### **Dileptons at low energies: Prospects**

### and challenges

### CPOD 2017 Stony Brook University, August 7-11, 2017

### Itzhak Tserruya





### Motivation

What did we learn from almost 25 years of dilepton measurements at SPS and RHIC

Prospects and challenges at low energies
 Summary

## **Motivation**

Dileptons (e<sup>+</sup>e<sup>-</sup>, µ<sup>+</sup>µ<sup>-</sup>) are sensitive probes of the two fundamental properties of the QGP:

- > Deconfinement
- Chiral Symmetry Restoration

 <u>Thermal radiation</u> emitted in the form of real photons or virtual photons (dileptons) provides a direct fingerprint of the matter formed (QGP and HG) and a measurement of its temperature.

QGP: 
$$q\overline{q} \longrightarrow \gamma^* \longrightarrow l^+l^-$$

$$\mathsf{HG:} \quad \pi^+\pi^- \longrightarrow \rho \longrightarrow \gamma^* \longrightarrow I^+I^-$$

## NA60 Acceptance corrected invariant mass spectrum

NA60 dimuon excess corrected for acceptance in m -  $p_T$ 

Rapp and Hees PLB 753, 586 (2016) NA60 data: Eur. Phys. J. C 59 (2009) 607



## LOW MASSES (m < 1 GeV/c<sup>2</sup>) SPS: CERES, NA60 RHIC: PHENIX, STAR

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#### First CERES result PRL 75, 1272 (1995)

(renowned paper: 579 citations)







Eur. Phys J. C41, 475 (2005) Itzhak Tserruya



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PRL 91, 042301 (2003)

<dN<sub>ee</sub>/dm<sub>ee</sub>>/<N<sub>ch</sub>>(100 MeV/c<sup>2</sup>)<sup>-1</sup>  $(d^2N_{ee}/d\eta dm)$  /  $(dN_{ch}/d\eta)$  (100 MeV/c<sup>2</sup>)<sup>-1</sup> CERES/NA45 Pb-Au 158 A GeV CERES/NA45 S-Au 200 GeV/u  $\sigma_{trid} \sigma_{tot} \approx 7 \%$ 10 2.1 < n < 2.65p.>200 MeV/c p, > 200 MeV/c ⊖<sub>aa</sub>>35 mrad  $\Theta_{ee} > 35 \text{ mrad}$ 10-5  $\langle dN_{ch}/d\eta \rangle = 125$ 2.1<n<2.65 10 First CERES result Last CERES result (a) PRL 75, 1272 (1995) PLB 666, 425 (2008) 10<sup>-6</sup> 10 (renowned paper: 579 citations) 10 10 charm 10-10 0 0.2 0.4 0.6 0.8 1.2 1.4 1.6 1 m<sub>ee</sub> (GeV/c<sup>2</sup>) m\_ (GeV/c<sup>2</sup>) 10 Pb-Au 158 AGeV σ/σ<sub>geo</sub>≈ 28% CERES/NA45 Pb-Au 40 AGeV <dN<sub>ee</sub>/dm<sub>ee</sub>>/<N<sub>ch</sub>> (100 MeV/c<sup>2</sup>)<sup>-1</sup>  $<dN_{ee}/dm_{ee}>/<N_{ch}>$  (100 MeV/c<sup>2</sup>)<sup>-1</sup>  $<dN_{ch}/d\eta>=245$ σ/σ<sub>αeo</sub>≈ 30 % 10-5 2.1<η<2.65 <dN<sub>ch</sub>/dη>=210 combined 95/96 data 10 p,>0.2 GeV/c 2.1<η<2.65 p,>200 MeV/c  $\Theta_{oo}$ >35 mrad ⊖<sub>∞</sub>>35 mrad 10 "een 10 eev 10-7 €ee ee à ee. 10 10-8 10 0 0 0.2 0.6 0.8 1.2 1.4 1.6 0.2 0.4 0.6 0.8 1 1.2 0.4 m<sub>ee</sub> (GeV/c<sup>2</sup>) m<sub>ee</sub> (GeV/c<sup>2</sup>) Eur. Phys J. C41, 475 (2005) PRL 91, 042301 (2003) Itzhak Tserruya

S-Au 200 GeV/u

2.1 < n < 2.65

p, > 200 MeV/c

 $\langle dN_{ch}/d\eta \rangle = 125$ 

m<sub>ee</sub> (GeV/c<sup>2</sup>)

 $\Theta_{ee} > 35 \text{ mrad}$ 

CERES/NA45

charm

10

10

10

10

<dN<sub>ee</sub>/dm<sub>ee</sub>>/<N<sub>ch</sub>>(100 MeV/c<sup>2</sup>)<sup>-1</sup>

10

10-6

10

10-

0 0.2 0.4

CERES/NA45

0.6 0.8

Pb-Au 158 A GeV

 $\sigma_{trid} \sigma_{tot} \approx 7 \%$ 

p>200 MeV/c

⊖<sub>aa</sub>>35 mrad

2.1<n<2.65

(a)

1.2 1.4 1.6

1

 $(d^2N_{ee}/d\eta dm) / (dN_{ch}/d\eta) (100 \text{ MeV/c}^2)^{-1}$ (renowned paper: 579 citations)

First CERES result

PRL 75, 1272 (1995)



Eur. Phys J. C41, 475 (2005) Itzhak Tserruya

Strong enhancement of low-mass e<sup>+</sup>e<sup>-</sup> pairs in all A-A systems studied

First evidence of thermal radiation from the HG  $\pi^+\pi^- \longrightarrow \rho \longrightarrow \gamma^* \longrightarrow e^+e^-$ 







PRL 91, 042301 (2003)

## Dropping Mass or Broadening (I) ?

#### \* Interpretations invoke: $\pi^+\pi^- \rightarrow \rho \rightarrow \gamma^* \rightarrow e^+e^-$



### thermal radiation from HG

 vacuum ρ not enough to reproduce data

### CERES Pb-Au 158 A GeV 95/96 data



## **Dropping Mass or Broadening (I)**?

#### \* Interpretations invoke: $\pi^{+}\pi^{-} \rightarrow \rho \rightarrow \gamma^{*} \rightarrow e^{+}e^{-}$



### thermal radiation from HG

vacuum p not enough to reproduce data

### \* in-medium modifications of $\rho$ : $\diamond$ broadening $\rho$ spectral shape

(Rapp and Wambach)

(Brown et al)

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### CERES Pb-Au 158 A GeV 95/96 data



## Low-mass dileptons and in-medium p



NA60, PRL 96, 162302 (2006)

Excess shape in agreement with broadening of the ρ mainly due to the scattering of ρ off baryons (Rapp-Wambach)

**Dropping mass of the** ρ (Brown-Rho) ruled out

Confirmed by CERES results (PLB 666, (2008) 425) Melting of the ρ

### Low-mass e<sup>+</sup>e<sup>-</sup> Pairs: Prospects at RHIC

- At SPS energies, the ρ-meson broadening, that explains both the CERES and NA60 data, relies on the high baryon density.
- □ What was expected at RHIC?

	SPS (Pb-Pb)	RHIC (Au-Au)
dN( <del>p</del> ) / dy	6.2	20.1
Produced baryons ( <del>p</del> , p, <del>n</del> , n )	24.8	80.4
$p - \overline{p}$	33.5	8.6
Participants nucleons ( $\overline{p} - p$ )A/Z	<mark>85</mark>	21.4
Total baryon density	110	102

 Baryon density is almost the same at RHIC and SPS (the decrease in the participating nucleons transported to mid-rapidity is compensated by the copious production of nucleon-antinucleon pairs)

Strong enhancement of low-mass pairs predicted to persist at RHIC



## PHENIX vs. STAR



#### Enhancement factor in 0.15<M<sub>ee</sub><0.75 Gev/c<sup>2</sup>

	Minimum Bias	<b>Central collisions</b>
PHENIX	4.7 ± 0.4 ± 1.5	$7.6 \pm 0.5 \pm 1.3$
STAR	$1.40 \pm 0.06 \pm 0.38$	$1.54 \pm 0.09 \pm 0.45$

Large quantitative differences

## Last PHENIX results

#### PRC 93, 014904 (2016)



#### □ HBD upgrade:

- Improved hadron rejection:  $30\% \rightarrow 5\%$
- Improved signal sensitivity

#### □ New improved analysis

- Neural network for e-id
- Flow modulation incorporated in the mixed event using an exact analytical method
- Absolutely normalized correlated BG

### Minimum bias data/cocktail

0.3-0.76 (GeV/c²)	Data/cocktail ±stat ±syst ±model
PHENIX 2010	$2.3 \pm 0.4 \pm 0.4 \pm 0.2$ (Pythia) $1.7 \pm 0.3 \pm 0.3 \pm 0.2$ (MC@NLO)
STAR	$1.76 \pm 0.06 \pm 0.26 \pm 0.29$

### Consistent results between PHENIX and STAR

## Comparison to Rapp's model





Mass and p<sub>T</sub> dependencies of excess well reproduced by Rapp's model

- In-medium ρ broadening due to the scattering of the ρ off baryons in the HG
- Significant contribution from the QGP at low masses

Centrality dependence consistent with  $N_{part}^{1.45}$  as predicted by model.



## STAR beam energy scan



Systematic study of the dielectron continuum at:

200, 62.4, 39, 27 and 19.6 GeV

Low mass excess observed at all energies

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Systematic study of the dielectron continuum at:
200, 62.4, 39, 27 and 19.6 GeV

Low mass excess observed at all energies

 Rapp's model reproduces the excess at all energies.

invariant dielectron mass, Mee (GeV/c<sup>2</sup>)

## STAR beam energy scan



### **Connection with CSR?**

- In-medium broadening of the ρ meson (mainly by scattering off baryons) explains the dilepton excess in the LMR – The ρ meson "melts" in the high density medium.
- Is this connected to CSR? The measurement of the chiral partner a<sub>1</sub> is very difficult
- Recent calculations by Hohler and Rapp (PLB 731 (2014) 103) show that ρ and a<sub>1</sub> become degenerate at high temperatures: the ρ broadens as T increases, whereas a<sub>1</sub> mass drops and the spectral shapes of ρ and a<sub>1</sub> coincide at high T.



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## Intermediate masses (m = 1-3 GeV/c<sup>2</sup>) SPS: NA60 RHIC: PHENIX and STAR

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## **Origin of the IMR Excess**

NA60, PRL 96, 162302 (2006)



## Origin of the IMR Excess

NA60, PRL 96, 162302 (2006)

Renk/Ruppert, PRL 100,162301 (2008)



Dominant process in mass region  $m > 1 \text{ GeV/c}^2$ :

qq annihilation – thermal radiation from the QGP

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### PHENIX and STAR results are now consistent

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## <u>cc in cocktail</u>

#### PHENIX, PRC 93, 014904 (2016)



Cross section derived using IMR in d+Au collisions and extrapolating to m~0
 → uncertainty in cross section

PHENIX, PRC 91, 014907 (2015)

	d $\sigma^{pp}_{c\overline{c}}$ /dy (µb)
PYTHIA	$106 \pm 9^{stat} \pm 33^{syst}$
MC@NLO	$287 \pm 29^{stat} \pm 100^{syst}$

Hadronic decays of D mesons
 STAR, PRL 113, 22301 (2014)
 dσ/dy = 171 ± 26 μb (PYTHIA)

## <u>cc in cocktail</u>

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Hadronic decays of D mesons
 STAR, PRL 113, 22301 (2014)
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- ➤ c quarks suffer energy loss in the medium → effect on the cc correlation?
- Lack of appropriate modeling of cc correlation
  - $\rightarrow$  uncertainty in shape



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Rapp and Hees PLB 753, 586 (2016) NA60 data: Eur. Phys. J. C 59 (2009) 607





STAR PRC 90, 64904 (2014)



Inclusive dielectron v<sub>2</sub> consistent with simulated v<sub>2</sub> from cocktail sources



#### STAR PRC 90, 64904 (2014)



## Challenge: isolate the v<sub>2</sub> of the excess dileptons



### Inclusive dielectron v<sub>2</sub> consistent with simulated v<sub>2</sub> from cocktail sources

# Lower – energies:

BM@N at Nuclotron CBM at FAIR HADES at GSI MPD at NICA NA60+ at SPS ? STAR – BES-II at RHIC

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## Matter at low energy collisions

Rapp, Wambach, Adv.Nucl.Phys. 25, 1 (2000)



Baryon dominated matterVery low pion density

### Evolution of average $\rho_{\text{B}}$



- $\Box$  Sizable densities 3-6  $\rho_0$
- □ Long lifetime

## Matter at low energy collisions

Rapp, Wambach, Adv.Nucl.Phys. 25, 1 (2000)



Evolution of average  $\rho_{\rm B}$ 



Study dileptons under highest baryon density

- Unveil onset of excess?
- Critical point? First order phase transition?

T (MeV)

- Baryon dominated matter
- Very low pion density

- $\Box$  Sizable densities 3-6  $\rho_0$
- □ Long lifetime

## IMR as thermometer

Rapp and Hees, PLB 753, 586 (2016)

T given by inverse slope of the acceptance corrected mass spectrum in the IMR.



## IMR as thermometer

## and LMR as chronometer

Rapp and Hees, PLB 753, 586 (2016)

T given by inverse slope of the acceptance corrected mass spectrum in the IMR.



The thermal radiation integrated in the LMR m = 0.3 - 0.7 GeV/c<sup>2</sup> tracks the fireball lifetime quite well



## Thermal yields at low energies

R. Rapp – private communication



 Cross sections decrease by almost two orders of magnitude between central Au+Au at 200 GeV and central Pb+Pb at 6.3 GeV at m=2 GeV/c<sup>2</sup>
 Challenging measurements

## Charm cross section in pp



Cross sections down by ~3 orders of magnitude between RHIC and NICA energies.

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## **Dilepton experiments – energy map**





## **Comparison to other facilities**



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## **MPD detector at NICA**

- □ 9 m long 6 m diameter
- □ Low material budget
- Tracking (TPC):
   up to |η|<2, 2π in azimuth</li>
- PID (TOF, TPC, ECAL): hadrons, e, γ





## MPD detector at NICA



Estimate of dilepton yield in central Au+Au at  $m = 2 - 2.5 \text{ GeV/c}^2$  $\sqrt{s_{NN}} = 8 \text{ GeV}$  410 pairs/10 d

BM@I



- □ All systems at all energies studied show an enhancement of dileptons.
- □ A single model consistently reproduces the observed enhancement.
- The thermal radiation from the QGP dominates the dilepton excess in the IMR. Provides a measurement of the average temperature of the medium in the QGP phase.
- The thermal radiation from the HG dominates the dilepton excess in the LMR. Seems to track the medium lifetime.
- **□** Emerging picture for the realization of CSR: the ρ meson broadens in the medium, the  $a_1$  mass drops and becomes degenerate with the ρ.
- □ Missing:
  - ✤ precise measurements of IMR at RHIC energies.
  - $\bullet$  v<sub>2</sub> measurements of the excess dileptons.
- Clear predictions and strong experimental program to study dileptons at low energies.