INTT Commissioning Plan

RIKEN/RBRC

Itaru Nakagawa

Short Term Plan

- Establish pre-measurement routine for the day.
 - Check RACK power status
 - Check cooling
- Establish summary report of the day format. (Cheng-Wei & Genki)
- Complete all S&N ROC calibration and address in newly discovered problem in the latest data. (Cheng-Wei)
- LV/Bias monitor (plot voltage/current as a function of time) to be developed within ignition scheme for LV.

RCDAQ on Felix developments ···

- Raul's test is the highest priority. Raul can test whenever he needs. Other data taking will be carried out while he is not doing test in IR.
- Raul works on the GTM integration and test it on Monday afternoon.
- Long time (~10 hours) high threshold noise data taking (optimize threshold not to get the file too big) for the stability test of the optical link. Genki will execute while his shift on Monday.

Data Taking Plan (Week of April 24)

- Noise data for all S+N ROCs by self-trigger for environment dependence study
 - Noise study with light off (3+3 hours) -> Rates analysis Joseph
 - Noise study with light on/off for a few ROCs. DAC0=15. (1hour)
 - Noise study with magnet door closed (3+3 hours)
- DAC0 scan for noise rate study <u>for threshold optimization practice</u> (3hours)
- DAC scan (First attempt to observe MIP in IR) with self-trigger mode. (3hours)
- Test 2 vs 1 ROCs modes calibration data taking with the latest DAQ version. (1h)
 - Make sure no difference in results between two modes. ✓
- Take calib data North and South fibers together in 2 ROCs mode.
 (1h)
 - Make sure nothing odd happen.

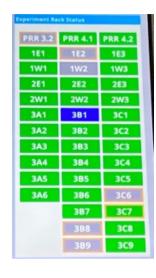
Check List for the daily measurements

- Establish pre-measurement routine for the day.
 - Check RACK power status
 - Check cooling (30GPH/ROC, 10GPH/ladder), GUI is under bug fix by Steve and Rob. (Rachid/Itaru needs to check with them)
 - Let shift leader know the activities. (make e-log entry at the end)
 - Make sure turn off the ROC/FPHX after the measurement

A 18 15 A 8 6 6 6	
TA TO TO A TO TO THE TAIL THE	
SUPPLY INTT DETECTOR COOLING	





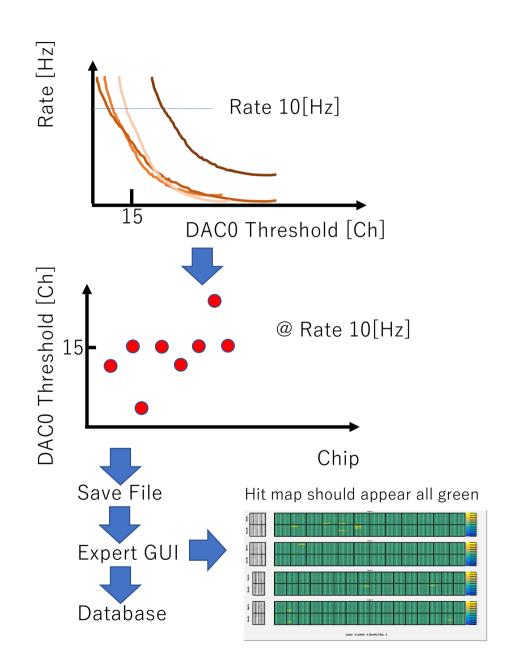




DACO Scan

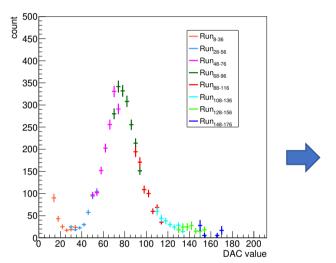
- Scan DAC0 value one-by-one [20, 19, 18, 17, ···] w/ self-trigger.
- Need develop analysis software to evaluate rates/chip.
- Optimize the DAC0 threshold for each Chip to give the same given rate. Save threshold values (52 x 56) in file loadable to Expert GUI.
- Customize DAC0 threshold setting chip-by-chip (Expert-GUI)
- This is the practice for the DAC0 threshold scan with the beam during commissioning. (S/N is also considered w/ beam)

Joseph develops data converter to feed non-prdf data to the online monitor

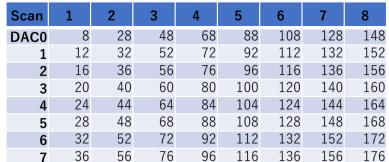


DAC Scan

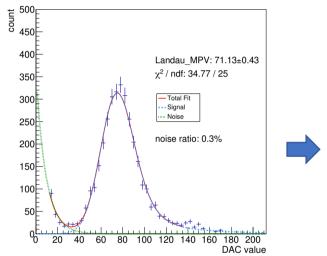
w/ Cosmic Ray
w/ self-trigger



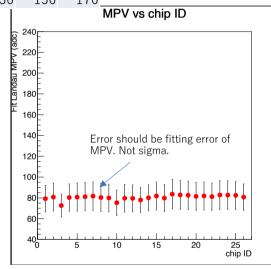
- Chip-by-Chip Base
- Clustering (Optimize offset value)
- Normalization btwn adjacent runs
- Concatenate all runs



Analyzer:
onsite quick analysis
+ Yuka for detail?



 Fitting with Landau+Gaussian convolution function.



Half ladder by half ladder



All ladders

Save all fitting parameters: MPV, Width, ...

Commissioning without beam Status

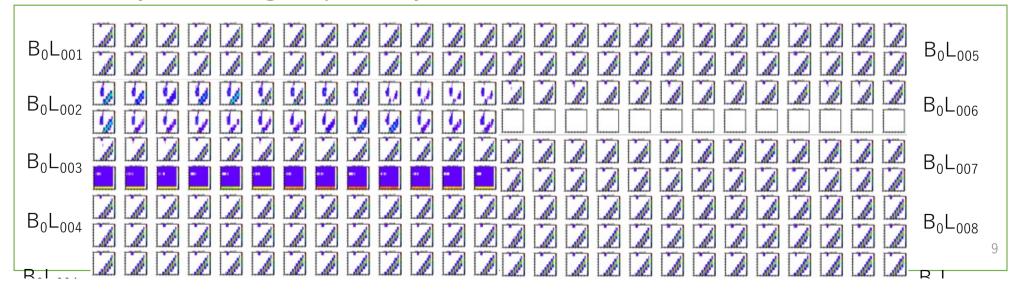
- 1. Apply 100V bias (HV GUI). Diagnose any over current channels. ✓
- 2. Power on a ladder by ladder (LV GUIs) and apply 100V bias. Run the calibration. Make sure the results appears in the expected ladder map in the Calibration Display/Analyzer/Monitor. (Misaki/Cheng-Wei) ✓
- 3. Diagnose missing channels and try to recover. ✓
- 4. Tune the alert range of LV/HV voltage/current control panels (alert features of LV/HV GUI). (Maya/Mai, still need fine tuning) ✓
- 5. Random trigger noise run (random external trigger). Debug any large noise half ladder or channels (online monitor). (Joseph, ongoing)
- 6. Save dead/hot channels in the database. (Expert GUI) (Hikaru, Itaru, ongoing)
- 7. Cosmic ray trajectory observation by the event display with the INTT self-trigger and standalone DAQ. If the big-partition and calorimeter trigger are available, we try to take data with the calorimeter external trigger.

Calibration Display



Done ✓

- Should have a calibration results at a glance. At least the results of $\frac{1}{2}$ barrel appears in a single page. Is it possible?
- Perhaps a calibration mode can be implemented to the OnlineMonitor, but #of hits/strip is not sufficient. We definitely need ADC vs. Amplitude 2D plots.
- Any warning capability like the online monitor?

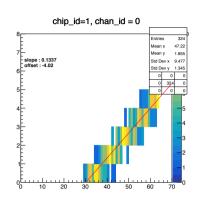


Calibration Result Analyzer Done v

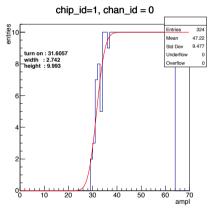
Cheng-Wei

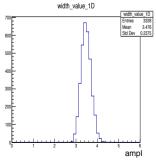
Calibration results are to be simplified into key characters and textualized such as:

- # of entries /chip (half entry, partial data drop to be detected) ✓
- ADC vs. Amp slope ✓
- Thresholds ✓
- Noise levels ✓
- # of bad channels/chip ✓
- These textualized results are to be saved in the calibration database. ✓
- Export the database to sPHENIX official db. (End of April?)



Fit function: error function





Cheng-Wei's slide 2021/8/11

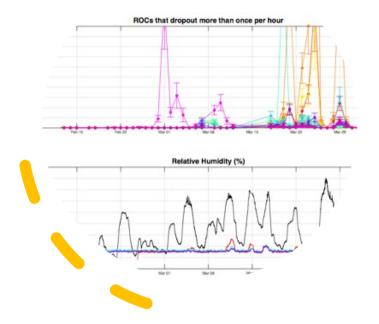
10

Calibration Monitor

Wei-Che?

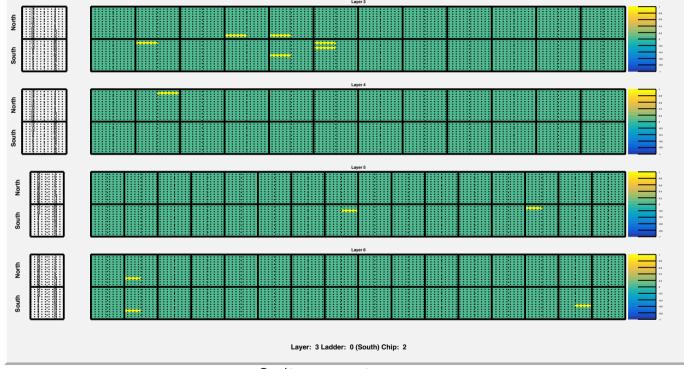
- Calibration run is to be executed regularly a couple of times a day.
- Results are to be monitored by the calibration monitor. The code is to be developed with the same framework with the stability monitor something like

%> StabilityMonitor.sh --calib



Noise rates

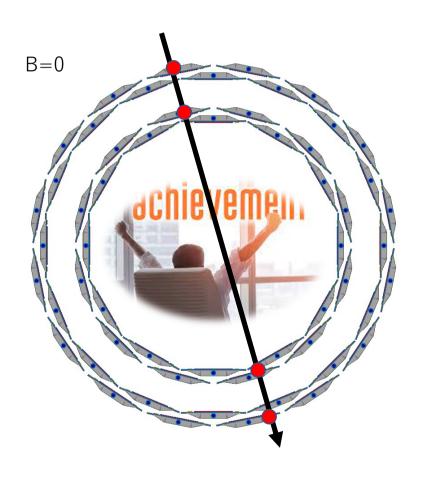




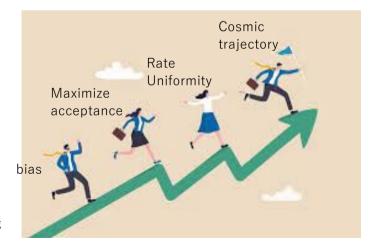
Online monitor

- Data should be taken with a random external trigger.
- Noise rate are to be compared with the <u>online monitor</u> and check the uniformity of the noise.
- Spot noisy channels and too cold channels and diagnose.
- The conversion software from Genki's roottree to Fun4All roottree is under development (4/24~).
- Once this is completed, the online monitor can spit out observed rates strip-by-strip.

Cosmic Ray Trajectory.

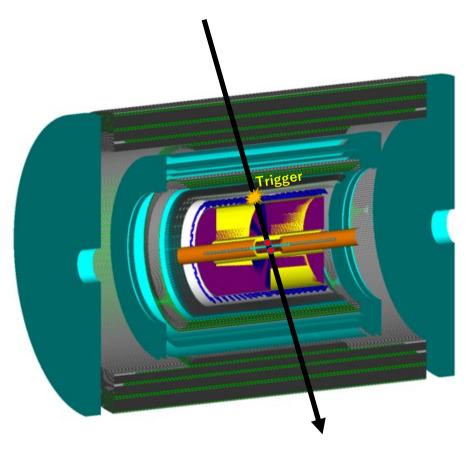


- If the calorimeter trigger becomes available, we may be able to take cosmic data.
- Without the fancy tracking algorithm, we should observe a straight track **by eye** with typical corresponding 4 clusters in each layer using the event display.
- The trigger can be the self trigger with the INTT standalone, but we may need to set relatively high DACO threshold to avoid pile up even after DACO threshold optimization.
- The latency should be set around the same setting with source/cosmic rays.





Cosmic Ray Trajectory (Advanced)



- More sophisticated data taking with:
 - Calorimeter external trigger
 - Big partition DAQ? Can be done with the INTT standalone rcdaq?
 - Latency scan with respect to the trigger timing.
 - event display

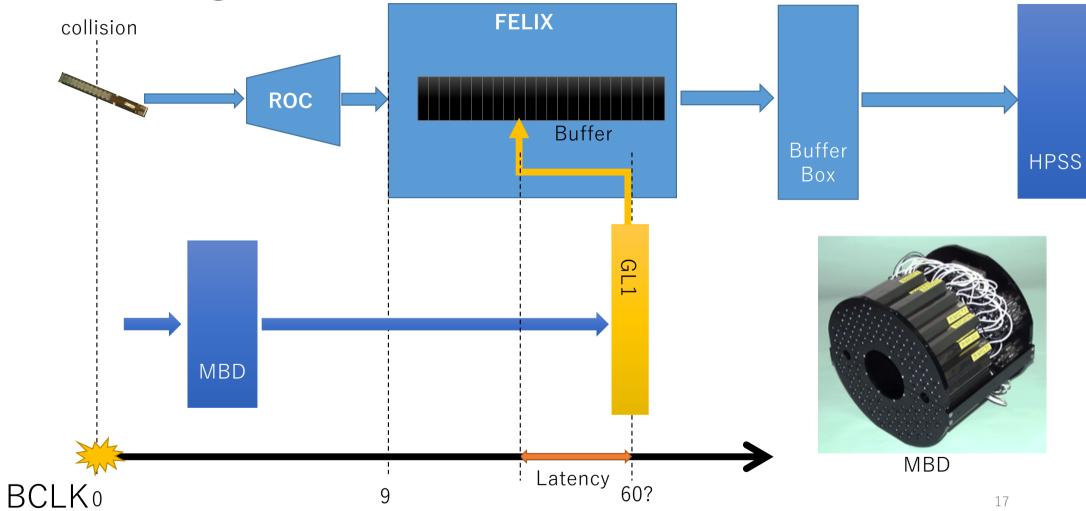


Commissioning with beam

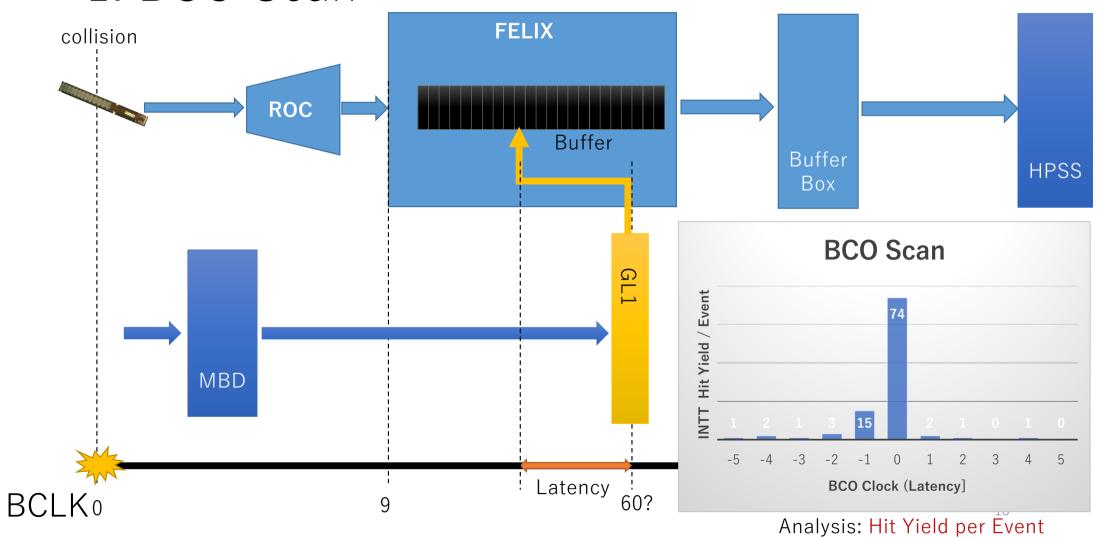
- 1. BCO Timing Scan (INTT Hit Yield/Event).
- 2. BCO Phase Scan fine tweak the timing with respect to BCO.
- 3. Mis-cabling check by the geometry (Event Display)
- 4. Diagnose missing channels and try to recover
- 5. Check yield uniformity (Online Monitor, Stability Monitor)
- 6. Gain matching between ladders or fine tweak noise
- 7. DAC Scan at HV=100V (DAC Scan Analyzer)
- 8. Bias Voltage Scan (MIP/MPV Fitter)
- 9. DAC0 threshold optimization. S/N evaluation chip by chip.

Timing Tune

Timing Tune



1. BCO Scan



2. BCO Phase Scan

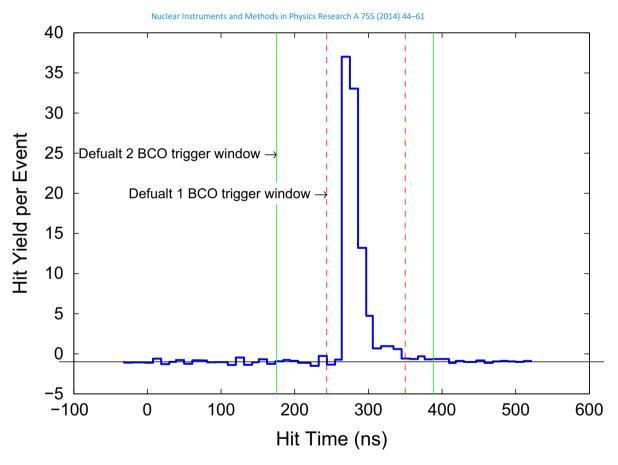


Fig. 32. Timing distribution of the FVTX hits relative to the RHIC beam clock.

6.1. Timing

The distribution in time of FVTX hits is studied relative to the RHIC collision time by comparing the hit rate at different FVTX delay values relative to the RHIC beam clock. The timing distribution for two sectors of wedges in the south arm is shown in Fig. 32. Most hits fall in a window ~ 30 ns wide.

Two standard trigger timing configurations were used during FVTX operation, as shown by the vertical lines in Fig. 32: during relatively low trigger rate running (in heavy ion systems) hits arriving in a time window two RHIC beam clocks (BCO) wide (1 BCO \sim 106 ns) are accepted. In high trigger rate p+p running, a 1 BCO-wide window is used to avoid recording accidental hits from neighboring beam crossings (1 BCO apart).

On 2023/01/12 22:22, Huang, Jin wrote:

That was exactly how it was done and highly recommended for intt too. It took few hours of a special low bunch fill to perform this scan, shifting BCO phase 19-20ns at a time. That appears the only way to set timing for the sub-bco delay Jin

2. BCO Phase Scan

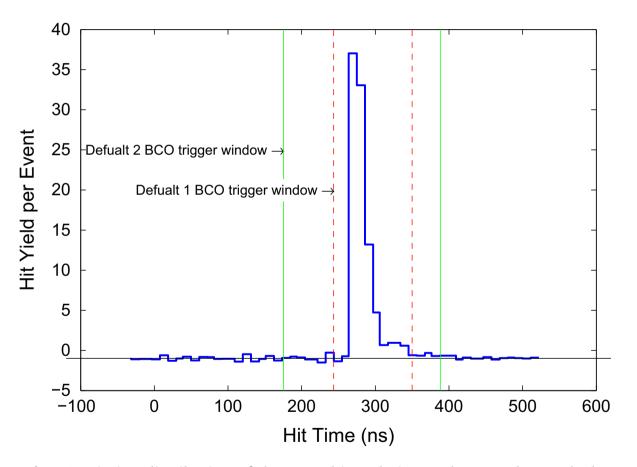
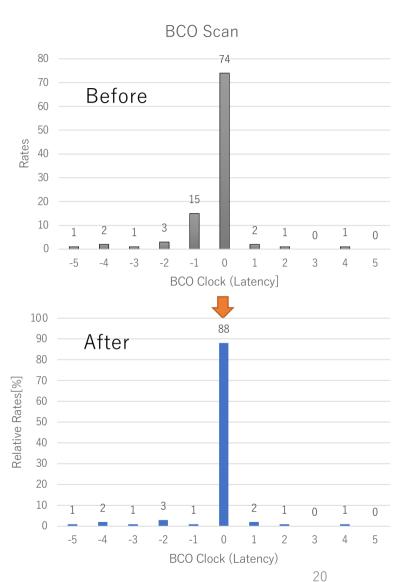
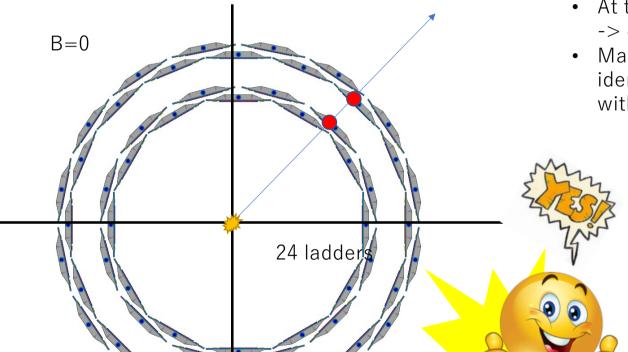


Fig. 32. Timing distribution of the FVTX hits relative to the RHIC beam clock.

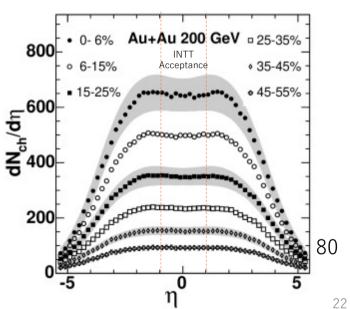


Ladder Geometrical Check

3. Hit Matching



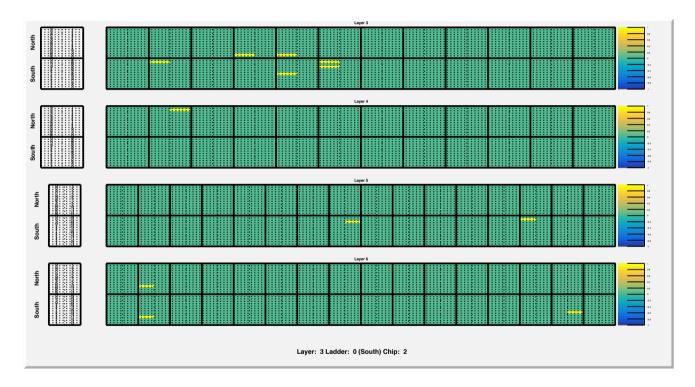
- In early stage of the commissioning, sPHENIX is operated with magnetic field off.
 - Tracks are expected to be straight.
 - At the 45-55% centrality collision, $80/|\eta|$ tracks -> 4 tracks/half ladder -> 0.15 hit/chip.
 - Matching hits between L0 and L1 can be identified by eye using the event display without fancy tracking algorithm.



https://wiki.sphenix.bnl.gov/index.php/INTT_GEANT_model/geometry#/media/File:2020-05-30-160330_940x871_scrot.png

Hit Rate Uniformity Check

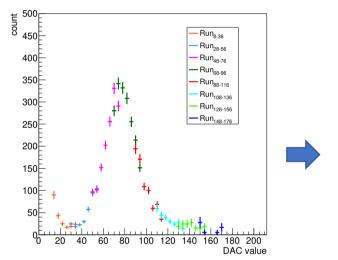
Hit Rate Uniformity Check



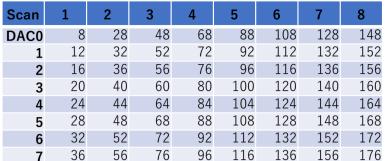
- Definitely need the online monitor working.
- Some non—uniformity can be observed by:
 - Bad cable contact
 - Dead channel
 - Hot channel
 - Gain variation
 - ...
- Dead/Hot channel detector is to be implemented.
- This check suppose to be executed periodically after the DAC0 threshold optimization

DAC Scan

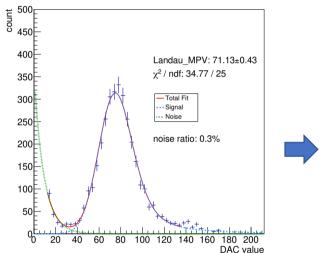
DAC Scan



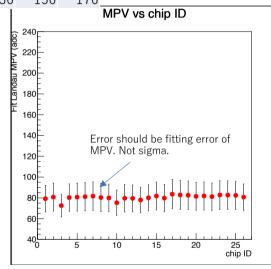
- Chip-by-Chip Base
- Clustering (Optimize offset value)
- Normalization btwn adjacent runs
- Concatenate all runs



ビームテストと同じ設定で良いか? オーバーラップ binは二つ?



 Fitting with Landau+Gaussian convolution function.



Half ladder by half ladder

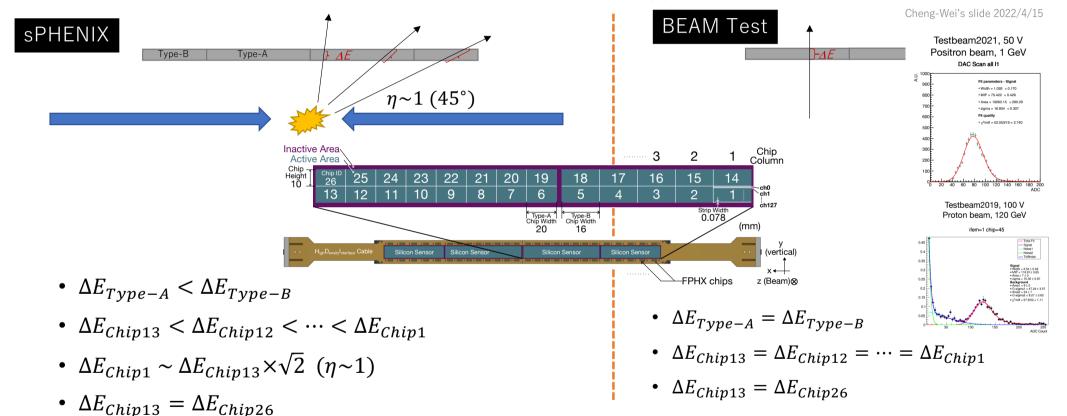


All ladders

Save all fitting parameters: MPV, Width, ...

How Energy Deposit Looks like in sPHENIX?

DACのオーバーフローが出ちゃうか?



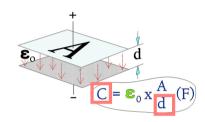
 $\Delta E_{Chip13}^{sPHENIX} @50V \sim \Delta E_{Chip13}^{ELPH2021} @50V$, $\Delta E_{Chip13}^{sPHENIX} @100V \sim \Delta E_{Chip13}^{FNAL2019} @100V$? To be studied by a **simulation** in advance.

Bias Voltage Scan

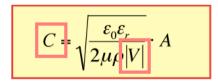
Bias Voltage Dependence

Cheng-Wei's slide 2022/4/15

sPHENIX INTT



$$W = \sqrt{2\epsilon (V + V_{bi})/Ne} = \sqrt{2\rho\mu\epsilon(V + V_{bi})}$$



$$\frac{dE/dx}{I_0} \cdot \frac{d}{3.87 \cdot 10^6 \text{ eV/cm} \cdot 0.03 \text{ cm}}{3.62 \text{ eV}} \approx 3.2 \cdot 10^4 \text{ e}^-\text{h}^+\text{-pairs}$$
Signal

Based on the theory : $C \propto \frac{1}{d} \propto \frac{1}{\sqrt{V}} \propto \frac{1}{signal}$

C: capacitance

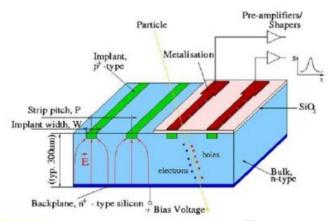
d: the distance of the depletion region

V : supply bias voltage

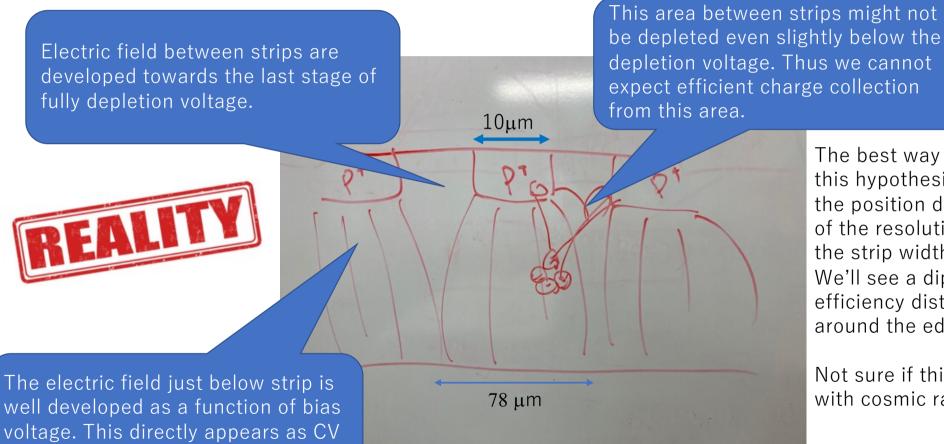
signal: edep

Itaru's Slide 2022/06/22

Principles of operation



Electric field at non-fully depleted voltage



response.

The best way to prove this hypothesis is to see the position dependence

the strip width.

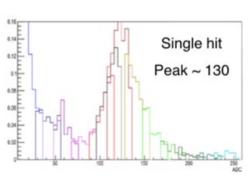
We'll see a dip in the efficiency distribution around the edge of a strip.

of the resolution within

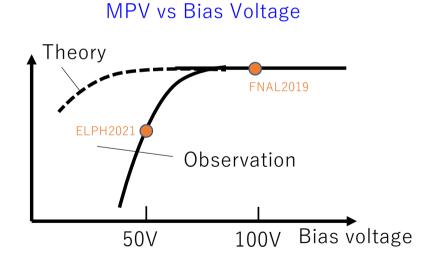
Not sure if this is doable with cosmic ray...

Bias Voltage Scan Plan

- Importance: It is likely we ends up with operating <100V due to over current of some silicons.
- We need to know the collecting # of electrons below 100V.

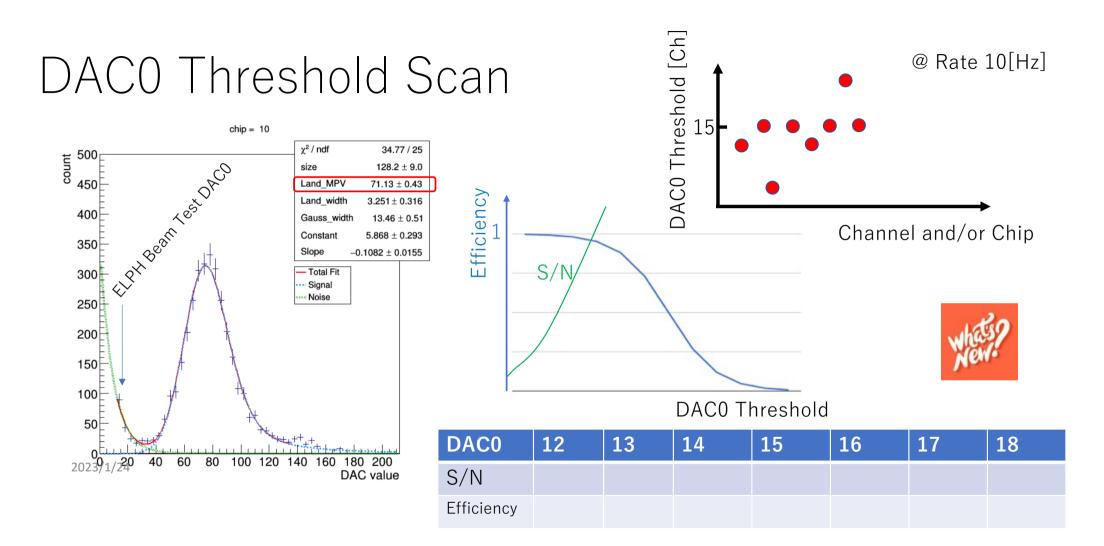


Scar	n	1	2	3	4	5	6	7	8
DAC	0	8	28	48	68	88	108	128	148
	1	12	32	52	72	92	112	132	152
	2	16	36	56	76	96	116	136	156
	3	20	40	60	80	100	120	140	160
	4	24	44	64	84	104	124	144	164
	5	28	48	68	88	108	128	148	168
	6	32	52	72	92	112	132	152	172
	7	36	56	76	96	116	136	156	176
								1	



- Scan at 90V, 80V, 70V, 60V, 50V only around MIP region.
- Need immediate semi-online analysis (DAC Scan code) if data is satisfactory to cover MPV peak.
- The goal is to make the plot of MPV vs. Bias voltage.
- Not sure if we can run a simulation.

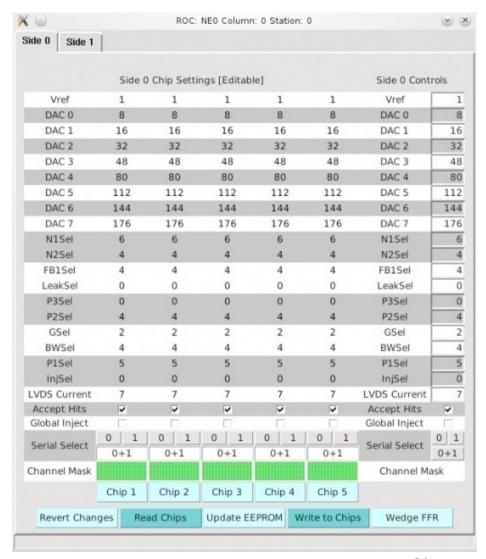
DAC0 Threshold Scan



• The final DACO threshold is to be determined by the S/N. The philosophy is to keep the signal fraction as high as possible within the tolerable noise rate. 33

FPHX Operation

FVTX Operating Condition



FPHX Parameters

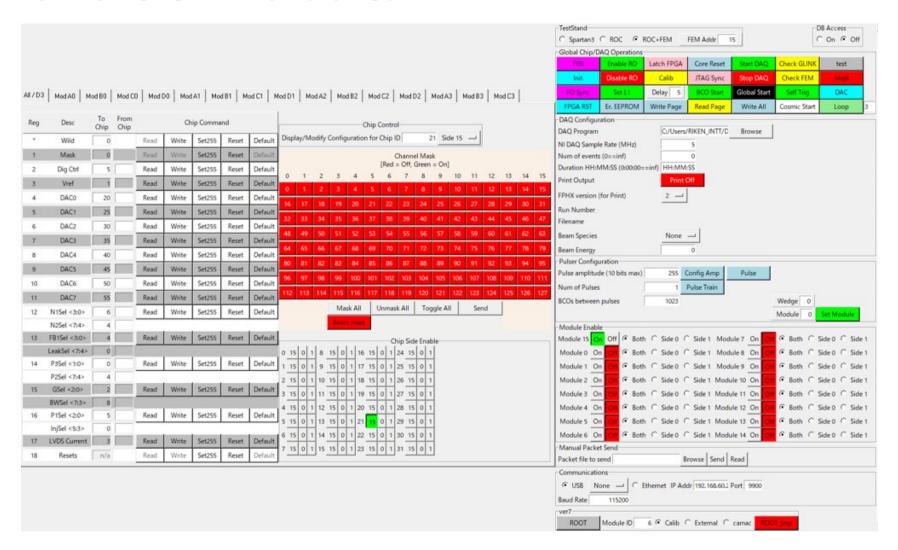
Name	Description	Default	FVTX	2021 Beam Test
Vref		1	1	1
DAC0		8	8	15
DAC1		16	16	30
DAC2		32	32	60
DAC3		48	48	90
DAC4		80	80	120
DAC5		112	112	150
DAC6		144	144	180
DAC7		176	176	210
N1Sel		6	6	6
N2Sel		4	4	4
FB1Sel		4	4	4
LeakSel		0	0	0
P3Sel		0	0	0
P2Sel		4	4	4
GSel	Gain Select	2	2	2
BWSel	Bandwidth Select	4	4	8
P1Sel		5	5	5
InjSel		0	0	0
LVDS Current [mA]		1	3	2







Nevis GUI Defaults

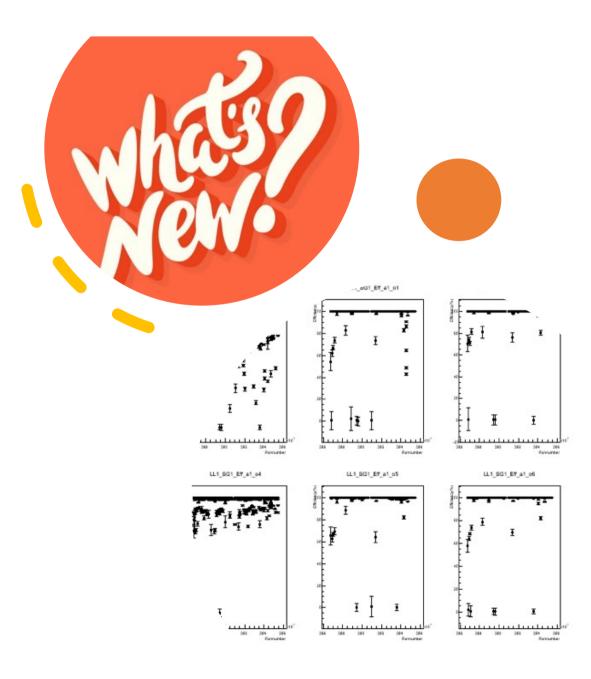


Monitoring (Advices from previous discussion)

- 1. Define online monitor. Develop and test anomaly (dead/hot channel) checker.
- 2. Establish flushing anomaly checker results to database.

Stability Monitor

- Normalized hit rates/Chip, ADC peak position, S/N, etc. are again, textualized (part of online monitor?) and be saved in the calibration database.
- Then we should develop the stability monitor by plotting above key performance parameters as a function of time.



Manpower operation during commissioning

Onsite Crews

- Rachid
- Genki
- Maya
- Itaru: April. 10th –
- Cheng-Wei: March 2nd -
- Joseph: Middle of Feb -
- Jaein: March 1st -
- Else?

Hardware debugging, data taking, change setting, logging incidents,...



Japan



Remote Crews









- Once we have a beam, the data taking is 24 hours.
- As long as we have a plan, we'll continue data taking with the INTT standalone DAQ independently from other subsystem data taking.
- The hall access will be coordinated with other subsystems.
- We will be blind immediately without the analysis code and cannot move forward!

Summary

- Various analyses and display codes are required for each measurement in commissioning. Need to prepare all necessary codes in advance!!
- Depending on the results, the INTT operating parameters are determined and the next measurement is made, so results are required immediately. Compared to single ladder analysis, the number of channels and the amount of data are overwhelmingly large.
- Data analysis is not an amount that can be handled on-site. Need feedback from remote analysis crews.
- Since we cannot have every INTT team members to be on-site, the commissioning program is designed to analyze the data remotely and provide feedback to the on-site.
- Volunteers are always welcome!

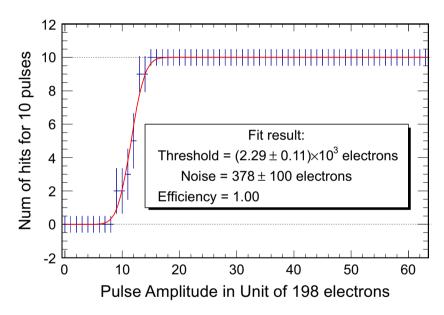


Fig. 35. Typical calibration data for a single channel (data points), fit with a normal cumulative distribution function.

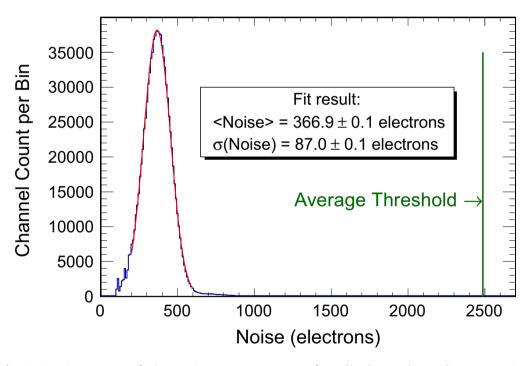


Fig. 36. Histogram of the noise parameter, σ , for all channels under operating conditions, in a typical calibration run. A Gaussian distribution fit to the data gives a mean noise level of 367 electrons. The nominal discriminator threshold at \sim 2500 electrons is shown by the vertical line.