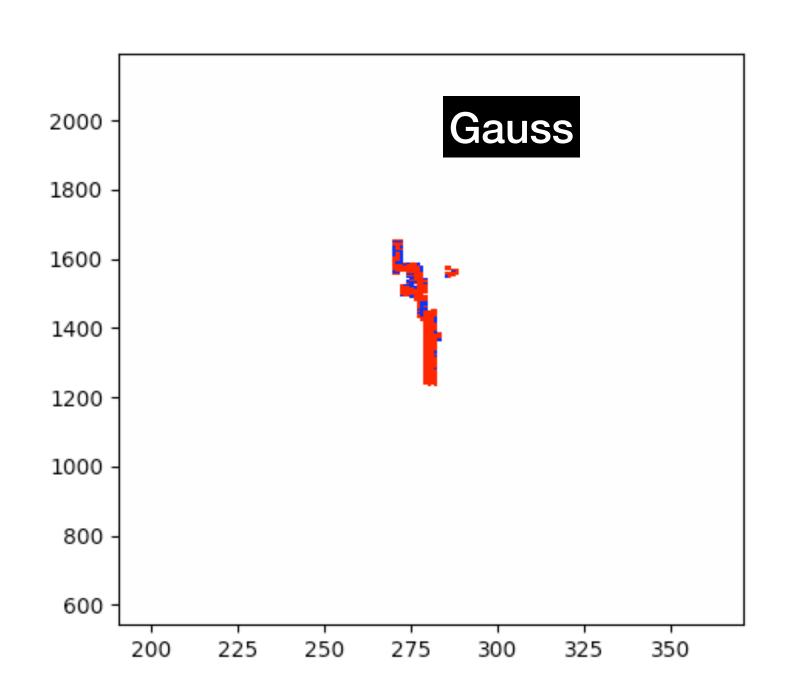
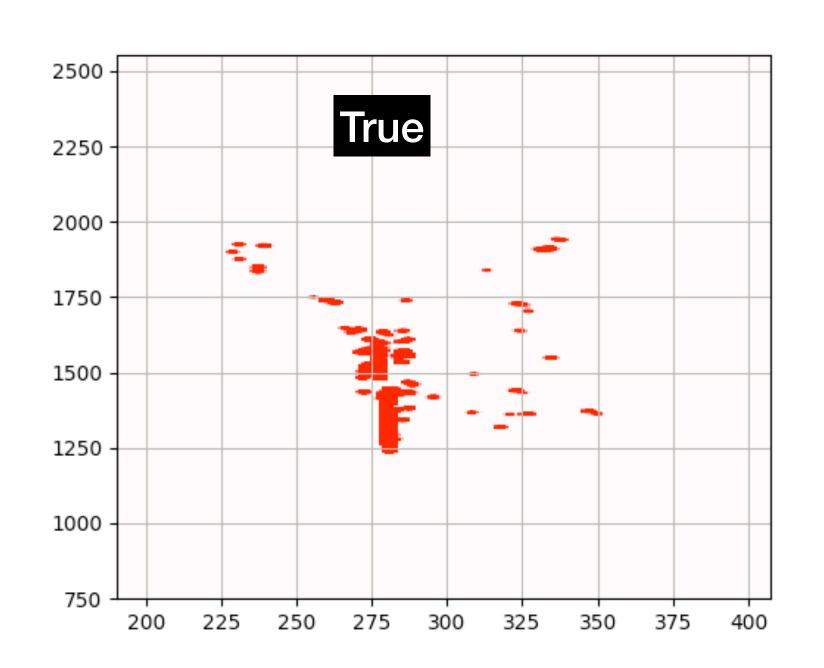
# **DNNROI** for PDHD [showers]

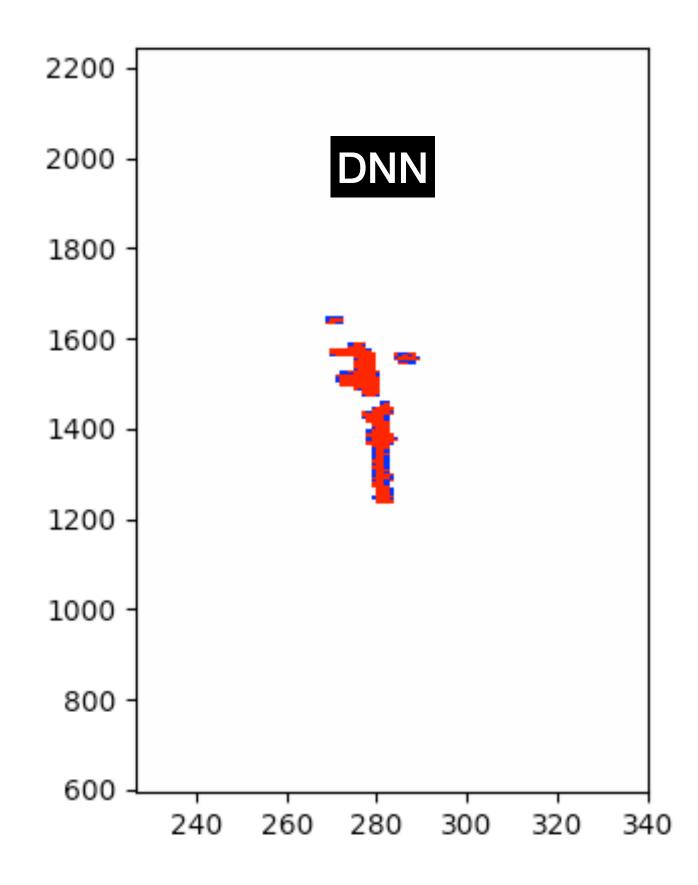
#### Evaluate showers

- Each event has single e- simulated in the ~middle of the detector
- Energy bins 100MeV, 500MeV, 1GeV, 1.5GeV
- Angles XY: 5deg, Angles XZ = 0,15,45,60,90deg [drift along X]
- Evaluate on total activity in the event per plane :
  - Simulate 20 event per Angle/Energy bin
  - Calculate quantile bias and resolution based on total energy seen by each plane ( $(E_{true}^{tot} E_{reco}^{tot})/E_{true}^{tot}$ )

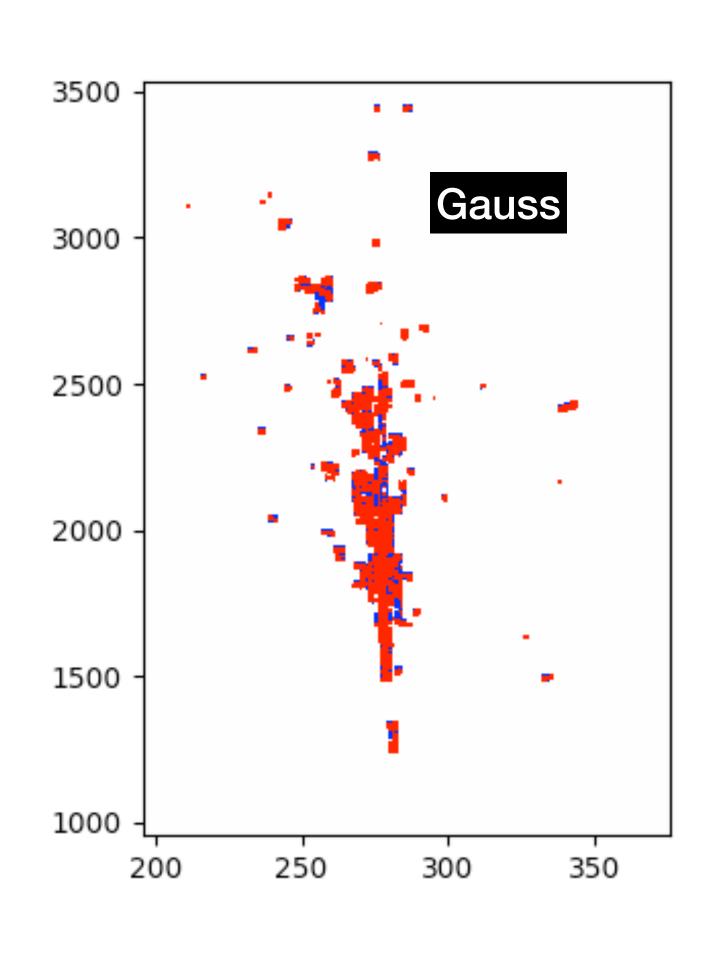
### Example of 100MeV 90 deg XZ electron U-plane

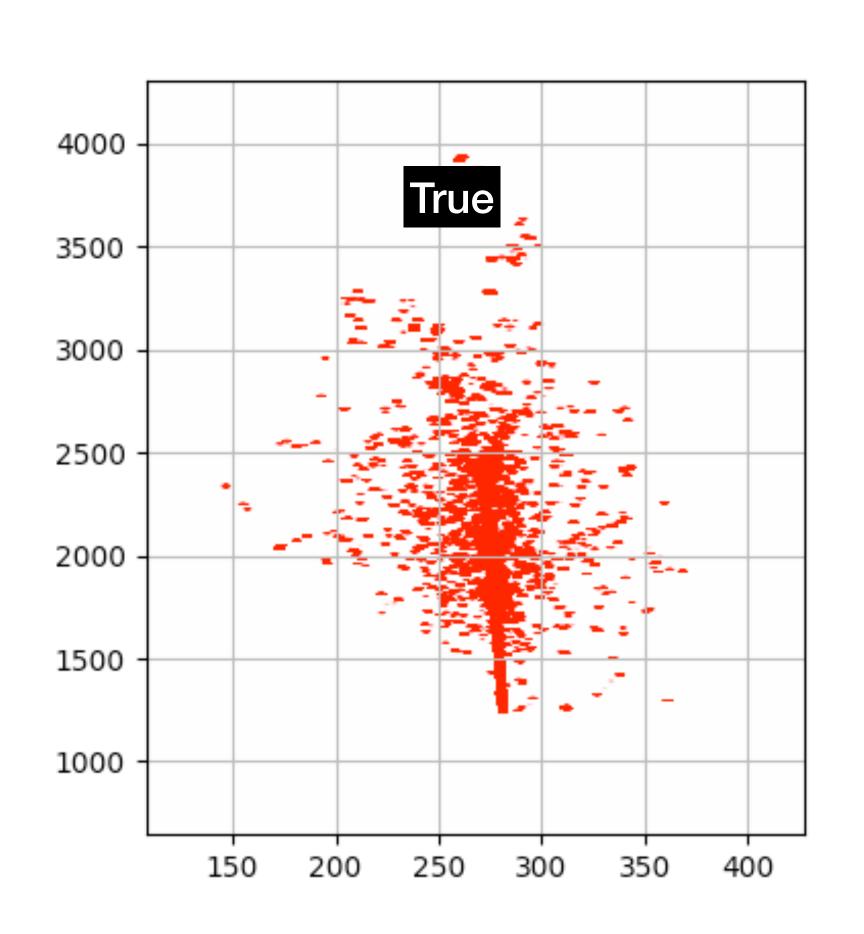


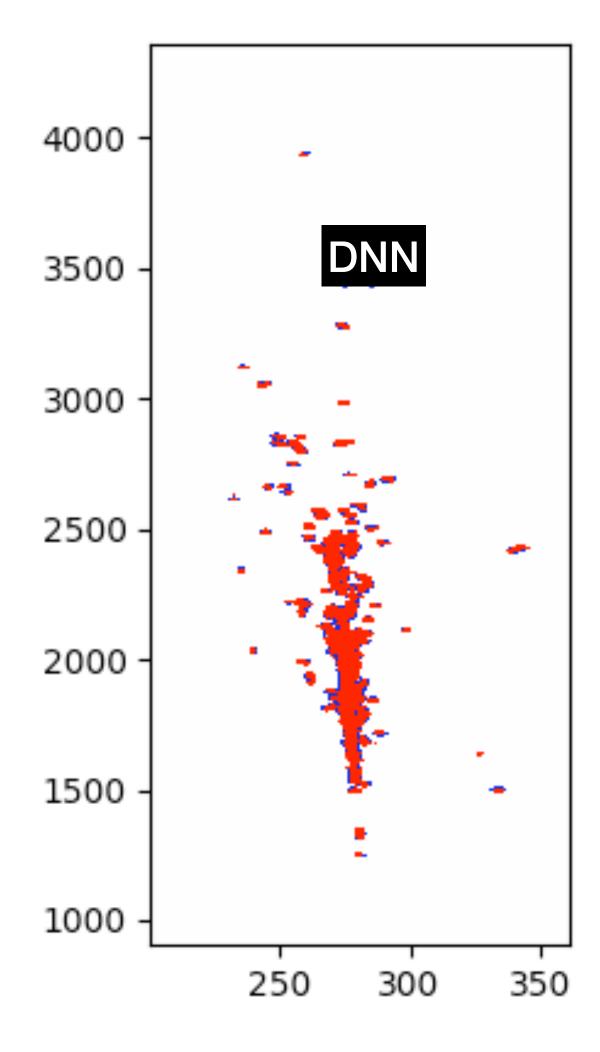


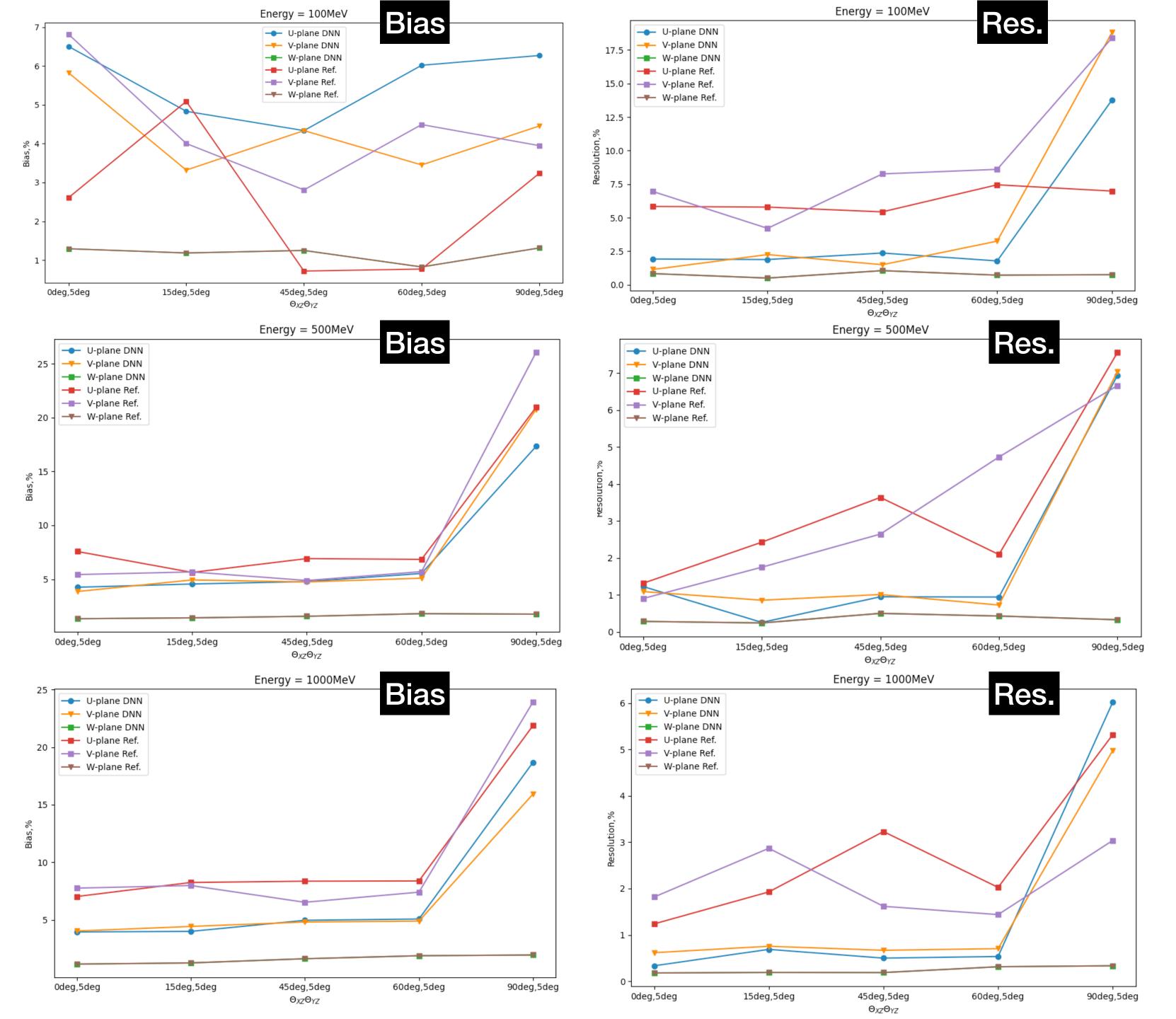


### Example of 1.5GeV 90 deg XZ electron U-plane

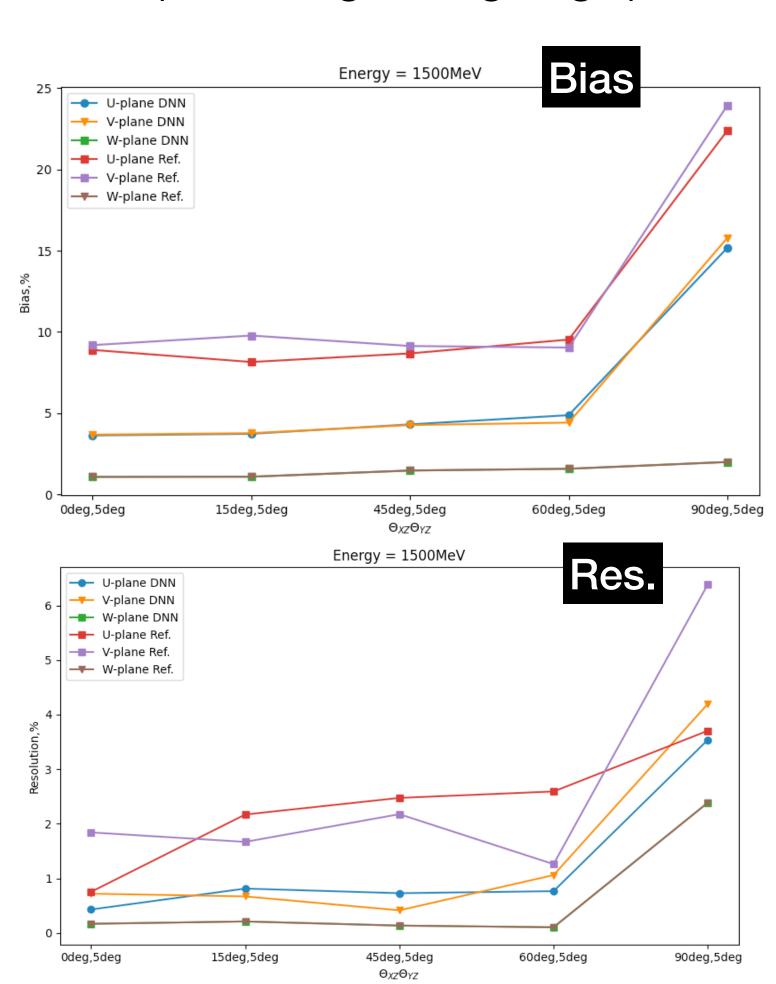






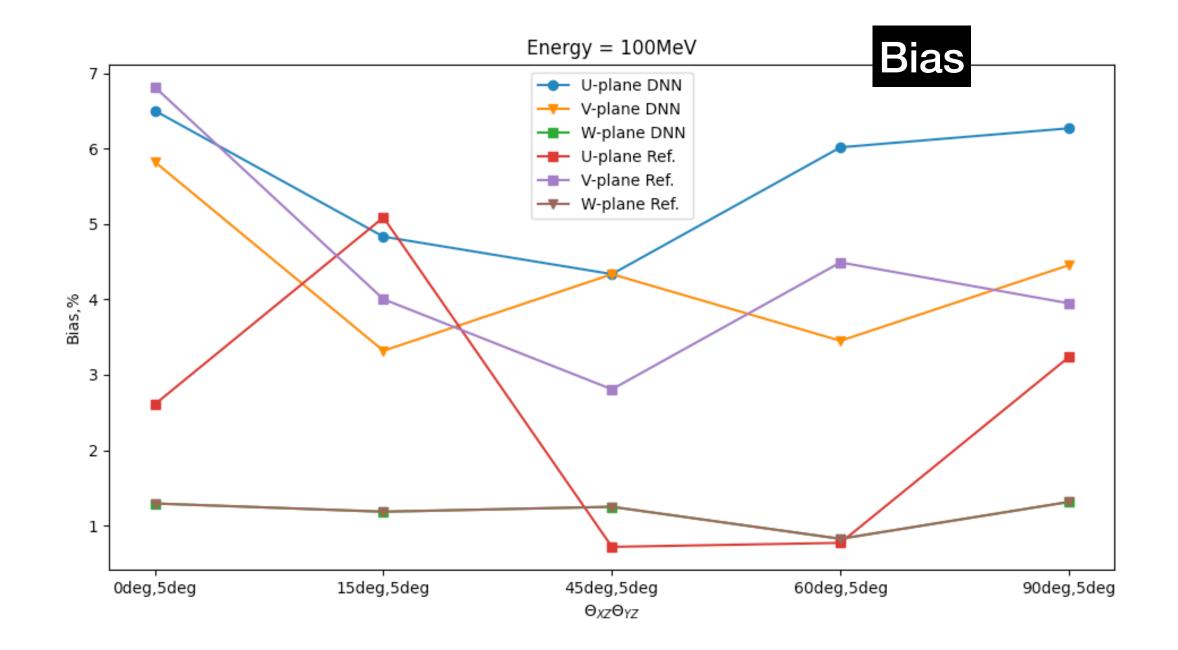


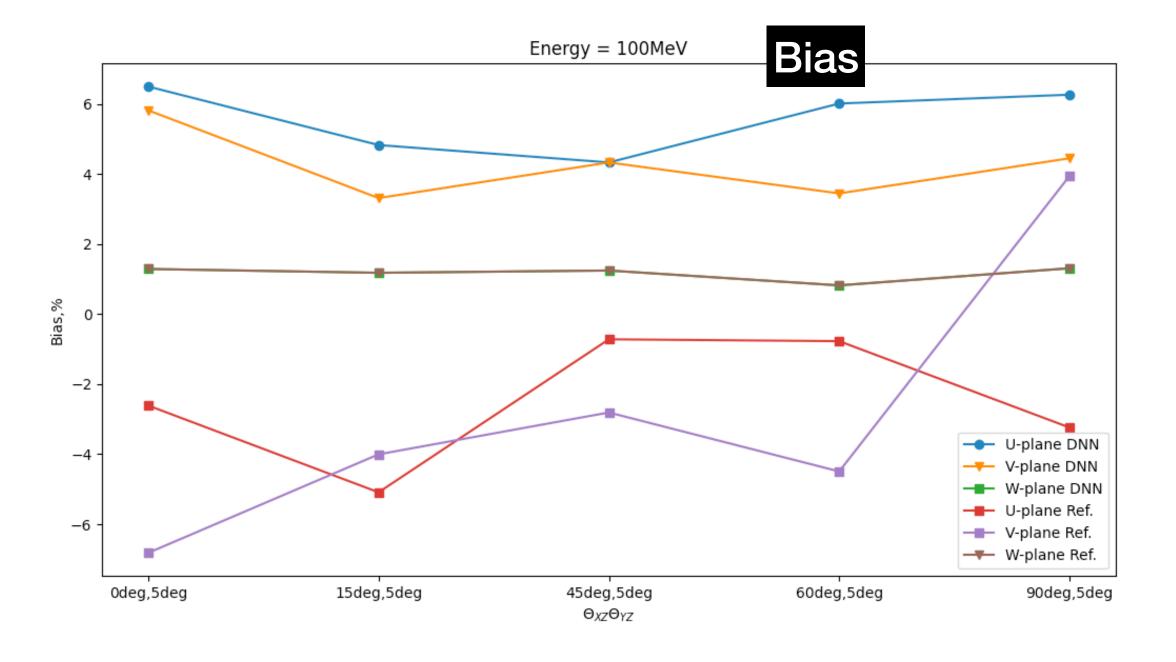
- Performance for DNN and Ref. are visually very close for all angles/ energies
- Bias is comparable between DNN and Ref. for lower energies with DNN outperforming Ref. at higher energies
- Resolution is generally smaller for DNN case (excluding 90deg angle)



#### Check sign of the bias

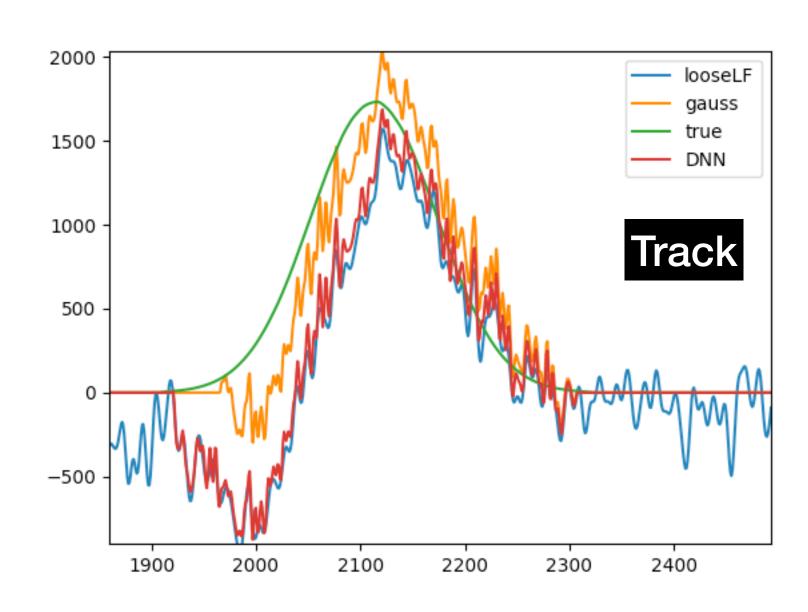
- Previously bias was plotted as absolute value
- Switching to signed value only 100MeV bin changes > Reference bias becomes negative
- Lets check WF of one 15deg event

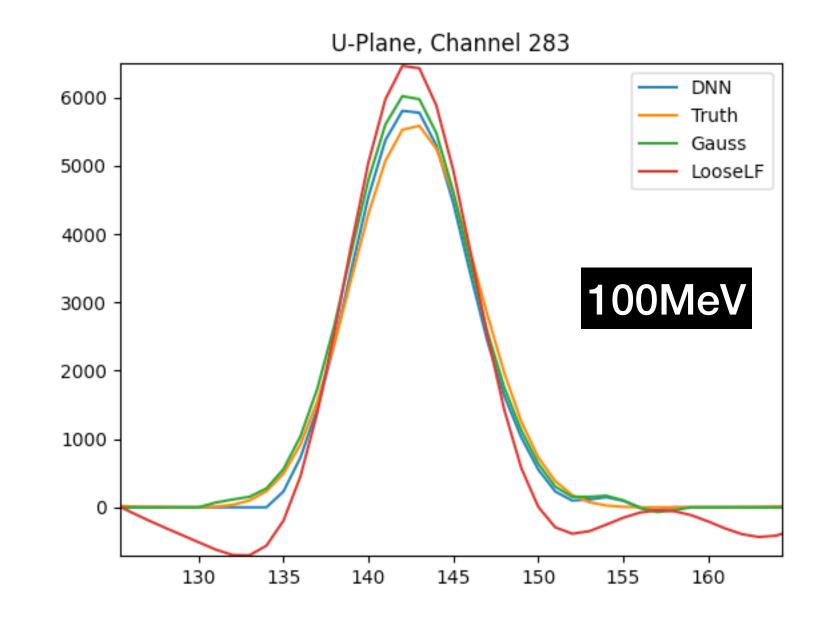


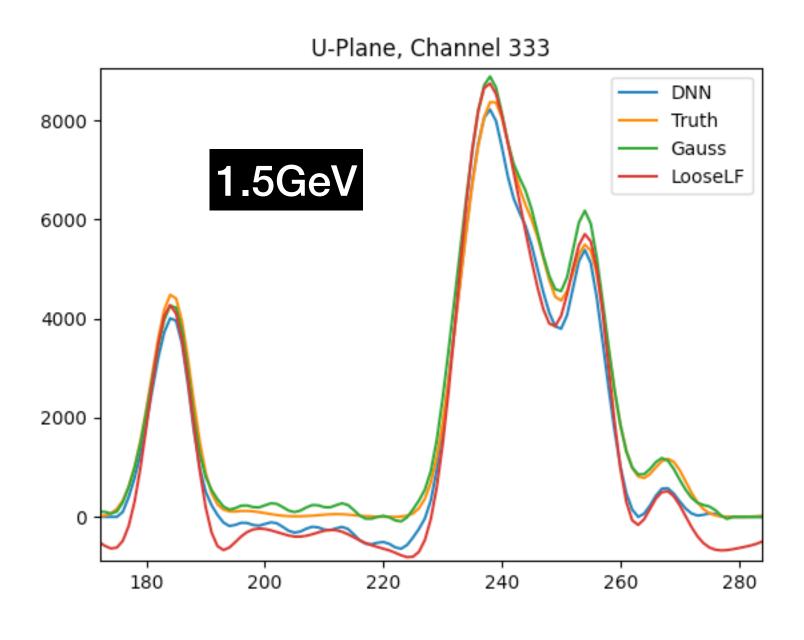


#### Look at WF on specific channel

- Lets check WF of one 15deg event
- Check 100MeV and 1.5GeV cases where channel has Reference reco. charge > True, but DNN reco. charge <= True</li>
  - DNN generally follows LooseLF WF much close, especially outside of the peak
- Similar check for tracks is also performed







## Summary

- Performance for DNN and Ref. are visually very close for all angles/ energies
- Bias is comparable between DNN and Ref. for lower energies with DNN outperforming Ref. at higher energies
- Resolution is generally smaller for DNN case (excluding 90deg angle)
- DNN does not give much improvement for a single shower

