Magnetic Field at BES Energies

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Contents

Introduction and motivation

- Magnetic filed evolving in vacuum
- > Magnetic field in conductive medium
- > Impact of the magnetic field

> Summary

Introduction

• Strong Magnetic field



Strongest magnetic field

RHIC: $eB_y \sim 10^{18}$ Gauss @ 200 GeV LHC: $eB_y \sim 10^{19}$ Gauss @ 2760 GeV

- Physics:
 - Hyperon spin polarization
 - Vector meson spin alignment
 - Anomalous transport, e.g. CME, CMW...
 - Breit-Wheeler process $\gamma \gamma \rightarrow l^+ l^-$ and photon polarization
 - Direct flow of D^0 meson
 - ...

Question:

Can we give some theoretical constrains on the magnetic field at the QGP evolution stage?

Review on magnetic field

For a moving point charge

Lienard-Wiechert formula:

$$\mathbf{E}(t, \mathbf{x}) = \frac{q}{4\pi} \frac{\gamma_q \mathbf{R}}{[R^2 + (\gamma_q \mathbf{v}_q \cdot \mathbf{R})^2]^{3/2}},$$
$$\mathbf{B}(t, \mathbf{x}) = \frac{q}{4\pi} \frac{\gamma_q \mathbf{v}_q \times \mathbf{R}}{[R^2 + (\gamma_q \mathbf{v}_q \cdot \mathbf{R})^2]^{3/2}},$$

For heavy-ion collisions:

- Charge distribution:
 - Wood-Saxon distribution
 - Glauber sampling
- After collision:
 - Unwounded nucleons keep moving
 - Wounded nucleons slow down (stopping effect)

 $\mathbf{R} = \mathbf{x} - \mathbf{x}_q$

- Full transport model or empirical formula
- Magnetic field evolves in:
 - Vacuum
 - Conductive medium





Skokov et al. 2009 Deng, Huang 2012 Tuchin 2013 McLerran, Skokov 2014 Gursoy, Kharzeev, Rajapopal 2014 Li, Sheng, Wang 2016

...

Magnetic field in vacuum

• Magnetic field drops rapidly in vacuum



In vacuum, and suppose nucleons are unwounded
→ symmetric evolution w.r.t. t = 0

Stopping effect

Simulate the EM field in vacuum by transport models HSD, UrQMD, ...



- It needs time to propagate the signal of nucleon's stopping (retarded potential)
- The stopping effect shows up only after t > 1 fm/c

V. Voronyuk, et.al, Phys. Rev. C 83, 054911

Electric conductivity of QGP medium

• Lattice calculations (T_c is critical temperature)

$$\sigma = (5.8 \pm 2.9) \frac{T}{T_c} \text{MeV},$$

• Transport simulation by BAMPS

$$\sigma = \frac{1}{3T} \tau \sum_{i} q_i^2 n_i \qquad \tau = \frac{3}{2} \frac{1}{\sum_{i} n_i \sigma_{22}} \qquad \sigma \propto \frac{1}{\sigma_{22}}$$

where au is the relaxation time for deviation from equilibrium

In the case of T = 255 MeV,

$\sigma_{22} = 1 \text{ mb}$	$\sigma = 11.6$ MeV
$\sigma_{22} = 2 \text{ mb}$	$\sigma=5.8$ MeV

It is small compared to typical QCD scale of ~200 MeV

But it has obvious effect on the lifetime of EM field.

H. T. Ding, et.al, Phys. Rev. D 83, 034504 Z. Y. Wang, et.al, Phys. Rev. C 105, L041901

Magnetic field in conductive medium

Analytical formula:



- $\succ \sigma = \text{const} \text{ at both } t > 0 \text{ and } t < 0 \text{ (not realistic)}$
- \blacktriangleright Delay the field increasing at t < 0 (not realistic)
- > Delay the field decreasing at t > 0 (expected, but over-estimated)

BAMPS simulation

- BAMPS simulates the response of quark matter to EB field, and solves Maxwell's equations.
- Significant suppression of the magnetic filed compared to the analytical formula.



Z. Y. Wang, J. X. Zhao, C. Greiner, Z. Xu, P. F. Zhuang, Phys. Rev. C 105, L041901

Solve Maxwell equation numerically



- Numerical solution of Maxwell's equations
- \succ σ is switched on at t = 0
- > Effect of σ is not obvious at early time (t < 1 fm/c) for $\sigma = 5.8$ MeV
- L. McLerran, V. Skokov, Nucl. Phys. A 929 184-190
- B. G. Zakharov, Phys. Lett. B, 737 262-266

Our model setup

Before collision (t<0):

Calculate the EM field in vacuum by L-W formula

After collision (t>0):

Numerically solve Maxwell's equations with non-zero σ

 $\nabla \cdot \mathbf{E} = \rho,$ $\nabla \cdot \mathbf{B} = 0,$ $\nabla \times \mathbf{E} = -\partial_t \mathbf{B},$ $\nabla \times \mathbf{B} = \mathbf{j} + \sigma \mathbf{E} + \partial_t \mathbf{E},$ We consider two cases

> $\sigma = \sigma_0 \theta(t)$ $\sigma = \sigma_0 \theta(t) / (1 + t/t_0)^{1/3}$

In both, σ is non-zero only after collision

woods-Saxon distribution f(r) = $\frac{\rho_0}{1 + e^{[(r-R)/a]}}$ h non-zero σ $\rho^{\pm} = \gamma f \left(\sqrt{(x \pm b/2)^2 + y^2 + \gamma(z \pm v_z t)^2} \right)$ $j_x^{\pm} = j_y^{\pm} = 0$ $j_z^{\pm} = \pm \gamma v_z f \left(\sqrt{(x \pm b/2)^2 + y^2 + \gamma(z \pm v_z t)^2} \right)$

Simplification:

Nucleons fly freely without stopping

Analytical vs. Numerical



- > Effect of σ is not obvious at early time (t < 1.5 fm/c)
- > At early time, the analytical formula over estimates the field
- At very late time (t > 10 fm/c), it agrees with the numerical solution,
- but the time is too late (the field has became very small)

Analytical vs. Numerical



The analytical formula over estimates the field at all energies

Compare with BAMPS calculation



- Numerical calculation agrees with BAMPS results (apple-to-apple comparation)
- Ohm's law is valid in heavy ion collisions

Z. Y. Wang, J. X. Zhao, C. Greiner, Z. Xu, P. F. Zhuang, Phys. Rev. C 105, L041901

Magnetic field with $\sigma = \sigma_0 \theta(t)$



- At late time, the field is sensitive to the σ value, and less sensitive to v_z
- We have used the same σ value at different energies
- For t > 4 fm/c, the field is at the order of $10^{-3} \sim 10^{-2}$ ($\sigma = 5.8$ MeV)

Magnetic field with $\sigma = \sigma_0 \theta(t)/(1 + t/t_0)^{1/3}$



- The field also depends on σ 's time behavior.
- When necessary, the numerical method can be applied to more accurate model.

Impact of the magnetic field

Effects at different time stage:



17

Global A polarization







Note: Λ and $\overline{\Lambda}$ flow is different at low energy

Spin anti-alignment

Quark and anti-quark are anti-polarized by magnetic field

• ρ_{00} with respect to different directions:

$$\rho_{00}^{(x)} = \frac{1 - P_x^q P_x^{\bar{q}} + P_y^q P_y^{\bar{q}} + P_z^q P_z^{\bar{q}}}{3 + \mathbf{P}^q \cdot \mathbf{P}^{\bar{q}}}$$
$$\rho_{00}^{(y)} = \frac{1 - P_y^q P_y^{\bar{q}} + P_x^q P_x^{\bar{q}} + P_z^q P_z^{\bar{q}}}{3 + \mathbf{P}^q \cdot \mathbf{P}^{\bar{q}}}$$
$$\rho_{00}^{(z)} = \frac{1 - P_z^q P_z^{\bar{q}} + P_x^q P_x^{\bar{q}} + P_y^q P_y^{\bar{q}}}{3 + \mathbf{P}^q \cdot \mathbf{P}^{\bar{q}}}$$



Y. Liu, C. Greiner, C. M. Ko, arxiv: 1403.4317 Y. G. Yang, et al, Phys. Rev. C 97, 034917

X. L. Xia, et al, Phys. Lett. B 817, 136325

 ρ_{00} is effect of $\langle P_q P_{\bar{q}} \rangle$, instead of $\langle P_q \rangle \langle P_{\bar{q}} \rangle$

- $\langle P_q \rangle$ and $\langle P_{\overline{q}} \rangle$ are small (can be constrained by Λ polarization)
- $\langle P_q P_{\bar{q}} \rangle$, arising from fluctuation of magnetic/vorticity fields, may be not small

Our results do not rule out the magnetic-induced spin anti-alignment

$$\langle B_y \rangle = 0$$
 does not put any limit on $\langle B_x^2 \rangle$, $\langle B_y^2 \rangle$, and $\langle B_z^2 \rangle$

Summary

We have numerically solved Maxwell's equations in conductive medium,

• The previous analytical formula over-estimated the magnetic field.

From the numerical solution,

- At early time: the field is similar with that in vacuum;
- At late time: the field decreases slower, but it is already very small.

It is important to distinguish effects at different time-stage.

- Can not lead to observable global Λ polarization
- Does not rule out spin (anti-)alignment
- For CME, depends on early time dynamics
- For direct flow of charged particle, more accurate model is needed

Thank you for attention!