

A DARK FORCE FOR BARYONS

Ian Shoemaker

BF2011

October 19th, 2011

with Michael Graesser and Luca Vecchi



OR:

WHY YOU SHOULD LEARN TO
STOP WORRYING AND GAUGE
BARYON NUMBER

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Dark matter à la Occam



lex parsimoniae

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Visible sector $\sim 17\%$

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THE STANDARD MODEL

	Fermions			Bosons	
Quarks	u up	c charm	t top	γ photon	Force carriers
	d down	s strange	b bottom	Z Z boson	
Leptons	ν_e electron neutrino	ν_μ muon neutrino	ν_τ tau neutrino	W W boson	
	e electron	μ muon	τ tau	g gluon	
			Higgs*		

*Yet to be confirmed

Source: AAAS

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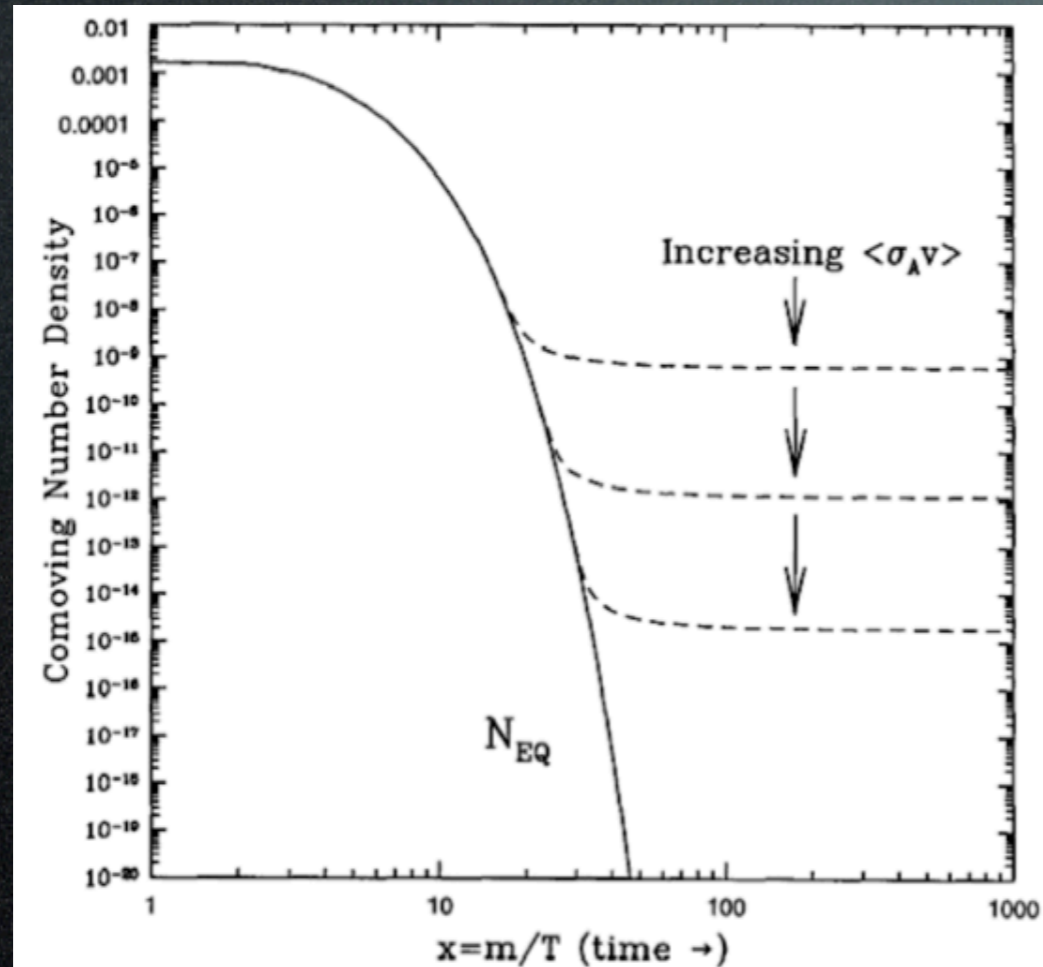
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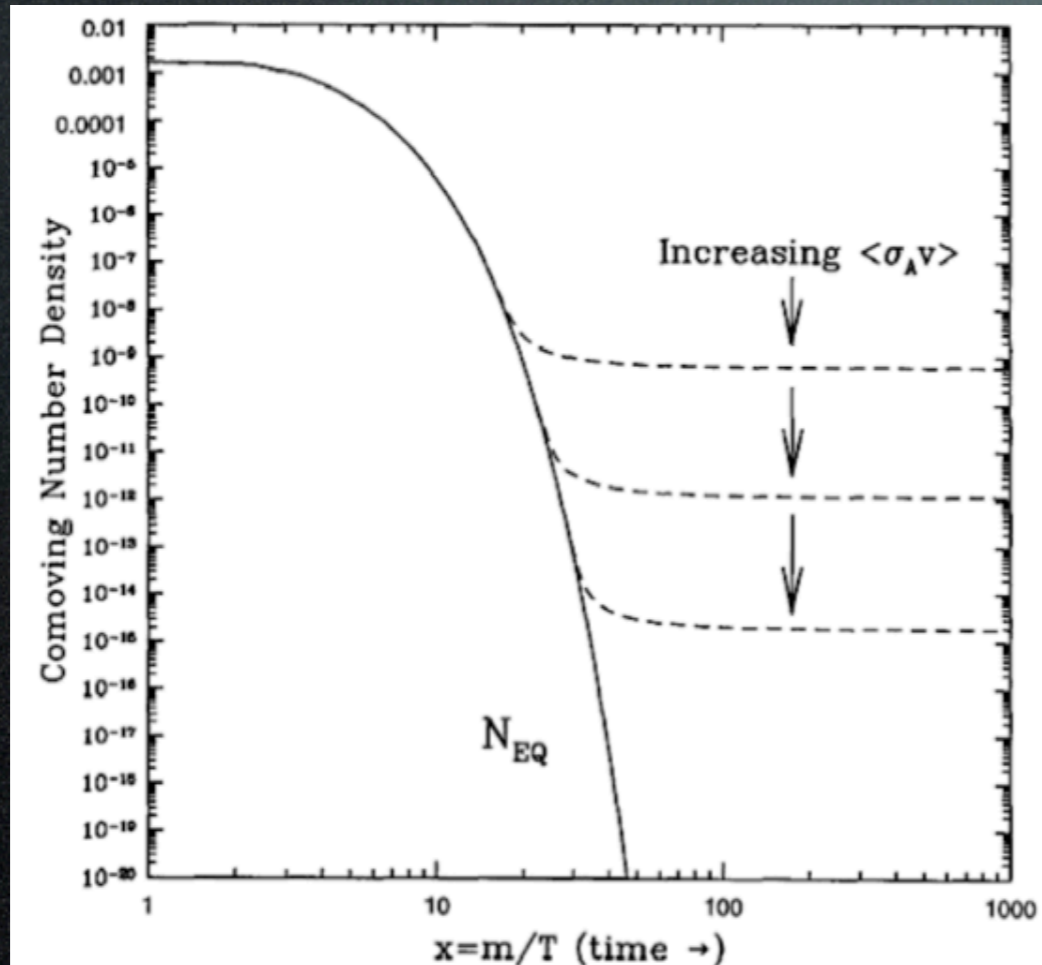
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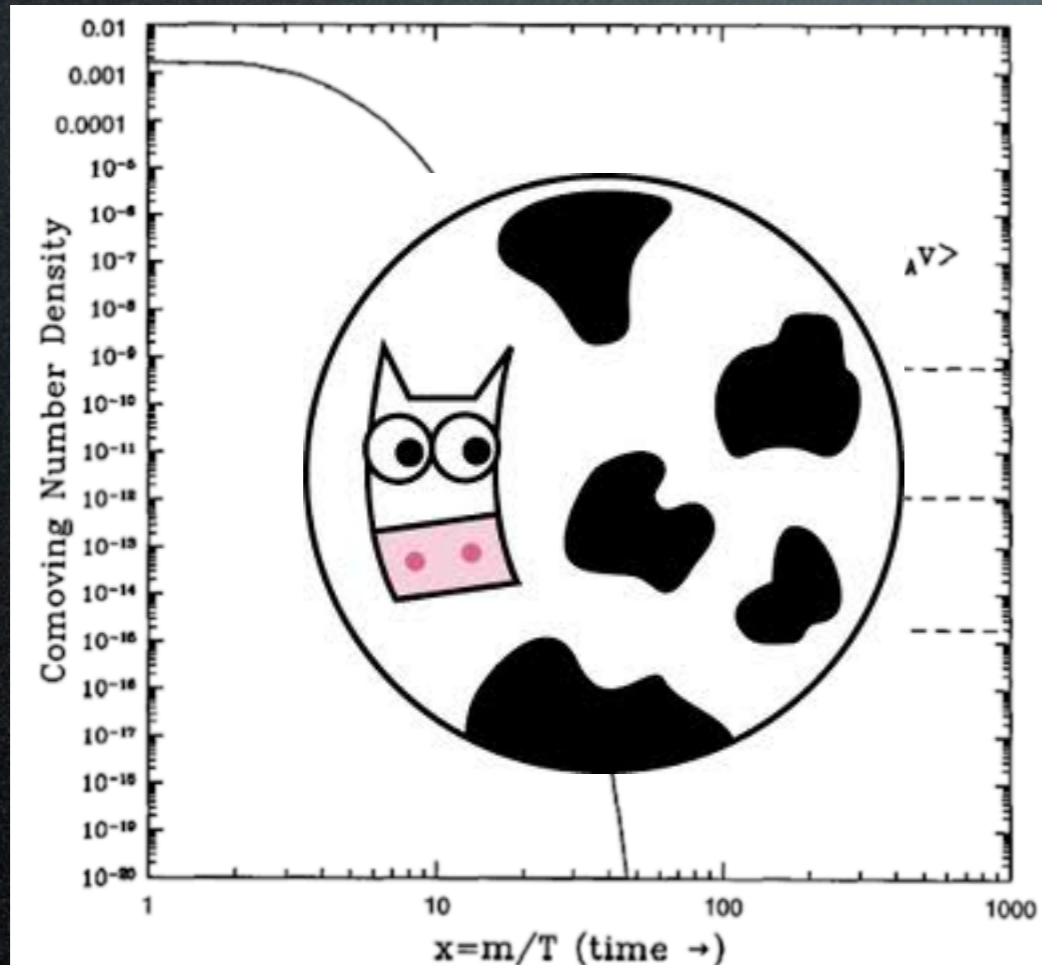
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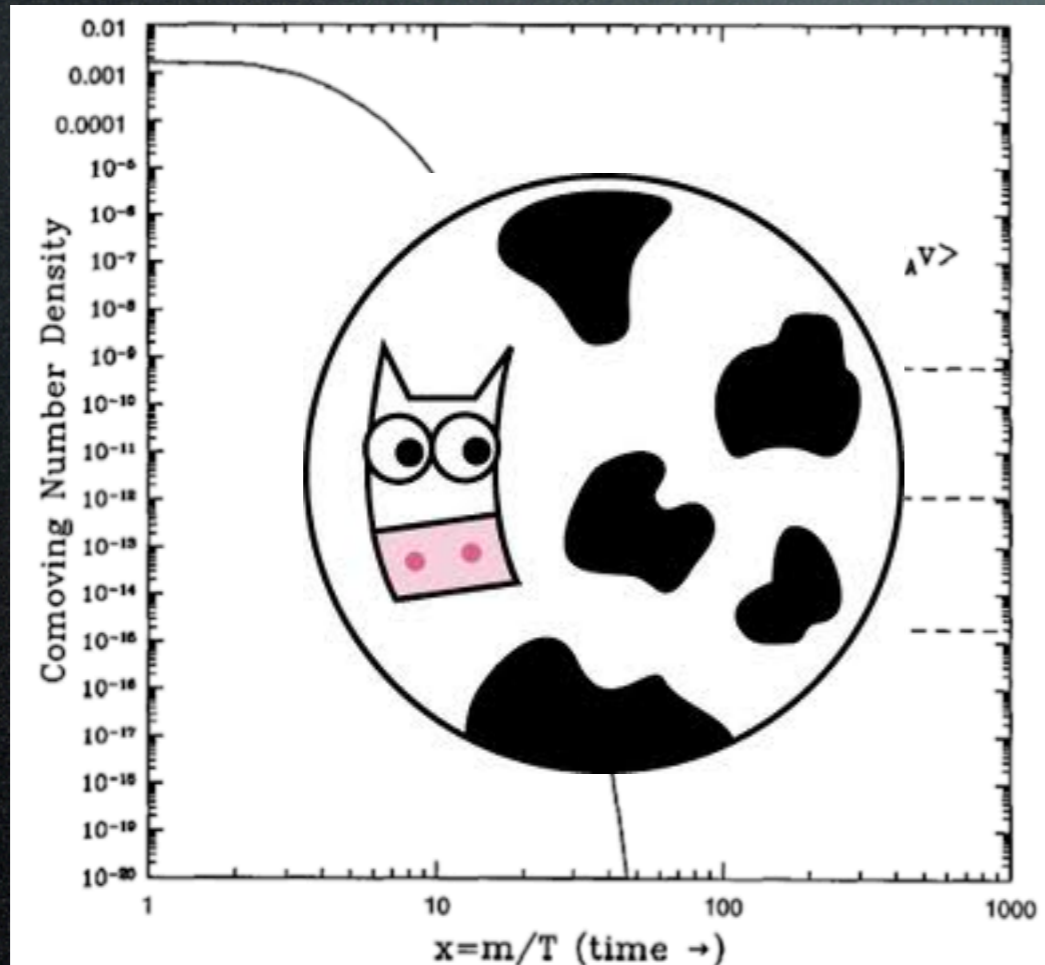
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1. Cosmological abundance.
2. It's stable (or at least very long-lived).

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 3. An indication of an underlying origin.

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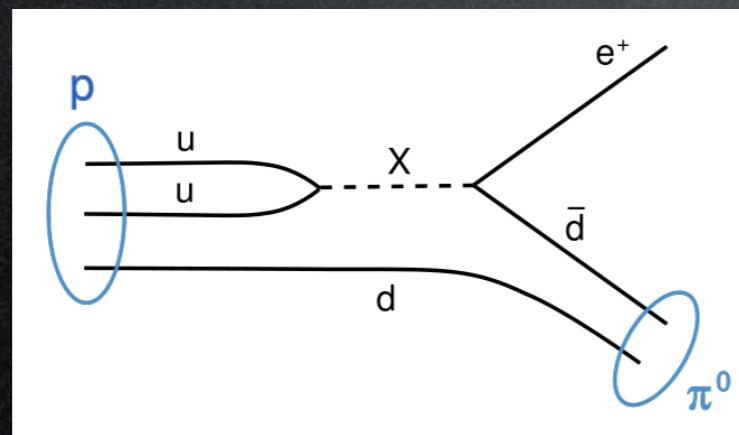
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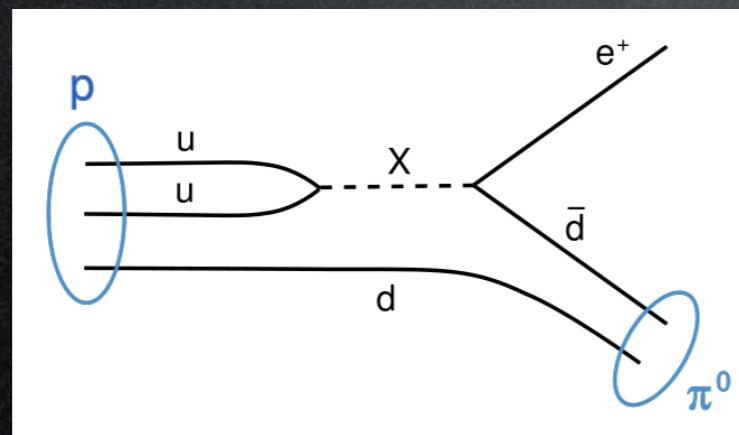
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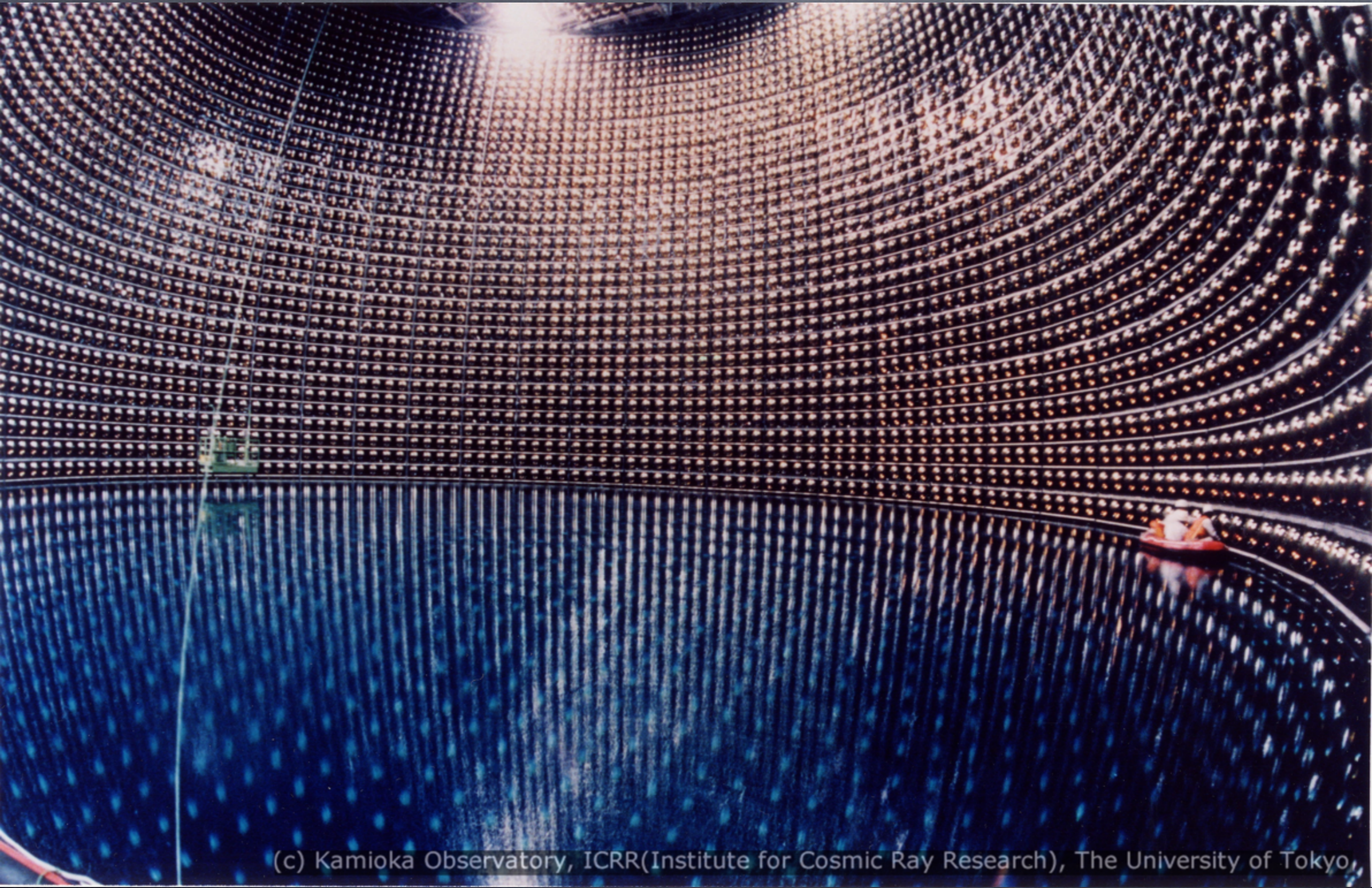
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The only problem is...

Super Kamiokande says:



(c) Kamioka Observatory, ICRR(Institute for Cosmic Ray Research), The University of Tokyo,

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Baryon number is an
unreasonably good symmetry

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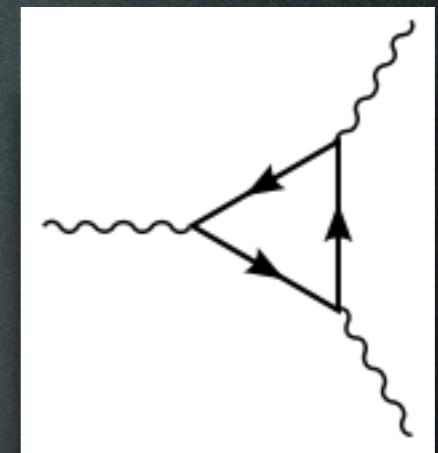
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- Baryogenesis *requires* a DM asymmetry.
- Shared gauge interactions with baryons predict novel signatures: monojets and low mass DD.

Gauging baryon number

- Older examples:
 - Carone and Murayama 1994; Bailey and Davidson 1995; Aranda and Carone 1998.
- More recently:
 - Dulaney, Fileviez-Perez and Wise (2010); Buckley, Fileviez-Perez, Hooper, and Neil (2011).

An anomaly-free example



- New chiral states

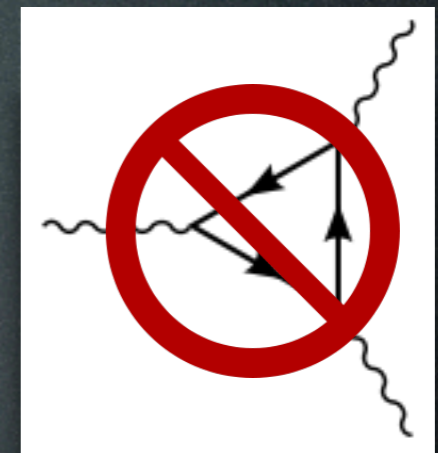
	$SU(3)_C$	$SU(2)_W$	$U(1)_Y$	$U(1)_B$
Q'_i	3	2	$+\frac{1}{6}$	$-\frac{1}{N}$
u'_{ci}	$\bar{3}$	1	$-\frac{2}{3}$	$+\frac{1}{N}$
d'_{ci}	$\bar{3}$	1	$+\frac{1}{3}$	$+\frac{1}{N}$
L'_i	1	2	$-\frac{1}{2}$	0
ν'_{ci}	1	1	0	0
e'_{ci}	1	1	+1	0

N dark generations

- Spontaneously break $U(1)_B$

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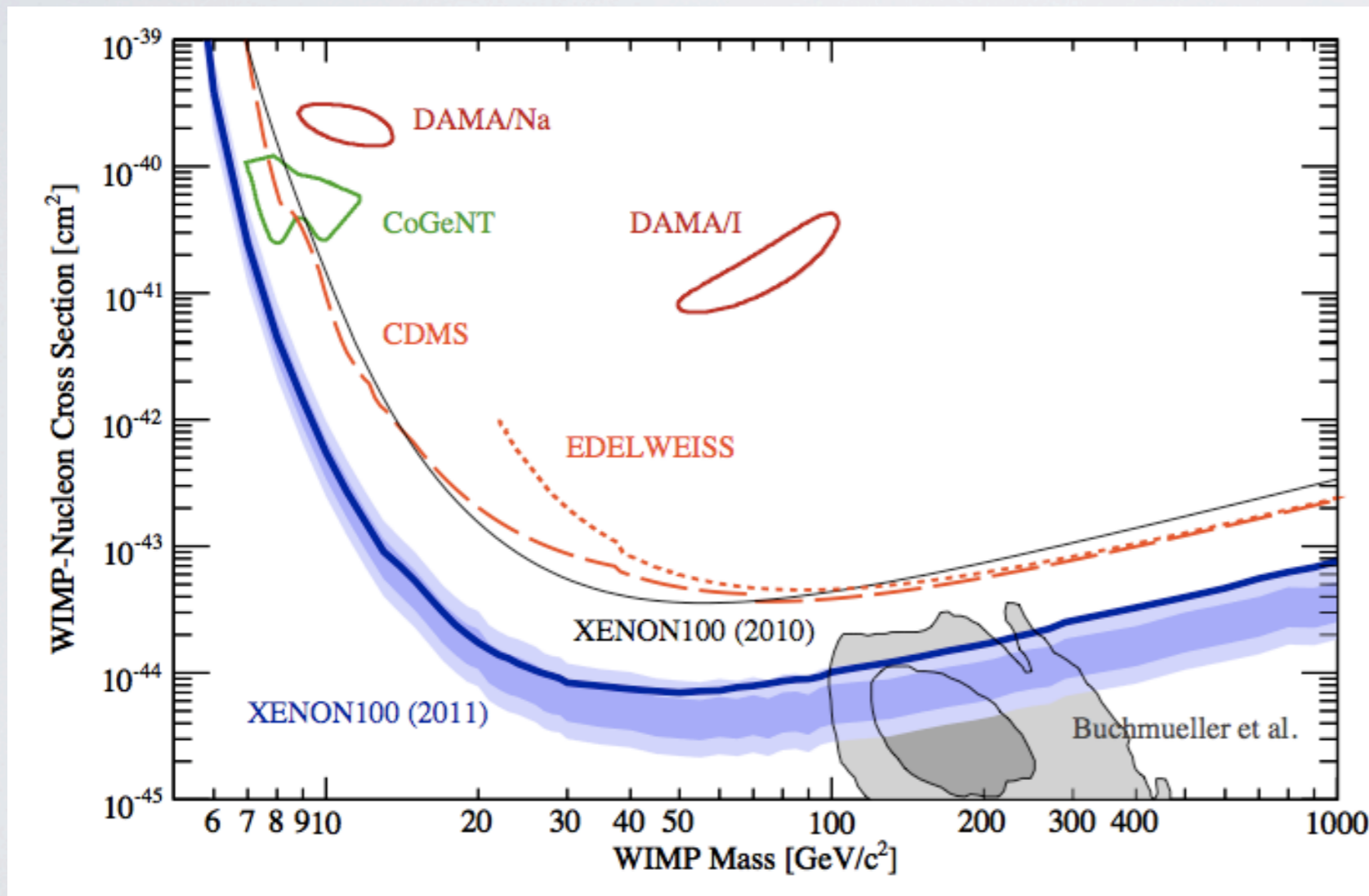
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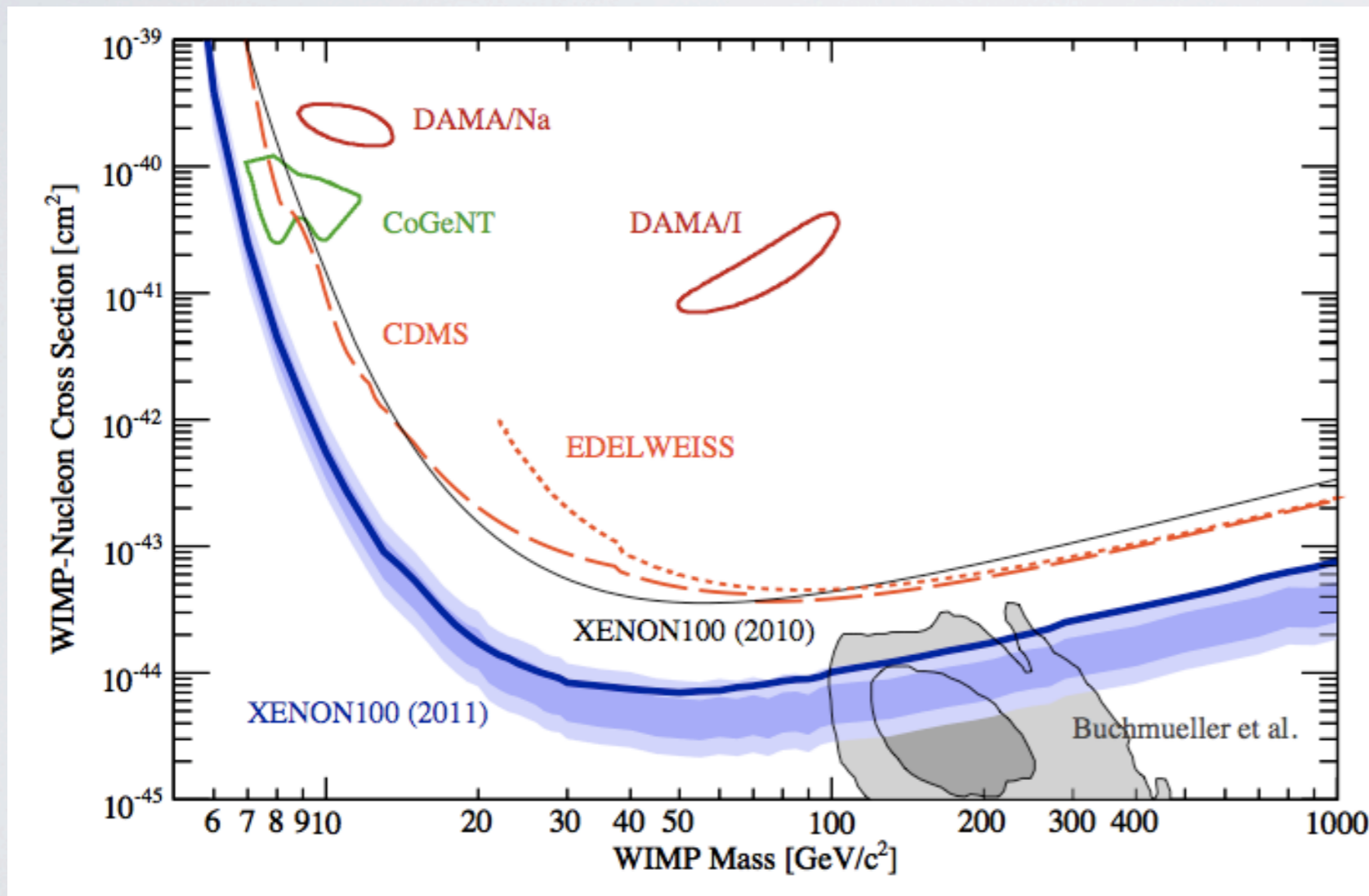
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- Recent work by: Bell, Petraki, IMS, Volkas [1105.3730].

DIRECT DETECTION BOUNDS



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annihilation physics
↕
DM-quark
scattering

RECOIL SPECTRUM

$$\frac{dR}{dE_R} = \frac{N_T \rho_{\odot}}{m_X} \int_{|\vec{v}| > v_{min}} d^3v \, v f(\vec{v}, \vec{v}_{\oplus}) \frac{d\sigma}{dE_R}$$

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[Lisanti, Strigari, Wacker,
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High-velocity tail is important for light DM.

RECOIL SPECTRUM

VECTOR CASE:

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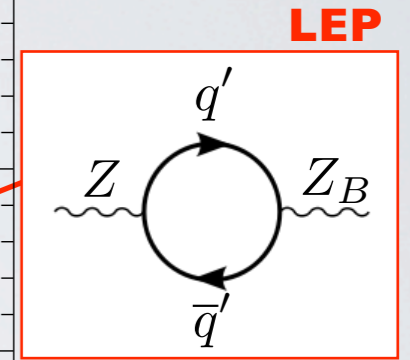
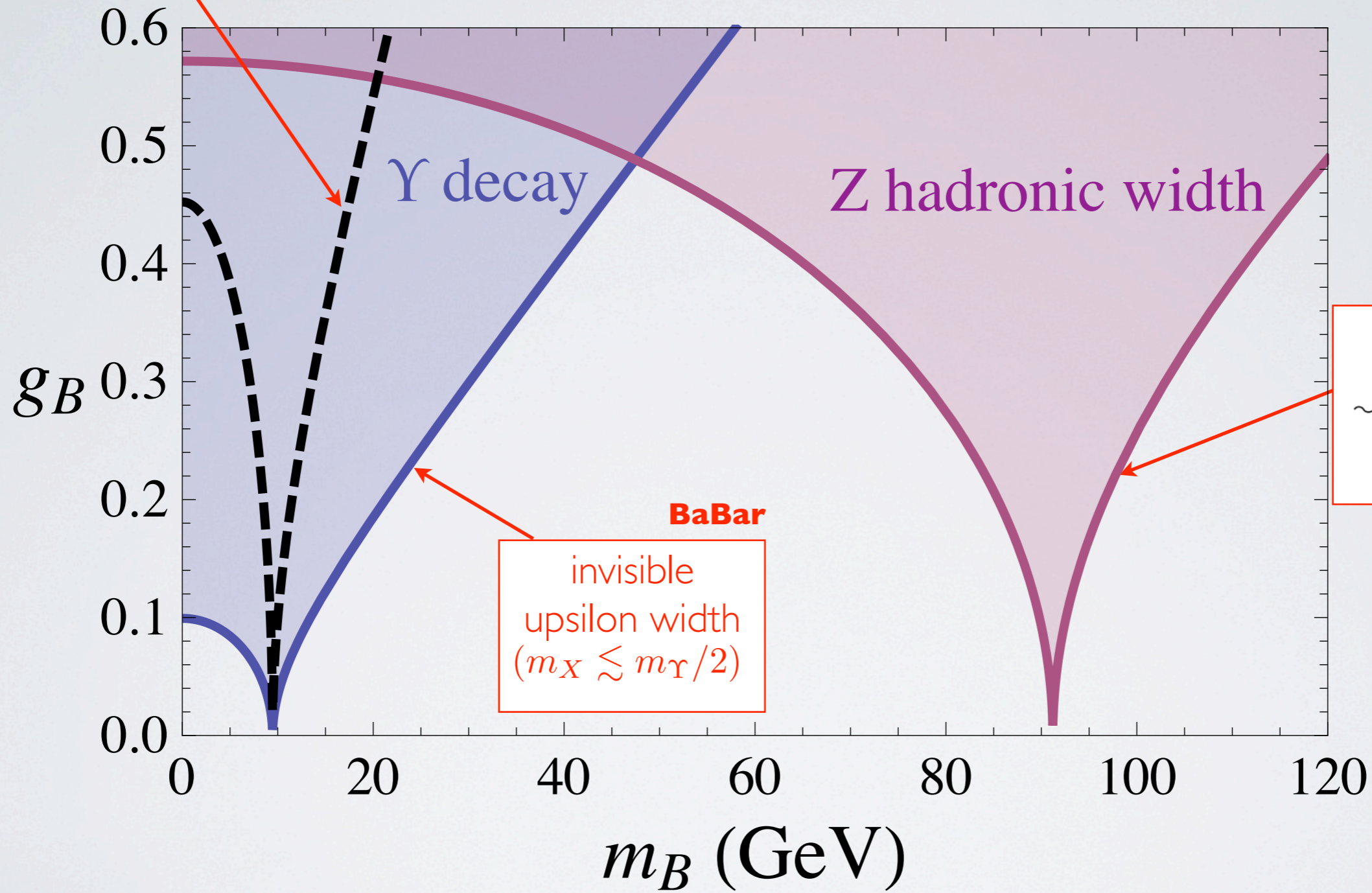
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DD imposes:
no bound

LEP + B-factories

BaBar
hadronic upsilon
width
($m_X \gtrsim m_\Upsilon/2$)

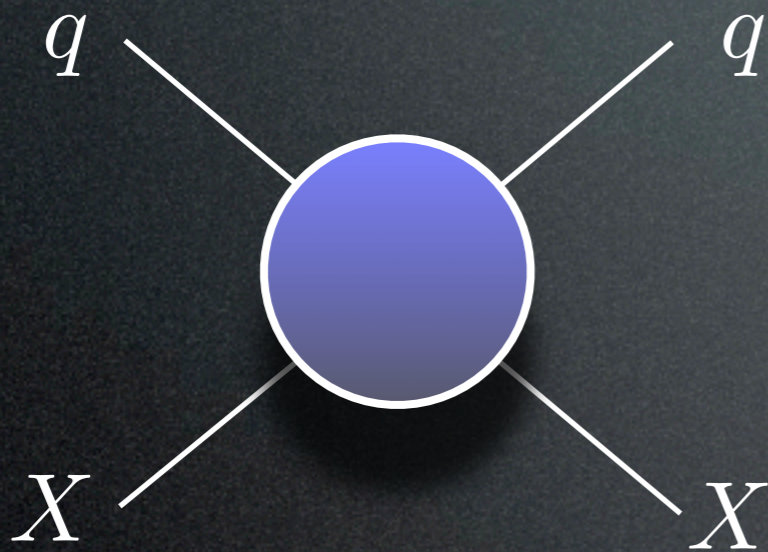


Monojets at the Tevatron

- For light DM, the Tevatron and the LHC are the world's best DD experiments [Goodman, et al. (2010); Bai, Fox, Harnik (2010)].

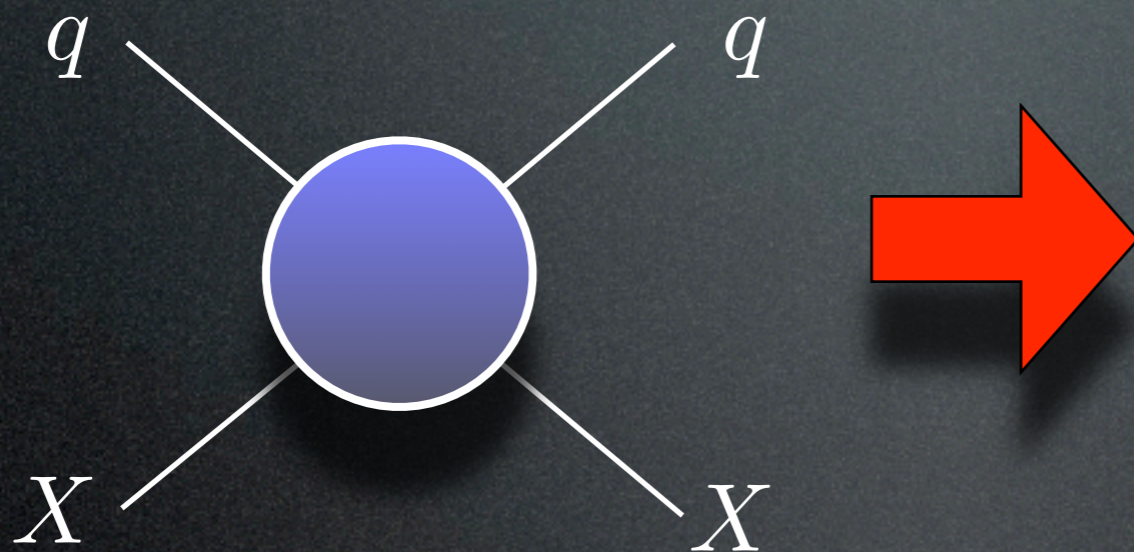
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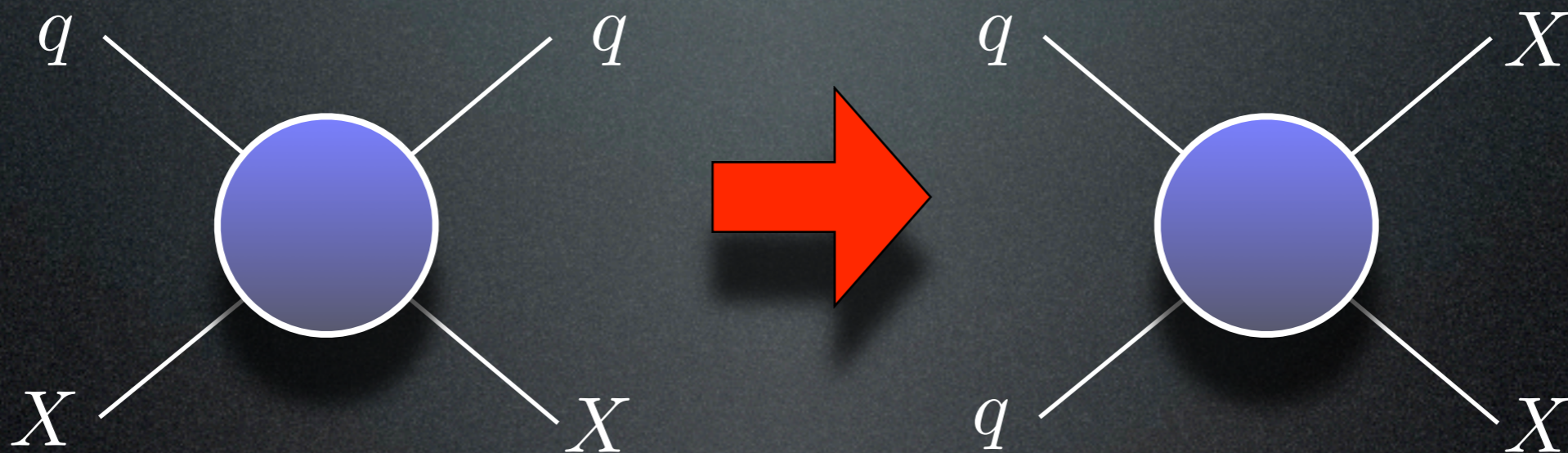
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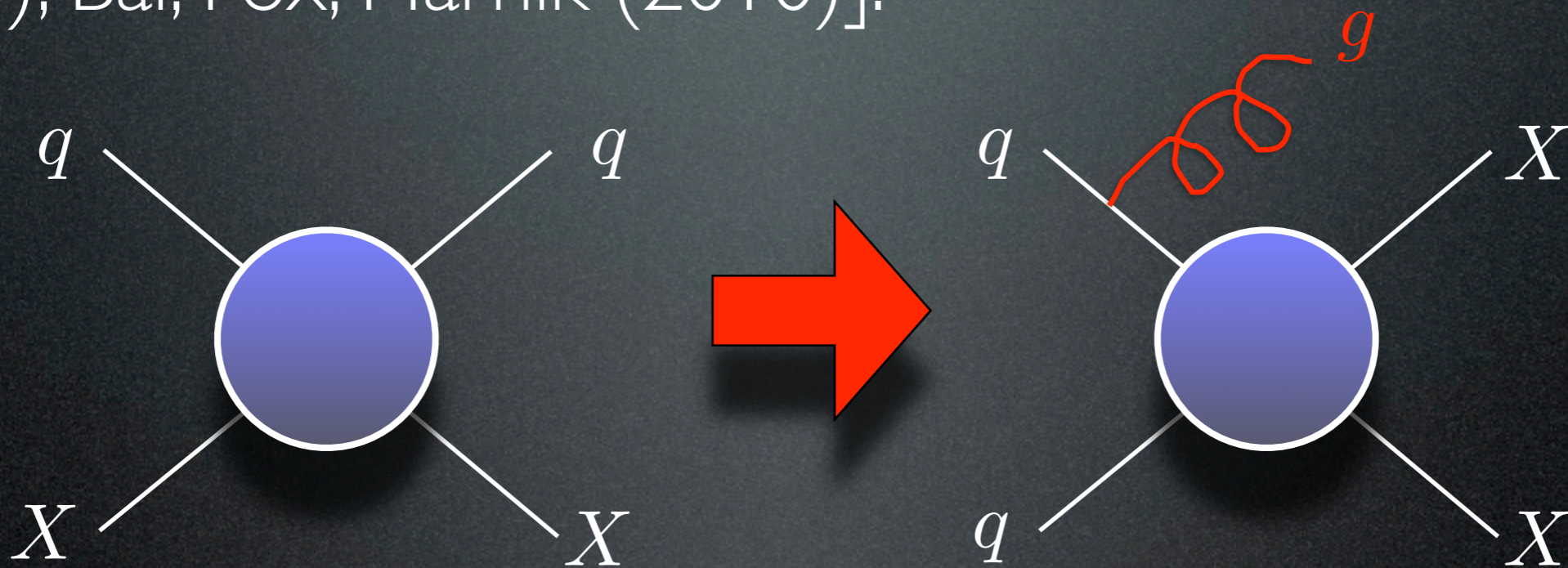
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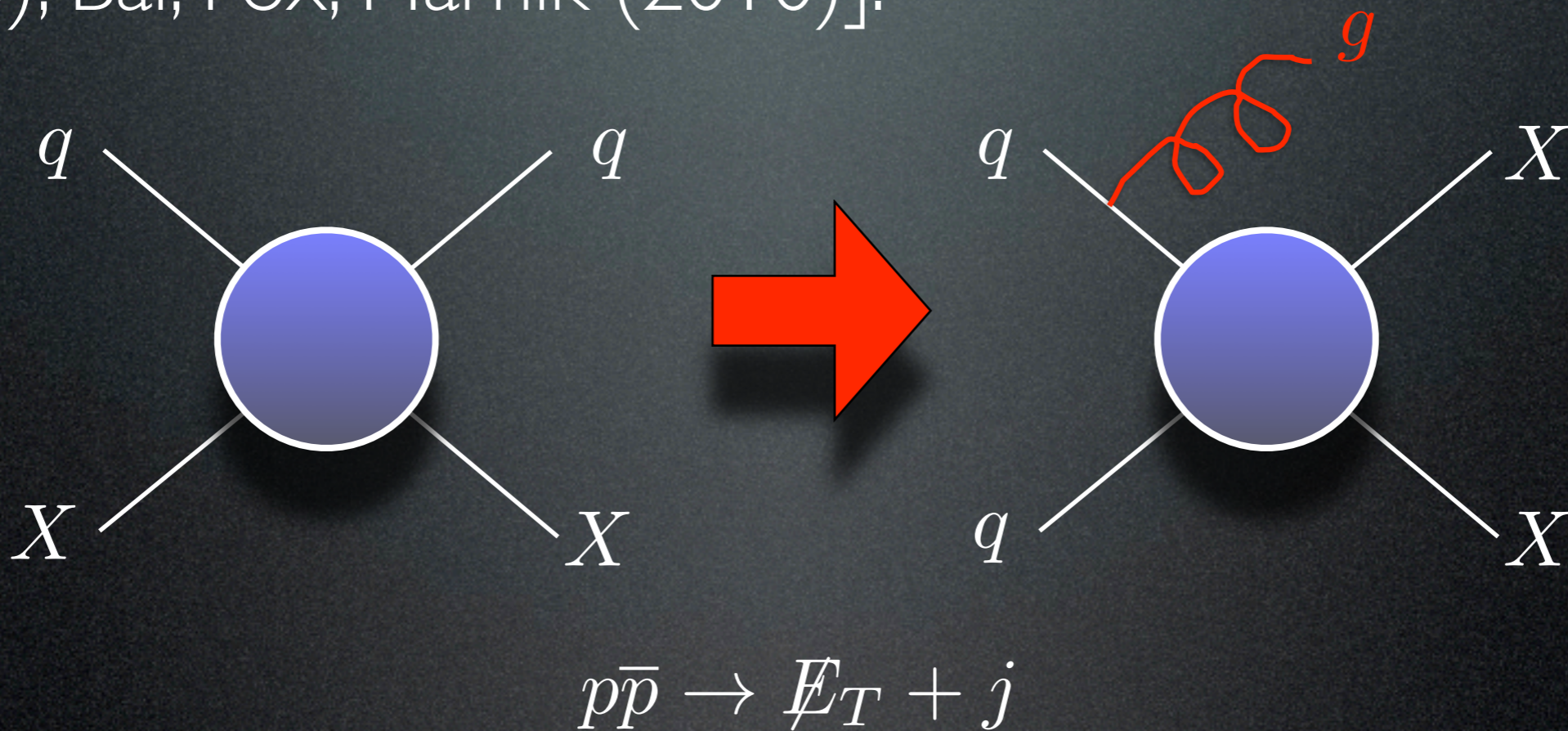
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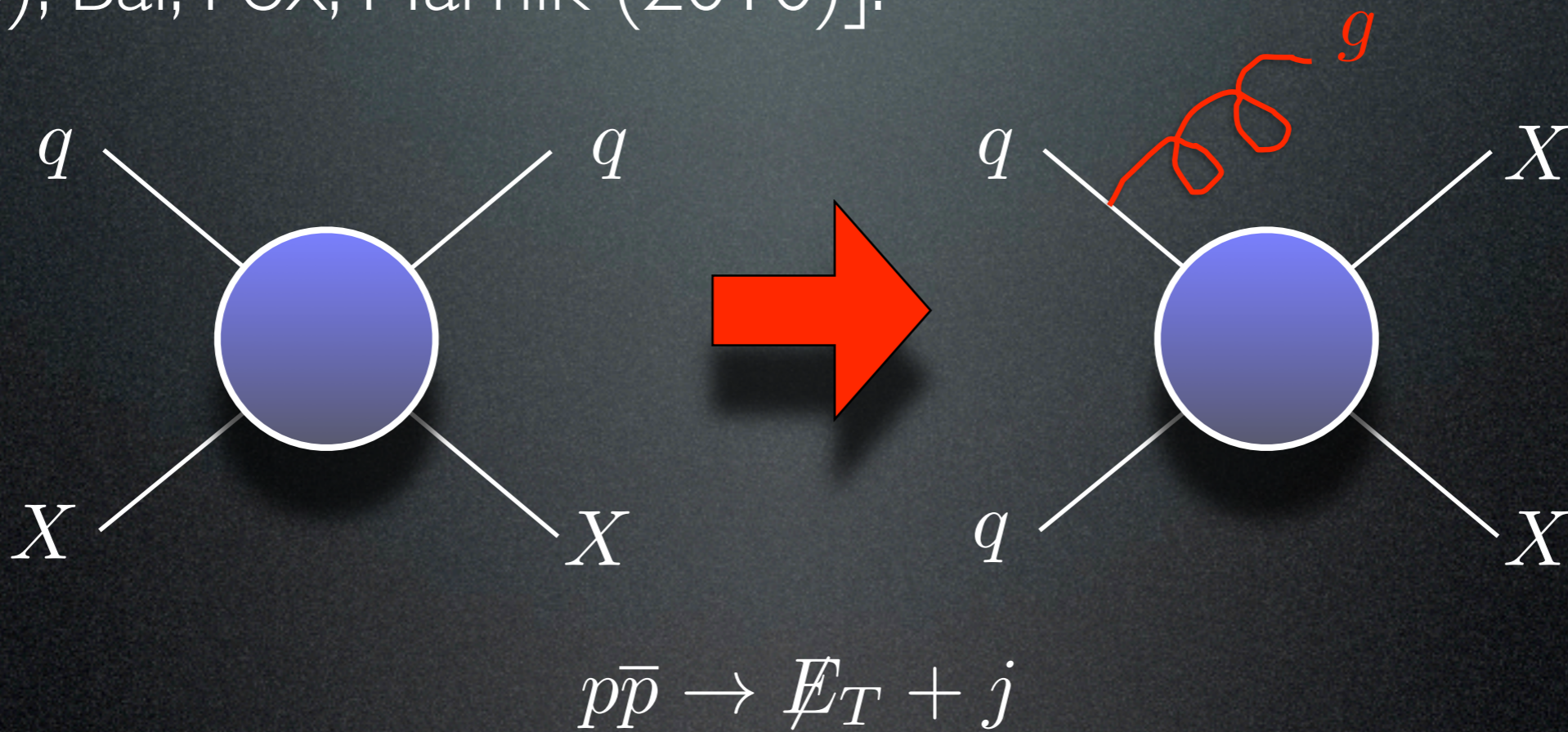
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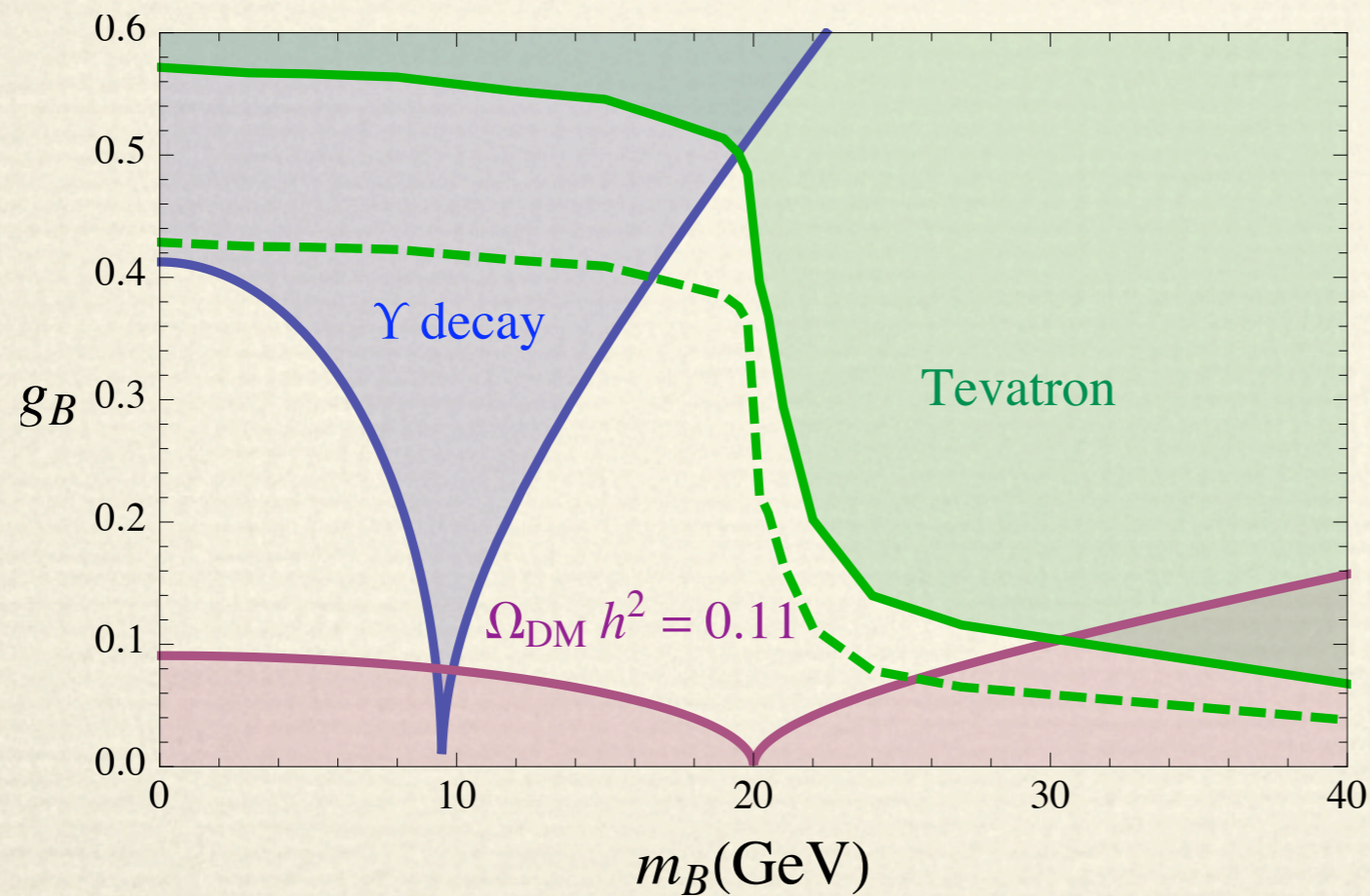
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See Luca's talk.

Combined constraints: axial case

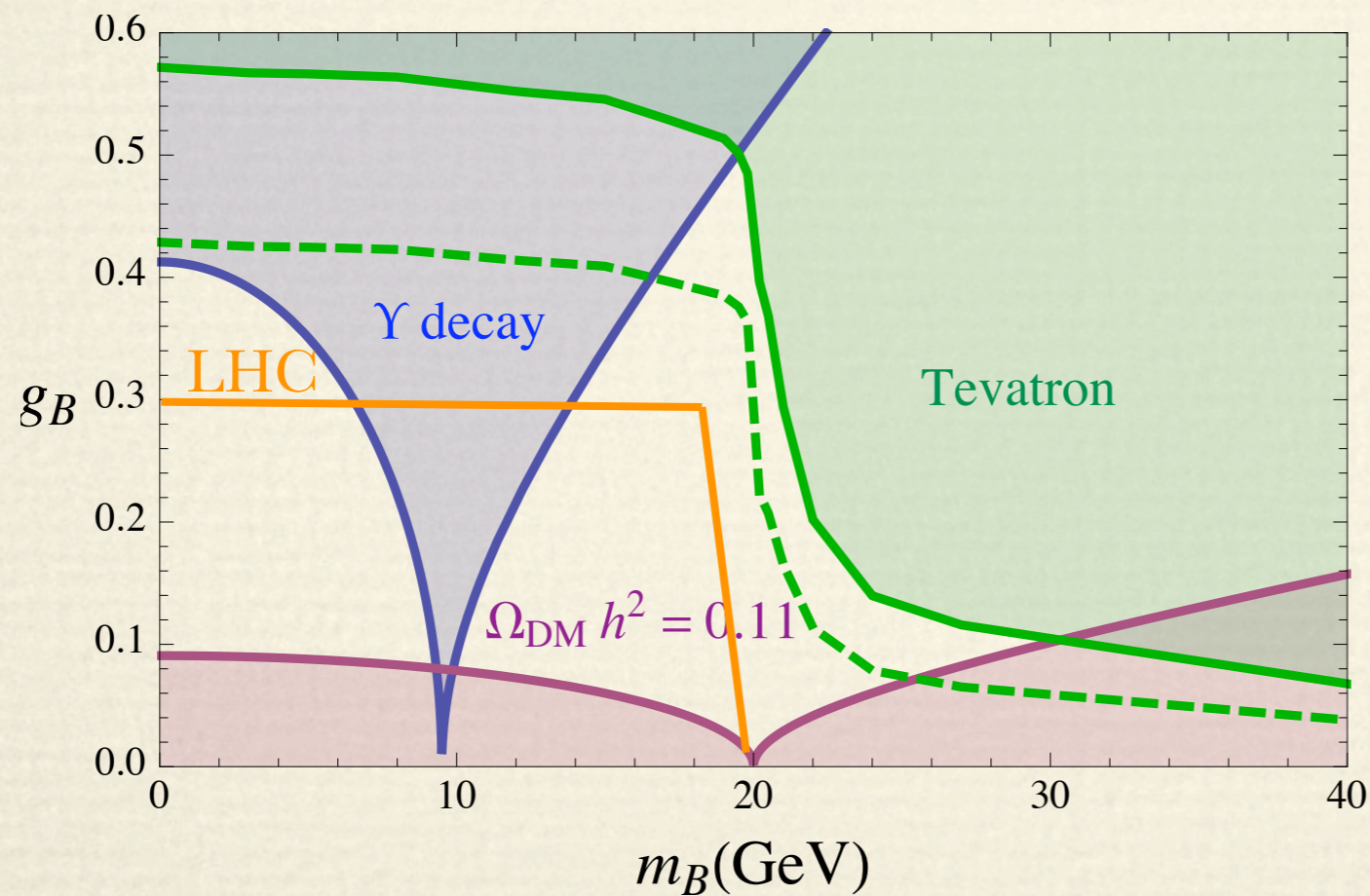
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CONCLUSIONS

- Gauging baryon number saves the proton + automatic DM candidate charged under baryonic force.
- Simultaneous generation of dark and visible asymmetries.
- Consistent with bounds from B-factories, LEP, mono-jet Tevatron searches, and direct detection for:
 - GeV-scale DM with a GeV-scale mediator.
- LHC and direct detection will probe much of the remaining parameter space.

EXTRAS

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Decay operator \leftrightarrow asymmetry transfer operator