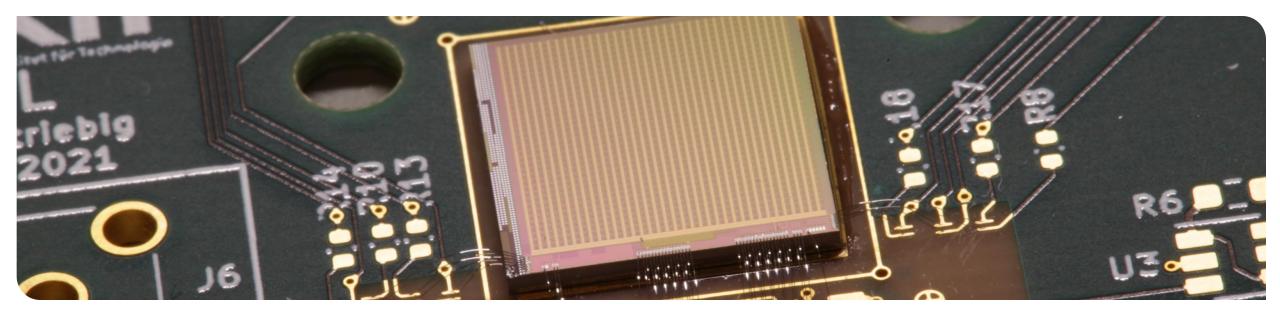




## **AstroPix Questions EIC Workshop**

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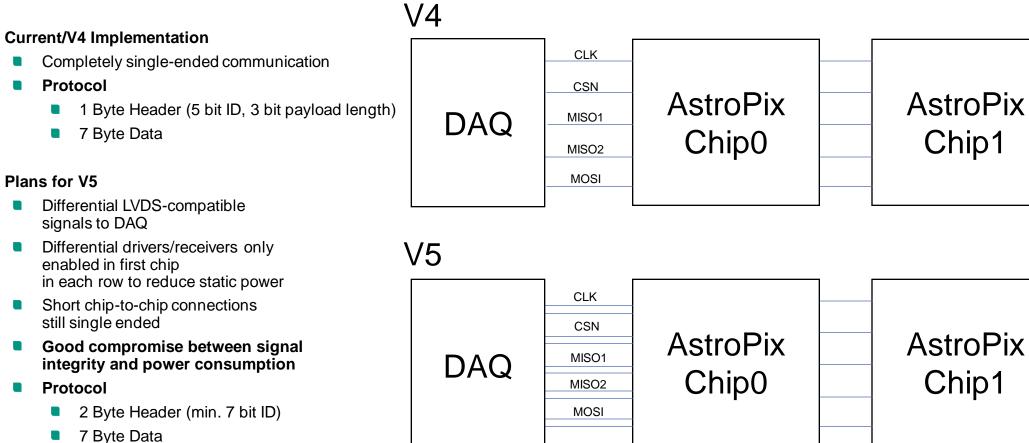
#### AstroPix Design Meeting, Sep 2023

# **SPI Data Transmission**

Protocol

Protocol

Plans for V5









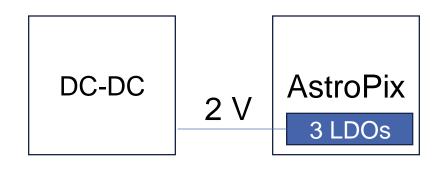
# **SPI Chain Reliability**

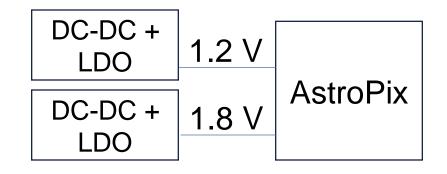
- High probability
  - Noisy pixels
    - Pixels can be enabled/disabled individually
    - Does not disturb readout chain
- Low probability
  - Failure in the digital logic especially in the SPI interface
    - Could potentially disturb the entire chain or a part of it
    - Bypass switches for SPI lines possible, but should not be controlled via SPI -> Second Control connection to the chip needed -> Shift Register Interface?
    - After Manufacturing: Wafer level testing -> Implement scan chain?



### Power

- Especially analog circuits connected to VDDA and VSSA are susceptible to noise
- Usually DC-DC step-down converters are used together with low noise linear regulators
- Voltages needed by the chip:
  - VDDA = 1.8 V
  - VSSA = 1.2 V
  - VDDD = 1.8 V
  - Vminuxpix = 0-1V
- On-chip regulator
  - Advantages
    - Simpler module design
- Off-chip regulator
  - Advantages
    - No additional chip design complexity
    - One regulator can supply several chips
    - Higher efficiency due to lower dropout voltage of commercial off-the-shelf LDOs









# **Integrated Voltage Regulator**



- Could be based on MuPix/ATLASpix design
- Single 2 2.2V input from DC-DC on module to generate voltages internally
  - VDDA,VSSA = 1.8 V
  - VDDD = 1.8 V
  - Vminuxpix = 0-1V
- Option to use internal regulators or external supplies
- More design time needed
- Not guaranteed to work better than external solution



# **Power supply specifications**



- Higher voltages than nominal are generally no problem as long as they don't exceed 1.95V
- Digital VDDD should work down to 1.62 V
- Lower VDDA and VSSA can have impact on performance
  - Reduced VDDA decreases amplifier output swing -> lower dynamic range
  - Characterisation with respect to different operating temperatures needed
- For impact of ripple and noise, characterisation is needed
  - As reference our lab supplies from Toellner (8733) are specified to 50 uV ripple



# **Dynamic Range**



- Amplifier is optimized for high dynamic range required by AMEGO-X 20 700 keV
- Closed loop gain V ~ Q/C
  - High gain seen by small signals
    - Small feedback capacitance ~500aF
  - Low gain seen by signals
    - 10 fF VNCAP
    - Configbit to disable low gain mode by disconnecting the capacitor through a switch
- If needed an additional capacitor and switch could be added for EIC
  - Low risk, circuits already tested in AstroPix2-4
- Can we even absorb 3 MeV with a 500um depletion?

