



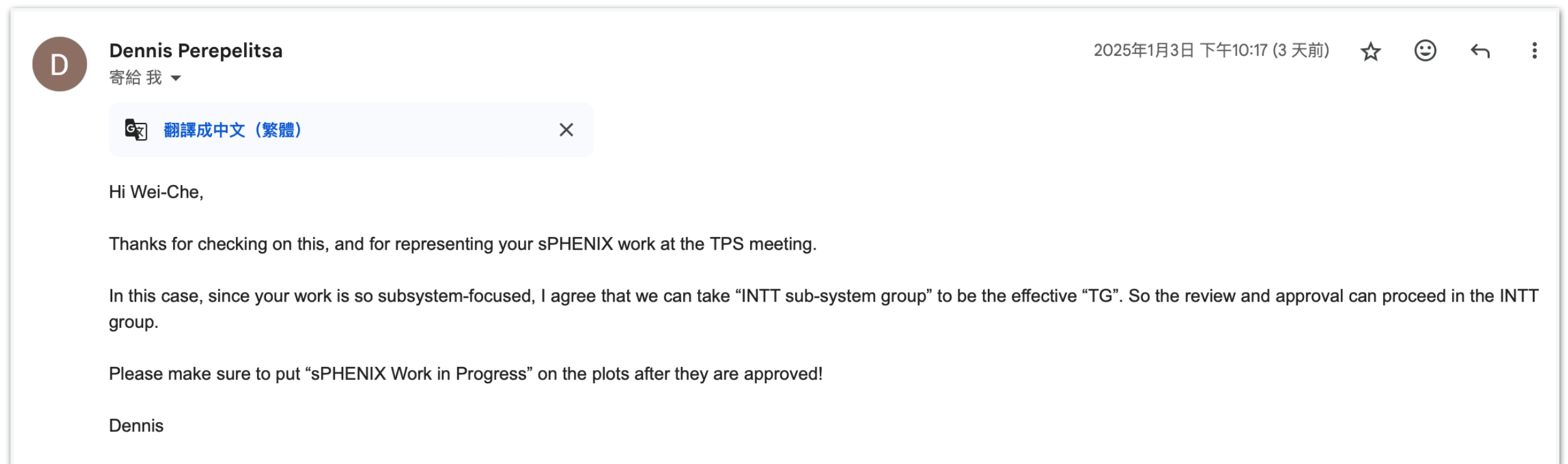
# INTT cosmic analysis update

Wei-Che Tang, NCU

2025/01/10 INTT meeting

# Aim for this presentation

- The TPS Annual Meeting (Annual Meeting of the Physical Society of Taiwan) is approaching on Jan. 14.
- I would like to present the current status of my comics analysis on the TPS meeting.
- I confirmed with Dennis that I could get the approval for “work in progress” plots within the INTT group.

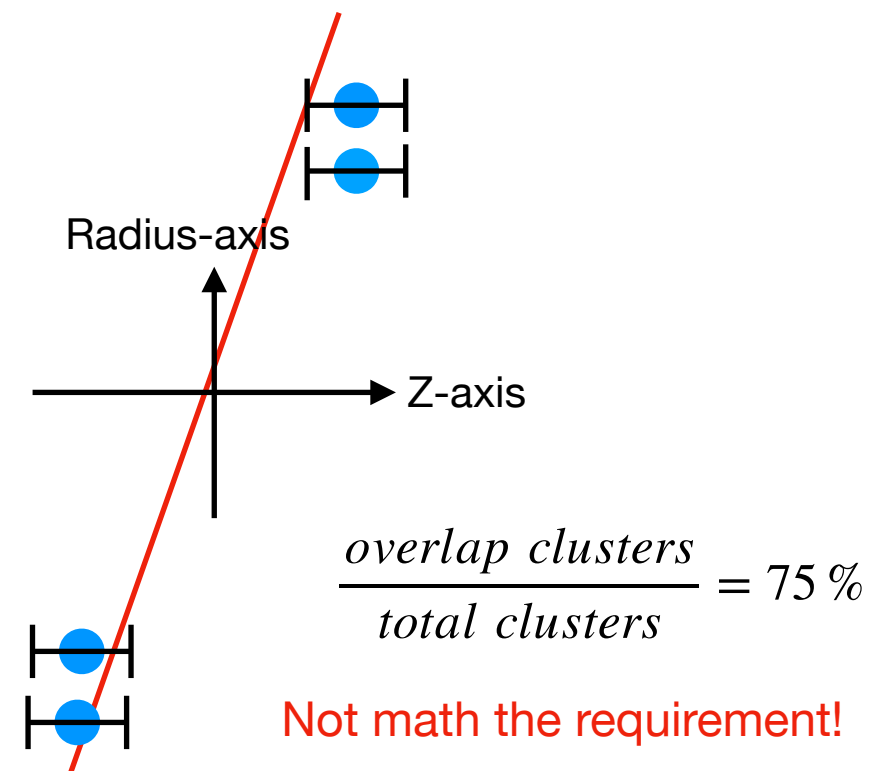
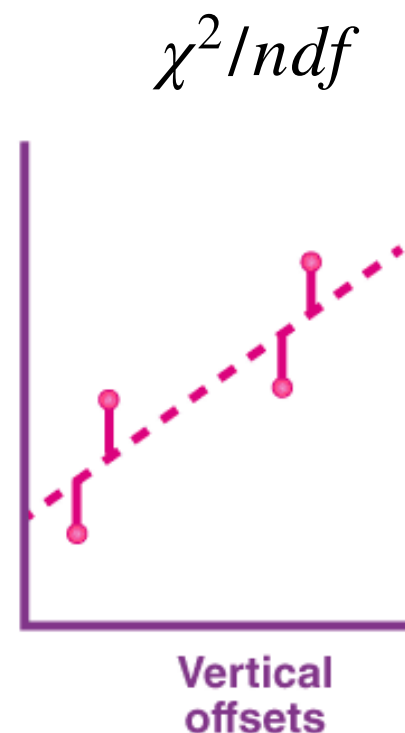


# Cosmic Analysis status

- Run number: 39365~39530, ~30 runs if not otherwise specified
  - Trigger: HCal Vertical trigger
  - Data taking with 0 magnetic field.
  - Total number of events: 4.2M
  - Threshold: 15
  - Data taking period: Apr. 12 ~ May. 1 (Run2024 cosmic data)
- INTTRawHit DSTs are from official production.
- INTTTrkrCluster DSTs are from private production using Fun4All framework.
  - To apply the private hot channel masking map.
- Good clusters required
  - '*cluster Adc > 15*' (Adc0 is set to be 15) & '*cluster phisize < 9*'
- Two approaches are used in this analysis.
  - 4 good clusters only.
    - Minimal requirement to reconstruct a pure cosmic track.
  - 4~7 good clusters included.
    - Find rare comic tracks with more ladders involved.

# 4 good clusters only

- Event selection: # of good clusters = 4
- Fit the clusters with a slope in X-Y and Z-Radius plane.
- Reduced residual is calculated in X-Y plane to evaluate the fitting quality.
  - To minimize the track angle dependence.
- Definition of a good track:
  - $X$ - $Y$  residual  $< 0.02$  cm
  - Fitting line requires to overlap with each cluster (100% overlap) in Z-Radius plane.

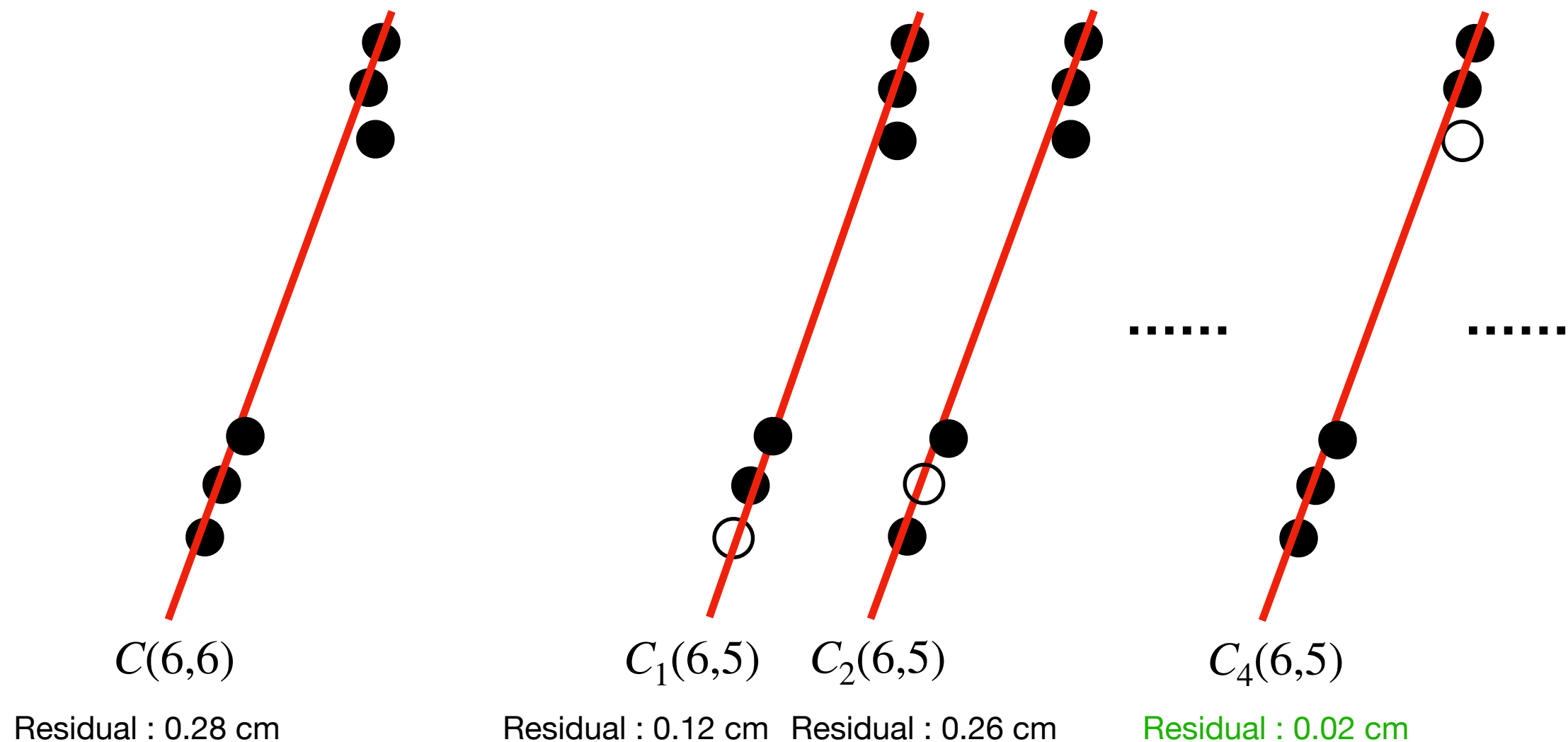


# 4~7 good clusters

- Event selection:  $4 \leq \# \text{ of good clusters} \leq 7$
- Prepare all the combinations based on  $C(n, k)$ 
  - $n$  is total number of clusters in an event
  - $k$  is the number of clusters involved in the fitting
  - i.e. if  $n = 5 \rightarrow \{4 \times k=4\} + \{1 \times k=5\}$
- Fit each combination with a slope in X-Y and Z-Radius plane.
- Reduced residual is calculated in X-Y plane to evaluate the fitting quality
  - To minimize the track angle dependence.
- Fitting line requires to overlap with each cluster in Z-Radius plane.
- In each  $k$ , pick up the one with the least reduced residual.
- Definition of a good track:
  - The combination with larger  $k$  whose reduced residual smaller than 0.02 cm is selected.

# Clusters combination

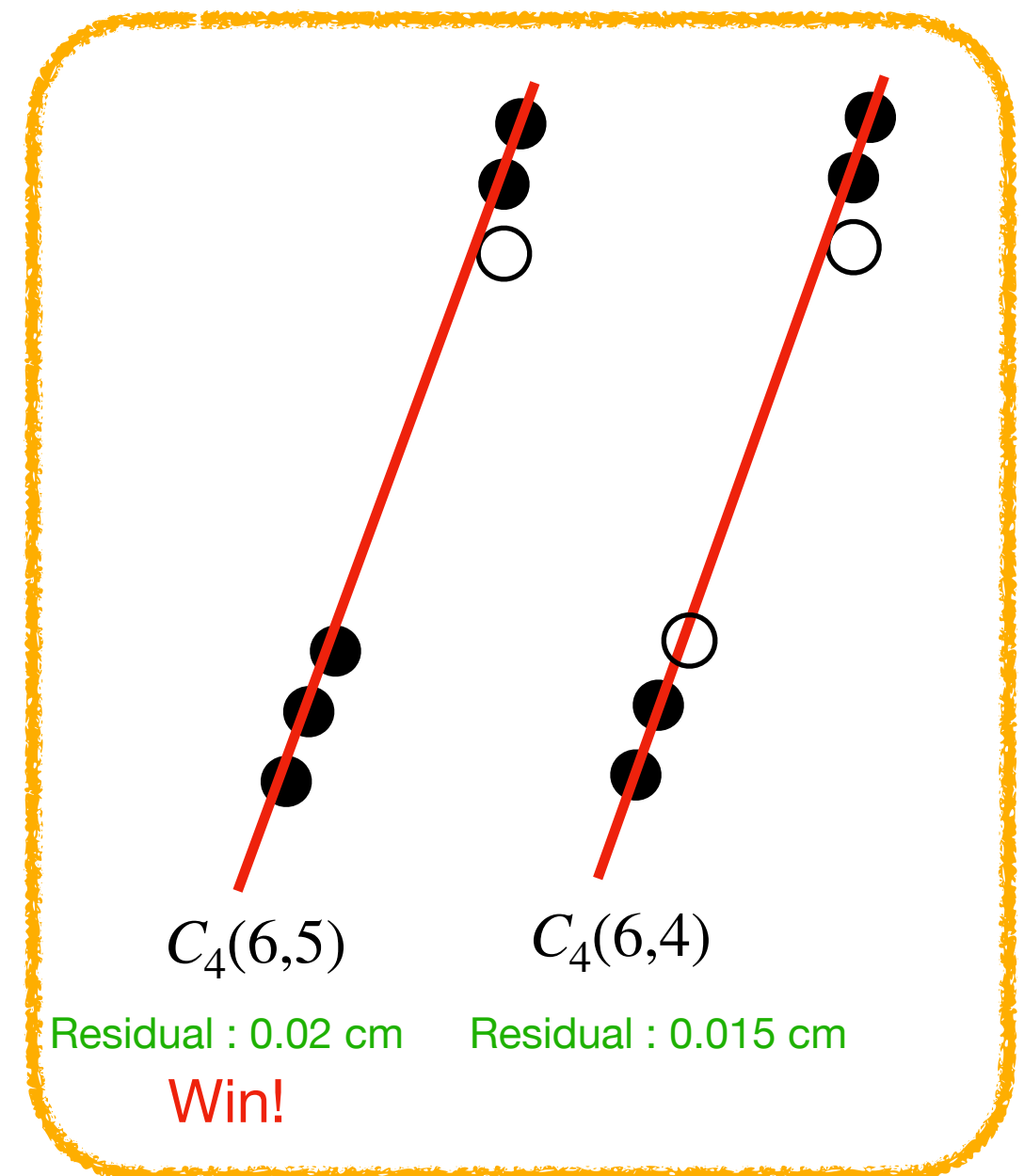
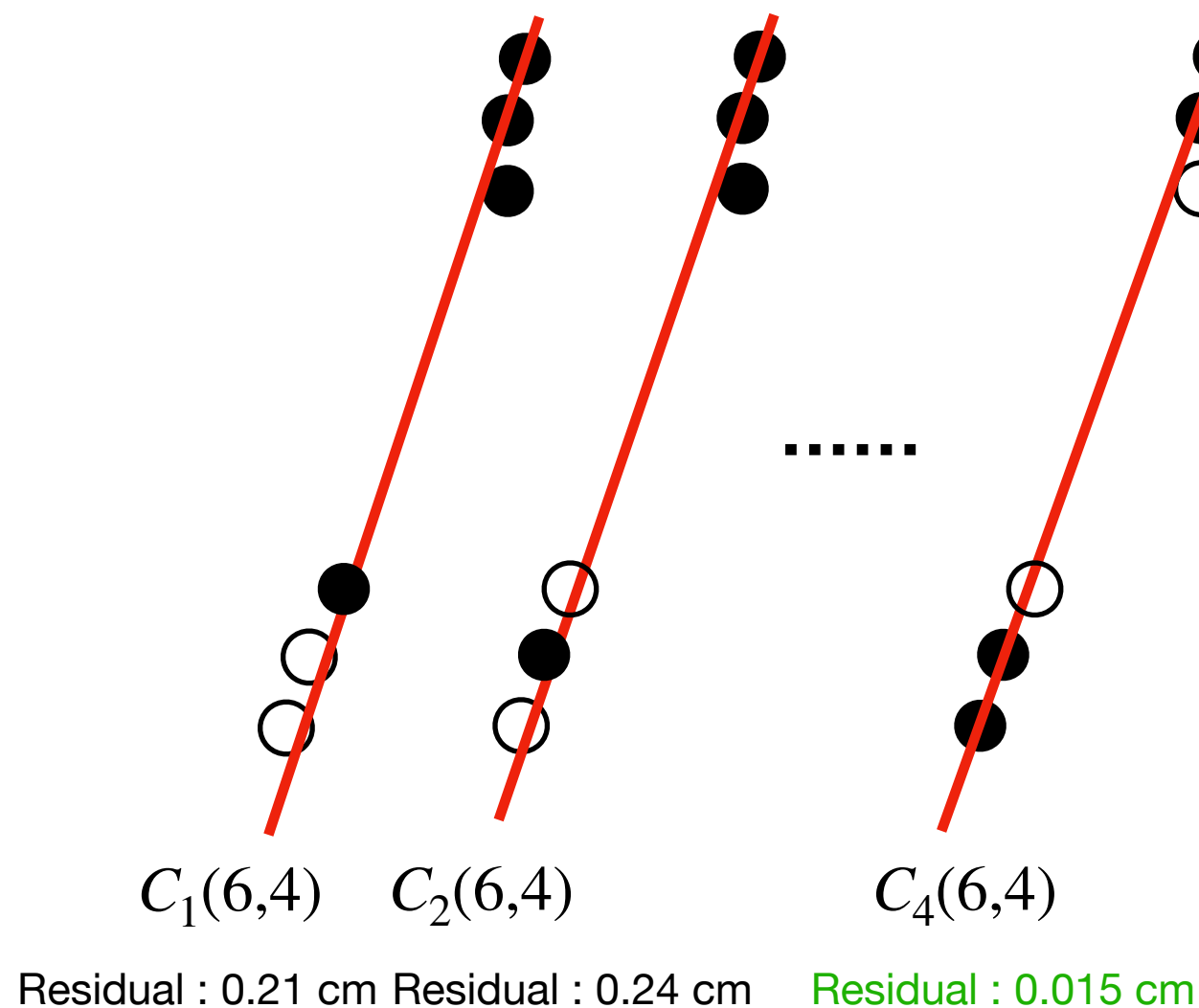
- Prepare all the combinations based on  $C(n, k)$ 
  - $k$  ranges from 6 to 4,  $n = 6$
  - The reduced residual smaller than 0.02 cm can be the candidate.



# Clusters combination

- Prepare all the combinations based on  $C(n, k)$ 
  - Although two combinations' residual is smaller than 0.02cm, the combination of larger clusters is selected.

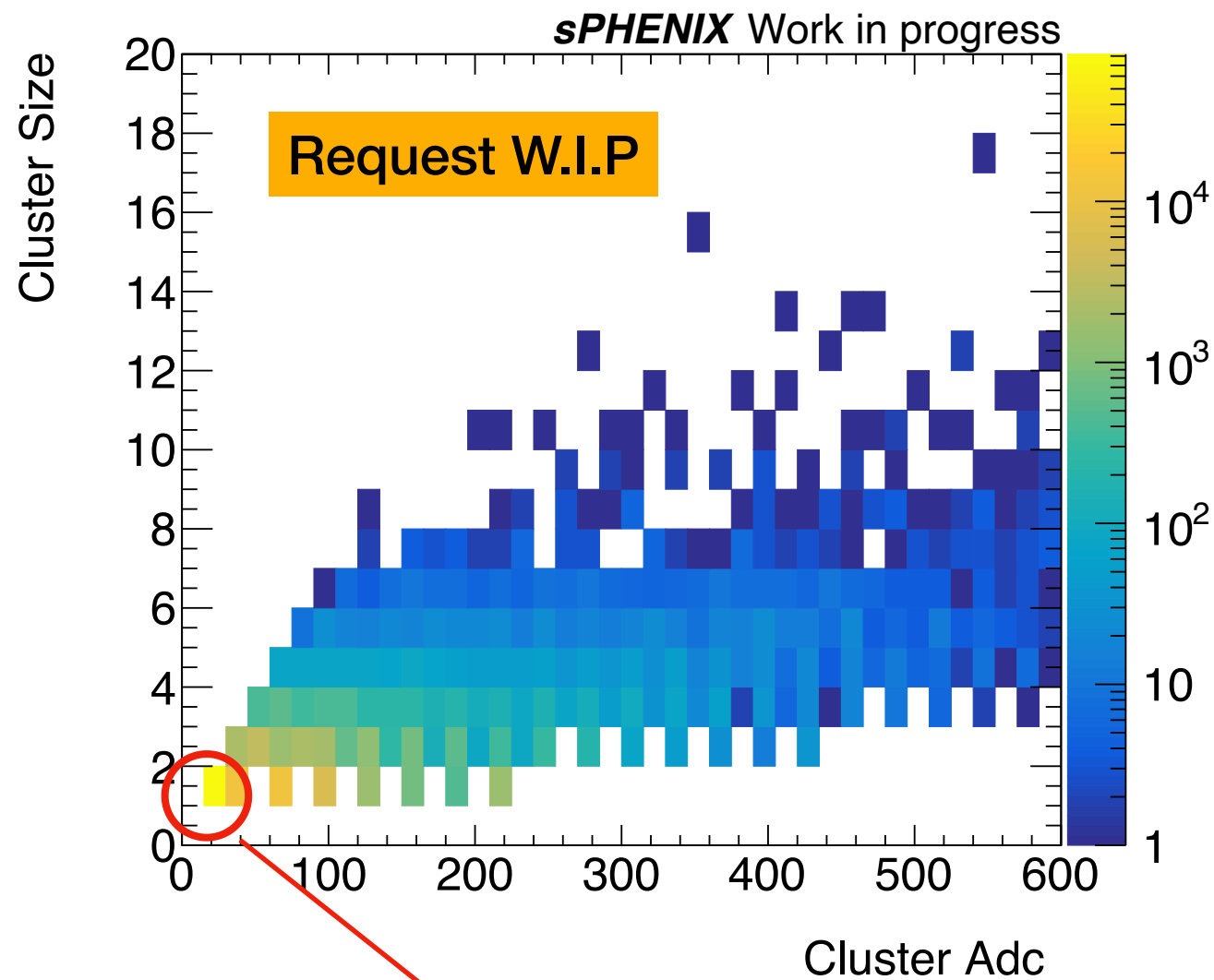
Two candidates



# Good cluster required

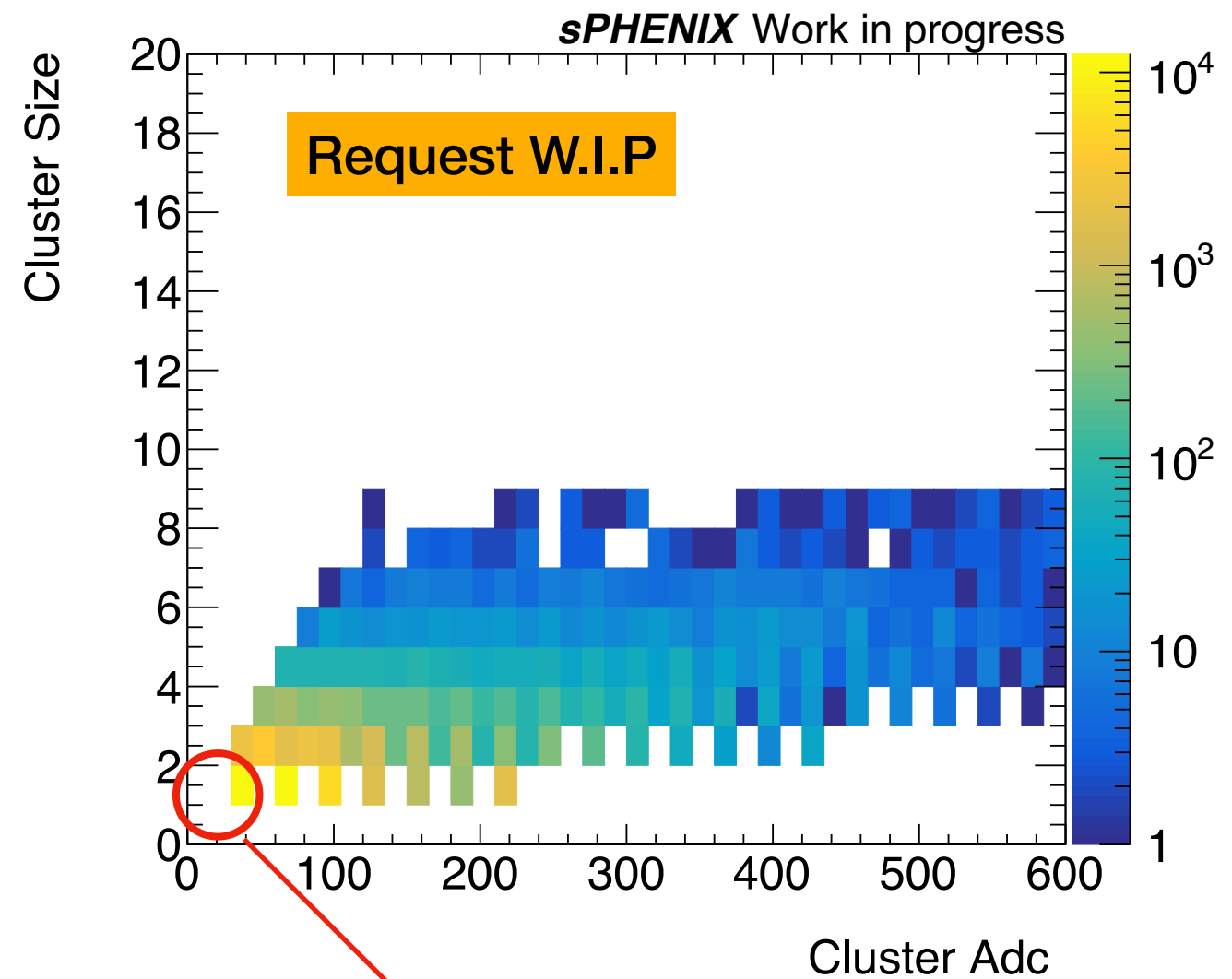
- '*cluster Adc > 15*' & '*cluster phisize < 9*'
- The plots use single run data.

Inclusive



Cluster size = 1 && Adc = 15

Good cluster only

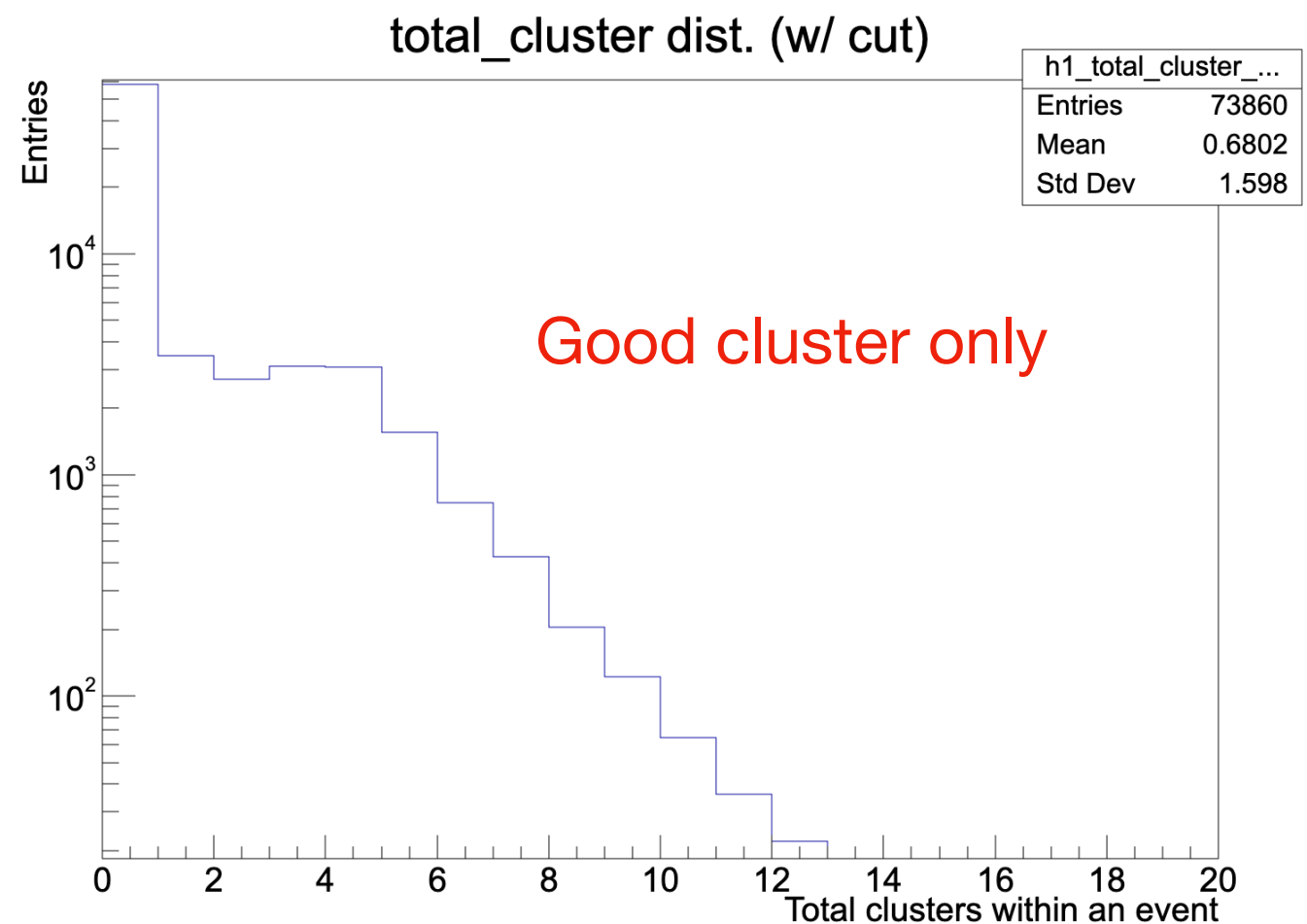
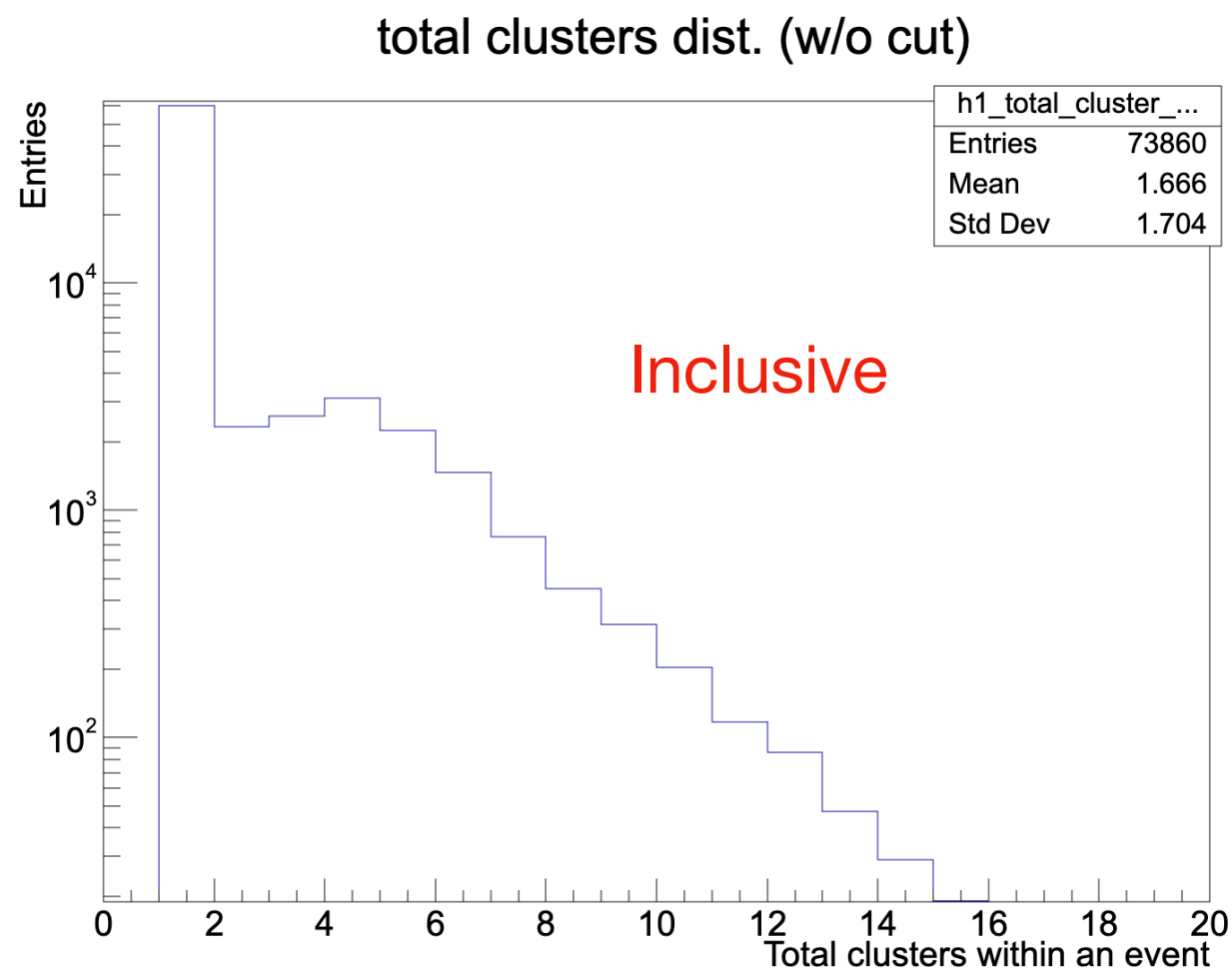


Absence of cluster size = 1 && Adc = 15



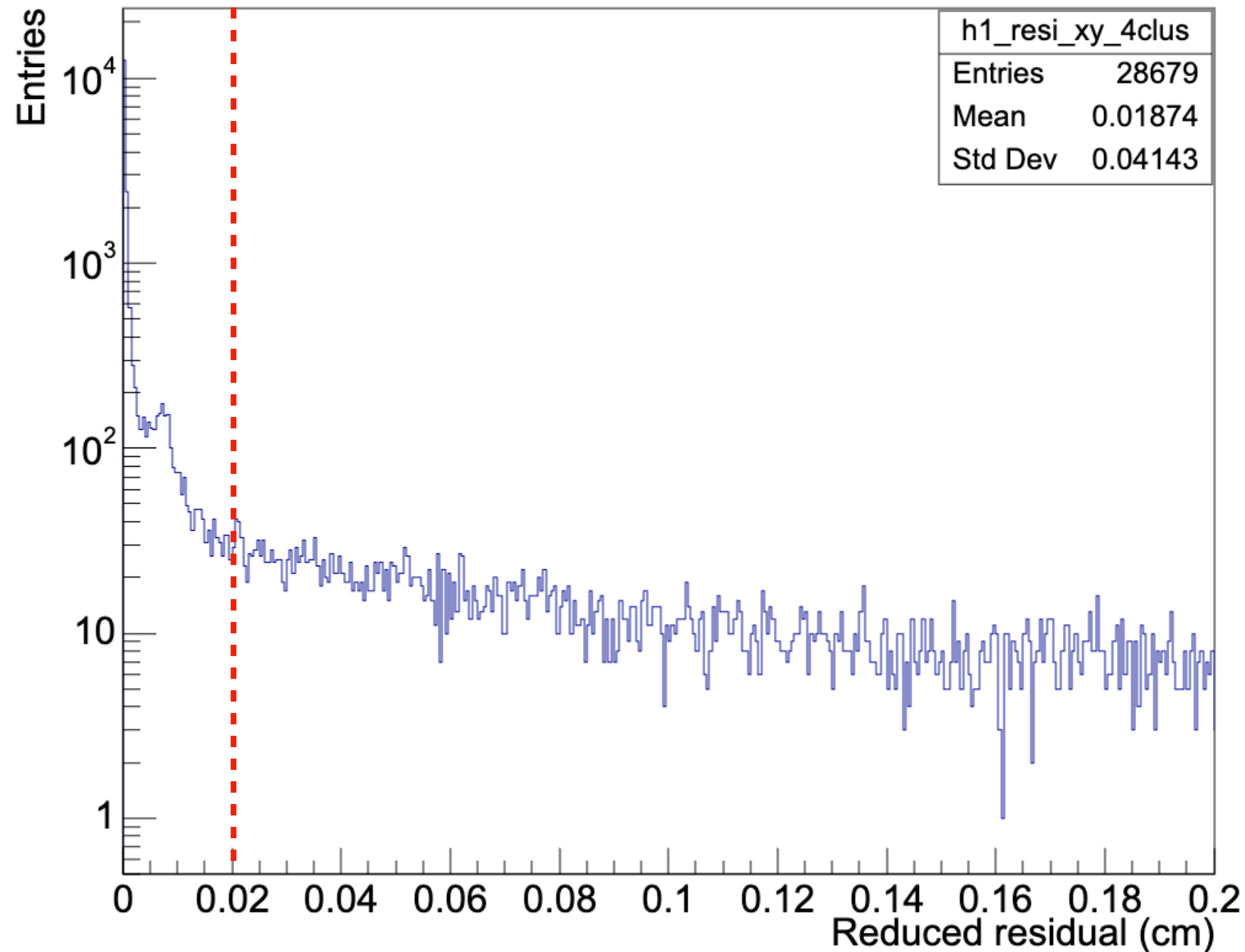
# Good cluster required

- The plots use single run data.
- Empty events are excluded.
- A large fraction of '*cluster size = 1*' clusters are excluded.



# Reduced residual distribution

- The reduced residual distribution for **4 clusters only** in **X-Y plane**.  
residual dist.



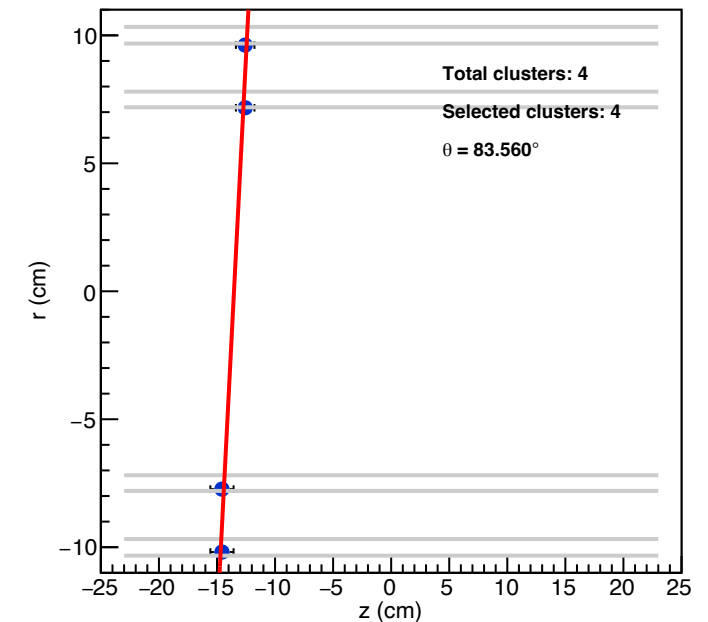
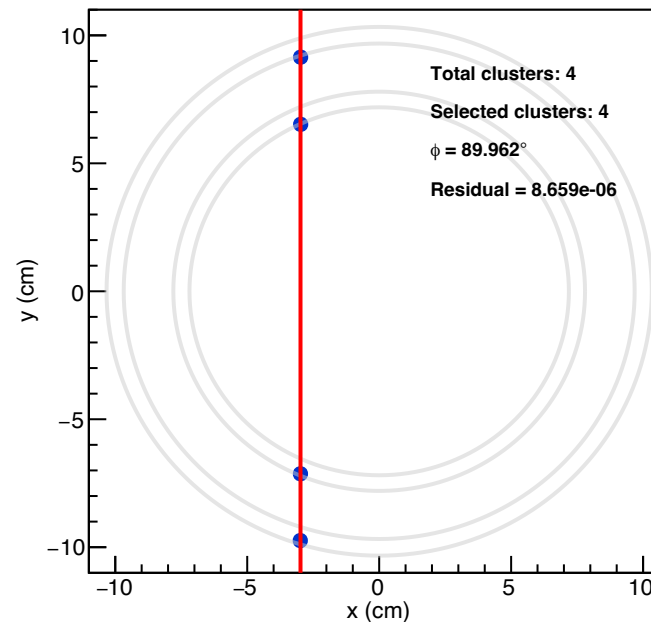
# Event displays for 4 clusters only

Request W.I.P

Event : 23

sPHENIX Work in progress

- Good track!
- Reduced residual < 0.02 cm

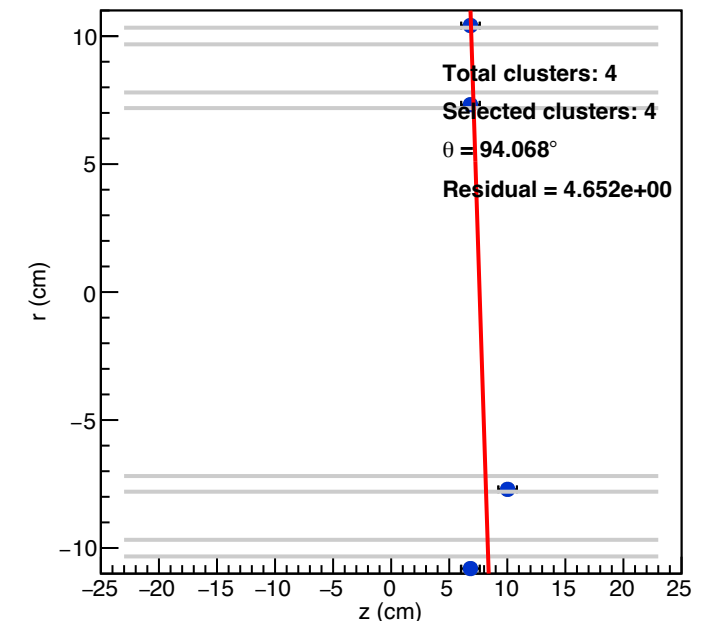
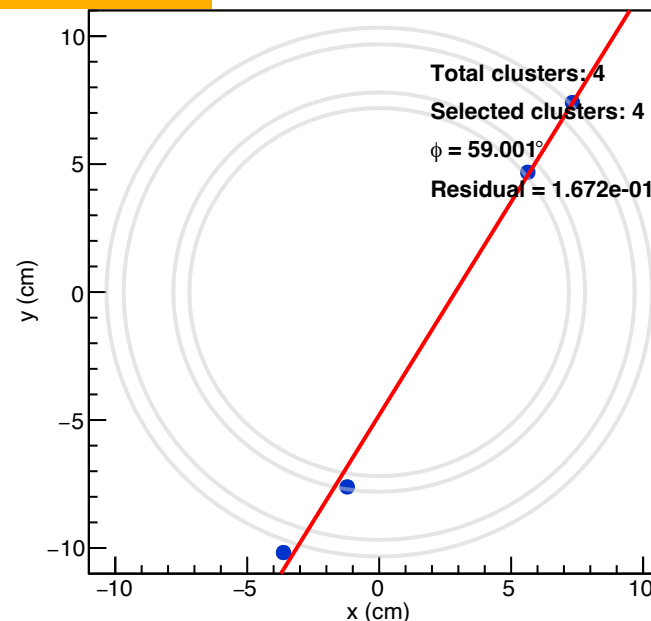


Request W.I.P

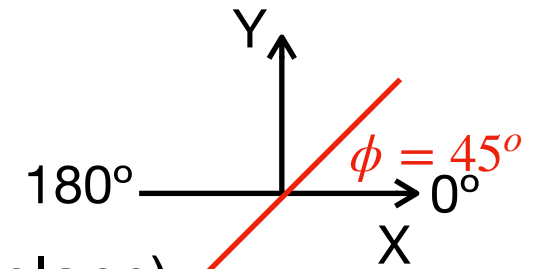
Event : 5270 (C(4,4) combination 0)

sPHENIX Work in progress

- Bad track!
- Reduced residual > 0.02 cm

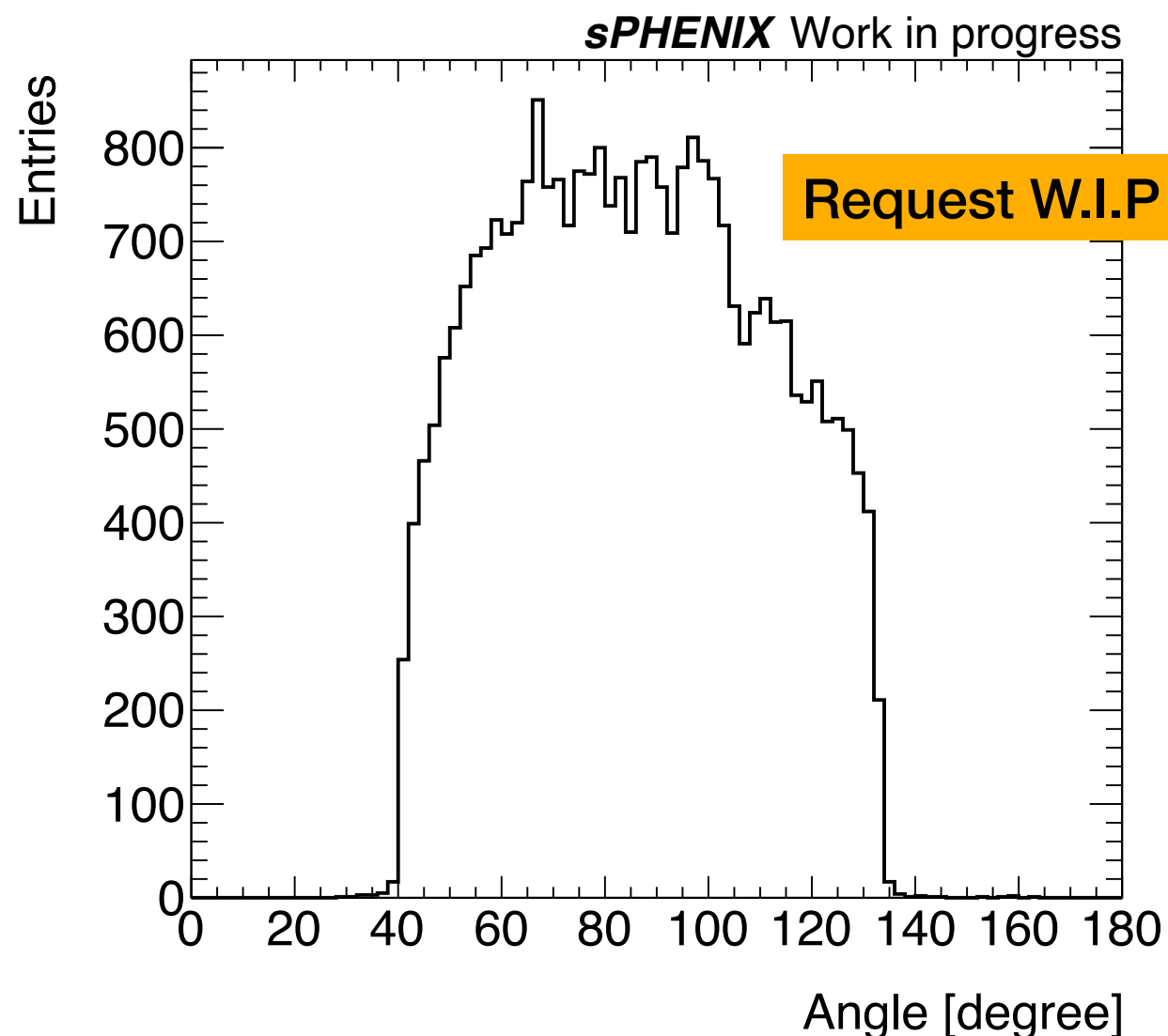


# Angle distribution again

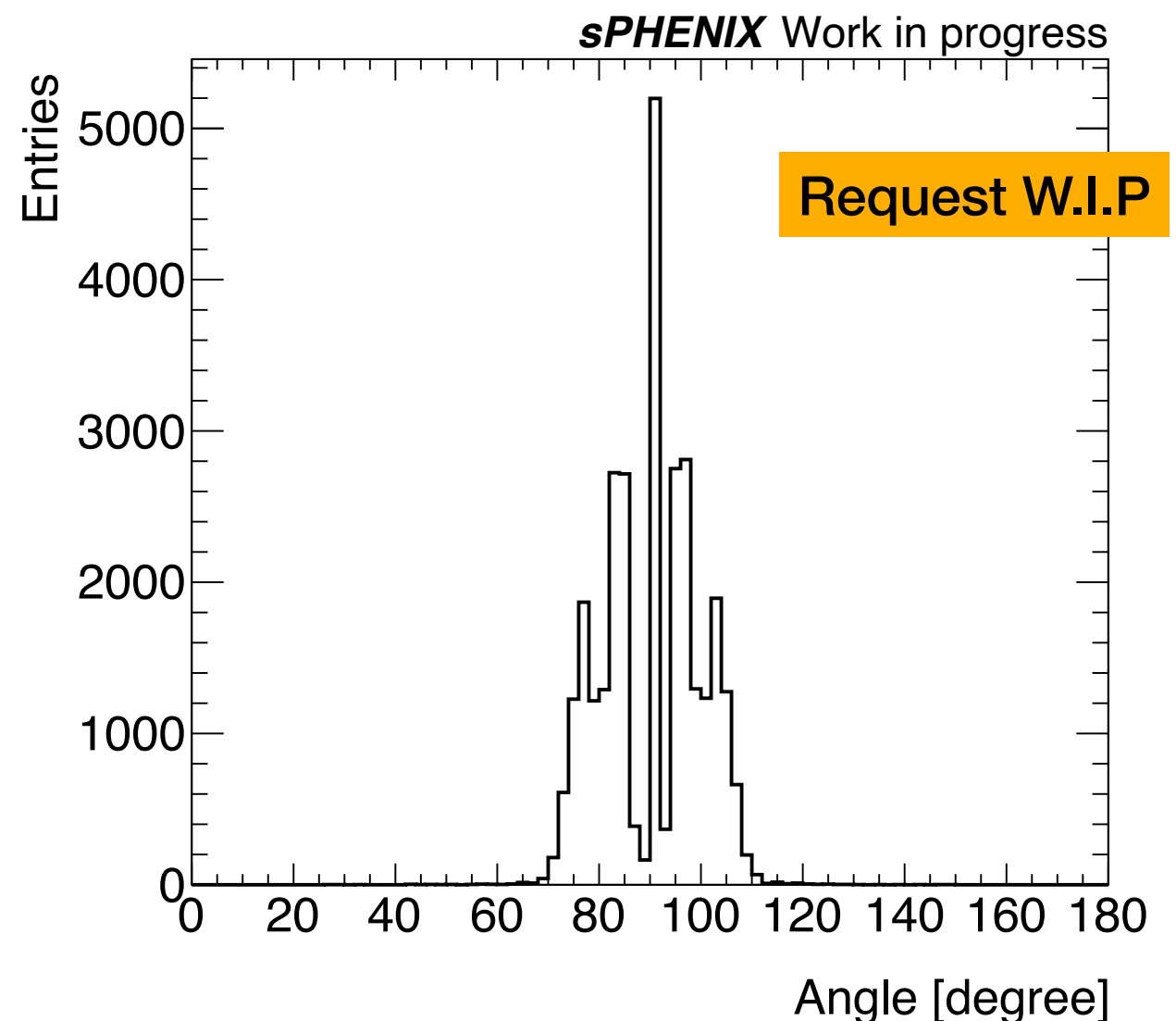


- The distribution of the  $\phi$  angle (X-Y plane) and  $\theta$  angle (Z-Radius plane)
- X-Y residual  $< 0.02\text{cm}$ , all 4 clusters cover the fitting line in Z-Radius plane.

$\phi$  angle distribution in xy plane

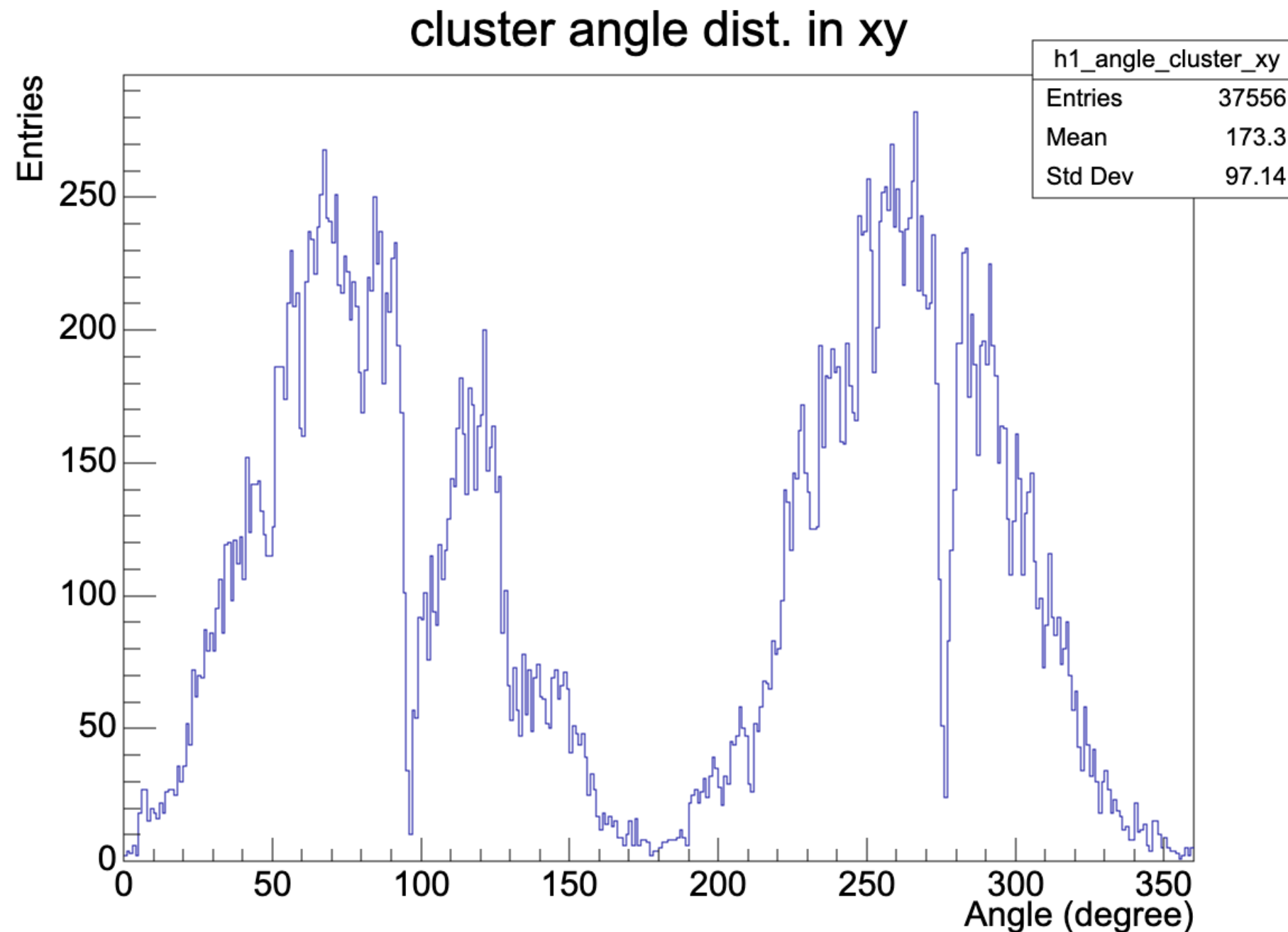
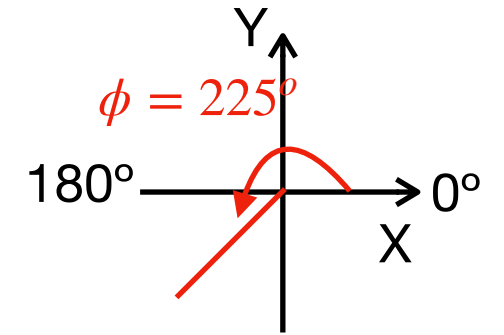


$\theta$  angle distribution in zr plane



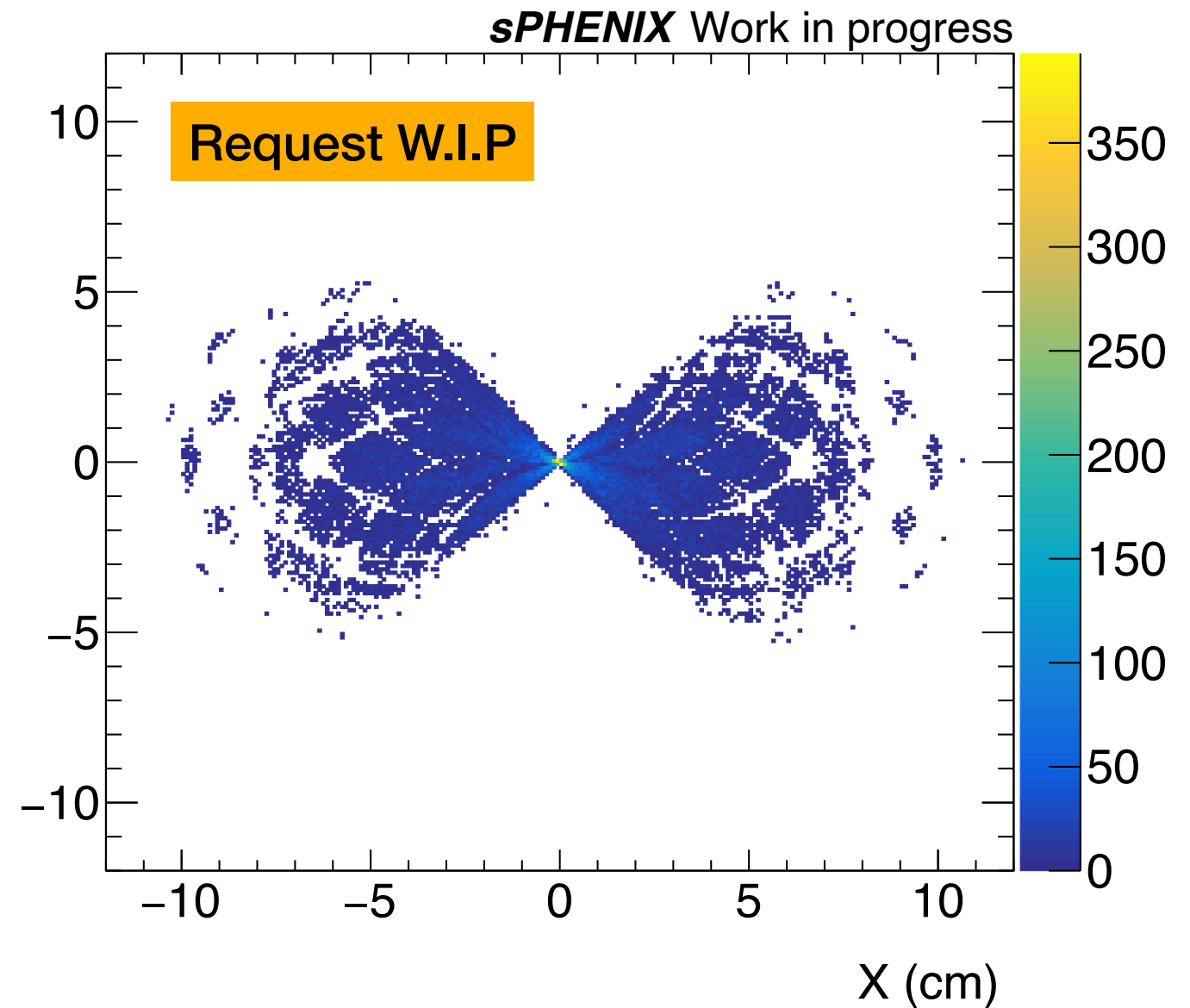
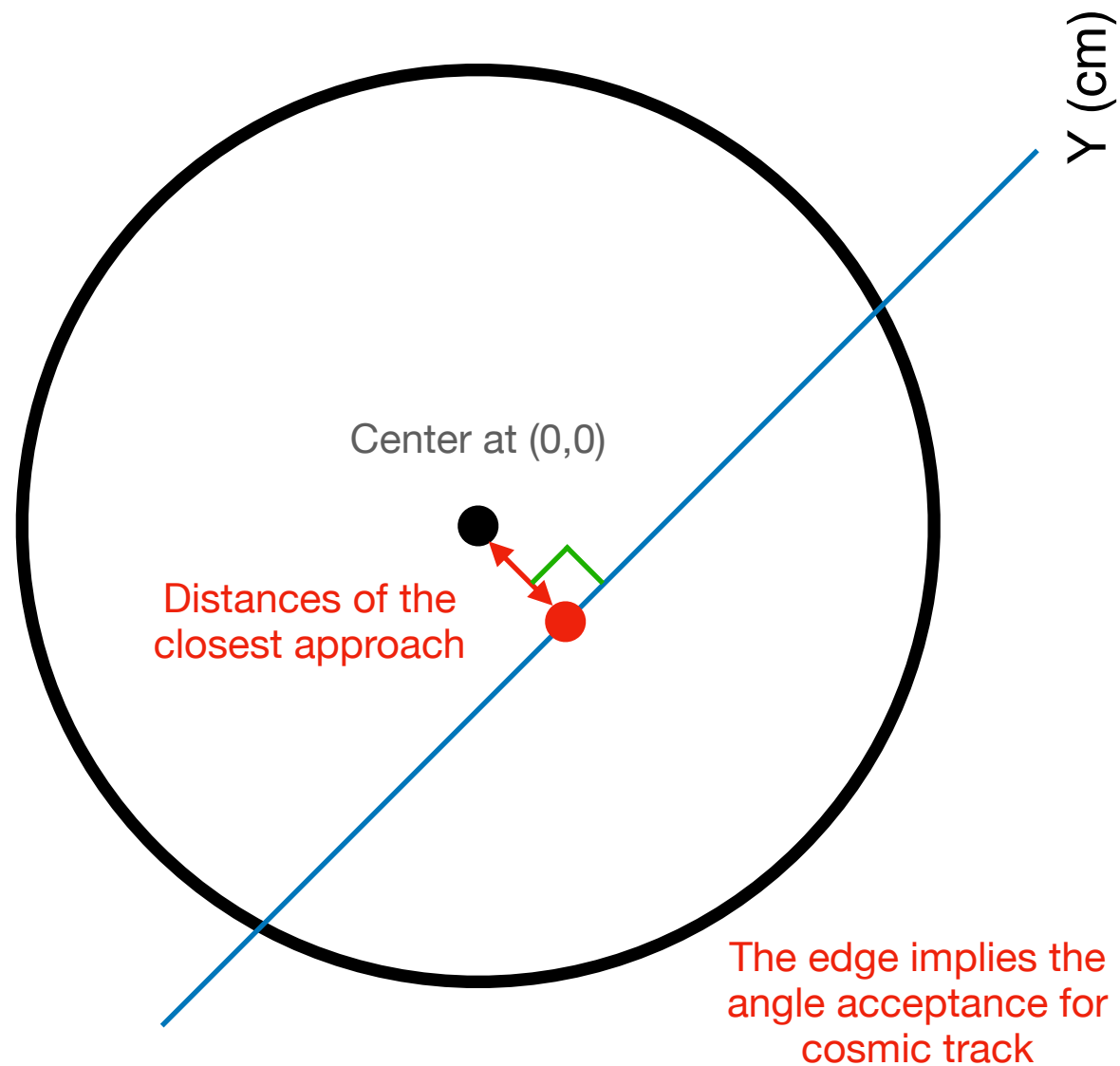
# Cluster angle distribution

- The clusters of good tracks in  $\phi$  angle distribution in X-Y plane.
- The distribution is in symmetry shape.



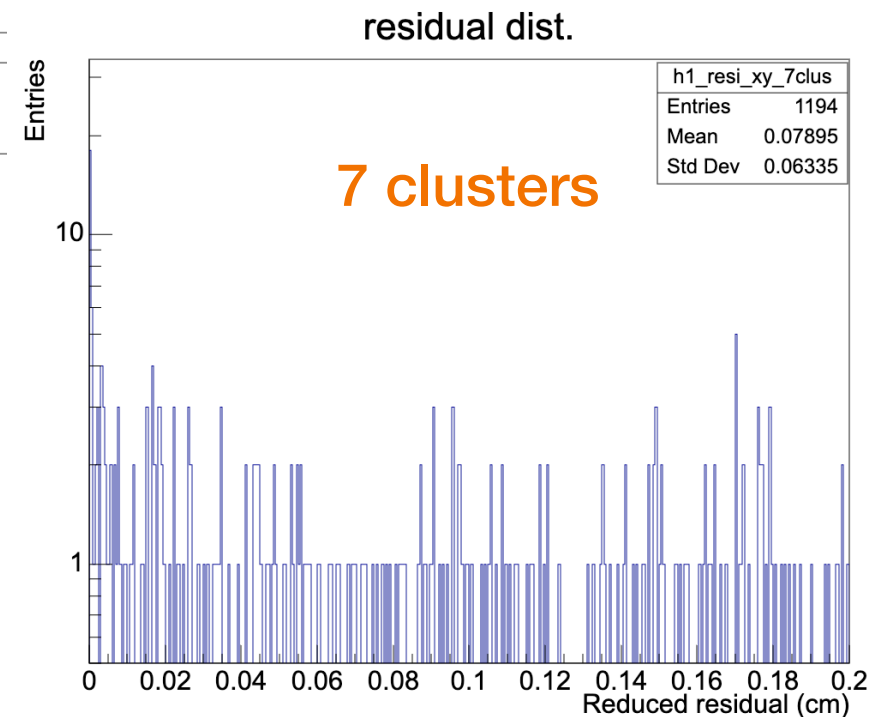
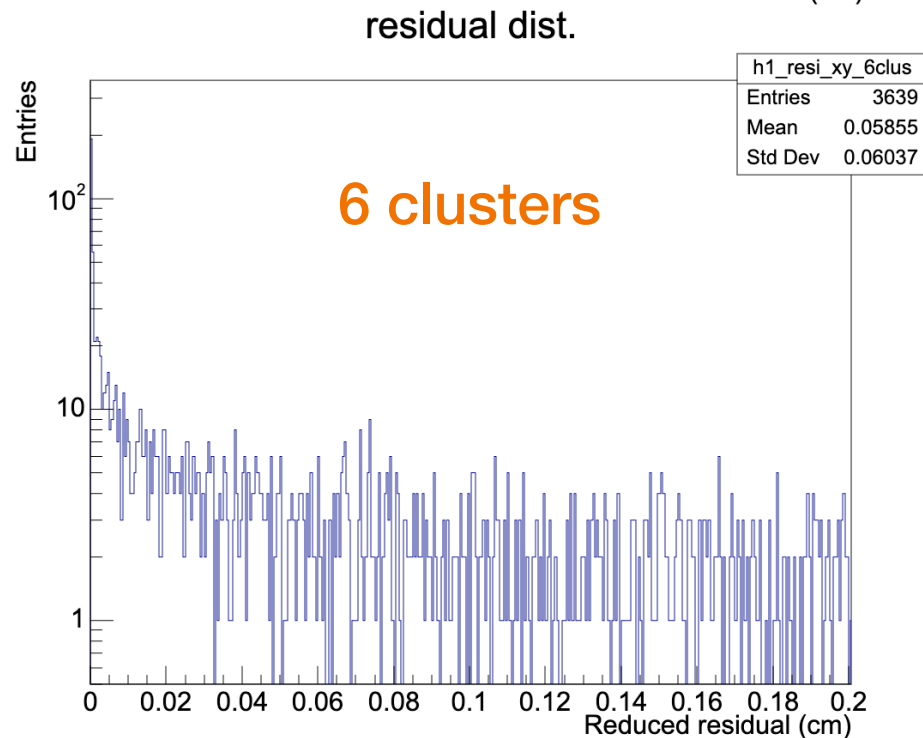
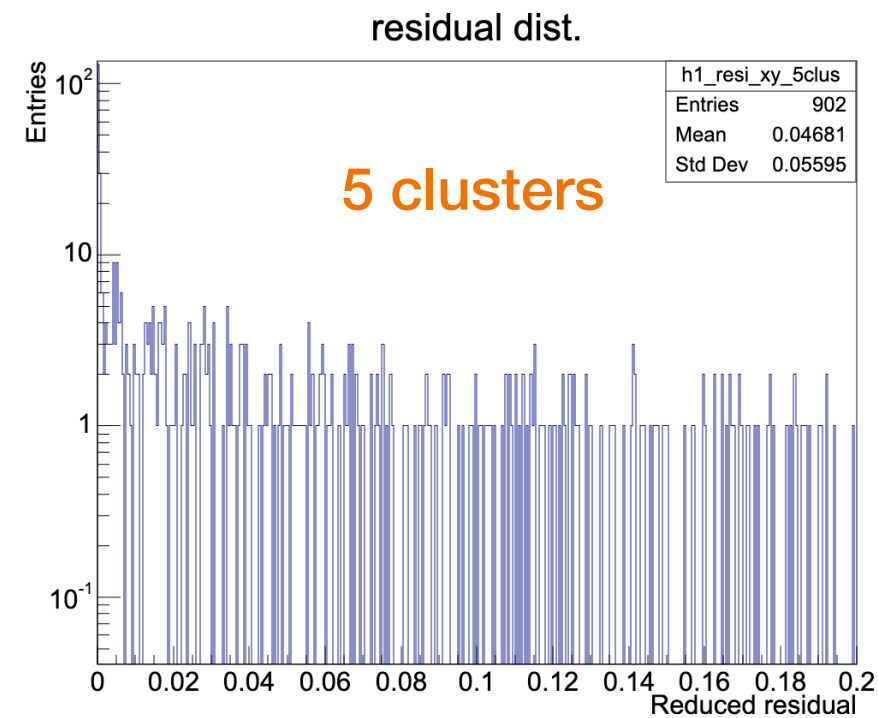
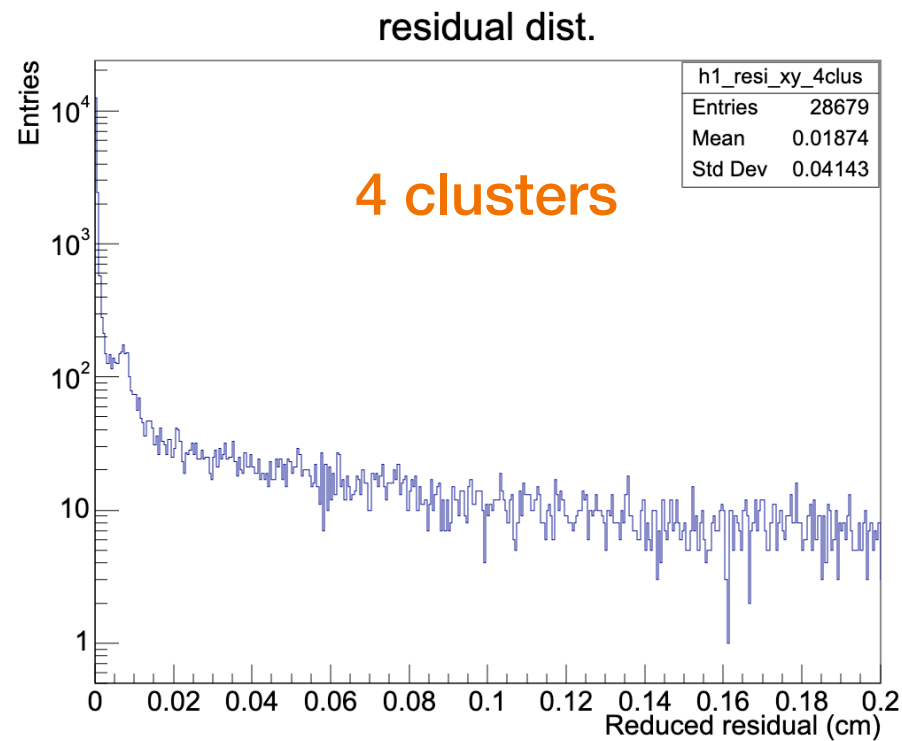
# The DCA points for tracks

- The plot also implies the acceptance angle of cosmic tracks.



# Reduced residual

- The reduced residual of different clusters tracks.

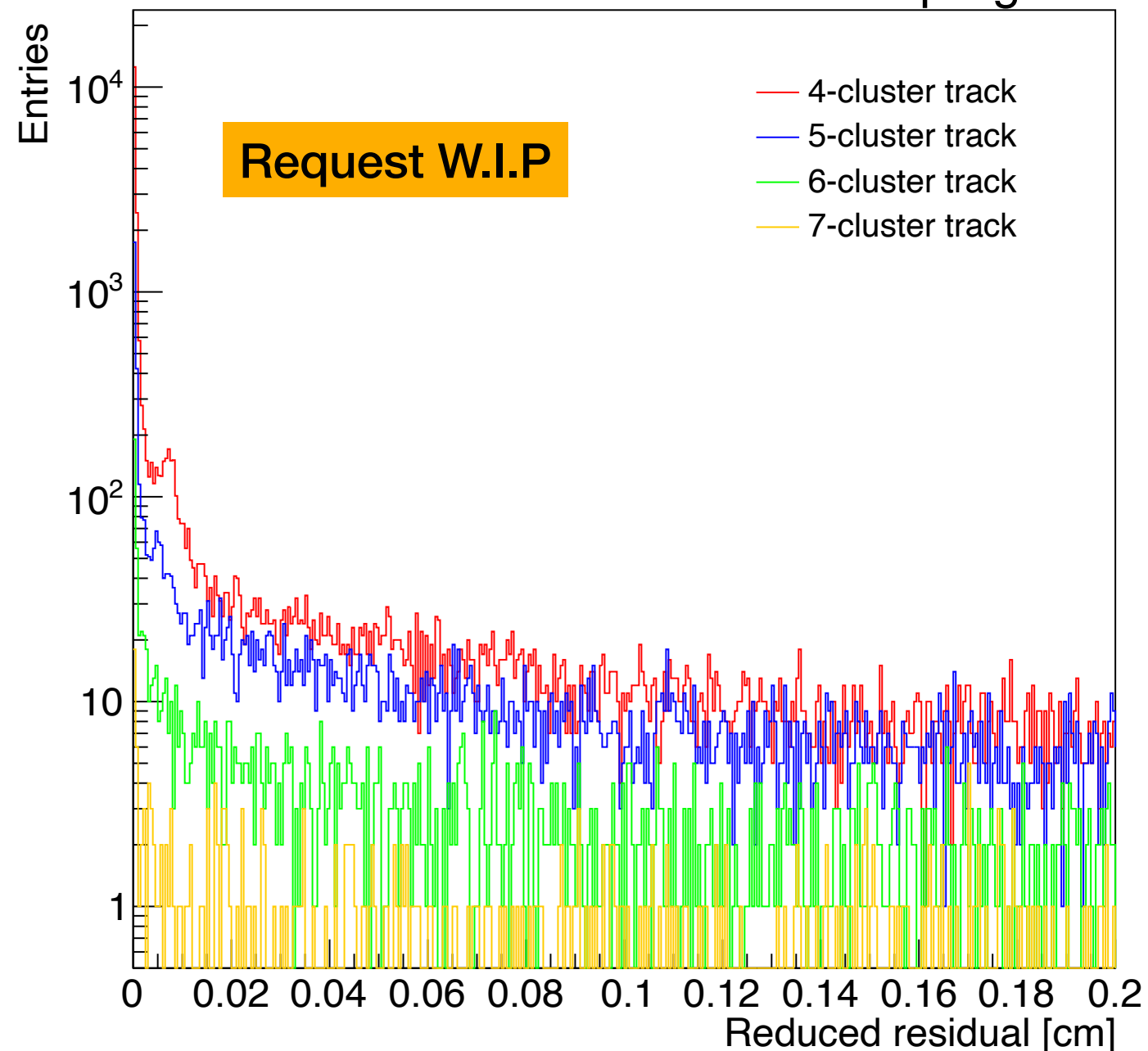


# Reduced residual overlap

- No significant differences between the distributions of the clusters.

Reduced residual distribution

*s*PHENIX Work in progress

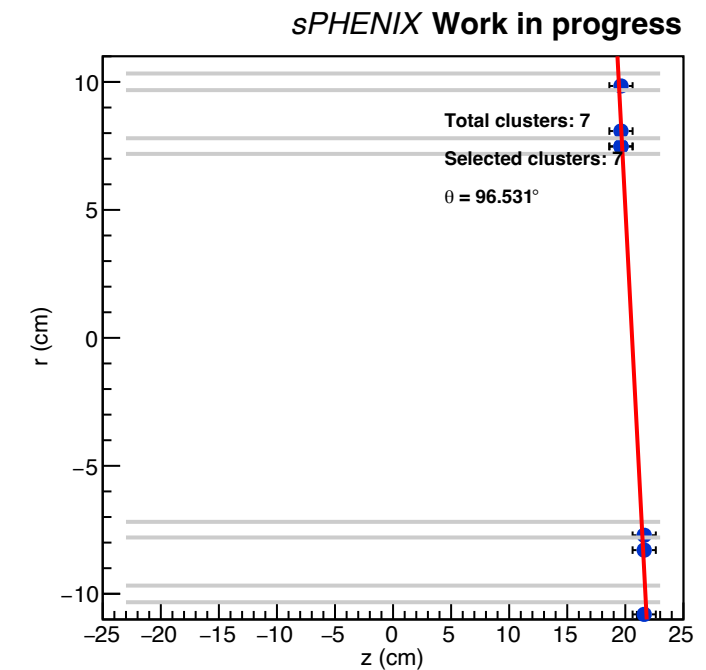
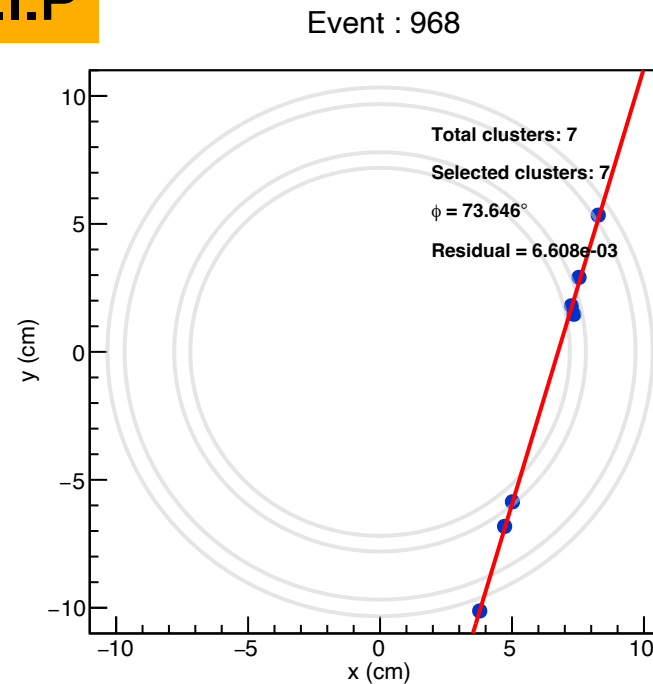




# Event displays for good tracks

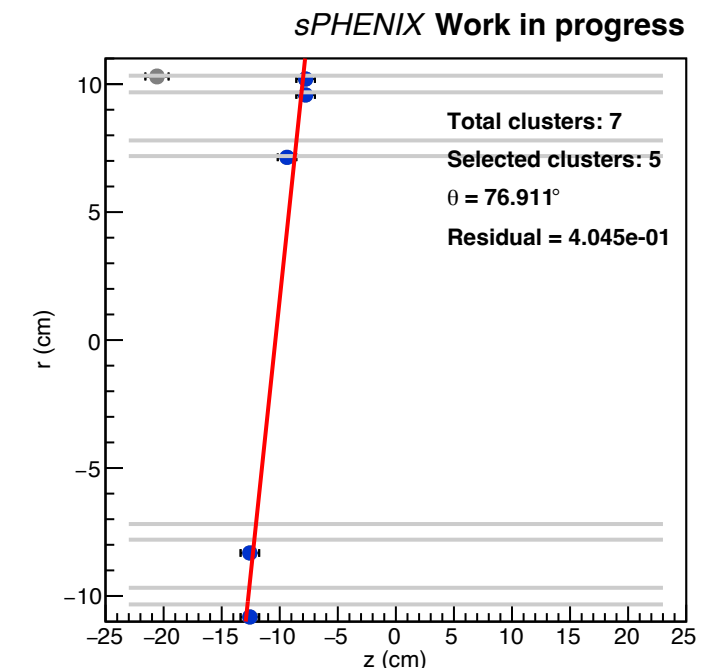
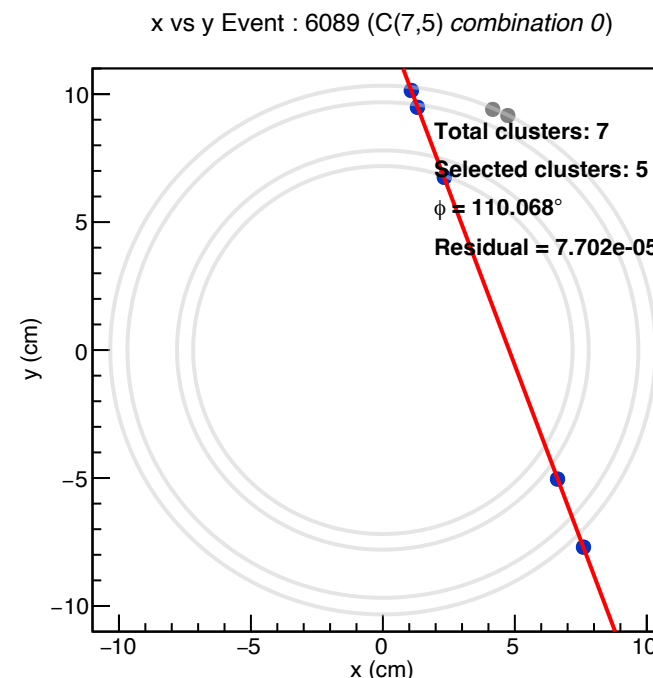
Request W.I.P

- Good track!
- Reduced residual  $< 0.02$  cm with 5 clusters



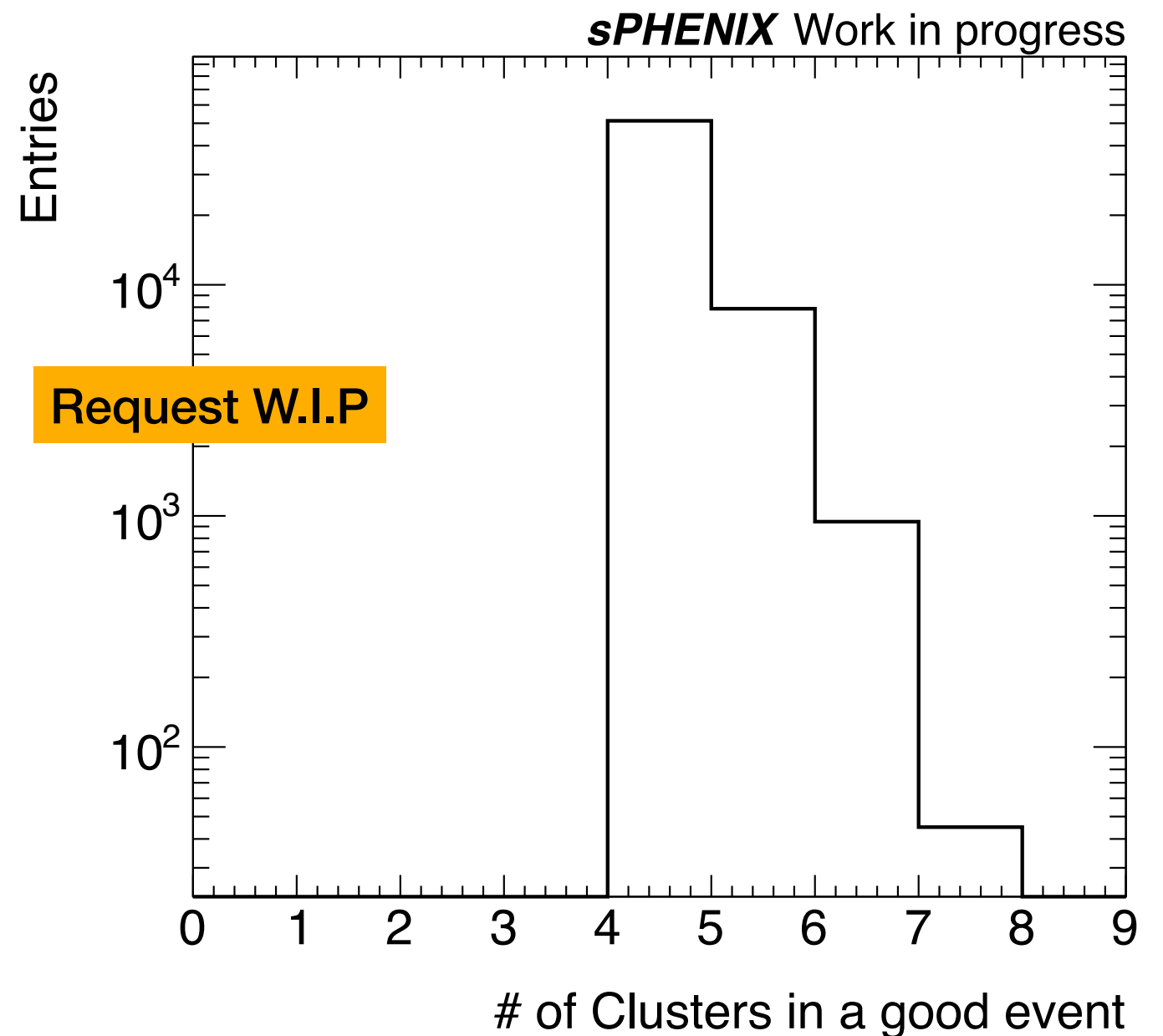
Request W.I.P

- Good track!
- Reduced residual  $< 0.02$  cm with 5 clusters (2 noise hits excluded)

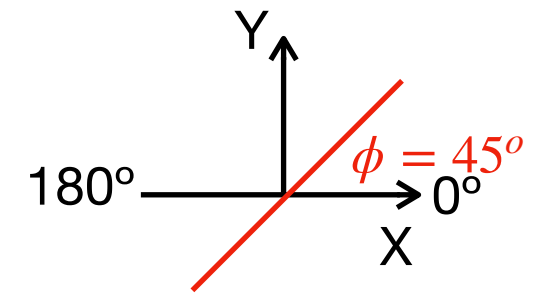


# # of clusters in a good track

- Reconstruction efficiency :  $\frac{\text{Good tracks}}{\text{Total events}} = \frac{30K}{4200K} = 0.7\%$ 
  - 4 clusters/total good clusters = 85 %
  - 5 clusters/total good clusters = 13 %
  - 6 clusters/total good clusters = 1.5 %
  - 6 clusters/total good clusters = 0.5 %

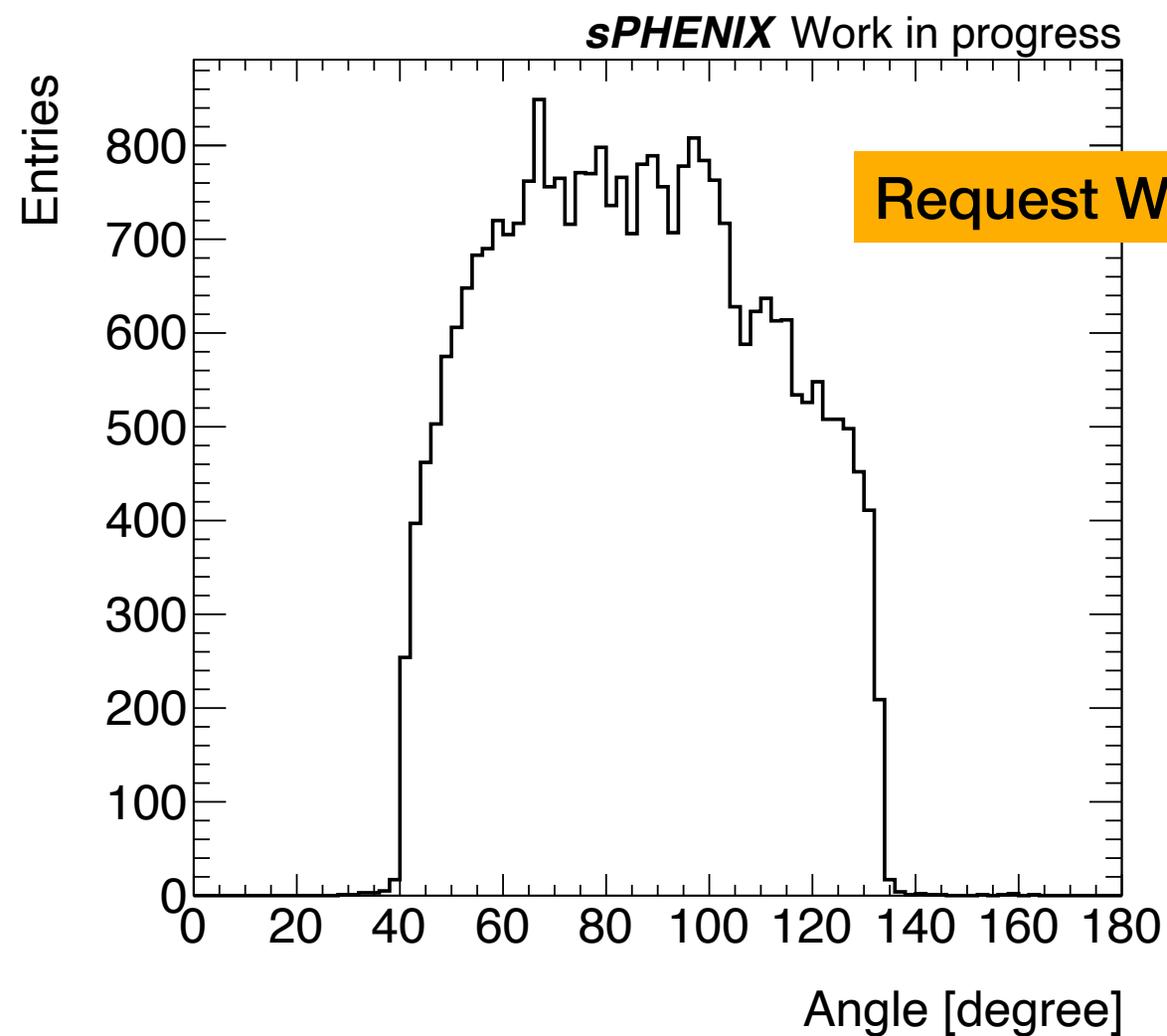


# Angle distribution

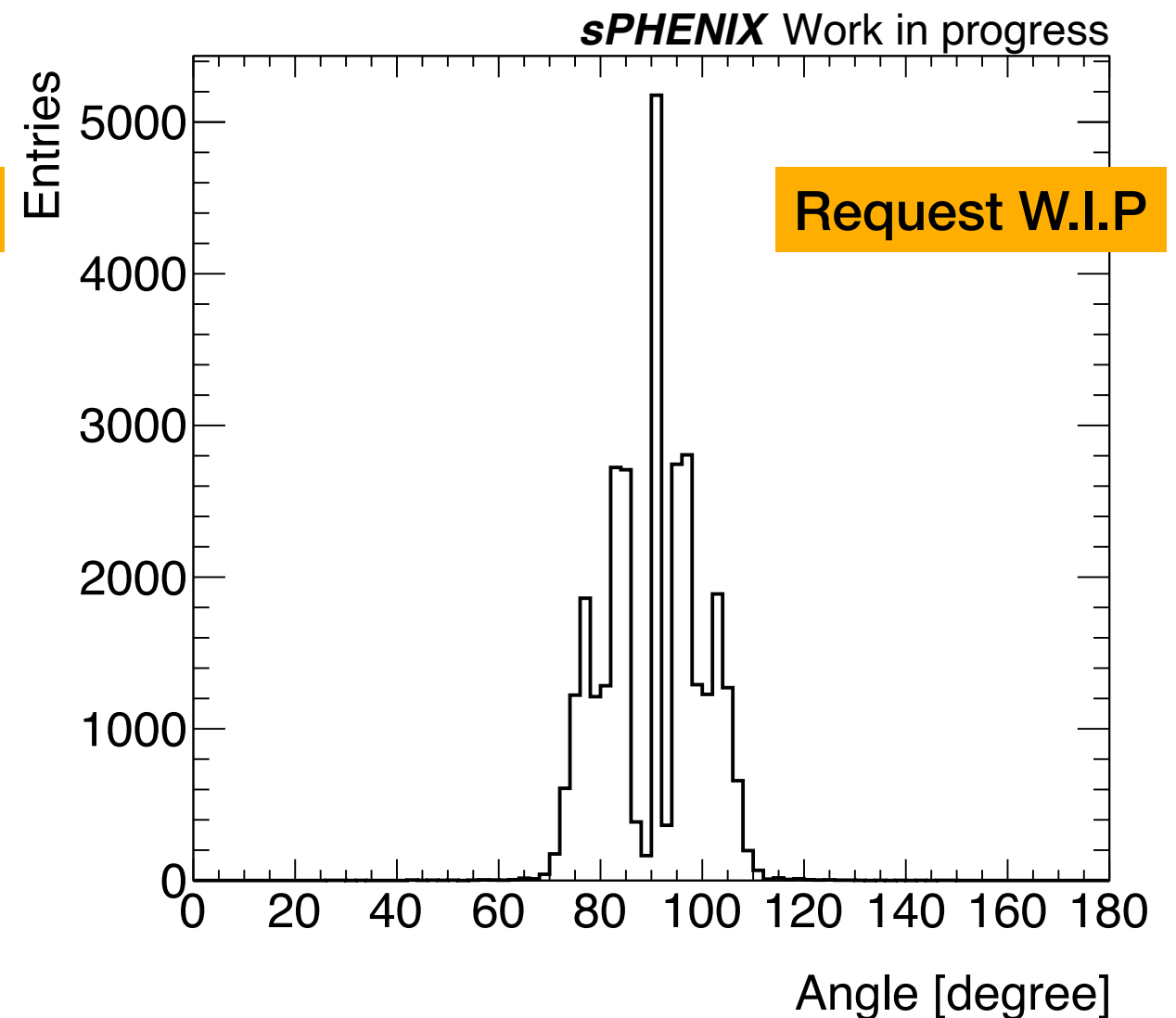


- The distribution of the  $\phi$  angle (in X-Y plane) and  $\theta$  angle (in Z-Radius plane).
- X-Y residual  $< 0.02\text{cm}$ , full cluster coverage of the fitting line in Z-Radius plane.

$\phi$  angle distribution in xy plane



$\theta$  angle distribution in zr plane



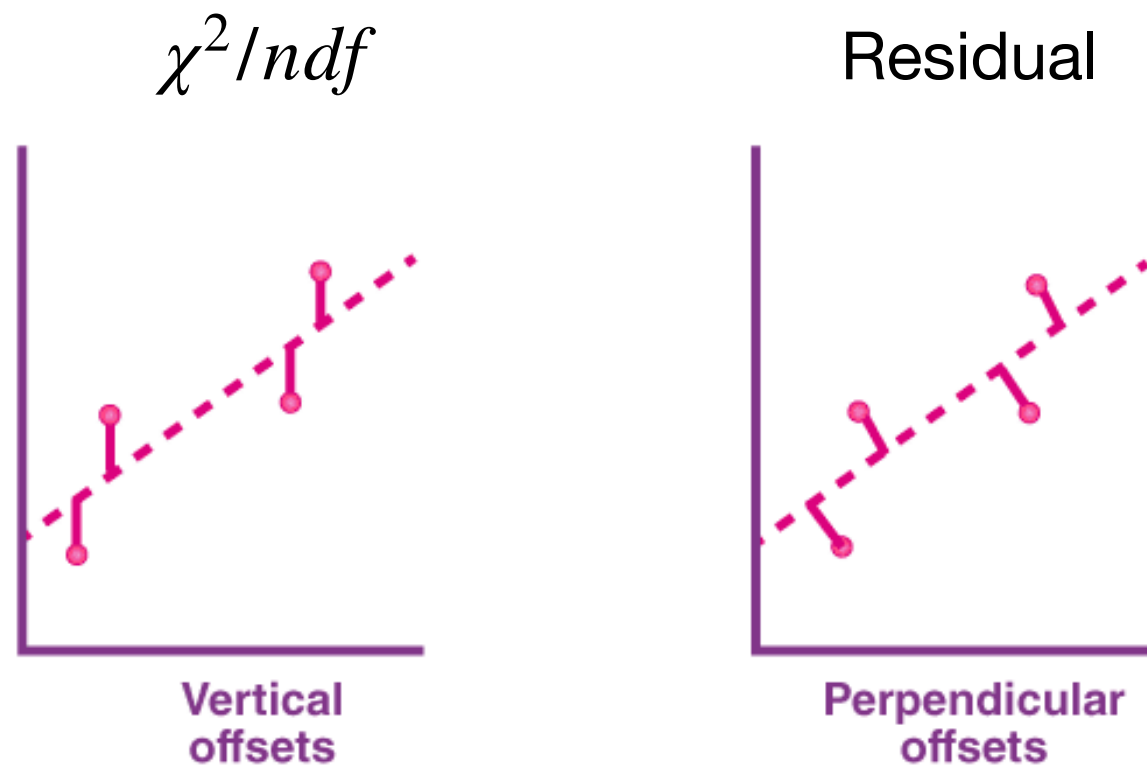
# Summary

- Roughly 25 runs from 2024 cosmic data is analyzed.
  - INTTRawHit DSTs are from official production.
  - INTTTrkrCluster DSTs are from private production using Fun4All framework.
  - Good cluster required.
  - Two different approaches are performed to this analysis.
- Look into the distribution of cluster size, cluster Adc and total clusters.
  - Compare the distribution before and after the cluster cut
    - > Those distributions seem to be normal.
- Residual distribution:
  - Distinguish good clusters with reduced residual smaller than 0.02cm.
  - Event displays support the choice of reduced residual cut.
- Angle distributions are presented.

Back up

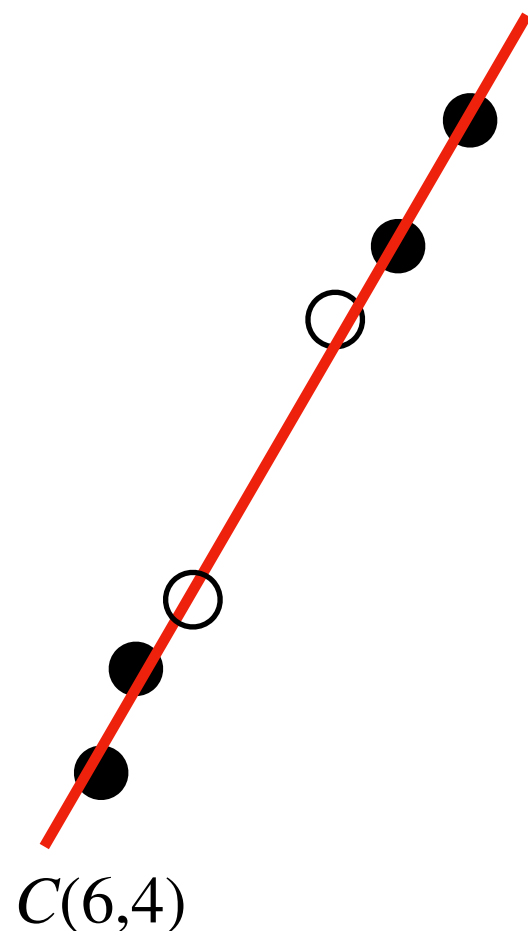
# Residual distribution check

- The cosmic tracks are fitted with a straight line,  $y = c + mx$
- Residual distribution of the tracks. (instead of  $\chi^2/ndf$  distribution)
- Because the cosmic tracks mostly come from vertically, calculating the residual by perpendicular way is a better method.



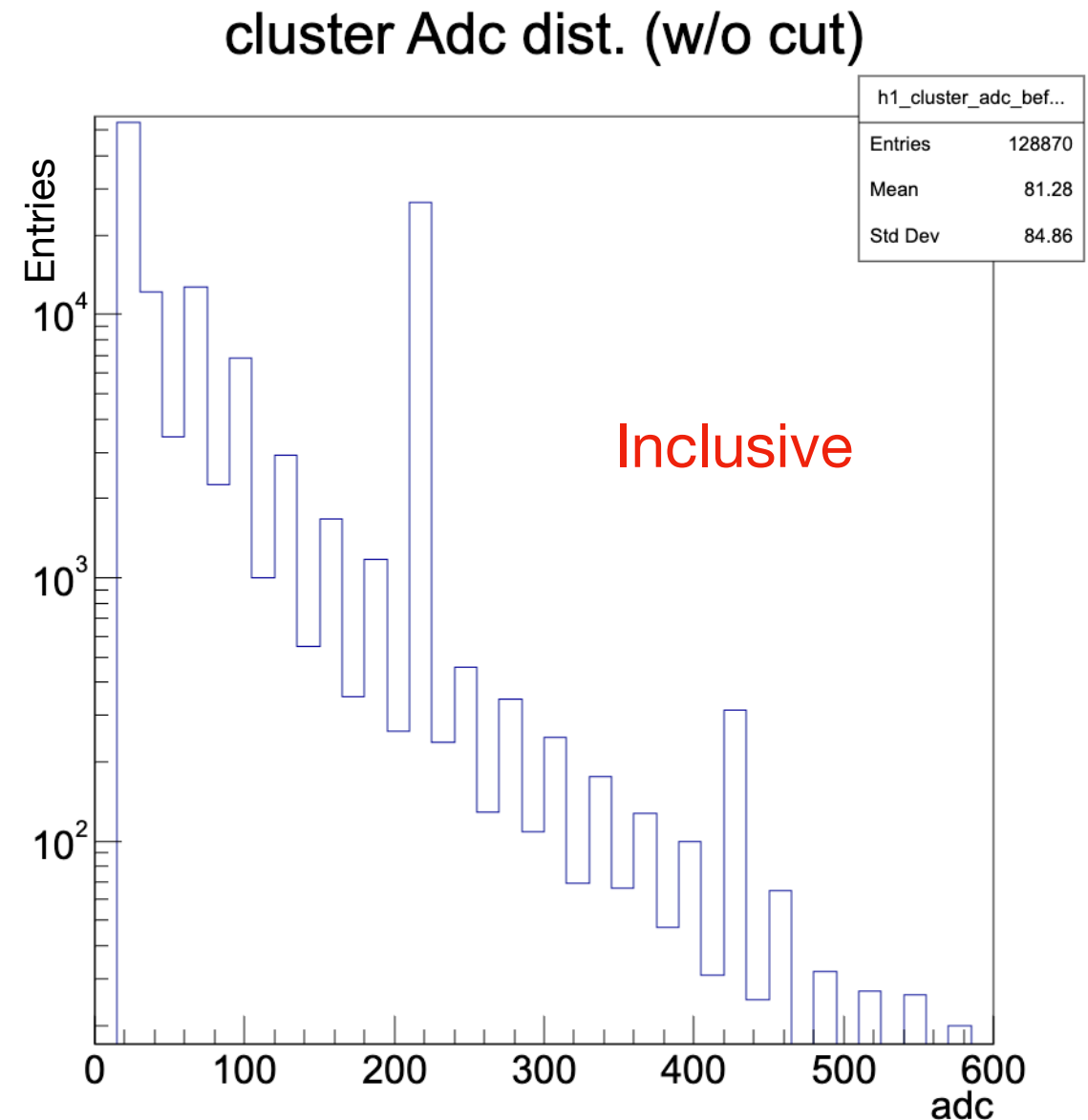
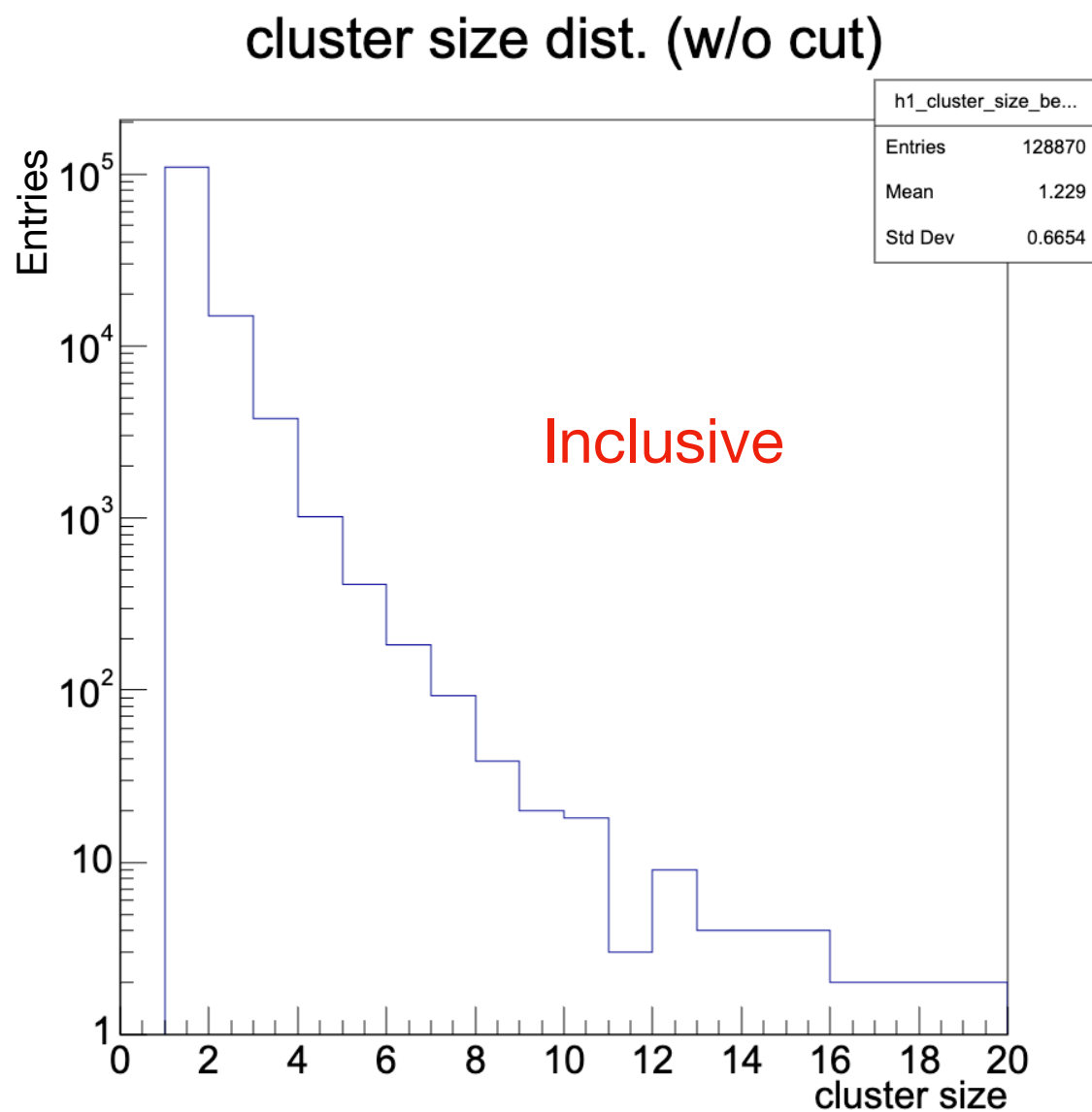
# Residual distribution check

- There are some events are not only have 4 clusters, but even more.
- There are combinations of clusters  $C(n, k)$  for those events.
- For the tracks that have more than 4 clusters, if the residual of them is similar or close to ones with 4 clusters, they may can be accepted.



# Run data check

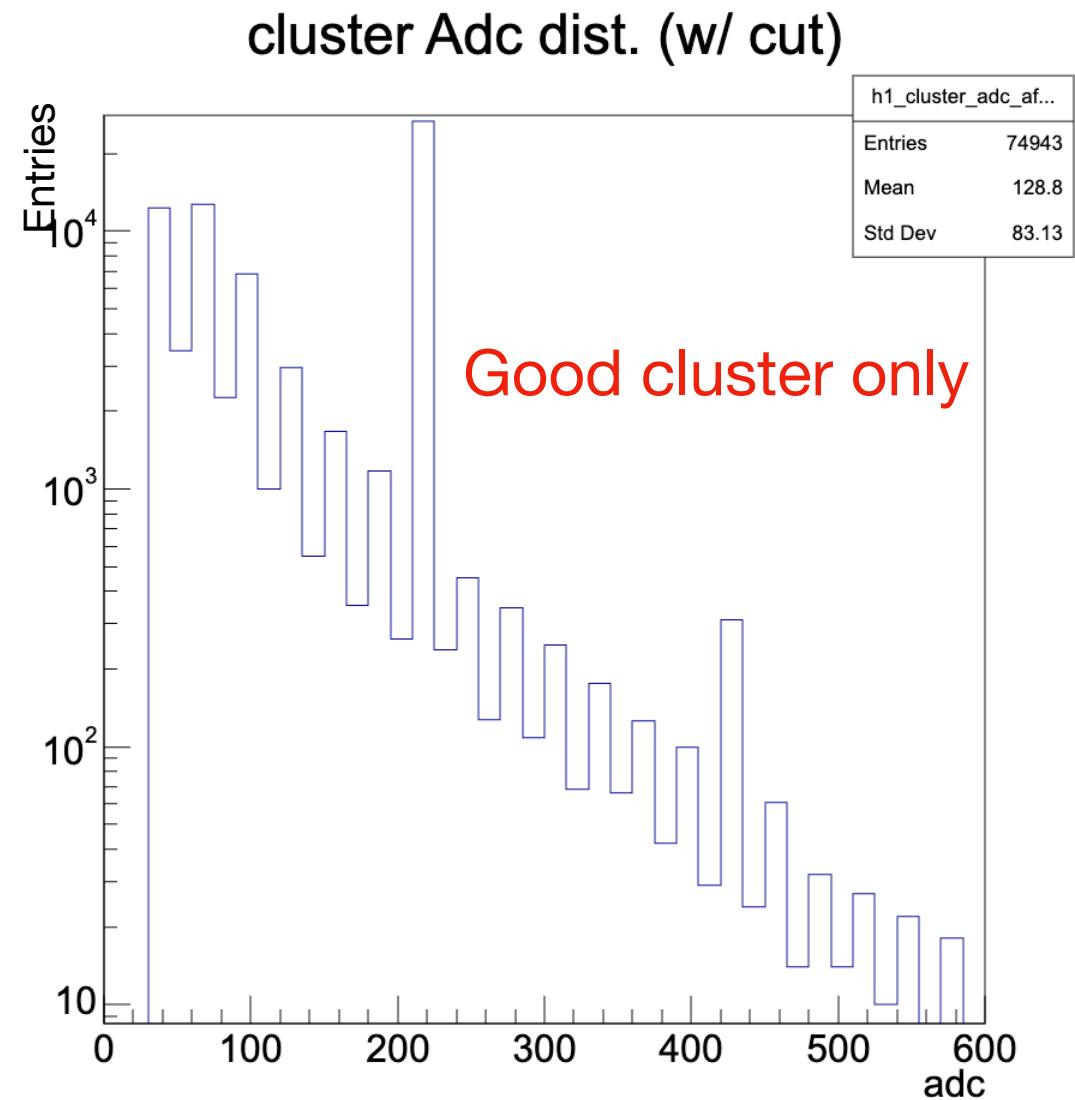
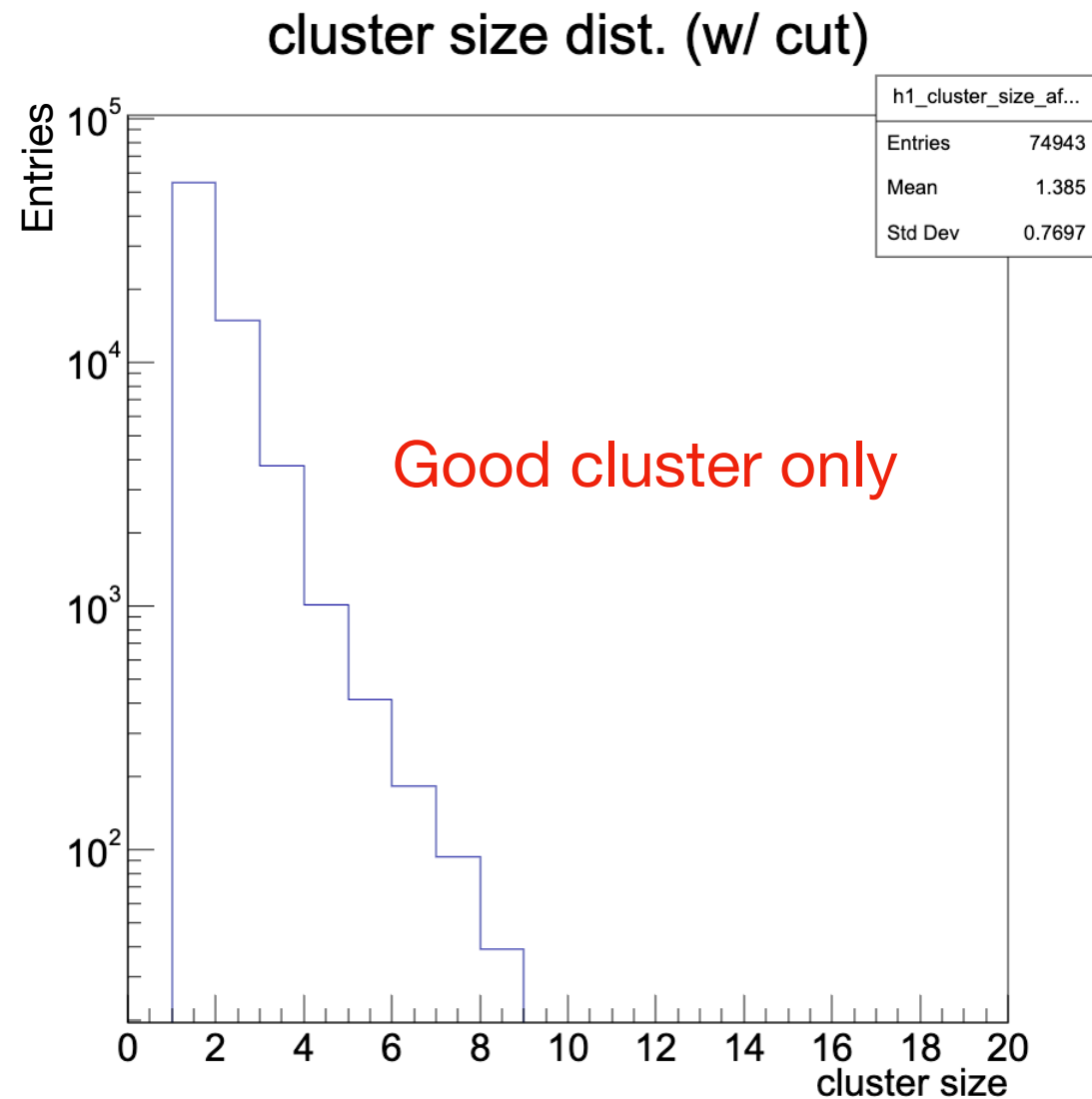
- Without cluster cut:
  - Left: Cluster size distribution
  - Right: Cluster adc distribution





# Run data check

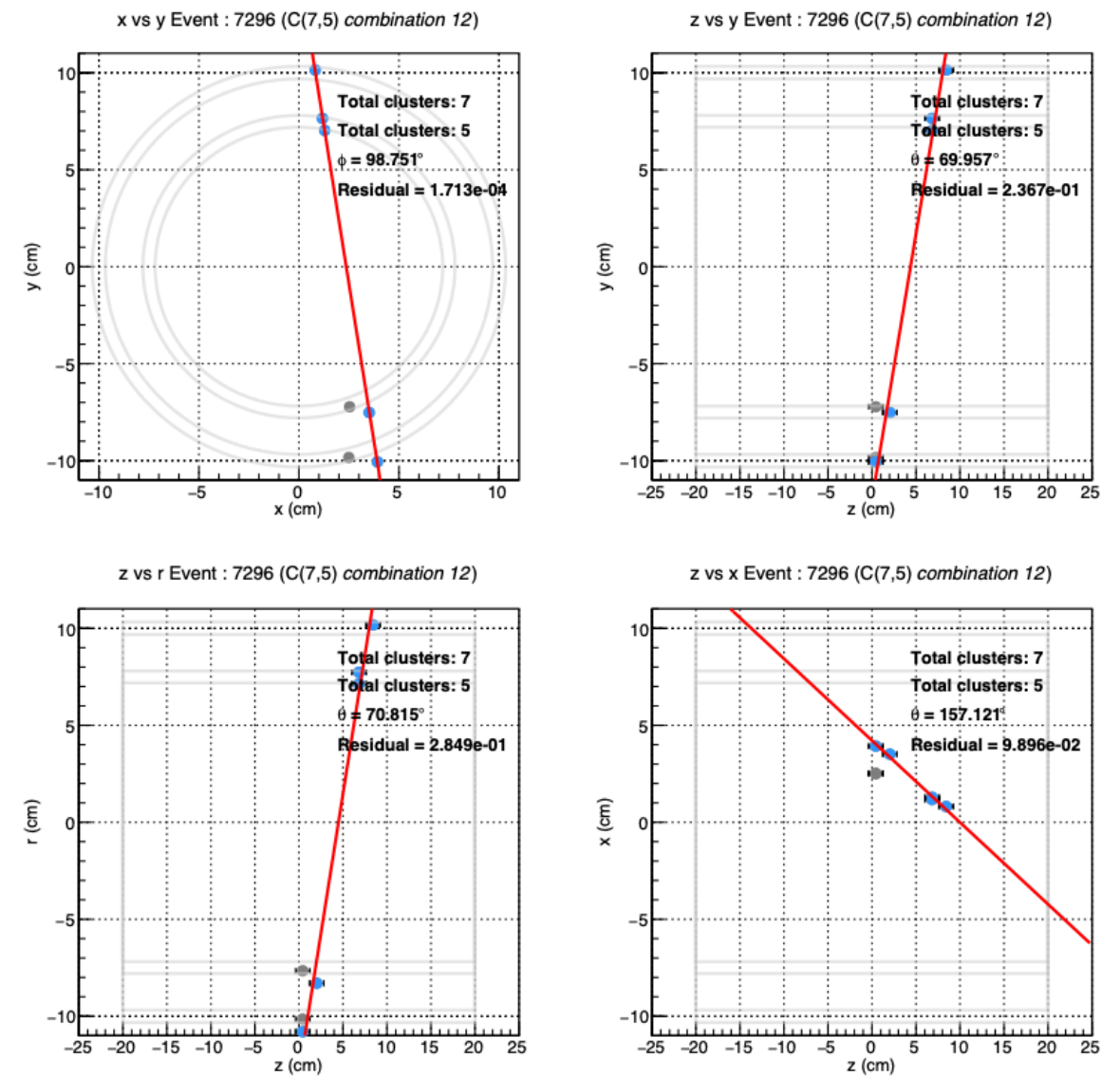
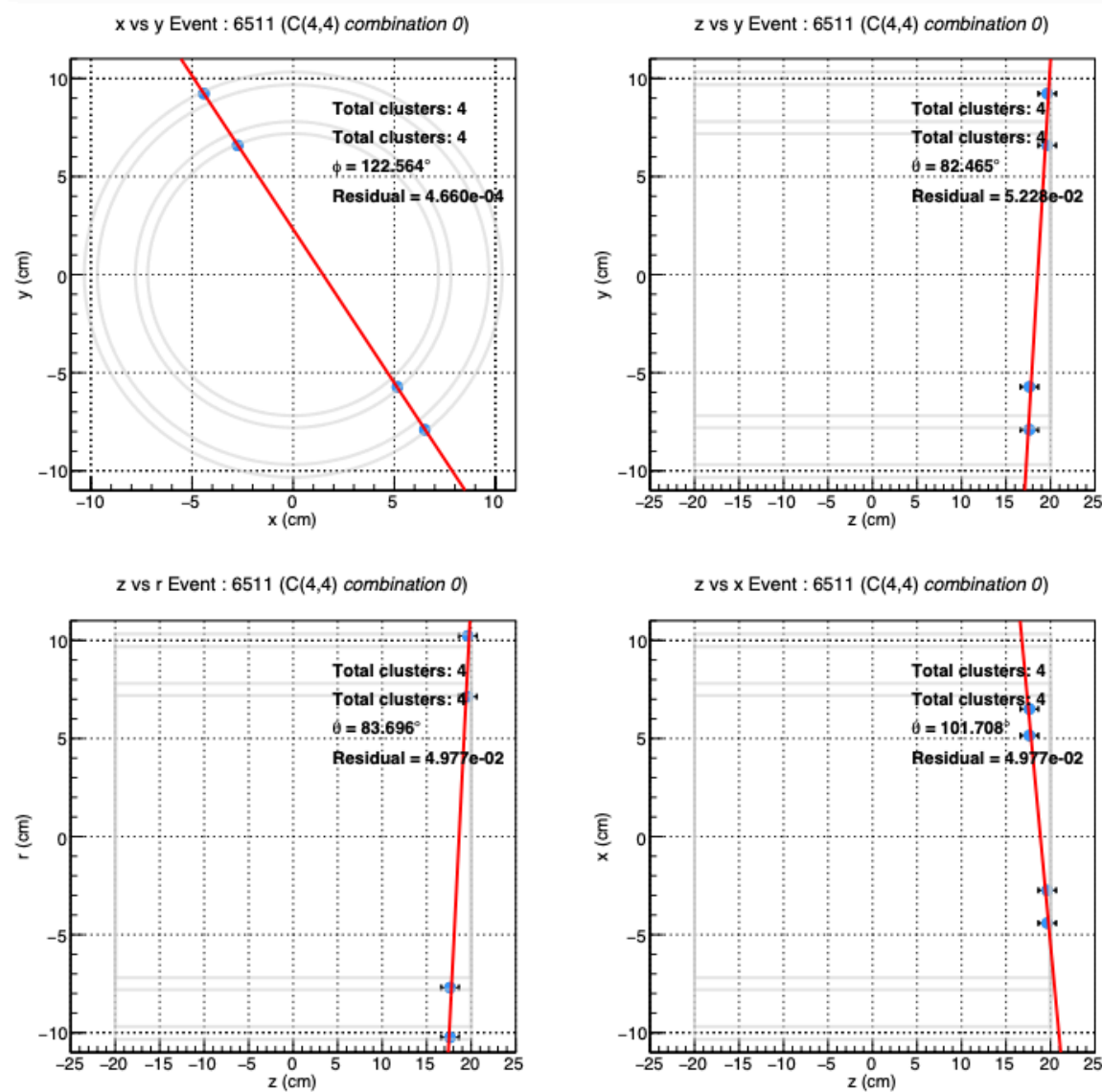
- Pre-selection comparison:
  - Left: Cluster size distribution
  - Right: Cluster adc distribution



# Cosmic tracks event displays

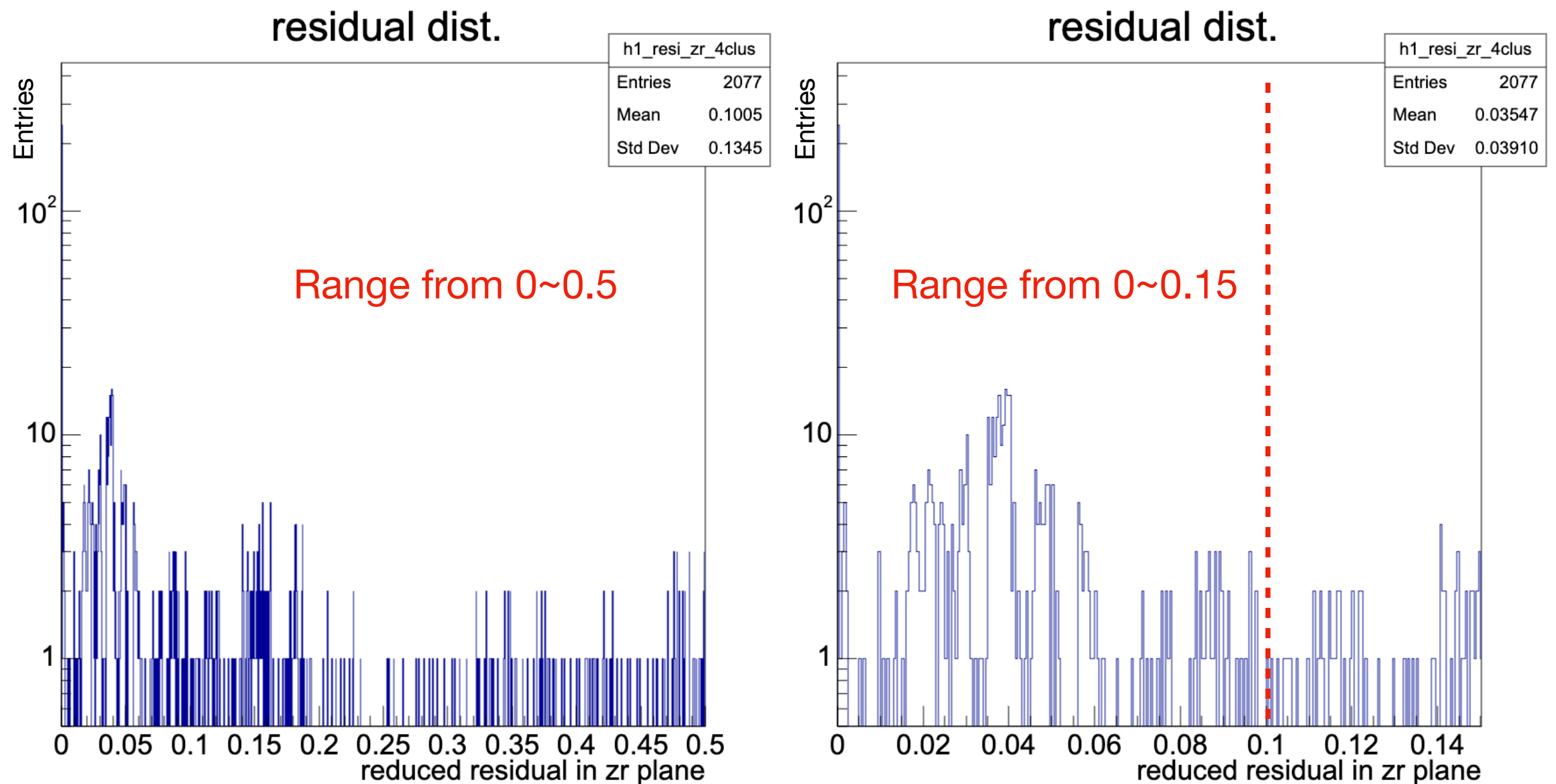
Total clusters = 4, selected clusters = 4  
Residual xy = 0.0046, residual zr = 0.0497

Total clusters = 7, selected clusters = 5  
Residual xy = 0.0017, residual zr = 0.2849



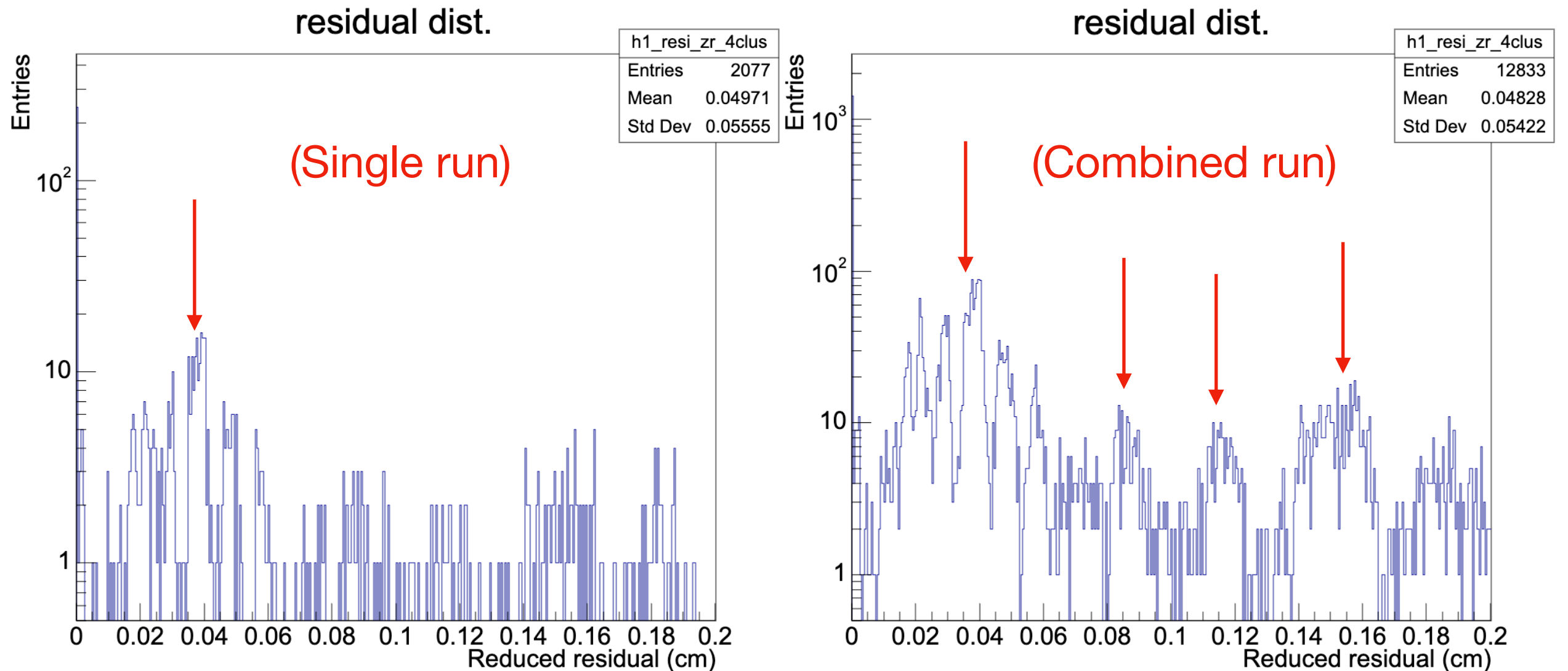
# Fitting quality check

- The reduced residual distribution in **zr plane**.
- Selected the tracks whose reduced residual is smaller than 0.1 as a good track.



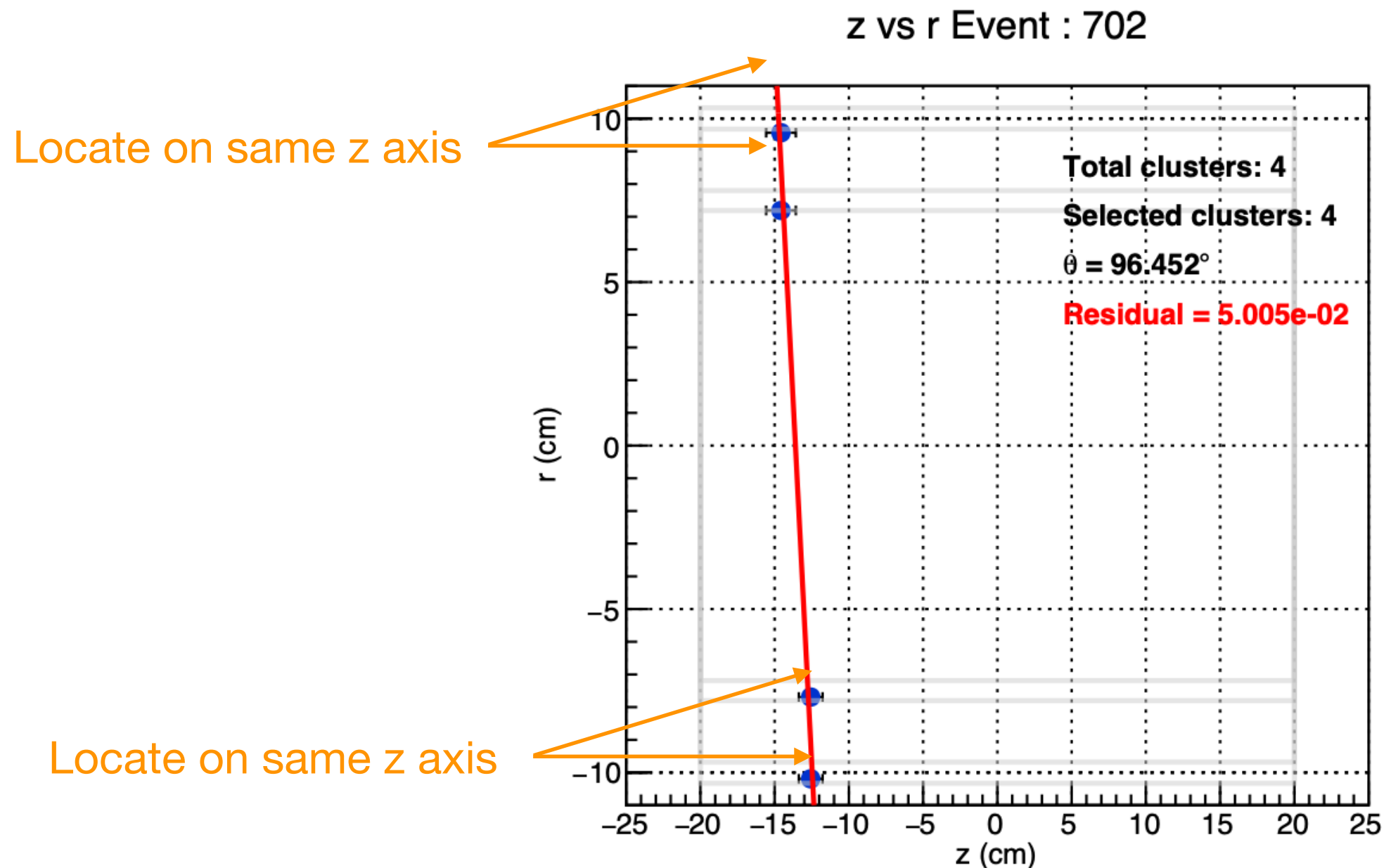
# Fitting quality check (multi runs)

- The reduced residual distribution in **zr plane**.
- More peaks appear in zr residual distribution in combined run.



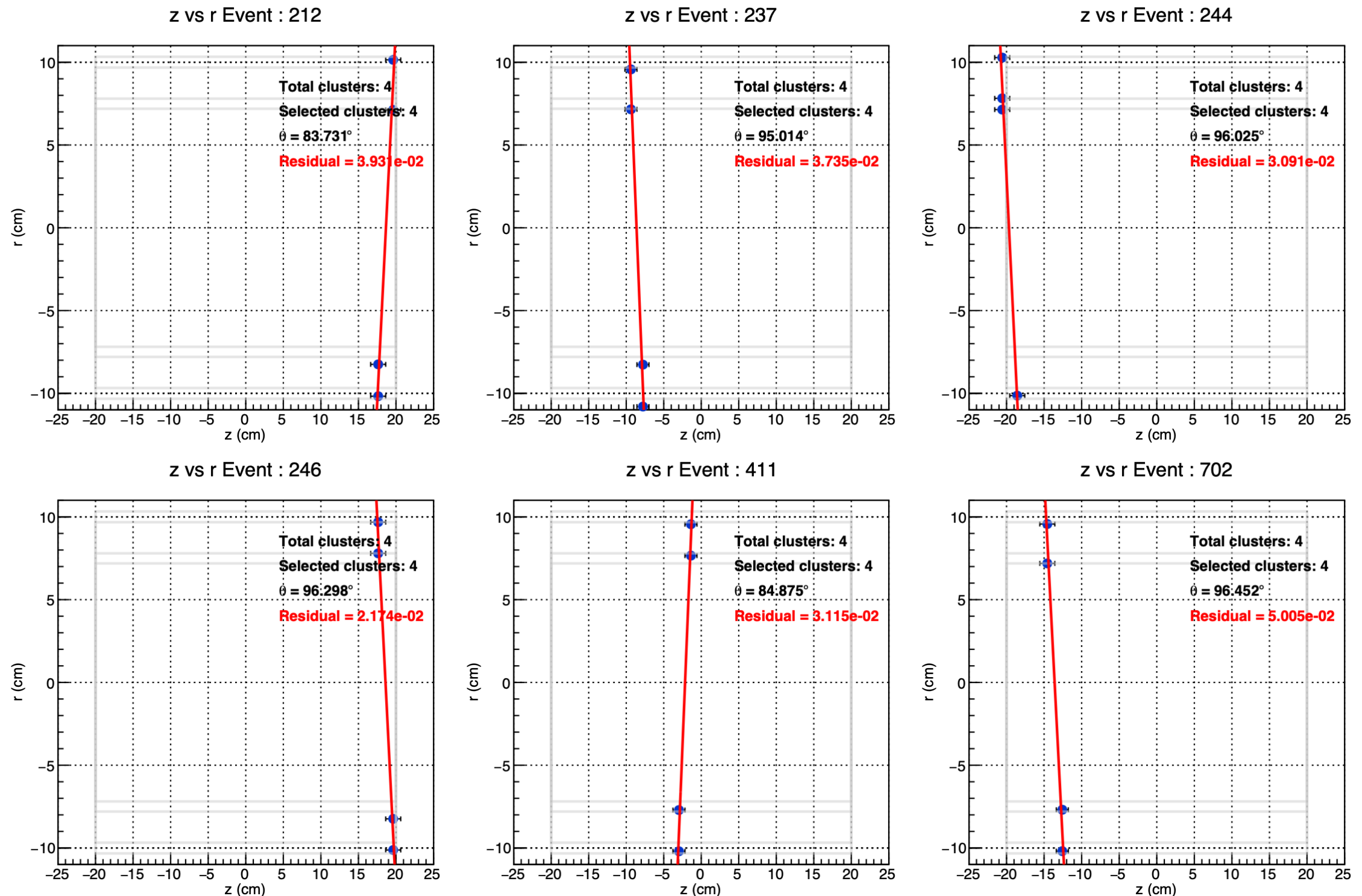
# Run data check

- The event displays show that there is a pattern among those events.
- There is a z axis shift between upper hand clusters and lower hand clusters.
- The angle of these event is fix around  $96^\circ$  and  $84^\circ$ .



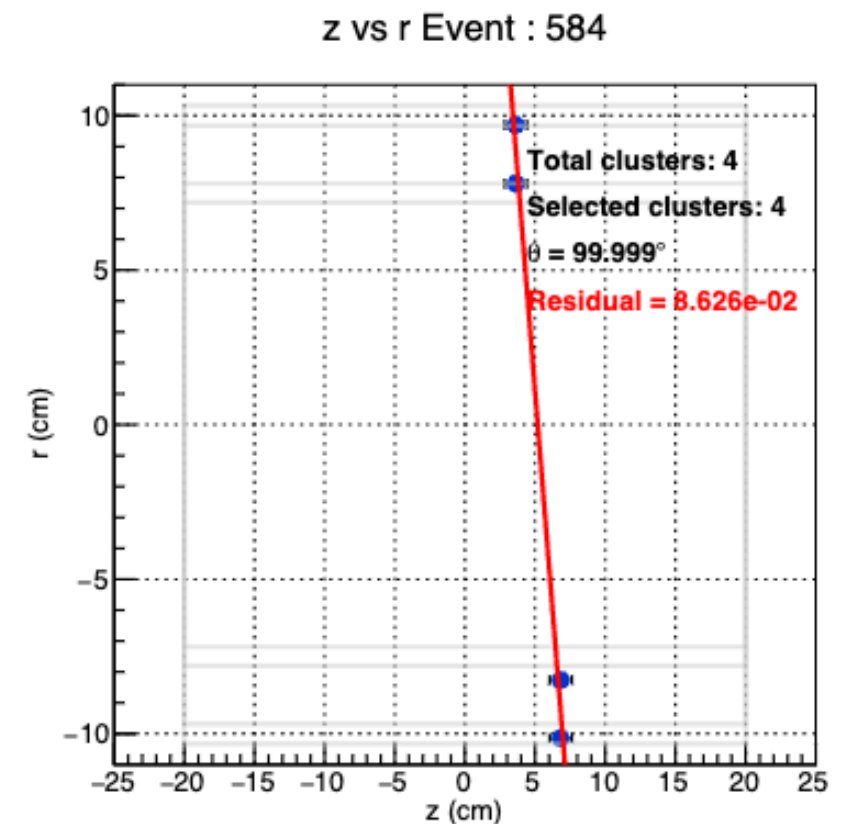
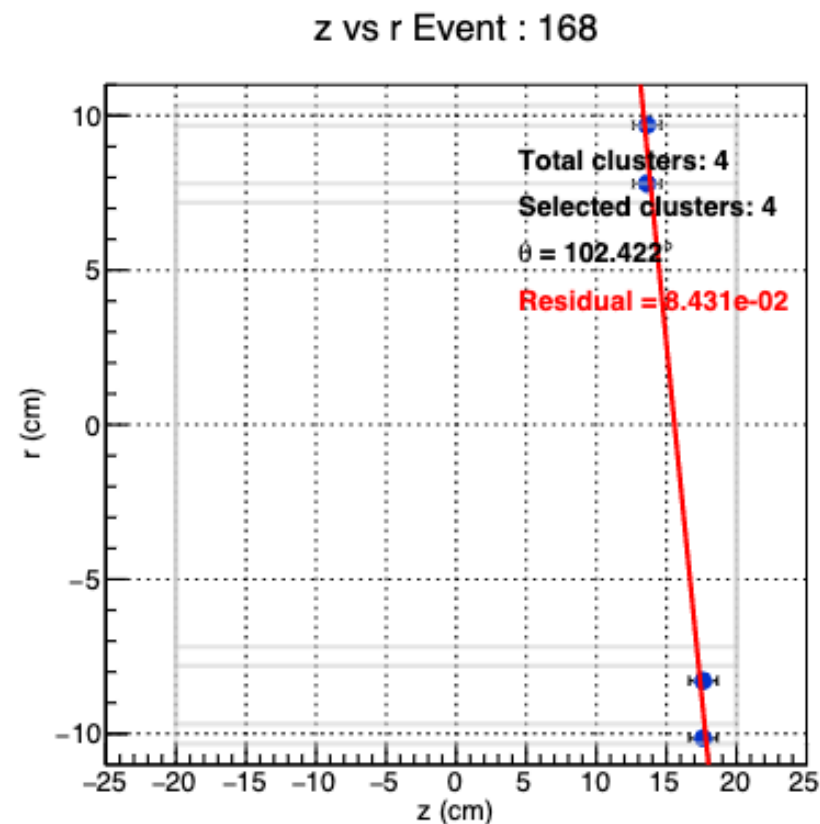
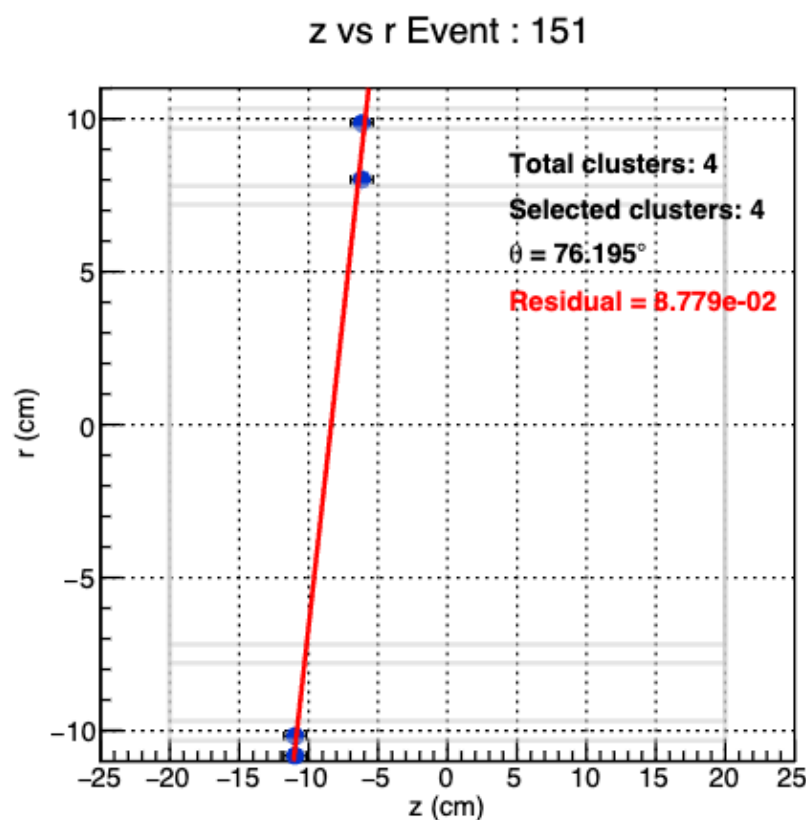
# Event display check

- Take a look of event display where zr residual located at **0.02~0.06**.



# Event display check

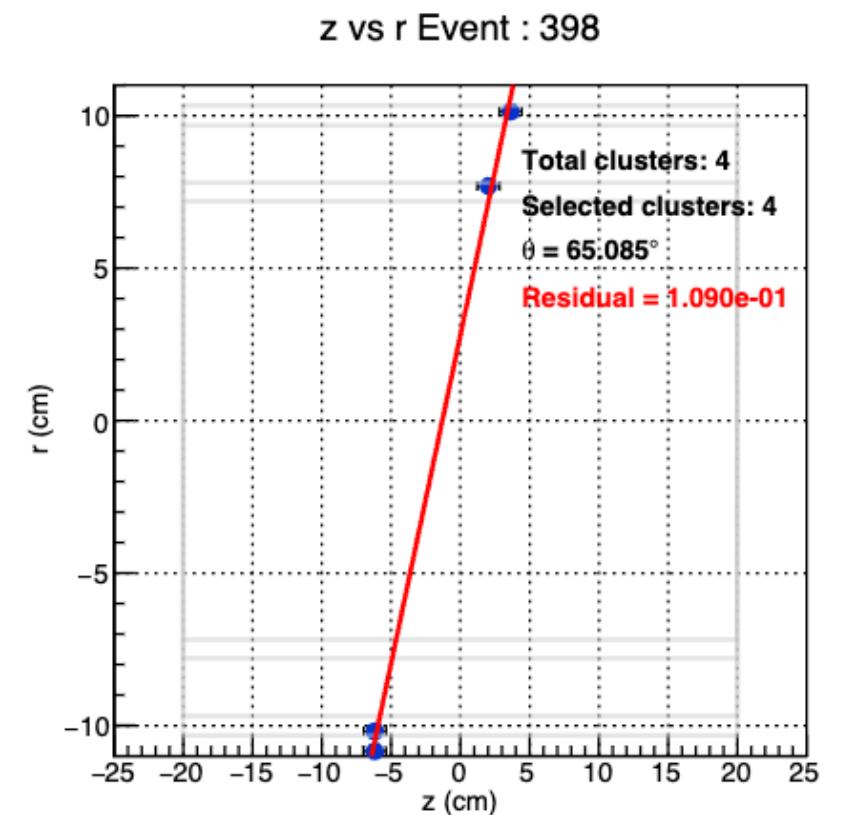
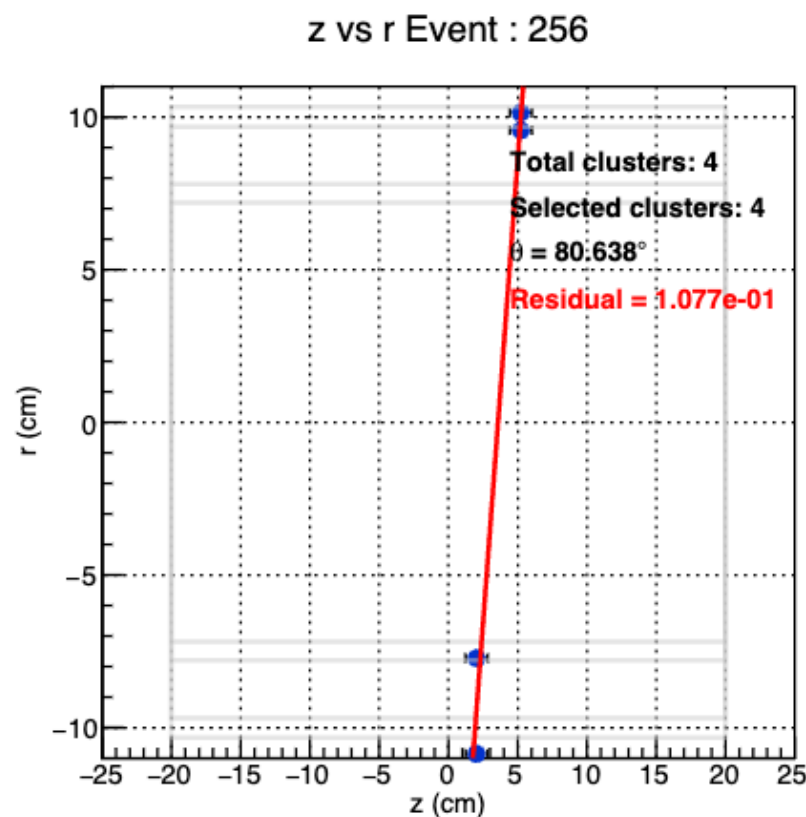
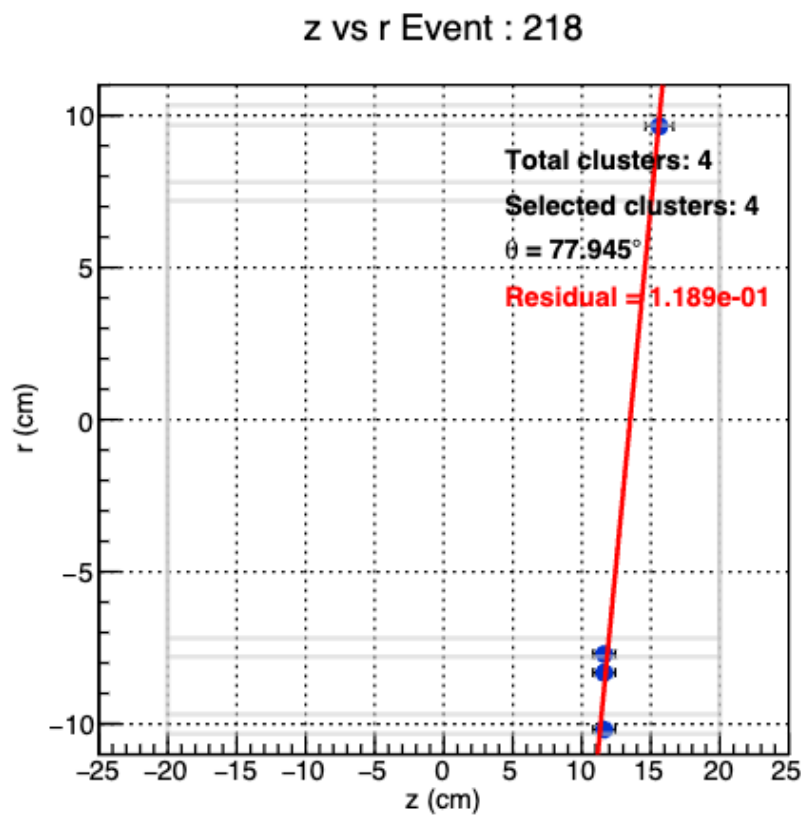
- Another look of event display where zr residual located at 0.08~0.1.
- There is a z axis shift between upper hand clusters and lower hand clusters.  
-> The shift is bit larger comparing to the previous event display.





# Event display check

- Another look of event display where zr residual located at 0.1~0.12.
- There is a z axis shift between upper hand clusters and lower hand clusters.  
-> Although there is larger residual in zr plane, all the clusters are still within the z acceptance.

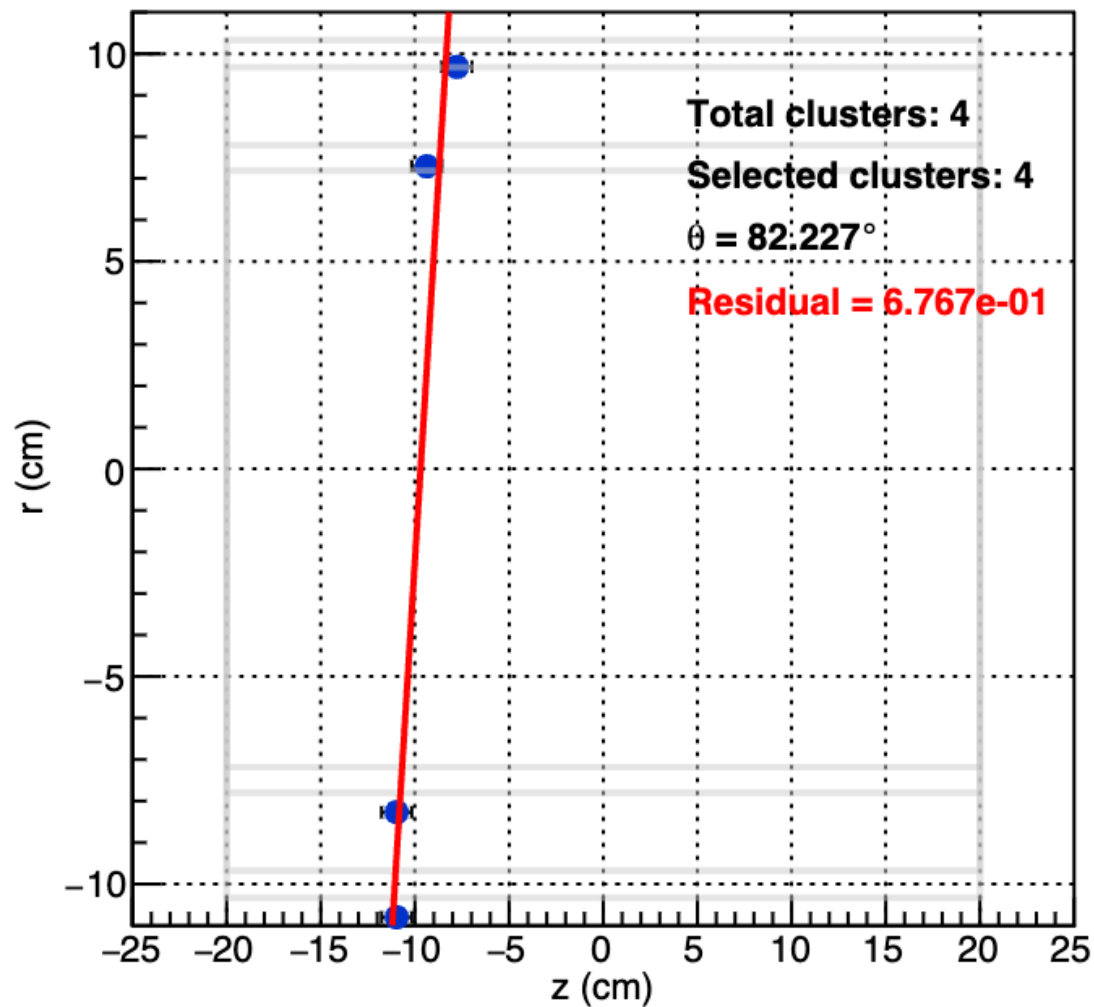




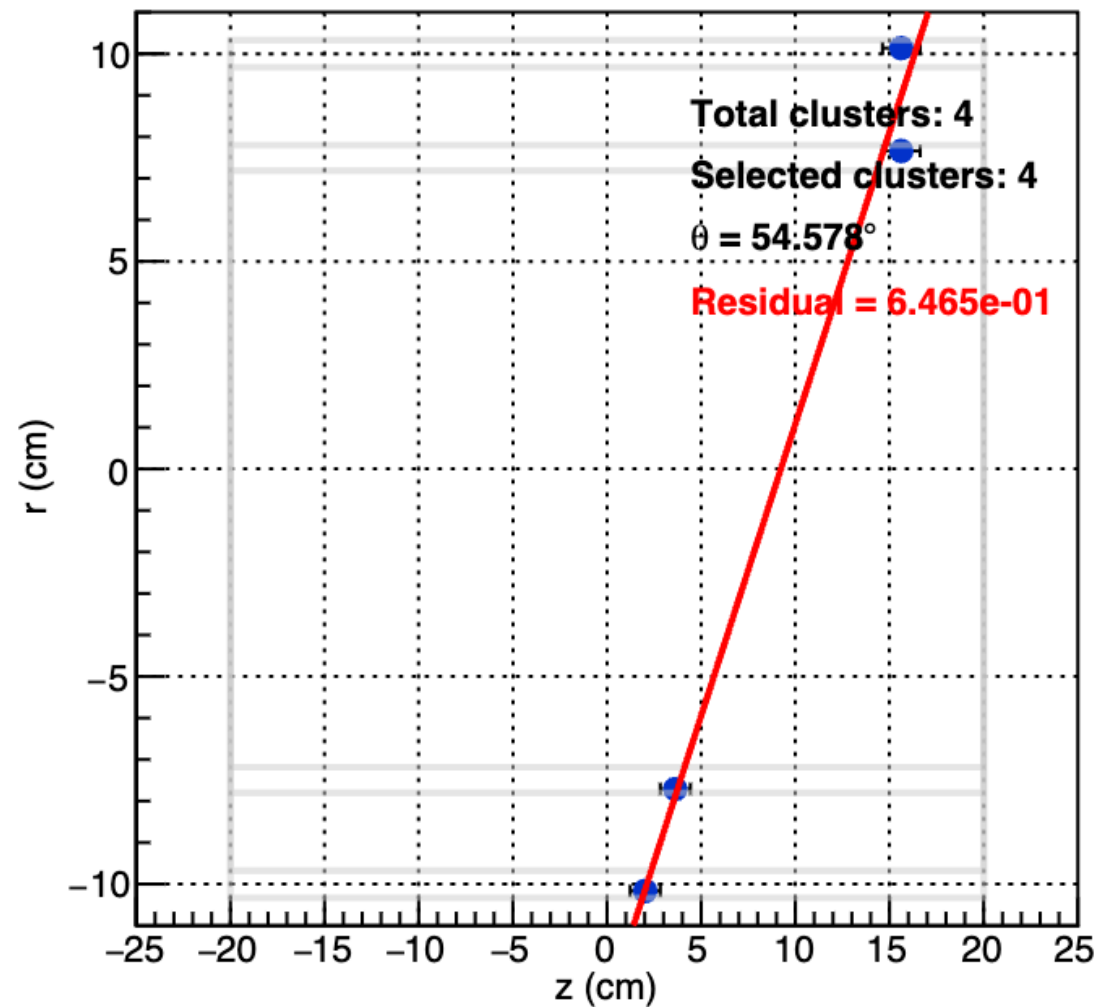
# Event display check

- Another look of event display where zr residual located at 0.5~1.

z vs r Event : 20

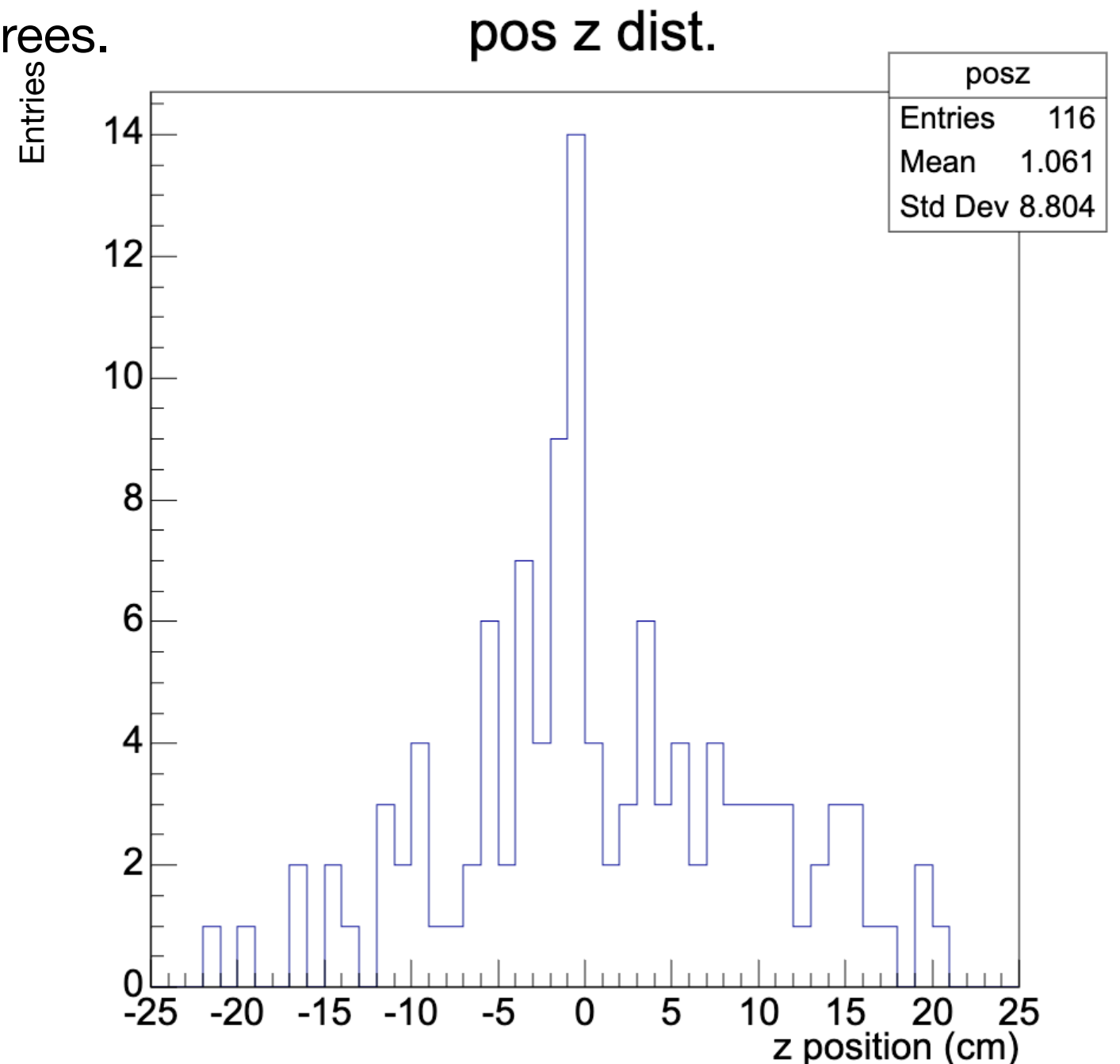
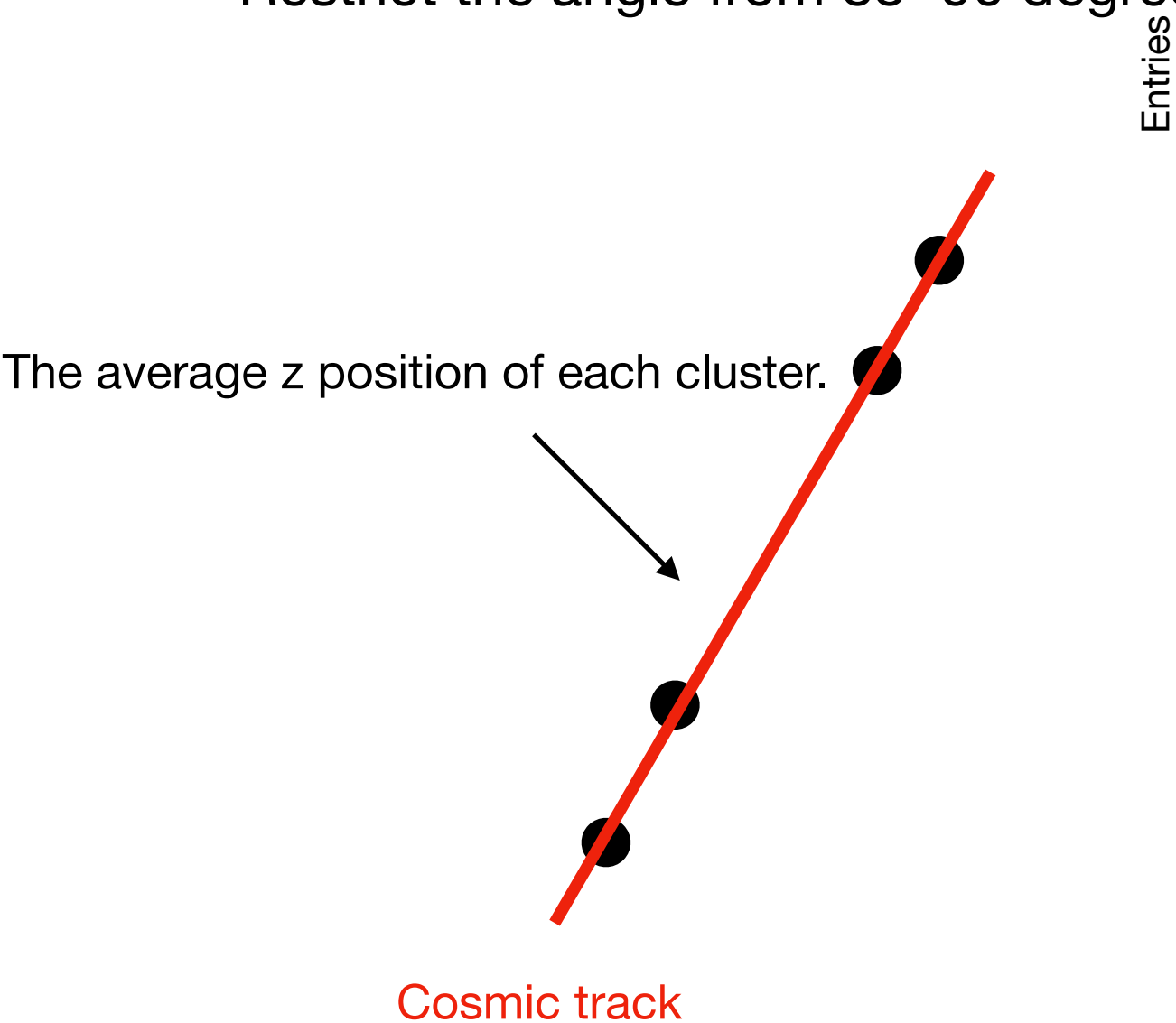


z vs r Event : 570



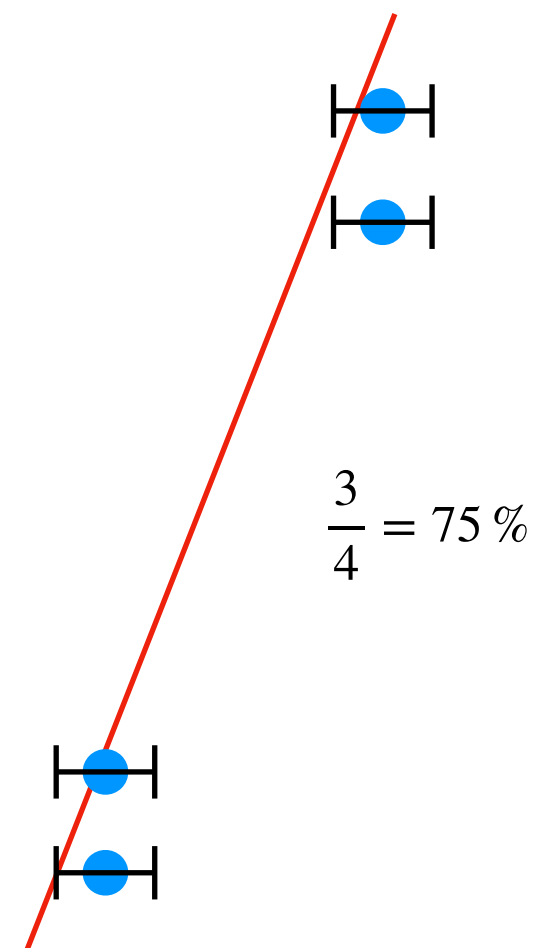
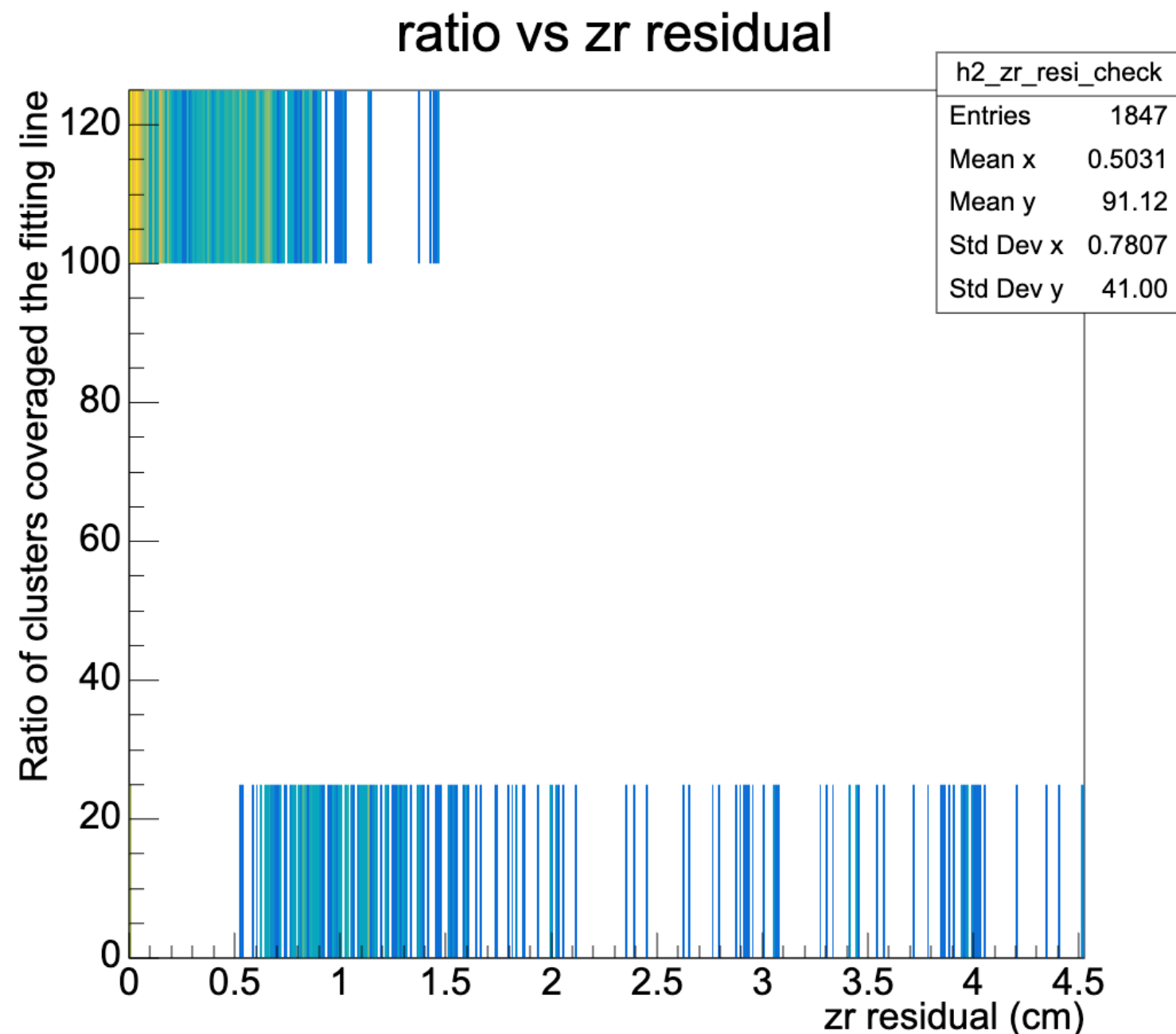
# Z position distribution

- The position distribution along z axis.
- Residual cut applied.
- Restrict the angle from 88~90 degrees.



# Clusters coverage

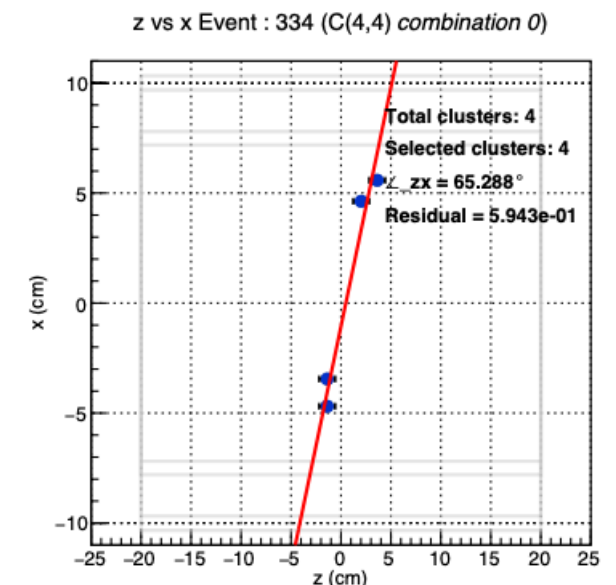
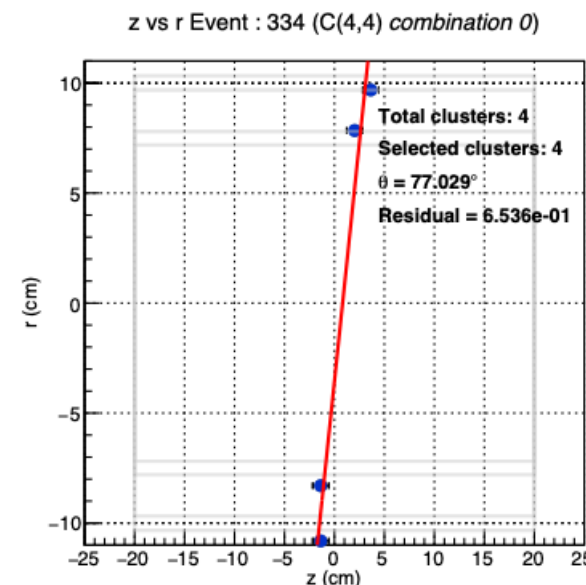
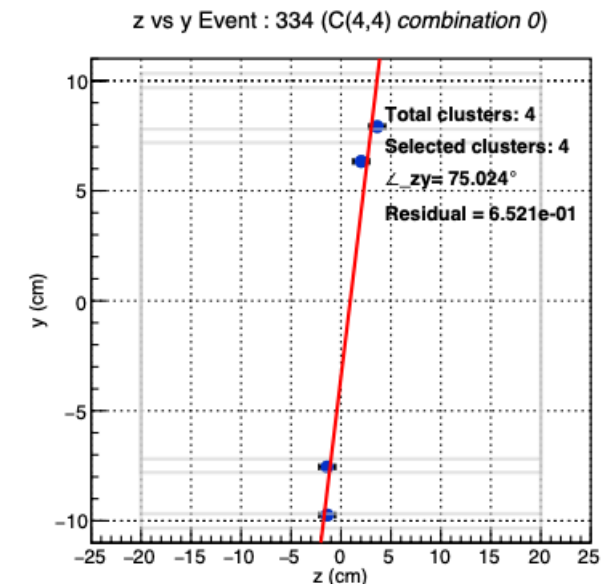
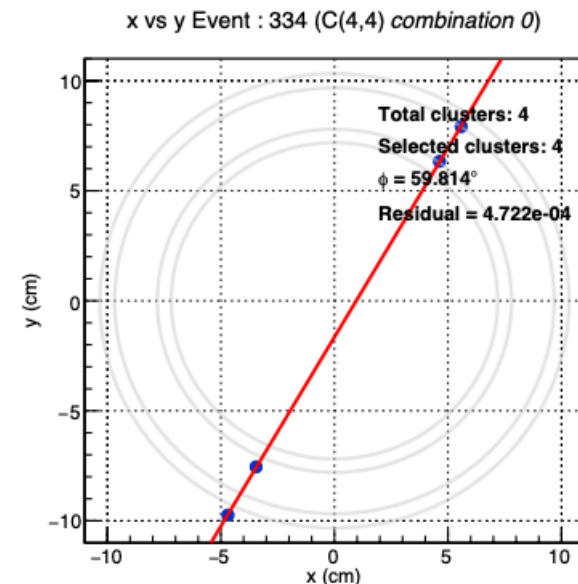
- The x axis is the zr residual.
- The Y axis is ratio of the number of clusters whose coverage is within the fitting line. (Ratio from 0% ~ 100%)



# Cosmic tracks selection

- Only focus on the event whose total good clusters equals to 4.
- Track angle determined by reduced  $\chi^2$  method.
- Use the reduced residual to evaluate the fitting quality (to minimize the angle dependence).

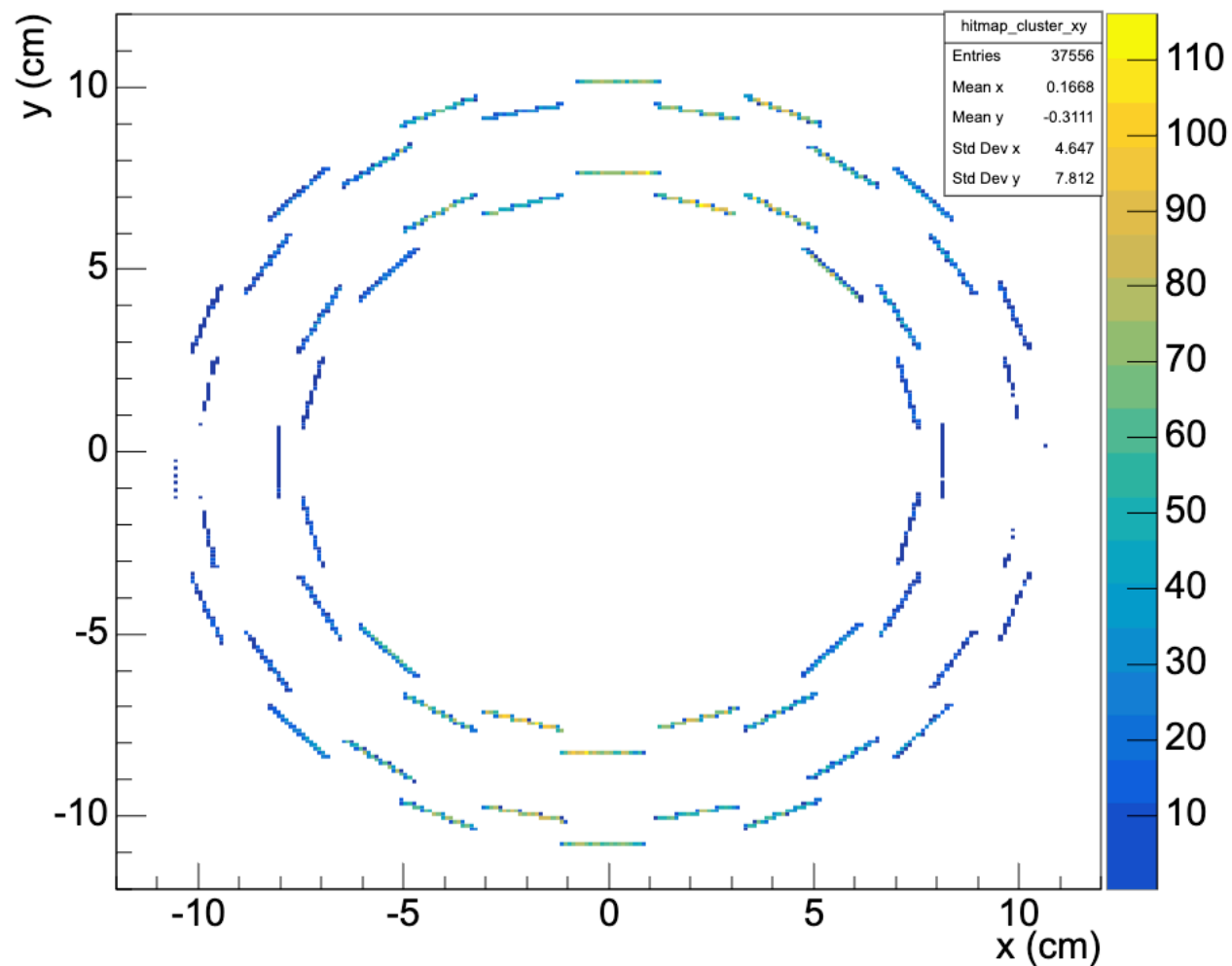
- The event display is shown in xy, zy, zx and zr planes. (The units are in cm)
- Mainly focus on the xy and zr planes in the analysis.



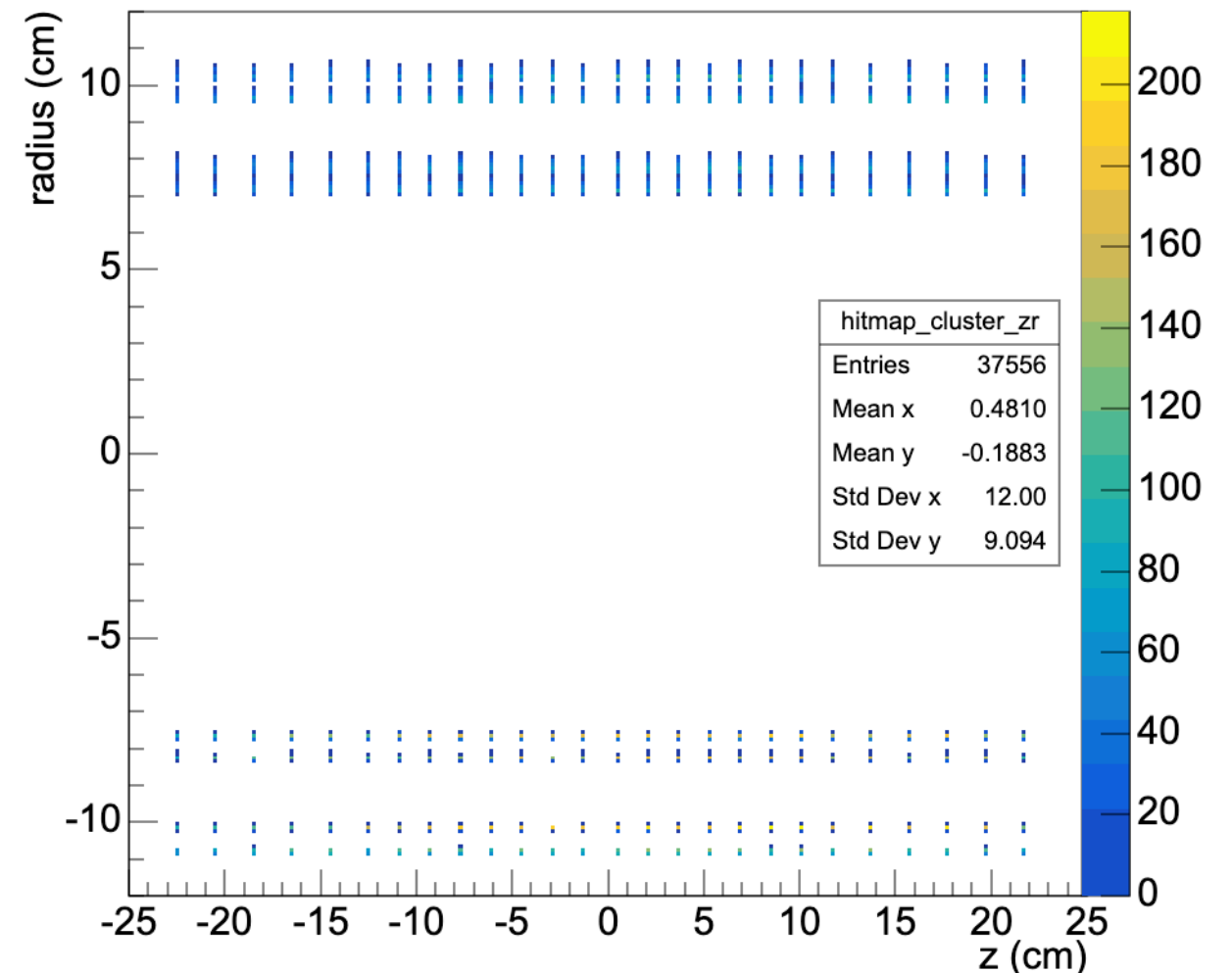
# Clusters Hitmap

- The cluster hitmap in both xy and zr plane
  - There is a hot region on xy plane. → This also implies the track distribution.
  - No hot region on zr plane.

The clusters hitmap of xy plane

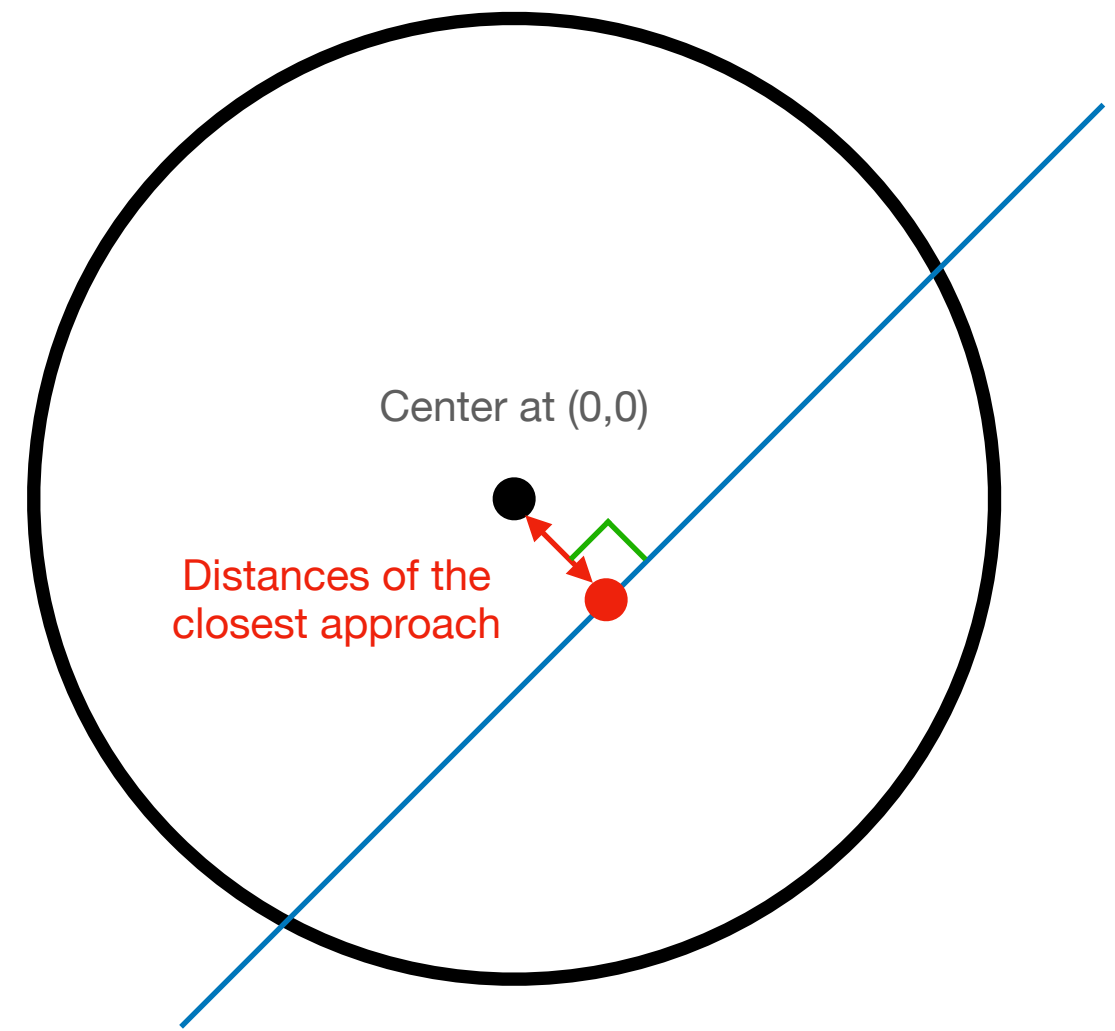
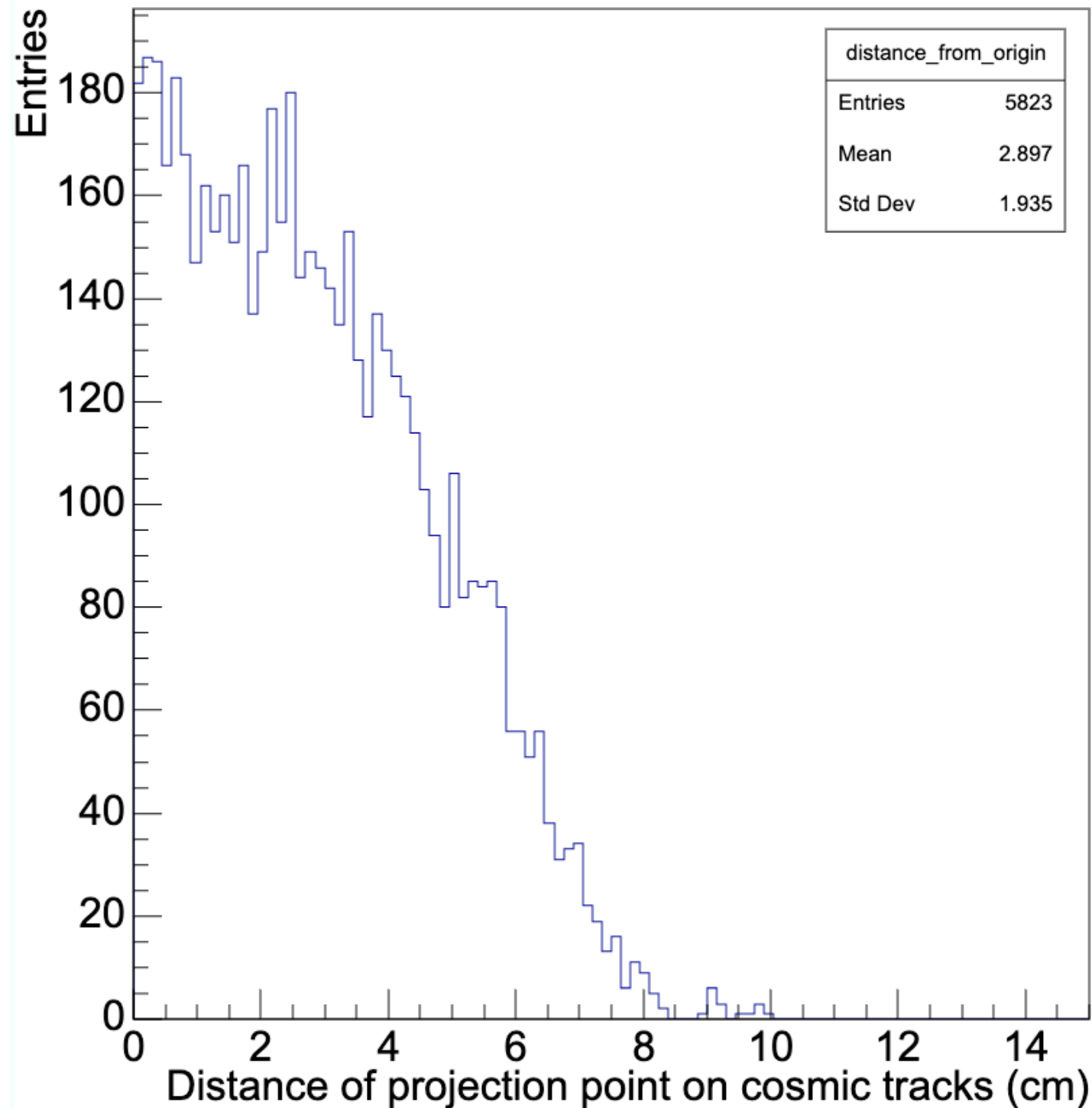


The clusters hitmap of zr plane

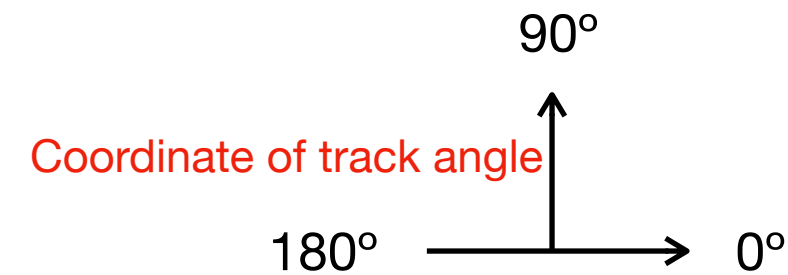


# The distribution of the projection point of the origin on the cosmic tracks

The distance from the origin

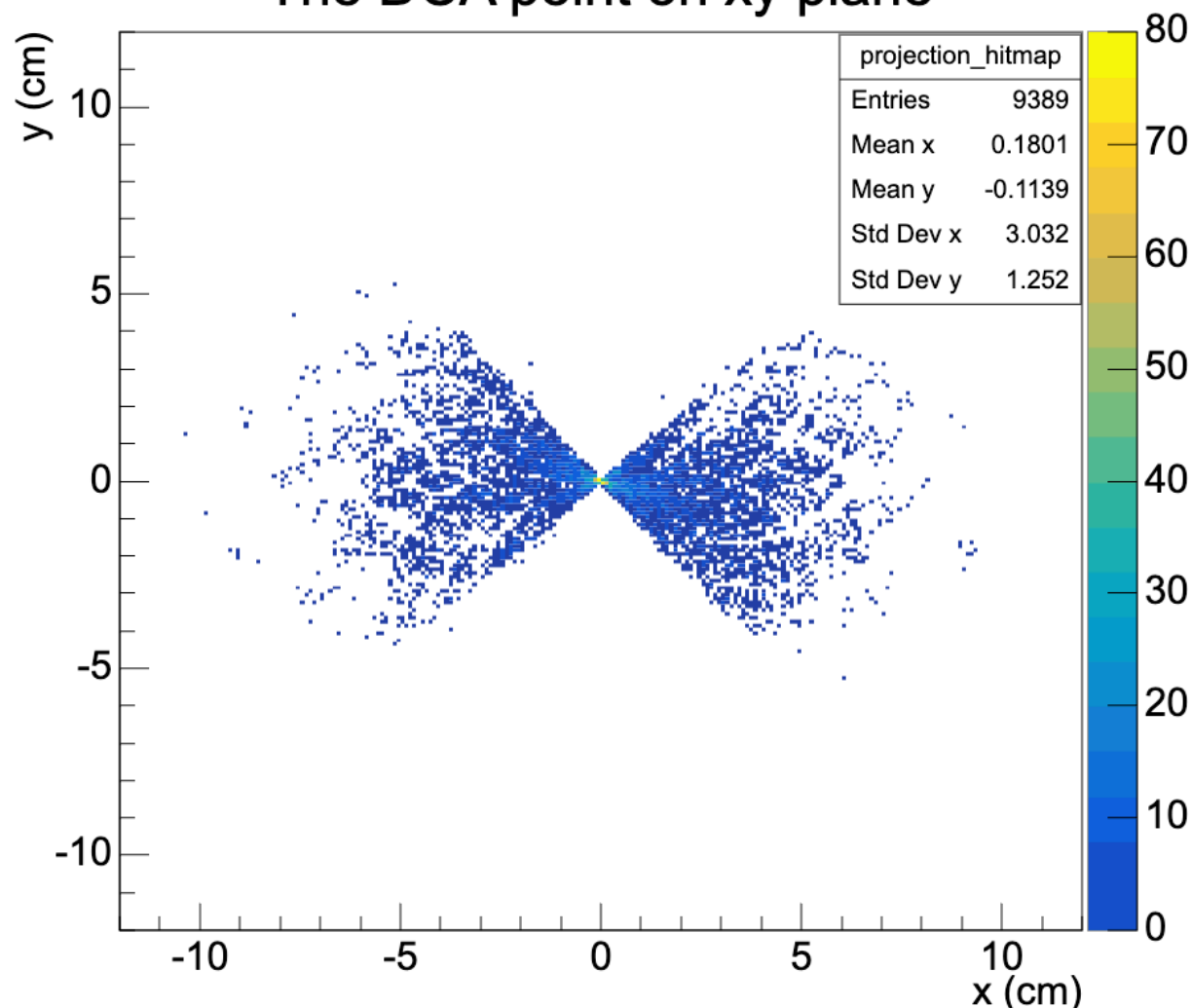


# DCA point



- The right figure shows the correlation between DCA and tracking angle.
  - Also shows the acceptance angle.
  - The more track angle close to 90 degrees, the larger DCA range can have.
- Expected geometry. Those plots can check whether the trigger conditions (high correlation with track angle acceptance) have been changed or not.

The DCA point on xy plane



Correlation b/w DCA and track angle

