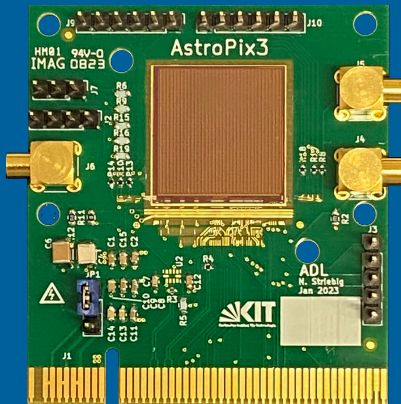
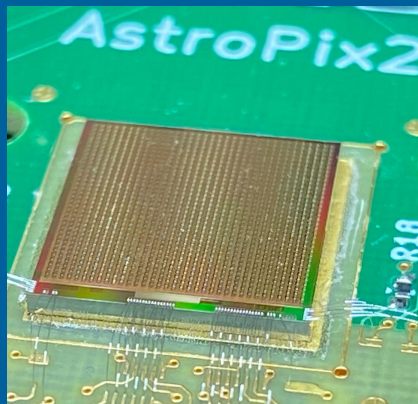


AstroPix v3



Manoj Jadhav
Argonne National Laboratory

Jan 10, 2024

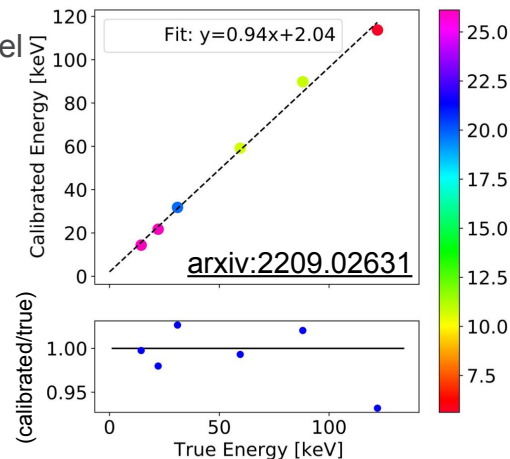
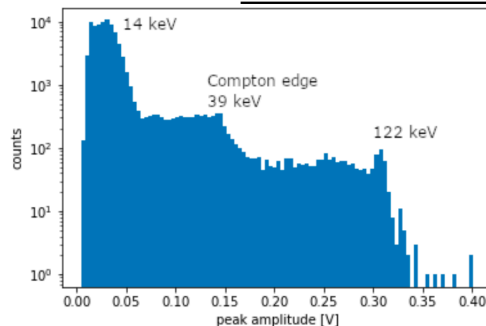
AstroPix Developments

Amanda Steinhebel

[arxiv:2302.00101](https://arxiv.org/abs/2302.00101)

AstroPix v1 - 0.45 cm × 0.45 cm chip

- 18 × 18 pixel matrix, 175 μm pitch
- Tested for analog output

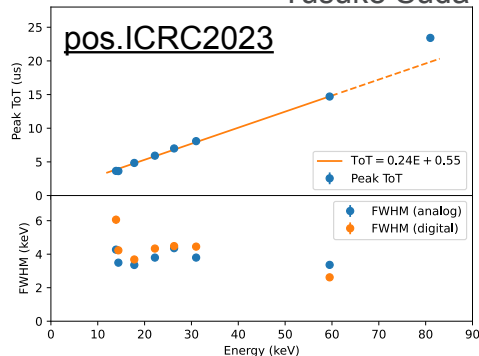


Energy Calibration using analog data

AstroPix v2 - 1 × 1 cm² chip (MPW)

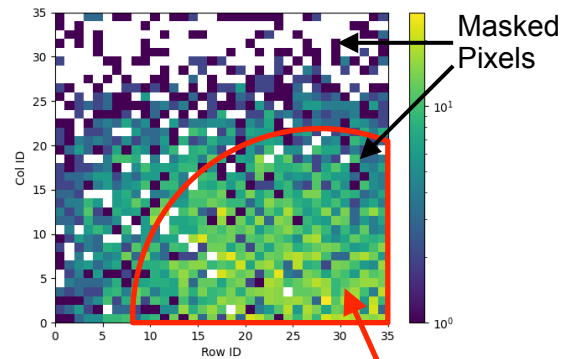
- 35 × 35 pixel matrix, 250 μm pitch
- Row/Column hit identification
- Analog + digital output

Yusuke Suda



Energy Calibration using digital data

⁹⁰Sr Source - bench test

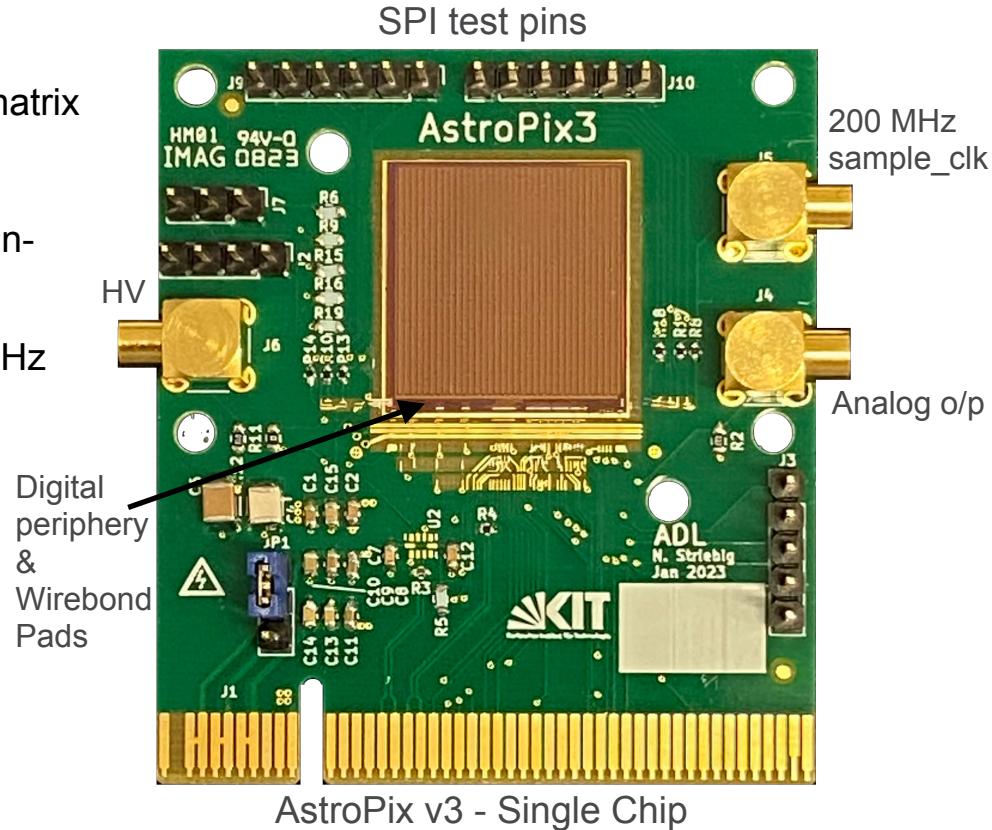


Source spot

AstroPix v3 Specifications

AstroPix v3

- $2 \times 2 \text{ cm}^2$ full-size chip with 35×35 pixel matrix
- $300 \mu\text{m}$ pixel size, $500 \mu\text{m}$ pixel pitch
- Increased spacing between the outer pixel n-well and the chip edge
- Timestamp clock 2.5MHz, ToT clock 200 MHz
- 10-byte data frame per hit
- Chip-generated injection signal
- The first 3 columns are implemented with PMOS amplifier



AstroPix v3 Measurements

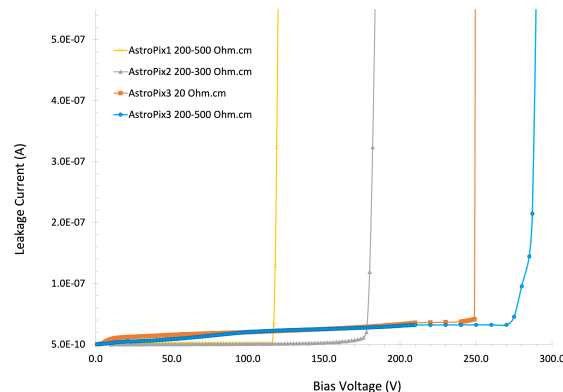
Full-size chip (ongoing testing)

- Sensor characterization
 - IV/CV measurements
 - TCT measurements
- Test bench validation and optimization
- FTBF testbeam performance studies
- Active and passive irradiation $\sim 10^{15} n_{\text{equivalent}}/\text{cm}^2$
- **Quad-chip readout** (under testing) for NASA's hosted payload mission (A-Step) - **(Taylor's talk)**
- Integration with Pb/SciFi **(Henry's talk)**

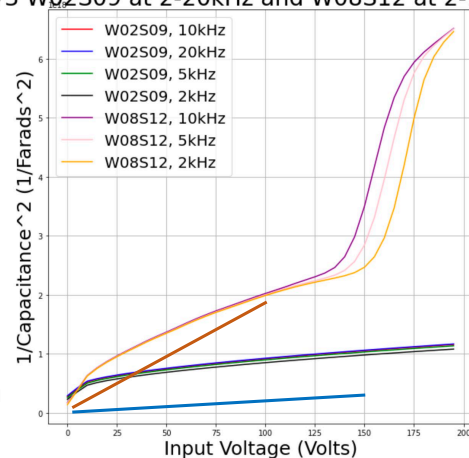
W02 - 20 Ohm.cm
W08 - 200-500 Ohm.cm

AstroPix

–IV measurements for different versions–



APXV3 W02S09 at 2-20kHz and W08S12 at 2-10kHz CV

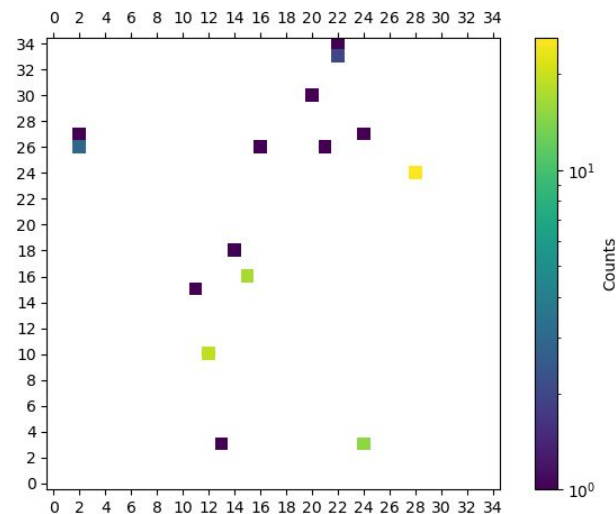
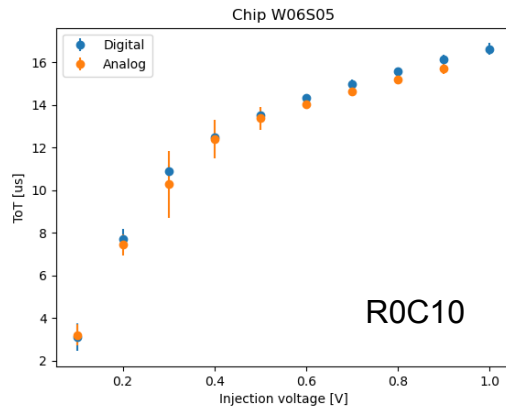
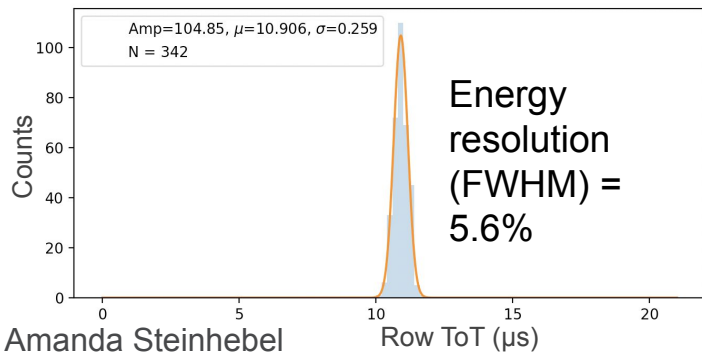


UCSC group

AstroPix v3 Results

Test Bench Measurements

- Injection voltage scan shows that the analog and digital ToT agree well
- The energy resolution of 5.6% is measured using an injected pulse
- Noise scan shows <1% of noisy pixels
- resolves issue with noisy pixels observed with v2
- Noisy pixels can be masked by disabling the comparator



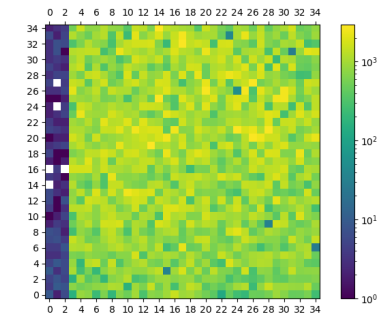
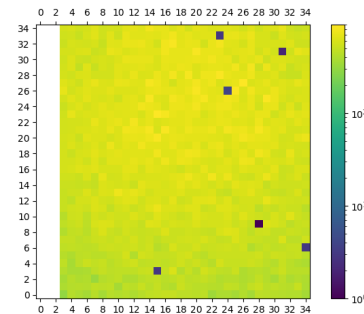
AstroPix v3 Results

Source Scan Measurements

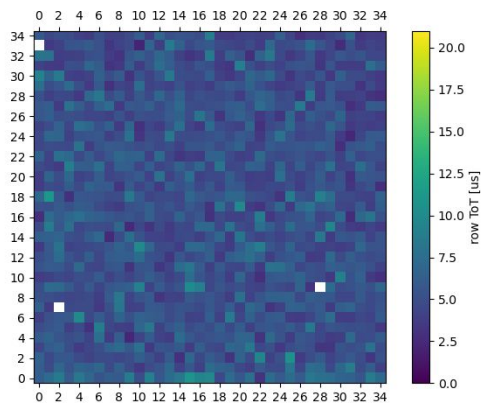
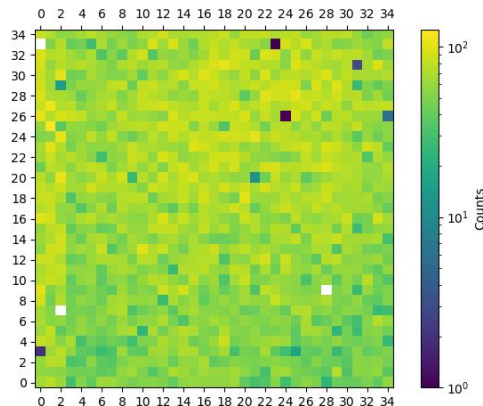
- Ba133, 30min, ~8 nCi
- Am241, 10min, ~106 nCi
- 200 mV threshold, $\Delta(TS) \leq 1$
- First 3 cols (PMOS) disabled
- Ba133, 5min, ~8 nCi, enabling PMOS amplifier col
- higher hit rate in PMOS cols

Ba133 - 93.6% of hits paired

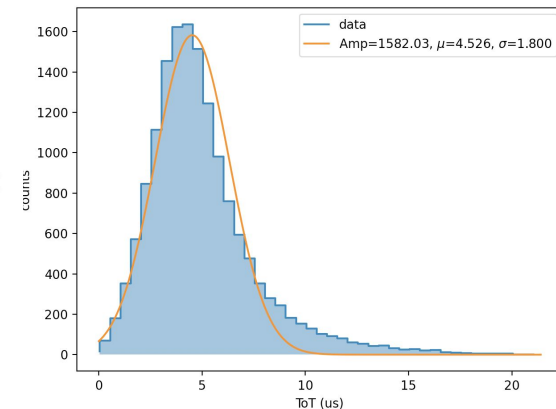
Am241 - 82.1% of hits paired



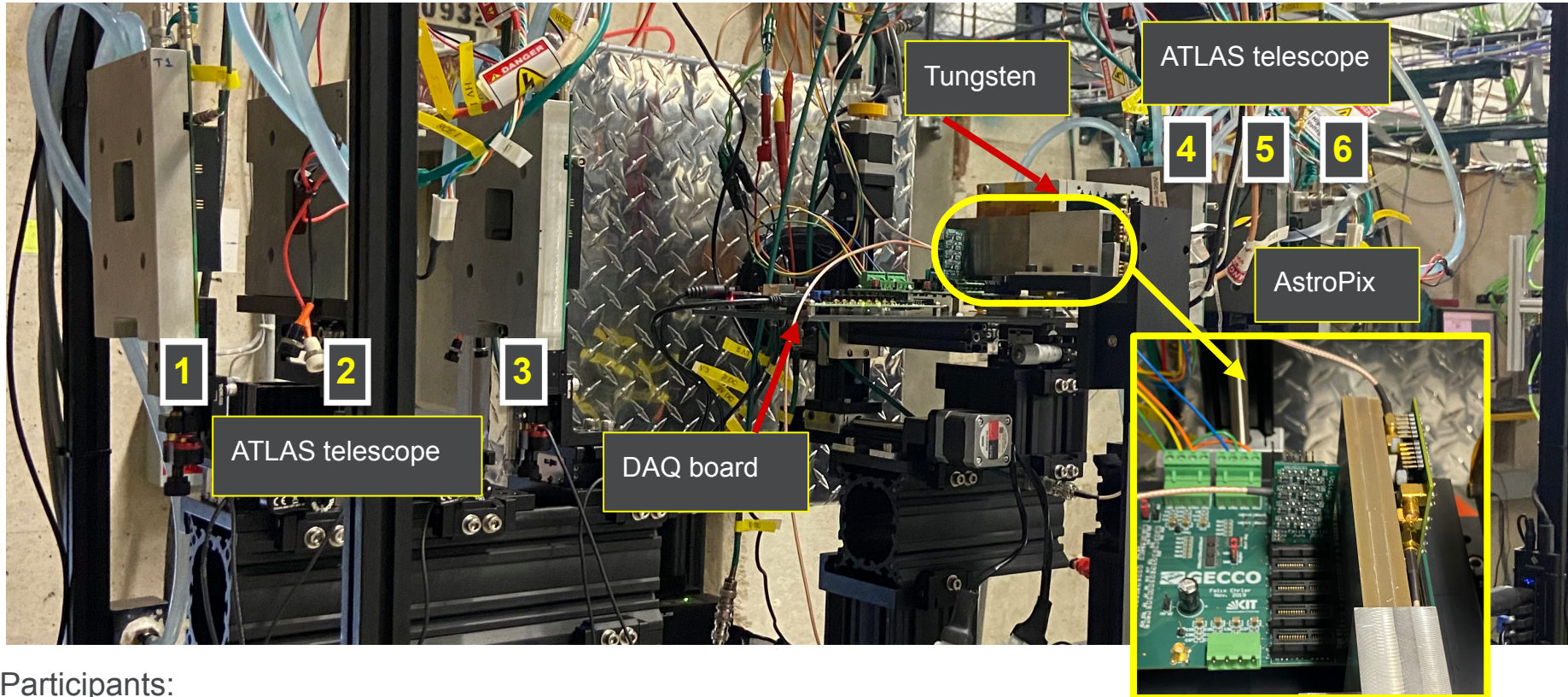
Ba133



Mean ToT distribution



AstroPix v3 TestBeam

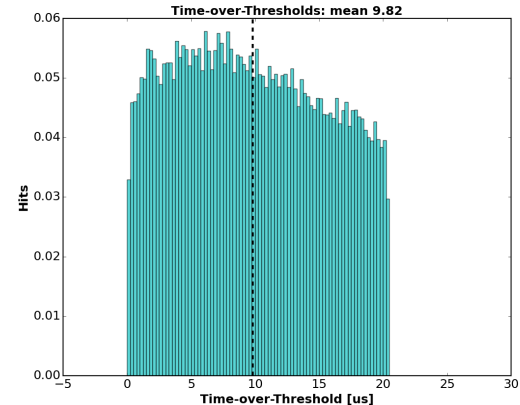
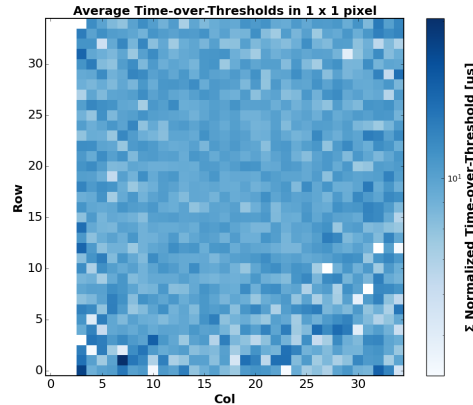
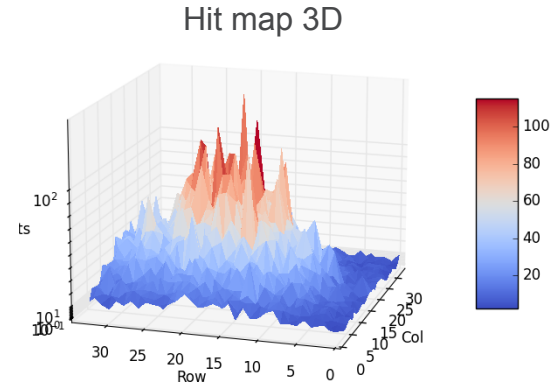
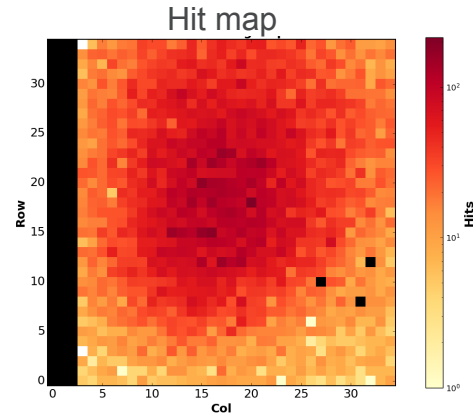


Participants:
Jihee Kim, Maria Zurek, Manoj Jadhav, Jessica Metcalfe

AstroPix v3 TestBeam Results

120 GeV Proton Low intensity beam

- 120 GeV Proton
 - 5000 protons/spill
 - 4.7 mm × 4.8 mm beam spot
- Data acquisition
 - Total 8 hours
 - 300 mV threshold
 - HV bias voltage 150 V
- Total 37,472 raw events
 - 96.67 % of events were decodable
 - 44,742 pixels* were fired
 - Among 91.1 % of active pixels, 91.02 % of pixels were fired
- Uniform pixel response for Minimum Ionizing Particle



Average ToT in Pixel array

ToT distribution

*Matching hits with exact time timestamp + ToT matching

Software and Firmware Development

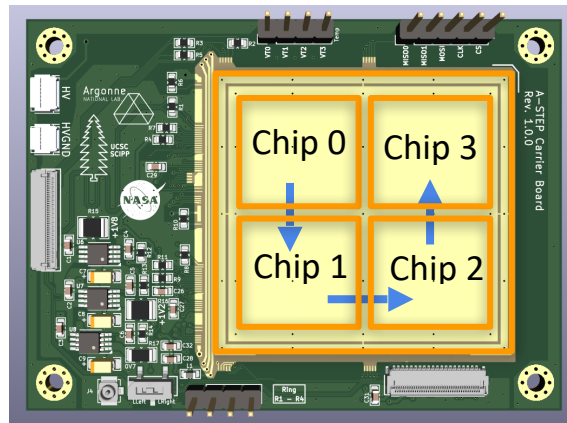
SW led by N. Striebig (KIT) and A. Steinhebel (GSFC), FW led by R. Leys and N. Striebig (KIT)

- FW-driven SPI readout to reduce deadtime
 - The self-trigger readout when there is data in buffers without SW check
 - Sensor data frame detection, IDLE discard, Tagging/reframing, routing to single Readout Buffer
- FW Scale-ability
 - Read through the daisy chain in FW rather than SW
 - Up to 20 daisy-chained SPI inputs have their own interfaces, which feed into the global buffer
- SW speedup to match FW
 - Reduce the chance of incomplete data return
 - Speed-up in analysis scripts, esp. when probing every pixel individually

Slide from Amanda's Talk in TIC

Next Steps

- Energy calibration of AstroPix v3
- Optimize configuration settings
- Depletion depth measurements are ongoing
 - TCT data collected by Amanda with the help of UCSC colleagues
- Irradiated AstroPix v3 with different 400 MeV Proton dose
 - Data analysis and sensor characterization of irradiated samples
- AstroPix v3 quad-chip development
 - Carrier board if designed by Taylor (check his talk this afternoon)
 - We have assembled the first few boards
 - They are under testing
- Module-like board design with 8 single chips is undergoing



AstroPix v3 quad-chip carrier board

- Demonstrate required services
- Daisy chaining

Thank you

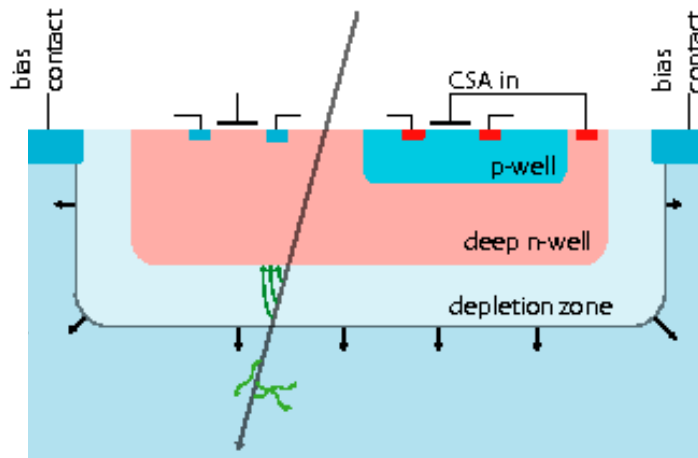
AstroPix

HV-CMOS Monolithic Active Pixel Sensor (MAPS):

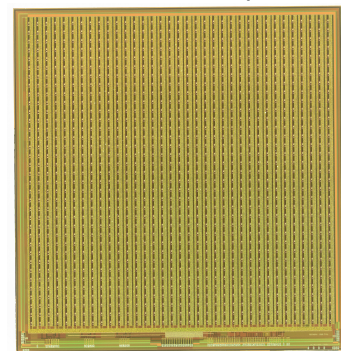
- Combination of silicon pixel and front-end ASIC
- On-pixel charge amplification and digitization
- The technology uses more typical CMOS wafer processing for cost-effective mass production
- Fabrication on a single wafer enables a shorter design cycle
- No need to bump-bond to each pixel - improves yield

AstroPix (based on ATLASPix3 [arXiv:2109.13409](https://arxiv.org/abs/2109.13409))

- 180 nm HV-CMOS MAPS sensor designed at KIT (also designed ATLASPix, MuPix, etc.)
- Developed for AMEGO-X GSFC/NASA mission (Upgrade to the Fermi's LAT)



AstroPix v3 (Under test)



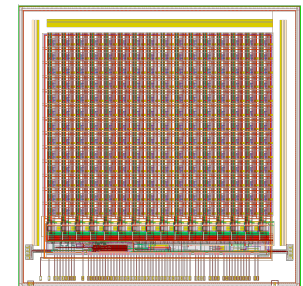
AstroPix

~2 cm

&

~2 cm

~1 cm



~1 cm

AstroPix v4
(Ready to test,
Nicolas' Talk)

AstroPix v3 Irradiation

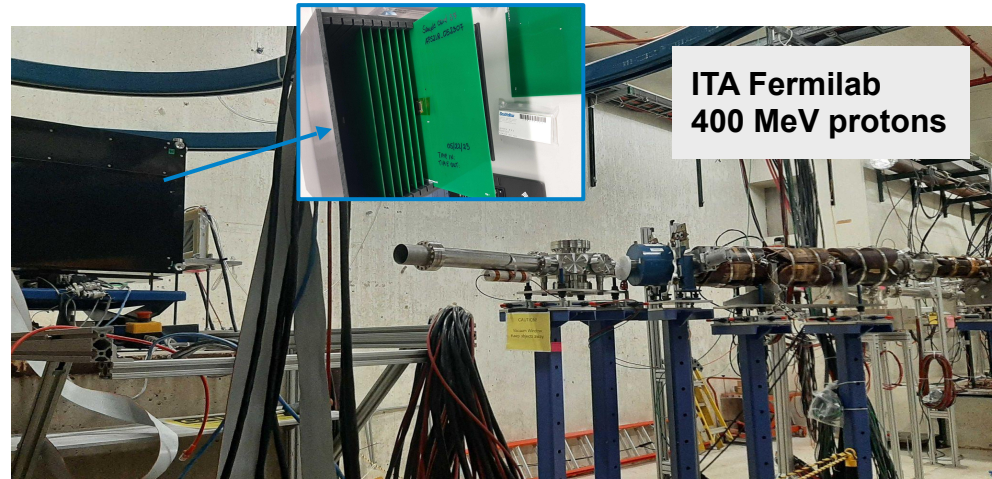
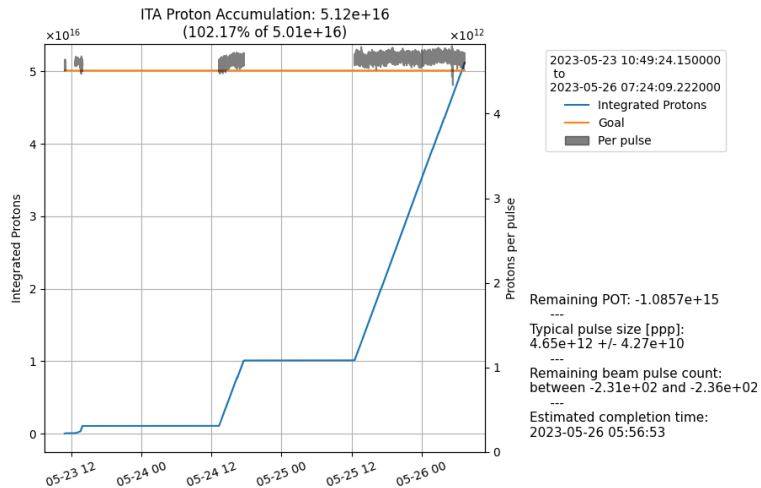
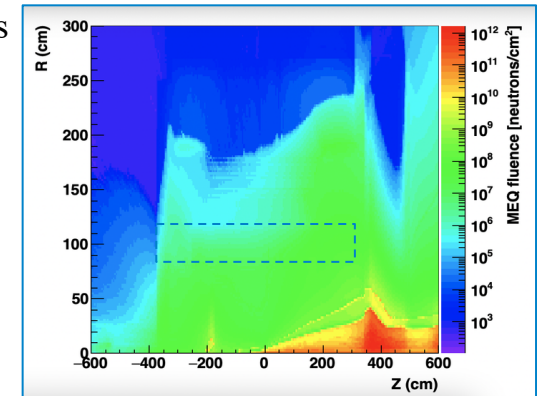
- IV and CV measurements performed for the v2/v3 chips before irradiations
 - Same measurements will be repeated post irradiation
- 9 v2 & 6 v3 chips irradiated for Passive Irradiation (Al-foil dosimetry)
- Active Irradiation for Latch-up (and SEE) is planned - week of 26th May

V2 Irradiation

Nb of samples	Doses (400 MeV protons)
3	4.50E+13
3	1.08E+15
2	1.01E+16
1	5.02E+16

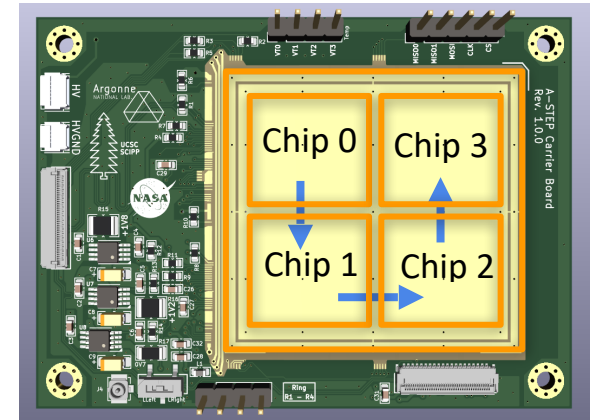
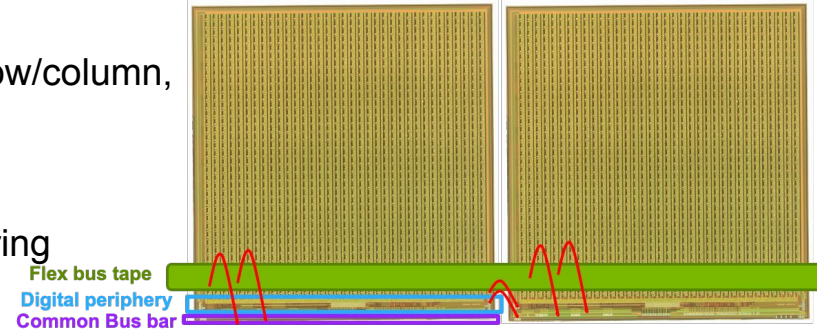
V3 Irradiation (low and high ResChips)

Nb of samples	Doses (400 MeV protons)
2	4.50E+13
1	5.04E+15



AstroPix Readout

- 10 bytes of data per hit - header (chipID, payload), row/column, timestamp, ToT
- SPI I/O daisy chained - chip-to-chip signal transfer
 - signals are digitized and routed out to the neighboring chip using 5 SPI lines via wire bond
- Power/Logic I/O distribution on the module (through a bus tape)
 - 4 power lines (LV, HV), ~20 Logic I/O (SPI, clk, timestamp, interrupt, digital Injection, etc.)
 - HV, VDDA/VDDD(1.8V), VSSA(1.2V), Vminuspix(0.7V)
 - power distribution can be controlled using voltage regulators
 - mostly part of the end of the stave services
- Data will be received by FPGA at the end of the stave
 - FPGA aggregates data before sending off-detector
- Low heat load at chip, only cooling of end of the stave card
- The operational temperature for AstroPix is at room temperature and considered to be operated at 22 °C



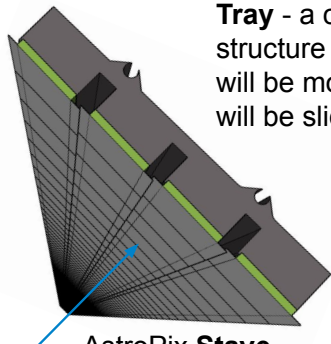
AstroPix v3 quad-chip carrier board

- Demonstrate required services
- Daisy chaining

AstroPix Assembly

AstroPix v5 (Production version)

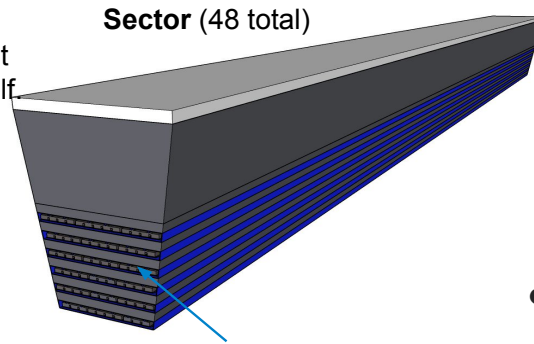
- Full size chip - $2 \times 2 \text{ cm}^2$, pixel pitch $500 \mu\text{m}$,
- 35×35 pixel matrix \rightarrow 1225 hit buffers
- Fix any bug from v4



Tray - a carbon fiber structure the staves will be mounted on. It will be slid into a shelf

AstroPix **Stave**

Consists of 1×108 chips with the support structure, "turbofanned"
AstroPix **Module**
Subset of chips



Sector (48 total)

Shelf - a carbon fiber structure that is glued to the Pb/ScFi layers, that we will slide trays with AstroPix staves on.

Module Strategy

- QC testing with wafer probing + Module and stave level QC testing and tuning
- "Baseline" model of Modules on Stave
 - Module - 8 single chips
 - Stave - 13 Modules - 104 chips
 - 12 or 14 Staves per AstroPix layer per Calorimeter Sector
 - Total 249600 chips
- All staves are identical and get combined in a separate production step
- Data is transmitted to the end of the Stave card using flex base tape
- Institutions - ANL, GSFC/NASA, KIT, UCSC, Korea, Oklahoma State

*The designs presented on these slides are not final but for illustration only

