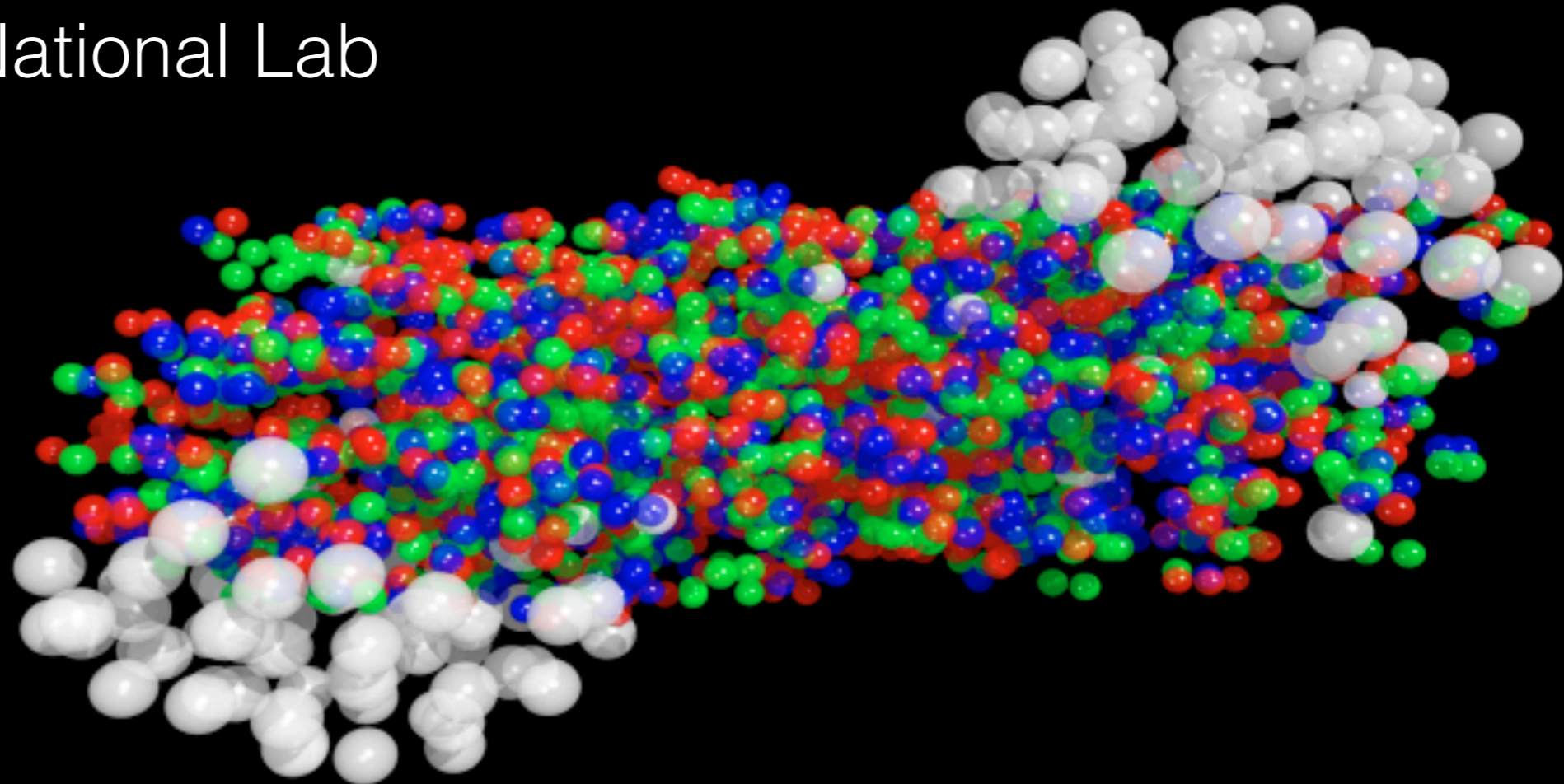
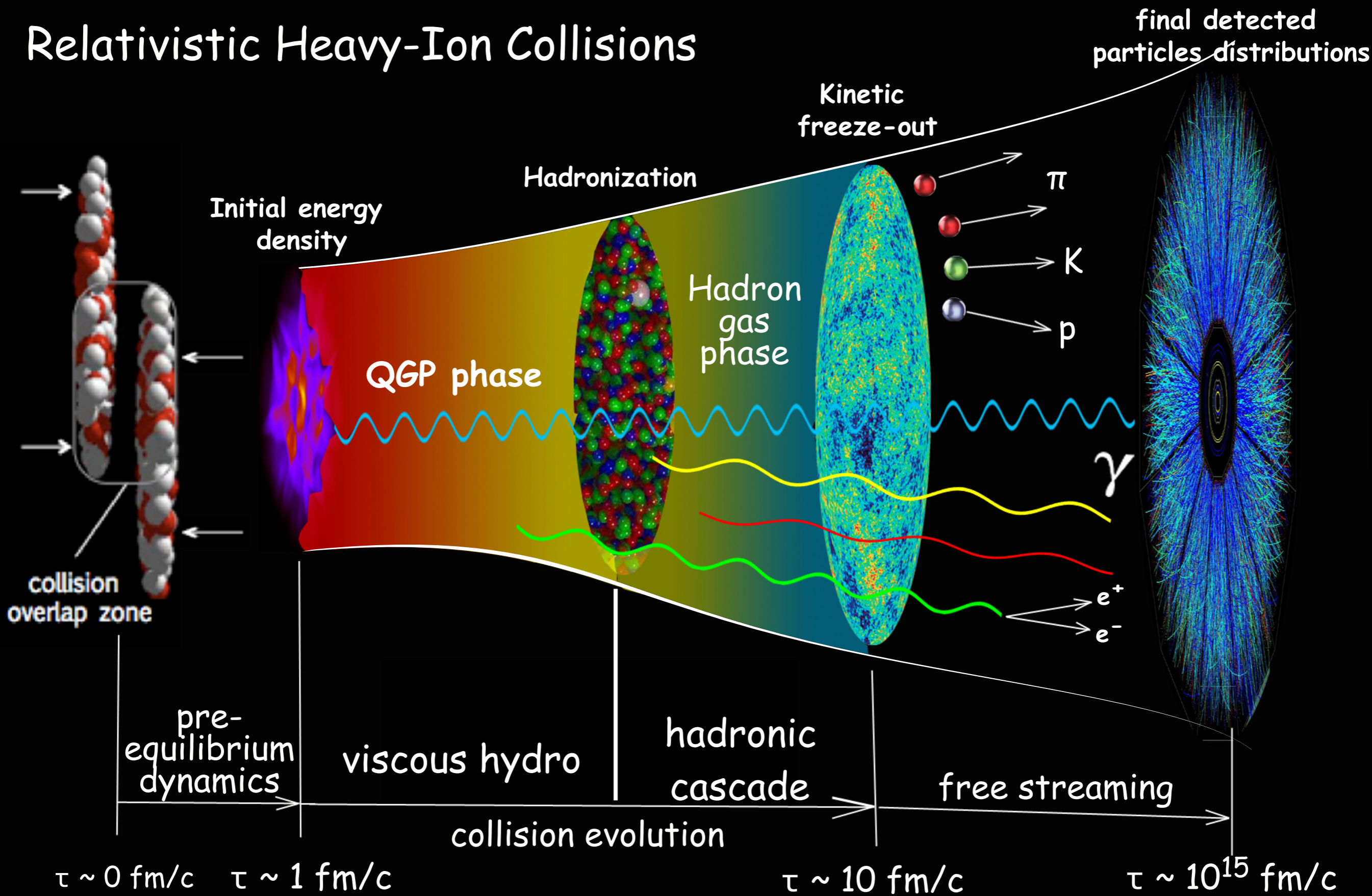


Hydrodynamic modeling of RHIC BES

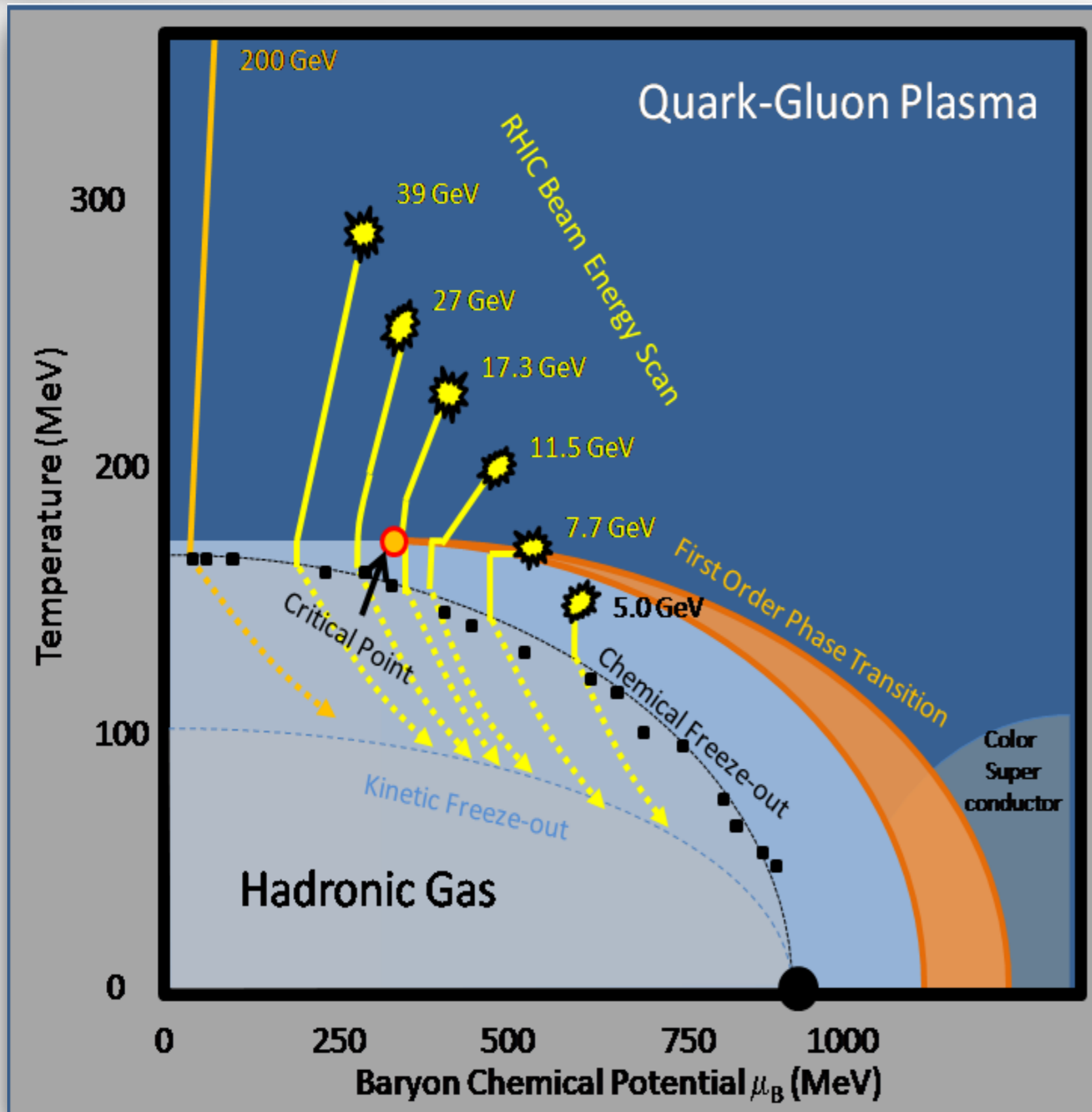
Chun Shen
Brookhaven National Lab



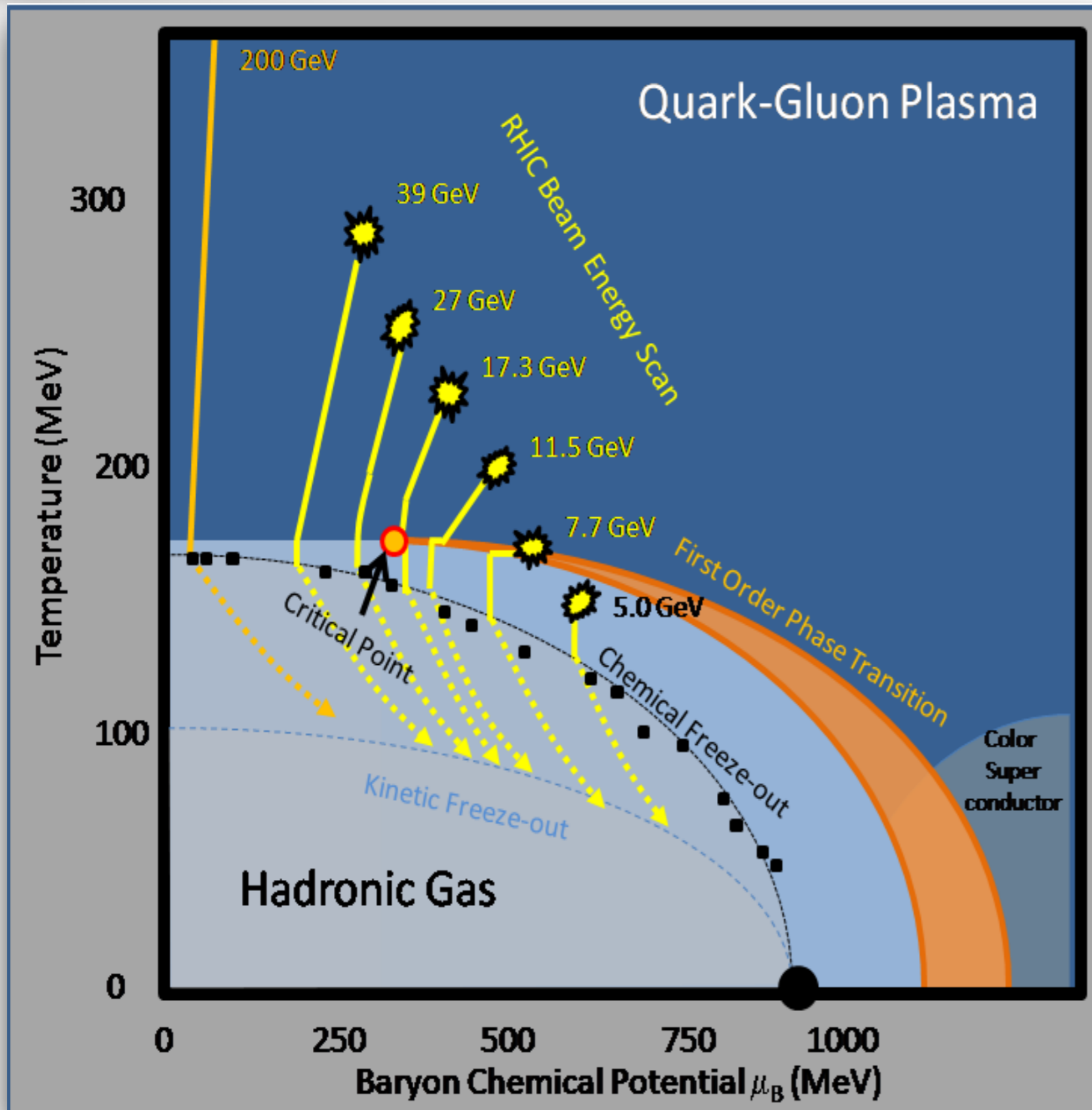
Relativistic Heavy-Ion Collisions



Exploring the phases of QCD

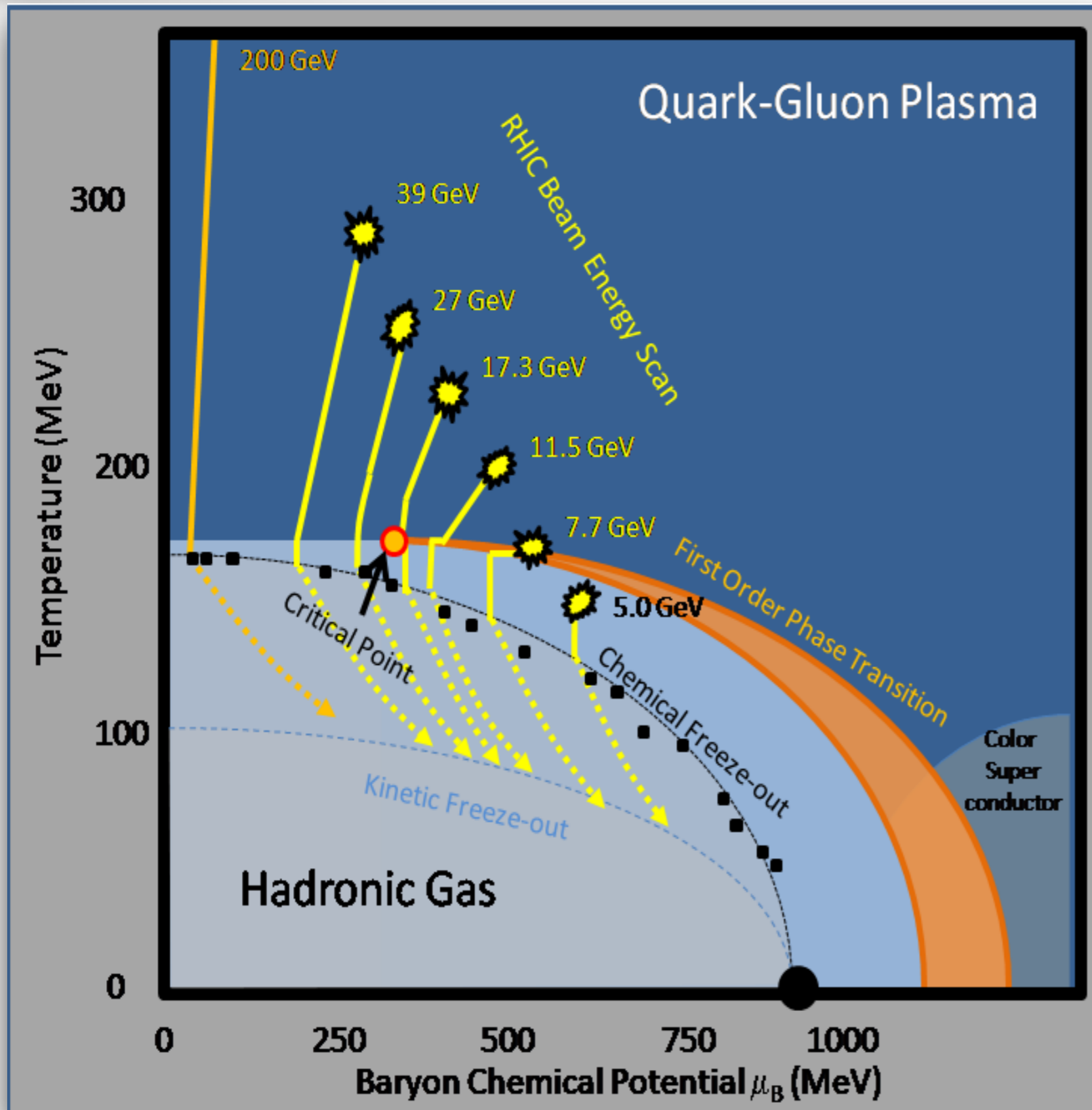


Exploring the phases of QCD



- Event-by-event fluctuating initial conditions
- (3+1)-d dissipative hydrodynamic modelling of the QGP
- Microscopic description for hadronic phase

Exploring the phases of QCD



- Event-by-event fluctuating initial conditions

Glauber-LEXUS

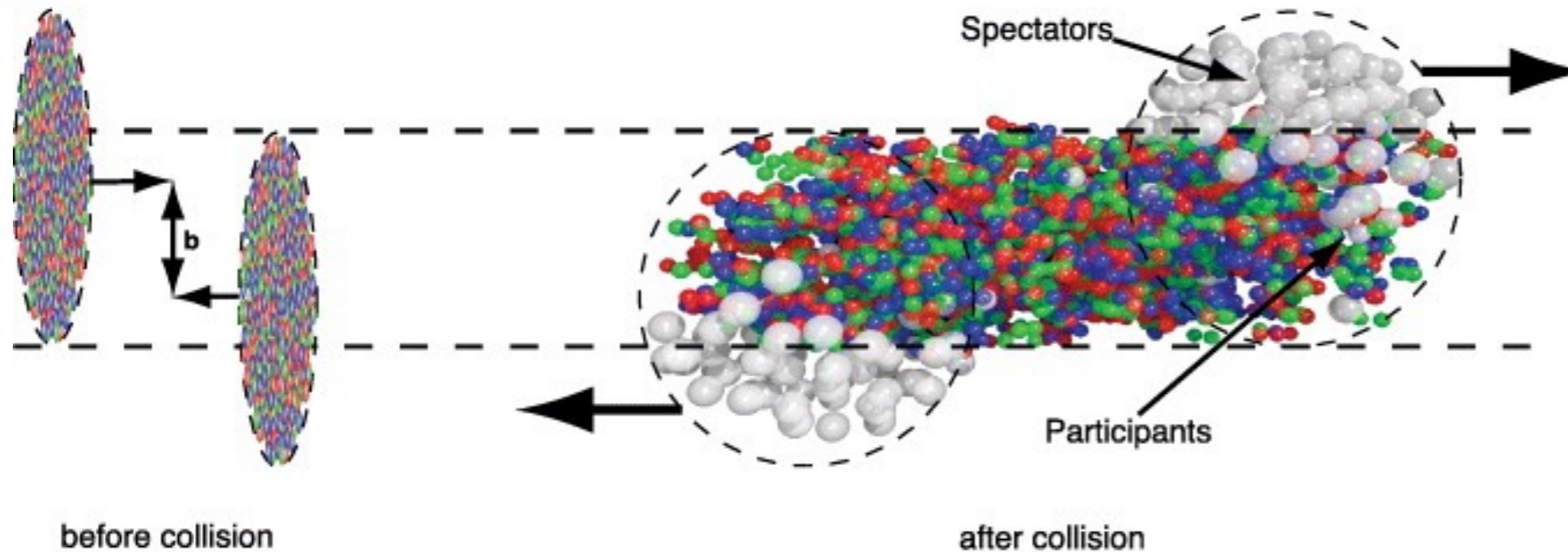
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MUSIC

- Microscopic description for hadronic phase

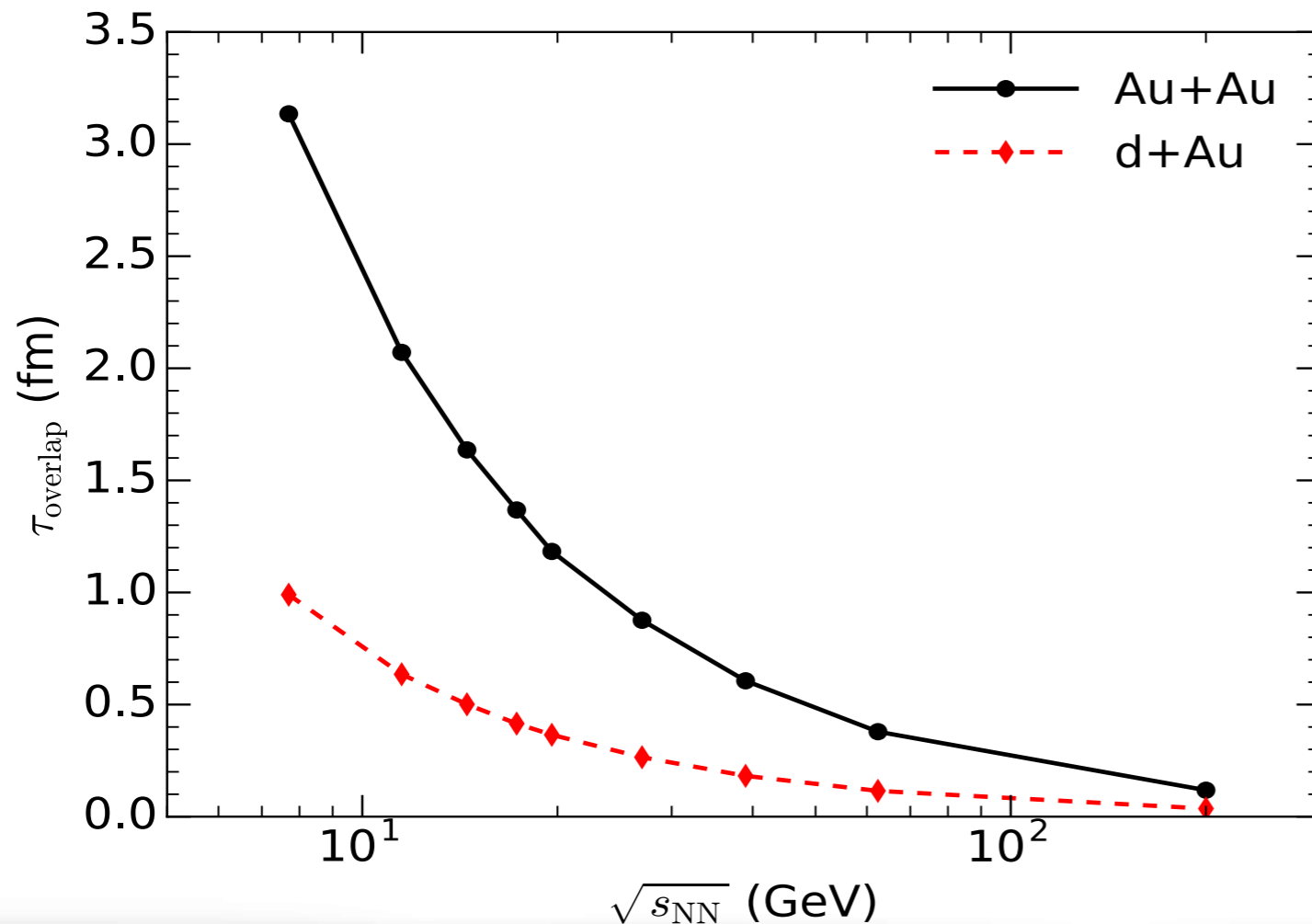
UrQMD/JAM

When to start hydrodynamics?



Two nuclei
overlapping time

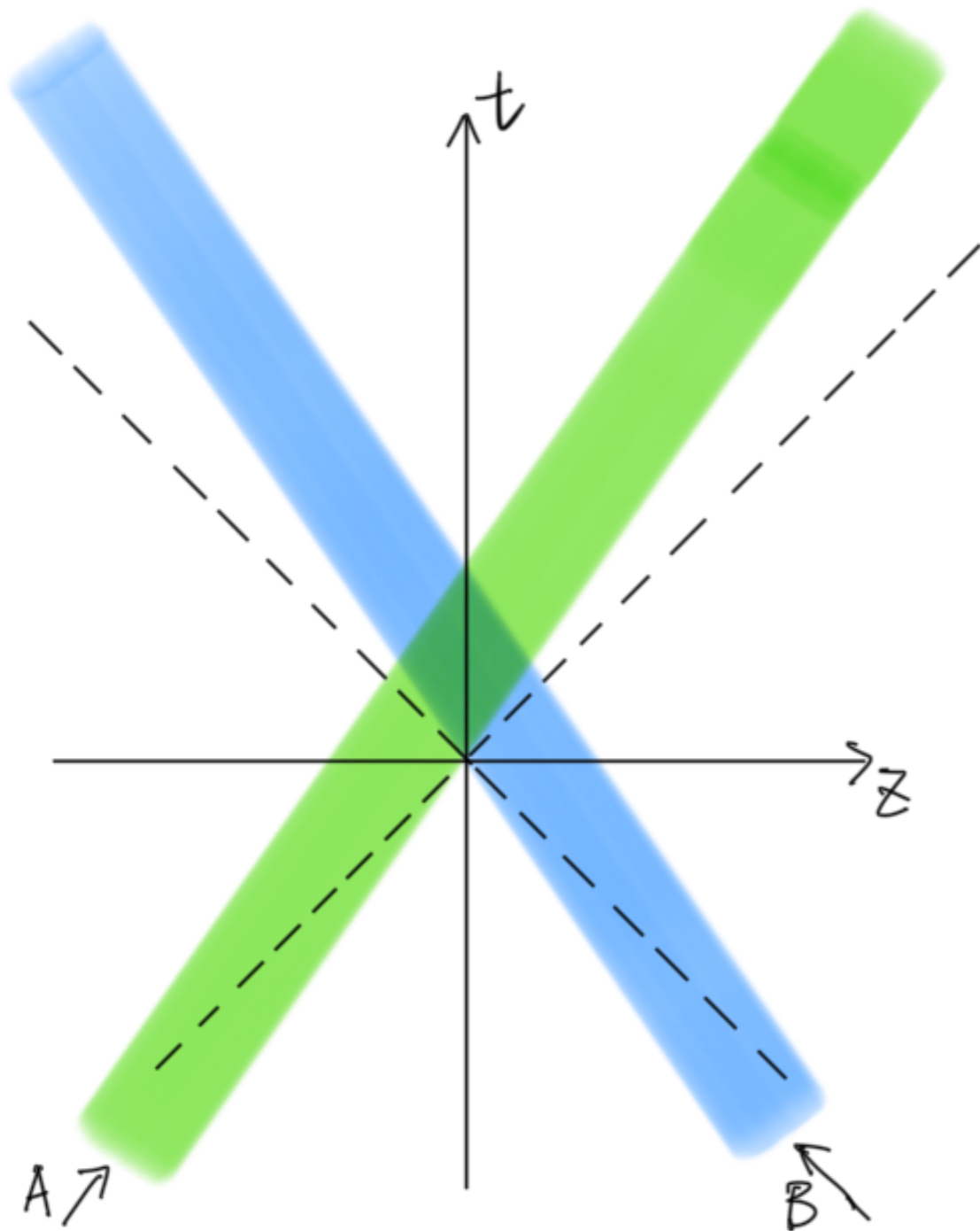
$$\tau \sim \frac{2R}{\gamma v_z}$$



- Nuclei overlapping time is **large** at low collision energy
- **Pre-equilibrium dynamics** can play an important role

Go beyond the Bjorken approximation

- The finite widths of the colliding nuclei are taken into account

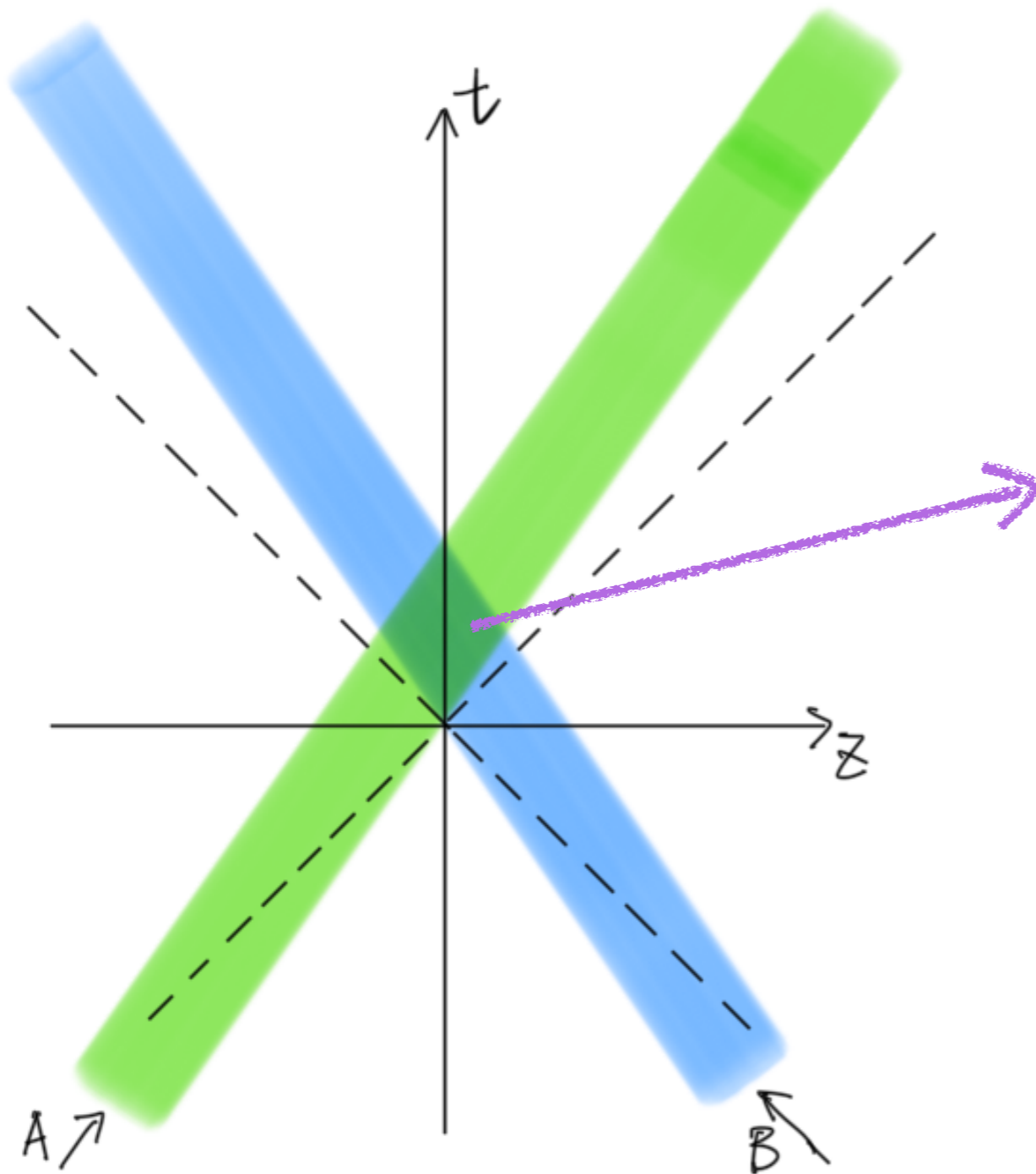


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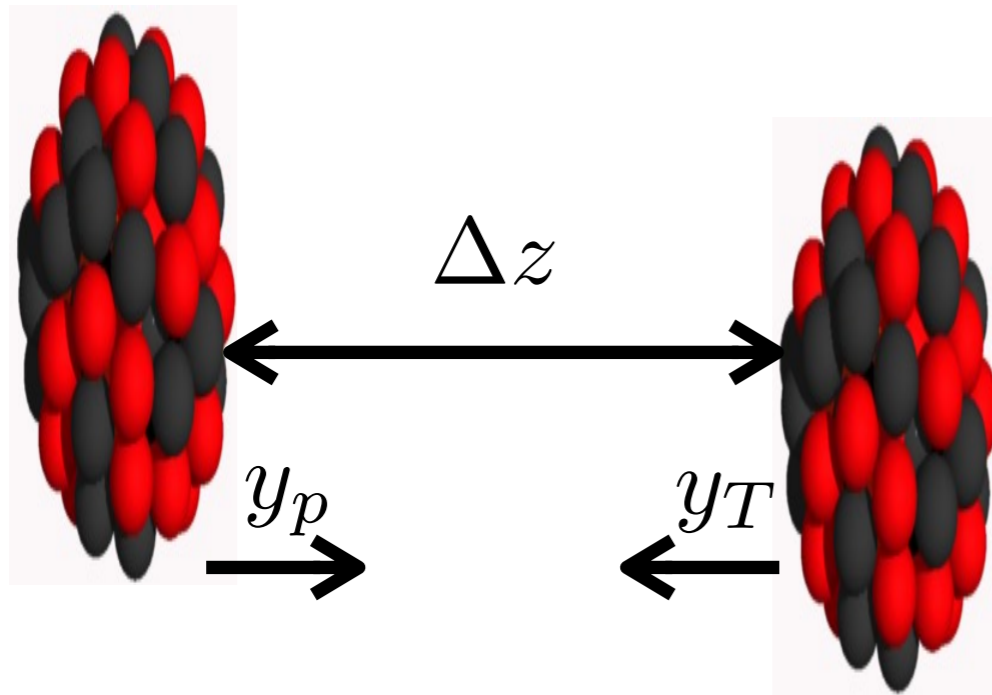
The interaction zone is not point like

$$y \neq \eta_s$$



The 3D MCGlauber-LEXUS model

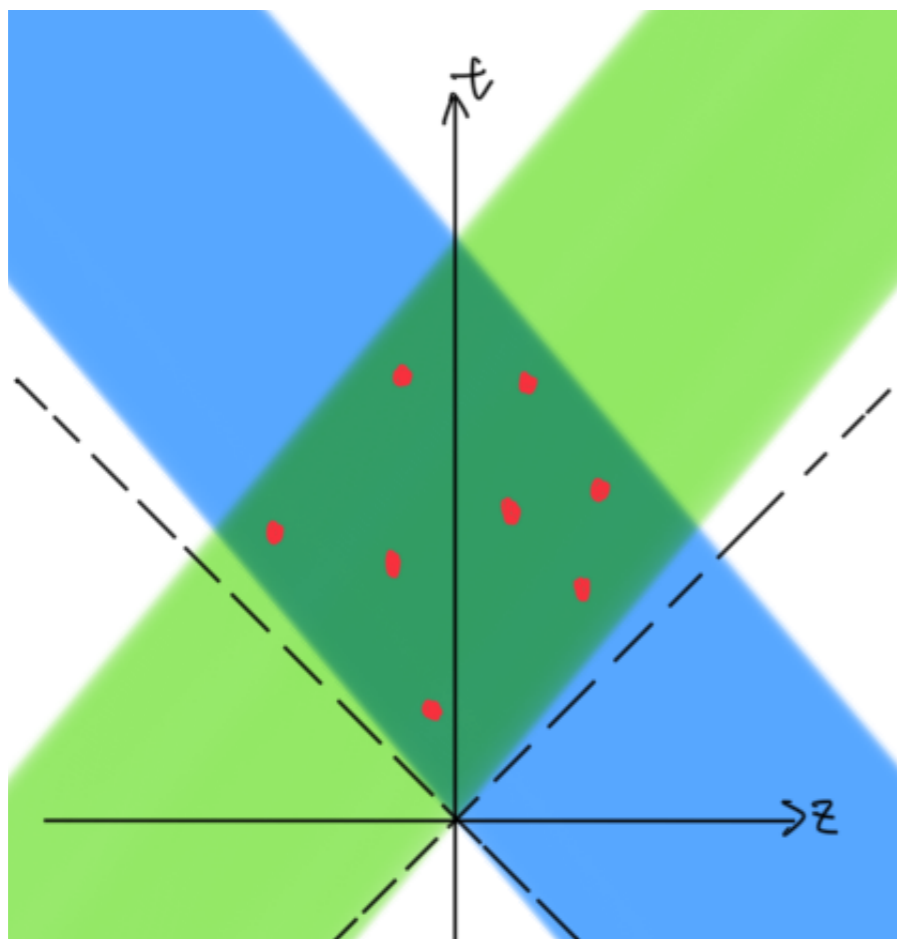
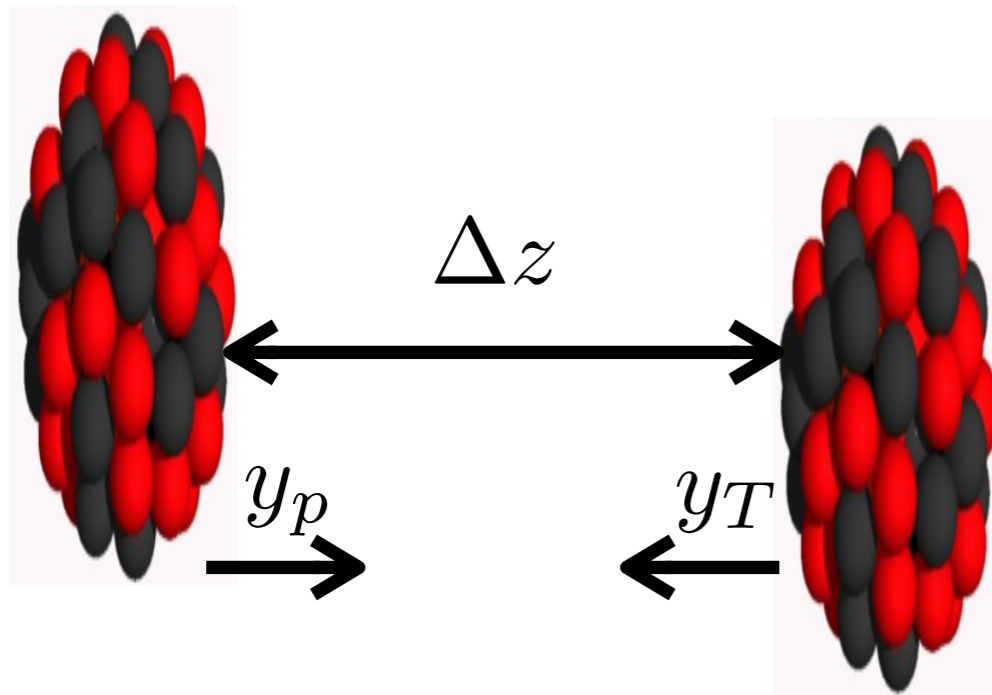
C. Shen, B. Schenke, in preparation



- Collision time and 3D spatial position are determined for every binary collision

The 3D MCGlauber-LEXUS model

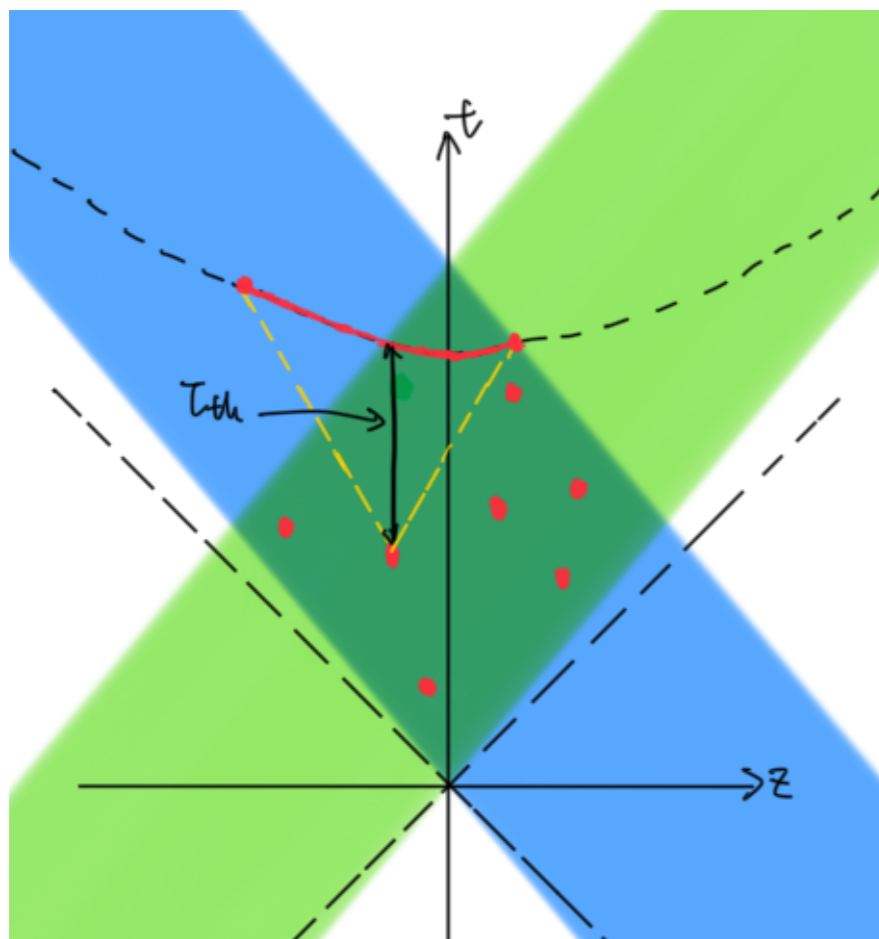
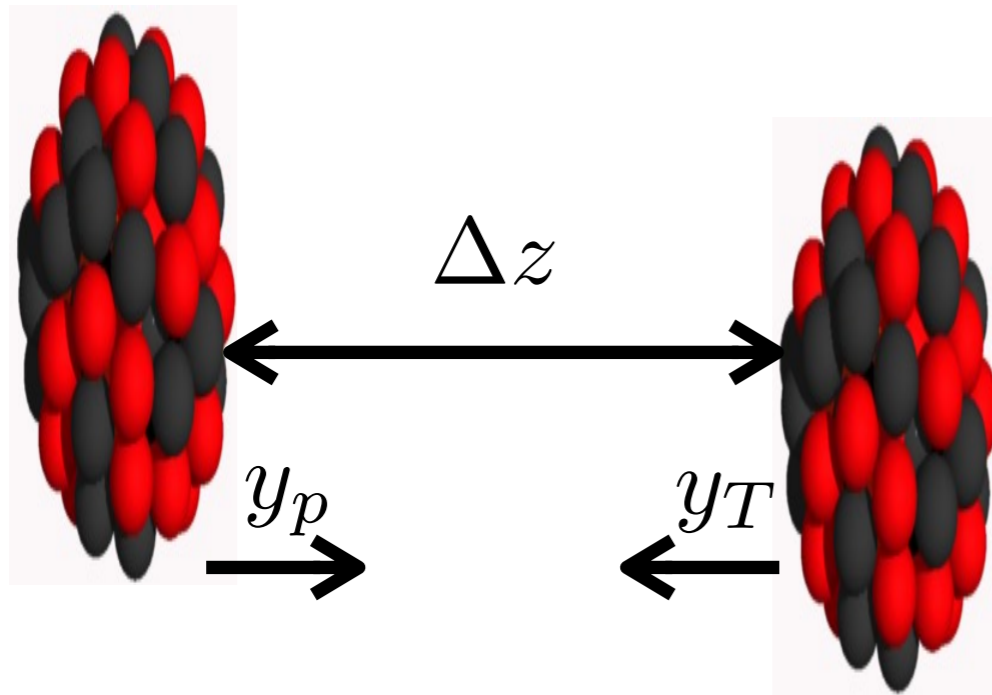
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The 3D MCGlauber-LEXUS model

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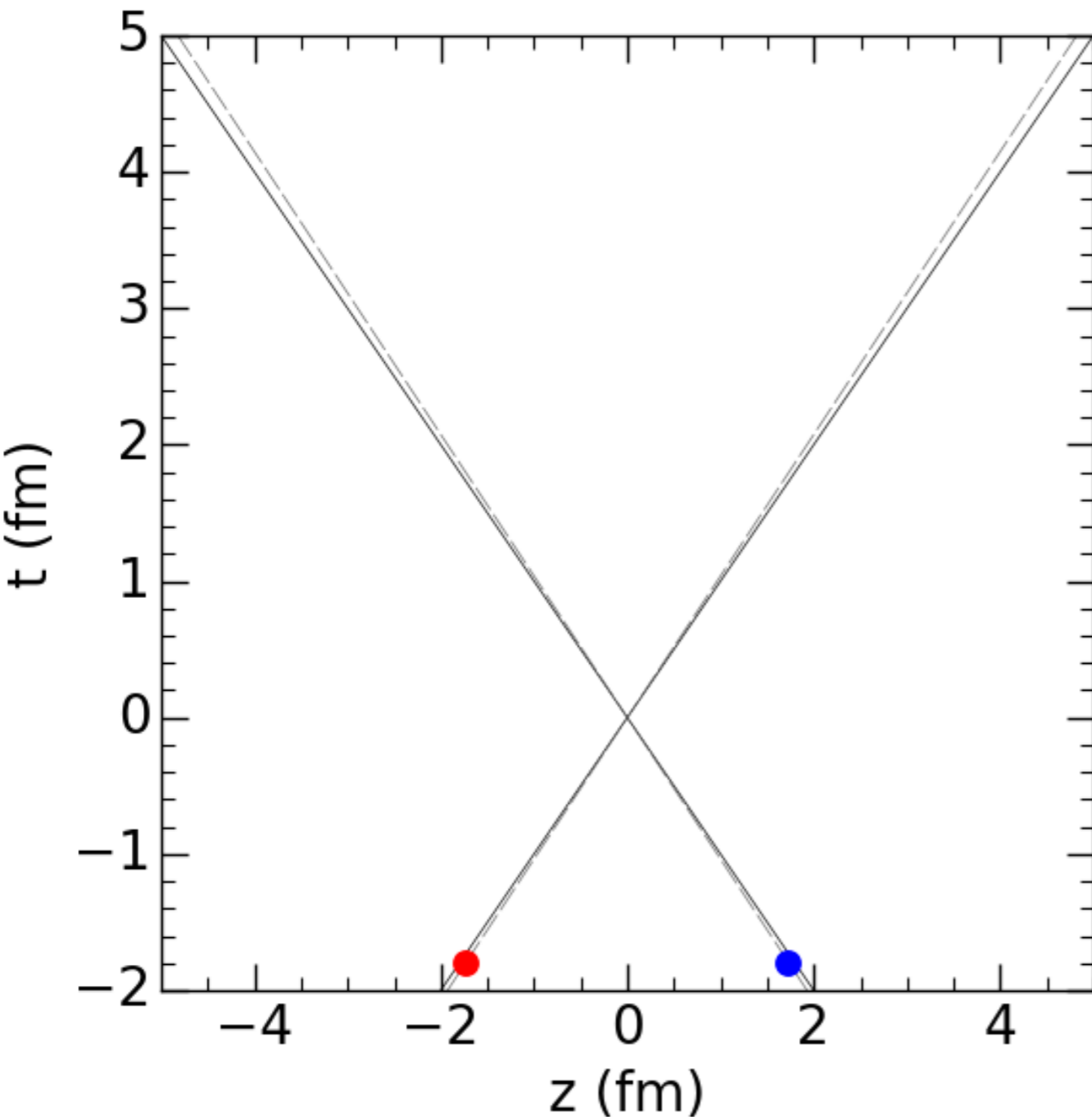
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A. Bialas, A. Bzdak and V. Koch,
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- These strings are decelerated with a constant string tension $\sigma = 1 \text{ GeV}/\text{fm}$ before thermalized to medium

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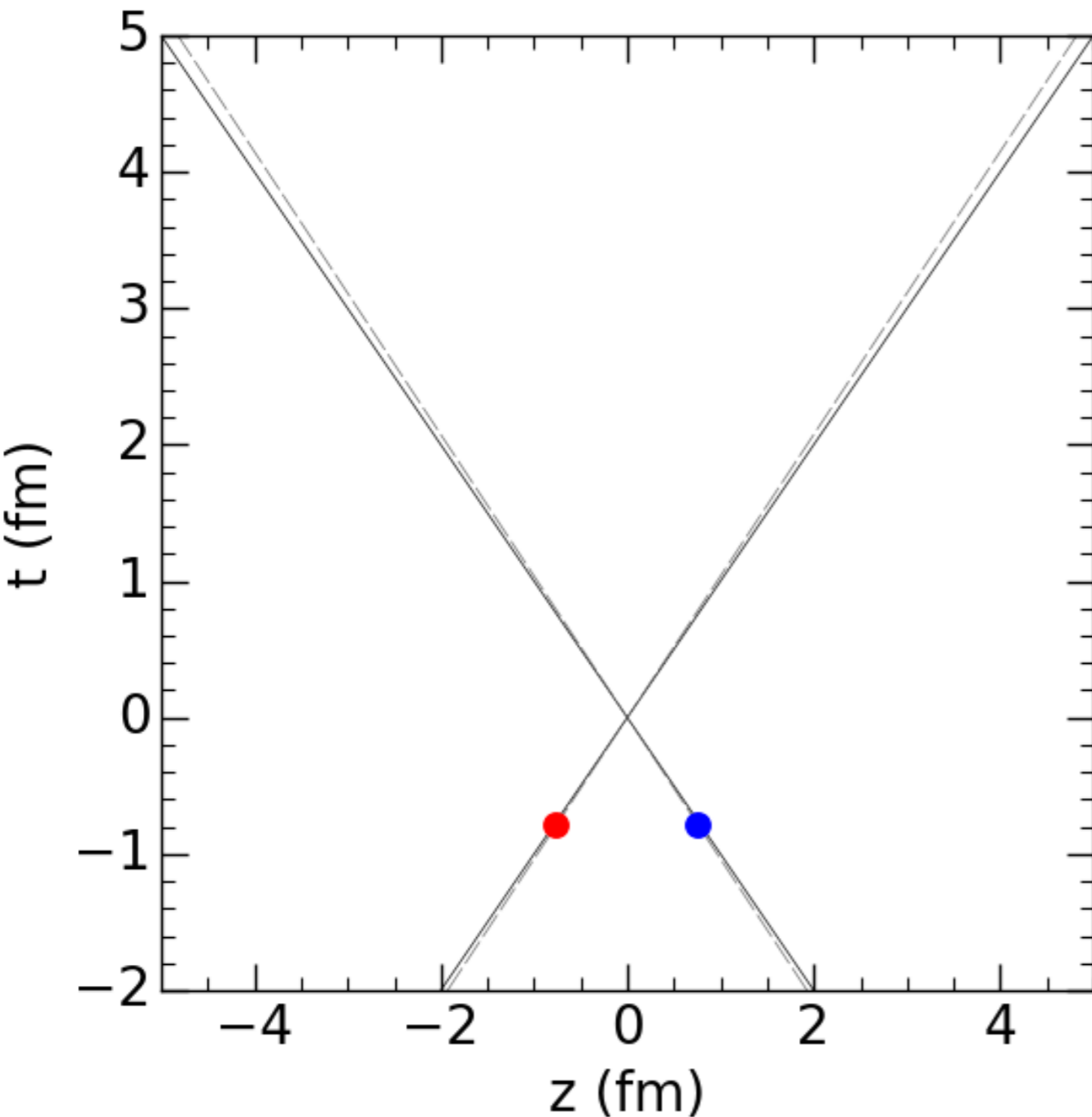
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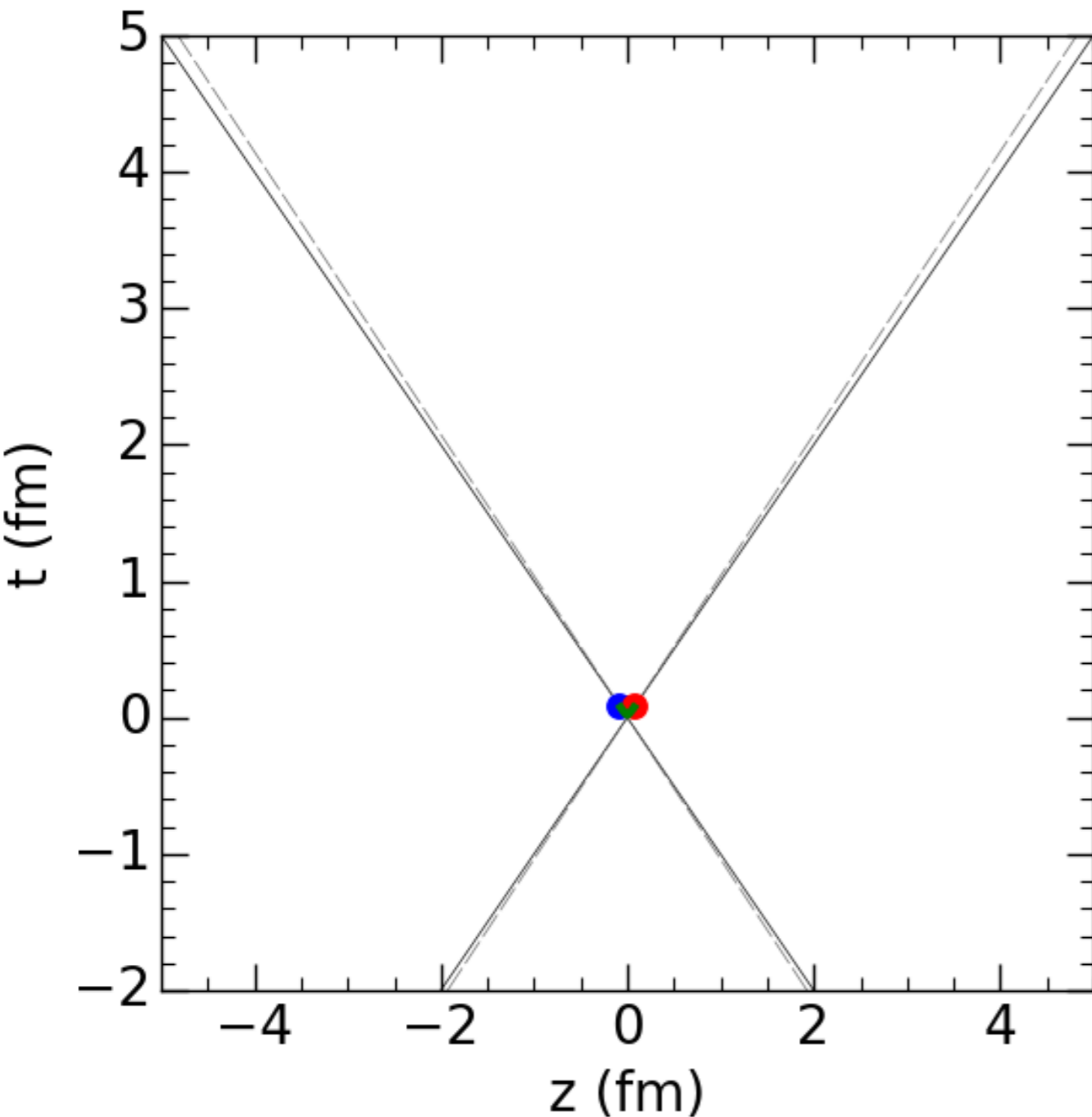
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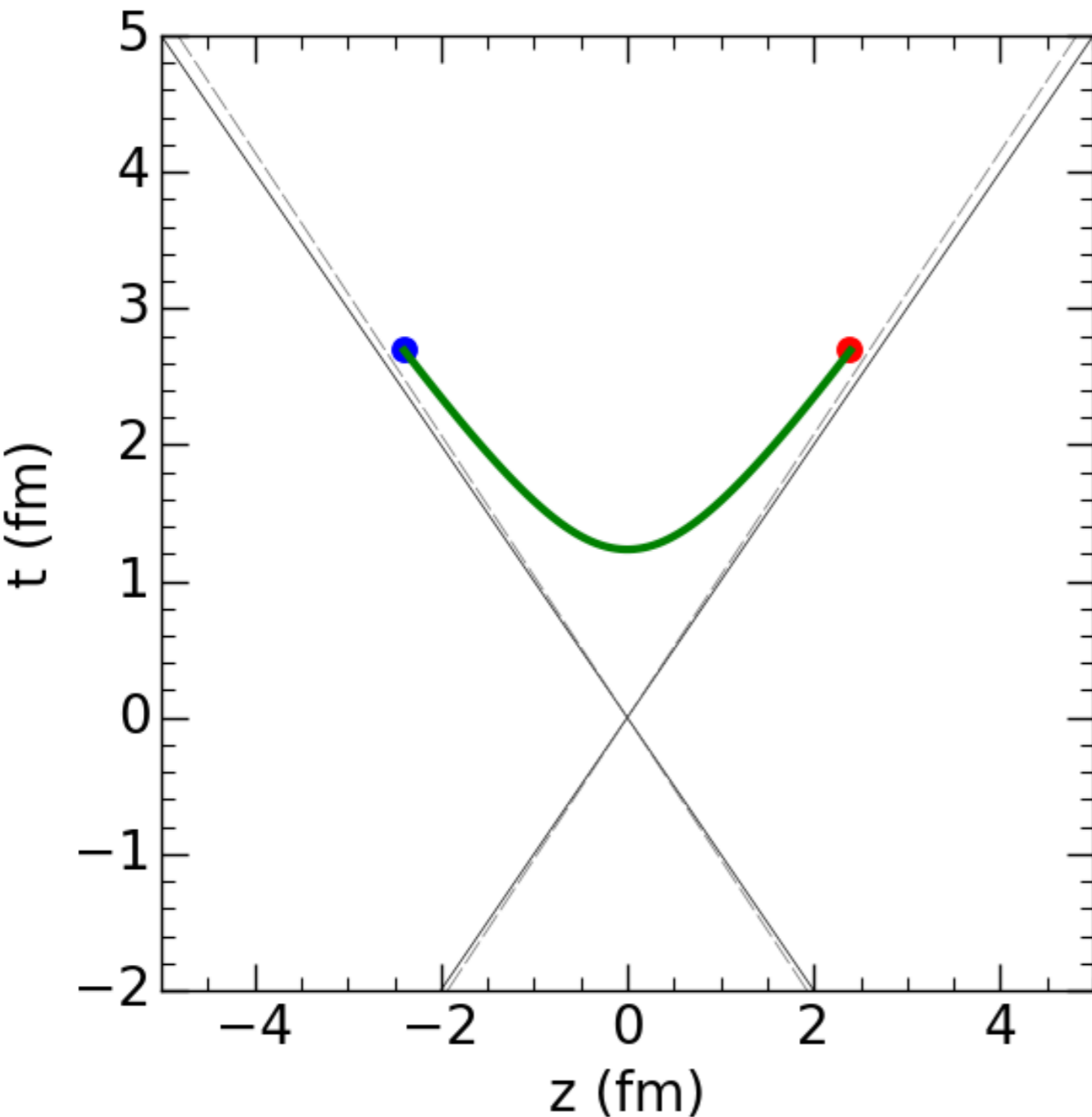
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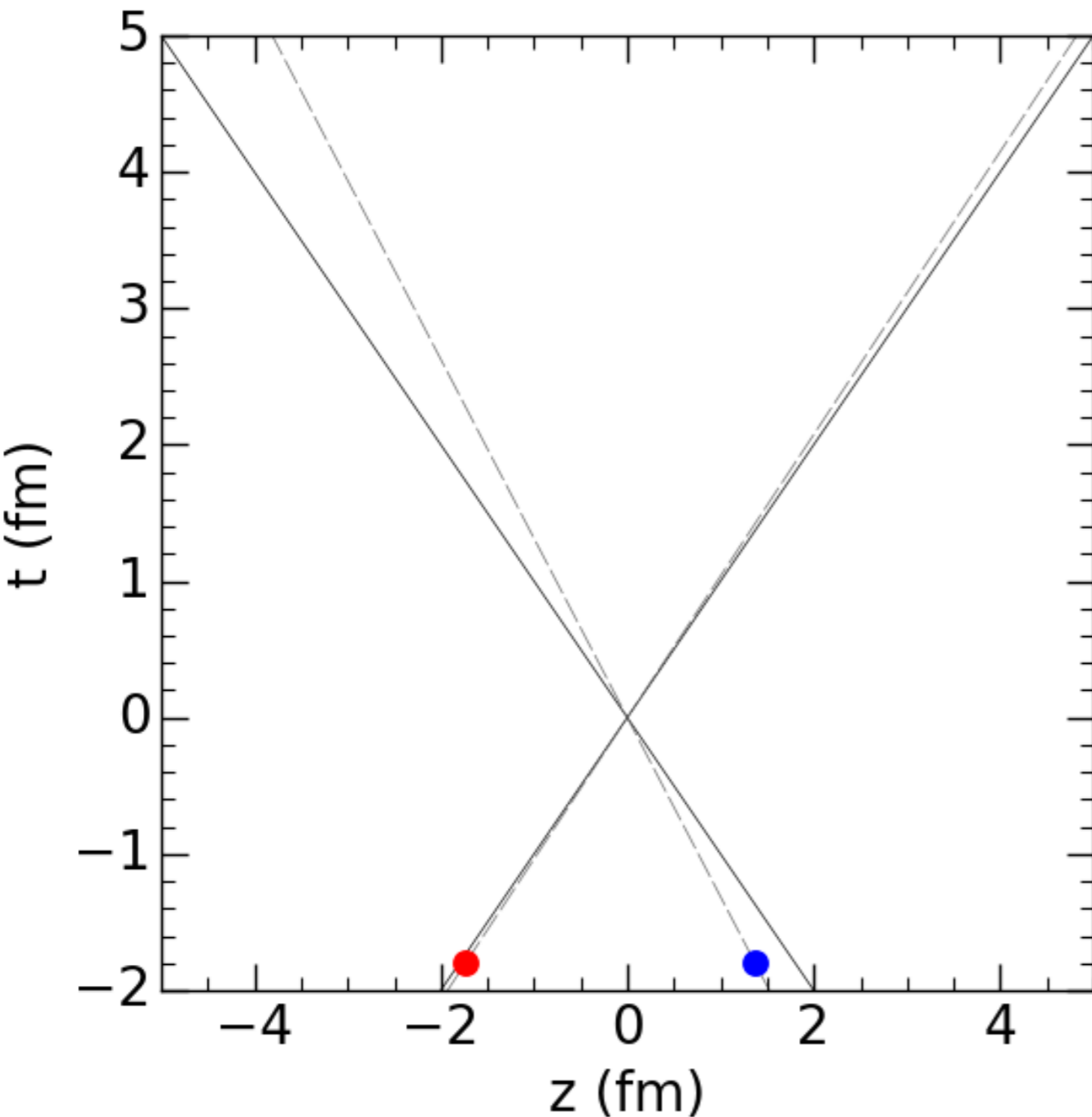
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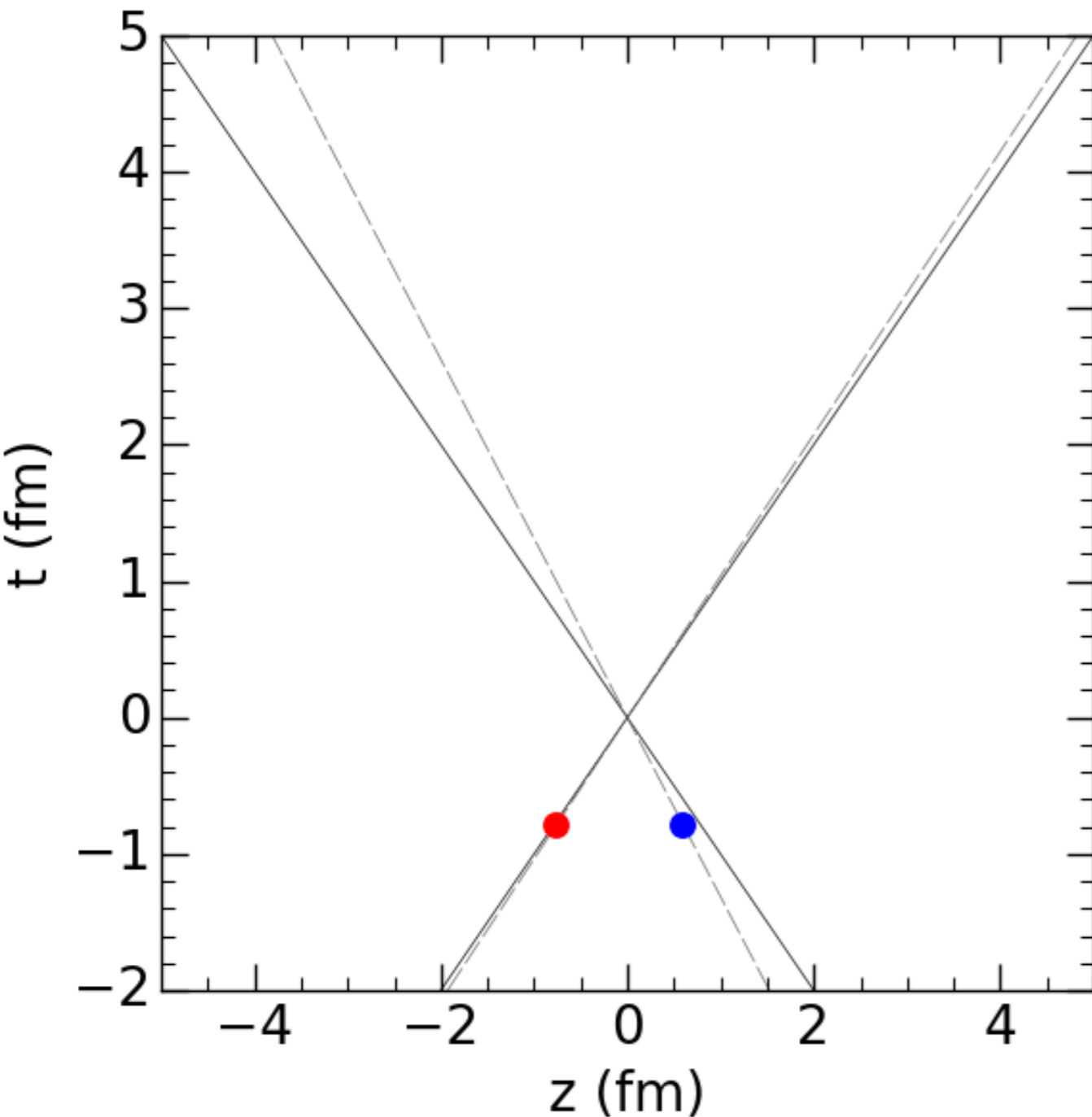
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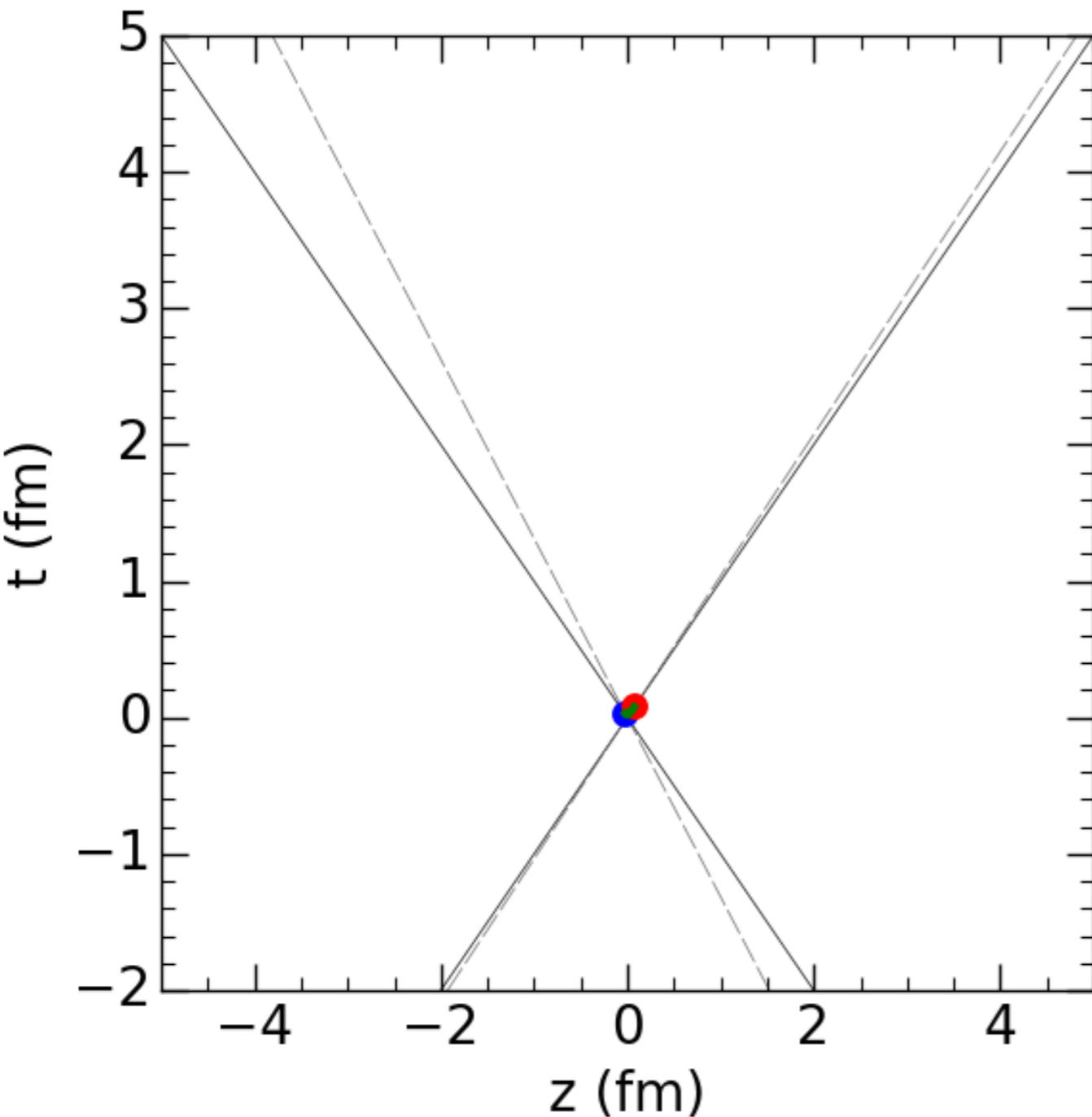
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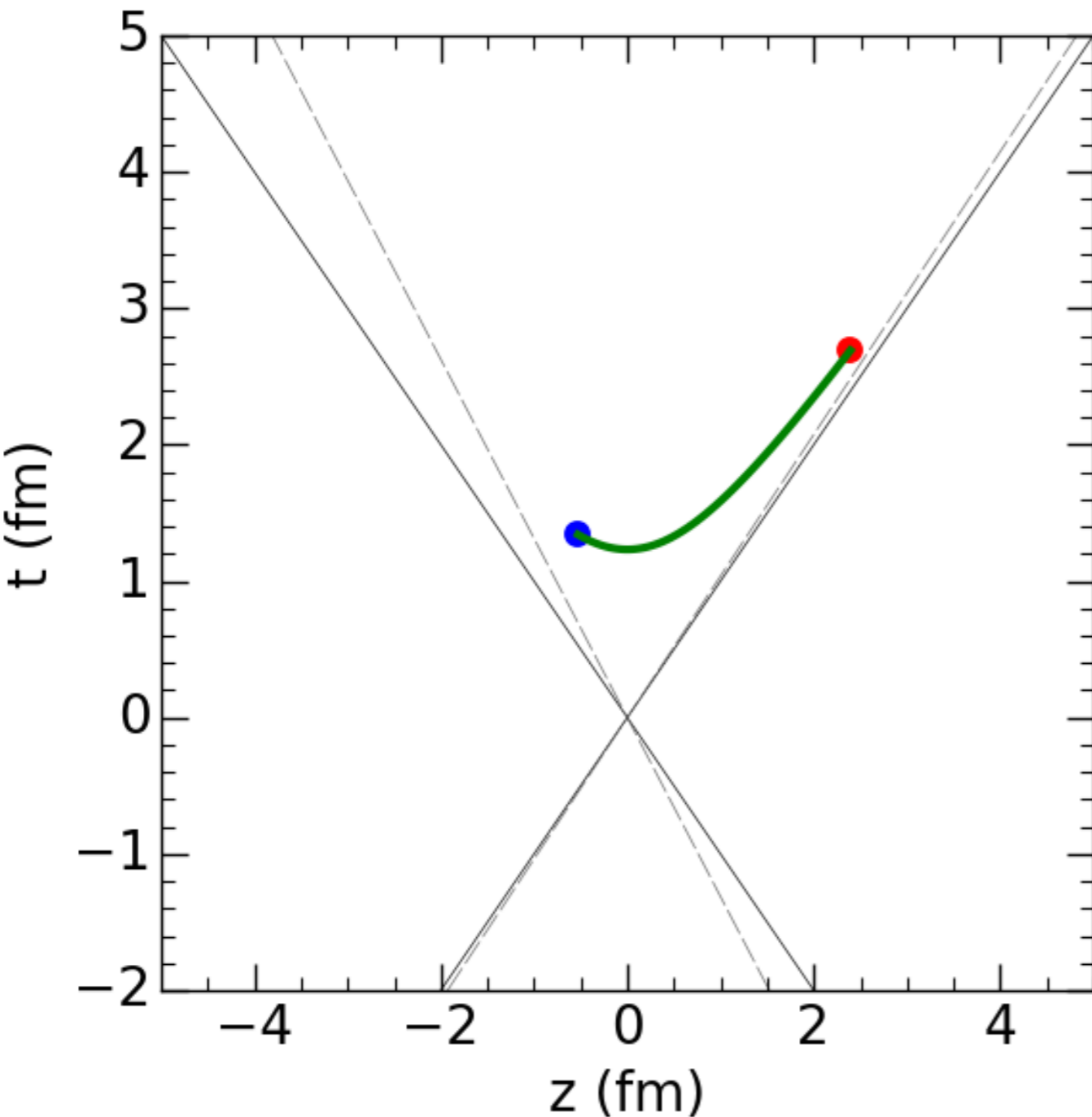
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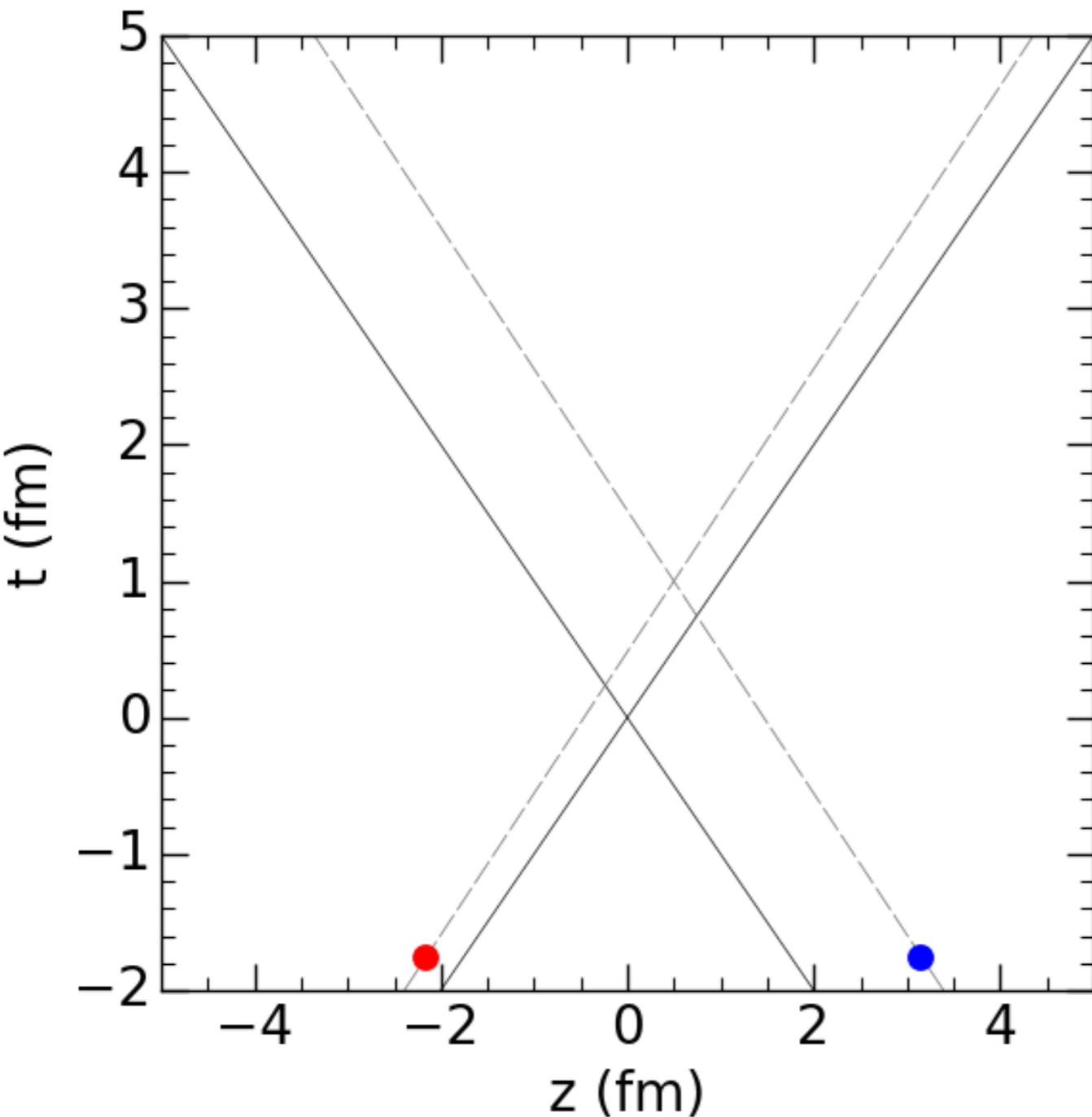
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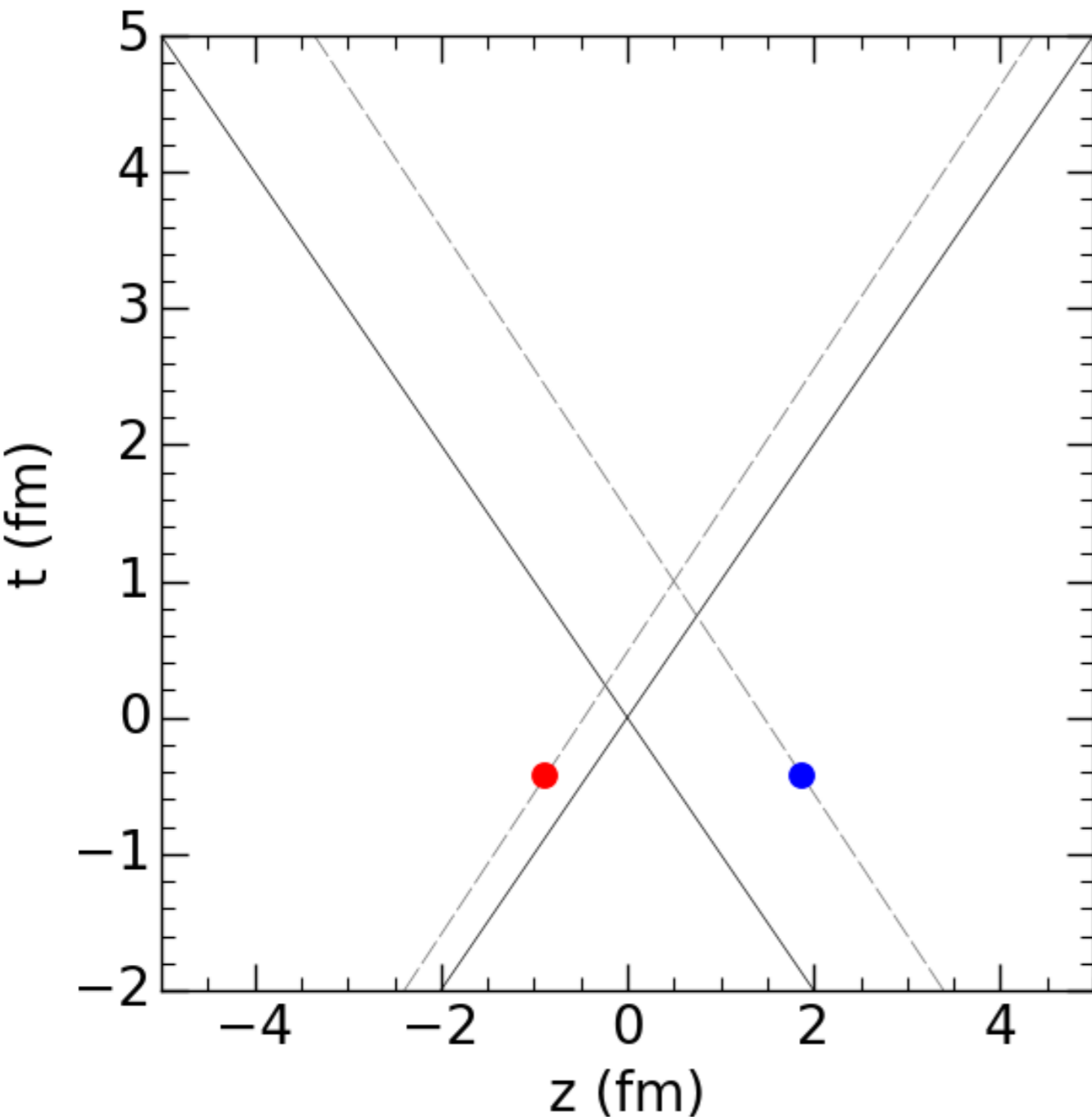
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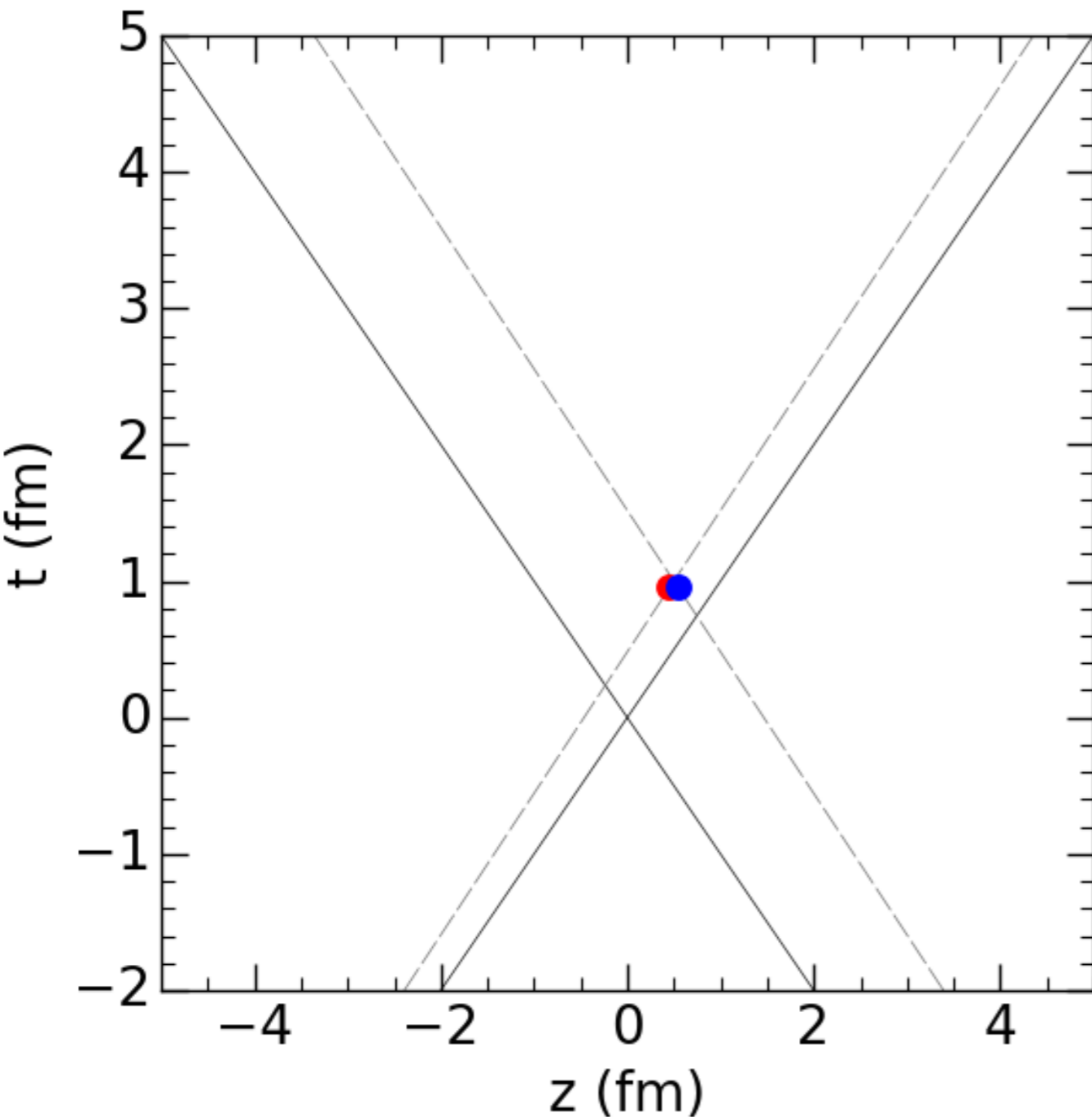
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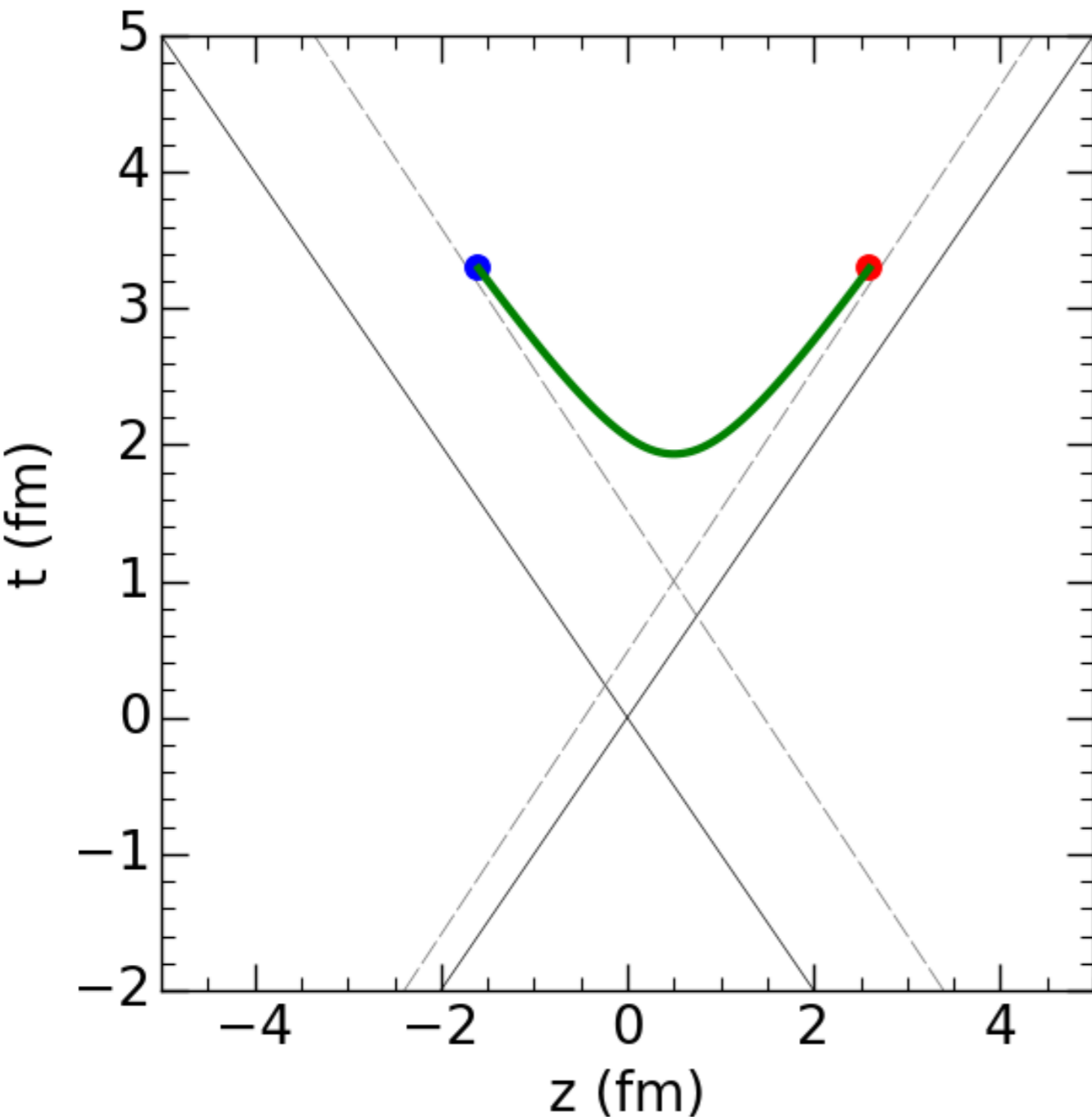
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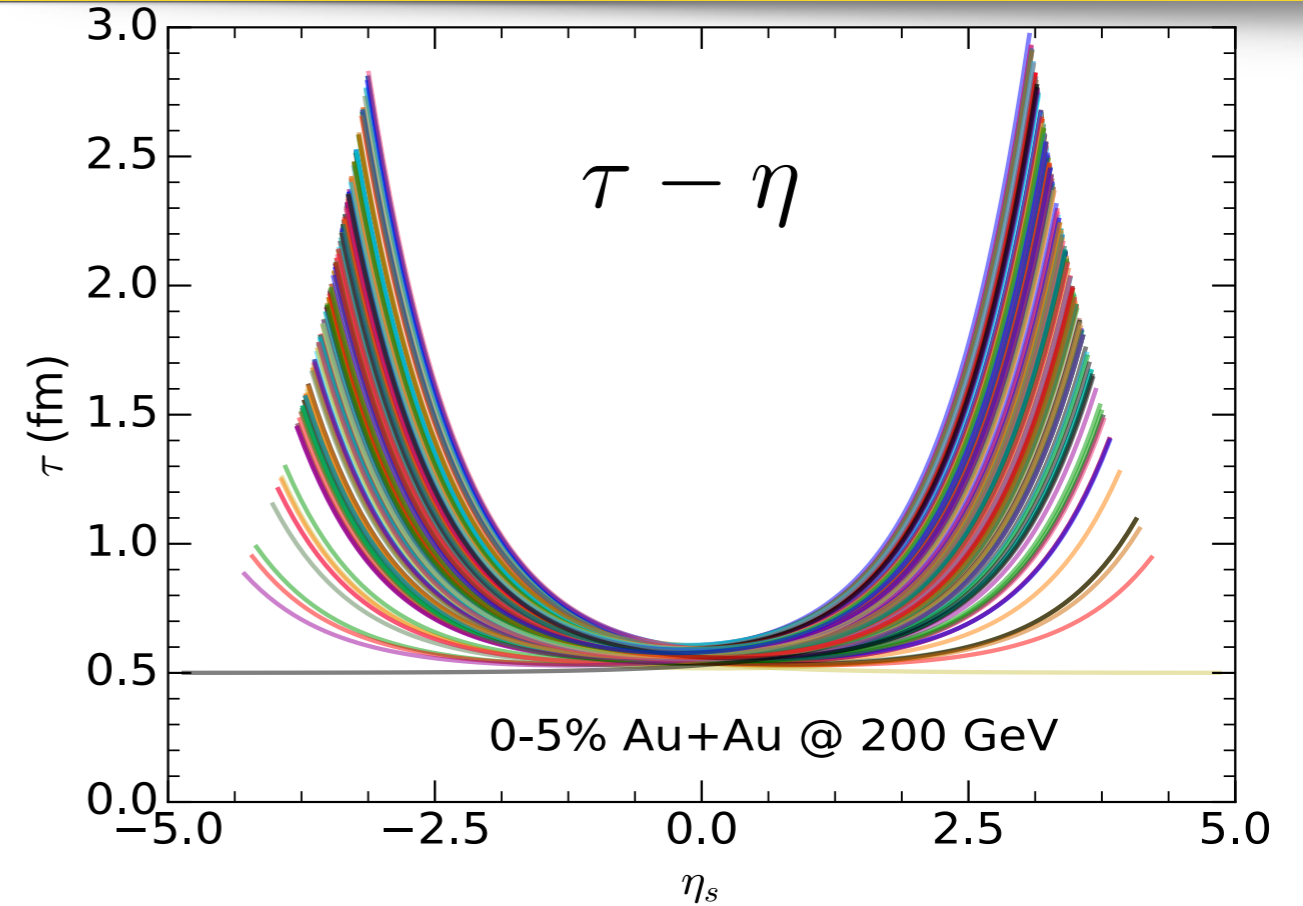
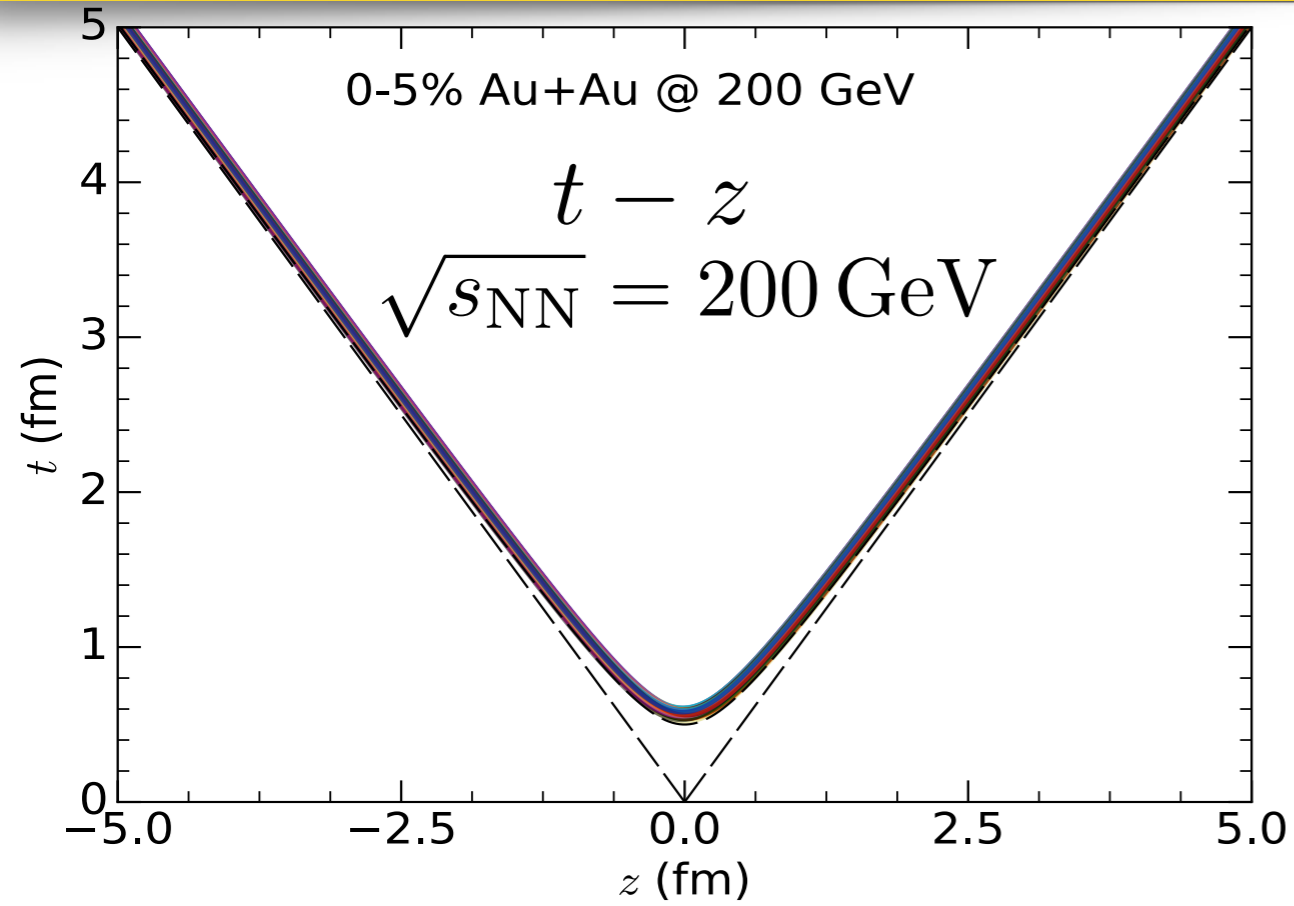


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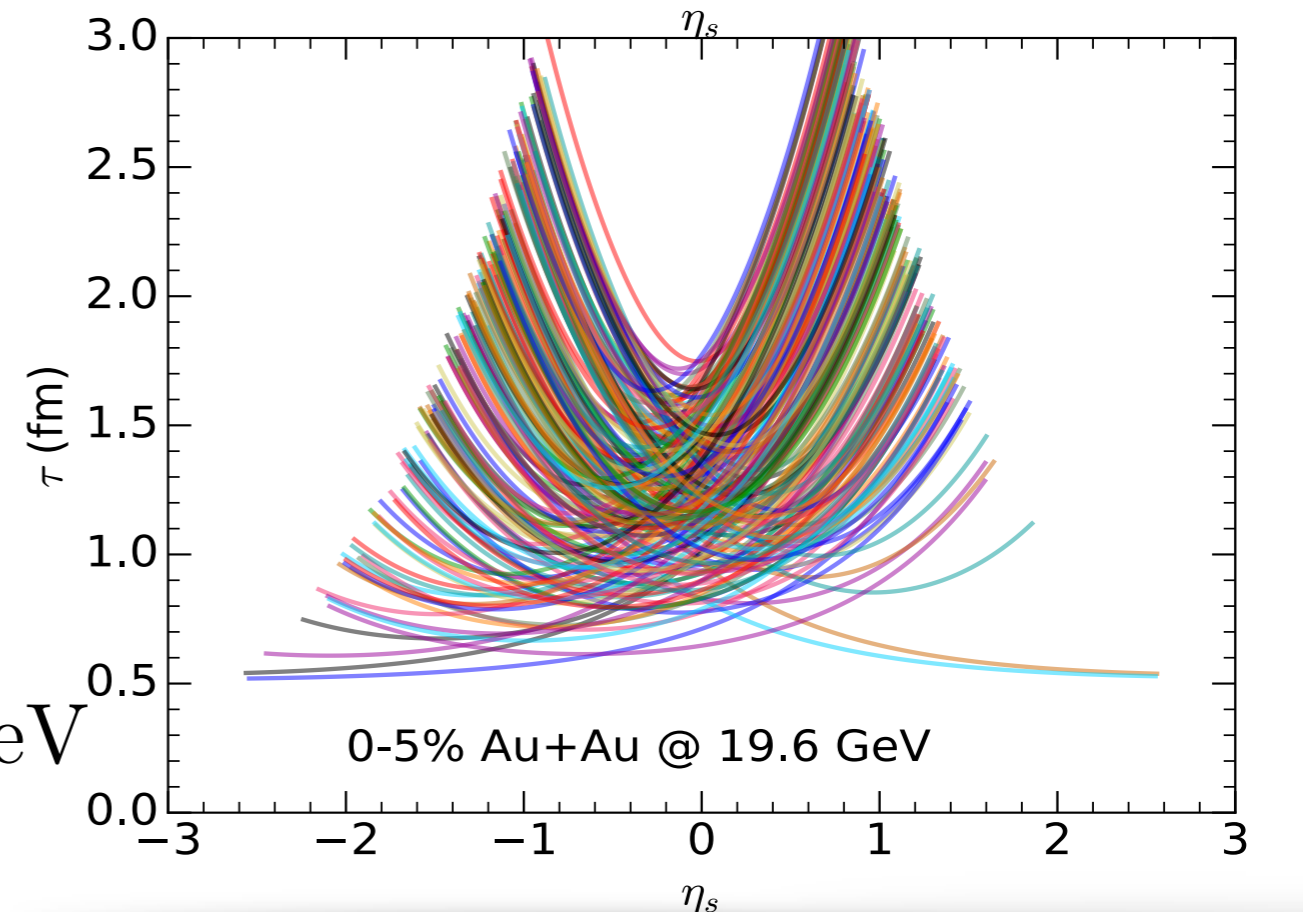
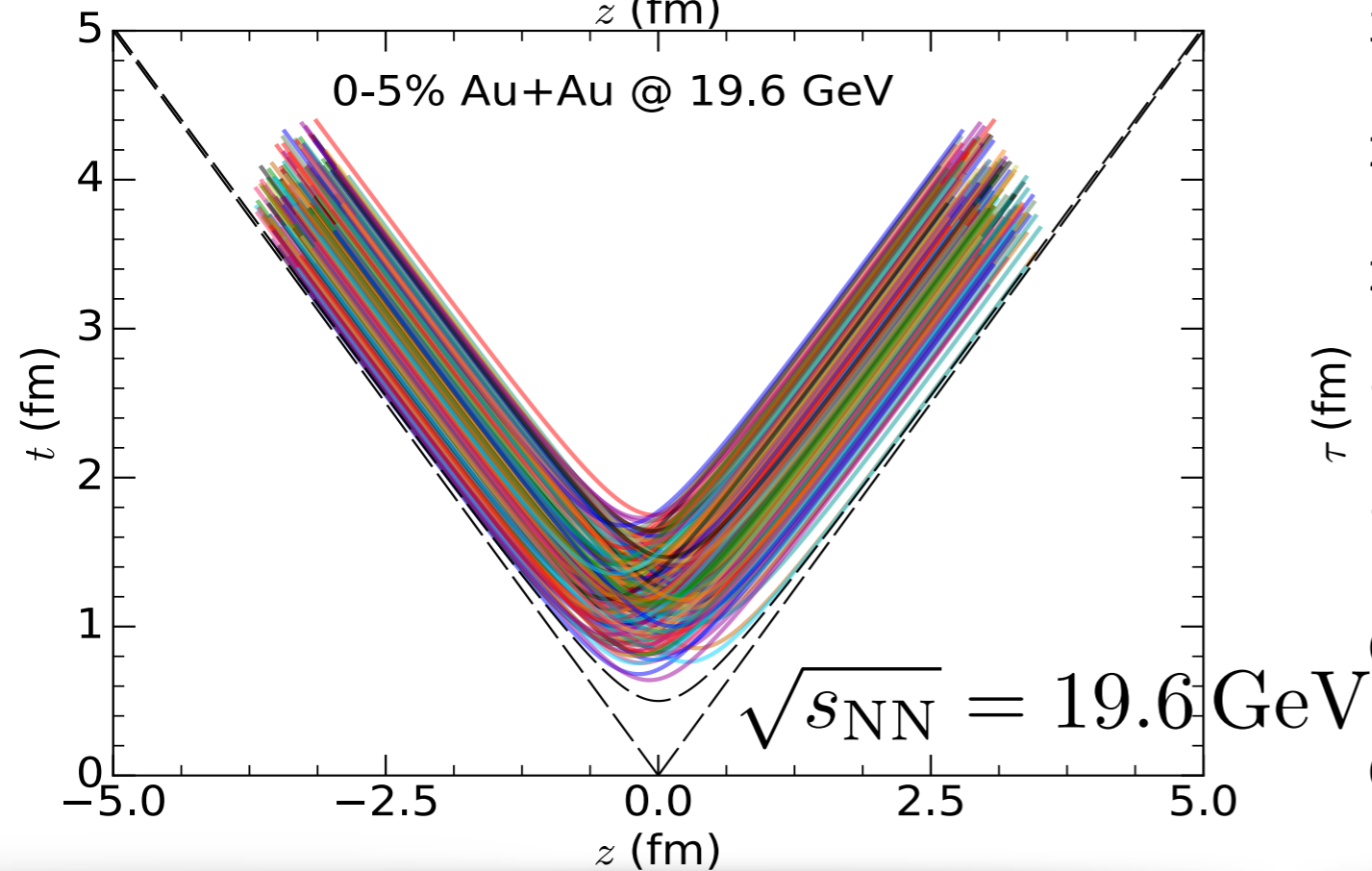
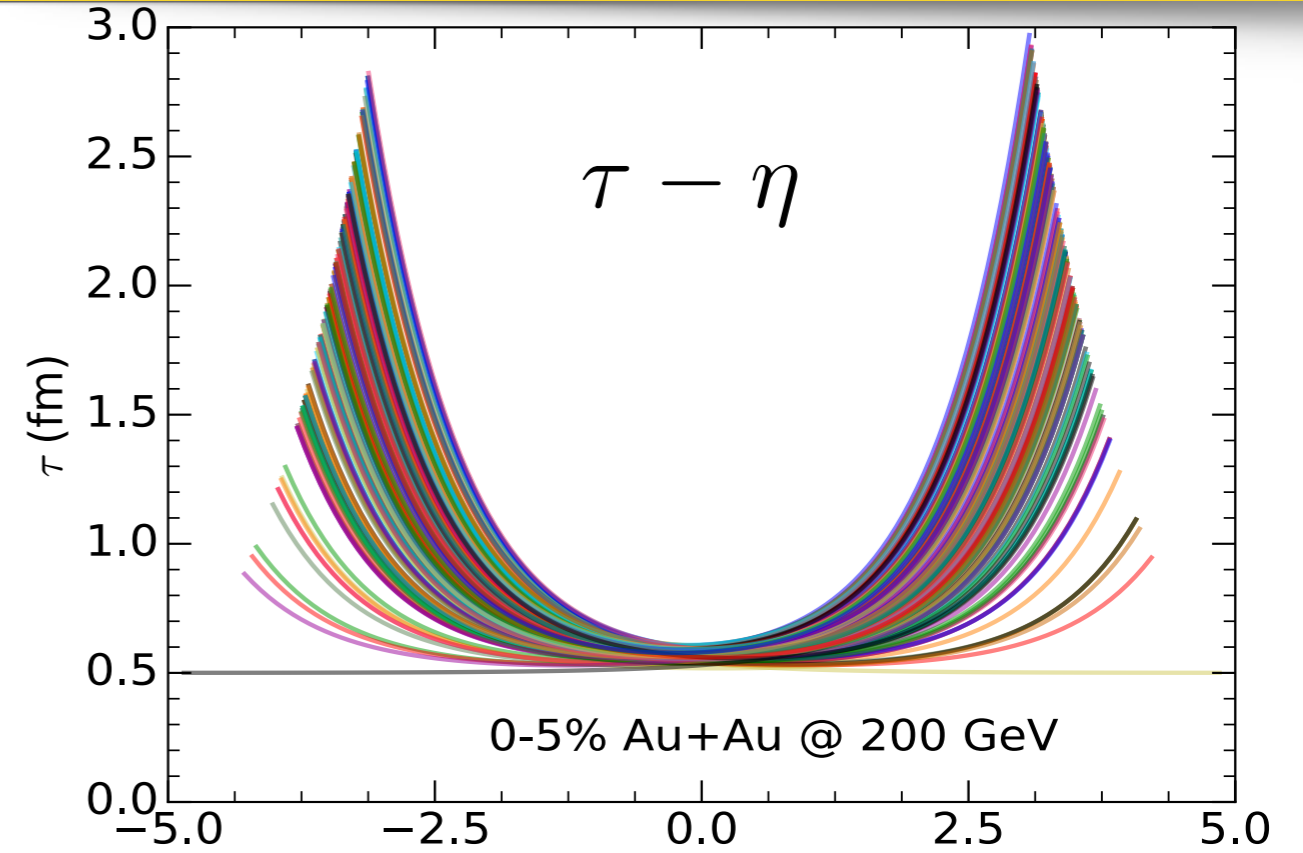
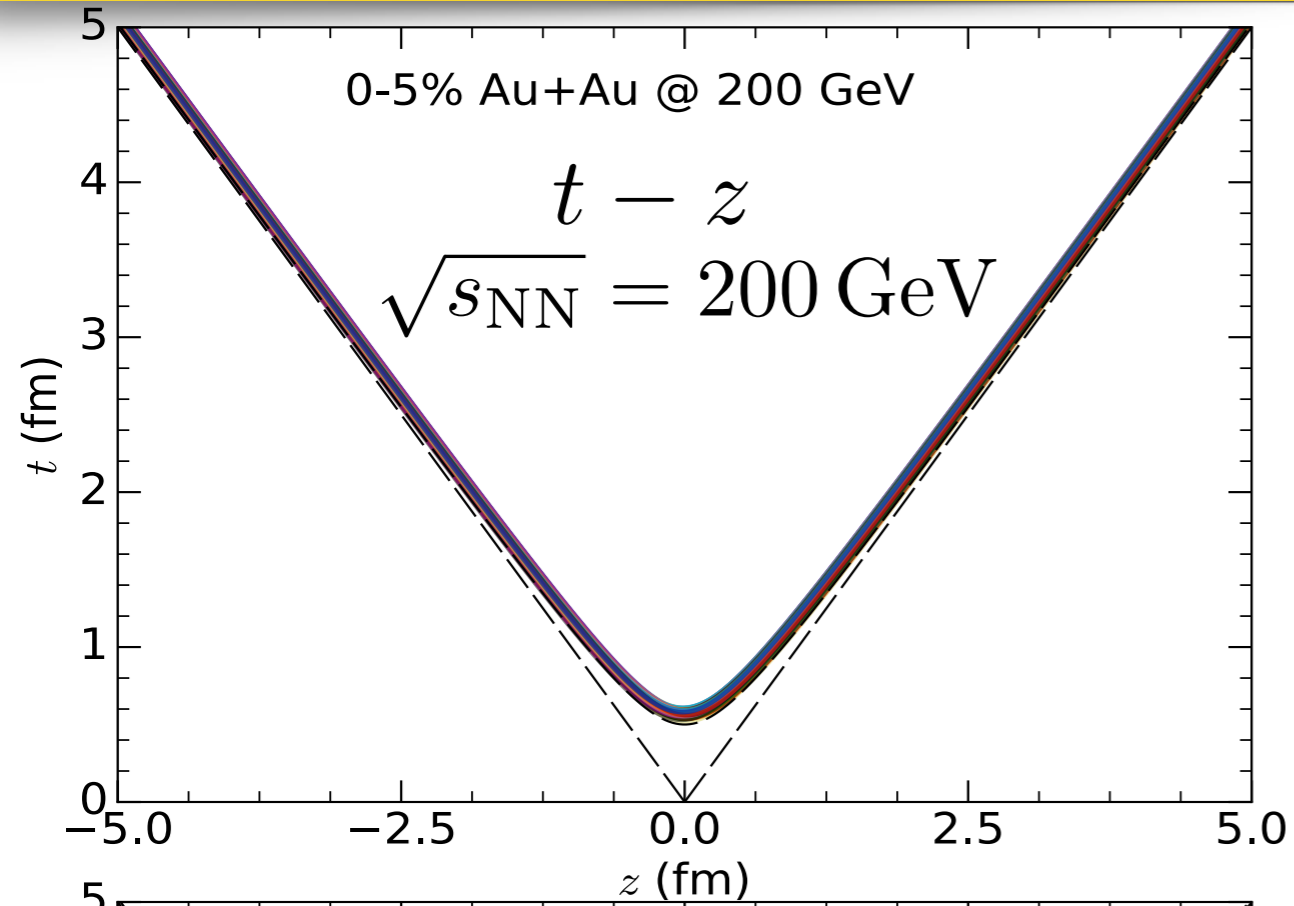
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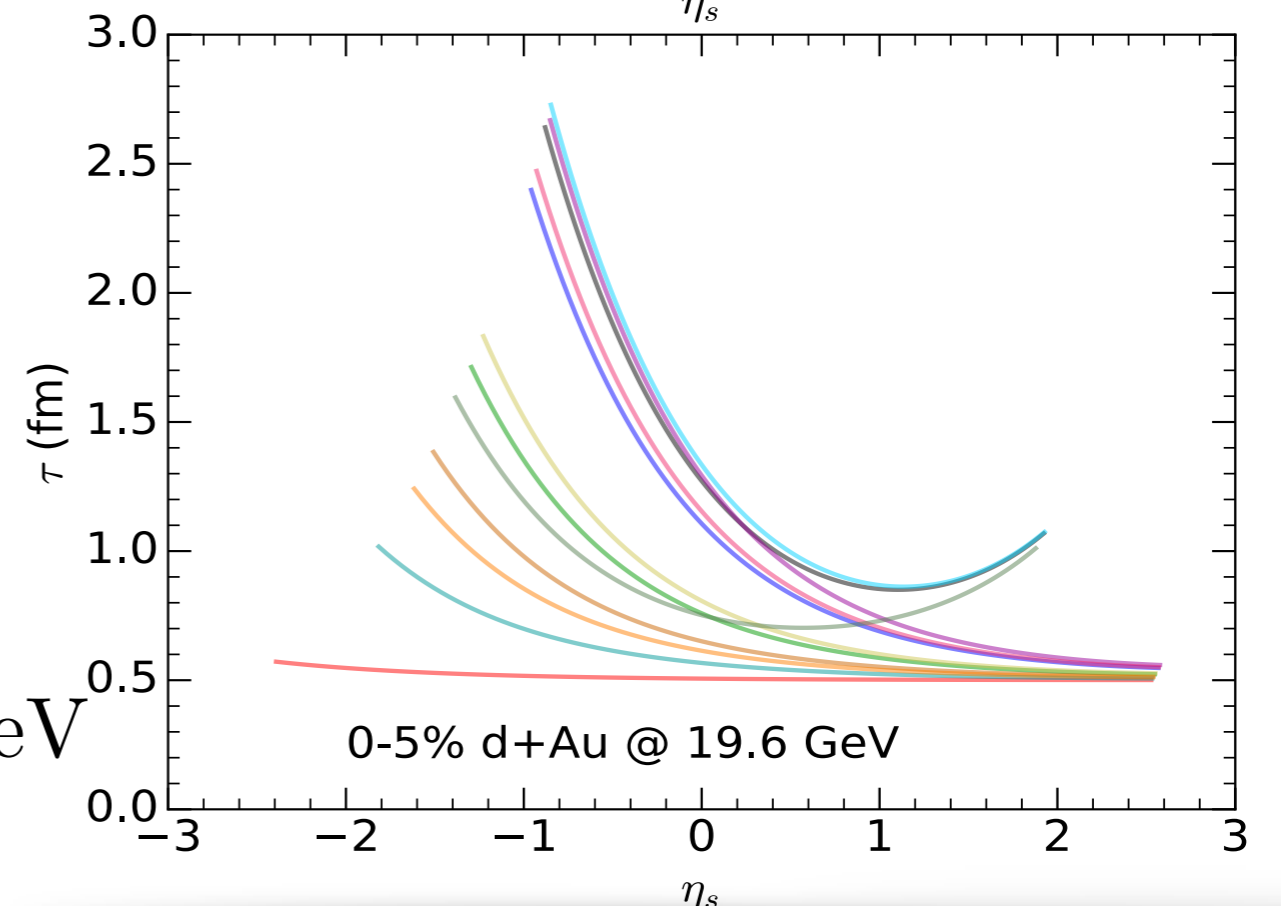
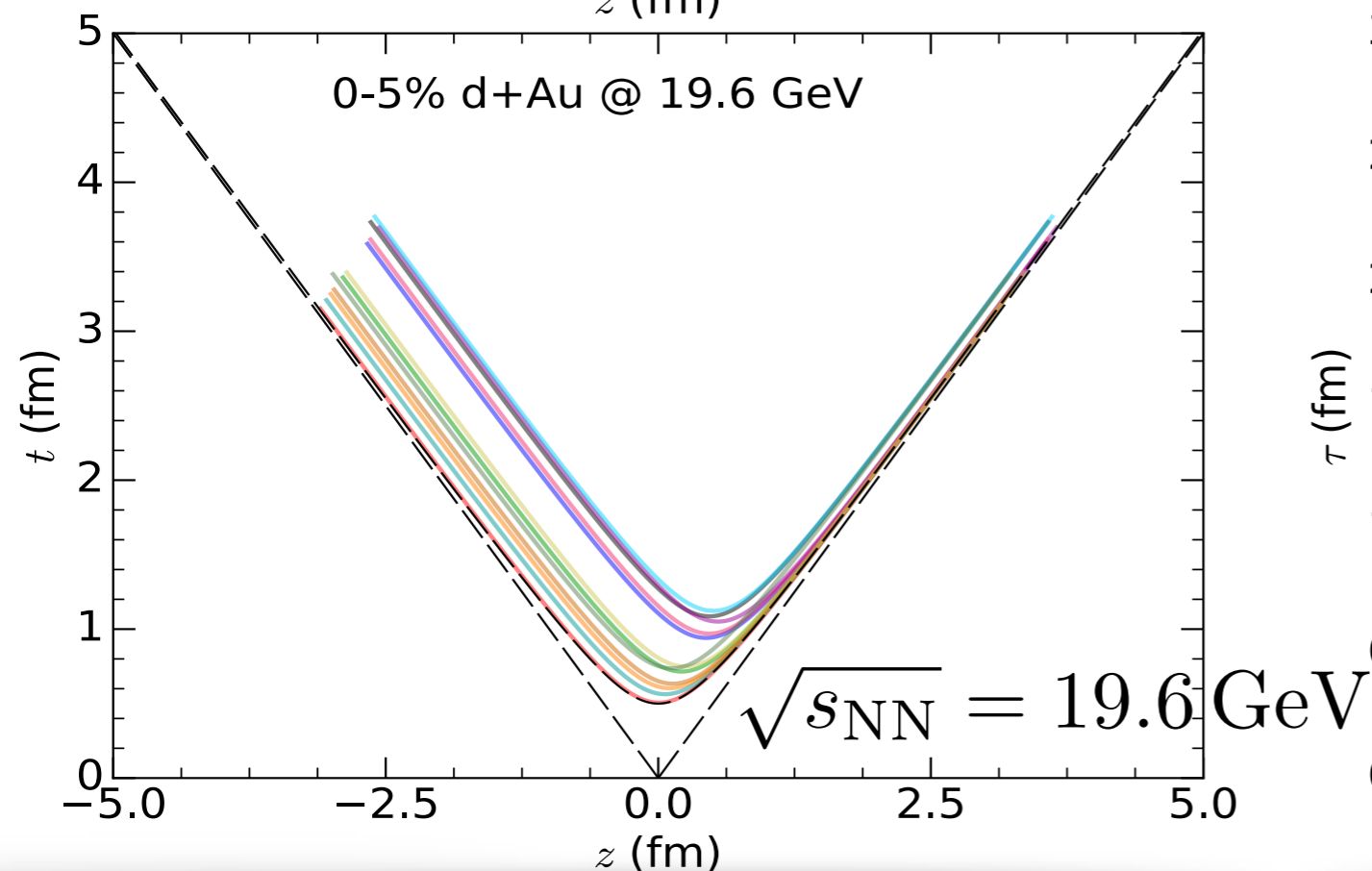
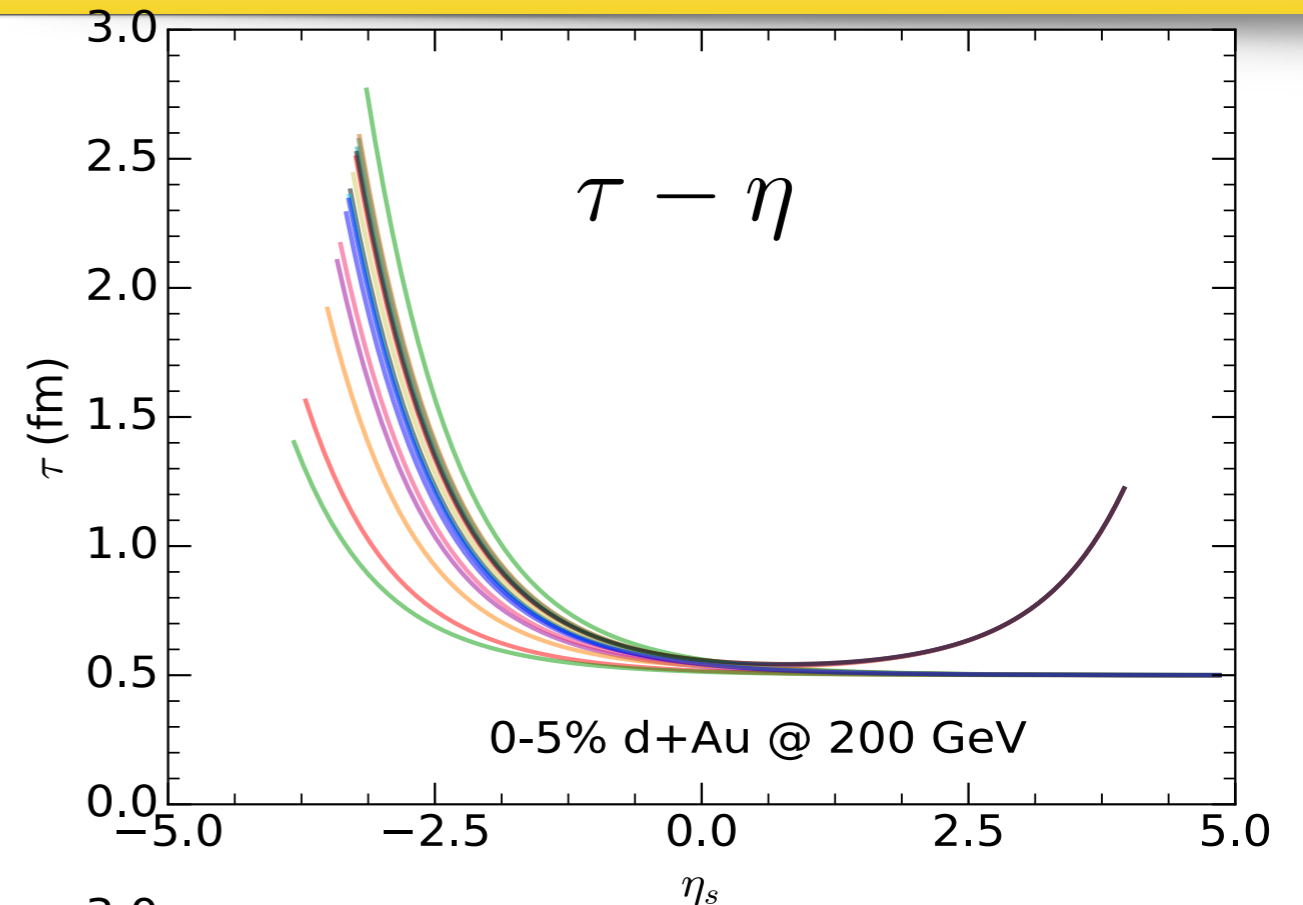
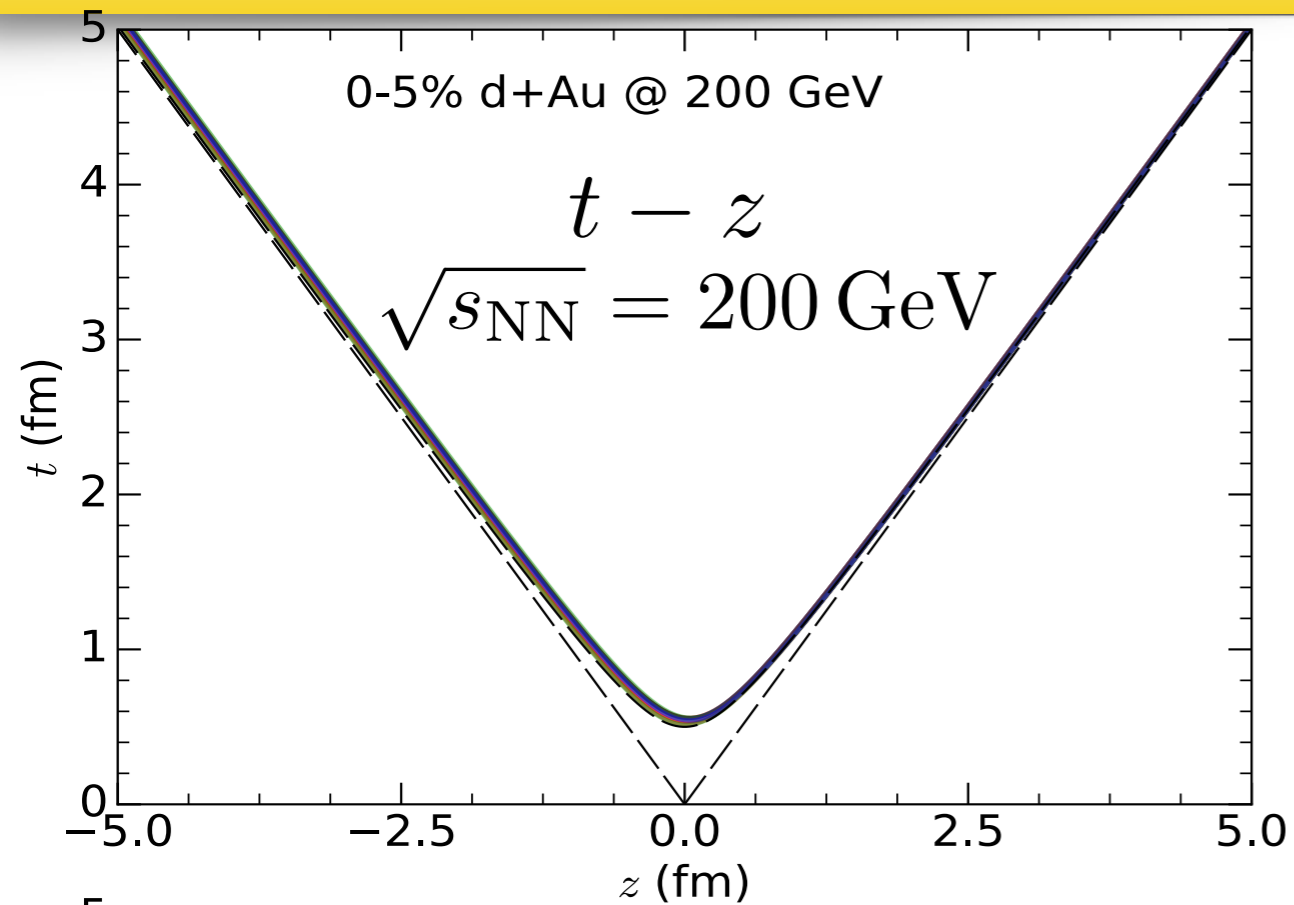
A 3D MCGlauber model



A 3D MCGlauber model



A 3D MCGlauber model

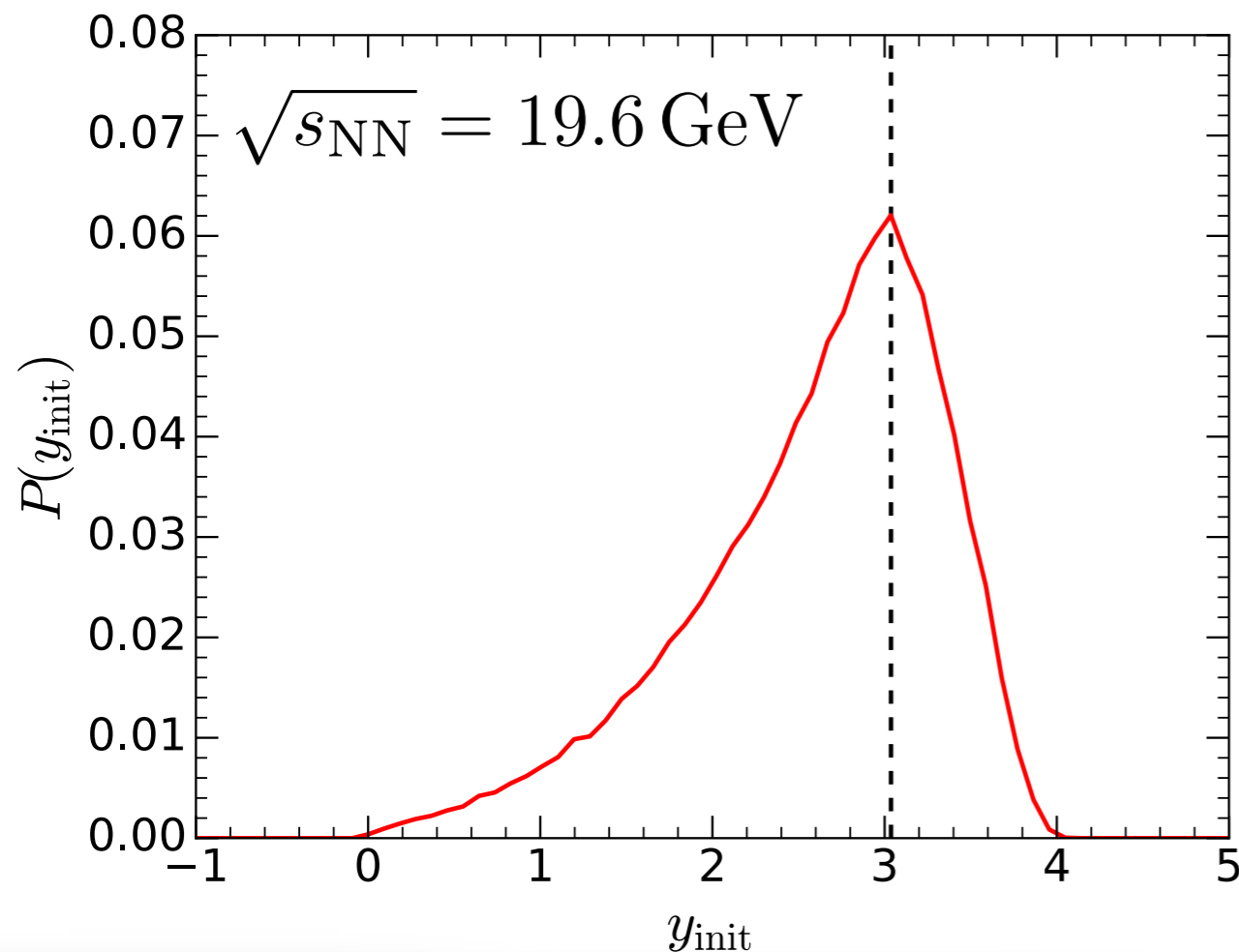


Introducing longitudinal fluctuations

- Sample valence quarks from the incoming participants

$$y_q = \operatorname{arcsinh} \left(x_q \sqrt{\frac{s}{4m_q^2} - 1} \right)$$

$$y_q = \log \left(\frac{x\sqrt{s}}{2m_q} \right)$$



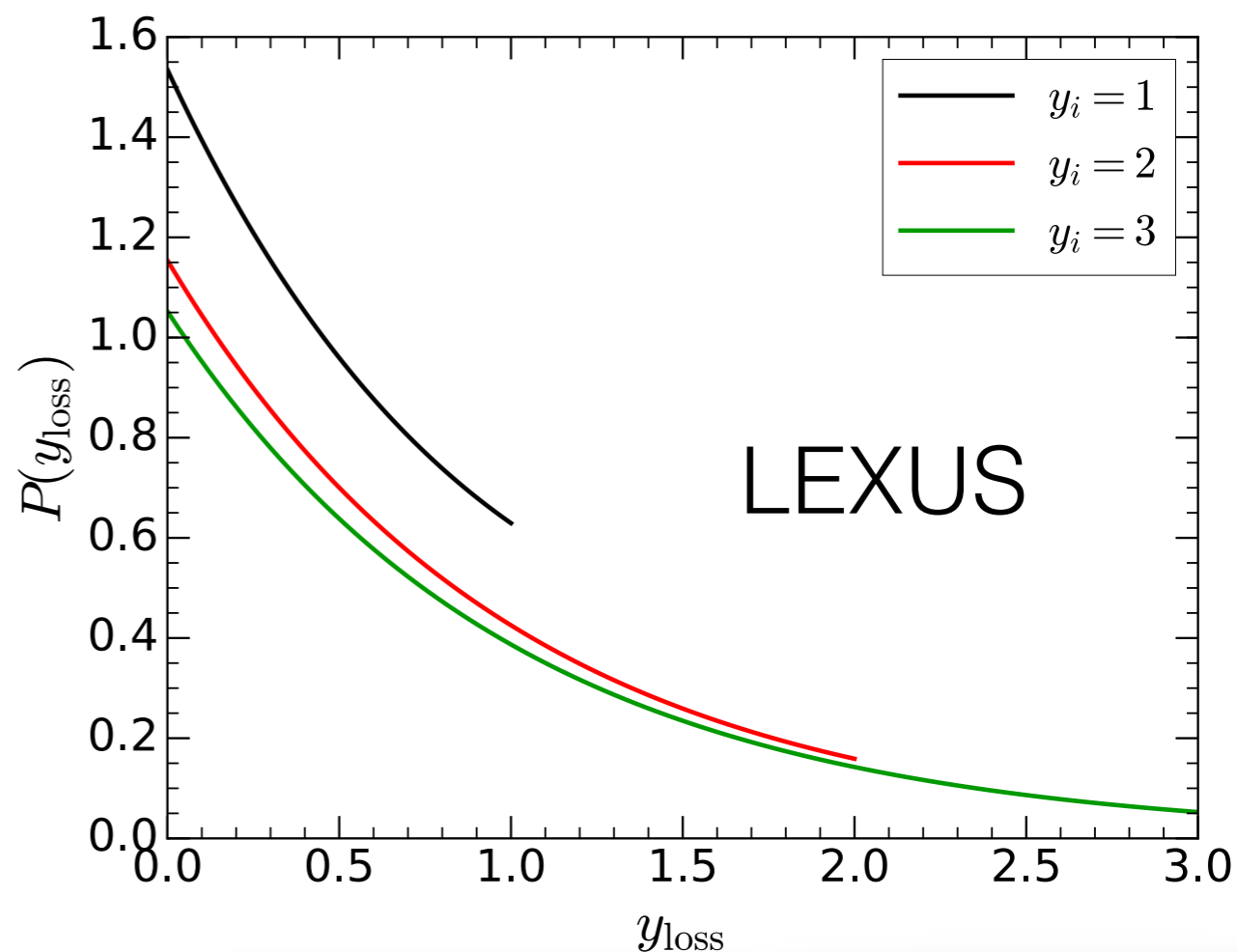
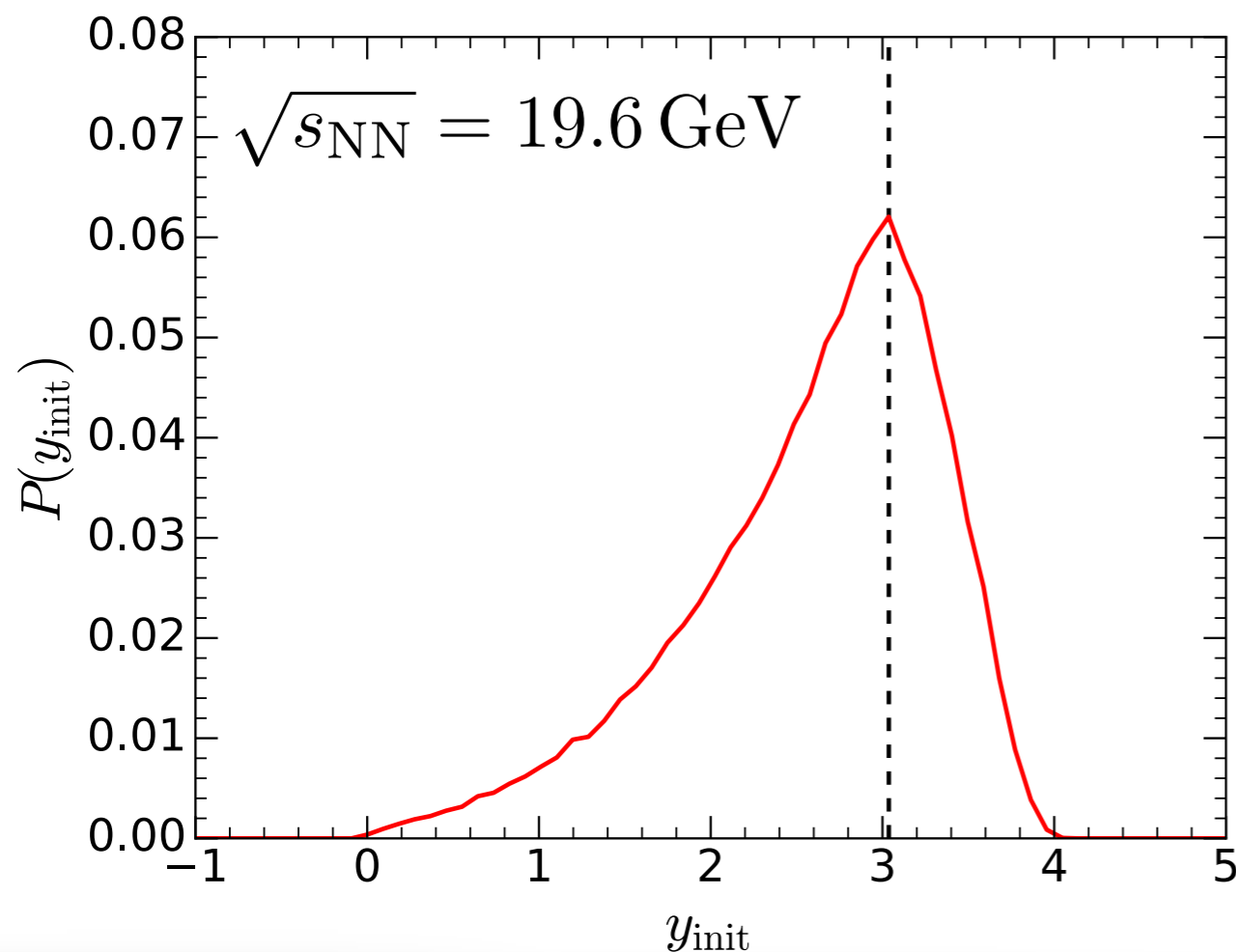
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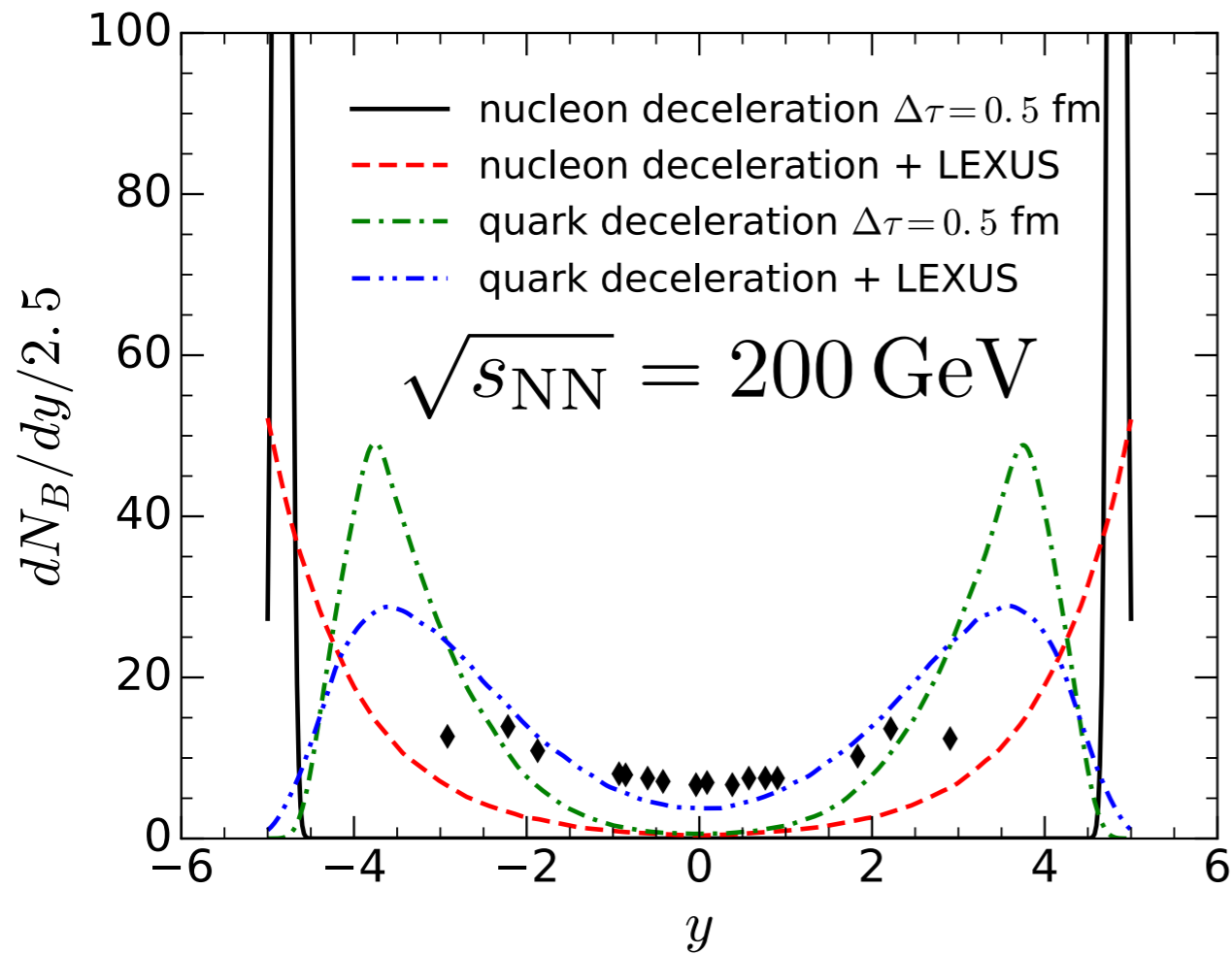
- Sample the rapidity loss according to the LEXUS model

$$P(y_{\text{loss}}) = \frac{\cosh(2y_{\text{init}} - y_{\text{loss}})}{\sinh(2y_{\text{init}}) - \sinh(y_{\text{init}})} \quad y_{\text{loss}} \in [0, y_{\text{init}}]$$



Net baryon rapidity distribution

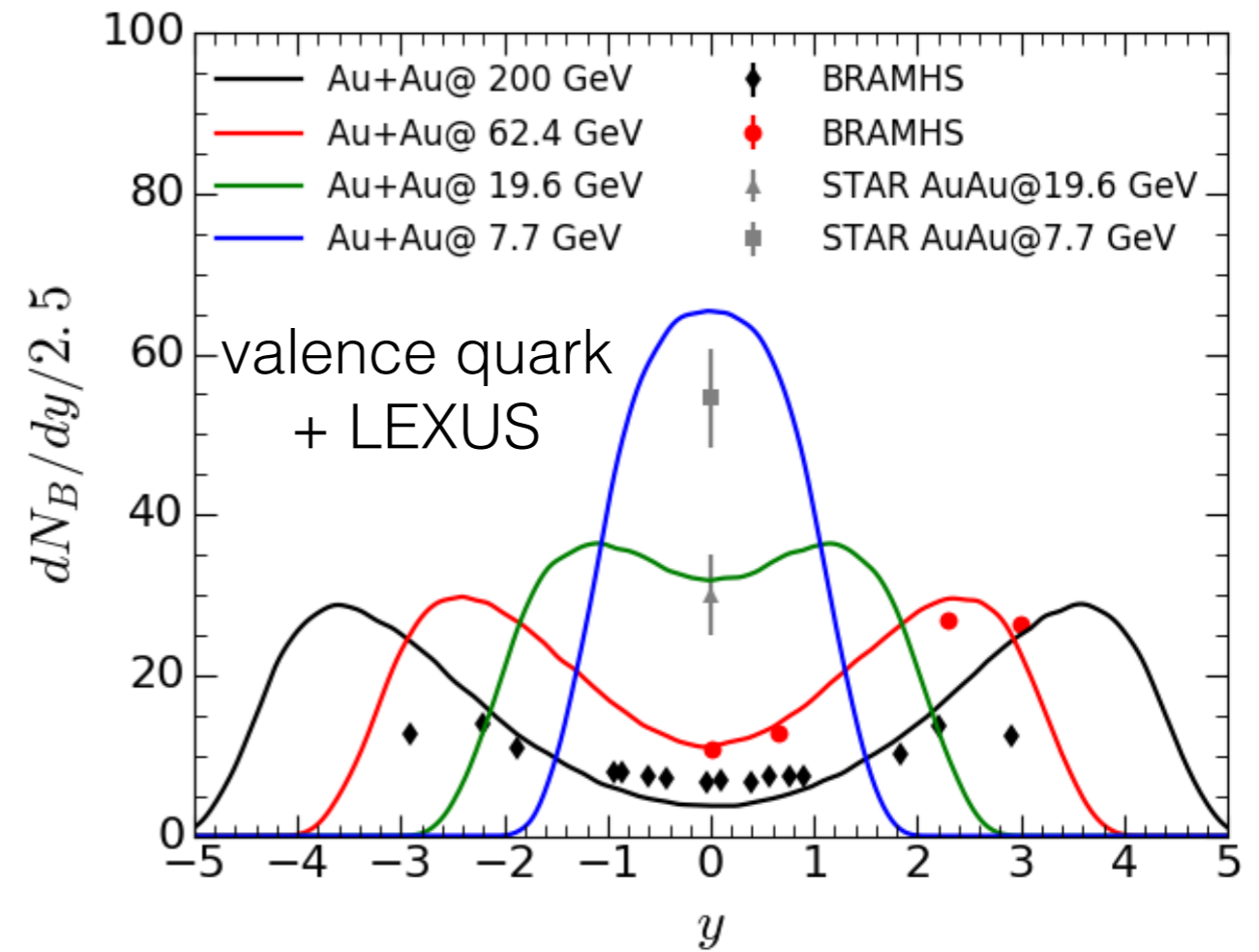
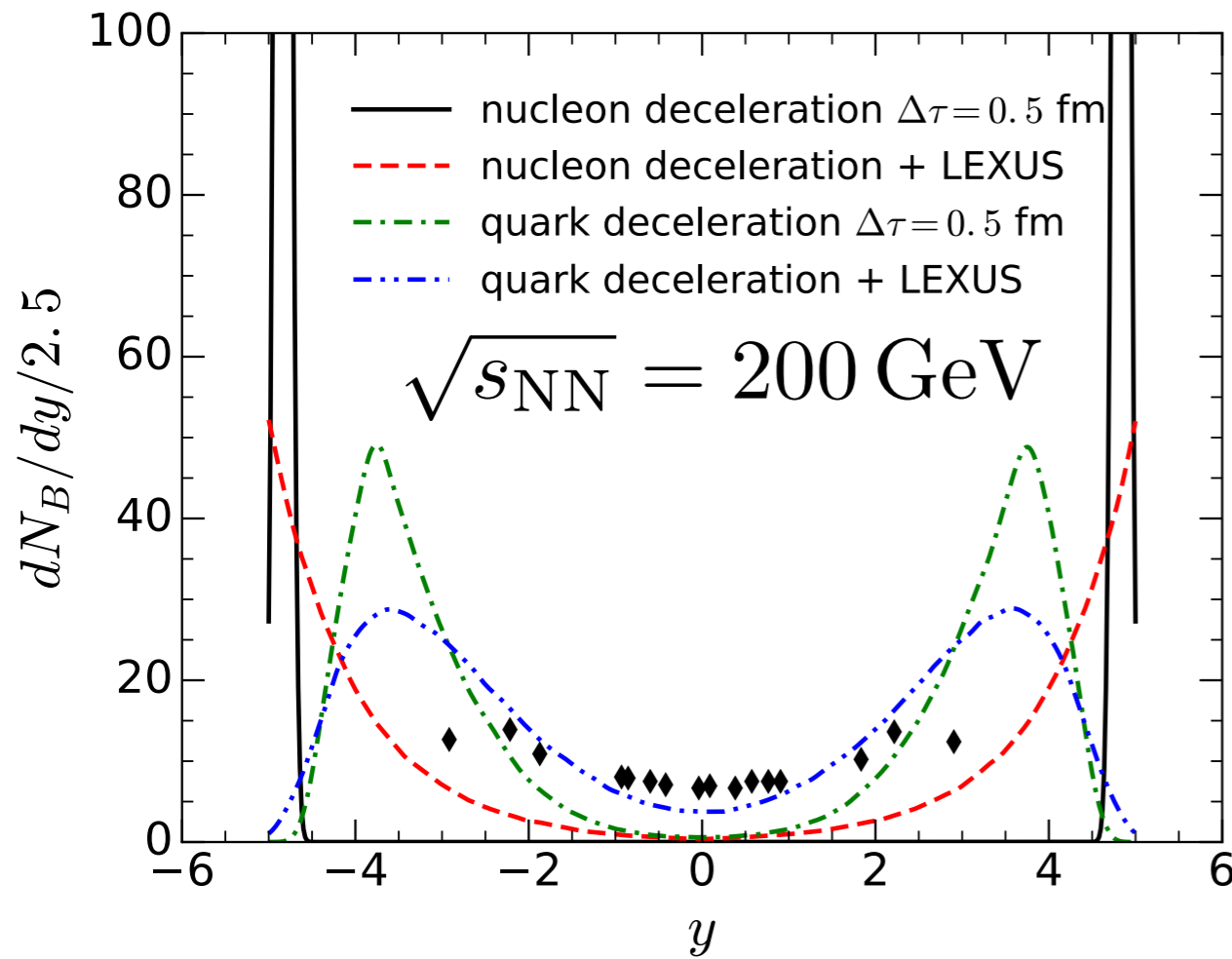
C. Shen, B. Schenke, in preparation



- Different rapidity fluctuation results different net baryon rapidity distribution

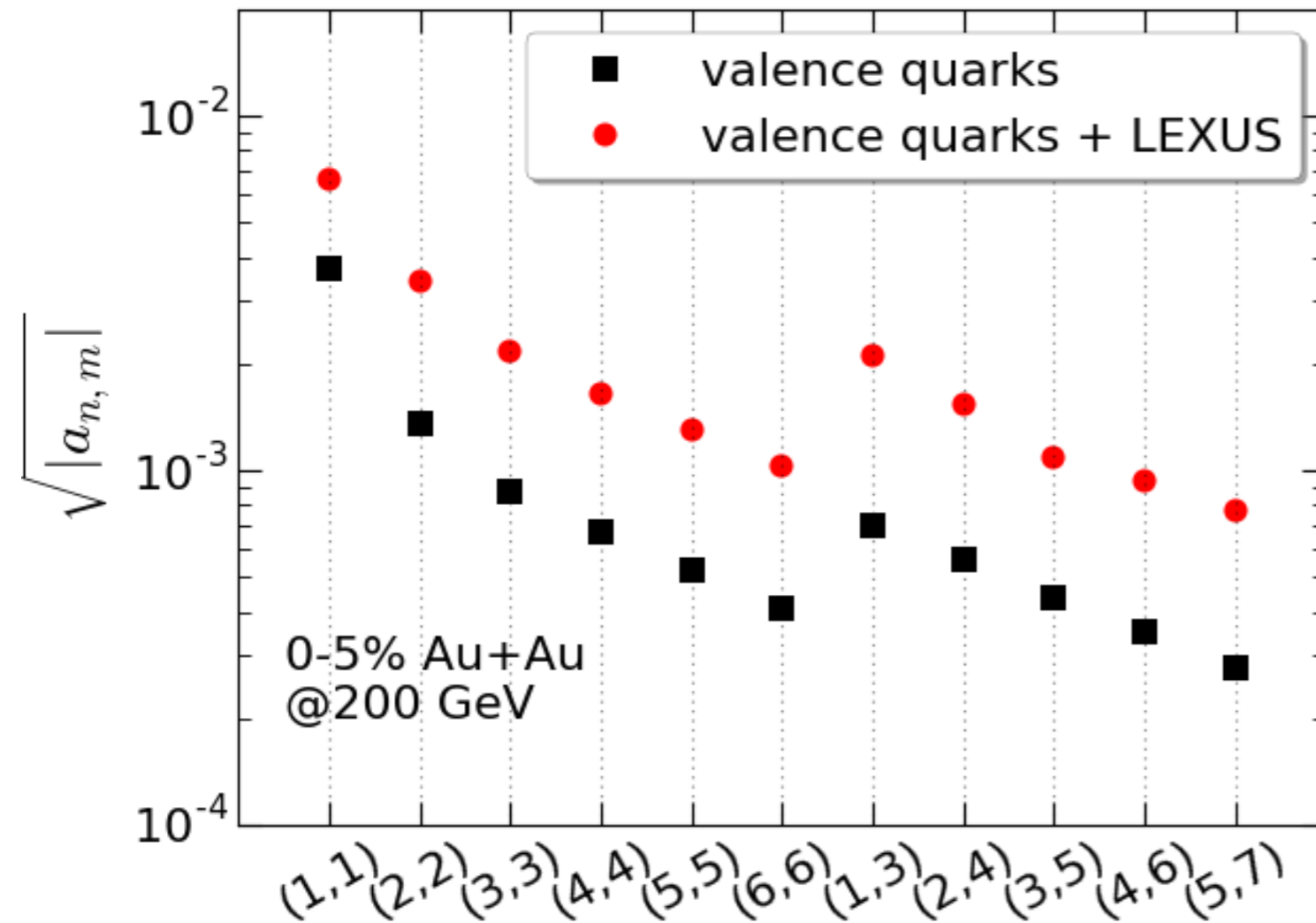
Net baryon rapidity distribution

C. Shen, B. Schenke, in preparation



- Different rapidity fluctuation results different net baryon rapidity distribution
- The valence quark + LEXUS model provides a reasonable net baryon rapidity distribution compared to the RHIC BES data

Quantify rapidity fluctuation



$$C(\eta_1, \eta_2) = \frac{\langle \frac{dN}{d\eta}(\eta_1) \frac{dN}{d\eta}(\eta_2) \rangle}{\langle \frac{dN}{d\eta}(\eta_1) \rangle \langle \frac{dN}{d\eta}(\eta_2) \rangle},$$

$$C_N(\eta_1, \eta_2) = \frac{C(\eta_1, \eta_2)}{C_p(\eta_1)C_p(\eta_2)},$$

$$C_p(\eta) = \frac{1}{2Y} \int_{-Y}^Y C(\eta_1, \eta_2) d\eta_2$$

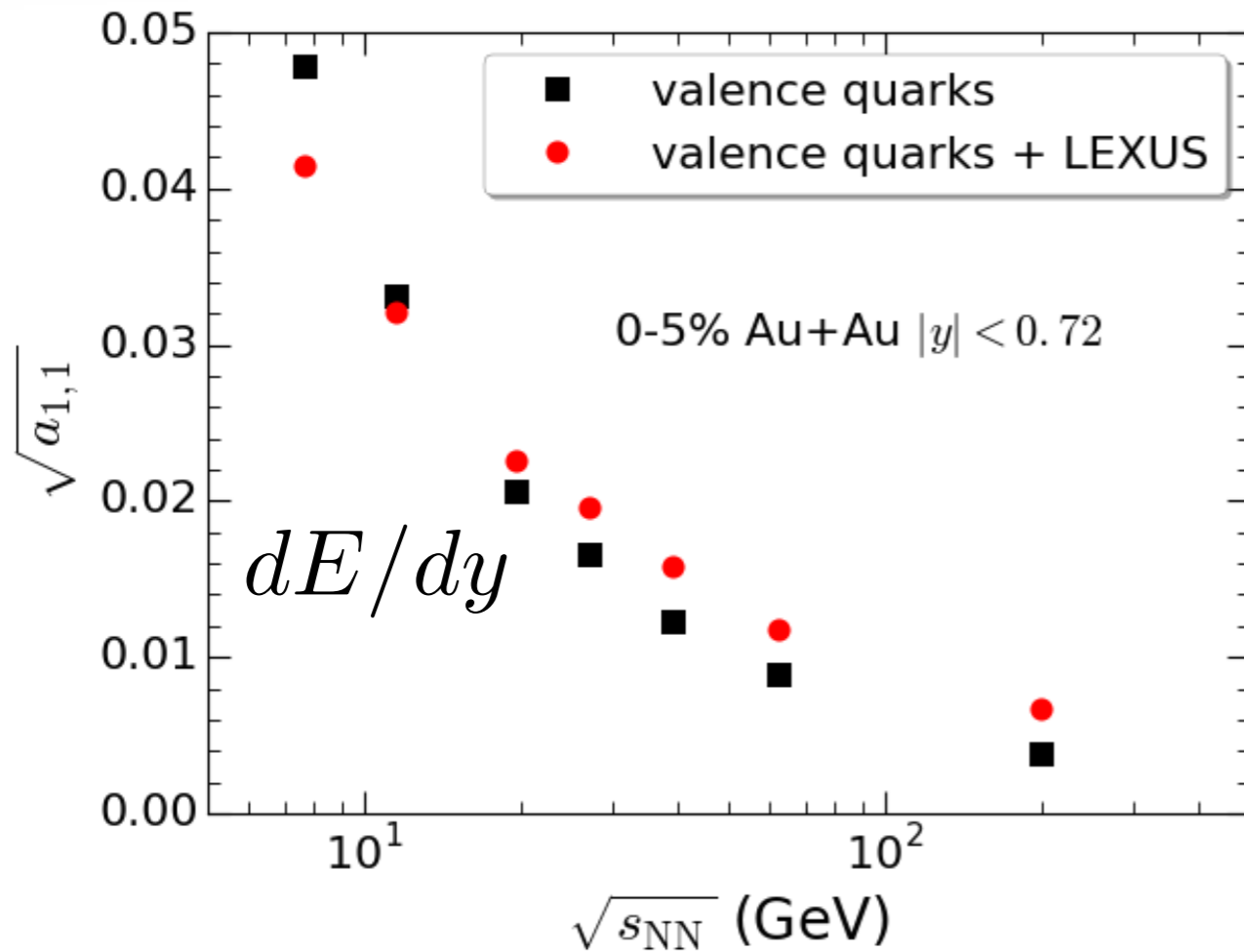
$$a_{n,m} = \int \frac{d\eta_1}{Y} \frac{d\eta_2}{Y} C_N(\eta_1, \eta_2) \frac{T_n(\eta_1)T_m(\eta_2) + T_n(\eta_2)T_m(\eta_1)}{2}$$

$$T_n(\eta) = \sqrt{n + \frac{1}{2}P_n(\frac{\eta}{Y})}$$

- The size of the $a_{n,m}$ coefficient can quantify the amount of longitudinal fluctuations

Quantify rapidity fluctuation

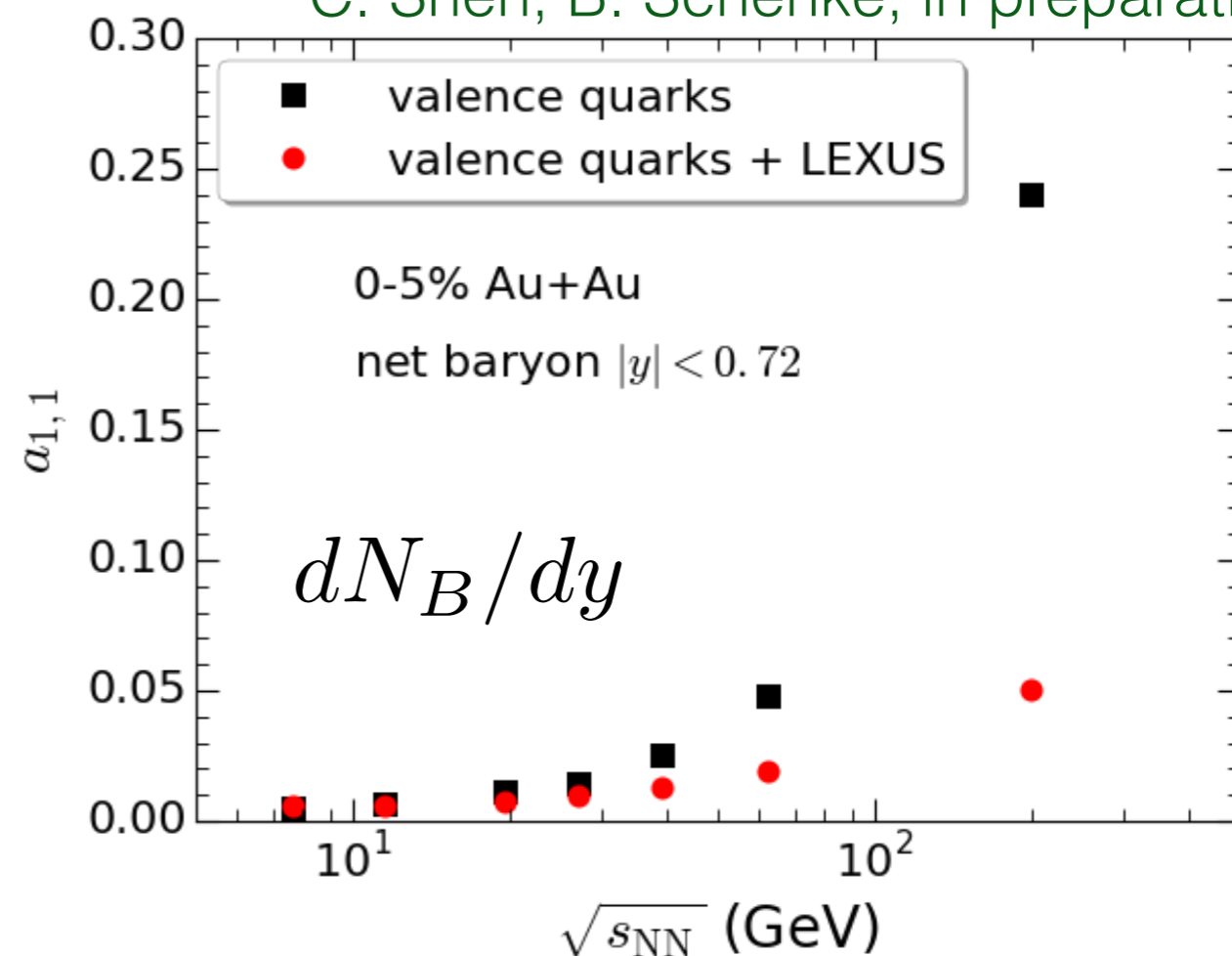
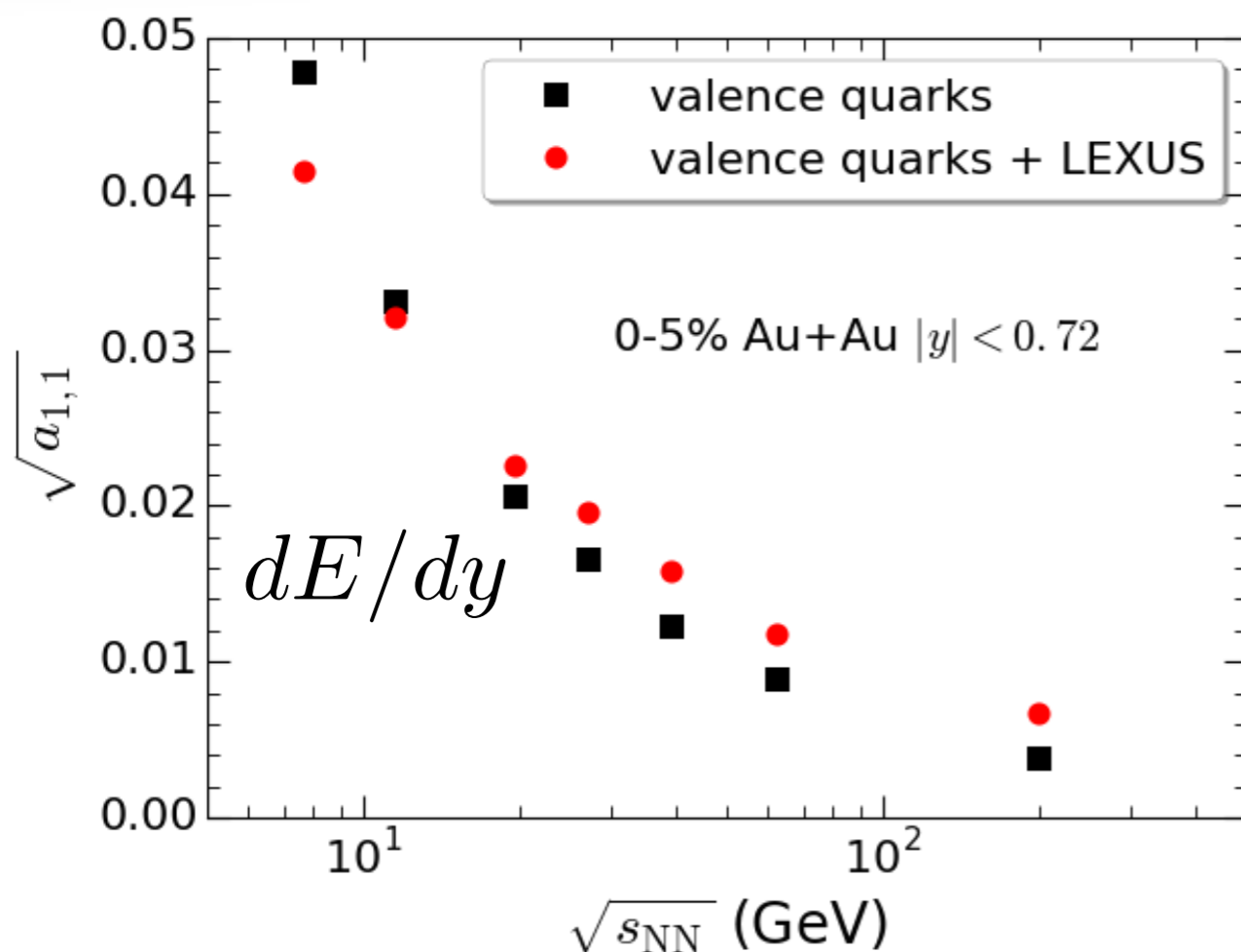
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- The a_{11} coefficient for dE/dy decreases at high collision energy because the system becomes more boost-invariant

Quantify rapidity fluctuation

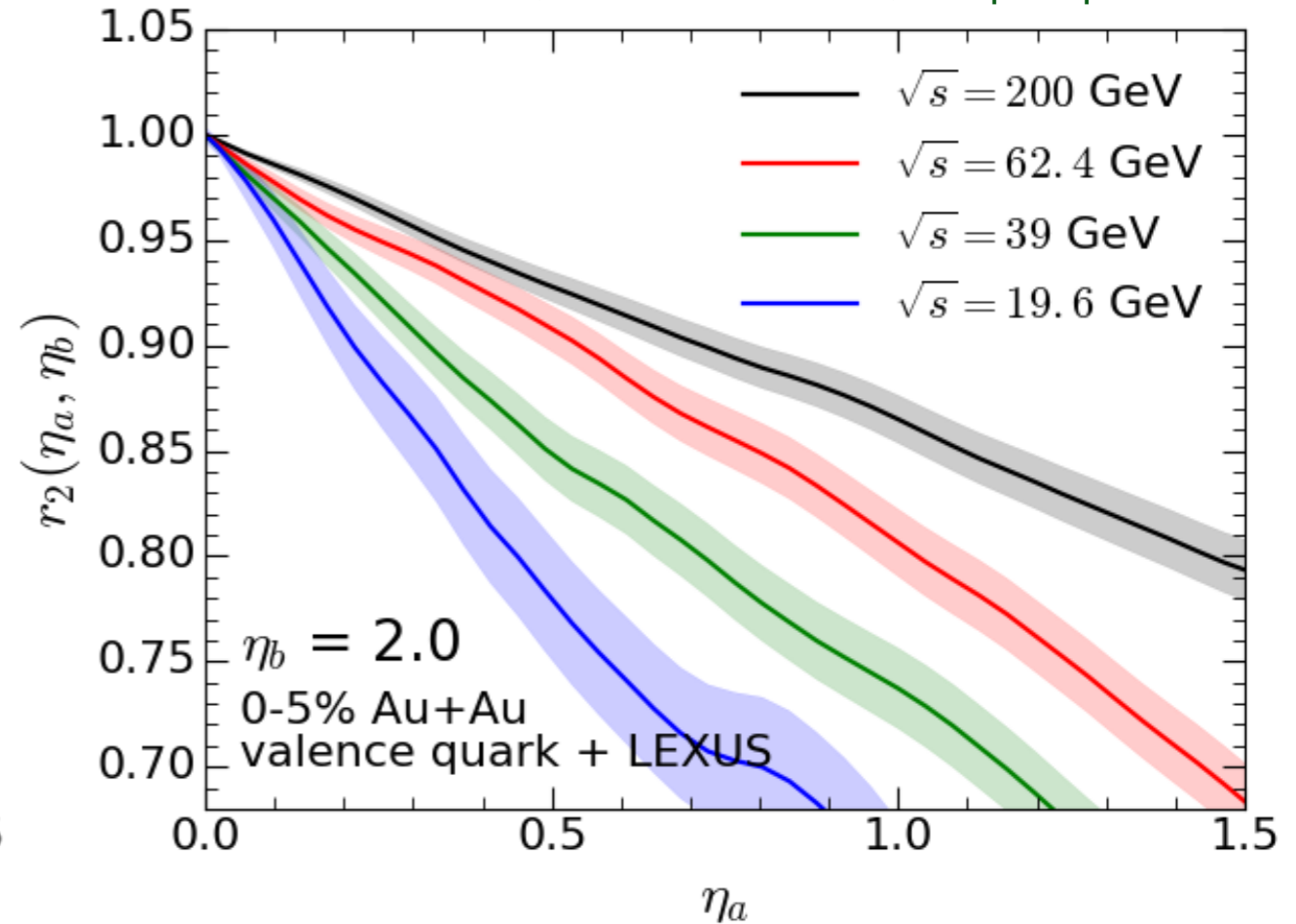
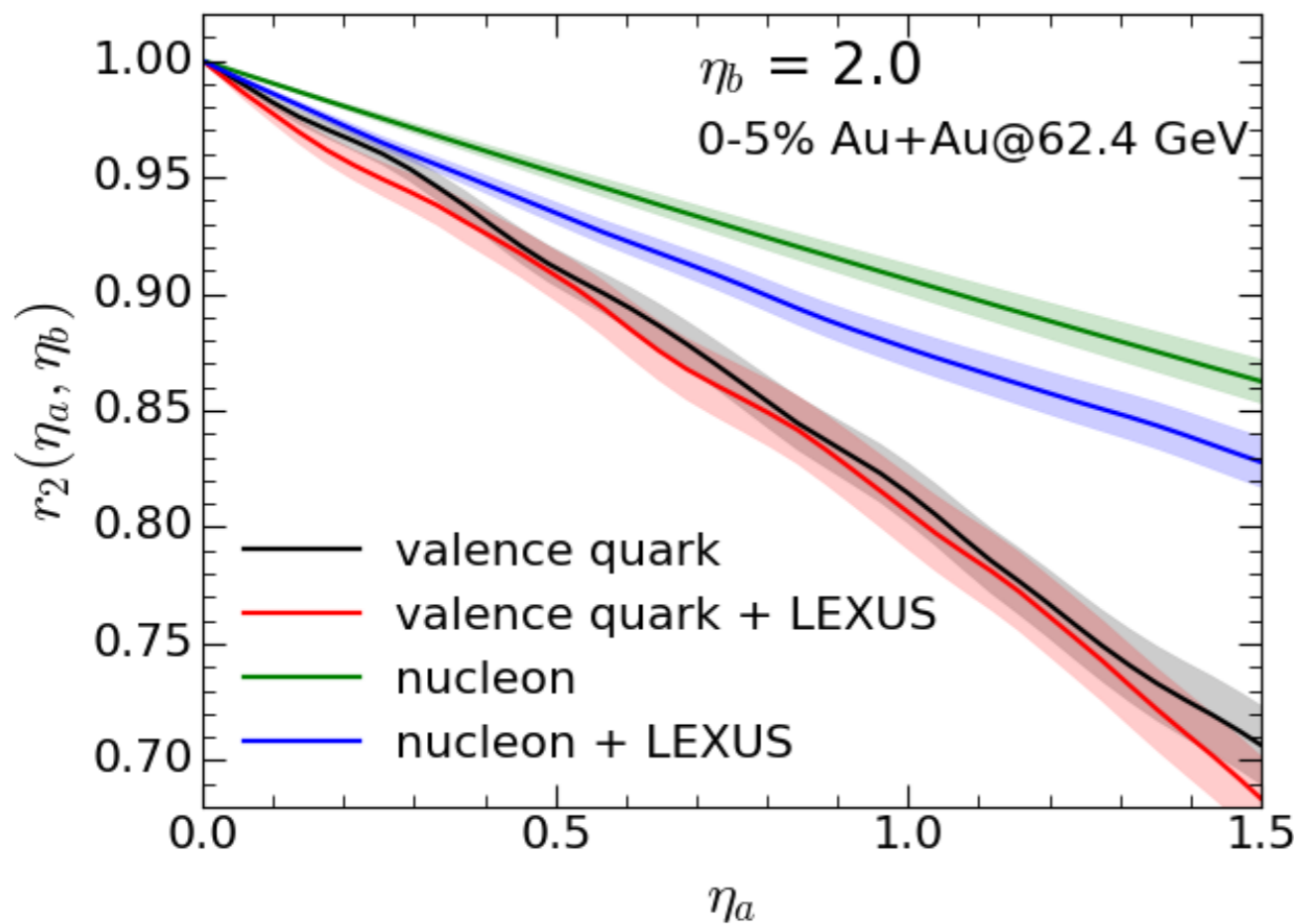
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- The a_{11} coefficient for dE/dy decreases at high collision energy because the system becomes more boost-invariant
- The a_{11} coefficient for dN_B/dy increases at high collision energy because less net baryon number at mid-rapidity

Quantify rapidity fluctuation

C. Shen, B. Schenke, in preparation



$$r_n(\eta_a, \eta_b) = \frac{\langle \Re \{ \mathcal{E}_n(-\eta_a) \cdot \mathcal{E}_n^*(\eta_b) \} \rangle_{ev}}{\langle \Re \{ \mathcal{E}_n(\eta_a) \cdot \mathcal{E}_n^*(\eta_b) \} \rangle_{ev}}$$

$$\mathcal{E}_n(\eta) = - \frac{\int r dr d\phi r^n e(r, \phi, \eta) e^{in\phi}}{\int r dr d\phi r^n e(r, \phi, \eta)}$$

- The initial eccentricities decorrelate along η direction faster with more longitudinal fluctuation and at lower collision energy

Hydrodynamics with sources

Energy-momentum current and net baryon density are feed into hydrodynamic simulation as source terms

$$\partial_\mu T^{\mu\nu} = J_{\text{source}}^\nu$$

$$\partial_\mu J^\mu = \rho_{\text{source}}$$

where

$$J_{\text{source}}^\nu = \delta e u^\nu + (e + P) \delta u^\nu$$

$$\delta u^\nu = \frac{\Delta_{\mu}^{\nu} J_{\text{source}}^{\mu}}{e + P}$$

δe heats up the system

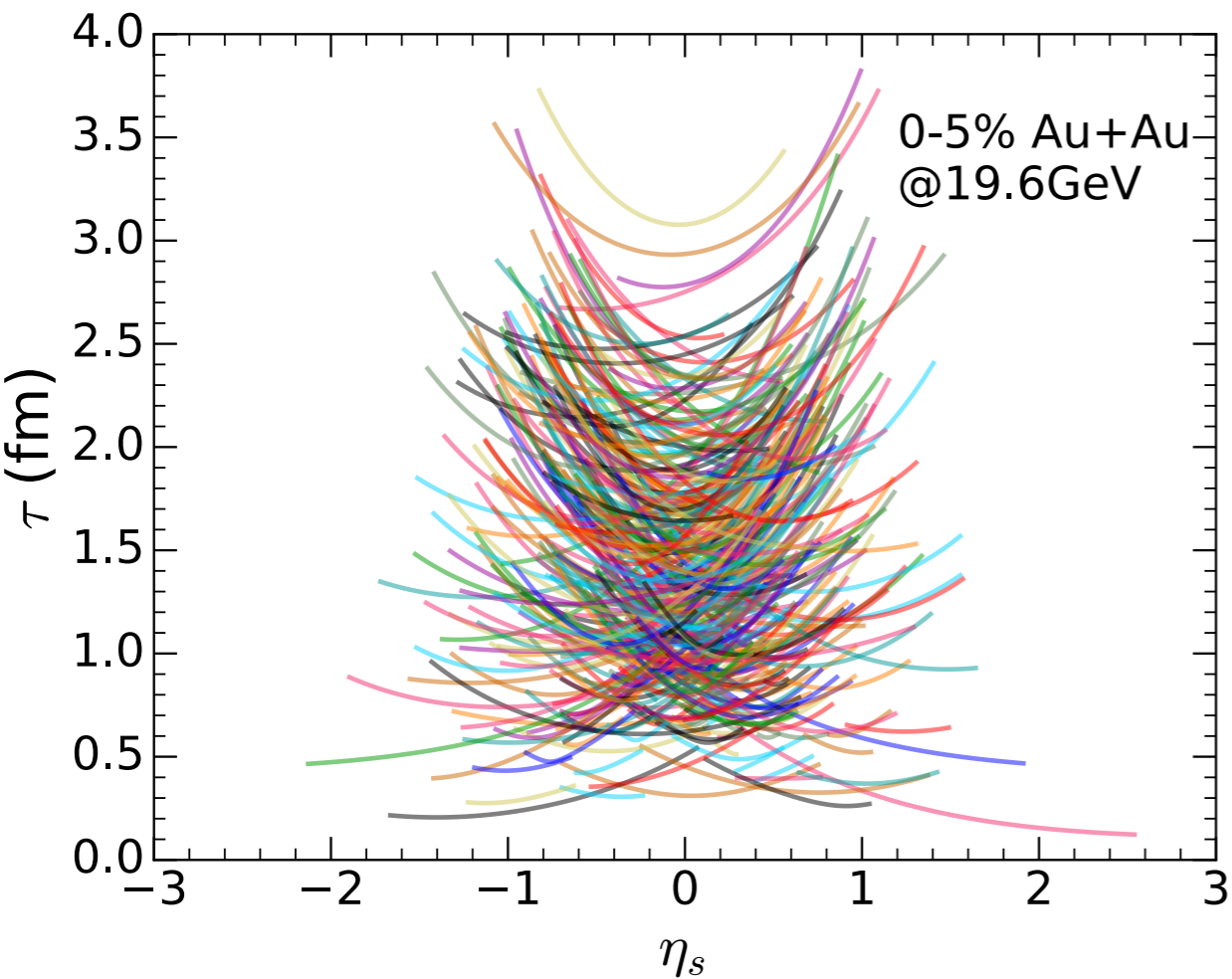
δu^ν accelerates the flow velocity

ρ_{source} dopes baryon charges into the system

- Source terms are smeared with Gaussians in space and time

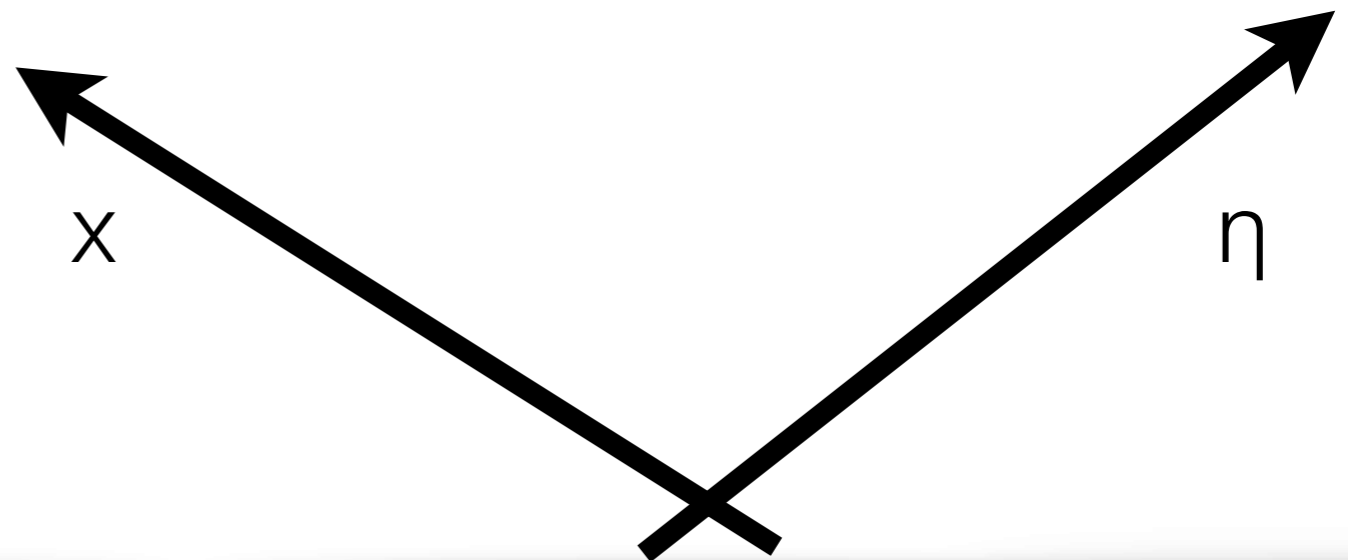
Hydrodynamical evolution with sources

energy density



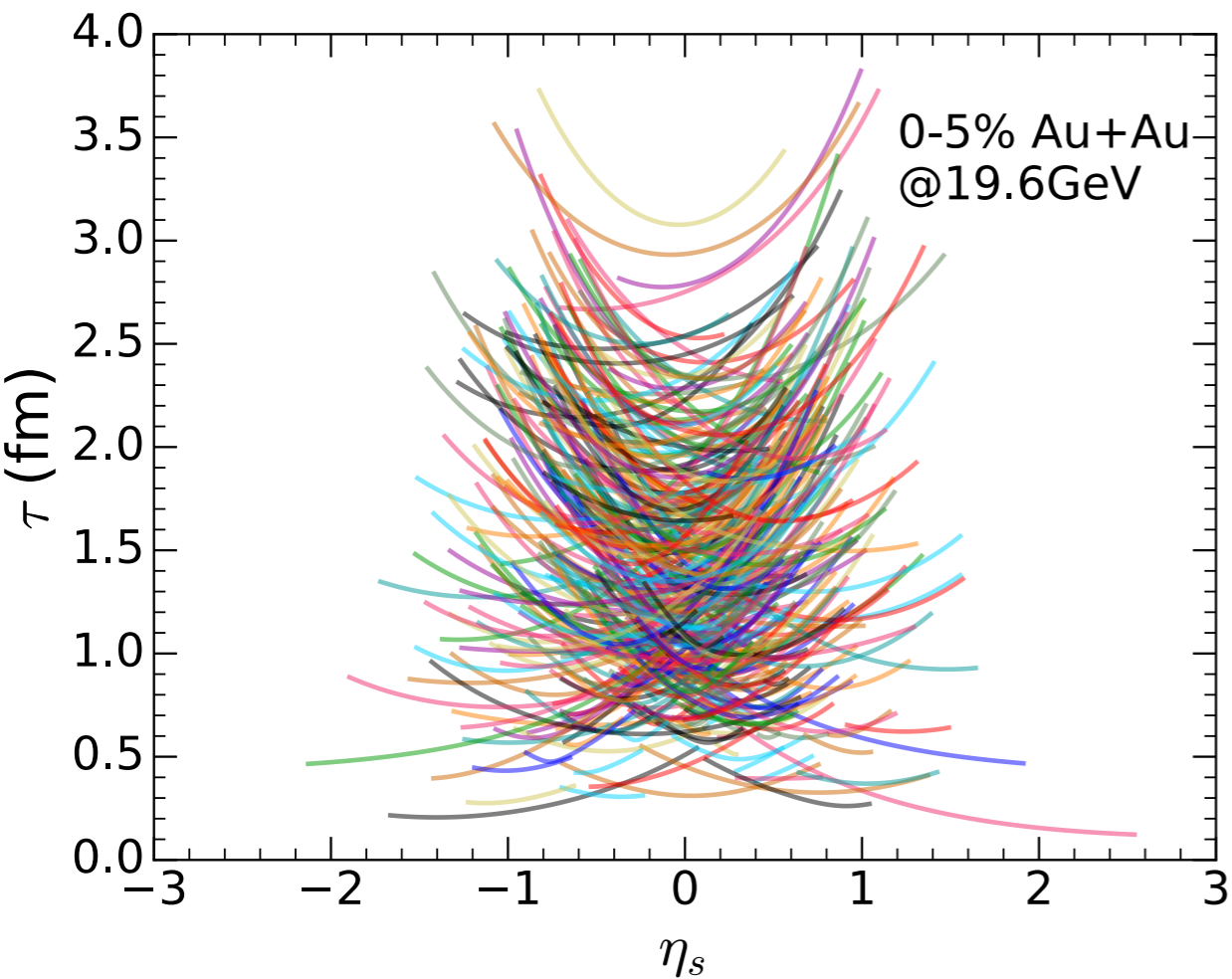
$\tau = 0.11$ fm

$\sqrt{s_{NN}} = 19.6$ GeV
valence quark + LEXUS



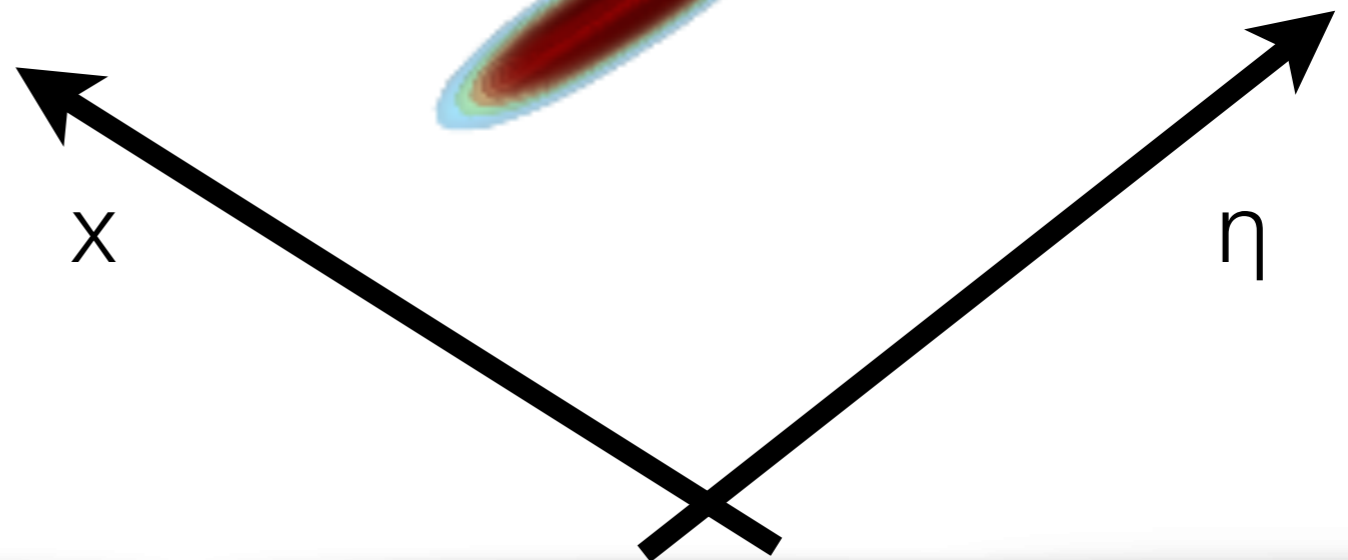
Hydrodynamical evolution with sources

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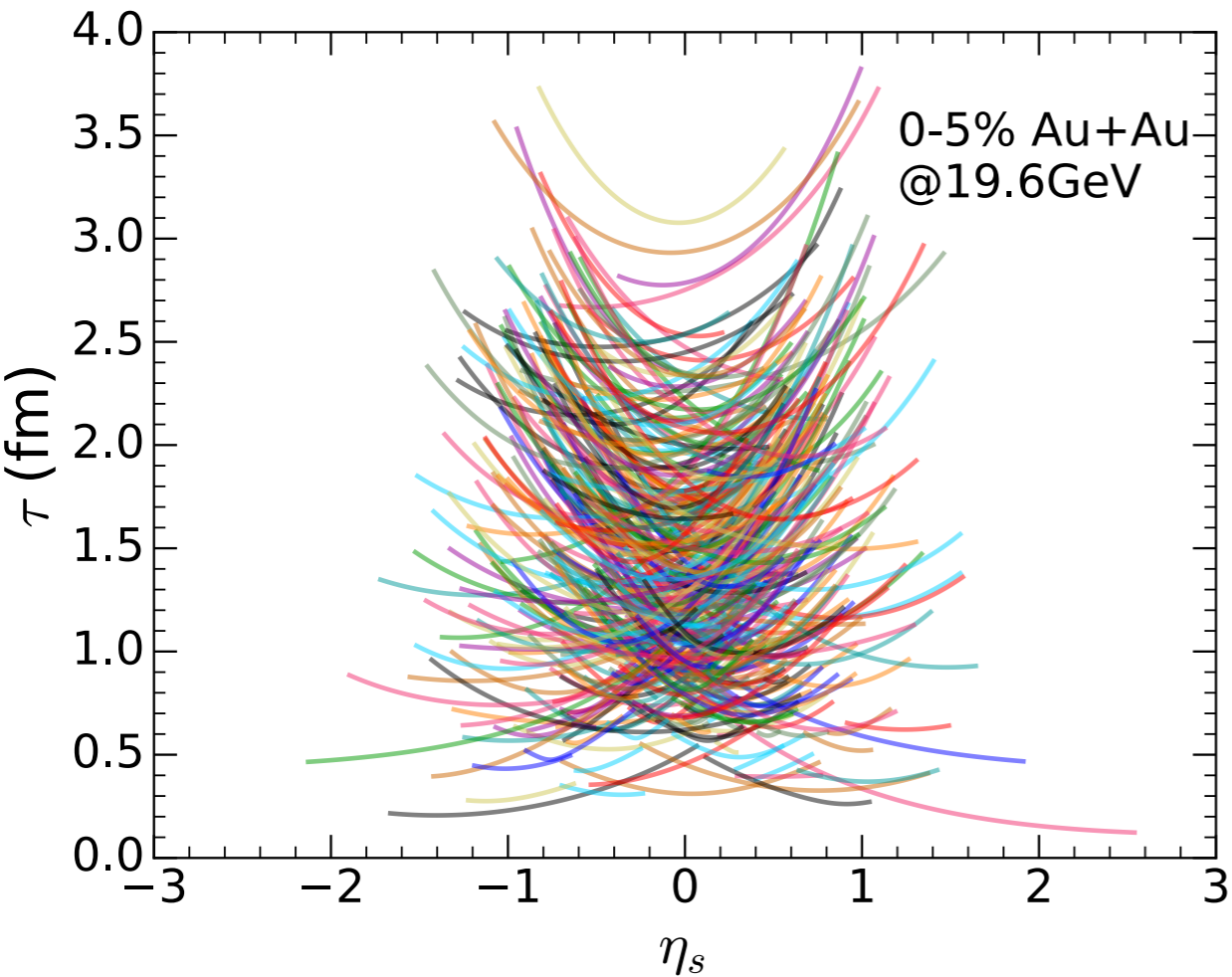
$\tau = 0.21 \text{ fm}$

$\sqrt{s_{NN}} = 19.6 \text{ GeV}$
valence quark + LEXUS

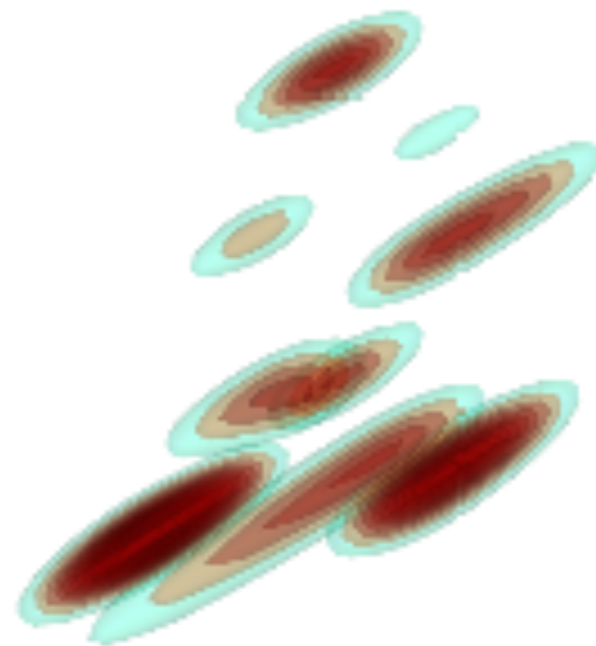


Hydrodynamical evolution with sources

energy density

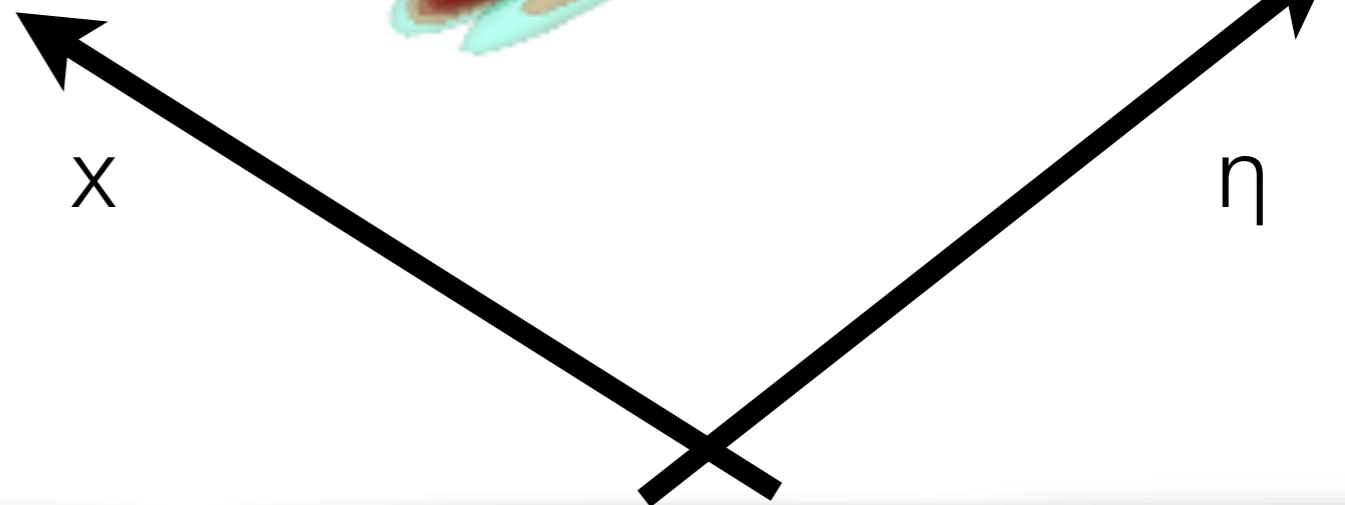


$\tau = 0.41 \text{ fm}$



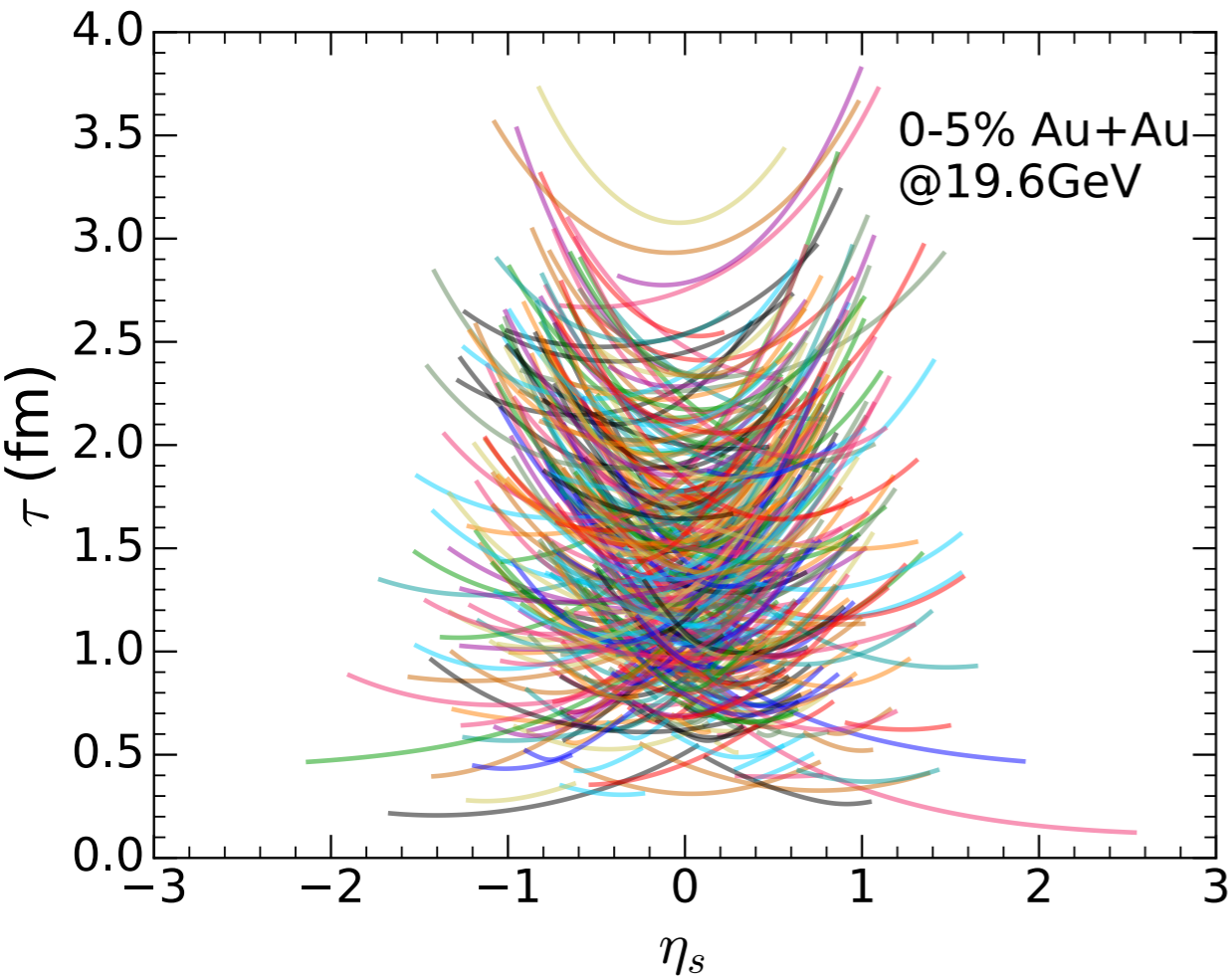
$$\sqrt{s_{NN}} = 19.6 \text{ GeV}$$

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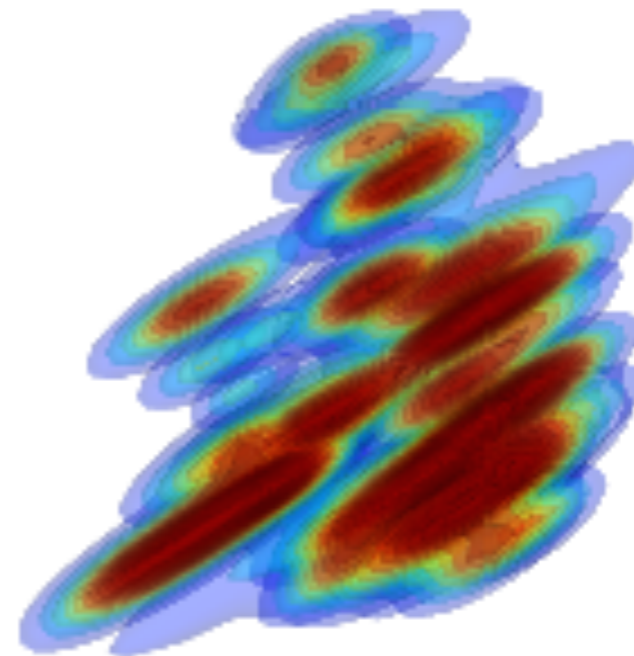


Hydrodynamical evolution with sources

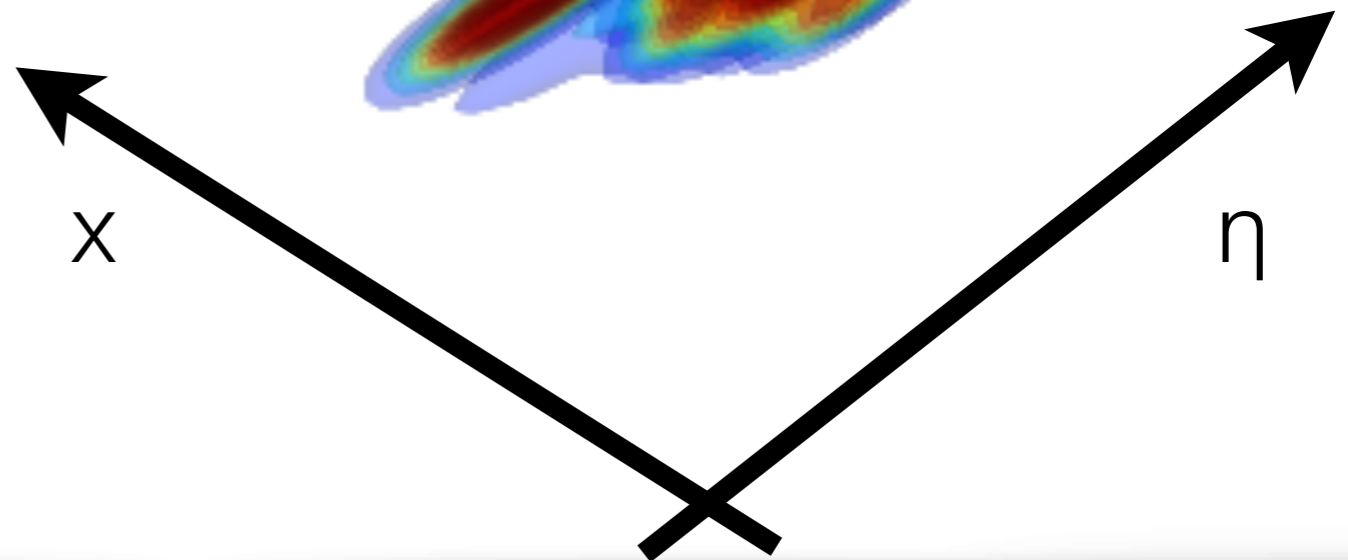
energy density



$\tau = 0.71$ fm

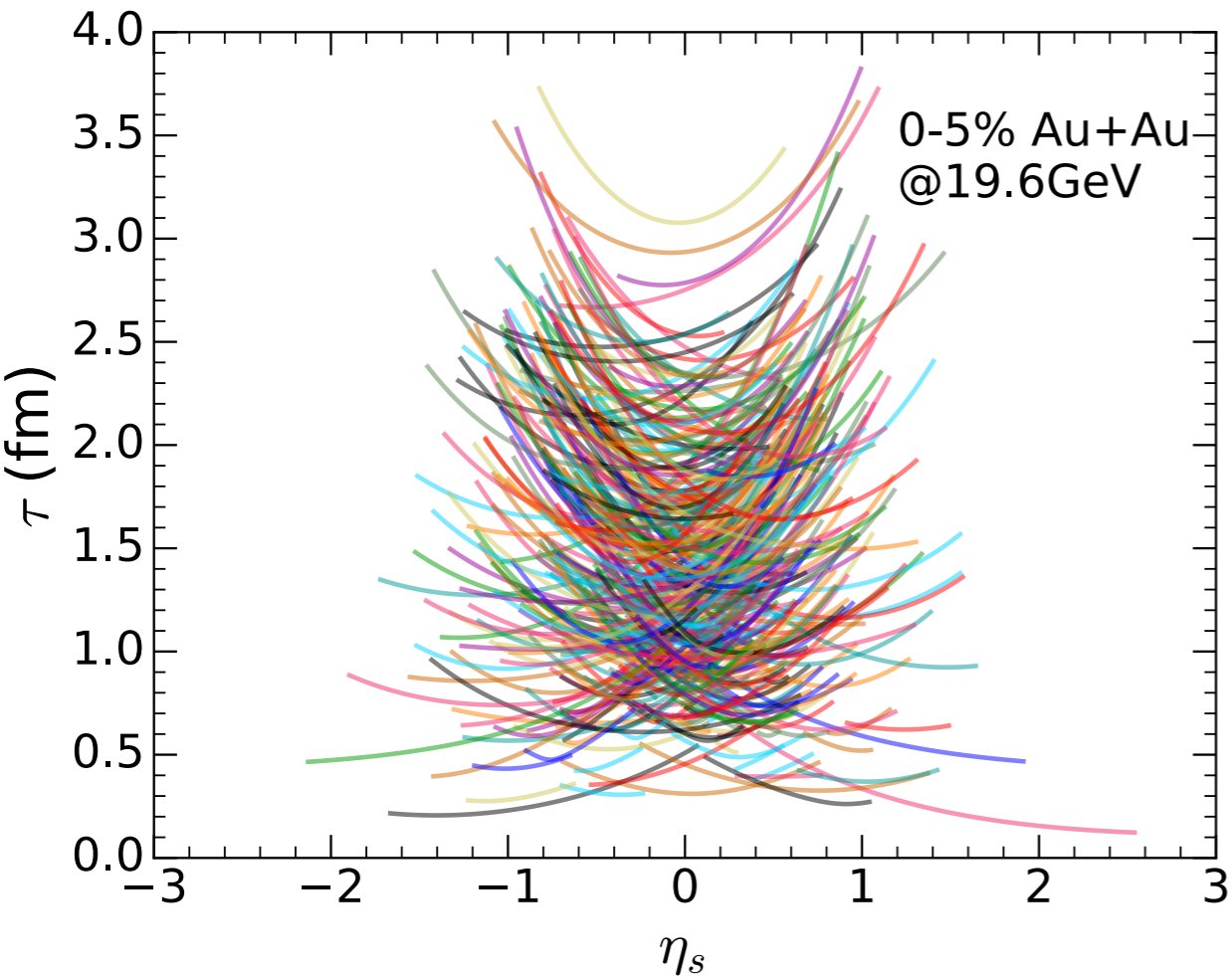


$\sqrt{s_{NN}} = 19.6$ GeV
valence quark + LEXUS

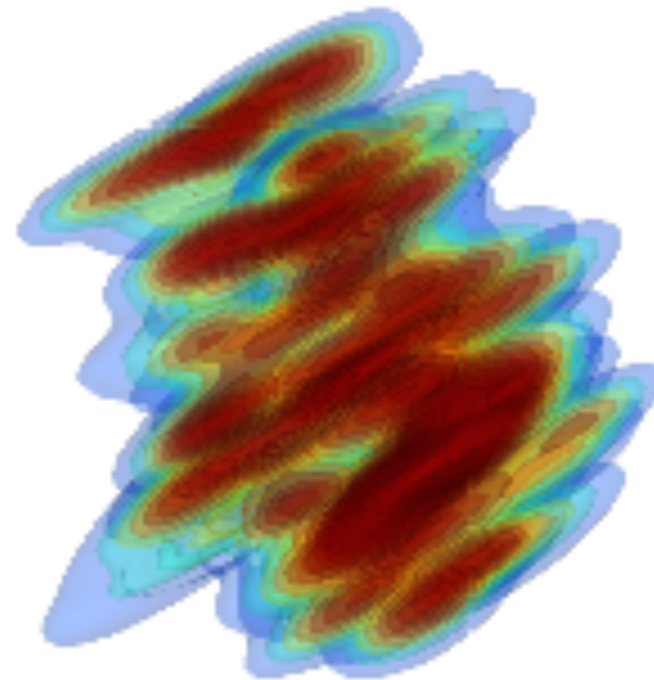


Hydrodynamical evolution with sources

energy density

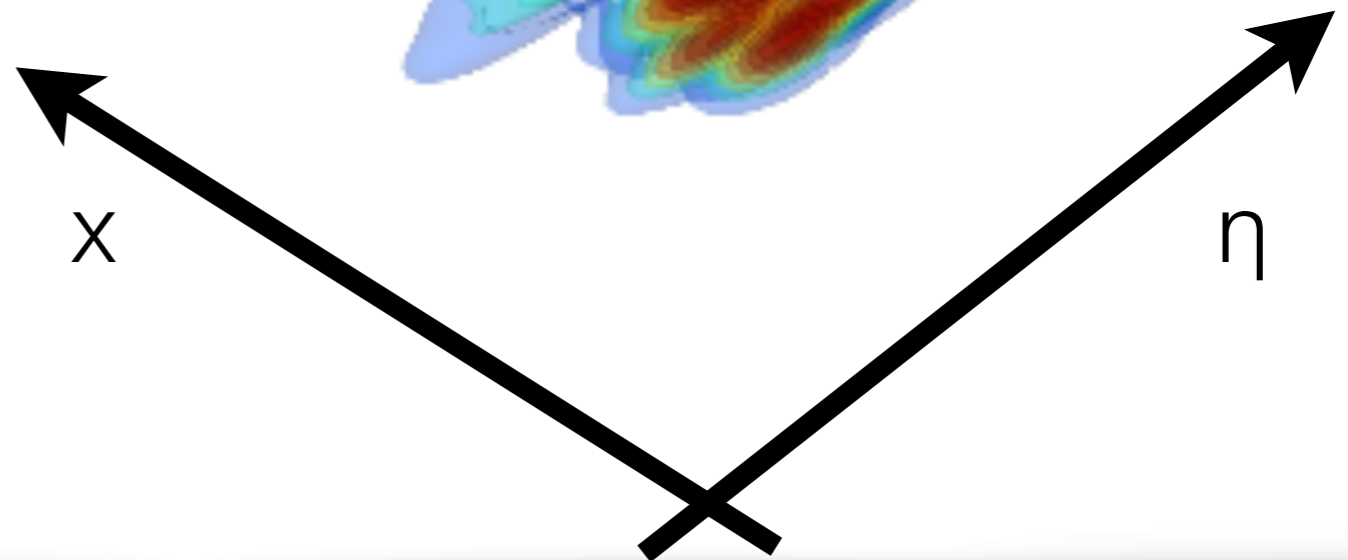


$\tau = 1.11 \text{ fm}$



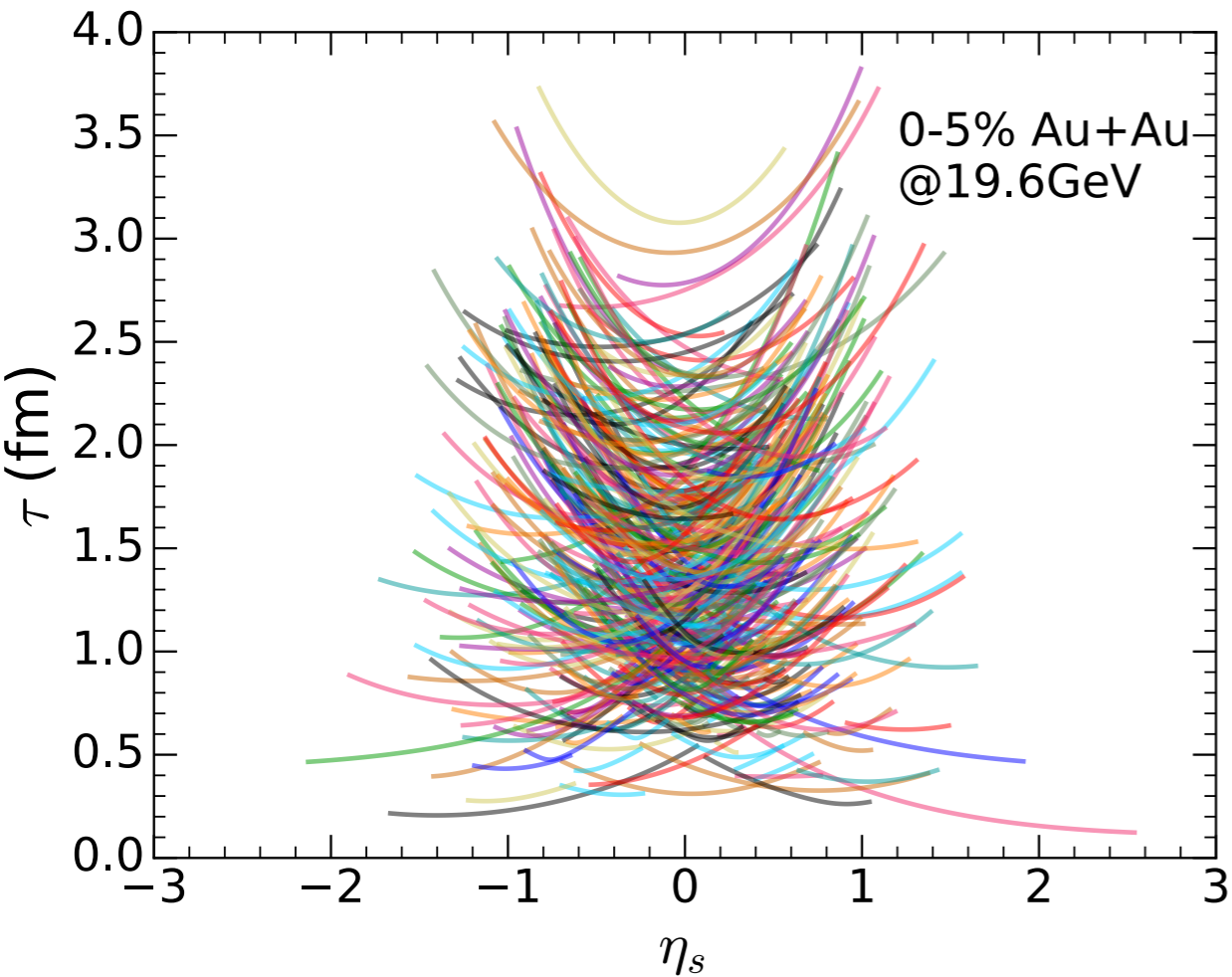
$$\sqrt{s_{NN}} = 19.6 \text{ GeV}$$

valence quark + LEXUS

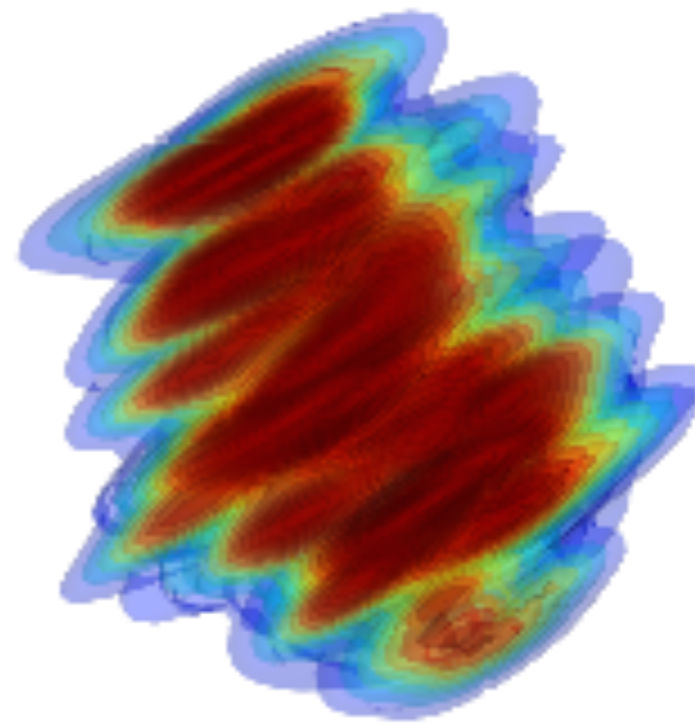


Hydrodynamical evolution with sources

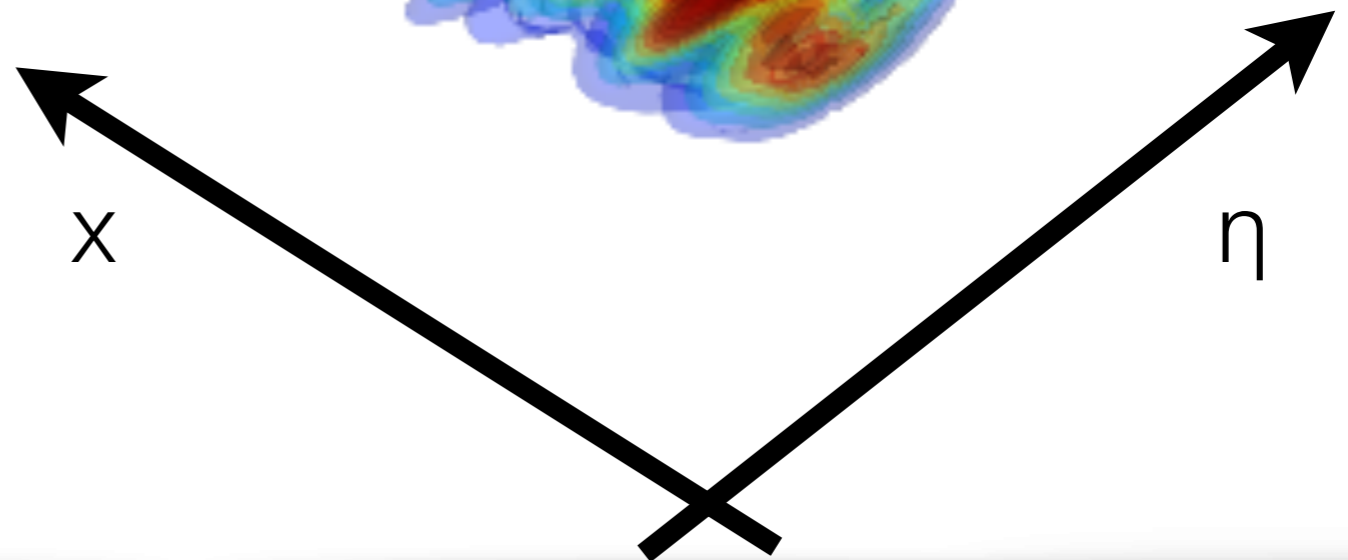
energy density



$\tau = 1.51 \text{ fm}$

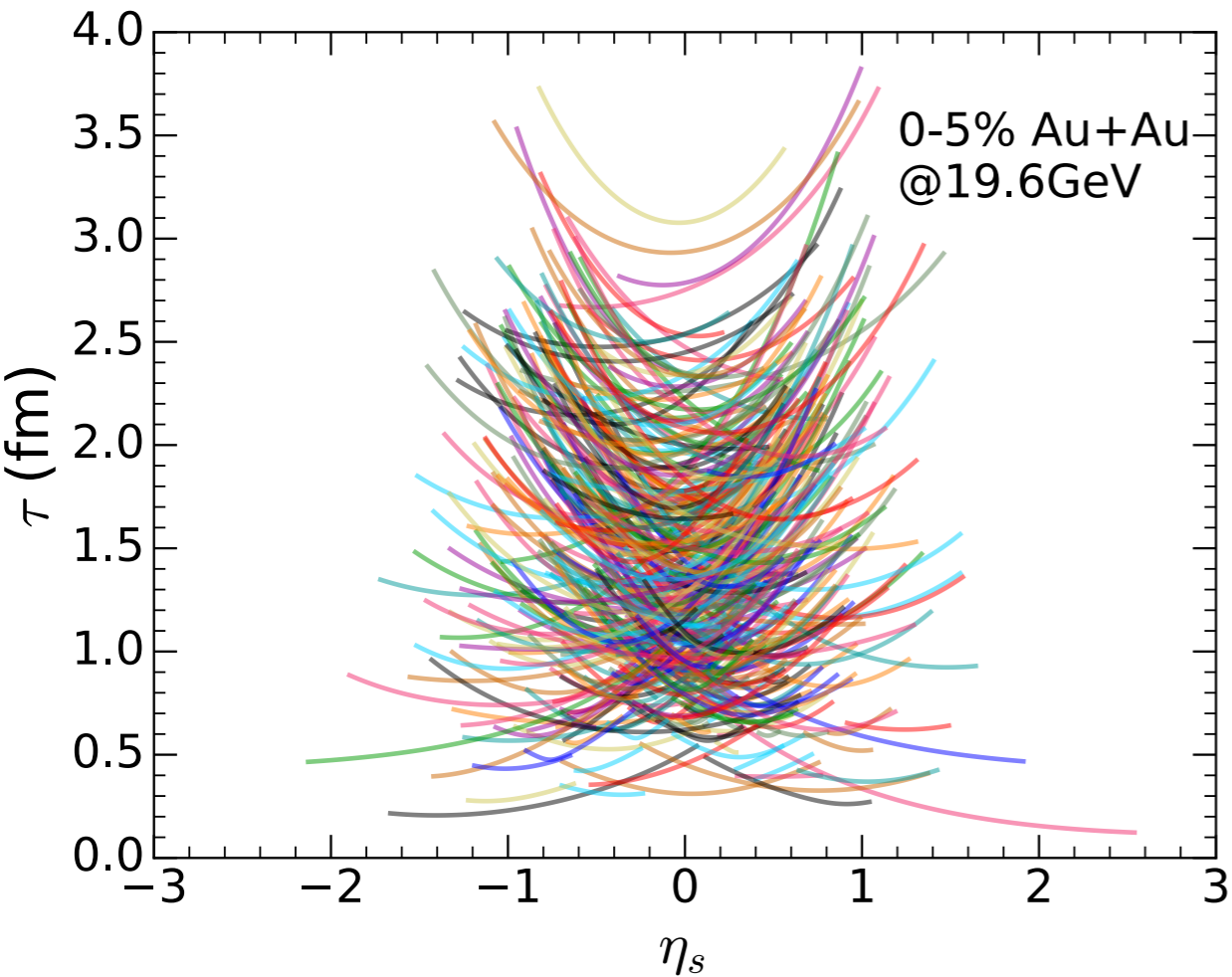


$\sqrt{s_{NN}} = 19.6 \text{ GeV}$
valence quark + LEXUS

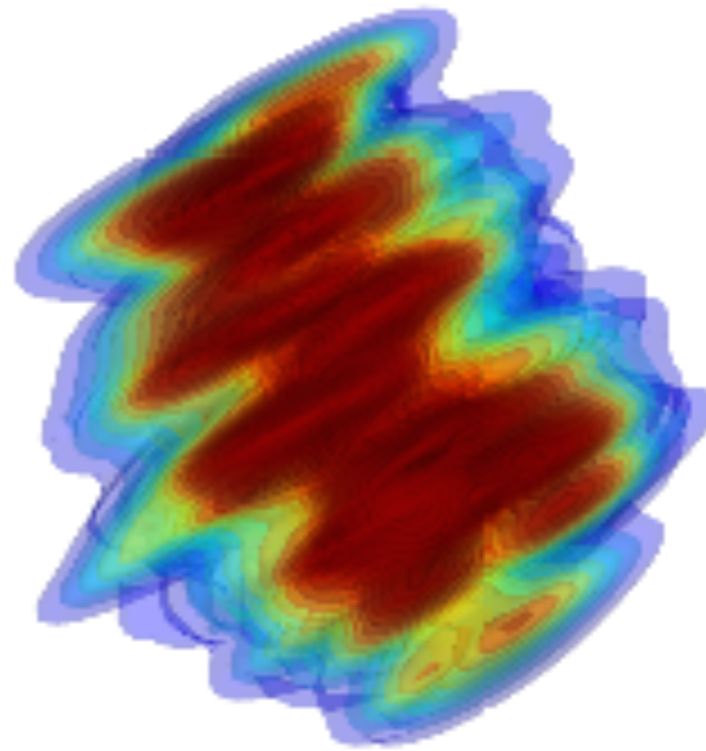


Hydrodynamical evolution with sources

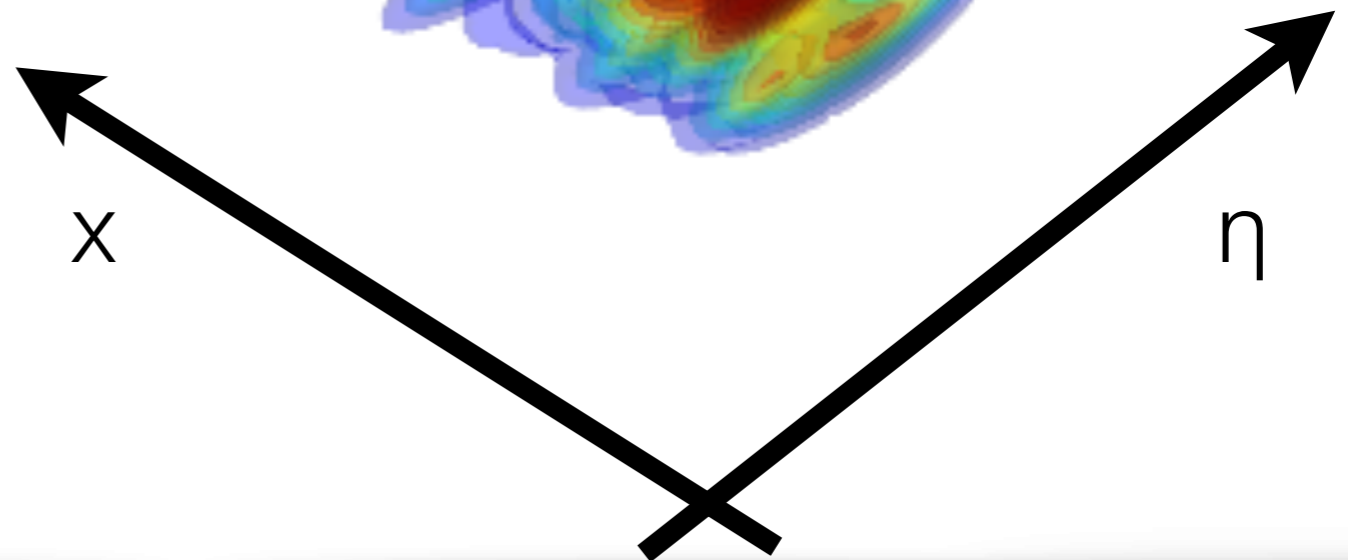
energy density



$\tau = 1.91 \text{ fm}$

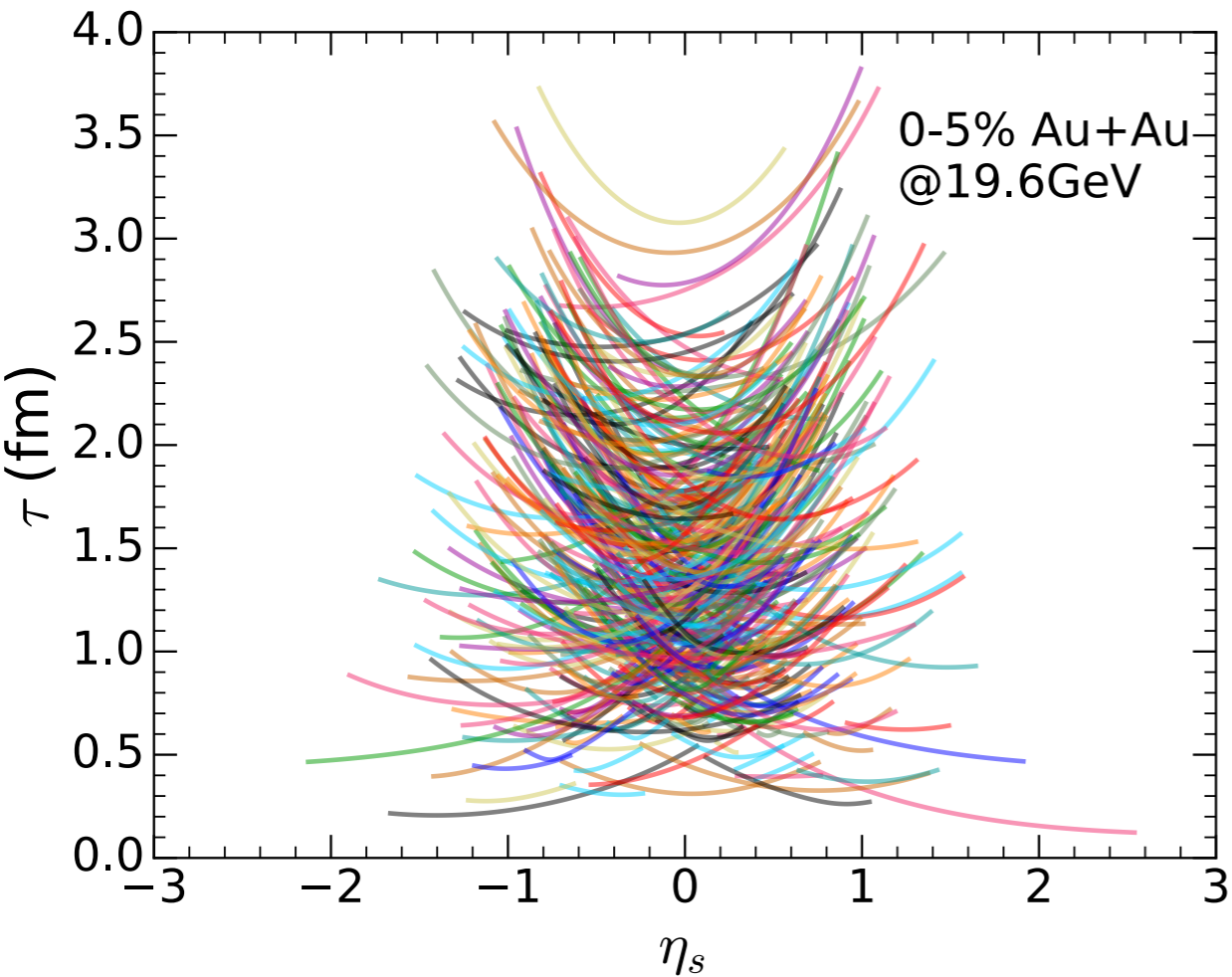


$\sqrt{s_{NN}} = 19.6 \text{ GeV}$
valence quark + LEXUS

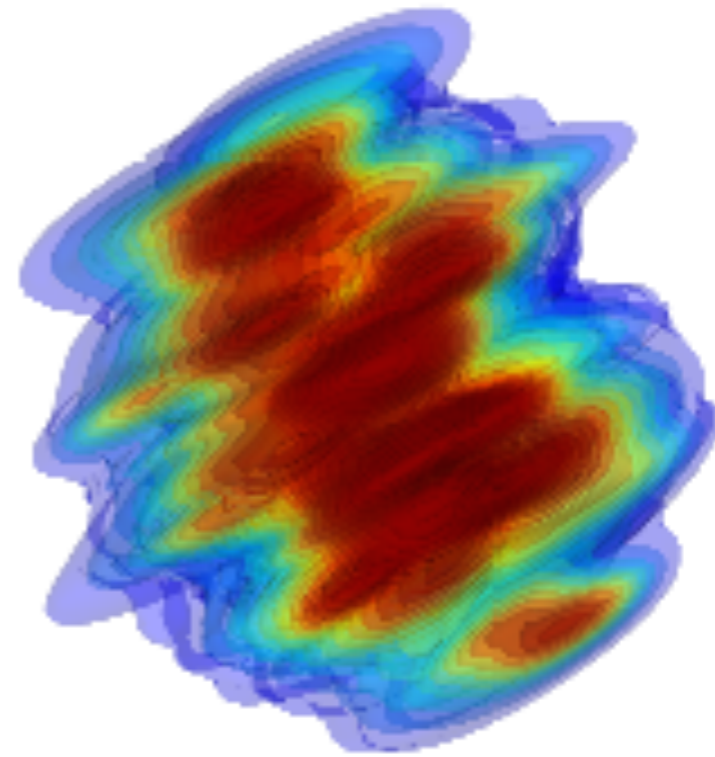


Hydrodynamical evolution with sources

energy density



$\tau = 2.31$ fm

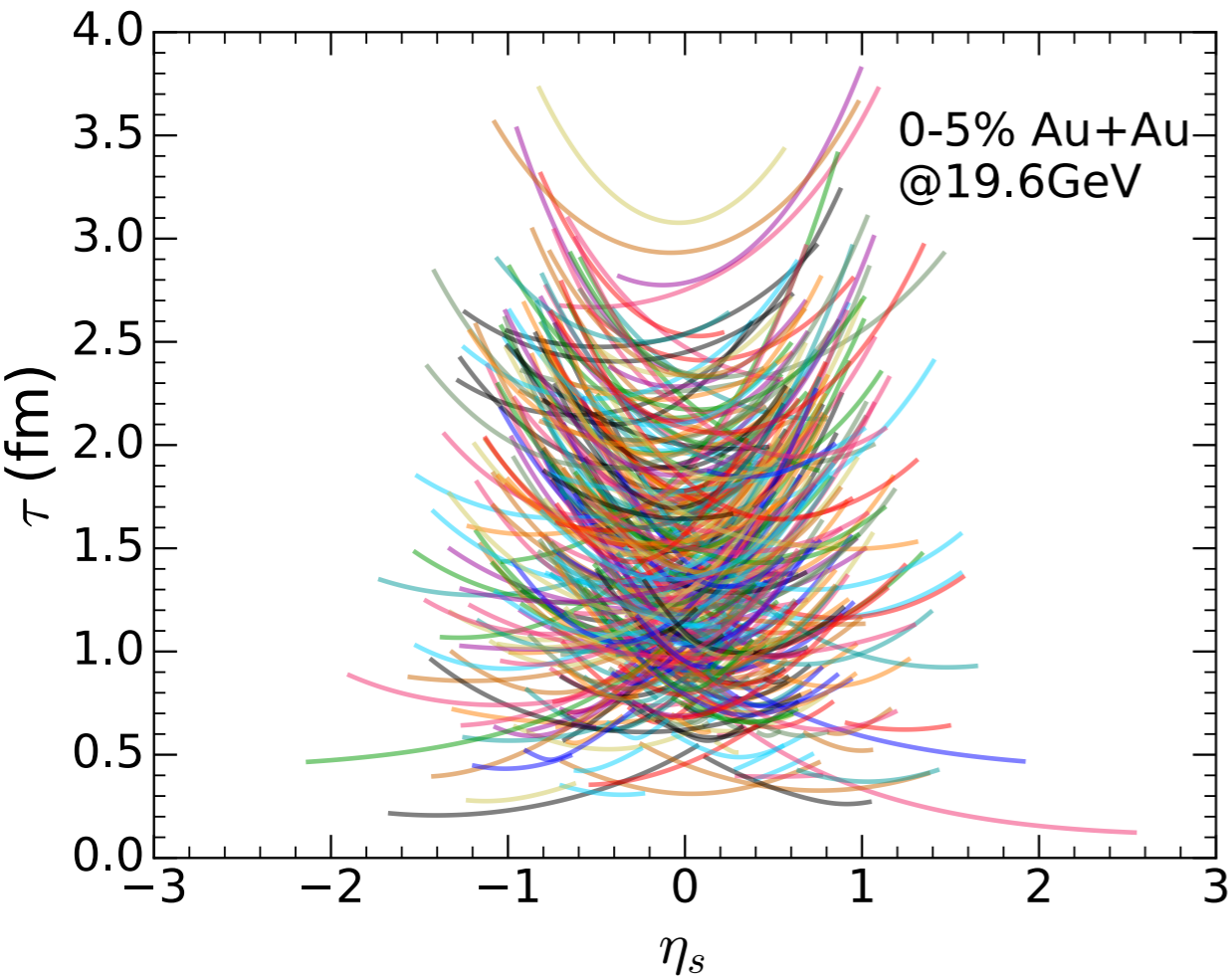


$$\sqrt{s_{NN}} = 19.6 \text{ GeV}$$

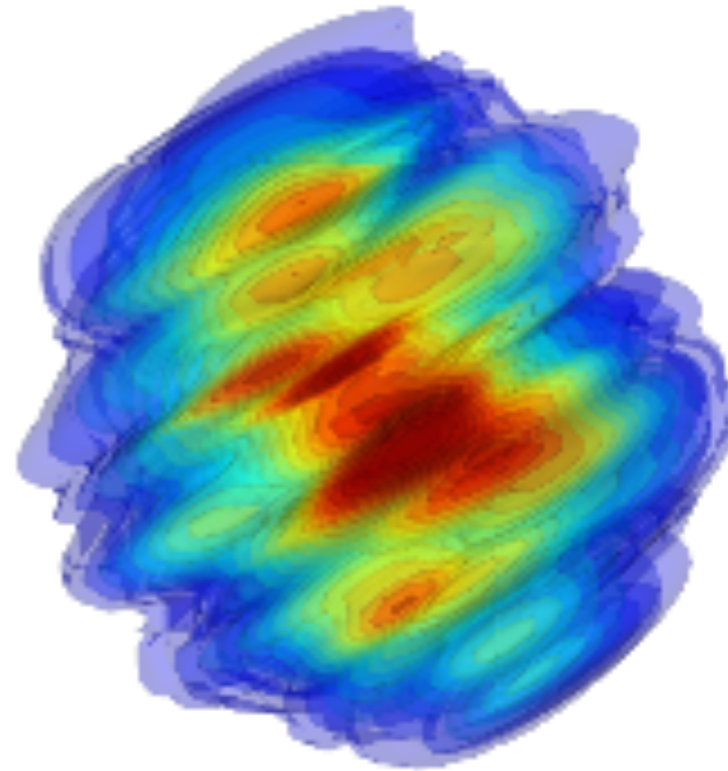
valence quark + LEXUS

Hydrodynamical evolution with sources

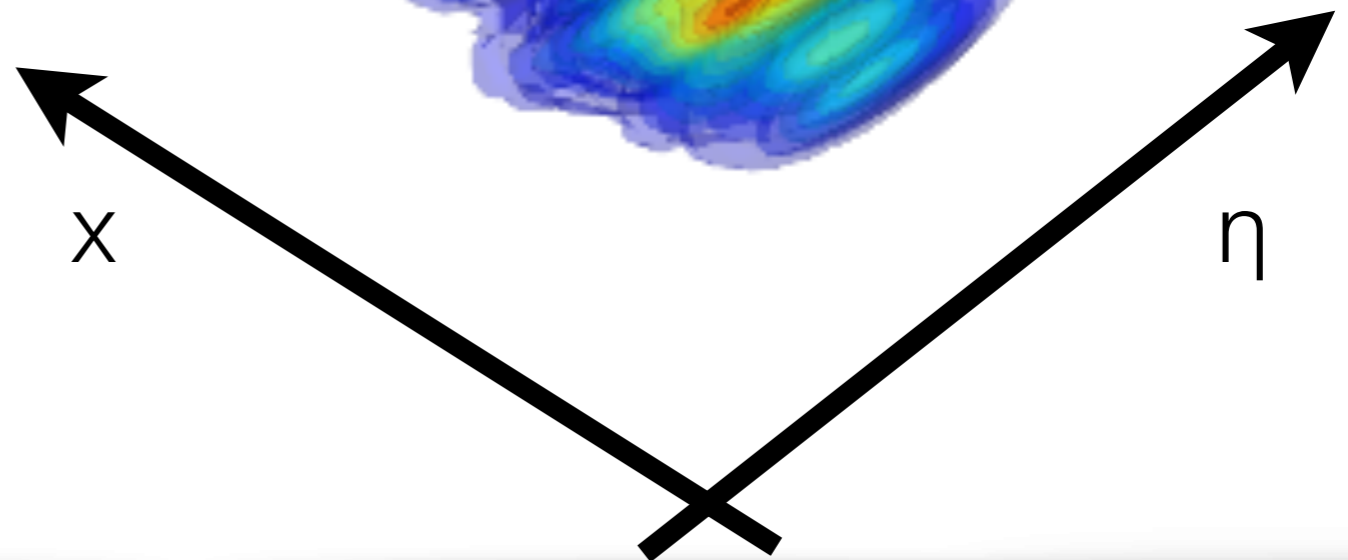
energy density



$\tau = 3.11 \text{ fm}$

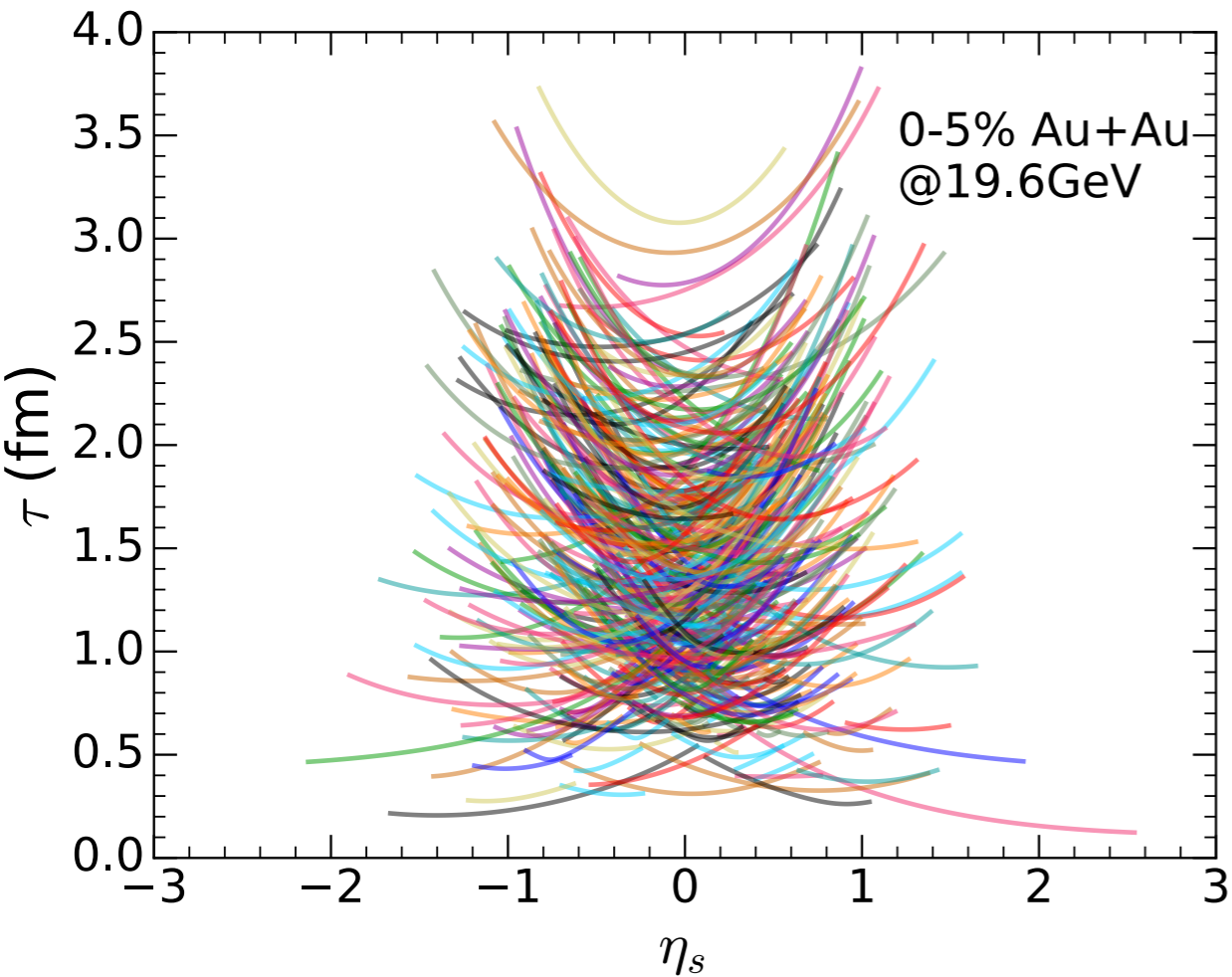


$\sqrt{s_{NN}} = 19.6 \text{ GeV}$
valence quark + LEXUS

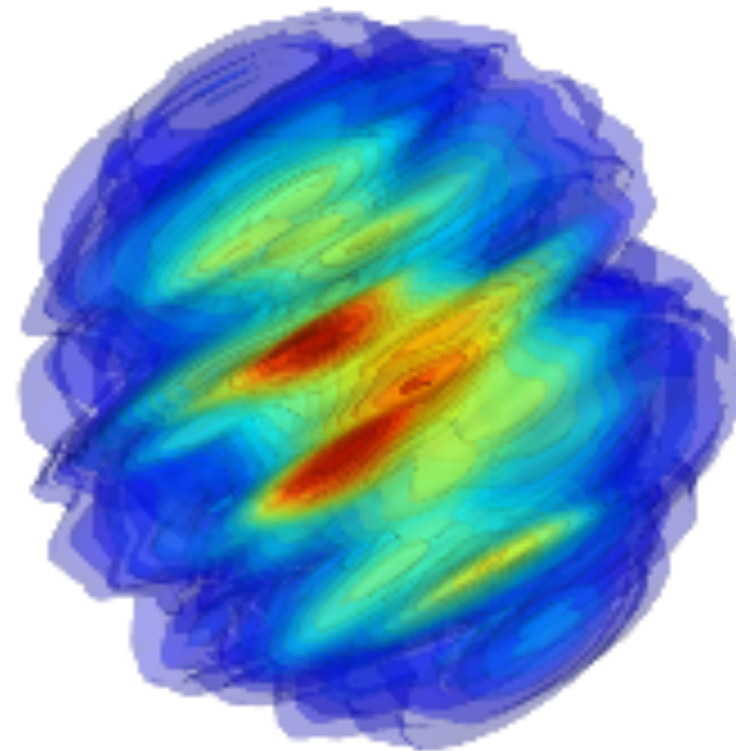


Hydrodynamical evolution with sources

energy density



$\tau = 3.71 \text{ fm}$

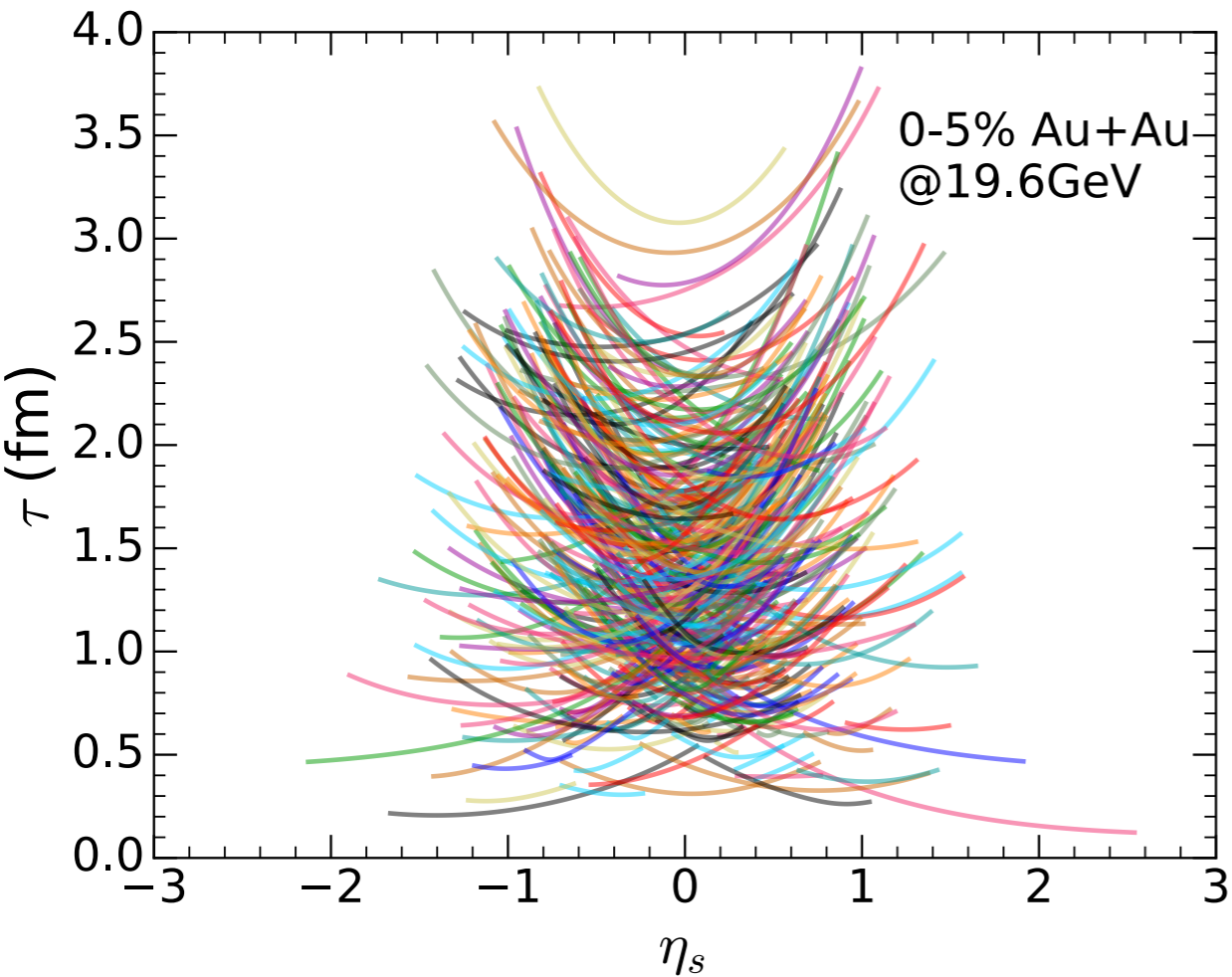


$\sqrt{s_{NN}} = 19.6 \text{ GeV}$
valence quark + LEXUS

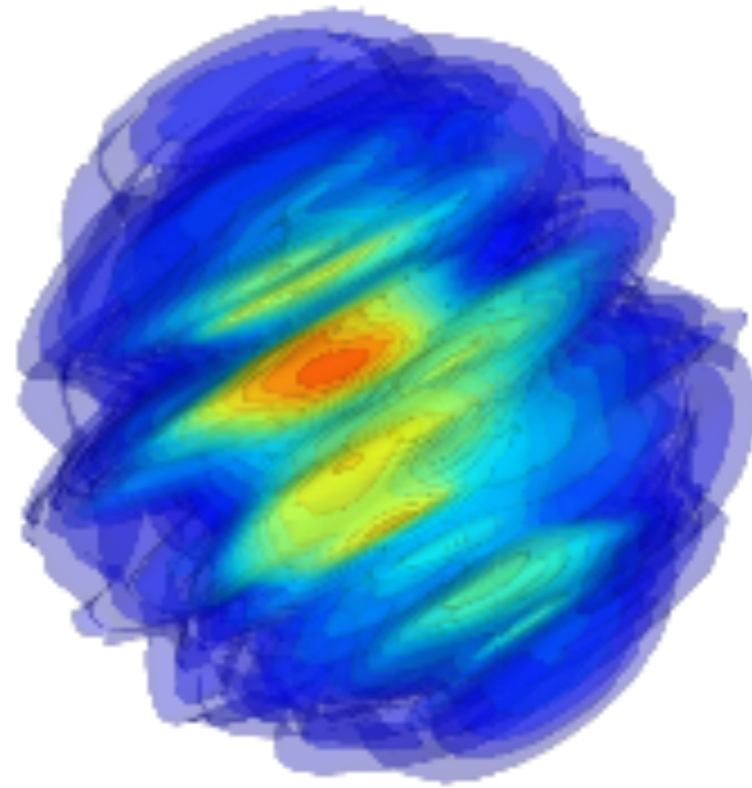


Hydrodynamical evolution with sources

energy density



$\tau = 4.31$ fm

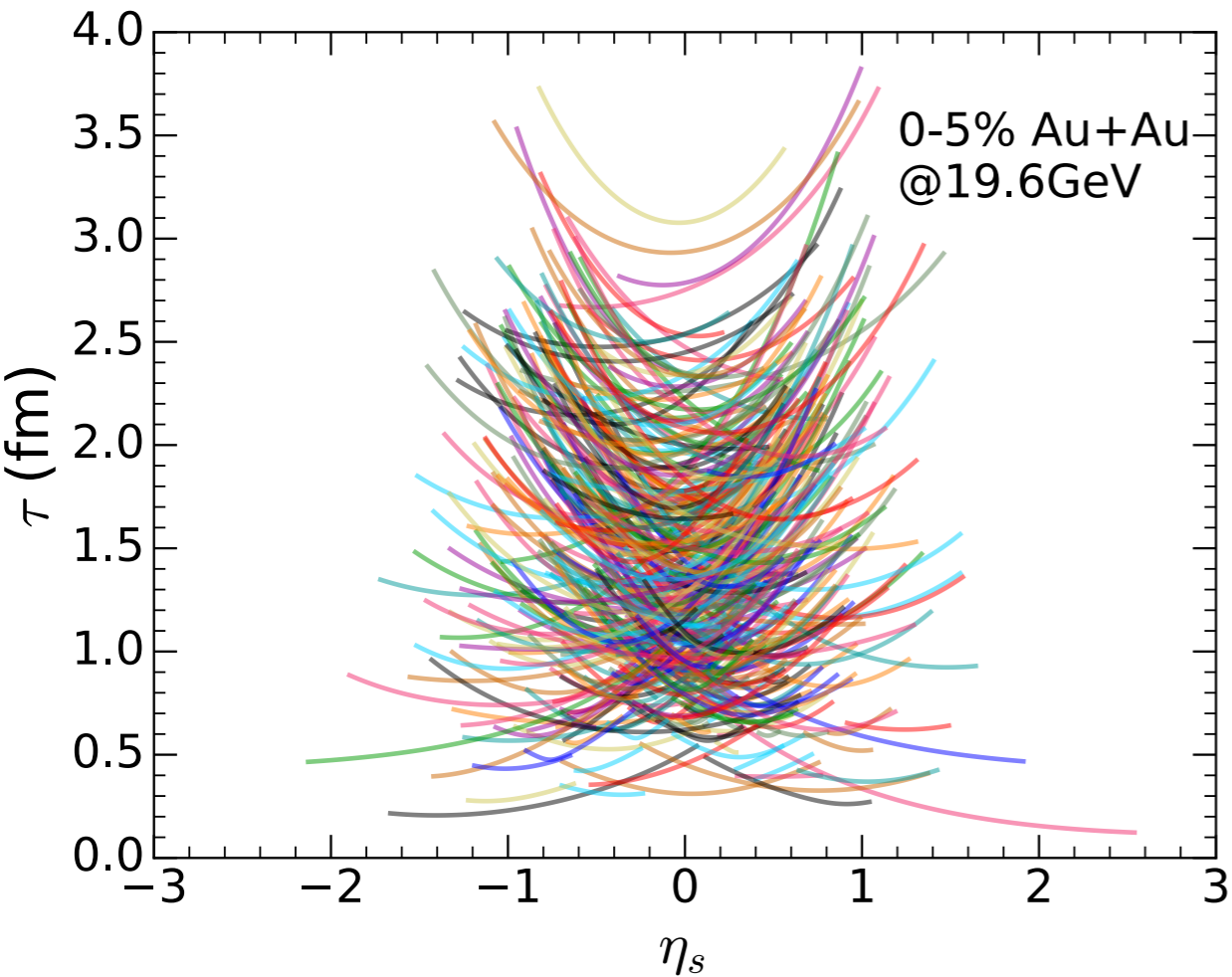


$\sqrt{s_{NN}} = 19.6$ GeV
valence quark + LEXUS

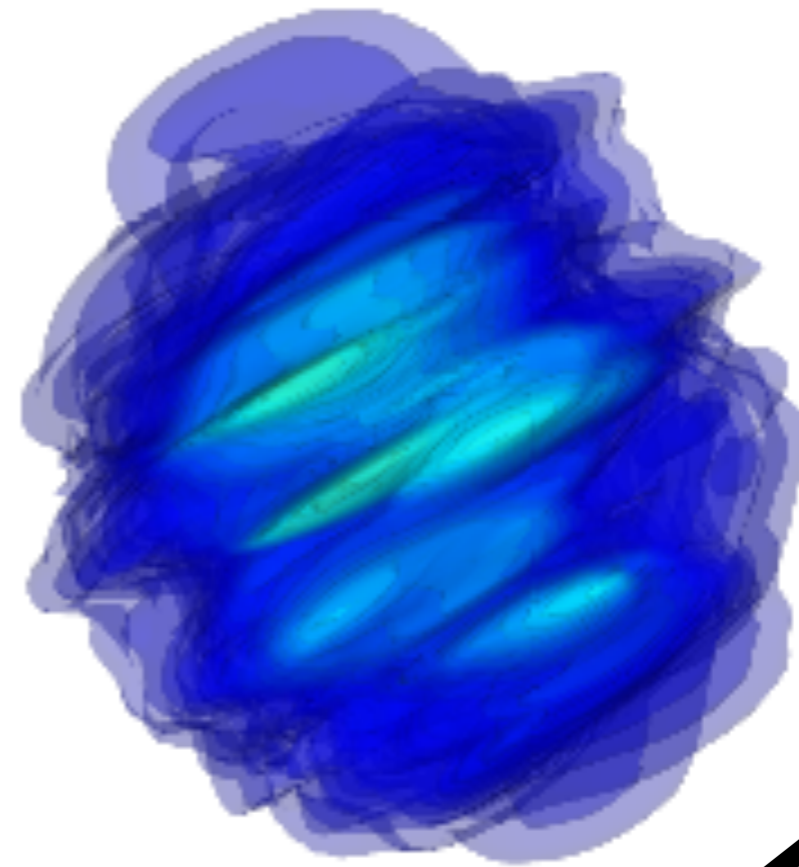


Hydrodynamical evolution with sources

energy density



$\tau = 5.91$ fm

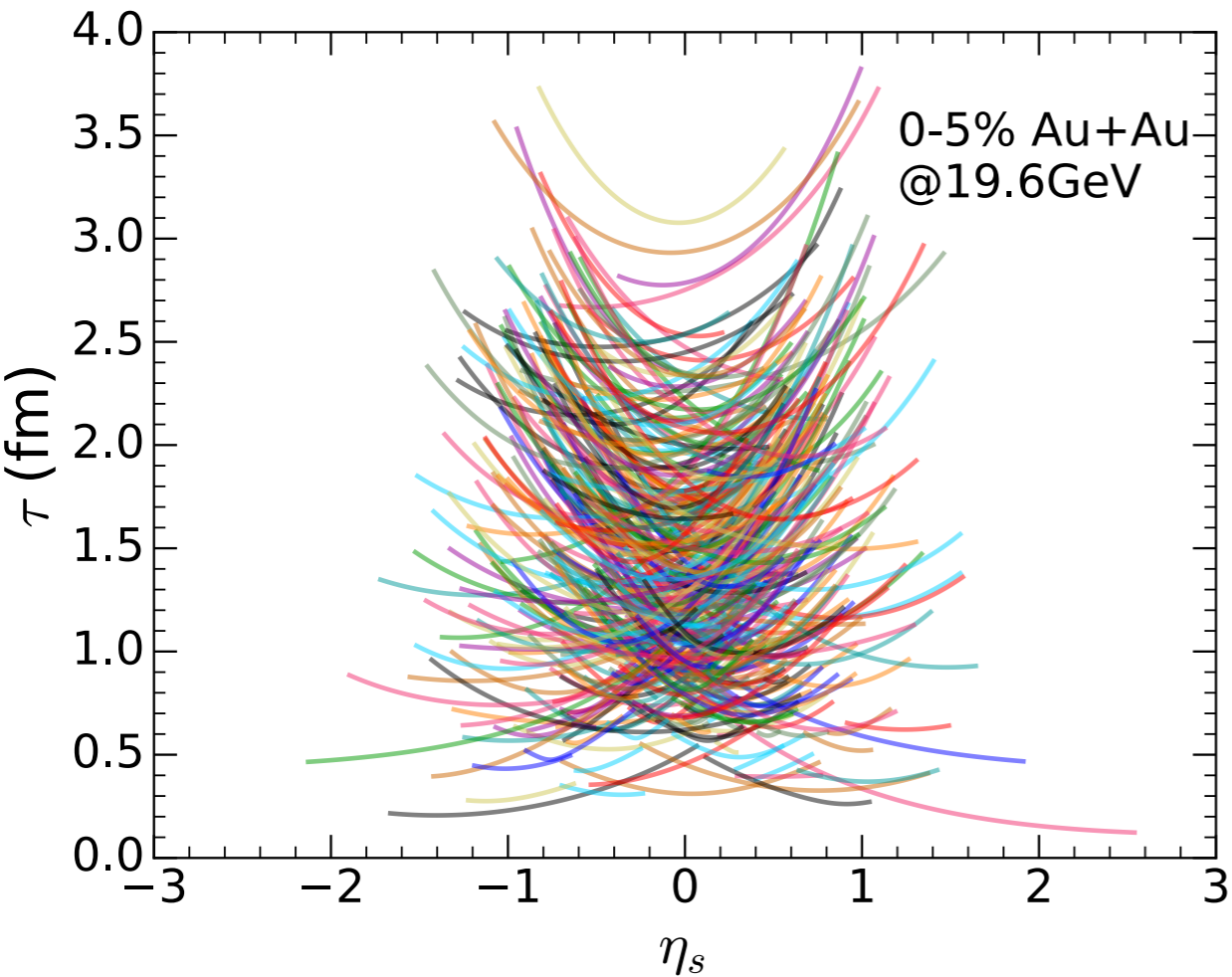


$$\sqrt{s_{NN}} = 19.6 \text{ GeV}$$

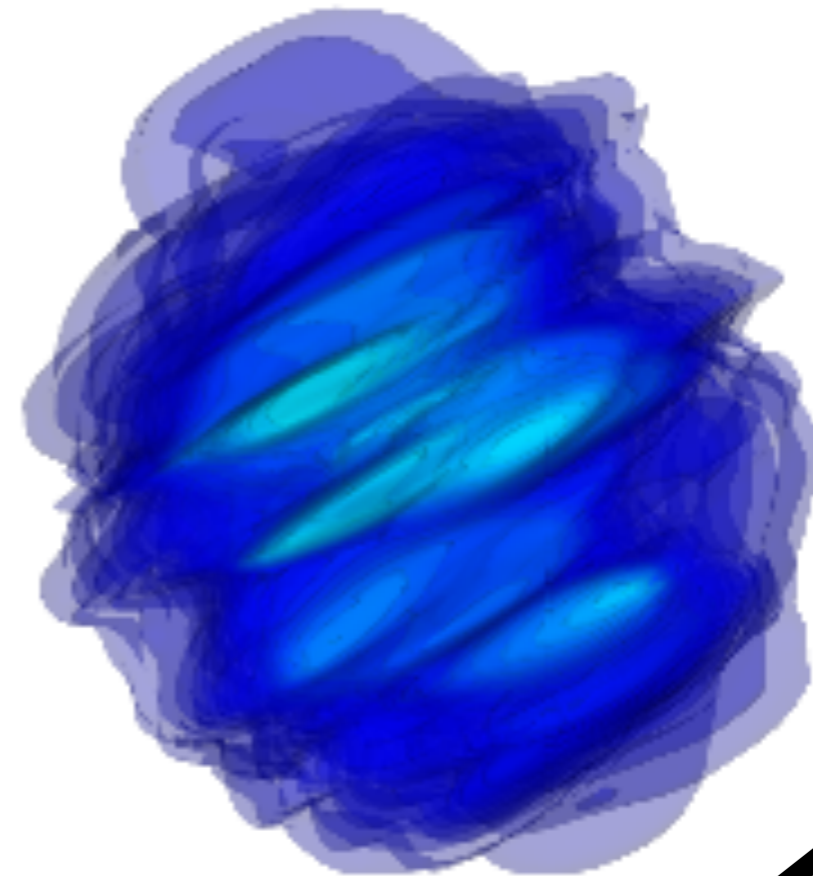
valence quark + LEXUS

Hydrodynamical evolution with sources

energy density



$\tau = 6.19$ fm

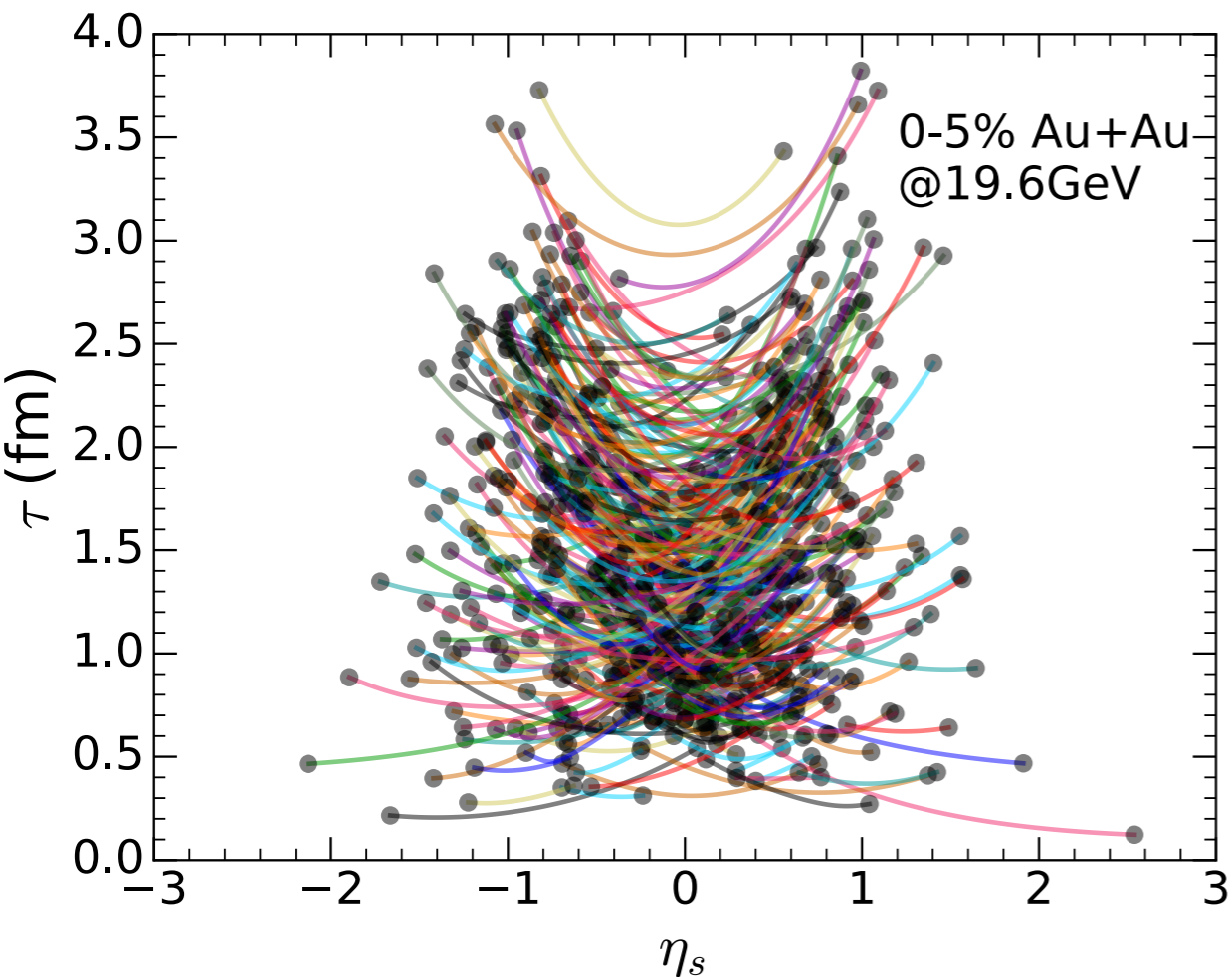


$$\sqrt{s_{NN}} = 19.6 \text{ GeV}$$

valence quark + LEXUS

Hydrodynamical evolution with sources

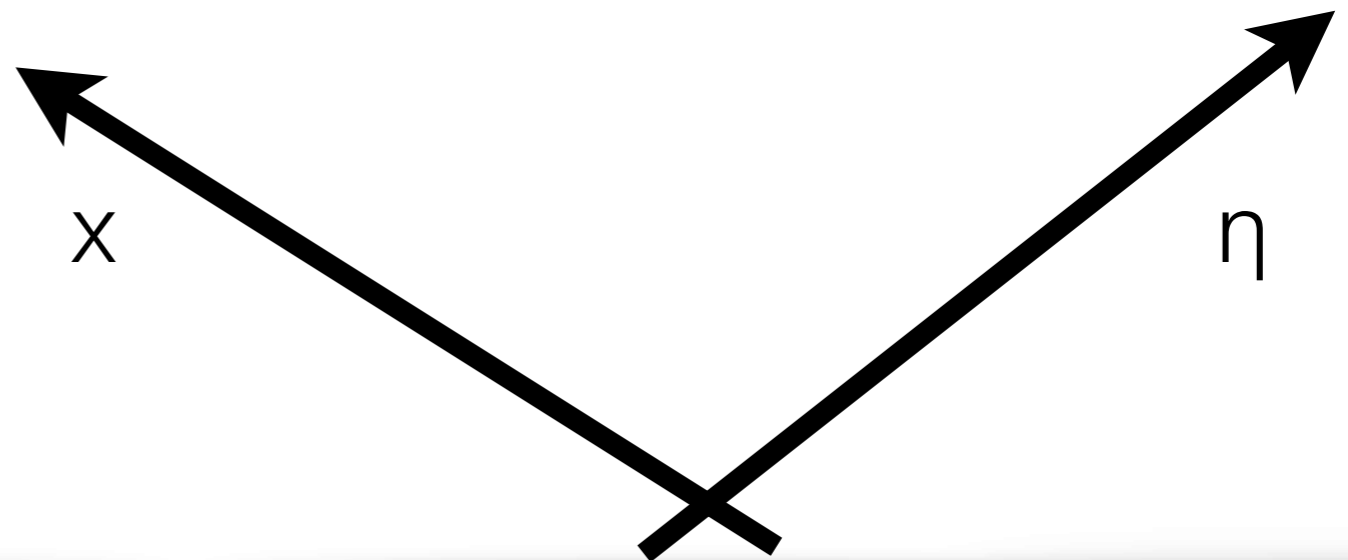
net baryon density



$\tau = 0.11 \text{ fm}$

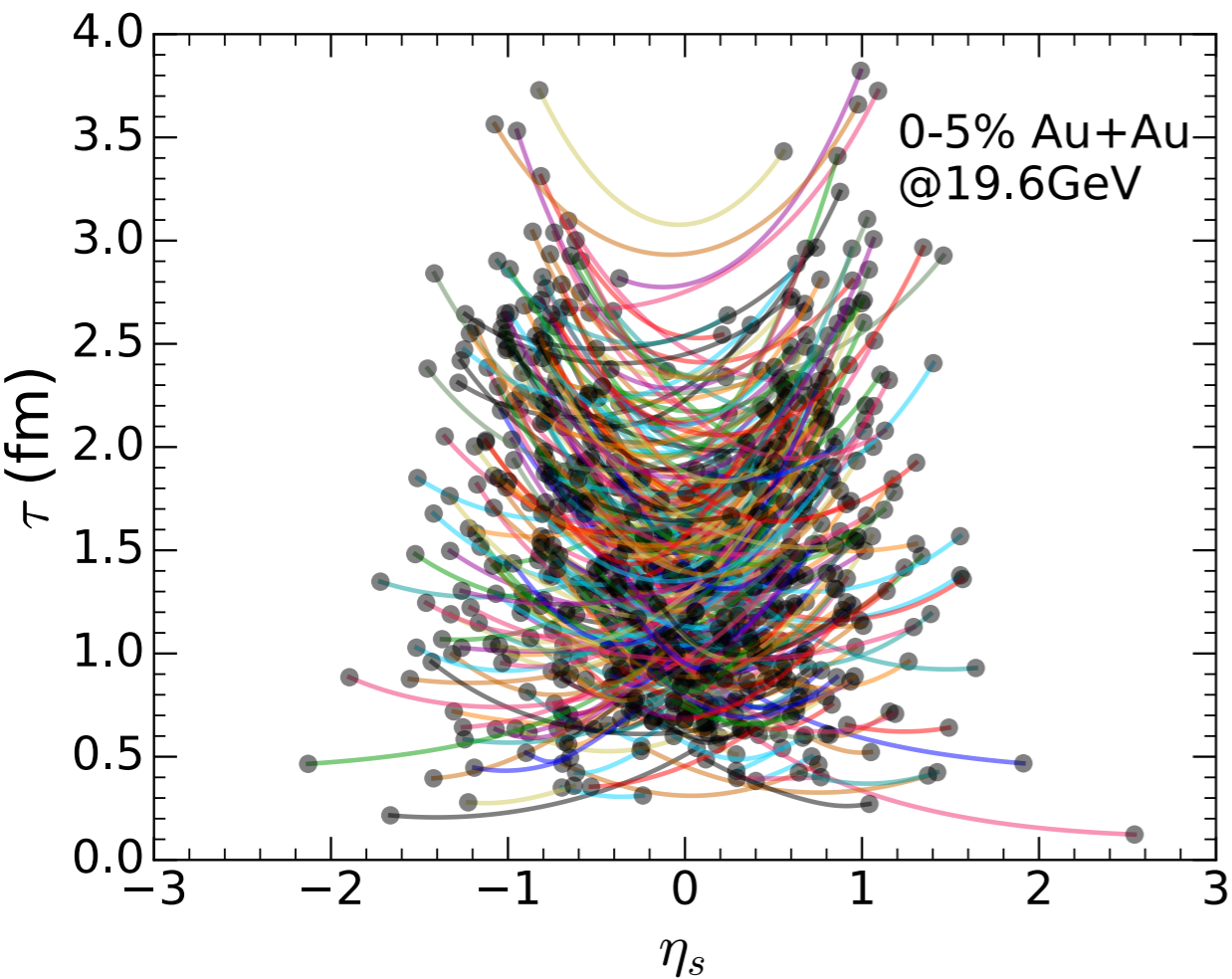
$$\sqrt{s_{NN}} = 19.6 \text{ GeV}$$

valence quark + LEXUS



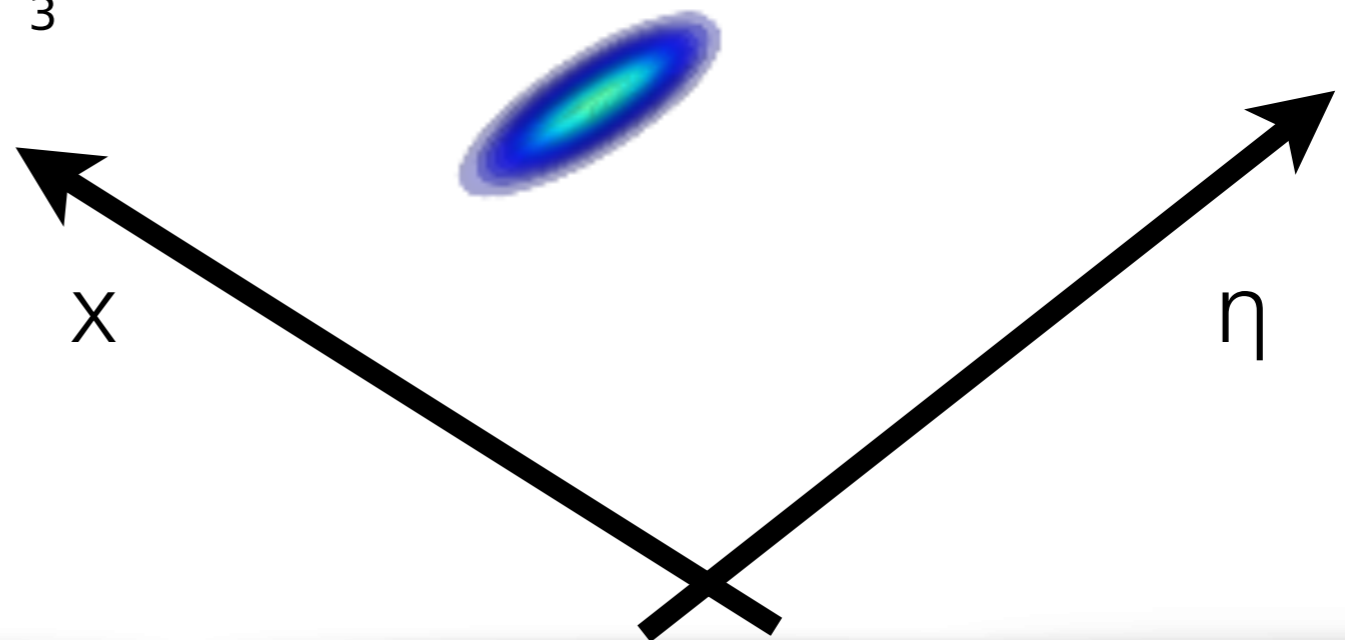
Hydrodynamical evolution with sources

net baryon density



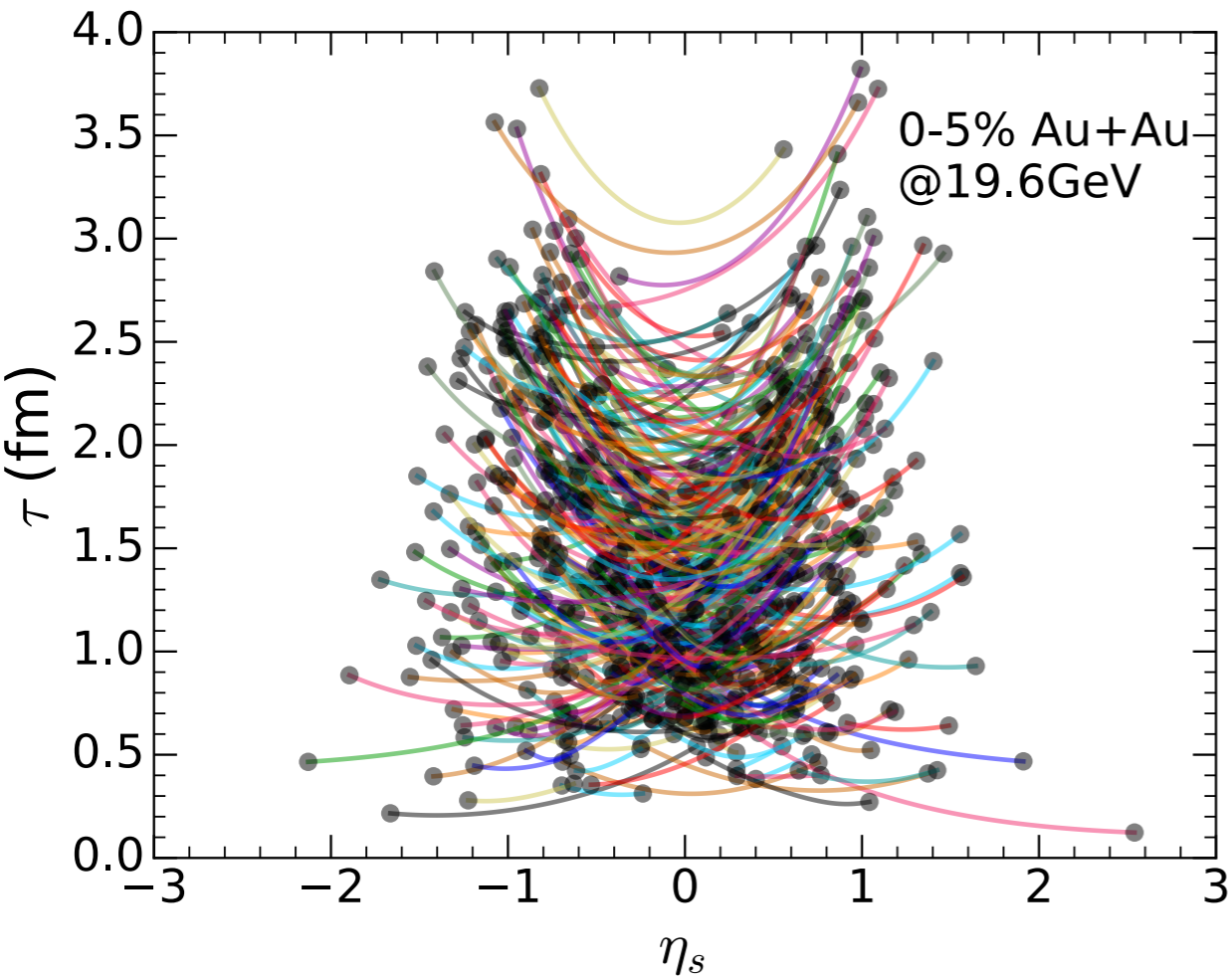
$\tau = 0.21$ fm

$\sqrt{s_{NN}} = 19.6$ GeV
valence quark + LEXUS

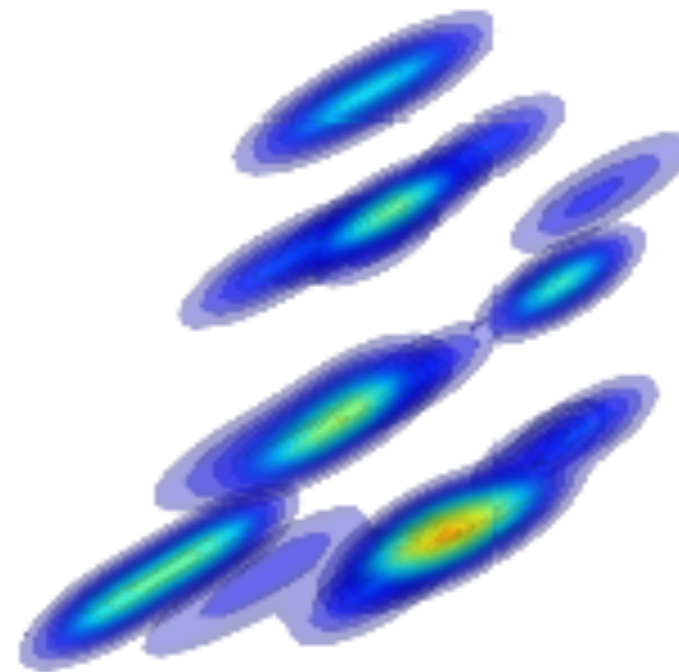


Hydrodynamical evolution with sources

net baryon density



$\tau = 0.51$ fm



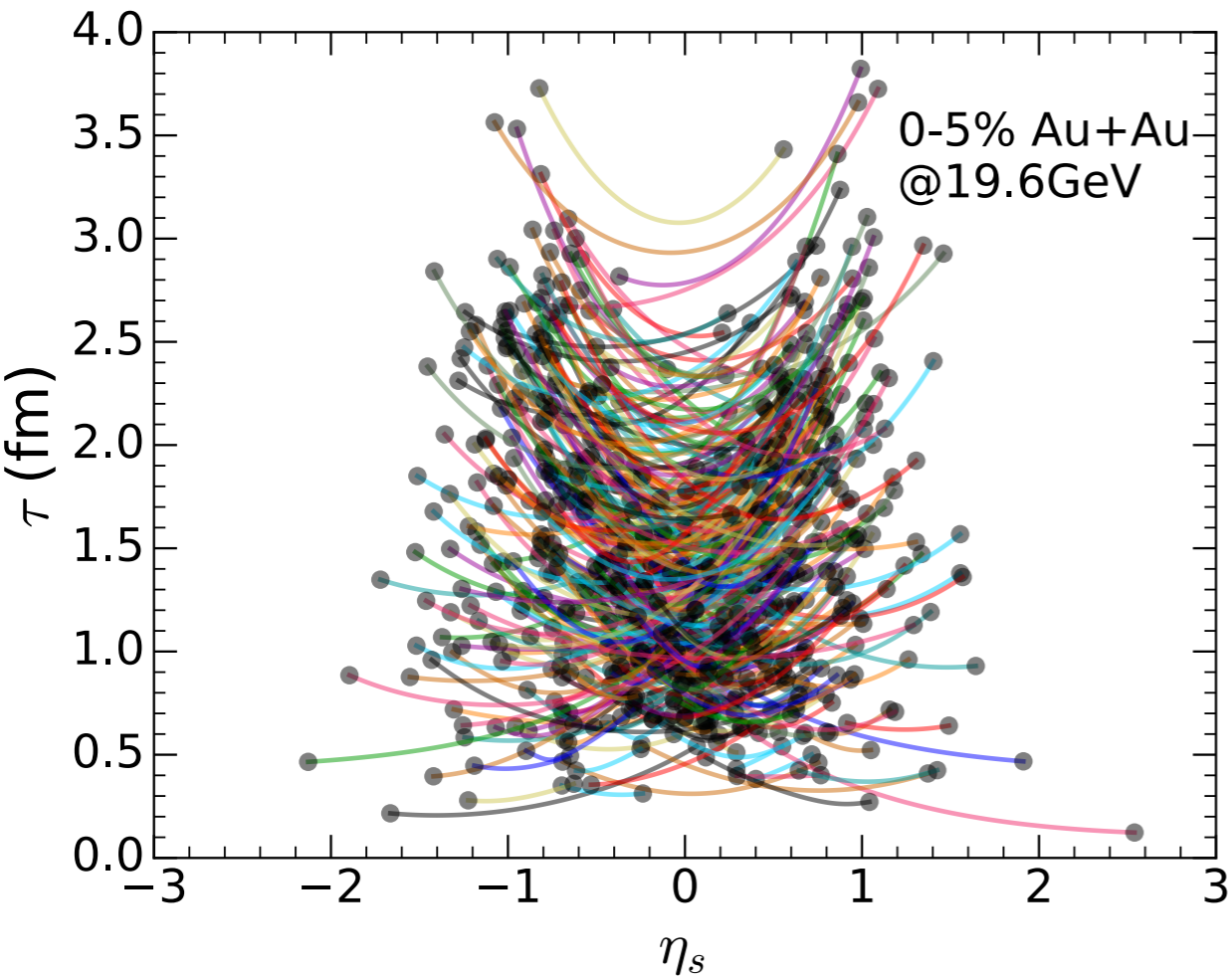
$$\sqrt{s_{NN}} = 19.6 \text{ GeV}$$

valence quark + LEXUS

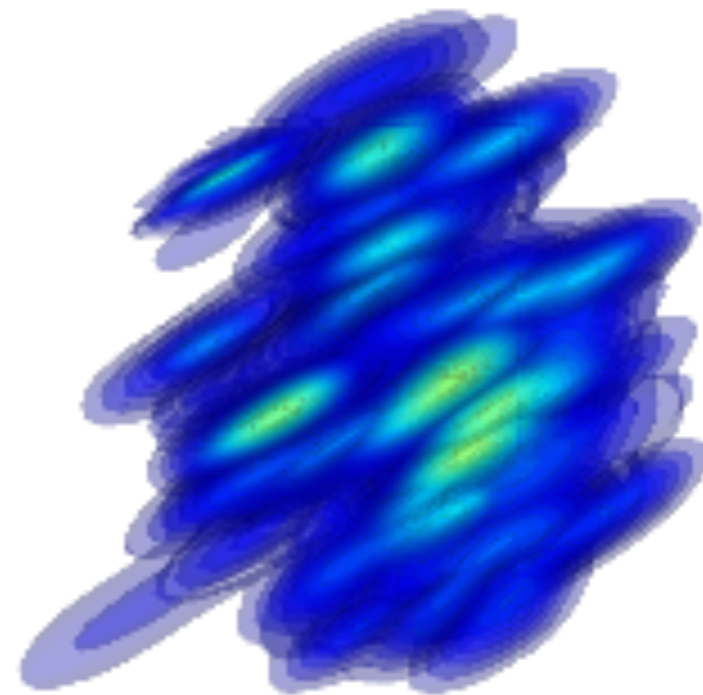


Hydrodynamical evolution with sources

net baryon density

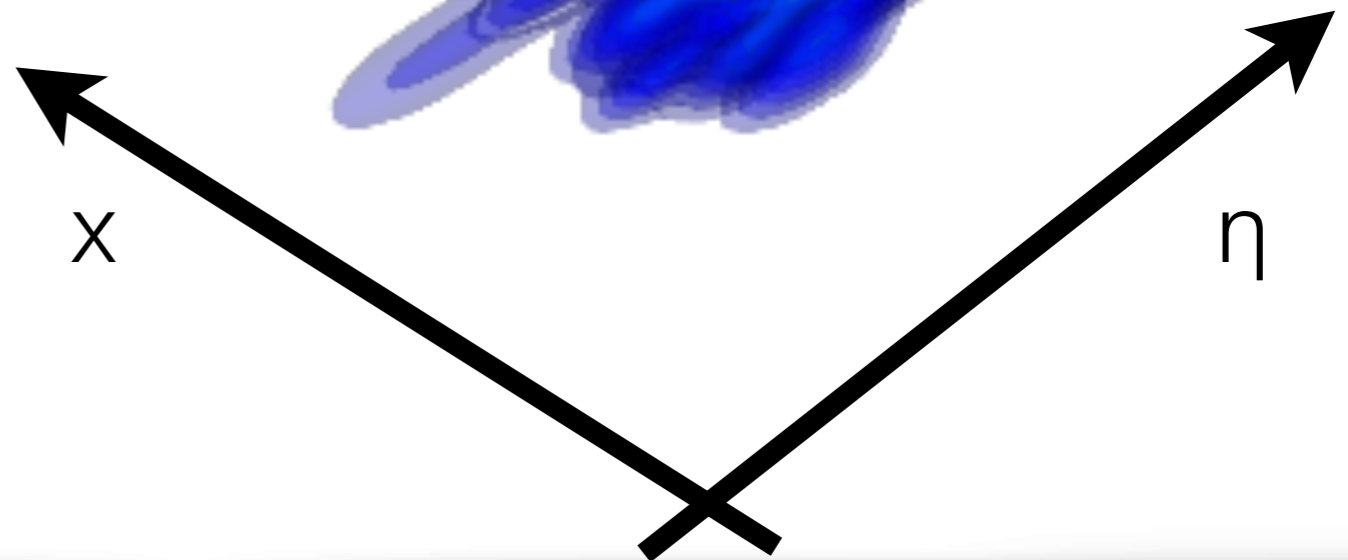


$\tau = 1.11$ fm



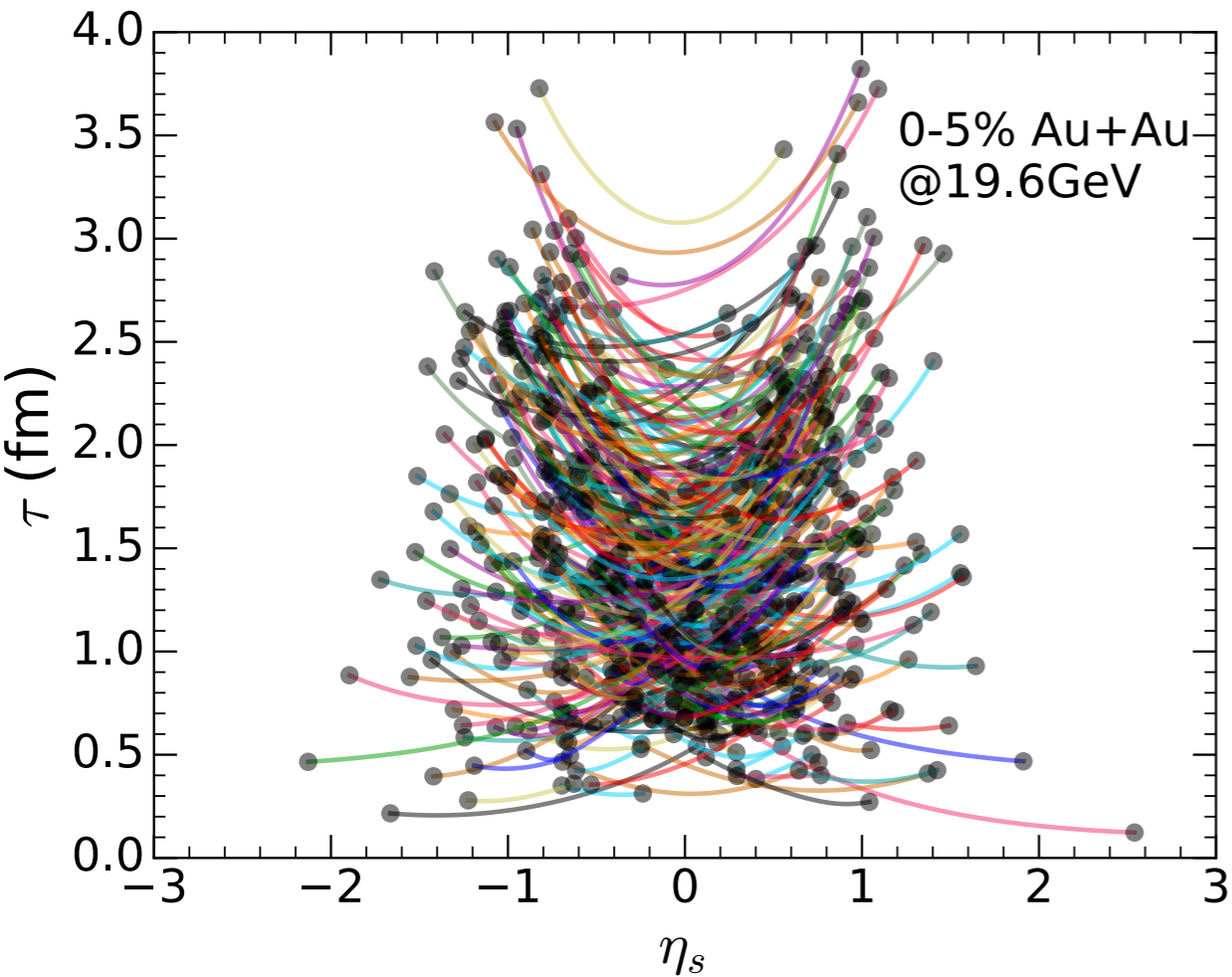
$$\sqrt{s_{NN}} = 19.6 \text{ GeV}$$

valence quark + LEXUS

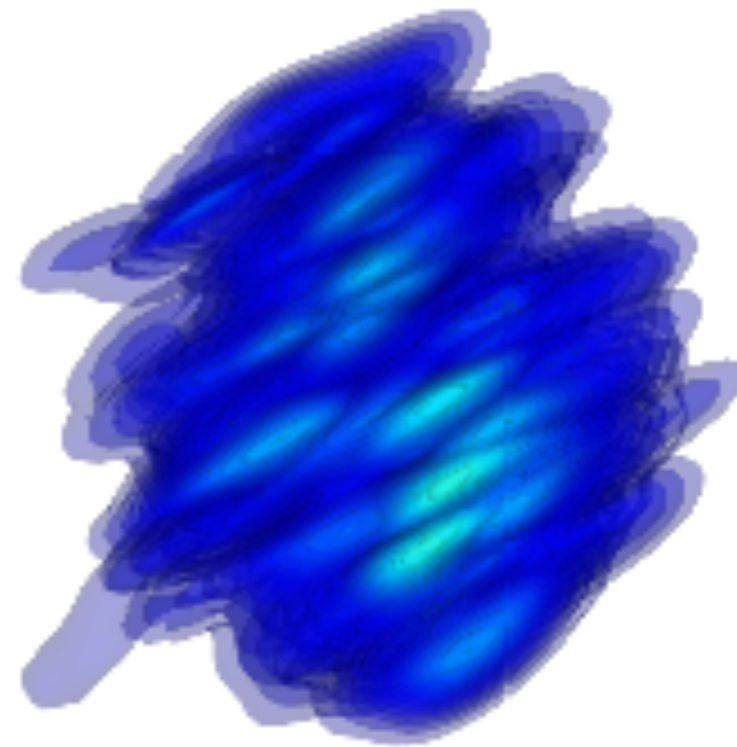


Hydrodynamical evolution with sources

net baryon density



$\tau = 1.51 \text{ fm}$



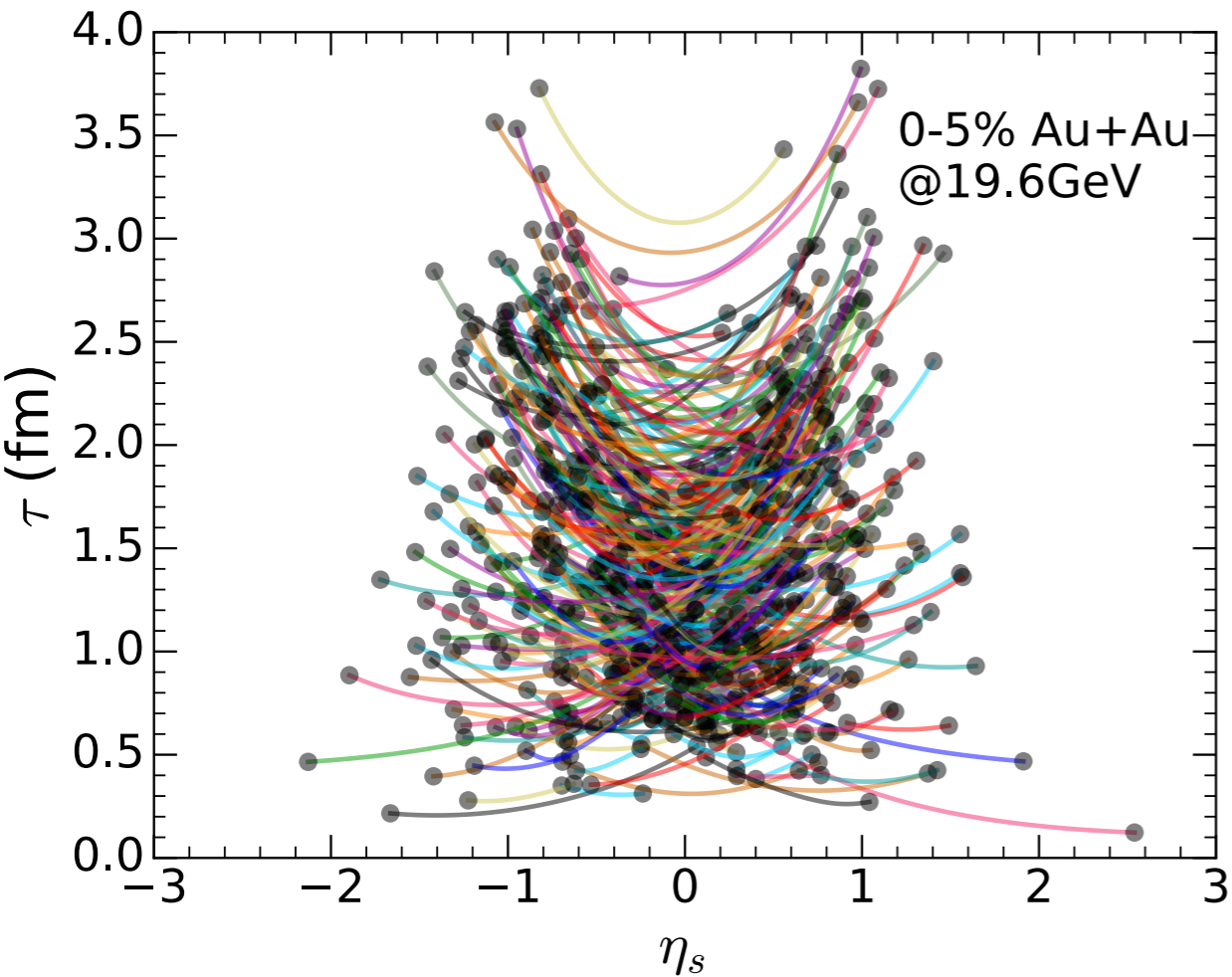
$$\sqrt{s_{NN}} = 19.6 \text{ GeV}$$

valence quark + LEXUS

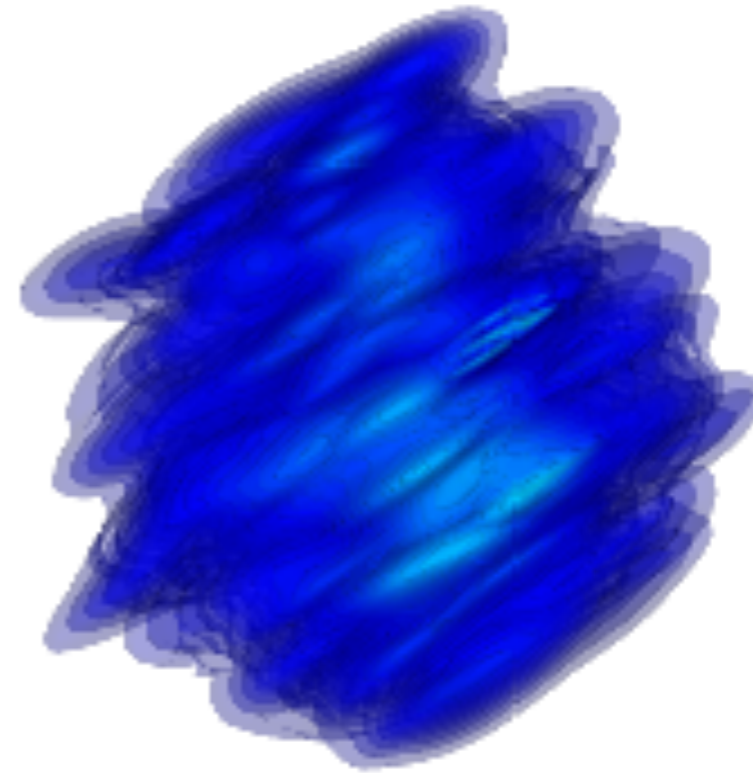


Hydrodynamical evolution with sources

net baryon density

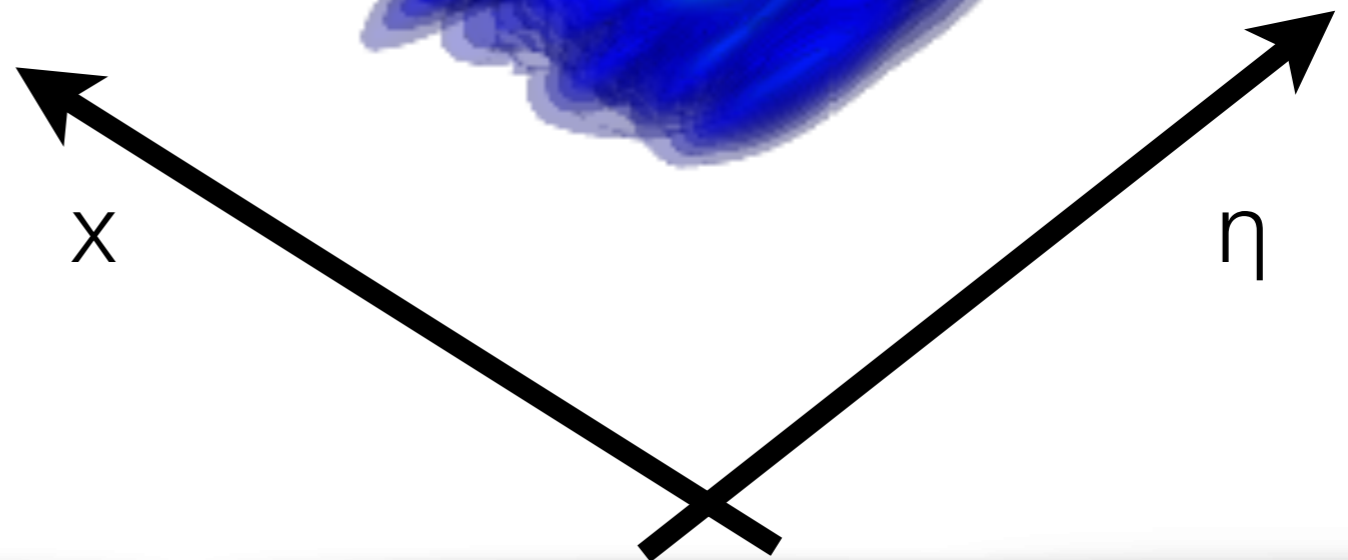


$\tau = 1.91 \text{ fm}$



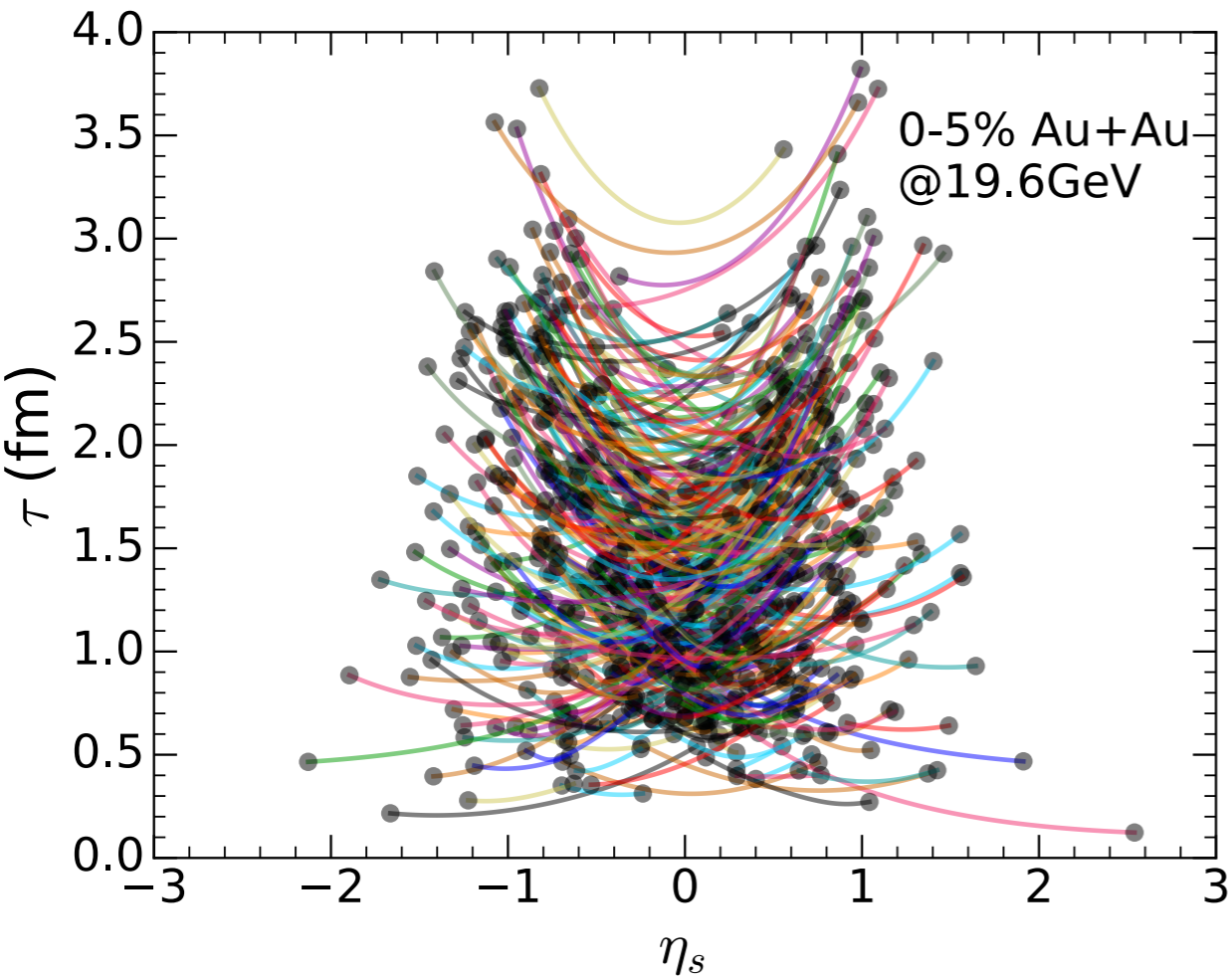
$$\sqrt{s_{NN}} = 19.6 \text{ GeV}$$

valence quark + LEXUS

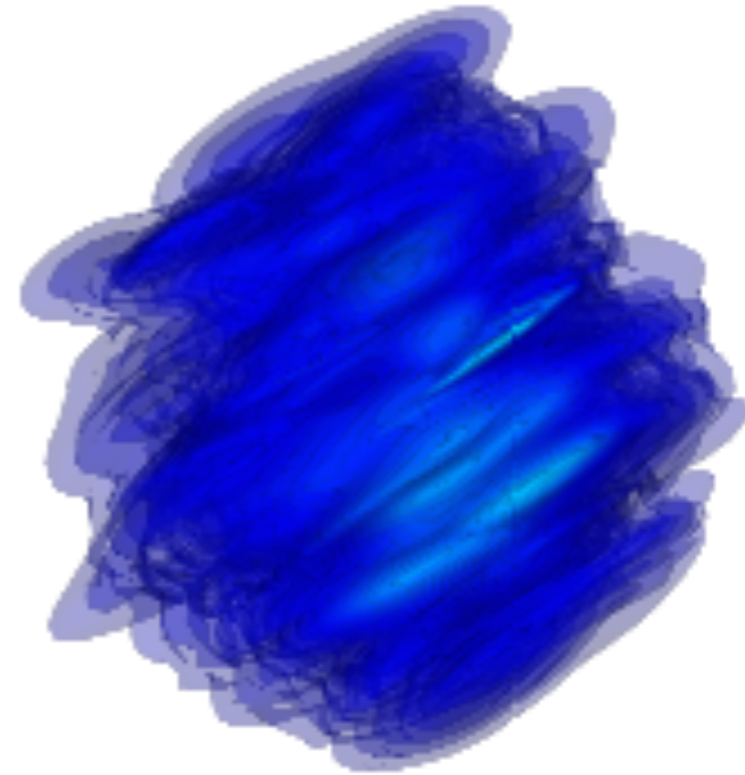


Hydrodynamical evolution with sources

net baryon density

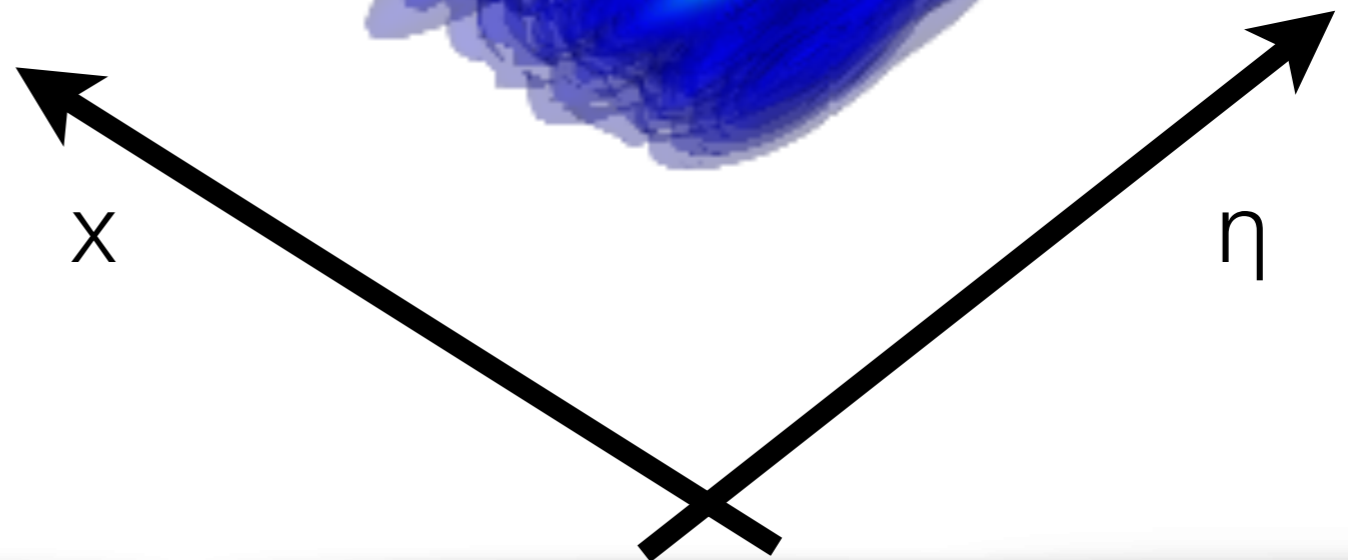


$\tau = 2.31 \text{ fm}$



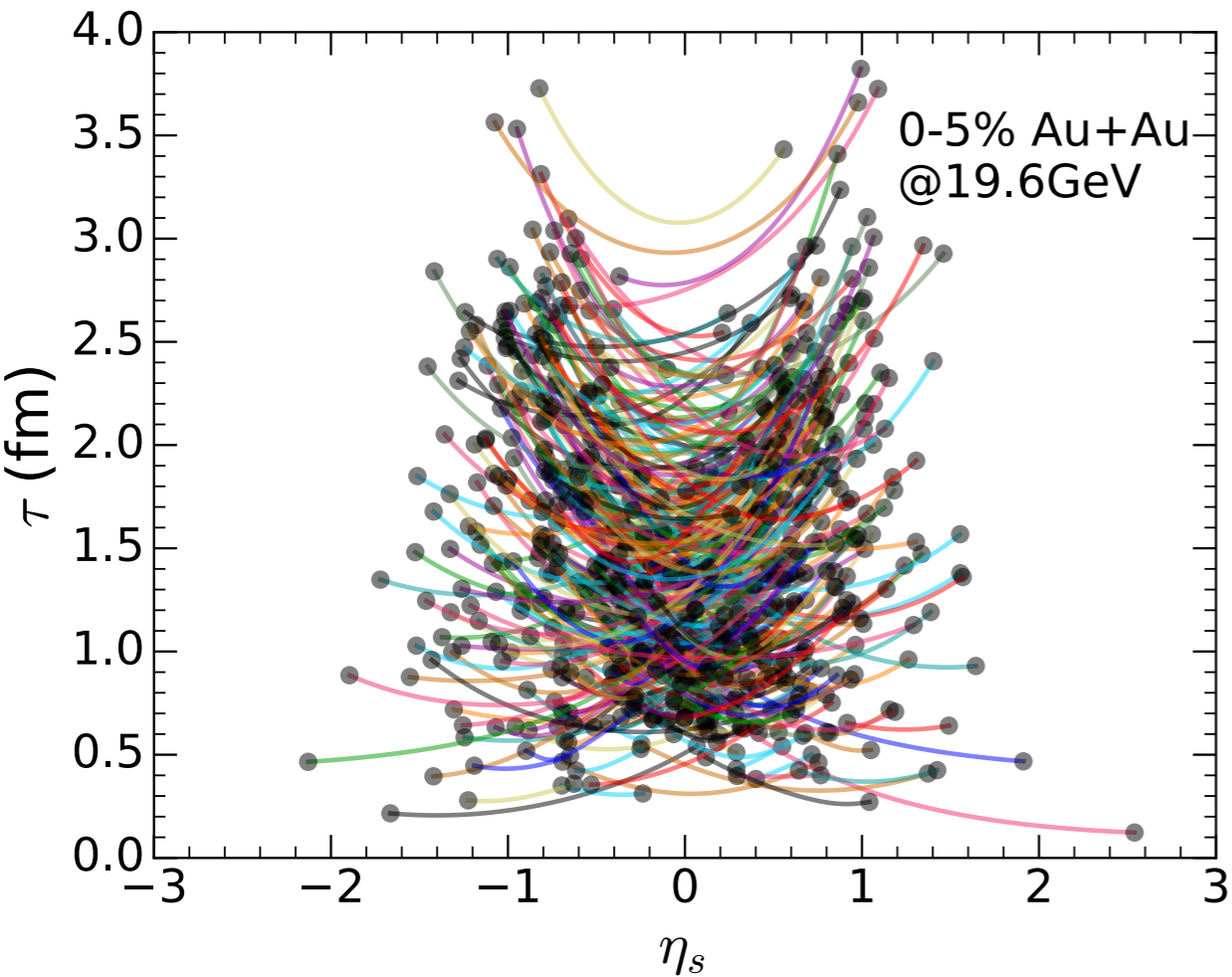
$$\sqrt{s_{NN}} = 19.6 \text{ GeV}$$

valence quark + LEXUS

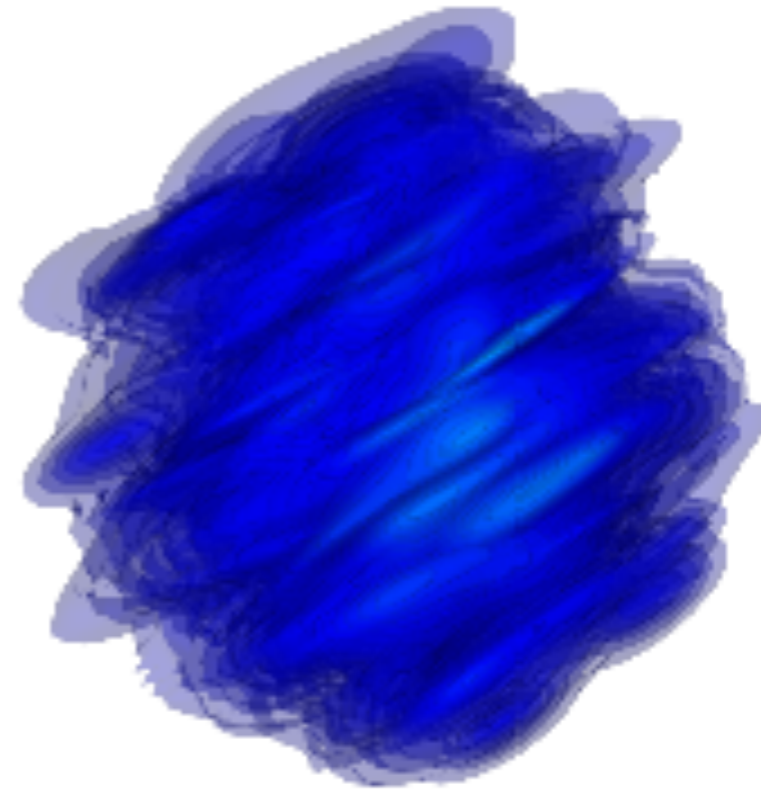


Hydrodynamical evolution with sources

net baryon density



$\tau = 2.71$ fm

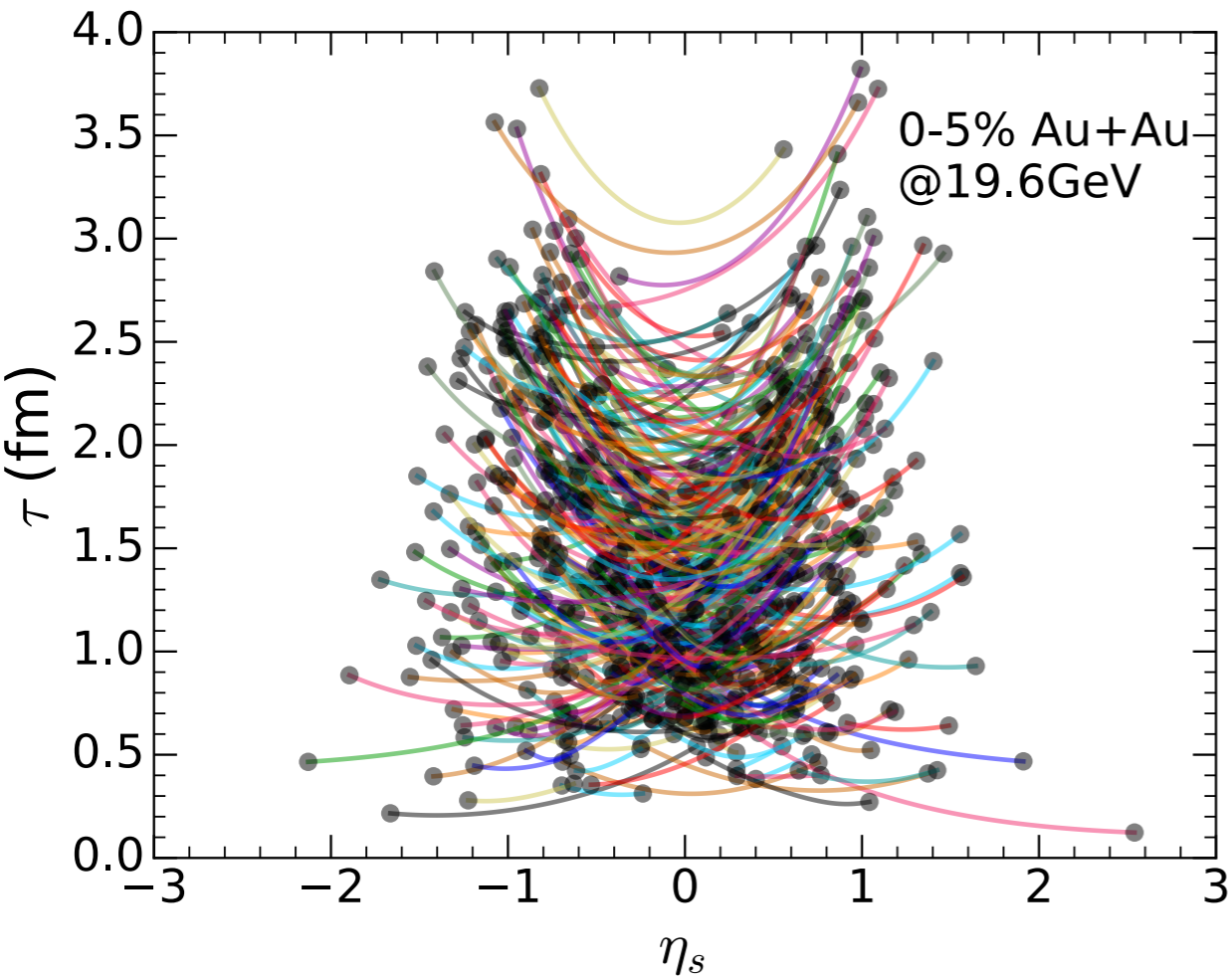


$\sqrt{s_{NN}} = 19.6$ GeV
valence quark + LEXUS

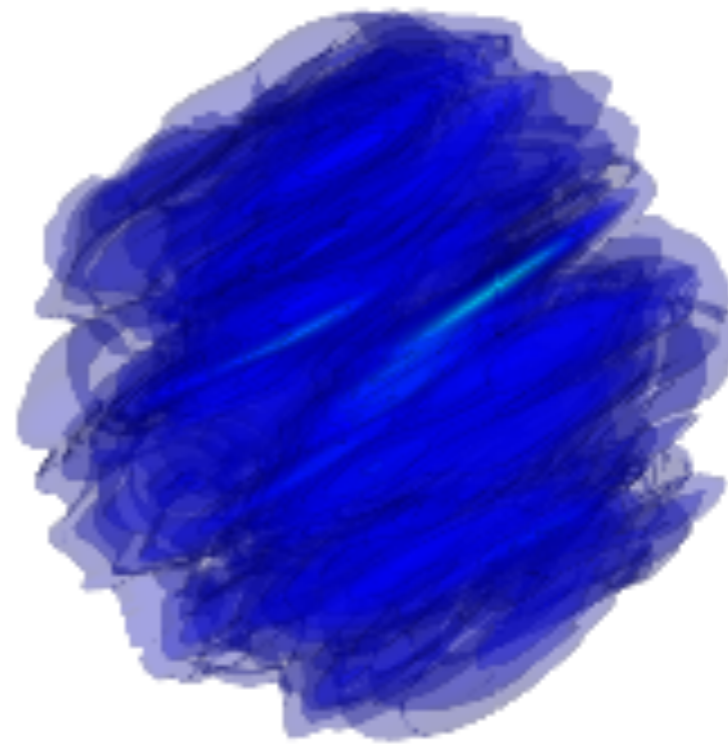


Hydrodynamical evolution with sources

net baryon density



$\tau = 3.51$ fm

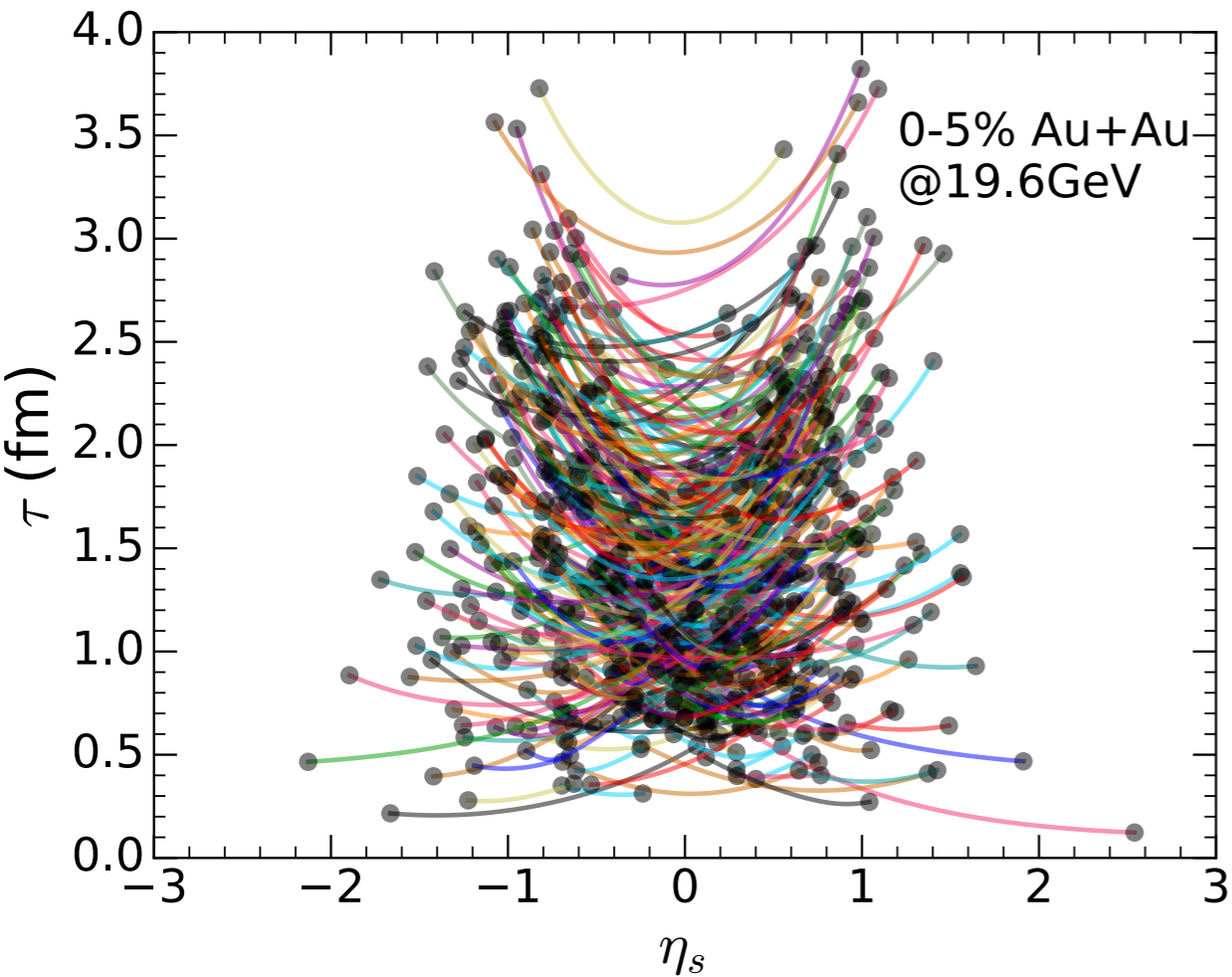


$\sqrt{s_{NN}} = 19.6$ GeV
valence quark + LEXUS

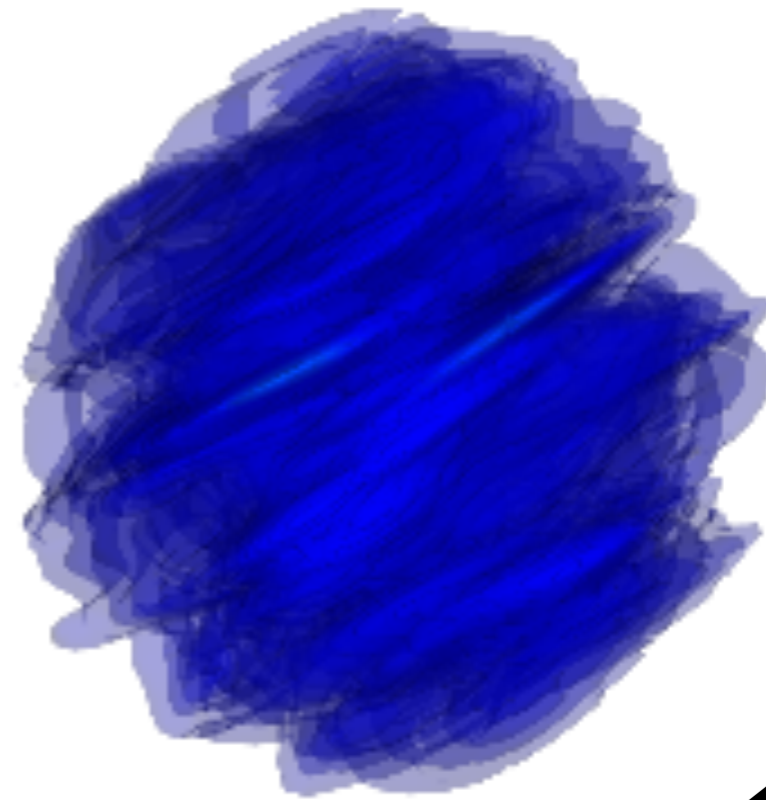


Hydrodynamical evolution with sources

net baryon density



$\tau = 4.11$ fm

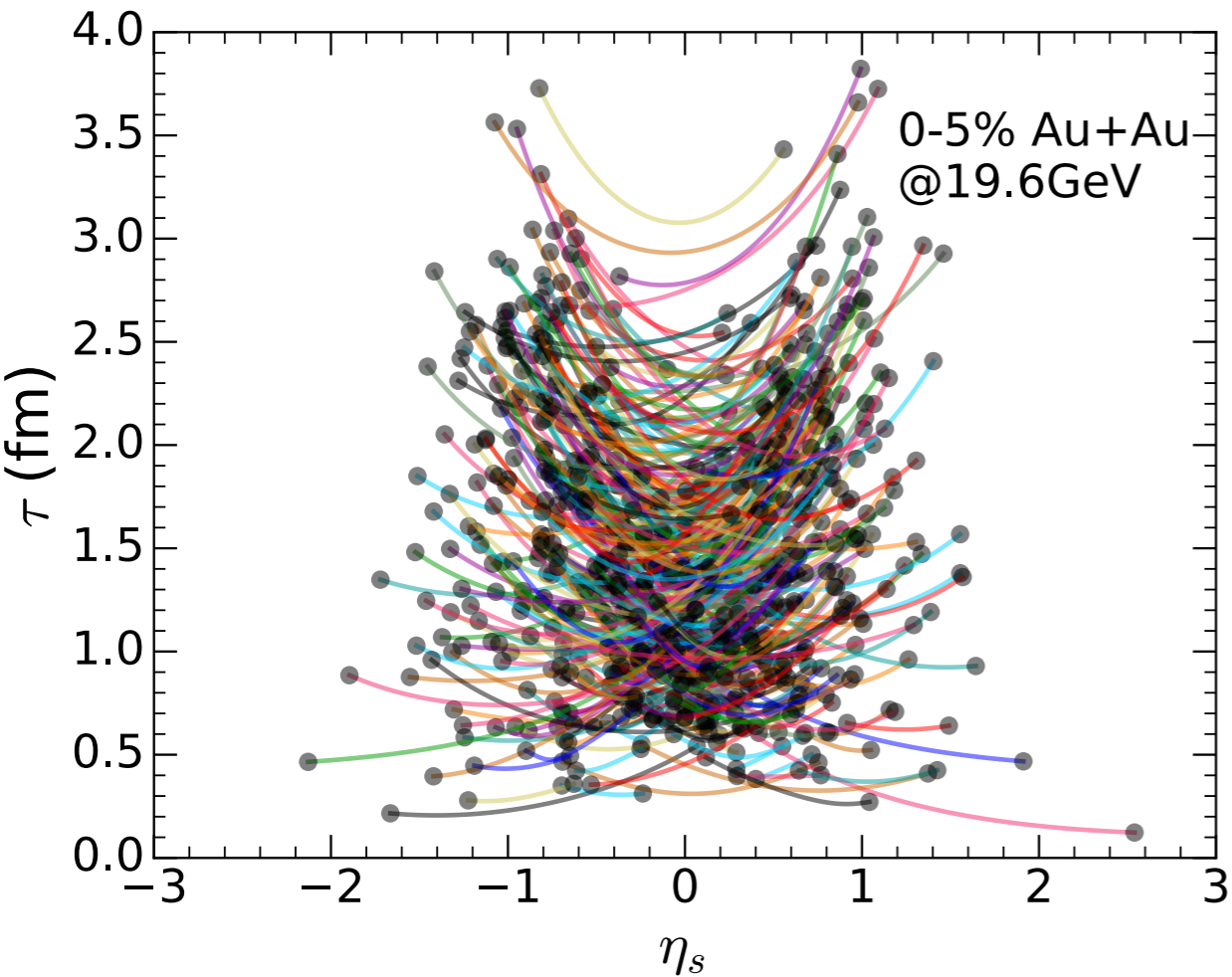


$$\sqrt{s_{NN}} = 19.6 \text{ GeV}$$

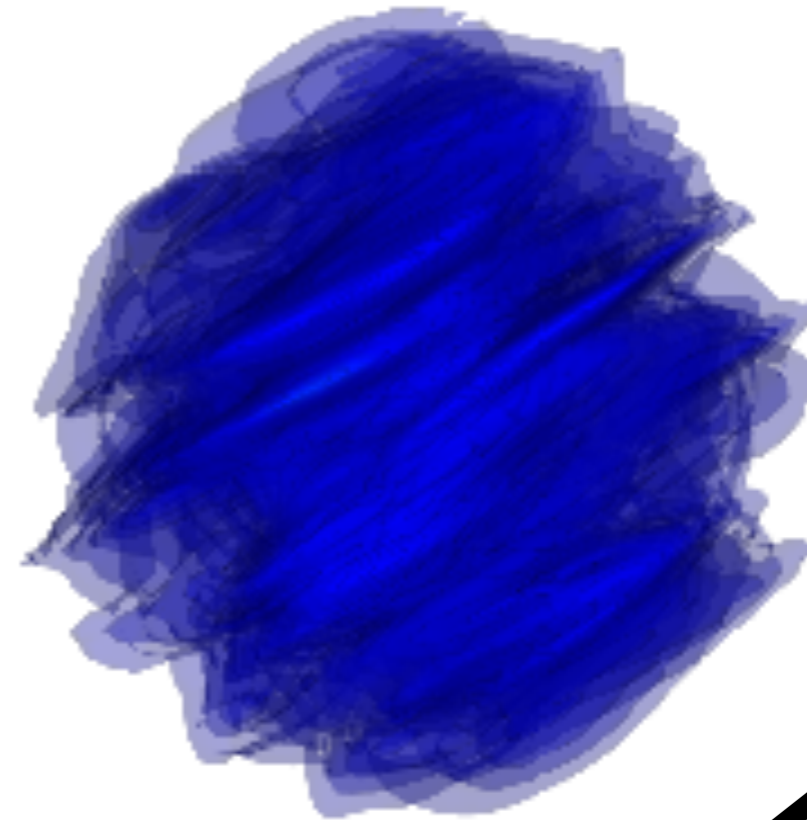
valence quark + LEXUS

Hydrodynamical evolution with sources

net baryon density



$\tau = 4.51$ fm

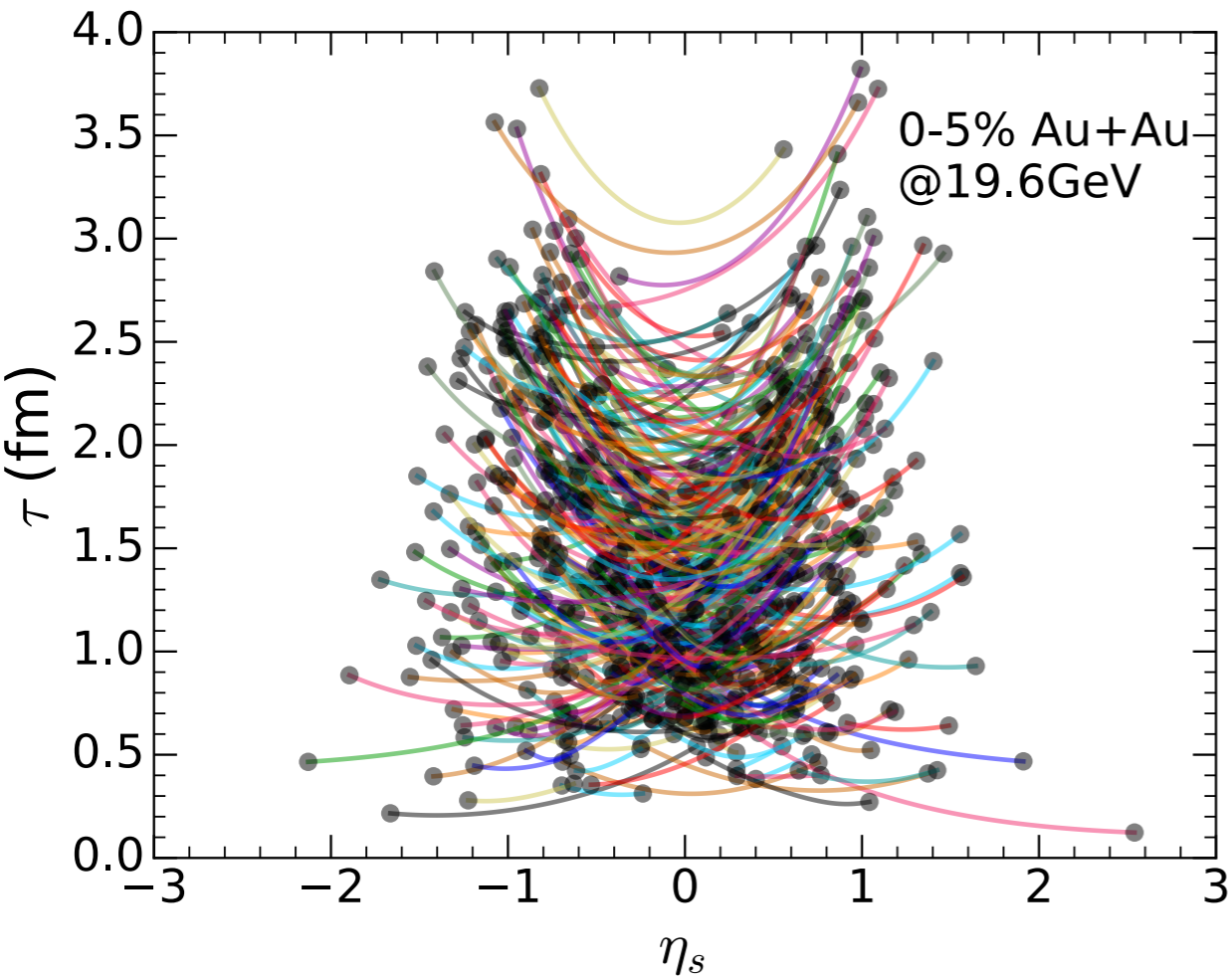


$\sqrt{s_{NN}} = 19.6$ GeV
valence quark + LEXUS

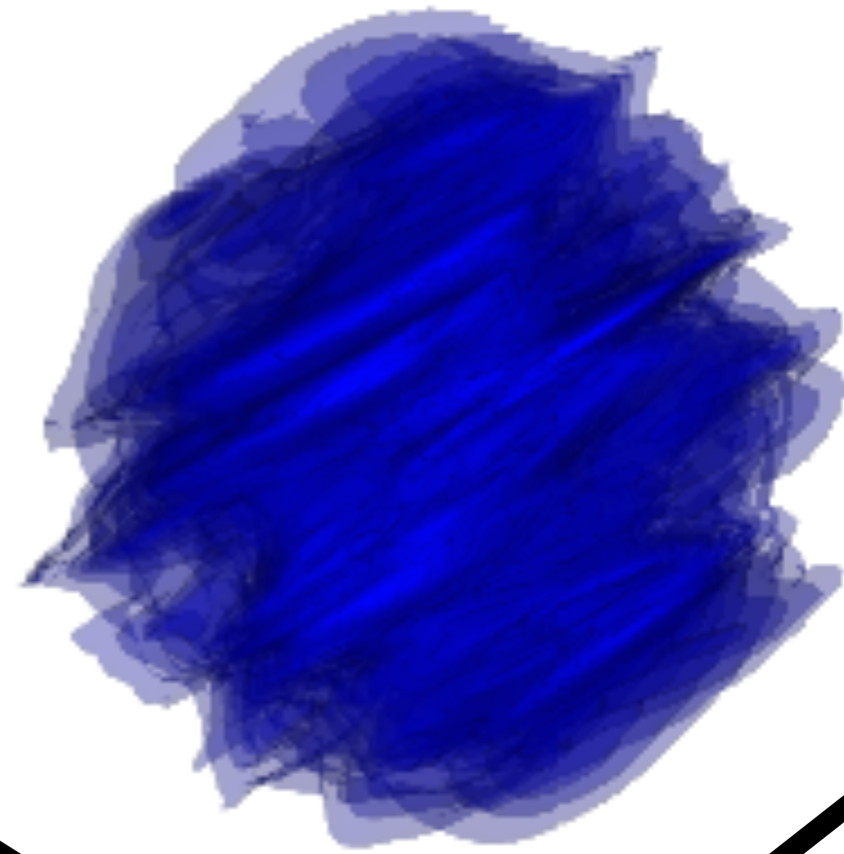


Hydrodynamical evolution with sources

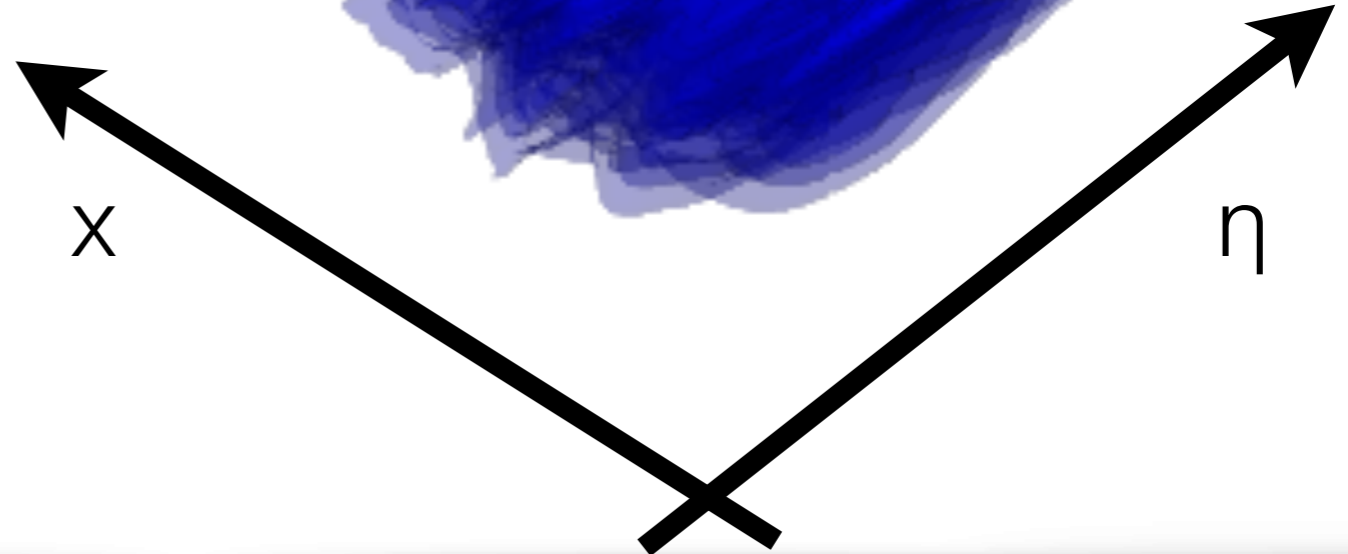
net baryon density



$\tau = 4.91$ fm

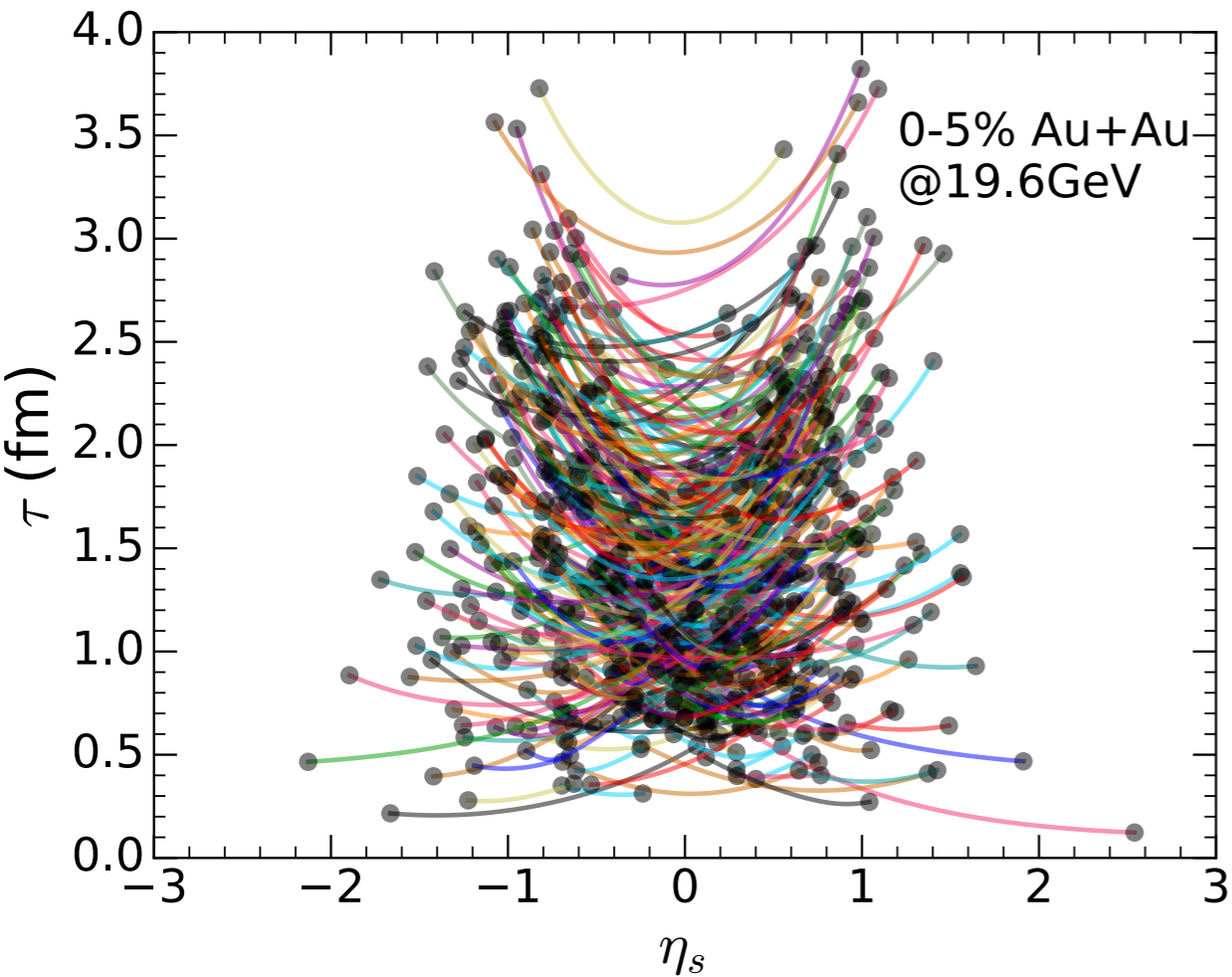


$\sqrt{s_{NN}} = 19.6$ GeV
valence quark + LEXUS

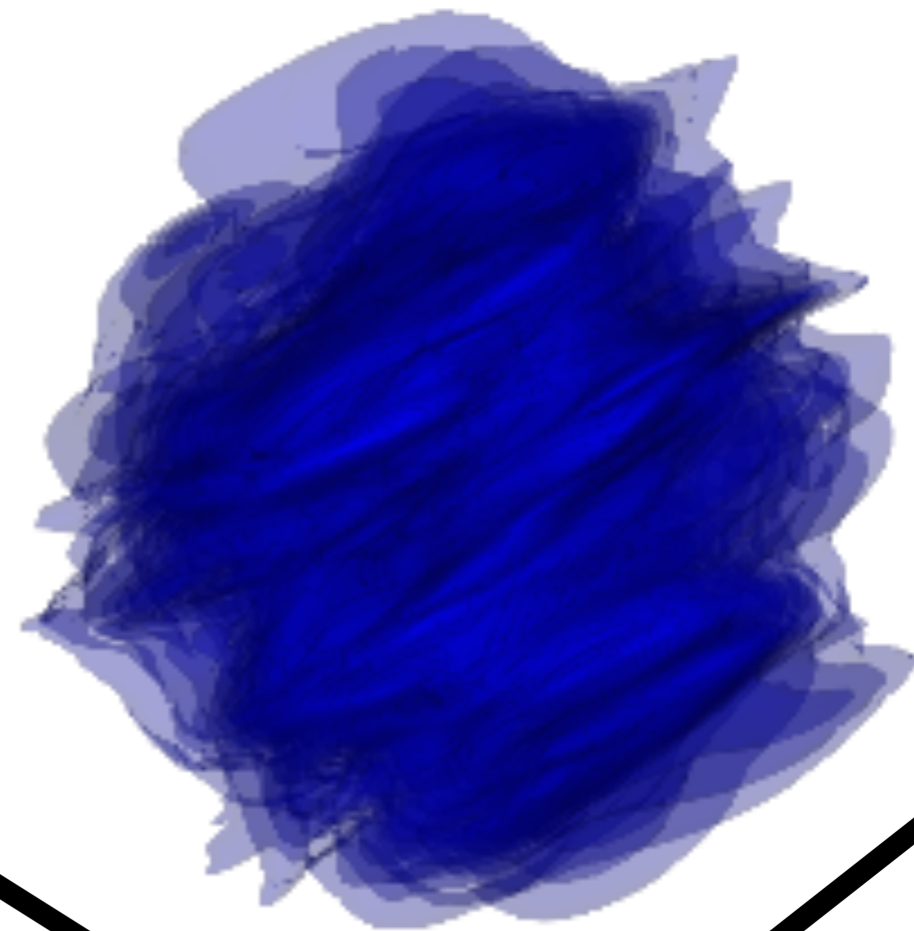


Hydrodynamical evolution with sources

net baryon density



$\tau = 6.11$ fm



$$\sqrt{s_{NN}} = 19.6 \text{ GeV}$$

valence quark + LEXUS



Progress in hydrodynamics

Dissipative hydrodynamics

Energy momentum tensor

$$T^{\mu\nu} = e u^\mu u^\nu - (P + \Pi) \Delta^{\mu\nu} + \pi^{\mu\nu} \quad \Delta^{\mu\nu} = g^{\mu\nu} - u^\mu u^\nu$$

Conserved currents

$$J^\mu = n u^\mu + q^\mu$$

Equations of motion

$$\begin{aligned} \partial_\mu T^{\mu\nu} &= 0 \\ \partial_\mu J^\mu &= 0 \end{aligned} \quad + \quad P(e, n)$$

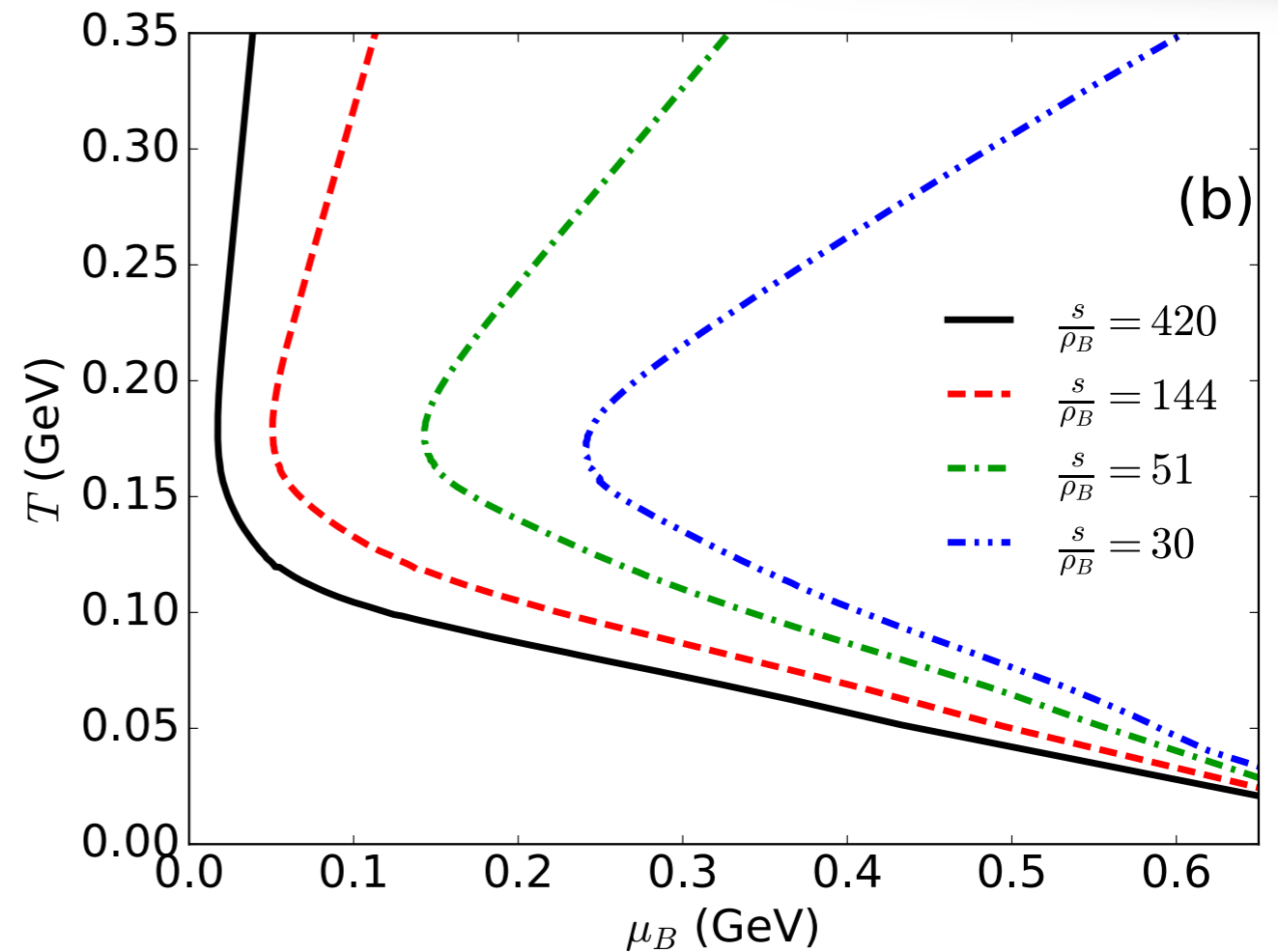
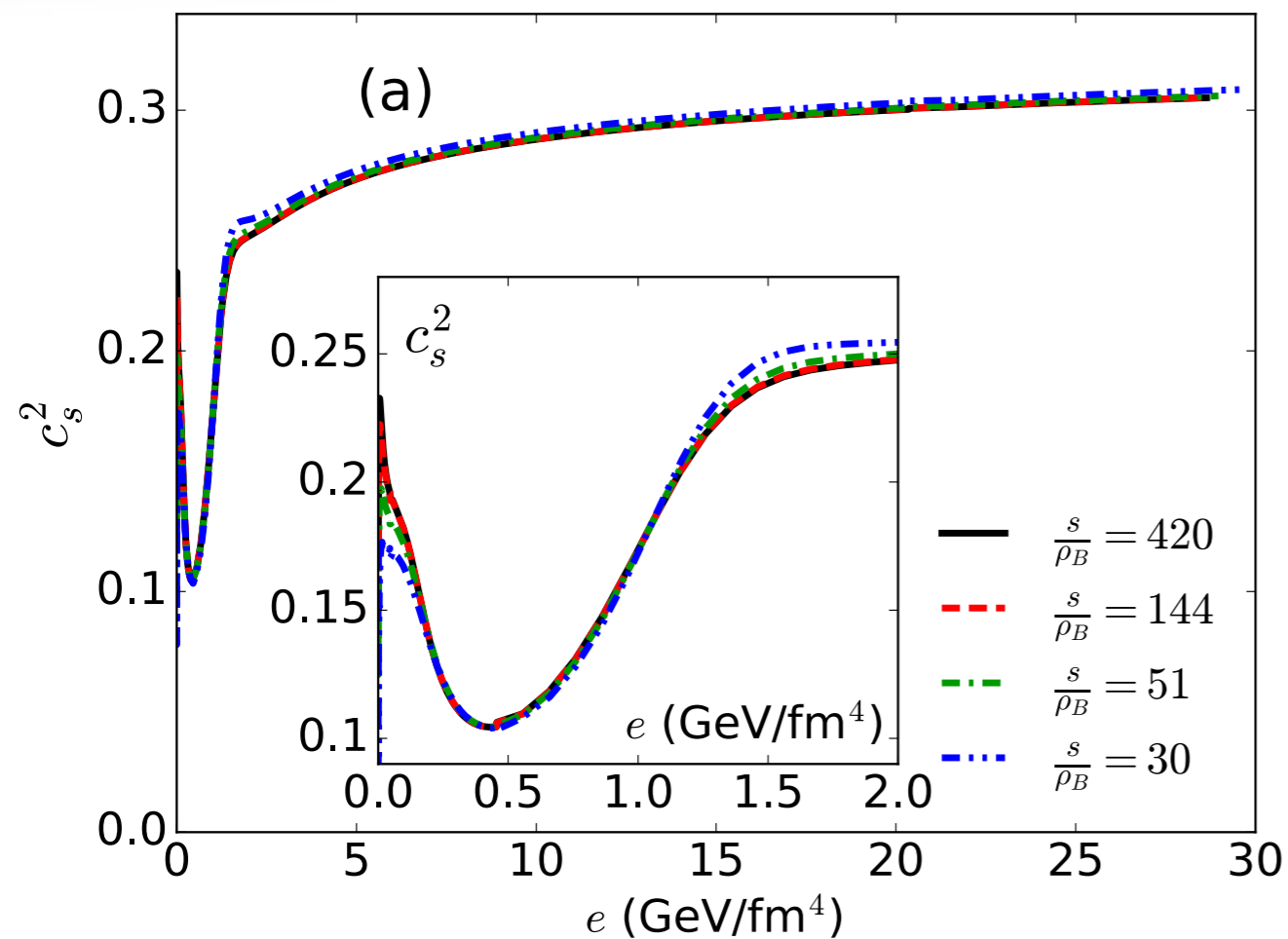
Dissipative quantities are evolved with 2nd order Israel-Stewart type of equations

At Navier-Stokes limit,

$$\pi^{\mu\nu} \sim 2\eta \nabla^{\langle\mu} u^{\nu\rangle} \quad \Pi \sim -\zeta \partial_\mu u^\mu \quad q^\mu \sim \kappa \nabla^\mu \frac{\mu}{T}$$

$$\nabla^\mu = \Delta^{\mu\nu} \partial_\nu$$

EoS at finite μ_B



High temperature:

- Lattice QCD EoS up to $\mathcal{O}(\mu_B^4)$

Low temperature:

- Glued with hadron resonance gas EoS

Transport coefficients

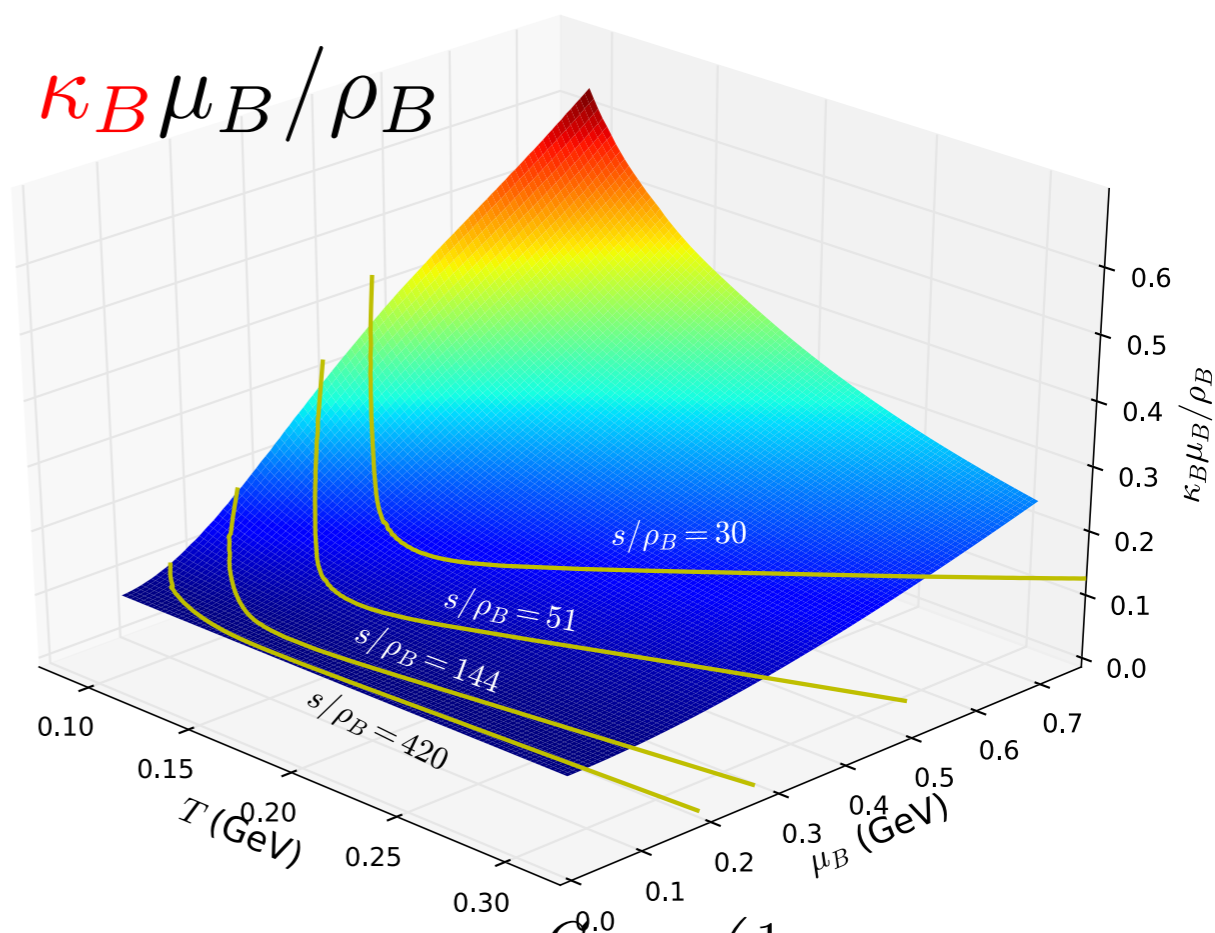
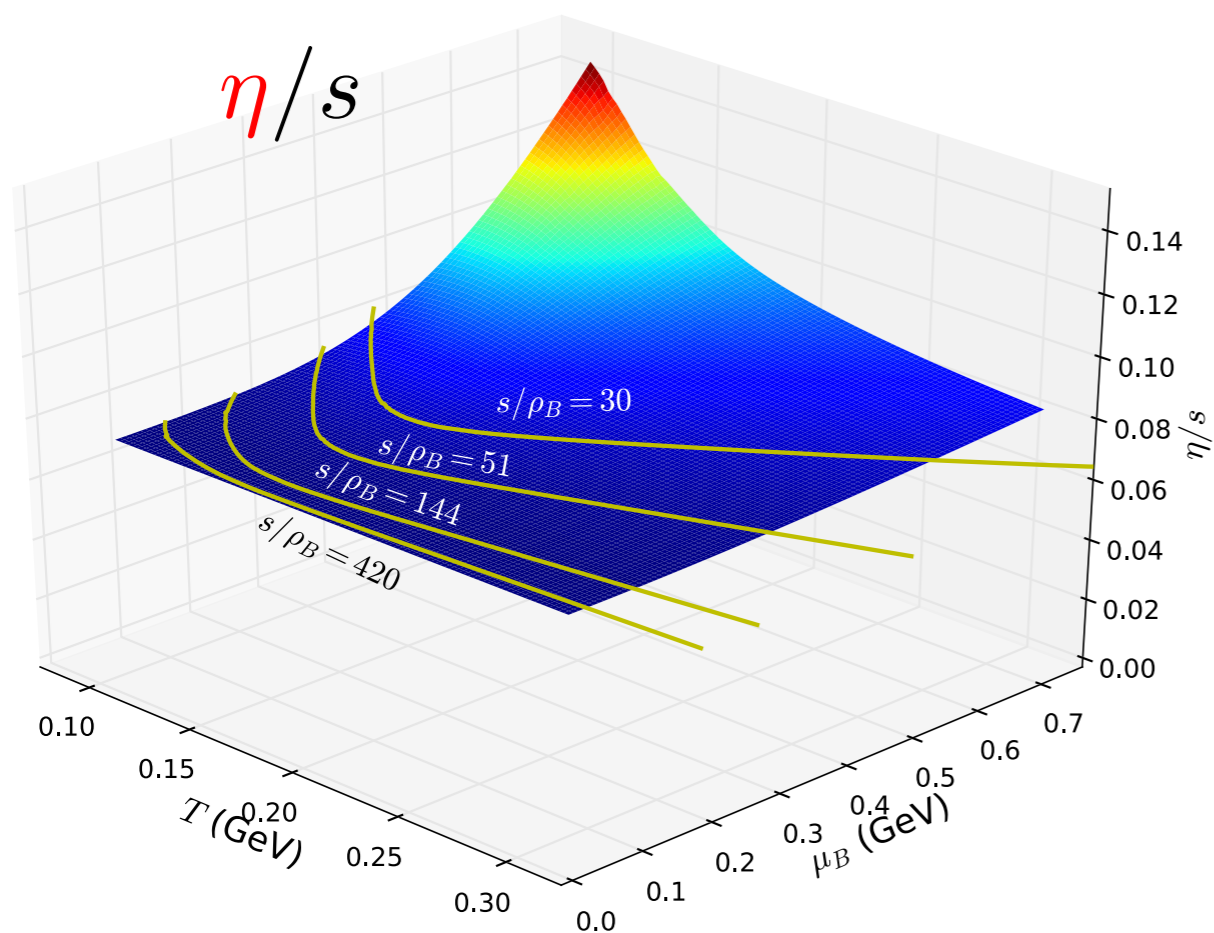
Dissipative part:

(relaxation time approximation)

$$\Delta_{\alpha\beta}^{\mu\nu} D\pi^{\alpha\beta} = -\frac{1}{\tau_\pi} (\pi^{\mu\nu} - 2\eta\sigma^{\mu\nu}) - \frac{\delta_{\pi\pi}}{\tau_\pi} \pi^{\mu\nu} \theta - \frac{\tau_{\pi\pi}}{\tau_\pi} \pi^\lambda \langle \mu \sigma^\nu \rangle_\lambda + \frac{\phi_7}{\tau_\pi} \pi_\alpha \langle \mu \pi^\nu \rangle_\alpha$$

$$\Delta^{\mu\nu} Dq_\nu = -\frac{1}{\tau_q} (q^\mu - \kappa \nabla^\mu \frac{\mu_B}{T}) - \frac{\delta_{qq}}{\tau_q} q^\mu \theta - \frac{\lambda_{qq}}{\tau_q} q_\nu \sigma^{\mu\nu}$$

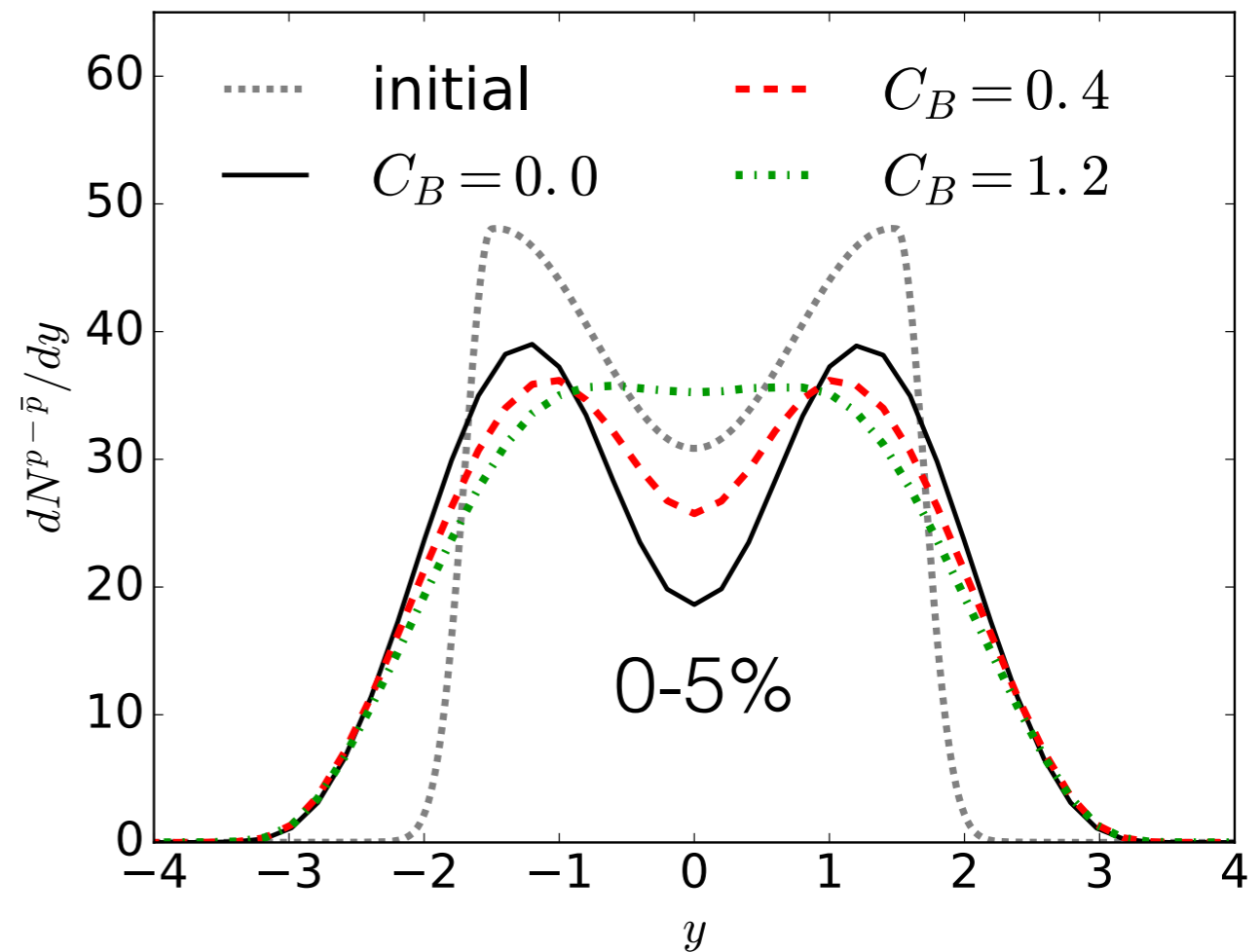
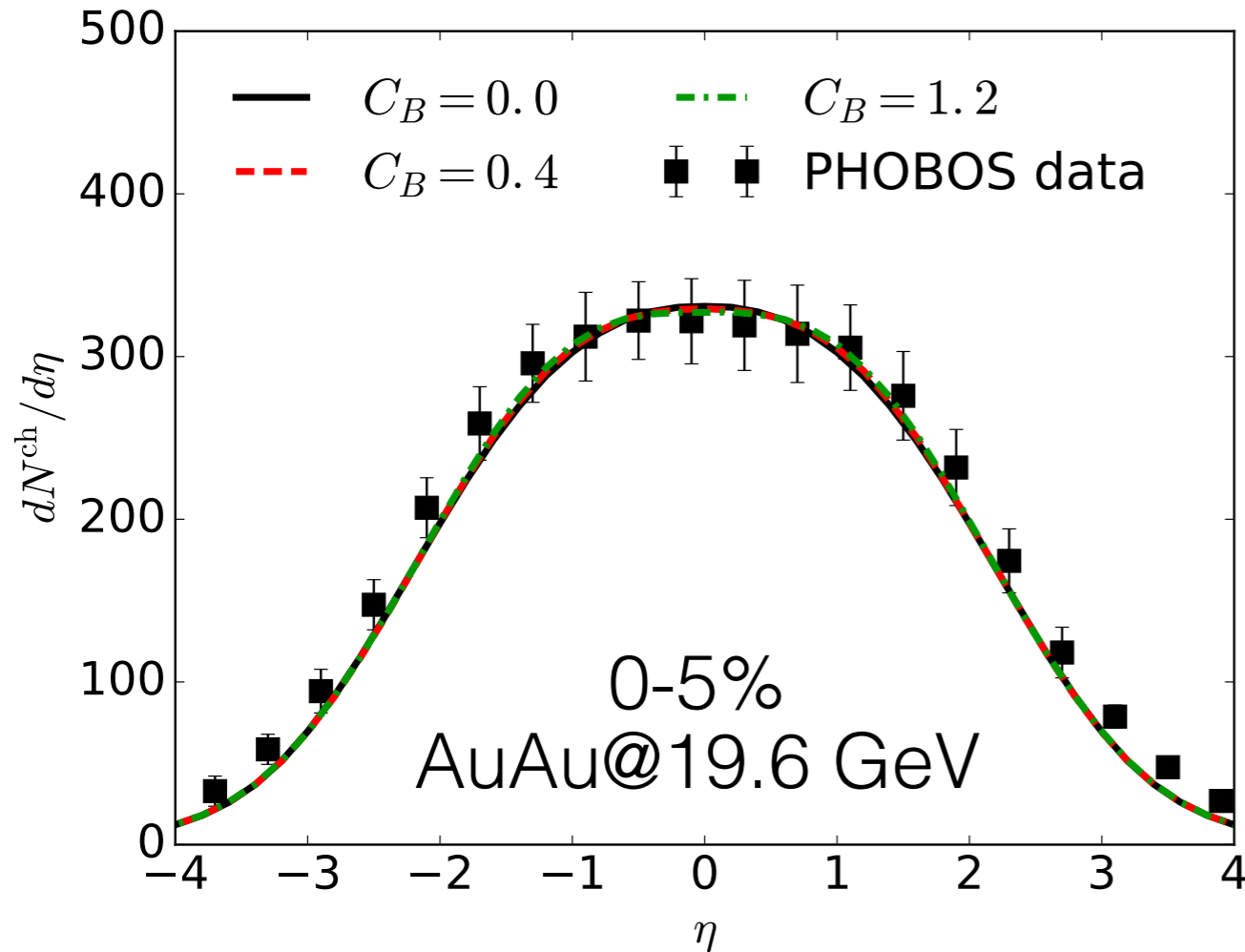
With non-zero μ , we choose $\tau_\pi = \tau_q = \frac{0.4}{T}$ $\frac{\eta T}{e + \mathcal{P}} = 0.08$



$$\kappa_B = \frac{C_B^{0,0}}{T} \rho_B \left(\frac{1}{3} \coth \left(\frac{\mu_B}{T} \right) - \frac{\rho_B T}{e + P} \right)$$

Effects of net baryon diffusion on particle yields

C. Shen, G. Denicol, C. Gale, S. Jeon, A. Monnai, B. Schenke, in preparation

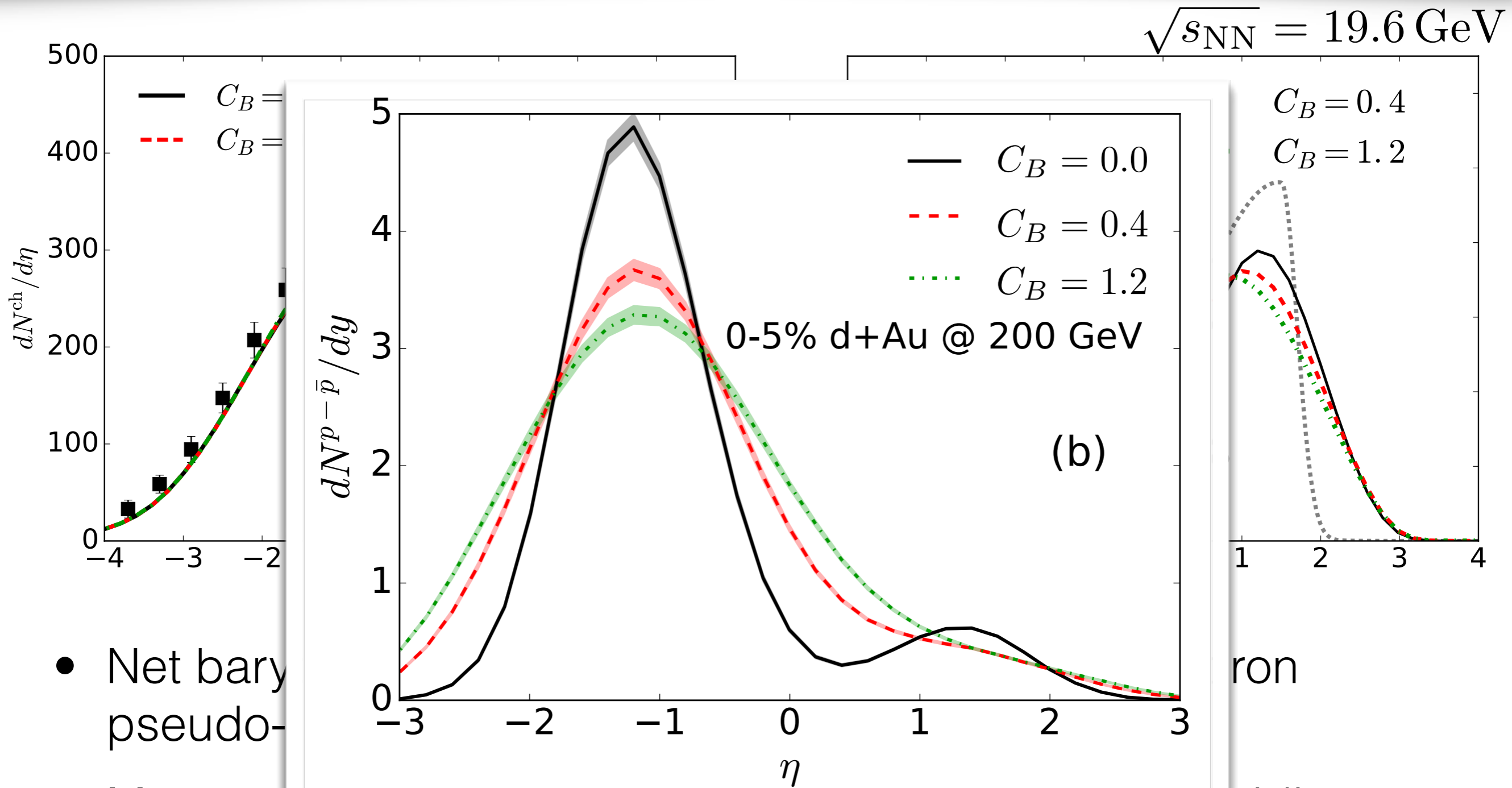


$$\kappa_B = \frac{C_B}{T} \rho_B \left(\frac{1}{3} \coth \left(\frac{\mu_B}{T} \right) - \frac{\rho_B T}{e + P} \right)$$

- More net baryon numbers are transported to mid-rapidity with a larger diffusion constant

Constraints on net baryon diffusion and initial condition

Effects of net baryon diffusion on particle yields

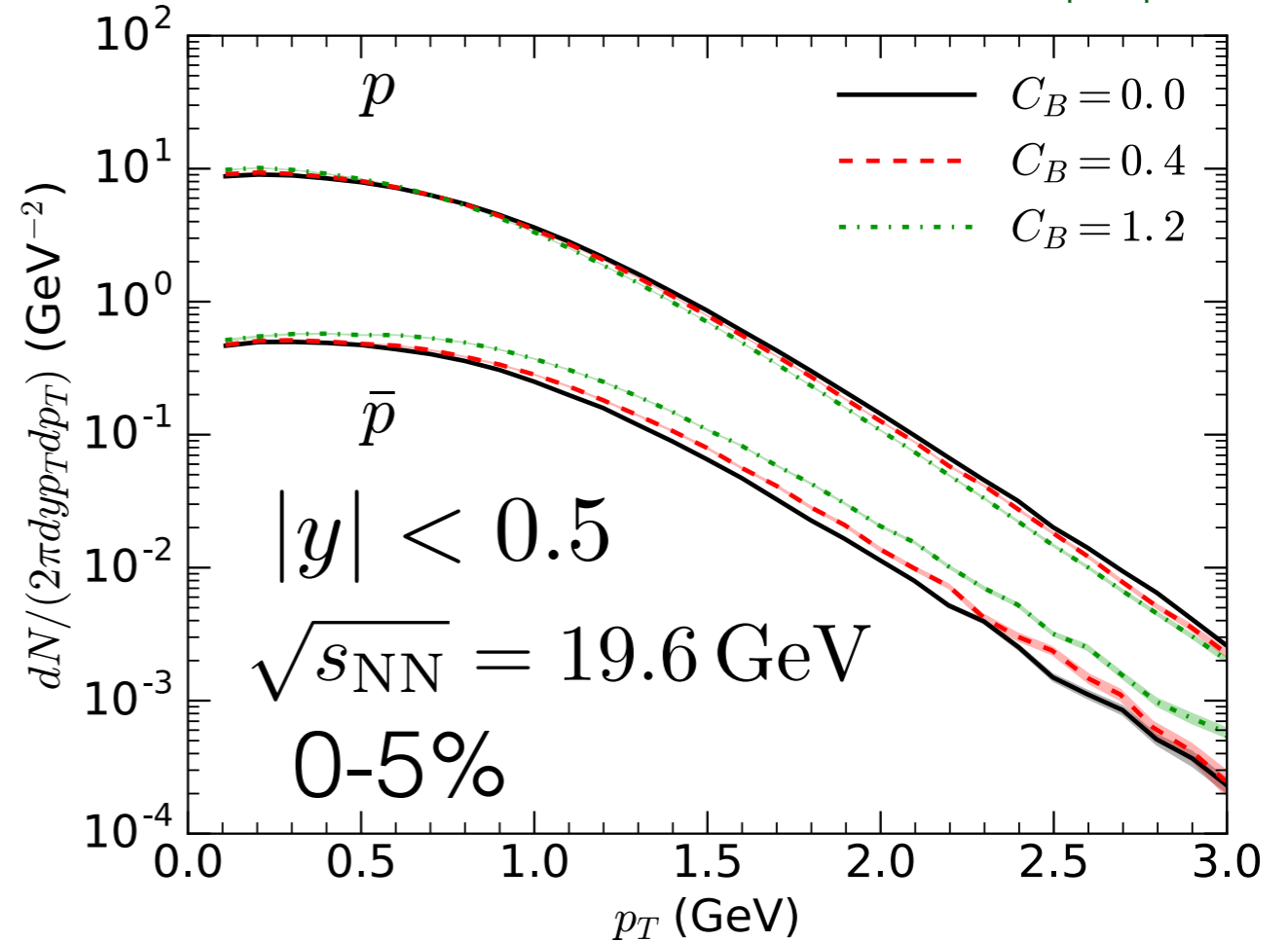
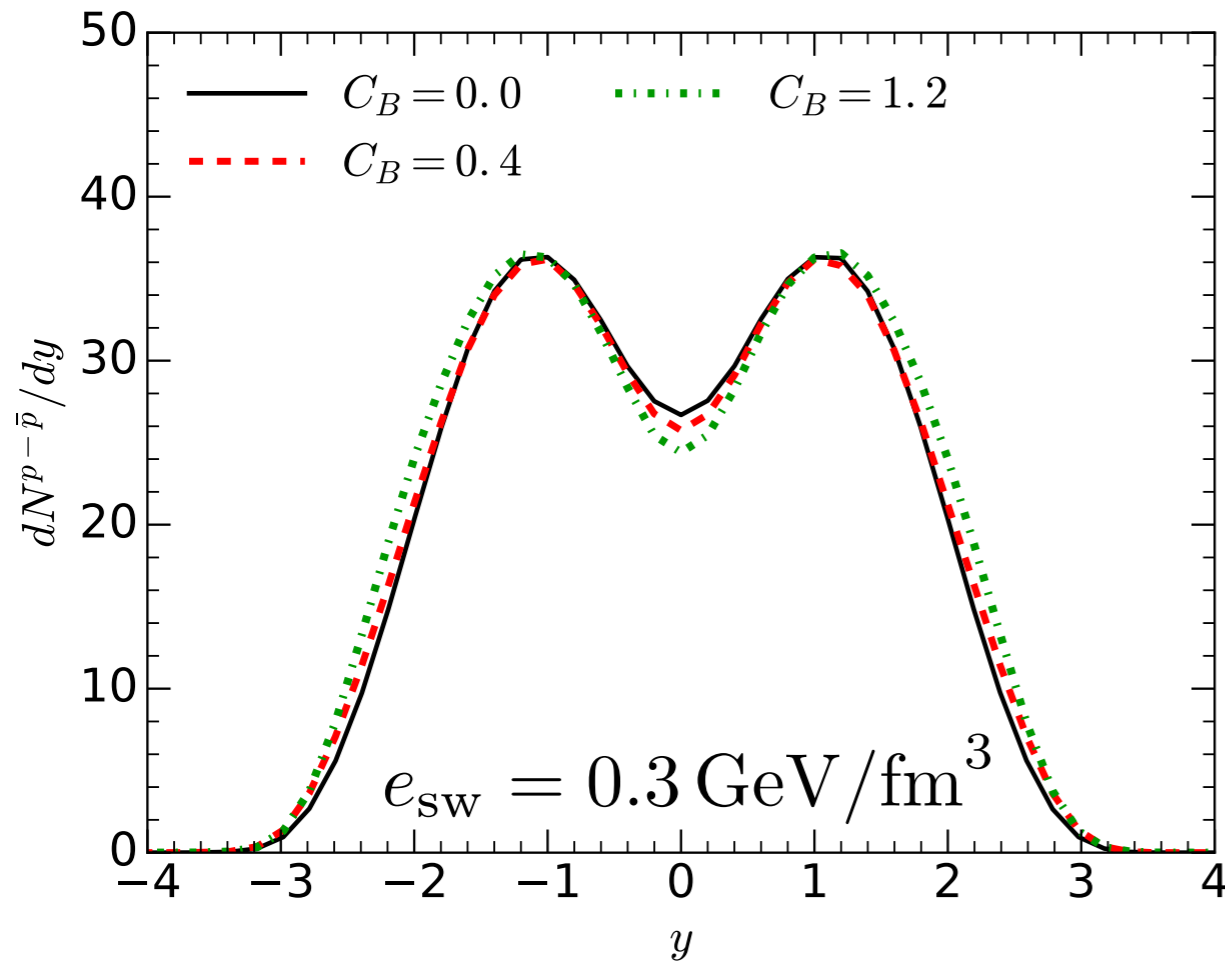


- Net baryon pseudo-rapidity distribution
- More net baryon numbers are transported to mid-rapidity with a larger diffusion constant

Constraints on net baryon diffusion and initial condition

Effects of net baryon diffusion on pid spectra

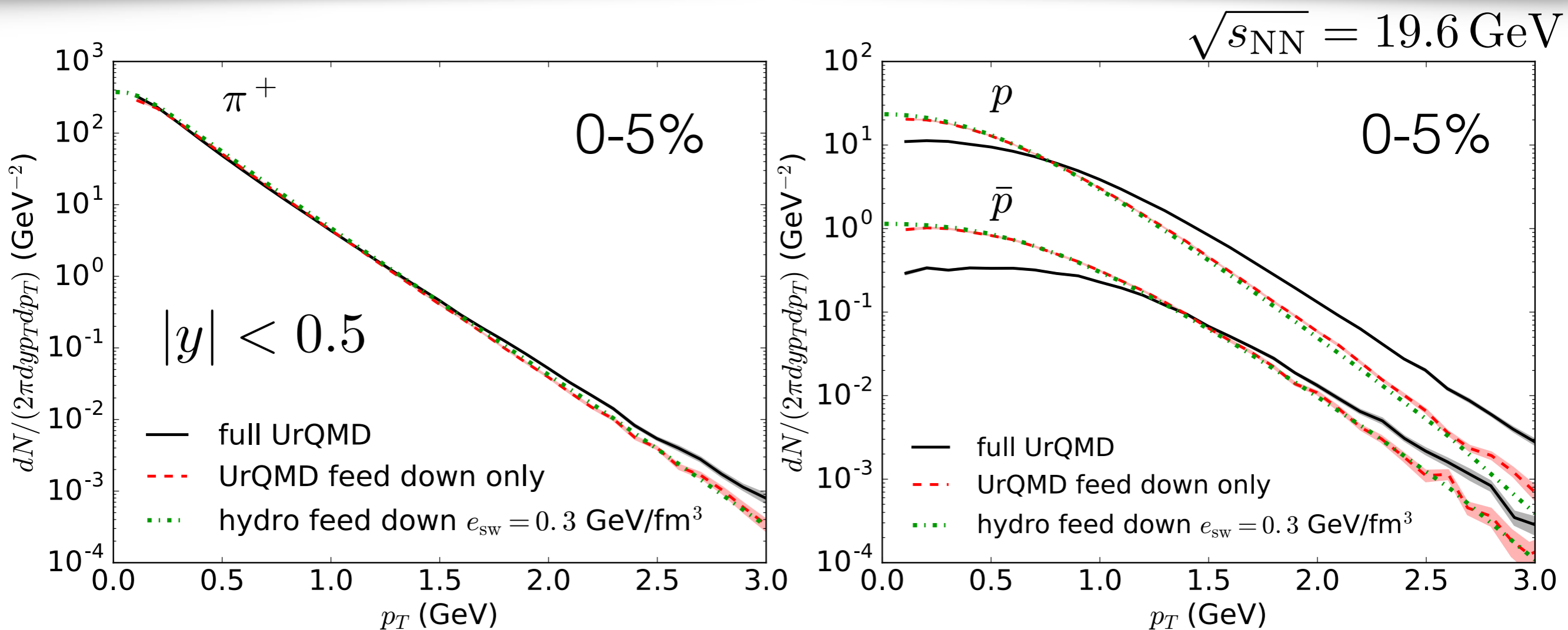
C. Shen, G. Denicol, C. Gale, S. Jeon, A. Monnai, B. Schenke, in preparation



- Net baryon diffusion results a flatter spectra for anti-proton compared to proton's

	$C_B = 0.0$	$C_B = 0.4$	$C_B = 1.2$
$\langle p_{\perp} \rangle^{\bar{p}} - \langle p_{\perp} \rangle^p$ (GeV)	0.046	0.091	0.158

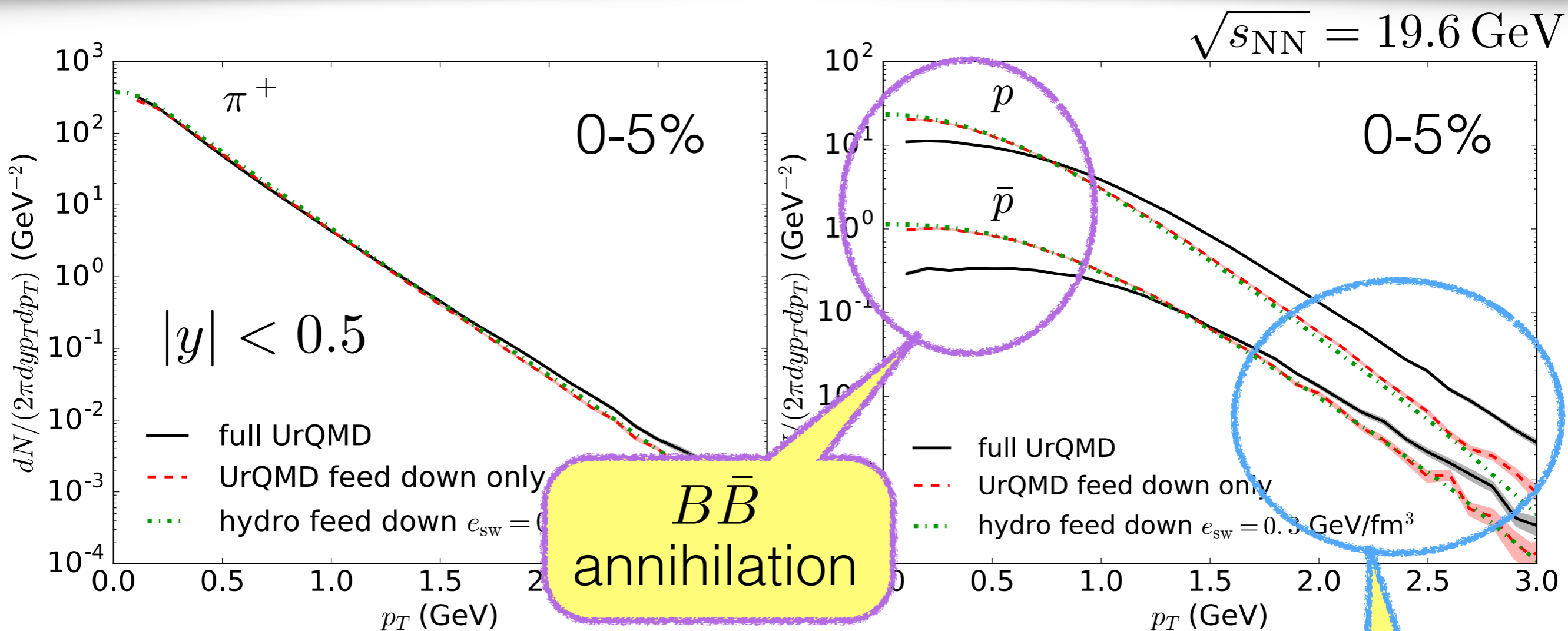
Effects of hadronic afterburner on pid spectra



- Hadronic afterburner harden pion spectra at high p_T
- Heavy baryon spectra are largely affected

hadronic afterburner is essential for baryon spectra

Effects of hadronic afterburner on pid spectra



- Hadronic afterburner harden pion spectra and baryon spectra
- Heavy baryon spectra are largely affected

hadronic rescatterings

hadronic afterburner is essential for baryon spectra

Conclusion

- We develop a **dynamical initialization** model to study the early time evolution of heavy-ion collisions at the BES energies
 - full **(3+1)-d** event-by-event with **net baryon current**
- With a hybrid approach, we identified a few experiment observables that could constrain the **net baryon diffusion**

$$dN^{p-\bar{p}}/dy \quad \langle p_{\perp} \rangle^{\bar{p}} - \langle p_{\perp} \rangle^p$$

- Future combining with the **Bayesian analysis** will help us to constrain the initial state and transport coefficients of the QGP in a baryon-rich environment



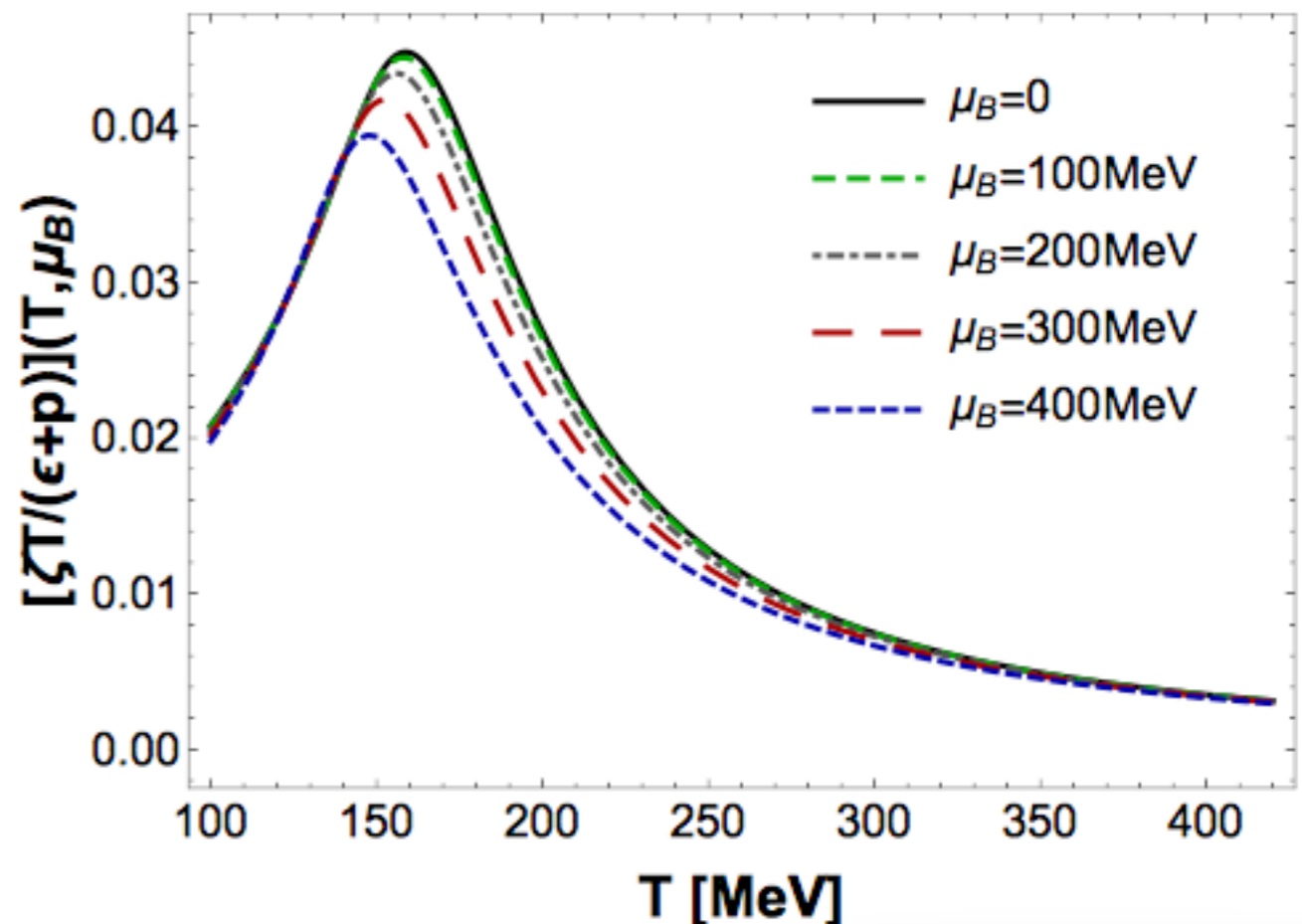
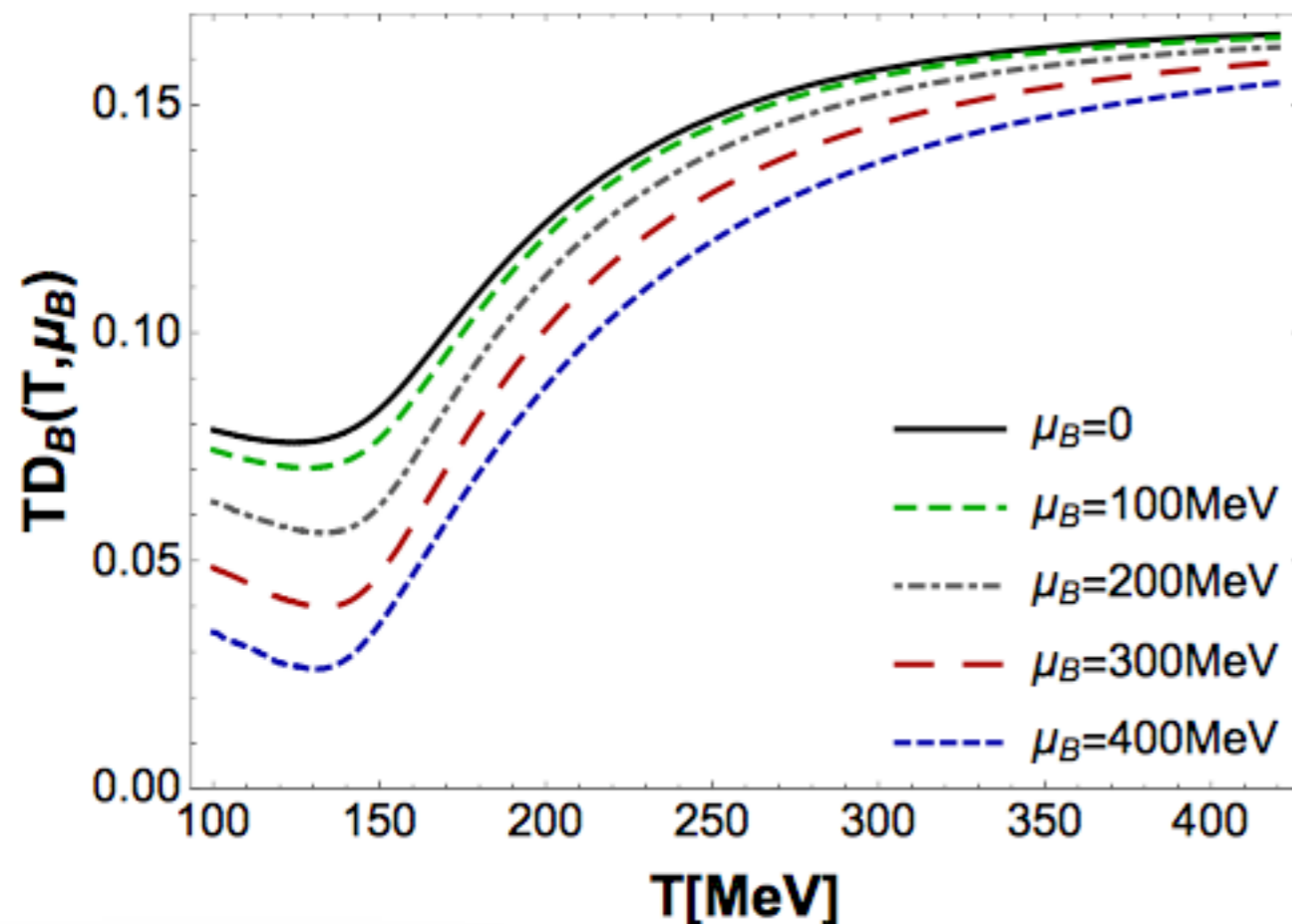
BACKUP

Transport coefficients

R. Rougemont, R. Critelli, J. Noronha-Hostler, J. Noronha and C. Ratti, Phys. Rev. D 96, 014032 (2017)

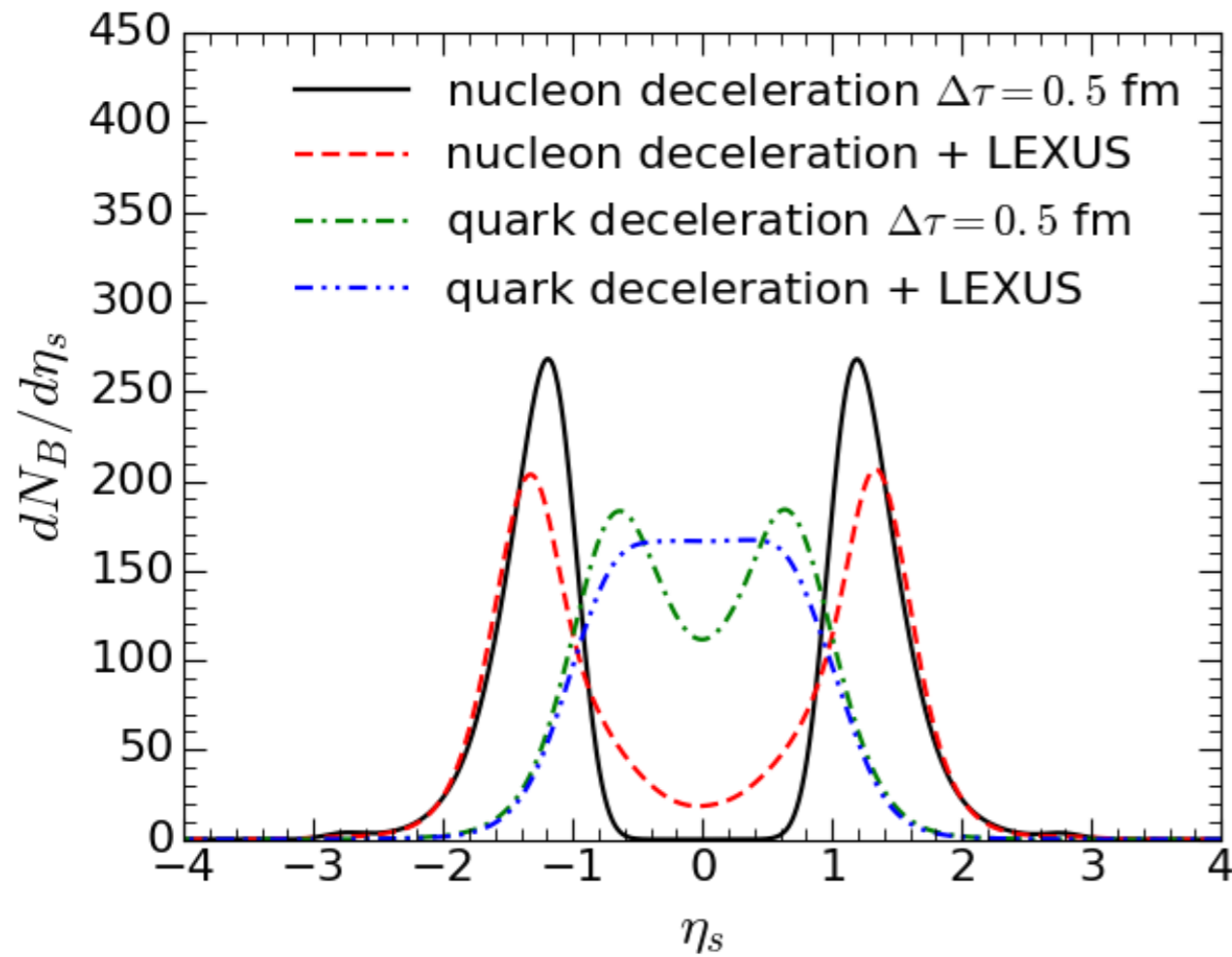
The holographic Einstein-Maxwell-Dilation (EMD) model is fit to the lattice results on thermodynamic quantities at $\mu_B = 0$

Predictions are made for thermodynamic variables at finite μ_B and for the temperature and μ_B dependence of various transport coefficients

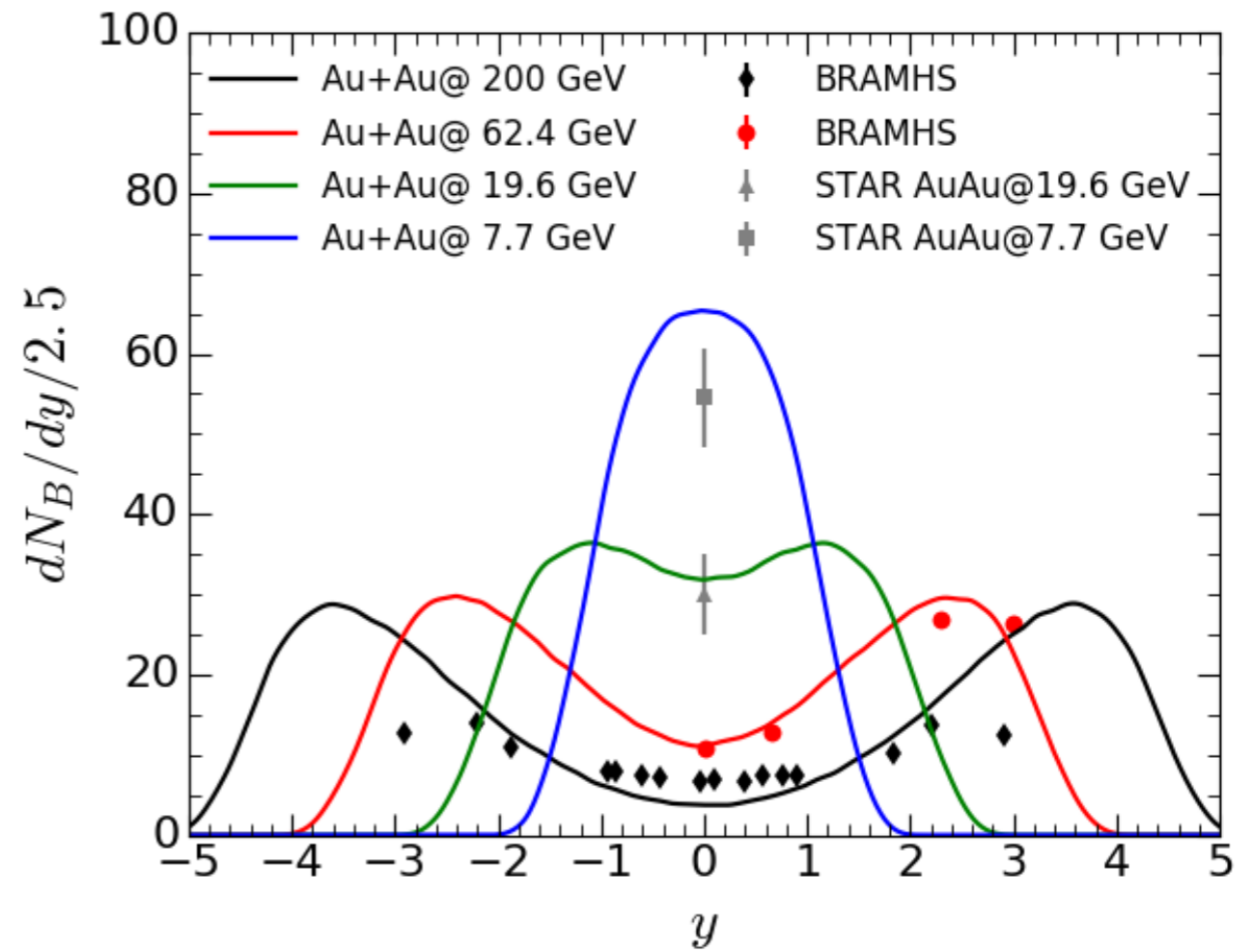


Net baryon number distribution

space-time rapidity

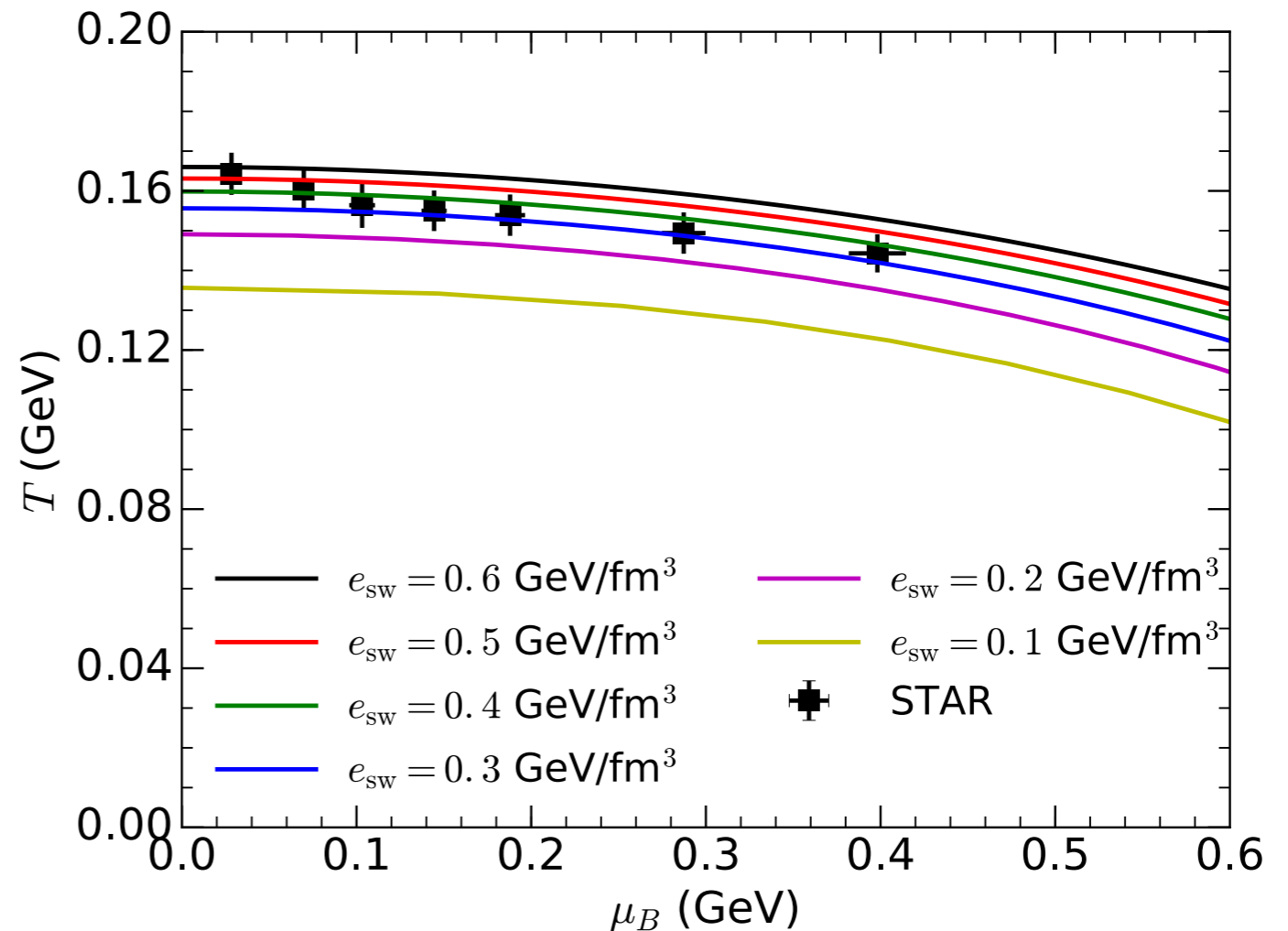
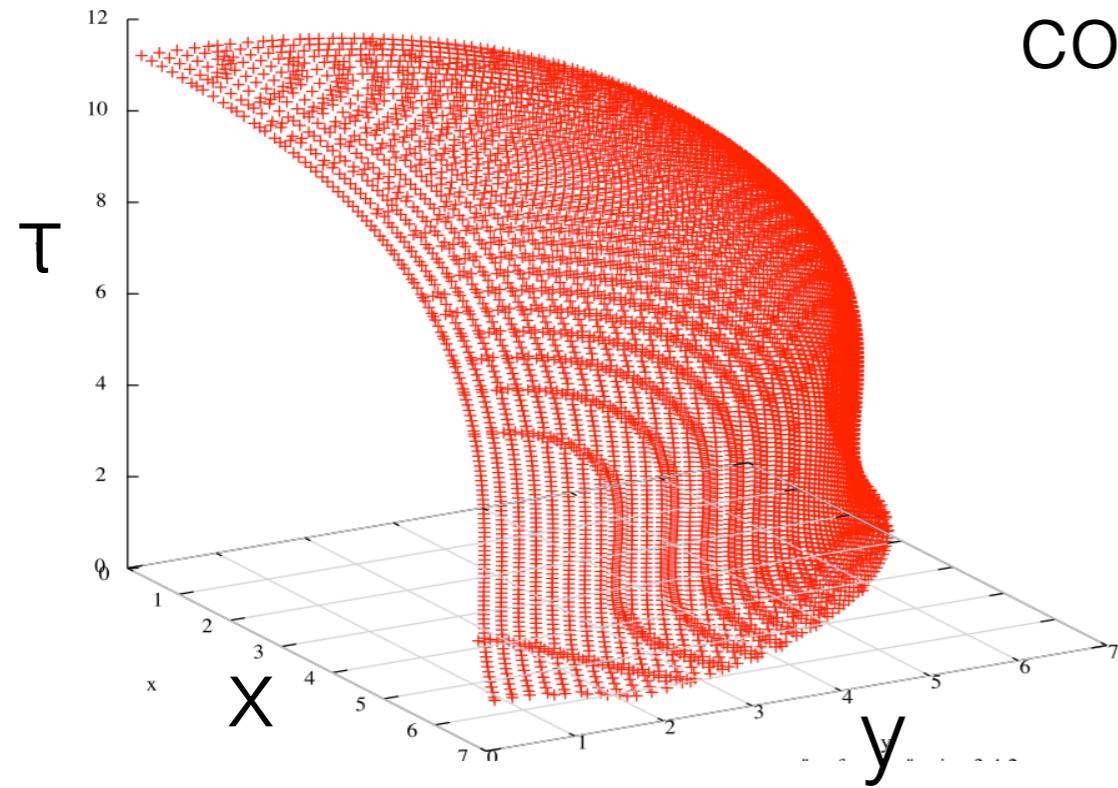


momentum rapidity



Convert to particles

Conversion surface is determined at a constant energy density



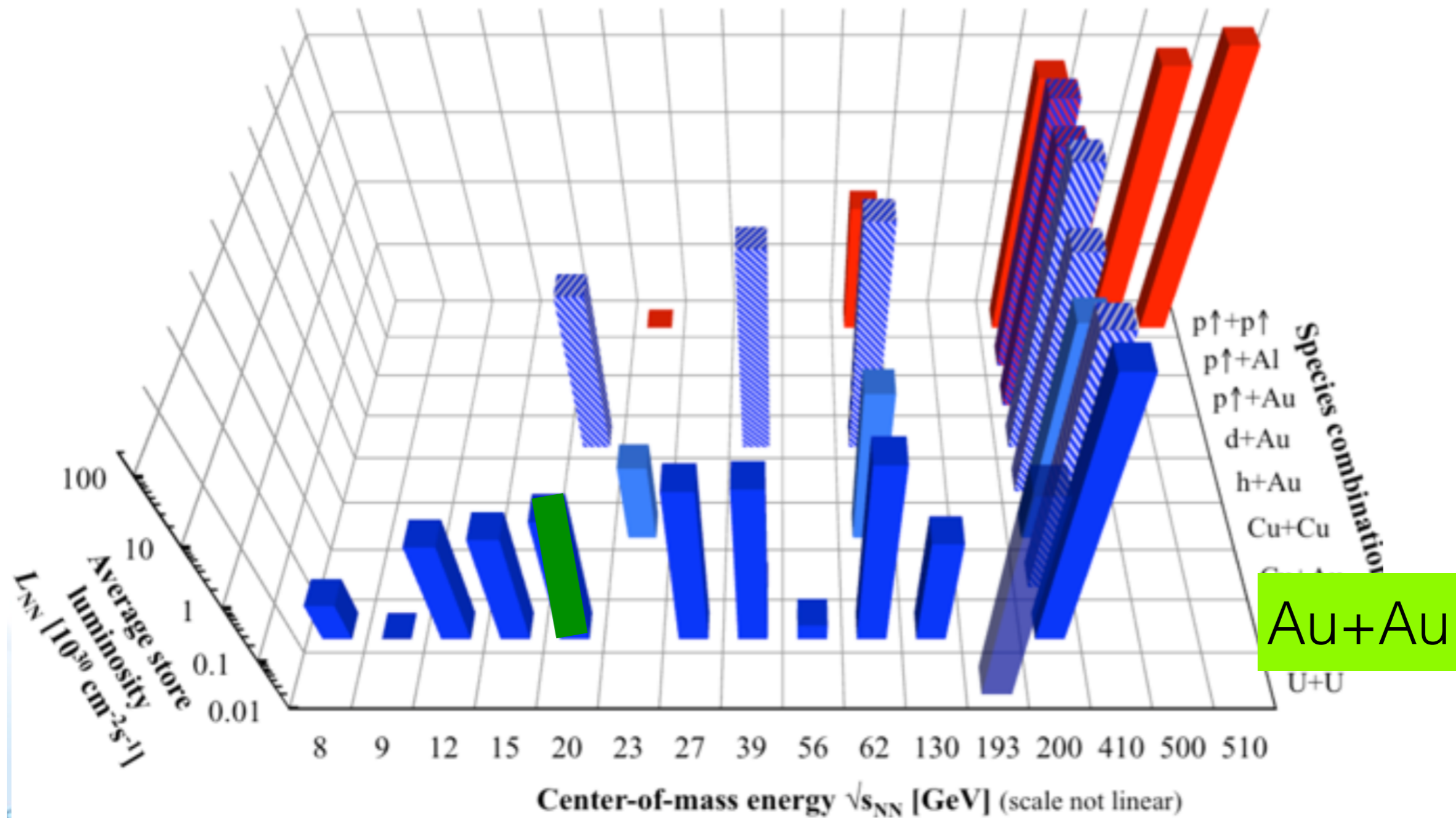
Cooper-Frye conversion:

$$E \frac{dN_i}{d^3p} = \frac{g_i}{(2\pi)^3} \int p^\mu d^3\sigma_\mu(x) (f_0(x, p) + \delta f(x, p))$$

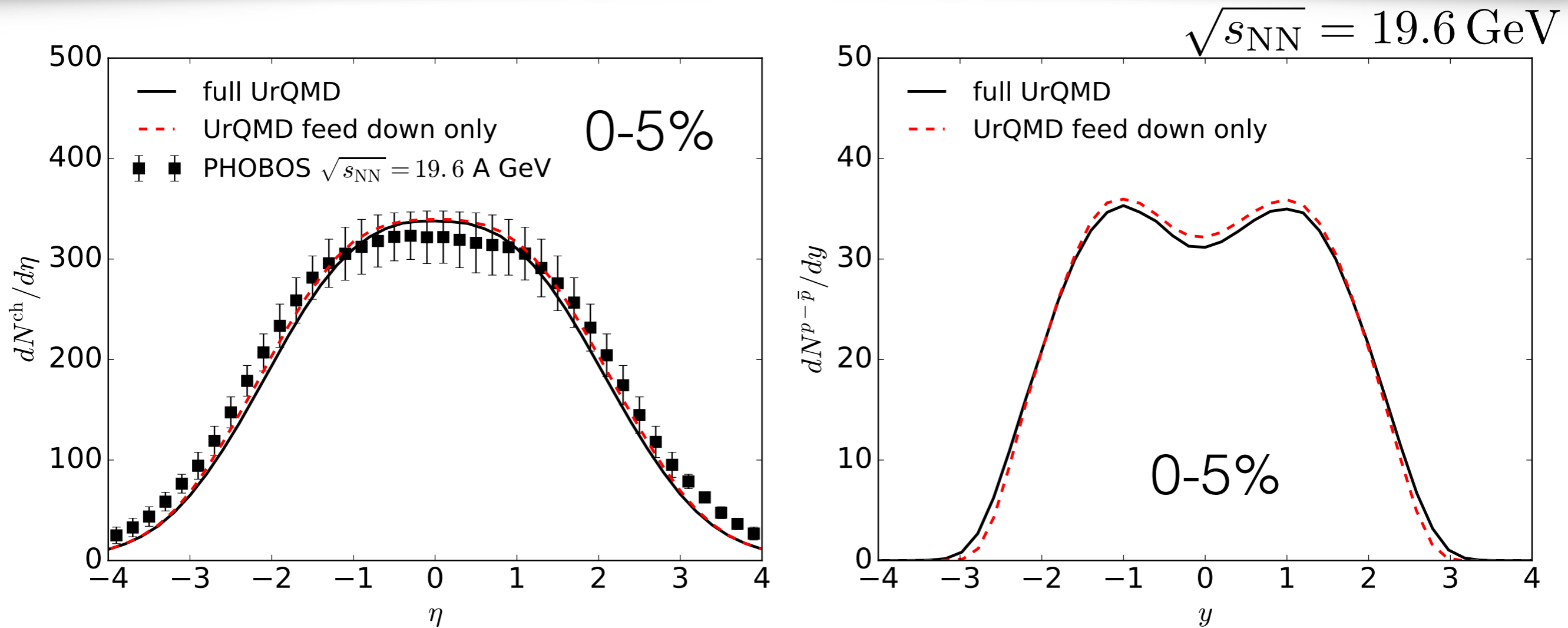
$$f_0^i(x, p) = \frac{1}{e^{(E - b_i \mu_B(x))/T(x)} \pm 1}$$

Heavy-ion collisions at BES energies...

RHIC energies, species combinations and luminosities (Run-1 to 16)



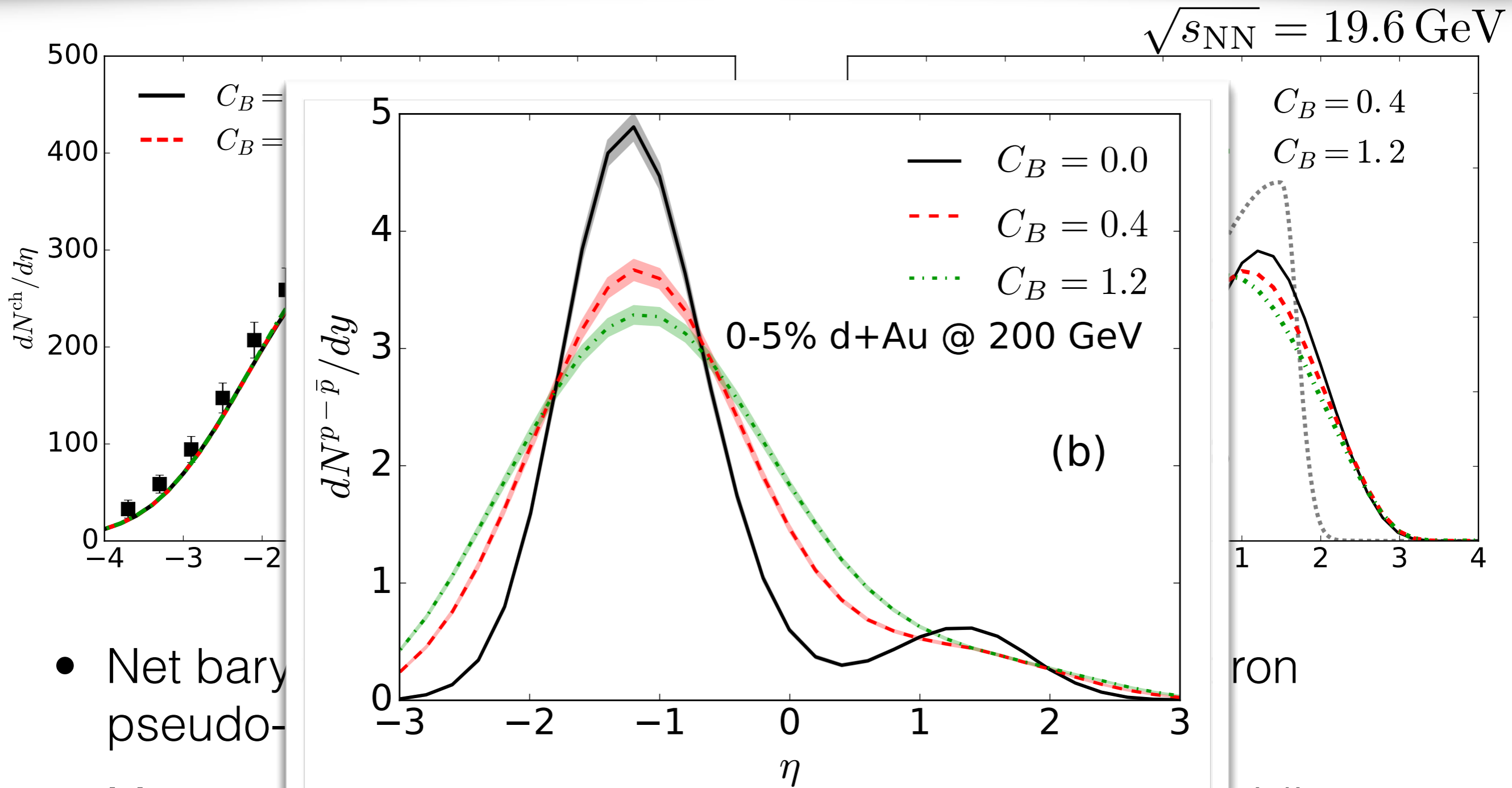
Effects of hadronic afterburner on particle yields



- Hadronic afterburner has little effect on charged hadron pseudo-rapidity distribution
- Net proton rapidity profile is slightly flatter after hadronic scatterings

more sensitive to early stage dynamics

Effects of net baryon diffusion on particle yields

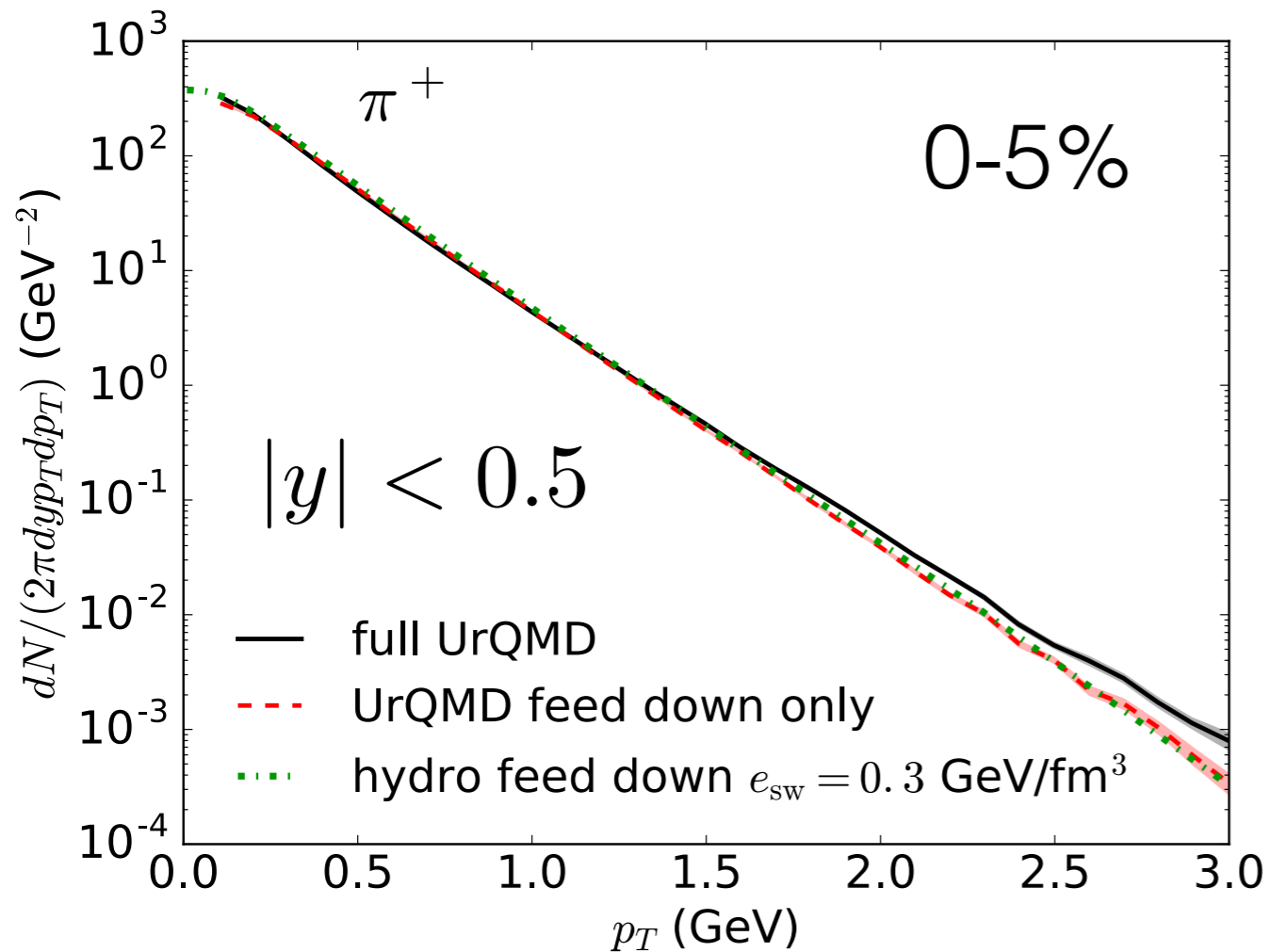


- Net baryon pseudorapidity
- More net baryon numbers are transported to mid-rapidity with a larger diffusion constant

Constraints on net baryon diffusion and initial condition

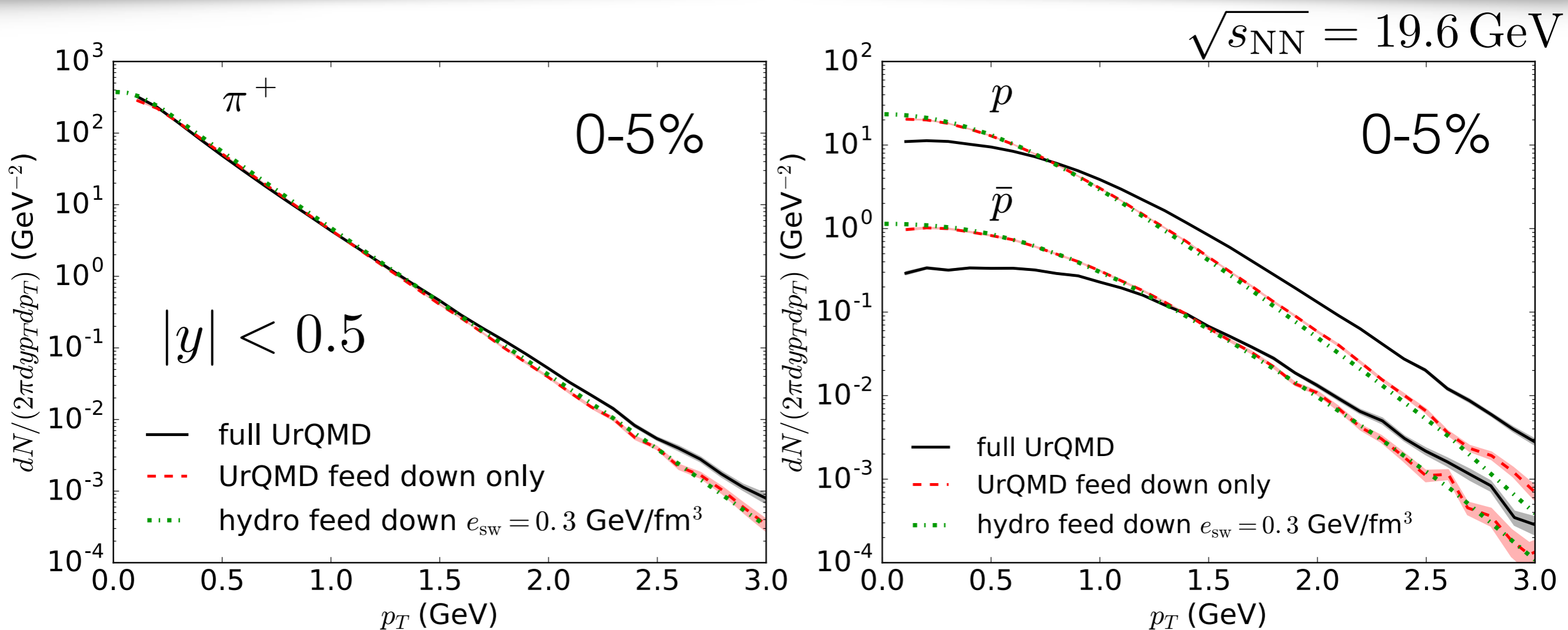
Effects of hadronic afterburner on pid spectra

$$\sqrt{s_{NN}} = 19.6 \text{ GeV}$$



- Hadronic afterburner harden pion spectra at high p_T

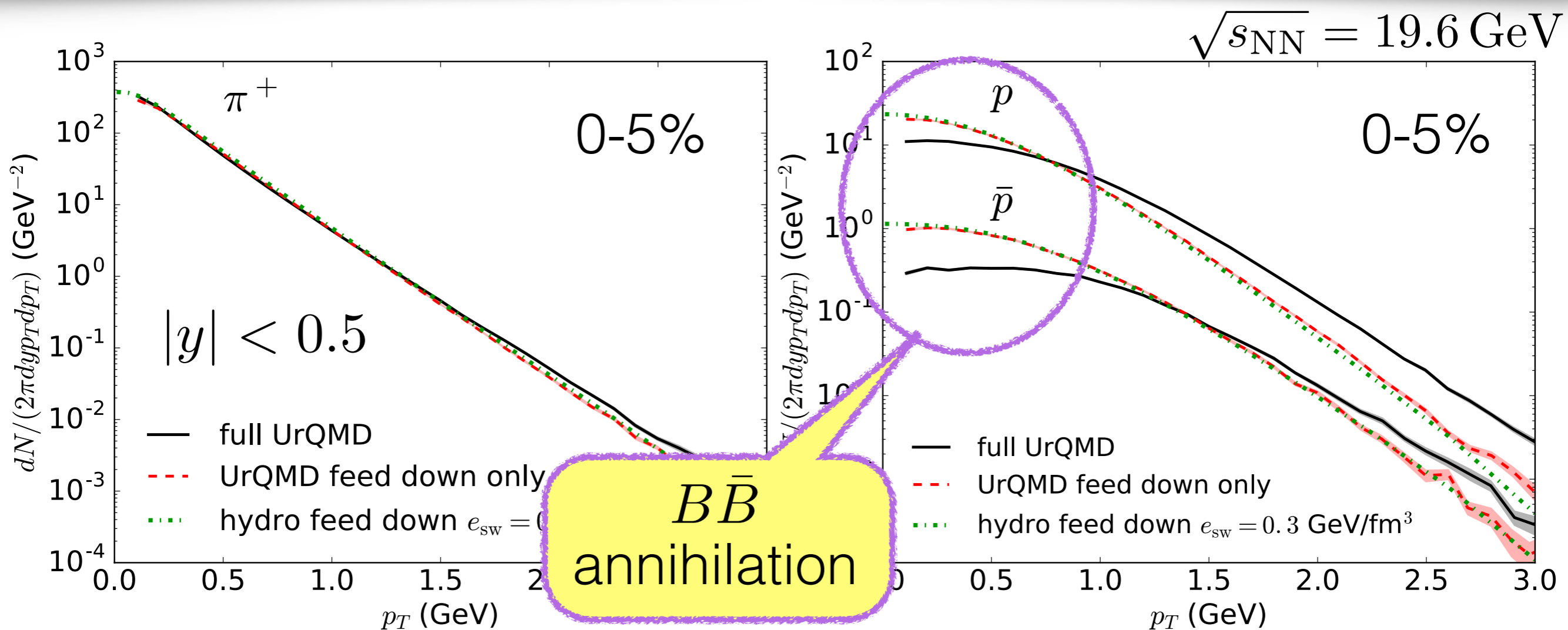
Effects of hadronic afterburner on pid spectra



- Hadronic afterburner harden pion spectra at high p_T
- Heavy baryon spectra are largely affected

hadronic afterburner is essential for baryon spectra

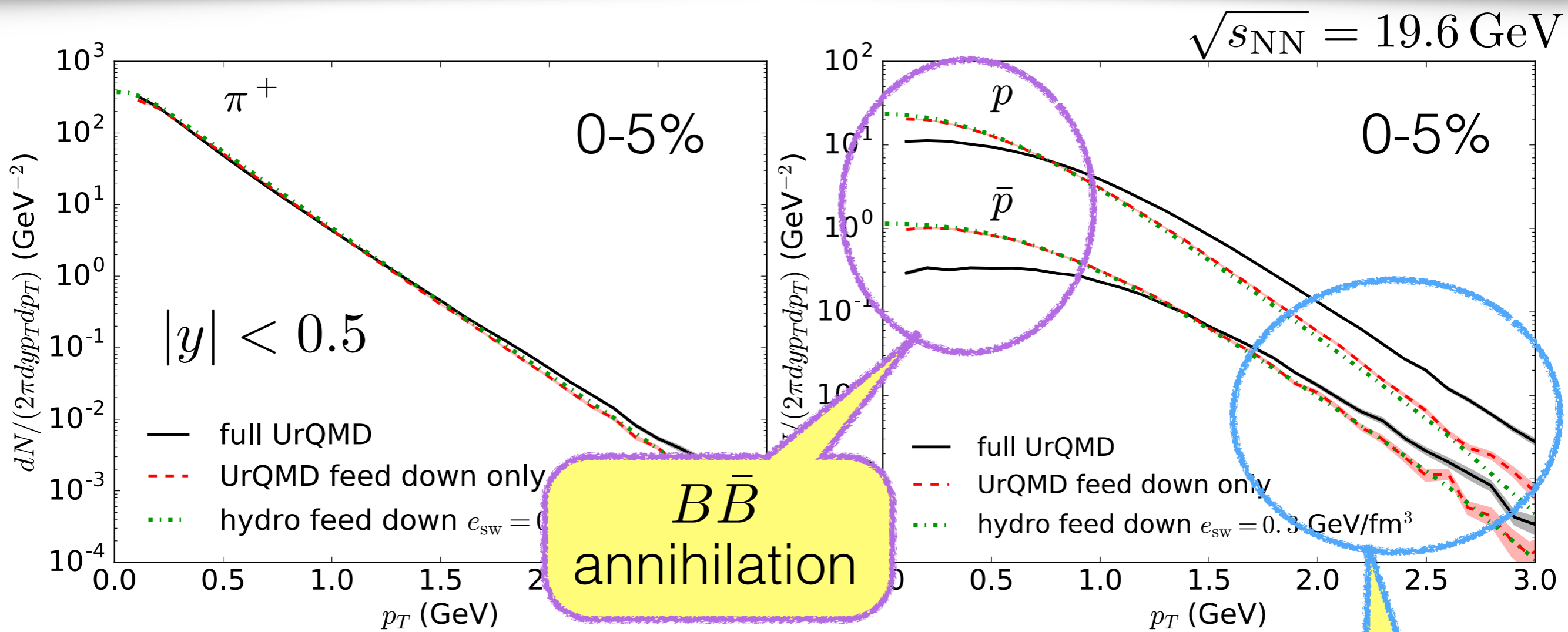
Effects of hadronic afterburner on pid spectra



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Effects of hadronic afterburner on pid spectra

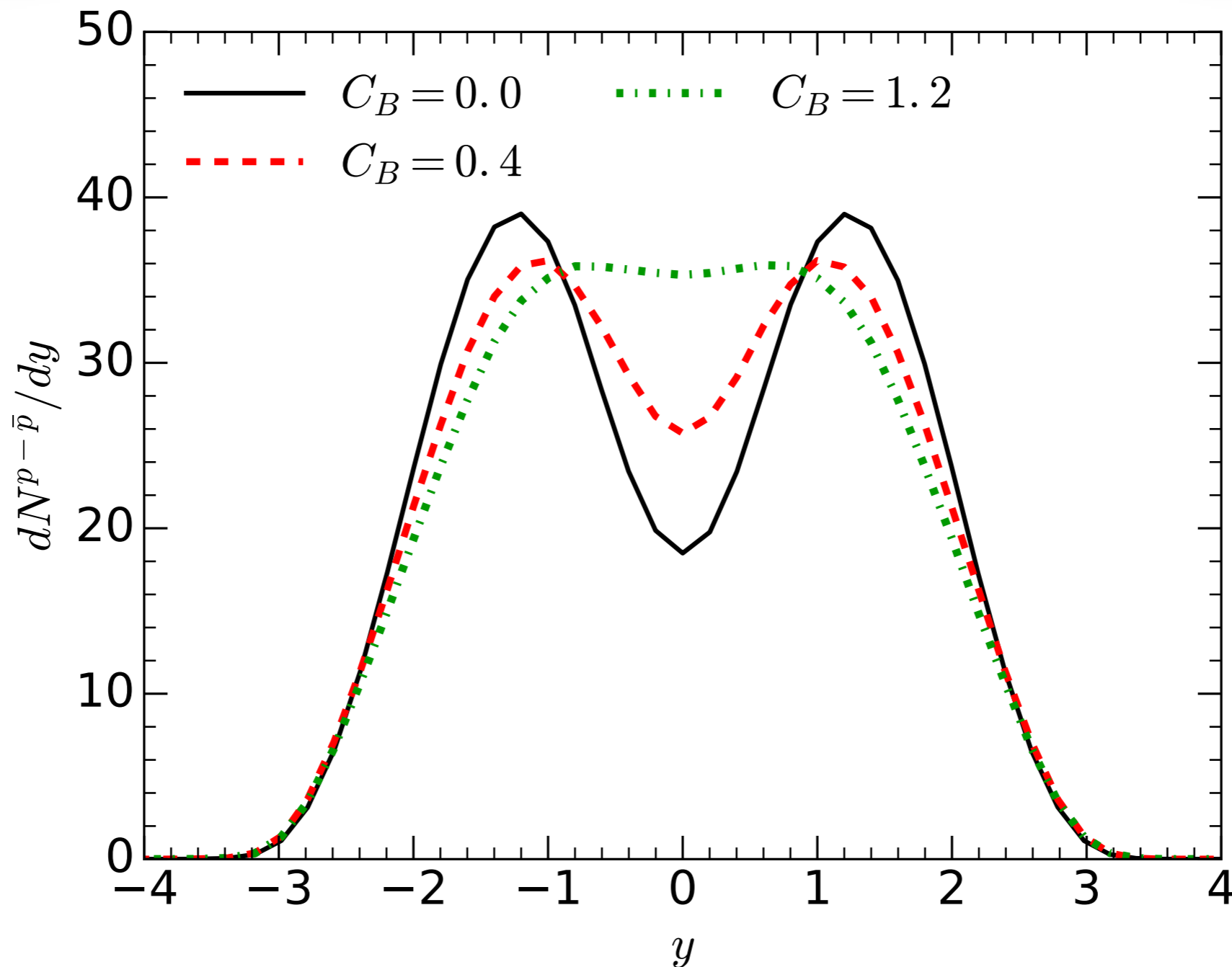


- Hadronic afterburner harden pion spectra and baryon spectra
- Heavy baryon spectra are largely affected

hadronic rescatterings

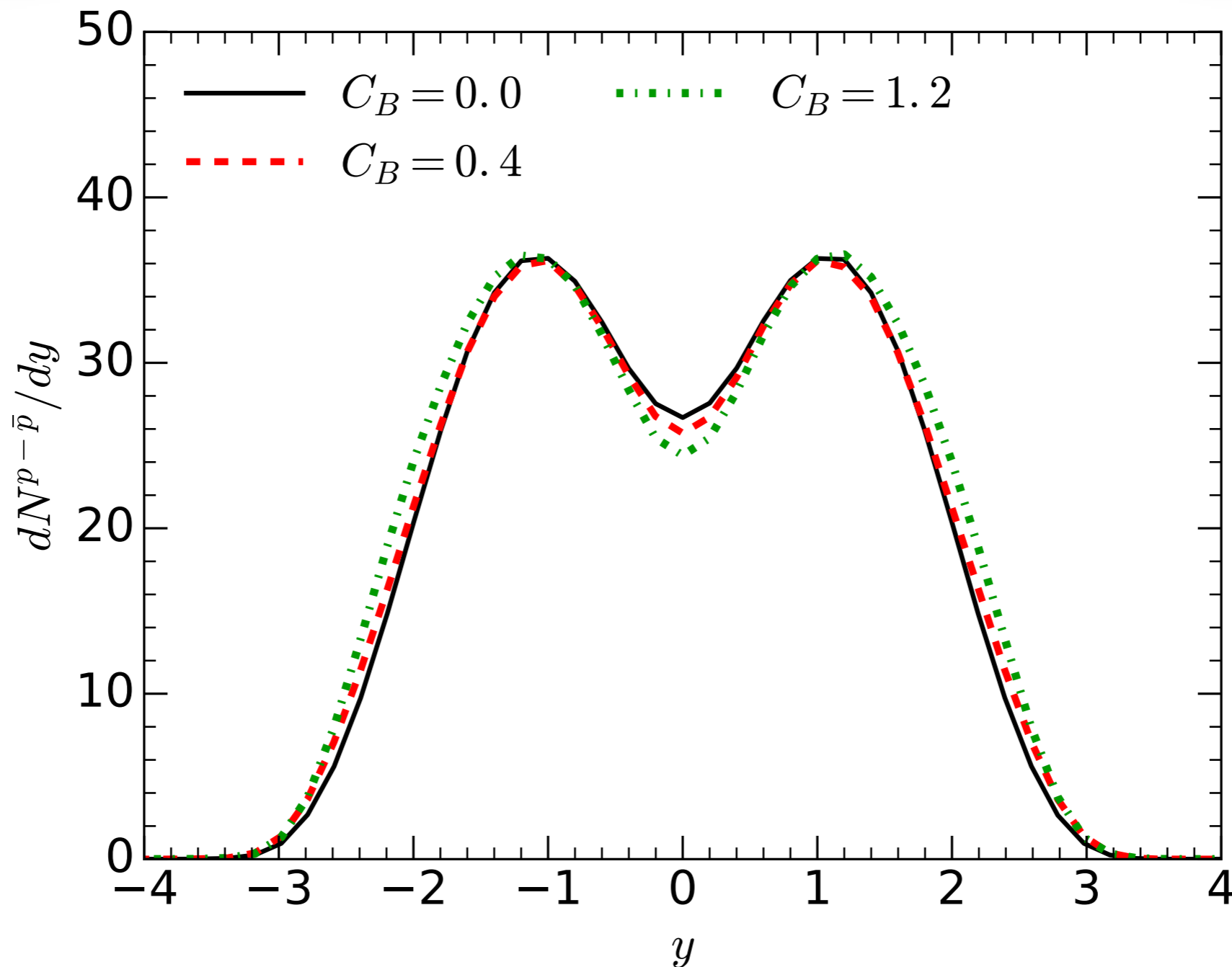
hadronic afterburner is essential for baryon spectra

Effects of net baryon diffusion on pid spectra



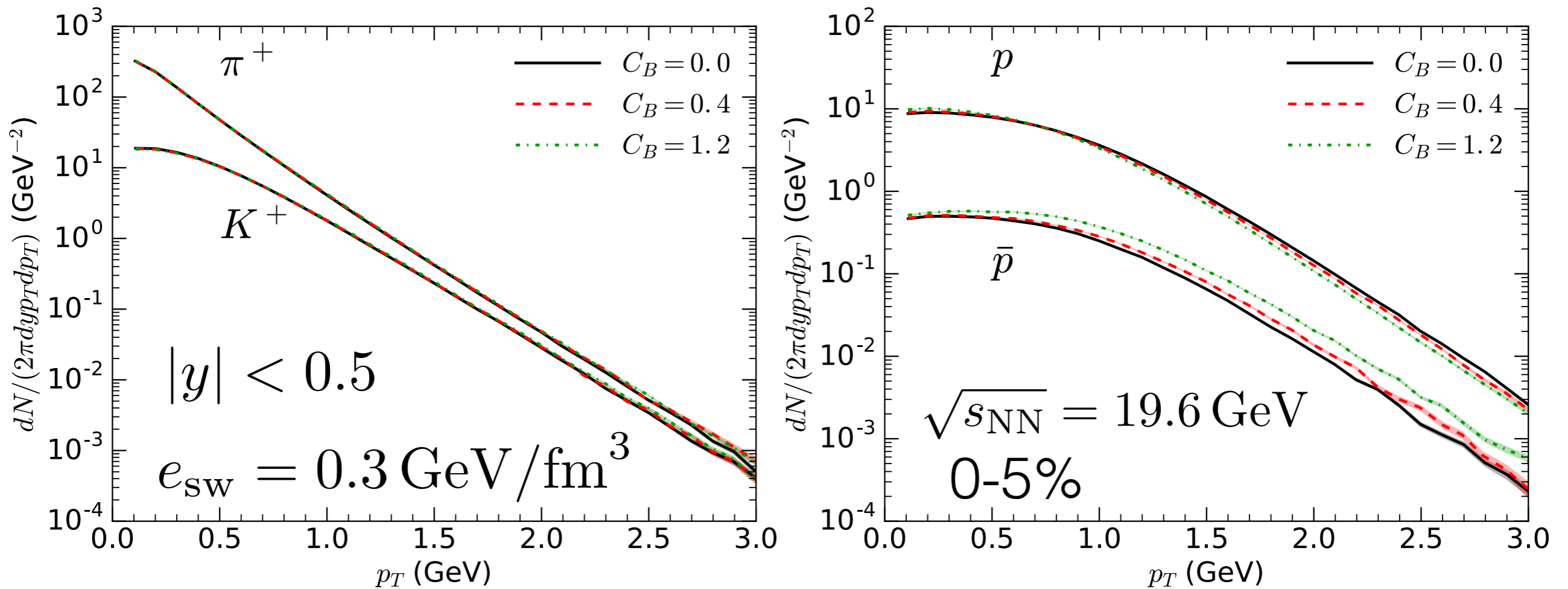
- To study the net baryon diffusion effects on transverse observables, we first tune the initial longitudinal profiles to produce the same $dN^{p-\bar{p}}/dy$

Effects of net baryon diffusion on pid spectra



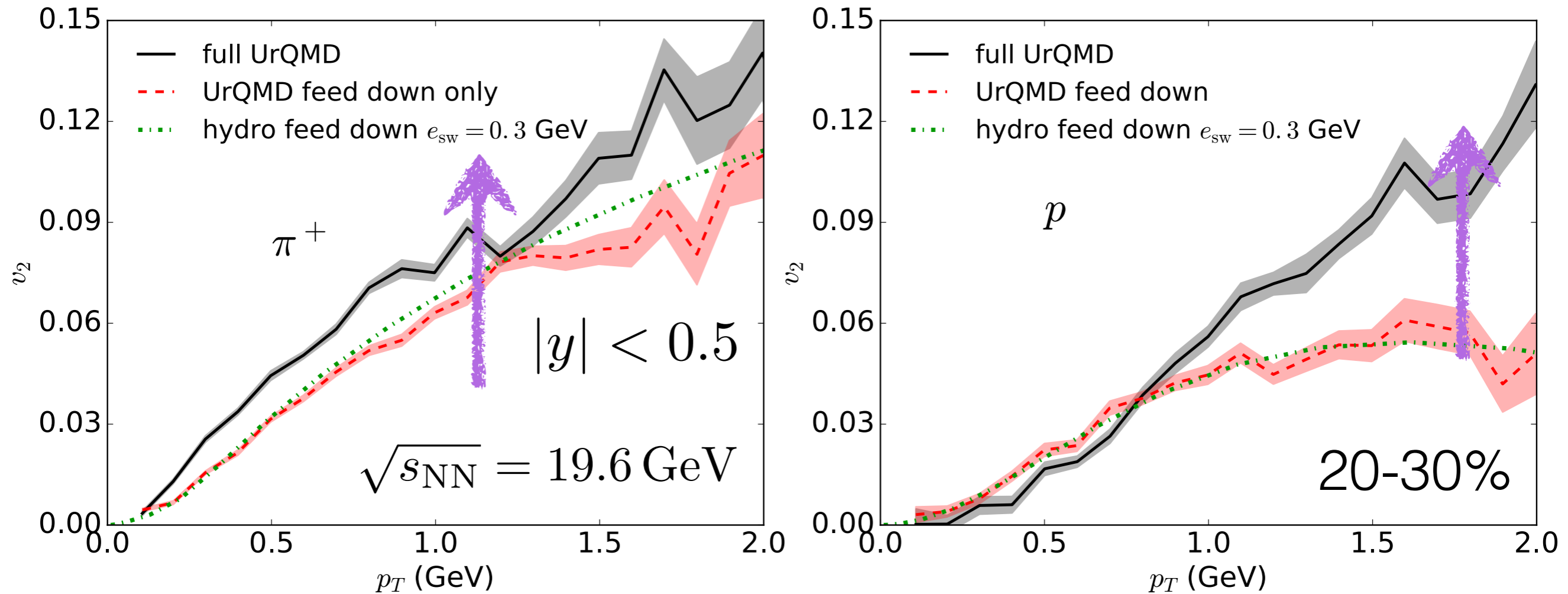
- To study the net baryon diffusion effects on transverse observables, we first tune the initial longitudinal profiles to produce the same $dN^{p-\bar{p}}/dy$

Effects of net baryon diffusion on pid spectra



- Net baryon diffusion results a flatter spectra for anti-proton compared to proton's

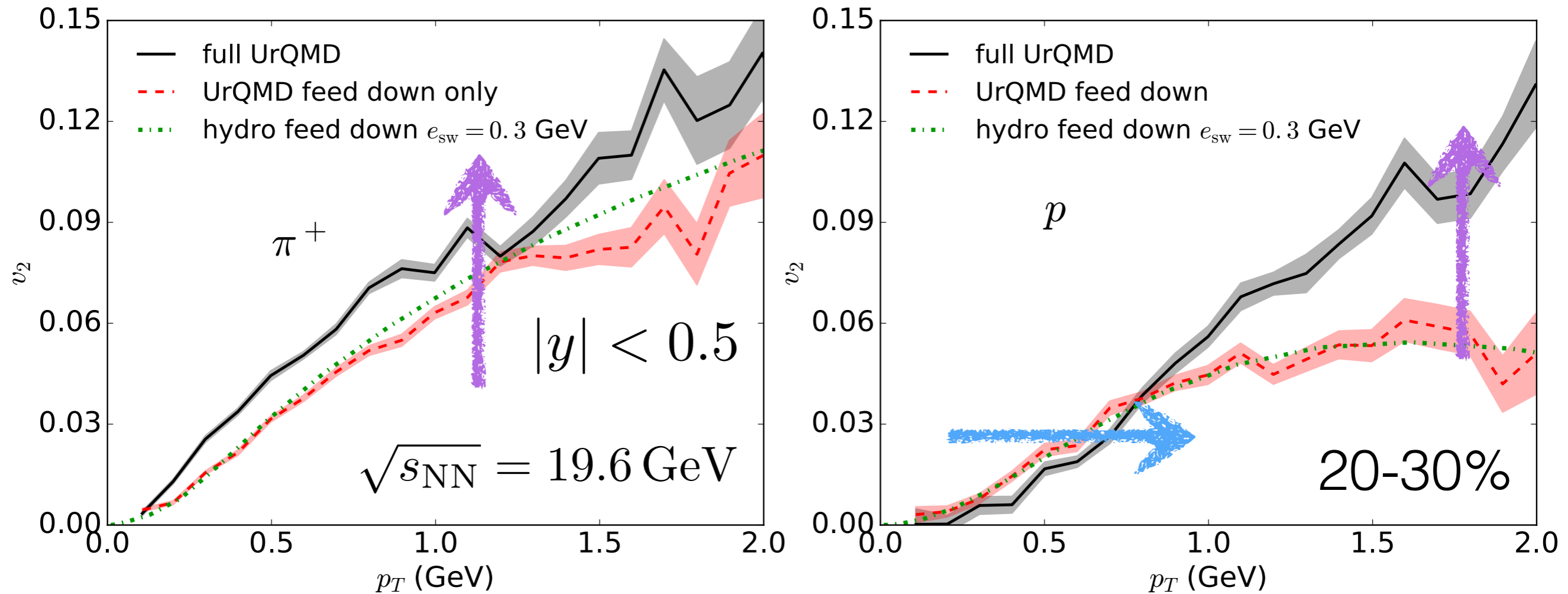
Effects of hadronic afterburner on pid v_2



- Momentum anisotropy keeps developing in the UrQMD phase

hadronic afterburner is essential

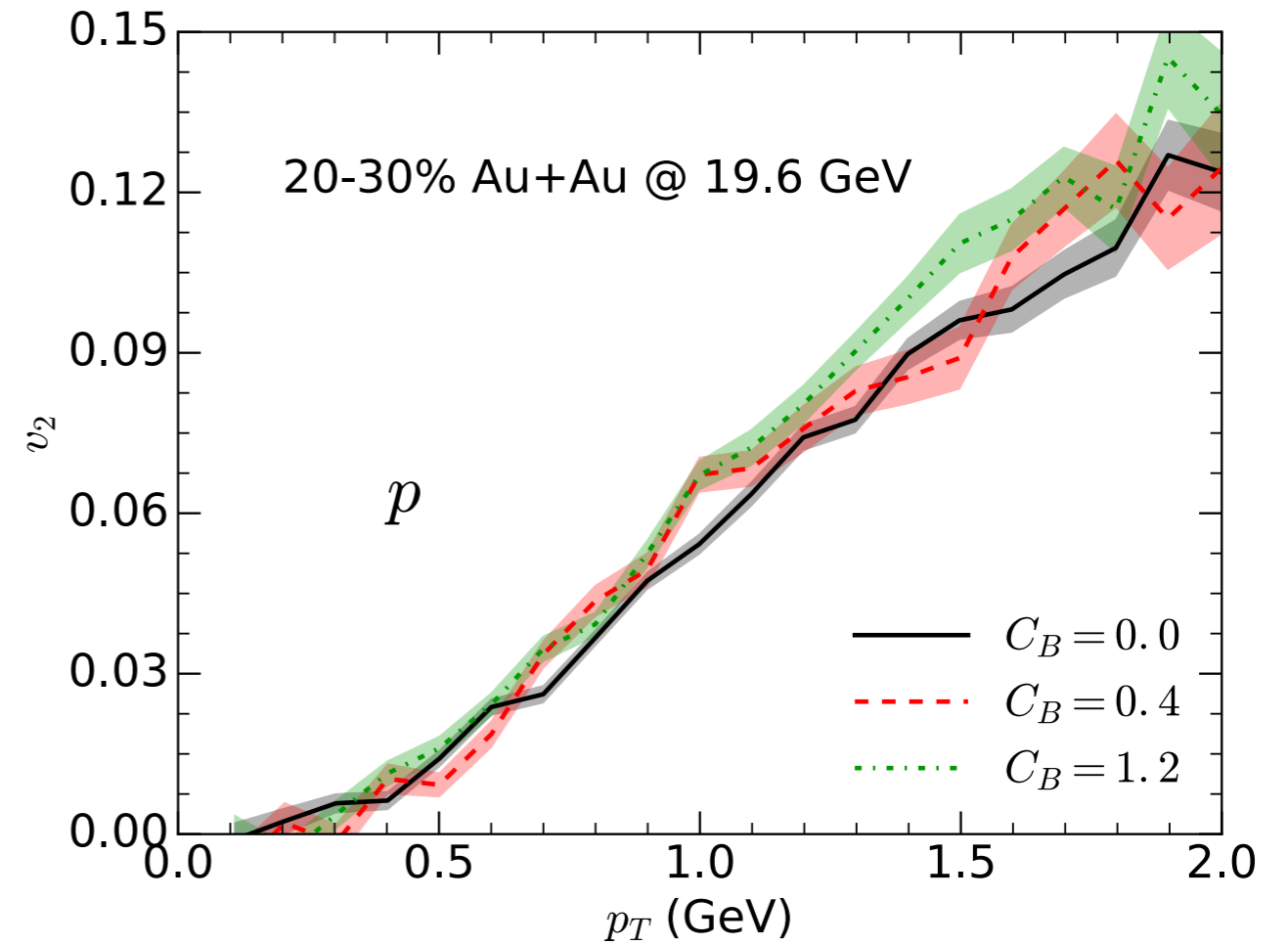
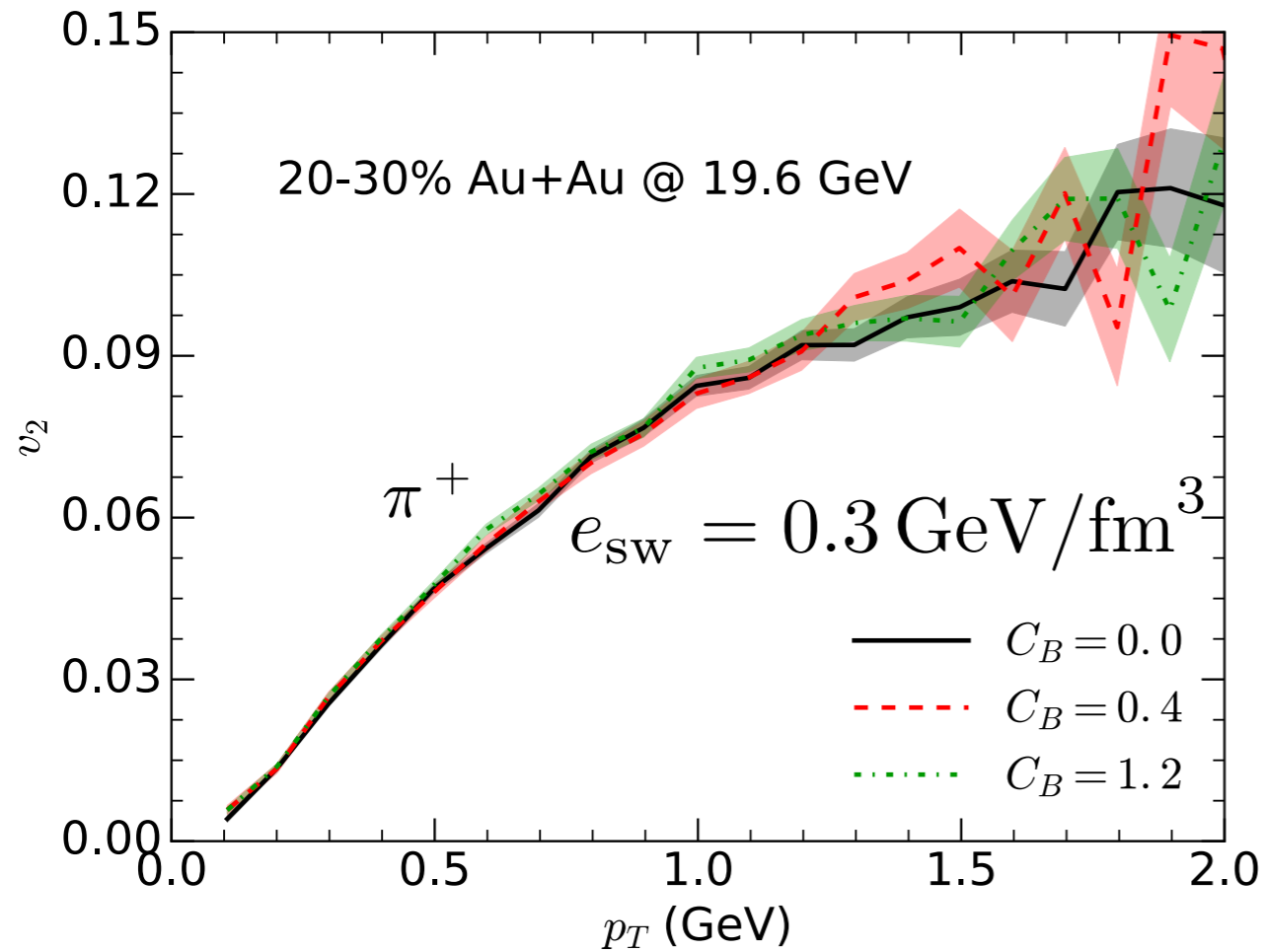
Effects of hadronic afterburner on v_2



- Momentum anisotropy keeps developing in the UrQMD phase
- Low p_T proton v_2 is **blue shifted** because of the pion “wind”

hadronic afterburner is essential

Effects of net baryon diffusion on v_2

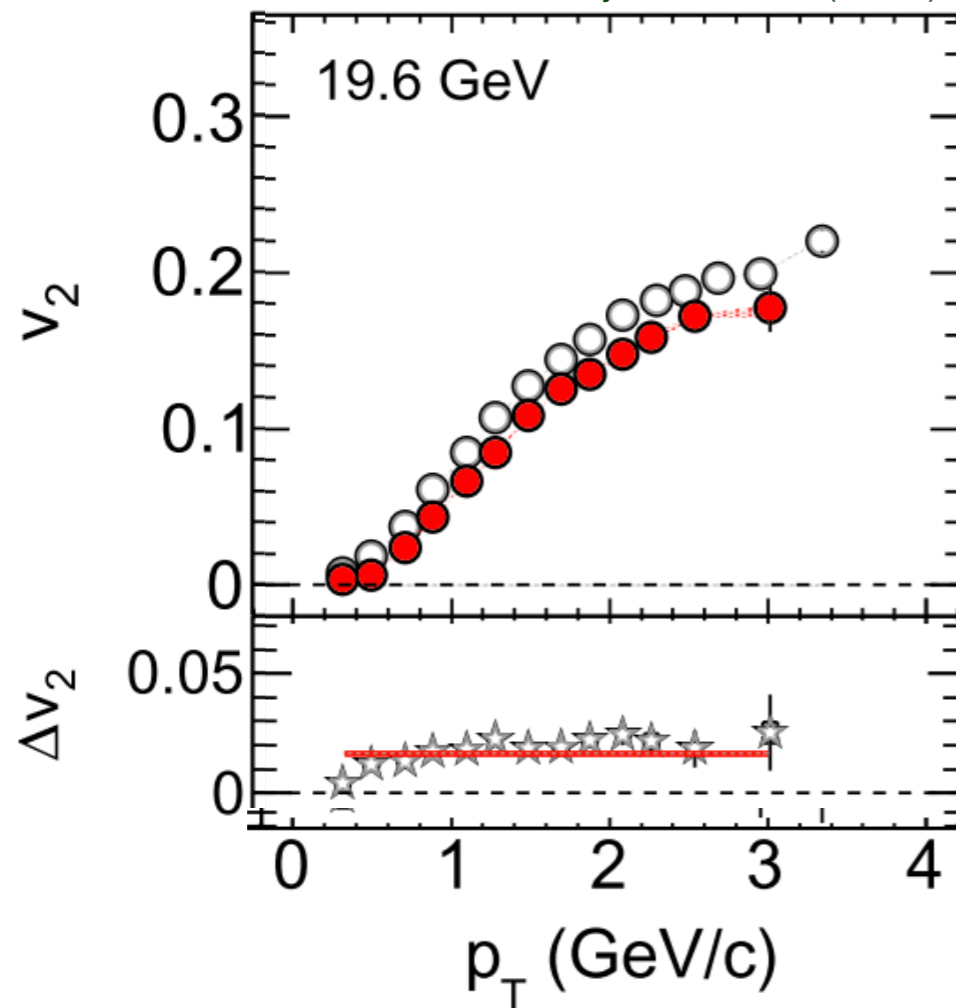


- Hadronic scatterings wash out most of the diffusion effects on v_2

$v_2(p)$ vs $v_2(\bar{p})$

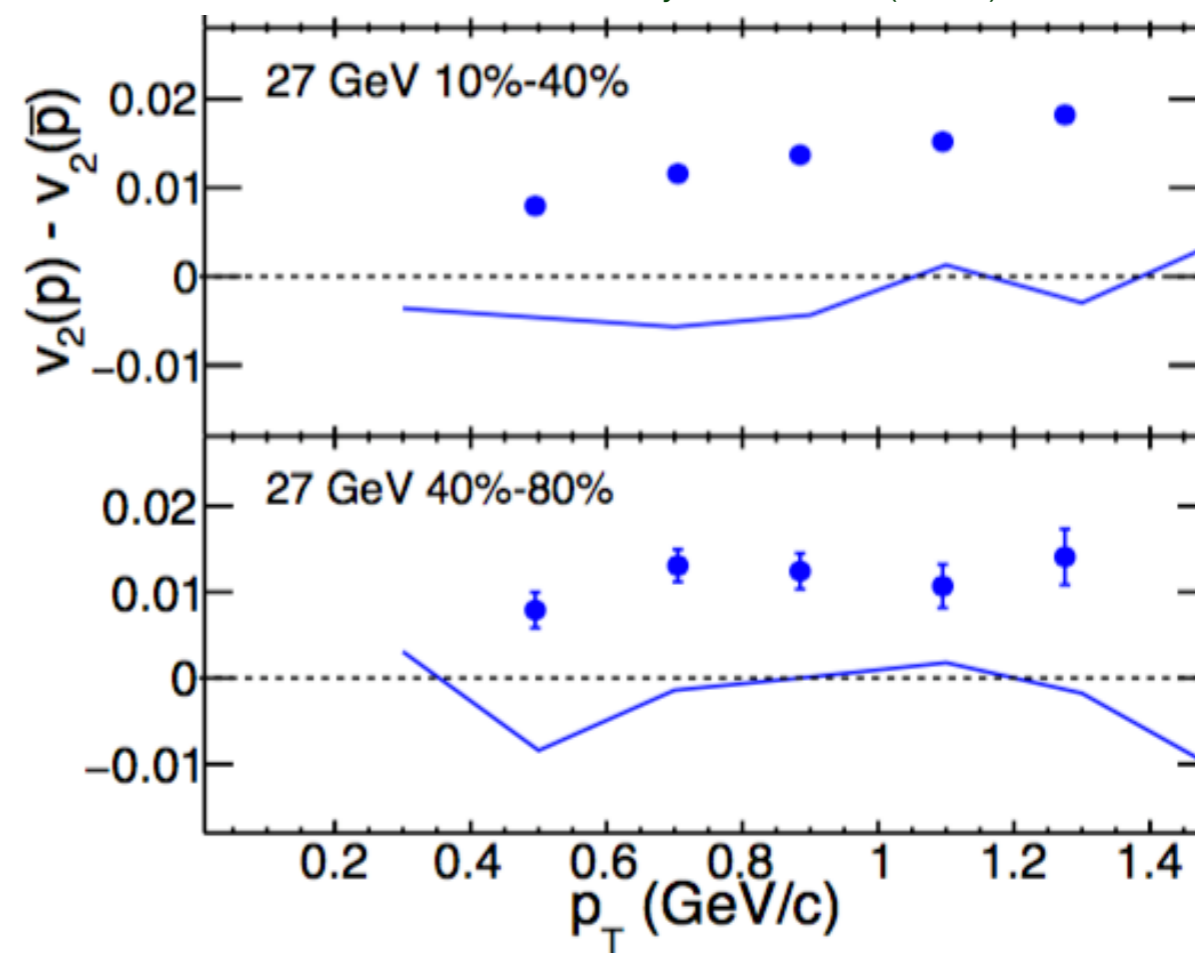
STAR

Phys.Rev. C88 (2013) 014902



AMPT

Phys.Rev. C93 (2016) no.1, 014907

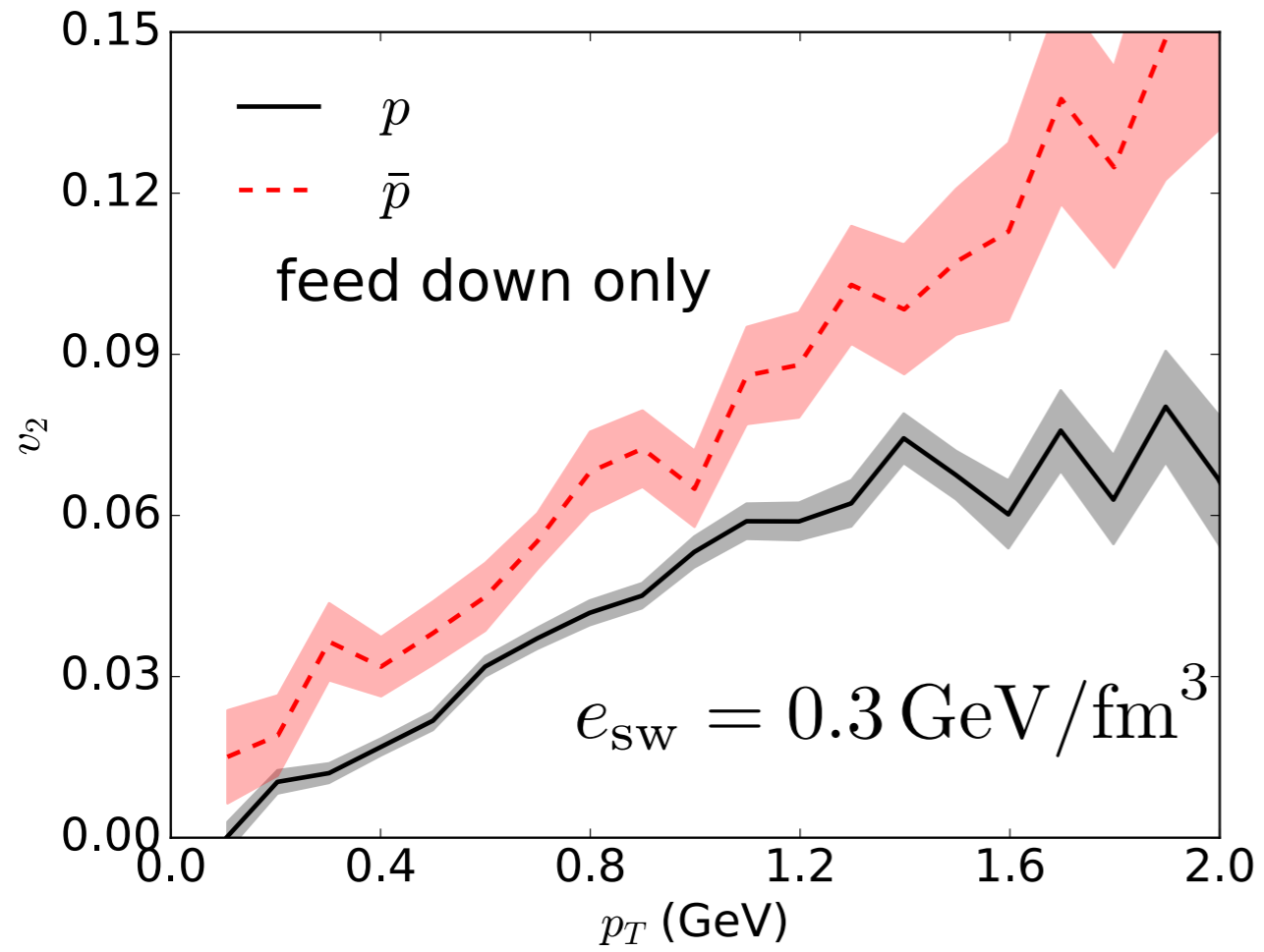


- Transport simulations give $v_2(\bar{p}) > v_2(p)$, which is **opposite** compared to the STAR data

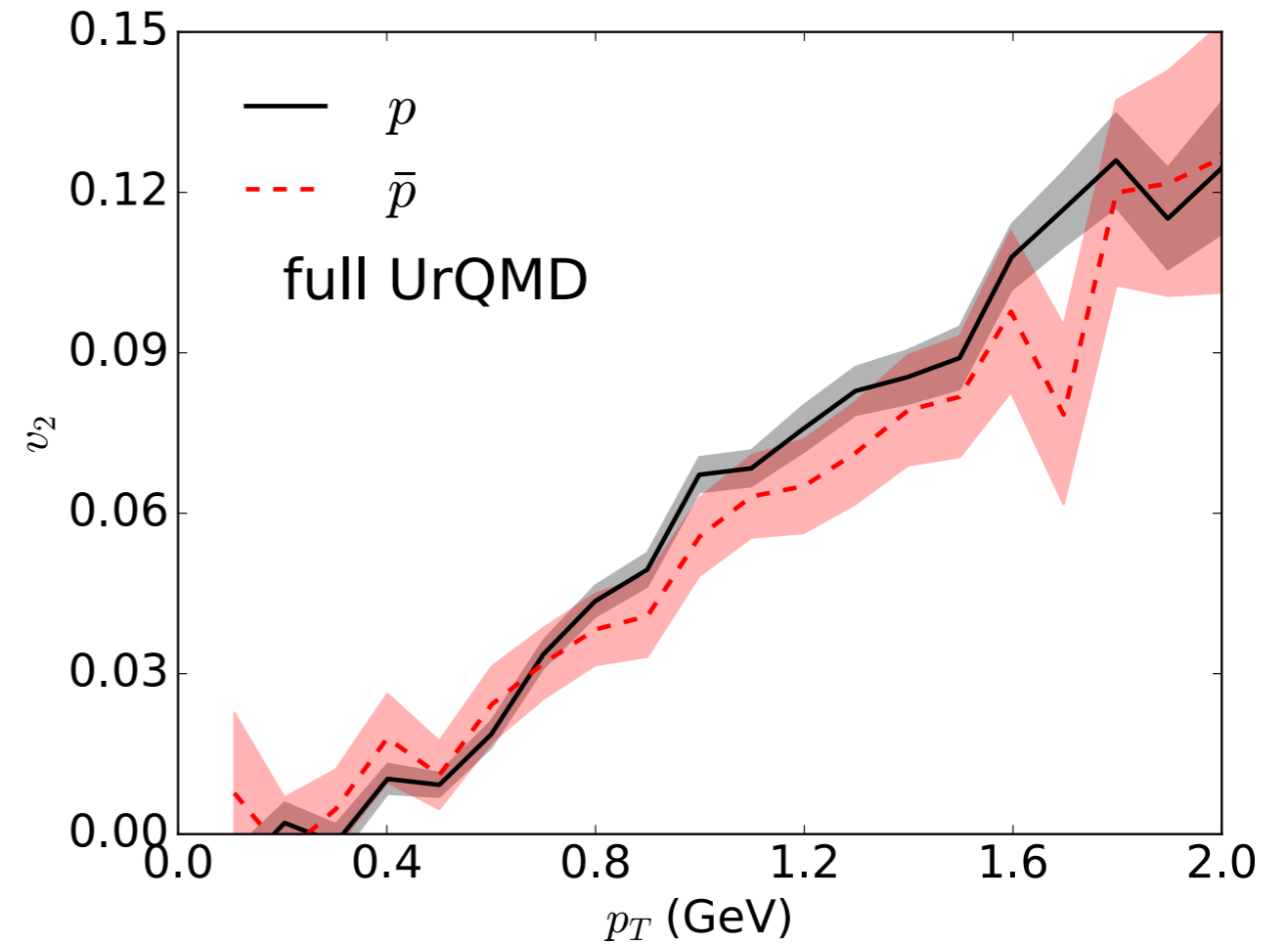
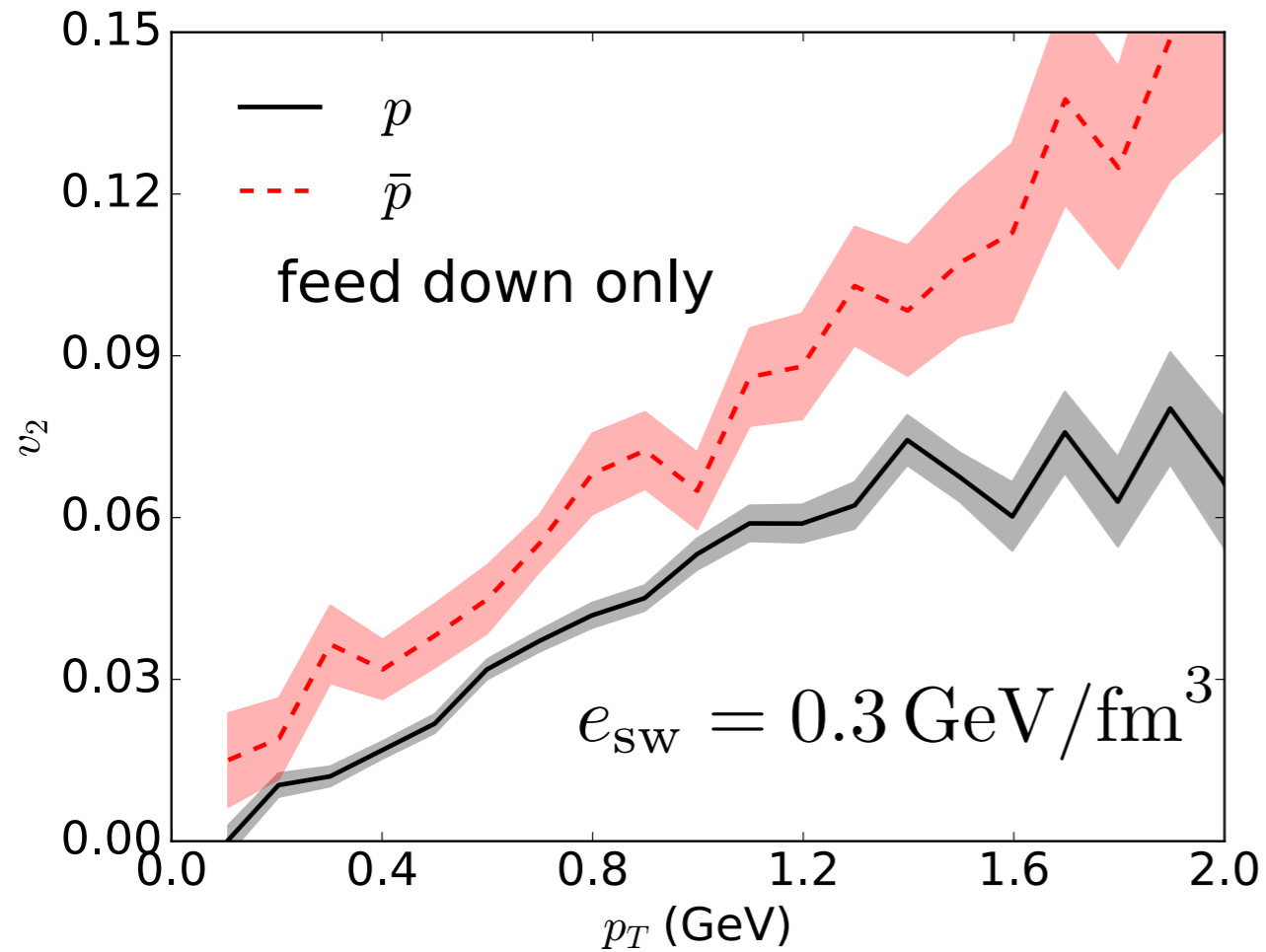
mean field effects in the hadronic phase?



$v_2(p)$ vs $v_2(\bar{p})$



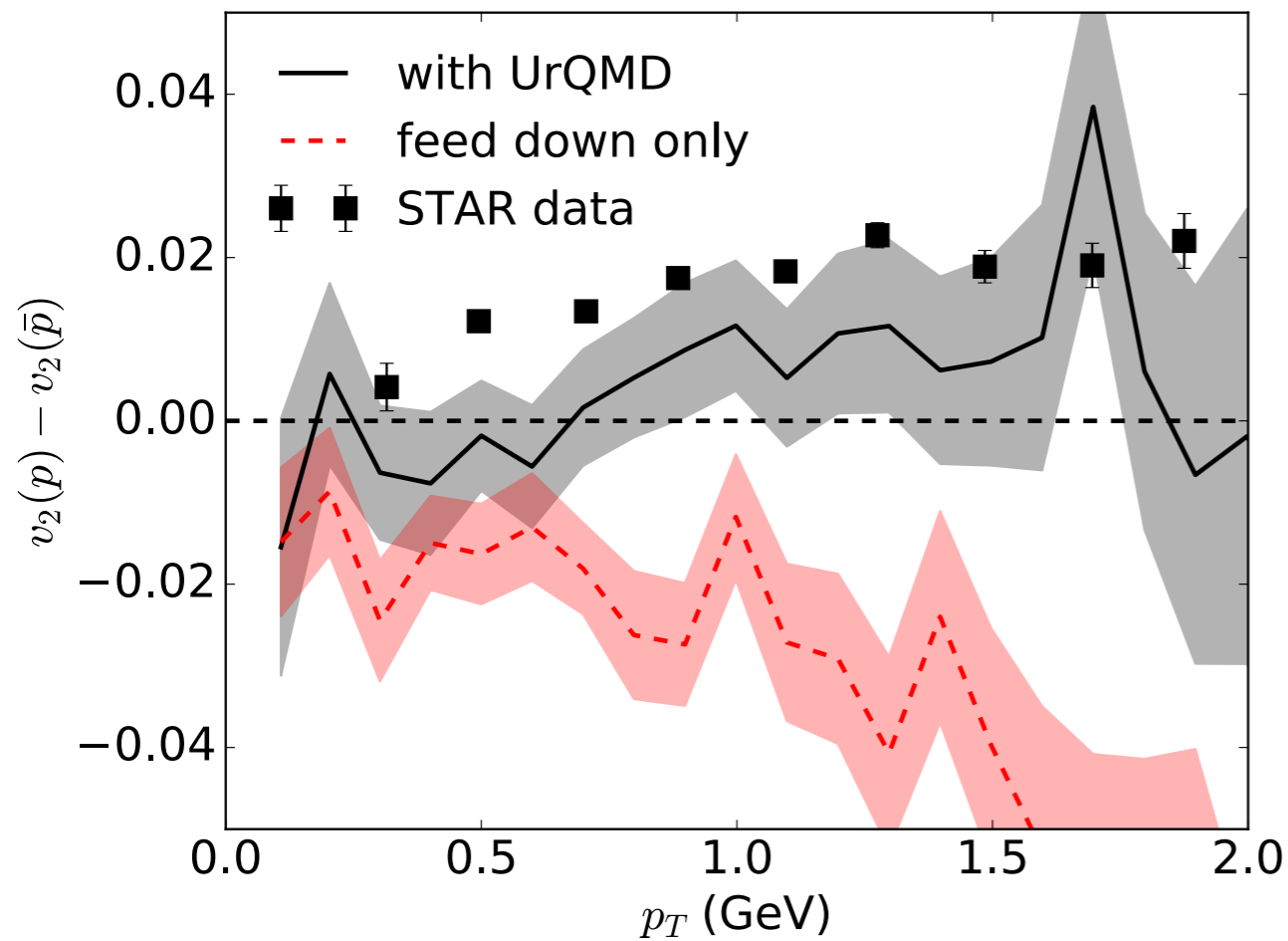
$v_2(p)$ vs $v_2(\bar{p})$



- Late stage hadronic scatterings correct the $v_2(p_T)$ ordering in our hybrid simulations

more statistics is needed

$v_2(p)$ vs $v_2(\bar{p})$

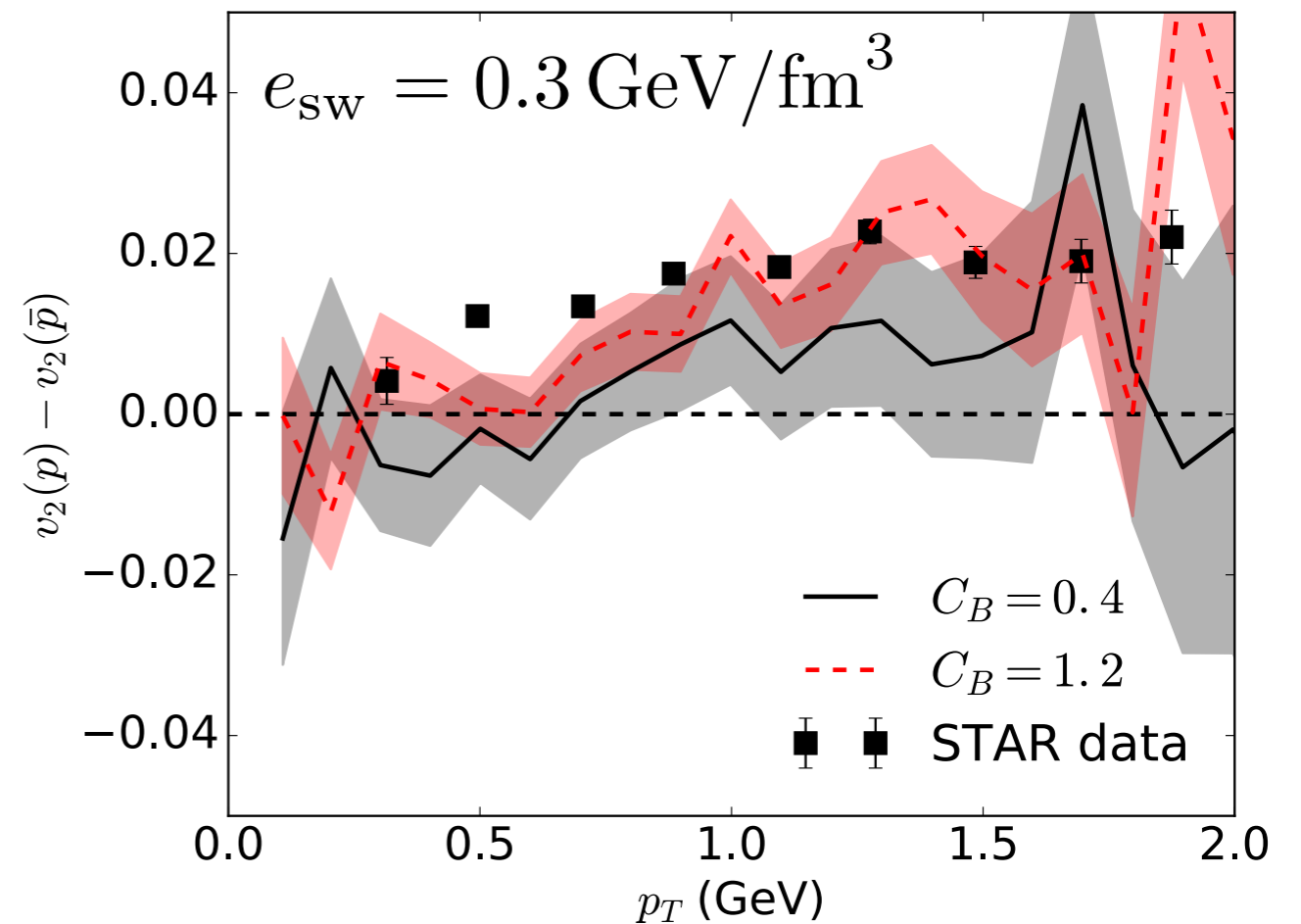
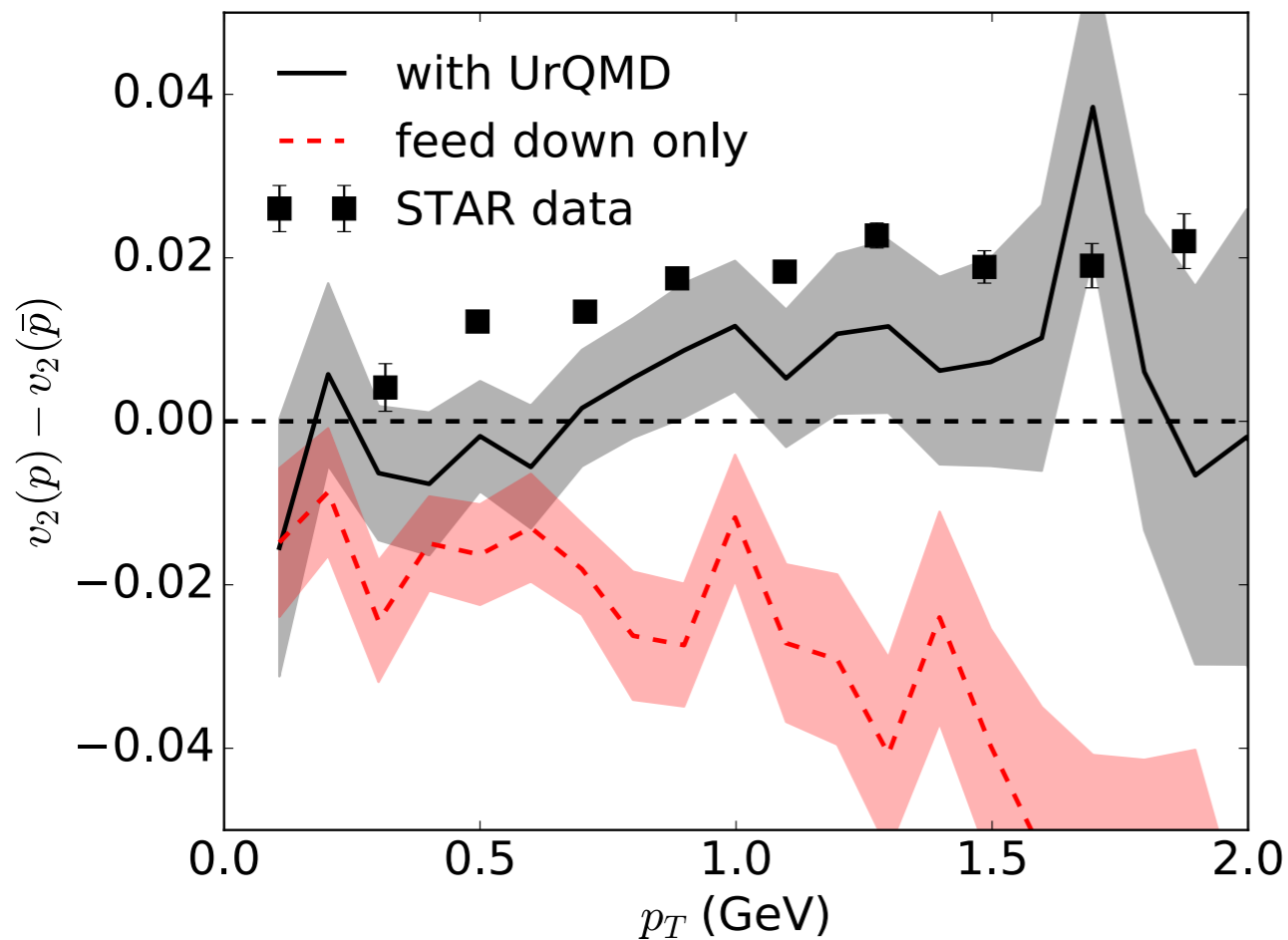


$$e_{sw} = 0.3 \text{ GeV}/\text{fm}^3$$

- Late stage hadronic scatterings correct the $v_2(p_T)$ ordering in our hybrid simulations

more statistics is needed

$v_2(p)$ vs $v_2(\bar{p})$

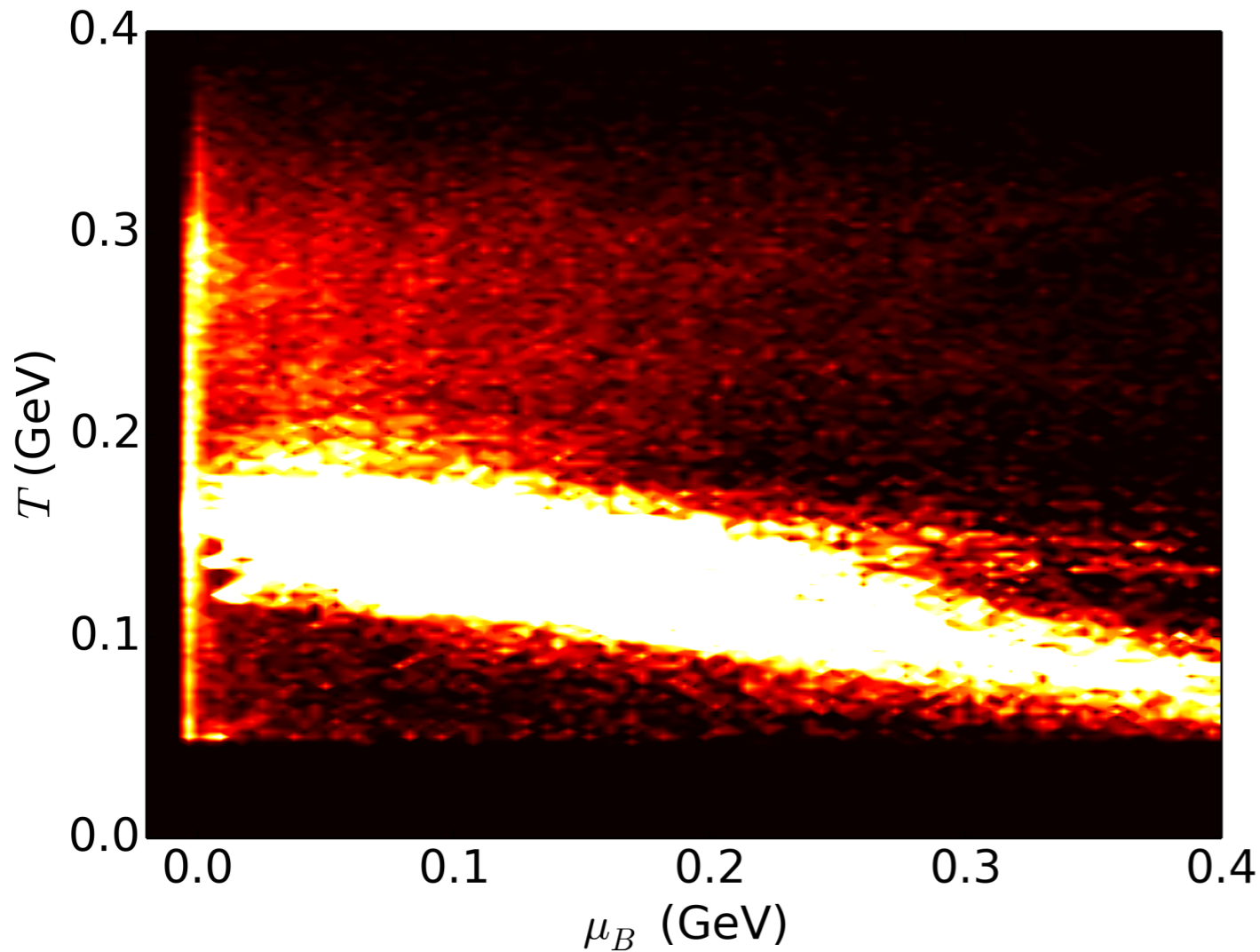


- Late stage hadronic scatterings correct the $v_2(p_T)$ ordering in our hybrid simulations
- A larger net baryon diffusion constant results a larger $v_2(p) - v_2(\bar{p})$

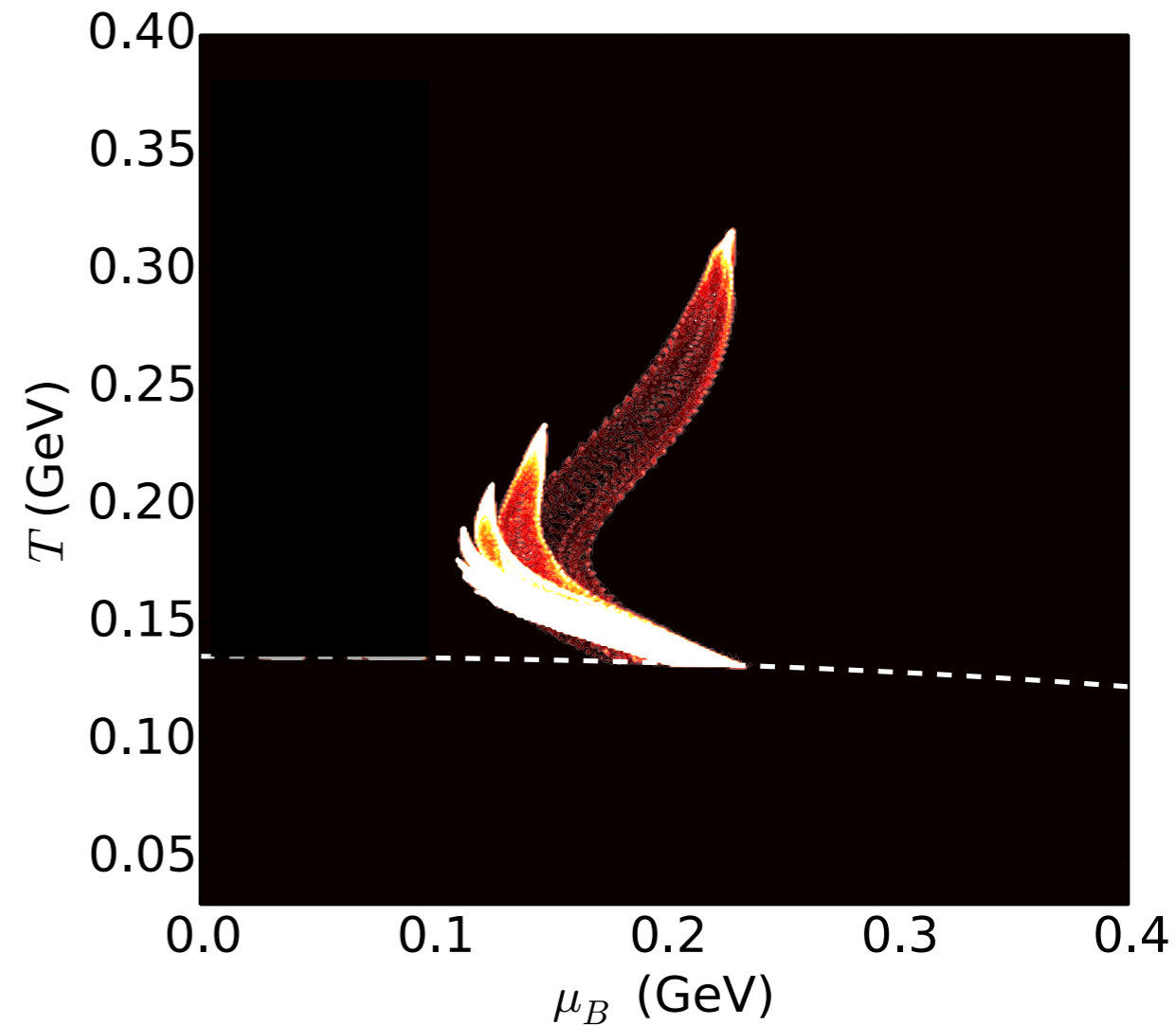
more statistics is needed

Fireball evolution in the QCD phase diagram

Fluctuating



Smooth



- In the fluctuating case, the hydrodynamic fluid cells are scattered on the phase diagram

A challenge to the search of critical point

Diffusion δf is essential

$$E \frac{dN_i}{d^3 p} = \frac{g_i}{(2\pi)^3} \int p^\mu d^3 \sigma_\mu(x) (f_0(x, p) + \delta f(x, p))$$

$$f_0^i(x, p) = \frac{1}{e^{(E - b_i \mu_B(x))/T(x)} \pm 1}$$

$$\delta f_0^i(x, p) = f_0^i(x, p) (1 \pm f_0^i(x, p)) \left(\frac{n_B}{e + \mathcal{P}} - \frac{b_i}{E} \right) \frac{p \cdot q}{\hat{\kappa}}$$

$$N^B - N^{\bar{B}} = \int d^3 \sigma_\mu \sum \frac{g_i}{(2\pi)^3} \int_p p^\mu \left[(f_0^B(x, p) - f_0^{\bar{B}}(x, p)) + (\delta f^B(x, p) - \delta f^{\bar{B}}(x, p)) \right]$$

$$= \int d^3 \sigma_\mu (n_B u^\mu + q^\mu)$$

$$\partial_\mu (n_B u^\mu + q^\mu) = 0$$



$N^B - N^{\bar{B}}$
is conserved

- With diffusion, δf is **essential** to ensure net baryon number conservation

Diffusion δf is essential

$$E \frac{dN_i}{d^3 p} = \frac{g_i}{(2\pi)^3} \int p^\mu d^3 \sigma_\mu(x) (f_0(x, p) + \delta f(x, p))$$

$$f_0^i(x, p) = \frac{1}{e^{(E - b_i \mu_B(x))/T(x)} \pm 1}$$

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$$= \int d^3 \sigma_\mu (n_B u^\mu + q^\mu)$$

$$\partial_\mu (n_B u^\mu + q^\mu) = 0$$

Conservation laws (net baryon number and total energy) are checked at every time step; the relative violation is below 1×10^{-5}

- With diffusion, δf is **essential** to energy conservation

Baryon diffusion constant

$$N_Q^\mu = \sum_{i=1}^N g_i Q_i \int dK_i k_i^\mu f_k^i,$$

Relaxation time
approximation

$$q_B^\mu = \sum_{i=1}^N g_i b_i \int dK_i k_i^{\langle \mu \rangle} \delta f_k^i$$

$$q_B^\mu = \kappa_B \nabla^\mu \alpha_B = \tau_R \hat{\kappa}_B \nabla^\mu \alpha_B$$

$$\hat{\kappa}_B = \frac{1}{3} \Delta_{\mu\nu} \sum_{i=1}^N g_i b_i \int dK_i k_i^\nu k_i^\alpha f_{ik}^{(0)} \tilde{f}_{ik}^{(0)} \left[\frac{n_B}{e + P} - \frac{b_i}{E_{ik}} \right]$$

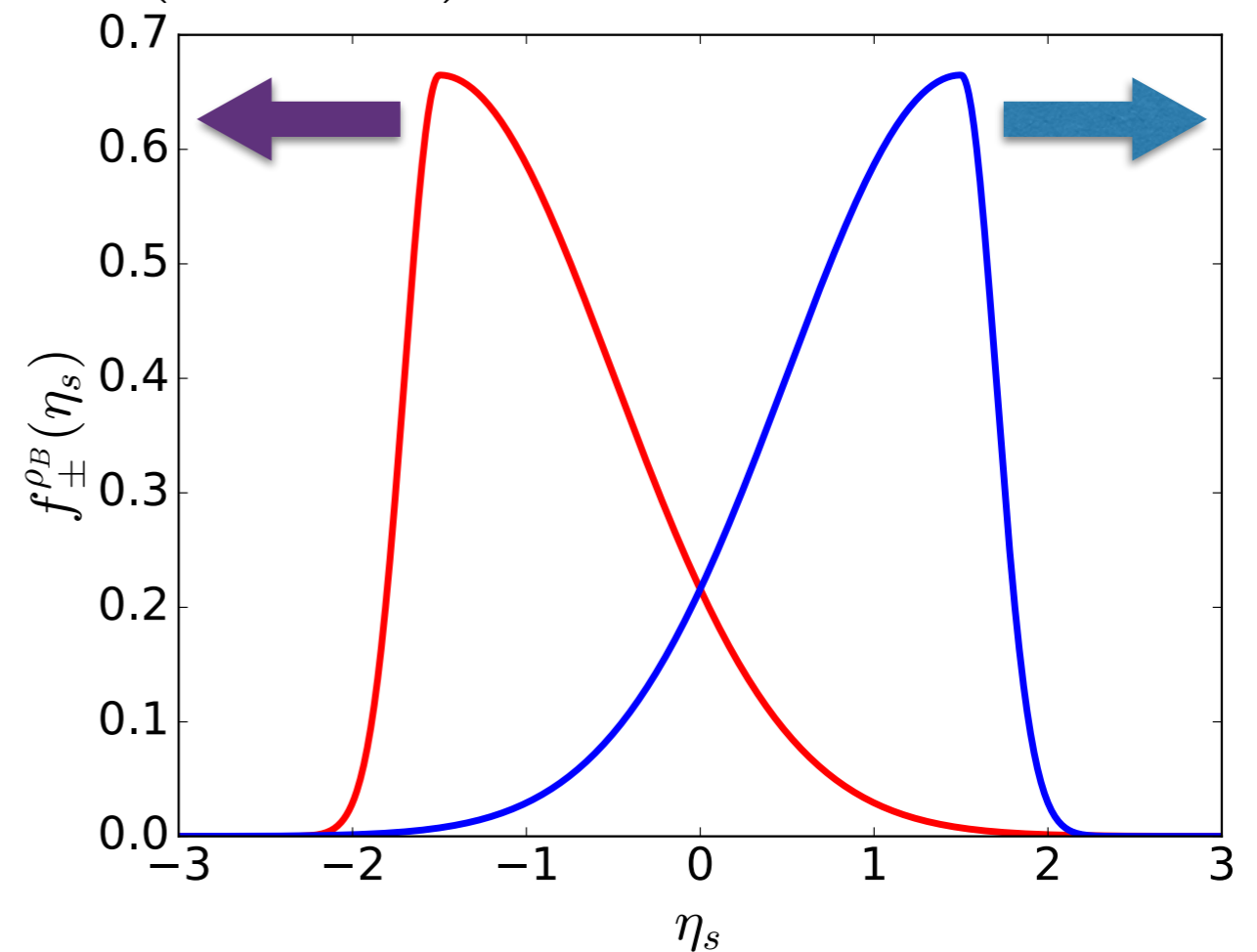
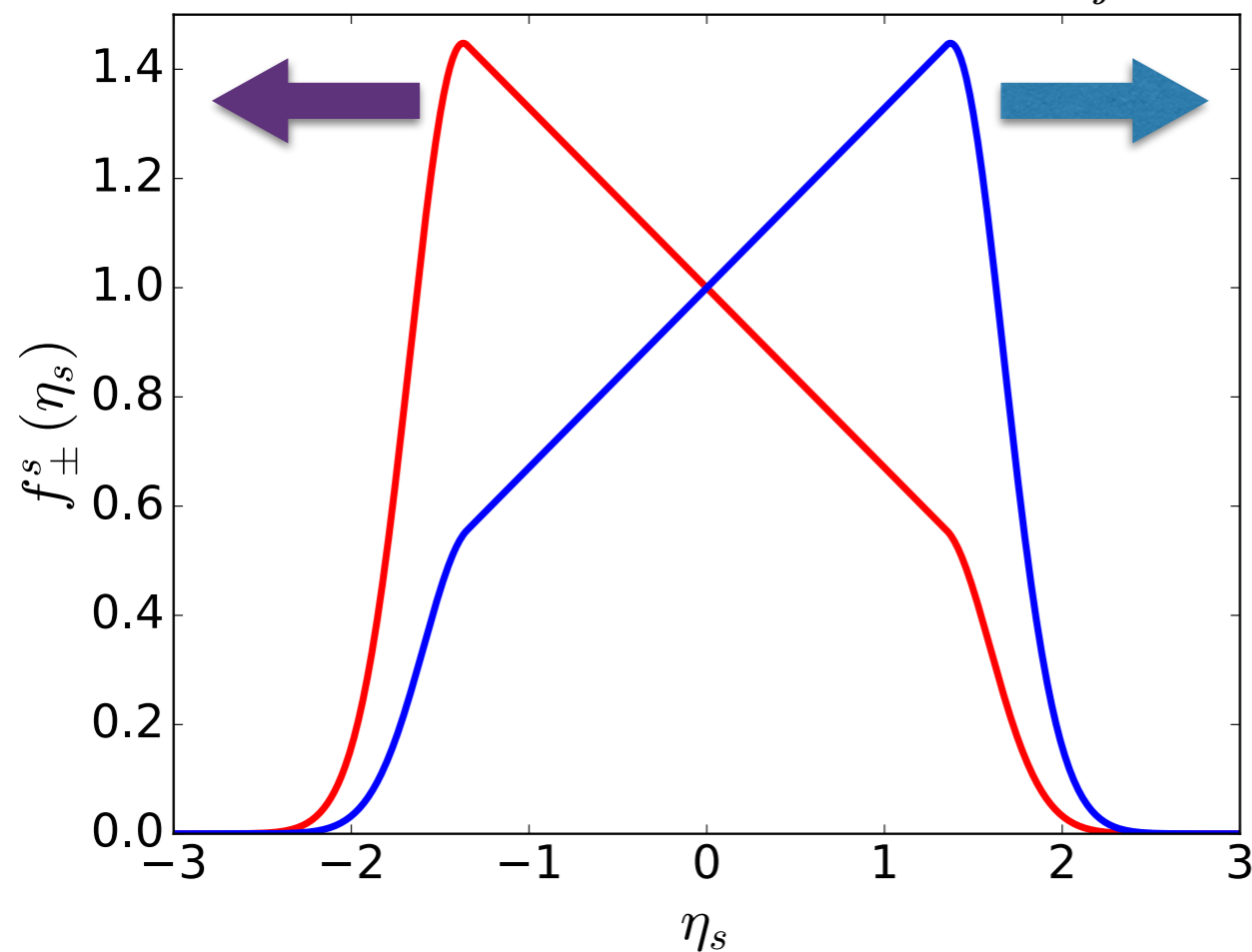
$$\kappa_B = \frac{C_B}{T} n_B \left(\frac{1}{3} \coth(\alpha_B) - \frac{n_B T}{e + \mathcal{P}} \right).$$

Initialize MUSIC with 3D density profiles

MC-Glauber model:

$$s(x, y, \eta; \tau_0) = \frac{s_0}{\tau_0} \sum_{i=\pm} f_i^s(\eta) s_i(x, y), \quad \rho_B(x, y, \eta; \tau_0) = \frac{1}{\tau_0} \sum_{i=\pm} f_i^{\rho_B}(\eta) s_i(x, y).$$

$$s_i(r, \phi) = \sum_{j=1}^{N_{\text{part}}^i} \frac{1}{2\pi\sigma^2} \exp\left(-\frac{(\mathbf{r} - \mathbf{r}_j)^2}{2\sigma^2}\right)$$

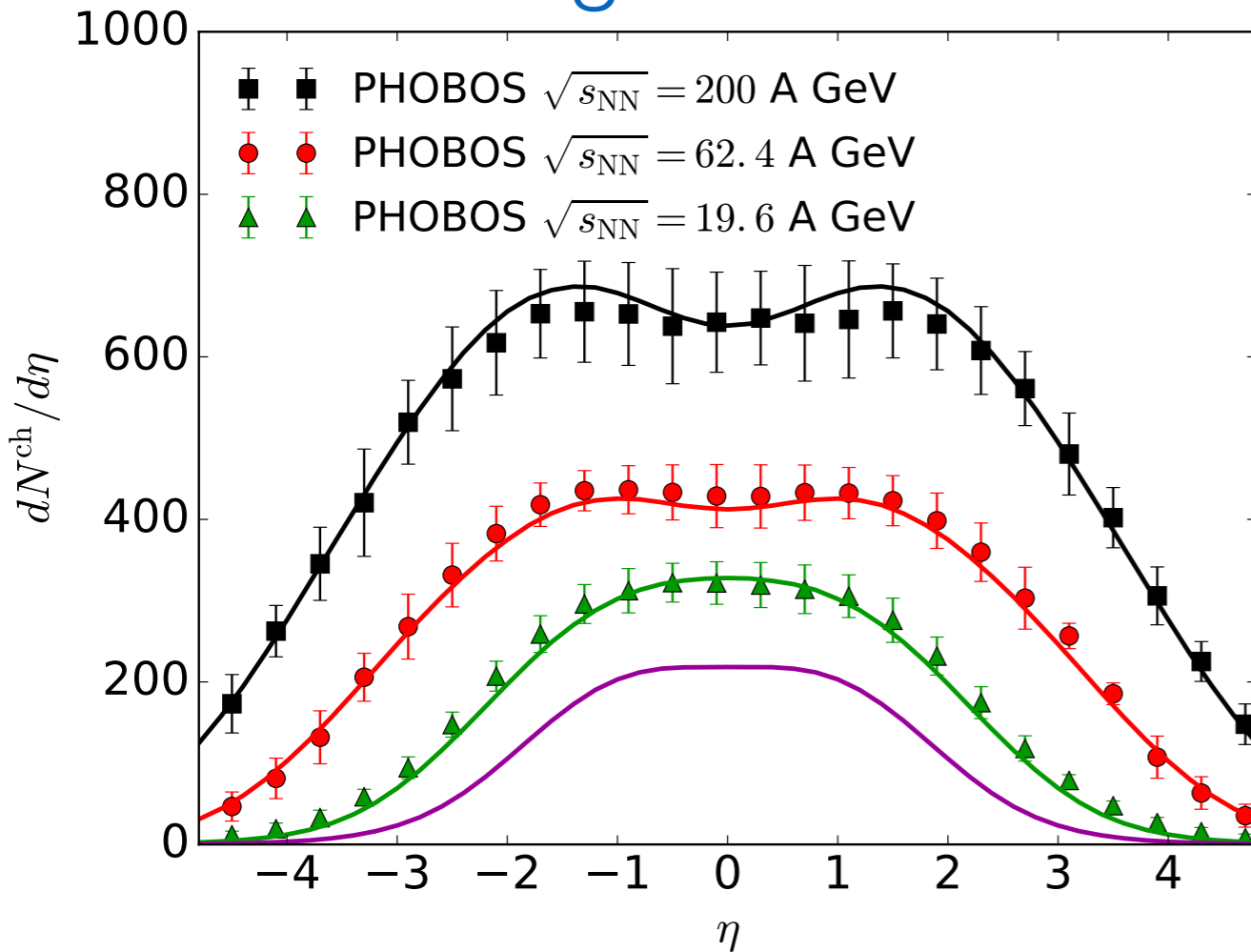


In every event, the net baryon number is

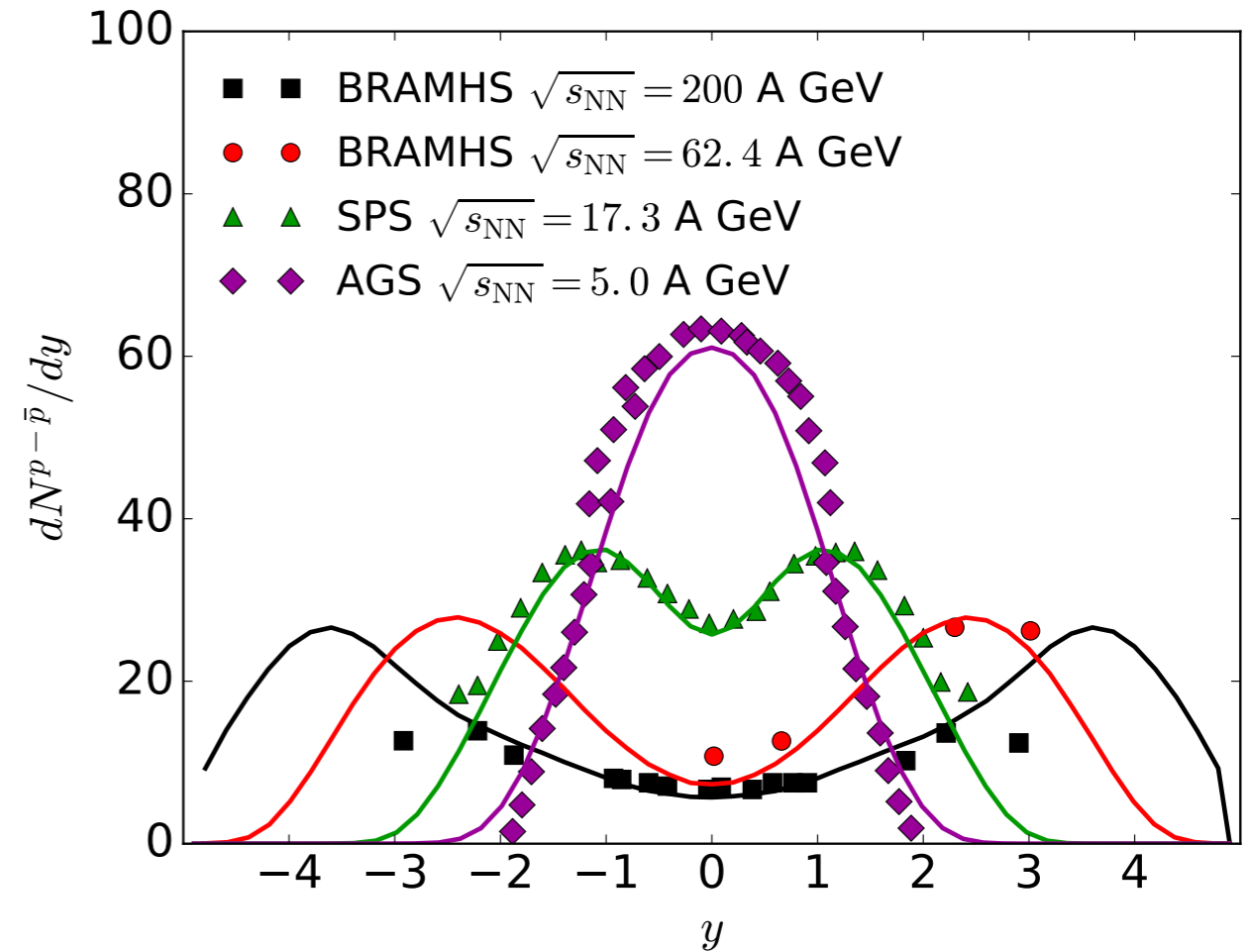
$$\int \tau_0 d\eta_s \int d^2 \mathbf{x}_{\perp} \rho_B(\mathbf{x}_{\perp}, \eta_s) = N_{\text{part}}.$$

Rapidity dependence of particle yield

charged hadrons

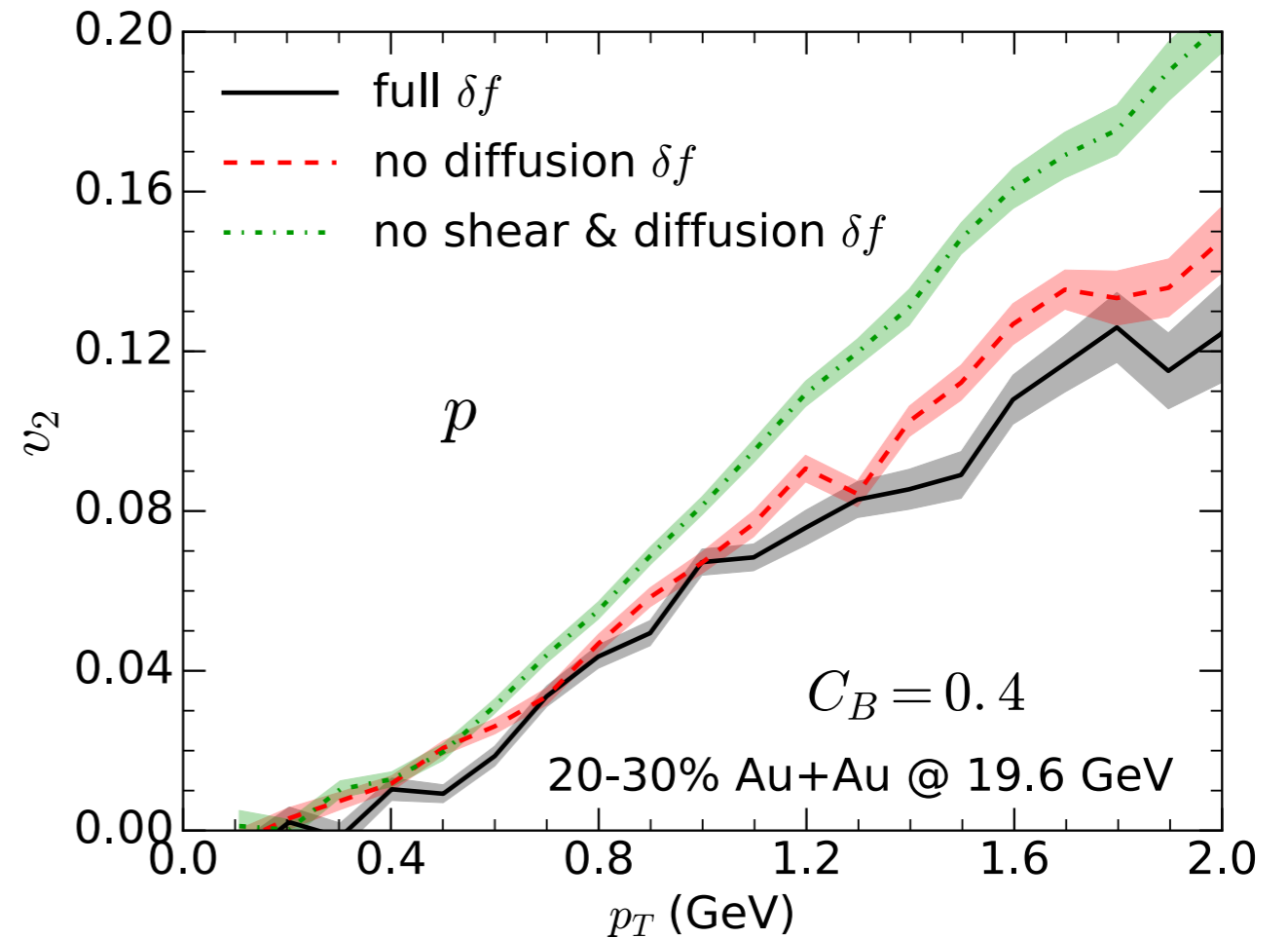
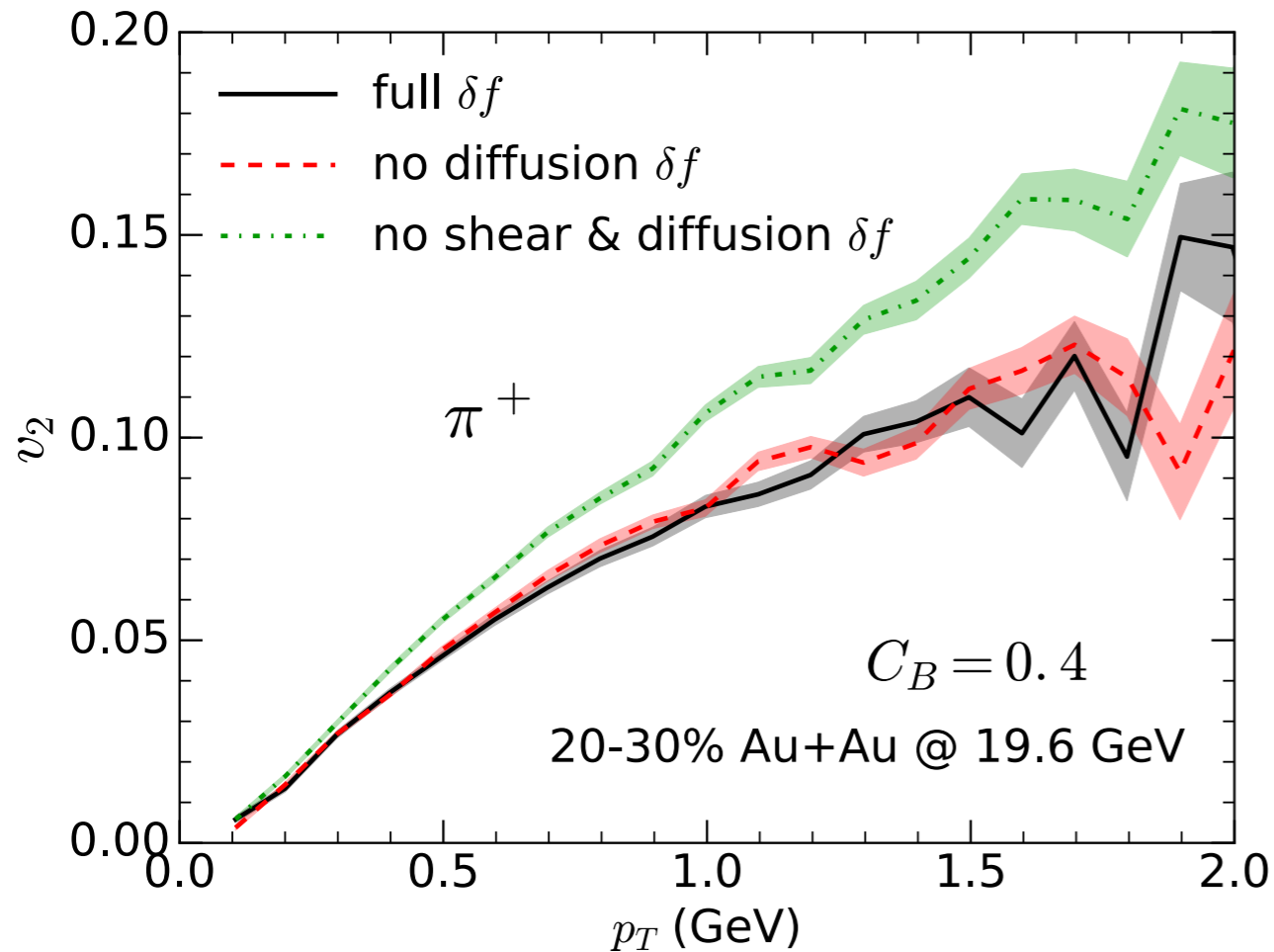


net protons



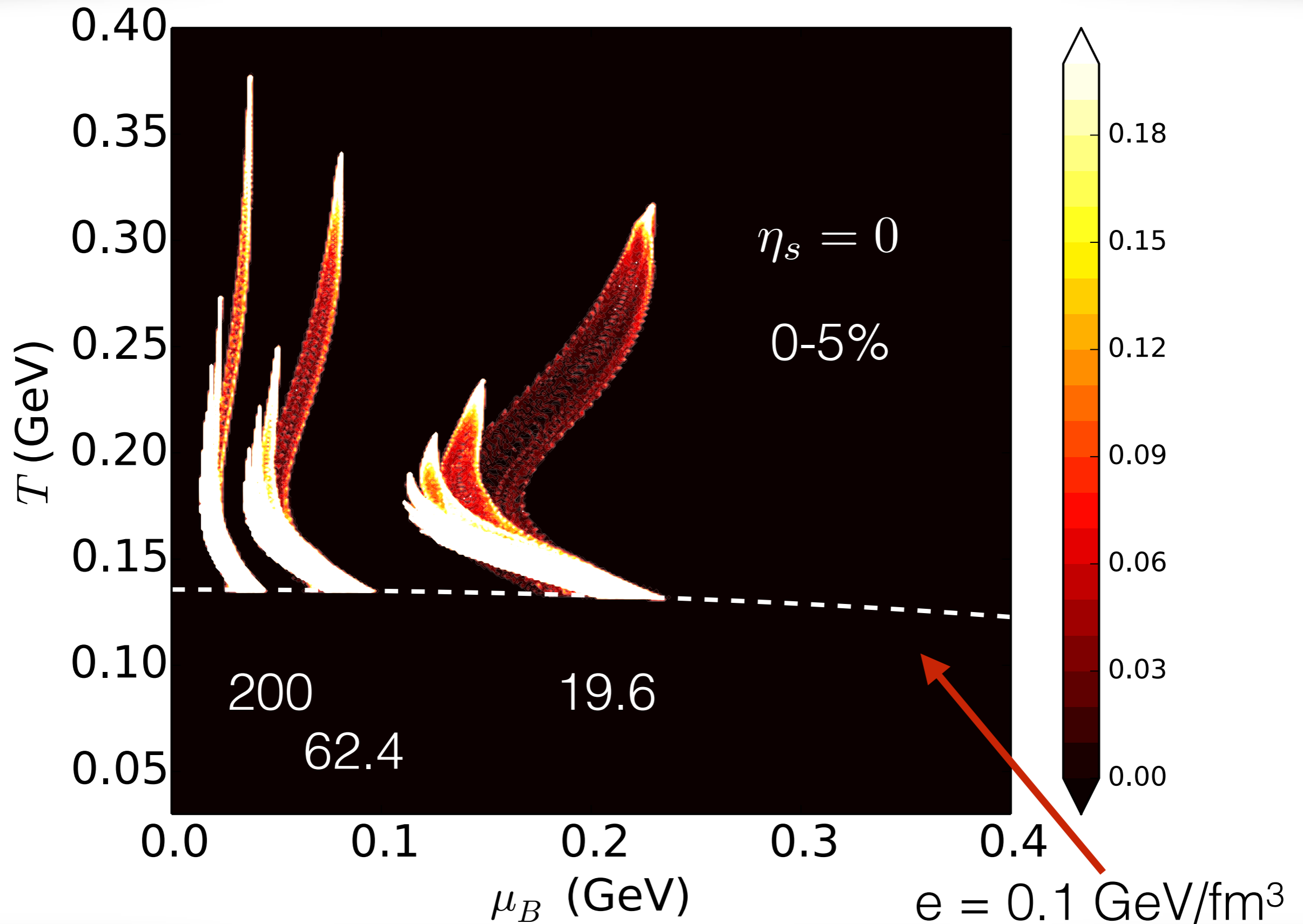
- Full (3+1)-d simulations with net baryon current and its diffusion
- Initial longitudinal profiles are tuned to reproduce the measured (pseudo)-rapidity distributions for charged hadrons and net protons

Effects of net baryon diffusion on v_2

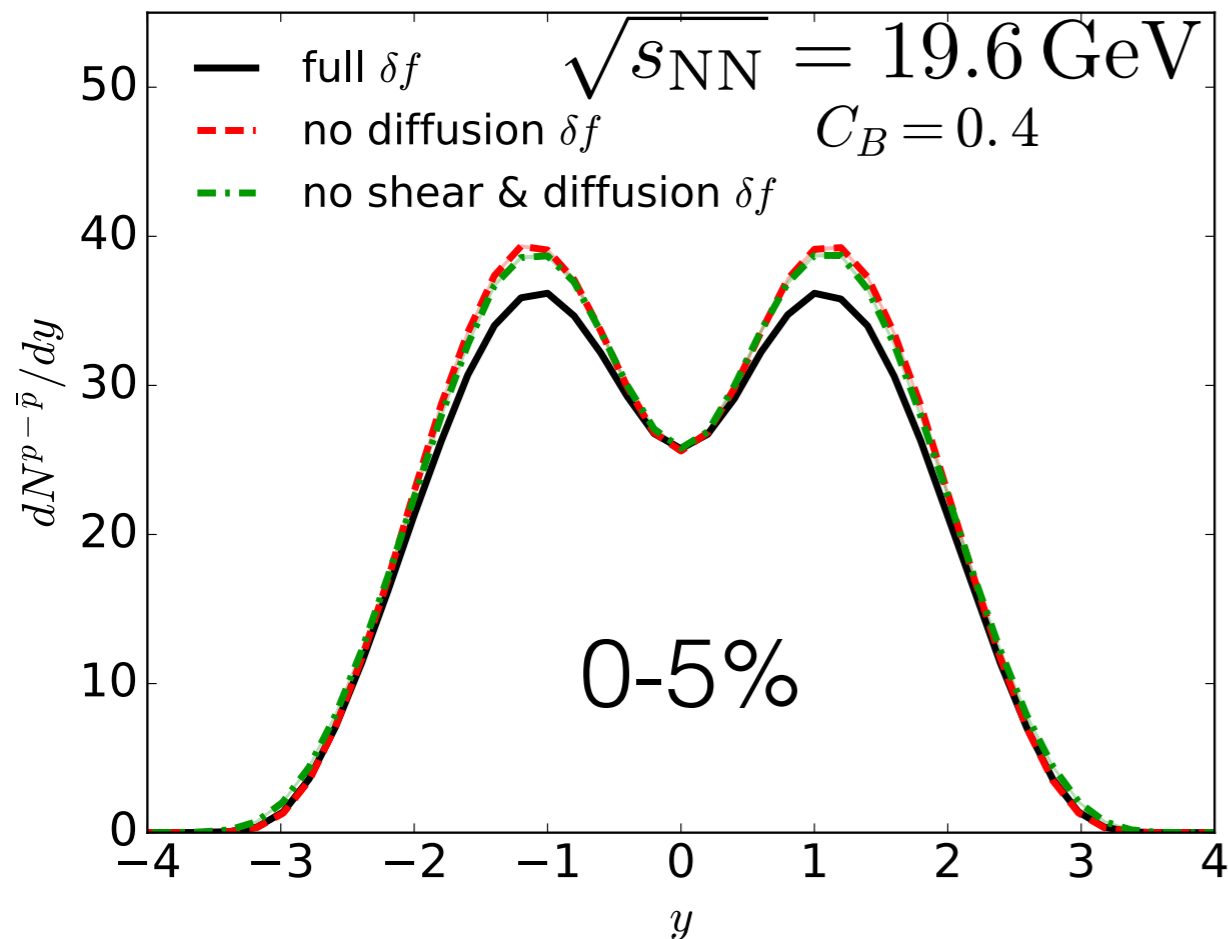


- Identified particles v_2 are more sensitive to shear out-of-equilibrium corrections

Compass for the QCD phase diagram



Effects of net baryon diffusion on particle yields



On the switching hyper-surface,

	$C_B = 0.0$	$C_B = 0.4$	$C_B = 1.2$
$\langle \mu_B \rangle$ (GeV)	0.245	0.260	0.276
σ_{μ_B} (GeV)	0.092	0.075	0.069

- The diffusion δf **changes** the net proton number
- Larger diffusion constant results a ~ 30 MeV **larger** averaged chemical potential on the switching hyper-surface; It **reduces** the standard deviation of μ_B by $\sim 25\%$

mapping the QCD phase diagram in precision