
Current and Future Neutrino Experiments

Dan Dwyer

Brookhaven Forum

6 Oct. 2023

Outline

- The Era of Three-Flavor Mixing
- Pursuing Neutrino Mass
- Beyond Three-flavor Mixing

Neutrinos: The Era of Three-Flavor Mixing

History of Neutrino Masses and Mixing: Initial Evidence

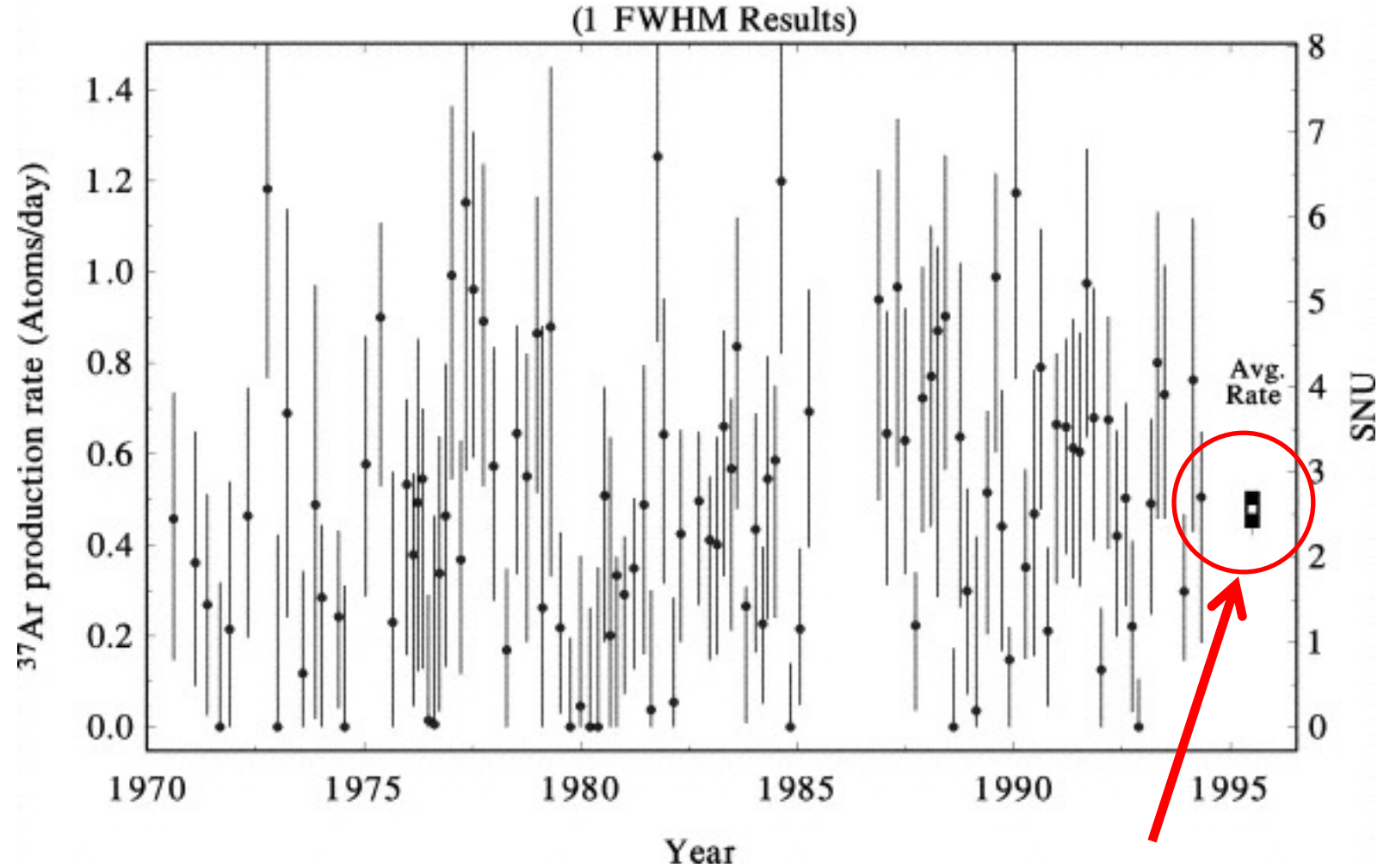
Homestake Experiment:

Ambitious radiochemical experiment

Search for argon atoms in 100,000 g of Cl



Observed deficit relative to prediction



Average rate: 1/3 of that expected from solar models.

History of Neutrino Masses and Mixing: Initial Evidence

Kamiokande Experiment:

1-kton Water Cherenkov detector

Observed electrons scattered by solar neutrinos, with deficit similar to Homestake Experiment.

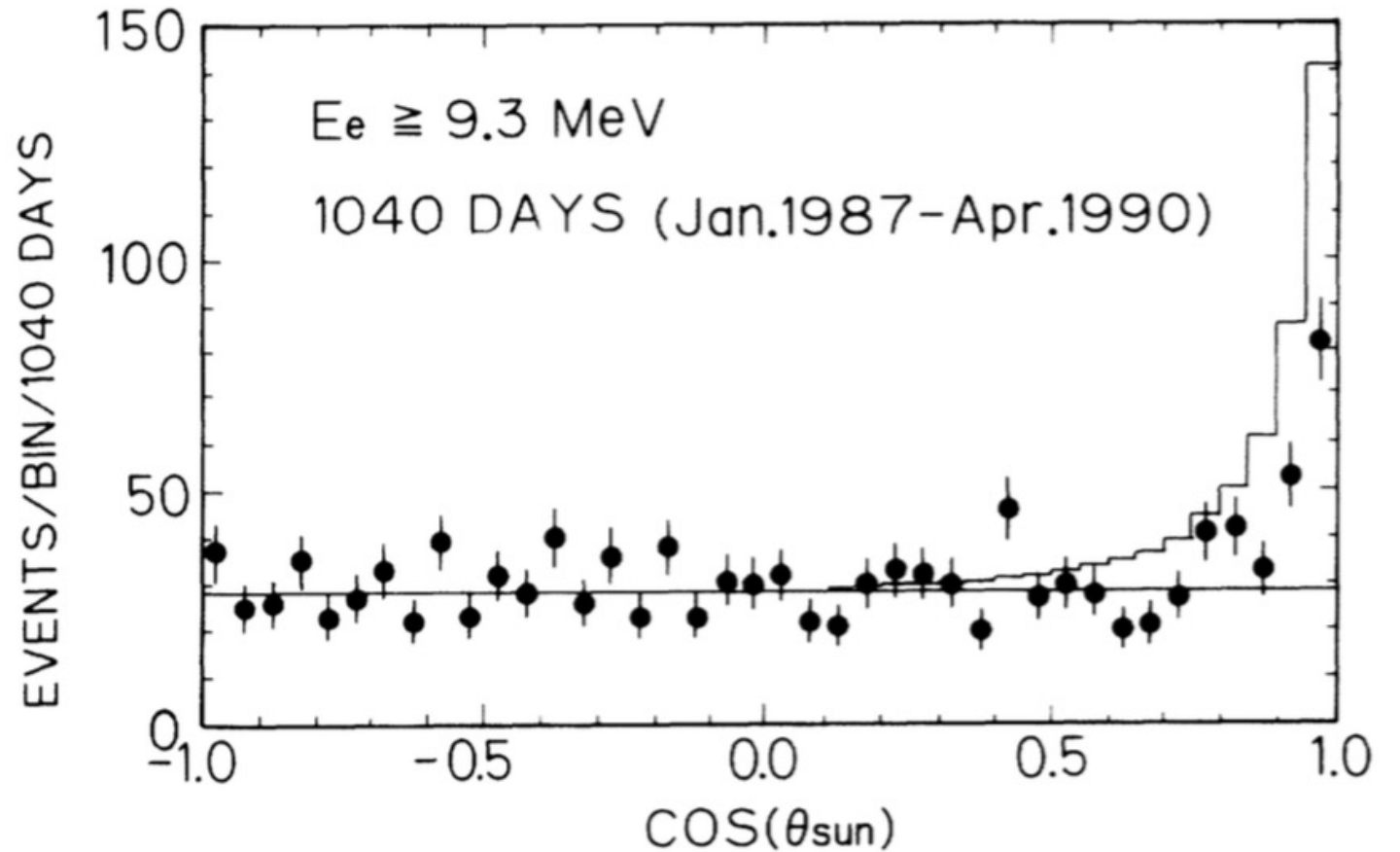
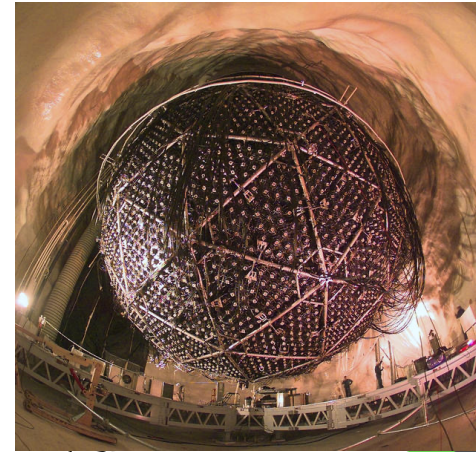
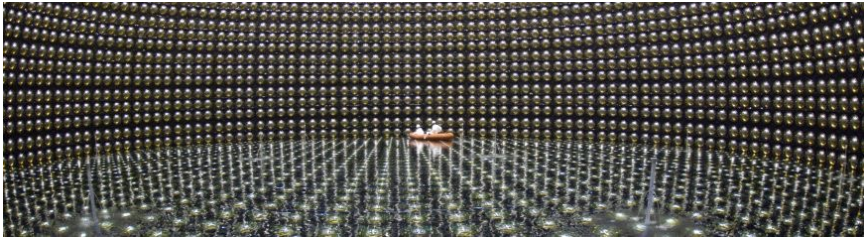
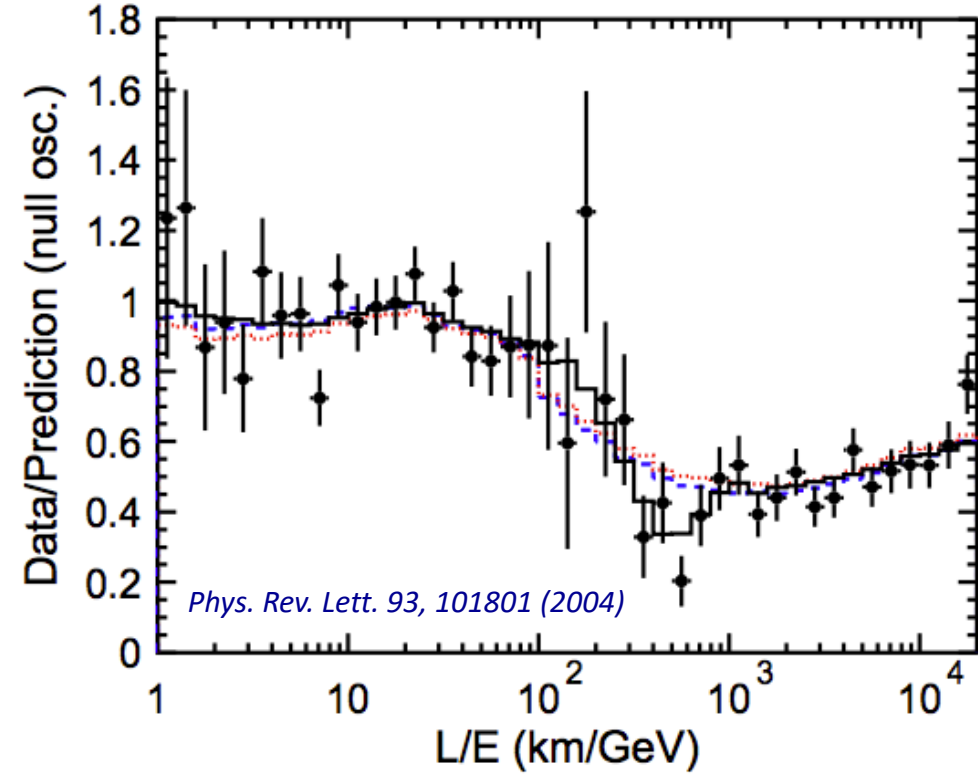


FIG. 3. Distribution in $\cos\theta_{\text{Sun}}$ of the combined 1040-day sample for $E_e \geq 9.3 \text{ MeV}$. The value of the ratio data/SSM from this figure is 0.43 ± 0.06 .

History of Neutrino Masses and Mixing

Super-Kamiokande:

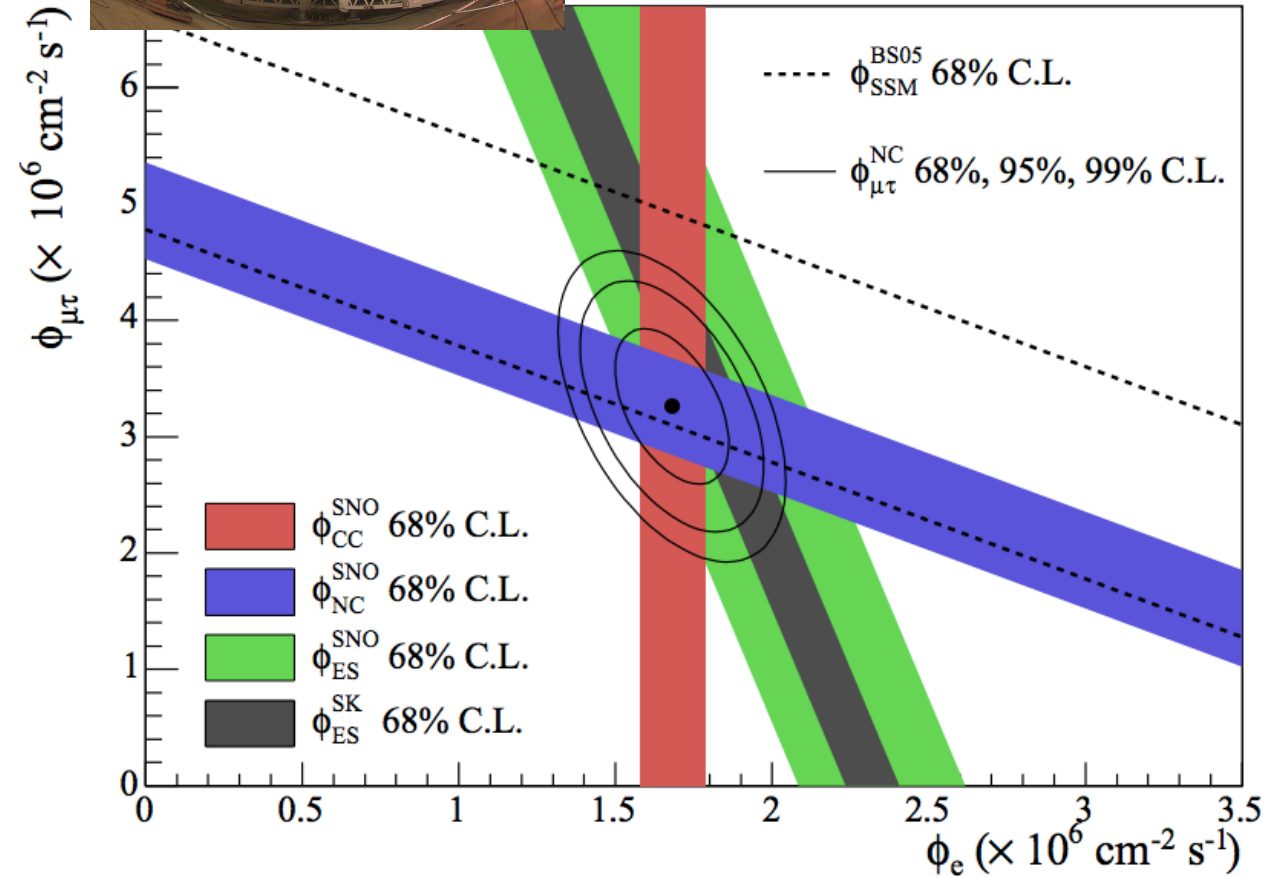
Observes disappearance of neutrinos produced in atmosphere



Sudbury Neutrino Observatory:

Observes flavor change for neutrinos from the sun.

Phys. Rev. C 72, 055502 (2005)



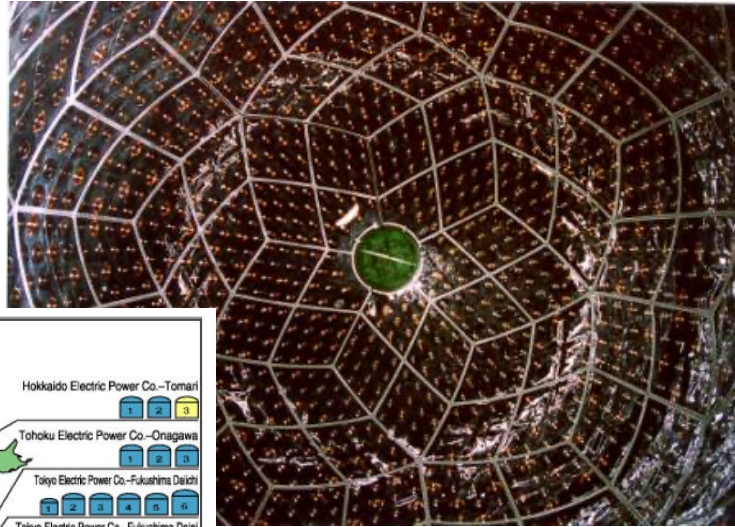
History of Neutrino Masses and Mixing

KamLAND:

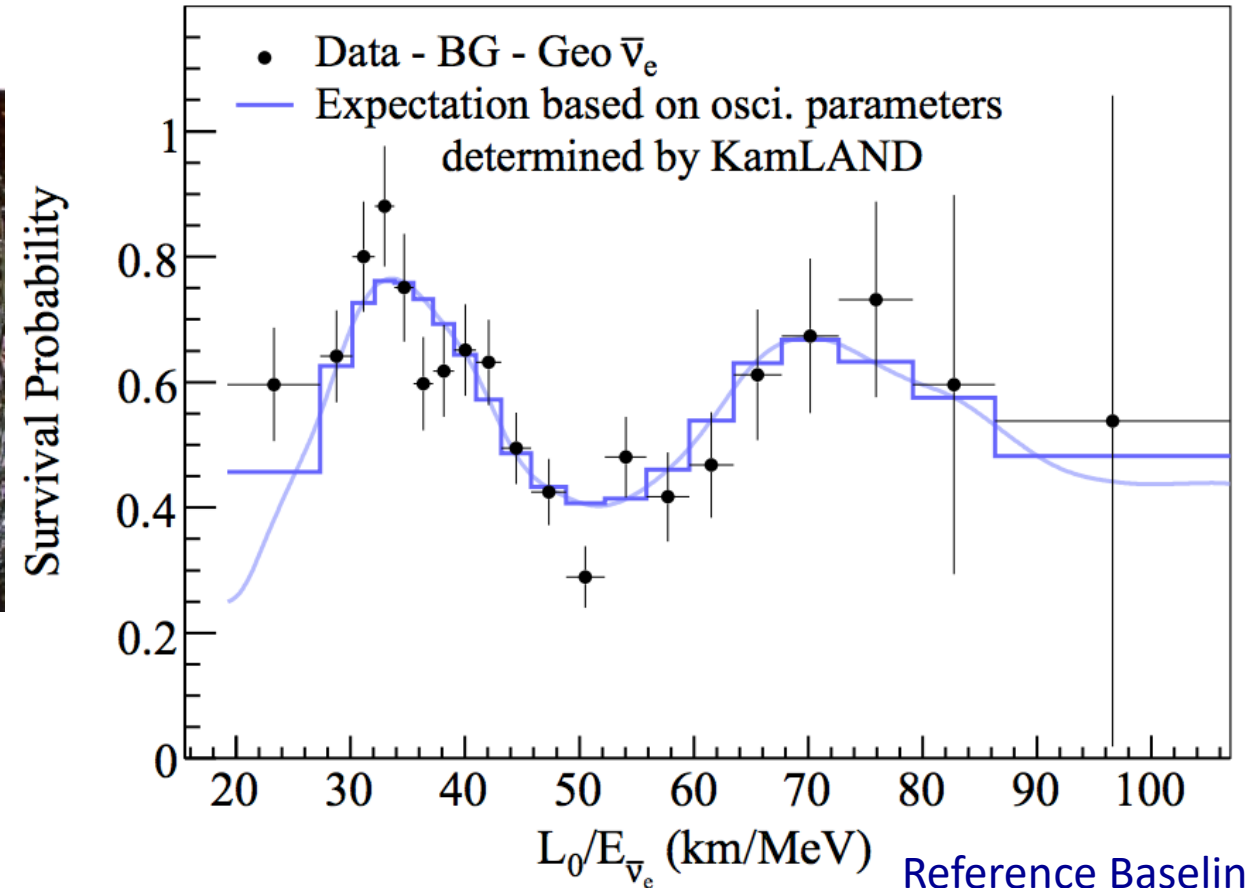
1 kton of liquid scintillator detects antineutrinos emitted by nuclear reactors

Observes energy-dependent disappearance / reappearance

-> *Neutrino flavor oscillation!*

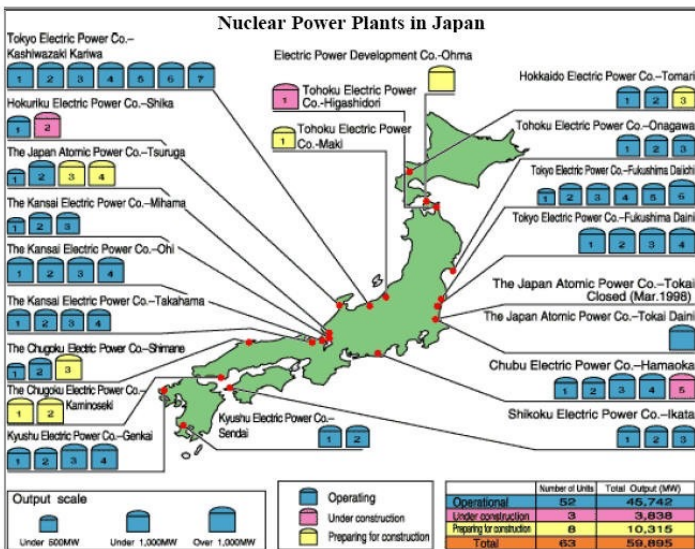


$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \sin^2(2\theta_{12}) \sin^2 \frac{\Delta m_{12}^2 L}{4E}$$



Phys. Rev. Lett. 100, 221803 (2008)

Reference Baseline:
 $L_0 = 180 \text{ km}$



History of Neutrino Masses and Mixing

Reactor Antineutrinos

Daya Bay

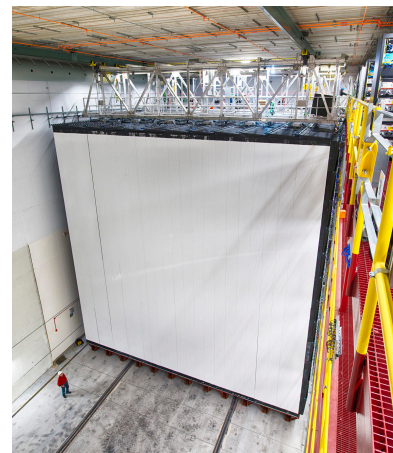
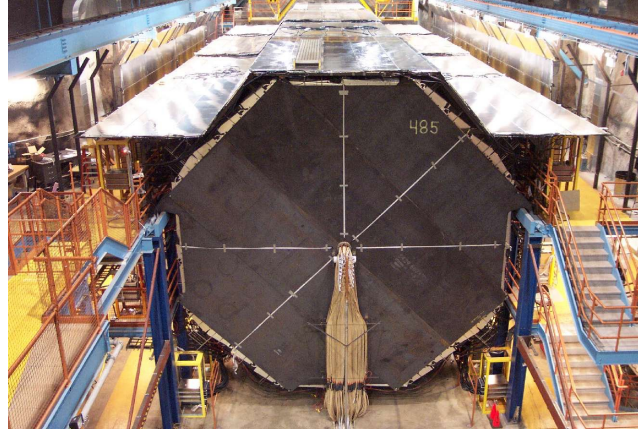
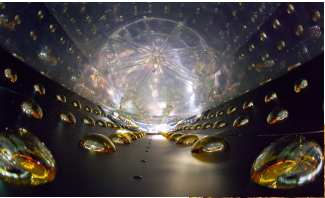
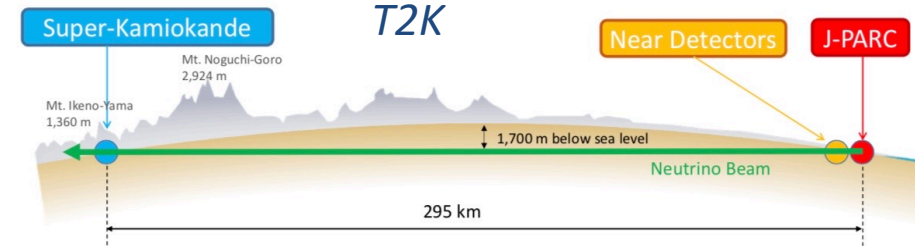
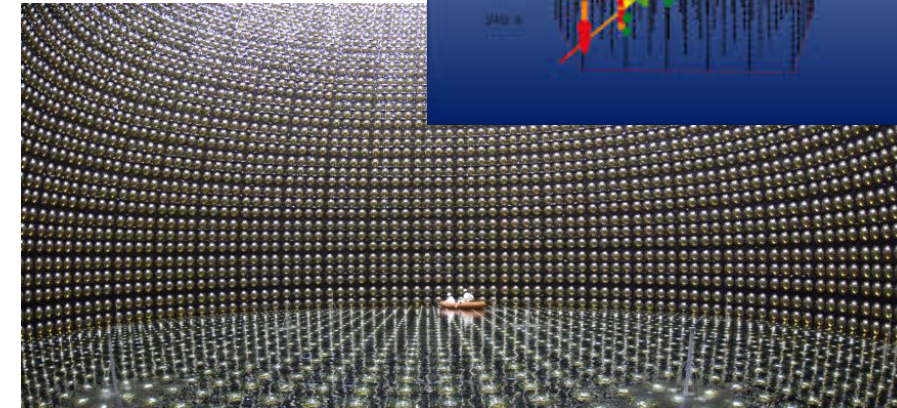
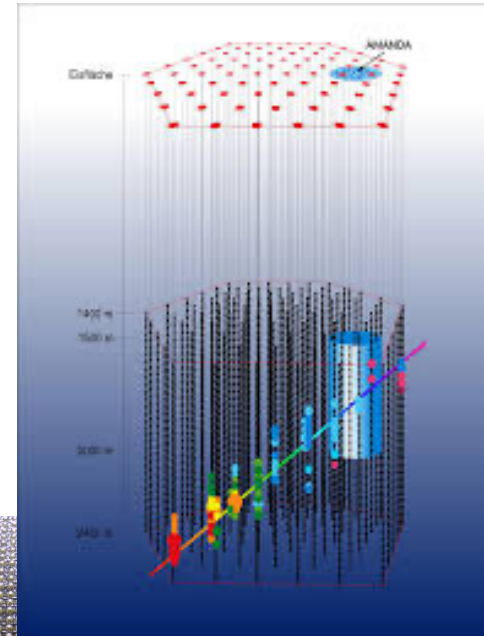
RENO

MINOS

Accelerator Neutrinos

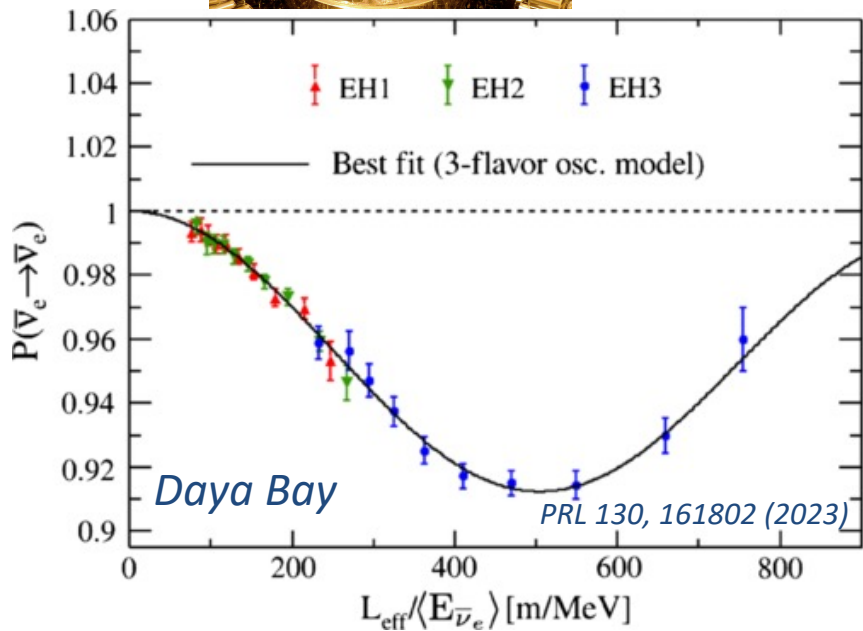
Atmospheric Neutrinos

IceCUBE

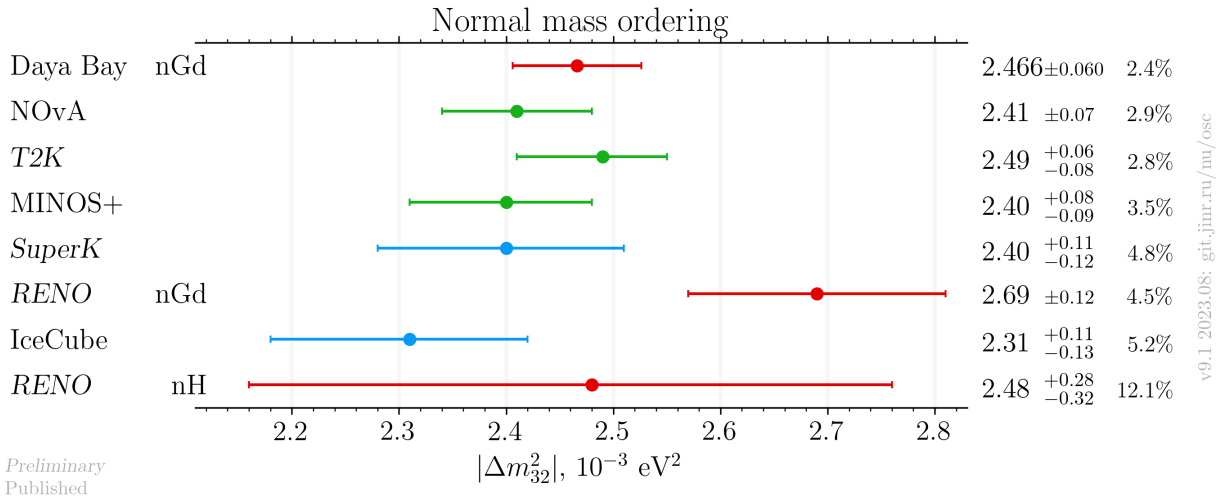


NOvA

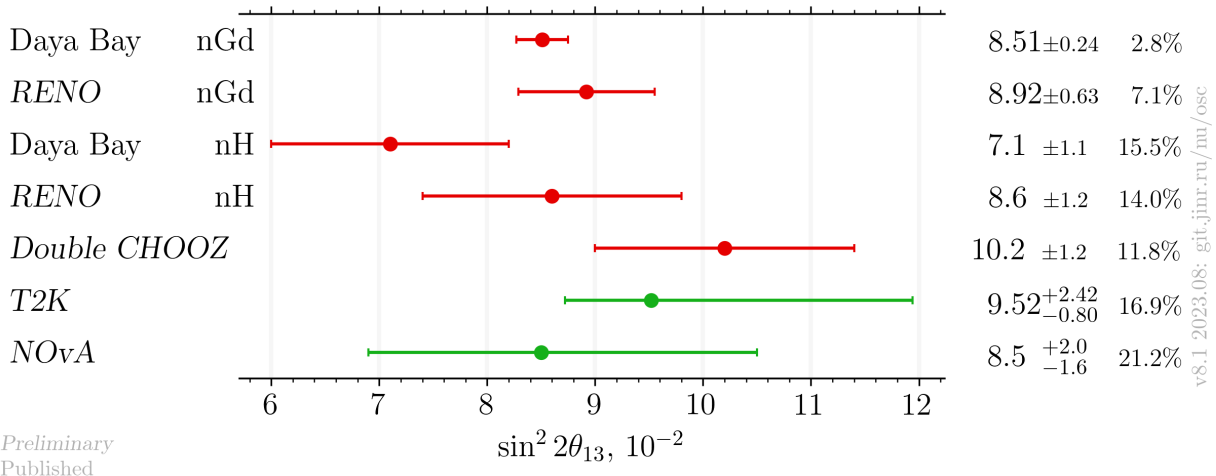
Double CHOOZ



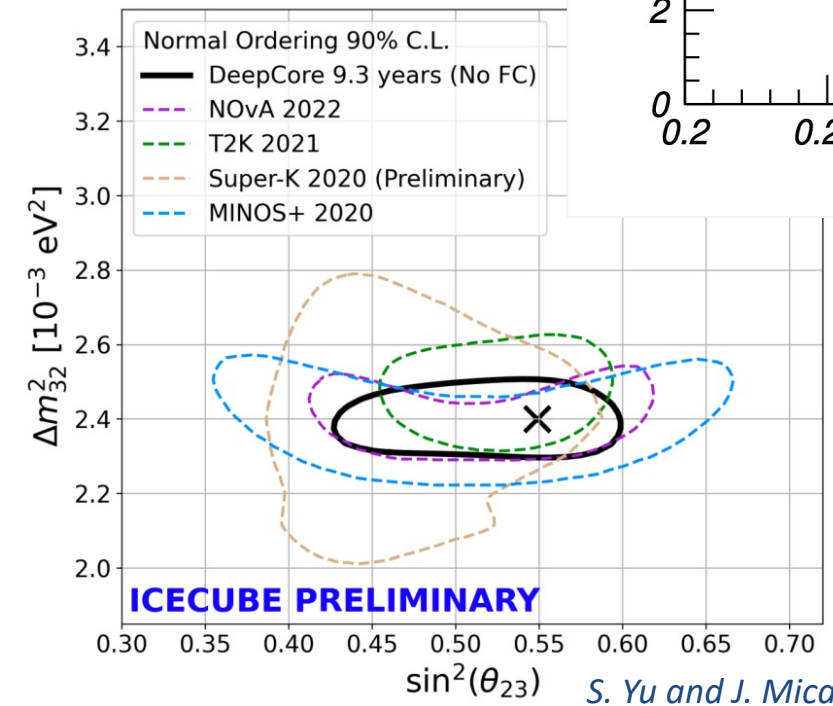
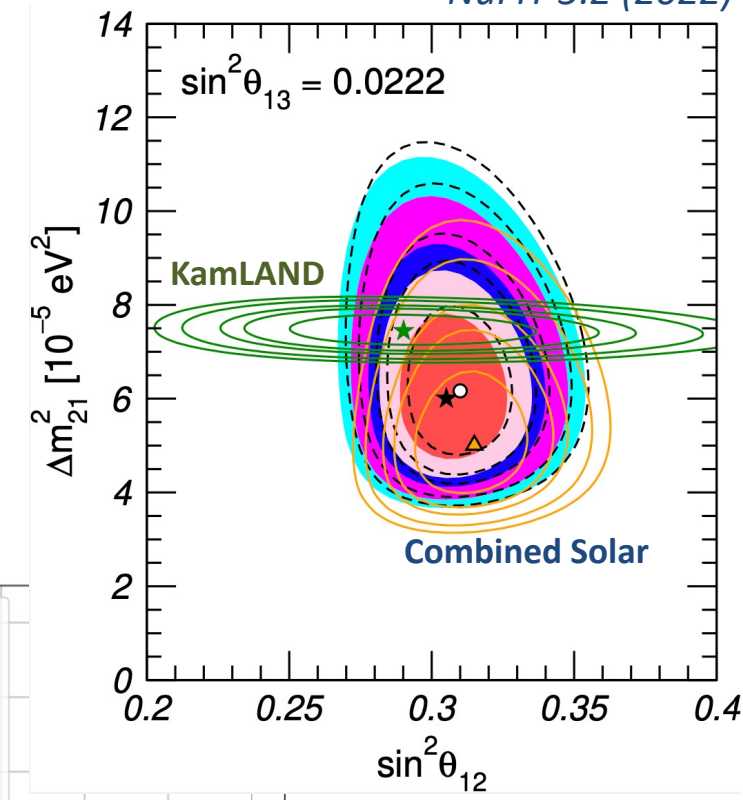
Neutrino Masses and Mixing: Results Summary



Preliminary
Published



Preliminary
Published

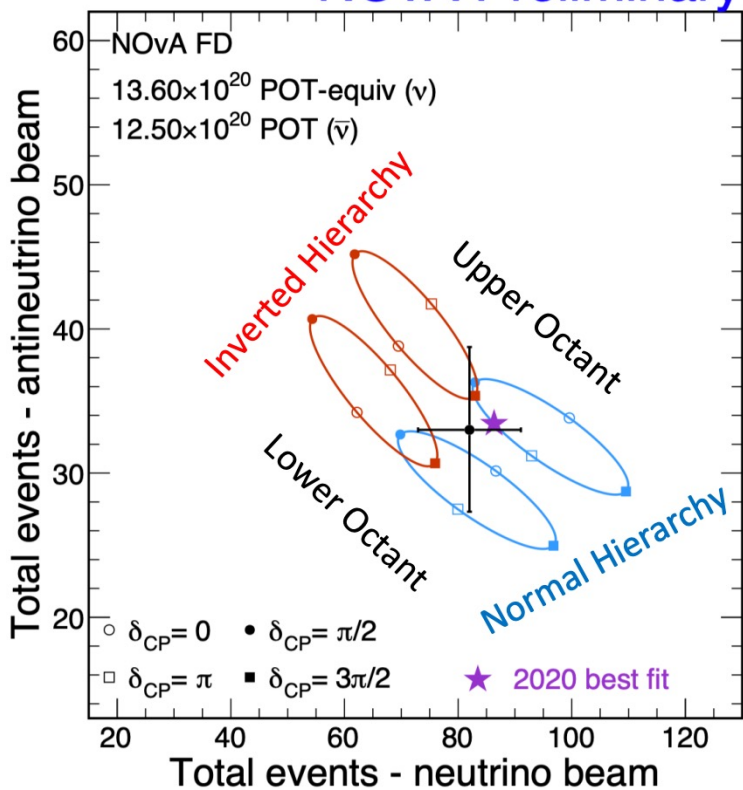


S. Yu and J. Micallef, ICRC 2023

Neutrino Masses and Mixing: Latest Results

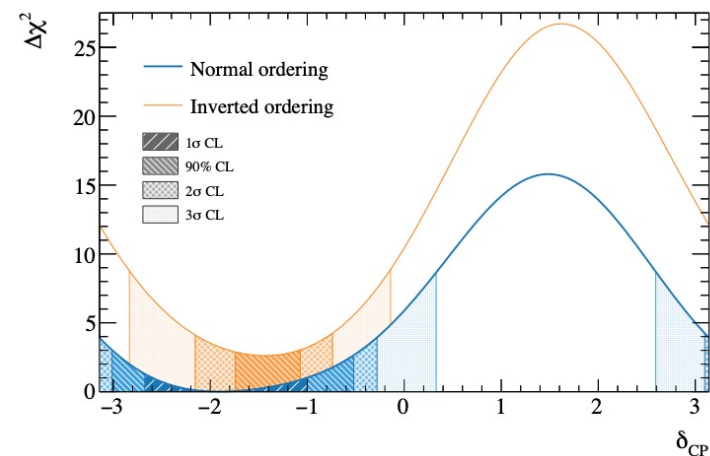
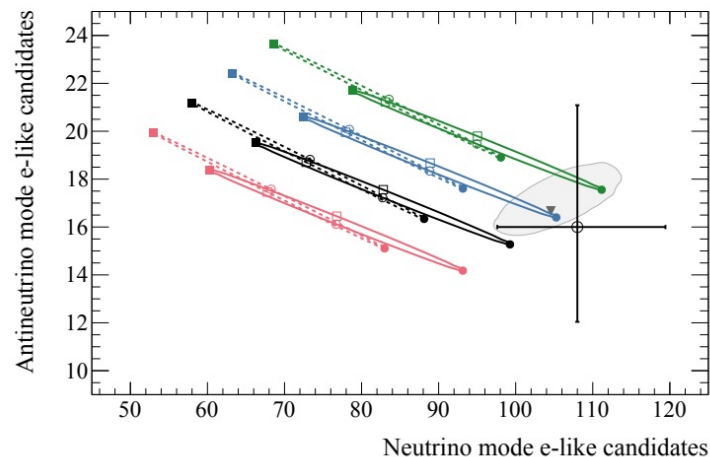
Gaining sensitivity to the neutrino mass ordering and possible CP Violation

NOvA Preliminary

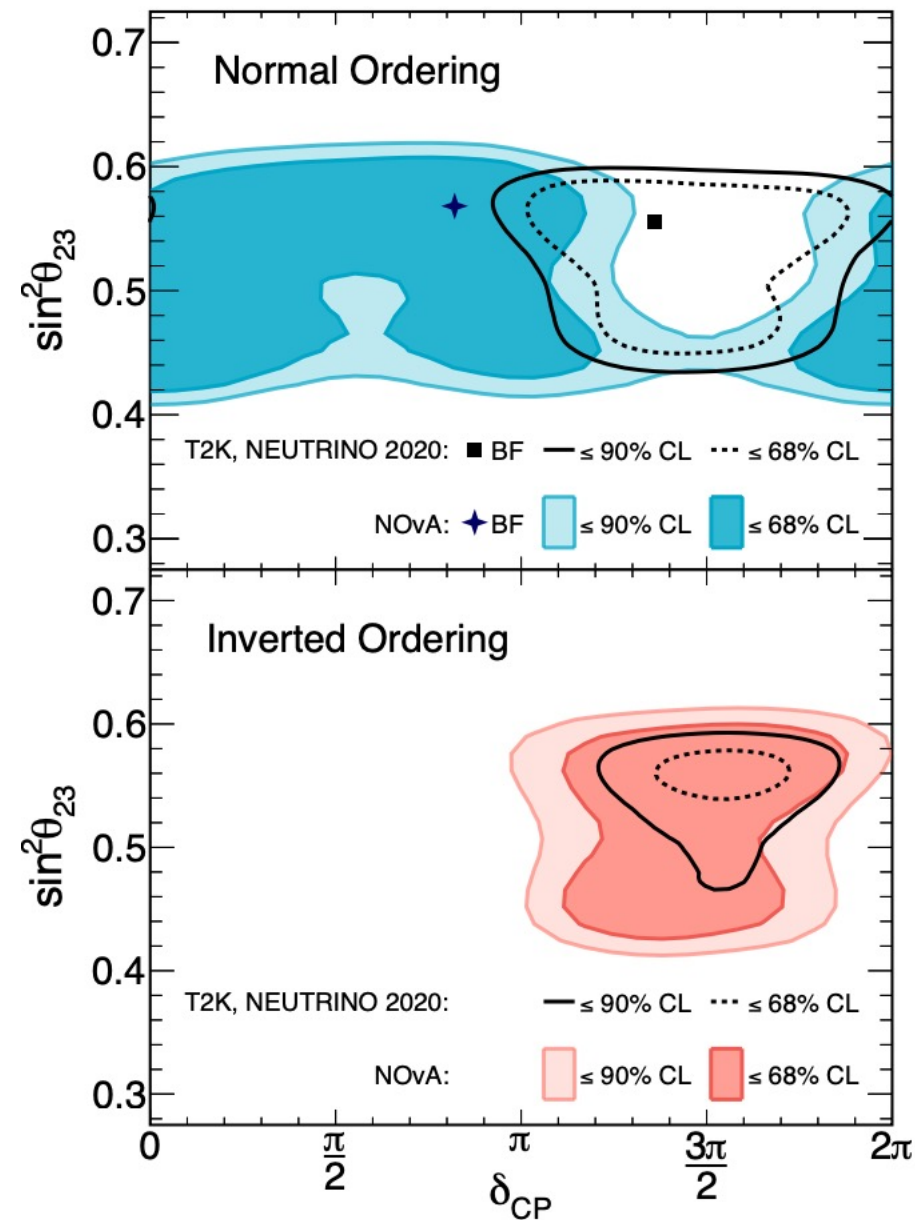


M. Sanchez, TAUP 2023

 68% syst err. at best-fit
▼ Best-fit
 Data (68% stat err.)
 $\sin^2\theta_{23} = 0.45, 0.50, 0.55, 0.60$
 $\Delta m_{32}^2 = 2.49 \times 10^{-3} \text{ eV}^2 \text{ (NO)}$
 $\Delta m_{31}^2 = -2.46 \times 10^{-3} \text{ eV}^2 \text{ (IO)}$
 $\delta_{CP} = \pi$
 $\delta_{CP} = +\pi/2$
 $\delta_{CP} = 0$
 $\delta_{CP} = -\pi/2$

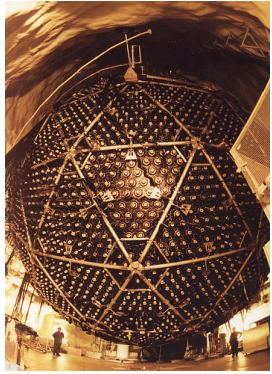


arxiv:2303.03222



M. Sanchez, TAUP 2023

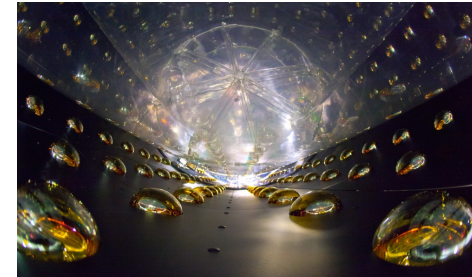
Current Status of Three-flavor Neutrino Oscillation Model



Neutrino Mixing

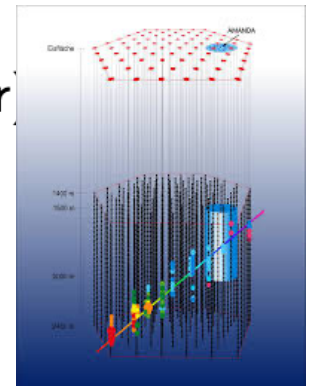
*R.L. Workman et al. (Particle Data Group),
Prog. Theor. Exp. Phys. 2022, 083C01 (2022) and 2023 update*

The following values are obtained through data analyses based on the 3-neutrino mixing scheme described in the review “Neutrino Masses, Mixing, and Oscillations.”



Leading Measurements:

- SNO, SuperK* → $\sin^2(\theta_{12}) = 0.307 \pm 0.013$
- KamLAND* → $\Delta m_{21}^2 = (7.53 \pm 0.18) \times 10^{-5} \text{ eV}^2$
- T2K, NOvA, IceCUBE* → $\sin^2(\theta_{23}) = 0.534^{+0.021}_{-0.024}$ (Inverted order)
- T2K, NOvA, IceCUBE* → $\sin^2(\theta_{23}) = 0.547^{+0.018}_{-0.024}$ (Normal order)
- Daya Bay, IceCUBE, T2K, NOvA* → $\Delta m_{32}^2 = (-2.519 \pm 0.033) \times 10^{-3} \text{ eV}^2$ (Inverted order)
- Daya Bay, IceCUBE, T2K, NOvA* → $\Delta m_{32}^2 = (2.437 \pm 0.033) \times 10^{-3} \text{ eV}^2$ (Normal order)
- Daya Bay* → $\sin^2(\theta_{13}) = (2.20 \pm 0.07) \times 10^{-2}$
- T2K, NOvA* → $\delta, CP \text{ violating phase} = 1.23 \pm 0.21 \pi \text{ rad}$ (S = 1.3)



Neutrino Masses and Mixing: Future Experiments

Neutrino Mixing

*R.L. Workman et al. (Particle Data Group),
Prog. Theor. Exp. Phys. 2022, 083C01 (2022) and 2023 update*

The following values are obtained through data analyses based on the 3-neutrino mixing scheme described in the review “Neutrino Masses, Mixing, and Oscillations.”

**Upcoming
Measurements:**

Mass Ordering:

*T2K, NOvA, JUNO, T2HK,
IceCUBE upgrade, DUNE*

Search for CP-Violation:

T2K, NOvA, T2HK, DUNE

$$\sin^2(\theta_{12}) = 0.307 \pm 0.013$$

$$\Delta m_{21}^2 = (7.53 \pm 0.18) \times 10^{-5} \text{ eV}^2$$

$$\sin^2(\theta_{23}) = 0.534^{+0.021}_{-0.024} \quad (\text{Inverted order})$$

$$\sin^2(\theta_{23}) = 0.547^{+0.018}_{-0.024} \quad (\text{Normal order})$$

$$\Delta m_{32}^2 = (-2.519 \pm 0.033) \times 10^{-3} \text{ eV}^2 \quad (\text{Inverted order})$$

$$\Delta m_{32}^2 = (2.437 \pm 0.033) \times 10^{-3} \text{ eV}^2 \quad (\text{Normal order})$$

$$\sin^2(\theta_{13}) = (2.20 \pm 0.07) \times 10^{-2}$$

$$\delta, \text{ CP violating phase} = 1.23 \pm 0.21 \pi \text{ rad} \quad (S = 1.3)$$

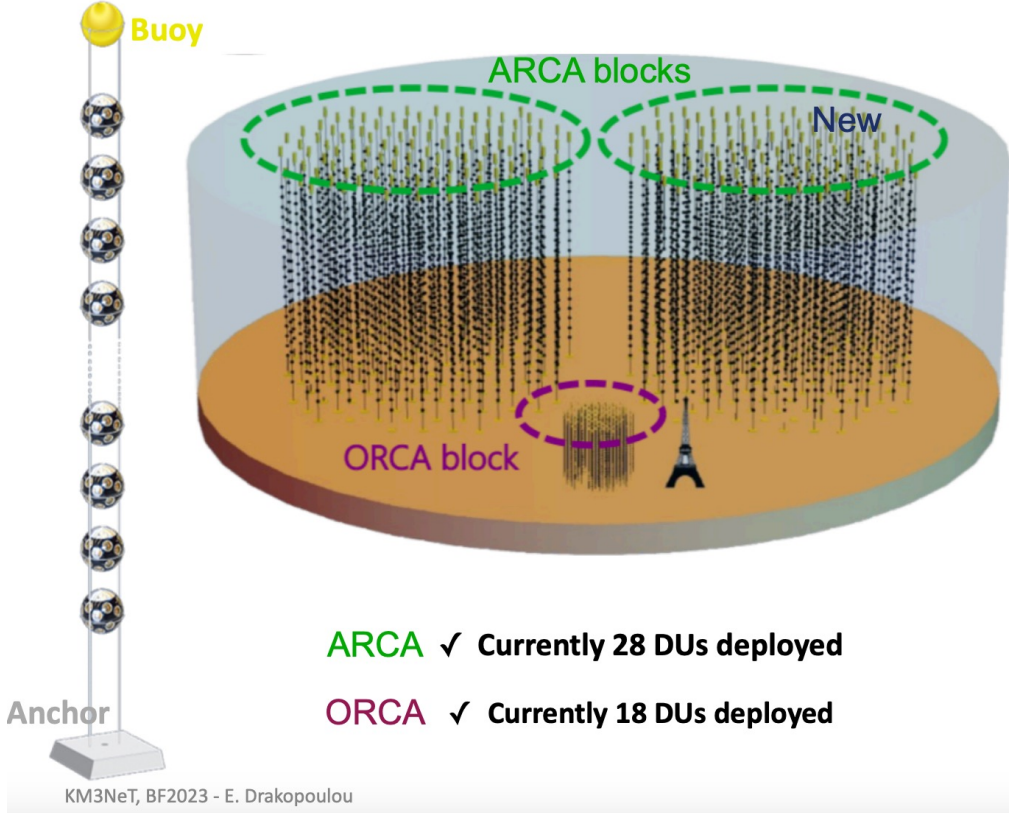
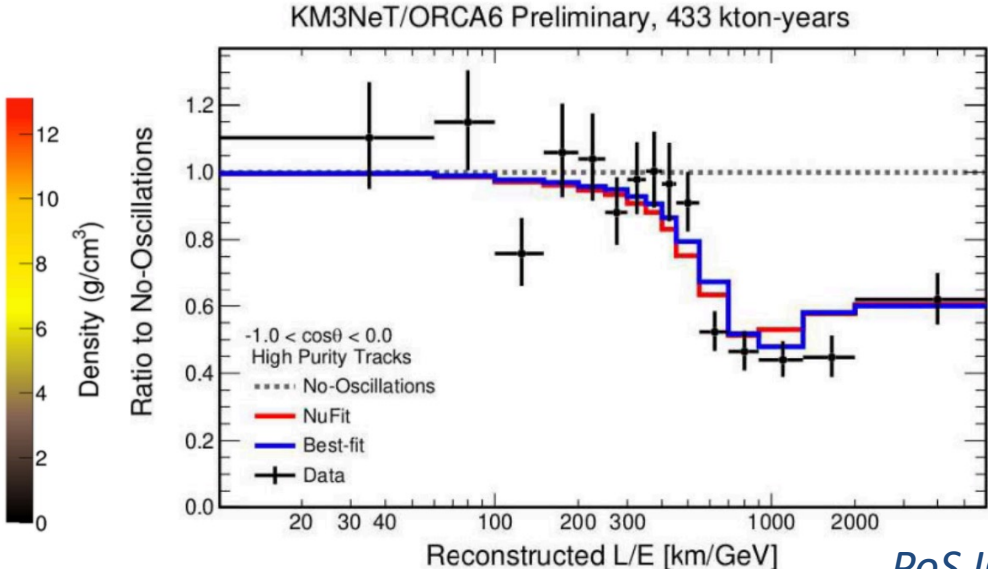
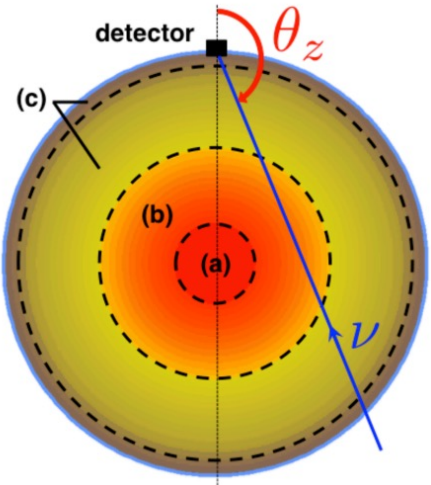
Neutrino Masses and Mixing: Future Experiments

KM3Net-ORCA:

- 7 Mton water Cherenkov detector in the Mediterranean Sea
- **ORCA**: Dense distribution of photodetectors for ~GeV threshold
- Atmospheric neutrinos with few-GeV energy particularly sensitive to neutrino mass ordering due to matter-enhanced oscillation in earth.

Status:

- 18 (of 155) DUs installed, and initial measurement of atmospheric neutrino disappearance released



ARCA ✓ Currently 28 DUs deployed
ORCA ✓ Currently 18 DUs deployed

PoS ICRC2023 (2023) 996

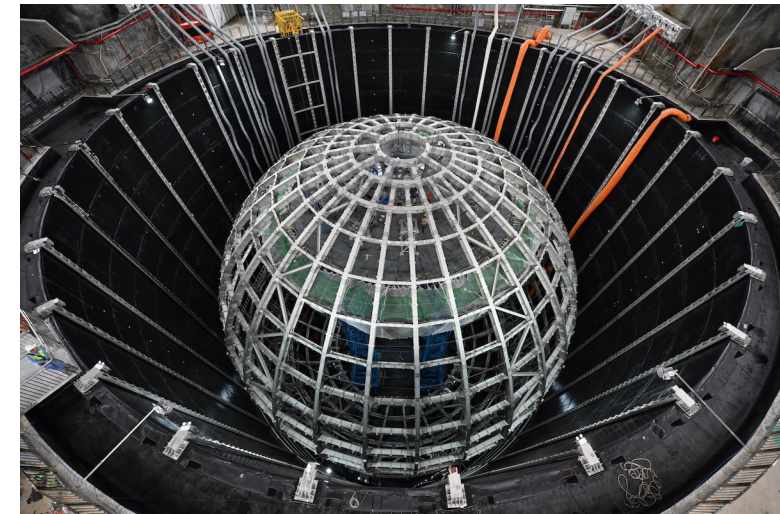
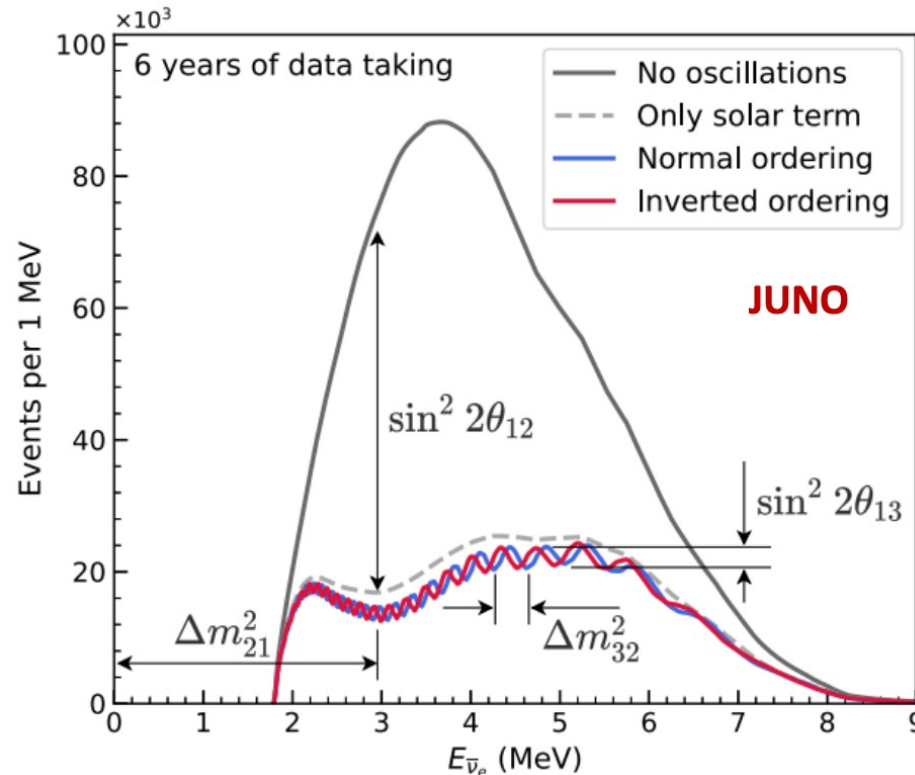
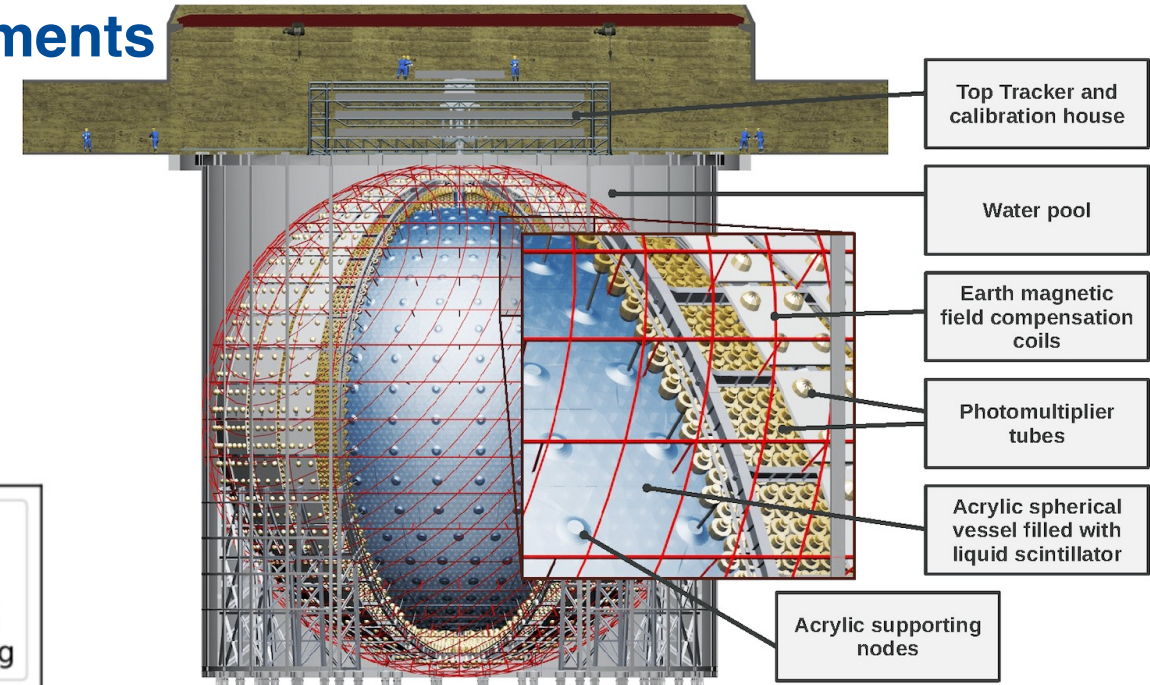
Neutrino Masses and Mixing: Future Experiments

JUNO:

- Measure reactor electron antineutrino disappearance at 52.5 km
- Detect via inverse-beta decay in 20 kt liquid scintillator detector

Status:

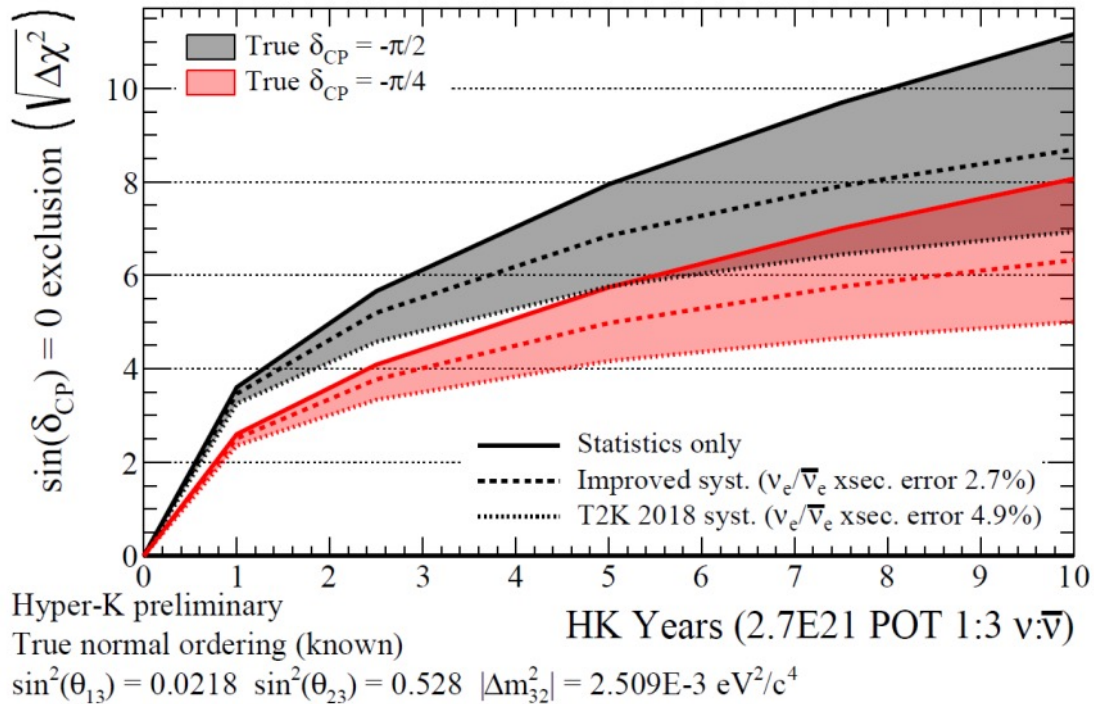
- Under construction, aiming for first data in 2024



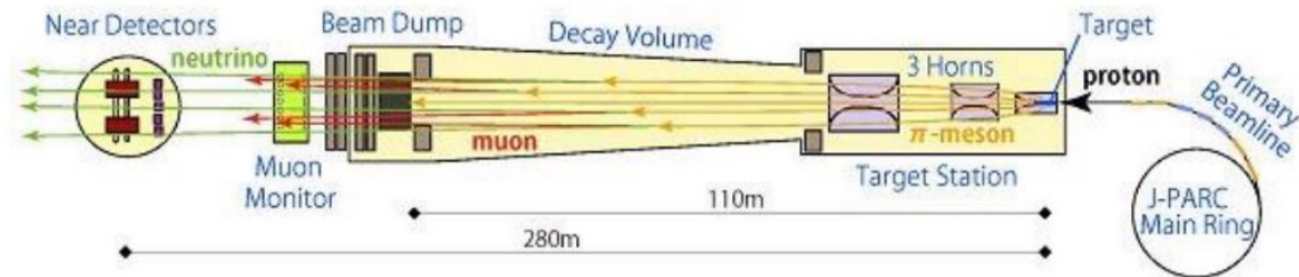
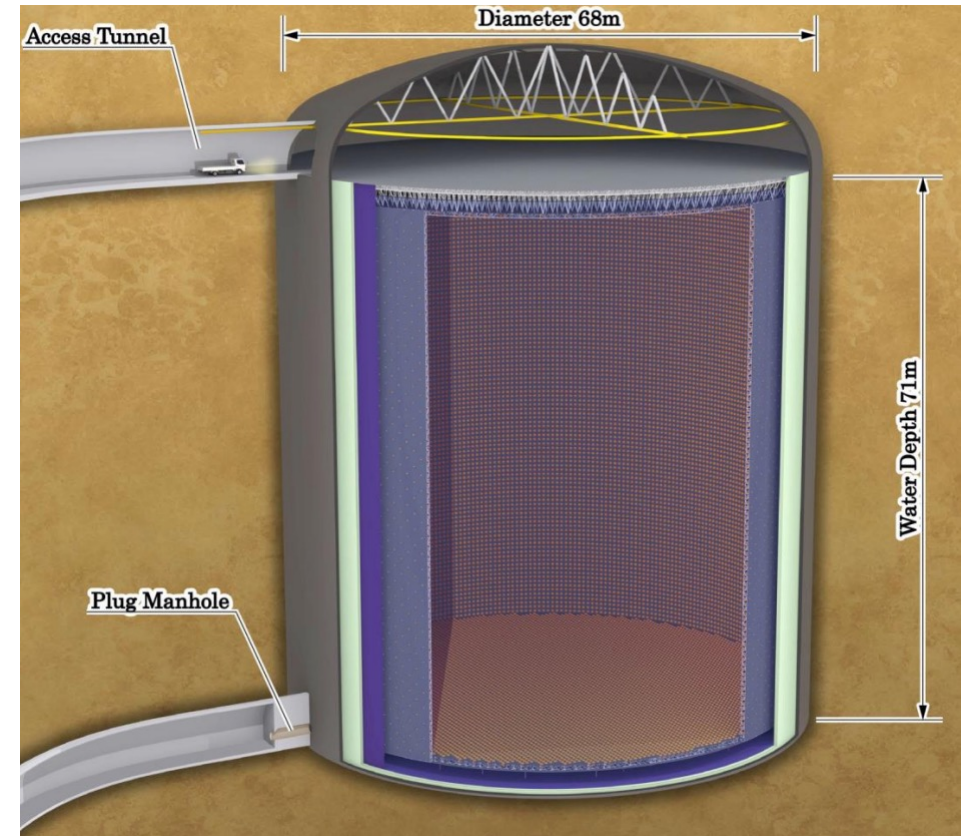
Neutrino Masses and Mixing: Future Experiments

Hyper-Kamiokande:

- 258 kton Water Cherenkov detector
- Upgraded off-axis neutrino beam from J-PARC, ~ 0.6 GeV, 295 km away
- Status: Excavation in progress, aiming for start of operations in 2027



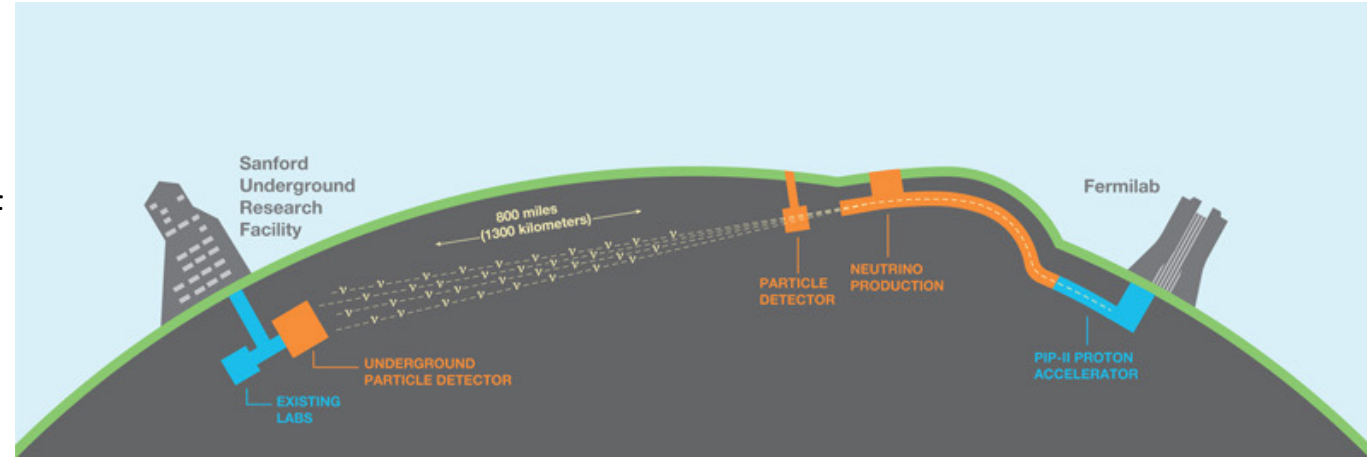
N. McCauley, NuFact 2023



Neutrino Masses and Mixing: Future Experiments

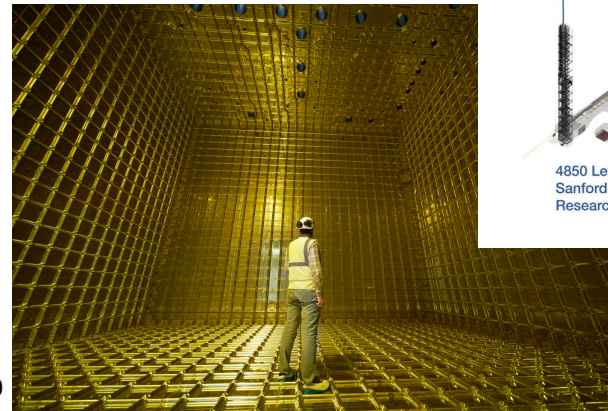
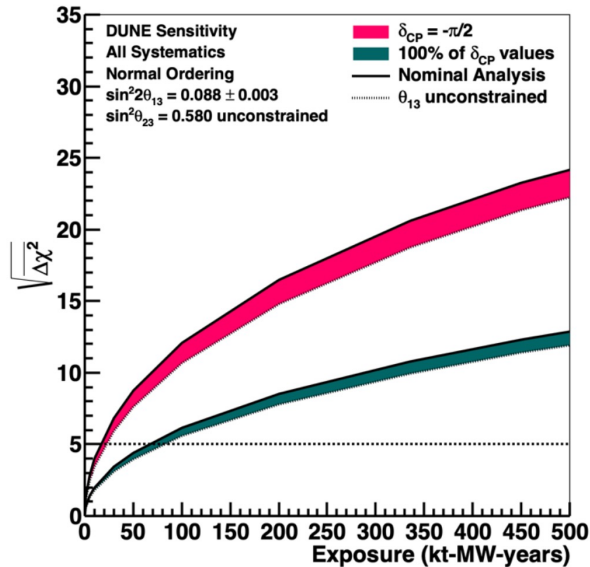
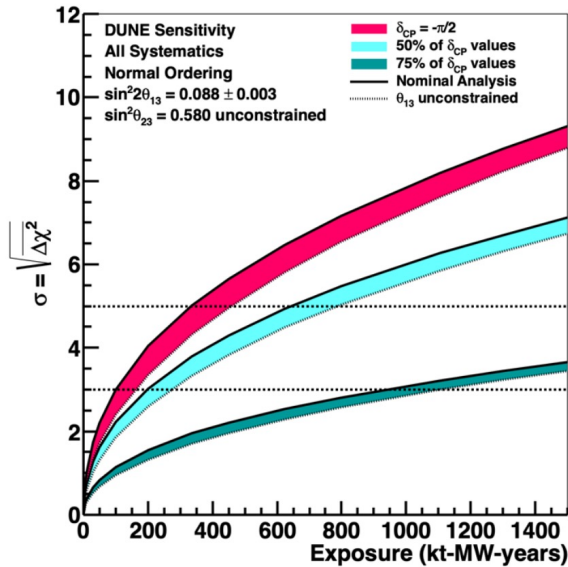
Deep Underground Neutrino Experiment (DUNE)

- Multiple 10 kton Liquid Argon Time-Projection Chambers at SURF
- New on-axis neutrino beam from FNAL, ~1-5 GeV, 1300 km
- Physics goals:
 - Mass ordering, CP-Violation, Precision Oscillation
 - Supernova Neutrinos, Nucleon Decay, BSM Searches
- Status: Far site excavation nearly complete, First 10 kton TPC production starting now

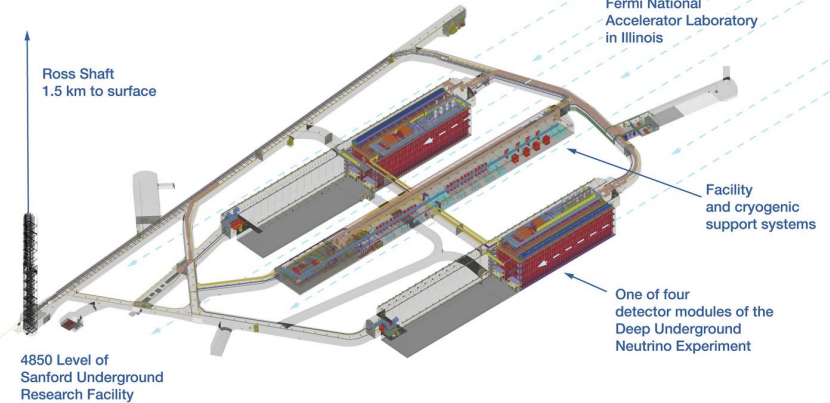


CP Violation Sensitivity

Mass Ordering Sensitivity



Long-Baseline Neutrino Facility South Dakota Site



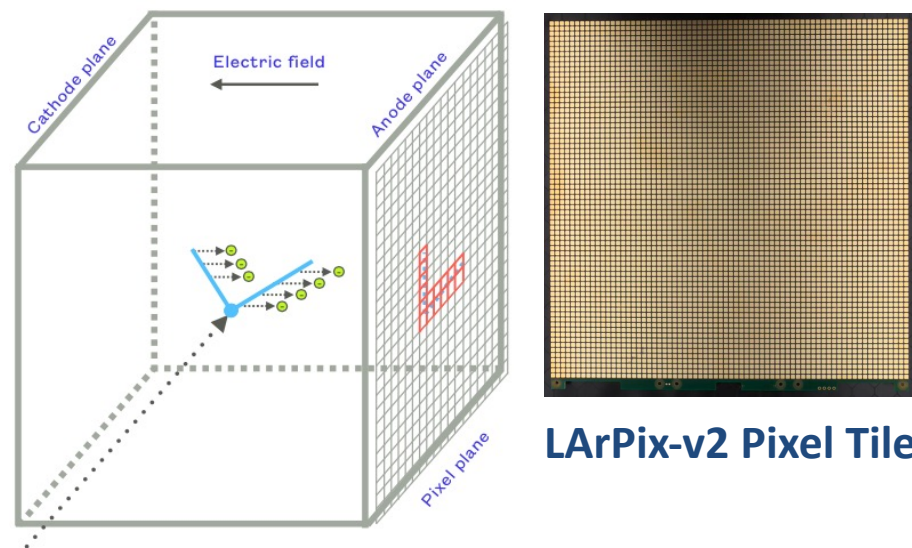
Interlude: Development of 3D LArTPC

3D Pixel LArTPC (LArPix):

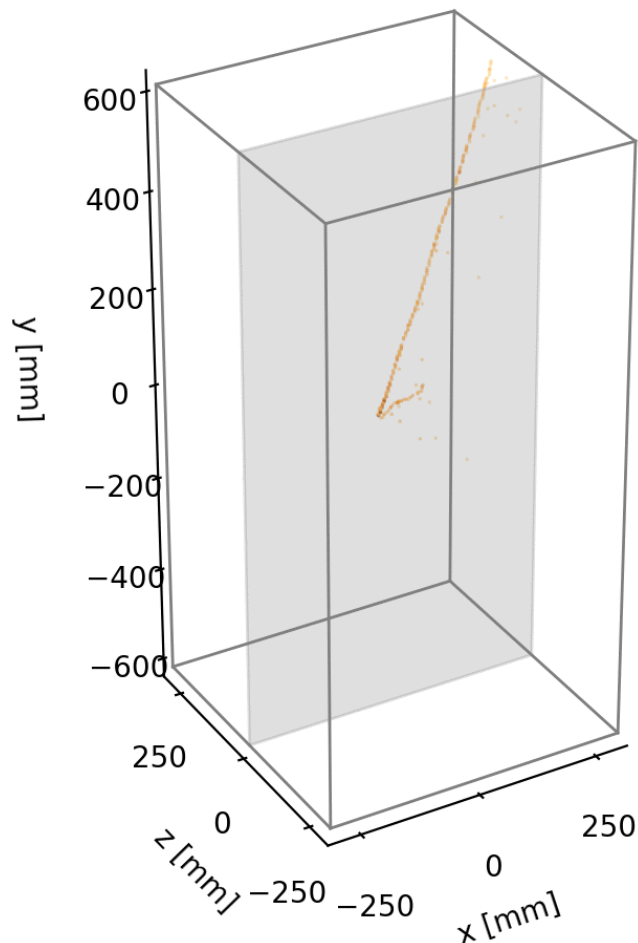
- True 3D imaging
- Continuous readout, ~100% uptime
- Intrinsically sparse data, low data volume

Recent Progress:

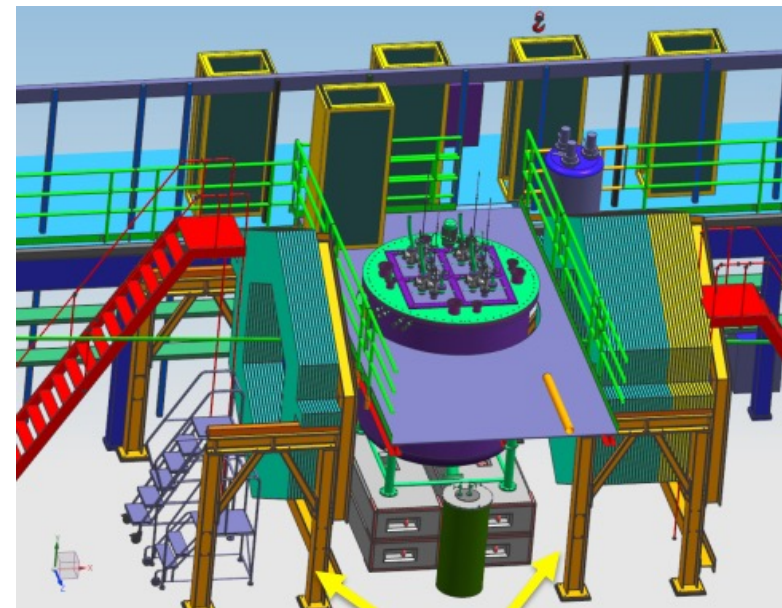
- Operation of four ton-scale 100k-pixel prototype LArTPCs



LArPix-v2 Pixel Tile



Actual raw cosmic ray data imaged in ton-scale prototype pixel TPC



Next Steps:

- Currently installing a 300k-pixel LArTPC in NuMI neutrino beam
- Use to characterize neutrino-nuclear interactions in the few-GeV energy regime, relevant for DUNE program
- Serve as prototype for the DUNE Liquid Argon Near Detector (ND-LAr)

Pursuing Neutrino Mass

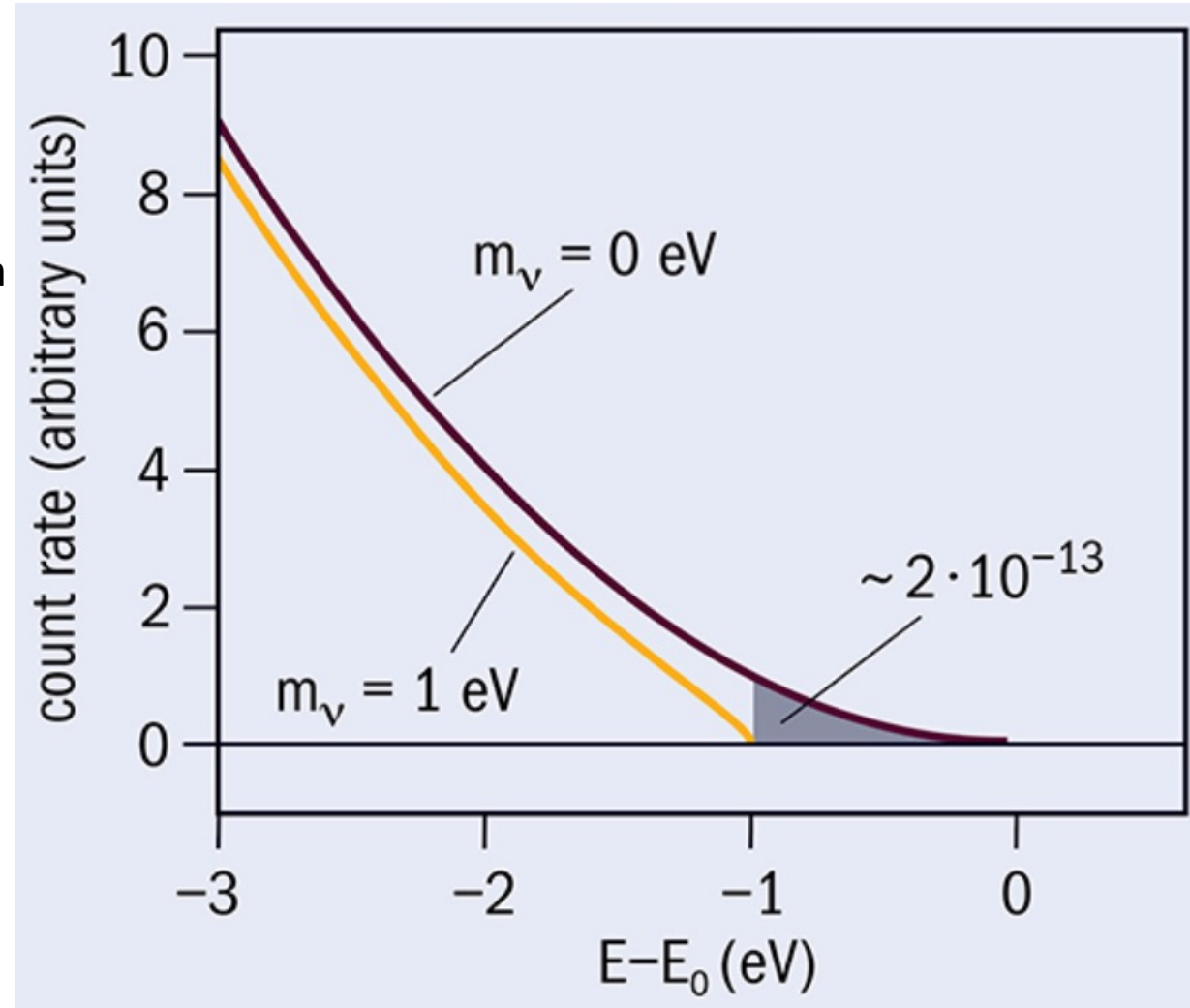
Neutrino Masses and Mixing: Direct Mass Measurements

Kinematic Measurement:

- Neutrino mass distorts the spectrum of beta-decay electrons
- Small neutrino mass only visible in very tail of decay spectrum

Effective Neutrino Mass:
$$m_\beta = \sqrt{\sum_i |U_{ei}^2| m_i^2}$$

- ^3H (beta-decay), ^{163}Ho (EC) are preferred isotopes, due to lower half-life and decay endpoint.
- So far, experiments have placed limits at the < 1 eV scale, approaching oscillation measurement mass regime.

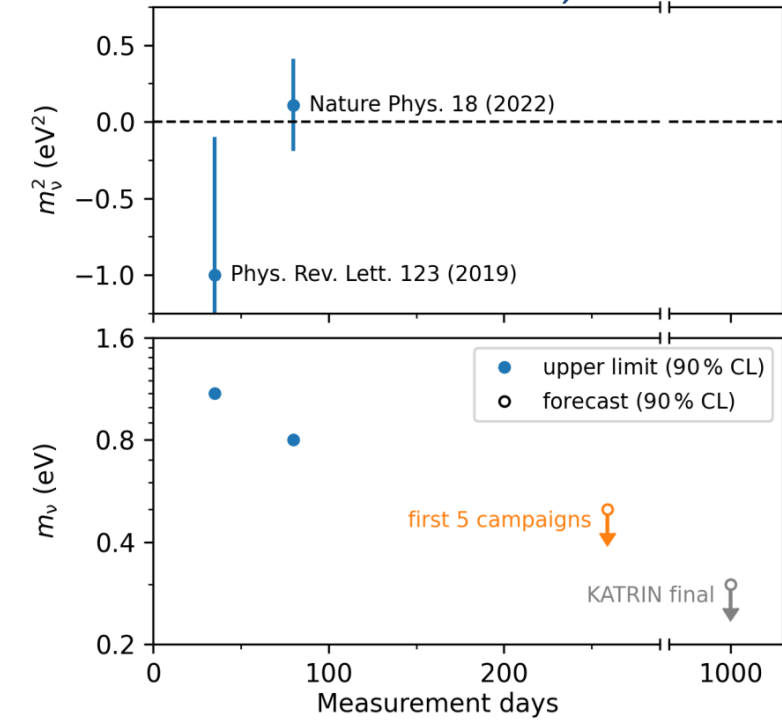
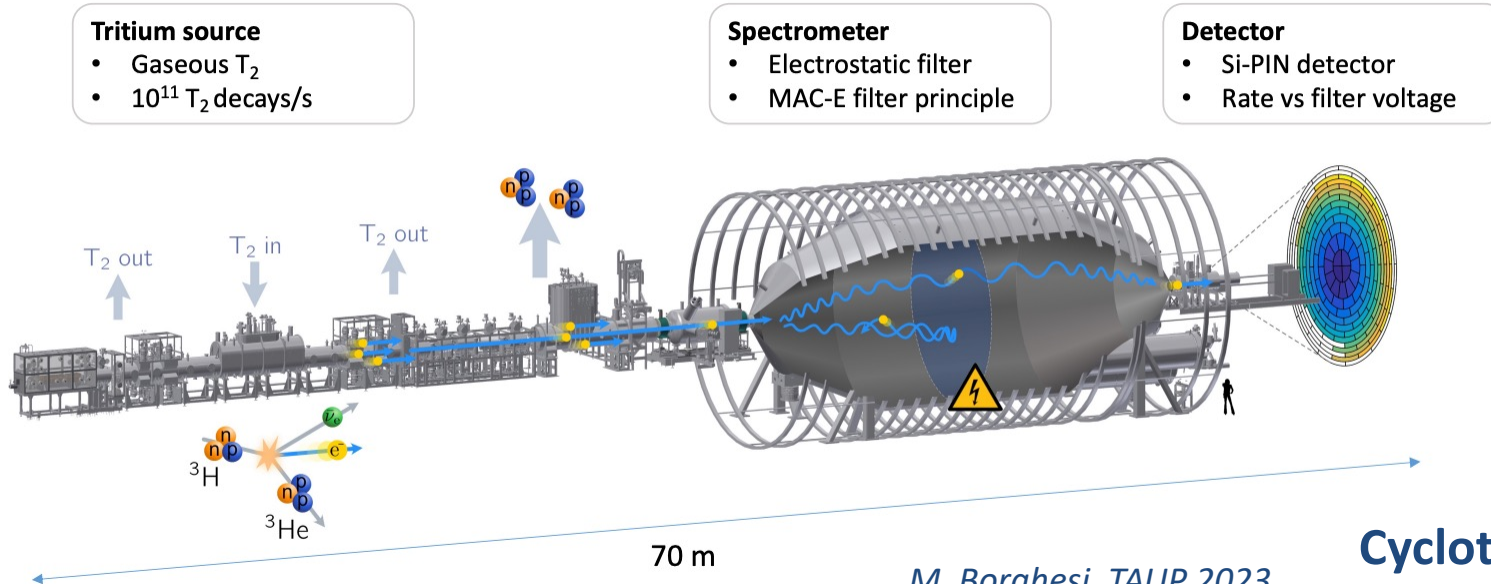


CERN Courier, Jan. 2020

Neutrino Masses and Mixing: Direct Mass Measurements

Spectrometer:

KATRIN: Electrostatic filtering spectrometer measuring ^3H decay



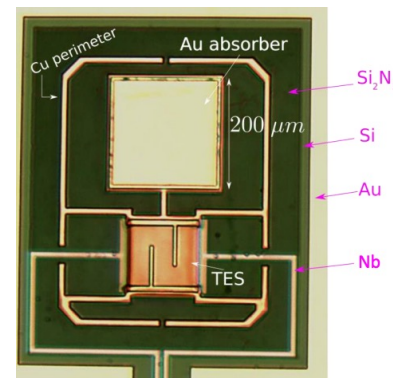
Microcalorimeters:

HOLMES: ^{163}Ho embedded in Au absorber, coupled to TES

ECHO: ^{163}Ho embedded in magnetic micro-calorimeter (MMC)

Eur. Phys. J. Special Topics 226, 1623–1694 (2017)

M. Borghesi, TAUP 2023



Cyclotron Resonance Spectrometer: *PRD 80, 051301 (2009)*

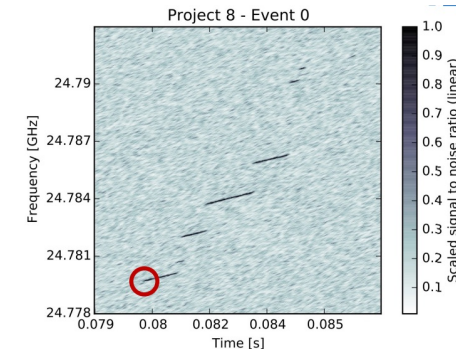
Cyclotron emission frequency sensitive to electron energy

Project-8: *Y.H. Sun, BNL Forum 2023*

^3H in atom trap resonant cavity

QTNM:

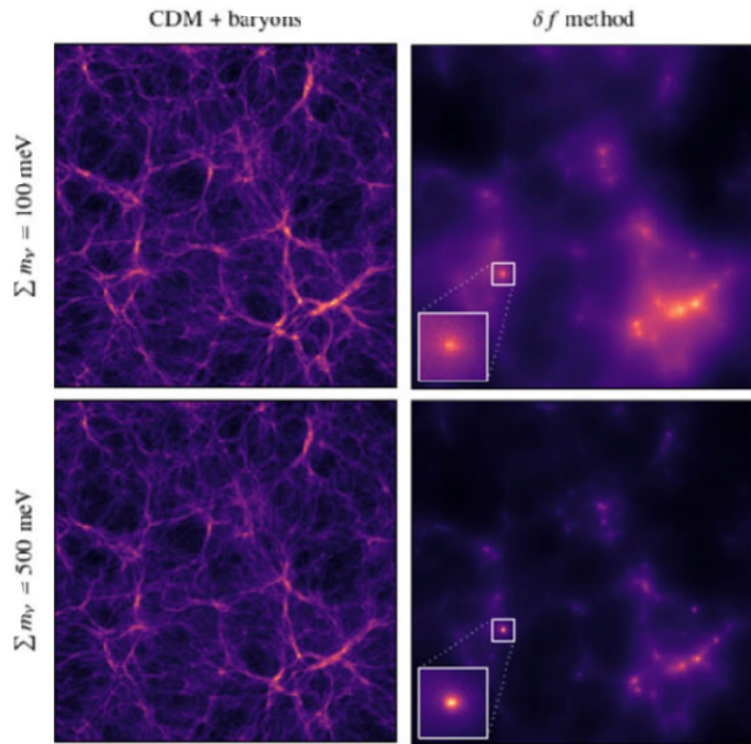
^3H storage ring



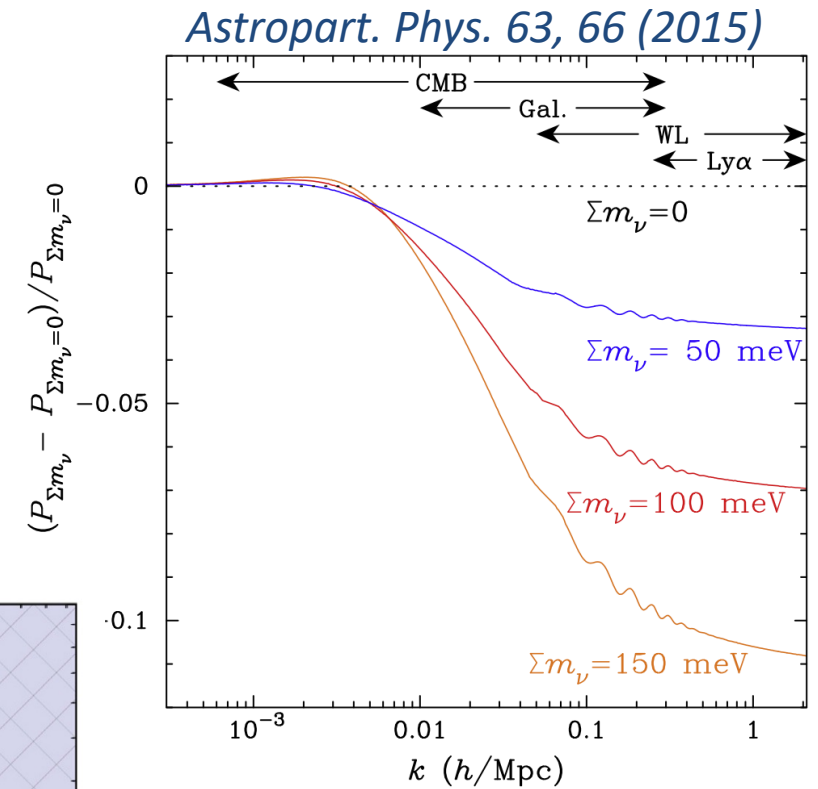
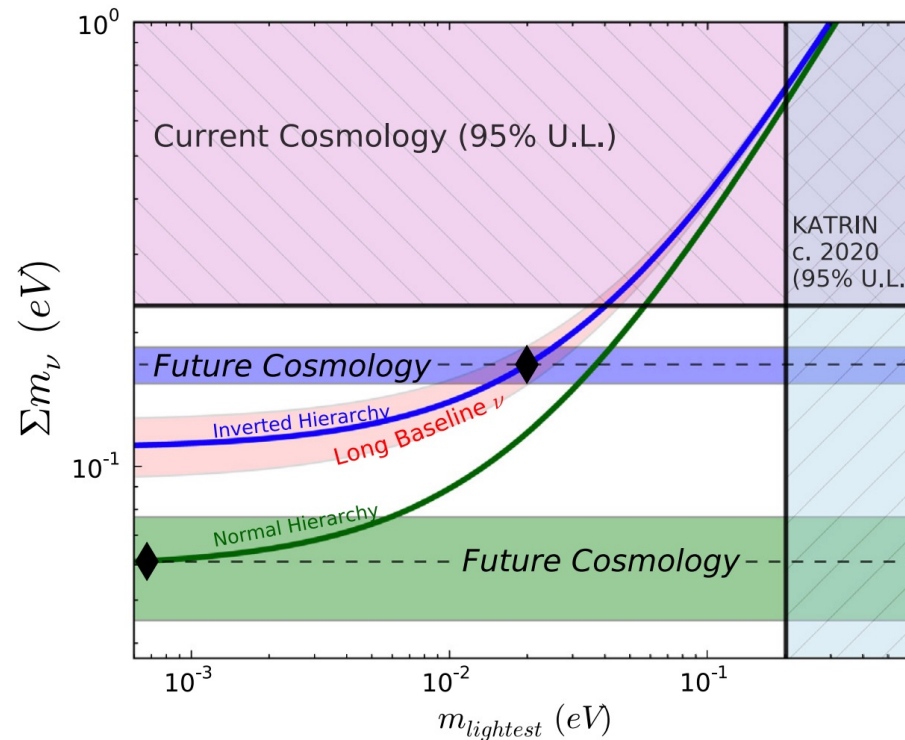
Neutrino Masses and Mixing: Cosmological Probes

Large Scale Structure:

- Free-streaming of massive neutrinos suppresses clustering in early universe
- Measurable using CMB, Galaxy surveys, Weak Lensing, and Lyman-alpha forest
- Sensitive to the sum of neutrino masses



MNRAS 507(2) 2614 (2015)



Current combined limit:

$$\text{Sum}(m_{\text{nu}}) < 0.12 \text{ eV}$$

Next-generation experiments (CMB-S4 and Spectroscopic surveys) aim for a resolution of 0.01 eV

Implications of Neutrino Masses and Mixing

Lepton flavor is not a conserved quantity.

Neutrinos have mass.

- Mass is at least 6 orders of magnitude smaller than known particles.
- Currently unknown how to correctly incorporate neutrino mass into the Standard Model.

Dirac Mass

$$- m\psi_R\psi_L - m\psi_R\psi_L$$

- **Dirac mass:**

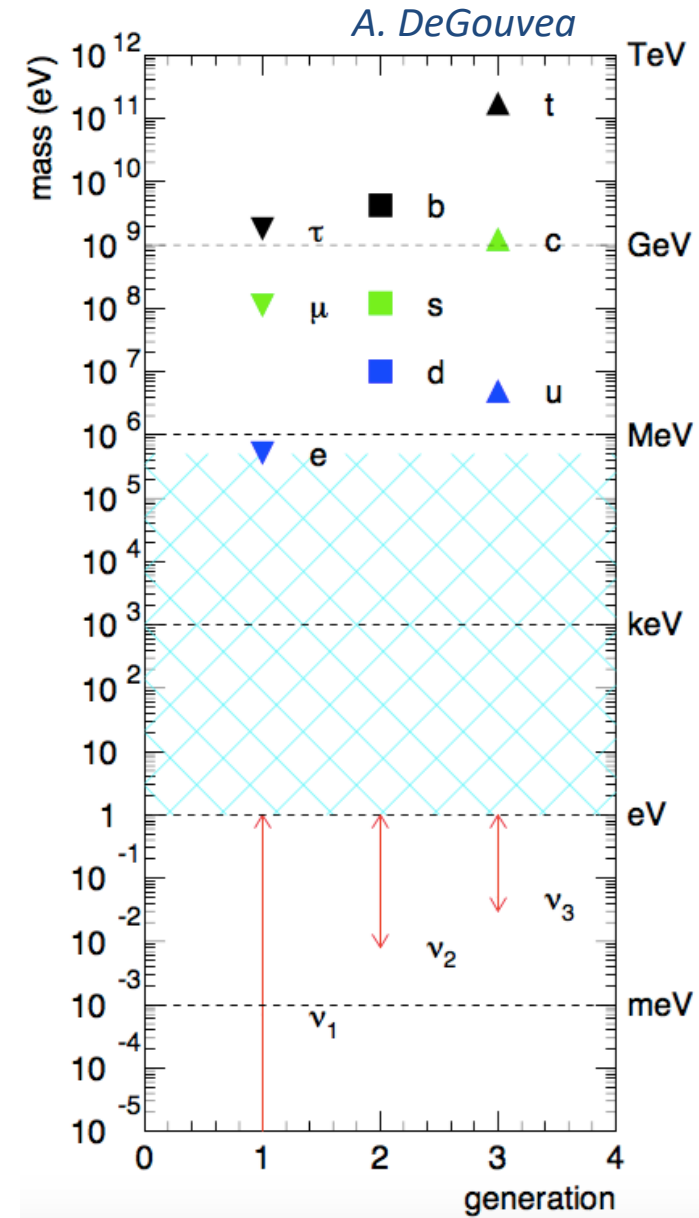
Similar to other known particles, but difficult to motivate small mass

- **Majorana mass:**

Possible for neutral neutrino, and seesaw mechanism can generate small mass, but violates lepton number conservation

Majorana Mass

$$- m_L\chi_L\chi_L - m_R\chi_R\chi_R$$



The Nature of Neutrino Mass

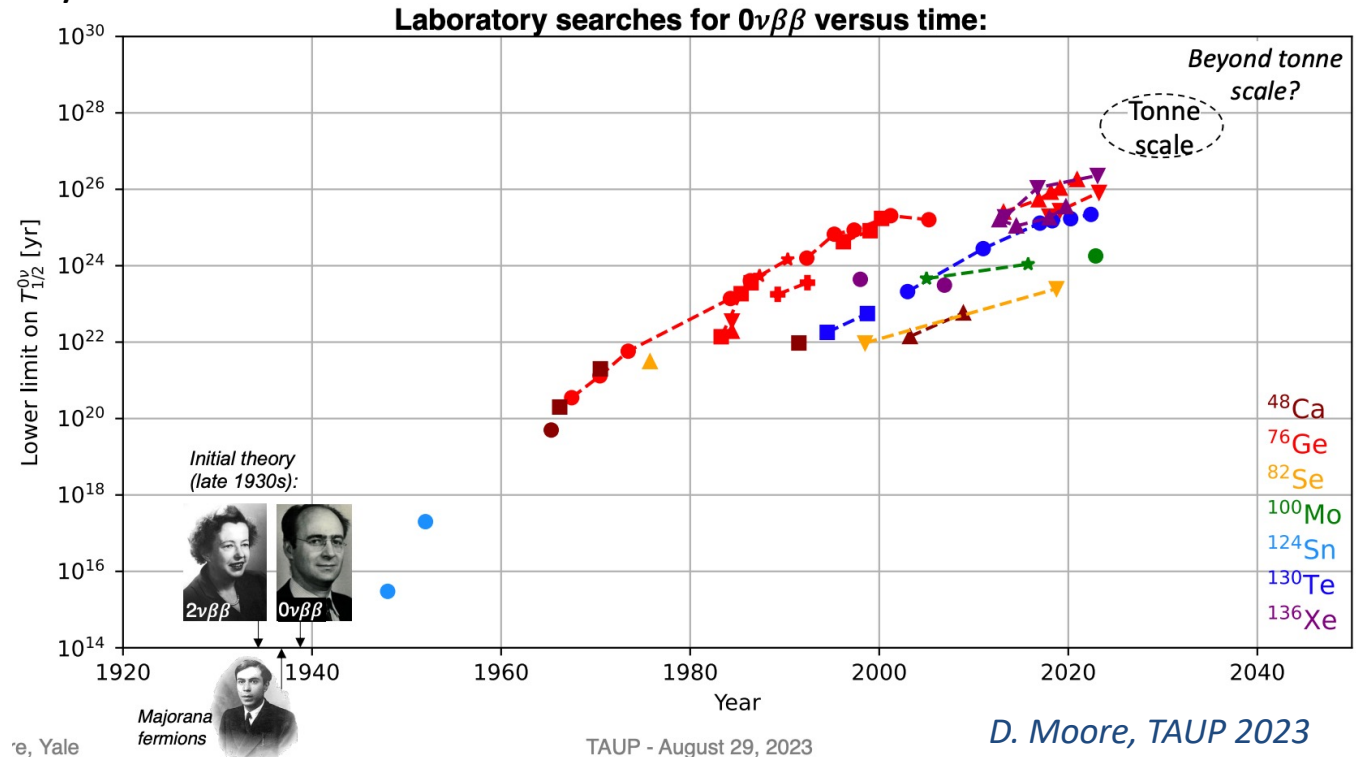
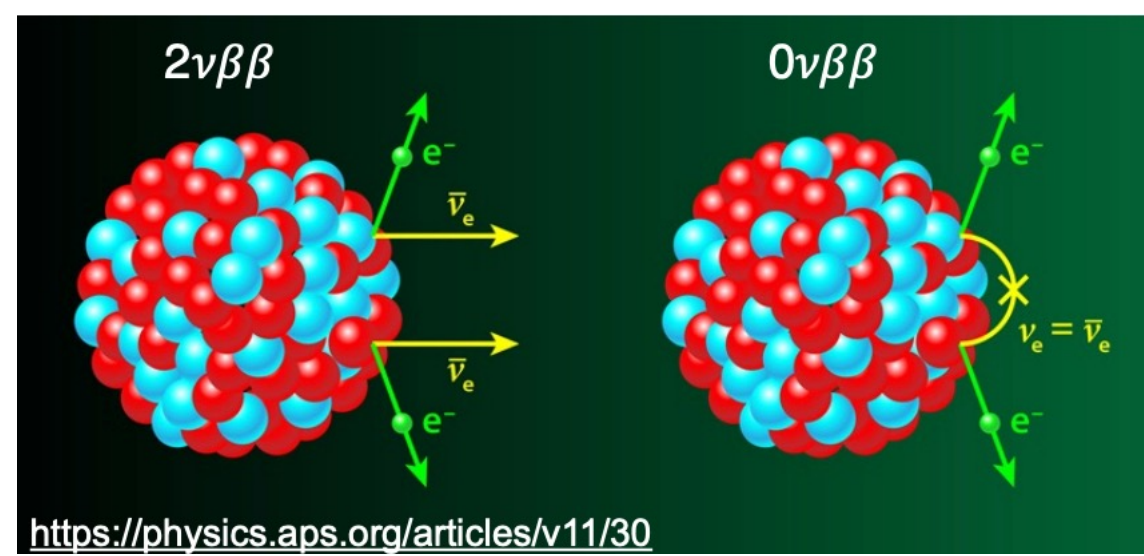
Neutrino mass and Lepton Number Violation:

- Neutrinoless double beta decay:
Currently the most sensitive probe of possible LNV
- Experiments aim to measure/limit half-life of isotope decay
- In the simplest SM extension:
e.g. Exchange of a light Majorana neutrino

$$\frac{1}{T_{1/2}^{0\nu}} = G_{0\nu} g_A^4 \mathcal{M}^2 |f|^2$$

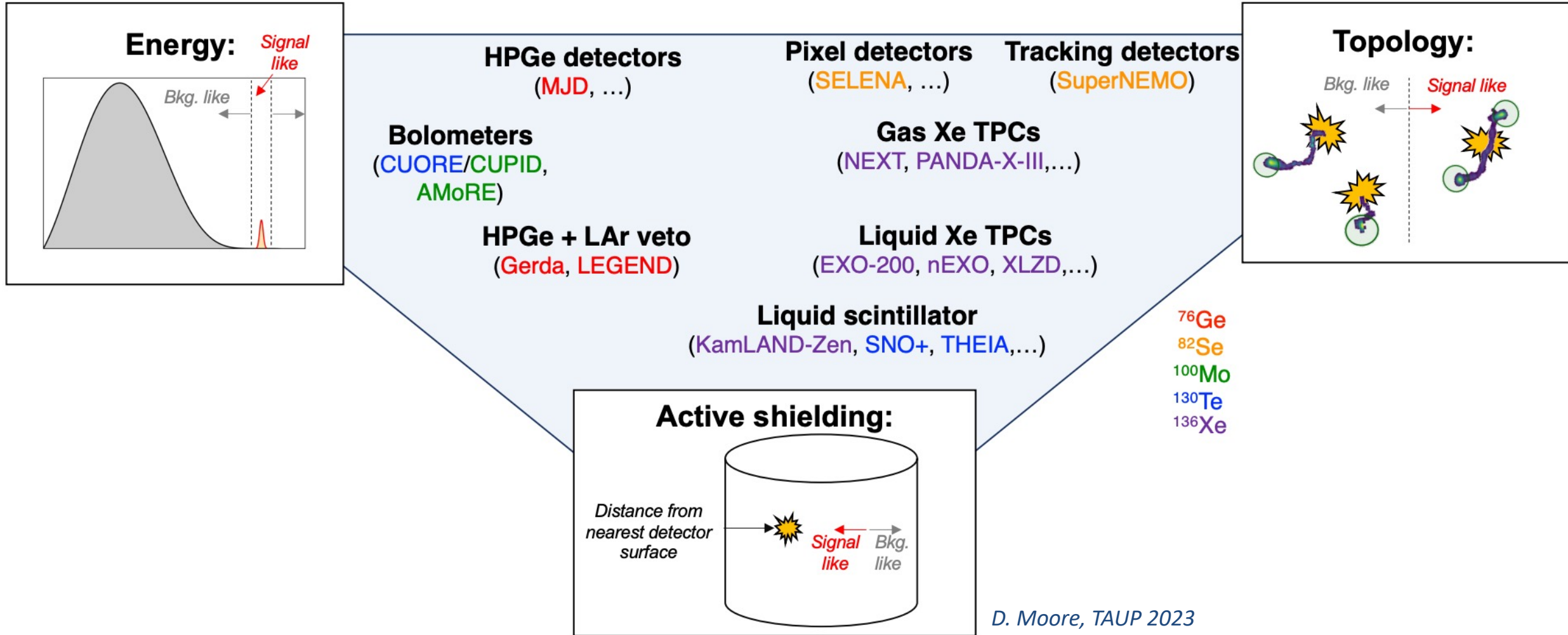
$$|f| = \frac{m_{\beta\beta}}{m_e} = \frac{|\sum_{i=1}^3 U_{ei}^2 m_i|}{m_e}$$

-> Sensitive to coherent sum of neutrino masses



The Nature of Neutrino Mass: Techniques

Rich experimental field leveraging a wide range of detection technologies.



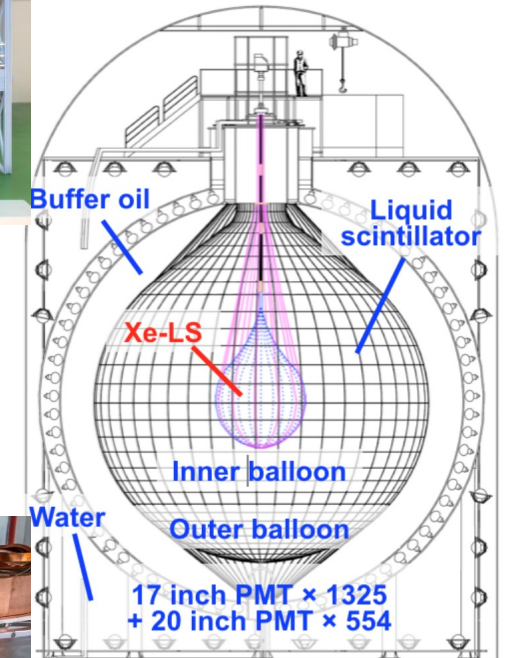
The Nature of Neutrino Mass: Current Limits



GERDA

^{76}Ge

PRL 125, 252502 (2020)



KamLAND-Zen 800

^{136}Xe

JPCS 1468 (2020) 012142

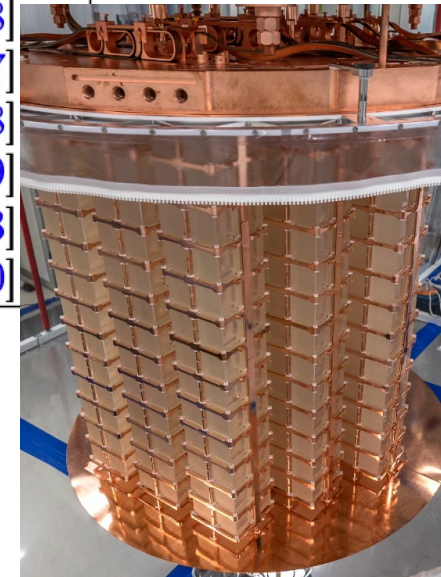
| Isotope | Technique | $T_{1/2}^{0\nu}$ | $m_{\beta\beta}$ (eV) | Year Published |
|-------------------|---|--------------------------|-----------------------|----------------|
| ^{48}Ca | CaF ₂ scint. crystals | $> 5.8 \times 10^{22}$ y | $< 3.5-22$ | 2008 [65] |
| ^{76}Ge | ^{76}Ge detectors | $> 1.8 \times 10^{26}$ y | $< 0.079-0.180$ | 2020 [12] |
| ^{82}Se | Zn ^{82}Se bolometers | $> 4.6 \times 10^{24}$ y | $< 0.263-0.545$ | 2022 [19] |
| ^{96}Zr | Thin metal foil within TPC | $> 9.2 \times 10^{21}$ y | $< 3.9 - 19.5$ | 2009 [66] |
| ^{100}Mo | Li ₂ $^{100}\text{MoO}_4$ bolometers | $> 1.8 \times 10^{24}$ y | $< 0.28-0.49$ | 2022 [18] |
| ^{116}Cd | $^{116}\text{CdWO}_4$ scint. crystals | $> 2.2 \times 10^{23}$ y | $< 1.0-1.7$ | 2018 [67] |
| ^{128}Te | TeO ₂ bolometers | $> 3.6 \times 10^{24}$ y | $< 1.5-4.0$ | 2022 [68] |
| ^{130}Te | TeO ₂ bolometers | $> 2.2 \times 10^{25}$ y | $< 0.090-0.305$ | 2022 [69] |
| ^{136}Xe | Liquid Xe scintillators | $> 2.3 \times 10^{26}$ y | $< 0.036-0.156$ | 2022 [13] |
| ^{150}Nd | Thin metal foil within TPC | $> 2 \times 10^{22}$ y | 1.6-5.3 | 2016 [70] |

arXiv:2212.11099

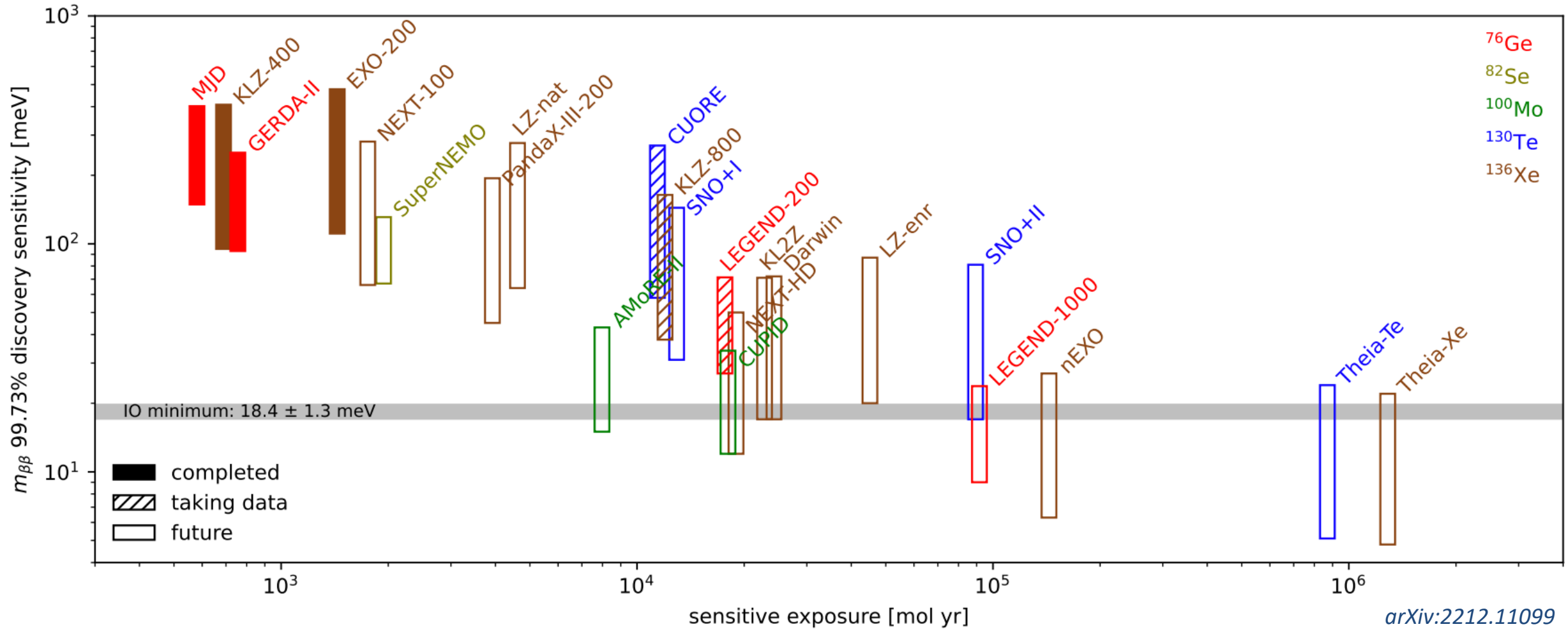
CUORE

^{130}Te

Nature 604 53 (2022)



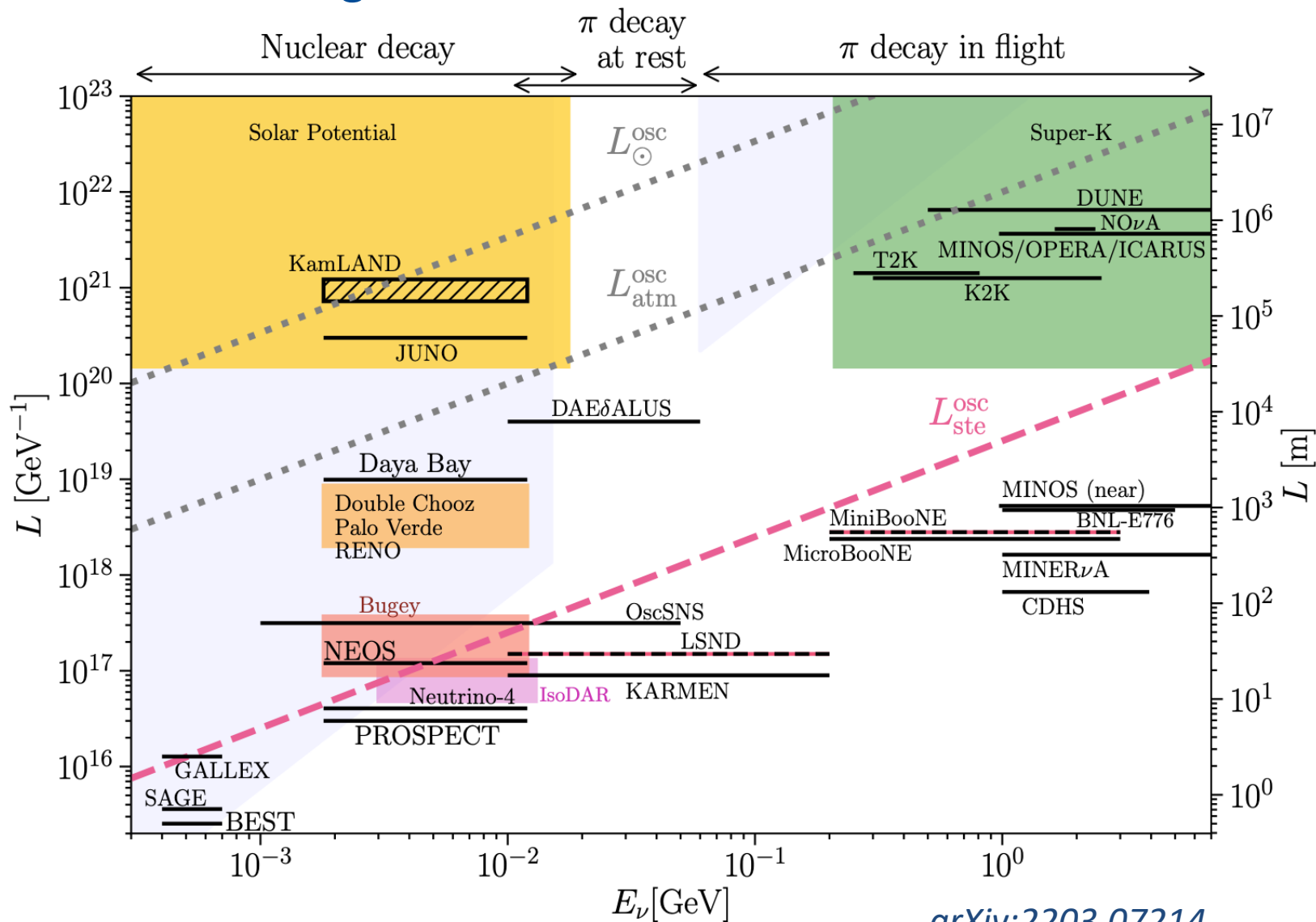
The Nature of Neutrino Mass: Current and Future Reach



arXiv:2212.11099

Beyond Three-flavor Mixing

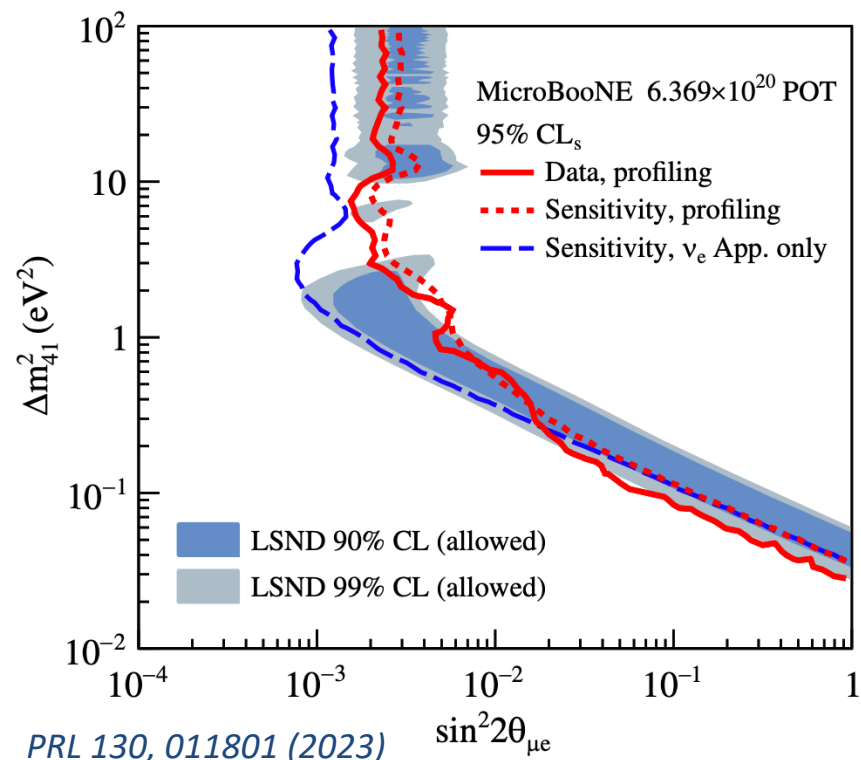
Beyond Three Flavor Mixing



Beyond Three Flavors: Accelerator and Radiochemical Experiments

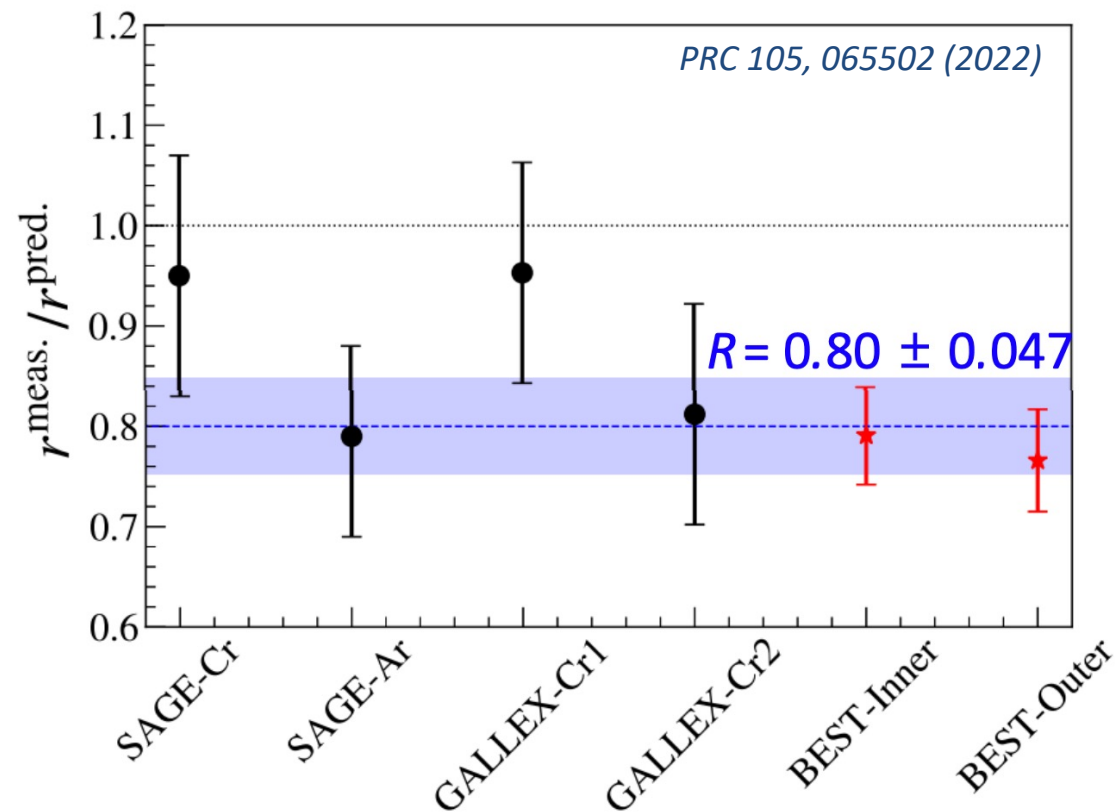
Short-baseline Accelerator Experiments

- LSND and MiniBooNE observed an excess of electron neutrino-like events from muon neutrino sources
- Data is in strong tension with other oscillation results
- Recent MicroBooNE measurement does not observe any excess
- Expect more results soon from the Short-Baseline Neutrino Program (SBN)



Radiochemical Experiments

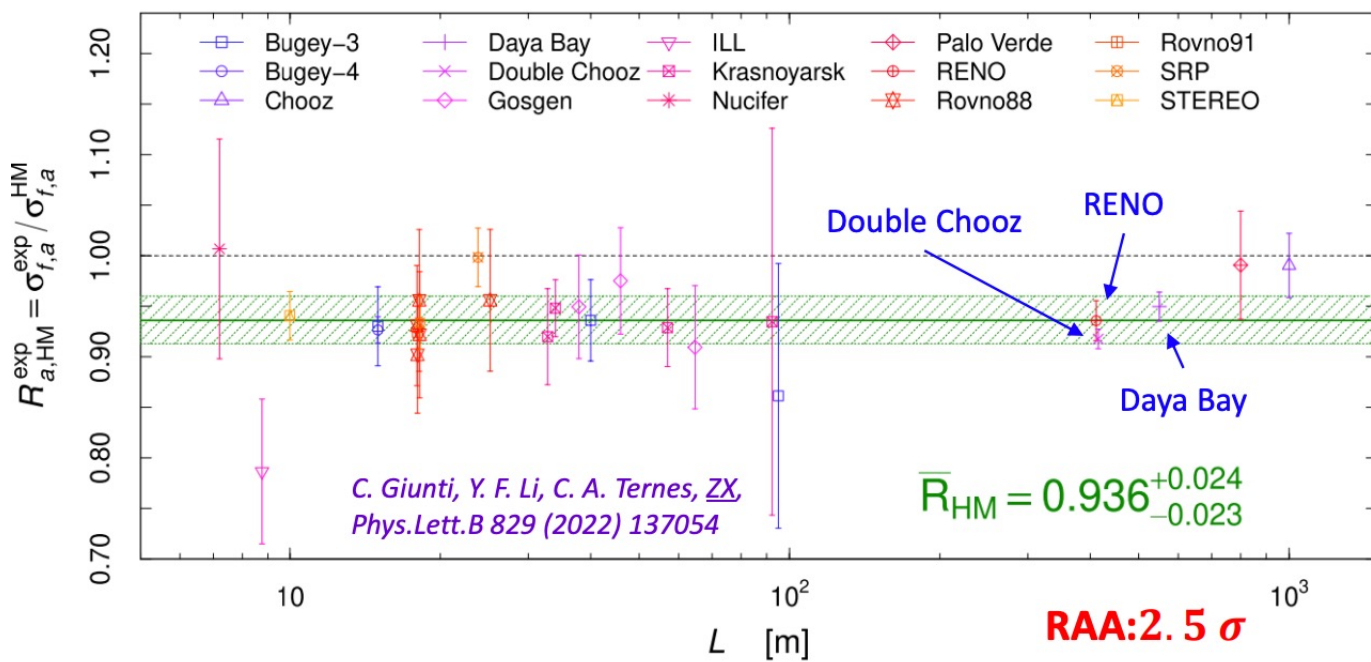
- Look for the conversion of ^{71}Ga to ^{71}Ge when exposed to an intense neutrino source (e.g. ^{51}Cr , ^{37}Ar)
- Less ^{71}Ge measured than predicted



Beyond Three Flavors: Reactor Antineutrinos

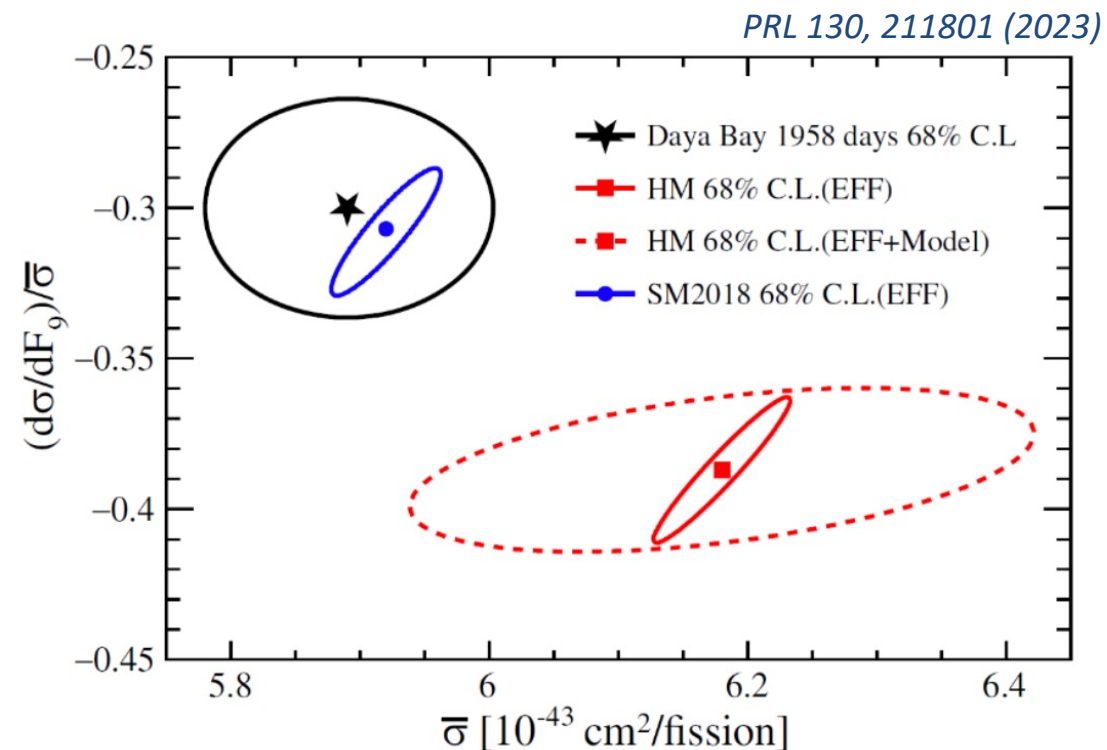
Discrepancy between observed vs predicted antineutrino rate

- Considered potential evidence for a light (eV-scale) sterile neutrino
- Consistent across multiple experiments
- Model prediction was semi-empirical; based on fission electron spectra
- Predicted energy spectrum also inconsistent with data (5 MeV shoulder)

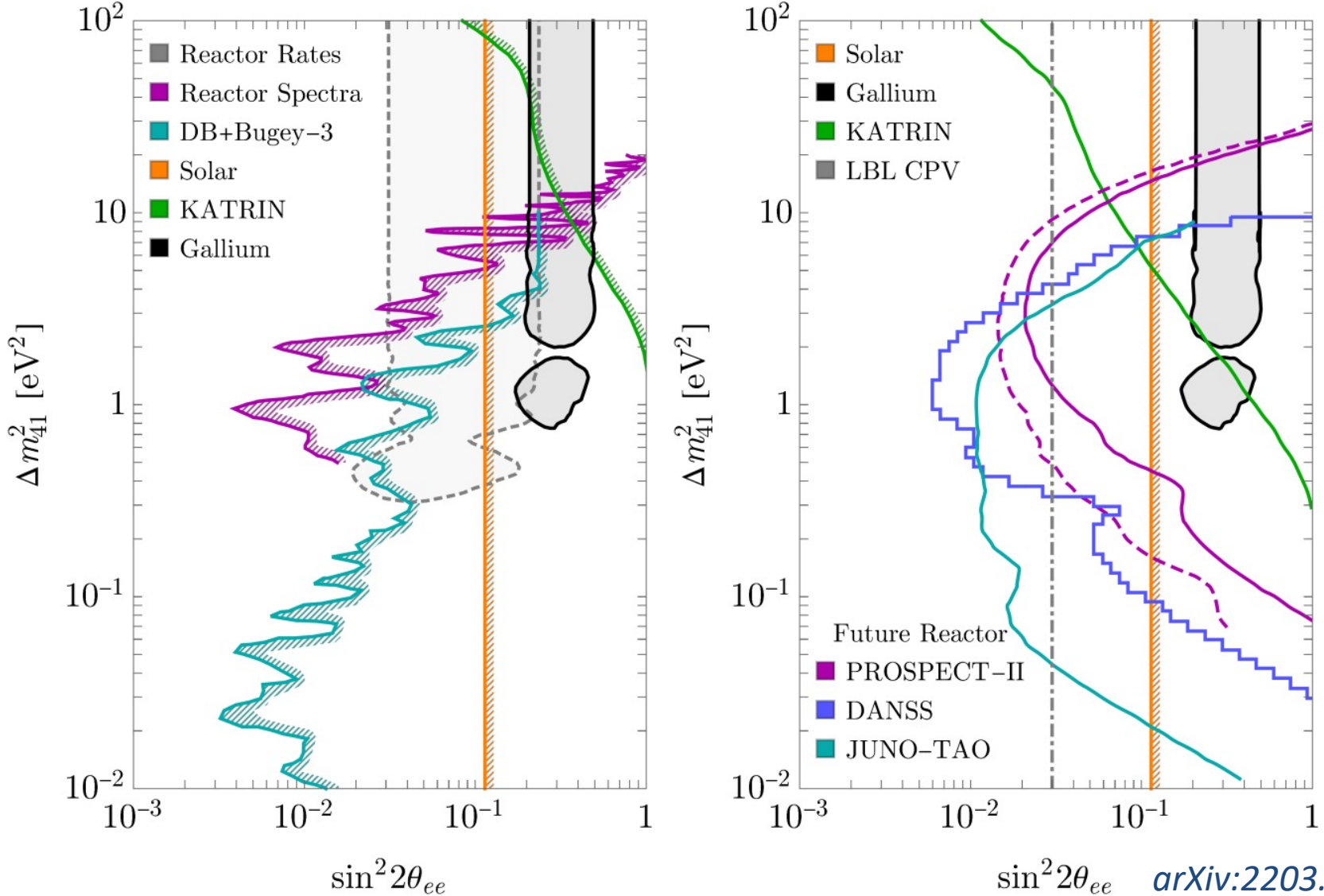


Recent Progress

- Variation of rate vs. reactor fuel composition inconsistent with semi-empirical model
- New models based on summation of nuclear decay data are consistent



Sterile Neutrinos: Current and Future Sensitivities



arXiv:2203.07214

Neutrino Experiments: Summary

Neutrino Masses and Mixing:

- Measurements sensitive to neutrino mass differences $<10^{-2} \text{ eV}^2$ present a remarkably consistent picture of three-flavor neutrino mixing
- Upcoming measurements from KM3NeT-ORCA, JUNO, Hyper-K, and DUNE will pursue the neutrino mass ordering and CP-violation, while also providing a new level of precision of the three-flavor model

Pursuing Neutrino Mass

- Kinematic and cosmic probes of neutrino mass are entering the regime suggested by oscillation measurements
- Neutrino oscillation proves neutrinos have mass, but we do not know the correct manner for inclusion in a new Standard Model
- Searches for neutrinoless double beta decay may provide direction, as well as determine if neutrinos violate lepton number conservation

Beyond Three Flavors

- Oscillation searches at $O(1) \text{ eV}^2$ present an inconsistent picture, but current and upcoming accelerator and reactor neutrino experiments are providing clearer answers.

