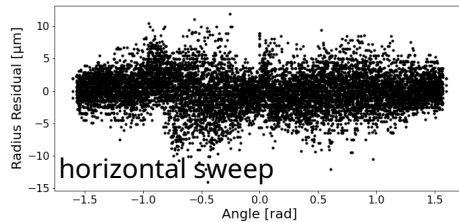
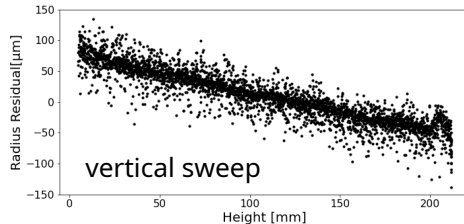
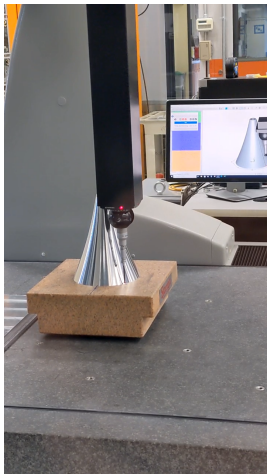
RF Simulation



A Reflector Designed to Fit Inside Large Solenoid Magnets

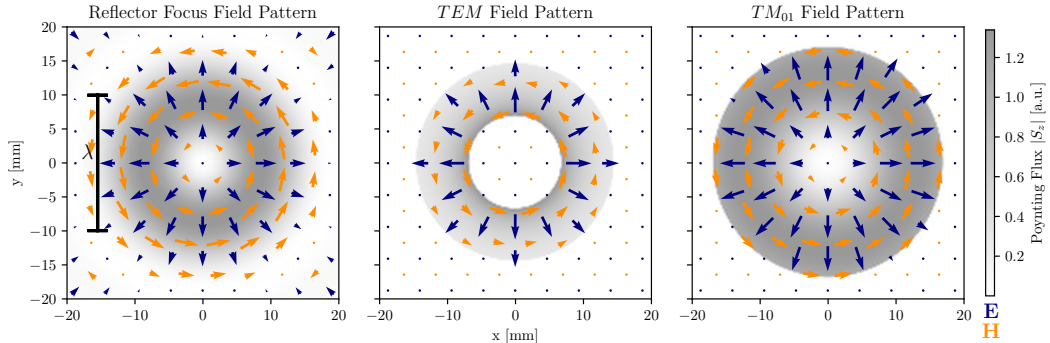


Reflector Surface Characterization with CMM



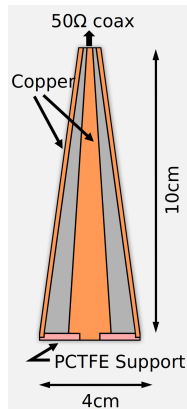
Matching Antenna to the Reflector Beam Shape

- In order to get the best performance, we look for an antenna with a near-field pattern similar to that of the reflector.



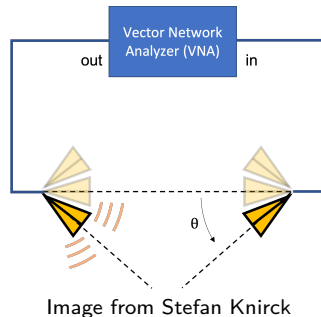
Coaxial Antenna

- Coaxial horn design to be used for GigaBREAD



Far-Field Measurement Setup

- ▶ Horns are mounted on robotic arms in the RF isolation chamber.
- ▶ The robotic arms can be made to rotate together as shown on the right which allows us to determine the far-field transmission at different angles.



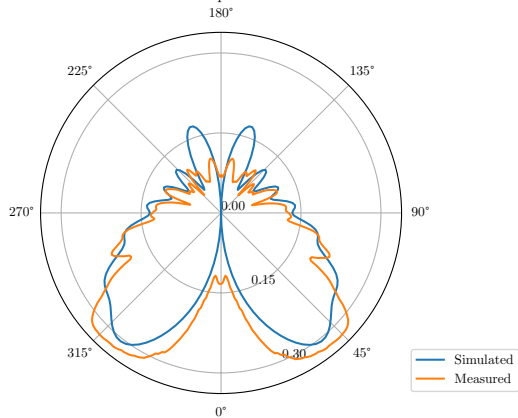
Far-Field Measurement Setup



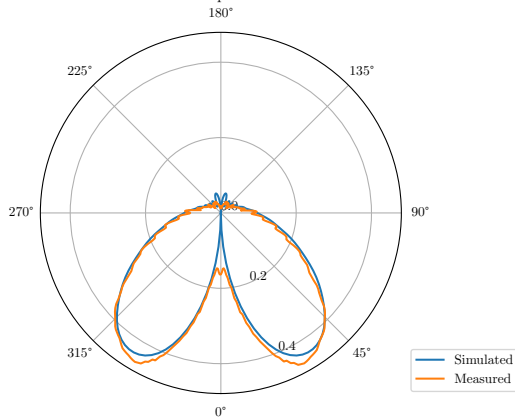
Image from Kristin Dona

Far-Field Measurement Results

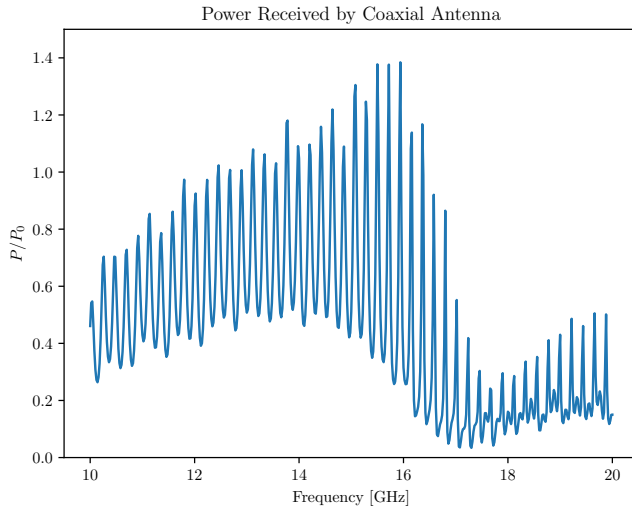
Far-Field Comparison 10 GHz



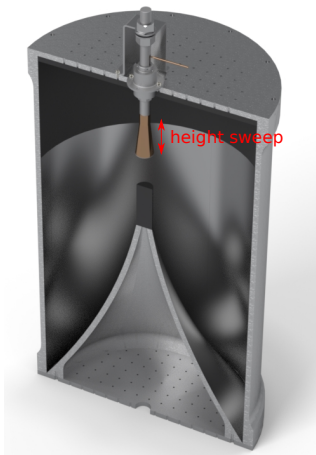
Far-Field Comparison 15 GHz



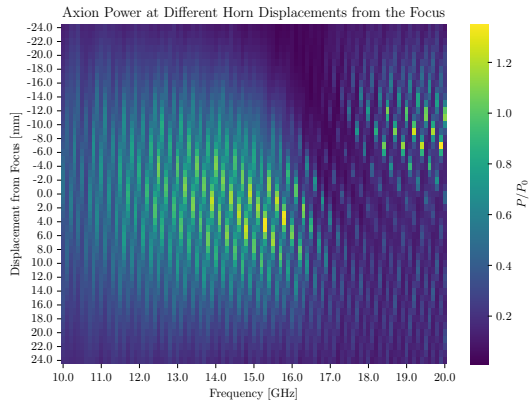
Horn Efficiency



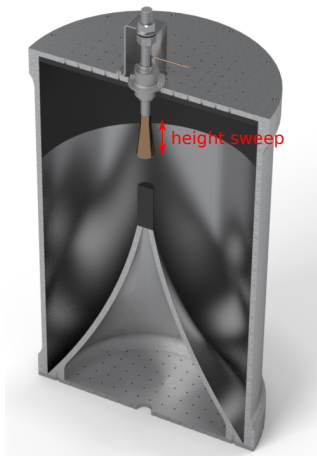
Horn Position Calibration



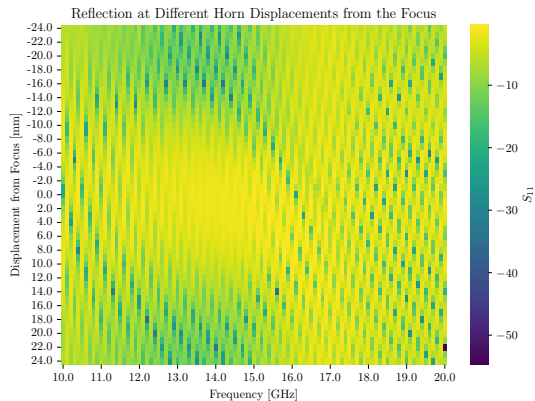
Horn Efficiency with Horn as Receiver



Horn Position Calibration

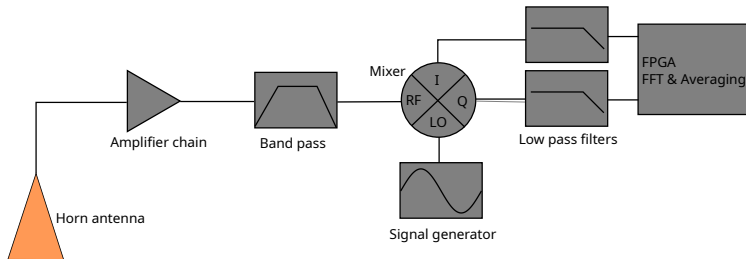


Horn Reflection with Horn as Emitter



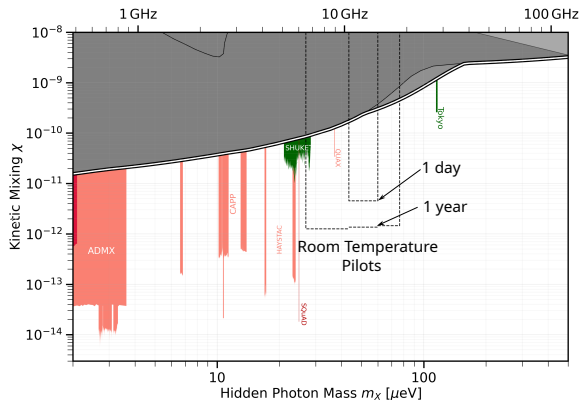
DAQ

- ▶ The signal from the antenna (10-15 GHz) is amplified and mixed down to the 0-3 GHz range used by the Xilinx RFSOC ZCU-111 board.
- ▶ FFT and averaging implemented in FPGA firmware with a python interface



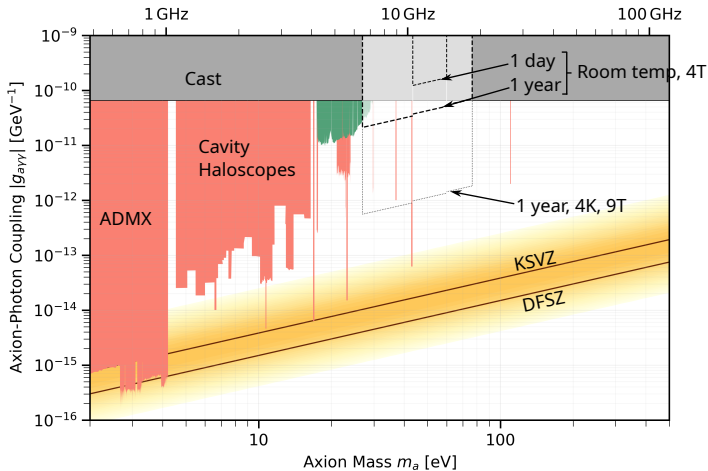
Dark Photon Sensitivity

- ▶ The dark photon kinetically mixes with the standard model photon.
- ▶ The GigaBREAD detector should be sensitive to the dark photon at room temperature without any magnet.



Adapted from cajohare.github.io/axionlimits

Axion Sensitivity



Thanks to the BREAD collaboration!

Pete Barry, Clarence Chang, Juliang Li, *Argonne National Laboratory*

Christina Wang, *Caltech*

Jesse Liu, *University of Cambridge*

Kristin Dona, Gabe Hoshino, Alex Lapuente, David Miller, Max Olberding, *University of Chicago*

Daniel Bowring, Gustavo I Cancelo, Claudio Chavez, Aaron Chou, Mohamed Hassan, Stefan Knirck, Samantha Lewis, Matthew Malaker, Cristian Pena, Andrew Sonnenschein, Leonardo Stefanazzi, Kevin Zvonarek, *Fermilab*

Rakshya Khatiwada, *Fermilab and Illinois Institute of Technology*

Gianpaolo Carosi, *Lawrence Livermore National Laboratory*

Karl Berggren, Dip Joti Paul, Tony (Xu) Zhou, *Massachusetts Institute of Technology*

Omid Noroozian, *NASA Goddard Space Flight Center*

Sae Woo Nam, *National Institute of Standards and Technology*

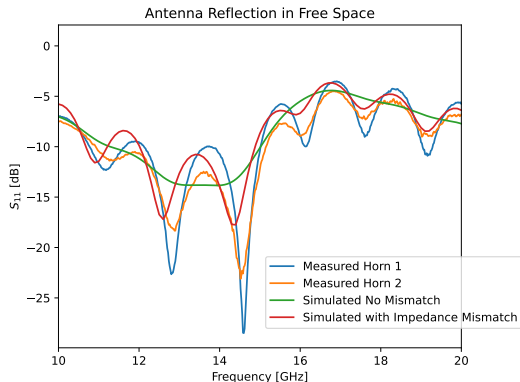
Huma Jafree, *Randolph-Macon College*

Noah Kurinsky, *SLAC*

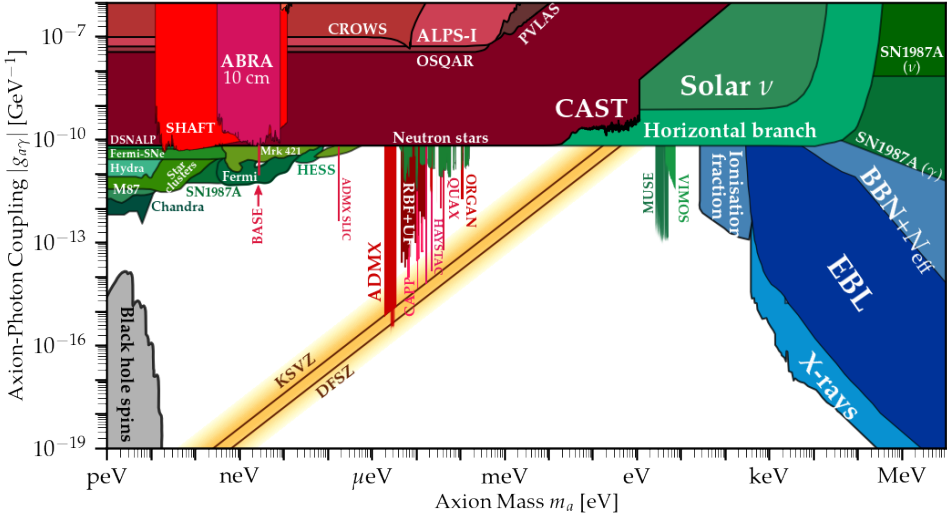
Backup

Free Space Coupling

- ▶ Free space coupling of the horns were measured by pointing each horn toward the foam absorbers in the RF isolation chamber.
- ▶ With an impedance mismatch added between the horn and the connector in simulation, the simulated data resembles what we measure more closely.

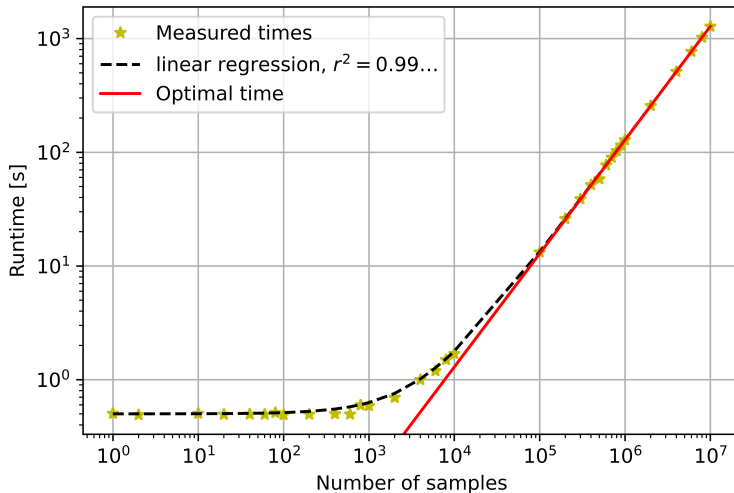


Axion Parameter Space



DAQ Firmware Deadtime

New Firmware Runtime



DAQ Firmware Signal Injection

