

Clusters and higher moments of proton number fluctuations

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Critical Point and the Onset of Deconfinement
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- 1 (Net) Proton number fluctuations and the influence of deuterons
- 2 Production mechanism of deuterons and their scaled moments

Part 1: Proton number fluctuations

- particle number fluctuations are sensitive to the phase transition
- higher-order cumulants are sensitive to the vicinity of the critical point
- however, cumulants can be influenced:
 - conservation laws
 - final state hadronic interactions
 - acceptance effects
 - efficiency effects
 - ...
- proton number serves as proxy for baryon number in measurements
- How well is the baryon number (multiplicity) distribution represented by proton number?

The influence of deuteron formation

- deuterons are formed **after** the fireball breakup
- some of the originally produced protons are eaten up by deuterons
- deuteron numbers scales with n_p^2 – non-linear coupling to proton number
- therefore, if proton number in an event is high, many protons disappear in deuterons
- this must be seen in higher moments of multiplicity distributions

The measures

n_p is the number of protons

$\langle \dots \rangle =$ averaging over events

variance

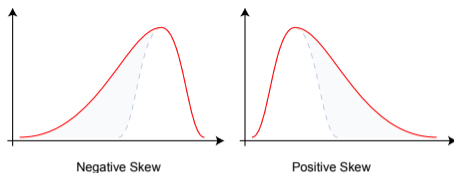
$$\sigma^2 = \langle (n_p - \langle n_p \rangle)^2 \rangle$$

skewness

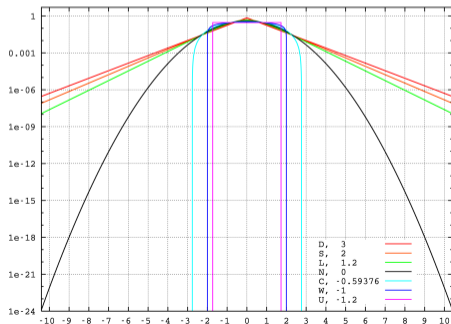
$$S = \frac{\langle (n_p - \langle n_p \rangle)^3 \rangle}{\sigma^3}$$

kurtosis

$$\kappa = \frac{\langle (n_p - \langle n_p \rangle)^4 \rangle}{\sigma^4} - 3$$



kurtosis measures the tails:



Underlying distributions of protons and deuterons

- **initial** proton number distribution Poissonian

$$P_i(n_i) = \lambda_p^{n_i} \frac{e^{-\lambda_p}}{n_i!}$$

NB: n_i is **not** measurable – it still includes protons which go into deuterons

- **average** number of deuterons is proportional to n_i^2

$$\lambda_d = Bn_i^2$$

- number of deuterons fluctuates according to Poisson distribution

$$P_d(n_d | n_i) = \lambda_d^{n_d} \frac{e^{-\lambda_d}}{n_d!} = (Bn_i^2)^{n_d} \frac{e^{-Bn_i^2}}{n_d!}$$

The observed distributions of protons and deuterons

- the **observed** number of deuterons is distributed as

$$P_d(n_d) = \sum_{n_i \geq n_d} P_d(n_d | n_i) P_i(n_i)$$

- the measured proton number is obtained after subtracting protons in deuterons

$$n_p = n_i - n_d$$

- observed** proton number distribution

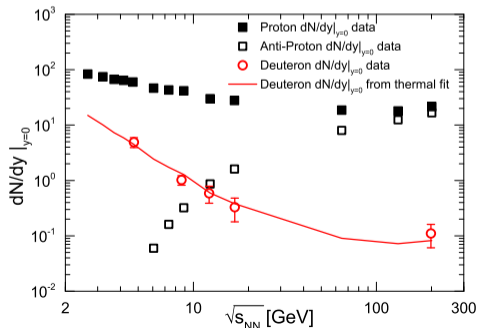
$$P(n_p) = \sum_{n_i \geq n_p} P_i(n_i) P_d(n_i - n_p | n_i)$$

Parameters fixed by observables

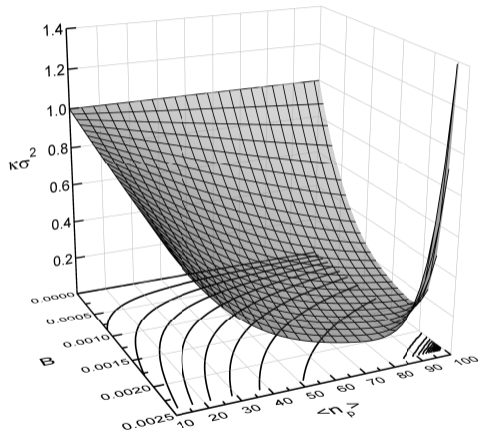
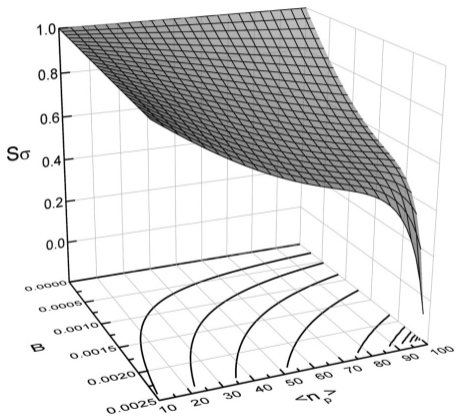
- parameters are:
 - mean initial proton number λ_p
 - coalescence factor B
- can be fixed with the help of observed multiplicities:
 - $\langle n_p \rangle = \sum_{n_p} n_p P(n_p)$
 - $\langle n_d \rangle = \sum_{n_d} n_d P(n_d)$

parametrisation of the mean deuteron/proton ratio

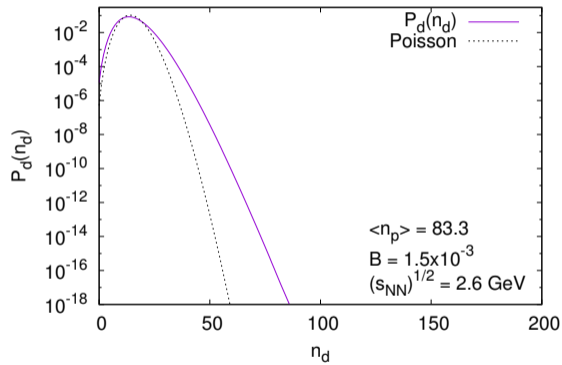
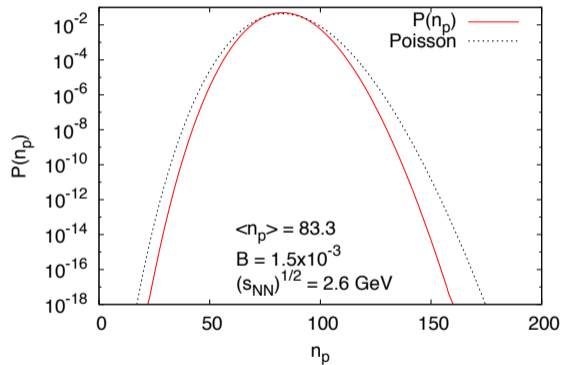
$$\frac{d}{p} = 0.8 \left[\frac{\sqrt{s_{NN}}}{1 \text{ GeV}} \right]^{-1.55} + 0.0036$$



Skewness and kurtosis after subtraction of deuterons

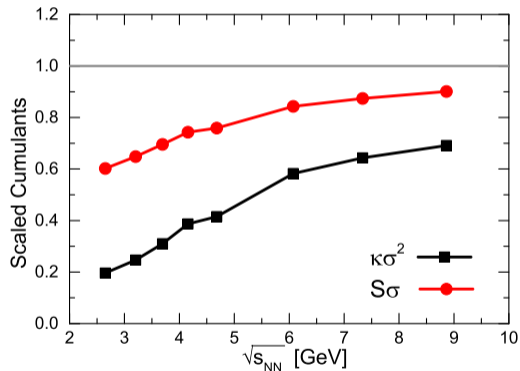


Example: distributions of p and d for $\langle n_p \rangle = 83.3$ and $B = 1.5 \times 10^{-3}$

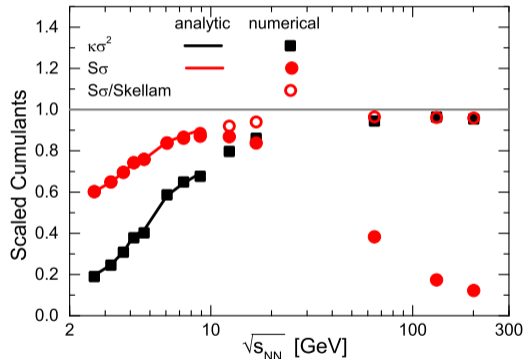


Results for scaled cumulants as functions of $\sqrt{s_{NN}}$

low energy: no antiprotons



higher energy:
also antiprotons fluctuate



Formation of deuterons has an important impact on proton number fluctuations.

Part 2: Thermal production vs coalescence of deuterons

- due to their fragility, deuterons can hardly exist in the dense hadronic system
- mean production numbers are well described by [coalescence](#)
- good description of data is also obtained with the help of [Statistical model](#)

Such models could be distinguished by fluctuations

- in Statistical model fluctuations of all species are Poissonian
- coalescence leads to non-Poissonian fluctuations of clusters (deuterons)

Deuteron number distribution

Model A: fully correlated proton and neutron numbers (as previously)

$$\begin{aligned}\lambda_d &= Bn_i^2 \\ P_d(n_d|n_i) &= \lambda_d^{n_d} \frac{e^{-\lambda_d}}{n_d!} = (Bn_i^2)^{n_d} \frac{e^{-Bn_i^2}}{n_d!} \\ P_d(n_d) &= \sum_{n_i \geq n_d} P_d(n_d|n_i) P_i(n_i)\end{aligned}$$

Model B: independent proton number n_i and neutron number n_j

$$\begin{aligned}\lambda_d &= Bn_i n_j \\ P_d(n_d|n_i, n_j) &= \lambda_d^{n_d} \frac{e^{-\lambda_d}}{n_d!} = (Bn_i n_j)^{n_d} \frac{e^{-Bn_i n_j}}{n_d!} \\ P_d(n_d) &= \sum_{n_i, n_j \geq n_d} P_d(n_d|n_i, n_j) P_i(n_i) P(n_j)\end{aligned}$$

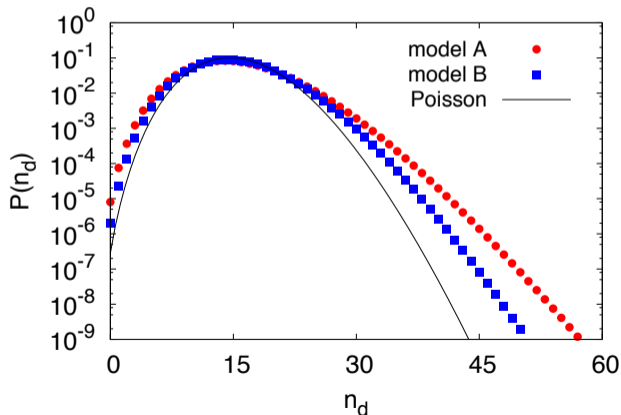
An example of deuteron number distribution

calculated for $\sqrt{s_{NN}} = 2.6$ GeV

correlated p and n: $\sigma^2/\langle n_d \rangle = 1.609$, $S\sigma = 2.218$, $\kappa\sigma^2 = 6.915$

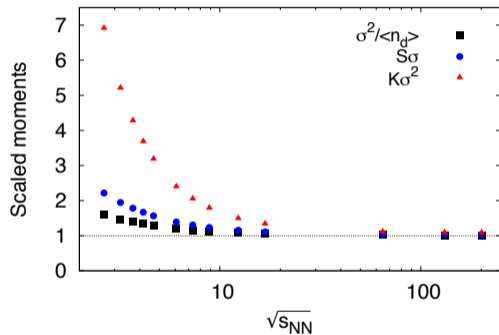
independent p and n: $\sigma^2/\langle n_d \rangle = 1.308$, $S\sigma = 1.616$, $\kappa\sigma^2 = 3.422$

Poissonian values are 1.

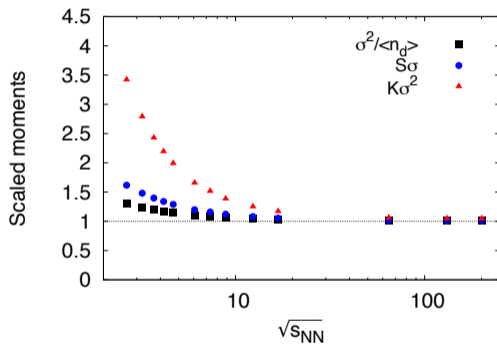


Predictions for the deuteron scaled moments

correlated p and n number



independent p and n number



Poissonian values (Statistical model) would be 1!

Higher moments in particular can clearly distinguish production by coalescence!

Deuteron formation has large influence on proton number fluctuations, especially at NICA/FAIR energies.

Z. Fecková, J. Steinheimer, B. Tomášik, M. Bleicher, Phys. Rev. C **92**, 064908 (2015)

Higher moments of deuteron number distribution can help to distinguish between statistical production and coalescence.

Z. Fecková, J. Steinheimer, B. Tomášik, M. Bleicher, Phys. Rev. C **93**, 054906 (2016)