

# Observation of antihypertriton in Au+Au collisions at 200 GeV -- the antinuclei journey with Prof. Keane

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STAR Col. Science **328**, 58 (2010)

STAR Col. Nat. Phys. **16**, 409 (2020)

J. Chen, D. Keane, Y.-G. Ma, A. Tang, Z. Xu, Phys. Rept. **760**, 1 (2018)

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# Outline

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- ★ First of all, Happy 70<sup>th</sup> Birthday!
- ★ Introduction to the antinuclei study
- ★ Focus on the antihypertriton
- ★ Summary



# Introduction

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in 2009, we said:

- ★ **Hypernuclei**: ideal lab for YN and YY interaction
  - Baryon-baryon interaction with strangeness sector
  - Input for theory describing the nature of neutron stars
  
- ★ No **anti-hypernuclei** have ever been observed
  
- ★ **Coalescence mechanism** for production: depends on overlapping wave functions of Y+N at final stage
  
- ★ Anti-hypernuclei and hypernuclei ratios: sensitive to **anti-matter and matter** profiles in HIC
  - Extension of the nuclear chart into anti-matter with S <sup>[1]</sup>

[1] W. Greiner, *Int. J. Mod. Phys. E* 5 (1995) 1



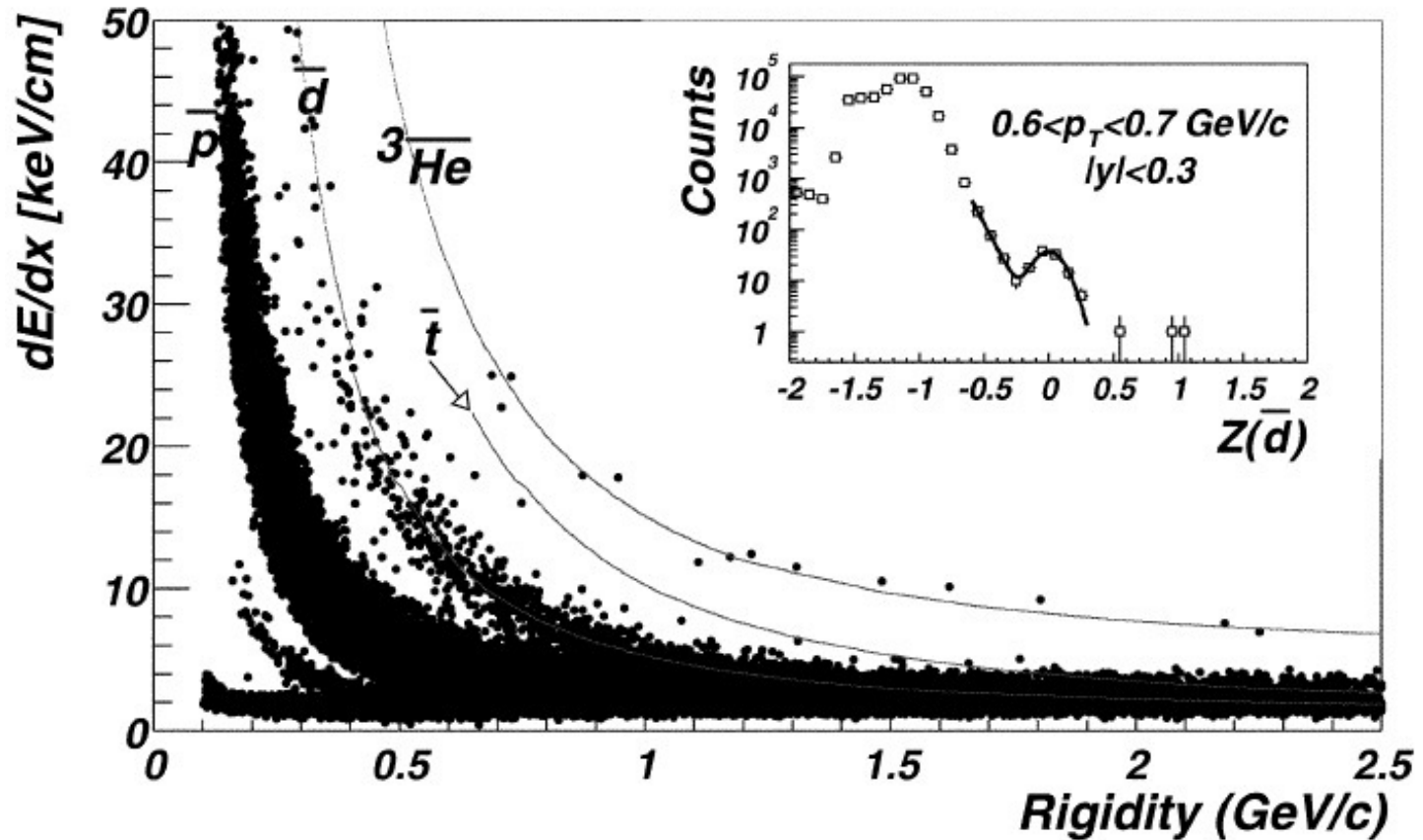
# STAR has measured antinuclei in run1

VOLUME 87, NUMBER 26

PHYSICAL REVIEW LETTERS

24 DECEMBER 2001

$\bar{d}$  and  ${}^3\bar{\text{He}}$  Production in  $\sqrt{s_{NN}} = 130$  GeV Au + Au Collisions

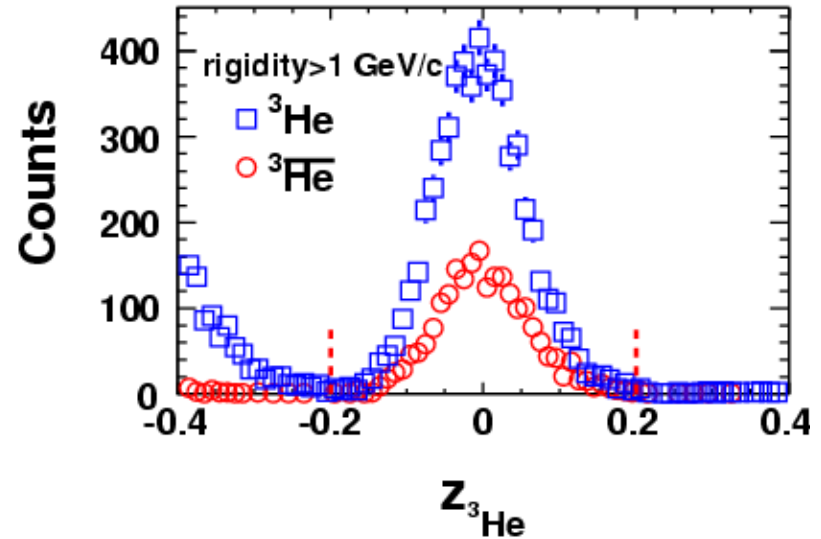
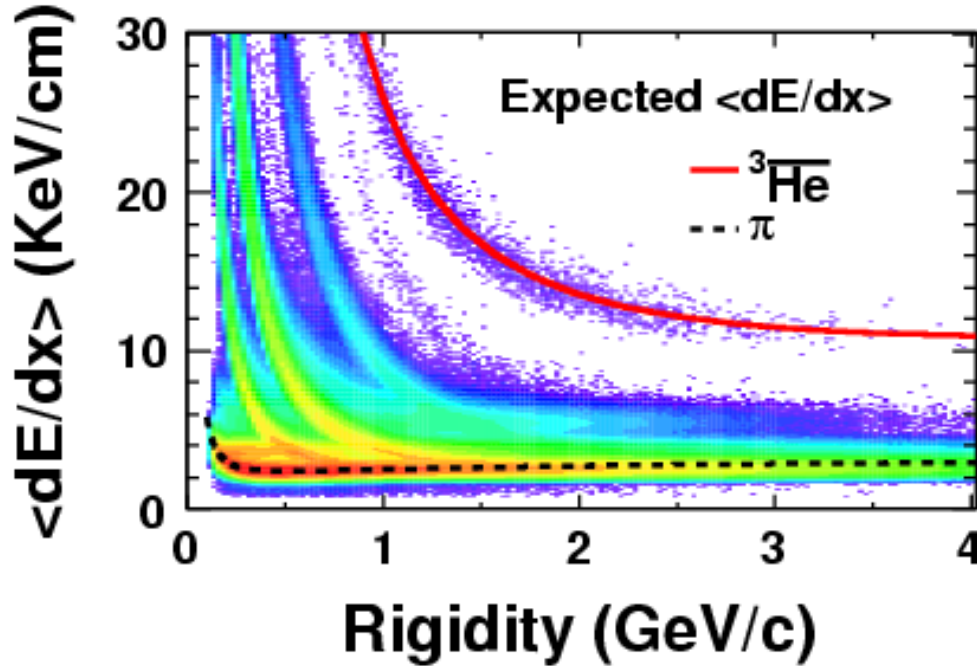


★ 14 anti- ${}^3\text{He}$  based on 0.6M central Au+Au at 130 GeV





# Collect giant data with run4 + run7



$$z = \ln\left(\frac{\langle dE / dx \rangle}{\langle dE / dx \rangle^{th}}\right)$$

Theory curve: *Phys. Lett. B 667 (2008) 1*

★ Select pure  ${}^3\text{He}$  sample:  **${}^3\text{He}$ : 5810 counts**

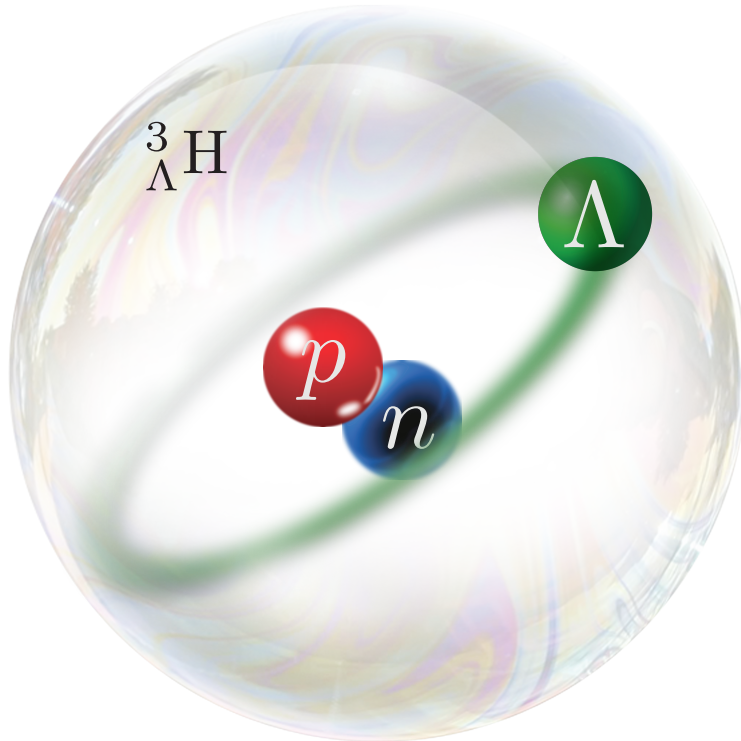
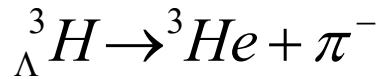
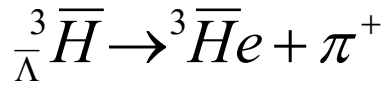
**anti- ${}^3\text{He}$ : 2168 counts**

condition:  $-0.2 < z < 0.2$  &  $dca < 1.0\text{cm}$  &  $p > 2\text{ GeV/c}$ ...

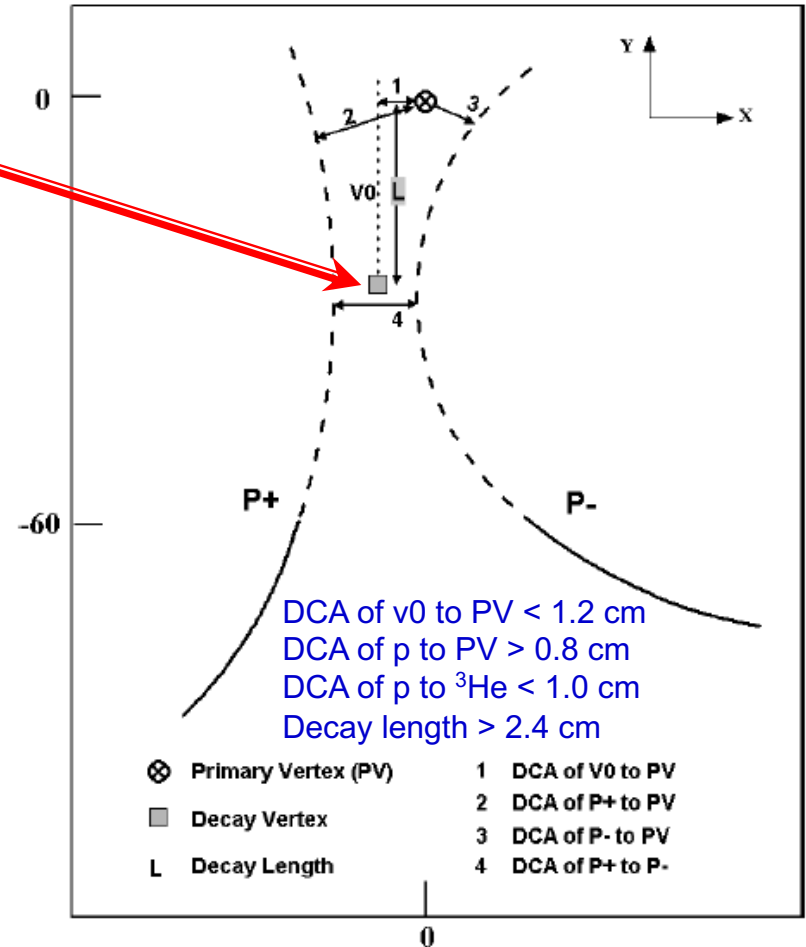


# Extend the antinuclei study with v0

${}^3_{\Lambda}H$  mesonic decay,  $m=2.991$  GeV, B.R. 0.25;

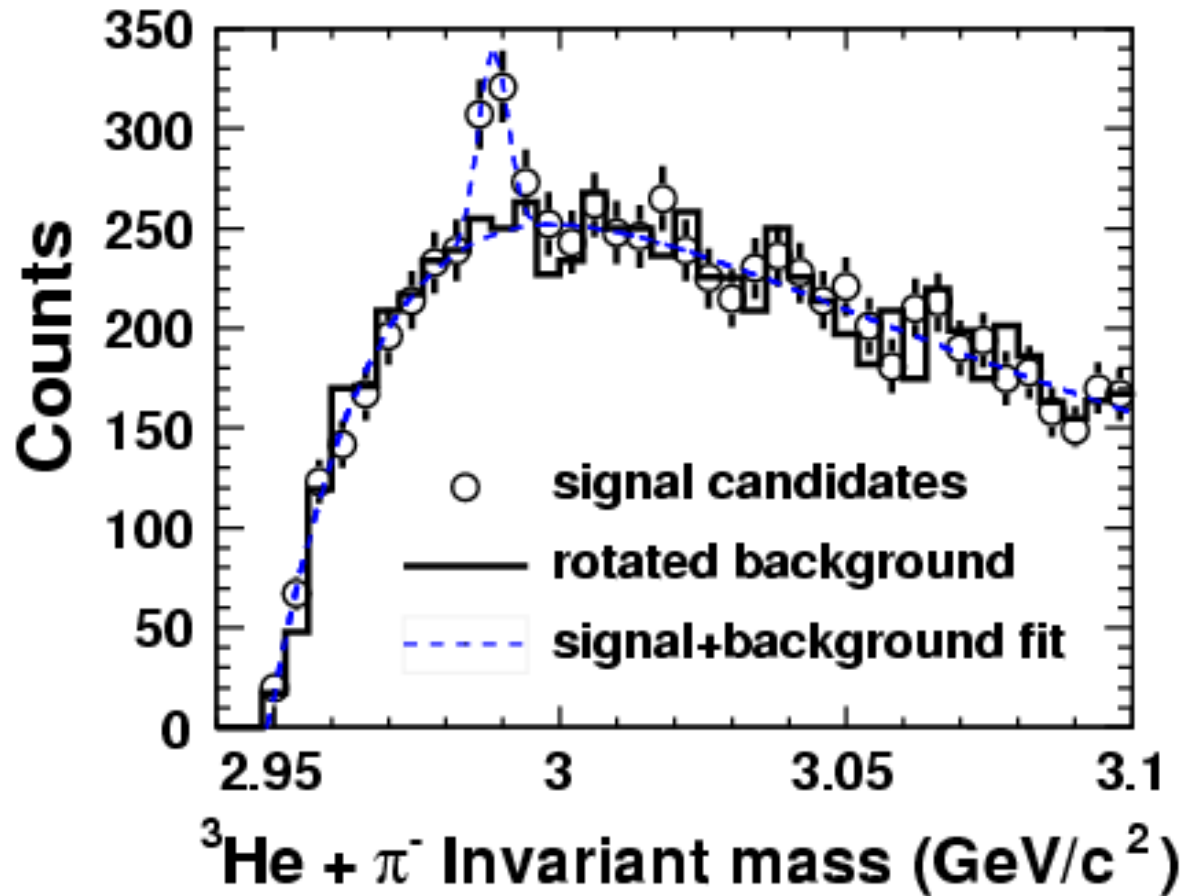


## Secondary vertex finding technique





# ${}^3_{\Lambda}\text{H}$ signal from the data



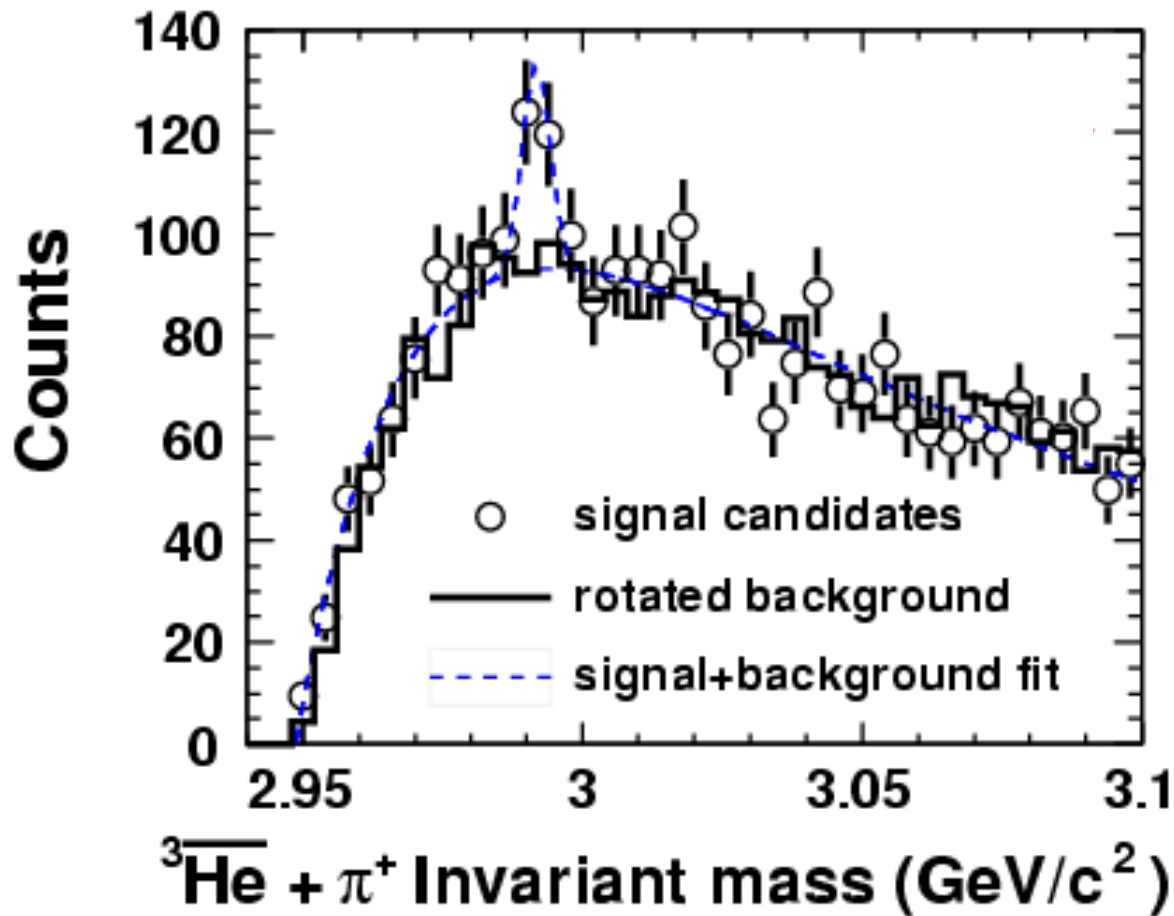
★ Signal observed from the data (bin-by-bin counting):  **$157 \pm 30$** ;

Mass:  $2.989 \pm 0.001 \pm 0.002$  GeV; Width (fixed): 0.0025 GeV.

★ Projection on anti-hypertriton yield:  $\frac{{}^3\bar{\text{H}}_{\Lambda}}{\Lambda} = \frac{{}^3\bar{\text{H}}_{\Lambda} \times {}^3\bar{\text{He}}}{{}^3\text{He}} = 157 \times 2168 / 5810 = \mathbf{59 \pm 11}$



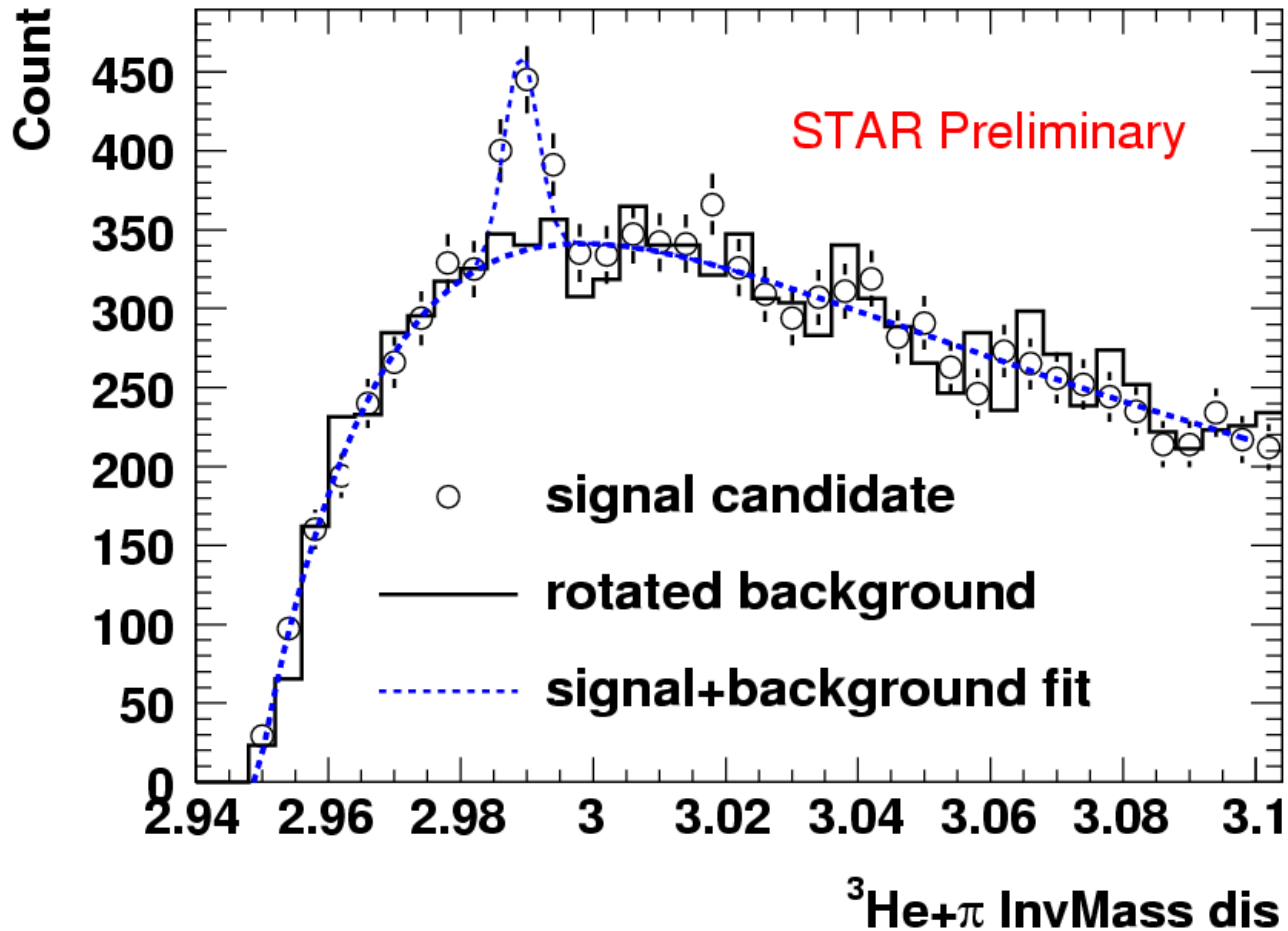
# $\frac{3}{\Lambda} \bar{\text{H}}$ signal from the data



- ★ Signal observed from the data (bin-by-bin counting):  $70 \pm 17$ ;  
Mass:  $2.991 \pm 0.001 \pm 0.002$  GeV; Width (fixed): 0.0025 GeV.



# Combined the signal

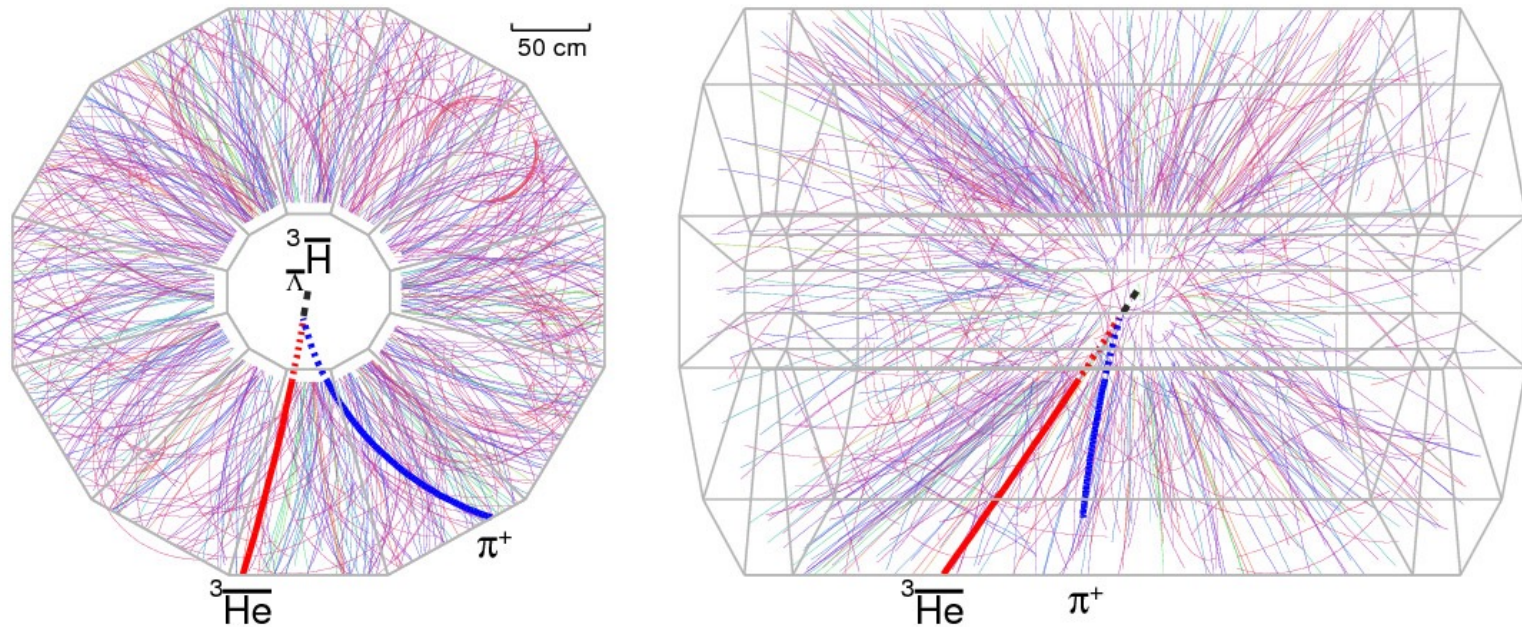


★ Combined hyperT and anti-hyperT signal :  $225 \pm 35$ ;

It provides a  $>6\sigma$  significance for discovery.



# A beautiful event and the PR

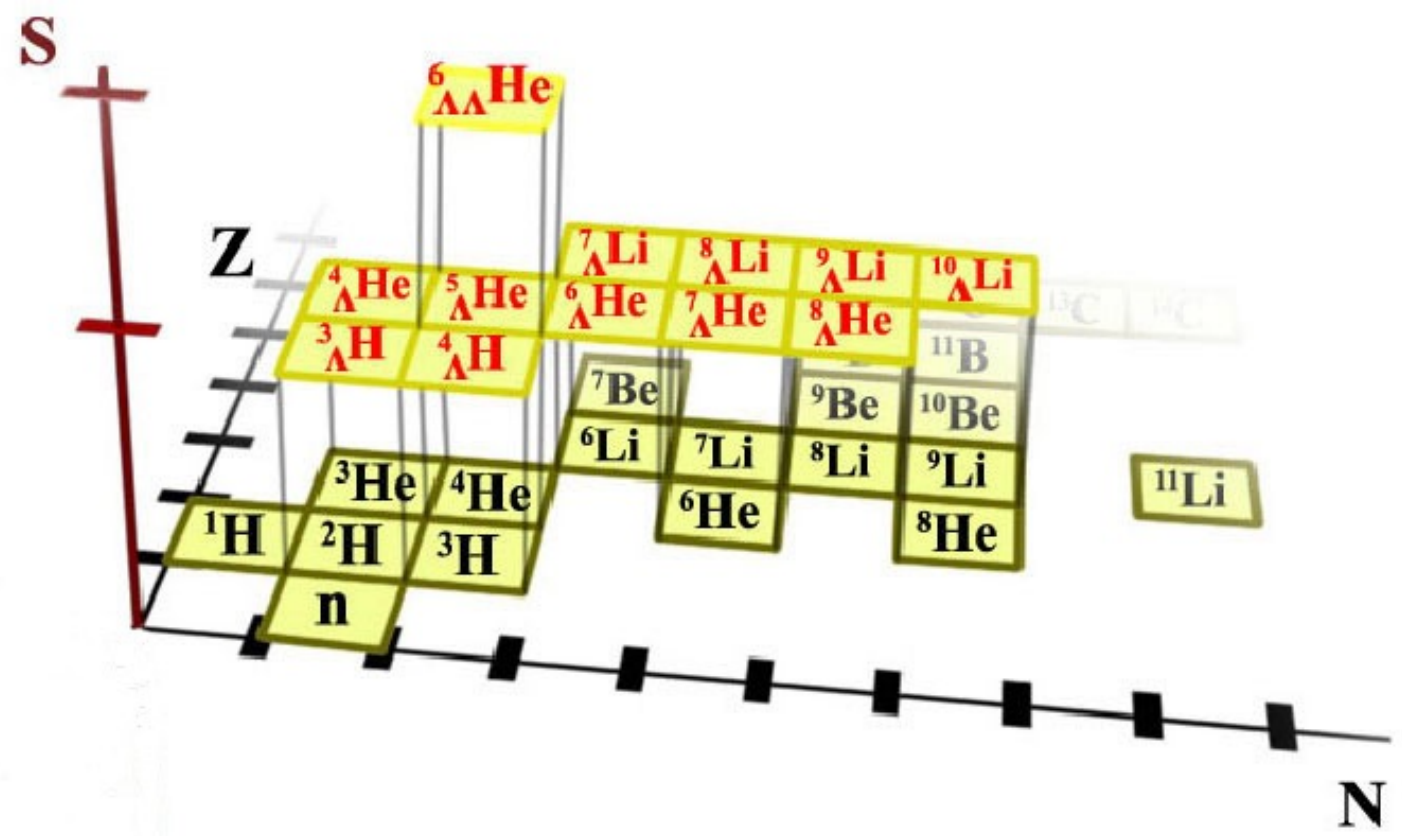


- ★ Anti-hyperT : anti-proton, anti-neutron & anti- $\Lambda$  – the first antinucleus with strangeness, and the heaviest antinucleus until 2011.
- ★ After searching >100 million AuAu collisions, found 70 anti-hyperT.
- ★ Published in *Science* in March 2010; much favorable PR for STAR & RHIC. News stories in *Nature*, *Scientific American*, *National Geographic*, many news outlets worldwide.



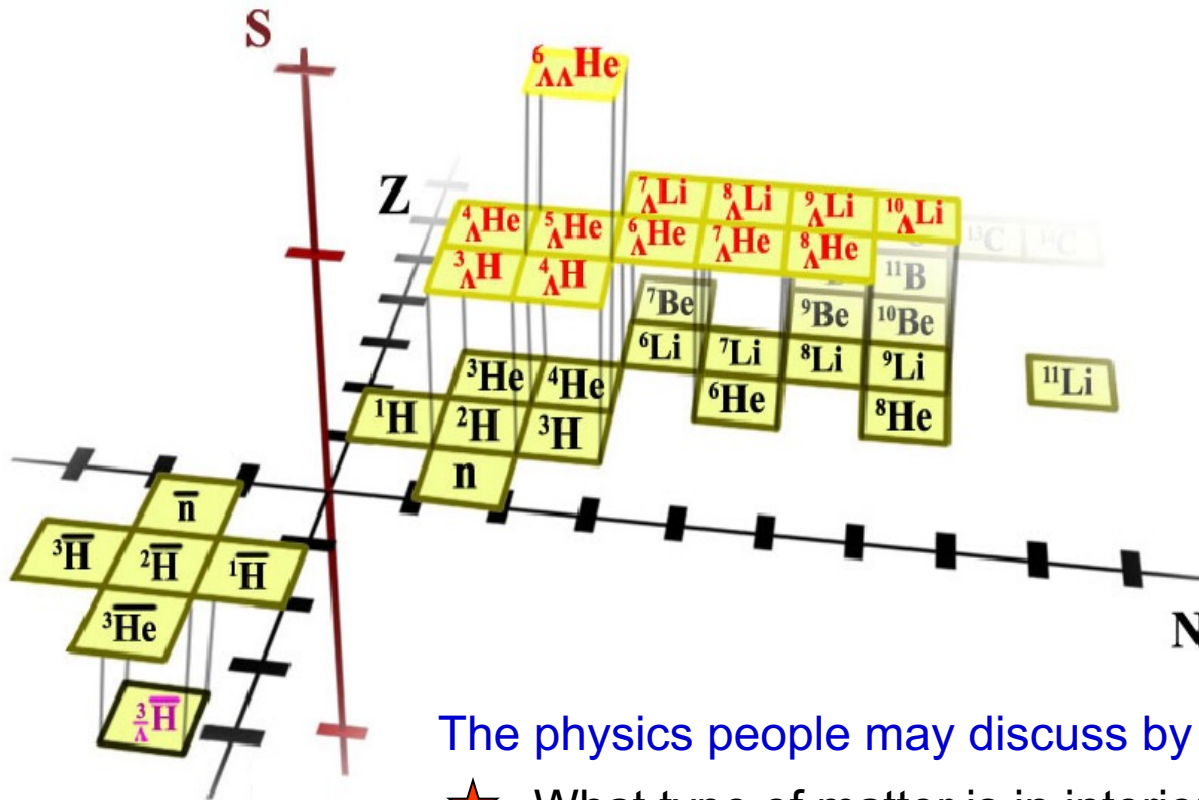


# 3-D Chart of the Nuclides





# Extension of the Chart of the Nuclides into antimatter with Strangeness sector



Out of >800 peer-reviewed papers from all Brookhaven programs in 2010, antimatter paper in *Science* was named by lab management in Jan 2011 as one of the “top 5 for 2010”

The physics people may discuss by measuring antinuclei:

- ★ What type of matter is in interior of collapsed stars?
- ★ What happened to antimatter created in the Big Bang?
- ★ Implications for cosmic ray searches for new physics.

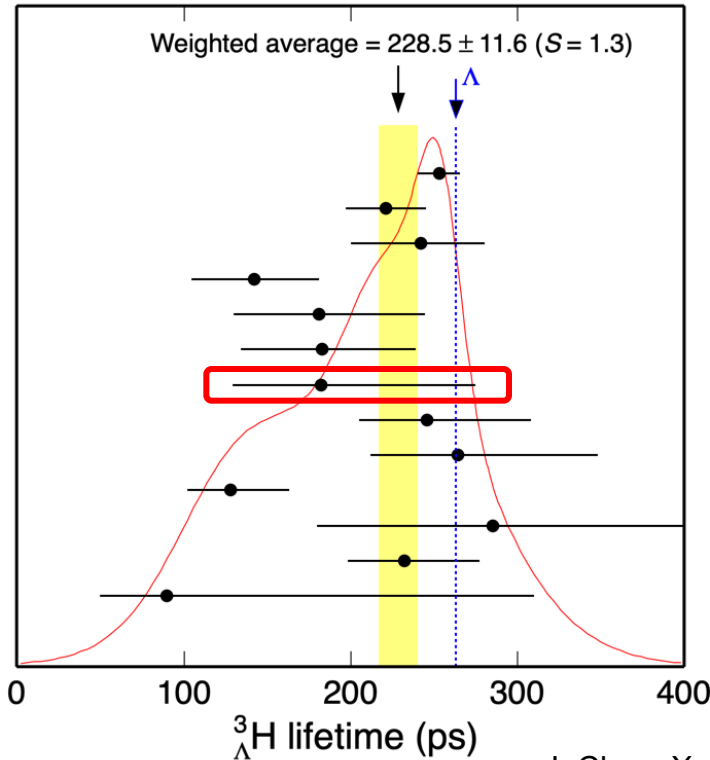




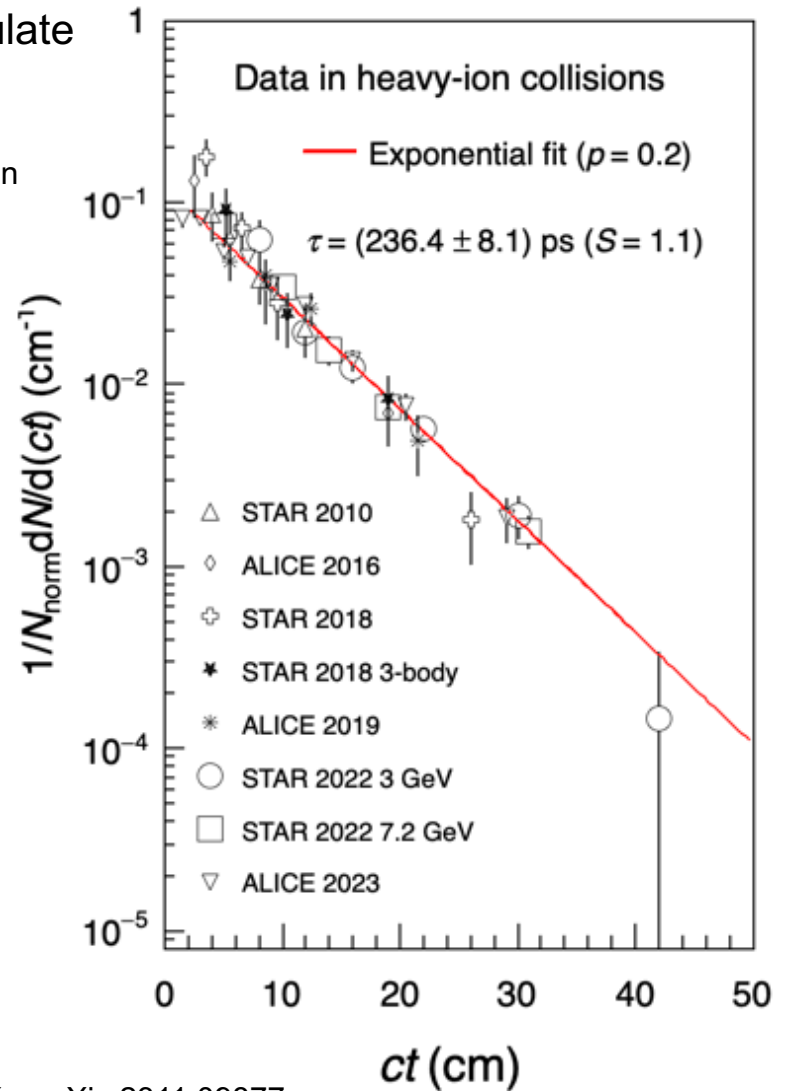
# The lifetime quest

★ First measurement of the lifetime in HIC and stimulate studies from other experiments.

“a measurement to a precision of a few percent will guide and constrain the theoretical input leading to a more precise determination of the YN interaction, eventually contributing to solving the hyperon puzzle”



|               | $\chi^2$ |
|---------------|----------|
| ALICE 2023    | 3.8      |
| STAR 2022     | 0.1      |
| ALICE 2019    | 0.1      |
| STAR 2018     | 4.9      |
| ALICE 2016    | 0.6      |
| HypHI 2013    | 0.7      |
| STAR 2010     | 0.3      |
| Keyes 1973    | 0.1      |
| Keyes 1970    | 0.4      |
| Bohm 1970     | 8.2      |
| Phillips 1969 |          |
| Keyes 1968    | 0.0      |
| Prem 1964     |          |
|               | 19.3     |
| C.L.          | = 0.036  |



J. Chen, X. Dong, Y.-G. Ma, Z. Xu, arXiv:2311.09877



# From $\tau$ to $\Lambda$ separation energy

Dalitz's comment on the lifetime "I feel that we are far from seeing the end of this road. A good deal of theoretical work on this 3-body system would still be well justified."

Nucl. Phys. A 754, 14 (2005)

★ The early data suffers from large statistical uncertainty!

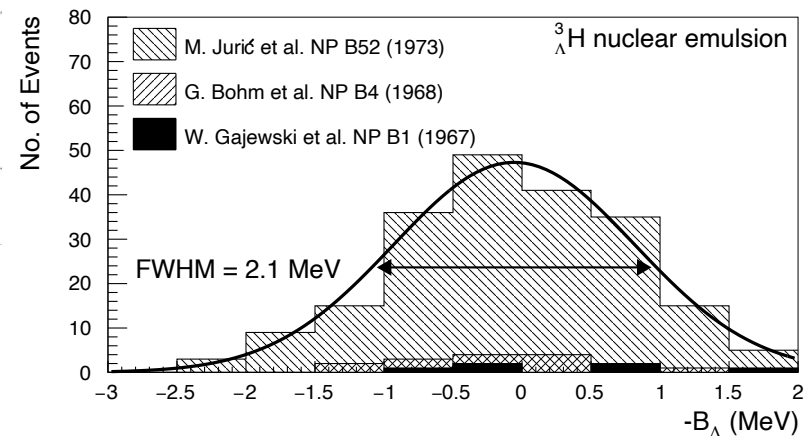
| Hypernuclide        | Decay mode                          | Number of events | $\bar{B}_\Lambda \pm \Delta\bar{B}_\Lambda$ (MeV) <sup>a)</sup> |
|---------------------|-------------------------------------|------------------|---|
| $\Lambda^3\text{H}$ | $\pi^- \ ^3\text{He}$ <sup>b)</sup> | 26               | $0.13 \pm 0.15$   |
|                     | $\pi^- \ ^1\text{H} \ ^2\text{H}$   | 6                | $0.33 \pm 0.21$   |
|                     | total                               | 32               | $0.20 \pm 0.12$   |

W. Gajewski et al., Nucl. Phys. B 1, 105 (1967)

| $B_\Lambda \pm \Delta B_\Lambda$ (MeV) |                           |                 |
|--|---------------------------|-----------------|
|  | Bohm et al. <sup>a)</sup> | This work       |
| $\Lambda^3\text{H}$                    | $0.01 \pm 0.07$           | $0.15 \pm 0.08$ |

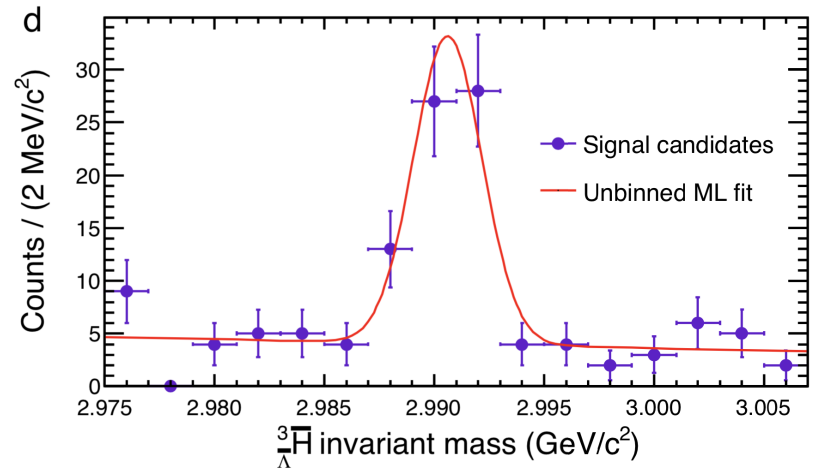
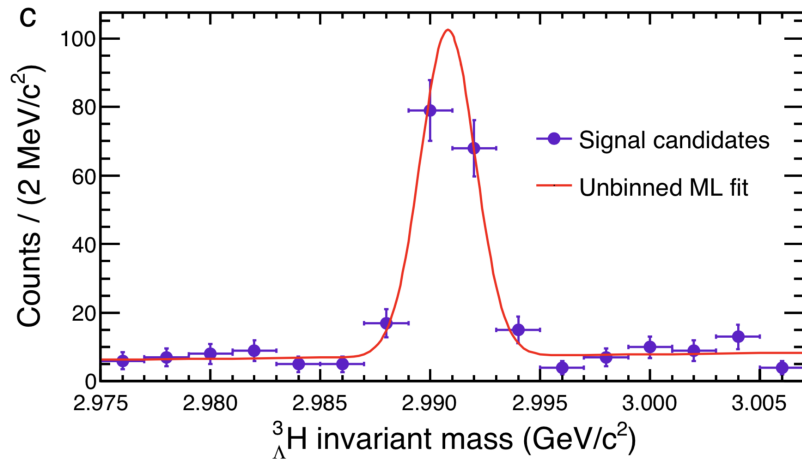
a) G. Bohm et al., Nucl. Phys. B4, 511 (1968)

b) This work : M. Juric, G. Bohm et al., Nucl. Phys. B52,1 (1973)



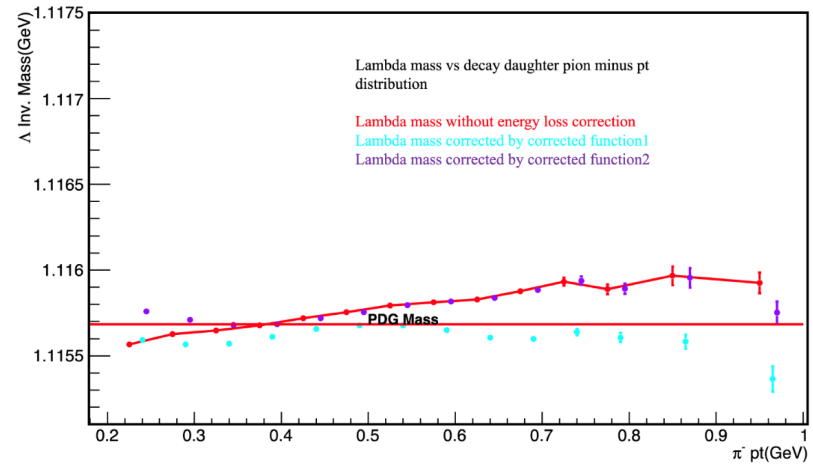
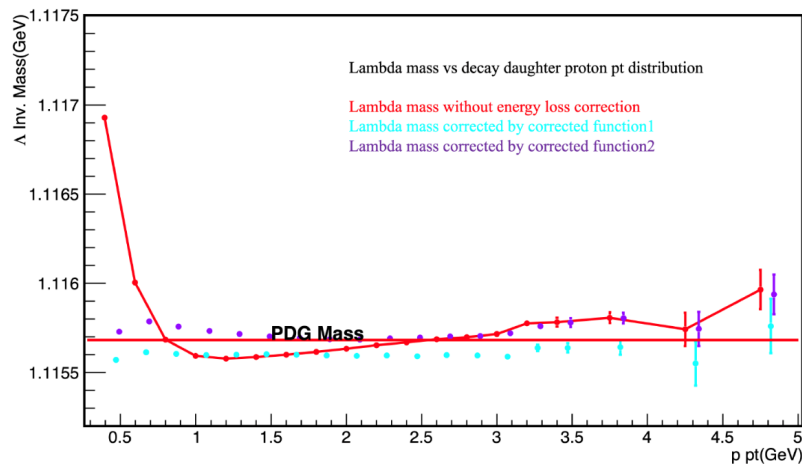


# STAR measurements with detector upgrade



$$m_{{}^3_{\Lambda}H} = 2990.95 \pm 0.13(\text{stat.}) \pm 0.11(\text{syst.}) \text{ MeV}/c^2$$

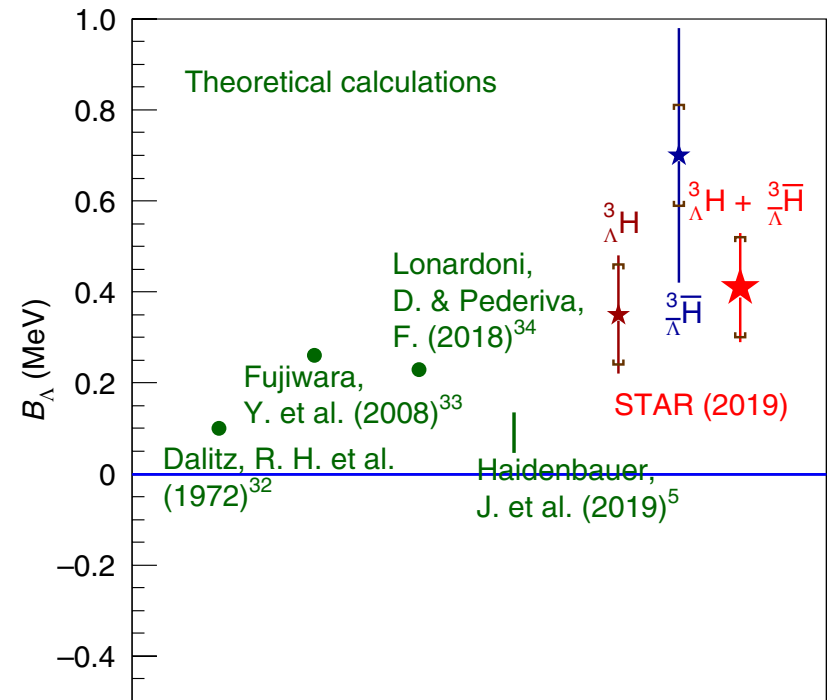
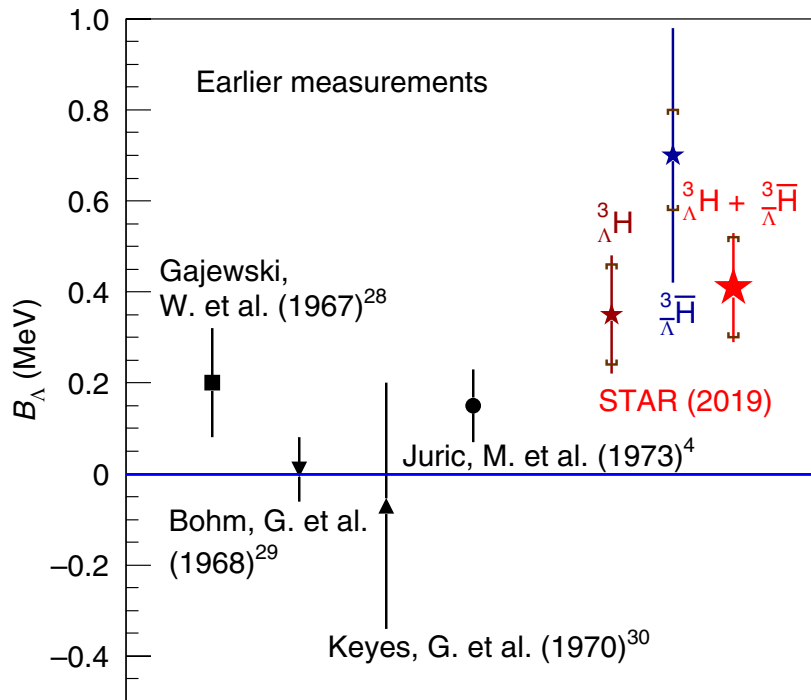
$$m_{{}^3_{\Lambda}\bar{H}} = 2990.60 \pm 0.28(\text{stat.}) \pm 0.11(\text{syst.}) \text{ MeV}/c^2$$



$\Lambda$  serves as a calibrated probe to understand the momentum distortion



# And the results

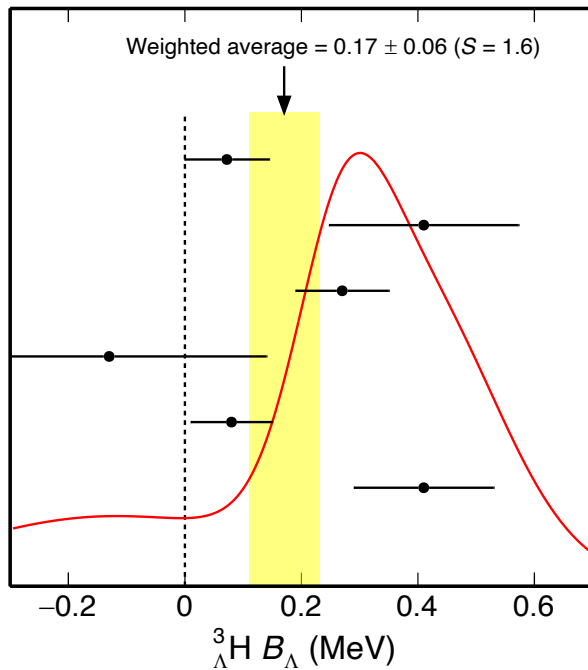


$$B_\Lambda = 0.41 \pm 0.12(\text{stat.}) \pm 0.11(\text{syst.}) \text{ MeV}$$

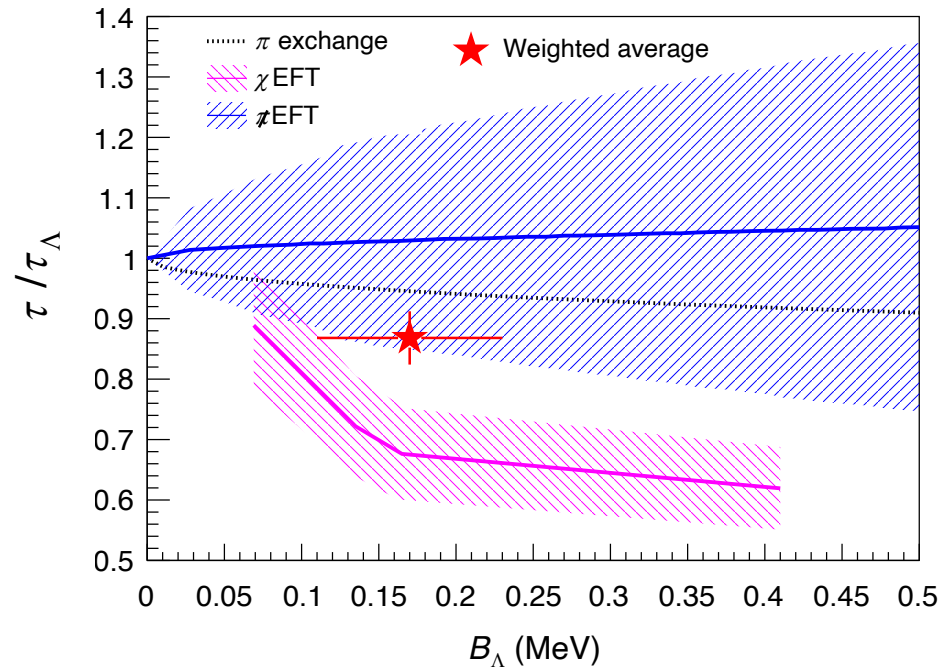
- ★ STAR data differs from zero ( $3.4\sigma$ ) and larger than the prior measurements from 1973
- ★ Theoretical calculations span in a wide range



# Correlated the $\tau$ and the $B_\Lambda$



|            | $\chi^2$ |
|------------|----------|
| ALICE 2023 | 1.8      |
| STAR 2020  | 2.2      |
| NPB52 1973 | 1.6      |
| PRD1 1970  | 1.2      |
| NPB4 1968  | 1.7      |
| NPB1 1967  | 4.0      |
|            | 12.4     |
| C.L.       | = 0.029  |



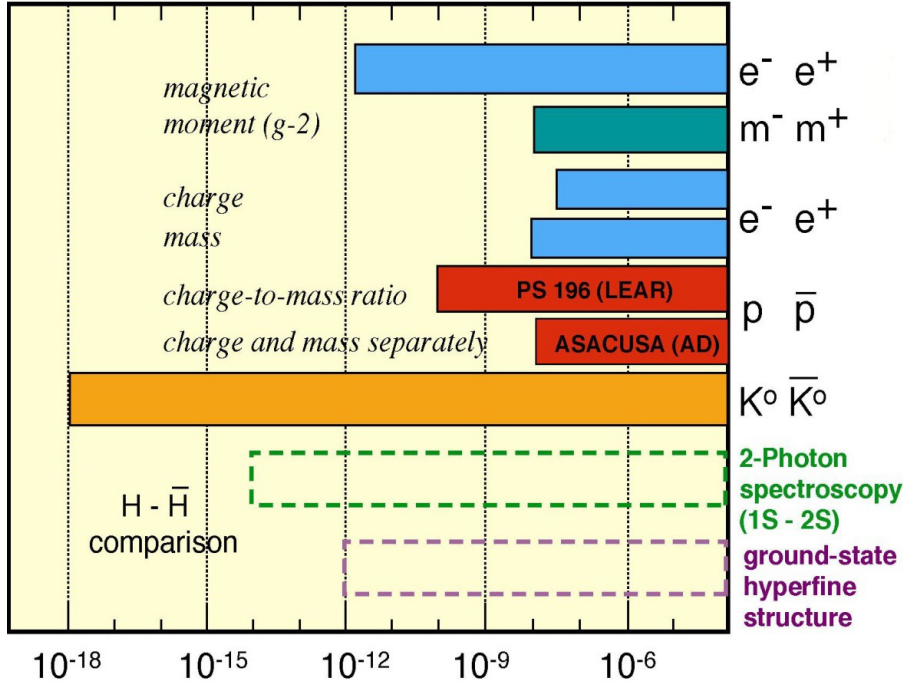
J. Chen, X. Dong, Y.-G. Ma, Z. Xu, arXiv:2311.09877

- ★ The ideogram presentation of the data points shows a large spread of the values
- ★ The weighted average value is consistent with predictions, considering the uncertainties associated in calculations



# The CPT test

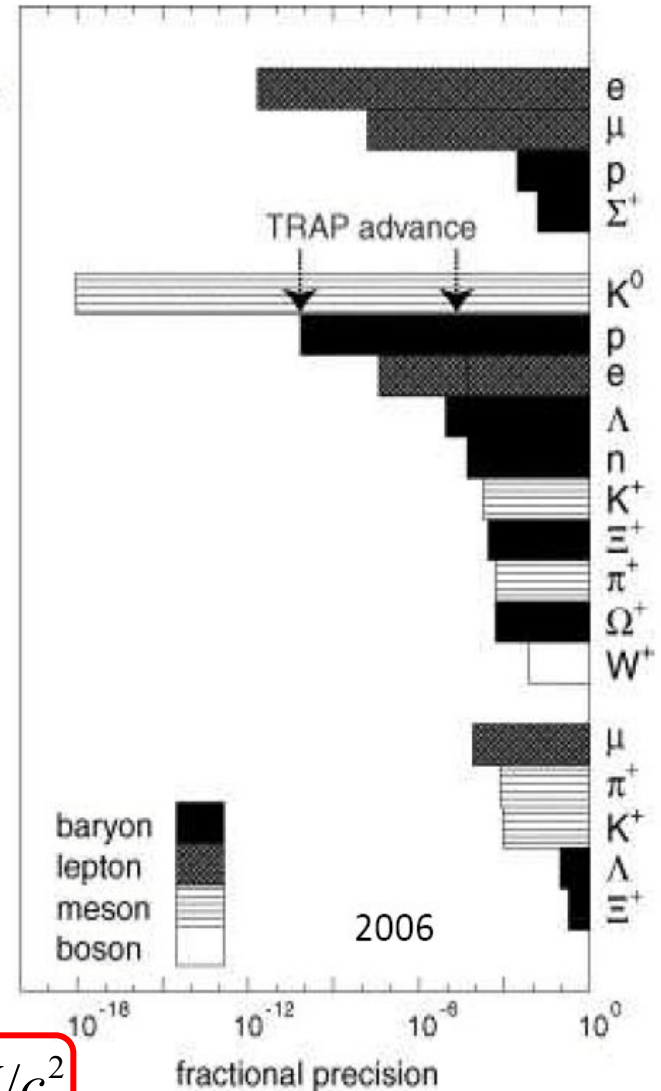
Precision of some CPT Tests



magnetic moment  
 $\frac{\Delta g}{g}$

mass  
 $\frac{\Delta m}{m}$

mean life  
 $\frac{\Delta \tau}{\tau}$

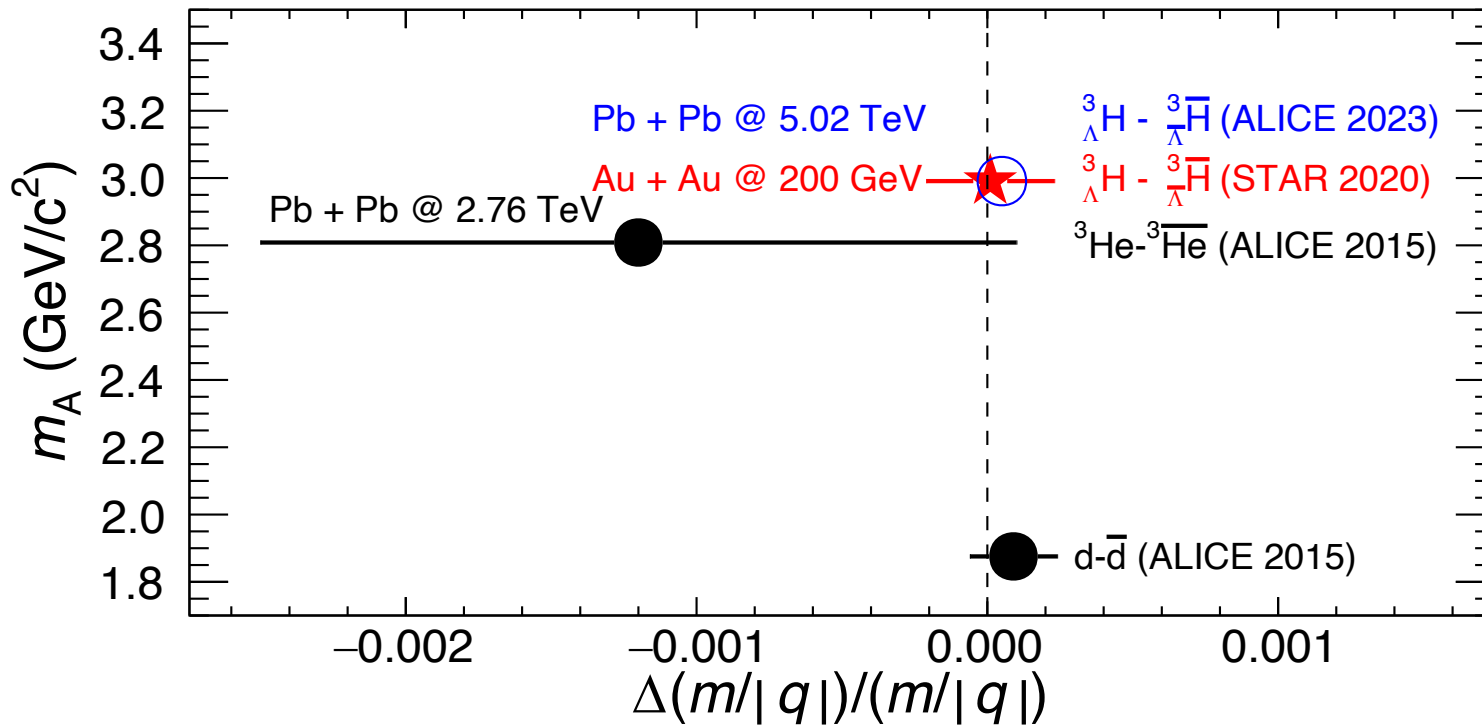


★ antinuclei mass difference to probe any diff. in the interaction between nucleons and antinucleons encoded in the antinuclei masses, which can not be directly derived from QCD yet.

Declan in GHP 2011  $\frac{3}{\Lambda} \bar{H} - \frac{3}{\Lambda} H \quad \Delta m = 0 \pm 1 \pm 2 \text{ MeV}/c^2$



# Improved precision on the test



★ From antihypertrion to anti<sup>3</sup>He, two order of magnitude improved

$$\frac{\Delta\mu_{3\text{He}^3\bar{\text{He}}}}{\mu_{3\text{He}}} = (-1.2 \pm 0.9(\text{stat.}) \pm 1.0(\text{syst.})) \times 10^{-3}$$

ALICE Col. Nat. Phys. 11 (2015) 811  
 STAR Col. Nat. Phys. 16 (2020) 409  
 ALICE Col. PRL 131 (2023) 102302

$$\frac{m_{\Lambda^3\text{H}} - m_{\Lambda^3\bar{\text{H}}}}{m} = (0.1 \pm 2.0(\text{stat.}) \pm 1.0(\text{syst.})) \times 10^{-4} \quad \frac{m_{\Lambda^3\text{H}} - m_{\Lambda^3\bar{\text{H}}}}{m_{\Lambda^3\text{H}}} = [5 \pm 5(\text{stat}) \pm 3(\text{syst})] \times 10^{-5},$$



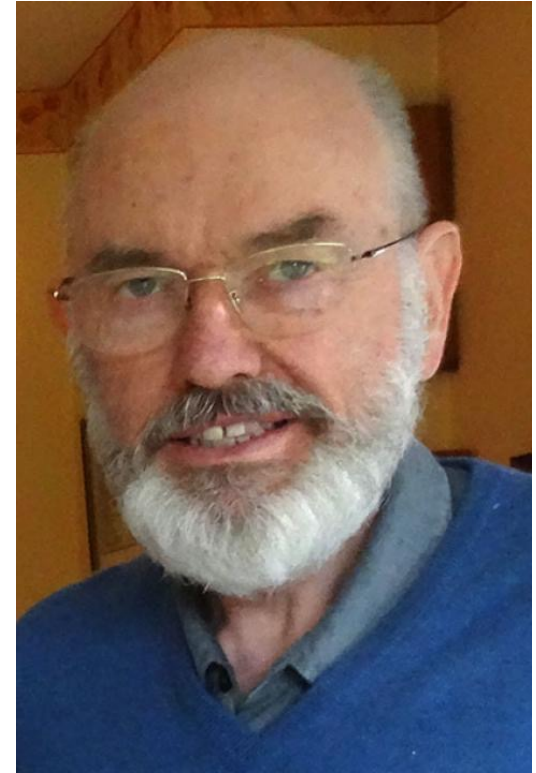


# Summary

photo from web page of ~2009



Our captain of the  
antinuclei journey



- ★ We have observed the 1<sup>st</sup> antihypernucleus, we have seen many interesting physics
- ★ How does his mentorship shape my career?
  - ❖ the 2009 project guaranteed a professor position
  - ❖ the 2020 measurement won the competition of NSFC for Distinguished Young Scholars
- ★ Wish you good health and looking forward to another 15-year journey