

Superconducting Nanowire Single Photon Detectors

CPAD, Stony Brook
29 November 2022

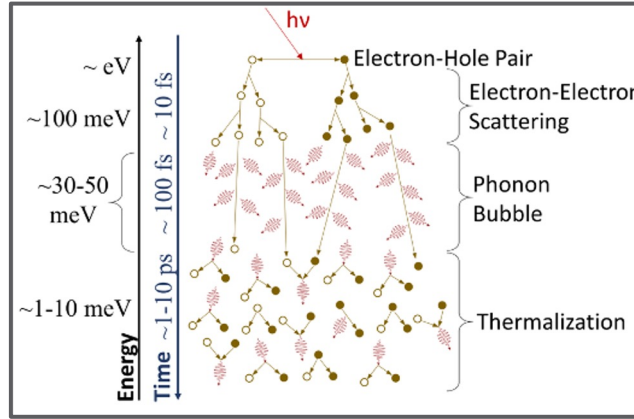
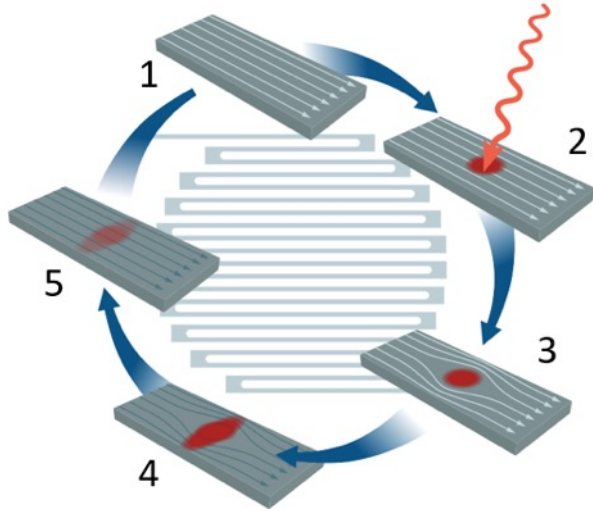
Matt Shaw

Supervisor

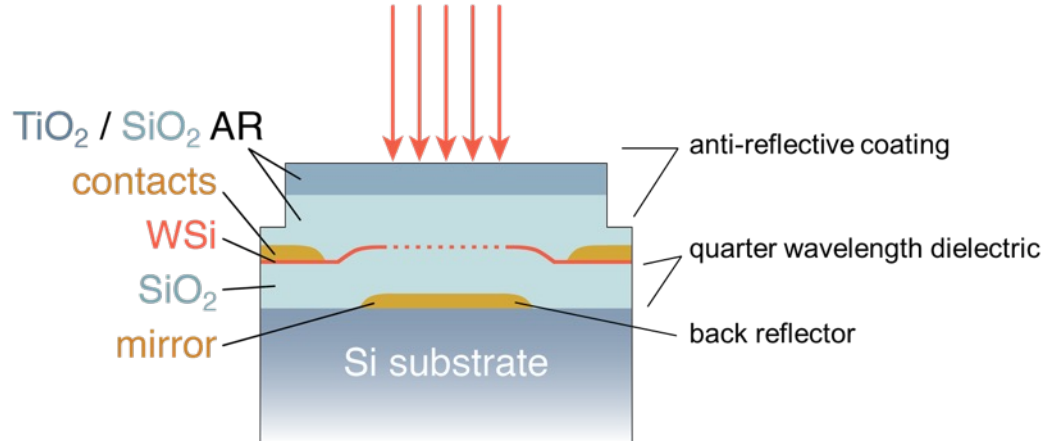
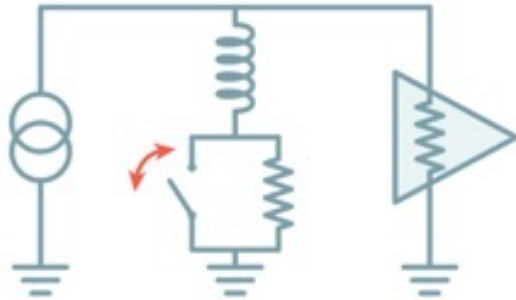
Superconducting and Quantum Devices Group,
Jet Propulsion Laboratory



Superconducting Nanowire Single Photon Detectors



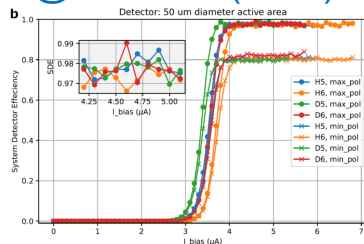
- Time-resolved single photon counting from UV to mid-IR
- World-leading detector performance
- Operating temperature 1-4 K in most cases



Present State of The Art in SNSPDs

High Efficiency

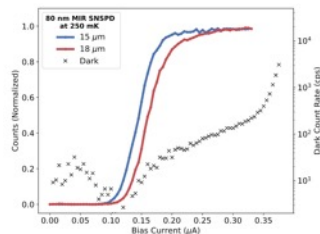
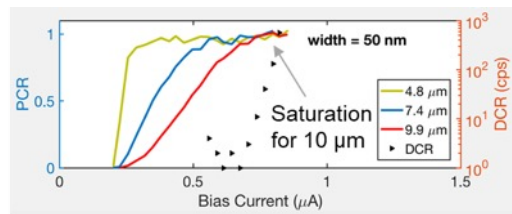
98% SDE @ 1550 nm (NIST)



Reddy et al, *Optica* (2018)

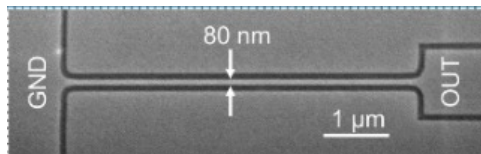
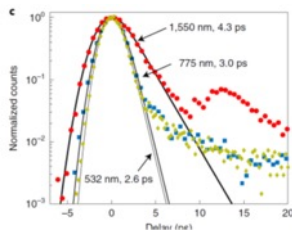
UV – Mid-IR Operation

Photon counting to 18 μm (JPL/MIT/NIST)



High Time Resolution

2.6 ps FWHM (MIT/JPL/NIST)

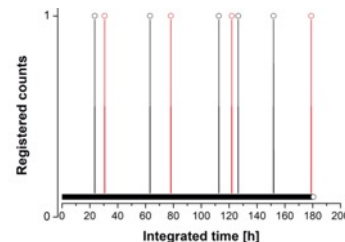
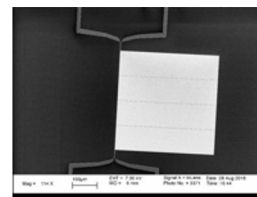


Korzh et al, *Nature Photonics* (2020)

Low Dark Counts

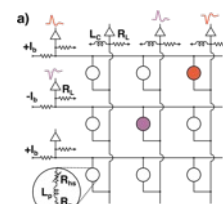
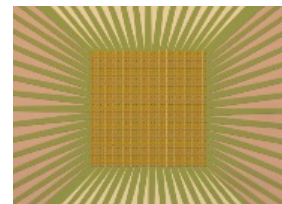
6e-6 cps (MIT/NIST)

Chiles et al, *Phys. Rev. Lett.* (2022)



Kilopixel Array Formats

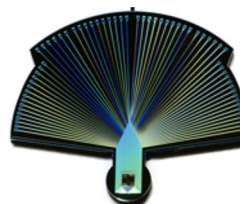
32x32 “row-column” array (NIST/JPL)



High Event Rate

1.4 Gcps in 32-element array (JPL)

Wollman et al, *Optics Express* (2019)



SNSPD Applications

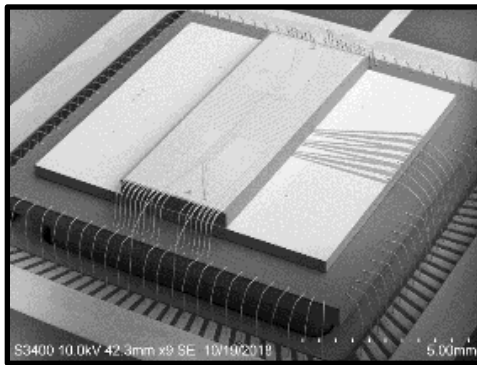
Free-Space Optical Communication

- *Deep Space Optical Communication (Psyche)*
- *Optical-to-Orion*
- *Lunar Laser Comm Demo*
- *Space-to-Ground Quantum Communication*



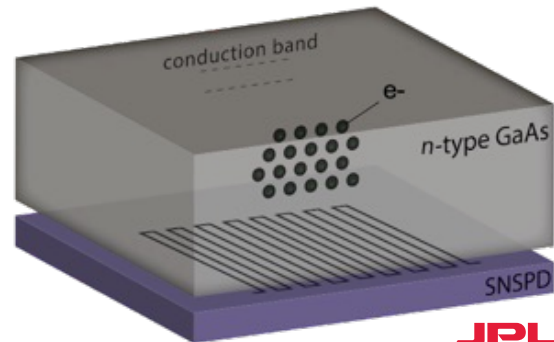
Quantum Information Science

- *Quantum Communication*
- *Trapped Ion Quantum Computing*
- *Linear Optical Quantum Computing*



Fundamental Physics

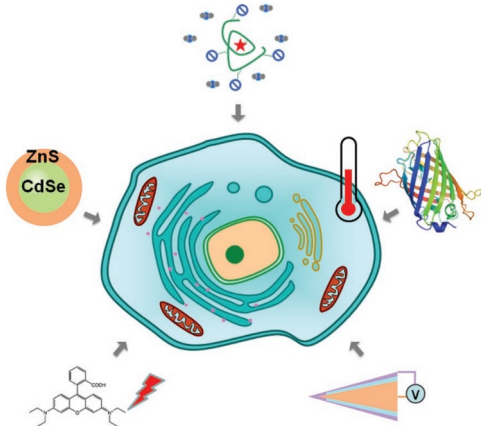
- *Dark Matter searches (scintillation from GaAs, dielectric haloscopes, BREAD)*
- *Tabletop tests of quantum gravity*



SNSPD Applications

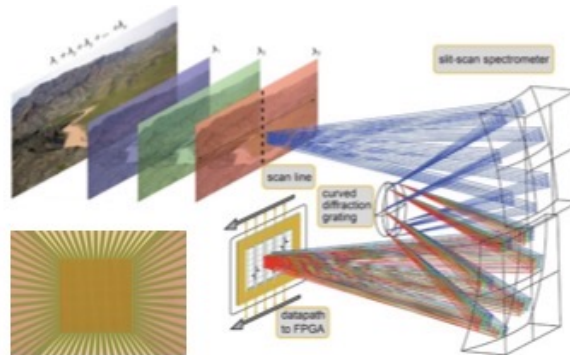
Biomedical

- *Diffuse Correlation Spectroscopy*
- *Fluorescence Lifetime Imaging Spectroscopy*
- *Singlet Oxygen Detection*
- *Thermal microscopy*



Remote Sensing

- *Passive Thermal Ranging*
- *Photon-counting Lidar*
- *Time-resolved Raman spectroscopy*
- *Entangled-photon spectroscopy*



Astronomy and Heliophysics

- *Exoplanet transit spectroscopy (IR)*
- *Interferometry for exoplanets (IR)*
- *Direct exoplanet imaging (vis/IR)*
- *UV astronomy*
- *Optical counterparts of fast radio bursts (NIR)*
- *Metastable helium lidar (NIR)*

SNSPD Advantages for Fundamental Physics

Dark Matter Detection

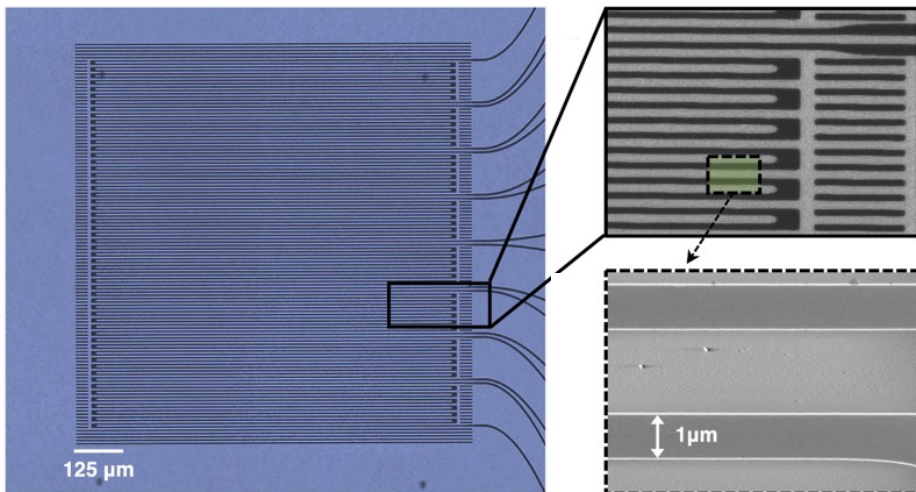
- Low dark counts (10^{-5} cps)
- Low energy threshold (70 meV)
- Large active area ($\text{mm}^2 \rightarrow \text{cm}^2$)

Tests of Quantum Gravity

- Low dark counts (10^{-5} cps)
- High efficiency (98% @ $1.5 \mu\text{m}$)
- Photon number resolution (1, 2, or many)

Nuclear Physics and Collider Physics

- High time resolution (3 ps)
- Low dark counts (10^{-5} cps)
- Radiation hardness



Technology Development for Dark Matter

Active Area

- Currently $\sim\text{mm}^2$ area
- Targeting cm^2 area and beyond
- Fab process development and readout

Dark Counts

- Currently at $<10^{-5}$ cps at 1550 nm
- Physical origin of dark counts is not well understood
- Limits are not well understood as energy threshold decreases

Energy Threshold

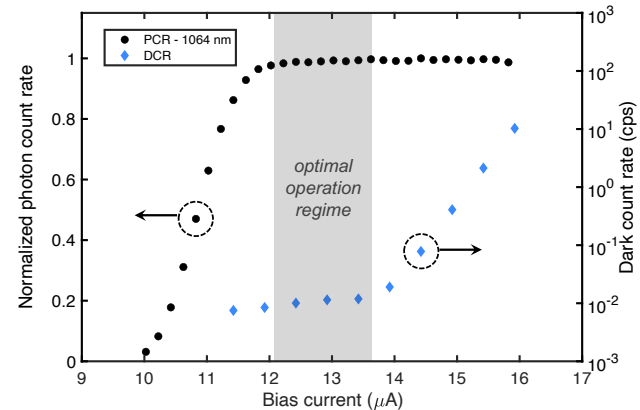
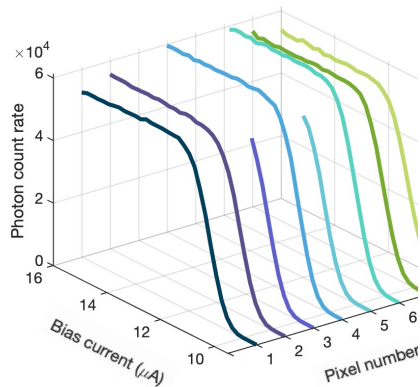
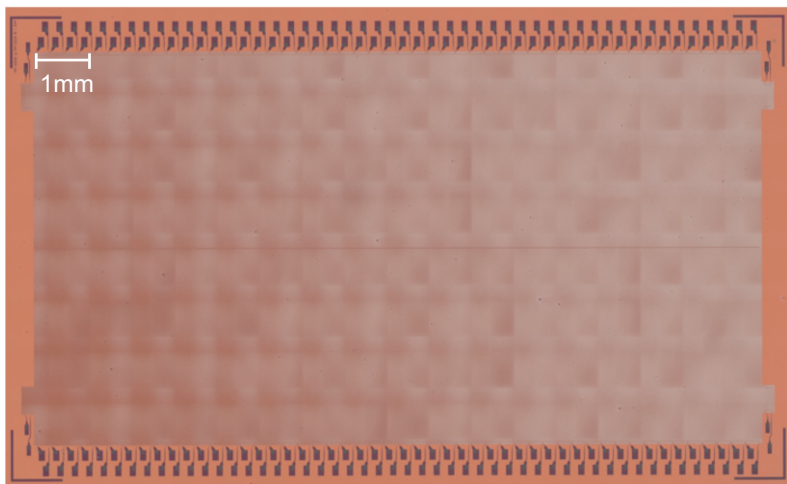
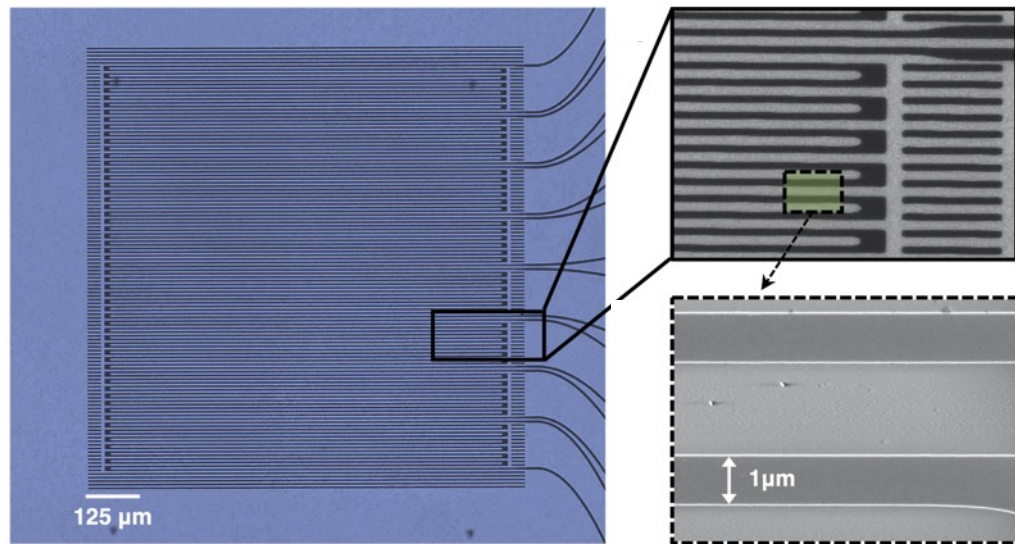
- Currently at 18 μm (70 meV)
- Likely possible to reach 30 μm (40 meV) and possibly lower
- Exploring strategies to efficiently couple (antennas, dielectric stacks, photonics)

UV Operation

- Characterization and optimization of SNSPDs at VUV energies
- Currently preparing cryostat for SNSPD measurements at NIST SURF synchrotron (3 – 300 eV photons, ~ 1 ns pulses)

Scaling to Larger Areas

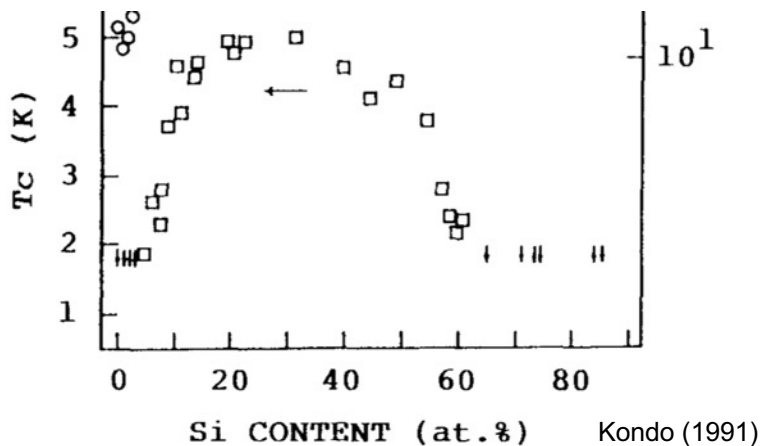
- Currently fabricating mm^2 and cm^2 SNSPD arrays
- Micron-wide wires enable larger area with photolithography
- Investigating multiple approaches to multiplexing many pixels
- Frequency-domain, thermal coupling, row-column readouts, SFQ



Strategies for Lower Energy Thresholds

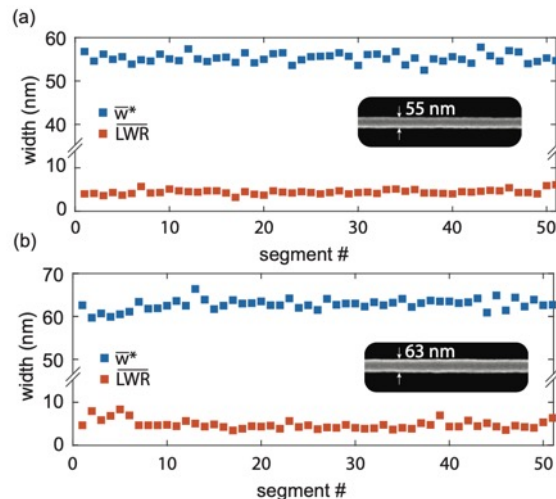
Reduced Superconducting Gap Energy

- Now using Si-rich WSi to reduce T_c to 1.3-2.1 K (depending on thickness)
- “Conventional” WSi for NIR devices has $T_c = 3.1 - 3.6$ K



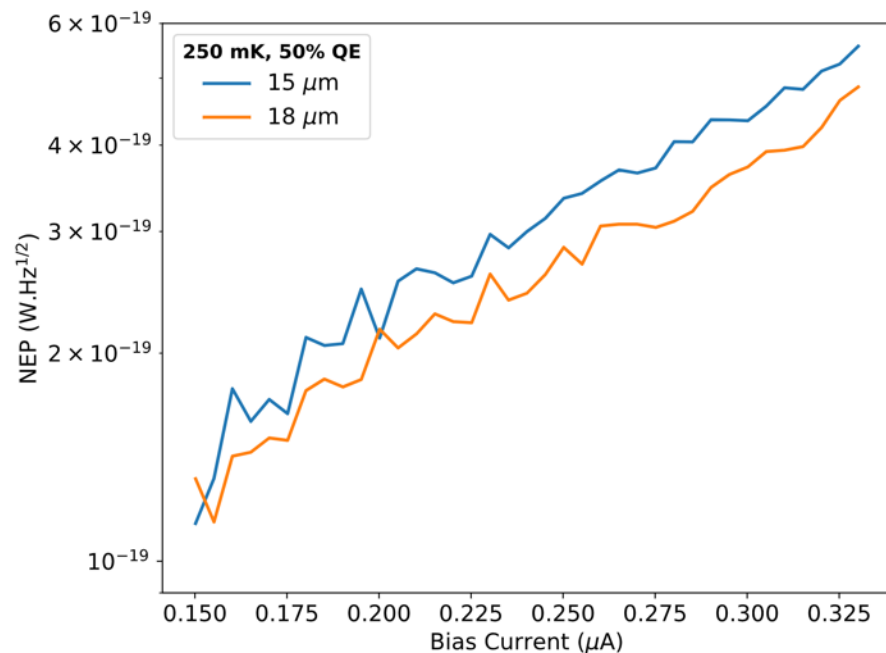
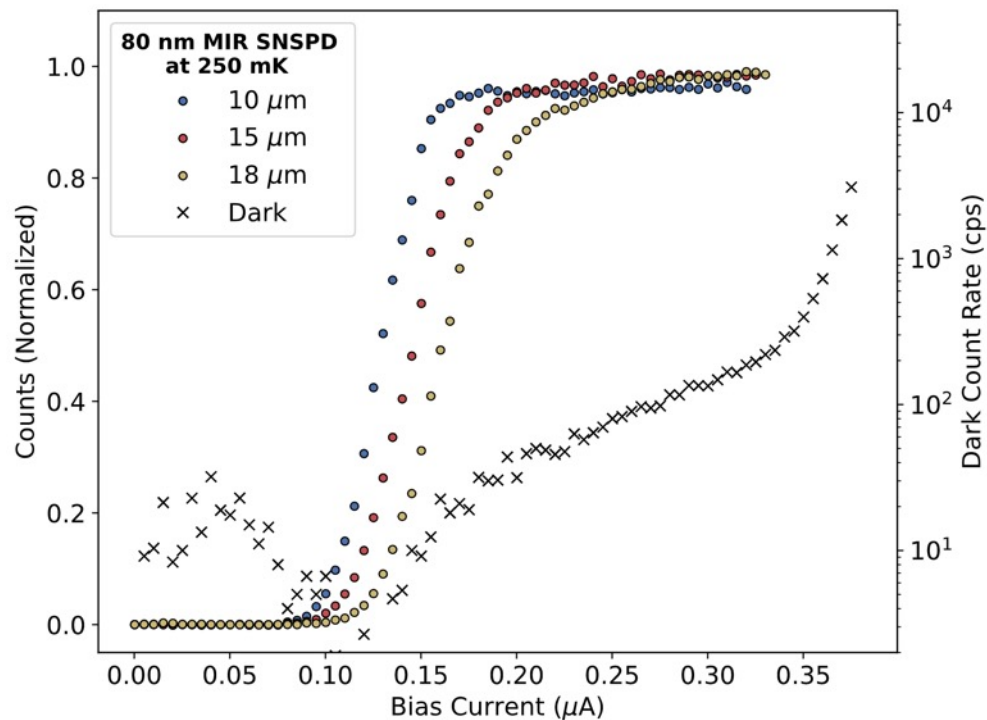
Narrow Nanowires

- Narrower nanowires enhance IR sensitivity by constraining hotspot growth
- Reliably fabricating SNSPDs with 50-60 nm wires using electron-beam lithography



Tradeoffs: Lower operating temperature (< 1 K) and smaller readout currents (< 2 μ A)

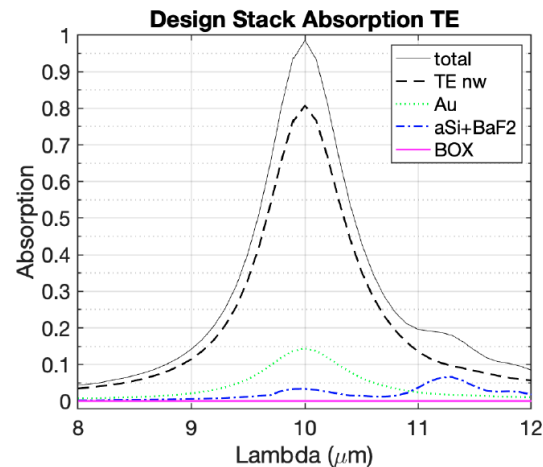
Saturated Internal Efficiency from 10 - 18 μm



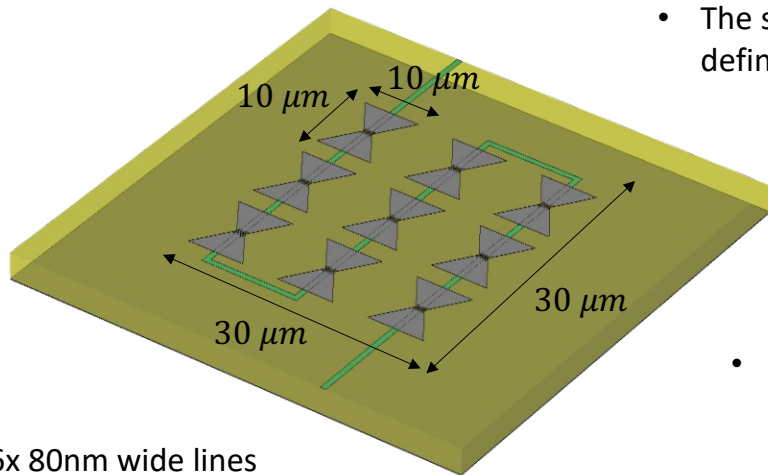
WSi SNSPD nanobridge with 3 nm thickness: $T_c = 1.3$ K

Approach to Efficient Coupling: Optical Stacks

- Optical stack approach has proven very effective in the NIR (98% system detection efficiency observed at NIST)
- SNSPDs do not face the same restrictions on amorphous dielectrics as KIDs (TLS noise)
- Conventional materials will work to $\sim 7\text{ }\mu\text{m}$
- Adapting the approach for wavelengths as long as $18\text{ }\mu\text{m}$ requires careful selection of dielectric materials
- Investigating Ge, a-Si, c-Si, BaF_2 , CdTe/PbSe as low-loss candidates at long wavelengths



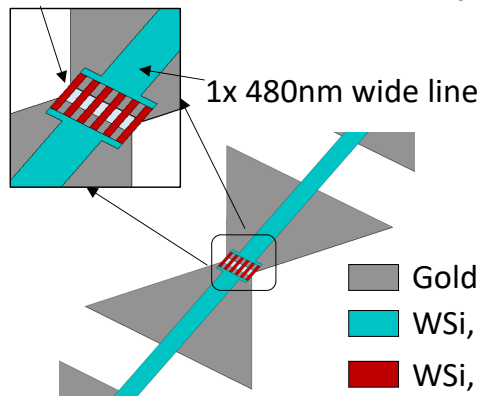
Approach to Efficient Coupling: Dipole Antennas



- The structure is illuminated by a plane-wave at broadside incidence. It is defined with unitary amplitude $|E_x| = 1 \text{ V/m}$
- The power incident on the 3×3 active area ($30 \mu\text{m} \times 30 \mu\text{m}$) is:

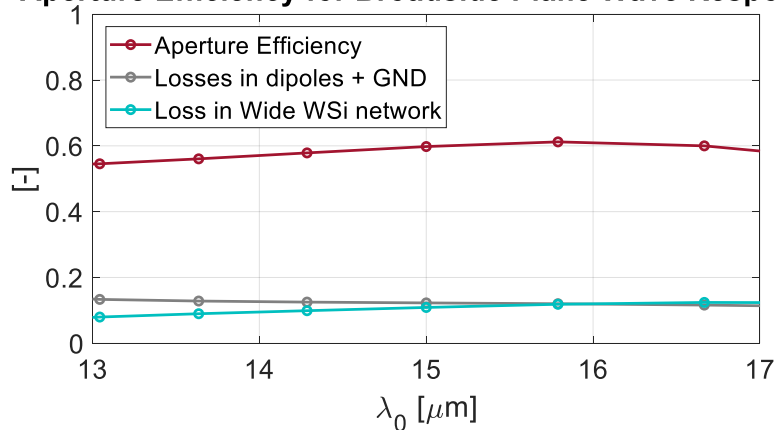
$$P_{inc}^{PW} = \frac{1}{2\zeta_0} |E_x|^2 A_{active}$$
- The aperture efficiency is defined as: $\eta_{ape} = \frac{P_{inc}^{PW}}{P_{abs}^{WSi, Active}}$
- The loss is the power dissipated in the antennae or the wide WSi network.

6x 80nm wide lines



- Gold ($\sigma = 2 \cdot 10^7 \text{ S/m}$ is assumed)
- WSi, $R_s = 1.1 \text{ k}\Omega/\text{sq}$
- WSi, $R_s = 1.1 \text{ k}\Omega/\text{sq}$, Active area

Aperture Efficiency for Broadside Plane Wave Response



Conclusions

- *SNSPDs are the highest performing detectors available for time-resolved single photon counting, from the UV to the Infrared*
- *SNSPDs are rapidly advancing with significant room for further gains*
- *SNSPDs can be a powerful new tool for fundamental physics*

