

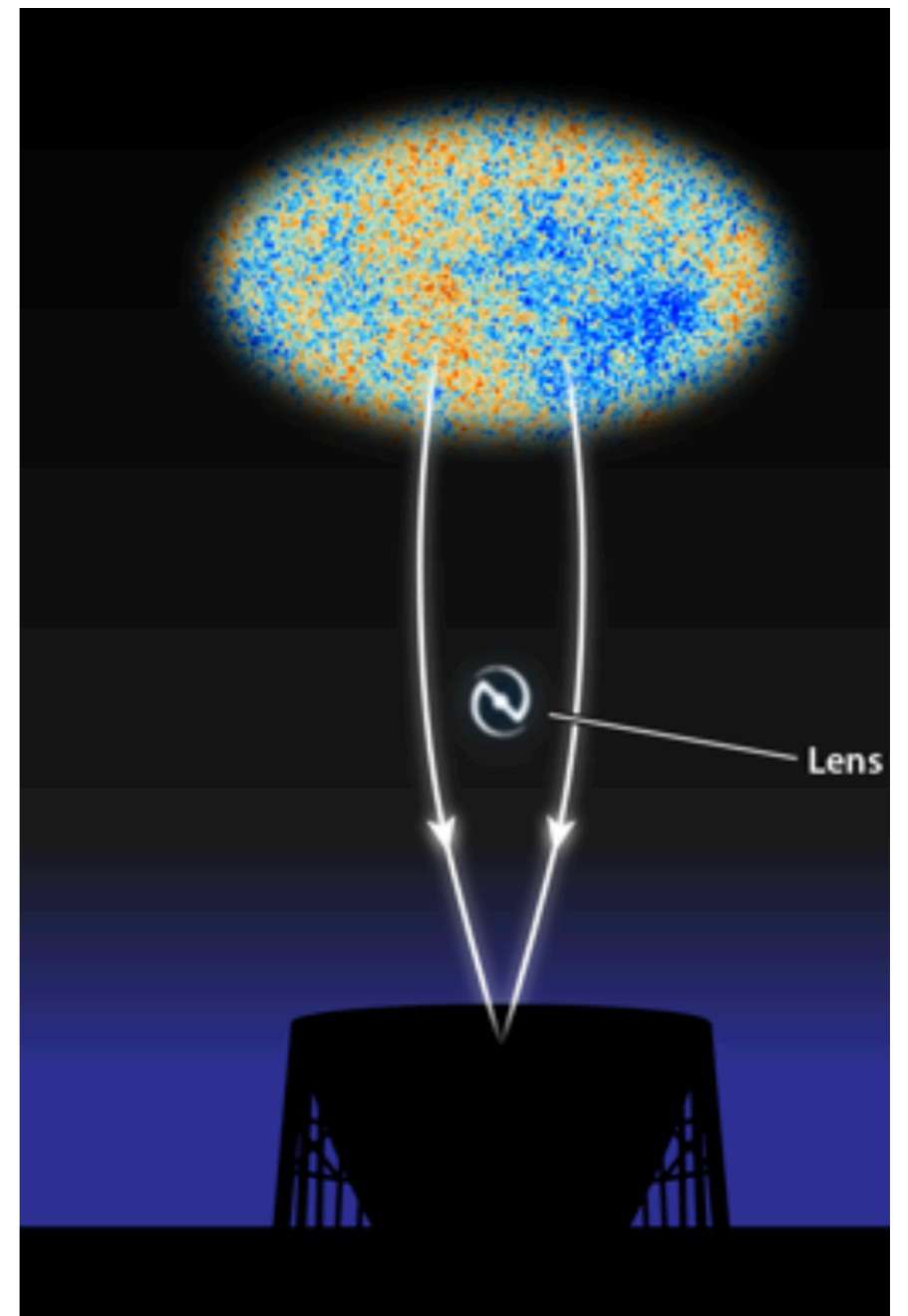
# Measuring the Small-Scale Matter Power Spectrum with Ultra-High-Resolution CMB Lensing

Neelima Sehgal

BNL Dark Interactions

Oct. 3rd, 2018

Ho Nam Nguyen, NS, Mathew Madhavacheril, 2017,  
arXiv:1710.03747



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**Key Question: What do matter fluctuations look like on small-scales?**

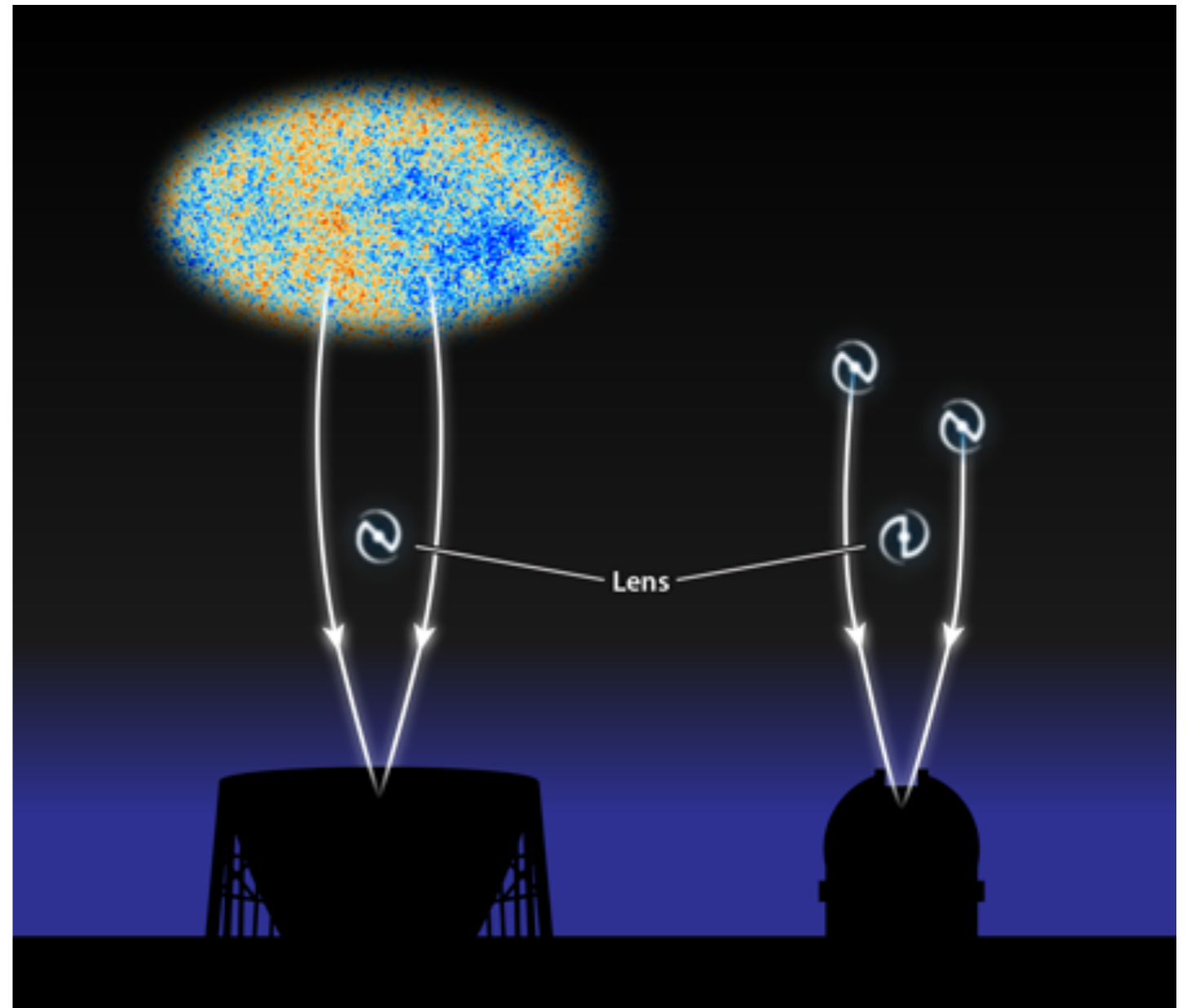
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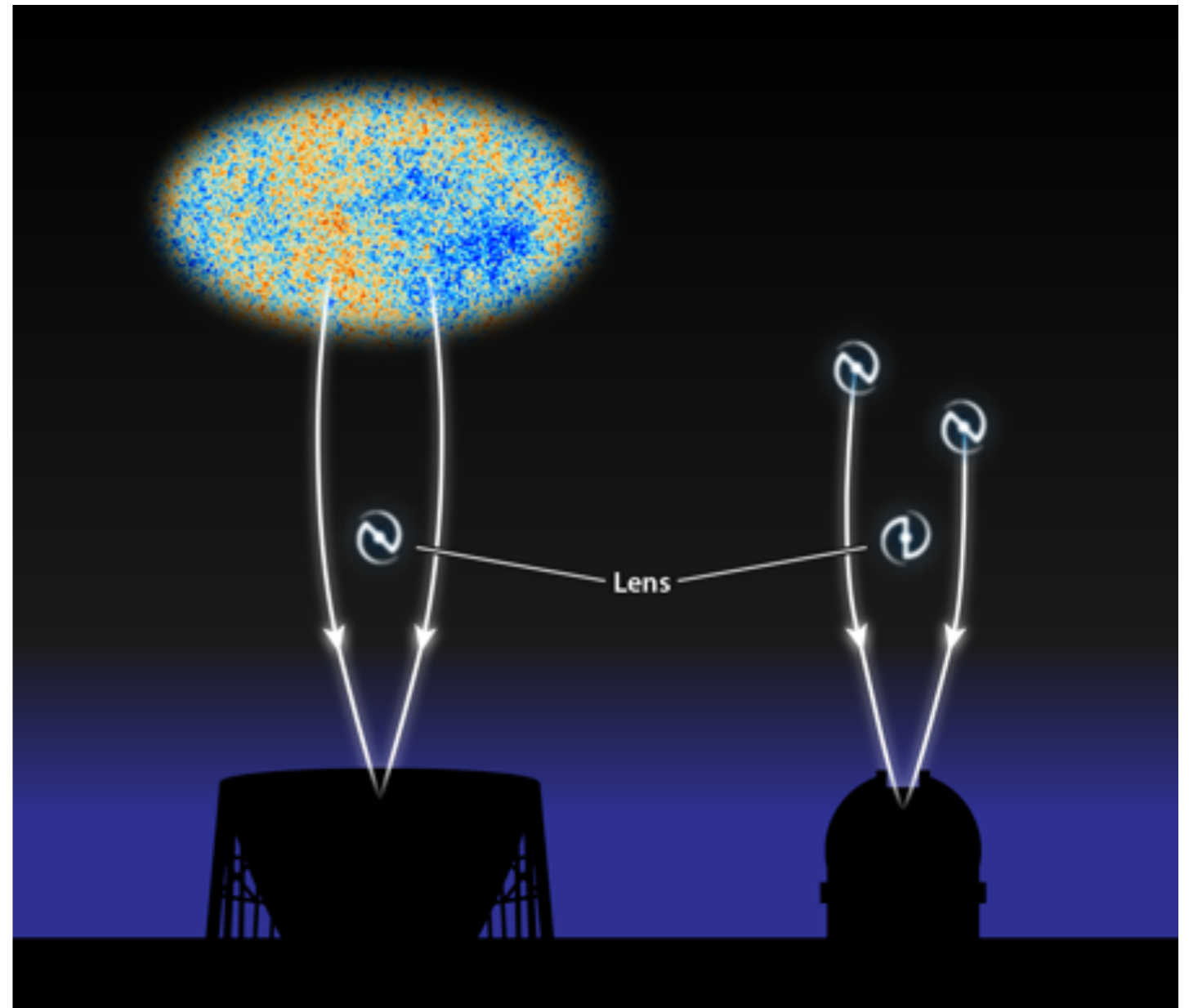
**A technique that relies on lensing avoids complications of baryonic tracers.**

# Gravitational Lensing of the Cosmic Microwave Background (CMB)



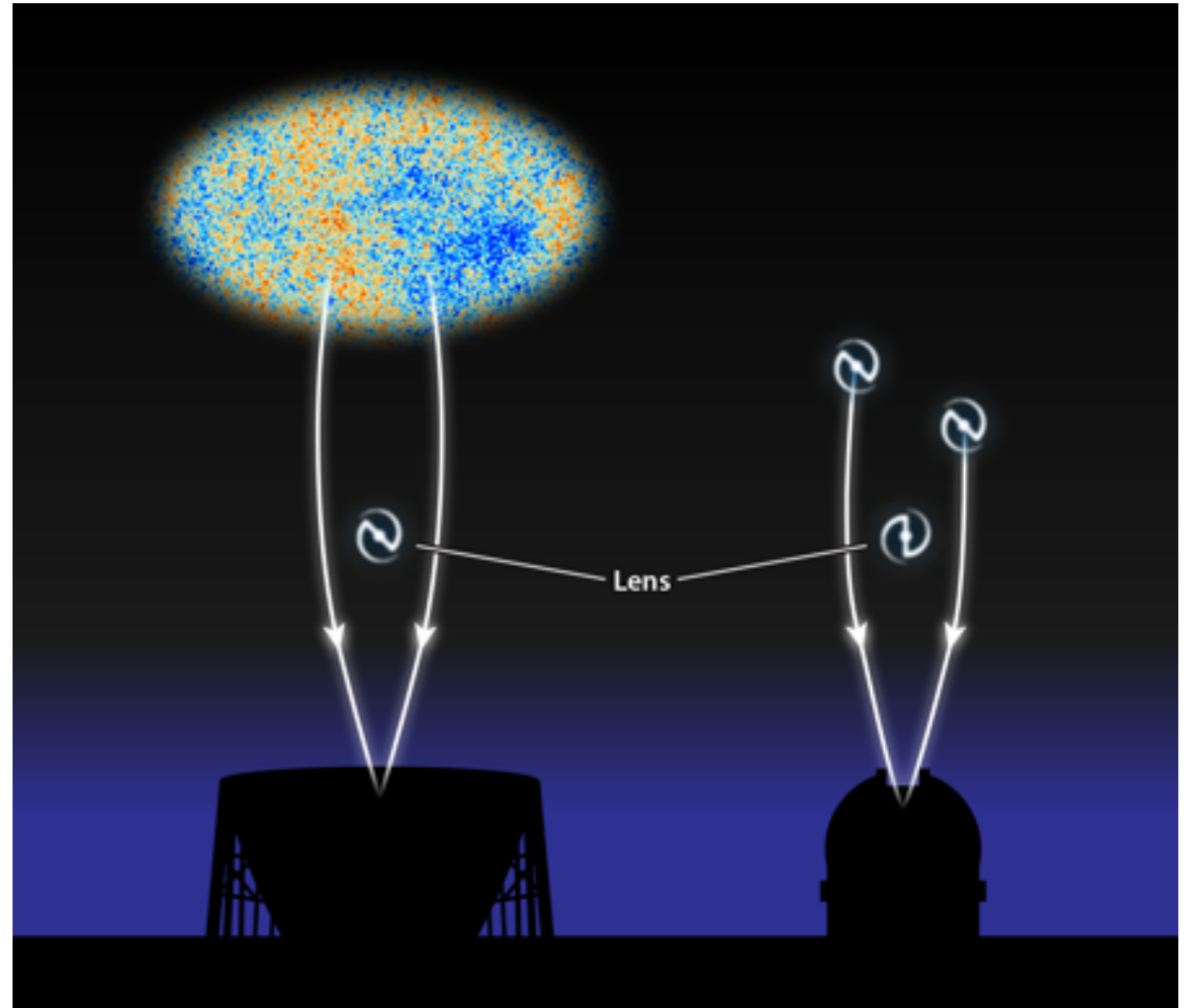
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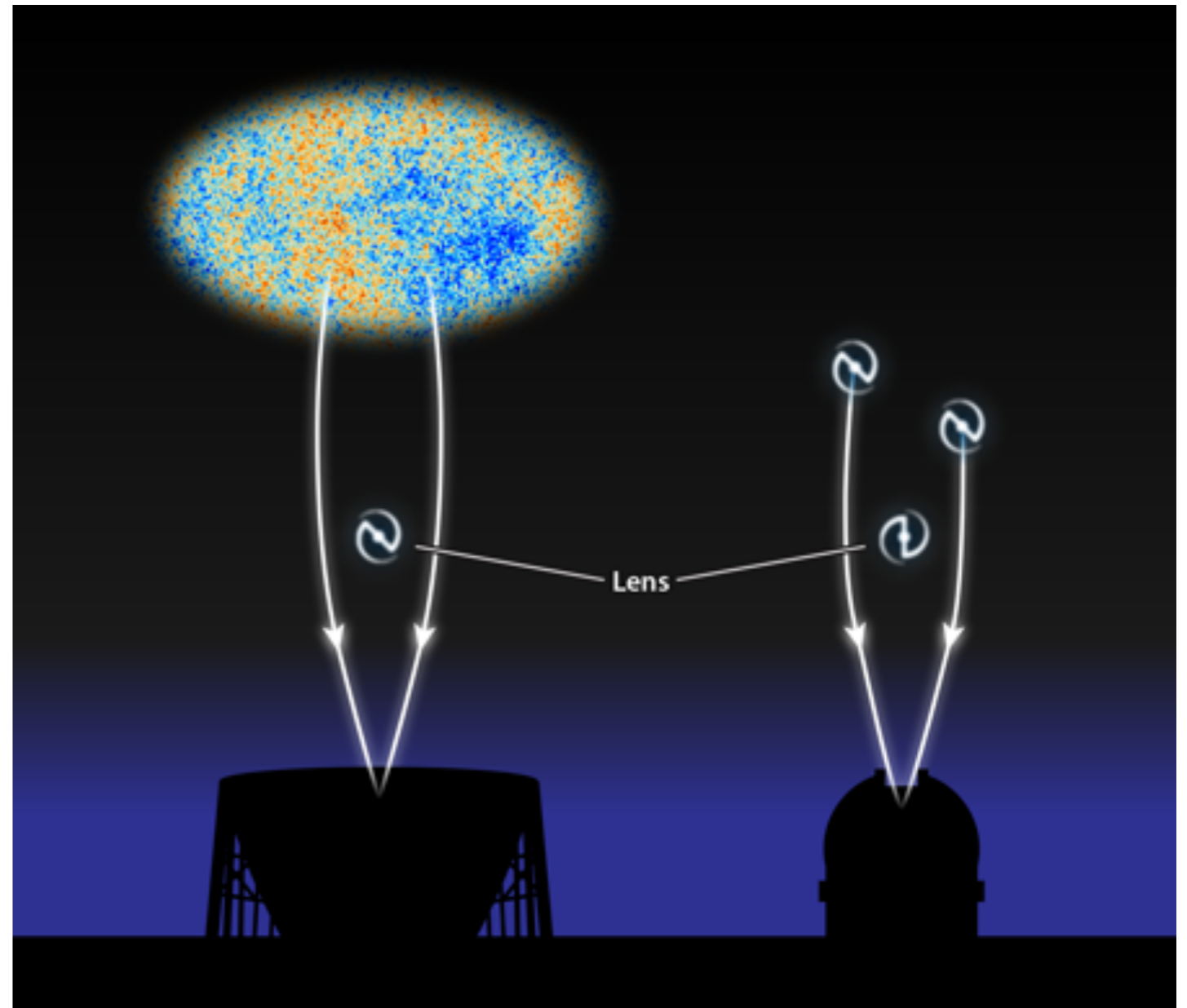
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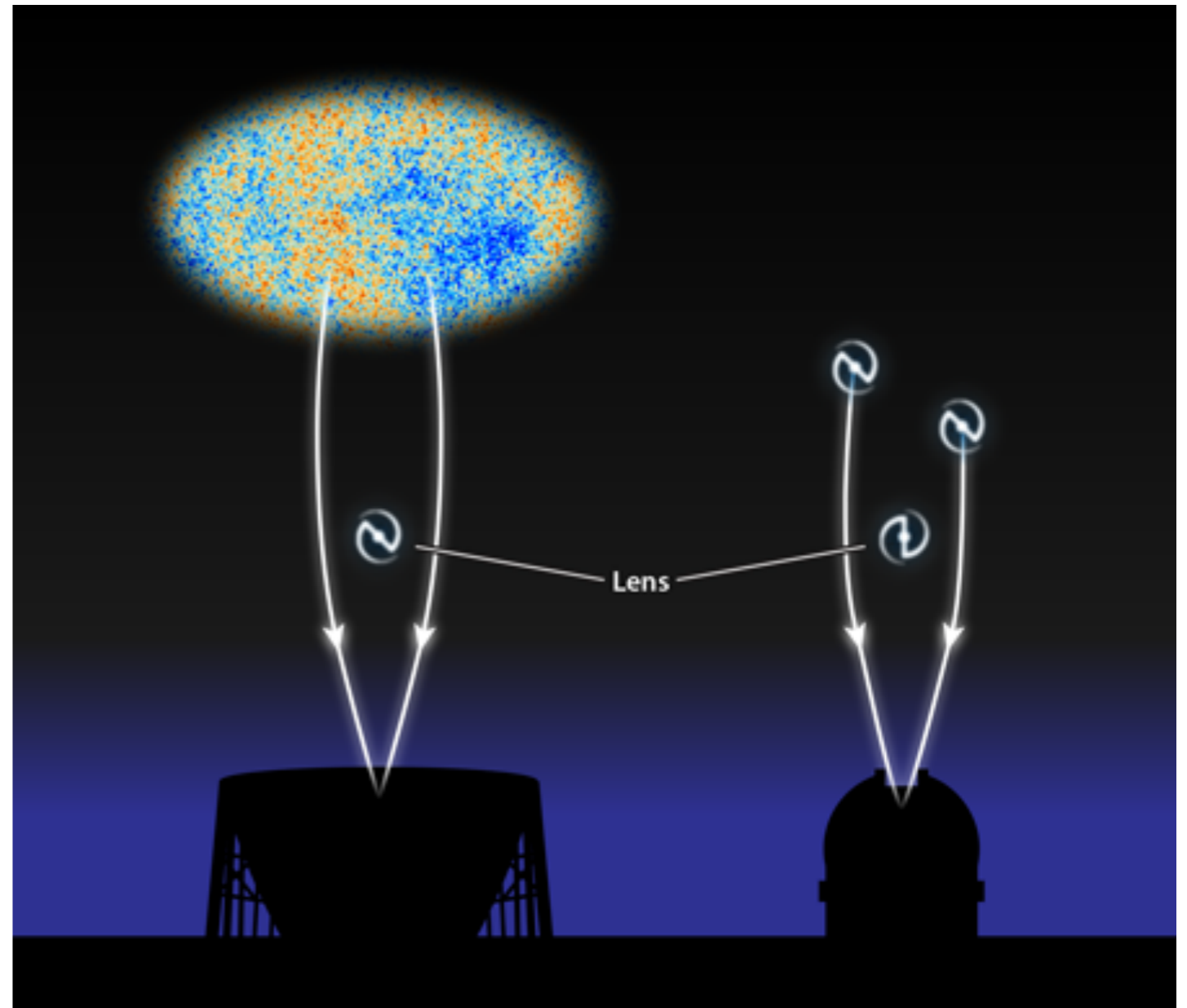
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First Measurement of CMB Lensing on Halo Scales  
Madhavacheril, NS, for the ACT Collaboration  
PRL, 114, 2015

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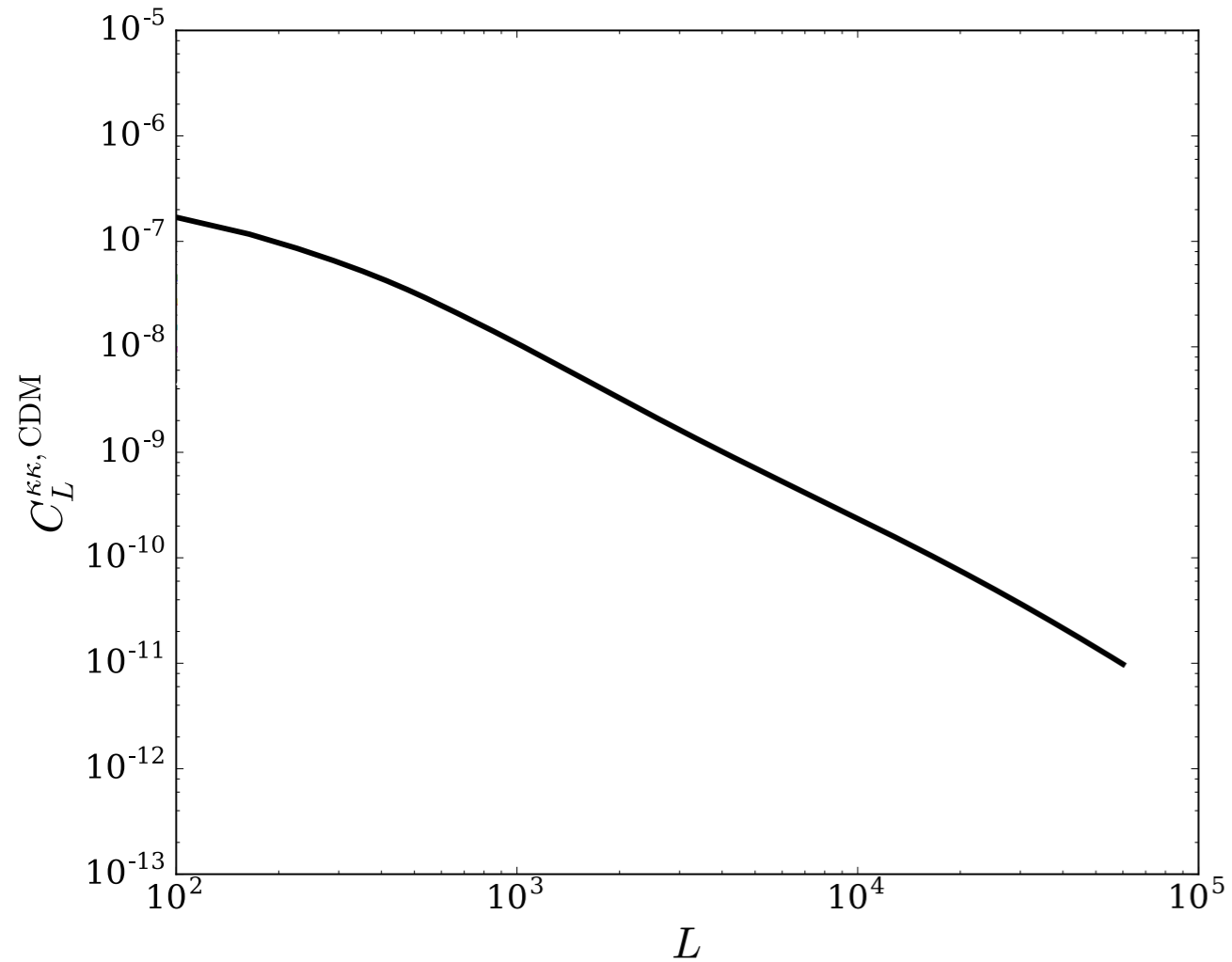
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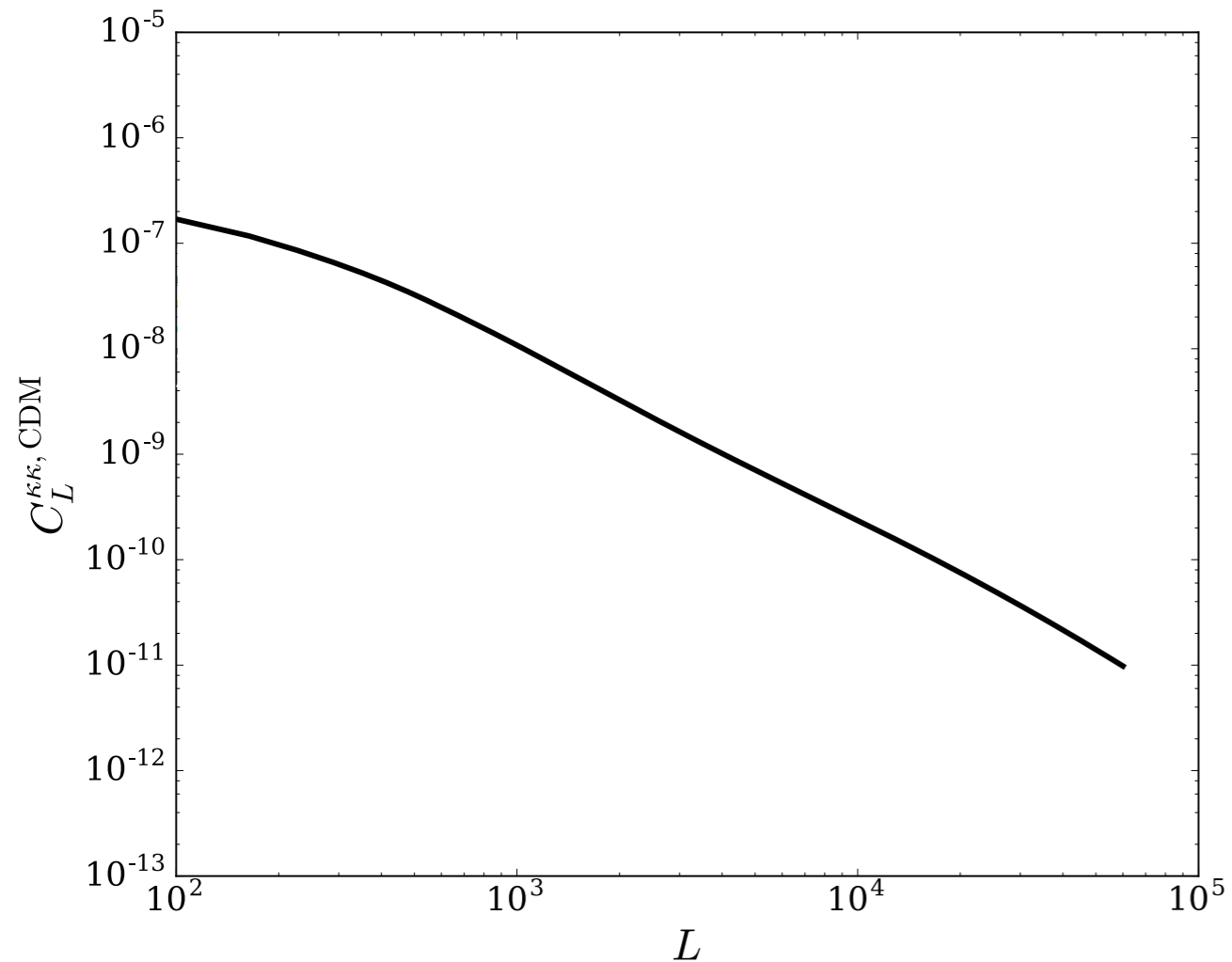
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4. Sensitive to structure at higher redshifts than other gravitational lensing probes; this makes it more sensitive to FDM/WDM-type models

# CMB Lensing Power Spectrum

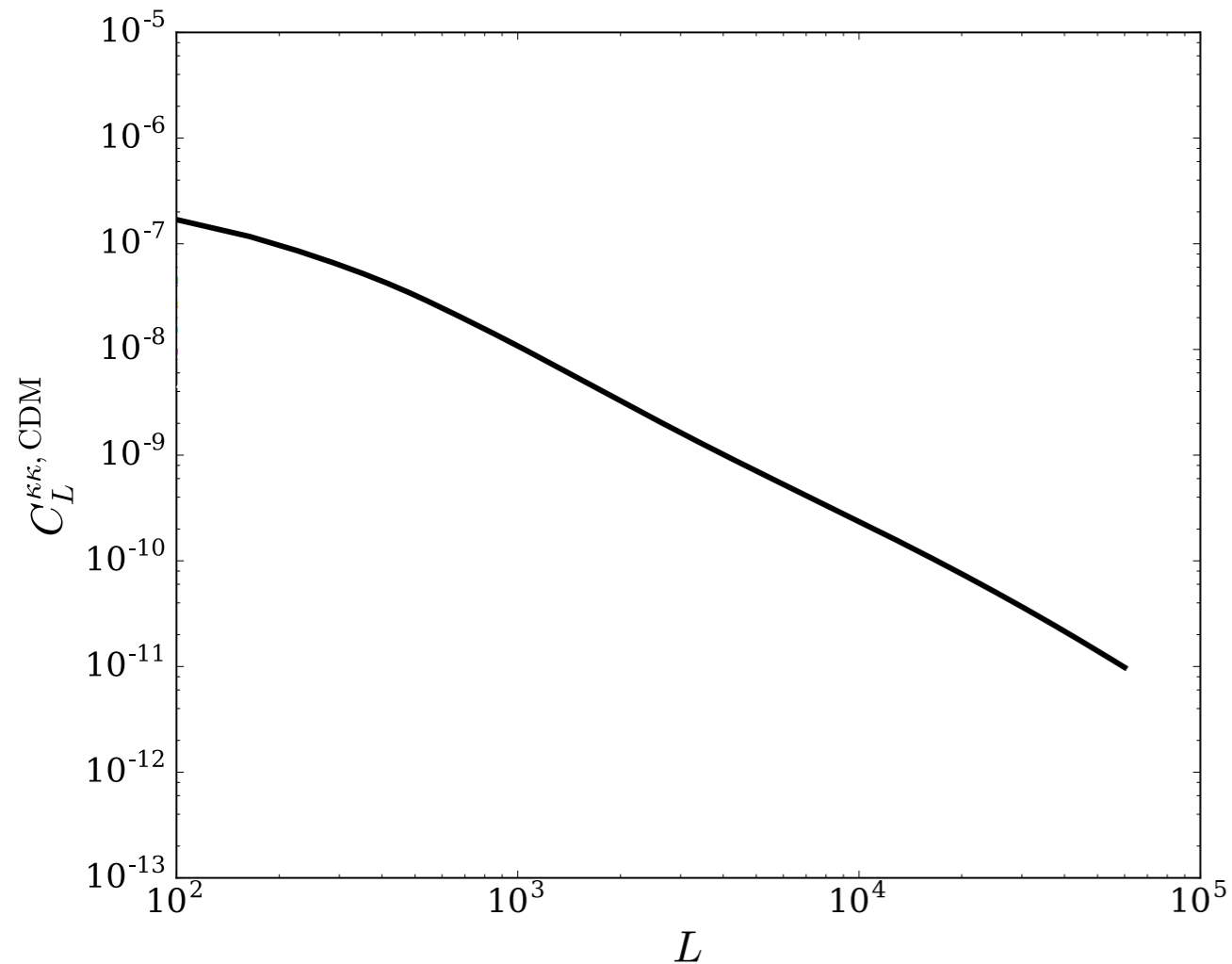


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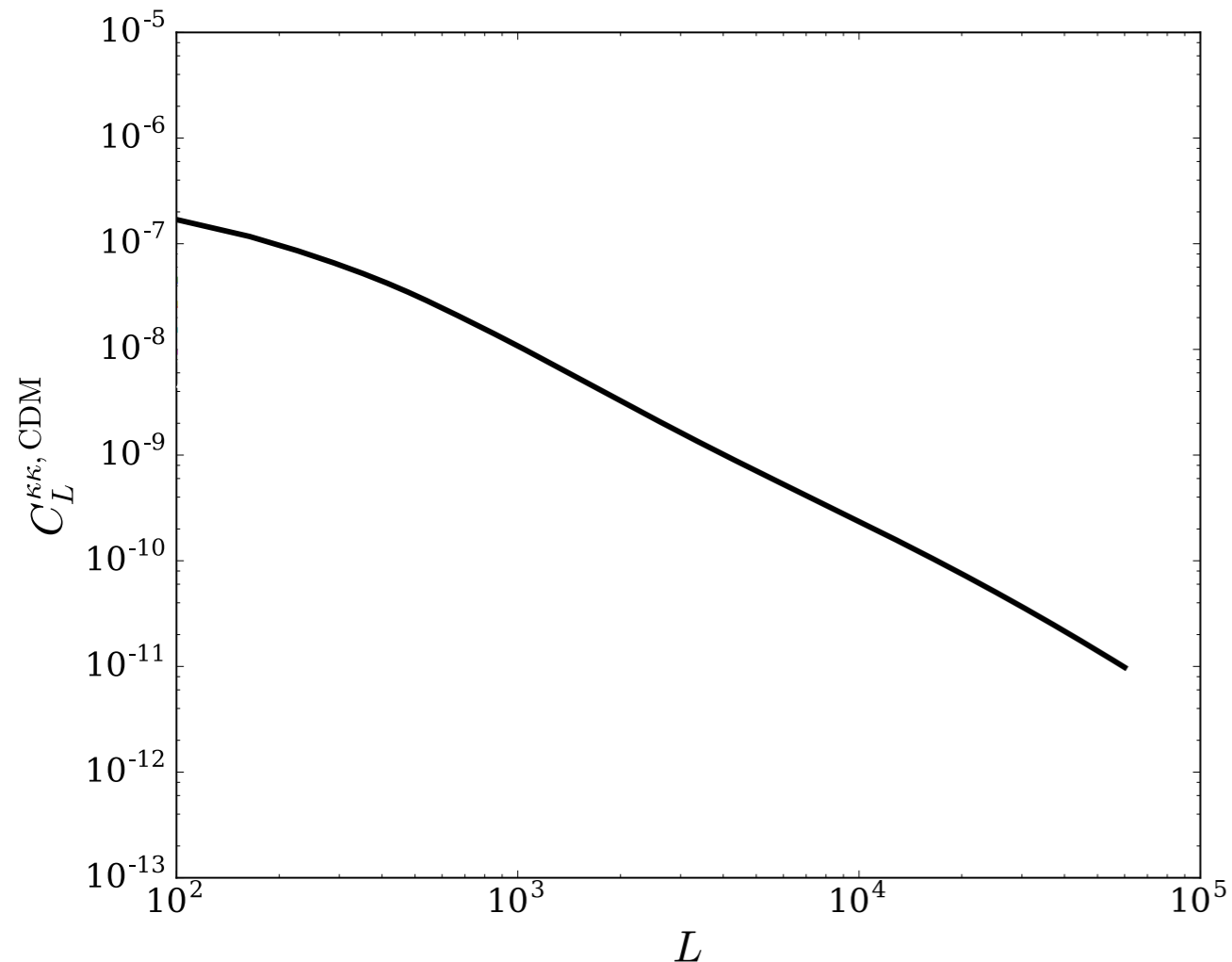
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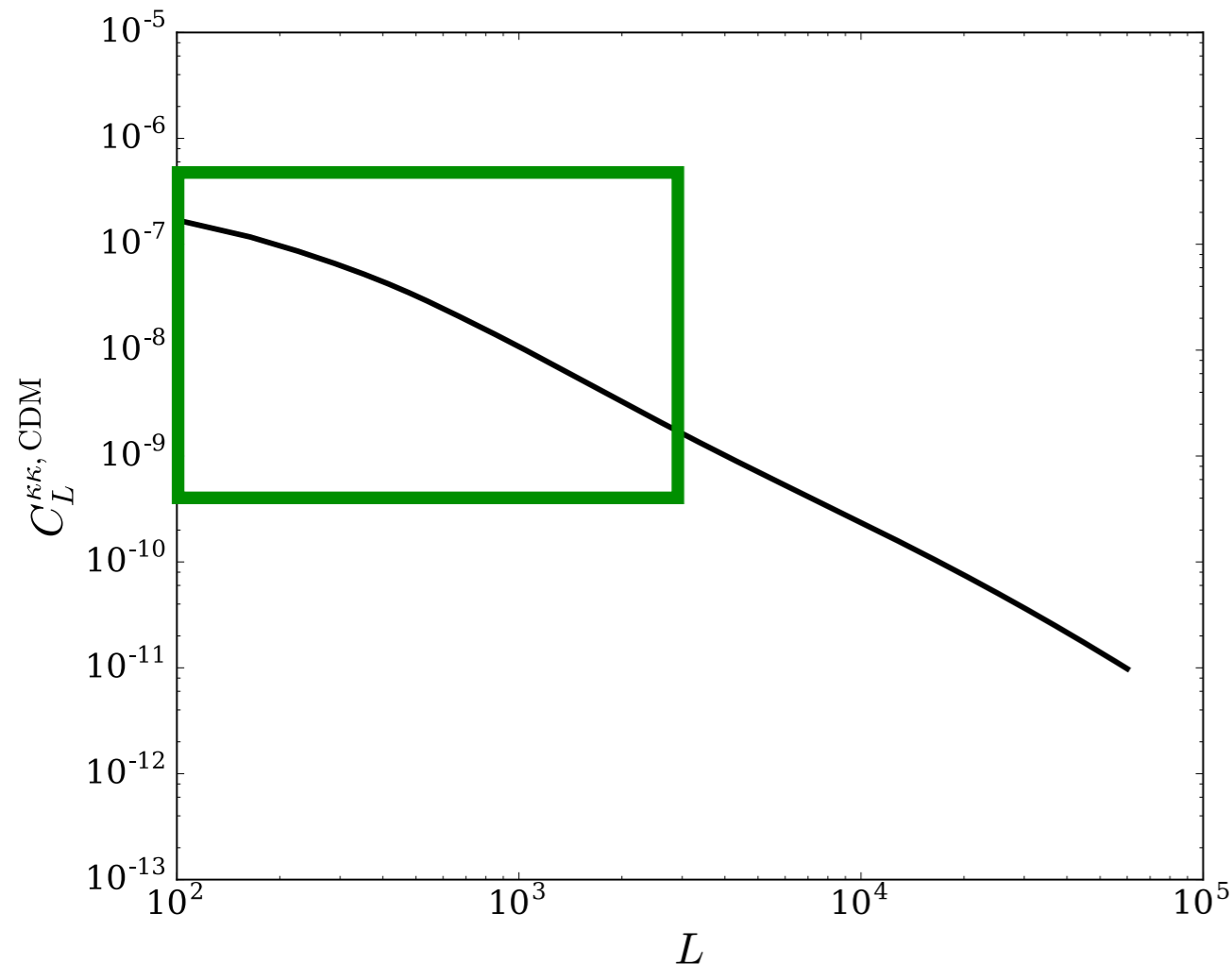
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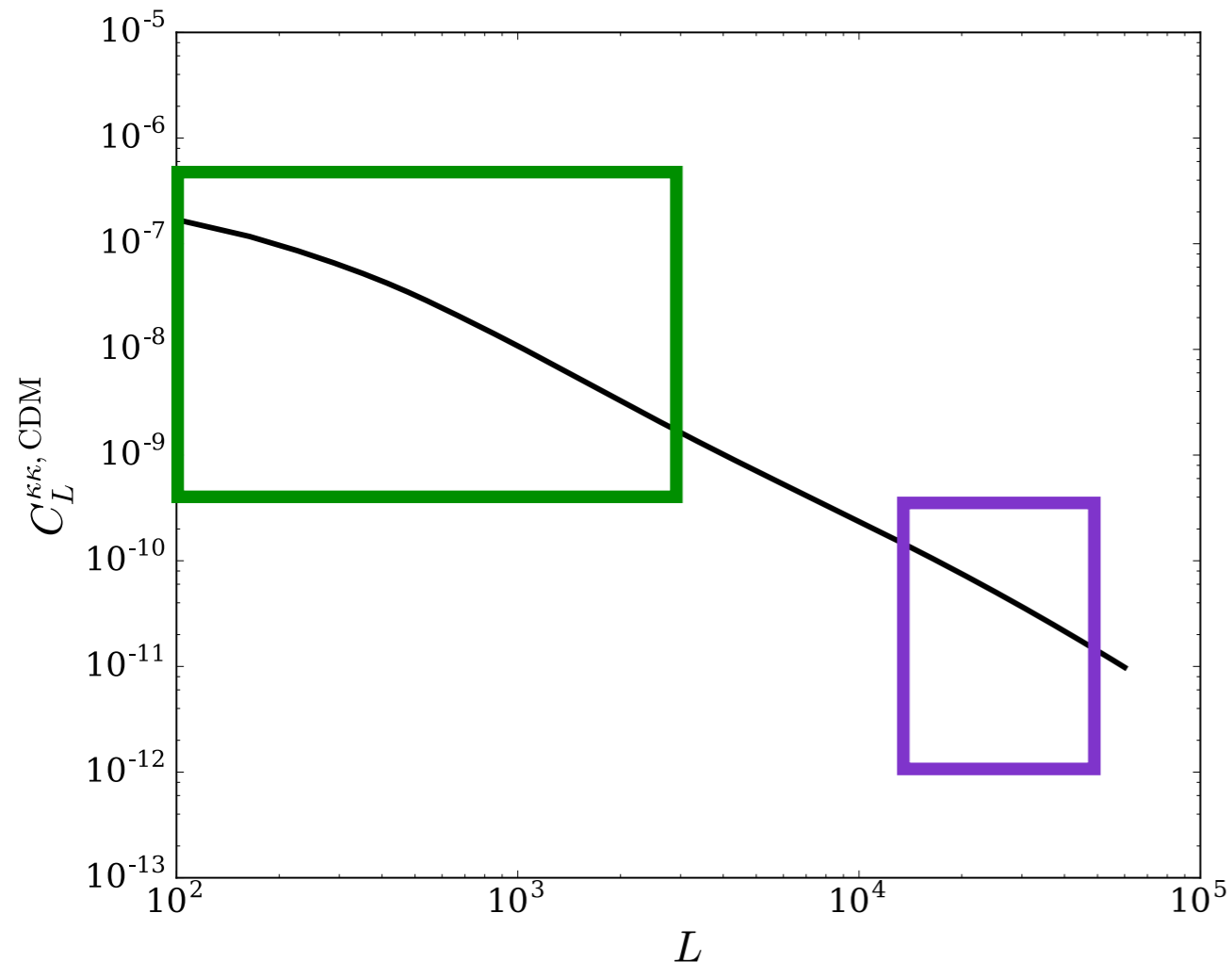
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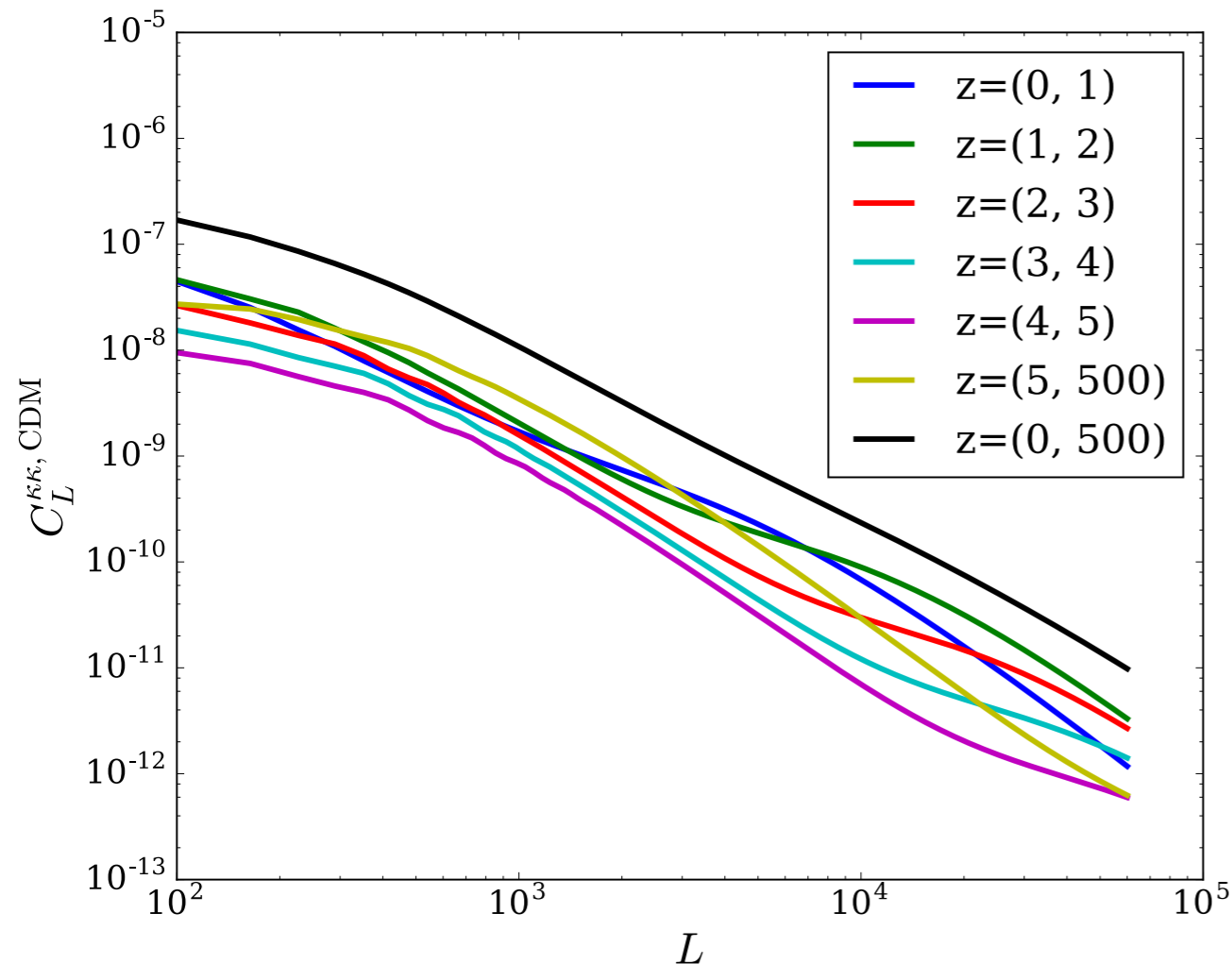
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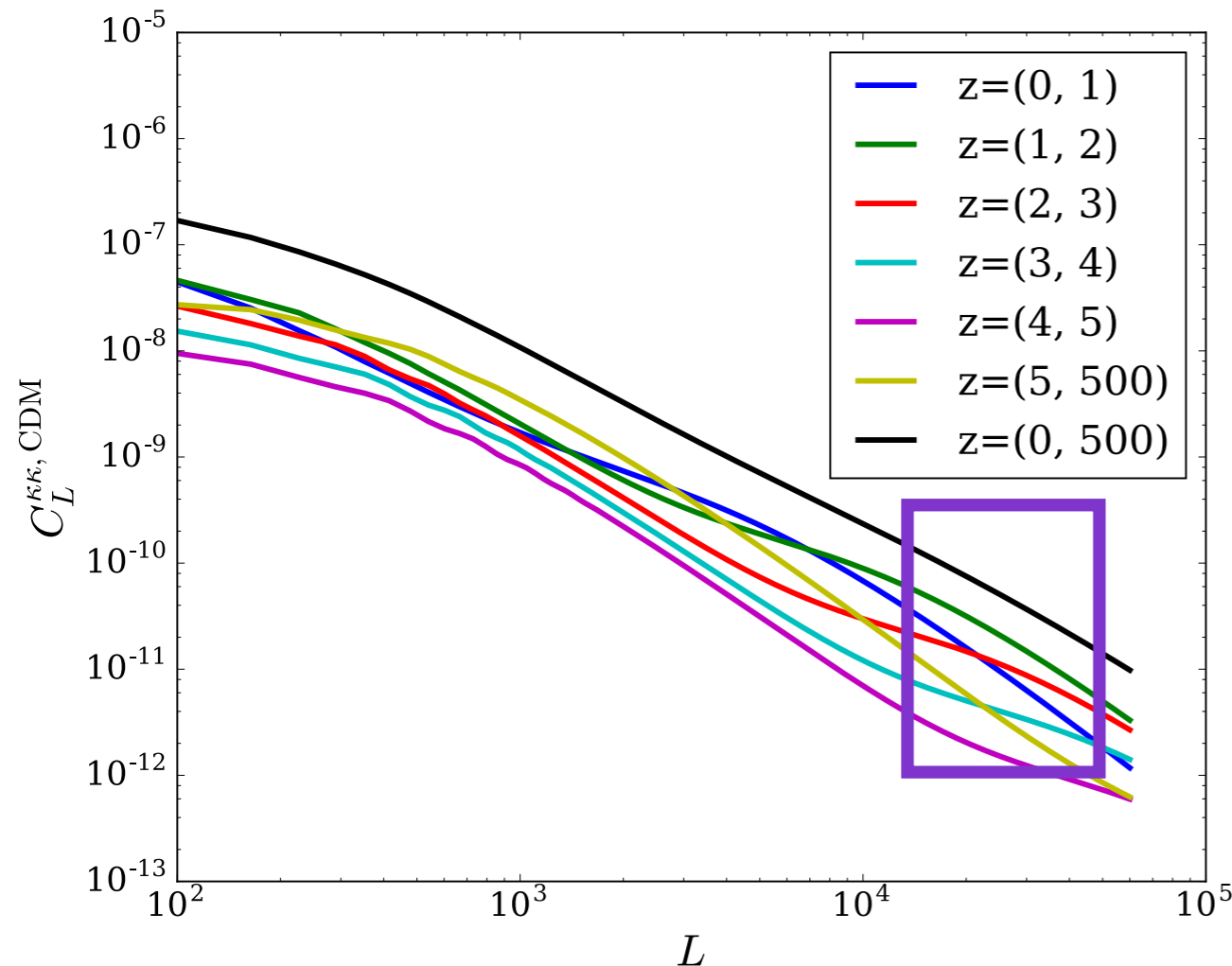
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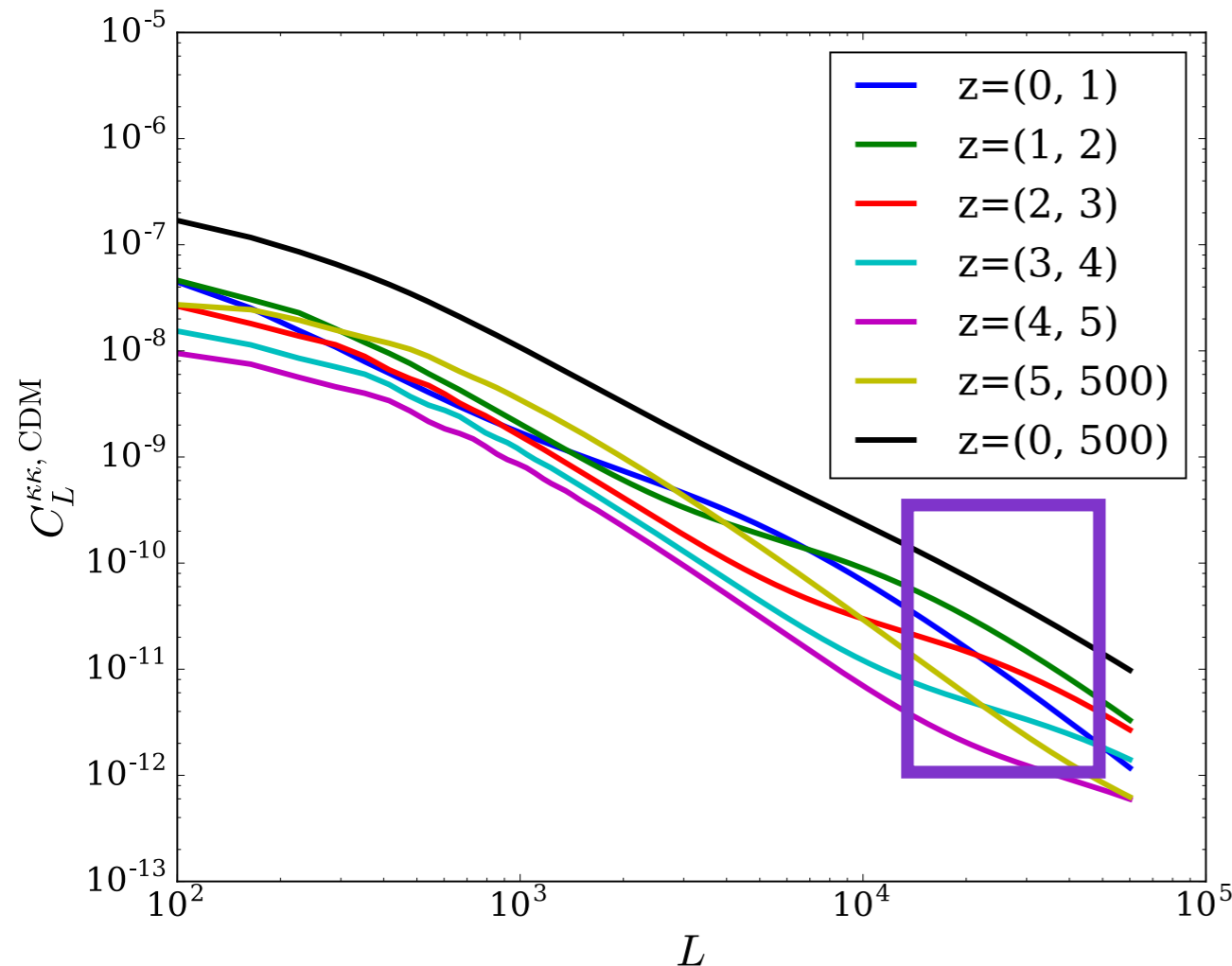
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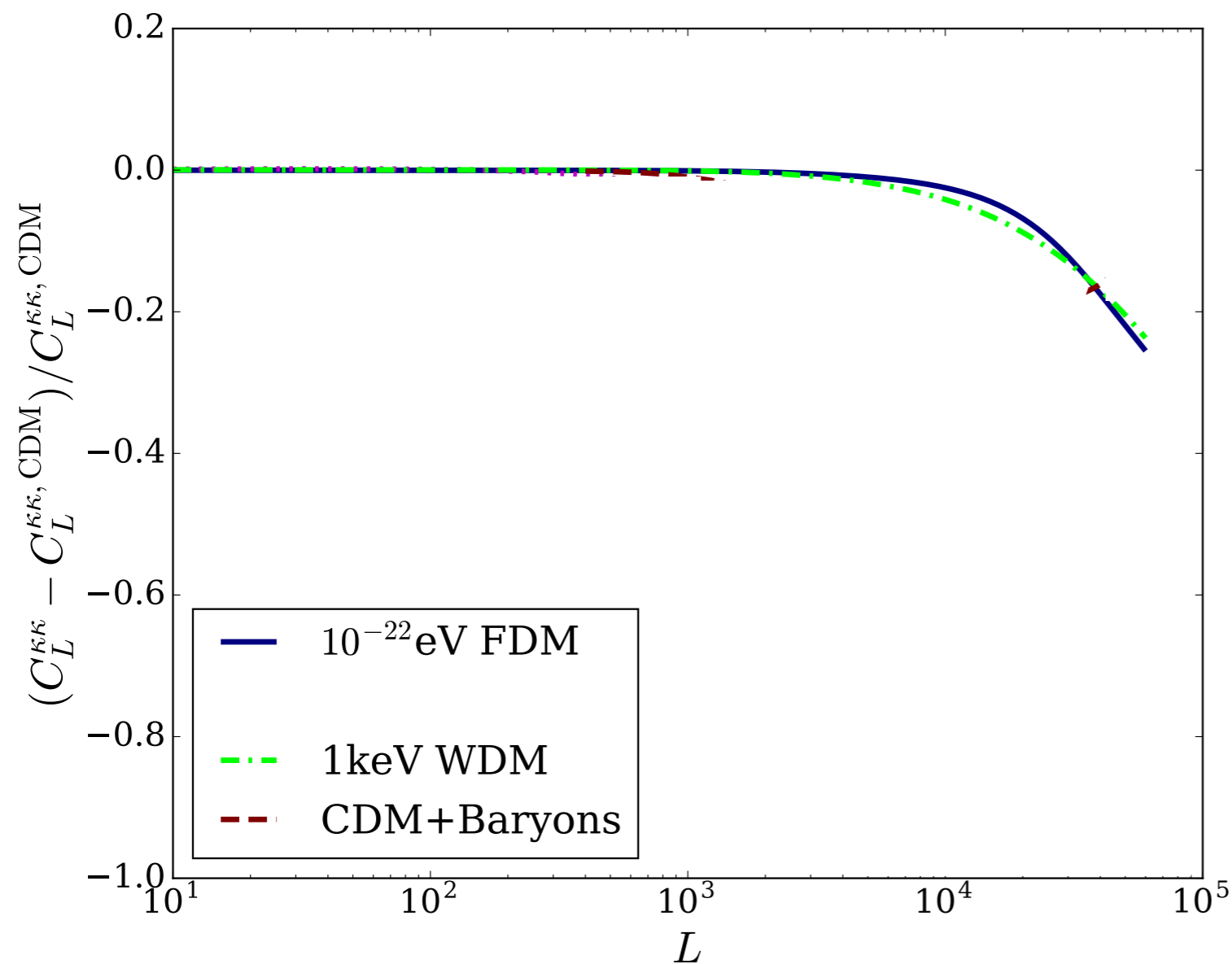
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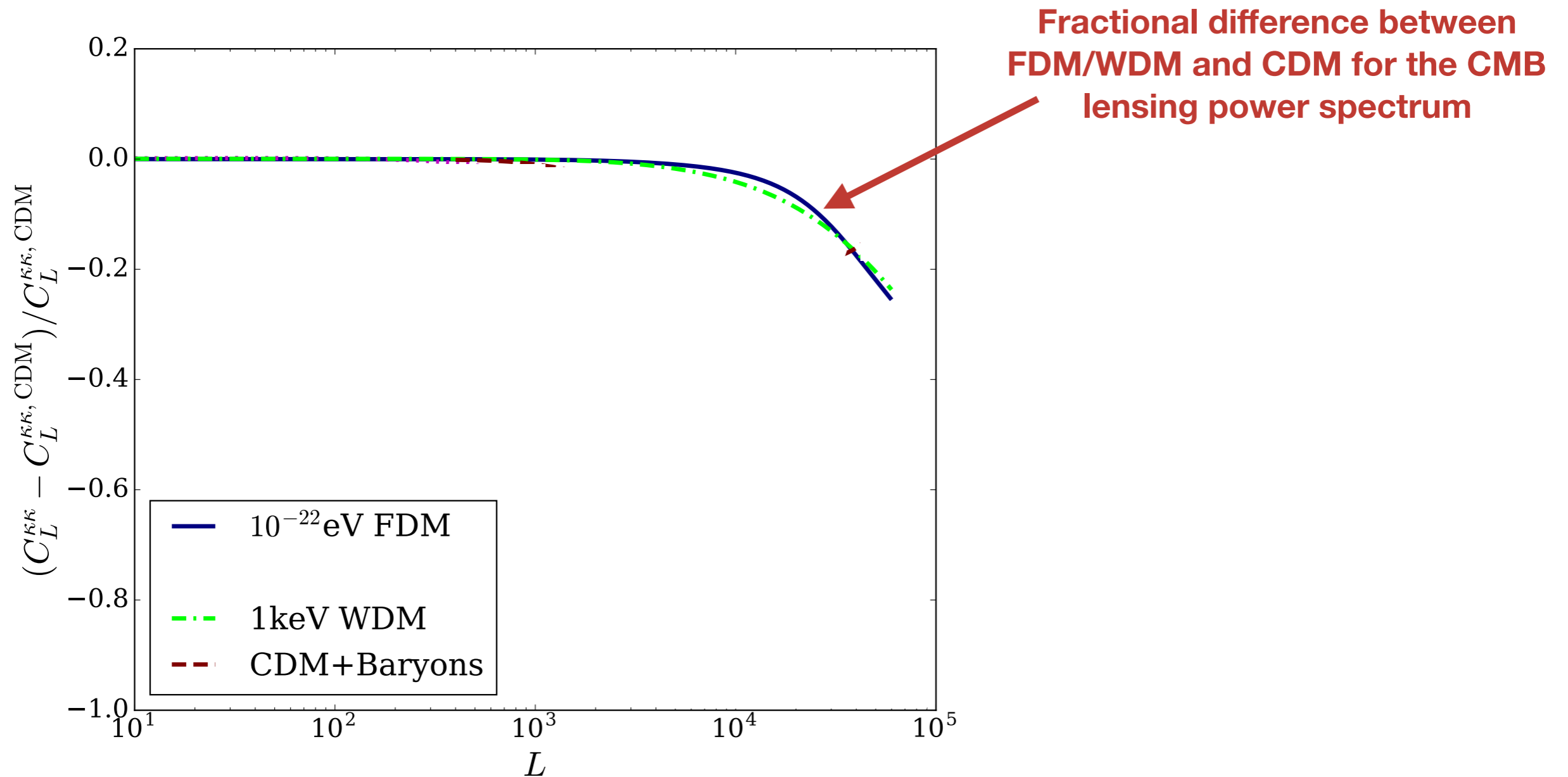
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**Contrast between CDM and DM models that wash out small-scale structure is larger at higher redshifts**

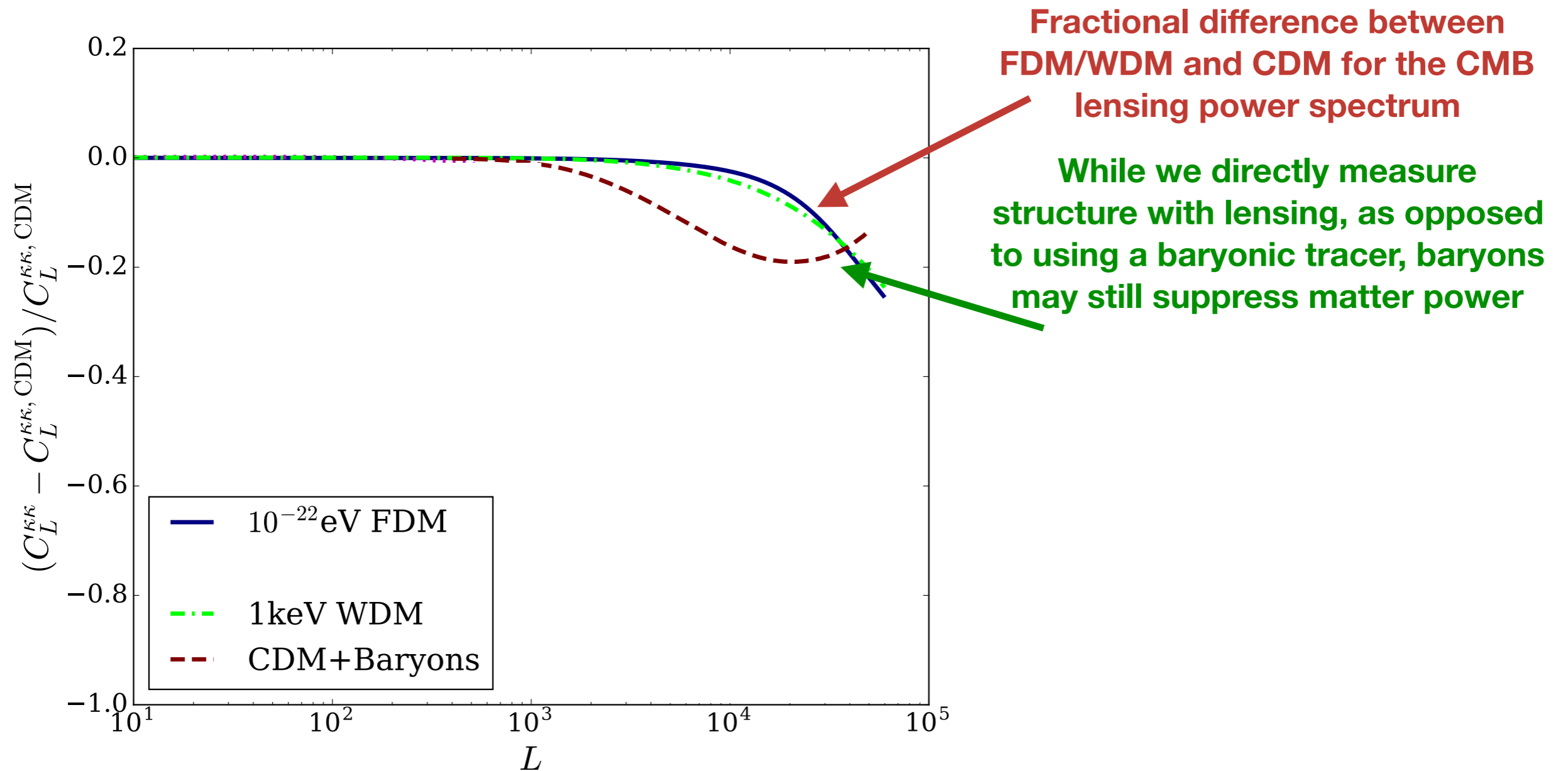
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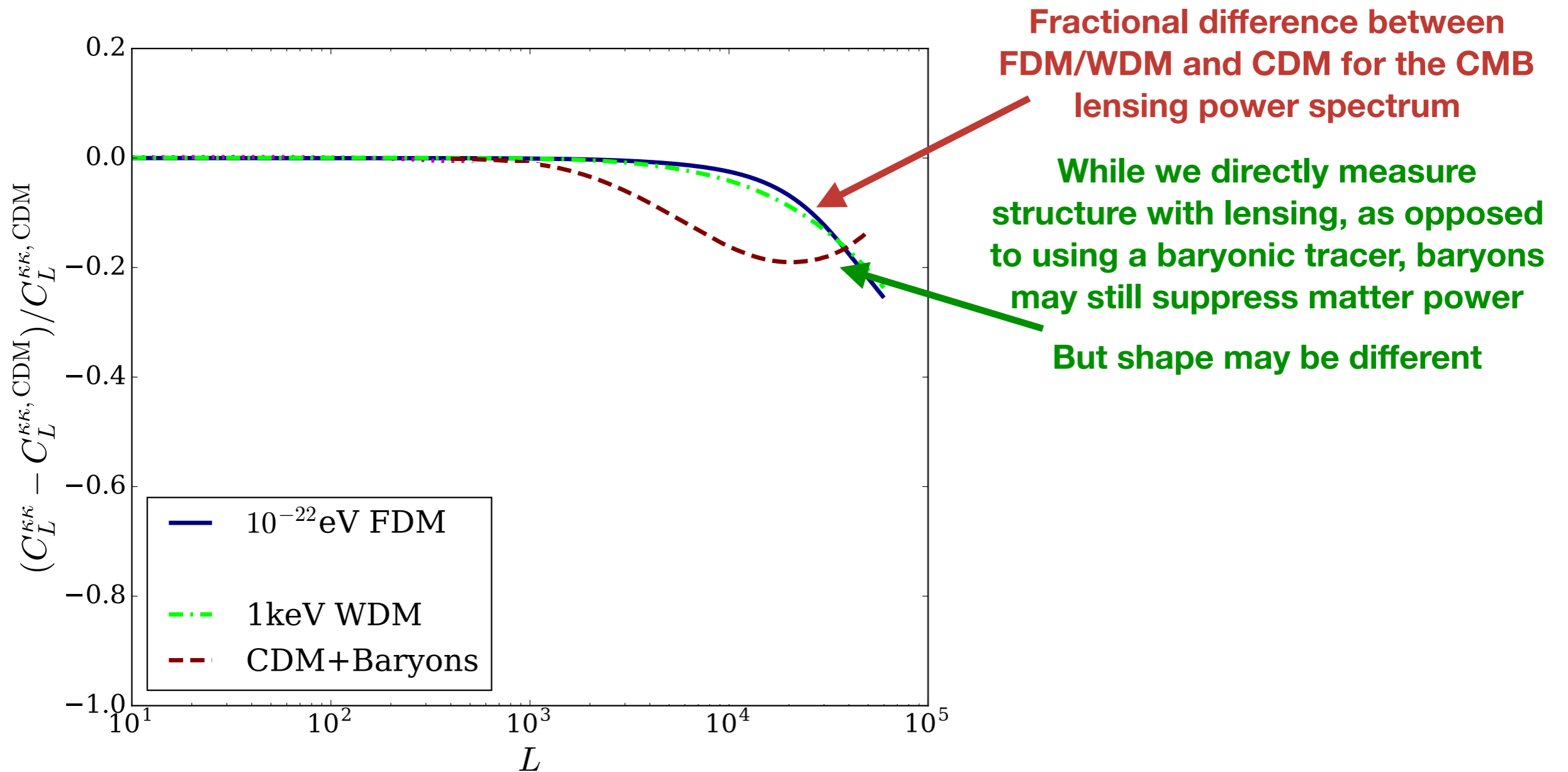


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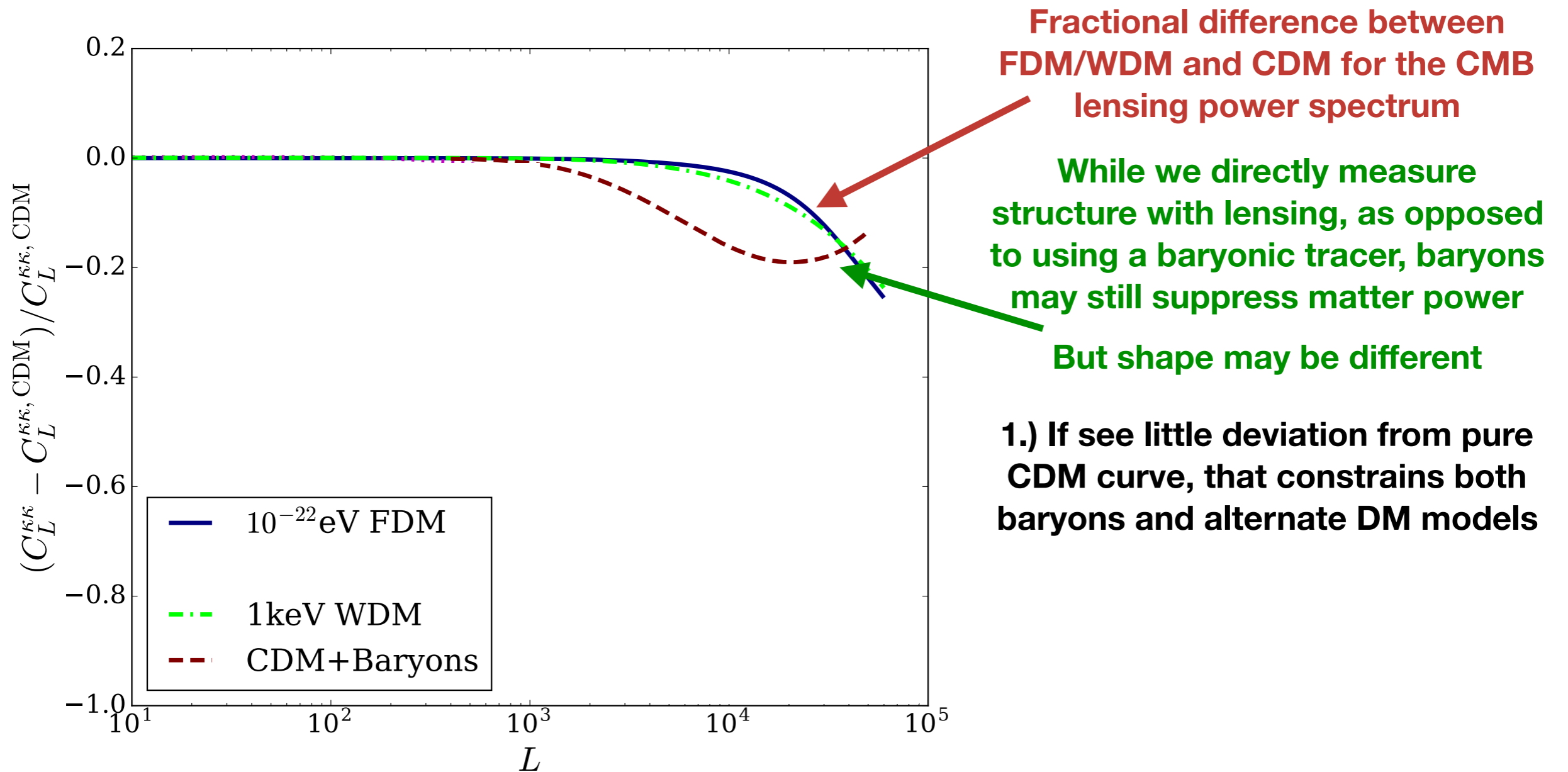




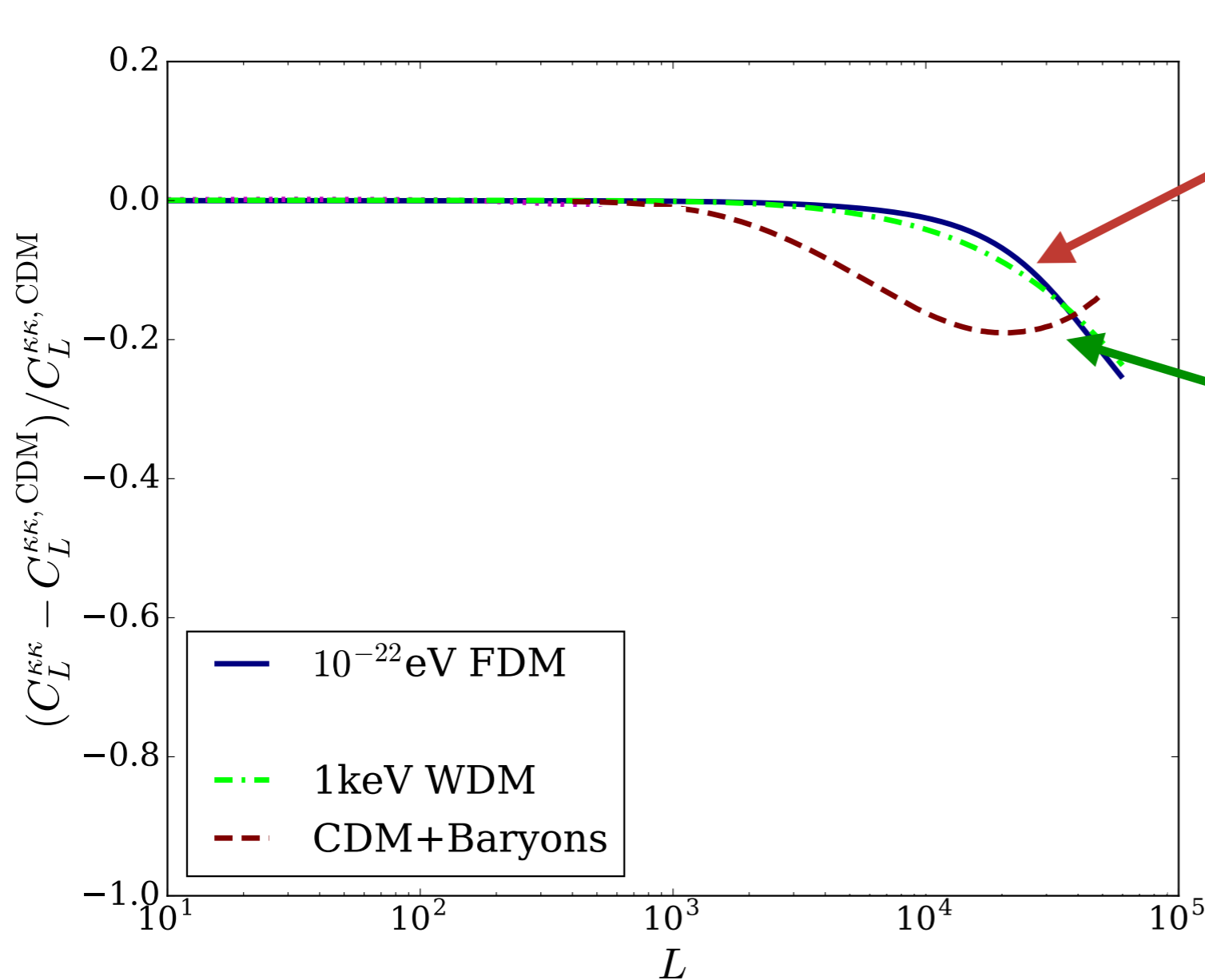
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Fractional difference between FDM/WDM and CDM for the CMB lensing power spectrum

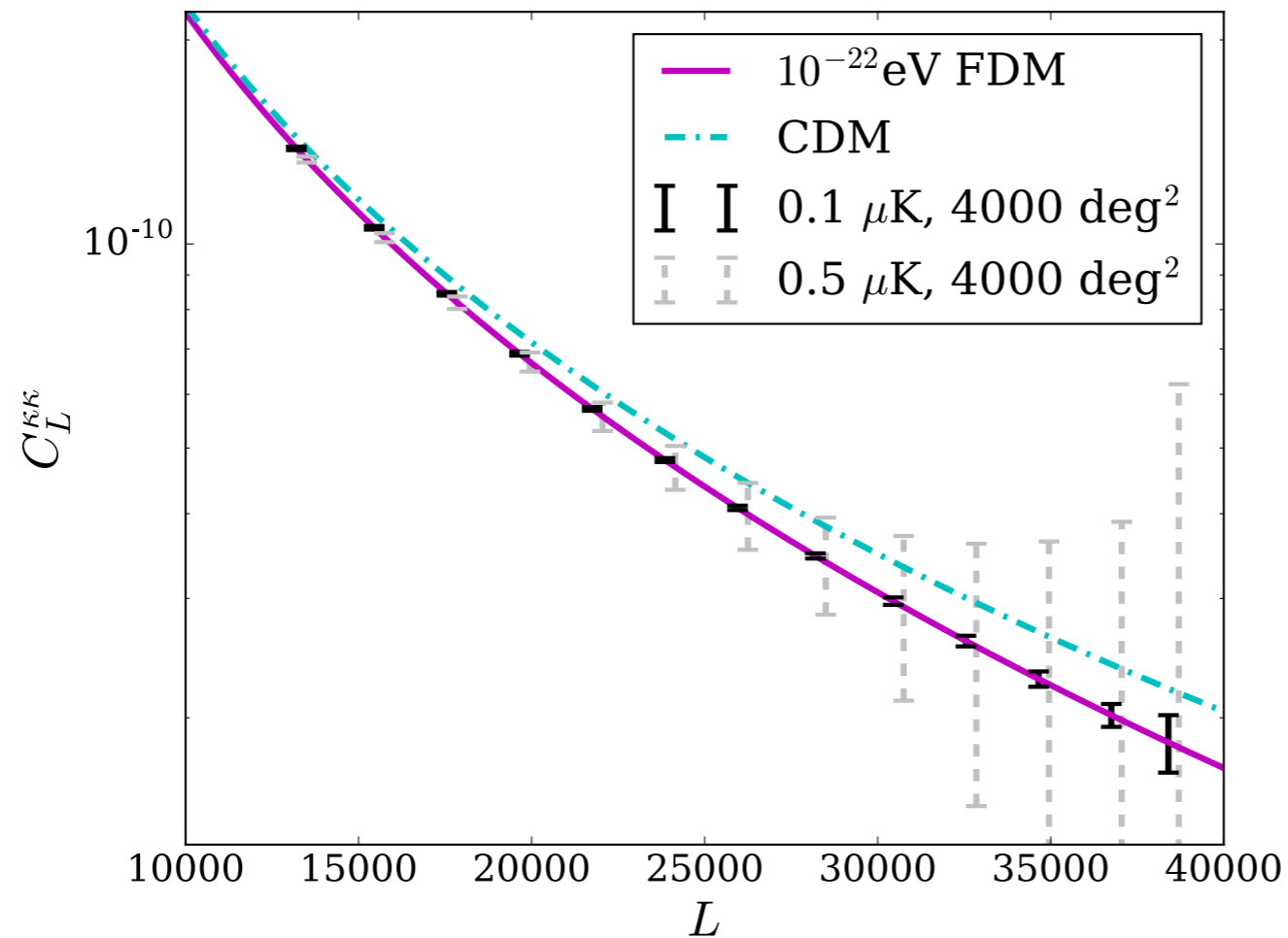
While we directly measure structure with lensing, as opposed to using a baryonic tracer, baryons may still suppress matter power

But shape may be different

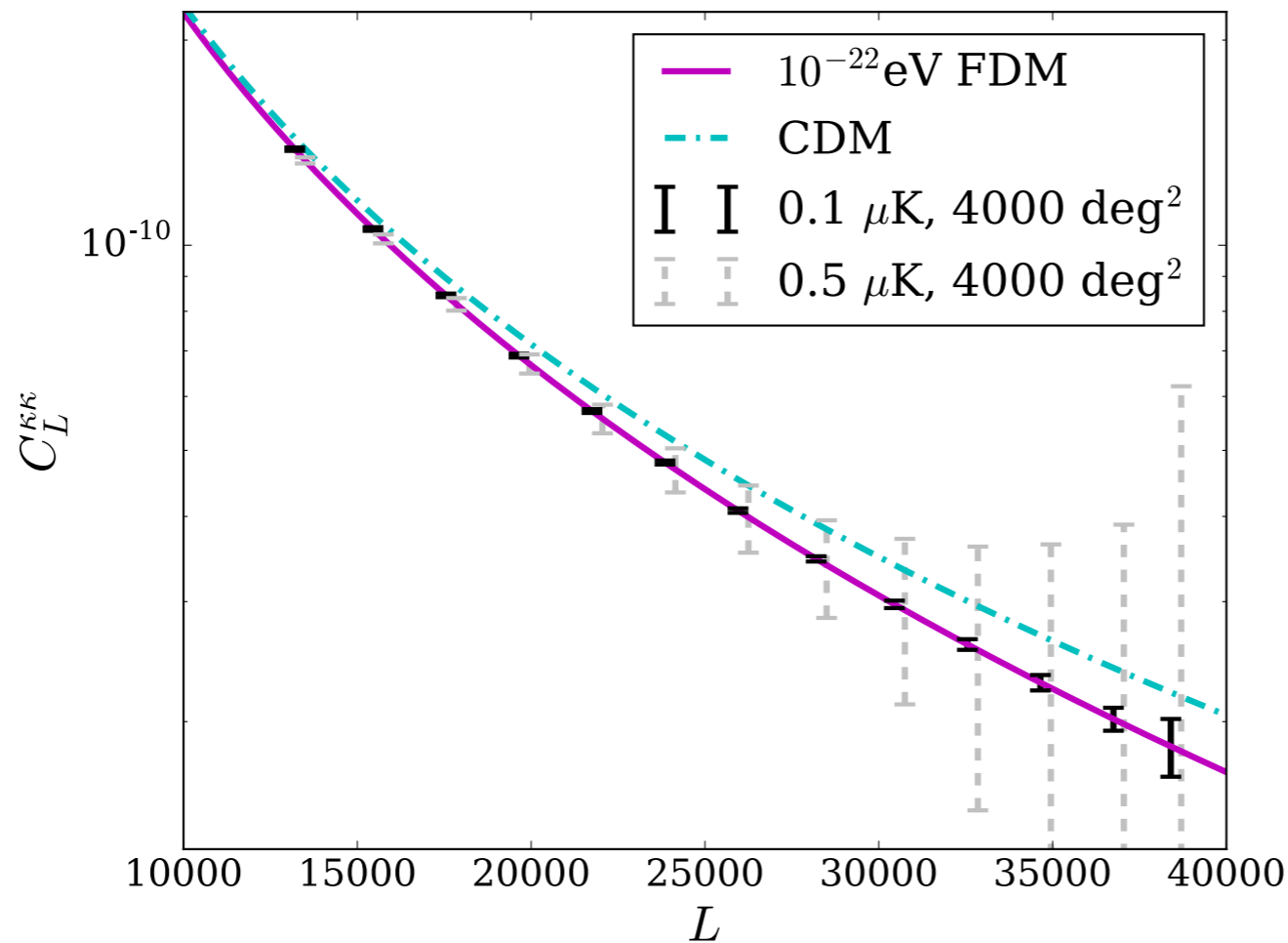
1.) If see little deviation from pure CDM curve, that constrains both baryons and alternate DM models

2.) If see significant deviation, then can potentially use shape of curve to determine whether it is due to baryons or alternative to CDM

# Potential Ability to Distinguish Between Dark Matter Models

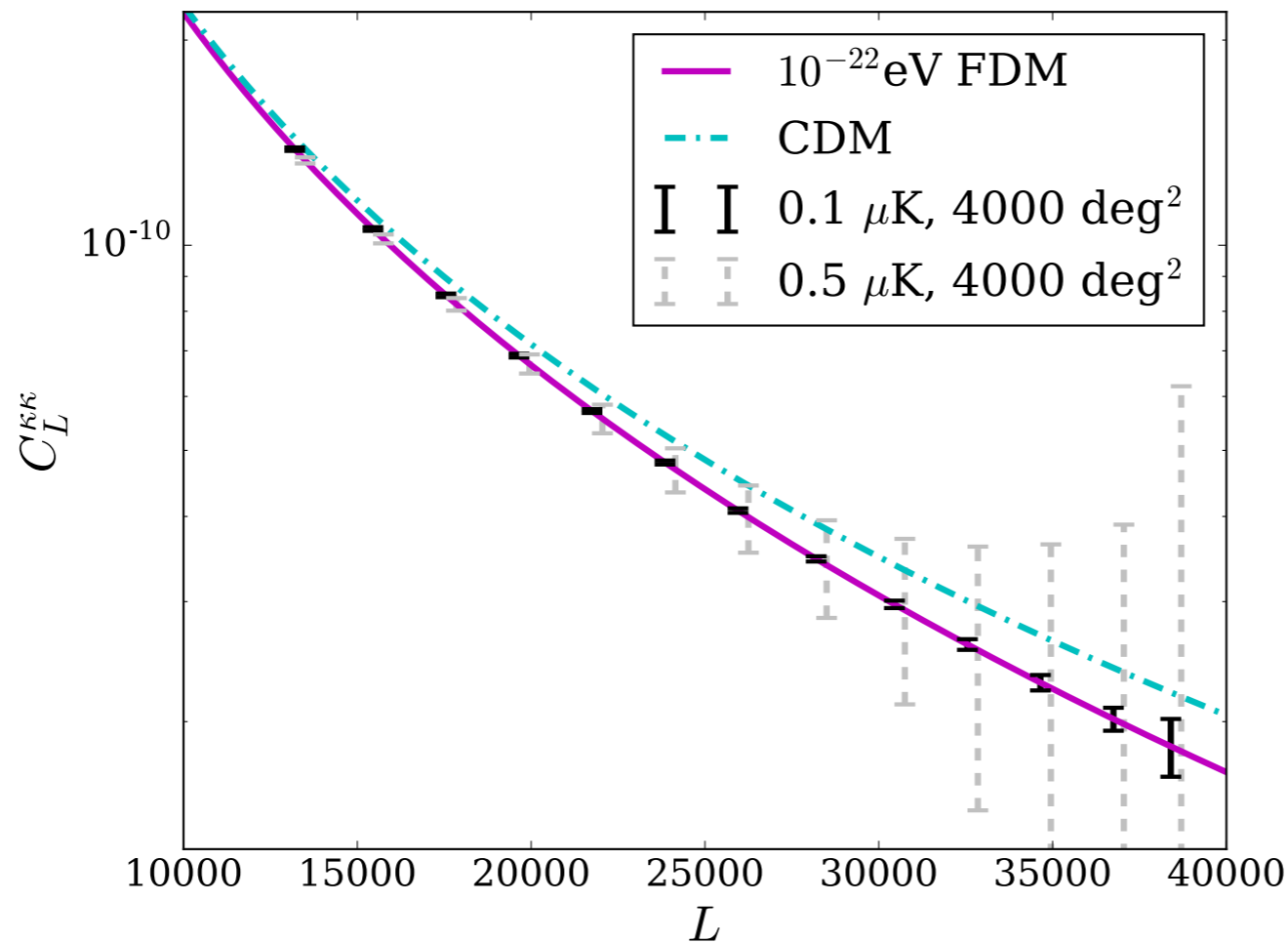


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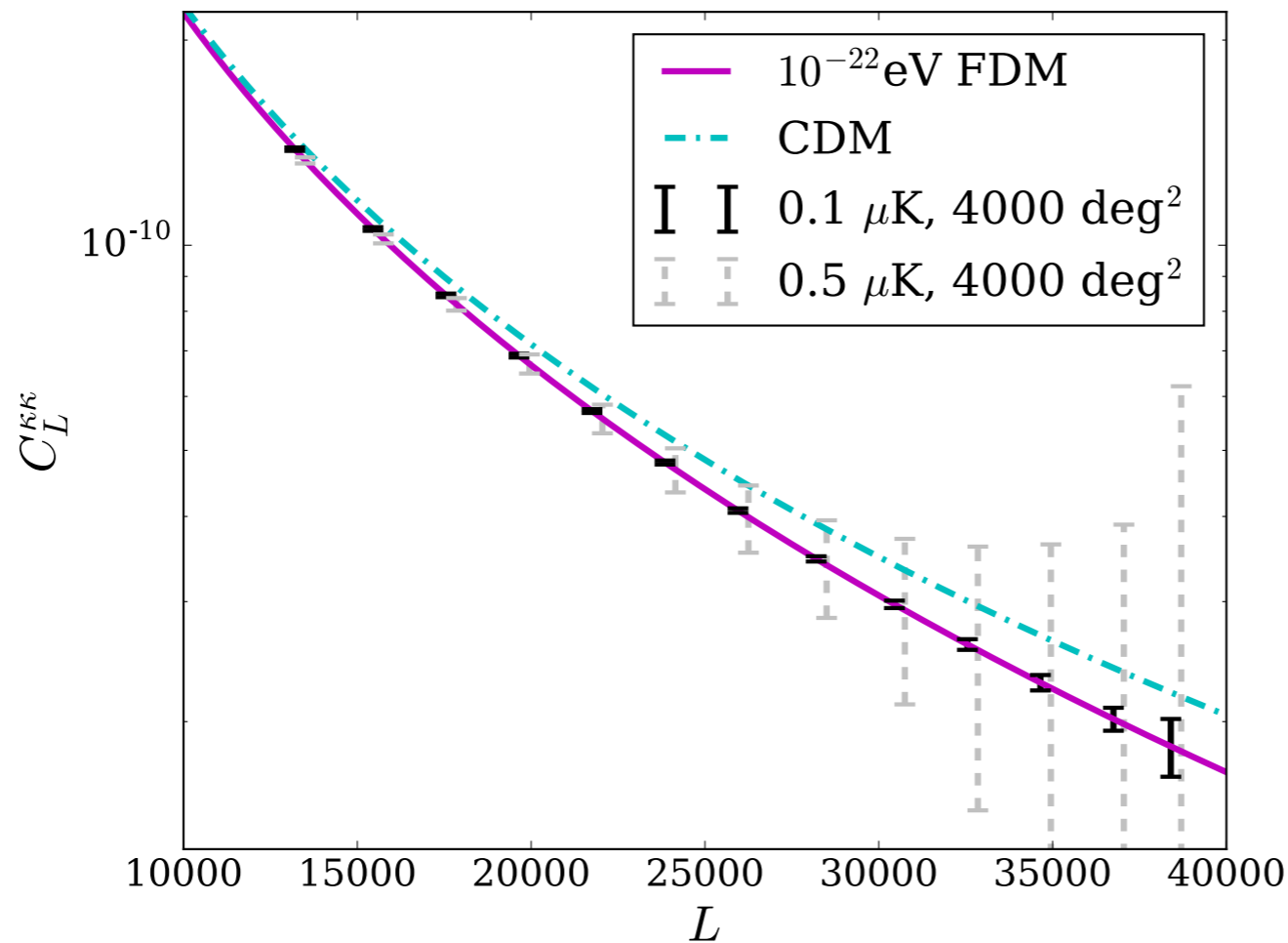
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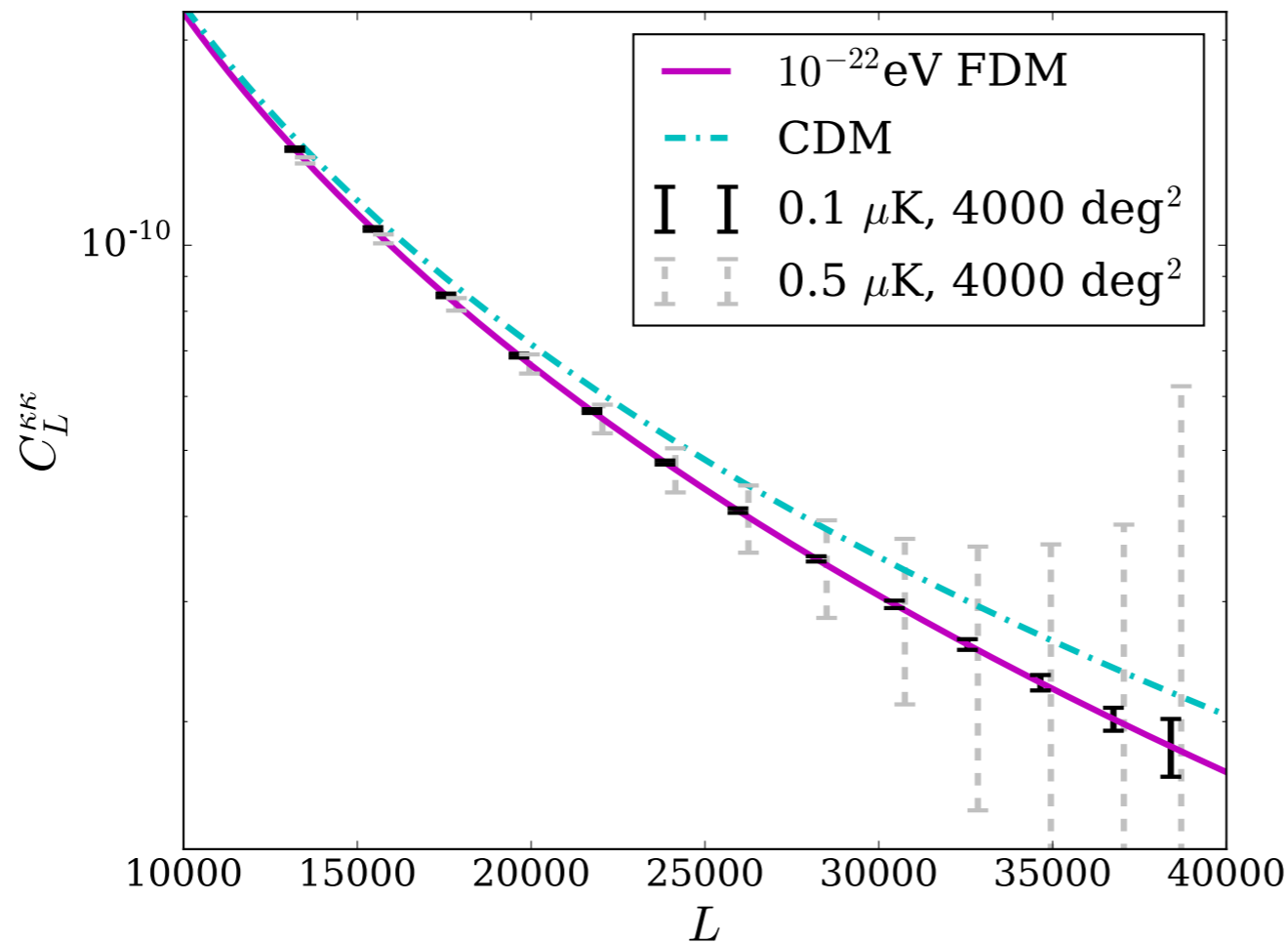
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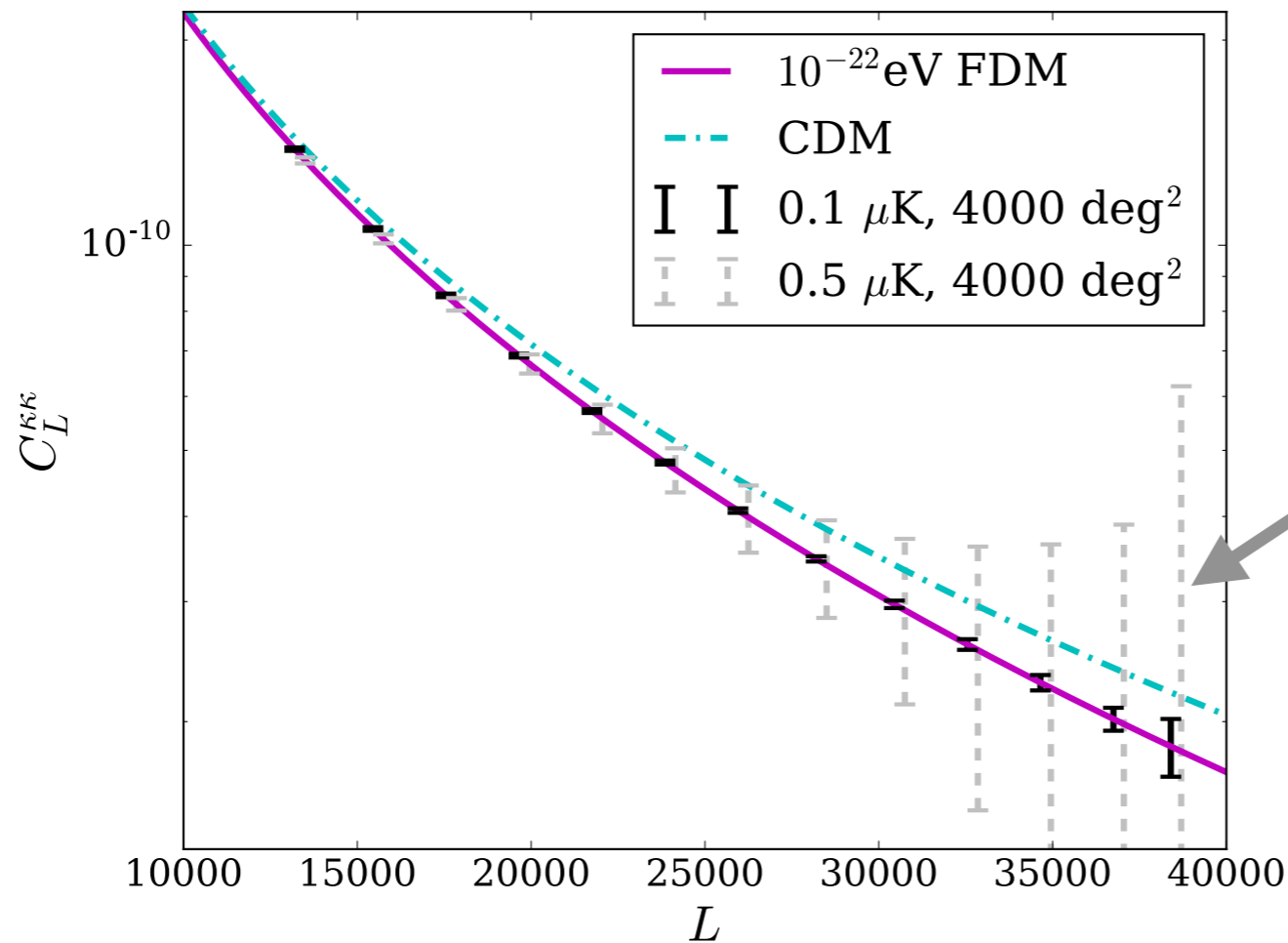
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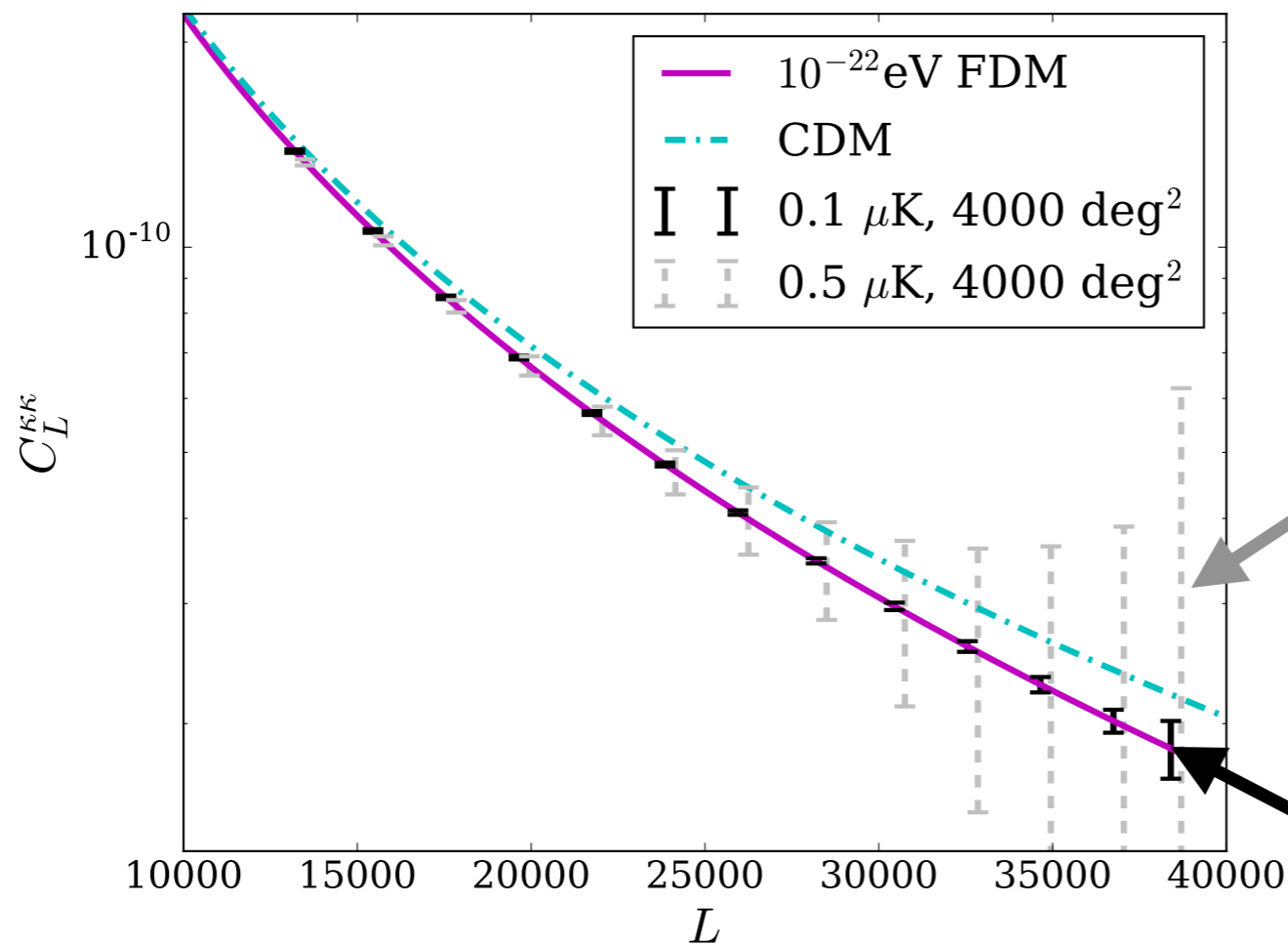


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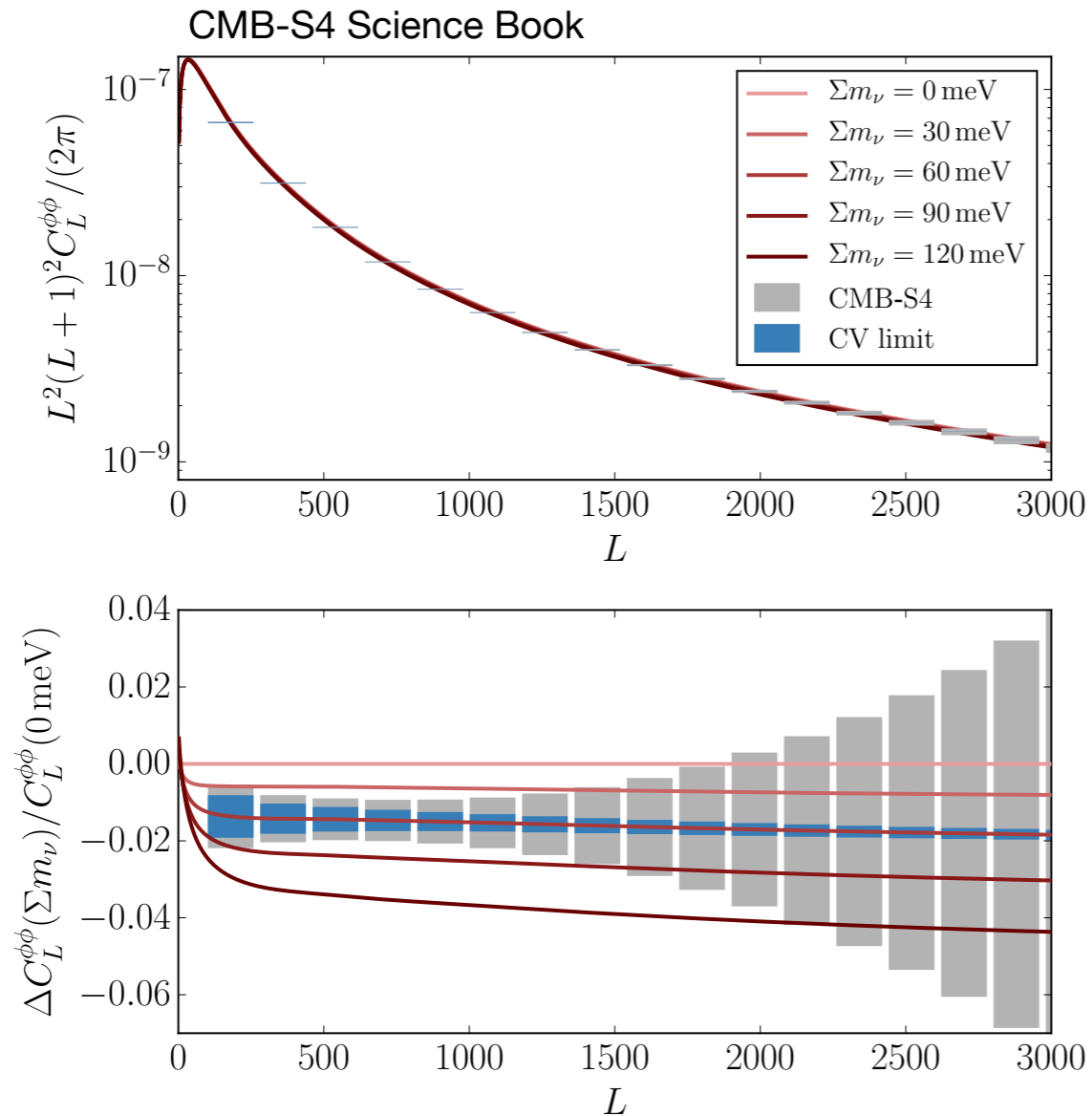
**Black:** S/N ~ 30 for distinguishing between CDM and FDM/WDM

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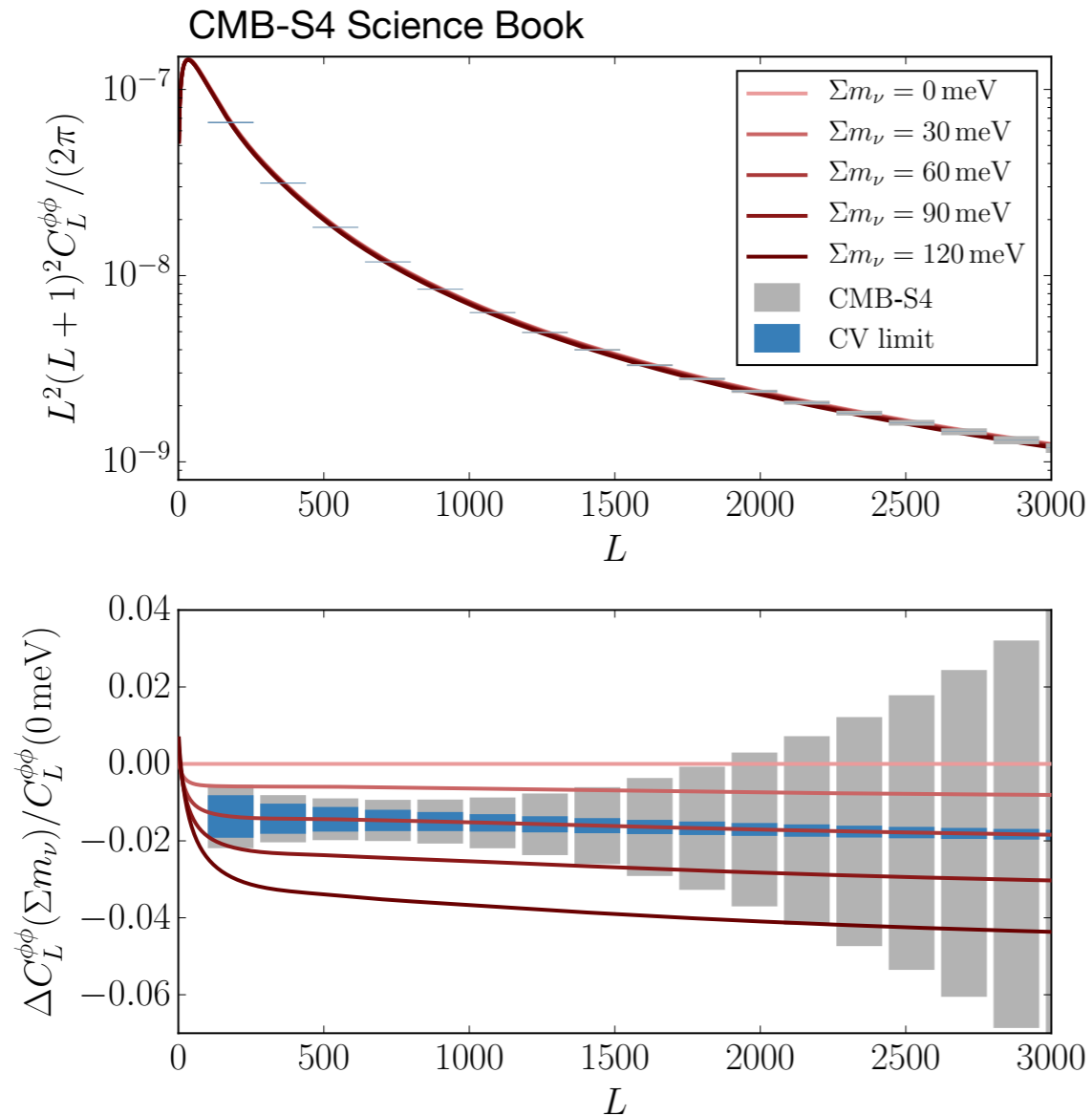
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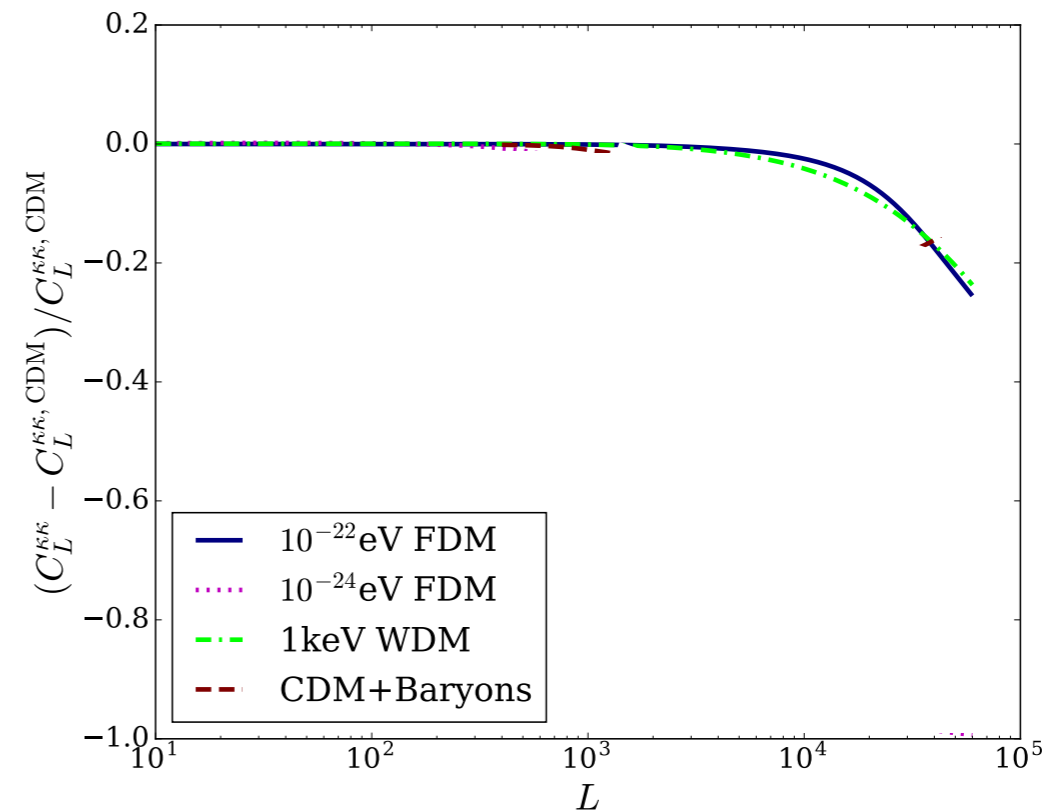
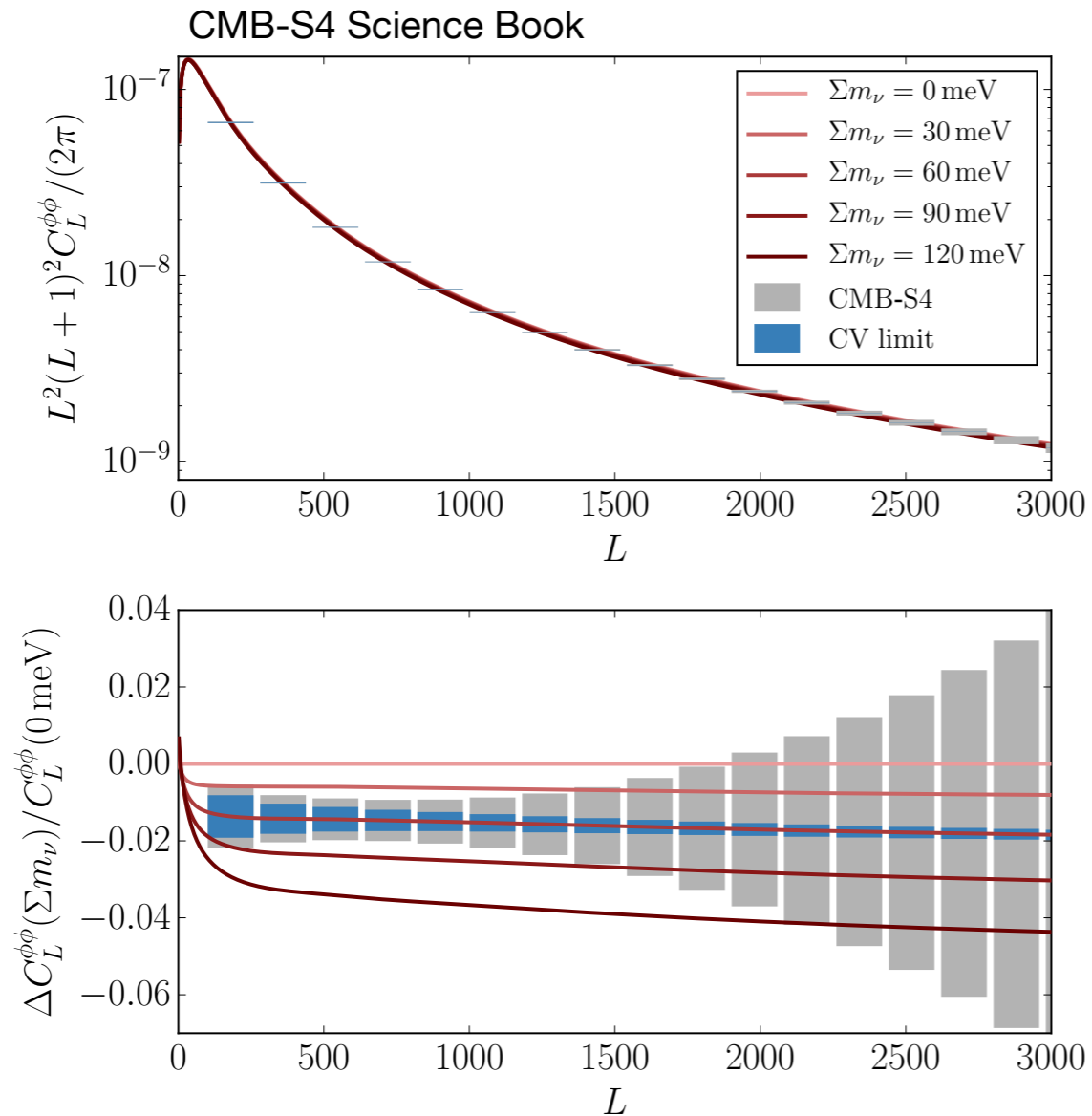


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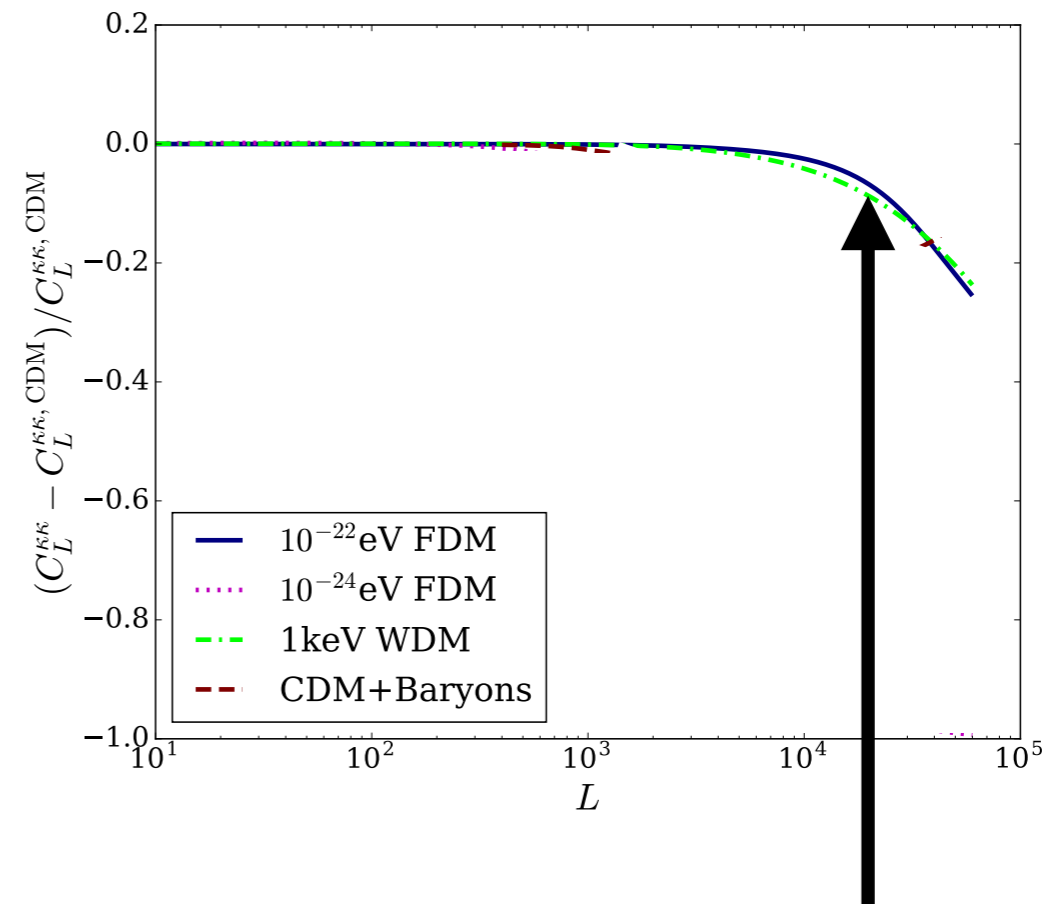
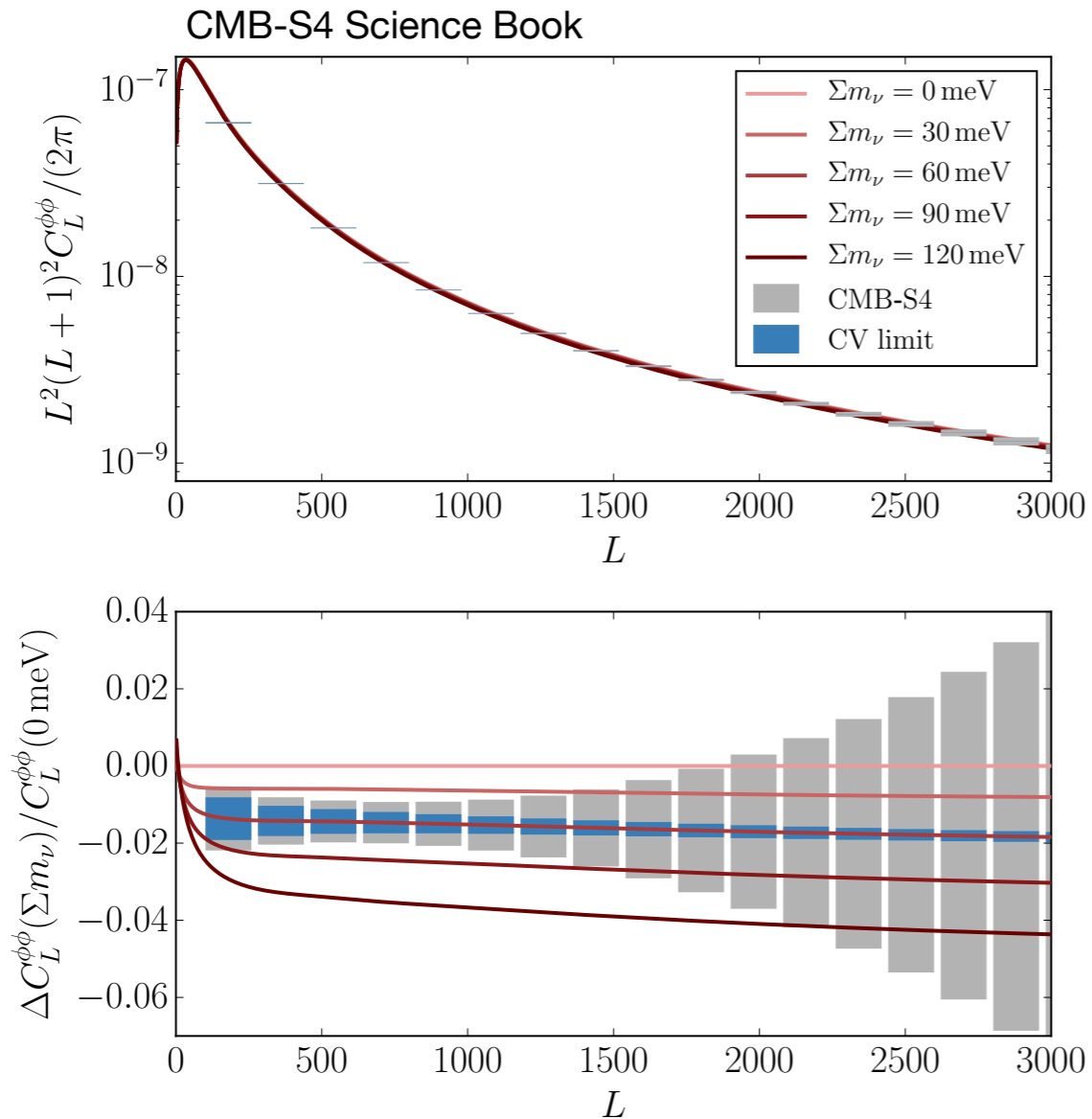
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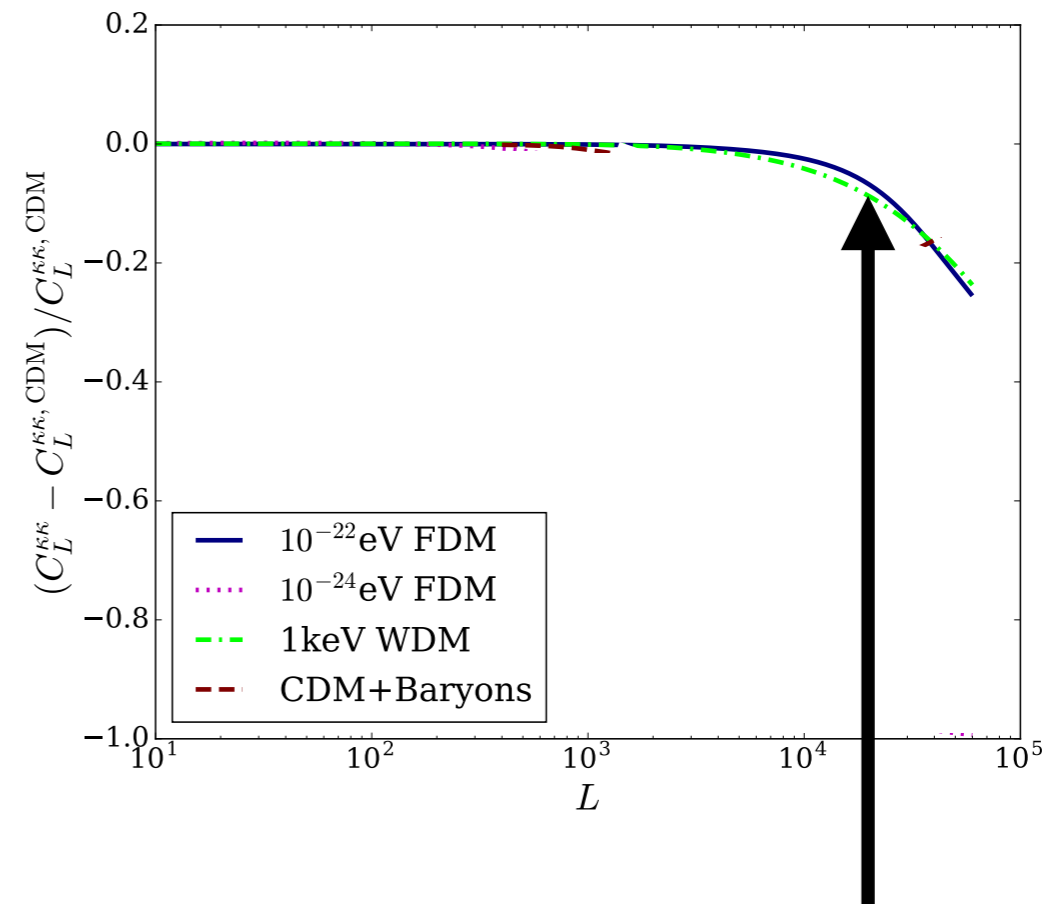
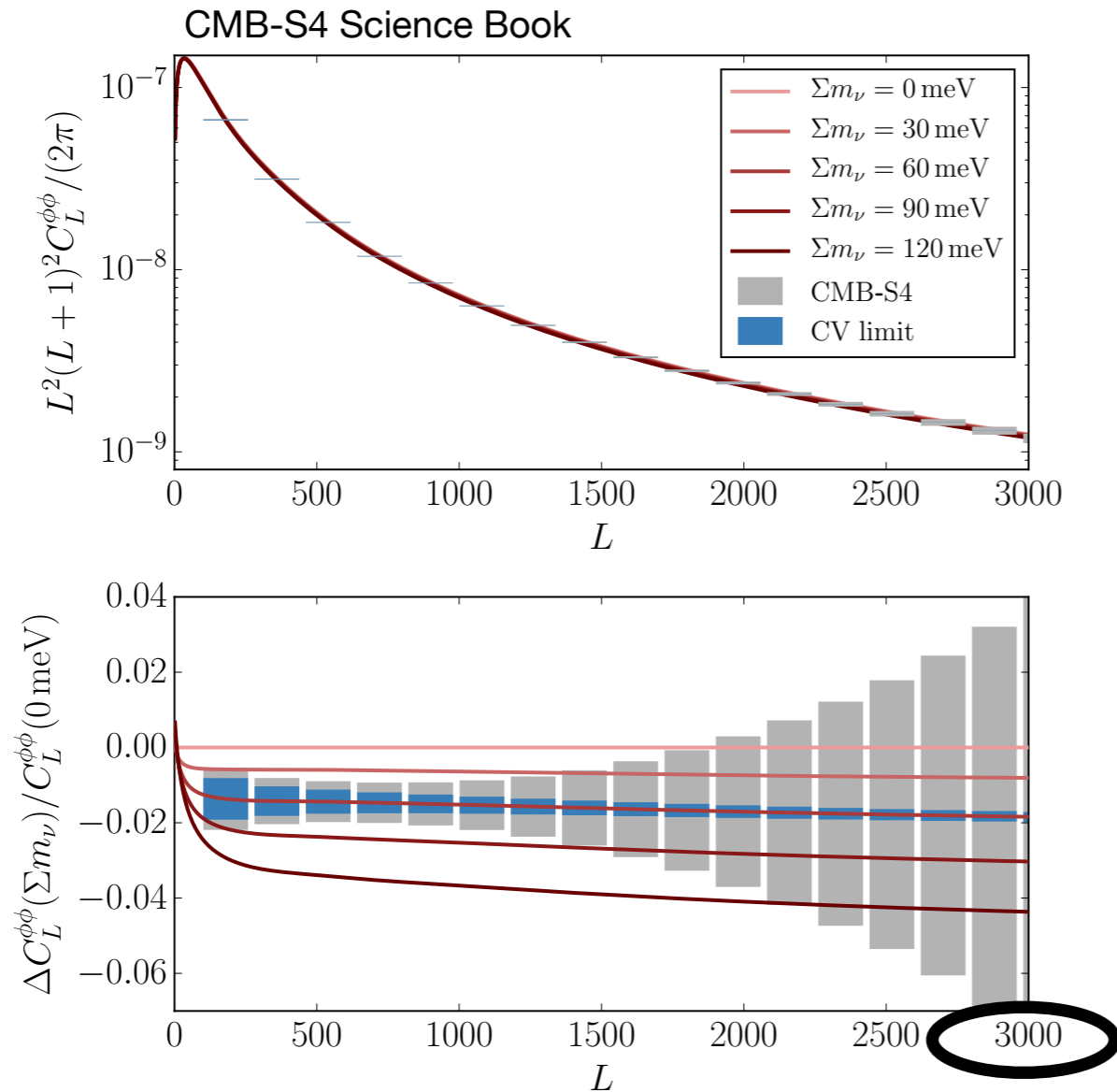
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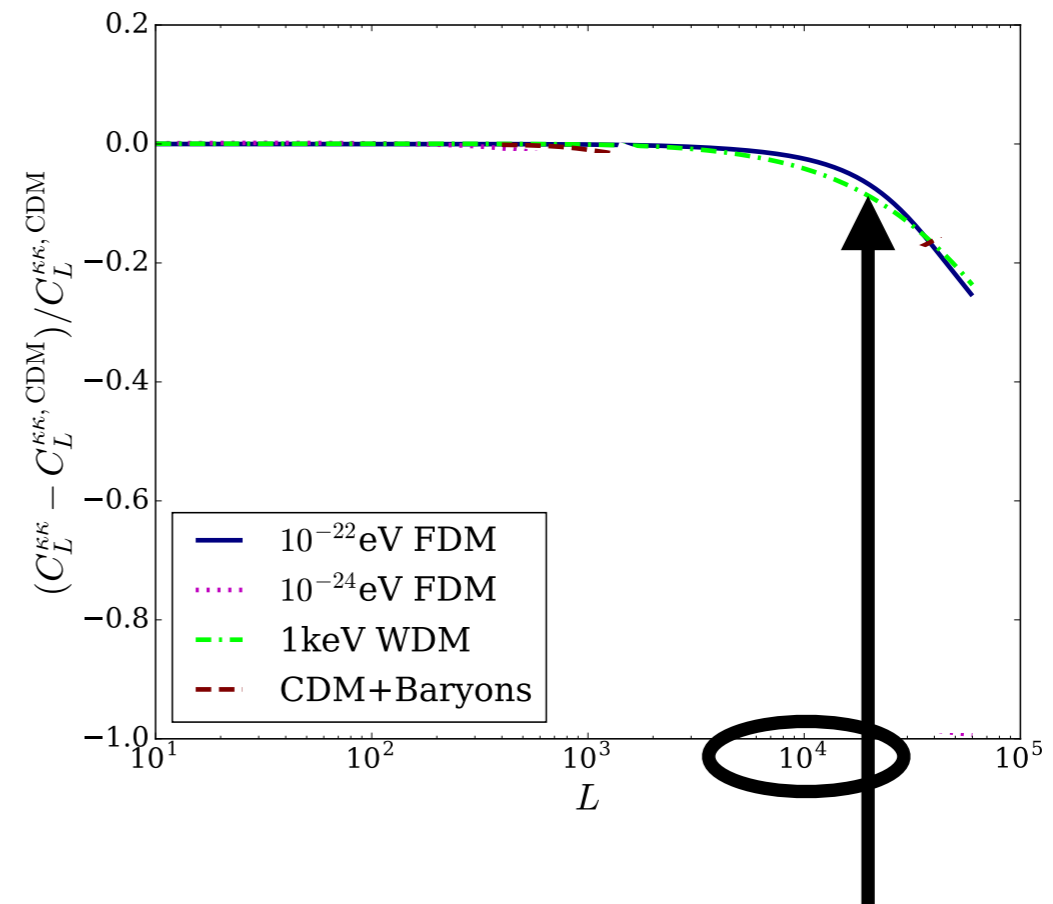
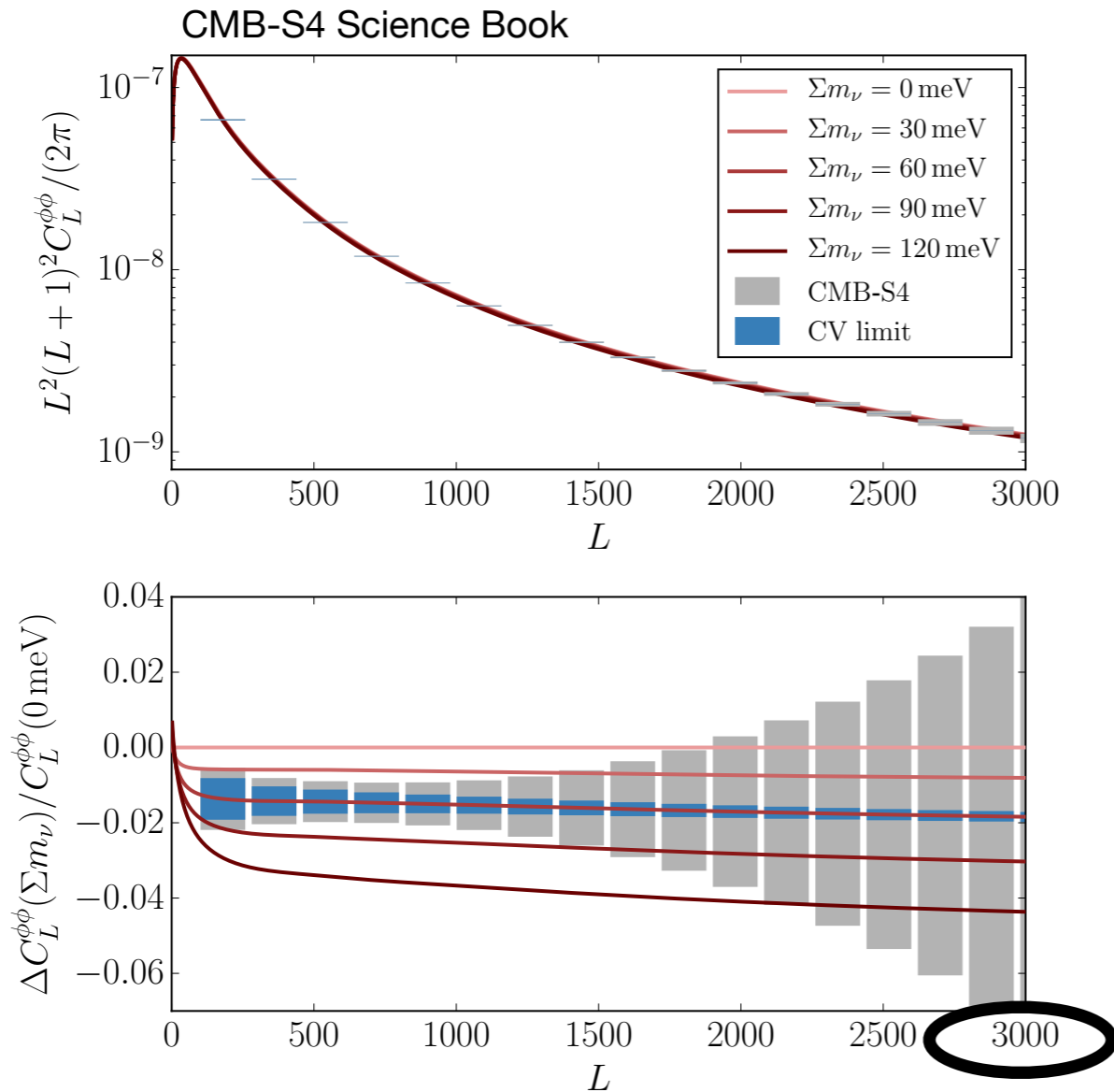


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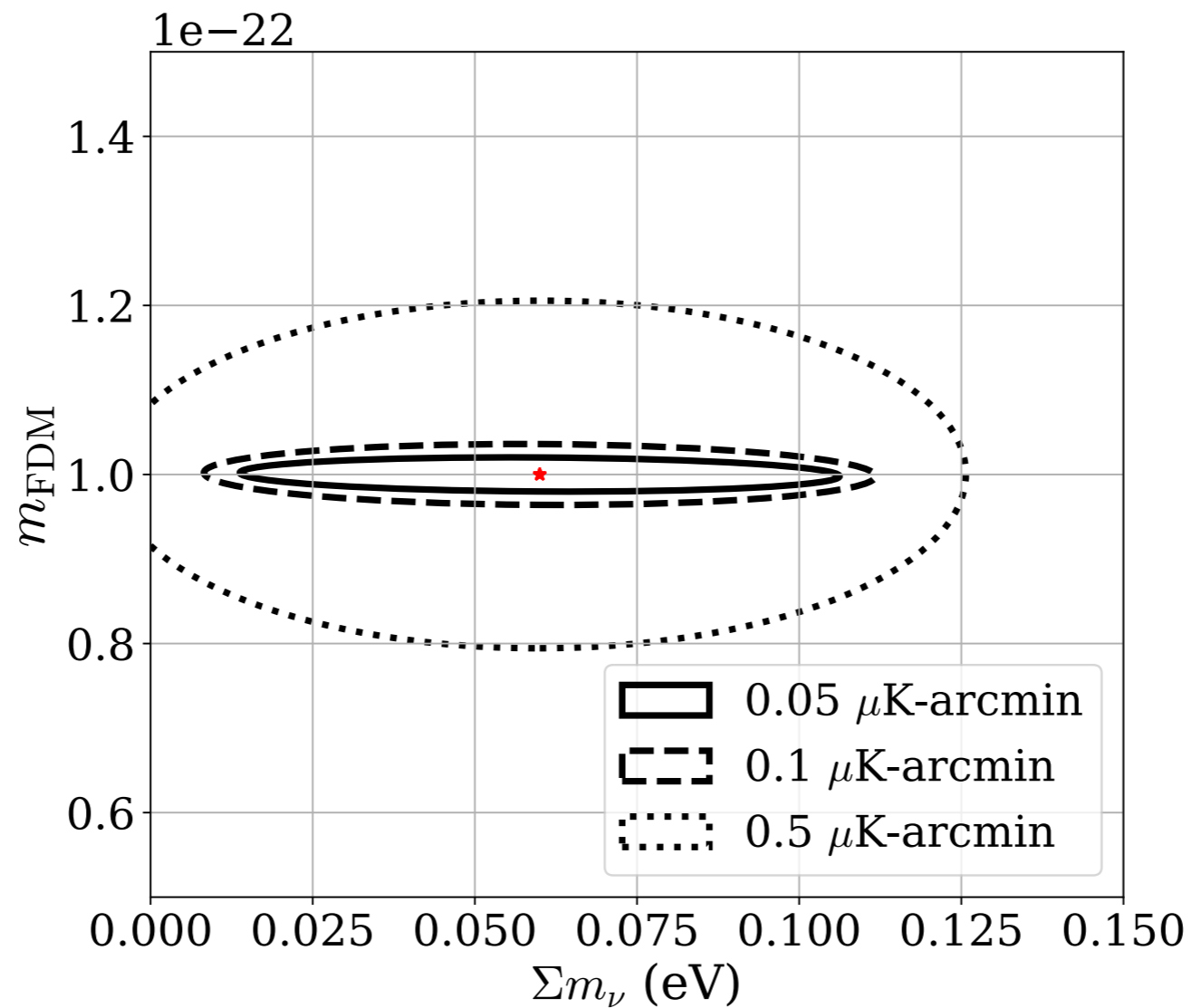
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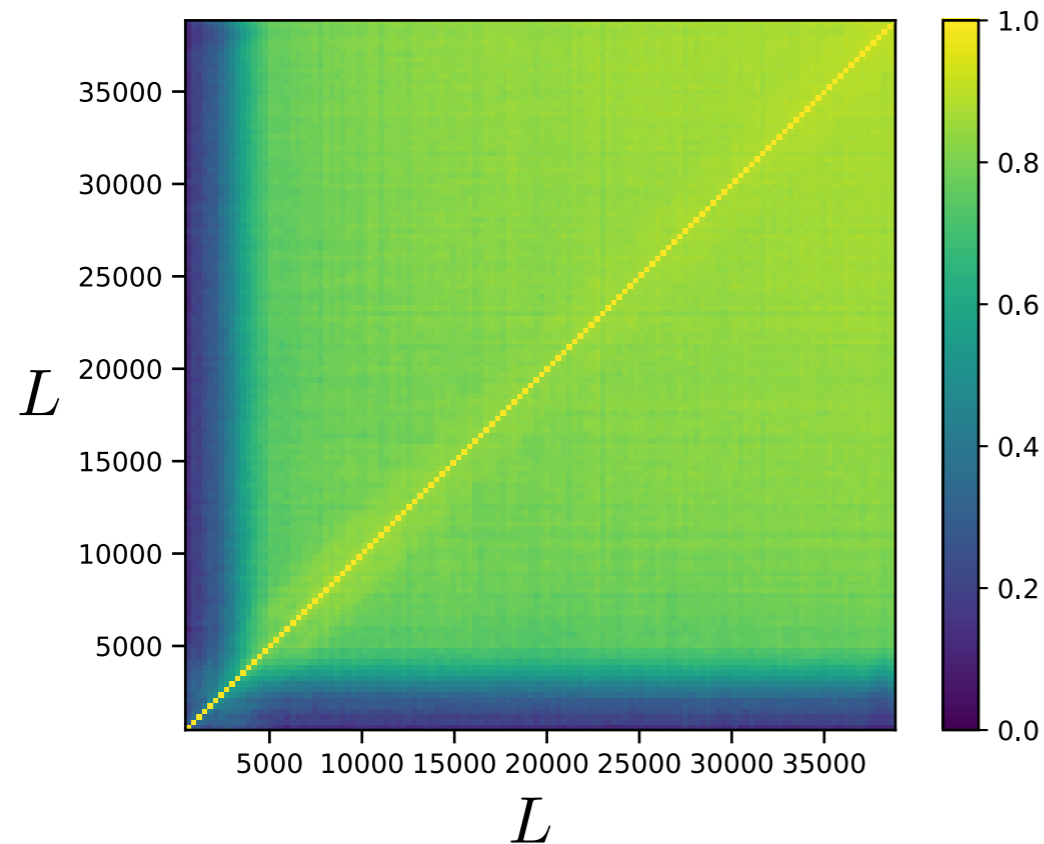
- **Easier to remove correlated modes on small scales?**

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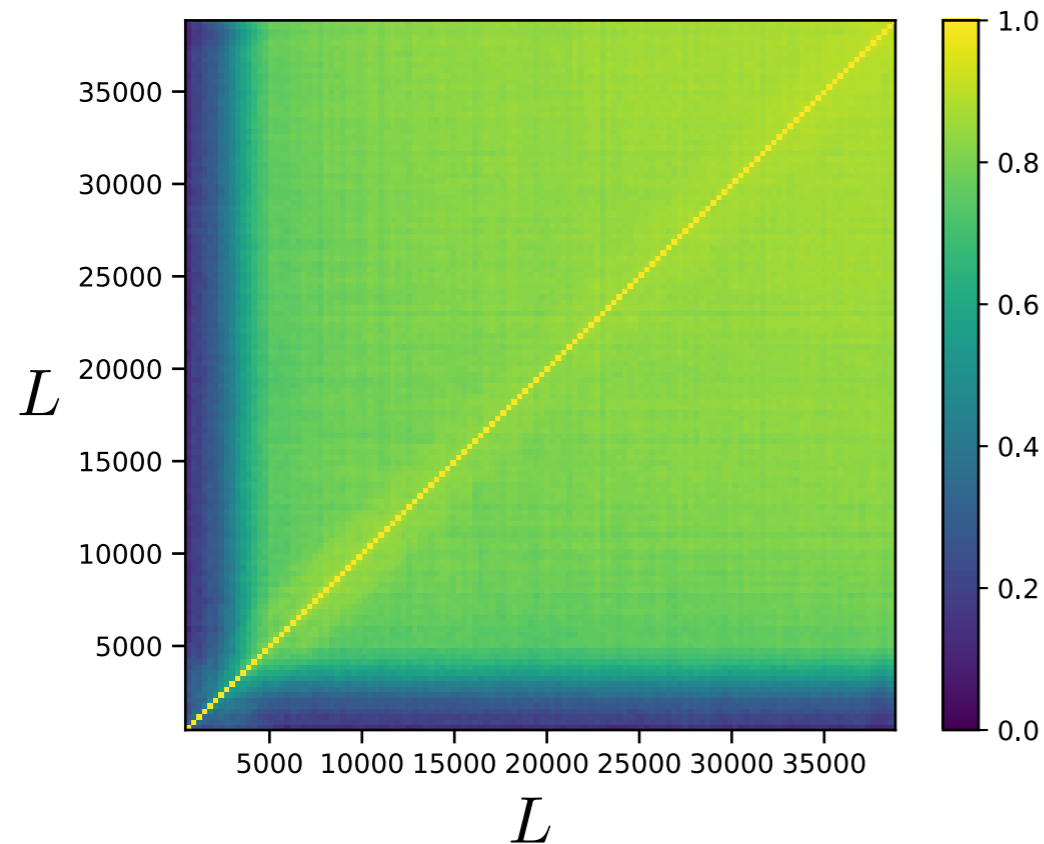
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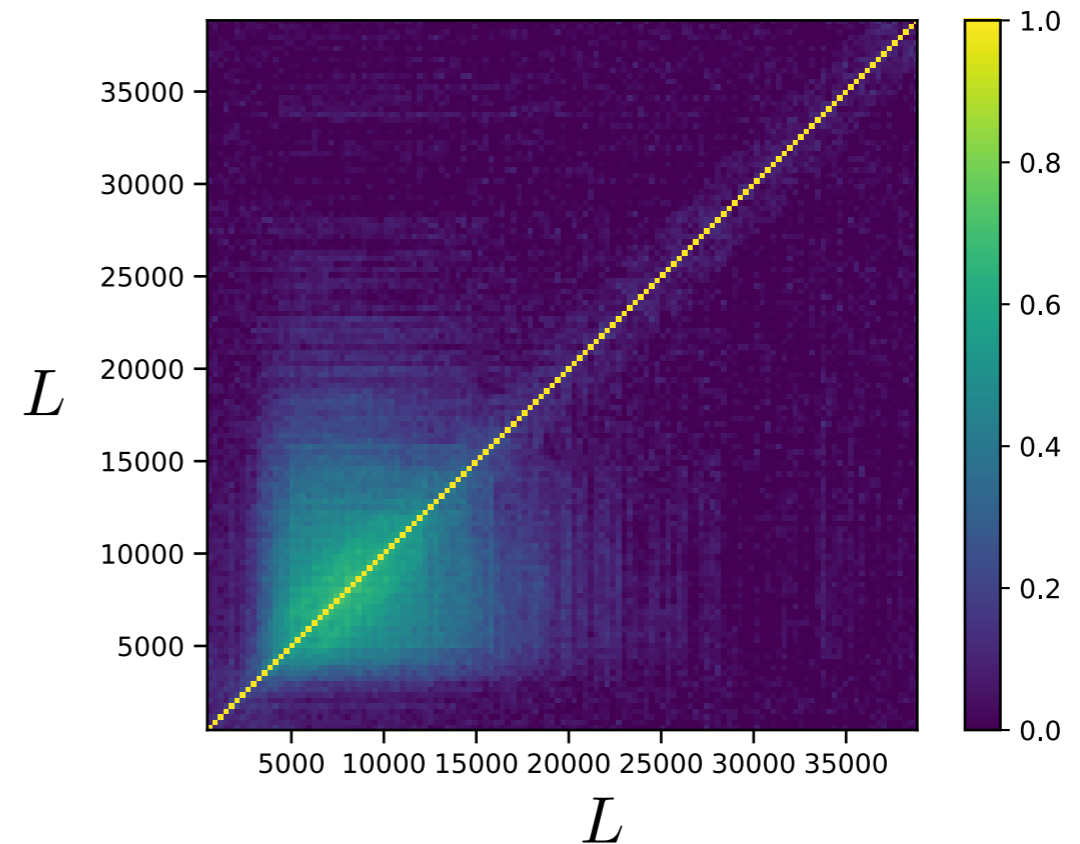
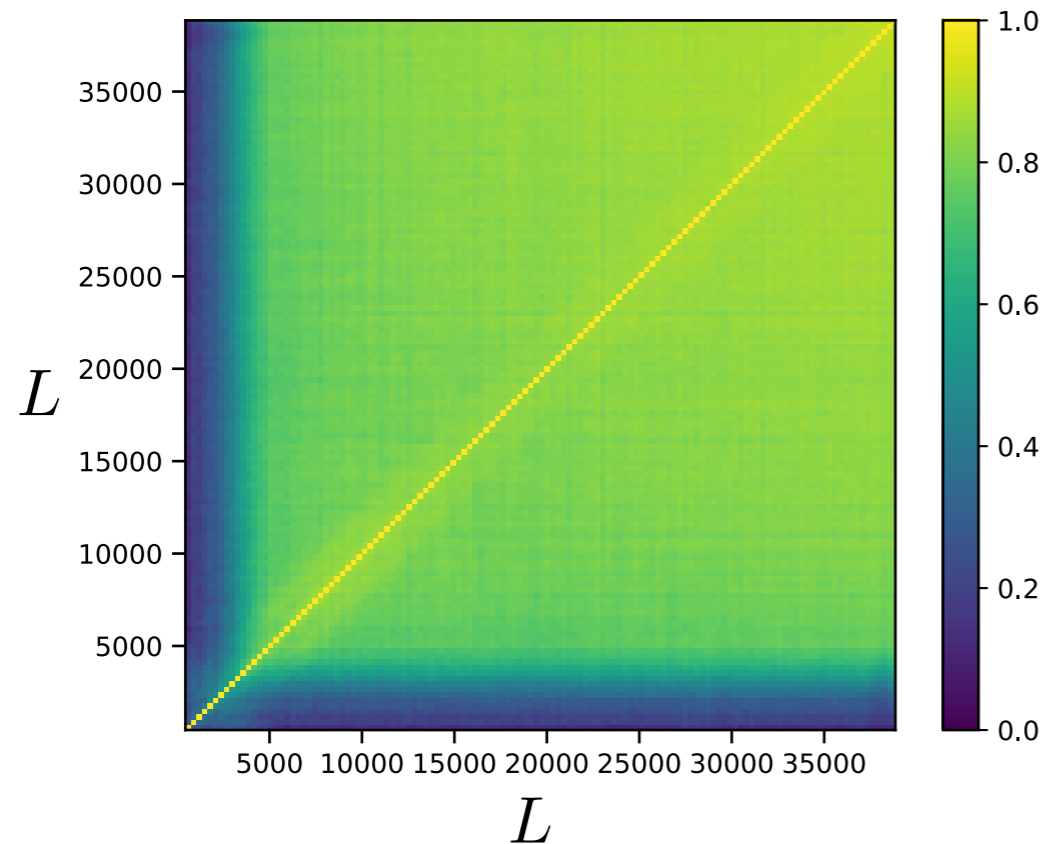
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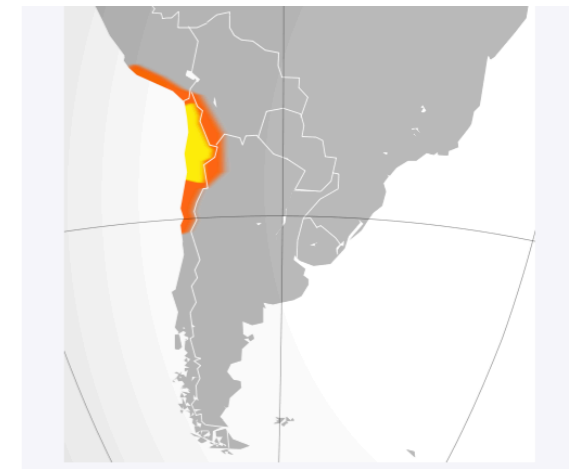
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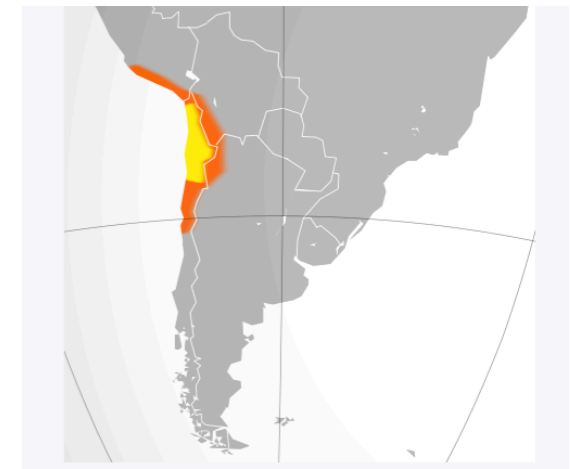


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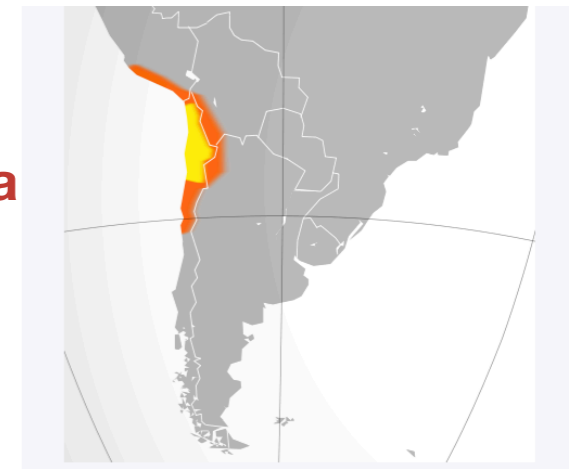


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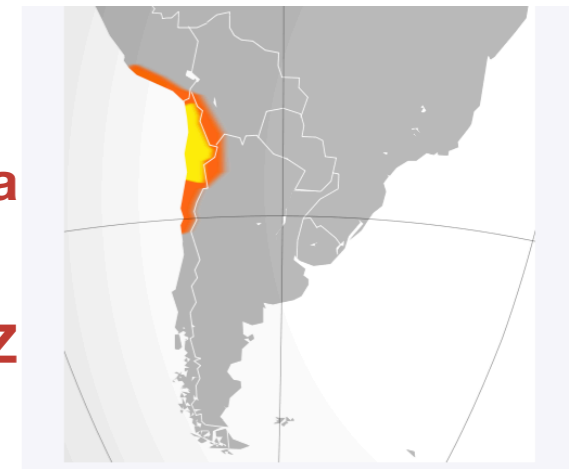


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**New regime of CMB: tons of astrophysics, tSZ and kSZ science, and excellent synergy with optical surveys**



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- Organizing workshop mid-Dec at CCA to explore science case and instrumental feasibility for ultra high-res, low-noise CMB lensing survey - <https://www.simonsfoundation.org/event/the-cmb-in-hd-the-low-noise-high-resolution-frontier/>