

High-energy dipole scattering amplitude off a proton

from low-energy light-cone wave function & BK evolution

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for the SURGE collaboration

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BFKL / BK/ JIMWLK evolution eqs are differential eqs, require initial condition:
dipole in Large- N_c approximation

$$\partial_Y S(\vec{x}, \vec{y}) = \frac{\alpha_s N_c}{2\pi^2} \int d^2 z \frac{(\vec{x} - \vec{y})^2}{(\vec{x} - \vec{z})^2 (\vec{z} - \vec{y})^2} [S(\vec{x}, \vec{z}) S(\vec{z}, \vec{y}) - S(\vec{x}, \vec{y})]$$

- issues with current implementations of I.C. (next slide)
- body of empirical knowledge of “large-x” structure of proton, can it be incorporated ?
- in the future: proton LCwf from the lattice? (though only $x \gtrsim 0.1$)
If available, how can we use it to construct I.C. for small-x ?

Initial conditions for (rc)BK fits to small-x inclusive HERA data :

$$N(r; x_0) = 1 - \exp \left[-\frac{1}{4} (r^2 Q_{s0}^2)^\gamma \log \left(\frac{1}{r\Lambda} + ee_c \right) \right]$$

AAMQS: PRD 80 (2009)
EPJ-C 71 (2011)
Lappi, Mäntysaari: PRD 88 (2013)

- fits prefer $\gamma \sim 1.1 - 1.2$; violates pQCD prediction $N(r) \sim r^2$ for small r^2 ,
positivity of Fourier transform (UGD)
issues with generalization to nuclei (Albacete et al, 1209.2001)
- imposed at $x_0=0.01$, requires re-fit for other x_0
no prediction for dependence on x_0 as required for NLO (e.g. Ducloue et al, 1902.06637)
- no connection to “large”-x proton structure; parameters fitted to small-x data itself,
rcBK + LO photon w.f. $\rightarrow \chi^2/N_{\text{dof}} \sim 1$, \rightarrow “perfect fit” !?
no statistical evidence for NLO corrections!
(they’d either make matters worse or they could be absorbed into a retune
of the initial condition)

The proton on the light front

The proton on the light front (three quark Fock state; L.C. time $x^+ = 0$)

$$|P\rangle = \frac{1}{\sqrt{6}} \int \frac{dx_1 dx_2 dx_3}{\sqrt{x_1 x_2 x_3}} \delta(1 - x_1 - x_2 - x_3) \int \frac{d^2 k_1 d^2 k_2 d^2 k_3}{(16\pi^3)^3} 16\pi^3 \delta(\vec{k}_1 + \vec{k}_2 + \vec{k}_3) \\ \times \psi(k_1, k_2, k_3) \sum_{i_1, i_2, i_3} \epsilon_{i_1 i_2 i_3} |p_1, i_1; p_2, i_2; p_3, i_3\rangle$$

+ higher Fock states

P. Lepage & Brodsky, 1979 -
Brodsky, Pauli, Pinsky, PR (1998)

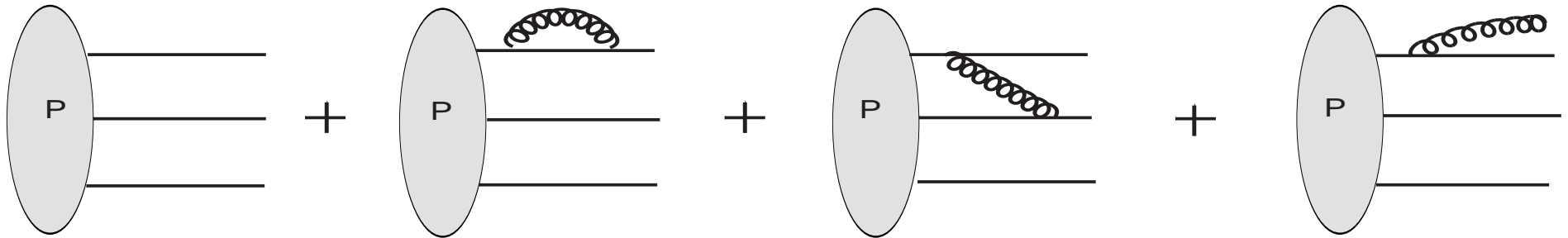
* Fock space amplitude ψ is gauge invariant, universal, and process independent, encodes the non-perturbative structure of hadrons (QCD eigenstates)

* **fix 3q wave function Ψ_{qqq} from low-energy / “large”-x proton structure, augment by Ψ_{qqqg} via LCpt**

* evolution adds additional soft gluons into wave function

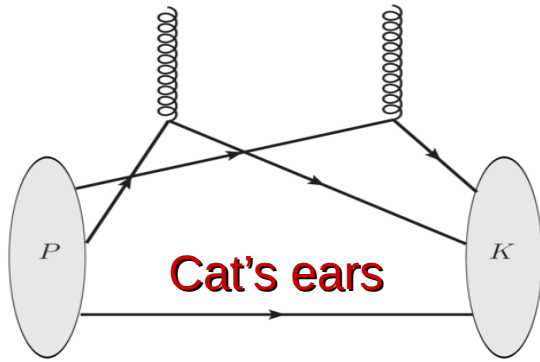
Now to $|P\rangle \sim \psi_{qqq}|qqq\rangle + \psi_{qqqg}|qqqg\rangle$

computed in perturbation theory, 1-gluon emission / exchange,
w/o employing $x_g \rightarrow 0$ approximation;
built on top of qqq “low-energy structure” of the proton



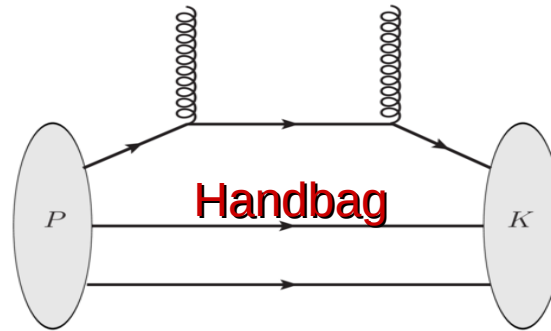
Color charge correlators:

$$\langle \rho^a(\vec{q}_1) \rho^b(\vec{q}_2) \rangle \sim g^2 \delta^{ab} G_2(\vec{q}_1, \vec{q}_2)$$



2-body GTMD / Wigner

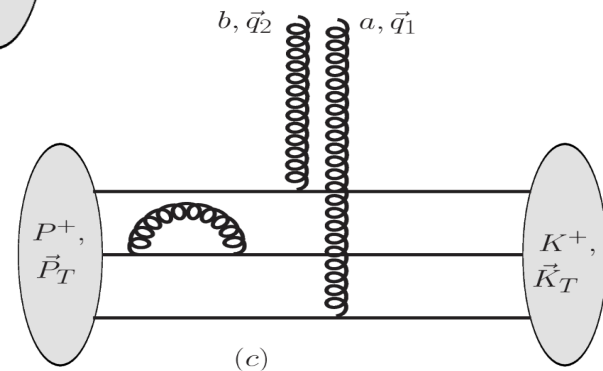
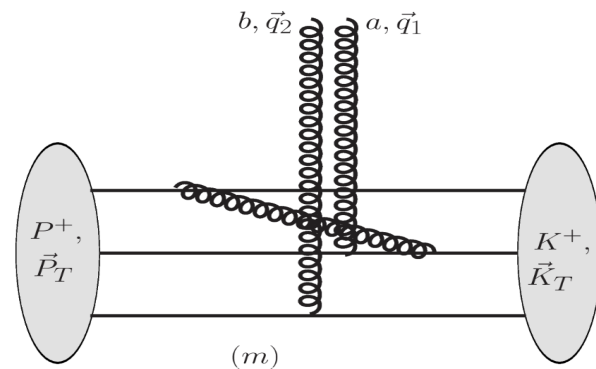
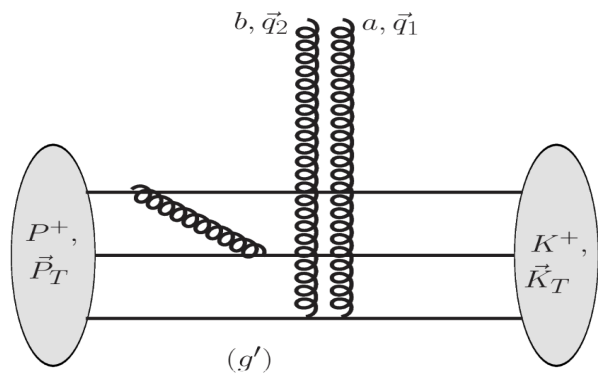
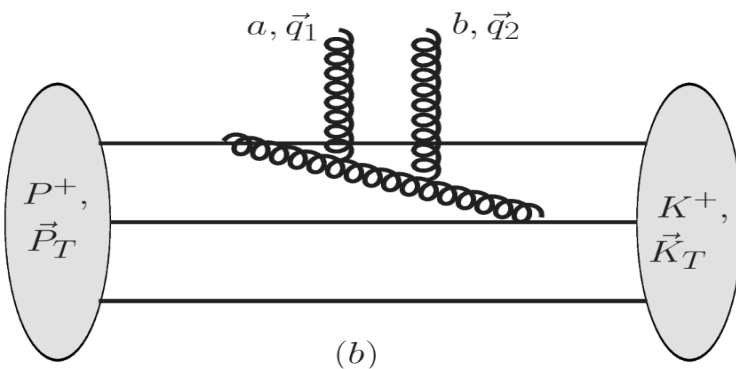
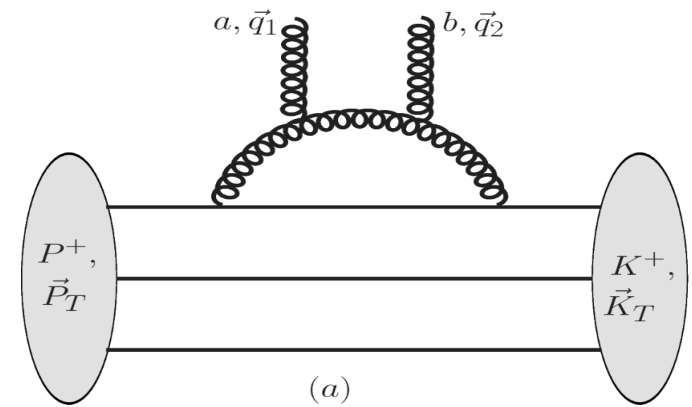
$$\rho^a(\vec{x}) = \int dx^- J^{+a}(x^+ = 0, x^-, \vec{x})$$



1-body GTMD / Wigner

LO diagrams

NLO color charge correlators



notes:

- * soft & collinear div.
- * UV divergences cancel
- * Ward satisfied

+ many more

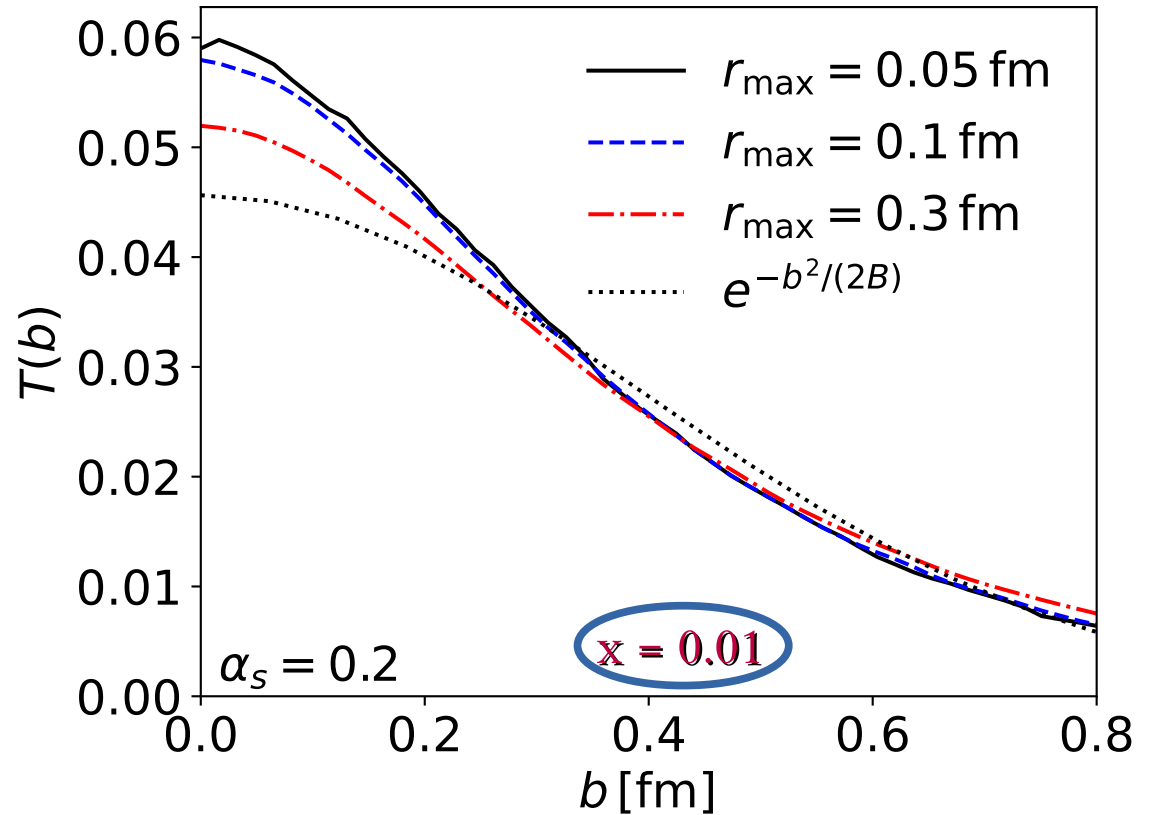
Twenty thousand leagues under the seas

from $x = 0.1 \rightarrow x = 10^{-4}$ (and below) ?

selected results from arXiv:2303.16339
(w/ Heikki Mäntysaari & Risto Paatelainen)

Effective transverse “profile”

- $T(b) \sim \int d^2r N(\vec{r}, \vec{b}; x)$
- r and x -dependent
- close to the Gaussian profile that was fitted to data



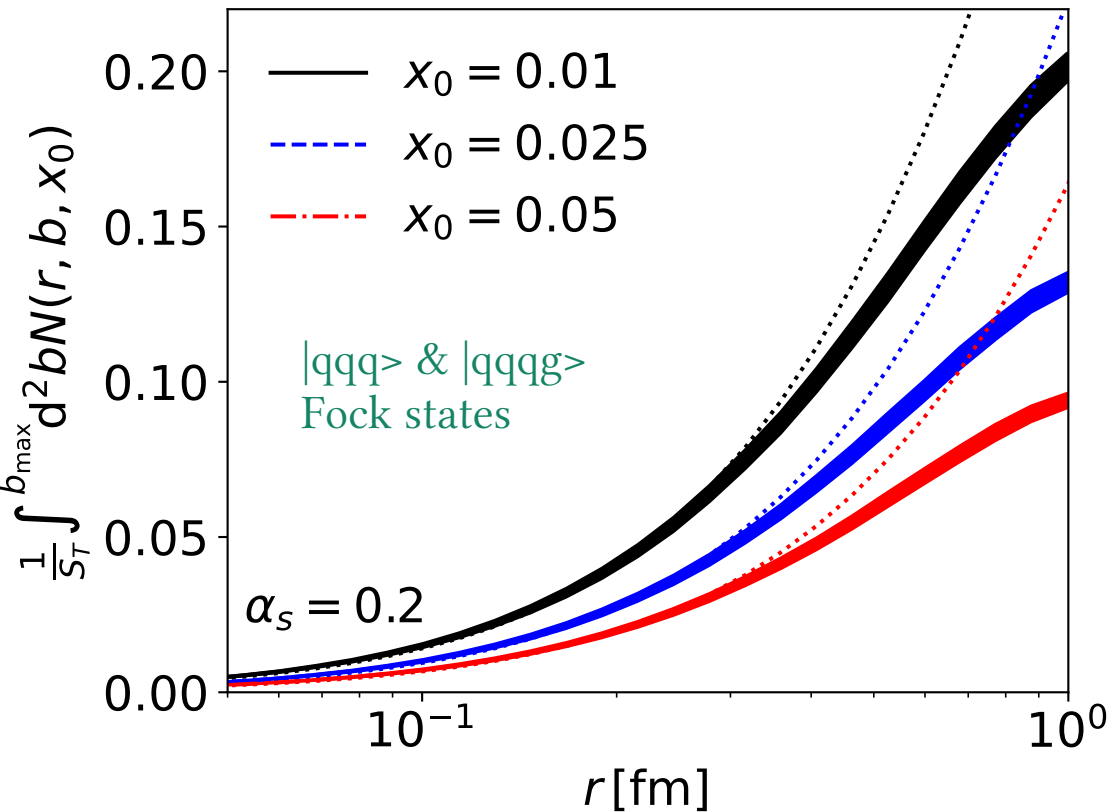
r-dependence of $N(r;x)$

- $\sim r^2$ in perturbative regime, as expected
- analytic parameterization for $r < \sim 0.25$ fm:

$$N(r) = 1 - \exp \left[-\frac{1}{4} (r^2 Q_{s0}^2)^\gamma \log \left(\frac{1}{r\Lambda} + 1 \right) \right] \frac{1}{S_T} \int_{b_{\min}}^{b_{\max}} d^2 b N(r, b, x_0)$$

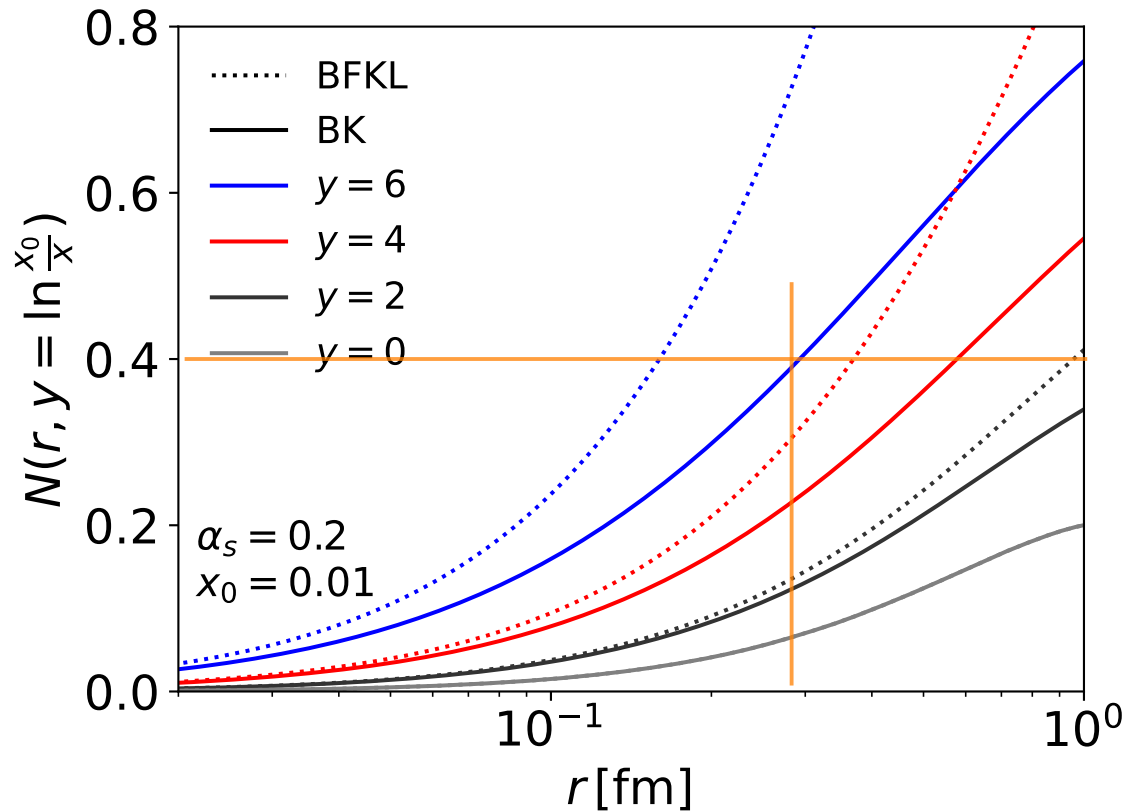
Q_{s0} , γ fitted to numerical data,
 $\Lambda = 0.24$ GeV was fixed

x	$Q_{s,0}^2$ [GeV ²]	γ
0.01	$0.100^{+0.004}_{-0.004}$	$1.001^{+0.001}_{-0.001}$
0.025	$0.066^{+0.003}_{-0.003}$	$0.998^{+0.001}_{-0.001}$
0.05	$0.047^{+0.002}_{-0.002}$	$0.997^{+0.001}_{-0.001}$



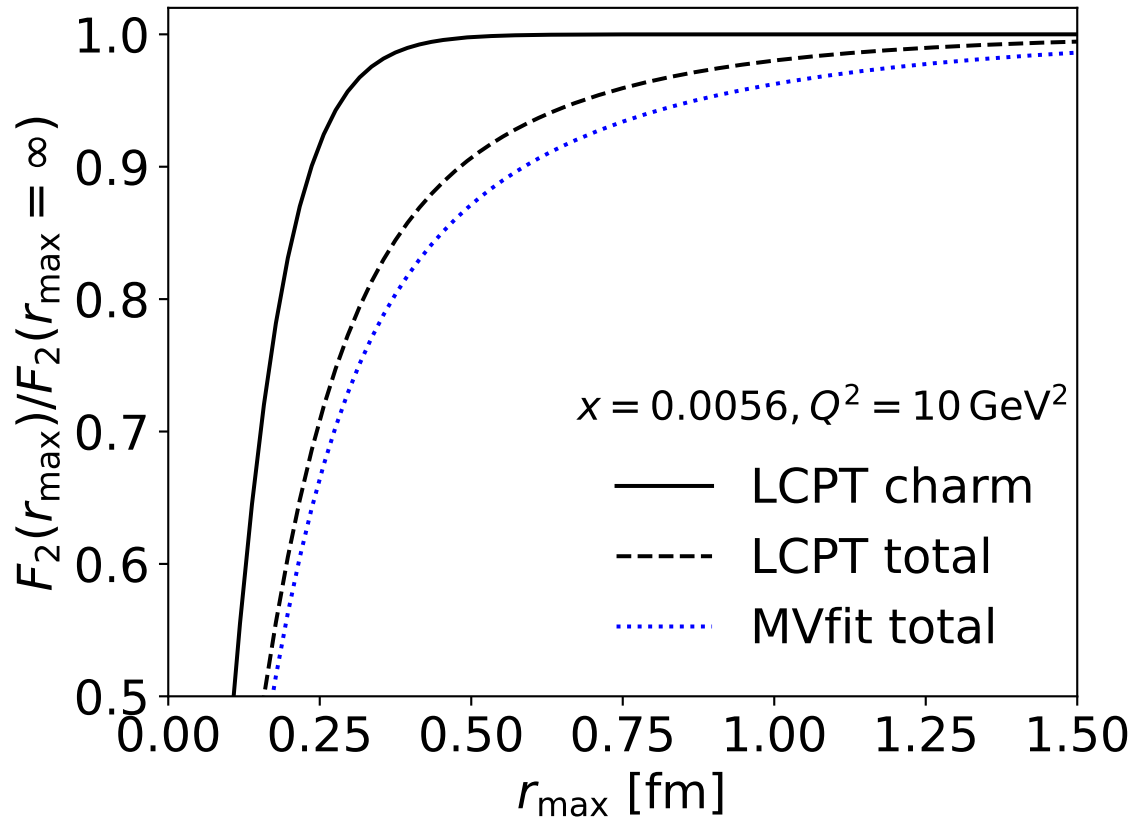
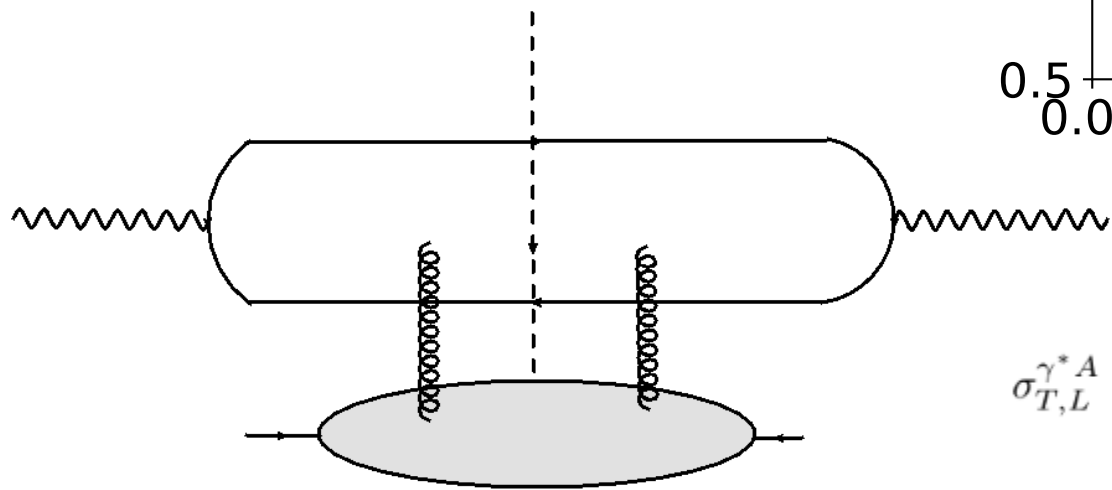
LL evolution of $N(r;x)$ from $x_0 = 0.01$

- adds Fock states with additional soft gluons
- BFKL only for $y < \sim 2$
- saturation requires $y \sim 6 \leftrightarrow x \sim 2E-5: r_s = 0.3\text{fm}, Q_s = 1\text{GeV}$



DIS cross section

- total (light q) X-sec sensitive to large dipoles
- charm X-sec saturated by perturbative dipoles \rightarrow focus on charm



$$\sigma_{T,L}^{\gamma^* A} = 2 \sum_f \int d^2\mathbf{b} d^2\mathbf{r} dz \left| \Psi^{\gamma^* \rightarrow q\bar{q}}(\mathbf{r}, z, Q^2) \right|^2 N(\mathbf{r}, \mathbf{b}, \bar{x})$$

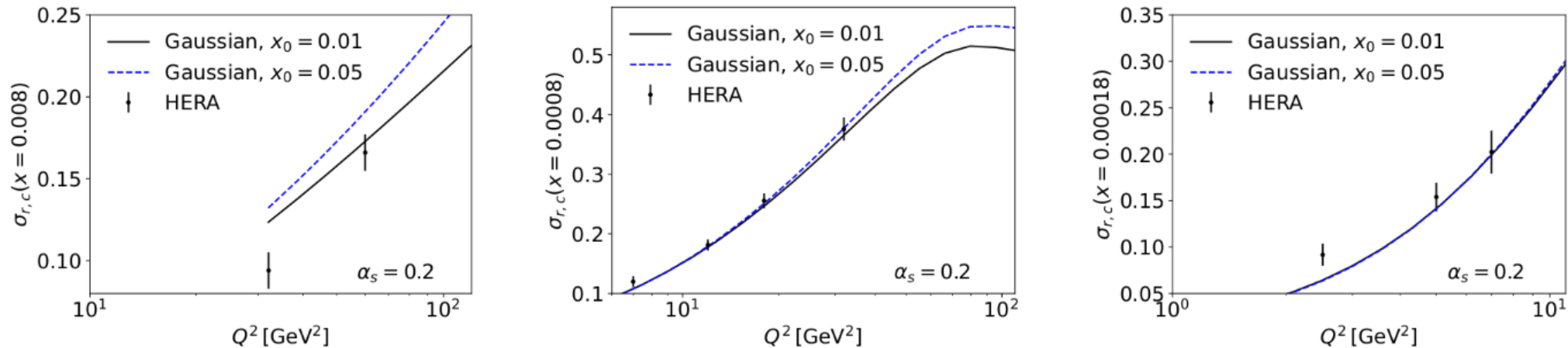
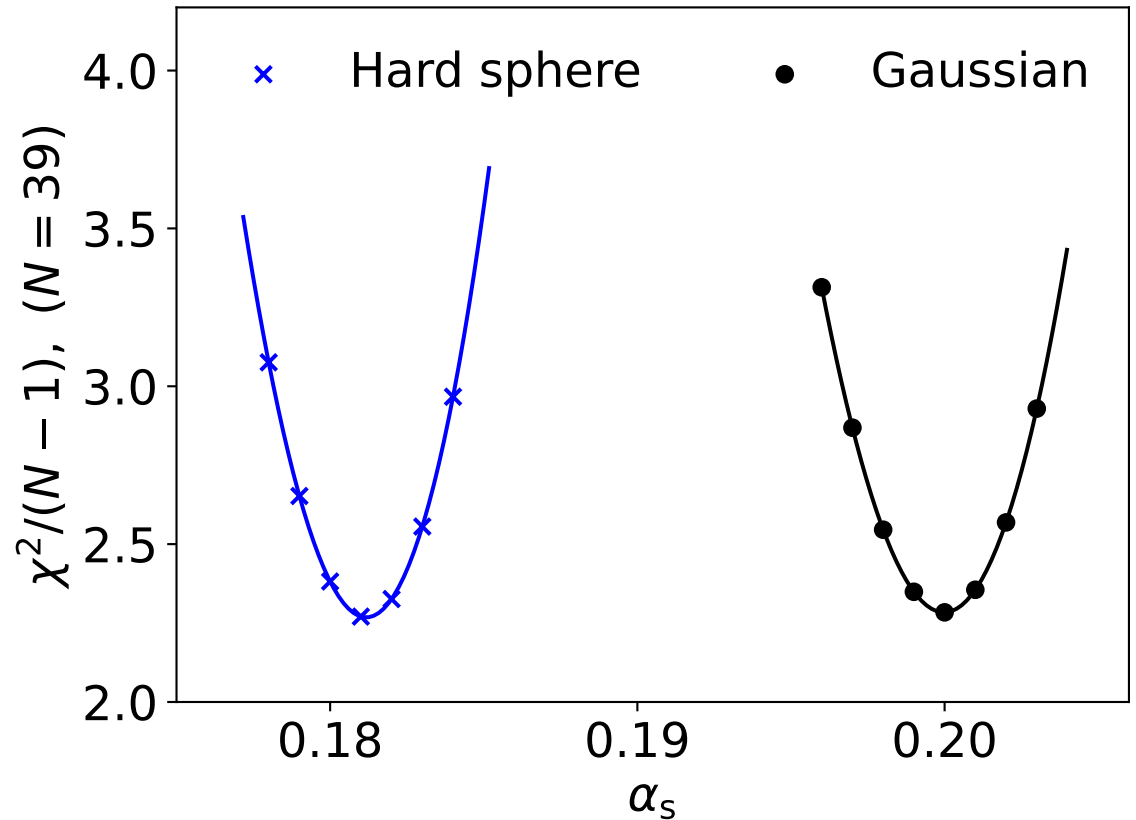


FIG. 9: Charm production reduced cross section at $\sqrt{s} = 318$ GeV compared to the HERA data. Results are shown in the region where $\bar{x} \leq 0.01$.

- decent agreement, only fit parameter is (fixed) α_s ,
no re-tune of Brodsky & Schlumpf non-perturbative qqq wave function was performed !
- we covered 3 decades in energy from $x \sim 0.1 \rightarrow x \sim 10^{-4}$

Statistical evidence for systematic corrections :

- HERA data is very accurate!
- $\chi^2/N_{\text{dof}} = 2.27$ for $N_{\text{dof}}=38$:
p-value (integral over χ^2 distribution) $\sim 10^{-5}$!
- evidence for moderate but *systematic* corrections:
photon w.f. (here: LO)
evolution eqn (here: LL-BK)
initial condition



Done:

- tables & analytical parameterizations for b-averaged $N_0(r; x)$
using Brodsky & Schlumpf non-pert. qqq LCwf + LCpt one gluon correction
- tables & interpolation routines for $N_0(\vec{r}, \vec{b}; x)$, incl. angular dependence !

Later on (if requested):

- option to try other non-pert. qqq wave functions
- imaginary part due to perturbative C-odd ggg exchange
- etc...

Backup

Model LFwf for the proton (Brodsky & Schlumpf, PLB 329, 1994)

$$\psi_{\text{H.O.}}(x_1, \vec{k}_1; x_2, \vec{k}_2; x_3, \vec{k}_3) = N_{\text{H.O.}} \exp(-\mathcal{M}^2 / 2\beta^2) ,$$

$$\psi_{\text{Power}}(x_1, \vec{k}_1; x_2, \vec{k}_2; x_3, \vec{k}_3) = N_{\text{Power}} (1 + \mathcal{M}^2 / \beta^2)^{-p} .$$

$$\mathcal{M}^2 = \sum_{i=1}^3 \frac{\vec{k}_{\perp i}^2 + m^2}{x_i}$$

$$m = 0.26 \text{ GeV}, \quad \beta = 0.55 \quad \text{for H.O. wf}$$

$$m = 0.263, \quad \beta = 0.607, \quad p = 3.5 \quad \text{for PWR wf}$$

With these parameters they fit:

- proton radius $R^2 = -6 \left. \frac{dF_1(Q^2)}{dQ^2} \right|_{Q^2=0} = (0.76 \text{ fm})^2$

- proton / neutron magnetic moments $1 + F_2(Q^2 \rightarrow 0) = 2.81 / -1.66$

- axial vector coupling $g_A = 1.25$

The definition of χ^2

If ν independent variables x_i are each normally distributed with mean μ_i and variance σ_i^2 , then the quantity known as *chi-square*² is defined by

$$\chi^2 \equiv \frac{(x_1 - \mu_1)^2}{\sigma_1^2} + \frac{(x_2 - \mu_2)^2}{\sigma_2^2} + \dots + \frac{(x_\nu - \mu_\nu)^2}{\sigma_\nu^2} = \sum_{i=1}^{\nu} \frac{(x_i - \mu_i)^2}{\sigma_i^2} \quad (1)$$

Note that ideally, given the random fluctuations of the values of x_i about their mean values μ_i , each term in the sum will be of order unity. Hence, if we have chosen the μ_i and the σ_i correctly, we may expect that a calculated value of χ^2 will be approximately equal to ν . If it is, then we may conclude that the data are well described by the values we have chosen for the μ_i , that is, by the hypothesized function.

The χ^2 distribution

The quantity χ^2 defined in Eq. 1 has the probability distribution given by

$$f(\chi^2) = \frac{1}{2^{\nu/2} \Gamma(\nu/2)} e^{-\chi^2/2} (\chi^2)^{(\nu/2)-1}$$

This is known as the χ^2 -*distribution with ν degrees of freedom*. ν is a positive integer.