

Friday Presentations

FRI, JULY 18



10:00 AM → 12:00 PM **Students give presentations**

🕒 2h 📍 small seminar room



12:00 PM → 1:30 PM **Pizza party**

🕒 1h 30m



1:30 PM → 2:30 PM **Students give presentations**

🕒 1h 📍 small seminar room



2:30 PM → 2:40 PM **Picture taking**

🕒 10m 📍 small seminar room



2:40 PM → 3:00 PM **Open discussion**

🕒 20m



- ❑ 13-minute presentation + 2-minute Q&A
- ❑ Can be anything you learned at BNL or from the whole NuSTEAM/NuPUMAS program, found interesting, or want to do in the future



Reactor Neutrinos

Chao Zhang
czhang@bnl.gov

Outline

- ❑ Nuclear Reactors
- ❑ Reactor Neutrinos
- ❑ Experimental Detection
- ❑ The Daya Bay Experiment



My experience with reactor neutrinos

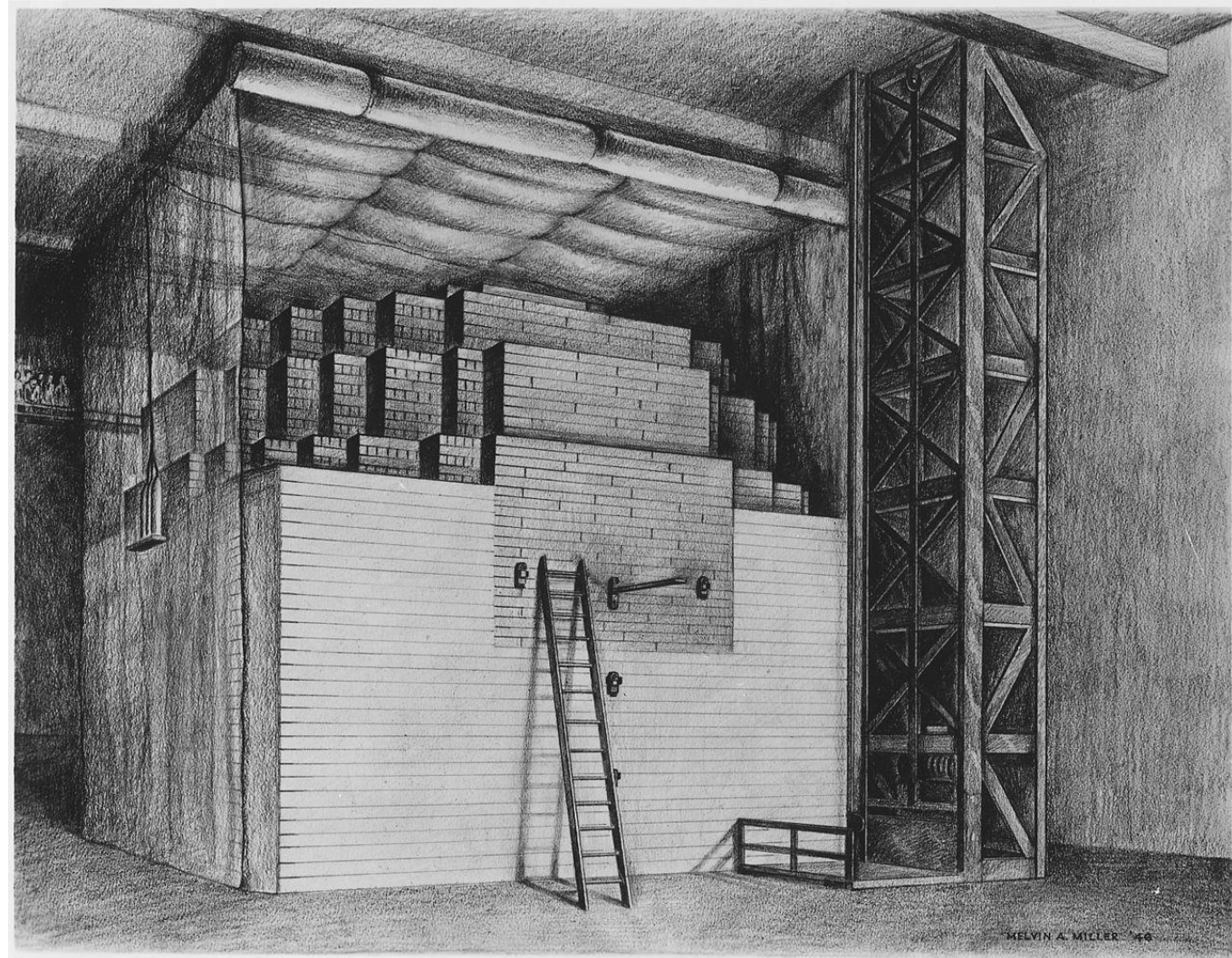
- ❑ PhD at CalTech (2011)
 - KamLAND experiment (Japan)
- ❑ Came to BNL as a postdoc then became a staff scientist
 - Daya Bay experiment (China)
 - PROSPECT experiment (US)
- ❑ Love cats
 - My YouTube Channel (The Furry Crew):
<https://www.youtube.com/@czczc-25>

World's First Nuclear Reactor

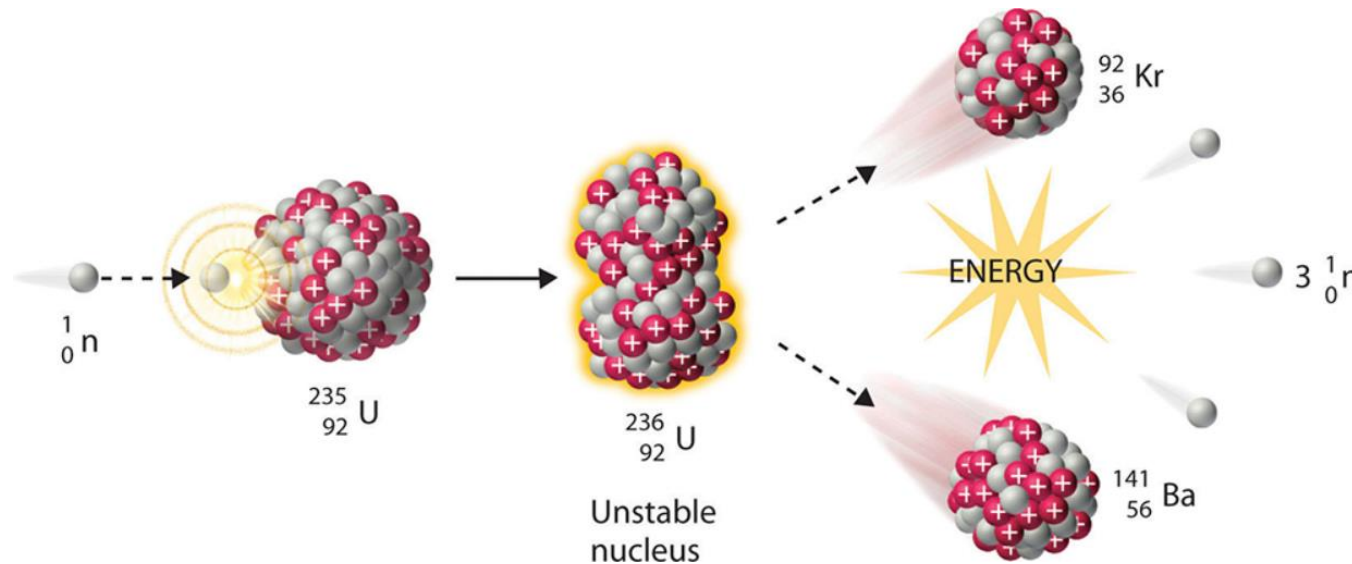
Chicago Pile-1 (CP-1)

- ❑ Part of the Manhattan Project
- ❑ Designed and tested by a team of 49 scientists led by Enrico Fermi at U. Chicago
- ❑ Dec 2, 1942: first human-made self-sustaining **nuclear chain reaction**
 - Fuel: 40-ton Uranium Oxide and 6-ton Uranium Metal
 - Neutron moderator: 380-ton Graphite blocks
 - Control rods: Cadmium sheets

CP-1 in a squash court under the stands of Stagg Field at U. Chicago



Fission and Energy Release



1 fission = 200 Mev

1 gram U-235 fissioned = 8.6×10^{10} joules = 24,000 kwh

(Equivalent to lighting a small city for overnight)

24,000 kwh requires 3.2 tons of coal

12.6 bbls oil

Energy Density (energy / mass)

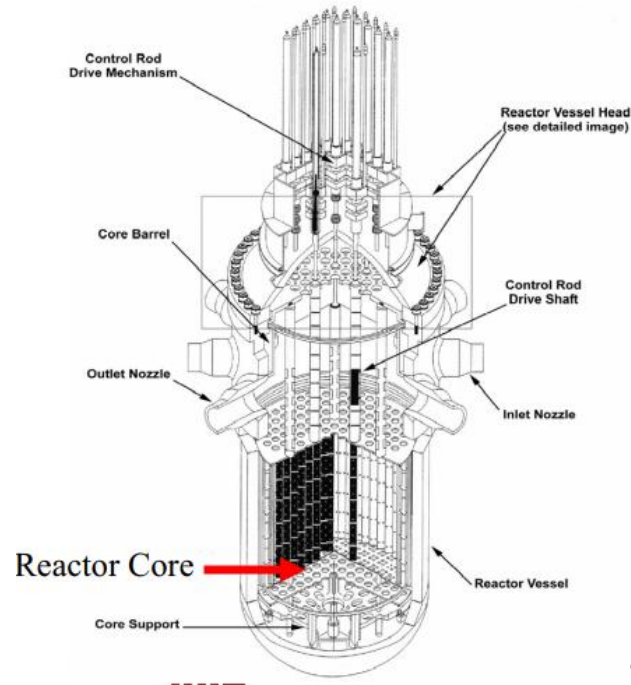
Energy Density of U-235 = 28,000 times energy density of coal

Key properties of fission

- ❑ Release substantial energy (mostly as kinetic energy of the fission fragments)
- ❑ Release excess neutrons: possibility of chain reaction.

Reactor design requirement

- ❑ Fission requires thermal neutron: needs “moderator” to slow down neutrons
 - Water, heavy water (D_2O), Graphite, etc.
- ❑ Controllable fission: reactor engineering to make output neutron = 1 (critical condition)



- ❑ Reactor Core Design
 - Core Power Distribution
 - Ability to shutdown plant
 - no fuel failure or melting

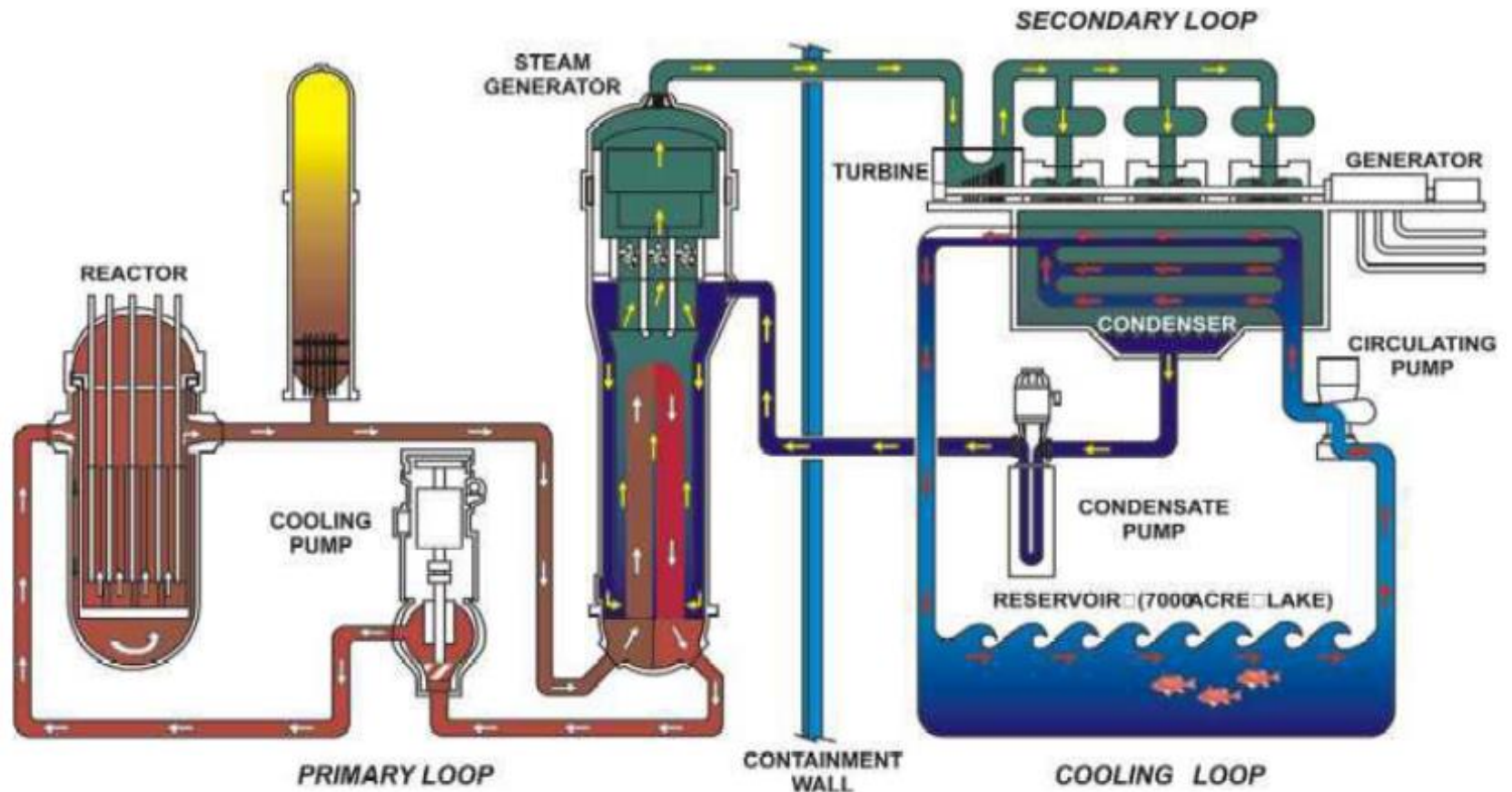
❑ Core Heat Removal

- Coolant: Heat Transfer
- Safety Systems (Emergency)

❑ Confinement of Radioactivity

- ❑ Electricity Production
- ❑ Spent fuel processing

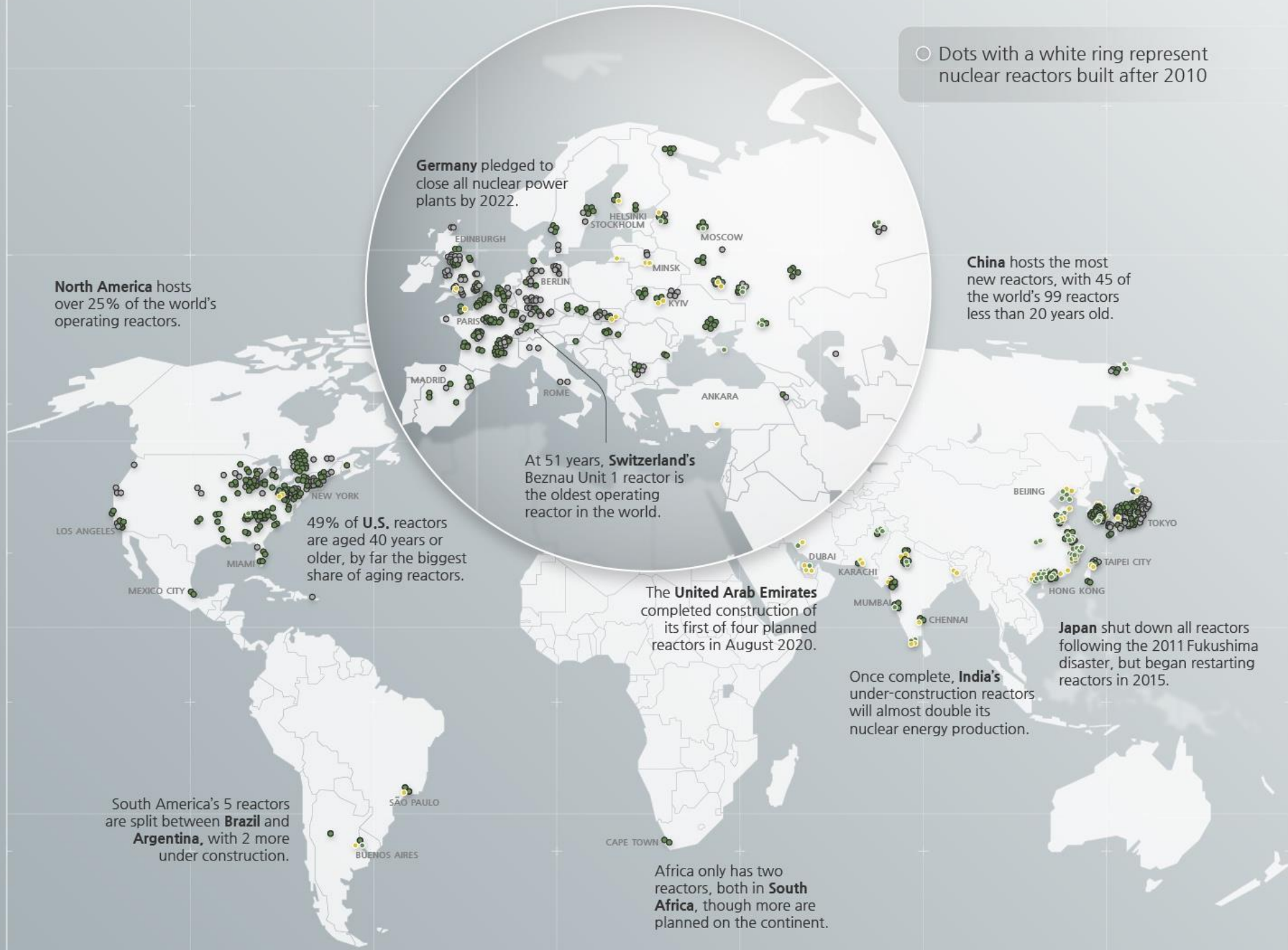
Schematic of a Pressurized Water Reactor (PWR)



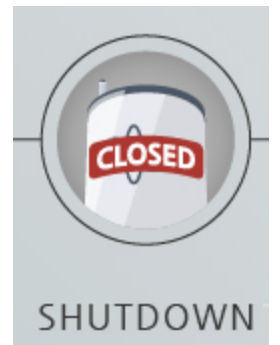
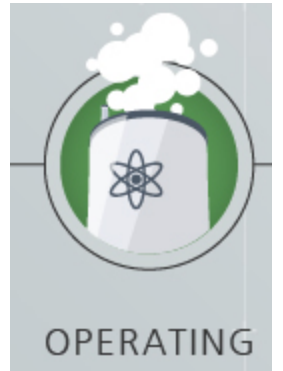


Video with annotations:

[Breazeale Nuclear Reactor Start up, 500kW, 1MW, and Shut Down \(ANNOTATED\) - YouTube](#)



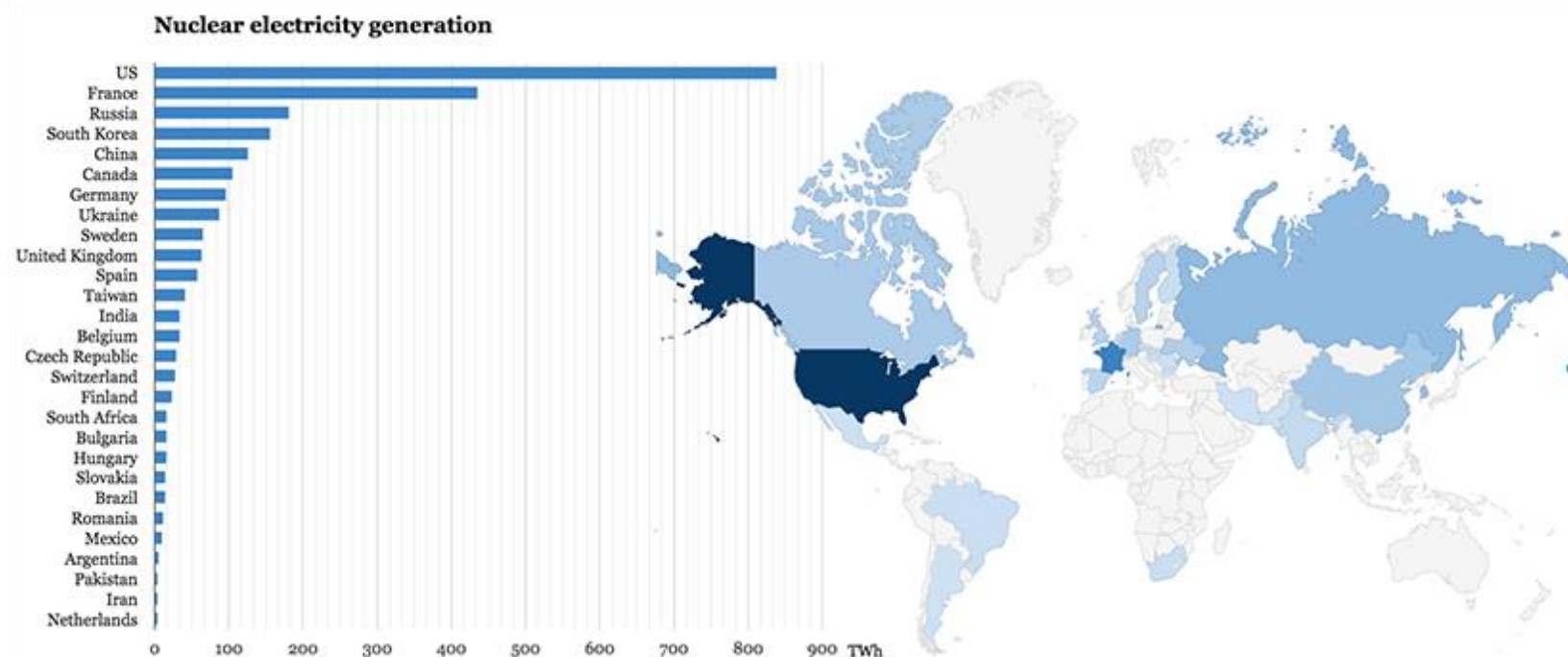
○ Dots with a white ring represent nuclear reactors built after 2010



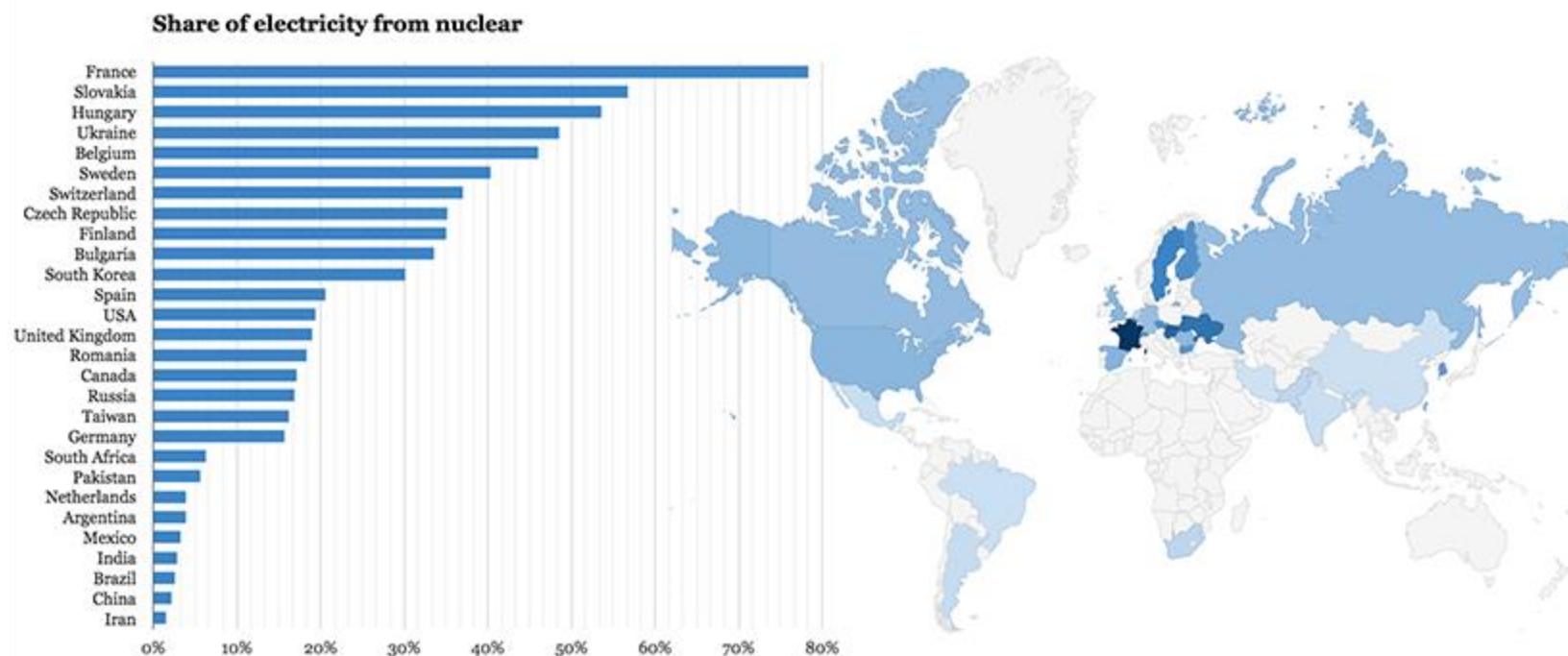
[Map link](#)

Country Ranking in 2015

Total Electricity
from nuclear



Percentage
from nuclear



Nuclear Reactor as a Research Tool: Neutron Source



GRAPHITE RESEARCH REACTOR

Operated: 1950 to 1969

World's first peacetime research reactor. Fuel placed in 700-ton graphite "pile" that moderated fission. Scientists exposed experiments to neutrons by inserting them into slots on top and three sides of the core.

Initially ran on natural uranium, but in 1958 fuel was switched to enriched uranium, with reactor operating at 20 megawatts.

Scientific advances

- The radioactive isotope Technetium-99m, used as a medical tracer and similar to X-rays for diagnostic imaging, first detected here.
- Multi-grade motor oils developed as a result of studying engine piston rings in the reactor.
- Irradiated seeds used to produce the Star Ruby grapefruit, a sweet and nearly seedless variety with deep red flesh.

Cost to close: \$114 million, with \$92 million already spent. Stimulus money will pay about 60 percent of remaining \$22 million cost.

HIGH-FLUX BEAM REACTOR

Operated: 1965 to 1996
Permanently shut in 1999

Provided neutrons for research in material science, chemistry, biology and physics. Scientists conducted experiments with external neutron beams delivered through ports placed around reactor core.

Enriched uranium fueled the reactor. "Heavy" water — in which deuterium replaces the two hydrogen atoms — moderated fission and served as main coolant. Operated at 30, 40 or 60 megawatts.

Scientific advances

- Structure of cell's "protein factory" — the 16-part ribosome — first discerned here.
- New uses of radioactive isotopes developed for treating illnesses such as cancer, heart disease and arthritis.
- Advanced understanding of life span and decays of isotopes such as zinc-80, which astro-physicists use to study supernovas.
- Magnet experiments led to Nobel Prize-winning theories of cooperative ordering in large collections of atoms.
- Scientists using the high-flux beam reactor determined structures of the 23 amino acids, which make up every protein in every cell in living things.

Cost to close: \$64 million, with \$32 million already spent. Stimulus money will pay about 90 percent of the remaining cost, which excludes taking it apart after 65 years.



MEDICAL RESEARCH REACTOR

Operated: 1959 to 2000

The smallest of the lab's reactors, it was the first in the nation built just for medical research. Large objects were irradiated at one of the reactor's four faces; holes in another face permitted irradiation of samples and production of short-lived radioisotopes. Neutron streams traveled from two remaining ports to treatment rooms for animal and clinical studies.

Reactor operated at 3 megawatts but could generate 5 megawatts for short periods of time. Core was water cooled.

Scientific advances

- Boron neutron capture therapy, developed to treat a deadly form of brain cancer, was pioneered here.

Cost to close: Decommissioning plan and budget **not yet developed**.

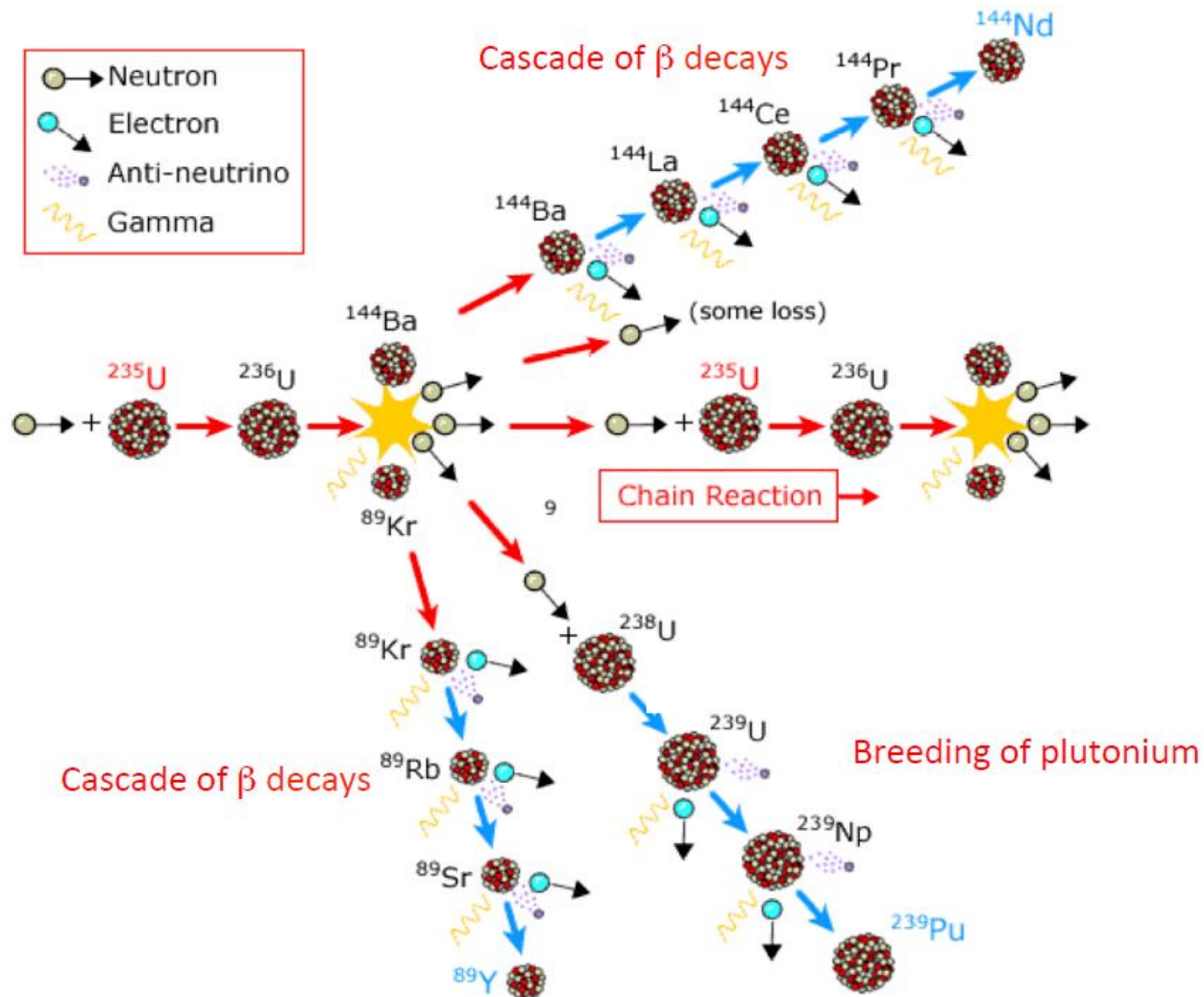
Source: Brookhaven National Laboratory

NEWSDAY, MONDAY, MAY 4, 2009 www.newsday.com

- ❑ research reactors typically ~10 MW
- ❑ BNL's past 3 reactors
 - BGRR, HFBR, BMRR

<https://www.bnl.gov/about/history/reactors.php>

Nuclear Reactor as (anti)Neutrino Source



- ❑ Pure $\bar{\nu}_e$ from beta decays of fission daughters
 - neutrino energy: < 10 MeV, peak ~ 4 MeV
- ❑ 2×10^{20} ν / sec / gigawatt
- ❑ free for physicists

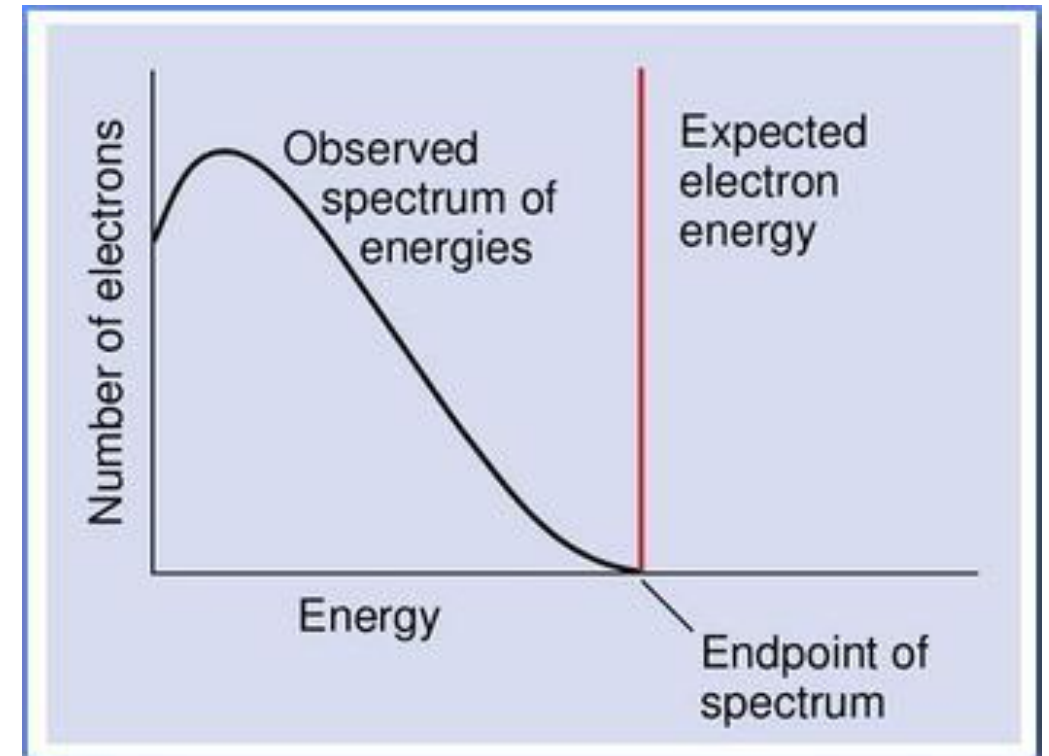
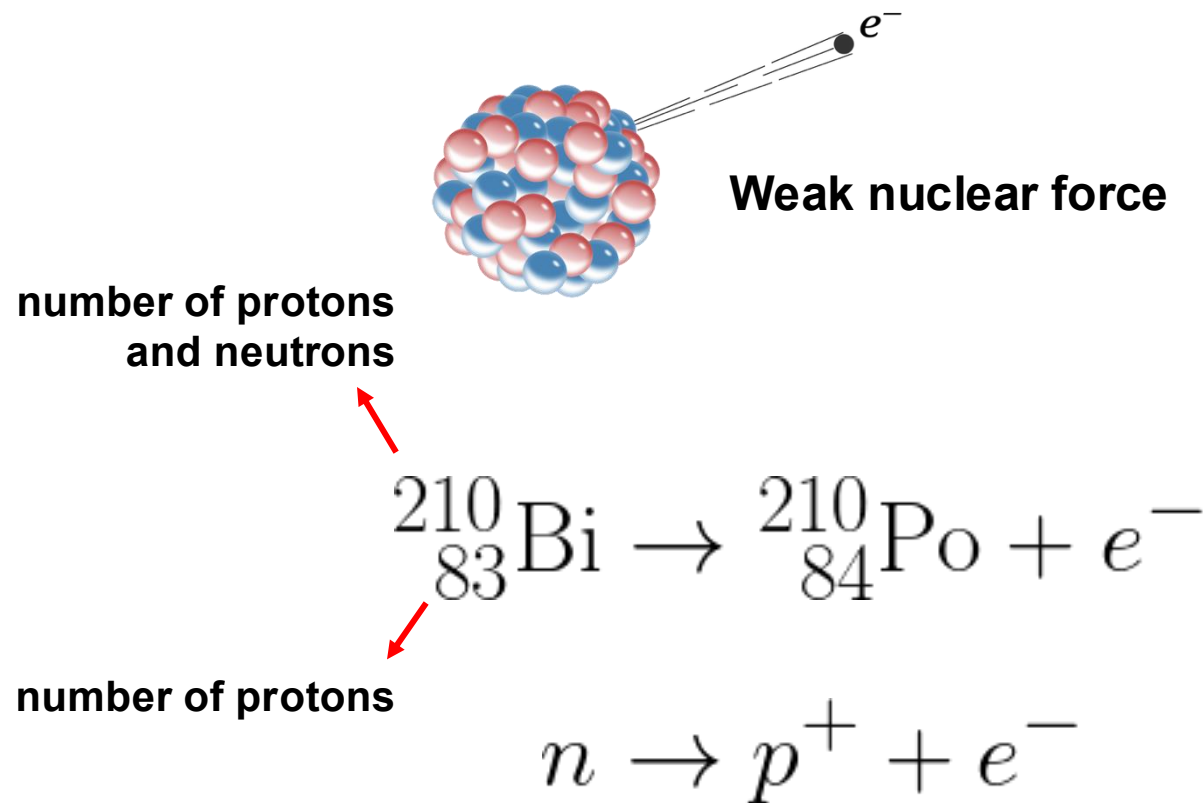
Fission fractions in a typical power reactor

^{235}U	55%
^{239}Pu	30%
^{238}U	10%
^{241}Pu	5%

Homework Problem 1

- ❑ How many antineutrinos are produced per second for a typical 3-gigawatt (thermal) commercial reactor?
 1. Each fission releases ~ 200 MeV energy. How many fissions are produced per second?
 2. Each fission produce ~ 6 antineutrinos on average from the beta-decay chains. How many antineutrinos are produced per second?

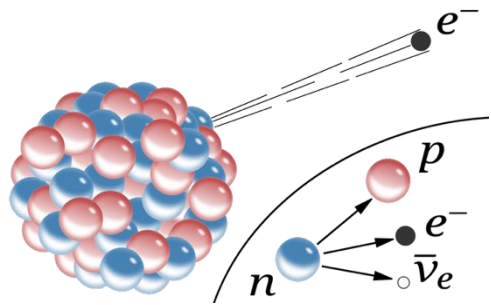
Beta Decay and Neutrino History



1899 – 1927

Rutherford, Meitner, Hahn, Chadwick, Ellis, Mott, *et. al*

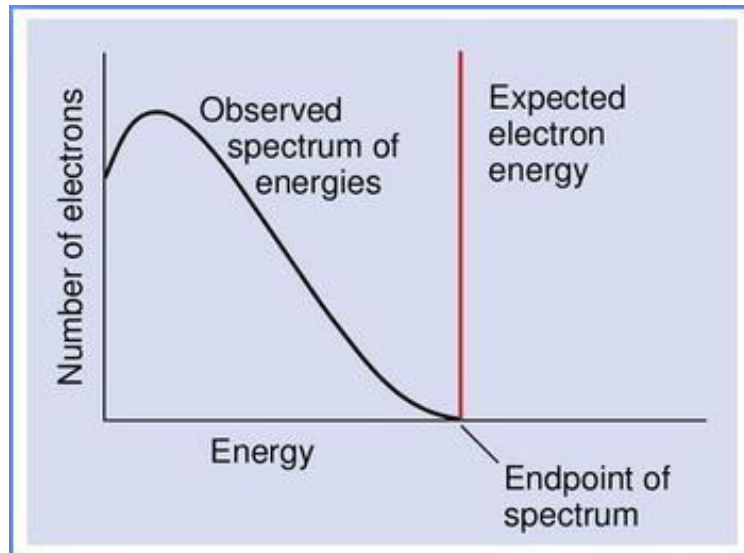
Neutrino: Proposed as a hypothetical particle



1930: Pauli's letter to physicists
at a workshop in Tübingen



Wolfgang Pauli



Dear Radioactive Ladies and Gentlemen,

....., I have hit upon a desperate remedy to save the "exchange theorem" of statistics and the law of conservation of energy. Namely, the possibility that there could exist in the nuclei electrically neutral particles, that I wish to call neutrons.... The mass of the neutrons should be of the same order of magnitude as the electron mass and in any event not larger than 0.01 proton masses. The continuous beta spectrum would then become understandable by the assumption that in beta decay a neutron is emitted in addition to the electron such that the sum of the energies of the neutron and the electron is constant.....

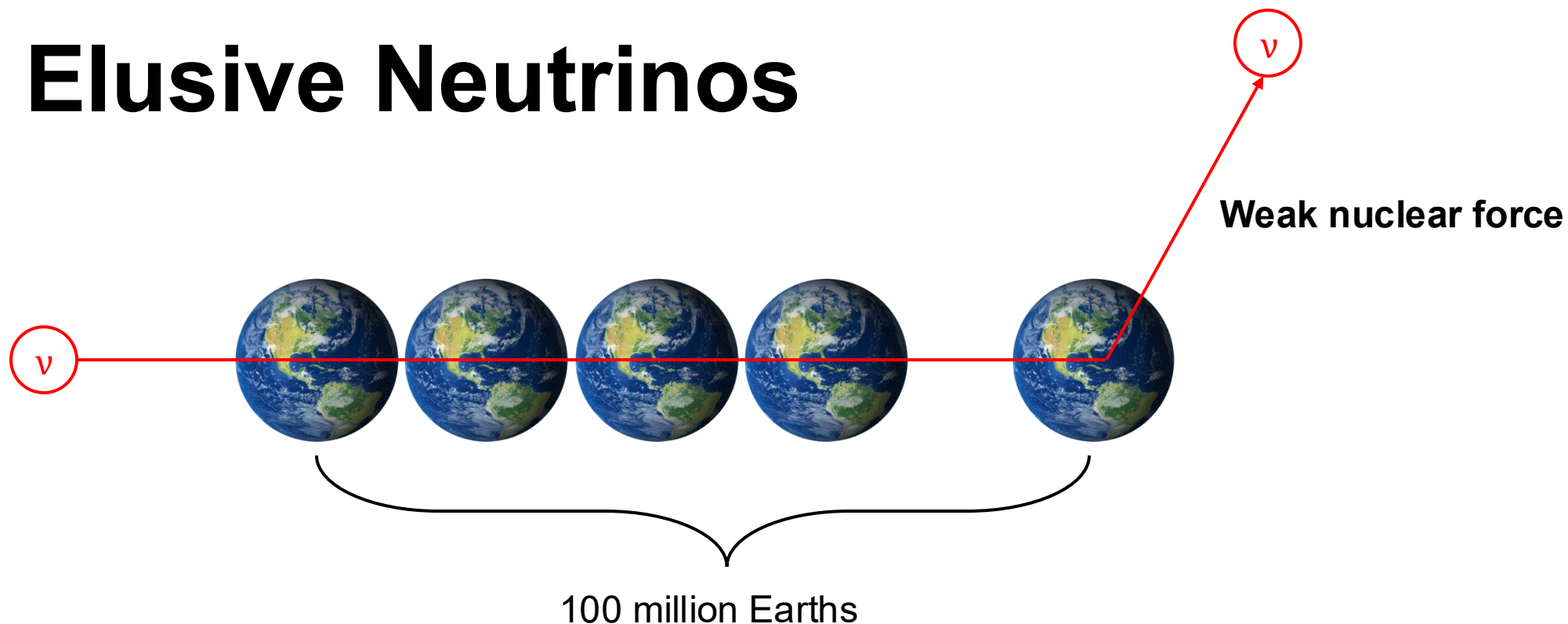
Unfortunately, I cannot appear in Tübingen personally since I am indispensable here in Zurich because of a ball on the night of 6/7 December. With my best regards to you, and also to Mr Back.

Your humble servant

. W. Pauli

"I have done a terrible thing. I have postulated a particle that cannot be detected."

The Elusive Neutrinos

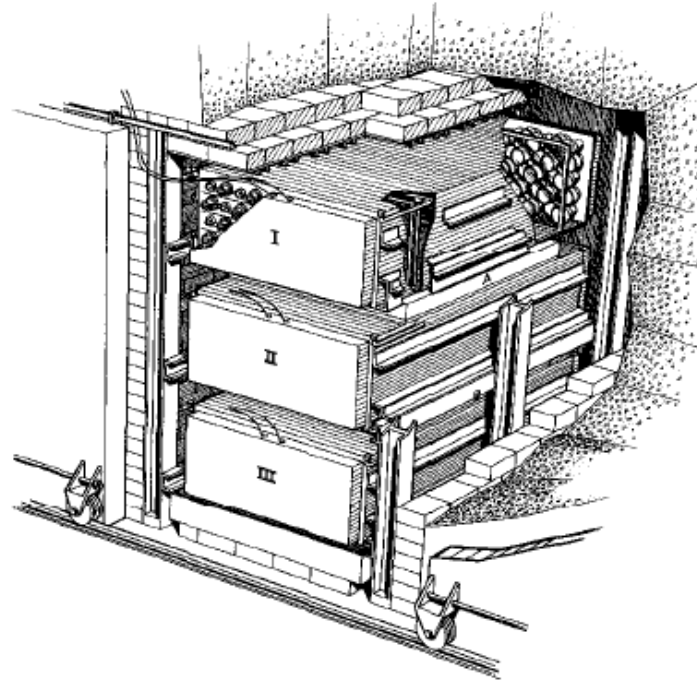
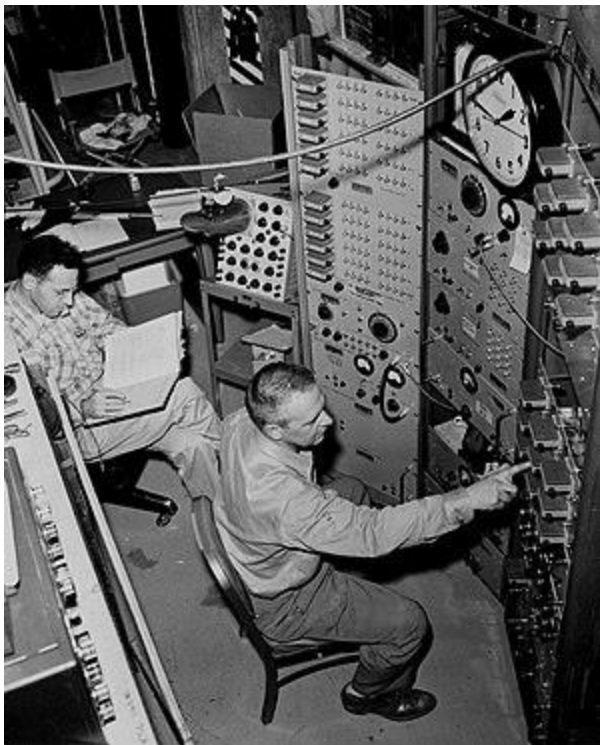


Neutrino detection requires:

- ❑ An intensive neutrino source: **a billion trillion ($\sim 10^{21}$) ν per second**
- ❑ A huge neutrino detector: **tons to kilotons of target material**
- ❑ **A distinctive method** to tell “neutrino interactions” from other backgrounds

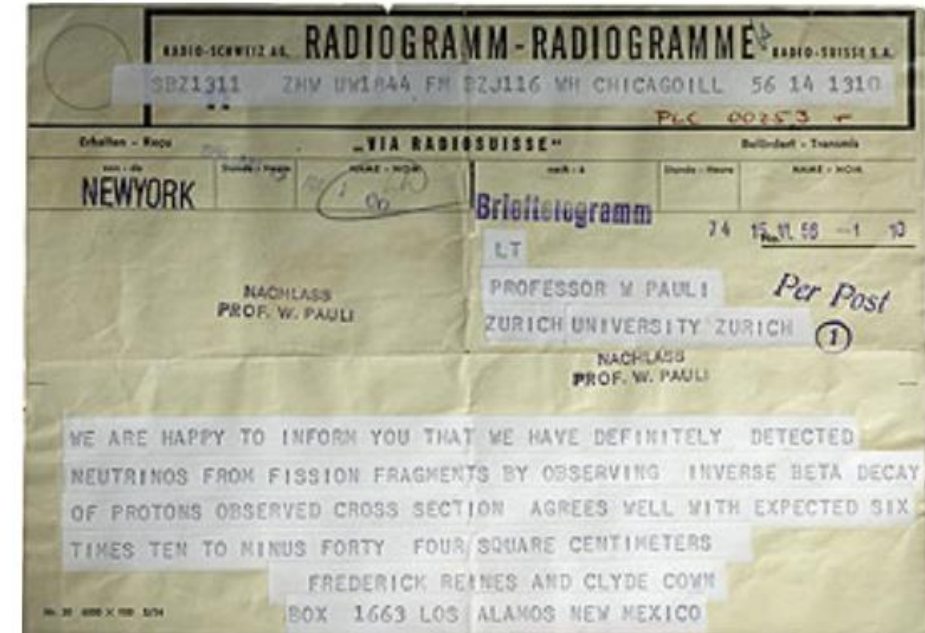
Neutrinos: First Detection

Frederick Reines and Clyde Cowan first detected (anti)neutrinos using the [Savannah River nuclear reactor](#) in South Carolina in **1956**. (26 years after Pauli's proposal)



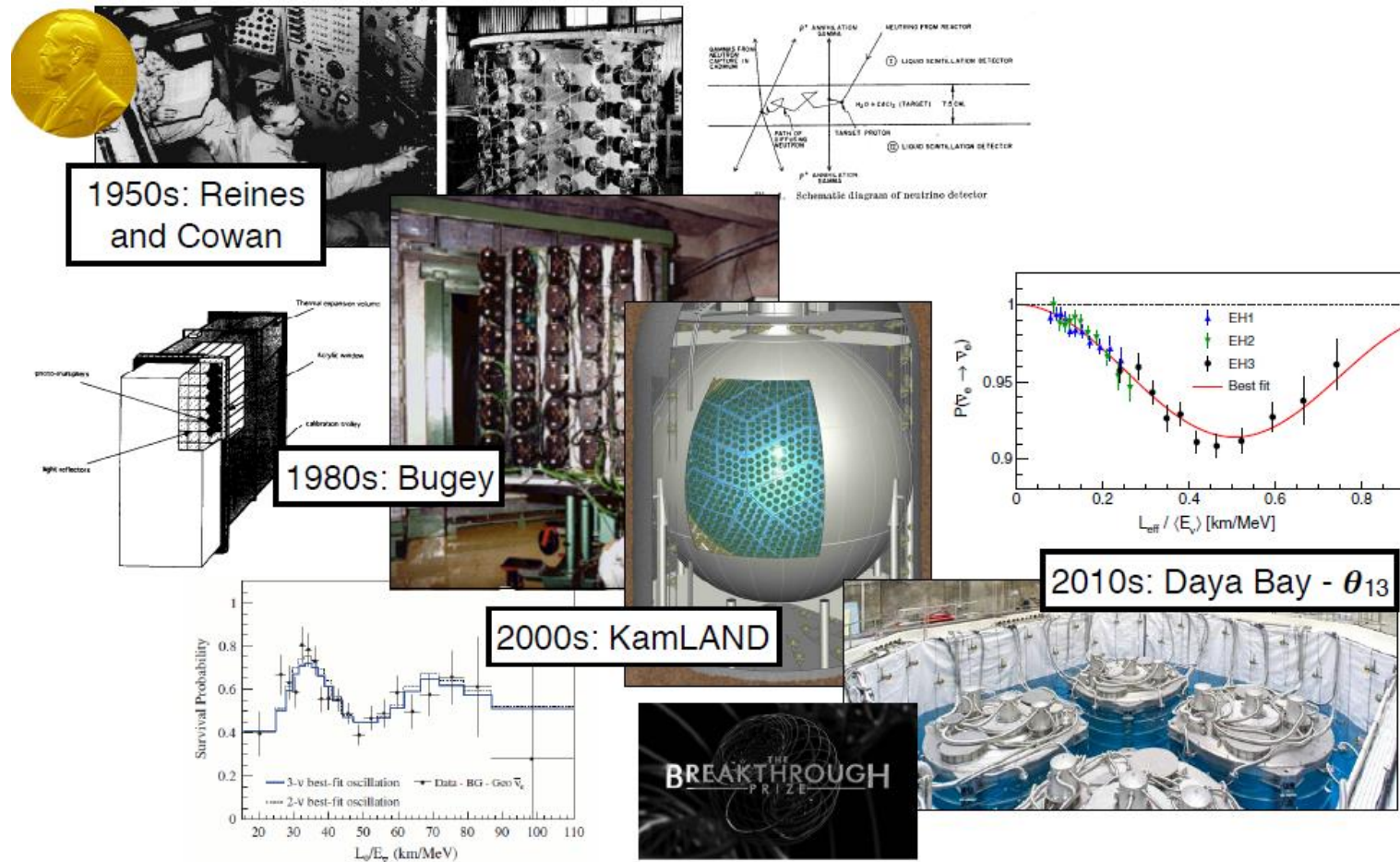
Target: Water + CdCl_2
Detector: Liquid Scintillator + PMTs

Reines and Cowan's
telegram to W. Pauli
(1956)



"We are happy to inform you that we have definitely detected neutrinos from fission fragments by observing inverse beta decay of protons"

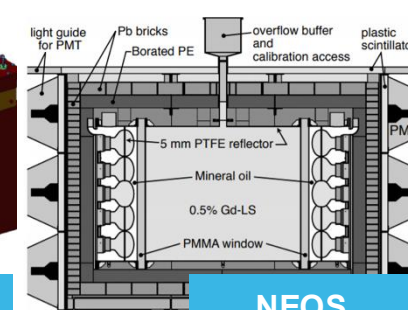
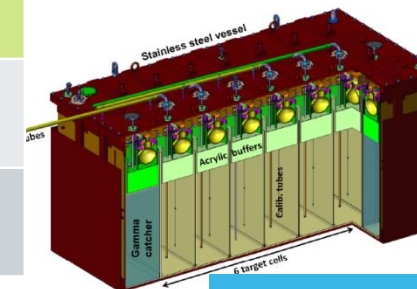
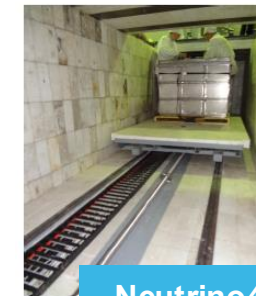
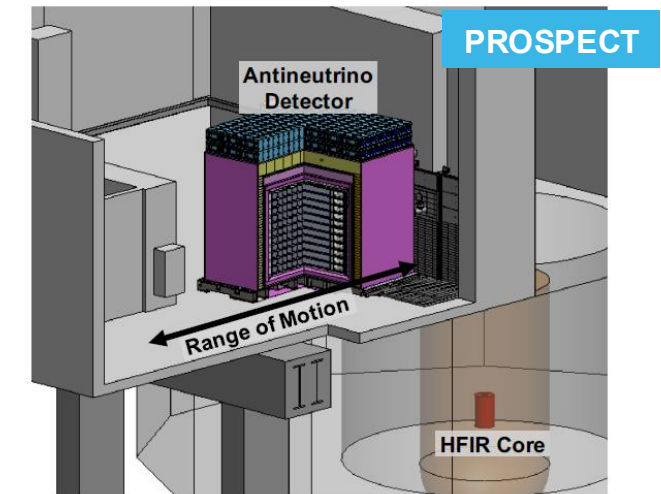
History of Reactor Neutrino Experiments



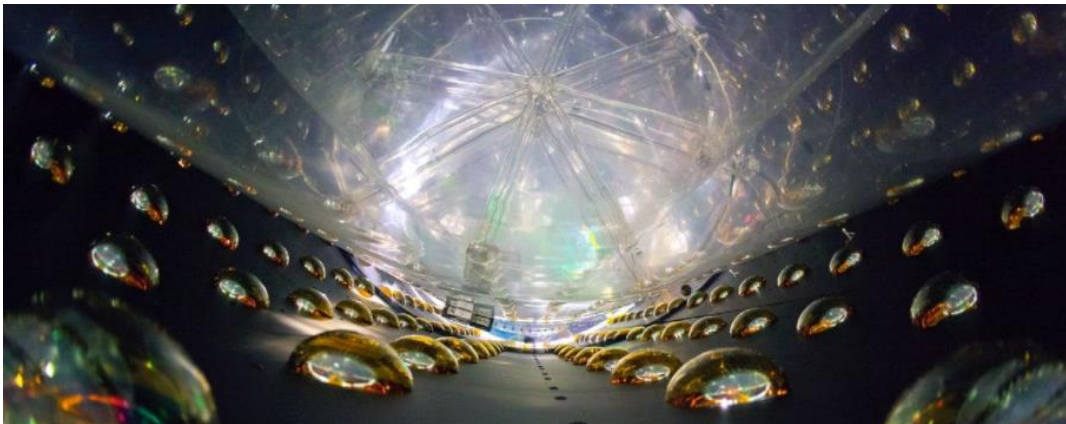
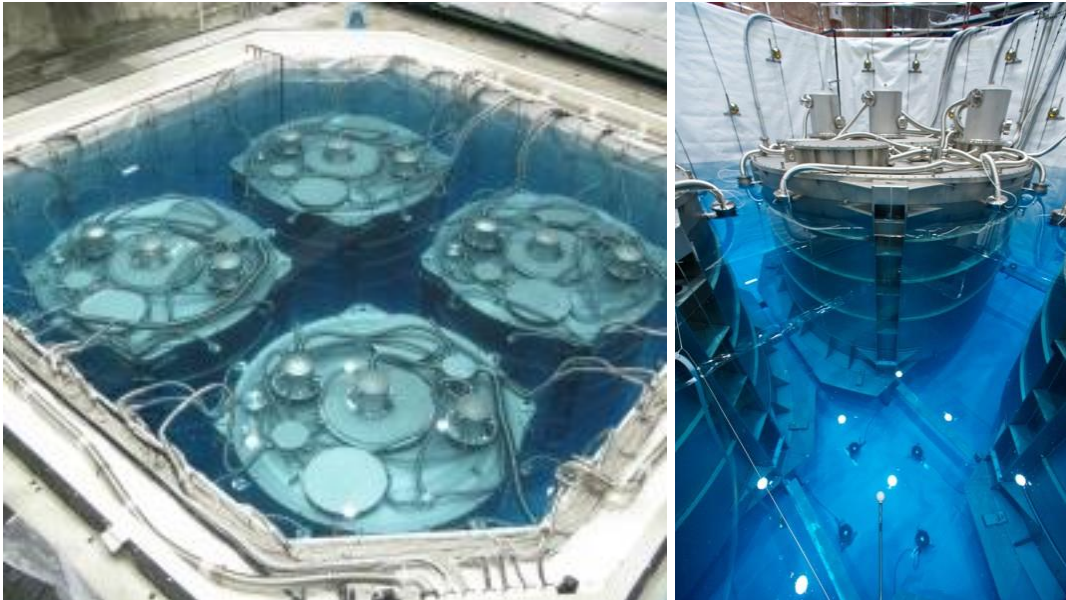
- ❑ Discovery of ν
- ❑ Solving solar ν problem on Earth
- ❑ Discovery of smallest oscillation angle θ_{13}
- ❑ Currently hold the best precision of
 - Δm_{21}^2 (KamLAND)
 - θ_{13} (Daya Bay)
- ❑ Comparable precision to accelerator-based experiments
 - $|\Delta m_{32}^2|$ (Daya Bay)

A lot of recent short-baseline reactor experiments (2010 – now)

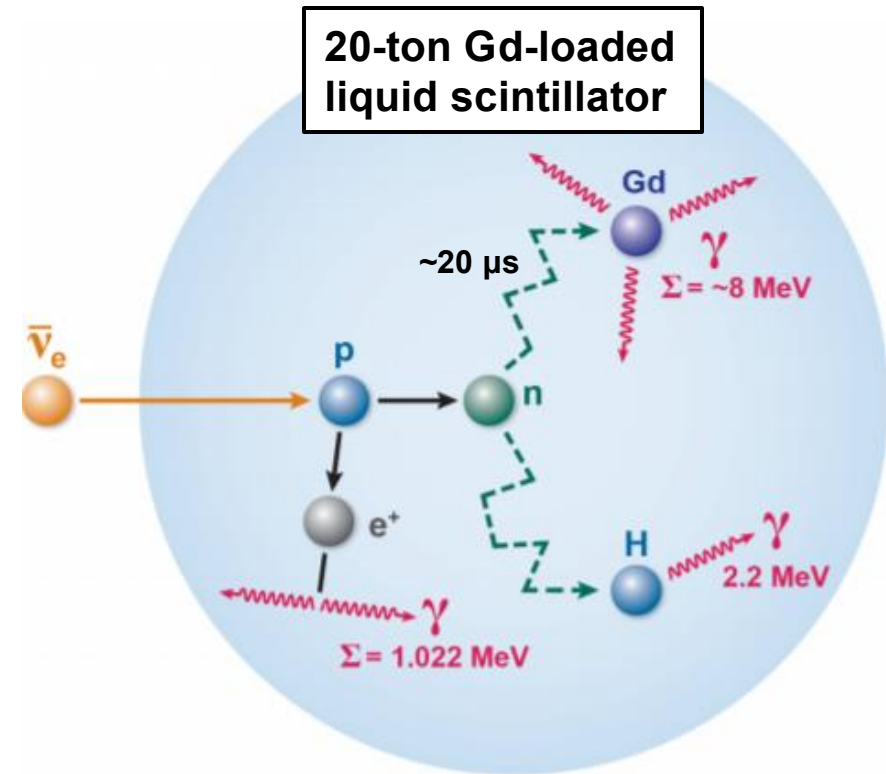
Experiment	Reactor	Baseline (m)	Overburden (m.w.e)	Mass (ton)	Segmentation	Energy res. (@ 1 MeV)
NEOS (South Korea)	LEU 2.8 GW	23.7	~20	1.0	none	5%
Nucifer (France)	HEU 70 MW	7.2	~12	0.6	none	10%
NEUTRINO4 (Russia)	HEU 100 MW	6 - 12	~10	0.3	2D	
DANSS (Russia)	LEU 3.1 GW	10.7 - 12.7	~50	1.1	2D	17%
STEREO (France)	HEU 58 MW	9 – 11	~15	1.6	1D 25 cm	8%
PROSPECT (USA)	HEU 85 MW	7 - 12	< 1	1.5	2D 15cm	4.5%
SoLid (UK Fr Bel US)	HEU 70 MW	6 - 9	~10	1.6	3D 5cm	14%
CHANDLER (USA)	HEU 75 MW	5.5 - 10	~10	1.0	3D 5cm	6%
NuLAT (USA)	HEU 20 MW	4	few	1	3D 5cm	4%



Detecting Reactor Neutrinos



Inverse Beta Decay: $\bar{\nu}_e + p \rightarrow e^+ + n$

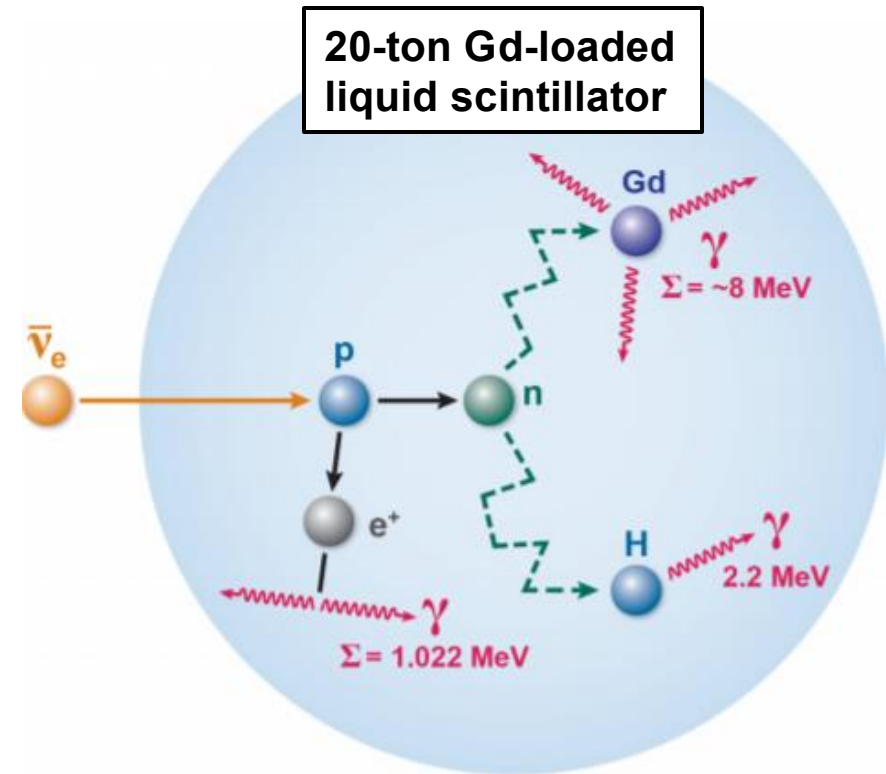
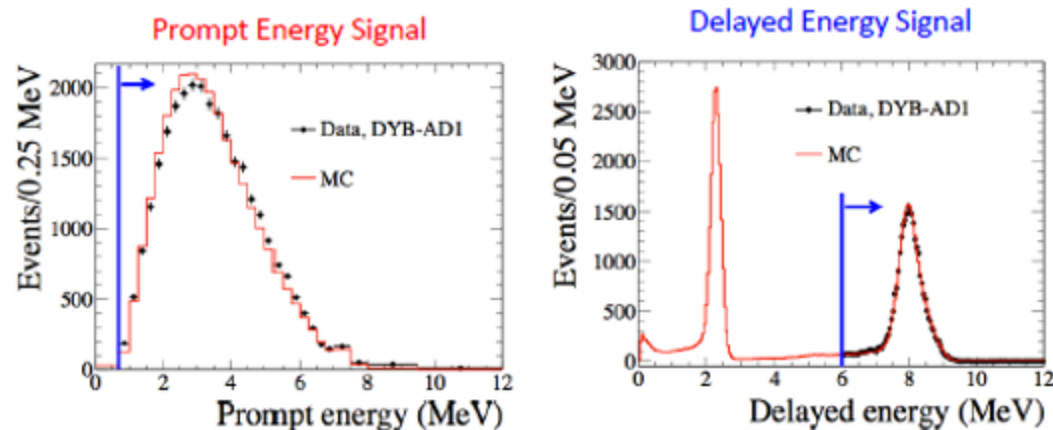
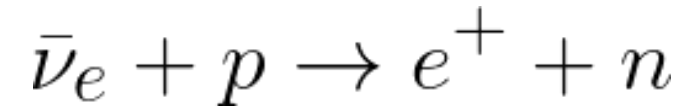


Daya Bay Reactor Neutrino Experiment

Detecting Reactor Neutrinos

Inverse Beta Decay (IBD)

- $E_{\text{threshold}} = 1.8 \text{ MeV}$
- 'Large' cross section $\sigma \sim 10^{-42} \text{ cm}^2$
- Distinctive coincidence signature in a large liquid scintillator detector



Daya Bay Reactor Neutrino Experiment

Reactor $\bar{\nu}_e$ Flux Prediction: Summation method

- Calculate each beta-decay spectrum using nuclear databases:



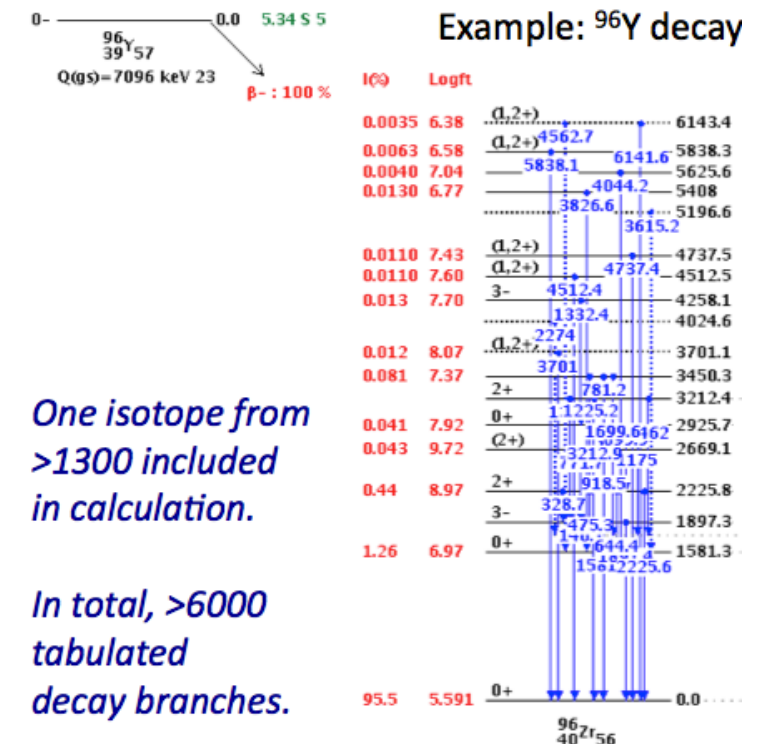
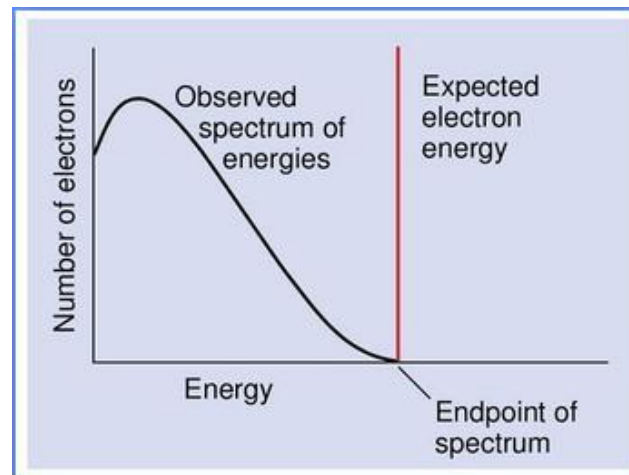
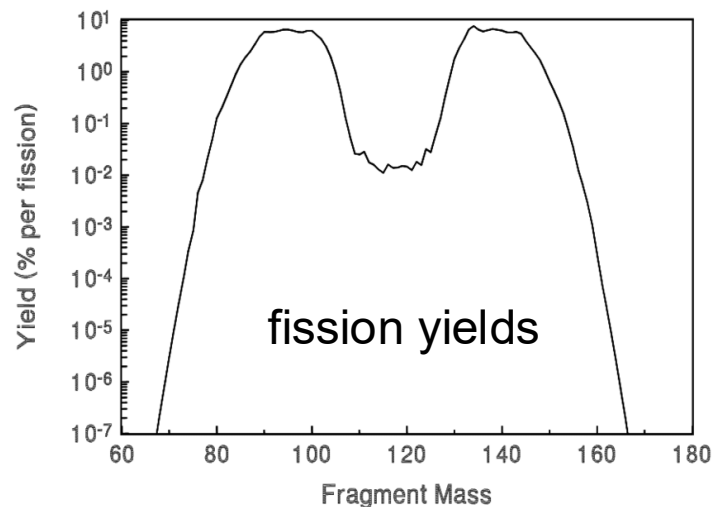
ENDF, JEFF, JENDL, CENDL, ROSFOND ...

$$\frac{d\phi_i}{dE_\nu} = \sum_n Y_n(Z, A, t) \cdot (\sum_m b_{n,m} \cdot P(E_\nu, E_0, Z)),$$

fission products

beta spectra
($E_\nu = E_0 - E_e$)

beta-decay branches



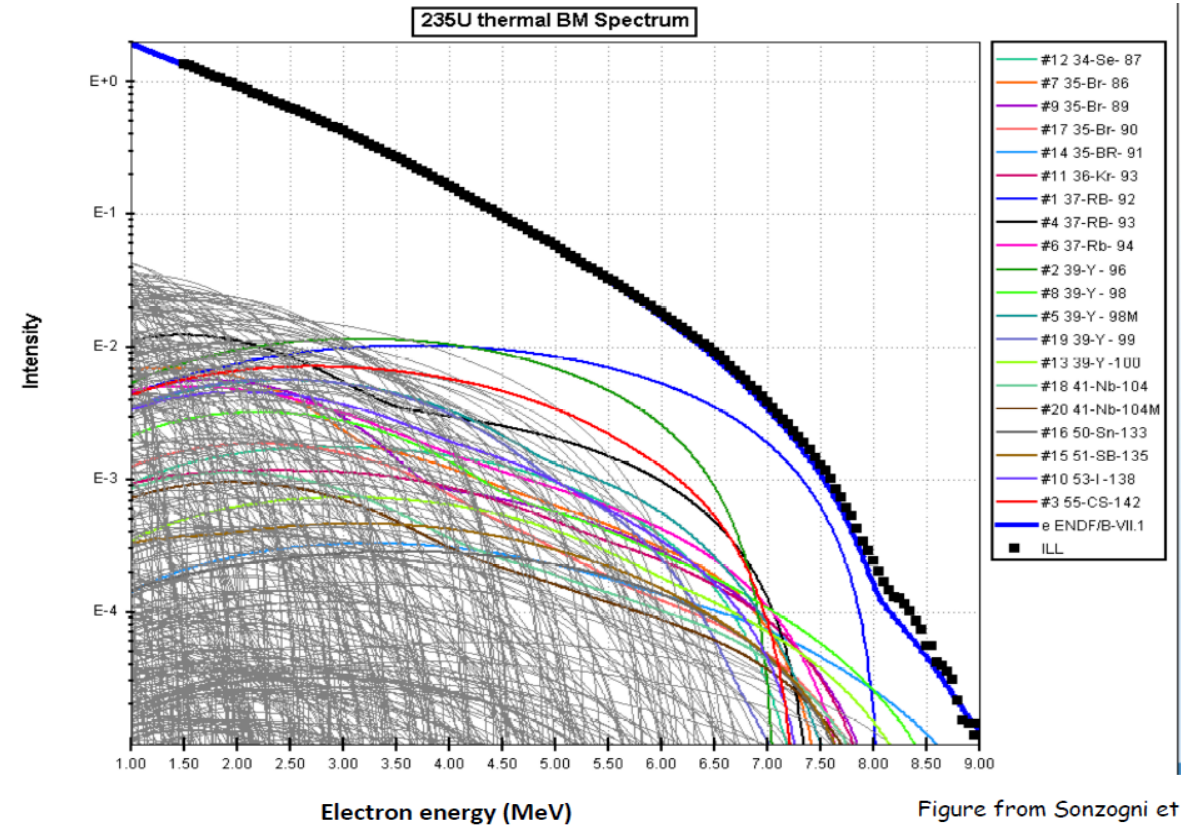
Reactor $\bar{\nu}_e$ Flux Prediction: Summation method

❑ Challenges

- **Incomplete databases** for beta-decay branches (~10% missing)
- Known systematic bias in some beta decay data with large Q-values (**pandemonium effect**)
- ~30% of beta decays are **forbidden decays** where shape corrections are necessary but not easy to calculate theoretically

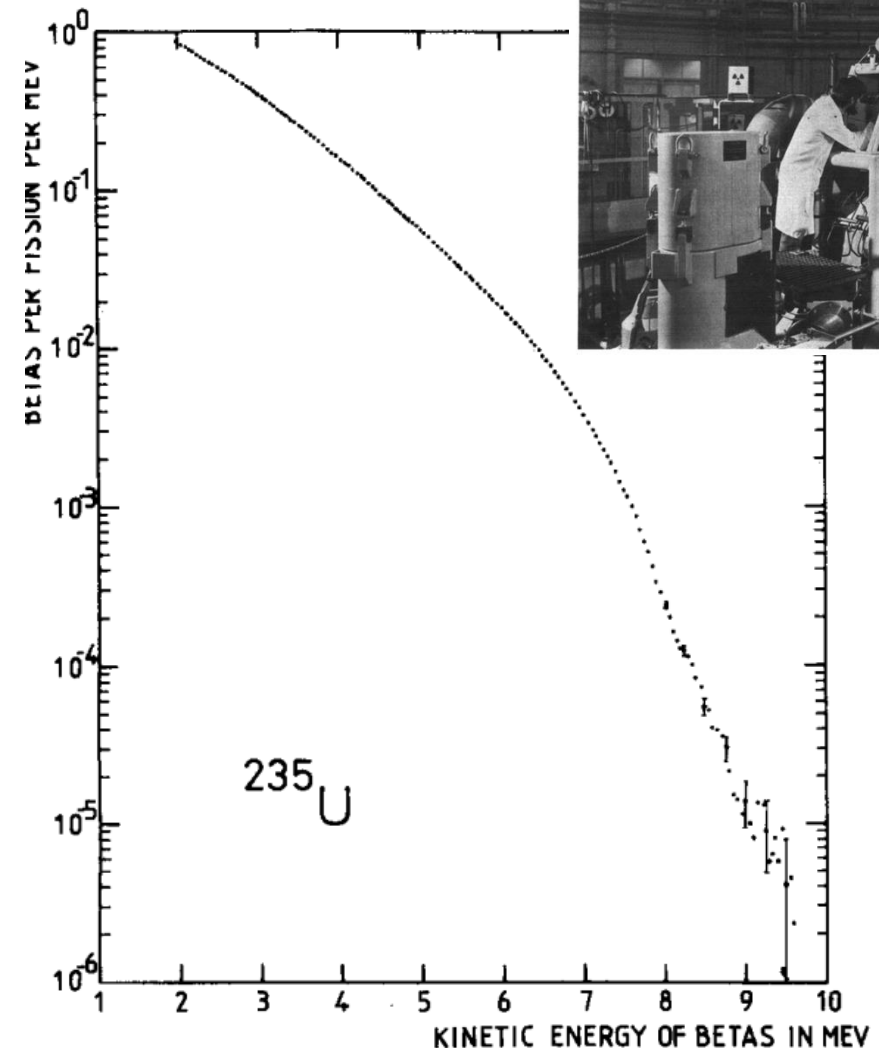
❑ Large uncertainty (~10%)

- Historically only used to predict ^{238}U flux (~10% fissions in a commercial reactor)
 - *Vogel et.al, PRC 24, 1543 (1981)*



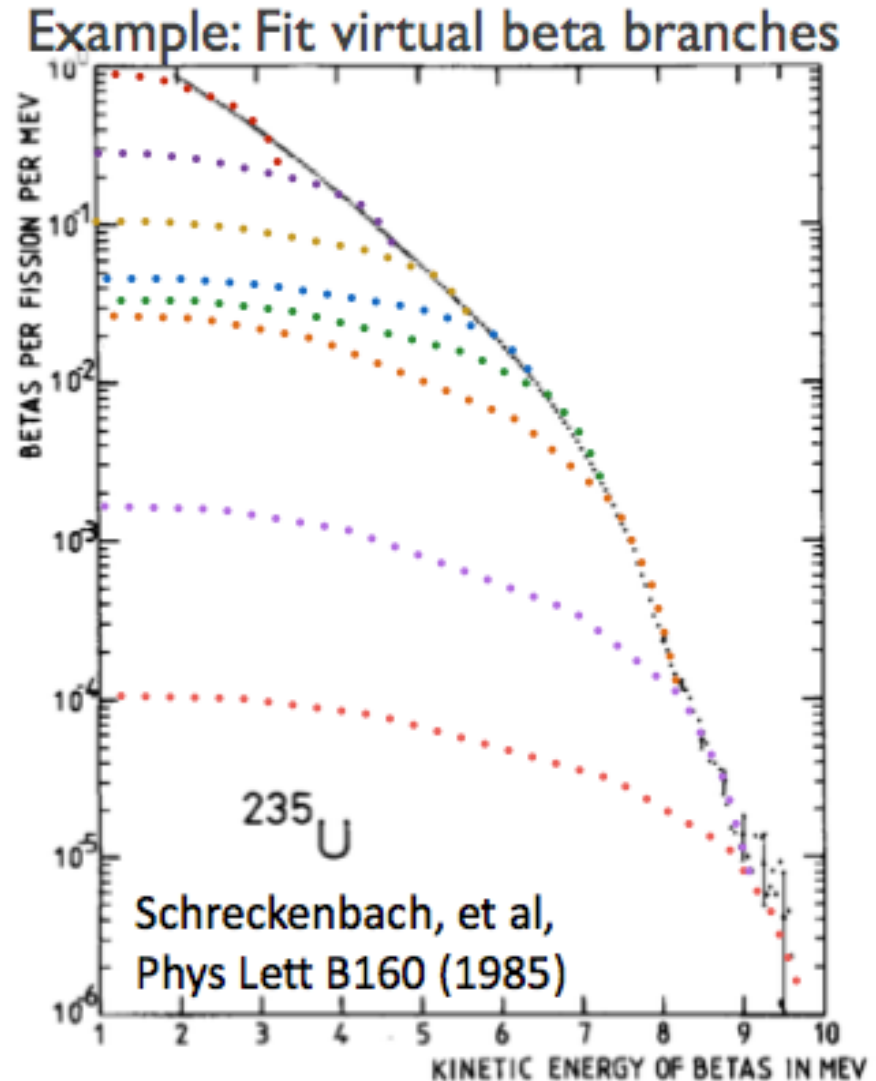
Reactor $\bar{\nu}_e$ Flux Prediction: Conversion method

- ❑ Experiments at ILL in Grenoble, France in the 1980s for ^{235}U , ^{239}Pu , ^{241}Pu
 - Irradiate fission isotope target (e.g. thin foil of $^{235}\text{UO}_2$) in a high flux of thermal neutrons for tens of hours.
 - Measure total outgoing beta-decay electron energy spectrum.
 - Used a high resolution, double focusing e-spectrometer “BILL”:
NIMA 154, 127 (1978)
 - Calibration with conversion electron sources (^{207}Pb , ^{197}Au , ^{113}Cd , ^{115}In)
 - High statistics in bins of 50 keV.
- ❑ ^{238}U was not measured (only fission with fast neutrons) until 2014 at FRM-II in Garching, Germany



Reactor $\bar{\nu}_e$ Flux Prediction: Conversion method

- ❑ Convert total electron spectrum to total antineutrino spectra with **fit to ~30 virtual beta-decay branches**
 - equidistant end-point energy
 - assume allowed beta-decay shape $P(E_\nu, E_0, Z)$
 - empirical function of Z vs Q -value
- ❑ Does not rely on fission yields or beta decay data. Considered much more precise and can reach **~2% uncertainty**
- ❑ Standard reactor $\bar{\nu}_e$ flux model (**ILL-Vogel model**)
 - ILL conversion for ^{235}U , ^{239}Pu , ^{241}Pu
 - Vogel's summation for ^{238}U
 - agree with ~20 reactor flux measurements from 1980 -1990s



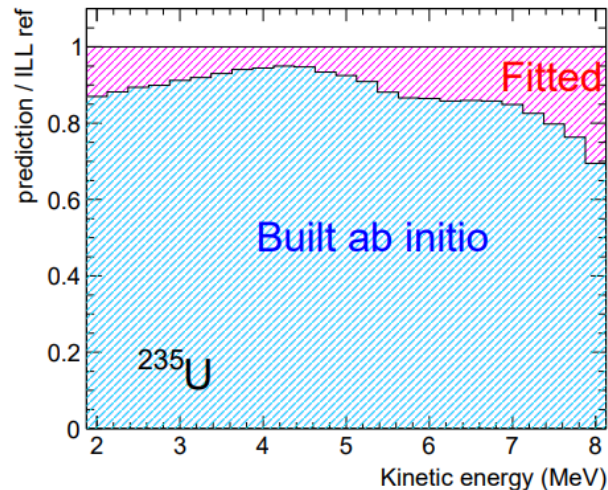
Re-evaluation: Huber-Mueller Model

[Submitted on 13 Jan 2011 (v1), last revised 11 Mar 2011 (this version, v3)]

Improved Predictions of Reactor Antineutrino Spectra

Th. A. Mueller, D. Lhuillier, M. Fallot, A. Letourneau, S. Cormon, M. Fechner, L. Giot, T. Lasserre, J. Martino, G. Mention, A. Porta, F. Yermia

- ❑ Hybrid method: **+3%**
 - Updated summation calculation from the ENSDF database (for ^{235}U , ^{239}Pu , ^{241}Pu , ^{238}U)
 - Conversion method for the missing 10% contribution (for ^{235}U , ^{239}Pu , ^{241}Pu)
 - Correct for non-equilibrium effect

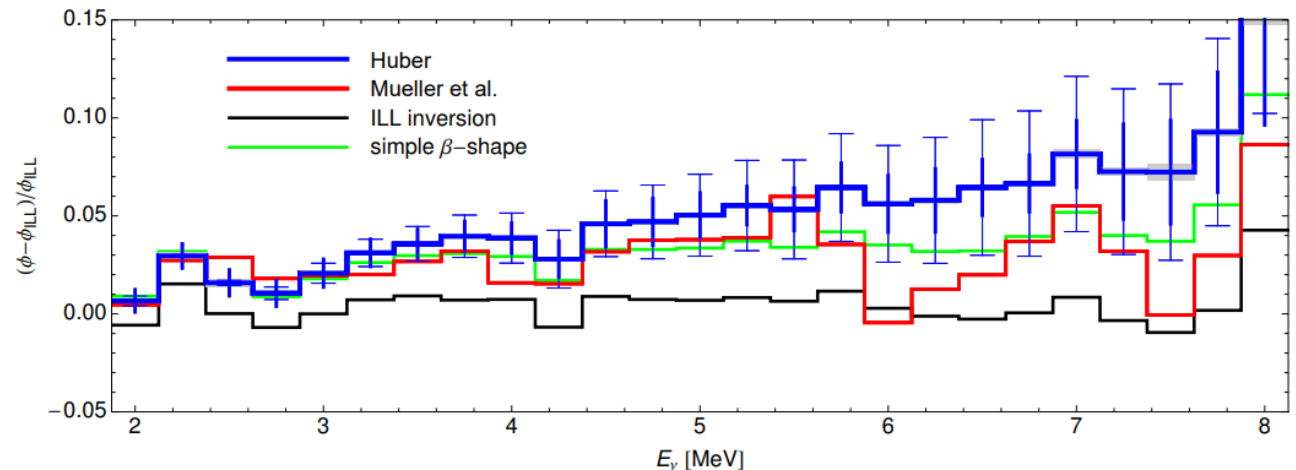


[Submitted on 3 Jun 2011 (v1), last revised 17 Jan 2012 (this version, v4)]

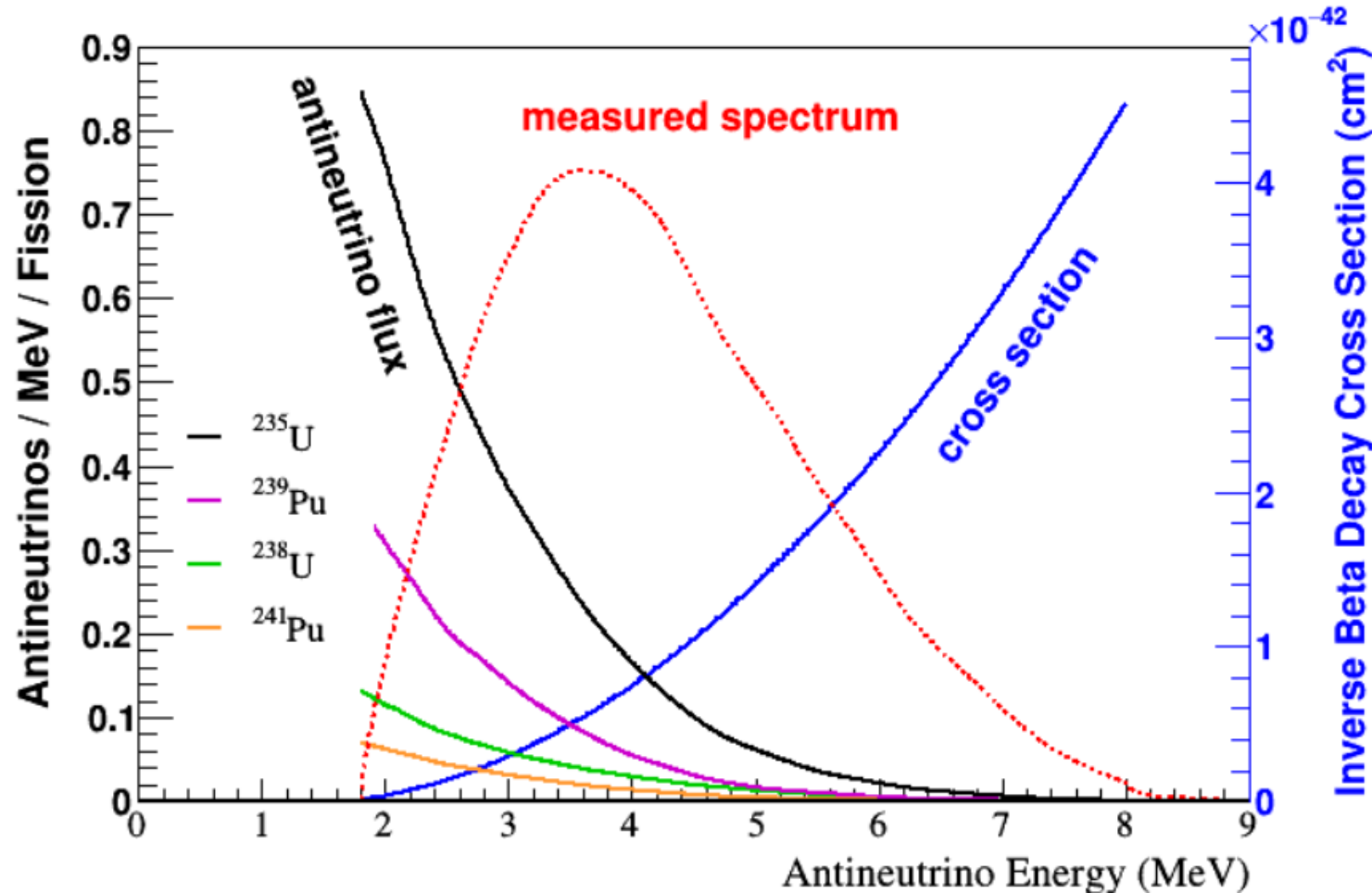
On the determination of anti-neutrino spectra from nuclear reactors

Patrick Huber

- ❑ Improved conversion method using ILL data (for ^{235}U , ^{239}Pu , ^{241}Pu):
 - Reevaluated nuclear effects in correcting the beta-spectrum shape **+3%**
 - effective Z as a function of Q-value for virtual branches
 - finite-size, radiative correction, weak magnetism
 - Non-equilibrium effect **+1-2%**
 - New neutron lifetime measurement **+1%**



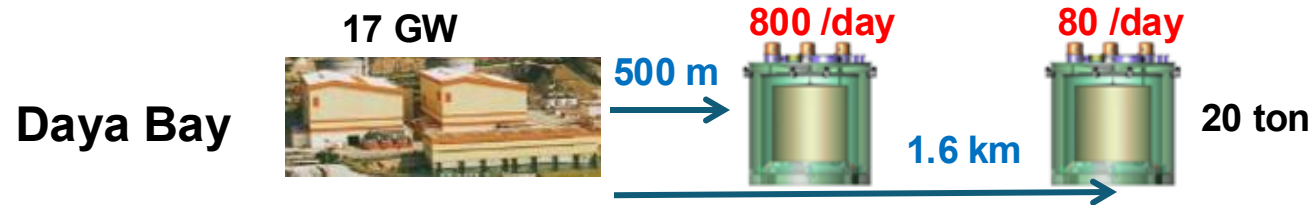
Antineutrino Signal Prediction



- Antineutrino flux predicted from reactor theory (with large uncertainties)
- IBD threshold: $E > 1.8$ MeV (why?)
 - only about 1/3 of the reactor antineutrinos can be detected
- Flux-weighted cross section $\sim 3 \times 10^{-43} \text{ cm}^2$

Event Rate

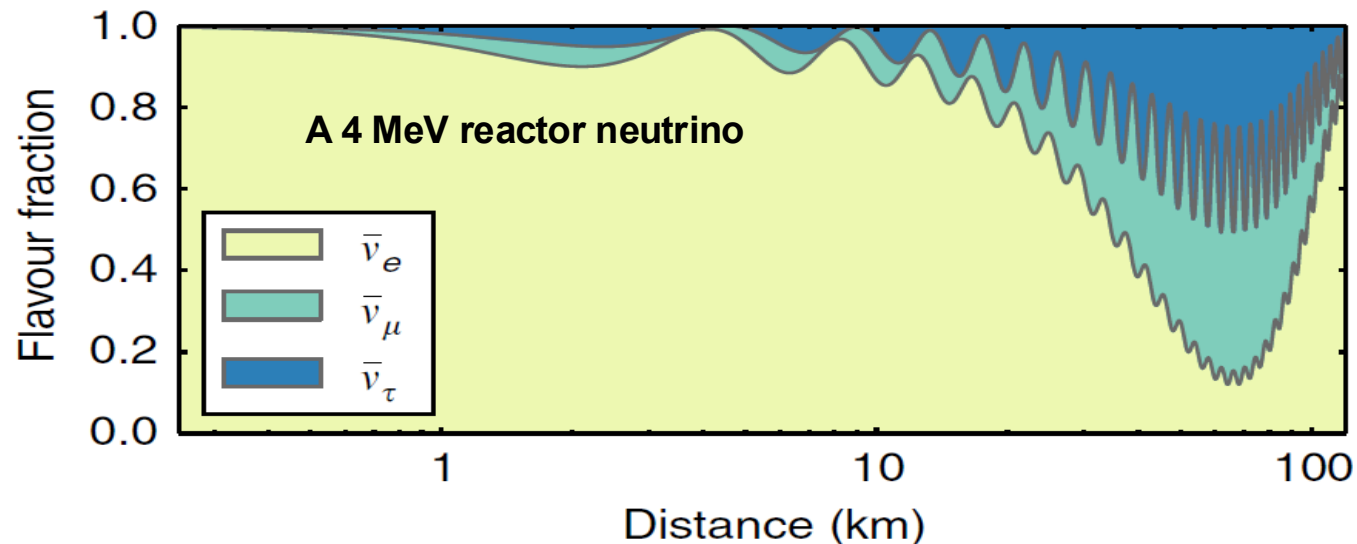
- Depending on the power of the reactors, size of the detectors, and the distance between them



Homework Problem 2

- Daya Bay's antineutrino detector (AD) is a 20-ton liquid scintillator detector. The far ADs are placed at ~ 1.6 km away from reactors with a total power of ~ 17 GW. How many inverse beta decay (IBD) reactions are expected per day in each far AD?
1. The hydrogen mass fraction in the AD is $\sim 12\%$. How many free protons (from hydrogen) are there in each AD?
 2. How many reactor antineutrinos per cm^2 per second is expected at the AD? (this is referred as the "flux")
 3. The average IBD cross section is $\sim 3 \times 10^{-43} \text{ cm}^2$, calculate **event rate = flux * cross section * number of protons**, assuming 100% detection efficiency. (remember that IBD can only detect 1/3 of all reactor antineutrinos)

Neutrino Oscillations with Reactors



$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U_{\text{PMNS}} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

↓

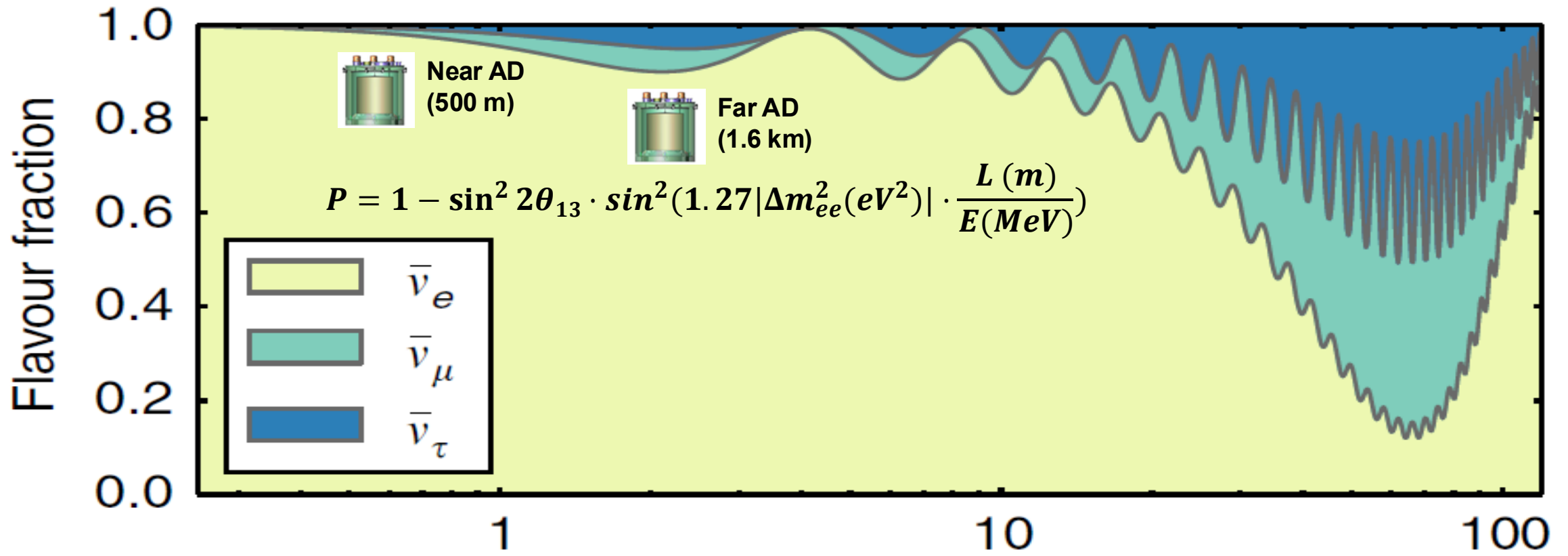
$$\begin{pmatrix} \cos \theta_{12} & \sin \theta_{12} & 0 \\ -\sin \theta_{12} & \cos \theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \cos \theta_{13} & 0 & \sin \theta_{13} e^{-i\delta} \\ 0 & 1 & 0 \\ -\sin \theta_{13} e^{i\delta} & 0 & \cos \theta_{13} \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & \sin \theta_{23} \\ 0 & -\sin \theta_{23} & \cos \theta_{23} \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{-i\alpha_1/2} & 0 \\ 0 & 0 & e^{-i\alpha_2/2} \end{pmatrix}$$

Solar /
Long baseline reactor

Short baseline reactor /
Long baseline accelerator

Atmospheric /
Long baseline accelerator

Neutrinoless
double beta decay



$$\frac{N_{far}}{N_{near}} = \left(\frac{N_{p,f}}{N_{p,n}} \right) \cdot \left(\frac{L_n}{L_f} \right)^2 \cdot \left(\frac{\epsilon_f}{\epsilon_n} \right) \cdot \left(\frac{P_{survival}(E, L_f)}{P_{survival}(E, L_n)} \right)$$

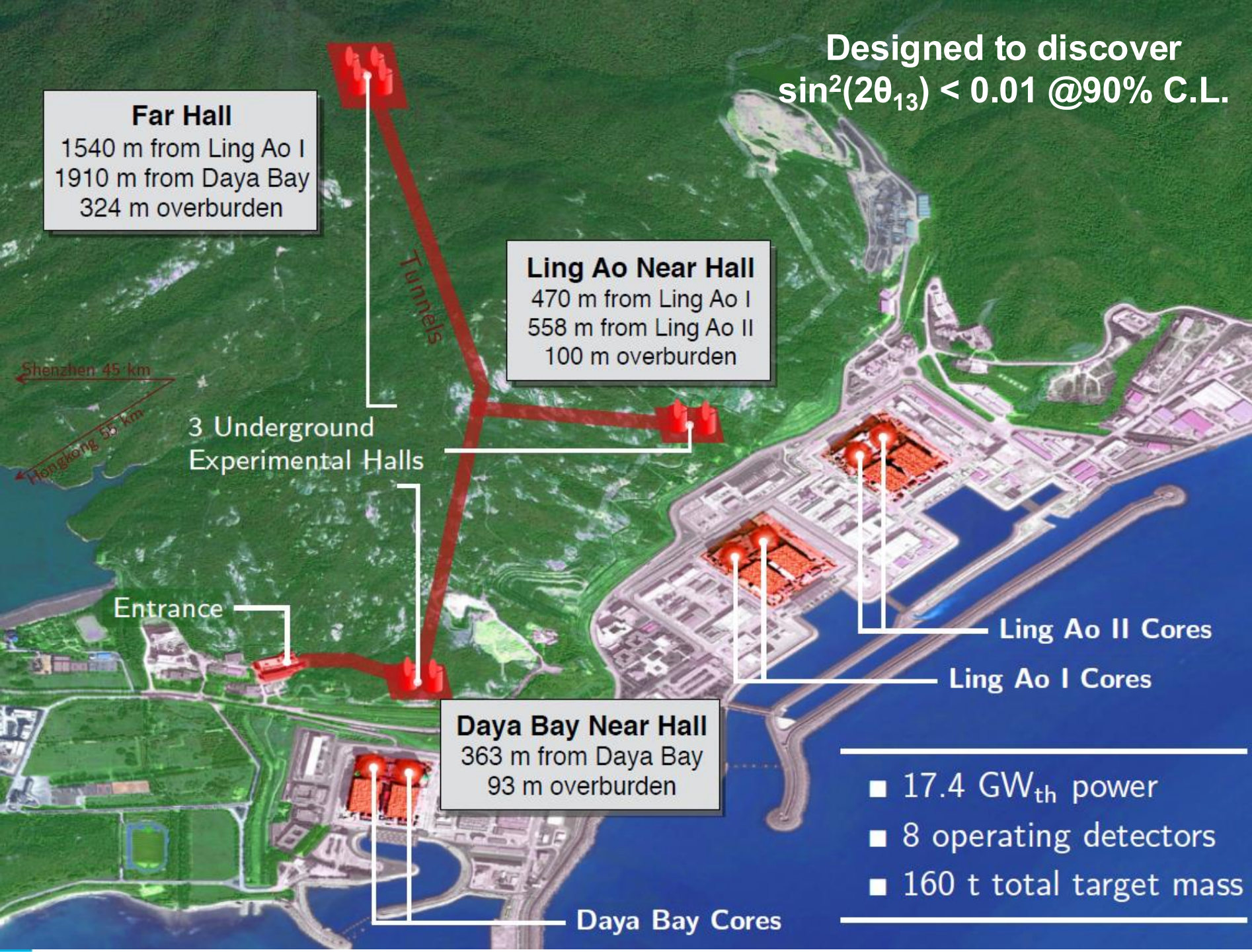
Far/Near
Neutrino
Ratio

Detector
Target
Mass

Distance
from
Reactor

Detector
Efficiency

Survival Probability
(θ_{13})



Statistics

- powerful reactors (17.4 GW_{th}) + large detectors (80 ton at Far site)

Systematics

□ Reactor

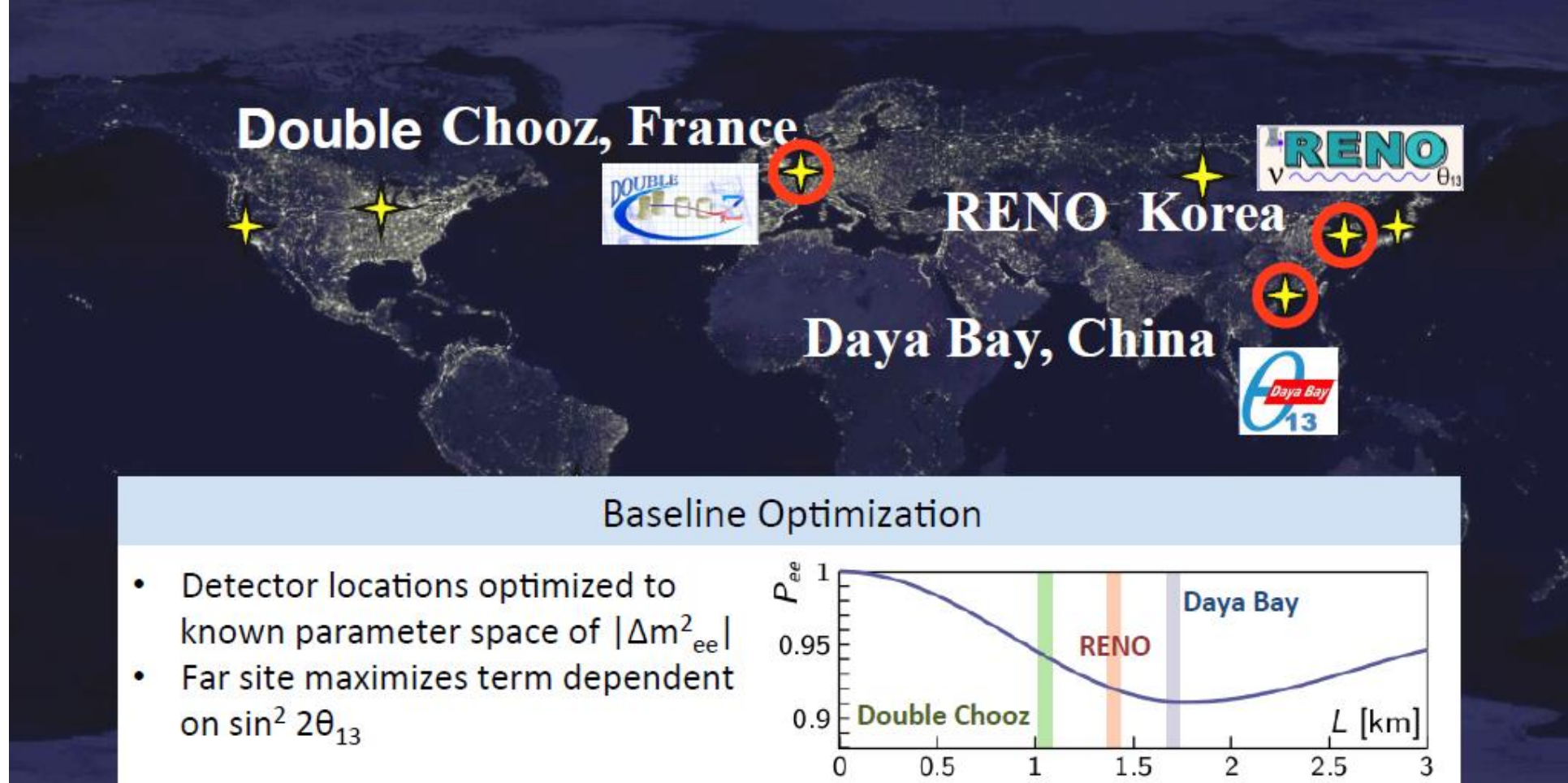
- Far/Near relative measurement

□ Detector

- multiple functionally identical detectors (4 Near + 4 Far)

□ Background

- deep underground (860 m.w.e at far site)



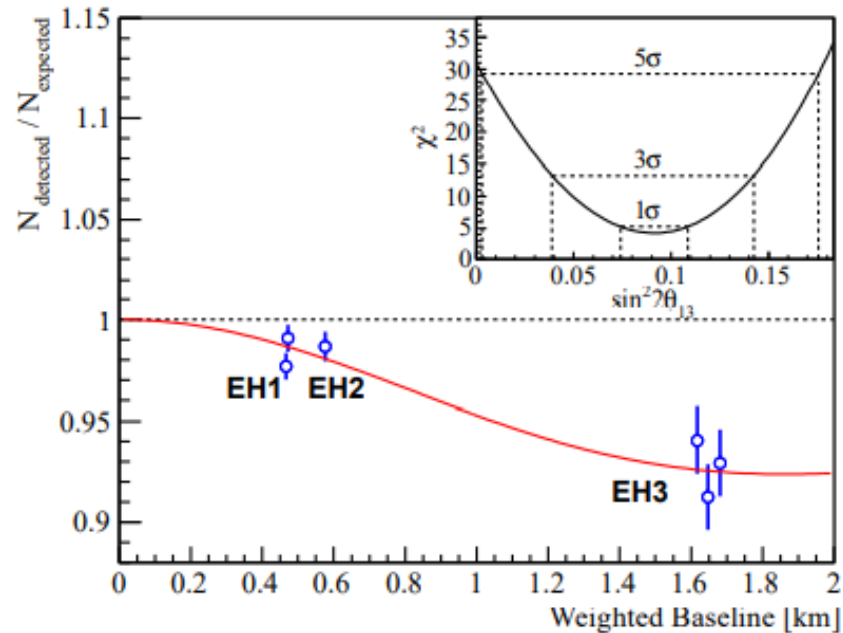
Go strong, big and deep!

	Reactor [GW _{th}]	Target [tons]	Depth [m.w.e]
Double Chooz	8.6	16 (2 × 8)	300, 120 (far, near)
RENO	16.5	32 (2 × 16)	450, 120
Daya Bay	17.4	160 (8 × 20)	860, 250
	Large Signal		Low Background



Discovery of non-zero θ_{13} at 5.2σ

- 2011/12/24 – 2012/2/17 (55 days)
- 6 detectors in operation first



Phys. Rev. Lett. 108, 171803 (2012)

In fact, in the **first 5 days** we already knew that θ_{13} is large from the data. In the homework I'll give you all the inputs to do a simplified analysis.

End of operation ceremony (Dec 24, 2011 - Dec 12, 2020)



BNL virtual mini-symposium: The Daya Bay Reactor Neutrino Experiment and the Discovery of Non-zero Theta13

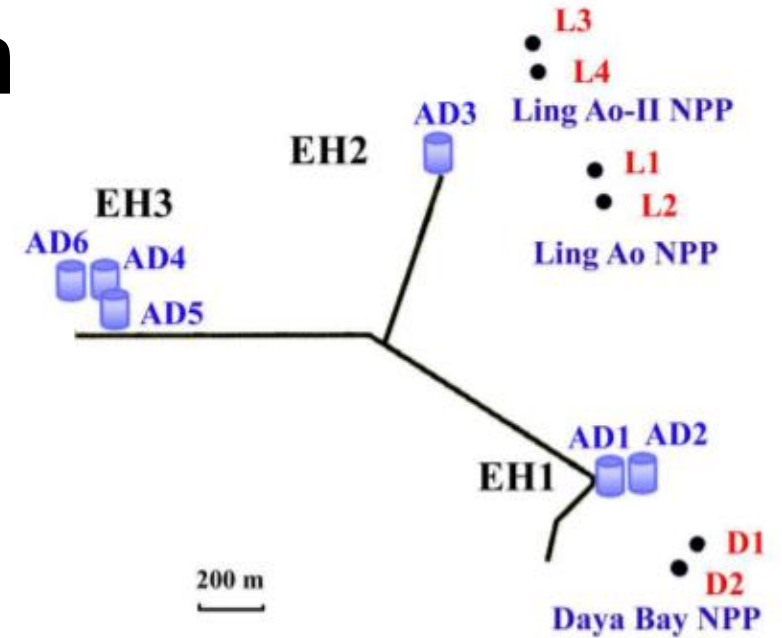
<https://indico.bnl.gov/event/9947/>

Homework Problem 3

- ☐ How to discover the smallest neutrino oscillation with 5 days of Daya Bay reactor neutrino data?

Reactor and Detector Location

Reactor	D1	D2	L1	L2	L3	L4
x (m)	43.0	-44.6	856.0	792.3	1143.6	1076.5
y (m)	-7.0	6.9	830.9	767.9	1206.1	1138.5
z (m)	-12.0	-12.0	-12.0	-12.0	-12.0	-12.0



AD	1	2	3	4	5	6
x (m)	94.5	97.8	584.1	-254.3	-259.5	257.3
y (m)	350.2	345.2	1216.2	1892.6	1889.6	897.8
z (m)	-20.0	-20.0	-16.6	-15.4	-15.4	-15.4

- All reactor cores operated at approximately **equal power** for the 5 days
 - **L2 was powered off** during the 5 days

Summary of event selection for the first 5 days

	EH1 AD1	EH1 AD2	EH2 AD1	EH3 AD1	EH3 AD2	EH3 AD3	
IBD Candidates	3278	3194	2193	338	350	348	Signal + Backgrounds
DAQ Live Time [days]	5.39	5.39	4.97	5.20	5.20	5.20	
Accidentals	60.9	59.6	49.3	20.5	19.4	19.3	} Backgrounds
Li9	43	42	28	4	4	4	
Fast Neutron	6	6	6	0.6	0.6	0.6	
Efficiency	0.8144	0.8120	0.8510	0.9515	0.9501	0.9508	

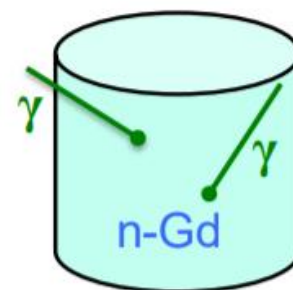
- ❑ Calculate the signal rate per day after efficiency correction
 - EH1-AD1: $(3278 - 60.9 - 43 - 6) / 0.8144 / 5.39 = 721.7$ events/day
- ❑ Calculate the statistical error on the signal rate
 - EH1-AD1: $\sqrt{3278} / 0.8144 / 5.39 = 13.0$ events/day

Background to IBD

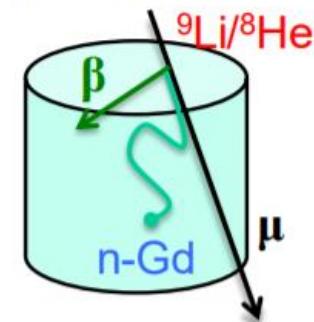
- ❑ **Accidentals**: statistically calculate from uncorrelated singles
- ❑ **Li9 / He8**: measure time distribution of after-muon events
- ❑ **Fast neutron**: measure energy spectrum from AD/water/RPC tagged muon events

Background	Near	Far	Uncertainty
Accidentals	1.4%	2.3%	negligible
Li-9 / He-8	0.4%	0.4%	~30%
Fast neutron	0.1%	0.1%	~30%

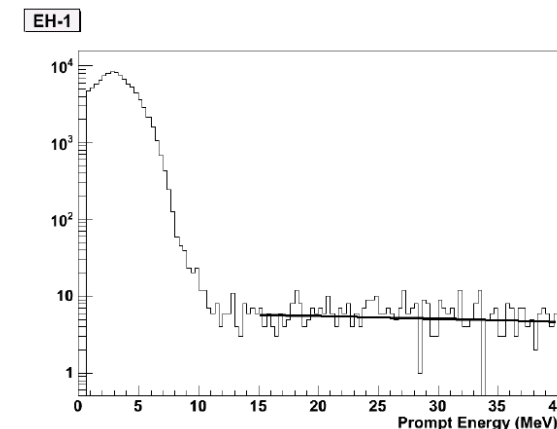
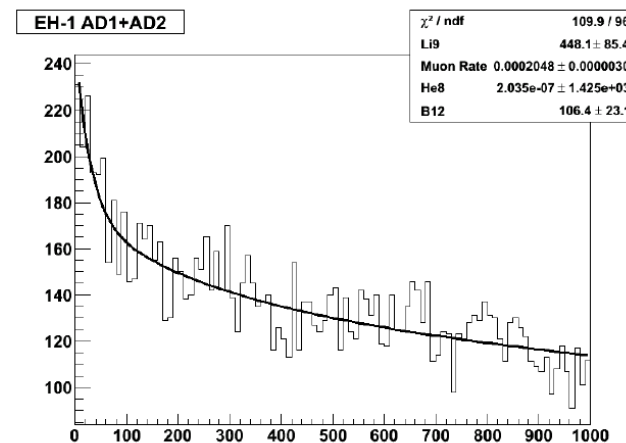
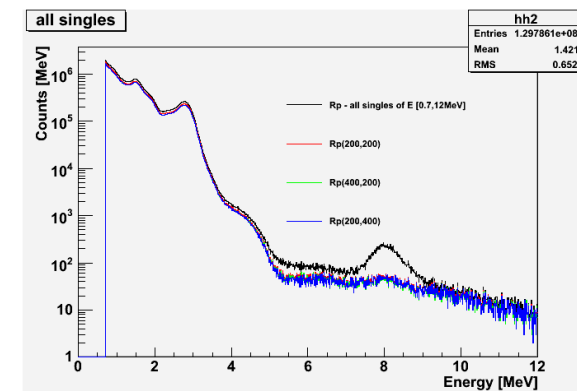
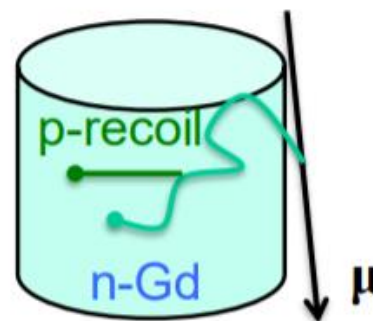
Accidentals



β -n isotopes



Fast neutrons



□ How to discover the smallest neutrino oscillation with 5 days of Daya Bay reactor neutrino data?

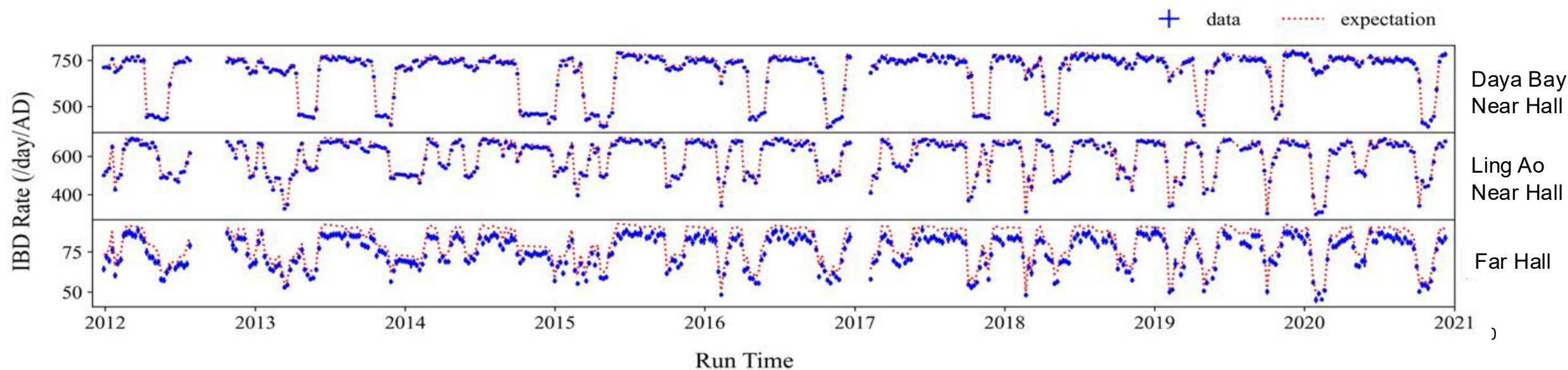
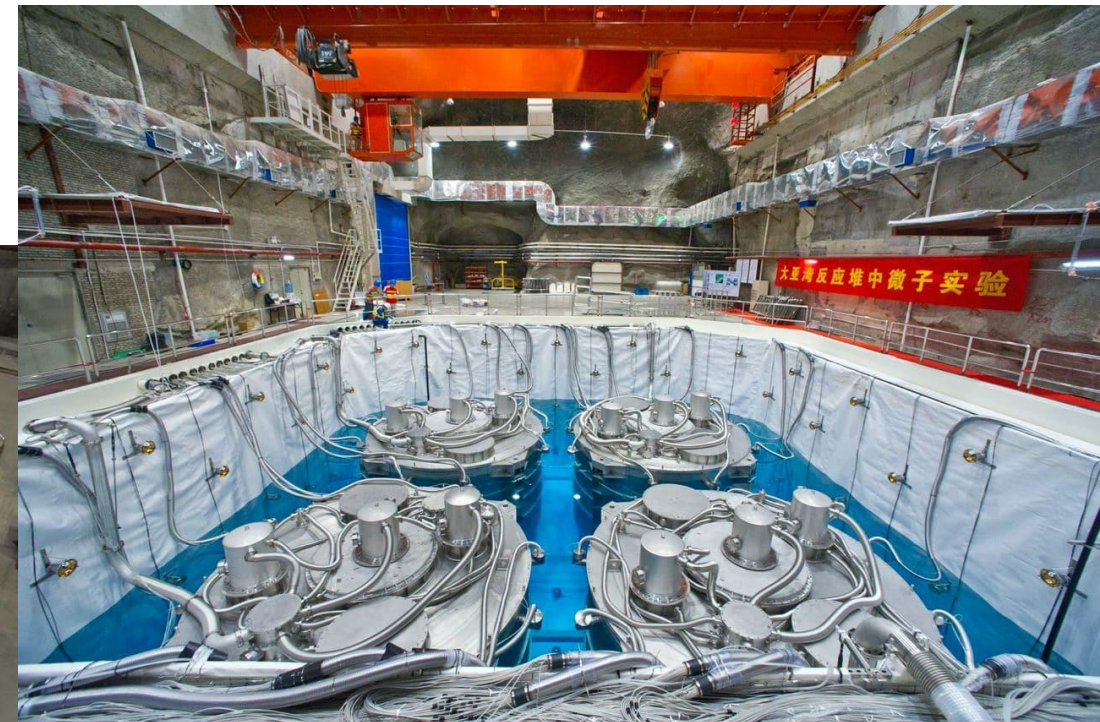
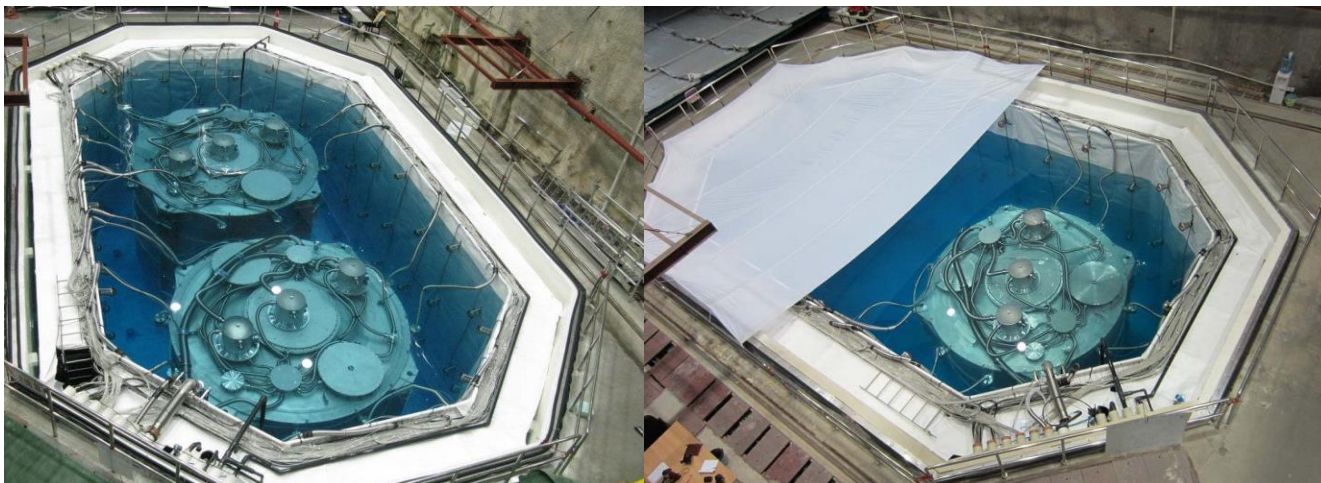
1. Plot the measured antineutrino signal rate of each AD vs. the expected flux, assuming each AD has the same size, and each reactor has the same power.
2. Fit the data (what function to use?) with the near ADs and extrapolate to the far ADs. What do you see?
3. What is the “survival probability” in the far ADs relative to the near ADs? What is the statistical significance of this observation?
4. What is the size of θ_{13} using the oscillation formula?

$$P = 1 - \sin^2 2\theta_{13} \cdot \sin^2(1.27|\Delta m_{ee}^2(\text{eV}^2)| \cdot \frac{L(\text{m})}{E(\text{MeV})})$$

Δm^2	$2.4 \times 10^{-3} (\text{eV}^2)$
L	$1.66 \times 10^3 (\text{m})$
E	3.5 (MeV)

□ Data taking (12/24/2011 – 12/12/2020)

- 3275 days, 5.5M $\bar{\nu}_e$ events
largest reactor neutrino data sample in the world



Precision Oscillation

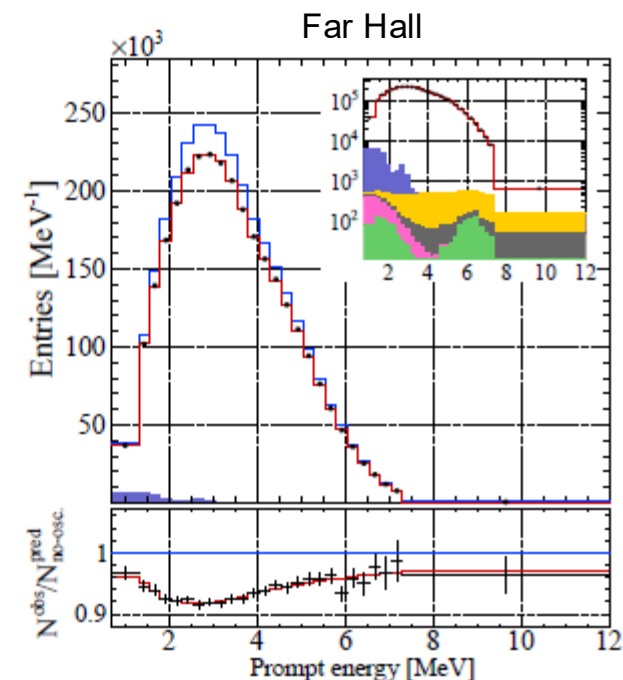
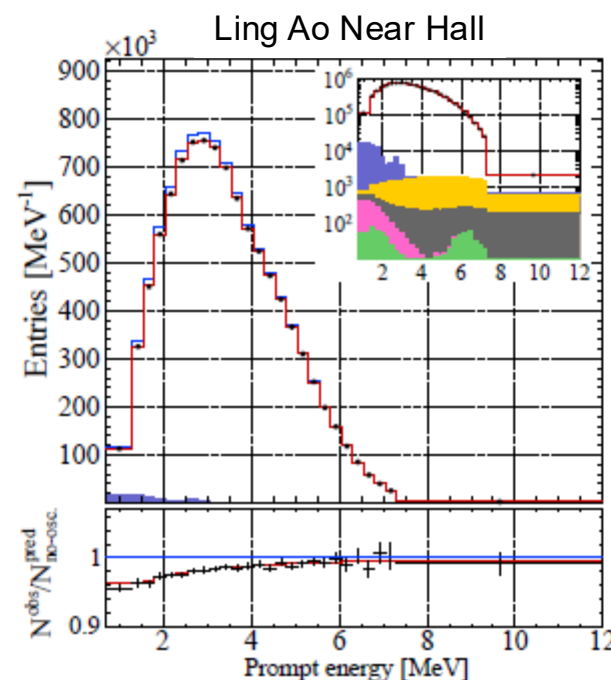
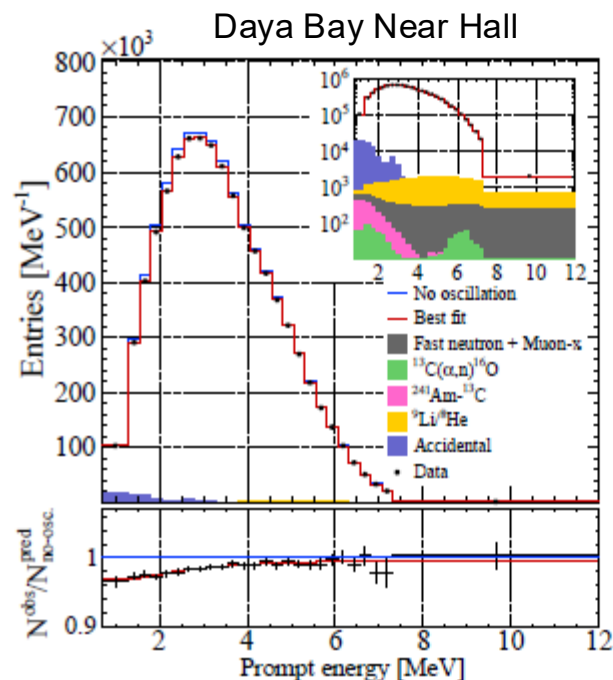
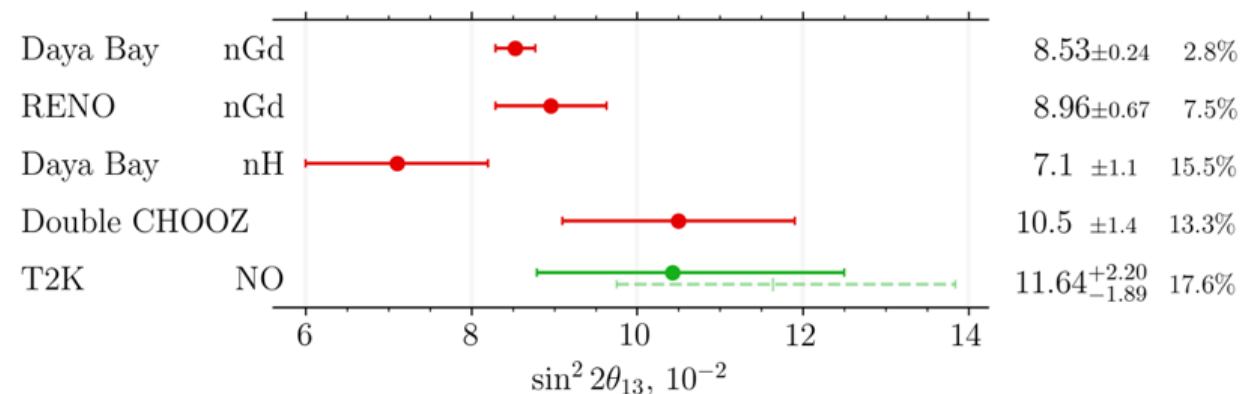
□ Final results with the full data set

Phys. Rev. Lett. 130, 161802 (2023)

$$\sin^2 2\theta_{13} = 0.0853^{+0.0024}_{-0.0024}$$

(2.8% precision)

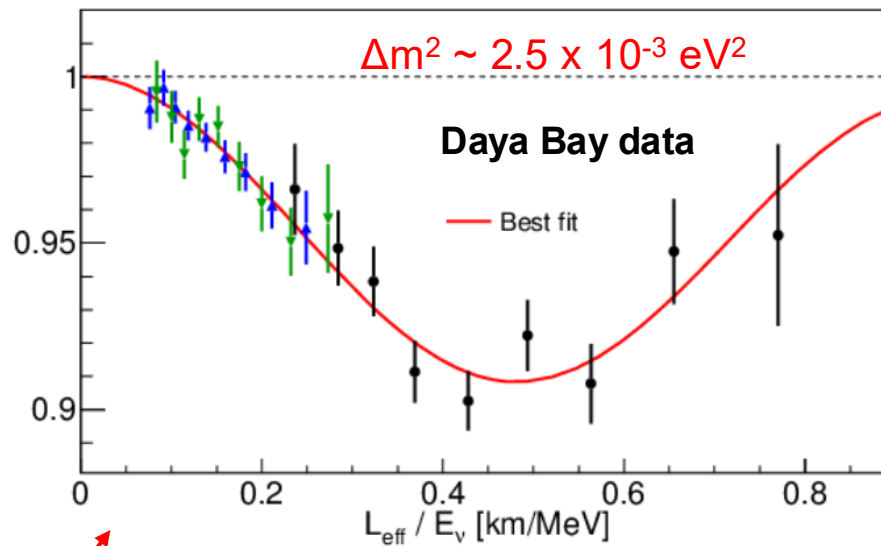
- Likely to be the best measurement in the foreseeable future
- Critical input to the current and future long-baseline experiments (DUNE)



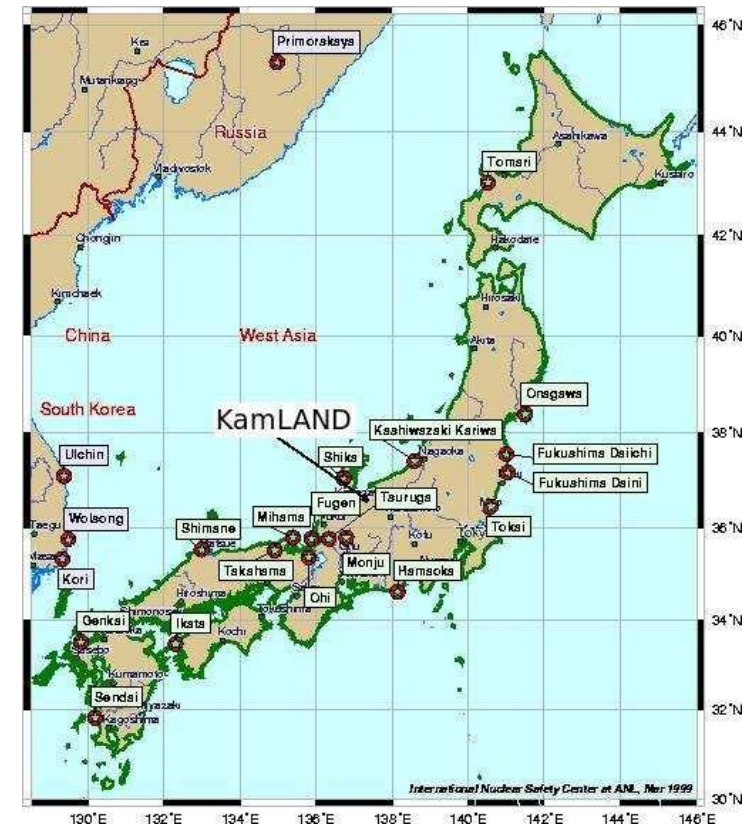
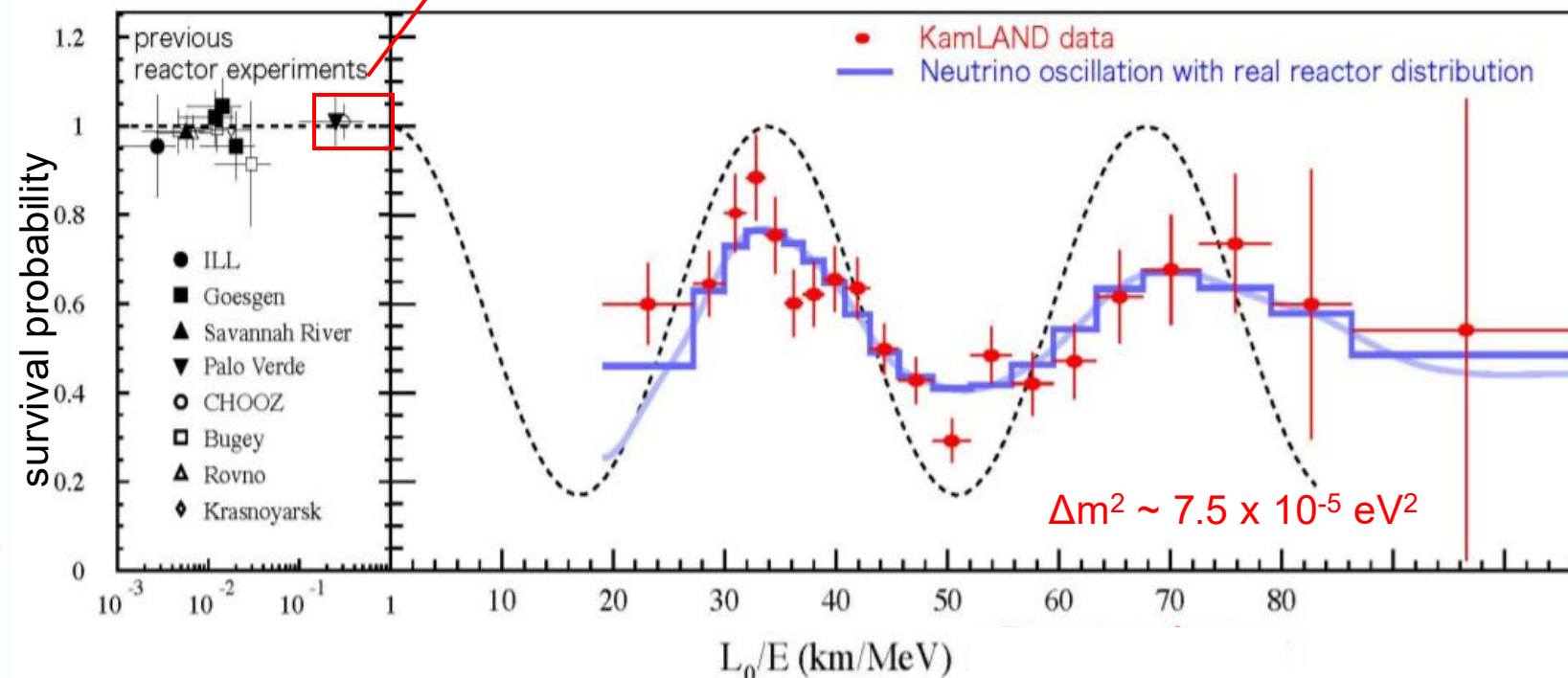


Questions?

Backup Slides

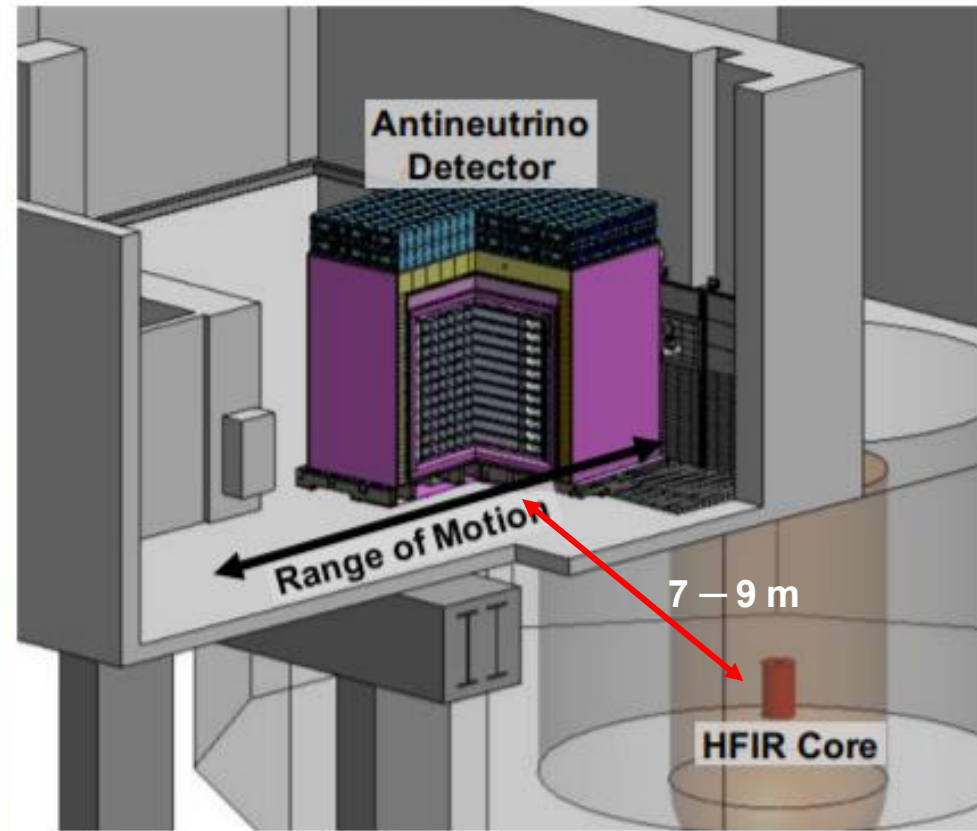


My postdoc work at BNL

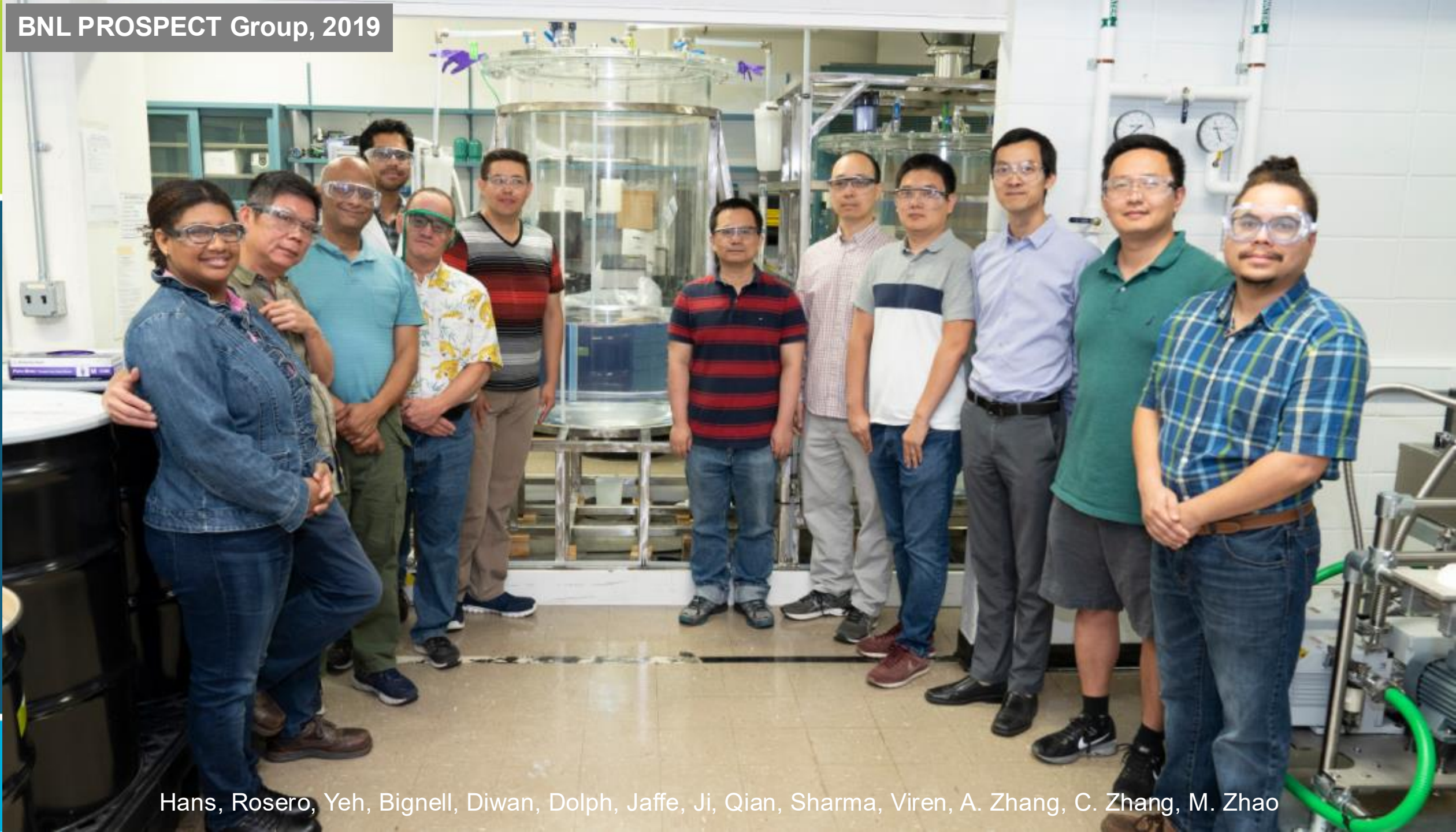


My Ph.D. thesis

Go Closer: The Precision Reactor Oscillation and Spectrum Experiment



Search for “oscillation patterns” from eV-scale sterile neutrinos
independent of reactor models at < 10 m



Hans, Rosero, Yeh, Bignell, Diwan, Dolph, Jaffe, Ji, Qian, Sharma, Viren, A. Zhang, C. Zhang, M. Zhao

Reactor, Detector, Data

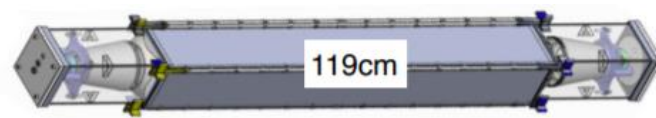
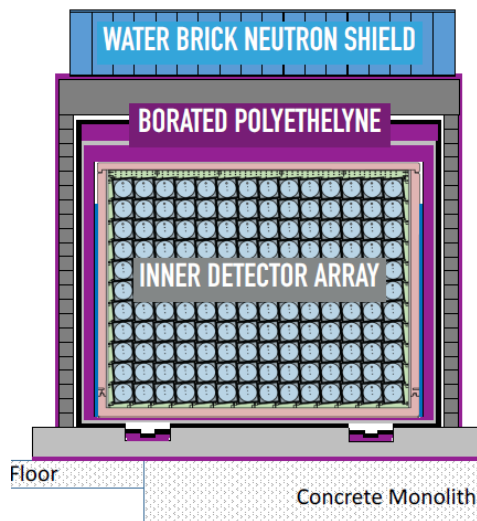


HFIR research reactor at ORNL

Power: 85 MW

Size: $\Phi=44\text{cm}$, $h=51\text{cm}$

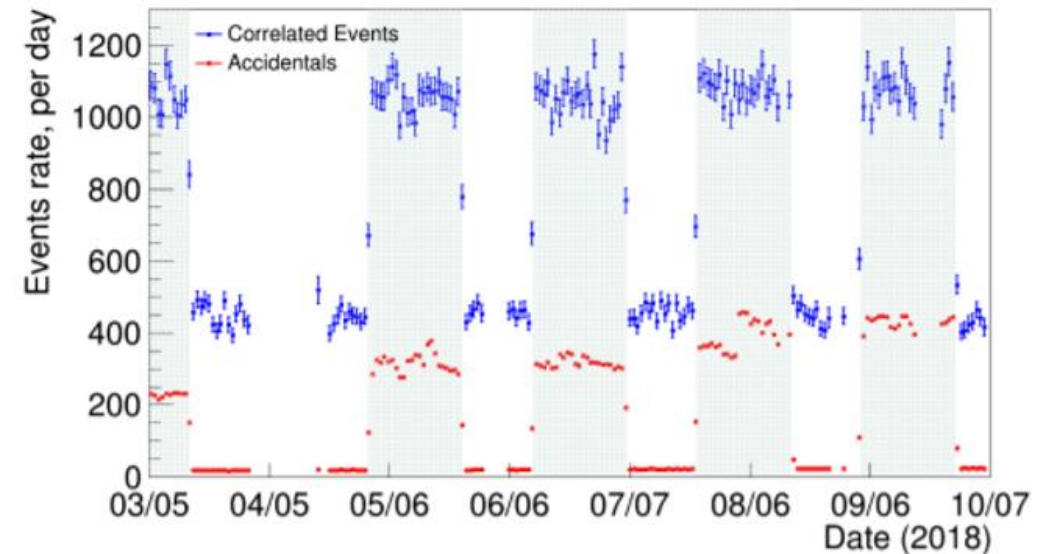
Fuel: highly-enriched ^{235}U
(^{235}U fission fraction > 99%)



Detector near surface,
Segmented to 11 x 14 “cells”

^6Li -loaded liquid scintillator (4 ton)
developed and produced at BNL

Reactor on 95 days; off: 73 days

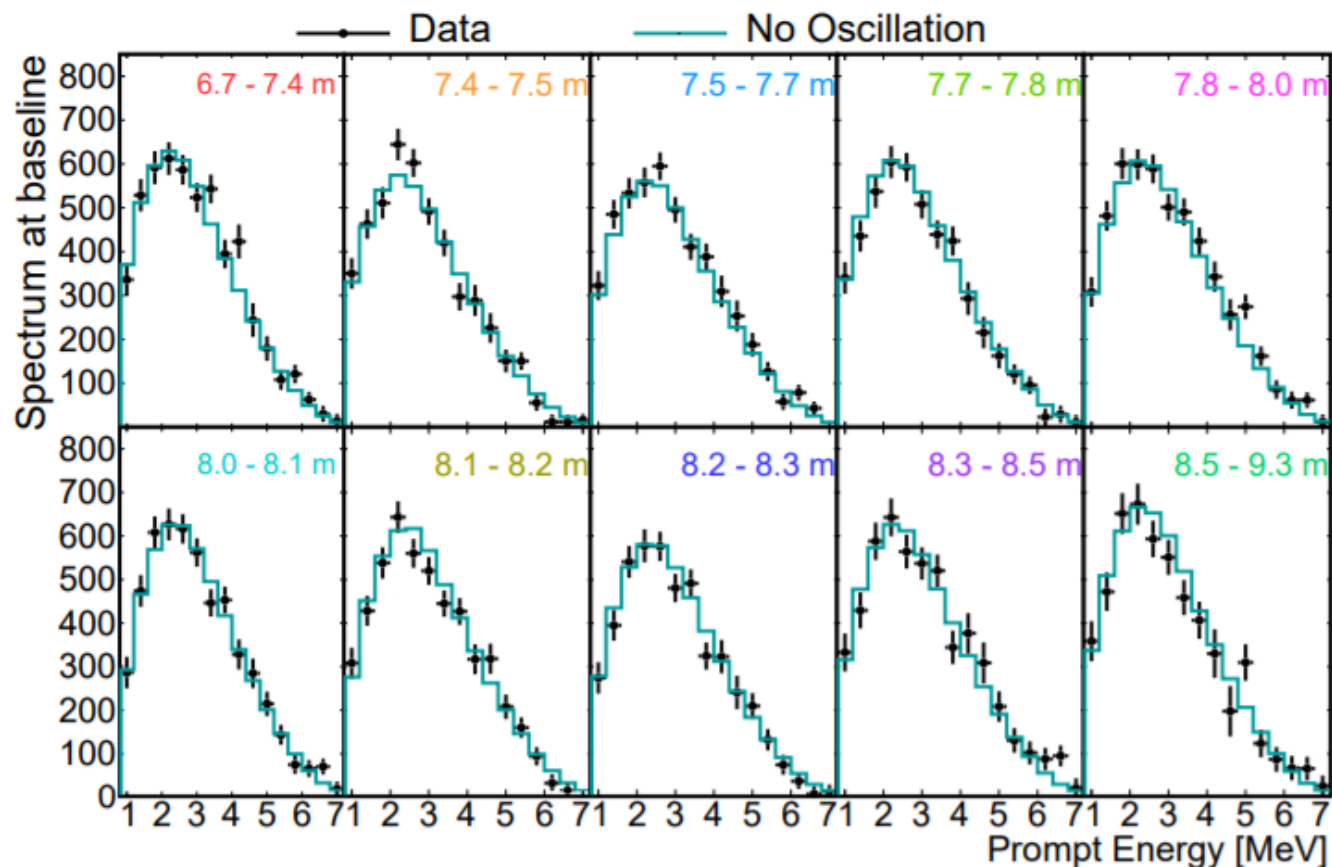


❑ > 50,000 antineutrinos from
 ^{235}U fission collected

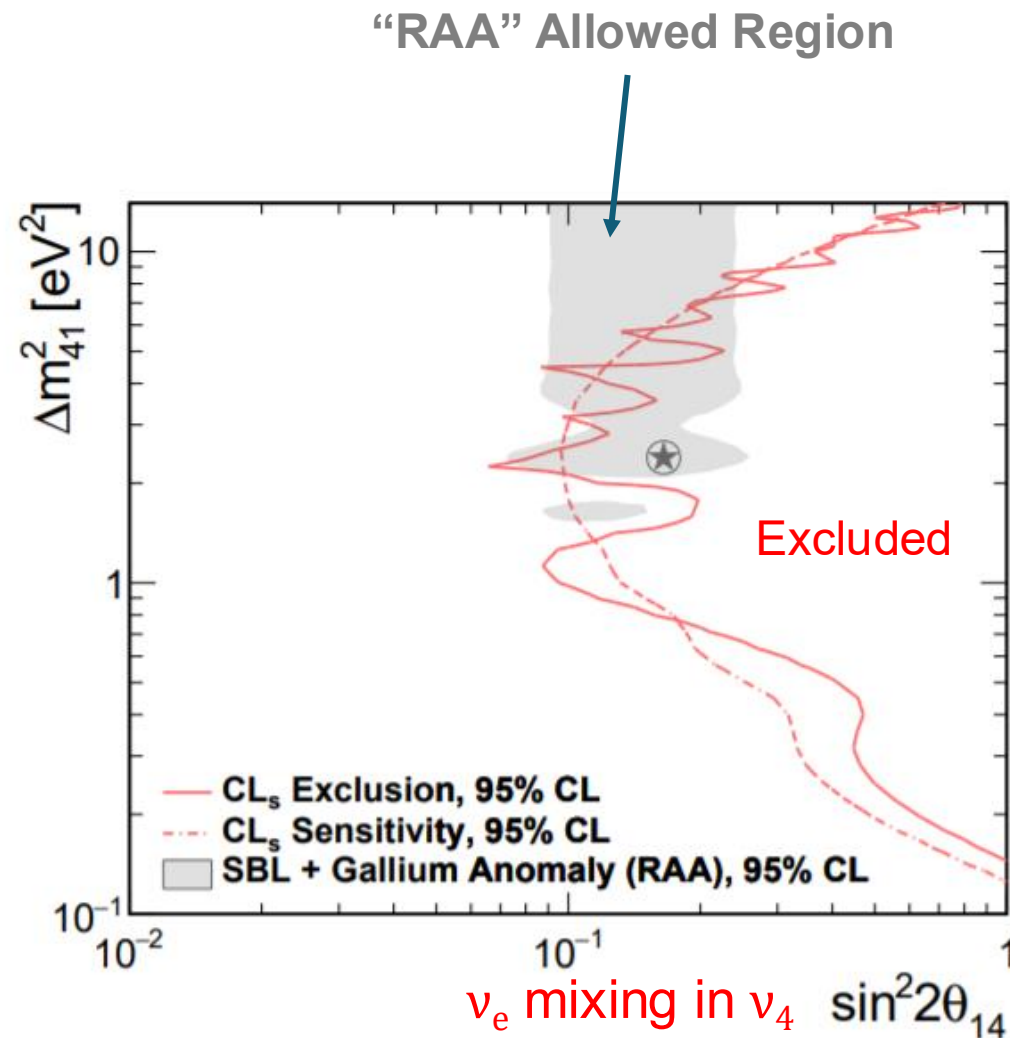
(10 times more than the previous record
from ILL in 1981)

❑ First surface-based
detector to achieve $S:B > 1$
near a research reactor

No “Wiggles” Found



Energy spectra at different distances all agree with “no oscillation” prediction => No indication of “eV-scale” sterile neutrinos



“RAA” allowed region largely excluded (independent of reactor models)