

Weaker Gravity and Thermal Relic Abundance

Adagio for Thermal Relics, arxiv:2308.10928

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- ① Hubble rate as the 'cosmic clock'
- ② Thermal relics in an example model
- ③ Changing the cosmic clock by changing the strength of gravity
- ④ How this changes the story for GeV scale thermal relics

- On cosmological scales, our universe is described by FLRW metric:

$$ds^2 = dt^2 - a(t)^2(dr^2 + r^2 d\Omega^2)$$

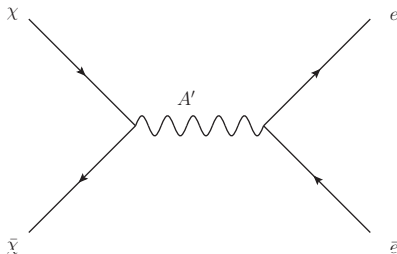
- Time dependence of the scale factor $a(t)$ describes expansion
- The Hubble rate $H(t)$ parametrizes the rate of expansion:

$$H(t) \equiv \frac{\dot{a}(t)}{a(t)}$$

- The Hubble rate is the standard to which rates are compared in cosmology
- Scalar fields behave as massless when $H(t) \gg m$, after which fluctuations redshift like matter, e.g.
 - Axions
 - Moduli
- Species fall out of equilibrium when equilibrating scattering rates are small compared to the Hubble rate, e.g.
 - Proton to neutron ratio for Big Bang nucleosynthesis
 - Thermal freeze-out of dark matter
- Models can change particle physics rates to a certain extent, but there are limits (e.g. unitarity)

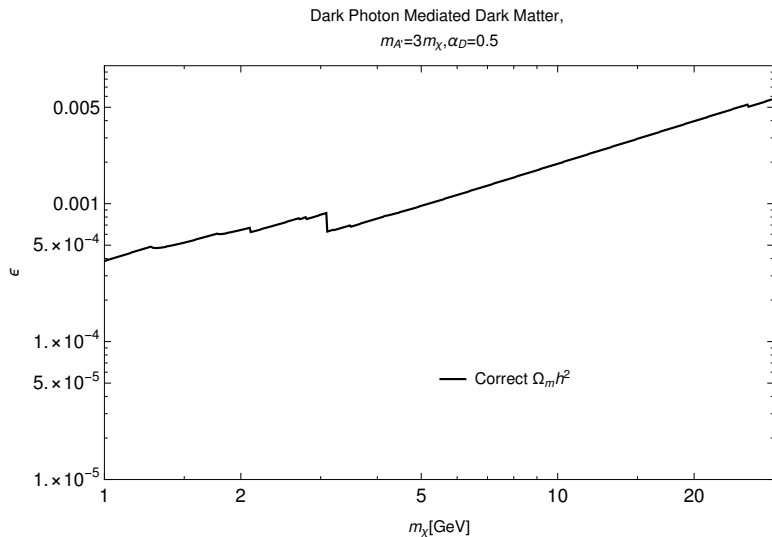
Dark Photon Mediated Dark Matter

- Simple fermionic dark matter with a dark photon mediator



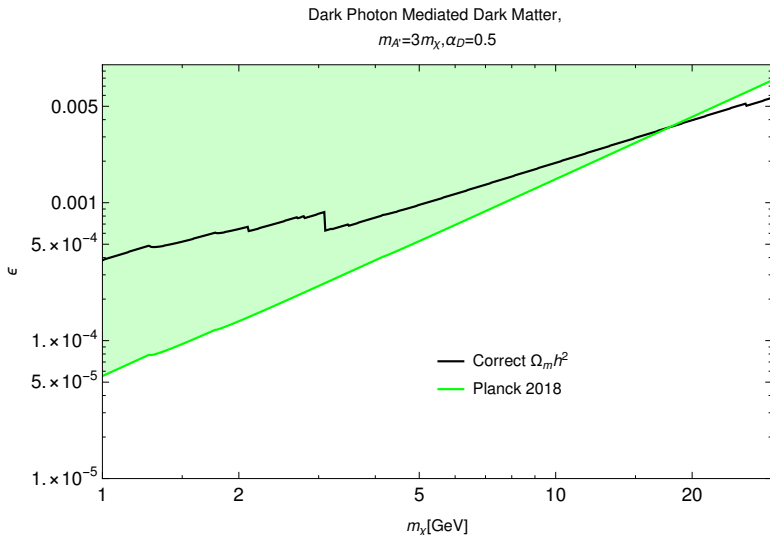
- Dark matter annihilates to SM particles via kinetic mixing ϵ between photon and dark photon
- Relic abundance set by standard thermal freezeout

Correct Relic Abundance



Confronting CMB Measurements

- Planck is sensitive to DM annihilation at later times



Changing the Hubble Rate?

- The Friedmann equations relate the Hubble rate to the energy density of the universe:

$$H^2 = \frac{8\pi G}{3}\rho$$

- There is at least one way to change H : change the particle content, thus changing the history of ρ
- We suggest a different way: change G (equivalently, M_P) with time
- Smaller G (larger M_P) leads to smaller H , which we dub 'Adagio' cosmology

M_P With Extra Dimensions

- Separation between fundamental scale of gravity M_F and the effective 4D M_P is well-known in extra dimensional models
- With e.g. n large extra dimensions of size R :

$$M_P^2 \approx M_F^{2+n} R^n$$

- In a $4 + n$ -dimensional model of gravity, the size of the extra dimensions R is dynamical
- If R was a factor of κ larger at early times, then M_P was larger by a factor of $\kappa^{n/2}$

Changing the Size of Extra Dimensions

- Size of extra dimensions is set by radion potential
- Phase transition of radion potential leads to changing the equilibrium size of the extra dimensions
- Nima et al.: when n extra dimensions grow, the macroscopic dimensions shrink (approximately Kasner solutions)

$$ds^2 = dt^2 - a_i \left(\frac{t}{t_i}\right)^{2k} d\vec{x}_3^2 - b_i \left(\frac{t}{t_i}\right)^{2l} d\vec{y}_n^2$$

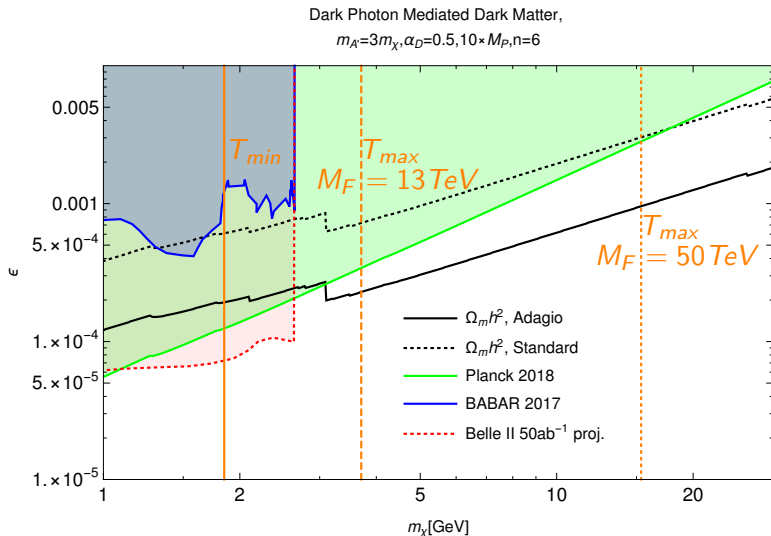
$$k = \frac{3 \pm \sqrt{3n(n+2)}}{3(n+3)}$$

$$l = \frac{n \mp \sqrt{3n(n+2)}}{n(n+3)}$$

- In reverse: when extra dimensions shrink, the macroscopic dimensions grow (+ solution for k , - solution for l)

- Must not mess up Big Bang nucleosynthesis
 - Temperature redshifts alongside the expansion of the macroscopic dimensions
 - If 6 extra dimensions shrink by a factor of κ , macroscopic dimensions grow substantially by a factor of κ^5
 - Must start Kasner phase sufficiently early so that it finishes before BBN
- If universe is reheated to too high of a temperature, things like KK graviton production can be a problem
 - Need to reheat at least above the dark matter freezeout temperature
 - Depends on absolute size of extra dimensions
 - If we want M_F to be 'close' to electroweak scale, then the extra dimensions do have to be large

Applying to the Dark Photon Mediated Model



- GeV-scale dark matter without conflicting with Planck

Potential Signals of the 'Adagio' Scenario

- For the extra dimensions:
 - Direct evidence of the phase transition may not be achievable
 - If M_F is relatively low to address the hierarchy problem: quantum black holes at colliders!
 - However: this mechanism works regardless of the size of M_F
- For this application to thermal relic dark matter:
 - Thermal relics in regions of parameter space you wouldn't naively expect
 - GeV-scale dark matter with s wave annihilation that's not ruled out by Planck - search for it

- In extra dimensional models, the size of extra dimensions is dynamical
- Changing the size of the extra dimensions means changing the effective strength of gravity in the large dimensions
- Cosmic history before BBN is largely unknown
- Thermal relic dark matter could have frozen out when M_P was different
- Slowing Hubble can also be used for other things in cosmology

Thank you!