



RHIC Beam Use Request For Runs 17 and 18

The STAR Collaboration



Zhangbu Xu (Brookhaven National Lab)

- Highlights
- Performance in run16
- BUR on Spin (sign change)
- BUR on Isobars (chiral effect)
- BUR on Au+Au @ 62GeV (Jet, NPE)
- BUR on Au+Au @ 27 GeV (global Lambda polarization)
- Summary

https://drupal.star.bnl.gov/STAR/starnotes/public/sn0657



May 26, 2016



5.1.3

5.2 AU+AU COLLISIONS AT 27 GEV

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BUR charges

From: Mueller, Berndt Sent: Tuesday, February 23, 2016 9:12 PM To: Xu, Zhangbu; Morrison, David; Gunther M Roland; Yasuyuki Akiba; Takashi SAKO Cc: Dunlop, James C Subject: 2016 Beam Use Request

Dear Zhangbu, Takashi, Yasuyuki, Dave & Gunther:

I am writing to solicit the STAR beam use request for RHIC Run-17 and to request presentations from the RHIC collaborations at this year's PAC meeting, which is scheduled to be held on June 16+17, 2016.

Following the recommendations of the 2015 PAC, we anticipate the combination of a 510 GeV polarized p+p run and a 200 GeV run of a suitable isobar pair. Following current guidance from DOE, we are planning for a 24 cryo-week run, but the beam use request should also consider options for shorter runs of 20 weeks and 15 weeks to allow flexibility in case of other budget scenarios. The beam use request should be submitted no later than May 27 in order to allow the PAC members to study it in detail before the meeting.

In addition to the STAR/ from Runs 14 and 15, in detectors. This is one of update from the PHENIX findings to the PAC.

Best regards Berndt

Future RHIC Run Plans: The plan for RHIC runs before the BES-II has recently been refined to include independent runs in 2017 and 2018. The change is driven by the desire to permit a sufficiently long run with transversely polarized p+p collisions at 510 GeV in sPHENIX collaboration (Run-17 (up to 19 cryo-weeks depending on budgetary constraints) to "test unique QCD predictions for relations between single-transverse spin phenomena in p-p scattering and those observed in deep-inelastic scattering" (NP Milestone HP13), and the plan to collide isobars (⁹⁶Zr+⁹⁶Zr and ⁹⁶Ru+⁹⁶Ru) at 200 GeV in Run-18 (13 cryo-weeks) as a critical test of the contribution from the possible Chiral Magnetic Effect to the various observed charge separation effects.

Executive Summary Table

Run	Energy	Duration	System	Goals	priority	Sequence
	√s _{NN} =500 GeV	13-wk	Transverse p+p	A_N of W [±] , γ , Drell-Yan, L=400 pb ⁻¹ , 55% pol	1	1
		1-wk	p+p	RHICf		2
17		2-wk	CeC			
	$\sqrt{s_{NN}}$ =62.4 GeV	4-wk	Au+Au	Jets, dileptons, NPE 1.5B MB	3	3
18	√s _{NN} =200 GeV	3.5-wk	Ru+Ru	1.2B MB	2	4
	√s _{NN} =200 GeV	3.5-wk	Zr+Zr	1.2B MB	2	5
	√s _{nn} =27 GeV	2-wk	Au+Au	>500M MB	3	6

Options from guidance:

- 1) 24 cryo-weeks in run 17, 13 weeks in run 18
- 2) 19 cryo-weeks in run 17, 13 weeks in run 18
- 3) If only 15 weeks in run 17, all for pp500

STAR Detector System_{15 fully functioning detector systems}



X10³ increases in DAQ rate since 2000, most precise Silicon Detector (HFT 2014-16)

Highlights of Recent Results

Jets
Heavy-Flavor
Dileptons
V_n
Spin

Experimental overview on run14/15: Frank Geurts (Rice) BES-II (I): Helen Caines (Yale)

CME Task Force: Cold QCD Plan: Paul Sorensen (BNL) Carl Gagliardi (TAMU)

y-jet correlation



- Golden probe of Jet Energy Loss calibrated energy from γ for FF (Z_T)
- Cover large range of Z
- γ and π^0 triggered hadron I_{AA} similar
- Suppression independent of jet energy
- Lost energy reappear at low p_T



Semi-inclusive hadron jets



Quarkonia production and probes



0

3. Beam Energy dependence of $J/\psi R_{AA}$

√s_{NN} (GeV)

10³

10²

(STAR QM15) Penetrating Probes



- Heavy Flavor Tracker (HFT) delivers its first results
- First result of quarkonia suppression from the Muon Telescope Detector (MTD)
- Charm flows at RHIC top energy
- Extracted diffusion coefficient compared to theory

Low-mass di-electron production

- Measured in many systems (Au+Au, U+U, p+p) and different energies (19.6, 27, 39, 62, 200 GeV)
- Quantifying how vector mesons evolve in the medium
- The yields probe timescale of collisions



Flows in nth order



Flows in Rapidity



Flow in asymmetric collisions sensitive to the existence of E-field and quark content Flow (vn) vs rapidity: η/s and initial stage; baryon stopping at BES-II (require upgrades)



NSAC Milestone (HP12)



NSAC Milestone (HP8)



2015 BUR -> RUN 16

Run	Energy	Duration	System	Goals	priority	Sequence
16	√s _{NN} =200 GeV	13-wk	Au+Au	Λ_c , D v ₂ , R _{AA} , Y R _{AA} 10nb ⁻¹ , 2billion MB	1	1
	√s _{NN} =62 GeV	4-wk	Au+Au	1.5B MB (1B w/ HFT)	4	2
	√s _{NN} =19.6 GeV	1-wk	d+Au	100M MB	4	3
17	√s = 510 GeV	11 wk	Transverse p+p	A_N of W [±] , γ , Drell-Yan, L=360 pb ⁻¹ , 55% pol	2	1
	$\sqrt{s_{NN}}$ =19.6 GeV	1-wk	p+p	400M MB	4	2
	$\sqrt{s_{NN}} = 200 \text{ GeV}$	3-wk	Ru+Ru	1.2billion MB	3	3
	$\sqrt{s_{NN}}$ =200 GeV	3-wk	Zr+Zr	1.2billion MB	3	4

In each scenario, the data requirements for STAR's two top priority scientific goals will be met. In no scenario can all the data needs for all programs be met in runs 16 and 17

Run 16 Heavy-Flavor Program - Completion of DM12

2016	DM12 (new)	Measure production rates, high p_T spectra, and correlations in heavy-ion collisions at $\sqrt{s_{NN}} = 200$ GeV for identified hadrons with heavy flavor valence quarks to constrain the mechanism for parton energy loss in the quark-gluon plasma.
	()	valence quarks to constrain the mechanism for parton energy loss in the quark-gluon plasma.

increase in RHIC luminosity that is part of the RHIC luminosity upgrade and associated detector upgrades to study rare particles with charm quarks, and possibly particles with bottom quarks, as a demanding way to learn how matter flow and energy loss are established in the partonic phase at RHIC.

Au-Au 200 GeV Highest Priority of STAR and PAC for Run-16

DOE Milestones for High Temperature/High Density Hadronic Matter <u>http://science.energy.gov/~/media/np/nsac/pdf/docs/perfmeasevalfinal.pdf</u> DM12 uses the

Run 16 proposed by STAR

Run	Energy	Duration	System	Goals	Priority	Sequence
16	√s _{NN} =200 GeV	13-wk	Au+Au	Λ_c , D v ₂ , R _{AA} , Y R _{AA} 10nb ⁻¹ ,2billion MB	1	1
	√s _{NN} =19.6 GeV	1-wk	d+Au	100M MB	2	2
	√s _{NN} =39 GeV	1-wk	d+Au	400M MB	2	3

PAC recommendation of 10 weeks of Au+Au; additional 2 floating weeks

Could have reached goals: 95 hours*10wk*3600seconds*600Hz=2B Or we could have just scaled down our goals by 30%

RHIC Machine Efficiency not Luminosity key STAR Operation and Optimization - key to achieving this goal.

Presented at 01/19/2016 schedule meeting

Run16 schedule and main events

Run 16 plan based on 23.5 weeks cryo operation

and Fischer et.al. RHIC Collider Projections (FY 2016 - FY 2022), 19 April 2015

Today, June 15th

- 19 Jan, Begin cool-down to 4.5K
- 25 Jan, Beam in Yellow
- 22 26 Jan, Beam in Blue
- 29 Jan, Feb 3, First Collisions
- 5 7 Feb, Begin <u>10 week</u> **√s=200 GeV/n AuAu** physics run
- 7 am March 18th, RHIC Operations halted for Blue ring Diode issue
- 7 pm April 6th, RHIC Operations resumed (19.5 days offline)
- Decision made to add 19.5 days to the AuAu running to account for the Diode related down time
- 9 May, End 10 week (+ 2 days) vs=200 GeV/n AuAu physics run
- 12 May, Begin <u>1 week</u> Vs=200 GeV/n dAu physics run
- 20 May, End ~1 week Vs=200 GeV/n dAu physics run
- 21 May, Begin <u>1 week</u> Vs=62 GeV/n dAu physics run
- 27 May, End ~1 week Vs=62 GeV/n dAu physics run
- 28 May, Begin 1.5 week Vs=19.6 GeV/n dAu physics run
- 8 June, End 1.5 week Vs=19.6 GeV/n dAu physics run
- 10 June, Begin <u>1 week</u> **vs=39 GeV/n dAu** physics run
- 17 June, End 1 week Vs=39 GeV/n dAu physics run
- 19 June, Begin <u>~7 day</u> Return to 200 GeV/n AuAu physics run
- 27 June, End of 200 GeV/n AuAu physics run
- 27 June, begin cryo warm-up
- 1 July, Cryo warm-up complete, <u>23.5 cryo weeks</u> of operation

- February 7, start Au+Au physics run
- March 4, switching to run14 HFT firmware
- March 17, new online vertex implemented
- May 9, end of 1st 200GeV Au+Au run 75% goal reached for HFT (open charm) 90% goal reached for MTD (Quarkonia)
- May 12, begin d+Au Run
- June 19, begin 2nd Au+Au datataking

Schedule in Orange text not updated yet

See http://www.rhichome.bnl.gov/AP/RHIC2016/ for the Run Coordinator's detailed plan

Substantial Investments to reach our BUR goals

-- Many improvements since Run14 Open Charm (HFT) related: (MB events)

- **Cables:** Cu => AI Cable for HFT readout: up to x2 better S/B low- $p_T D^0$
- Refurbished PXL and SSD firmware: ~18% PXL dead in Run-14

SSD improves tracking 10% (20% for Ds)

Overall factor of 3.6 improvement for D^o

- Vertex Cut quality improvement (~15%)
- Pile-up protection study w/o 30% more data volume and 10% worse efficiency optimize protection (10%)
- Re-populate TPC ASIC and RDO, DAQ software optimization, online disk and network, +50% faster readout speed, reduced deadtime
- Bring up detector at RHIC Flattop and detector ramp down for beam dump Run 16: 7 (5) minutes vs Run 14: 9 (11) minutes

Quarkonia (MTD) related: (triggered/luminosity)

- High-Level Trigger dedicated to online dimuon selection
- Express stream of Upsilon candidates x10 reduction
- Reduce monitoring triggers to minimum required

Setbacks in First few weeks of run16

- Online Vertex Selection Efficiency (-5-10%)
 A version of hardware has been commissioned during the first 5 weeks, final implementation in March 16.
 improve efficiency 17%
- Uptime (-15%)
 Run 14 over 95 DAQ hours per week (110 CAD hours)
 Run 16 about 85 DAQ hours per week (94 CAD hours)
 April 6 after CAD Diode Fix, 95 DAQ hours (107 CAD hours)
- HFT readout Firmware (-20%)
 Low efficiency due to incorrect time latch
 Affects data from first 3 weeks
 A report assembled by a Detector Operation Task Force
- 4. RHIC Machine Diode Failure (March 18-April 6) Repaired and came back more efficient (Thanks!)

Typical Data-taking Mode



Designed and implemented by Jeff Landgraf. Linked from Trigger versioning page http://online.star.bnl.gov/RTS/plots/storedPlots.php Last 4 weeks' performance: 95 hours*8.3wk*3600seconds*700Hz=2B

Excellent Machine Performance after Diode Repair

Au+Au Dataset Goals



2000

Thu May 12 15:02:16 2016

07/Feb

06/Mar

(mostly) reach our goals.

29/May

day

01/May

03/Apr

HFT Preliminary Performance (Run 16)

With preliminary calibrations, Run16 HFT matching efficiency and DCA resolution show a comparable or slightly better performance w.r.t Run14

- ➡ different luminosity level
- different trigger setup for pileup protection



d+Au goals and projections



BUR Executive Summary Table

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- 2) 19 cryo-weeks in run 17, 13 weeks in run 18
- 3) If only 15 weeks in run 17, all for pp500

Run 17 Spin Program BUR and Projections

In this Beam Use Request the STAR Collaboration presents four compelling and prioritized scientific programs for the 2017 and 2018 RHIC runs.

STAR's **highest scientific priority** is the first significant measurement of the sign change of the Sivers function, as compared to the value measured in semi-inclusive deep inelastic scattering experiments, through measurements of single spin asymmetries in $W^{+/-}$, Z, direct photon and Drell-Yan production in transversely polarized $\sqrt{s} = 500 \text{ GeV}$ p+p collisions. This measurement will also shed light on the size and nature of the evolution of these transverse momentum dependent distributions. The sign change measurement is a fundamental test of QCD and is being pursued by other experiments, making a timely measurement imperative. We therefore request **13 weeks of 500 GeV** p+p running in Run17.

Year	#	Milestone
2015	HP13 (new)	Test unique QCD predictions for relations between single-transverse spin phenomena in p-p scattering and those observed in deep-inelastic lepton scattering.



NSAC Milestone (HP13)

Year	#	Milestone					
2013	HP8	Measure flavor-identified q and \overline{q} contributions to the spin of the proton via the longitudinal-spin asymmetry of W production.					
2013	HP12 (update of HP1, met in 2008)	Utilize polarized proton collisions at center of mass energies of 200 and 500 GeV, in combination with global QCD analyses, to determine if gluons have appreciable polarization over any range of momentum fraction between 1 and 30% of the momentum of a polarized proton.					
2015	HP13 (new) 	Test unique QCD predictions for relations between single-transverse spin phenomena in p-p scattering and those observed in deep-inelastic lepton scattering.					

 A_N of W^{\pm},γ , DY in pp 500 are all sensitive to the Sivers sign-change. STAR can access all three world-class measurements in the proposed 2017 Run – **Significant discovery potential before 2020!**

Measure QCD interaction sign and **TMD** evolution More in-depth discussions this afternoon: See C. Gagliardi's talk STAR p-p 500 GeV (L = 25 pb⁻¹) STAR p-p 500 GeV (L = 25 pb⁻¹) **0.8** $0.5 < P_{\tau}^{W} < 10 \ GeV/c$ **0.8** $0.5 < P_T^W < 10 \ GeV/c$ 0.6 0.6 0.4 0.4 0.2 0.2 -0.2 -0.2 $W^{\dagger} \rightarrow I^{\dagger} v$ $W \rightarrow I \gamma$ -0.4 unpolarized polarized beam beam KQ (assuming "sign change") KQ (no ``sign change'') proton proton Global χ^2 /d.o.f. = 7.4 /6 Global x²/d.o.f. = 19.6 /6 -0.8 3.4% beam pol. uncertainty not shown -0.8 3.4% beam pol. uncertainty not shown recoil jet -0.5 -0.5 0 0.5 a 0.5 vW vW

How to measure certain color interactions are repulsive and others attractive: A View of the Colorful Microcosm Within a Proton (foundation for run 2017) https://www.bnl.gov/rhic/news2/news.asp?a=1824&t=pr STAR paper Phys. Rev. Lett. **116**, 132301 (2016), Editors' Suggestion

Run 17 Projection $W^{\pm} A_N$





First measurement of sign change (a big deal)

Quantifying TMD evolution

Constraint sea quark Sivers

Another TMD Observable: Drell-Yan e⁺e⁻

The orange square: achievable statistical precision for A_N DY asymmetry with 400 pb⁻¹

the theoretical prediction for the Sivers asymmetry A_N as a function of *DY* leptonpair rapidity at \sqrt{s} =500 GeV, before any TMD evolution is applied.





Same evolution scheme as in W-Boson

Theoretical predictions for DY for $0 < p_T < 1$ GeV and 4 < Q < 9 GeV **after TMD evolution is applied**. The yellow bands represent the uncertainties for the asymmetry.

QCD background rejection: Drell-Yan e⁺e⁻

Add post-shower (run 17) to refurbished Pb-Glass FMS and scintillator pre-shower detector (run 15).

Simulate response of detector package to pions and electrons, collecting energy deposited in 3 pre-shower layers, FMS tower and post-shower layer.

Use multivariate analysis techniques to achieve hadron rejection powers of 800-14,000 for hadrons of 15 to 60 GeV while maintaining 90% electron detection efficiency.





Twist-3 Observable: Direct photon A_N

Measurements of twist-3 observables explore the consistency between TMD and Collinear Twist-3 formalism.

Statistical and systematic uncertainties for the direct photon $A_{\rm N}$ after background subtraction compared to theoretical predictions for $\sqrt{s} = 500$ GeV.



Summary of Sivers Function Tests

Planned for STAR

	$A_{N}(W^{+/-},Z^{0})$	A _N (DY)	$A_{N}(\gamma)$
Sensitive to Sivers fct.	Yes	Yes	No
sign change through			
TMDs			
Sensitive to Sivers fct.	No	No	Yes
sign change through			
Twist-3 $T_{q,F}(x,x)$			
Sensitive to TMD	Yes	Yes	No
evolution			
Sensitive to sea quark	Yes	Yes	No
Sivers function		for x ~ 10 ⁻⁴	
Detector upgrade needed	No	Yes	No
		FMS post-shower	
Biggest experimental	Integrated luminosity	Background suppression	
challenge	`	Integrated luminosity	

pp500 Operation mode for W-boson

STAR TPC event pile-up affects tracking efficiency

The W-boson reconstruction efficiency was obtained from the data measured in 2011 to 2013 and with improved Tracking Algorithm.

The highest FoM is reached at a ZDC rate of 330 kHz corresponding to a luminosity of $1.5 \times 10^{32} \text{ cm}^{-2}\text{s}^{-1}$.

Optimization of delivered luminosity with dynamic beta* squeeze.

Requires 13 weeks to reach 400pb⁻¹



Forward Detector preparation for run 17

Forward Meson Spectrometer (FMS) + FMS Pre-shower + FMS Poster-Shower

Installation of pre-shower for run 15



FMS radiation cure by UV lights



- 1. All three detector subsystems: existing FMS, Pre-shower Add new post-shower
- 2. Add UV lights to cure FMS radiation

https://drupal.star.bnl.gov/STAR/system/files/STAR.FM S_.Postshower.v2.pdf



Isobar

Run	Energy	Duration	System	Goals	priority	Sequence
	$\sqrt{s_{NN}}$ =500 GeV	13-wk	Transverse p+p	A_N of W [±] , γ , Drell-Yan, L=400 pb ⁻¹ , 55% pol	1	1
		1-wk	p+p	RHICf		2
17		2-wk	CeC			
	$\sqrt{s_{NN}}$ =62.4 GeV	4-wk	Au+Au	Jets, dileptons, NPE 1.5B MB	3	3
18	√s _{NN} =200 GeV	3.5-wk	Ru+Ru	1.2B MB	2	4
	√s _{NN} =200 GeV	3.5-wk	Zr+Zr	1.2B MB	2	5
	√s _{nn} =27 GeV	2-wk	Au+Au	>500M MB	3	6

STAR's **second scientific priority** is to clarify the interpretation of measurements related to the chiral magnetic effect, chiral magnetic wave and chiral vortical effect. We therefore request **two 3.5 week runs in Run 18** with collisions of isobaric nuclei, Ruthenium-96 (Ru+Ru) and Zirconium-96 (Zr+Zr). Ru nuclei have an atomic charge of 44 compared to 40 for Zr. Ru+Ru collisions will therefore generate a magnetic field approximately 10% larger than Zr+Zr collisions while all else remains virtually fixed. Comparisons of charge separation in Ru+Ru and Zr+Zr collisions will isolate the magnetic field dependence of the observed charge separation thereby determining what fraction of those measurements are related to the chiral magnetic effect. These results will greatly advance our understanding of the chiral magnetic effect and have fundamental impact beyond the field of high-temperature QCD.

Results from Chiral Effects



Editors' Suggestion

Observation of Charge Asymmetry Dependence of Pion Elliptic Flow and the Possible Chiral Magnetic Wave in Heavy-Ion Collisions

L. Adamczyk *et al.* (STAR Collaboration) Phys. Rev. Lett. **114**, 252302 (2015) – Published 26 June 2015



A possible signature of chiral symmetry restoration, in the form of a chiral magnetic wave in the quark-gluon plasma, has been observed in heavy-ion collisions at RHIC.

Show Abstract +

PRL on possible Chiral Magnetic Wave

More in-depth discussions this afternoon: Talk by Paul Sorensen from CME Task Force

- We have published a few papers on possible Chiral Magnetic Effect and potential background
- U+U collisions
 Better understanding
- BES-I results on CME
 14.5GeV
- BES II with more statistics
- Chiral Magnetic Wave
 14.5GeV
- Chiral Vortical Effect

A decisive test with Isobars



CAD was not able to locate enriched Ruthenium-96 source Possible with the refurbished Oakridge Isotope Facility, Run with natural abundance reduces luminosity by x5 Can reach 1.2 Billion events within 3.5 weeks of operation

Projections for Isobar



Au+Au at 62.4GeV

Run	Energy	Duration	System	Goals	priority	Sequence
	$\sqrt{s_{NN}}$ =500 GeV	13-wk	Transverse p+p	A_N of W [±] , γ , Drell-Yan, L=400 pb ⁻¹ , 55% pol	1	1
		1-wk	p+p	RHICf		2
17		2-wk	CeC			
	$\sqrt{s_{NN}}$ =62.4 GeV	4-wk	Au+Au	Semi-inclusive Jets, dileptons, NPE, 1.5B MB	3	3
10	√s _{NN} =200 GeV	3.5-wk	Ru+Ru	1.2B MB	2	4
	√s _{NN} =200 GeV	3.5-wk	Zr+Zr	1.2B MB	2	5
	√s _{NN} =27 GeV	2-wk	Au+Au	>500M MB	3	6

The first of these requests is driven by theoretical calculations suggesting that the partonic coupling to the medium is stronger when temperatures are close to the critical value. Results from Run 10 at 62.4 GeV provide tantalizing evidence that this is true, but the data are statistically limited. We therefore request a **4 weeks 62.4 GeV Au+Au run** in Run 17 to significantly enhance the statistical precision of these results. Precise measurements of non-photonic electron v_2 will confirm that this enhanced coupling results in the partial thermalization of charmed quarks even at this low energy, while our semi-inclusive recoil jet studies will probe possible differences in the jet energy redistribution. Finally, measurements of direct (virtual) photons will improve our understanding of thermal photon production and their medium coupling.

Au+Au at 62GeV



Resolving possible charm puzzle

Large charm (NPE) suppression in Au+Au at 200GeV with large elliptic flow Indication of enhancement in Au+Au at 62.4GeV with (~)zero elliptic flow



Au+Au at 27GeV

Run	Energy	Duration	System	Goals	priority	Sequence
	$\sqrt{s_{NN}}$ =500 GeV	13-wk	Transverse p+p	A_N of W [±] , γ , Drell-Yan, L=400 pb ⁻¹ , 55% pol	1	1
		1-wk	p+p	RHICf		2
17		2-wk	CeC			
	$\sqrt{s_{NN}}$ =62.4 GeV	4-wk	Au+Au	Jets, dileptons, NPE 1.5B MB	3	3
18	√s _{NN} =200 GeV	3.5-wk	Ru+Ru	1.2B MB	2	4
	√s _{NN} =200 GeV	3.5-wk	Zr+Zr	1.2B MB	2	5
	√s _{nn} =27 GeV	2-wk	Au+Au	>500M MB	3	6

The second request is derived from STAR's recent report of the observation of the *global* polarization of hyperons (GPH) in non-central Au+Au collisions at energies below 40GeV. This is the first direct probe of the plasma vorticity, a fundamental characteristic of any fluid that must be quantified in ordrer to understand the physics in detail. It also provides a measure of the magnetic field present over the evolution of the QGP. In addition to being of fundamental interest on their own accounts, these measurements provide critical context for recent high-profile studies of exotic phenomenon. In

QCD phase transition is a chiral

phase transition

- 1. Charge separation (14.5GeV)
- 2. Bulk charge dependence of $\pi^{\pm} v_2$
- 3. Low-mass dilepton excess
- 4. Global polarization of hyperons







QCD fluid responds to external field



June 7 Xin-Nian Wang Polarization of Fermions in Vorticular Field June 8 Issal Upsal Lambda Global Polarization

Request Au+Au 27GeV in run18 with EPD

https://drupal.star.bnl.gov/STAR/system/files/EPD_Construction_Proposal.pdf

- Clearly, very exciting development
- Signal and BES dependence need more data
- Request Au+Au 27GeV in run 18 with EPD or run17 without EPD, but earlier results
- To establish whether there is a difference
- Result will guide further studies in BES-II





Plan presented in 2011

Summary of the plan

	Near term (Runs 11-13)	Mid-decade (Runs 14-16)	Long term (Runs 17–)	
Colliding systems	<i>p+p</i> , A+A	<i>p</i> + <i>p</i> , A+A	<i>p</i> + <i>p</i> , <i>p</i> +A, A+A, e+ <i>p</i> , e+A	
Upgrades	FGT, FHC, RP, DAQ10K, Trigger	HFT, MTD, Trigger	Forward Instrum, eSTAR, Trigger	
(1) Properties of sQGP	Y, J/ ψ $ ightarrow$ ee, $m_{ m ee}$, v_2	Y, $J/\psi \rightarrow \mu\mu$, Charm v_2 , R_{CP} , corr, Λ_c/D ratio, μ -atoms	<i>p</i> +A comparison	
(2) Mechanism of energy loss	Jets, γ-jet, NPE	Charm, Bottom	Jets in CNM, SIDIS, <i>c/b</i> in CNM	
(3) QCD critical point	Fluctuations, correlations, particle ratios	Focused study of critical point region		
(4) Novel symmetries	Azimuthal corr, spectral function	e-μ corr, μ-μ corr		
(5) Exotic particles	Heavy anti-matter, glueballs			
(6) Proton spin structure	$W A_L$, jet and di-jet A_{LL} , intra-jet corr, (Λ + $\overline{\Lambda}$) D_{LL}/D_{TT}		$\overline{\Lambda} D_{LL}/D_{TT},$ polarized DIS & SIDIS	
(7) QCD beyond collinear fact	Forward A _N		Drell-Yan, F-F corr, polarized SIDIS	
(8) Properties of			Charm corr,	
initial state	Measurements lis	ted when they first	Drell-Yan, <i>Jlψ</i> , F-F corr, Λ, DIS, SIDIS	
	Many will continu	Many will continue in future periods		

STAR Decadal Plan – June, 2011 PAC Meeting C. Gagliardi for the STAR Collaboration

Plan presented in 2011

Summary of the plan

	Near term (Runs 11-13)	Mid-decade (Runs 14-16)	Long term (Runs 17–)	
Colliding systems	<i>p+p</i> , A+A	<i>p+p</i> , A+A	<i>p+p</i> , <i>p+</i> A, A+A, e+ <i>p</i> , e+A	
Upgrades	FGT, FHC, RP, DAQ10K, Trigger	HFT, MTD, Trigger	Forward Instrum, eSTAR, Trigger	
(1) Properties of sQGP	Y, J/ ψ $ ightarrow$ ee, $m_{ m ee}$, v_2	Y, $J/\psi \rightarrow \mu\mu$, Charm v_2 , R_{CP} , corr, Λ_c/D ratio, μ -atoms	<i>p</i> +A comparison	
(2) Mechanism of energy loss	Jets, γ-jet, NPE	Charm, Bottom	Jets in CNM, SIDIS, c/b in CNM	
(3) QCD critical point	Fluctuo ti ans correlations, Da fici Dinos	Formedistativ of crit 🗢 point gran		
(4) Novel symmetries	Azimuthal corr, spectral function	e-μ corr, μ-μ corr		
(5) Exotic particles	Heavy anti-matter, glueballs			
(6) Proton spin structure	WA_L , jet and di-jet A_{LL} , intra-jet corr, (Λ + $\overline{\Lambda}$) D_{LL}/D_{TT}		$\overline{\Lambda} D_{LL}/D_{TT},$ polarized DIS & SIDIS	
(7) QCD beyond collinear fact	Forward A _N		Drell-Yan, F-F corr, polarized SIDIS	
(8) Properties of			Charm corr,	
initial state	Measurements lis	ted when they first	Drell-Yan, <i>Jlψ</i> , F-F corr, Λ, DIS, SIDIS	
	Many will continu	Many will continue in future periods		

STAR Decadal Plan – June, 2011 PAC Meeting C. Gagliardi for the STAR Collaboration

RHIC has been adaptable to science needs

2010-13 2014-16 BES-I BES-II

RHIC has been adaptable to science needs

2010- 2013	2014	2015	2016	2017	2018	2019	2020	2022+
Au+Au p+p	Au+Au	p+p p+A	Au+Au d+Au	p+p Au+Au	Isobar Au+Au	Au+Au	Au+Au	Au+Au pp,pA
∆G, QGP property	Charm flow	Ref. A _N	D _c , Λ _c Υ, Jets	Fc sign	CME, Λ↑	Critical Point, Phase Transition		Jets, Υ forward A_N
BES-I	200, 14.5	200	200- 19.6	500, 62.4	200, 27	BES-II 11-20	BES-II 7-11	200

BES-I

BES-II

Expand to include several programs: p+A in run 15, pp500 in run17, Isobar (Zr, Ru-96) in run 18 BES-II more compelling, detector and machine upgrades in 2018 Future high-luminosity jets and Upsilon in 2020+ 3+1D hydrodynamics and Unique Cold QCD (DY) portal to EIC

Summary

- Successful run16 (thanks CAD)
- Compelling Heavy-Ion Programs for run 16 and run 18
 - Completion of Heavy-Flavor program for HFT+MTD
 - Decisive test of Chiral Magnetic Effect
 - Partonic coupling to QGP close to T_c
 - Quantifying the role of external field in Global Hyperon Polarization
- Compelling Spin Program in run17 Three measurements related to TMD evolution and sign change (A_N W^{+/-}, γ, DY)
- Maintain track record in Results and Publications
- Preparation for BES II and beyond (3+1D hydro and Cold QCD)

Run	Energy	Duration	System	Goals	priority
17	√s _{NN} =500 GeV	13-wk	Transverse p+p	A_N of W [±] , γ , Drell-Yan, L=400 pb ⁻¹ , 55% pol	1
17	$\sqrt{s_{NN}}$ =62.4 GeV	4-wk	Au+Au	Jets, dileptons, NPE 1.5B MB	3
	√s _{NN} =200 GeV	3.5-wk	Ru+Ru	1.2B MB	2
18	√s _{NN} =200 GeV	3.5-wk	Zr+Zr	1.2B MB	2
	√s _{NN} =27 GeV	2-wk	Au+Au	>500M MB	3

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1 Executive Summary

In this Beam Use Request the STAR Collaboration presents four compelling and prioritized scientific programs for the 2017 and 2018 RHIC runs.

STAR's highest scientific priority is the first significant measurement of the sign change of the Sivers function, as compared to the value measured in semi-inclusive deep inelastic scattering experiments, through measurements of single spin asymmetries in $W^{+/-}$, Z, direct photon and Drell-Yan production in transversely polarized $\sqrt{s} = 500 \text{ GeV}$ p+p collisions. This measurement will also shed light on the size and nature of the evolution of these transverse momentum dependent distributions. The sign change measurement is a fundamental test of QCD and is being pursued by other experiments, making a timely measurement imperative. We therefore request 13 weeks of 500 GeV p+p running in Run17.

STAR's second scientific priority is to clarify the interpretation of measurements related to the chiral magnetic effect, chiral magnetic wave and chiral vortical effect. We therefore request two 3.5 week runs in Run 18 with collisions of isobaric nuclei, Ruthenium-96 (Ru+Ru) and Zirconium-96 (Zr+Zr). Ru nuclei have an atomic charge of 44 compared to 40 for Zr. Ru+Ru collisions will therefore generate a magnetic field approximately 10% larger than Zr+Zr collisions while all else remains virtually fixed. Comparisons of charge separation in Ru+Ru and Zr+Zr collisions will isolate the magnetic field dependence of the observed charge separation thereby determining what fraction of those measurements are related to the chiral magnetic effect. These results will greatly advance our understanding of the chiral magnetic effect and have fundamental impact beyond the field of high-temperature QCD.

Our next scientific priorities involve taking data at two beam energies that are lower than the nominal energies, but are not part of STAR's proposed Beam-Energy-Scan Phase-II (BES-II) program in 2019-2020 [1].

The first of these requests is driven by theoretical calculations suggesting that the partonic coupling to the medium is stronger when temperatures are close to the critical value. Results from Run 10 at 62.4 GeV provide tantalizing evidence that this is true, but the data are statistically limited. We therefore request a 4 weeks 62.4 GeV Au+Au run in Run 17 to significantly enhance the statistical precision of these results. Precise measurements of non-photonic electron v_2 will confirm that this enhanced coupling results in the partial thermalization of charmed quarks even at this low energy, while our semi-inclusive recoil jet studies will probe possible differences in the jet energy redistribution. Finally, measurements of direct (virtual) photons will improve our understanding of thermal photon production and their medium coupling.

The second request is derived from STAR's recent report of the observation of the global polarization of hyperons (GPH) in non-central Au+Au collisions at energies below 40GeV. This is the first direct probe of the plasma vorticity, a fundamental characteristic of any fluid that must be quantified in ordrer to understand the physics in detail. It also provides a measure of the magnetic field present over the evolution of the QGP. In addition to being of fundamental interest on their own accounts, these measurements provide critical context for recent high-profile studies of exotic phenomenon. In particular, the Chiral Vortical Effect (CVE) and Chiral Magnetic Effect (CME) are among the RHIC program's most exciting and visible topics today. However, they are nontrivial and require several conditions to be simultaneously met; "extraordinary claims require extraordinary evidence." GPH has nothing "chiral" or exotic about it. In a theoretically well-grounded way, it measures the average vorticity (ω) and field (B). Hence, GPH must be understood first, to put more "extraordinary" claims on firmer ground. A solid understanding of GPH and the magnetic field, however, requires the increased statistics available with a 2-week run at 27 GeV. While similar high-statistics measurements of GHP will eventually be performed in BES-II, the scientific impact of getting firm measures of vorticity and B much sooner is high, given the current intensity of research and attention on CME and CVE.

Table 1.1 and Table 1.2 summarize the above requests, along with our scientific priorities and proposed running sequence. We have considered two scenarios:

Scenario 1: 19 cryo-weeks for Run 17 and 13 cryo-weeks for Run 18 (Table 1-1) Scenario 2: 24 cryo-weeks for Run 17 and 13 cryo-weeks for Run 18 (Table 1-2)

In this document we first present highlights from papers published since the last PAC meeting and on-going analyses. Next we summarize the analysis status of data from STAR's two recent upgrades, the Heavy Flavor Tracker (HFT) and the Muon Telescope Detector (MTD), and discuss the detector's performance in Run 16. In Sections 4 and 5, we detail the STAR Collaboration's compelling physics program driving our Run 17 and Run 18 beam-use request. Collection of these data will allow us to achieve our spin and relativistic heavy ion physics goals on a timescale consistent with the current intense international interest while utilizing RHIC beams effectively and taking full advantage of recent improvements in machine and detector capability. Finally, we outline the planned upgrades for Run 17 and 18 and the BES-II.

Run	Energy	Duration	System	Goals	priority	Seq
17	√s _{NN} =500 GeV	13-wk	Transverse p+p	A_N of W^{\pm} , γ , Drell- Yan, \mathcal{L} =360 pb ⁻¹ ,	1	1
17		1-wk	p+p	55% pol RHICf		2
		2-wk	CeC			
	√s _{NN} =200 GeV	3.5-wk	Ru+Ru	1.2B MB	2	3
18	√s _{NN} =200 GeV	3.5-wk	Zr+Zr	1.2B MB	2	4
	√s _{NN} =27 GeV	2-wk	Au+Au	500M MB	3	5

Table 1.1: Scenario I: 19 cryo-weeks for Run 17 and 13 cryo-weeks for Run 18. All luminosities requested are actual sampled luminosity. See text for details about pp500 run conditions and optimizations.



