

The BEST plan for lattice QCD

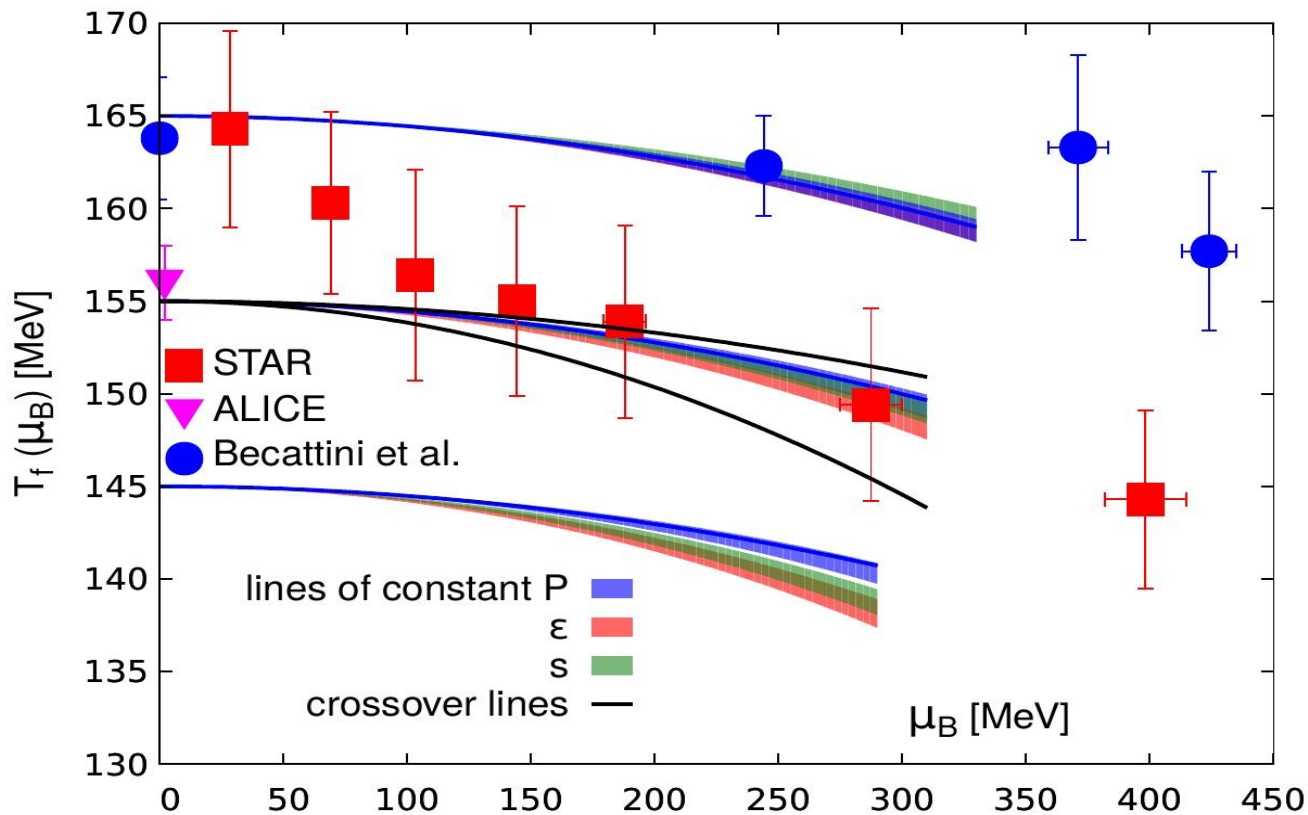
Frithjof Karsch

Brookhaven National Laboratory & Bielefeld University

- characterize bulk thermodynamics and fluctuations of conserved charges on the freeze-out line
- provide the equation of state for BES
- (constrain) the location of the critical point

Chiral transition, hadronization and freeze-out

- pseudo-critical temperature $T_c = 154(9)\text{MeV}$
- hadronization temperatures $T_h = 164(2)\text{ MeV}$
- freeze-out temperatures: $T_{fo} = 156(3)\text{ MeV}$
 $T_{fo} = [164(5) - 168(4)]\text{ MeV}$



Where does hadronization set in?

physics is quite different at lower and upper end of the current error bar on T_c

probed with net-charge correlations & fluctuations

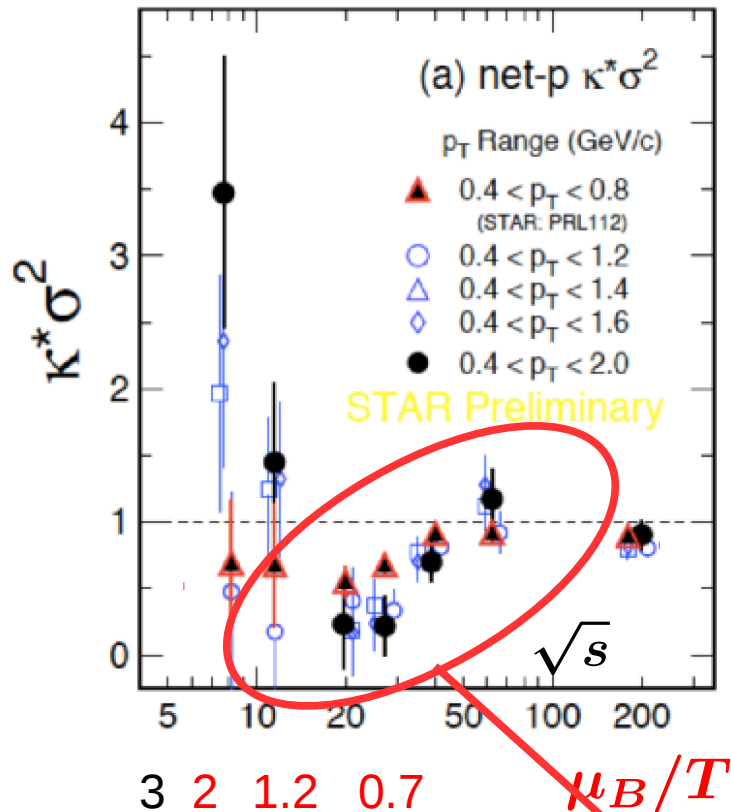
crossover transition lines:

O. Kaczmarek et al., arXiv:1011.3130, C. Bonati et al., arXiv:1507.03571

R. Bellwied et al., arXiv:1507.07510, P. Cea et al., arXiv:1508.07599

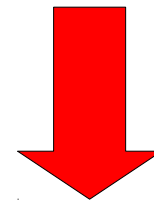
lines of constant physics
A. Bazavov et al.
(BI-BNL-CCNU...),
arXiv:1701.04325

Exploring the QCD phase diagram



More moderate questions:

- Can we understand the systematics seen in cumulants of charge fluctuations in terms of **QCD thermodynamics** ?
- How far do we get with low order Taylor expansions of **QCD** in explaining the obvious deviations from HRG model behavior ?



- For $\sqrt{s} \geq 19.6$ GeV :
Structure of net-electric charge and net-proton cumulants is inconsistent with HRG thermodynamics, but can eventually be understood in terms of **QCD thermodynamics in a next-to-leading order Taylor expansion**

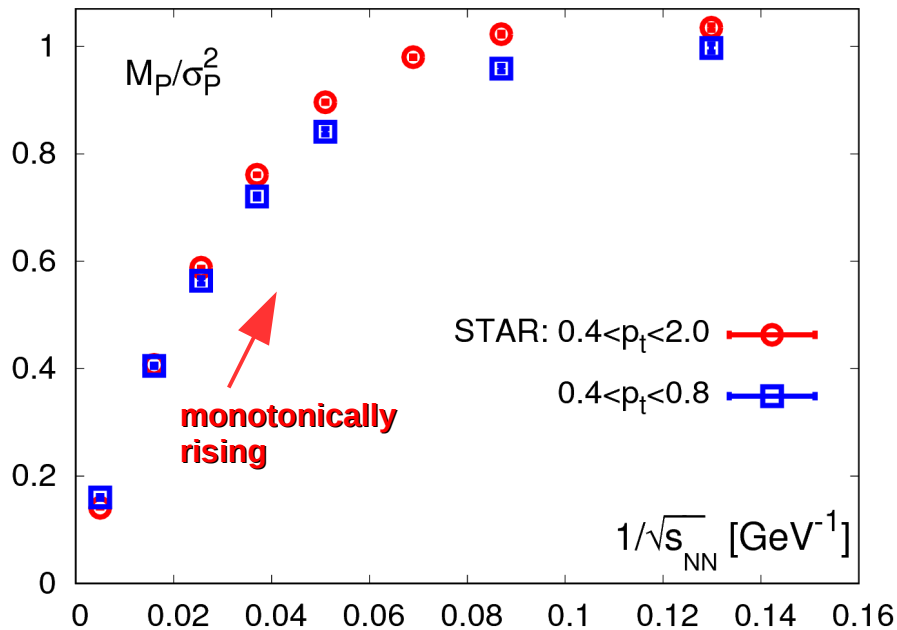
?

Cumulant ratios of conserved net-charge fluctuations

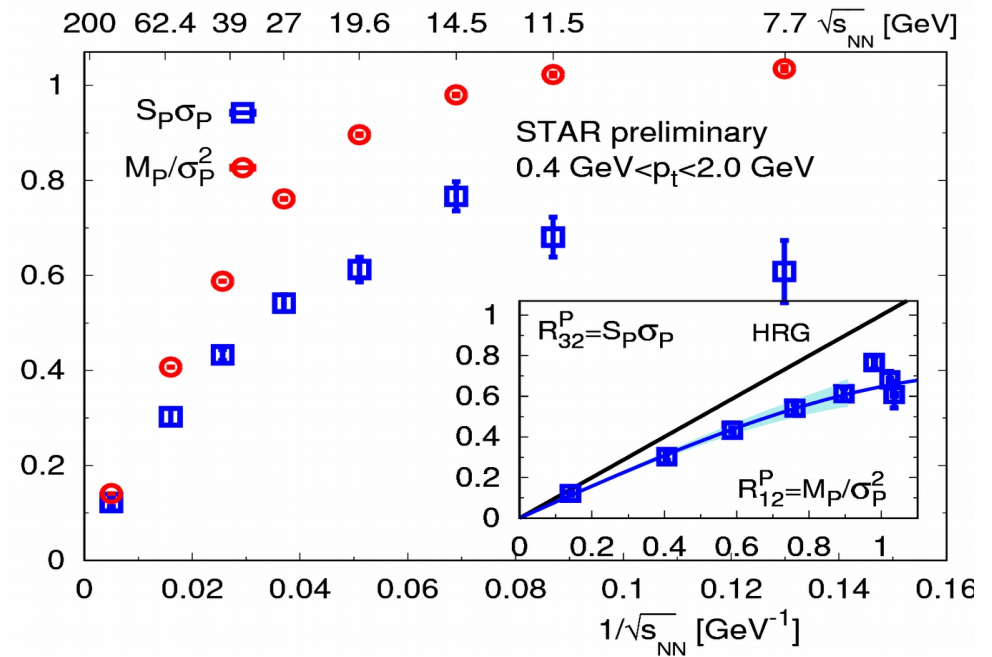
cumulant ratios: $\frac{\chi_n^X}{\chi_m^X}$, $X = B, Q, S$ $(n, m) = (1, 2) : M_X/\sigma_X^2$

$$\chi_n^X = \frac{\partial^n P/T^4}{\partial(\mu_X/T)^n}$$

$(3, 2) : S_X \sigma_X$
 $(4, 2) : \kappa_X \sigma_X^2$



- M_P/σ_P^2 is monotonically rising



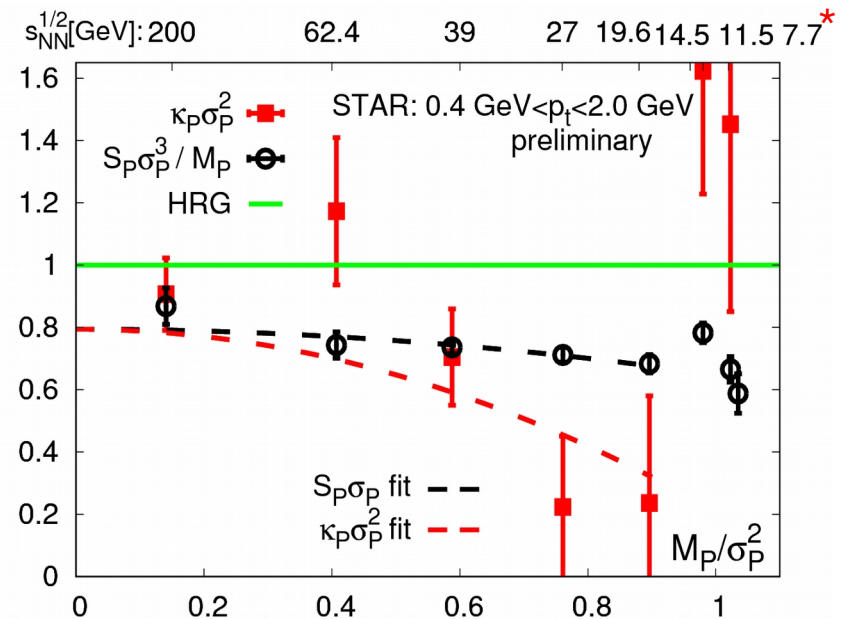
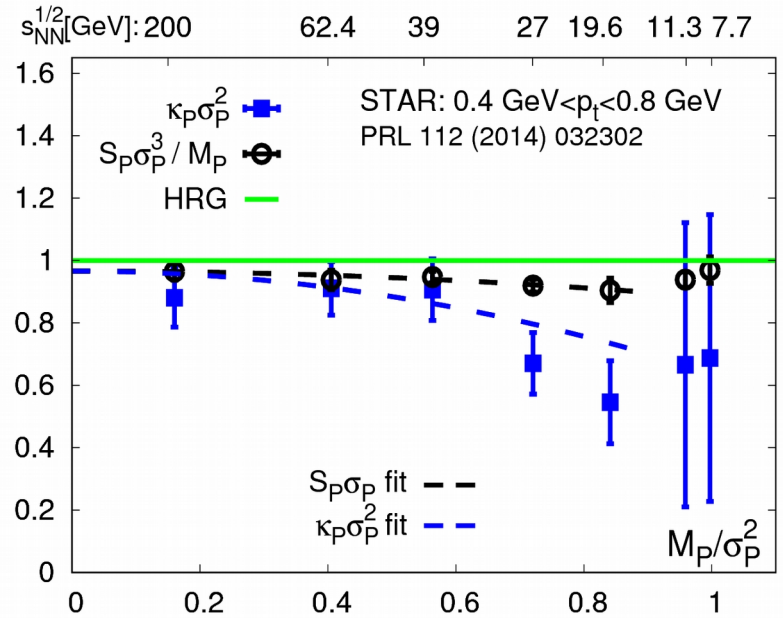
- $S_P \sigma_P < M_P/\sigma_P^2$

replace $\sqrt{s_{NN}}$ in favor of M_P/σ_P^2

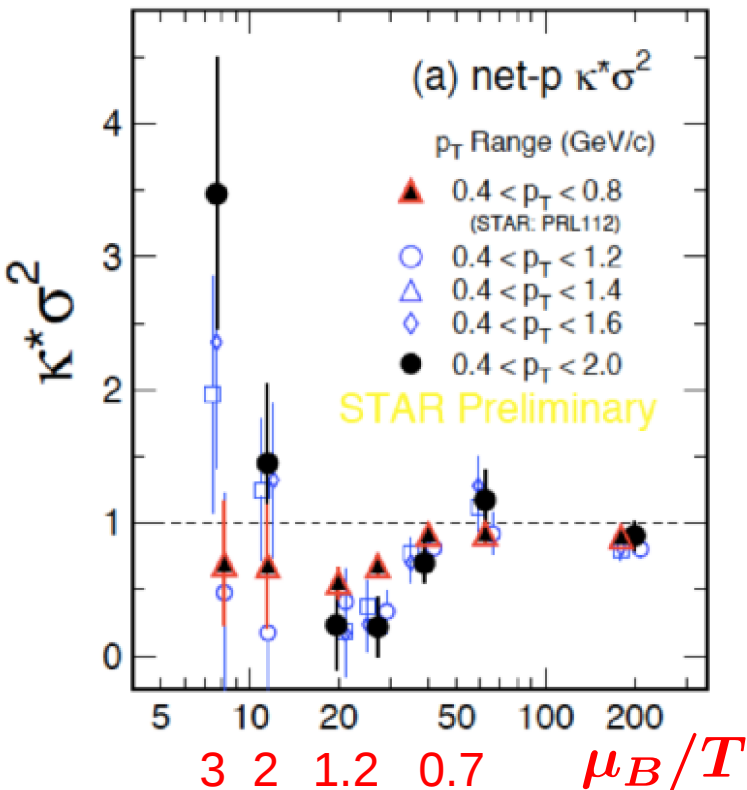
Cumulant ratios of conserved net-charge fluctuations

kurtosis*variance: $R_{42}^X = (\kappa\sigma^2)_X = \frac{\chi_4^X}{\chi_2^X}$

skewness*variance^{1/2}: $R_{32}^X = (S\sigma)_X = \frac{\chi_3^X}{\chi_2^X}$



* not shown



at $\mu_B = 0$

$$(\kappa\sigma^2)_P \simeq S_P\sigma_P^3/M_P$$

slope of $(\kappa\sigma^2)_P$

$$\simeq 3S_P\sigma_P^3/M_P$$

X. Luo (STAR Collaboration),
PoS CPOD2014 (2014) 019

Conserved charge fluctuations and freeze-out

$$\frac{\Delta(T, \mu_B)}{T^4} = \frac{P(T, \mu_B) - P(T, 0)}{T^4} = \frac{\chi_2^B}{2} \left(\frac{\mu_B}{T}\right)^2 \left(1 + \frac{1}{12} \frac{\chi_4^B}{\chi_2^B} \left(\frac{\mu_B}{T}\right)^2\right)$$

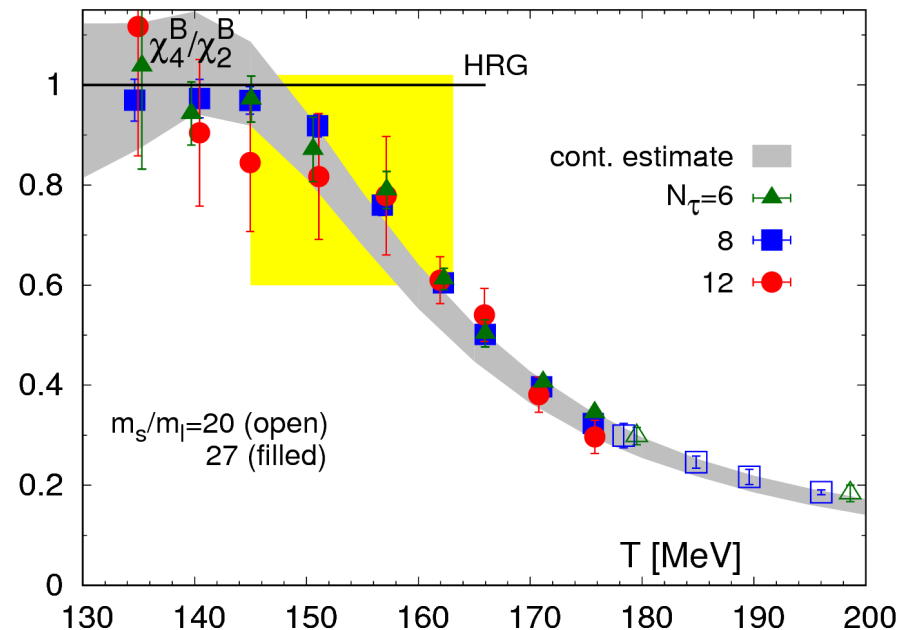
$$\frac{M_B}{\sigma_B^2} = \frac{\mu_B}{T} + \mathcal{O}(\mu_B^3)$$

$$S_B \sigma_B = \frac{\mu_B}{T} \frac{\chi_4^B}{\chi_2^B} + \mathcal{O}(\mu_B^3)$$



no need for talking
about a chemical
potential

$$S_B \sigma_B = \frac{M_B}{\sigma_B^2} \frac{\chi_4^B}{\chi_2^B} + \mathcal{O}(\mu_B^3)$$



formulas are given for the case $\mu_S = \mu_Q = 0$

However, entire analysis is done for $n_S = 0$, $n_Q/n_B = 0.4$

Conserved charge fluctuations and freeze-out

$$\frac{\Delta(T, \mu_B)}{T^4} = \frac{P(T, \mu_B) - P(T, 0)}{T^4} = \frac{\chi_2^B}{2} \left(\frac{\mu_B}{T}\right)^2 \left(1 + \frac{1}{12} \frac{\chi_4^B}{\chi_2^B} \left(\frac{\mu_B}{T}\right)^2\right)$$

$$\frac{M_B}{\sigma_B^2} = \frac{\mu_B}{T} + \mathcal{O}(\mu_B^3)$$

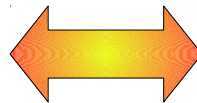
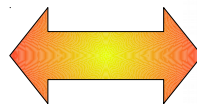
$$S_B \sigma_B = \frac{\mu_B}{T} \frac{\chi_4^B}{\chi_2^B} + \mathcal{O}(\mu_B^3)$$



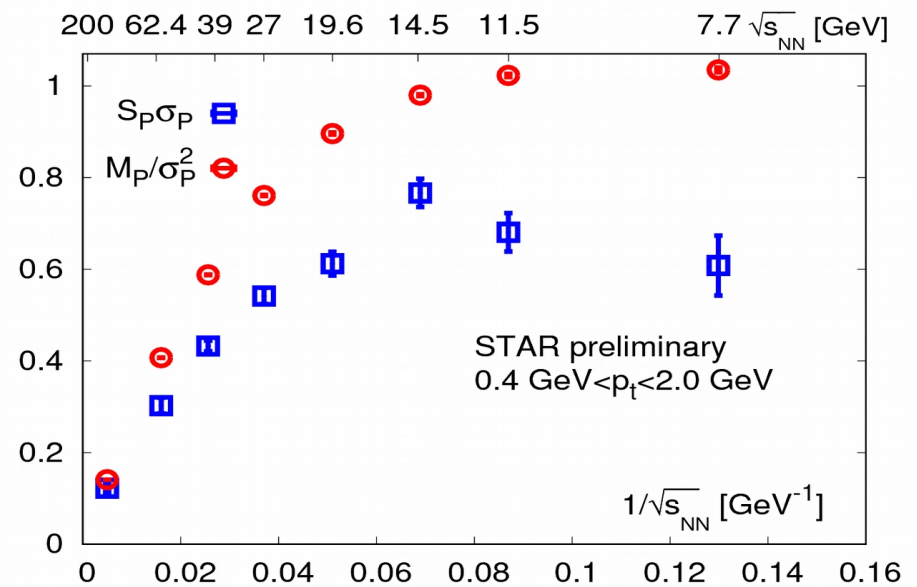
no need for talking about a chemical potential

$$S_B \sigma_B = \frac{M_B}{\sigma_B^2} \frac{\chi_4^B}{\chi_2^B} + \mathcal{O}(\mu_B^3)$$

$$\frac{\chi_4^B}{\chi_2^B} < 1$$



$$S_B \sigma_B < M_B / \sigma_B^2$$



Conserved charge fluctuations and freeze-out mean, variance, **skewness and kurtosis**

in a NLO Taylor expansion $R_{31}^B \equiv S_B \sigma_B^3 / M_B$ } are closely related
 $R_{42}^B \equiv \kappa_B \sigma_B^2$

$$\left. \begin{aligned} R_{31}^B &= r_{31}^{B,0} + r_{31}^{B,2} (R_{12}^B)^2 \\ R_{42}^B &= r_{42}^{B,0} + r_{42}^{B,2} (R_{12}^B)^2 \end{aligned} \right\}$$

$$\underline{\mu_S = \mu_Q = 0 :}$$

$$r_{42}^{B,0} = r_{31}^{B,0} = \frac{\chi_4^B}{\chi_2^B}$$

$$r_{42}^{B,2} = 3r_{31}^{B,2} = \frac{1}{2} \left(\frac{\chi_6^B}{\chi_2^B} - \left(\frac{\chi_4^B}{\chi_2^B} \right)^2 \right)$$

Conserved charge fluctuations and freeze-out mean, variance, **skewness** and **kurtosis**

in a NLO Taylor expansion $R_{31}^B \equiv S_B \sigma_B^3 / M_B$ } are closely related
 $R_{42}^B \equiv \kappa_B \sigma_B^2$

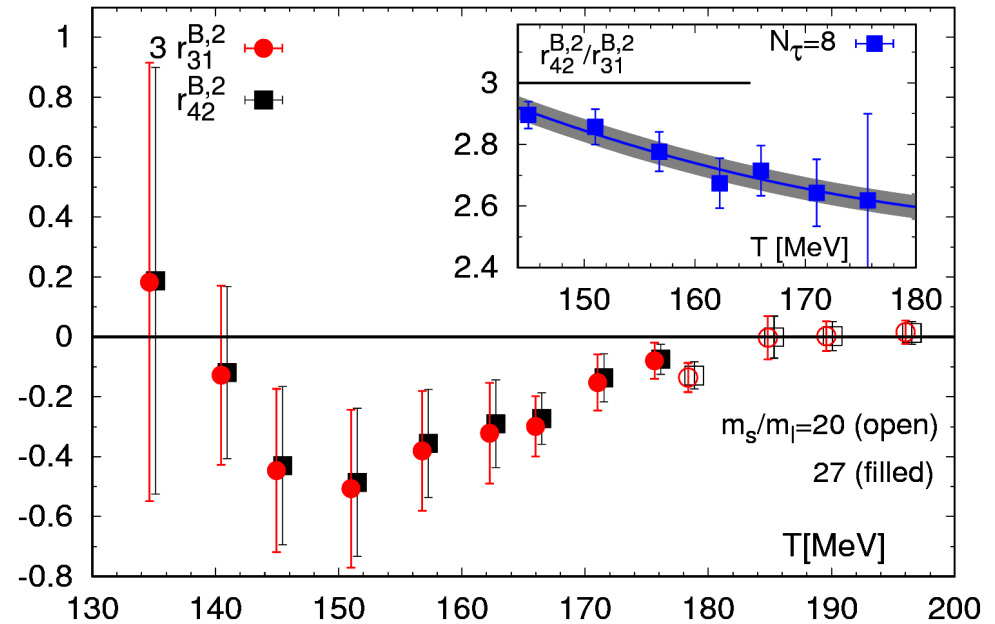
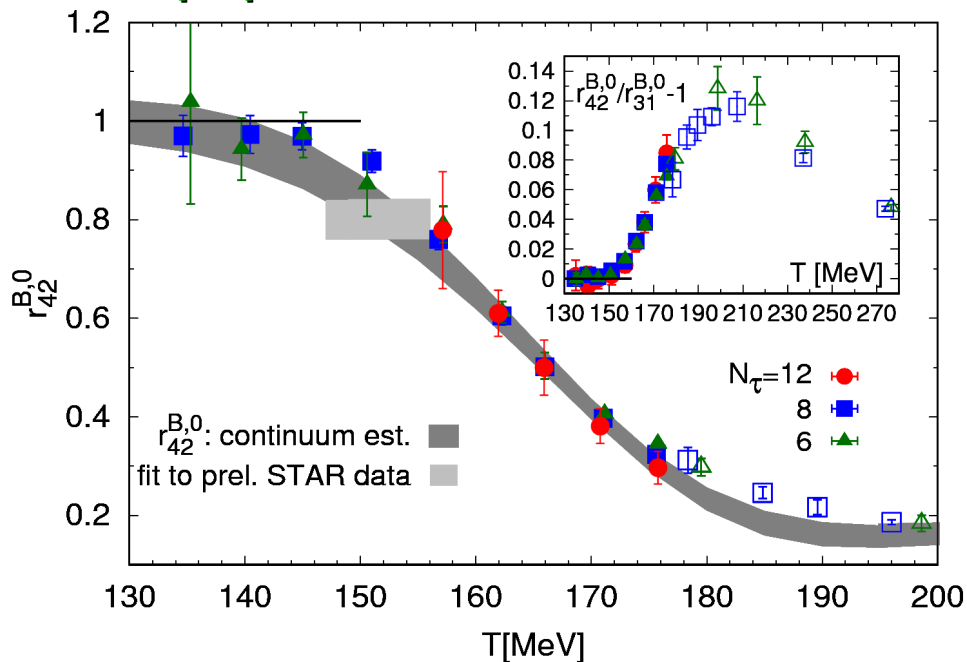
$$\left. \begin{aligned} R_{31}^B &= r_{31}^{B,0} + r_{31}^{B,2} (R_{12}^B)^2 \\ R_{42}^B &= r_{42}^{B,0} + r_{42}^{B,2} (R_{12}^B)^2 \end{aligned} \right\}$$

$n_S = 0, n_Q/n_B = 0.4 :$

$$r_{42}^{B,0} \simeq r_{31}^{B,0}$$

$$r_{42}^{B,2} \simeq 3r_{31}^{B,2}$$

A. Bazavov et al. (HotQCD Collaboration), in preparation

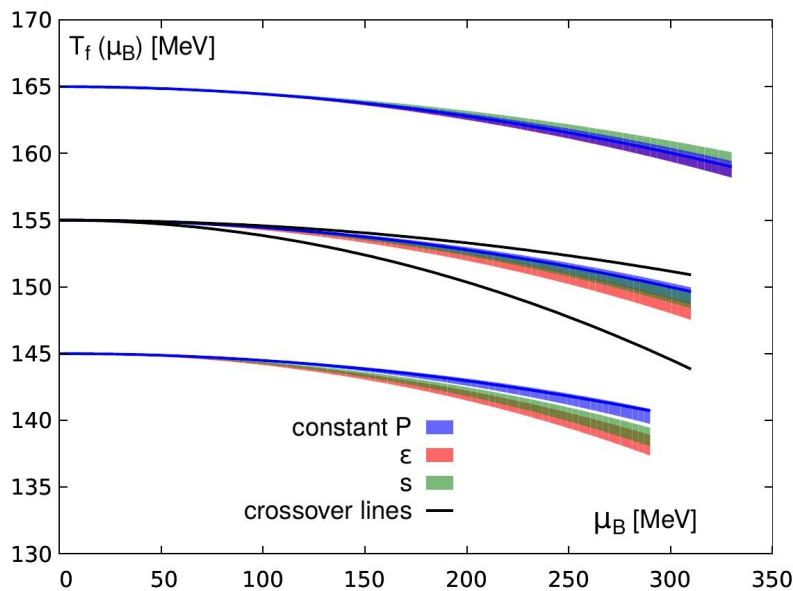


Skewness and kurtosis on a "line of constant physics"

- Temperature on the "freeze-out" line changes with increasing μ_B
- consider cumulant ratios on lines $T_f(\mu_B) = T_f(0)(1 - \kappa_2^f(\mu_B/T_f(0))^2) + \mathcal{O}(\mu_B^4)$
- Taylor expansion in T and μ_B

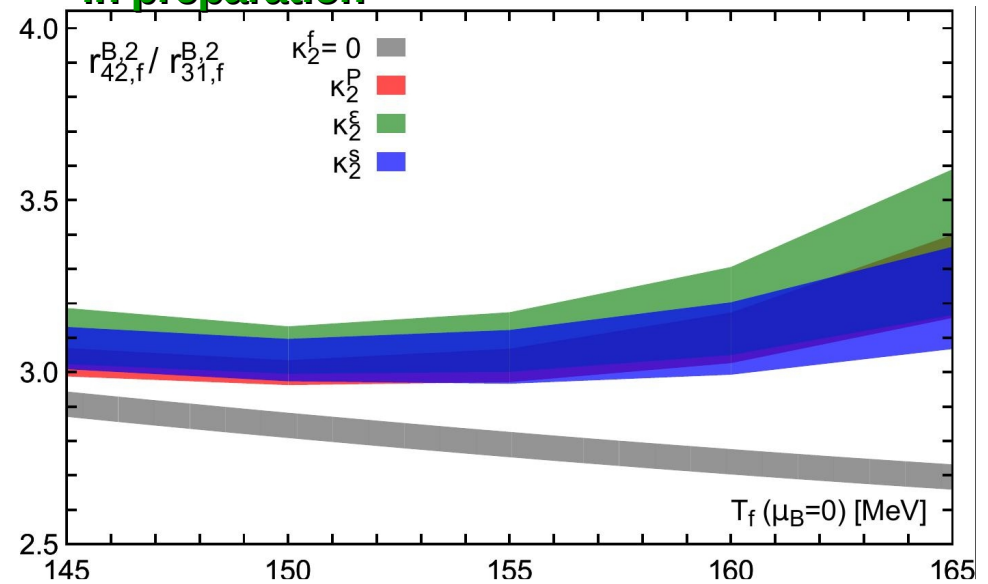
$$r_{nm}^{B,k} \longrightarrow r_{nm,f}^{B,k} \equiv r_{nm}^{B,k}(T_{f,0}) - \kappa_2^f T_{f,0} \left. \frac{dr_{nm}^{B,k-2}}{dT} \right|_{T=T_{f,0}}$$

A. Bazavov et al. arXiv:1701.04325



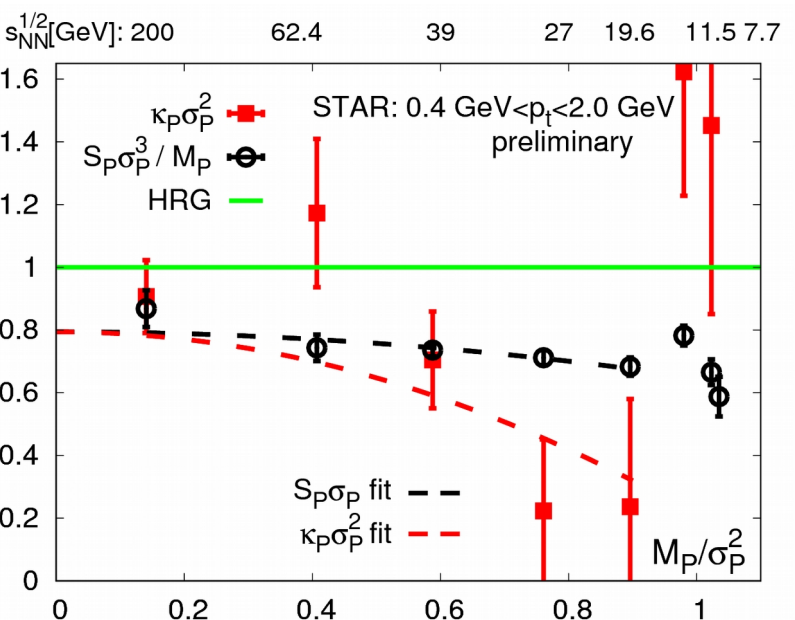
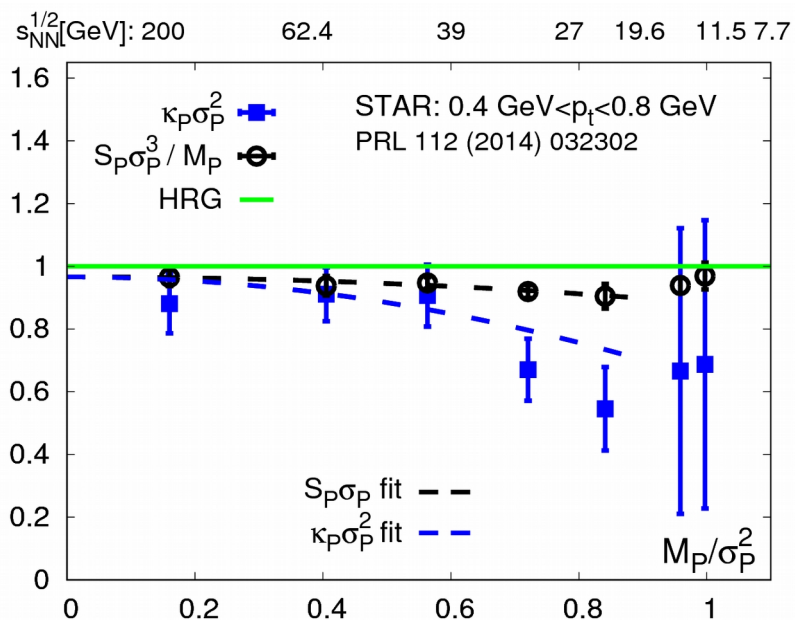
$$0.006 < \kappa_2^f < 0.015$$

A. Bazavov et al. (HotQCD Collaboration), in preparation

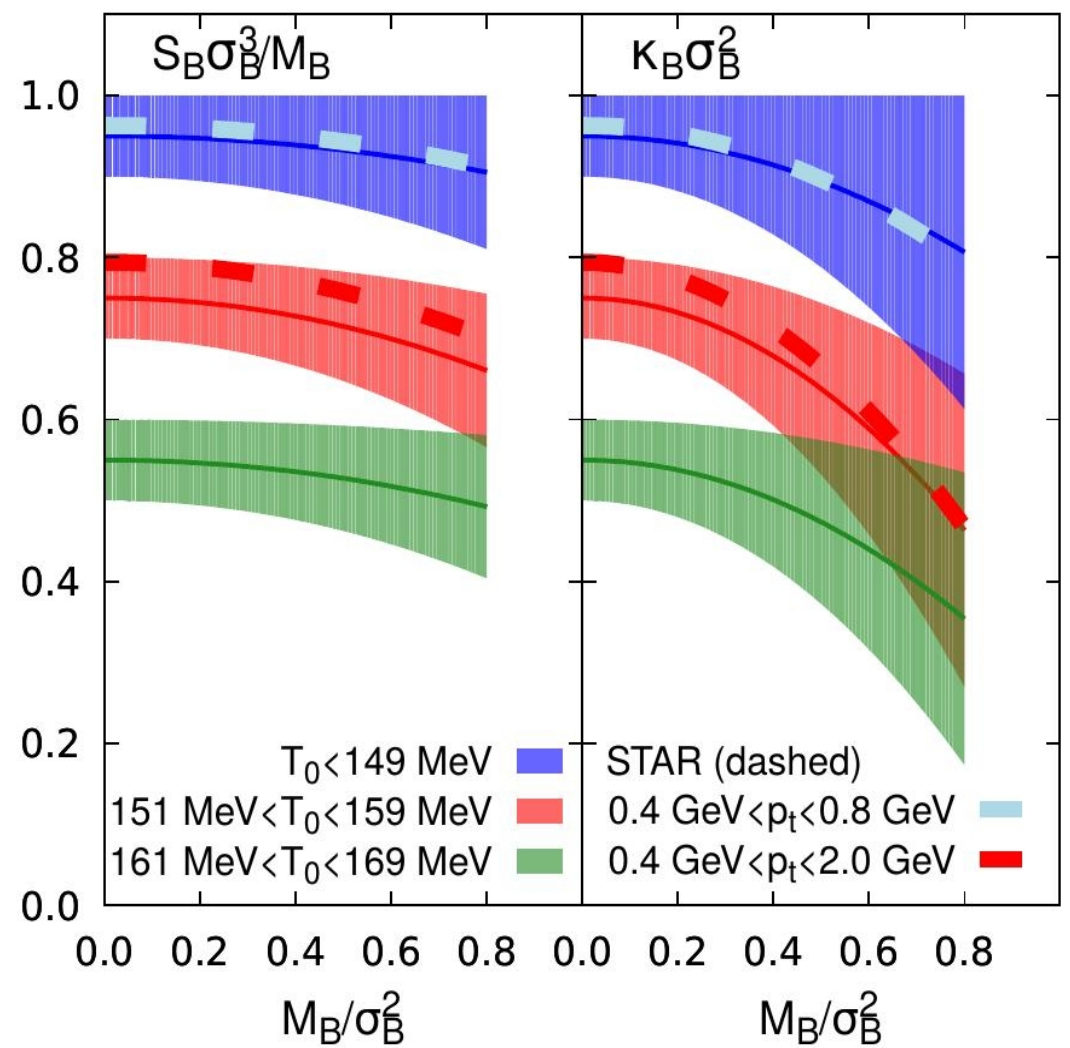


$$\frac{r_{42,f}^{B,2}}{r_{31,f}^{B,2}} = 3 - 3.6$$

Conserved charge fluctuations and freeze-out mean, variance, skewness and kurtosis



NLO lattice QCD
(not yet continuum extrapolated)

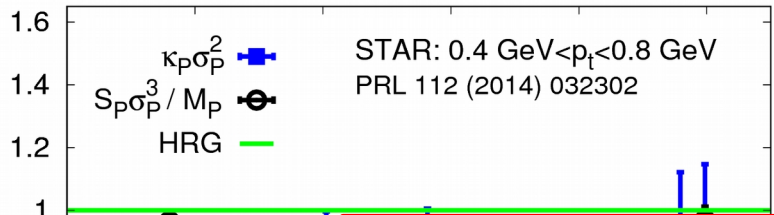


hotQCD in preparation

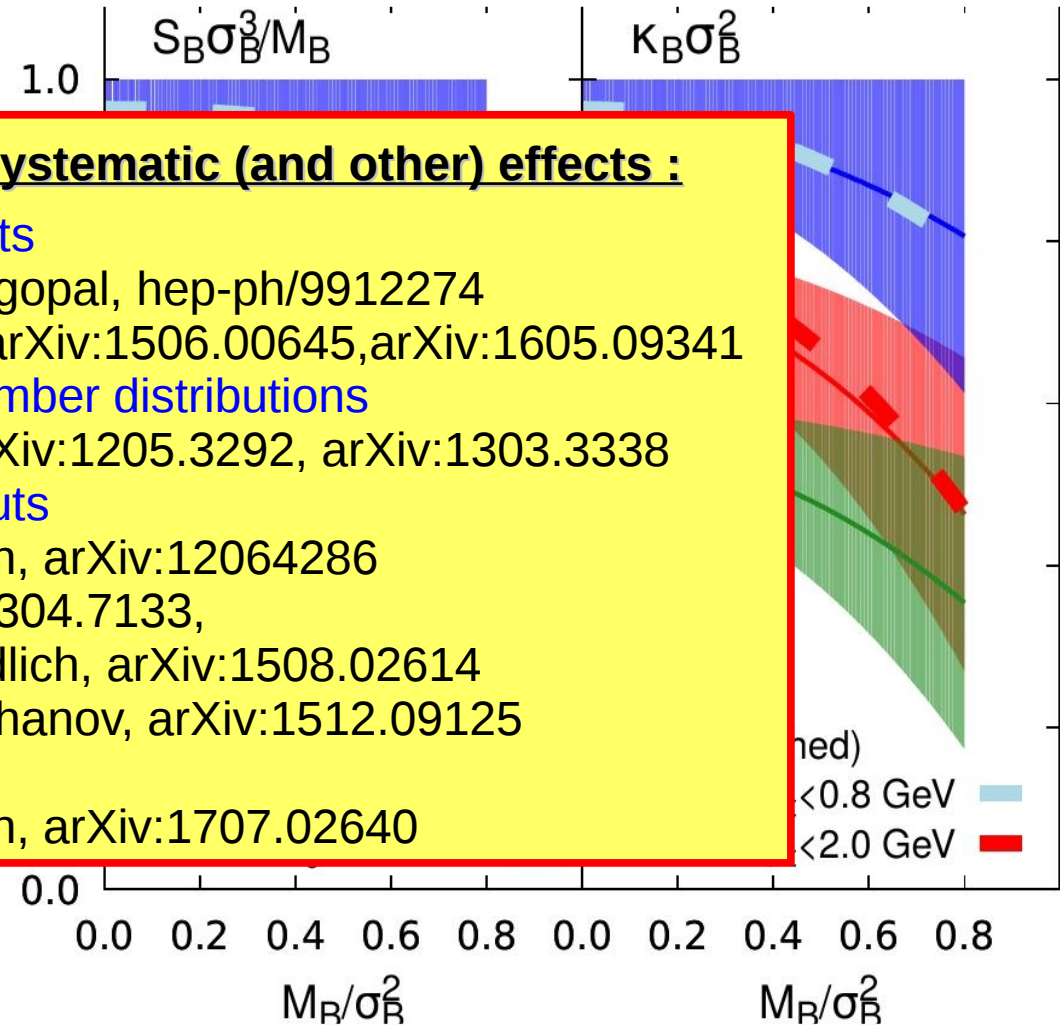
Conserved charge fluctuations and freeze-out

mean, variance, skewness and kurtosis

$s_{NN}^{1/2}$ [GeV]: 200 62.4 39 27 19.6 11.5 7.7

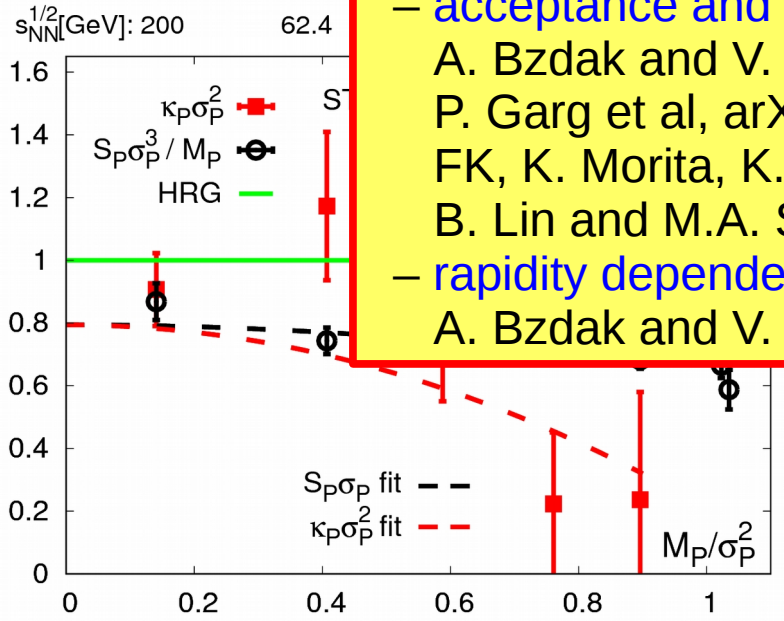


NLO lattice QCD
(not yet continuum extrapolated)



need to understand systematic (and other) effects :

- non-equilibrium effects
B. Berdnikov, K. Rajagopal, hep-ph/9912274
S. Mukherjee et al., arXiv:1506.00645, arXiv:1605.09341
- proton vs. baryon number distributions
M. Kitazawa et al, arXiv:1205.3292, arXiv:1303.3338
- acceptance and p_t -cuts
A. Bzdak and V. Koch, arXiv:12064286
P. Garg et al, arXiv:1304.7133,
FK, K. Morita, K. Redlich, arXiv:1508.02614
B. Lin and M.A. Stephanov, arXiv:1512.09125
- rapidity dependence
A. Bzdak and V. Koch, arXiv:1707.02640



hotQCD in preparation

Equation of state of (2+1)-flavor QCD: $\mu_B/T > 0$

$$\frac{P}{T^4} = \sum_{i,j,k=0}^{\infty} \frac{1}{i!j!k!} \chi_{i,j,k}^{BQS}(T) \left(\frac{\mu_B}{T}\right)^i \left(\frac{\mu_Q}{T}\right)^j \left(\frac{\mu_S}{T}\right)^k$$

the simplest case: $\mu_S = \mu_Q = 0$

$$\frac{P(T, \mu_B)}{T^4} = \frac{P(T, 0)}{T^4} + \frac{\chi_2^B(T)}{2} \left(\frac{\mu_B}{T}\right)^2 + \frac{\chi_4^B(T)}{24} \left(\frac{\mu_B}{T}\right)^4 + \mathcal{O}((\mu_B/T)^6)$$

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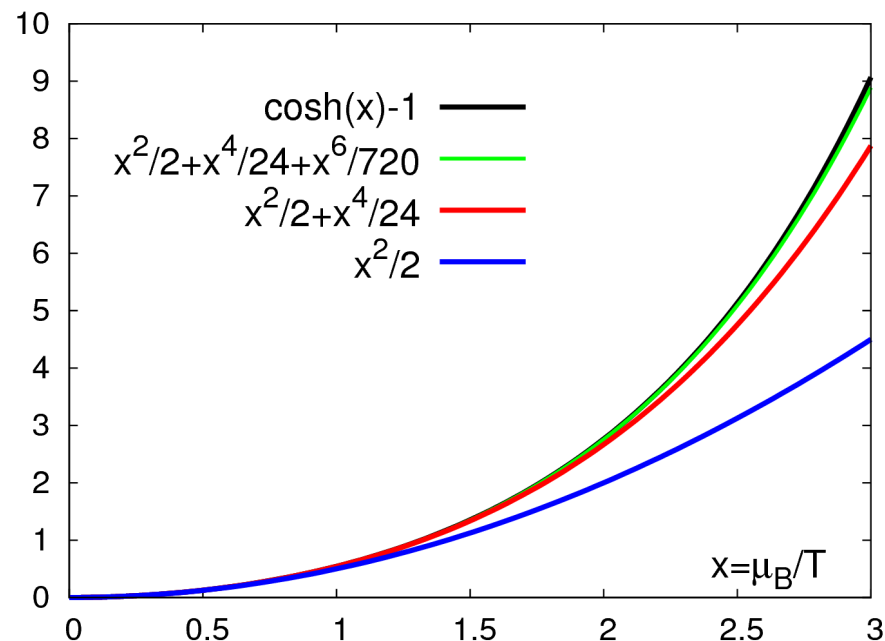
$$\frac{P(T, \mu_B)}{T^4} = \frac{P(T, 0)}{T^4} + \frac{\chi_2^B(T)}{2} \left(\frac{\mu_B}{T}\right)^2 + \frac{\chi_4^B(T)}{24} \left(\frac{\mu_B}{T}\right)^4 + \mathcal{O}((\mu_B/T)^6)$$

An $\mathcal{O}((\mu_B/T)^4)$ expansion is exact in a QGP up to $\mathcal{O}(g^2)$

HRG vs. QCD:

$\mathcal{O}((\mu_B/T)^4)$: difference is less than 3% at $\mu_B/T = 2$

$\mathcal{O}((\mu_B/T)^6)$: difference is less than 2% at $\mu_B/T = 3$



Equation of state of (2+1)-flavor QCD: $\mu_B/T > 0$

$$\frac{P}{T^4} = \sum_{i,j,k=0}^{\infty} \frac{1}{i!j!k!} \chi_{i,j,k}^{BQS}(T) \left(\frac{\mu_B}{T}\right)^i \left(\frac{\mu_Q}{T}\right)^j \left(\frac{\mu_S}{T}\right)^k$$

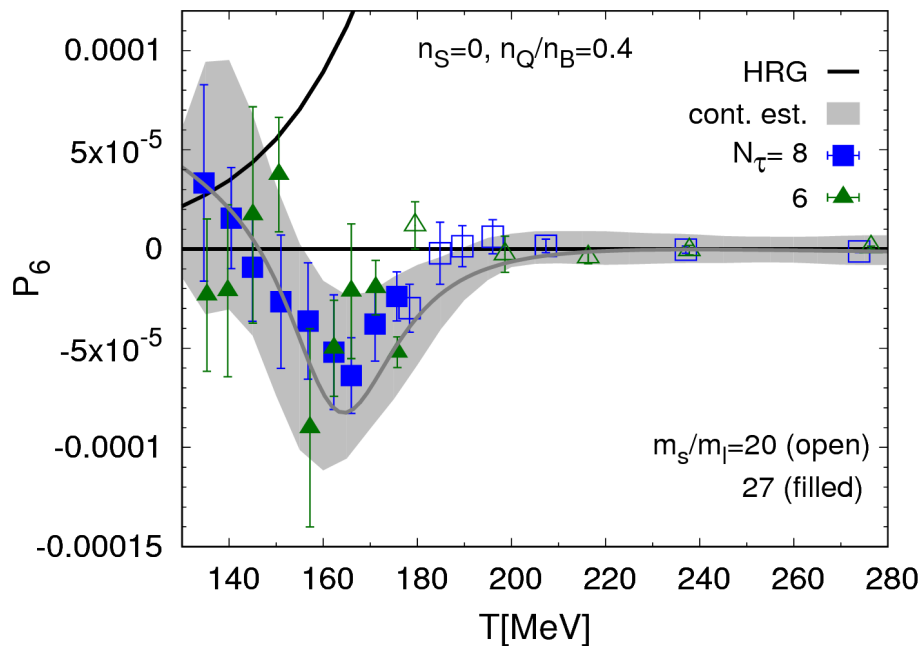
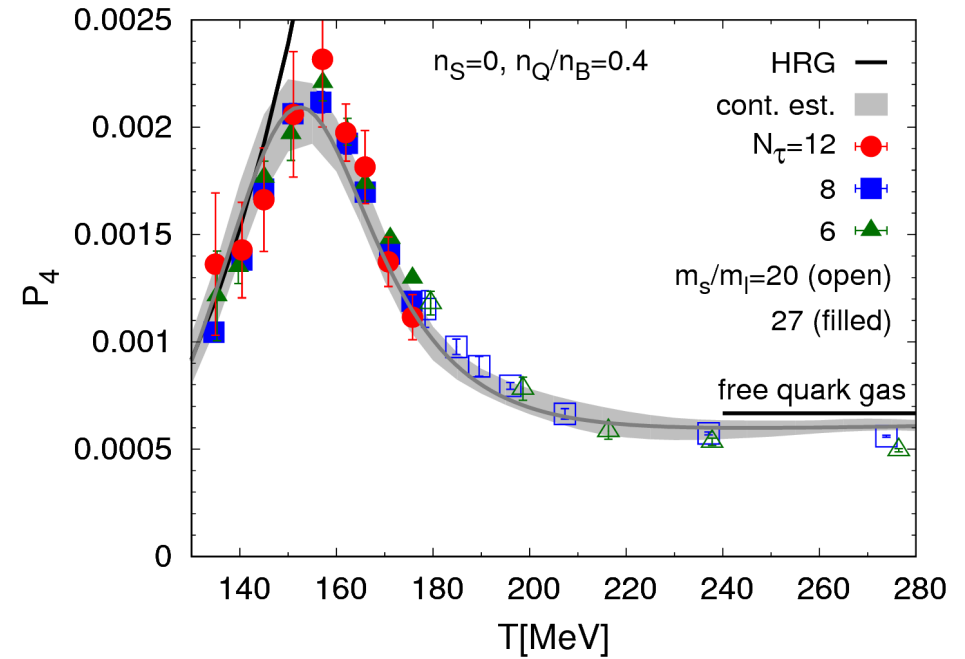
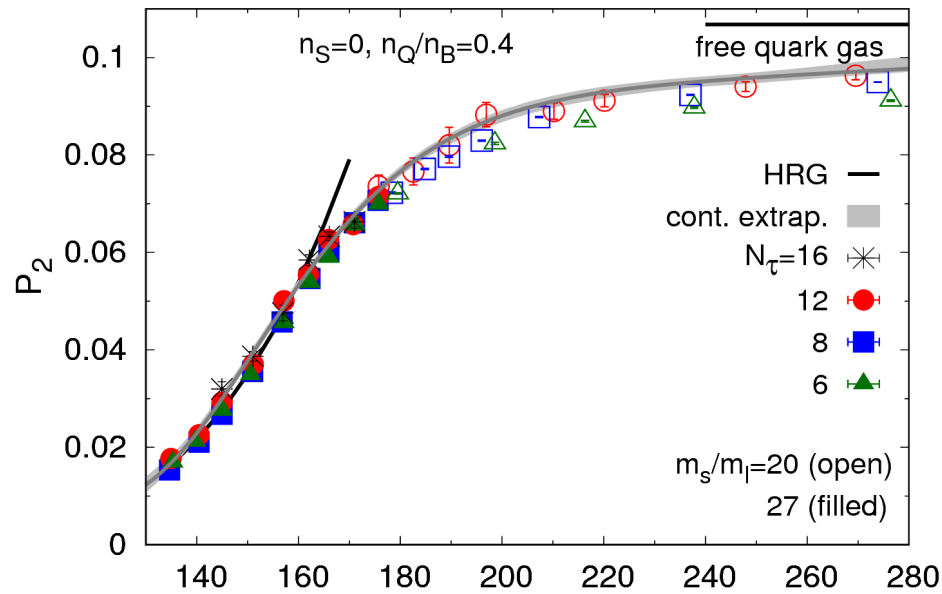
the simplest case: $\mu_S = \mu_Q = 0$

$$\frac{P(T, \mu_B)}{T^4} = \frac{P(T, 0)}{T^4} + \frac{\chi_2^B(T)}{2} \left(\frac{\mu_B}{T}\right)^2 + \frac{\chi_4^B(T)}{24} \left(\frac{\mu_B}{T}\right)^4 + \mathcal{O}((\mu_B/T)^6)$$

strangeness neutral: $n_S = 0$, $n_Q/n_B = 0.4$

$$\frac{P(T, \mu_B)}{T^4} = P_0(T) + P_2(T) \left(\frac{\mu_B}{T}\right)^2 + P_4(T) \left(\frac{\mu_B}{T}\right)^4 + \mathcal{O}((\mu_B/T)^6)$$

Taylor expansion coefficients

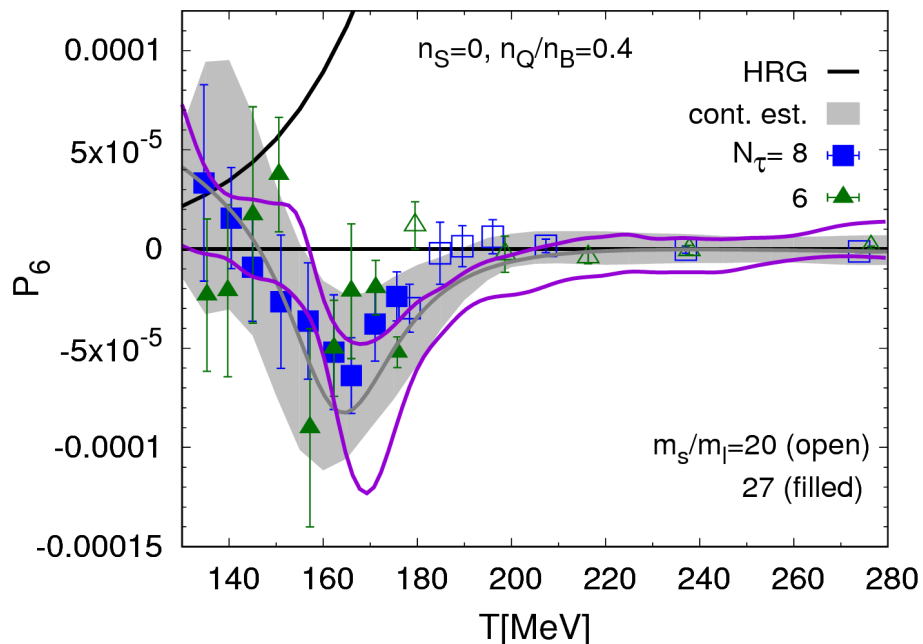
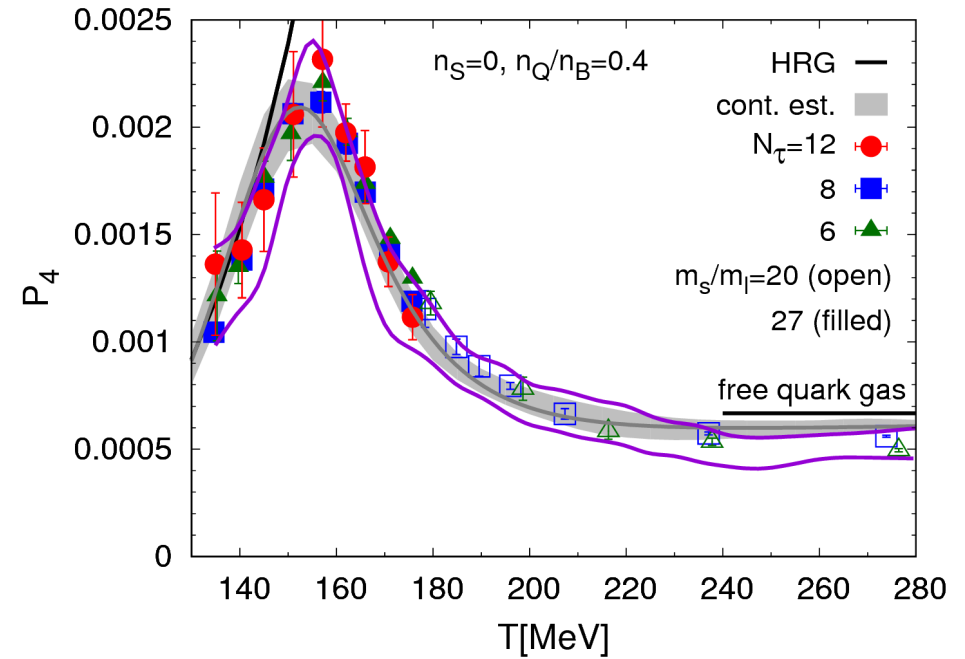
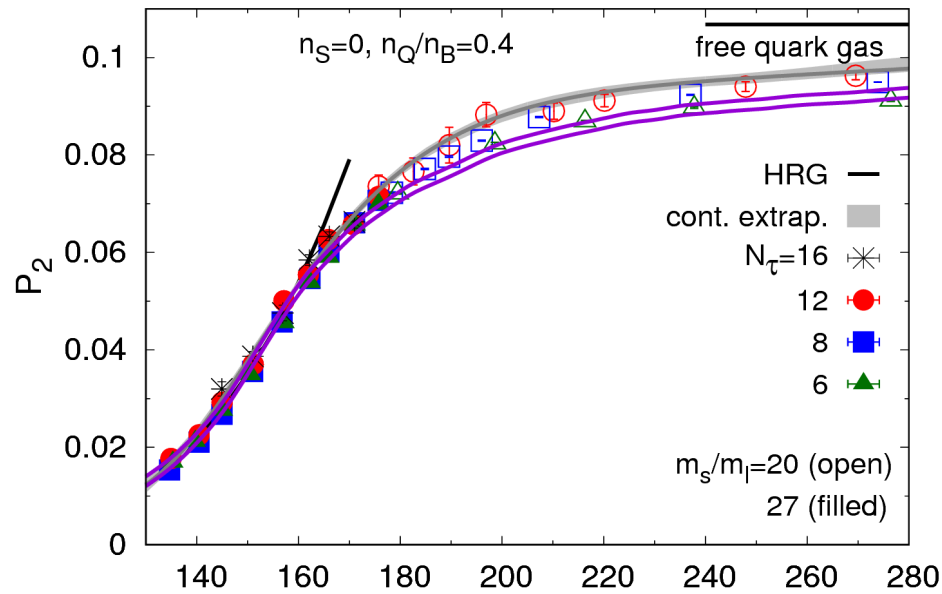


fits from: [A. Bazavov et al., Phys. Rev. D 95 \(2017\) 054504](#)

data are updated: hotQCD 2017

$P_6 < 0$ for $T \gtrsim 150$ MeV

Taylor expansion coefficients



fits from: [A. Bazavov et al., Phys. Rev. D 95 \(2017\) 054504](#)

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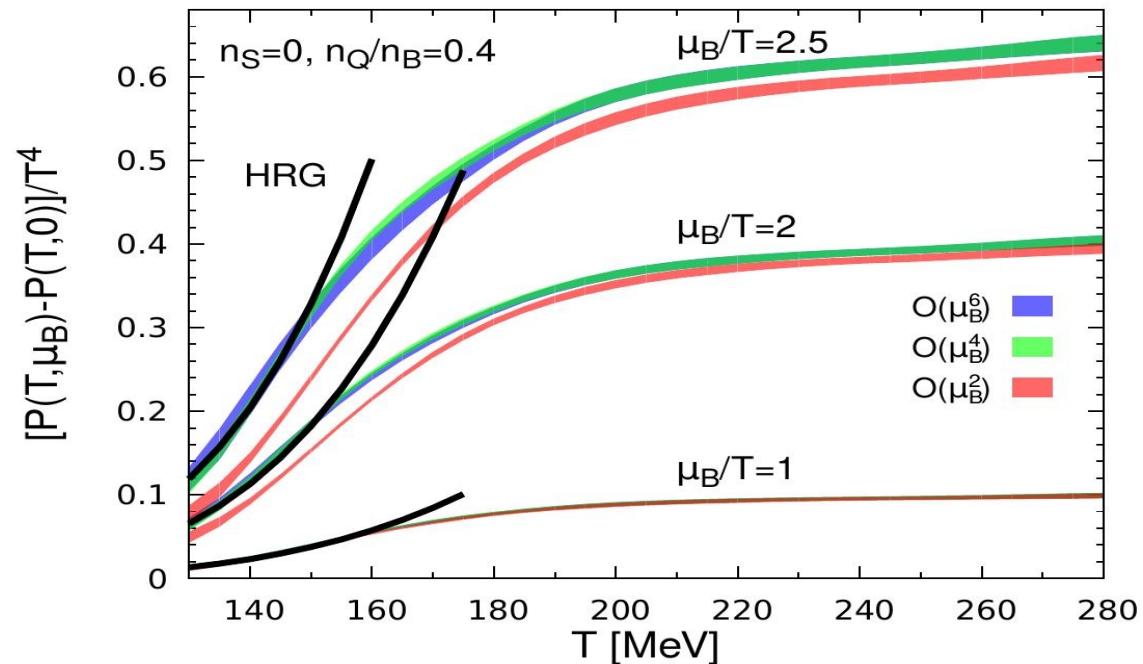
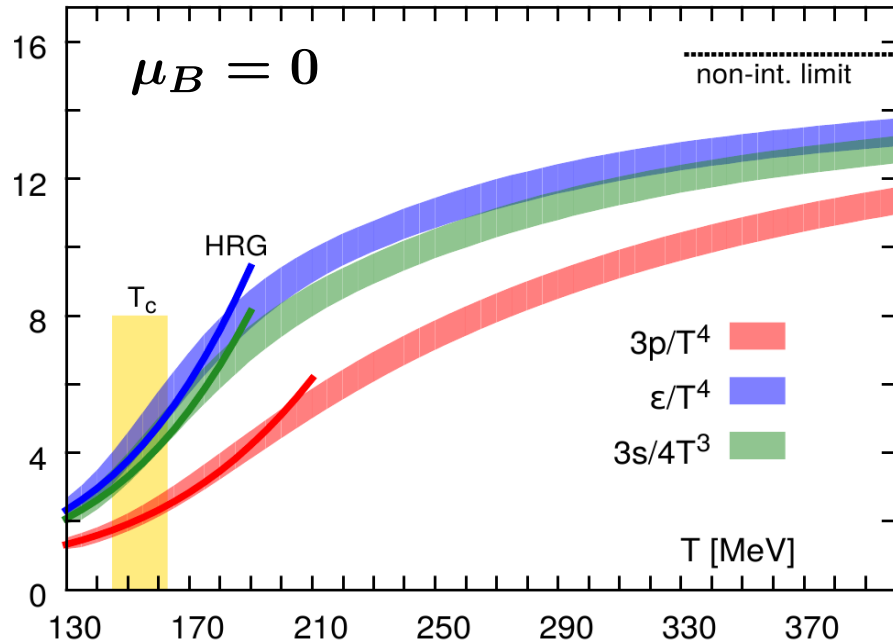
$P_6 < 0$ for $T \gtrsim 150$ MeV

consistent with results obtained from analytic continuation:
[J. Gunther et al., EPJ Web Conf. 137 \(2017\) 07008](#)

Equation of state of (2+1)-flavor QCD: $\mu_B/T > 0$

$$\frac{\Delta(T, \mu_B)}{T^4} = \frac{P(T, \mu_B) - P(T, 0)}{T^4} = \frac{\chi_2^B}{2} \left(\frac{\mu_B}{T}\right)^2 + \frac{\chi_4^B}{24} \left(\frac{\mu_B}{T}\right)^4 + \frac{\chi_6^B}{720} \left(\frac{\mu_B}{T}\right)^6 + \dots$$

(10-30)% contribution to total pressure at $\mu_B/T = 2$



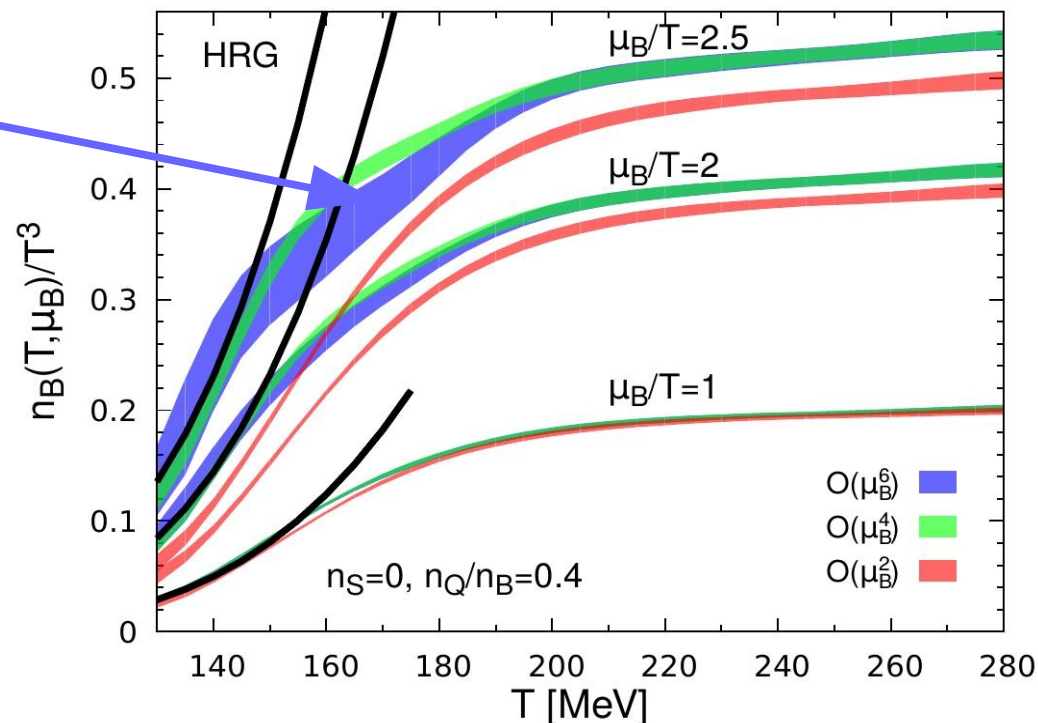
The EoS is well controlled for $\mu_B/T \leq 2$ or equivalently $\sqrt{s_{NN}} \geq 19.6$ GeV

Equation of state of (2+1)-flavor QCD: $\mu_B/T > 0$

net baryon-number density:

$$\frac{n_B}{T^3} = \chi_2^B \frac{\mu_B}{T} + \frac{\chi_4^B}{6} \left(\frac{\mu_B}{T}\right)^3 + \frac{\chi_6^B}{120} \left(\frac{\mu_B}{T}\right)^5 + \dots$$

dip in $\chi_6^B(T)$, $P_6(T)$
calls for inclusion of
higher order corrections
and/or better statistics!!



The EoS is well controlled for $\mu_B/T \leq 2$
or equivalently $\sqrt{s_{NN}} \geq 19.6$ GeV

HRG vs. QCD

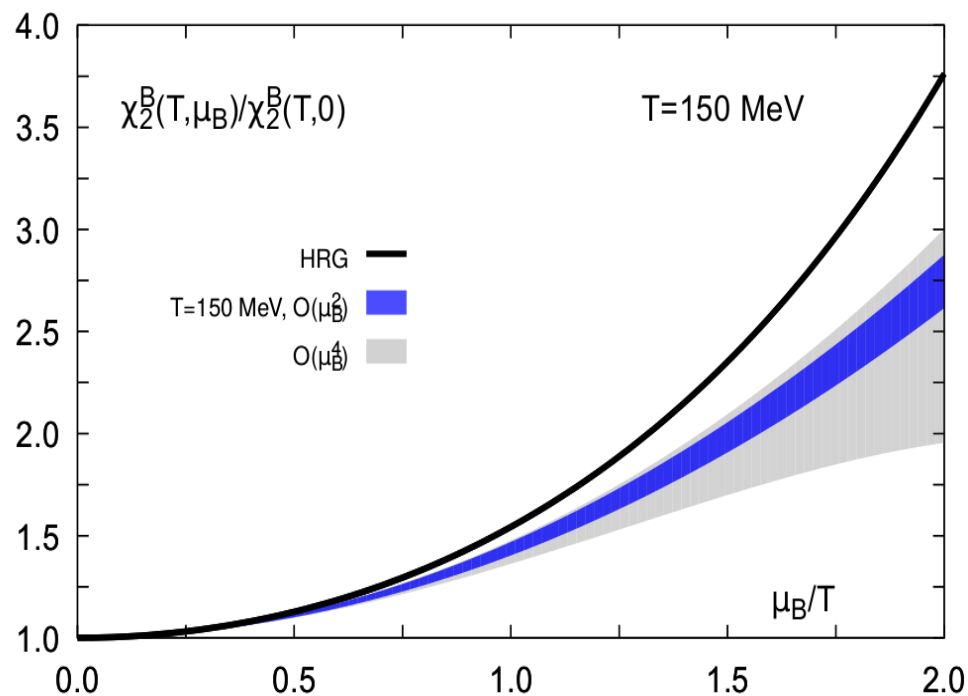
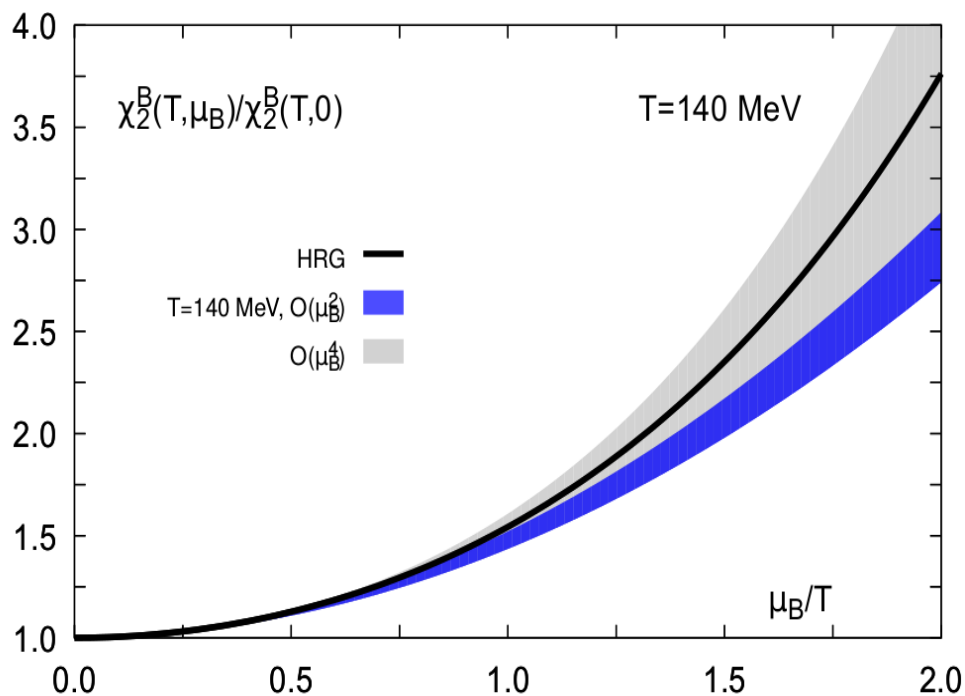
net baryon-number fluctuations

$$\mu_B/T > 0$$

for simplicity: $\mu_Q = \mu_S = 0$

$$\chi_2^B(T, \mu_B) = \chi_2^B + \frac{1}{2} \chi_4^B \left(\frac{\mu_B}{T} \right)^2 + \frac{1}{24} \chi_6^B \left(\frac{\mu_B}{T} \right)^4 + \mathcal{O}(\mu_B^6)$$

- agreement between HRG and QCD will start to deteriorate for $T > 150$ MeV
- net baryon-number fluctuations in QCD always smaller than in HRG for $T > 150$ MeV



HRG vs. QCD

net baryon-number fluctuations

$$\mu_B/T > 0$$

for simplicity: $\mu_Q = \mu_S = 0$

$$\chi_2^B(T, \mu_B) = \chi_2^B + \frac{1}{2} \chi_4^B \left(\frac{\mu_B}{T} \right)^2 + \frac{1}{24} \chi_6^B \left(\frac{\mu_B}{T} \right)^4 + \mathcal{O}(\mu_B^6)$$

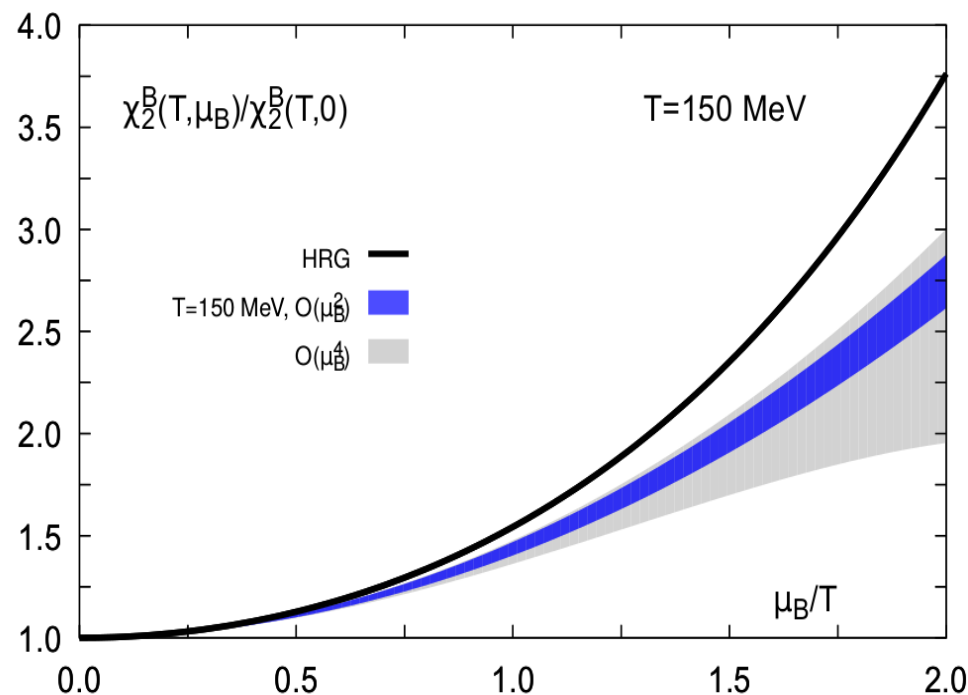
- agreement between HRG and QCD will start to deteriorate for $T > 150$ MeV
- net baryon-number fluctuations in QCD always smaller than in HRG for $T > 150$ MeV

no evidence for enhanced net baryon-number fluctuations for

$$T \geq 135 \text{ MeV}, \mu_B \leq 2T$$



no evidence for getting closer to a "critical region"



Taylor expansion of the pressure and critical point

$$\frac{P}{T^4} = \sum_{n=0}^{\infty} \frac{1}{n!} \chi_n^B(T) \left(\frac{\mu_B}{T}\right)^n$$

estimator for the radius of convergence:

$$\left(\frac{\mu_B}{T}\right)_{crit,n}^{\chi} \equiv r_n^{\chi} = \sqrt{\left| \frac{n(n-1)\chi_n^B}{\chi_{n+2}^B} \right|}$$

for simplicity : $\mu_Q = \mu_S = 0$

– radius of convergence corresponds to a critical point **only**, iff

$$\chi_n > 0 \text{ for all } n \geq n_0$$

forces P/T^4 and $\chi_n^B(T, \mu_B)$ to be monotonically growing with μ_B/T

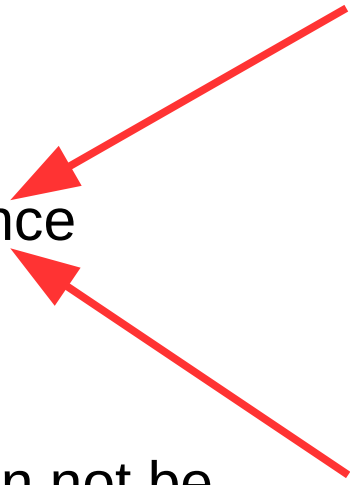


$$\text{at } T_{CP} : \kappa_B \sigma_B^2 = \frac{\chi_4^B(T, \mu_B)}{\chi_2^B(T, \mu_B)} > 1$$

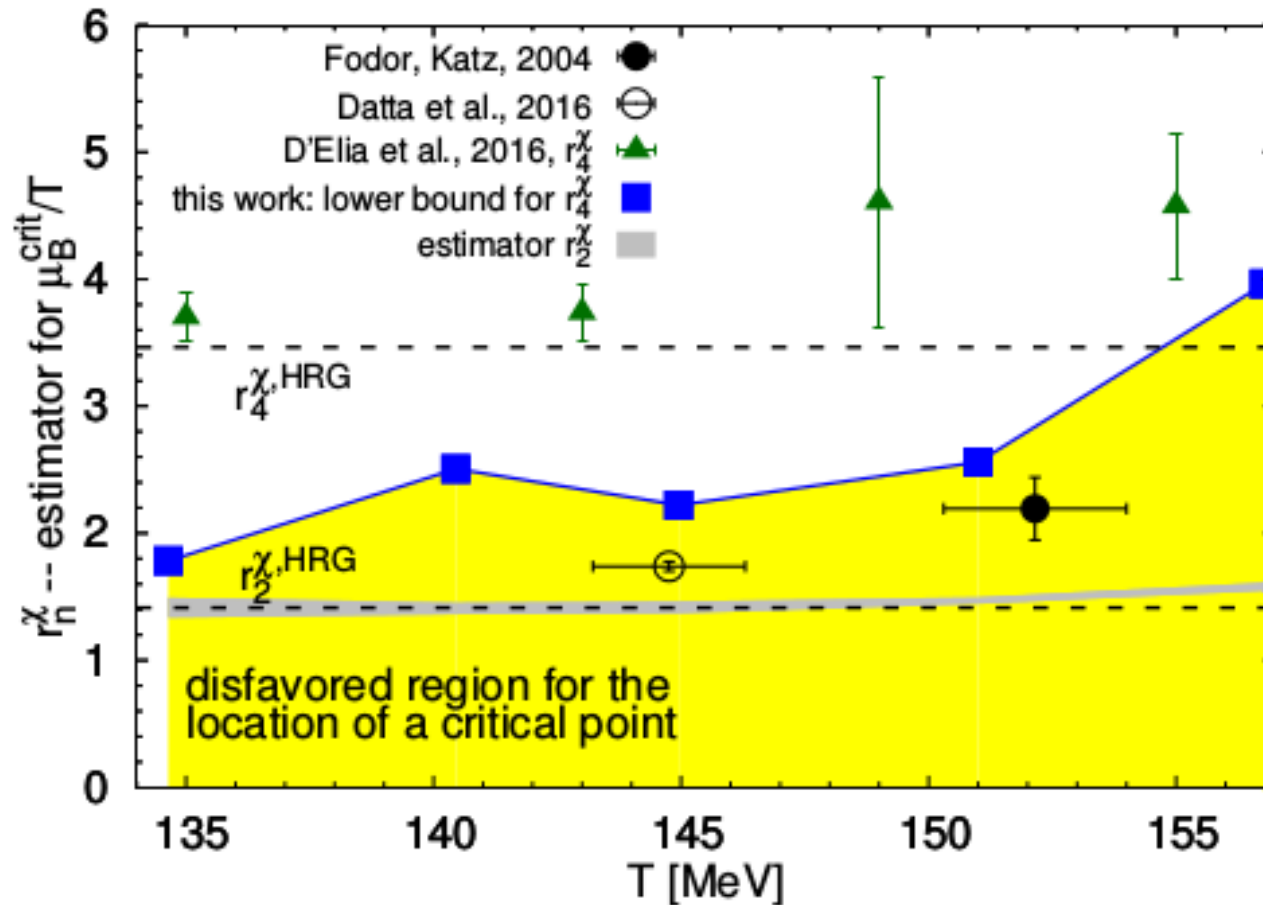
if not:

– radius of convergence does not determine the critical point

– Taylor expansion can not be used close to the critical point



estimates/constraints on critical point location

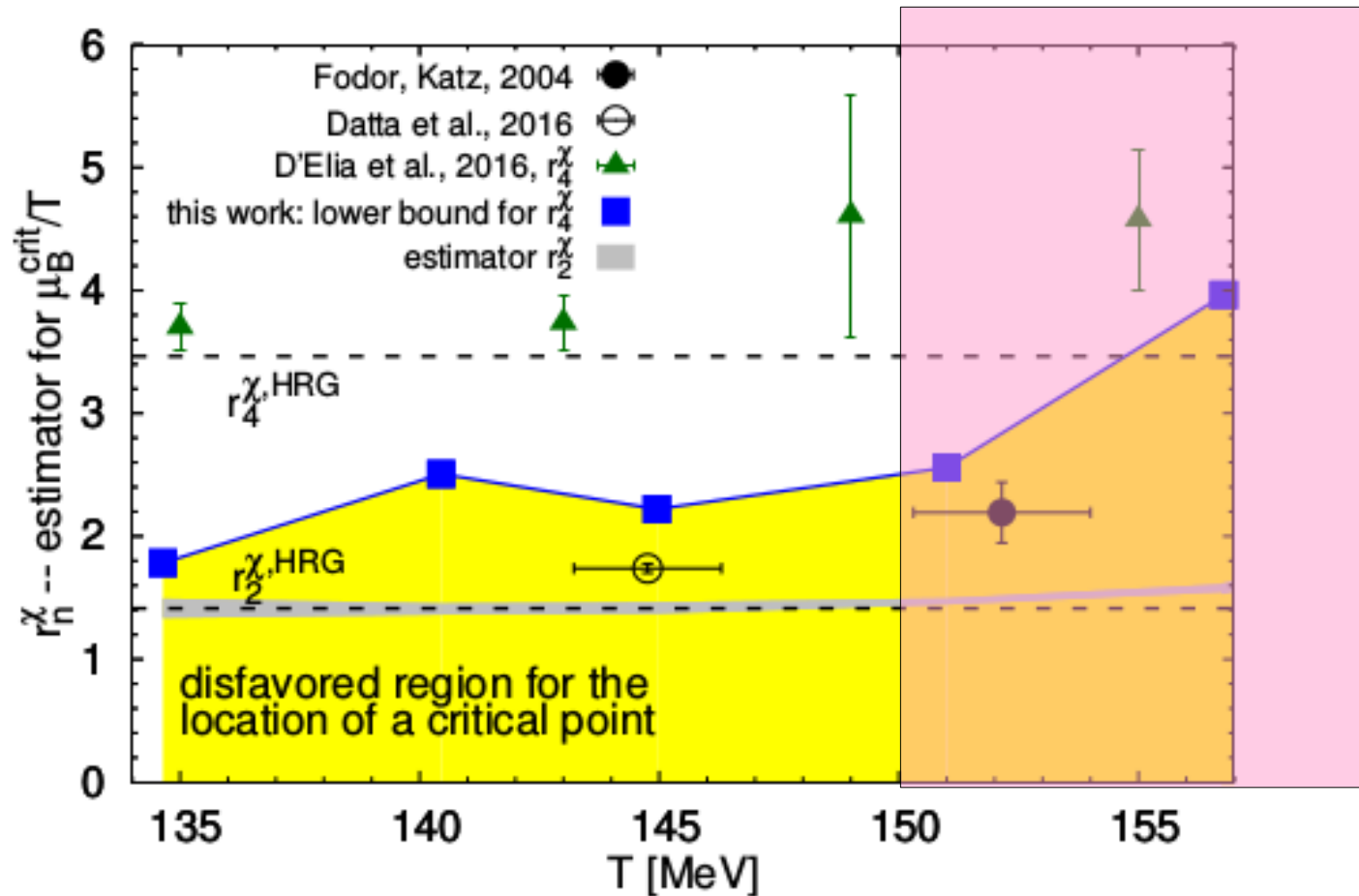


01/01/17:

based on ongoing calculations of 6th order Taylor expansion coefficients performed by the Bielefeld-BNL-CCNU collaboration

A. Bazavov et al., arXiv:1701.04325

estimates/constraints on critical point location



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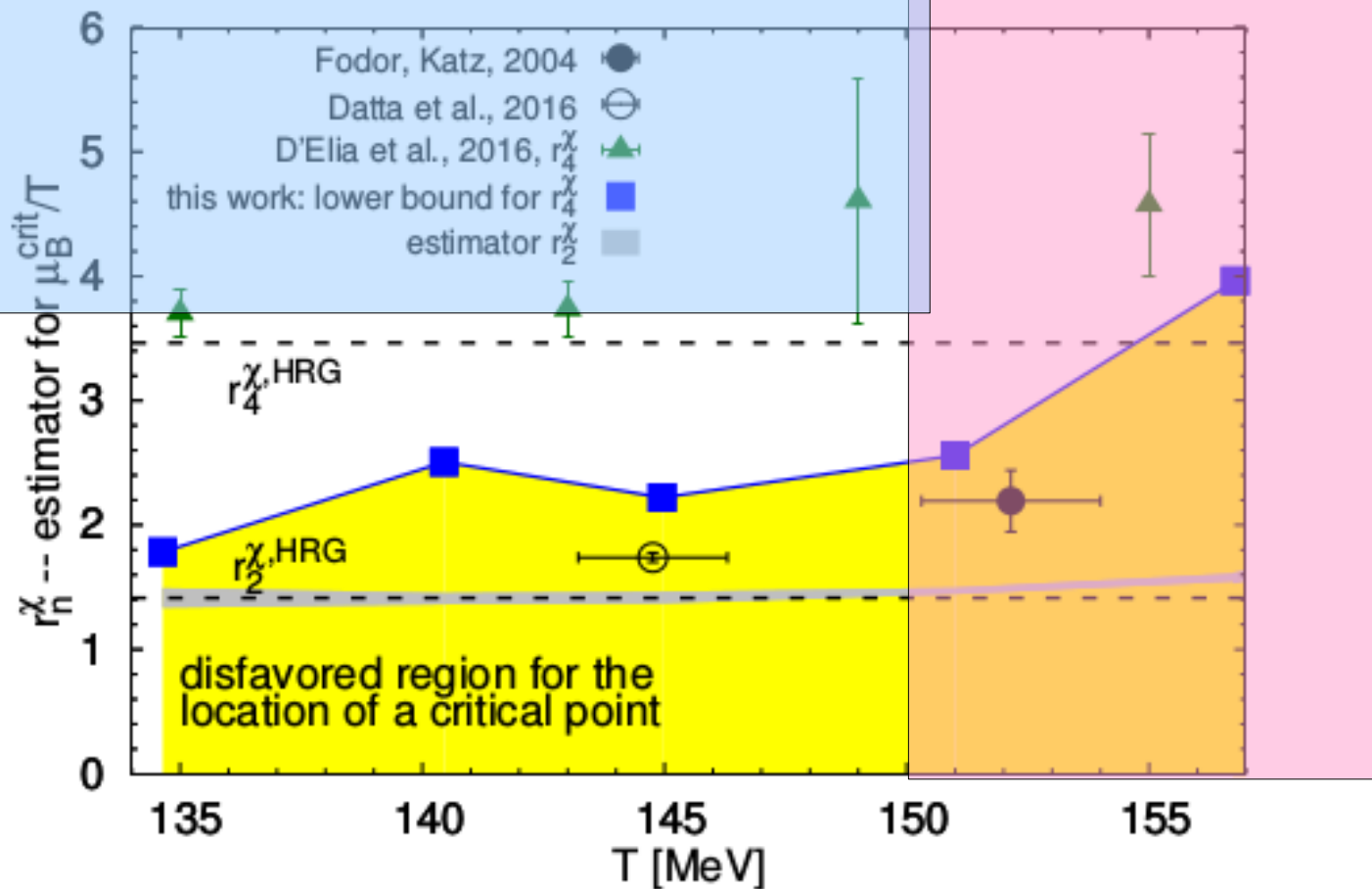
A. Bazavov et al., arXiv:1701.04325

strongly disfavored
as

$$\chi_6^B < 0$$

estimates/constraints on critical point location

not accessible
in BES@RHIC
collider mode



01/01/17:

based on ongoing calculations of 6th order Taylor expansion coefficients performed by the Bielefeld-BNL-CCNU collaboration

A. Bazavov et al., arXiv:1701.04325

strongly disfavored
as

$$\chi_6^B < 0$$

To do list

What is needed to understand equilibrium properties of conserved charge fluctuations on the freeze-out line?

- improve lattice QCD results on 6th (and 8th) order cumulants of conserved charge fluctuations
- self-consistent determination of freeze-out parameters based on equilibrium QCD calc.: $T_f(\mu_B)$, μ_B , $[\mu_S(\mu_B), \mu_Q(\mu_B)]$

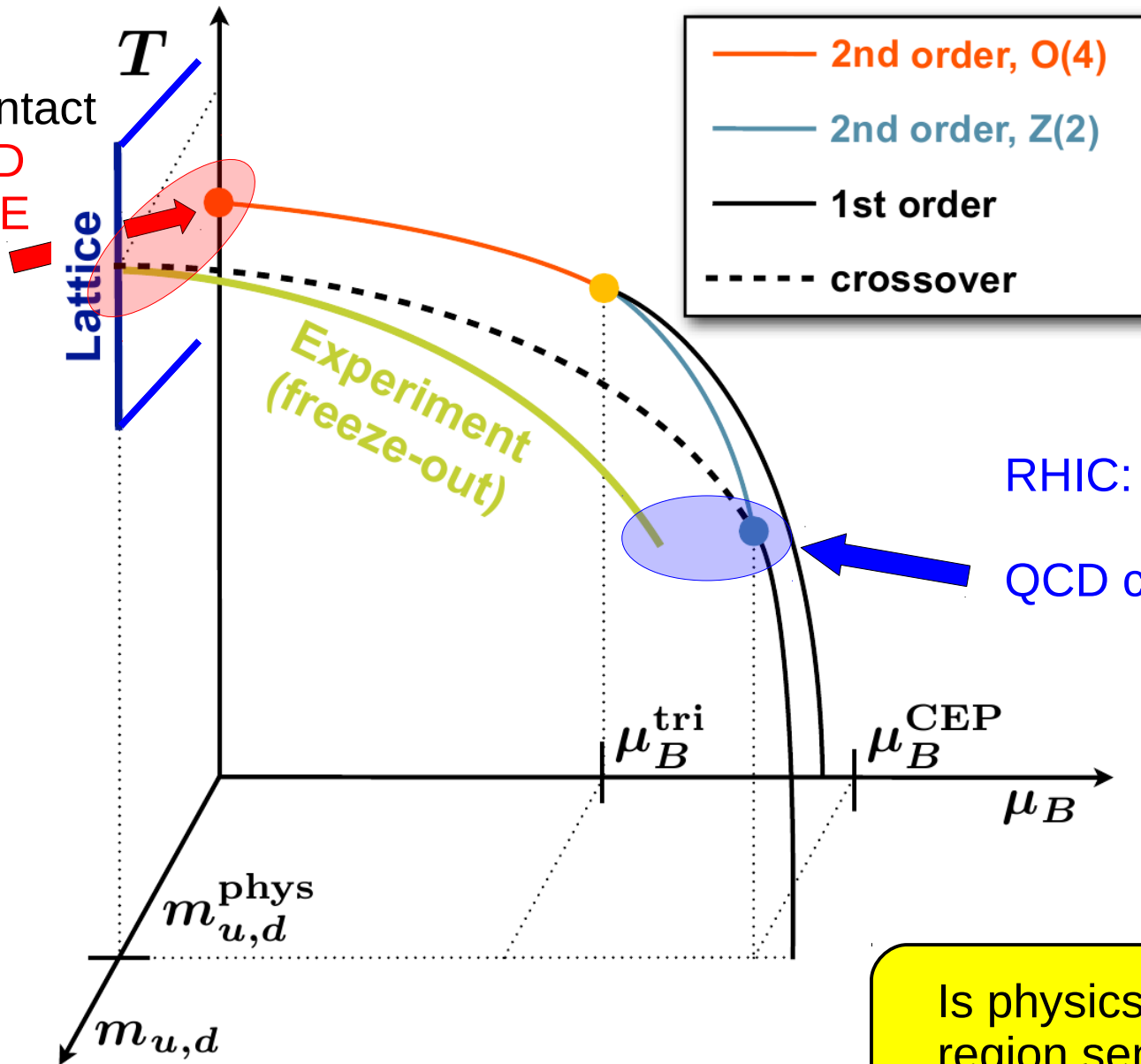
What can be done about "locating the critical point"?

- use 6th (and 8th) order cumulants to put bounds on its location
- keep working on new simulation techniques

By-product: EoS in the entire parameter range accessible to the RHIC BES-II

Chiral critical point and QCD critical endpoint

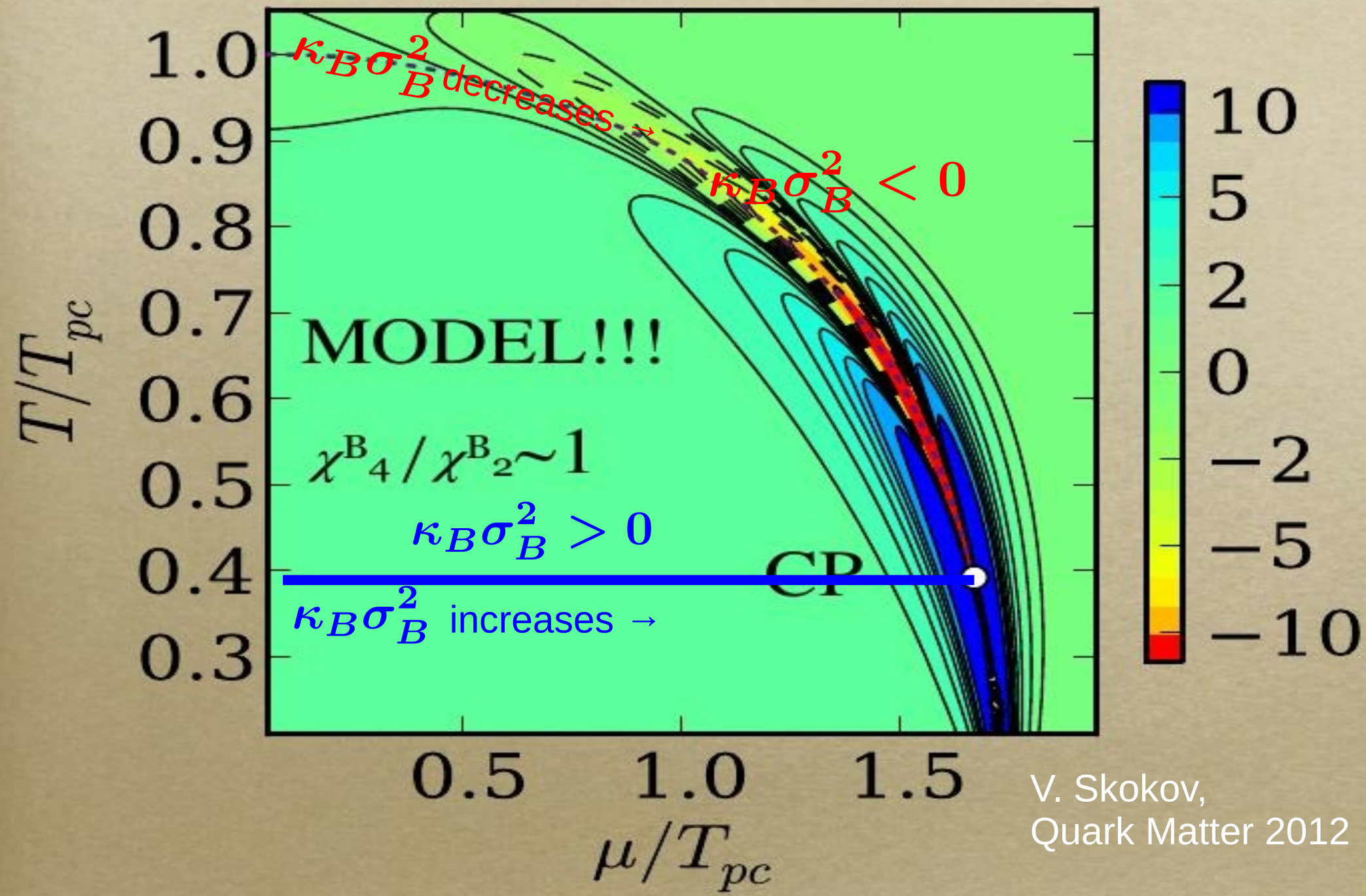
LHC: may establish contact with the QCD chiral PHASE transition



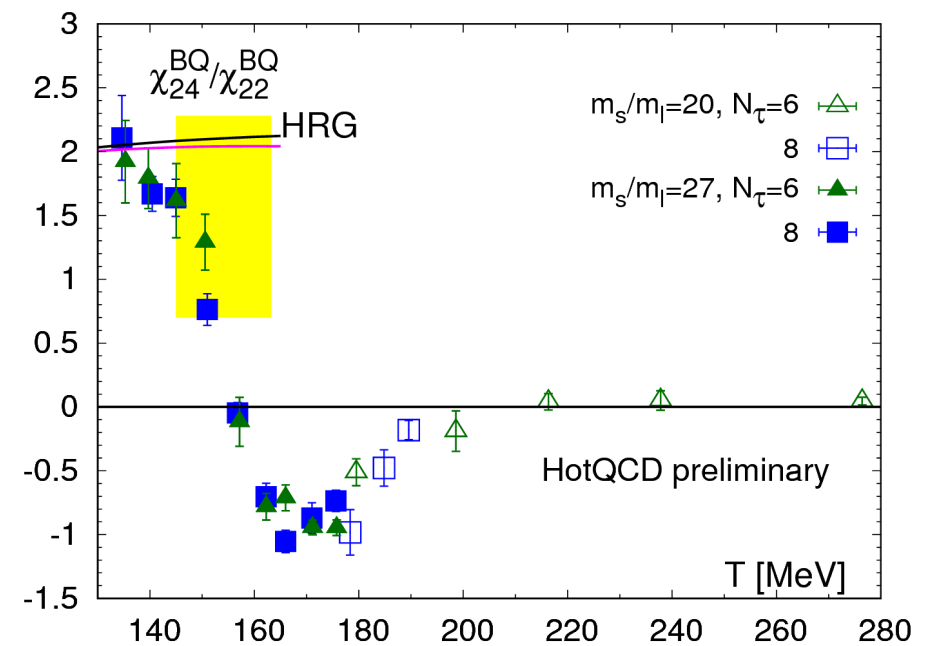
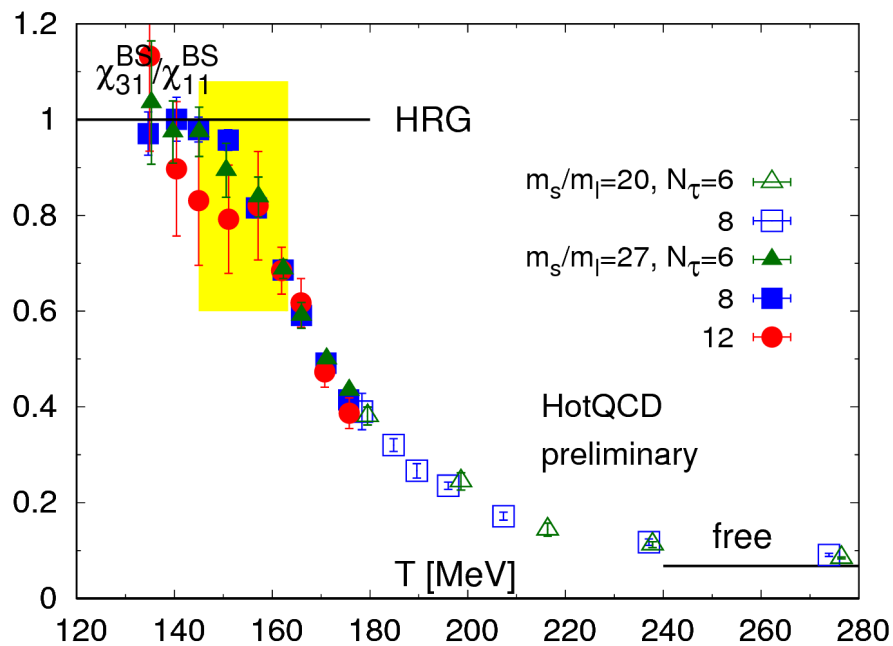
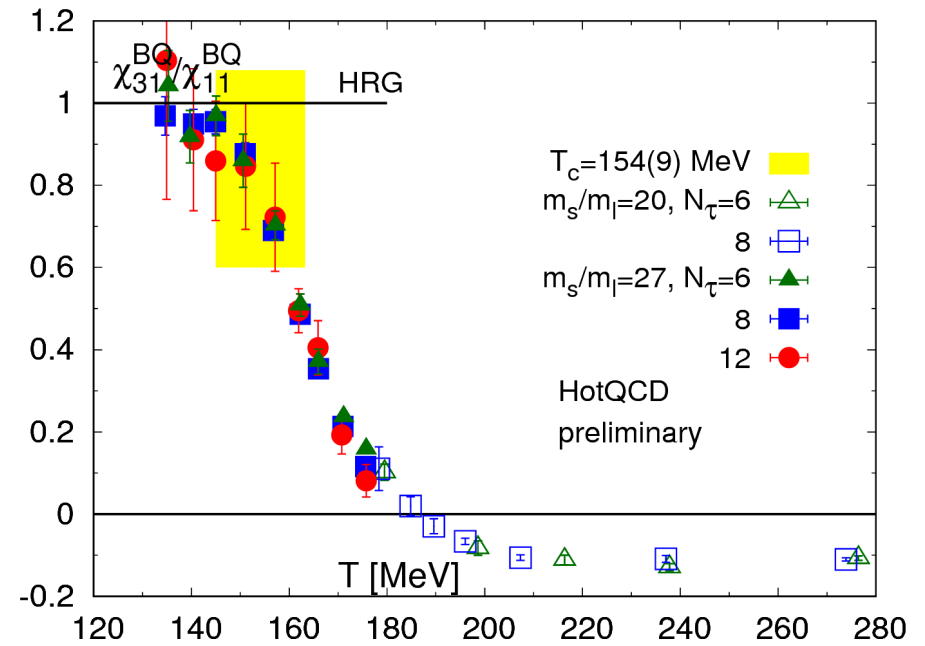
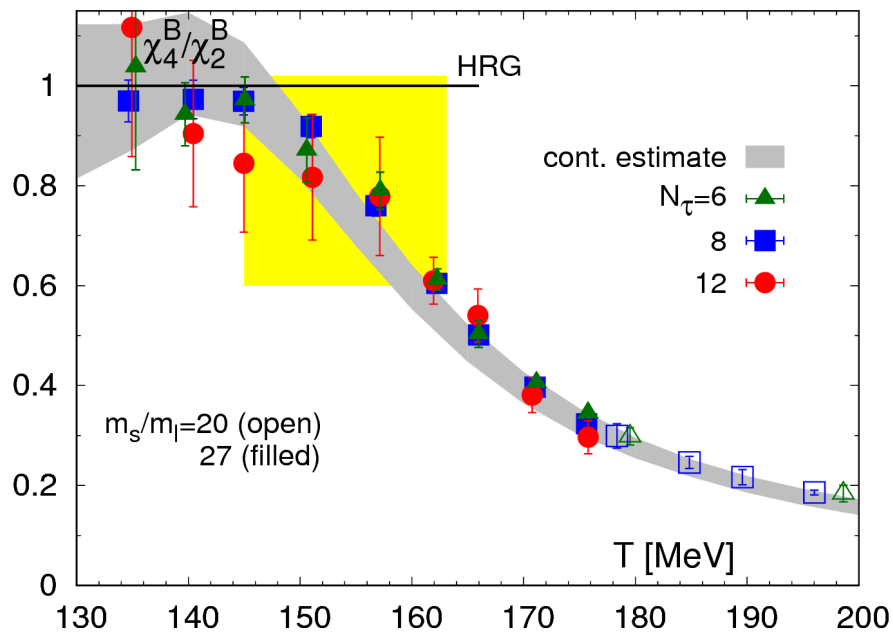
RHIC: may establish evidence for the QCD critical end point

Is physics in the freeze-out region sensitive to critical behavior?

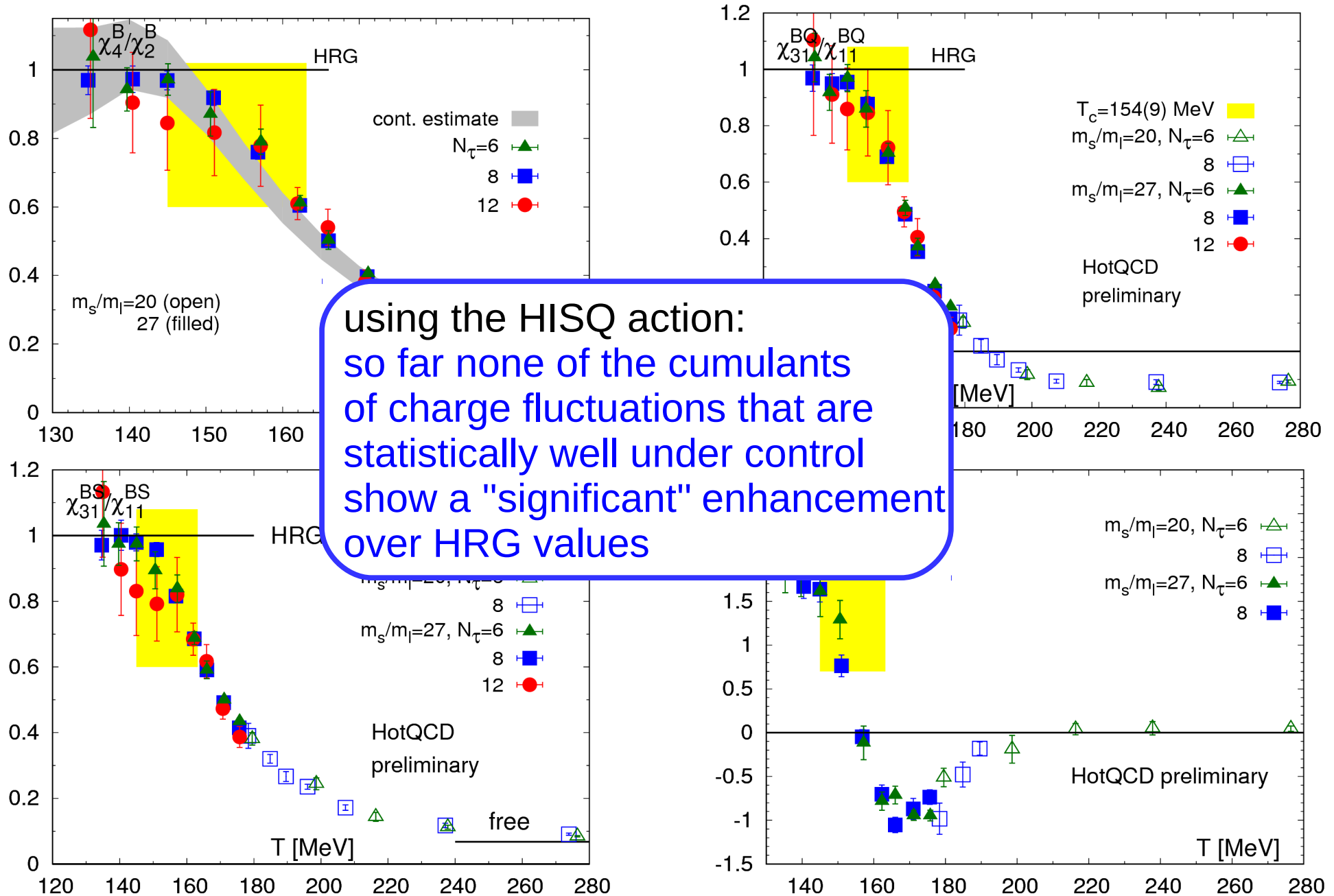
Chiral model and negative χ^B_4 / χ^B_2 :



Some 4th and 6th order cumulants



Some 4th and 6th order cumulants



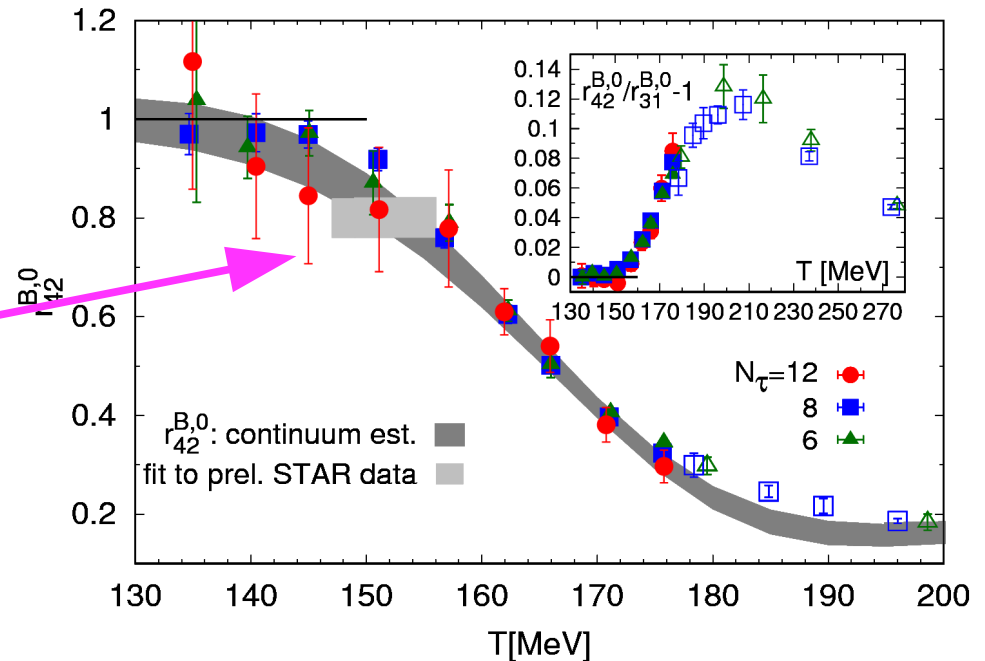
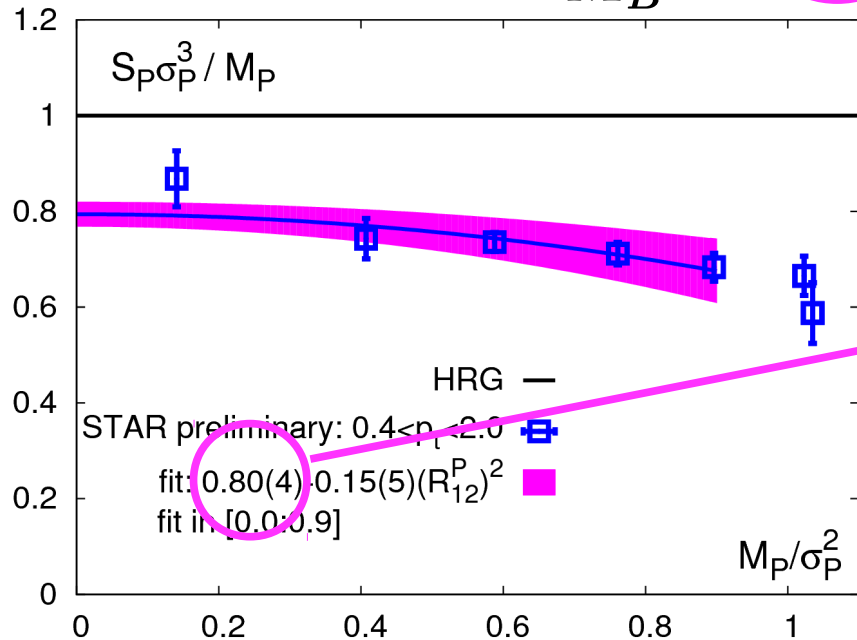
Conserved charge fluctuations and freeze-out

mean, variance and skewness

NLO Taylor expansion

$$S_B \sigma_B = \frac{\chi_4^B}{\chi_2^B} \frac{M_B}{\sigma_B^2} + \frac{1}{6} \left(\frac{\chi_6^B}{\chi_2^B} - \left(\frac{\chi_4^B}{\chi_2^B} \right)^2 \right) \left(\frac{M_B}{\sigma_B^2} \right)^3 + \dots$$

$$\Leftrightarrow R_{31}^B \equiv \frac{S_B \sigma_B^3}{M_B} = r_{31}^{B,0} + r_{31}^{B,2} (R_{12}^B)^2$$



intercept consistent with QCD result,

– used $T_{f,0}$ determined from fit to $(M_Q / \sigma_Q^2) / (M_P / \sigma_P^2)$

Conserved charge fluctuations and freeze-out

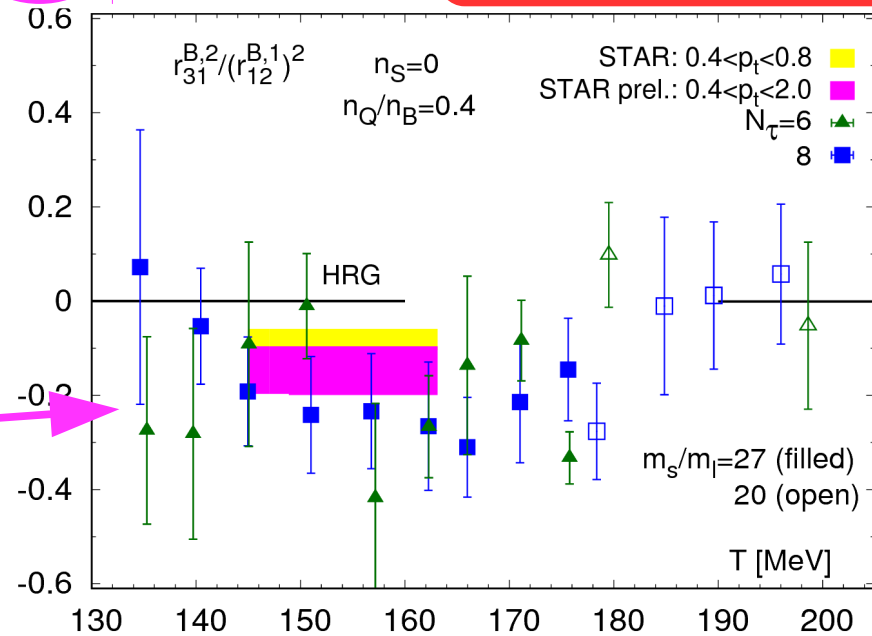
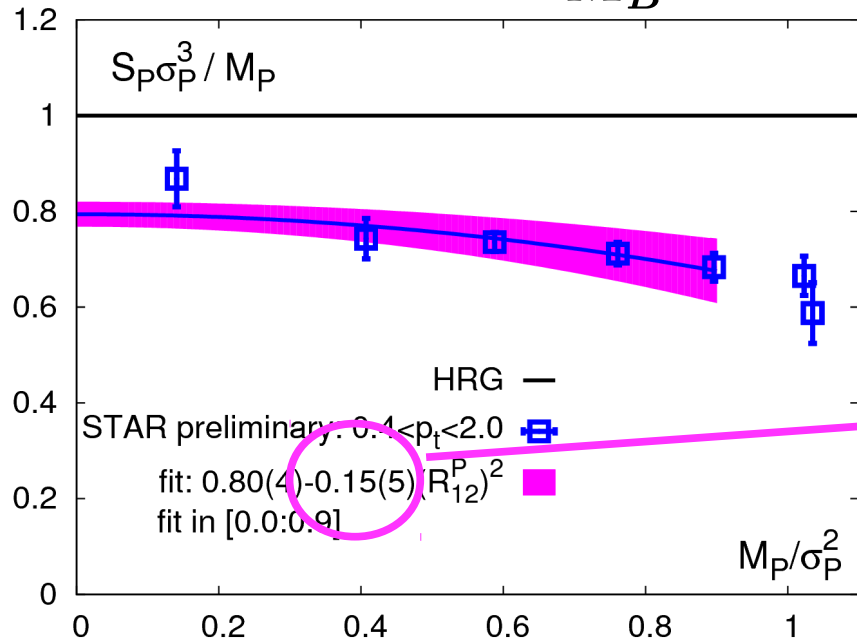
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$$\Leftrightarrow R_{31}^B \equiv \frac{S_B \sigma_B^3}{M_B} = r_{31}^{B,0} + r_{31}^{B,2} (R_{12}^B)^2$$

lattice QCD calculation involves 6th order cumulants



curvature consistent with (still preliminary) QCD result (noisy, coarse lattice)

used $T_{f,0}$ determined from fit to $(M_Q/\sigma_Q^2)/(M_P/\sigma_P^2)$