





EIC Collider Performance

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With thanks to E. Aschenauer, W. Bergan, other EIC colleagues

Uncovering New Laws of Nature at the EIC November 20, 2024

Electron-Ion Collider

Outline: EIC Collider Performance

- Present Design Approach
- Parameters for Luminosity
- Luminosity Optimization Parameters
 - Luminosity demonstration/game
- CD-1 Design Luminosity Curve
 - Low E_{cm} : Space charge beam dynamics
 - Mid E_{cm}: Beam-beam dynamics
 - High E_{cm} : Synchrotron radiation power/RF power
- Proposed Early Science Program
 - Luminosity curves and limits
- Luminosity Evolution and Complexity



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EIC

iosity [10³³

80 100 120

Center of Mass Energy [GeV]

0 MW SB limit

The gap was due to ncomplete understanding of several beam physics

issues and a number technical difficulties

2001 2002 2003 2004 2005 2006 2007 2008 2009 201

Present EIC Design Approach

Ultimate EIC Performance Parameters:

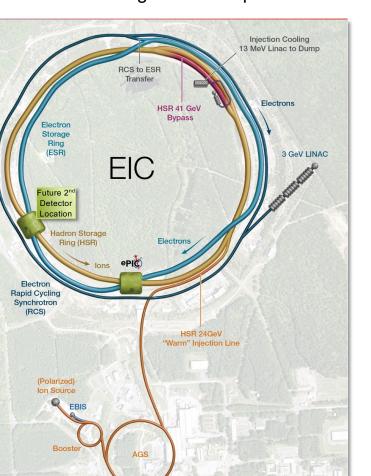
- High Luminosity: L= 10³³ 10³⁴ cm⁻² s⁻¹
- Highly Polarized Beams: 70%
- Large Center of Mass Energy Range: E_{cm} = 28 140 GeV
- Large Ion Species Range: protons Uranium
- Large Detector Forward Acceptance and Good Background Conditions
- Possibility to Implement a Second Interaction Region (IR)/Detector

Accelerator Status in a glance:

- ✓ Polarized ion/proton source
- ✓ Ion injection and initial acceleration systems Linac (200 MeV), Booster (1.5 GeV), AGS (25 GeV)
- multi Hadron Storage Ring (40-275 GeV) HSR
- Electron Pre-Injector (3 GeV) EPI
- Electron Rapid Cycling Synchrotron (3 GeV top energy) RCS
- Electron Storage Ring (5 GeV 18 GeV) ESR
- Interaction Region(s) IR
- Strong Hadron Cooler System SHC

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S. Nagaitsev – Sep 2024 MAC

Luminosity Optimization Parameters

 $L \propto f_{coll} N_1 N_2 / \sigma_x^* \sigma_y^*$ f_{coll} : collision frequency $N_{1,2}$: particles per bunch $\sigma_{x,y}^*$: (equal) beam sizes at IP

Every parameter optimized separately and collectively in the EIC design

Try multiplying out the given numbers – should be very close to 10^{34} cm⁻²s⁻¹

- Maximize collision frequency (~90 MHz)
 - · Limited by kicker rise times
 - Limited by parasitic collisions (crabbing)
- Maximize particles per bunch (~10¹¹)
 - Limited by sources, space charge
 - Limited by collective effects
 - Interaction of beam with impedances
 - Also total currents: $I_{1,2}=q_{1,2}N_{1,2}f_{coll} \sim 1-3A$
- Minimize beam sizes at IP (~250/25 um)
 - Limited by IR focusing, magnets
 - Limited by chromatic dynamic aperture
 - Limited by emittance growth (IBS)

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Just For Fun: Final Focus/Luminosity Game

 $L \propto f_{\rm coll} N_1 N_2 / \sigma_x^\star \sigma_y^\star$

 f_{coll} : collision frequency $N_{1,2}$: particles per bunch $\sigma_{x,y}^{\star}$: (equal) beam sizes at IP

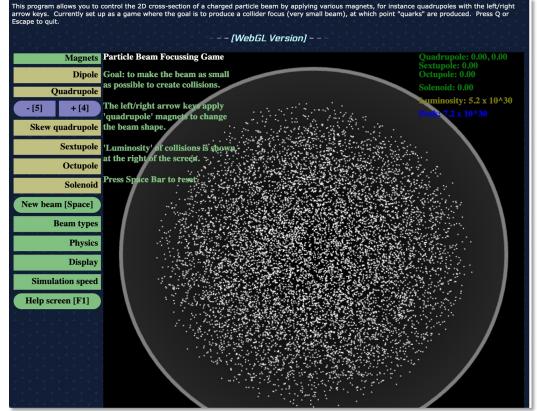
http://stephenbrooks.org/ap/beam2d/

Can **you** achieve a peak luminosity of 10^{34} cm⁻² s⁻¹?

- Use the [5] and [4] keys to horizontally focus/defocus)
- Click Help Screen for other keys/magnets



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CD-1 Design Luminosity Curve

	Electrons	Protons				
Beam energies	2.5 - 18 GeV	41- 275 GeV				
Center of mass energy range	E _{Cm} = 20-140 GeV					
	Electrons	Protons				
Beam energies	10 GeV	275 GeV				
Center of mass energy	E _{Cm} =	105 GeV				
number of bunches	nb =	1160				
crossing angle	25 mrad					
Bunch Charge	1.7·10 ¹¹ e	0.7·10 ¹¹ e				
Total beam current	2.5 A	1 A				
Beam emittance, horizontal	20 nm	9.5 nm				
Beam emittance, vertical	1.2 nm	1.5 nm				
β - function at IP, horizontal	43 cm	90 cm				
β - function at IP, vertical	5 cm	4 cm				
Beam-beam tuneshift, horizontal	0.073	0.014				
Beam-beam tuneshift, vertical	0.1	0.007				
Luminosity at E _{cm} = 105 Gev	1·10 ³⁴ cm ⁻² s ⁻¹					

10 **Beam-beam limited** Luminosity [1033 cm⁻² s⁻¹] beam size lumi scaling 10 MW SR limited Space charge limited 10 e [GeV] 5 5 10 18 275 p [GeV] 275 41 100 100 0.1 20 40 60 80 100 120 140 160 Center of Mass Energy [GeV]

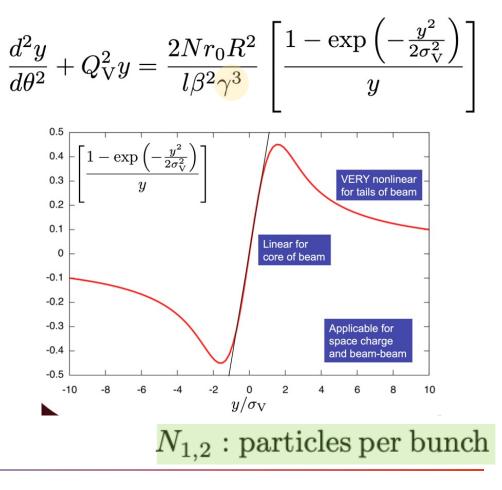
- Result of collective parameter optimization
- CM energy and luminosity limiting factors are correlated
 - Three fundamental luminosity limiting factors
- Later: acceptance and luminosity inversely correlated

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Luminosity at Low E_{cm}: Hadron Space Charge

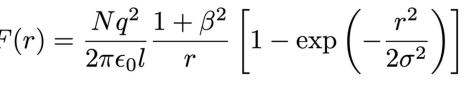
- Dense charged particle bunches electrostatically repel in rest frame
- Creates **nonlinear** space charge force and equation of motion in lab frame
- Fortunately scales with 1/γ³ so worst at low energies
 - Great example of time dilation
 - Limits high-intensity injector emittances
 - · Force applies continuously within beam
- Tolerable linear "space charge tune spread" of 0.05 limits total current of 41 GeV proton beam to ~0.4A.
- (IBS: intra-beam hard scattering also contributes)

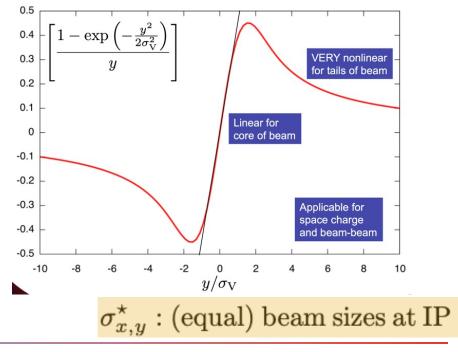


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Luminosity at Central E_{cm}: Colliding Beam-Beam Forces

- Colliding beams see each other's collectiv charge distributions $F(r) = \frac{Nq^2}{2\pi\epsilon_0 l} \frac{1+\beta^2}{r} \left[1-\exp\left(-\frac{r^2}{2\sigma^2}\right)\right]$
- Creates **nonlinear** beam-beam force and equation of motion similar to space charge
 - **BUT** now the fields and force are in the lab frame already
 - NO benefit of relativistic scaling
 - Force applies only once per turn
- Tolerable "beam-beam tune spead" of 0.015 limits highest EIC luminosity





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Luminosity at High E_{cm}: Electron Synchrotron Radiation Power

- Accelerated charged particles emit photons
 - Electrons in synchrotron: radially accelerated
 - Synchrotron radiation emitted in forward cone
 - Cone opening angle $\propto 1/\gamma$

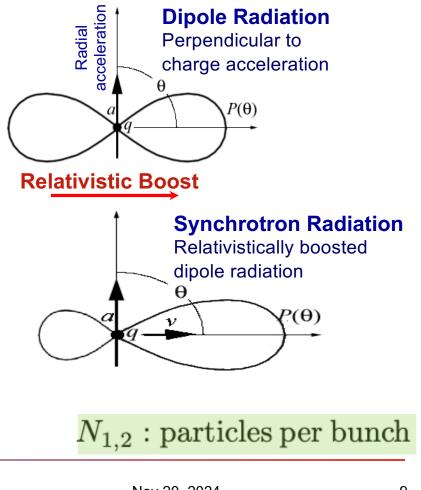
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Radiated power

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$$P_{\gamma} = \frac{2}{3} \frac{e^2 c}{4\pi\epsilon_0} \frac{(\gamma\beta)^4}{\rho^2}$$

- γ scaling **much** worse for electrons
 - 18 GeV e: γ = 3.5x10⁴ vs 255 GeV p: γ = 3x10²
- Design: 9 MW @ 18 GeV (facility limit 10 MW)
- Expensive: Power must be provided by SRF
- Raise electron energy last (e- current limit)



Accelerator at Day One and Ultimate

EIC for Early Science

RCS:

7nC / bunch 5 – 10 GeV polarized e-



change injection scheme to reach 28 nC / bunch and 18 GeV electrons

add more RF cavities to reach 28 nC / bunch

Ultimate EIC

ESR:

7nC / bunch 5 – 10 GeV polarized e-



and 18 GeV electrons

HSR:

100 – 250 GeV polarized p 100 GeV/u nuclear beams pre-cooling at injection

upda add

update PS to reach 275 GeV protons and 110 GeV/u nuclei add 41 GeV bypass to get full HSR beam energy capabilities

Proposal for the Day-One Physics and the initial years of science is driven by

- Start of the promised NSAC/NAS science program
- Alignment with expected order in commissioning the collider and ramp up of performance that comes with gain of operational experience
- Having access to new physics results early to get high impact publications, i.e. PRLs,

Proposed Early Science Program incorporates comments from EIC-UG/ePIC Summer Meeting

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Proposal for EIC Science Program in the First Years

Year - 1

Start with Phase 1 EIC New Capability:

Commission electron polarization in parallel Run:

10 GeV electrons on 115 GeV/u heavy ion beams (Ru or Cu) Physics:

gives world-wide new data on nPDFs, NFF and a first look on saturation

0 111

Year - 2

Start with Phase 1 EIC Commission electron polarization in parallel

New Capability:

Commission hadron polarization in parallel Run:

10 GeV electrons on 130 GeV/u Deuterium Physics:

gives world-wide new data → critical baseline for nPDFs and

saturation free vs. bound proton structure

Run:

Last weeks 10 GeV electrons and 130 GeV polarized protons Physics:

first look to 3d imaging of the proton

Year - 3

Start with Phase 1 EIC Commission electron polarization in parallel Commission hadron polarization in parallel

New Capability:

Commission running with hadron spin rotators

10 GeV electrons on 130 GeV transverse polarized protons Physics:

3d imaging of the proton / mass of the nucleon

Run:

Last weeks switch to longitudinal proton polarization

Physics:

first look helicity structure of the proton – unravel quark, gluon and orbital angular contributions

in parallel Commission hadron polarization in parallel Commission running with hadron spin rotators New Capability: Commission hadron accelerator to operate with not centered orbits Run: 10 GeV electrons on 100 GeV Au Physics: gives world-wide new data on Adep.of nPDFs, nFF and saturation

Commission electron polarization

Year - 4

Start with Phase 1 EIC

Run:

10 GeV electrons on 250 GeV transverse and longitudinal polarized protons

Physics:

- 3d imaging of the proton at low x
- helicity structure of the proton unravel quark, gluon and orbital angular contributions

Year - 5

Start with Phase 1 EIC Commission electron polarization in parallel Commission hadron polarization in parallel Commission running with hadron spin rotators Commission hadron accelerator to operate with not centered orbits Run: 10 GeV electrons on 100 GeV Au **Physics:** gives world-wide new data on A-dep.of nPDFs, nFF and saturation Run: 10 GeV electrons on 166 GeV transverse and longitudinal polarized He-3 **Physics:** • 3d imaging of the nucleons \rightarrow flavor

- separation
 helicity structure of the nucleon– unravel helicities for different quark
- flavors
- first look to nuclear binding

Time to install additional ESR RF and HSR PS to reach design Current and max Energies

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Proposal for EIC Science Program in the First Years

Year - 5

Start with Phase 1 EIC Commission electron polarization in parallel Commission hadron polarization in parallel Commission running with hadron spin rotators Commission hadron accelerator to operate with not centered orbits **Run:**

10 GeV electrons on 100 GeV Au Physics:

gives world-wide new data on A-dep.of nPDFs, nFF and saturation Run:

10 GeV electrons on 166 GeV transverse and longitudinal polarized He-3 Physics:

- 3d imaging of the nucleons → flavor separation
- helicity structure of the nucleon– unravel helicities for different quark flavors

first look to nuclear binding

Year - 6

Start with Phase 1 EIC

Commission electron polarization in parallel Commission hadron polarization in parallel Commission running with hadron spin rotators

Commission hadron accelerator to operate with not centered orbits

New Capability:

Commission ESR & HSR at max. energy and beam currents

Run:

18 GeV polarized electrons on 275 GeV/u polarized (longitudinal & transverse) proton beams

Physics:

- 3d imaging of the proton at low x
- helicity structure of the proton unravel quark, gluon and orbital angular contributions

Year - 7

Start with Phase 1 EIC Commission electron polarization in parallel

Commission hadron polarization in parallel Commission running with hadron spin rotators

Commission hadron accelerator to operate with not centered orbits

Commission ESR & HSR at max. energy and beam currents

New Capability:

Operate HSR with 41 GeV bypass Run:

5 GeV polarized electrons on 41 GeV transverse polarized proton beams **Physics:**

first look to moderate-x 3d imaging and Kaon structure/ mass

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Assumptions and Luminosity Calculation

- 7 nC electron bunch charge compared to 28 nC (CDR)
- Constant proton beam IP divergencies are maintained throughout the store by gradual increase of proton IP beta-functions as the beam emittance increases.
- The electron IP beta-functions are adjusted accordingly to match electron and proton transverse beam size.
- Ion beam is cooled at low energy (24 GeV/u) but no stochastic cooling is used in the store
- 1 Run is ½ year operation at 80% uptime
- 2 h store turnaround time
- 30 min at the beginning of the store is taken by the ESR fill and detector turn-on
- Not yet included a ramp of of luminosity through the Run

 at RHIC: 1st week 25% of projected lumi / week
 2nd week 50% of projected lumi / week
 3rd week 75% of projected lumi / week
 4th week to X week 100% of projected lumi / week
 First guess for EIC early years (1 to ~5?)
 1st week 10% of projected lumi of the run / week
 increase by 10% every → 10 weeks to reach projected lumi / week

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Luminosity eA in Early Years

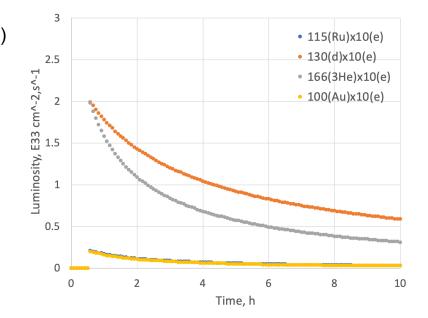
Possible Beam energies:

electron: 5 GeV, 10 GeV and ultimately 18 GeV proton: 41 GeV, 100 GeV to 255 GeV ultimately 275 GeV Au: 41 GeV, 100 GeV to 110 GeV ultimately A: 41 GeV, 100 GeV to Max ~255 / (A/Z) ultimately ~275 / (A/Z)

	Lumi per Fill (5 h)	Lumi per Year
10 GeV e x 115 GeV Ru	1.3 pb ⁻¹	0.9 fb ⁻¹
10 GeV e x 100 GeV Au	1.2 pb ⁻¹	0.84 fb ⁻¹
10 GeV e x 130 GeV d	16 pb ⁻¹	11.4 fb ⁻¹
10 GeV e x 166 GeV ³ He	12 pb ⁻¹	8.65 fb ⁻¹

Note:

eA luminosity is per nucleon



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Luminosity ep in Early Years

Possible Beam energies:

electron: 5 GeV, 10 GeV and ultimately 18 GeV proton: 41 GeV, 100 GeV to 255 GeV ultimately 275 GeV Au: 41 GeV, 100 GeV to 110 GeV ultimately A: 41 GeV, 100 GeV to Max ~255 / (A/Z) ultimately ~275 / (A/Z)

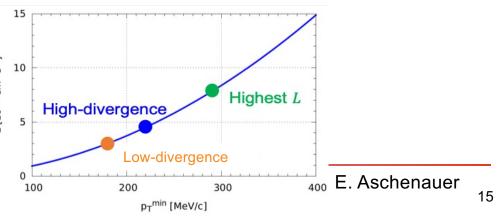
High Divergence	Lumi per Fill (5 h)	Lumi per Year	Low Divergence	Lumi per Fill (5 h)	Lumi per Year
5 GeV e x 250 GeV p	9.26 pb ⁻¹	6.48 fb ⁻¹	5 GeV e x 250 GeV p	6.81 pb ⁻¹	4.78 fb ⁻¹
10 GeV e x 250 GeV p	13.12 pb ⁻¹	9.18 fb ⁻¹	10 GeV e x 250 GeV p	8.8 pb ⁻¹	6.19 fb ⁻¹
5 GeV e x 130 GeV p	6.3 pb ⁻¹	4.36 fb-1	5 GeV e x 130 GeV p	5.8 pb ⁻¹	4.1 fb ⁻¹
10 GeV e x 130 GeV p	7.6 pb ⁻¹	5.33 fb ⁻¹	10 GeV e x 130 GeV p	7.1 pb ⁻¹	4.95 fb ⁻¹

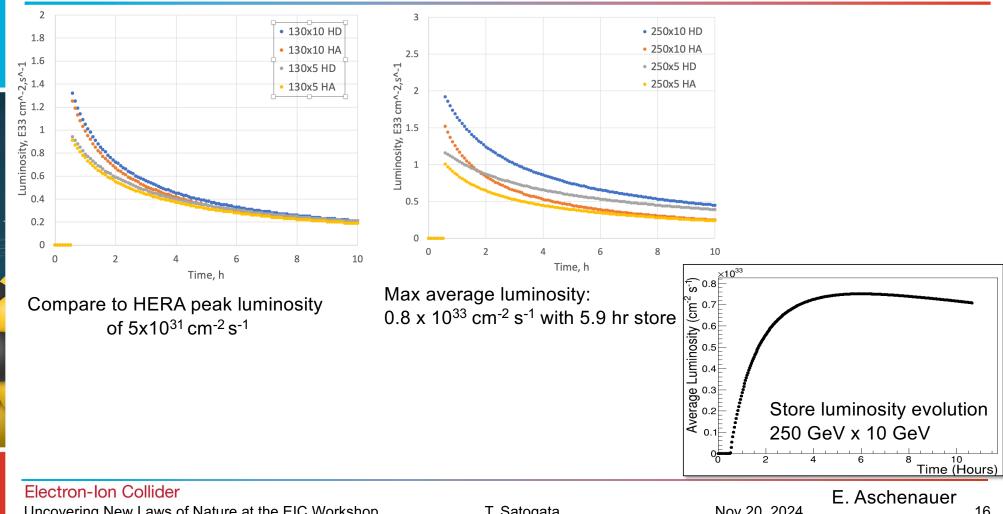
Remember:

high divergence: higher lumi, but reduced acceptance for low forward particle p_T^{min} low divergence: lower lumi, but increased acceptance for low forward particle p_T^{min}

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Luminosity ep in Early Years

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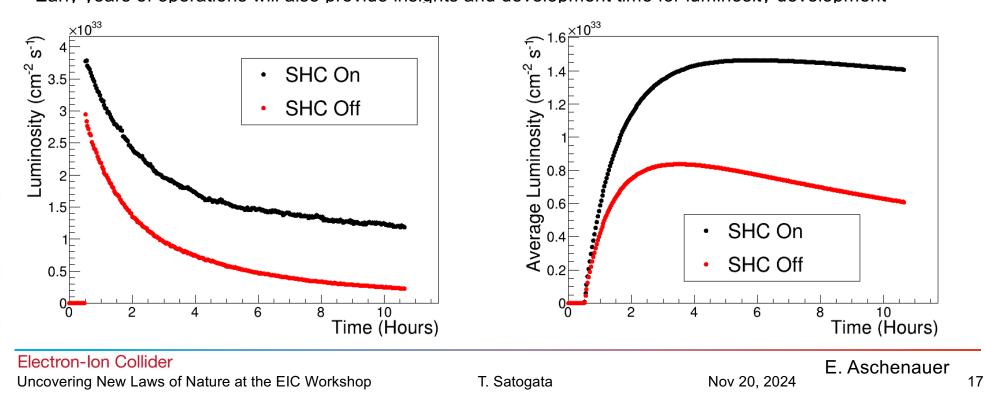
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Onwards to Full Luminosity Maturity

After Year 5 will have full bunch current for electrons 7 nA to 28 nA → increase in luminosity ~4

Adding Strong Hadron Cooling will increase the integrated current by fill another factor of 2 Example: 10 GeV electrons x 100 GeV protons

Early years of operations will also provide insights and development time for luminosity development



Luminosity Lessons

Tevatron Run-II: design 275x10³⁰

C = 8.6

The gap was due to

incomplete understanding of several beam physics

issues and a number technical difficulties

51/12=4.12 over 4 yrs

C=4/ln(4.12)=2.85

2000 2001 2002 2003 2004 2005

2002 2003 2004 2005 2006 2007 2008 2009 2010

HERA: design 15x10³⁰

8661 999

16/8=2 over 2 yrs

C=2/ln(2)=2.9

966

L601

- Luminosity ramp-up to design takes years
 - Useful paper: arxiv 1202.3950 (V. Shiltsev)
 - Contextualizes Tevatron Run-II and early LHC
 - Luminosity ramp-up parameter C: complexity
 - C: time (years) to increase luminosity by e
 - C=2: factor of e luminosity increase in 2 years
 - Early commissioning can make quick strides
 - C<1 (or <<1) but do not get too exuberant
 - Long-term commissioning is usually C~2-3

• EIC is an exceptionally complex collider

- EIC will most likely take years to reach design luminosities
 - We will get there!

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1995

1992 1993

Tevatron Peak Luminosity, 10^{30} cm^{1.5} rough 10^{30} rough 10

2001

60

50

40

30

20

Peak Luminosity, e30 cm⁻²s⁻¹

C=2.



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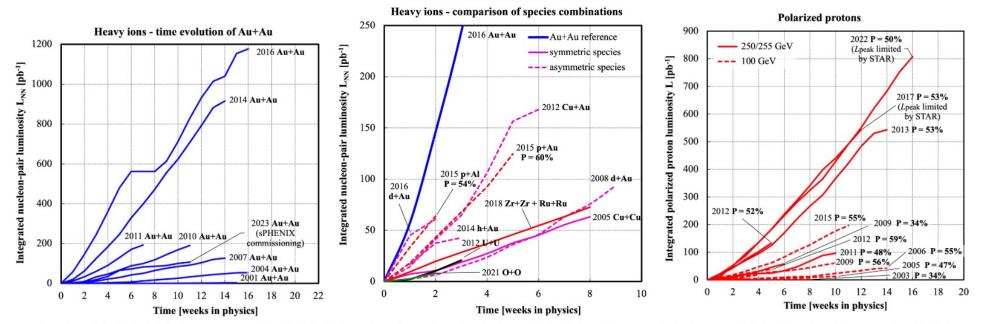
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Species	proton	electron									
Energy [GeV]	275	18	275	10	100	10	100	5	41	5	
CM energy [GeV]	14	0.7	10	4.9	63	3.2	44	l.7	28	3.6	27
Bunch intensity [10 ¹⁰]	19.1	6.2	6.9	17.2	6.9	17.2	4.8	17.2	2.6	13.3	
No. of bunches	2	90	11	.60	11	.60	11	60	11	.60	
Beam current [A]	0.69	0.227	1	2.5	1	2.5	0.69	2.5	0.38	1.93	
RMS norm. emit., h/v [μm]	5.2/0.47	845/71	3.3/0.3	391/26	3.2/0.29	391/26	2.7/0.25	196/18	1.9/0.45	196/34	EIC CDR
RMS emittance, h/v [nm]	18/1.6	24/2.0	11.3/1.0	20/1.3	30/2.7	20/1.3	26/2.3	20/1.8	44/10	20/3.5	CD-1 with
β*, h/v [cm]]	80/7.1	59/5.7	80/7.2	45/5.6	63/5.7	96/12	61/5.5	78/7.1	90/7.1	196/21.0	SHC
IP RMS beam size, h/v [µm]	119	9/11	95,	/8.5	138	6/12	125	/11	198	8/27	1
K_x	1	1.1	11	1.1	11	1.1	11	.1	7	.3	
RMS $\Delta \theta$, h/v [µrad]	150/150	202/187	119/119	211/152	220/220	145/105	206/206	160/160	220/380	101/129	
BB parameter, $h/v [10^{-3}]$	3/3	93/100	12/12	72/100	12/12	72/100	14/14	100/100	15/9	53/42	
RMS long. emittance $[10^{-3}, eV \cdot s]$	36		36		21		21		11		•
RMS bunch length [cm]	6	0.9	6	0.7	7	0.7	7	0.7	7.5	0.7	
RMS $\Delta p / p [10^{-4}]$	6.8	10.9	6.8	5.8	9.7	5.8	9.7	6.8	10.3	6.8	
Max. space charge	0.007	neglig.	0.004	neglig.	0.026	neglig.	0.021	neglig.	0.05	neglig.	
Piwinski angle [rad]	6.3	2.1	7.9	2.4	6.3	1.8	7.0	2.0	4.2	1.1	
Long. IBS time [h]	2.0		2.9		2.5		3.1		3.8		
Transv. IBS time [h]	2.0		2		2.0/4.0		2.0/4.0		3.4/2.1		_
Hourglass factor H	0.	91	0.	94	0.	90	0.	88	0.	93	
Luminosity $[10^{33} \text{cm}^{-2} \text{s}^{-1}]$	1.	54	10	.00	4.	48	3.	68	0.	44	21
					-						21

Table 3.3: EIC beam parameters for different center-of-mass energies \sqrt{s} , with strong hadron cooling. High divergence configuration.

RHIC Runs



Center-of-mass energy $\sqrt{s_{NN}}$ [GeV] (scale not linear)

Note: The nucleon-pair luminosity is defined as $L_{NN} = A_1 A_2 L$, where L is the luminosity, and A_1 and A_2 are the number of nucleons of the ions in the two beams respectively. The proton polarization is intensity and time averaged over the whole run, as measured by the H-jet.

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Tabulated Complexities from Shiltsev Paper

	С	years
CESR $e+e-$	4.3	1883-1988
LEP I $e+e-$	3.3	1989-1995
SLC <i>e+e-</i>	1.5	1989-1997
HERA I, II <i>p-e</i>	2.9	1992-00-2005
ISR <i>p-p</i>	3.0	1972-1982
SppS <i>p-pbar</i>	2.0	1982-1990
Tevatron Run II <i>p-pbar</i>	2.0	2002-2007
RHIC <i>p-p</i>	2.2	2000-2004
Tevatron startup	0.03	1987
LHC startup	0.06	2010

Table II: "Complexities" of colliding beam facilities.

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Shiltsev Paper Conclusions

One should not expect that the period of incredibly fast growth of the LHC luminosity - as in 2010- will last long. At some point the progress will most probably turn to the rate corresponding to complexity of C~2. Such a period of exploration and fight for ultimate performance with $C\approx 2$ might take as short as 3-4 years and as long as 6-10 years It will be followed by relative stabilization of performance (either running out of ideas or preparing for a major upgrade). A numerical example: progress from $L=3\times10^{33}$ cm⁻²s⁻¹ to $L=5\times10^{34}$ cm⁻²s⁻¹ might take 6-9 years if C=2-3.

Expectations management is crucial. As in the case of the Tevatron, the LHC goals may need to be expressed in terms of two goals: "base" goal – that is believed has very high degree of certainty of being achieved and the "design" or "stretched" goal that represents your "best estimate" of the limit of performance to which the facility can be pushed. The goals and the ratio of "base" to "design" goals will depend on the level of understanding of the machine, e.g. the ratio might change from larger to smaller to reflect lower level of uncertainty in later years.

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