



STAR Beam Use Requests for Runs 24-25

Lijuan Ruan (BNL)



Lijuan Ruan, BNL



28 cryo-weeks for p+p, p+Au in Run 24 28 cryo-weeks in Run 25 and 6 additional cryo-weeks in Run 24 for Au+Au

$\sqrt{S_{\rm NN}}$	Species	Number Events/	Year	
(GeV)	~p color	Sampled Luminosity	20002	Transversely polarized pp and p+Au with
200	p+p	$142 { m ~pb^{-1}/12w}$	2024	equal nucleon-nucleon
200	$p{+}\mathrm{Au}$	$0.69~{ m pb}^{-1}/10.5{ m w}$	2024	to optimize several
200	Au+Au	$18B / 32.7 \text{ nb}^{-1}/40w$	2023 + 2025	critical measurements

p+p: enable detailed evolution studies, critical for precise factorization and universality tests, essential baseline for p+Au

p+Au: probe gluon saturation, quark-gluon structure of heavy nuclei, propagation and hadronization of colored partons

Au+Au: probe the inner workings of the QGP



Physics program

quantitative comparisons of the validity and the limits of factorization and universality in lepton-proton and proton-proton collisions for initial and final state TMDs

Test of Sivers non-universality: Sivers_{SiDIS} = -- Sivers_{DY, W+/-,Z0}; Full jet and dijet Sivers asymmetry

Probe final state TMDs: Collins asymmetry for hadrons in jet

- Requirement:
 - > large data sets \sqrt{s} = 200 and 500 GeV p⁺p
 - \rightarrow low to high x, highest and lowest x with fSTAR
 - \rightarrow A_{UT} for W^{+/-} Z⁰, A_{UT} for hadrons in jet
- $\Box \quad \text{First look Gluon GPD} \rightarrow E_{g}$
 - Requirement:
 - > data sets $\sqrt{s} = 500 \text{ GeV } p^{\uparrow}p$ and $\sqrt{s} = 200 \text{ GeV } p^{\uparrow}A$
- $\hfill\square$ Physics driving the large A_N at forward rapidities and high x_F
 - > Requirement:
 - ➢ large data sets √s = 200 and 500 GeV p[↑]p
 → low to highest x_F → fSTAR
 - > charge hadron A_N at forward rapidities
- Nuclear dependence of PDFs, FF, and TMDs
 - > Requirement:
 - ➢ large equal data set of √s = 200 p[↑]p and p[↑]Au
 → low to high x, highest and lowest x with fSTAR
 - ightarrow R_{pA} direct photons and DY, hadrons in jet A_{UT}
- non-linear effects in QCD
 - Requirement:
 - ➢ large equal data set of √s = 200 p[↑]p and p[↑]Au
 → lowest-x through fSTAR
 - > dihadron correlations for $h^{+/-}$, γ -jet, di-jets

To address important questions about the inner workings of the QGP

- What is the nature of the 3-dimensional initial state at RHIC energies? r_n over a wide rapidity, $J/\psi \ v_{1,}$ photon Wigner distributions
- What is the precise temperature dependence of shear and bulk viscosity? v_n as a function of η
- What can be learned about confinement from charmonium measurements? $J/\psi \; v_2$
- What is the temperature of the medium? Different Y states, $\psi(2S),$ thermal dileptons
- What are the electrical, magnetic, and chiral properties of the medium? Λ , Ξ , Ω P_H and K*, ϕ , J/ ψ ρ_{00} , thermal dileptons, CME observables
- What are the underlying mechanisms of jet quenching at RHIC energies? What do jet probes tell us about the microscopic structure of the QGP as a function of resolution scale? γ_{dir} +jet I_{AA} , γ_{dir} +jet acoplanarity, jet substructure
- What is the precise nature of the transition near $\mu_B \mbox{=} 0 \mbox{? Net-proton C_6/C_2}$
- What can we learn about the strong interaction? Correlation functions

To inform EIC physics with photon induced processes:

- Probe gluon distribution inside the nucleus: vector mesons $(J/\psi),$ dijets (\ref{scalar})
- Search for collectivity and signatures of baryon junction: inclusive charge particles and cross sections, v_n, identified particle spectra



Plans for Run-24

28 cryo-weeks for p+p, p+Au in Run 24

28 cryo-weeks in Run 25 and 6 additional cryo-weeks in Run 24 for Au+Au

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1.2 (1.5) times the total luminosity in Run-15 p+p (p+Au)

2.7 (1.5) times the transverse lumi. in Run-15

Kinematic coverage for Collins and Sivers Asymmetry STAR covers 0.005<x<0.5

Polarization:60%



Transversely polarized pp and p+Au with equal nucleon-nucleon luminosities essential to optimize several critical measurements



Physics Opportunities in 2024

Central role played by 200 GeV pp:

- In most cases, similar measurements will be performed with 510 GeV and 200 GeV pp
- Very wide *x* coverage (0.005 < *x* < 0.5) by combining 200 and 510 GeV pp
 - 510 (200) GeV pp with the Forward Upgrade provides access to the lowest (highest) x value with jets and hadrons in jets over a wide range of perturbative scales
 - 200 GeV pp provides best coverage for the intermediate x range
 - provides best overlap with the x-Q² coverage of EIC
- Overlapping *x* coverage **enables detailed evolution studies**
- 200 GeV pp critical for precise factorization and universality tests
 - Best statistical precision for much of the kinematics overlapping with EIC
- 200 GeV pp essential baseline for 200 GeV p+Au
 - Must investigate gluon saturation in both pA and eA to verify universality
 - Precise probe of the quark-gluon structure of heavy nuclei
 - Explore the propagation and hadronization of colored partons

Must measure non-perturbative part of TMD experimentally!



Inclusive transverse single spin asymmetries at forward





Sivers and ETQS function

forward upgrade



Mid-rapidity Collins effect at 200 vs 510 GeV



 A_{UT} vs jet (p_{T},η) measures the collinear transversity distribution



 A_{UT} vs hadron (z,j_{T}) maps the Collins fragmentation function



Precision measurements at both energies probe TMD evolution and provide important cross-checks and essential x-Q² overlap with EIC

A_{UT} in p+Au: an alternative universality test and a unique look at spindependent hadronization

• Run-24 will reduce these uncertainties at 200 GeV by a factor of 2, enabling the most sensitive universality test with EIC data





Nuclear PDF



STAR



Small DY cross section $(10^{-6}-10^{-5} \text{ of} hadron)$: need suppress hadron to the order of 0.1% while maintaining a decent electron efficiency

With forward upgrades: hadron rejection power: 200-2000 for hadrons of 15-50 GeV electron efficiency: 80%

Drell-Yan : constrain nuclear sea quark distribution in a broad x range

Essential in testing fundamental universality properties of nPDFs combined with data from EIC



QCD non-linear effects

STAR, PRL129 (2022) 92501







Run-15 di- π^0 correlation:

away side area suppressed significantly, while the pedestal and away side widths change very little.

probe x down to 10⁻³



QCD non-linear effects



counting experiment of Di-jets in pp and pA Saturation: Disappearance of backward jet in pA



Forward rapidities at STAR provide an absolutely unique opportunity to have very high gluon densities → proton – Au collisions combined with an unambiguous observable STAR forward upgrade characterize non-linear effects with charged di-hadrons, γ -jet, di-jet



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STAR



RHIC is in the ideal kinematic region to measure nuclear effects compared to LHC

Novel QGP droplet substructure: toroidal vorticity



Opportunity to discover a novel vortical configuration in the subatomic fluid



Summary of 2024

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200 GeV p+Au:

- Gluon saturation in both pA and eA to verify universality
- Precise probe of quark-gluon structure of heavy nuclei
- Explore the propagation and hadronization of colored partons
- A unique opportunity to discover toroidal vorticity

Equal nucleon-nucleon luminosities in pp and pAu in Run-24 essential to optimize several critical measurements

Fully utilize forward upgrades and excellent PID over extended η coverage



Low material, PID capability over extended η and p_T , improved trigger capability forward π^0 , γ , e, Λ , charged hadron, jets

40 weeks physics data taking for 200 GeV Au+Au: 10 (Run 23) + 4 (Run 24) + 26 (Run 25)

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iTPC + EPD + eTOF Forward upgrades DAQ5k

A factor of 9 more minimum bias data compared to Run-14 + Run-16 A factor of 1.2 more luminosity for high-p_T trigger



Physics Opportunities for 2023 + 2025

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Time

Constrain longitudinal structure of initial state



$$V_n(\eta_a, \eta_b) = V_{n\Delta}(-\eta_a, \eta_b)/V_{n\Delta}(\eta_a, \eta_b)$$

 $V_{n\Delta}$ the Fourier coefficient calculated with pairs of particles in different rapidity regions

r_n sensitive to different initial state inputs:

- 3D glasma model: weaker decorrelation, describes CMS r₂ but not r₃
- Wounded nucleon model: stronger decorrelation than data

Precise measurement of r_n over a wide rapidity window will provide a stringent constraint



Impact parameter dependence of transverse momentum distribution of EM production is the key component to describe data;

 p_T broadening and azimuthal correlations of e^+e^- pairs sensitive to electro-magnetic (EM) field.

Is there a sensitivity to final magnetic field in QGP?

Precise measurement of p_T broadening and angular correlation will tell at >3 σ for each observable.

Fundamentally important and unique input to CME phenomenon. Lijuan Ruan, BNL



Constrain temperature dependence of η/s



Flow measurements at forward rapidity sensitive to η /s as a function of T.

Much more precise than previous PHOBOS measurements.



Deconfinement and thermalization





Global vorticity transfer

improved PID, extended η coverage by iTPC, and forward tracking



How exactly the global vorticity is dynamically transferred to fluid? How does the local thermal vorticity of the fluid gets transferred to the spin angular momentum?

Rapidity dependence of Λ , Ξ , Ω P_H at STAR, probe the nature of global vorticity transfer: Initial geometry and local thermal vorticity + hydro predict opposite trends.

Can we reconcile P_H with vector meson spin alignment ρ_{00} ? Strong force field effect?

Precise measurements of ρ_{00} of K^{*}, φ , J/ ψ will tell.



Charge dependent directed flow







0.8

Run 23 (Projection)

0

 $\overline{\Omega}^+$ (sss)

0.2 0.4 0.6

Chiral property



Low-mass dielectron measurement: lifetime indicator and provide a stringent constraint for theorists to establish chiral symmetry restoration at μ_B =0

Intermediate mass: direct thermometer to measure temperature

Enable dielectron v₂ and polarization, and solve direct photon puzzle (STAR vs PHENIX) Lijuan Ruan, BNL



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Chiral cross-over transition



Lattice QCD predicts a sign change of susceptibility ratio χ_6^B/χ_2^B at T_C The cumulants of net-proton distribution sensitive to chiral cross over transition at μ_B =0

Observed a hint of a sign change from peripheral to central collisions at 200 GeV $C_6/C_2 < 0$ at central collisions

High statistics measurements (10% statistical error for C_6/C_2 in central) will pin down the sign change



Significant cos2 $\Delta \phi$ azimuthal modulation in $\pi^+\pi^-$ pairs from photonuclear ρ^0 and continuum Modulation vs. p_T , shows a diffractive pattern structure

Theory (linear polarized photon + saturated gluons), sensitive to nuclear geometry and gluon distribution, closest to the gluon 3D tomography at EIC

Run23+25:

multi-differential measurements (vs. mass, rapidity, p_T): provide strong theoretical constraints, separate ρ^0 from continuum (Drell-Soding), investigate how double-slit interference mechanism affects the structure

Enable a similar measurement for J/ψ , a cleaner probe for gluon spatial distribution



Search for collectivity and signatures of baryon junction in photo-nuclear processes



 γ +Au process in UPC associated with a large rapidity asymmetry:

- Search for collectivity
- Study bulk observables

Further understand the origin of collectivity observed in small systems



Search for collectivity and signatures of baryon junction in photo-nuclear processes

improved PID, extended n coverage by iTPC, and forward tracking



³¹



Summary of 2023 + 2025

STAR is in an excellent position to address important questions about the inner workings of the QGP and inform EIC physics with photon induced processes

- What is the nature of the 3-dimensional initial state at RHIC energies? r_n over a wide rapidity, $J/\psi v_1$, photon Wigner distributions
- What is the precise temperature dependence of shear and bulk viscosity? v_n as a function of η
- What can be learned about confinement from charmonium measurements? $J/\psi v_2$
- What is the temperature of the medium? Different Υ states, ψ (2S), thermal dileptons
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- What is the precise nature of the transition near μ_B=0? Net-proton C₆/C₂
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- Probe gluon distribution inside the nucleus: vector mesons (J/ψ) , dijets (?)
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Proposed measurements based on our detector performances in past years and/or forward capabilities. STAR will be ready for physics data taking within a week.



Runs 24-25: 24 cryo-weeks each

Data taking:
9.5 weeks pp
9 weeks p+Au
22 weeks Au+Au

$\sqrt{s_{ m NN}}$	Species	Number Events/
(GeV)		Sampled Luminosity
200	$p{+}p$	$108 { m ~pb^{-1}}/9.5 { m w}$
200	$p{+}\mathrm{Au}$	$0.58 { m ~pb^{-1}/9w}$
200	Au+Au	$16B / 29.4 \text{ nb}^{-1}/36w$

Equal nucleon-nucleon luminosities in pp and p+Au essential to optimize several critical measurements

Au+Au: Decrease statistics by at least 11%, hard probes (jets and quarkonia), thermal dilepton, photon-induced processes (di-lepton and J/ψ) most impacted

p+p and p+Au: Decrease statistics by at least 24-16%, even larger impact on nuclear PDFs, fragmentation functions, and gluon saturation measurements since these require comparisons of the same observables measured in both p+p and p+Au

Run 24: 20 cryo-weeks

$\sqrt{s_{ m NN}}$	Species	Number Events/
(GeV)		Sampled Luminosity
200	p+p	$81 \ { m pb^{-1}}/7.5 { m w}$
200	$p+\mathrm{Au}$	$0.42~{ m pb^{-1}/7w}$



The current goals of data analysis call for timely analysis and publication of data taken. The data processing will have to keep pace with data taking, therefore the CPU processing will also have to scale with the data taken.

We updated computing plan in November 2021 and discussed with NPP management and SDCC.

After 2025, we could do with what we have in 2025 but it must be preserved for +3 years for CPU. HPSS storage does not change. The Xrootd/central storage will be needed for at least a decade; need to maintain software and computing person power for at least 5 years, for STAR data calibration, production, simulation, embedding, and analysis.





Nuclear data for space radiation protection

Not part of STAR physics program in the final RHIC phase but represents an opportunity for RHIC to contribute with some important nuclear data

- The Space Radiation Protection community has identified 3-50 GeV/n region as an area of need. https://doi.org/10.3389/fphy.2020.565954
- STAR has excellent light fragment capabilities.
- RHIC can deliver the ion beam species (C, AI, Fe) and energies (3-50 GeV/n) of need to the Space Radiation Protection community. STAR installed the targets of interest (C, AI, Ni) and is ready to take FXT data when opportunities arise.





Figure 112: Left panel: Installation of the targets and holder on the East side of the STAR detector. Right Panel: A view down the beam pipe showing the three targets (C, Al, and Ni) installed at STAR.

In total, two weeks of running including machine setup

Lijuan Ruan, BNL



Future opportunity II

Imaging shape and radial profile of atomic nuclei

$$\rho(r,\theta,\phi) = \frac{\rho_0}{1 + e^{(r-R(\theta,\phi)/a)}} \qquad R(\theta,\phi) = R_0 \left(1 + \beta_2 \left[\cos\gamma Y_{2,0} + \sin\gamma Y_{2,2}\right] + \beta_3 Y_{3,0} + \beta_4 Y_{4,0}\right)$$

- Collective flow measurements sensitive to nuclear deformation
- Understanding of the nuclear shape of current available systems not ideal: impact η /s extraction



 Step1: calibrate systematics using two species around ¹⁹⁷Au: ²⁰⁸Pb & ¹⁹⁸Hg (β₂=-0.11) at 200 GeV Pb: control on effects of Au deformation; precision on initial state and pre-equilibrium dynamics (energy dependence) vs. LHC Hg: two systems with known β₂ can triangulate the consistency of β_{2Au}.

Constrain η/s with improved understanding of initial state.

- Step2: explore more exotic regions for triaxiality and octuple
 Scan an isotopic chain: ¹⁴⁴Sm (β₂=0.08),¹⁴⁸Sm (β₂=0.14,triaxial),¹⁵⁴Sm (β₂=0.34)
 - large octuple expected/predicted for Z~56/N~88; compare ¹⁵⁴Sm (β_2 = 0.34) with ¹⁵⁴Gd (β_2 = 0.31)

Use hydrodynamics and flow measurements to perform precision cross-check of low energy nuclear physics.



Summary

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 - \blacktriangleright A_{UT} for J/ ψ in UPC
- $\hfill\square$ Physics driving the large A_N at forward rapidities and high x_F
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Impact of STAR science goals with no pAu data in Run 2024

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- \Box First look at gluon GPD \rightarrow E_g
 - **Requirement:**
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 - **Correlations for h**^{+/-}, γ -jet, di-jets

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Without pAu data, STAR's forward upgrade will not be fully utilized for its discovery potential and RHIC will lose important physics opportunities on the following:

- First look at gluon GPD $\rightarrow E_g$
- Probe Nuclear dependence of PDFs, FF, and TMDs
- Study Non-linear effects in QCD
- Discover a novel vortical configuration