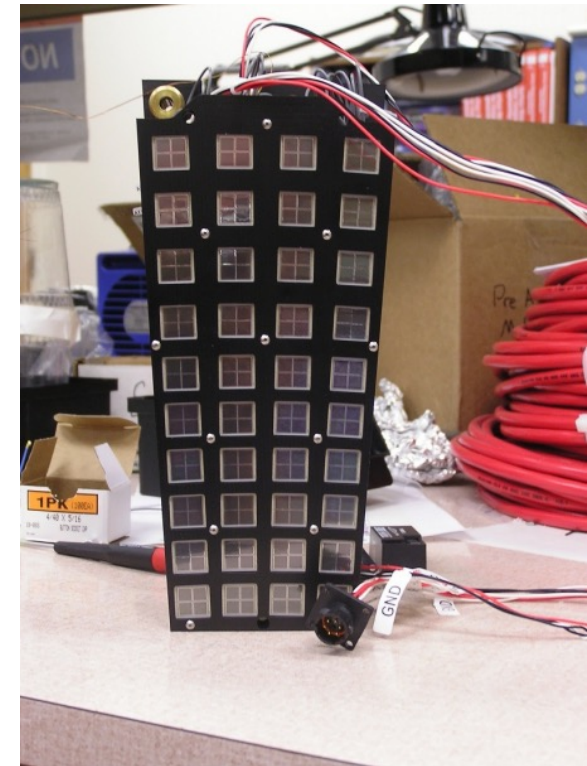


# “Baby ” BCAL Progress Report

Jon Zarling

5/14/24





# Primary Goals

Using 58 cm “baby” BCAL sector, improve determination of:

- Resolution ( $\sigma_E/E$ )
- Number of photoelectrons per GeV ( $N_{p.e.}$ )

GlueX BCAL: coverage up to 2 GeV

Pair Spectrometer in Hall D: **3-6 GeV**



# $e^+$ Beam Setup

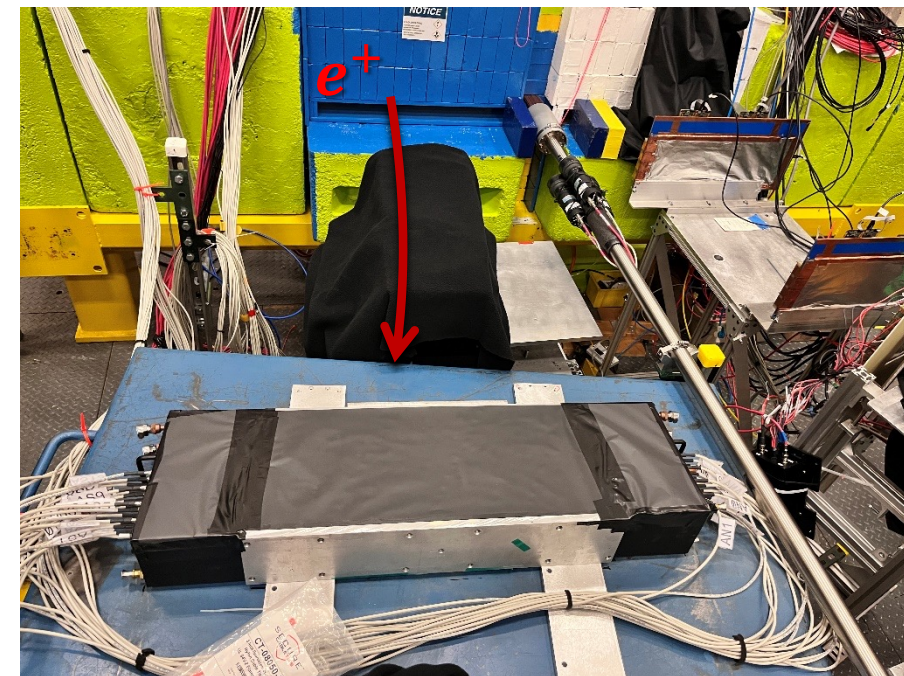
Good source of EM showers with known energy

Essentially identical to GlueX lead SciFi wedge, except 58 cm in length rather than 390 cm

Hall D pair spectrometer setup:

- $e^+$  tagged upstream roughly 3-6 GeV
- Unfortunately, SciGlass blocks portions
- About 1 kHz event rate BBCAL
- About 12 days (on and off) running

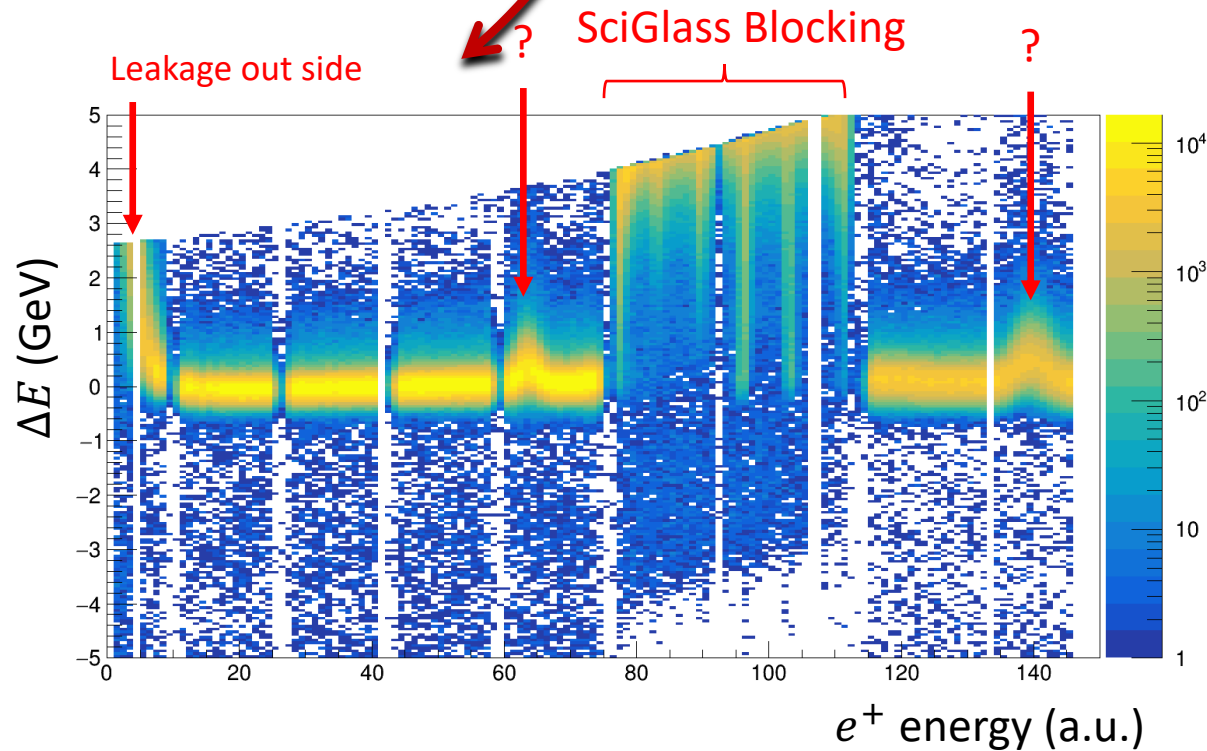
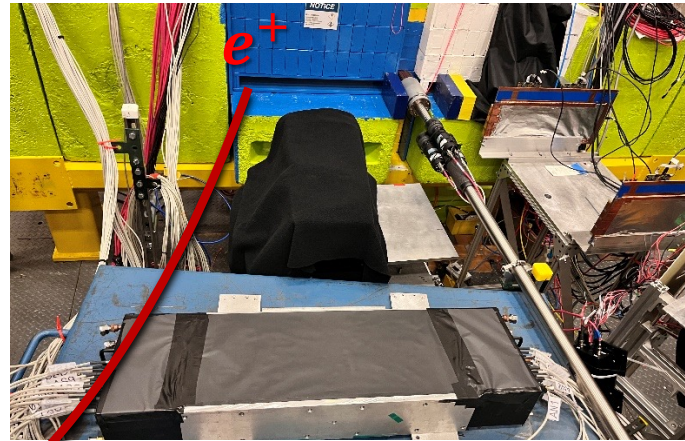
View from above



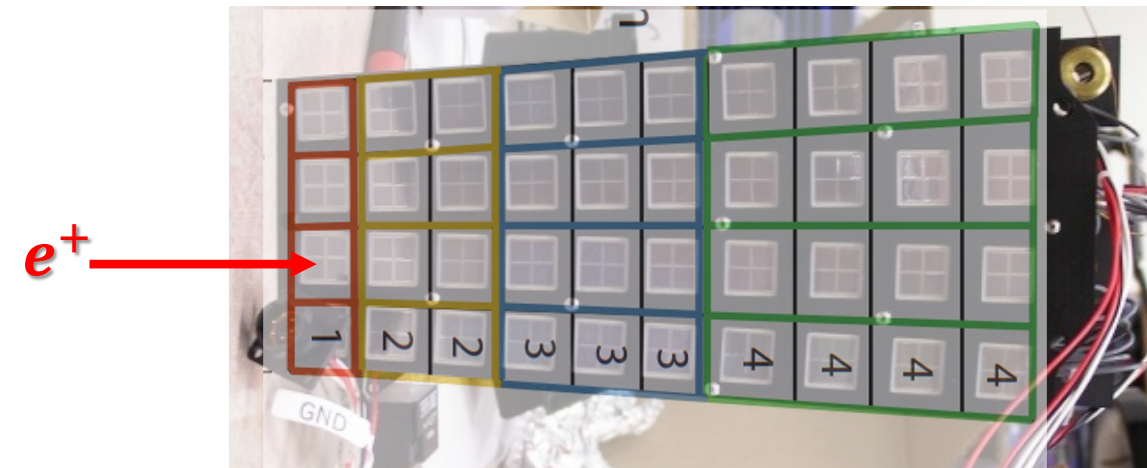


# Pair Spectrometer Setup

View from above



View from side



(summation for bottom channels 1, 2, 3, 4)



# Energy Calibration

Not calibrated prior! Fortunately  $e^+$  energy is known.

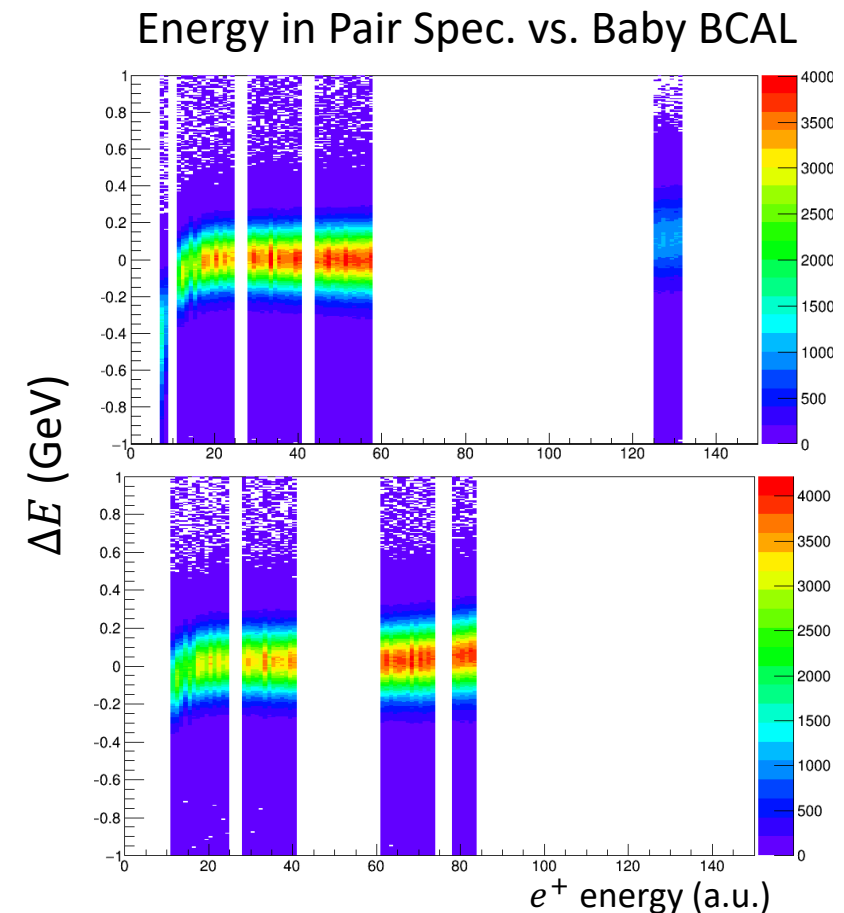
Can solve for (linear-only) gain factors  $c_k$

- Minimize  $F = \sum_i^{events} (E_i^{BCAL} - E^{PS})^2$

$E^{BCAL}$  is sum of north and south readouts

$$E^{BCAL} = \sum_j^{16} c_j (A_j^N + A_j^S) = \sum_j^{16} c_j A_j^{sum}$$

- Calculate  $c_k$  with  $\frac{\partial F}{\partial c_k} = 0$
- Math in backup slides





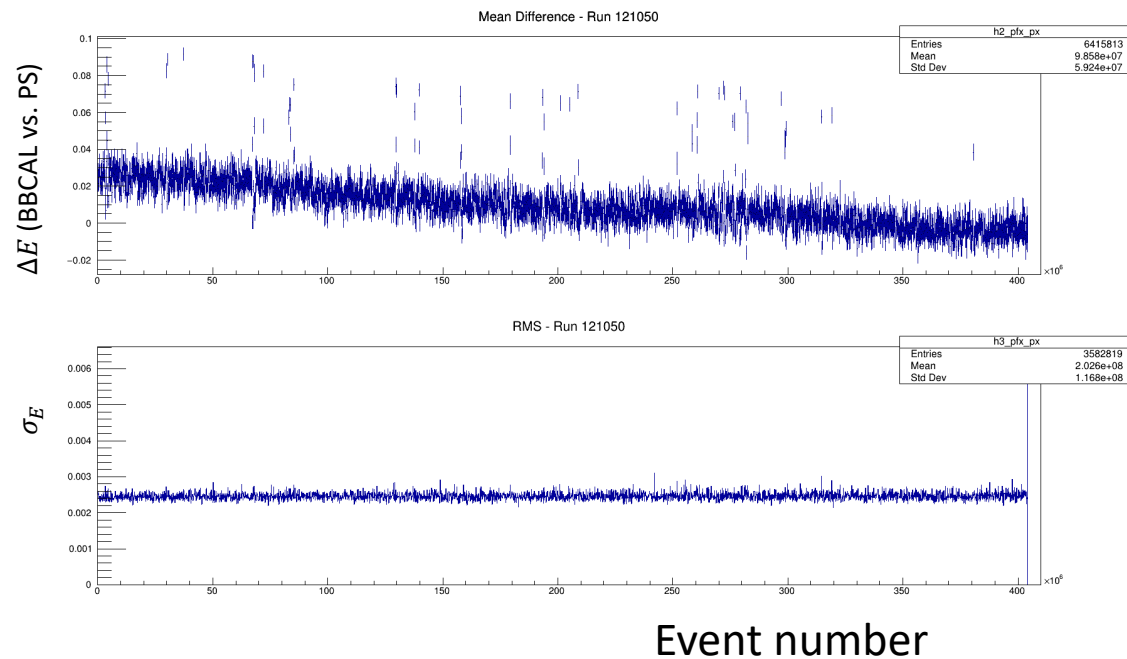
# Quality Checks

Within run (one run ~ 2 hours)

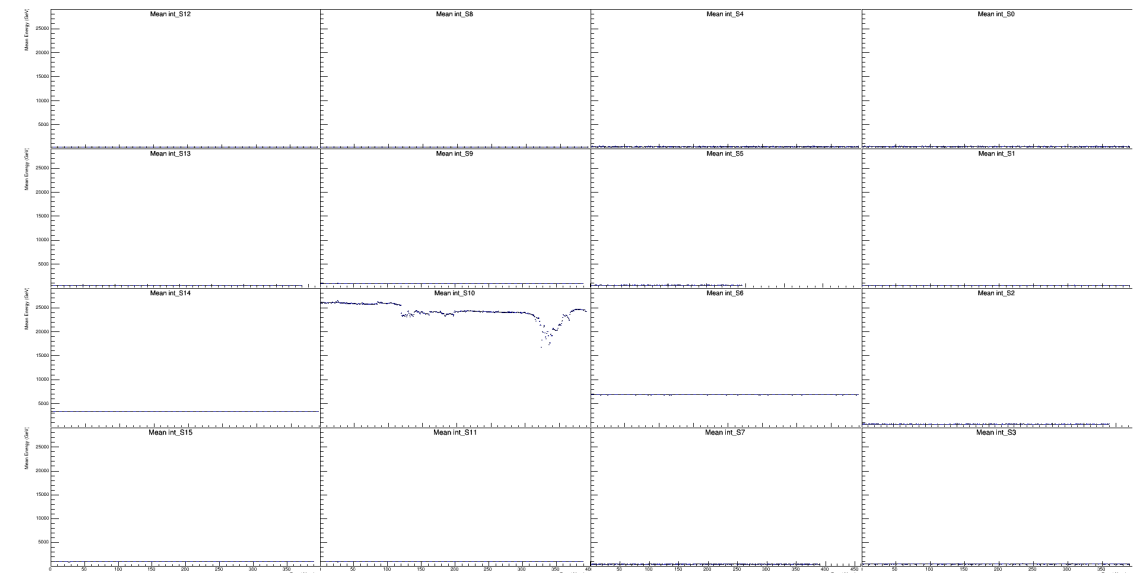
- Check individual channel stability
- Check overall energy stability

## Broad features:

- Channel S10 stopped working early in March 23 data
- Resolution fairly stable over ~ few hours
- Some slow drift in overall energy, maybe temperature related?

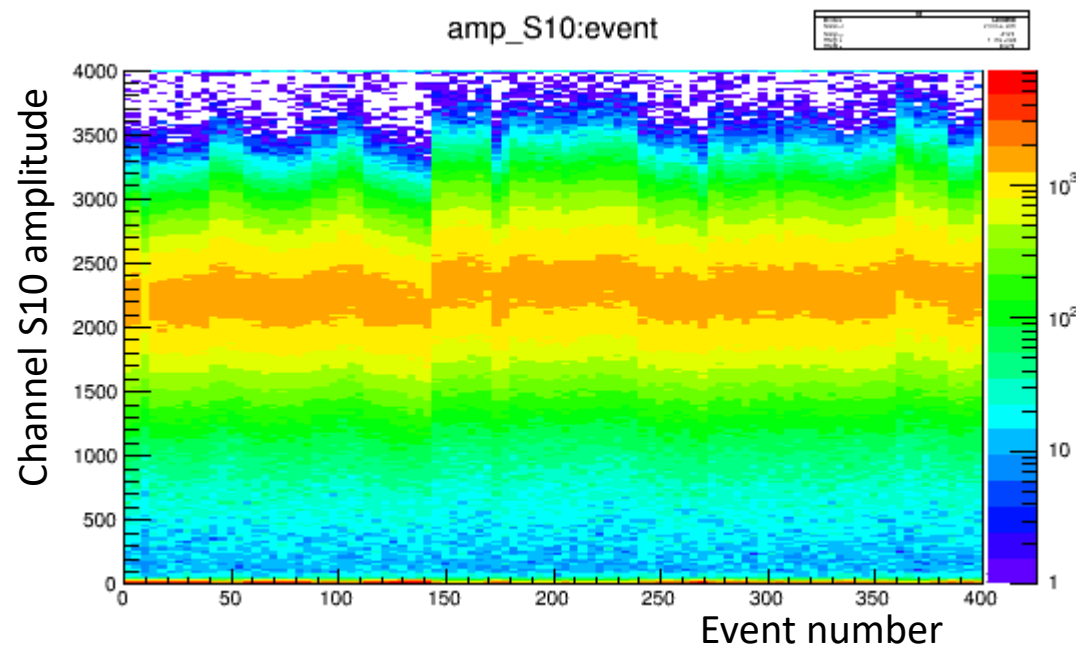


## Individual Channels

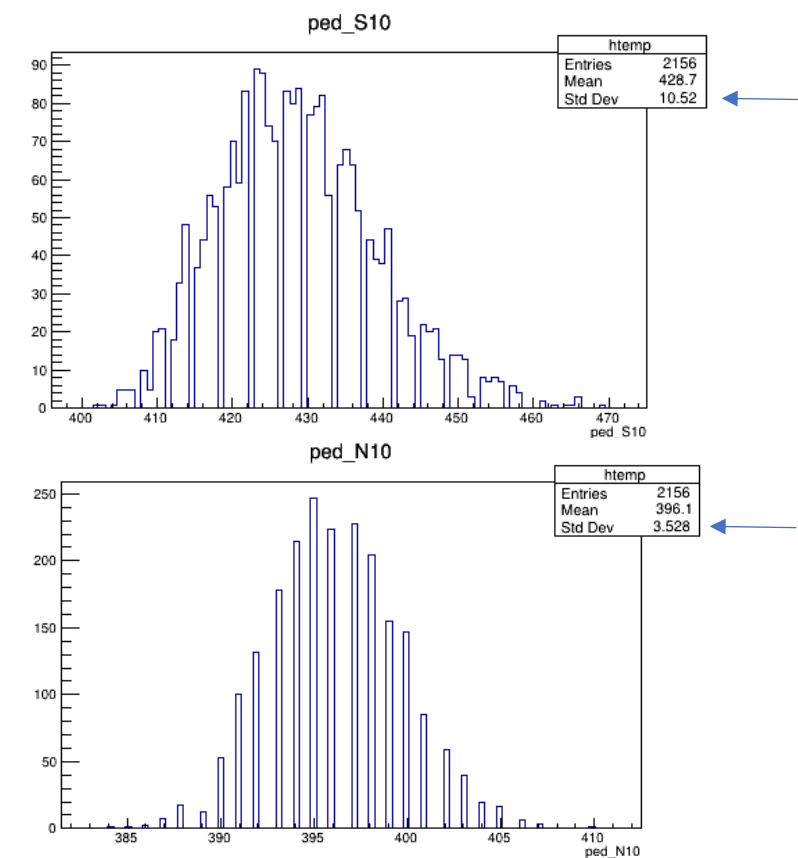


# Quality Checks, cont.

- Channel S10 demonstrates time-instability
- Also higher dark rate (pedestal RMS)



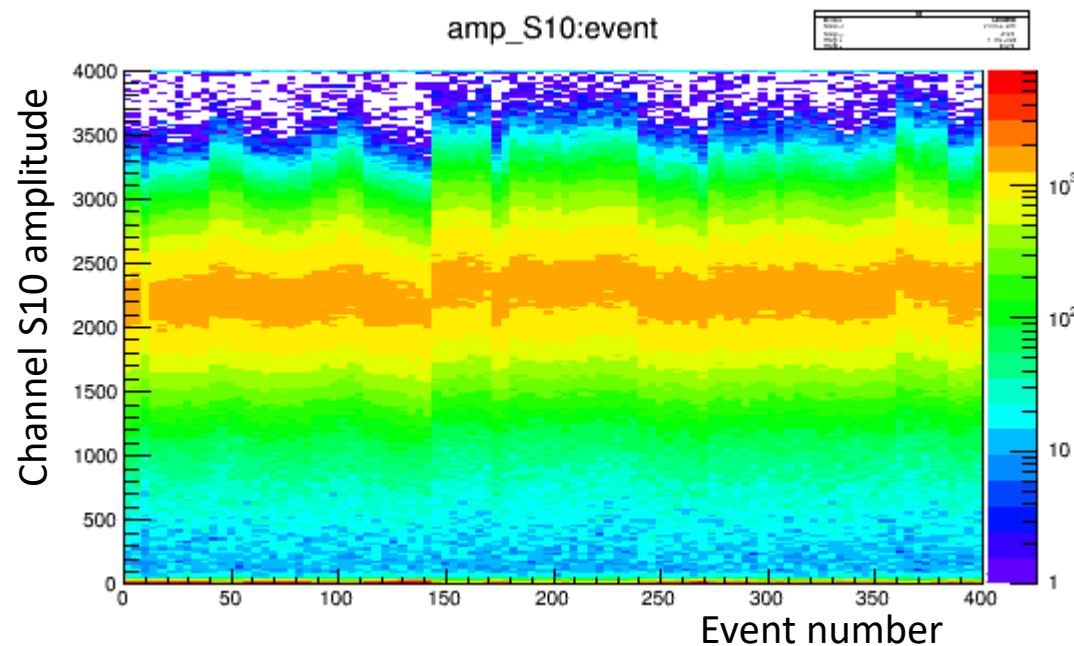
Measured Pedestal  
(4 sample sum)



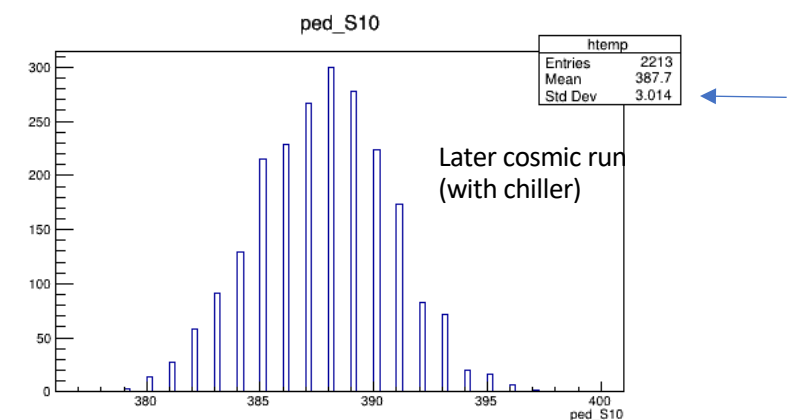
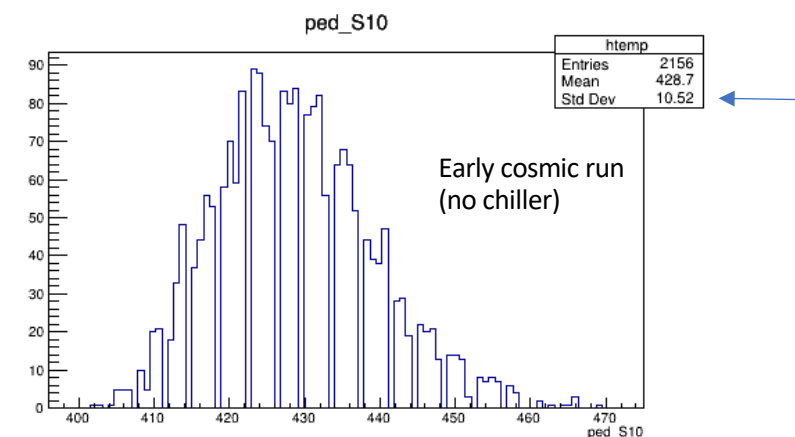


# Quality Checks, cont.

- Channel S10 demonstrates time-instability
- Also higher dark rate (pedestal RMS)
- Somehow, recovered when chiller was put in (late cosmic runs)



Measured Pedestal  
(4 sample sum)







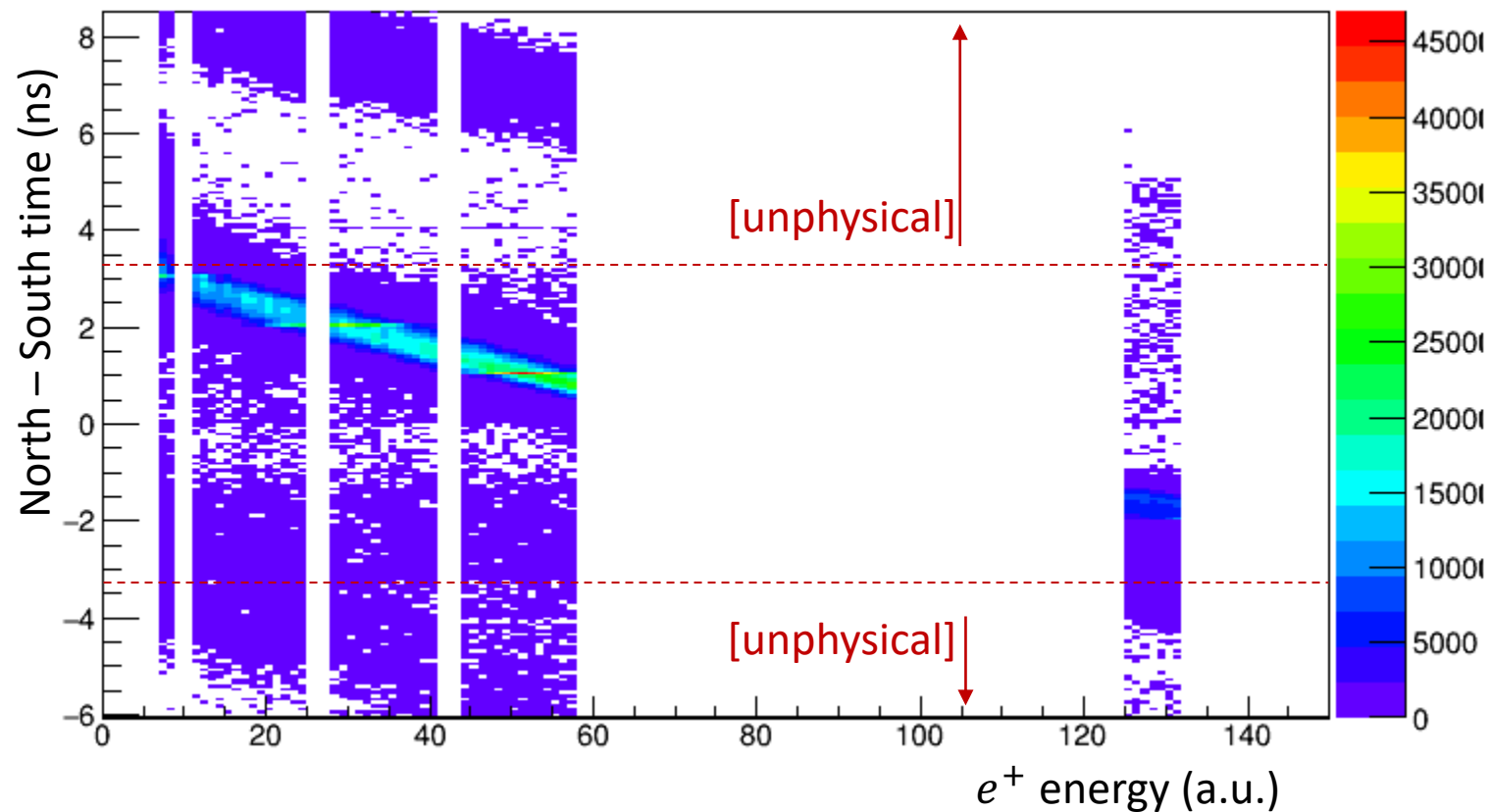
# Suspicious Channel Summary

- S10: noisy, time-dependent, bad dark rate
  - Good: earliest runs in March  
+ final cosmic runs (after chiller added)
  - Bad: everything in between
- S2 & S3: no signal in some cosmic runs
- Recommended checks:
  - Check ratio of N/S sides for each channel
  - Check pedestal RMS for dark rate

# Energy Leakage

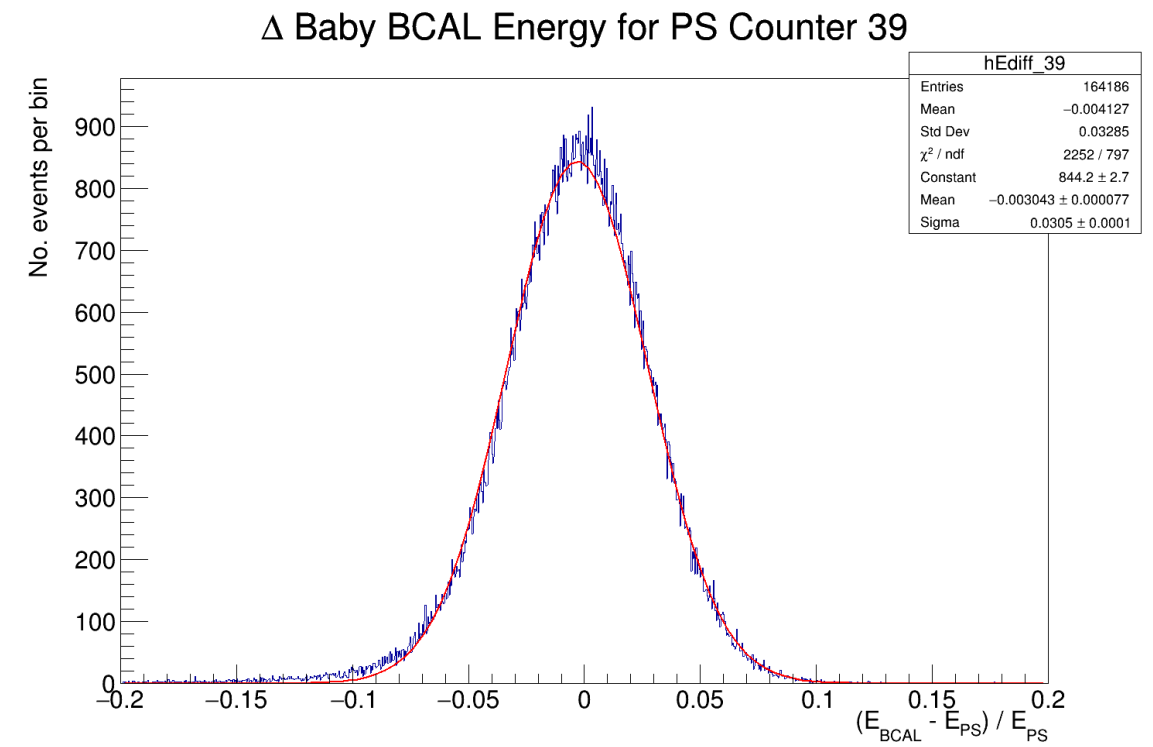
Timing difference confirms lowest energy  $e^+$  hits at edge of baby BCAL

Dashed lines: time for signal to travel full length (58 cm)



# Energy Resolution Extraction

- Fit for single PS counter  
(positrons energy  $\pm 10$  MeV at most)
- Fit to get  $\sigma_E/E$
- Repeat for all good PS counters

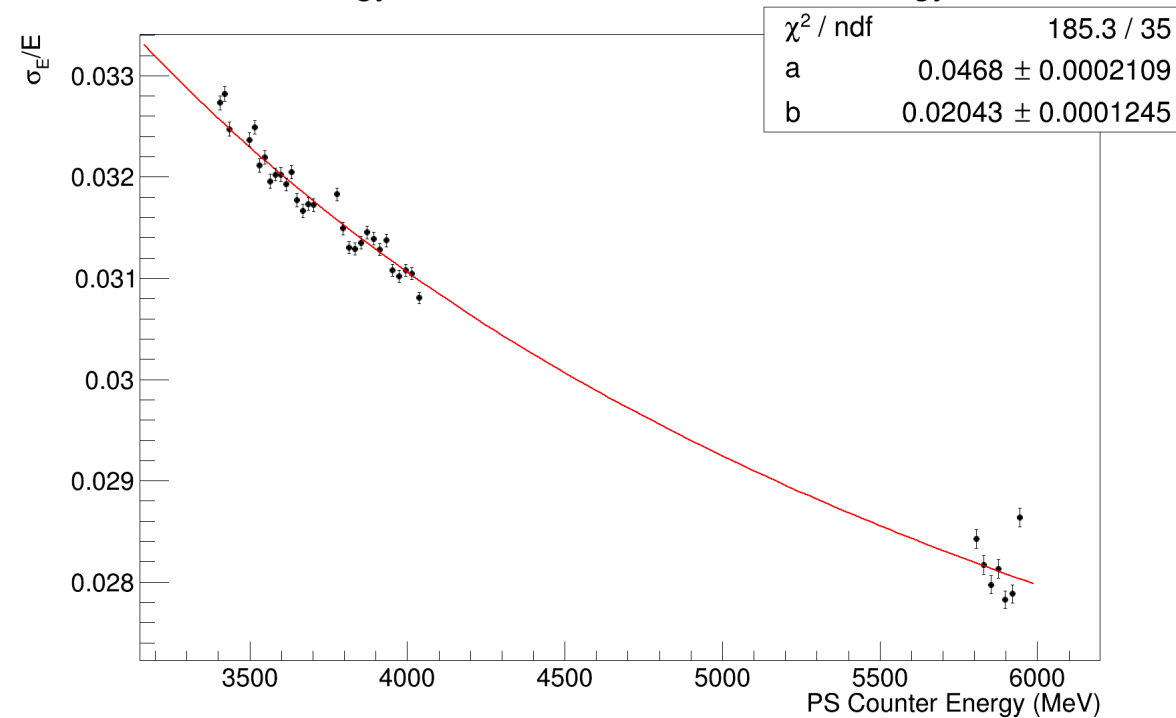


# Energy Resolution

Fit function

$$\frac{\sigma_E}{E} = \frac{a}{\sqrt{E} \text{ (GeV)}} \oplus b$$

Energy Resolution vs PS Counter Energy



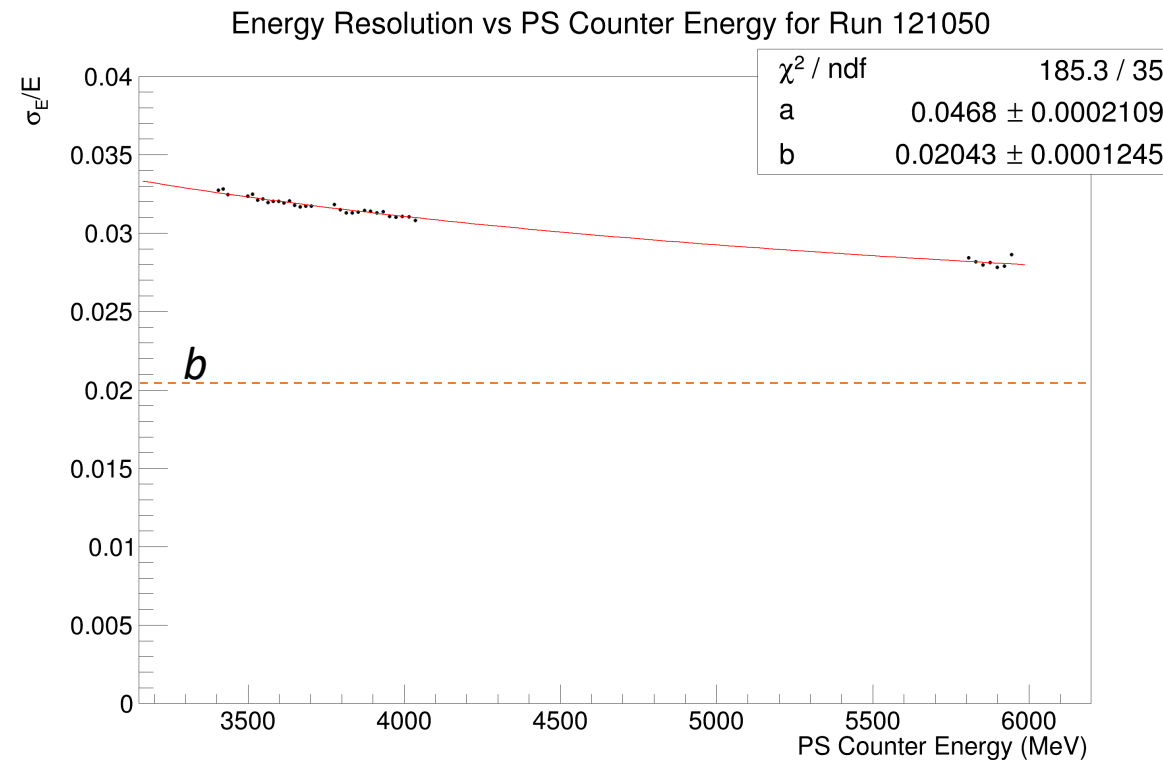




# Energy Resolution (zoomed out)

Fit function

$$\frac{\sigma_E}{E} = \frac{a}{\sqrt{E} \text{ (GeV)}} \oplus b$$



**BCAL NIM (2018):**

**a: 0.052**

**b: 0.036**

(measured via symmetric etas in  $\gamma p \rightarrow \eta \pi^+ \pi^- p$ )



# Energy Resolution

Run Number	$a$ (stochastic term, $\sqrt{\text{GeV}}$ )	$b$ (flat term, unitless)
121050	0.047	0.020
121051	0.047	0.020
121128	0.047	0.023
121129	0.047	0.024
121185	0.047	0.23
121186	0.047	0.024
121197	0.046	0.021
121199	0.047	0.021
121200	0.047	0.021
121201	0.047	0.021
121204	0.046	0.021
121206	0.047	0.021

Reasonably consistent across runs



# Number of Photoelectrons (per GeV)

- Form the ratio of energy measured on north/south sides
  - Reduces uncertainty due to sampling fraction
  - Poisson-type statistical process on each end
  - $N_{PE}$  extracted from fit

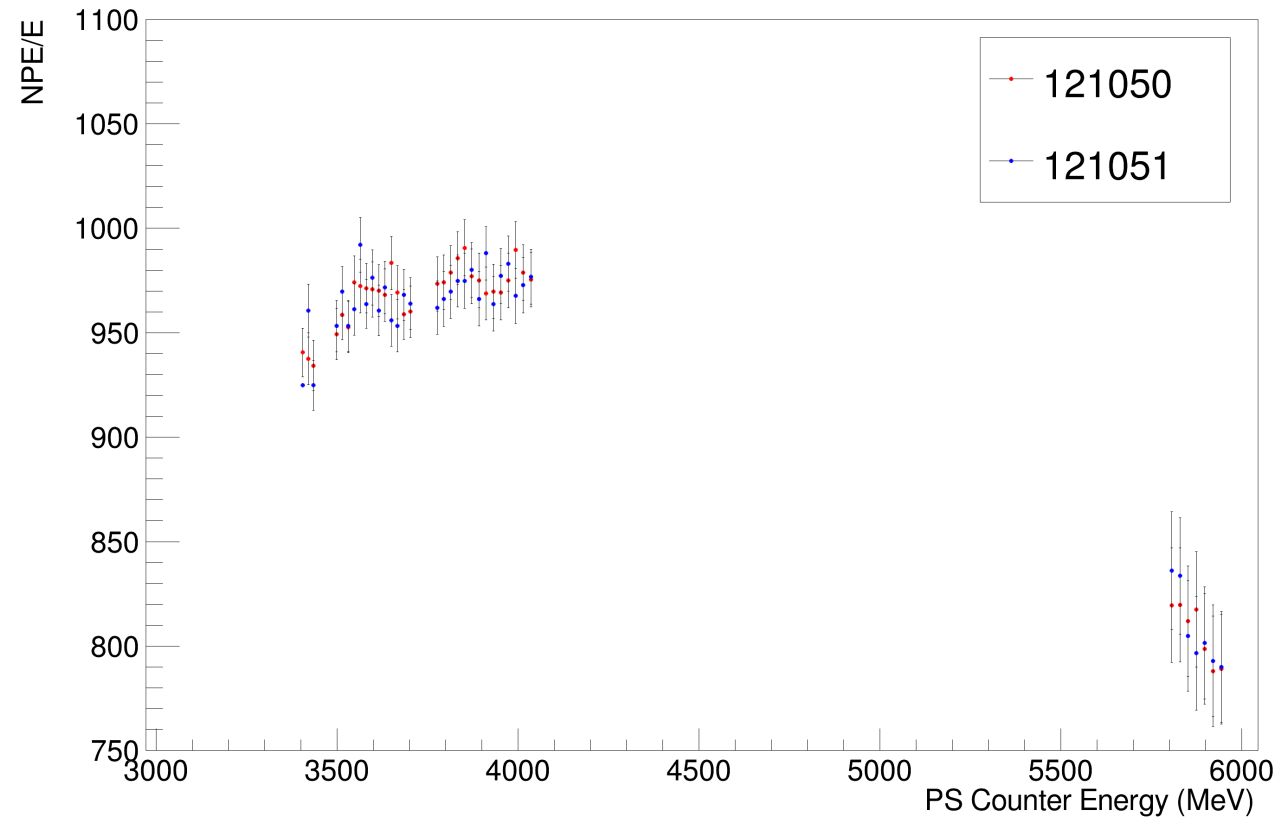
$$f(r) \sim \int P(x, N_{pe} \cdot \sqrt{R}) \cdot \frac{1}{r} P\left(\frac{x}{r}, \frac{N_{pe}}{\sqrt{R}}\right) \left[\frac{x}{r} dx\right]$$

$$P(x, N) = \frac{e^{-N} N^x}{\Gamma(x + 1)}$$



# $N_{PE}$ Over Positron Energy

NPE/E vs PS Counter Energy for Runs 121050 and 121051



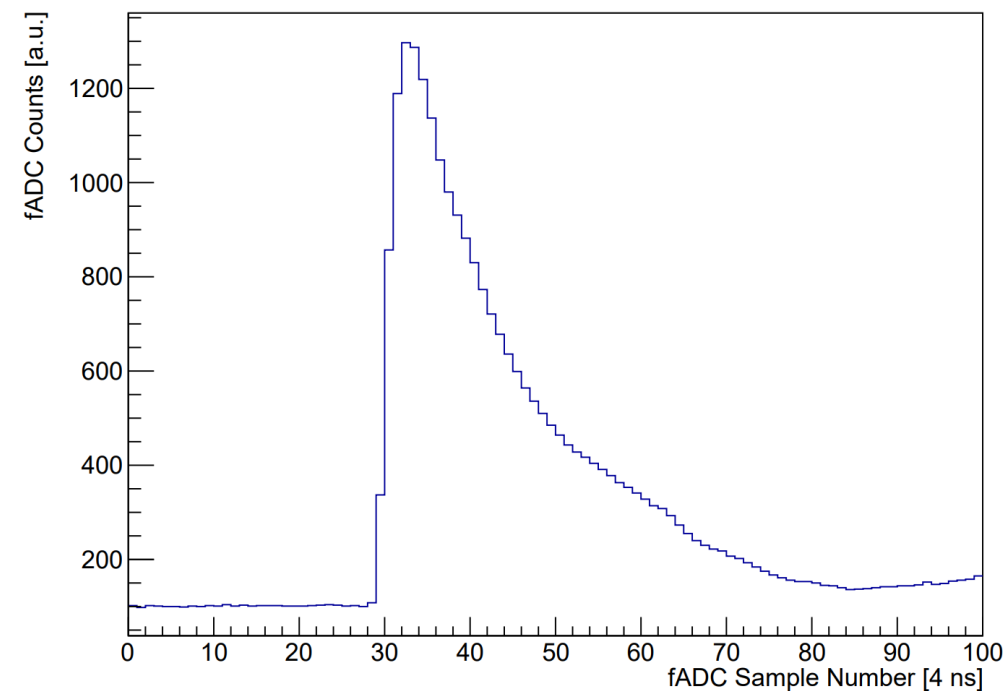
[S10 channel prevents us from easily extracting in later runs]





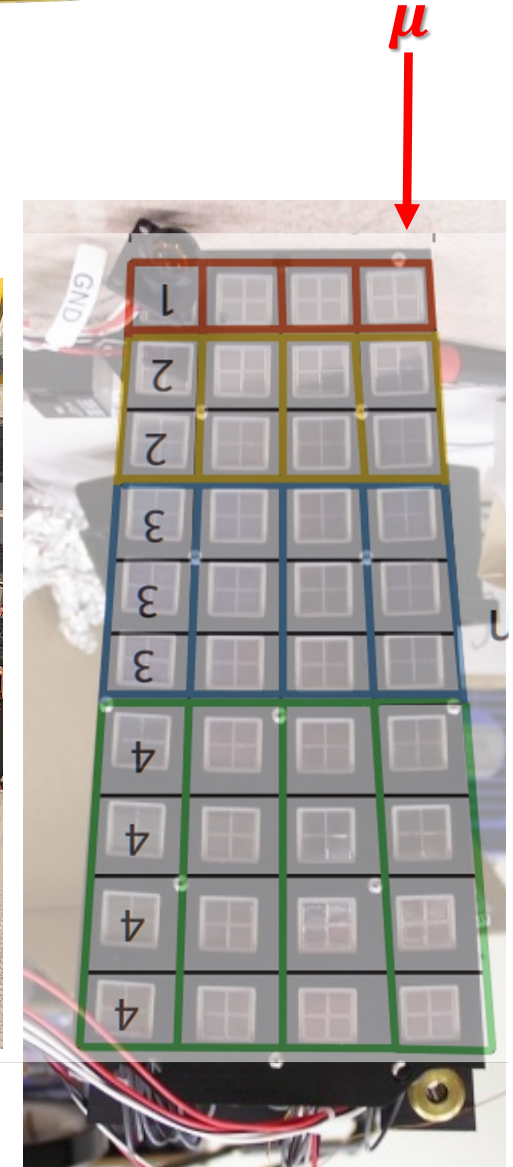
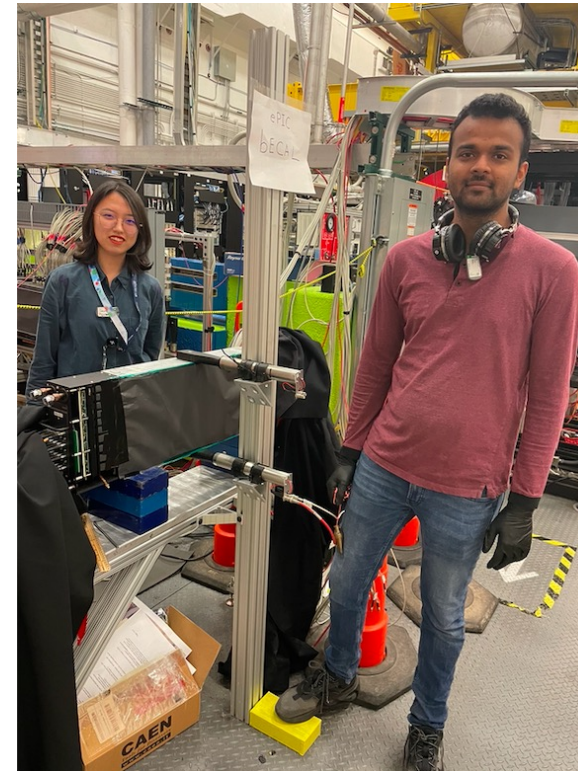
# Raw Waveforms

- A number of runs with full waveforms
- Could look for double pulses?
  - Might be rare, analysis challenge to isolate
  - May have enough data though



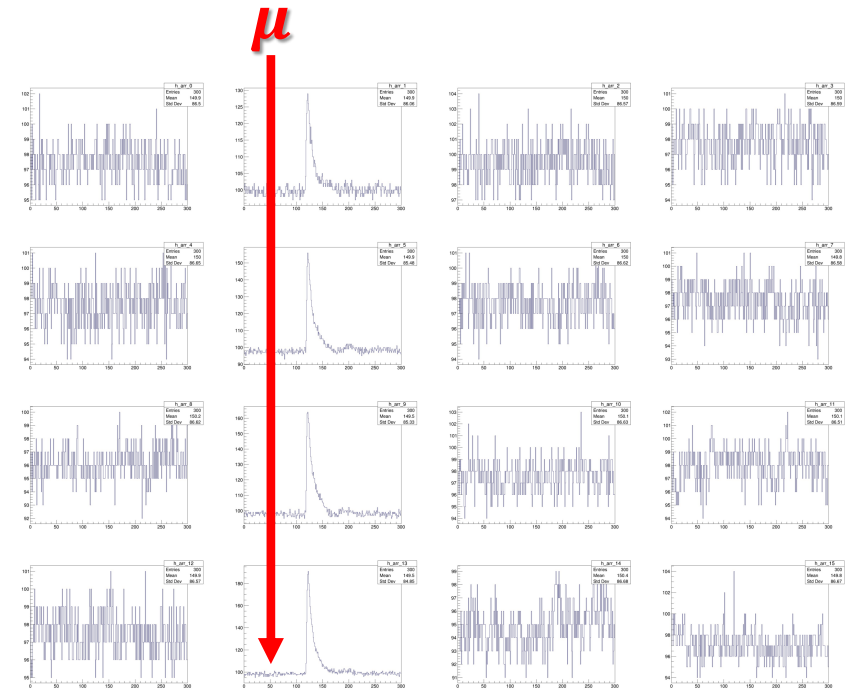
# Cosmic Runs

- Unfortunately,  $e^+$  always struck the same spot on the baby BCAL
- 99.5% of energy deposited in 12 channels
  - Easy gain determination
- Other 4 channels had very little energy deposition
  - Cosmics for gain determination?

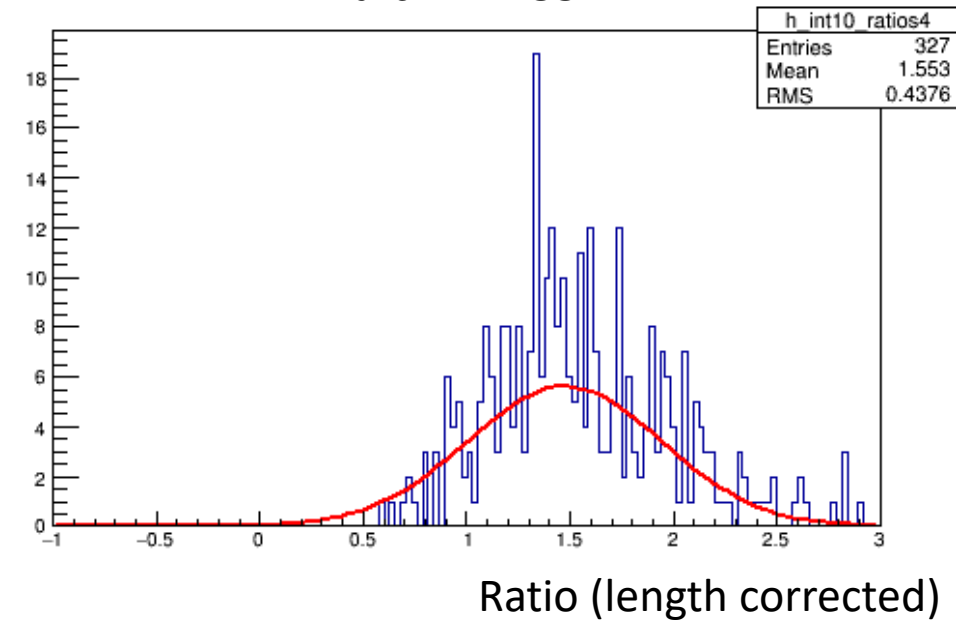


# Cosmic Runs, cont.

- Match voltage settings during March running
- Good cosmic events, but rate is extremely low
- Unclear if relative gains would be useful



South 6 / South 14  
Cosmics: **1.47**  
March: **1.55**





# Publication Plans

- $\frac{\sigma_E}{E}$  essentially ready
- $N_{p.e.}$  a few additional checks to perform, consistency looks good
- Anything else high priority?

Hope to have publication draft this summer



# Backup: Two Setups in JLab Hall D

## March $e^+$ Beam

- GlueX fADCs and DAQ
  - $e^+$  energy 3-6 GeV
- 
- Goal: resolution studies &  $N_{p.e.}$  extraction
  - Upstream hodoscope to measure  $e^+$  energy and trigger
  - Largely uncalibrated prior



## Fall Cosmics

- GlueX fADCs and DAQ
  - Trigger on cosmics
- 
- Goal: better gain determination for low occupancy channels
  - Scintillator paddles above/below provide trigger
  - Better geometric coverage for calibrations

- GlueX fADCs and DAQ
- $e^+$  energy 3-6 GeV

- Goal: resolution studies &  $N_{p.e.}$  extraction
- Upstream hodoscope to measure  $e^+$  energy and trigger
- Largely uncalibrated prior

# Backup: $e^+$ Beam in Hall D

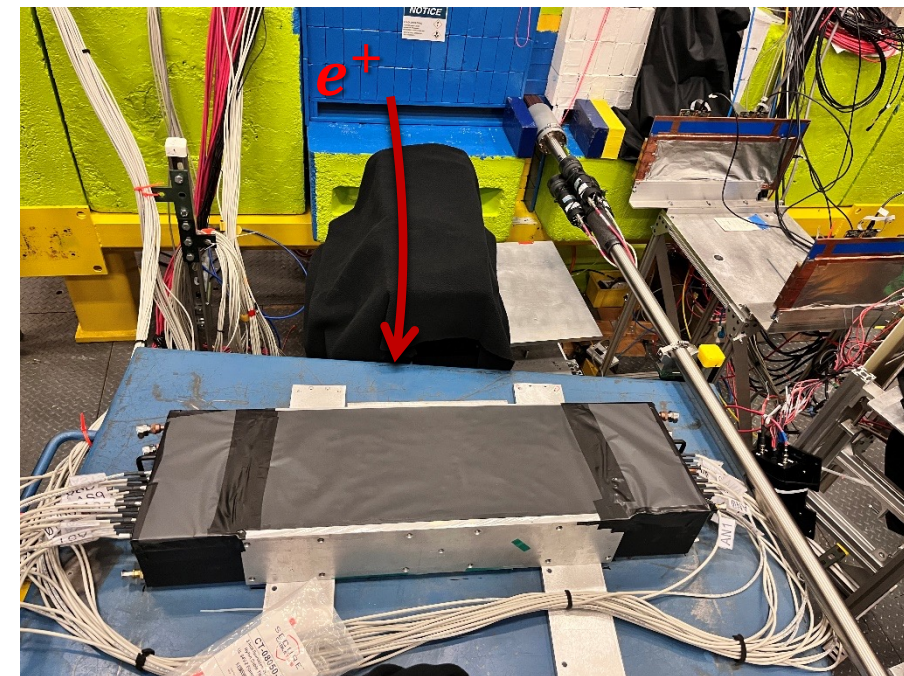
Good source of EM showers with known energy

Essentially identical to GlueX lead SciFi wedge, except 58 cm in length rather than 390 cm

Hall D pair spectrometer:

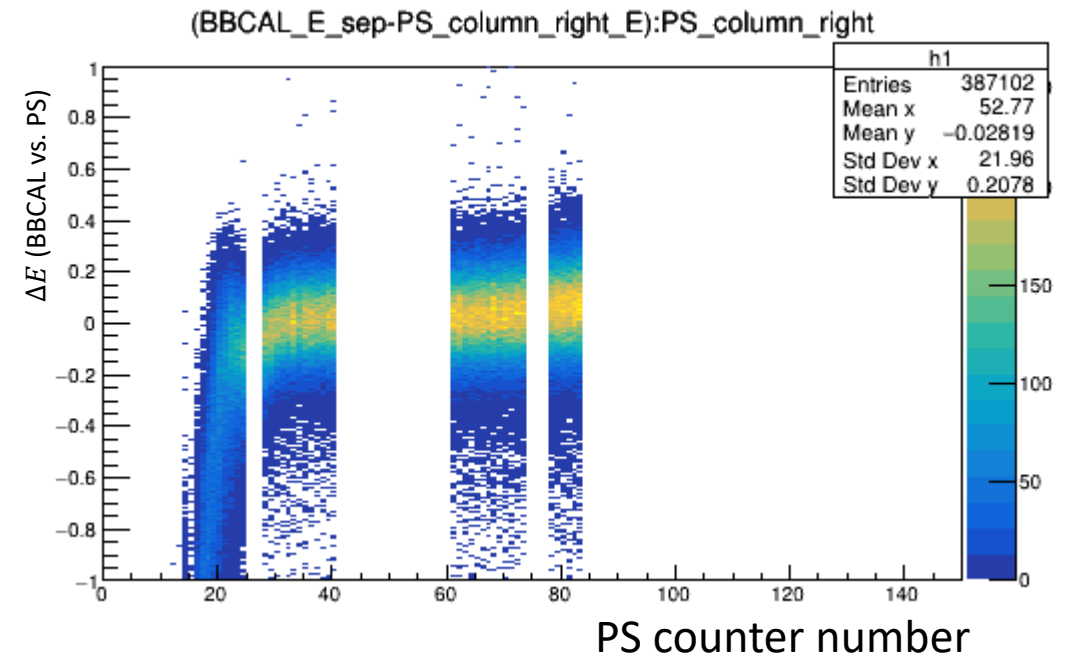
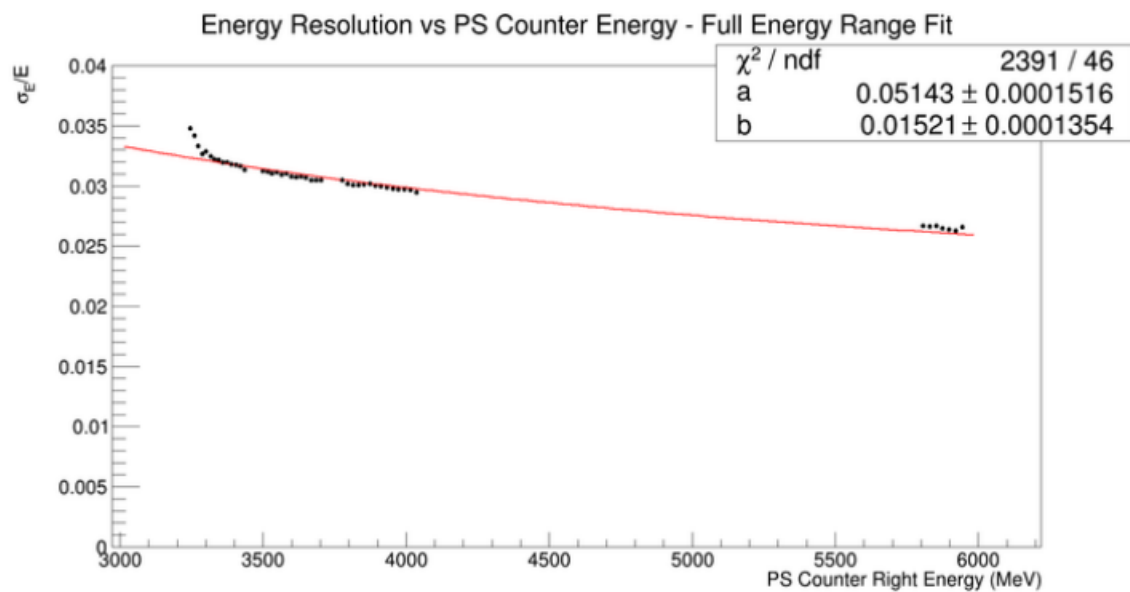
- $e^+$  energy roughly 3-6 GeV
- Unfortunately, SciGlass blocks portions
- About 1 kHz event rate BBCAL
- About 12 days (on and off) running

View from above





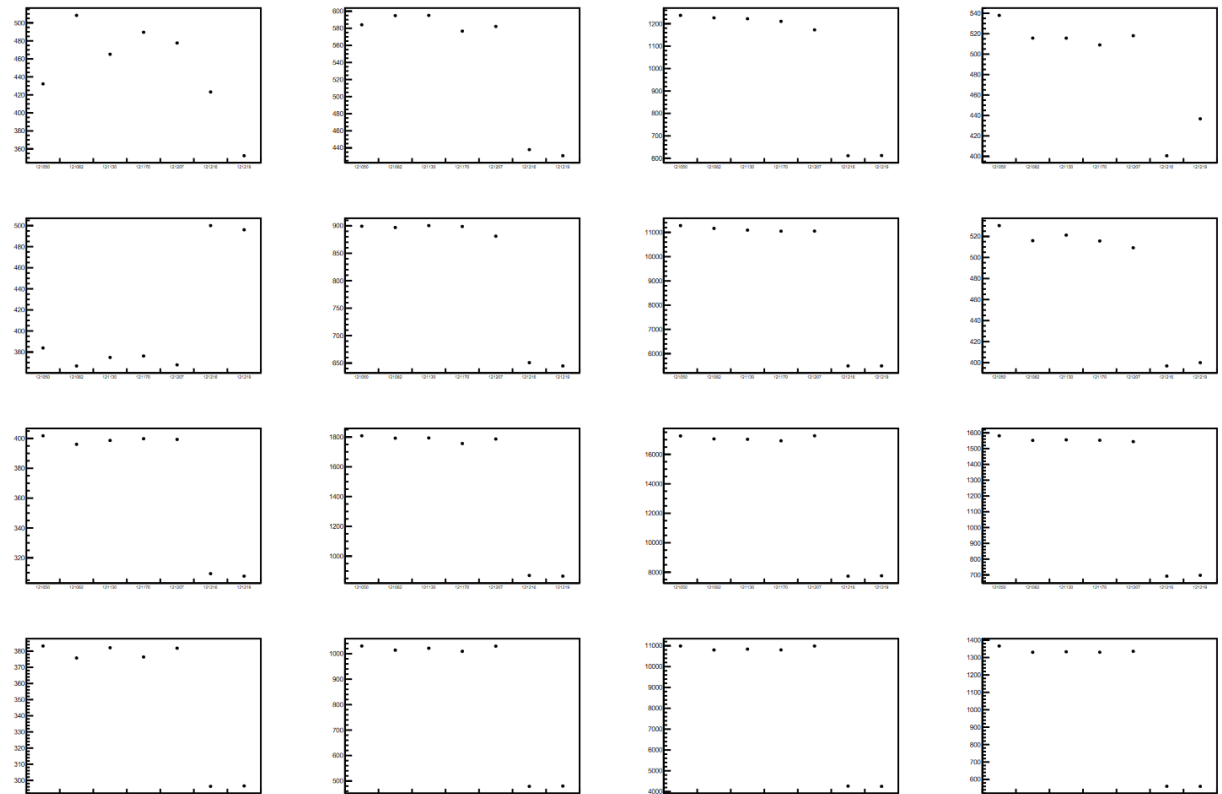
# Backup: More Energy Leakage Plots





# Backup: Longer Term Stability

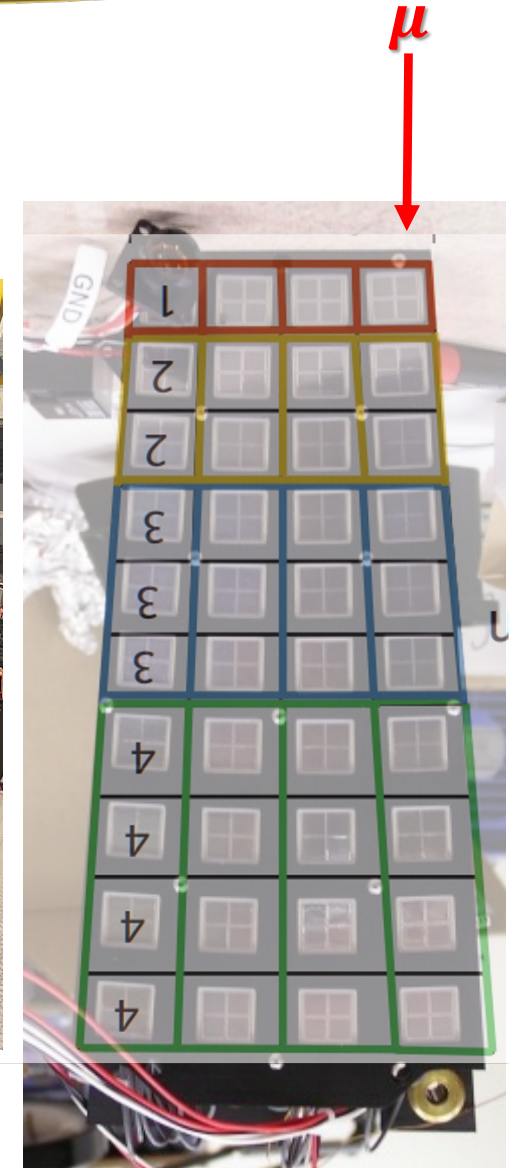
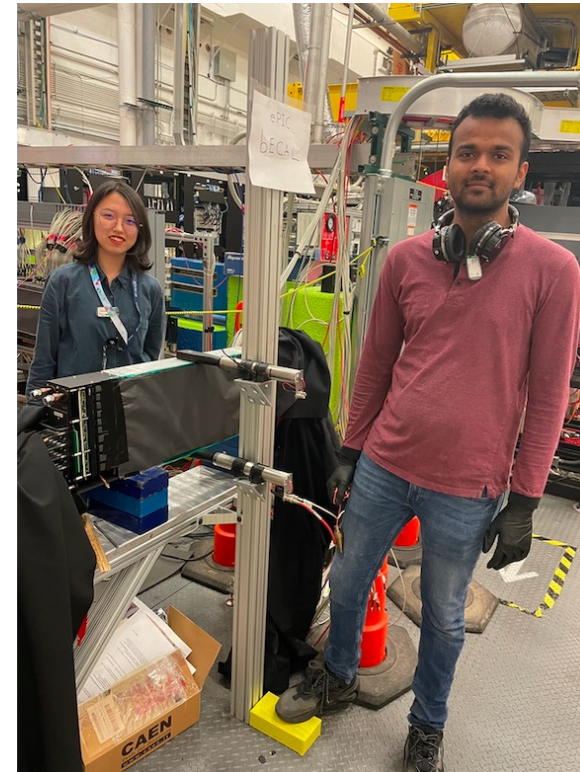
- Check individual channels over 12 day period
- Mostly checking for rad. damage
- Average fADC integral in un-shadowed region
- SiPM bias voltage change for last two points, jump expected





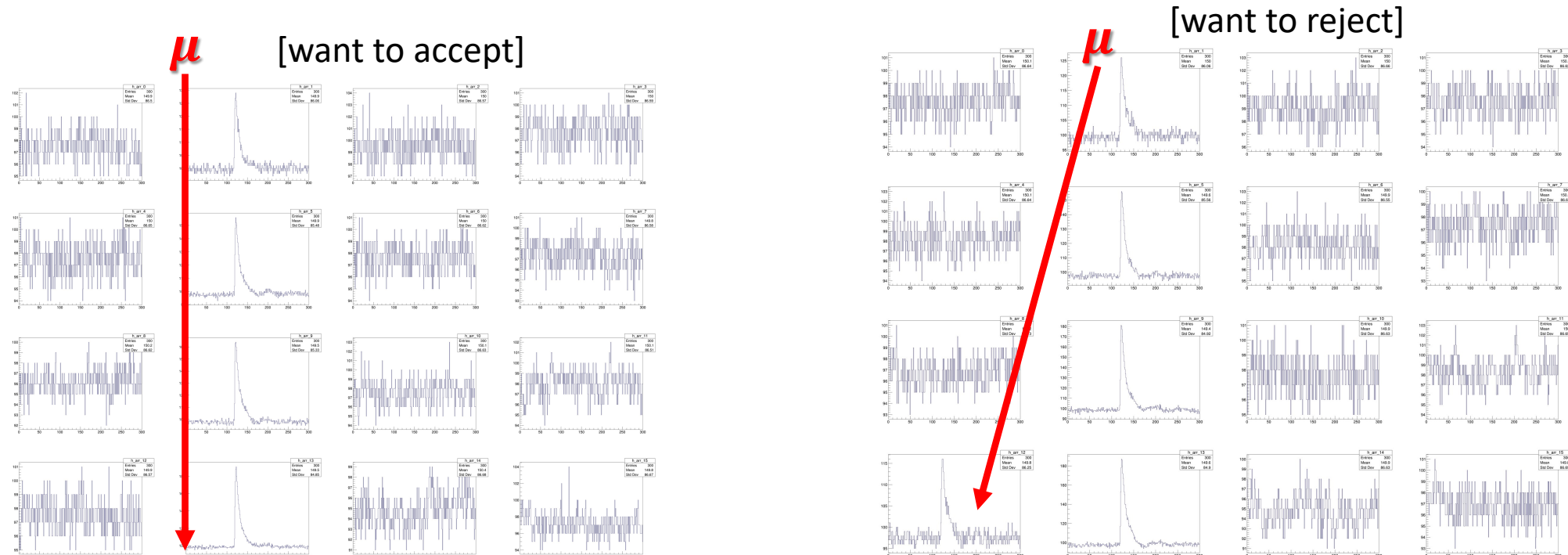
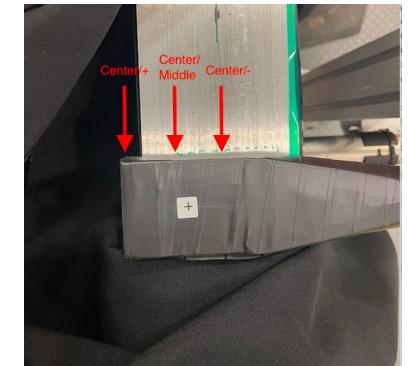
# Backup: Cosmic Setup Details

- Unfortunately,  $e^+$  always struck the same spot on the baby BCAL
- Very low occupancy in some channels
- Use MIPs to calibrate?
- Cosmic setup ran from August – October collecting data
  - Close to 30,000 muons collected



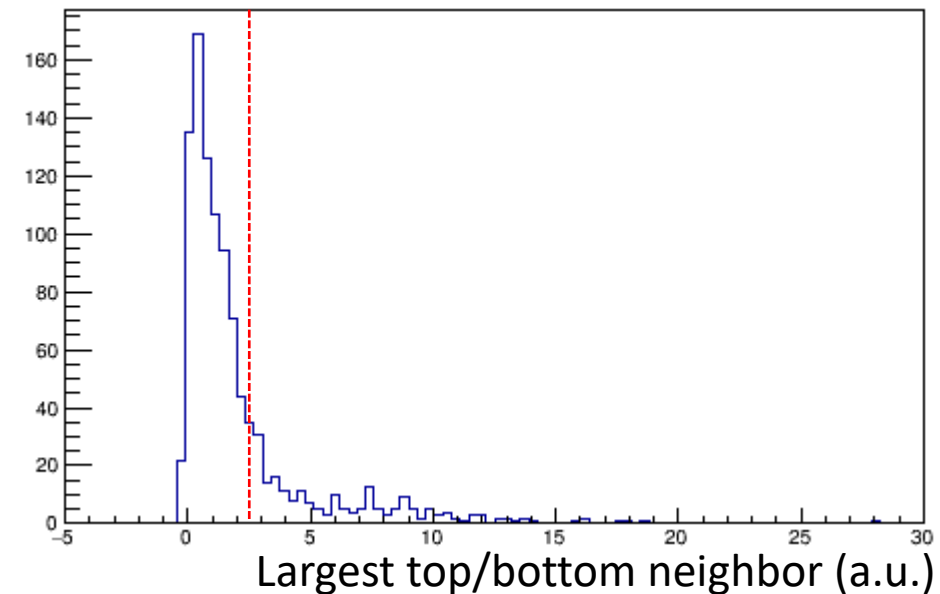
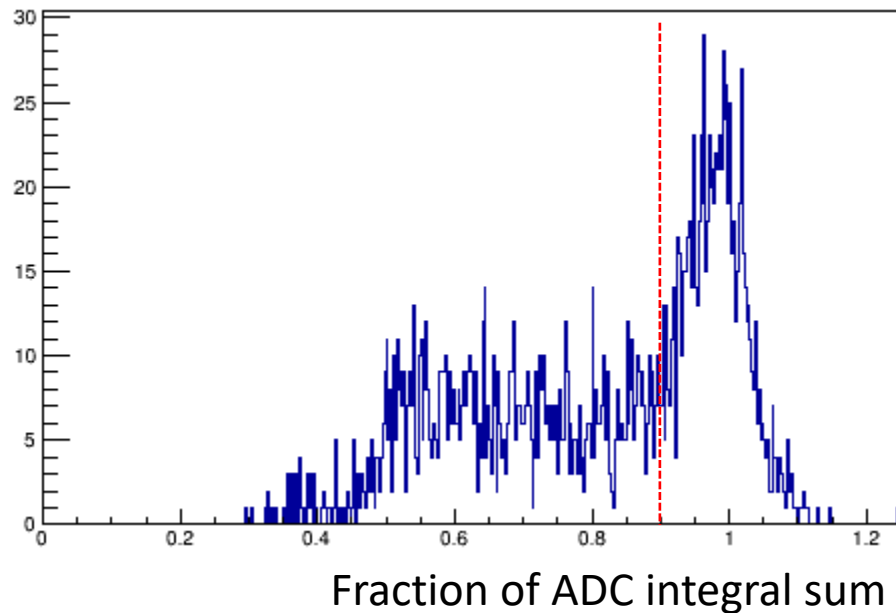
# Backup: Cosmic Analysis and Event Selection

- Match voltage settings during March running
- Paddles ensure that tracks are fairly vertical
- Can still improve with some event selection



# Backup: Cosmic Event Selection, cont.

- Find column with most deposition
  - Require  $> 90\%$
- Find largest neighboring top/bottom cell
  - Require this to be very small ( $< 2.5$  ADC units)
- Afterwards: less than  $\pm 1.5\%$  length traversed





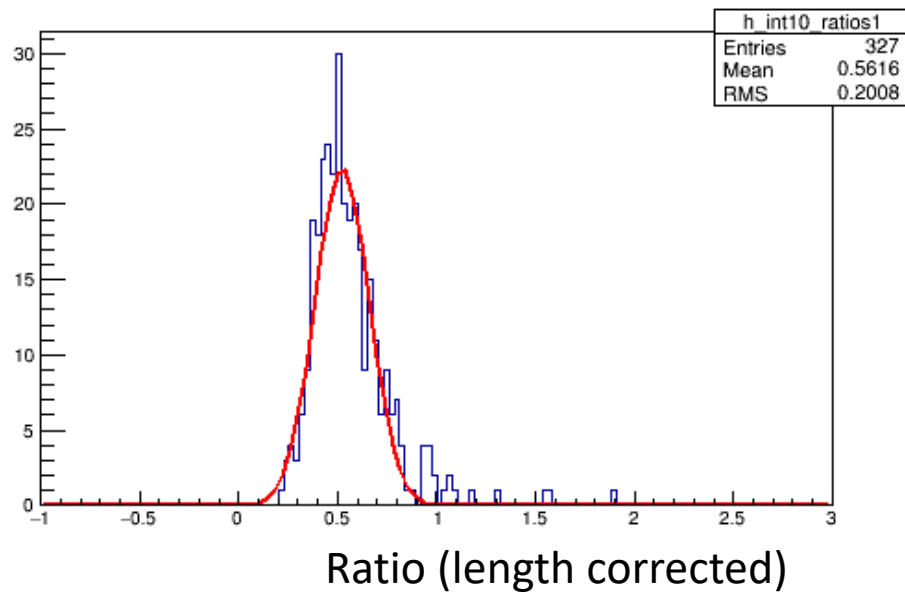
# Backup: Cosmic Initial Checks

- Check channels that were well determined in March run

North 6 / North 14

Cosmics: **0.524**

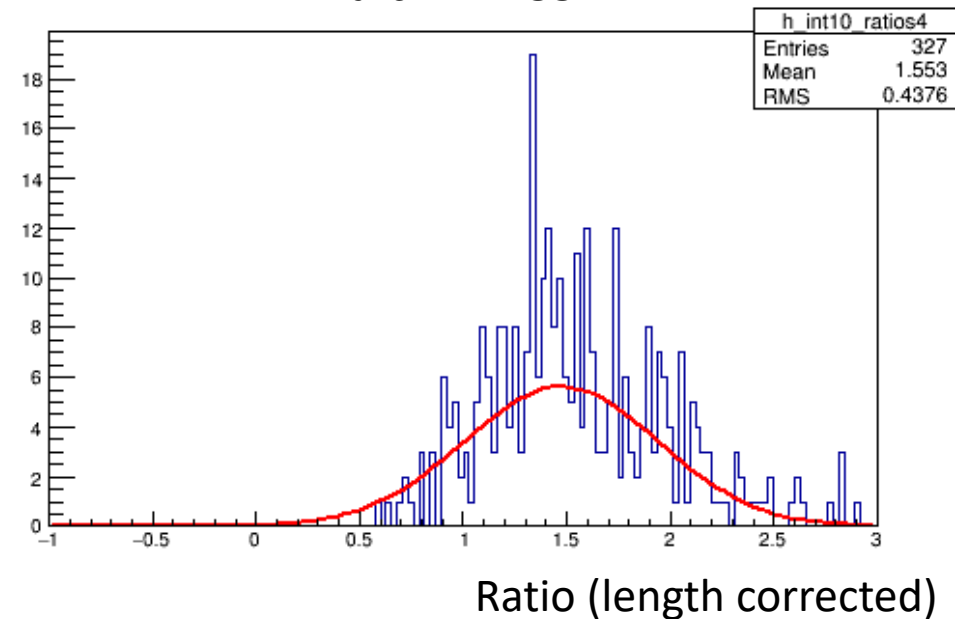
March: **0.733**



South 6 / South 14

Cosmics: **1.47**

March: **1.55**

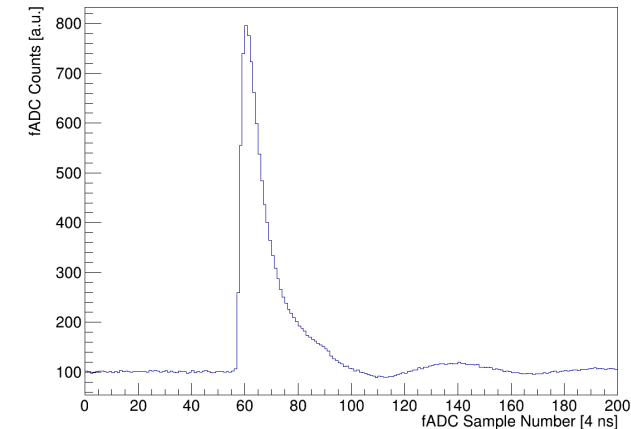




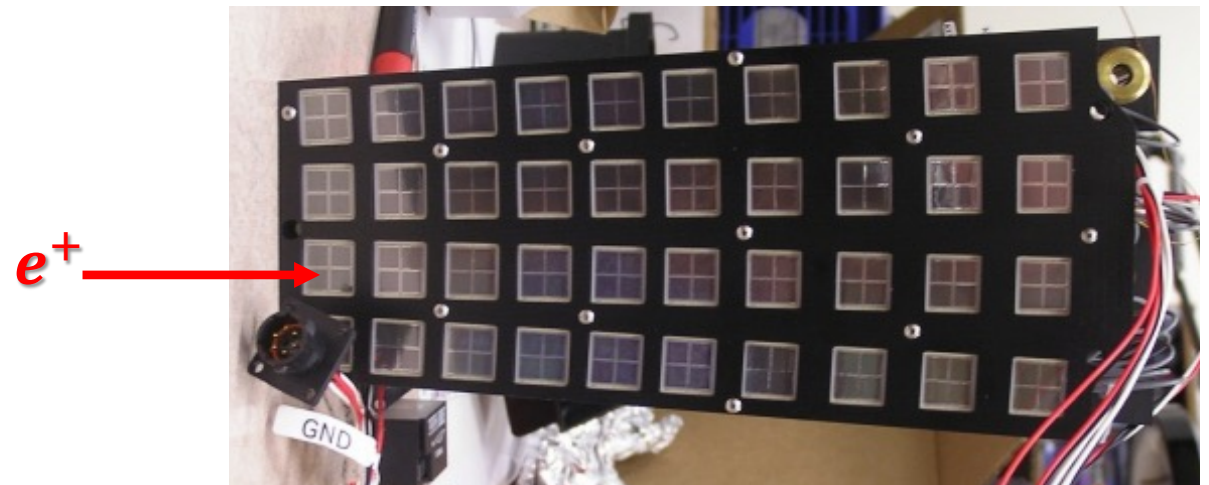
# Backup: Segmentation and Readout

## Flash ADC readout

- 250 MHz readout frequency (or every 4 ns)
- Up to 200 samples, if storing full waveform
- Some  $\sim 25$  sample integral used otherwise



View from side

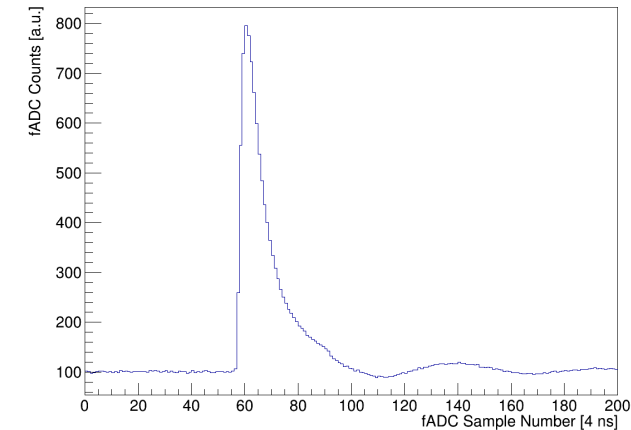




# Backup: Segmentation and Readout

## Flash ADC readout

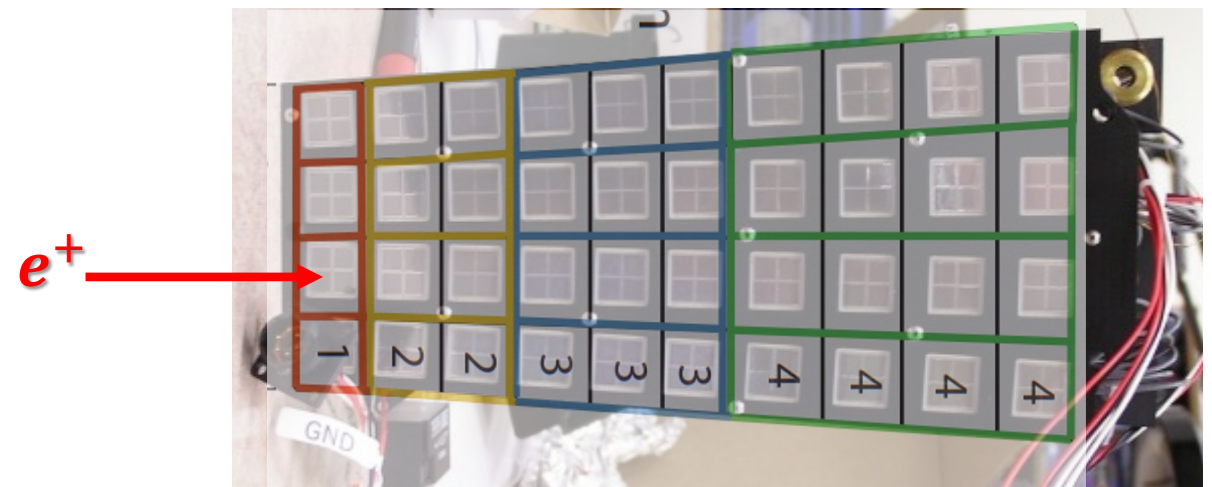
- 250 MHz readout frequency (or every 4 ns)
- Up to 200 samples, if storing full waveform
- Some  $\sim 25$  sample integral used otherwise



View from side

## Two-ended readout:

- 4 SiPM readouts up/down
- 10 SiPMs left/right, some summed pre-readout
- **16 fADC readouts per side** (32 in total)



(summation for bottom channels 1, 2, 3, 4)

- Strikes a little low of center
- Upper left & rightmost layers: very little energy deposited

$e^+$  



# Backup: Baby BCAL Gains

- Minimize  $F = \sum_i^{events} \left( \sum_j^{16} c_j A_{ij}^{sum} - E^{PS} \right)^2$
- $\frac{\partial F}{\partial c_k} = 0 = 2 \sum_i^{events} A_{ik}^{sum} \left( \sum_j^{16} c_j A_{ij}^{sum} - E^{PS} \right)$
- $\Rightarrow \sum_i^{events} A_{ij}^{sum} E^{PS} = \sum_i^{events} A_{ij}^{sum} \left( \sum_j^{16} c_j A_{ij}^{sum} - E^{PS} \right)$



# Backup: Baby BCAL Gains

- $\sum_i^{events} A_{ik}^{sum} E^{PS} = \sum_i^{events} A_{ik}^{sum} \left( \sum_j^{16} c_j A_{ij}^{sum} \right)$

- Define vectors  $\mathbf{A}_i = \begin{bmatrix} A_{i0}^{sum} \\ \dots \\ A_{i15}^{sum} \end{bmatrix}$  and  $\mathbf{C} = \begin{bmatrix} c_0 \\ \dots \\ c_{15} \end{bmatrix}$

- Then in matrix form  $\sum_i E_i^{PS} \mathbf{A}_i = \left( \sum_i \mathbf{A}_i \mathbf{A}_i^T \right) \mathbf{C}$

- Now define  $\mathbf{W} = \sum_i E_i^{PS} \mathbf{A}_i$ , and  $\mathbf{Z} = \left( \sum_i \mathbf{A}_i \mathbf{A}_i^T \right)$

- $\Rightarrow \mathbf{W} = \mathbf{Z} \mathbf{C}$

- Solving for gain constants  $\mathbf{C}$ :

- $\mathbf{C} = \mathbf{Z}^{-1} \mathbf{W}$



# Backup: Baby BCAL Gains

- Determine separate north/south

- Let  $f_j = \frac{\sum_i A_j^N}{\sum_i (A_j^N + A_j^S)} \Rightarrow (1 - f_j) = \frac{\sum_i A_j^S}{\sum_i (A_j^N + A_j^S)}$

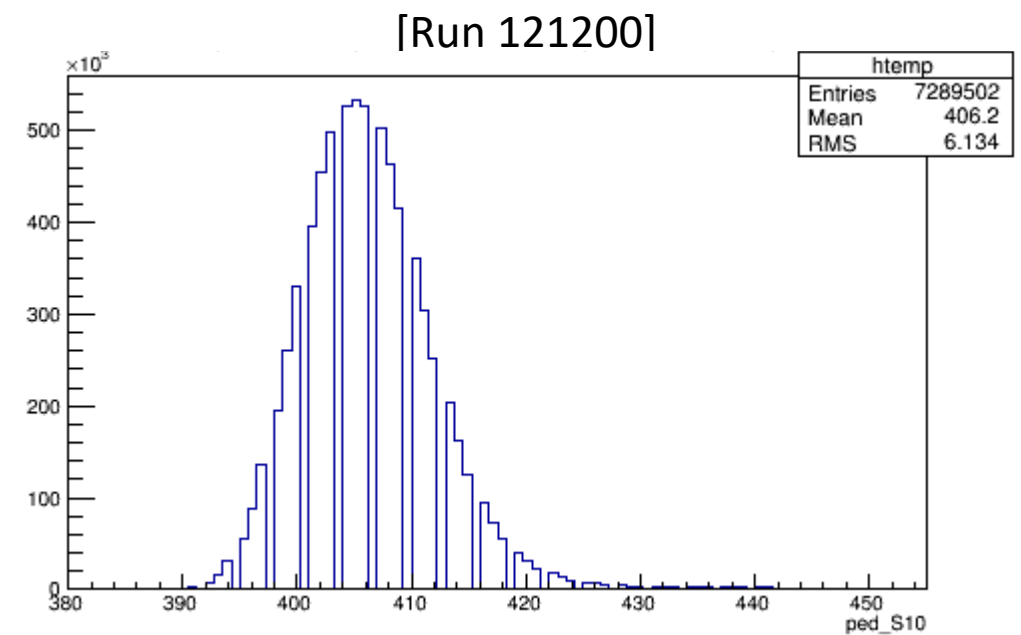
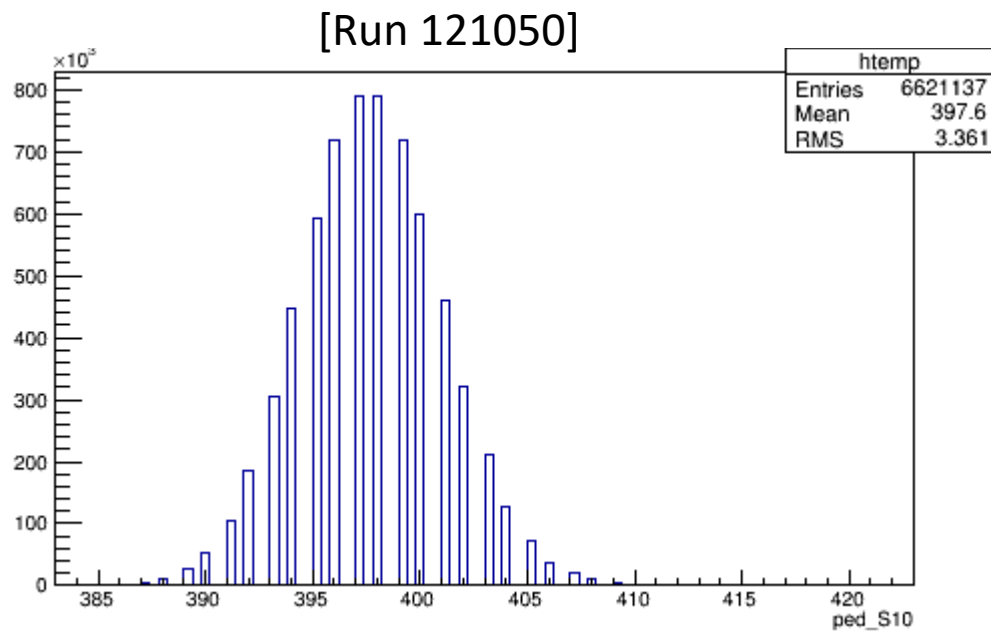
- Final gain factors:  $c_j^N = \frac{c_j}{2f_j}$  and  $c_j^S = \frac{c_j}{2(1-f_j)}$





# Backup: Channel S10 Degrading over March Runs

## Measured Pedestal (4 sample sum)



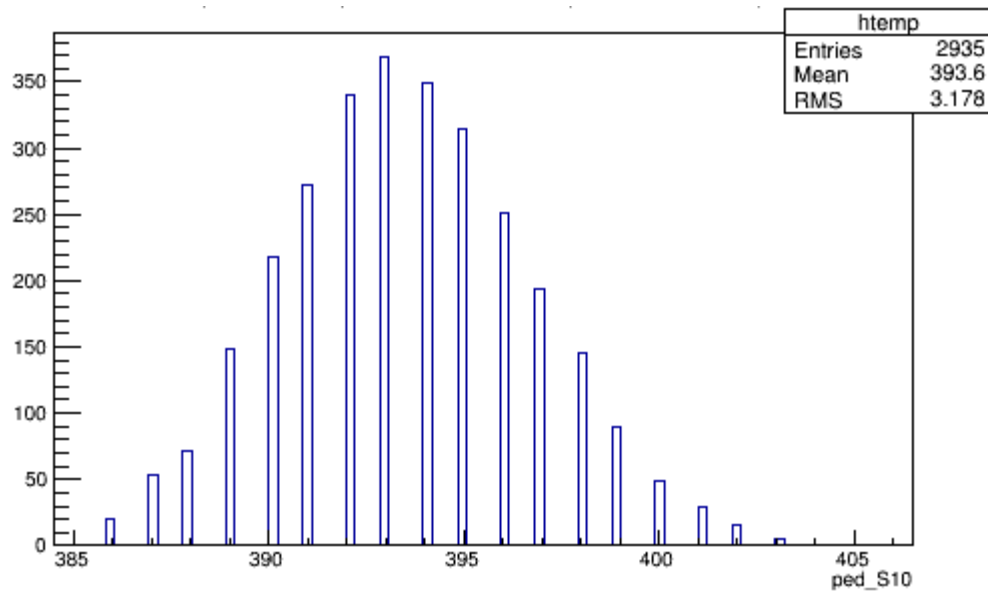


# Backup: Channel S10 in Cosmic Data

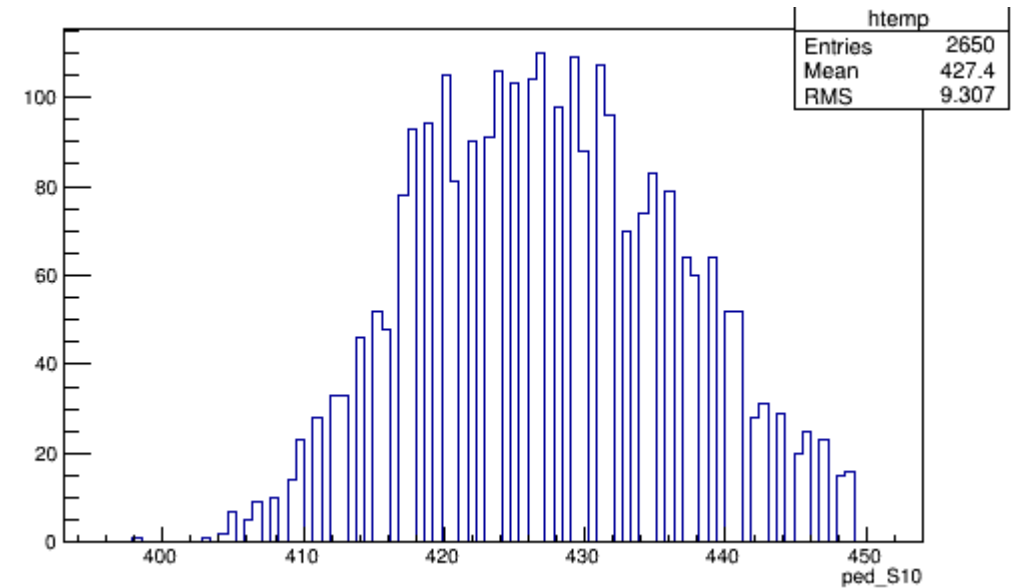
Before chillers

Measured Pedestal  
(4 sample sum)

[Run 121303]



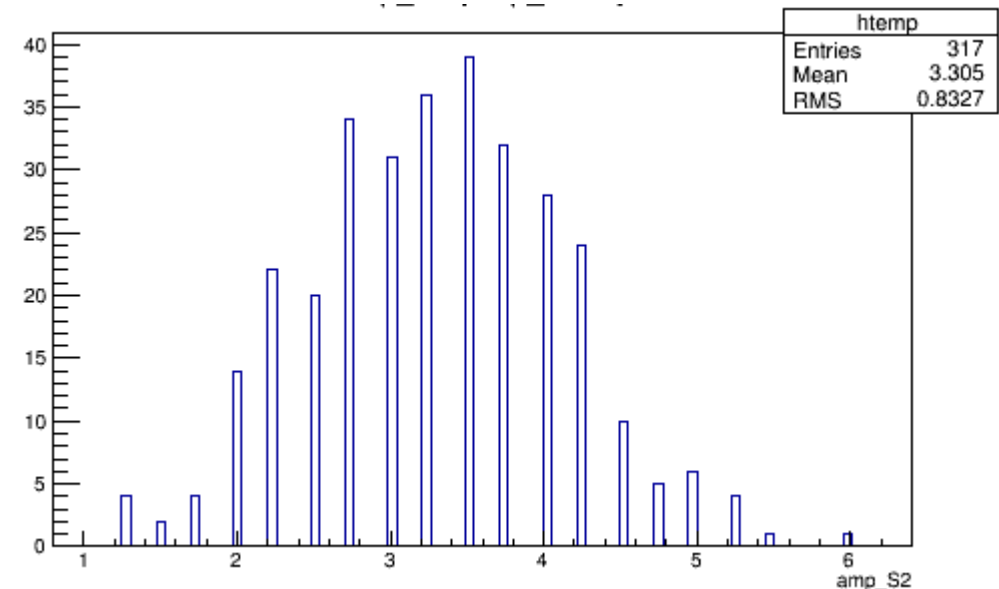
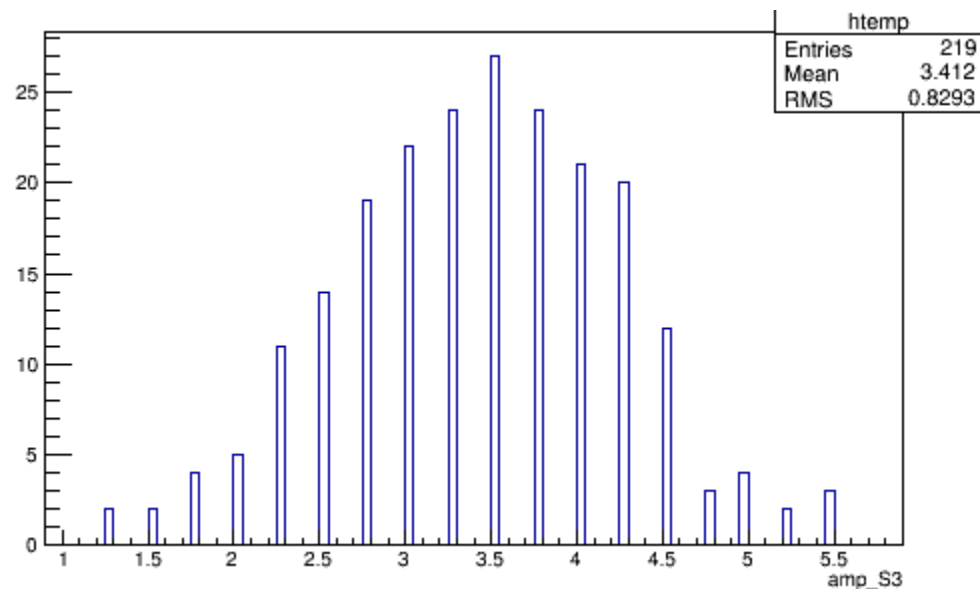
[Run 121304]



# Backup: Also curious

## 99.9% of the time:

- Either S10 is in a bad state
  - (pedestal too large, integral too high, varies with time)
- Or S2 & S3 are BOTH in bad state
  - Looks like just fADC noise, nothing from BBCAL

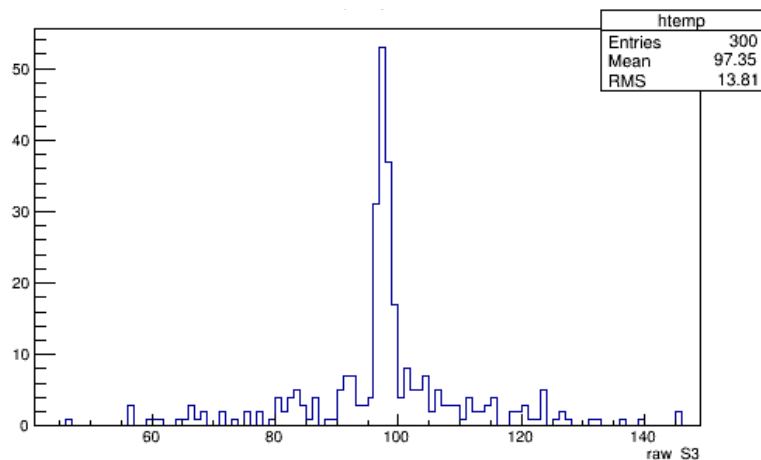




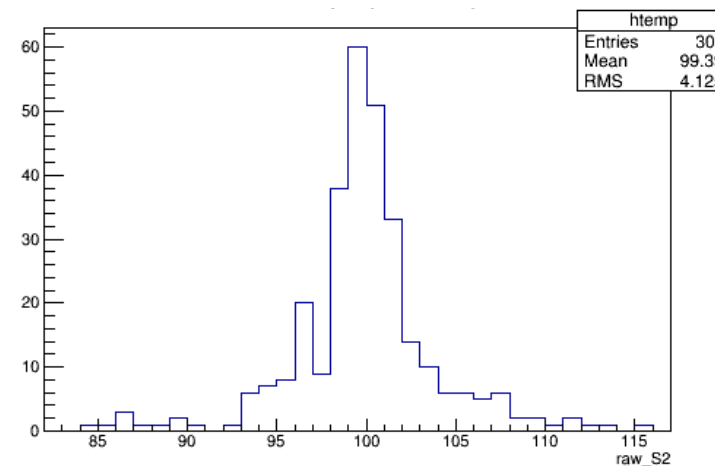
# Backup: Exactly Two Events

After adding chillers, found two events where S2 and S3 seem to read out above pedestal

[Run 121316]



[Run 121318]

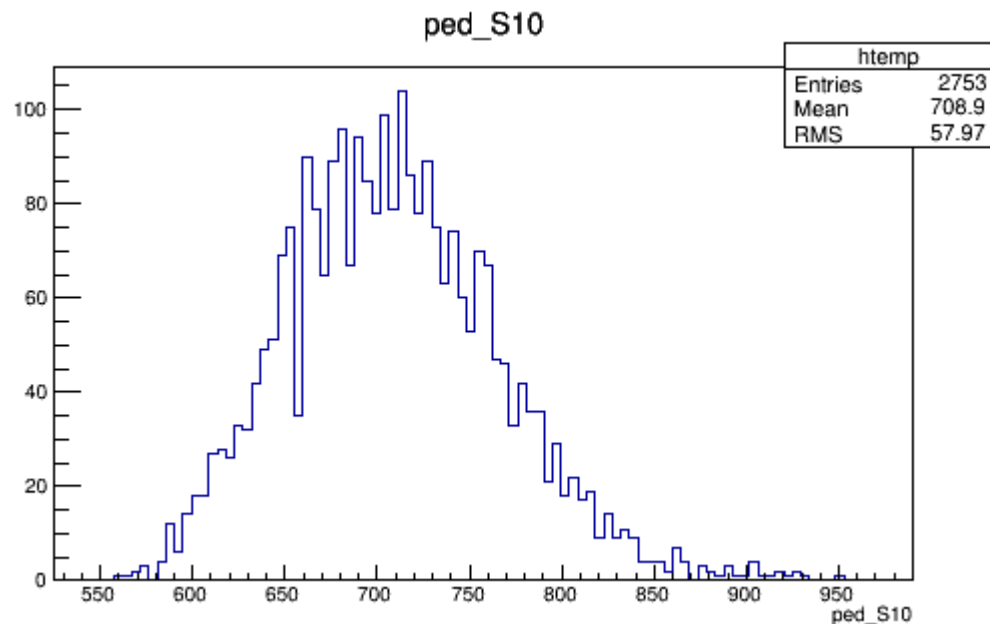


Points to some combination of voltage setting, connectors, or power distribution?

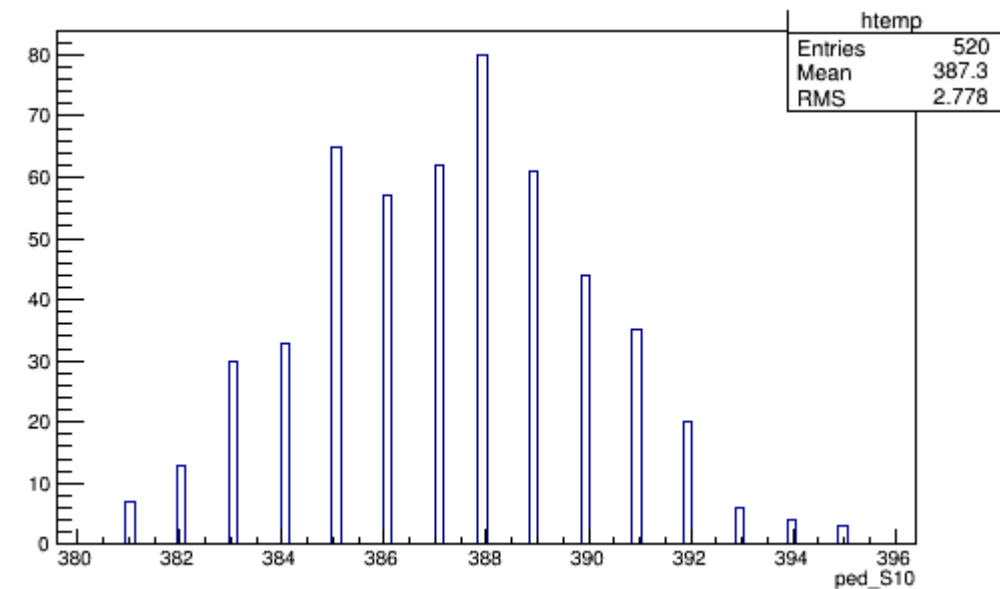
# Backup: Channel S10 in Cosmic Data

- Runs 121308 and 121309: significantly worse
- Only two 75.5 / 74.0 V setting runs. Is 75.5 the south or north side?
- Better when chillers added in run 121312

[Run 121308]



[Run 121312]



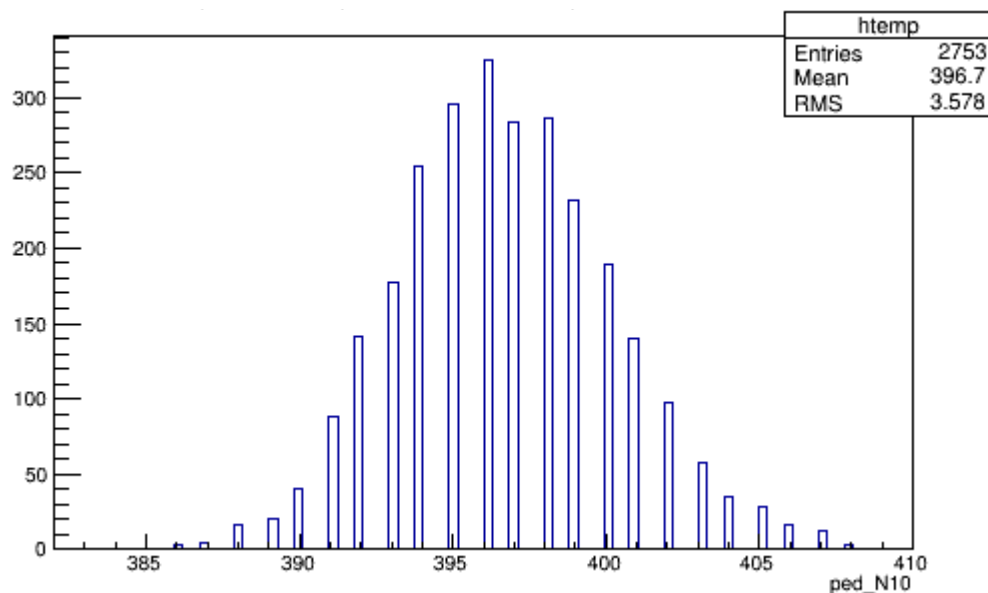




# Backup: Chillers in Channel N10 in Cosmic Data

- When chillers added, good channels saw pedestal go down too
- Here: channel N10

[Run 121308]



[Run 121312]

ped\_N10 {ped\_N10>380&&ped\_N10<450}

