

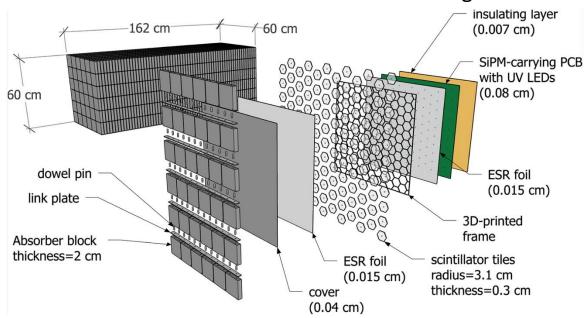
Sean Preins
University of California, Riverside
7/30/25



Background

- This test serves as the pre-production effort for the SiPM-on-tile Zero Degree Calorimeter, building off previous prototype tests
- Assess the performance and scalability of the design
- Determine assembly steps, QA check procedures, and other final improvements

Initial SiPM-on-tile ZDC design



Design and simulation of a SiPM-on-tile ZDC for the future EIC, and its performance with graph neural networks

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Previous Test Beams

- Gen I prototype consisted of 10 sampling layers, 40 channels
 - Tested with 4 GeV positrons at Jefferson Lab in Jan 2023
- Gen II prototype consists of 9 sampling layers, 192 channels
 - Tested in the STAR experiment hall at RHIC, run from April – October 2024
 - Currently running now



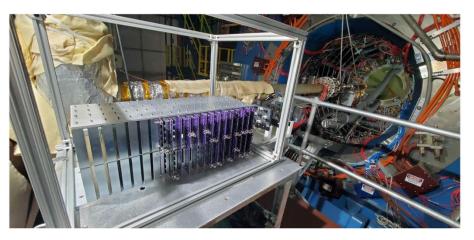




Article

Beam Test of the First Prototype of SiPM-on-Tile Calorimeter Insert for the EIC Using 4 GeV Positrons at Jefferson Laboratory

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First-ever deployment of a SiPM-on-tile calorimeter in a collider: a parasitic test with 200 GeV *p p* collisions at RHIC.

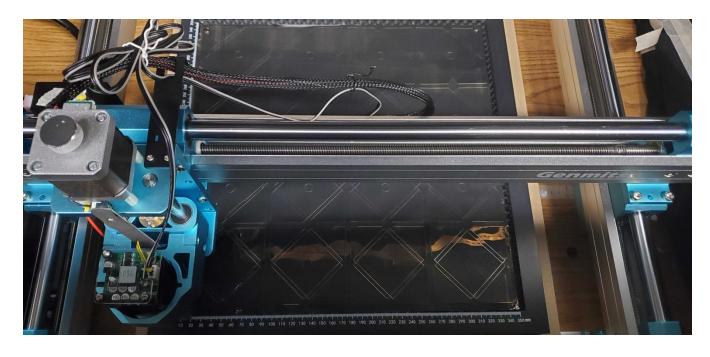
Generation III Prototype Design

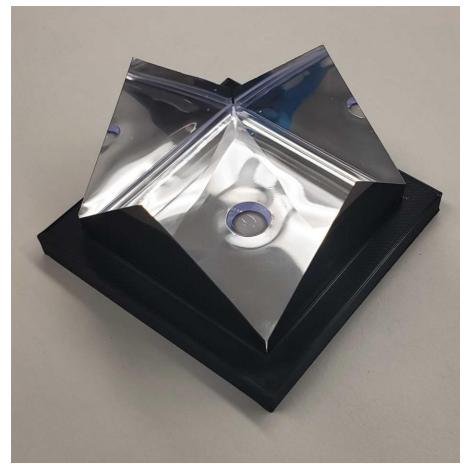
- Area of 29.4 cm x 28.8 cm
- Each layer consists of 5 x 5 square scintillating tiles, shifted diagonally every other layer
- 15 layers, 25 channels per layer
- 375 channels total
- 7 dead channels, 98.1% channels functional. No dead channels in shower core region.



2 cm of iron = $1.1 X_0$

Tile Construction





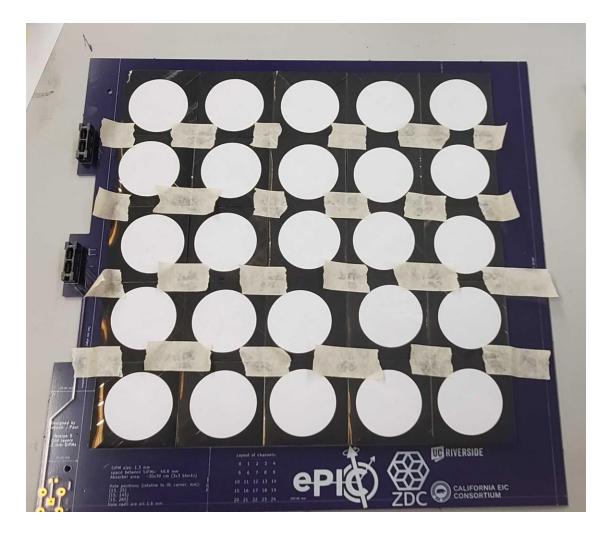
- ESR foil is cut using a CNC laser, and individually folded around each tile
- Tile dimensions 48.8 mm x 48.8 mm x 4 mm

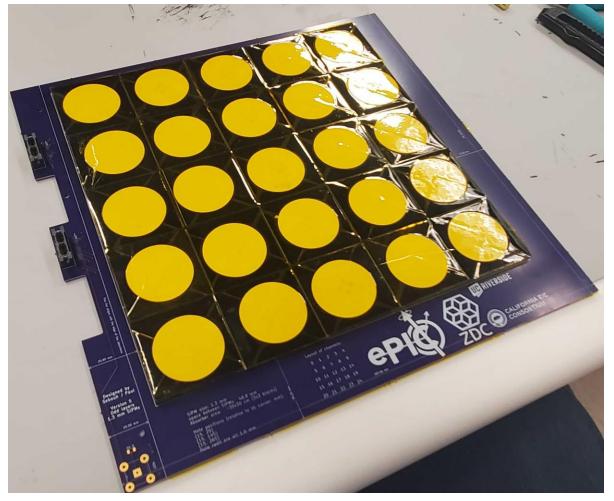




PCB Construction

 Hamamatsu s14160-1315PS 1.3 mm
 SiPMs are soldered to the PCB using an IR oven

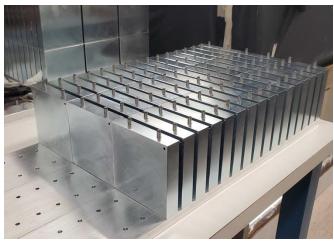




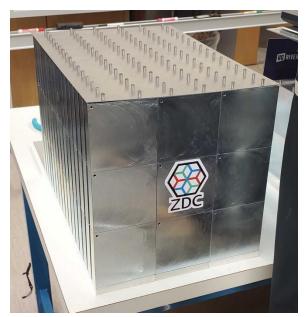
- After QA checks, tiles are affixed to the board via double-sided tape
- Polyimide film is placed on the front and back of the board to protect it from damage

Iron Structure Construction





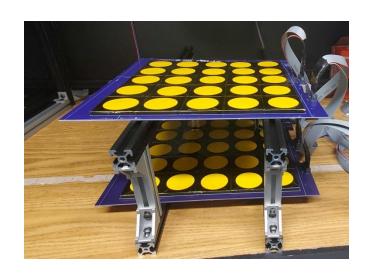
LEGO-style pin-and-block assembly

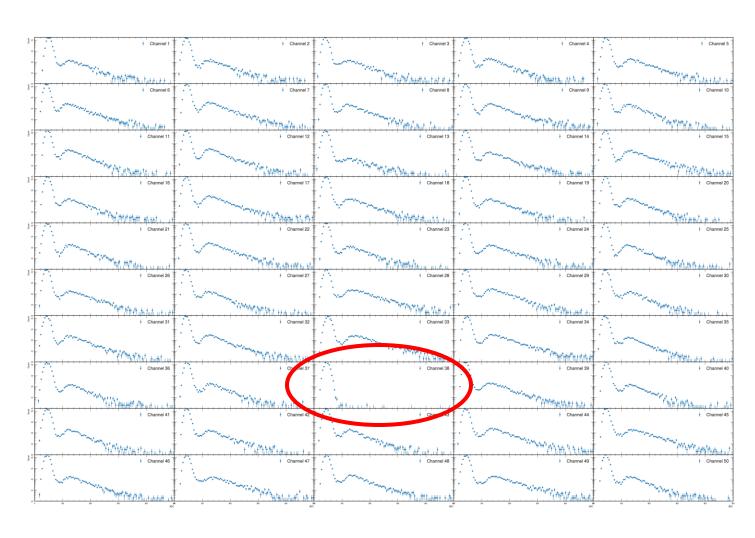


 Active components can be placed in their respective slots post-assembly

Benchtop QA Checks

- Finalized layers are checked with a pedestal test, IV scan test, and cosmic test
- Layers are then tested with cosmic rays two at a time with a MAJ=2 trigger, to confirm that it begins to record a landau curve



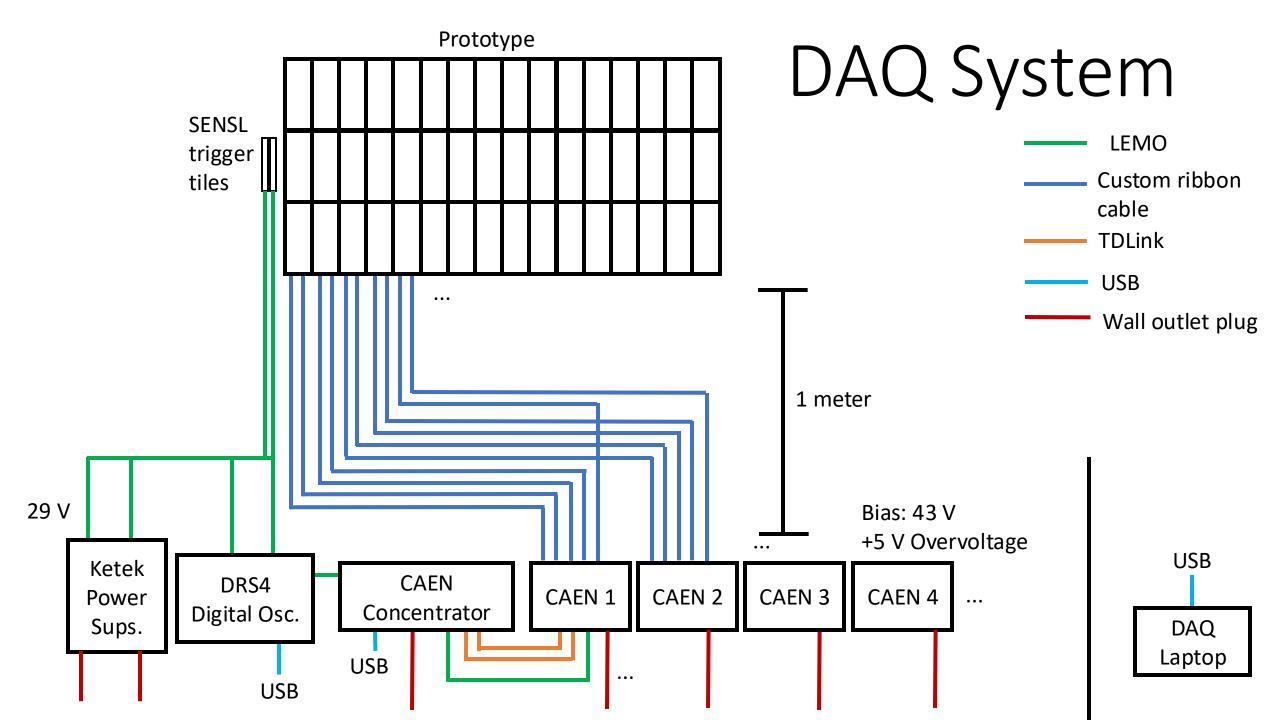


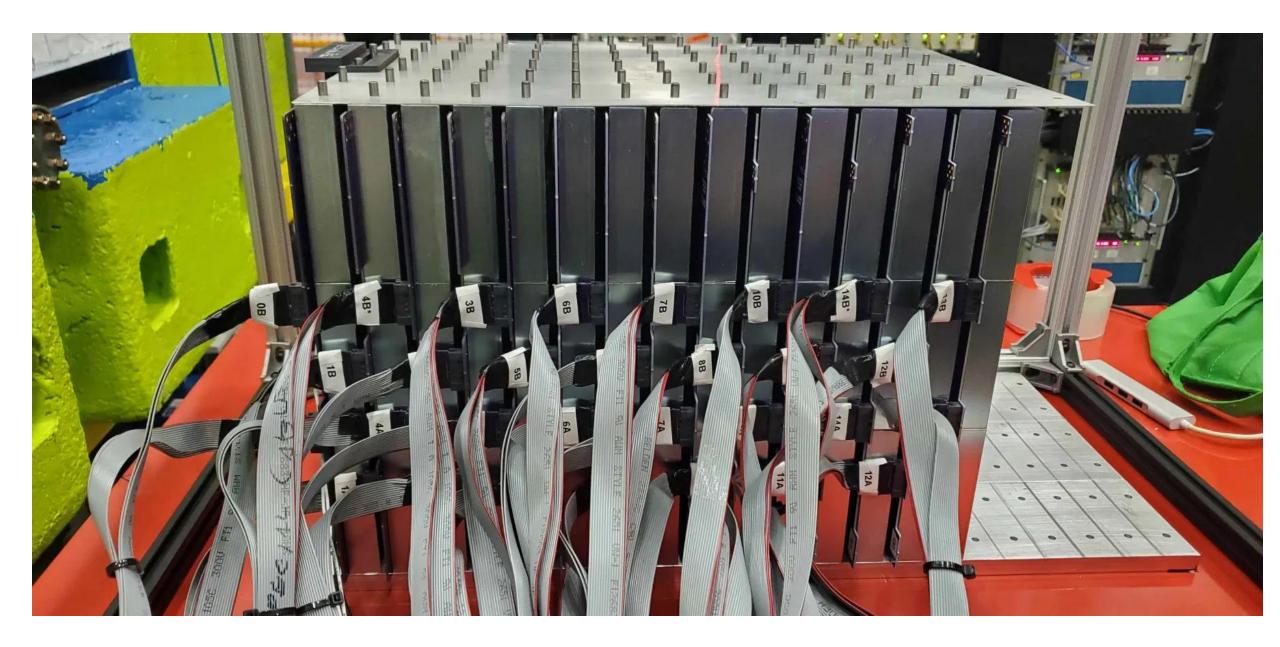
DAQ System

- PCBs send signals to 6 CAEN DT5202 units using custom soldered ribbon cables
- Signals are then collected by a CAEN DT5215 Concentrator Board via fiber optic cables
- Data is collected on a laptop being controlled remotely via VNC Viewer
- System is triggered by an external trigger system at the front center of the module





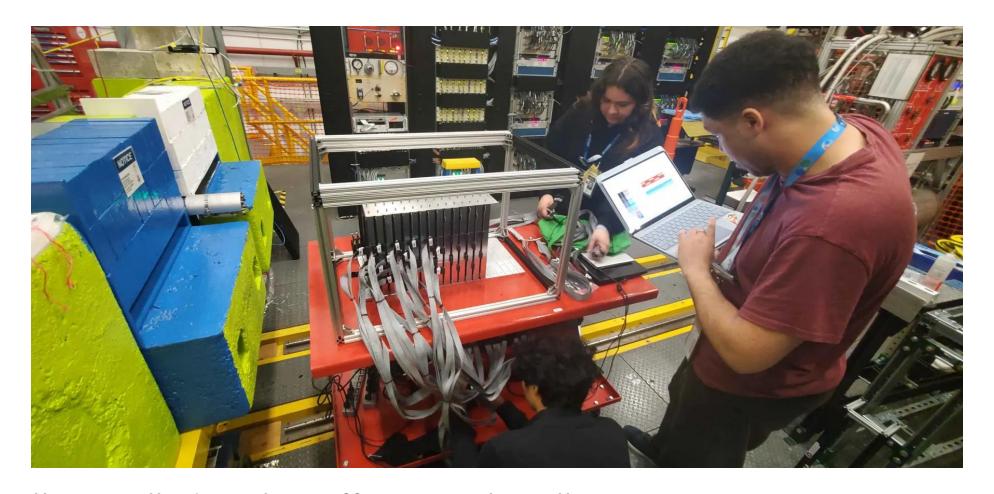




Trigger System

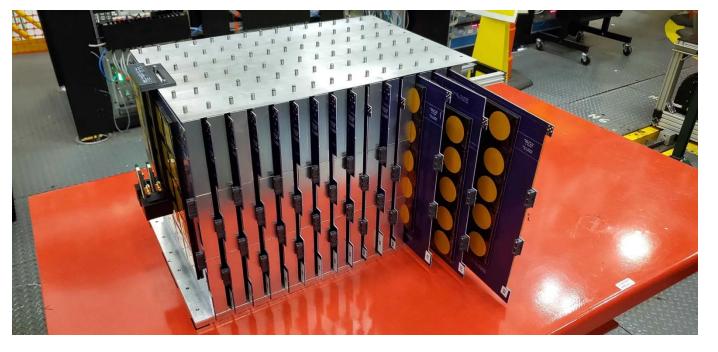
- Trigger system consists of two thin scintillating strips air-coupled to SiPMs on SENSL boards
- A DRS4 digital oscilloscope is set to trigger on AND logic for the two boards with a threshold of 5 mV.
- System is held at the front center of the prototype using a 3D printed mount





- Initially installed at the Jefferson Lab Hall D pair spectrometer on February 6
- Final adjustments made on April 23
- Positioned 5.4 degrees from the beamline, estimated to have received 5 GeV positrons



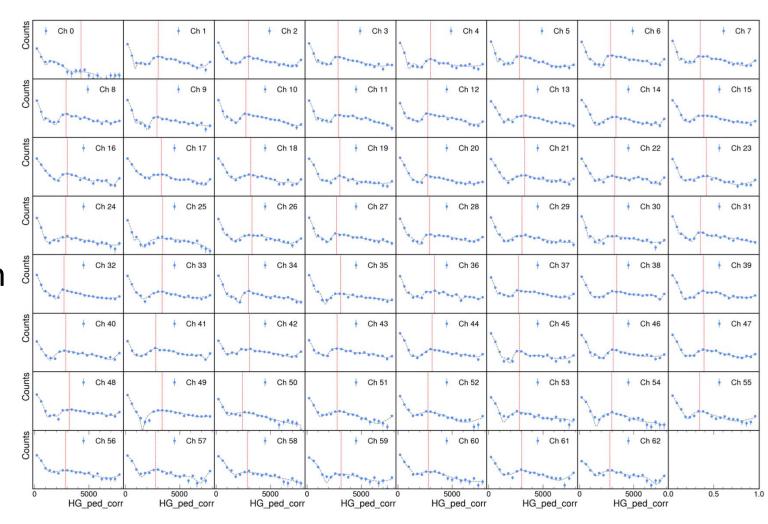


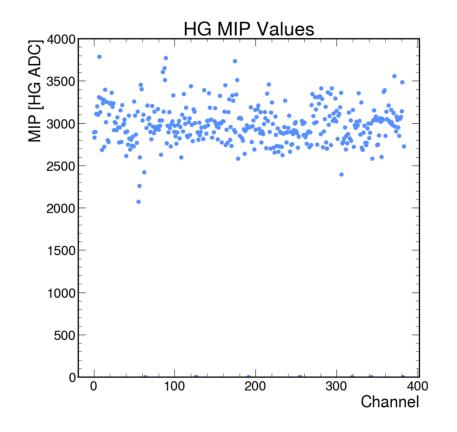


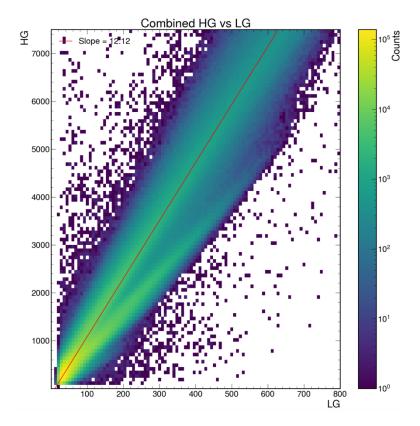
- Aligned using a laser system
- Module housed in a black-out fabric dark box
- Collected 15.7 million events
- Recorded with 14 different setting configurations, 6M events collected with one set

Calibration

- Devoted pedestal runs were recorded for each setting used
- 12 hours of cosmic ray data were collected before the beam was delivered
- The MIP scale in the simulation was determined similarly







 After pedestal subtraction, the LG readout of each channel is brought to the MIP scale using

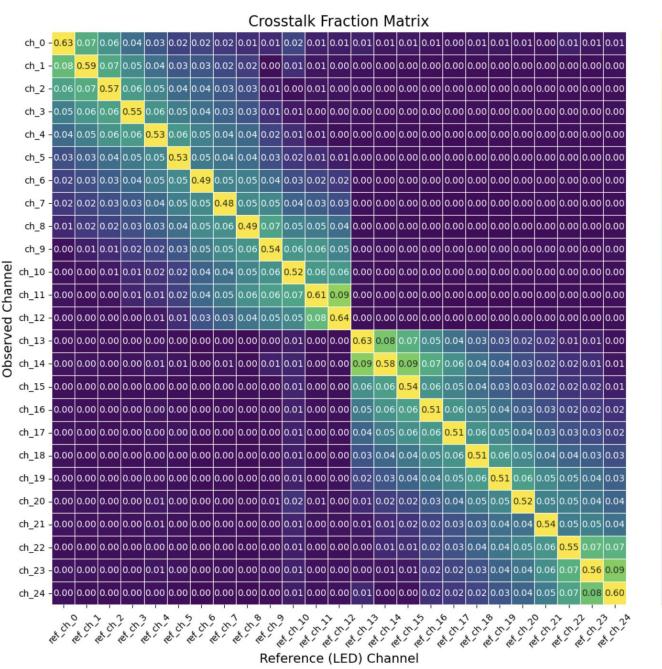
• Factor of 0.8 was applied to the MIP calibration to agree with simulation, likely from large angular distribution of cosmic rays

Electrical Cross-Talk

 After a careful inspection of the pedestal region and HG/LG ratios, we found electrical cross-talk to have a very significant effect

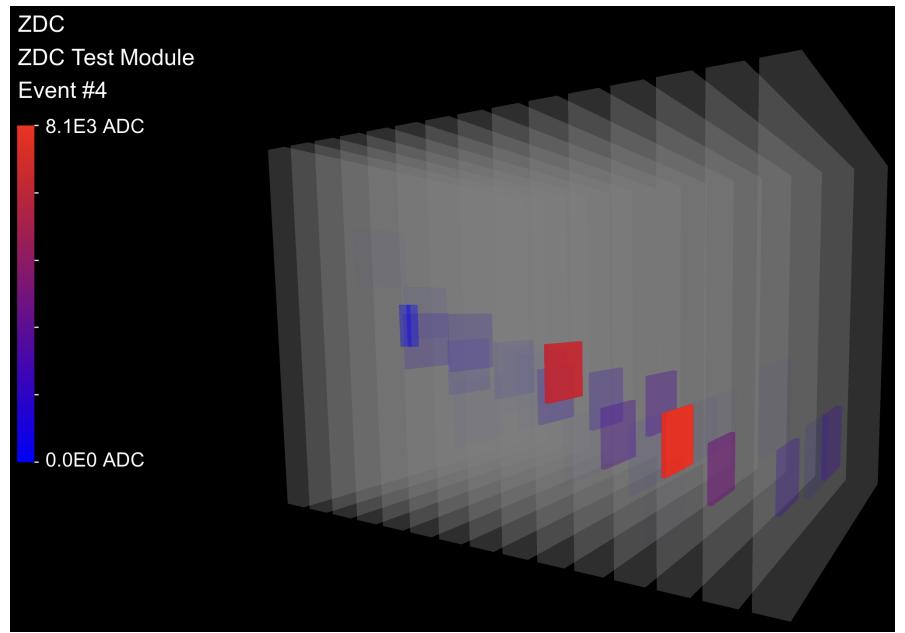
• A "map" of cross-talk effects was determined from bench-top LED tests, imposed on simulation

 The unshielded ribbon cables were found to be the source



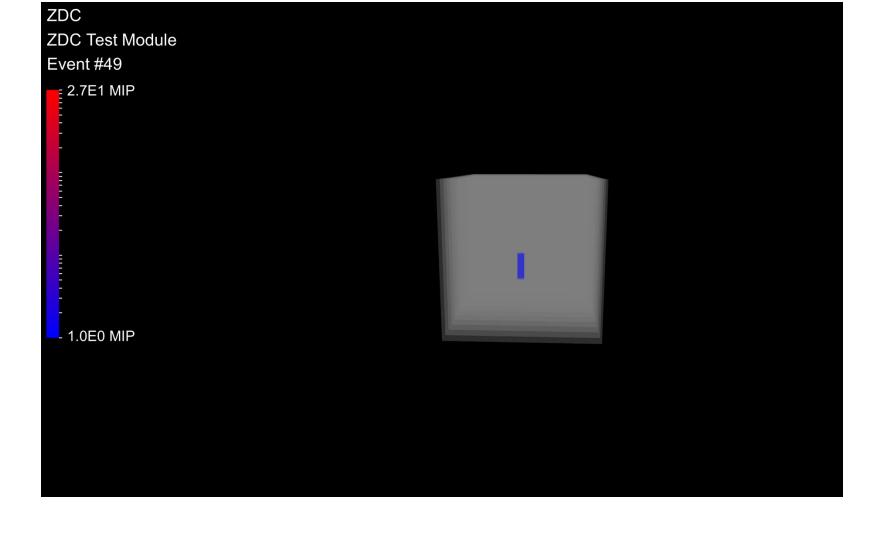
- An LED light source was pulsed into each channel to determine what fraction of the signal is distributed into the other channels
- The other channels were covered with black out fabric to prevent optical cross-talk
- When the ribbon cable was peeled apart, the cross-talk entirely went away





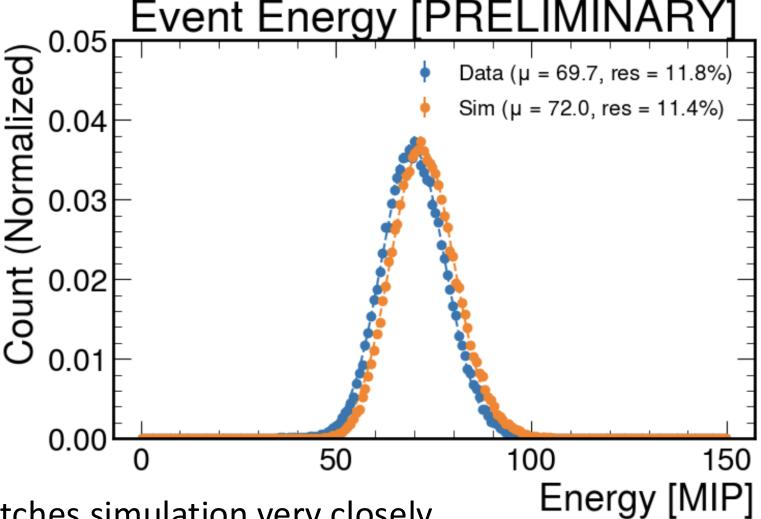
 Cosmic ray track reconstruction verifies event building

Preliminary Results

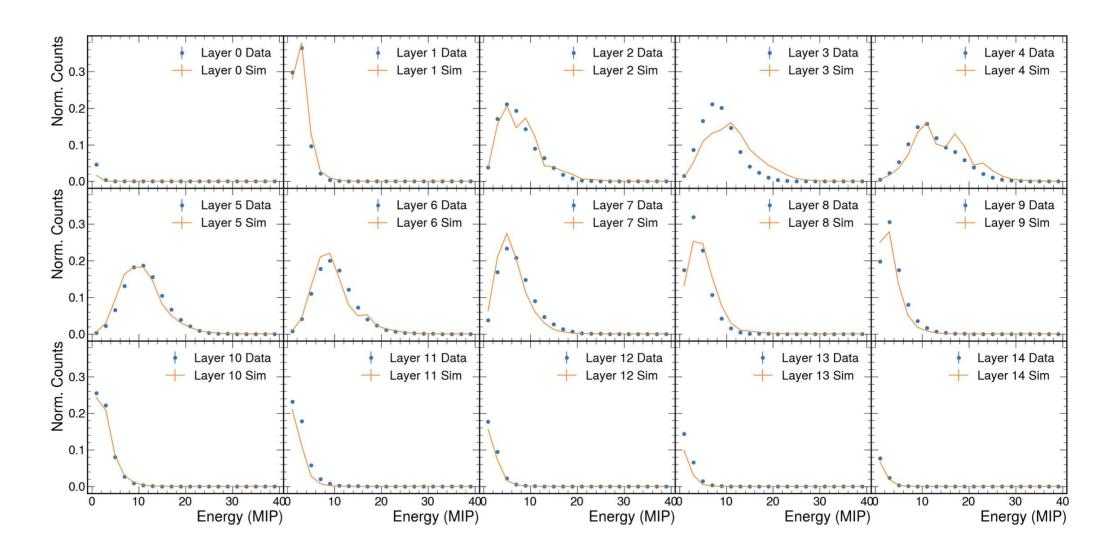


- Results are compared with a DD4HEP simulation of 5 GeV positrons
- 1.0 MIP hit energy cut applied to both simulation and data
- Initial estimate of 3-degree incident angle, 1 cm vertical shift

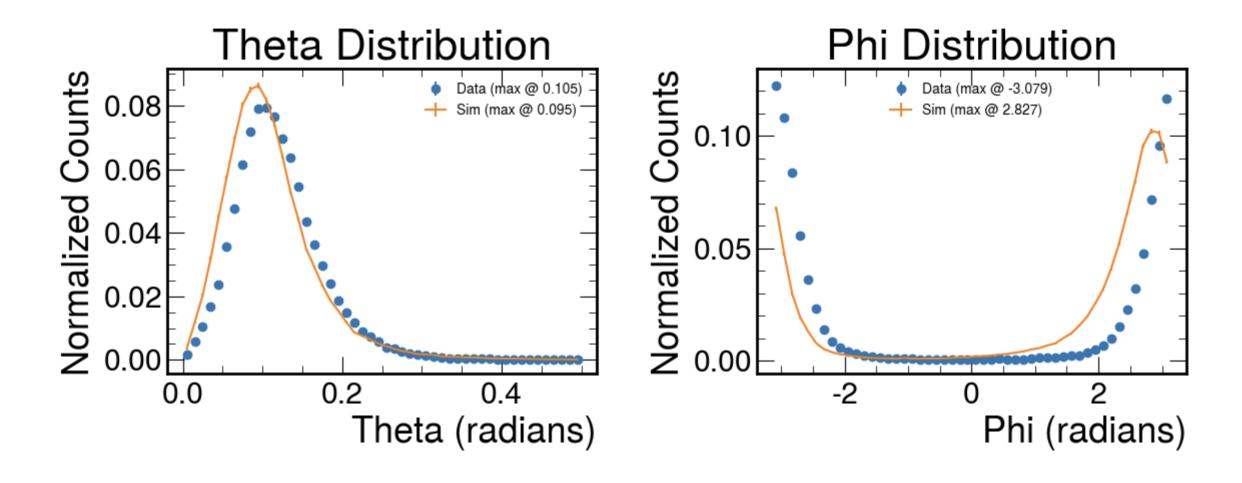
Preliminary Results



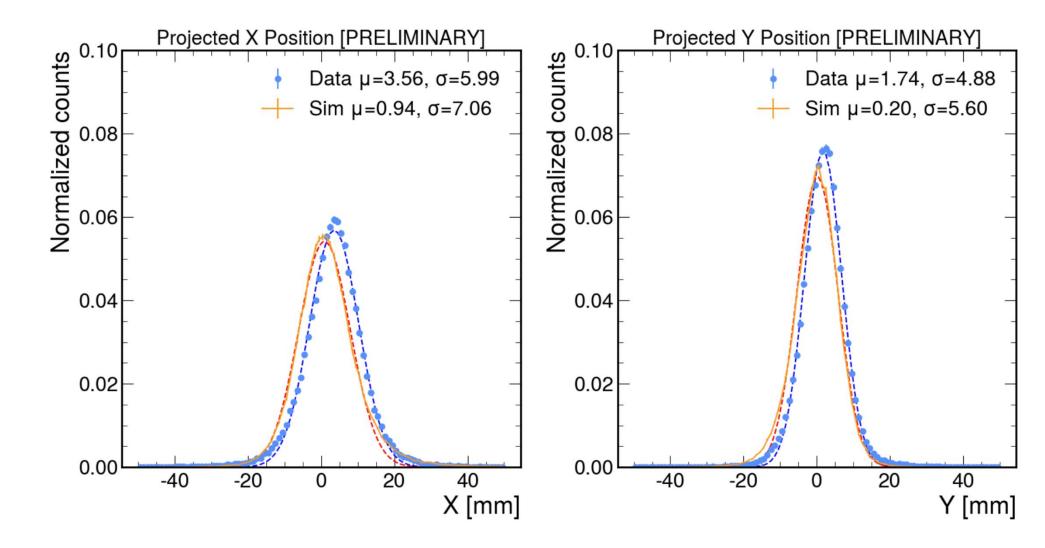
- Energy resolution matches simulation very closely
- Extra MIP calibration factor of 0.8 needs to be verified



• Layer energy distributions match reasonably well, need to inspect layer 3 more



- Theta distribution used to determine a 3-degree incident angle
- Phi angle can be tuned further in the simulation



- Horizontal trigger window / vertical beam spread effects subtracted in quadrature from the data
- Position resolution agrees to ~1 mm

Next Steps

- Implement HEXPLIT clustering algorithm to take advantage of the staggered layer design of the detector
- Simulate angular distribution of cosmic rays for proper MIP calibration in the simulation
- Analyze full dataset, analyze different settings configurations
- Design final ZDC PCB
 - Connectors will be routed out of the top instead of the side
 - Each layer will be three vertical strips instead of four quadrants