Chiral vortical effect for bosons

Andrey V. Sadofyev

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Stony Brook, August, 2017

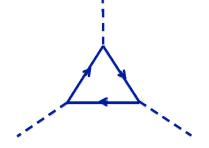
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$$\partial_\mu ar{\psi} \gamma^\mu \gamma_5 \psi = rac{1}{2\pi^2} {\sf E} \cdot {\sf B}$$

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In medium it results in

$$\mathcal{J}' = \sigma_B^V B_\mu + \sigma_\omega^V \omega_\mu \quad , \quad J_\mu^A = \sigma_B^A B_\mu + \sigma_\omega^A \omega_\mu$$

 $\sigma_B^V = \frac{\mu_A}{2\pi^2} \quad , \quad \sigma_\omega^V = \frac{\mu_V \mu_A}{\pi^2}$
 $\sigma_B^A = \frac{\mu_V}{2\pi^2} \quad , \quad \sigma_\omega^A = \left(\frac{\mu_V^2 + \mu_A^2}{2\pi^2} + \frac{T^2}{6}\right)$

and the conductivities appear to be pretty universal.

Let's concentrate on the CVE which gives an additional contribution to the axial charge:

$$J_5^\mu = \sigma^A_\omega \omega^\mu ~,~~ Q_{fh} \sim \int \sigma^A_\omega \left(\mathbf{v} \cdot \Omega
ight) d^3 x \,,$$

which survives in the absence of the EM fields

and we naively allow $Q_{fh} o Q_5^{(0)}$.

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• A simple explanation may be obtained in an EFT description¹

$$\delta S = \int \mu \bar{\psi} \gamma^0 \psi d^4 x \Rightarrow \int \mu u_\mu \bar{\psi} \gamma^\mu \psi d^4 x$$

where we introduced a slowly varying boost field.

• Then, the axial charge conservation is modified

$$\partial_t \left(Q_5^{(0)} + \frac{C_a}{2} \int_x A \cdot B + \int_x \sigma_B^A \mathbf{v} \cdot B + \int_x \sigma_\sigma^A \mathbf{v} \cdot \Omega \right) = \mathbf{0} \,.$$

where helicities are known to correspond to the linkage of the flow and/or field lines.

¹AS, V.I. Shevchenko, V.I. Zakharov, 2011

 If the generalized axial charge picture is valid it may considerably enrich the dynamics, for instance through the new instabilities¹

$$Q_5^{(0)} \to \mathcal{H}_{mh}, \mathcal{H}_{fmh}, \mathcal{H}_{fh}$$

However it is NOT obvious that there is a unified conservation.
 Indeed, in ideal hydrodynamics all helicities are conserved separately

$$\sigma_E \to \infty \ , \ \eta \to 0 \Longrightarrow \partial_t \mathcal{H} = 0$$

while one has to suggest a microscopic mechanism to transfer the medium motion to fermionic modes in the presence of the dissipation.

¹A. Avdoshkin, V. Kirilin, AS, V.I. Zakharov, 2014

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• Recently we found an example of such a mechanism¹ with an intermediate generation of \mathcal{H}_{mh}

$$\int \sigma_{\omega}^{\mathcal{A}} \mathbf{v} \cdot \mathbf{\Omega} \, d^{3}x \to \int \mathcal{A} \cdot \mathcal{B} d^{3}x \to \mathcal{N}_{\mathcal{A}}$$

The first step may be seen as a consequence of a novel chiral effect

 CVE for photons

$$\langle K^{\mu}
angle \sim T^2 \omega^{\mu} \,,$$

which results in a helicity transfer along the vorticity.

¹A. Avkhadiev, AS, 2017

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But let's firstly review the thermal contribution to the common CVE (tCVE). In holography it is related to the **gravitational axial anomaly**¹:

$$d = 5 \qquad d = 4$$

$$\chi \int \epsilon_{MNPQR} A^{M} R_{A}^{BNP} R_{B}^{AQR} \implies \nabla_{\mu} J_{5}^{\mu} \sim \chi R \tilde{R}$$

where χ controls the anomalous coefficient.

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$$J^{\mu}_5\sim\chi {\cal T}^2\omega^{\mu}$$

¹K. Landsteiner et al., 2011

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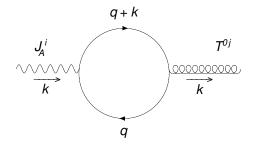
▶ 《母》 《ヨ》 《王》 王 のへの Stony Brook, August, 2017 8 / 26 However the relation is less transparent on the field theory side.

 Indeed, the gravitational anomaly appears to be of higher order in derivatives:

and cannot play the same role as the axial anomaly in EM fields at least within hydrodynamics.

• In the weakly interacting limit, the tCVE can be obtained by the Kubo formula with gravitational perturbation $g_{0i} \sim v_i$:

$$J^{\mu}_{A}(x) = \sigma^{A}_{\omega}\omega^{\mu} \quad , \quad \sigma^{A}_{\omega} = \lim_{q \to 0} \frac{-i}{q_{k}} \epsilon_{ijk} \left\langle J^{i}_{A} T^{0j} \right\rangle |_{\omega = 0}$$



K. Landsteiner et al., 2011

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, $\sigma^{A}_{\omega} = \lim_{q \to 0} \frac{-i}{q_{k}} \epsilon_{ijk} \left\langle J^{i}_{A} T^{0j} \right\rangle |_{\omega=0}$

 The coefficient in front of the gravitational anomaly and the coefficient in the conductivity coincide

$$\nabla_{\mu}J_{5}^{\mu} = \frac{\#}{768\pi^{2}} \cdot \epsilon^{\mu\nu\rho\lambda} R^{\alpha}_{\beta\mu\nu} R^{\beta}_{\alpha\rho\lambda} \quad , \quad J_{5}^{\mu} = \frac{\#}{12} T^{2} \omega^{\mu}$$

K. Landsteiner et al., 2011

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Despite the discussion mentioned above there is **one particular** field theoretical **setup** where the anomalous origin of the tCVE is transparent and it requires two ingredients:

• the tCVE as a P-odd contribution to the Hawking radiation¹:

$$\frac{dN}{dtd\omega do} \sim \frac{\omega^2}{T^2} \left(1 - \frac{L\Omega}{4T} \cos \theta \right)$$

• the relation of the Hawking radiation with the anomalies of the effective 1 + 1d theory at the horizon² (the chiralness of the dimensionally reduced theory at the horizon).

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¹A. Vilenkin, 1978-1979

²S. Robertson, F. Wilczek, 2006

- The tCVE can be seen as a polarization effect in a rotating thermal radiation.
- It has an analog in the Hawking radiation of a rotating BH.
- It is related to the gravitational anomaly in the holographic considerations.
- There is a connection of the tCVE and the global anomalies.

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One expects polarization effects for any non-zero spin.

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- It has an analog in the Hawking radiation of a rotating BH. The spin-gravity interaction is "trivial" and it is natural to expect a similar effect for $s \neq \frac{1}{2}$
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- The spin-gravity interaction is "trivial" and it is natural to expect a similar effect for $s \neq \frac{1}{2}$
- It is related to the gravitational anomaly in the holographic considerations.

There are well known examples of gravitational anomalies for theories with $s>\frac{1}{2}$

• There is a connection of the tCVE and the global anomalies.

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Gravitational anomaly for photons¹:

$$\begin{split} F'_{\mu\nu} &= F_{\mu\nu} \cos \alpha + \tilde{F}_{\mu\nu} \sin \alpha \\ & \downarrow \\ K^{\mu} &= \frac{1}{\sqrt{-g}} \epsilon^{\mu\nu\alpha\beta} A_{\nu} \partial_{\alpha} A_{\beta} \ , \ \nabla_{\mu} K^{\mu} = \frac{1}{2} F \tilde{F} \ , \ \langle \nabla_{\mu} K^{\mu} \rangle_{\text{naive}} = 0 \\ & \downarrow \\ \langle \nabla_{\mu} K^{\mu} \rangle_{\text{triangle}} = -\frac{1}{96\pi^{2}} R \tilde{R} \end{split}$$

¹A.D. Dolgov et al., 1988 Andrey V. Sadofyev (MIT)

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- One expects polarization effects for any non-zero spin.
- The spin-gravity interaction is "trivial" and it is natural to expect a similar effect for $s \neq \frac{1}{2}$
- There are well known examples of gravitational anomalies for theories with $s > \frac{1}{2}$
- There is a connection of the tCVE and the global anomalies. The relation of some chiral effects to the global anomalies could be extended to $s \neq \frac{1}{2}$ (some examples for $d \neq 4$ in the literature).

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- There are well known examples of gravitational anomalies for theories with $s > \frac{1}{2}$
- The relation of the tCVE to the global anomalies could be extended to s ≠ ¹/₂ (some examples for d ≠ 4 in the literature¹).

¹S.D. Chowdhury, J.R. David, 2016

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$$J_{A}^{\mu} = C \epsilon^{\mu\nu\alpha\beta} A_{\nu} \partial_{\alpha} A_{\beta}$$

$$\Downarrow$$

$$J_{A}^{\mu} = \frac{T^{2}}{6} \left(1 + \frac{e^{2}}{4\pi^{2}} \right) \omega^{\mu}$$

D.-F. Hou at el., 2012; S. Golkar, D. T. Son, 2012

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Combining these insights one expects a bosonic analogue of CVE (bCVE)



$K^{\mu} \sim T^2 \omega^{\mu}$

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and it is supported by a simple direct calculation¹.

¹ A. Avkhadiev, AS, 2017			500
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Combining these insights one expects a bosonic analogue of CVE (bCVE)

$$\sigma_{b} = \lim_{q \to 0} \frac{-i}{q_{k}} \epsilon_{ijk} \left\langle K_{A}^{i} T^{0j} \right\rangle |_{\omega=0}$$
$$K^{\mu} = \epsilon^{\mu\nu\alpha\beta} A_{\nu} \partial_{\alpha} A_{\beta}$$
$$T^{\mu\nu} = F^{\mu\lambda} F^{\nu}_{\ \lambda} - g^{\mu\nu} \frac{1}{4} F^{2}$$

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• This effect gives a microscopic mechanism for $\mathcal{H}_{fh}
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 $\partial_t H_{mh} \propto \partial_\mu K^\mu \propto \partial_\mu T^2 \omega^\mu$

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 While the tCVE relation to the anomalies is under discussion, it is clear that the bCVE and tCVE have the same origin (polarization by Ω, BH radiation, g. anomaly, global anomalies, etc.).

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- While the tCVE relation to the anomalies is under discussion, it is clear that the bCVE and tCVE have the same origin (polarization by Ω, BH radiation, g. anomaly, global anomalies, etc.).
- Photons can have non-trivial topological phase¹ and one may think about a generalization of the chiral kinetic theory².

¹V. Liberman et al., 1992; K. Bliokh et al., 2004; M. Onoda et al., 2004;

²N. Muller, R. Venugopalan, 2017; N. Yamamoto, 2017; → <♂→ < ≧→ < ≧→ · ≧

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- Photons can have non-trivial topological phase¹ and one may think about a generalization of the chiral kinetic theory².
- Possible **phenomenological applications** in condensed matter, primordial plasma, QGP, etc.

¹V. Liberman et al., 1992; K. Bliokh et al., 2004; M. Onoda et al., 2004; ²N. M. H. D. Y. (2017) N. Y. (2017) N. Y. (2017)

²N. Muller, R. Venugopalan, 2017; N. Yamamoto, 2017; A. Karamoto, 2017; A. Karamato, 2017; A. Karamoto, 20

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