



# The small system flow measurements from the STAR experiment

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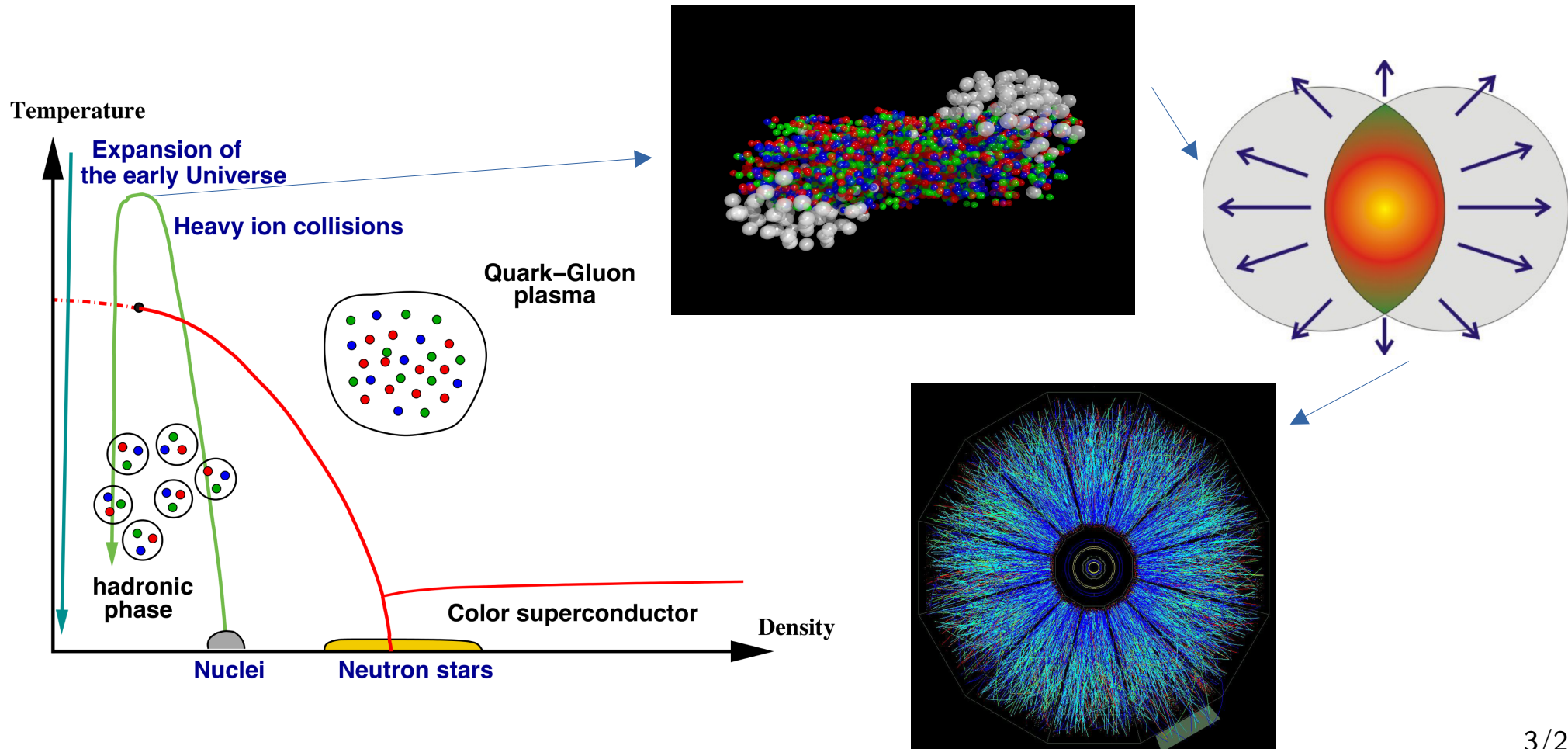
The State University of New York

# Outline

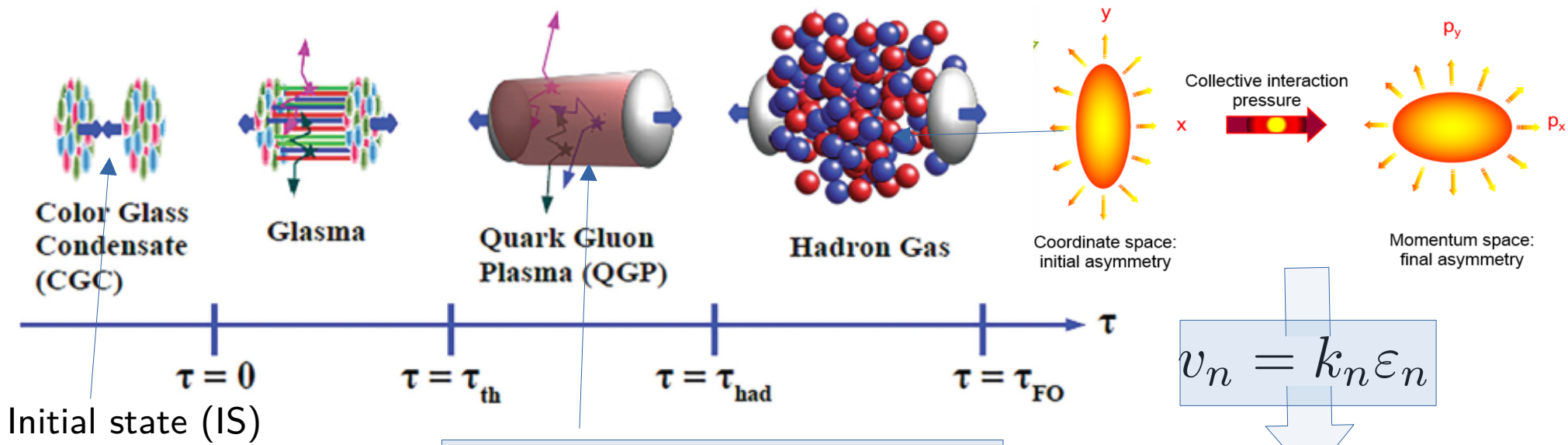
- Introduction
- Nucleon and subnucleon fluctuation
- Longitudinal decorrelation & non-flow
- Many-body correlation
  - O+O and d+Au data
- Collision geometry
- Summary and outlook

# Physics context

- Highly compressed, hot matter forms Quark-Gluon Plasma(QGP) with quark and gluon degrees of freedom.
- Heavy-ion collision  $\rightarrow$  study the properties of QGP.
  - Rapid expansion  $\rightarrow$  converts spacial non-uniformity in QGP to transverse momentum anisotropy of emitted particles.
  - For large-ion collision, QGP evolution is well described by the relativistic viscous hydrodynamic models as liquid with minimal viscosity.



# Flow in small system

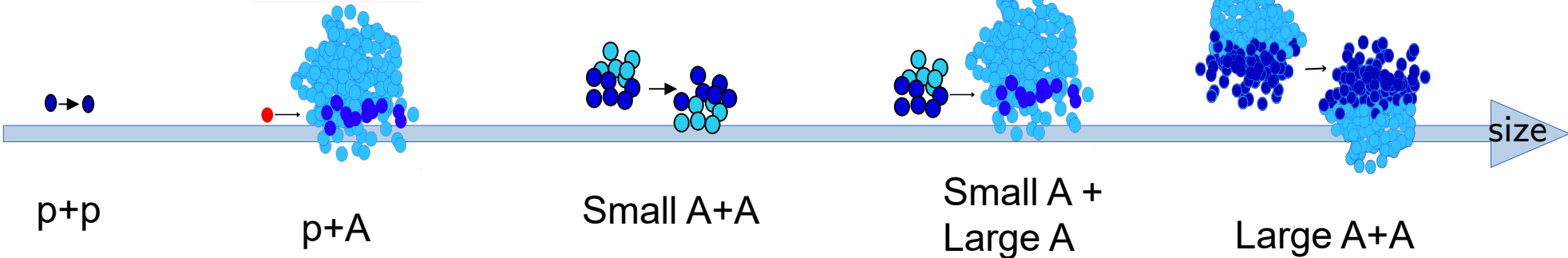


Final State (FS) collective response to geometry fluctuation

Particle distribution

- QGP evolution process. How much contributions from IS and FS?  $\rightarrow$  system size dependence?
- Collision frequency can depend on QGP size  $\rightarrow$  probe thermalization and limit of hydro

$$\frac{dN}{d\phi} \propto 1 + 2 \sum_{n=1}^{\infty} v_n(p_T) \cos(n(\phi - \Phi_n))$$



STAR scans whole range in size to study those effect!



## 2-particle correlation $v_n$

- Transverse hadron distribution in Fourier series:

$$\frac{dN}{d\phi} \propto 1 + 2 \sum_{n=1}^{\infty} v_n(p_T) \cos(n(\phi - \Phi_n))$$

- Associate and trigger particle pair distribution in azimuthal separation.

- Corrected for acceptance using mixed-event distribution.
- Pair distribution,

$$\frac{dN_{\text{pairs}}}{d\Delta\phi} \propto 1 + 2 \sum_{n=1}^{\infty} c_n(p_T^t, p_T^a) \cos(n\Delta\phi)$$

- $p_T$  dependence for trigger particles,

$$v_n(p_T^t) = \frac{c_n^{sub}(p_T^t, p_T^a)}{\sqrt{c_n^{sub}(p_T^a, p_T^a)}}$$

- Where non-flow is extracted using  $c_1$  method. Subtraction is robust with different methods..

$$c_n^{sub} = c_n - \frac{c_1}{c_1^{periph}} \times c_n^{periph}$$

PRC 58,1671

PRL. 130, 242301

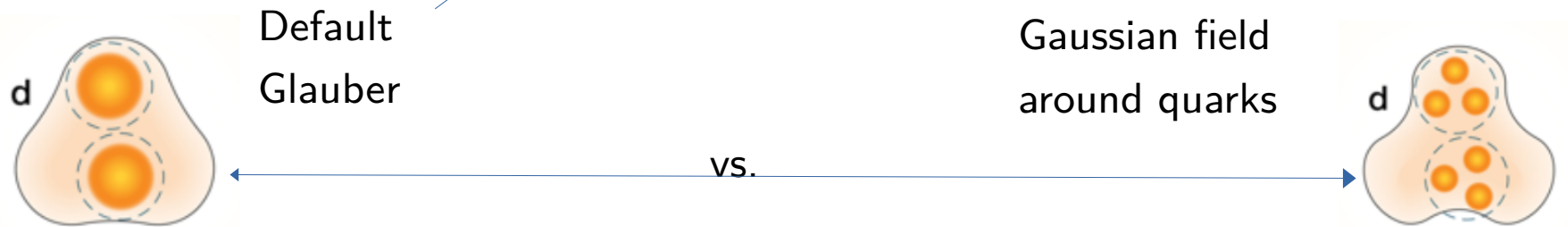
$v_n = k_n \varepsilon_n$  up to  $n = 3$ . Trace the eccentricities though particle correlation

# The role of (sub)nucleonic fluctuation on eccentricity

Geometry fluctuation response in FS depends on the degrees of freedom in the nucleus.

Model	a [33, 43] $\varepsilon_2^a(\varepsilon_3^a)$	c [31] $\varepsilon_2^c(\varepsilon_3^c)$
${}^3\text{He}+\text{Au}$	0.50(0.28)	0.53(0.38)
$d+\text{Au}$	0.54(0.18)	0.53(0.36)
$p+\text{Au}$	0.23(0.16)	0.41(0.34)

PRL 130, 242301



$$\begin{aligned} \varepsilon_2^{3\text{He}+\text{Au}} &\approx \varepsilon_2^{d+\text{Au}} > \varepsilon_2^{p+\text{Au}} \\ \varepsilon_3^{3\text{He}+\text{Au}} &> \varepsilon_3^{d+\text{Au}} \approx \varepsilon_3^{p+\text{Au}} \end{aligned}$$

Where RMS  
 $\varepsilon_n = \sqrt{\langle \varepsilon_n^2 \rangle}$

$$\begin{aligned} \varepsilon_2^{3\text{He}+\text{Au}} &\approx \varepsilon_2^{d+\text{Au}} > \varepsilon_2^{p+\text{Au}} \\ \varepsilon_3^{3\text{He}+\text{Au}} &\approx \varepsilon_3^{d+\text{Au}} \approx \varepsilon_3^{p+\text{Au}} \end{aligned}$$

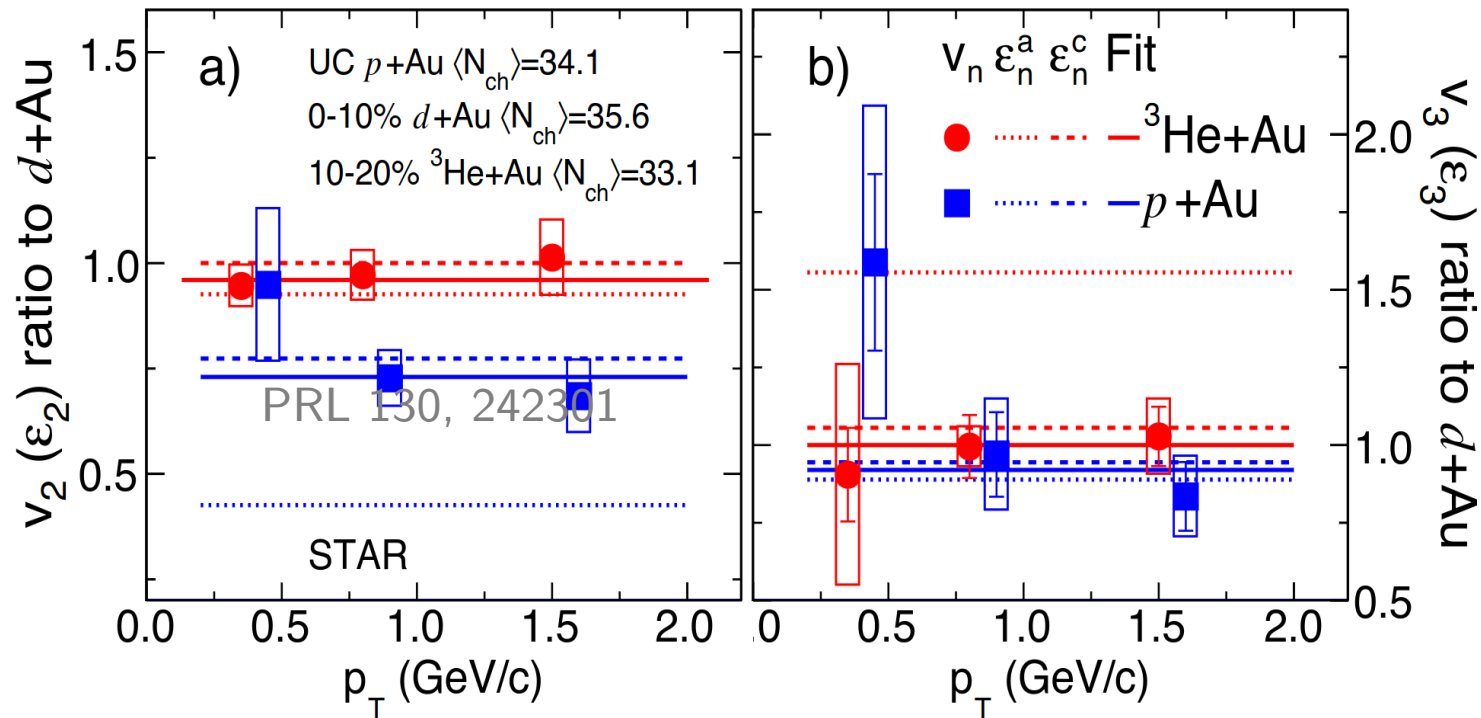
${}^3\text{He}+\text{Au}$  to have largest  $\varepsilon_3$  due to nucleon geometry.

Similar triangular flow with sub-nucleon fluctuation.

Nucleon vs, quark picture different hierachy → Check by ratio between systems

# The role of (sub)nucleonic fluctuation on eccentricity

Previous STAR result



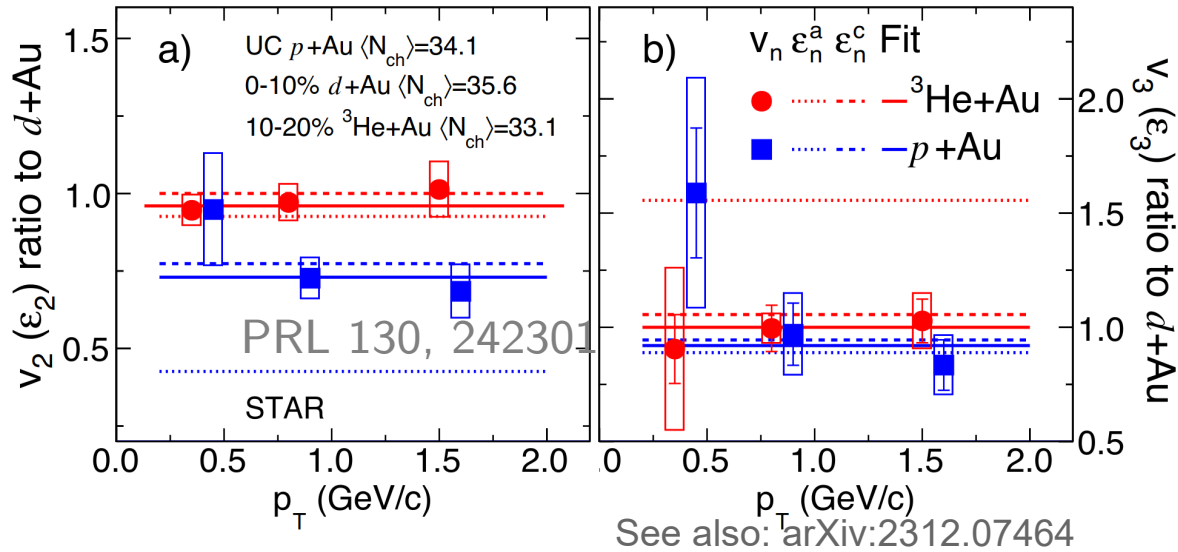
$v_n \rightarrow$  STAR data

Model c  $\rightarrow$  Gaussian field around quarks

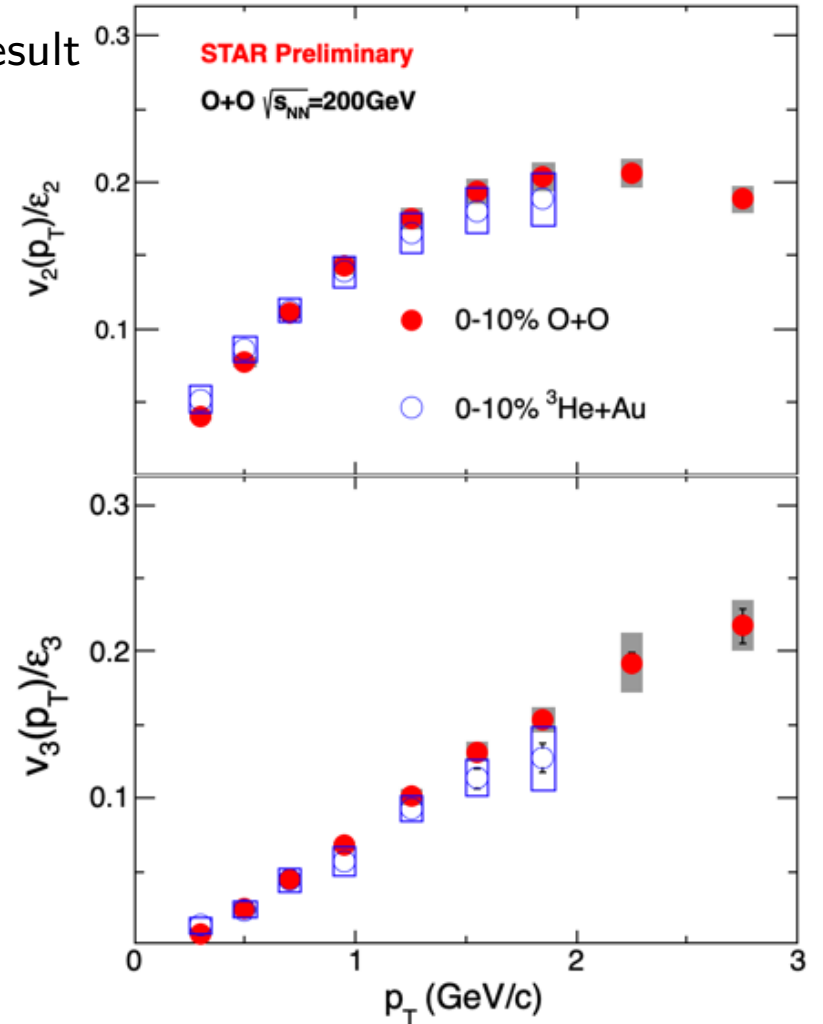
Compare  $v_n$  ratio in data and  $\epsilon_n$  ratio in model

# The role of (sub)nucleonic fluctuation on eccentricity

Previous STAR result



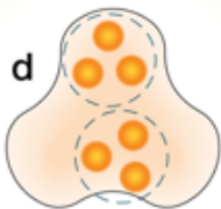
O+O result



Model c  $\rightarrow$  Gaussian field around quarks

Compare  $v_n$  ratio in data and  $\epsilon_n$  ratio in model

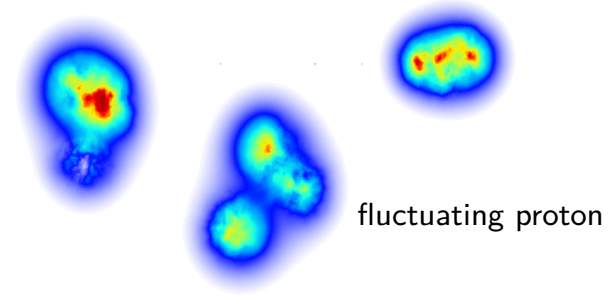
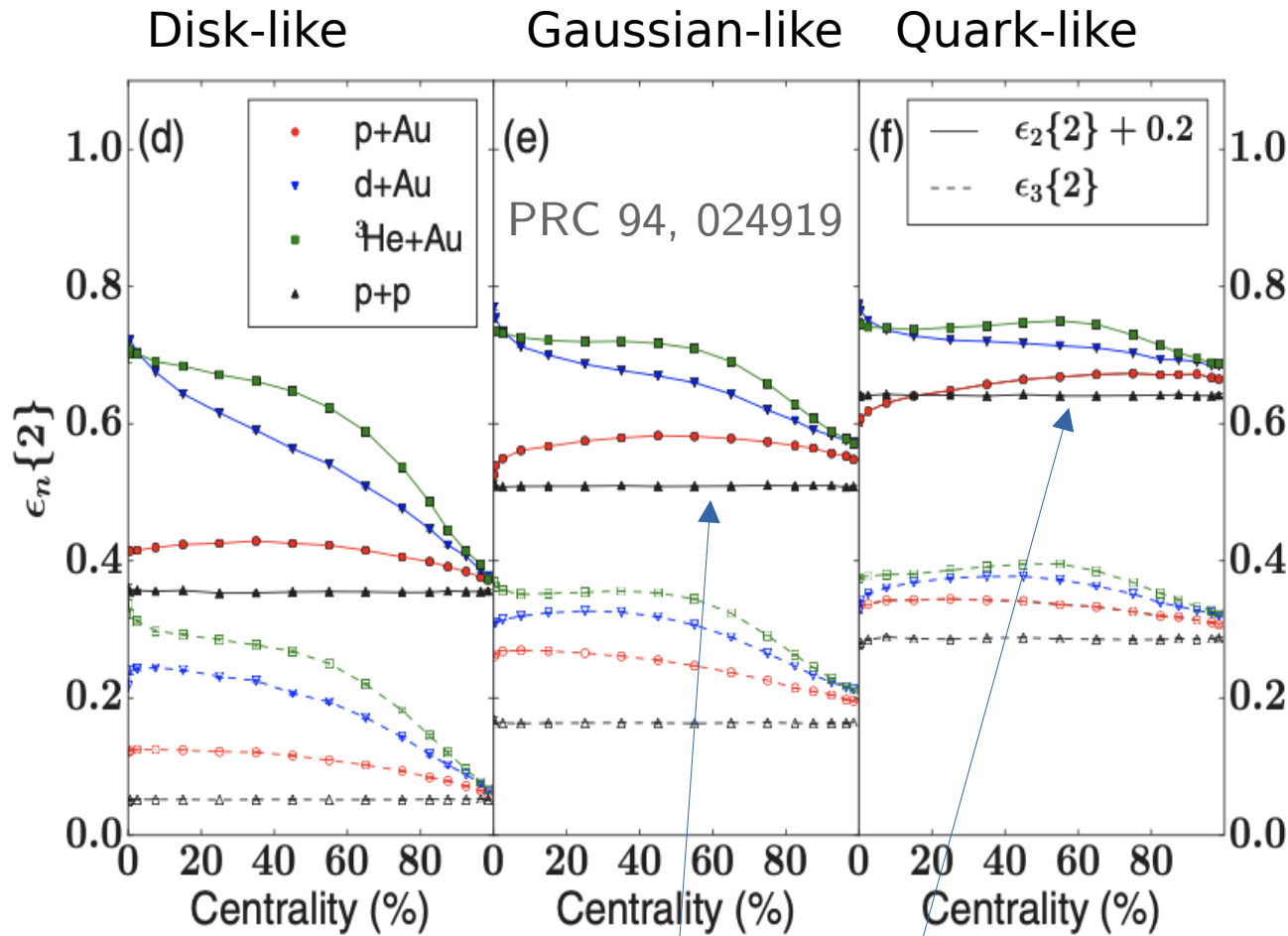
- $\epsilon_n$  with quark Glauber
- Same  $v_n/\epsilon_n$  scaling to  ${}^3He+Au$ .



No evidence of larger triangular flow in  $He+Au$  than  $p+Au$ ,  $d+Au$   
 Scaling in O+O agrees with quark Glauber  
 Data prefer model with subnucleon fluctuation.

# The role of (sub)nucleonic fluctuation on eccentricity

- p+p 200 GeV data taken in 2024.

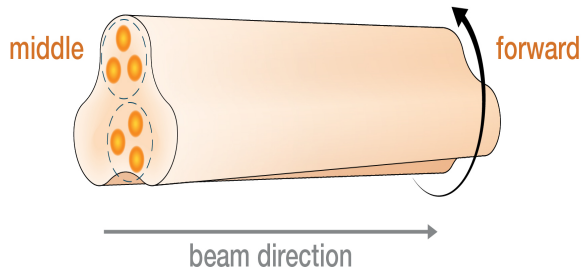


Large  $\epsilon_2$  difference, easier to measure compared to  $\epsilon_3$

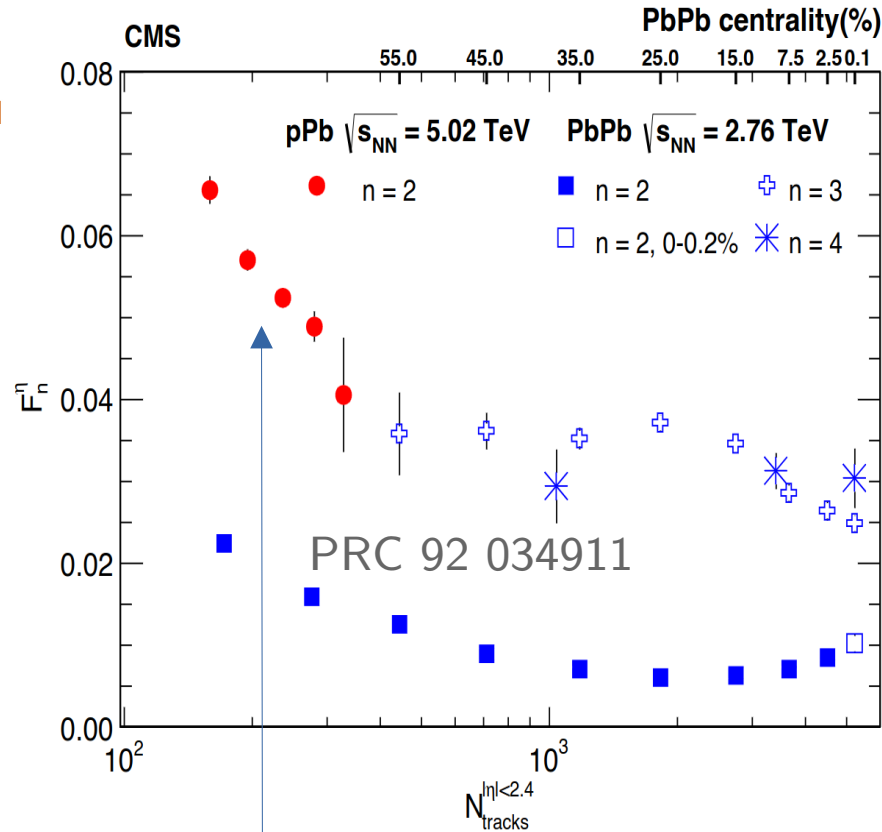
p+p can better distinguish (sub)nucleonic fluctuation

# Longitudinal decorrelation & nonflow

- Eccentricity fluctuation in  $\eta$  + particle pair from different  $\eta$  = decorrelation in  $v_n$

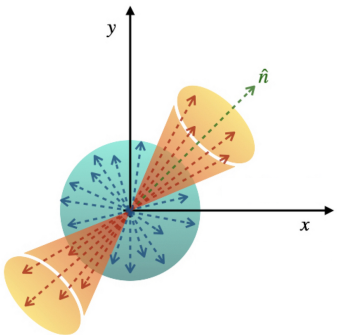


Unequal eccentricity vectors due to fluctuation



Rapidity factorization breaking is more evident for smaller system.

$$\cos \left\{ n \left[ \Psi_n (\eta^a) - \Psi_n (\eta^b) \right] \right\} = e^{-F_n^\eta |\eta^a - \eta^b|}$$

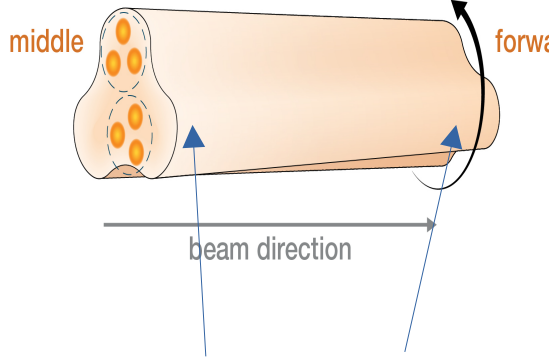


Measure 2-particle correlation with large  $\eta$  gap  
 $\rightarrow$  Suppress non-flow correlation such as particle decay.

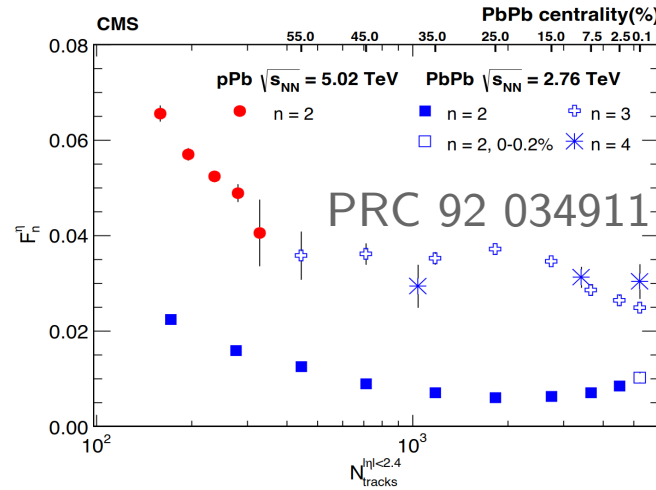


# Longitudinal decorrelation & nonflow

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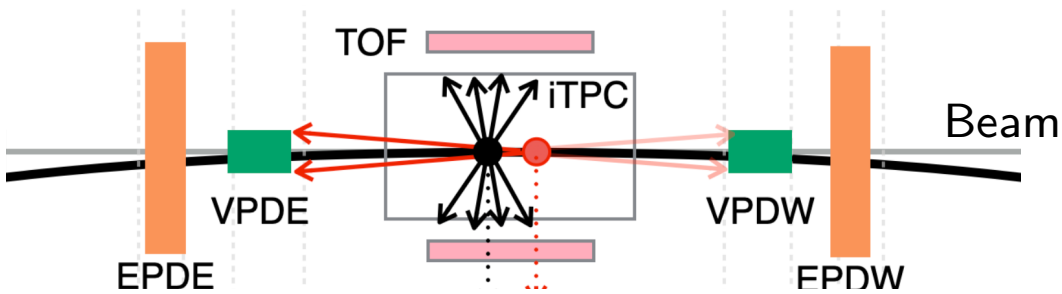


Unequal eccentricity vectors due to fluctuation

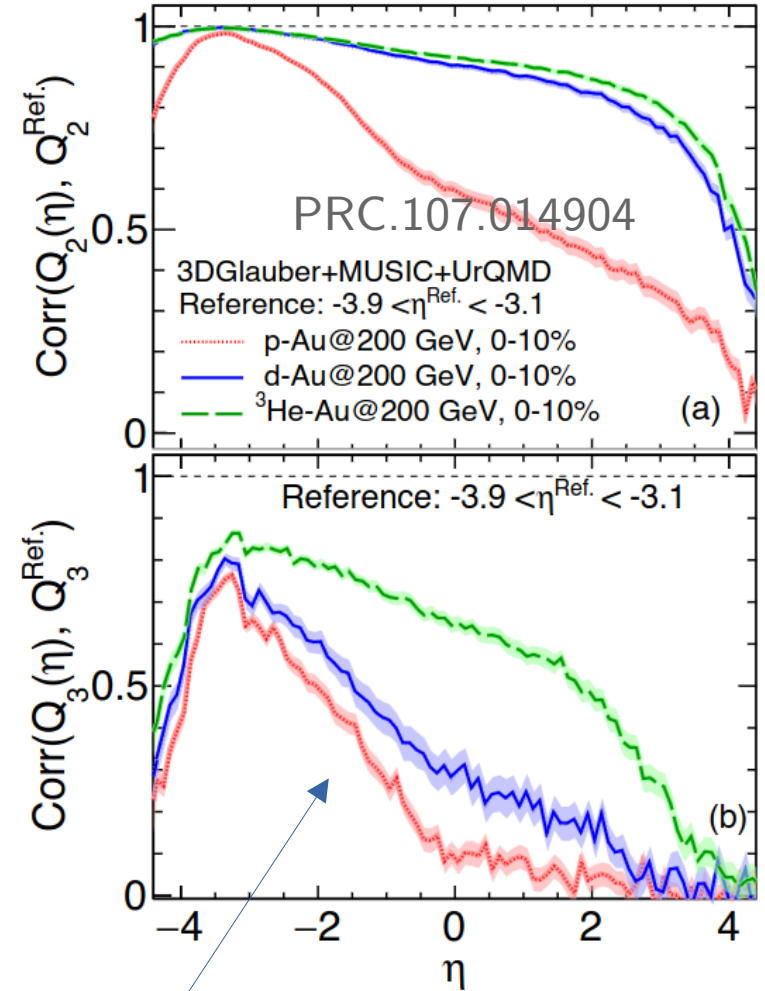


Rapidity factorization breaking is more evident for smaller system.

Wider  $\eta$  coverage helps suppressing nonflow



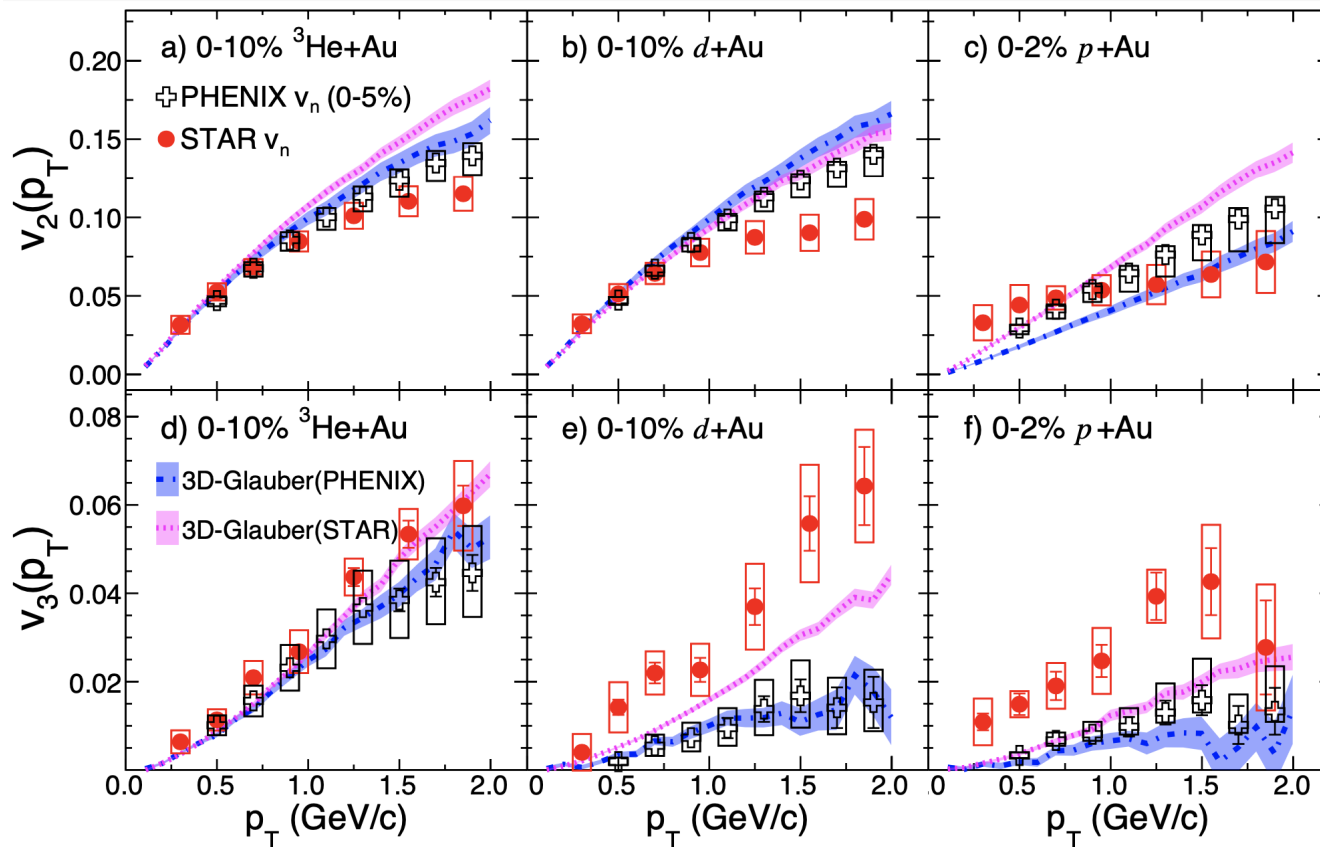
Upgraded STAR detector with wider  $\eta$  coverage is ideal for quantifying the decorrelation.



3+1D model shows non-boost-invariance

# Longitudinal decorrelation & nonflow

Previous publications



STAR: PRL 130, 242301  
 arXiv:2312.07464  
 PHENIX: Nature Phys. 15, 214  
 3D-Glauber: PRC 107, 014904

Large difference in  $v_3$  for STAR and PHENIX in data and models.

STAR use mid-rapidity:  $|\eta| < 0.9$

PHENIX use mid-forward rapidity:  $|\eta| < 0.35$  to  $-3.0(3.9) < \eta < -1(-3.1)$

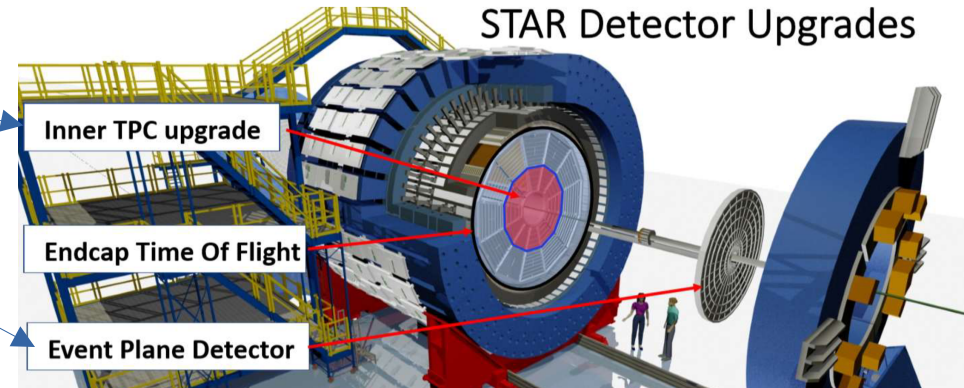
→ Due to decorrelation?

# Flow in O+O

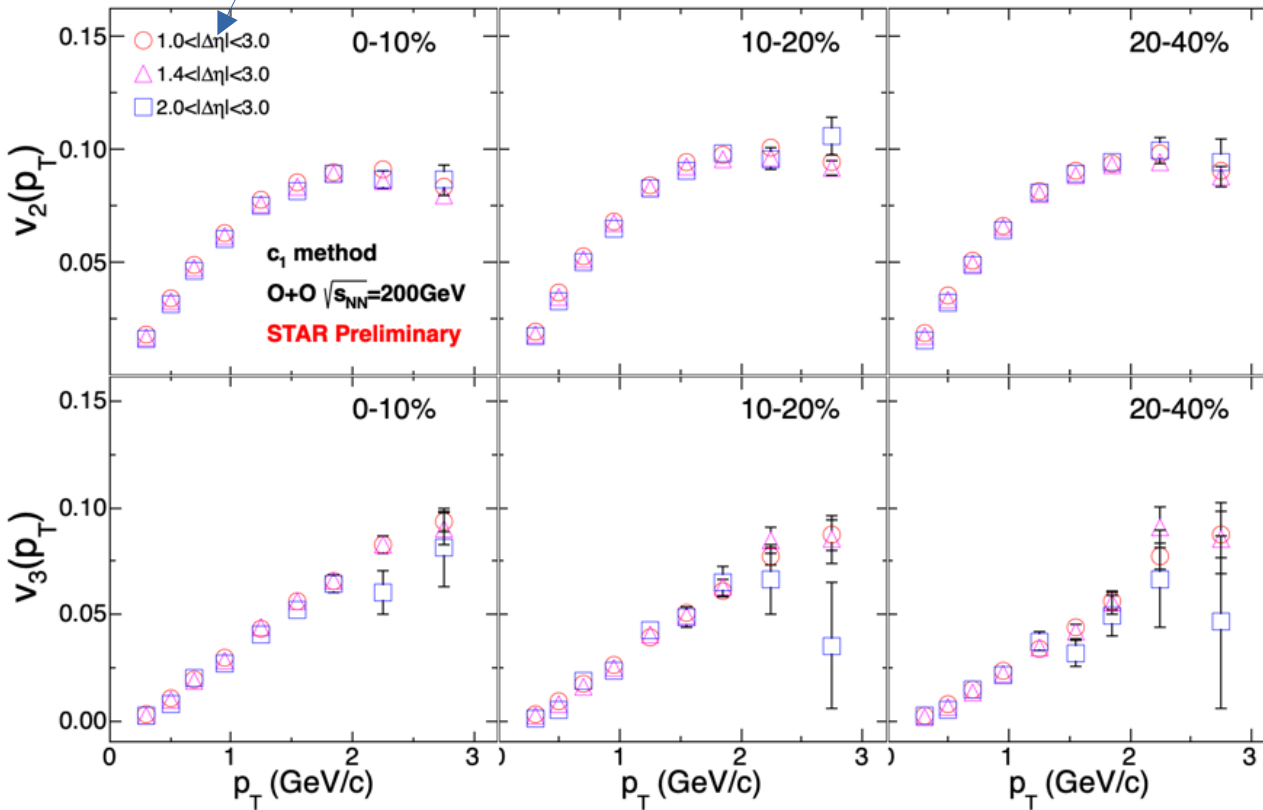
- New iTPC  $-1.5 < \eta < 1.5$
- New EPD  $2.1 < |\eta| < 5.1$

Further separation in  $\eta \rightarrow$

- More decorrelation?
- Suppress short range non-flow?



Change  $\eta$  gap for two-particle correlation in TPC



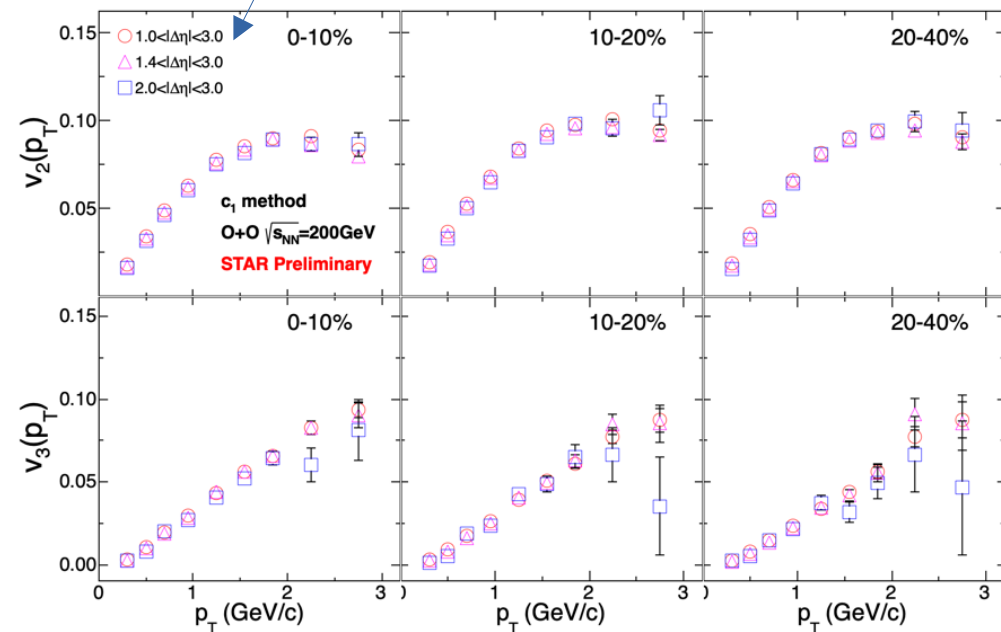
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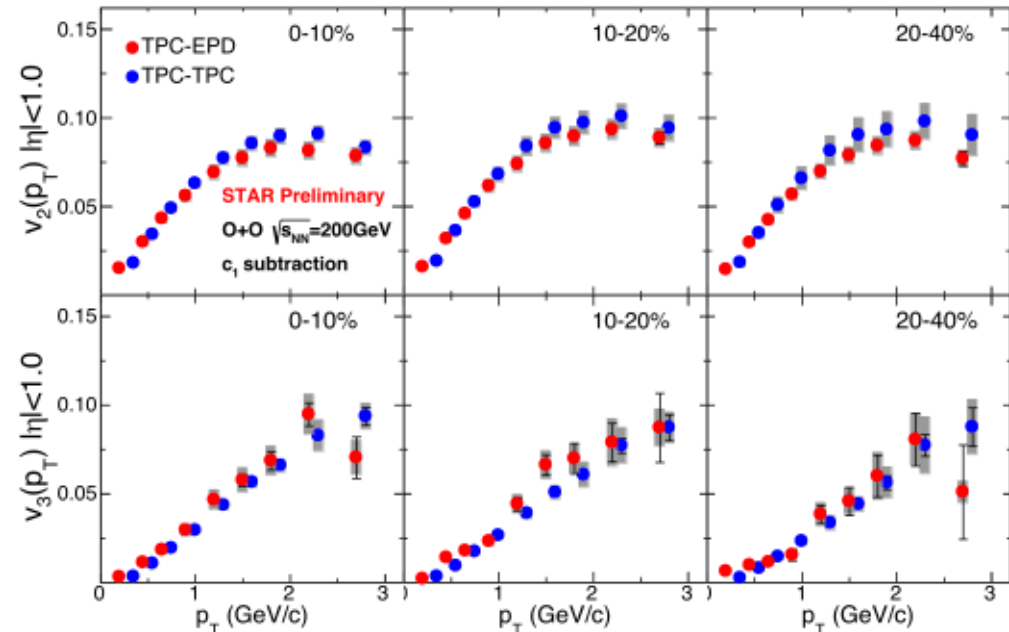
Change  $\eta$  gap for two-particle correlation in TPC



TPC-EPD two particle correlation

- TPCa:  $|\eta| < 1.0$
- TPCb:  $0.5 < \eta < 1.5$
- TPCc:  $-1.5 < \eta < -0.5$
- EPD:  $2.1 < |\eta| < 5.3$

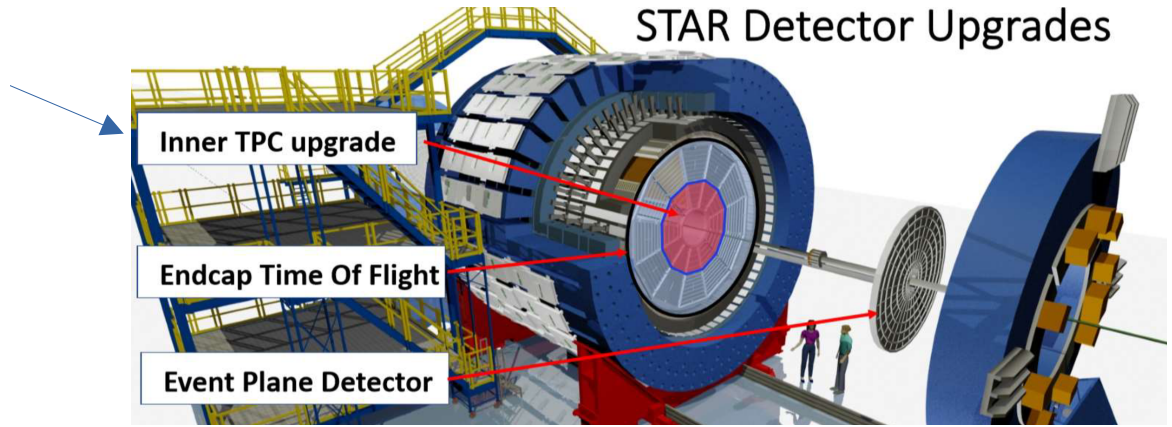
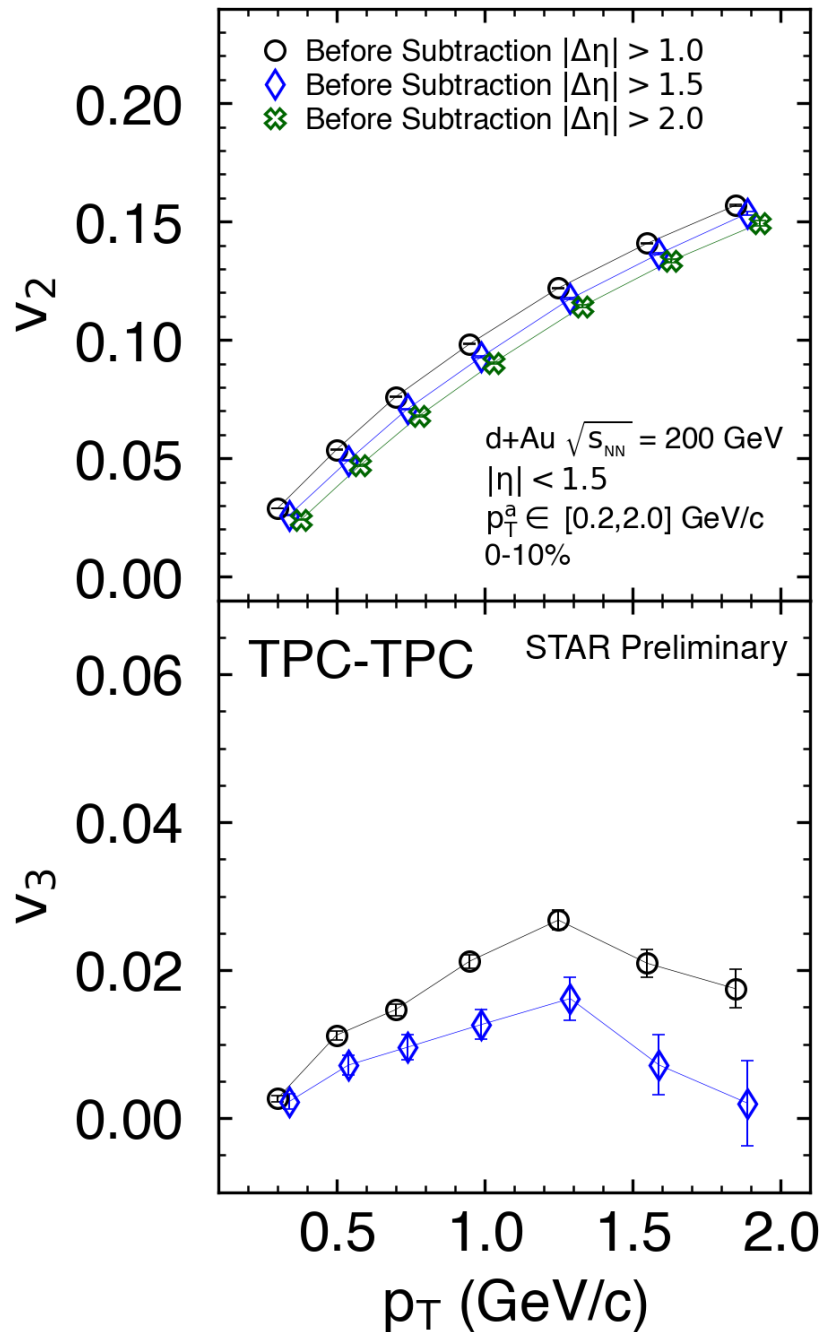
$$v_n(\text{TPCa}, p_T) = \frac{Q_n^{\text{TPCa}} Q_n^{*\text{EPD}}}{\sqrt{\frac{(Q_n^{\text{TPCb}} Q_n^{*\text{EPD}}) \times (Q_n^{\text{TPCc}} Q_n^{*\text{EPD}})}{Q_n^{\text{TPCb}} Q_n^{*\text{TPCc}}}}$$



Do not observe significant change in flow magnitudes due to different  $\eta$  selections.

# New: Flow in d+Au from Run 21

- Use new iTPC  $-1.5 < \eta < 1.5$

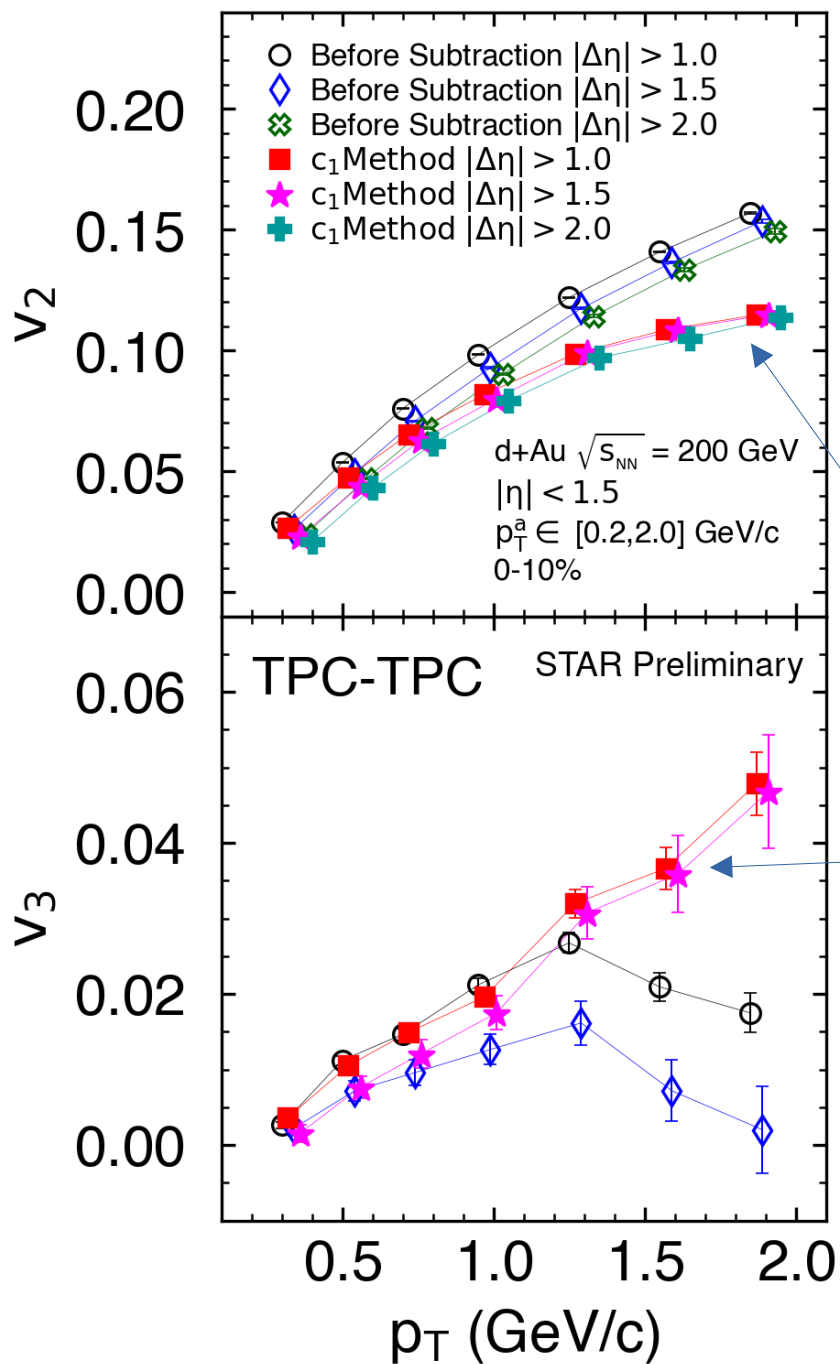
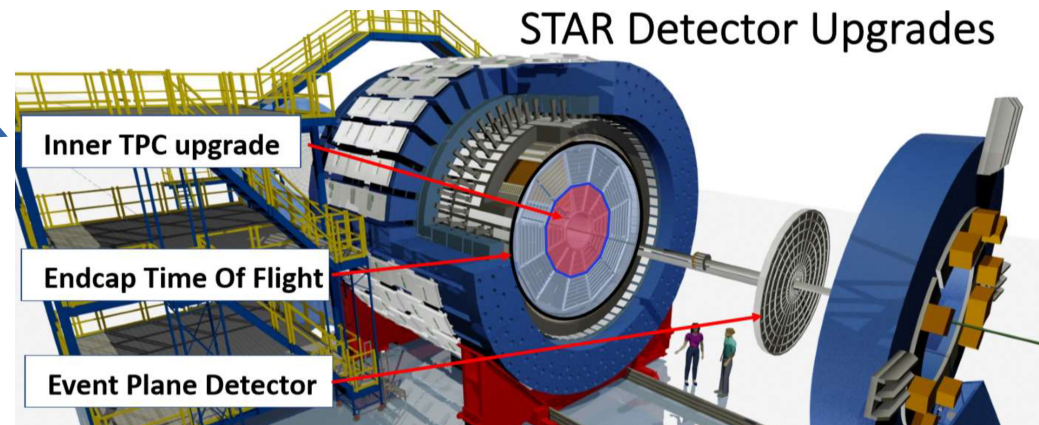


Raw flow magnitude affected by changing  $\eta$  acceptance.

# New: Flow in d+Au from Run 21

- Use new iTPC  $-1.5 < \eta < 1.5$

STAR Detector Upgrades

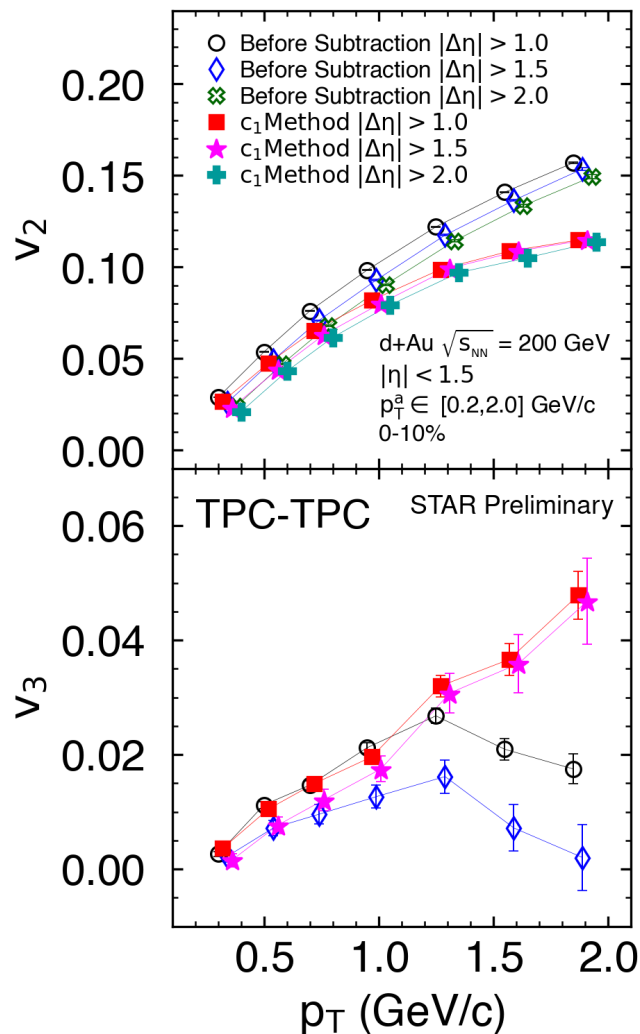


Consistent magnitudes after non-flow subtraction between difference acceptances.



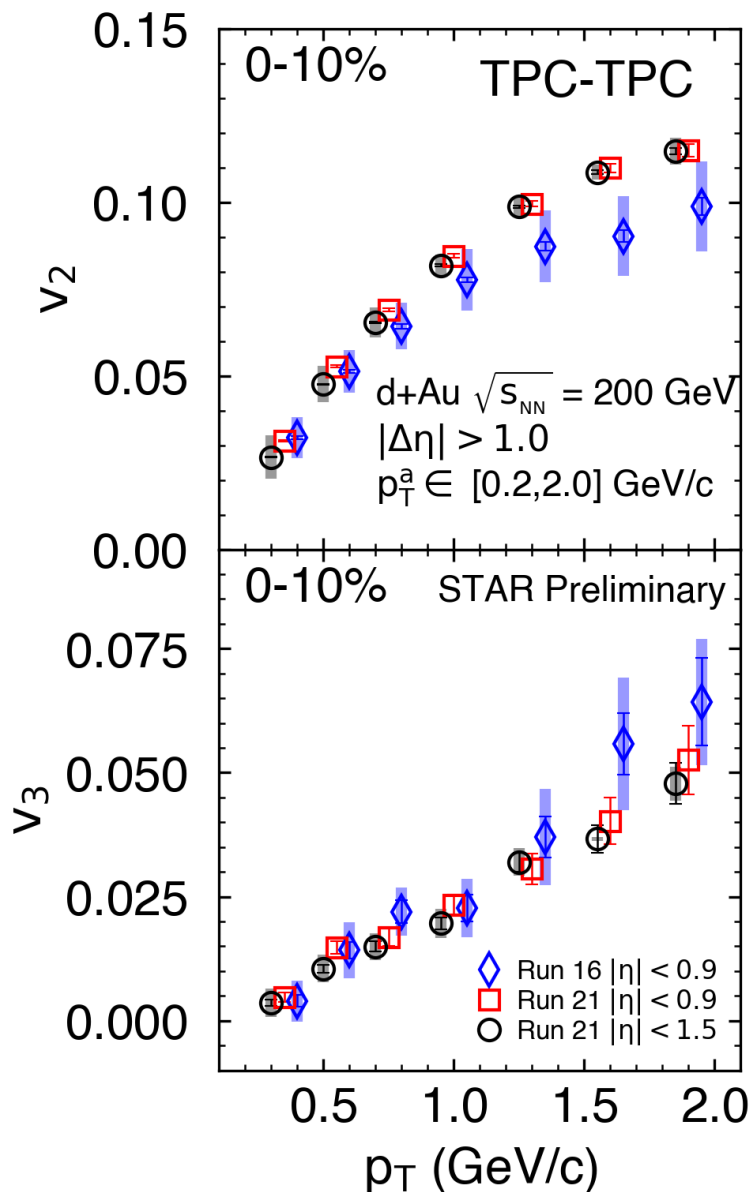
# New: Flow in d+Au from Run 21

- Use new iTPC  $-1.5 < \eta < 1.5$



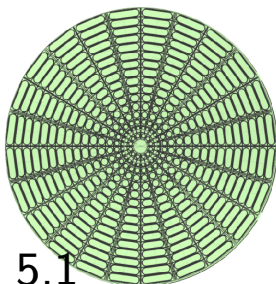
Compare to previous run:

- with same  $\eta$  acceptance
- or
- with larger  $\eta$  acceptance



New EPD

$2.1 < |\eta| < 5.1$

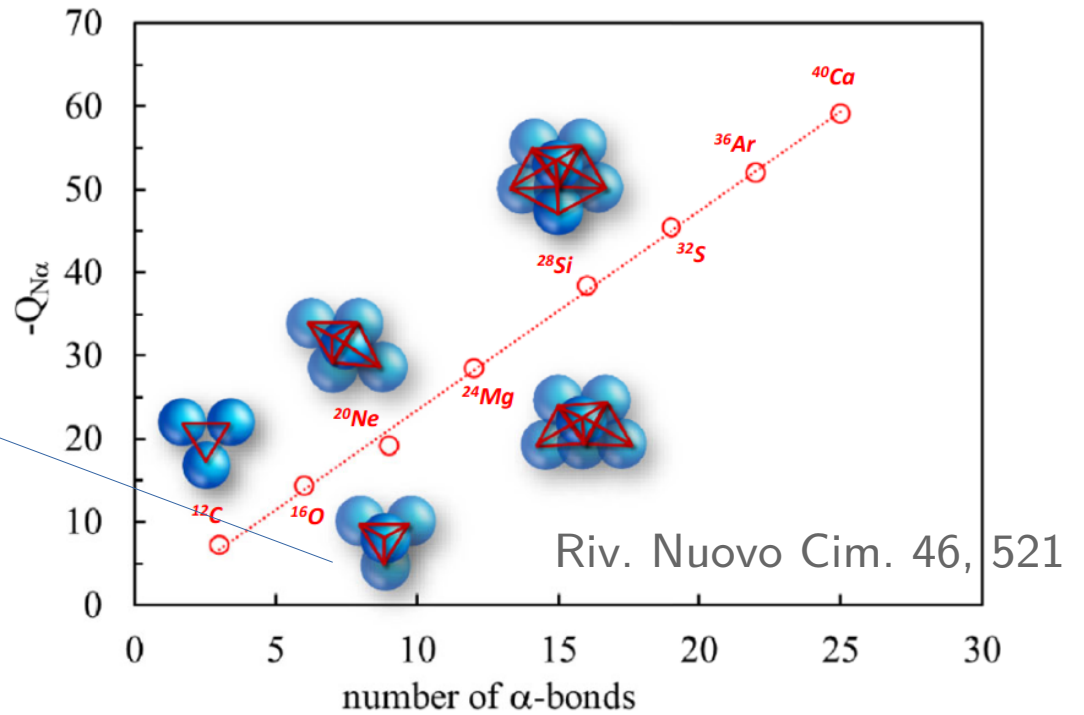
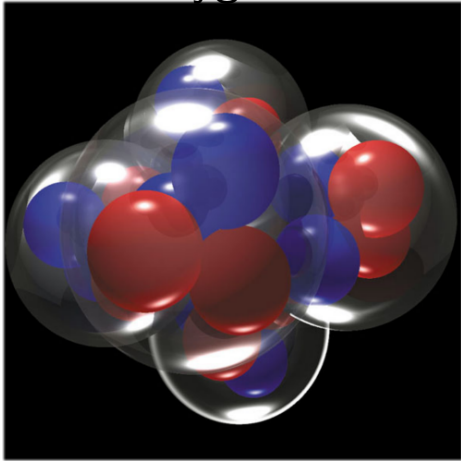


STAR results are robust with  $\eta$  gap change.  
 Measure wider  $\eta$  range with EPD. Work in progress.

# Many-body correlation

- Inter-nucleon multi-body correlation and clustering in nucleus

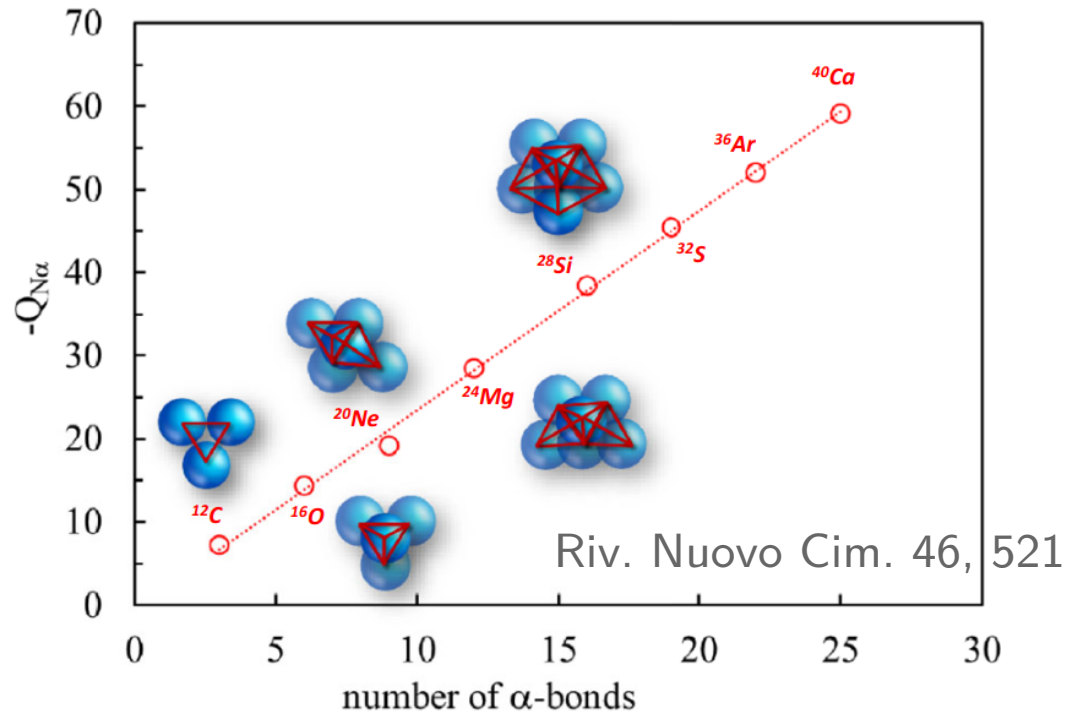
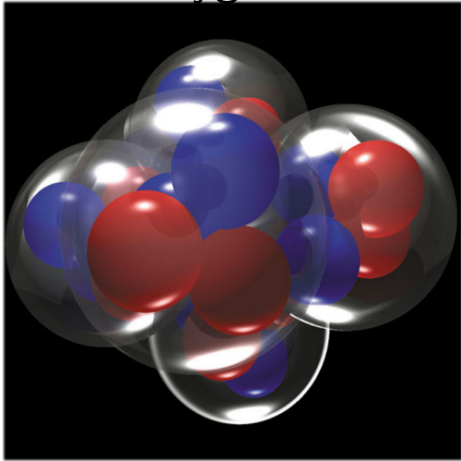
<sup>16</sup>Oxygen



# Many-body correlation

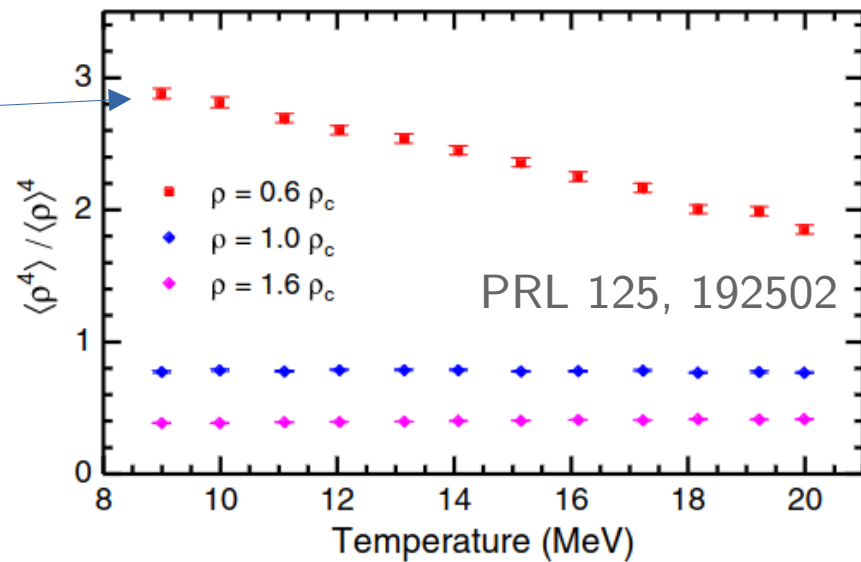
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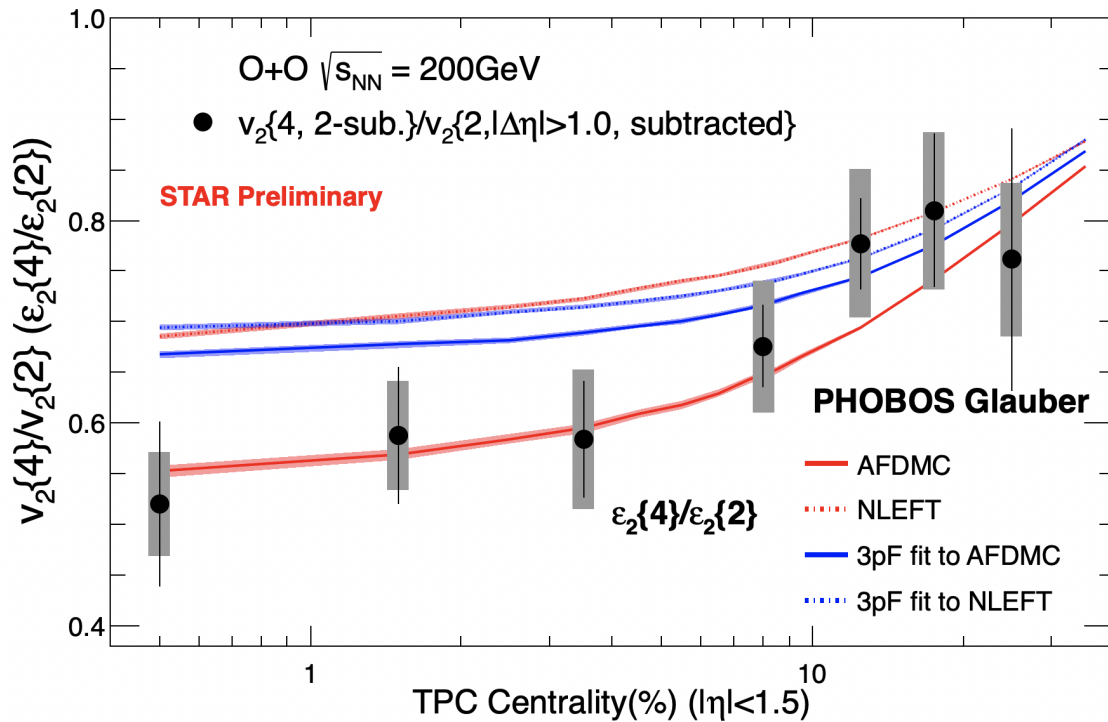
Ab initio thermodynamics calculation shows four-body clustering.

Also see O+O AMPT:  
arXiv:2404.08385

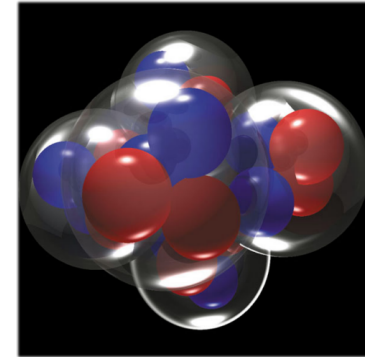


Can we see effect using multi-particle correlation observables?

# Many-particle correlation in O+O



<sup>16</sup>Oxygen



- Direct model input **with** many-body correlation
- 3-parameter-Fermi **without** many-body correlation

Ratio  $\epsilon_2\{4\}/\epsilon_2\{2\}$  decrease, amount dependent on model.

→ many body correlation enhances the flow fluctuation?

→ change in average geometry?

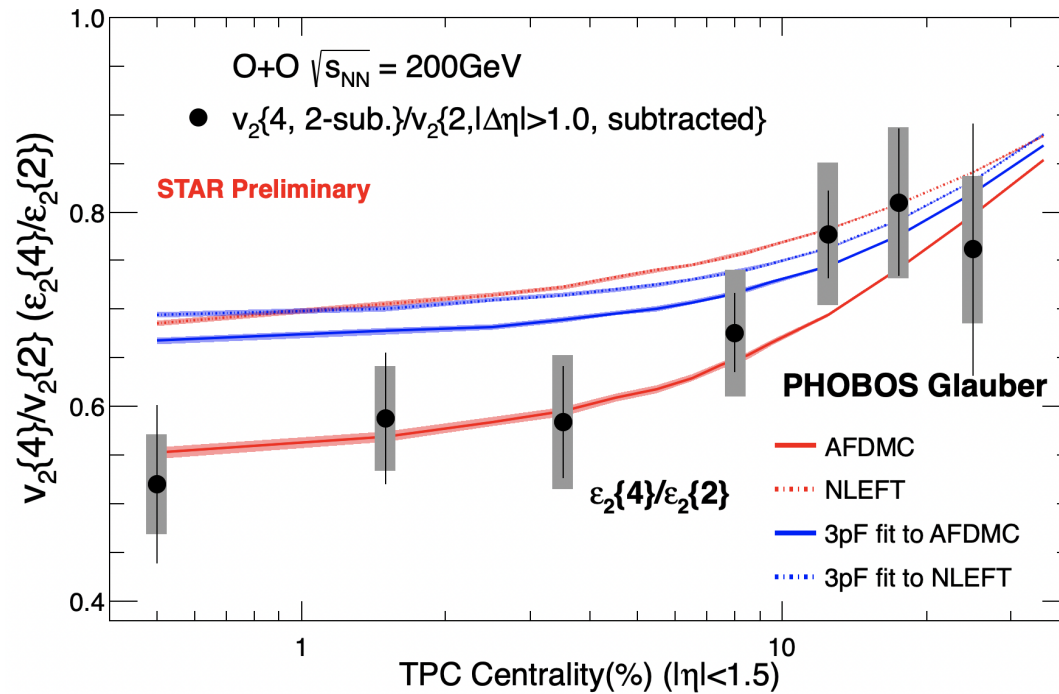
$$\sigma^2 = \langle \epsilon_n^2 \rangle - \langle \epsilon_n \rangle^2$$

$$(\epsilon_n\{2\})^2 = \langle \epsilon_n^2 \rangle = \langle \epsilon_n \rangle^2 + \sigma^2$$

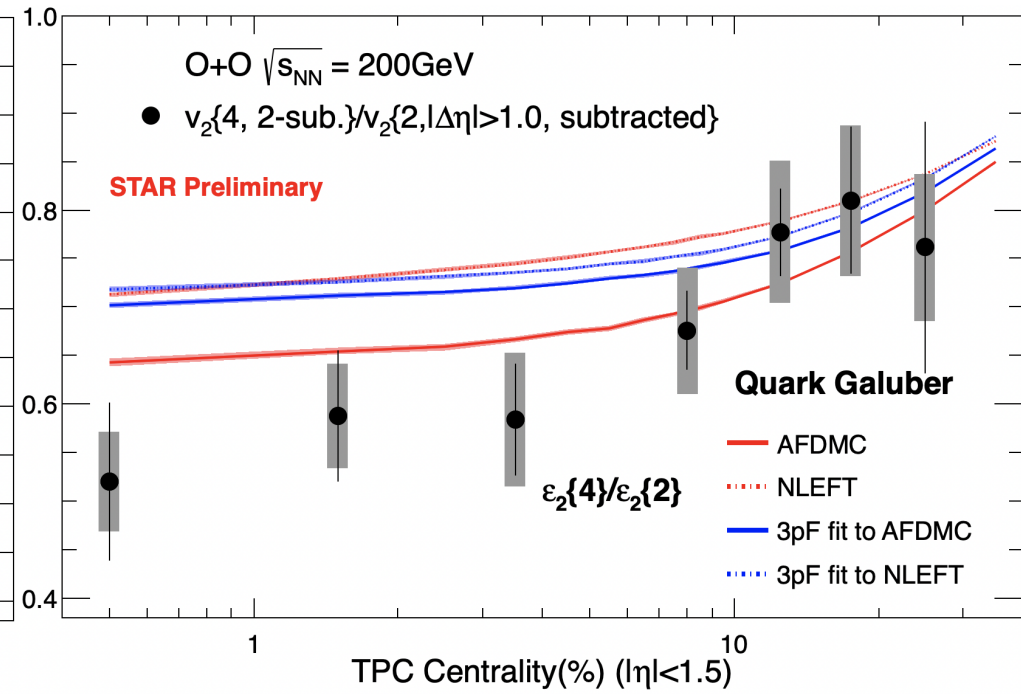
$$(\epsilon_n\{4\})^2 = \sqrt{2 \langle \epsilon_n^2 \rangle - \langle \epsilon_n^4 \rangle} \approx \langle \epsilon_n \rangle^2 - \sigma^2$$

# Many-particle correlation in O+O

Default Glauber



Gaussian field around quarks

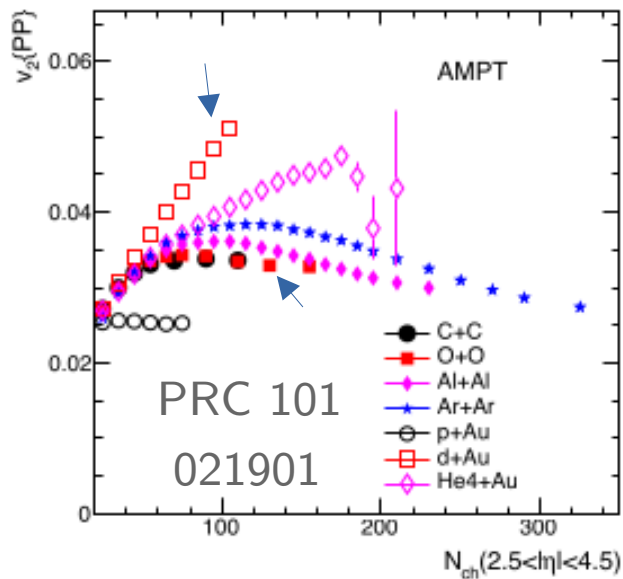
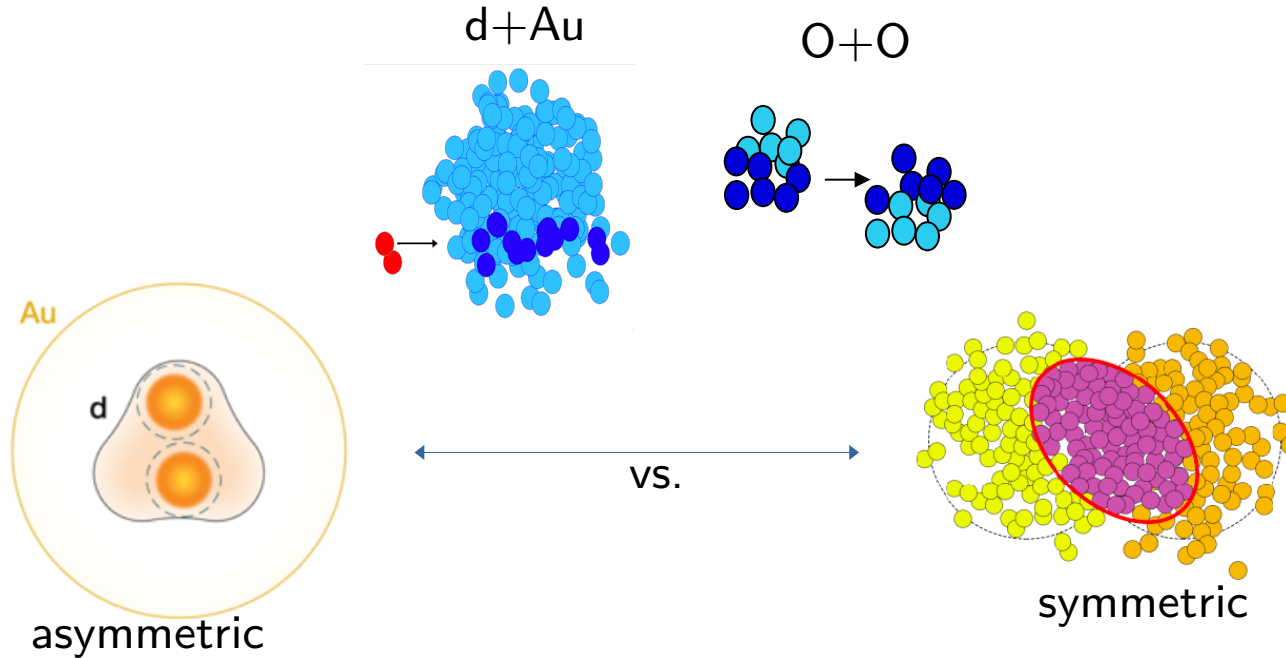


Sub-nucleon fluctuation reduce the effect of many-body correlation.

Multi-particle correlation in O+O can distinguish if many body correlation is present. Model dependent results.

# Collision geometry

Similar size, drastically different collision geometry



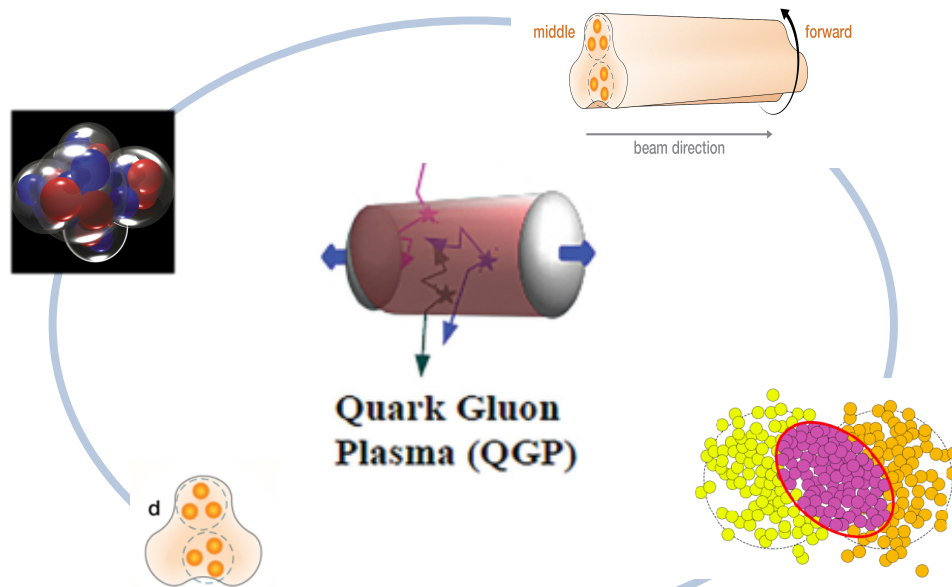
Different collision geometry captured by  $N_{ch}$  dependence  $v_2$

Compare d+Au and O+O flow to explore geometry. Work in progress



# Summary and outlook

- New data (d+Au, O+O, p+p) taken with upgraded STAR detector provide wealth of opportunity to study different effects in the evolution of QGP
  - $^3\text{He}+\text{Au}$  is similar to O+O and other systems, indicating no nucleon scale geometry transmutation.
  - Robust flow measurements with variations in  $\eta$  gap and acceptance indicate consistent non-flow estimates and no evident decorrelation.
  - $v_2\{4\}/v_2\{2\}$  could be used to probe multi-particle correlation in  $^{16}\text{oxygen}$ .



- Future flow measurements in d+Au and p+p to better distinguish (sub)nucleon geometry fluctuation response.
- Look for decorrelation in d+Au with wider  $\eta$  range using EPD.
- Explore unique collision geometries by comparing d+Au and O+O.

**WORK IN PROGRESS**

**Thank you for your attention**