

# Millicharged Particles in Neutrino Experiments

arXiv:1806.03310 (submitted to PRL)

Yu-Dai Tsai, **Fermilab**

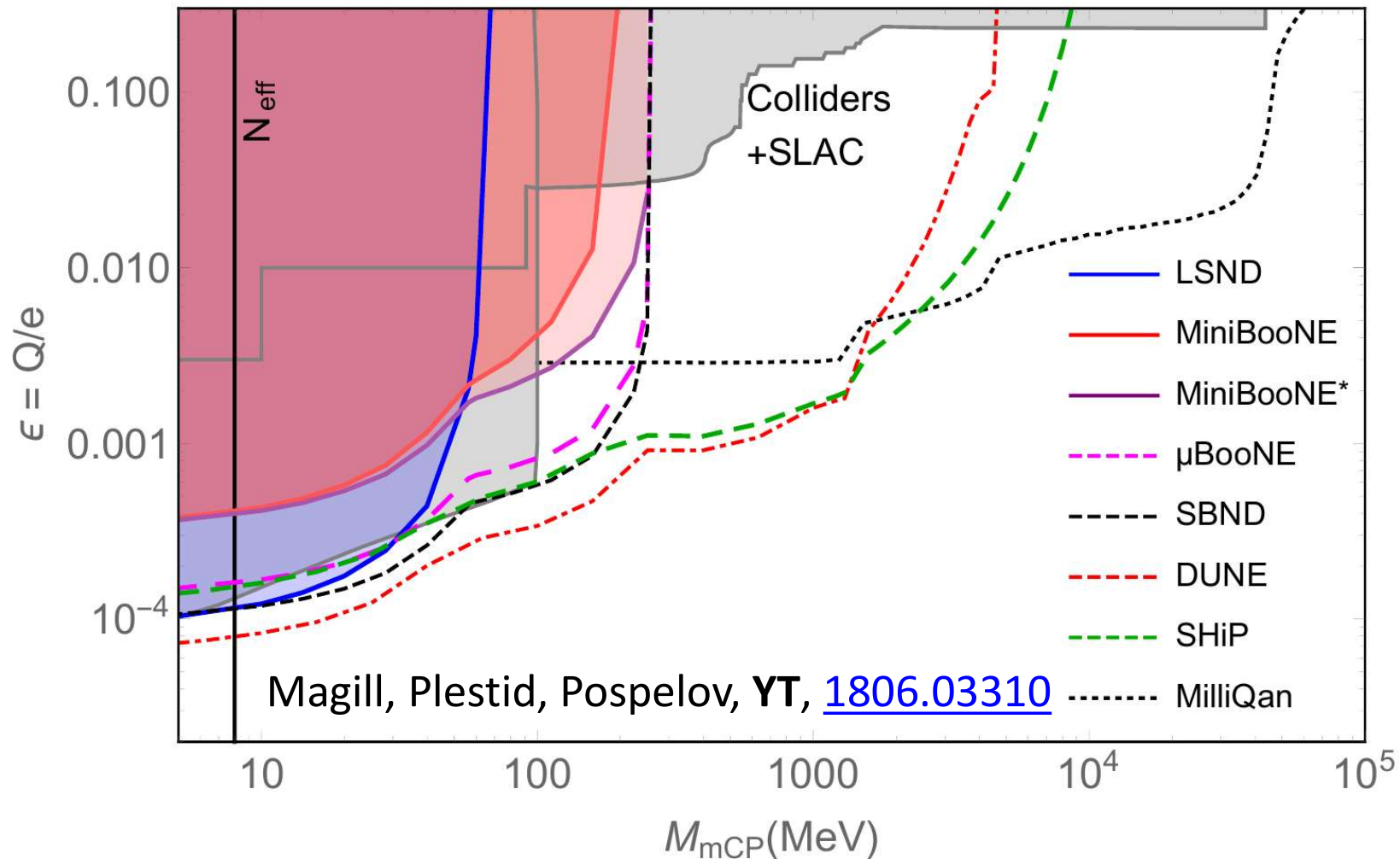
with Gabriel Magill, Ryan Plestid, and Maxim Pospelov

# Outline

- **Motivations**
- **Millicharged Particle (mCP) & Neutrino Experiments**
- **Bounds & Sensitivity Reaches**
- **Discussion**

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# Preview



- Solid: current bounds
- Dashed: future sensitivity
- General review on other bounds: [Andy Haas, Fermilab, 2017](#)

# Millicharged Particles

Is electric charge quantized?

Other Implications

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# Finding Minicharge

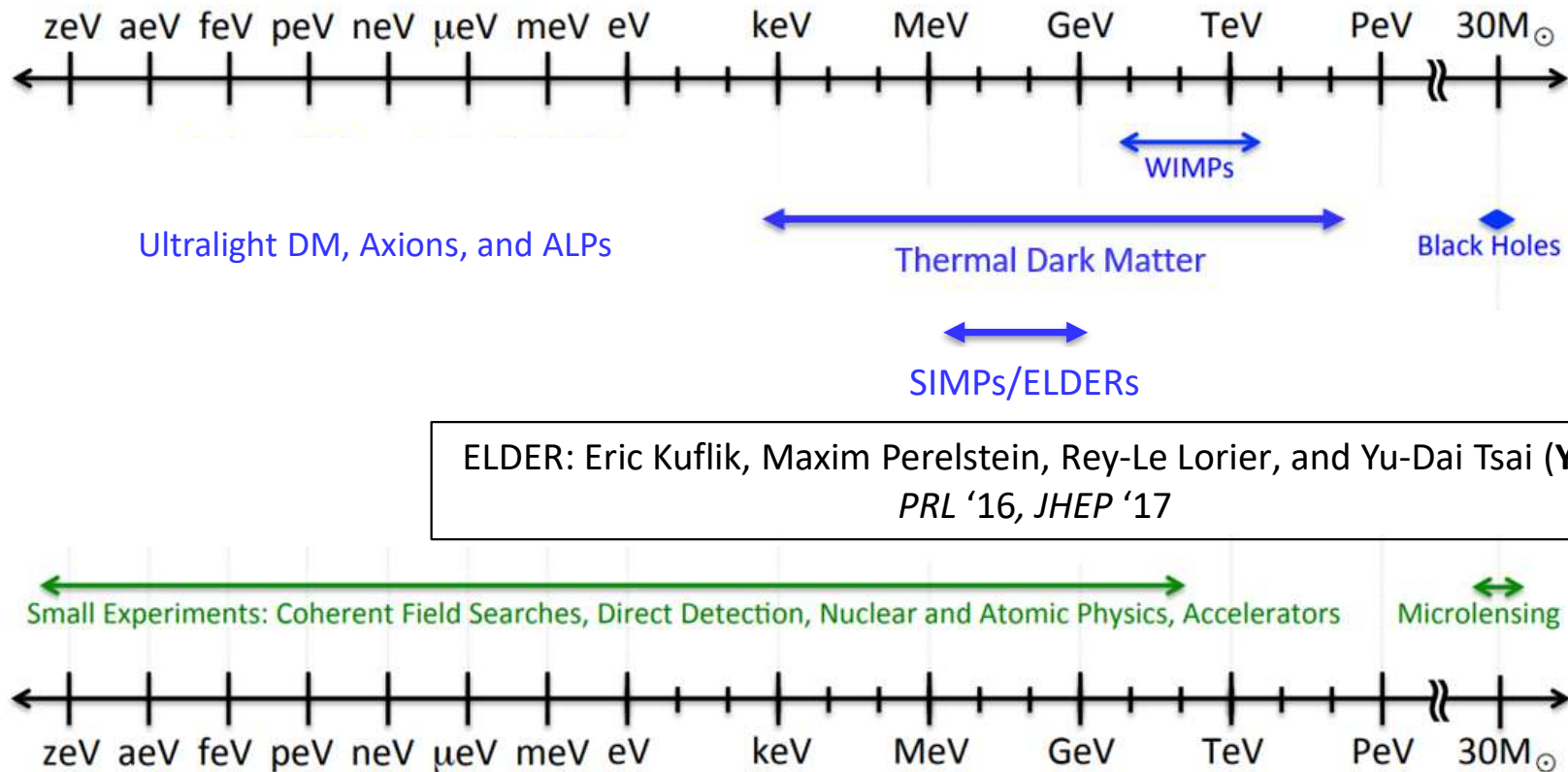
- **Is electric charge quantized?**
- U(1) group allows arbitrarily small charges. Why don't we see them in electric charges? This motivated **Dirac monopole**, **Grand Unified Theory (GUT)**, etc
- Practically, searching for millicharge is a test of  **$e/3$  charge quantization**, and this could imply something more significant
- Could have natural link to **dark sector** (e.g. dark photon)
- Could account for DM (WIMP-like or other scenarios)
- Used to explain the cooling of gas temperature to explain the EDGES result [**EDGES collab., Nature, (2018)**, **Barkana, Nature, (2018)**]. Only  $\sim 1\%$  of the DM allowed to explain the “anomaly” given other constraints

# Neutrino Experiments

- Neutrinos are weakly interacting particles. Just like **Millicharged particles**
- **High statistics**, e.g. LSND has  $10^{23}$  **Protons on Target** (POT)
- Shielded/underground: low background (e.g. solar  $\nu$  programs)
- There are many of them existing and many to **come: strength in numbers**
- Produce hidden particles without DM assumptions: more “direct” than cosmology/astrophysics probes, DM direct detections, etc.

# Dark Matter/Hidden Particles Exploration

## Dark Sector Candidates, Anomalies, and Search Techniques



## US Cosmic Visions 2017

- Proton fix-target/neutrino experiments are important for MeV ~ GeV!
- Golowich and Robinett, PRD 87
- Babu, Gould, and Rothstein, PLB 94
- Gninenko, Krasnikov, and Rubbia, PRD 07, ...

# $\nu$ Hopes for New Physics: Personal Trilogy

⋮

- **Light Scalar/Dark Photon** at Borexino & LSND

(Pospelov & YT, [1706.00424](#))

- **Dipole Portal Heavy Neutral Lepton**

(Magill, Plestid, Pospelov & YT, [1803.03262](#))

- **Millicharged Particles** in Neutrino Experiments

(Magill, Plestid, Pospelov & YT, [1806.03310](#))

Inspired by ...

deNiverville, Pospelov, Ritz, '11,

Kahn, Krnjaic, Thaler, Toups, '14, ...

⋮



# $\nu$ Hopes for New Physics & **Anomalies**

⋮

- **Light Scalar**/Dark Photon at Borexino & LSND

Proton charge radius anomaly

- **Dipole Portal Heavy Neutral Lepton**

LSND/MiniBooNE anomaly

- **Millicharged Particles** in Neutrino Experiments

EDGES 21-cm measurement anomaly

We can chat about  
these over coffee

⋮

# Millicharged Particle: Models

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# mCP Model

- Small charged particles under U(1) hypercharge

$$\mathcal{L}_{\text{mCP}} = i\bar{\psi}(\not{\partial} - i\epsilon' e\not{B} + M_{\text{mCP}})\psi$$

- Can just consider this effective lagrangian term by itself (no extra mediator, i.e., dark photon)
- Or this could be from **Kinetic Mixing**
  - give a nice origin to this term
  - an example that gives rise to **dark sector**
  - easily compatible with GUT
  - I will not spend too much time on the model

# Kinetic Mixing



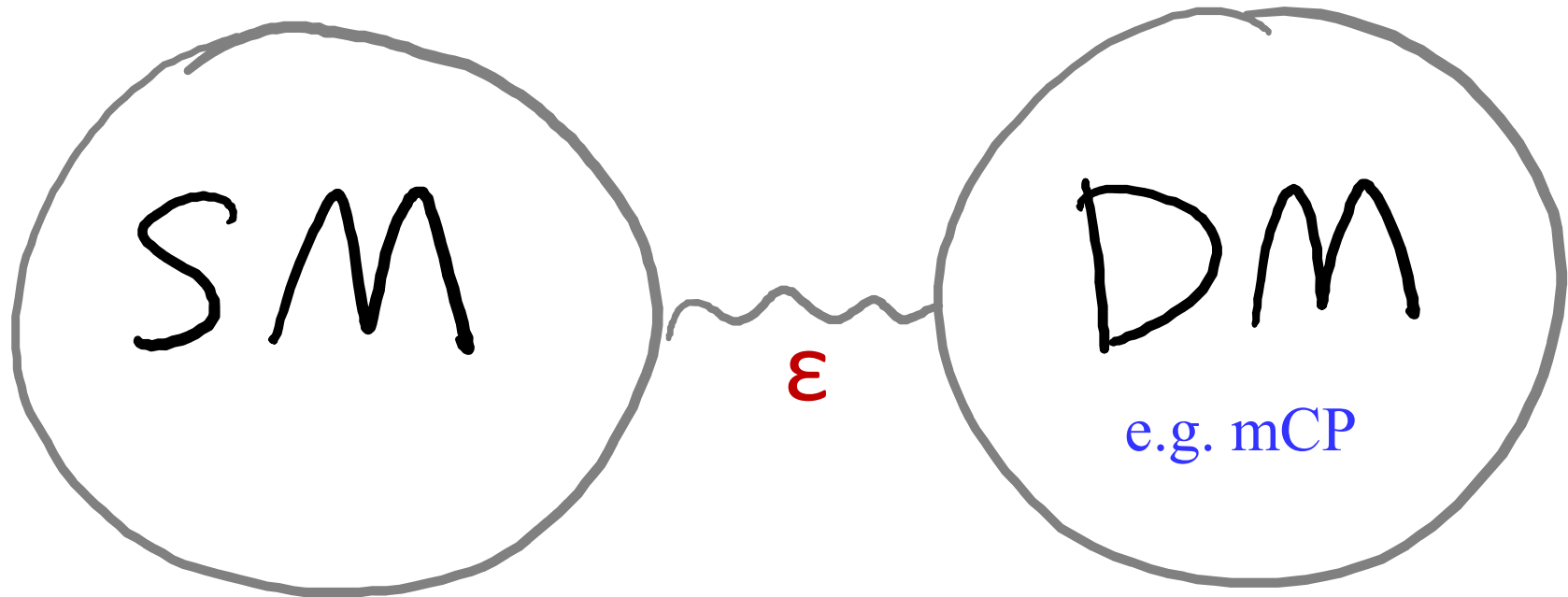
See, Holdom, 1985

$$\mathcal{L} = \mathcal{L}_{\text{SM}} - \frac{1}{4} B'_{\mu\nu} B^{\mu\nu'} - \frac{\kappa}{2} B'_{\mu\nu} B^{\mu\nu} + i\bar{\psi}(\not{\partial} + ie'B' + iM_{\text{mCP}})\psi$$

- Field redefinition into a more convenient basis for massless  $B'$ ,  $B' \rightarrow B' + \kappa B$
- Getting rid of the mixing term,  $B'$  decouple from SM
- After EWSB the new fermion acquires an small EM charge  $Q$  (the charge of mCP  $\psi$ ):

$$Q = \kappa e' \cos \theta_W. \quad \epsilon \equiv \kappa e' \cos \theta_W / e.$$

# The Rise of Dark Sector



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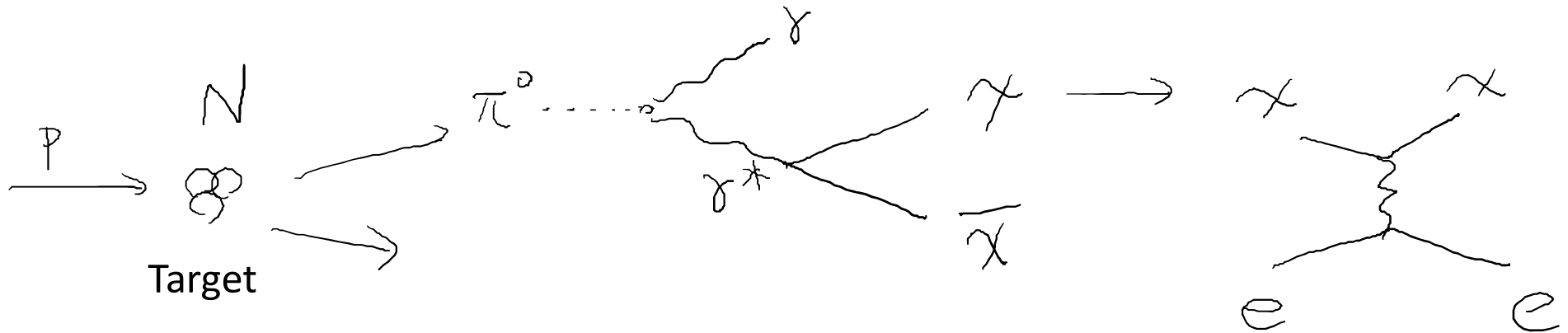
# IMPORTANT NOTE

- Our search is simply a search for particles (fermion  $\chi$ ) with {mass, electric charge} =  $\{m_\chi, \epsilon e\}$
- Minimal theoretical inputs/parameters
- **mCPs do not have to be DM in our searches**
- The bounds we derive **still put constraints on DM as well as dark sector scenarios.**
- **Not considering bounds on dark photon**  
(not necessary for mCP particles)
- **Similar bound/sensitivity applies to scalar mCPs**

# Millicharged Particle: Signature

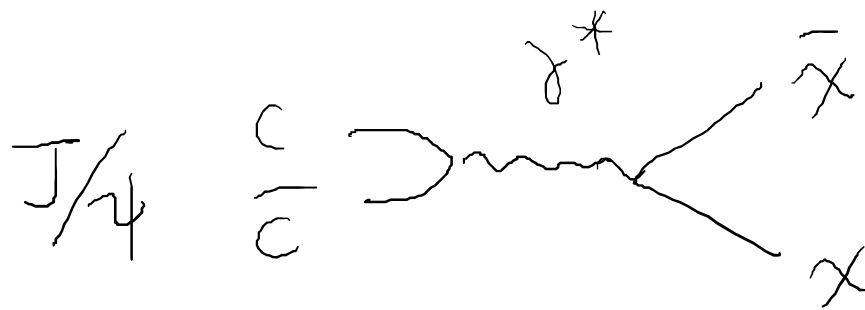
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# MCP: production & detection



□ production:  
meson decays

□ detection:  
scattering electron



$$\text{BR}(\pi^0 \rightarrow 2\gamma) = 0.99$$

$$\text{BR}(\pi^0 \rightarrow \gamma e^- e^+) = 0.01$$

$$\text{BR}(\pi^0 \rightarrow e^- e^+) = 6 * 10^{-6}$$

$$\text{BR}(J/\psi \rightarrow e^- e^+) = 0.06$$

□ Heavy mesons are important  
for higher mass mCP's in high  
enough beam energy



# MCP Signals

- **signal events**  $S_{event}$

$$sig \simeq \sum_{\text{Energies}} N_{\chi}(E_i) \times \frac{N_e}{\text{Area}} \times \sigma_{e\chi}(E_i; m_{\chi}) \times \mathcal{E}.$$

detection efficiency

- $N_{\chi}(E_i)$  represents the number of mCPs with energy  $E_i$  arriving at the detector.  $N_{\chi}(E_i)$  is a function of both the **branching ratio** and **geometric losses** which can vary significantly between experiments
- $N_e$ : **total number of electrons** inside the active volume of the detector
- Area: the active volume divided by the average length traversed by particles inside the detector.
- $\sigma_{e\chi}(E_i)$  is the **detection cross section consistent** with the angular and recoil cuts in the experiment

# MCP productions

- For  $\eta$  &  $\pi^0$ , Dalitz decays:  $\pi^0/\eta \rightarrow \gamma \chi \bar{\chi}$  dominate
- For  $J/\psi$  &  $Y$ , direct decays:  $J/\psi, Y \rightarrow \chi \bar{\chi}$  dominate.  
**Important for high-mass mCP productions!**
- The branching ratio for a meson,  $\mathcal{M}$ , to mCPs is given roughly by

$$\text{BR}(\mathcal{M} \rightarrow \chi \bar{\chi}) \approx \epsilon^2 \times \text{BR}(\mathcal{M} \rightarrow X e^+ e^-) \times f\left(\frac{m_\chi}{M}\right),$$

- M: the mass of the parent meson, X: any additional particles,  $f(m_\chi/M)$ : phase space factor as a function of  $m_\chi/M$ .
- Also consider **Drell-Yan production of mCP** from **q q-bar annihilation**.

# MCP Detection

- Detection signature: **elastic scattering with electrons.**
- Look for **single-electron events**
- **The dominance of electron scattering as a detection signal** is related to the **low-  $Q^2$  sensitivity of the scattering cross section.** ( $Q^2$  is the squared 4-momentum transfer)
- Explicitly, in the limit of small electron mass, we have

$$\frac{d\sigma_{e\chi}}{dQ^2} = 2\pi\alpha^2\epsilon^2 \times \frac{2(s - m_\chi^2)^2 - 2sQ^2 + Q^4}{(s - m_\chi^2)^2 Q^4}.$$

# MCP Detection

- Integrate over momentum transfers, the total cross section will be dominated by the small  $Q^2$  contribution, we have  $\sigma_{e\chi} = 4\pi \alpha^2 \epsilon^2 / Q_{min}^2$ .
- We can relate  $Q_{min}$  in the lab frame to the recoil energy of the electron via  $Q^2 = 2m_e (E_e - m_e)$ .
- An experiment's **recoil energy threshold**,  $E_e^{(min)}$  sets the scale of the detection cross section

$$\sigma_{e\chi} = 2.6 \times 10^{-25} \text{cm}^2 \times \epsilon^2 \times \frac{1 \text{ MeV}}{E_e^{(min)} - m_e}.$$

# MCP Detection

$$\sigma_{e\chi} = 2.6 \times 10^{-25} \text{cm}^2 \times \epsilon^2 \times \frac{1 \text{ MeV}}{E_e^{(\text{min})} - m_e}.$$

- Sensitivity to mCPs can be greatly enhanced by accurately **measuring low energy electron recoils**
- An important feature for **search strategies at future experiments for mCP's and LDM-electron scattering**
- Demonstrated in  
[Magill, Plestid, Pospelov, YT, 1806.03310](#) &  
[\(for sub-GeV DM\) deNiverville, Frugiuele, 1807.06501](#)

# MCP Signals Recap

$$sig \simeq \sum_{\text{Energies}} N_{\chi}(E_i) \times \frac{N_e}{\text{Area}} \times \sigma_{e\chi}(E_i; m_{\chi}) \times \mathcal{E}.$$

Detection efficiency

- $sig \propto \epsilon^4$
- For most of the mCP parameter space under consideration, **electromagnetic decays of mesons** provide the dominant flux contribution
- **Drell-Yan production (DYP) dominates for the large mCP masses** that are only accessible at **DUNE** and **SHiP**.
- For DYP, we integrate over the full production phase-space using **MSTW parton distribution functions**  
[Martin, Stirling, Thorne, Watt, 2009]

# MCP Bound/Sensitivity

- **signal events**  $s_{event}$

$$sig \simeq \sum_{\text{Energies}} N_{\chi}(E_i) \times \frac{N_e}{\text{Area}} \times \sigma_{e\chi}(E_i; m_{\chi}) \times \mathcal{E}.$$

- Our sensitivity curves are obtained by performing a standard sensitivity analysis [PDG, PLB 2010]:
- Given a number of background events  $b$  and data  $n$ , the number of signal events  $s_{event}$ . The  $(1 - \alpha)$  credibility level is found by solving the equation  $\alpha = \Gamma(1 + n, b + sup)/\Gamma(1 + n, b)$ , where  $\Gamma(x, y)$  is the upper incomplete gamma function.
- Throughout this paper, we choose a credibility interval of  $1 - \alpha = 95\%$  ( $\sim 2$  sigma)
- Roughly,  $\epsilon_{sensitivity} \propto E_{e,R,min}^{1/4} Bg^{1/8}$

# Background Estimation for Future Measurements

- Single-electron background for ongoing/future experiments for **MicroBooNE, SBND, DUNE, and SHiP?**
- We consider two classes of backgrounds:
  - 1) Coming from neutrino fluxes (calculable)  
[i.e.  $\nu_e \rightarrow \nu_e$  and  $\nu_n \rightarrow \nu_e$ ], we sum over the neutrino-flux contributions from each collaboration and account for the detection efficiencies.
  - 2) Other sources such as  
beam related: **dirt related events, mis-id particles**  
external: **cosmics**, etc



# Background Reduction

- For neutrino-caused backgrounds, large background reduction is obtained by **imposing the maximum electron recoil energy cuts  $E_e(\text{max})$** .
- This do not significantly affect the mCP signal (which is dominated by low electron recoils from low- $Q^2$ ), **but significantly reduce charged and neutral current neutrino backgrounds.**

# Estimation of Other Background

- **Liquid Argon Time Projection Chamber (LArTPC) can use timing information as vetoes to reduce backgrounds**, Therefore, we multiply our neutrino induced backgrounds by a factor of **10 for LArTPC detectors (MicroBooNE, SBND, and DUNE)**
- For a nuclear emulsion chamber detector, we times a factor of **25 for the background (SHiP)**;
- These increase in the backgrounds decreases our sensitivity to  $\varepsilon$  by 20 – 30%
- **Our results can be easily revised for different background assumptions**, roughly,  $\varepsilon \propto Bg^{1/8}$ .

# Summary Table

|          | Exp. (Beam Energy, POT)                      | $N [\times 10^{20}]$ |        | $A_{\text{geo}}(m_\chi)[\times 10^{-3}]$ |         | Cuts [MeV]         |                    | Bkg |
|----------|--|----------------------|--------|--|---------|--------------------|--------------------|-----|
|          |  | $\pi^0$              | $\eta$ | 1 MeV                                    | 100 MeV | $E_e^{\text{min}}$ | $E_e^{\text{max}}$ |     |
| Existing | LSND (0.8 GeV, $1.7 \times 10^{23}$ )        | 130                  | —      | 20                                       | —       | 18                 | 52                 | 300 |
|          | mBooNE (8.9 GeV, $2.4 \times 10^{21}$ )      | 17                   | 0.56   | 1.2                                      | 0.68    | 130                | 530                | 2k  |
|          | mBooNE* (8.9 GeV, $1.9 \times 10^{20}$ )     | 1.3                  | 0.04   | 1.2                                      | 0.68    | 75                 | 850                | 0   |
| Future   | $\mu$ BooNE (8.9 GeV, $1.3 \times 10^{21}$ ) | 9.2                  | 0.31   | 0.09                                     | 0.05    | 2                  | 40                 | 16  |
|          | SBND (8.9 GeV, $6.6 \times 10^{20}$ )        | 4.6                  | 0.15   | 4.6                                      | 2.6     | 2                  | 40                 | 230 |
|          | DUNE (80 GeV, $3.0 \times 10^{22}$ )         | 830                  | 16     | 3.3                                      | 5.1     | 2                  | 40                 | 19k |
|          | SHiP (400 GeV, $2.0 \times 10^{20}$ )        | 4.7                  | 0.11   | 130                                      | 220     | 100                | 300                | 140 |

- $\varepsilon \propto E_{e,R,\text{min}}^{1/4} Bg^{1/8}$
- $\cos \theta > 0$  is imposed (\*except for at MiniBooNE's dark matter run where a cut of  $\cos \theta > 0.99$  effectively reduces backgrounds to zero [Dharmapalan, MiniBooNE, (2012)]).
- efficiency of 0.2 for Cherenkov detectors, 0.5 for nuclear emulsion detectors, and 0.8 for liquid argon time projection chambers.

# Recasting Existing Analysis: LSND, MiniBooNE, and MiniBooNE\* (DM Run)

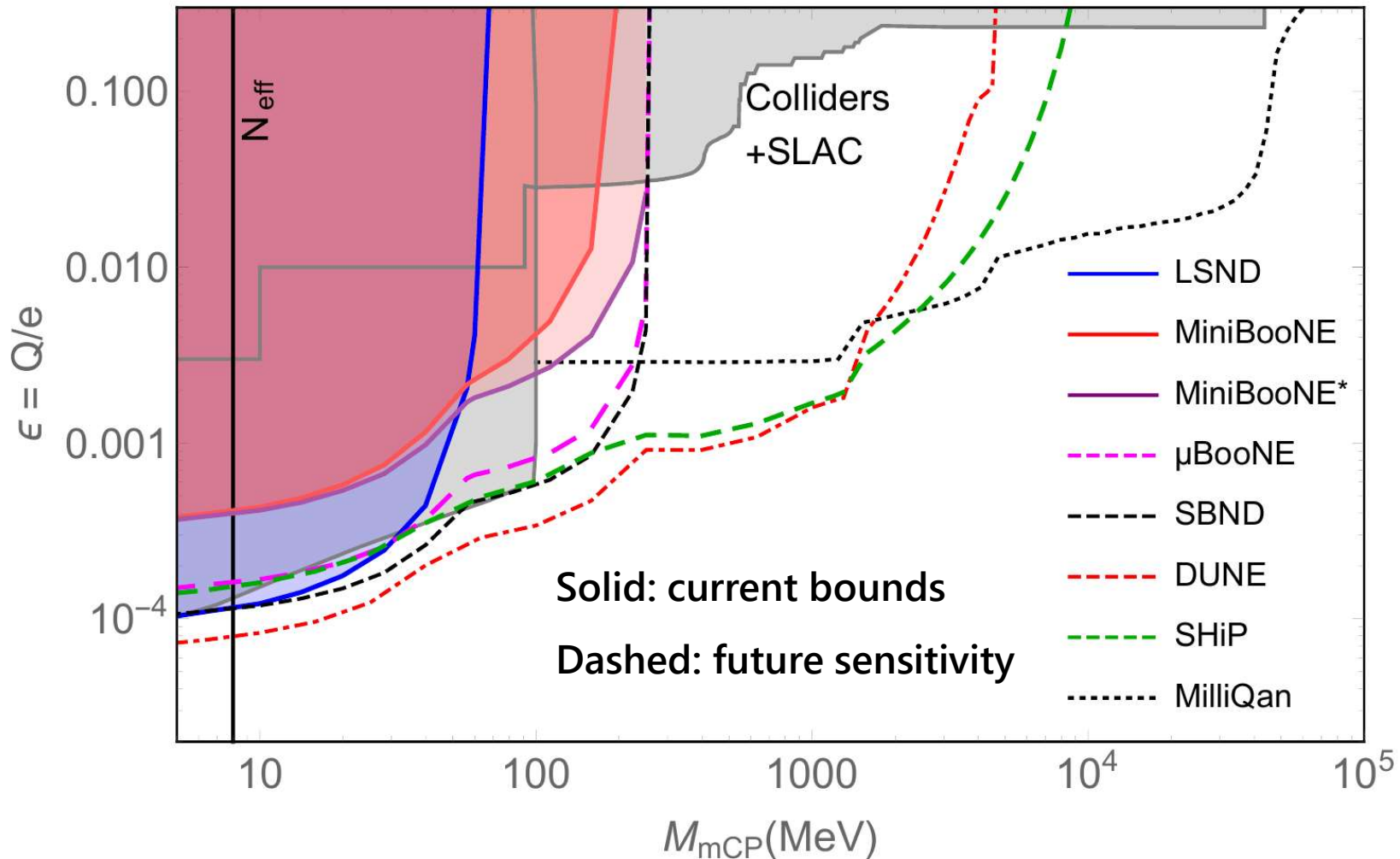
- **LSND**: [hep-ex/0101039](#). Measurement of **electron-neutrino electron elastic scattering**
- **MiniBooNE**: [arXiv:1805.12028](#).  
**Electron-Like Events in the MiniBooNE Short-Baseline Neutrino Experiment**, combines data from both **neutrino and anti-neutrino runs** and consider a sample of  $2.4 \times 10^{21}$  POT for which we take the **single electron background to be  $2.0 \times 10^3$  events** and the **measured rate to be  $2.4 \times 10^3$**
- **MiniBooNE\* (DM run)**: [arXiv:1807.06137](#) (came out after our v1).  
**Electron recoil analysis, details see talk by Dr. Remington**  
We did not include their timing cuts in our calculations, since they were optimized by the MiniBooNE collaboration to the signal's timing profile.

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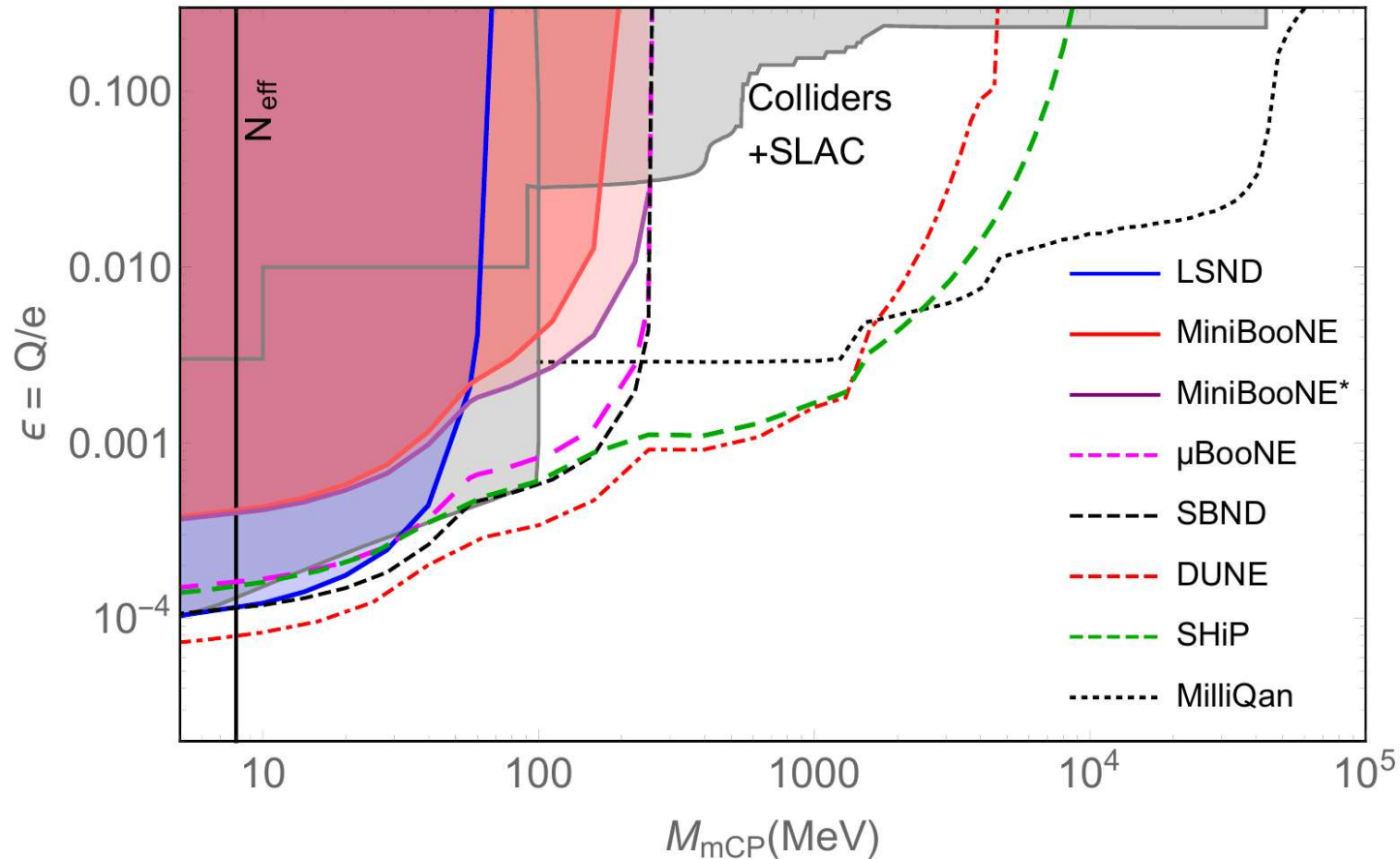
- $\varepsilon \propto E_{e,R,\text{min}}^{1/4} Bg^{1/8}$
- At **LArTPC**, the **wire spacing is about 3 mm**, the ionization stopping power is approximately **2.5 MeV/cm**: electrons with total energy larger than at least **2 MeV** produce tracks long enough to be reconstructed across two wires.

# Probes of mCPs



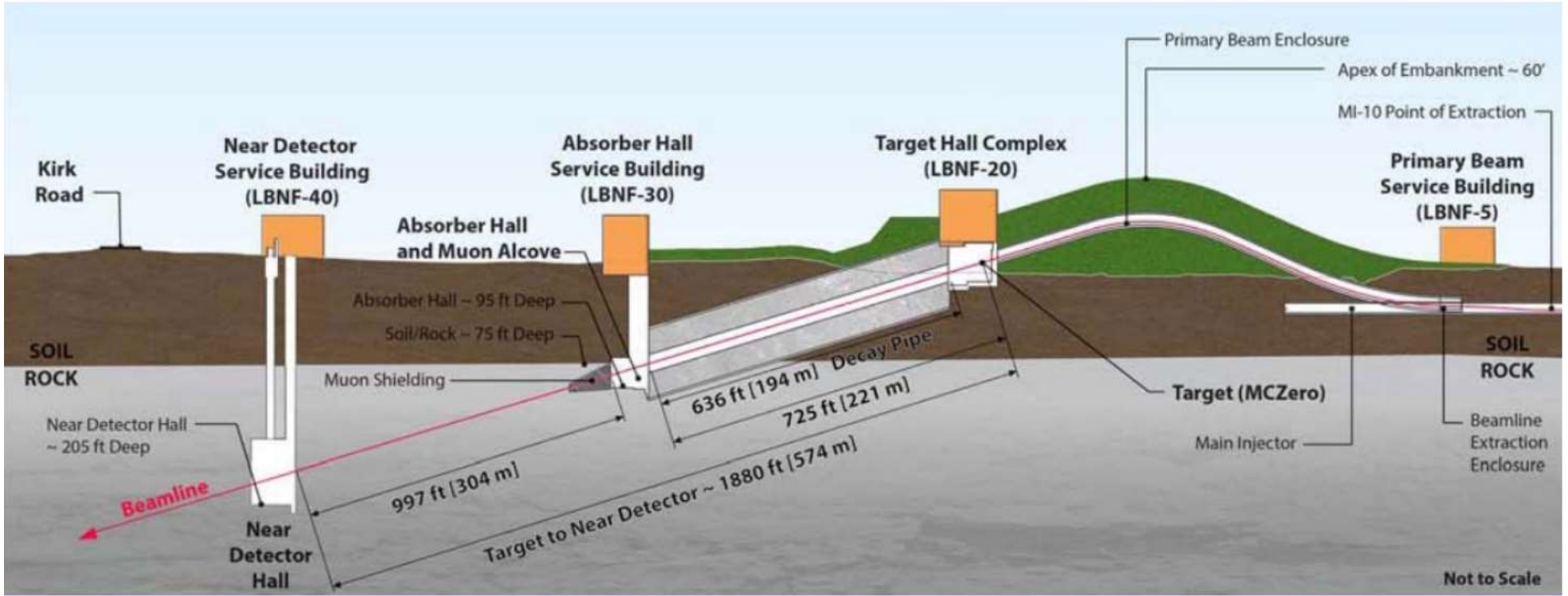
- Heavy neutral meson production turns on in large enough beam energy. Extend mCP mass above 200 MeV
- $\epsilon \propto E_{e,R,\min}^{1/4} B g^{1/8}$  for future experiments

# Existing Bounds & MilliQan



- MilliQan: Haas, Hill, Izaguirre, Yavin, 15, + (LOT arXiv:1607.04669)
- $N_{eff}$ : Böhm, Dolan, and McCabe, (2013)
- Colliders/Accelerator: Davidson, Hannestad, Raffelt 2000 + refs within.
- SLAC mQ: Prinz et al, PRL (1998); Prinz, Thesis, (2001).

# DUNE Near Detector



Jonathan Asaadi – University of Texas Arlington



# Remarks and Outlook

- Our technique can be applied to more generic **light dark matter** and other **weakly interacting particles**
- For **MCP**, and generically **light dark matters**
  - Production from **heavy neutral mesons** important
  - Signature favor **low electron-recoil energy threshold**
- Discussing with **DUNE/SBN experimentalists** about realistic background,  $E_{e,R,min}$  cut, and better simulation (adding **heavy mesons**)
- Sensitivity (along with other new physics scenarios) from **NuMI beam in off-axis detectors**? High energy, lower neutrino background, but also lower statistics ([Yun-Tse Tsai, YT, + ... just started](#))

# To Do List

- **Millicharged particles in electron-based experiments** (electron collisions + fixed target experiments), ([Blinov, Krnjaic, YT, in progress](#)) including BaBar, Belle II, BDX, etc.  
  
(LDMX sensitivity set by [Berlin, Blinov, Krnjaic, Schuster, Toro, 18](#) soon after our paper)
- A lot more to do! Let's chat more during the coffee time.

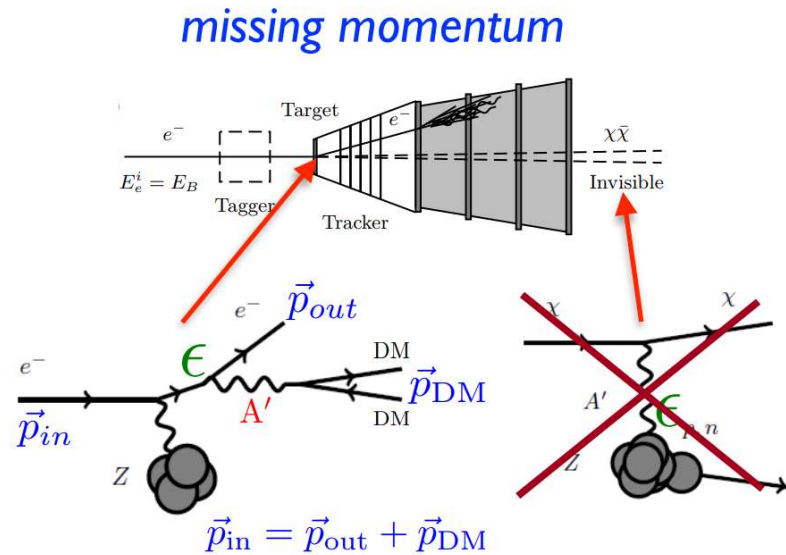
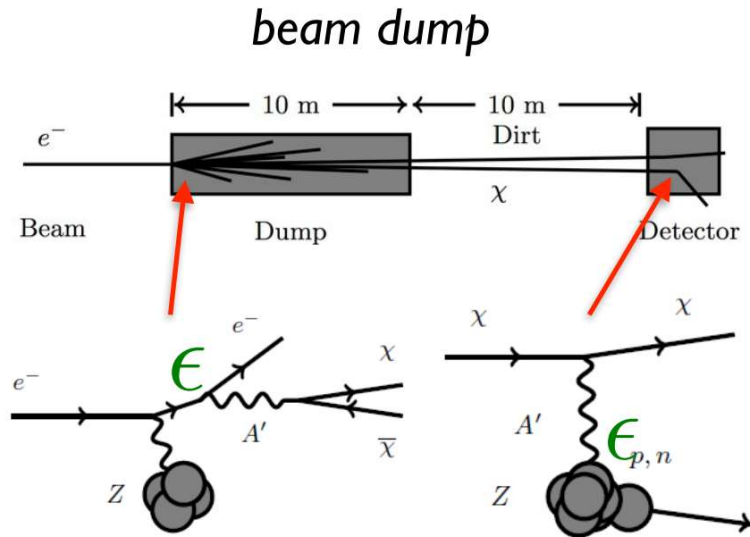
Thank You!  
Thanks for the conference!

Yu-Dai Tsai, Fermilab @ BNL 2018

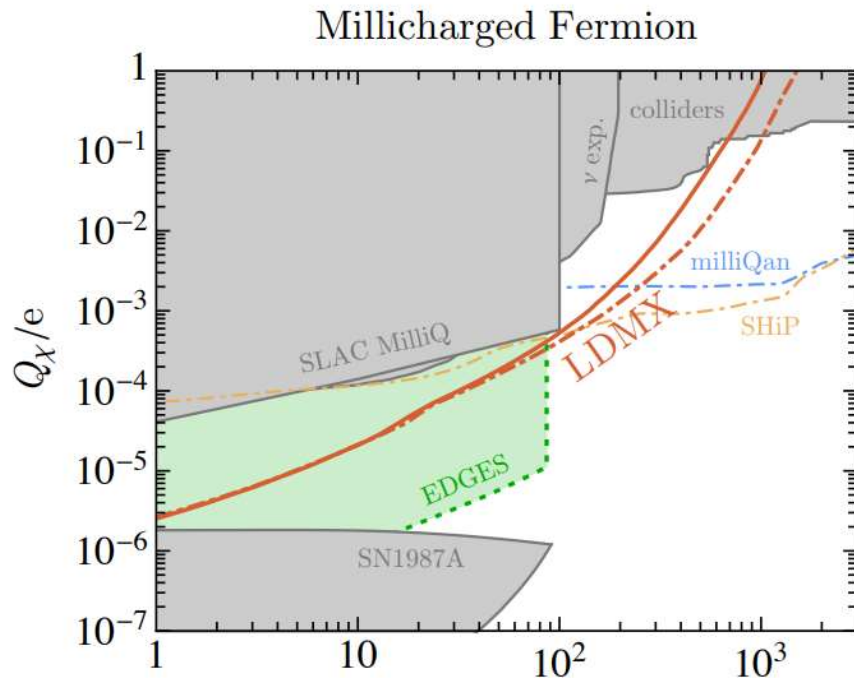
# MCP Existing bounds

- A recent analysis looking for low ionizing particles in **CMS** excluded particles: charge  $\pm e/3$  for  $M_{mCP} < 140$  GeV & particles: charge  $\pm 2e/3$  for  $M_{mCP} < 310$  GeV [CMS, PRD (2013)].
- mCP coupling to the Z is suppressed by  $\sin(\theta_W)$   
 $\sin^2 \theta_W = 1 - (m_W/m_Z)^2 = 0.2223(21)$ .
- **LEP invisible Z width:** mCP not contribute more than the  $2\sigma$  width at LEP. [Davidson, Hannestad, Raffelt, 2000]:  $\epsilon < 0.24$   $m_\epsilon > 45$  GeV,

# LDMX @ SLAC



Tim Nelson, for LDMX, 2017



From Berlin, Blinov, Krnjaic, Schuster, Toro, 18

# BDX @ BNL

Beam Dump eXperiment: Light Dark Matter (LDM) direct detection in a  $e^-$  beam, fixed-target setup<sup>1</sup>

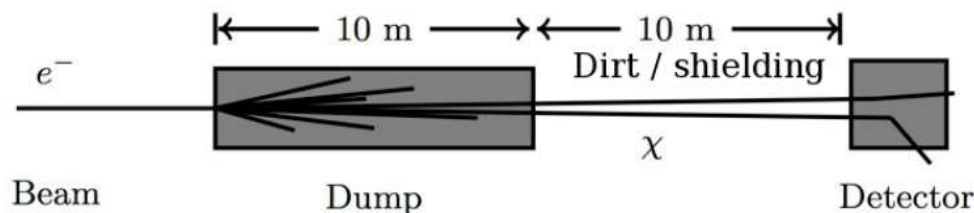
## $\chi$ production

- High-energy, high-intensity  $e^-$  beam impinging on a dump
- $\chi$  particles pair-produced radiatively, through  $A'$  emission

## $\chi$ detection

- Detector placed behind the dump,  $\simeq 20m$
- Neutral-current  $\chi$  scattering on atomic  $e^-$  through  $A'$  exchange, recoil releasing visible energy
- Signal: high-energy EM shower,  $E > .3$  GeV

Number of events scales as:  $N \propto \frac{\alpha_D \varepsilon^4}{m_A^4}$

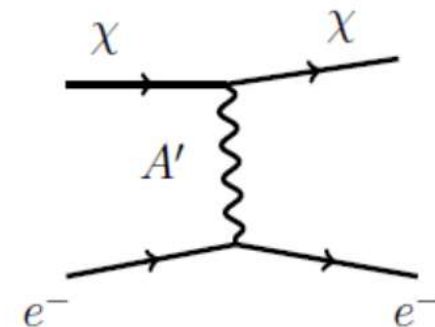
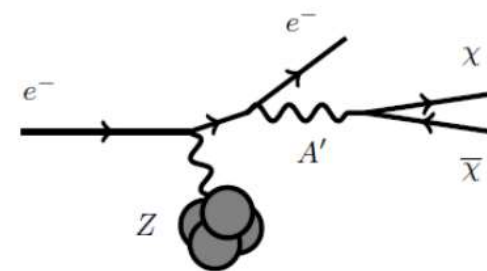


LDM parameters space:

$$M'_A, M_\chi, \varepsilon, \alpha_D$$

$$M'_A \simeq 10 \div 1000 \text{ MeV}$$

$$M_\chi \simeq 1 \div 100 \text{ MeV}$$



Andrea Celentano, INFN-Genova, 2017

# Meson Production List!

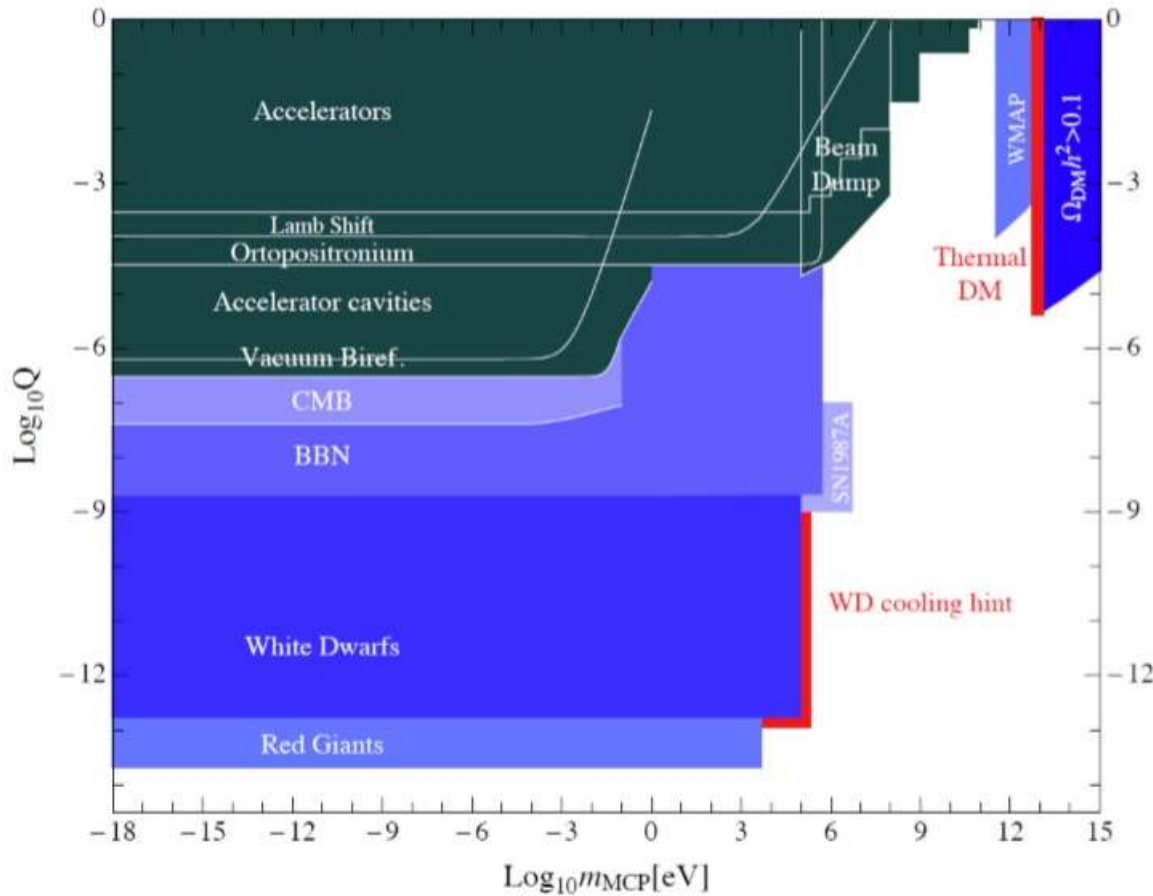
- At LSND, the  $\pi^0$  (135 MeV) spectrum is modeled using a Burman-Smith distribution
- Fermilab's Booster Neutrino Beam (BNB):  $\pi^0$  and  $\eta$  (548 MeV) mesons.  $\pi^0$ 's angular and energy spectra are modeled by the **Sanford-Wang distribution**.  $\eta$  mesons by the Feynman Scaling hypothesis.
- SHiP/DUNE: pseudoscalar meson production using the **BMPT distribution**, as before, but use a beam energy of 80 GeV
- $J/\psi$  (3.1 GeV), we assume that their energy production spectra are described by the distribution from **Gale, Jeon, Kapusta, PLB '99**, nucl-th/9812056.
- Upsilon,  $Y$  (9.4 GeV): Same dist. , normalized by data from HERA-B, I. Abt et al., PLB (2006), hep-ex/0603015.

# Meson Production Check!

- Using this, we have reproduced the Pb rates in [NA50, EPJ '06, nucl-ex/0612012] for  $J/\psi$ , and for  $Y$ , we reproduced the Pt rates in [Herb et al., PRL '77].



# Other Broader Constraints



- Andy Haas, Fermilab, [2017](#)
- Beware of the different assumptions

$Br(o - Ps \rightarrow \text{invisible}) < 4.2 \times 10^{-7}$  (90% C.L.) Badertscher et al, PRD (2007)

$$\delta E = E(2S_{1/2}) - E(2P_{1/2}) \quad \delta E_{VP} \simeq -\frac{\alpha^5 m_e}{30\pi} \left(\frac{m_e}{\mu}\right)^2 \varepsilon^2. \quad \text{Glück, Rakshit, Reya, PRD (2007)}$$

# Other Constraints

- Astrophysics: Cooling/energy loss bounds from stars and SN
- Cosmology: BBN/CMB  $N_{\text{eff}}$
- Laboratory:
  - Invisible decay of ortho-positronium
  - Lamb Shift
  - Accelerators: E613, ASP, LEP, etc
- Andy Haas, Fermilab, [2017](#)

# MCP production

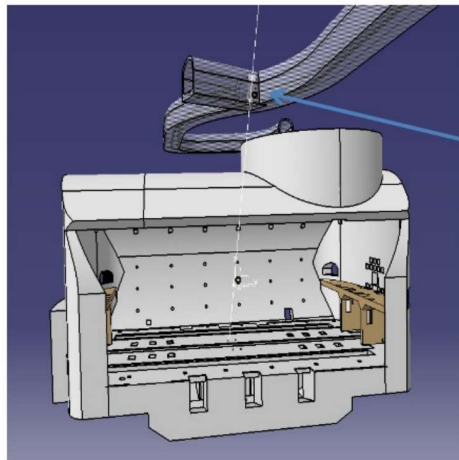
- **Significant branching ratios of electromagnetic decays to lepton pairs** implies an corresponding **decay to pairs of mCPs**, resulting in a **significant flux of mCPs, even with small charges**.
- Fixed target neutrino experiments: produce neutrinos from **weak decays of charged pions**. Similar spectrum of  $\pi^0$  is produced (inferred by comparing the production of  $\pi^0$  and  $\pi^+$  in pp and pn collisions ( [Teis et al, 1996]), or by using counting arguments on the number of production [Norbury & Townsend, NIM B, 2007]).
- For large enough beam energies, other neutral mesons (e.g.  $\eta$ ,  $\Upsilon$ ,  $J/\psi$ ) are also produced.

# Backup: MilliQan at CERN

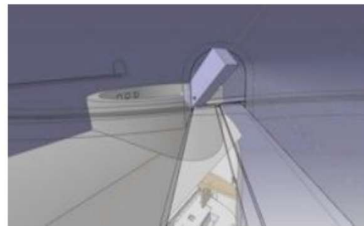
# MilliQan Brief Review

- ~ 1 photo-electron observed per 1 long scintillator
- Require triple incidence in time window
- Moving in CMS “drainage gallery” at P5

- “Drainage Gallery” - an interlocked tunnel above CMS Point 5

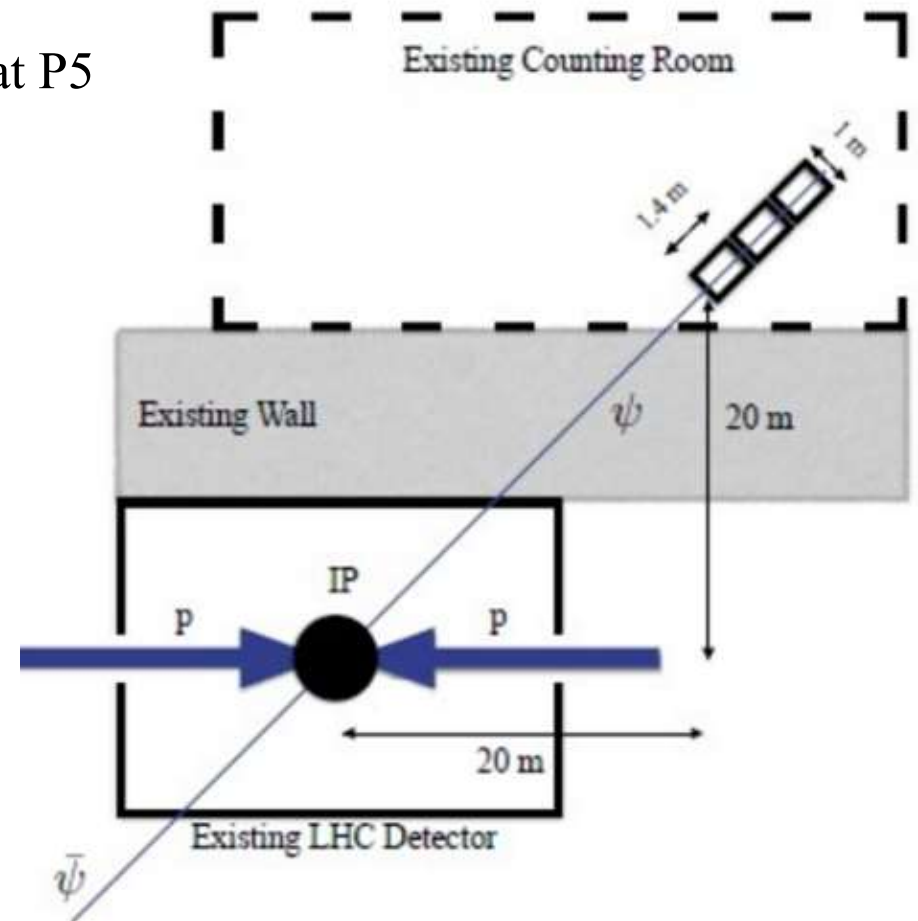


Beam backgrounds shielded by 14m of rock



30m from interaction point  
Small angle from vertical

Andrew Haas, Fermilab (2017)



- MilliQan: Haas, Hill, Izaguirre, Yavin, 15, + (LOT arXiv:1607.04669)