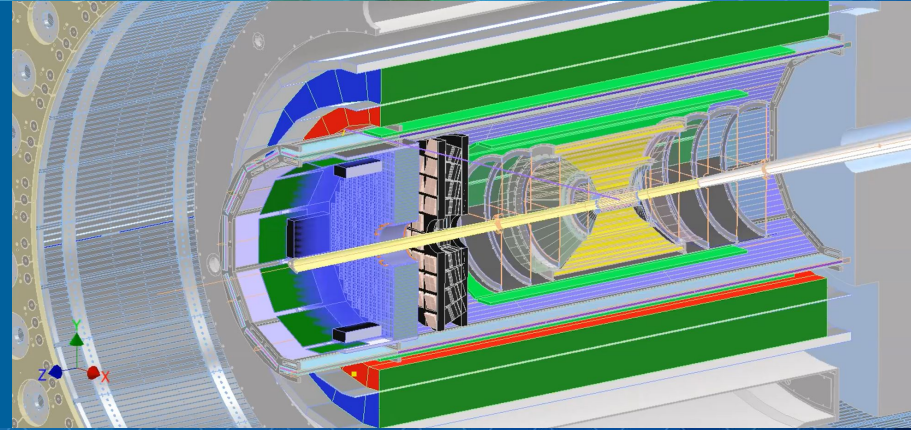


BIC Simulation Meeting

BIC Dynamic Range Simulation Updates + Layer Combination Studies

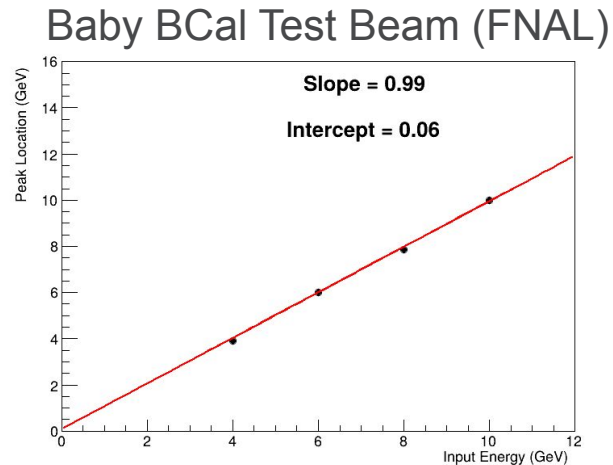


Maria Zurek

SiPM Considerations - Dynamic Range

Estimate for
review

- On the high energy side of the spectrum, we have F-DET-ECAL-BAR.2
 - *Shall provide photon measurements up to 10 GeV*
- SiPM readout consists of 57324 pixels per channel
 - When a pixel fires, it cannot register another signal for some time
 - Large signals can cause non-linearity due to multiple photons striking the same pixel
- 10 GeV photon will yield ~7700 p.e. at max per channel
 - Corresponds to ~14% of max occupancy
 - Slight non-linearity expected, can be corrected for
- Baby BCal beam tests showed no sign of saturation-induced non-linearity for 10 GeV electrons
 - This demonstrates linearity for BIC up to ~6 GeV, after scaling for different photon detection efficiency

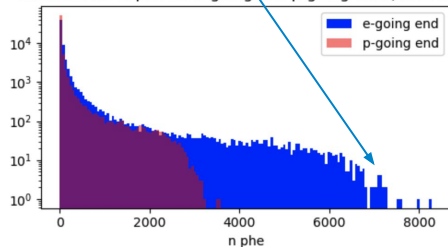


Max Number of Photoelectrons

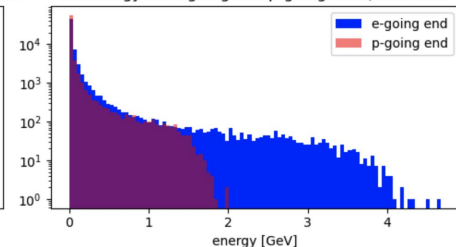
Full simulations

~7000 phe

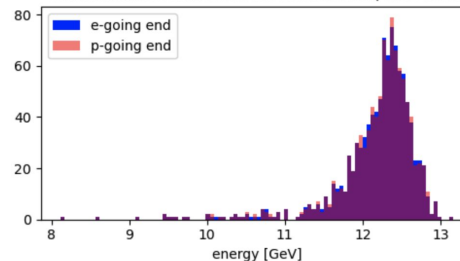
Smearred nb of phe for e-going and p-going ends, NKS fiber



Hit energy for e-going and p-going ends, NKS fiber

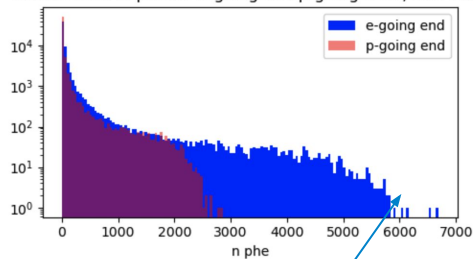


Summed hit E corrected for attenuation, NKS fiber

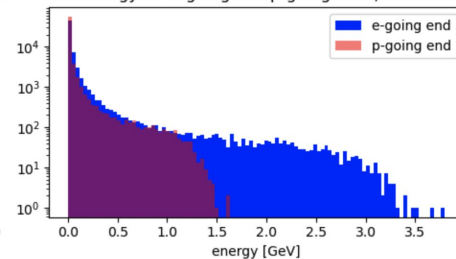


10 GeV photons
at $\eta \sim -1.7$
**fsam tuned to 5
GeV e^- at $\eta = 0$**

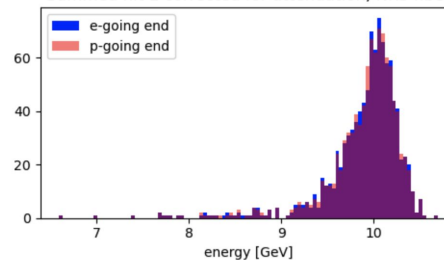
Smearred nb of phe for e-going and p-going ends, NKS fiber



Hit energy for e-going and p-going ends, NKS fiber



Summed hit E corrected for attenuation, NKS fiber



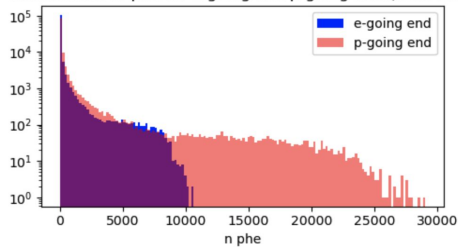
10 GeV photons
at $\eta \sim -1.7$
**fsam tuned to
this sample**

~6000 phe

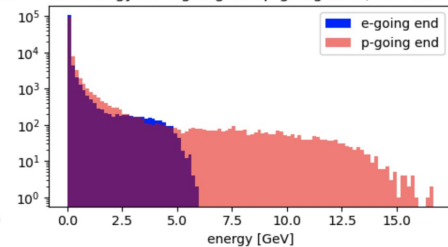
Max Number of Photoelectrons

Full simulations

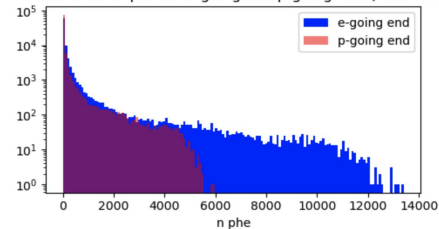
Smearred nb of phe for e-going and p-going ends, NKS fiber



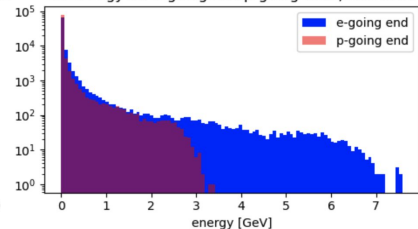
Hit energy for e-going and p-going ends, NKS fiber



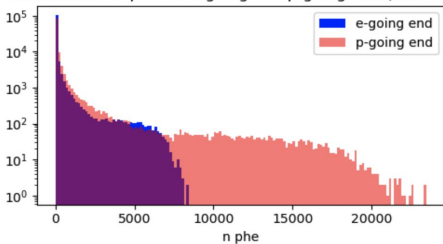
Smearred nb of phe for e-going and p-going ends, NKS fiber



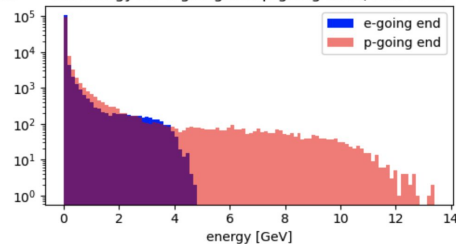
Hit energy for e-going and p-going ends, NKS fiber



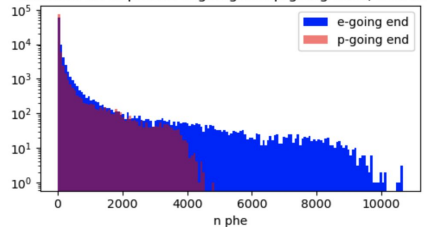
Smearred nb of phe for e-going and p-going ends, NKS fiber



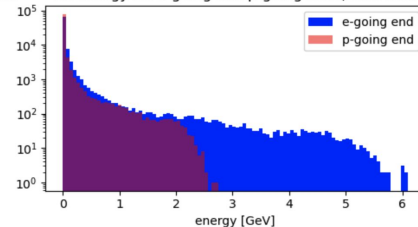
Hit energy for e-going and p-going ends, NKS fiber



Smearred nb of phe for e-going and p-going ends, NKS fiber



Hit energy for e-going and p-going ends, NKS fiber



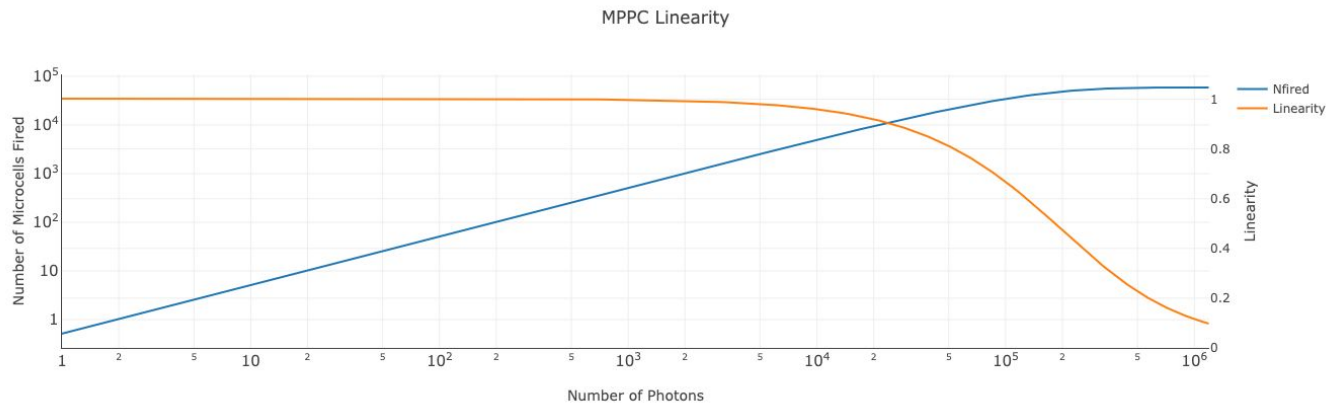
50 GeV electrons at $\eta \sim 1.4$

- top: fsam tuned to 5 GeV e^- at $\eta = 0$
- bottom: fsam tuned to this sample

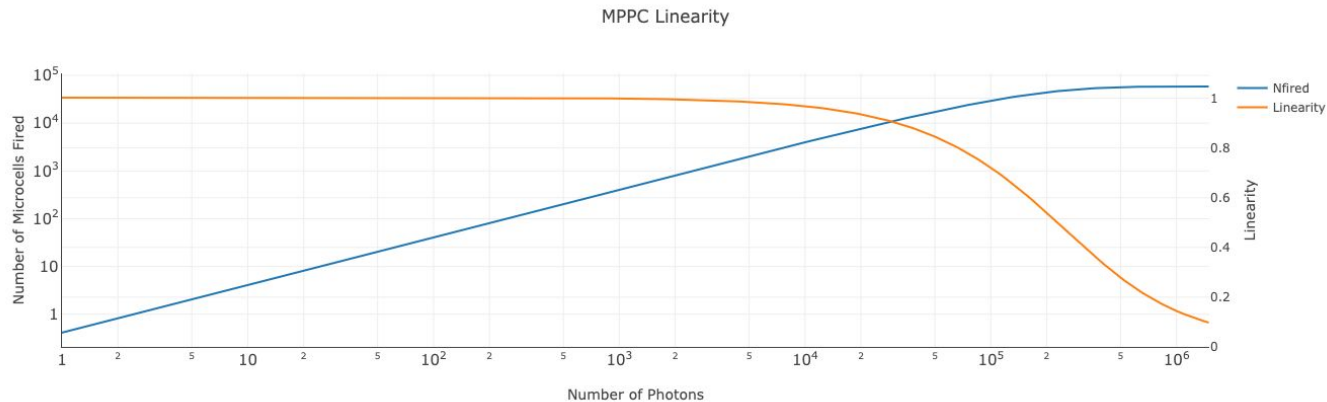
19 GeV electrons at $\eta \sim -1.7$

- top: fsam tuned to 5 GeV e^- at $\eta = 0$
- bottom: fsam tuned to this sample

Nonlinearity



50% PDE folded in
(S14 Family)



40% PDE folded in
(S13 Family)

Pixel Size and Number of Pixels

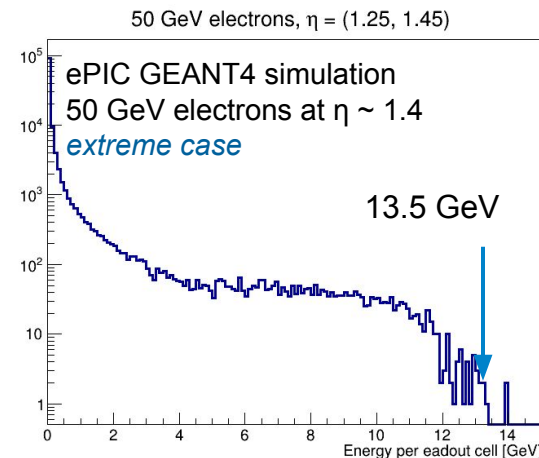
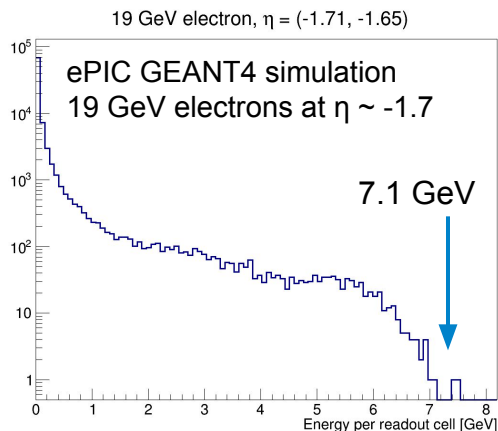
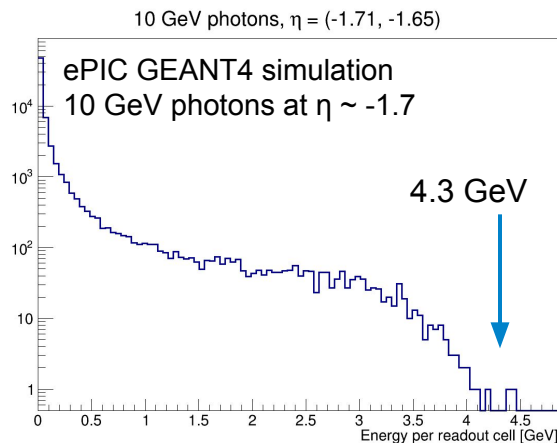
FDR
simulations

Defined by **photoelectron statistics** and **energy range** to be measured

Energy measurement ranges in BECal:

- Shall provide photon measurements up to 10 GeV (F-DET-ECAL-BAR.2) and down to 100 MeV (F-DET-ECAL.9)
- Shall provide electron ID up to 50 GeV and down to 1 GeV and below (F-DET-ECAL-BAR.1)
 - Electron energy measurement needed for e/π separation only (straightforward at high energies)
- Reasonable performance for MIPs needed for calibration and for muon ID

Largest energy deposit occurs for particles at large η (steep angle) where the pathlength in a cell is maximal and the attenuation is minimal.



Photoelectron statistics

From our 2023 Hall D tests using GlueX SiPMs and double-clad Kuraray fibers: **1000 phe/GeV** per side for showers at the center of the Baby BCAL prototype

- Corrected for attenuation: **1077 phe/GeV*** per side

We can scale these results for the **ePIC Barrel ECal***:

- x 1.5 factor improvement in **SiPM photon detection efficiency**
- x 1.16 factor to account for **better optical coupling**
- x 0.69 reduction accounting for **single-clad Kuraray fibers**

This gives **~ 1239 phe/GeV** per side (fully corrected for attenuation)

- **10 GeV γ at $\eta \sim -1.7$: 5560 phe \rightarrow **9.8 % max SiPM occupancy****
- **19 GeV e^- at $\eta \sim -1.7$: 9181 phe \rightarrow **16.1 % max SiPM occupancy****
- **50 GeV e^- at $\eta \sim 1.4$ (most extreme case): 17456 phe \rightarrow **30.1% max SiPM occupancy****

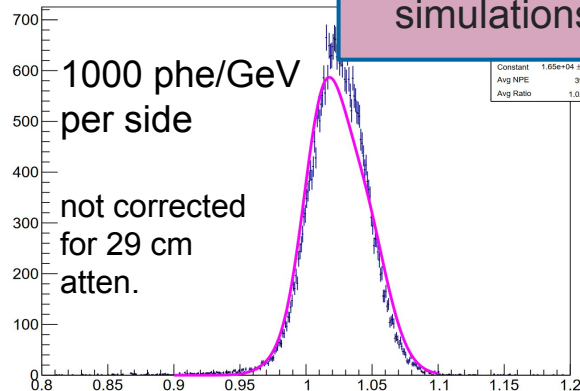
Well below the region where large nonlinearities in the SiPM response are expected in almost all cases.

Small non-linear effects possible for some ultra-high energy electrons, which is acceptable ($e-\pi$ separation straightforward).

* See backup slide for the attenuation length measurement and extraction of those factors

2023 Hall D, Baby BCAL

FDR simulations



2008 Hall B beam test, photons

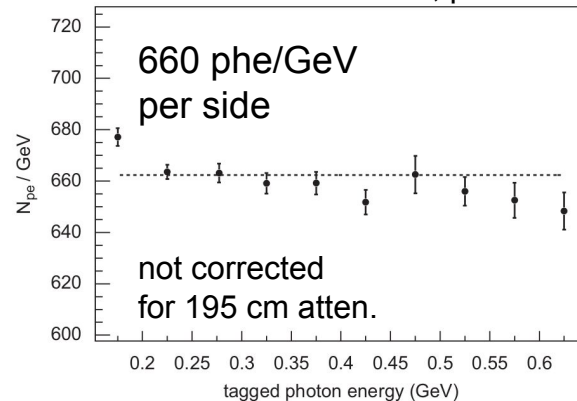


Fig. 16. The number of photoelectrons per GeV per end of the BCAL module is shown as a function of energy. A one parameter fit is plotted (dashed line). For more details see the text.

Layer Configuration

AstroPix Layer Placement

Layer Placement (1-3-4-6): General Motivation

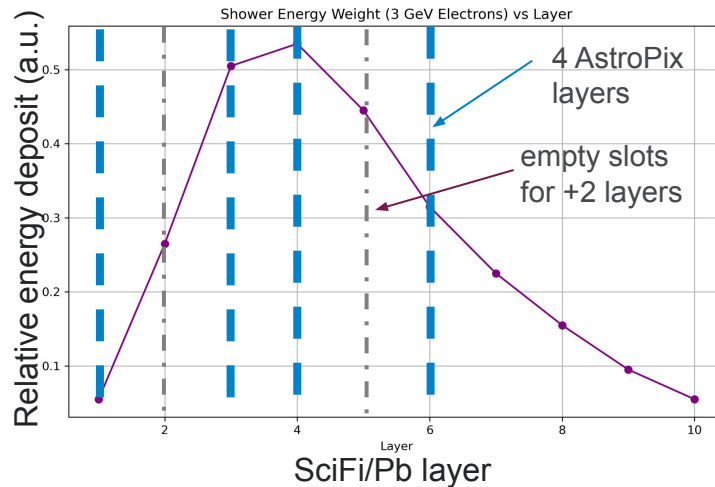
6 imaging layers separated by $1.45X_0$ at $\eta = 0$ of Pb/ScFi

All layers important for the e/π separation for mid energy particles <5 GeV and overall sampling of shower energy for SciFi/Pb close shower energy splitting

Front layers important for γ/π^0 separation and position resolution

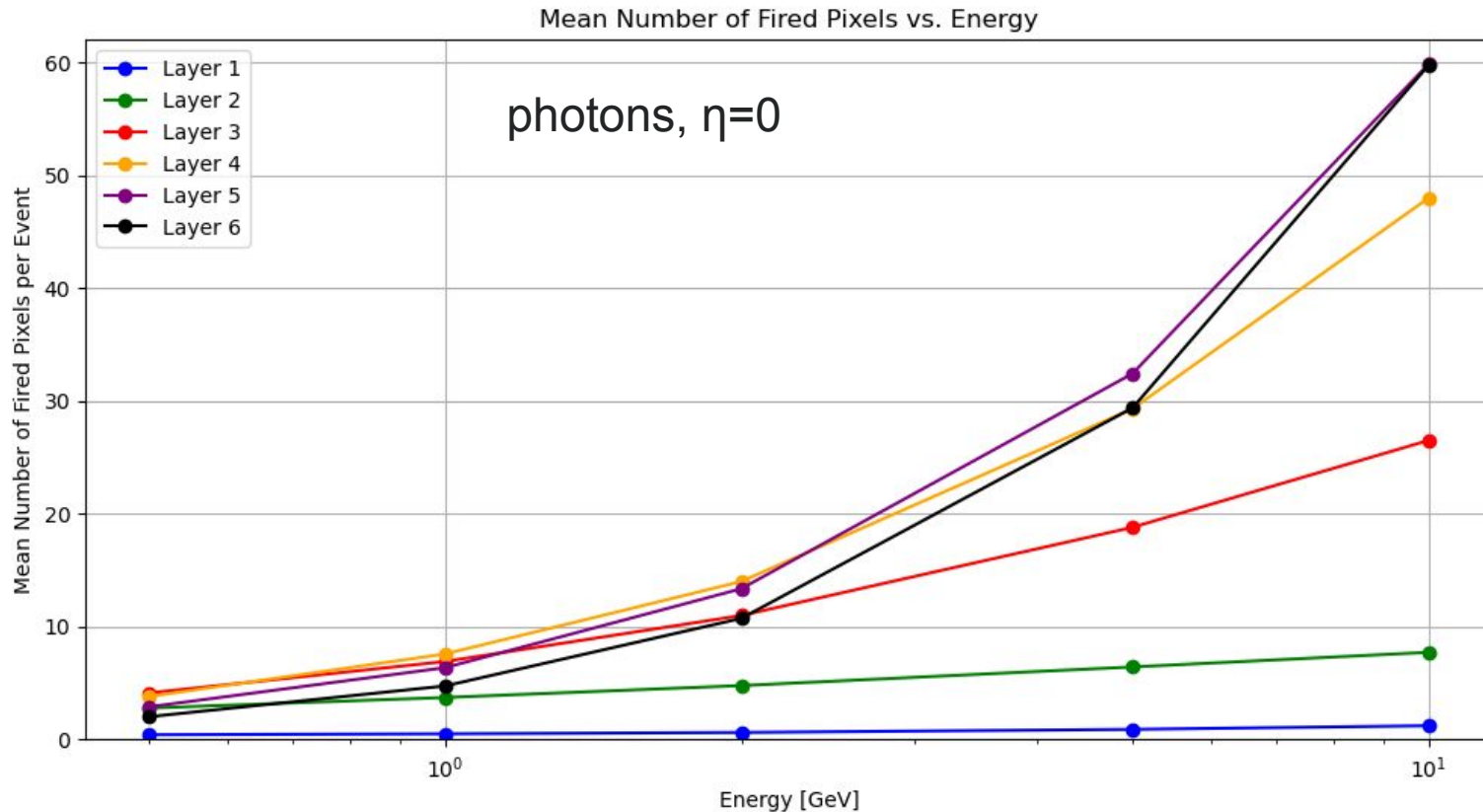
- 1st layer in front of the calorimeter: effectively a tracking layer for charged particles to support DIRC PID - very little “calorimetric” performance
- 1 pre-shower slot empty (impact on γ/π^0 separation and position resolution)
- 2 layers around shower max (sample much total of shower for energy reconstruction and shower separation and e/π separation)
- 1 post-shower slot empty (important sampling overall shower energy, e/π)
- 1 layer in tail (deeper in the tail for larger η to catch e/π separation and still sample important part of shower energy)

3 GeV electron shower profile at $\eta = 0$



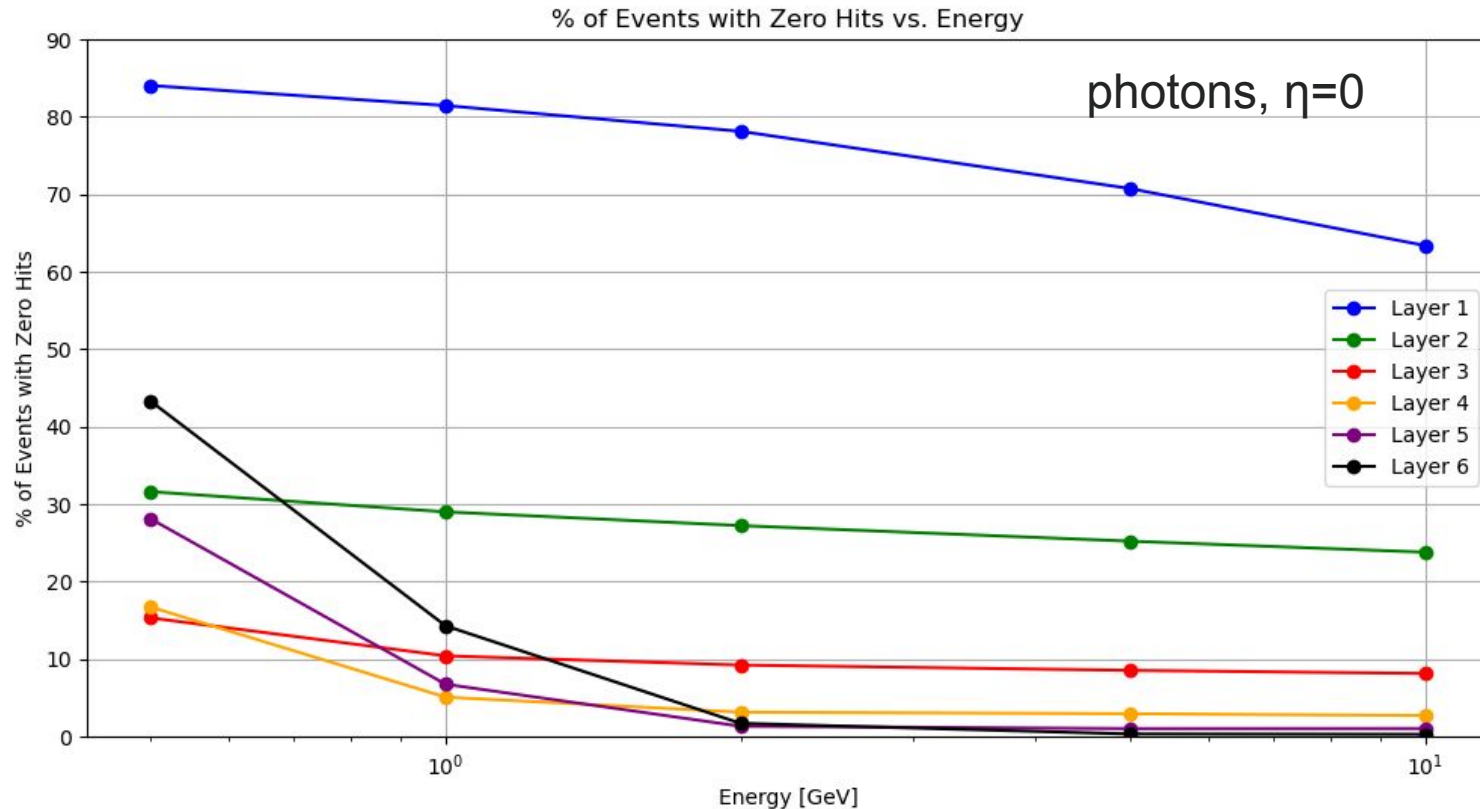
Optimized for preserving e/π separation for mid energy particles and max shower sampling for effective 3 calorimetric layers only

Mean hit multiplicity per AstroPix Layer vs Energy



Note that mean includes the cases when there is no hits at all

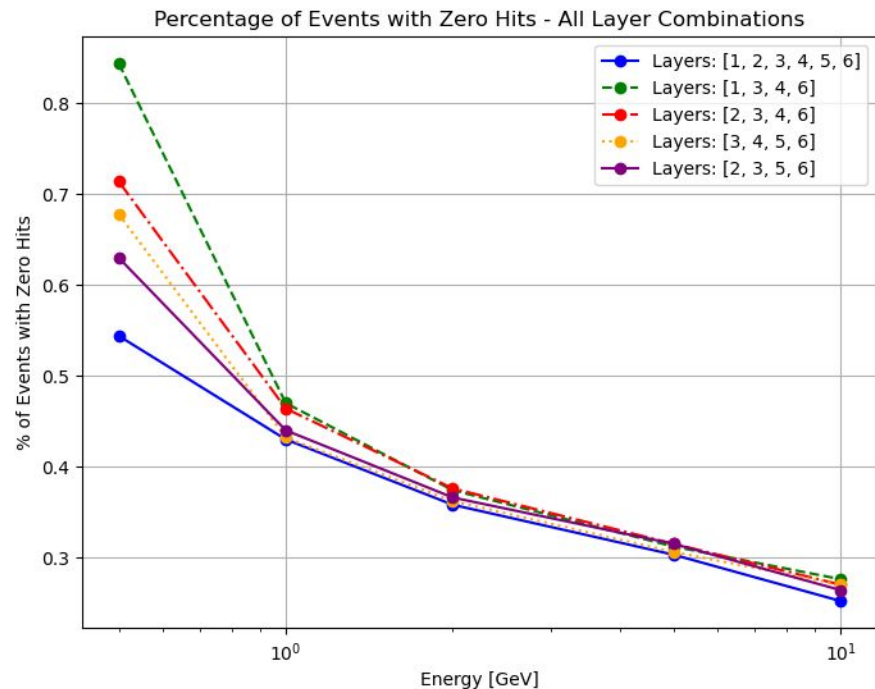
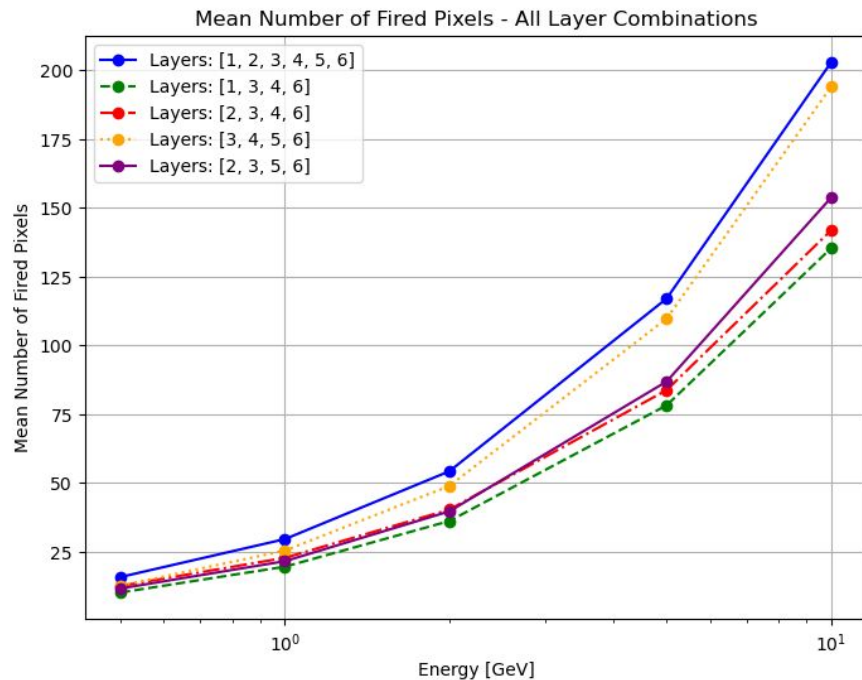
% of events with zero hits in the layers



See backup for example distributions of nb of hits per layer

Mean hit multiplicity and % of zero hits in all AstroPix layers

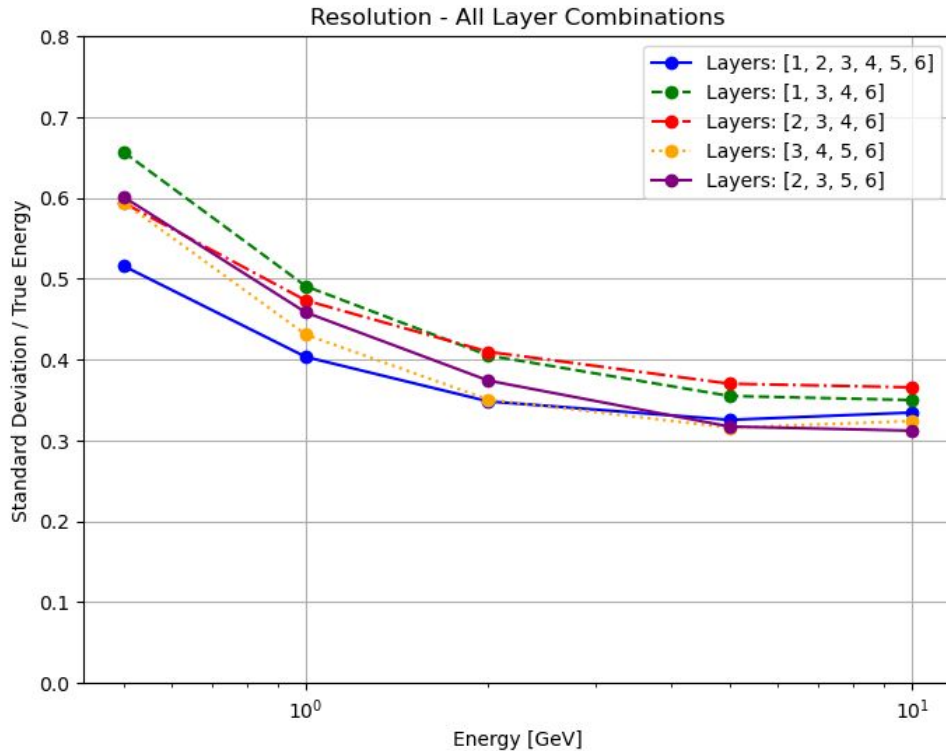
Different layer configurations



Note that this is for photons at $\eta=0$, different η will differ

Energy Resolution

Different layer configurations



Plot show Standard Deviation of energy deposit in AstroPix layers

Photons, $\eta=0$: for low energy response at this rapidity, 3-4-5-6, 2-3-5-6, 2-3-4-6 look preferable

For high energy, overall energy reconstruction affected by longitudinal shower (and it's shower max) fluctuations.

Extreme example for $\eta=0$, at larger η , more confinement