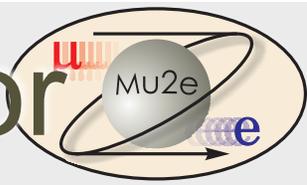


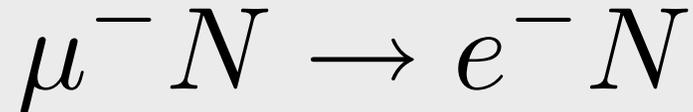
# A New Charged Lepton Flavor Violation Experiment: Muon-Electron Conversion at Sensitivity $< 10^{-16}$



for the Mu2e Collaboration

# What is $\mu e$ Conversion?

muon converts to electron in the field of a nucleus

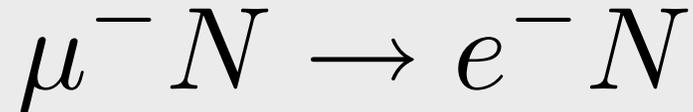


$$R_{\mu e} = \frac{\Gamma(\mu^- + N(A, Z) \rightarrow e^- + N(A, Z))}{\Gamma(\mu^- + N(A, Z) \rightarrow \text{all muon captures})}$$

- Standard Model Background of  $10^{-54}$
- Charged Lepton Flavor Violation (CLFV)
  - can measure a signal with SES of  $\sim 3 \times 10^{-17}$
- Related Processes:  $\mu$  or  $\tau \rightarrow e\gamma$ ,  $\tau \rightarrow 3l$ ,  $K_L \rightarrow \mu e$  and more

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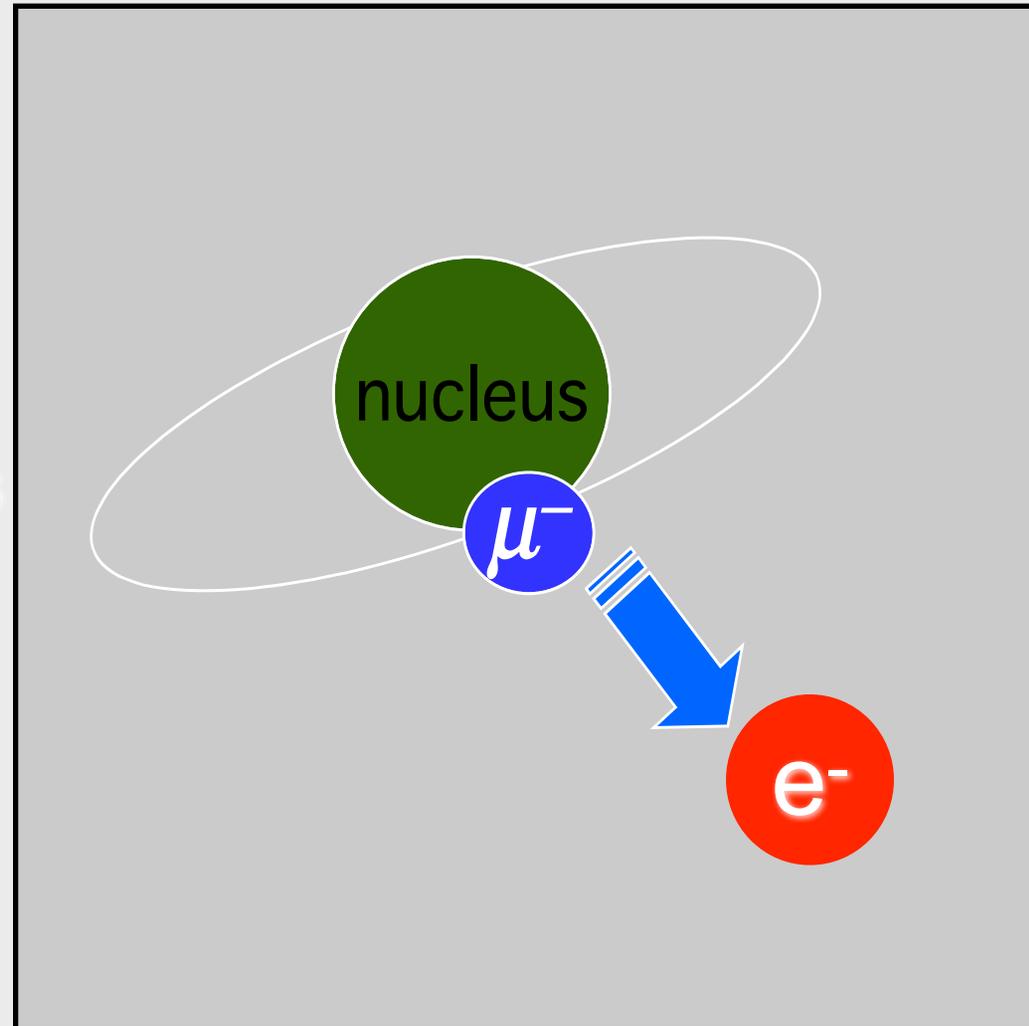
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# Experimental Signal



- A Single Monoenergetic Electron
- If  $N = \text{Al}$ ,  $E_e = 105. \text{ MeV}$ 
  - electron energy depends on  $Z$
- Nucleus coherently recoils off outgoing electron, no breakup





# “Who ordered that?”

– I.I. Rabi

After the  $\mu$  was discovered, it was logical to think the  $\mu$  is just an excited electron:

- expect  $\text{BR}(\mu \rightarrow e\gamma) \approx 10^{-4}$
- Unless another  $\nu$ , in Intermediate Vector Boson Loop, cancels (Feinberg, 1958)

➔ same as GIM mechanism!

•

# “Who ordered that?”

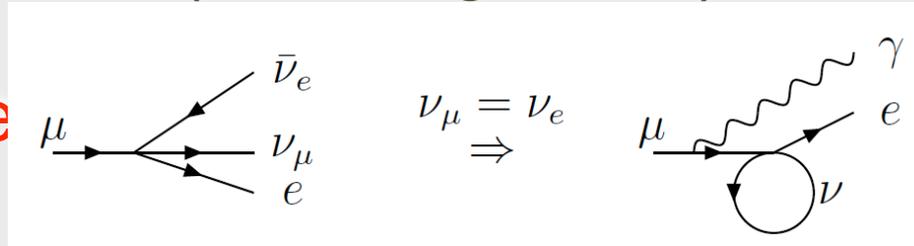


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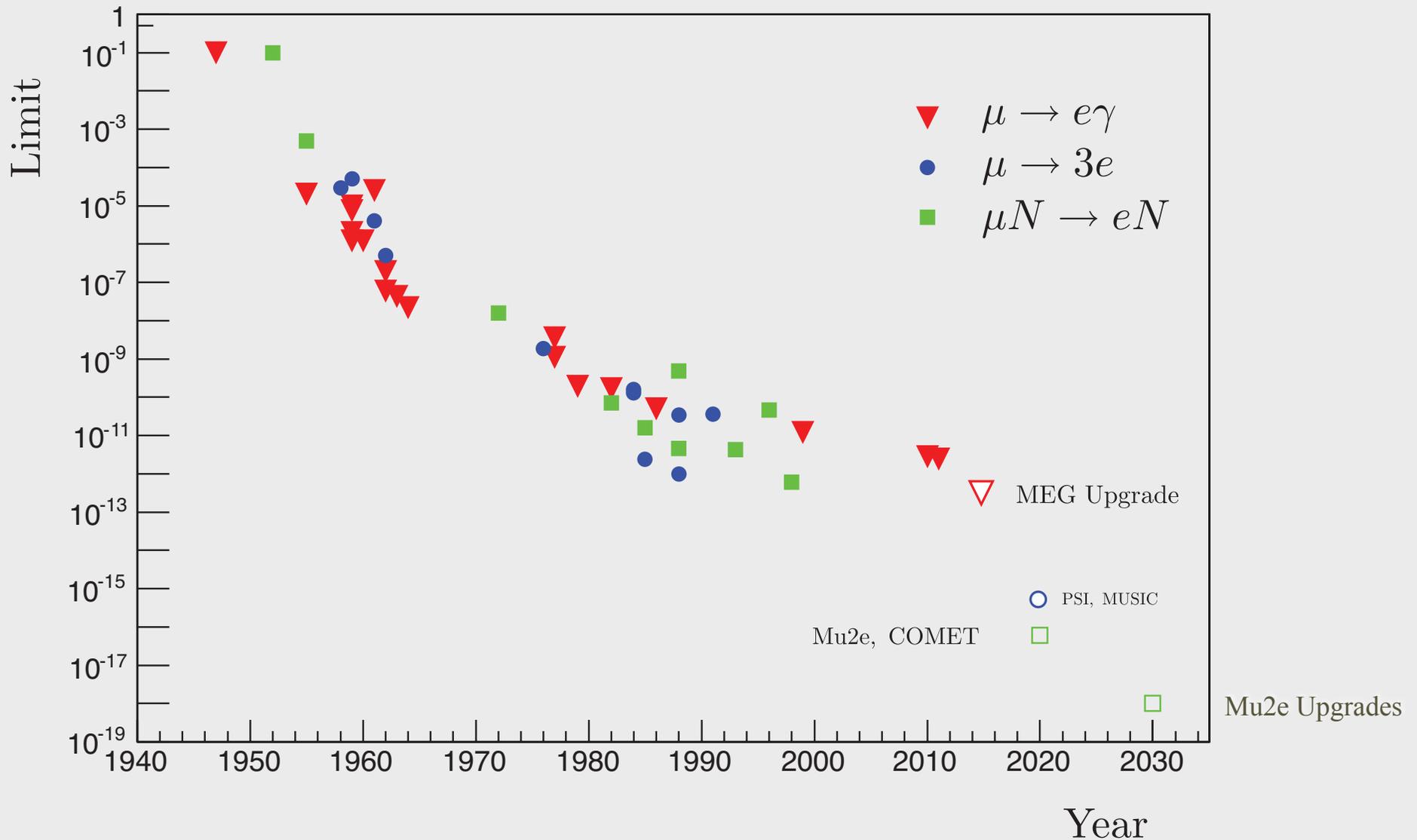
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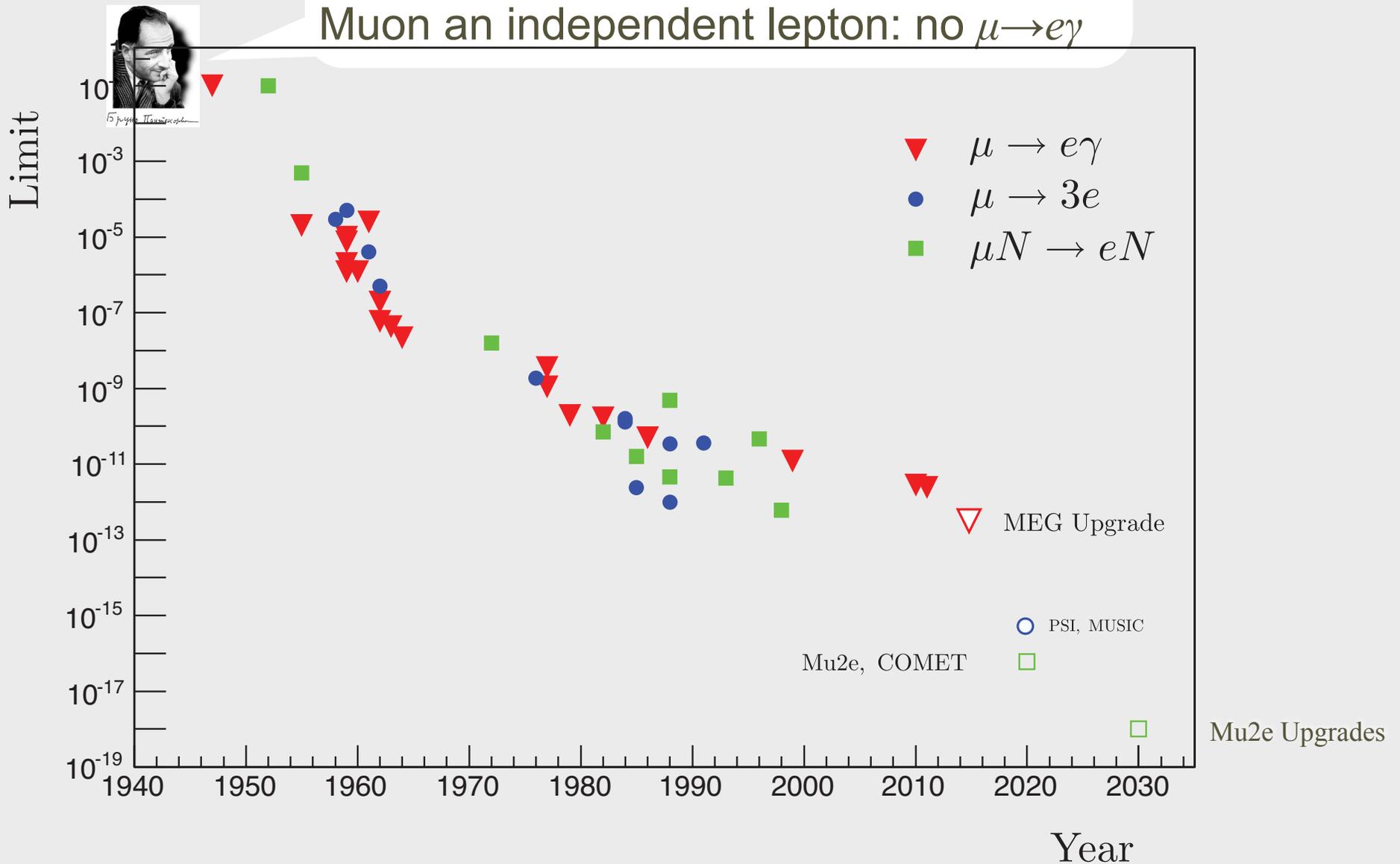
<sup>1</sup>Unless we are willing to give up the 2-component neutrino theory, we know that  $\mu \rightarrow e + \nu + \bar{\nu}$ .

# History of $\mu \rightarrow e\gamma$ , $\mu N \rightarrow eN$ , and $\mu \rightarrow 3e$



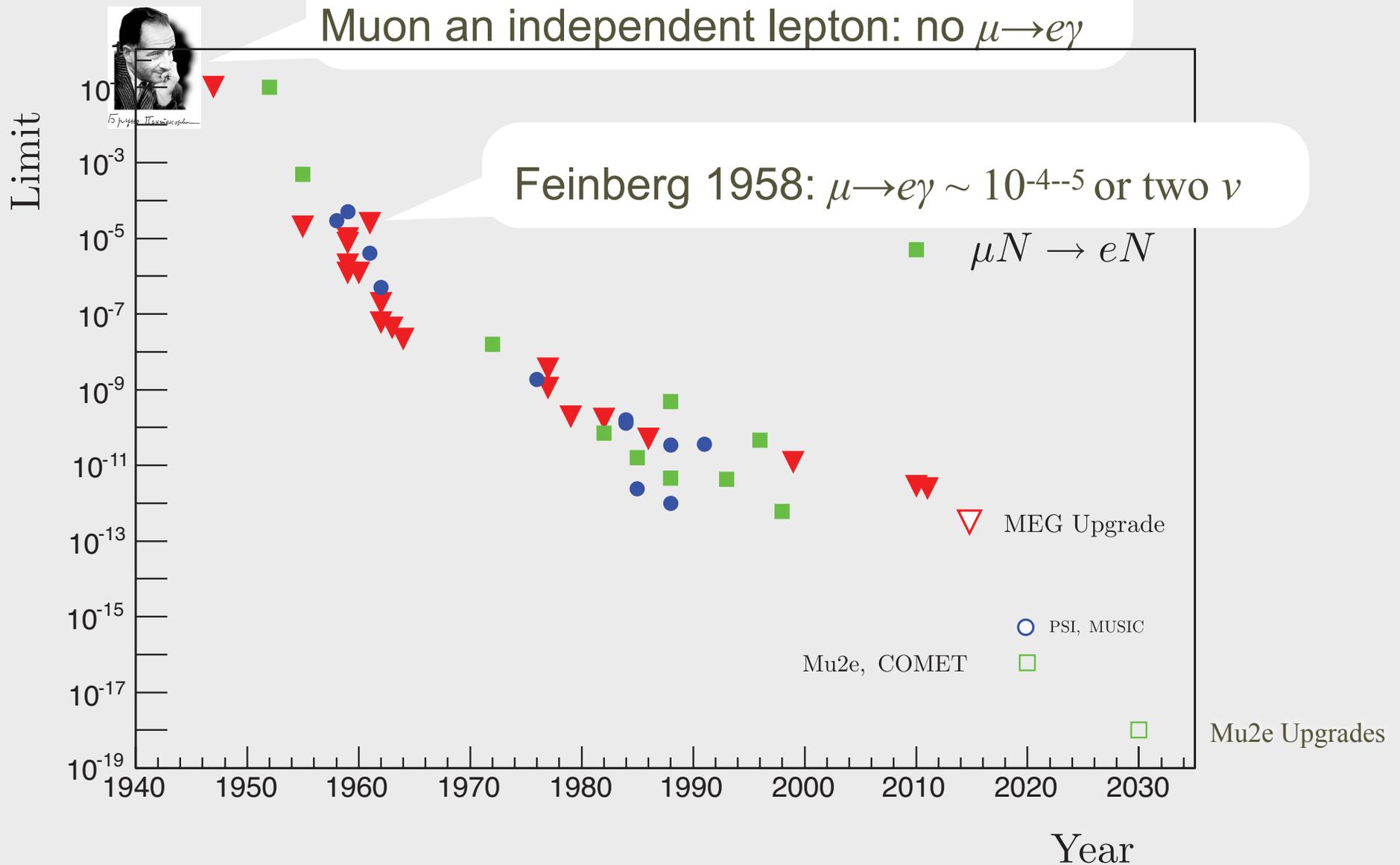
RHB and P.S. Cooper, Phys Rept C (1307.5787)

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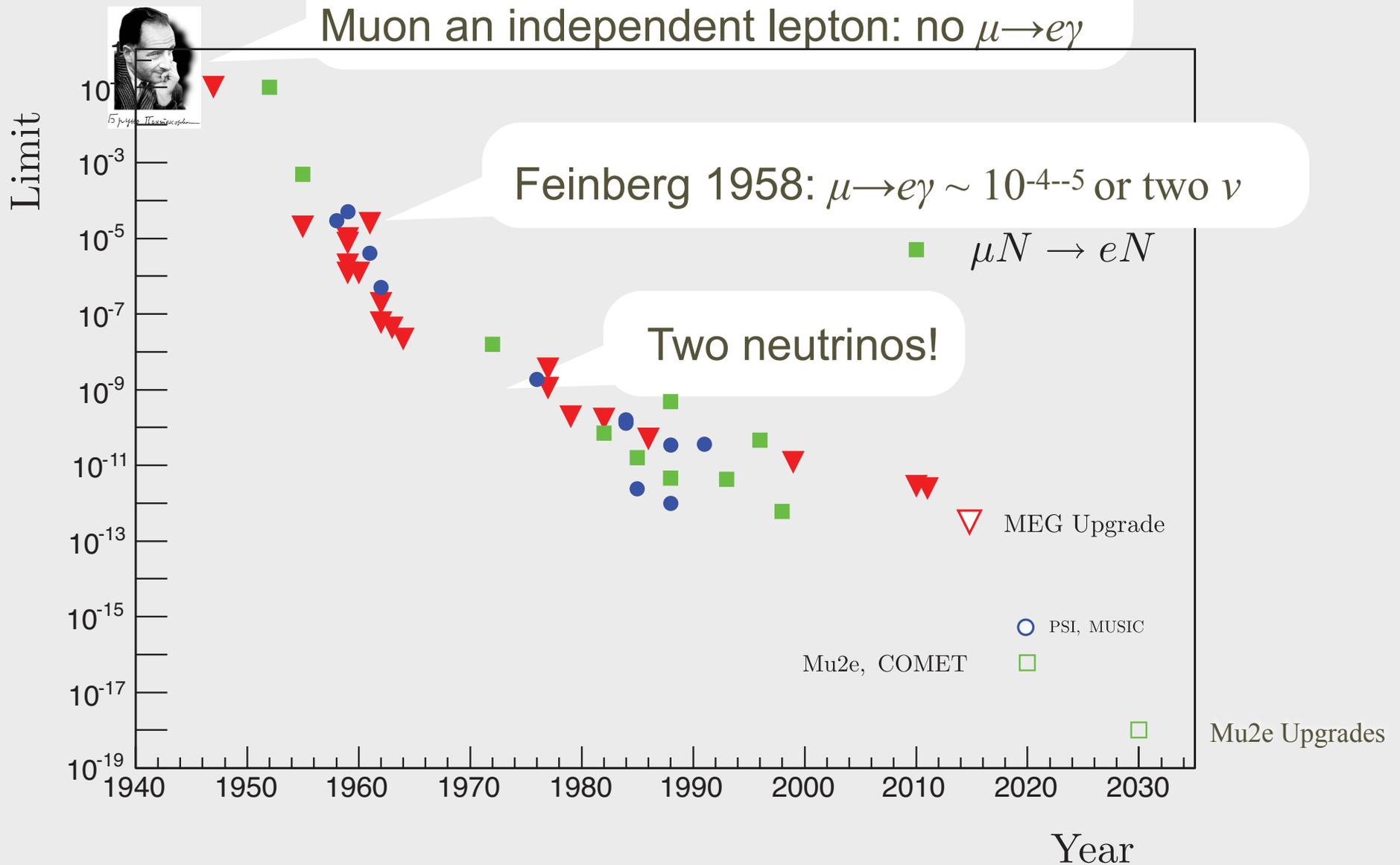
RHB and P.S. Cooper, Phys Rept C (1307.5787)

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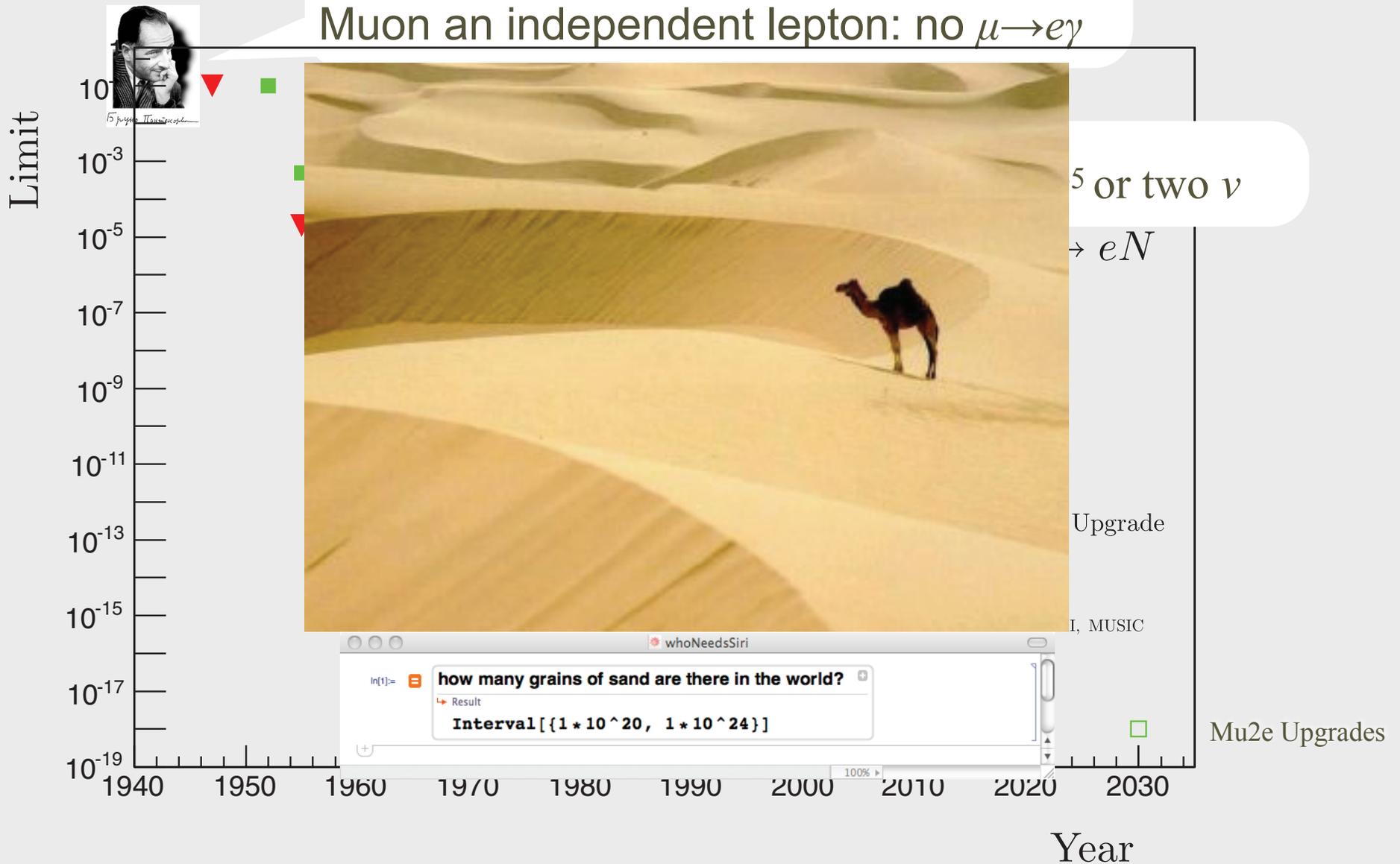
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RHB and P.S. Cooper, Phys Rept C (1307.5787)

# History of $\mu \rightarrow e\gamma$ , $\mu N \rightarrow eN$ , and $\mu \rightarrow 3e$



RHB and P.S. Cooper, Phys Rept C (1307.5787)

# How Rare is That?

- Pretty Rare: let us know if this happens to you!

Probability of...	
rolling a 7 with two dice	1.67E-01
rolling a 12 with two dice	2.78E-02
getting 10 heads in a row flipping a coin	9.77E-04
drawing a royal flush (no wild cards)	1.54E-06
getting struck by lightning in one year in the US	2.00E-06
winning Pick-5	5.41E-08
winning MEGA-millions lottery (5 numbers+megaball)	3.86E-09
your house getting hit by a meteorite this year	2.28E-10
drawing two royal flushes in a row (fresh decks)	2.37E-12
your house getting hit by a meteorite today	6.24E-13
getting 53 heads in a row flipping a coin	1.11E-16
your house getting hit by a meteorite AND you being struck by lightning both within the next six months	1.14E-16
your house getting hit by a meteorite AND you being struck by lightning both within the next three months	2.85E-17

thanks to Eric Prebys

# CLFV Muon Processes

- $\mu \rightarrow e\gamma$

- oldest studied, most powerful limits, and the best experiment so far: MEG at PSI

- $\mu N \rightarrow eN$

- muon to electron conversion: muon converts in field of nucleus, leaving nucleus unchanged

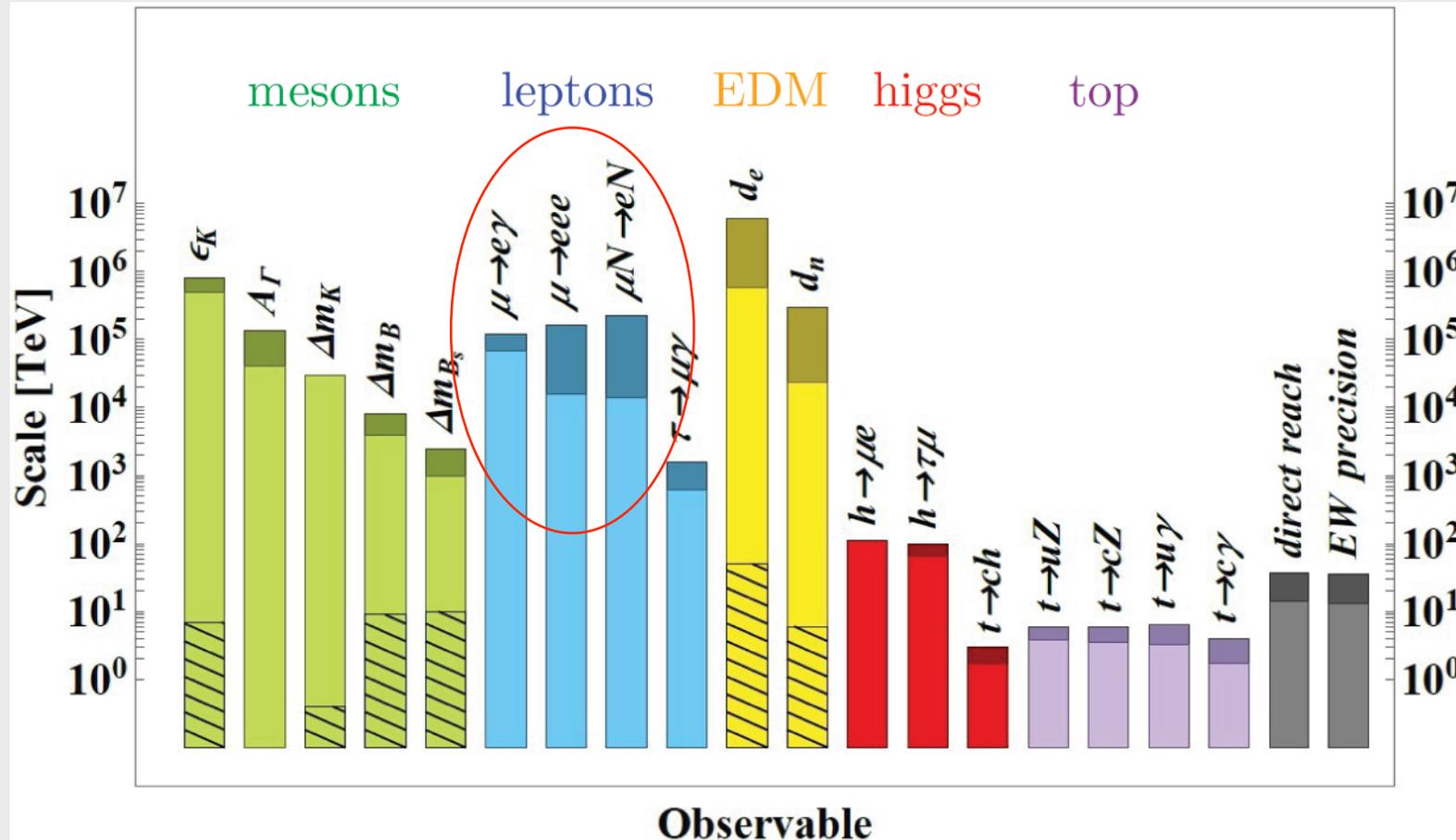
$$R_{\mu e} = \frac{\Gamma(\mu^- + N(A,Z) \rightarrow e^- + N(A,Z))}{\Gamma(\mu^- + N(A,Z) \rightarrow \text{all muon captures})}$$

- two experiments upcoming at FNAL and JPARC

- $\mu \rightarrow eee$

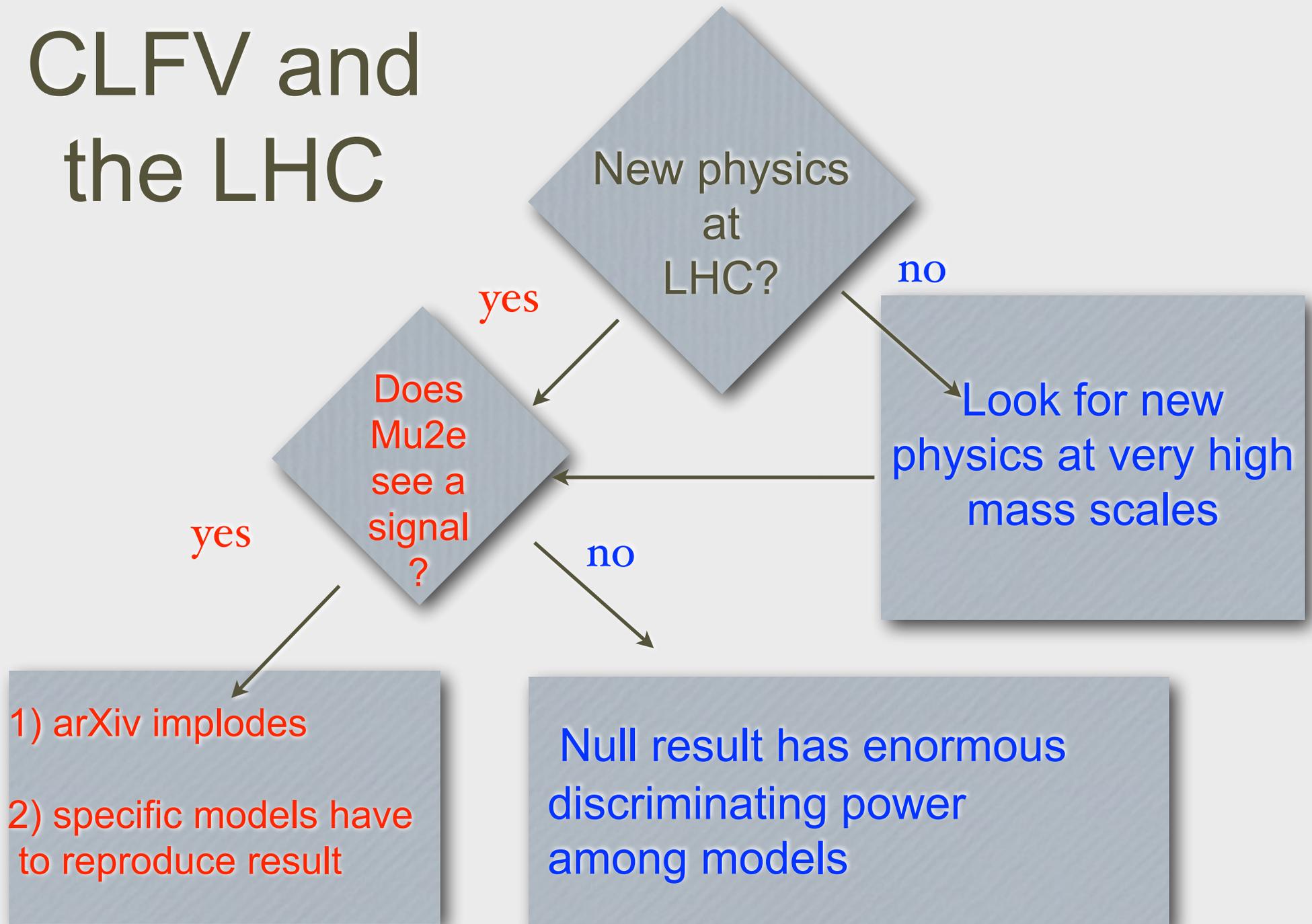
- ambitious and unique, excellent partner to other two (at PSI)

# Mass Scales of Muon CLFV Searches



operator coefficients = 1, from Physics Briefing Book, 1910.11775

# CLFV and the LHC

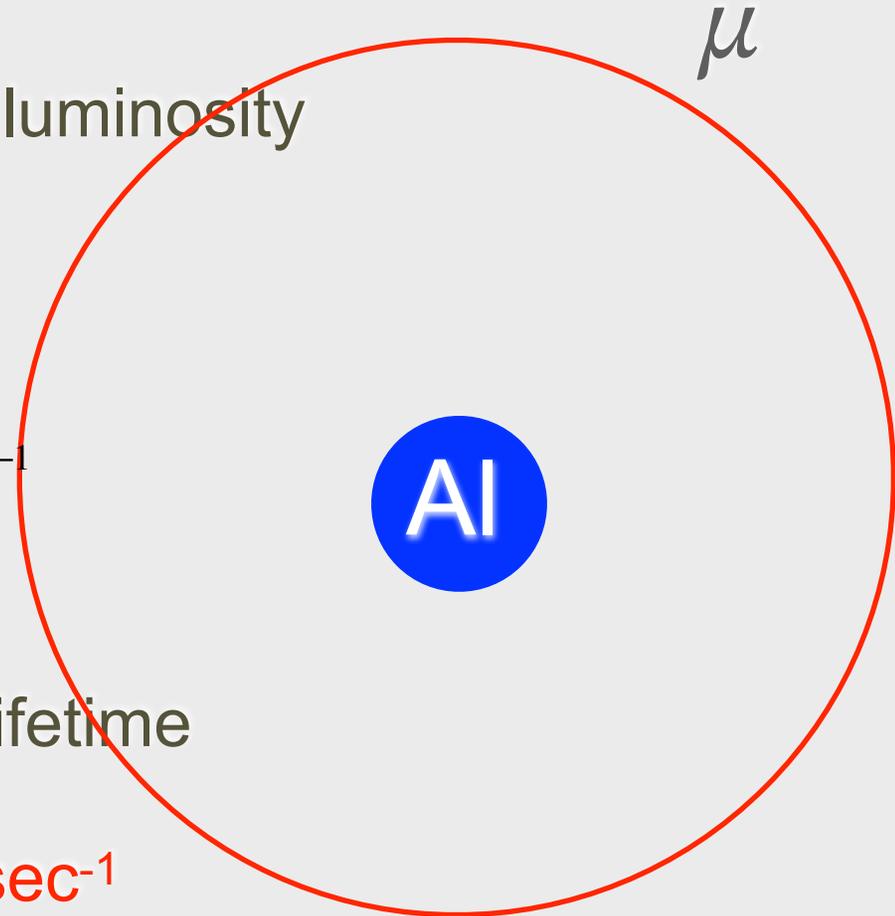


# Measuring $10^{-17}$ in Collider Units

- The captured muon is in a 1s state and the wave function overlaps the nucleus (*picture ~ to scale*)
- We can turn this into an effective luminosity
- Luminosity = density x velocity

$$|\psi(0)|^2 \times \alpha Z = \frac{m_\mu^3 Z^4 \alpha^4}{\pi} = 8 \times 10^{43} \text{ cm}^{-2} \text{ sec}^{-1}$$

- Times  $10^{10}$  muons/sec X 2  $\mu$ sec lifetime
- **Effective Luminosity of  $10^{48} \text{ cm}^{-2} \text{ sec}^{-1}$**

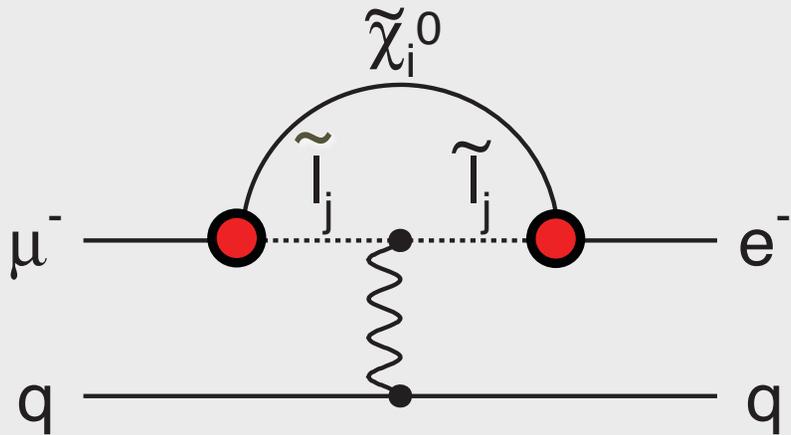


Andrzej Czarnecki

# LFV, SUSY and the LHC

## Supersymmetry

rate  $\sim 10^{-15}$



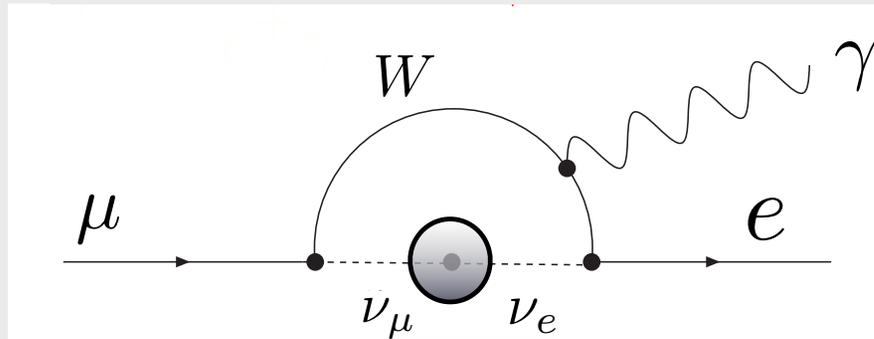
***Access SUSY  
through loops:***

***signal of  
Terascale at LHC  
implies***

***$\sim 40$  event signal /  
 $< 0.5$  bkg in this  
experiment***

# Neutrino Oscillations and Muon-Electron Conversion

- $\nu$ 's have mass! *individual lepton numbers are not conserved*
- Therefore Lepton Flavor Violation occurs in Charged Leptons as well



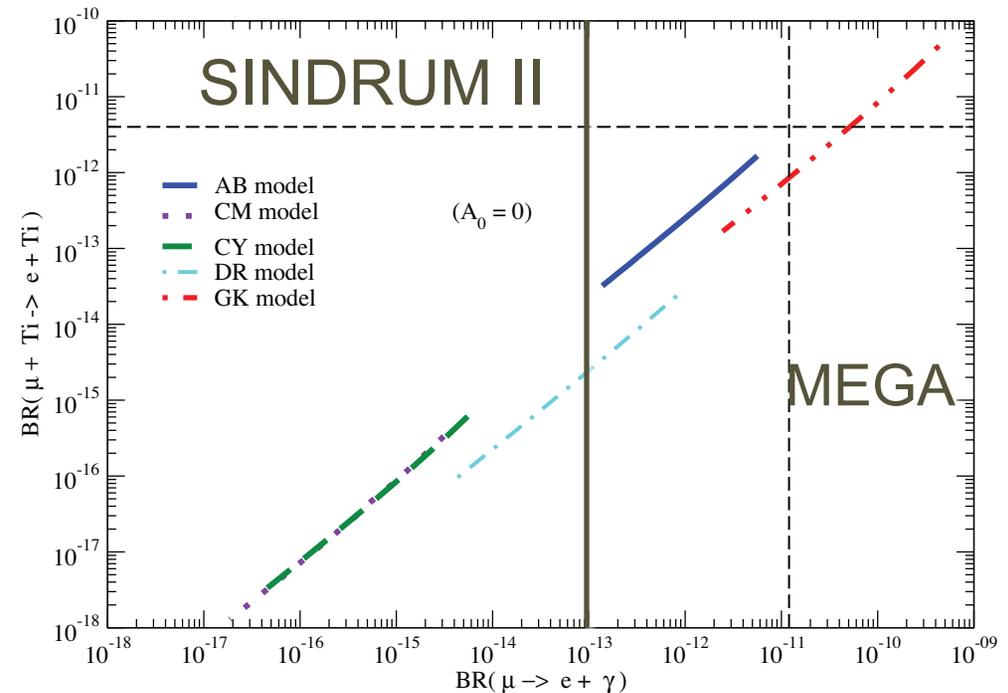
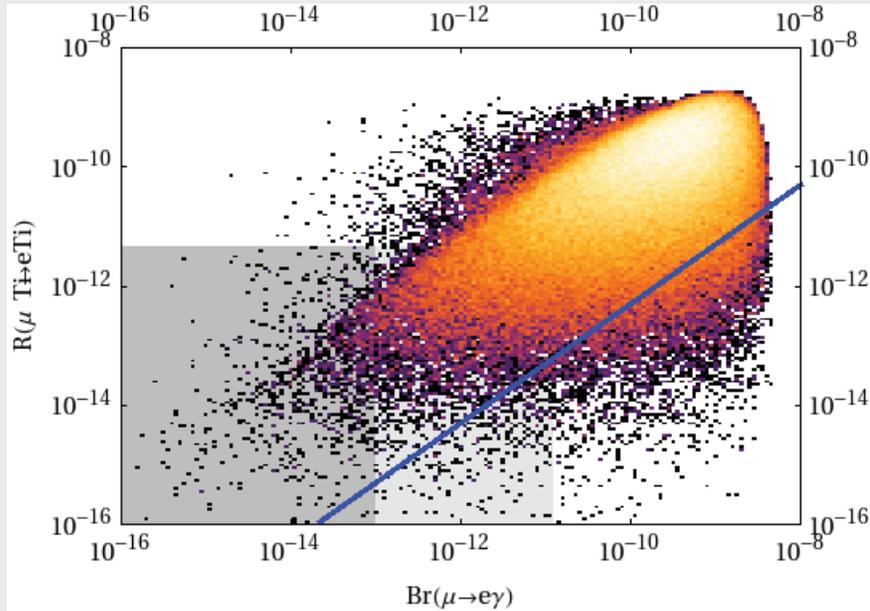
**NO STANDARD  
MODEL  
BACKGROUND**

$$\text{BR}(\mu \rightarrow e\gamma) = \frac{3\alpha}{32\pi} \left| \sum_{i=2,3} U_{\mu i}^* U_{ei} \frac{\Delta m_{1i}^2}{M_W^2} \right|^2 < 10^{-54}$$

# Combining $\mu \rightarrow e \gamma$ with $\mu \rightarrow e$ Conversion

## Littlest Higgs

Monika Blanke, Andrzej J. Buras, Bjoern Duling, Stefan Recksiegel, Cecilia Tarantino, Acta Phys.Polon.B41:657,2010,arXiv:0906.5454v2 [hep-ph]



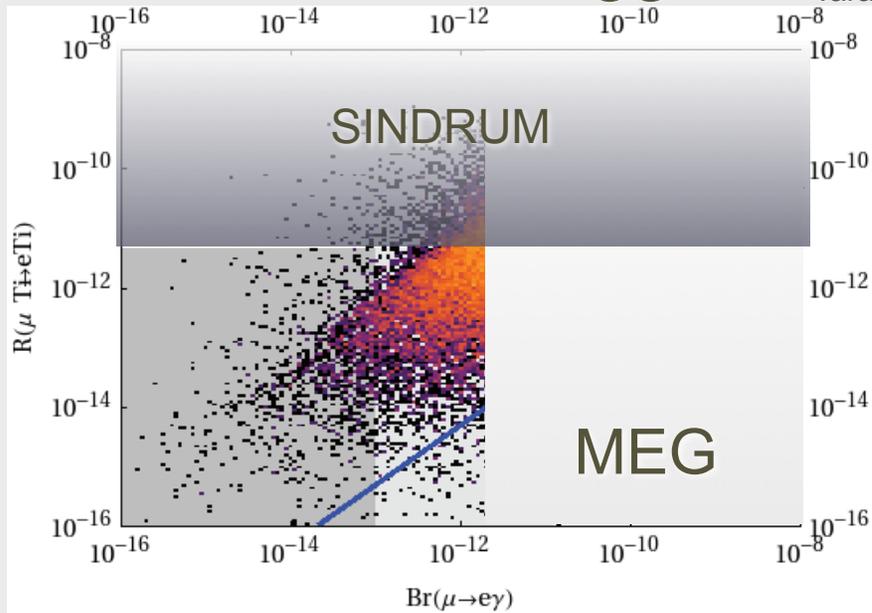
## SO(10) models:

C. Albright and M. Chen, arXiv:0802.4228, PRD D77:113010, 2008.

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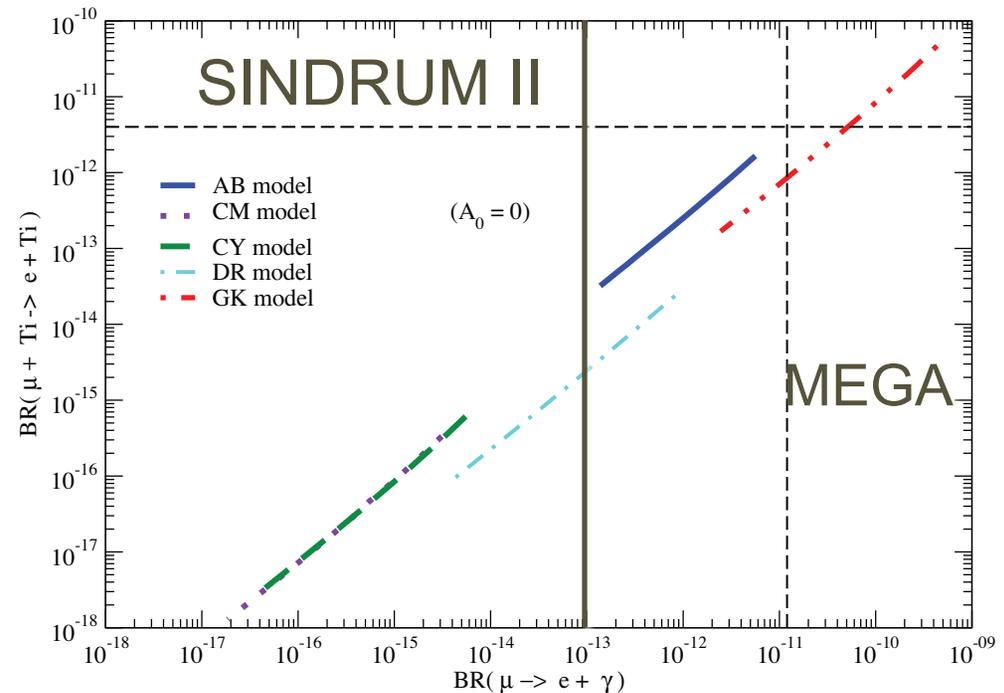
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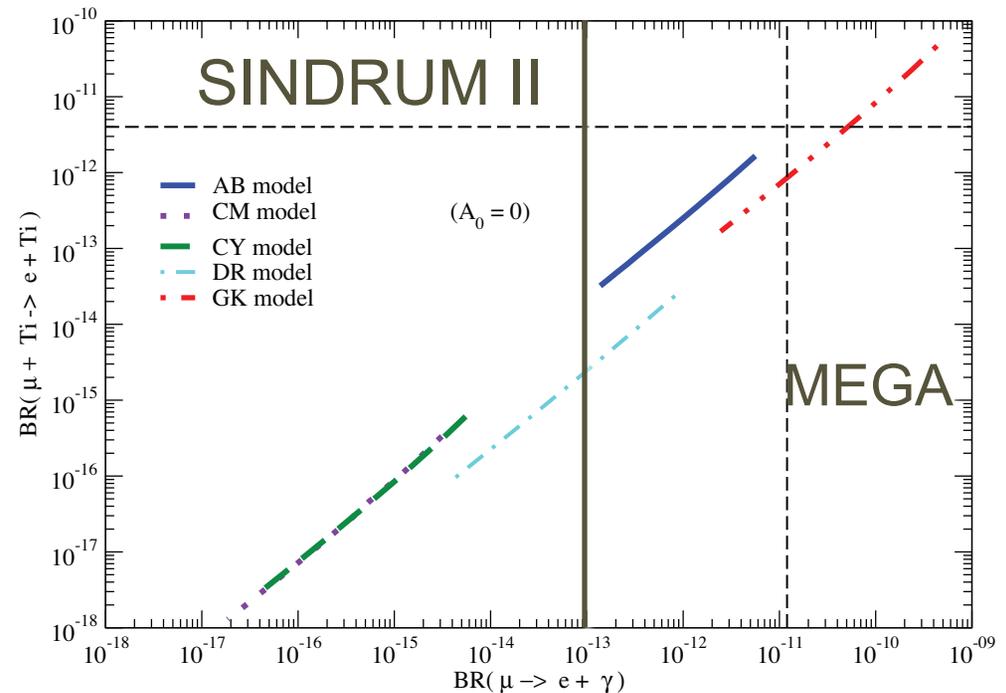
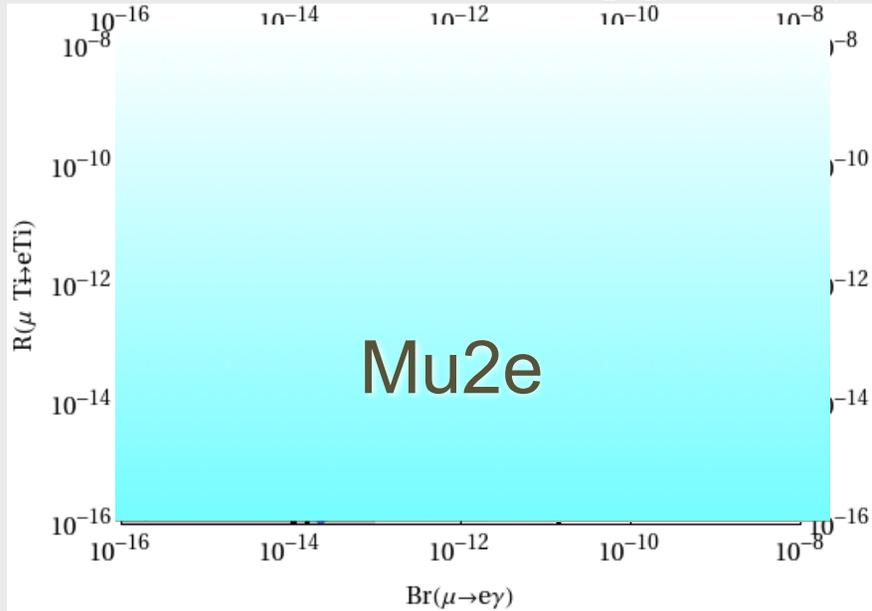
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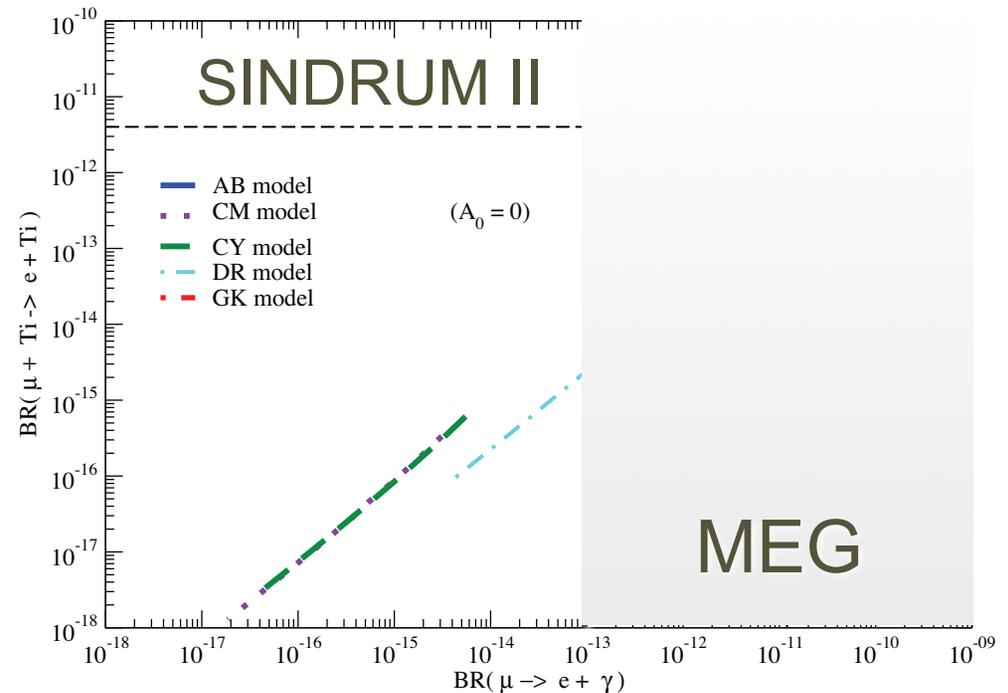
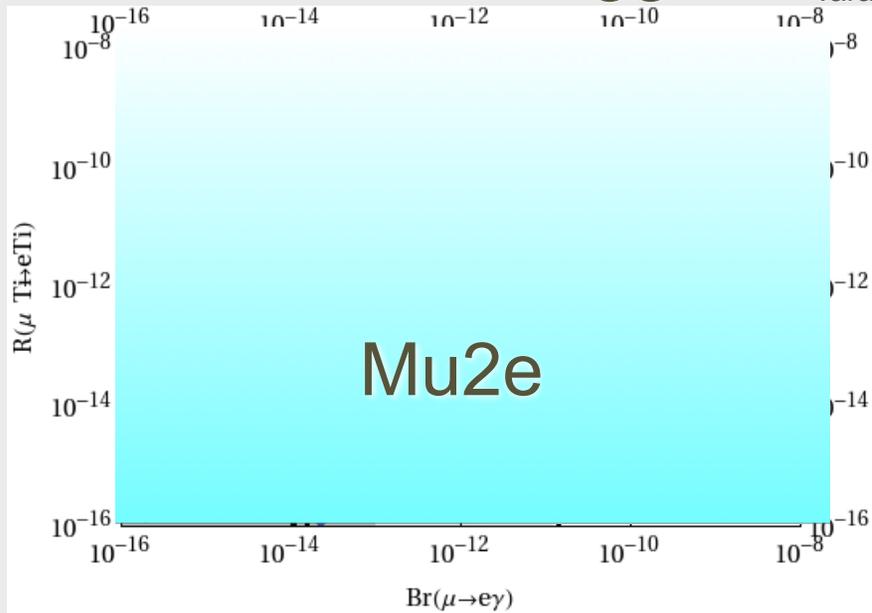
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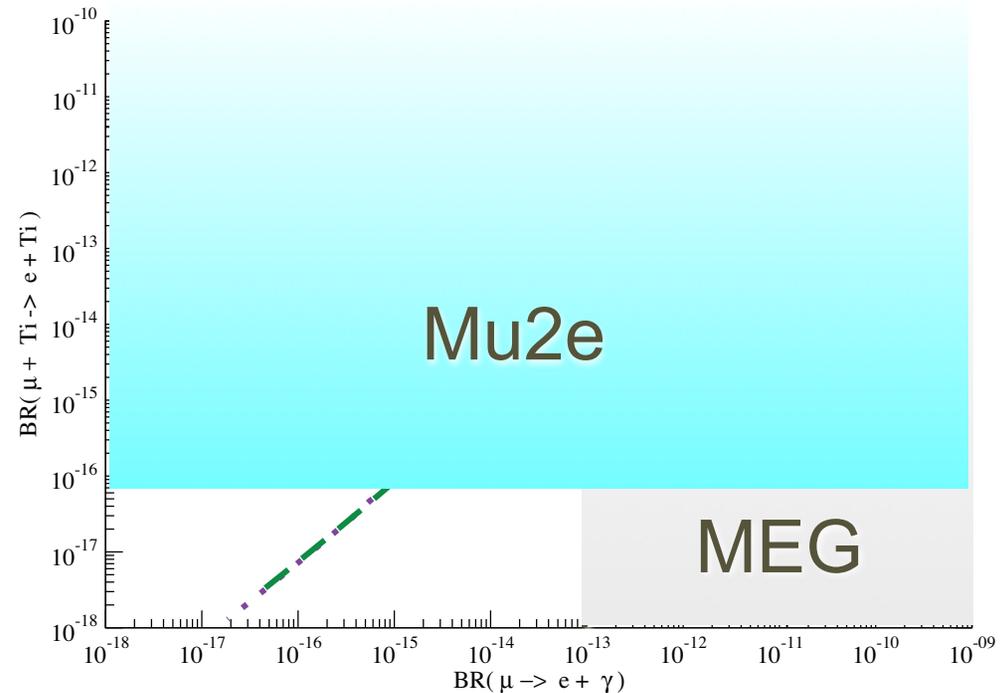
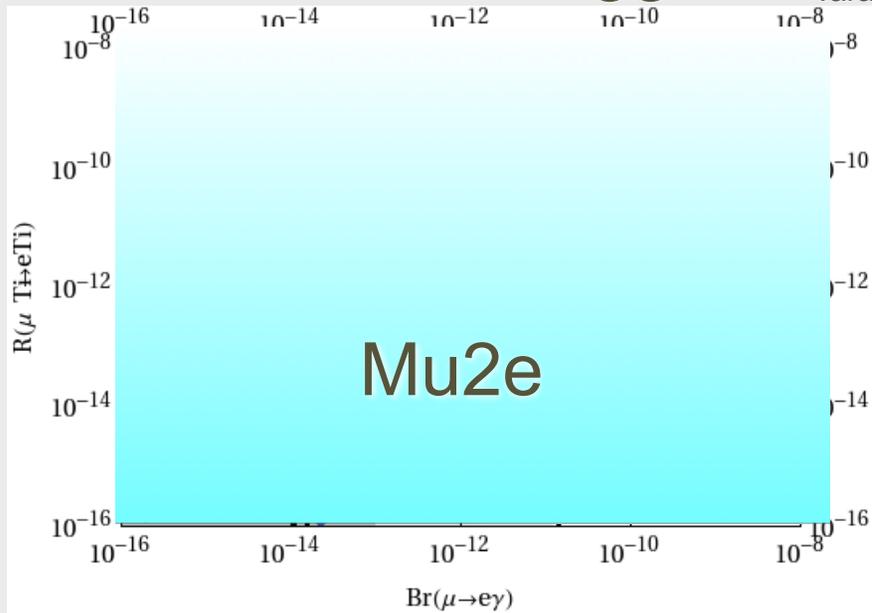
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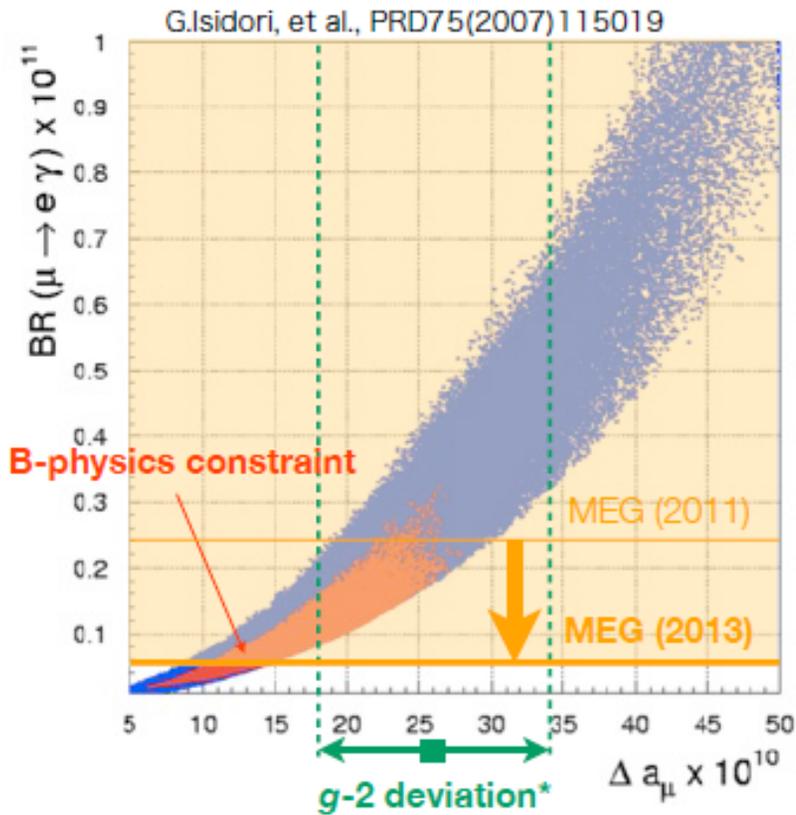
## SO(10) models:

C. Albright and M. Chen, arXiv:0802.4228, PRD D77:113010, 2008.

# Outside of CLFV As Well

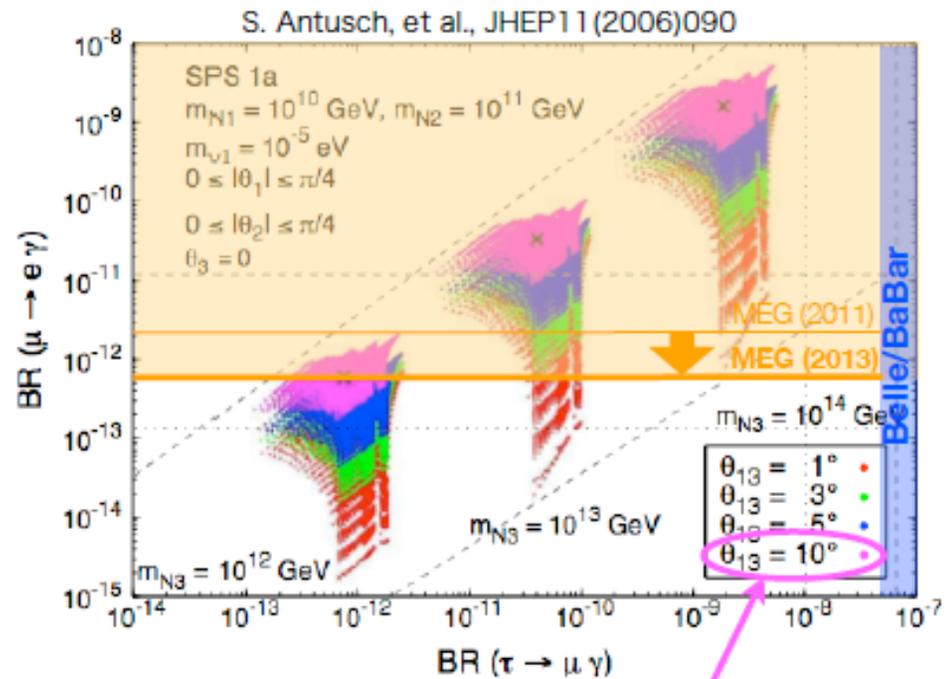
$\mu \rightarrow e \gamma$  and  $g-2$

SUSY-GUT



\*  $a_\mu(\text{EXP}): \text{PRD73}(2006)072,$   
 $a_\mu(\text{SM}): \text{Hagiwara et al., JPG38}(2011)085003$

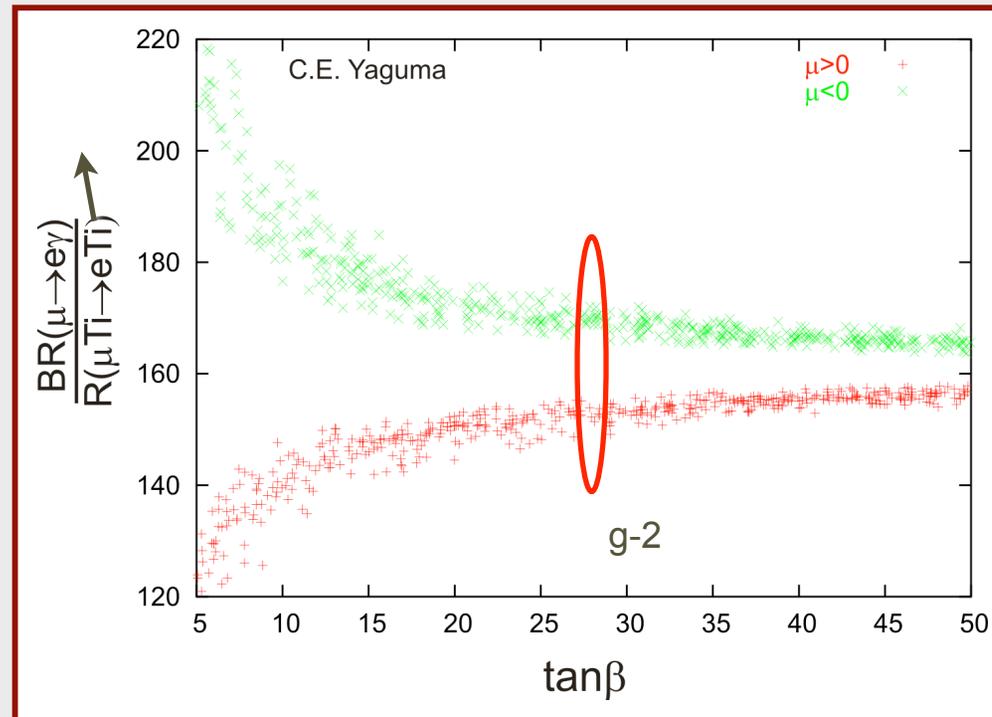
SUSY-Seesaw



Large  $\theta_{13}$  measured ( $\sim 9^\circ$ )!  
 $\mu \rightarrow e \gamma$  and  $\tau \rightarrow \mu \gamma$  and  $\theta_{13}$

# Mu2e, g-2, and $\mu \rightarrow e\gamma$

Yaguna,  
hep-ph/0502014v2  
MSSM w mSUGRA



- Need:

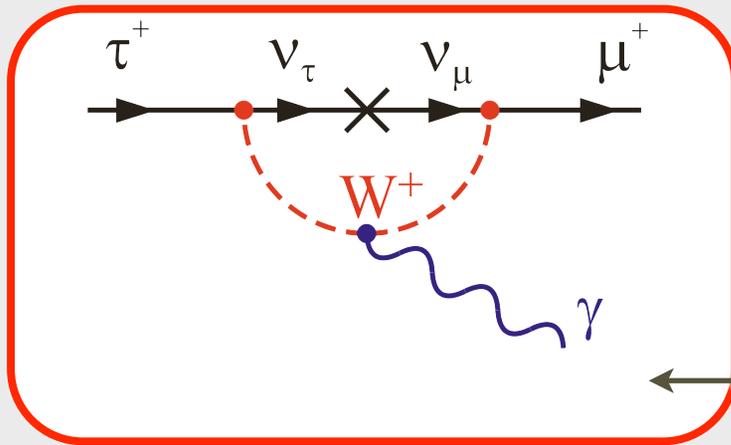
- observation of CLFV in more than one channel, and/or
- evidence from LHC, g-2, or elsewhere

to allow discrimination among different models

# CLFV and Tau Decays

$\tau$  processes also suppressed in Standard Model but less:

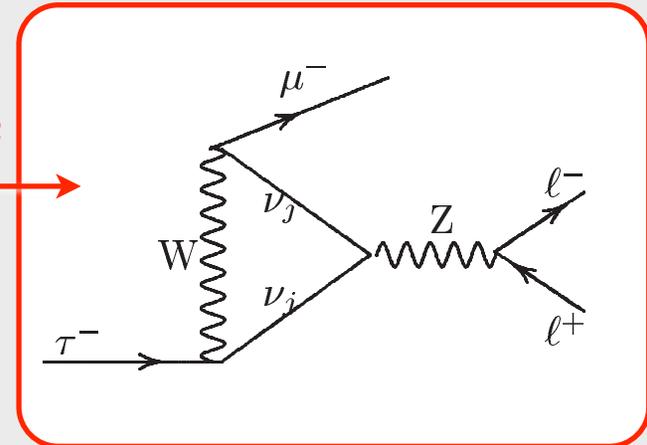
Lee, Shrock  
Phys.Rev.D16:1444,1977



SM  $\sim 10^{-49}$

$$\ln \left( \frac{m_3^2}{M_W^2} \right)^2$$

$$\left( \frac{\Delta m_{23}^2}{M_W^2} \right)^2$$



SM  $\sim 10^{-14}$  ?

Pham, hep-ph/9810484

## Good News:

Beyond SM rates can be orders of magnitude larger than in associated muon decays

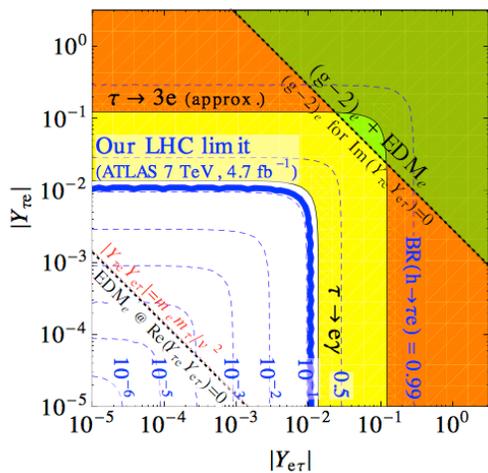
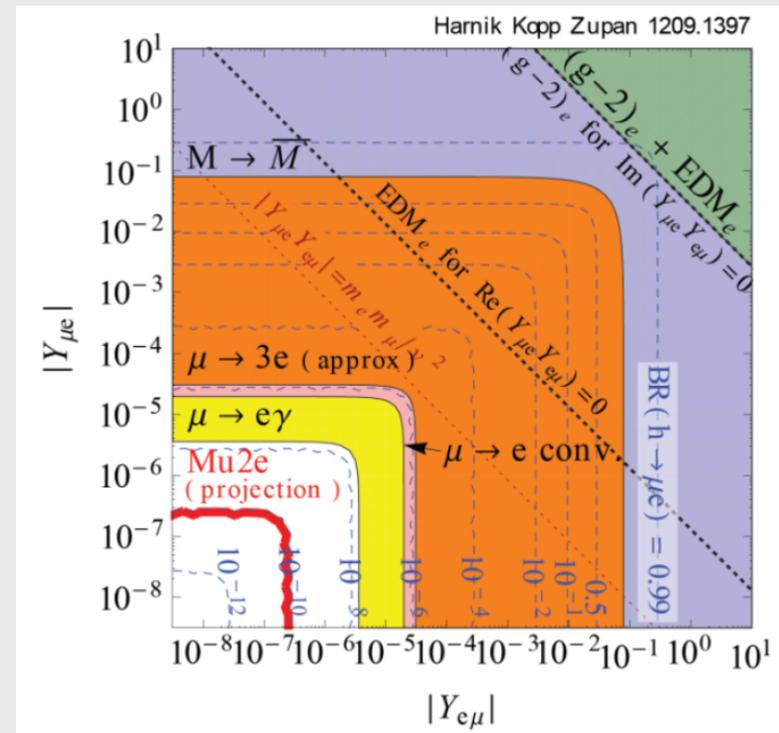
## Bad News:

$\tau$ 's hard to produce:  
 $\sim 10^{10} \tau/\text{yr}$  vs  $\sim 10^{11} \mu/\text{sec}$  in upcoming muon experiments

$\tau$ 's help pin down models and sometimes biggest BR

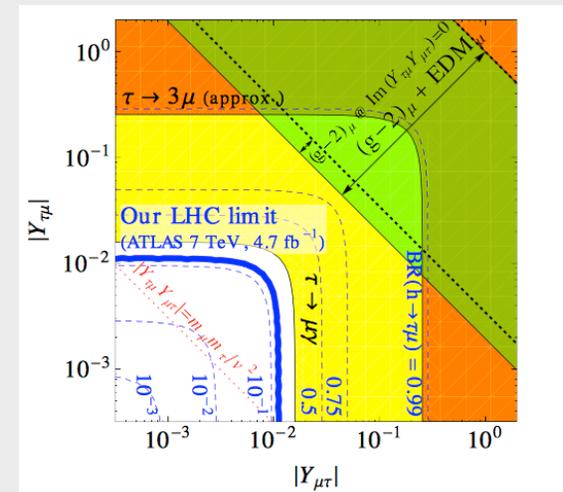
# Constraints on Higgs:

- Very strong limits on LFV Higgs decays for 1st-2nd generation



R. Bernstein, FNAL

- But not if  $\tau$  involved: 1st-3rd or 2nd-3rd

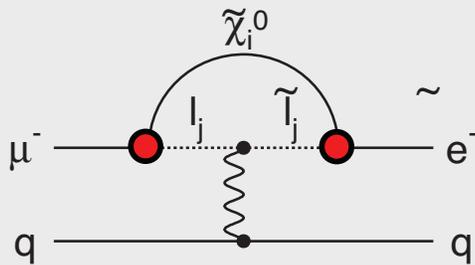


Mu2e

# Contributions to $\mu \rightarrow e$ Conversion

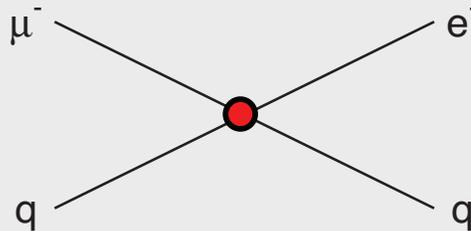
## Supersymmetry

rate  $\sim 10^{-15}$



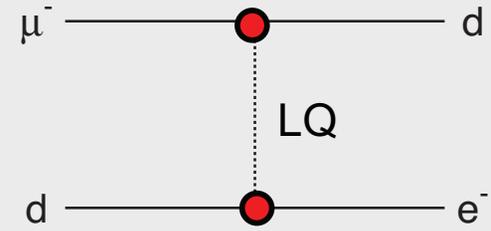
## Compositeness

$\Lambda_c \sim 3000$  TeV



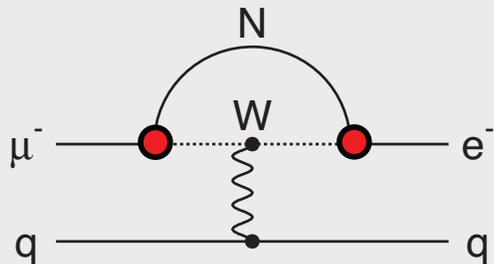
## Leptoquark

$M_{LQ} = 3000 (\lambda_{\mu d} \lambda_{e d})^{1/2}$  TeV/c<sup>2</sup>



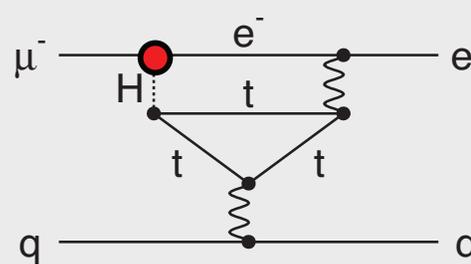
## Heavy Neutrinos

$|U_{\mu N} U_{e N}|^2 \sim 8 \times 10^{-13}$



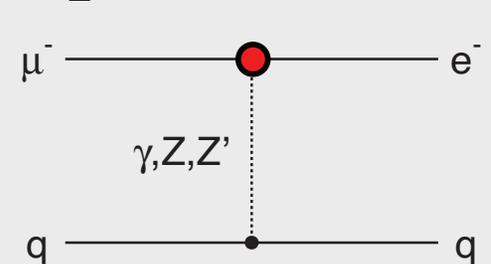
## Second Higgs Doublet

$g(H_{\mu e}) \sim 10^{-4} g(H_{\mu\mu})$



## Heavy Z' Anomal. Z Coupling

$M_{Z'} = 3000$  TeV/c<sup>2</sup>

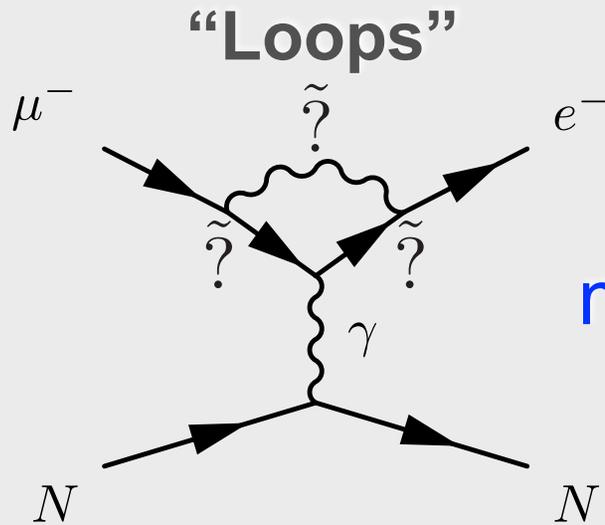


also see Flavour physics of leptons and dipole moments, [arXiv:0801.1826](https://arxiv.org/abs/0801.1826) ;

Marciano, Mori, and Roney, Ann. Rev. Nucl. Sci. 58, doi:[10.1146/annurev.nucl.58.110707.171126](https://doi.org/10.1146/annurev.nucl.58.110707.171126) ;

# De Gouvea Lagrangian

$$\mathcal{L}_{\text{CLFV}} = \frac{m_\mu}{\Lambda^2} \bar{\mu}_R \sigma_{\mu\nu} e_L F^{\mu\nu} + \frac{1}{\Lambda^2} \bar{\mu}_L \gamma^\mu e_L (\bar{u}_L \gamma_\mu u_L + \bar{d}_L \gamma_\mu d_L)$$



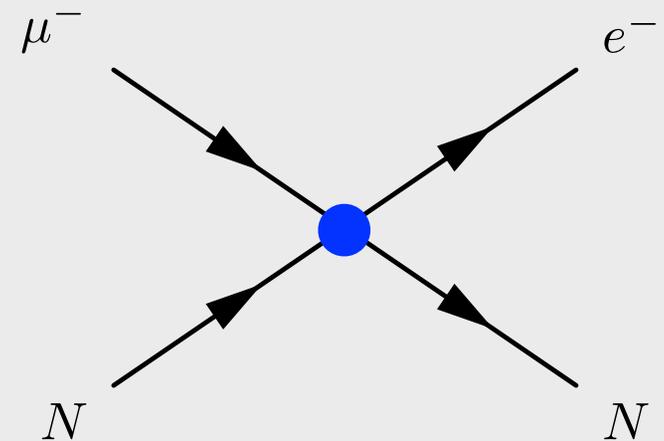
Supersymmetry and Heavy Neutrinos

**Contributes to  $\mu \rightarrow e\gamma$**

(just imagine the photon is real)

Mu2e/COMET/MEG/Mu3e  
at  $10^3$  TeV

**“Contact Terms”**



New Particles at High Mass Scale  
(leptoquarks, heavy Z,...)

**Does not produce  $\mu \rightarrow e\gamma$**

Mu2e/COMET at  $10^4$  TeV

# EFT: Beyond $\Lambda$ and $\kappa$

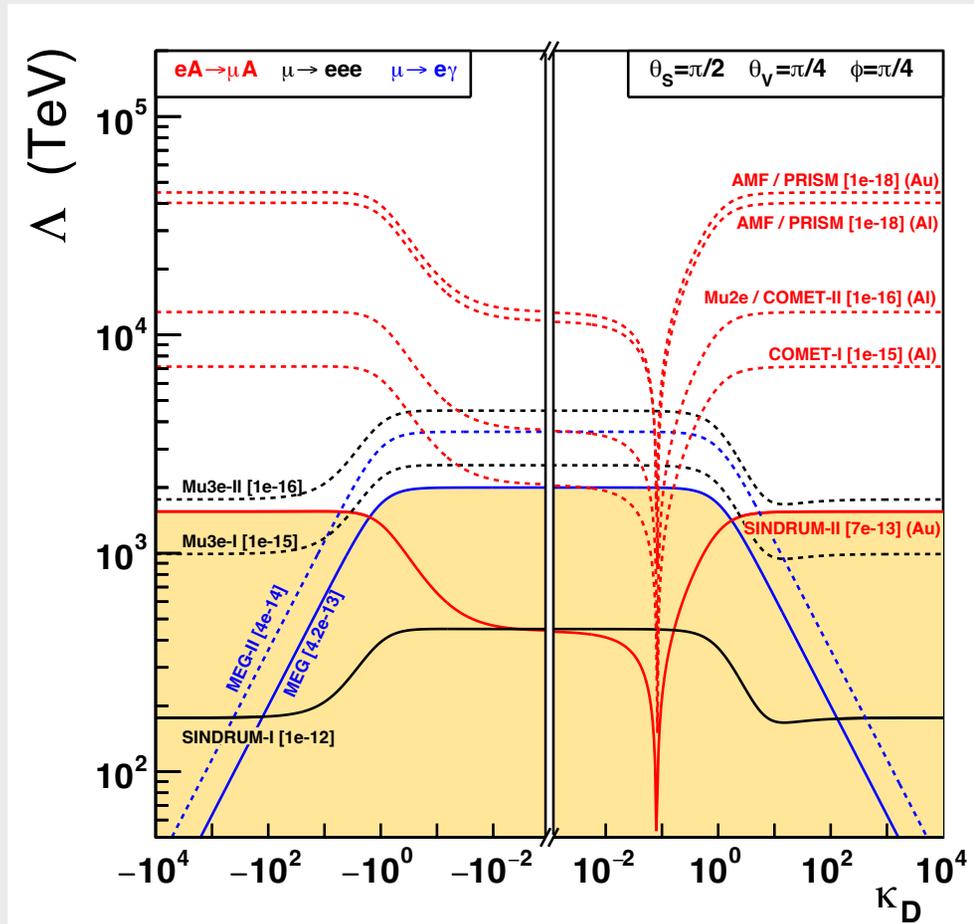
S. Davidson and B. Echenard, [2010.00317](#) [hep-ph]

- Write EFT Lagrangian:
  - Dipole ( $\mu \rightarrow e\gamma$ ) +  
Contact Scalar ( $\mu \rightarrow 3e$ )<sub>L</sub> +  
Contact Vector ( $\mu \rightarrow 3e$ )<sub>R</sub> +  
Contact  $\mu N \rightarrow eN$  (light nuclei) +  
Contact  $\mu N \rightarrow eN$  (heavy nuclei)
- Parameterize coefficient space with spherical coordinates: *lets you express constraints on all three processes simultaneously*
- Will show you “slices” in the multi-dimensional space

# Complementarity

S. Davidson and B. Echenard, [2010.00317](#) [hep-ph]

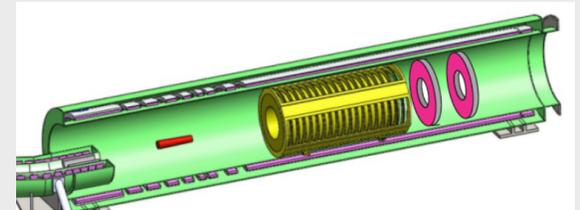
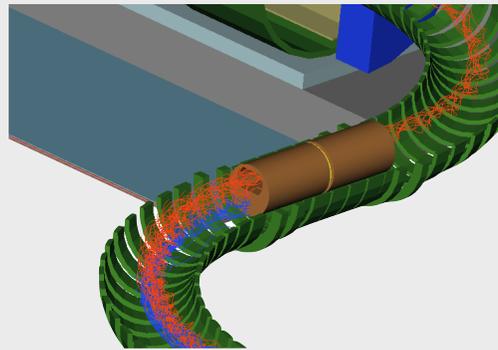
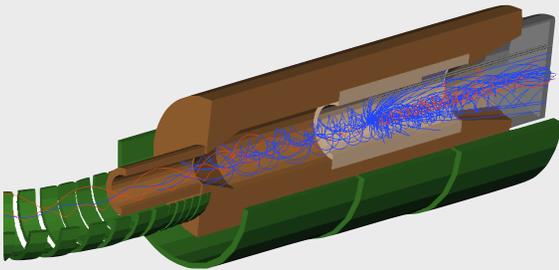
- All three channels have strengths; we need the combination



- $\mu \rightarrow e\gamma$  and  $\mu \rightarrow 3e$  at  $\mathcal{O}(10^{-15})$  are a next-gen target

# Next Part

- *An overview of experiment and walk you through the solenoids*

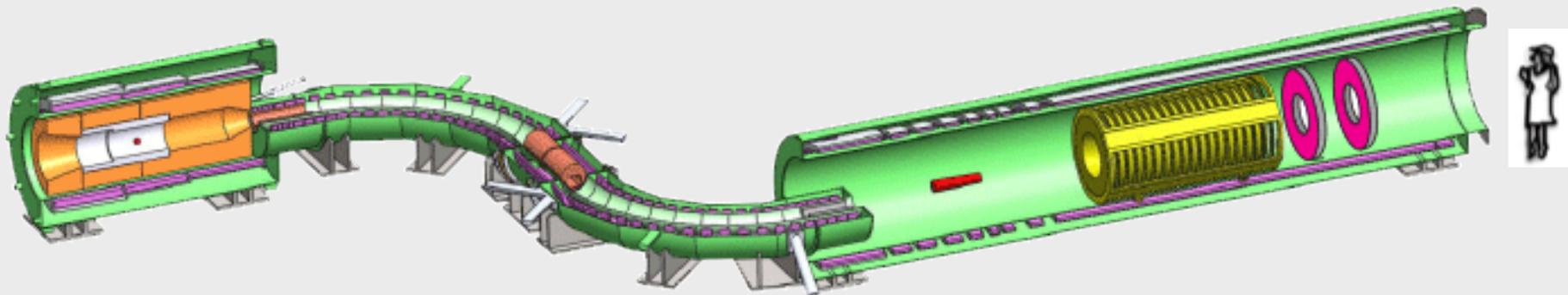
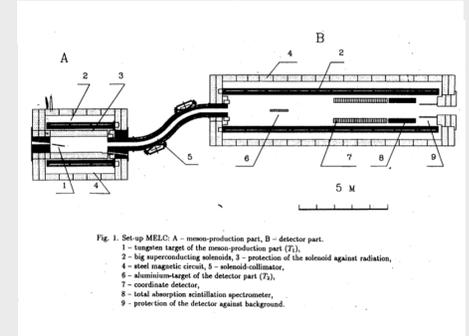


- Describe what happens when muons reach the stopping target
- Explain detector, signal, and backgrounds

# Mu2e Overview

- Production: protons hit a target, making  $\pi$ 's, which decay into accepted  $\mu$ 's

V. Lobashev, MELC 1992:



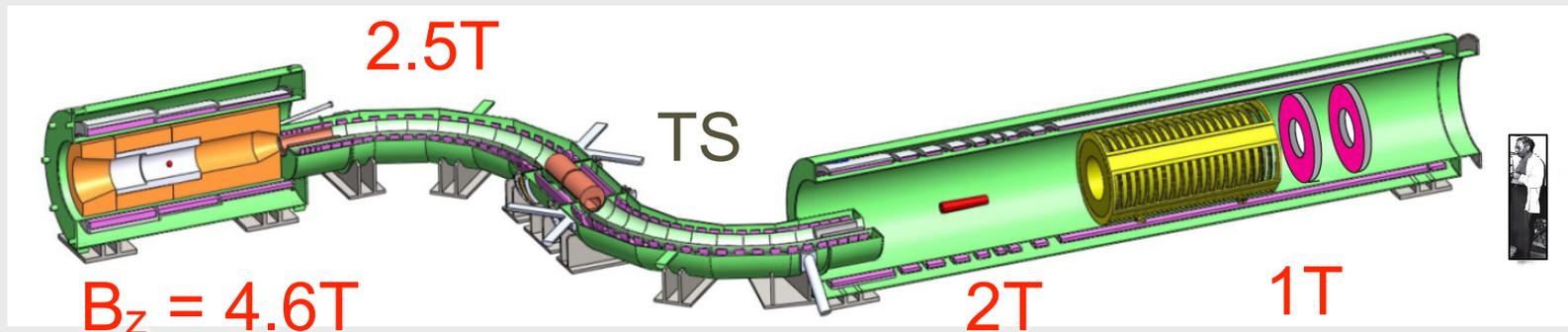
- *Transport*: S-curve eliminates backgrounds and sign-selects  $\mu^-$  vs.  $\mu^+$

entire system in vacuum  $< 10^{-4}$  torr

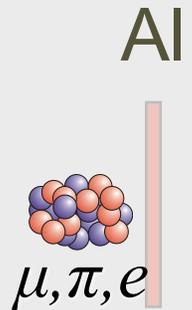
- *Detector*: Stopping Target, Tracking and Calorimeter

# Mu2e Muon Beam: Three Solenoids and Gradient

PS 4.6T  $\longrightarrow$  B-field gradient  $\longrightarrow$  1T DS



- Target protons at 8 GeV inside superconducting solenoid
- Capture muons and guide through S-shaped region to Al stopping target
- Gradient fields used to collect and transport muons

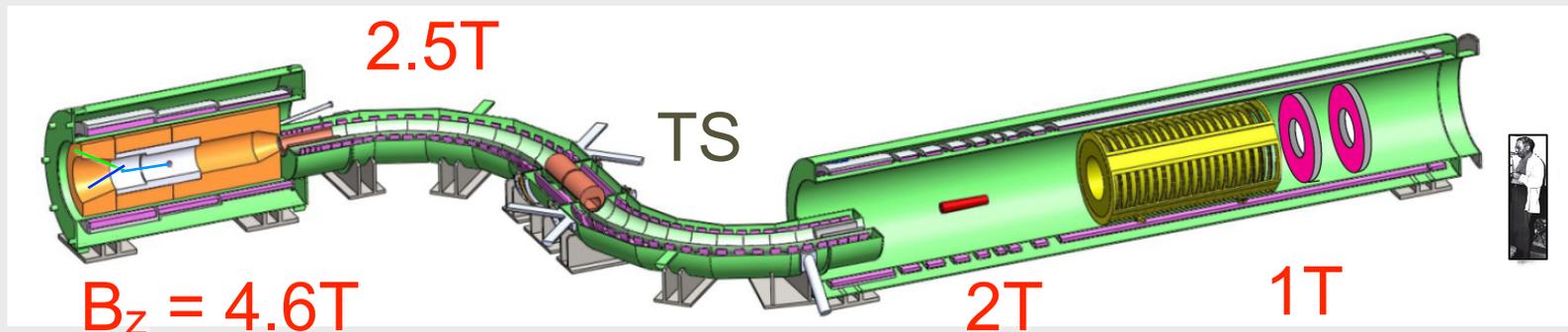


Muon K.E  $\sim$  7 MeV  
muons range out by  $dE/dx$   
in Aluminum

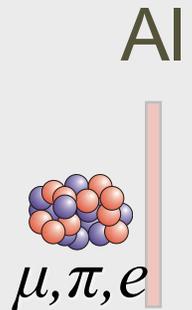
Mu2e

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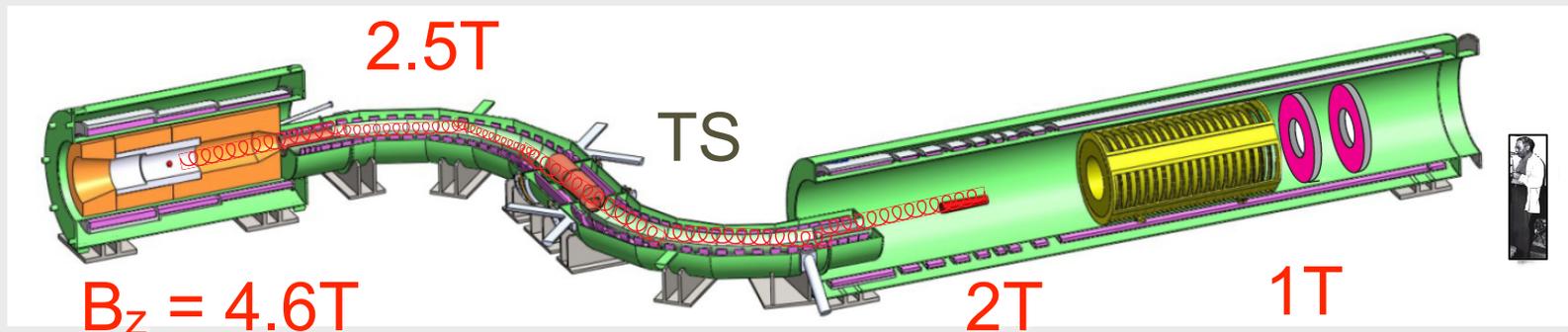


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Mu2e

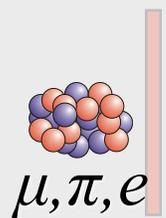
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Al

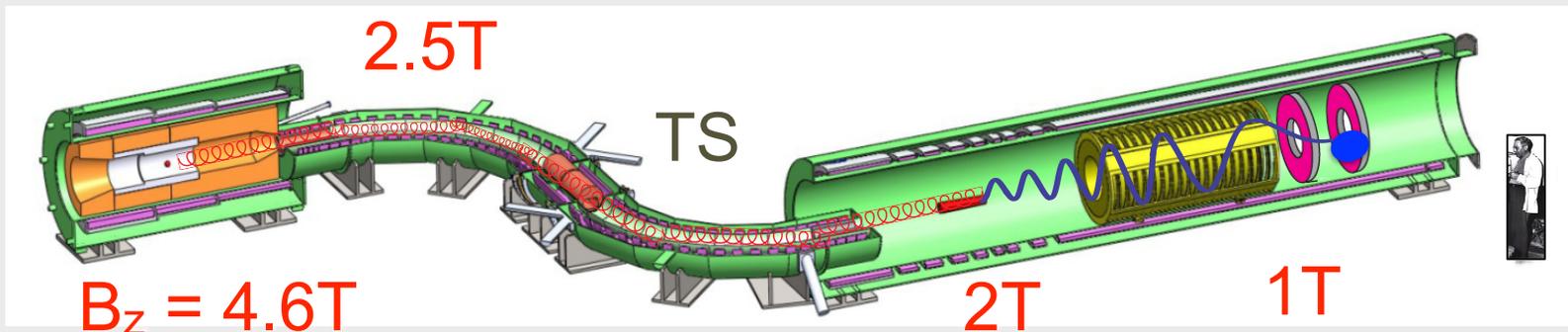


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Mu2e

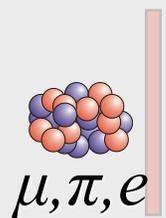
# Mu2e Muon Beam: Three Solenoids and Gradient

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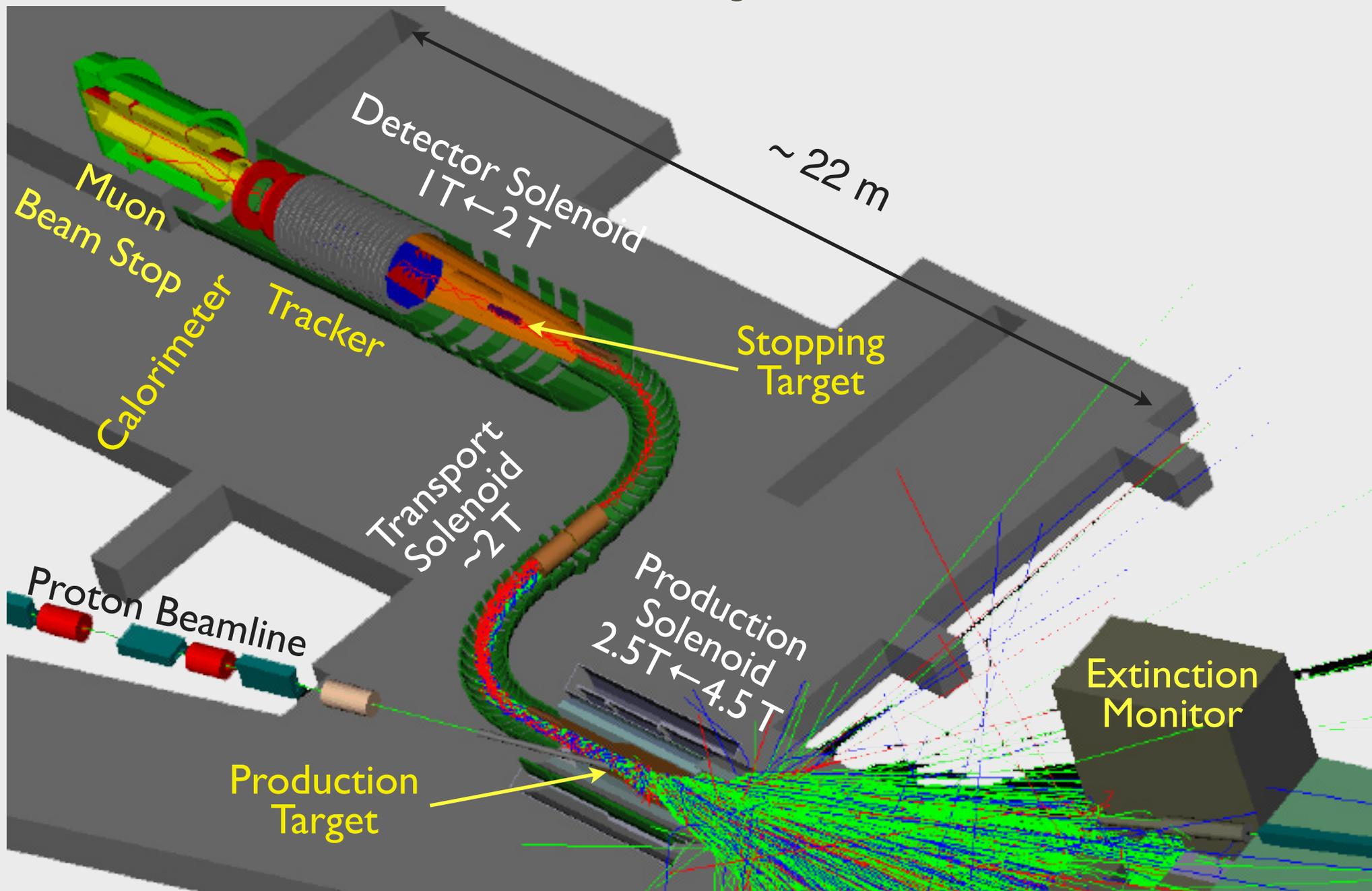
Al



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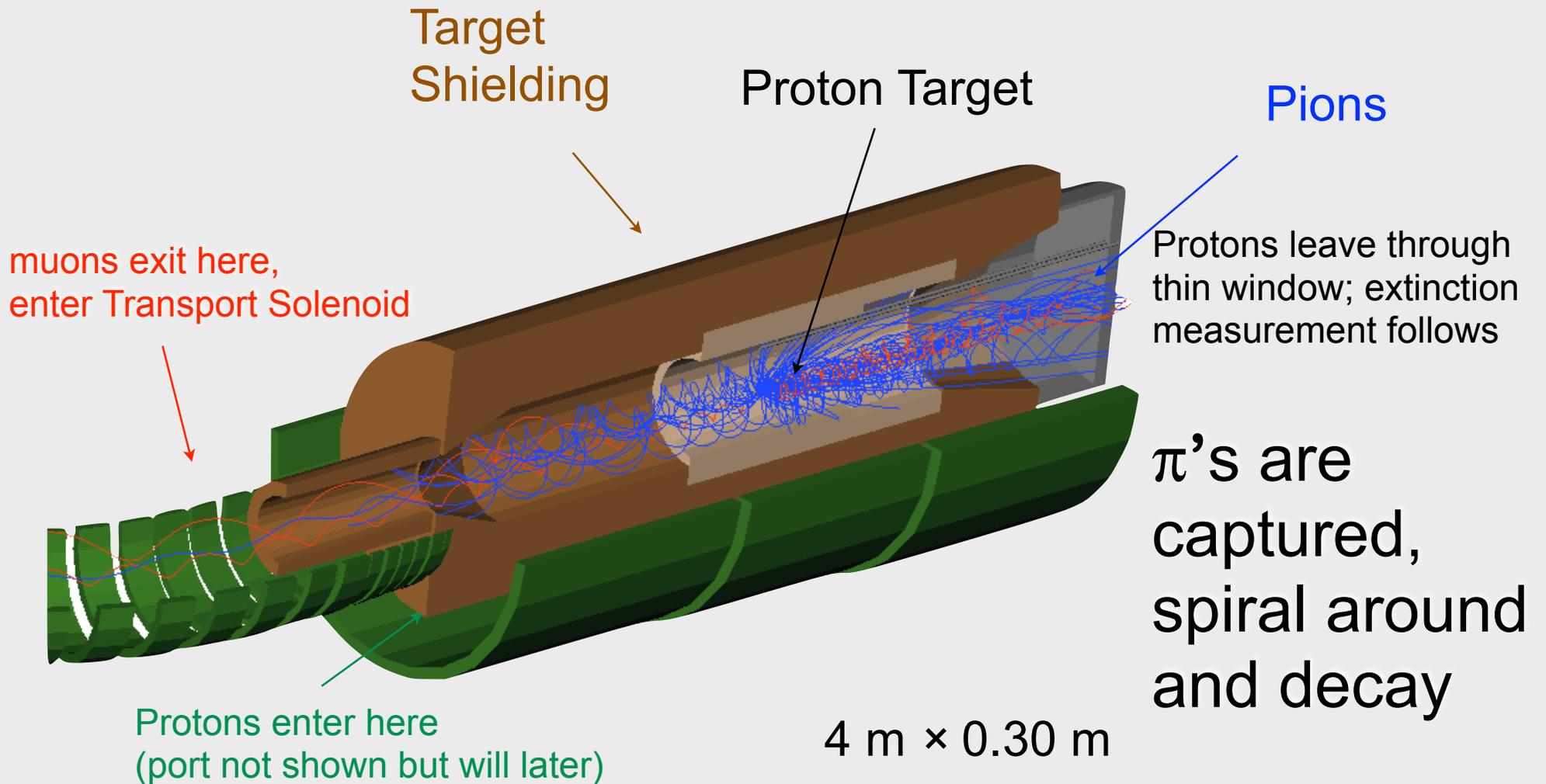
Mu2e

# Beam's Eye View



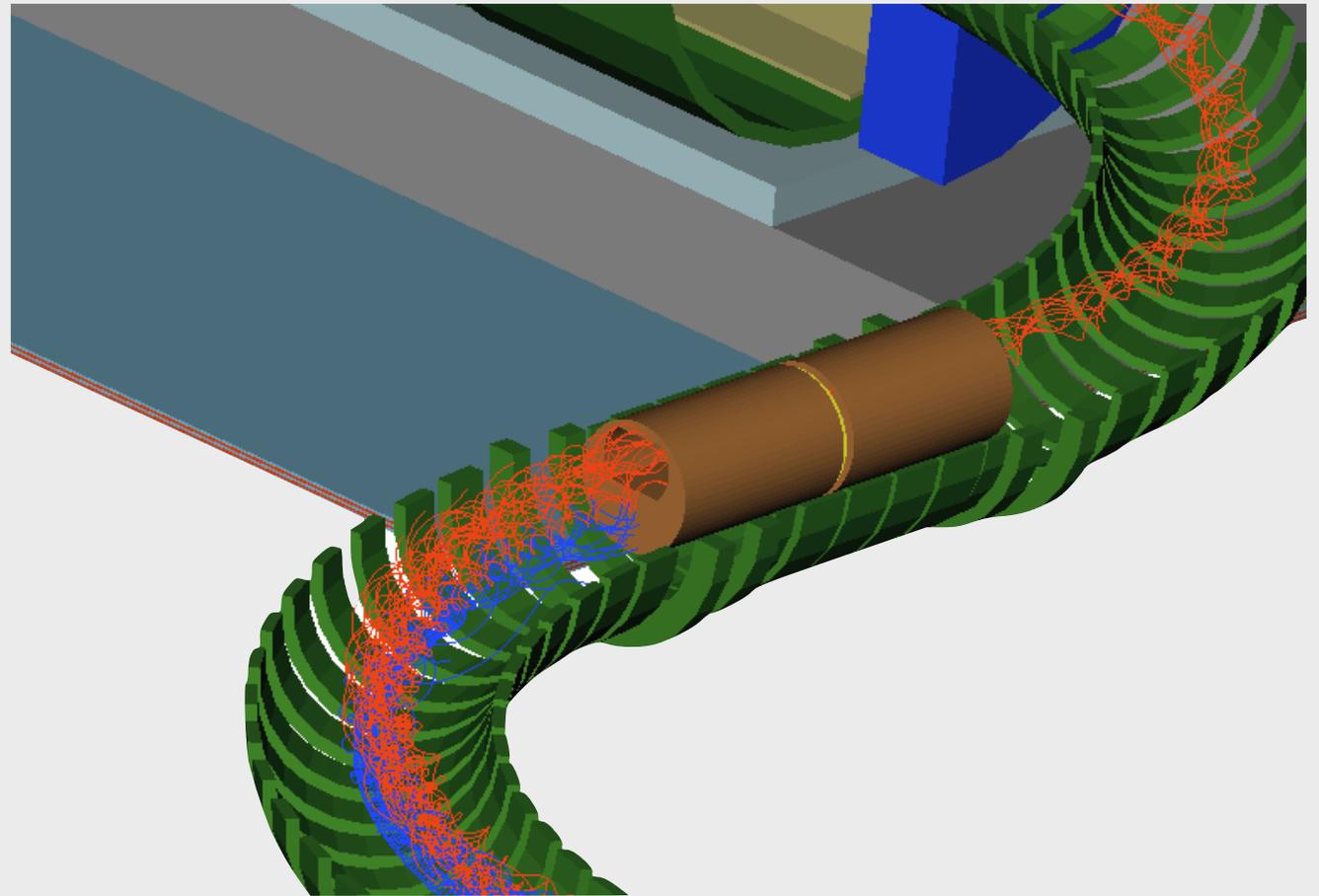
# Production Solenoid:

Protons enter opposite to outgoing muons



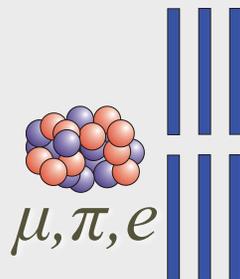
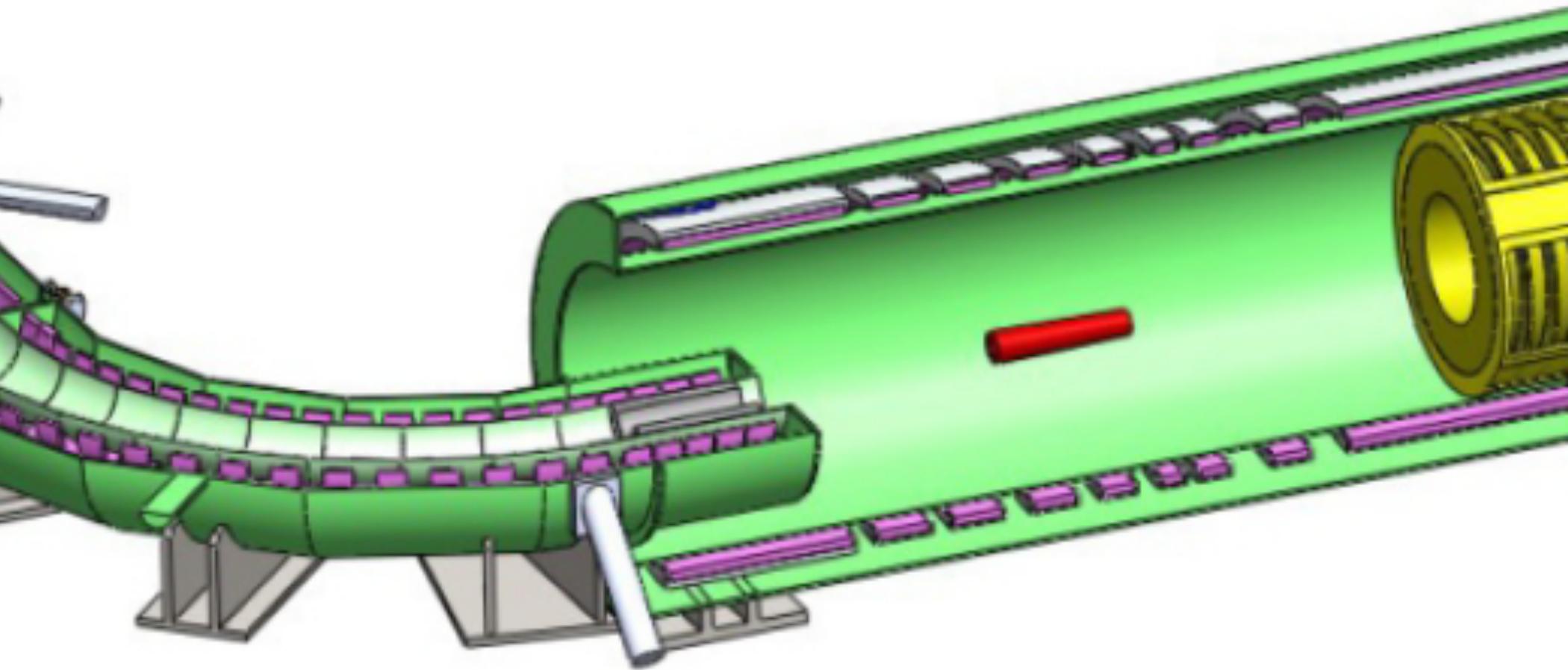
# Transport Solenoid

- Curved solenoid eliminates line-of-sight transport of photons and neutrons
- Curvature drift and collimators sign and momentum select beam



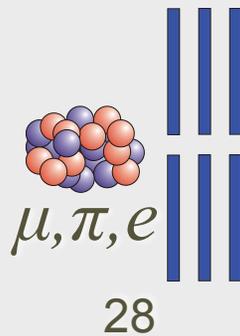
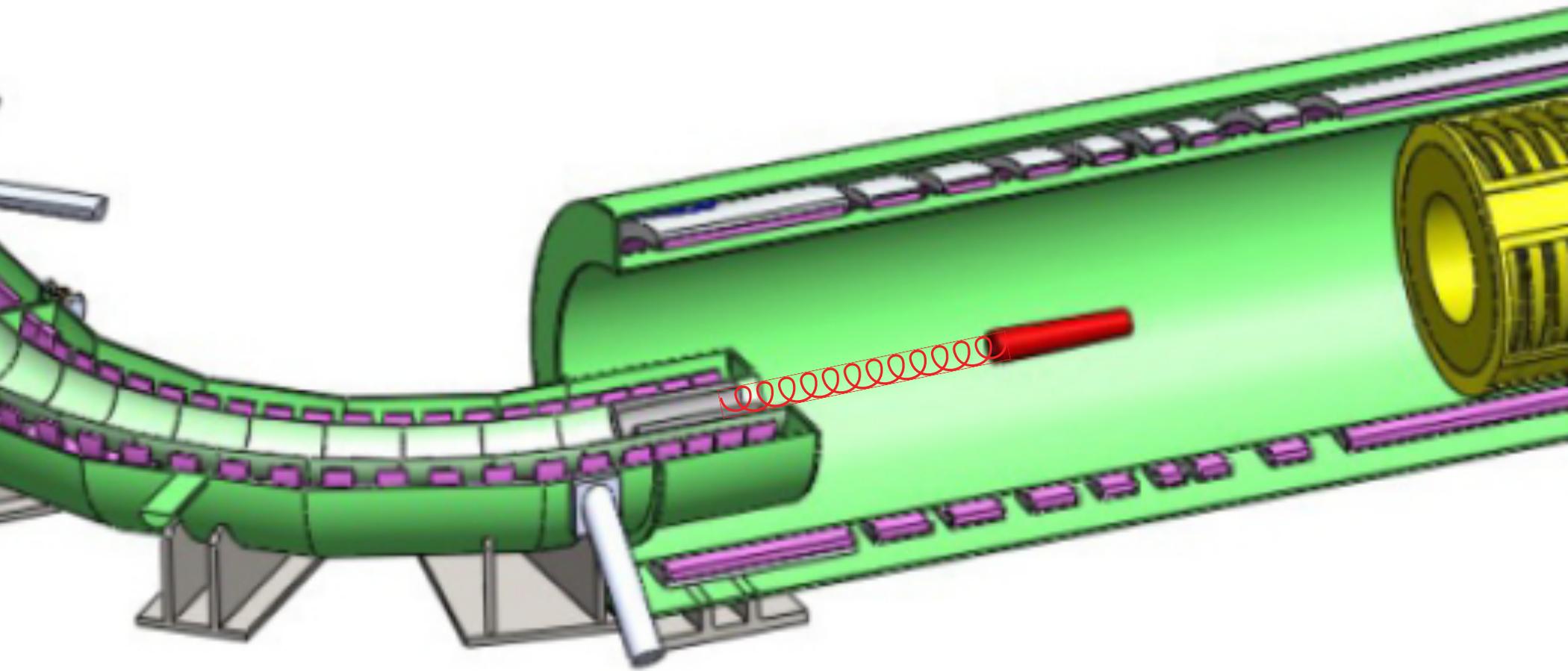
13.1 m along axis  $\times$   $\sim 0.25$  m

# Now Enter Detector Solenoid



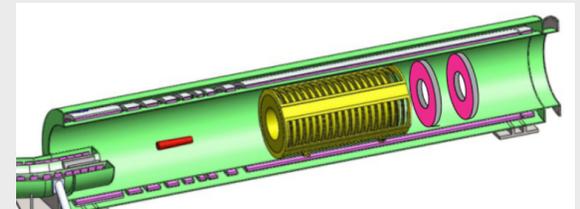
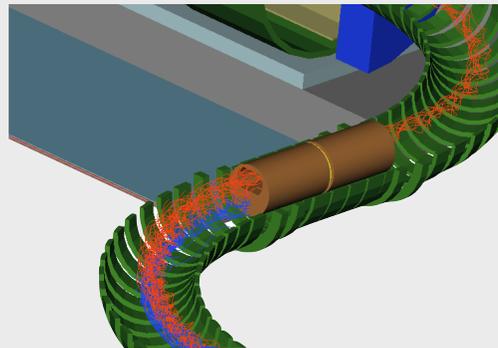
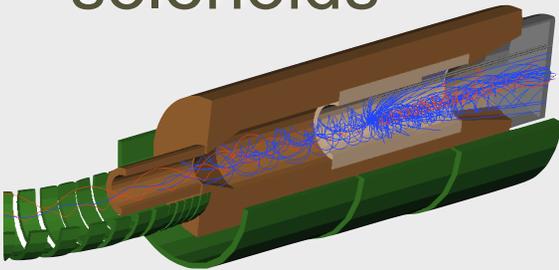
28

# Now Enter Detector Solenoid



# Next Part

- An overview of experiment and walk you through the solenoids



- *Describe what happens when muons reach the stopping target*
- Explain detector, signal, and backgrounds

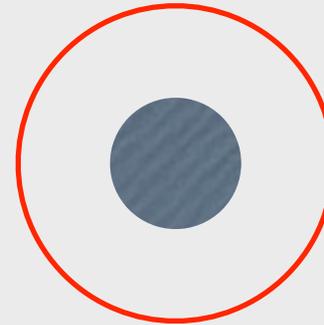
# Overview Of Processes

$\mu^-$  stops in thin Al foil



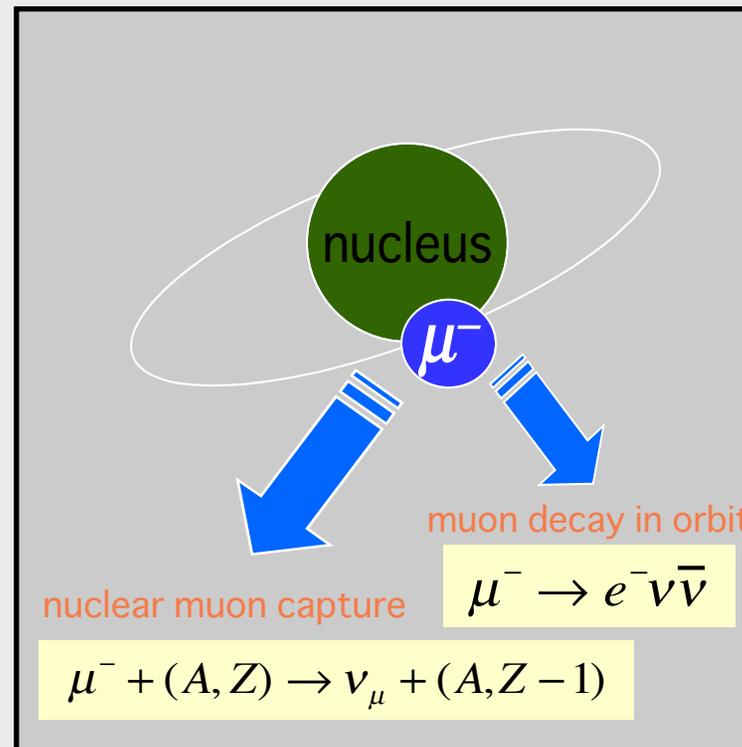
the Bohr radius is  $\sim 20$  fm,  
so the  $\mu^-$  sees the nucleus

$\mu^-$  in 1s state



Al Nucleus  
 $\sim 4$  fm

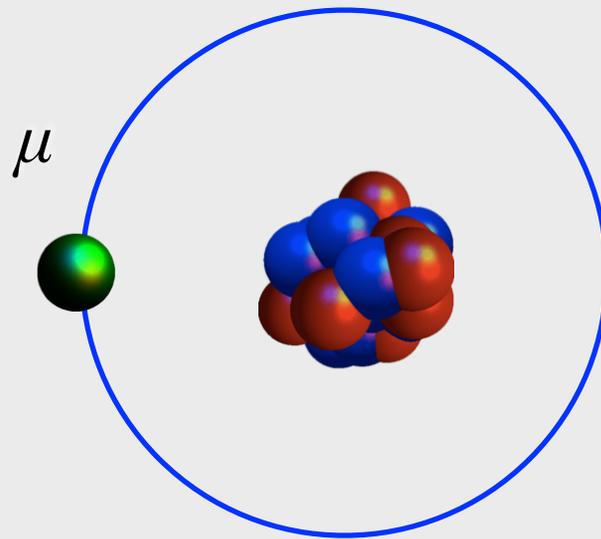
muon capture,  
muon “falls into”  
nucleus:  
**normalization**



60% capture  
40% decay

Decay in Orbit:  
**background**

# Decay-in-Orbit (DIO) Background



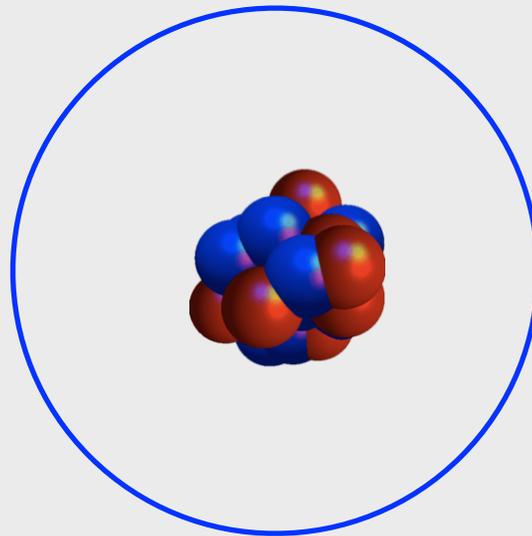
# Decay-in-Orbit (DIO) Background

this electron can be background;  
let's see how



$e$

$\nu_{\mu}$



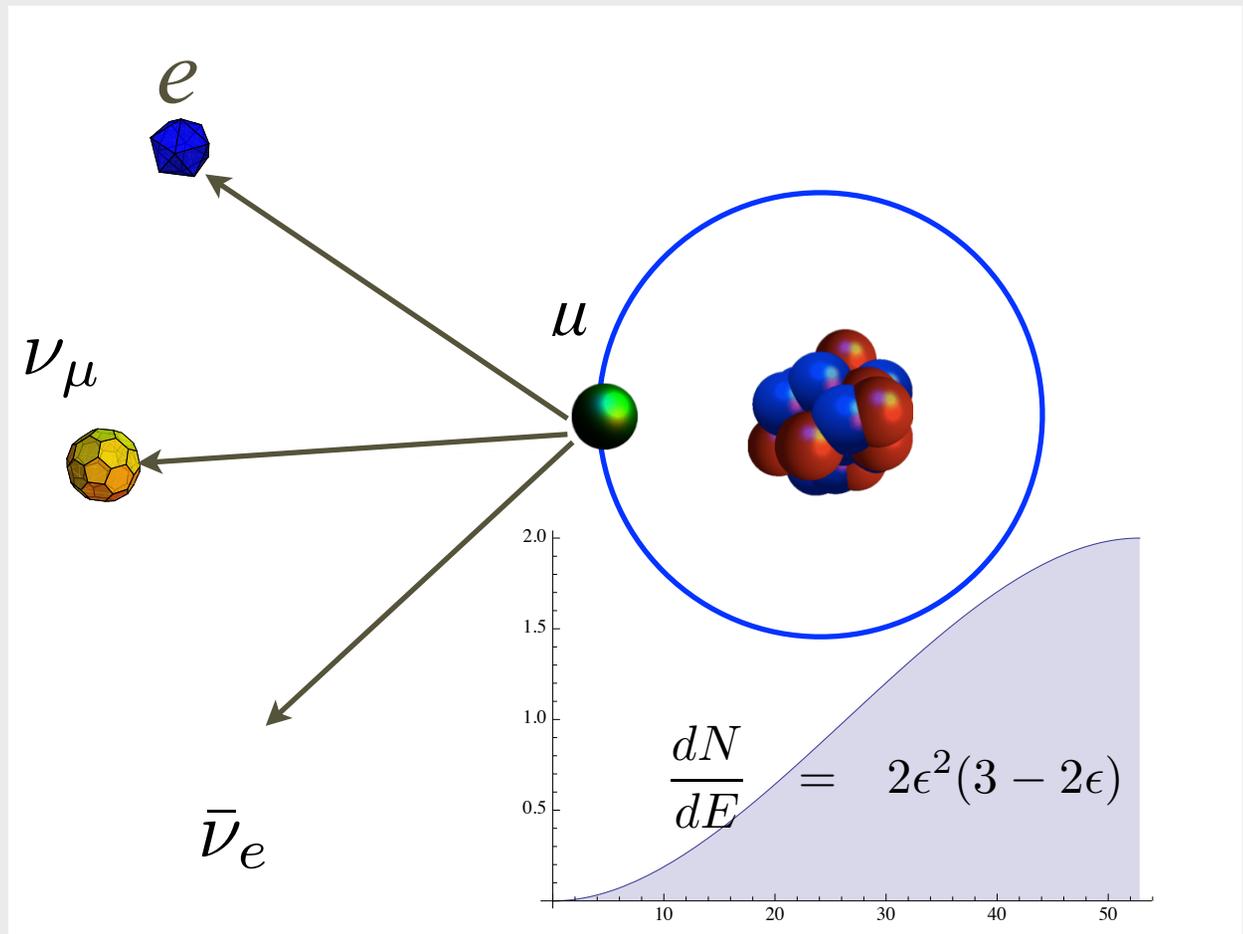
$\bar{\nu}_e$

# DIO: usually not background

- Peak and Endpoint of Michel Spectrum is at

$$E_{\max} = \frac{m_{\mu}^2 + m_e^2}{2m_{\mu}} \approx 52.8 \text{ MeV}$$

- Detector will be insensitive to electrons at this energy
- Recall *signal* at  $105 \text{ MeV} \gg 52.8 \text{ MeV}$

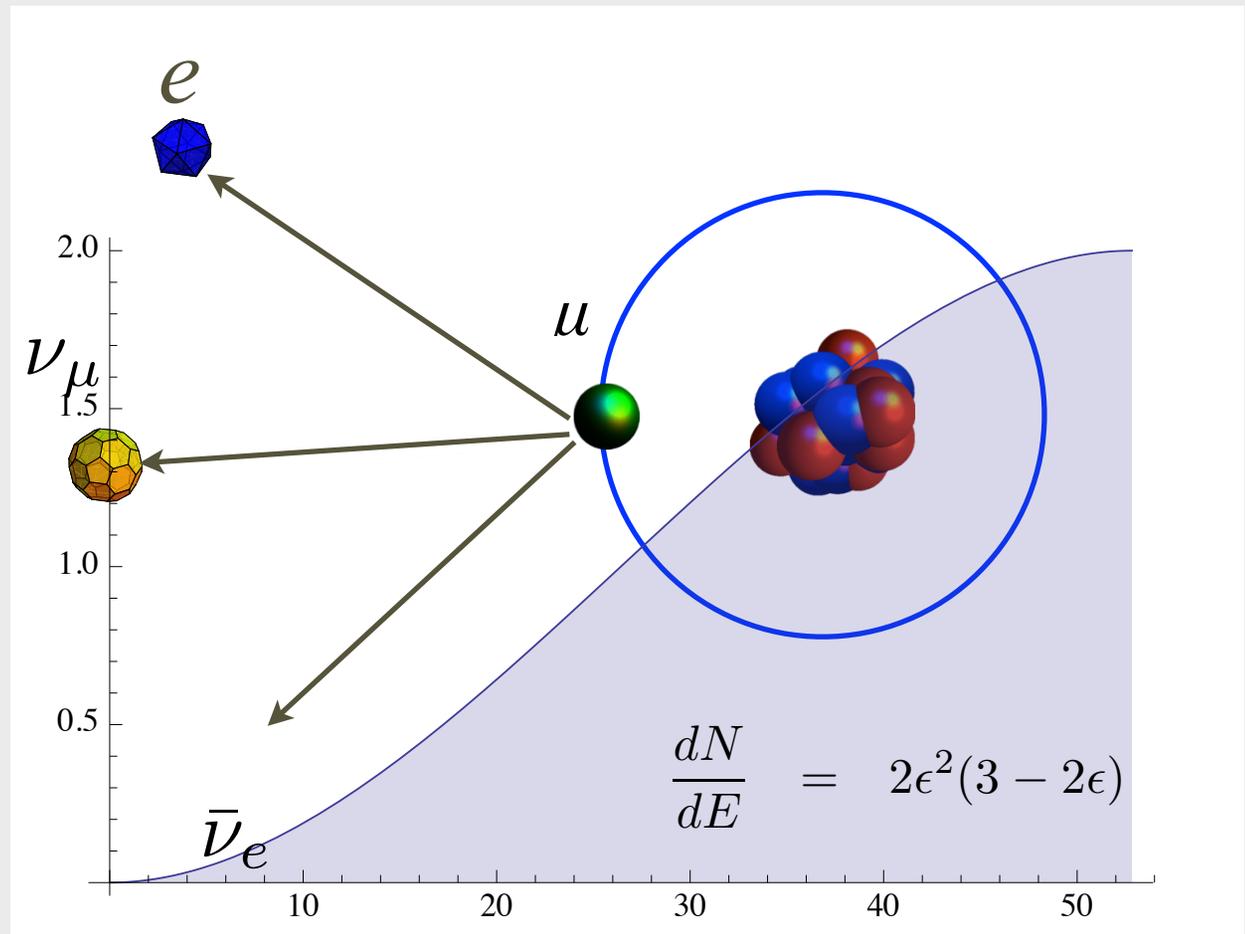


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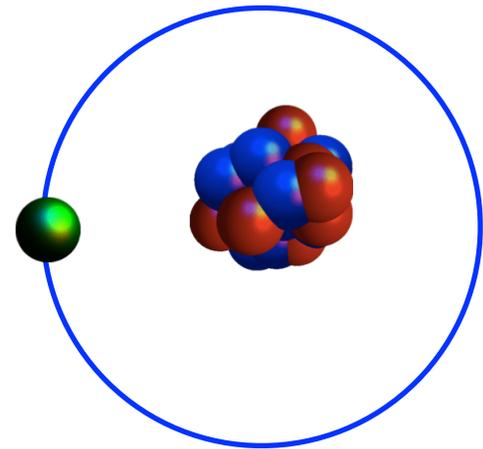
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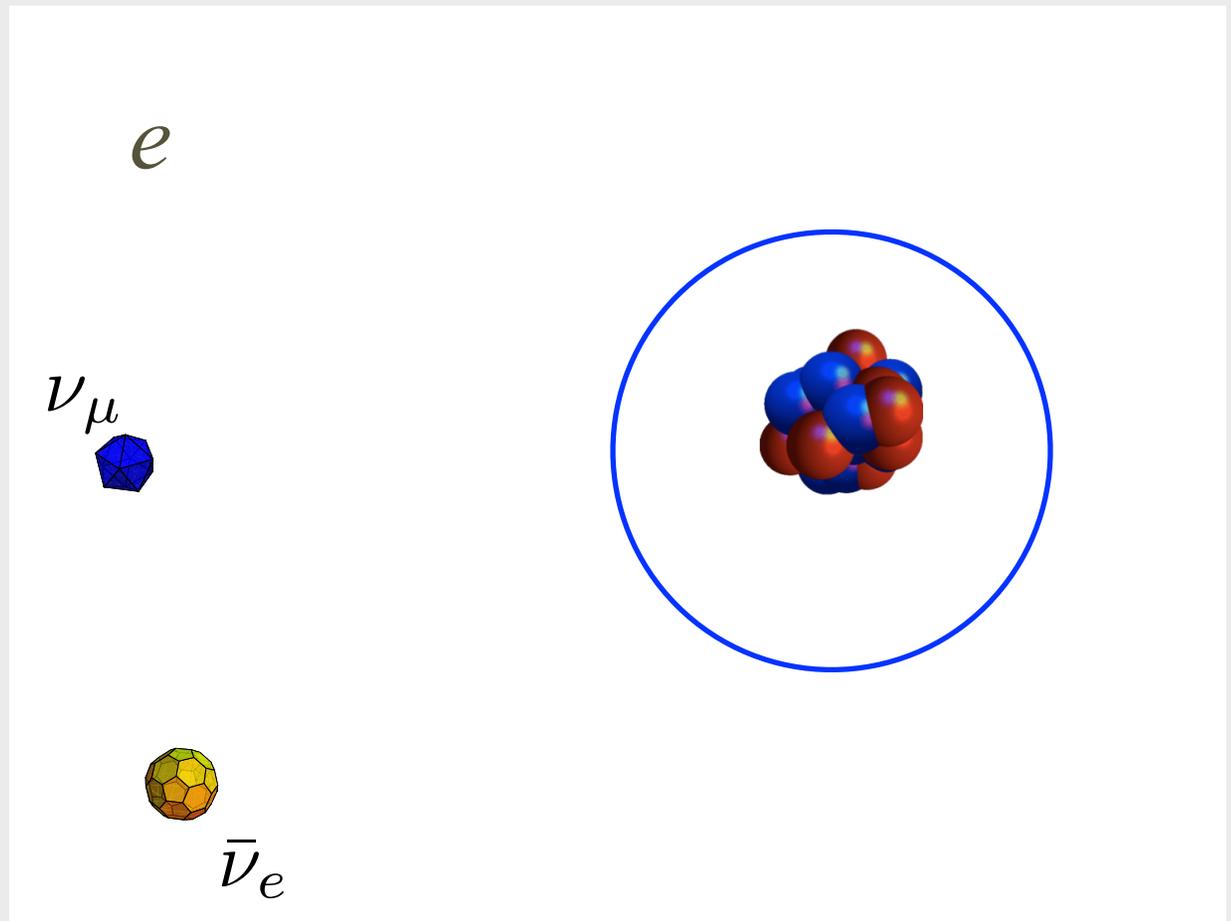
# Decay-In-Orbit Background

- Same process as before
- But this time, include electron recoil off nucleus
- If neutrinos are at rest, **the DIO electron can be exactly at conversion energy** (up to neutrino mass)



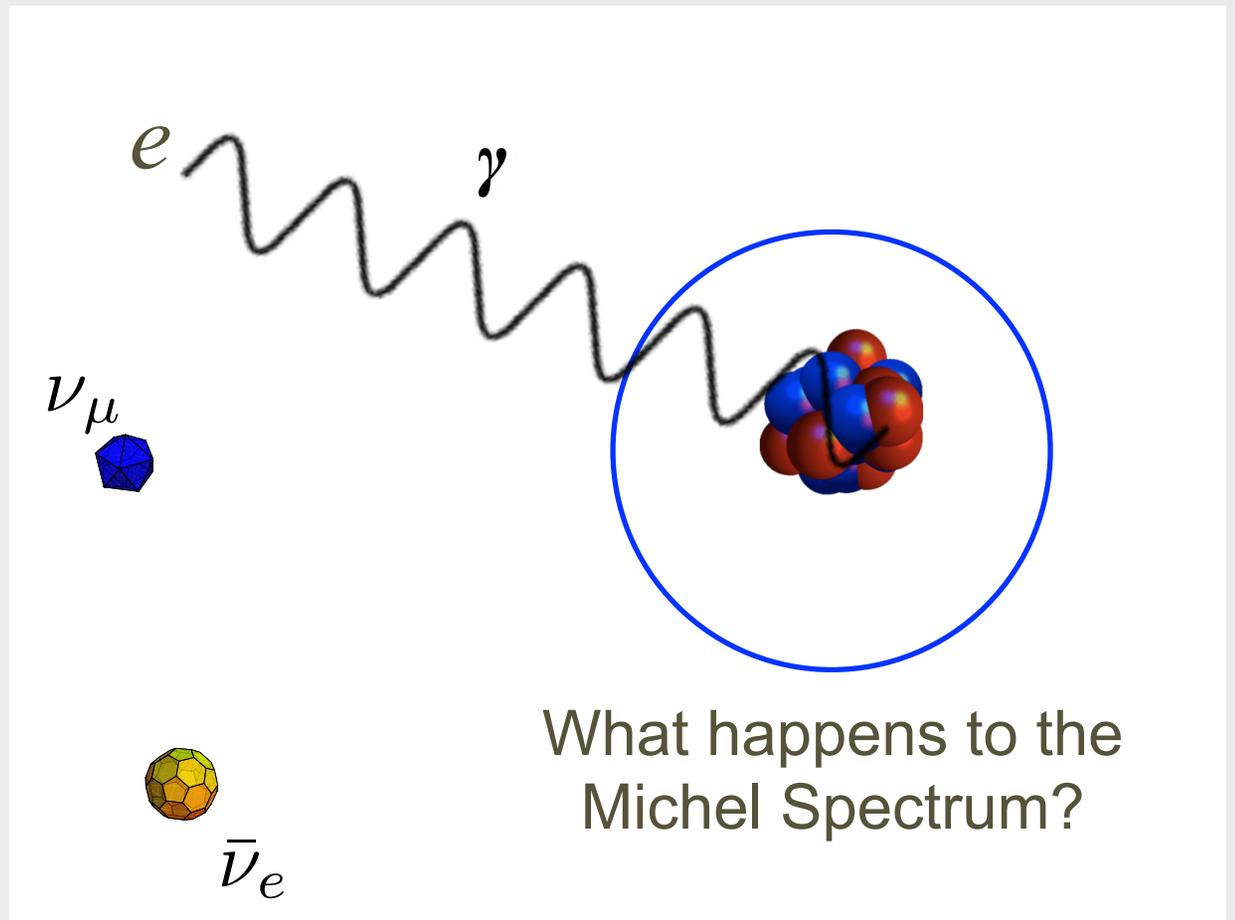
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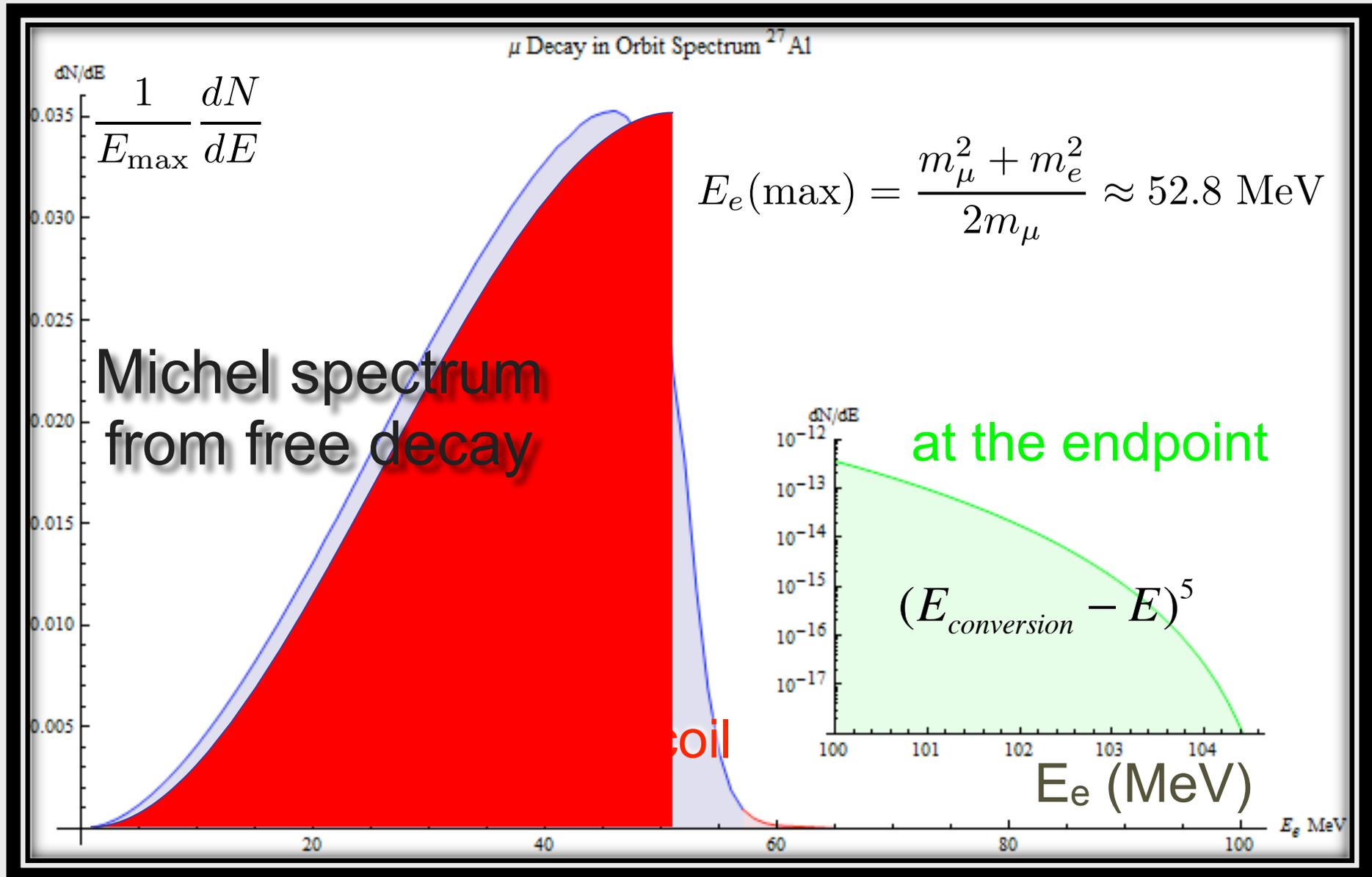
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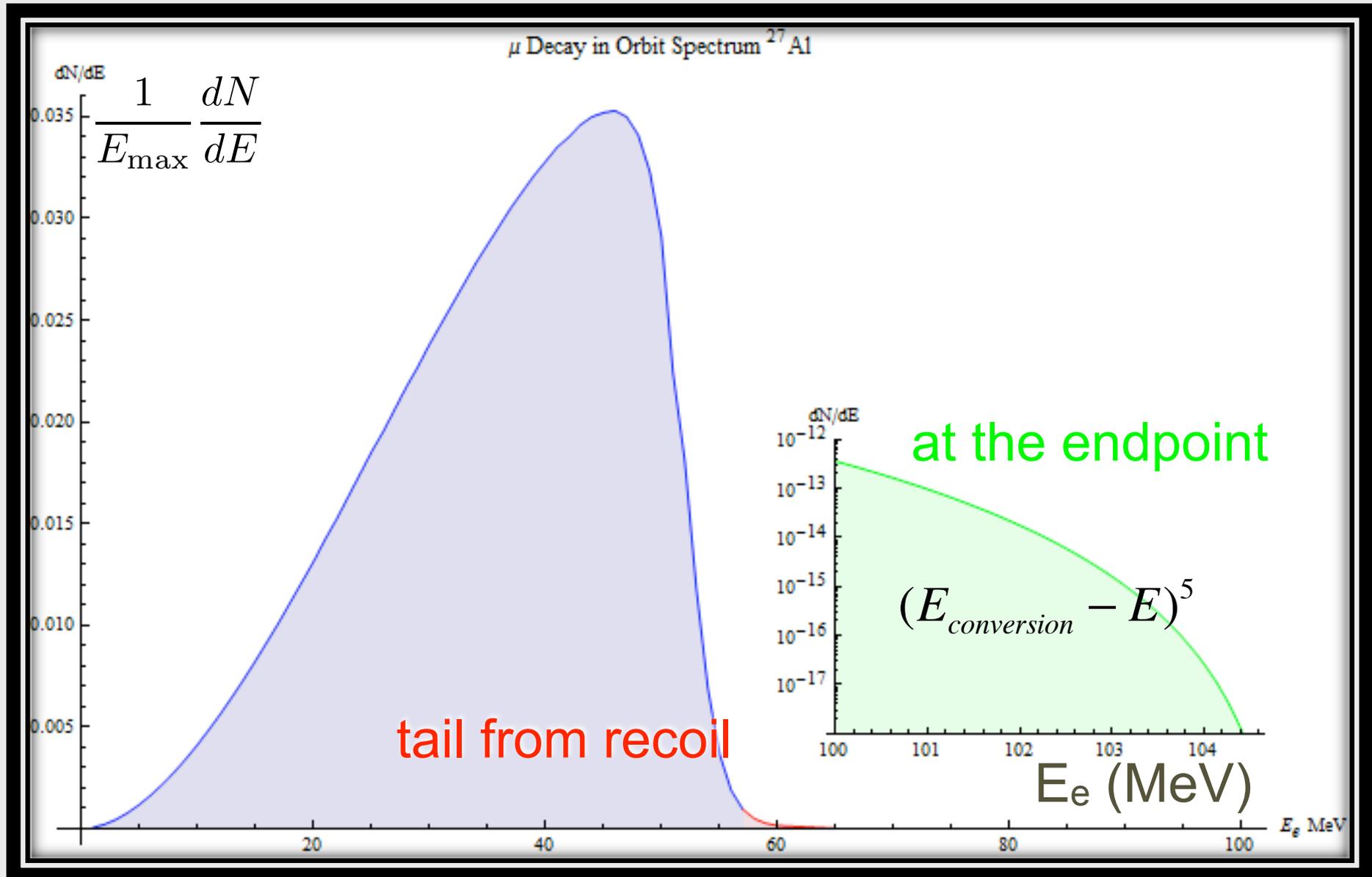
# Decay-in-Orbit Shape

Szafron and Czarnecki , [1608.05447\[hep-ph\]](#) [10.1103/PhysRevD.94.051301](#)



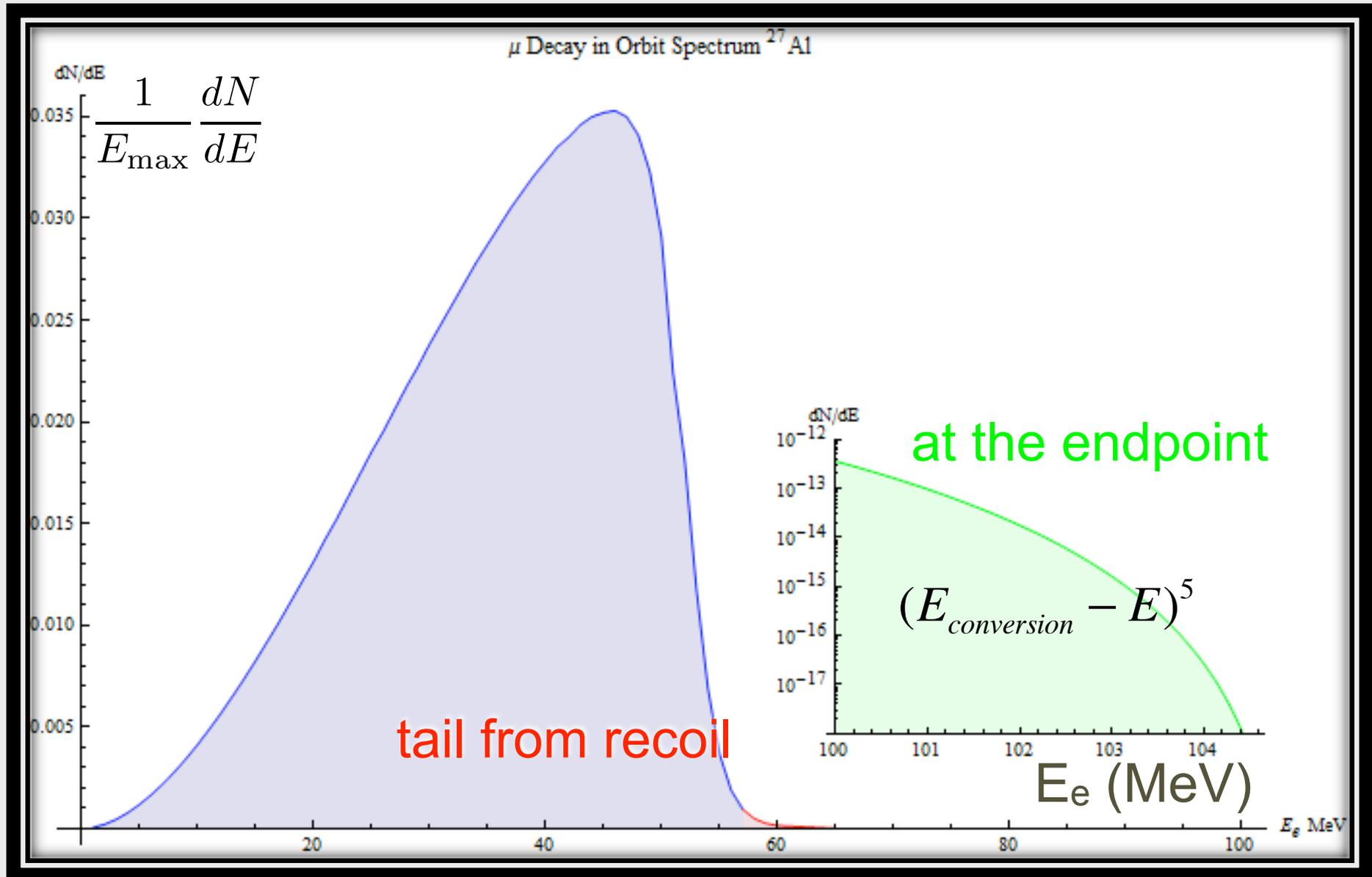
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# Decay-in-Orbit Shape

Szafron and Czarnecki , [1608.05447\[hep-ph\]](#) [10.1103/PhysRevD.94.051301](#)



# How Do We Suppress DIO?

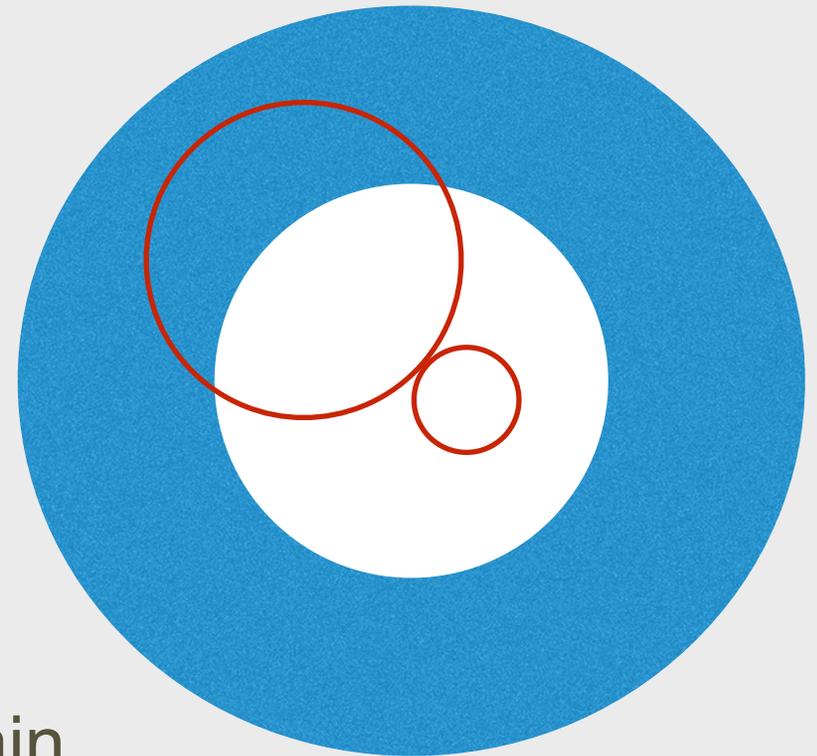
- Best possible energy resolution: we do not want DIO events near the endpoint resolution smeared upwards and promoted into the signal region, so we are sensitive to both the “gaussian width” and especially to “high side tails”
- We use a solenoidal field and annular detectors
  - $p=qBR$ ;  $p$  for Michel edge at 52.8 MeV is about 1/2 of conversion energy of 105 MeV.

# Annular Detectors

Looking along detector axis

- This design gives us a few  $10^5$  muons to reconstruct instead of  $\sim 10^{18}$  from the distorted DIO spectrum
- Making it possible to have a small DIO background

105 MeV signal



$p=qBR$  again

52.8 MeV/c  
background

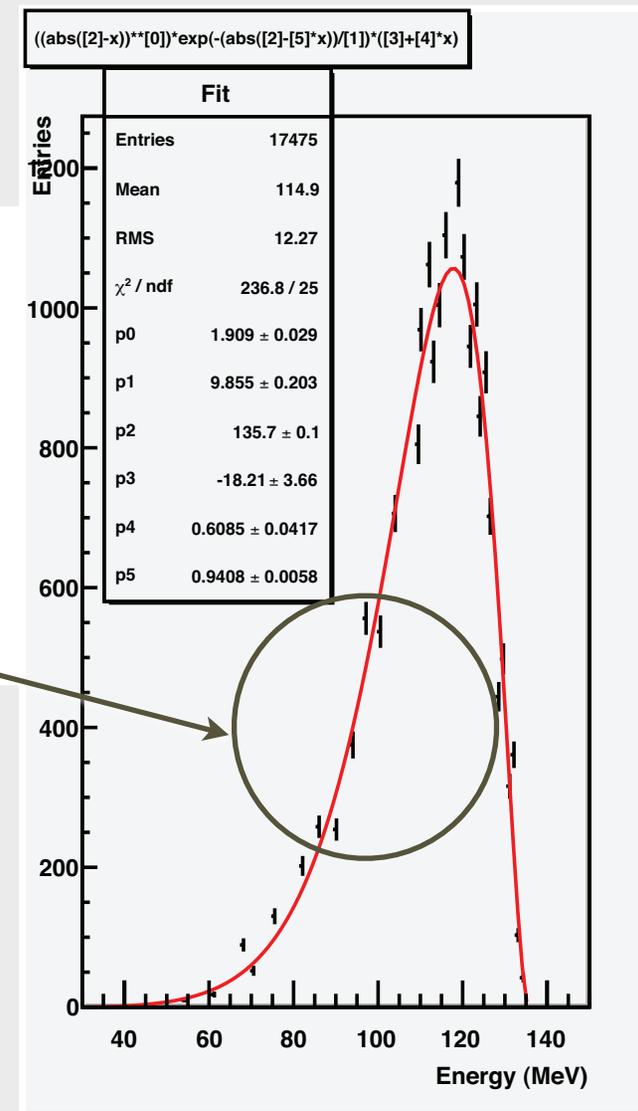
# Prompt Backgrounds

*Particles produced by proton pulse which interact almost immediately when they enter the detector:  $\pi$ , neutrons, pbars*

- **Radiative pion capture,  $\pi^- + A(N,Z) \rightarrow \gamma + X$ .**
  - $\gamma$  up to  $m_\pi$ , peak at 110 MeV;  $\gamma \rightarrow e^+e^-$ ; if one electron  $\sim 100$  MeV in the target, looks like signal: **limitation in best existing experiment, SINDRUM II?**

energy spectrum of  $\gamma$  measured on Mg  
 J.A. Bistirlich, K.M. Crowe et al., Phys Rev C5, 1867 (1972)

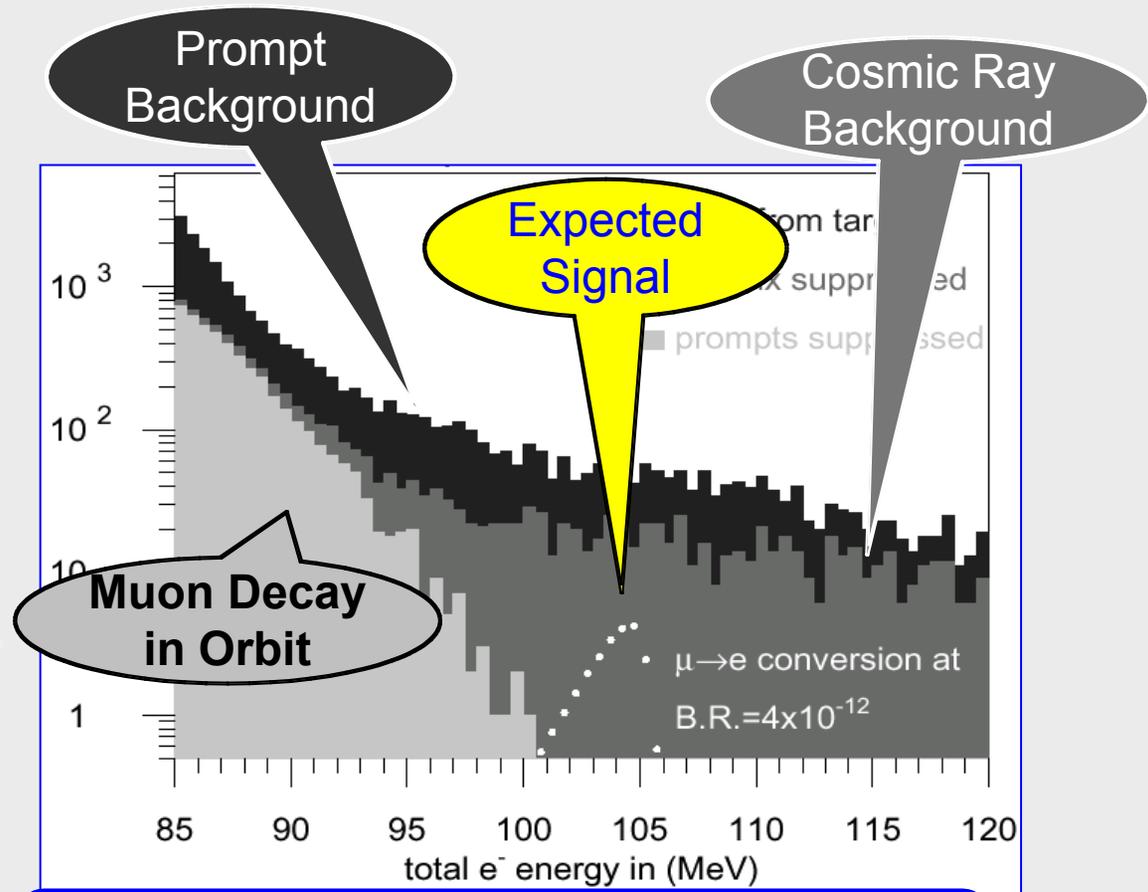
also included internal conversion,  $\pi^- N \rightarrow e^+ e^- X$



# Previous Best Experiment

## *SINDRUM-II*

- $R_{\mu e} < 6.1 \times 10^{-13}$  in Au
- Want to probe to  $6 \times 10^{-17}$
- $\approx 10^4$  improvement



**Experimental signature is 105 MeV  $e^-$  originating in a thin Ti stopping target**

# SINDRUM-II Results

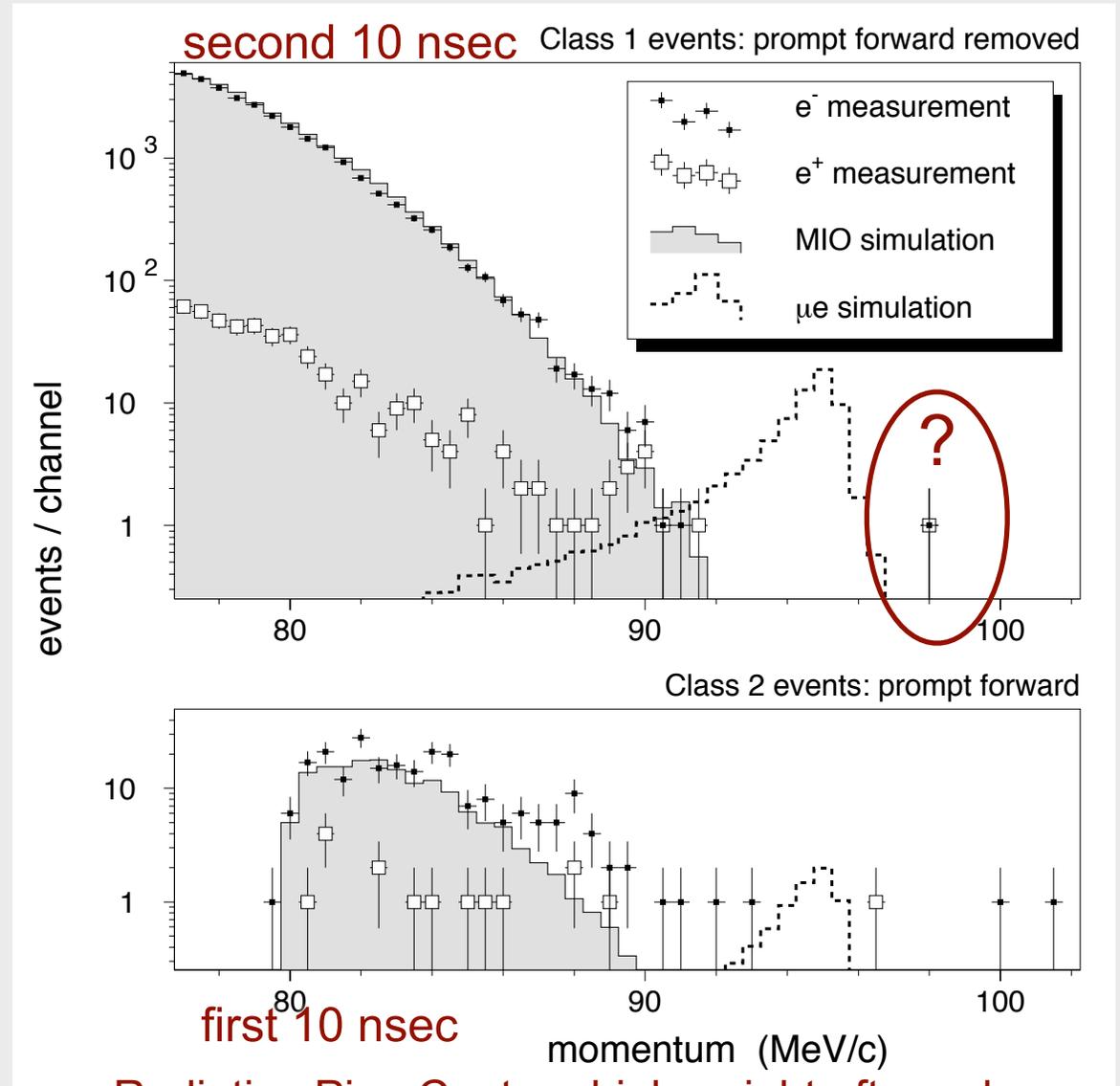
W. Bertl et al., Eur. Phys. J. C 47, 337–346 (2006)

- Final Results on Au:

$$B_{\mu e}^{\text{Au}} < 7 \times 10^{-13} \text{ @ 90\% CL}$$

**51 MHz (20 nsec)**  
 repetition rate,  
 width of pulse  
 ~0.3 nsec

not enough time separation  
 between  
 signal and prompt  
 background:  
 can't scale this method up  
 10<sup>4</sup>



Radiative Pion Capture higher right after pulse

# How Can We Do Better?

>10<sup>3</sup> increase in muon intensity from SINDRUM

## *Requiring*

Pulsed Beam to Eliminate prompt backgrounds like radiative  $\pi$  capture and CR

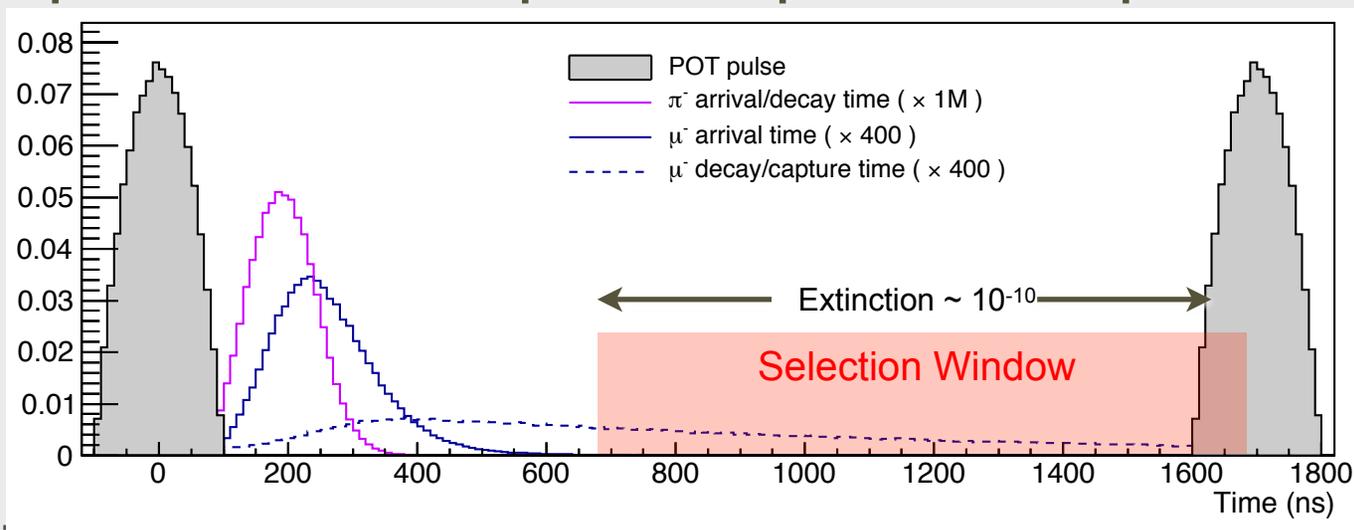
wait for pions to decay after pulse

protons out of beam pulse/ protons in beam-pulse < 10<sup>-10</sup>  
*and we must measure it*

# Pulsed Beam Structure

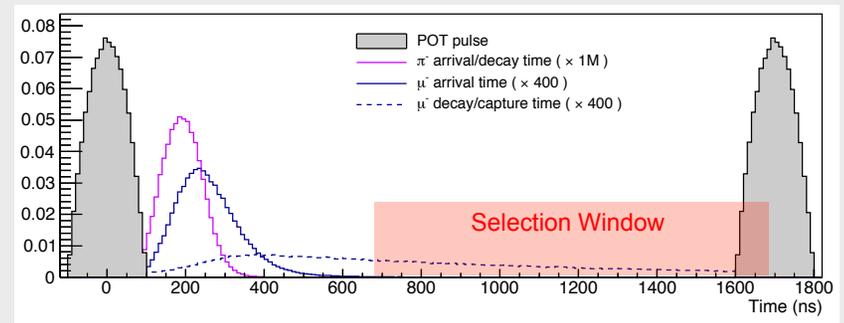
- Tied to prompt rate and machine: FNAL “perfect”
- Want **pulse duration**  $\ll \tau_{\mu}^{Al}$ , **pulse separation**  $\approx \tau_{\mu}^{Al}$ 
  - FNAL Ring has circumference  **$1.7\mu\text{sec}$** ,  $\sim \times 2\tau_{\mu}^{Al}$
- Extinction between pulses  $\leq 10^{-10}$  needed

= # protons out of pulse/# protons in pulse



# Prompt Background and Choice of Z

choose Z based on tradeoff  
between rate and lifetime:  
longer lived reduces prompt  
backgrounds



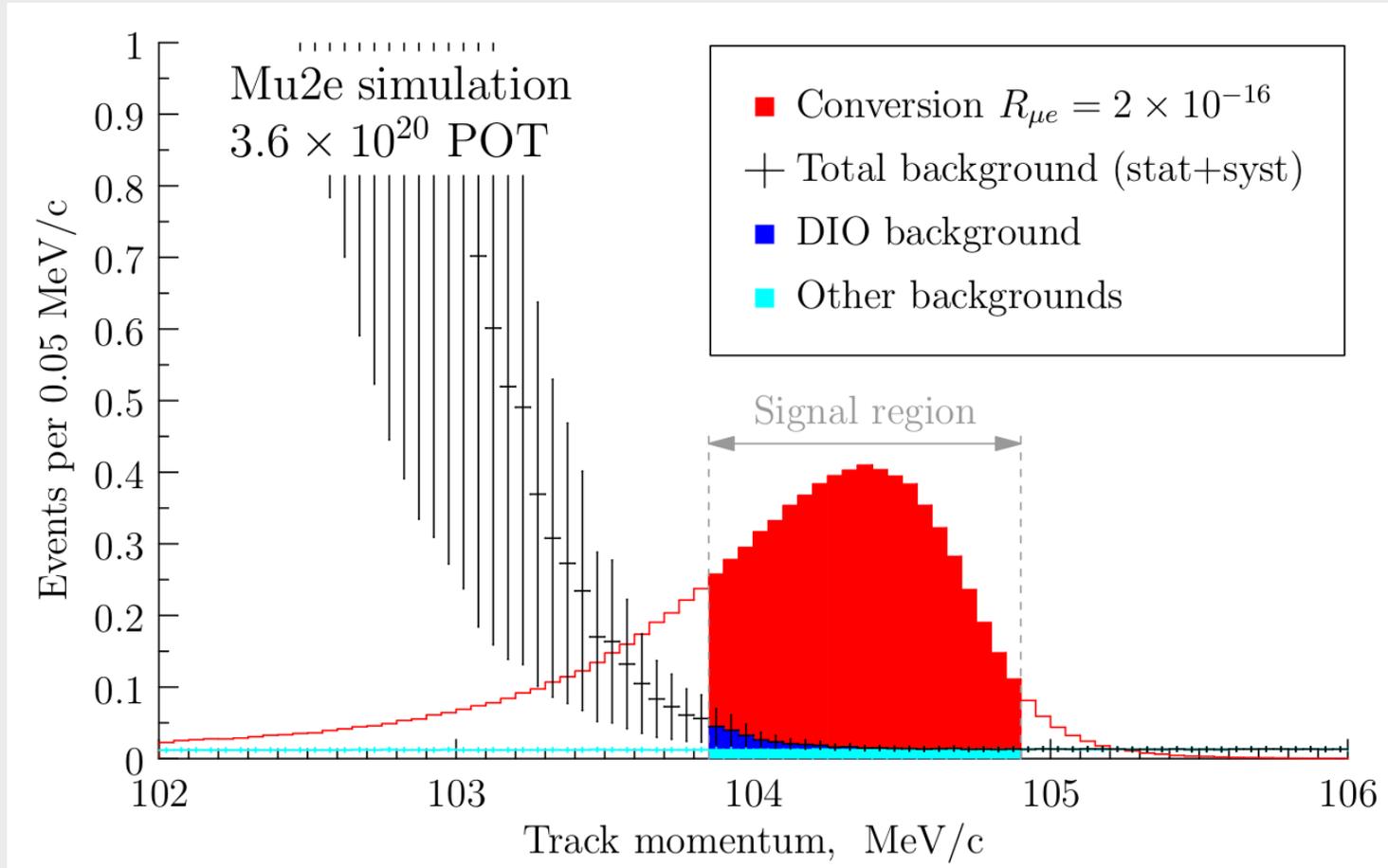
Nucleus	$R_{\mu e}(Z) / R_{\mu e}(Al)$	Bound Lifetime	Conversion Energy
Al(13,27)	1	864 nsec	104.96 MeV
Ti(22,~48)	1.7	328 nsec	104.18 MeV
Au(79,~197)	~0.8-1.5	72.6 nsec	95.56 MeV

# Signal and Background

- Full GEANT4 modeling and reconstruction without any truth input

$$5\sigma \sim 2 \times 10^{-16}$$

$$90\% \text{ CL} \sim 8 \times 10^{-17}$$



**typical SUSY at  $10^{-15}$ : 40 events vs 0.4 bkg**

# Final Backgrounds

• For  $R_{\mu e} = 10^{-15}$   
 $\sim 40$  events / 0.40 bkg

• For  $R_{\mu e} = 10^{-16}$   
 $\sim 4$  events / 0.40 bkg

Process	Expected event yield
Cosmic ray muons	$0.21 \pm 0.02(\text{stat}) \pm 0.06(\text{syst})$
DIO	$0.14 \pm 0.03(\text{stat}) \pm 0.11(\text{syst})$
Antiprotons	$0.040 \pm 0.001(\text{stat}) \pm 0.020(\text{syst})$
Pion capture	$0.021 \pm 0.001(\text{stat}) \pm 0.002(\text{syst})$
Muon DIF	$< 0.003$
Pion DIF	$0.001 \pm < 0.001$
Beam electrons	$(2.1 \pm 1.0) \times 10^{-4}$
RMC	$0.000^{+0.004}_{-0.000}$
Total	$0.41 \pm 0.13(\text{stat+syst})$

# Outline

- The search for muon-electron conversion
- Experimental Technique
- *Mu2e Upgrades*
- Conclusions

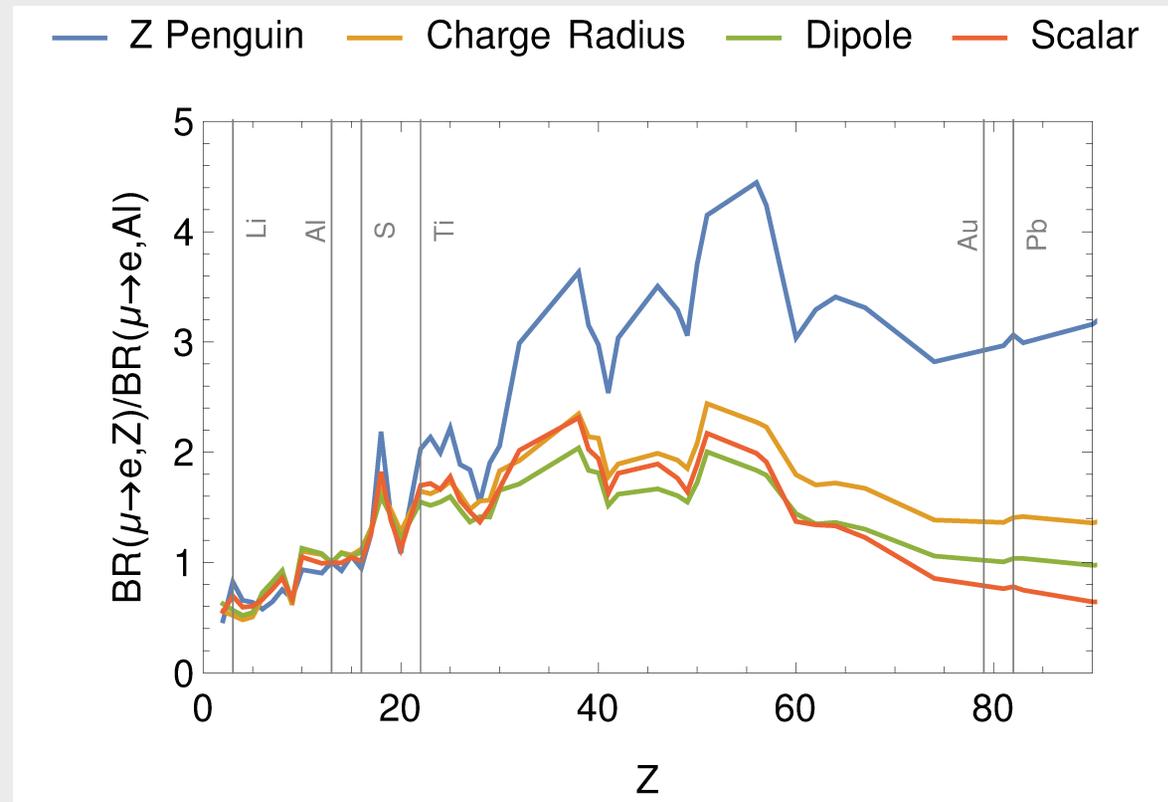
# Mu2e Upgrades

- Next Step in cLFV Program:



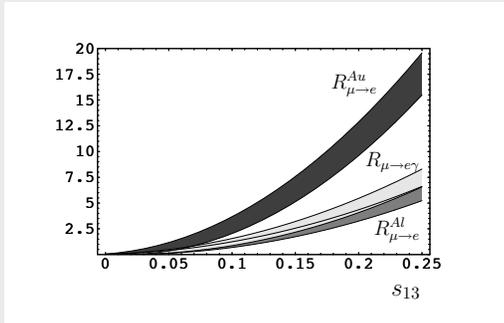
# Conversion at Higher Atomic Number

- Model Discrimination and Possibly Larger Signal at high  $Z$
- if Mu2e sees a signal, this is the obvious next step
- if not, we should try for another  $\times 10$ - $100$  better constraints



adapted from V. Cirigliano, B. Grinstein, G. Isidori, M. Wise **Nucl.Phys.B728:121-134,2005**

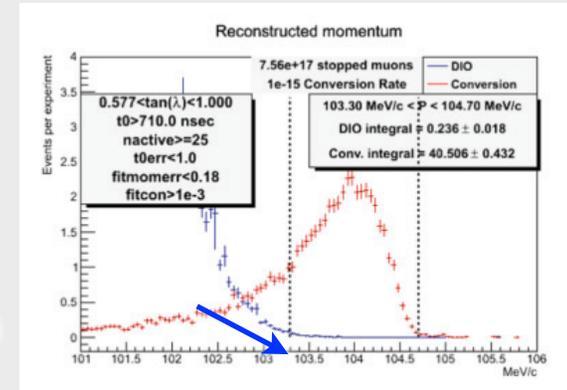
# Upgrade Plans...



Yes

Signal?

No



1. Change  $Z$  of Target to determine source of new physics

2. Prompt Rates will go up at higher  $Z$ , have to redesign detector and muon transport

1. Both Prompt and DIO backgrounds must drop to measure  $R_{\mu e} \sim 10^{-18}$

2. Detector, Muon Transport, Cosmic Ray Veto, Calorimeter

# Mu2e-II

- x10 upgrade of Mu2e, probably with Ti target
- Task Force formed
  - LOIs at Snowmass (13 on various aspects)
    - new calorimeter, tracker, upgraded CRV
- And LOIs for a muon program post Mu2e-II with PSI and FNAL experiments

[https://www.snowmass21.org/docs/files/summaries/RF/SNOWMASS21-RF5\\_RF0\\_Frank\\_Porter-106.pdf](https://www.snowmass21.org/docs/files/summaries/RF/SNOWMASS21-RF5_RF0_Frank_Porter-106.pdf)

[https://www.snowmass21.org/docs/files/summaries/RF/SNOWMASS21-RF5\\_RF0-AF5\\_AF0\\_Robert\\_Bernstein-027.pdf](https://www.snowmass21.org/docs/files/summaries/RF/SNOWMASS21-RF5_RF0-AF5_AF0_Robert_Bernstein-027.pdf)

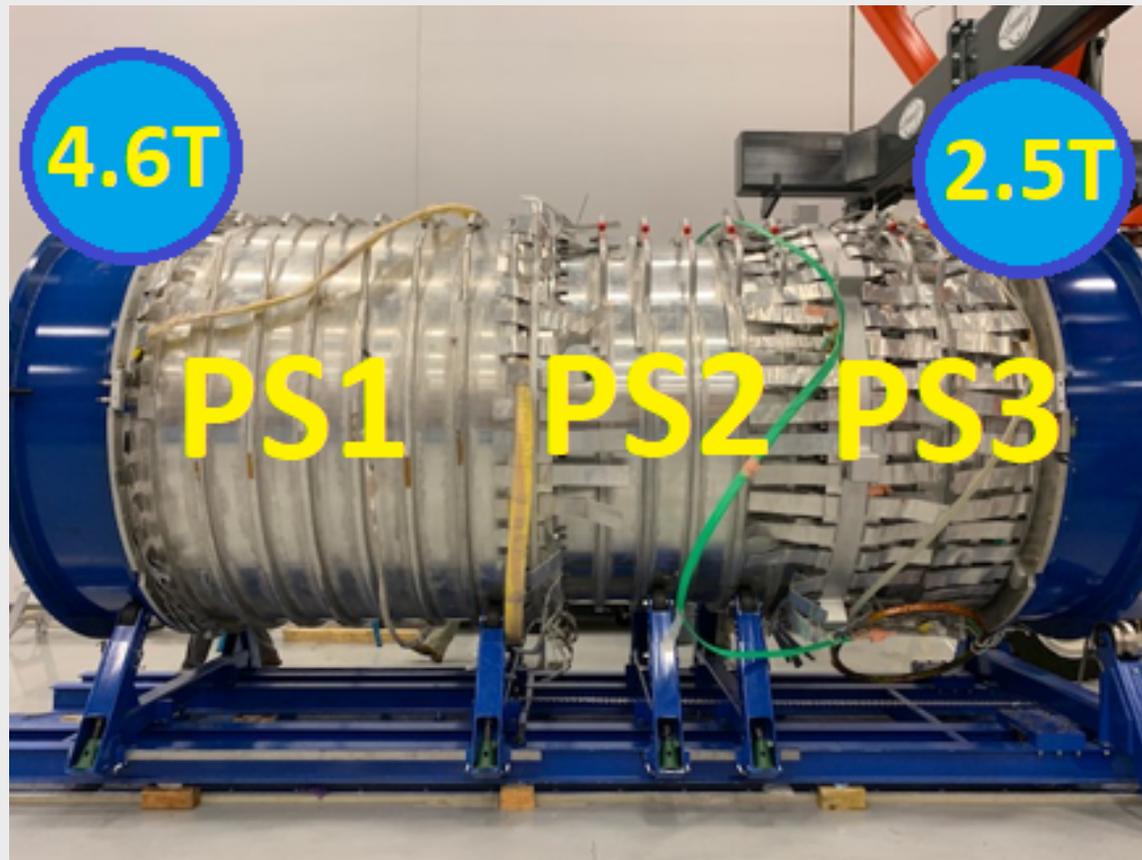
# Solenoid Status

- Transport Solenoid: about ready to go in Hall



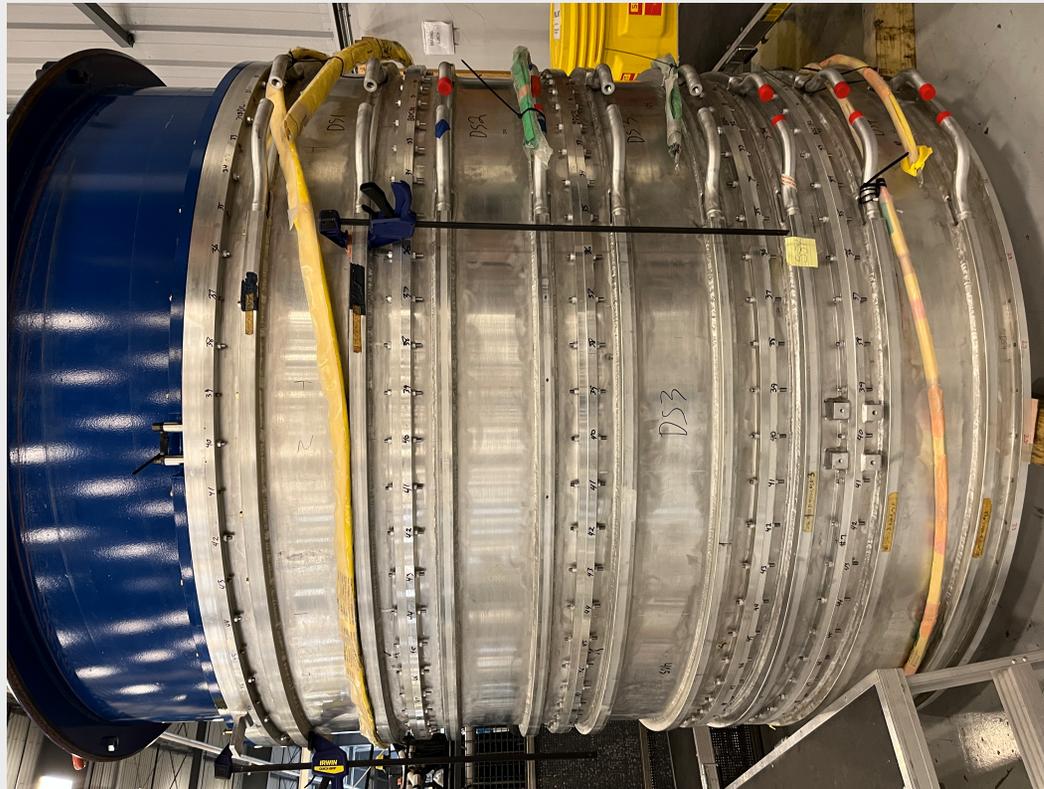
# Solenoid Status

- Production Solenoid



# Detector Solenoid

- Coils wound, in assembly (3 like this): tracker, calorimeter will fit inside this



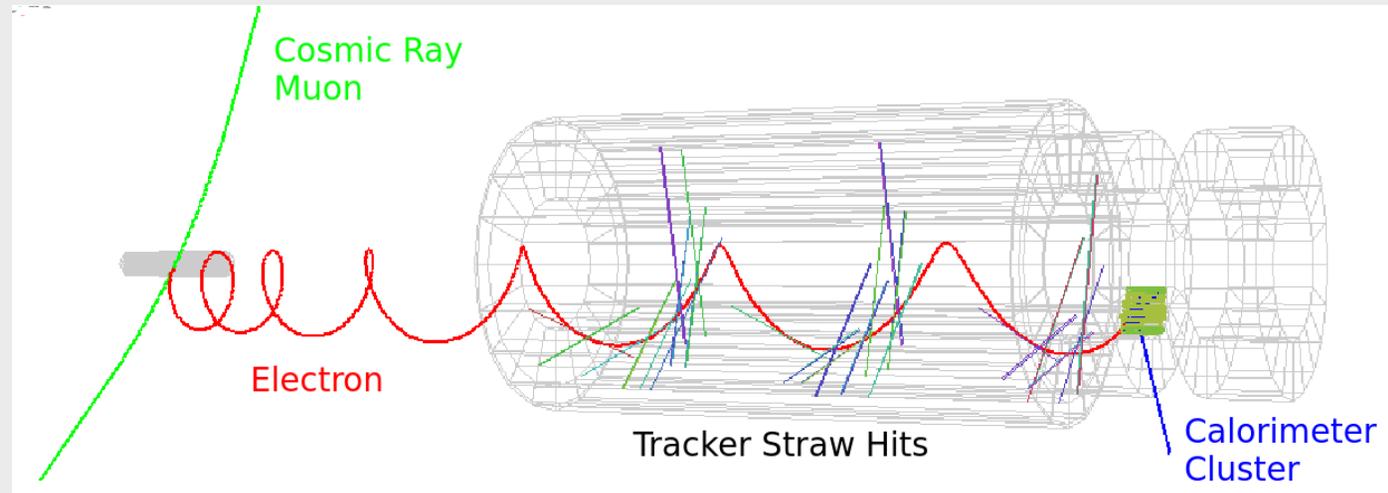
# Calorimeter

- Two disks, each 674 CsI crystals
- Fast energy measurement
  - for the trigger
  - combined with tracker momentum for PID
  - clusters can be used to seed track fit

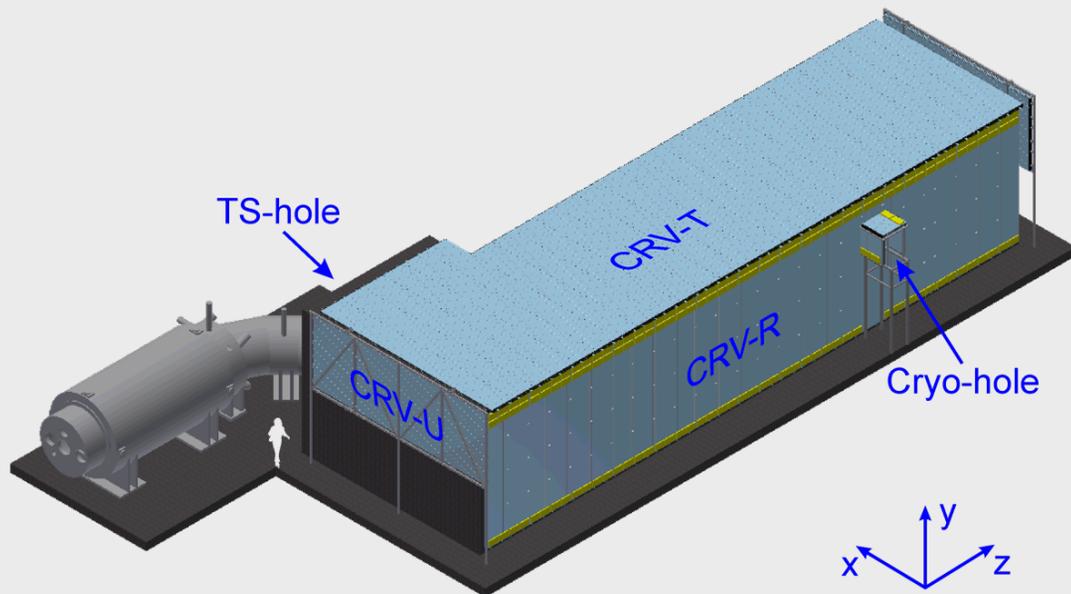


# Cosmic Ray Veto

- Would expect ~1 background/day from cosmic muons; CRV must be 99.99% efficient



finished and shipped to FNAL; debugging readout



# Tracker

- High Resolution (180 keV FWHM) momentum measurement
    - minimize energy loss with low mass straws
    - dual-end readout for location along straw
- 5mm diameter, 15  $\mu$  thick



36 planes = 216 panels

# Schedule

- Data-taking 2025-6
- x100 existing in a few weeks of running
- x1000 in under a year before LBNF/PIP-II Shutdown
- 4-5 years of running for full data set

# Conclusions

- Mu2e will:
  - Reduce the limit for  $R_{\mu e}$  by more than four orders of magnitude, x10 in mass reach ( $R_{\mu e} < 8 \times 10^{-17}$  @ 90% C.L.)
  - Discover unambiguous proof of new physics or
  - Set powerful constraints on a wide variety of models
- Mu2e will therefore both complement LHC results and independently probe up to  $10^4$  TeV/ $c^2$
- With upgrades, we could extend the limit by up to two additional orders of magnitude, study the details of new physics, and build a new rare muon process program

**WHY DO WHALES JUMP**  
**WHY ARE WITCHES GREEN**  
**WHY ARE THERE MIRRORS ABOVE BEDS**  
**WHY DO I SAY UH**  
**WHY IS SEA SALT BETTER**  
**WHY ARE THERE TREES IN THE MIDDLE OF FIELDS**  
**WHY IS THERE NOT A POKEMON MMO**  
**WHY IS THERE LAUGHING IN TV SHOWS**  
**WHY ARE THERE DOORS ON THE FREEWAY**  
**WHY ARE THERE SO MANY SUCROSTEXE RUNNING**  
**WHY AREN'T THERE ANY COUNTRIES IN ANTARCTICA**  
**WHY ARE THERE SCARY SOUNDS IN MINECRAFT**  
**WHY IS THERE KICKING IN MY STOMACH**  
**WHY ARE THERE TWO SLASHES AFTER HTTP**  
**WHY ARE THERE CELEBRITIES**  
**WHY DO SNAKES EXIST**  
**WHY DO OYSTERS HAVE PEARLS**  
**WHY ARE DUCKS CALLED DUCKS**  
**WHY DO THEY CALL IT THE CLAP**  
**WHY ARE KYLE AND CARTMAN FRIENDS**  
**WHY IS THERE AN ARROW ON AANG'S HEAD**  
**WHY ARE TEXT MESSAGES BLUE**  
**WHY ARE THERE MUSTACHES ON CLOTHES**  
**WHY ARE THERE MUSTACHES ON CARS**  
**WHY ARE THERE MUSTACHES EVERYWHERE**  
**WHY ARE THERE SO MANY BIRDS IN OHIO**  
**WHY IS THERE SO MUCH RAIN IN OHIO**  
**WHY IS OHIO WEATHER SO WEIRD**  
**WHY ARE THERE MALE AND FEMALE BIKES**  
**WHY ARE THERE BRIDESMAIDS**  
**WHY DO DYING PEOPLE SCREAM UP**  
**WHY AREN'T THERE UNKOSSE AIRBOTS**  
**WHY ARE OLD KUNGUIS DIFFERENT**  
**WHY ARE THERE SQUIRRELS**  
**WHY ARE THERE TINY SPIDERS IN MY HOUSE**  
**WHY DO SPIDERS COME INSIDE**  
**WHY ARE THERE HUGE SPIDERS IN MY HOUSE**  
**WHY ARE THERE LOTS OF SPIDERS IN MY HOUSE**  
**WHY ARE THERE SPIDERS IN MY ROOM**  
**WHY ARE THERE SO MANY SPIDERS IN MY ROOM**  
**WHY DO SPIDER BITES ITCH**  
**WHY IS DYING SO SCARY**  
**WHY IS THERE NO GPS IN LAPTOPS**  
**WHY DO KNEES CLICK**  
**WHY IS PROGRAMMING SO HARD**  
**WHY IS THERE A 0.0MM RESISTOR**  
**WHY DO AMERICANS HATE SOCCER**  
**WHY DO RHYMES SOUND GOOD**  
**WHY DO TREES DIE**  
**WHY IS THERE NO SOUND ON QM**  
**WHY AREN'T POKEMON REAL**  
**WHY AREN'T BULLETS SHARP**  
**WHY DO DREAMS SEEM SO REAL**  
**WHY DO TESTICLES MOVE**  
**WHY ARE THERE PSYCHICS**  
**WHY ARE HATS SO EXPENSIVE**  
**WHY IS THERE CAFFEINE IN MY SHIRT**  
**WHY DO YOUR BOOBS HURT**  
**WHY ARE THERE SLAVES IN THE BIBLE**  
**WHY DO TWINS HAVE DIFFERENT FINGERPRINTS**  
**WHY ARE AMERICANS AFRAID OF DRAGONS**  
**WHY IS HTTPS CROSSED OUT IN RED**  
**WHY IS THERE A LINE THROUGH HTTPS**  
**WHY IS THERE A RED LINE THROUGH HTTPS ON FACEBOOK**  
**WHY IS HTTPS IMPORTANT**  
**WHY AREN'T MY ARMS GROWING**  
**WHY ARE THERE WEBS IN**  
**WHY DO I FEEL DIZZY**  
**WHY ARE THERE SO MANY CROWS IN ROCHESTER**  
**WHY IS PSYCHIC WEAK TO BUG**  
**WHY DO CHILDREN GET CANCER**  
**WHY IS POSEIDON ANGRY WITH ODYSSEUS**  
**WHY IS THERE ICE IN SPACE**  
**WHY AREN'T ECONOMISTS RICH**  
**WHY DO AMERICANS CALL IT SOCCER**  
**WHY ARE MY EARS RINGING**  
**WHY ARE THERE SO MANY AVENGERS**  
**WHY ARE THE AVENGERS FIGHTING THE X MEN**  
**WHY IS WOLVERINE NOT IN THE AVENGERS**  
**WHY ARE THERE ANTS IN MY LAPTOP**  
**WHY IS EARTH TILTED**  
**WHY IS SPACE BLACK**  
**WHY IS OUTER SPACE SO COLD**  
**WHY ARE THERE PYRAMIDS ON THE MOON**  
**WHY IS NASA SHUTTING DOWN**  
**WHY ARE THERE GHOSTS**  
**WHY IS THERE AN OWL IN MY BACKYARD**  
**WHY IS THERE AN OWL OUTSIDE MY WINDOW**  
**WHY IS THERE AN OWL ON THE DOLLAR BILL**  
**WHY DO OWLS ATTACK PEOPLE**  
**WHY ARE AK 47s SO EXPENSIVE**  
**WHY ARE THERE HELICOPTERS CIRCLING MY HOUSE**  
**WHY ARE THERE GODS**  
**WHY ARE THERE TWO SPOOKS**  
**WHY IS MT VESUVIUS THERE**  
**WHY DO THEY SAY T MINUS**  
**WHY ARE THERE OBELISKS**  
**WHY ARE WRESTLERS ALWAYS WET**  
**WHY ARE OCEANS BECOMING MORE ACIDIC**  
**WHY IS ARWEN DYING**  
**WHY AREN'T MY QUAIL LAYING EGGS**  
**WHY AREN'T MY QUAIL EGGS HATCHING**  
**WHY AREN'T THERE ANY FOREIGN MILITARY BASES IN AMERICA**  
**WHY ARE MY BOOBS ITCHY**  
**WHY ARE CIGARETTES LEGAL**  
**WHY ARE THERE DUCKS IN MY POOL**  
**WHY IS JESUS WHITE**  
**WHY IS THERE LIQUID IN MY EAR**  
**WHY DO Q TIPS FEEL GOOD**  
**WHY DO GOOD PEOPLE DIE**  
**WHY AREN'T THERE GUNS IN HARRY POTTER**  
**WHY ARE ULTRASOUNDS IMPORTANT**  
**WHY ARE ULTRASOUND PROBED EXPENSIVE**  
**WHY IS STEALING WRONG**  
**WHY IS LIFE SO BORING**  
**WHY ARE THERE RED DOTS ON MY THIGHS**  
**WHY IS SEX SO IMPORTANT**  
**WHY IS THERE ALWAYS A JAMA UPDATE**  
**WHY ARE THERE RED DOTS ON MY THIGHS**  
**WHY IS LYING GOOD**  
**WHY ARE THERE DINOSAUR GHOSTS**  
**WHY ARE THERE FEMALE MR NIMES**  
**WHY ARE THERE DUCKS AFRAID OF FIREWORKS**  
**WHY IS THERE NO KING IN ENGLAND**  
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**QUESTIONS**  
 FOUND IN GOOGLE AUTOCOMPLETE