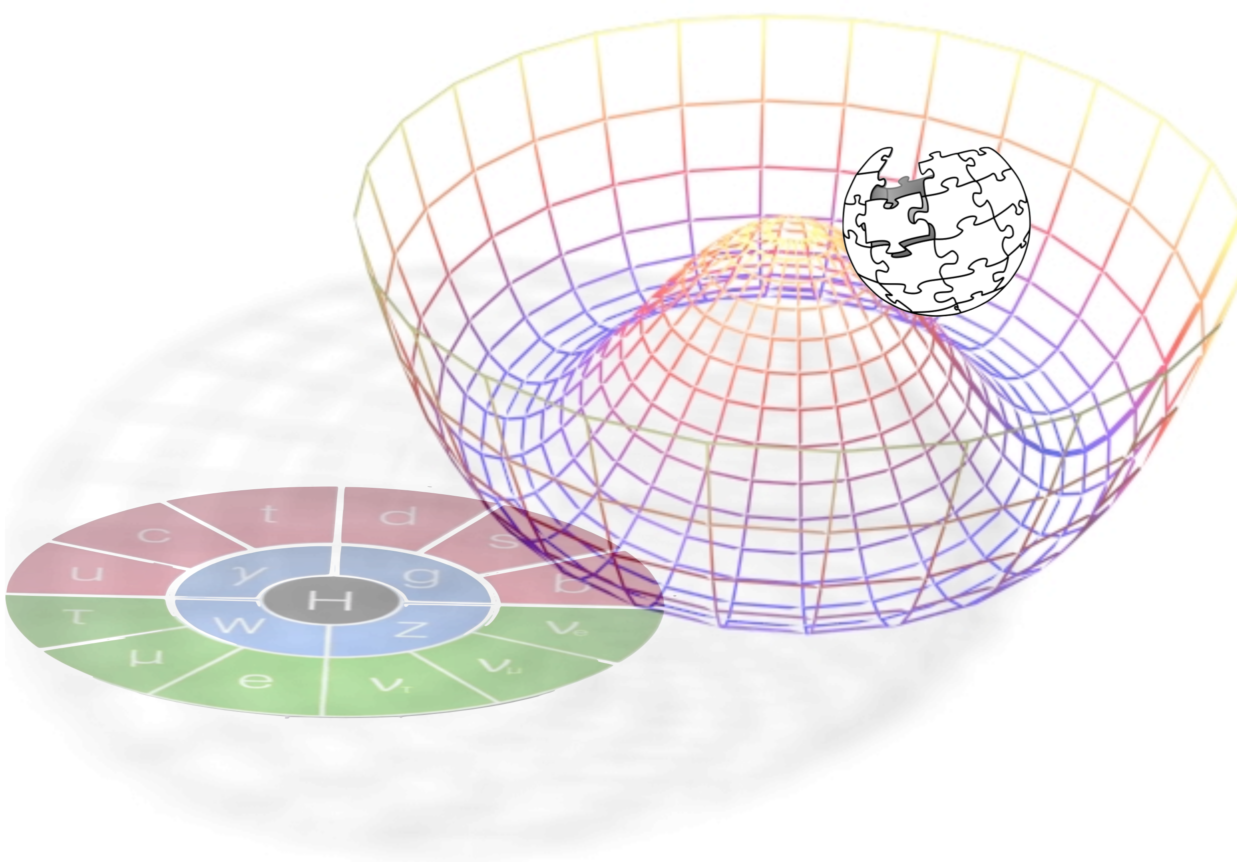
Higgs Physics
Behind and Beyond the SM

Brookhaven Forum 2017, Oct. 12

In Search of New Paradigms



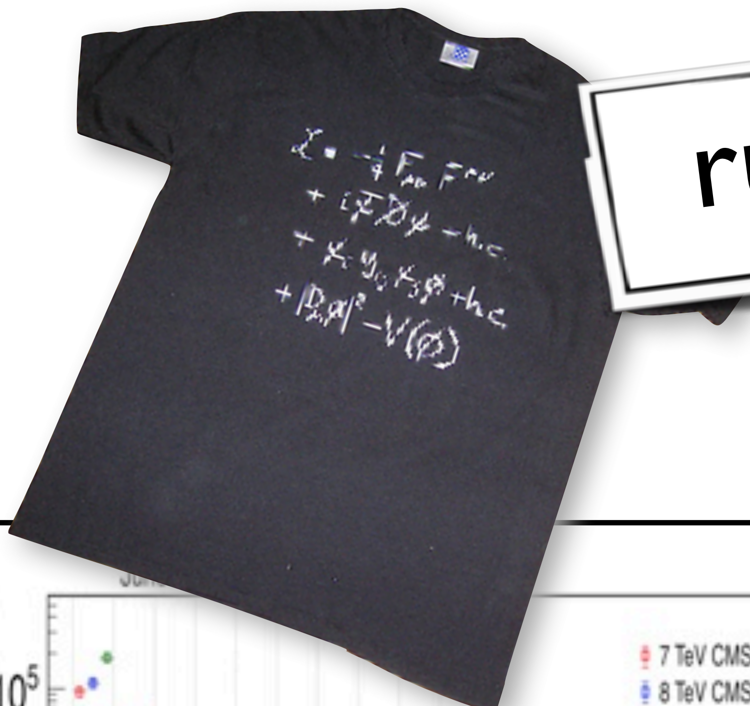
Christophe Grojean

DESY (Hamburg)

Humboldt University (Berlin)

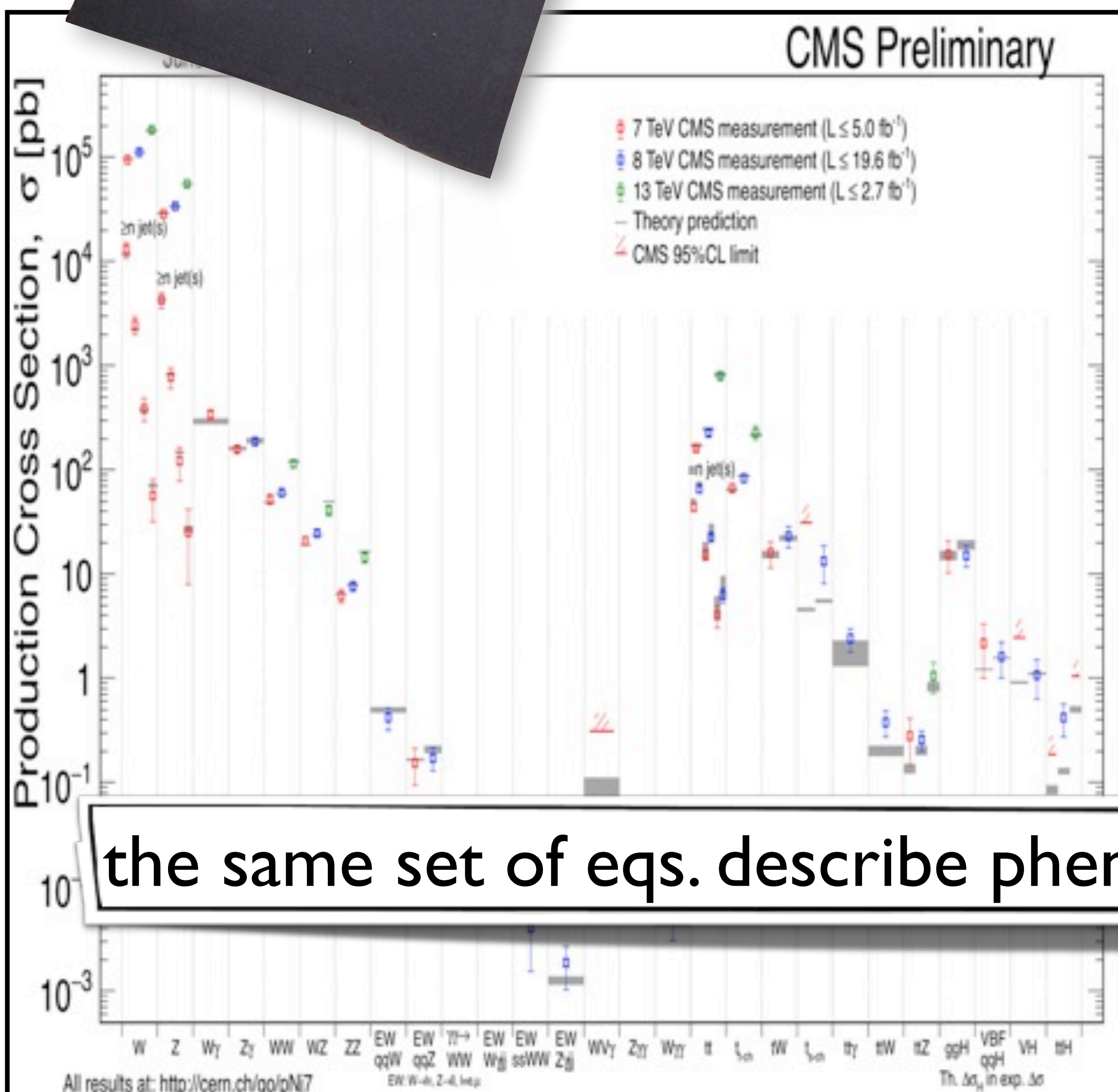
(christophe.grojean@desy.de)

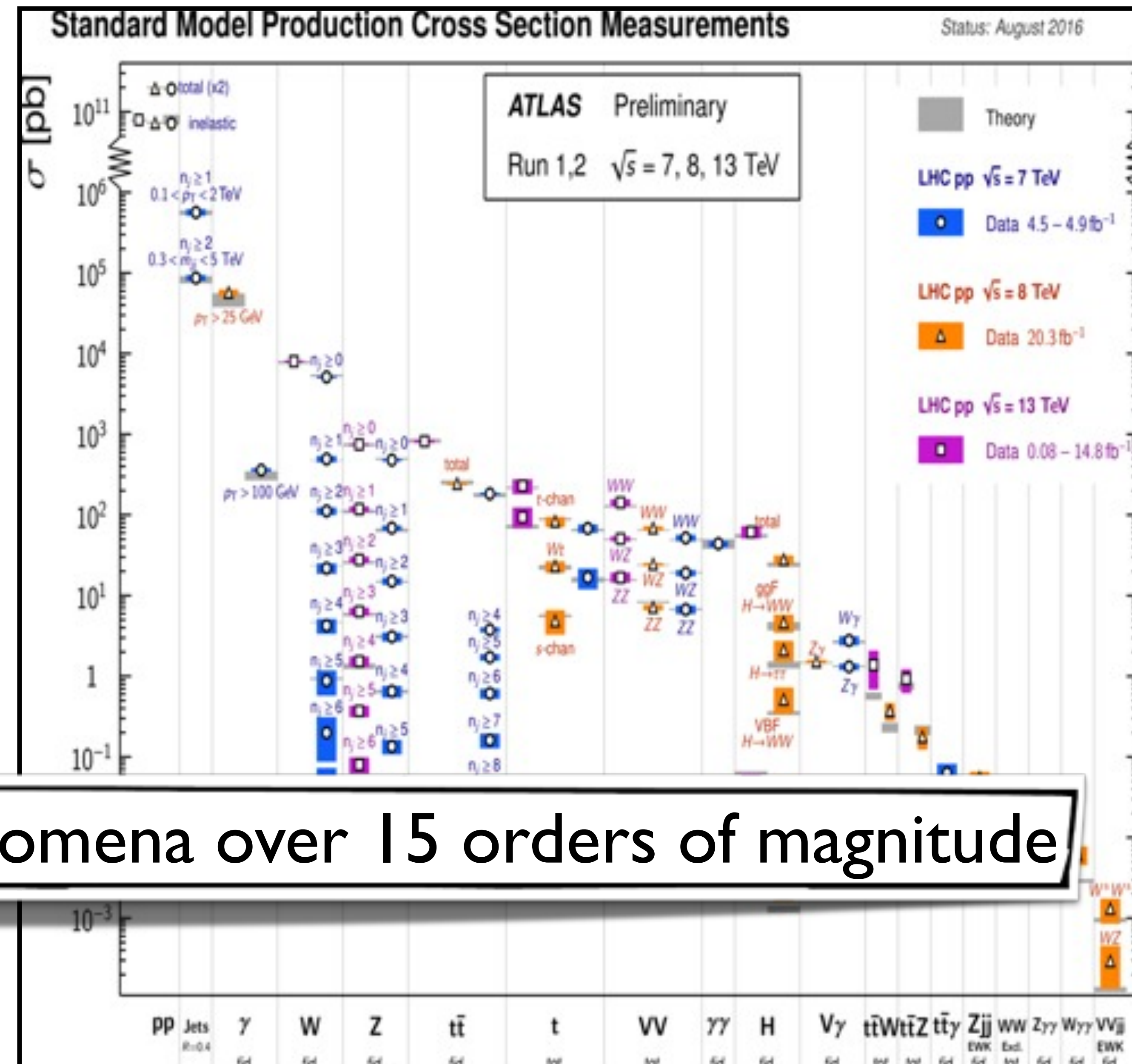
The SM and... the LHC data so far



rules the world!

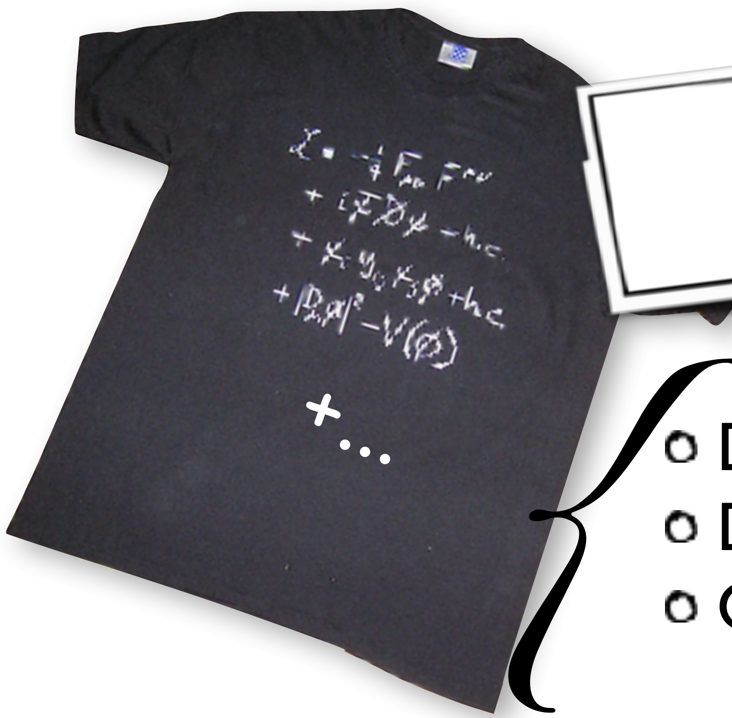
[and we, HEP practitioners, are all entitled for some royalties!]





the same set of eqs. describe phenomena over 15 orders of magnitude

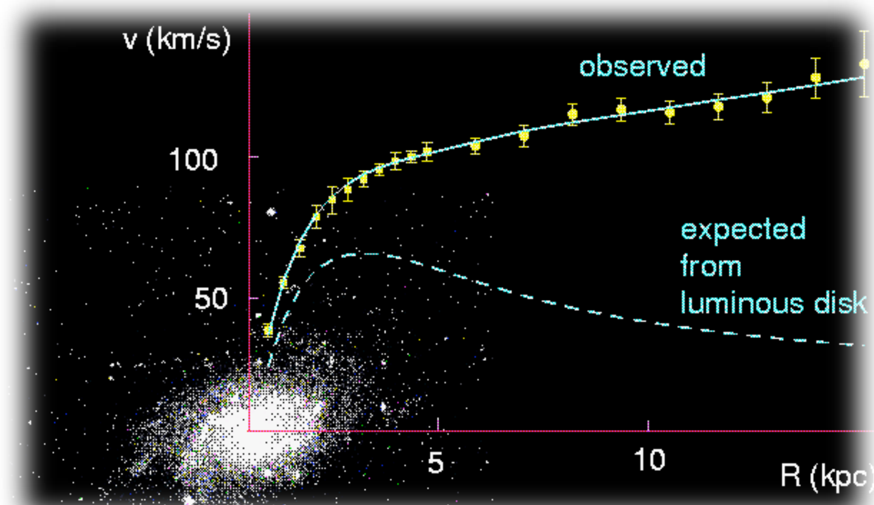
The SM and... the rest of the Universe

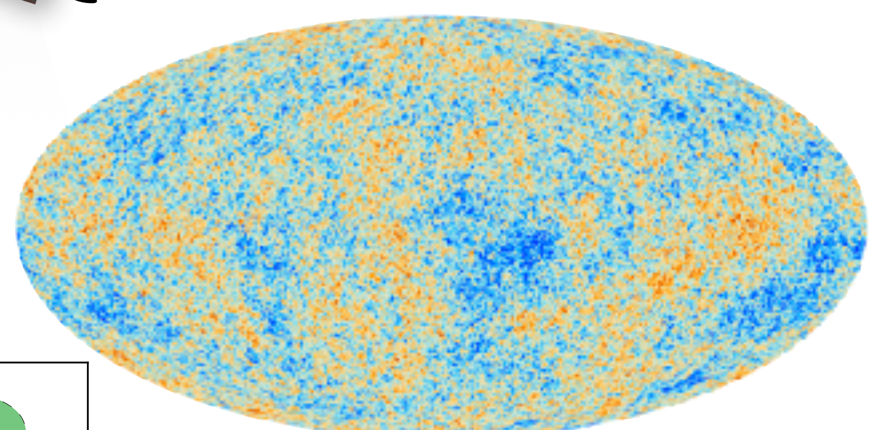


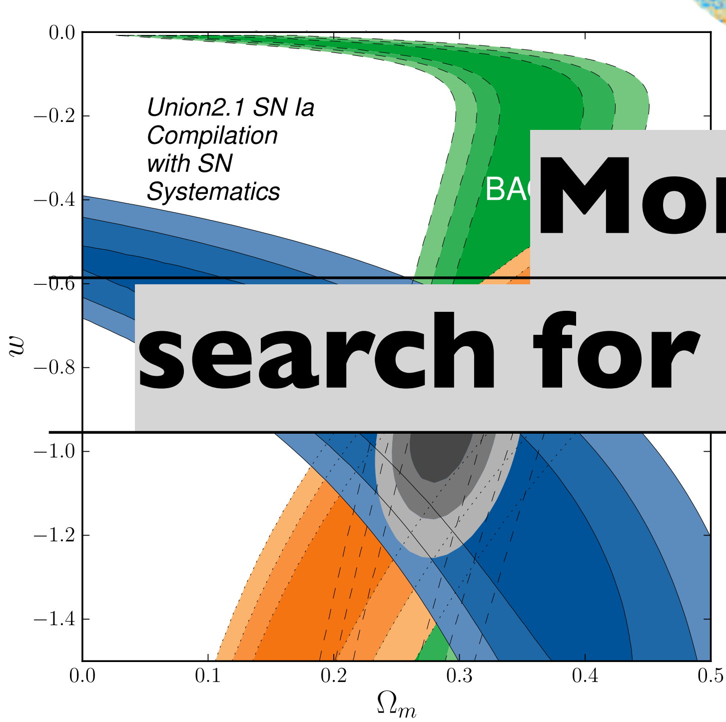
is not enough

[and we all have to return our royalties!]

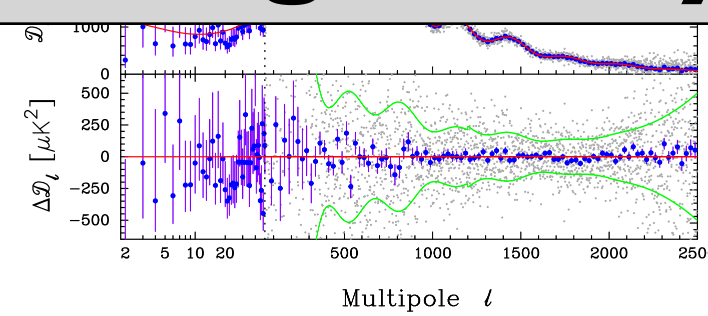
- Dark Matter
- Dark Energy
- Quantum gravity

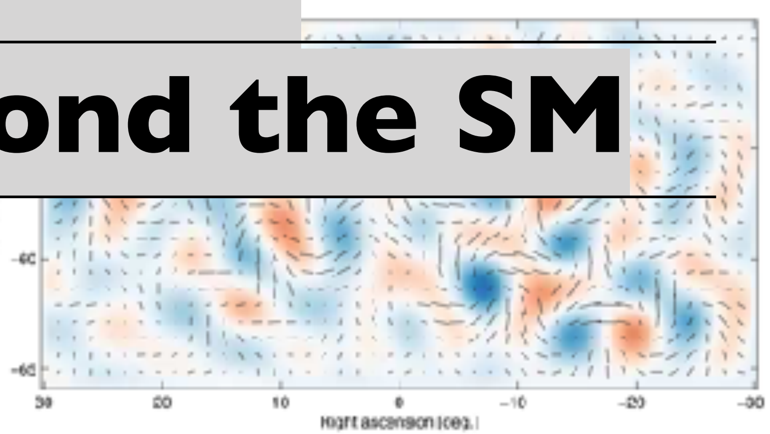






More than ever, it is time to search for New Paradigms Beyond the SM



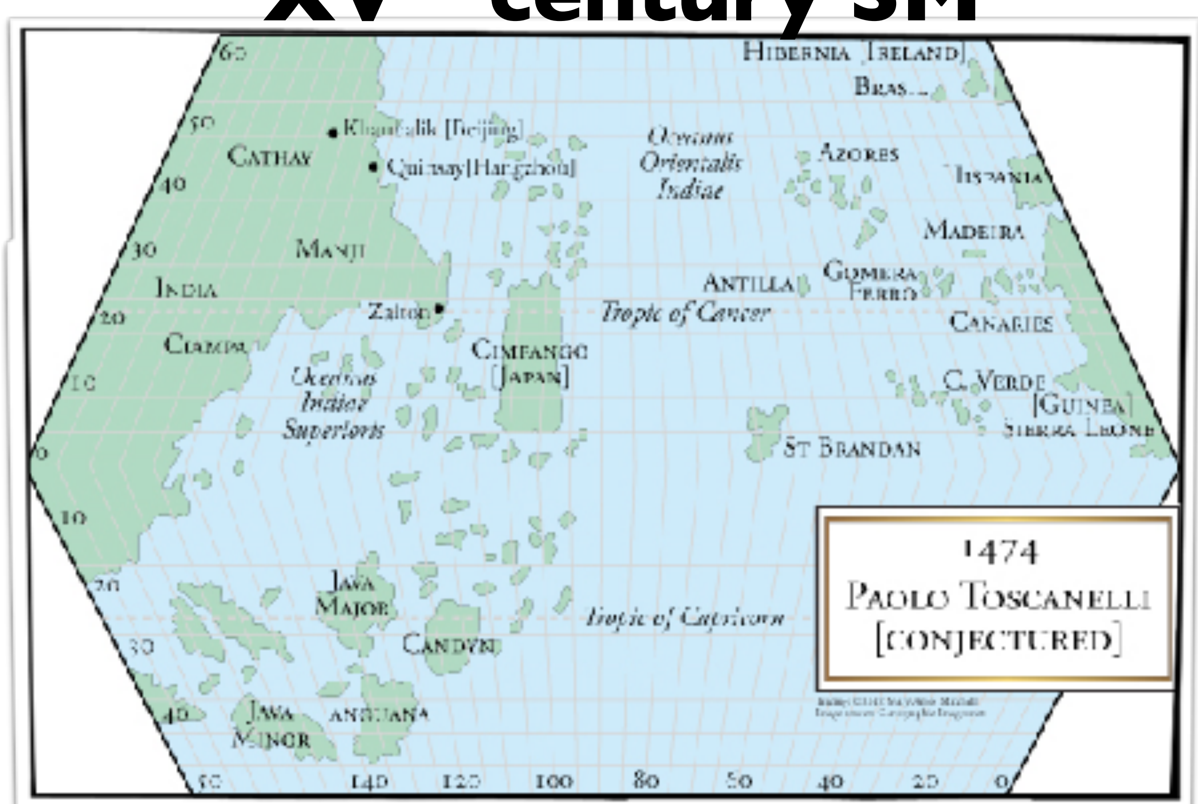


500 years ago: In Search for New Paradigms

Columbus had a great proposal: "reaching India by sailing from the West"

- He had a theoretical model
  - the Earth is round,
  - Eratosthenes of Cyrene first estimated its circumference to be 250'000 stadia
  - other measurements later found smaller values ☞Toscanelli's map
  - lost in unit-conversion or misled by post-truth statements, Columbus thought it was only 70'000 stadia, so he believed he could reach India in 4 weeks
- He had the right technology

XVth century SM



500 years ago: In Search for New Paradigms

Columbus had a great proposal: "reaching India by sailing from the West"

- He had a theoretical model
  - the Earth is round,
  - Eratosthenes of Cyrene first estimated its circumference to be 250'000 stadia
  - other measurements later found smaller values ☞Toscanelli's map
  - lost in unit-conversion or misled by post-truth statements, Columbus thought it was only 70'000 stadia, so he believed he could reach India in 4 weeks
- He had the right technology
  - Caravels were the only ships at that time to sail against the wind, necessary tool to fight the prevailing winds, aka Alizée. NB: this technology was known by the Vikings, but got lost to new generations

His proposal was scientifically rejected twice (by Portuguese's & Salamanca U.)
by the decision was overruled by Isabel ... and America became great (already)

Moral(s)

"if your proposal is rejected, submit it again"

"you need the right technology to beat your competitors"

"theorists don't need to be right!
but progress needs theoretical models to motivate exploration"

High Energy Physics with a Higgs boson

The successes have been breathtaking

- in 4 years, the Higgs mass has been measured to 0.2% (vs 0.5% for the 20-year old top)
- some of its couplings, e.g. κ_γ, have been measured with 1-loop sensitivity (as EW physics at LEP)

The meaning of the Higgs

Particle physics is not so much about particles but more about fundamental principles

- About 10^{-10}s after the Big Bang, the Universe filled with the Higgs substance because it saved energy by doing so:

“the vacuum is not empty”

(even when $\hbar \to 0$, not a Casimir effect)

- The masses are **emergent** quantities due to a non-trivial **vacuum** structure
- There are only a **finite number** of particles (the SM ones) that acquire their mass via the Higgs vev
- There exists a **new type** (non-gauged) of fundamental **forces**: matter-dependent forces ($e \neq \mu$), e.g. familon, relaxion, Higgs portals...

High Energy Physics with a Higgs boson

The successes have been breathtaking

- in 4 years, the Higgs mass has been measured to 0.2% (vs 0.5% for the 20-year old top)
- some of its couplings, e.g. κ_γ, have been measured with 1-loop sensitivity (as EW physics at LEP)

Higgs agenda for the LHC-II, HL-LHC, ILC/CLIC, FCC, CepC, SppC, SHiP

multiple independent, synergetic and complementary approaches to achieve **precision** (couplings), **sensitivity** (rare and forbidden decays) and **perspective** (role of Higgs dynamics in broad issues like EWSB and vacuum stability, baryogenesis, inflation, naturalness, etc)

M.L. Mangano, Washington '15

- rare Higgs decays: $h\to\mu\mu$, $h\to\gamma Z$
- Higgs flavor violating couplings: $h\to\mu\tau$ and $t\to hc$
- Higgs CP violating couplings
- exclusive Higgs decays (e.g. $h\to J/\Psi+\gamma$) and measurement of couplings to light quarks
- exotic Higgs decay channels:

  $h\to \not{E}_T$, $h\to 4b$, $h\to 2b2\mu$, $h\to 4\tau, 2\tau 2\mu$, $h\to 4j$, $h\to 2\gamma 2j$, $h\to 4\gamma$, $h\to \gamma/2\gamma + \not{E}_T$,

  $h\to$ isolated leptons$+ \not{E}_T$, $h\to 2l + \not{E}_T$, $h\to$ one/two lepton-jet(s)+X, $h\to bb + \not{E}_T$, $h\to \tau\tau + \not{E}_T$ ...
- searches for extended Higgs sectors ($H, A, H^\pm, H^{\pm\pm}$...)
- Higgs self-coupling(s)
- Higgs width
- Higgs/axion coupling?
- ...

High Energy Physics with a Higgs boson

The Higgs discovery has been an important milestone for HEP
but it hasn't taught us much about **BSM** yet

typical Higgs coupling deformation: $\frac{\delta g_h}{g_h} \sim \frac{v^2}{f^2} = \frac{g_*^2 \, v^2}{\Lambda_{\rm BSM}^2}$

current (and future) LHC sensitivity
O(10-20)% ⇔ $\Lambda_{\rm BSM}$ > 500(g_*/g_{SM}) GeV

not doing better than direct searches unless in the case of strongly coupled new physics
(notable exceptions: when New Physics breaks some structural features of the SM
e.g. flavor number violation as in h→$\mu\tau$)

Higgs precision program is very much wanted
to probe BSM physics

1

Higgs Portrait

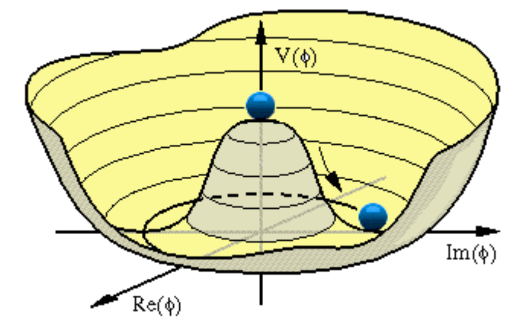
Higgs physics vs BSM

Several deformations away from the SM affecting Higgs properties are already probed in the vacuum

(assuming EW symmetry linearly realized and that new physics is heavy)

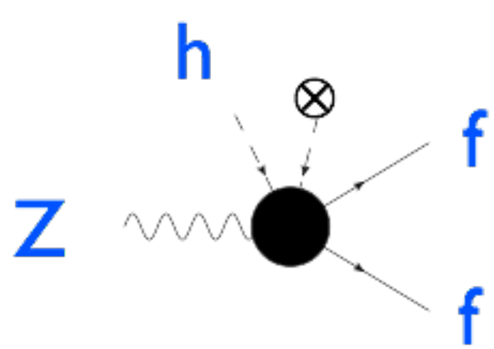
ϕ = v+h

↪ vacuum

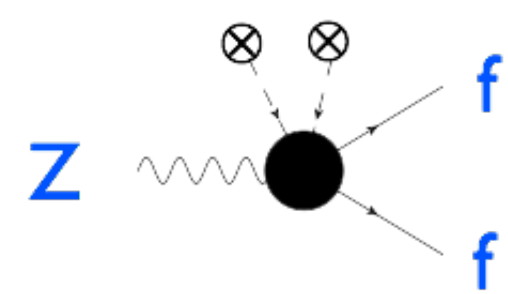


Potentially new BSM-effects in h physics could have been already tested in the vacuum

e.g.



$$= \frac{1}{2v} \times$$



$H^\dagger D_\mu H \bar{f} \gamma^\mu f$

(assuming that the Higgs boson is part of a doublet)

Modifications in h→Zff related to Z→ff

consistency check not discovery mode

One can use h→ZZ→4l to probe this deformation but hard time to compete with LEP bounds

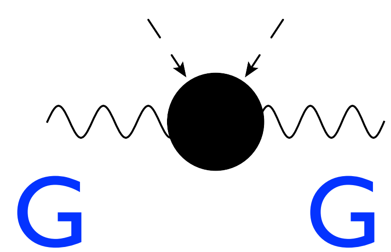
(courtesy of A. Pomarol@Moriond2014)

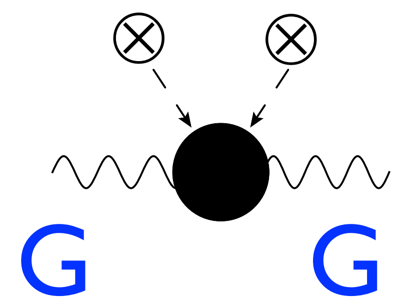
Higgs/BSM Primaries

There are others deformations away from the SM that are harmless in the vacuum and need a Higgs field to be probed

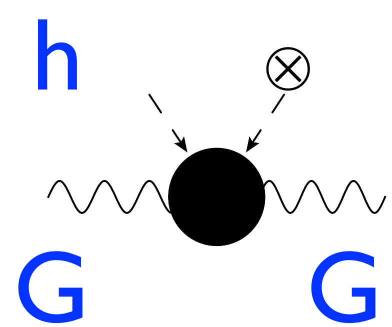
e.g. $$\frac{1}{g_s^2}G_{\mu\nu}^2 + \frac{|H|^2}{\Lambda^2}G_{\mu\nu}^2 \to \left(\frac{1}{g_s^2}+\frac{v^2}{\Lambda^2}\right)G_{\mu\nu}^2$$

operator not visible in the vacuum
(redefinition of input parameter)





But can affect h physics:



affects GG →h!

(courtesy of A. Pomarol@HiggsHunting2014)

Higgs/BSM Primaries

How many of these effects can we have?

Pomarol, Riva '13
Elias-Miro et al '13
Gupta, Pomarol, Riva '14

As many as parameters in the SM: **8** for one family

(assuming CP-conservation)

(courtesy of A. Pomarol@HiggsHunting2014)

g_s	$\lvert H\rvert^2 G^A_{\mu\nu} G^{A\mu\nu}$	→	GGh coupling
g	$\lvert H\rvert^2 B_{\mu\nu} B^{\mu\nu}$	→	hγγ coupling
g'	$\lvert H\rvert^2 W^a_{\mu\nu} W^{\mu\nu\, a}$	→	hZγ coupling
m_W	$\lvert H\rvert^2 \lvert D_\mu H\rvert^2$	→	hVV* (custodial invariant)
m_h	$\lvert H\rvert^6$	→	h^3 coupling
m_f (f=t,b,τ)	$\lvert H\rvert^2 \bar{f}_L H f_R + h.c.$	→	htt, hbb, $h\tau\tau$

yet to be measured at the LHC

the 6 others have been measured (~15%)

Higgs/BSM Primaries

How many of these effects can we have?

Pomarol, Riva '13

Elias-Miro et al '13

Gupta, Pomarol, Riva '14

Almost a 1-to-1 correspondence with the 8 κ's in the Higgs fit

Coupling	300 fb^{-1} Theory unc.: All	Half	None	3000 fb^{-1} Theory unc.: All	Half	None
κ_Z	8.1%	7.9%	7.9%	4.4%	4.0%	3.8%
κ_W	9.0%	8.7%	8.6%	5.1%	4.5%	4.2%
κ_t	22%	21%	20%	11%	8.5%	7.6%
κ_b	23%	22%	22%	12%	11%	10%
κ_τ	14%	14%	13%	9.7%	9.0%	8.8%
κ_μ	21%	21%	21%	7.5%	7.2%	7.1%
κ_g	14%	12%	11%	9.1%	6.5%	5.3%
κ_γ	9.3%	9.0%	8.9%	4.9%	4.3%	4.1%
$\kappa_{Z\gamma}$	24%	24%	24%	14%	14%	14%

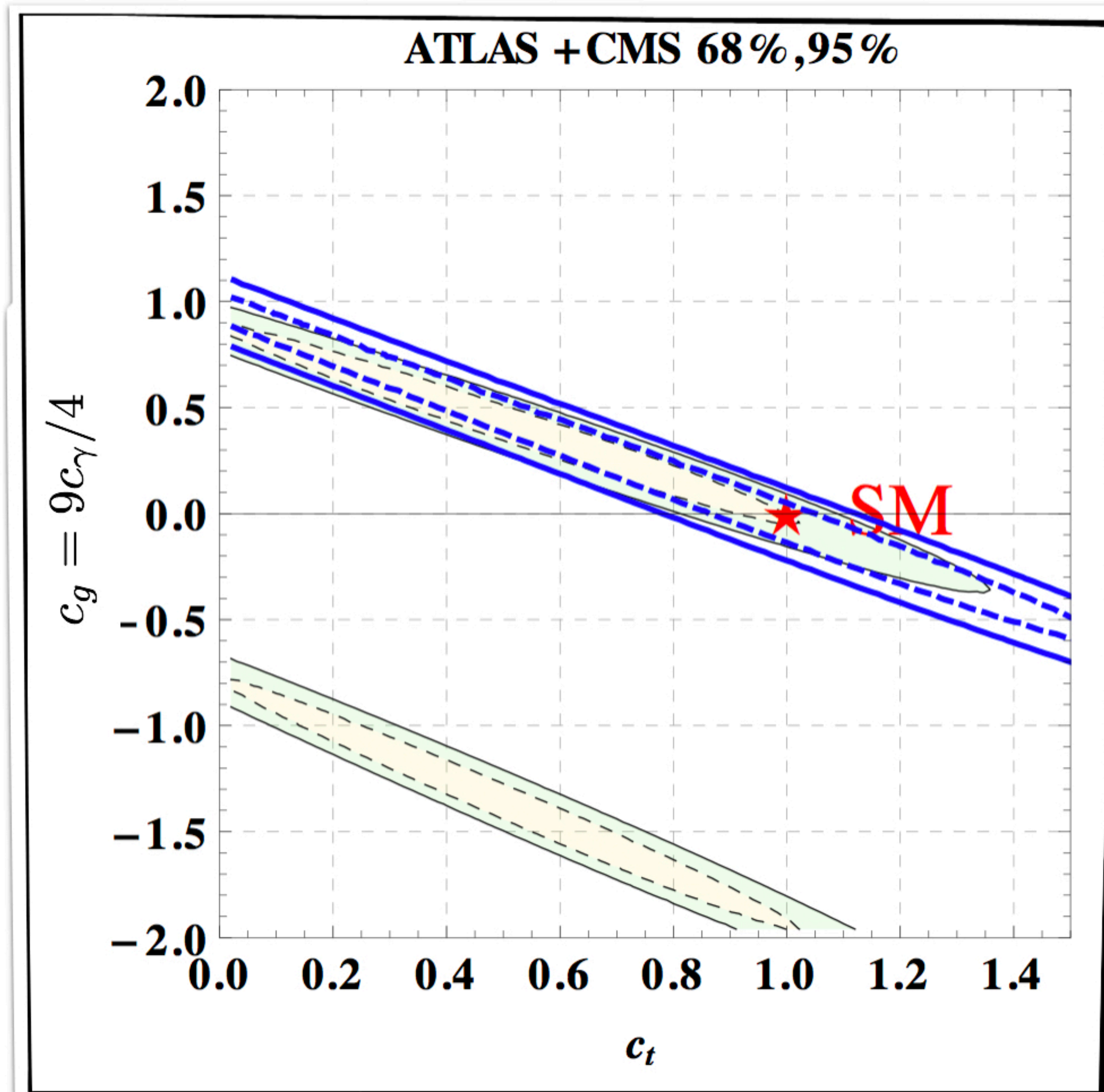
Atlas projection

With some important differences:

1) width hypothesis built-in

2) κ_W/κ_Z is not a primary (constrained by $\Delta\rho$ and TGC)

3) κ_g, κ_γ, $\kappa_{Z\gamma}$ do not separate UV and IR contributions

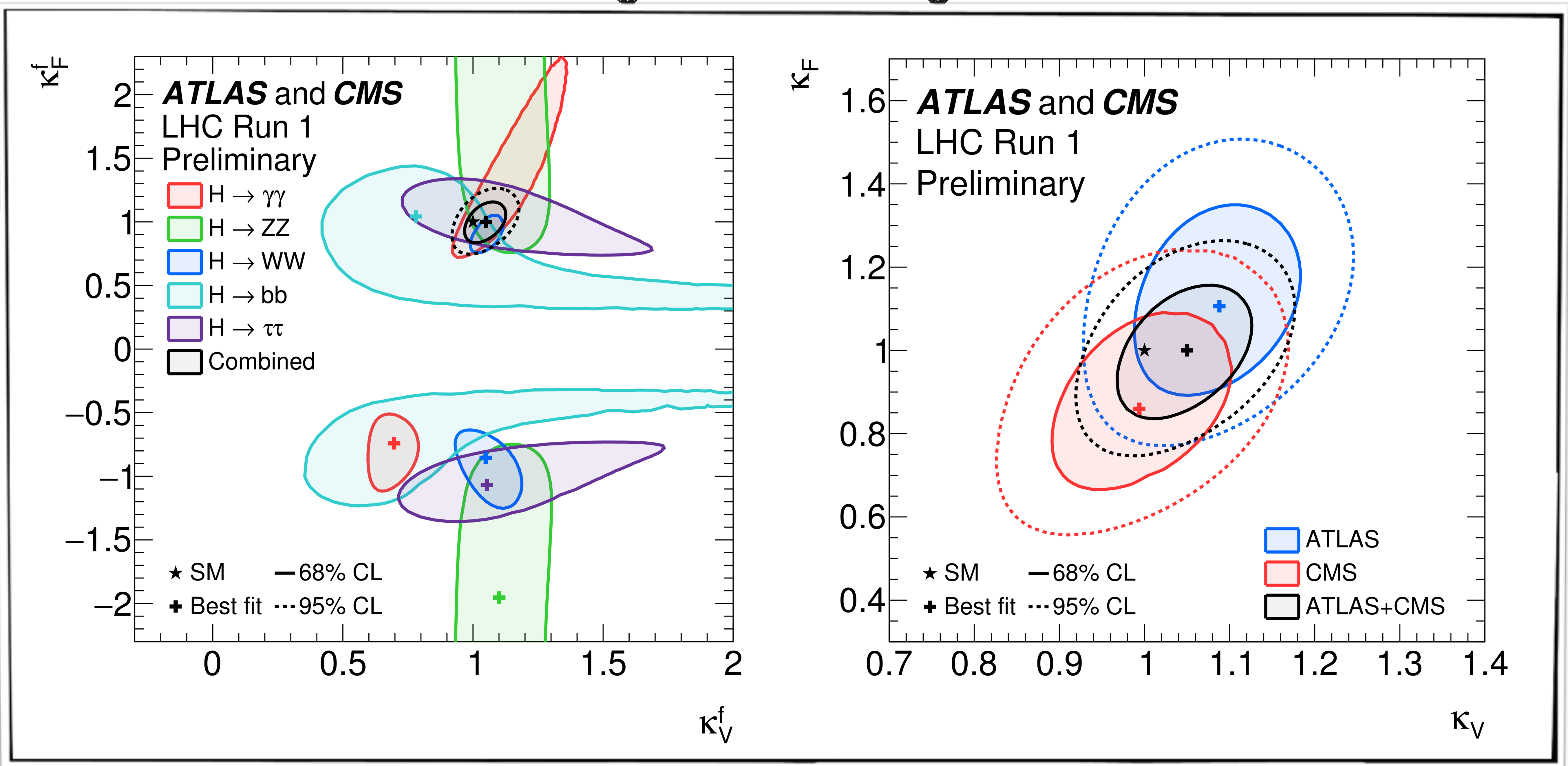


Azatov '15

the 6 others have been measured (~15%) up to a flat direction between between the top/gluon/photon couplings

Why going beyond inclusive Higgs processes?

So far the LHC has mostly produced Higgses on-shell
in processes with a characteristic scale $\mu \approx m_H$



Why going beyond inclusive Higgs processes?

So far the LHC has mostly produced Higgses on-shell
in processes with a characteristic scale $\mu \approx m_H$

⇩ ⇩

access to Higgs couplings @ m_H

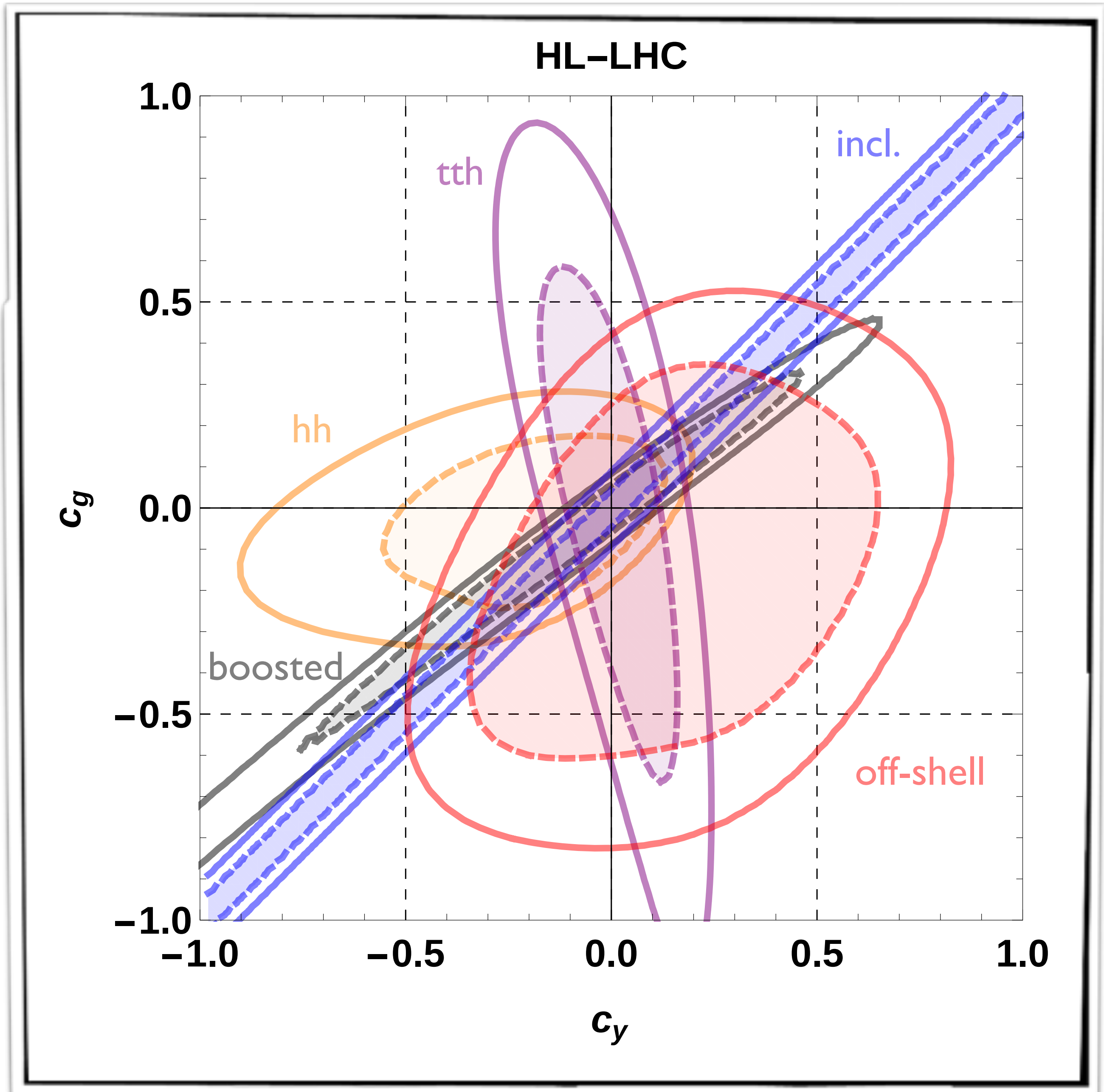
Producing a Higgs with boosted additional particle(s)
probe the Higgs couplings @ large energy
(important to check that the Higgs boson ensures perturbative unitarity)

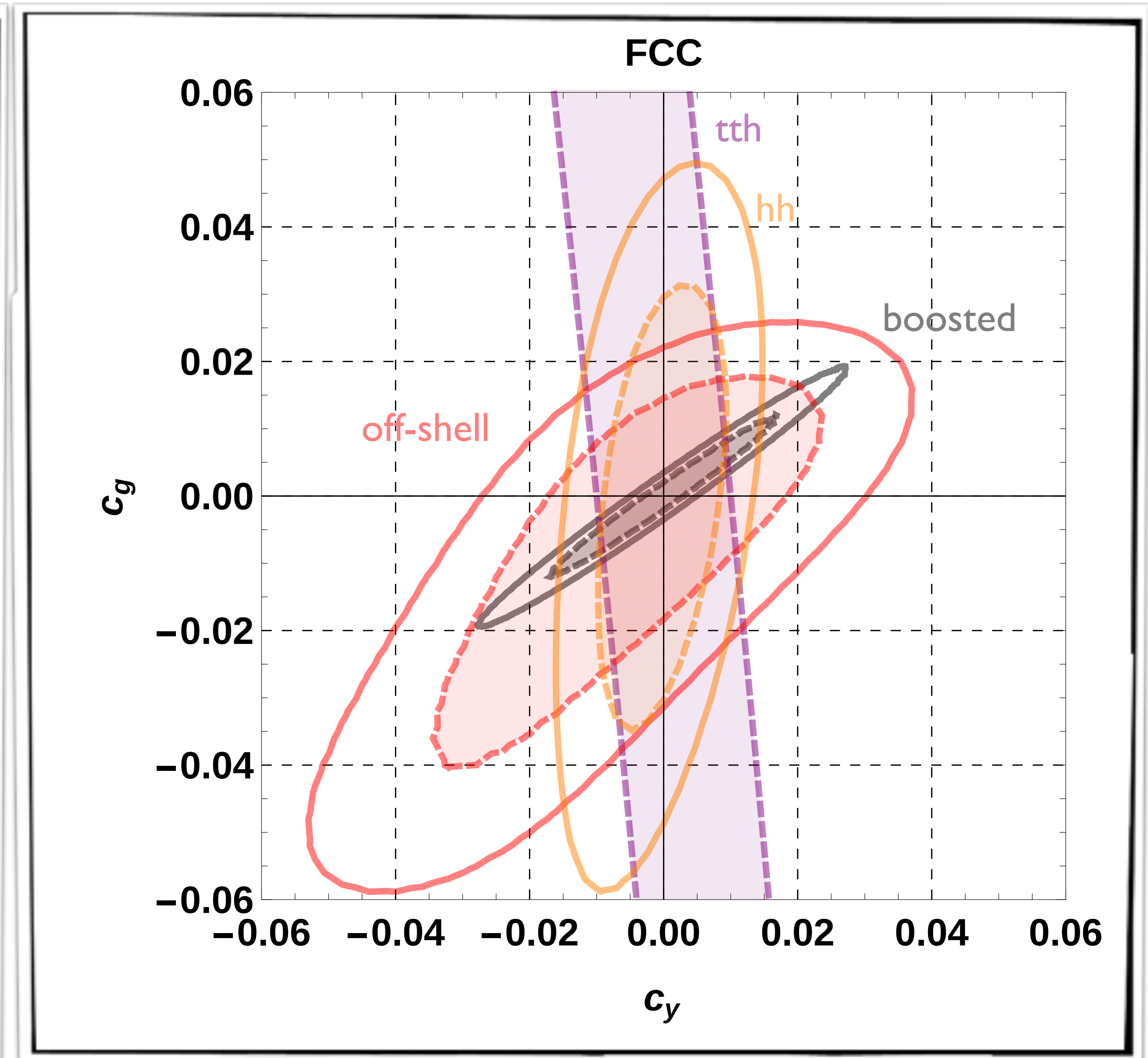
Examples of interesting channels to explore further:

1. off-shell $gg \to h^* \to ZZ \to 4l$
2. boosted Higgs: Higgs+ high-p_T jet
3. double Higgs production

Why going beyond inclusive Higgs processes?

So far the LHC has mostly produced Higgses on-shell
in processes with a characteristic scale $\mu \approx m_H$





Azatov, Grojean, Paul, Salvioni ’16

See also S. Dawson, I.M. Lewis, M. Zeng ‘15

2

Higgs and BSM physics

Higgs & BSM: a love story

In the context of the SM, there is nothing more to learn from the Higgs

- **This is a blessing and a curse:**
  - **A curse, since we might spend the rest of our lives confirming what we already know**
  - **A blessing, since we now have all ingredients required to assess the (in)consistency of exptl data with the SM itself**

Two extreme BSM scenarios...

- EWSB is intrinsically BSM (e.g. composite Higgs)
  - Higgs properties are directly modified
- EWSB is basically SM, it is not affected by BSM
  - Higgs properties are not visibly modified, but BSM particles manifest themselves through the Higgs (e.g. $\chi_2 \to h\chi_1$)

- … plus every scenario in between

(courtesy of M. Mangano@HiggsCouplings2016)

This makes Higgs physics immensely rich, diverse and challenging

Higgs & BSM: a love story

There are many ways to look for BSM physics in the Higgs sector...

BSM Higgs searches

- **Neutral heavy Higgs**
- **Charged Higgs**
- **Di-Higgs production**
- **Higgs as a portal to hidden sectors**
- **BSM constraints from coupling measurements**
- **Higgs in decay chains**
- **Higgs exotic decays**
- **Higgs lepton flavor violating decays**

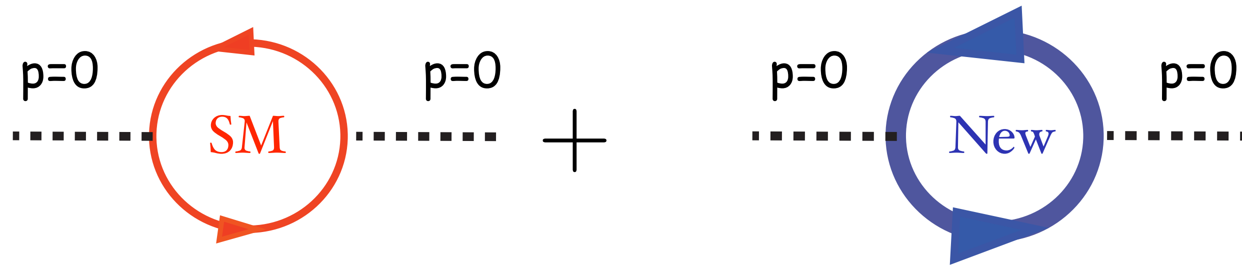
T. Guilemin@Search'16

Plenty of opportunities to discover new particles...
but we also want to learn about new structures and new principles

Higgs couplings as a test of naturalness

$$-(125\ \text{GeV})^2\left(\frac{\Lambda}{600\ \text{GeV}}\right)^2 \qquad\qquad \frac{g_*^2}{16\pi^2}\Lambda^2$$

$$\delta m_H^2 \quad = \quad (\text{SM loop}) \quad + \quad (\text{New loop}) \quad \sim m_H^2$$

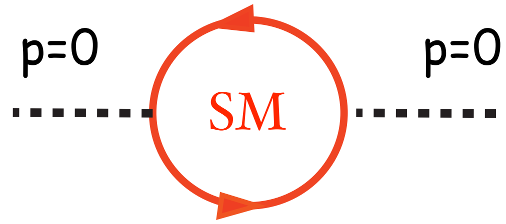


generically

Higgs couplings as a test of naturalness

$$-(125\ \mathrm{GeV})^2\left(\frac{\Lambda}{600\ \mathrm{GeV}}\right)^2 \qquad\qquad \frac{g_*^2}{16\pi^2}\Lambda^2$$

$\delta m_H^2 \quad =$



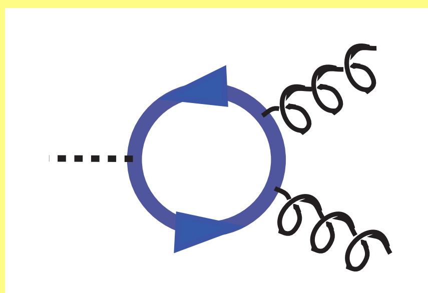
$+$

$\sim m_H^2$

generically

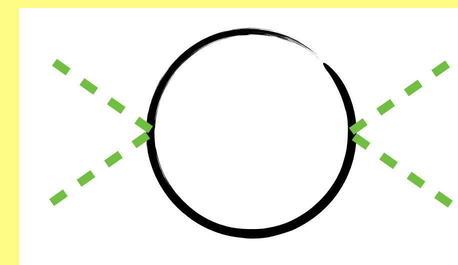
charged particles

neutral particles



$$\frac{g_s^2 g_*^2}{16\pi^2}\frac{1}{m_*^2}|H|^2 G_{\mu\nu}^2 \qquad\qquad \frac{e^2 g_*^2}{16\pi^2}\frac{1}{m_*^2}|H|^2 F_{\mu\nu}^2$$

$$\frac{\Delta BR(h\to\gamma\gamma, Z\gamma, gg)}{\mathrm{SM}} \sim \frac{g_*^2 v^2}{m_*^2}$$



$$\frac{g_*^2}{16\pi^2}\frac{1}{m_*^2}\left(\partial_\mu |H|^2\right)^2$$

$$BR(h\to ii) = BR_{\mathrm{SM}} \qquad \Gamma = \left(1-\frac{g_*^2 v^2}{16\pi^2 m_*^2}\right)\Gamma_{\mathrm{SM}}$$

$$\delta\sigma_{Zh} = -\frac{g_\star^2}{8\pi^2}\frac{v^2}{m_\star^2}$$

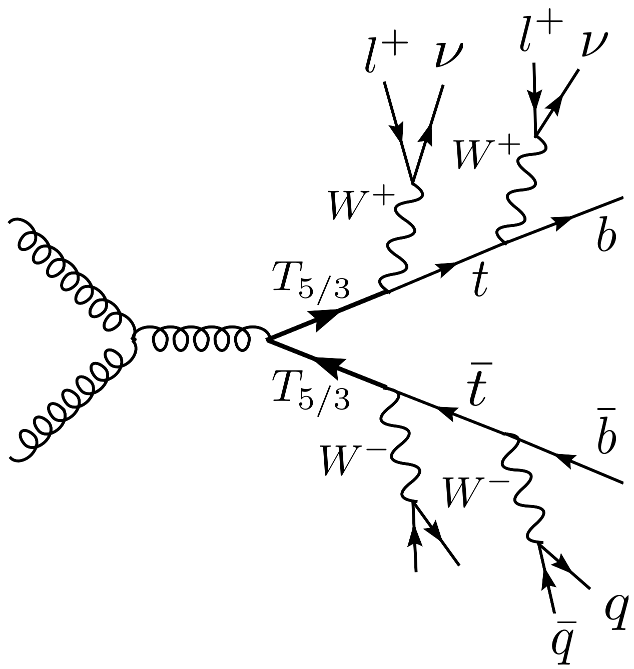
nice to be able to measure Zh & Γ

Colorful naturalness probed @ LHC

Neutral naturalness (invisible?) @ LHC

aka twin Higgs

Top partners in Composite Higgs models



- Search in same-sign dilepton events
- $\ell^\pm$ + 4b final state Aguilar-Saavedra '09

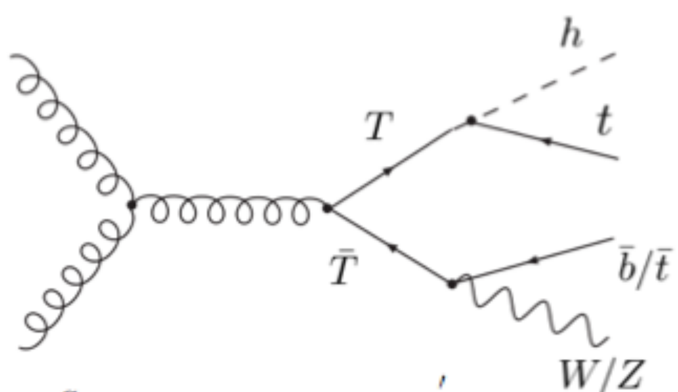
$$T\bar{T} \to HtW^-\bar{b} \to HW^+bW^-\bar{b} \qquad H \to b\bar{b}, WW \to \ell\nu q\bar{q}',$$
$$T\bar{T} \to HtV\bar{t} \to HW^+bVW^-\bar{b} \qquad H \to b\bar{b}, WW \to \ell\nu q\bar{q}', V \to q\bar{q}/\nu\bar{\nu}$$

- $\ell^\pm$ + 6b final state Aguilar-Saavedra '09

$$T\bar{T} \to HtH\bar{t} \to HW^+bHW^-\bar{b} \qquad H \to b\bar{b}, WW \to \ell\nu q\bar{q}'$$

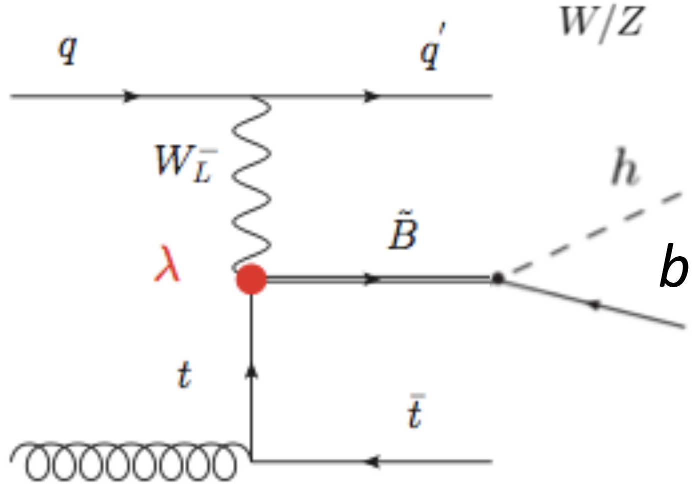
- $\gamma\gamma$ final state Azatov et al '12

$$thbW/thtZ/thth,\ h \to \gamma\gamma$$



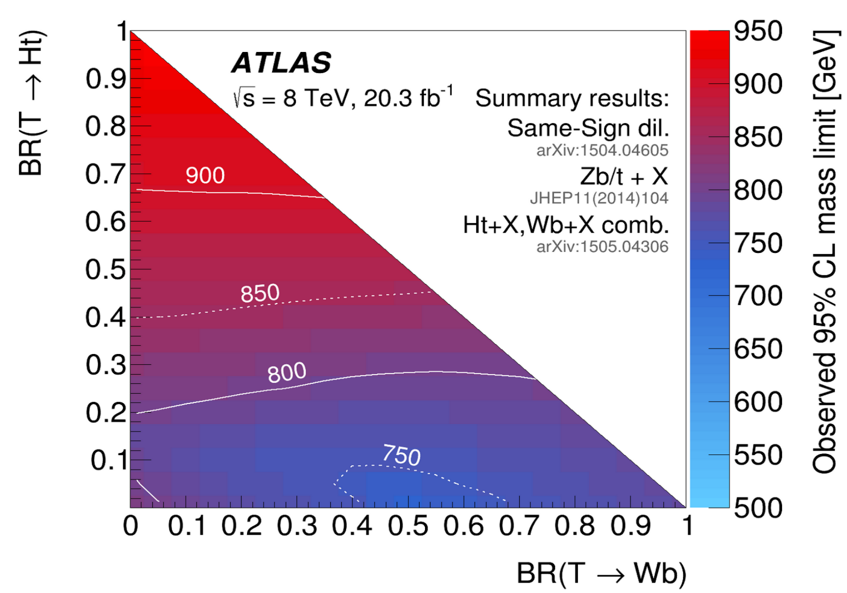
- $\ell^\pm$ + 4b final state Vignaroli '12

$$pp \to (\tilde{B} \to (h \to bb)b)t + X$$



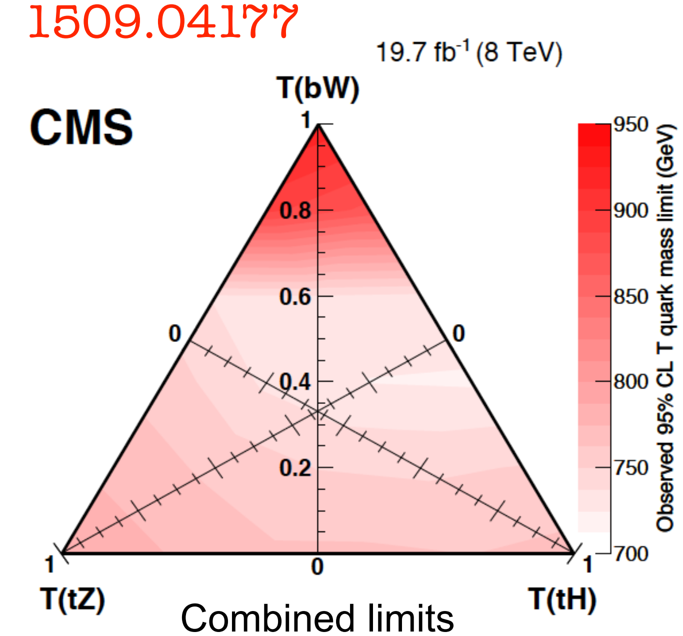
Moriond'17 update bounds above 1 TeV!

1505.04306



(*) Not a combination. Only most restrictive individual bounds shown.

1509.04177



Searching for the missing top partners

" Looking and not finding is different than not looking"

giving the null search results, the top partners should either be

- **heavy** (harder to produce because of phase space)
- **stealthy** (easy to produce but hard to distinguish from background, e.g. $m_{stop} \sim m_{top}$)
- **colorless** (hard to produce, unusual decay)

need to go beyond traditional searches

(C. Verhaaren@NKPI'16)

	Scalar Top Partner	Fermion Top Partner
All SM Charges	SUSY	pNGB/RS
EW Charges	Folded SUSY	Quirky Little Higgs
No SM Charges	???	Twin Higgs

only little corner of theory/model space has been explored so far

require **hidden QCD** with a higher confining scale:

⇒ 1) hidden glueball (0^{++}) that can mix with Higgs
$h \to G_0 G_0 \to 4l$ with displaced vertices

Curtin, Verhaaren '15

⇒ 2) emerging jets

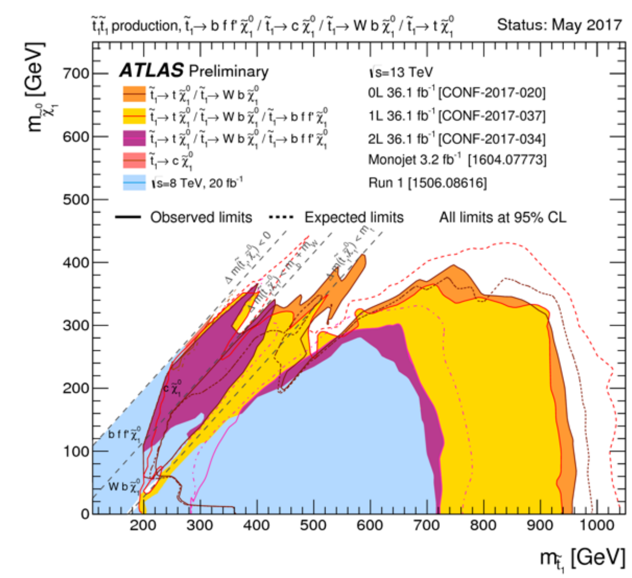
Schwaller, Stolarski, Weiler '15

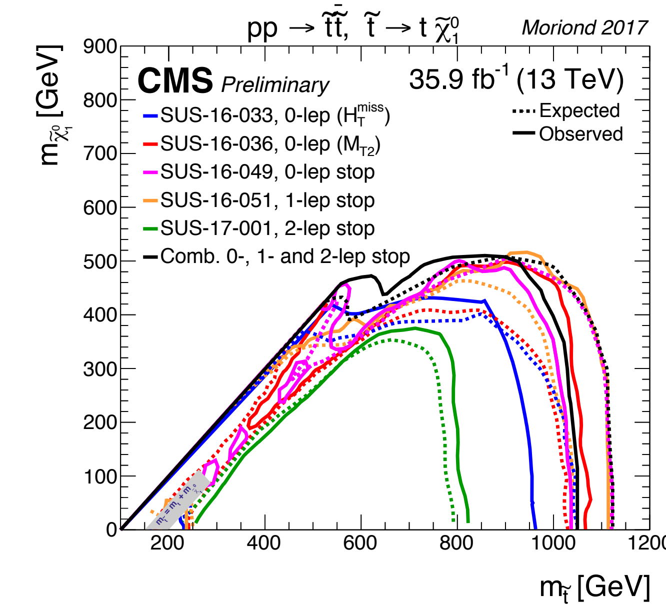
Evading EXP bounds

the stringent bounds from ATLAS/CMS on susy/composite models
force theorists to go beyond minimal/simple models

**SUSY is Natural
but not plain vanilla**

- ~~CMSSM~~
- pMSSM
- NMSSM
- Hide SUSY, e.g. smaller phase space
  - reduce production (eg. split families)
    Mahbubani et al
  - reduce MET (e.g. ~~R-parity,~~ compressed spectrum)
    Csaki et al
  - dilute MET (decay to invisible particles with more invisible particles)
  - soften MET (stealth susy, stop -top degeneracy)
    Fan et al





Generic limit > 1.1 TeV
→ **Δ~50 (2% tuning)**
(2 stops, Λ = 100 TeV)

Compressed limit
> (600 GeV, 500 GeV)
→ **Δ~32 (3% tuning)**

N. Craig @ EPS '17

Evading EXP bounds

the stringent bounds from ATLAS/CMS on susy/composite models
force theorists to go beyond minimal/simple models

SUSY is Natural but not plain vanilla	composite models involving top partners with really exotic charges

- ~~CMSSM~~
- pMSSM
- NMSSM
- Hide SUSY, e.g. smaller phase space
  - reduce production (eg. split families)
    Mahbubani et al
  - reduce MET (e.g. ~~R-parity~~, compressed spectrum)
    Csaki et al
  - dilute MET (decay to invisible particles with more invisible particles)
  - soften MET (stealth susy, stop -top degeneracy)
    Fan et al

"Hyperfolded Composite Higgs"

or how to get spin-1/2 partners with unconventional charges preliminary

Symmetry breaking pattern:

$$\mathrm{SU}(3)_G \times \mathrm{SU}(2)_X \times \mathrm{U}(1)_Z \to \mathrm{SU}(2)_L \times \mathrm{SU}(2)_X \times \mathrm{U}(1)_Y$$

$$\Phi \sim (\bar{3},1)_{\frac{1}{3}} = \exp\left(-i\frac{\pi^a T_G^a}{f}\right)\begin{pmatrix}0\\0\\f\end{pmatrix} \approx \begin{pmatrix}H\\ f-\frac{H^\dagger H}{2f}\end{pmatrix}$$

SM electroweak group generators:

$$T_L^{1,2,3} = T_G^{1,2,3} \qquad Y = Z - \frac{T_G^8}{\sqrt{3}} + \left(\frac{2}{3} - Y_T\right) T_X^3$$

free parameter, to become the top-partner hypercharge

Since the charge-Y_T partner does not mix with the SM quarks, the usual decays to W/Z/h + quark are absent.

Instead, the decay may proceed via a higher-dimensional operator. For example, the operator

$$\mathcal{L} \propto \bar{X}_\alpha^\dagger \bar{u}_{i\beta}^\dagger \bar{d}_j^\alpha \bar{d}_k^\beta + \mathrm{h.c.}$$

may give the potentially elusive decays

$$X \to jjj,\ tjj$$

Y. Kats @ DESY '16

3

BSM Higgs couplings: Baryogenesis

EW scale flavons for EW baryogenesis

Baldes, Konstandin, Servant '16

Electroweak baryogenesis requires:

- A strong first order phase transition
- Sufficient CP violation

However in the SM:

- The Higgs mass is too large
- Quark masses are too small

These negative results are tied to the fact that
Yukawa couplings during EW phase transition are identical the ones afterwards
What if they were larger?
E.g. flavor structure emerges during the EW transition

$$y_{ij} \bar{f}_L^i H f_R^j \quad \Rightarrow \quad y_{ij} \left(\frac{\chi}{M}\right)^{q_H + q_j - q_i} \bar{f}_L^i H f_R^j$$

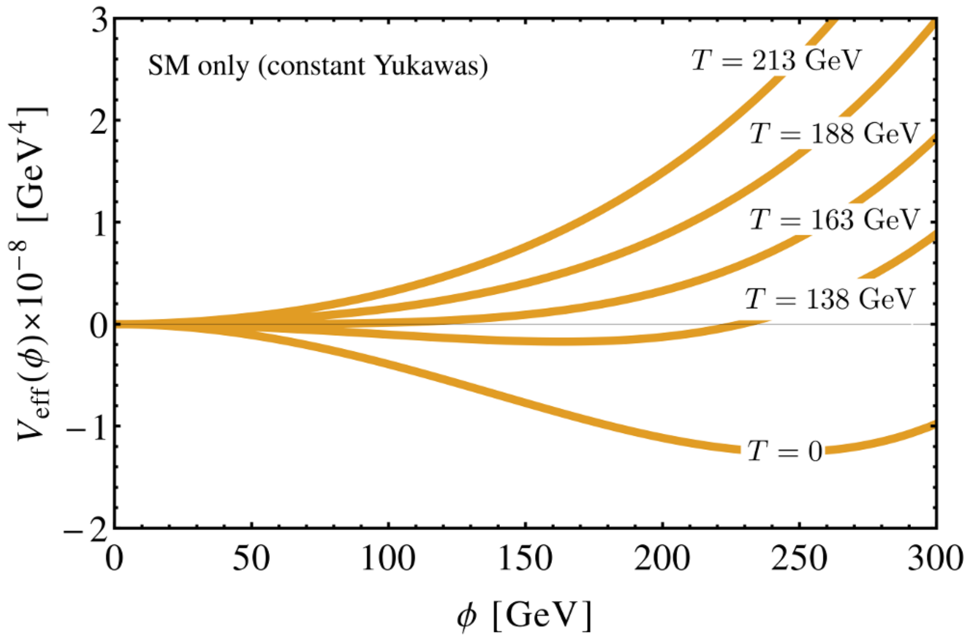
traditionally, M >> v and χ is frozen during EWSB

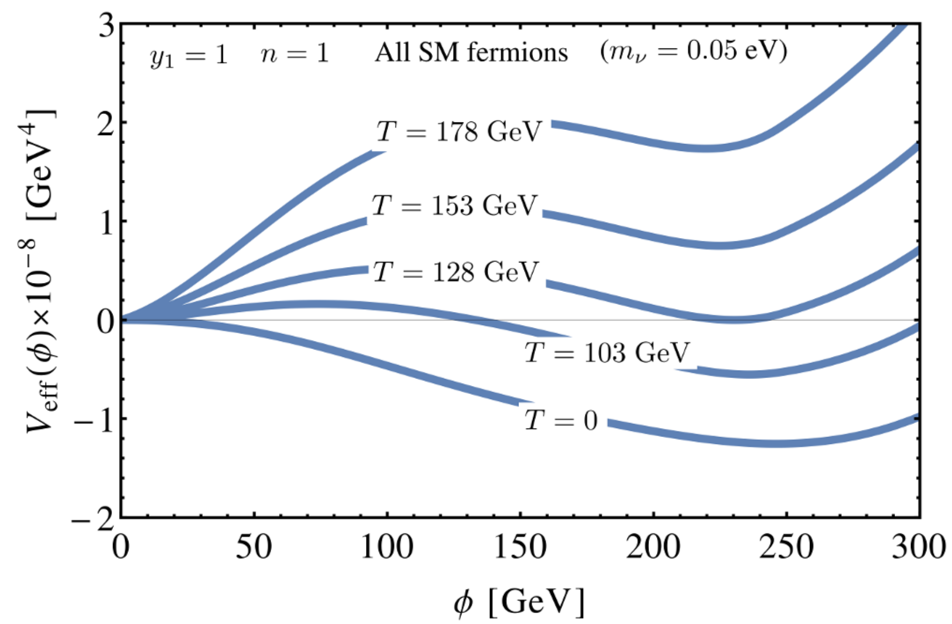
lowering M and allowing χ to vary leads to totally different phenomenology

EW scale flavons for EW baryogenesis

Baldes, Konstandin, Servant '16

$$y(\phi) = \begin{cases} y_1 \left(1 - \left[\frac{\phi}{v}\right]^n\right) + y_0 & \text{for} \quad \phi \le v, \\ y_0 & \text{for} \quad \phi \ge v. \end{cases}$$





The evolution of the effective potential with temperature in the SM (left) and with varying Yukawas (right) The varying Yukawa calculation includes all SM fermions with y1=1, n=1 and their respective y0, chosen to return the observed fermion masses today (the neutrinos are assumed to have a Dirac m=0.05eV).

In the varying Yukawa case, there is a first-order phase transition with ϕ_c=230GeV and Tc=128GeV (vs. second order transition at Tc=163GeV for the constant Yukawa case).

1st order phase transition + enhanced source of CP

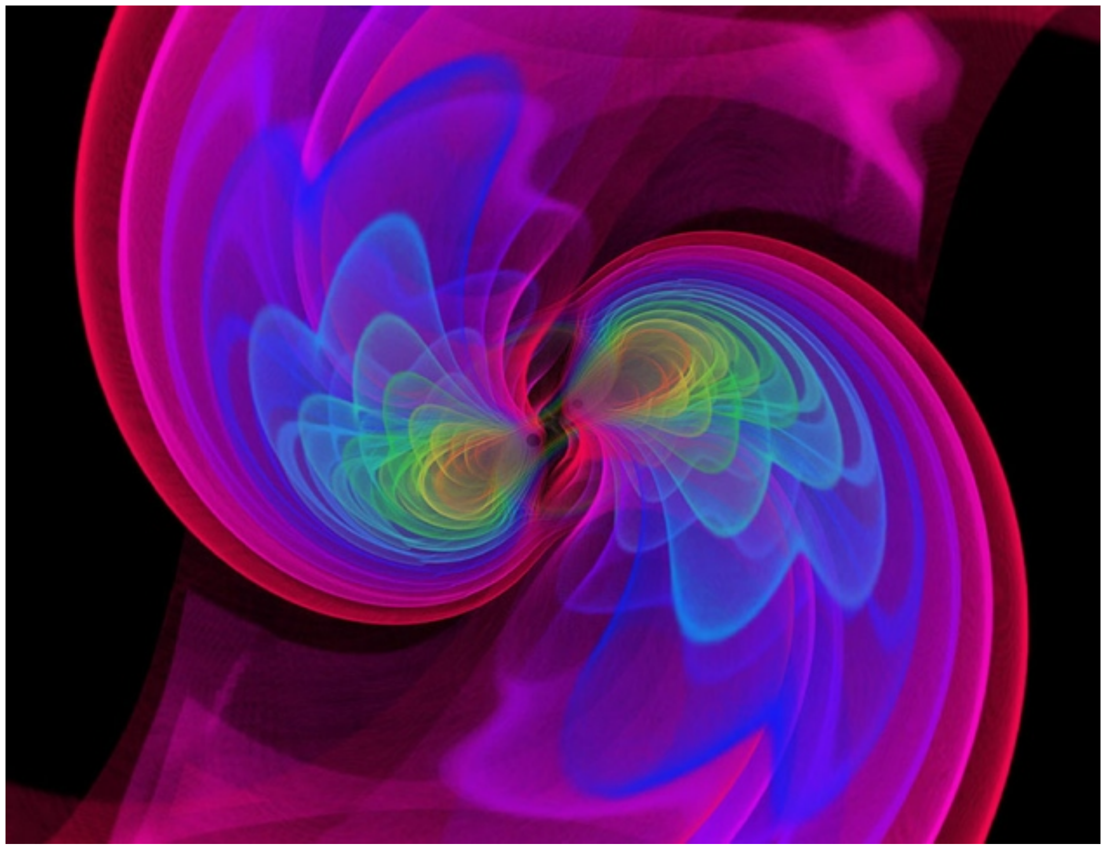
See also nice talk by D. Egana Ugrinovic yesterday

4

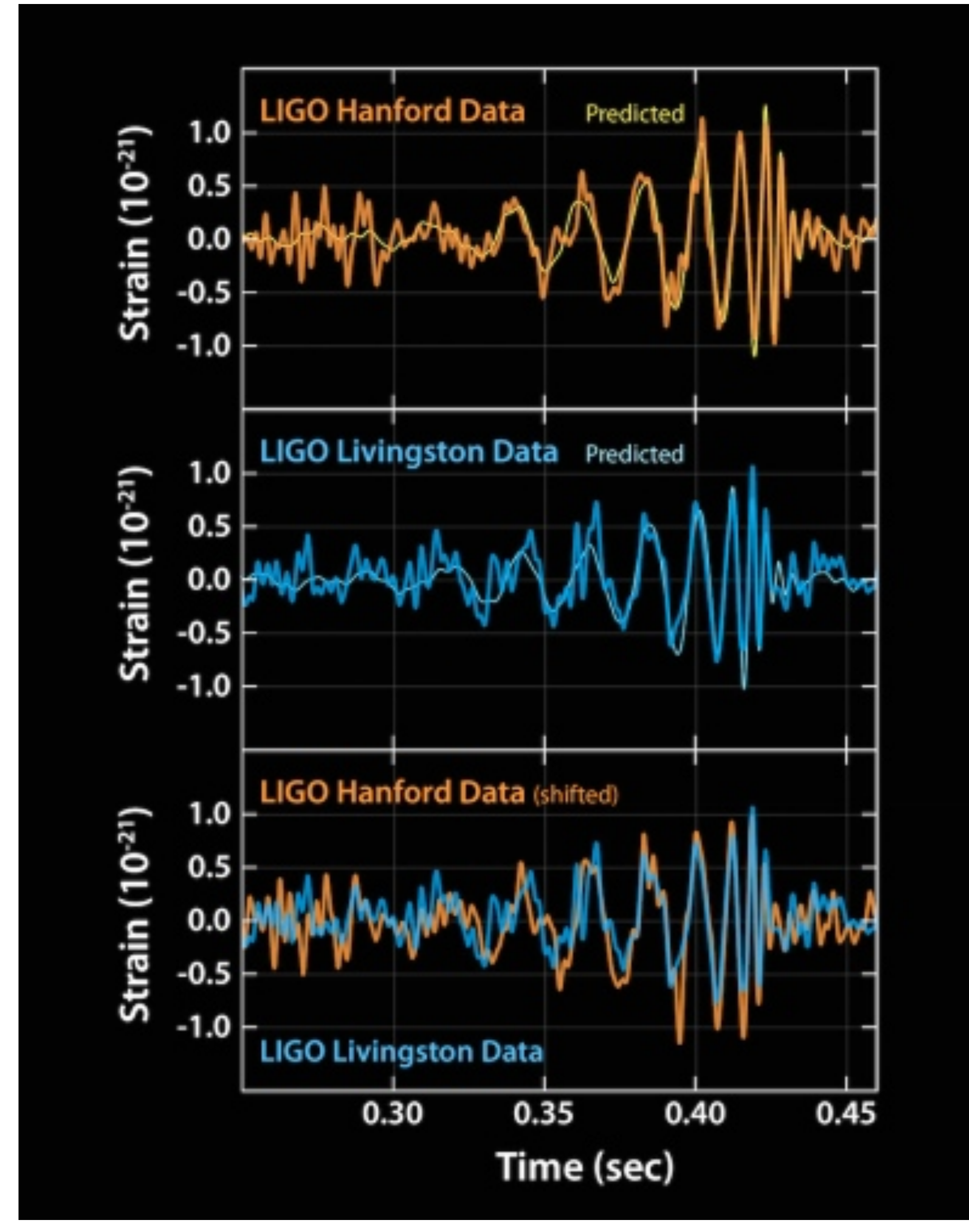
Fun with GW: stochastic GW background from phase transitions

The picture of 2016 (many more to come)

GW150914



1.3 billion years later on earth

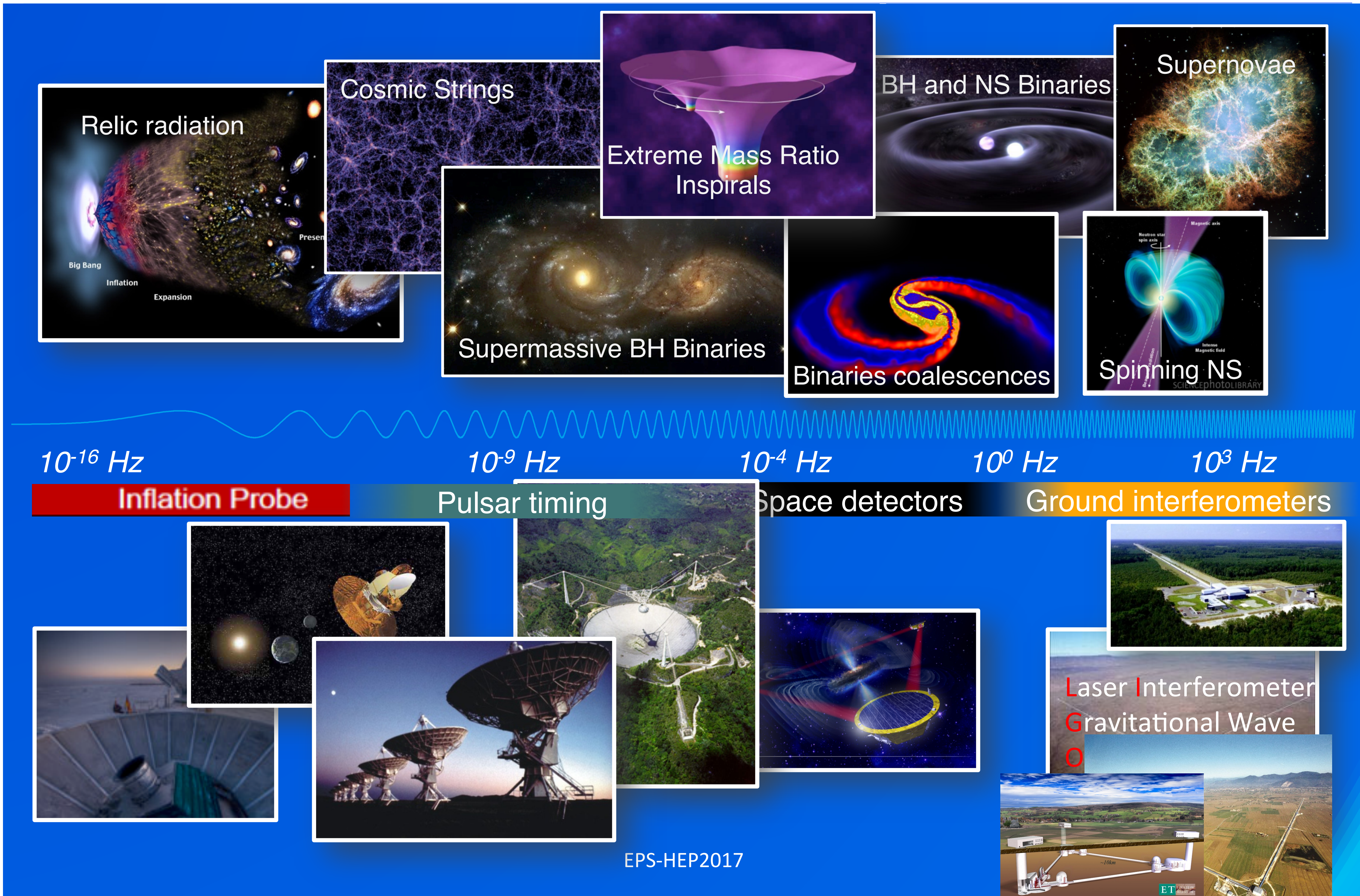


what did it teach us?

- never give up against strong background when you know you are right
- $m_g < 10^{-22}$ eV ($c_g - c_\gamma < 10^{-17}$ GRB observed together with GW with the same origin?)
- no spectral distortions: scale of quantum gravity > 100 keV

See beautiful BSA lecture by N. Mavalvala yesterday

GW and astrophysics/cosmology



GW and the ElectroWeak Phase Transition

GW interact very weakly and are not absorbed

⇓ ⇓

direct probe of physical process of the very early universe

possible cosmological sources:

inflation, vibrations of topological defects, excitations of xdim modes, 1st order phase transitions...

ElectroWeak Phase Transition (if 1st order)

typical freq. ~ (size of the bubble)$^{-1}$ ~ (fraction of the horizon size)$^{-1}$

$$@\, T = 100\ \text{GeV}, \qquad H = \sqrt{\frac{8\pi^3}{45}\frac{T^2}{M_{Pl}}} \sim 10^{-15}\ \text{GeV}$$

redshifted ⇓ freq.

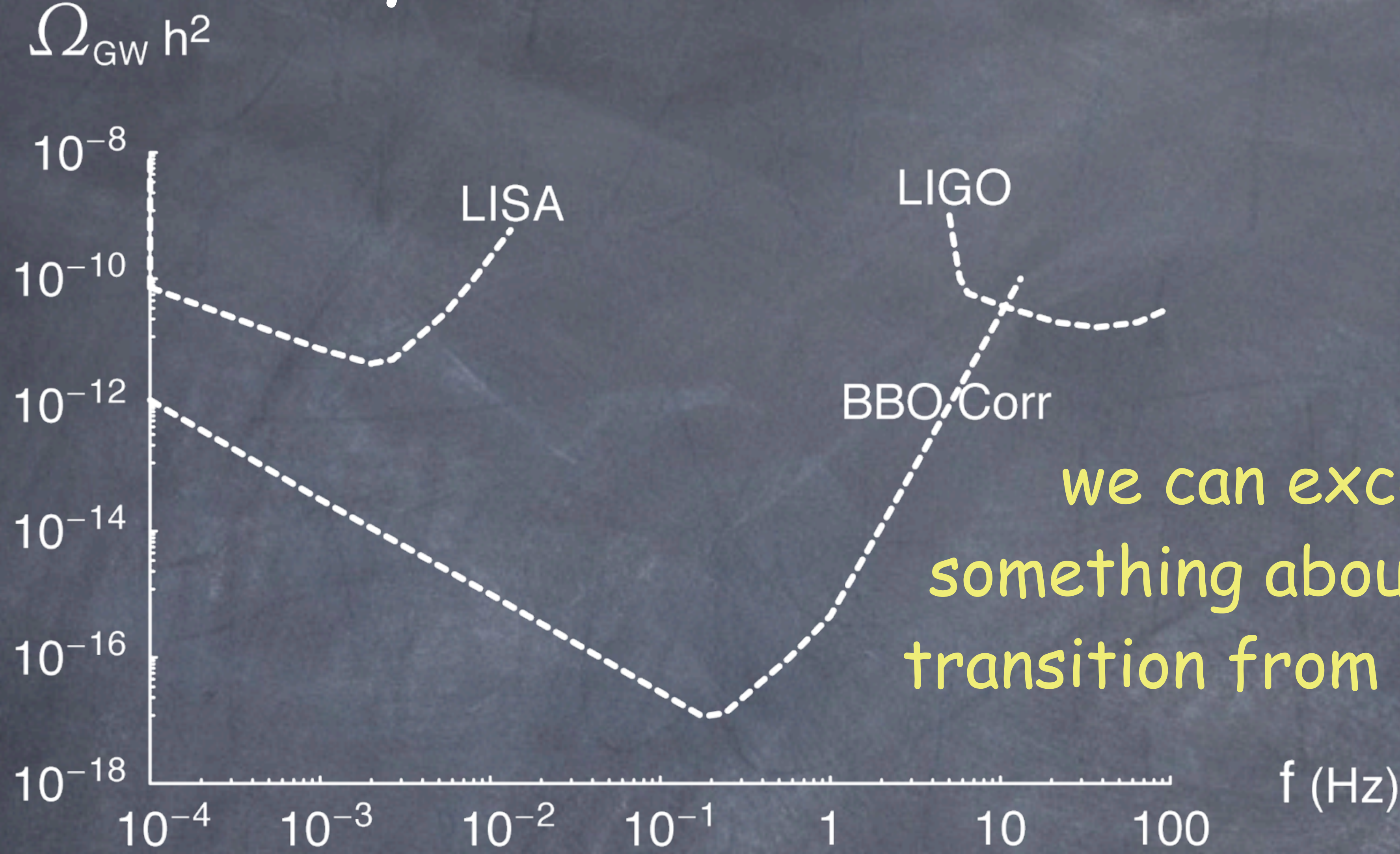
~ today ~

$$f \sim \# \frac{2\cdot 10^{-4}\ \text{eV}}{100\ \text{GeV}}\, 10^{-15}\ \text{GeV} \sim \#\, 10^{-5}\ \text{Hz}$$

Grojean, Servant '06

The GW spectrum from a 1st order electroweak PT
is peaked around the milliHertz frequency

Why should you be excited about mHZ freq.?



we can except to learn something about the EW phase transition from GW experiments

- test of the dynamics of the phase transition
  (quite important to analyze models of EW baryogenesis!)

redshift

$$\Omega^{\star}_{GW} \Rightarrow \Omega_{GW} = \left(\frac{a_\star}{a_0}\right)^4 \left(\frac{H_\star}{H_0}\right)^2 \Omega^{\star}_{GW} \sim 2{\cdot}10^{-5} h^{-2} \left(\frac{100}{g_\star}\right)^{1/3} \Omega^{\star}_{GW}$$

$$H_0 \sim h \times 2{\cdot}10^{-42}\ \text{GeV}$$

Hunting for phase transitions with GW

P. Schwaller '15

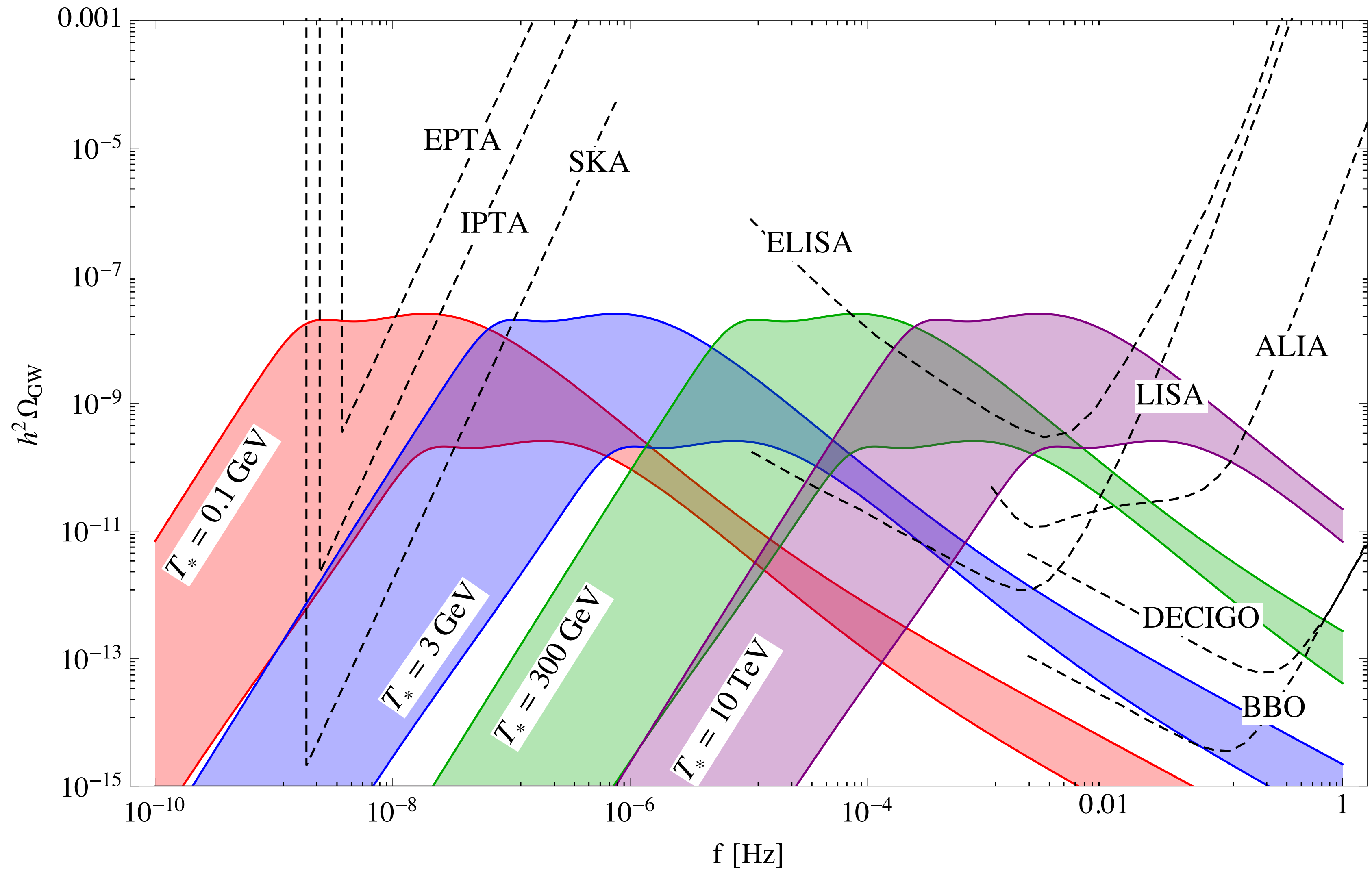


Figure 3: GW spectra $\Omega(f)h^2$ for $T_* = 0.1$ GeV (SIMP), $T_* = 3$ GeV (CDM1, TH models), $T_* = 300$ GeV and $T_* = 10$ TeV (CDM2 models). The upper (lower) edges of the contours correspond to $\beta = \mathcal{H}$ ($\beta = 10\mathcal{H}$), and furthermore $v = 1$ and $\Omega_{S*} = 0.1$ for all curves. The red band $T_* = 0.1$ GeV indicates where a signal of the QCD PT would lie if it was strong. The projected reach of several planned GW detection experiments is shown (dashed).

5

Naturalness without TeV-scale New Physics: relax!

Naturalness principle @ work

Following the arguments of Wilson, ‘t Hooft (and others):
only small numbers associated to **breaking of a symmetry** survive quantum corrections
(others are not necessarily theoretically inconsistent
but they require some conspiracy at different scales)

Beautiful examples of naturalness to understand the need of “new” physics

see for instance Giudice ’13 (and refs. therein) for a recent account

- the need of the positron to screen the **electron** self-energy: $\Lambda < m_e/\alpha_{\rm em}$
- the rho meson to cutoff the EM contribution to the **charged pion** mass: $\Lambda^2 < \delta m_\pi^2/\alpha_{\rm em}$
- the **kaon mass** difference regulated by the charm quark: $\Lambda^2 < \frac{\delta m_K}{m_K} \frac{6\pi^2}{G_F^2 f_K^2 \sin^2\theta_C}$
- the **light Higgs** boson to screen the EW corrections to gauge bosons self-energies
- ...
- **new physics** at the weak scale to cancel the UV sensitivity of the Higgs mass?

The Darwinian solution to the Hierarchy

Other origin of small/large numbers according to Weyl and Dirac:
hierarchies are induced/created by time evolution/the age of the Universe

Can this idea be formulated in a QFT language?
In which sense is it addressing the stability of small numbers at the quantum level?

Graham, Kaplan, Rajendran '15

Espinosa et al '15

- Higgs mass-squared promoted to a field
- The field evolves in time in the early universe and scans a vast range of Higgs mass
- The Higgs mass-squared relaxes to a small negative value
- The electroweak symmetry breaking stops the time-evolution of the dynamical system

Self-organized criticality

dynamical evolution of a system is stopped at a critical point due to back-reaction

hierarchies result from **dynamics** not from **symmetries** anymore!

important consequences on the spectrum of new physics

Higgs-axion cosmological relaxation

Graham, Kaplan, Rajendran '15

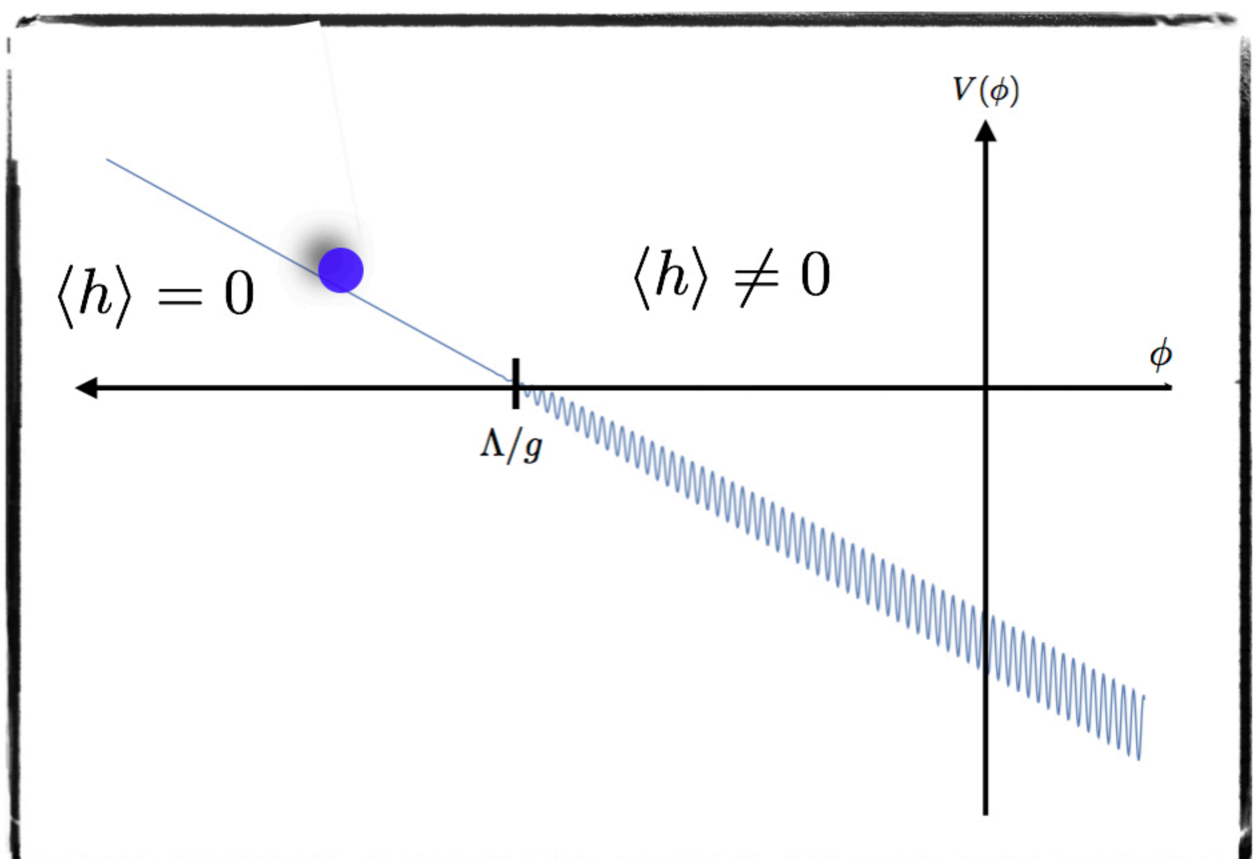
ϕ slowly rolling field (inflation provides friction) that scans the Higgs mass

$$\Lambda^2\left(-1+f\left(\frac{g\phi}{\Lambda}\right)\right)|H|^2+\Lambda^4 V\left(\frac{g\phi}{\Lambda}\right)+\frac{1}{32\pi^2}\frac{\phi}{f}\tilde{G}^{\mu\nu}G_{\mu\nu}$$

Higgs mass depends on ϕ

potential needed to force ϕ to roll-down in time (during inflation)

axion-like coupling that will seed the potential barrier stopping the rolling when the Higgs develops its vev



$$\Lambda_{\rm QCD}^3\, h \cos\frac{\phi}{f}$$

If ϕ continues rolling, the Higgs vev increases, the potential barrier increases and ultimately prevents ϕ from rolling down further

Higgs-axion cosmological relaxation

Graham, Kaplan, Rajendran '15

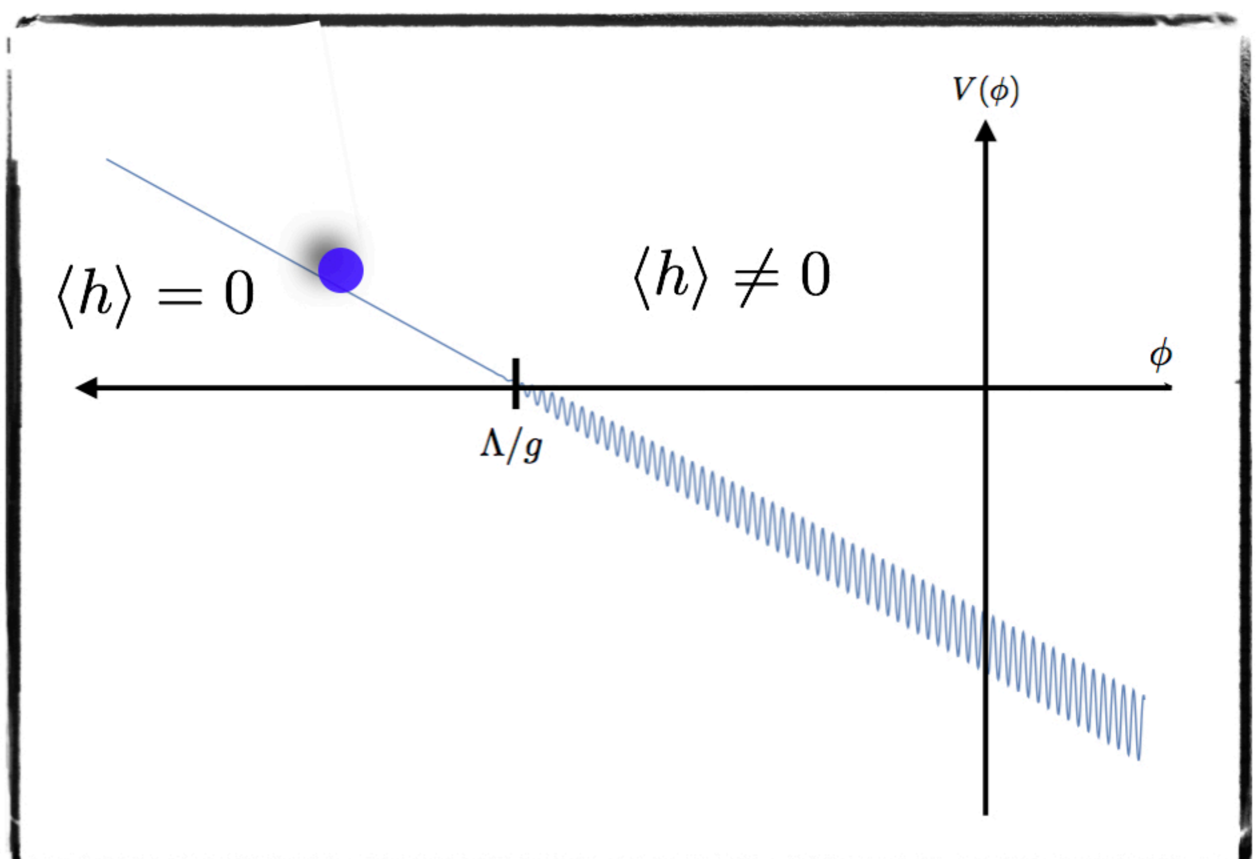
ϕ slowly rolling field (inflation provides frictio… s

$$\Lambda^2\left(-1+f\left(\frac{g\phi}{\Lambda}\right)\right)|H|^2+\Lambda^4 V\left(\frac{g\ldots}{\Lambda}\right.\ \ldots\ \nu$$

Higgs mass depends on ϕ

potential needed to force ϕ to roll-down in time (during inflation)

that … opping the … its vev



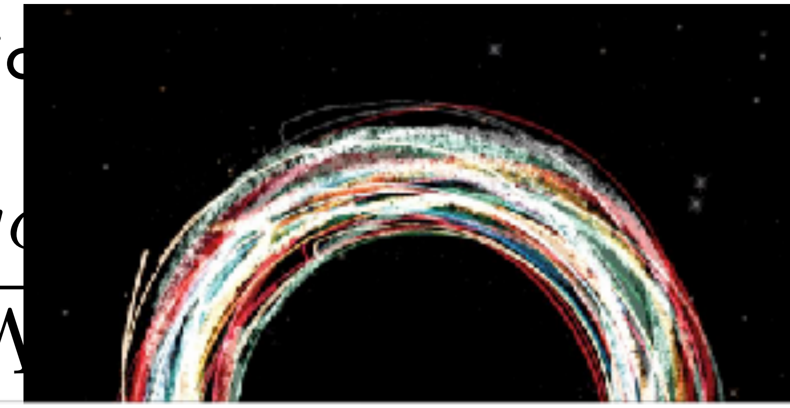
If ϕ continues rolling, the Higgs vev increases, the potential barrier increases and ultimately prevents ϕ from rolling down further

Higgs-axion cosmological relaxation

Graham, Kaplan, Rajendran '15

ϕ slowly rolling field (inflation provides frictio… s

$$\Lambda^2\left(-1+f\left(\frac{g\phi}{\Lambda}\right)\right)|H|^2+\Lambda^4 V\left(\frac{g\phi}{\Lambda}\right. \ldots$$



Hierarchy problem solved
by light weakly coupled new physics
and not by TeV scale physics

~interesting atomic physics~

◎ change of atom sizes

Espinosa et al '15 Flacke et al '16

~interesting cosmology signatures~

◎ BBN constraints
◎ decaying DM signs in γ-rays background
◎ ALPs
◎ superradiance

Choi and Im '16

~interesting signatures @ SHiP~

◎ production of light scalars
by B and K decays

Phenomenological signatures

Nothing to be discovered at the LHC/ILC/CLIC/CepC/SppC/FCC!

only BSM physics below Λ

two (very) light and very weakly coupled axion-like scalar fields

$$m_\phi \sim (10^{-20} - 10^{2})\,\mathrm{GeV}$$

$$m_\sigma \sim (10^{-45} - 10^{-2})\,\mathrm{GeV}$$

interesting signatures in cosmology

Conclusions

Conclusions: Executive Summary

The LHC leaves us with the deepest mathematical pb:

$$\infty \cdot 0 = ?$$

Dissertori, ECFA ’13

- ∞: number of already performed BSM searches
- 0: number of significant/ interesting/exciting deviations from SM predictions
- $?$: general state of (our) mind (?)

Conclusions: Executive Summary

The LHC leaves us with the deepest mathematical pb:

$$\infty \cdot 0 = ?$$

Dissertori, ECFA '13

solitary Higgs boson with NO new physics at TeV scale
challenges our understanding of the quantum world
and forces a paradigm shift

The Higgs boson is the Santa Maria of the 21st century:
understanding the scalar sector of the SM
will help us grasping what lays beyond the SM

We also need the right **technological tools** (HL-LHC, ILC, CLIC, CepC, FCC...)
to continue exploring the unknown

Conclusions: Executive Summary

The LHC leaves us with the deepest mathematical pb:

$$\infty \cdot 0 = ?$$

Dissertori, ECFA '13

" A ship is always safe at the shore
but that is not what it is built for "

A. Einstein

The Higgs boson is the Santa Maria of the 21st century:
understanding the scalar sector of the SM
will help us grasping what lays beyond the SM

We also need the right **technological tools** (HL-LHC, ILC, CLIC, CepC, FCC...)
to continue exploring the unknown