

CERN TB Preparation

26.7.21 BIC ESB/System Testing meeting

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Outline

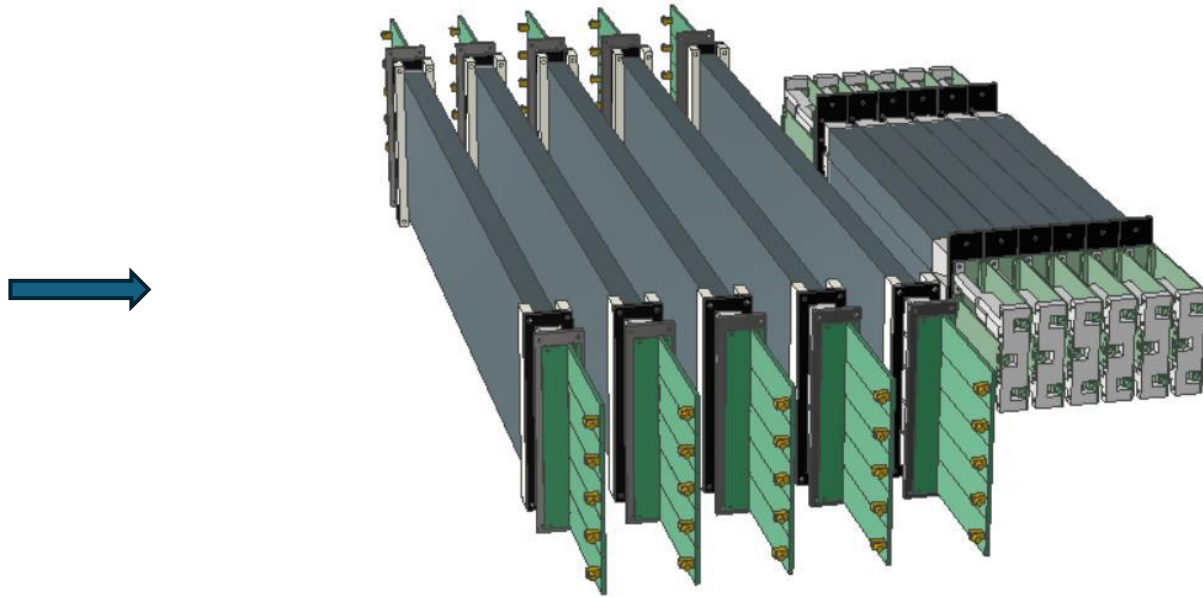
1. Preparation for CERN beam test
 - a) Lessons from KEK TB
 - b) Detector preparation
2. Timestamp matching at CERN
 - a) Trigger-H2GCROC timestamp matching at KEK in May
 - b) Idea for AstroPix-H2GCROC matching

1. Preparation for CERN TB

Lessons from KEK in May

- We can find optimal parameters to obtain signals through H2GCROC
- We are still studying H2GCROC data
- Event matching between Auxiliary detectors (FADC) and H2GCROC was successful within $\pm 1\mu\text{s}$
- AstroPix 9-chip was OK under 2.7kHz rate

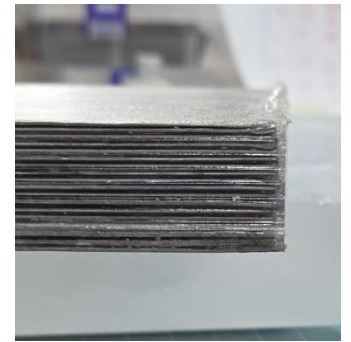
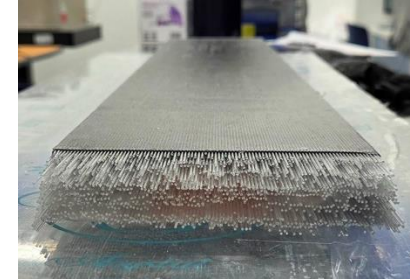
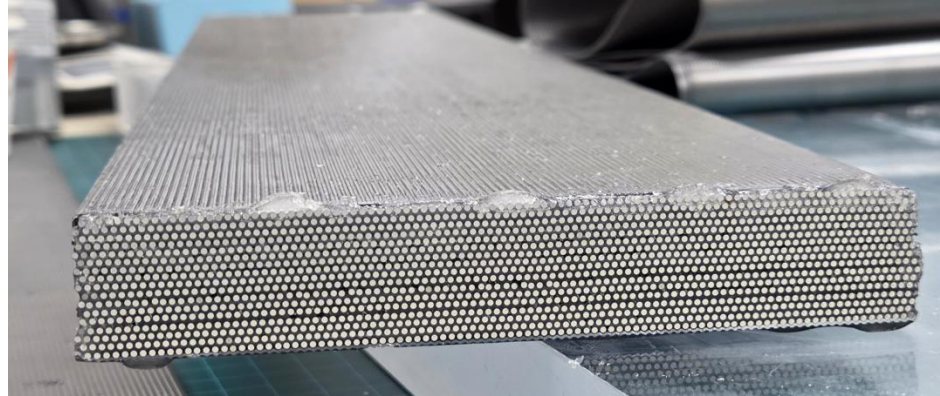
Possible Pb/SciFi setup at CERN



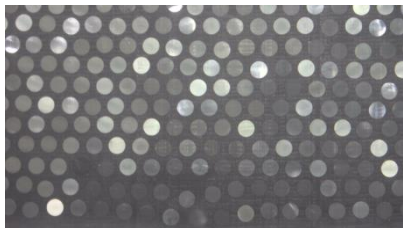
- 2~5 SFILs in front of
- 32x9x18cm module (reuse of previous prototype)
- AstroPix 9 chip module in various location
- Auxiliary: Cherenkov, Trigger, hodoscope

SFIL production at KNU

(Minsum Shim, Wonjun Ko, Huisu Lim, ki Seung, Hui, Changhui Lee)



First SFIL had with some defect <5% when we applied optical cement by brush.
Second SFIL using improved procedure



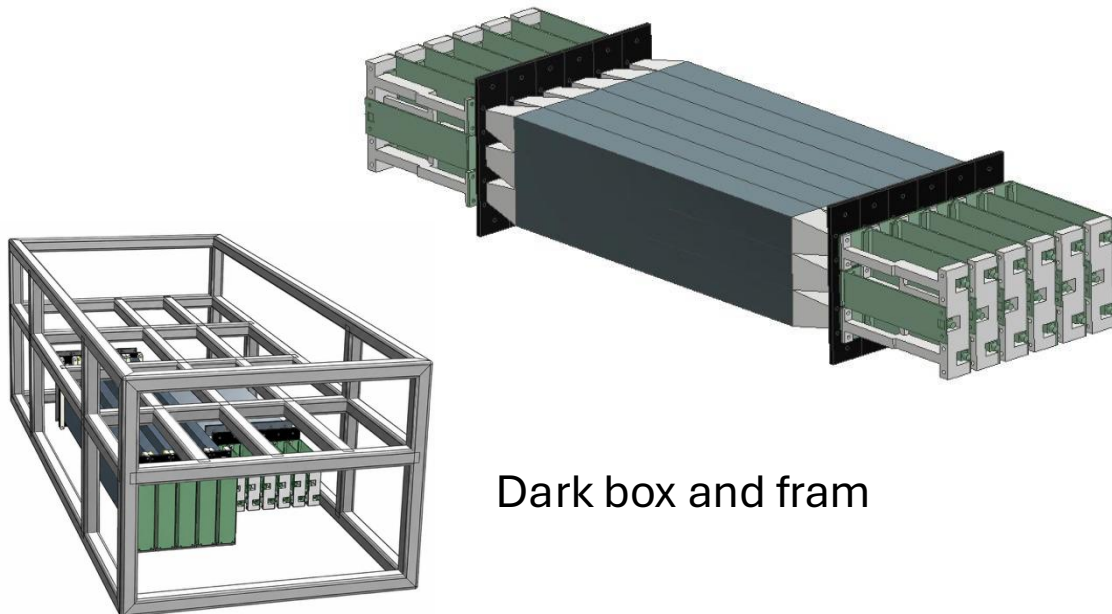
After milling 32x3x3cm³ module (fibers are cured firmly)



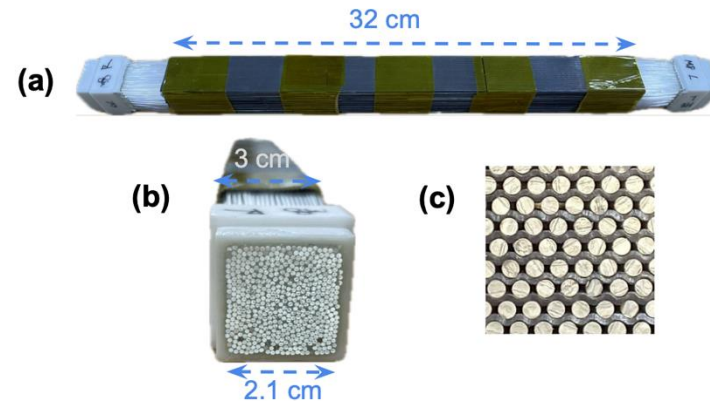
2 SFILs are ready, cutting at 2mm extra fiber+cement.
Dimension and thickness are under control.
We expect maximum 5 SFILs. a week for a SFIL production.
We contacted local company which provide cutting with Pb layers, as well as cutting for trapejoidal cross section.

Bulk section at CERN

- Reuse of $32 \times 3 \times 3 \text{ cm}^3$ modules, 3×6 array. $32 \times 9 \times 18 \text{ cm}^3$
- So far, we have used bundled fiber mached to the PMT dimension.
- This time, we will use Light Guide ($3 \times 3 \text{ cm}^2$ to SiPM array) for all 36 channels.

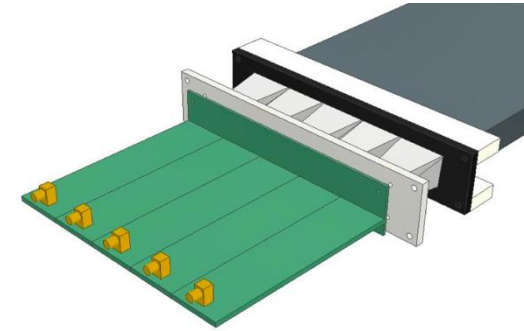
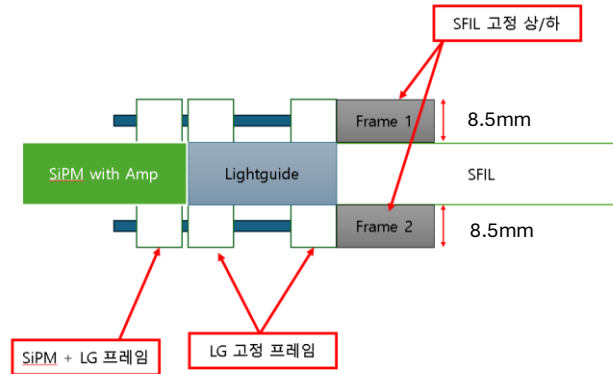
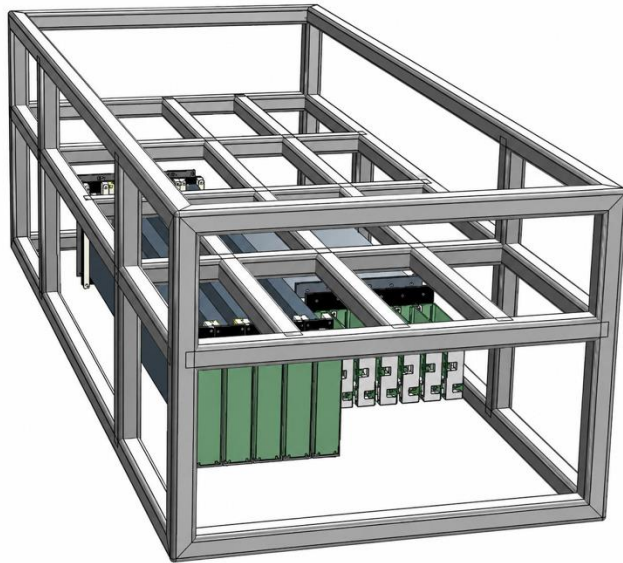


Dark box and fram



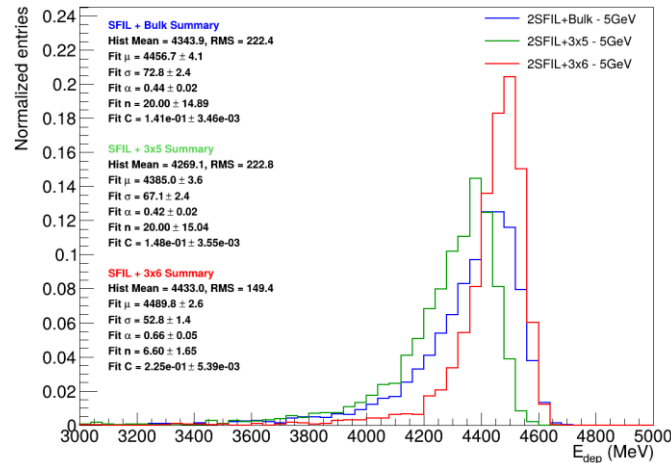
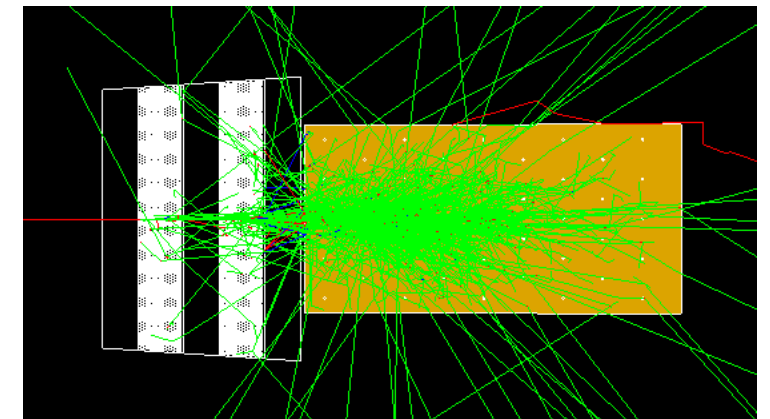
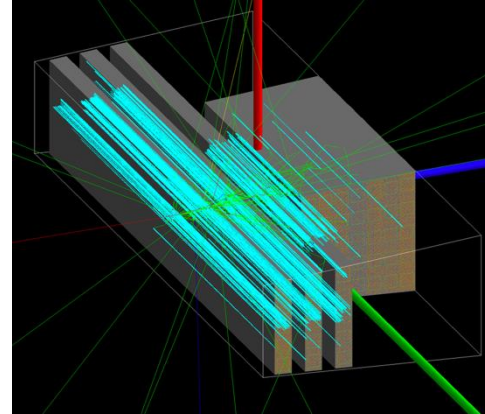
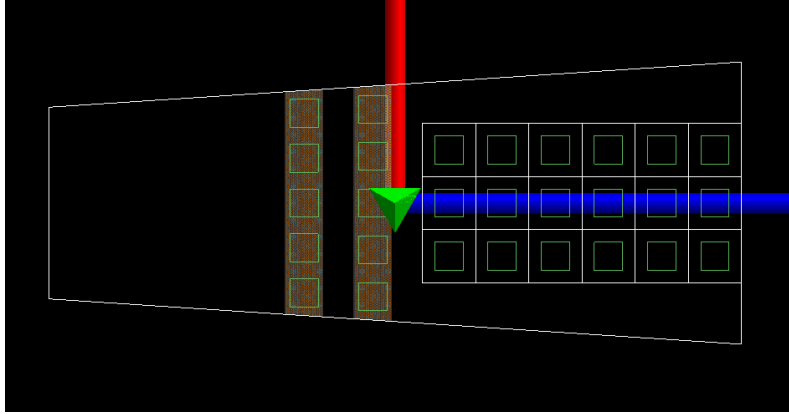
milling after cutting fiber bundle

Frame preparation (Junsub Shin)



- SFIL support to connect LG, SiPM, and FEE board, as well as 32x3x3cm³
- Frame to support
 - 5 SFILs(17mm distance) + bulk
 - H2GCROC
 - A-step and FPGA board for AstroPix

Simulation (Changhui Lee at KNU)



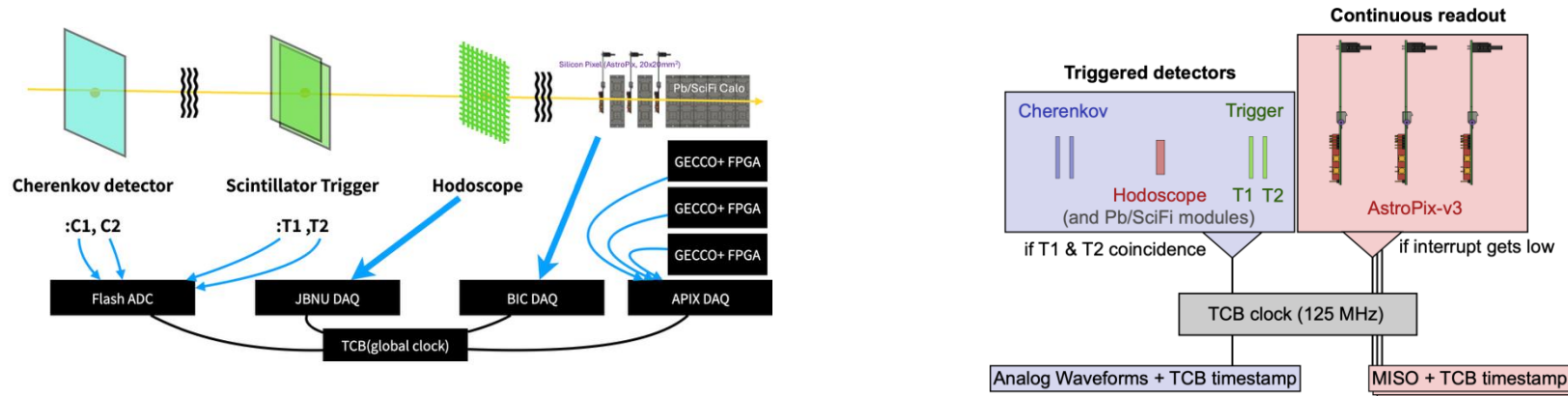
- Detector geometry were implemented. Electron energy deposit is being calculated for 2-5 SFILs

Goals

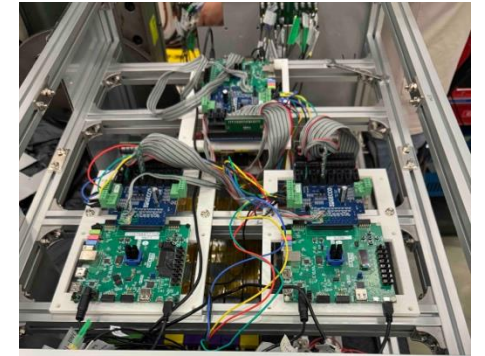
- H2GCROC data taking for the almost full depth with SFILs
- Confirm Npe for modules with exact material composition and single-clad fiber
- Confirm performance with FADC
- AstroPix data taking between SFILs

2. Timestamp matching at CERN

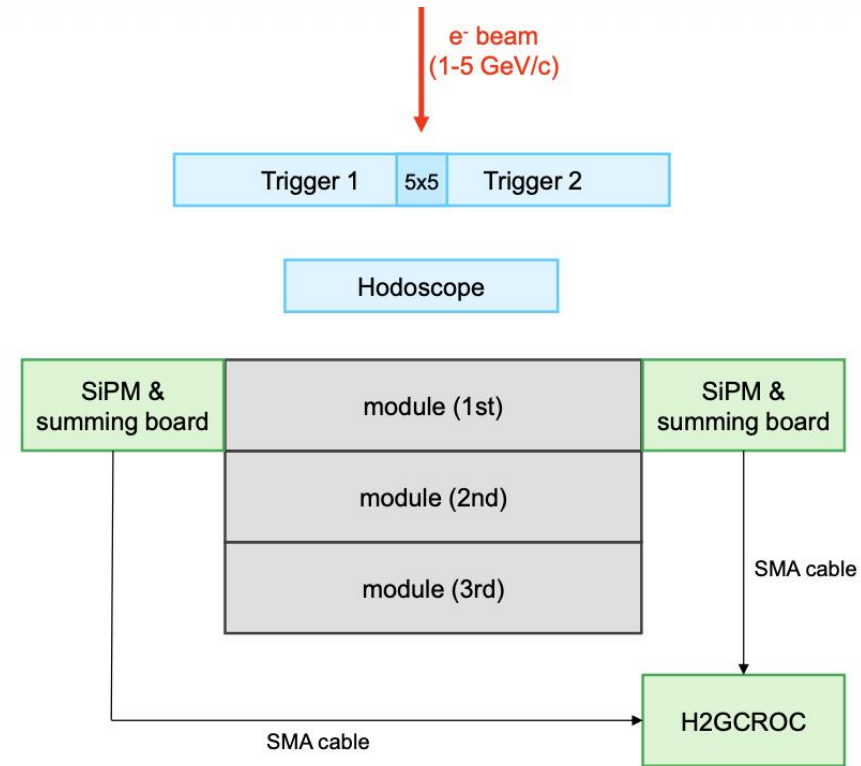
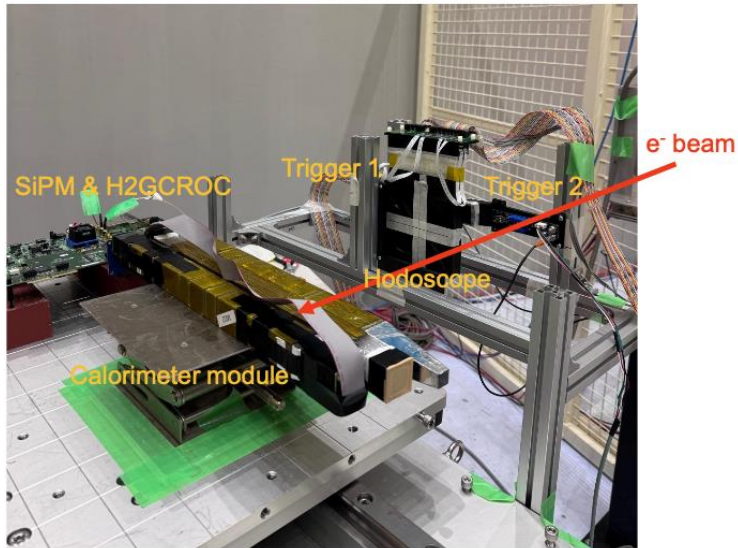
Setup at CERN in 2025



- Tested at KEK and used at CERN.
- When the interrupt signal is asserted, a dedicated DAQ module records the corresponding global timestamp from the Trigger Control Board (TCB) together with the SPI data stream.
- Not available this time
 - It was a temporary solution in 2025. A-step board interface does not support MISO, interrupt signal like last year.(?)

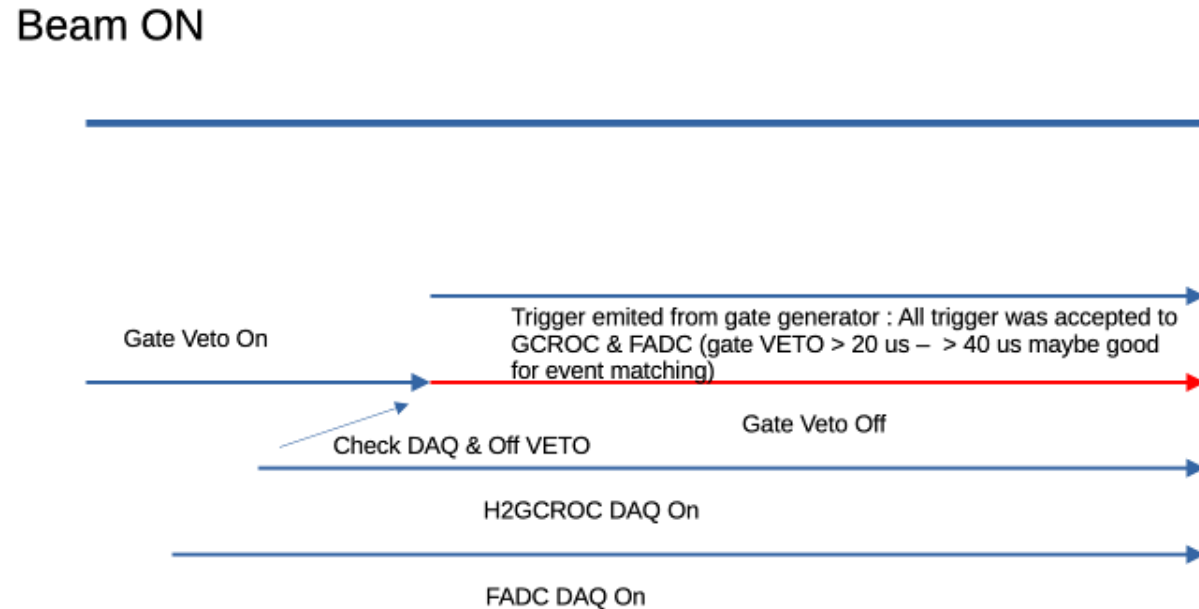


2-a) Trg-H2GCROC timestamp matching at KEK



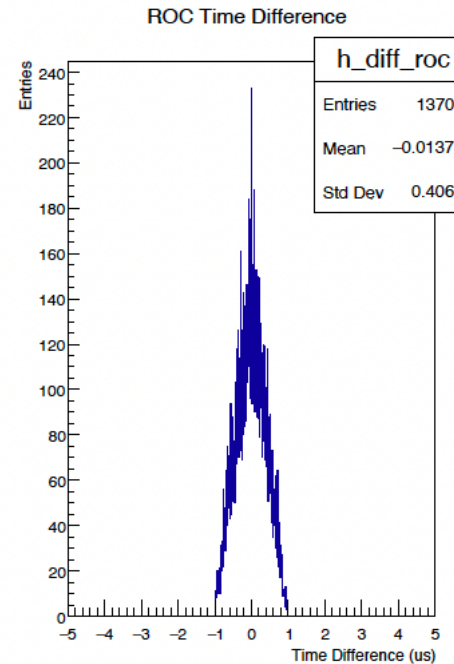
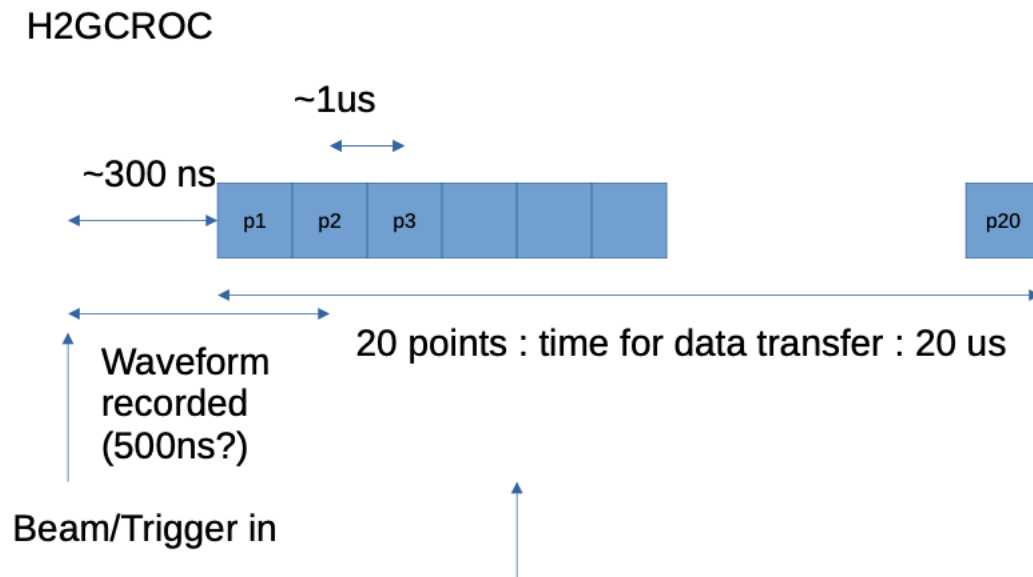
- Setup at KEK Data (Auxiliary, H2GCROC) recorded separately with common trigger.

2-a) Trg-H2GCROC timestamp matching at KEK



- Using veto, we made the same starting point for the H2GCROC and FADC readout.

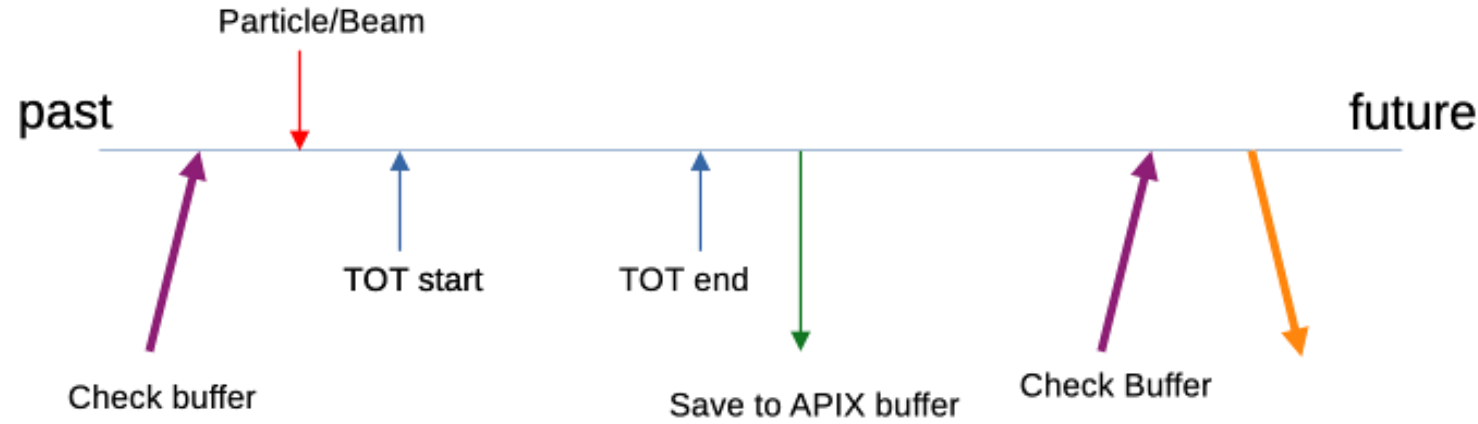
2-a) Trg-H2GCROC timestamp matching at KEK



- By using common trigger, we could match trigger and auxiliary detectors with H2GCROC FPGA time within ± 1 us

	Trg time	H2GCROC time			
Row	(ts_nk-ts_	(ts_roc-t	ts_nk-ts_	ts_nk_dif	ts_roc_di
0	*	0	*	0	*
1	*	2523.12	*	2523.1	*
2	*	5518.16	*	5518.15	*
3	*	10531.68	*	10531.65	*
4	*	15625.76	*	15625.725	*
5	*	22421.216	*	22421.15	*
6	*	46346.24	*	46346.125	*
7	*	58148.928	*	58148.8	*
8	*	60604.112	*	60603.975	*
9	*	61754.304	*	61754.15	*
10	*	64517.776	*	64517.65	*
11	*	65938.528	*	65938.4	*
12	*	78301.232	*	78301.05	*
13	*	84341.632	*	84341.45	*
14	*	88947.44	*	88947.25	*
15	*	103211.6	*	103211.37	*
16	*	109698.73	*	109698.5	*
17	*	112874.97	*	112874.72	*
18	*	119594.92	*	119594.67	*
19	*	126591.72	*	126591.45	*
20	*	134161.10	*	134160.82	*
21	*	145785.10	*	145784.77	*
22	*	156766.04	*	156765.7	*
23	*	158141.50	*	158141.15	*
24	*	160645.74	*	160645.4	*
25	*	170304.09	*	170303.72	*
26	*	173492.92	*	173492.55	*
27	*	192512.65	*	192512.22	*
28	*	210208.51	*	210208.05	*
29	*	211322.20	*	211321.75	*
30	*	211346.12	*	211345.67	*
31	*	220988.11	*	220987.62	*
32	*	226151.39	*	226150.9	*
33	*	227338.08	*	227337.6	*
34	*	229716.49	*	229716	*
35	*	231441.77	*	231441.27	*
36	*	235613.44	*	235612.92	*
37	*	242084.22	*	242083.7	*
38	*	243569.15	*	243568.62	*
39	*	251617.98	*	251617.45	*
40	*	266243.31	*	266242.72	*
41	*	278595.93	*	278595.32	*
42	*	284097.72	*	284097.12	*
43	*	286289.88	*	286289.27	*
44	*	293850.46	*	293849.82	*
45	*	294036.70	*	294036.07	*
46	*	299754.96	*	299754.3	*
47	*	306722.81	*	306722.15	*
48	*	309632.27	*	309631.6	*
49	*	310491.77	*	310491.1	*

AstroPix data and time



- In our point of view, it seems that
- Astropix periodically checks whether the buffer is full and retrieves the data.
- FPGA time looks like the time at which data arrives at the FPGA.

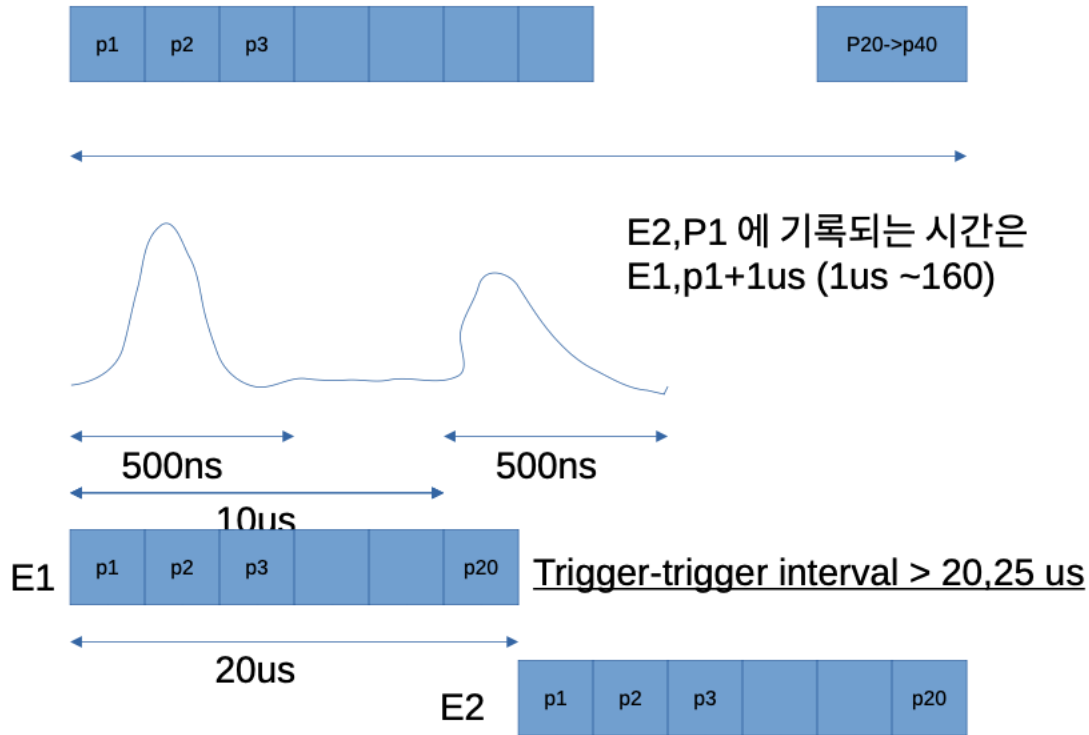
Idea for AstroPix-H2GCROC matching

- Recording the timestamp when the trigger arrives at AstroPix.
 - 0x0A[data] [time at trigger] 0x0A[data]
- Benefit
 - We don't have to worry about the starting time.
 - We don't have to worry about the non-trigger event at AstroPix

Summary

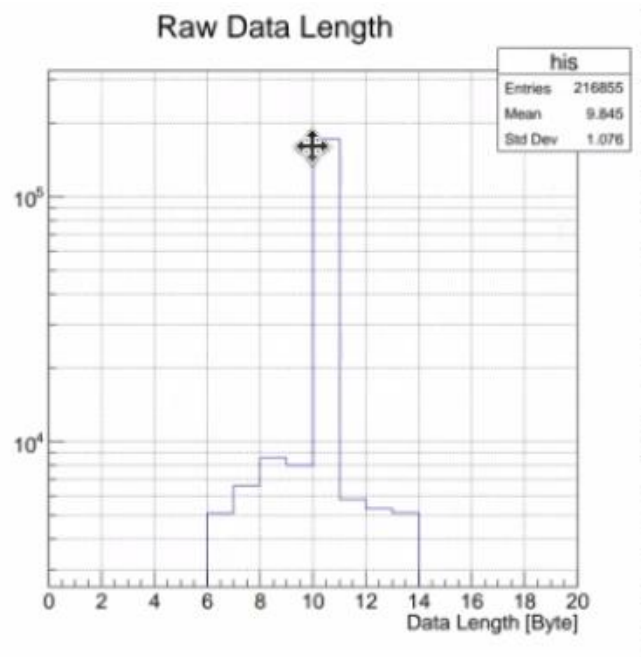
- We are being prepared for the CERN beam test
 - Pb/SciFi modules
 - Experimental setups
 - Event matching between Auxiliary detector and H2GCROC
- We need more careful approach and firmware update for AstroPix-H2GCROC synchronization

Backup: limitation in rate for H2GCROC



- When another trigger arrived while 20 points (20us in timestamp) are recorded, the next event is recorded after 20us. Maximum 50kHz, but we will keep it about a few kHz rate.

0x0A in AstroPix?



- About 20% of data are out of 10 byte length.