



# **Vorticity measurements**

# from the STAR experiment at RHIC

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#### 2024 RHIC/AGS Annual Users' Meeting

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2024 RHIC/AGS ANNUAL USERS' MEETING

**A New Era of Discovery** Guided by the New Long Range Plan for Nuclear Science

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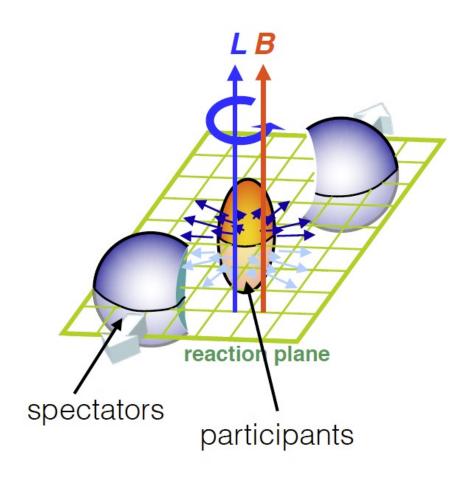


## **D** Brief introduction on vorticity and polarizaiton

- Analysis process
- □ Recent STAR experiment results
  - Hyperon global polarization
  - Hyperon polarization along beam direction
- **D** Summary and outlook

## Orbital angular momentum/magnetic field in heavy ion collisions





#### Orbital angular momentum

Z.-T. Liang and X.-N. Wang, PRL 94, 102301 (2005)

$$L \sim \frac{Ab\sqrt{s}}{2} \sim 10^6 \hbar$$

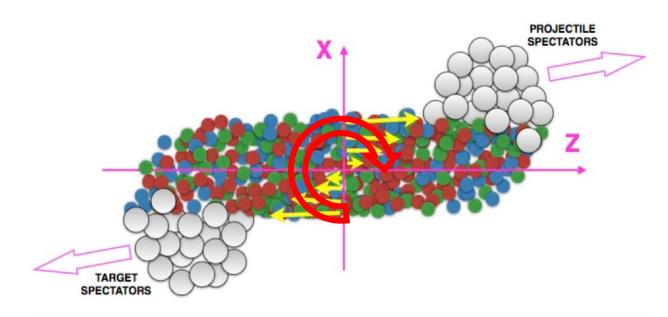
Expected to have strong magnetic field

$$eB \sim \gamma \alpha_{EM} \frac{Z}{b^2} \sim 10^{18} G \sim 10^{14} T$$

(RHIC Au+Au200 GeV, b=10 fm)

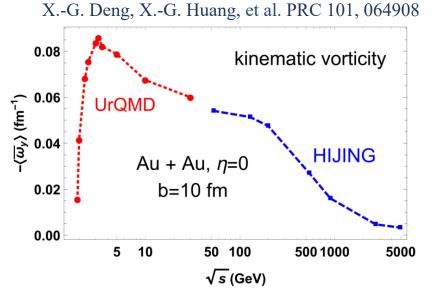
#### Vorticity in heavy ion collisions



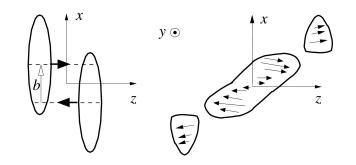


Orbital angular momentum

Local fluid vorticity 
$$\omega = \frac{1}{2} \nabla \times v$$
  
The most vortical fluid ~  $10^{20} - 10^{21} s^{-1}$   
(Au+Au@RHIC at **b**=10 fm)

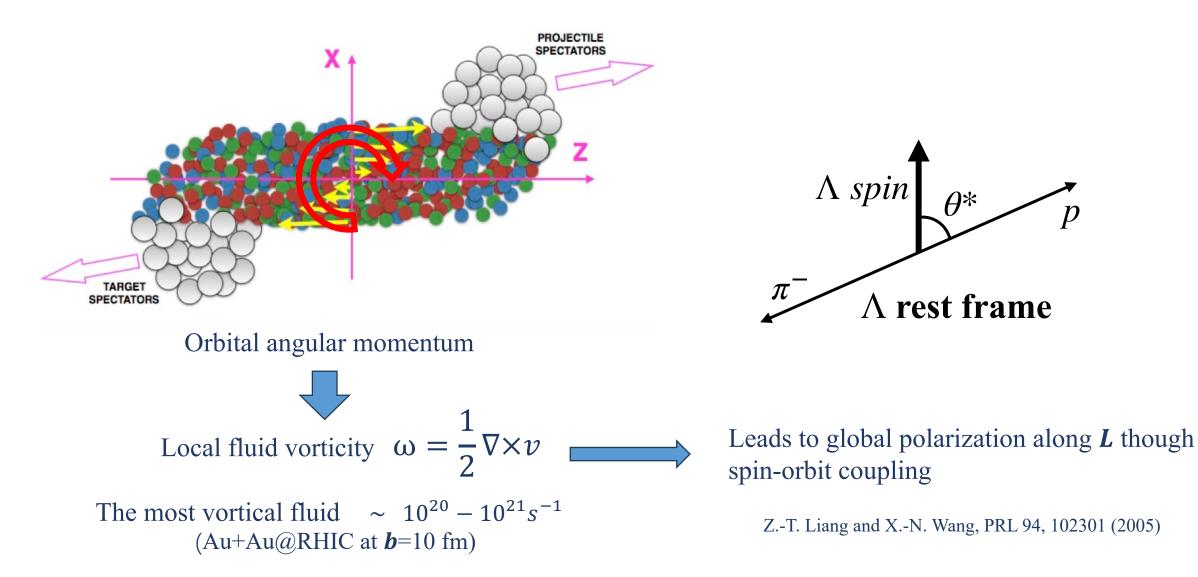


Energy dependence of initial vorticity



#### Vorticity and polarization





 $\theta^*$ 

 $\Lambda$  rest frame

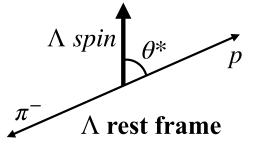
## Global polarization measurement

STAR

"Self-analyzing", parity-violating weak decay channel of hyperons
 Daughter baryon is preferentially emitted in the direction of the hyperon spin

 $\frac{\mathrm{d}N}{\mathrm{d}\Omega^*} = \frac{1}{4\pi} \left(1 + \alpha_H P_H \cos\theta^*\right)$ 

 $\alpha_H$ : hyperon decay parameter  $P_H$ : hyperon polarization  $\theta^*$ : polarization angle



 $\Lambda \rightarrow p + \pi^{-}$ (BR:63.9%,c $\tau$ ~7.9cm)

Hyperon	Deacy mode	$\alpha_H$	Spin
$\Lambda(uds)$	$\Lambda \to p + \pi^-$	0.732	1/2
$\Xi^{-}(dss)$	$\Xi^-  ightarrow \Lambda + \pi^-$	-0.401	1/2
$\Omega^{-}(sss)$	$\Omega^- \to \Lambda + K^-$	0.0157	3/2
			PDG20



□ "Self-analyzing", parity-violating weak decay channel of hyperons

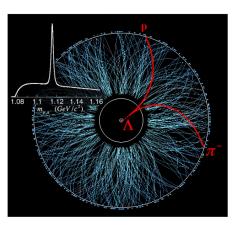
Daughter baryon is preferentially emitted in the direction of the hyperon spin

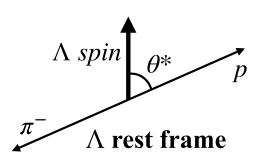
Measured via the distribution of the azimuthal angle of the hyperon decay baryon (in the hyperon rest frame) with respect to

the reaction plane.

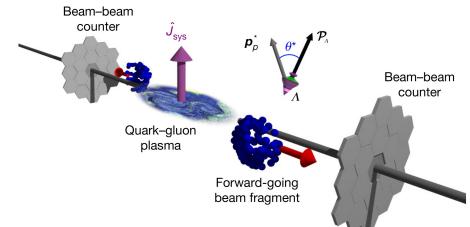
$$P_{\Lambda} = \frac{8}{\pi \alpha_{\Lambda}} \frac{1}{A_0} \frac{\left\langle \sin(\Psi_1 - \phi_p^*) \right\rangle}{Res(\Psi_1)}$$

 $\alpha_{\Lambda} = -\alpha_{\overline{\Lambda}} = 0.732 \pm 0.014$   $A_0$ : Acceptance correction factor  $\Psi_1$ : First-order event plane angle  $Res(\Psi_1)$ : Event plane resolution STAR, PRC76, 024915 (2007)





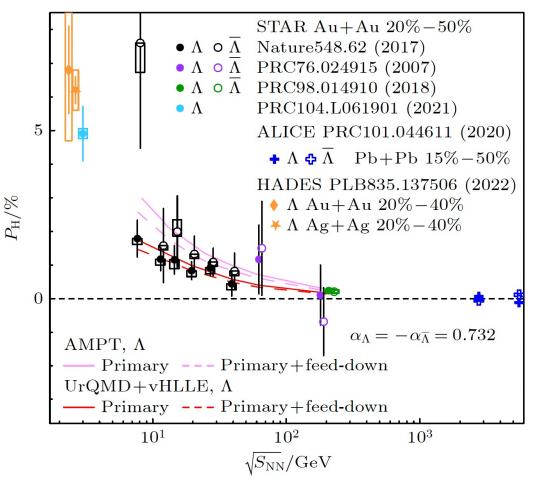
 $\Lambda \rightarrow p + \pi^-$ (BR:63.9%,c $\tau$ ~7.9cm)



#### Observation of $\Lambda$ global polarization



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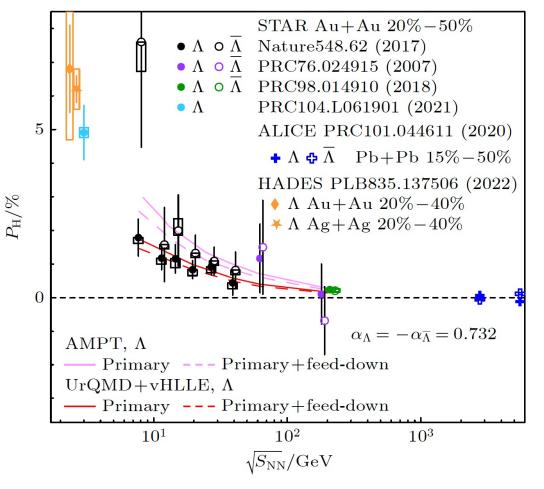


 $\square$  STAR, first measurement in AuAu 200 GeV,  $P_H < 2\%$ PRC 76, 024915 (2007) □ STAR, first observation in BES-I Nature 548, 62 (2017)  $\Box$  STAR, high precise  $P_H$  at 200 GeV PRC 90, 014910 (2018) □ ALICE, LHC energy region PRC 101, 044611 (2020)  $\square$  STAR,  $P_H$  at 3 GeV PRC 104, L061901 (2021) □ HADES energy region, consistent with STAR PLB 835,137506(2022)  $\square$  STAR,  $P_H$  at 19.6 and 27 GeV BES-II, no splitting PRC108,014910(2023) □ STAR, new recent results **BES-II Preliminary 7.7-17.3GeV**  $\Lambda, \Xi^{-}$  global polarization

## Energy dependence of $\Lambda$ global polarization



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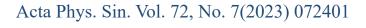
- Significant collision energy dependence, described well by various theoretical models
  - Liang and Wang, PRL 94,102301(2005),
  - Voloshin, nucl-th/0410089
  - Gao, Chen, Deng, Liang, QW, Wang, PRC 77, 044902(2008)
  - I. Karpenko and F. Becattini, EPJC(2017)77:213, UrQMD+vHLLE
  - H. Li et al., PRC 96, 054908 (2017), AMPT
  - Becattini, Lisa, Ann. Rev. Nucl. Part. Sci. 70, 395 (2020).
  - Huang, Liao, QW, Xia, Lect. Notes Phys. 987, 281 (2021).
  - Becattini, Rept. Prog. Phys. 85, No.12, 122301 (2022)
  - QW, Liang, Ma (editors), ActaPhys. Sin. 72, No. 7 & 11 (2023)

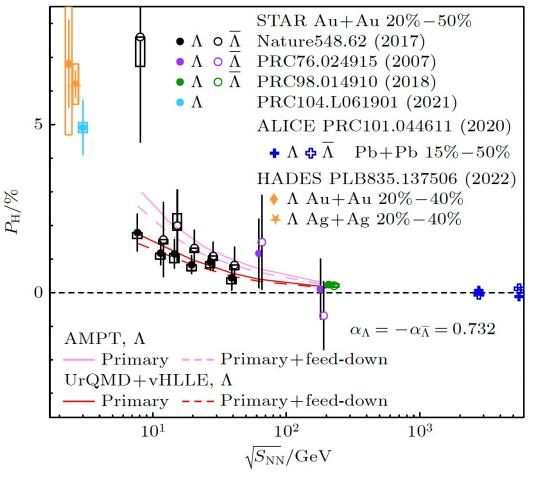
#### **D** Possible difference between $\Lambda$ and $\overline{\Lambda}$



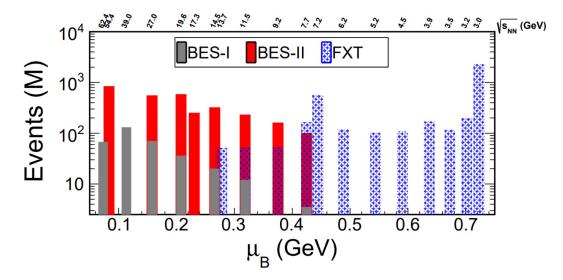
## Energy dependence of $\Lambda$ global polarization







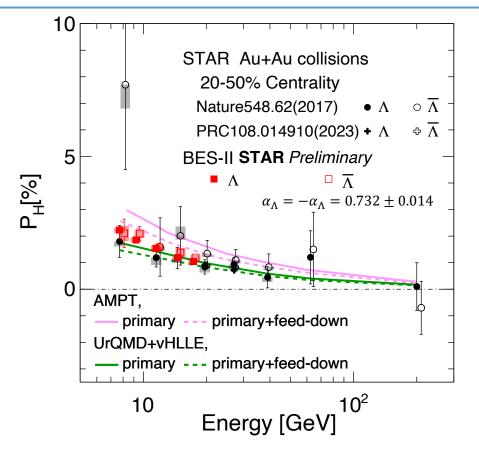
BES-I (2010-2017) and BES-II (2018-2021) statistics



Greatly improved precision from Beam Energy Scan phase-II

□ More significant global polarization in lower energies

#### Energy dependence of $\Lambda$ global polarization : from BES-II

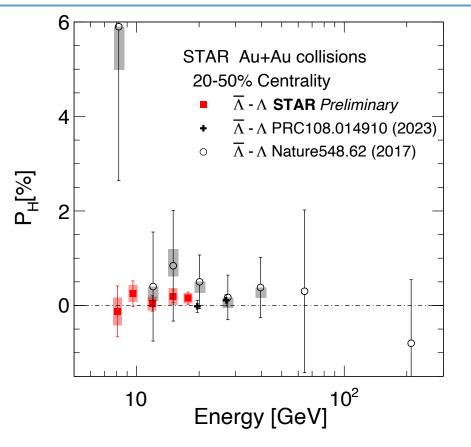


□ New STAR preliminary results at  $\sqrt{s_{NN}} = 7.7-17.3$  GeV from BES-II

Significant improvement in precision was achieved, collision energy dependence consistent with BES-I

#### Magnetic field effect of global polarization : from BES-II





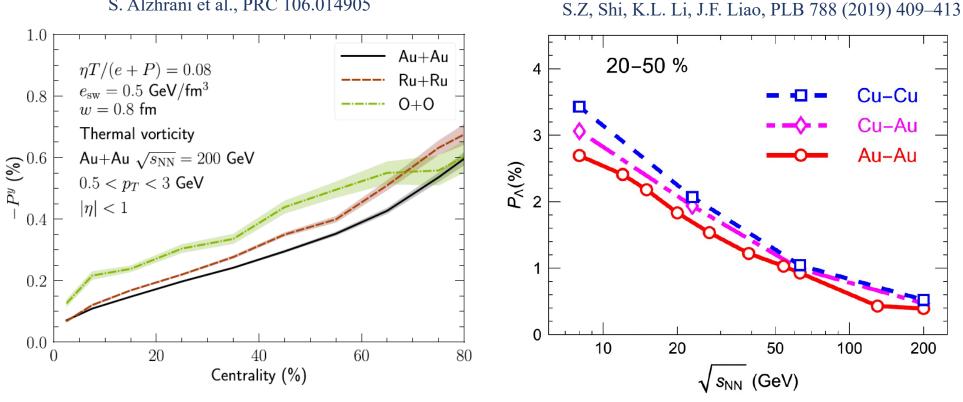
 $\square$  No splitting between  $\Lambda$  and  $\overline{\Lambda}$  global polarization within uncertainties

Upper limit on late stage magnetic field

- 95% confidence level STAR, PRC 108,014910(2023)
- $B < 9.4 \times 10^{12} T$  at 19.6 GeV
- $B < 1.4 \times 10^{13} T$  at 27 GeV

#### System size dependence of $\Lambda$ global polarization

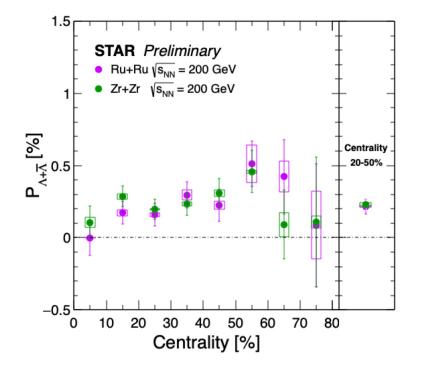




S. Alzhrani et al., PRC 106.014905

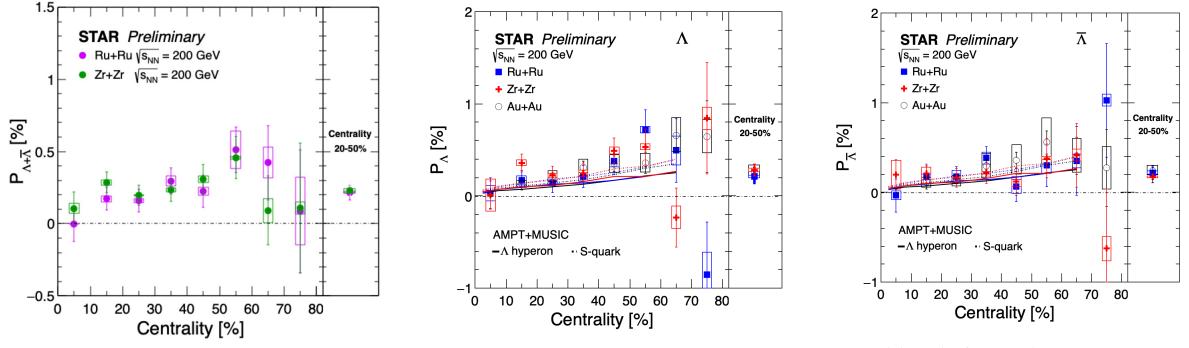
□ Longer system lifetime dilutes the vorticity/polarization Collision system size dependence of global polarization?  $^{197}_{79}Au > ^{96}_{44}Ru, ^{96}_{40}Zr > ^{63}_{29}Cu > ^{16}_{8}O$  $P_{\Lambda}^{Au} < P_{\Lambda}^{Ru} \approx P_{\Lambda}^{Zr} < P_{\Lambda}^{Cu} < P_{\Lambda}^{O}$ 





□ Significant global polarization observed in isobar collisions,  $P_{\Lambda}$  and  $P_{\overline{\Lambda}}$  increase with centrality

#### System size dependence of $\Lambda$ global polarization



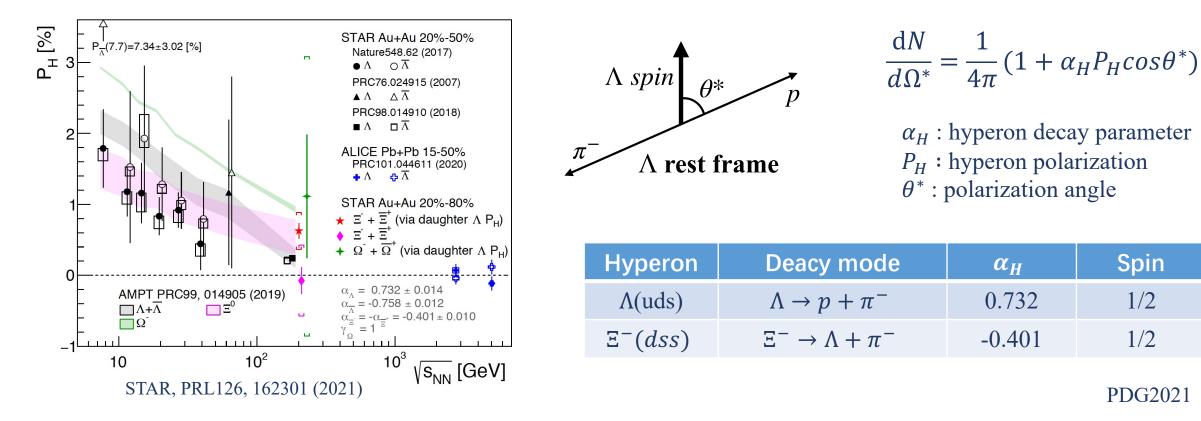
Model results from arXiv:2201.12970v1

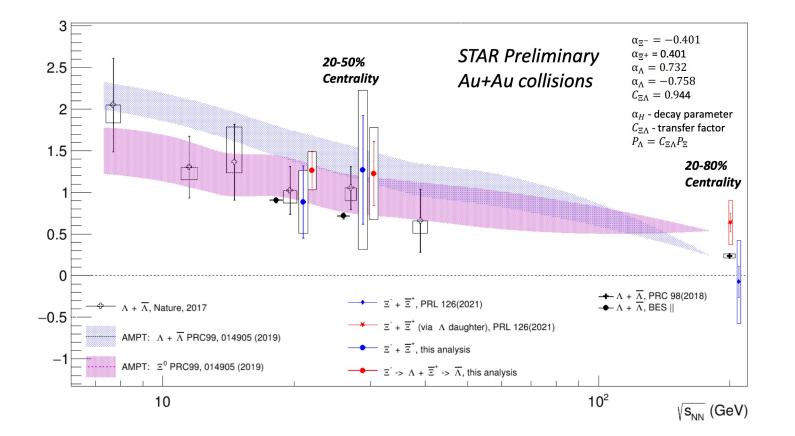
Significant global polarization observed in isobar collisions, P<sub>Λ</sub> and P<sub>Λ̄</sub> increase with centrality
 Global polarization of Λ + Λ̄ are consistent between Ru+Ru, Zr+Zr and Au+Au collisions within uncertainty

# $\Xi^- + \overline{\Xi}^+$ global polarization measurement

■ Possible larger Xi global polarization than Lambda due to earlier production and vorticity evolution

- Via daughter Lambda angle distribution in Xi rest frame
- Via daughter Lambda polarization with spin transfer factor( $C_{\Xi^- \to \Lambda} = 0.944$ )

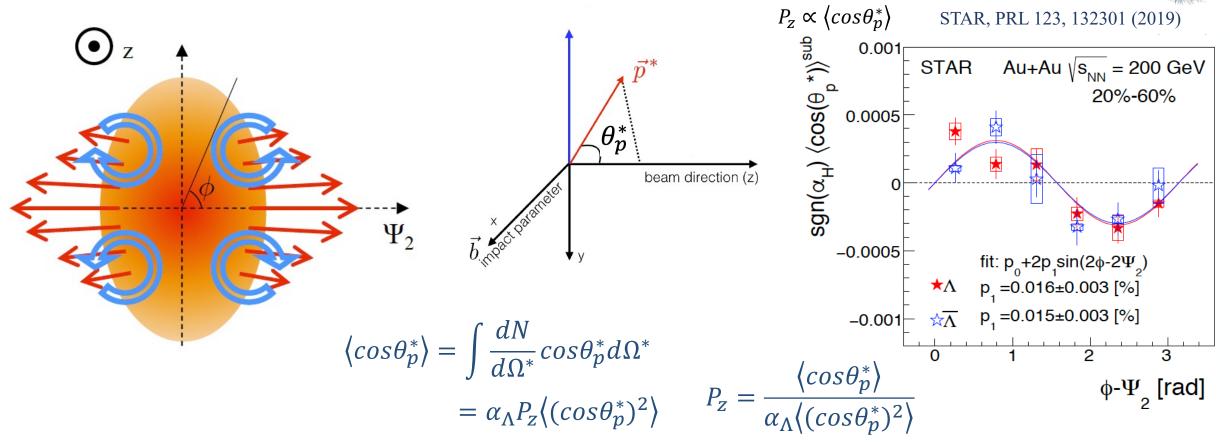




□ Significant  $\Xi^- + \overline{\Xi}^+$  global polarization observed in Au+Au at 19.6 and 27 GeV □  $\Xi^- + \overline{\Xi}^+$  global polarization measurement at lower BES-II energies underway

#### Local vorticity and polarization in heavy ion collisions

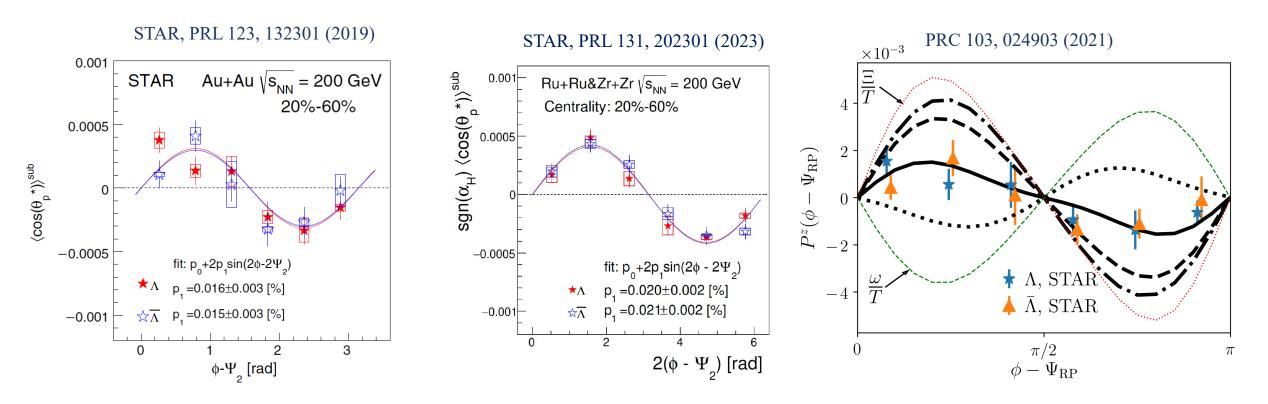




□ Elliptic flow indicates stronger expansion in-plane than out of plane

 $\implies$  Lead to polarization along the beam direction ( $P_z$ )

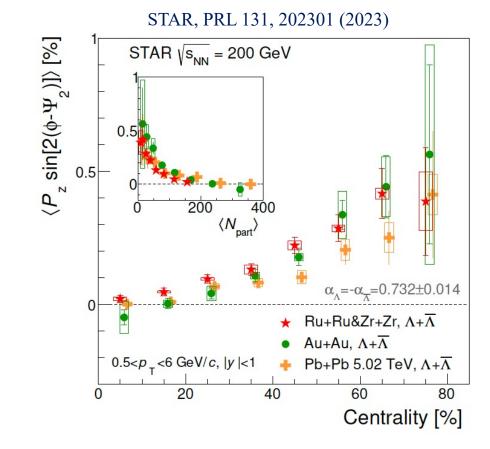




Clear azimuthal angle dependence observed in Au+Au and isobar collisions at 200 GeV
 New developments, Shear Induced Polarization(SIP), can describe trend of data

#### System size dependence of $P_z$



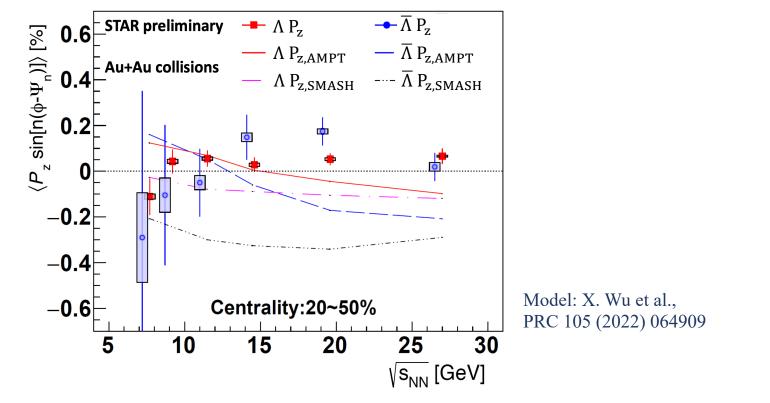


 $\square$  *P<sub>z</sub>* from isobar collision comparable to Au+Au and Pb+Pb

 $\checkmark\,$  No significant system size dependence observed at same energy

#### Energy dependence of $P_z$



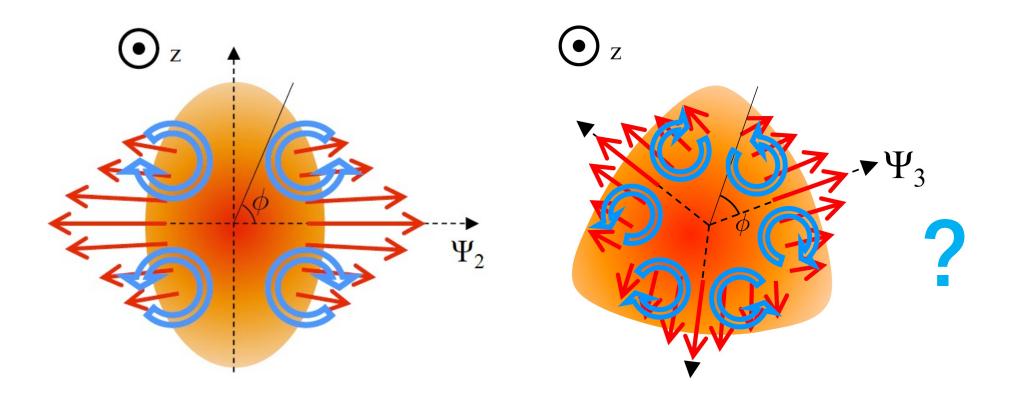


 $\square$  *P<sub>z</sub>* in Au+Au collisions comparable from 7.7 to 200 GeV, Pb+Pb collision at 5.02 TeV

- ✓ No significant collision energy dependence observed, hints of sign change at 7.7 GeV
  - $0.098 \pm 0.014(stat.)^{+0.019}_{+0.018}(syst.)$  in Au+Au 200 GeV STAR, PRL 123, 132301 (2019)
  - $0.082 \pm 0.011(stat.) \pm 0.014(syst.)$  in Pb+Pb 5.02 TeV ALICE, PRL128, 172005 (2022)

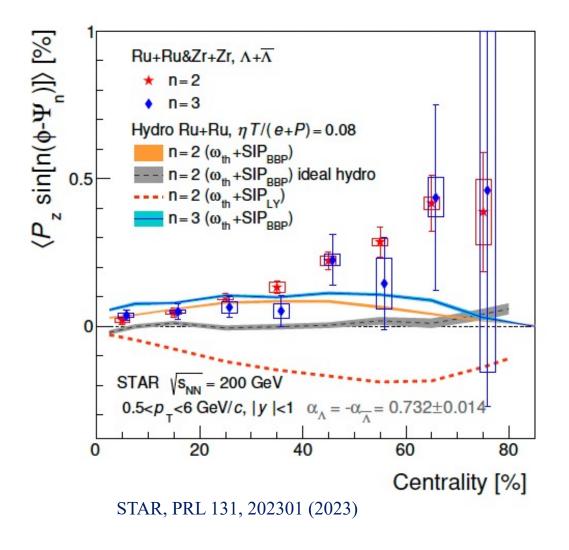
#### $P_z$ from higher harmonic flow





# Measurements $P_z$ relative to higher harmonic event planes provide new insights into polarization phenomena

## $P_z$ relative to third order event plane



- $\square Significant P_z w.r.t third-order event plane observed$
- $\square P_z \text{ w.r.t second-order event plane increases with centrality}$
- Comparable  $P_z$  w.r.t second and third order event plane, indicating  $v_3$ -driven polarization
- Hydrodynamic models with shear term reasonably describe the data for central collisions, but not for peripheral collisions

S. Alzhrani et al., PRC 106.014905

# Summary



#### Global polarization

- □ Significant improvement in precision was achieved in BES-II
- No splitting observed between  $\Lambda$  and  $\overline{\Lambda}$  global polarization in Au+Au collisions at 7.7 27 GeV and  $^{96}_{44}$ Ru +  $^{96}_{44}$ Ru,  $^{96}_{40}$ Zr +  $^{96}_{40}$ Zr collisions at 200 GeV
- □ No collision system size dependence between Ru+Ru, Zr+Zr and Au+Au collisions at 200 GeV
- □ Significant  $\Xi^- + \overline{\Xi}^+$  global polarization observed at 19.6, 27 GeV, measurements in lower energies underway

#### Polarization along beam direction $(P_z)$

- $\square$  First observation of polarization along beam direction ( $P_z$ ) w.r.t third-order event plane
- **\square** No significant system size dependence of  $P_z$  observed at same energy
- No significant collision energy dependence of  $P_z$  observed

# Summary



#### Global polarization

- □ Significant improvement in precision was achieved in BES-II
- No splitting observed between  $\Lambda$  and  $\overline{\Lambda}$  global polarization in Au+Au collisions at 7.7 27 GeV and  $^{96}_{44}$ Ru +  $^{96}_{44}$ Ru,  $^{96}_{40}$ Zr +  $^{96}_{40}$ Zr collisions at 200 GeV
- □ No collision system size dependence between Ru+Ru, Zr+Zr and Au+Au collisions at 200 GeV
- □ Significant  $\Xi^- + \overline{\Xi}^+$  global polarization observed at 19.6, 27 GeV, measurements in lower energies underway

#### Polarization along beam direction $(P_z)$

- $\square$  First observation of polarization along beam direction ( $P_z$ ) w.r.t third-order event plane
- □ No significant system size dependence of  $P_z$  observed at same energy
- No significant collision energy dependence of  $P_z$  observed

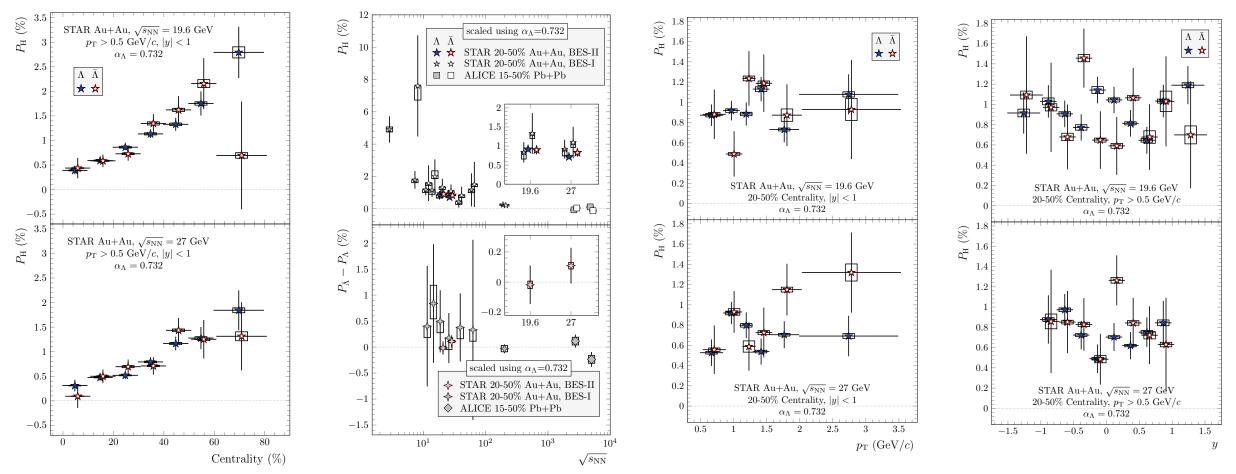






### Global polarization collision energy dependence

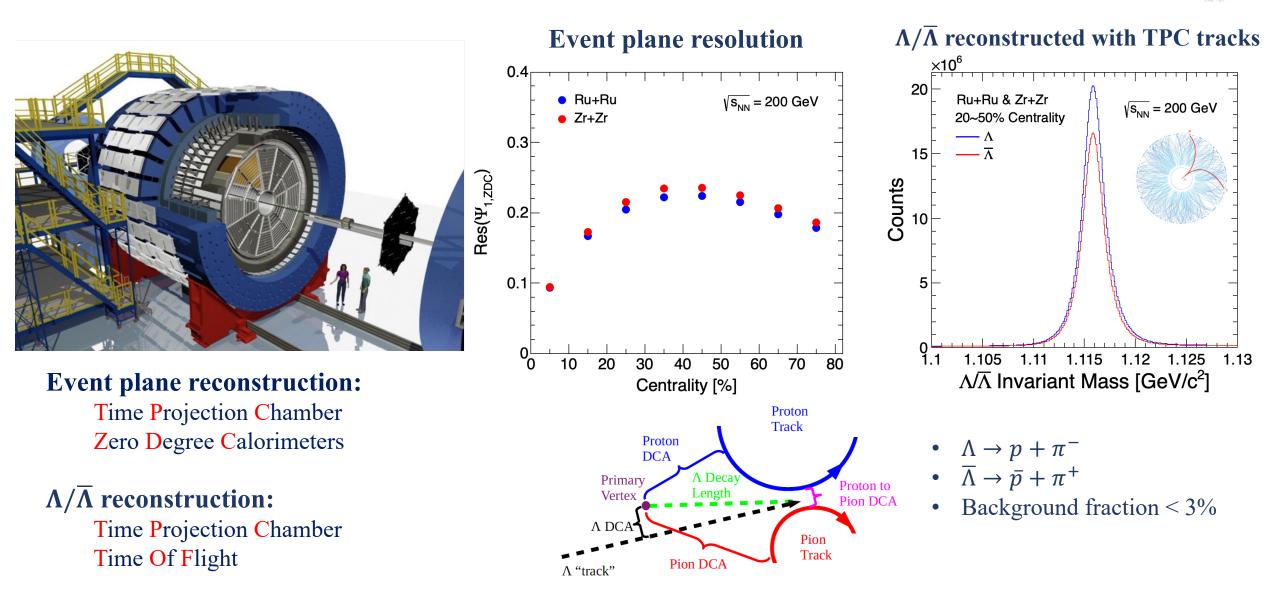




Significant global polarization centrality dependence observed
 Lambda and AntiLambda global polarization are consistent
 No observed dependence of global polarization on *p<sub>T</sub>*

## STAR detector and $\Lambda/\overline{\Lambda}$ reconstruction

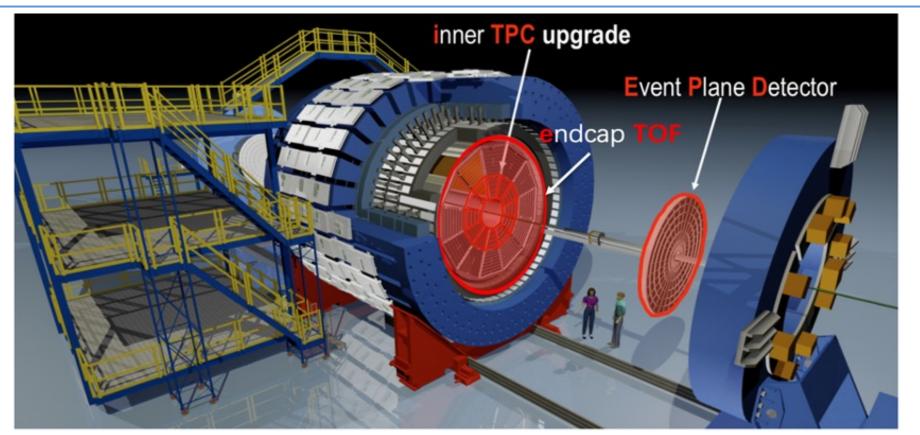




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#### STAR detector

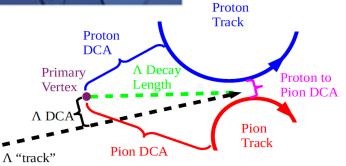


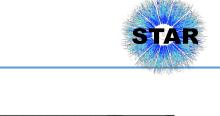


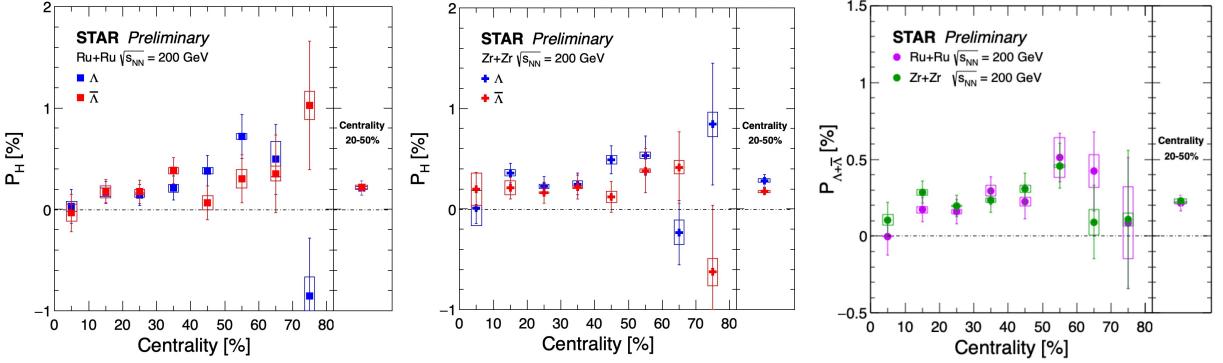
Event plane reconstruction: Time Projection Chamber Event Plane Detector

Zero Degree Calorimeters

 $\Box \ \Lambda/\overline{\Lambda} \ reconstruction:$ Time Projection Chamber
Time Of Flight



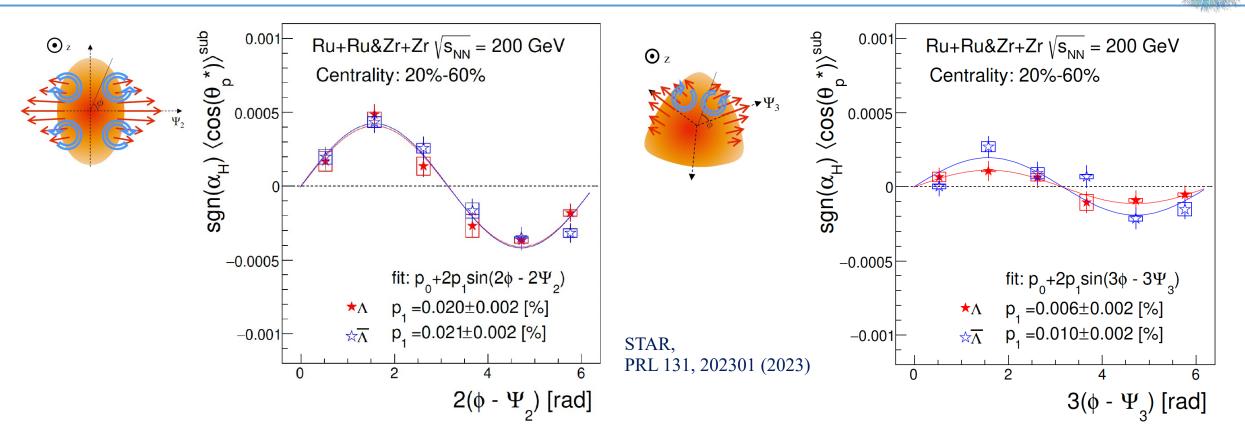




□ Significant global polarization observed,  $P_{\Lambda}$  and  $P_{\overline{\Lambda}}$  increase with centrality

- **D** No significant difference between  $P_{\Lambda}$  and  $P_{\overline{\Lambda}}$  in Ru+Ru and Zr+Zr collisions
- **G**lobal polarization of  $\Lambda + \overline{\Lambda}$  are consistent between Ru+Ru and Zr+Zr collisions

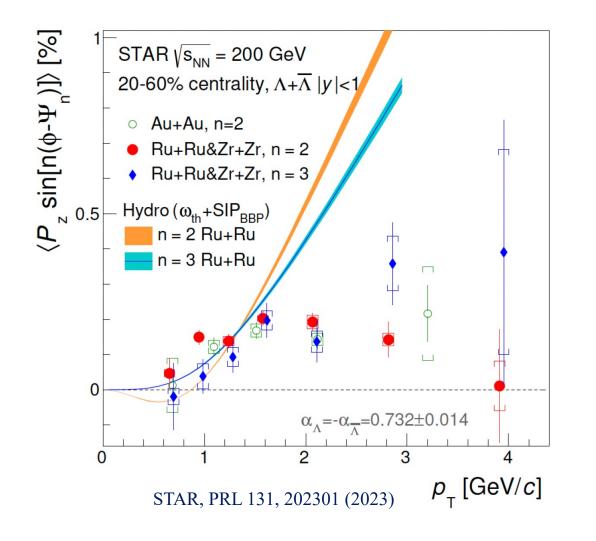
## Local polarization in Ru+Ru&Zr+Zr at 200 GeV



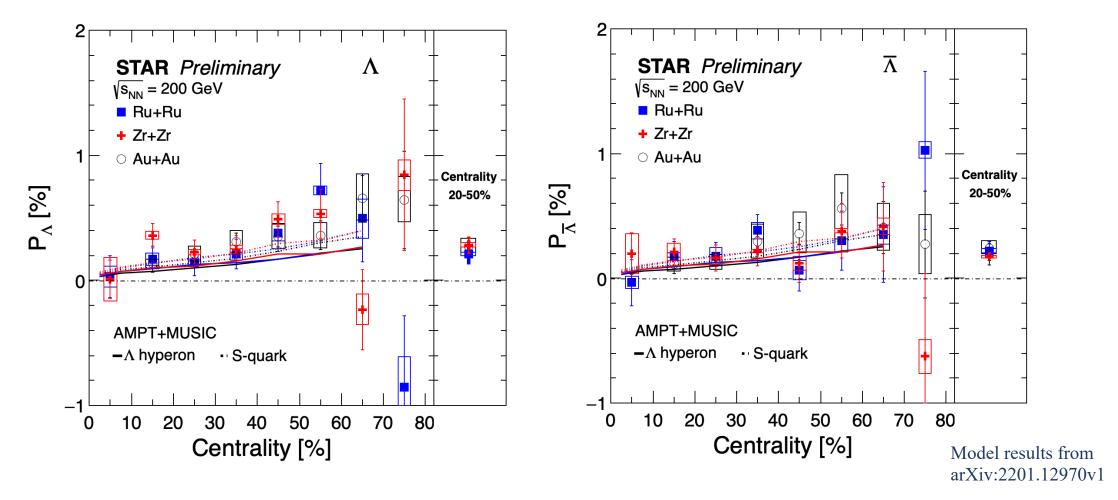
□ Significant local polarization w.r.t second-order event plane observed in isobar collisions

□ First observation of local polarization w.r.t the third-order event plane





- $\square$  Local polarization  $p_T$  dependence is observed
- □ Observed  $p_T$  dependence similar to that of elliptic  $(v_2)$  and triangular  $(v_3)$  flow
- Results are consistent between isobar and Au+Au collisions

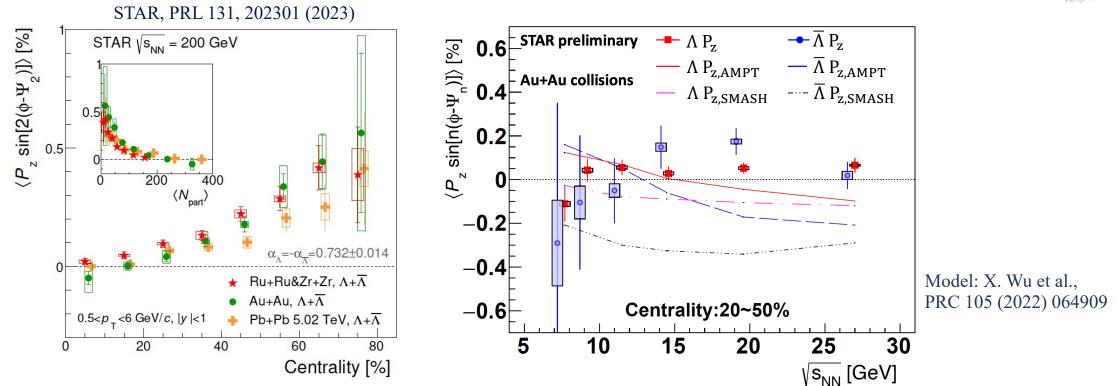


**G**lobal polarization of  $\Lambda$  and  $\overline{\Lambda}$  are consistent in isobar and Au+Au collision systems

STAR

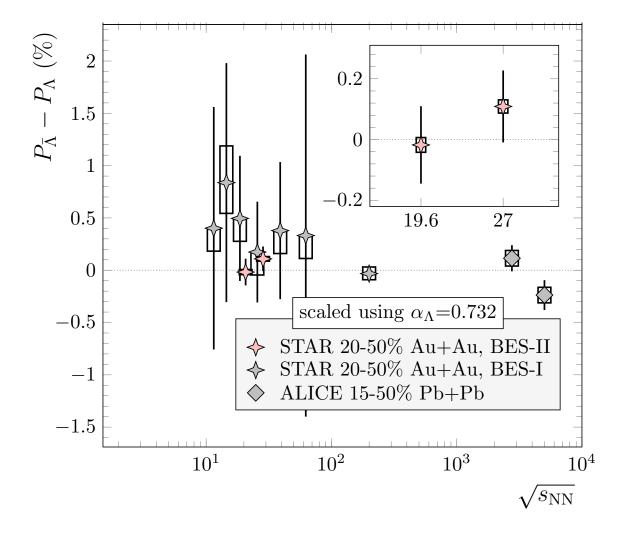
### Energy dependence of $P_z$





- $\square$  *P<sub>z</sub>* from isobar collision comparable to Au+Au and Pb+Pb
  - $\checkmark$  No significant system size dependence observed at same energy
- $\square$  *P<sub>z</sub>* in Au+Au collisions comparable from 7.7 to 200 GeV, Pb+Pb collision at 5.02 TeV
  - $\checkmark\,$  No significant collision energy dependence observed





**\square** No splitting of  $\Lambda / \overline{\Lambda}$  observed

Au+Au	19.6 GeV	27 GeV
$\begin{array}{c} P_{\overline{\Lambda}} - P_{\Lambda} \\ (\%) \end{array}$	$-0.018 \pm 0.127(stat.) \pm 0.024(sys.)$	0.109 ±0.118(stat.) ± 0.022(sys.)

 $\square |B| \approx \frac{T_s |P_{\overline{\Lambda}} - P_{\Lambda}|}{2|\mu_A|}, \text{ using hydrodynamics}$ 

 $T_s = 150 \text{ MeV}$ : the temperature of the emitting source  $\mu_A = -1.93 \times 10^{-14} \text{ MeV/T}$ : the magnetic moment of the  $\Lambda$  hyperon

#### Upper limit on late stage magnetic field

- 95% confidence level
- $B < 9.4 \times 10^{12} T$  at 19.6 GeV
- $B < 1.4 \times 10^{13} T$  at 27 GeV