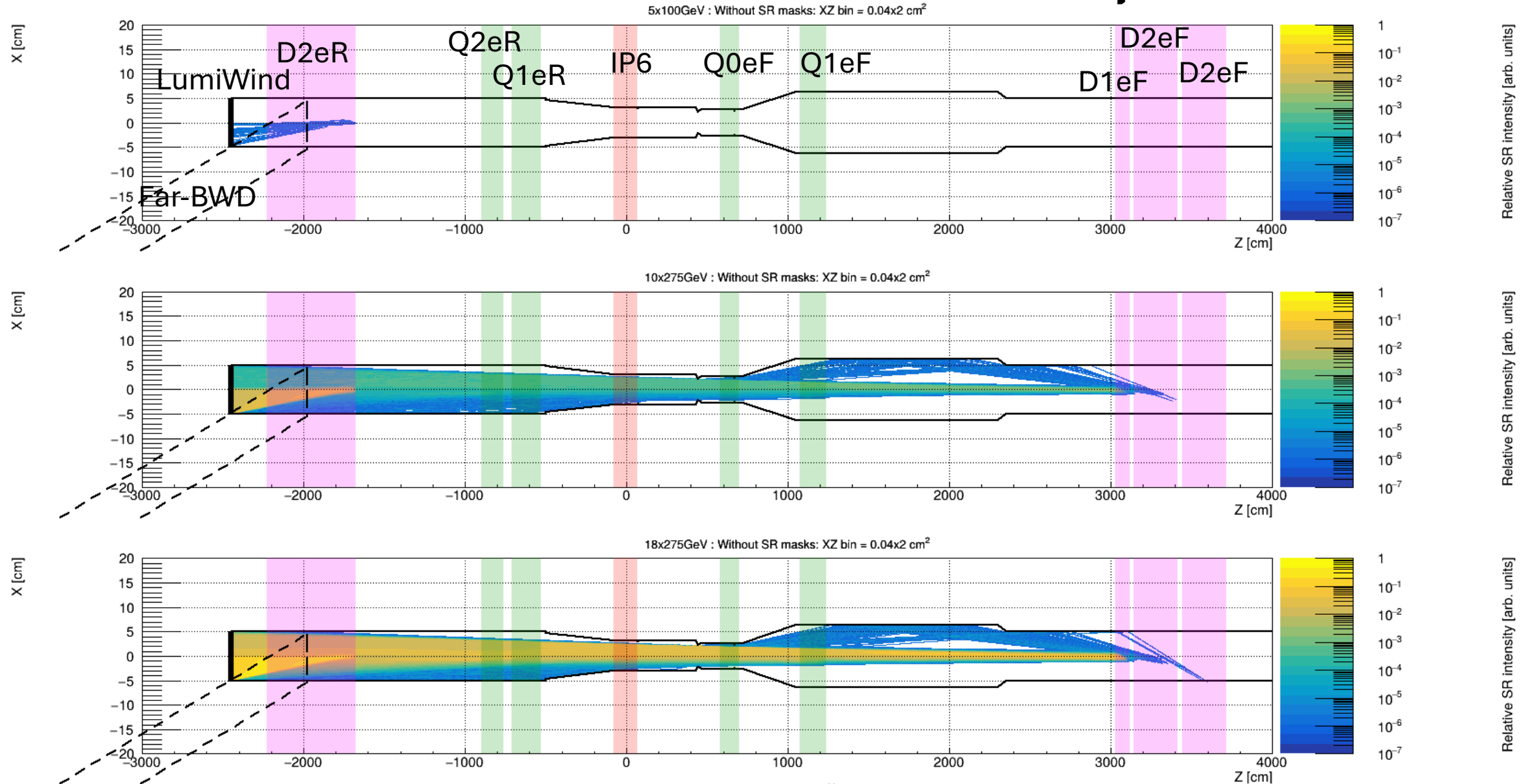


# Synchrotron Radiation on the Luminosity Exit Window

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[natochii@bnl.gov](mailto:natochii@bnl.gov)

**Acknowledgements:** Many thanks to D. Marx, C. Hetzel, and S. Kay for provided materials and discussions.

# SR Tracks Hitting the Lumi Window ( $E_\gamma > 10$ keV)



# SR Hitting the Lumi Window ( $E_\gamma > 30$ eV)

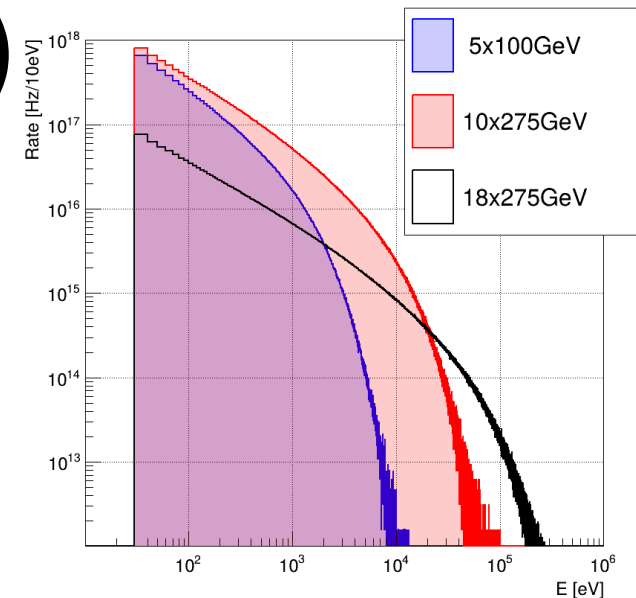
Extremely high density

SR power: Critical photon energy:

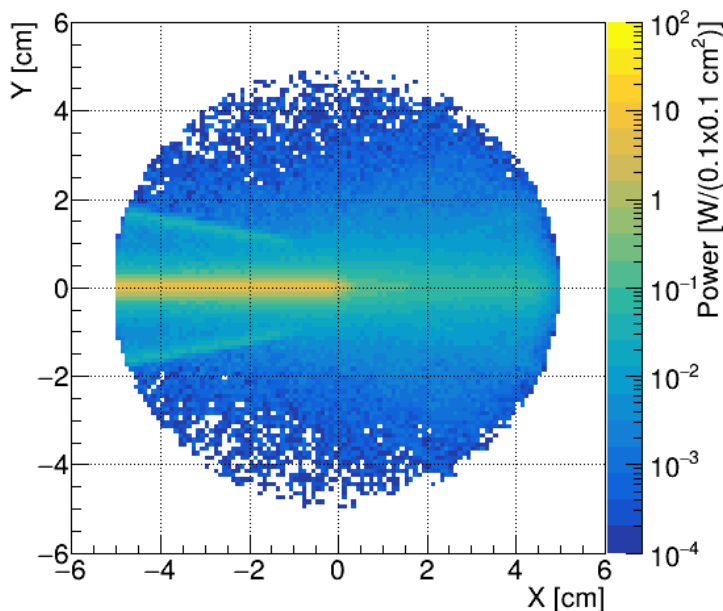
$$P_\gamma \sim \frac{\gamma^4}{\rho^2}$$

$$\varepsilon_c \sim \frac{\gamma^3}{\rho}$$

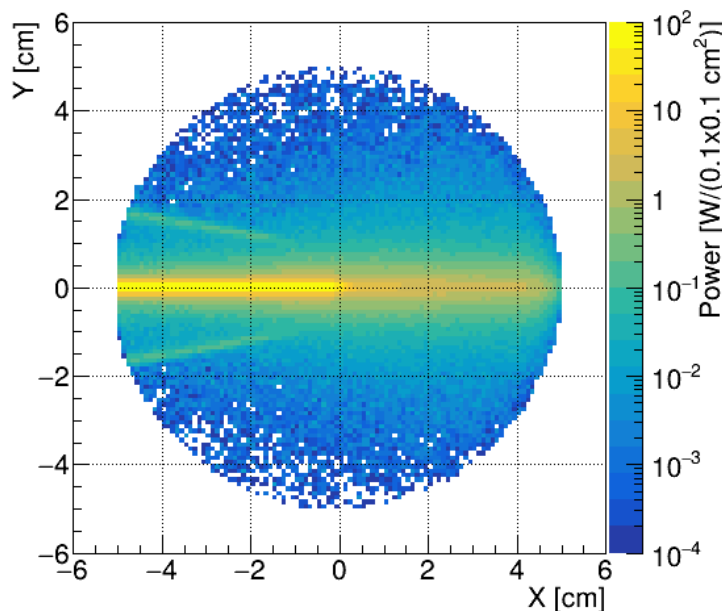
5 GeV @2.500A :  $P_{\max} = 5.9$  [W/mm<sup>2</sup>];  $P_{\text{tot}} = 676.0$  [W]  
 10 GeV @2.500A :  $P_{\max} = 93.5$  [W/mm<sup>2</sup>];  $P_{\text{tot}} = 9561.9$  [W]  
 18 GeV @0.227A :  $P_{\max} = 97.8$  [W/mm<sup>2</sup>];  $P_{\text{tot}} = 9605.8$  [W]



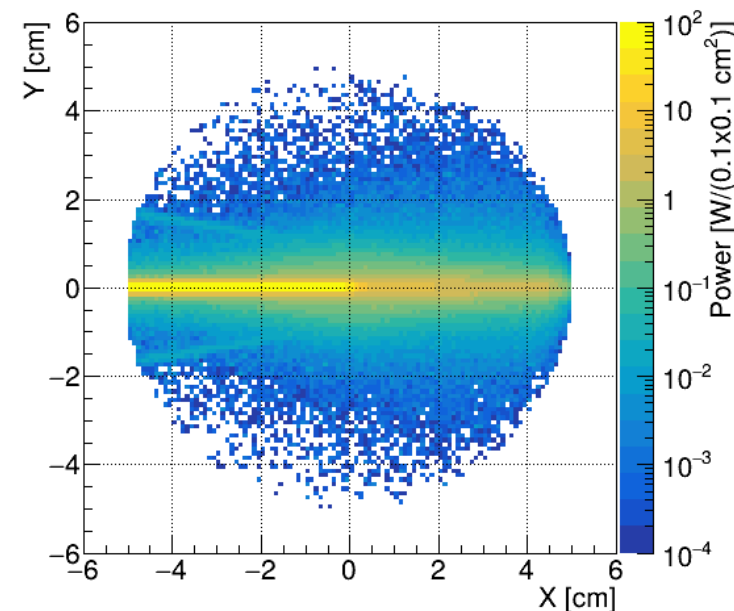
5x100GeV : Without SR masks



10x275GeV : Without SR masks



18x275GeV : Without SR masks



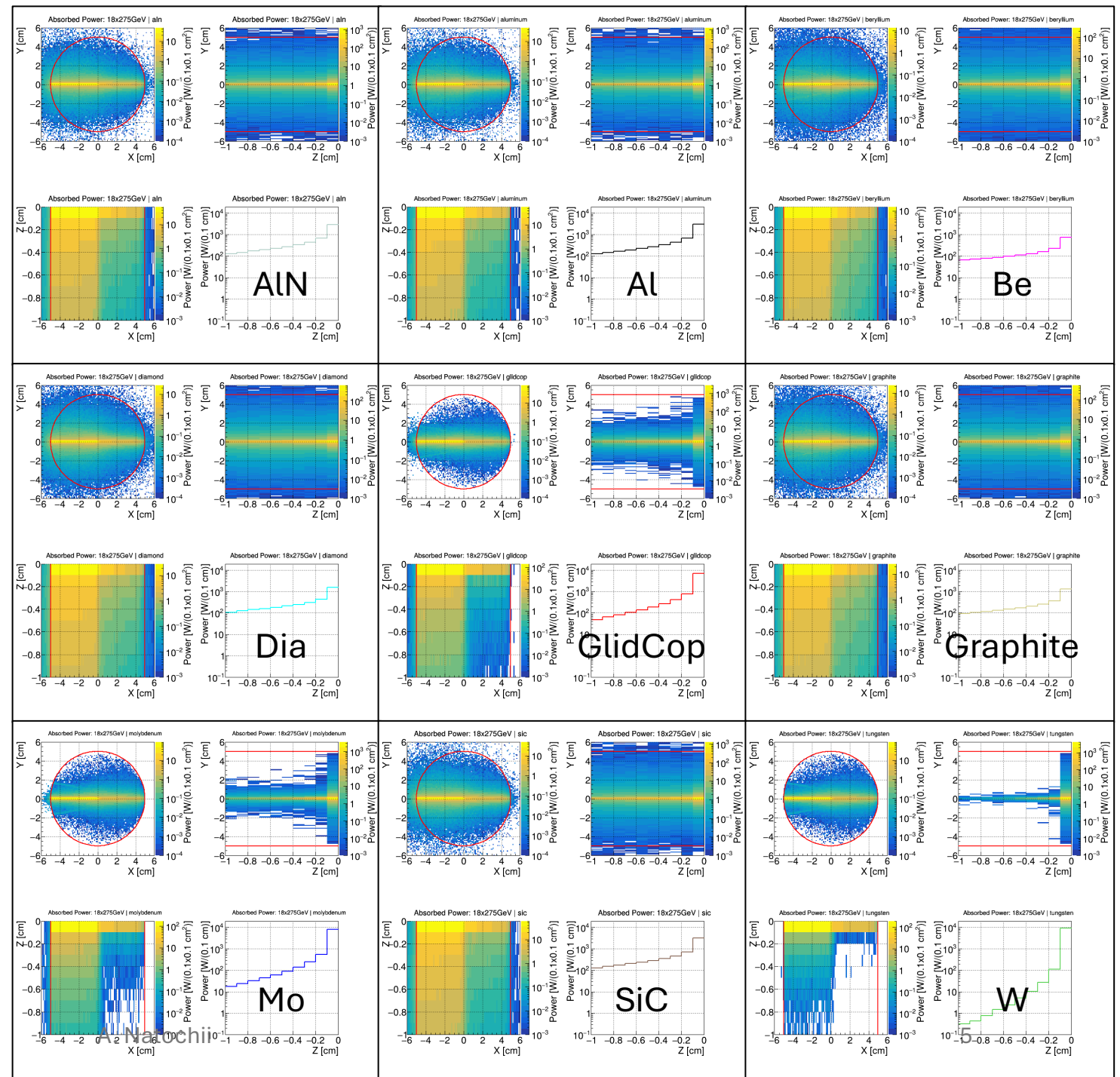
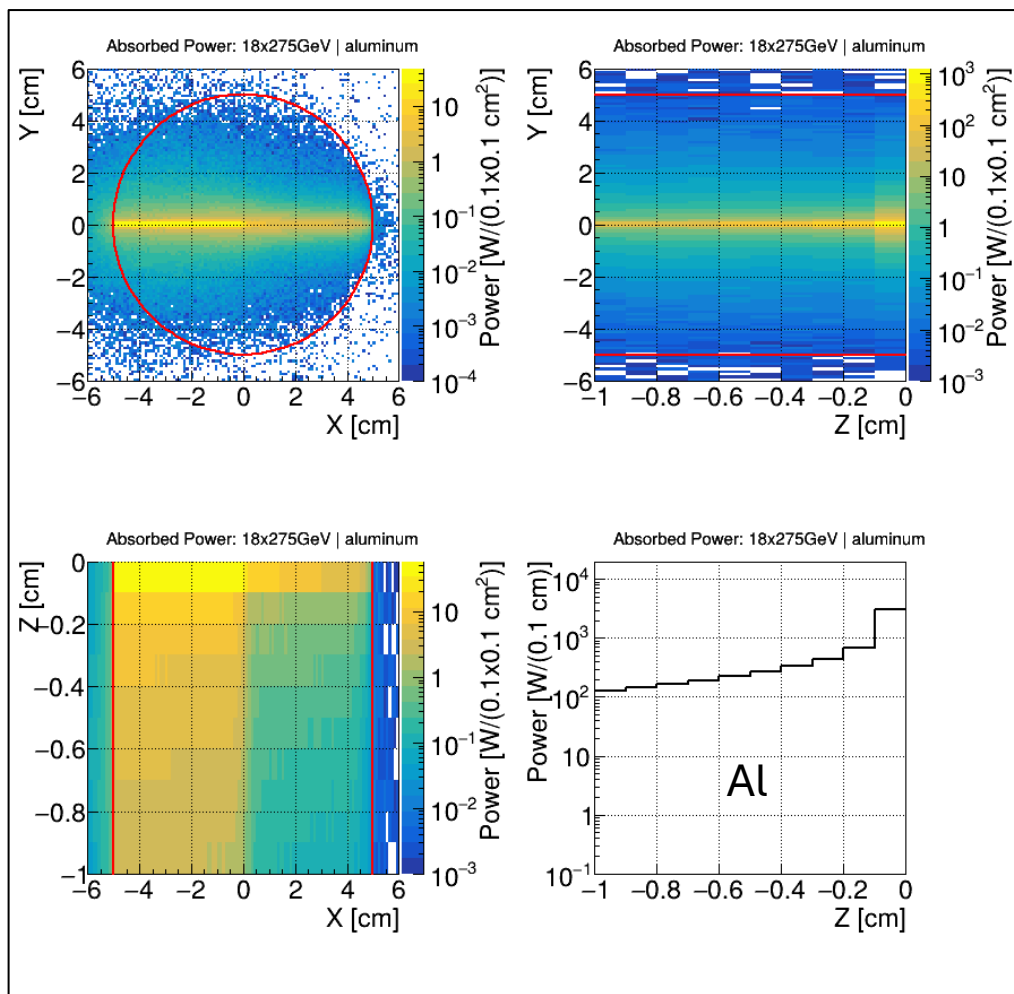
# Lumi Window Material Candidates

1. **Copper (Cu) & Alloys:** High thermal conductivity, melting point  $\sim 1,085^{\circ}\text{C}$ . Alloys like GlidCop enhance strength.
2. **GlidCop:** Copper alloy with aluminum ( $\text{Al}_2\text{O}_3$ ), high thermal conductivity, and improved resistance to high-temperature embrittlement.
3. **Tungsten (W):** Melting point  $\sim 3,422^{\circ}\text{C}$ , good heat resistance, but lower thermal conductivity than copper.
4. **Molybdenum (Mo):** High melting point ( $\sim 2,623^{\circ}\text{C}$ ), decent thermal conductivity, more brittle than Cu or W.
5. **Graphite:** Very high sublimation point ( $\sim 3,600^{\circ}\text{C}$ ), used in SR beam dumps, variable thermal conductivity.
6. **Beryllium (Be):** High thermal conductivity, melting point  $\sim 1,287^{\circ}\text{C}$ , lightweight, but toxic.
7. **Diamond:** Extremely high thermal conductivity, used as a coating to enhance heat dissipation.
8. **Ceramics (SiC, AlN):** High thermal stability, good thermal conductivity, used for thermal management.

The material choice also depends on factors such as cooling system design, mechanical load, and the duration of exposure.

# Absorbed SR Power

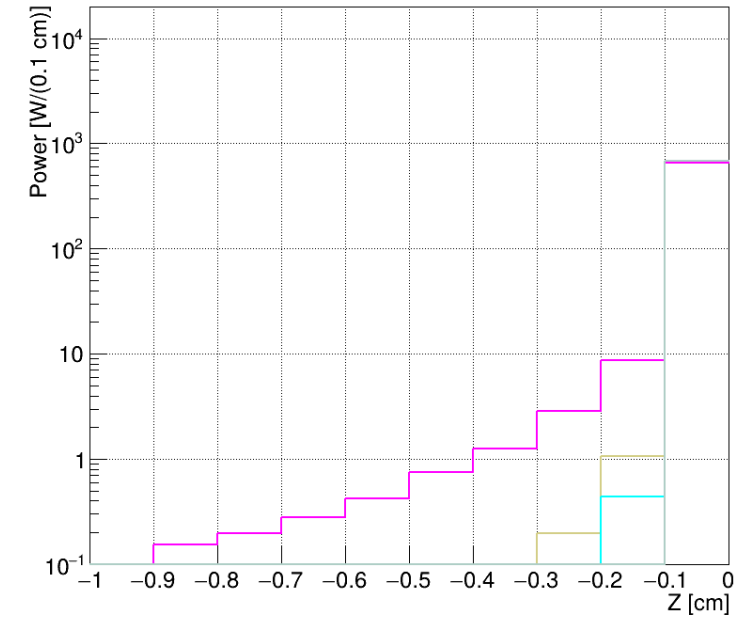
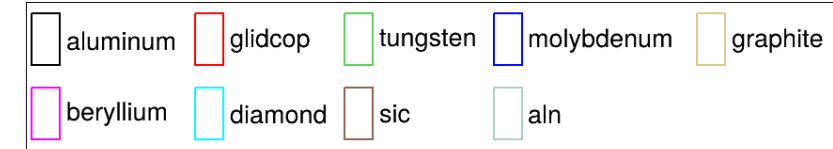
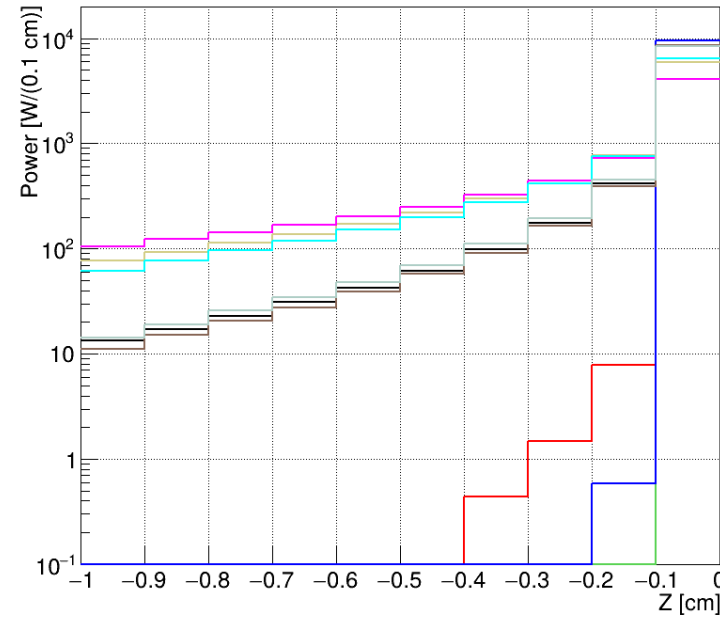
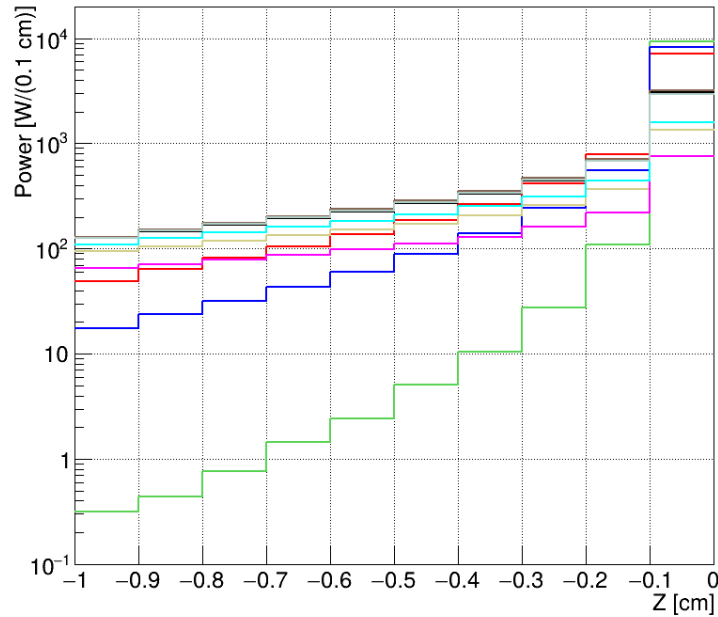
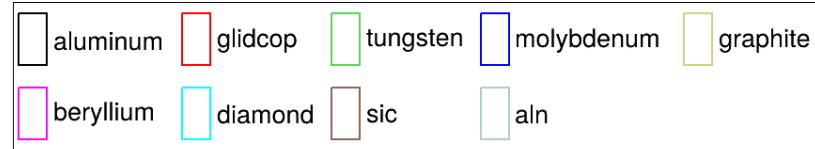
LumiWindow is  $12 \times 12 \text{ cm}^2$



## 18x275 GeV @ 0.227 A

March 20, 2025

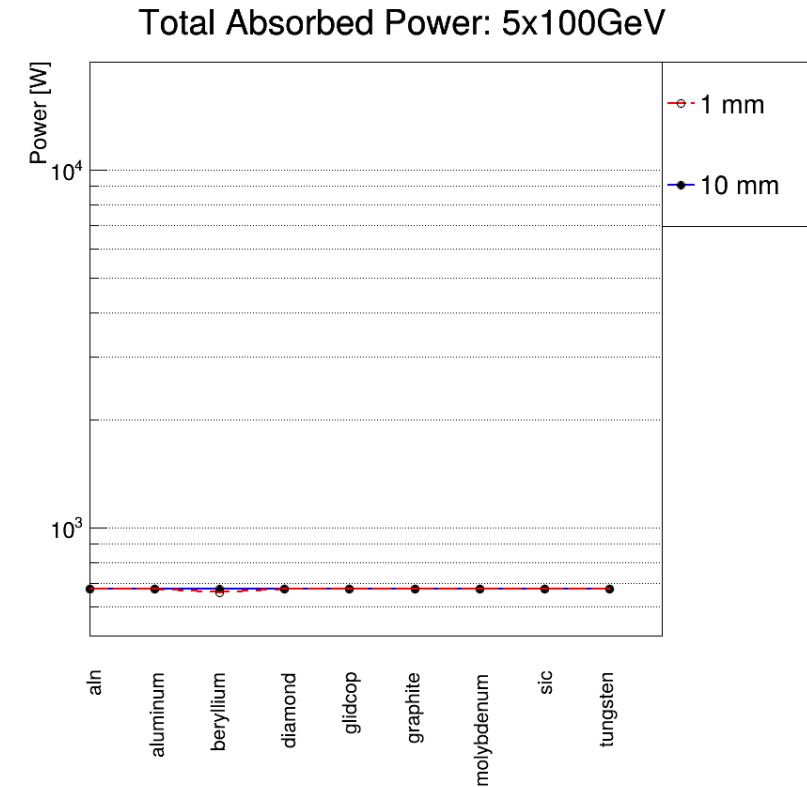
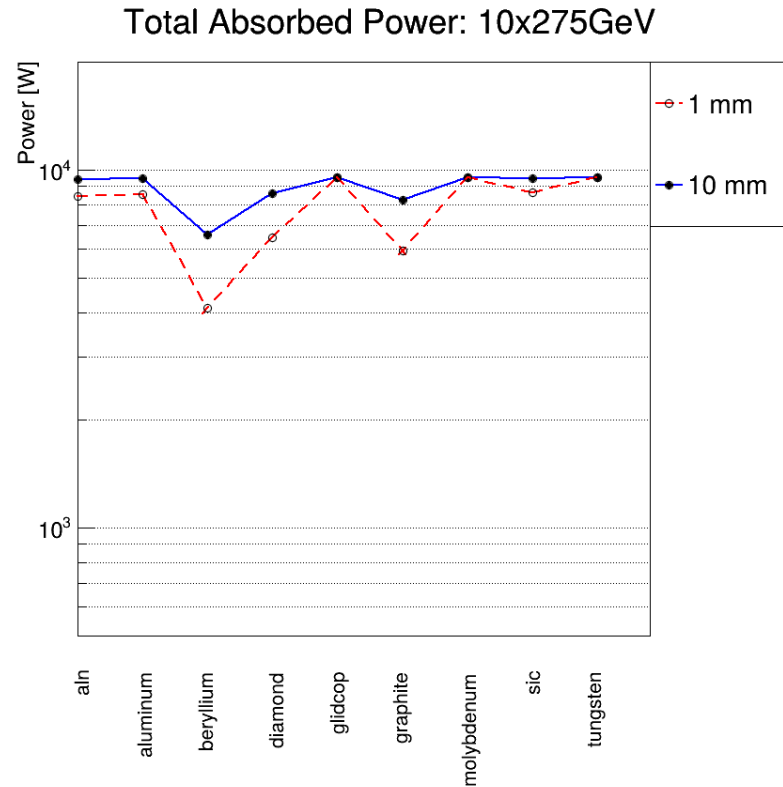
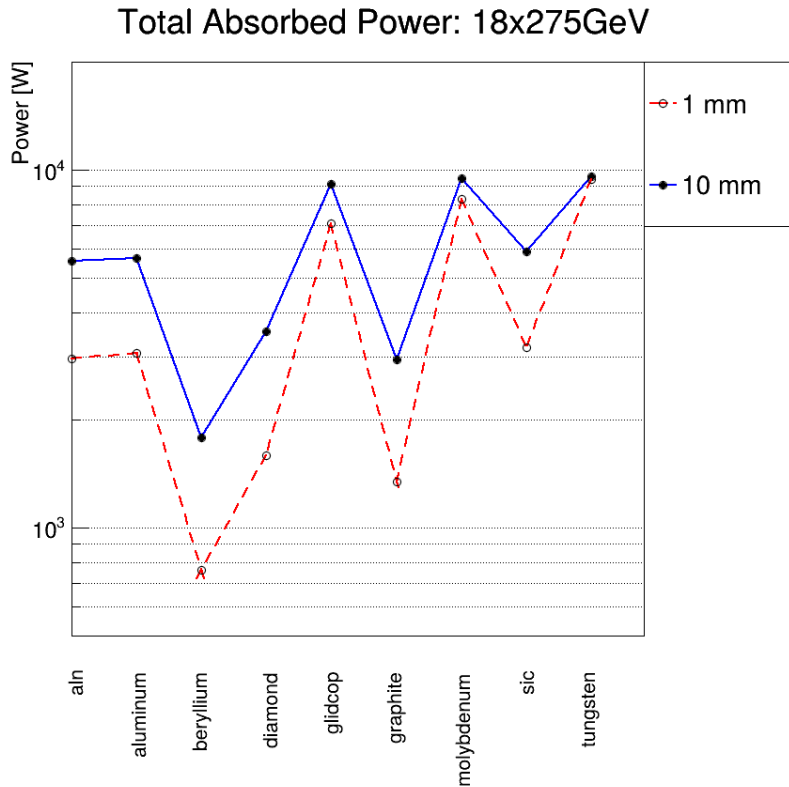
# Absorbed SR Power vs Lumi Window Thickness



**18x275 GeV @ 0.227 A    10x275 GeV @ 2.5 A    5x100 GeV @ 2.5 A**

# Total Absorbed SR Power

- Based on light source experience, **1-10kW** is a soft limit (manageable depending on the cooling system and thermal conductivity of the window material)
- For all materials at 10 GeV, we are close to the limit



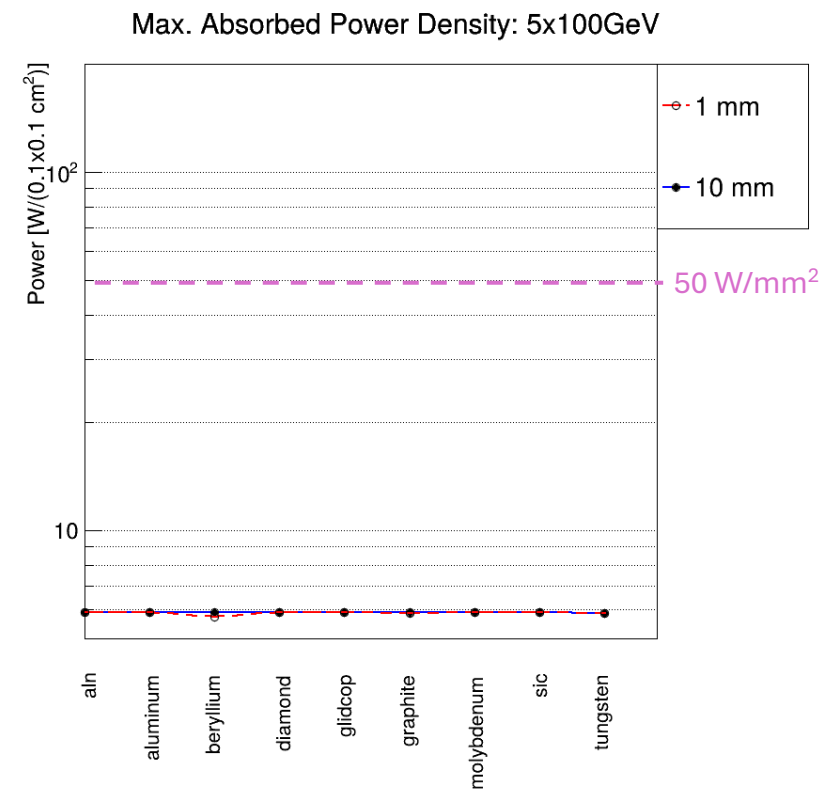
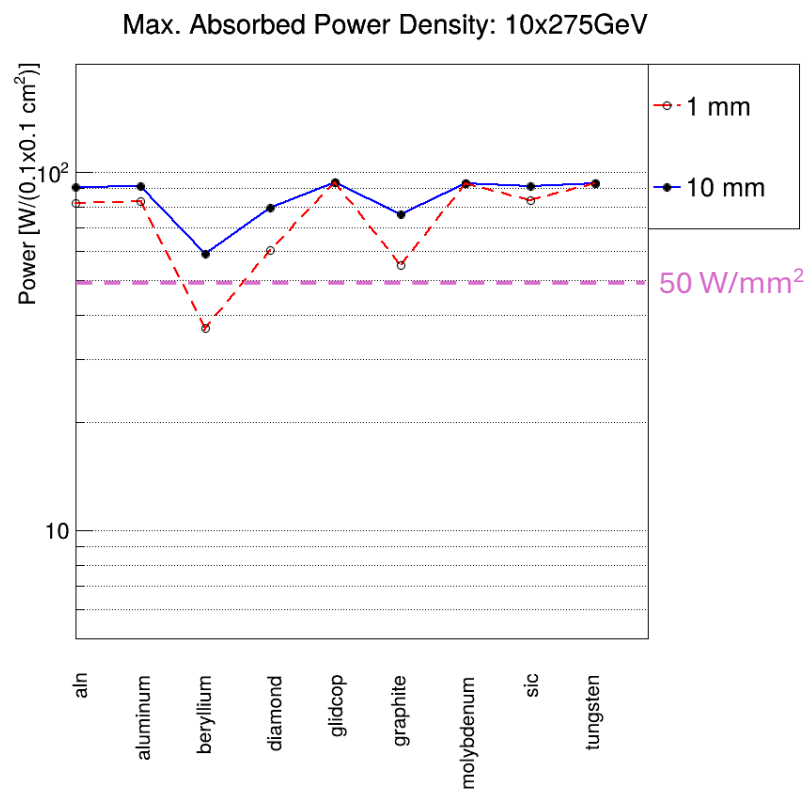
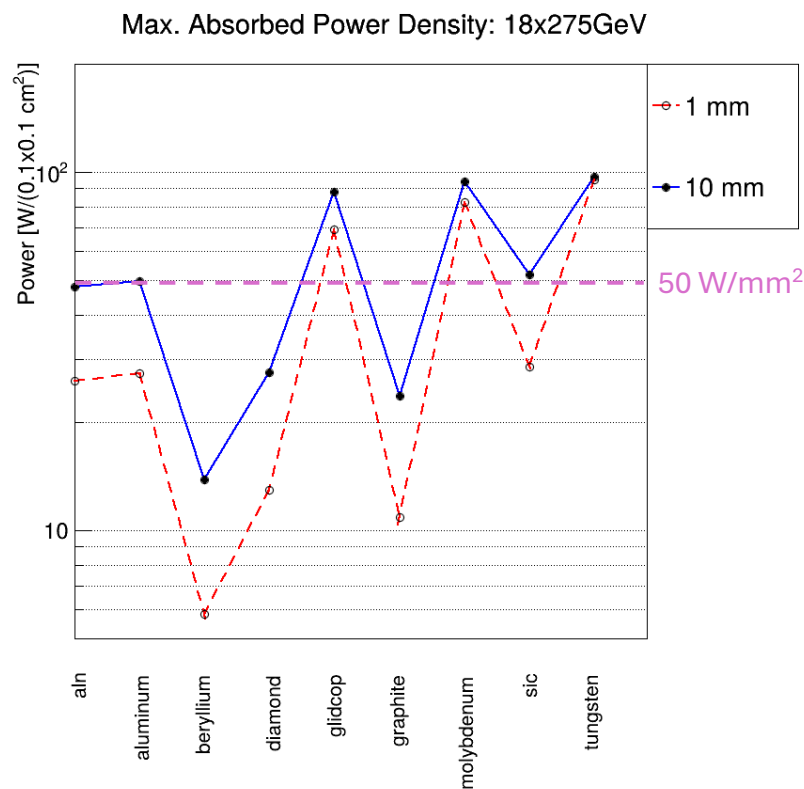
**18x275 GeV @ 0.227 A**

**10x275 GeV @ 2.5 A**

**5x100 GeV @ 2.5 A**

# Maximum Absorbed SR Power Density

- Based on light source experience, **50 W/mm<sup>2</sup>** is a hard limit
- At 10 GeV, we are very close or above the limit for all materials



**18x275 GeV @ 0.227 A**

**10x275 GeV @ 2.5 A**

**5x100 GeV @ 2.5 A**



# Proposal<sub>(part 1)</sub>:

Critical photon energy:

$$\varepsilon_c \sim \frac{\gamma^3}{\rho}$$

SR power:

$$P_\gamma \sim \frac{\gamma^4}{\rho^2}$$

Keep the SR fan from the second dipole outside the lumi window

Keep the beam core away from the window at  $>14\sigma_x$   
(see backup slides)

Lumi Exit Window

**B2eR** (~5m, 20 mrad) - *strong*

**B2AeR** (example: ~5m, ~2 mrad) - *weak*

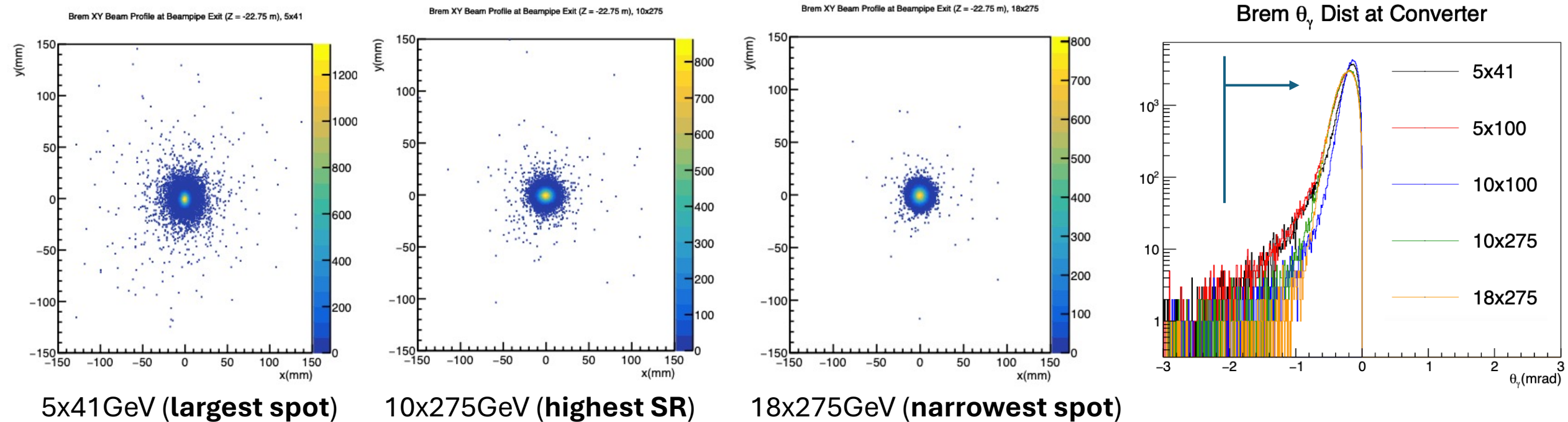
**B2BeR** (example: ~3m, ~18 mrad) - *strong*

IP6

IP6

- Making the dipole in front of the lumi window weaker (larger bending radius = small bending angle for the same length) reduces SR power and photon energy hitting the window.

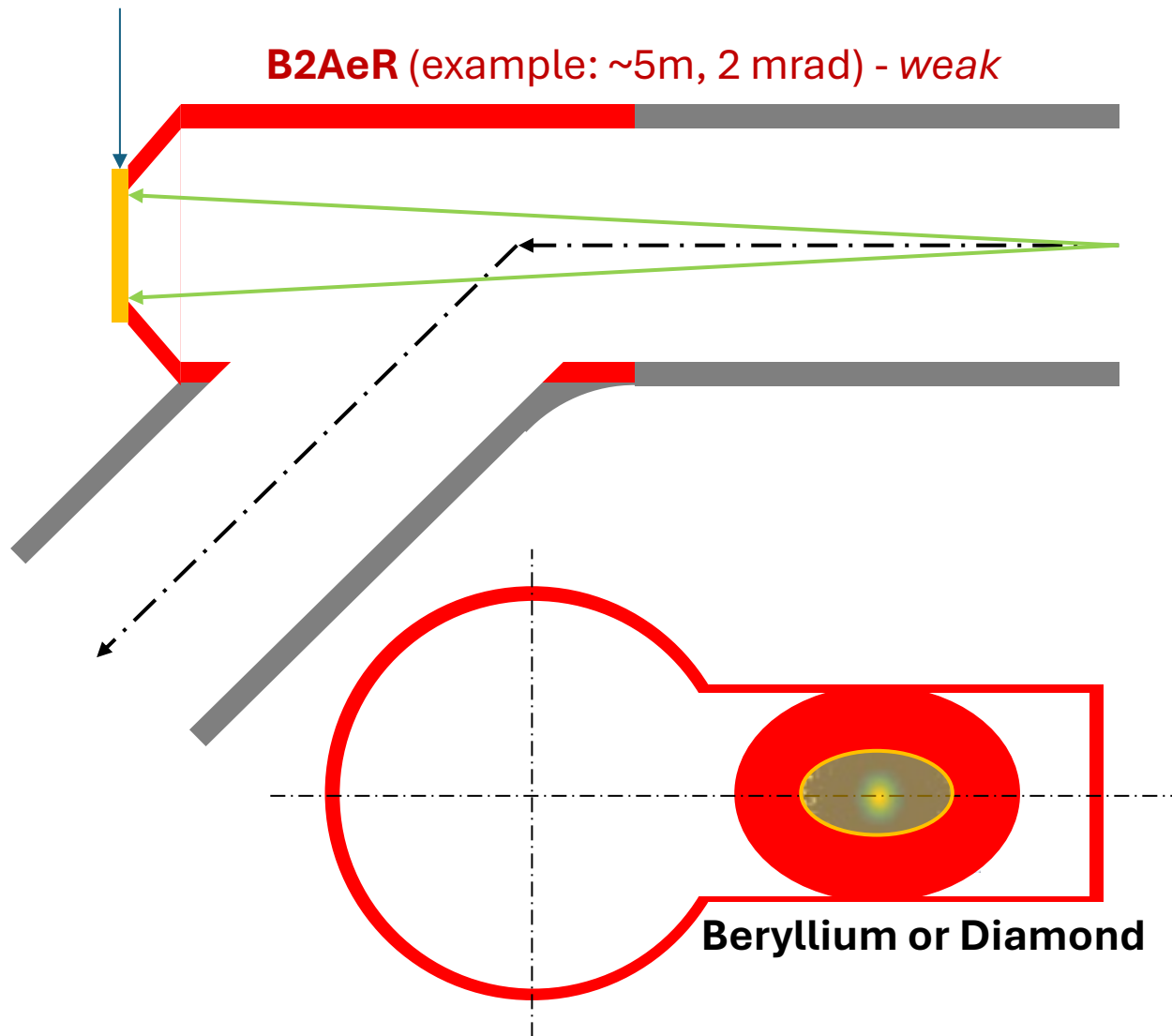
# Lumi Photon Distribution at the Window (~23 m from IP6)



- What is the minimum angular acceptance allowed for the lumi window at low (large spot/spread & low SR) and high (small spot/spread & high SR) energies?

# Proposal<sub>(part 2)</sub>:

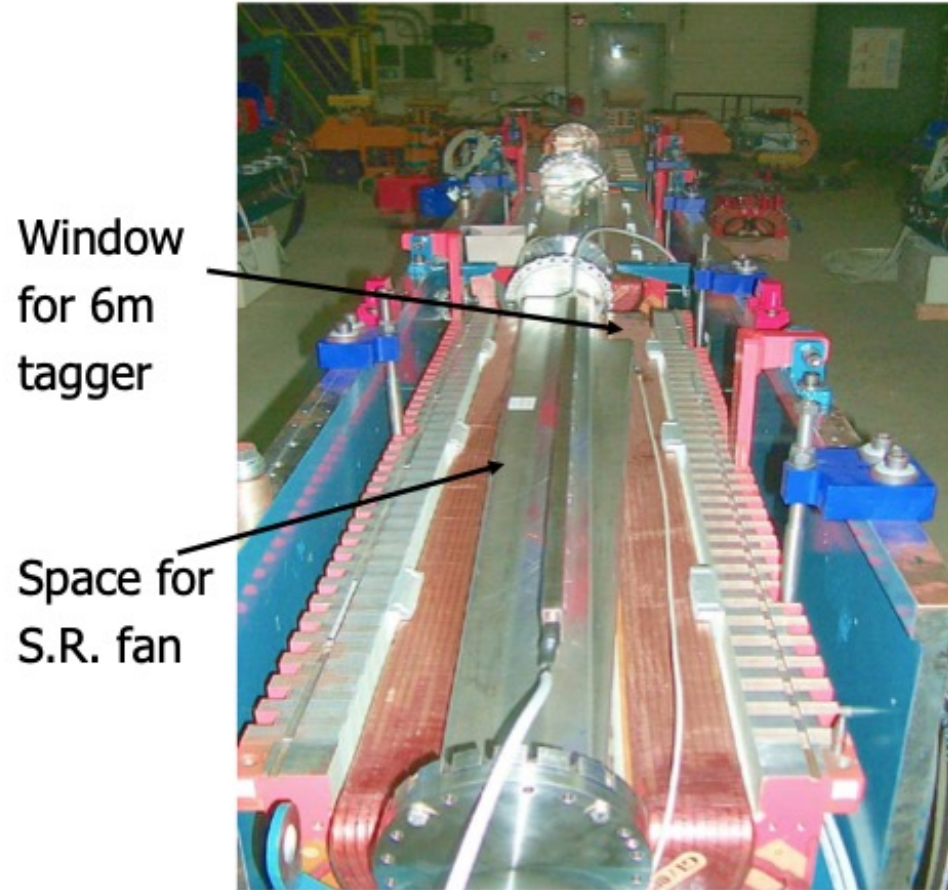
Lumi Exit Window



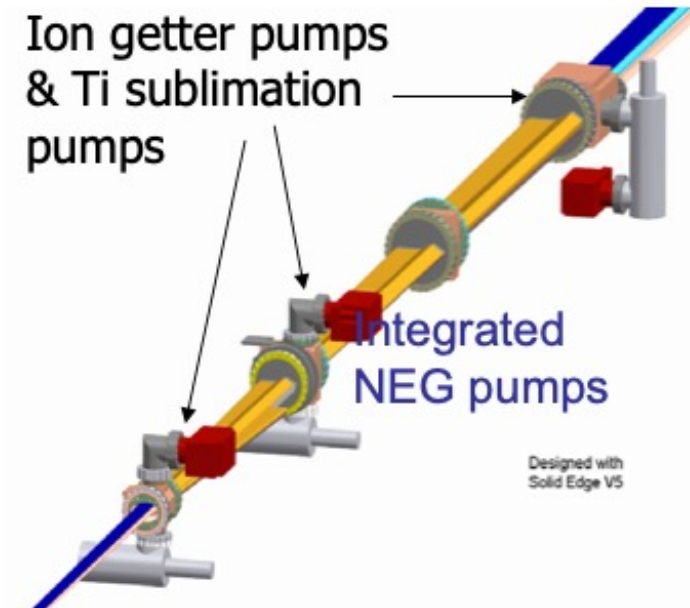
- Instead of using two beam pipes (one is for the main beam, another – for the luminosity photons), we could try building a beam pipe with an **ante-chamber**.
- The ante-chamber will reduce the beam pipe impedance.
  - Horizontal window and ante-chamber size is dictated by the distance between the electron and photon axes
  - The vertical window size is defined by the angular acceptance requirements.
  - For the acceptable impedance, the vertical ante-chamber size should be smaller than the beam pipe diameter.

# Vacuum System

Preamssembly in lab



Need quite complicated vacuum chambers to accommodate e and p beams and synchrotron radiation fan



# Lumi Exit Window Specs

- Is 50 mm exit window acceptable for the luminosity monitor?

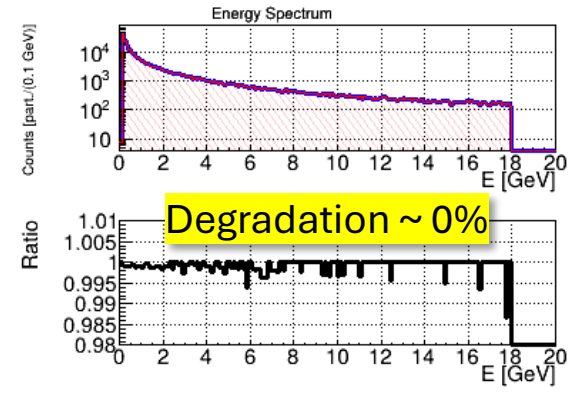
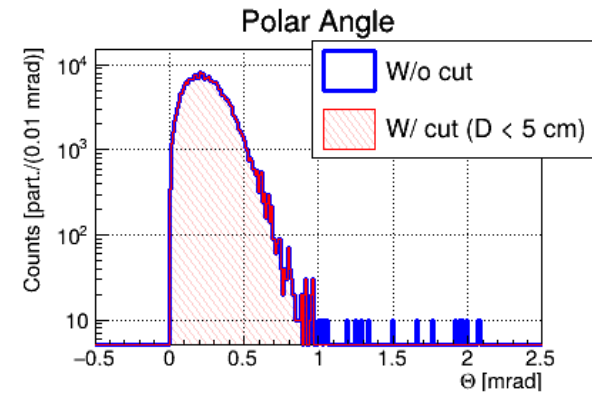
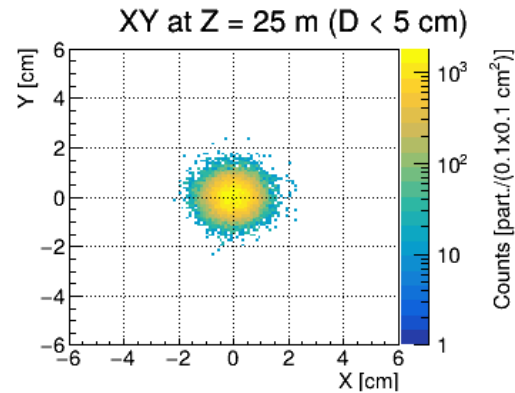
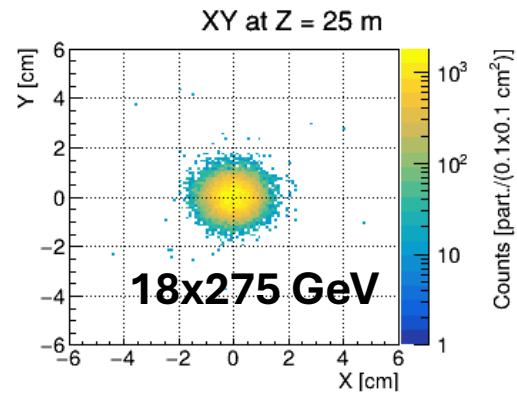
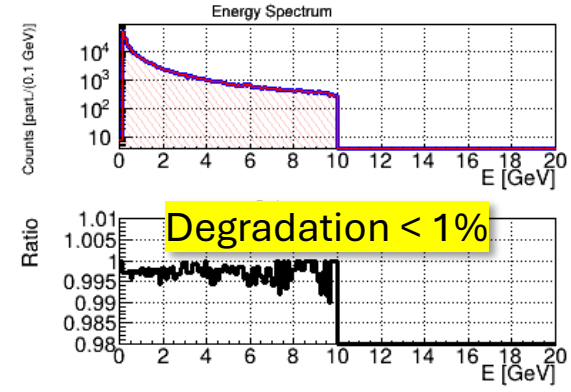
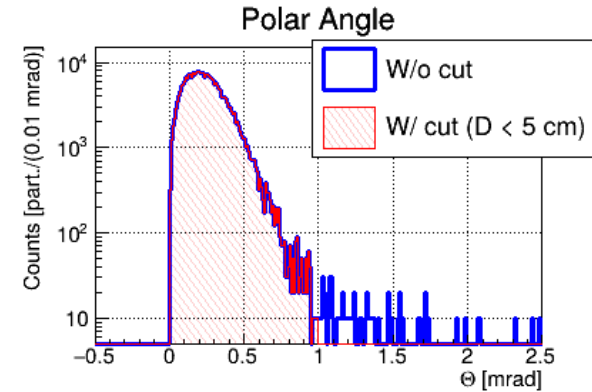
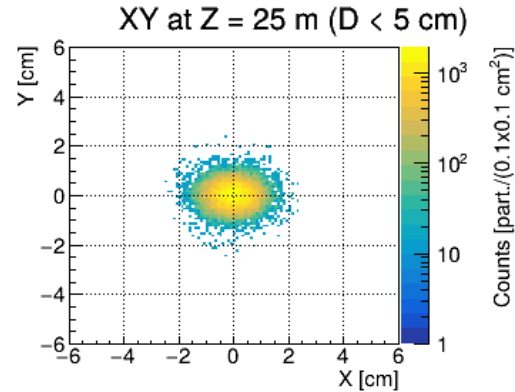
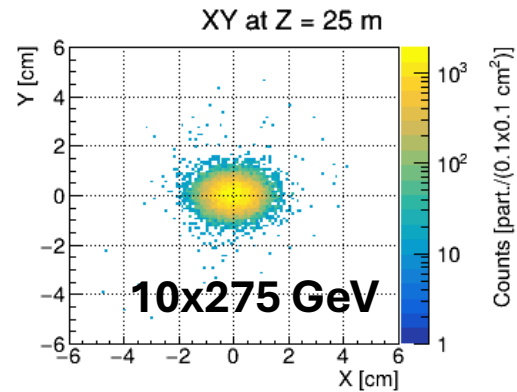
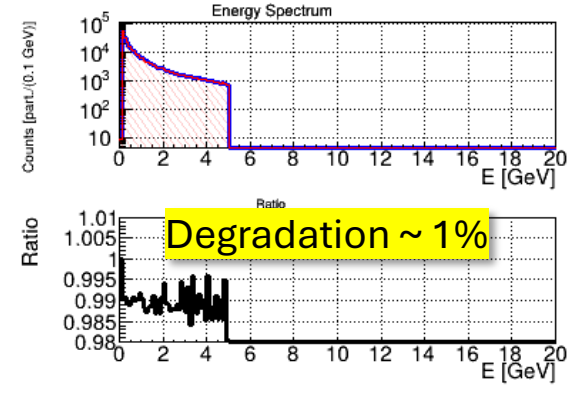
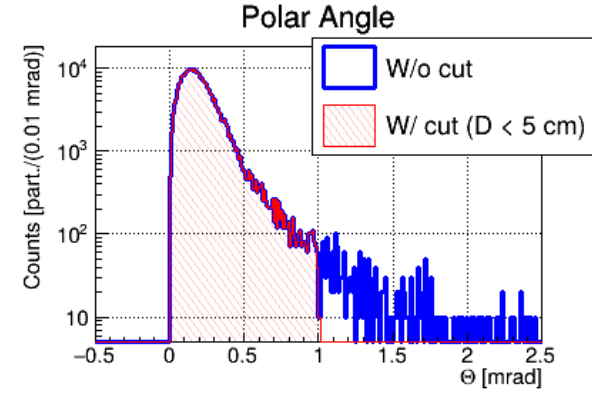
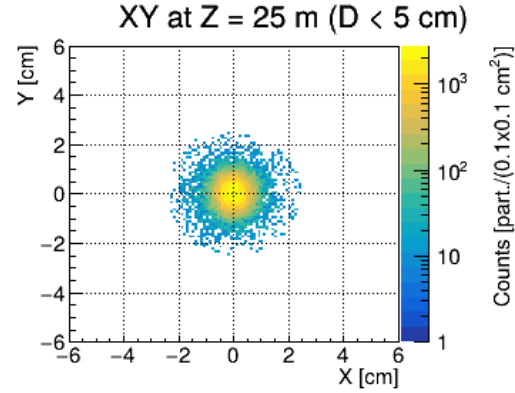
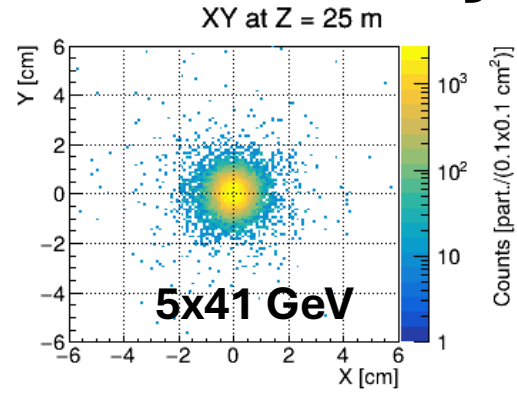
| <b>Diamond Infrared Windows: Specifications</b> |   |
|---|---|
| Thickness                                       | 300 – 1000 $\mu\text{m}$                  |
| Diameter  | 5 – 50 mm                                 |
| Wedge   | 0-1°                                      |
| Flatness  | < 1 fringe/cm @633 nm                     |
| Bakeable  | at up to 250°C                            |
| Vacuum tightness                                | He leak rate <10 <sup>-9</sup> mbar l/sec |



[https://www.diamond-materials.com/site/assets/files/1096/cvd\\_diamond\\_synchrotron\\_windows.pdf](https://www.diamond-materials.com/site/assets/files/1096/cvd_diamond_synchrotron_windows.pdf)



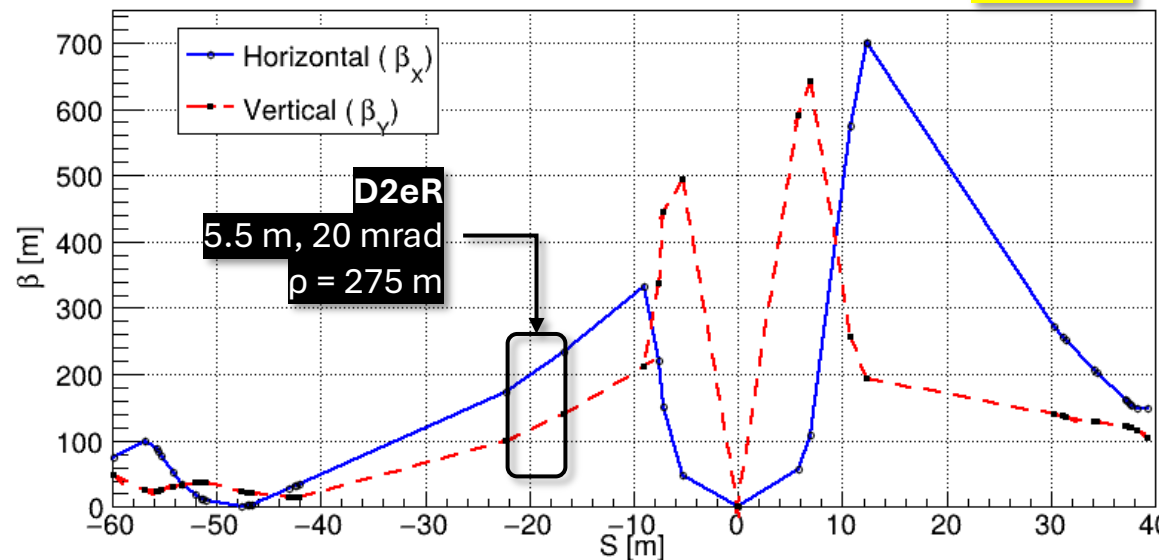
# Luminosity (Bethe-Heitler) Photon Distribution



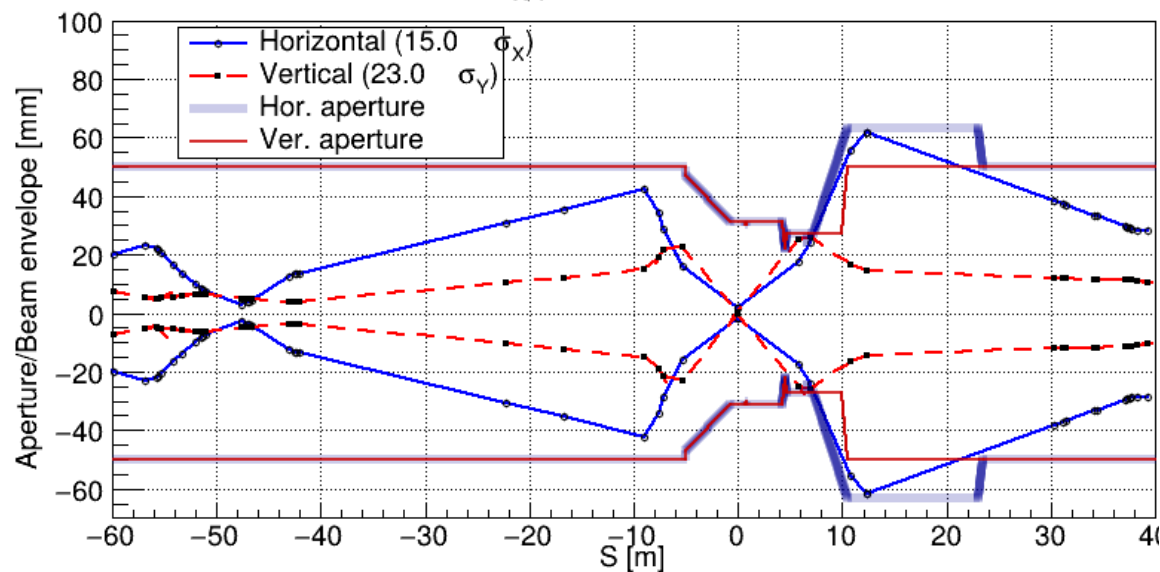
# D2eR Split-1

ESR v6.3.1 GeV

**Default**

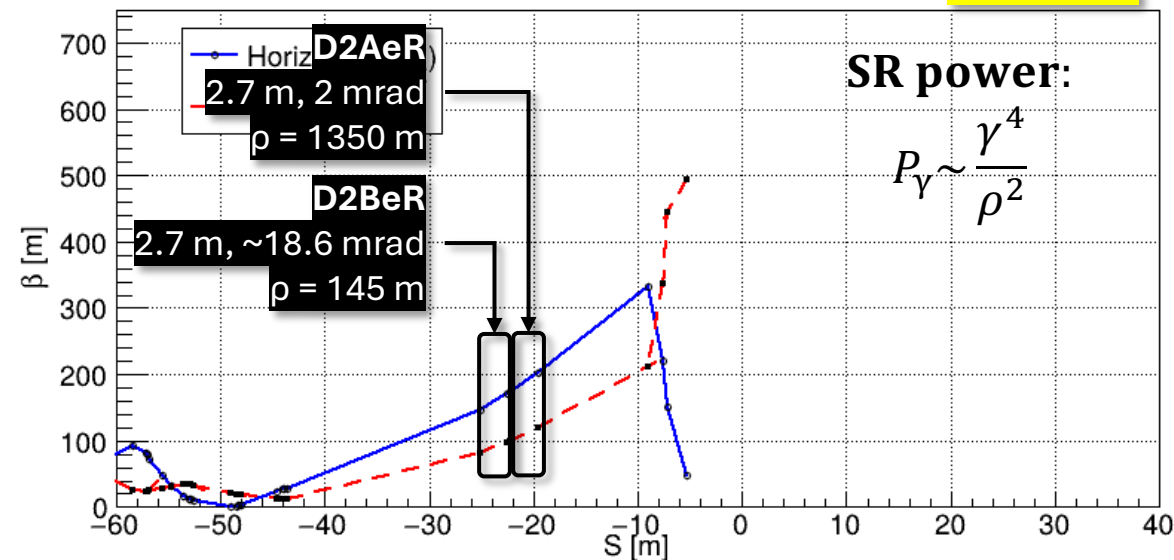


$\epsilon_{X/Y} = 24.0/2.0$  nm



ESR v6.3.1 GeV

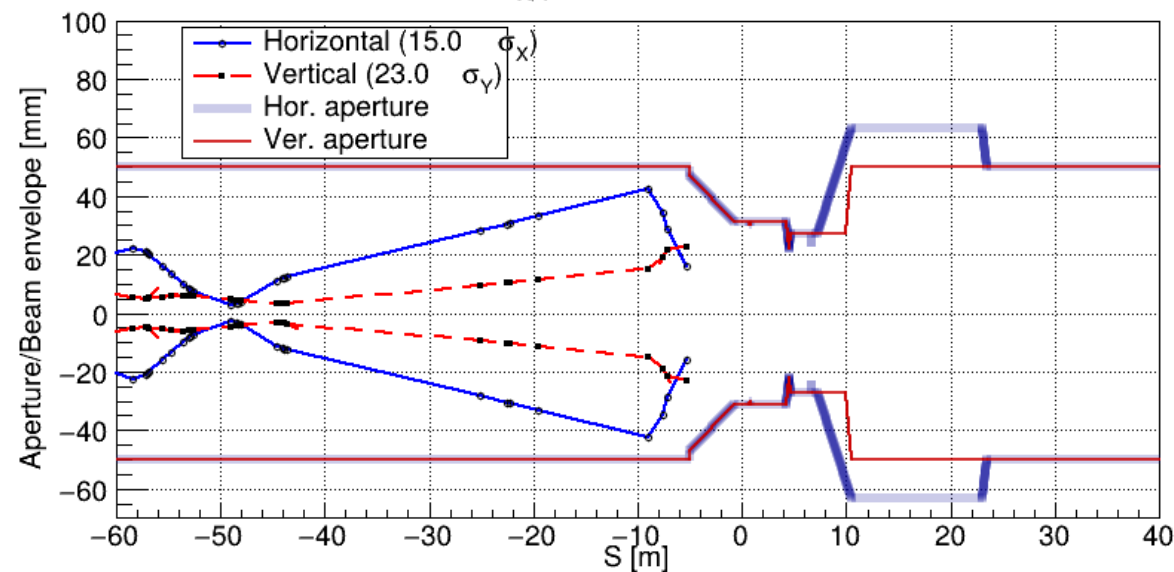
**Modified**



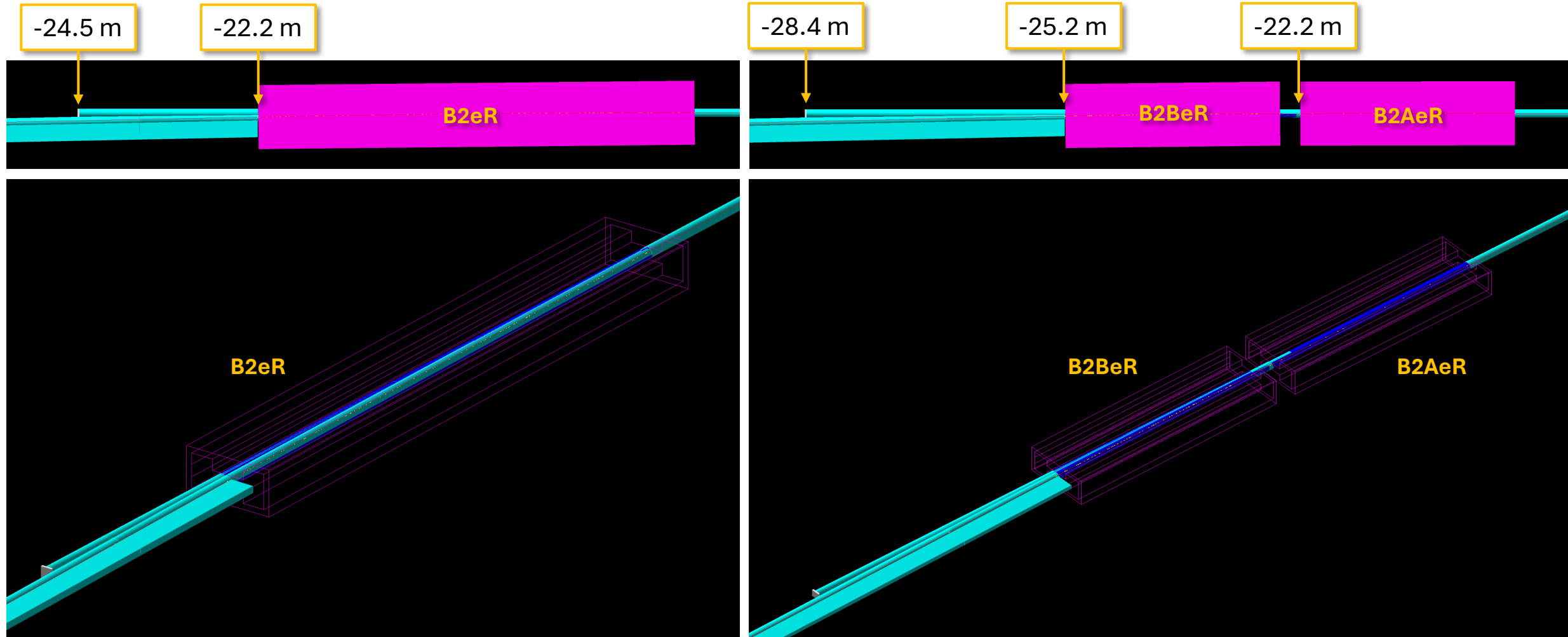
SR power:

$$P_\gamma \sim \frac{\gamma^4}{\rho^2}$$

$\epsilon_{X/Y} = 24.0/2.0$  nm

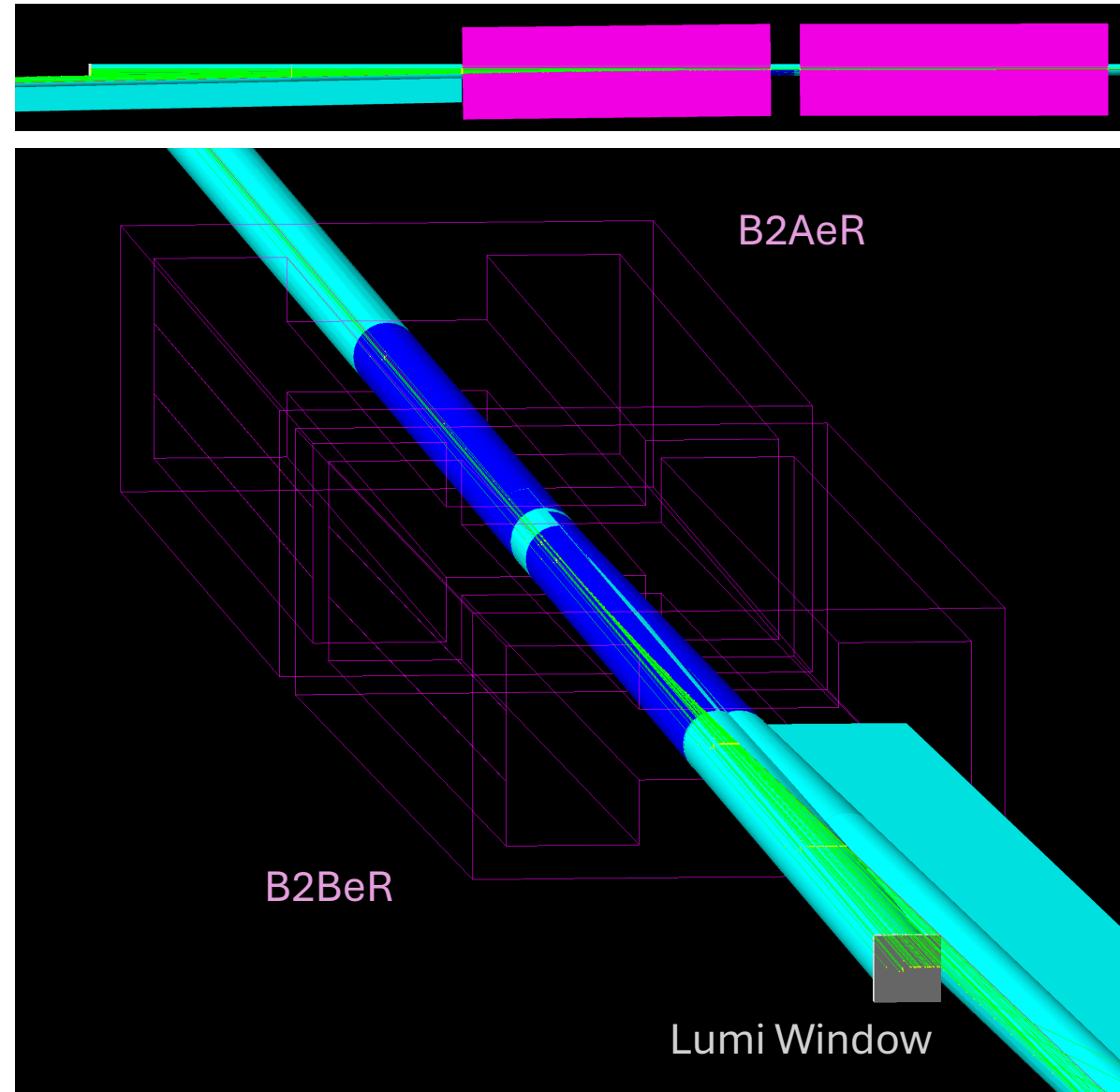
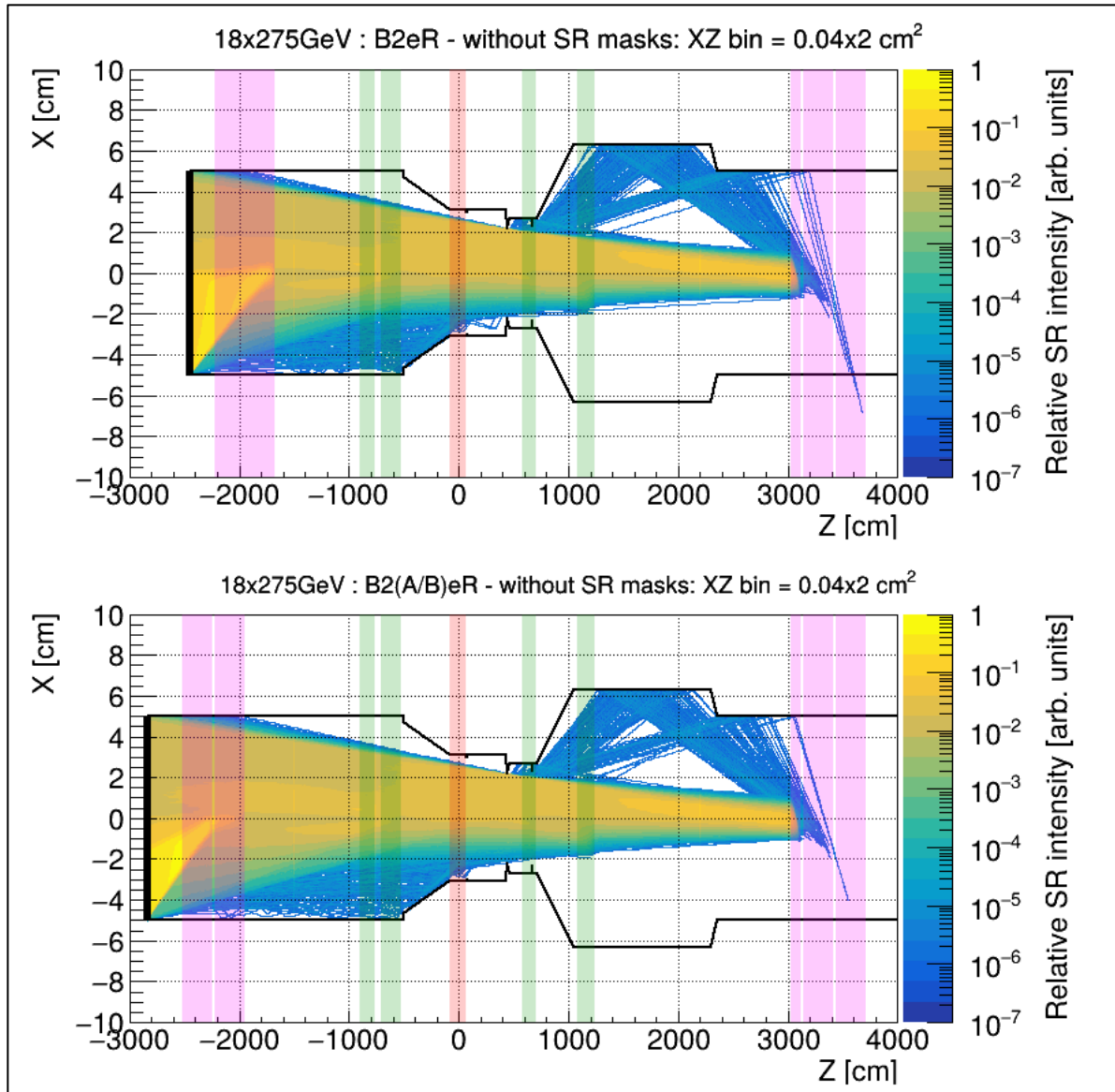


# SynradG4 Implementation-1





# X-ray Tracks Hitting the Window (for $E_\gamma > 10$ keV)



# SR Hitting the Lumi Window ( $E_\gamma > 30$ eV)

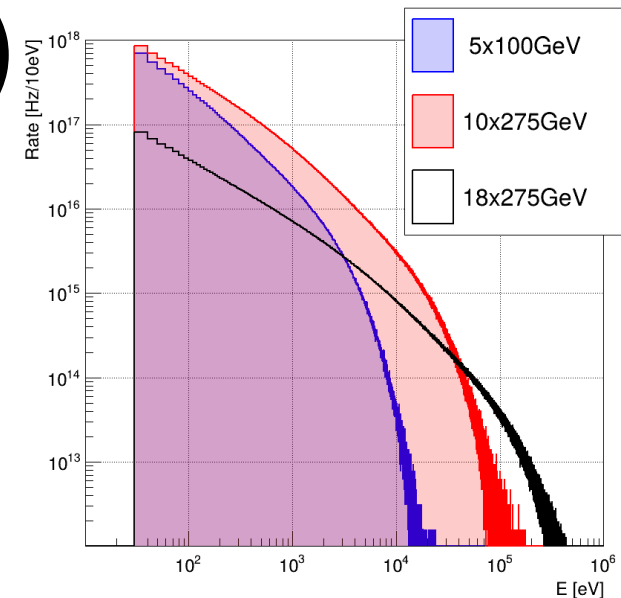
Still extremely high density

SR power: Critical photon energy:

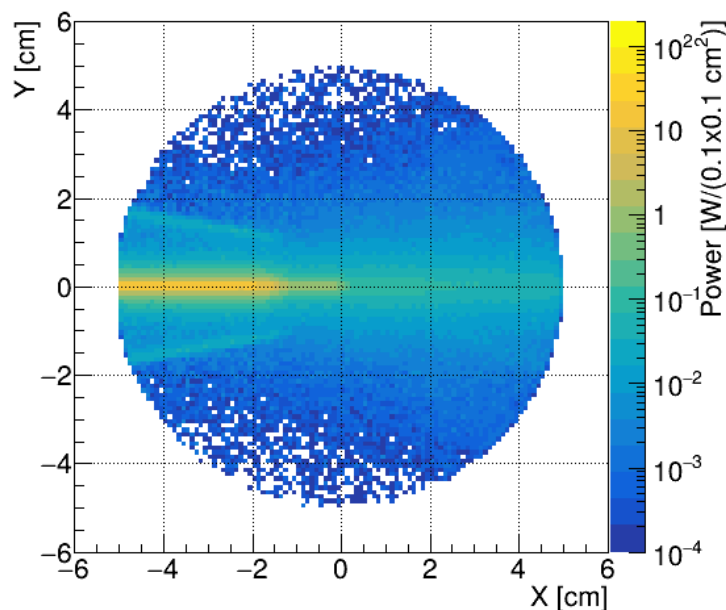
$$P_\gamma \sim \frac{\gamma^4}{\rho^2}$$

$$\varepsilon_c \sim \frac{\gamma^3}{\rho}$$

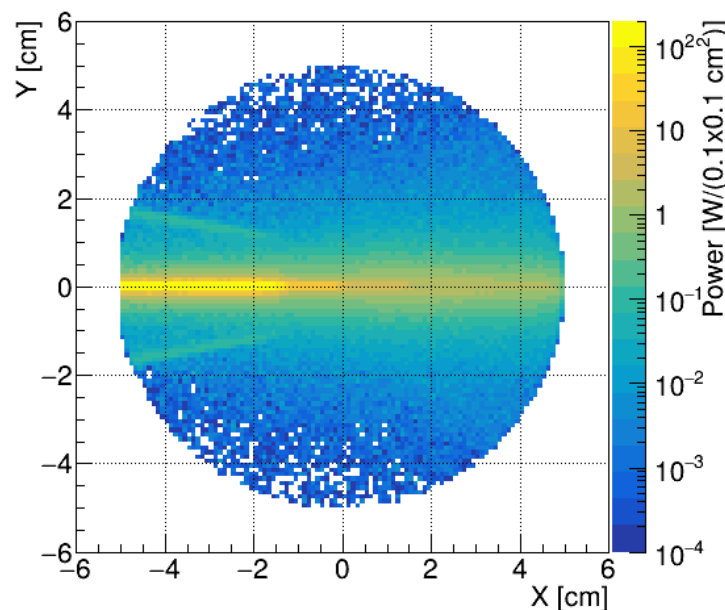
5 GeV @2.500A :  $P_{\max} = 13.9$  [W/mm<sup>2</sup>];  $P_{\text{tot}} = 1097.4$  [W]  
 10 GeV @2.500A :  $P_{\max} = 209.1$  [W/mm<sup>2</sup>];  $P_{\text{tot}} = 15405.7$  [W]  
 18 GeV @0.227A :  $P_{\max} = 215.4$  [W/mm<sup>2</sup>];  $P_{\text{tot}} = 15484.6$  [W]



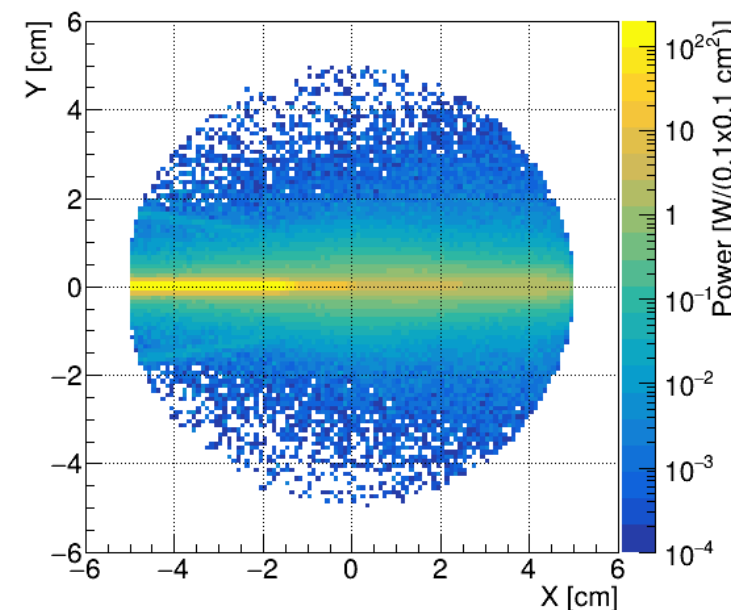
5x100GeV : Without SR masks



10x275GeV : Without SR masks

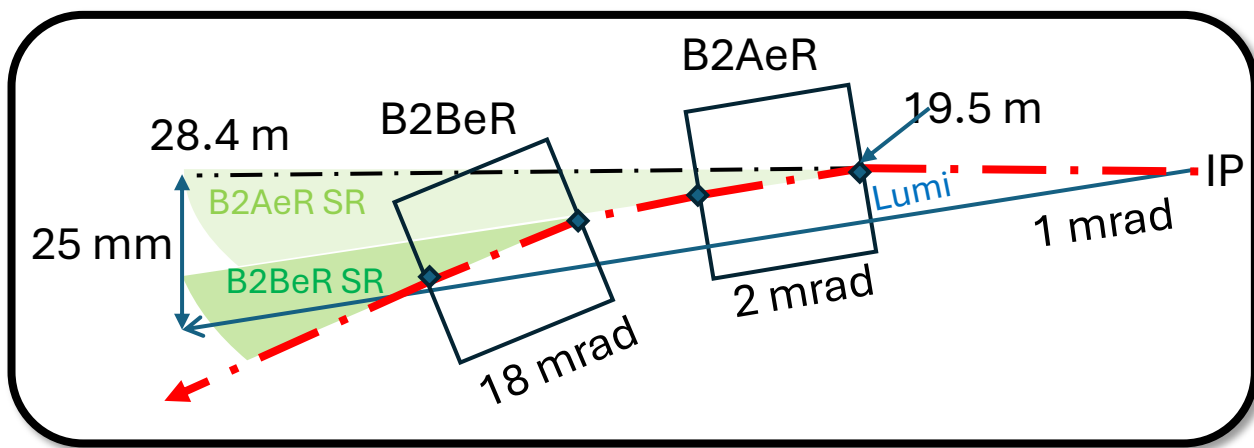


18x275GeV : Without SR masks

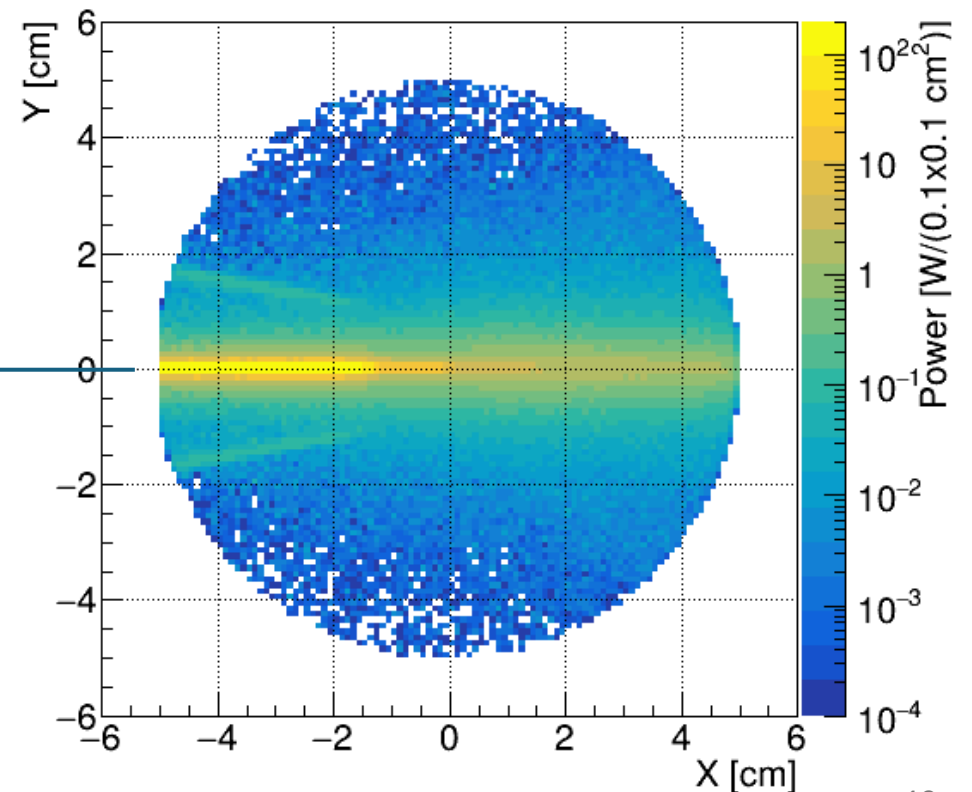
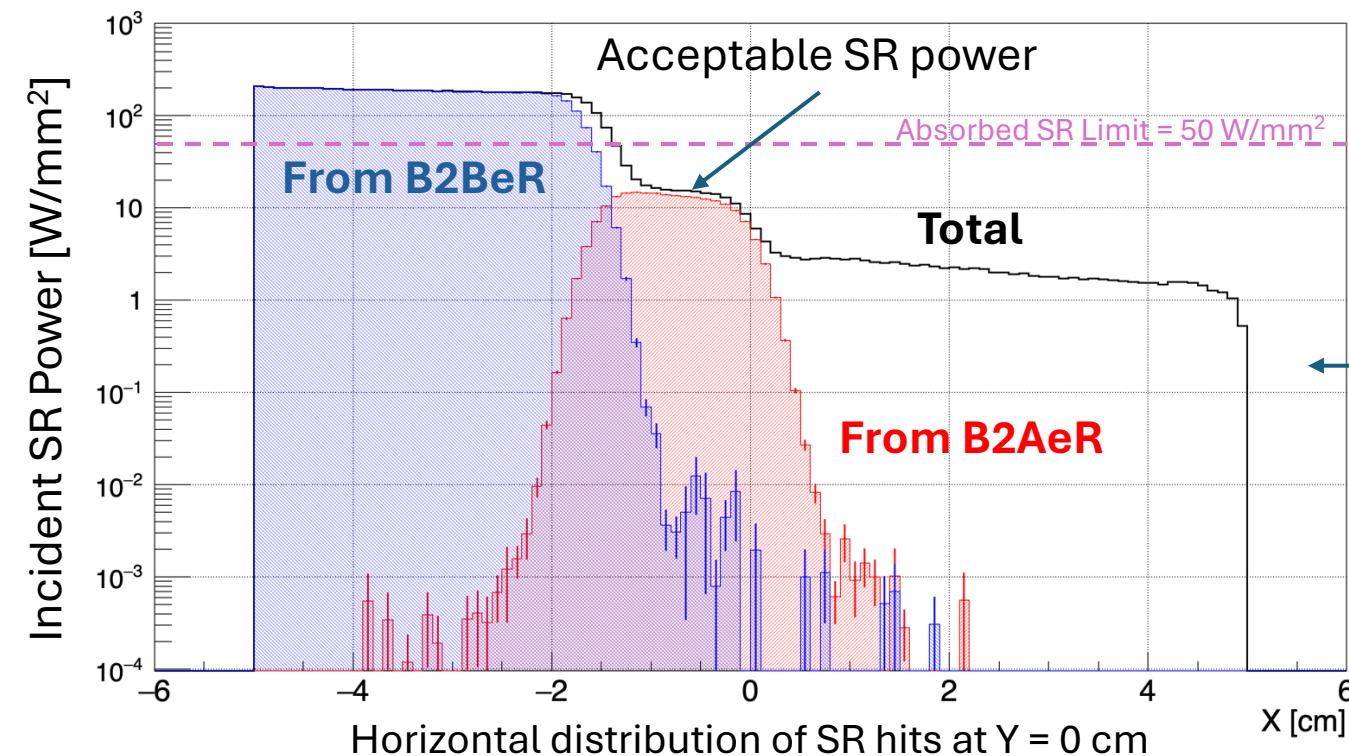


# Possible Improvements

- Let's widen the B2AeR distribution by enlarging the bending angle  $2 \rightarrow 3$  mrad
- This should move the B2BeR distribution further away from the luminosity window
- Additionally, placing B2AeR closer to the IP  $19.5 \rightarrow 16$  m allows moving the window closer towards the IP reducing the lumi photon spot on the window down to  $\varnothing 50$  mm



10x275GeV : Without SR masks

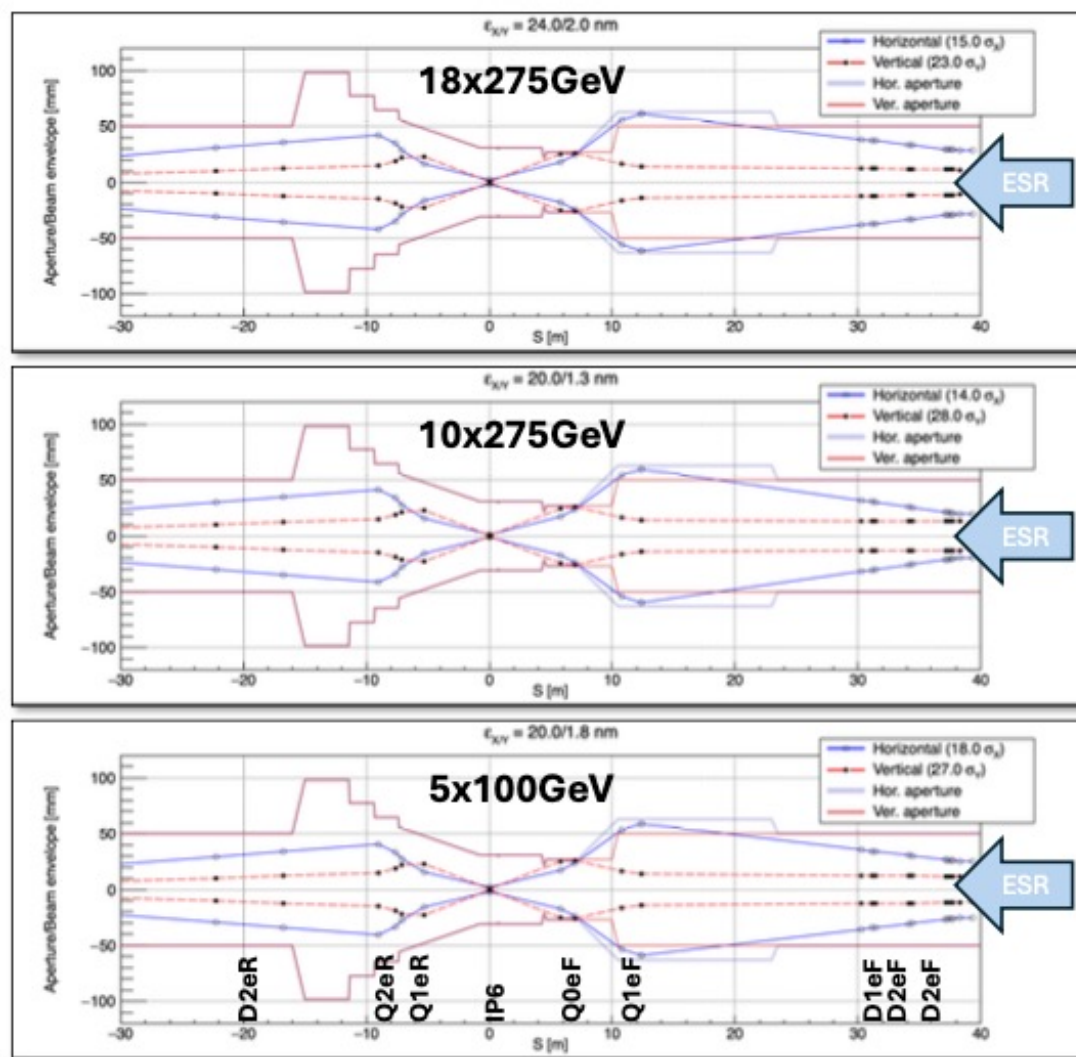




# Backup

# Aperture Change

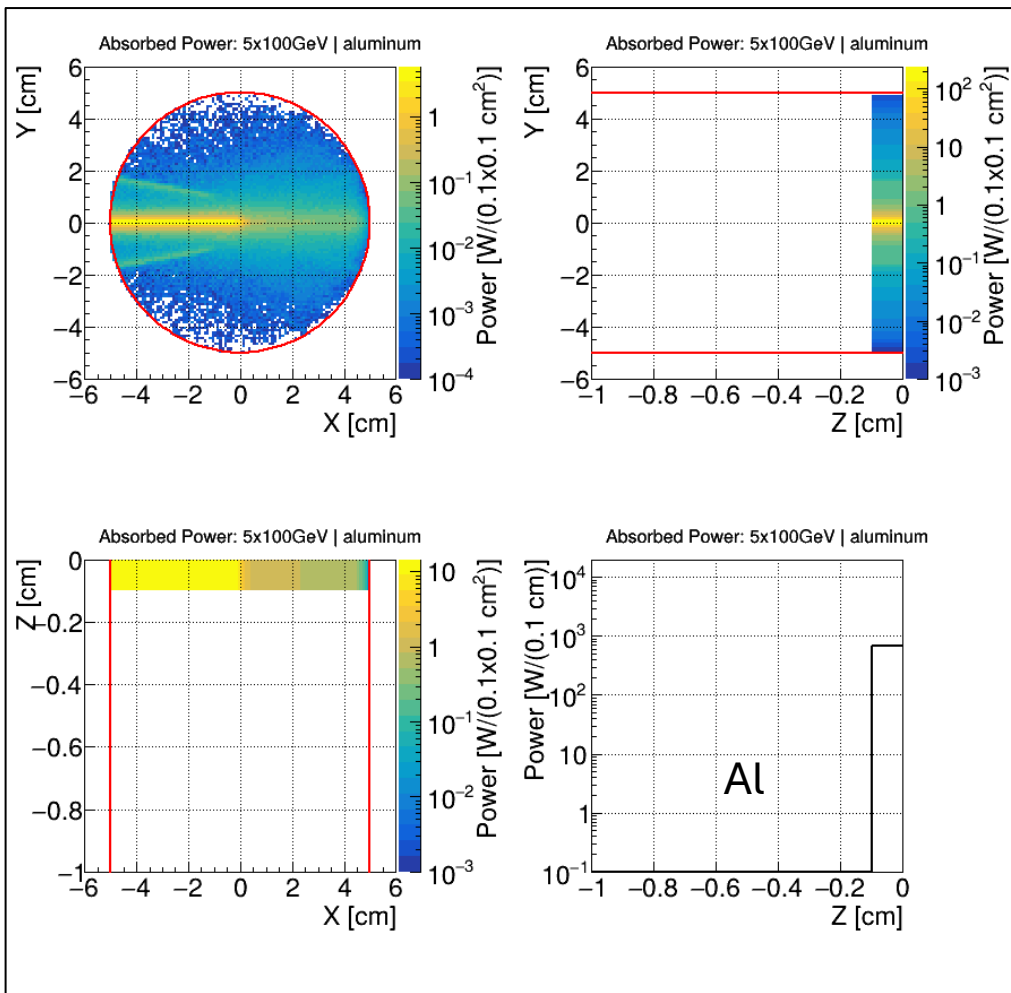
- Initially, the beam pipe aperture was made relatively large (approximately  $\varnothing 20$  cm) to allow SR to pass through.
- However, it is now feasible to reduce the aperture to  $\varnothing 10$  cm (the same as the CRAB cavities), as the **magnet can handle higher SR loads**, like the arc dipoles.
  - This aperture allows the luminosity photons exit the primary vacuum through the Lumi Window within a  $\pm 2$  mrad cone:
    - For  $\sim 20$  m (dipole end)  $\rightarrow \pm 4$  cm opening is needed.
- Despite this adjustment, the BWD beam pipe aperture within the **rear cryostat remains significantly larger** than the maximum beam size, which is dictated by the narrowest aperture in the FWD region.





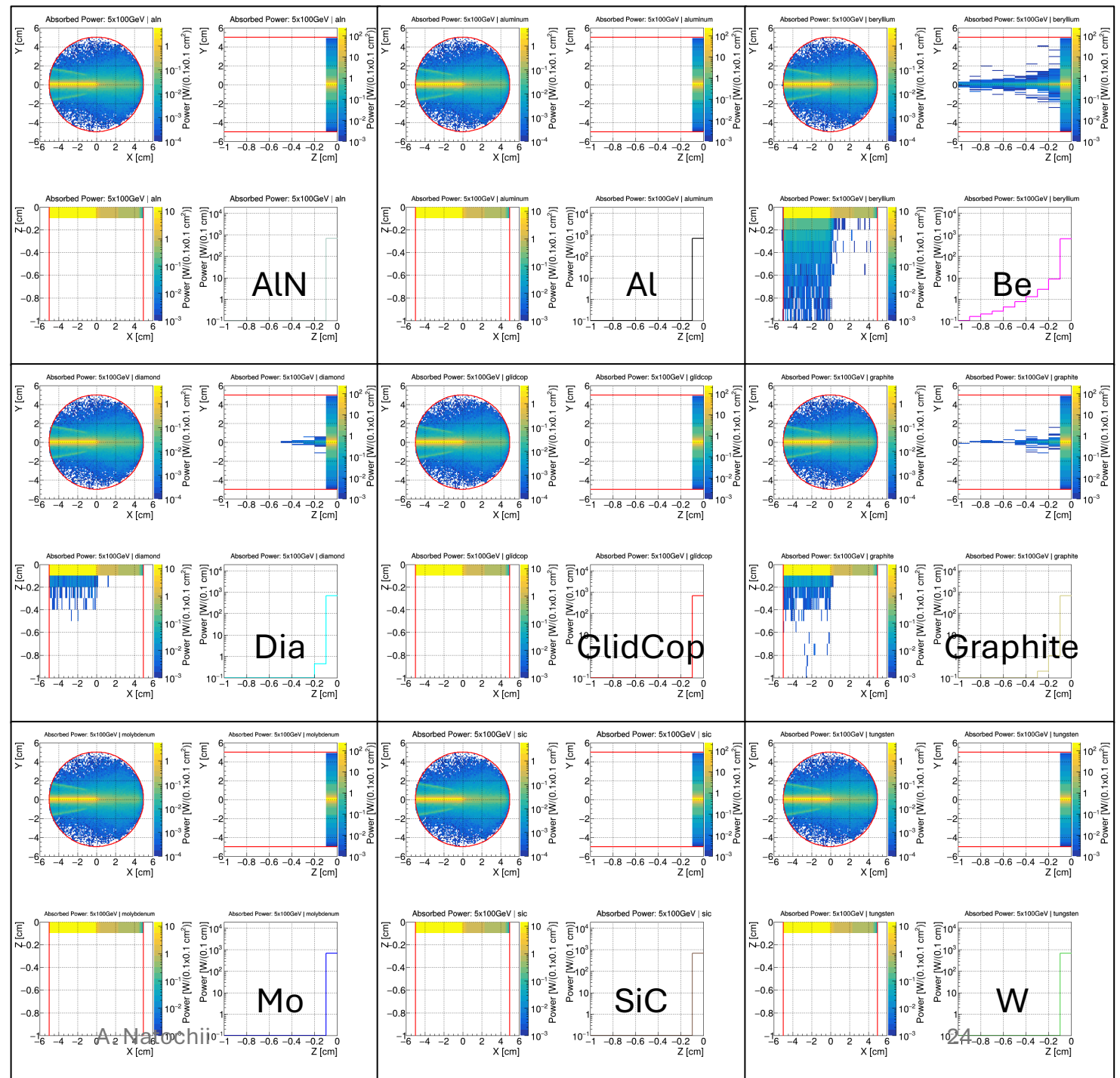
March 20, 2025

# Absorbed SR Power



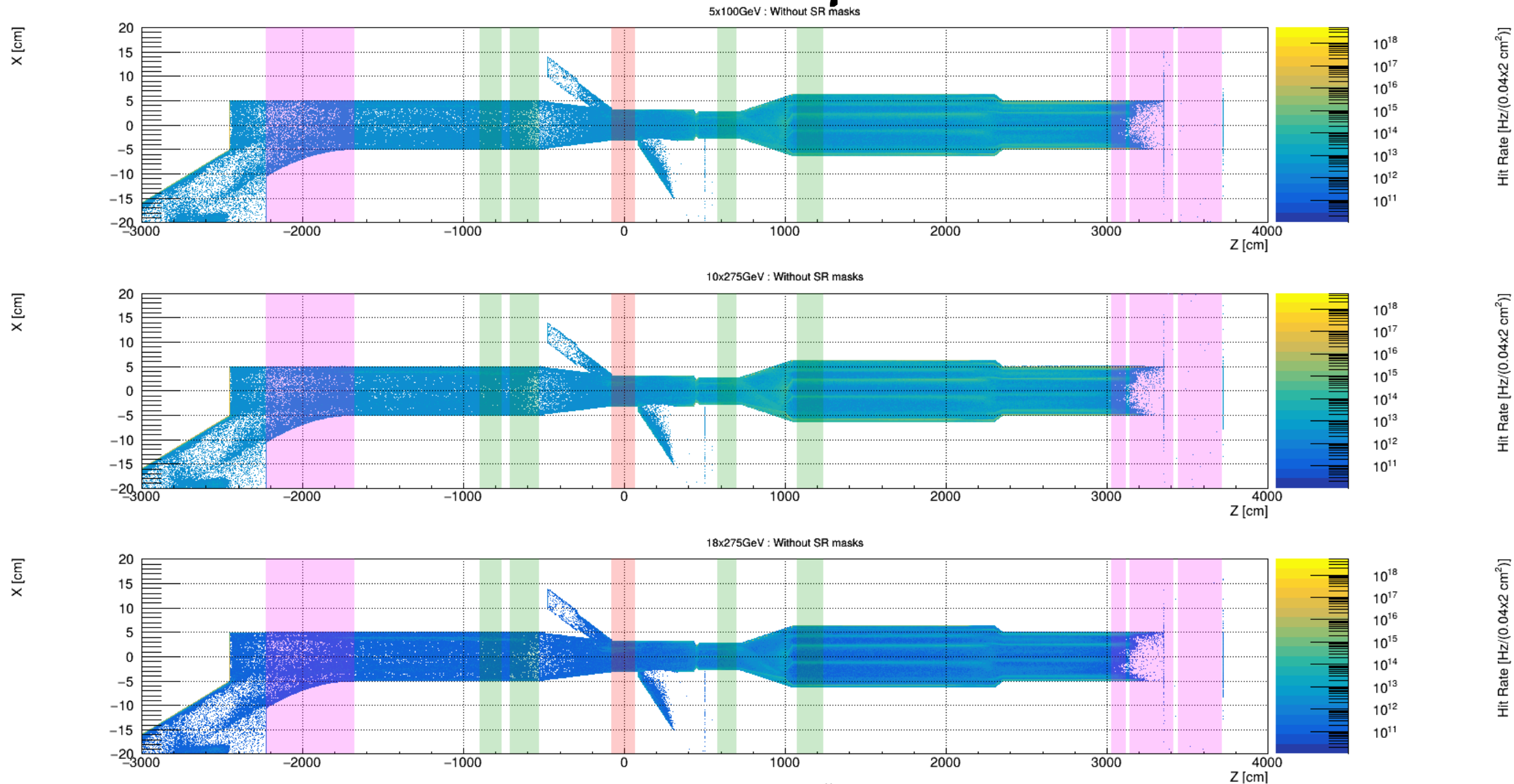
5x100 GeV @ 2.5 A

March 20, 2025





# SR Hits on the Beam Pipe ( $E_\gamma > 30$ eV)



# SR Hits on the Beam Pipe ( $E_\gamma > 30$ eV)-1

