

# Dark Sector Searches with *BABAR*

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Caltech  
for the *BABAR* Collaboration

Brookhaven Forum  
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# Dark Sector Searches with *BABAR* and LDMX

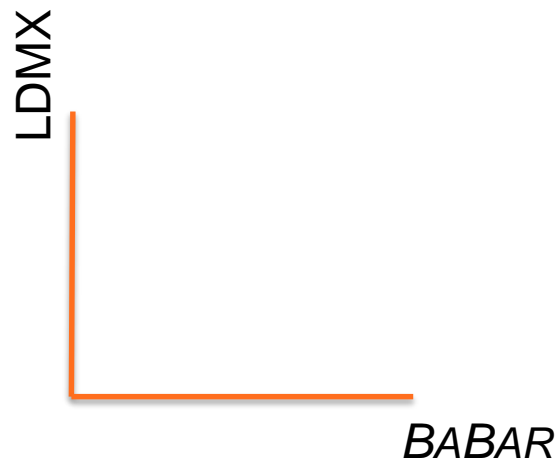
Since I will be giving two consecutive talks, the first on

- Dark Sector Searches with *BABAR*

and the second on

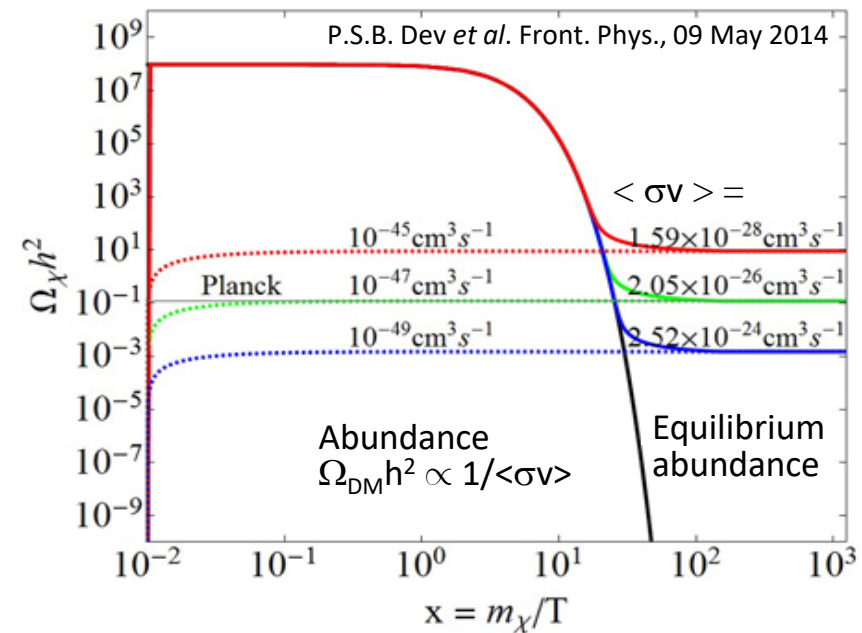
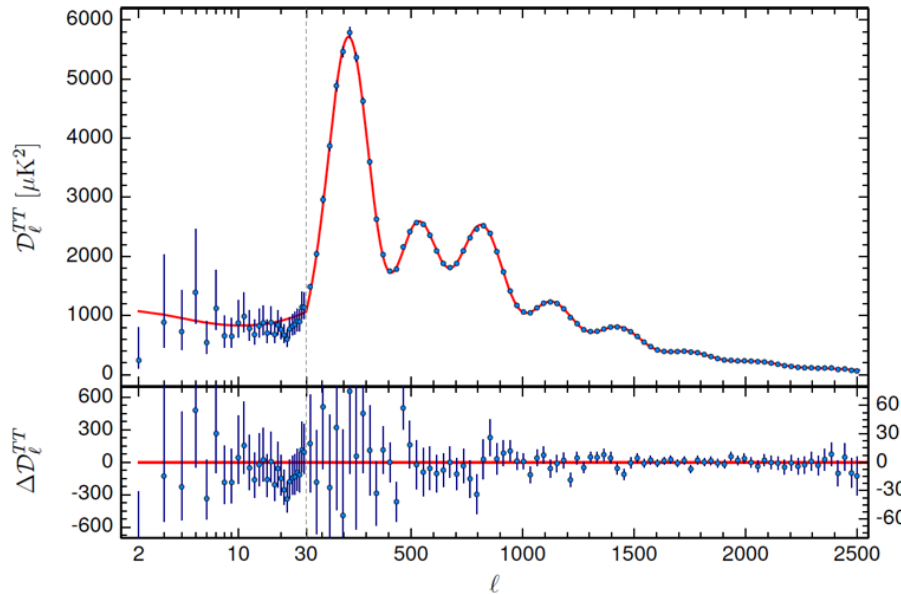
- Dark Sector searches with LDMX,

I will, for the clarity of the exposition, treat the two talks as a continuum



# Dark matter - the Thermal Relic Target

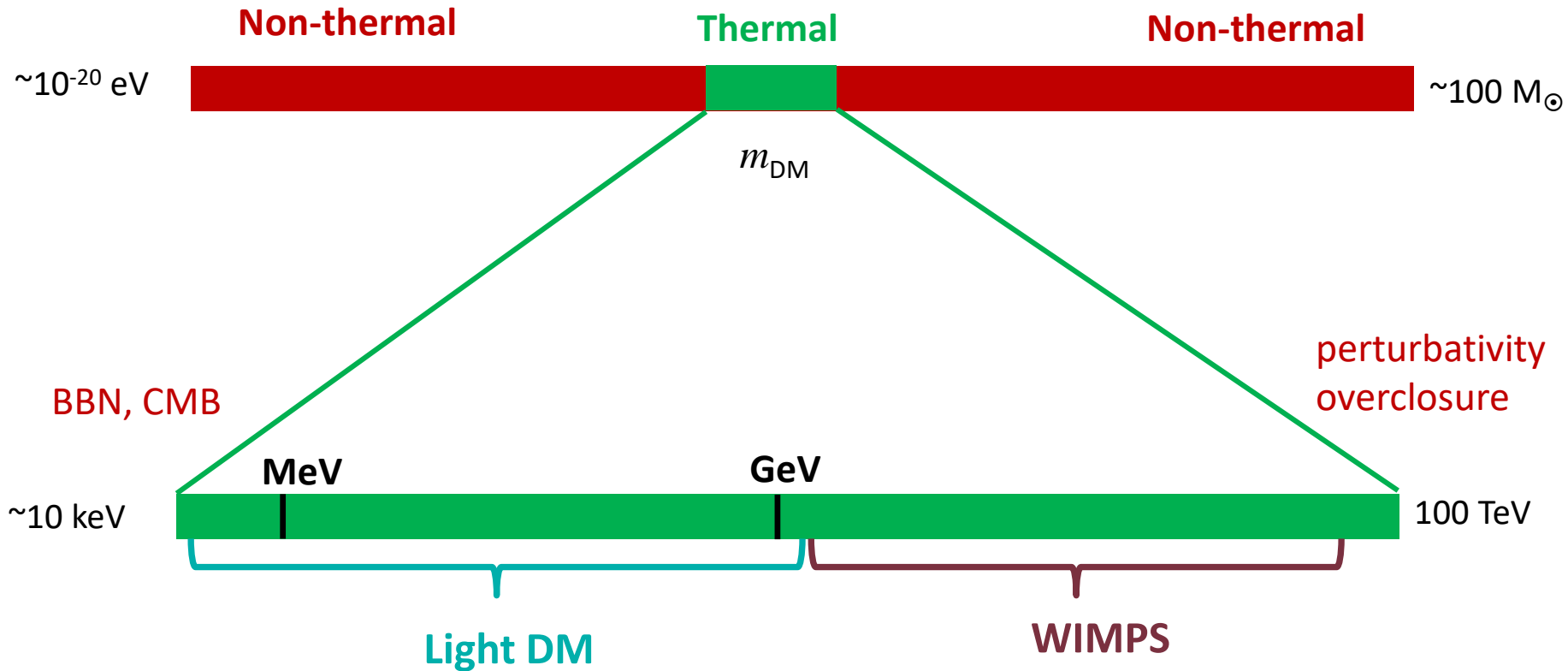
- Much effort has been expended to search for a particle physics connection to dark matter



- The cosmic microwave background establishes a matter-density target for dark matter
- If dark matter is produced thermally, the observed abundance sets a requirement for the coupling and particle mass

# Search range for thermal dark matter

The thermal relic hypothesis restricts the allowed range of DM masses

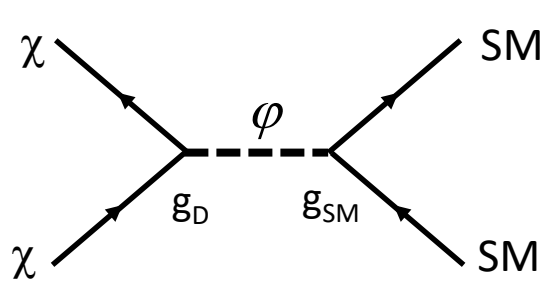


Thermal contact implies new mediator  
Hidden sector light DM well-motivated model

Thermal freeze-out for weak scale masses  
has driven DM searches for last  $\sim 30$  years

# Light thermal dark matter

A freeze-out scenario with light dark matter ( $\chi$ ) requires new light mediator to explain the relic density in order to avoid overproduction of the dark matter

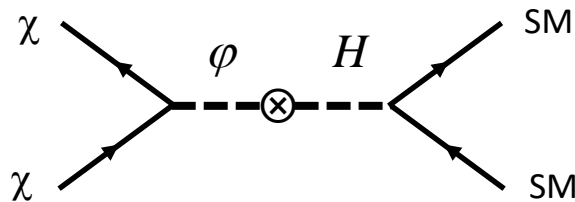


$$\langle \sigma v \rangle_{\text{relic}} \sim \frac{g_D^2 g_{SM}^2 m_\chi^2}{m_\phi^4} \quad (m_\phi \gg m_\chi)$$

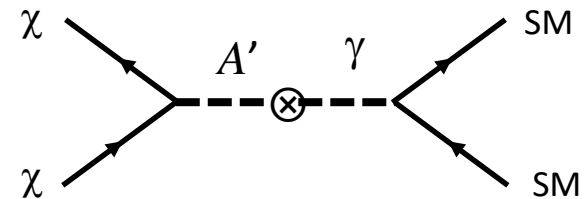
$$m_\phi^4 \sim \frac{g_D^2 g_{SM}^2 m_\chi^2}{\langle \sigma v \rangle} \leq \frac{m_\chi^2}{\langle \sigma v \rangle} \quad \text{since } g \leq O(1)$$

The mediator must be neutral under the SM and renormalizable. Simplest choices:

New scalar ( $\phi$ ) with Higgs coupling



New vector ( $A'$ ) with photon coupling



These are naturally realized in the context of hidden sectors

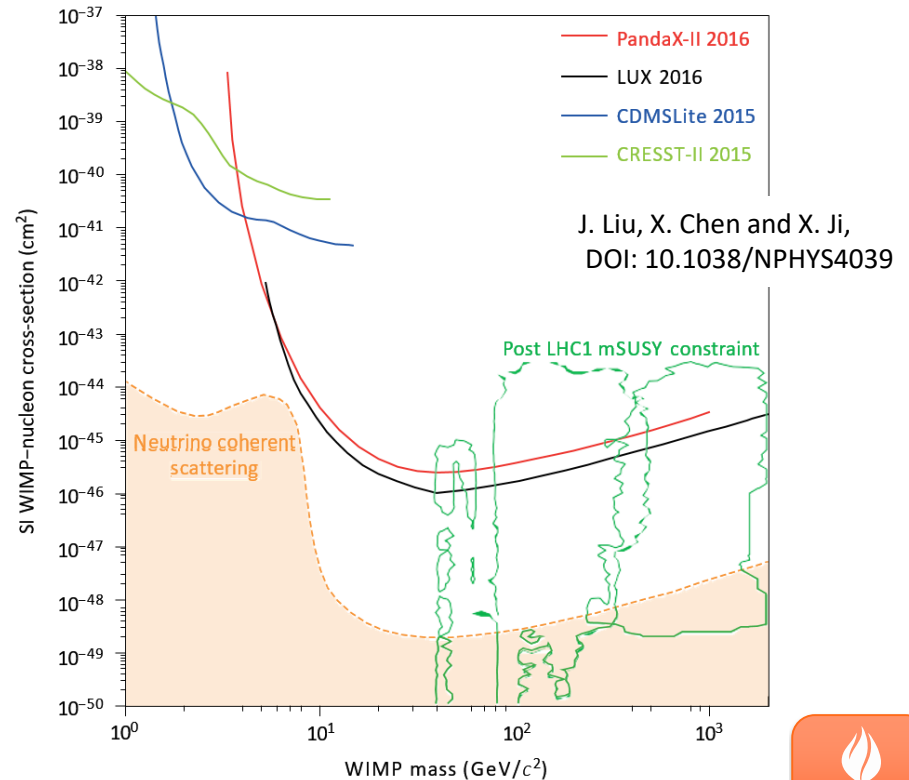
# WIMPs, Hidden Sectors, .....

- **WIMPs** provides an attractive realization of the thermal relic scenario
  - Tightly tied to EWSB
  - Identification with the LSP
  - large viable mass range
- However, direct dark matter and LHC searches now highly constrain this scenario

ATLAS SUSY Searches\* - 95% CL Lower Limits  
May 2017

Model	$e, \mu, \tau, \gamma$	Jets	$E_T^{\text{miss}}$	$\int \mathcal{L} d\Omega (h^{-1})$	Mass limit	
MSUGRA/CMSSM	0-3 $e, \mu$ / 1-2 $\tau$	2-10 jets/3 $b$	Yes	20.3	4.2	$m(\tilde{g})=m(\tilde{t}_1)$
$\tilde{g}\tilde{g}, \tilde{q}\tilde{q} \rightarrow q\bar{q}$	0	2-6 jets	Yes	36.1	4	$m(\tilde{t}_1) > 200 \text{ GeV}, m(\tilde{1}^{\text{st}} \text{ gen. } \tilde{g})=m(\tilde{2}^{\text{nd}} \text{ gen. } \tilde{g})$
$\tilde{g}\tilde{g}, \tilde{q}\tilde{q} \rightarrow q\bar{q}$ (compressed)	0	1-3 jets	Yes	3.2	4	$m(\tilde{g})=m(\tilde{t}_1) < 5 \text{ GeV}$
$\tilde{g}\tilde{g}, \tilde{q}\tilde{q} \rightarrow q\bar{q}$	0	2-6 jets	Yes	36.1	4	$m(\tilde{t}_1) > 200 \text{ GeV}$
$\tilde{g}\tilde{g}, \tilde{q}\tilde{q} \rightarrow q\bar{q}, \tilde{g}\tilde{g} \rightarrow \tilde{g}\tilde{g}$	0	2-6 jets	Yes	36.1	4	$m(\tilde{t}_1) > 200 \text{ GeV}, m(\tilde{t}_1) > 0.5(m(\tilde{t}_1)+m(\tilde{g}))$
$\tilde{g}\tilde{g}, \tilde{q}\tilde{q} \rightarrow q\bar{q}, \tilde{g}\tilde{g} \rightarrow \tilde{g}\tilde{g}$	3 $e, \mu$	4 jets	-	36.1	4	$m(\tilde{t}_1) < 400 \text{ GeV}$
$\tilde{g}\tilde{g}, \tilde{q}\tilde{q} \rightarrow WZ\gamma$	0	7-11 jets	Yes	36.1	4	$m(\tilde{t}_1) < 400 \text{ GeV}$
GMSB ( $\tilde{L}$ NLSP)	1-2 $\tau + 0-1 \ell$	0-2 jets	Yes	3.2	4	$m(\tilde{t}_1) < 200 \text{ GeV}$
GGM (bino NLSP)	2 $\gamma$	-	Yes	3.2	4	$c\tau(\text{NLSP}) < 0.1 \text{ mm}$
GGM (Higgsino-bino NLSP)	7	1 $b$	Yes	20.3	4	$m(\tilde{t}_1) < 350 \text{ GeV}, c\tau(\text{NLSP}) < 0.1 \text{ mm}, \mu < 0$
GGM (Higgsino-bino NLSP)	7	2 jets	Yes	13.3	4	$m(\tilde{t}_1) < 350 \text{ GeV}, c\tau(\text{NLSP}) < 0.1 \text{ mm}, \mu < 0$
GGM (Higgsino NLSP)	2 $e, \mu$ ( $Z$ )	2 jets	Yes	20.3	4	$m(\text{NLSP}) > 430 \text{ GeV}$
Gravitino LSP	0	mono-jet	Yes	20.3	4	$m(\tilde{g}) > 1.8 \times 10^{-4} \text{ eV}, m(\tilde{g})=m(\tilde{g}) > 1.5 \text{ TeV}$

- The lighter mass range is harder to access
  - Motivation for pushing direct searches to lower masses
  - Many new ideas in play
  - One of these approaches – **hidden sectors** - naturally accommodates lower masses
  - Hidden sectors are best studied in accelerator experiments
  - A combination of collider and fixed target accelerator experiments can fully probe directly annihilating thermal dark matter in the MeV to GeV range



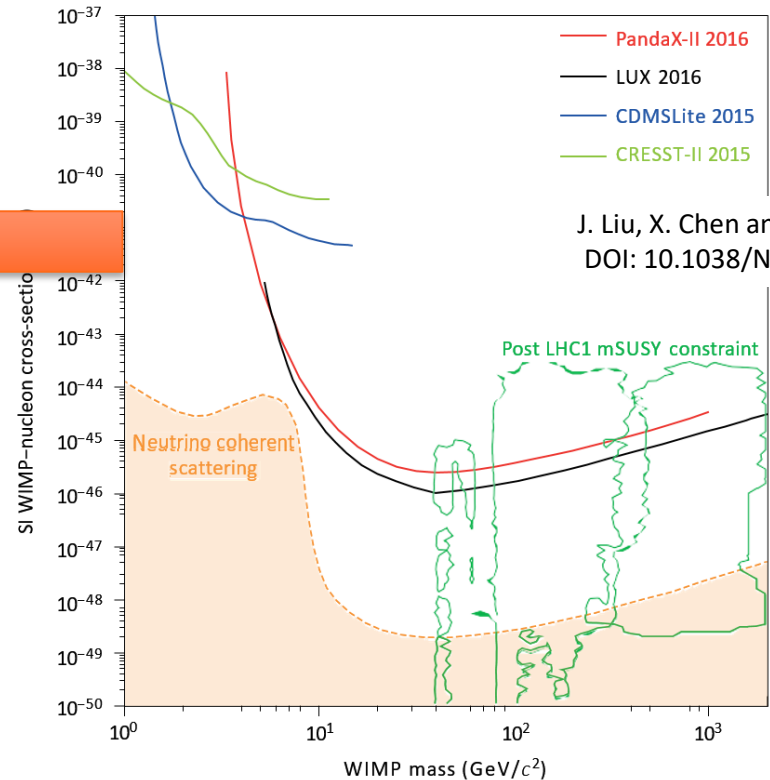
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$\tilde{g}\tilde{g}, \tilde{g}\rightarrow\tilde{g}^2$	0	2-6 jets	Yes	36.1	4	$m(\tilde{t}^1)=200 \text{ GeV}, m(\tilde{t}^2)=m(\tilde{b}^2)$ gen. 0
$\tilde{g}\tilde{g}, \tilde{g}\rightarrow\tilde{g}^2$ (compressed)	0	mono-jet	Yes	3.2	4	608 GeV
$\tilde{g}\tilde{g}, \tilde{g}\rightarrow\tilde{g}^2$	0	2-6 jets	Yes	36.1	4	$m(\tilde{g})=m(\tilde{t}^1)=5 \text{ GeV}$
$\tilde{g}\tilde{g}, \tilde{g}\rightarrow\tilde{g}^2$	0	2-6 jets	Yes	36.1	4	$m(\tilde{t}^1)=200 \text{ GeV}$
$\tilde{g}\tilde{g}, \tilde{g}\rightarrow\tilde{g}^2$	0	2-6 jets	Yes	36.1	4	$m(\tilde{t}^1)=200 \text{ GeV}, m(\tilde{t}^2)=0.5(m(\tilde{t}^1)+m(\tilde{g}))$
$\tilde{g}\tilde{g}, \tilde{g}\rightarrow\tilde{g}^2$	3 $e, \mu$	4 jets	-	36.1	4	1.625 TeV
$\tilde{g}\tilde{g}, \tilde{g}\rightarrow\tilde{g}^2$	0	7-11 jets	Yes	36.1	4	1.8 TeV
$\tilde{g}\tilde{g}, \tilde{g}\rightarrow\tilde{g}^2$	1-2 $\tau + 0-1 \ell$	0-2 jets	Yes	3.2	4	2.0 TeV
GGM (bino NLSP)	2 $\gamma$	-	Yes	3.2	4	1.55 TeV
GGM (Higgsino-bino NLSP)	7	1 $b$	Yes	20.3	4	1.37 TeV
GGM (Higgsino-bino NLSP)	7	2 jets	Yes	13.3	4	1.3 TeV
GGM (Higgsino NLSP)	2 $e, \mu$ ( $Z$ )	2 jets	Yes	20.3	4	900 GeV
Gravitino LSP	0	mono-jet	Yes	20.3	4	665 GeV

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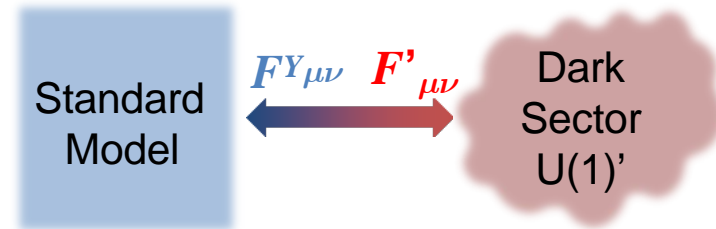


J. Liu, X. Chen and X. Ji,  
DOI: 10.1038/NPHYS4039



# Hidden sectors

- There are several viable new interactions involving a “hidden sector” that respect Standard Model symmetries and have dimensionless couplings, *i.e.* they are sizeable irrespective of their source
- Hidden sectors are generic in many BSM theories
- Dark matter could be part of a dark sector
- Let’s consider the so-called vector portal: a “**Dark Sector**” coupled to the SM via a low-mass spin 1 “dark photon” mediator, the gauge boson of a new U(1) symmetry
- These **dark photons**  $A'$  could be in the MeV to GeV mass range and mix with the SM photon with mixing strength  $\epsilon$
- The dark photon could decay to
  - SM fermions if other DM states are inaccessible, producing visible decays
  - a lighter **dark matter** state  $\chi$ 
    - If  $m_\chi < m_{A'}/2$ , then the dominant decay mode of the  $A'$  would then be invisible:  $A' \rightarrow \chi\bar{\chi}$
- A dark sector could potentially explain the proton charge radius puzzle and the muon  $g-2$  anomaly
- A dark sector is more general than light dark matter, but



heavy particle  $\psi$  with both dark and EM charges.

$\epsilon \sim 10^{-4} - 10^{-2}$

GUT (2 loops)

$\epsilon \sim 10^{-5} - 10^{-3}$

→  $10^{-7}$  if both U(1)'s are in unified groups

e.g. Arkani-Hamed & Weiner;  
Cheung, Ruderman, Wang, Yavin;  
Morrissey, Poland, Zurek;  
Essig, Schuster, Toro;

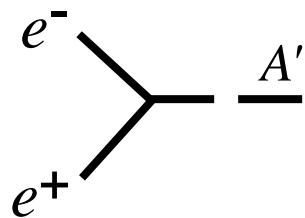




# Dark sector searches

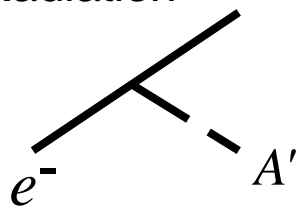
- Dark photons can be produced in accelerator experiments
  - Photons in any process can be replaced by dark photons (with an extra factor of  $\epsilon^2$ )

Annihilation

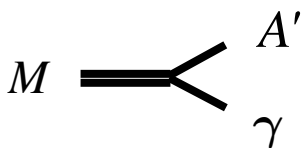


colliders

Radiation



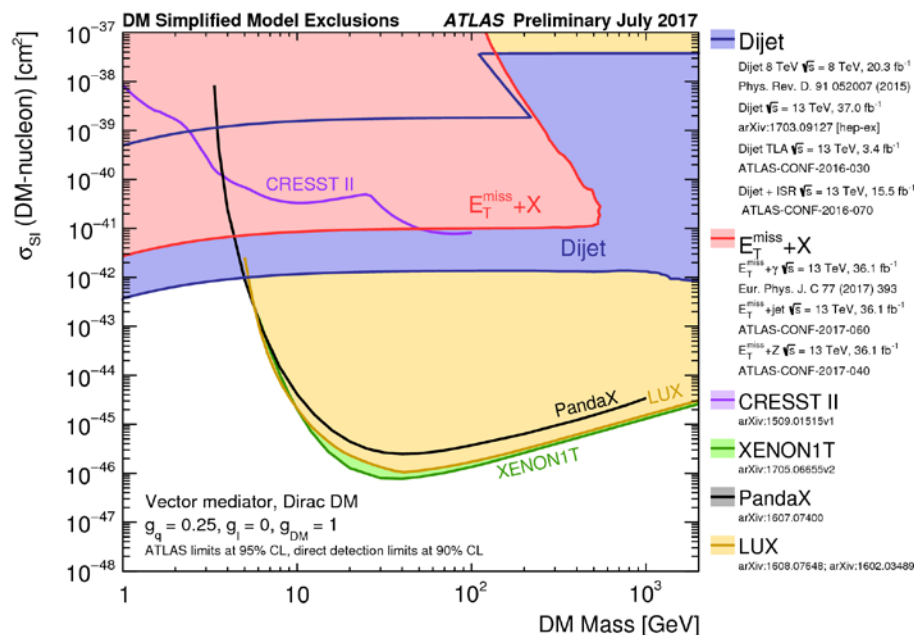
Meson decay



$p$  beam dump, hadron-rich environment

Require high intensity  
+ high energy for high mass  $A'$   
low energy for low mass  $A'$

- LHC can explore the high mass regime
  - Limits are not very constraining below the GeV regime



## ATLAS Exotics Searches\* - 95% CL Upper Exclusion I

Status: July 2017

Model	$\ell, \gamma$	Jets $^\dagger$	$E_T^{\text{miss}}$	$\int \mathcal{L} dt [\text{fb}^{-1}]$	Limit
DM	Axial-vector mediator (Dirac DM)	$0 e, \mu$	1-4 j	Yes	36.1
	Vector mediator (Dirac DM)	$0 e, \mu, 1 \gamma$	$\leq 1 j$	Yes	36.1
	VV $\chi\chi$ EFT (Dirac DM)	$0 e, \mu$	1 J, $\leq 1 j$	Yes	3.2

$\sqrt{s} = 8 \text{ TeV}$	$\sqrt{s} = 13 \text{ TeV}$
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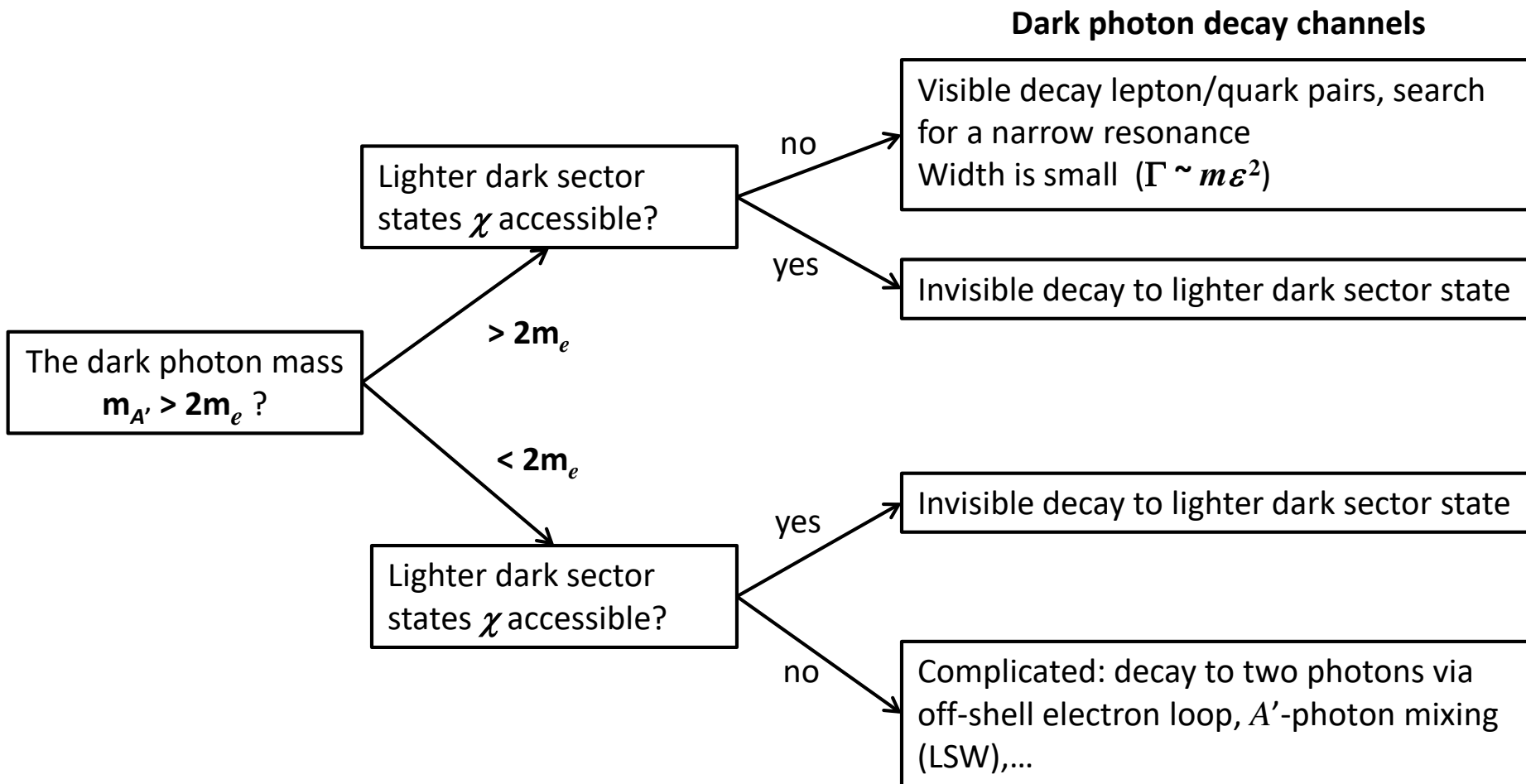
$m_{\text{med}} = 1.5 \text{ TeV}$	$m(\chi) < 400 \text{ GeV}$
$m_{\text{med}} = 1.2 \text{ TeV}$	$m(\chi) < 480 \text{ GeV}$
$M_s = 700 \text{ GeV}$	$m(\chi) < 150 \text{ GeV}$

Mass scale [TeV]



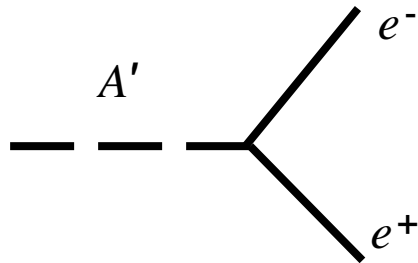
# The light dark sector at colliders

- A (simplified) roadmap



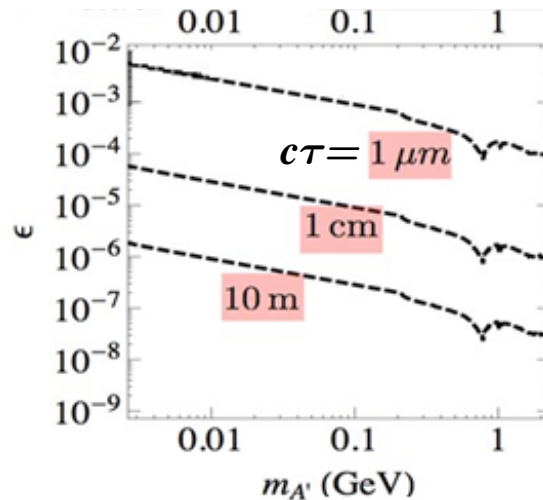
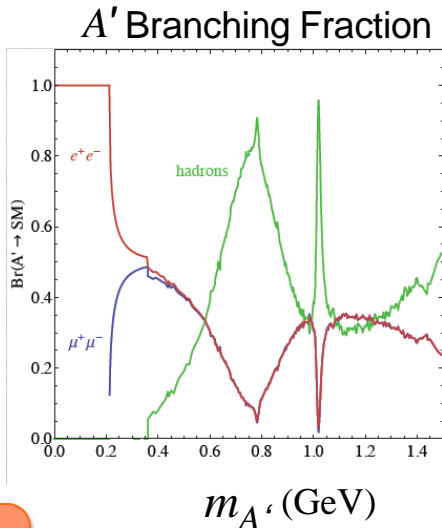
# MeV – GeV dark sector

## Lepton / quark decays

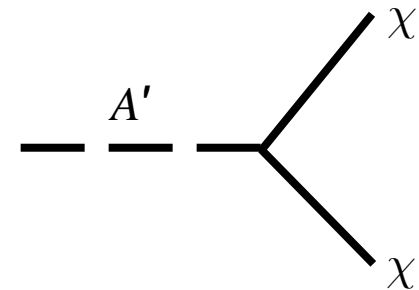


decay via kinetic mixing  
 $\rightarrow$  small width  $\Gamma \sim \alpha m \varepsilon^2$

Prompt or displaced decays



## Light dark sector states $\chi$



Not  $\varepsilon$  suppressed

Dominates if  $m_\chi < 2m_{A'}$

Two cases:

- $\chi$  is stable  $\Rightarrow$  invisible decays
- $\chi$  decays back to SM particles

Coupling is often characterized as

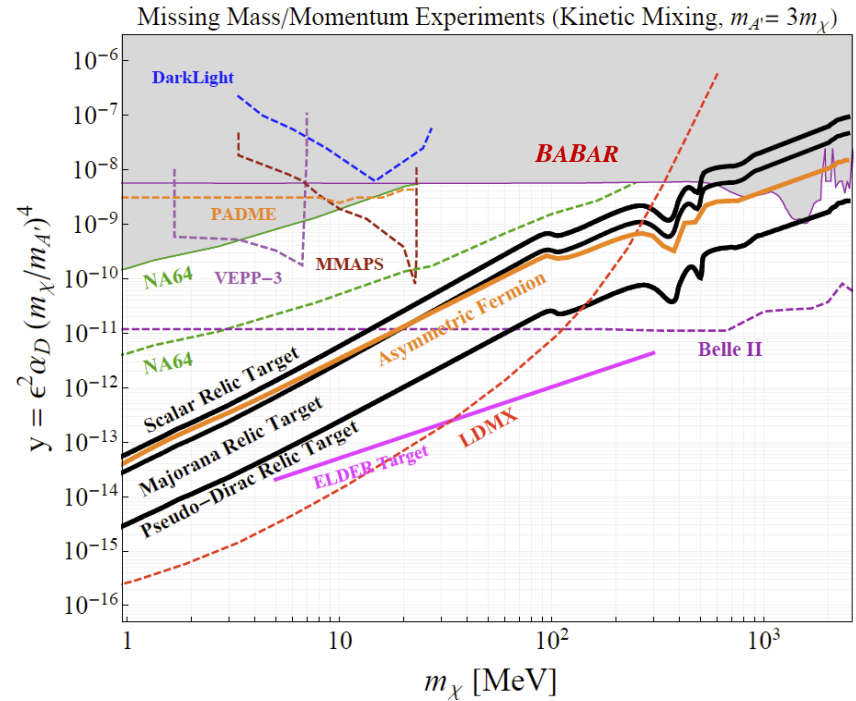
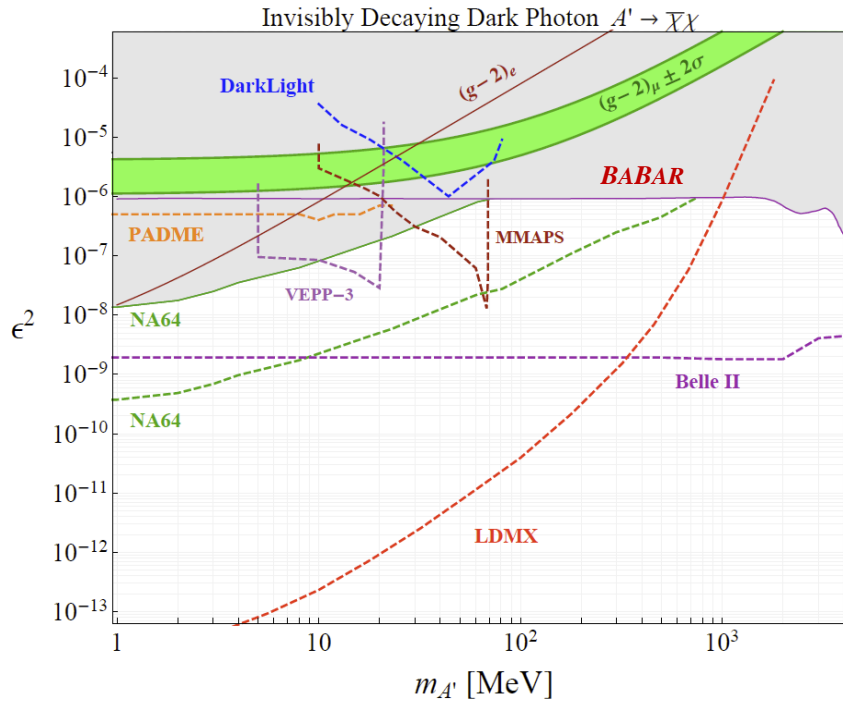
$$y = \varepsilon^2 \alpha_D \left( \frac{m_\chi}{m_{A'}} \right)^4$$



# New ideas abound – Cosmic Visions Workshop

## US Cosmic Visions: New Ideas in Dark Matter 2017 Community Report

arXiv:1707.04591v1 [hep-ph] 14 Jul 2017



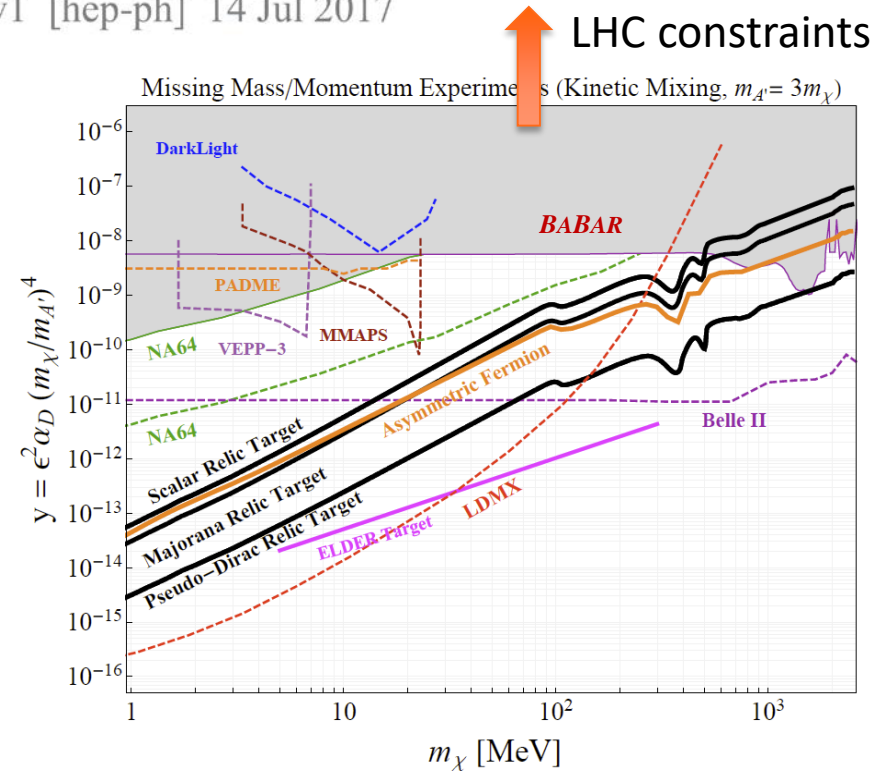
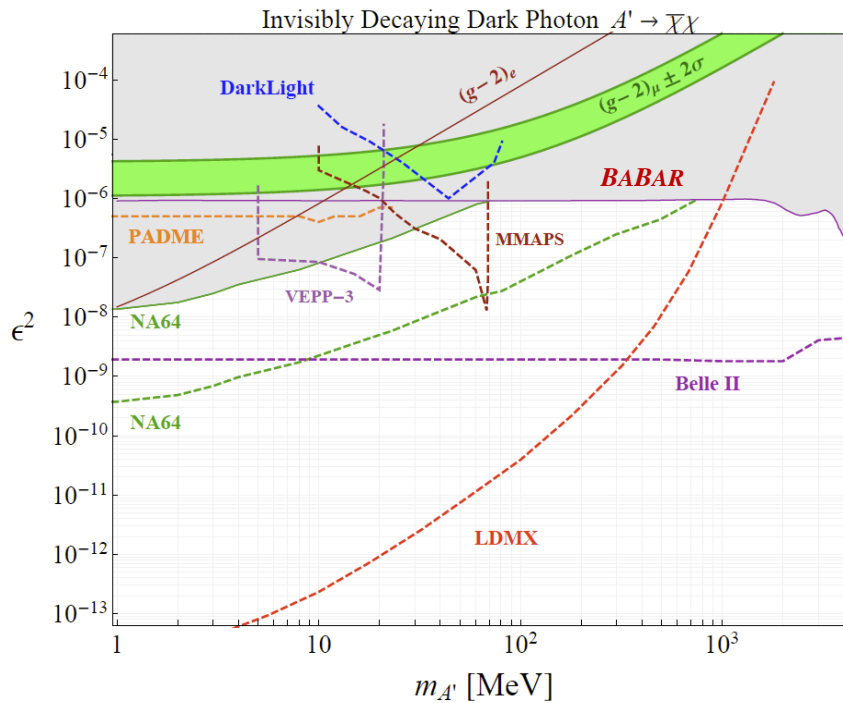
- The workshop considered WIMP, Hidden Sector and Ultralight Dark Matter and a suite of experiments to extend the sensitivity of direct detection and accelerator-based searches
- I will discuss the *BABAR* contribution to current dark matter/dark photon search limits and, in the next talk, a potential future experiment to improve sensitivity, LDMX



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# Dark sector studies at *BABAR*

*BABAR* has conducted an extensive Dark Sector search program

Search for dark photon

$$e^+e^- \rightarrow \gamma A', A' \rightarrow e^+e^-, \mu^+\mu^-$$

$$e^+e^- \rightarrow \gamma A', A' \rightarrow \text{invisible}$$

$$\pi^0 \rightarrow \gamma \ell^+\ell^-, \eta \rightarrow \gamma \ell^+\ell^-, \phi \rightarrow \eta \ell^+\ell^-$$

Search for dark Higgs boson

$$e^+e^- \rightarrow h' A', h' \rightarrow A' A'$$

Search for dark boson(s)

$$e^+e^- \rightarrow \gamma A' \rightarrow W' W''$$

Search for dark hadrons

$$e^+e^- \rightarrow \pi_D + X, \quad \pi_D \rightarrow e^+e^-, \mu^+\mu^-$$

Search for dark scalar (*s*)  
and dark pseudoscalar (*a*)

$$B \rightarrow K^{(*)} s \rightarrow K^{(*)} \ell^+\ell^- / B \rightarrow K^{(*)} a \rightarrow K^{(*)} \ell^+\ell^-$$

$$B \rightarrow ss \rightarrow 2(\ell^+\ell^-)$$

Search for “muonic dark force”

$$e^+e^- \rightarrow \mu^+\mu^- Z', Z' \rightarrow \mu^+\mu^-$$

Search for leptophilic dark scalar

$$e^+e^- \rightarrow \tau^+\tau^- h', h' \rightarrow \mu^+\mu^- \text{ (4 leptons + MET)}$$

Search for self-interacting dark matter

$$e^+e^- \rightarrow \gamma A' A' A', A' \rightarrow \ell^+\ell^-, \pi^+\pi^-$$

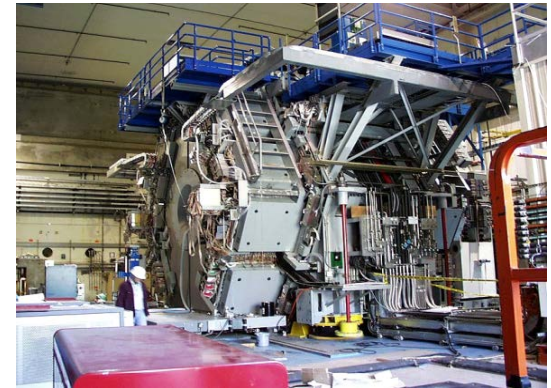
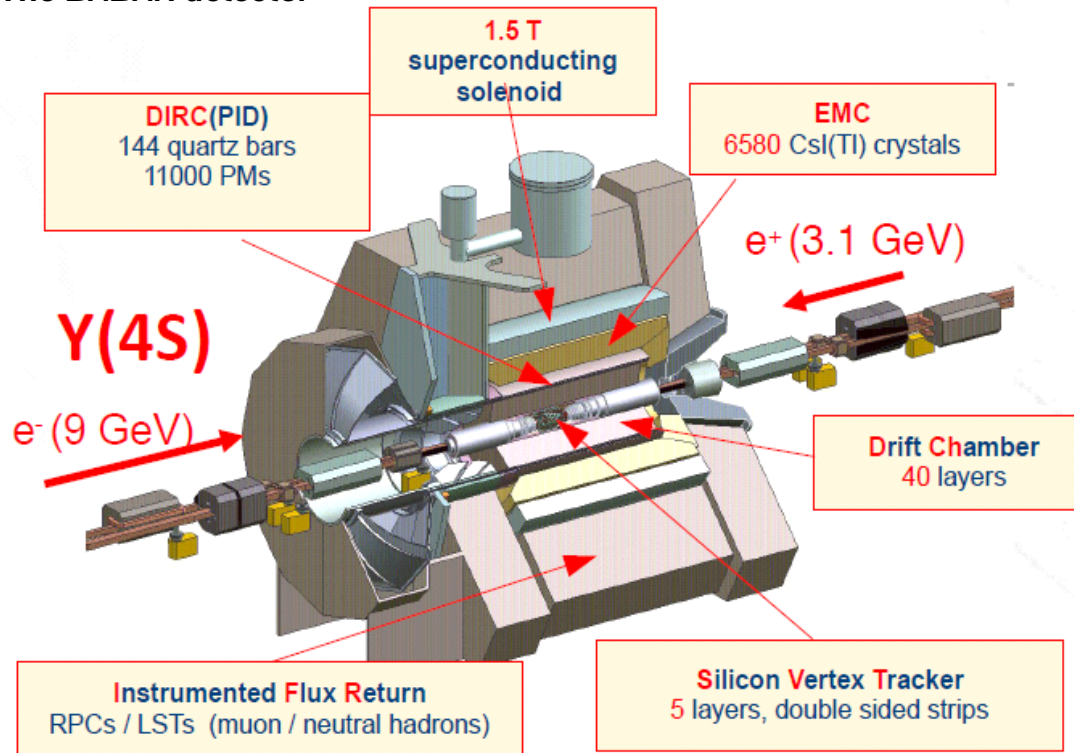
Ongoing / published  
**Preliminary studies**



# The *BABAR* experiment

*BABAR* at PEP-II collected  $\sim 500 \text{ fb}^{-1}$  of data, mainly at and around the  $\Upsilon(4S)$ ,  $\Upsilon(3S)$  and  $\Upsilon(2S)$  resonances

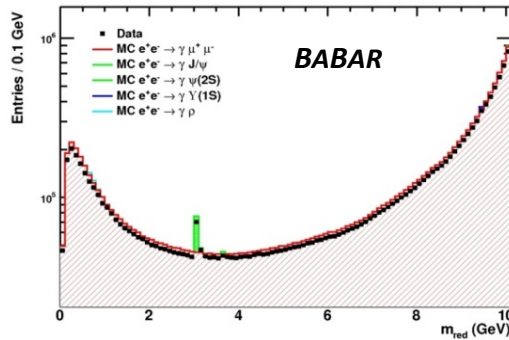
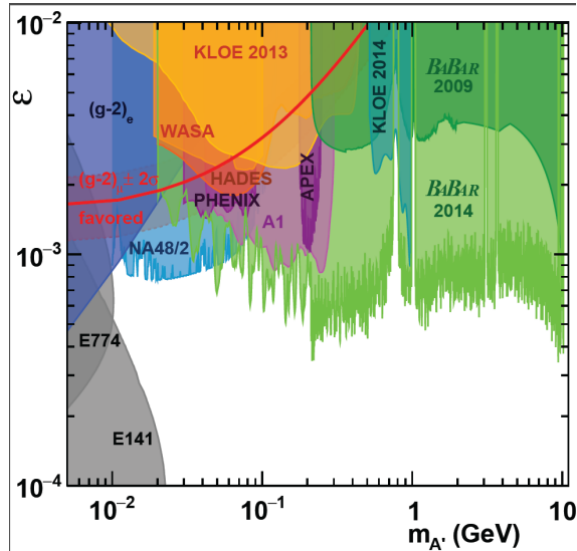
*The BABAR detector*



*B* factories offer an ideal environment to search for dark sector particles, provided there is an appropriate trigger, distinct from the usual multiparticle trigger

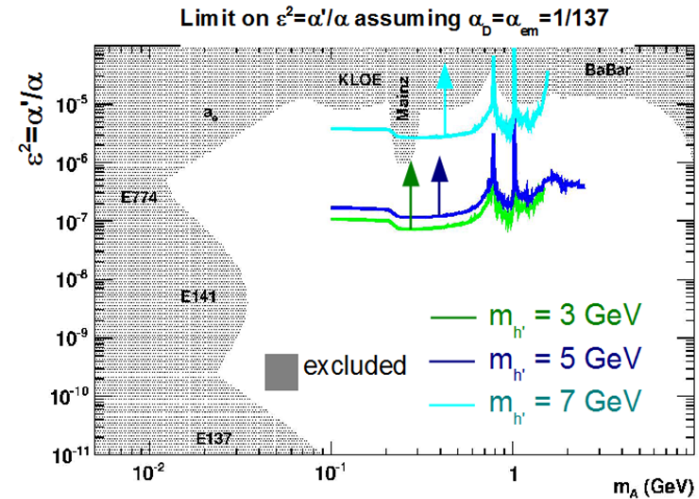
# Some highlights of the *BABAR* program

Search for dark photons in  $e^+e^- \rightarrow \gamma A'$ ,  $A' \rightarrow e^+e^-, \mu^+\mu^-$



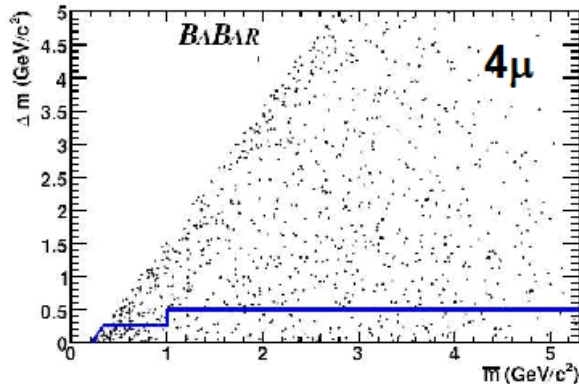
*PRL 113 (2014) 201801*  
*PRL viewpoint*

Search for dark Higgs bosons in  $e^+e^- \rightarrow h' A'$ ,  $h' \rightarrow A' A'$



*PRL 108 (2012) 211801*

Search for dark bosons in  $e^+e^- \rightarrow \gamma A' \rightarrow W' W'$



*arXiv:0908.2821*

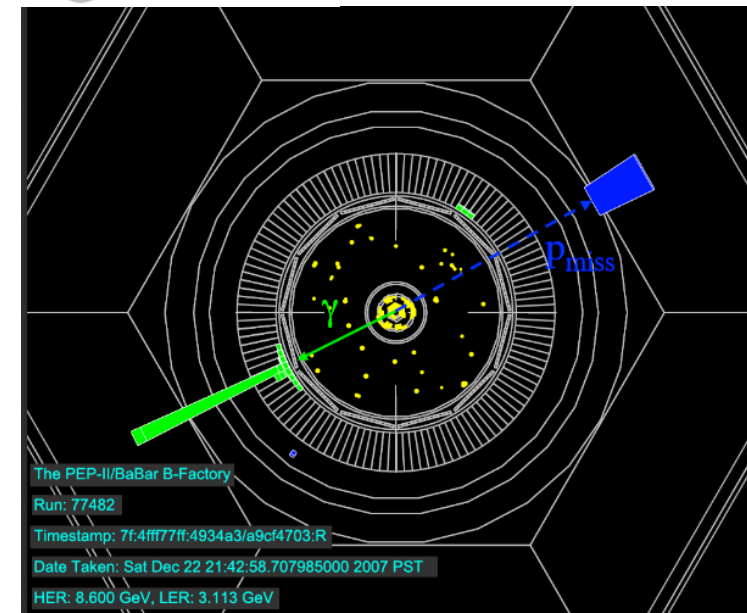
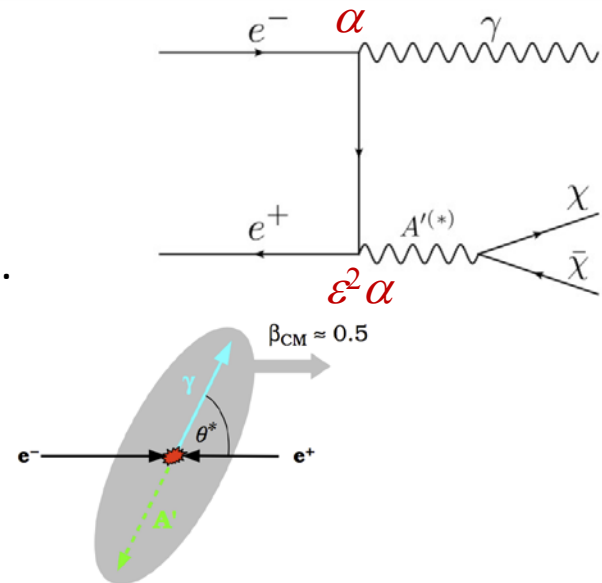
This talk will cover two recent measurements

Search for invisible dark photon decay  
Search for muonic dark force



# Search for invisible $A'$ decay

- The dark photon will decay invisibly if there are lighter dark sector states
- At  $e^+e^-$  colliders, we can search for  $e^+e^- \rightarrow \gamma A'$ ,  $A' \rightarrow$  invisible by tagging the recoil photon in “single photon” events.
- The signature of the invisible dark photon is a monochromatic photon and missing energy/mass
 
$$M_X^2 = s - 2E_\gamma^* \sqrt{s}$$
- Hermeticity is key, but we need to allow some machine background
- The search strategy:
  - select a single-photon final state, then look for a bump in missing mass  $M_X$  (or  $E_\gamma$ )
- Main backgrounds:  $e^+e^- \rightarrow \gamma\gamma$  and  $e^+e^- \rightarrow \gamma e^+e^-$  with particles outside detector acceptance
- *BABAR* collected  $\sim 53 \text{ fb}^{-1}$  of data with **dedicated single photon triggers** during its last year of data taking, mostly collected at the  $\Upsilon(3S)$  and  $\Upsilon(2S)$  energies.



# Single photon trigger

- Search requires a dedicated single-photon trigger
  - Level-1 Hardware trigger: 1 or more calorimeter clusters with  $E_{\text{lab}} > 0.8 \text{ GeV}$
  - Level-3 Software trigger: Two different software triggers were used

## Low $M_X$

- Require  $E_\gamma^* > 2 \text{ GeV}$
- No tracks originating from  $e^+e^-$  interaction region
- Trigger active for full  $53 \text{ fb}^{-1}$  data sample

## High $M_X$

- Require  $E_\gamma^* > 1 \text{ GeV}$
- No tracks originating from  $e^+e^-$  interaction region
- Trigger active for  $35.9 \text{ fb}^{-1}$  subset of  $53 \text{ fb}^{-1}$  data sample

	Low Mass	High Mass
$\Upsilon(4S)$	$5.9 \text{ fb}^{-1}$	-----
$\Upsilon(3S)$	$28 \text{ fb}^{-1}$	$20 \text{ fb}^{-1}$
$\Upsilon(2S)$	$14.4 \text{ fb}^{-1}$	$14.4 \text{ fb}^{-1}$
off-peak	$4.2 \text{ fb}^{-1}$	$1.5 \text{ fb}^{-1}$
Total	$53 \text{ fb}^{-1}$	$35.9 \text{ fb}^{-1}$

$$E_\gamma^* = \frac{s - M_X^2}{2\sqrt{s}}$$



# Event selection

$$\text{Low } M_X \\ 4 \text{ GeV}^2 < M_X^2 < 36 \text{ GeV}^2$$

- Dominant background from  $e^+e^- \rightarrow \gamma\gamma$  events where a photon escapes detection
- 1 Electromagnetic calorimeter (EMC) cluster
- Require  $E_\gamma^* > 3 \text{ GeV}$
- No drift chamber tracks with momentum  $p^* > 1 \text{ GeV}$
- Multivariate discriminator cut

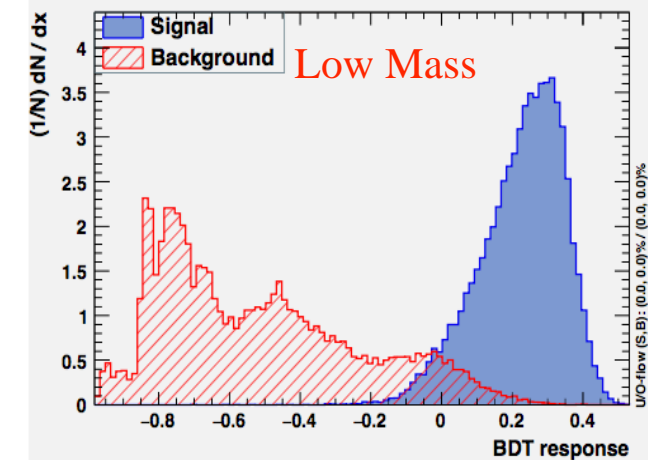
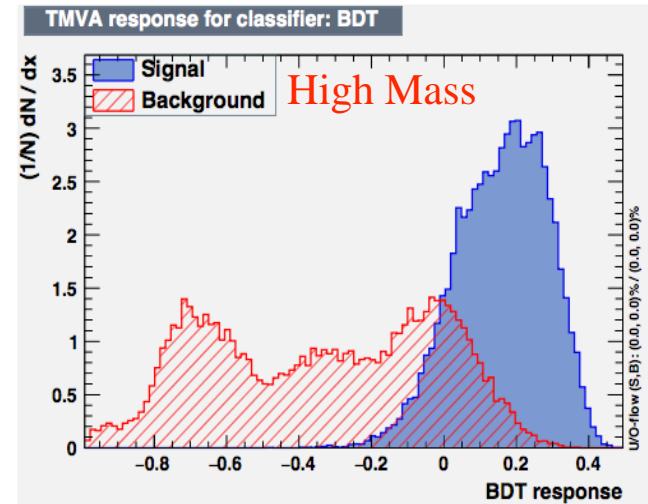
$$\text{Low } M_X \\ 24 \text{ GeV}^2 < M_X^2 < 69 \text{ GeV}^2 \quad \gamma(3S) \\ 63.5 \text{ GeV}^2 \quad \gamma(2S)$$

- Dominant background from radiative Bhabha events  $e^+e^- \rightarrow e^+e^- \gamma$  events where the electron and positron escape detection
- 1 EMC cluster with transverse profile consistent with an electromagnetic shower
- Require  $E_\gamma^* > 1.5 \text{ GeV}$
- No drift chamber tracks with momentum  $p^* > 0.1 \text{ GeV}$
- Multivariate discriminator cut



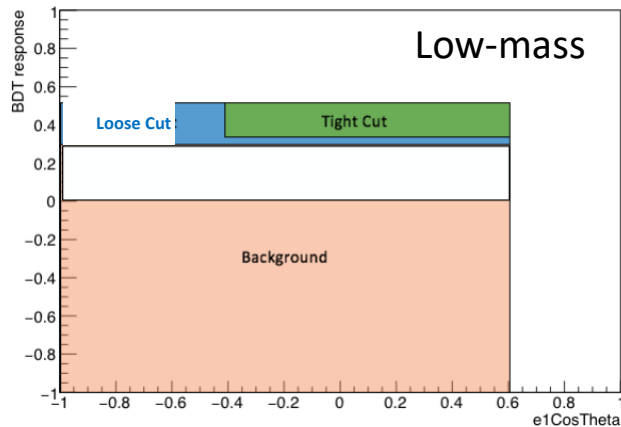
# The BDT discriminator

- Apply additional selection criterion using Boosted Decision Tree (BDT) multivariate discriminator
- The BDT is trained separately for the Low Mass and High Mass samples
- There are 12 discriminating variables including:
  - Shape parameters for the most energetic EMC cluster
  - Total EMC energy without the most energetic cluster
  - $E^*, \theta^*, \Delta\phi^*(E_1)$  of the second most energetic EMC cluster
  - $E^*, \theta^*, \Delta\phi^*(E_1)$  of the Instrumented Flux Return (IFR) cluster closest to the missing momentum direction
- BDT is trained using:
  - 25K simulated signal events with uniformly distributed  $A'$  masses
  - 25K background events from  $\Upsilon(3S)$  data sample



# Search for invisible $A'$ decay

To further reduce the residual peaking contribution from  $e^+e^- \rightarrow \gamma\gamma$  events near  $M_X \sim 0$ , instead of simply relying on the BDT output, we define several signal regions in the bi-dimensional space of BDT output vs photon angle to optimize the analysis



- Tight region: further remove the residual peaking background from  $e^+e^- \rightarrow \gamma\gamma$  near  $M_X \sim 0$
- Loose region: optimize for evidence of observation (smooth background)
- Background: pure background sample

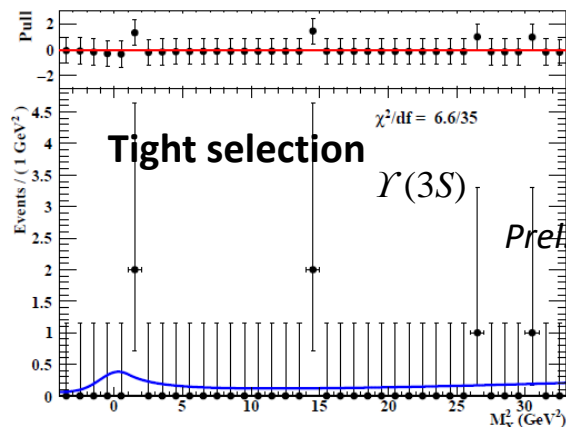
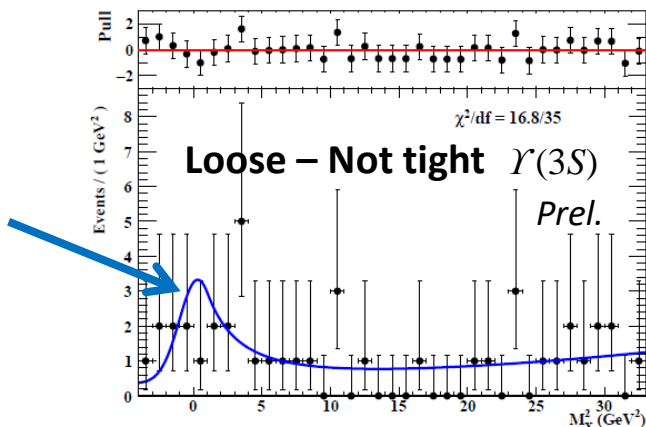
Split data into four non-overlapping regions for each dataset taken at different energies:  
Low-mass + tight, low-mass + loose and Not-tight, high-mass + loose, background

There are a total of 9 low-mass datasets and 4 high-mass datasets

# Search for invisible $A'$ decay

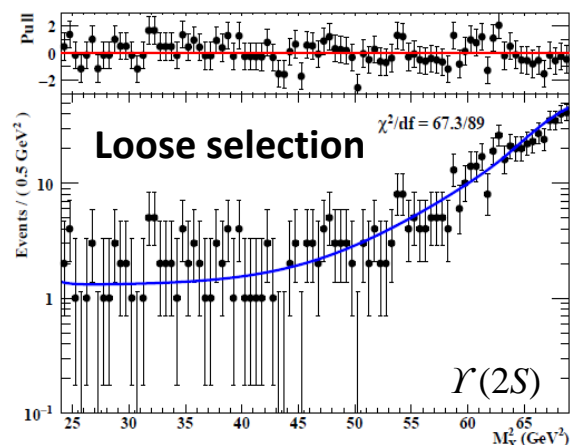
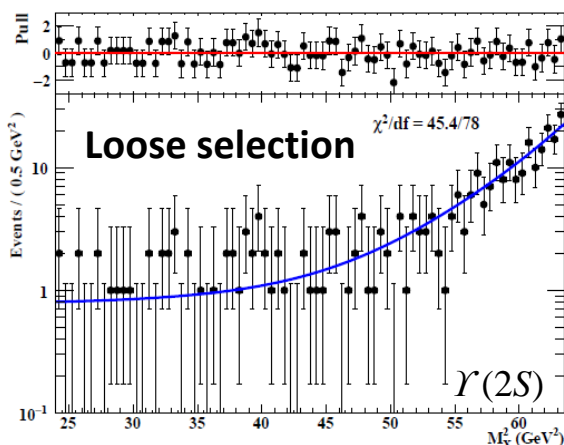
Low-mass region: data distributions at  $\Upsilon(3S)$

Peaking background  
 $e^+e^- \rightarrow \gamma\gamma$   
 (irreducible)



High-mass region: data distributions at  $\Upsilon(2S)$

In blue: bkg-only fit

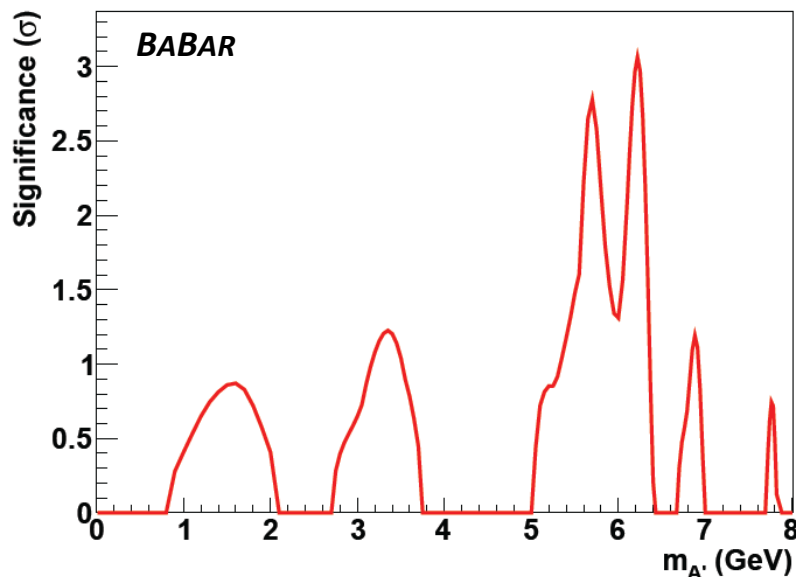


# Search for invisible $A'$ decay

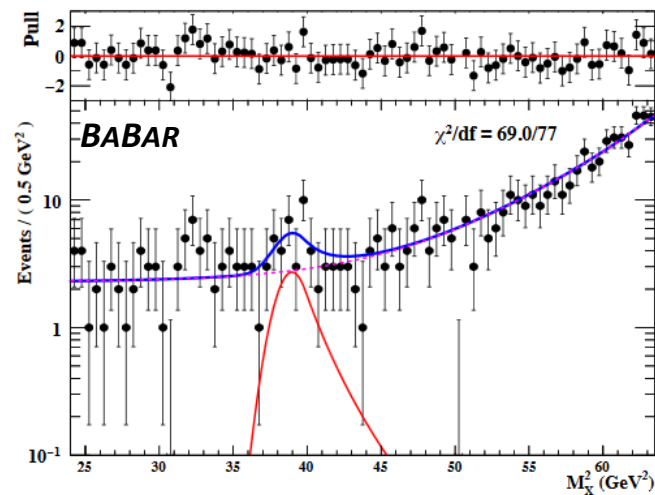
We extract the signal by a simultaneous fit to these independent regions for each beam energy, for a total of 166 mass hypotheses

For each fit, we fix the background shape using the background region, and float the signal yield, peaking and continuum background contributions

Signal significance distribution



Most significant fit  $m_A = 6.22$  GeV



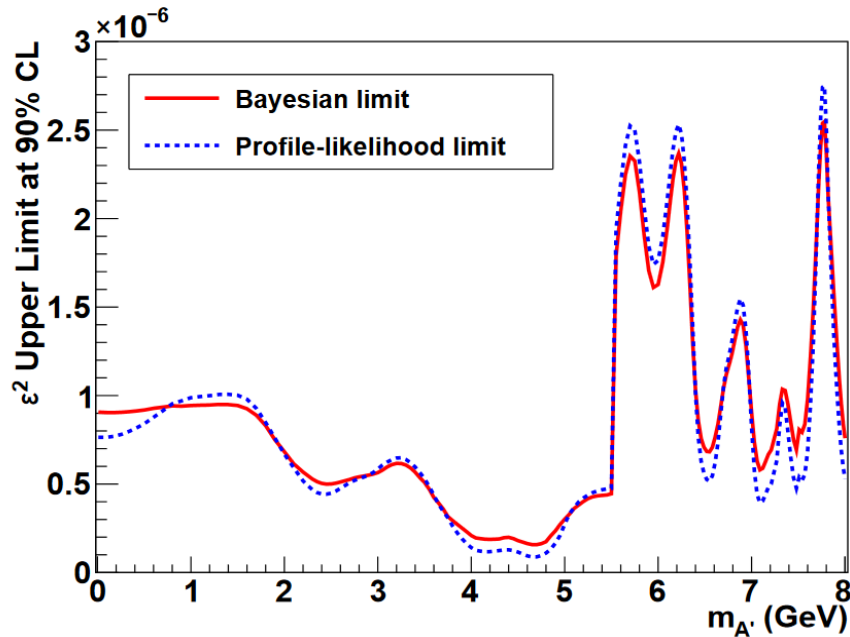
Local (global) significance:  $3.1\sigma$  ( $2.6\sigma$ )  
Global p-value  $\sim 1\%$

We find no significant signal

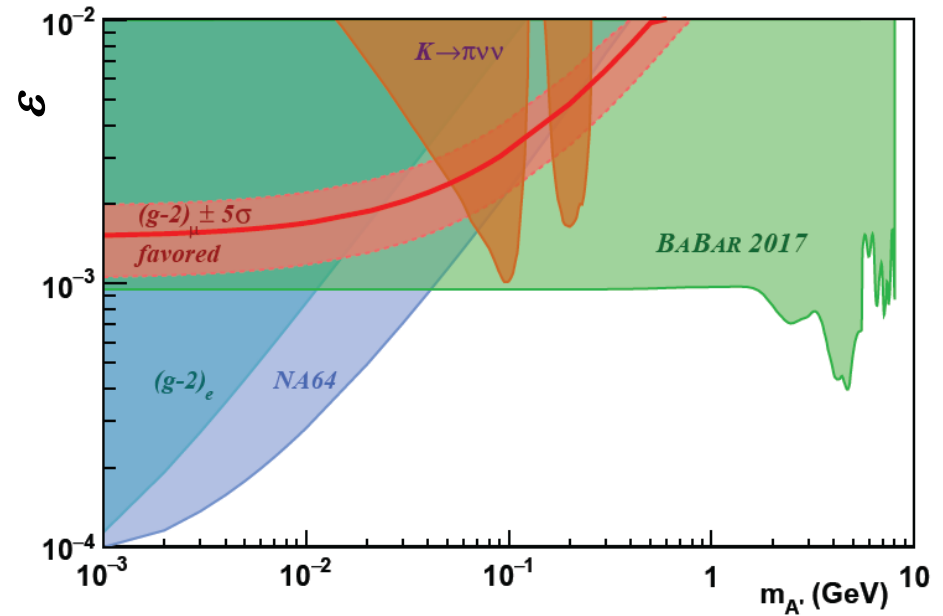
# Search for invisible $A'$ decay

Phys.Rev.Lett. 119, 131804 (2017)

Limits on mixing parameter  $\varepsilon^2$



Limits (90% CL) on mixing parameter



This represents a significant improvement over previous measurements

Rules out the entire region preferred by  $(g-2)_\mu$  anomaly



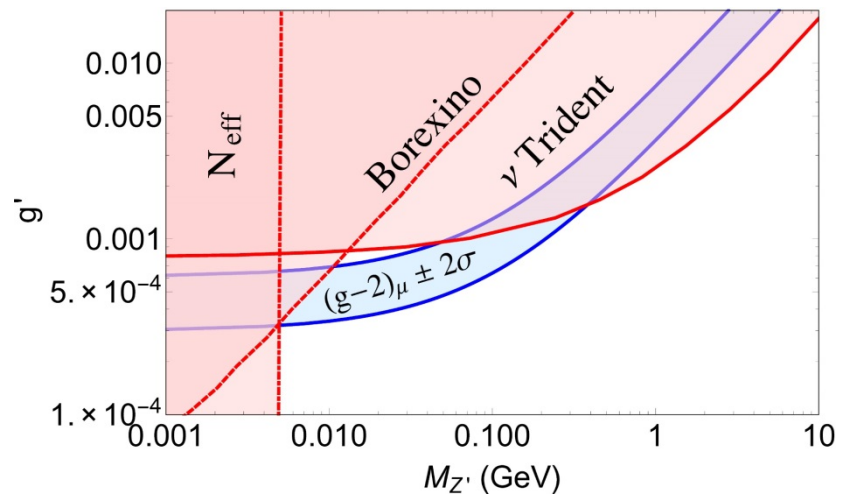
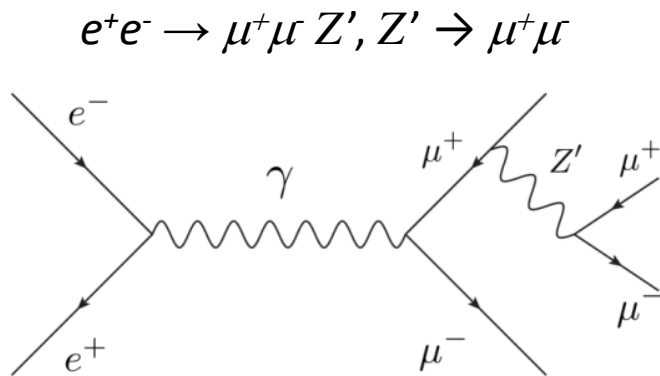
# Muonic dark force

Consider a new dark force that couples only to the second and third generation of leptons with a corresponding gauge boson  $Z'$  (arXiv:1401.2459)

Such a force could explain various anomalies observed in the muon sector (“ $g-2$ ” discrepancy, proton radius puzzle, ...), and account for dark matter as sterile neutrinos by increasing their cosmological abundance via new interactions with SM neutrinos

Some constraints from neutrino physics have already been derived, but they only indirectly probe the existence of  $Z'$  (with large systematic uncertainties)

We can search for direct  $Z'$  production at colliders via  $Z'$  bremsstrahlung :



# Muonic dark force

Search for  $Z'$  in  $e^+e^- \rightarrow \mu^+\mu^- Z'$ ,  $Z' \rightarrow \mu^+\mu^-$  events

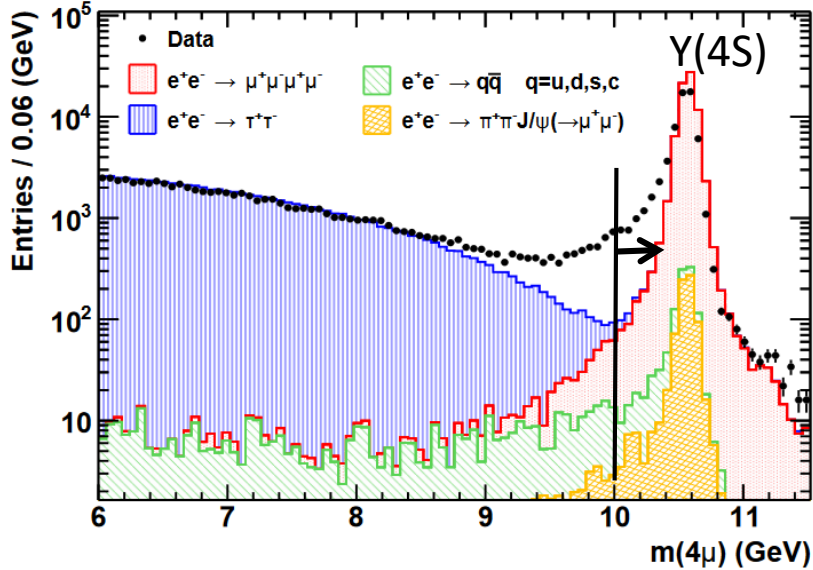
## Analysis overview

- Analysis based on  $514 \text{ fb}^{-1}$  of data collected at  $Y(4S)$ ,  $Y(3S)$  and  $Y(2S)$
- Requirements
  - Four tracks and no extra neutral energy ( $E_{\text{extra}} < 200 \text{ MeV}$ )
  - Particle identification: at least 2 same-sign tracks identified as muon
  - Invariant mass of the four  $\mu$ s within 500 MeV of the nominal CM-energy
  - Events having a dimuon candidate within 10 MeV of the  $Y(1S)$  mass for the  $Y(2S)$  and  $Y(3S)$  dataset are vetoed to reject
$$Y(2S,3S) \rightarrow \pi\pi Y(1S), Y(1S) \rightarrow \mu\mu$$
- A kinematic fit imposing the beam-energy constraint is performed, but no constraints on the  $\chi^2$  are applied.

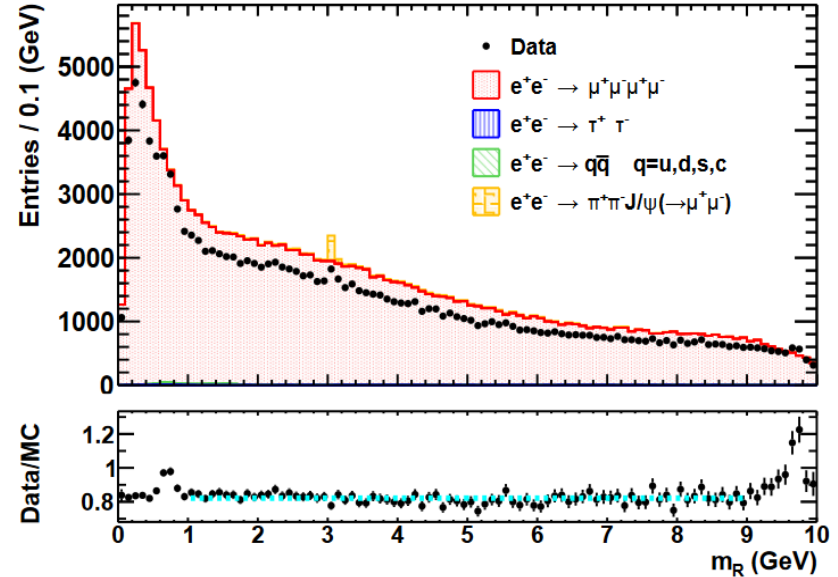
This is a blind analysis, with the selections criteria optimized on a small subset (5%) of the data, which is subsequently discarded

# Muonic dark force

Four-muon invariant mass -  $Y(4S)$  dataset



Reduced di-muon mass (2 entries / event)



$$m_R = (m_{\mu\mu}^2 - 4 m_\mu^2)^{1/2}$$

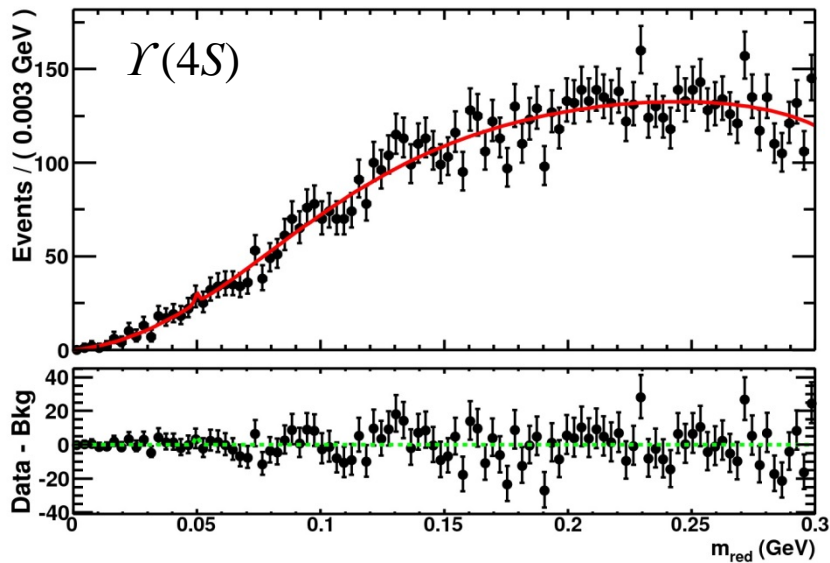
- The signal region is dominated by  $e^+e^- \rightarrow 4\mu$  background
- Discrepancies arise primarily from ISR, which is not in the Monte Carlo, as well as from differences in particle identification efficiencies, track reconstruction,...
- The ratio data/MC is used to correct the signal efficiency
- The low  $4\mu$  mass region is well-reproduced by the Monte Carlo
- Factoring in the ISR contribution, the correction factors derived from this region agree with those determined at high masses

# Muonic dark force

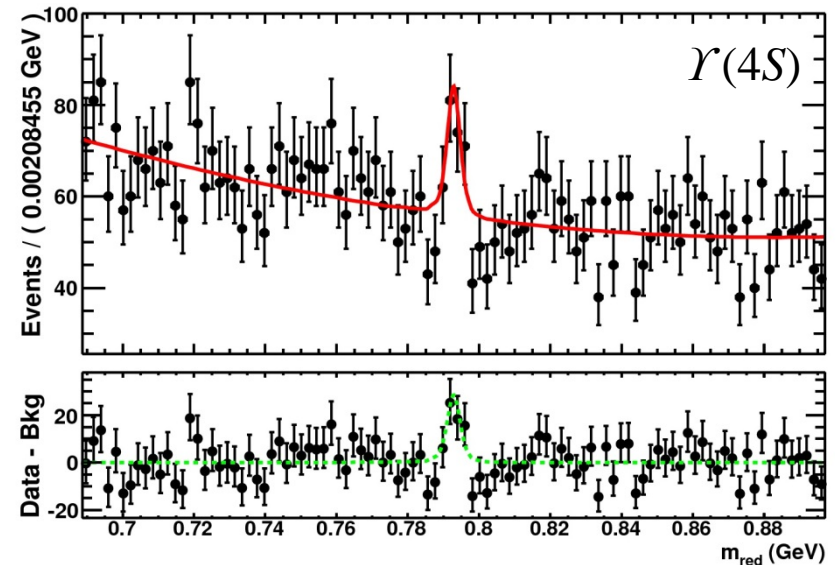
We extract the signal separately for the data at the  $\Upsilon(4S)$ ,  $\Upsilon(3S)$  and  $\Upsilon(2S)$  by performing a series of fits to the reduced dimuon mass for each sample

For each mass hypothesis, we fit over a fixed range of 0-0.3 GeV ( $m_R < 0.2$  GeV) or a window corresponding to  $5\sigma$  signal resolution ( $m_R > 0.2$  GeV). A region of  $\pm 30$  MeV around the  $J/\psi$  is excluded

Fit  $m_R = 0.05$  GeV



Most significant fit  $m_R = 0.79$  GeV



Local /global significance:  $4.3\sigma$  /  $1.6\sigma$

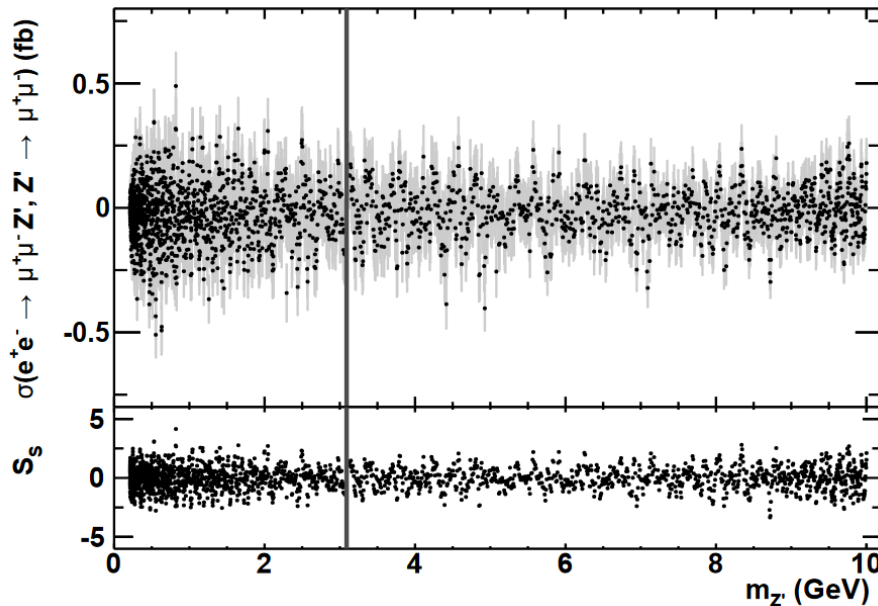
# Muonic dark force

We extract the signal separately for the data at the  $\Upsilon(4S)$ ,  $\Upsilon(3S)$  and  $\Upsilon(2S)$  by performing a series of fits to the reduced dimuon mass for each sample

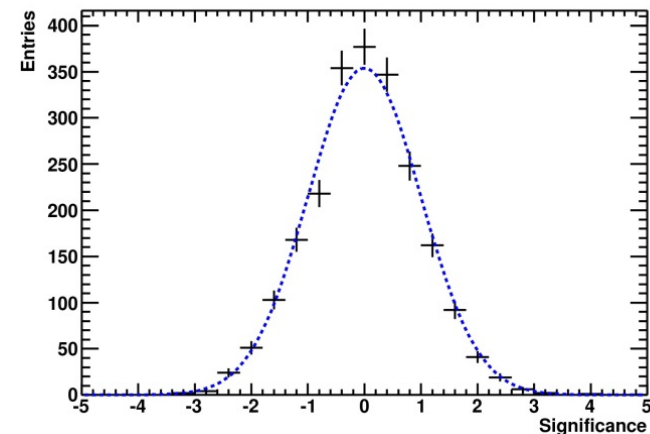
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We extract the cross-section separately for the data at the  $\Upsilon(4S)$ ,  $\Upsilon(3S)$  and  $\Upsilon(2S)$  and combine the results together

Combined cross-section and significance



Significance distribution

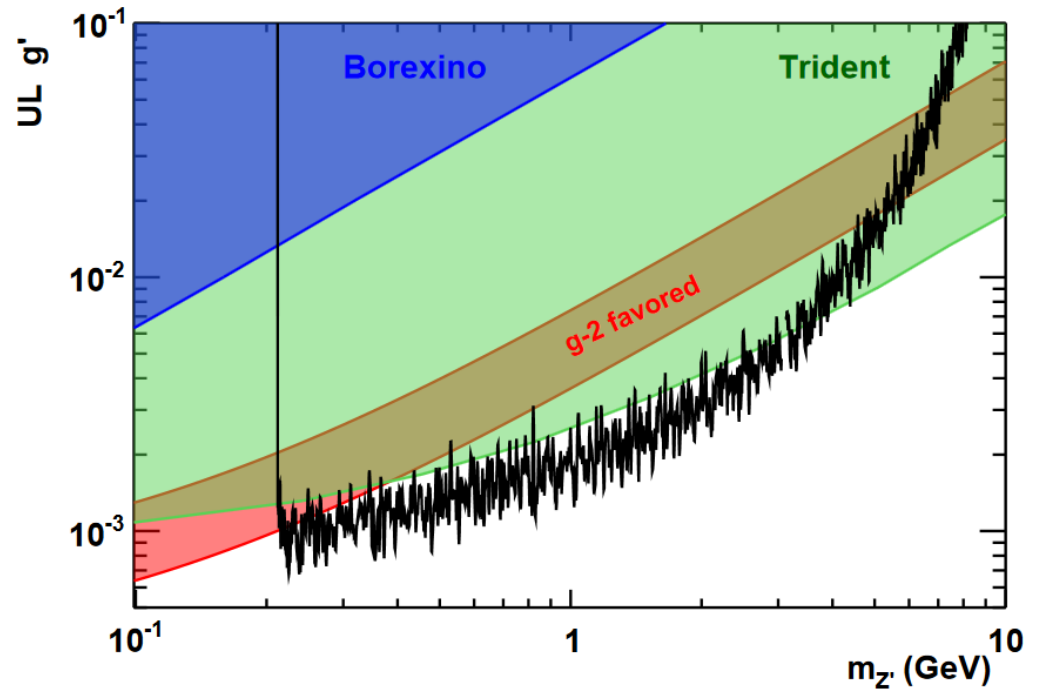
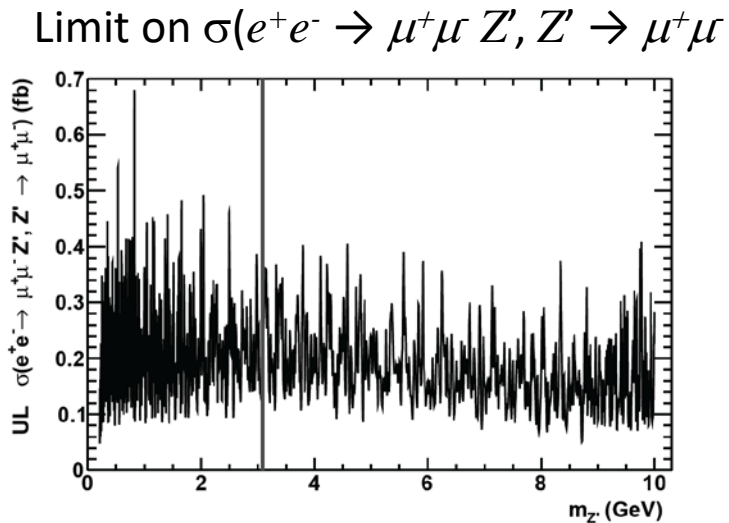


Compatible with null hypothesis

# Muonic dark force

PRD 94, 011102 (2016)

Limits (90% CL) on  $Z'$  coupling



These results improve previous constraints from the neutrino experiments and exclude all but a sliver of the region favored by the “ $g-2$ ” anomaly.

# More to come...

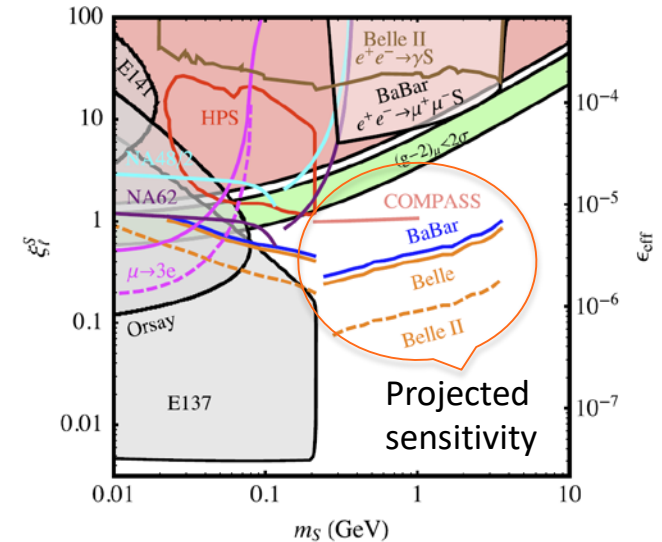
## On-going searches for dark sector at BABAR

Search for dark scalar ( $\phi$ ): A light dark scalar could couple to SM fermions via its mixing with the Higgs. Since the coupling are proportional to the mass, the search strategy is to look for a dimuon resonance in  $e^+e^- \rightarrow \tau^+\tau^- \phi$ ,  $\phi \rightarrow \mu^+\mu^-$ ,  $e^+e^-$  events. In this manner *BABAR* should be able to probe the remaining “ $g-2$ ” preferred region at low masses.

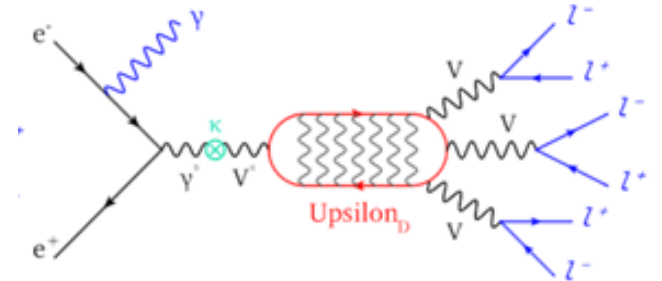
Self-interacting dark matter: if the dark-sector coupling is strong, dark sector bound states (darkonium) could be formed. These states have a striking multi-muon final state signature at *BABAR*.

And hopefully more...

## Limits on dark scalar ( $\phi$ ) coupling



B. Batell, N. Lange, D. McKeen, M. Pospelov, A. Ritz  
Phys.Rev. D95, 075003 (2017)



# Summary

Light dark sectors having a rich phenomenology have emerged as a fertile new regime for dark matter searches, as the WIMP paradigm is now highly constrained

Low-energy, high-intensity colliders provide a sensitive probe for dark sectors

*BABAR* has an extensive program to search for dark sector signatures, and set stringent limits on their existence

Belle II will over the next decade improve these collider limits

Extending the search region to the thermal relic limit at low masses requires fixed target accelerator experiments .....

